



Neutrino Mass Measurement with Cosmic Gravitational Focusing

Shao-Feng Ge

葛韶锋

gesf@sjtu.edu.cn



Pedro Pasquini



Tan Liang

SFG, Pedro Pasquini, Liang Tan, **JCAP 05 (2024) 108** [arXiv:2312.16972]



上海交通大学

SHANGHAI JIAO TONG UNIVERSITY

青岛高能大会
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李政道研究所
Tsung-Dao Lee Institute

1) Overview of ν Mass

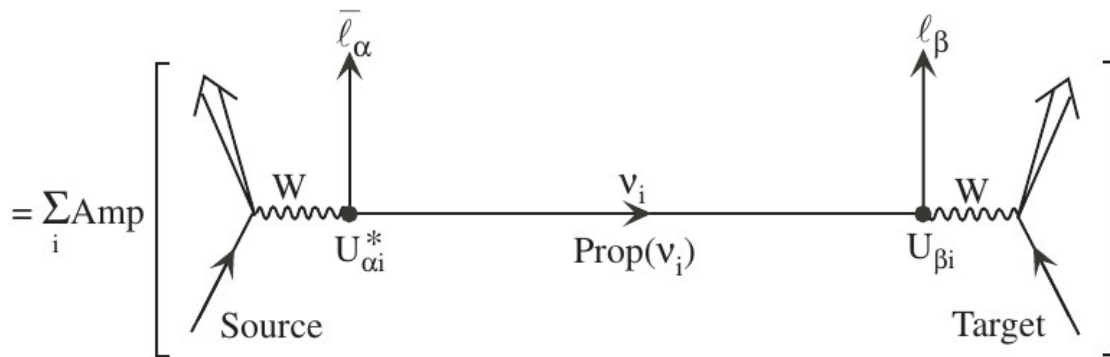
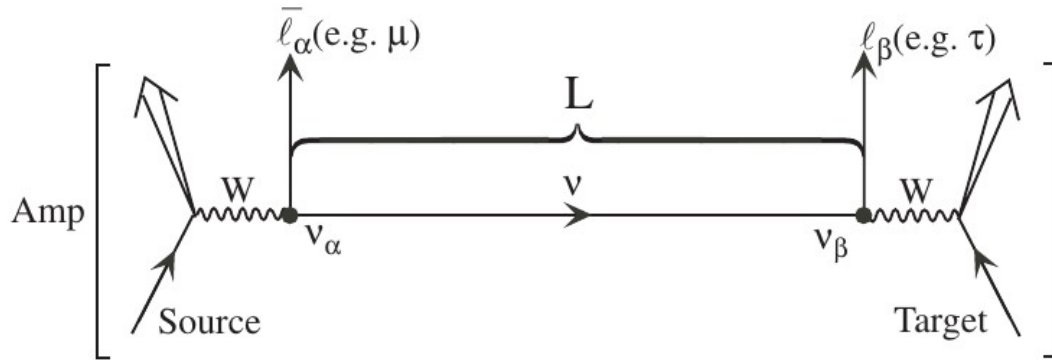
2) 3rd Cosmological Way

Cosmic Gravitational Focusing

Dipole Structure in Galaxy Correlation Function

3) Summary

Neutrino Oscillation & Mass



$$\nu_\alpha = \sum_i U_{\alpha i} \nu_i$$

$$\rightarrow \sum_i U_{\alpha i} e^{i(E_i t - \vec{P}_i \cdot \vec{x})} \nu_i$$

$$= \sum_i U_{\alpha i} P_i U_{\beta i}^\dagger \nu_\beta$$

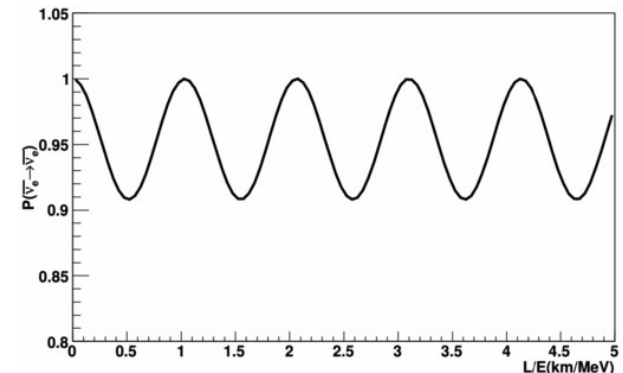
$$\equiv \sum_\beta A_{\alpha\beta} \nu_\beta$$

B. Kayser, [hep-ph/0506165]

$$P_{\alpha\beta} |_{\alpha \neq \beta} \equiv |A_{\alpha\beta}|^2 = \sin^2 2\theta \sin^2 \left(\delta m^2 \frac{L}{4E} \right)$$

1st New Physics

Mass



Global Fit of Oscillation Parameters

(for NO)	-1σ	Best Value	$+1\sigma$
$\Delta m_s^2 \equiv \Delta m_{21}^2$ (10^{-5}eV^2)	7.30	7.50	7.72
$ \Delta m_a^2 \equiv \Delta m_{31}^2 $ (10^{-3}eV^2)	2.52	2.56	2.59
$\sin^2 \theta_s$ ($\theta_s \equiv \theta_{12}$)	0.302 (33.3°)	0.318 (34.3°)	0.334 (35.3°)
$\sin^2 \theta_a$ ($\theta_a \equiv \theta_{23}$)	0.544 (47.54°)	0.566 (48.79°)	0.582 (49.72°)
$\sin^2 \theta_r$ ($\theta_r \equiv \theta_{13}$)	0.02147 (8.43°)	0.02225 (8.58°)	0.02280 (8.69°)
δ_D	191°	216°	257°
δ_{Mi}	??	??	??

Salas, Forero, Gariazzo, Martinez-Mirave, Mena, Ternes, Tortola & Valle, [arXiv:2006.11237]

- Heavy neutrinos (N)

$$\bar{\nu} M_D \mathcal{N} + h.c. + \bar{\mathcal{N}} M_N \mathcal{N} = \begin{pmatrix} \bar{\nu} & \bar{\mathcal{N}} \end{pmatrix} \begin{pmatrix} 0 & M_D \\ M_D^T & M_N \end{pmatrix} \begin{pmatrix} \nu \\ \mathcal{N} \end{pmatrix}$$

required by **Grand Unification Theory (大统一理论)**

- Seeaw Mechanism

The diagonalization of the full mass matrix

$$\begin{pmatrix} 0 & M_D \\ M_D^T & M_N \end{pmatrix} \Rightarrow M_\nu = -M_D \frac{1}{M_N} M_D^T$$

$$M_D \sim O(100) \text{ GeV}, \quad M_N \sim O(10^{15}) \text{ GeV} \quad \rightarrow \quad M_\nu \sim O(0.01) \text{ eV}$$

Light neutrino mass M_ν is suppressed by the heavy ones



Tsutomu Yanagida、
Xiao-Gang He :
Let's play seesaw

ν Oscillation

Beta Decay

Radiative Emission of ν Pairs

SFG, Pedro Pasquini, **JHEP 12 (2023) 083** [arXiv:2306.12953]
Phys.Lett.B 841 (2023) 137911 [arXiv:2206.11717]
Eur.Phys.J.C 82 (2022) 3, 208 [arXiv:2110.03510]

Supernova ν Time Delay

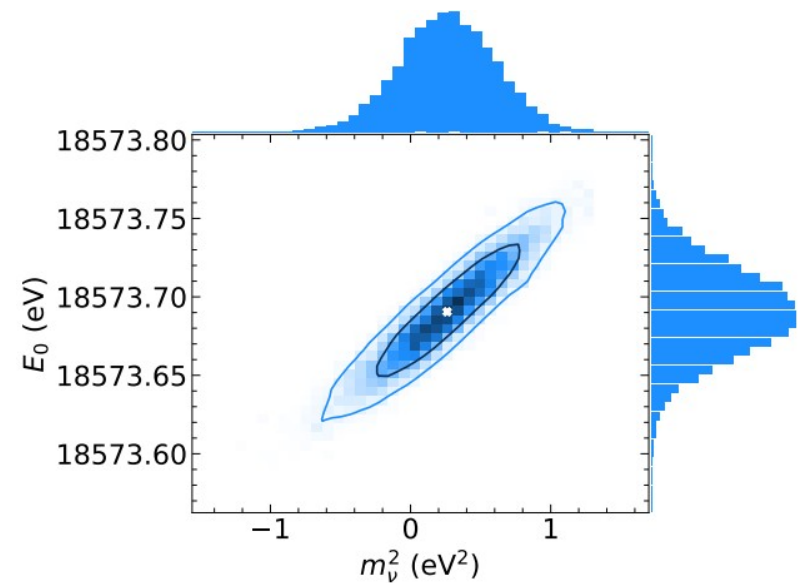
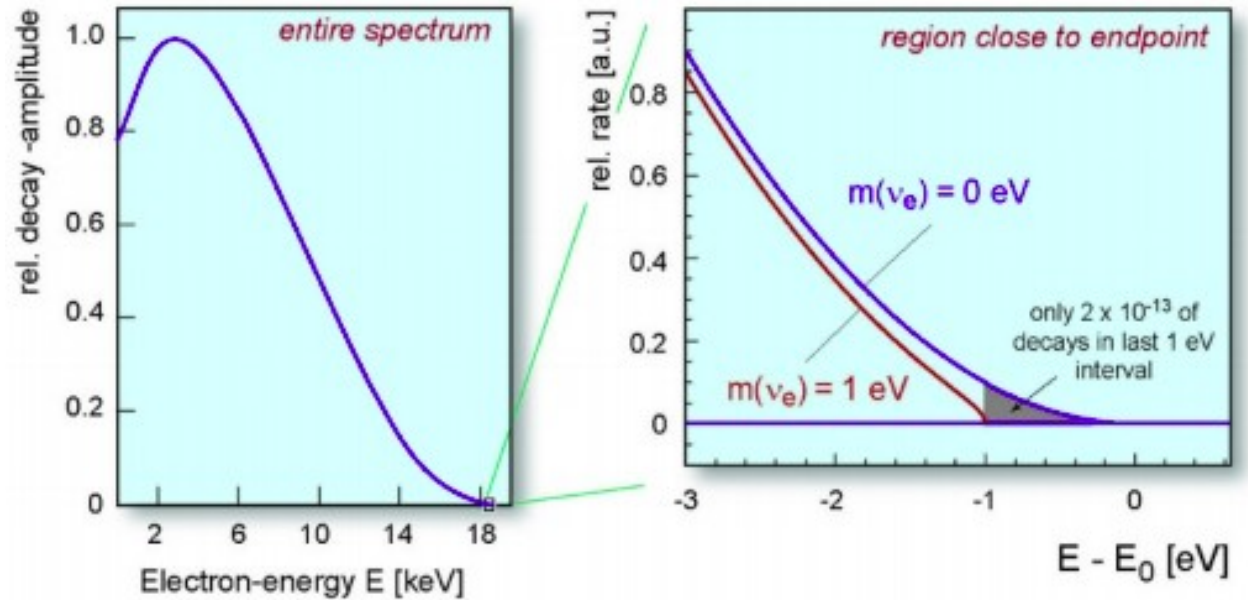
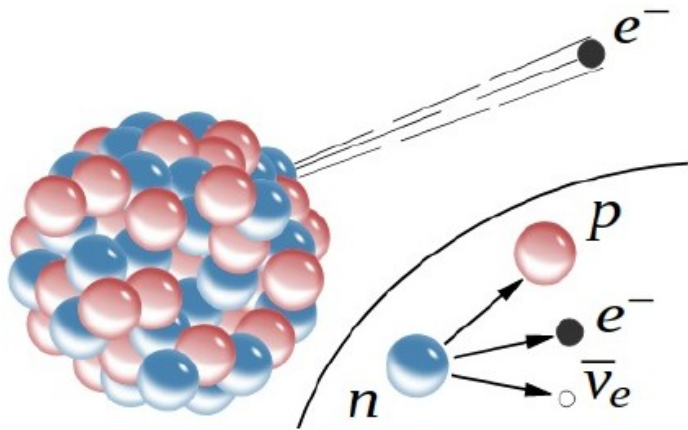
SFG, Chui-Fan Kong, Alexei Smirnov, **accepted by PRL** [arXiv:2404.17352]

C ν B Detection

CMB & LSS

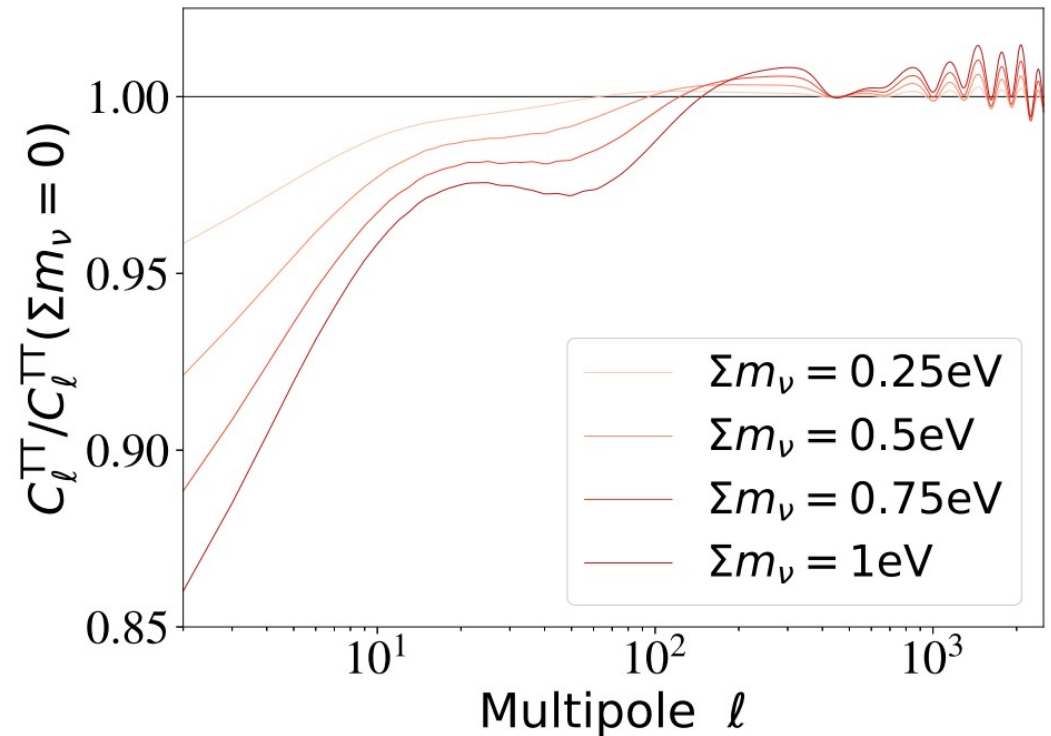
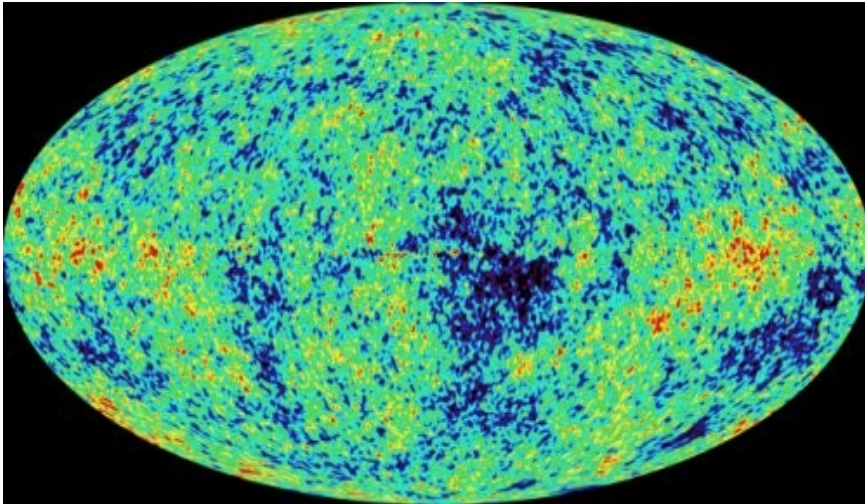
SFG, Pedro Pasquini, Liang Tan, **JCAP 05 (2024) 108** [arXiv:2312.16972]

Neutrino Mass @ β Decay



Nature Phys. 18 (2022) 2, 160-166

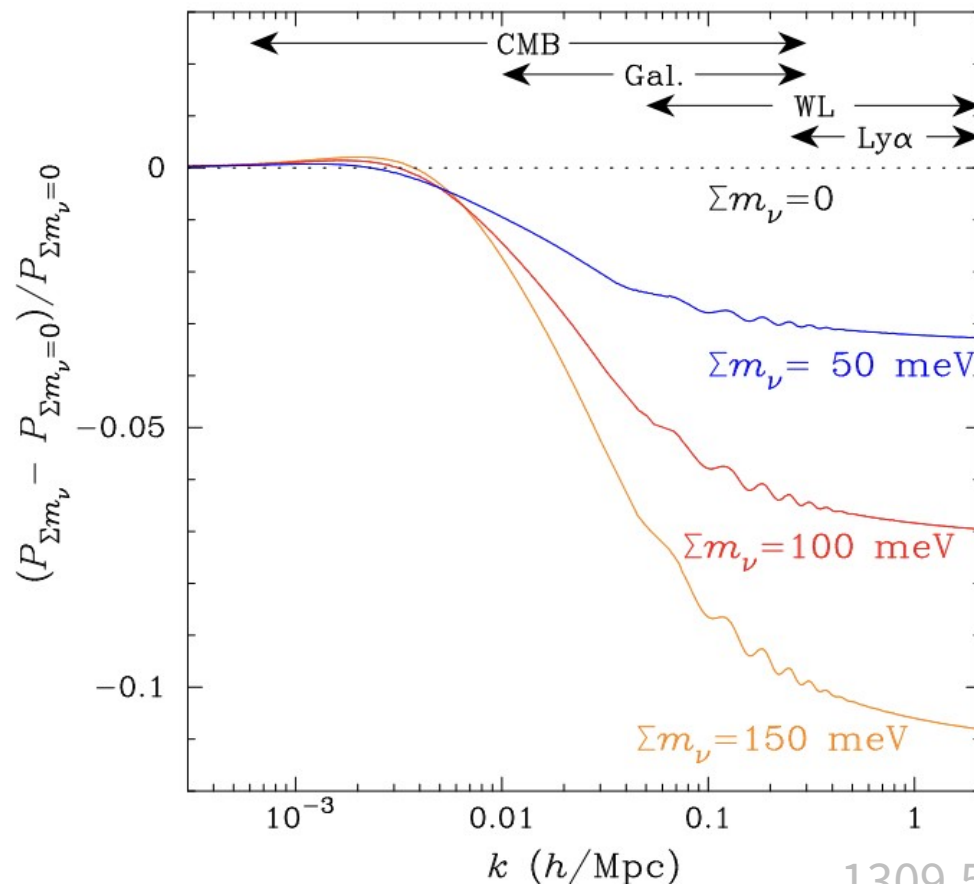
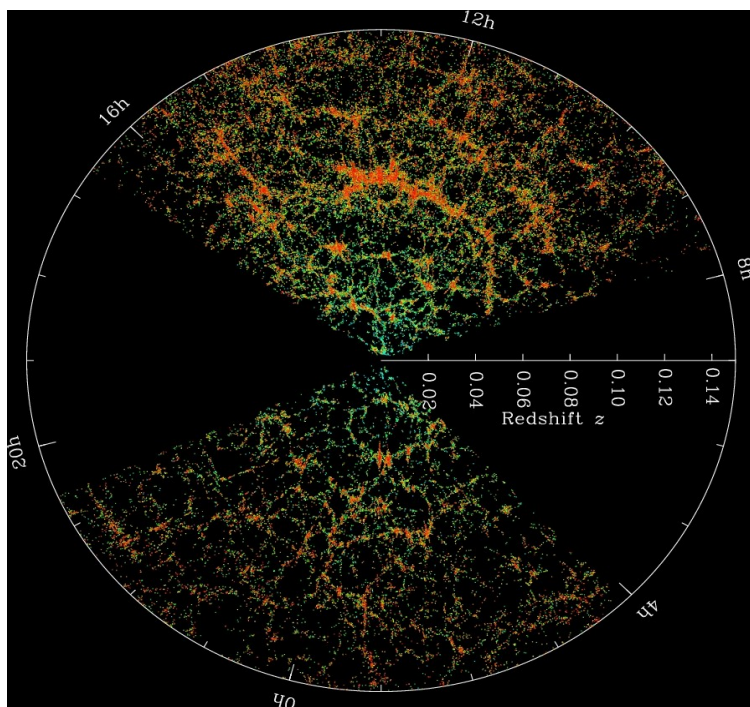
Massive neutrino decreasing CMB power spectrum



$$\rho_\nu = n_\nu \sum_i m_i$$

PDG 2022 Neutrino in Cosmology

Suppression matter power spectrum below neutrino free-streaming scale



1309.5383

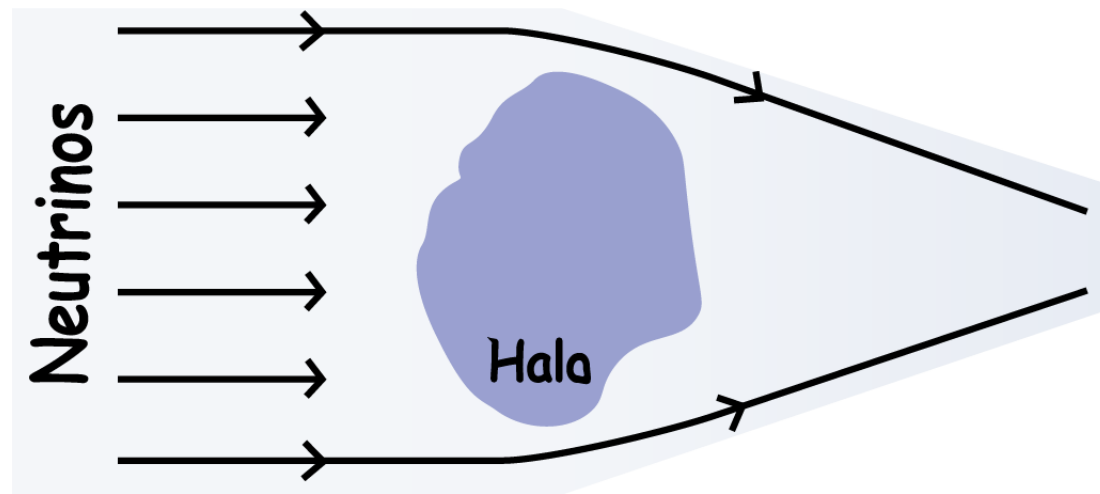
$$\rho_\nu = n_\nu \sum_i m_i$$

$$\sum m_\nu < 0.13 \text{ eV}$$

Planck18+BAO

Cosmic ν Fluid (CvF) vs DM Halo

Gravitational attraction between DM halo & CvF



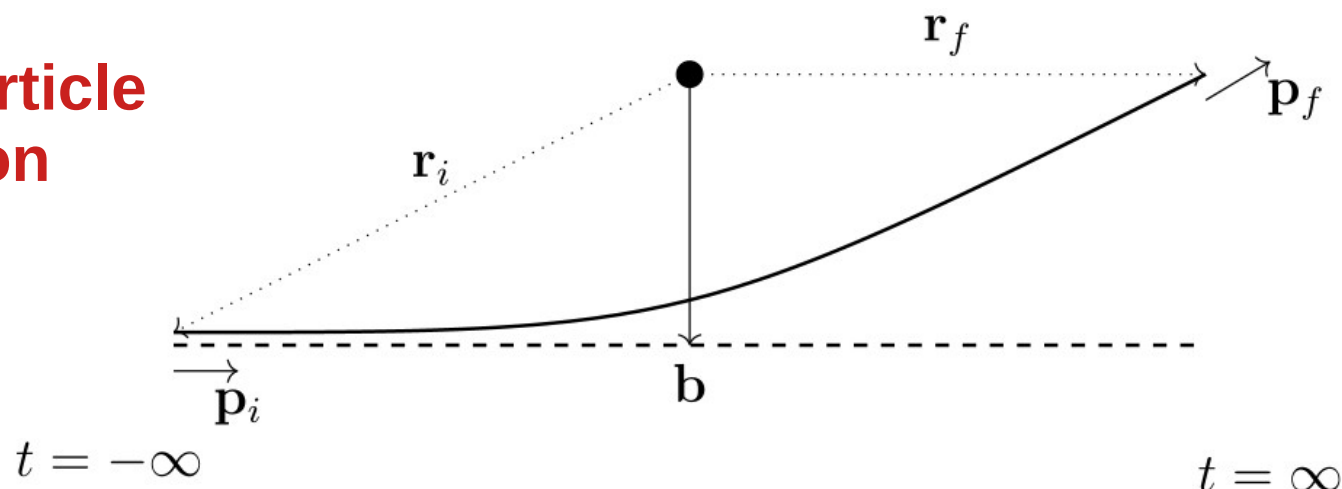
CvF

DM
Halo

vs Stone in Water Flow

Gravitational Deflection

Single-particle description

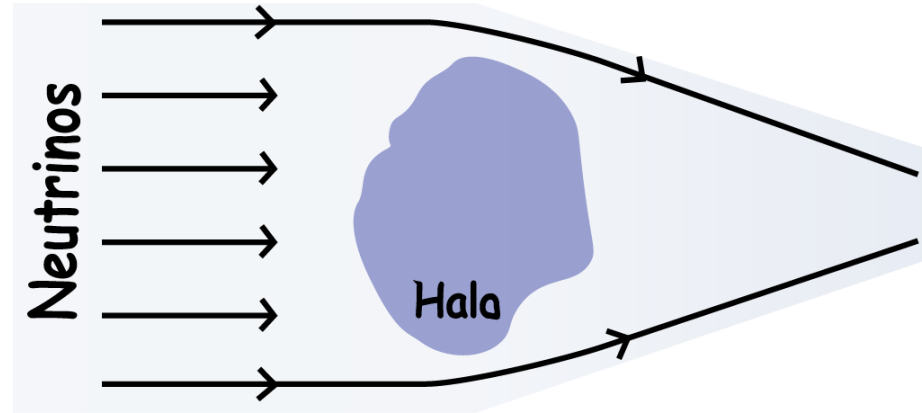


$$ds^2 \equiv - \left(1 - \frac{2GM}{r} \right) dt^2 + \left(1 - \frac{2GM}{r} \right)^{-1} dr^2 + r^2 d\Omega^2$$

$$\Delta\phi = 2|\mathbf{b}| \int_{r_{\min}}^{\infty} \frac{dr}{r^2} \left[1 + \frac{2b_{90}}{r} - \frac{|\mathbf{b}|^2}{r^2} + \frac{2GM|\mathbf{b}|^2}{r^3} \right]^{-1/2}$$

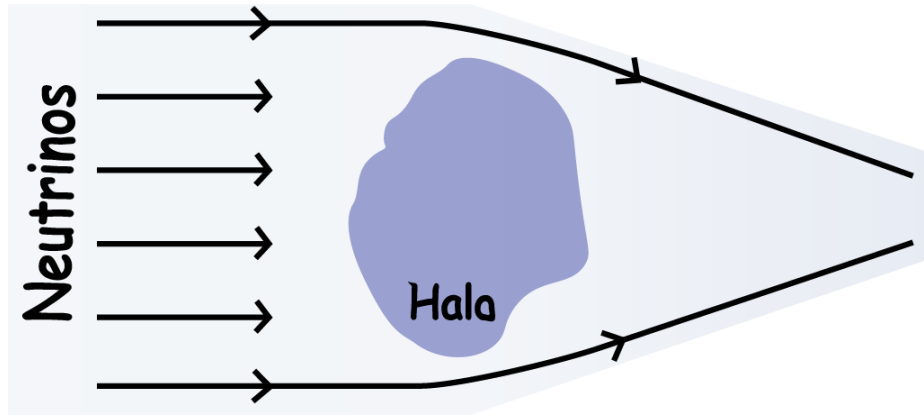
$$b_{90} \equiv GMm_{\nu}^2/|\mathbf{p}_i|^2 \quad r_{\min}^3 + 2b_{90}r_{\min}^2 - |\mathbf{b}|^2r_{\min} + 2GM|\mathbf{b}|^2 = 0$$

Single-particle description



$$\Delta\phi \approx \pi + 2\frac{GM}{|\mathbf{b}|} \left(\frac{m_\nu^2}{|\mathbf{p}_i|^2} + 2 \right) \quad f_\nu(\mathbf{p}_i, \mathbf{v}) \approx \frac{2}{e^{|\mathbf{p}_i - E_{\mathbf{p}_i} \mathbf{v}/c|/T} + 1}$$

$$\Delta\mathbf{p}^\parallel \equiv \mathbf{p}_f^\parallel - \mathbf{p}_i = (-\cos \Delta\phi - 1) \mathbf{p}_i \approx -\frac{2G^2 M^2}{|\mathbf{b}|^2} \left(\frac{m_\nu^2}{|\mathbf{p}_i|^2} + 2 \right)^2 \mathbf{p}_i$$



**Boltzmann
description**

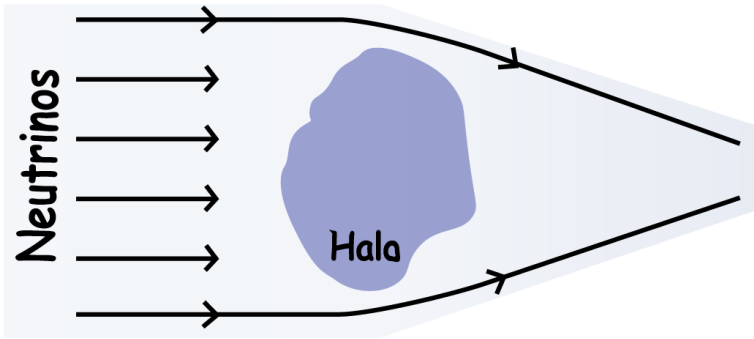
$$f_\nu(\mathbf{x}, \mathbf{p}) \equiv \bar{f}_\nu(\mathbf{p}) + \delta f_\nu(\mathbf{x}, \mathbf{p})$$

$$\left\{ \partial_t + \frac{\mathbf{p} \cdot \nabla_{\mathbf{x}}}{aE_{\mathbf{p}}} - \left[(H + \dot{\Phi})\mathbf{p} + \frac{E_{\mathbf{p}}}{a} \nabla_{\mathbf{x}} \Psi - \frac{|\mathbf{p}|^2 \nabla_{\mathbf{x}} \Phi - \mathbf{p}(\mathbf{p} \cdot \nabla_{\mathbf{x}} \Phi)}{aE_{\mathbf{p}}} \right] \cdot \nabla_{\mathbf{p}} \right\} f_\nu(\mathbf{x}, \mathbf{p}) = 0$$

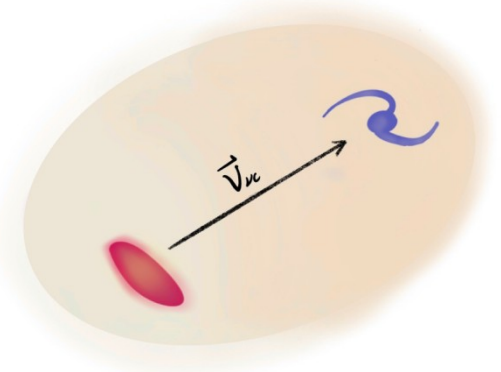
$$\delta \tilde{f}_\nu(\mathbf{k}, \mathbf{p}) = \tilde{\Psi}(\mathbf{k}) \left(\frac{m_\nu^2 + 2\mathbf{p}^2}{\mathbf{p} \cdot \mathbf{k}} \mathbf{k} - \mathbf{p} \right) \cdot \nabla_{\mathbf{p}} \bar{f}_\nu(\mathbf{p})$$

$$\text{Im}[\delta \tilde{\rho}_\nu] = - \frac{(\mathbf{v}_{\nu c} \cdot \hat{\mathbf{k}}) \tilde{\Psi}}{4\pi} \int d|\mathbf{p}'| (m_\nu^4 + 3m_\nu^2 |\mathbf{p}'|^2 + 2|\mathbf{p}'|^4) \frac{d\bar{f}_\nu}{d|\mathbf{p}'|} \Theta(|\mathbf{p}'| - E_{\mathbf{p}'} |\mathbf{v}_{\nu c} \cdot \hat{\mathbf{k}}|)$$

Cosmic Gravitational Focusing



$$\delta_m = \delta_{m0} + \sum_{i=1}^3 \frac{\delta\rho_{\nu_i}}{\rho_m}$$



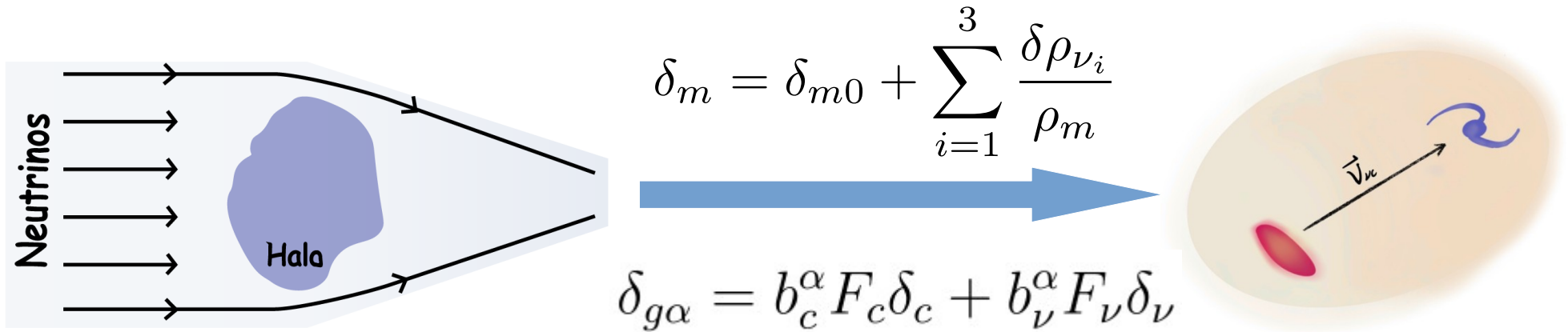
Density enhancement downwind!

$$\delta\rho_\nu(-\mathbf{x}) = -\delta\rho_\nu(\mathbf{x})$$

$$\tilde{\delta}_m \equiv \tilde{\delta}_{m0}(1 + i\tilde{\phi})$$

$$\begin{aligned} [\tilde{A}(\mathbf{k})]^* &= \int d\mathbf{x} e^{i\mathbf{k}\cdot\mathbf{x}} A(\mathbf{x}) = \int d\mathbf{x} e^{-i\mathbf{k}\cdot\mathbf{x}} A(-\mathbf{x}) \\ &= - \int d\mathbf{x} e^{-i\mathbf{k}\cdot\mathbf{x}} A(\mathbf{x}) = -\tilde{A}(\mathbf{k}) \end{aligned}$$

SFG, Pedro Pasquini, Liang Tan [arXiv:2312.16972]

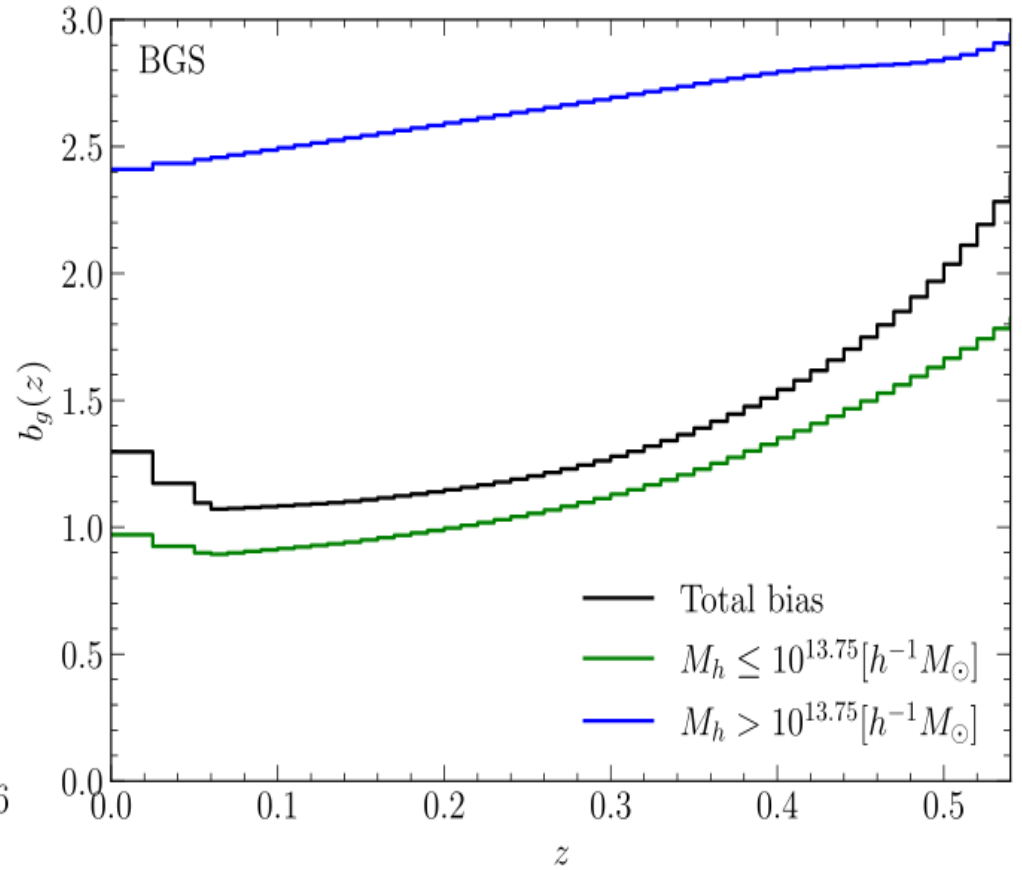
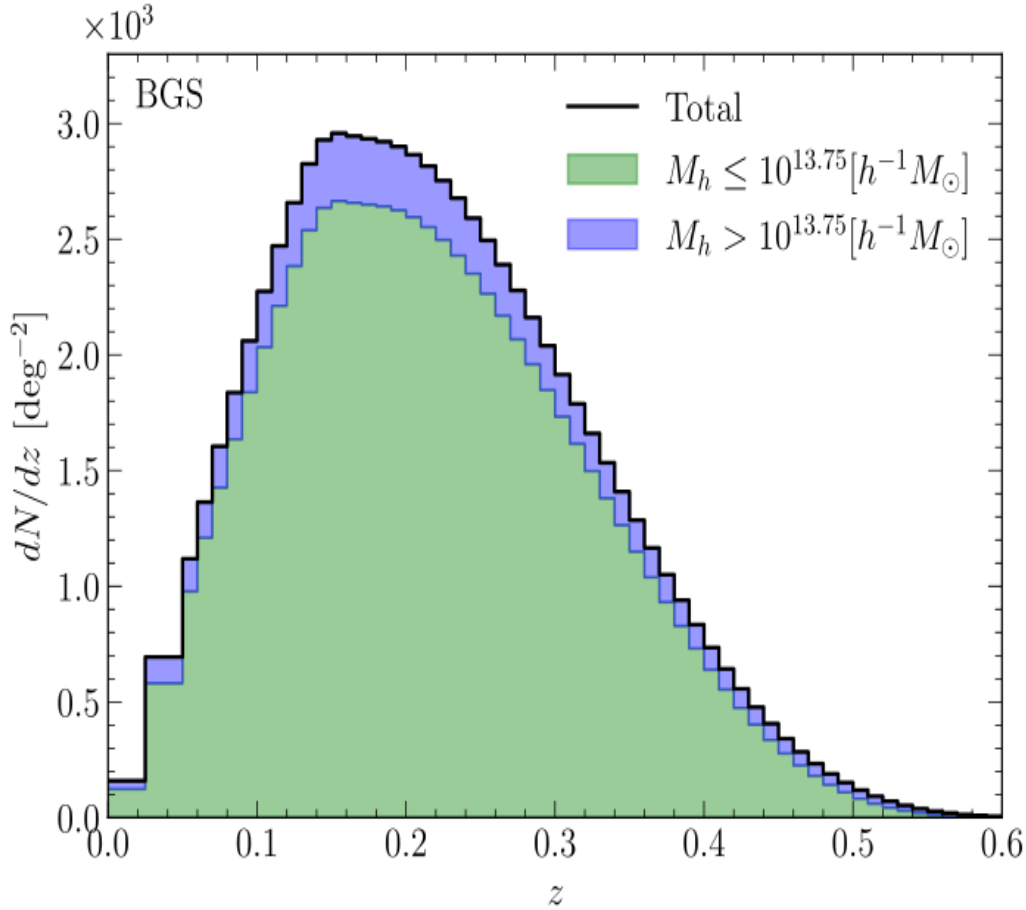


$$\delta_{g\alpha, \text{RSD}}(\mathbf{x}) \equiv \delta_{g\alpha}(\mathbf{x}) - \frac{\partial}{\partial x} \left(\frac{\mathbf{u}_m \cdot \hat{\mathbf{x}}}{aH} \right)$$

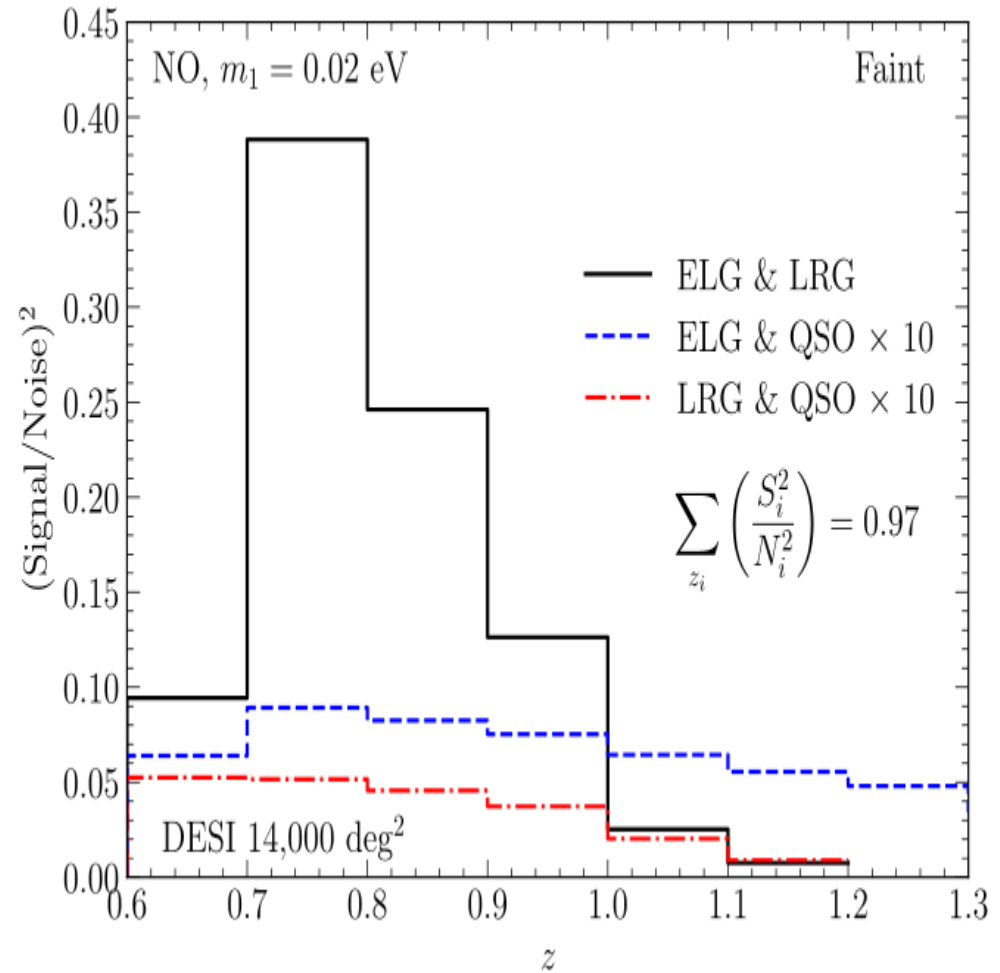
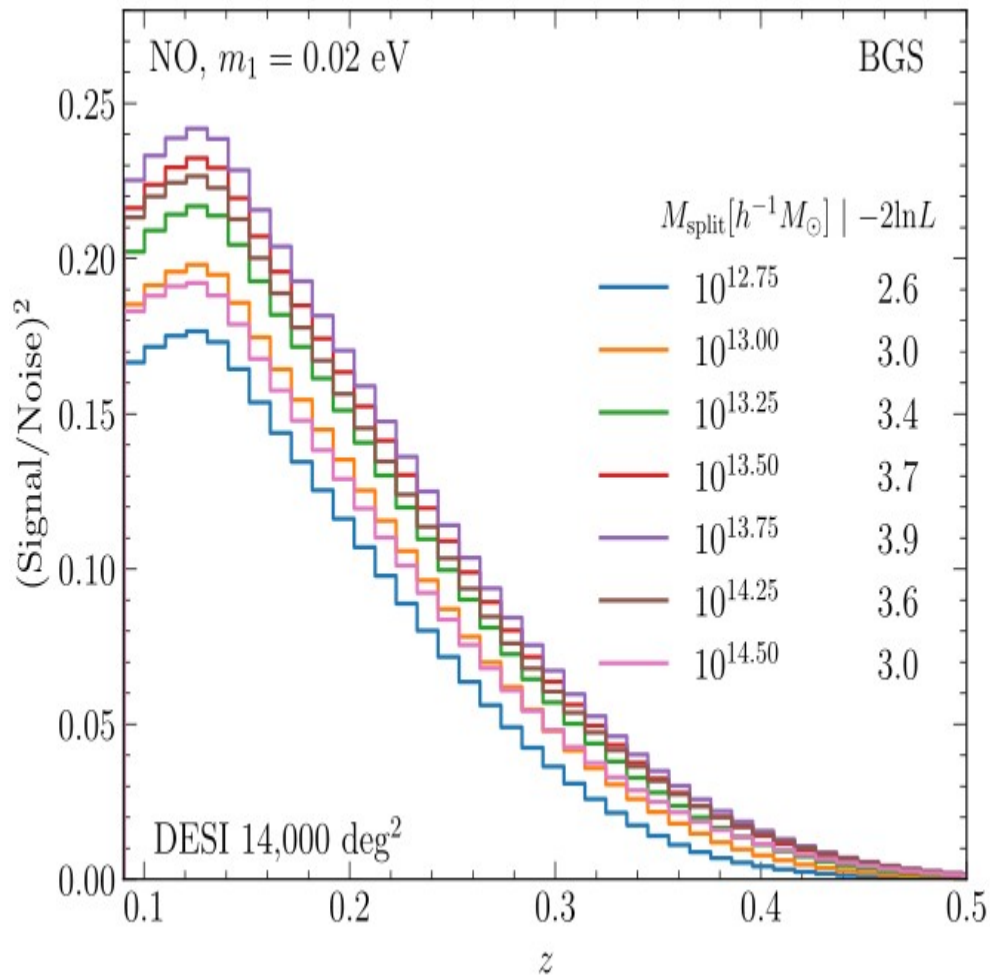
$$\text{Im}[\tilde{\delta}_{g\alpha, \text{RSD}} \tilde{\delta}_{g\beta, \text{RSD}}^*] = -i\Delta b \left[\mu_{\mathbf{k}}^2 \frac{\dot{\tilde{\phi}}}{H} + (f\mu_{\mathbf{k}}^2 + 1)\tilde{\phi} \right] \tilde{\delta}_{m0}^2$$

Galaxies with different bias

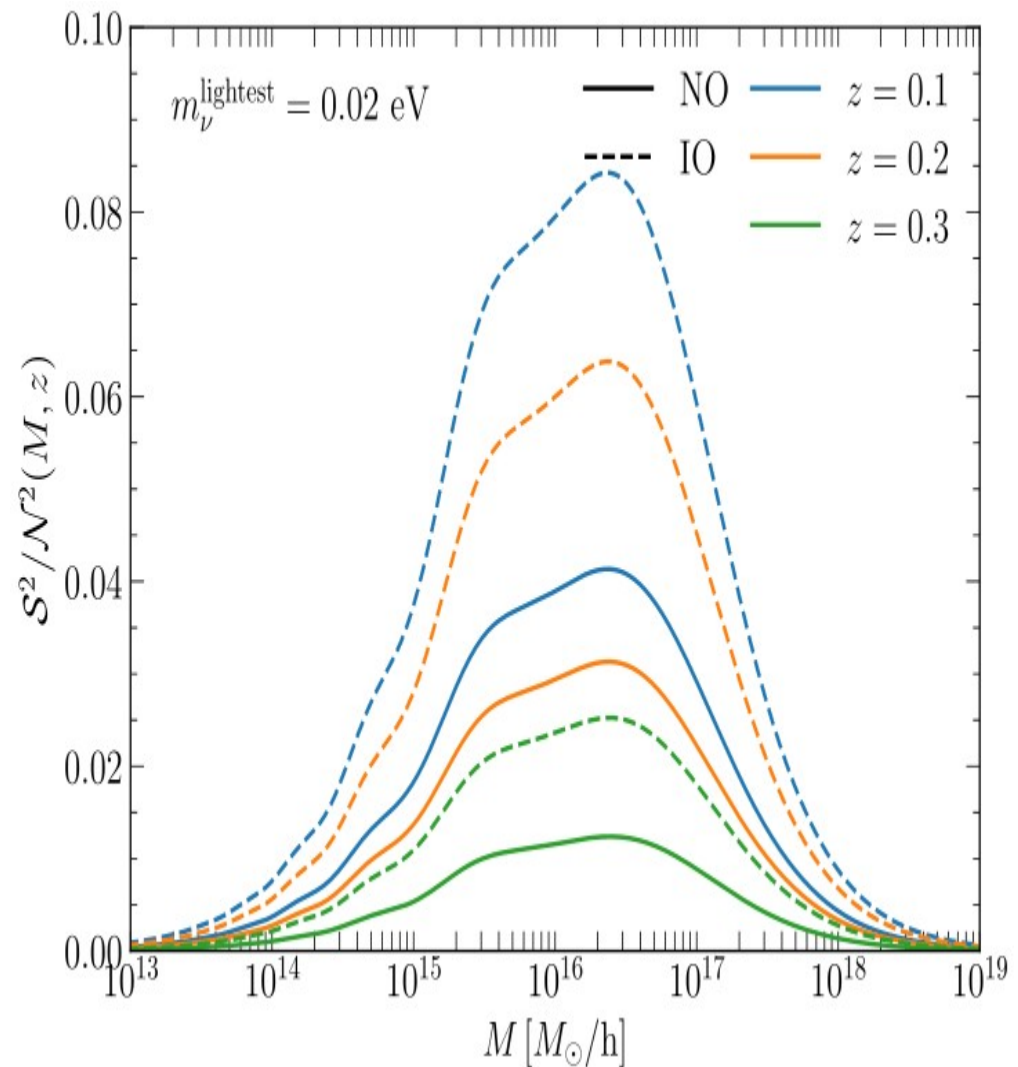
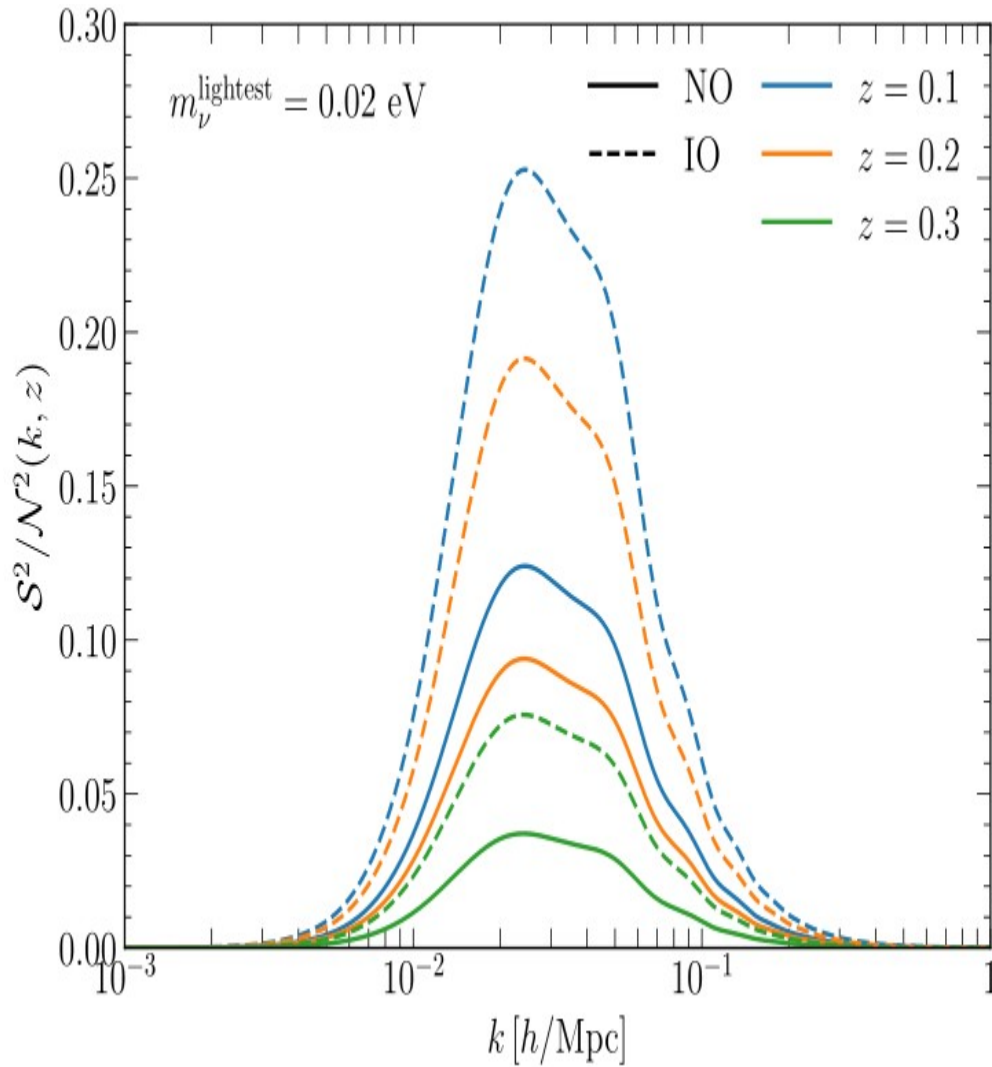
$$\Delta b \equiv b_c^\alpha - b_c^\beta$$



$$n_g(z) \equiv \int d \ln M_h \frac{dn(z)}{d \ln M_h} \langle N(M_h) \rangle \quad b_g(z) \equiv \frac{1}{n_g} \int d \ln M_h \frac{dn(z)}{d \ln M_h} \langle N(M_h) \rangle b_h(M_h, z)$$

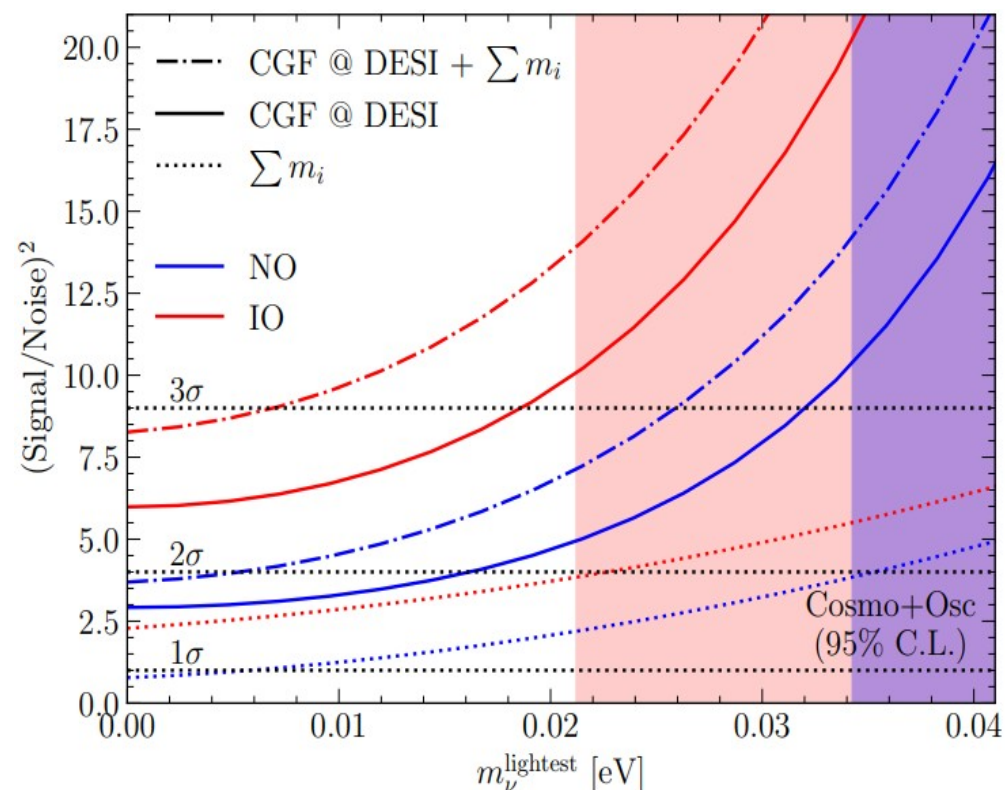
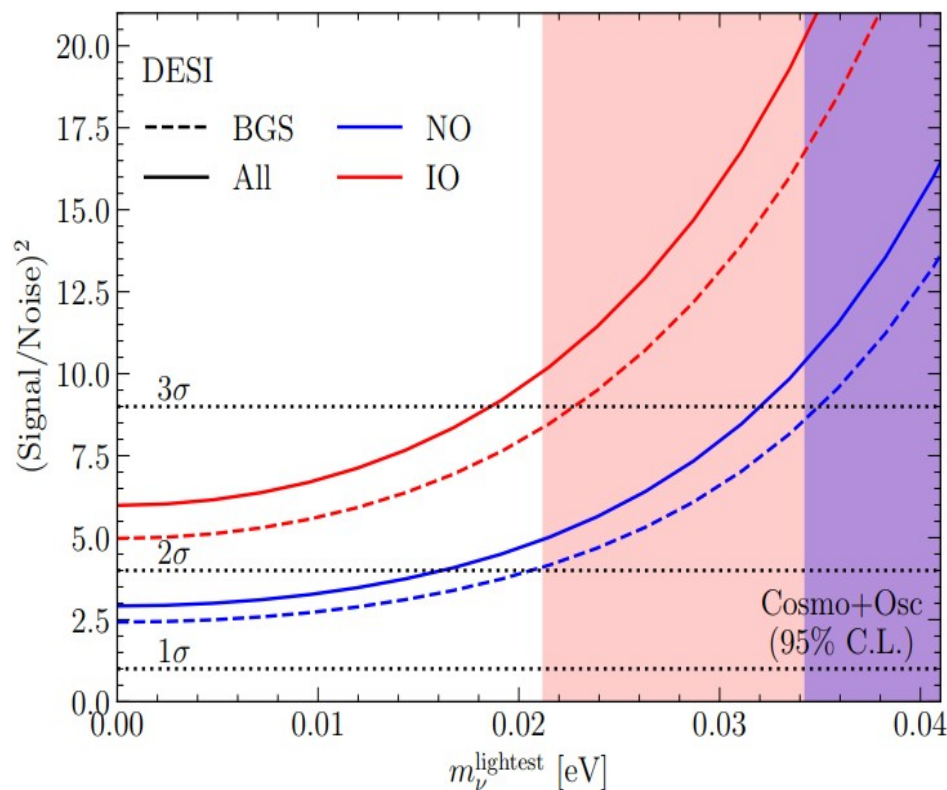


Redshift Distribution



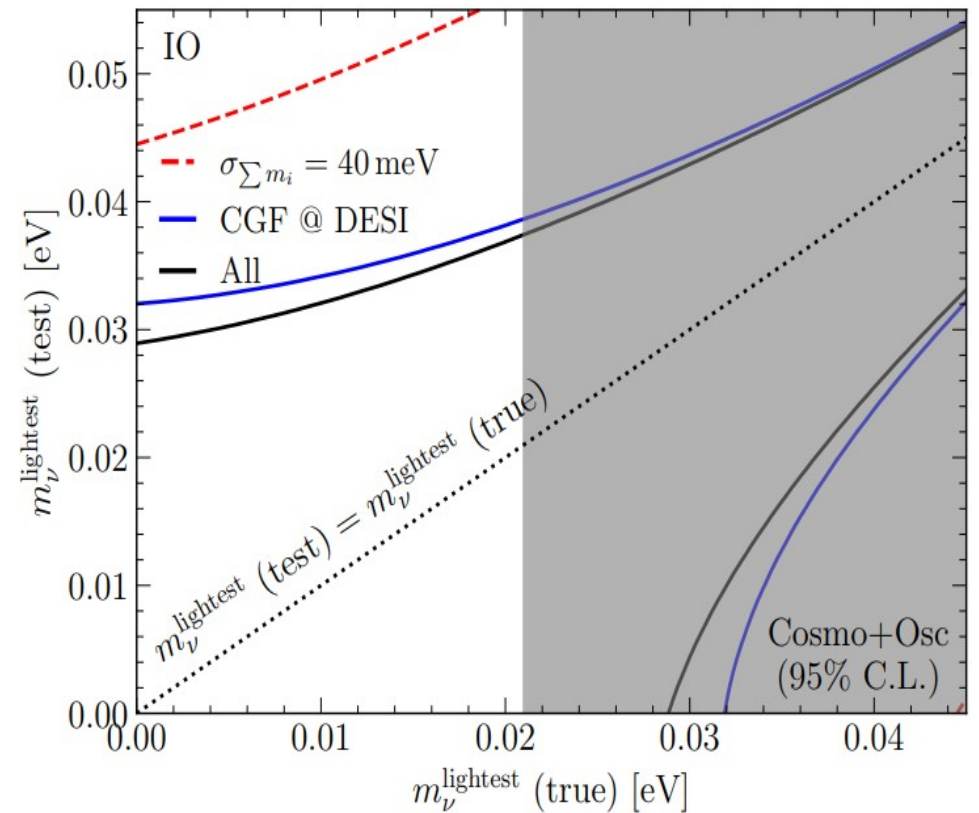
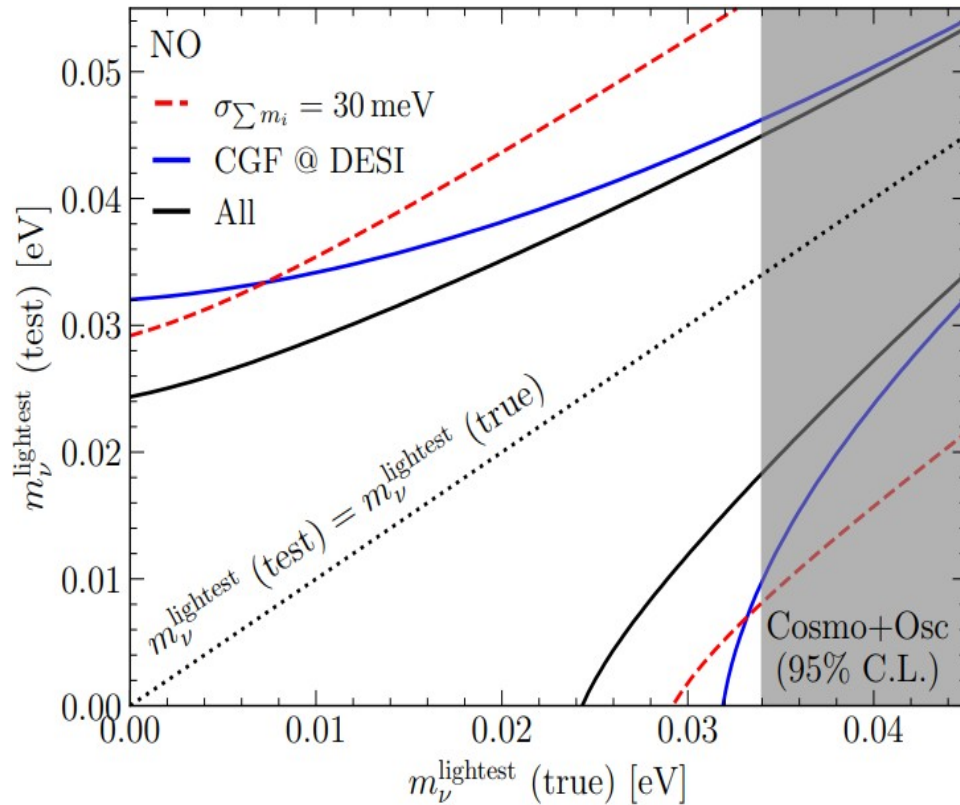
3rd Cosmic Measurement of ν Masses

$$\text{Im} \left[\tilde{\delta}_{g\alpha} \left(\tilde{\delta}_{g\beta} \right)^* \right] \propto \left(\mathbf{v}_{\nu c} \cdot \hat{\mathbf{k}} \right) \left(f_0 m_\nu^4 + f_1 m_\nu^2 T^2 + f_2 T^4 \right)$$



SFG, Pedro Pasquini, Liang Tan [arXiv:2312.16972]

$$\text{Im} \left[\tilde{\delta}_{g\alpha} \left(\tilde{\delta}_{g\beta} \right)^* \right] \propto \left(\mathbf{v}_{\nu c} \cdot \hat{\mathbf{k}} \right) \left(f_0 m_\nu^4 + f_1 m_\nu^2 T^2 + f_2 T^4 \right)$$



DESI, Euclid, Subaru PFS, **CSST**

SFG, Pedro Pasquini, Liang Tan [arXiv:2312.16972]

1) Overview of Neutrino Mass

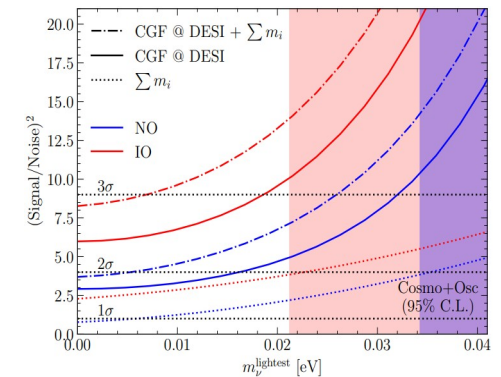
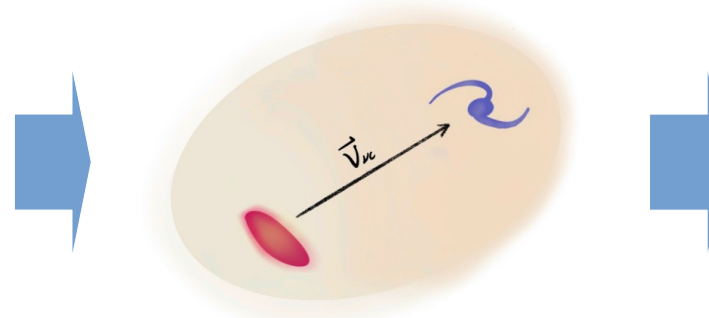
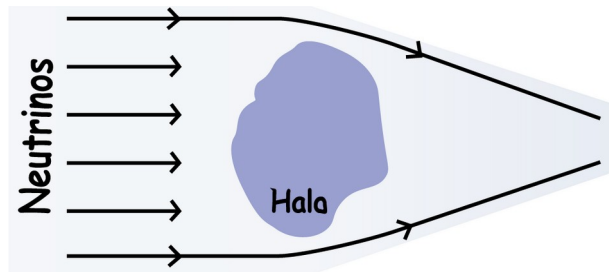
2) 3rd Cosmological Way

Dynamical Friction

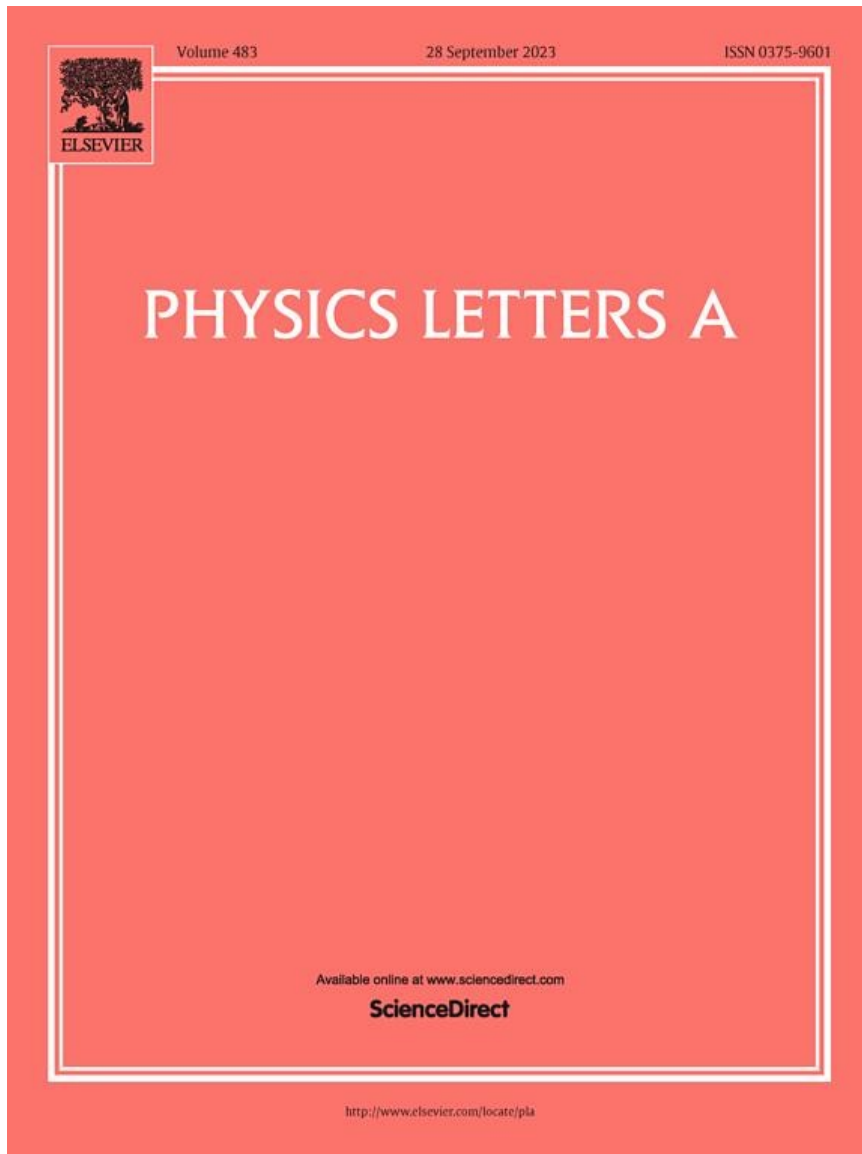
Cosmic Gravitational Focusing

Dipole Structure in Galaxy Correlation Function

Neutrino Mass Measurement @ DESI & CSST



3) Summary



Aims & Scope

- Nonlinear science,
- Statistical physics,
- Mathematical and computational physics,
- AMO and physics of complex systems,
- Plasma and fluid physics,
- Optical physics,
- General and cross-disciplinary physics,
- Biological physics and nanoscience,
- Astrophysics, Particle physics and Cosmology.

Thank You

Georg G. Raffelt

Stars as Laboratories for Fundamental Physics

The Astrophysics of Neutrinos, Axions, and Other
Weakly Interacting Particles

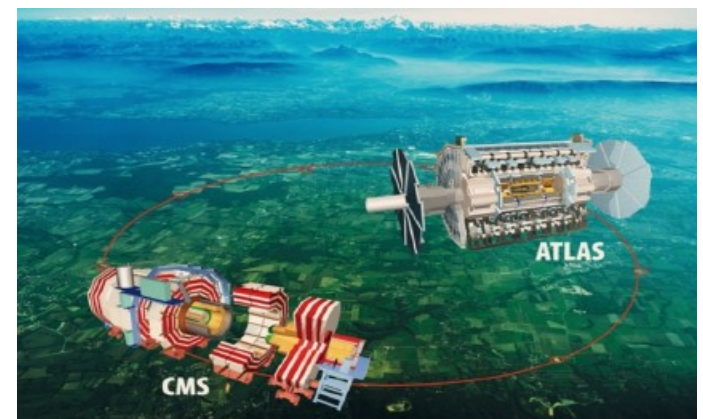
In the standard model, neutrinos have been assigned the most minimal properties compatible with experimental data: zero mass, zero charge, zero dipole moments, zero decay rate, zero almost everything.

Neutrino vs Higgs

- **Higgs boson** \Rightarrow electroweak symmetry breaking & mass. $\sim O(100)\text{GeV}$
- **Chiral symmetry breaking** \Rightarrow majority of mass.
- The world seems not affected by the tiny neutrino mass?
 - Neutrino mass \Rightarrow Mixing
 - 3 Neutrino \Rightarrow possible **CP violation**
 - CP violation \Rightarrow **Leptogenesis**
 - \Rightarrow **Matter-Antimatter Asymmetry**
 - There is something left in the Universe.
 - **EW Baryogenesis** is not enough.

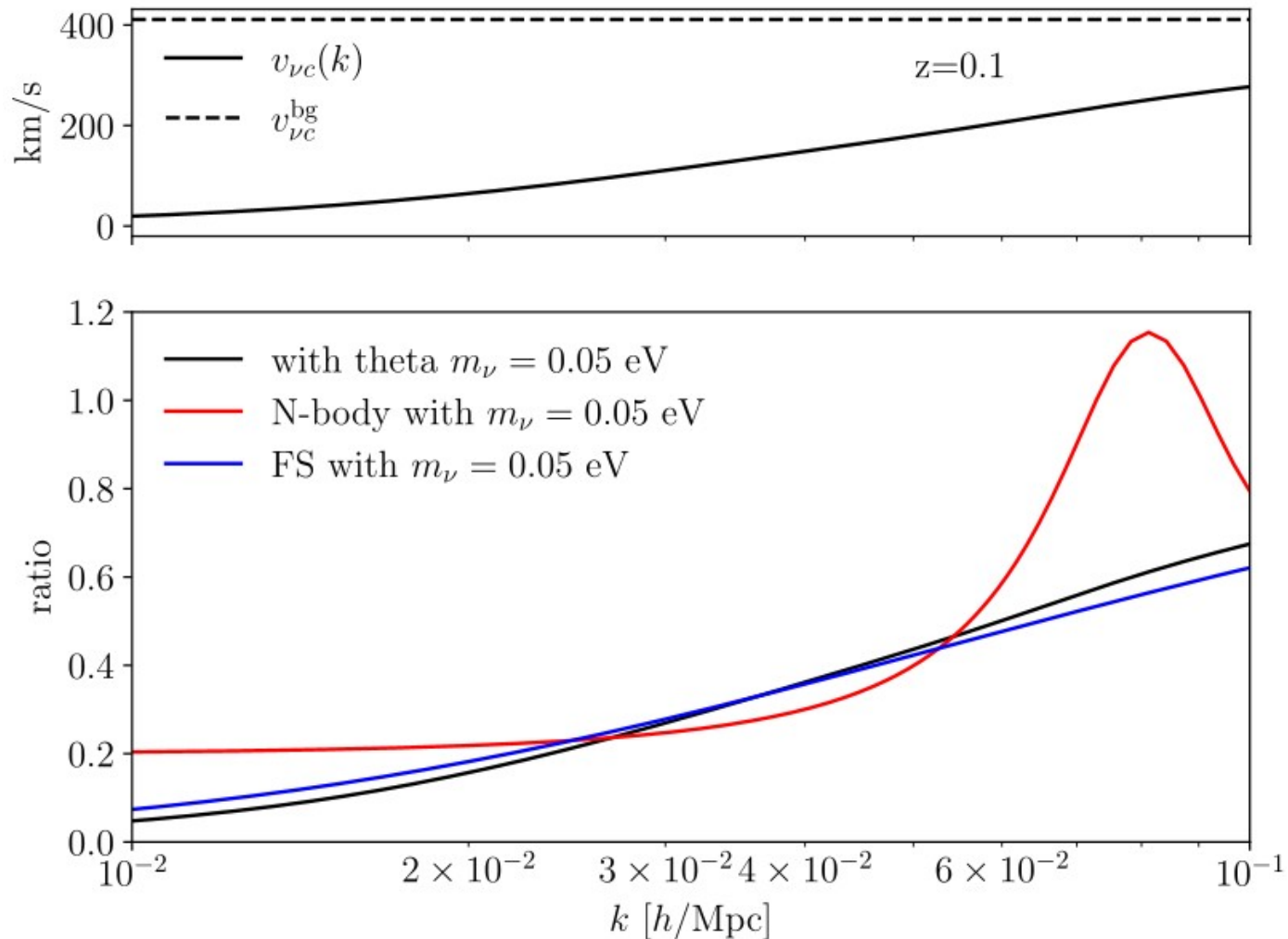


Daya Bay @ **March 8, 2012**

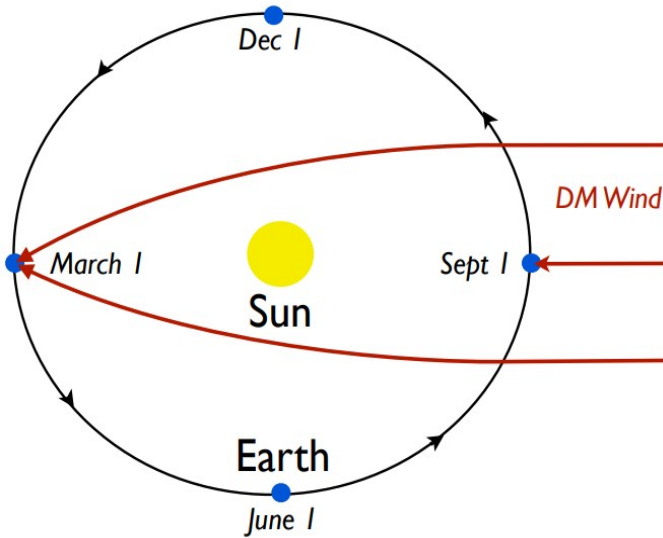


LHC @ **July 4, 2012**

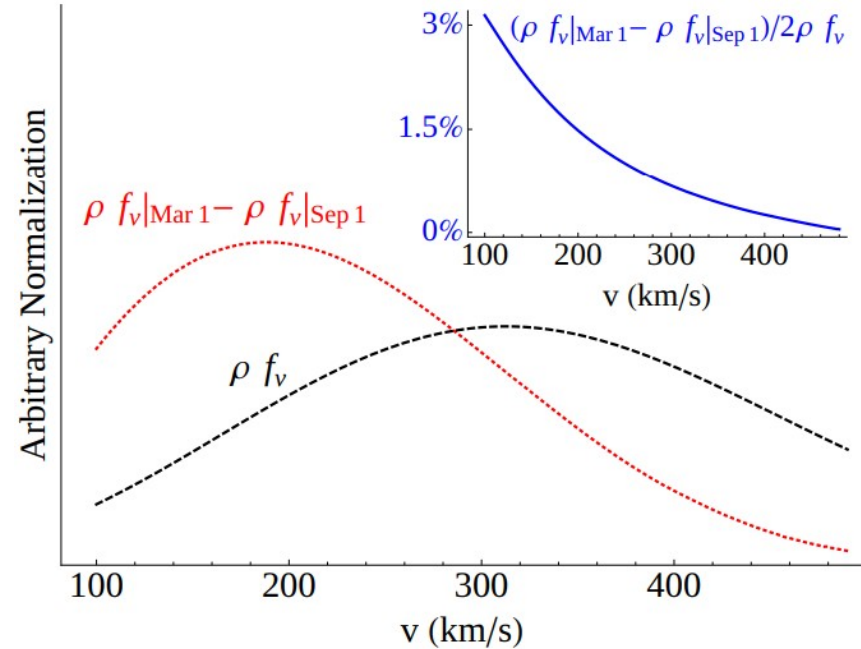
$$\langle (\mathbf{v}_{\nu c} \cdot \hat{\mathbf{k}})^2 \rangle = \frac{1}{3} \int \frac{d|\mathbf{k}'|}{|\mathbf{k}'|} \Theta(|\mathbf{k}| - |\mathbf{k}'|) \left| \widetilde{W}(|\mathbf{k}'|R) \right|^2 \Delta_{\zeta}^2(\mathbf{k}') \left| \frac{T_{\theta_{\nu ic}}(\mathbf{k}', z)}{|\mathbf{k}'|} \right|^2$$



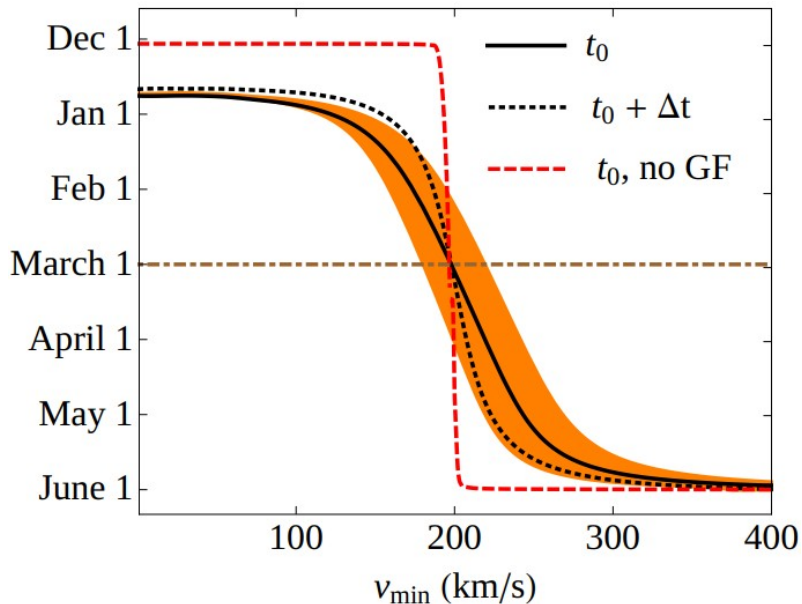
DM Gravitational Focusing



- Density enhancement



- Modulation phase shift

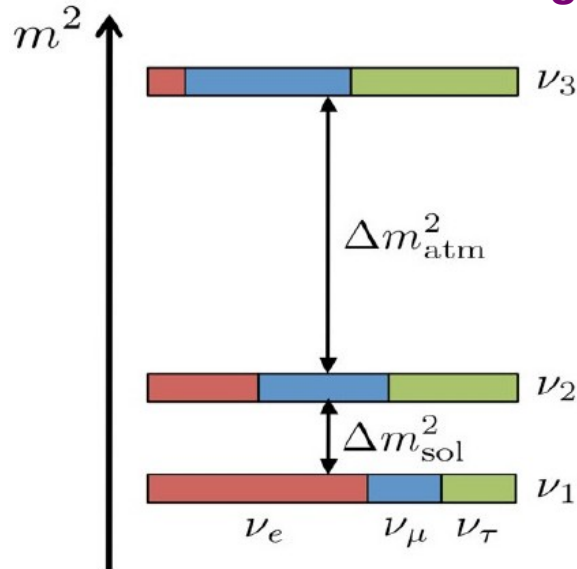


Lee, Lisanti, Peter, Safdi, Phys. Rev. Lett. 112, 011301 (2014) [arXiv:1308.1953]

Bozorgnia & Schwetz, JCAP 08 (2014) 013 [arXiv:1405.2340]

ν Mass Ordering

Normal Ordering (正序)



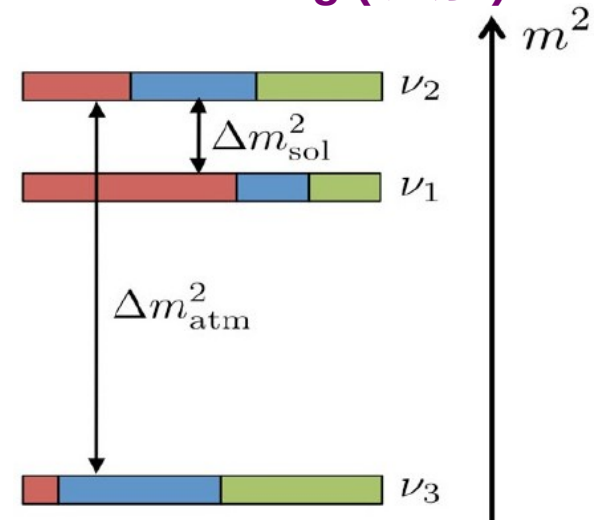
$$m_1 \lesssim m_2 < m_3$$

m_1

$$m_2 = \sqrt{m_1^2 + \Delta m_s^2}$$

$$m_3 = \sqrt{m_1^2 + \Delta m_a^2}$$

Inverted Ordering (反序)



$$m_3 < m_1 \lesssim m_2$$

$$m_1 = \sqrt{m_3^2 + \Delta m_a^2}$$

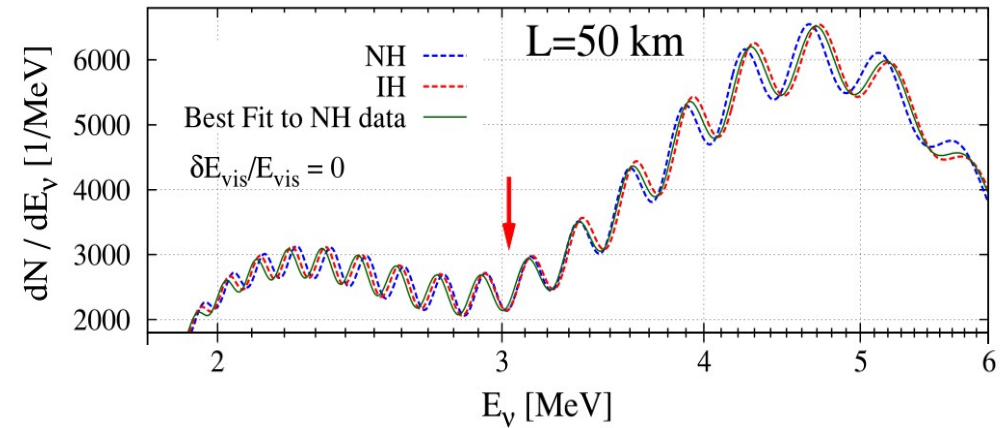
$$m_2 = \sqrt{m_3^2 + \Delta m_a^2 + \Delta m_s^2}$$

m_3

江门中微子实验 (JUNO)

However, ν oscillation experiment cannot measure the absolute mass!

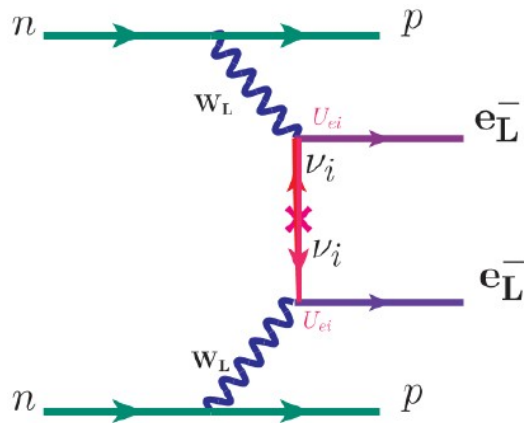
ν Mass Hierarchy @ JUNO



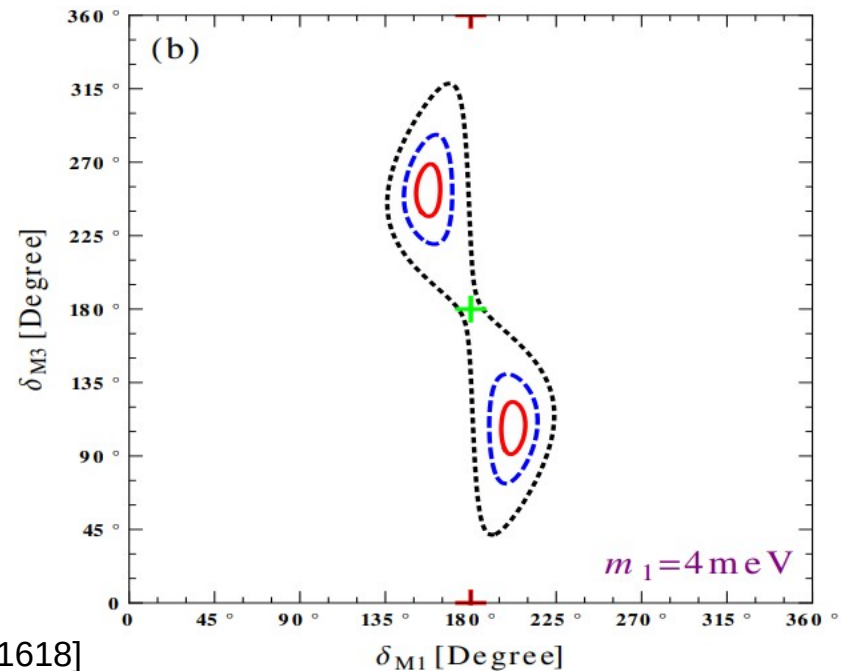
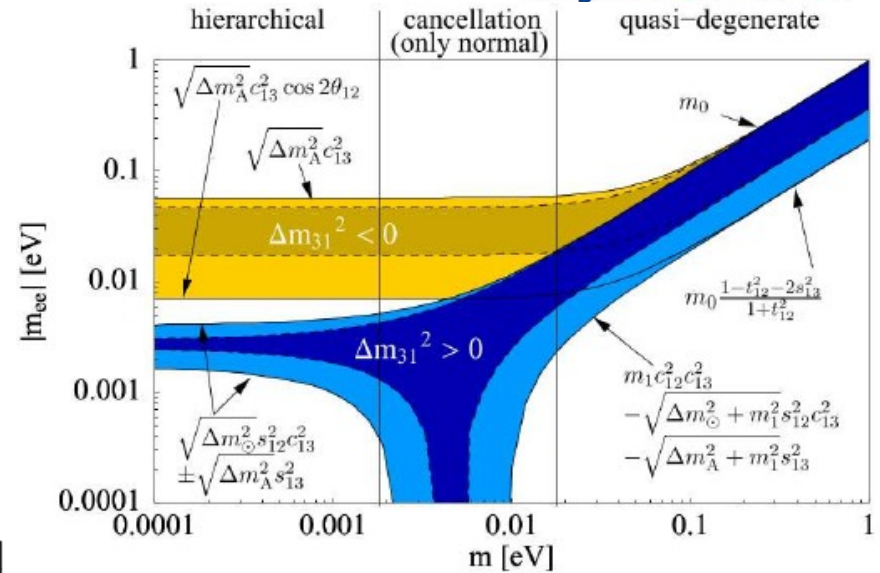
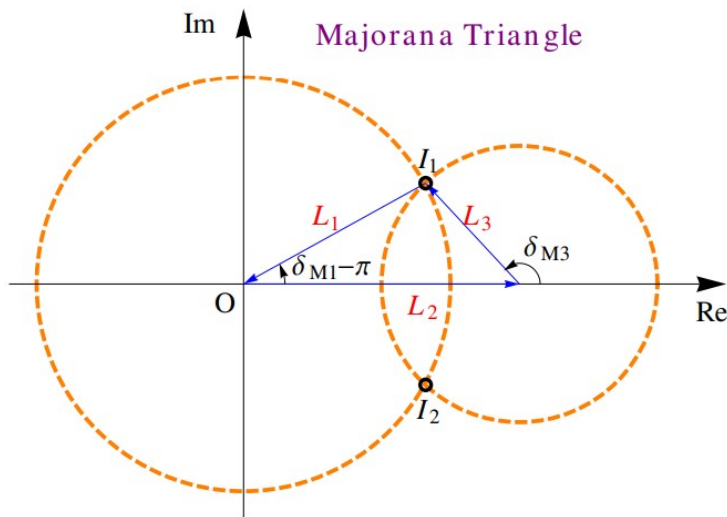
$$P_{ee} = 1 - \cos^4 \theta_{13} \sin^2 2\theta_{12} \sin^2 (\Delta_{21}) \\ - \cos^2 \theta_{12} \sin^2 2\theta_{13} \sin^2 (\Delta_{31}) \\ - \sin^2 \theta_{12} \sin^2 2\theta_{13} \sin^2 (\Delta_{32})$$

3 frequencies:

Neutrinoless Double Beta Decay

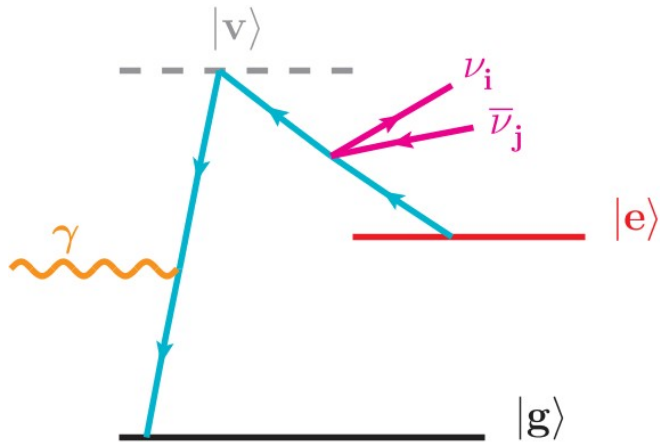


$$\langle m \rangle_{ee} \equiv \left| \sum_i m_i U_{ei}^2 \right| = \left| c_s^2 c_r^2 m_1 e^{i\delta_{M1}} + s_s^2 c_r^2 m_2 + s_r^2 m_3 e^{i\delta_{M3}} \right|$$



SFG & Manfred Lindner, PRD 95 (2017) No.3, 033003 [arXiv:1608.01618]

Radiative Emission of ν Pairs

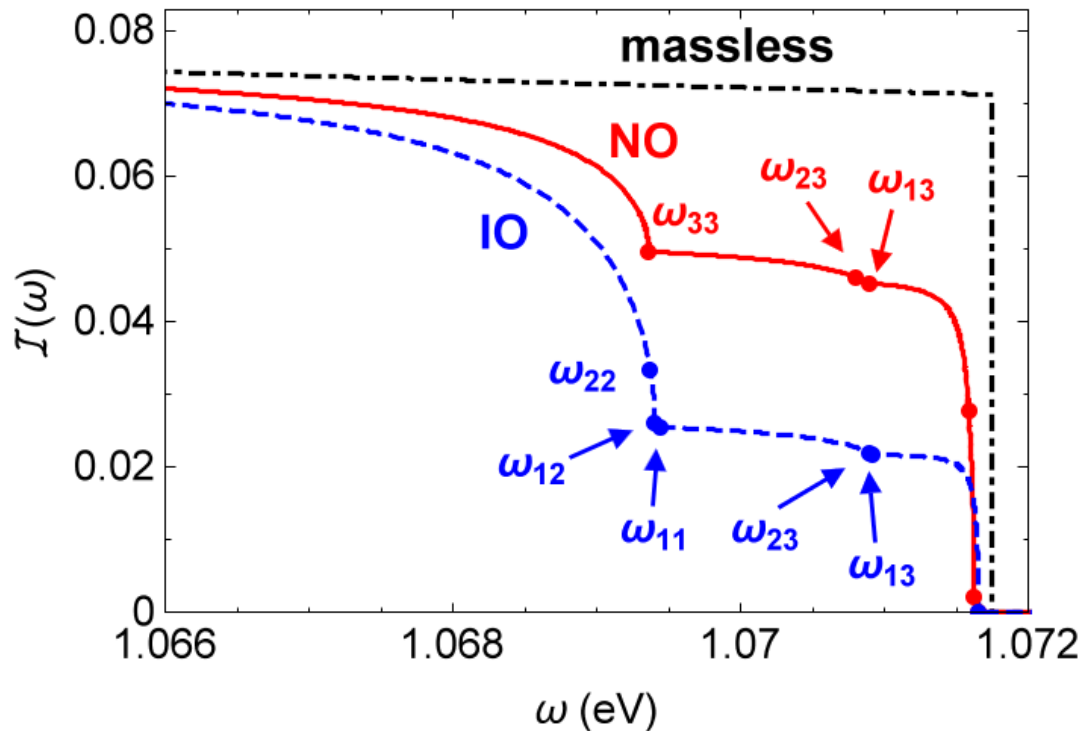


$$|e\rangle \rightarrow |\nu\rangle + \nu\bar{\nu}$$

$$E1 \times M1$$

M. Yoshimura, Phys. Rev. D 75, 113007 (2007)

Jue Zhang & Shun Zhou, Phys.Rev.D 93 (2016) 11, 113020

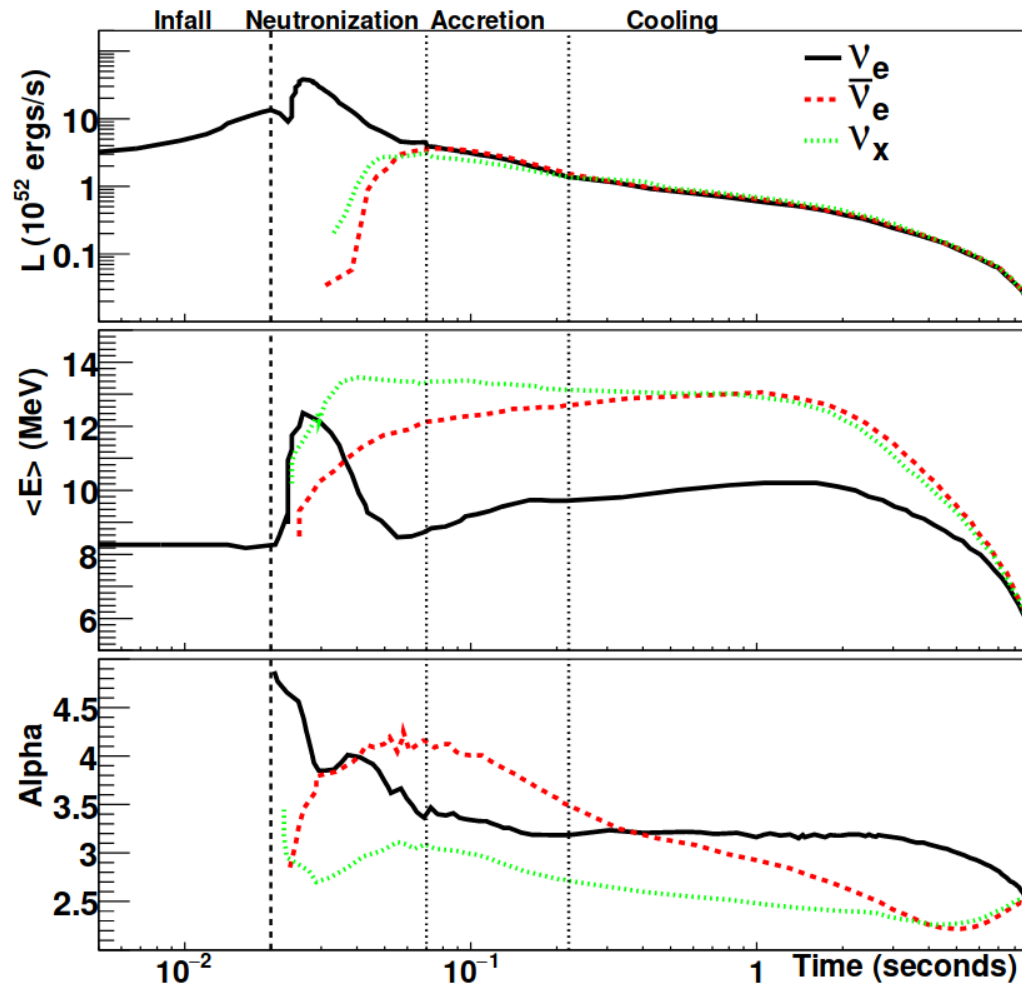


$$\omega_{ij}^{\max} \equiv \frac{E_e - E_g}{2} - \frac{1}{2} \frac{(m_i + m_j)^2}{(E_e - E_g)}$$

SFG & Pasquini

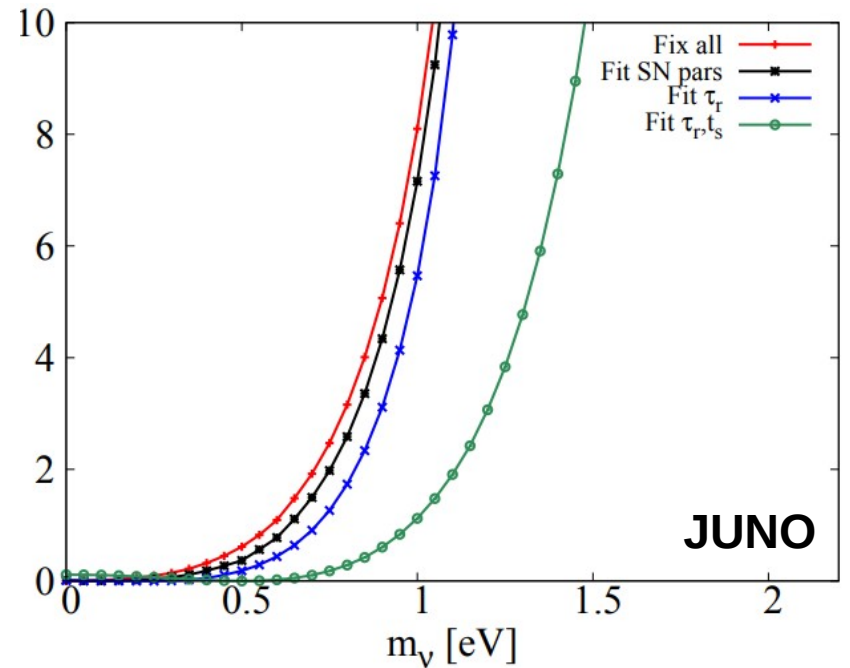
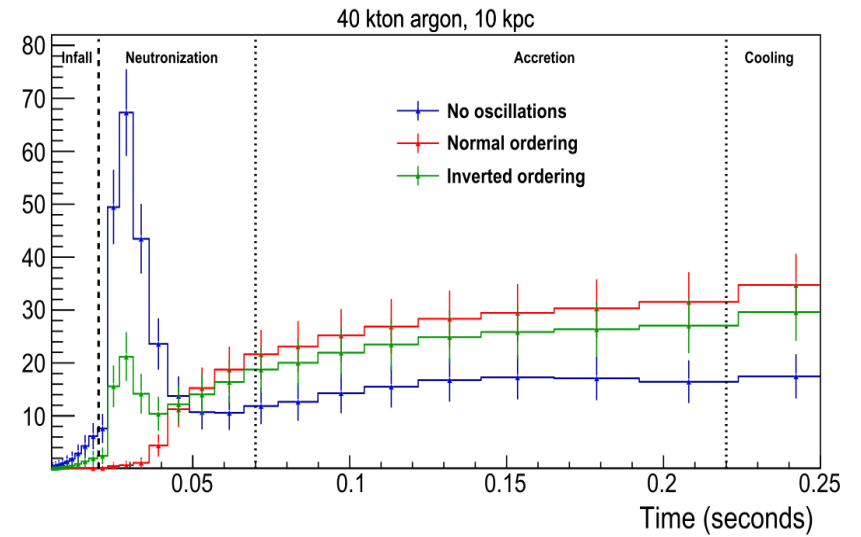
Eur.Phys.J.C 82 (2022) 3, 208;
Phys.Lett.B 841 (2023) 137911;
JHEP 12 (2023) 083

Time Delay in Supernova ν 's



DUNE, 2008.06647

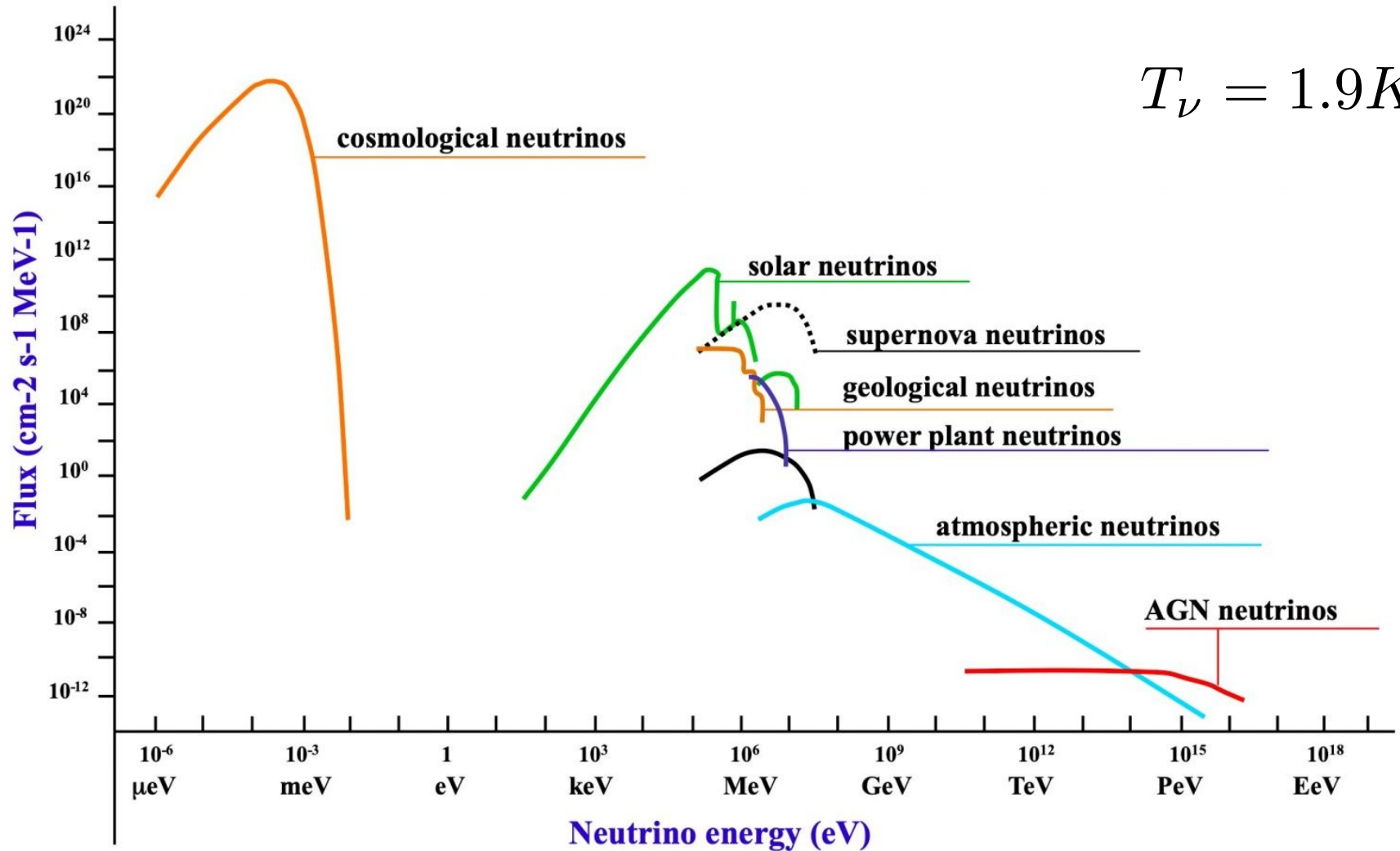
SFG, Chui-Fan Kong, Alexei Smirnov,
accepted by PRL [arXiv:2404.17352]



JUNO

Lu, Cao, Li & Zhou [JCAP 05 (2015) 044]

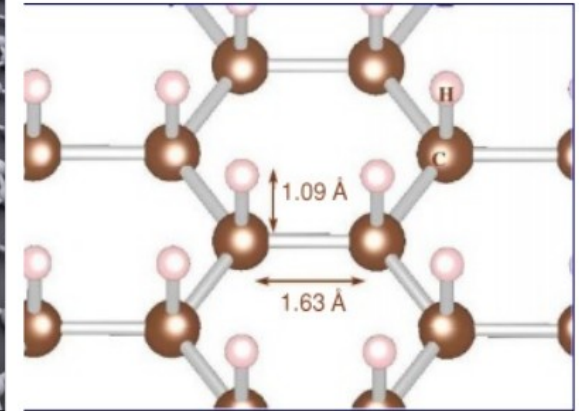
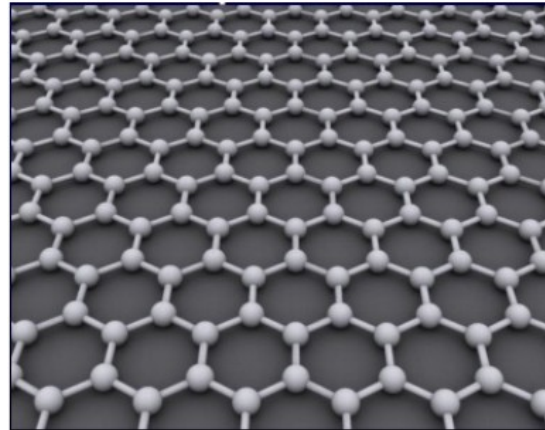
Cosmic ν Background



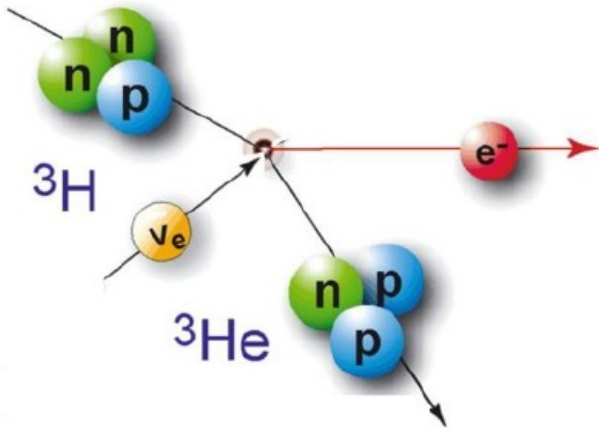
PTOLEMY Experiment



P on-
T ecorvo
O bservatory for
L ight,
E arly-universe,
M assive-neutrino
Y ield

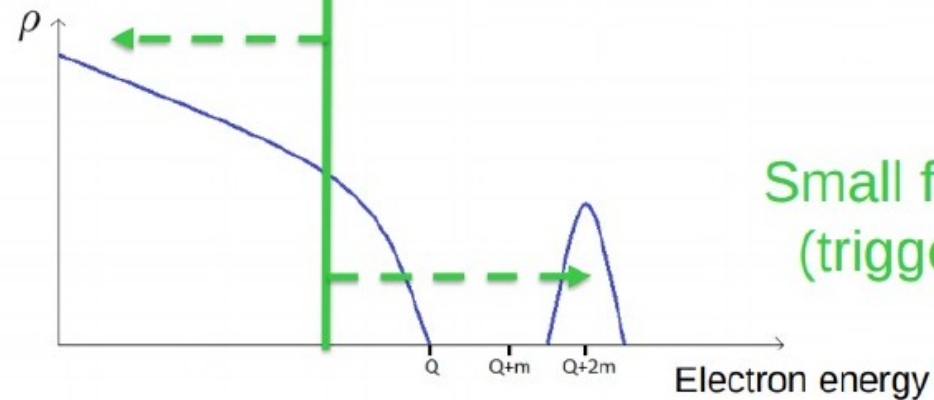


Graphene Substrate



Too much rate
(need to filter)

Need very high energy
resolution ($\sigma \sim m_\nu$)



SFG & Pasquini, Phys.Lett.B 811 (2020) 135961