



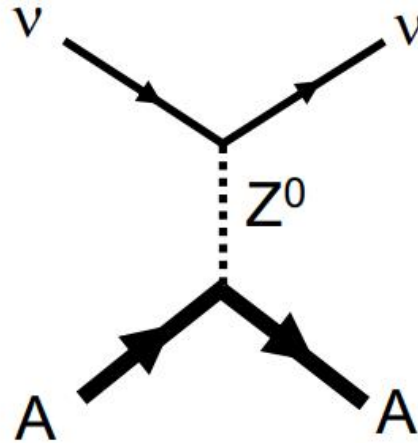
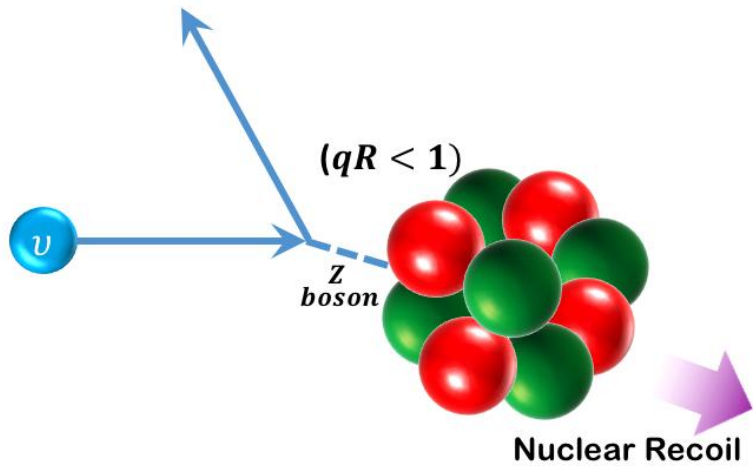
CICENNS: 300-kg CsI(Na) Detector for Coherent Elastic Neutrino-Nucleus Scattering(CE ν NS)

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on behalf of the CICENNS Collaboration

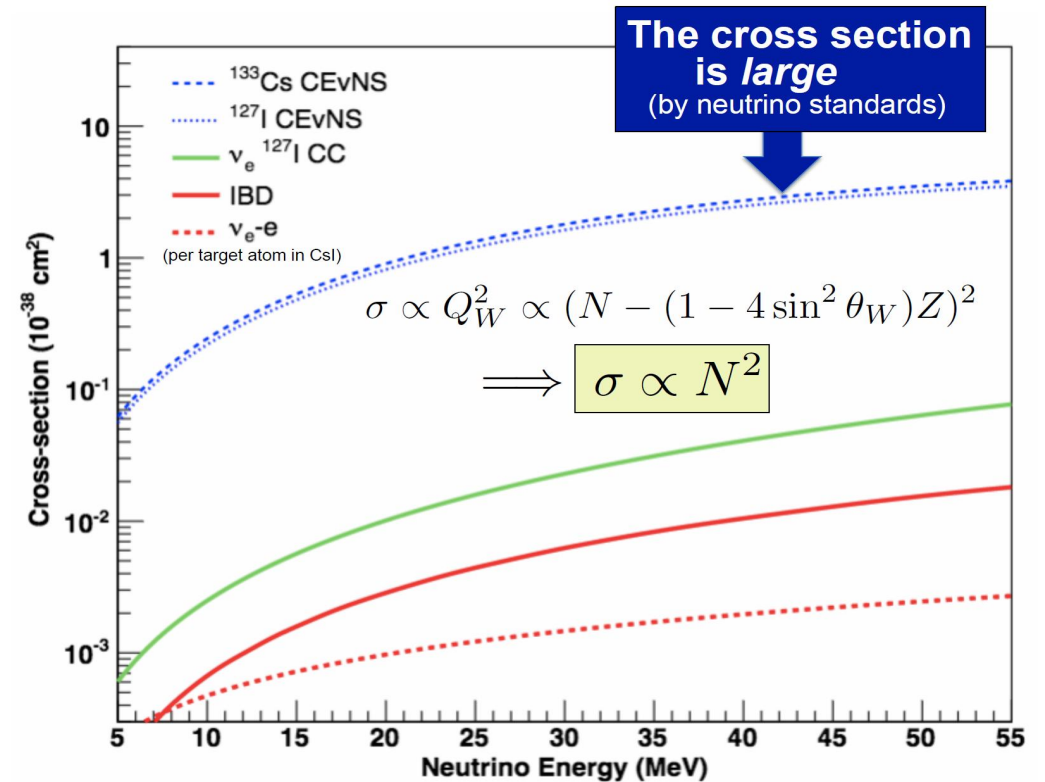
August 15, 2024

Coherent Elastic Neutrino-Nucleus Scattering (CEvNS)



$$\left(\frac{d\sigma}{dT}\right)_{\text{coh}} = \frac{G_F^2 M}{4\pi} Q_W^2 F^2(Q) \left[1 - \frac{MT}{2E_\nu^2}\right]$$

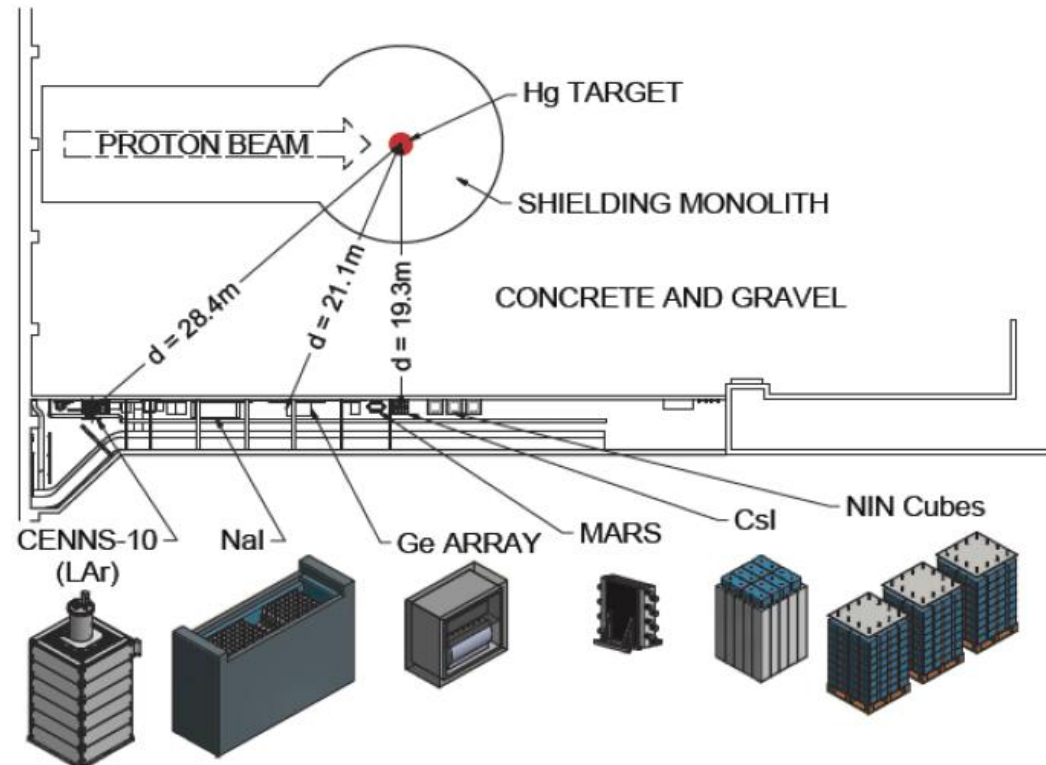
- $qR < 1$
- Z-exchange between neutrino and nucleus
- Nucleus recoils as a whole
- Coherent up to $E_\nu \sim 50$ MeV



First experimental observation by COHERENT



SCIENCE 357, no. 6356,
1123-1126 (2017)



- 14.6 kg CsI(Na) detector
- 308.1 day live-time (beam-on)
- 134 ± 22 events observed
- 6.7σ observation

Goals of CICENNS experiment

Development of 300kg CsI(Na) detector For Coherent Elastic Neutrino Nucleus Scattering (CEvNS)

- Obtain sufficient CEvNS event rate with substantial target mass
- Establish a new field of studying unexplored physics using CEvNS



*Detection of neutrinos from pion/muon decays at rest (at *China Spallation Neutron Source*)*



(1) Precise measurements (world-most accurate)

- CEvNS cross section → **weak couplings** at low momentum transfer ($\pm 3\%$)
- Mean **radius of neutron distribution** inside nucleus ($\pm 2\%$)
- Understanding of dark-matter background and detection of solar neutrinos

(2) New physics searches

- **Non-standard (new) neutrino interactions** (world-best search)
- New particle searches: dark photons or sub-GeV dark matter
- A new region of neutrino magnetic moment
- Efficient search for sterile neutrino oscillation by neutral current

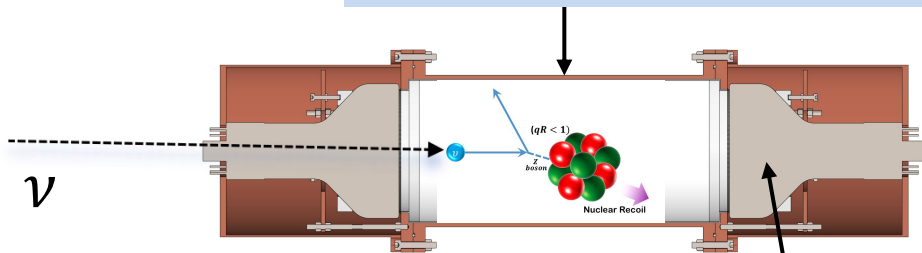
“Research contents”

“Scientific goals”

Schematic view of CsI(Na) Crystal Detector

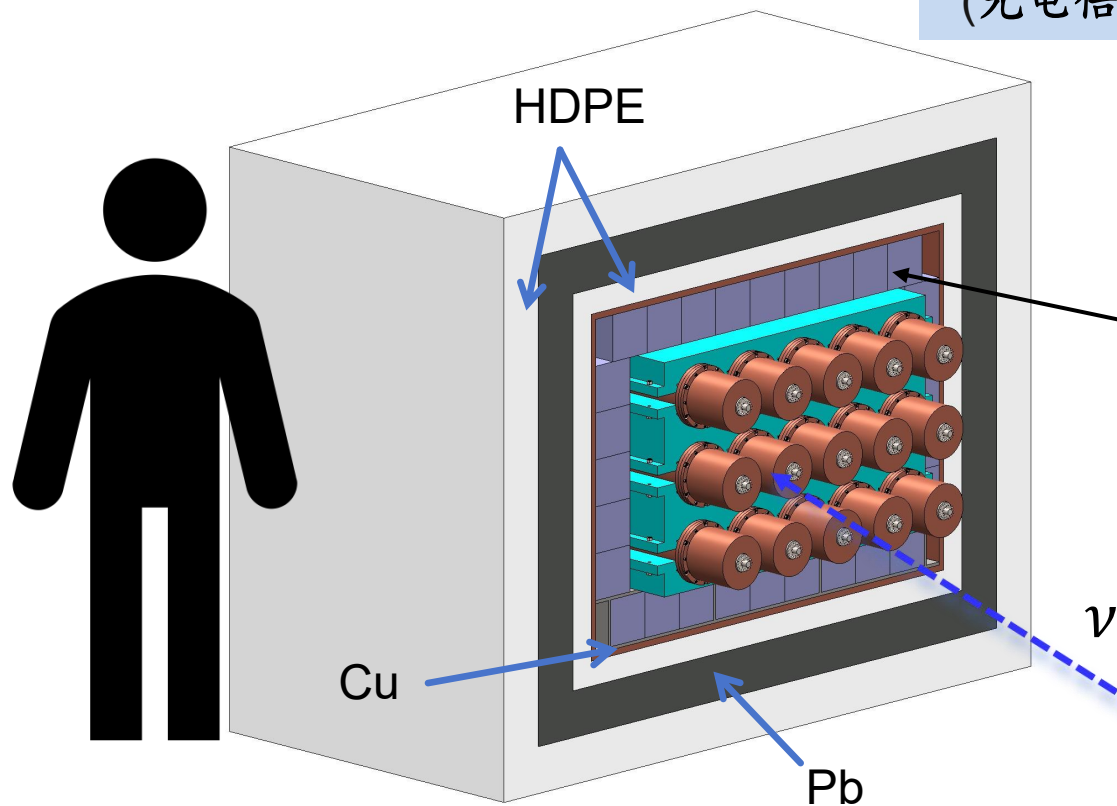
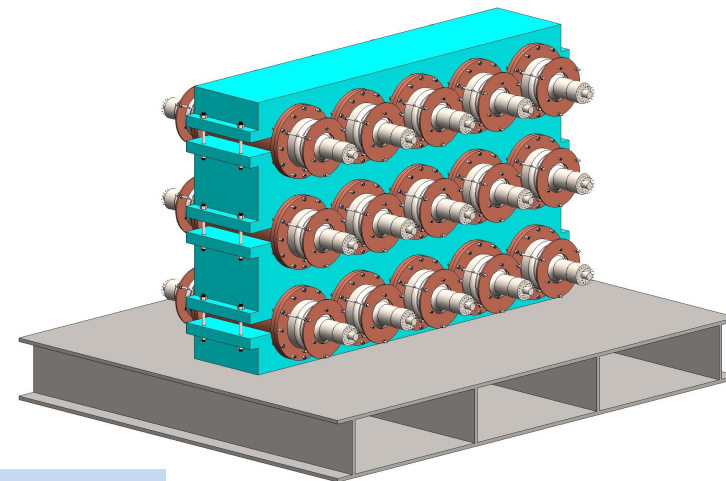
CsI(Na) crystal (闪烁晶体)

[* Building a good detector essential for this research]

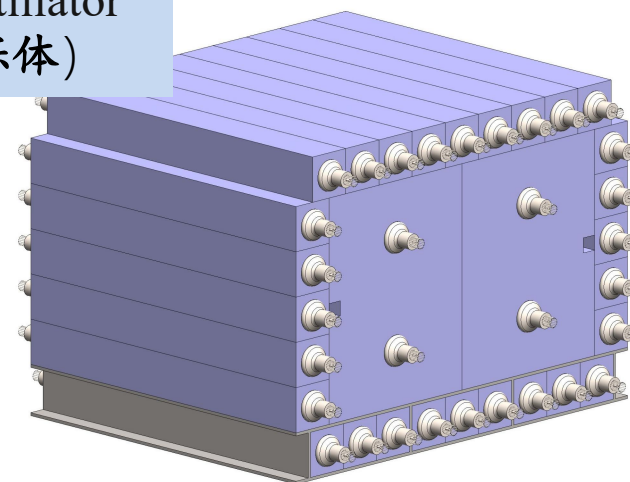


- 20 kg CsI x 15
- 14 cm (ϕ) x 28.7 cm each

5-inch PMT
(光电倍增管)



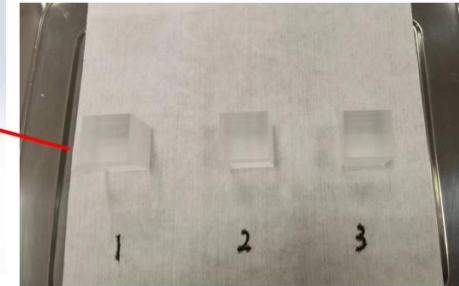
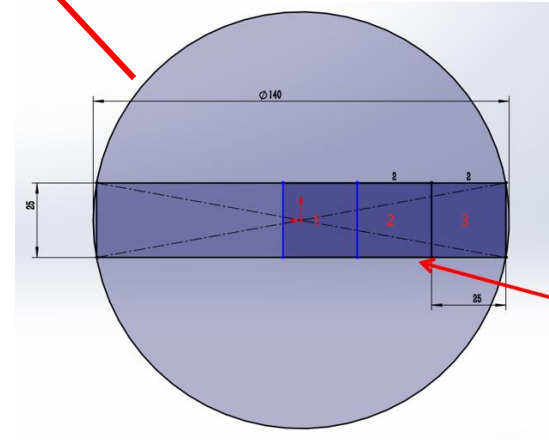
Plastic scintillator
(塑料闪烁体)



CsI(Na) crystal



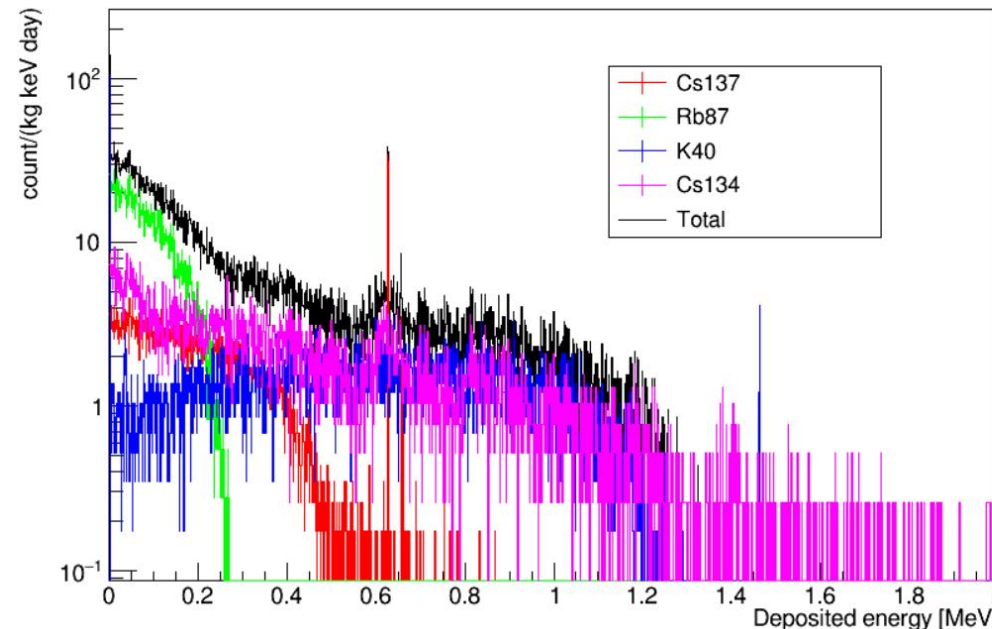
放射性测试样品 4个



- The first full-size CsI(Na) crystal sample of has been produced by SICCAS
- Radioisotopes measurement
- optical measurement

Internal Radioactivity of Crystal

Isotopes	²³⁸ U (ppt)	²³² Th (ppt)	⁸⁷ Rb (ppb)	¹³⁷ Cs (mBq/kg)	¹³⁴ Cs (mBq/kg)
CICENNS	50 ± 10	47 ± 10	~2	??	??
KIMS	0.75 ± 0.23	0.38 ± 0.07	1.3 ± 0.4	6.3 ± 0.7	14.1 ± 1.1
COHERENT	<1000	<1000	~20	28 ± 3	26 ± 2

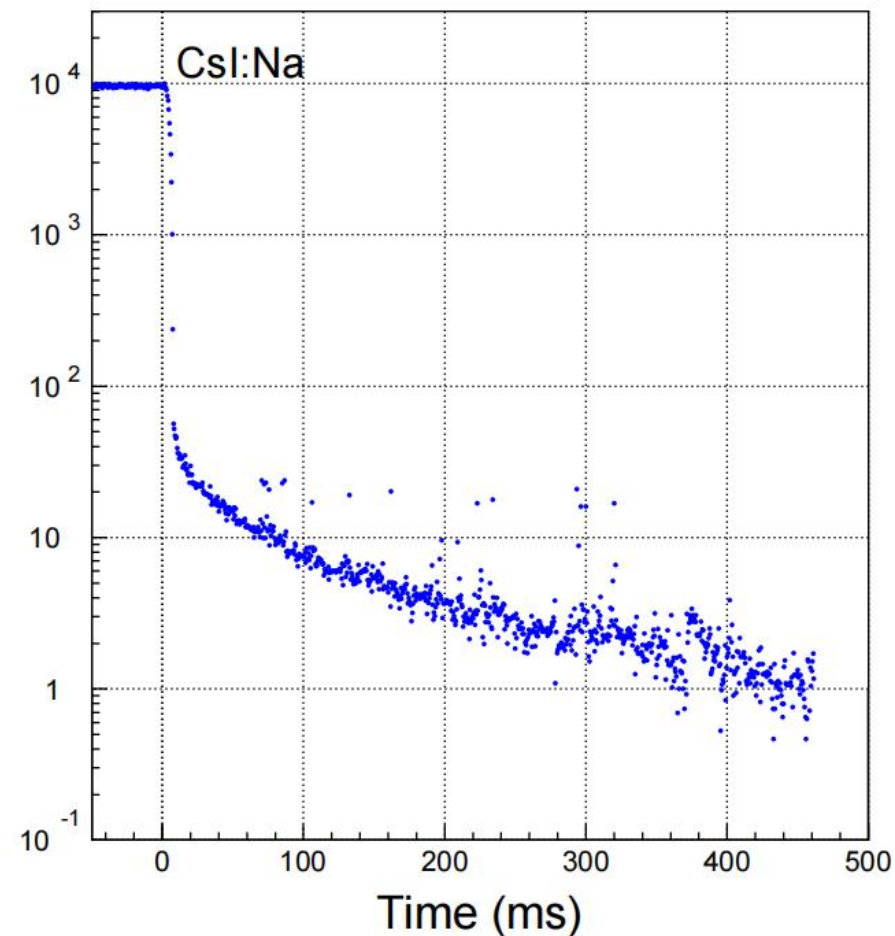
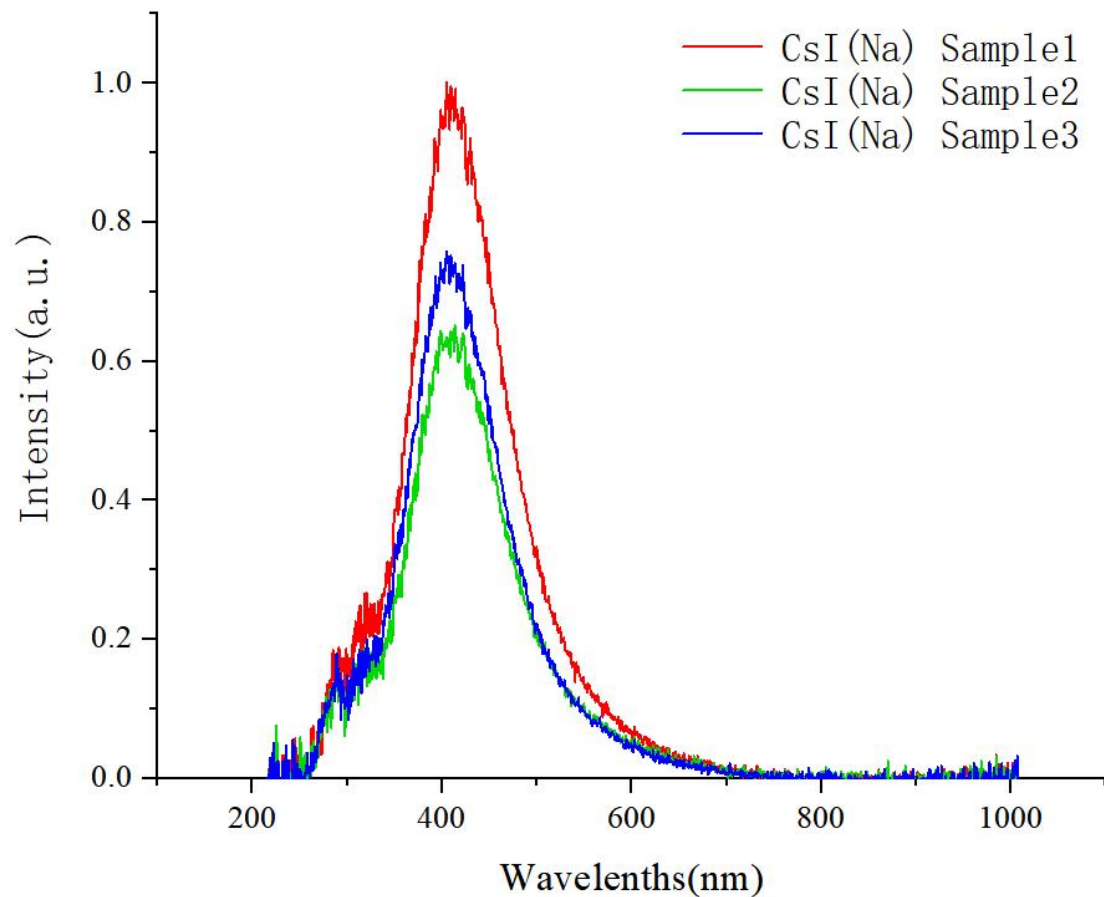


Background spectra obtained using GEANT4 simulation for a 20kg CsI(Na) with:
¹³⁷Cs ~10mBq/Kg, ¹³⁴Cs ~30mBq/Kg, ⁴⁰K ~20mBq/Kg, ⁸⁷Rb~10ppb.

- The top part of the CsI(Na) crystal sample is the part of the entire crystal with a higher radioactive concentration
- Previous measurements of the same type of CsI(Na) sample (6.7cm (φ) × 1.6 cm (L)): ¹³⁷Cs < 5.9 mBq/kg, ¹³⁴Cs ~ 23 mBq/kg



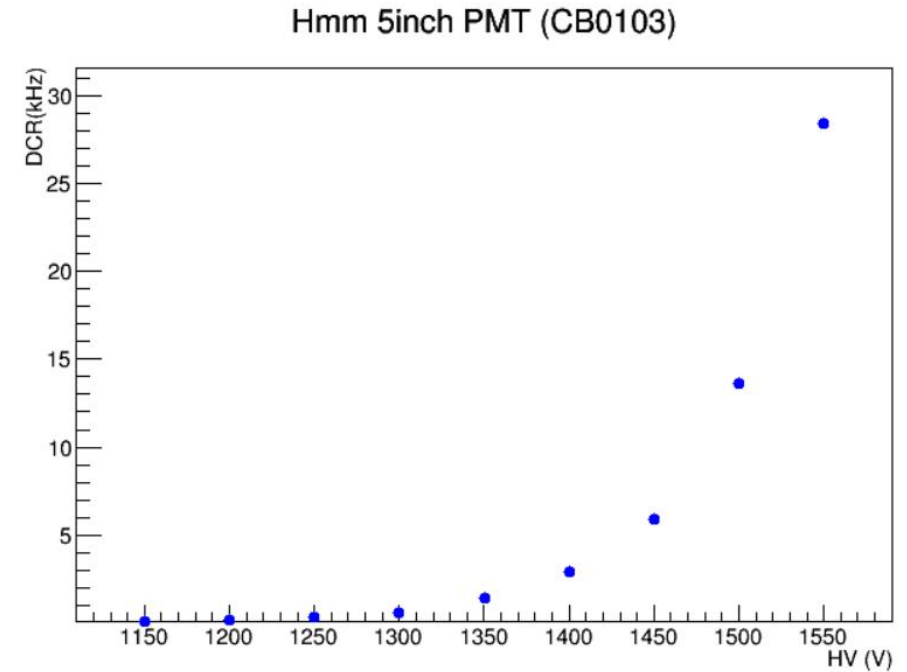
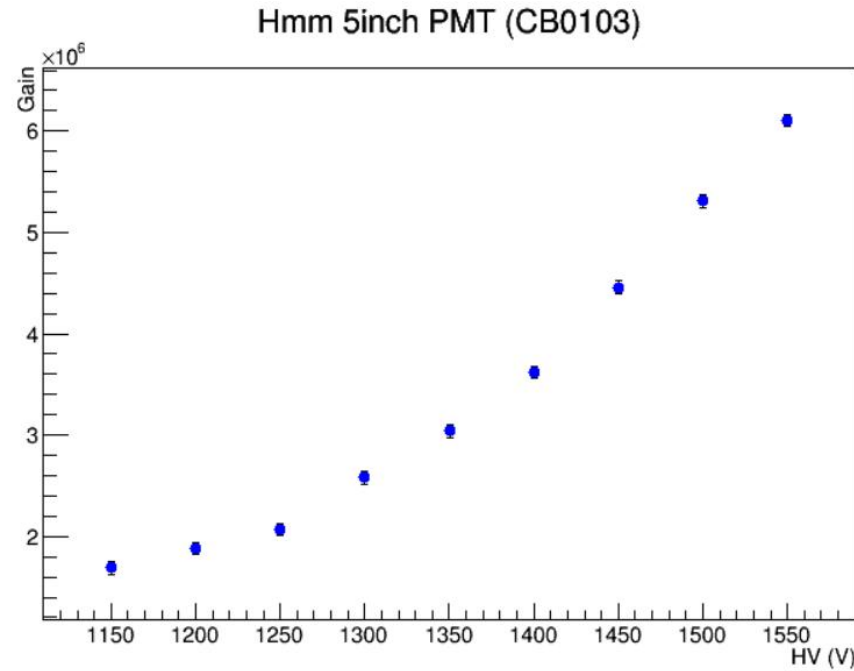
Optical properties of CsI(Na) crystals



- Emission spectrum of CsI(Na) crystal
- The peak emission spectrum of crystal doped sodium is 405nm, which is consistent with PMT detection(290~650nm)

- Crystal doped sodium afterglow:
~0.5% @ 6ms

Performance of 5-inch PMT



R877-100 PMTs (super-bialkali)

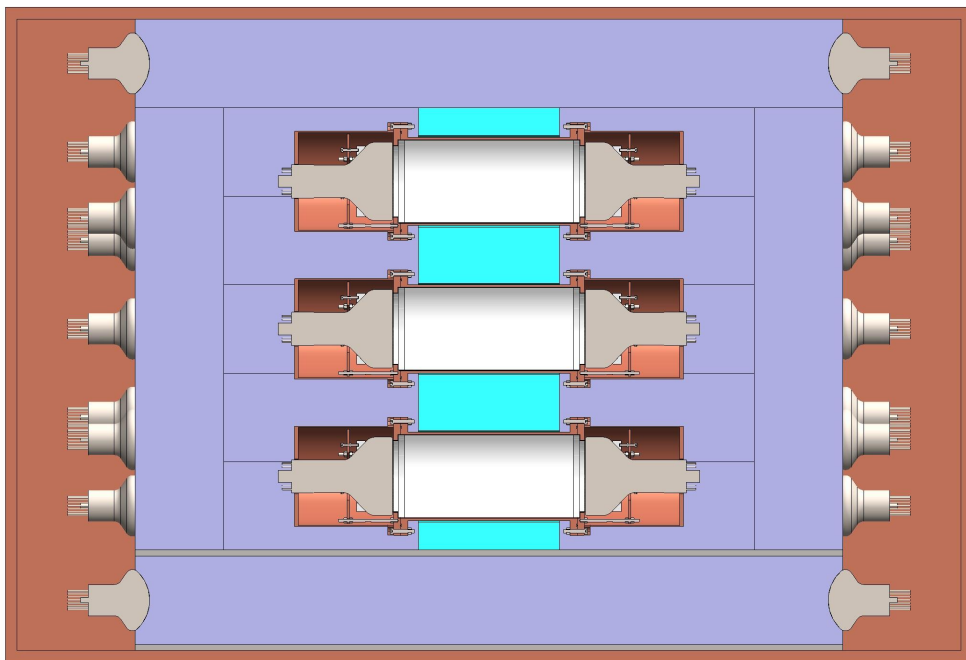
➤ QE: 37% @420nm

Gain and DCR of PMT at different operating voltages.

➤ Gain: 3.6×10^6 @1400V

➤ DCR: 2.9KHz @1400V

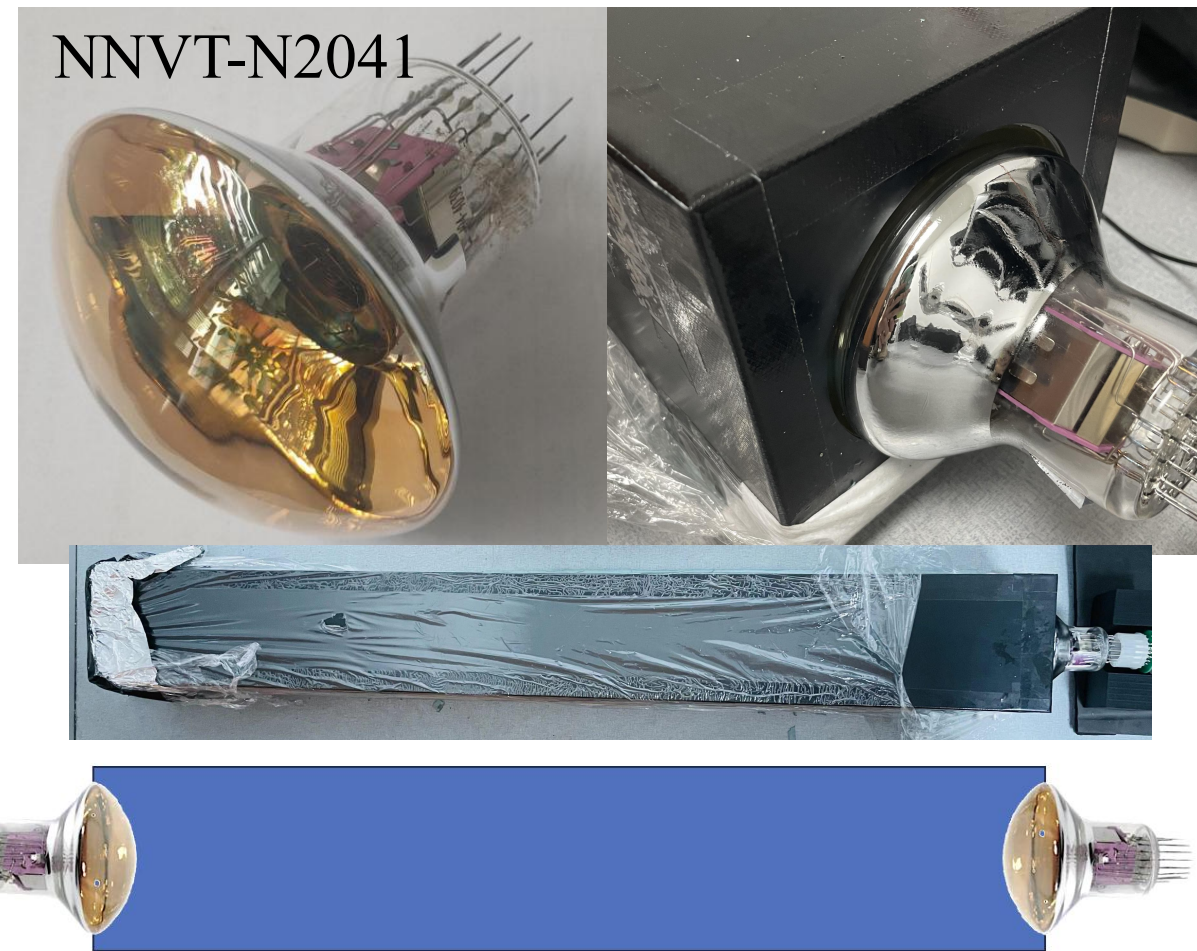
Plastic Scintillator Veto



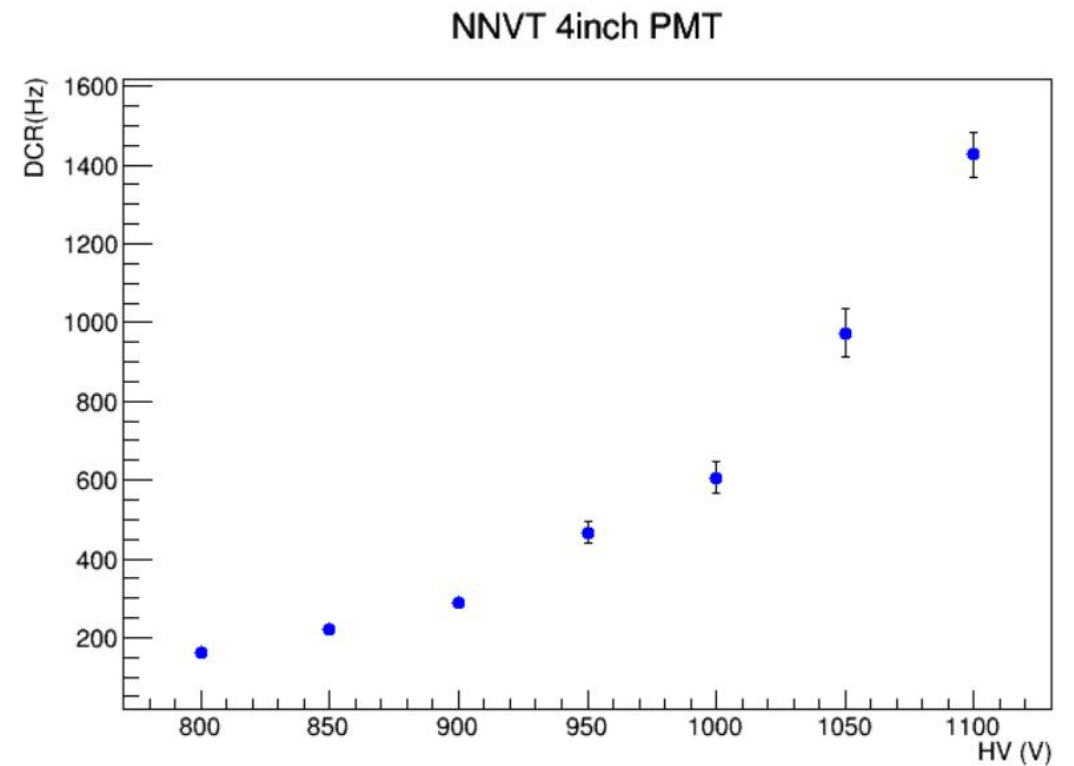
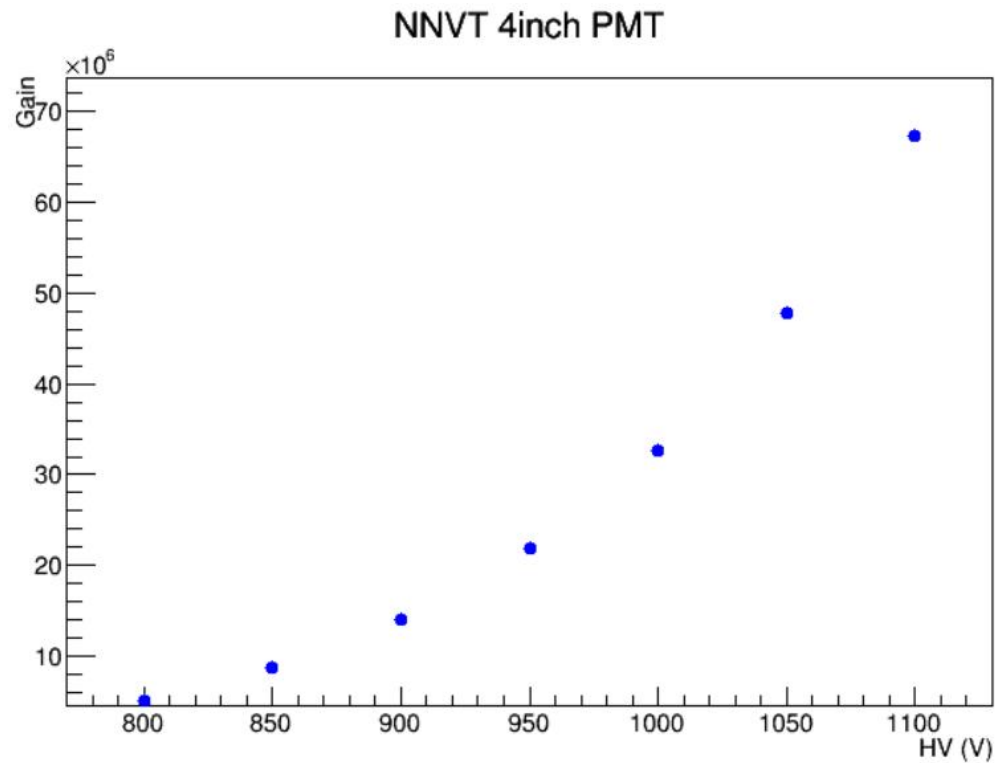
PSV detector

- Plastic Scintillator × 32
- 150 mm × 150mm × 1200 mm (L) each
- Two 4-inch N2041 PMTs

- Purpose: Identify and veto beam fast neutrons, cosmic muons, and gamma rays



- Coupling of camber PMT to plastic scintillator with light collection efficiency of 14.5% (by GEANT4)

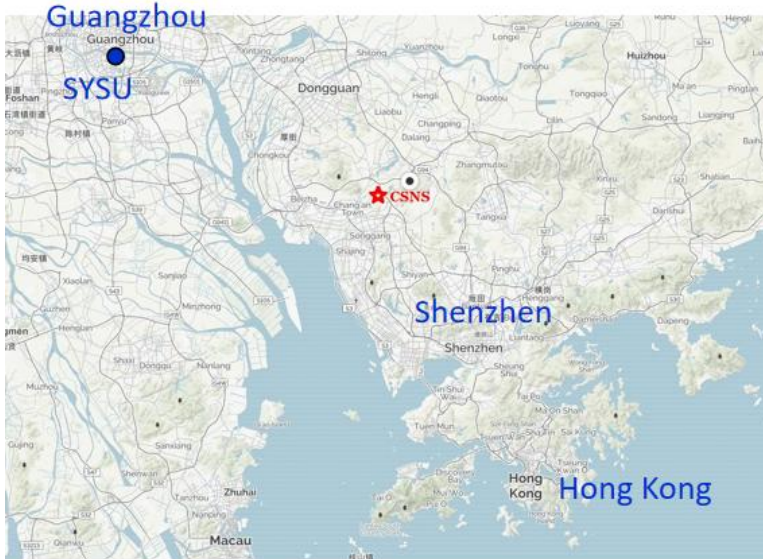


Gain and DCR of PMT at different operating voltages.

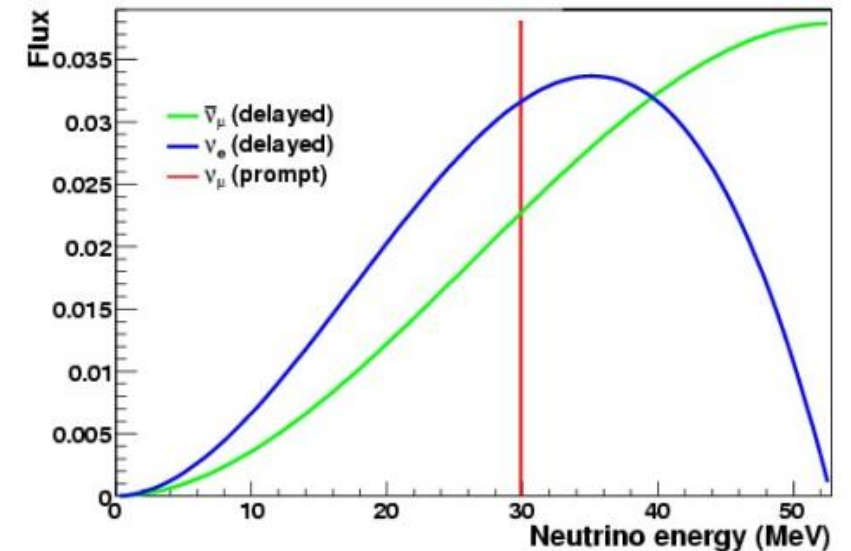
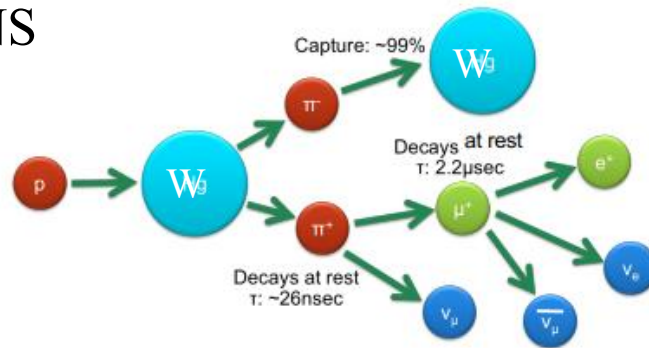
➤ Gain: 1.4×10^7 @900V

➤ DCR: 290Hz @900V

China Spallation Neutron Source (CSNS)



- The CICENNS experiment is initially planned to be conducted in the CSNS

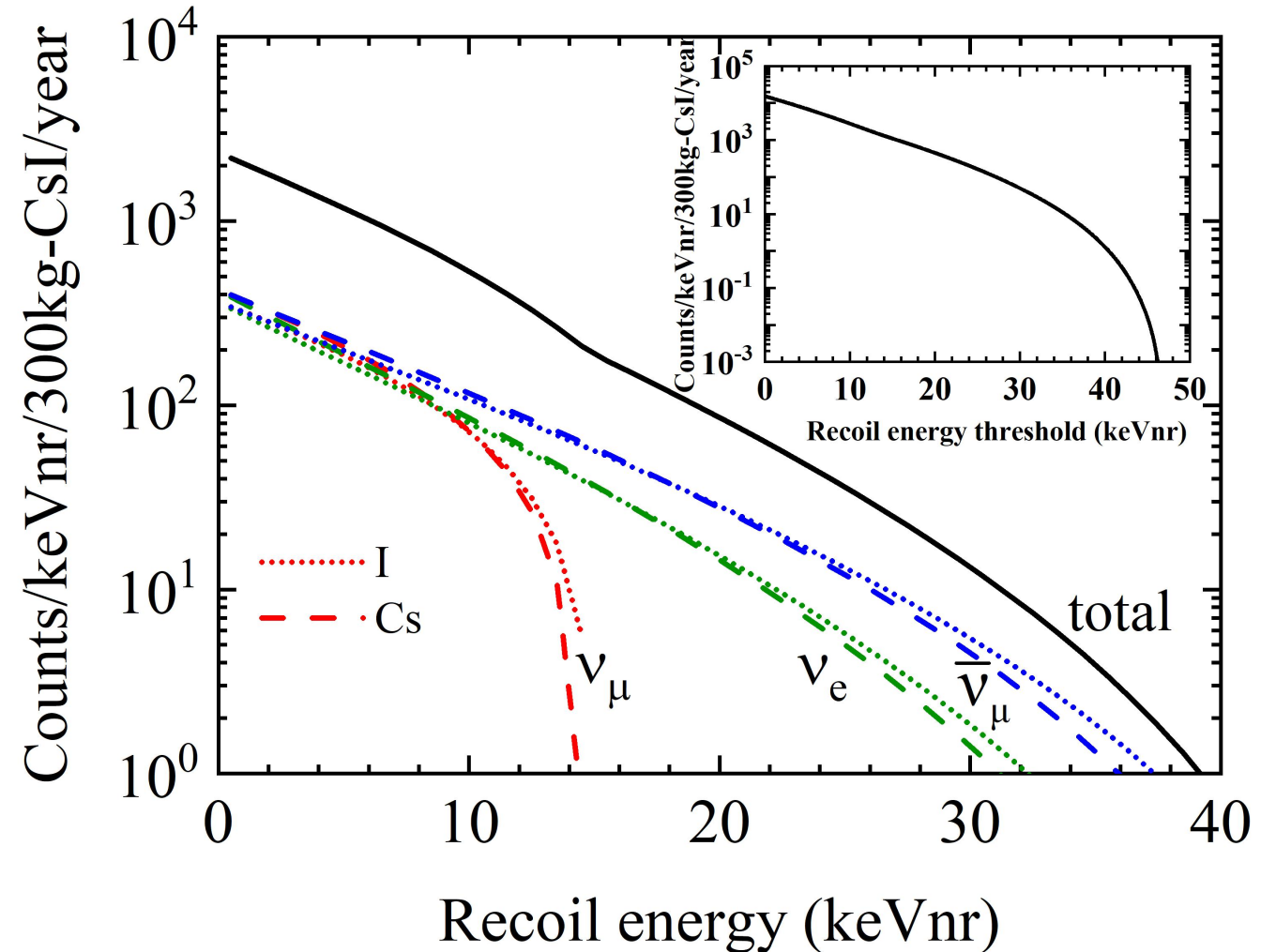


- 1.6 GeV proton beam
- 140 kW (→ 500 kW)
- 25 Hz repetition rate

Neutrino flux at 10.5 m: $6.71 \times 10^6 / \text{cm}^2 / \text{s} / \text{flavor}$ (~40% of COHERENT)

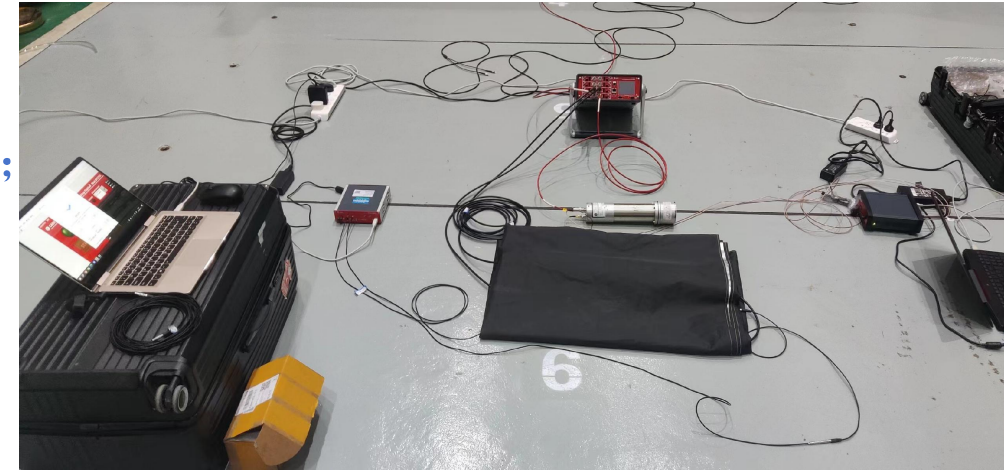
Expected CEvNS Signal of CICENNS at CSNS

- The maximum nuclear recoil energy is roughly 35 keVnr.
- 14.6 kg CsI(Na) crystal for ~ 300 CEvNS events (PRL 129, 081801, 2022)
- The expected CEvNS production rate per 300 kg for a year is about **4950 events** at the energy threshold of 5 keVnr and **6450 events** at 3 keVnr. (~ 20 times of COHERENT)

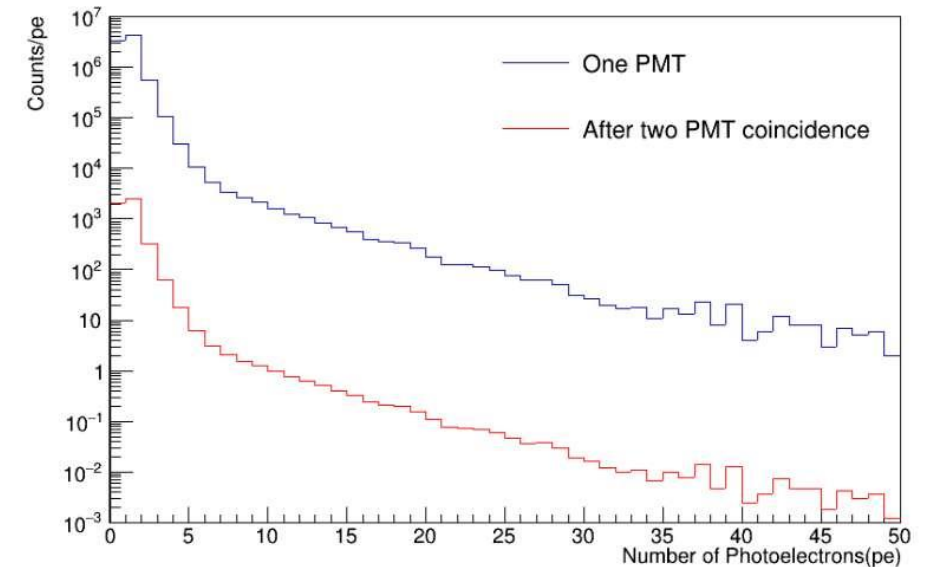


Background

- Beam related fast neutron
 - The most significant background;
 - On-site measurement on-going, joint effort from SYSU and UCAS;
 - Identification and veto with PSV
- PMT dark-current noise
 - By attaching two PMTs at both ends of a crystal, reduce PMT dark-current noise events by 3 orders of magnitude.
- Radioactivity in detector materials
 - Low internal radioactivity in CsI(Na) crystal by purification;⁸⁷Rb, ¹³⁷Cs, ...
 - Background from PMTs and surrounding materials.
- Environmental radioactivity
 - Most of gamma rays from the environmental radioactivity are expected to be attenuated by lead and other shielding.
- Cosmic muon induced neutrons and gammas
 - Full Geant4-based MC on-going



Dark pulse spectrum



Sensitivity: flavored CEvNS cross section and weak couplings

The CEvNS rate is a clean SM prediction

$$\frac{d\sigma}{dT} = \frac{G_F^2 M Q_W^2}{2\pi \cdot 4} F^2(Q) \left(2 - \frac{MT}{E_\nu^2} \right)$$

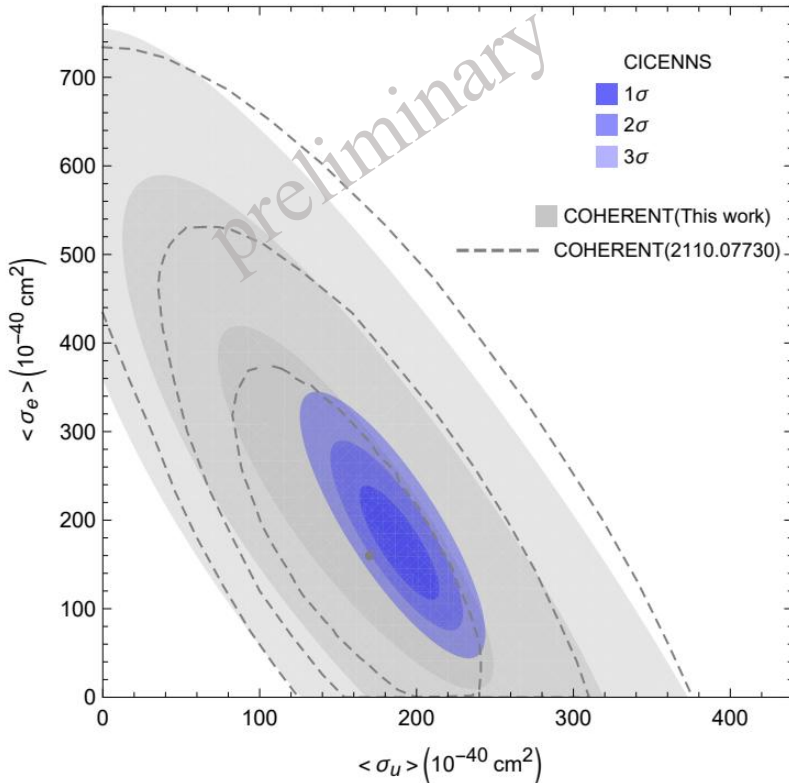
small nuclear uncertainties (±4.8%)

At momentum transfer $Q=50 \text{ MeV}/c$

- COHERENT: $\sin^2 \theta_W = 0.220_{-0.026}^{+0.028}$ (±10%)

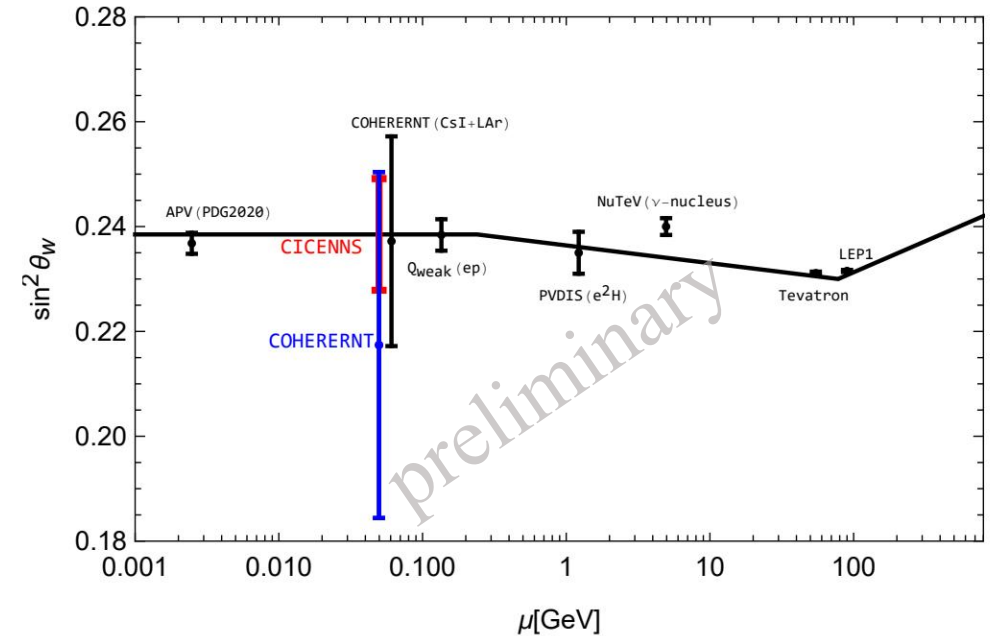
→ CICENNS: ~1%

* SM prediction: 0.23857(5)



Uncertainty of CEvNS cross section:

- COHERENT: +18/-15%
- CICENNS: 6.5%



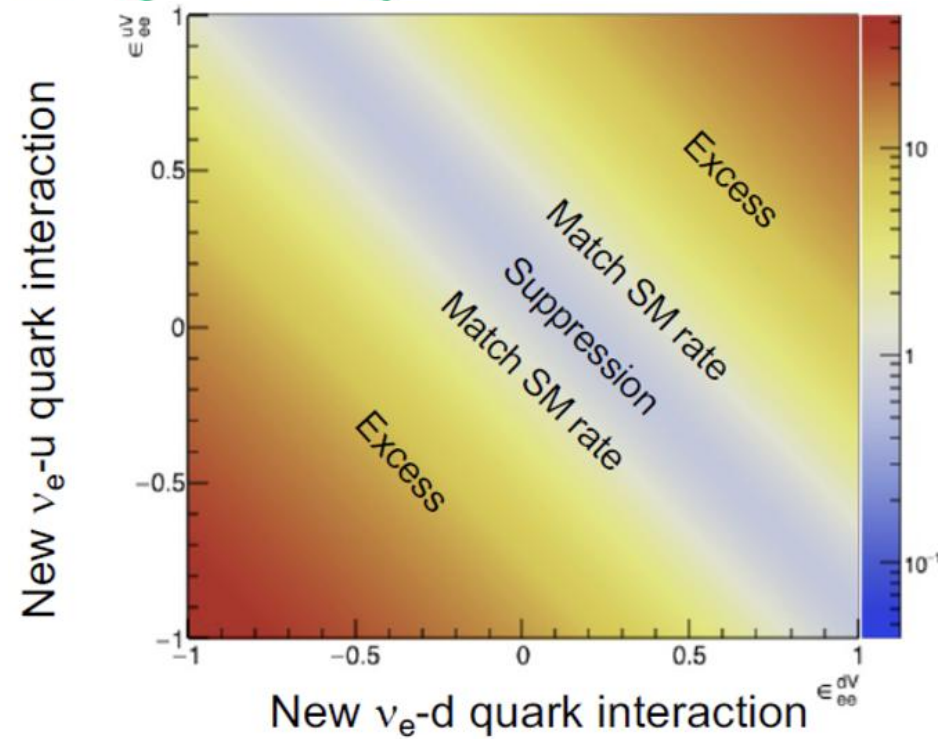
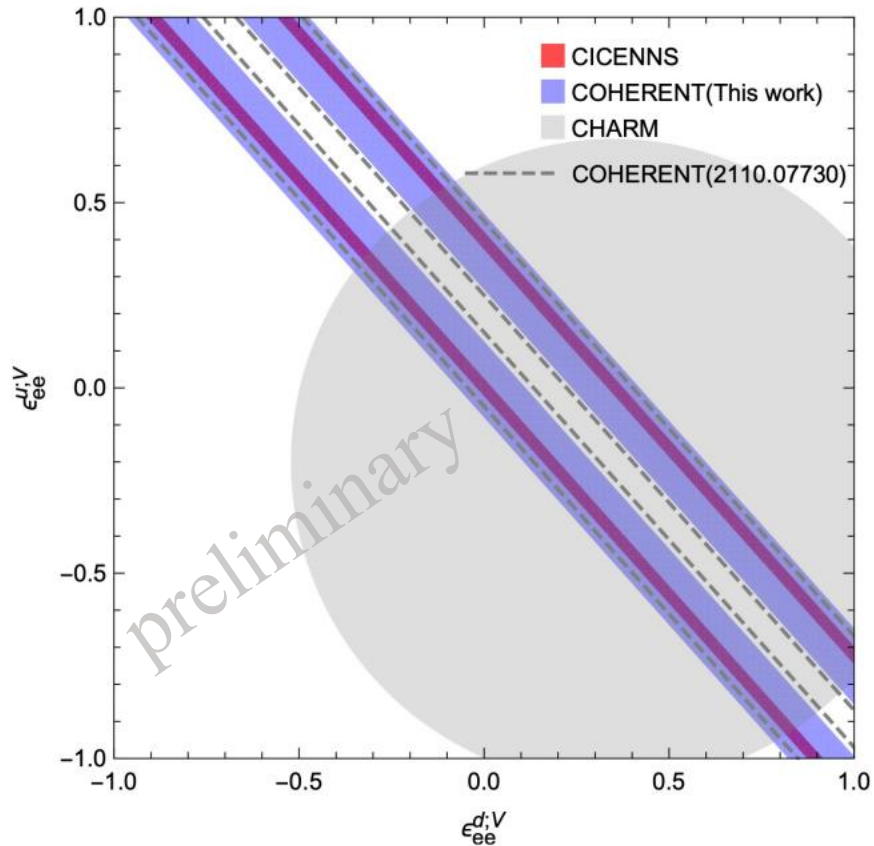
Sensitivity: Non-standard neutrino interaction

Look for a CEvNS **excess** or **deficit** wrt SM expectation

$$\mathcal{L}_{\nu H}^{NSI} = -\frac{G_F}{\sqrt{2}} \sum_{\substack{q=u,d \\ \alpha,\beta=e,\mu,\tau}} [\bar{\nu}_\alpha \gamma^\mu (1 - \gamma^5) \nu_\beta] \times (\epsilon_{\alpha\beta}^{qL} [\bar{q} \gamma_\mu (1 - \gamma^5) q] + \epsilon_{\alpha\beta}^{qR} [\bar{q} \gamma_\mu (1 + \gamma^5) q])$$

Csl

Ratio
wrt SM



- A 300-kg CsI(Na) crystal detector for CEvNS measurement at CSNS has been proposed, and is being built now.
- Through the above efforts, the threshold value of the recoil energy of the detector will be reduced($2 \text{ keV}_{\text{nr}}$).
- We wish to finalize the detector design and a TDR, and complete the detector construction in 2025.

Thanks!