

# Dynamics and radiation of vortex electrons in magnetic fields

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# Transit of vortex electrons through magnetic fields





# **Scalar cat waves**



#### Plane waves

$$\Psi(p_x, p_y, p_z)$$



# Cylindrical waves $\Psi(p_z, p_\perp, l)$



# Spherical waves $\Psi(\omega, j, l)$



in vacuum

#### **Vortex electrons in vacuum**

 $\hbar = c$ 

$$\begin{split} & \left( \begin{array}{c} \hbar = c = 1 \\ \hline \partial \Psi(\mathbf{r}, t) \\ \partial t \end{array} \right) = \hat{\mathcal{H}}\Psi(\mathbf{r}, t) \\ & \left( \begin{array}{c} \partial \Psi(\mathbf{r}, t) \\ \partial t \end{array} \right) = \hat{\mathcal{H}}\Psi(\mathbf{r}, t) \\ & \left( \begin{array}{c} \partial \Psi(\mathbf{r}, t) \\ \partial t \end{array} \right) = \hat{\mathcal{H}}\Psi(\mathbf{r}, t) \\ & \left( \begin{array}{c} \partial \Phi(\mathbf{r}, t) \\ \partial t \end{array} \right) = \hat{\mathcal{H}}\Psi(\mathbf{r}, t) \\ & \left( \begin{array}{c} \partial \Phi(\mathbf{r}, t) \\ \partial t \end{array} \right) = \hat{\mathcal{H}}\Psi(\mathbf{r}, t) \\ & \left( \begin{array}{c} \partial \Phi(\mathbf{r}, t) \\ \partial t \end{array} \right) = \hat{\mathcal{H}}\Psi(\mathbf{r}, t) \\ & \left( \begin{array}{c} \partial \Phi(\mathbf{r}, t) \\ \partial t \end{array} \right) = \hat{\mathcal{H}}\Psi(\mathbf{r}, t) \\ & \left( \begin{array}{c} \partial \Phi(\mathbf{r}, t) \\ \partial \theta \end{array} \right) = \hat{\mathcal{H}}\Psi(\mathbf{r}, t) \\ & \left( \begin{array}{c} \partial \Phi(\mathbf{r}, t) \\ \partial \theta \end{array} \right) + \frac{1}{\rho^2} \frac{\partial^2 \Psi(\mathbf{r}, t)}{\partial \phi^2} + \frac{\partial^2 \Psi(\mathbf{r}, t)}{\partial z^2} \right) \\ & \left( \begin{array}{c} \partial \Phi(\mathbf{r}, t) \\ \partial \Phi(\mathbf{r}, t) \\ \partial \Phi(\mathbf{r}, t) \\ \partial \Phi(\mathbf{r}, t) \\ & \left( \begin{array}{c} \partial \Phi(\mathbf{r}, t) \\ \partial \theta \end{array} \right) + \frac{1}{\rho^2} \frac{\partial^2 \Psi(\mathbf{r}, t)}{\partial \phi^2} + \frac{\partial^2 \Psi(\mathbf{r}, t)}{\partial z^2} \right) \\ & \left( \begin{array}{c} \partial \Phi(\mathbf{r}, t) \\ \partial \Phi(\mathbf{r}, t) \\ \partial \Phi(\mathbf{r}, t) \\ \partial \Phi(\mathbf{r}, t) \\ & \left( \begin{array}{c} \partial \Phi(\mathbf{r}, t) \\ \partial \Phi(\mathbf{r}, t) \\ \partial \Phi(\mathbf{r}, t) \\ \partial \Phi(\mathbf{r}, t) \\ & \left( \begin{array}{c} \partial \Phi(\mathbf{r}, t) \\ \partial \Phi(\mathbf{r}, t) \\ \partial \Phi(\mathbf{r}, t) \\ \partial \Phi(\mathbf{r}, t) \\ & \left( \begin{array}{c} \partial \Phi(\mathbf{r}, t) \\ \partial \Phi(\mathbf{r}, t) \\ \partial \Phi(\mathbf{r}, t) \\ \partial \Phi(\mathbf{r}, t) \\ & \left( \begin{array}{c} \partial \Phi(\mathbf{r}, t) \\ \partial \Phi(\mathbf{r}, t) \\ \partial \Phi(\mathbf{r}, t) \\ \partial \Phi(\mathbf{r}, t) \\ & \left( \begin{array}{c} \partial \Phi(\mathbf{r}, t) \\ \partial \Phi(\mathbf{r}, t) \\ \partial \Phi(\mathbf{r}, t) \\ \partial \Phi(\mathbf{r}, t) \\ & \left( \begin{array}{c} \partial \Phi(\mathbf{r}, t) \\ \partial \Phi(\mathbf{r}, t) \\ \partial \Phi(\mathbf{r}, t) \\ \partial \Phi(\mathbf{r}, t) \\ & \left( \begin{array}{c} \partial \Phi(\mathbf{r}, t) \\ \partial \Phi(\mathbf{r}, t) \\ \partial \Phi(\mathbf{r}, t) \\ \partial \Phi(\mathbf{r}, t) \\ & \left( \begin{array}{c} \partial \Phi(\mathbf{r}, t) \\ \partial \Phi(\mathbf{r}, t) \\ \partial \Phi(\mathbf{r}, t) \\ \partial \Phi(\mathbf{r}, t) \\ & \left( \begin{array}{c} \partial \Phi(\mathbf{r}, t) \\ \partial \Phi(\mathbf{r}, t) \\ \partial \Phi(\mathbf{r}, t) \\ \partial \Phi(\mathbf{r}, t) \\ & \left( \begin{array}{c} \partial \Phi(\mathbf{r}, t) \\ \partial \Phi(\mathbf{r}, t) \\ \partial \Phi(\mathbf{r}, t) \\ \partial \Phi(\mathbf{r}, t) \\ & \left( \begin{array}{c} \partial \Phi(\mathbf{r}, t) \\ \partial \Phi(\mathbf{r}, t) \\ \partial \Phi(\mathbf{r}, t) \\ & \left( \begin{array}{c} \partial \Phi(\mathbf{r}, t) \\ \partial \Phi(\mathbf{r}, t) \\ \partial \Phi(\mathbf{r}, t) \\ & \left( \begin{array}{c} \partial \Phi(\mathbf{r}, t) \\ \partial \Phi(\mathbf{r}, t) \\ & \left( \begin{array}{c} \partial \Phi(\mathbf{r}, t) \\ \partial \Phi(\mathbf{r}, t) \\ & \left( \begin{array}{c} \partial \Phi(\mathbf{r}, t) \\ \partial \Phi(\mathbf{r}, t) \\ & \left( \begin{array}{c} \partial \Phi(\mathbf{r}, t) \\ & \left($$



# **Paraxial Laguerre-Gaussian packet**





# **Comparison of Bessel and LG beams**

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 $\varepsilon$  – energy

 $n_{\perp}$  – transverse momentum

Bessel vortex electrons



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$\Psi = N\Psi_{\parallel}(z)J_l(p_{\perp}\rho)e^{-i\varepsilon t + it\varphi}$	l - z-projection of OAM				
Paraxial LG vortex electrons					
$\Psi = N \frac{\rho^{ l }}{\sigma^{ l +1}(z)} L_n^{ l } \left(\frac{\rho^2}{\sigma^2(z)}\right) \exp\left[-\frac{\rho^2}{2\sigma^2(z)}\right]$	$\left \right] \exp\left[-i\varepsilon t + il\varphi + i \times \text{phase}(z)\right]$				
Nonstationary LG (NSLG	i) vortex electrons				
$\Psi = N\Psi_{\parallel}(z) \frac{\rho^{ l }}{\sigma^{ l +1}(t)} L_n^{ l } \left(\frac{\rho^2}{\sigma^2(t)}\right) \exp\left[-\frac{1}{2}\left(\frac{\rho^2}{\sigma^2(t)}\right)\right]$	$\frac{\rho^2}{2\sigma^2(t)} \right] \exp\left[il\varphi + i \times \text{phase}(t)\right]$				

	Bessel	Paraxial LG	NSLG
Definite OAM	$\checkmark$		
Normalizable	×		$\checkmark$
Exact solution	$\checkmark$	×	
Stationary	V		×

# NSLG wave packet in free space



\* Parameters are as in the work G. Guzzinati, et al. Phys. Rev. Lett., **110**, 093601, 2013 Observation of the Larmor and Gouy rotations with electron vortex beams





- Keeping vortex electrons from spreading
- Controlling parameters of vortexelectron wave packets
- Accounting for boundary conditions

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#### **Problem statement**





#### Two approaches







Liping Zou, Pengming Zhang, and Alexander J. Silenko Phys. Rev. A **103**, L010201 – Published 6 January 2021

Paraxial stationary Schrödinger equation in magnetic field 
$$\left(\nabla_{\perp}^{2} - ieH\frac{\partial}{\partial\phi} - \frac{e^{2}H^{2}r^{2}}{4} + 2es_{z}H + 2ik\frac{\partial}{\partial z}\right)\Psi = 0$$

$$\begin{pmatrix} \text{Nonstationary Schrödinger equation in magnetic field} \\ \left( \nabla_{\perp}^2 - ieH \frac{\partial}{\partial \phi} - \frac{e^2 H^2 r^2}{4} + 2im \frac{\partial}{\partial t} \right) \Psi = 0 \end{pmatrix}$$

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## System of optical equations



Paraxial stationary Schrödinger equation in magnetic field

$$\left(\nabla_{\perp}^{2} - ieH\frac{\partial}{\partial\phi} - \frac{e^{2}H^{2}r^{2}}{4} + 2im\frac{\partial}{\partial t}\right)\Psi = 0$$

$$\begin{aligned} & \frac{1}{R(t)} = \frac{\sigma'(t)}{\sigma(t)}, \\ & \frac{1}{\lambda_{\rm C}^2 R^2(t)} + \frac{1}{\lambda_{\rm C}^2} \left[ \frac{1}{R(t)} \right]' = \frac{1}{\sigma^4(t)} - \frac{1}{\sigma_{\rm L}^4}, \\ & \frac{1}{\lambda_{\rm C}} \Phi_{\rm G}'(t) = \frac{l}{\sigma_{\rm L}^2} + \frac{2n + |l| + 1}{\sigma^2(t)}. \end{aligned} \qquad \begin{array}{l} \lambda_{\rm C} = m^{-1} - \\ & \text{Compton wavelength} \\ & \sigma_{\rm L} = \sqrt{2/|eH|} - \\ & \text{magnetic length} \end{aligned}$$

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# **Initial conditions**





An NSLG vortex electron enters a magnetic field region and propagates inside it,  $z_g$  – location of generating device,  $z_0$  – position of the boundary dividing vacuum and the magnetic field region.

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#### **NSLG state in the field**

$$\begin{aligned} & \mathsf{Dispersion} \\ & \sigma(t) = \sigma_{\rm st} \sqrt{1 + \sqrt{1 - \left(\frac{\sigma_{\rm L}}{\sigma_{\rm st}}\right)^4} \sin\left[s(\sigma_0, \sigma_0')\omega(t - t_0) - \theta\right]} \end{aligned}$$

Rms radius  

$$\rho(t) \equiv \sqrt{\langle \rho^2 \rangle - \langle \rho \rangle^2} = \sigma(t)\sqrt{2n + |l| + 1}$$

$$\sigma_{\rm st} \ge \sigma_{\rm L}$$
  $\rightarrow$   $\rho_{\rm st} = \sqrt{(2n + |l| + 1)}\sigma_{\rm st} \ge \rho_{\rm L}$ 



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#### **Rms radius oscillations**



Rms radius oscillations (in red).  $\rho_L = 52.7$  nm (in blue). Black dashed lines correspond to the square of the period-averaged transverse radius squared. Magnetic field strength H=1.9 T, n=0, l=3,  $\rho'_0 = 0$ . a)  $\rho_0 = 54$  nm, b)  $\rho_0 = 25$  nm, c)  $\rho_0 = 111.1$  nm, d)  $\rho_0 = 1 \ \mu$ m.



Transmission of

VEs through magnetic fields - VE in vacuum - Nonstationary states approach - Properties of the electron state in the field Radiation of VEs in magnetic fields - Evolved-state approach - Emitted photons' properties

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# **Experiment confirms?**





\* From the work P. Schattschneider et al., Nature Comm., **5**:4586, 2014 Imaging the dynamics of free-electron Landau states.

## **Experimental feasibility**







# Radiation of vortex electrons in magnetic fields





### **Motivation**





# **Applications**



- Generation of vortex photons in laboratory Transmission of VEs through Astrophysics: neutron stars et al.  $H \leq H_c = 4.4 \times 10^9 T_c$ magnetic fields - VE in vacuum - Nonstationary states approach A vortex photon in a star is born - Properties of the electron state in the field Radiation of VEs The photon interacts with matter in magnetic fields (not like a plane-wave photon!): - Evolved-state
  - Higher Landau states can be excited
  - Multipole transitions in atoms become significant
  - Vortex cross-section ≠ plane-wave cross-section

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approach - Emitted

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#### State of the system





# **Disentangling the final state**



Final electron is measured:  $|f_e
angle|$ 

Photon is **not** measured

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#### **Evolved state of the photon**



Final disentangled state 
$$|f'
angle = |f_e
angle \otimes |\gamma
angle_{ev}$$

$$\begin{aligned} & \left( \begin{aligned} & \text{State of the photon} \\ & |\gamma\rangle_{ev} = \sum_{\lambda=\pm 1} \int \frac{d^3k}{(2\pi)^3} \frac{1}{2\omega} S^{(1)}_{fi} | \mathbf{k}, \lambda \right) \end{aligned} \end{aligned}$$

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# Vector potential of the photon



$$\mathbf{A}(t, \mathbf{r}) = \langle 0 | \hat{\mathbf{A}}(t, \mathbf{r}) | \gamma \rangle_{ev} = \sum_{\lambda} \int \frac{d^3k}{(2\pi)^3} \frac{1}{\sqrt{2\omega}} \mathbf{A}_{\lambda}(t, \mathbf{r}; \mathbf{k}) S_{fi}^{(1)}$$

$$\boldsymbol{A}_{\lambda}(t,\boldsymbol{r};\boldsymbol{k}) = \frac{\boldsymbol{e}_{\lambda}(\boldsymbol{k})}{\sqrt{2\omega}} e^{-i(\omega t - \boldsymbol{k}\cdot\boldsymbol{r})}$$

Vector potential in the momentum representation  $A(m{k}) = \sum_{\lambda=\pm 1} S_{fi}^{(1)} m{e}_{\lambda}(m{k})$ 

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$$\begin{split} & \Psi_{s,\ell}^{\uparrow}(x) = N^{\uparrow} \begin{pmatrix} (m+\varepsilon) \Phi_{s,\ell-1/2}(\rho) e^{-i\varphi/2} \\ 0 \\ p_z \Phi_{s,\ell-1/2}(\rho) e^{-i\varphi/2} \\ -ieH \Phi_{s,\ell+1/2}(\rho) e^{i\varphi/2} \end{pmatrix} e^{-it\varepsilon_{s,\ell}+i\ell\varphi+ip_z z} \end{split}$$

$$\left( \begin{split} & \text{S-matrix element} \\ S_{fi}^{(1)} = \delta(k_{\perp} - \kappa) \delta(k_z + p'_z - p_z) e^{i(\ell - \ell')\varphi_k} \mathcal{F} \end{split} \right)$$



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#### **Photon state**





# **Emission probability and intensity**

$$\begin{split} \left\langle \gamma | \gamma \right\rangle = \sum_{\lambda = \pm 1} \int \frac{d^3 \mathbf{k}}{(2\pi)^3} \left| S_{fi}^{(1)} \right|^2 \equiv W_{s',\ell'}^{(1)}(p'_z) \end{split}$$

$$\begin{array}{l} & \text{Intensity} \\ I_{s',\ell'}^{(1)} \equiv \int L \frac{dp'_z}{2\pi} \; \omega \dot{W}^{(1)}_{s',\ell'}(p'_z) \end{array} \end{array}$$

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#### **Evolved-state approach**





Probability and intensity of emission with definite angular momentum (with no spin-flip).  $H = 10^{-2}H_c, \ s = s' = 5, \ p_z = 10^{-3}m$  $H_c = 4.4 \times 10^9T$ 

# Conclusion

- NSLG states provide a single-mode description of a vortex electron transfer through a solenoid
- The NSLG states approach enable to easily account for the boundary conditions
- NSLG states highlight the dynamics of vortex electrons on magnetic fields
- Electrons in a constant magnetic field spontaneously emit twisted photons
- Total angular momentum is conserved in this process.
- The majority of the emitted photons turn out to be twisted, with  $\ell-\ell'\gtrsim 1$
- The obtained results can offer a new perspective on the synchrotron radiation, generation of twisted photons and astrophysical processes in extreme magnetic field.



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Sizykh, G. K., Chaikovskaia, A. D., Grosman, D. V., Pavlov, I. I., & Karlovets, D. V. (2024). Transmission of vortex electrons through a solenoid. *Physical Review A*, *109*(4), L040201.



Pavlov, I., & Karlovets, D. (2024). Emission of twisted photons by a Dirac electron in a strong magnetic field. *Physical Review D*, *109*(3), 036017.

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# Why study vortex electrons?



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# How to produce VE?



\* From Knyazev, B. A., & Serbo, V. G. (2018). Beams of photons with nonzero projections of orbital angular momenta: new results. *Physics-Uspekhi*, *61*(5), 449.

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#### Gouy phase of a free wave packet

$$\Phi_G(t) = (2n + |l| + 1) \arctan\left(\frac{t}{t_d}\right)$$

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#### **NSLG state in the field**

#### General form of an NSLG state

$$\Psi_{n,l}(\boldsymbol{\rho},t) = N \frac{\rho^{|l|}}{\sigma^{|l|+1}(t)} L_n^{|l|} \left(\frac{\rho^2}{\sigma^2(t)}\right) \exp\left[il\varphi - i\Phi_G(t) - \frac{\rho^2}{2\sigma^2(t)} \left(1 - i\frac{\sigma^2(t)}{\lambda_C R(t)}\right)\right]$$

Dispersion

$$\sigma(t) = \sigma_{\rm st} \sqrt{1 + \sqrt{1 - \left(\frac{\sigma_{\rm L}}{\sigma_{\rm st}}\right)^4}} \sin\left[s(\sigma_0, \sigma_0')\omega(t - t_0) - \theta\right]}$$

Stationary dispersion

$$\sigma_{\rm st}^2 = \frac{\sigma_0^2}{2} \left( 1 + \left(\frac{\sigma_{\rm L}}{\sigma_0}\right)^4 + \left(\frac{\sigma_0' \sigma_{\rm L}^2}{\lambda_{\rm C} \sigma_0}\right)^2 \right)$$

Initial phase

$$heta = rcsin rac{1 - (\sigma_0/\sigma_{
m st})^2}{\sqrt{1 - (\sigma_{
m L}/\sigma_{
m st})^4}}$$

Sign function  $s(\sigma_0, \sigma'_0) = \begin{cases} \operatorname{sgn}(\sigma'_0), \ \sigma'_0 \neq 0, \\ \operatorname{sgn}(\sigma_{\mathrm{L}} - \sigma_0), \ \sigma'_0 = 0, \\ 0, \ \sigma_0 = \sigma_{\mathrm{L}} \text{ and } \sigma'_0 = 0. \end{cases}$ 

$$\begin{split} & Gouy \, \text{phase} \\ \Phi_{\rm G}(t) = \Phi_0 + \frac{l\omega(t-t_0)}{2} + (2n+|l|+1)s(\sigma_0,\sigma_0') \\ \times \left[ \arctan\left(\frac{\sigma_{\rm st}^2}{\sigma_{\rm L}^2} \tan \frac{s(\sigma_0,\sigma_0')\omega(t-t_0) + \theta}{2} + \frac{\sigma_{\rm st}^2}{\sigma_{\rm L}^2}\sqrt{1 - \left(\frac{\sigma_{\rm L}}{\sigma_{\rm st}}\right)^4}\right) \\ & - \arctan\left(\frac{\sigma_{\rm st}^2}{\sigma_{\rm L}^2} \tan \frac{\theta}{2} + \frac{\sigma_{\rm st}^2}{\sigma_{\rm L}^2}\sqrt{1 - \left(\frac{\sigma_{\rm L}}{\sigma_{\rm st}}\right)^4}\right) \right]. \end{split}$$



Curvature radius

 $R(t) = \frac{\sigma(t)}{\sigma'(t)}$ 

#### **Gouy phase**





Figure 4: Gouy phase  $\Phi_{\rm G}(t)$  of the NSLG<sub>H</sub> wave packet. The field strength H = 1.9 T (corresponding  $\rho_{\rm L} \approx 37$  nm),  $\rho_0 \approx 71$  nm,  $\rho'_0 = 0$ ,  $T_{\rm c} \approx 0.02$  ns, and  $\Phi_0 = 0$ . The quantum numbers are: n = 0, l = 0 (red line); n = 0, l = 1 (blue); and n = 1, l = 1 (green).

#### **Decomposition into Landau states**



n = 2, l = 35, (l) n = 3, l = 35.

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work [17]. The parameters of the NSLG packet are: energy  $\varepsilon = 200$  KeV (corresponding velocity  $v \approx 0.7c$ ),  $n = 0, l = 1, \rho_0 \approx 67.5$  nm,  $\rho'_0 \approx -4.4 \times 10^{-4}$ . Magnetic field strength B = 1.9 T (corresponding  $\sigma_L \approx 26$  nm).

### **Oscillations parameters**





Setup	$E_{\parallel}$	v	H	$ ho_{ m L}$	d	$z_{ m R}$	$ ho_0$	$ d\rho/dz _{z=z_0}$	$\xi_1$	$\xi_2$
SEM	100  eV	0.02c	1 T	72.6 nm	$5.16 \mathrm{~cm}$	$5.16 \mathrm{~cm}$	$2.82 \ \mu m$	$27~\mathrm{pm}/\mathrm{\mu m}$	0.025	$6.6 \times 10^{-4}$
TEM	200  keV	0.70c	1.9 T	52.7 nm	10 cm	179 cm	$2 \ \mu \mathrm{m}$	62  pm/mm	0.026	$3.9 \times 10^{-5}$
Medical linac	1 MeV	0.94c	0.1 T	$0.23~\mu{ m m}$	10 cm	243 cm	$2 \ \mu \mathrm{m}$	0.34  nm/cm	0.115	$5.5 \times 10^{-4}$
Linac	$1 { m GeV}$	c	0.01 T	$0.72~\mu{ m m}$	100 cm	$258 \mathrm{~cm}$	$2.14 \ \mu \mathrm{m}$	$0.28 \ \mu \mathrm{m/m}$	0.339	0.045

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#### **Transversely relativistic motion**



Free NSLG: 
$$2n + |l| + 1 \ll \frac{\rho_{\rm w}}{\lambda_{\rm C}}$$

However, being refocused to a 1 nm waist size, electrons with quantum numbers of the order of  $10^3$  become transversely relativistic.

Landau: 
$$\sqrt{(2n+|l|+l+1)} \ll \frac{\sigma_{\rm L}}{\lambda_{\rm C}}$$

such a state remains nonrelativistic for any attainable values of nand |l|. However, for l > 0, the relativistic regime cannot be achieved either, as it would require OAM of the order of  $10^{10}$ .

NSLG in a field: 
$$\sqrt{\left[(2n+|l|+1)\frac{\sigma_{\rm st}^2}{\sigma_{\rm L}^2}+l\right]} \ll \frac{\sigma_{\rm L}}{\lambda_{\rm C}} \qquad \sqrt{(2n+|l|+1)} \ll \frac{\sigma_{\rm L}}{\lambda_{\rm C}} \sigma_{\rm L}.$$

From here it follows that for wave packets with  $\sigma_0/\sigma_L \ge \sigma_L/\lambda_C$ , even a Gaussian mode with n = l = 0 is relativistic. For a field strength of the order of 1 T, this happens when  $\sigma_0 \sim 1$  mm, which can also be decreased if the divergence rate  $\sigma'_0$  in (34) is taken into account.

#### Slight misalignment





Figure 7: Free twisted electron entering a magnetic lens at a small angle  $\alpha$  with respect to the field direction.

$$\begin{aligned} |c_{nn'll'}| = &\delta_{n,n'}\delta_{l,l'} + \frac{\alpha \langle p_z \rangle \sigma(t_0)}{4\pi} \delta_{|l'|,|l|-1} \left[ \delta_{n',n}\sqrt{n+|l|} + \delta_{n',n+1}\sqrt{n+1} \right] \\ &+ \frac{\alpha \langle p_z \rangle \sigma(t_0)}{4\pi} \delta_{|l'|,|l|+1} \left[ \delta_{n',n}\sqrt{n+|l|+1} + \delta_{n',n-1}\sqrt{n} \right]. \end{aligned}$$

From Eq. (44), we see that the actual dimensionless parameter defining the magnitude of the coefficients is  $\alpha \langle p_z \rangle \sigma(t_0)$ . In real life, the value of  $\sigma(t_0)$  is of the order of several  $\mu$ m or less. Provided that currently  $n \sim 1$ ,  $l \leq 10^4$ , even for 10 GeV-electrons with  $\langle p_z \rangle \sim 10^{-3} \mu \text{m}^{-1}$ , we obtain  $|c_{nn'll'}| \leq 10^{-2} \alpha$ . This means that the off-axis corrections are negligible for any feasible experimental scenario.

#### Quantum rms emittance





Figure 9: Emittances of the NSLG<sub>H</sub> (in red), the Landau (in blue), and the NSLG<sub>f</sub> (in green) states. B = 1.9 T, n = 0,  $\sigma_0 = 25$  nm,  $\sigma' = 0$ . (a) l = -3, (b) l = 3.



# **Classical radiation of vortex photons**





Epp V., Guselnikova U., Angular momentum of radiation from a charge in circular and spiral motion, Physics Letters A, 2019

#### **Relativistic Landau states**





#### Photon transverse momentum

Transverse momentum of the photon 
$$\kappa = \sqrt{(\varepsilon - \varepsilon')^2 - (p_z - p_z')^2} \ge 0$$



## **Radiation in the critical field**





Figure 1: The emission probability (32) (left) and the corresponding intensity (33) (right) for  $H = H_c$ ,  $p_z = 10^{-3}mc$ , and no spin-flip. For the solid lines s = s' = 20, the dashed lines correspond to the twisted photons with a simultaneous change of the radial quantum number  $s \to s' \neq s$ , the dash dotted lines correspond to the untwisted photons with the TAM  $j_z = \ell - \ell' = 0$ .

#### More radiation in the critical field



Figure 2: The same as in Fig 1, but for  $H = 0.1H_c$  and s = 5.

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#### **Sokolov-Ternov effect**





Figure 4: The ratio of four possible types of transition probabilities to the probability of emission by a scalar charge derived in [15].  $H = 10^{-3}H_c$  (left) and  $H = H_c$  (right);  $\ell - \ell' = 3$ , s = s' = 20 for all transitions.



#### **Dependence on** *p*<sub>*z*</sub>



Figure 5: The dependence of the emission probability (left) and the intensity (right) on the electron momentum  $p_z$  for  $H = 0.1H_c$ , s = s' = 20. The transition  $20\frac{1}{2} \rightarrow 19\frac{1}{2}$  means  $\ell = 20\frac{1}{2}$ ,  $\ell' = 19\frac{1}{2}$ , s = s' = 20; those with  $\ell : 20\frac{1}{2} \rightarrow 20\frac{1}{2}$  correspond to the untwisted photons with  $j_z = 0$ . The green line overlaps with the pink dashed one on the left; the cyan line on the left overlaps with the blue one on the right. The magenta dash-dotted line corresponds to increase of the electron OAM during the emission (so that the photon TAM is  $\ell - \ell' = -1$ ).