

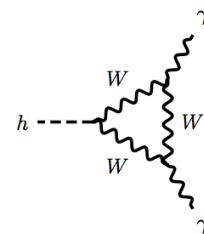
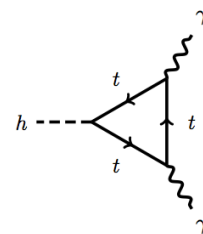
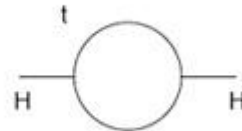
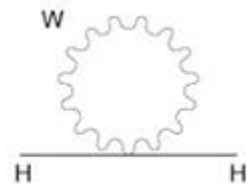
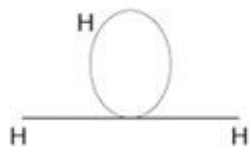
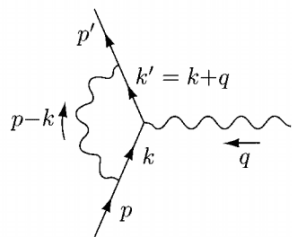
UV divergences of loops, the Higgs boson's low mass and the graviton loop in quantizing Einstein gravity

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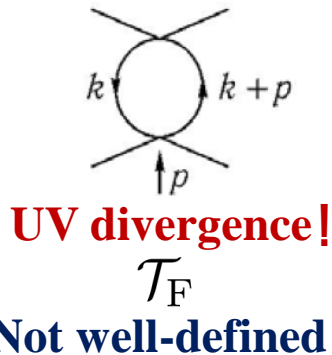
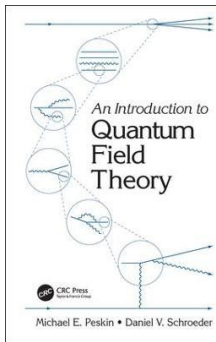
Based on arXiv:2305.18104, 2403.09487



Outline:

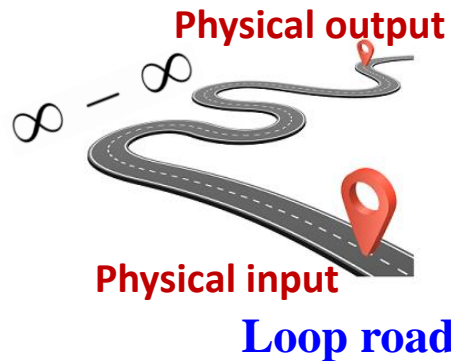
- **I. Background: UV divergences of loops**
- **II. Free flow of ideas --- UV-free scheme**
- **III. The hierarchy problem of Higgs mass**
- **IV. Graviton loop in Einstein gravity**
- **V. Summary and outlook**

I. Background: UV divergences of loops



Paradigm procedure

Devil or Angel?



Not well-defined

Regularization Renormalization

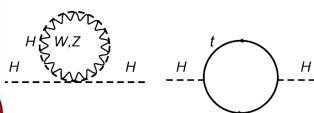
Renormalization



**Log divergence is OK,
power-law divergence
is problematic**

Two devils over the renormalization building

Large Devil (Higgs mass)



Huge Devil (Gravity)



Renormalization

$$\infty - \infty$$



QFT (SM)

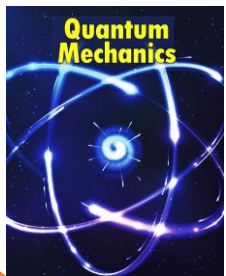


Next



II. Free flow of ideas --- UV-free scheme

Newton's Laws of Motion



Negligible

百
花
齊
放
各
展
其
新
奇

百花齊放
百家爭鳴

Low-energy
corrections

Loop

UV regions
(Planck scale)

Loop

Negligible?!

Regularization &
renormalization

本故事纯属虚构

UV-free scheme

arXiv:2305.18104

A presumption:



The physical contributions of loops are finite with contributions from UV regions being insignificant.

To obtain the physical results of loops, an equation is introduced

$$\mathcal{T}_F \rightarrow \mathcal{T}_P = \left[\int d\xi_1 \cdots d\xi_i \frac{\partial \mathcal{T}_F(\xi_1, \dots, \xi_i)}{\partial \xi_1 \cdots \partial \xi_i} \right]_{\{\xi_1, \dots, \xi_i\} \rightarrow 0} + C$$

(primary antiderivative + boundary constant)

or an equivalent form
$$\mathcal{T}_P = \left[\int (d\xi)^n \frac{\partial^n \mathcal{T}_F(\xi)}{\partial \xi^n} \right]_{\xi \rightarrow 0} + C,$$

ξ的自白
悄悄的我走了
正如我悄悄的来
我带走紫外发散
不带走一片云彩

今日种种，似水无痕
明夕何夕，君已陌路

\mathcal{T}_F

\mathcal{T}_P



除非 ...

UV-free scheme:

assume that the physical transition amplitude \mathcal{T}_P with propagators can be described by an equation of

$$\mathcal{T}_P = \left[\int d\xi_1 \cdots d\xi_i \frac{\partial \mathcal{T}_F(\xi_1, \dots, \xi_i)}{\partial \xi_1 \cdots \partial \xi_i} \right]_{\{\xi_1, \dots, \xi_i\} \rightarrow 0} + C, \quad (1)$$

a. Tree-level:

the photon propagator $\frac{-ig_{\mu\nu}}{p^2+i\epsilon}$,

$$\mathcal{T}_F(\xi) = \frac{-ig_{\mu\nu}}{p^2+\xi+i\epsilon}, \quad \frac{\partial \mathcal{T}_F(\xi)}{\partial \xi} = \frac{-ig_{\mu\nu}(-1)}{(p^2+\xi+i\epsilon)^2},$$

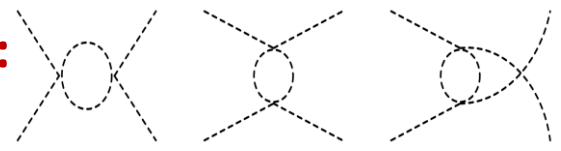
$$\left[\int d\xi \frac{\partial \mathcal{T}_F(\xi)}{\partial \xi} \right] = \frac{-ig_{\mu\nu}}{p^2+\xi+i\epsilon}, \quad \text{with } C = 0$$

$$\mathcal{T}_P = \left[\int d\xi \frac{\partial \mathcal{T}_F(\xi)}{\partial \xi} \right]_{\xi \rightarrow 0} = \frac{-ig_{\mu\nu}}{p^2+i\epsilon}$$

the gauge field propagator restored

b. Loop-level Log:

ϕ^4 theory



The physical scattering amplitude

$$\begin{aligned} \mathcal{T}_P(s) &= \left[\int d\xi \frac{\partial \mathcal{T}_F(\xi)}{\partial \xi} \right]_{\xi \rightarrow 0} + C_1 \\ &= \left[\frac{-\lambda^2}{2} \int d\xi \int \frac{d^4k}{(2\pi)^4} \frac{-i}{(k^2-m^2+\xi)^2} \frac{i}{(k+q)^2-m^2} \right]_{\xi \rightarrow 0} + C_1, \end{aligned}$$

$$\mathcal{T}_P(s) = \frac{-i\lambda^2}{32\pi^2} \int_0^1 dx \log[m^2 - x(1-x)s] + C_1.$$

A freedom of ξ in propagators

Considering the renormalization conditions, $s = 4m^2$,

$$t = u = 0. \quad \Rightarrow \quad C_1 = \frac{i\lambda^2}{32\pi^2} \int_0^1 dx \log[m^2 - 4m^2x(1-x)].$$

**No troublesome UV divergence
in loop calculations!**



In massless limit $\mathcal{T}_P = \mathcal{T}_P(s) + \mathcal{T}_P(t) + \mathcal{T}_P(u)$

$$s = -t = -u = \mu^2 = \frac{i\lambda^2}{32\pi^2} \left(\log \frac{\mu^2}{s} + \log \frac{\mu^2}{-t} + \log \frac{\mu^2}{-u} \right)$$

the n -point physical correlation function $G_P^{(n)}$ can be set by the physical field $\phi_P(x)$ with $\phi_P(x) = Z^{1/2}\phi(x, \mu)$, and the rescaling factor Z is finite here. The local correlation function $G^{(n)}$ (shorthand for a full expression $G^{(n)}(\phi, \lambda, m, \dots, \mu)$) in the perturbation expansion can be written as $G^{(n)} = Z^{-n/2}G_P^{(n)}$. Considering $\frac{dG_P^{(n)}}{d\mu} = 0$, the variation of μ in the massless limit can be described by a relation

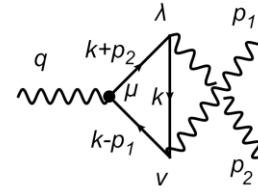
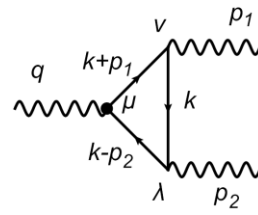
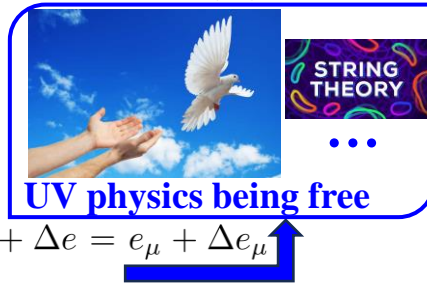
$$\left(\mu \frac{\partial}{\partial \mu} + \beta \frac{\partial}{\partial \lambda} + n\gamma \right) G^{(n)} = 0.$$

This is the form of the Callan-Symanzik equation [5, 6], and we have another picture about it in UV-free scheme. The μ -dependent term in UV-free scheme is from the boundary constant C . For the ϕ^4 theory in the massless limit, the one-loop result of the parameter γ is zero ($\mathcal{T}_P^{2p} = 0$). The beta function can be derived by Eq. (10), with the result

$$\begin{aligned} \beta &= -i\mu \frac{\partial}{\partial \mu} \mathcal{T}_P \\ &= \frac{3\lambda^2}{16\pi^2} + \mathcal{O}(\lambda^3). \end{aligned}$$

An illustration:

electron physical charge $e = e_0 + \Delta e = e_\mu + \Delta e_\mu$



γ^5 the original form

$$\partial_\mu j^{\mu 5} = iq_\mu \mathcal{T}_P^{\mu\nu\lambda} \epsilon_\nu^*(p_1) \epsilon_\lambda^*(p_2)$$

$$= -\frac{e^2}{16\pi^2} \left(\frac{2}{3} - 2 \log r \right) \epsilon^{\alpha\nu\beta\lambda} F_{\alpha\nu} F_{\beta\lambda}$$

Taking $C_0 = 2 \log r$

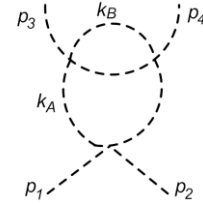
If $C_0 = \frac{2}{3}$

SM self-consistent

charge values of quarks coincidence, or correlation?

two-loop transition

$$\begin{aligned} \mathcal{T}_P &= \left[\int d\xi \frac{\partial \mathcal{T}_F(\xi)}{\partial \xi} \right]_{\xi \rightarrow 0} + C \\ &= \left[\frac{(-i\lambda)^3}{2} \int d\xi \int \frac{d^4 k_A}{(2\pi)^4} \frac{d^4 k_B}{(2\pi)^4} \frac{i}{k_A^2 - m^2} \frac{i}{(k_A + q)^2 - m^2} \right. \\ &\quad \left. \times \frac{-i}{(k_B^2 - m^2 + \xi)^2} \frac{i}{(k_B + k_A + p_3)^2 - m^2} \right]_{\xi \rightarrow 0} + C \end{aligned}$$

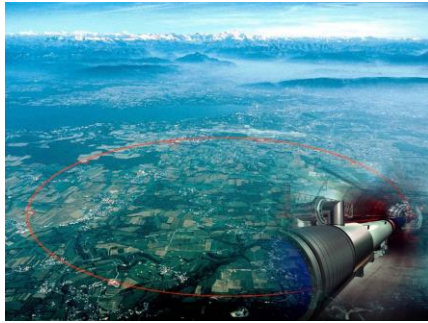


with $q = p_1 + p_2$

Log divergences are OK



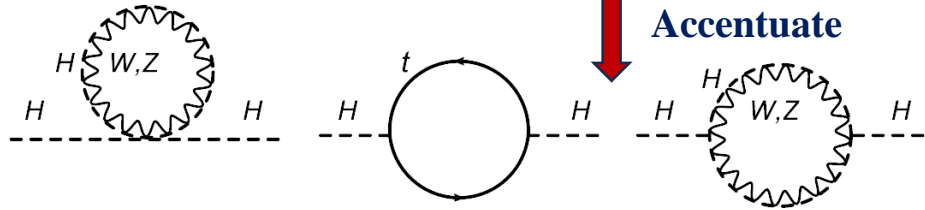
III. The hierarchy problem (c. Loop-level Λ^2, Λ^4)



LHC
Higgs boson

125 GeV

Accentuate



Power-law divergences (Λ^2, Λ^4)

For W, Z

In Feynman-'t Hooft gauge (Λ^2)

$$\mu \leftarrow \overset{\leftarrow}{k} \nu = \frac{-ig^{\mu\nu}}{k^2 - m_A^2}; \quad \text{---} \overset{\leftarrow}{k} \text{---} = \frac{i}{k^2 - m_A^2}.$$



In unitarity gauge (Λ^4)



$$\mu \leftarrow \overset{\leftarrow}{k} \nu = \frac{-i}{k^2 - m_A^2} \left(g^{\mu\nu} - \frac{k^\mu k^\nu}{m_A^2} \right)$$

Supersymmetry?

The hierarchy problem

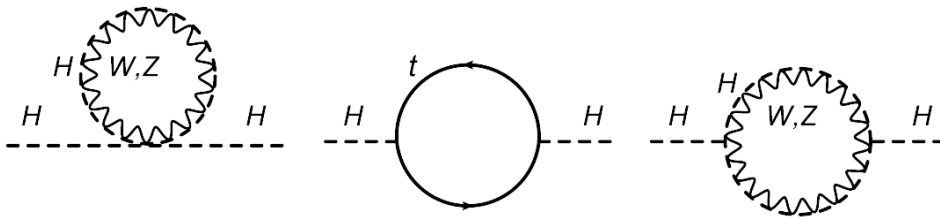
$$M_H^2 = (M_H^0)^2 + \frac{3\Lambda^2}{8\pi^2 v^2} [M_H^2 + 2M_W^2 + M_Z^2 - 4m_t^2]$$

Fine-tuning!



Large Devil
(Higgs mass)

A real problem for renormalization!



Power-law divergences (Λ^2 , Λ^4)

In UV-free scheme

Higgs in the first diagram

$$\begin{aligned} \mathcal{T}_P^{H1} &= \left[\int d\xi_1 d\xi_2 \frac{\partial \mathcal{T}_F^{H1}(\xi_1, \xi_2)}{\partial \xi_1 \partial \xi_2} \right]_{\{\xi_1, \xi_2\} \rightarrow 0} + C \\ &= \left[(-3i) \frac{m_H^2}{2v^2} \int d\xi_1 d\xi_2 \int \frac{d^4 k}{(2\pi)^4} \right. \\ &\quad \left. \times \frac{2i}{(k^2 - m_H^2 + \xi_1 + \xi_2)^3} \right]_{\{\xi_1, \xi_2\} \rightarrow 0} + C. \end{aligned}$$

After integral, one has

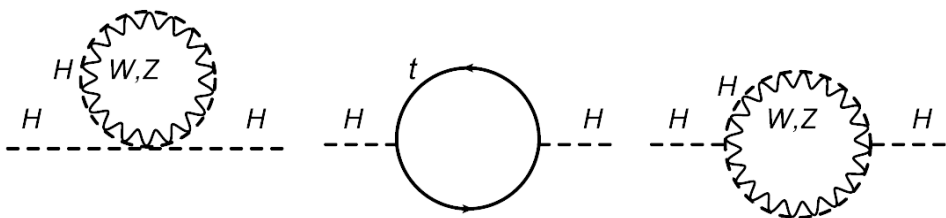
$$\begin{aligned} \mathcal{T}_P^{H1} &= i \frac{3m_H^4}{32\pi^2 v^2} \left(\log \frac{1}{m_H^2} + 1 \right) + C \\ &= i \frac{3m_H^4}{32\pi^2 v^2} \left(\log \frac{\mu^2}{m_H^2} + 1 \right). \end{aligned}$$

V (V=W,Z) in **unitary gauge**

$$\begin{aligned} \mathcal{T}_P^{V1} &= \left[\int d\xi_1 d\xi_2 d\xi_3 \frac{\partial \mathcal{T}_F^{V1}(\xi_1, \xi_2, \xi_3)}{\partial \xi_1 \partial \xi_2 \partial \xi_3} \right]_{\{\xi_1, \xi_2, \xi_3\} \rightarrow 0} + C \\ &= \left[i \frac{2m_V^2}{v^2 s_V} \int d\xi_1 d\xi_2 d\xi_3 \int \frac{d^4 k}{(2\pi)^4} g_{\mu\nu} \right. \\ &\quad \left. \times \frac{6i(g^{\mu\nu} - k^\mu k^\nu / m_V^2)}{(k^2 - m_V^2 + \xi_1 + \xi_2 + \xi_3)^4} \right]_{\{\xi_1, \xi_2, \xi_3\} \rightarrow 0} + C, \end{aligned}$$

where the symmetry factor s_V is $s_V = 1, 2$ for W, Z respectively. After integral, one has

$$\begin{aligned} \mathcal{T}_P^{V1} &= i \frac{2m_V^2}{v^2 s_V} \frac{m_V^2}{16\pi^2} \left(3 \log \frac{1}{m_V^2} + \frac{5}{2} \right) + C \\ &= i \frac{2m_V^2}{v^2 s_V} \frac{3m_V^2}{16\pi^2} \left(\log \frac{\mu^2}{m_V^2} + \frac{5}{6} \right). \end{aligned}$$



Power-law divergences (Λ^2 , Λ^4)

top quark loop

$$\begin{aligned} \mathcal{T}_P^t &= -\frac{3m_t^2}{v^2} \frac{i}{4\pi^2} \int_0^1 dx [m_t^2 - p^2 x(1-x)] \\ &\quad \times \left(3 \log \frac{1}{m_t^2 - p^2 x(1-x)} + 2 \right) + C \\ &= -\frac{3m_t^4}{v^2} \frac{3i}{4\pi^2} \int_0^1 dx \left[1 - \frac{p^2}{m_t^2} x(1-x) \right] \\ &\quad \times \left(\log \frac{\mu^2}{m_t^2 - p^2 x(1-x)} + \frac{2}{3} \right). \end{aligned}$$

Higgs in the third diagram

$$\begin{aligned} \mathcal{T}_P^{H3} &= \frac{9m_H^4}{2v^2} \frac{i}{16\pi^2} \int_0^1 dx \log \frac{1}{m_H^2 - x(1-x)p^2} + C \\ &= i \frac{9m_H^4}{32\pi^2 v^2} \int_0^1 dx \log \frac{\mu^2}{m_H^2 - x(1-x)p^2}. \end{aligned}$$

V (V=W,Z) in the third diagram

$$\begin{aligned} \mathcal{T}_P^{V3} &= \frac{4m_V^4}{v^2 s_V} \frac{6i}{16\pi^2} \int_0^1 dx \left(\left[\frac{1}{2} - \frac{p^2}{m_V^2} (x-x^2 + \frac{1}{12}) \right. \right. \\ &\quad \left. \left. + \frac{p^4}{m_V^4} \frac{x(1-x)}{12} (20x-20x^2-1) \right] \log \frac{1}{m_V^2 - x(1-x)p^2} \right. \\ &\quad \left. + \frac{1}{12} - \frac{p^2}{12m_V^2} (22x(1-x)-1) \right. \\ &\quad \left. - \frac{p^4 x(1-x)}{12m_V^4} (-21x(1-x)+1) \right) + C \\ &= \frac{m_V^4}{v^2 s_V} \frac{3i}{2\pi^2} \int_0^1 dx \left(\left[\frac{1}{2} - \frac{p^2}{m_V^2} (x-x^2 + \frac{1}{12}) \right. \right. \\ &\quad \left. \left. + \frac{p^4}{m_V^4} \frac{x(1-x)(20x-20x^2-1)}{12} \right] \log \frac{\mu^2}{m_V^2 - x(1-x)p^2} \right. \\ &\quad \left. + \frac{1}{12} - \frac{p^2(22x(1-x)-1)}{12m_V^2} - \frac{p^4 x(1-x)(-21x(1-x)+1)}{12m_V^4} \right) \end{aligned}$$

Considering μ in the electroweak scale,

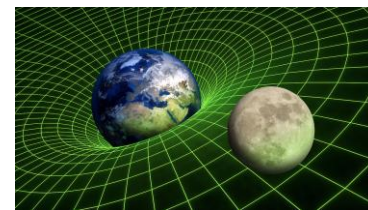
125 GeV Higgs can be obtained without fine-tuning, i.e., an alternative interpretation within SM.

Power-law divergences are OK in UV-free scheme!



除非 ...

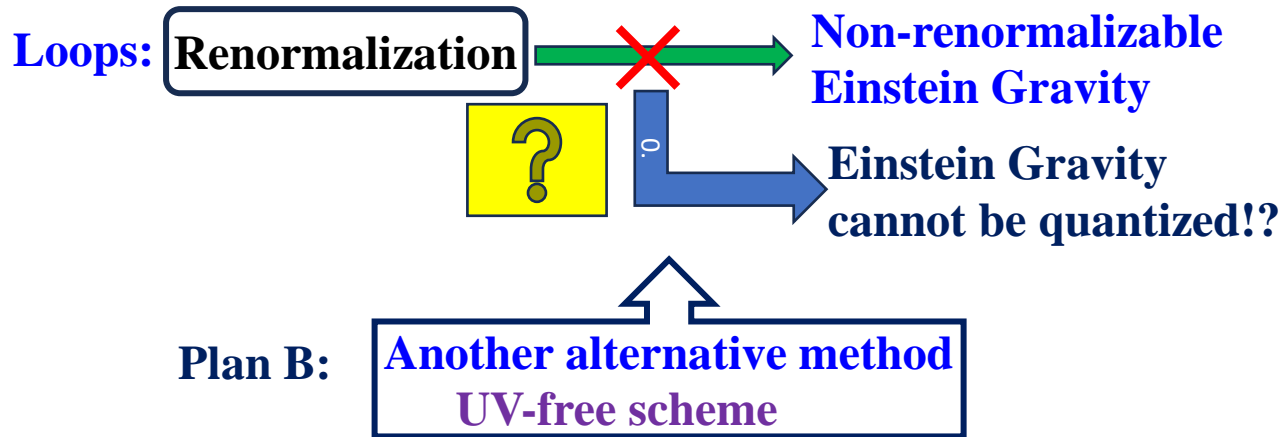
IV. Graviton loop in Einstein gravity



Huge Devil (Gravity)

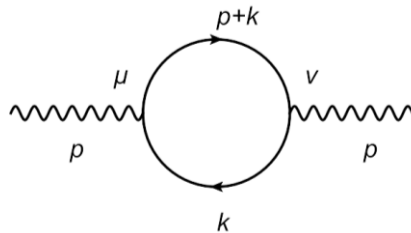


$$\mathcal{S} = \int d^4X \sqrt{-g} \left[-\frac{2}{\kappa^2} R + \mathcal{L}_M \right] \quad g_{\mu\nu} = \eta_{\mu\nu} + \kappa h_{\mu\nu}$$



For the primary antiderivative ξ -dependent choice

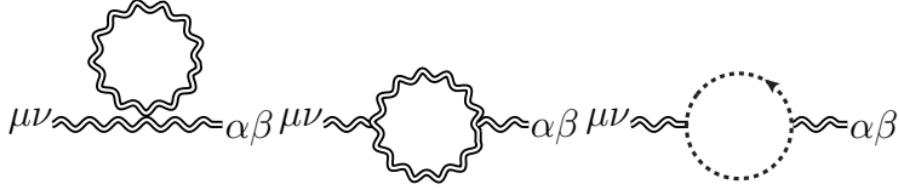
$$\begin{aligned} \mathcal{T}_P^{t2n} &= A \left[\frac{(\xi + \Delta)^n}{n!} (\log|\xi + \Delta| - (\sum_{l=1}^n \frac{1}{l})) \right]_{\xi \rightarrow 0} + C_1 \\ &= A \frac{\Delta^n}{n!} \log|\Delta| + C. \end{aligned}$$



$$\begin{aligned} \mathcal{T}_P^{\mu\nu} &= -\frac{ie^2}{2\pi^2} \int_0^1 dx (p^\mu p^\nu - g^{\mu\nu} p^2) x(1-x) \\ &\quad \times \log(m^2 - p^2 x(1-x)) + C^{\mu\nu}, \end{aligned}$$

with the Ward identity automatically preserved by the primary antiderivative.

One-loop propagator



The $\mu\nu \leftrightarrow \alpha\beta$ asymmetry involved at one-loop level in a particle propagation means that time reversal is not invariant in quantum gravity, i.e. an arrow of time at the microscopic level.

$$\mathcal{T}_P^a = \left[i\kappa^2 \frac{i\Pi_{\mu_3\nu_3\mu_4\nu_4}}{2} \frac{i}{16\pi^2} \left(V^{\mu_3\nu_3\mu_4\nu_4|\lambda_1\mu\nu\lambda_2\alpha\beta} p_{\lambda_1} p_{\lambda_2} \right. \right. \\ \left. \left. \times (\xi_1 - \xi_1 \log \xi_1) + \frac{V^{\mu\nu\alpha\beta|\lambda_3\mu_3\nu_3\lambda_4\mu_4\nu_4} \eta_{\lambda_3\lambda_4}}{4} \right) \right]_{\xi_1 \rightarrow 0} \\ \left. \times \left(\xi_1^2 \log \xi_1 - \frac{3}{2} \xi_1^2 \right) \right]_{\xi_1 \rightarrow 0} + C_a^{\mu\nu\alpha\beta} .$$

$$= 0.$$

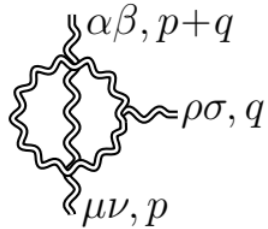
$$\mathcal{T}_P^b = \frac{(2i\kappa)^2}{2} \frac{i}{16\pi^2} \int_0^1 dx \left(-\frac{1}{4} \right) \left\{ \frac{1}{16} [40x^2(1-x)^2 p^\mu p^\nu p^\alpha p^\beta \right. \\ + 2p^2((1-2x)^2(15x^2-15x-2)(p^\mu p^\nu \eta^{\alpha\beta} + p^\alpha p^\beta \eta^{\mu\nu}) \\ + (10x^4-20x^3+17x^2-7x+2)(p^\nu p^\beta \eta^{\mu\alpha} + p^\mu p^\beta \eta^{\nu\alpha} \\ + p^\nu p^\alpha \eta^{\mu\beta} + p^\mu p^\alpha \eta^{\nu\beta})) + p^4((115x^4-230x^3+103x^2 \\ + 12x+1)\eta^{\mu\nu} \eta^{\alpha\beta} + (85x^4-170x^3+139x^2-54x+3) \\ \left. \times (\eta^{\mu\alpha} \eta^{\nu\beta} + \eta^{\mu\beta} \eta^{\nu\alpha})) \right\} \log \frac{1}{-p^2 x(1-x)} \Big\} + C_b^{\mu\nu\alpha\beta} .$$

$$\mathcal{T}_P^c = (-1)(i\kappa)^2 \frac{4i}{16\pi^2} \int_0^1 dx \left(-\frac{1}{4} \right) \left\{ \frac{1}{4} [4(4x^4-8x^3+2x^2 \right. \\ + 2x+1)p^\mu p^\nu p^\alpha p^\beta + p^2((8x^4-16x^3+4x^2+4x-1) \\ \times (p^\nu p^\beta \eta^{\mu\alpha} + p^\mu p^\beta \eta^{\nu\alpha} + p^\nu p^\alpha \eta^{\mu\beta} + p^\mu p^\alpha \eta^{\nu\beta}) \\ + 2x(14x^3-24x^2+13x-4)p^\mu p^\nu \eta^{\alpha\beta} + 2p^\alpha p^\beta \eta^{\mu\nu} \\ \times (14x^4-32x^3+25x^2-6x-1)) + p^4(2x(11x^3-22x^2 \\ + 13x-2)(\eta^{\mu\alpha} \eta^{\nu\beta} + \eta^{\mu\beta} \eta^{\nu\alpha}) + (12x^4-24x^3+16x^2 \\ - 4x+1)\eta^{\mu\nu} \eta^{\alpha\beta})] \log \frac{1}{-p^2 x(1-x)} \Big\} + C_c^{\mu\nu\alpha\beta} ,$$

n -loop with overlapping divergences

superficial degree of divergence $2n+2$ $\mathcal{T}_P^{t2n} = A \frac{\Delta^n}{n!} \log |\Delta| + C$

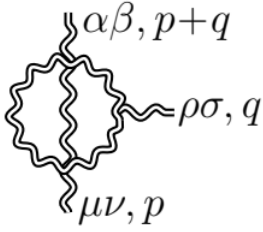
$$\mathcal{T}_P^{\text{total}} = \mathcal{T}_P^{t2(n+1)} + \mathcal{T}_P^{t2n} + \dots + \mathcal{T}_P^{t2} \\ + \mathcal{T}_P(\log) + \mathcal{T}_P(\text{finite}) ,$$



$$\mathcal{T}_P^V = (2i)^3 \kappa^5 \left(\frac{i}{16\pi^2} \right)^2 \int_0^1 dx \int_0^1 dy \int_0^1 dz \frac{z}{2^4(1-z)^4} \\ \times \left\{ A_3 \frac{\Delta_0^3}{3!} + A_2 \frac{\Delta_0^2}{2!} + A_1 \Delta_0 + A_0 \right\} \log \frac{1}{\Delta_0} + C^{\mu\nu\alpha\beta\rho\sigma}$$

Here Δ_0 is $\Delta_0 = b^2 - ac$, with $a = z + (1-z)x(x-1)$, $b = yzq + (1-z)x(x-1)p$, $c = yzq^2 + (1-z)x(x-1)p^2$. A_3, A_2, A_1, A_0 are coefficients related to sextic, quartic, quadratic, logarithmic divergence inputs respectively.

arXiv: 2403.09487



$$\begin{aligned}
 A_3 = & \frac{z-1}{64a^8} \left([440a^2 + a(1564x^2 + 1300x + 23)](z-1) \right. \\
 & + 4(281x^4 - 562x^3 + 683x^2 - 402x + 273)(z-1)^2] \\
 & \times \eta^{\mu\nu} (\eta^{\alpha\rho} \eta^{\beta\sigma} + \eta^{\alpha\sigma} \eta^{\beta\rho}) + [744a^2 + a(1932x^2 + 44x \\
 & + 1203)](z-1) + 4(297x^4 - 594x^3 + 1563x^2 - 1266x \\
 & + 673)(z-1)^2] \eta^{\rho\sigma} (\eta^{\alpha\nu} \eta^{\beta\mu} + \eta^{\alpha\mu} \eta^{\beta\nu}) + [440a^2 \\
 & + a(1564x^2 - 1100x + 2423)](z-1) + 4(281x^4 \\
 & - 562x^3 + 683x^2 - 402x + 273)(z-1)^2] \eta^{\alpha\beta} (\eta^{\mu\rho} \eta^{\nu\sigma} \\
 & + \eta^{\mu\sigma} \eta^{\nu\rho}) + [1032a^2 + a(3396x^2 - 3020x + 801) \\
 & \times (z-1) + 4(591x^4 - 1182x^3 + 1101x^2 - 510x + 215) \\
 & \times (z-1)^2] (\eta^{\alpha\rho} \eta^{\beta\nu} \eta^{\mu\sigma} + \eta^{\alpha\nu} \eta^{\beta\rho} \eta^{\mu\sigma} + \eta^{\alpha\nu} \eta^{\beta\sigma} \eta^{\mu\rho} \\
 & + \eta^{\alpha\sigma} \eta^{\beta\nu} \eta^{\mu\rho} + \eta^{\alpha\rho} \eta^{\beta\mu} \eta^{\nu\sigma} + \eta^{\alpha\mu} \eta^{\beta\rho} \eta^{\nu\sigma} + \eta^{\alpha\mu} \eta^{\beta\sigma} \eta^{\nu\rho} \\
 & + \eta^{\alpha\sigma} \eta^{\beta\mu} \eta^{\nu\rho}) + [1696a^2 + a(4844x^2 + 848x + 4147) \\
 & \times (z-1) + 4(787x^4 - 1574x^3 + 2521x^2 - 1734x \\
 & + 795)](z-1)^2] \eta^{\alpha\beta} \eta^{\mu\nu} \eta^{\rho\sigma} \Big).
 \end{aligned}$$

Parameter A_0 ($l = p + q$)

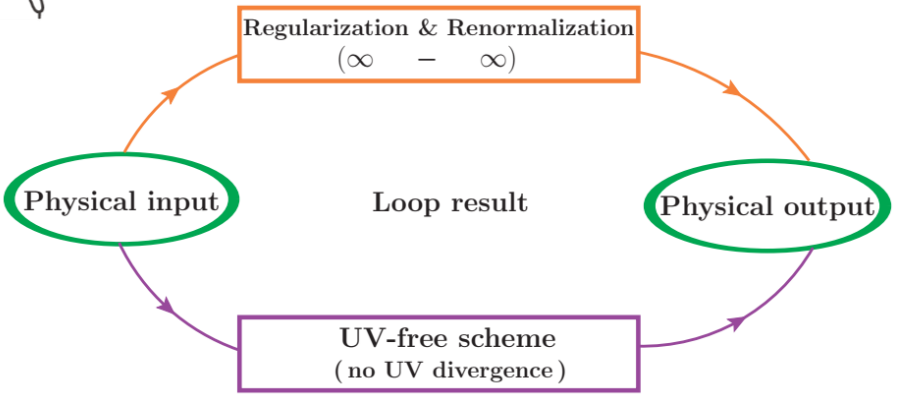
In the case of $p^2 = l^2 = 0$, the result is

$$\begin{aligned}
 A_0 = & \frac{(z-1)^3}{64a^8} \left\{ 16y^2z^3[a^3(8x^2 - 8x + 7) - 2a^2(4x^4 - 8x^3 + 16x^2 - 12x + 11)]yz + a(14x^4 - 28x^3 + 53x^2 - 39x + 28)y^2z^2 \right. \\
 & - 14(x^2 - x + 1)^2y^2z^3[\eta^{\alpha\rho} \eta^{\beta\sigma} \eta^{\mu\nu} \eta^{\rho\sigma} - 8y^2z^2[a^4(6 - 9x + 9x^2) + a^2(x-1)x(47 - 75x + 83x^2 - 16x^3 + 8x^4)y(1-z) - \\
 & - 2a(x-1)x(21 - 25x + 32x^2 - 14x^3 + 7x^4)y^2(1-z)z^2 + 28(x-1)x(1-x+x^2)^2y^3(1-z)z^3 + a^4((x-1)x(-12+41x \\
 & - 41x^2)(1-z) + (-7+14x-6x^2-16x^3+8x^4)y)] + [\eta^{\alpha\rho} \eta^{\beta\sigma} \eta^{\mu\nu} \eta^{\rho\sigma} + \eta^{\alpha\rho} \eta^{\beta\sigma} \eta^{\mu\nu} \eta^{\rho\sigma} + 8y^2z^2[a^4(7-9x+9x^2) - 28 \\
 & \times (x-1)x(1-x+x^2)^2y^3(1-z)z^3 + ay^2z^2(x-1)x(49-57x+59x^2-4x^3+2x^4)(1-z) + (-14+45x-73x^2+56x^3 \\
 & - 28x^4)yz] + a^3((x-1)x(-3+29x-29x^2)(1-z) + (-7-9x+x^2+16x^3-8x^4)yz) + a^2yz((x-1)x(-9-21x+13x^2 \\
 & + 16x^3-8x^4)(1-z) + (12-17x+37x^2-40x^3+20x^4)yz)] + [\eta^{\alpha\rho} \eta^{\beta\sigma} \eta^{\mu\nu} \eta^{\rho\sigma} + \eta^{\alpha\rho} \eta^{\beta\sigma} \eta^{\mu\nu} \eta^{\rho\sigma} + 4yz[2a^3(-56(x-1)^2 \\
 & \times z^2 + a^2(2(x-1)x^2y^2(1-z)z^2 + a^4(9(x-1)x(1-z) + 2(3-7x+7x^2)yz) + 2a(x-1)x(1-x+x^2)^2y^3(1-z)z^3 - 19x^2 \\
 & - 17x^2 - 24x^3 + 12x^4)(1-z) + (14-45x+73x^2-56x^3+28x^4)yz] + a^2(x-1)x(1-x+x^2)(10(x-1)x(5-6x+6x^2) \\
 & \times (1-z) + (-38+31x+9x^2-80x^3+40x^4)yz) + 2a^2(-3(x-1)^2z^2(1-z)^2 + (x-1)x(17-18x+18x^2)y(1-z)z \\
 & + (2-3x+11x^2-16x^3+8x^4)y^2z^2)] + [\eta^{\alpha\rho} \eta^{\beta\sigma} \eta^{\mu\nu} \eta^{\rho\sigma} + \eta^{\alpha\rho} \eta^{\beta\sigma} \eta^{\mu\nu} \eta^{\rho\sigma} + \eta^{\alpha\rho} \eta^{\beta\sigma} \eta^{\mu\nu} \eta^{\rho\sigma} + \eta^{\alpha\rho} \eta^{\beta\sigma} \eta^{\mu\nu} \eta^{\rho\sigma} - 8yz[a^5(6-6x \\
 & + 6x^2) + 28(x-1)^2z^2(1-x+x^2)^2y^3(1-z)z^3 + 2a^4((x-1)x(7-x+x^2)(1-z) + (-21+34x-30x^2-8x^3+4x^4) \\
 & \times yz) + 2a(x-1)x(1-x+x^2)(3(x-1)x(-7+6x-2x^2-8x^3+4x^4)(1-z) + (14-45x+73x^2-56x^3+28x^4)yz) \\
 & + a^3(2(x-1)^2z^2(3+2x-2x^2)(1-z)^2 + (x-1)x(-47+107x-91x^2-32x^3+16x^4)y(1-z)z) + (53-50x+30x^2 \\
 & + 40x^3-20x^4)y^2z^2] + a^2yz((x-1)^2z^2(-30+87x-79x^2-16x^3+8x^4)(1-z)^2 + (x-1)x(11-30x+34x^2-8x^3 \\
 & + 4x^4)y(1-z)z) + 2(-8-11x+25x^2-28x^3+14x^4)y^2z^2)] + [\eta^{\alpha\rho} \eta^{\beta\sigma} \eta^{\mu\nu} \eta^{\rho\sigma} - 8[a^5 + 28(x-1)^3x^2(1-x+x^2)^2y^3(1-z)^3 \\
 & \times z^3 + a^4(2(x-1)x(1-x+x^2)(1-z) + (-2-5x+5x^2)yz) + 2a(x-1)^2z^2(1-x+x^2)^2y^3(1-x-1x(-14+4x+15x^2 \\
 & - 38x^3+19x^4)(1-z) + (14-45x+73x^2-56x^3+28x^4)yz] + a^3((1-2x)(x-1)^2z^2(1-x)^2 + (x-1)x(7-8x+8x^2) \\
 & \times y(1-z)z) + (1+8x-16x^3+8x^4)y^2z^2] + a^2(x-1)x(1-x+x^2)(-1+12x-12x^2)(1-z)^2 + 3(x-1) \\
 & \times (-1-11x+39x^2-56x^3+28x^4)y(1-z)z) + 2(-8-11x+25x^2-28x^3+14x^4)y^2z^2] + a^3(x-1)x(1-x+x^2)(2(x-1)^2 \\
 & \times z^2(1-x+x^2)(1-z)^2 + (x-1)x(10+9x-9x^2)y(1-z)z) + (18-25x+79x^2-108x^3+54x^4)y^2z^2)] + [\eta^{\alpha\rho} \eta^{\beta\sigma} \eta^{\mu\nu} \eta^{\rho\sigma} \\
 & + \eta^{\alpha\rho} \eta^{\beta\sigma} \eta^{\mu\nu} \eta^{\rho\sigma} + 8y^2z^2[8a^4(1-2x+2x^2) - 28(x-1)x(1-x+x^2)^2y^3(1-z)z^3 + ay^2z^2((x-1)x(49-88x+142x^2 \\
 & - 108x^3+54x^4)(1-z) + (-14+45x-73x^2+56x^3-28x^4)yz) + a^3(12(x-1)x(1-x+x^2)(1-z) + (-31+88x-104x^2 \\
 & + 32x^3-16x^4)yz) + 2a^2yz((x-1)x(-16+21x-29x^2+16x^3-8x^4)(1-z) + (17-51x+69x^2-36x^3+18x^4)y(1-z) \\
 & \times [\eta^{\alpha\rho} \eta^{\beta\sigma} \eta^{\mu\nu} \eta^{\rho\sigma} + \eta^{\alpha\rho} \eta^{\beta\sigma} \eta^{\mu\nu} \eta^{\rho\sigma} - 4yz[2a^5(6-11x+11x^2) + 56(x-1)^2z^2(1-x+x^2)^2y^3(1-z)z^3 + a^4((x-1) \\
 & \times y(1-z)z)((x-1)x(68-113x+129x^2-32x^3+16x^4)(1-z) + (-26+117x-213x^2+192x^3-96x^4)yz) + a^4((x-1) \\
 & \times (x-11+52x-52x^2)(1-z) + 2(-4+7x+x^2-16x^3+8x^4)yz) + 2a(x-1)x(1-x+x^2)(-5(-1+x)x(7-12x+20x^2 \\
 & - 16x^3+8x^4)(1-z) + (14-45x+73x^2-56x^3+28x^4)yz) + a^3(-2(x-1)^2z^2(12-37x+37x^2)(1-z)^2 + (x-1)x \\
 & \times (27-31x+63x^2-64x^3+32x^4)y(1-z)z) - 2(2-3x+11x^2-16x^3+8x^4)y^2z^2)] + [\eta^{\alpha\rho} \eta^{\beta\sigma} \eta^{\mu\nu} \eta^{\rho\sigma} + \eta^{\alpha\rho} \eta^{\beta\sigma} \eta^{\mu\nu} \eta^{\rho\sigma} \\
 & + \eta^{\alpha\rho} \eta^{\beta\sigma} \eta^{\mu\nu} \eta^{\rho\sigma} + \eta^{\alpha\rho} \eta^{\beta\sigma} \eta^{\mu\nu} \eta^{\rho\sigma} - 4yz[2a^5(3-5x+5x^2) + 56(x-1)^2x^2(1-x+x^2)^2y^3(1-z)z^3 + a^4((x-1)x \\
 & \times (-11-32x+32x^2)(1-z) + 2(27-44x+52x^2-16x^3+8x^4)yz) + 2a(x-1)x(1-x+x^2)((x-1)x(-42+67x-95x^2 \\
 & + 56x^3-28x^4)(1-z) + 2(14-45x+73x^2-56x^3+28x^4)yz) + a^2yz((x-1)^2z^2(7+36x-20x^2-32x^3+16x^4)(1-z) - \\
 & - 8(x-1)x(13-29x+43x^2-28x^3+14x^4)y(1-z)z) + 4(10-21x+35x^2-28x^3+14x^4)y^2z^2] - 2a^3((x-1)^2z^2(-3 \\
 & + 31x-31x^2)(1-z)^2 + (x-1)x(-21+17x-33x^2+32x^3-16x^4)y(1-z)z) + (35-57x+85x^2-56x^3+28x^4)y^2z^2] \\
 & \times [\eta^{\alpha\rho} \eta^{\beta\sigma} \eta^{\mu\nu} \eta^{\rho\sigma} + \eta^{\alpha\rho} \eta^{\beta\sigma} \eta^{\mu\nu} \eta^{\rho\sigma} + \eta^{\alpha\rho} \eta^{\beta\sigma} \eta^{\mu\nu} \eta^{\rho\sigma} + \eta^{\alpha\rho} \eta^{\beta\sigma} \eta^{\mu\nu} \eta^{\rho\sigma} - 4[56(x-1)^3x^2(1-x+x^2)^2y^3(1-z)z^3 + 2a^5 \\
 & \times (x-1)^2z^2(1-z) + (1-x+x^2)yz) - a^4(x-1)x(1-z)((x-1)x(1-z) + 4(x-1)x^2(1-z) - 4(x-1)x^3(1-z) \\
 & + (-10-31x+31x^2)yz) + 2a(x-1)^2z^2y^3((x-1)x(-28+39x-53x^2+28x^3-14x^4)(1-z) + 2(14-45x \\
 & + 73x^2-56x^3+28x^4)yz) + a^3(x-1)x(1-x+x^2)(37-41x+41x^2)(1-z)^2 - 2(x-1)x(38-71x+99x^2 \\
 & - 56x^3+28x^4)y(1-z)z) + 4(10-21x+35x^2-28x^3+14x^4)y^2z^2] + a^3(x-1)x(1-z)((x-1)^2z^2(-5-2x+2x^2) \\
 & \times (1-z)^2 + 8(x-1)x(6-x+x^2)y(1-z)z) - 2(25-36x+50x^2-28x^3+14x^4)y^2z^2)] + [\eta^{\alpha\rho} \eta^{\beta\sigma} \eta^{\mu\nu} \eta^{\rho\sigma} + \eta^{\alpha\rho} \eta^{\beta\sigma} \eta^{\mu\nu} \eta^{\rho\sigma} \\
 & + \eta^{\alpha\rho} \eta^{\beta\sigma} \eta^{\mu\nu} \eta^{\rho\sigma} + \eta^{\alpha\rho} \eta^{\beta\sigma} \eta^{\mu\nu} \eta^{\rho\sigma} + \eta^{\alpha\rho} \eta^{\beta\sigma} \eta^{\mu\nu} \eta^{\rho\sigma} + \eta^{\alpha\rho} \eta^{\beta\sigma} \eta^{\mu\nu} \eta^{\rho\sigma} + 4[4a^5(1-x \\
 & + x^2) - 56(x-1)^3x^2(1-x+x^2)^2y^3(1-z)z^3 + 2a^4((x-1)x(3-8x+8x^2)(1-z) + (2-21x+13x^2+16x^3-8x^4) \\
 & \times yz) + 2a(x-1)^2z^2y^3((x-1)x(35-46x+48x^2-4x^3+2x^4)(1-z) - 3(14-45x+73x^2-56x^3+28x^4) \\
 & \times yz) + a^4(4(x-1)^2z^2(1-5x+5x^2)(1-z)^2 + (x-1)x(61-104x+56x^2+96x^3-48x^4)(1-z)z) + 2(1+9x+19x^2 \\
 & - 56x^3+28x^4)y^2z^2] - 2a^2(x-1)x(1-z)((x-1)^2z^2(-29+75x-67x^2-16x^3+8x^4)(1-z)^2 + (x-1)x(-35+66x
 \end{aligned}$$

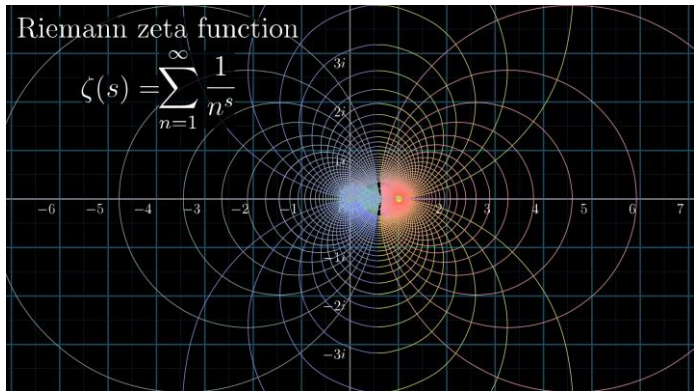
21 pages



Why does the UV-free scheme seem effective for power-law divergences and give another picture?



Two alternative routes of concern



The hierarchy problem

(a) New particles (TeV) needed to cancel out UV contributions of loops to the Higgs mass

(b) An interpretation within SM



(a) *Equivalent transformation* of the loop integral from UV divergence to UV divergence mathematically expressed form (regularization), with renormalization required to remove the UV divergence.

(b) *Analytic continuation* of the transition amplitude from UV divergent \mathcal{T}_F to UV converged \mathcal{T}_P (the UV-free scheme here), without UV divergences in calculations.

UV-free scheme Analytic continuation

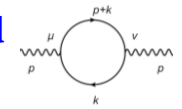
$$\mathcal{T}_F \longrightarrow \mathcal{T}_P = \left[\int (d\xi)^n \frac{\partial^n \mathcal{T}_F(\xi)}{\partial \xi^n} \right]_{\xi \rightarrow 0} + C,$$

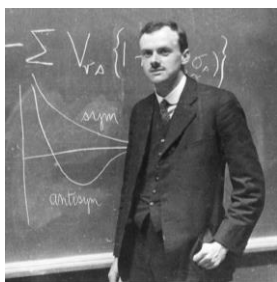
Finite input		UV divergence input (continuation)	
Tree level	Loop finite	Loop Log	Loop $\Lambda^2, \Lambda^4, \Lambda^6, \dots$

Originally well-defined

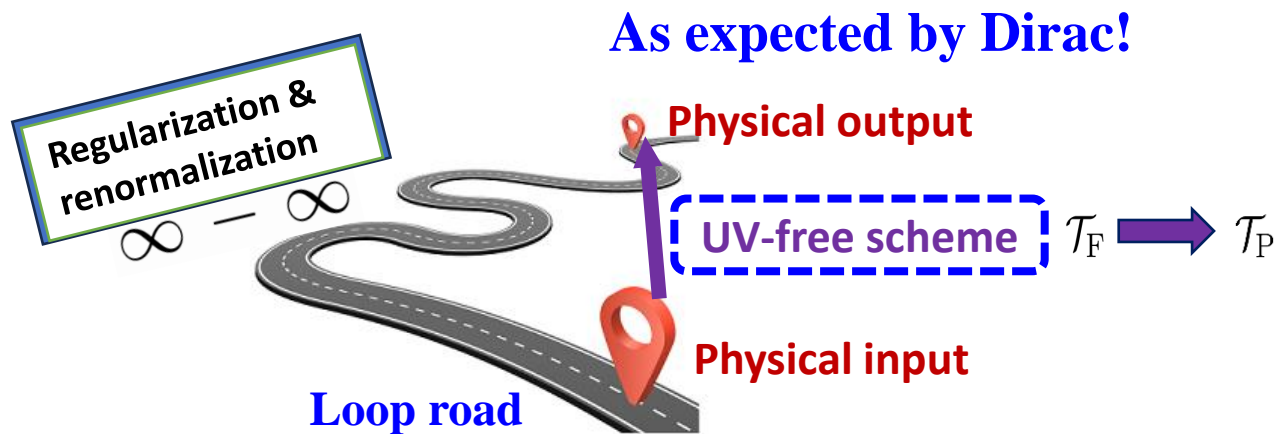
Verified

To be verified





P. A. M. Dirac I believe the successes of the renormalization theory will be on the same footing as the successes of the Bohr orbit theory applied to one-electron problems.



Schemes	Tree level	Loop finite	Loop Log	Loop $\Lambda^2, \Lambda^4, \Lambda^6, \dots$
Regularization & renormalization			OK	Problematic
UV-free scheme	OK	OK	OK	OK



Both loops of the renormalizable **Standard Model** and non-renormalizable **Einstein gravity** being OK!



V. Summary and outlook

A. An alternative method --- **UV-free scheme**:
Finite loop results obtained without UV divergences, the original γ^5 matrix, and effective for loop Log and power-law divergence inputs.

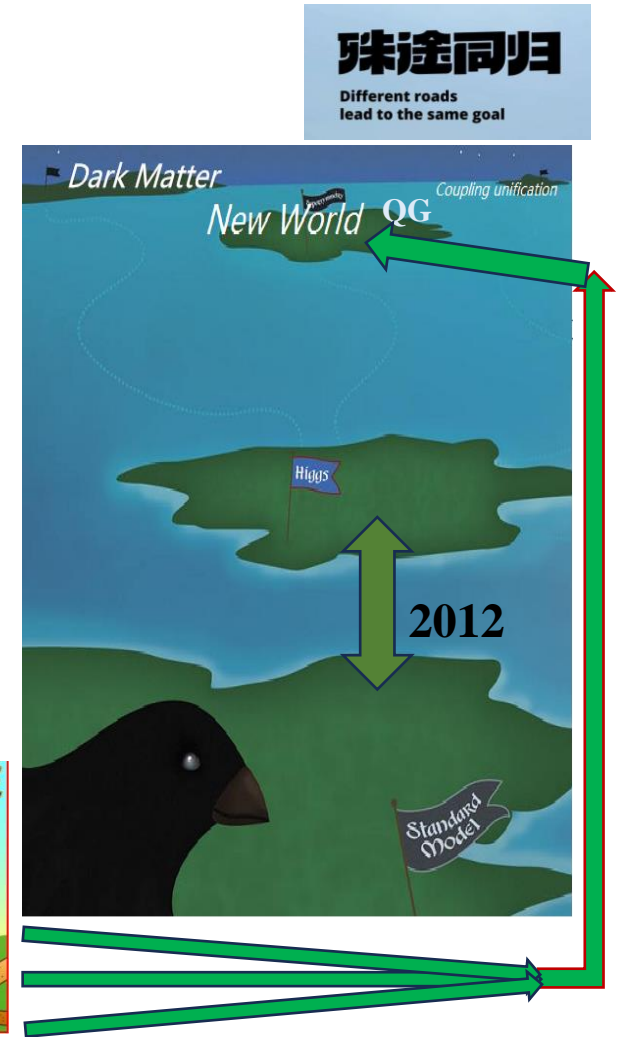
B. **To the hierarchy problem** of the 125 GeV Higgs, an alternative interpretation without fine-tuning within SM.

C. It is possible to incorporate **Einstein gravity** into the framework of QFT.

Outlook:

It is the beginning of a new alternative method.

Thank you!



殊途同归

Different roads
lead to the same goal