

A systematic investigation on dark matter-electron scattering in effective field theories

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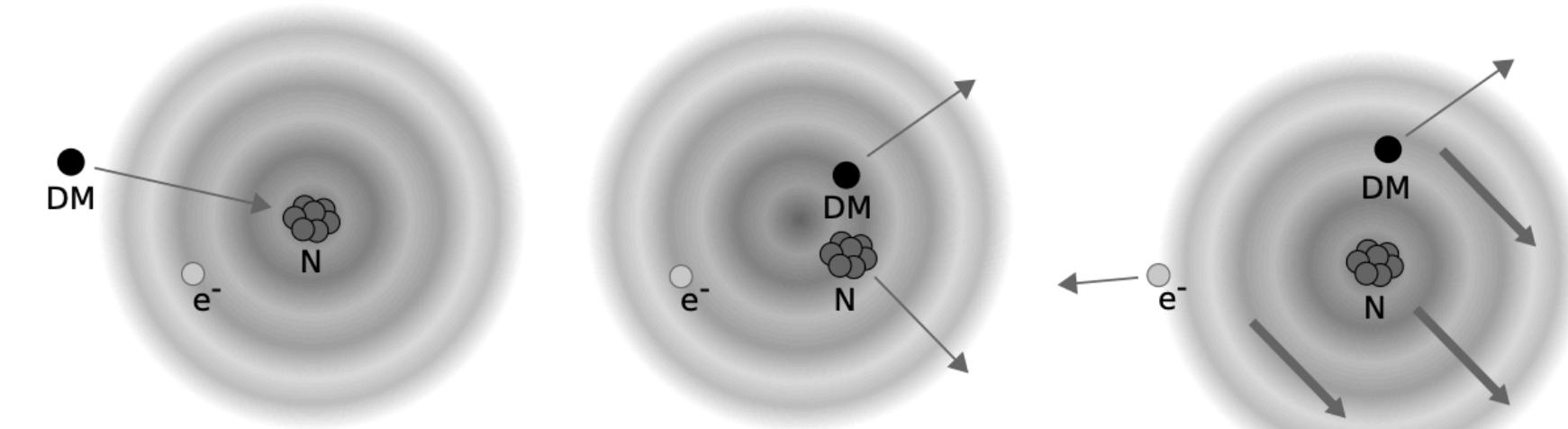
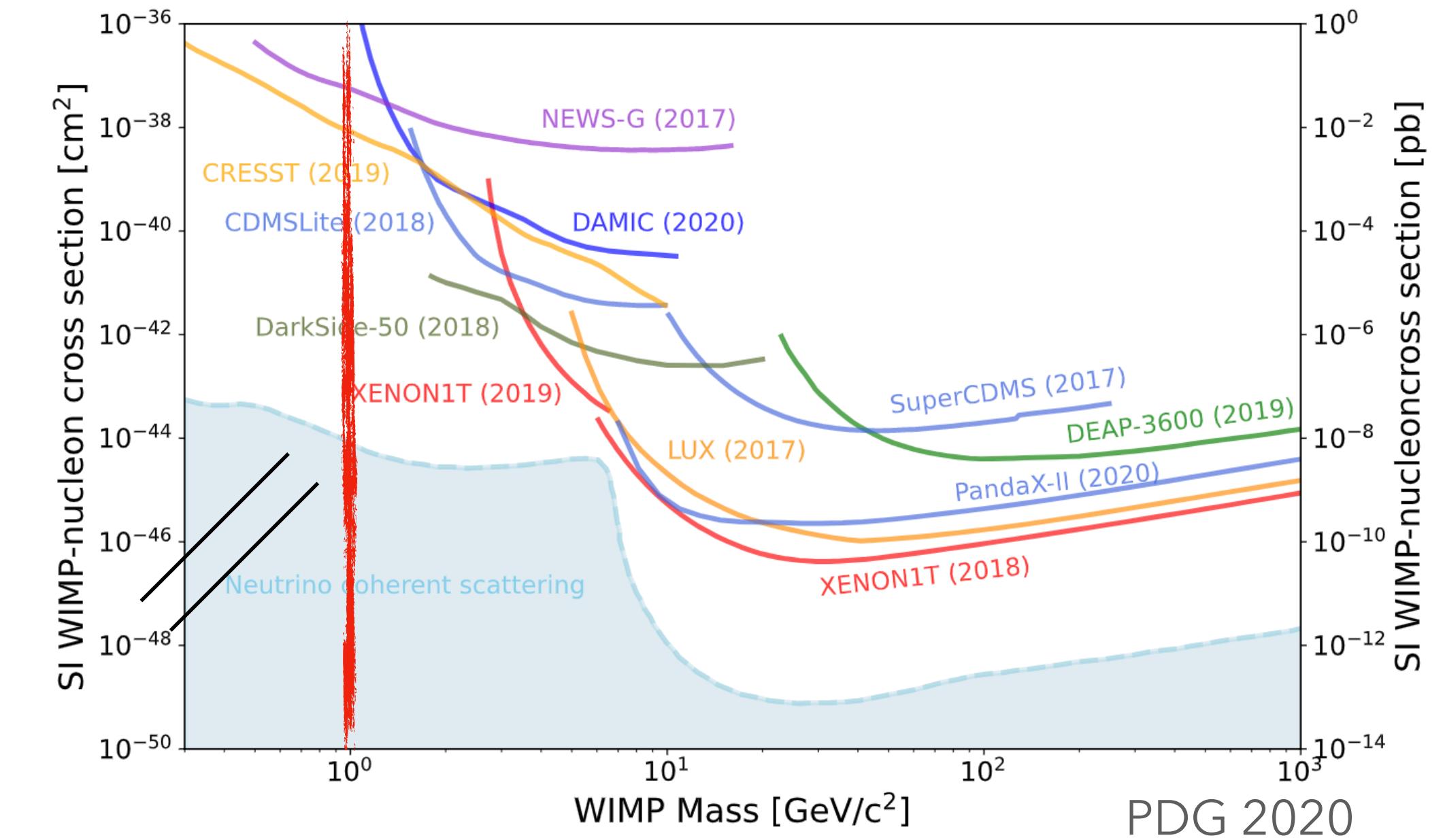
In collaboration with Jin-Han Liang, Yi Liao and Xiao-Dong Ma

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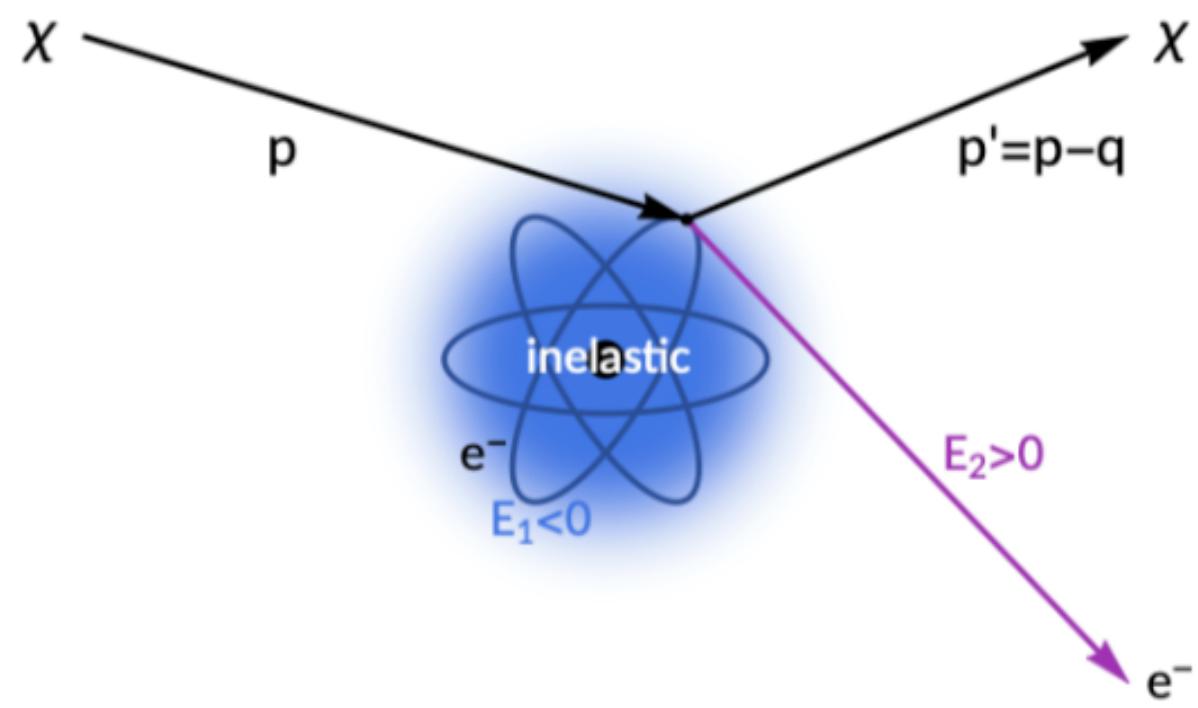
DM direct detection (DMDD) and electron recoil (ER)

- Nuclear recoil
 - ✓ Conventional direct DM searches
 - ✗ Lose sensitivity to DM masses below ~ 1 GeV
- Migdal effects
 - ✓ Observe the electron ionised after DM induced but unobservable nuclear recoil
 - ✗ The Migdal effect has never been observed
- **Electron recoil**
 - ✓ Sensitive to sub-MeV DM
- ...



1711.09906

Formalism for DM-atom scattering



Catena et al., 1912.08204

Summing over the quantum numbers
of the initial and final state electrons

$$\frac{d\mathcal{R}_{\text{ion}}^{n\ell}}{d \ln E_e} = \frac{n_x}{128\pi m_x^2 m_e^2} \int dq q \int \frac{d^3v}{v} f_x(v) \Theta(v - v_{\min}) \overline{\left| \mathcal{M}_{\text{ion}}^{n\ell} \right|^2}$$

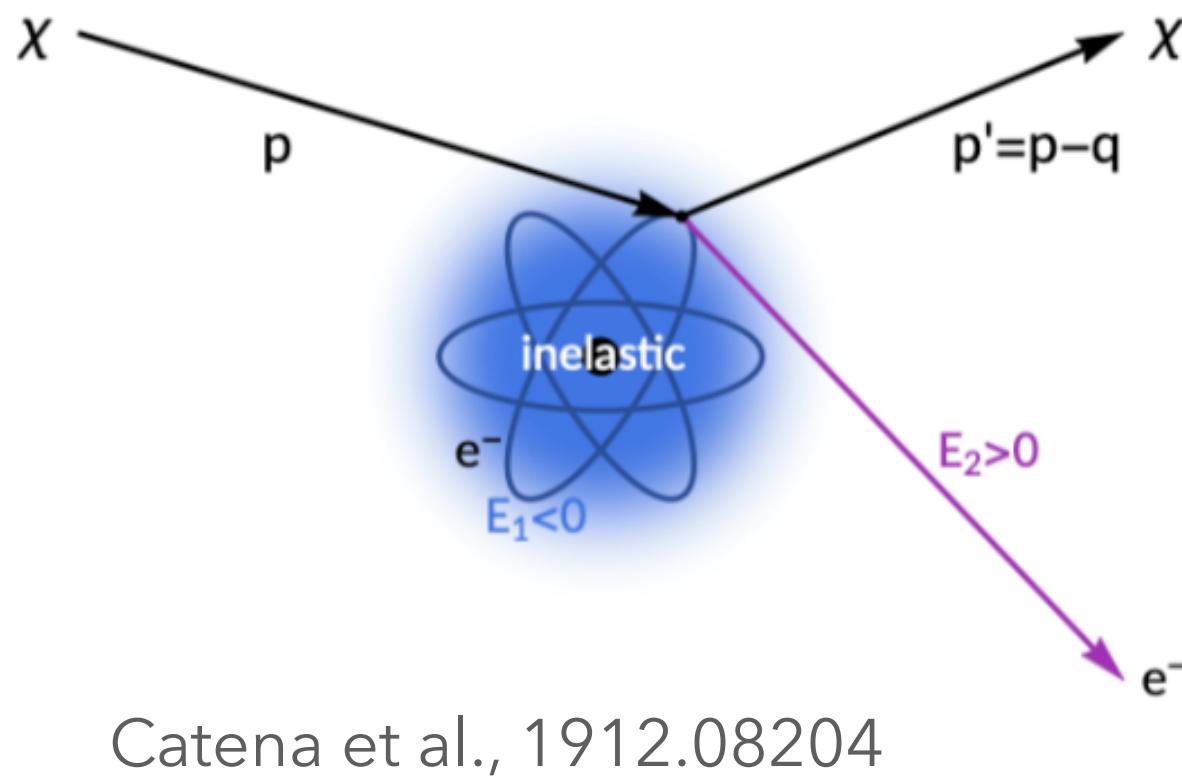
$$\overline{\left| \mathcal{M}_{\text{ion}}^{n\ell} \right|^2} \equiv \frac{4Vk'^3}{(2\pi)^3} \sum_{m=-\ell}^{\ell} \sum_{\ell'=0}^{\infty} \sum_{m'=-\ell'}^{\ell'} \overline{\left| \mathcal{M}_{1 \rightarrow 2} \right|^2}$$

$$\mathcal{M}_{1 \rightarrow 2} = \int \frac{d^3k}{(2\pi)^3} \tilde{\psi}_2^*(\mathbf{k} + \mathbf{q}) \mathcal{M}(\mathbf{q}, \mathbf{v}_{\text{el}}^\perp) \tilde{\psi}_1(\mathbf{k})$$

Free electron-DM amplitude

Atomic wave-functions

Formalism for DM-atom scattering



$$\frac{d\mathcal{R}_{\text{ion}}^{n\ell}}{d \ln E_e} = \frac{n_x}{128\pi m_x^2 m_e^2} \int dq q \int \frac{d^3v}{v} f_x(v) \Theta(v - v_{\min}) \overline{\left| \mathcal{M}_{\text{ion}}^{n\ell} \right|^2}$$

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Free electron-DM amplitude

Atomic wave-functions

$$\mathcal{M}(\mathbf{q}, \mathbf{v}_{\text{el}}^\perp) \rightarrow \mathcal{M}(\mathbf{q}) \quad \rightarrow \quad \mathcal{M}_{1 \rightarrow 2} = \mathcal{M}(q = \alpha m_e) \frac{\mathcal{M}(q)}{\mathcal{M}(q = \alpha m_e)} \int \frac{d^3k}{(2\pi)^3} \tilde{\psi}_2^*(\mathbf{k} + \mathbf{q}) \tilde{\psi}_1(\mathbf{k})$$

$$\frac{d\mathcal{R}_{\text{ion}}^{n\ell}}{d \ln E_e} = \frac{n_\chi \bar{\sigma}_e}{8\mu_{\chi e}} \times \int q dq \eta(v_{\min}) |f_{\text{ion}}^{n\ell}(q)|^2 |F_{\text{DM}}(q)|^2$$

J. Kopp et al., 0907.3159,

R. Essig, et al., 1108.5383,

R. Essig, et al., 1206.2644

$$\eta(v_{\min}) \equiv \int d^3\vec{v} f_\chi(\vec{v}) \frac{1}{v} \Theta(v - v_{\min})$$

$$|f_{\text{ion}}^{n\ell}(q)|^2 \propto \sum_{l'm'm} |f_{1 \rightarrow 2}(\mathbf{q})|^2$$

How about the general case?

Based on the EFT approach

$$v_{\text{DM}} \sim 10^{-3}, v_e \sim \alpha_{\text{em}} \sim 10^{-2}$$



Non-relativistic (NR) EFT framework

NR operators	Power counting	DM type		
		scalar	fermion	vector
$\mathcal{O}_1 = \mathbb{1}_x \mathbb{1}_e$	1	✓	✓	✓
$\mathcal{O}_3 = \mathbb{1}_x \left(\frac{i\mathbf{q}}{m_e} \times \mathbf{v}_{\text{el}}^\perp \right) \cdot \mathbf{S}_e$	qv	✓	✓	✓
$\mathcal{O}_4 = \mathbf{S}_x \cdot \mathbf{S}_e$	1	—	✓	✓
$\mathcal{O}_5 = \mathbf{S}_x \cdot \left(\frac{i\mathbf{q}}{m_e} \times \mathbf{v}_{\text{el}}^\perp \right) \mathbb{1}_e$	qv	—	✓	✓
$\mathcal{O}_6 = \left(\mathbf{S}_x \cdot \frac{\mathbf{q}}{m_e} \right) \left(\frac{\mathbf{q}}{m_e} \cdot \mathbf{S}_e \right)$	q^2	—	✓	✓
$\mathcal{O}_7 = \mathbb{1}_x \mathbf{v}_{\text{el}}^\perp \cdot \mathbf{S}_e$	v	✓	✓	✓
$\mathcal{O}_8 = \mathbf{S}_x \cdot \mathbf{v}_{\text{el}}^\perp \mathbb{1}_e$	v	—	✓	✓
$\mathcal{O}_9 = -\mathbf{S}_x \cdot \left(\frac{i\mathbf{q}}{m_e} \times \mathbf{S}_e \right)$	q	—	✓	✓
$\mathcal{O}_{10} = \mathbb{1}_x \frac{i\mathbf{q}}{m_e} \cdot \mathbf{S}_e$	q	✓	✓	✓
$\mathcal{O}_{11} = \mathbf{S}_x \cdot \frac{i\mathbf{q}}{m_e} \mathbb{1}_e$	q	—	✓	✓
$\mathcal{O}_{12} = -\mathbf{S}_x \cdot (\mathbf{v}_{\text{el}}^\perp \times \mathbf{S}_e)$	v	—	✓	✓
$\mathcal{O}_{13} = (\mathbf{S}_x \cdot \mathbf{v}_{\text{el}}^\perp) \left(\frac{\mathbf{q}}{m_e} \cdot \mathbf{S}_e \right)$	qv	—	✓	✓
$\mathcal{O}_{14} = (\mathbf{S}_x \cdot \frac{\mathbf{q}}{m_e})(\mathbf{v}_{\text{el}}^\perp \cdot \mathbf{S}_e)$	qv	—	✓	✓
$\mathcal{O}_{15} = \mathbf{S}_x \cdot \frac{\mathbf{q}}{m_e} \left[\frac{\mathbf{q}}{m_e} \cdot (\mathbf{v}_{\text{el}}^\perp \times \mathbf{S}_e) \right]$	q^2v	—	✓	✓

$$\{\mathbb{1}_e, \mathbf{S}_e\} \otimes \{\mathbb{1}_x, \mathbf{S}_x, \tilde{\mathbf{S}}_x\} \otimes \{i\mathbf{q}, \mathbf{v}_{\text{el}}^\perp\}$$

$$\mathbf{v}_{\text{el}}^\perp = \mathbf{v} - \mathbf{q}/(2\mu_{xe}) - \mathbf{k}/m_e$$

$\mathcal{O}_{17} = \frac{i\mathbf{q}}{m_e} \cdot \tilde{\mathbf{S}}_x \cdot \mathbf{v}_{\text{el}}^\perp \mathbb{1}_e$	qv	—	—	✓
$\mathcal{O}_{18} = \frac{i\mathbf{q}}{m_e} \cdot \tilde{\mathbf{S}}_x \cdot \mathbf{S}_e$	q	—	—	✓
$\mathcal{O}_{19} = \frac{\mathbf{q}}{m_e} \cdot \tilde{\mathbf{S}}_x \cdot \frac{\mathbf{q}}{m_e} \mathbb{1}_e$	q^2	—	—	✓
$\mathcal{O}_{20} = -\frac{\mathbf{q}}{m_e} \cdot \tilde{\mathbf{S}}_x \cdot \left(\frac{\mathbf{q}}{m_e} \times \mathbf{S}_e \right)$	q^2	—	—	✓
$\mathcal{O}_{21} = \mathbf{v}_{\text{el}}^\perp \cdot \tilde{\mathbf{S}}_x \cdot \mathbf{S}_e$	v	—	—	✓
$\mathcal{O}_{22} = \left(\frac{i\mathbf{q}}{m_e} \times \mathbf{v}_{\text{el}}^\perp \right) \cdot \tilde{\mathbf{S}}_x \cdot \mathbf{S}_e + \mathbf{v}_{\text{el}}^\perp \cdot \tilde{\mathbf{S}}_x \cdot \left(\frac{i\mathbf{q}}{m_e} \times \mathbf{S}_e \right)$	qv	—	—	✓
$\mathcal{O}_{23} = -\frac{i\mathbf{q}}{m_e} \cdot \tilde{\mathbf{S}}_x \cdot (\mathbf{v}_{\text{el}}^\perp \times \mathbf{S}_e)$	qv	—	—	✓
$\mathcal{O}_{24} = \frac{\mathbf{q}}{m_e} \cdot \tilde{\mathbf{S}}_x \cdot \left(\frac{\mathbf{q}}{m_e} \times \mathbf{v}_{\text{el}}^\perp \right)$	q^2v	—	—	✓
$\mathcal{O}_{25} = \left(\frac{\mathbf{q}}{m_e} \cdot \tilde{\mathbf{S}}_x \cdot \mathbf{v}_{\text{el}}^\perp \right) \left(\frac{\mathbf{q}}{m_e} \cdot \mathbf{S}_e \right)$	q^2v	—	—	✓
$\mathcal{O}_{26} = \left(\frac{\mathbf{q}}{m_e} \cdot \tilde{\mathbf{S}}_x \cdot \frac{\mathbf{q}}{m_e} \right) (\mathbf{v}_{\text{el}}^\perp \cdot \mathbf{S}_e)$	q^2v	—	—	✓

How about the general case?

Catena et al., Phys. Rev. Res. 2, 033195 (2020) (105 citations)

$$\mathcal{M}(\mathbf{q}, \mathbf{v}_{\text{el}}^\perp) = \mathcal{M}(\mathbf{q}, \mathbf{v}_{\text{el}}^\perp) |_{\mathbf{k}=0} + \mathbf{k} \cdot \nabla_{\mathbf{k}} \mathcal{M}(\mathbf{q}, \mathbf{v}_{\text{el}}^\perp)$$

→ $\mathcal{M}_{1 \rightarrow 2} = f_{1 \rightarrow 2}(\mathbf{q}) \mathcal{M}(\mathbf{q}, \mathbf{v}_{\text{el}}^\perp) |_{\mathbf{k}=0} + m_e \mathbf{f}_{1 \rightarrow 2}(\mathbf{q}) \cdot \nabla_{\mathbf{k}} \mathcal{M}(\mathbf{q}, \mathbf{v}_{\text{el}}^\perp) |_{\mathbf{k}=0}$

$$f_{1 \rightarrow 2}(\mathbf{q}) = \int \frac{d^3k}{(2\pi)^3} \tilde{\psi}_2^*(\mathbf{k} + \mathbf{q}) \tilde{\psi}_1(\mathbf{k})$$

$$\mathbf{f}_{1 \rightarrow 2}(\mathbf{q}) = \int \frac{d^3k}{(2\pi)^3} \tilde{\psi}_2^*(\mathbf{k} + \mathbf{q}) \frac{\mathbf{k}}{m_e} \tilde{\psi}_1(\mathbf{k})$$

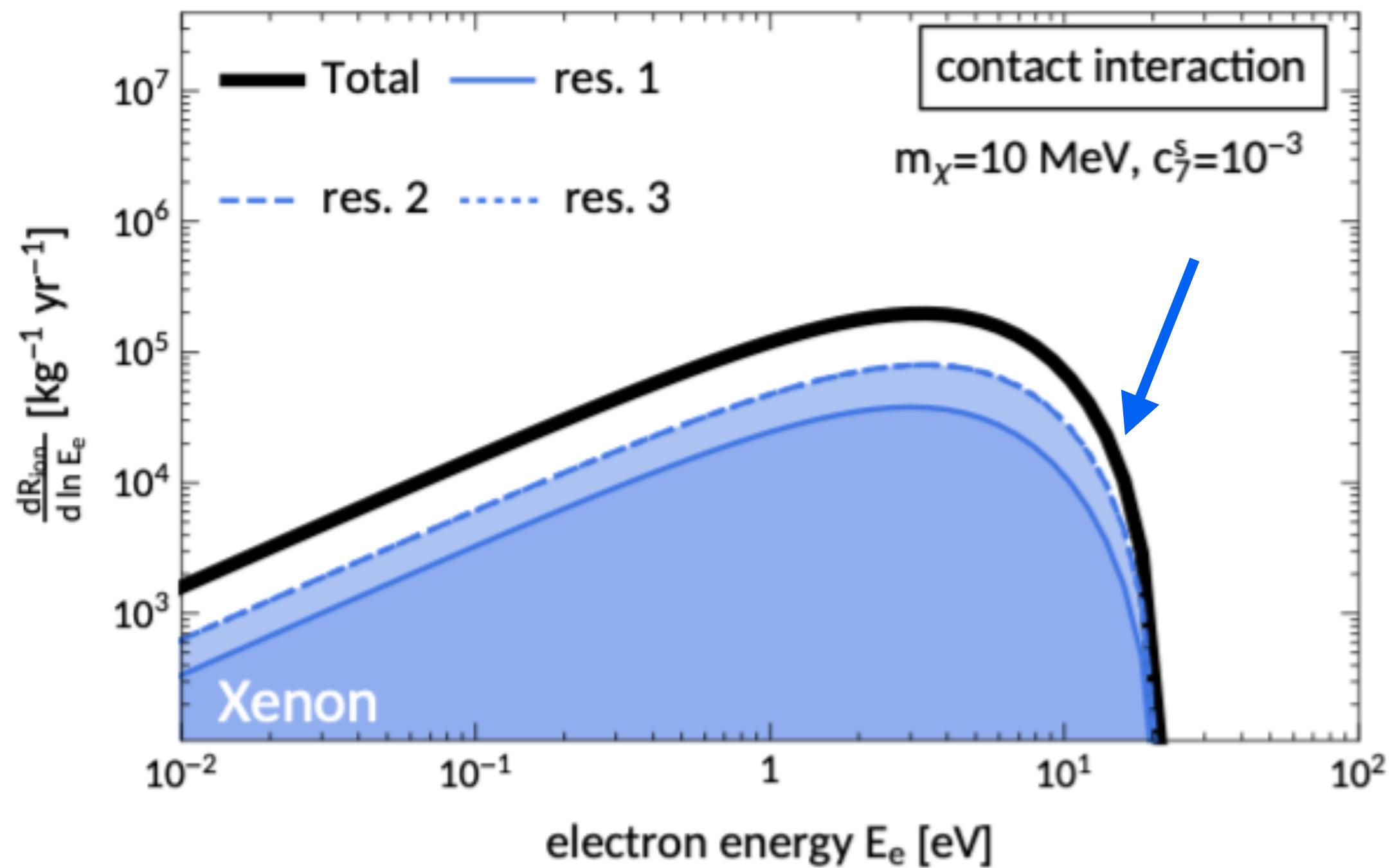
$$\overline{\left| \mathcal{M}_{\text{ion}}^{n\ell} \right|^2} = \sum_{i=1}^4 R_i^{n\ell}(q, \nu) W_i^{n\ell}(k', q)$$

1 usual $|f_{\text{ion}}^{n\ell}(q)|^2 (W_1)$ + 3 new atomic response functions ($W_{2,3,4}$)

Comparison

$$\mathcal{O}_7 = \mathbb{1}_x \mathbf{v}_{\text{el}}^\perp \cdot \mathbf{S}_e$$

Contributions from W_2 and W_3 dominate



Catena et al., 1912.08204

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$$f_{1 \rightarrow 2}(\mathbf{q}) = \int \frac{d^3k}{(2\pi)^3} \tilde{\psi}_2^*(\mathbf{k} + \mathbf{q}) \tilde{\psi}_1(\mathbf{k})$$

$$\mathbf{f}_{1 \rightarrow 2}(\mathbf{q}) = \int \frac{d^3k}{(2\pi)^3} \tilde{\psi}_2^*(\mathbf{k} + \mathbf{q}) \frac{\mathbf{k}}{m_e} \tilde{\psi}_1(\mathbf{k})$$

A minus sign is missed

$$\mathbf{f}_{1 \rightarrow 2}(\mathbf{q}) = \int d^3\mathbf{r} \psi_2^*(\mathbf{r}) e^{i\mathbf{q} \cdot \mathbf{r}} \frac{-i\nabla}{m_e} \psi_1(\mathbf{r})$$

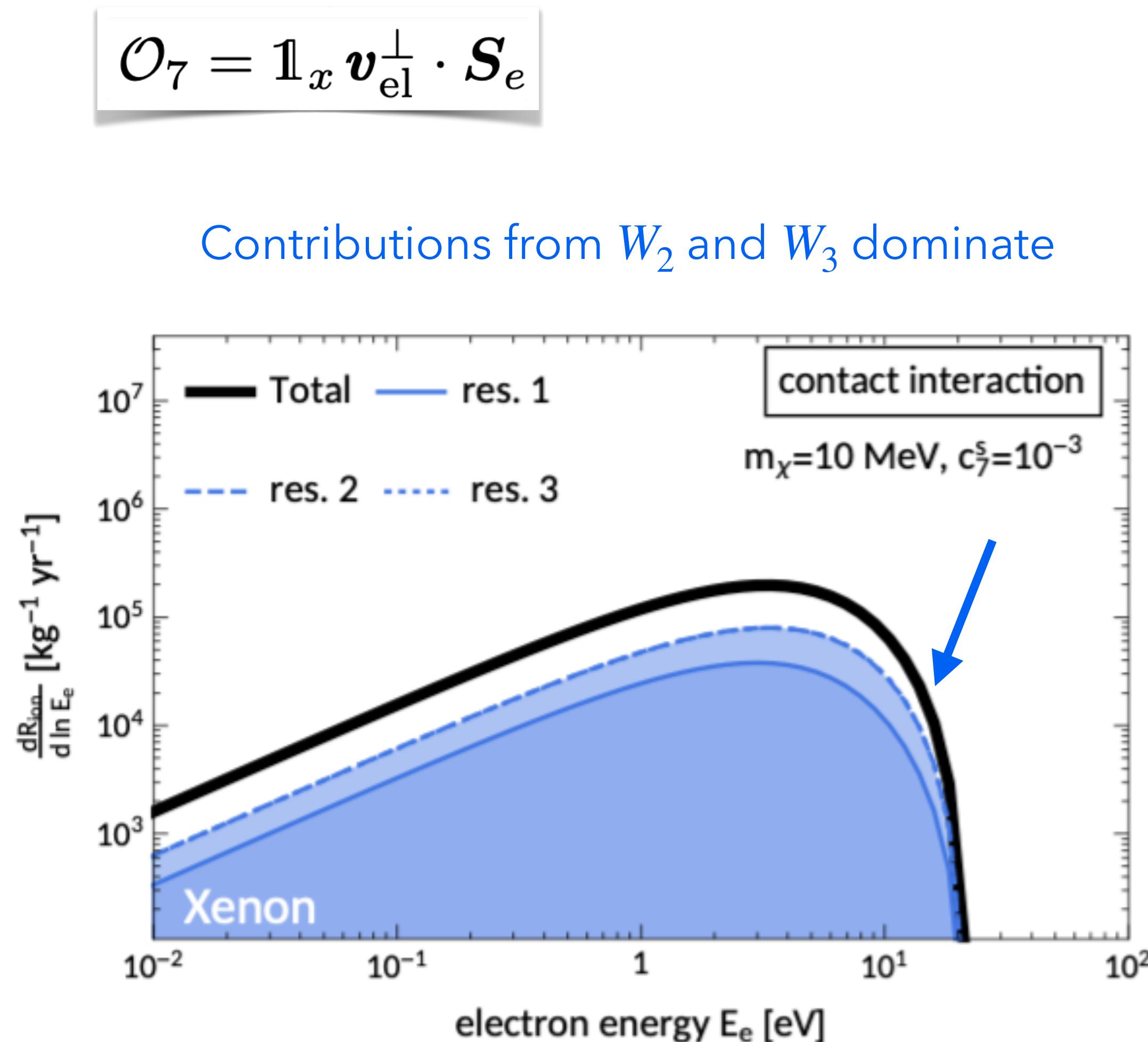
$$\overline{\left| \mathcal{M}_{\text{ion}}^{nl} \right|^2} = \sum_{i=1}^4 R_i^{nl}(q, \nu) W_i^{nl}(k', q)$$

$$W_2 \rightarrow -W_2$$

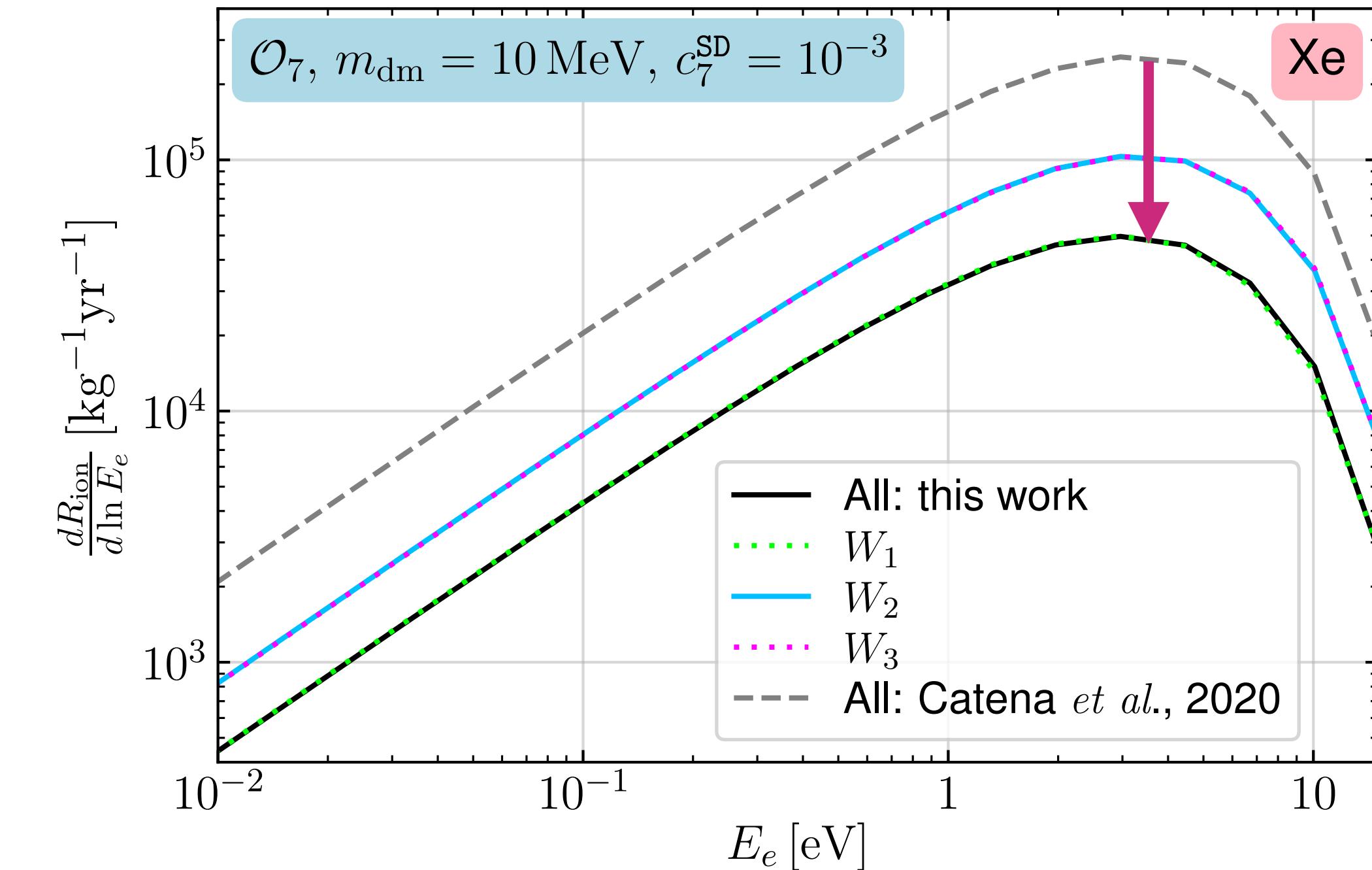
1 usual $|f_{\text{ion}}^{nl}(q)|^2 (W_1)$ + 3 new atomic response functions ($W_{2,3,4}$)

Comparison

arXiv: 2405. 04855, JL, YL, XM, and HW



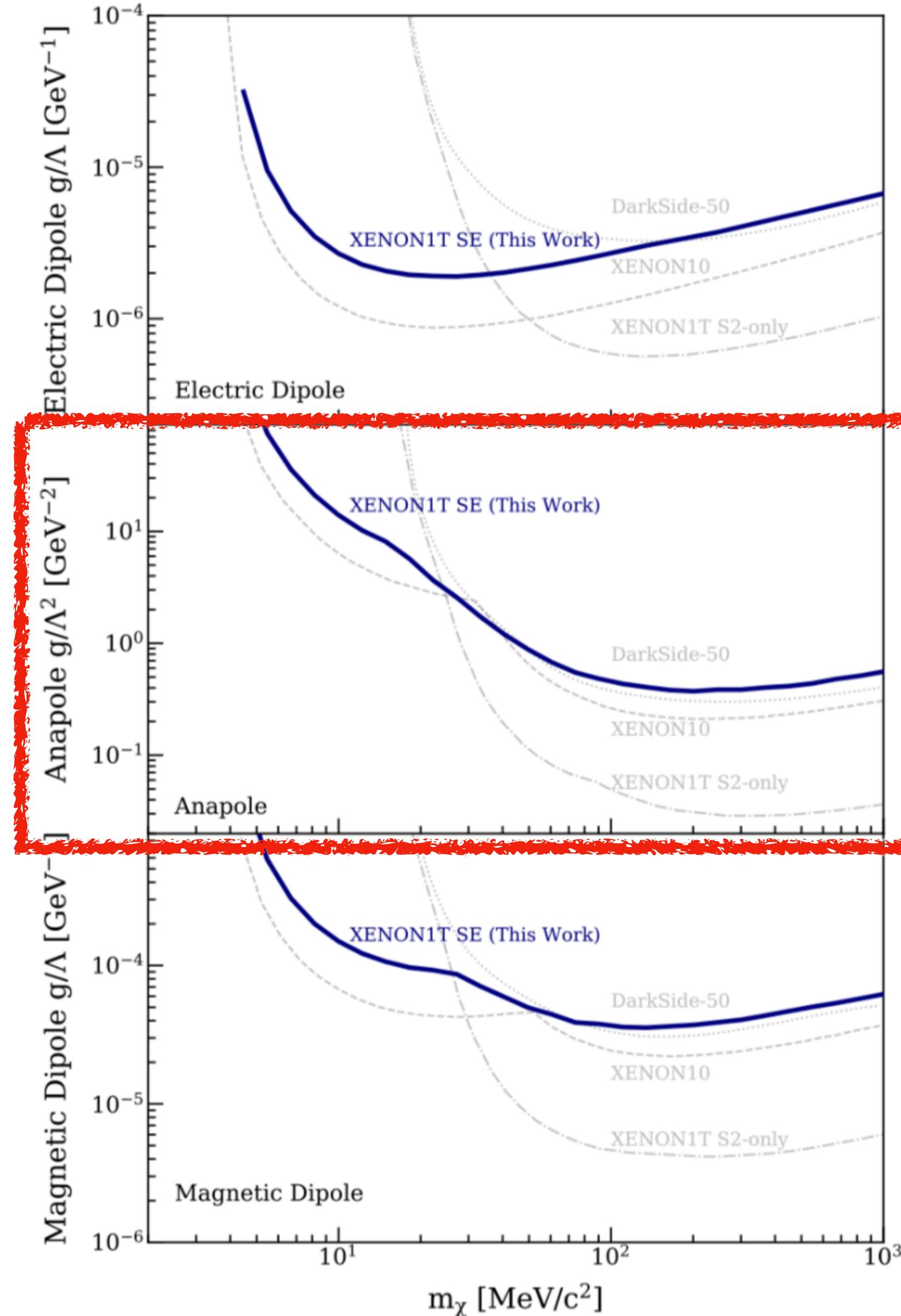
Catena et al., 1912.08204



- The sign difference cause a factor of 5 difference in the differential rate
- Constructive interference between W_2 and W_3 , leaving W_1 contribution dominant

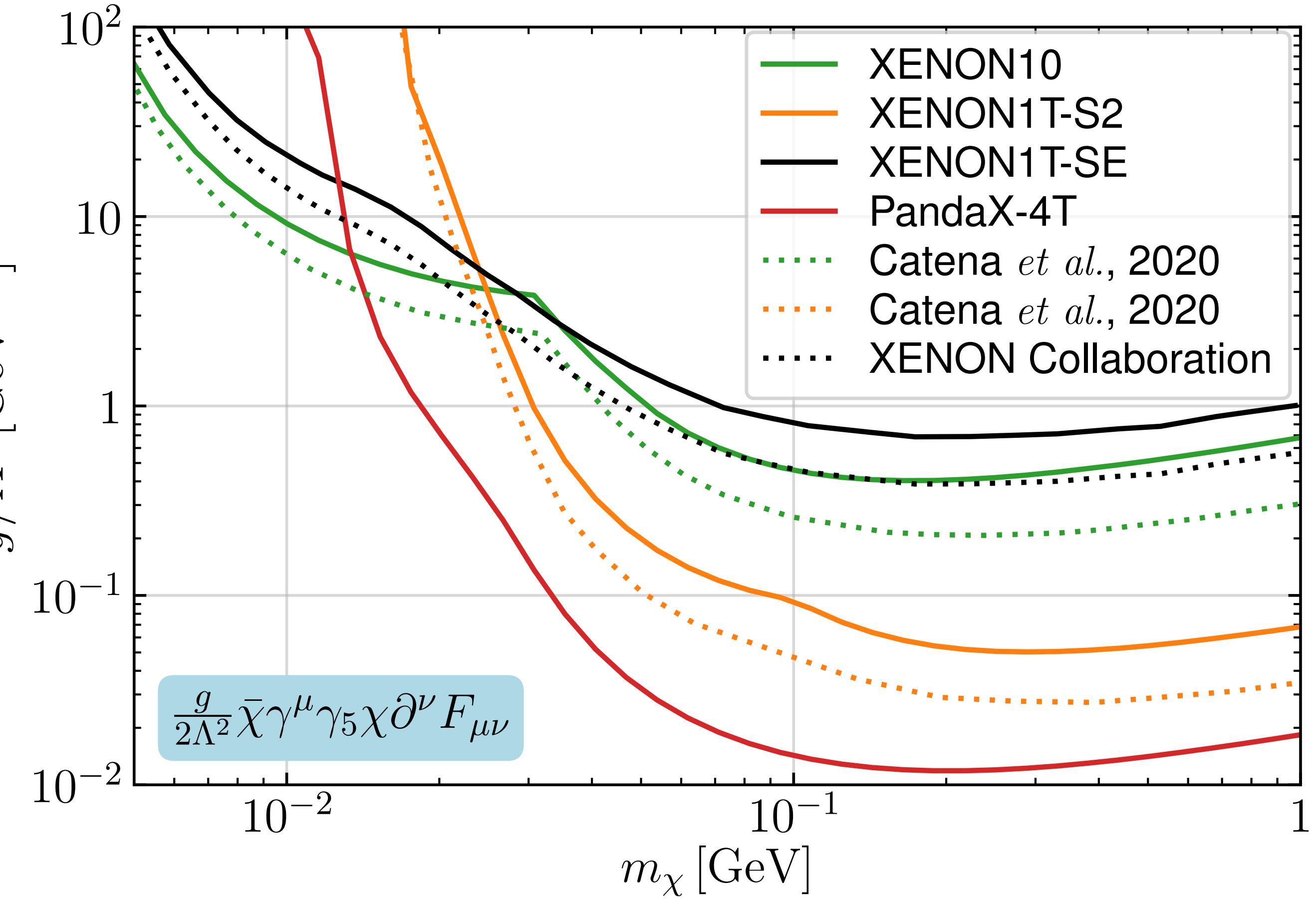
XENON1T SE constraints on EM form factors of DM

XENON Collaboration, 2112.12116



$$a_\chi (\bar{\chi} \gamma^\mu \gamma_5 \chi) \partial^\nu F_{\mu\nu} \rightarrow 8a_\chi e m_\chi m_e (\mathcal{O}_8 - \mathcal{O}_9)$$

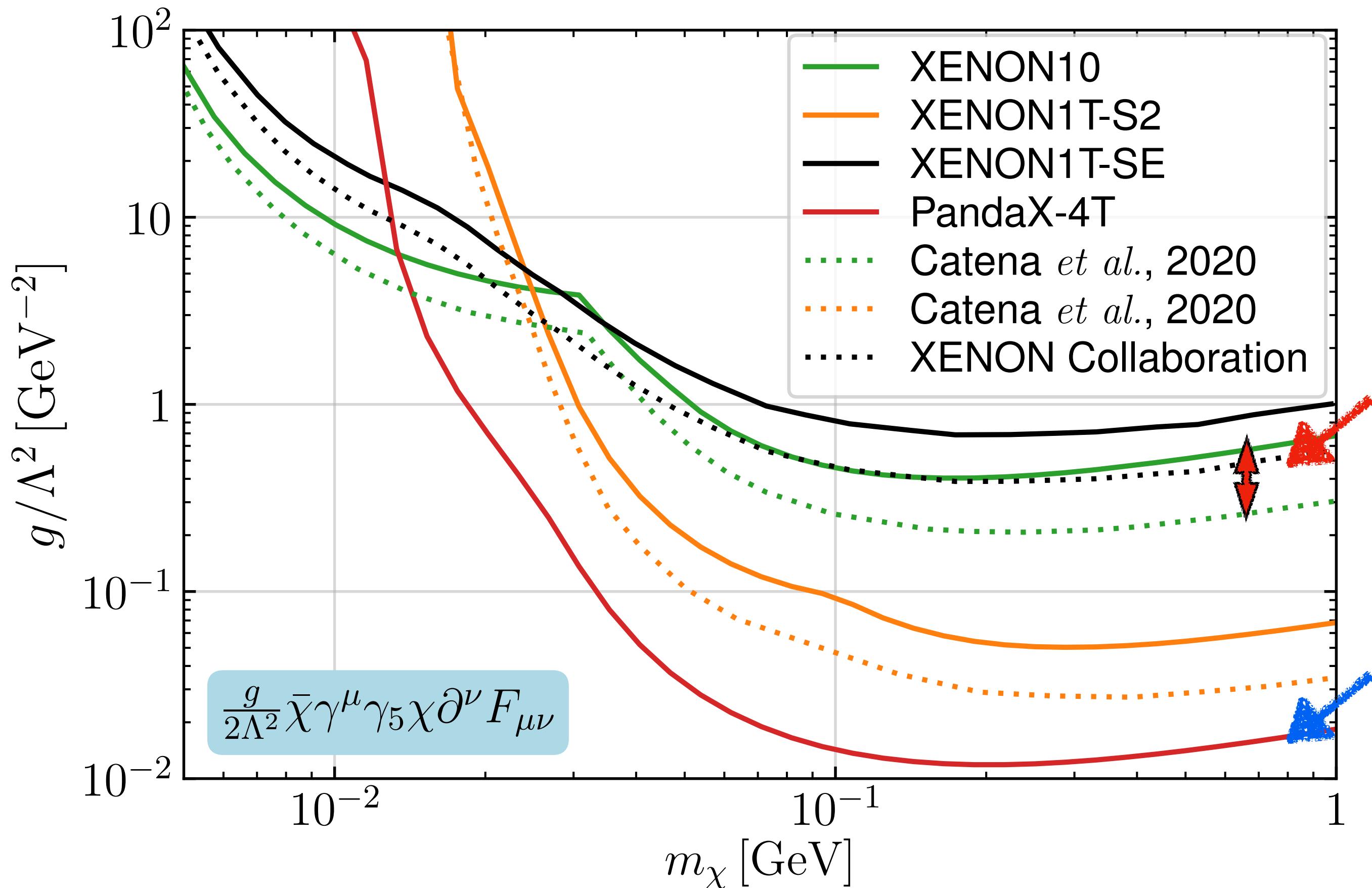
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XENON1T SE constraints on EM form factors of DM

$$a_\chi (\bar{\chi} \gamma^\mu \gamma_5 \chi) \partial^\nu F_{\mu\nu} \rightarrow 8a_\chi e m_\chi m_e (\mathcal{O}_8 - \mathcal{O}_9)$$

arXiv: 2405. 04855, JL, YL, XM, and HW



- The constraints are weakened by a factor of 2 if the sign of W2 is corrected
- PandaX-4T set the most stringent constraint when $m_\chi \gtrsim 20 \text{ MeV}$

Our formalism

JHEP 07 (2024) 279, JL, YL, XM, and HW

Based on the NR operator structure

$$\mathcal{M}(\mathbf{q}, \mathbf{v}_{\text{el}}^\perp) = \mathcal{M}(\mathbf{q}, 0) + \mathbf{v}_{\text{el}}^\perp \cdot \nabla_{\mathbf{v}_{\text{el}}^\perp} \mathcal{M}(\mathbf{q}, \mathbf{v}_{\text{el}}^\perp)$$

$$\rightarrow \mathcal{M}_{1 \rightarrow 2} = f_S(\mathbf{q}) \mathcal{M}_S + \mathbf{f}_V(\mathbf{q}) \cdot \mathcal{M}_V$$

$$\mathbf{v}_{\text{el}}^\perp = \mathbf{v} - \mathbf{q}/(2\mu_{xe}) - \mathbf{k}/m_e$$
$$\mathbf{v}_0^\perp \equiv \mathbf{v} - \mathbf{q}/(2\mu_{xe})$$

$$f_S(\mathbf{q}) = f_{1 \rightarrow 2}(\mathbf{q})$$

$$\mathbf{f}_V(\mathbf{q}) = \mathbf{v}_0^\perp f_{1 \rightarrow 2}(\mathbf{q}) - \mathbf{f}_{1 \rightarrow 2}(\mathbf{q})$$

$$\overline{\left| \mathcal{M}_{\text{ion}}^{n\ell} \right|^2} = a_0 \widetilde{W}_0 + a_1 \widetilde{W}_1 + a_2 \widetilde{W}_2$$

We need only three generalized atomic response functions

$$\widetilde{W}_0 = W_1,$$

$$\widetilde{W}_1 = |\mathbf{v}_0^\perp|^2 W_1 - 2 \frac{y_e}{x_e} W_2 + W_3,$$

$$\widetilde{W}_2 = \frac{y_e^2}{x_e} W_1 - 2 \frac{y_e}{x_e} W_2 + \frac{1}{x_e} W_4.$$

Our formalism

JHEP 07 (2024) 279, JL, YL, XM, and HW

Based on the NR operator structure

$$\mathcal{M}(\mathbf{q}, \mathbf{v}_{\text{el}}^\perp) = \mathcal{M}(\mathbf{q}, 0) + \mathbf{v}_{\text{el}}^\perp$$



$$\mathcal{M}_{1 \rightarrow 2} = f_S(\mathbf{q}) \mathcal{M}_S +$$

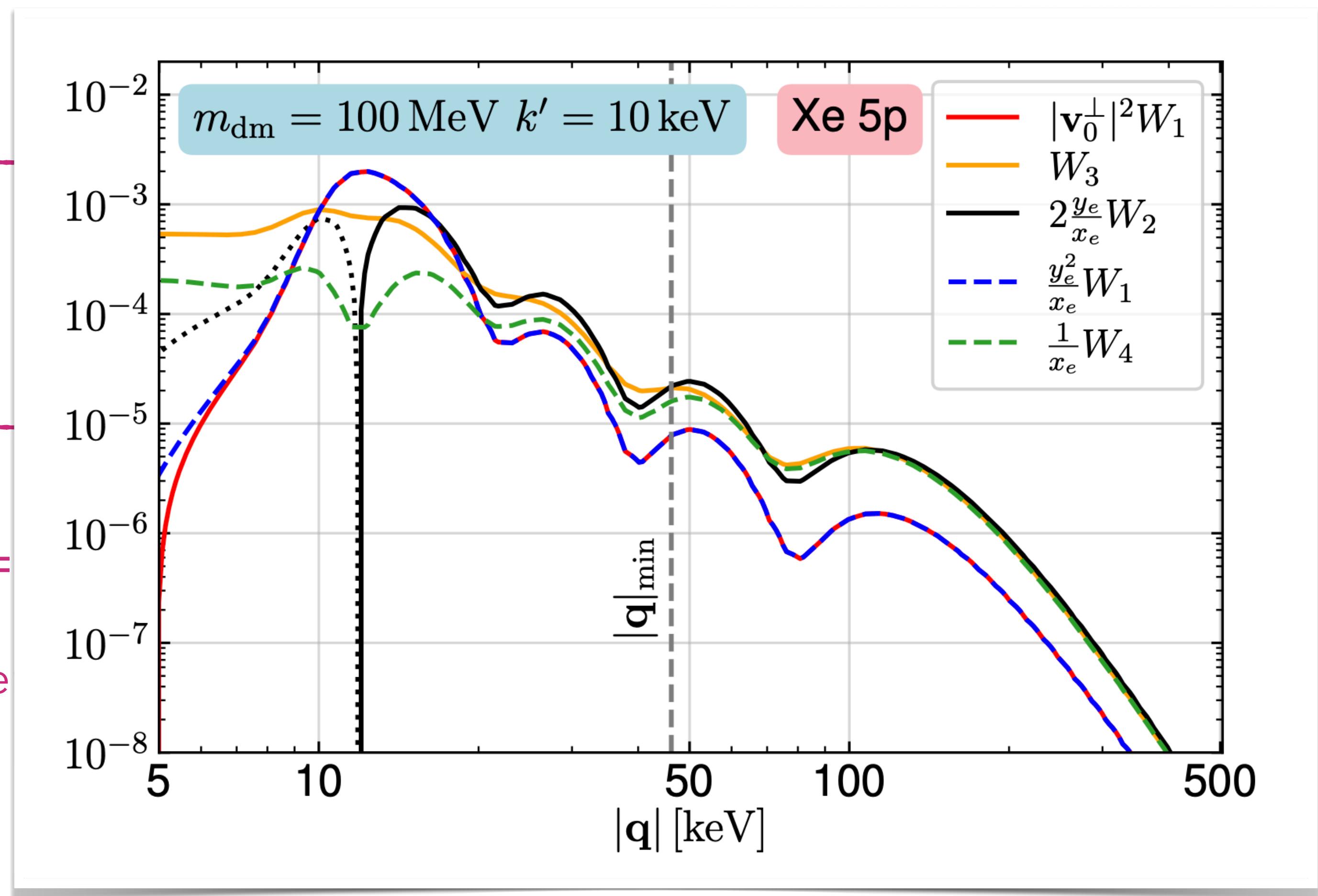
$$|\mathcal{M}_{\text{ion}}^{nl}|^2 =$$

We need only three ge

$$\tilde{W}_0 = W_1,$$

$$\tilde{W}_1 = |\mathbf{v}_0^\perp|^2 W_1 - 2 \frac{y_e}{x_e} W_2 + W_3,$$

$$\tilde{W}_2 = \frac{y_e^2}{x_e} W_1 - 2 \frac{y_e}{x_e} W_2 + \frac{1}{x_e} W_4.$$



The minus sign leads to a strong cancellation among W_2 and $W_{3,4}$

Our formalism

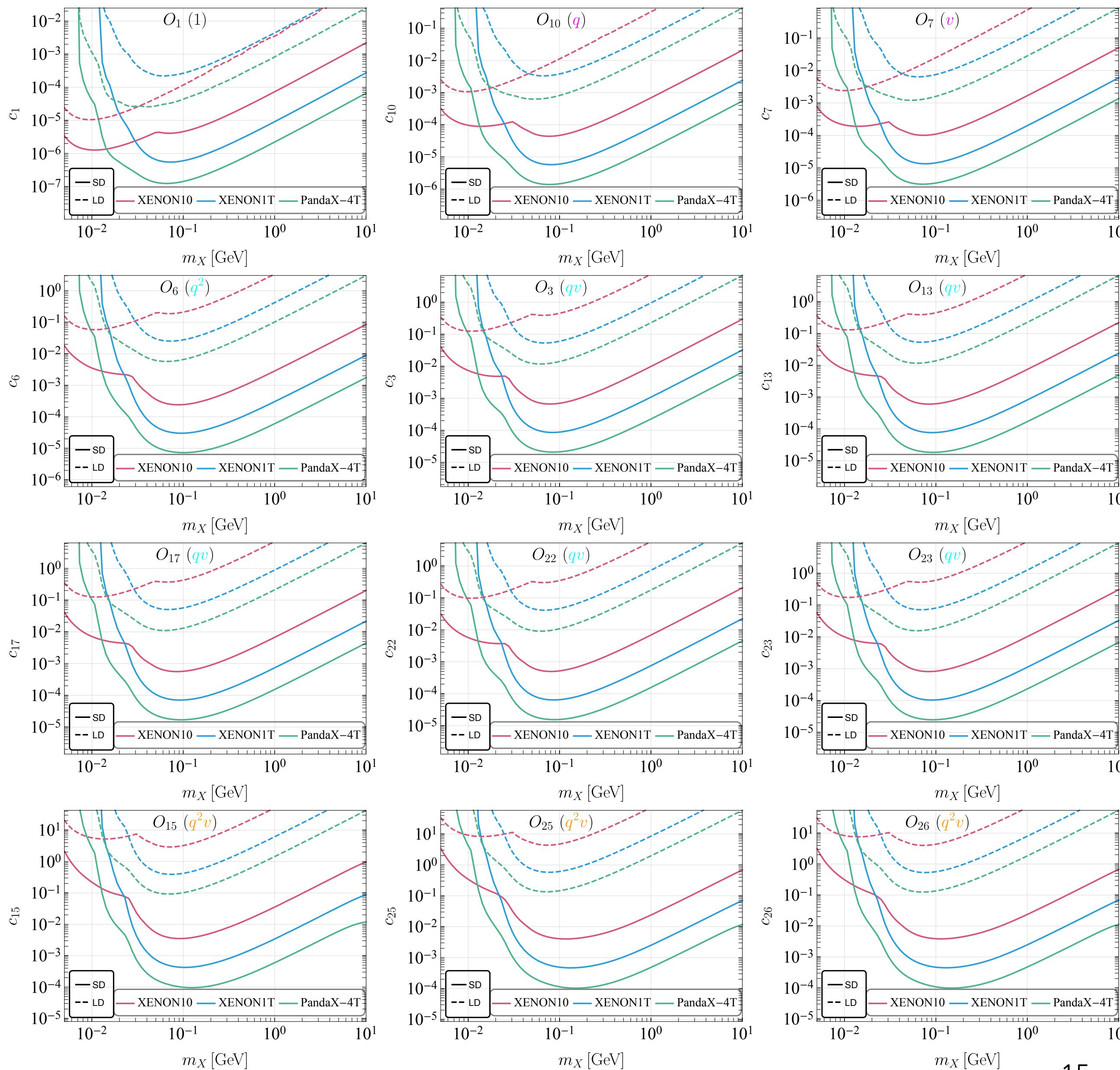
$$\overline{\left| \mathcal{M}_{\text{ion}}^{n\ell} \right|^2} = a_0 \widetilde{W}_0 + a_1 \widetilde{W}_1 + a_2 \widetilde{W}_2$$

- a_0 and $a_{1,2}$ involve different NR operators
- Clear power counting for q
- Do not contain any atomic properties

$$x_e \equiv q^2/m_e^2$$

Type	DM response functions
Scalar DM	$a_0 = c_1 ^2 + \frac{1}{4} c_{10} ^2 \textcolor{blue}{x}_e$ $a_1 = \frac{1}{4} c_7 ^2 + \frac{1}{4} c_3 ^2 \textcolor{blue}{x}_e$ $a_2 = -\frac{1}{4} c_3 ^2 \textcolor{blue}{x}_e$
Fermion DM	$a_0 = c_1 ^2 + \frac{3}{16} c_4 ^2 + \left(\frac{1}{8} c_9 ^2 + \frac{1}{4} c_{10} ^2 + \frac{1}{4} c_{11} ^2 + \frac{1}{8}\Re[c_4 c_6^*] \right) \textcolor{blue}{x}_e + \frac{1}{16} c_6 ^2 \textcolor{blue}{x}_e^2$ $a_1 = \frac{1}{4} c_7 ^2 + \frac{1}{4} c_8 ^2 + \frac{1}{8} c_{12} ^2 + \left(\frac{1}{4} c_3 ^2 + \frac{1}{4} c_5 ^2 + \frac{1}{16} c_{13} ^2 + \frac{1}{16} c_{14} ^2 - \frac{1}{8}\Re[c_{12} c_{15}^*] \right) \textcolor{blue}{x}_e + \frac{1}{16} c_{15} ^2 \textcolor{blue}{x}_e^2$ $a_2 = -\left(\frac{1}{4} c_3 ^2 + \frac{1}{4} c_5 ^2 - \frac{1}{8}\Re[c_{12} c_{15}^*] - \frac{1}{8}\Re[c_{13} c_{14}^*] \right) \textcolor{blue}{x}_e - \frac{1}{16} c_{15} ^2 \textcolor{blue}{x}_e^2$
Vector DM	$a_0 = c_1 ^2 + \frac{1}{2} c_4 ^2 + \left(\frac{1}{3} c_9 ^2 + \frac{1}{4} c_{10} ^2 + \frac{2}{3} c_{11} ^2 + \frac{5}{36} c_{18} ^2 + \frac{1}{3}\Re[c_4 c_6^*] \right) \textcolor{blue}{x}_e + \left(\frac{1}{6} c_6 ^2 + \frac{2}{9} c_{19} ^2 + \frac{1}{12} c_{20} ^2 \right) \textcolor{blue}{x}_e^2$ $a_1 = \frac{1}{4} c_7 ^2 + \frac{2}{3} c_8 ^2 + \frac{1}{3} c_{12} ^2 + \frac{5}{36} c_{21} ^2 + \left(\frac{1}{4} c_3 ^2 + \frac{2}{3} c_5 ^2 + \frac{1}{6} c_{13} ^2 + \frac{1}{6} c_{14} ^2 + \frac{1}{6} c_{17} ^2 + \frac{3}{8} c_{22} ^2 + \frac{7}{72} c_{23} ^2 - \frac{1}{3}\Re[c_{12} c_{15}^*] + \frac{1}{12}\Re[c_{21} c_{25}^*] - \frac{1}{18}\Re[c_{21} c_{26}^*] + \frac{1}{12}\Re[c_{22} c_{23}^*] \right) \textcolor{blue}{x}_e + \left(\frac{1}{6} c_{15} ^2 + \frac{1}{6} c_{24} ^2 + \frac{1}{24} c_{25} ^2 + \frac{1}{18} c_{26} ^2 \right) \textcolor{blue}{x}_e^2$ $a_2 = -\left(\frac{1}{4} c_3 ^2 + \frac{2}{3} c_5 ^2 - \frac{1}{18} c_{17} ^2 + \frac{7}{24} c_{22} ^2 + \frac{1}{72} c_{23} ^2 - \frac{1}{3}\Re[c_{12} c_{15}^*] - \frac{1}{3}\Re[c_{13} c_{14}^*] - \frac{1}{36}\Re[c_{21} c_{25}^*] - \frac{1}{6}\Re[c_{21} c_{26}^*] + \frac{1}{4}\Re[c_{22} c_{23}^*] \right) \textcolor{blue}{x}_e - \left(\frac{1}{6} c_{15} ^2 + \frac{1}{6} c_{24} ^2 - \frac{1}{72} c_{25} ^2 - \frac{1}{9}\Re[c_{25} c_{26}^*] \right) \textcolor{blue}{x}_e^2$

Constraints on the NR operators



$$M_i = (c_i^s + c_i^l \frac{q_{\text{ref}}^2}{q^2}) < O_i >$$

$$q_{\text{ref}} = \alpha m_e$$

- The constraints follows consistently with the power counting of q and/or v
- The suppression from v is comparable with that from q

PandaX-4T set the most stringent constraint when $m_X \gtrsim 20$ MeV

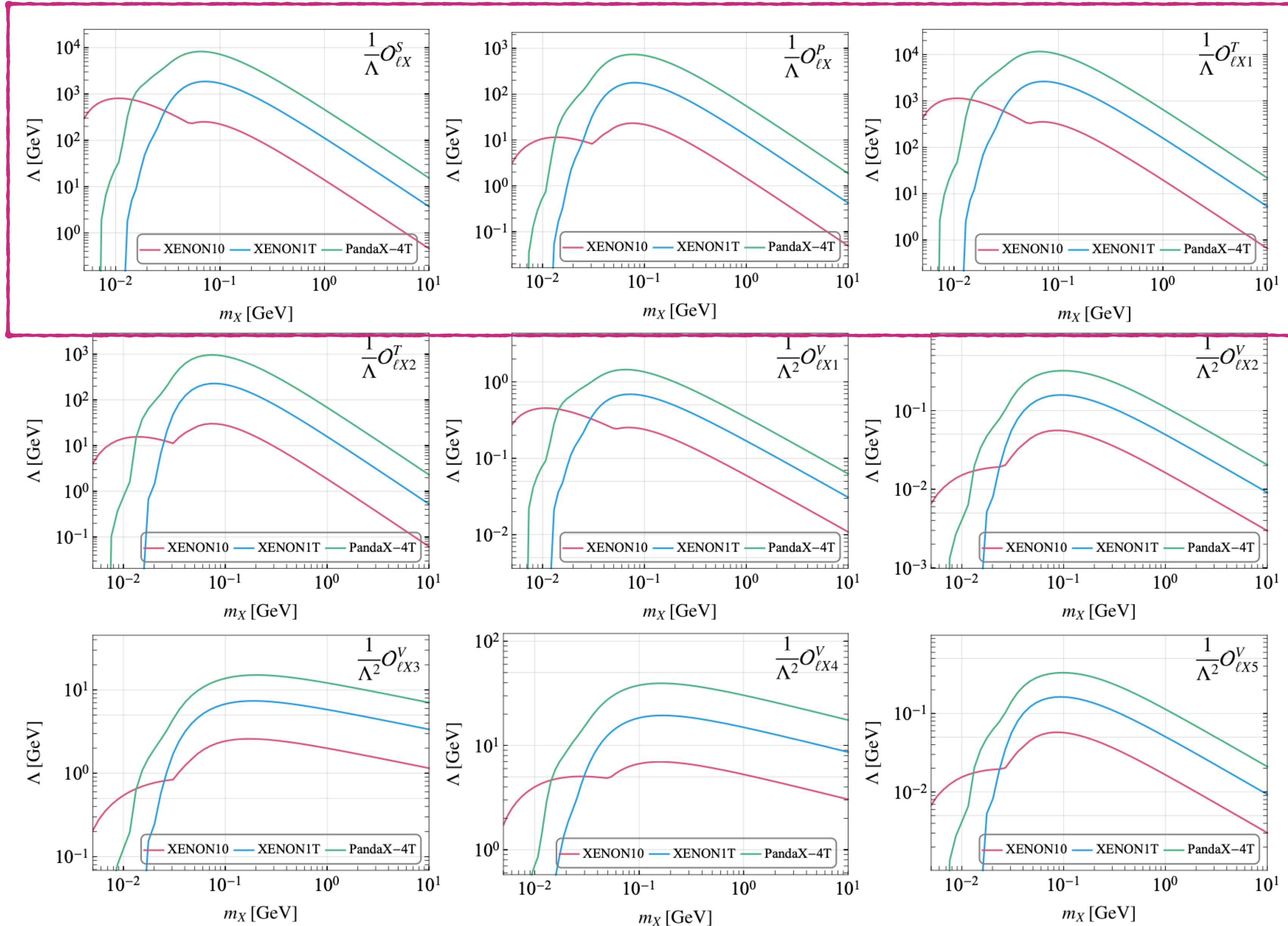
Constraints on the relativistic operators

All possible LO DM-electron and DM-photon operators

Dim	Relativistic operators	NR reduction
Scalar case		
dim-5	$\mathcal{O}_{\ell\phi}^S = (\bar{\ell}\ell)(\phi^\dagger\phi)$	$2m_e\mathcal{O}_1$
	$\mathcal{O}_{\ell\phi}^P = (\bar{\ell}i\gamma_5\ell)(\phi^\dagger\phi)$	$-2m_e\mathcal{O}_{10}$
dim-6	$\mathcal{O}_{\ell\phi}^V = (\bar{\ell}\gamma^\mu\ell)(\phi^\dagger i\overleftrightarrow{\partial}_\mu\phi) (\times)$	$4m_e m_\phi \mathcal{O}_1$
	$\mathcal{O}_{\ell\phi}^A = (\bar{\ell}\gamma^\mu\gamma_5\ell)(\phi^\dagger i\overleftrightarrow{\partial}_\mu\phi) (\times)$	$-8m_e m_\phi \mathcal{O}_7$
	$\mathcal{L}_\phi^Q = (\partial_\mu - iQ_\phi e A_\mu)\phi ^2 (\times)$	$-4Q_\phi e^2 \frac{m_e m_\phi}{q^2} \mathcal{O}_1$
	$\mathcal{L}_\phi^{\text{cr}} = b_\phi (\phi^\dagger i\overleftrightarrow{\partial}_\mu\phi) \partial^\nu F_{\mu\nu} (\times)$	$4b_\phi e m_e m_\phi \mathcal{O}_1$
Fermion case		
dim-6	$\mathcal{O}_{\ell\chi}^S = (\bar{\ell}\ell)(\bar{\chi}\chi)$	$4m_e m_\chi \mathcal{O}_1$
	$\mathcal{O}_{\ell\chi}^S = (\bar{\ell}\ell)(\bar{\chi}i\gamma_5\chi)$	$4m_e^2 \mathcal{O}_{11}$
	$\mathcal{O}_{\ell\chi}^P = (\bar{\ell}i\gamma_5\ell)(\bar{\chi}\chi)$	$-4m_e m_\chi \mathcal{O}_{10}$
	$\mathcal{O}_{\ell\chi}^P = (\bar{\ell}i\gamma_5\ell)(\bar{\chi}i\gamma_5\chi)$	$4m_e^2 \mathcal{O}_6$
	$\mathcal{O}_{\ell\chi}^V = (\bar{\ell}\gamma^\mu\ell)(\bar{\chi}\gamma_\mu\chi) (\times)$	$4m_e m_\chi \mathcal{O}_1$
	$\mathcal{O}_{\ell\chi}^V = (\bar{\ell}\gamma^\mu\ell)(\bar{\chi}\gamma_\mu\gamma_5\chi)$	$8m_e m_\chi (\mathcal{O}_8 - \mathcal{O}_9)$
	$\mathcal{O}_{\ell\chi}^A = (\bar{\ell}\gamma^\mu\gamma_5\ell)(\bar{\chi}\gamma_\mu\chi) (\times)$	$-8m_e (m_\chi \mathcal{O}_7 + m_e \mathcal{O}_9)$
	$\mathcal{O}_{\ell\chi}^A = (\bar{\ell}\gamma^\mu\gamma_5\ell)(\bar{\chi}\gamma_\mu\gamma_5\chi)$	$-16m_e m_\chi \mathcal{O}_4$
	$\mathcal{O}_{\ell\chi}^T = (\bar{\ell}\sigma^{\mu\nu}\ell)(\bar{\chi}\sigma_{\mu\nu}\chi) (\times)$	$32m_e m_\chi \mathcal{O}_4$
	$\mathcal{O}_{\ell\chi}^T = (\bar{\ell}\sigma^{\mu\nu}\ell)(\bar{\chi}i\sigma_{\mu\nu}\gamma_5\chi) (\times)$	$8m_e (m_e \mathcal{O}_{10} - m_\chi \mathcal{O}_{11} - 4m_\chi \mathcal{O}_{12})$
	$\mathcal{L}_\chi^Q = \bar{\chi}i\gamma^\mu(\partial_\mu - iQ_\chi e A_\mu)\chi (\times)$	$-4Q_\chi e^2 \frac{m_e m_\chi}{q^2} \mathcal{O}_1$
	$\mathcal{L}_\chi^{\text{mdm}} = \mu_\chi (\bar{\chi}\sigma^{\mu\nu}\chi) F_{\mu\nu} (\times)$	$4\mu_\chi e (m_e \mathcal{O}_1 + 4m_\chi \mathcal{O}_4 + \frac{4m_e^2 m_\chi}{q^2} (\mathcal{O}_5 - \mathcal{O}_6))$
	$\mathcal{L}_\chi^{\text{edm}} = d_\chi (\bar{\chi}i\sigma^{\mu\nu}\gamma_5\chi) F_{\mu\nu} (\times)$	$d_\chi e \frac{16m_e^2 m_\chi}{q^2} \mathcal{O}_{11}$
	$\mathcal{L}_\chi^{\text{cr}} = b_\chi (\bar{\chi}\gamma^\mu\chi) \partial^\nu F_{\mu\nu} (\times)$	$4b_\chi e m_e m_\chi \mathcal{O}_1$
	$\mathcal{L}_\chi^{\text{anap.}} = a_\chi (\bar{\chi}\gamma^\mu\gamma_5\chi) \partial^\nu F_{\mu\nu}$	$8a_\chi e m_e m_\chi (\mathcal{O}_8 - \mathcal{O}_9)$

Dim	Relativistic operators	NR reduction
Vector case A		
dim-5	$\mathcal{O}_{\ell X}^S = (\bar{\ell}\ell)(X_\mu^\dagger X^\mu)$	$-2m_e \mathcal{O}_1$
	$\mathcal{O}_{\ell X}^P = (\bar{\ell}i\gamma_5\ell)(X_\mu^\dagger X^\mu)$	$2m_e \mathcal{O}_{10}$
	$\mathcal{O}_{\ell X 1}^T = \frac{i}{2}(\bar{\ell}\sigma^{\mu\nu}\ell)(X_\mu^\dagger X_\nu - X_\nu^\dagger X_\mu), (\times)$	$-4m_e \mathcal{O}_4$
	$\mathcal{O}_{\ell X 2}^T = \frac{1}{2}(\bar{\ell}\sigma^{\mu\nu}\gamma_5\ell)(X_\mu^\dagger X_\nu - X_\nu^\dagger X_\mu), (\times)$	$-m_e (\mathcal{O}_{11} + 4\mathcal{O}_{12}) + 4\frac{m_e^2}{m_X} \left(\frac{1}{3} \mathcal{O}_{10} - \mathcal{O}_{18} \right)$
dim-6	$\mathcal{O}_{\ell X 1}^V = \frac{1}{2}[\bar{\ell}\gamma_\mu i\overleftrightarrow{D}_\nu\ell](X^{\mu\dagger} X^\nu + X^{\nu\dagger} X^\mu)$	$m_e^2 \mathcal{O}_1$
	$\mathcal{O}_{\ell X 2}^V = (\bar{\ell}\gamma_\mu\ell)\partial_\nu(X^{\mu\dagger} X^\nu + X^{\nu\dagger} X^\mu)$	$-4m_e^2 (\mathcal{O}_{17} + \mathcal{O}_{20}) + \frac{4}{3}m_e(i\mathbf{q} \cdot \mathbf{v}_{\text{el}}^\perp) \mathcal{O}_1$
	$\mathcal{O}_{\ell X 3}^V = (\bar{\ell}\gamma_\mu\ell)(X_\rho^\dagger \overleftrightarrow{\partial}_\nu X_\sigma) \epsilon^{\mu\nu\rho\sigma}$	$-4m_e m_X (\mathcal{O}_8 - \mathcal{O}_9)$
	$\mathcal{O}_{\ell X 4}^V = (\bar{\ell}\gamma^\mu\ell)(X_\nu^\dagger i\overleftrightarrow{\partial}_\mu X^\nu), (\times)$	$-4m_e m_X \mathcal{O}_1$
	$\mathcal{O}_{\ell X 5}^V = (\bar{\ell}\gamma_\mu\ell)i\partial_\nu(X^{\mu\dagger} X^\nu - X^{\nu\dagger} X^\mu), (\times)$	$2m_e^2 \left(\mathcal{O}_5 - \mathcal{O}_6 - \frac{m_e}{m_X} \mathcal{O}_{19} \right) + 2\mathbf{q}^2 \mathcal{O}_4 + \frac{2}{3} \frac{m_e}{m_X} \mathbf{q}^2 \mathcal{O}_1$
	$\mathcal{O}_{\ell X 6}^V = (\bar{\ell}\gamma_\mu\ell)i\partial_\nu(X_\rho^\dagger X_\sigma) \epsilon^{\mu\nu\rho\sigma}, (\times)$	$-2m_e^2 \mathcal{O}_{11}$
	$\mathcal{O}_{\ell X 1}^A = \frac{1}{2}[\bar{\ell}\gamma_\mu\gamma_5 i\overleftrightarrow{D}_\nu\ell](X^{\mu\dagger} X^\nu + X^{\nu\dagger} X^\mu)$	$-2m_e^2 \left(\frac{m_e}{m_X} \mathcal{O}_9 - 4\mathcal{O}_{21} + \frac{4}{3} \mathcal{O}_7 \right)$
	$\mathcal{O}_{\ell X 2}^A = (\bar{\ell}\gamma_\mu\gamma_5\ell)\partial_\nu(X^{\mu\dagger} X^\nu + X^{\nu\dagger} X^\mu)$	$-8m_e^2 \left(\frac{1}{3} \mathcal{O}_{10} - \mathcal{O}_{18} \right)$
	$\mathcal{O}_{\ell X 3}^A = (\bar{\ell}\gamma_\mu\gamma_5\ell)(X_\rho^\dagger \overleftrightarrow{\partial}_\nu X_\sigma) \epsilon^{\mu\nu\rho\sigma}$	$8m_e m_X \mathcal{O}_4$
	$\mathcal{O}_{\ell X 4}^A = (\bar{\ell}\gamma^\mu\gamma_5\ell)(X_\nu^\dagger i\overleftrightarrow{\partial}_\mu X^\nu)$	$8m_e m_X \mathcal{O}_7$
dim-6	$\mathcal{O}_{\ell X 5}^A = (\bar{\ell}\gamma_\mu\gamma_5\ell)i\partial_\nu(X^{\mu\dagger} X^\nu - X^{\nu\dagger} X^\mu), (\times)$	$4m_e^2 \mathcal{O}_9$
	$\mathcal{O}_{\ell X 6}^A = (\bar{\ell}\gamma_\mu\gamma_5\ell)i\partial_\nu(X_\rho^\dagger X_\sigma) \epsilon^{\mu\nu\rho\sigma}, (\times)$	$4m_e^2 \left(\mathcal{O}_{14} - \frac{m_e}{m_X} \mathcal{O}_{20} \right)$
	$\mathcal{L}_{\kappa_\Lambda} = i\frac{\kappa_\Lambda}{2}(X_\mu^\dagger X_\nu - X_\nu^\dagger X_\mu) F^{\mu\nu} (\times)$	$-2e\kappa_\Lambda \left[\frac{m_e}{m_X} \left(\frac{1}{3} \mathcal{O}_1 - \frac{m_e^2}{q^2} \mathcal{O}_{19} \right) - \mathcal{O}_4 - \frac{m_e^2}{q^2} (\mathcal{O}_5 - \mathcal{O}_6) \right]$
	$\mathcal{L}_{\tilde{\kappa}_\Lambda} = i\frac{\tilde{\kappa}_\Lambda}{2}(X_\mu^\dagger X_\nu - X_\nu^\dagger X_\mu) \tilde{F}^{\mu\nu} (\times)$	$2e\tilde{\kappa}_\Lambda m_e^2 \frac{1}{q^2} \mathcal{O}_{11}$
	$\mathcal{O}_{X\gamma 1} = \epsilon^{\mu\nu\rho\sigma}(X_\rho^\dagger \overleftrightarrow{\partial}_\nu X_\sigma) \partial^\lambda F_{\mu\lambda}$	$-4em_e m_X (\mathcal{O}_8 - \mathcal{O}_9)$
	$\mathcal{O}_{X\gamma 2} = \epsilon^{\mu\nu\rho\sigma}i\partial_\nu(X_\rho^\dagger X_\sigma) \partial^\lambda F_{\mu\lambda} (\times)$	$-2em_e^2 \mathcal{O}_{11}$
	$\mathcal{O}_{X\gamma 3} = (X_\nu^\dagger i\overleftrightarrow{\partial}^\mu X^\nu) \partial^\lambda F_{\mu\lambda}$	$-4em_e m_X \mathcal{O}_1$
	$\mathcal{O}_{X\gamma 4} = \partial_\nu(X^{\mu\dagger} X^\nu + X^{\nu\dagger} X^\mu) \partial^\lambda F_{\mu\lambda}$	$4e m_e \left[\frac{1}{3} (i\mathbf{q} \cdot \mathbf{v}_{\text{el}}^\perp) \mathcal{O}_1 - m_e (\mathcal{O}_{17} + \mathcal{O}_{20}) \right]$
	$\mathcal{O}_{X\gamma 5} = i\partial_\nu(X^{\mu\dagger} X^\nu - X^{\nu\dagger} X^\mu) \partial^\lambda F_{\mu\lambda} (\times)$	$e \left[2m_e^2 \left(\mathcal{O}_5 - \mathcal{O}_6 - \frac{m_e}{m_X} \mathcal{O}_{19} \right) + 2\mathbf{q}^2 \mathcal{O}_4 + \frac{2}{3} \frac{m_e}{m_X} \mathbf{q}^2 \mathcal{O}_1 \right]$
	Vector case B	
dim-7	$\tilde{\mathcal{O}}_{\ell X 1}^S = (\bar{\ell}\ell)X_{\mu\nu}^\dagger X^{\mu\nu}$	$4m_e m_X^2 \mathcal{O}_1$
	$\tilde{\mathcal{O}}_{\ell X 2}^S = (\bar{\ell}\ell)X_{\mu\nu}^\dagger \tilde{X}^{\mu\nu}$	$4m_e^2 m_X \mathcal{O}_{11}$
	$\tilde{\mathcal{O}}_{\ell X 1}^P = (\bar{\ell}i\gamma_5\ell)X_{\mu\nu}^\dagger X^{\mu\nu}$	$-4m_e m_X^2 \mathcal{O}_{10}$
	$\tilde{\mathcal{O}}_{\ell X 2}^P = (\bar{\ell}i\gamma_5\ell)X_{\mu\nu}^\dagger \tilde{X}^{\mu\nu}$	$4m_e^2 m_X \mathcal{O}_6$
	$\tilde{\mathcal{O}}_{\ell X 1}^T = \frac{i}{2}(\bar{\ell}\sigma^{\mu\nu}\ell)(X_{\mu\rho}^\dagger X_\nu^\rho - X_{\nu\rho}^\dagger X_\mu^\rho), (\times)$	$4m_e m_X^2 \mathcal{O}_4$
	$\tilde{\mathcal{O}}_{\ell X 2}^T = \frac{1}{2}(\bar{\ell}\sigma^{\mu\nu}\gamma_5\ell)(X_{\mu\rho}^\dagger X_\nu^\rho - X_{\nu\rho}^\dagger X_\mu^\rho), (\times)$	$\frac{1}{3}m_e m_X [3m_X(\mathcal{O}_{11} + 4\mathcal{O}_{12}) - 4m_e(2\mathcal{O}_{10} + 3\mathcal{O}_{18})]$
dim-6	$\tilde{\mathcal{O}}_{X\gamma 1} = i(X_{\mu\rho}^\dagger X_\nu^\rho - X_{\nu\rho}^\dagger X_\mu^\rho) F^{\mu\nu} (\times)$	$2e \left[\frac{2}{3} m_X (m_e + m_X) \mathcal{O}_1 + 2m_X^2 \mathcal{O}_4 \right]$
		$+ \frac{1}{q^2} (2m_e^2 m_X^2 (\mathcal{O}_5 - \mathcal{O}_6) - 2m_e^2 m_X (m_X - 2m_e) \mathcal{O}_{19}) \Big]$
	$\tilde{\mathcal{O}}_{X\gamma 2} = i(X_{\mu\rho}^\dagger X_\nu^\rho - X_{\nu\rho}^\dagger X_\mu^\rho) \tilde{F}^{\mu\nu} (\times)$	$-4em_e^2 m_X^2 \frac{1}{q^2} \mathcal{O}_{11}$

Constraints on the relativistic operators



– Vector DM case

$$\mathcal{O}_{\ell X}^S = (\bar{\ell}\ell)(X_\mu^\dagger X^\mu)$$

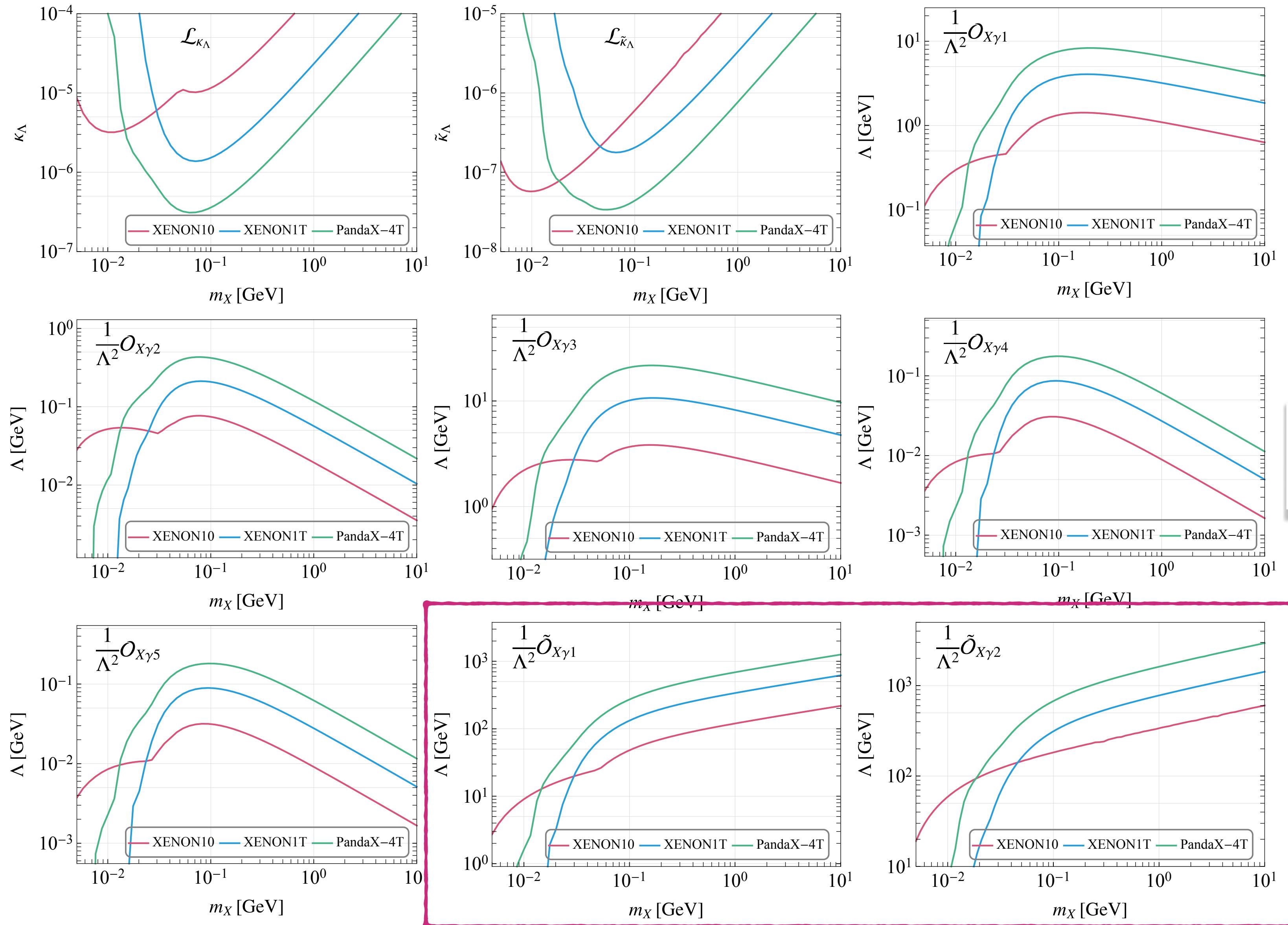
$$\mathcal{O}_{\ell X}^P = (\bar{\ell}i\gamma_5\ell)(X_\mu^\dagger X^\mu)$$

$$\mathcal{O}_{\ell X1}^T = \frac{i}{2}(\bar{\ell}\sigma^{\mu\nu}\ell)(X_\mu^\dagger X_\nu - X_\nu^\dagger X_\mu), (\times)$$

The PandaX-4T constraints on $\mathcal{O}_{\ell X}^S$ and $\mathcal{O}_{\ell X1}^T$ can up to about 10 TeV

Constraints on the relativistic operators

– Vector DM case



$$\begin{aligned}\tilde{\mathcal{O}}_{X\gamma 1} &= i(X_{\mu\rho}^\dagger X_\nu^\rho - X_{\nu\rho}^\dagger X_\mu^\rho) F^{\mu\nu} \quad (\times) \\ \tilde{\mathcal{O}}_{X\gamma 2} &= i(X_{\mu\rho}^\dagger X_\nu^\rho - X_{\nu\rho}^\dagger X_\mu^\rho) \tilde{F}^{\mu\nu} \quad (\times)\end{aligned}$$

$$2e \left[\frac{2}{3} m_X (m_e + m_X) \mathcal{O}_1 + 2m_X^2 \mathcal{O}_4 \right. \\ \left. + \frac{1}{q^2} (2m_e^2 m_X^2 (\mathcal{O}_5 - \mathcal{O}_6) - 2m_e^2 m_X (m_X - 2m_e) \mathcal{O}_{19}) \right] \\ - 4em_e^2 m_X^2 \frac{1}{a^2} \mathcal{O}_{11}\end{math>$$

Become stronger as m_X increases

Summary

- We find a crucial minus sign was missed for W_2 in 1912.08204, which has significant phenomenological consequences on some specific DM scenarios.
- A more compact amplitude squared is provided for the general DM-electron interactions for three DM scenarios.
- A matching dictionary between the relativistic and NR operators is given.
- The constraints from the xenon target experiments were studied, and we find the PandaX-4T set the most stringent constraints on the effective operators when $m_{\text{DM}} \gtrsim 20$ MeV.

Thank You!