# EW corrections at future colliders



2024 International Workshop on the High Energy CEPC Hangzhou, China 24-10-2024

## Davide Pagani



Grazzini, Kallweit, Lindert, Pozzorini, Wiesemann '19

$$p \to \ell^- \ell^+ \nu_{\ell'} \bar{\nu}_{\ell'} \qquad pp \to \ell^- \ell'^+ \nu_{\ell'} \bar{\nu}_{\ell'}$$

2

QCD) + SMEFT@LO and SM@(NLO QCD + approximate NLO EW) + SMEFT@LO. The expectation assumption is the same, SM@(NLO QCD + approximate NLO EW) for both theory assumptions. The other parameters have not been marginalised over and we consider  $\Delta \chi^2 = 3.84$ for a one-parameter  $\chi^2$  fit to present our allowed regions.

 $pp \rightarrow \ell^- \ell^+ \ell' \nu_{\ell'}$  LHC  $\sqrt{s} = 13 \,\text{TeV}$ 



## **Future Colliders vs. EW corrections**

### (1) Higher Precision



EW corrections become even more relevant. NLO EW is not sufficient and higher orders are necessary.

### (2) Higher Energy



EW corrections become larger (Sudakov).

EW corrections become relevant both for signal/bkg or BSM/SM, not only for precision studies.

For Precision: see (1).





# Future Colliders vs. EW corrections

### (1) Higher Precision



EW corrections become even more relevant. NLO EW is not sufficient and higher orders are necessary.

## **EW** is the new **QCD**

corrections cannot he neglected at futures collider.

### (2) Higher Energy



EW corrections become larger (Sudakov).

EW corrections become relevant both for signal/bkg or BSM/SM, not only for precision studies.

For Precision: see (1).





# Higher Precision



# **Precision: the Higgs case**

### Sub-percent precision is expected and so NLO EW is not enough.

### Contributions from fermion loops at NNLO



Small corrections, but necessary for matching the experimental precision.  $| \mathcal{M}_{\mathcal{A}} |$ The choice of the renormalisation scheme is relevant. +

This is only one of the effects e.g. PDFs and ISR are not taken into account.

 $e^+e^- \rightarrow ZH$ : NNLO EW  $\mathcal{O}(\alpha^2)$ 

Freitas, Song '22 Freitas, Song, Xie '24

	$\alpha(0)$ scheme	$G_{\mu}$ scheme
$\sigma^{\rm LO}$ [fb]	222.96	239.18
$\sigma^{\rm NLO}$ [fb]	+3.1% 229.89	-2.9% 232.08
$\sigma^{\rm NNLO}$ [fb]	+0.7% 231.55	+0.3% 232.74
$\mathcal{O}(\alpha_{N_f=2}^2)$	1.88	0.73
$\mathcal{O}(\alpha_{N_f=1}^{2^{*}})$	-0.23	-0.07

 $\sqrt{S} = 240 \text{ GeV}$ 



# $e^+e^- \rightarrow ZH$ : mixed NNLO QCD-EW $\mathcal{O}(\alpha_s \alpha)$

Gong, Li, Xu, Yang, Zhao '17 Sun, Feng, Jia, Sang '24

TABLE I.Total cross sections at various collider energies in the $\overline{\text{MS}}$  scheme.

	$\sqrt{s}$ (GeV)	$\sigma_{ m LO}~({ m fb})$	$\sigma_{\rm NLO}~({\rm fb})$	$\sigma_{\rm NNLO}~({\rm fb})$	$\sigma^{ m exp}_{ m NNLO}$
+1.1%	240	256.3(9)	228.0(1)	230.9(4)	230.9(4
+1.1%	250	256.3(9)	227.3(1)	230.2(4)	230.2(4
+1.1%	300	193.4(7)	170.2(1)	172.4(3)	172.4(
+1.3%	350	138.2(5)	122.1(1)	123.9(2)	123.6(2
+0.6%	500	61.38(22)	53.86(2)	54.24(7)	54.64

TABLE II. Total cross sections at various collider energies in the  $\alpha(m_Z)$  scheme.

$\perp$ $22222222222222222222222222222222222$		$\sqrt{s}$ (GeV)	$\sigma_{ m LO}~({ m fb})$	$\sigma_{ m NLO}~( m fb)$	$\sigma_{ m NNLO}$ (fb)	$\sigma^{ m exp}_{ m NNLO}$
	+1.1%	240	252.0	228.6	231.5	231.
	+1.1%	250	252.0	227.9	230.8	230.
	+1.1%	300	190.0	170.7	172.9	172.
$+ \mathbf{w}$	+1.3%	350	135.6	122.5	124.2	124.
7	+0.6%	500	60.12	54.03	54.42	54.





# $e^+e^- \rightarrow ZH$ : mixed NNLO QCD-EW $\mathcal{O}(\alpha_c \alpha)$

Gong, Li, Xu, Yang, Zhao '17 Sun, Feng, Jia, Sang '24

Convergence of the  $1/m_t^2$  expansion for the mixed TABLE III. QCD-EW corrections in the  $\overline{\text{MS}}$  scheme with  $\mu = \sqrt{s}/2$ .

$\sqrt{s}$ (GeV)	$\mathcal{O}(m_t^2)$	$\mathcal{O}(m_t^0)$	$\mathcal{O}(m_t^{-2})$	$\mathcal{O}(m_t^-)$
240	81.8%	16.2%	1.4%	0.49
250	81.7%	16.1%	1.5%	0.5%
300	80.0%	15.2%	2.1%	1.19
350	69.7%	12.6%	2.7%	2.19
500	137%	18.6%	17.3%	31.19

Expansion in powers of  $m_t$  works well only for energies up to  $2m_t$ 





### but necessary for Small matching the experimental precision.



Bertone, Cacciari, Frixione, Stagnitto, Zaro, Zhao '22' Bertone, Cacciari, Frixione, Stagnitto '19 Frixione '19



ISR radiation resummed within the PDFs of e in the e. NLO EW with NLL and LL PDFs can have ~ 1% differences: small but necessary for matching the 1% precision. Renormalisation scheme relevant too, factorisation scheme less relevant for x-sections.

PDFs at NLL accuracy in QED

$$\sigma(\tau_{min}) = \int d\sigma \,\Theta\left(\tau_{min} \le \frac{M_{p\bar{p}}^2}{s}\right) \,, \qquad p = q \,, t \,, W^2$$

# PDFs at NLL accuracy in QED

Bertone, Cacciari, Frixione, Stagnitto, Zaro, Zhao '22' Bertone, Cacciari, Frixione, Stagnitto '19 Frixione '19



ISR radiation resummed within the PDFs of e in the e. But  $\gamma$  in the e too! Regardless of the scheme used, photon initiated processes can be relevant: small but necessary for matching the 1% precision.

$$\sigma(\tau_{min}) = \int d\sigma \,\Theta\left(\tau_{min} \le \frac{M_{p\bar{p}}^2}{s}\right) \,, \qquad p = q \,, t \,, W^+$$





Process	$\lambda_3$	$\lambda_4$
$ZH, \nu_e \bar{\nu}_e H \text{ (WBF)}$	one-loop	two-loop
$ZHH, \nu_e \bar{\nu}_e HH \text{ (WBF)}$	tree	one-loop
$ZHHH, \nu_e \bar{\nu}_e HHH \text{ (WBF)}$	tree	tree

### HHH from single-H at one loop





$$\kappa_3 \equiv \frac{\lambda_3}{\lambda_3^{\rm SM}} = 1 + \frac{c_6 v^2}{\lambda \Lambda^2} \equiv 1 + \bar{c}_6,$$

### Maltoni, DP, Zhao '18



### HHHH from double-H at one loop







$$\kappa_{4} \equiv \frac{\lambda_{4}}{\lambda_{\mu^{2}}} = 1 + \frac{6c_{6}v^{2}}{\sqrt{\lambda_{\mu^{2}}}} + \frac{4c_{8}v^{4}}{\sqrt{\lambda_{4}}} = 1 + \frac{6c_{6}c_{6}}{\sqrt{\lambda_{4}}} + \frac{6c_{6}v^{2}}{\sqrt{\lambda_{4}}} + \frac{4c_{8}v^{4}}{\sqrt{\lambda_{4}}} = 1 + \frac{6c_{6}c_{6}}{\sqrt{\lambda_{4}}} + \frac{6c_{6}v^{2}}{\sqrt{\lambda_{4}}} +$$

12



# HHH from single Higgs

	$\sqrt{\hat{s}}  [{ m GeV}]$	process	$\epsilon$ [%]	$C_1 \ [\%]$	$\bar{c}_6(\pm 1\sigma)$	$\bar{c}_6(\pm 2\sigma)$
CEPC	250	ZH	0.51	1.6	$(-0.38, 0.42) \cup (8.0, 8.8)$	$(-0.73, 0.88) \cup (7.5, 9.1)$
	240	ZH	0.4	1.8	$(-0.26, 0.28) \cup (9.4, 9.9)$	$(-0.51.0.57) \cup (9.1, 10.2)$
FCC-ee	240	WBF $H$	2.2	0.66	(-2.81, 5.1)	(-4.3, 6.6)
	350	WBF $H$	0.6	0.65	(-1.15, 3.4)	(-1.89, 4.1)
	250	ZH	0.71	1.6	$(-0.52, 0.59) \cup (7.8, 8.9)$	$(-0.98, 1.3) \cup (7.1, 9.4)$
ILU	500	WBF $H$	0.23	0.63	(-0.56, 2.7)	(-0.97, 3.1)
	1000	WBF $H$	0.33	0.61	(-0.78, 2.7)	(-1.3, 3.3)
	350	ZH	1.65	0.59	(-2.48, 4.3)	(-3.80, 5.6)
CLIC	1400	WBF $H$	0.4	0.61	(-0.91, 2.9)	(-1.50, 3.5)
	3000	WBF $H$	0.3	0.59	(-0.75, 2.6)	(-1.26, 3.1)

### Lower energy, ZH larger and $C_1$ too.



$$\sigma_{\rm NLO}^{\rm pheno}(H) = \sigma_{\rm LO} + \sigma_1 \bar{c}_6 + \sigma_2$$

$$\delta\sigma(H) \equiv \frac{\sigma_{\rm NLO}^{\rm pheno} - \sigma_{\rm LO}}{\sigma_{\rm LO}} = \frac{\sigma_1 \bar{c}_6 + \sigma_2 \bar{c}_6^2}{\sigma_{\rm LO}} = (\kappa_3 - 1)C_1 + (\kappa_3 - 1)C_1 +$$

Maltoni, DP, Zhao '18 see also McCullough '13



### $(\kappa_3^2 - 1)C_2$ ,

Other operators enter via EW loop corrections. An example here:

 $(\varphi^{\dagger}\varphi)^{3}$ but also  $(\varphi^{\dagger}i\overleftrightarrow{D}_{\mu}\varphi)(\bar{u}_{p}\gamma^{\mu}u_{r})$  $(\bar{l}_p \gamma_\mu l_r) (\bar{q}_s \gamma^\mu q_t)$ 



 $Z/\gamma$ 

**Figure 9**. Contributions from modifications of the Higgs tri-linear coupling  $C_{\phi}$  on the cross-section for  $e^+e^- \to ZH$  correlated with those from  $C_{\phi u}[3,3]$ , which modifies the  $Zt\bar{t}$  vertex, and from  $C_{la}^{(1)}[1,1,3,3]$  vertex which modifies the  $e^+e^-t\bar{t}$  interaction. The sensitivity to a 0.5% measurement at  $\sqrt{s} = 240$  GeV is shown. Note that there is no sensitivity to  $C_{\phi}$ ,  $C_{\phi u}[3,3]$  or  $C_{la}^{(1)}[1,1,3,3]$  at tree level.

# Higher Energy (muon collider / 100 TeV pp)

# NLO EW corrections at high energies

NLO EW corrections for energies of the order of few TeVs are as large as (or even more than) NLO QCD corrections at the LHC. Origin: **EW Sudakov logarithms.** 

EW corrections should be considered not only for precision physics, since they give  $\mathcal{O}(10 - 100\%)$  effects. This includes also BSM scenarios.

$\mu^+\mu^- \to X, \sqrt{s} = 3 \text{ TeV}$	$\sigma_{ m LO}^{ m incl} \ [{ m fb}]$	$\sigma_{ m NLO}^{ m incl}$ [fb]	$\delta_{ m EW}~[\%]$
$W^+W^-Z$	$3.330(2)\cdot 10^1$	$2.568(8) \cdot 10^1$	-22.9(2)
$W^+W^-H$	$1.1253(5)\cdot 10^{0}$	$0.895(2)\cdot 10^{0}$	-20.5(2)
ZZZ	$3.598(2)\cdot 10^{-1}$	$2.68(1) \cdot 10^{-1}$	-25.5(3)
HZZ	$8.199(4) \cdot 10^{-2}$	$6.60(3) \cdot 10^{-2}$	-19.6(3)
HHZ	$3.277(1)\cdot 10^{-2}$	$2.451(5) \cdot 10^{-2}$	-25.2(1)
HHH	$2.9699(6) \cdot 10^{-8}$	$0.86(7) \cdot 10^{-8}$ *	
$W^+W^-W^+W^-$	$1.484(1) \cdot 10^0$	$0.993(6) \cdot 10^0$	-33.1(4)
$W^+W^-ZZ$	$1.209(1)\cdot 10^{0}$	$0.699(7)\cdot 10^{0}$	-42.2(6)
$W^+W^-HZ$	$8.754(8) \cdot 10^{-2}$	$6.05(4) \cdot 10^{-2}$	-30.9(5)
$W^+W^-HH$	$1.058(1)\cdot 10^{-2}$	$0.655(5) \cdot 10^{-2}$	-38.1(4)
ZZZZ	$3.114(2) \cdot 10^{-3}$	$1.799(7) \cdot 10^{-3}$	-42.2(2)
HZZZ	$2.693(2)\cdot 10^{-3}$	$1.766(6) \cdot 10^{-3}$	-34.4(2)
HHZZ	$9.828(7) \cdot 10^{-4}$	$6.24(2) \cdot 10^{-4}$	-36.5(2)
HHHZ	$1.568(1)\cdot 10^{-4}$	$1.165(4) \cdot 10^{-4}$	-25.7(2)

### 3 TeV Muon Collider

WHIZARD Bredt, Kilian, Reuter, Steinemeier '22



### How large are expected to be the EW Sudakov at 1 loop?

$$\mathcal{O}(1) \to \frac{\alpha}{4\pi s_w^2} \sim 0.3 \,\%, \quad \text{Single Log} \to \frac{\alpha}{4\pi s_w^2} \log(s/m_W^2),$$
  
Double Log  $\to \frac{\alpha}{4\pi s_w^2} \log^2(s/m_W^2)$ 



The estimate done via the variation of a factor of 10 is actually conservative.

$$\delta_{e^+e^- \to \mu^+\mu^-}^{RR,ew} = -2.58 L(s) - 5.15 \left( \log \frac{t}{u} \right) l(s) + 0.29 l_Z + 7.73 l_C + 8.80 l_{PR},$$
  
$$\delta_{e^+e^- \to \mu^+\mu^-}^{RL,ew} = -4.96 L(s) - 2.58 \left( \log \frac{t}{u} \right) l(s) + 0.37 l_Z + 14.9 l_C + 8.80 l_{PR},$$

order 1

**--** order 1 (times 10)

Single Log

Single Log (times 10)

Double Log

Double Log (times 10)

### Taking into account only DL, and not SL, is not safe for partonic energies up to 10 TeV.

Just a representative example of a process

Denner Pozzorini '01



Future Colliders: are EW Sudakov logarithms a good and robust approximation for EW corrections at high energies?

Currently: exact NLO EW automated for SM but not for BSM.

Since EW corrections are expected to be relevant also for BSM, can we safely use the high-energy Sudakov approximation?



## MadGraph5\_aMC@NLO: EW corrections for FC

NLO EW hadron colliders: Frederix, Frixione, Hirshi, DP, Shao, Zaro '18

**NLO EW**  $e^+e^-$  **colliders:** Bertone, Cacciari, Frixione, Stagnitto, Zaro, Zhao '22'

**One-loop EW Sudakov alone:** DP, Zaro '21

### one-loop EW virtual corrections $\mathcal{O}(\alpha)$

Having separately exact NLO EW and EW Sudakov logarithms is possible to study the goodness of the high-energy approximation(s). SM as a test case!

 $\alpha$  [Sudakov Logs  $\mathcal{O}(-\log^k(s/m_W^2), k = 1,2) +$ constant term  $\mathcal{O}(1)$  + mass-suppressed terms  $O(m_W^2/s)$ ]

# Master formula (Denner&Pozzorini)



### **ASSUMPTIONS:**

 $r_{kl} \equiv (p_k + p_l)^2 \simeq 2p_k p_l \gg M_W^2 \simeq M_H^2, m_t^2, M_W^2$ 

 $r_{kl}/r_{k'l'} \simeq 1$ 

All invariants  $\simeq s$ . Reasonable, but  $r_{kl} = s$  is impossible.

### Denner Pozzorini '01

$p_1,\ldots,$	$p_n)\delta_{i'_1i_1\dots i'_ni_n}$
e-level udes	the logs
	The logs inside the $\delta^i$ have always the form:
	$L( r_{kl} , M^2) \equiv \frac{\alpha}{4\pi} \log^2 \frac{ r_{kl} }{M^2}$
eter alis.	$l( r_{kl} , M^2) \equiv \frac{\alpha}{4\pi} \log \frac{ r_{kl} }{M^2}$
the sation ce is	$M = M_W, M_Z, m_f, \lambda, \dots$ $r_{kl} \equiv (p_k + p_l)^2$
$,M_{Z}^{z}$	the high-energy limit

**Our revisitation:** 

DP, Zaro '21

Logs among invariants: Logs like log(t/s) taken into account.

 $SDK_{Weak}$  scheme: A purely Weak (no QED) scheme for improving approximation of IR-finite physical observables. Different to the more common  $SDK_0$  scheme that has been used in the literature.





# ZZZ production at 100 TeV FCC-hh/SppC



 $p_T(Z_i) > 1 \text{ TeV}, \qquad |\eta(Z_i)| < 2.5, \qquad m(Z_i, Z_j) > 1 \text{ TeV},$  $\Delta R(Z_i, Z_j) > 0.5.$ 

**Orange:** NLO EW, (**dotted**: NLO EW no  $\gamma$  PDF) **Green =**  $SDK_0$ , **Red =**  $SDK_{weak}$ **Dashed**: standard approach for amplitudes. **Solid**: our formulation (more angular information)

**Reference Prediction:** 

 $SDK_{weak}$  and  $SDK_0$  not so relevant for neutral final state).

Only the solid lines, having more angular information, correctly capture NLO EW.

One cannot forget terms as  $\log^2[m^2(Z_2, Z_3)/s]$ 

Larger invariant -> larger correction



 $\mu^+\mu^- \longrightarrow F$ , where F is a generic final state involving W, Z, t, H. Thus we select direct production, with no VBF contributions.



than  $\mu$  are relevant. particle in F: particles.

Han, Ma, Xie '20, '21

# The muon collider case

We require  $m(F) > 0.8\sqrt{S}$ , so that neither VBF nor PDFs other

We apply further experimentally motivated cuts for each X, Y

 $p_T(X) > 100 \text{ GeV}, |\eta(X)| < 2.44, \Delta R(X, Y) > 0.4$ 

And we recombine photons with charged (also massive)

The  $\mu\,$  PDF in the  $\mu\,$  is peaked at **Bjorken-x=1**, therefore: Collider  $S \simeq$  partonic s

22













Sudakov logs capture NLO EW corrections up to the % level, but only if all the logs of the form log(t/s) are taken into account.

Green: logs of the form  $log^2(t/s)$  or log(t/s) ignored.

### Ma, DP, Zaro '24





Exponentiation as an approximation of proper resummation.

At 10 TeV resummation is unavoidable for sensible predictions, and it is necessary for precision at 3 TeV.





# What about extra radiation of Z (and H)?

logarithms.

But a cancellation is still present, how much large?

Is it really Heavy-Boson-Radiation (HBR) leading to  $\mathcal{O}(1)$  corrections?

- We know that unlike QCD in virtual+real there is not the exact cancellation of

  - **EW** is the new **QCD**, but it is not exactly as the QCD!



It is a general pattern: radiation of heavy bosons is much less important than loops!



Ma, DP, Zaro '24

# EW Sudakov and SMEFT: tt

### Only Four-Fermion operators are considered in the study.



### $\mathcal{O}_{tl} = (\bar{t}\gamma^{\mu}t)(\bar{l}_i\gamma_{\mu}l_i)$

### K-factors can be different in SM and BSM!

# 10 Tev $\mu$ -coll

QCD Both and corrections are different for SM, SM-SMEFT interference, and SMEFT^2 contributions of dim-6.

### **QCD** and **EW** cancel each other: both are important.

El Faham, Mimasu, DP, Severi, Vryonidou, Zaro: in preparation

27





## CONCLUSION

- energies. Not only for the SM also for BSM!
- mandatory.
- single(double) H.
- logs among invariants present and other features not discussed here in the plot.
- than the virtual contributions.
- general for precision.

- EW corrections are mandatory for phenomenology at future colliders, especially for high

- For precision: both NNLO ( $\alpha_s \alpha$  and  $\alpha^2$ ) corrections and NNL accuracy for PDFs are

-EW corrections open up sensitivity to new (BSM) interactions: HHH(H) in

- Sudakov logs are the dominant contribution of EW corrections at high energy (muon Collider) and they are a good approximation of them, but only IF: single logs present,

- Heavy-Boson Radiation has an impact, but not always so large and typically smaller

Resummation may be mandatory for sensible results in many configurations and in













EXTRA SLIDES





### Results

Maltoni, DP, Zhao '18

# NLO EW: some open questions/issues

### **Resummation?**

When is it necessary to resum EW (Sudakov) corrections?

### **BSM?**

What features of NLO EW corrections are universal and can be extended to the BSM case?

# **Heavy Boson Radiation (HBR)?** What should one do with Z,W radiat the calculation result.

**PDFs or VBF with matrix elements?** If PDFs involve weak effects, weak counter terms in NLO EW corrections should be included. Resum logs or keep power corrections? Both?

What should one do with Z,W radiation? Experimental set-up may impact

# What are EW Sudakov logarithms?

sections the contributions are combined and poles cancel.

poles  $\rightarrow \log(Q^2/\lambda^2)$ , where Q is a generic scale.

into account, which is anyway IR-safe.

Therefore, at high energies EW loops induce corrections of order

- **QCD**: virtual and real terms are separately IR divergent ( $1/\epsilon$  poles). In physical cross
- QED: same story, but I can also regularise IR divergencies via a photon-mass  $\lambda$ . So  $1/\epsilon$
- **EW**: with weak interactions  $\lambda \to m_W, m_Z$  and W and Z radiation are typically not taken

  - $-\alpha^k \log^n(s/m_W^2)$
- where k is the number of loops and  $n \leq 2k$ . These logs are physical. Even including the real radiation of W and Z, there is not the full cancellation of this kind of logarithms.













# e<sup>+</sup>e<sup>-</sup> production at 100 TeV FCC-hh

 $p_T(\ell^{\pm}) > 200 \text{ GeV}, \qquad |\eta(\ell^{\pm})| < 2.5, \qquad m(\ell^+, \ell^-) > 400 \text{ GeV}, \qquad \Delta R(\ell^+, \ell^-) > 0.5.$ 

**Orange:** NLO EW, (**dotted**: NLO EW no  $\gamma$  PDF) **Green =**  $SDK_0$ , **Red =**  $SDK_{weak}$ **Dashed**: standard approach for amplitudes. **Solid**: our formulation (more angular information)

**Reference Prediction:** 

Only the  $SDK_{weak}$  approach correctly captures the NLO EW prediction.

Solid and dashed very similar.

Photon PDF cannot be ignored.

Larger invariant -> larger correction





For smaller  $p_T$  , larger corrections.

Sudakov (in the SDK<sub>weak</sub> scheme) capture NLO EW corrections up to the % level.

If double logs are written in the form  $\log^2(s/m_W^2)$ , the shapes observed here are all arising from **single logs**.

tt

Ma, DP, Zaro TODAY





# Sudakov may completely fail: ZHH



Ma, DP, Zaro **TODAY** 

EW corrections are

Sudakov logarithms work very well at low pt and very bad at high pt.





# Sudakov may completely fail: ZHH



Ma, DP, Zaro **TODAY** 

For High pt of the Z boson, the two Higgs can have very small  $\Delta \vec{R}$  and so small  $\left[ \frac{2}{2} \right]^{10^{-6}}$ m(HH), recoiling against

In that configuration, formally mass suppressed can  $m(H_1H_2)$ become numerically sizeable, and the **DP** 



## Very small effects from Z and H radiation, especially in the bulk: $p_T(t) \simeq \sqrt{S/2}$



Notice that in order to allow more phase space we required just  $m(F) > 0.5\sqrt{S}$ . Still HBR << NLO EW in absolute value.



Ma, DP, Zaro '24

# EW Sudakov and SMEFT: *tt*

### Only Four-Fermion operators are considered in the study.

$$\mathcal{O}_{tu}^8 = (\overline{t}\gamma^\mu T^A t)(\overline{u}_i\gamma_\mu T^A u_i)$$



### K-factors can be different in SM and BSM!

# LHC

Both QCD and corrections are different for SM, SM-SMEFT interference, and SMEFT^2 contributions of dim-6.

### **QCD** and **EW** cancel each other: both are important.

El Faham, Mimasu, DP, Severi, Vryonidou, Zaro: in preparation

39





## Our revisitation and automation: Amplitude level

We have revisited and automated in aMG5 the Denner&Pozzorini **algorithm** for the evaluation of one-loop EW Sudakov corrections to amplitudes (Denner, Pozzorini '01). In particular we have introduced the following novelties.

- with strictly massless photons and light fermions.
- therefore **angular** dependences, are taken into account.
- loops on top of subleading LO terms.

**IR QED** divergencies are dealt with **via D**imensional **R**egularisation,

Additional logarithms that involve ratios between invariants, and

We correctly take into account an **imaginary term** that was **previously omitted** in the literature. Relevant for  $2 \rightarrow n$  processes with n > 2

Moving to the level of interferences of tree and one-loop amplitudes, we take into account NLO EW contributions originating from QCD

# **Derivation of LSC and SSC**



$$E_{l}|, M^{2}) = L(s, M^{2}) + 2l(s, M^{2}) \log \frac{|r_{kl}|}{s} + L(|r_{kl}|, s)$$

$$\equiv L(s) + 2l(s) \log \frac{M_{W}^{2}}{M^{2}} + 2l(s) \log \frac{|r_{kl}|}{s} + \cdots$$

$$LSC \qquad SSC$$

$$= L(s, M_{W}^{2}) \qquad \text{and} \qquad l(s) = l(s, M_{W}^{2})$$

### more on the very small effects from Z and H radiation

![](_page_41_Figure_1.jpeg)

![](_page_41_Figure_2.jpeg)

Ma, DP, Zaro '24

![](_page_42_Figure_0.jpeg)

# It is a general pattern: radiation of heavy bosons is much less important than loops!

### EW jets

 $\sigma_X(2j_{\rm EW}) \equiv \sigma_X(2V)$  for X = LO, NLO EW, SDK<sub>weak</sub>

$$\sigma_{\rm HBR}(2j_{\rm EW}) \equiv \sigma_{\rm LO}(3V) \,,$$
  
$$\sigma_{\rm NLO_{\rm EW}+HBR}(2j_{\rm EW}) \equiv \sigma_{\rm NLO_{\rm EW}}(2V) + \sigma_{\rm LO}(3V)$$

$$\sigma_{\rm nNLO_{EW}+HBR_{\rm NLO}}(2j_{\rm EW}) \equiv \sigma_{\rm LO}(2V) \left(1 + \delta_{\rm NLO_{EW}} + \frac{\delta_{\rm SDK_{weak}}^2}{2}\right) + \sigma_{\rm NLO_{EW}}(3V) + \sigma_{\rm LO}(4V).$$

$$\Delta_X(2V) \equiv \frac{\sigma_X(2V) - \sigma_{\rm LO}(2V)}{\sigma_{\rm LO}(2V)}$$
$$\Delta_X(3V) \equiv \frac{\sigma_X(3V)}{\sigma_{\rm LO}(2V)}$$
$$\Delta_X(4V) \equiv \frac{\sigma_X(4V)}{\sigma_{\rm LO}(2V)}$$

Ma, DP, Zaro '24

## Cross-sections: our approach.

FOR WHAT EW SUDAKOV ARE USEFUL? For providing a very **good approximation of NLO EW** in the **high-energy** limit.

HOW SHOULD ONE PERFORM THE CALCULATION IN THE HIGH-ENERGY LIMIT? Photons have to be always clustered with massless charged particle for IR-safety reasons. But from an experimental point of view, at high energy also clustering tops and W bosons with photons is very reasonable, either if you imagine to tag heavy object directly or via their massless decay products.

The QED Logs, involving s and  $\lambda^2$  (or  $Q^2$ ), cancel against their real-emission counterparts and PDF counterterms. The only one surviving are those from tops in vacuum polarisation for external (not tagged) photons, both in the initial and final state:

Almost all the contributions of QED are removed (e.g.  $C_{\rm EW}(k) \to C_{\rm EW}(k) - Q_k^2$ ,  $Q_k^2 = 0$ ), but NOT in the parameter renormalisation  $\delta^{\mathrm{PR}}$ .

DP, Zaro '21

![](_page_43_Picture_9.jpeg)

## Implementation

Born amplitude:

 $\mathcal{M}_0^{i_1\dots i_n}(p_1,\dots,p_n)$ 

**One-loop EW** Sudakov corrections:  $\delta \mathcal{M}^{i_1...i_n}(p_1,\ldots,p_n)$ 

Born process:

![](_page_44_Figure_6.jpeg)

![](_page_44_Picture_7.jpeg)

different external particles w.r.t. the Born have to be generated.

$$Z \longrightarrow \chi,$$
  
 $W^{\pm} \longrightarrow \phi^{\pm},$ 

GBE theorem for longitudinal W and Z bosons.

![](_page_44_Figure_14.jpeg)

$$\delta_{i_{1}'i_{1}...i_{n}'}(p_{1},...,p_{n})\delta_{i_{1}'i_{1}...i_{n}'i_{n}}$$

the logs

other tree-level amplitudes

$$\varphi_{i_1}(p_1)\ldots\varphi_{i_n}(p_n)\to 0$$

Relevant for SSC charged contributions.

# Organisation of the logs in the algorithm

Two examples: LSC and C for fermions

$$\begin{split} \delta_{i'_{k}i_{k}}^{\mathrm{LSC}}(k) &= -\frac{1}{2} \begin{bmatrix} C_{i'_{k}i_{k}}^{\mathrm{ew}}(k) L(s) - 2(I^{Z}(k))_{i'_{k}i_{k}}^{2} \log \frac{M_{Z}^{2}}{M_{W}^{2}} l(s) + \delta_{i'_{k}i_{k}} Q_{k}^{2} L^{\mathrm{em}}(s, \lambda^{2}, m_{k}^{2}) \end{bmatrix} \\ \mathbf{Casimir for the entire}_{SU(2)_{L} \times U(1)_{B}} & \mathbf{Charge for}_{U(1)_{QED}} \\ \delta_{f_{\sigma}f_{\sigma'}}^{\mathrm{C}}(f^{\kappa}) &= \delta_{\sigma\sigma'} \left\{ \begin{bmatrix} \frac{3}{2} C_{f^{\kappa}}^{\mathrm{ew}} - \frac{1}{8s_{W}^{2}} \left( (1 + \delta_{\kappa\mathrm{R}}) \frac{m_{f_{\sigma}}^{2}}{M_{W}^{2}} + \delta_{\kappa\mathrm{L}} \frac{m_{f_{-\sigma}}^{2}}{M_{W}^{2}} \right) \right\} l(s) + Q_{f_{\sigma}}^{2} l^{\mathrm{em}}(m_{f_{\sigma}}^{2}) \right\} \end{split}$$

$$\begin{split} f(k) &= -\frac{1}{2} \begin{bmatrix} C_{i'_{k}i_{k}}^{\text{ew}}(k) L(s) - 2(I^{Z}(k))_{i'_{k}i_{k}}^{2} \log \frac{M_{Z}^{2}}{M_{W}^{2}} l(s) + \delta_{i'_{k}i_{k}} Q_{k}^{2} L^{\text{em}}(s, \lambda^{2}, m_{k}^{2}) \end{bmatrix} \\ \begin{array}{c} \text{Casimir for the entire} \\ SU(2)_{L} \times U(1)_{B} \end{bmatrix} \\ \delta_{f_{\sigma}f_{\sigma'}}^{\text{C}}(f^{\kappa}) &= \delta_{\sigma\sigma'} \left\{ \begin{bmatrix} \frac{3}{2} C_{f^{\kappa}}^{\text{ew}} - \frac{1}{8s_{w}^{2}} \left( (1 + \delta_{\kappa R}) \frac{m_{f_{\sigma}}^{2}}{M_{W}^{2}} + \delta_{\kappa L} \frac{m_{f_{-\sigma}}^{2}}{M_{W}^{2}} \right) \end{bmatrix} l(s) + Q_{f_{\sigma}}^{2} l^{\text{em}}(m_{f_{\sigma}}^{2}) \right\} \end{split}$$

$$L(s) \equiv L(s, M_W^2) \quad \text{and} \quad l(s) \equiv l(s, M_W^2)$$
$$l^{\text{em}}(m_f^2) := \frac{1}{2}l(M_W^2, m_f^2) + l(M_W^2, \lambda^2) \quad L^{\text{em}}(s, \lambda^2, m_k^2) := 2l(s)\log\left(\frac{M_W^2}{\lambda^2}\right) + L(M_W^2, \lambda^2) - L(m_k^2, \lambda^2)$$

The full EW is present between s and  $M_W^2$ , while only QED is present between  $M_W^2$  and  $\lambda^2$ .

just a technical parameter and not a physical quantity.

So the QED contribution is split between the intervals  $(s, M_W^2) + (M_W^2, \lambda^2)$ . But the division at  $M_W^2$  is simply determined by convenience, in parallel with the weak case. In this case  $M_W^2$  is

# Cross-sections: standard approach in the literature **SDK**

Two examples: LSC and C for fermions

$$\delta_{i'_k i_k}^{\text{LSC}}(k) = -\frac{1}{2} \left[ C^{\text{ew}}_{i'_k i_k}(k) L(s) - 2(I^Z(k))_{i'_k}^2 \right]$$

**Casimir for the entire**  $SU(2)_L \times U(1)_B$ 

$$\delta_{f_{\sigma}f_{\sigma'}}^{\mathcal{C}}(f^{\kappa}) = \delta_{\sigma\sigma'} \left\{ \left[ \frac{3}{2} C_{f^{\kappa}}^{\mathrm{ew}} - \frac{1}{8s_{\mathrm{w}}^2} \left( (1 + \delta_{\kappa\mathrm{R}}) \frac{m_{f_{\sigma}}^2}{M_{\mathrm{W}}^2} + \delta_{\kappa\mathrm{L}} \frac{m_{f_{-\sigma}}^2}{M_{\mathrm{W}}^2} \right) \right] l(s) + \mathcal{Q}_{f_{\sigma}}^2 \mathcal{Q}_{f_{\sigma}}^2 \left( (1 + \delta_{\kappa\mathrm{R}}) \frac{m_{f_{\sigma}}^2}{M_{\mathrm{W}}^2} + \delta_{\kappa\mathrm{L}} \frac{m_{f_{-\sigma}}^2}{M_{\mathrm{W}}^2} \right) \right] l(s) + \mathcal{Q}_{f_{\sigma}}^2 \mathcal{Q}_{f_{\sigma}}^2 \left( (1 + \delta_{\kappa\mathrm{R}}) \frac{m_{f_{\sigma}}^2}{M_{\mathrm{W}}^2} + \delta_{\kappa\mathrm{L}} \frac{m_{f_{-\sigma}}^2}{M_{\mathrm{W}}^2} \right) \right] l(s) + \mathcal{Q}_{f_{\sigma}}^2 \mathcal{Q}_{f_{\sigma}}^2 \left( (1 + \delta_{\kappa\mathrm{R}}) \frac{m_{f_{\sigma}}^2}{M_{\mathrm{W}}^2} + \delta_{\kappa\mathrm{L}} \frac{m_{f_{-\sigma}}^2}{M_{\mathrm{W}}^2} \right) \right] l(s) + \mathcal{Q}_{f_{\sigma}}^2 \mathcal{Q}_{f_{\sigma}}^2 \left( (1 + \delta_{\kappa\mathrm{R}}) \frac{m_{f_{\sigma}}^2}{M_{\mathrm{W}}^2} + \delta_{\kappa\mathrm{L}} \frac{m_{f_{-\sigma}}^2}{M_{\mathrm{W}}^2} \right) \right] l(s) + \mathcal{Q}_{f_{\sigma}}^2 \mathcal{Q}_{f_{\sigma}}^2 \left( (1 + \delta_{\kappa\mathrm{R}}) \frac{m_{f_{-\sigma}}^2}{M_{\mathrm{W}}^2} + \delta_{\kappa\mathrm{L}} \frac{m_{f_{-\sigma}}^2}{M_{\mathrm{W}}^2} \right) \right] l(s) + \mathcal{Q}_{f_{\sigma}}^2 \mathcal{Q}_{f_{\sigma}}^2 \left( (1 + \delta_{\kappa\mathrm{R}}) \frac{m_{f_{-\sigma}}^2}{M_{\mathrm{W}}^2} + \delta_{\kappa\mathrm{L}} \frac{m_{f_{-\sigma}}^2}{M_{\mathrm{W}}^2} \right) \right] l(s) + \mathcal{Q}_{f_{\sigma}}^2 \mathcal{Q}_{f_{\sigma}}^2 \left( (1 + \delta_{\kappa\mathrm{R}}) \frac{m_{f_{-\sigma}}^2}{M_{\mathrm{W}}^2} + \delta_{\kappa\mathrm{L}} \frac{m_{f_{-\sigma}}^2}{M_{\mathrm{W}}^2} \right) \right] l(s) + \mathcal{Q}_{f_{\sigma}}^2 \mathcal{Q}_{f_{\sigma}}^2 \left( (1 + \delta_{\kappa\mathrm{R}}) \frac{m_{f_{-\sigma}}^2}{M_{\mathrm{W}}^2} + \delta_{\kappa\mathrm{L}} \frac{m_{f_{-\sigma}}^2}{M_{\mathrm{W}}^2} \right) \right] l(s) + \mathcal{Q}_{f_{\sigma}}^2 \mathcal{Q}_{f_{\sigma}}^2 \left( (1 + \delta_{\kappa\mathrm{R}}) \frac{m_{f_{-\sigma}}^2}{M_{\mathrm{W}}^2} + \delta_{\kappa\mathrm{L}} \frac{m_{f_{-\sigma}}^2}{M_{\mathrm{W}}^2} \right) \right] l(s) + \mathcal{Q}_{f_{\sigma}}^2 \mathcal{Q}_{f_{\sigma}}^2 \left( (1 + \delta_{\kappa\mathrm{R}}) \frac{m_{f_{-\sigma}}^2}{M_{\mathrm{W}}^2} \right) \right] l(s) + \mathcal{Q}_{f_{\sigma}}^2 \mathcal{Q}_{f_{\sigma}}^2 \left( (1 + \delta_{\kappa\mathrm{R}}) \frac{m_{f_{-\sigma}}^2}{M_{\mathrm{W}}^2} \right) \right] l(s) + \mathcal{Q}_{f_{\sigma}}^2 \mathcal{Q}_{f_{\sigma}}^2 \left( (1 + \delta_{\kappa\mathrm{R}}) \frac{m_{f_{-\sigma}}^2}{M_{\mathrm{W}}^2} \right) \right] l(s) + \mathcal{Q}_{f_{\sigma}}^2 \mathcal{Q}_{f_{\sigma}}^2 \left( (1 + \delta_{\kappa\mathrm{R}}) \frac{m_{f_{-\sigma}}^2}{M_{\mathrm{W}}^2} \right) \right] l(s) + \mathcal{Q}_{f_{\sigma}}^2 \mathcal{Q}_{f_{\sigma}}^2 \left( (1 + \delta_{\kappa\mathrm{R}}) \frac{m_{f_{\sigma}}^2}{M_{\mathrm{W}}^2} \right) \right] l(s) + \mathcal{Q}_{f_{\sigma}}^2 \mathcal{Q}_{f_{\sigma}}^2 \left( (1 + \delta_{\kappa\mathrm{R}}) \frac{m_{f_{\sigma}}^2}{M_{\mathrm{W}}^2} \right) \right] l(s) + \mathcal{Q}_{f_{\sigma}}^2 \mathcal{Q}_{f_{\sigma}}^2 \left( (1 + \delta_{\kappa\mathrm{R}}) \frac{m_{f_{\sigma}}^2}{M_{\mathrm{W}}^2} \right) \right] l(s) + \mathcal{Q}_{f_{\sigma}}^2 \mathcal{Q}_{f_{\sigma}}^2 \left( (1 + \delta_{\kappa\mathrm{R}}) \frac{m_{f_{\sigma}}$$

and  $l(s) \equiv l(s, M_W^2)$  $L(s) \equiv L(s, \frac{M_W^2}{M_W})$ 

case of DR, logarithms involving  $M_W^2$  and the IR regulator  $Q^2$ .

### Easy, but not very well motivated.

We will denote in the following this approach as  $SDK_0$ .

![](_page_46_Picture_9.jpeg)

The logarithms between  $M_W^2$  and the infrared scale are simply removed. Equivalently in the