

Plasmon-enhanced Direct Detection of sub-MeV Dark Matter



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Main Structure in one slide

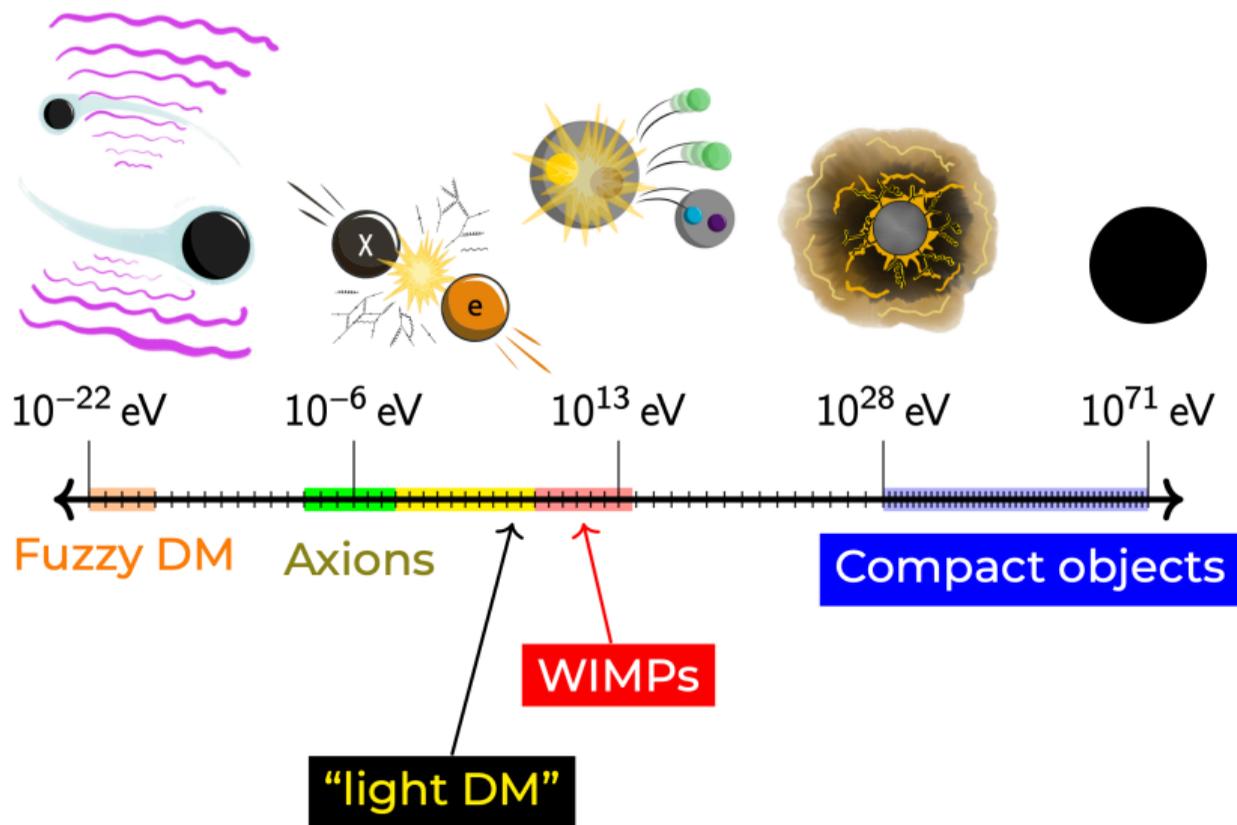
Objective:

Introduces an improved sensitivity approach to detect sub-MeV dark matter as detailed in the paper 2401.11971 with Zheng-Liang Liang, Liang-Liang Su and Lei Wu.

Overview:

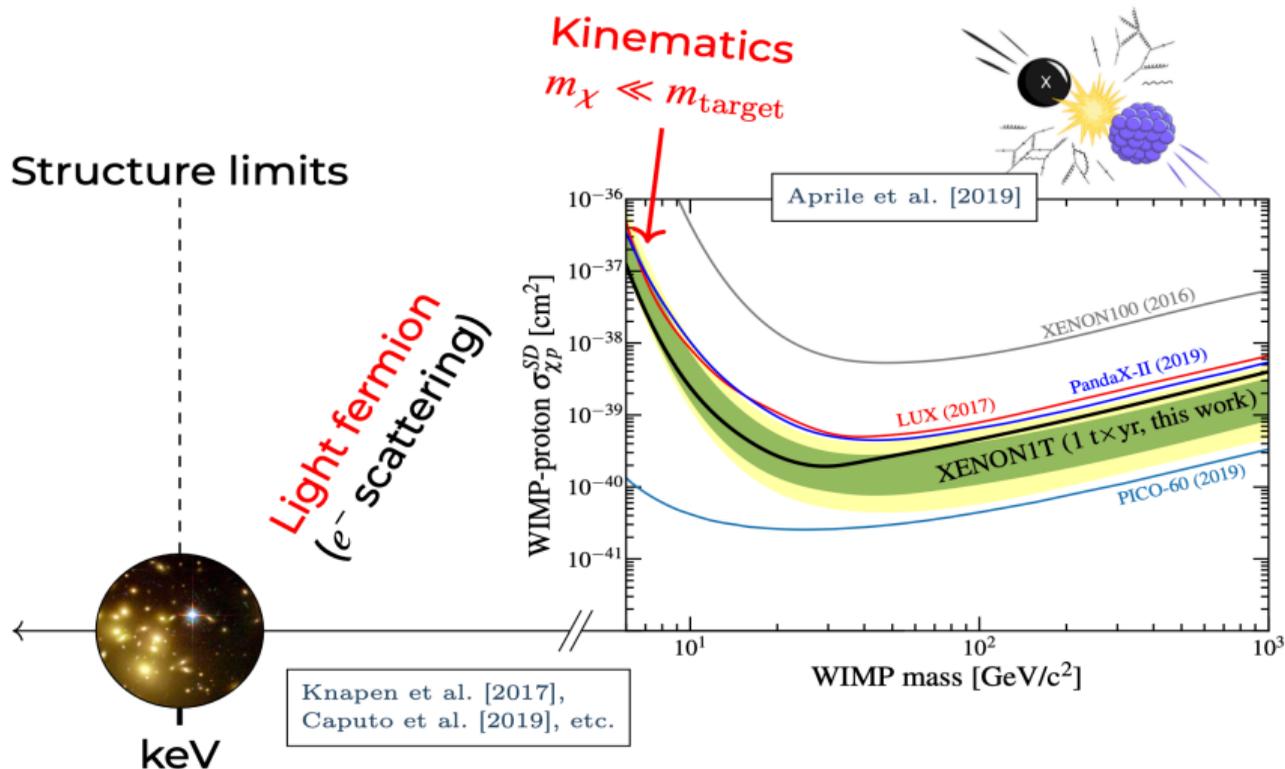
- ▶ Show why and what is sub-MeV dark matter and plasmon
- ▶ Explain why we need relativistic dark matter to excite plasmon
- ▶ Present the computational framework
- ▶ Demonstrate improved sensitivity in SENSEI experiment

Dark Matter Landscape



Why and What is Light Dark Matter?

probe keV needs significant detection analysis



A Broad Perspective

More than just nuclear recoil!

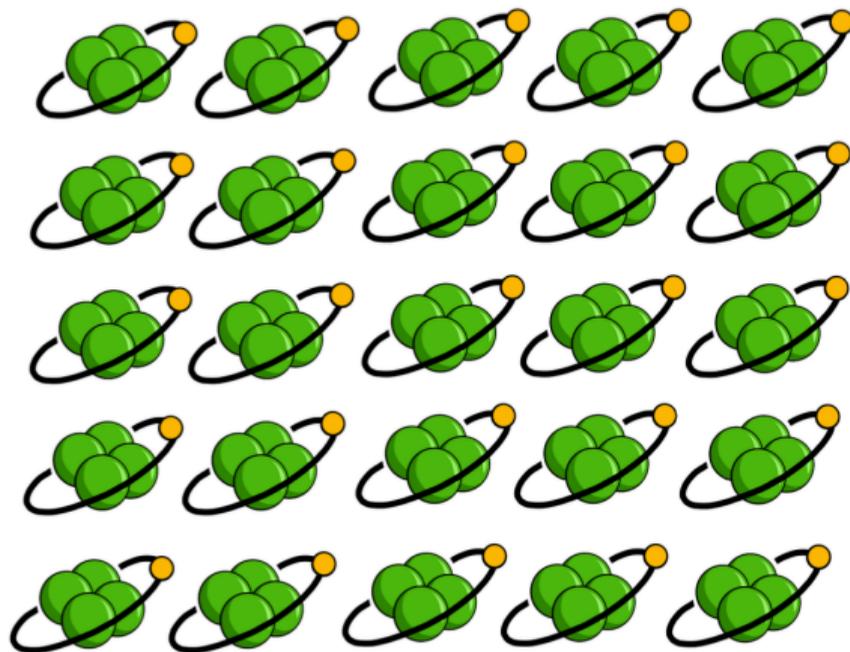
$$R \sim \int d^3\mathbf{v} f(\mathbf{v}) \int d^3\mathbf{q} F^2(\mathbf{q}) S(\mathbf{q}, \omega_{\mathbf{q}})$$

Material properties (e.g. dielectric function) for something must respond at the appropriate (q, ω) :

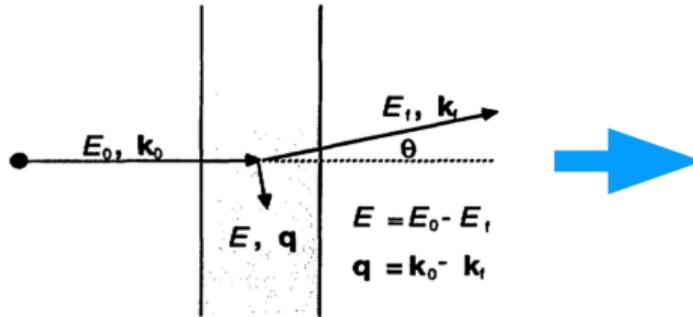
- ▶ electronic bands
- ▶ Migdal electron
- ▶ free nucleus
- ▶ phonons
- ▶ many more collective effects → Plasmons!

What is Plasmon?

A collective oscillation of electrons, like phonons being collective mode of nucleus

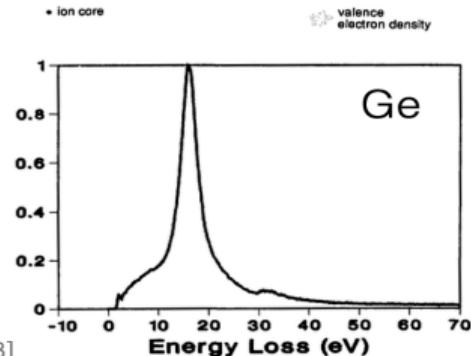
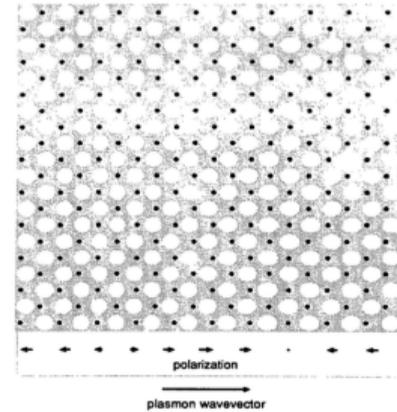


EELS and Plasmons



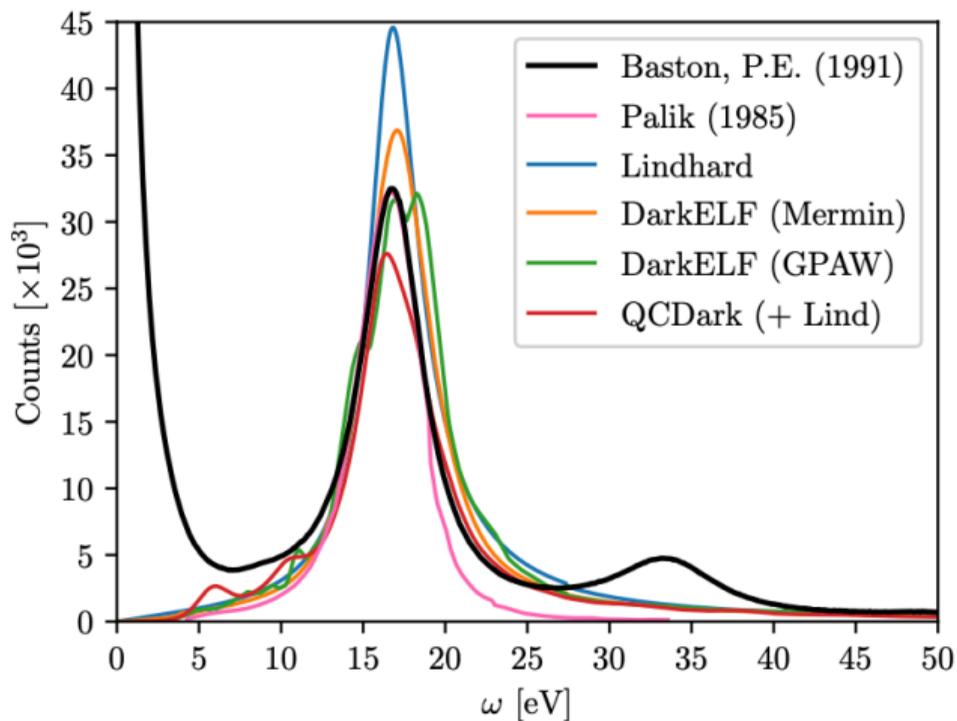
Semi-relativistic electron scattering **not** described by single-particle electron-electron scattering, but by a collective long-range charge wave (plasmon). Electron preferentially deposits ~ 15 eV of energy, **regardless of initial kinetic energy**

[M. Kundmann, Ph.D. thesis 1988]



Why Plasmon?

Shows up as a resonance in the loss function



Computational Framework

DM scattering in dielectrics

$$\Gamma = \int \frac{d^3\mathbf{q}}{(2\pi)^3} |V(q)|^2 \left[2 \frac{q^2}{e^2} \text{Im} \left(-\frac{1}{\epsilon(\mathbf{q}, \omega_{\mathbf{q}})} \right) \right]$$

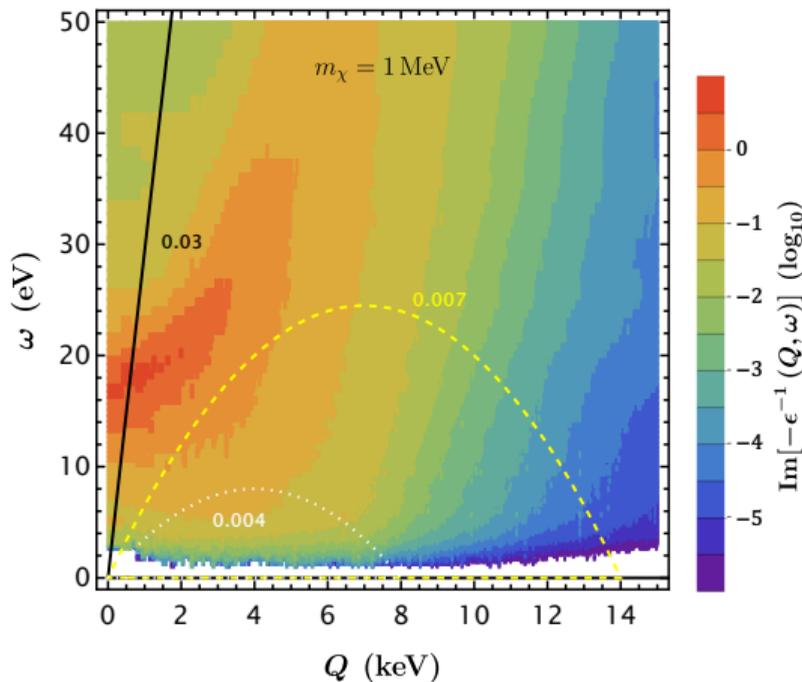
- ▶ Scattering potential, flexible for different dark models
- ▶ Dielectric function, directly measurable and predictable

Different from conventional electron ionization factor

ϵ contains all collective modes

Zeroth-order Consideration

Why halo dark matter fails exciting plasmon?



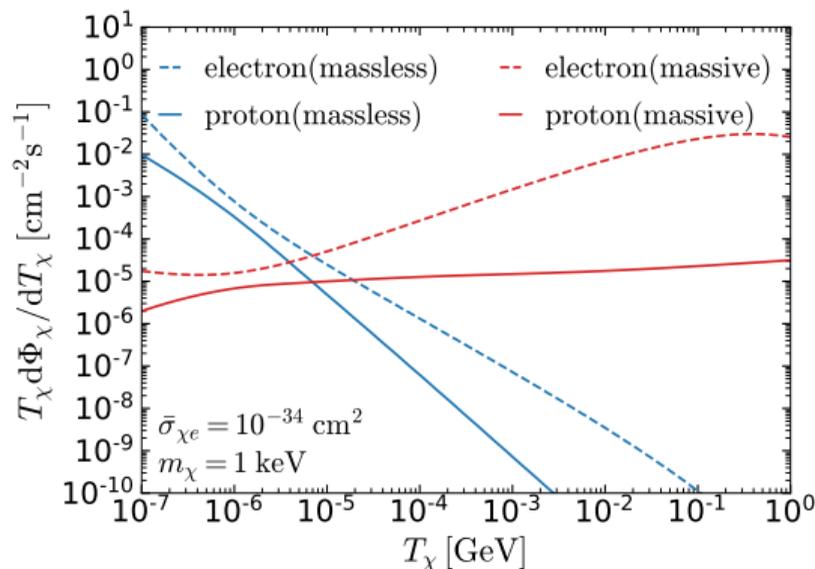
- ▶ $\text{Im}[-\epsilon^{-1}(\mathbf{Q}, \omega)]$ for silicon
- ▶ Resonance structure (plasmon excitation) exists ($|\mathbf{Q}| < 5\text{keV}, \omega \sim 15\text{eV}$)
- ▶ To excite plasmon, need small q for fixed ω from $\omega = \mathbf{q} \cdot \mathbf{v} - \frac{q^2}{2m_\chi}$

$$v_{\min} > q/\omega \sim 10^2$$

Natural Relativistic Source: CRDM

Since we assume dark matter scatters with electron, it must scatter with cosmic electro too!

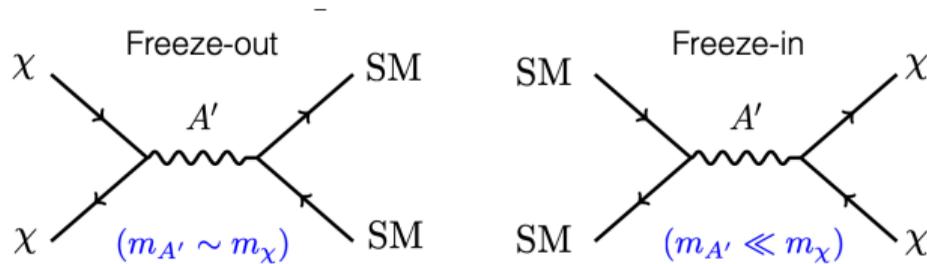
$$\frac{d\Phi_\chi}{dT_\chi} = D_{\text{eff}} \frac{\rho_\chi^{\text{local}}}{m_\chi} \sum_i \int_{T_i^{\text{min}}}^{\infty} dT_i \frac{d\sigma_{\chi i}}{dT_\chi} \frac{d\Phi_i^{\text{LIS}}}{dT_i}$$



Benchmark Model

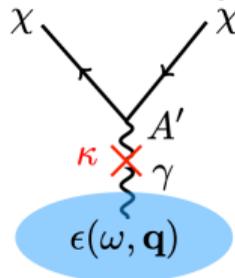
Dark Photon Mediator Model

$$\mathcal{L}_e^{\text{eff}} \supset g_e A'_0 \psi_e^* \psi_e + \frac{ig_e}{2m_e} \mathbf{A}' \cdot \left(\psi_e^* \vec{\nabla} \psi_e - \psi_e^* \overleftarrow{\nabla} \psi_e \right) + \dots,$$



Interactions in CM systems

DM interacts
with anything
electrically charged



(easiest to see this
in interaction basis
rather than mass basis;
valid if $m_{A'}^2 \ll \mathbf{q}^2$)

Our Computational Framework

$$\Gamma(\mathbf{p}_\chi) = \int \frac{d^3\mathbf{Q}}{(2\pi)^3} |V(\mathbf{Q}, \omega)|^2 \left[2 \frac{Q^2}{e^2} \text{Im} \left(-\frac{1}{\epsilon(\mathbf{Q}, \omega)} \right) \right]$$

- ▶ Similar Fermi's Golden Rule, but different kinematics

$$Q = |\mathbf{Q}| = |\mathbf{p}_\chi - \mathbf{p}'_\chi|, \quad \omega = E_\chi - E'_\chi = \sqrt{p_\chi^2 + m_\chi^2} - \sqrt{|\mathbf{p}_\chi - \mathbf{Q}|^2 + m_\chi^2}$$

- ▶ Scattering potential

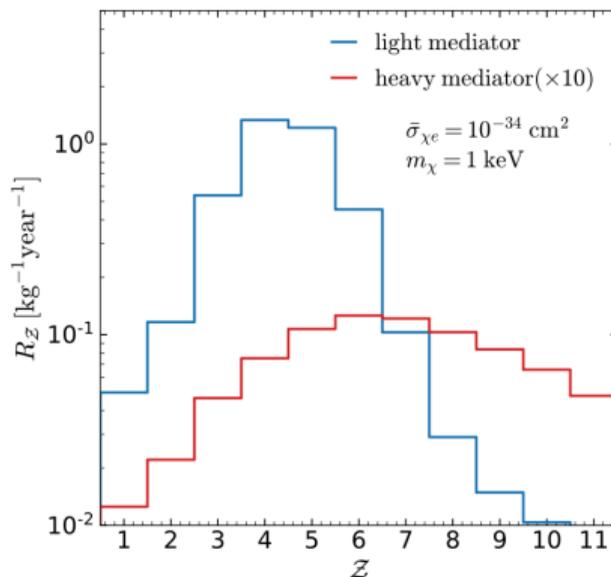
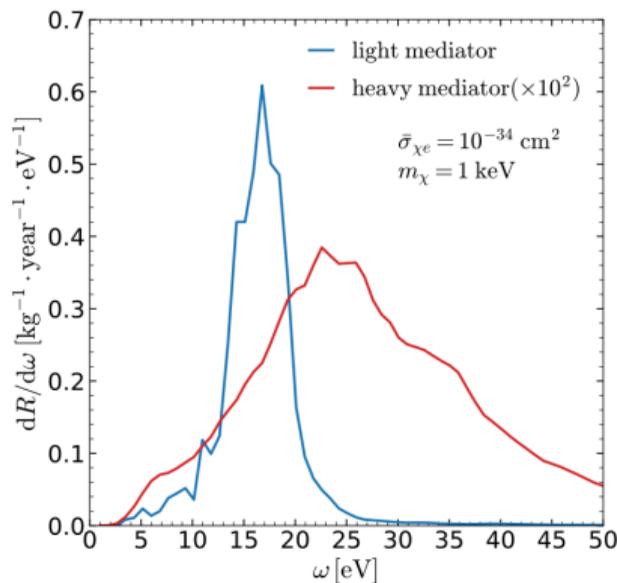
$$|V(\mathbf{Q}, \omega)|^2 = \frac{\pi \bar{\sigma}_{\chi e} \left[(2E_\chi - \omega)^2 - Q^2 \right]}{4\mu_{\chi e}^2 E_\chi (E_\chi - \omega)} |F_{\text{DM}}(q)|^2,$$

- ▶ Dielectric function remains the same
- ▶ Event rate

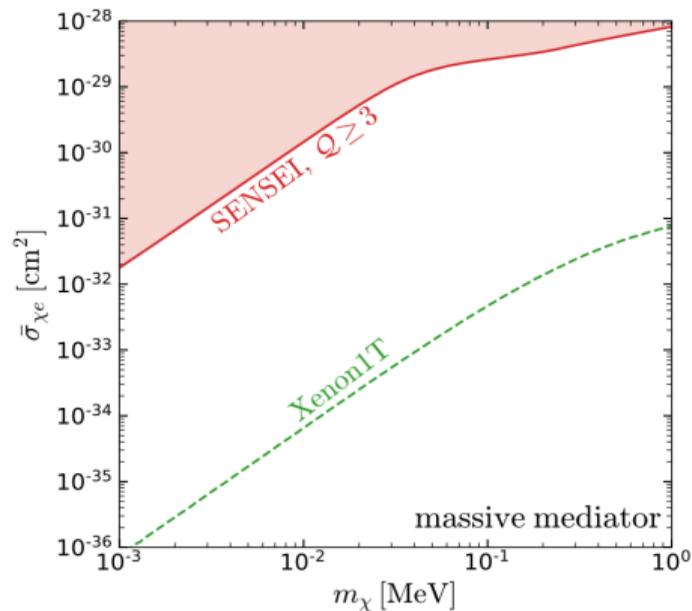
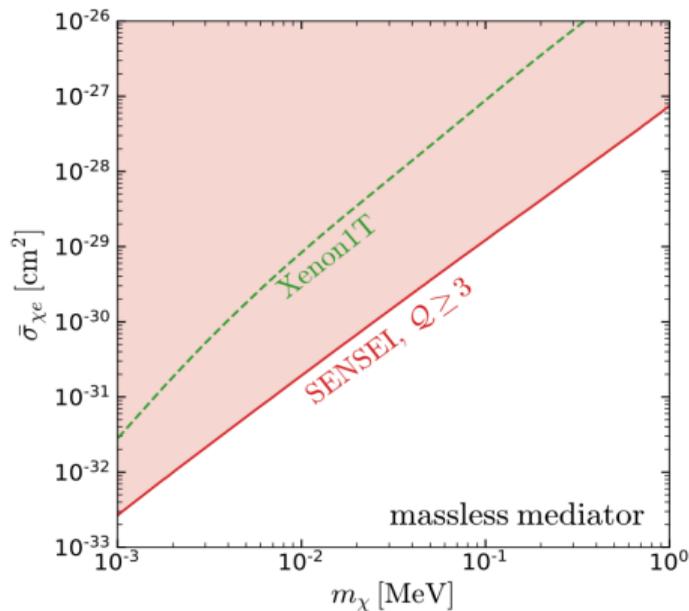
$$R = \frac{1}{\rho_T} \int dT_\chi \int \frac{d\Omega}{4\pi} \frac{d\Phi_\chi}{dT_\chi} \left(\frac{E_\chi}{p_\chi} \right) \Gamma(p_\chi)$$

Numerical Results

$$|F_{\text{DM}}(q)|^2 = \frac{(\alpha^2 m_e^2 + m_{A'}^2)^2}{(q^2 + m_{A'}^2)^2} = \begin{cases} 1 & \text{heavy mediator} \\ \frac{(\alpha m_e)^4}{q^4} & \text{light mediator} \end{cases}$$



Plasmon + DM with high velocity + massless mediator



Summary and Outlook

- ▶ Plasmon provides resonance enhancement to the event rate for relativistic dark matter
- ▶ SENSEI is now observing similar behavior like plasmon, which can be the signal of dark matter
- ▶ For now, only focus on the electron density operator, how to generalize the current-current operator?

Backup Slides

Rate predictions from QEDark and QCDark

electron recoil
spectrum

$$\frac{dR_{\text{crystal}}}{d \ln E_e} = \frac{\rho_\chi}{m_\chi} N_{\text{cell}} \bar{\sigma}_e \alpha \times \frac{m_e^2}{\mu_{\chi e}^2} \int d \ln q \left(\frac{E_e}{q} \eta(v_{\min}(q, E_e)) \right) F_{\text{DM}}(q)^2 |f_{\text{crystal}}(q, E_e)|^2$$

crystal form
factor

$$|f_{\text{crystal}}(q, E_e)|^2 = \frac{2\pi^2 (\alpha m_e^2 V_{\text{cell}})^{-1}}{E_e} \sum_{i i'} \int_{\text{BZ}} \frac{V_{\text{cell}} d^3 k}{(2\pi)^3} \frac{V_{\text{cell}} d^3 k'}{(2\pi)^3} \times E_e \delta(E_e - E_{i' \vec{k}'} + E_{i \vec{k}}) \sum_{\vec{G}'} q \delta(q - |\vec{k}' - \vec{k} + \vec{G}'|) |f_{[i \vec{k}, i' \vec{k}', \vec{G}']}|^2$$

wavefunction
overlap

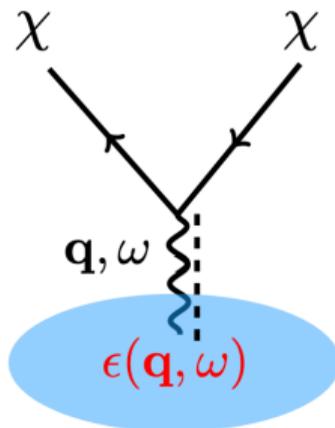
$$f_{[i \vec{k}, i' \vec{k}', \vec{G}']} = \sum_{\vec{G}} u_{i'}^*(\vec{k}' + \vec{G} + \vec{G}') u_i(\vec{k} + \vec{G})$$

This is pure theory: how well does it compare with data?
How do we calibrate charge yield from recoil spectrum?

Rate predictions from dielectric

$$\text{Im} \left(-\frac{1}{\epsilon(\mathbf{q}, \omega)} \right) = \frac{\pi e^2}{q^2} \sum_f |\langle f | \hat{\rho}(\mathbf{q}) | 0 \rangle|^2 \delta(\omega_f - \omega)$$

many-body
electron density



DM-electron interaction
(assumed spin-independent)

Dielectric function

$$\Gamma(\mathbf{v}) = \int \frac{d^3 \mathbf{q}}{(2\pi)^3} |V(\mathbf{q})|^2 \left[\frac{q^2}{e^2} 2 \text{Im} \left(-\frac{1}{\epsilon(\mathbf{q}, \omega)} \right) \right]$$

That's the answer, for any material.