

Latest result on searching for fractionally charged particles with the DAMPE

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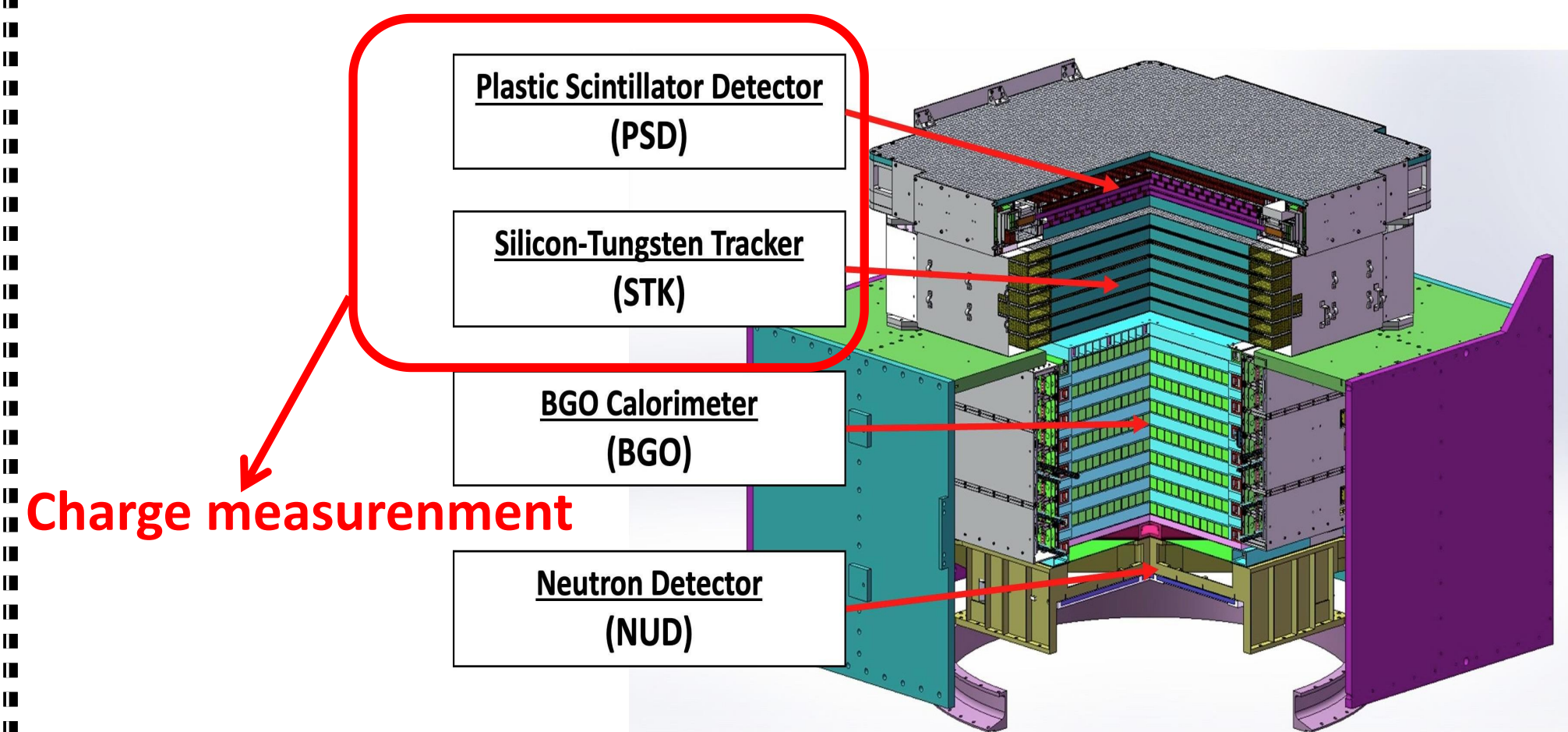
Outline

- Motivation and DAMPE
- Search for FCP with DAMPE
- Summary

Motivation and DAMPE

- All charged particles have **multiples of electron charge** except quarks. According to theories of quantum chromodynamics (QCD) free quarks do not exist, however **Fractionally Charged Particles (FCPs)** are of great interest. Our first try is based on **heavy lepton assumption**.
- There are three possible sources of FCP in cosmic rays:
 - early universe
 - high-energy astrophysical processes
 - extensive air shower

- DAMPE Experiment
DAMPE is a space experiment for detecting high energy cosmic rays launched on Dec.17th 2015.



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Search for FCP with DAMPE

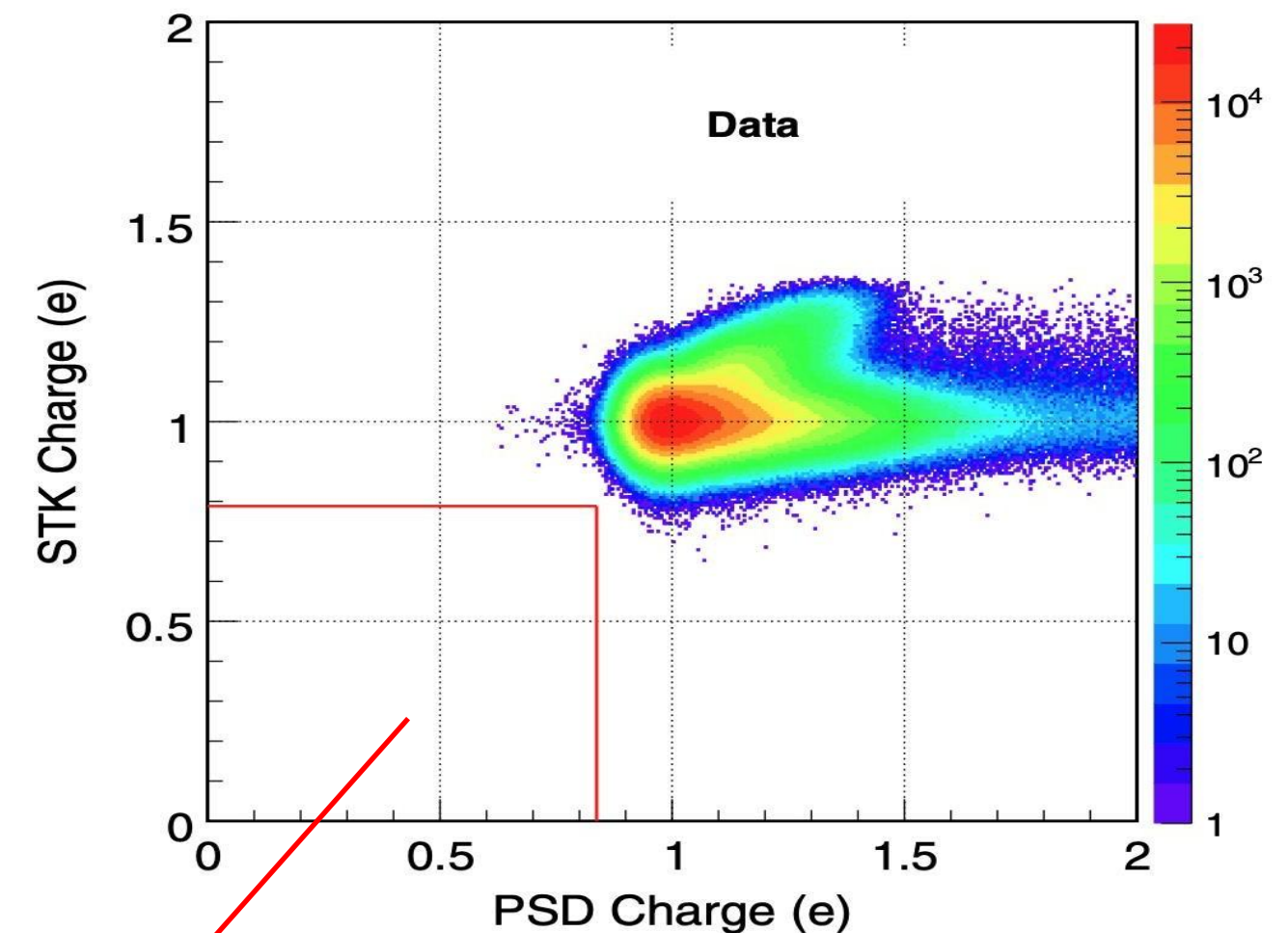
Charge reconstruction is based on **Bethe-Bloch Formula**. The energy deposited in the PSD and STK is **proportional** to the square of particle charge.

$$-\frac{dE}{dx} = K z^2 \frac{Z}{A} \frac{1}{\beta^2} \left[\frac{1}{2} \ln \frac{2m_e c^2 \beta^2 \gamma^2 T_{\max}}{I^2} - \beta^2 - \frac{\delta(\beta\gamma)}{2} \right]$$

	PSD Charge resolution (Charge unit, c.u.)	STK Charge resolution (Charge unit, c.u.)
Proton	0.06	0.04
Helium	0.10	0.07

Data Sample

- Simulation (**Assumption**):
 - like **a massive lepton** (e.g. muon)
 - The charge would be **2/3e**
- Flight Data: 2016.01.01 ~ 2020.12.31



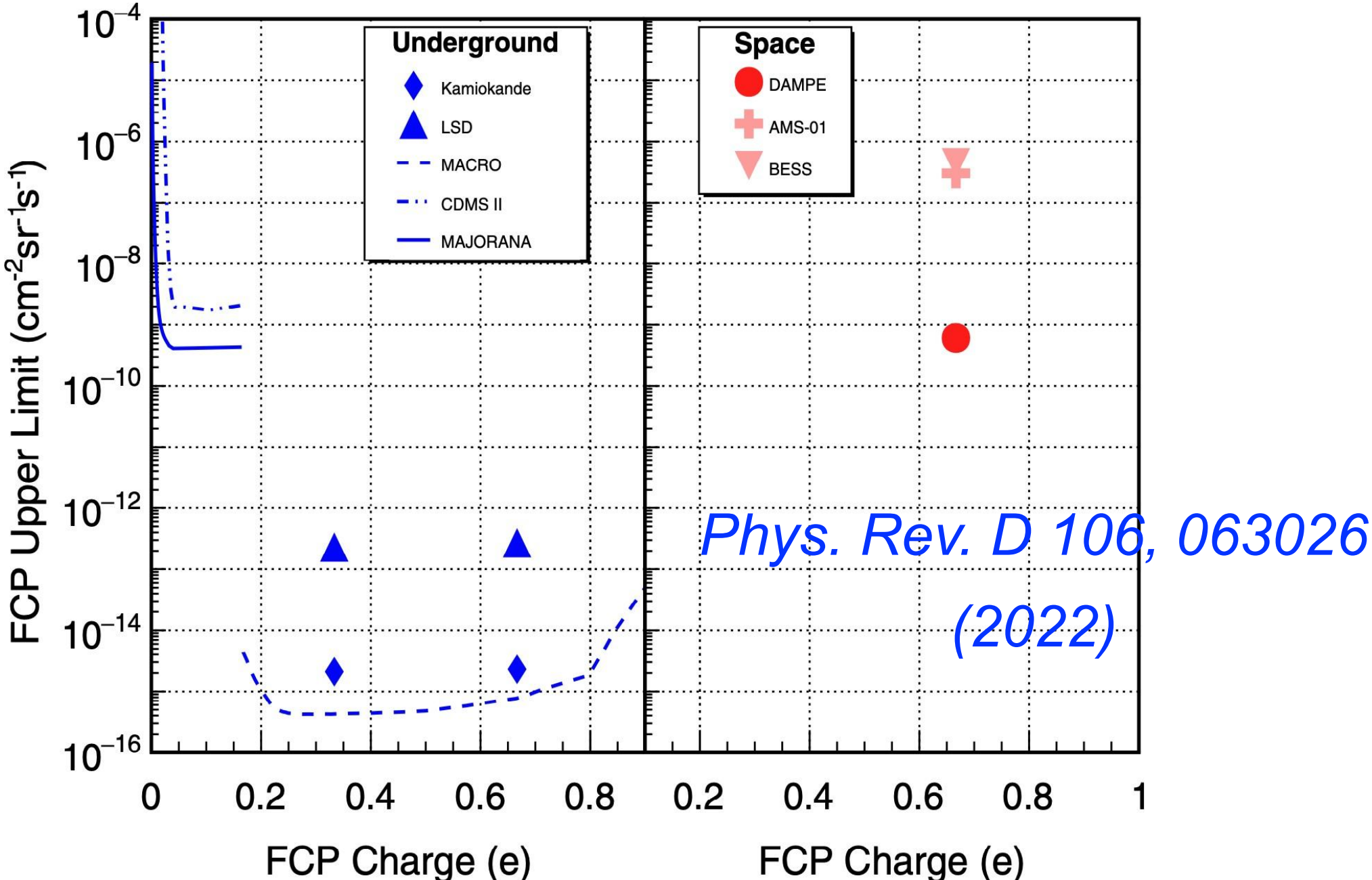
FCP Signal Region

Upper Limit of 2/3e FCP

TABLE I. The comparison between DAMPE and other similar types experiments.

Experiments	Geometric acceptance(cm ⁻² sr)	Exposure time (s)	Upper limit (cm ⁻² sr ⁻¹ s ⁻¹)
AMS-01	3000	3.6 × 10 ⁴	3.0 × 10 ⁻⁷ (95% CL)
BESS	1500	3.2 × 10 ⁵	4.5 × 10 ⁻⁷ (90% CL)
DAMPE	3000	2.3 × 10 ⁷	6.2 × 10 ⁻¹⁰ (90% CL)

$$\Phi < 6.2 \times 10^{-10} \text{cm}^{-2} \text{sr}^{-1} \text{s}^{-1}$$



- **Underground Experiment**
Energy loss when particle pass through rocks about **300 GeV** (1km depth)
 - **Space Experiment**
Cutoff by the earth's magnetic field **6 ~ 7 GeV**
- Lower kinetic energy limit for space experiment !**

Prospects for Searching for Light FCP

FCP should not be constrained to be heavy lepton

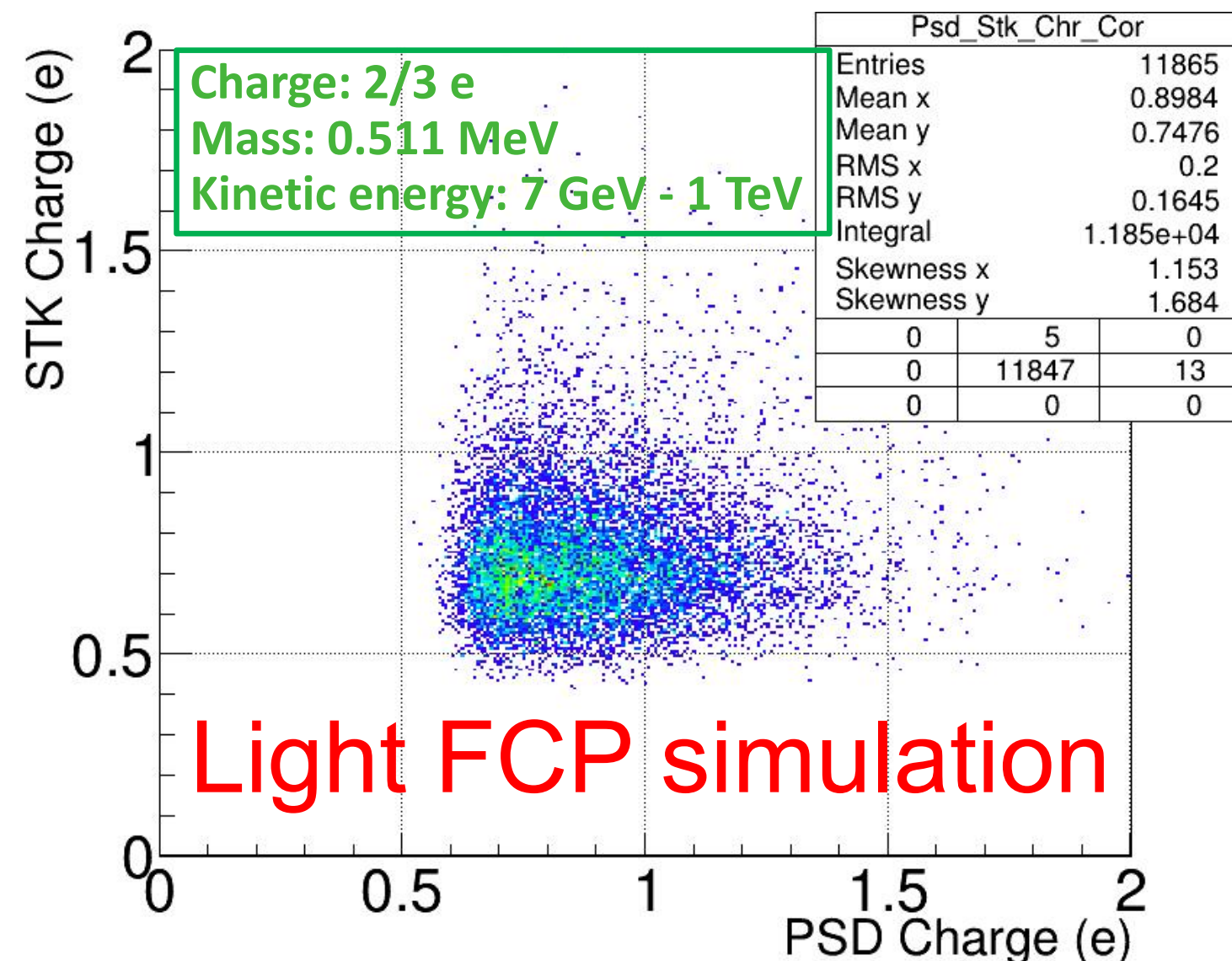
- Shower can happen
- Mass may be light
- Electron-like **light-mass particle**

Origin: nearby stars

Energy loss: **Bremsstrahlung**

$$-\frac{dE}{dx} = 4\alpha N_A \frac{Z^2}{A} z^2 \left(\frac{1}{4\pi\epsilon_0} \frac{e^2}{mc^2} \right)^2 E \ln \frac{183}{Z^{1/3}}$$

Based on the dataset accumulated by DAMPE, the **mass-, energy-dependent spectrum** are supposed to be observed.



Next step

- Light-mass FCP simulation.
- Evaluate the background contamination (electron, proton, gamma).
- Evaluate the selection efficiency and effective acceptance.

Summary

- We search for $2/3e$ FCP with DAMPE experiment
- Space experiments can detect FCPs with energy as low as a few GeV
- FCPs are assumed to be a type of heavy lepton
- No FCP signals are observed and a flux upper limit of $\Phi < 6.2 \times 10^{-10} \text{cm}^{-2} \text{sr}^{-1} \text{s}^{-1}$ is established at the 90% C.L..
- Result is published in [*Phys. Rev. D 106, 063026 \(2022\)*](#)
- Searching for Light FCP is on-going.

Thank you!