



Measurement of Iron Spectrum in Cosmic Rays with DAMPE

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Outline

- Motivation
- DAMPE Instrument
- Data sample
- Reconstruction and selections
- Template fit
- Systematic uncertainties
- Result : Iron flux





Motivation

- Iron provides a favorable condition to understand Systematic uncertaintie ATIC 02 (2003 CRN-spacelab2
- cosmic ray origin and propagation Most abundant heavy nuclei beyond Si 0
 - Contamination from spallation of heavier 0 elements is negligible





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- Previous measurements on iron flux follow single power law (SPL) above hundred GeV/n
- DAMPE has capability to extend the measurement of Iron to higher energy with high statistics





The DArk Matter Particle Explorer



DAMPE main goals

- Dark Matter indirect detection
- Origins and propagations of CRs
- High Energy γ astronomy

Orbit information

- Sun-synchronous Orbit
- Altitude: ~500 km
- Inclination: ~97 deg
- Launch Data: Dec.17th,2015

DAMPE sub-detectors

- PSD : Charge measurement and anticoincidence of γ
- STK : Tracking and additional charge measurement
- BGO : Energy measurement and e/p discrimination
- NUD : Further e/p discrimination





Data Sample

Flight Data

- 7 years : 2016. 01.01~ 2022. 12.31
- Live Time : 1.68×10⁸s
 - 1. Dead time of the instrument is excluded
 - 2. Data in South Atlantic Anomaly (SAA) region is excluded
 - 3. Data during 2017.09 Solar Flare is excluded

Simulation

• Geant4.10.5.p02(FTFP BERT)



5





Reconstruction and Selections

1.Not in SAA

2.High Energy Trigger

3.Fiducial : Machine-Learning based STK track

- STK track should be fully contained in (PSD + STK +BGO)
- Angle between STK and BGO Track < 15 deg
- Average distance from energy center is less than 25 mm at first 4 layers of BGO
- Track pass through the max energy bar or the bar near the max energy bar in PSD Layer 0 & 1

Fe

4.BGO :

- E_{dep} > 100 GeV
- ID of strip with max deposited energy is not at the edge of BGO

5.PSD :

• $|Q_0 - Q_1| < 3$







Template Fit





- Mn is the main source of the background
- Total contamination from other nuclei is less than 4%



Unfold and Effective Acceptance

Bayesian Unfolding Method

 $N_{inc,i} = \sum M_{ij} N_{obs,j}$

- *M_{ij}* : response matrix
- Nobs, j : Events in jth deposited energy bin
- $N_{inc,i}$: Events in *i*th incident energy bin









Systematic Uncertainties (1)



1.Uncertainty estimation $\epsilon_{i} = \frac{N_{Layer_0\&Layer_1}}{N_{Layer_0\&Layer_1}}$

$$-\frac{N_{Layer_i}}{N_{Layer_i}}$$

- ϵ_i : Charge efficiency of PSD layer *i*, *i* = 0,1
- N_{Layer₀&Layer₁}: Events passing the Selections of layer 0 and layer 1
- N_{Layer_i} : Events passing layer *i*
- Layer 0 : MC/Data difference <4%
- Layer 1 : MC/Data difference <7%







Systematic Uncertainties (2)









DAMPE Iron Spectrum

• Flux calculation : $\phi(E_i, E_i + \Delta E_i) = \frac{\Delta N_i}{\Delta E_i A_{eff,i} \Delta T}$



 Conclusion : DAMPE observed a hardening of the iron spectrum at several hundred GeV/n





Summary

- Seven years of DAMPE data are analyzed for Iron flux
- Preliminary analysis has yielded the iron flux up to 10 TeV/n, which is roughly consistent with result of AMS and observed spectral hardening at several hundred GeV/n
- With further refinement and more data, the measurement of iron flux can be extended up to tens of TeV/n in the near future

Thank you !