



中国科学院高能物理研究所
Institute of High Energy Physics
Chinese Academy of Sciences

CEPC NP Potential

-- White paper status

Xuai Zhuang (IHEP)

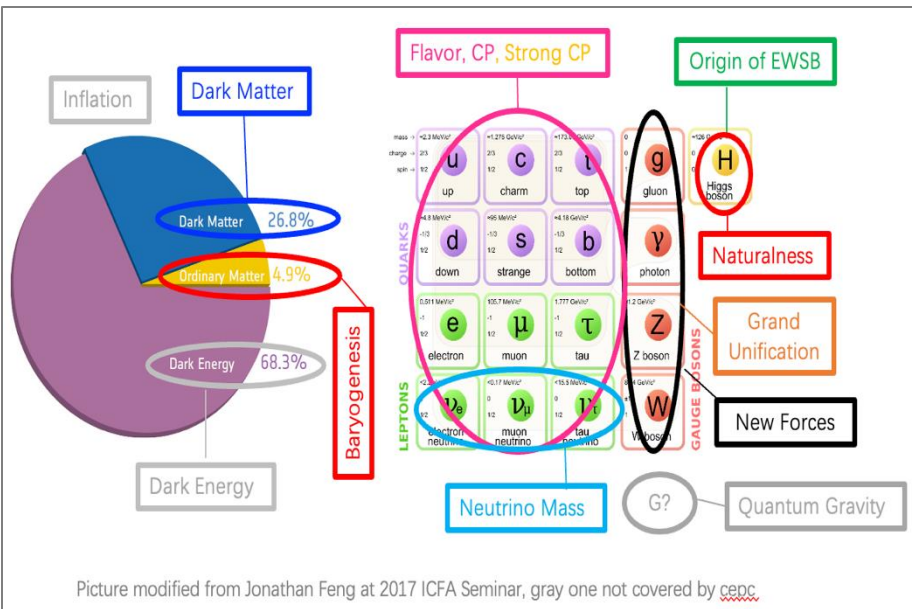
On behalf of CEPC NP team

zhuangxa@ihep.ac.cn

CEPC Internation Workshop 2024, Hangzhou, China

Oct. 22-26, 2024

Big Questions and Ideas in particle physics



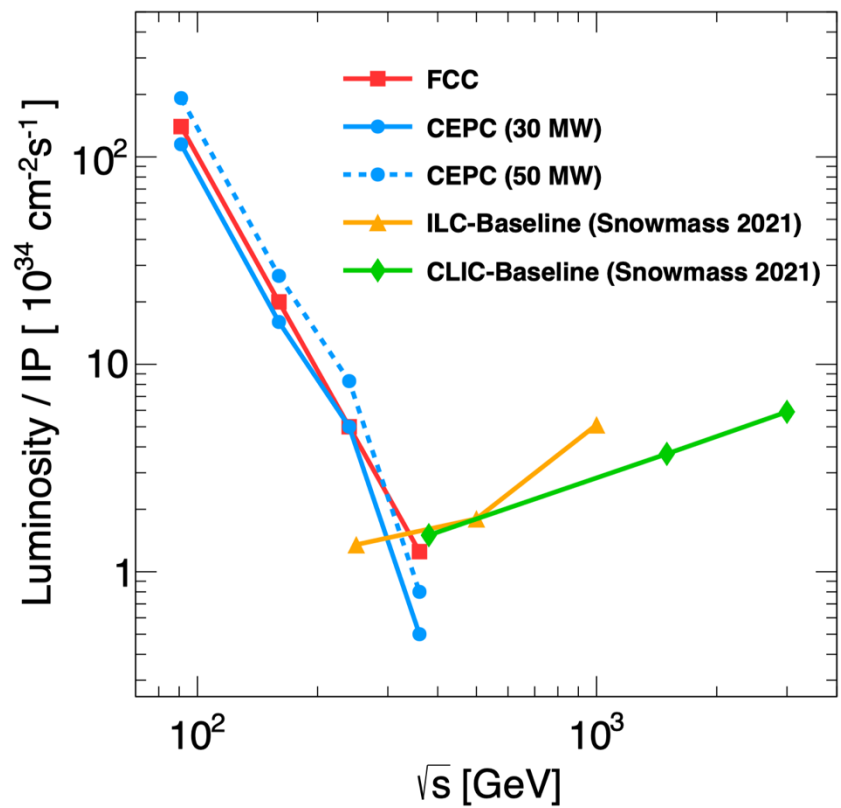
Big Questions

Big Ideas

Extended Higgs Sector Dark Sector LLP FCNC SUSY EWPT Exotic Neutrino EFT ALP

SM Higgs boson?	Red bar			Red bar			Red bar
Dark Matter		Blue bar		Blue bar			Blue bar
Origin of EWSB	Green bar			Green bar			
Unification				Yellow bar			
Origin of Flavor	Pink bar			Pink bar			
New Forces	Black bar			Black bar			Black bar
Neutrino Masses	Blue bar						Blue bar
Naturalness	Red bar			Red bar			
Strong CP	Yellow bar		Yellow bar				Yellow bar

CEPC operation scheme and NP Program

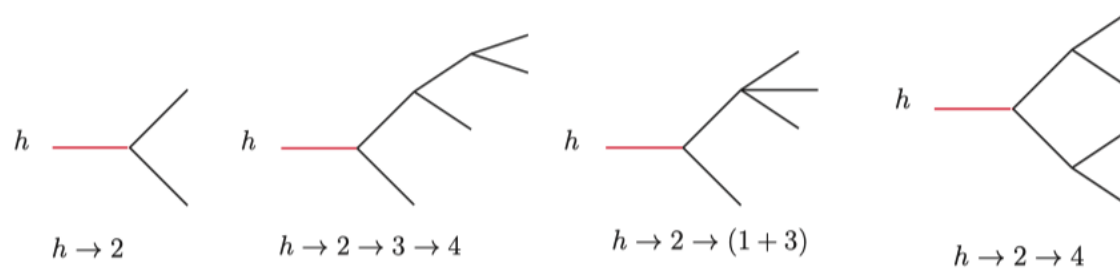


Operation mode	Z factory	WW threshold	Higgs factory	$t\bar{t}$
\sqrt{s} (GeV)	91.2	160	240	360
Run time (year)	2	1	10	5
Instantaneous luminosity ($10^{34} \text{ cm}^{-2} \text{ s}^{-1}$, per IP)	191.7	26.6	8.3	0.83
Integrated luminosity (ab^{-1} , 2 IPs)	100	6	20	1
Event yields	3×10^{12}	1×10^8	4×10^6	5×10^5

TABLE I: Nominal CEPC operation scheme, and the physics yield, of four different modes.

1. Exotic Higgs/Z/top decays

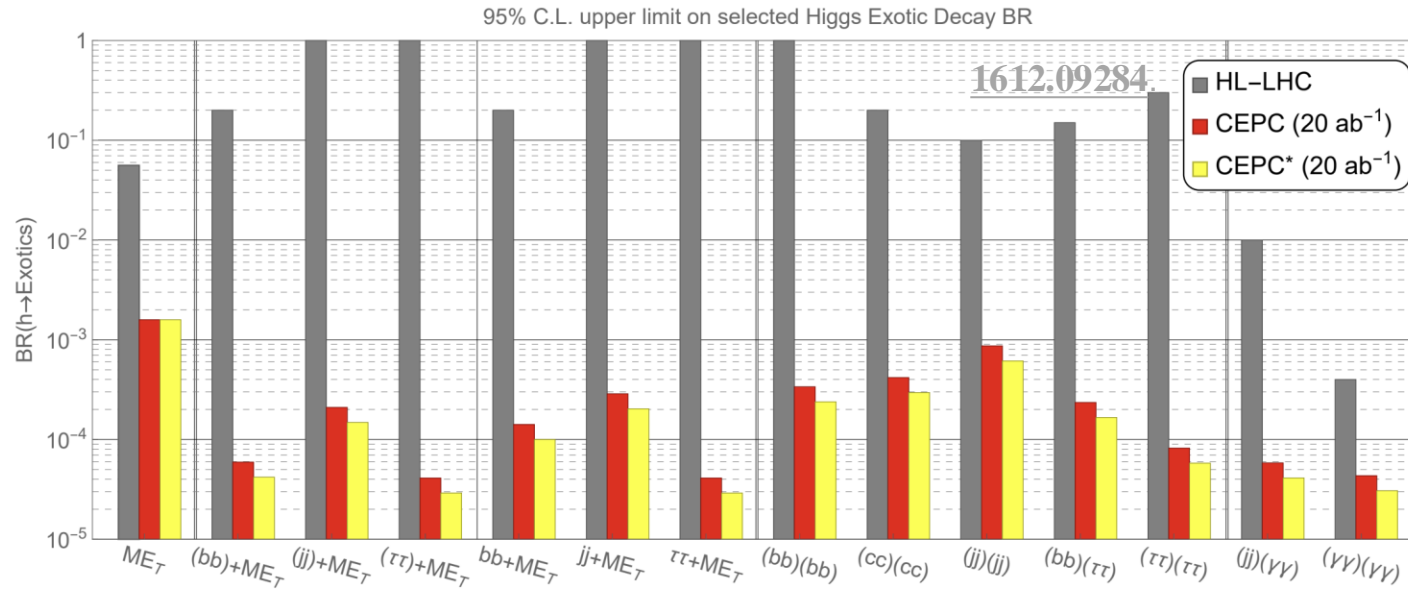
- *Higgs exotic decay* motivated by a large class of BSM physics, such as singlet extensions, two Higgs-doublet-models (2HDM), SUSY models, Higgs portals, gauge extensions of the SM ...



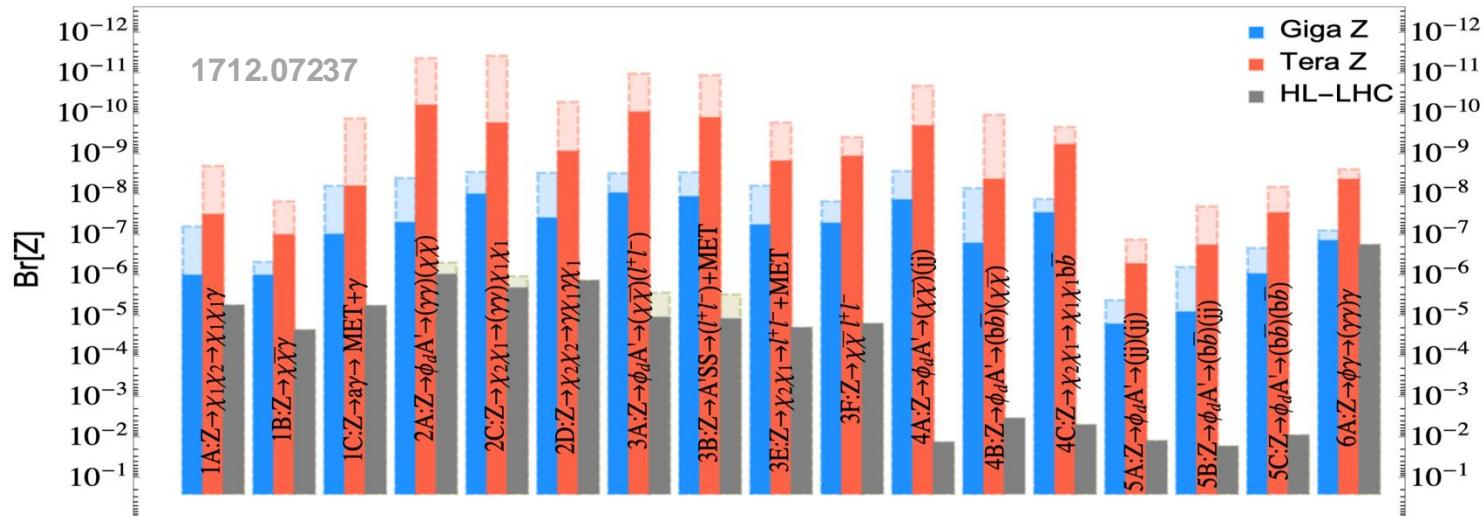
Representative topologies of the Higgs exotic decays

- *Exotic Z or top decays* are also motivated by many BSM models (ED, Heavy Vector Triplet, ...) and can also be searched at CEPC
- *Light Higgs* are motivated by 2HDM and Axion-like particle models, which can be searched at CEPC well if they exist.

Exotic Higgs/Z/top decays



The 95% C.L. upper limit on selected Higgs exotic decay BR

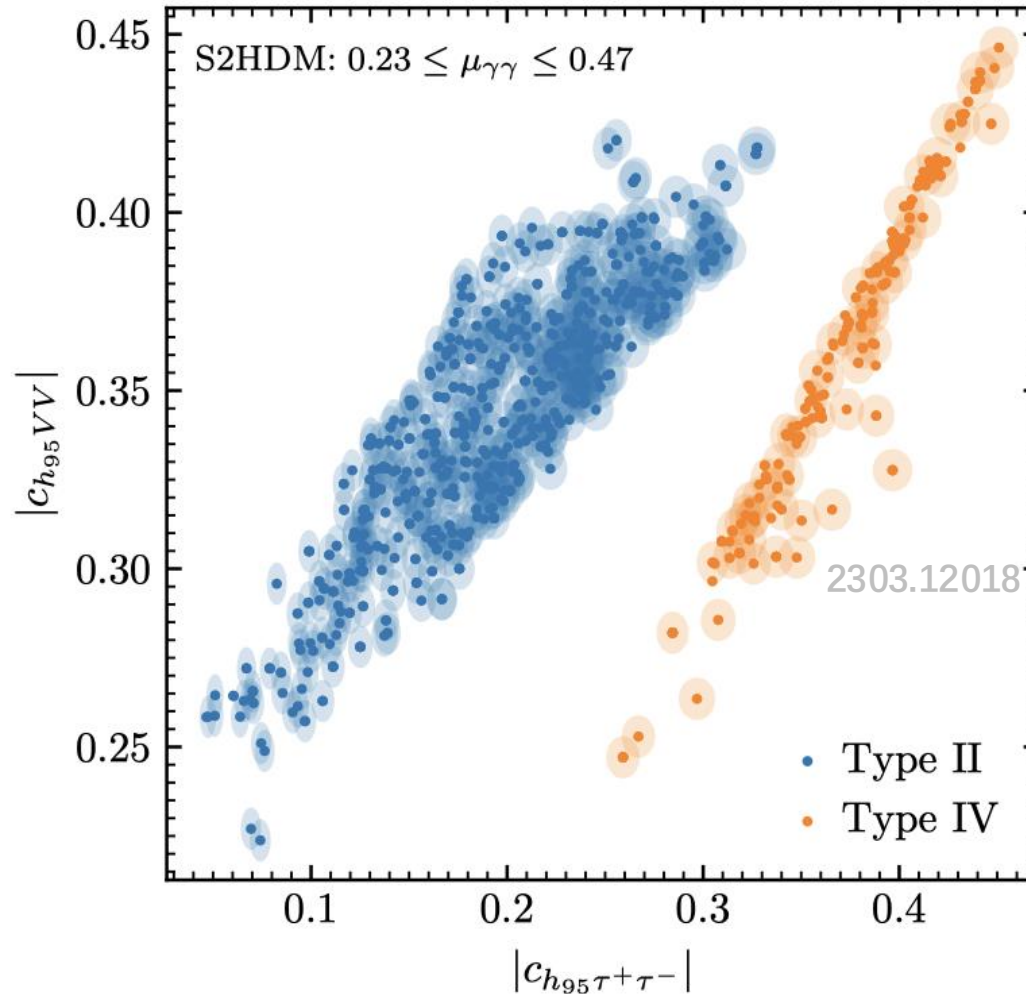
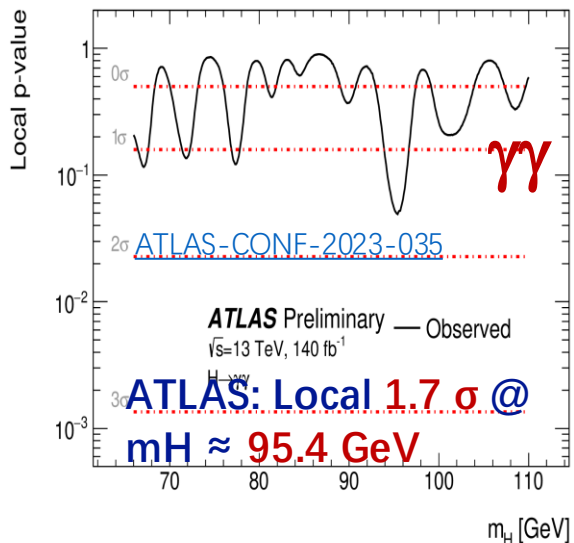
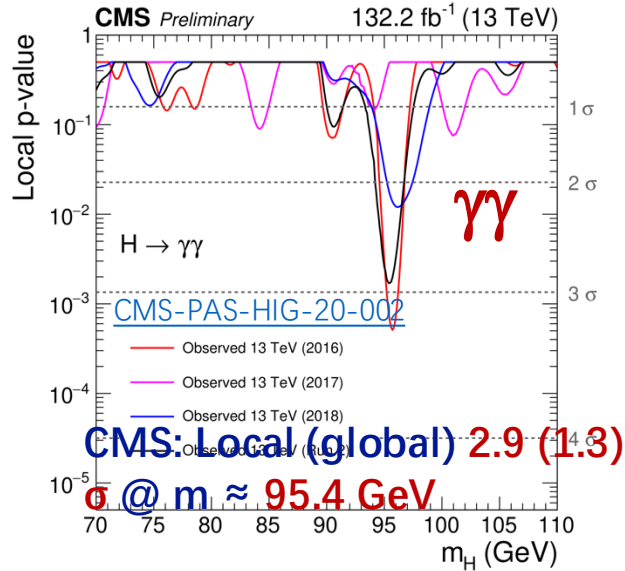


The reach for the branching ratio of various exotic Z decay modes

→ Good sensitivity of exotic Higgs/Z decay from CEPC

Light Higgs

- Light Higgs are motivated by 2HDM and Axion-like particle models



S2HDM parameter points passing the applied constraints for the di-photon signal strengths.

Light higgs can be searched at CEPC very well if exists.

2. Dark Matter and Dark Sector

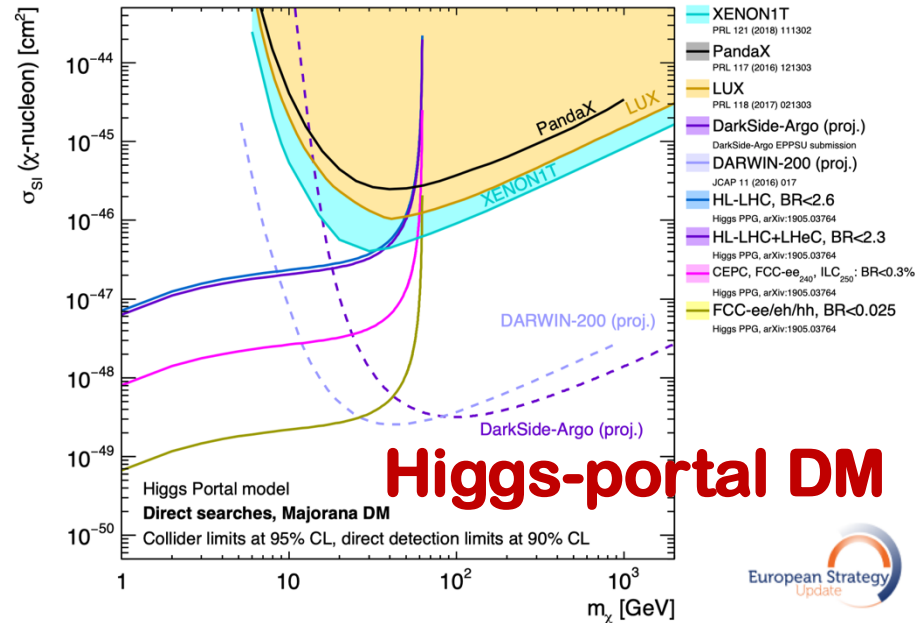
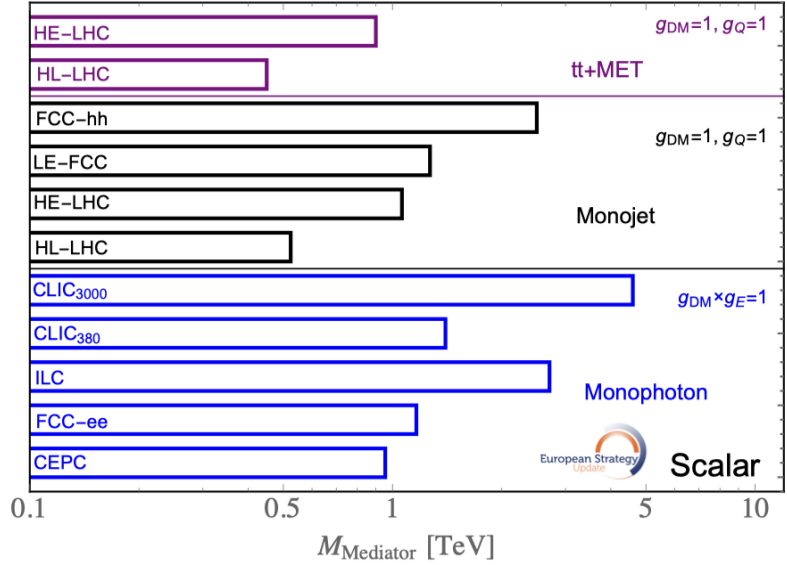
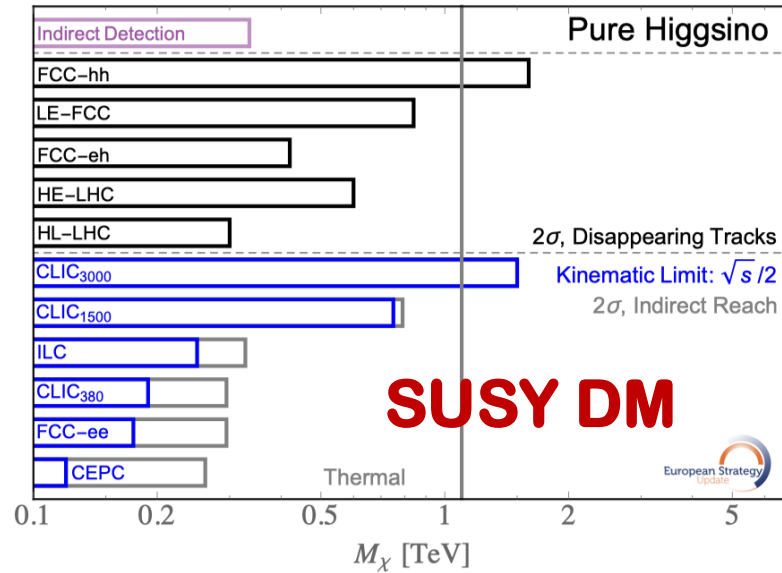
UV models DM:

- SUSY DM
- Double dark portal model
-

Simplified models DM:

- Scalar portal
- Fermion portal
- Vector portal

EFT DM

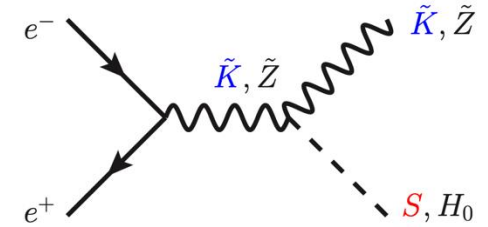


Simplified model DM

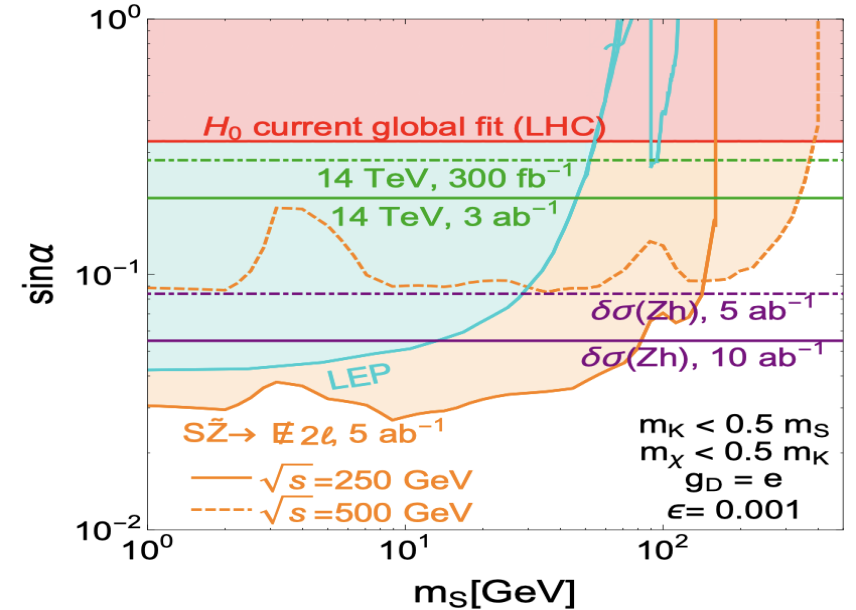
2. Dark Matter and Dark Sector

Portal	Effective operator	\sqrt{s} [GeV]	$\mathcal{L}[ab^{-1}]$	Sensitivity of CEPC (HL-LHC)	Figs.	Ref.
Scalar	$\lambda_{HP} H ^2S^2 \rightarrow$ scalar mixing $\sin\theta$	250	5	invisible S, $\sin\theta \approx 0.03$ (0.20 global-fits)	22	[108]
Fermion	$y\ell\bar{\chi}_L S^i \ell_R + \text{H.c.}$	250	5	covering $100 \text{ GeV} < m_S < 170 \text{ GeV}$	23	[56]
	$\kappa\Phi q_L^i \ell_R + \text{H.c.}$ (dark QCD)	250	5	$m_\Phi \sim 10 \text{ TeV}$ for $c\tau_{\text{darkpion}} \in [1, 10^3] \text{ cm}$ (Null)	25	[109]
	$y\Phi\bar{F}_L \ell_R + \text{H.c.}$	240	5.6	$y\theta_L \in [10^{-11}, 10^{-7}]$ ($\lesssim 10^{-8} - 10^{-9}$)	26	[110]
Vector	$A'_\mu (e\epsilon J_{\text{em}}^\mu + g_D \bar{\chi}\gamma^\mu\chi)$	250	5	$\epsilon \sim 10^{-3}$ for $g_D = e$ and $m_{A'} < 125 \text{ GeV}$ ($\epsilon \sim 0.02$)	27, 28	[108]
	$\epsilon A_\mu \bar{\chi}\gamma^\mu\chi$, (millicharge DM)	250	5	$\epsilon \sim 0.1$ for $m_\chi \sim 50 \text{ GeV}$	29	[111]
		91.2	2.6	$\epsilon \sim 0.02$ for $m_\chi \sim 5 \text{ GeV}$		
		160	16	$\epsilon \sim 0.5$ for $m_\chi \sim 10 \text{ GeV}$		
	$\frac{1}{2}\mu_\chi \bar{\chi}\sigma^{\mu\nu}\chi F_{\mu\nu} + \frac{i}{2}d_\chi \bar{\chi}\sigma^{\mu\nu}\gamma^5\chi F_{\mu\nu}$	91.2	100	$\mu_\chi, d_\chi \sim 4 \times 10^{-7}$ (4×10^{-6}) μ_B for $m_\chi < 25 \text{ GeV}$	30	[112]
$-a_\chi \bar{\chi}\gamma^\mu\gamma^5\chi\partial^\nu F_{\mu\nu} + b_\chi \bar{\chi}\gamma^\mu\chi\partial^\nu F_{\mu\nu}$	240	20	$a_\chi, b_\chi \sim 10^{-6}$ (2×10^{-6}) GeV^{-2} for $m_\chi < 80 \text{ GeV}$			
EFT	$\frac{1}{\Lambda^2} \sum_i (\bar{\chi}\gamma_\mu(1-\gamma_5)\chi)(\bar{\ell}\gamma^\mu(1-\gamma_5)\ell)$	250	5	$\Lambda_i \sim 2 \text{ TeV}$ ($m_\chi = 0$) (Null)	31	[113]
	$\frac{1}{\Lambda_A^2} \bar{\chi}\gamma_\mu\gamma_5\chi\bar{\ell}\gamma^\mu\gamma_5\ell$	250	5	$\Lambda_A \sim 1.5 \text{ TeV}$ (Null)	32	[111]
	$\sum_i \frac{1}{\Lambda_i^2} (\bar{e}\Gamma_\mu e)(\bar{\nu}_L\Gamma^\mu\chi_L) + \text{H.c.}$ $\Gamma_\mu = 1, \gamma_5, \gamma_\mu, \gamma_\mu\gamma_5, \sigma_{\mu\nu}$	240	20	$\Lambda_i \sim 1 \text{ TeV}$ ($m_\chi = 0$) (Null)	33	[114]

Dark Sector from Z/H associate production



Double dark portal model: Scale and Vector-portal DM

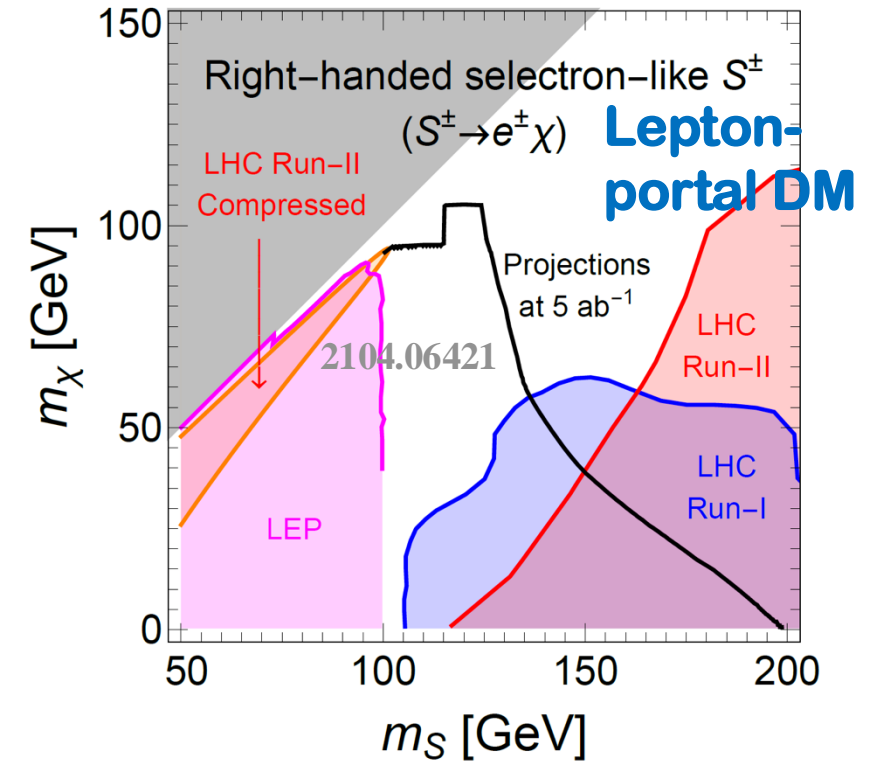


See details at Yongchao's talk this Wed.

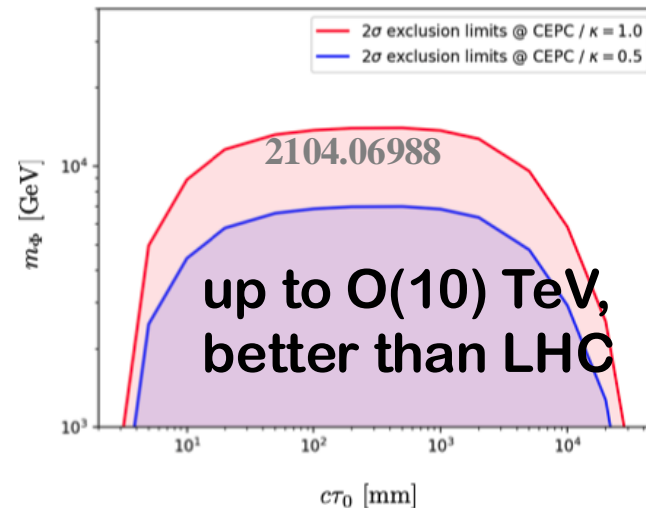
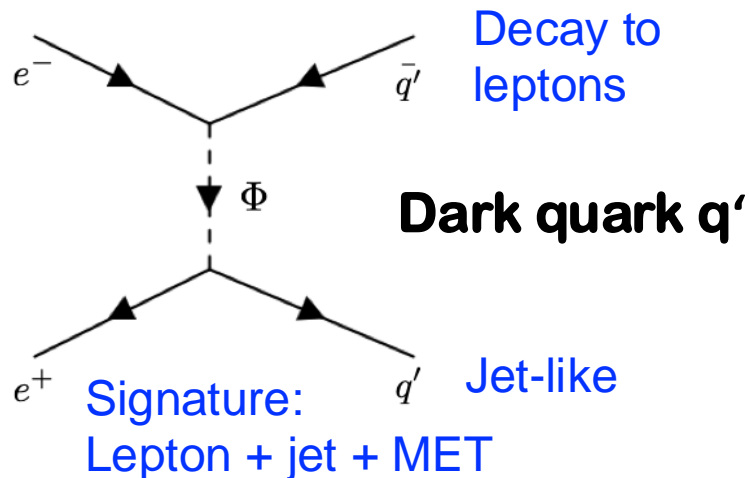
2. Dark Matter and Dark Sector

Portal	Effective operator	\sqrt{s} [GeV]	$\mathcal{L}[ab^{-1}]$	Sensitivity of CEPC (HL-LHC)	Figs.	Ref.
Scalar	$\lambda_{HP} H ^2 S^2 \rightarrow$ scalar mixing $\sin\theta$	250	5	invisible S, $\sin\theta \approx 0.03$ (0.20 global-fits)	22	[108]
Fermion	$y_\ell \bar{\chi}_L S^\dagger \ell_R + \text{H.c.}$	250	5	covering $100 \text{ GeV} < m_S < 170 \text{ GeV}$	23	[56]
	$\kappa \Phi \bar{q}'_L \ell_R + \text{H.c.}$ (dark QCD)	250	5	$m_\Phi \sim 10 \text{ TeV}$ for $c\tau_{\text{darkpion}} \in [1, 10^3] \text{ cm}$ (Null)	25	[109]
	$y\Phi \bar{F}_L \ell_R + \text{H.c.}$	240	5.6	$y\theta_L \in [10^{-11}, 10^{-7}]$ ($\lesssim 10^{-8} - 10^{-9}$)	26	[110]
Vector	$A'_\mu (e\epsilon J_{\text{em}}^\mu + g_D \bar{\chi} \gamma^\mu \chi)$	250	5	$\epsilon \sim 10^{-3}$ for $g_D = e$ and $m_{A'} < 125 \text{ GeV}$ ($\epsilon \sim 0.02$)	27, 28	[108]
	$\epsilon A_\mu \bar{\chi} \gamma^\mu \chi$, (millicharge DM)	250	5	$\epsilon \sim 0.1$ for $m_\chi \sim 50 \text{ GeV}$	29	[111]
		91.2	2.6	$\epsilon \sim 0.02$ for $m_\chi \sim 5 \text{ GeV}$		
		160	16	$\epsilon \sim 0.5$ for $m_\chi \sim 10 \text{ GeV}$		
	$\frac{1}{2} \mu_\chi \bar{\chi} \sigma^{\mu\nu} \chi F_{\mu\nu} + \frac{i}{2} d_\chi \bar{\chi} \sigma^{\mu\nu} \gamma^5 \chi F_{\mu\nu}$	91.2	100	$\mu_\chi, d_\chi \sim 4 \times 10^{-7}$ (4×10^{-6}) μ_B for $m_\chi < 25 \text{ GeV}$	30	[112]
$-a_\chi \bar{\chi} \gamma^\mu \gamma^5 \chi \partial^\nu F_{\mu\nu} + b_\chi \bar{\chi} \gamma^\mu \chi \partial^\nu F_{\mu\nu}$		240	20	$a_\chi, b_\chi \sim 10^{-6}$ (2×10^{-6}) GeV^{-2} for $m_\chi < 80 \text{ GeV}$		
EFT	$\frac{1}{\Lambda^2} \sum_i (\bar{\chi} \gamma_\mu (1 - \gamma_5) \chi) (\bar{\ell} \gamma^\mu (1 - \gamma_5) \ell)$	250	5	$\Lambda_i \sim 2 \text{ TeV}$ ($m_\chi = 0$) (Null)	31	[113]
	$\frac{1}{\Lambda_A^2} \bar{\chi} \gamma_\mu \gamma_5 \chi \bar{\ell} \gamma^\mu \gamma_5 \ell$	250	5	$\Lambda_A \sim 1.5 \text{ TeV}$ (Null)	32	[111]
	$\sum_i \frac{1}{\Lambda_i^2} (\bar{e} \Gamma_\mu e) (\bar{\nu}_L \Gamma^\mu \chi_L) + \text{H.c.}$ $\Gamma_\mu = 1, \gamma_5, \gamma_\mu, \gamma_\mu \gamma_5, \sigma_{\mu\nu}$	240	20	$\Lambda_i \sim 1 \text{ TeV}$ ($m_\chi = 0$) (Null)	33	[114]

$$e^+ e^- \rightarrow S^{\pm(*)} S^\mp \rightarrow \ell^+ \chi \ell'^- \chi$$

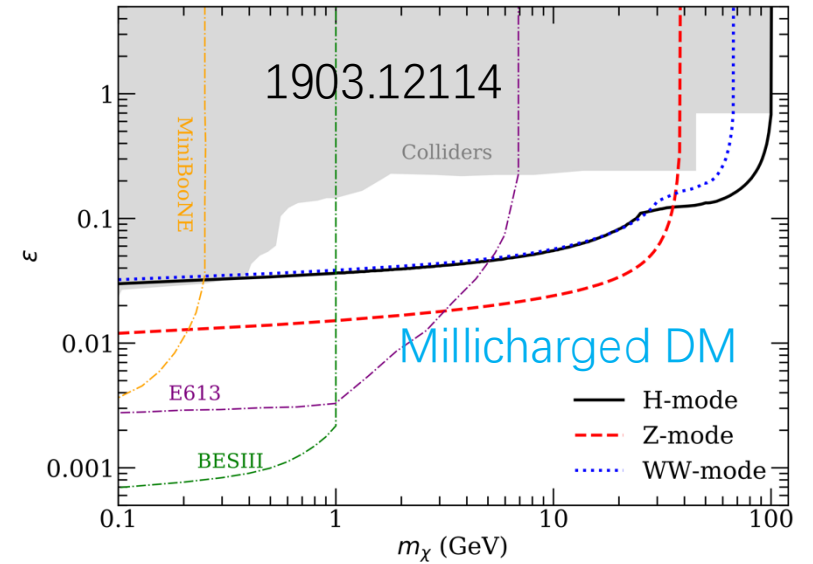


$m_S \sim 100 - 170 \text{ GeV}$



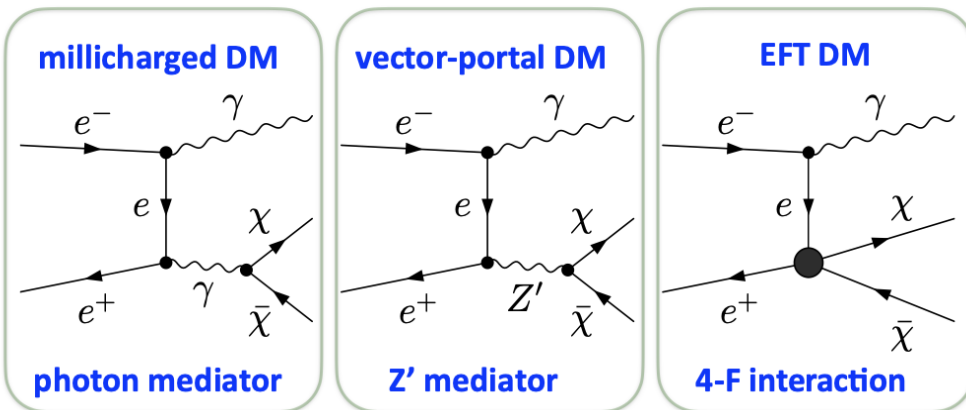
2. Dark Matter and Dark Sector

Vector-portal DM



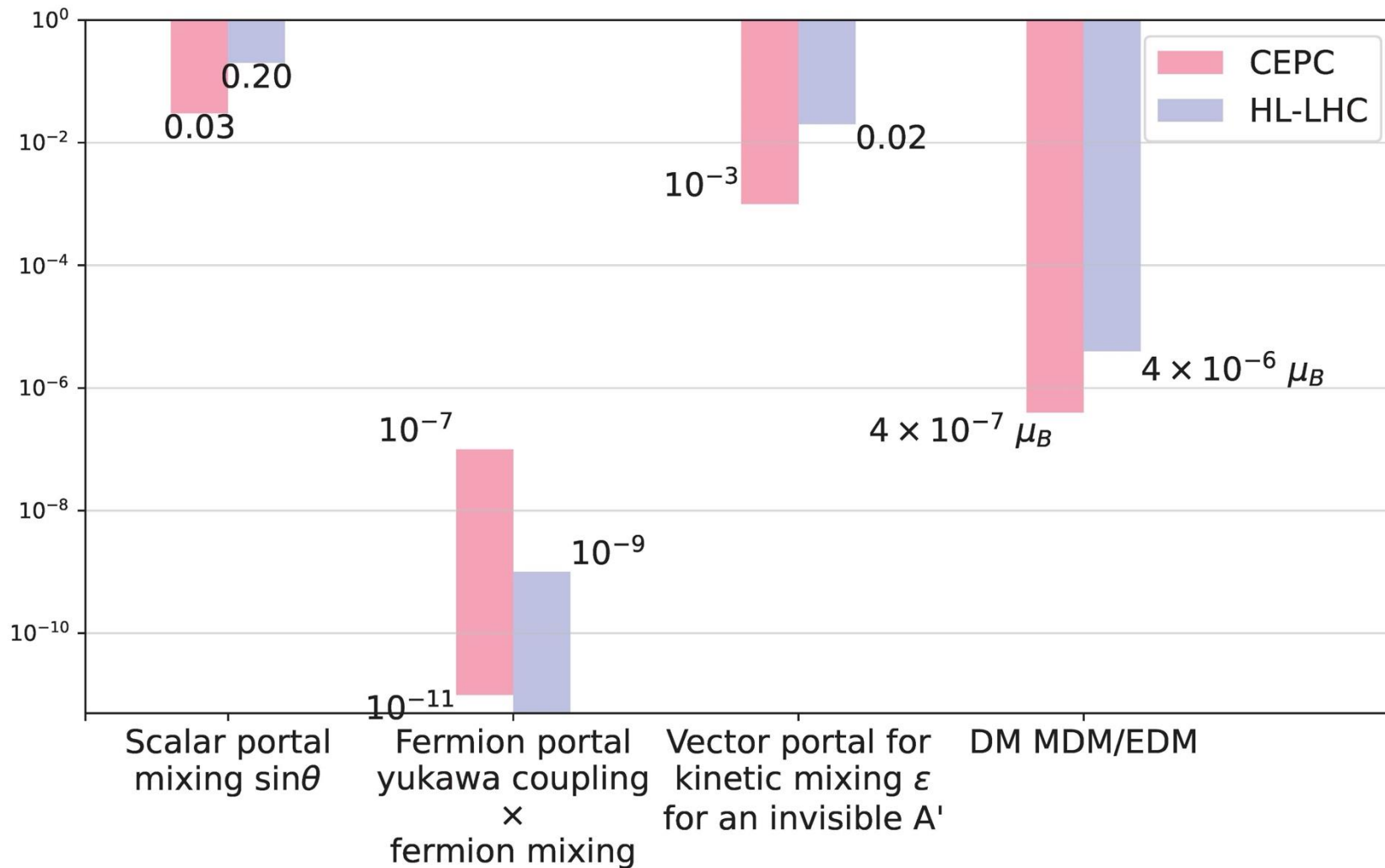
Portal	Effective operator	\sqrt{s} [GeV]	$\mathcal{L}[ab^{-1}]$	Sensitivity of CEPC (HL-LHC)	Figs.	Ref.
Scalar	$\lambda_{HP} H ^2 S^2 \rightarrow$ scalar mixing $\sin \theta$	250	5	invisible S, $\sin \theta \approx 0.03$ (0.20 global-fits)	22	[108]
Fermion	$y_\ell \bar{\chi}_L S^\dagger \ell_R + \text{H.c.}$	250	5	covering $100 \text{ GeV} < m_S < 170 \text{ GeV}$	23	[56]
	$\kappa \Phi q_L^i \ell_R + \text{H.c.}$ (dark QCD)	250	5	$m_\Phi \sim 10 \text{ TeV}$ for $c\tau_{\text{darkpion}} \in [1, 10^3] \text{ cm}$ (Null)	25	[109]
	$y\Phi \bar{F}_L \ell_R + \text{H.c.}$	240	5.6	$y\theta_L \in [10^{-11}, 10^{-7}]$ ($\lesssim 10^{-8} - 10^{-9}$)	26	[110]
Vector	$A'_\mu (e\epsilon J_{\text{em}}^\mu + g_D \bar{\chi} \gamma^\mu \chi)$	250	5	$\epsilon \sim 10^{-3}$ for $g_D = e$ and $m_{A'} < 125 \text{ GeV}$ ($\epsilon \sim 0.02$)	27, 28	[108]
	$\epsilon A_\mu \bar{\chi} \gamma^\mu \chi$, (millicharge DM)	250	5	$\epsilon \sim 0.1$ for $m_\chi \sim 50 \text{ GeV}$	29	[111]
		91.2	2.6	$\epsilon \sim 0.02$ for $m_\chi \sim 5 \text{ GeV}$		
		160	16	$\epsilon \sim 0.5$ for $m_\chi \sim 10 \text{ GeV}$		
	$\frac{1}{2} \mu_\chi \bar{\chi} \sigma^{\mu\nu} \chi F_{\mu\nu} + \frac{i}{2} d_\chi \bar{\chi} \sigma^{\mu\nu} \gamma^5 \chi F_{\mu\nu}$	91.2	100	$\mu_\chi, d_\chi \sim 4 \times 10^{-7}$ (4×10^{-6}) μ_B for $m_\chi < 25 \text{ GeV}$	30	[112]
240		20	$a_\chi, b_\chi \sim 10^{-6}$ (2×10^{-6}) GeV^{-2} for $m_\chi < 80 \text{ GeV}$			
EFT	$\frac{1}{\Lambda^2} \sum_i (\bar{\chi} \gamma_\mu (1 - \gamma_5) \chi) (\bar{\ell} \gamma^\mu (1 - \gamma_5) \ell)$	250	5	$\Lambda_i \sim 2 \text{ TeV}$ ($m_\chi = 0$) (Null)	31	[113]
	$\frac{1}{\Lambda_A^2} \bar{\chi} \gamma_\mu \gamma_5 \chi \bar{\ell} \gamma^\mu \gamma_5 \ell$	250	5	$\Lambda_A \sim 1.5 \text{ TeV}$ (Null)	32	[111]
	$\sum_i \frac{1}{\Lambda_i^2} (\bar{e} \Gamma_\mu e) (\bar{\nu}_L \Gamma^\mu \chi_L) + \text{H.c.}$ $\Gamma_\mu = 1, \gamma_5, \gamma_\mu, \gamma_\mu \gamma_5, \sigma_{\mu\nu}$	240	20	$\Lambda_i \sim 1 \text{ TeV}$ ($m_\chi = 0$) (Null)	33	[114]

new physics process: $e^+e^- \rightarrow \bar{\chi}\chi\gamma$



→ CEPC can probe low-mass light dark states.

2. Dark Matter and Dark Sector

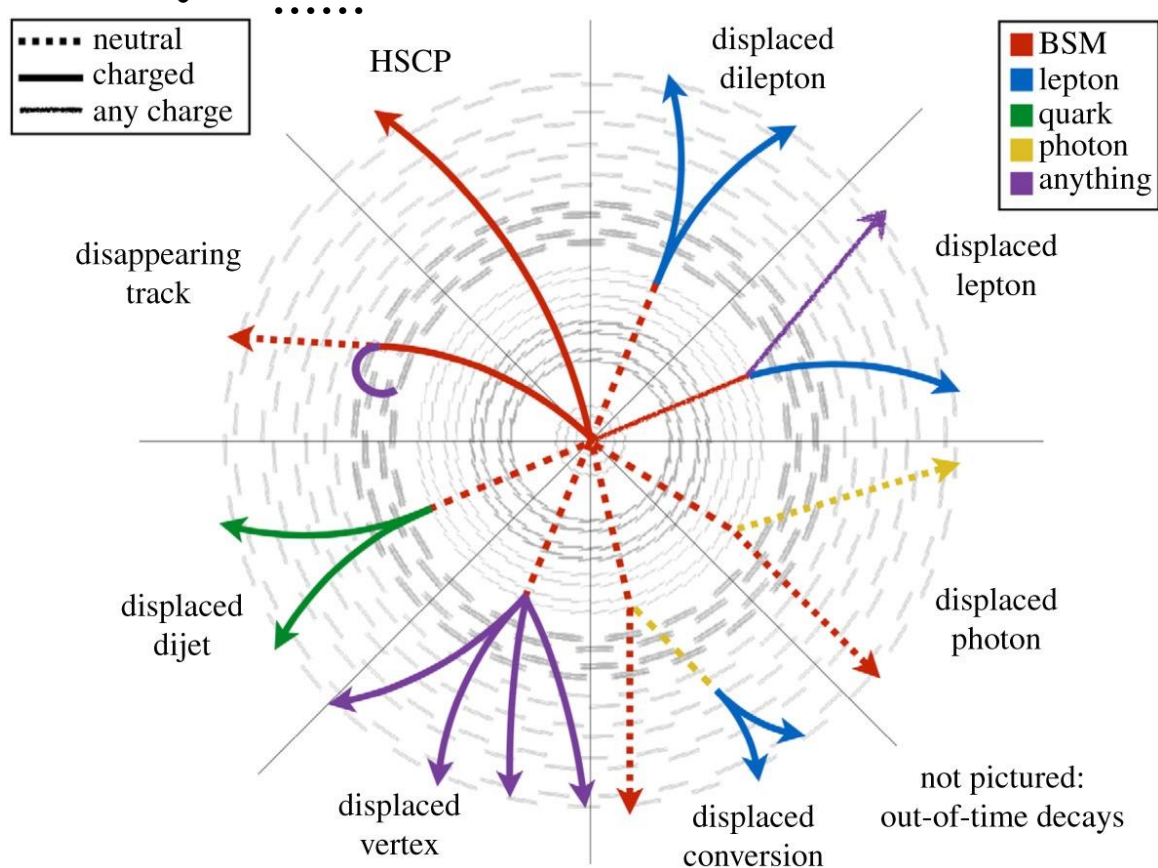


CEPC can improve the sensitivities by roughly one order of magnitude (vs LHC), for some cases.

3. Long-lived particles (LLP)

Long lifetimes result from a few simple physical mechanisms:

- Small couplings (ex. RPV SUSY)
- Limited phase space: small mass splitting (ex. compressed SUSY, ...)
- Heavy intermediate states

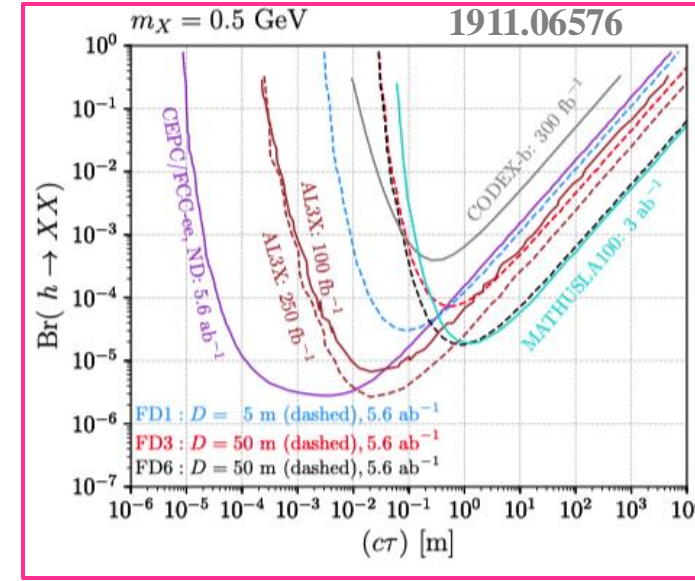


- New scale particles from higgs decay
- SUSY RPV N1 from Z-boson decays
- ALP
- Dark photons
- ...

→ Far Detector can help a lot!

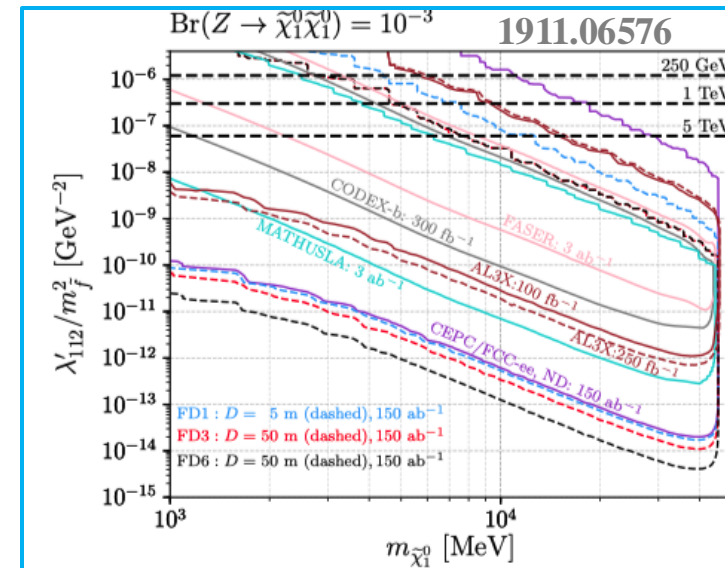
3. Long-lived particles (LLP)

LLP Type	Signal Signature	\sqrt{s} [GeV]	\mathcal{L} [ab ⁻¹]	Detector	Sensitivities on parameters [Assumptions]	Figs.	Refs.
New scalar particles (X)	$Z(\rightarrow \text{incl.}) h(\rightarrow XX), X \rightarrow q\bar{q}/\nu\bar{\nu}$	240	20	ND	$\text{Br}(h \rightarrow XX) \sim 10^{-6}$ [$m \in (1, 50)$ GeV, $\tau \in (10^{-3}, 10^{-1})$ ns]	37	[80]
	$Z(\rightarrow \text{incl.}) h(\rightarrow XX), X \rightarrow \text{incl.}$	240	5.6	ND	$\text{Br}(h \rightarrow XX) \sim 3 \times 10^{-6}$ [$m = 0.5$ GeV, $c\tau \sim 5 \times 10^{-3}$ m]	49	[86]
				FD3	$\text{Br}(h \rightarrow XX) \sim 7 \times 10^{-5}$ [$m = 0.5$ GeV, $c\tau \sim 1$ m]	49	[86]
				LAYCAST	$\text{Br}(h \rightarrow XX) \sim 5 \times 10^{-6}$ [$m = 0.5$ GeV, $c\tau \sim 10^{-1}$ m]	49	[241]
RPV-SUSY neutralinos ($\tilde{\chi}_1^0$)	$Z \rightarrow \tilde{\chi}_1^0 \tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow \text{incl.}$	91.2	150	ND	$\lambda'_{112}/m_f^2 \in (2 \times 10^{-14}, 10^{-8})$ GeV ⁻² [$m \sim 40$ GeV, $\text{Br}(Z \rightarrow \tilde{\chi}_1^0 \tilde{\chi}_1^0) = 10^{-3}$]	43	[86]
				FD3	$\lambda'_{112}/m_f^2 \in (10^{-14}, 10^{-9})$ GeV ⁻² [$m \sim 40$ GeV, $\text{Br}(Z \rightarrow \tilde{\chi}_1^0 \tilde{\chi}_1^0) = 10^{-3}$]	50	[86]
				LAYCAST	$\lambda'_{112}/m_f^2 \in (7 \times 10^{-15}, 10^{-9})$ GeV ⁻² [$m \sim 40$ GeV, $\text{Br}(Z \rightarrow \tilde{\chi}_1^0 \tilde{\chi}_1^0) = 10^{-3}$]	50	[241]
ALPs (a)	$Z^{(*)} \rightarrow \mu^- \mu^+ a$	91	150	ND	$f_a/C_{\mu\mu}^A \lesssim 950$ GeV	44	[85]
	$\gamma a, a \rightarrow \gamma\gamma$	91.2	150	ND	$C_{\gamma\gamma}/\Lambda \sim 10^{-3}$ TeV ⁻¹ [$C_{\gamma Z} = 0, m \sim 2$ GeV]	51	[241]
				FD3	$C_{\gamma\gamma}/\Lambda \sim 6 \times 10^{-3}$ TeV ⁻¹ [$C_{\gamma Z} = 0, m \sim 0.3$ GeV]	51	[242]
				LAYCAST	$C_{\gamma\gamma}/\Lambda \sim 2 \times 10^{-3}$ TeV ⁻¹ [$C_{\gamma Z} = 0, m \sim 0.7$ GeV]	51	[241]
Hidden valley particles (π_V^0)	$Z h(\rightarrow \pi_V^0 \pi_V^0), \pi_V^0 \rightarrow b\bar{b}$	350	1.0	ND	$\sigma(h) \times \text{BR}(h \rightarrow \pi_V^0 \pi_V^0) \sim 10^{-4}$ pb [$m \in (25, 50)$ GeV, $\tau \sim 10^2$ ps]	41	[243]
Dark photons (γ_D)	$Z(\rightarrow q\bar{q}) h(\rightarrow \gamma_D \gamma_D), \gamma_D \rightarrow \ell^- \ell^+ / q\bar{q}$	250	2.0	ND	$\text{Br}(h \rightarrow \gamma_D \gamma_D) \sim 10^{-5}$, [$m \in (5, 10)$ GeV, $\tau \sim 10^2$ ps, $\epsilon \in (10^{-6}, 10^{-7})$]	42	[83]



Light Scalars from Exotic Higgs Decays

FD can extend and complement the sensitivity to the LLPs compared with Near Detector

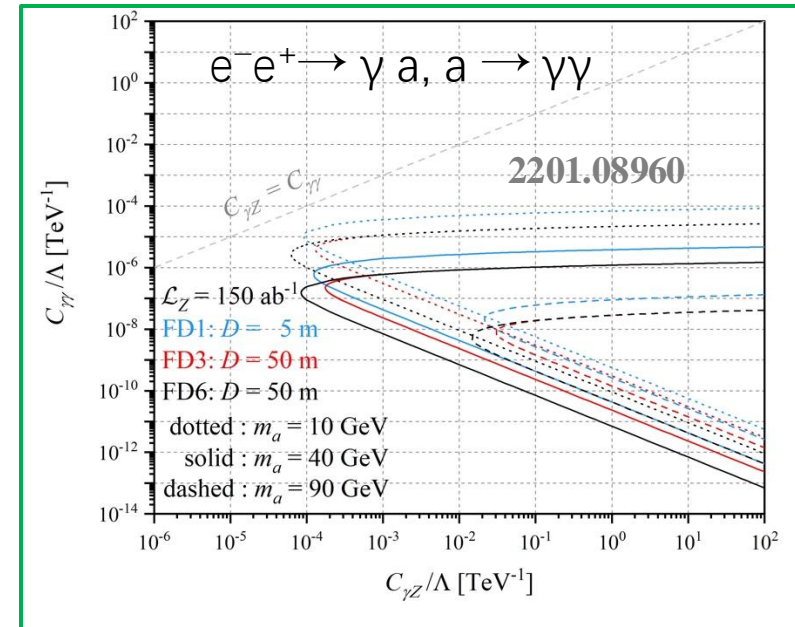


SUSY RPV Neutralino1 from Z Decays

3. Long-lived particles (LLP)

LLP Type	Signal Signature	\sqrt{s} [GeV]	\mathcal{L} [ab $^{-1}$]	Detector	Sensitivities on parameters [Assumptions]	Figs.	Refs.
New scalar particles (X)	$Z(\rightarrow \text{incl.}) h(\rightarrow XX),$ $X \rightarrow q\bar{q}/\nu\bar{\nu}$	240	20	ND	$\text{Br}(h \rightarrow XX) \sim 10^{-6}$ [$m \in (1, 50)$ GeV, $\tau \in (10^{-3}, 10^{-1})$ ns]	37	[80]
	$Z(\rightarrow \text{incl.}) h(\rightarrow XX),$ $X \rightarrow \text{incl.}$	240	5.6	ND	$\text{Br}(h \rightarrow XX) \sim 3 \times 10^{-6}$ [$m = 0.5$ GeV, $c\tau \sim 5 \times 10^{-3}$ m]	49	[86]
				FD3	$\text{Br}(h \rightarrow XX) \sim 7 \times 10^{-5}$ [$m = 0.5$ GeV, $c\tau \sim 1$ m]	49	[86]
				LAYCAST	$\text{Br}(h \rightarrow XX) \sim 5 \times 10^{-6}$ [$m = 0.5$ GeV, $c\tau \sim 10^{-1}$ m]	49	[241]
RPV-SUSY neutralinos ($\tilde{\chi}_1^0$)	$Z \rightarrow \tilde{\chi}_1^0 \tilde{\chi}_1^0,$ $\tilde{\chi}_1^0 \rightarrow \text{incl.}$	91.2	150	ND	$\lambda'_{112}/m_f^2 \in (2 \times 10^{-14}, 10^{-8})$ GeV $^{-2}$ [$m \sim 40$ GeV, $\text{Br}(Z \rightarrow \tilde{\chi}_1^0 \tilde{\chi}_1^0) = 10^{-3}$]	43	[86]
				FD3	$\lambda'_{112}/m_f^2 \in (10^{-14}, 10^{-9})$ GeV $^{-2}$ [$m \sim 40$ GeV, $\text{Br}(Z \rightarrow \tilde{\chi}_1^0 \tilde{\chi}_1^0) = 10^{-3}$]	50	[86]
				LAYCAST	$\lambda'_{112}/m_f^2 \in (7 \times 10^{-15}, 10^{-9})$ GeV $^{-2}$ [$m \sim 40$ GeV, $\text{Br}(Z \rightarrow \tilde{\chi}_1^0 \tilde{\chi}_1^0) = 10^{-3}$]	50	[241]
ALPs (a)	$Z^{(*)} \rightarrow \mu^- \mu^+ a$	91	150	ND	$f_a/C_{\mu\mu}^A \lesssim 950$ GeV	44	[85]
	$\gamma a,$ $a \rightarrow \gamma\gamma$	91.2	150	ND	$C_{\gamma\gamma}/\Lambda \sim 10^{-3}$ TeV $^{-1}$ [$C_{\gamma Z} = 0, m \sim 2$ GeV]	51	[241]
				FD3	$C_{\gamma\gamma}/\Lambda \sim 6 \times 10^{-3}$ TeV $^{-1}$ [$C_{\gamma Z} = 0, m \sim 0.3$ GeV]	51	[242]
				LAYCAST	$C_{\gamma\gamma}/\Lambda \sim 2 \times 10^{-3}$ TeV $^{-1}$ [$C_{\gamma Z} = 0, m \sim 0.7$ GeV]	51	[241]
Hidden valley particles (π_V^0)	$Z h(\rightarrow \pi_V^0 \pi_V^0),$ $\pi_V^0 \rightarrow b\bar{b}$	350	1.0	ND	$\sigma(h) \times \text{BR}(h \rightarrow \pi_V^0 \pi_V^0) \sim 10^{-4}$ pb [$m \in (25, 50)$ GeV, $\tau \sim 10^2$ ps]	41	[243]
Dark photons (γ_D)	$Z(\rightarrow q\bar{q}) h(\rightarrow \gamma_D \gamma_D),$ $\gamma_D \rightarrow \ell^- \ell^+ / q\bar{q}$	250	2.0	ND	$\text{Br}(h \rightarrow \gamma_D \gamma_D) \sim 10^{-5},$ [$m \in (5, 10)$ GeV, $\tau \sim 10^2$ ps, $\epsilon \in (10^{-6}, 10^{-7})$]	42	[83]

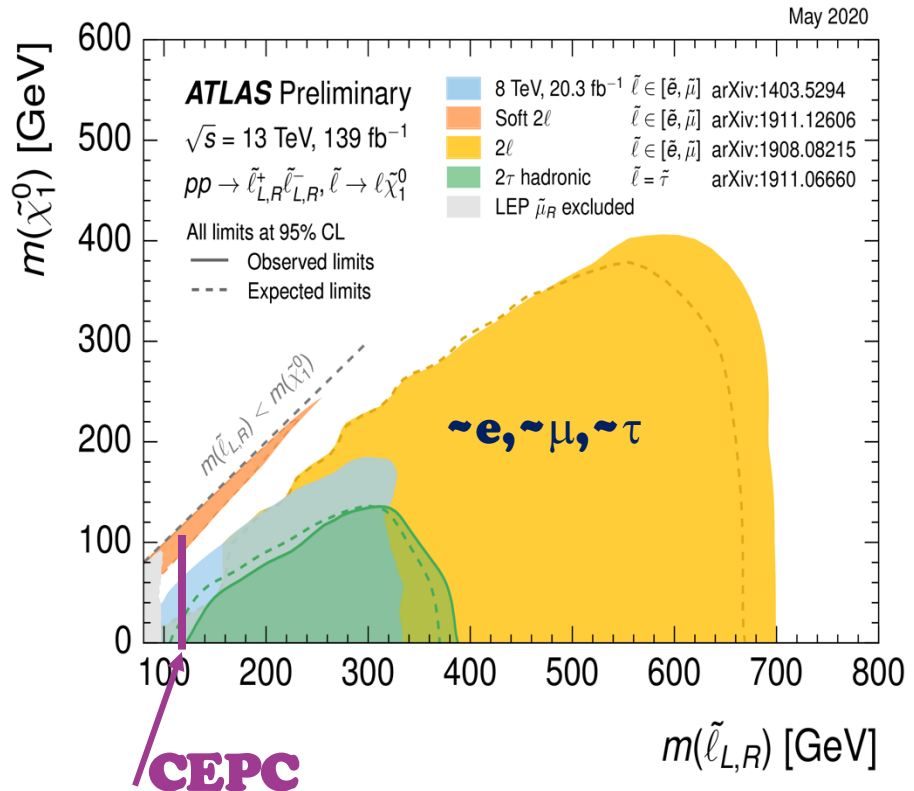
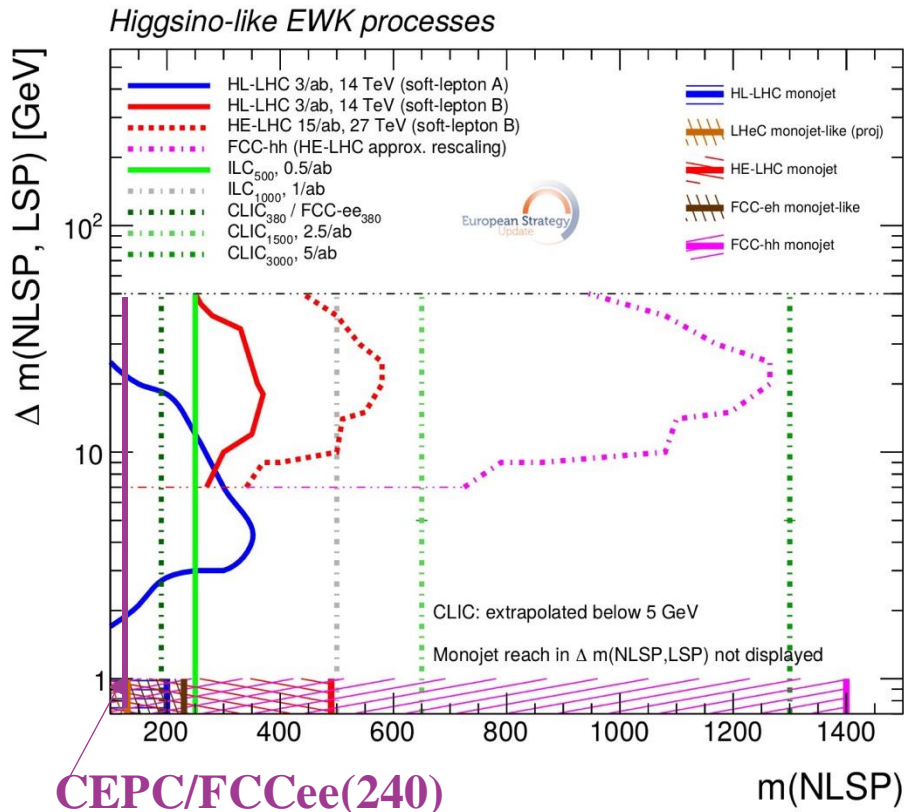
Good sensitivity for ALP



Axion-like Particles

4. SUSY Searches at CEPC

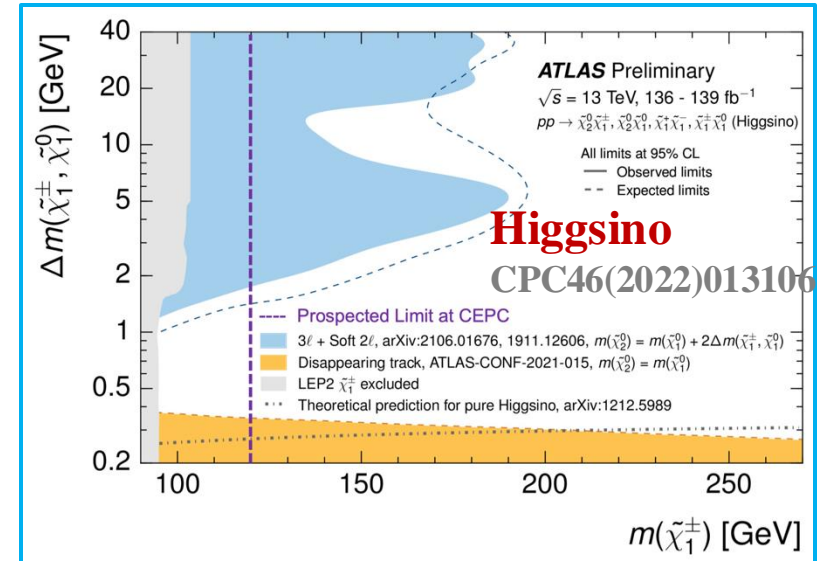
- **SUSY: establishes a symmetry between fermions and bosons, solve many big questions: unification, DM, Hierarchy,**
- **Complementary with LHC: lower mass/soft energy region**
 - ✓ **Mainly light EWKino and slepton for CEPC**



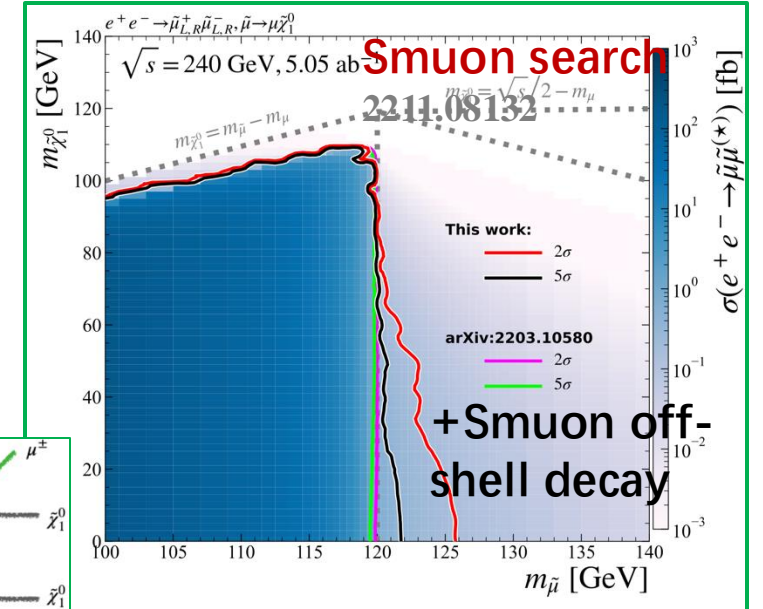
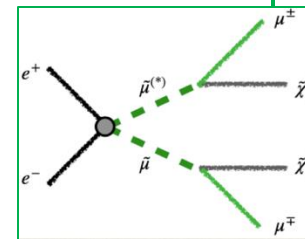
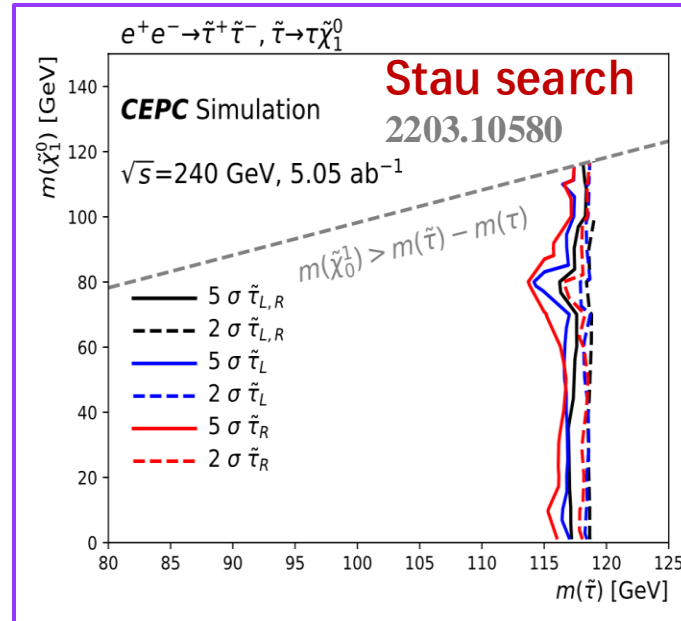
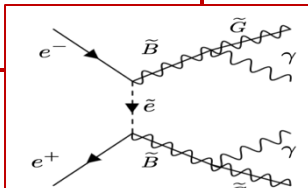
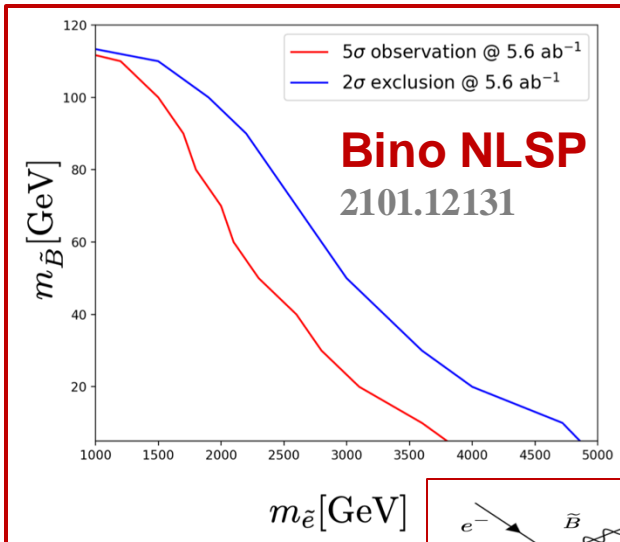
Lepton collider: discovery in all scenarios up to kinematic limit: $\sqrt{s}/2$

4. SUSY Searches at CEPC

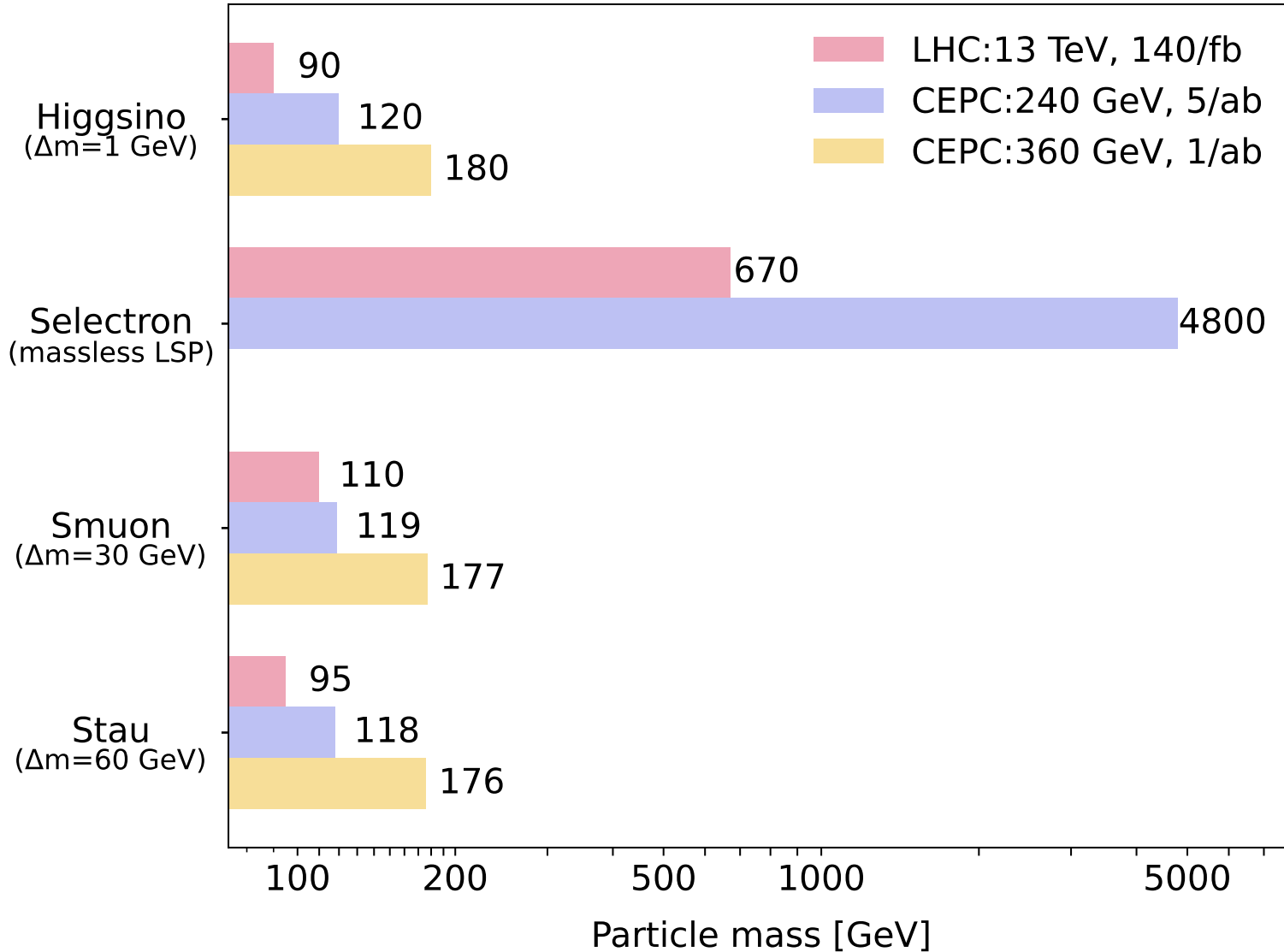
Search	Production	\sqrt{s} [GeV]	\mathcal{L} [ab ⁻¹]	Sensitivity	Figs.	Ref.
Light electroweakino	chargino pair	240	5.05	chargino excluded up to 120 GeV	57	[339]
	$e^+e^- \rightarrow \tilde{B}\tilde{B} \rightarrow \gamma\gamma\tilde{G}\tilde{G}$.	240	5.6	selectron excluded up to 4.5 TeV	58	[341]
Light slepton	smuon pair	240	5.05	smuon excluded up 118 GeV	59	[342]
	stau pair	240	5.05	stau excluded up 117 GeV	59	[342]
	smuon pair	360	1	smuon excluded up 178 GeV	59	□
	stau pair	360	1	stau excluded up 175 GeV	59	□
	off-shell smuon pair	240	5	smuon excluded up 126 GeV	61	[344]
	$e_R^+e_R^- \rightarrow \tilde{\chi}_1^0(\text{bino}) + \tilde{\chi}_1^0(\text{bino}) + \gamma$	240	3	right-handed selectron excluded up to 210 GeV	60	[343]
$\mathcal{F}\text{-}SU(5)$	-	-	upper limits on $\tilde{\tau}_1$ up to 115 GeV	62	[345]	
$\mathcal{F}\text{-}SU(5)$	-	-	upper limits on \tilde{e}_R up to 150 GeV	62	[345]	



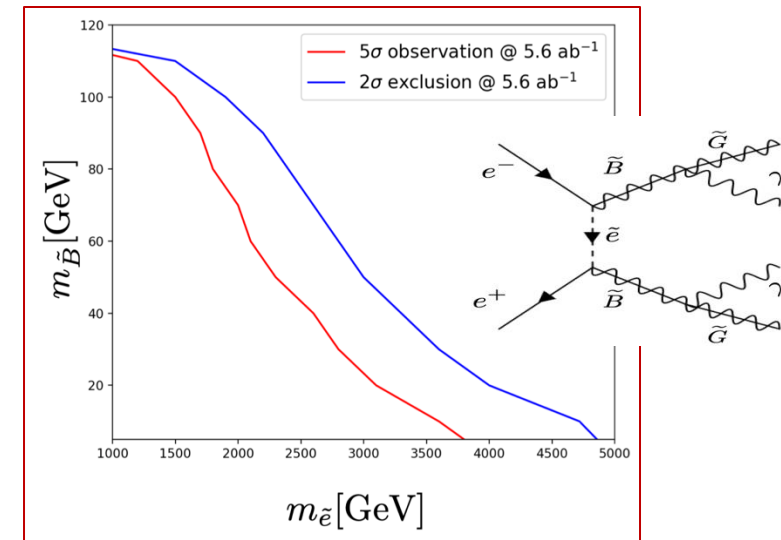
$$\tilde{B} \rightarrow \tilde{G}/\tilde{a} + \gamma$$



4. SUSY Searches at CEPC



- **Light EWKinos/sleptons: discovery in all scenarios up to kinematic limit $\sqrt{s}/2$**
- **Heavy selectron from t-channel**



5. Flavor portal NP

- CEPC is also a flavor factory (b,c,tau) when running at Z pole, which has a unique sensitivity for some rare processes due to suppression in SM
- The sensitivity of the flavor sector to new physics is underscored by several factors:
 - cLFV processes
 - Decays of b and c hadrons
 - Light BSM degrees of freedom from flavor transitions (cLFV or quark FCNC processes) with inv. BSM states or LLP

Measurement	Current Limit	CEPC [373]
BR($Z \rightarrow \tau\mu$)	$< 6.5 \times 10^{-6}$	$\mathcal{O}(10^{-9})$
BR($Z \rightarrow \tau e$)	$< 5.0 \times 10^{-6}$	$\mathcal{O}(10^{-9})$
BR($Z \rightarrow \mu e$)	$< 7.5 \times 10^{-7}$	$10^{-8} - 10^{-10}$
BR($\tau \rightarrow \mu\mu\mu$)	$< 2.1 \times 10^{-8}$	$\mathcal{O}(10^{-10})$
BR($\tau \rightarrow eee$)	$< 2.7 \times 10^{-8}$	$\mathcal{O}(10^{-10})$
BR($\tau \rightarrow e\mu\mu$)	$< 2.7 \times 10^{-8}$	$\mathcal{O}(10^{-10})$
BR($\tau \rightarrow \mu ee$)	$< 1.8 \times 10^{-8}$	$\mathcal{O}(10^{-10})$
BR($\tau \rightarrow \mu\gamma$)	$< 4.4 \times 10^{-8}$	$\mathcal{O}(10^{-10})$
BR($\tau \rightarrow e\gamma$)	$< 3.3 \times 10^{-8}$	$\mathcal{O}(10^{-10})$
BR($B_s \rightarrow \phi\nu\bar{\nu}$)	$< 5.4 \times 10^{-3}$	$\lesssim 1\%$ (relative)
BR($B^0 \rightarrow K^{*0}\tau^+\tau^-$)	-	$\lesssim \mathcal{O}(10^{-6})$
BR($B_s \rightarrow \phi\tau^+\tau^-$)	-	$\lesssim \mathcal{O}(10^{-6})$
BR($B^+ \rightarrow K^+\tau^+\tau^-$)	$< 2.25 \times 10^{-3}$	$\lesssim \mathcal{O}(10^{-6})$
BR($B_s \rightarrow \tau^+\tau^-$)	$< 6.8 \times 10^{-3}$	$\lesssim \mathcal{O}(10^{-5})$
BR($B^0 \rightarrow 2\pi^0$)	$\pm 16\%$ (relative)	$\pm 0.25\%$ (relative)
$C_{CP}(B^0 \rightarrow 2\pi^0)$	± 0.22 (relative)	± 0.01 (relative)
BR($B_c \rightarrow \tau\nu$)	$\lesssim 30\%$	$\pm 0.5\%$ (relative)
BR($B_c \rightarrow J/\psi\tau\nu$)/BR($B_c \rightarrow J/\psi\mu\nu$)	$\pm 0.17 \pm 0.18$	$\pm 2.5\%$ (relative)
BR($B_s \rightarrow D_s^{(*)}\tau\nu$)/BR($B_s \rightarrow D_s^{(*)}\mu\nu$)	-	$\pm 0.2\%$ (relative)
BR($\Lambda_b \rightarrow \Lambda_c\tau\nu$)/BR($B_c \rightarrow \Lambda_c\mu\nu$)	± 0.076	$\pm 0.05\%$ (relative)
BR($\tau \rightarrow \mu X_{\text{inv.}}$)	7×10^{-4}	$(3-5) \times 10^{-6}$
BR($B \rightarrow \mu X_{\text{LLP}}(\rightarrow \mu\mu)$)	-	$\mathcal{O}(10^{-10})$ (optimal)

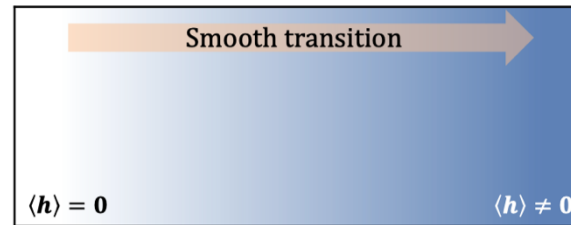
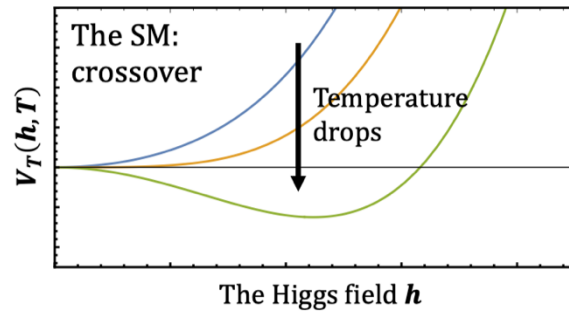
Preliminary sensitivities of BSM flavor physics probes at CEPC

> two orders of magnitude improv. 18

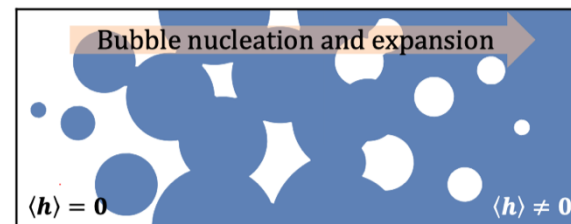
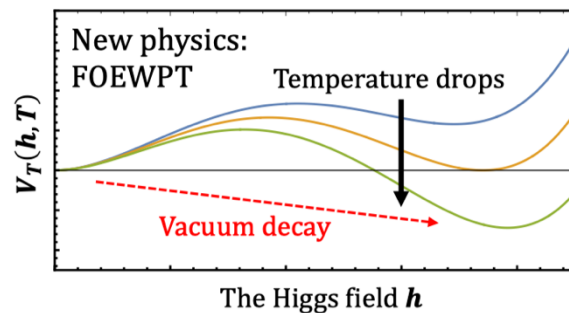
6. EWPT at CEPC

■ The nature of Electroweak Phase Transition (EWPT) deeply impacts the thermal history of the Universe, closely linked to puzzles of DM, matter-antimatter asymmetry

- Probing the nature of EWPT at colliders



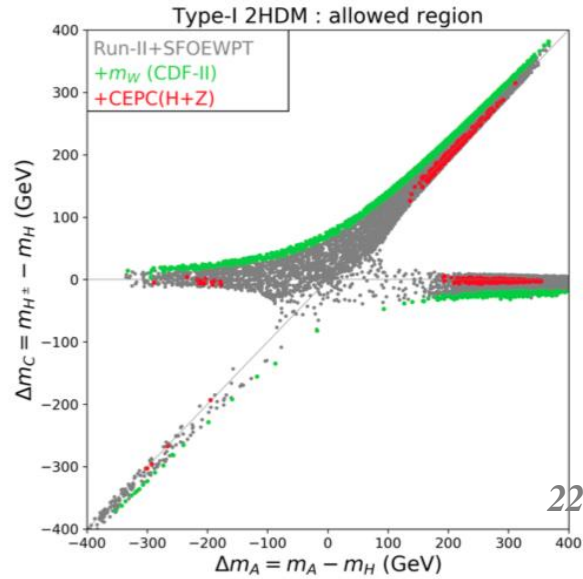
In the SM, the transition is a smooth crossover



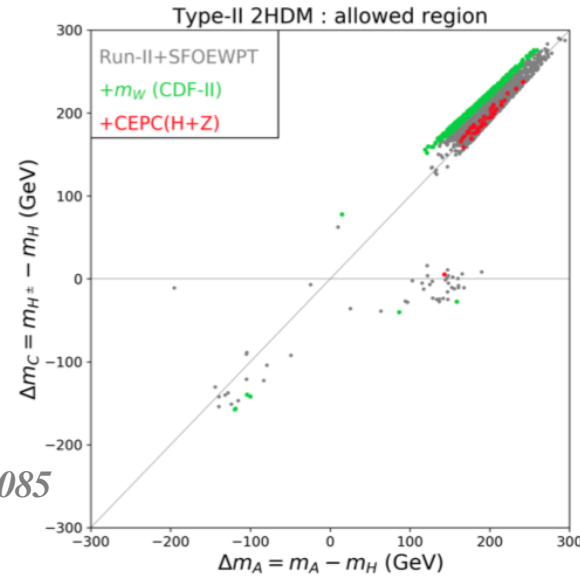
In NP, the scalar potential exhibits a barrier, allowing for a FOEWPT with bubble nucleation and expansion

- Higgs precision measurements
- Higgs exotic decay

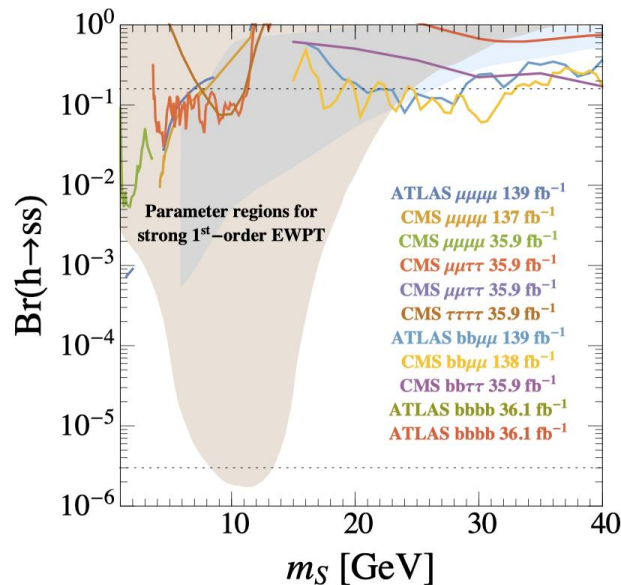
6. EWPT at CEPC



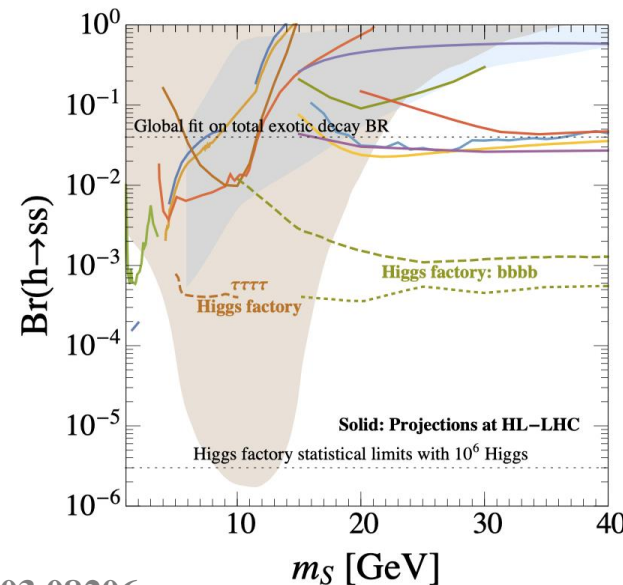
2204.05085



Under current constraints, both Type-I and Type-II 2HDM can explain the SFOEWPT, **Z-pole, Higgs precision measurements** and **mW precision measurement** of CDF-II at same time.



2203.08206

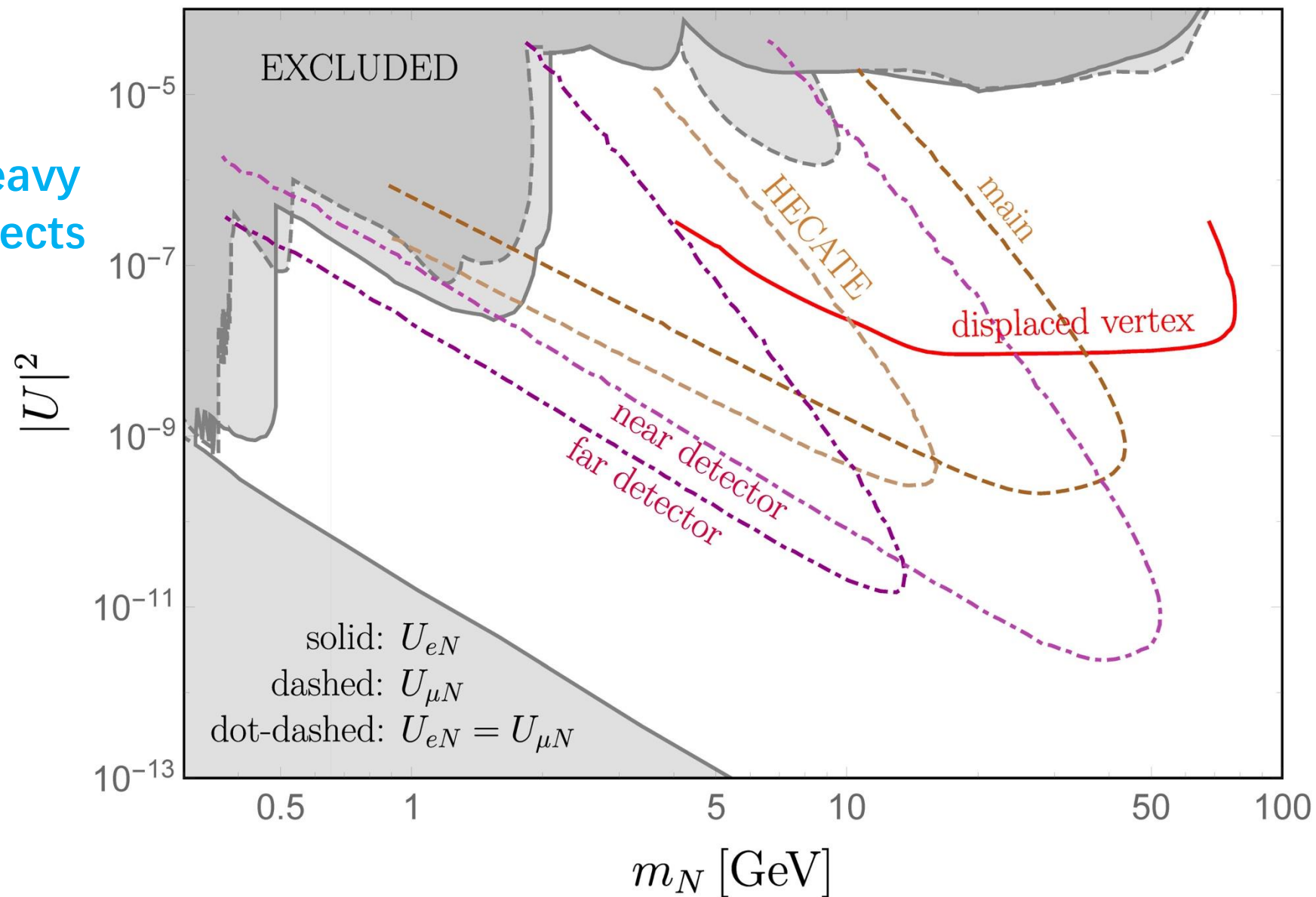


Higgs exotic decay $h \rightarrow ss \rightarrow XXYY$ as a probe for the FOEWPT:

CEPC has the potential to probe almost the entire FOEWPT parameter space for **4b** and **4tau** channels

7. Neutrino physics

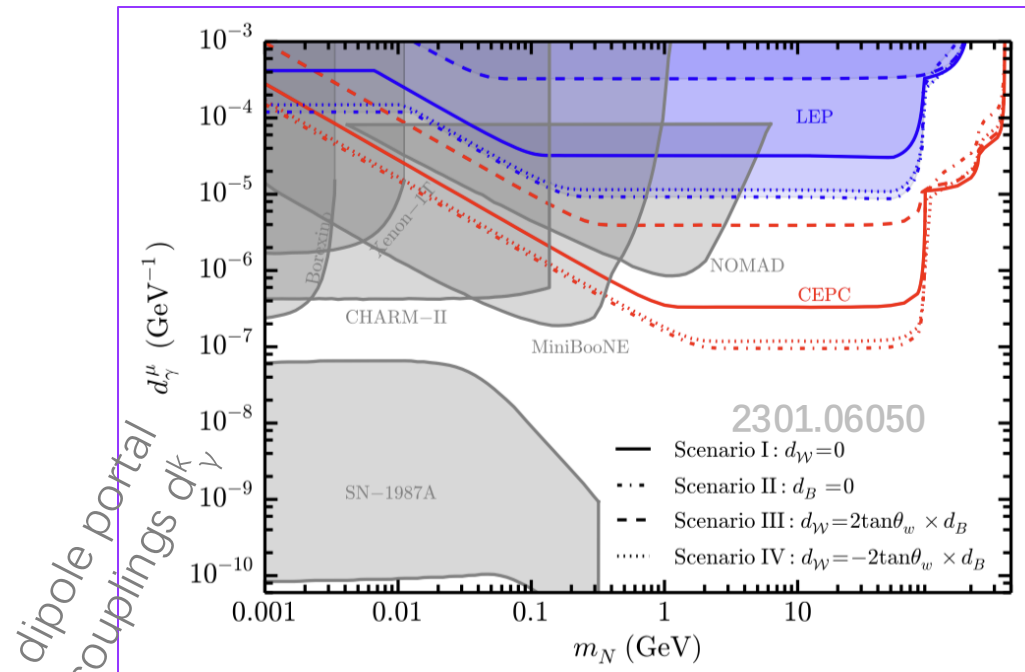
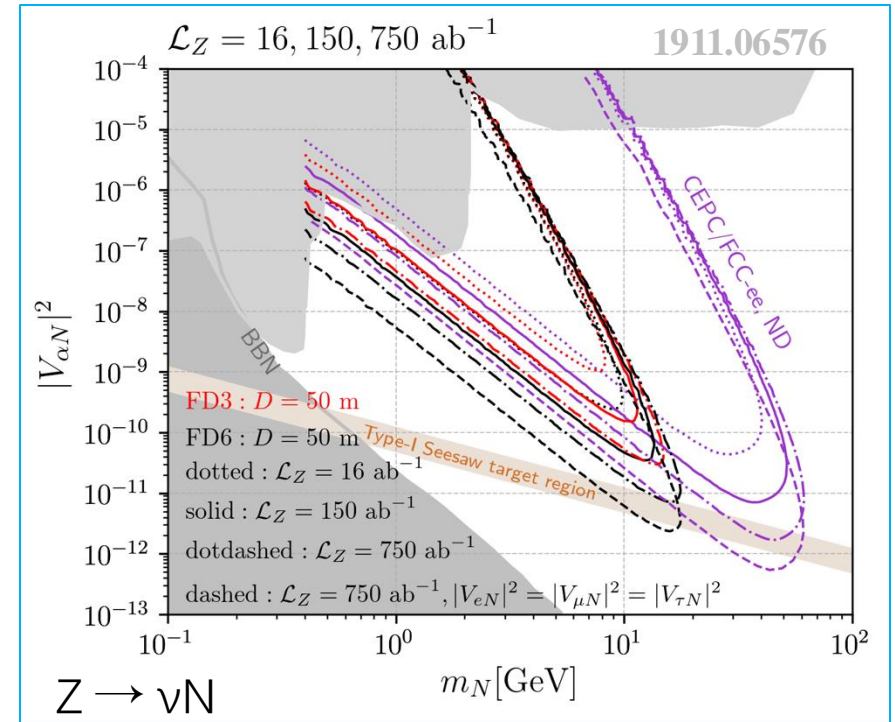
Summary of heavy neutrino prospects
@ CEPC



7. Neutrino physics

BSM related neutrino physics from neutrino mass mechanism, new messengers and interactions at EW scale:

- **Heavy neutrino (@ND, FD)**
- **Non-standard neutrino interactions**
- **Active-sterile neutrino transition magnetic moments**
- **Neutral and doubly-charged scalars in seesaw models**
- **Connection to leptogenesis (collider probes) and dark matter (sterile neutrino in the ν MSM)**

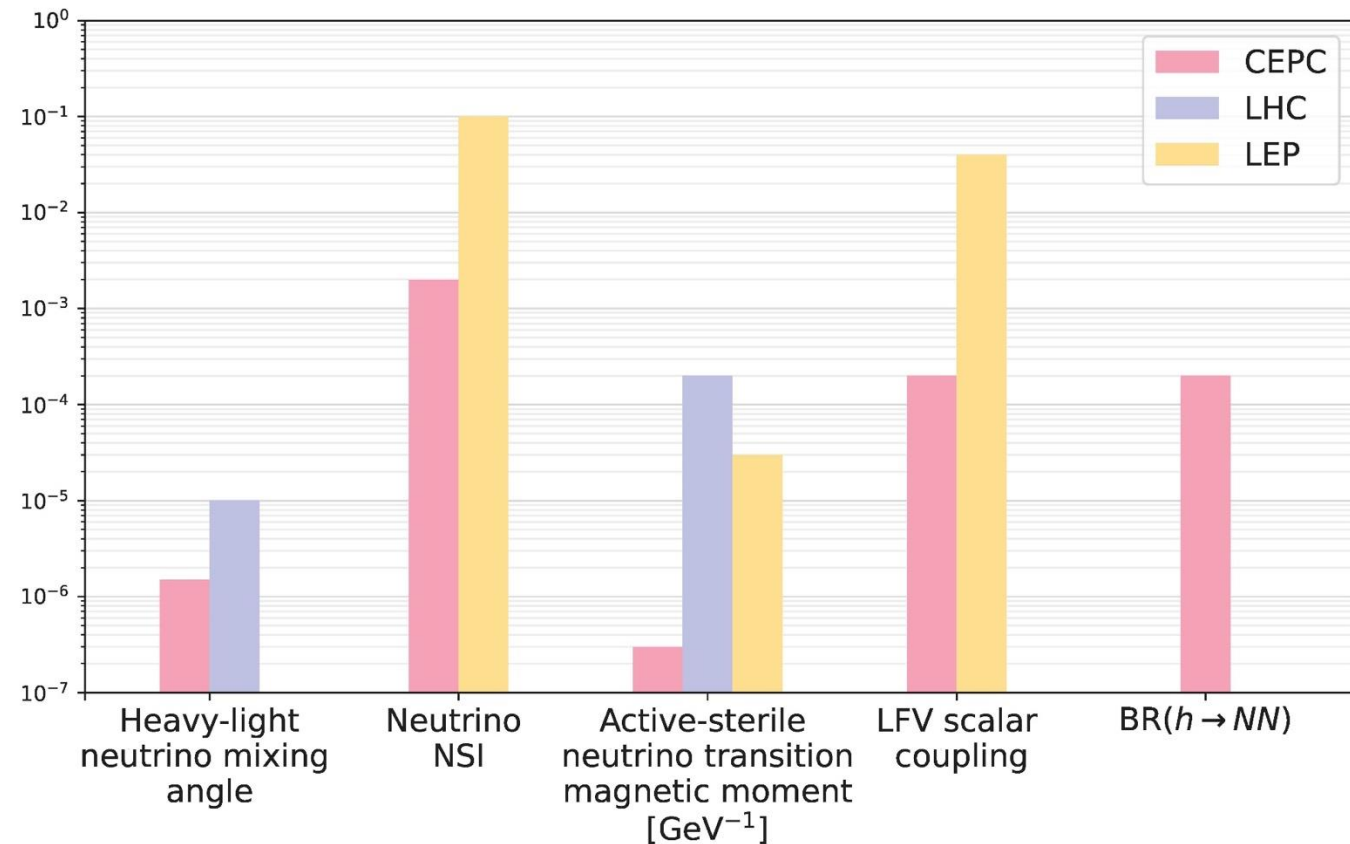


7. Neutrino physics

BSM related neutrino physics from neutrino mass mechanism, new messengers and interactions at EW scale:

- **Heavy neutrino (@ND, FD)**
- **Non-standard neutrino interactions**
- **Active-sterile neutrino transition magnetic moments**
- **Neutral and doubly-charged scalars in seesaw models**
- **Connection to leptogenesis (collider probes) and dark matter (sterile neutrino in the ν MSM)**

Summary plot of neutrino relevant models

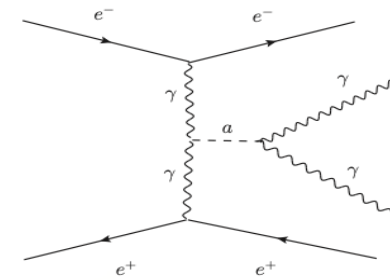


The sensitivities can be improved by roughly 1 to 2 (or more) orders of magnitude (vs LHC & LEP), for some cases.

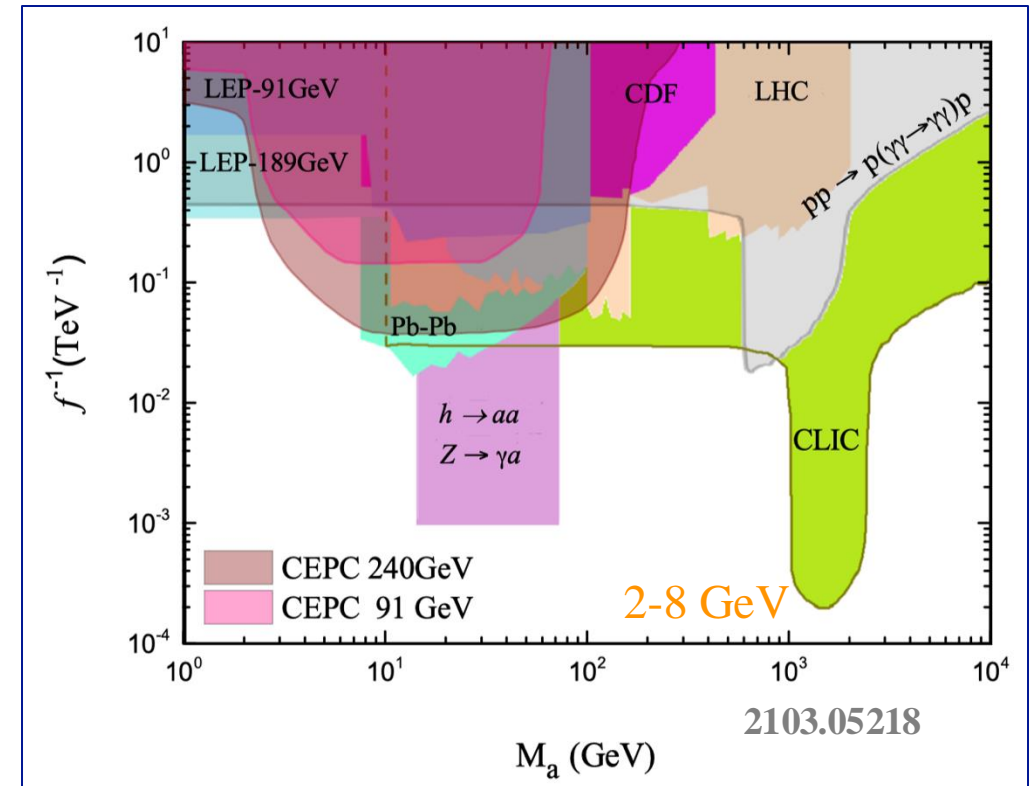
8. More exotics

High precision of Z, h width offers power test of exotics process of Lepton number/flavor violation, Sterile states, Axion-like particles ...

- Axion-like particles (solve “strong-CP” problem)
- Lepton form factors (μ / e g-2, μ / e dipole moments in SUSY, τ weak-electric dipole moments)
- Emergent Hadron Mass
- Exotic lepton mass models
- Spin entanglement
-



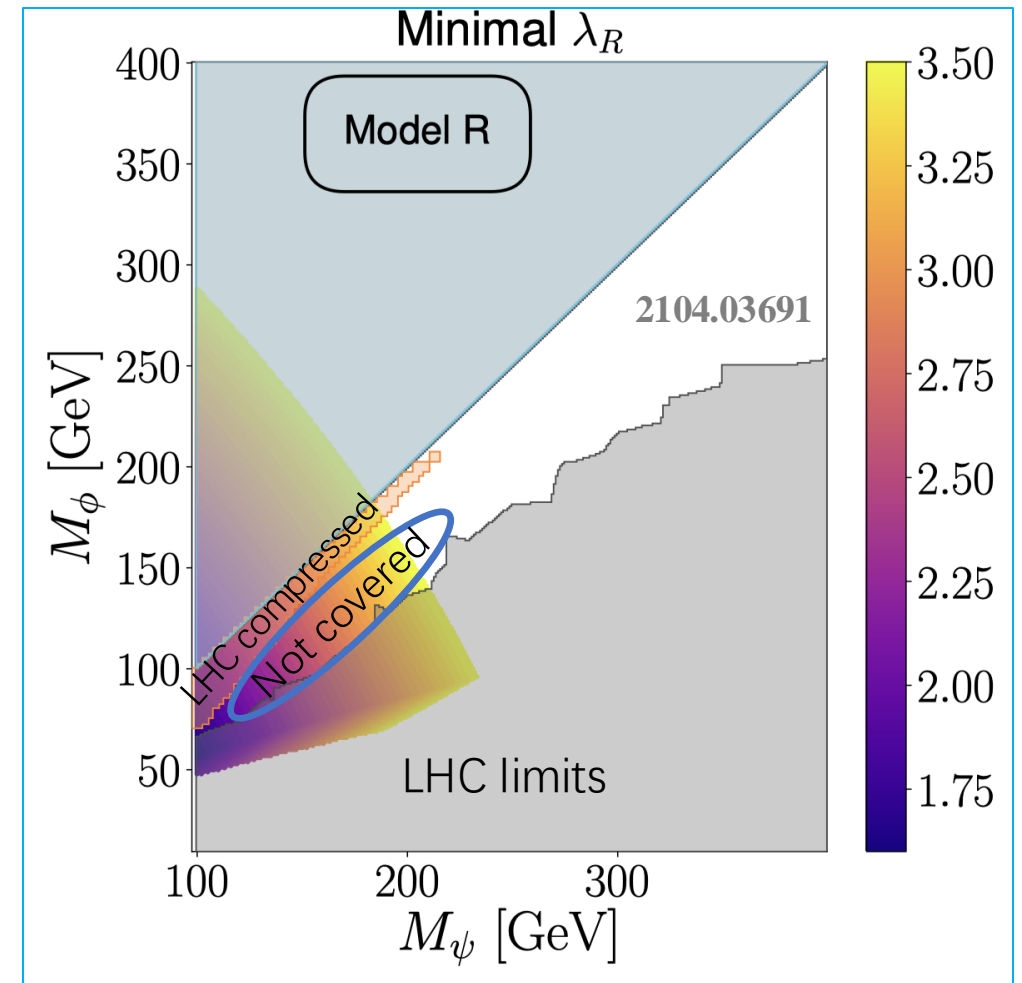
$$e^+e^- \rightarrow \gamma\gamma e^+e^-$$



8. More exotics

High precision of Z, h width offers power test of exotics process of Lepton number/flavor violation, Sterile states, Axion-like particles ...

- Axion-like particles (solve “strong-CP” problem)
- Lepton form factors ($\mu/e g-2$, μ/e dipole moments in SUSY, τ weak-electric dipole moments)
- Emergent Hadron Mass
- Exotic lepton mass models
- Spin entanglement
-
 - Light EWKinos, smuon, stau co-annihilation can explain mu g-2 excess
 - Gaps from LHC, can cover by CEPC



A simple model with a new scalar and and a new fermion

8. More exotics

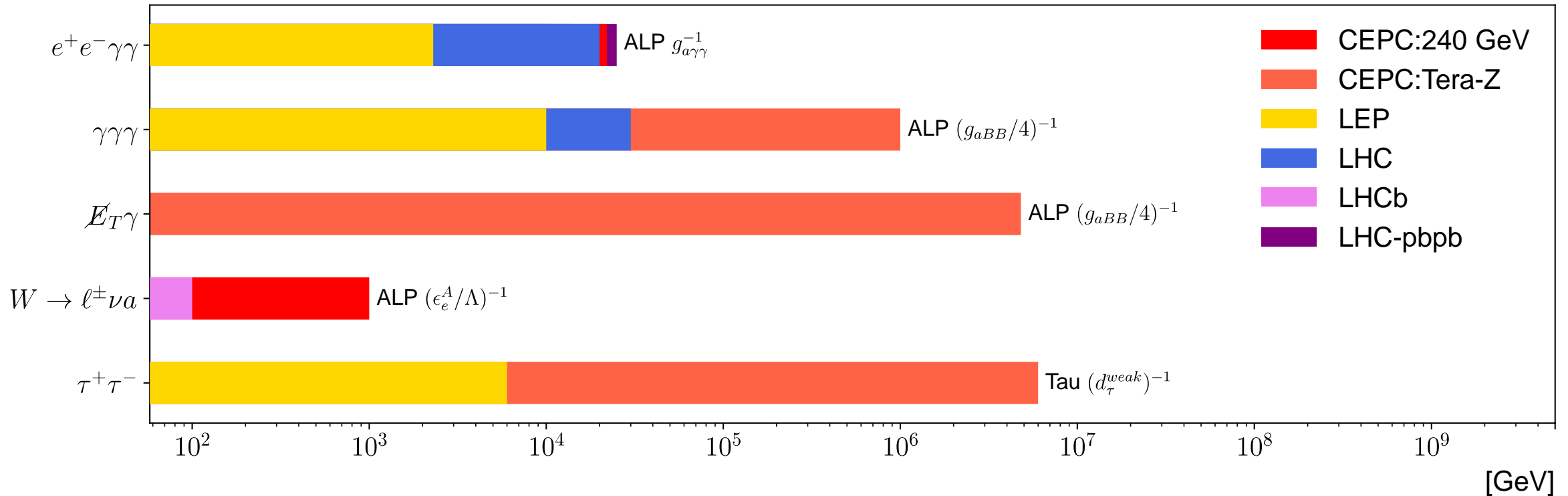
High precision of Z, h width offers power test of exotics process of Lepton number/flavor violation, Sterile states, Axion-like particles ...

- Axion-like particles (solve “strong-CP” problem)
- Lepton form factors (μ / e g-2, μ / e dipole moments in SUSY, τ weak-electric dipole moments)
- Emergent Hadron Mass
- Exotic lepton mass models
- Spin entanglement
-

Projected energy scale sensitivities via exotic searches at the CEPC

Quantity	Channel	Sensitivity scale (GeV)	CEPC Run
ALP $g_{a\gamma\gamma}^{-1}$	$e^+e^-\gamma\gamma$	6.7×10^3 [637]	Tera-Z
	$e^+e^-\gamma\gamma$	2.2×10^4 [637]	240 GeV
ALP $(g_{aBB}/4)^{-1}$	$\bar{f}fa$	6.5×10^3 [637]	250 GeV
	3γ	10^6 [72]	Tera-Z
	$\cancel{E}_T\gamma$	4.8×10^6 [72]	Tera-Z
ALP $(\epsilon_e^A/\Lambda)^{-1}$	$W \rightarrow \ell^\pm \nu a$	10^3 [639]	240 GeV
Tau $(d_\tau^{weak})^{-1}$	$\tau^+\tau^-$	6×10^6 [667]	Tera-Z
Bell Inequality	$Z, h \rightarrow \tau^+\tau^-$	1σ [694]	240 GeV

8. More exotics



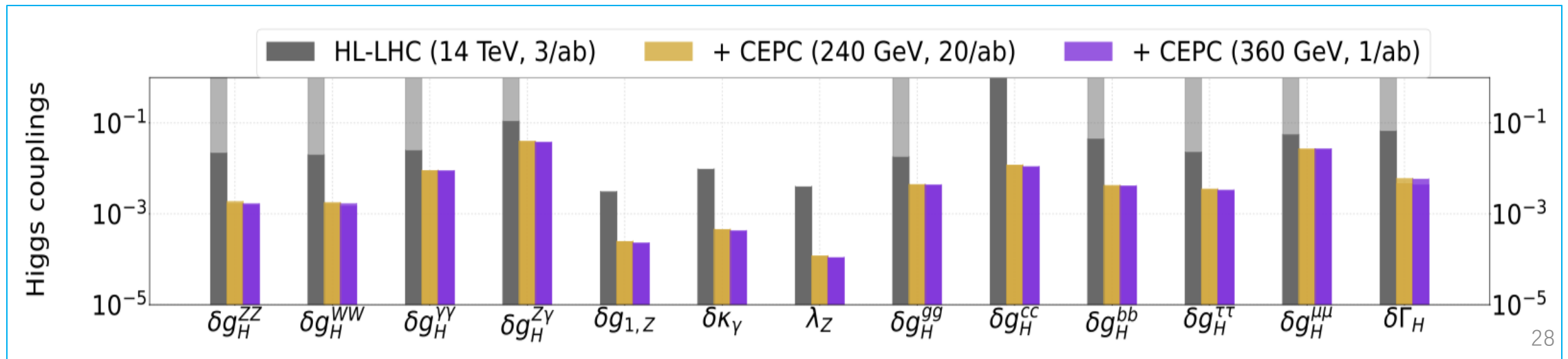
Energy reach in representative exotic search channels at the CEPC. Note the maximal energy reach may apply to different model parameter regions between experiments.

9. Global fits

Global fits: an essential tool to obtaining a thorough understanding of a NP model, and the implications and predictions of the models for future searches and experiments.

- **SMEFT**
- **2HDM**
- **SUSY global fits**

- **Global fit for SMEFT operators at future colliders**
- CEPC can improve the Higgs couplings by a factor of a few, or even orders of magnitude ($\delta g_{1,Z}$, $\delta \kappa_\gamma$, and λ_Z .)



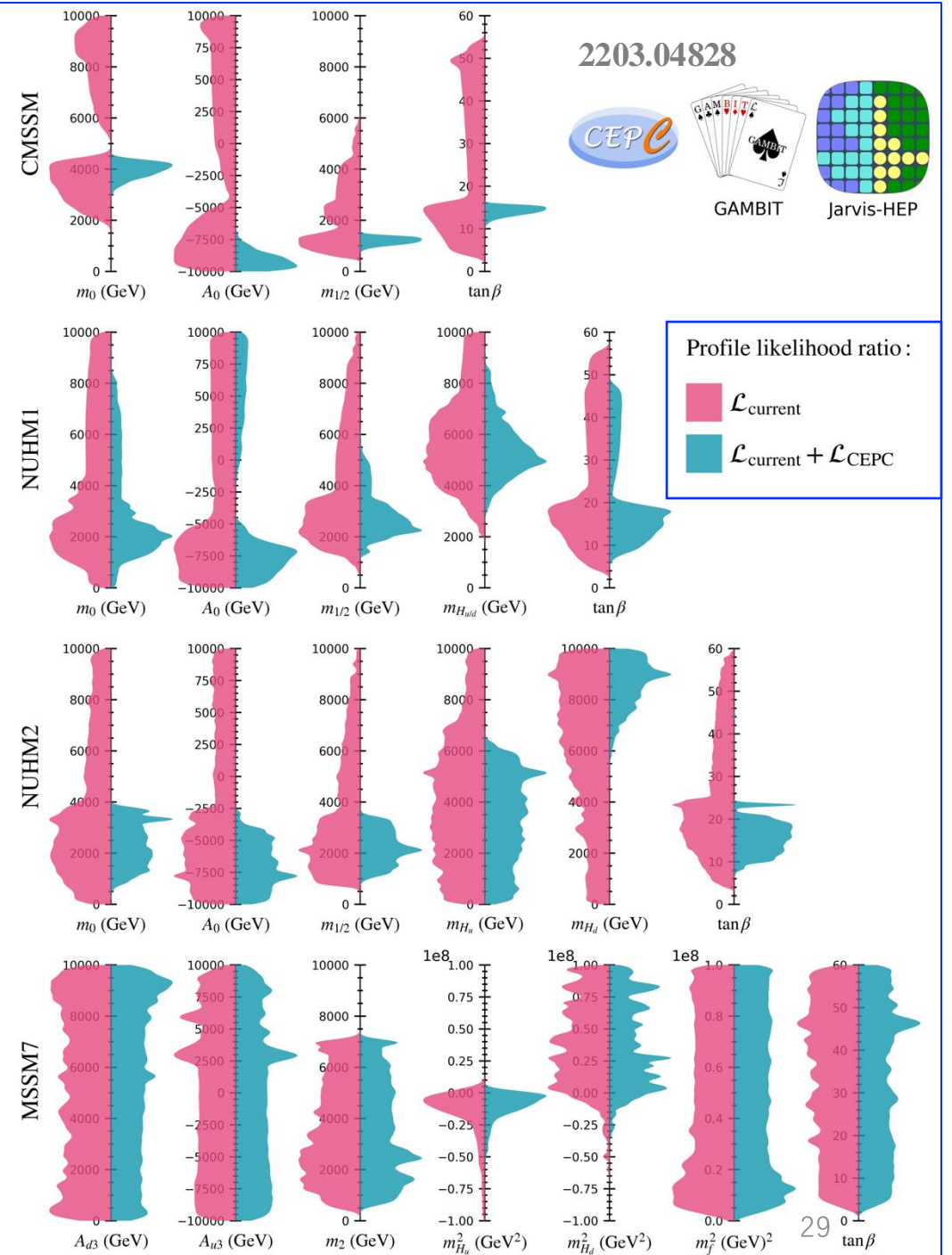
9. Global fits

Global fits: an essential tool to obtaining a thorough understanding of a NP model, and the implications and predictions of the models for future searches and experiments.

- SMEFT
- 2HDM
- **SUSY global fits**

CEPC has the potential to greatly enhance our understanding of the parameter space and mass spectrum in the MSSM.

One-dimensional profiled likelihood ratio for the global fit



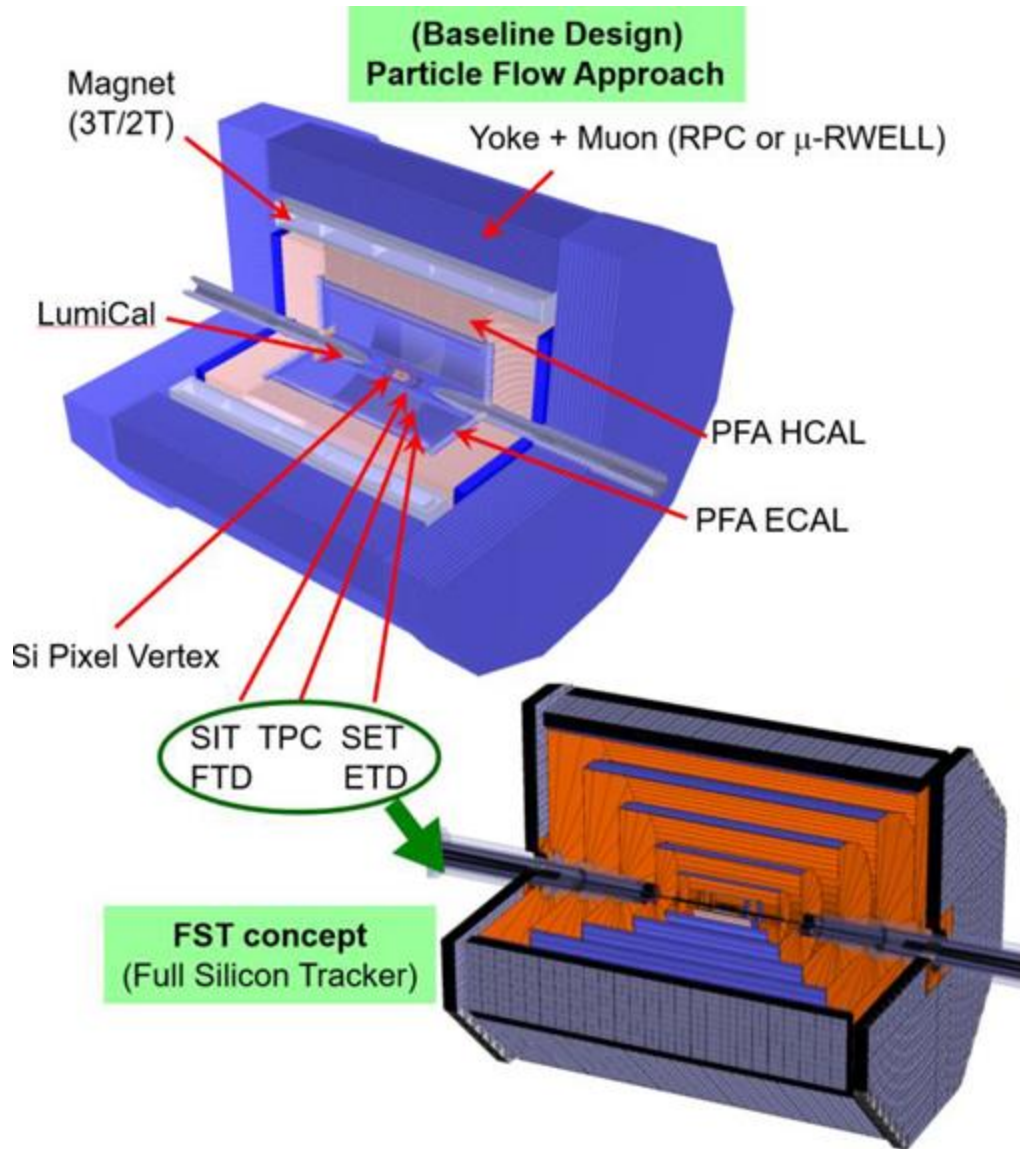
Summary and Outlook

- **CEPC has good discovery potential for NP, which is complementary to LHC and has big advantage at low energy/mass scale**
- **CEPC BSM white paper is preparing and to be ready for review by this year**
- **Please let us know if you would like to help to polish and review the BSM white paper !**

Thanks for your attention!

Backup

CEPC



- About CEPC

ECM=240GeV, higgs factory, 100 km circumference, 2 interaction points.

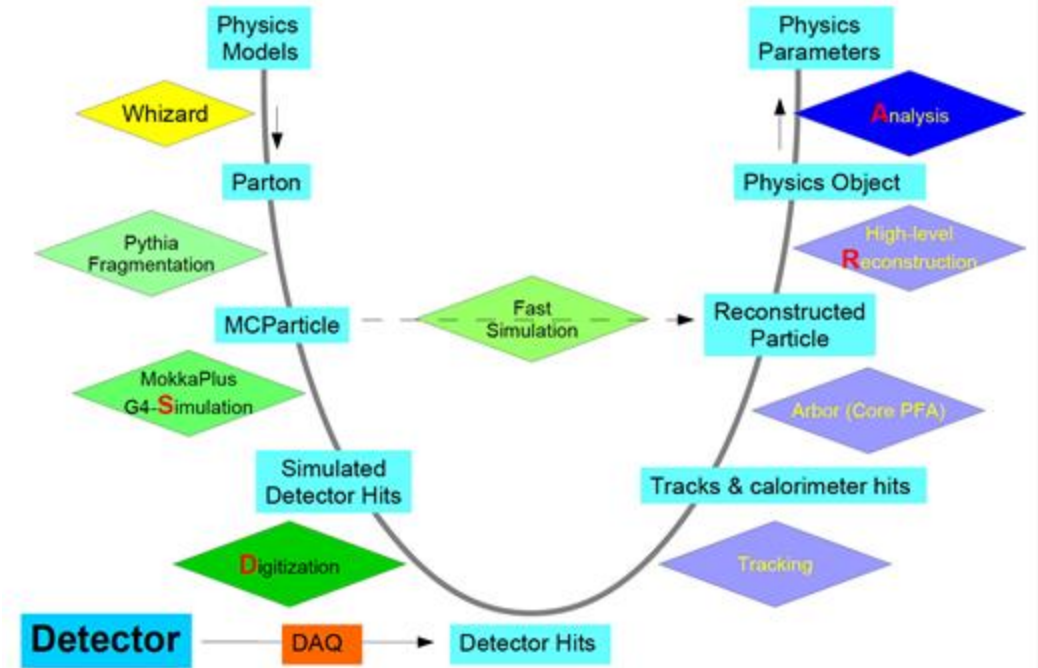
ILD-like detector

- Software

Signal samples: MadGraph+Pythia8

Simulation: Mokka

Reconstruction: Marlin



Full simulation reconstruction Chain with Arbor, iterating/validation with hardware studies