

Searching for Fractionally Charged Particles with DAMPE

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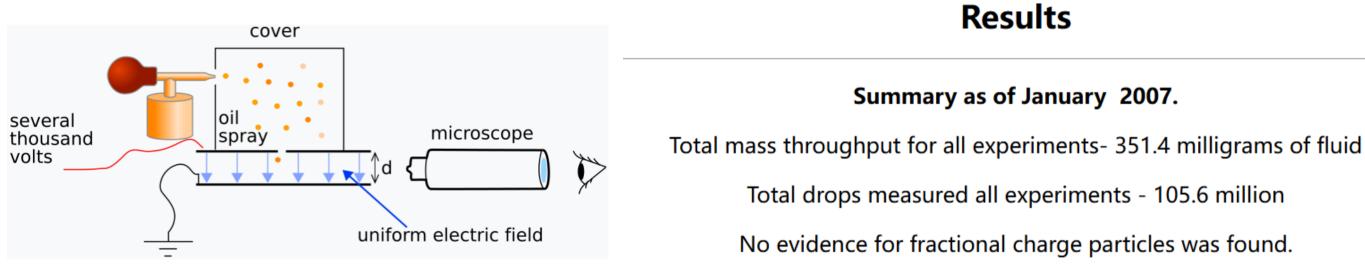


Outline

- Motivation
- Previous results of FCP
- DAMPE experiment
- Search for FCP with DAMPE
- Summary

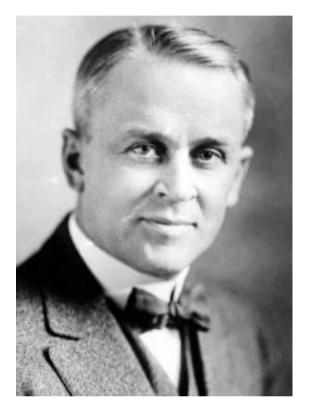
Motivation

•In 1909, Millikan conducted the oil drop experiment to measure electric charge, ultimately determining the charge of a single electron.

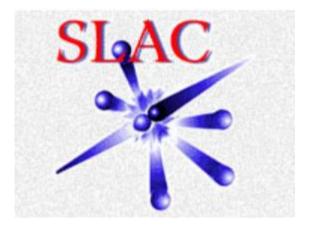


•The fractional charge search experiment at SLAC utilizes the Millikan method of determining the electric charge of falling fluid drops.



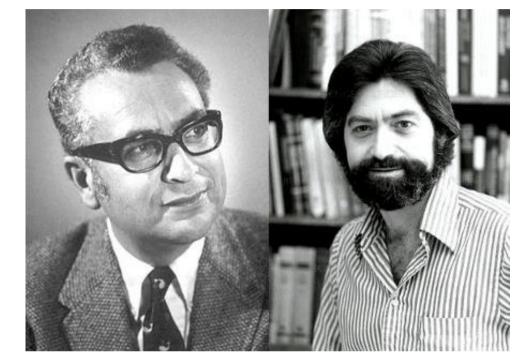


R. A. Millikan

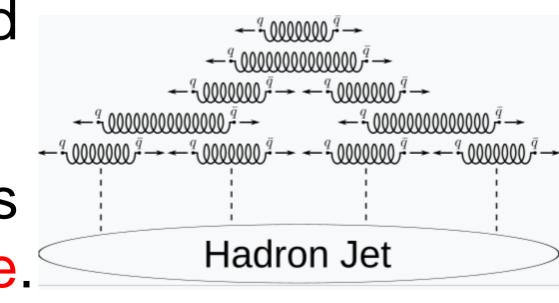


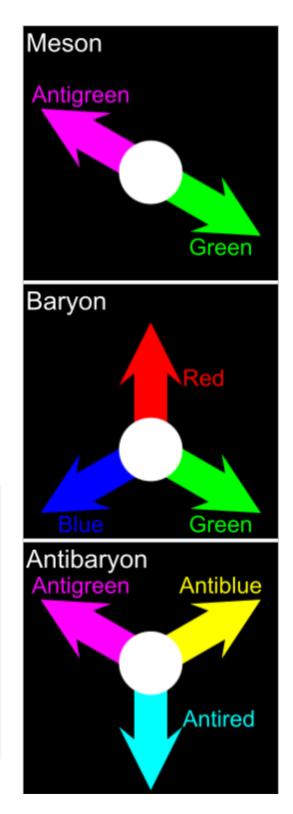
Motivation

- In 1964, quark model was proposed by Gell-mann and Zweig.
- •Hadrons composed of quarks and gluons.
- Due to the QCD theory, the quarks can not exist freely (asymptotic freedom and color confinement).
- Fractionally Charged Particle (FCP) is supposed to carry any non-integer charge.

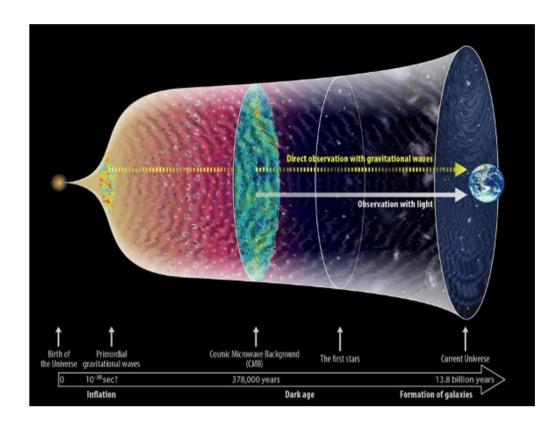


Gell-mann and Zweig





The possible origins of FCP





Early universe

Supernova explosion Extensive air shower

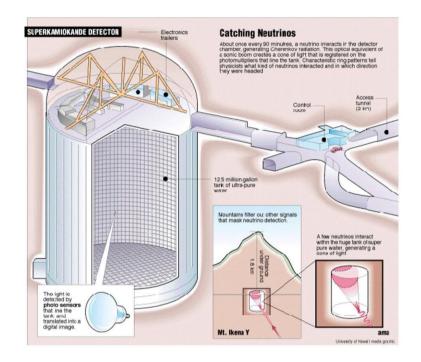
• There are three possible sources of FCP in cosmic rays:

- -It may be produced at the early Universe after the Big Bang
- -Or high-energy astrophysical processes
- -Or extensive air shower of cosmic-rays

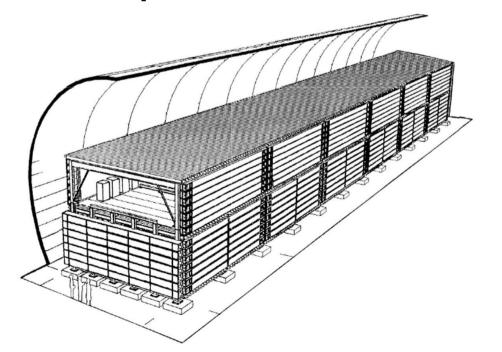


Searching for FCP with Underground Experiment

Kamiokande II depth: 1000 m



MACRO depth: 1400 m



LSD depth: 1800 m

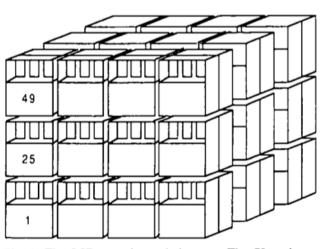
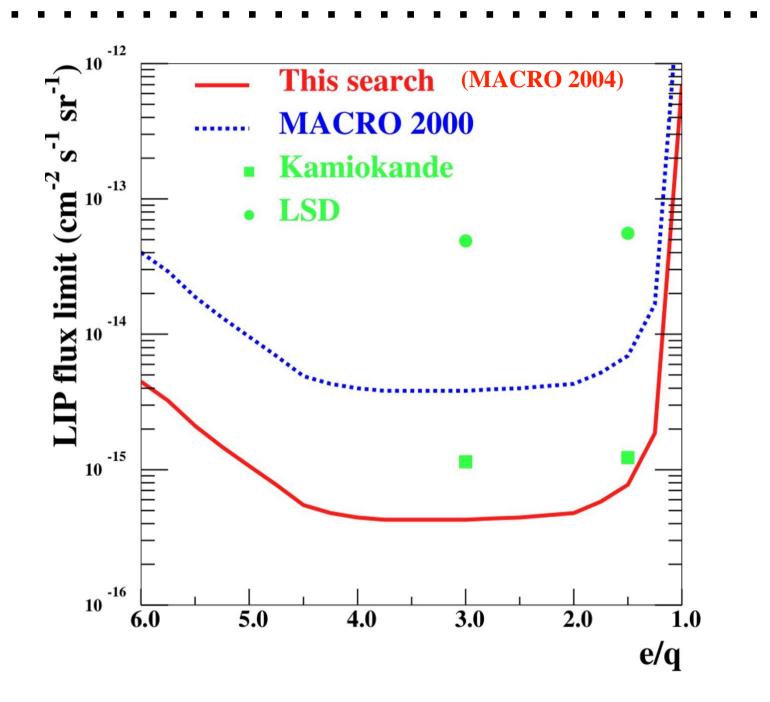


Fig. 1. The LSD experimental detector. The 72 tanks are considered as divided into 24 vertical columns (e.g. tanks 1-25-49 form the first telescope).



Current strictest upper limit given by MACRO (2004):

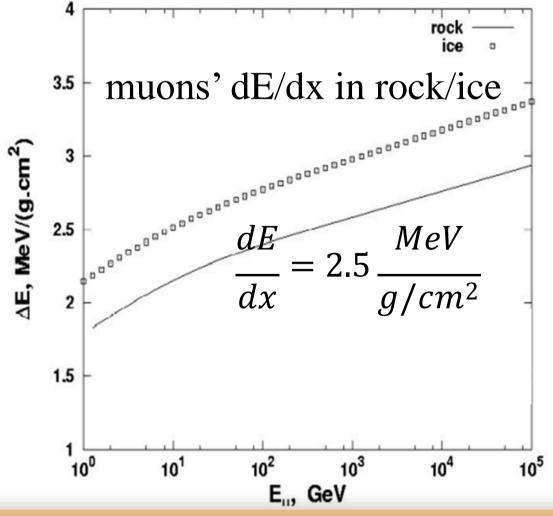
$$\Phi\left(\frac{1}{4} \sim \frac{2}{3}\right) = 6.1 \times 10^{-16} \text{ cm}^{-2} \text{ sr}^{-1} \text{ s}^{-1}$$

Comparison between experiments

Underground Experiment

Energy loss when a particle passes through the 1000m depth rock

- for muon: ~ 663 GeV
- for 2/3e FCP: ~ 300 GeV



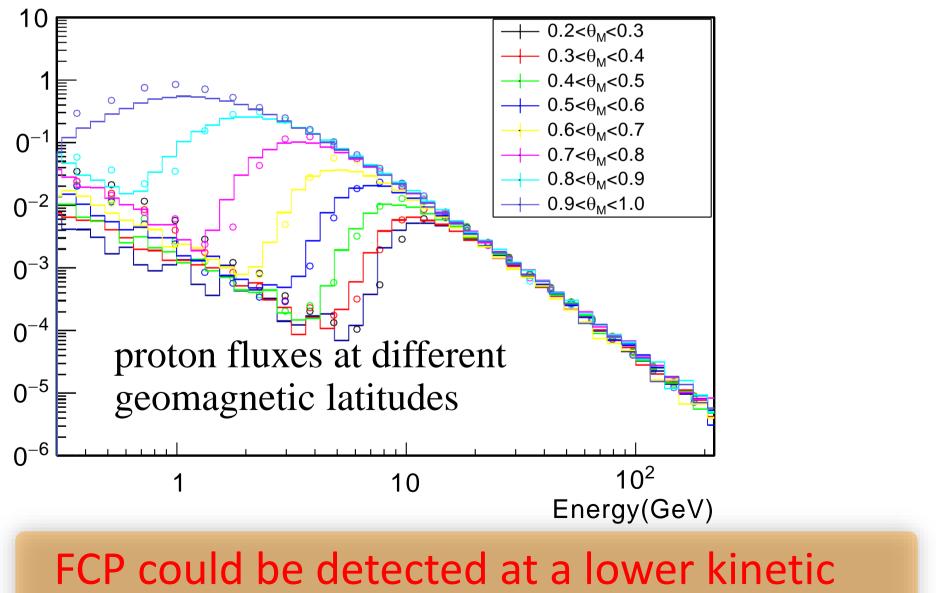
FCP should be with a high kinetic energy (> hundreds of GeV)



A cutoff structure is caused by the earth's magnetic field

Near the equator, proton flux cutoff ~ 10 GeV

2/3e FCP flux cutoff: 6 ~ 7 GeV

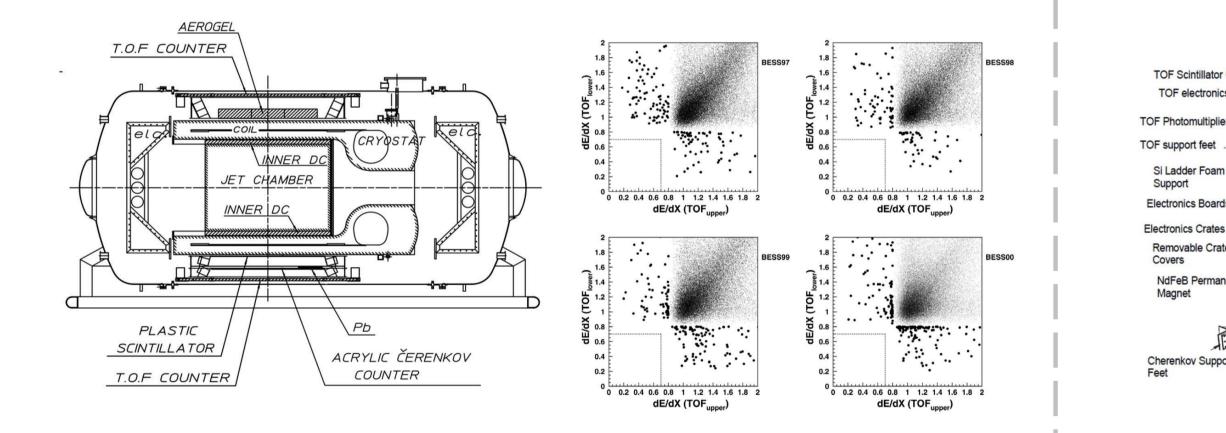


energy (tens of GeV)

Space Experiment

Searching for FCP with Space Experiment

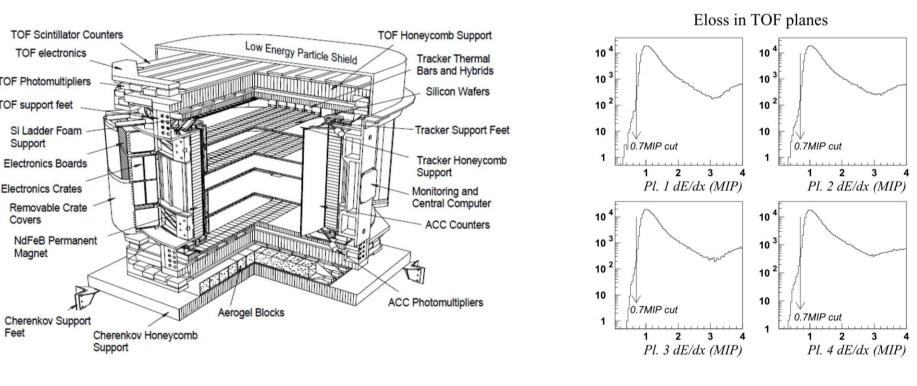
BESS



Upper limit (90% C.L.):

$$\Phi\left(\frac{2}{3}\right) = 4.5 \times 10^{-7} \text{ cm}^{-2} \text{sr}^{-1} \text{s}^{-1}$$

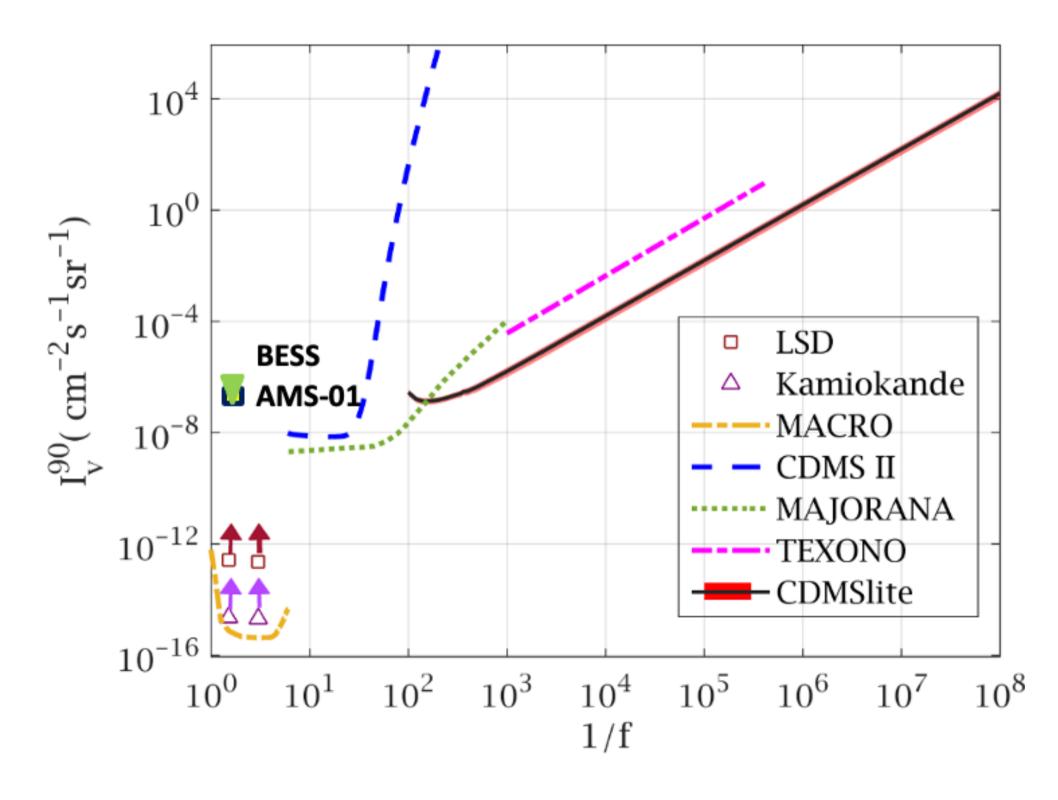
AMS01



Upper limit (95% C.L.):

$$\Phi\left(\frac{2}{3}\right) = 3.0 \times 10^{-7} \text{ cm}^{-2} \text{sr}^{-1} \text{s}^{-1}$$

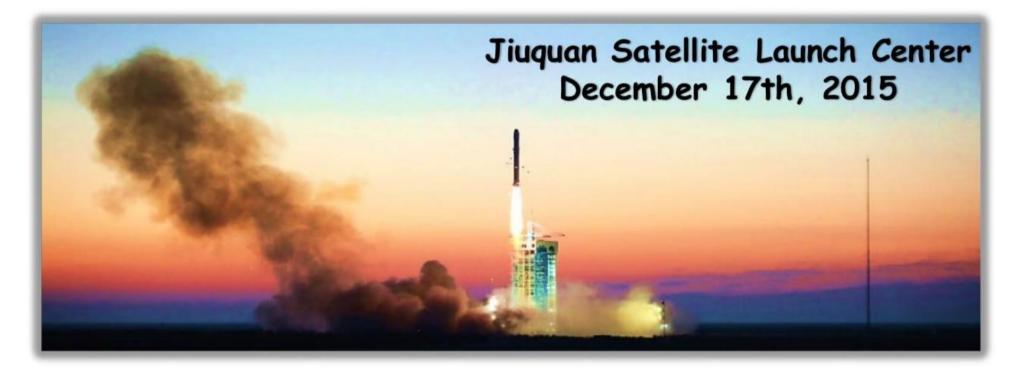
The results of previous experiments

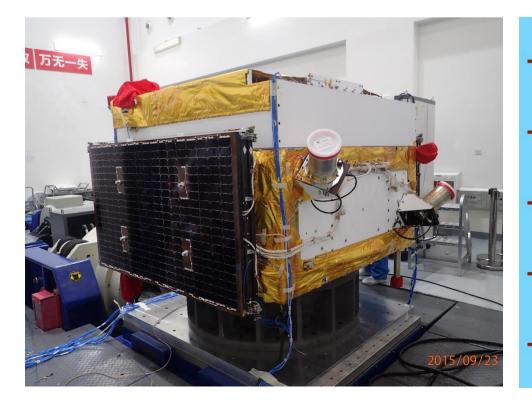


The flux upper limit versus the inverse charge value

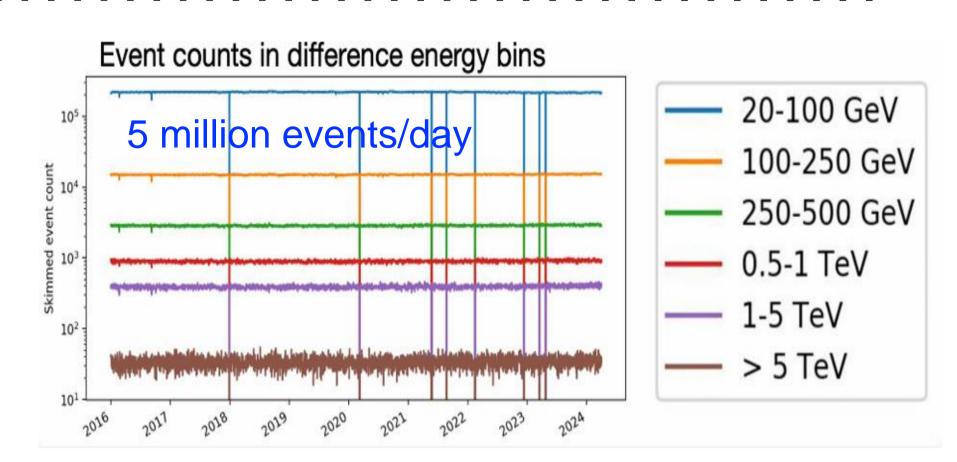
DArk Matter Explorer (DAMPE)

 DArk Matter Particle Explorer (DAMPE) is a space experiment for detecting high energy cosmic rays





Orbit: sun-synchronous Altitude: 500 km Period: 94 minutes 5 million events/day 16 GB/day downlink



CNINA

- -Purple Mountain Observatory, CAS
- University of Science and Technology of China
- -Institute of High Energy Physics, CAS
- -Institute of Modern Physics, CAS
- -National Space Science Center, CAS

ITALY

- INFN Perugia and University of Perugia
- INFN Bari and University of Bari
- -INFN Lecce and University of Salento
- -INFN LNGS and Gran Sasso Science Institute

SWITZERLAND

-University of Geneva

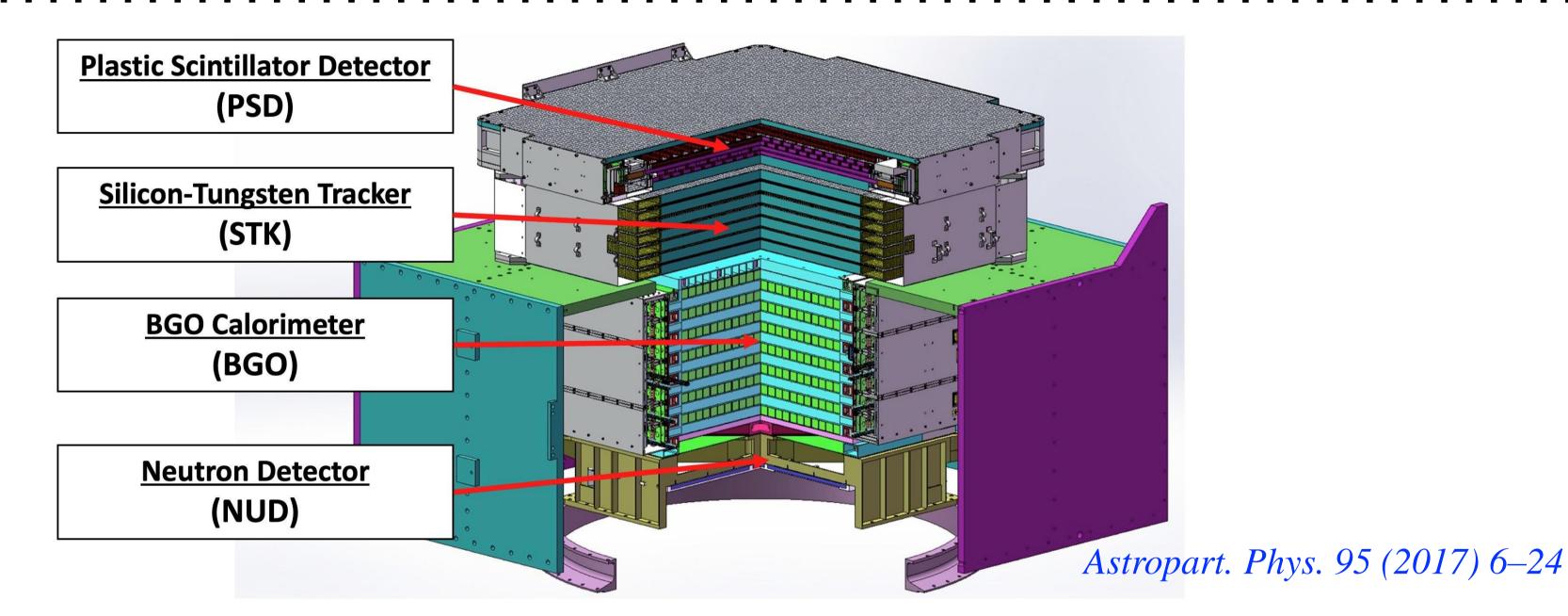






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DArk Matter Explorer (DAMPE)



- Charge measurement (dE/dx in PSD, STK)
- Gamma-ray converting and tracking (STK + BGO)
- Precise energy measurement (BGO)
- Hadron rejection (BGO + NUD)

dx in PSD, STK) I tracking (STK + BGO) ent (BGO) UD)

What can DAMPE do ?

- Advantages of DAMPE compared with other FCP experiments
 - -Observe FCPs with significantly lower energy (a few GeV)
 - -Relatively large acceptance
 - -Long exposure time

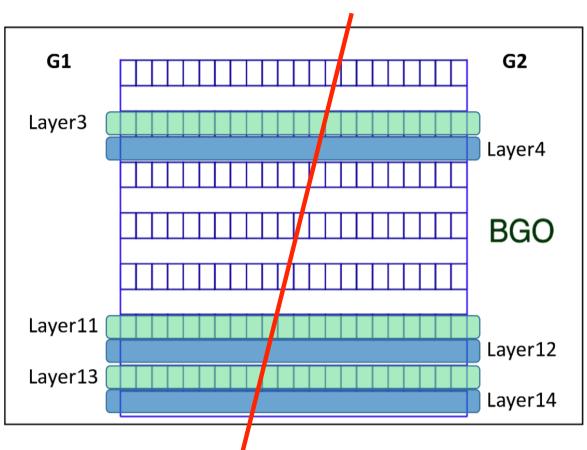
Experiment	Geometrical acceptance (cm ² sr)	Exposure time (s)	Upper limit (cm ⁻² sr ⁻¹ s ⁻¹)
AMS01	3000	3.6×10 ⁴	3.0×10⁻7 (95% C.L.)
BESS	1500	3.6×10 ⁵	4.5×10⁻7 (90% C.L.)
DAMPE	3000	2.3×10 ⁷	?

th other FCP experiments ower energy (a few GeV)

Assuming the nature of FCP

- Assuming the nature of FCP:
 - Massive
 - With electromagnetic interaction
 - W/O hadronic interaction
 - The charge would be 1/3e or 2/3e
 - -like a massive lepton (e.g. muon)

- Ionization energy loss for $1/3 \in FCP \approx 1/9 \text{ MIP}$, Not Pass Trigger
- Ionization energy loss for 2/3e FCP≈ 4/9 MIP,
 - Pass Trigger



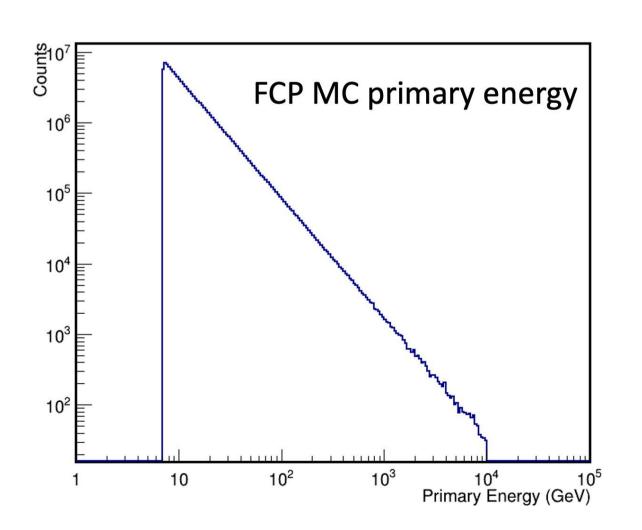
DAMPE MIPs trigger

Trigger threshold: 0.2 proton MIP

Our target

Data Sample

- Flight data: 2016.01.01 ~ 2020.12.30
- Simulation:
 - proton 10 GeV ~ 100 TeV
 - FCP 7 GeV ~ 10 TeV



- FCP simulation

 - Charge with 2/3 e

 - Mass 1200 MeV



 Created a virtual particle in Geant4 Add ionization and multi scattering process - Energy spectrum obey the E^{-2.7} - Spheric particle source

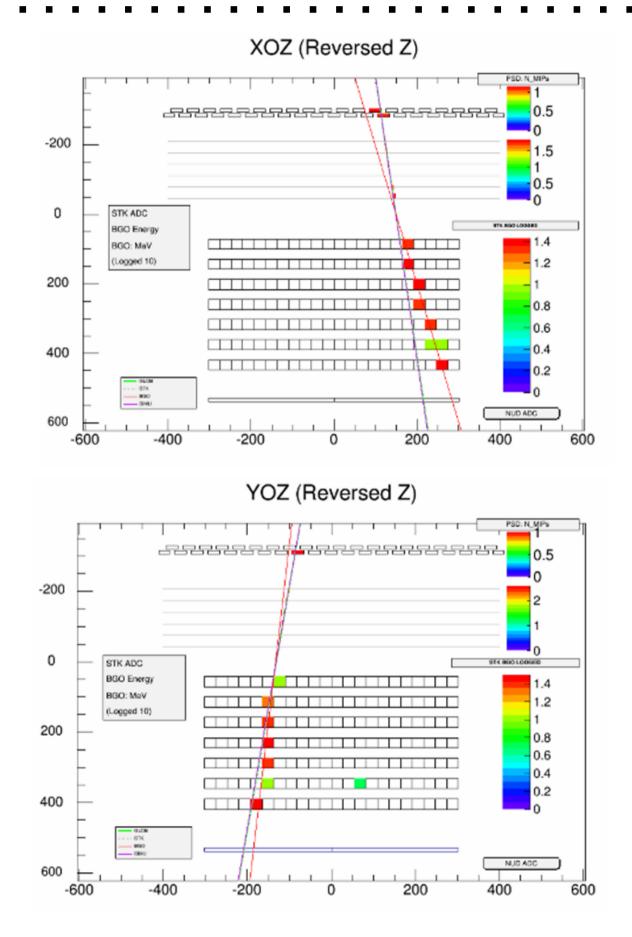
Sample Selections

Features of heavy lepton-like FCP

- Fractional charges
- Only ionization process

Identification heavy lepton-like FCP

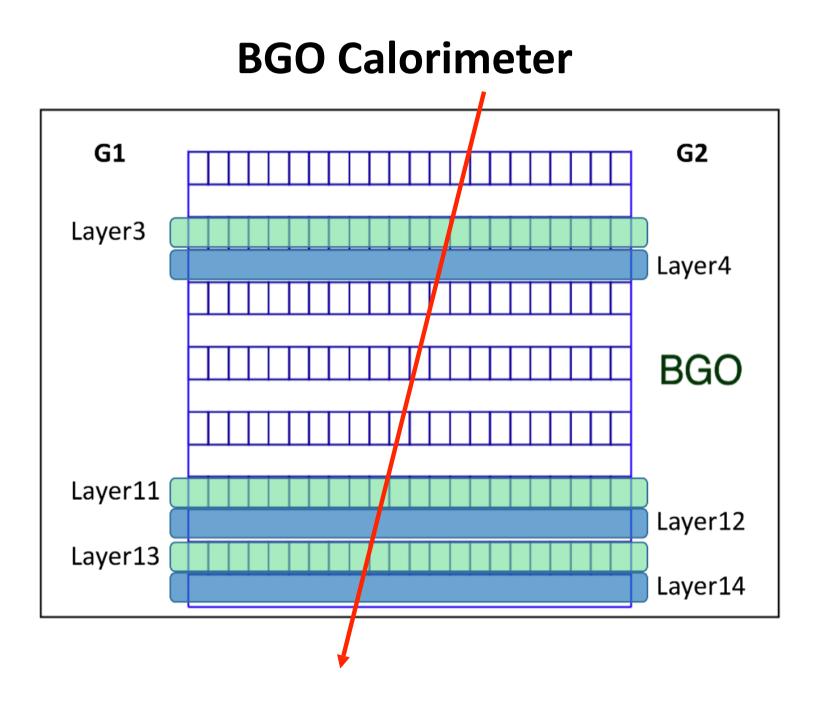
- Charge measurements
- MIP-like event selection



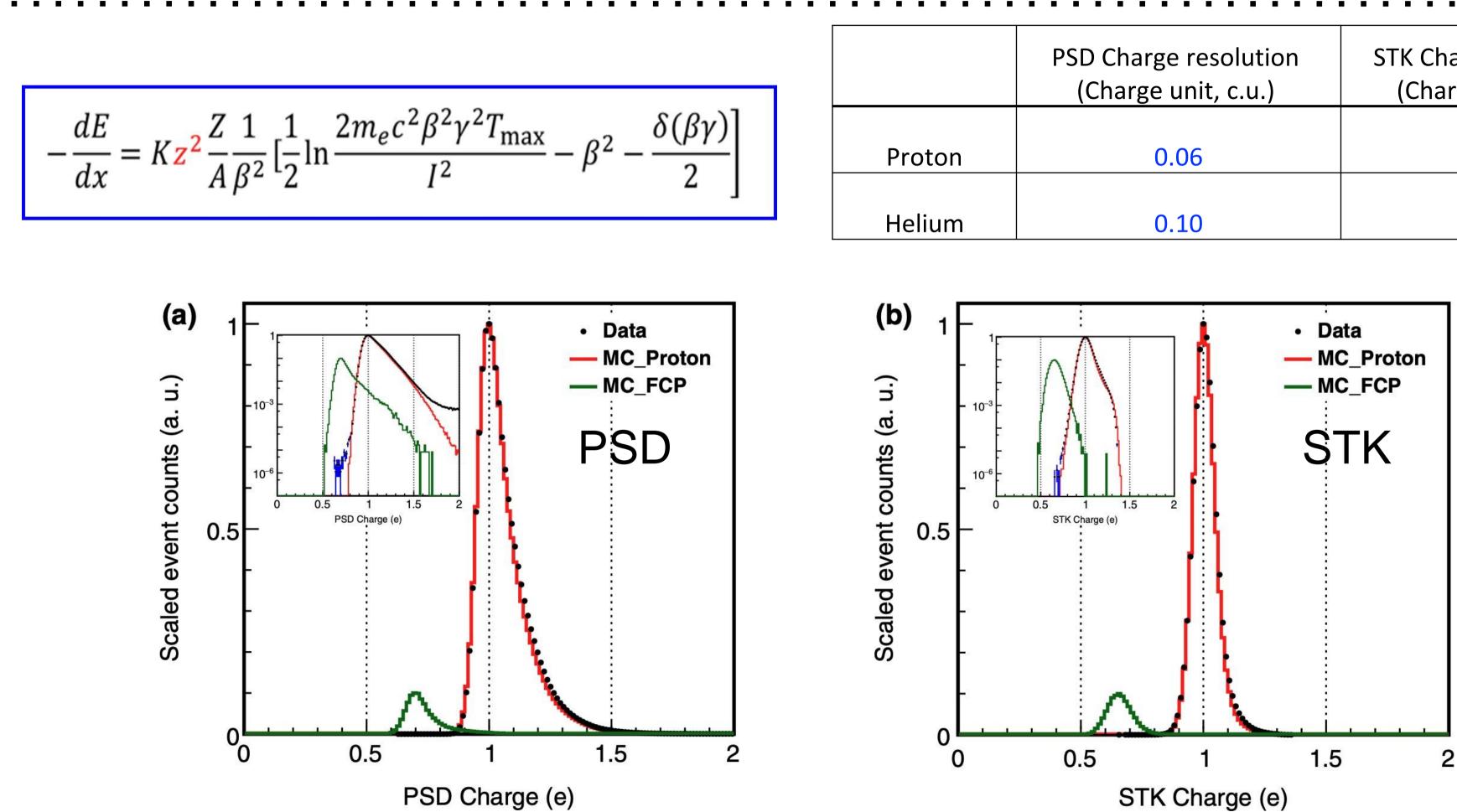
MIP-like Event Selection

- MIPs in BGO:
 - Over-threshold(2 MeV) hits no more than 2 in one layer along the track
 - More than 10 layers are required to have signals
 - One of last two layers should be fired





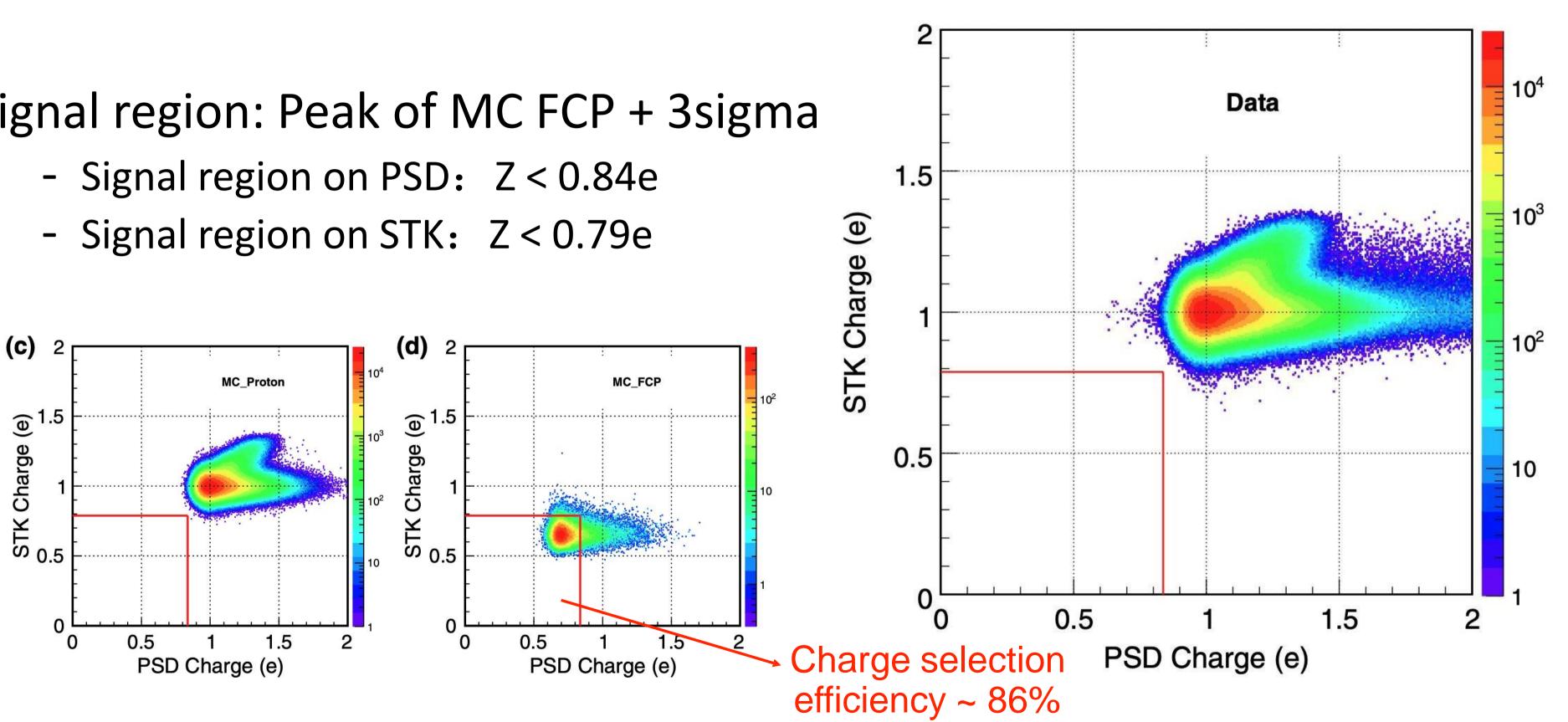
Charge Measurement



	PSD Charge resolution (Charge unit, c.u.)	STK Charge resolution (Charge unit, c.u.)
ton	0.06	0.04
um	0.10	0.07

Signal Region

Signal region: Peak of MC FCP + 3sigma



Upper Limit Calculation

$$\Phi = \frac{N_{\rm obs}}{T_{\rm exp}\epsilon_{\rm scale}\epsilon_{\rm trig}}$$

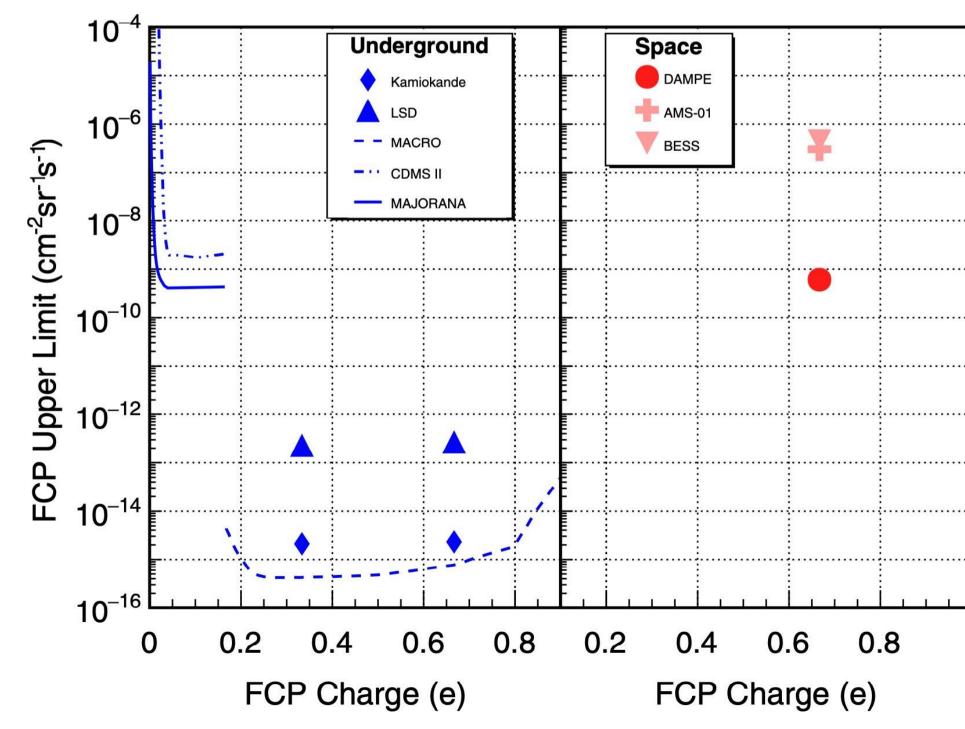
: Flux or flux upper limit $(cm^{-2}sr^{-1}s^{-1})$ Φ Nobs to be 2.44 at the 90% C.L. by Feldman-Cousins method T_{exp} : Exposure time 2.34×10^{7} s ϵ_{trig} : trigger efficiency, 85.5%, given by MC ϵ_{scale} : pre-scale factor 1/4 A_{eff} : effective acceptance 940 cm²sr ϵ_{reaion} : signal region efficiency 86%

- $A_{\rm eff}\epsilon_{\rm region}$
- : Number of observed FCP, no candidate is observed, N_{obs} is taken

Upper Limit of 2/3e FCP

TABLE I. The comparison between DAMPE and other similar types experiments.

Experiments	Geometric acceptance(cm ⁻² sr)	Exposure time (s)	Upper limit (cm ^{-2} sr ^{-1} s ^{-1})
AMS-01	3000	3.6×10^{4}	3.0×10^{-7} (95% CL)
BESS	1500	3.2×10^{5}	4.5×10^{-7} (90% CL)
DAMPE	3000	2.3×10^{7}	6.2×10^{-10} (90% CL)



$$\Phi < 6.2 \times 10^{-10} \text{ cm}^{-2} \text{ sr}^{-1} \text{ s}^{-1}$$

Improve three orders of magnitude

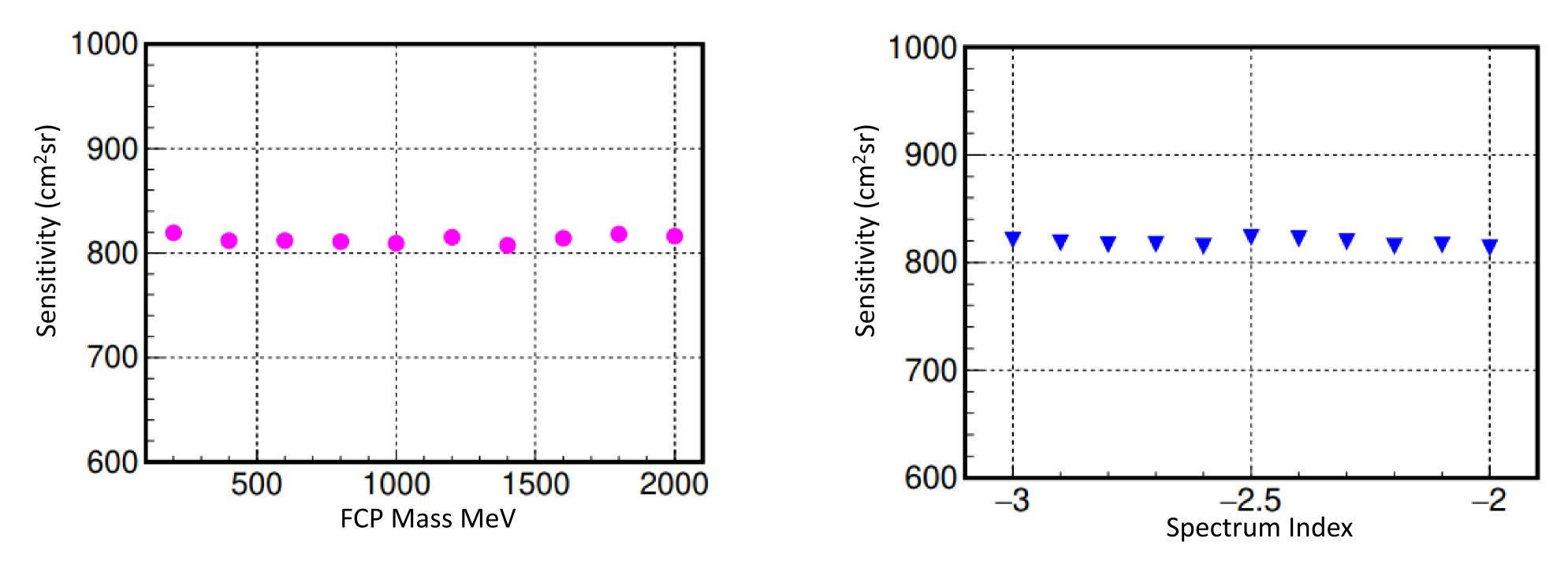
$$\delta = \sqrt{\delta_{\text{trigger}}^2 + \delta_{\text{track}}^2 + \delta_{\text{charge}}^2} = 3.1 \%$$

where $\delta_{\text{trigger}} = 1.1\%$, $\delta_{\text{track}} = 2.9\%$, and $\delta_{\text{charge}} = 0.5\%$

Phys. Rev. D 106, 063026 (2022)

Mass and Spectrum index scan

In the research: FCP mass is 1200 MeV and Spectrum index is 2.7



The mass varies from 200 MeV to 2000 MeV with a step of 200 MeV is applied to the simulation.

The spectrum index varies from -3 to -2 with a step of 0.1 is applied to the simulation. 21

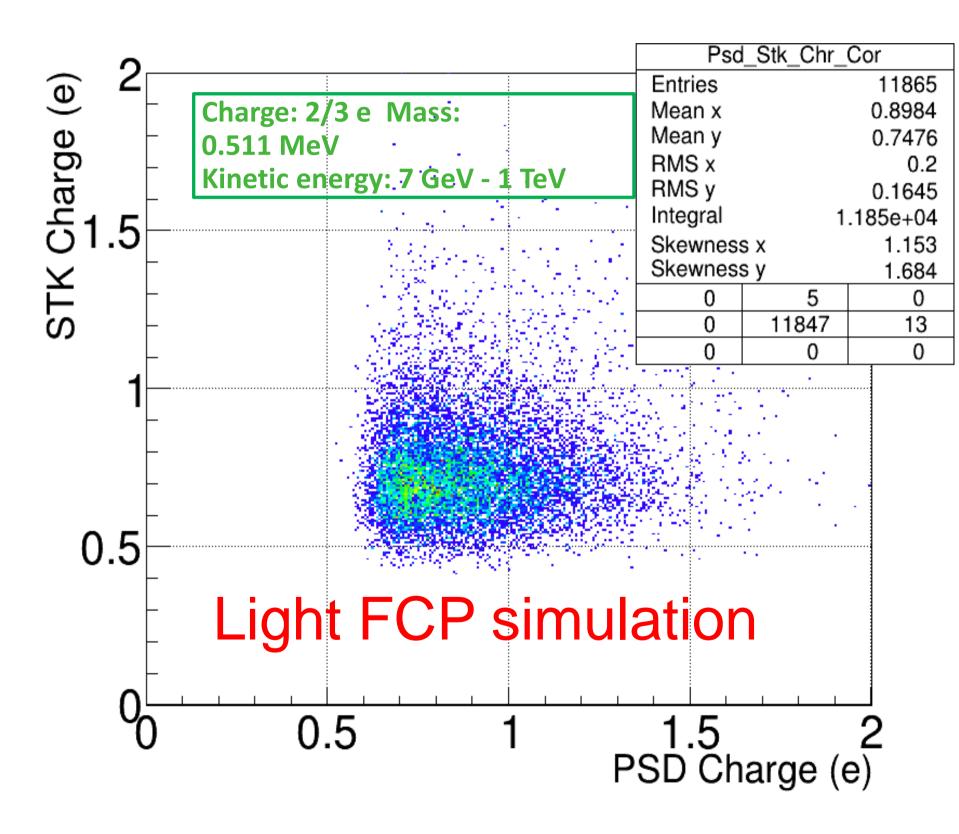
Prospects for Searching for Light FCP

FCP should not be constrained to be heavy lepton

- Shower can happen
- Mass may be light
- Electron-like light-mass particle

Energy loss: Bremsstrahlung

$$-\frac{dE}{dx} = 4\alpha N_A \frac{Z^2}{A} z^2 \left(\frac{1}{4\pi\epsilon_0} \frac{e^2}{mc^2}\right)^2 E \ln \frac{183}{Z^{1/3}}$$

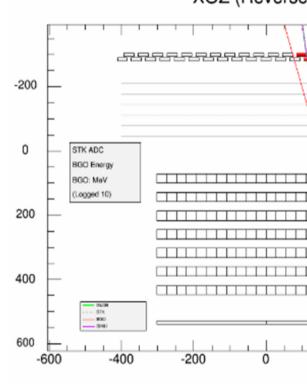


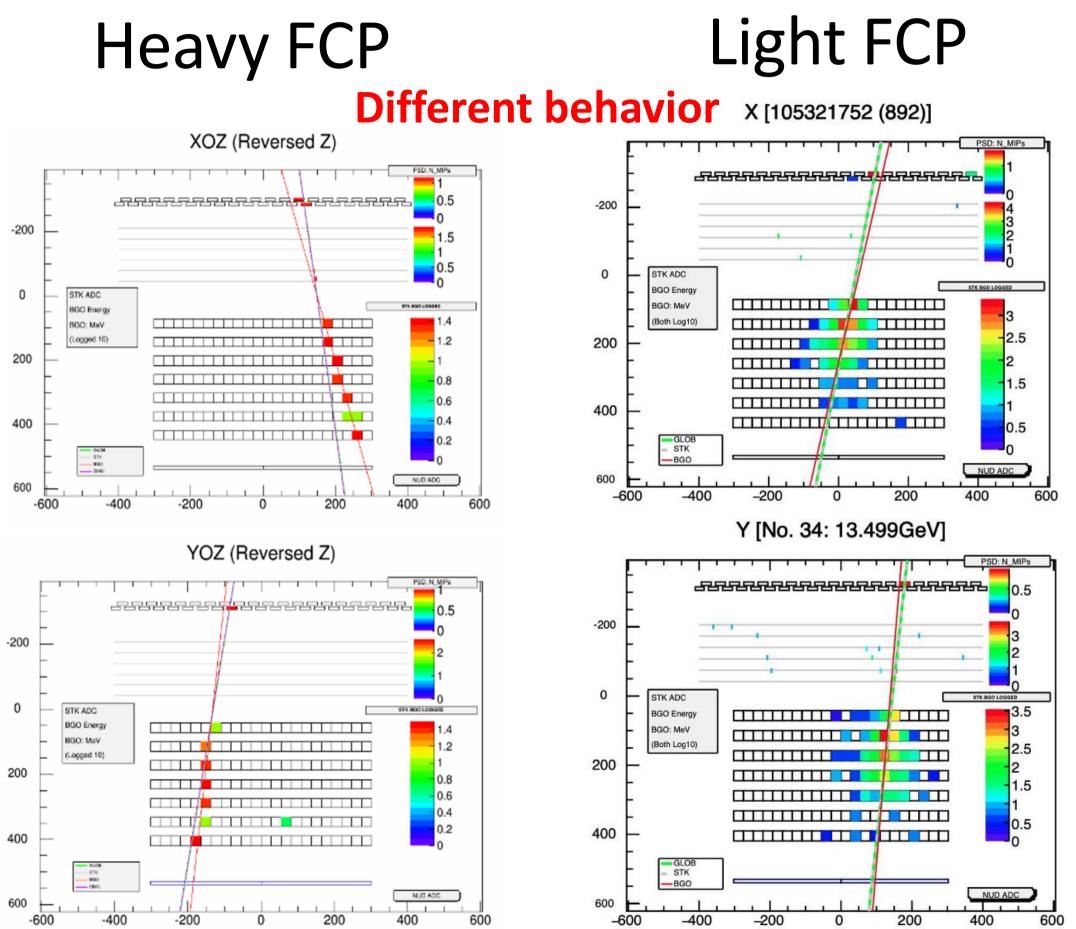
Prospects for Searching for Light FCP

Features of light lepton-like FCP

- Fractional charges
- **Ionization process**
- **Bremsstrahlung process**
- **Electromagnetic shower**

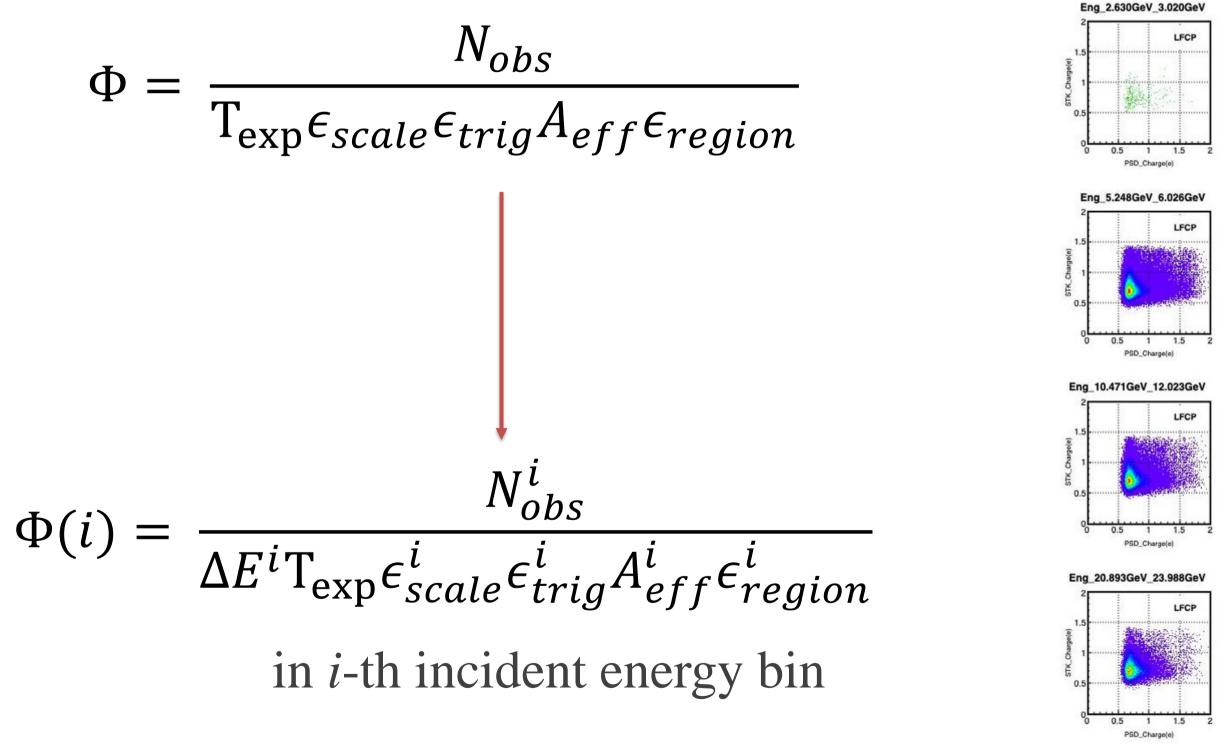
We can measure their energy by **BGO** calorimeter



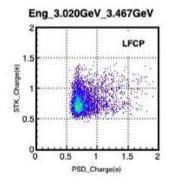


23

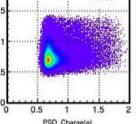
Prospects for Searching for Light FCP



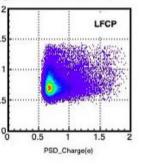
In this case, we can provide differential results with kinetic energy-dependent spectrum.



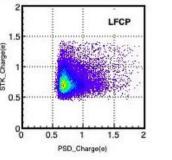
Eng 6.026GeV 6.918GeV



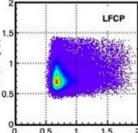
Eng 12.023GeV 13.804GeV



Eng 23.988GeV 27.542GeV

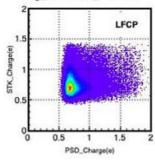




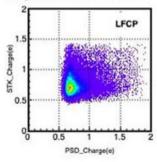




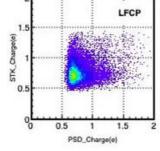
PSD Chargele



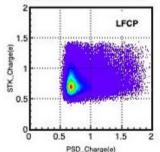
Eng_13.804GeV_15.849GeV

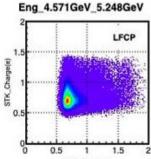


Eng 27.542GeV 31.623GeV



Eng 3.981GeV 4.571GeV

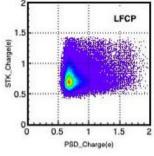




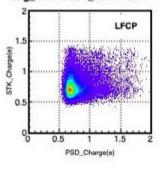
PSD Chargele

Eng 9.120GeV 10.471GeV

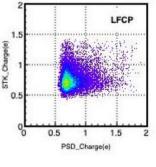
Eng 7.943GeV 9.120GeV

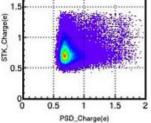


Eng 15.849GeV 18.197GeV

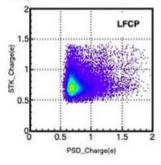


Eng 31.623GeV 36.308GeV

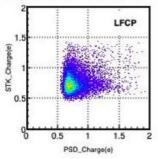




Eng 18,197GeV 20,893GeV



Eng 36 308GeV 41 687GeV



Summary

- Space experiments can detect FCPs with energy as low as a few GeV
- No 2/3e FCP signals are observed and a flux upper limit of $\Phi < 6.2 \times 10^{-10} \,\mathrm{cm}^{-2} \mathrm{sr}^{-1} \mathrm{sr}^{-1}$ is established at the 90% C.L.
- Result is published in *Phys. Rev. D 106, (2022) 6-15*
- Record in Particle Data Group
- Searching for Light FCP is on-going.

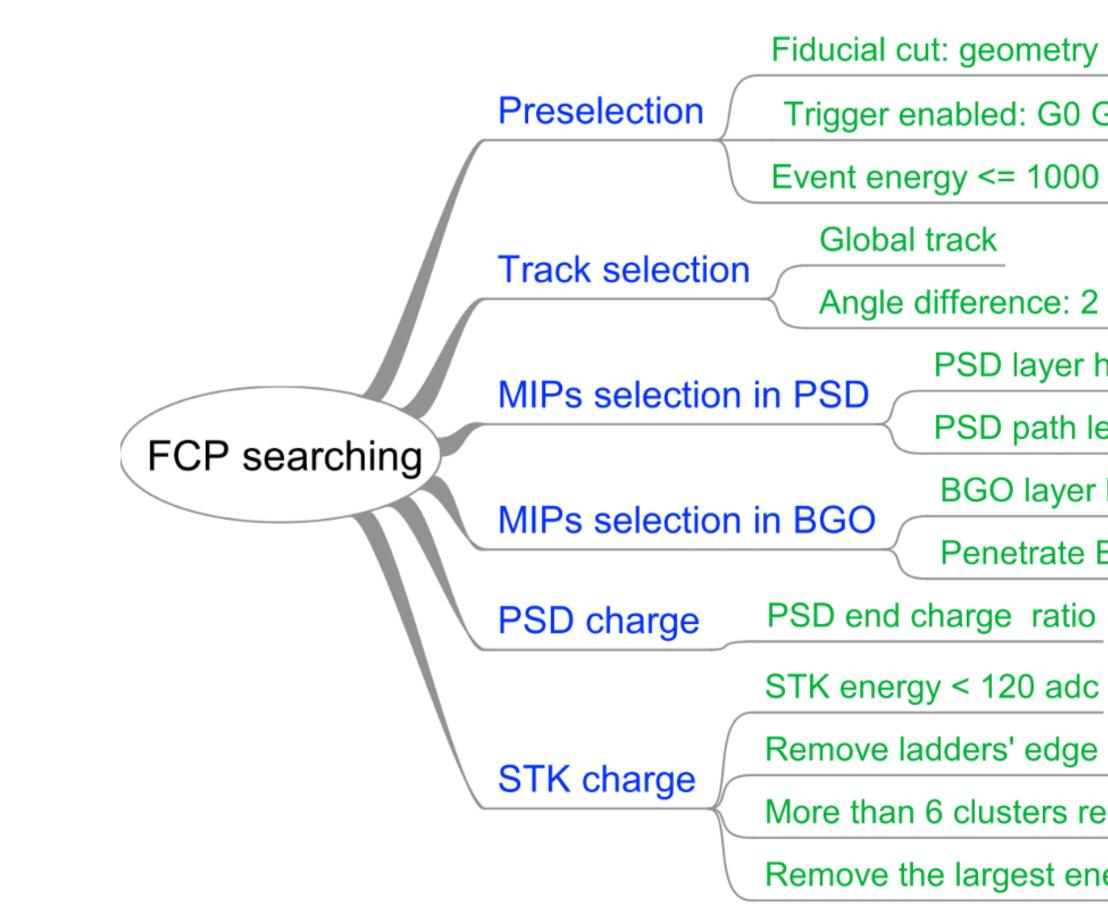
Quark Flux — Cosmic Ray Searches				PDGID:SC	027F INSPIRE Q	
Shielding values followed with an asterisk indicate altitude in km. Shielding values not followed with an asterisk indicate sea level in kg/cm ² .						
$\textit{FLUX}(\text{cm}^{-2}\text{sr}^{-1}\text{s}^{-1})$	$C\!H\!G\left(e/3 ight)$	$\textit{MASS}({\rm GeV})$	SHIELDING	DOCUMENT ID		TECN
< 6.2E-10	± 2			¹ ALEMANNO	2022	DAMP
< 1.E-8	±1/6-1/10			² AGNESE	2015	CDMS
<9.2E-15	± 1		3800	³ AMBROSIO	2000C	MCRO

Thank you!

Backup



Selections

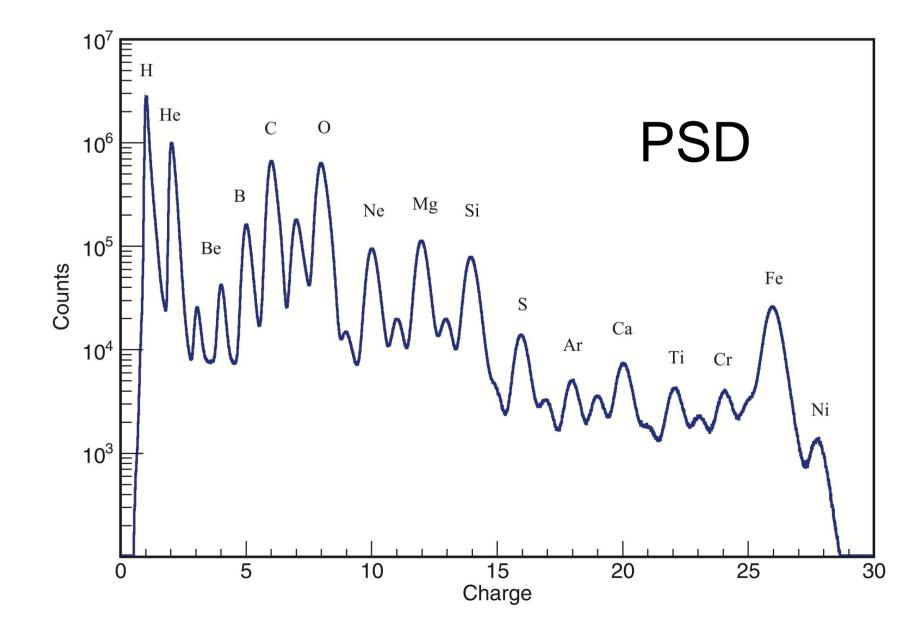


- Fiducial cut: geometry angle
- Trigger enabled: G0 G1 G2
- Event energy <= 1000 MeV

 - Angle difference: 2σ
 - PSD layer hit num <= 2
 - PSD path length = 10 mm
 - BGO layer hit num <= 2
 - Penetrate BGO: layer 12 or 13
- Remove ladders' edge cluster
- More than 6 clusters remain
- Remove the largest energy

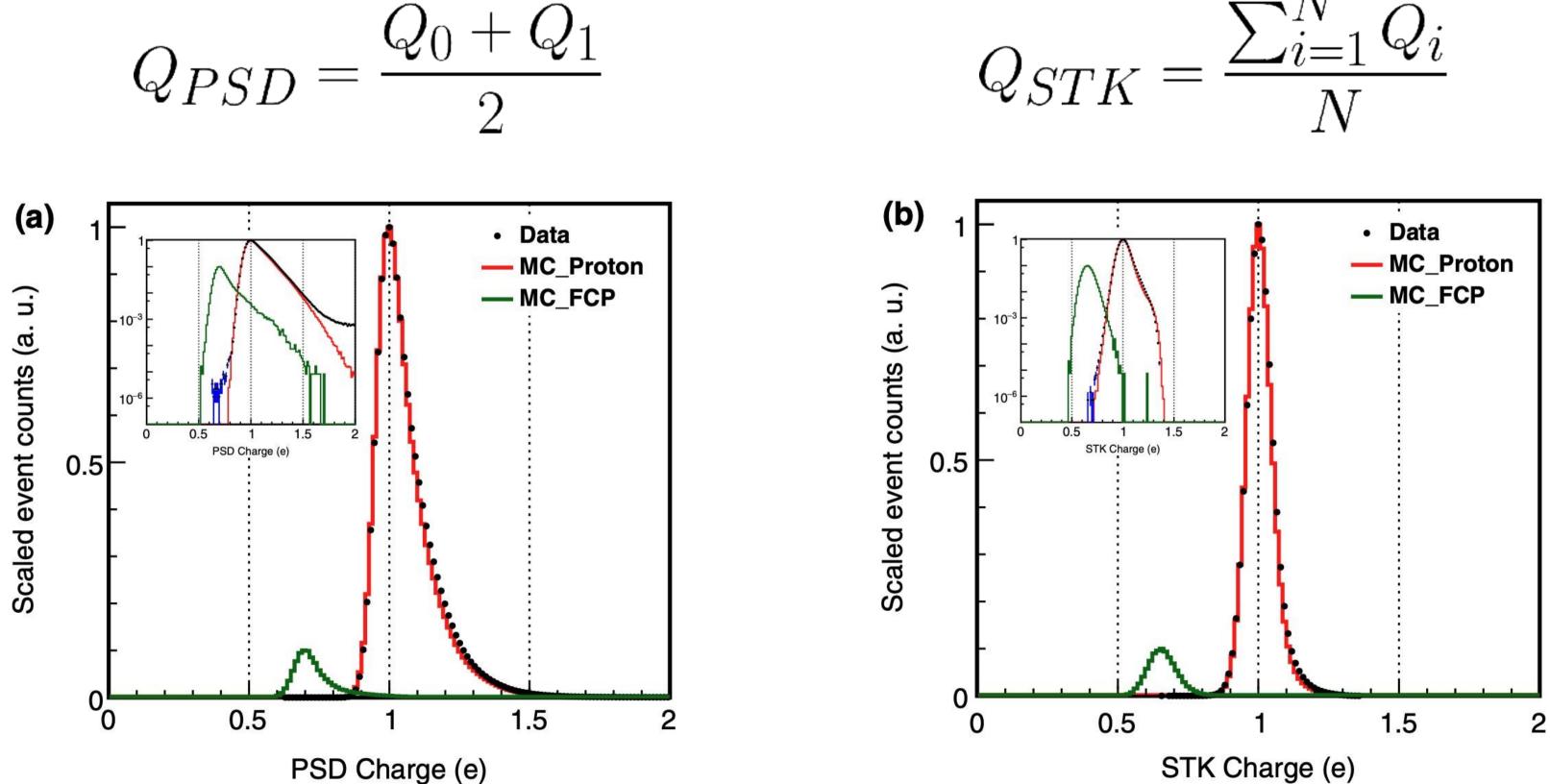
Charge Measurement

 $-\frac{dE}{dx} = K z^{2} \frac{Z}{A} \frac{1}{\beta^{2}} \left[\frac{1}{2} \ln \frac{2m_{e}c^{2}\beta^{2}\gamma^{2}T_{\max}}{I^{2}} - \beta^{2} - \frac{1}{2} \ln \frac{2m_{e}c^{2}\beta^{2}\gamma^{2}}{I^{2}} + \frac{1}{2} \ln \frac{2m_{e}c^{2}\beta^{2}}{I^{2}} + \frac{1}$ $\delta(\beta\gamma)$ 2

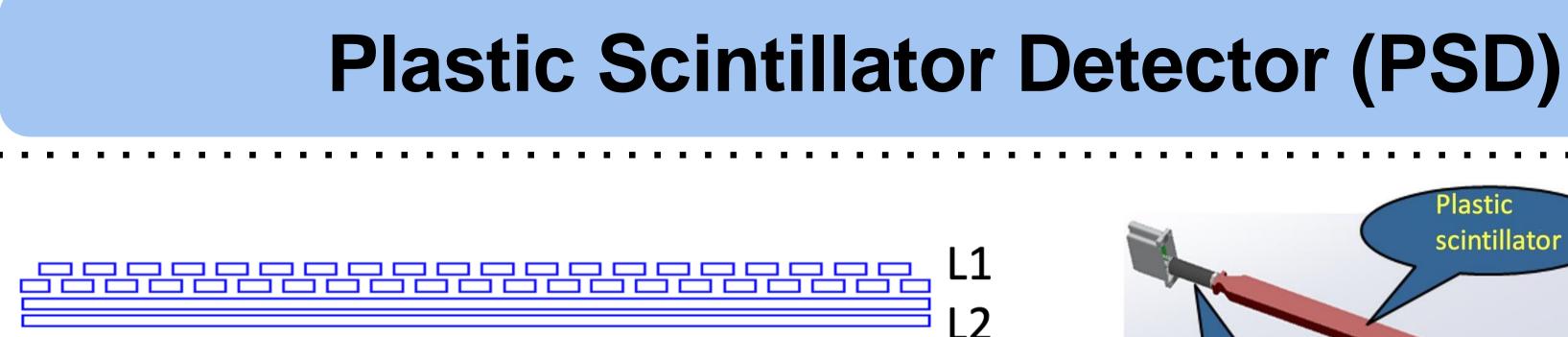


	PSD Charge resolution (Charge unit, c.u.)	STK Charge resolution (Charge unit, c.u.)		
Proton	0.06	0.04		
Helium	0.10	0.07		

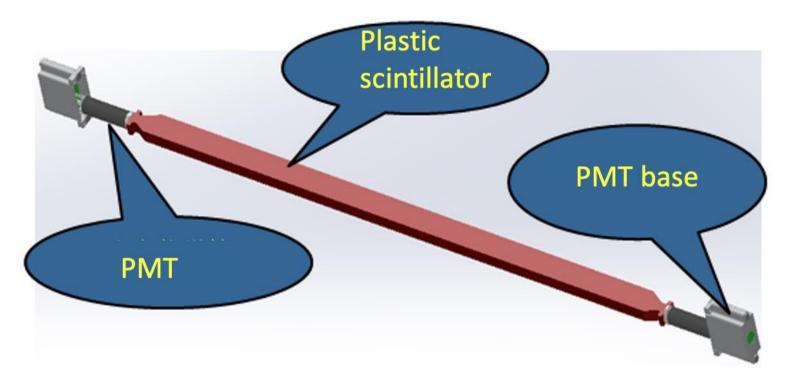
Charge Reconstruction

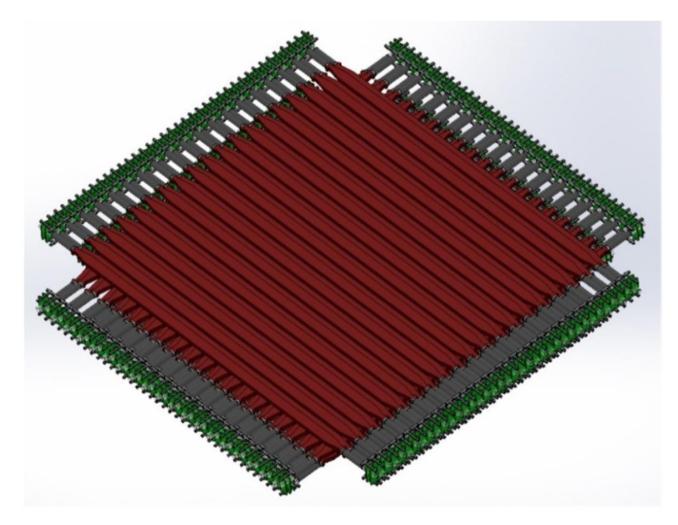


 $Q_{STK} = \frac{\sum_{i=1}^{N} Q_i}{N}$



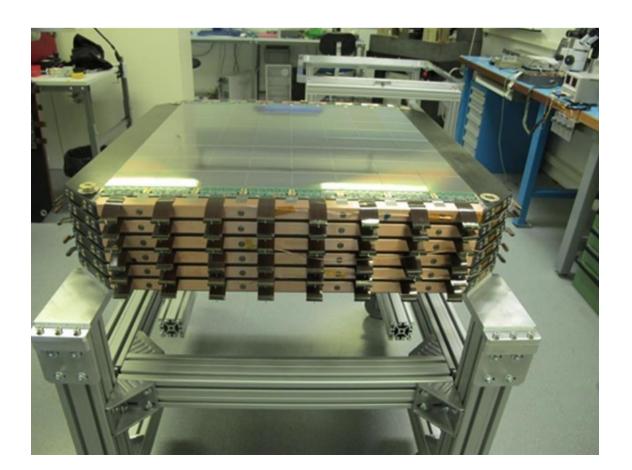
- PSD is located on the top of the payload
 - Active area: 82 cm × 82 cm
 - Number of planes: 2
 - -41 strips each layer
 - Overall efficiency ≥0.9975 for charged particles

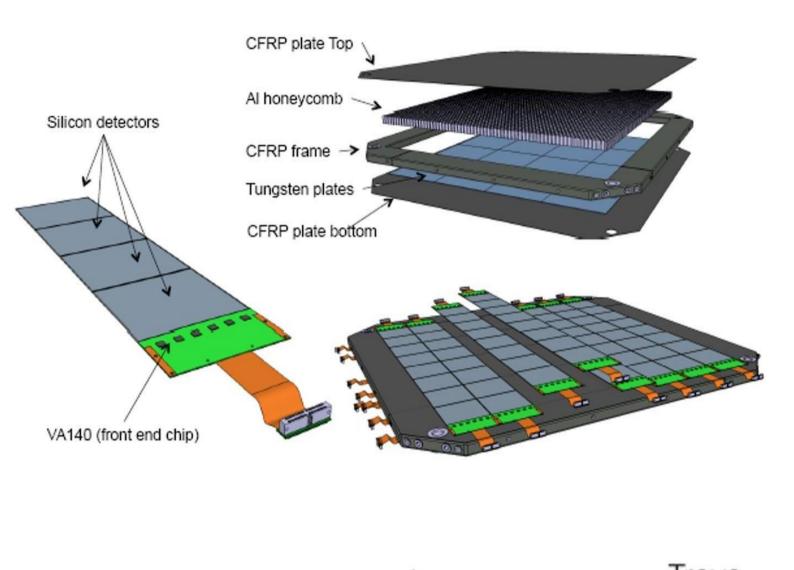


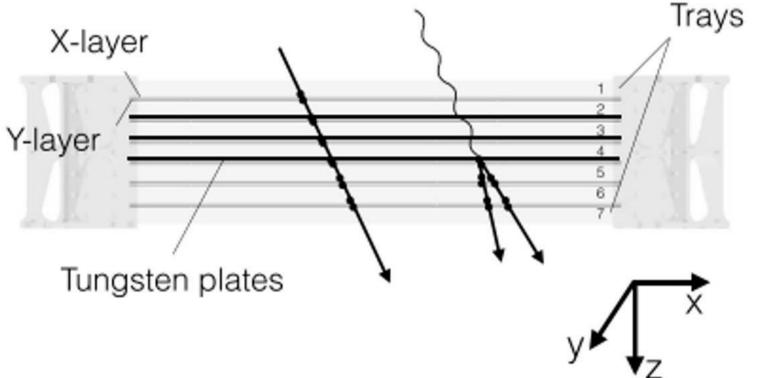


Silicon Tungsten tracKet (STK)

- Structure of STK
 - Pitch of silicon micro-strip: 121 µm
 - Active area: 75.8 × 75.8 cm²
 - -6 Planes (6X + 6Y)
 - Three 1 mm thickness tungsten layers embed in the STK







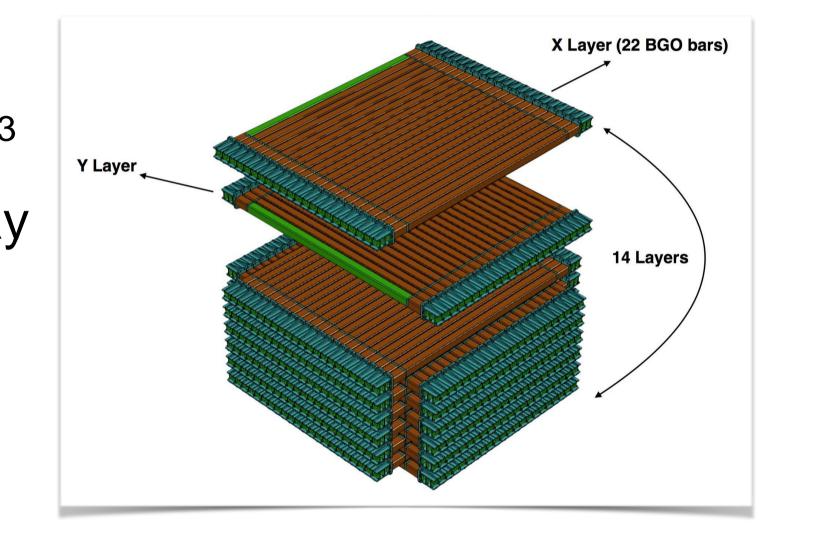
31

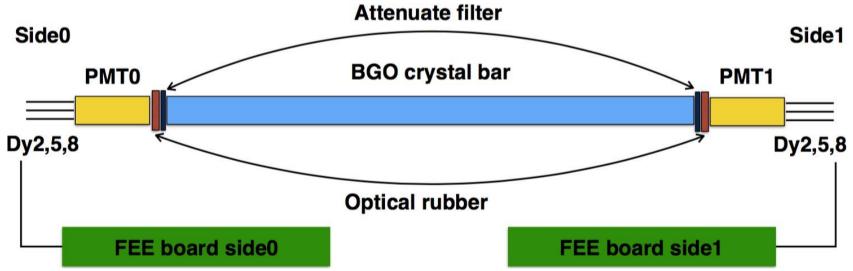
BGO calorimeter

- 14 × 22 BGO crystal array
 - Dimension of a BGO bar: 2.5×2.5×60cm³
 - Layers are alternated in an orthogonal way
- Thickness: $32X_0$, $1.6\lambda_I$
- Each end of BGO bar is coupled to a PMT



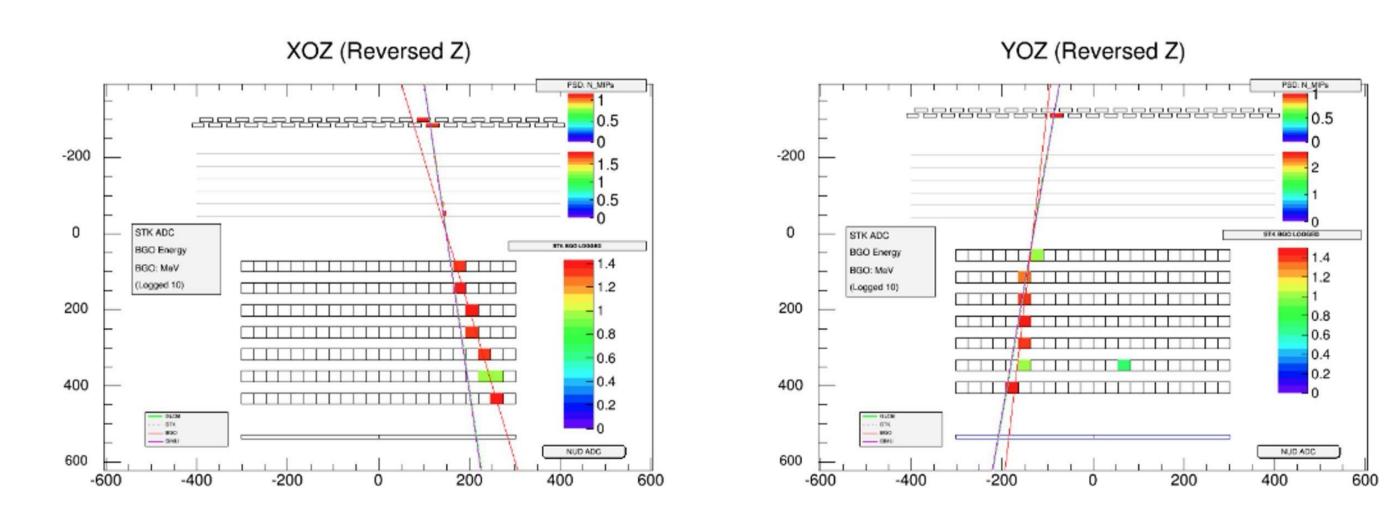


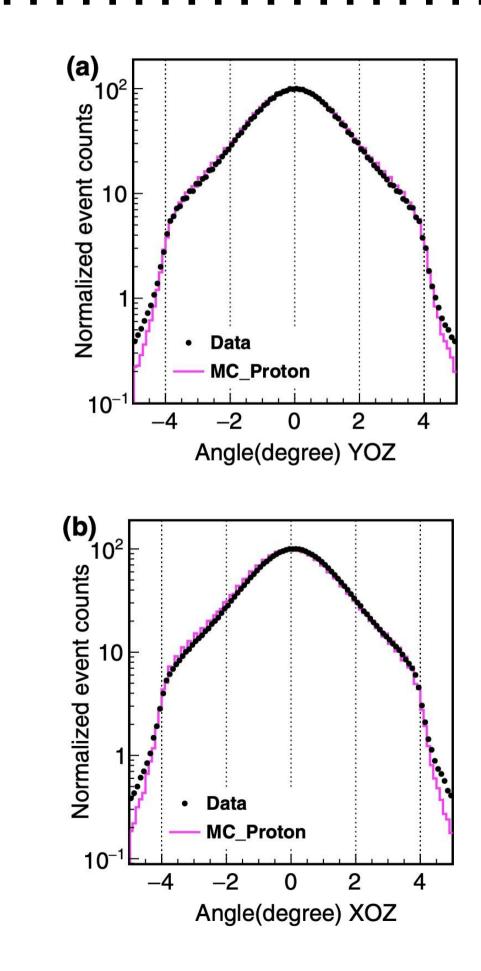




Track Selection

- Select STK tracks with good qualities
- Constrain the angle difference between STK Track and BGO Track
 - Angle difference $< 4^{\circ}$
 - Reject the scattering events

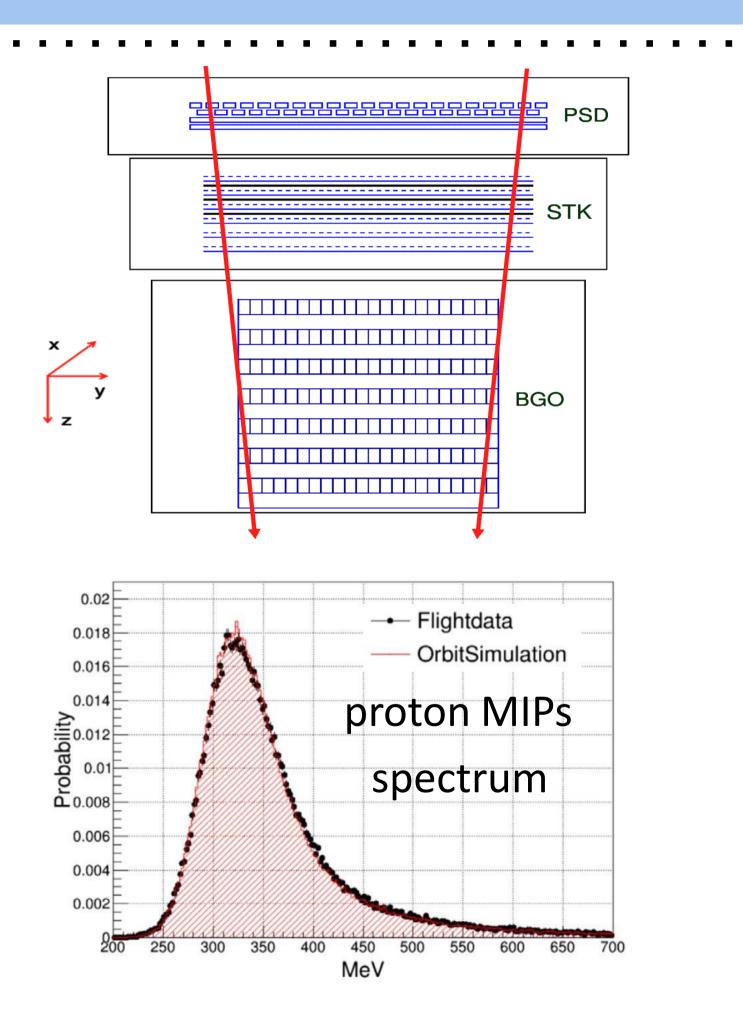




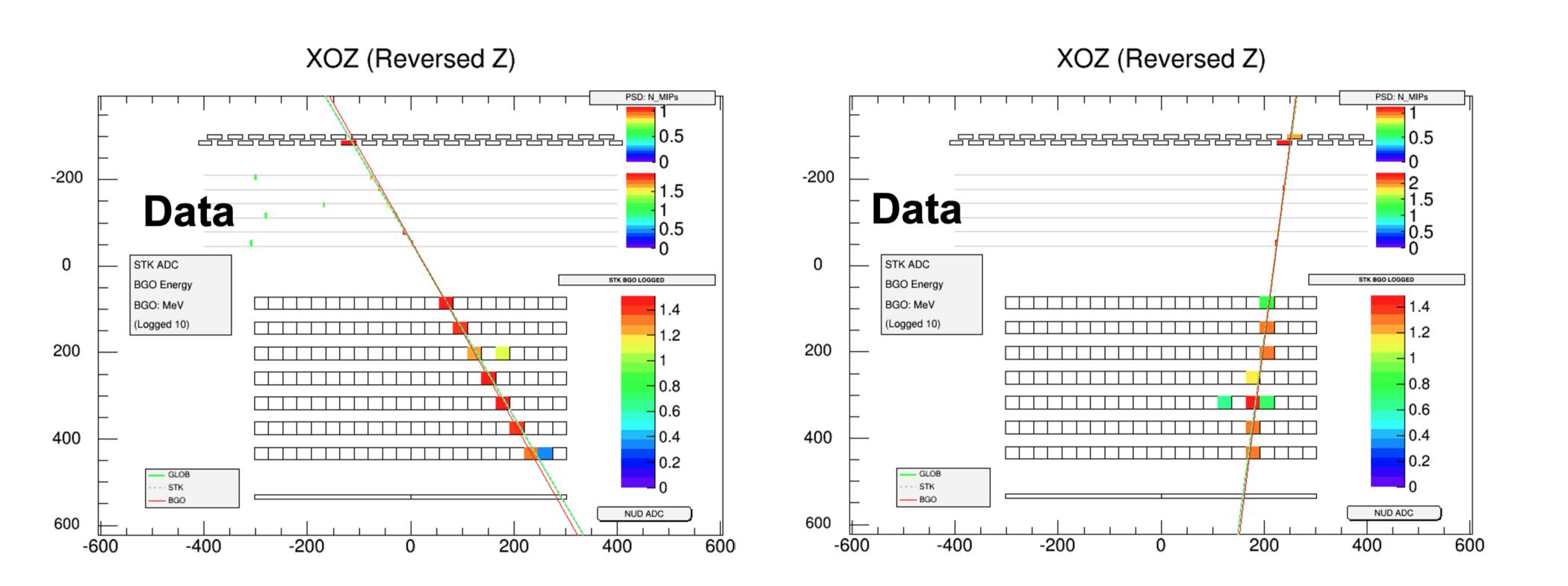
Pre-Selection

Pass MIPs trigger

- Fiducial selection: Constrain the positions of injection and ejection to maintain the event in the whole detector
- Total energy selection:
 - Energy deposition in ECAL < 1 GeV
 - Reject particles with charges higher than proton



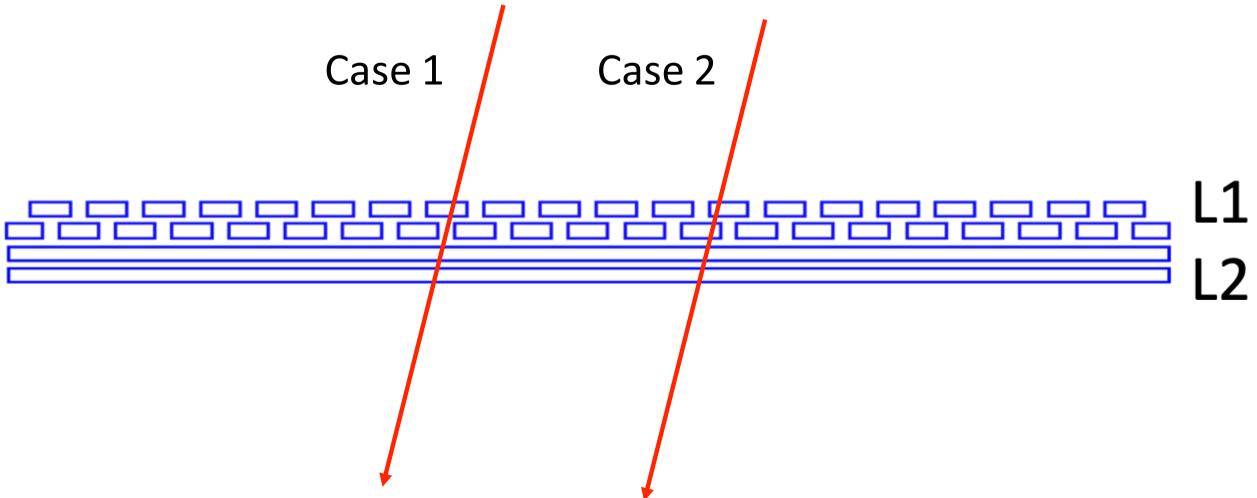
MIP-like Event Selection





MIP-like Event Selection

- MIPs in PSD:
 - The number of fired strip in one layer ≤ 2
 - The selected track should cross the strip with maximum energy

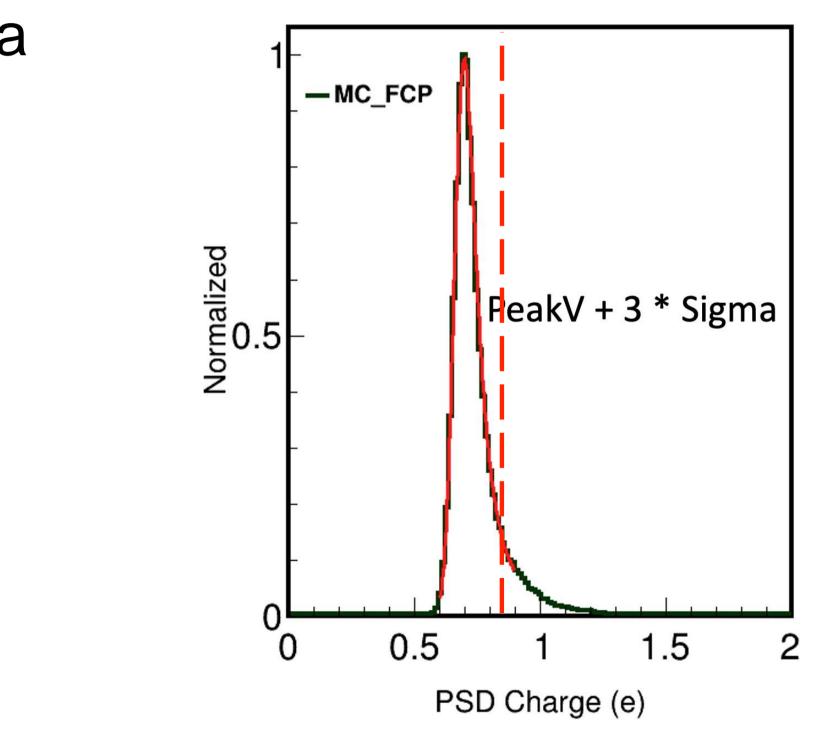




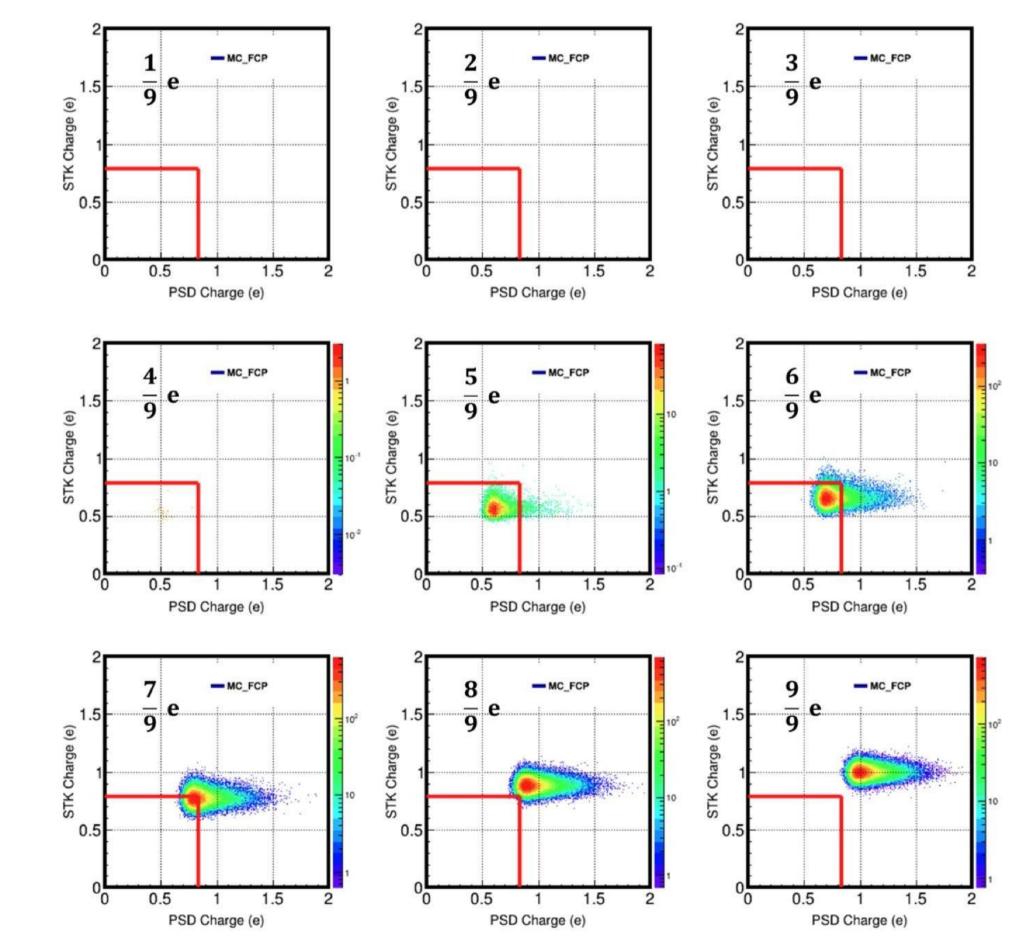
Signal Region Definition

• Signal region: Peak of MC FCP + 3sigma

- Signal region on PSD: Z < 0.84e
- Signal region on STK: Z < 0.79e



Charge scan



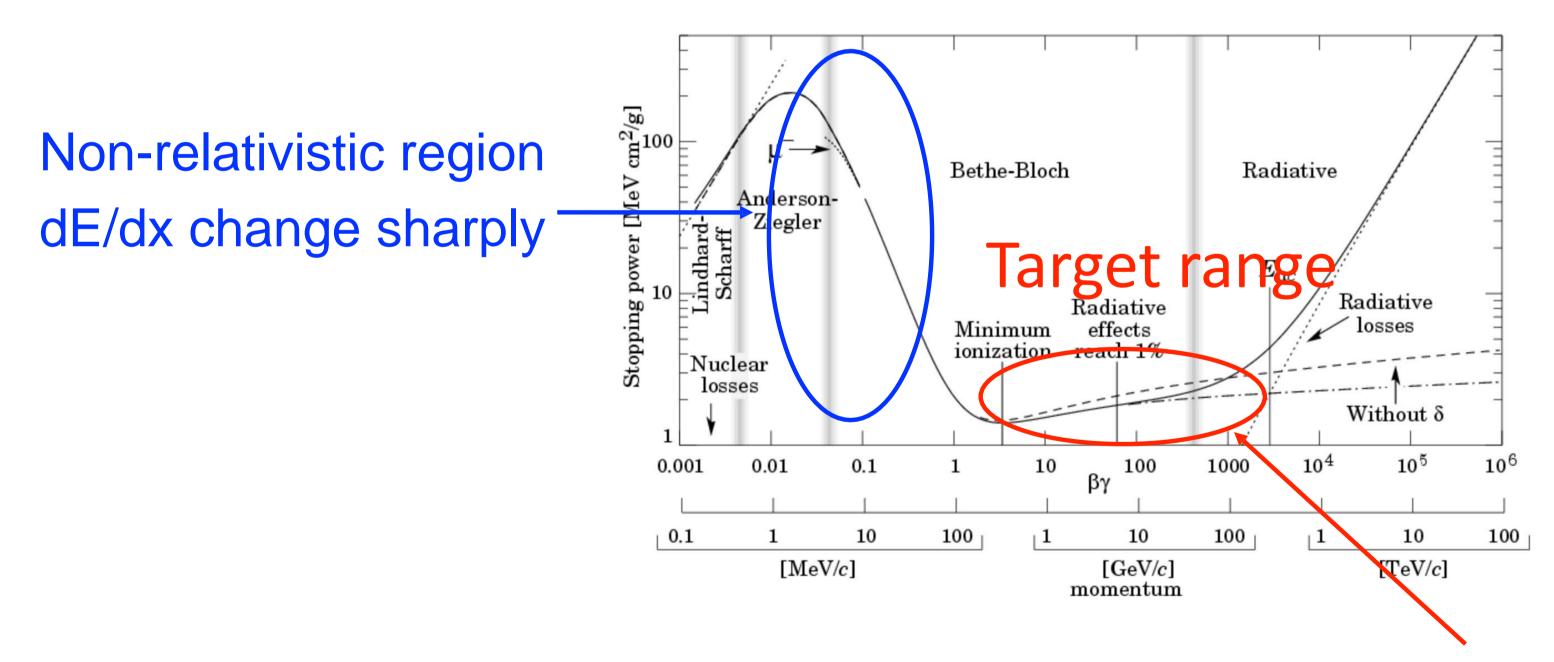
Low charge:

No response from detector High charge: Difficult to distinguish from

the background

Kinetic Energy Control

• Energy of MIP-like events cannot be measured by a calorimeter Kinetic energy should be constrained



dE/dx is stable

Why did N_{obs} choose 2.44?

According to the Feldman-Cousins method, the observed small signal events *n* obeys the Poisson distribution $p(n|s) = e^{-(s+b)} \frac{(s+b)^n}{n!}$, the real signal events *s* and background events *n* obey the Poisson distribution as well. Since *n* and *b* are zero, the *s* takes 2.44 for the calculation of upper limit within 90% confidence level.

	$n_0 b$	0.0	0.5
	0	0.00, 2.44	0.00, 1.94
90% C.L.	1	0.11, 4.36	0.00, 3.86
	2	0.53, 5.91	0.03, 5.41
	3	1.10, 7.42	0.60, 6.92
	4	1.47, 8.60	1.17, 8.10

95% C.L.



$n_0 \setminus b$	0.0	0.5
0	0.00, 3.09	0.00, 2.63
1	0.05, 5.14	0.00, 4.64
2	0.36, 6.72	0.00, 6.22
3	0.82, 8.25	0.32, 7.75
4	1.37, 9.76	0.87, 9.26