

A Novel Approach to Search for the Chiral Magnetic Effect from STAR and the Future Prospect

黄焕中 (Huan Zhong Huang)
University of California Los Angeles

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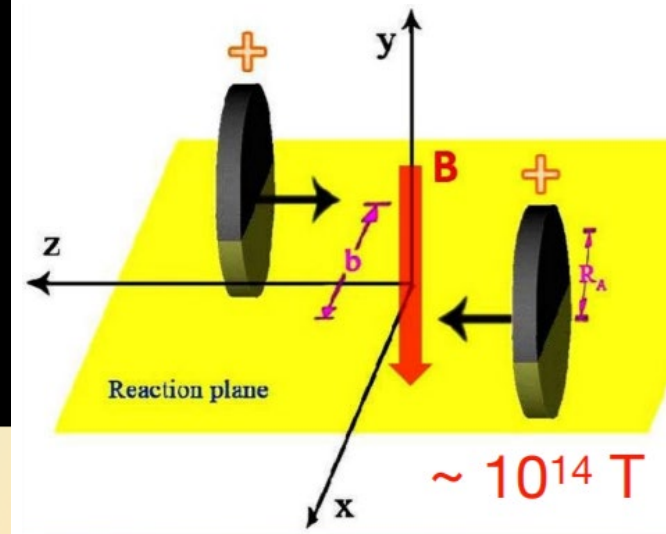
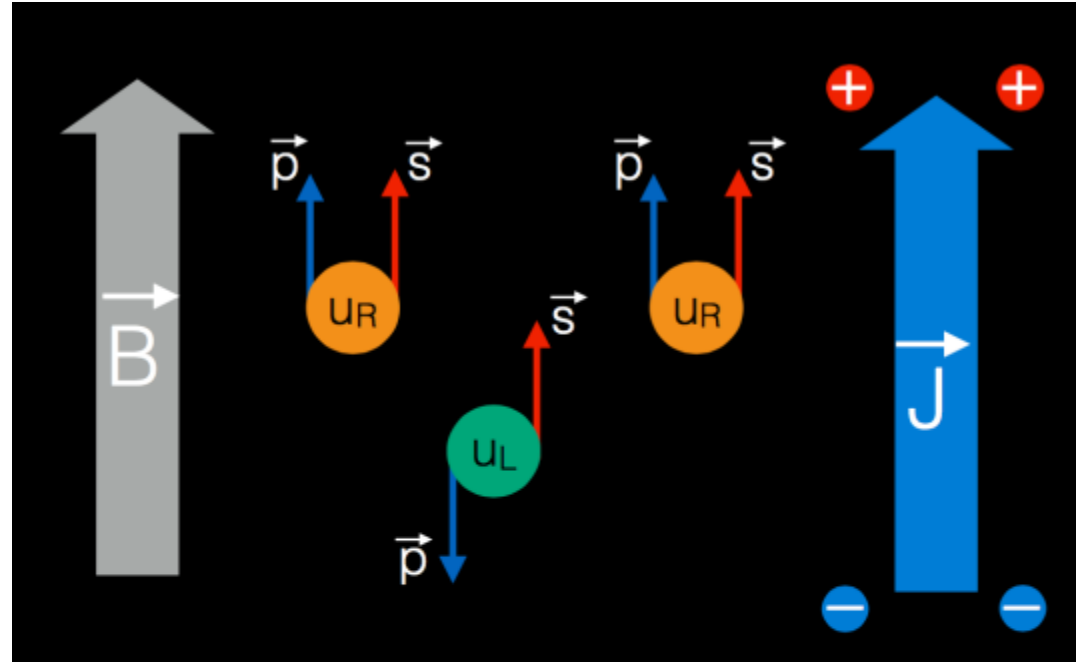
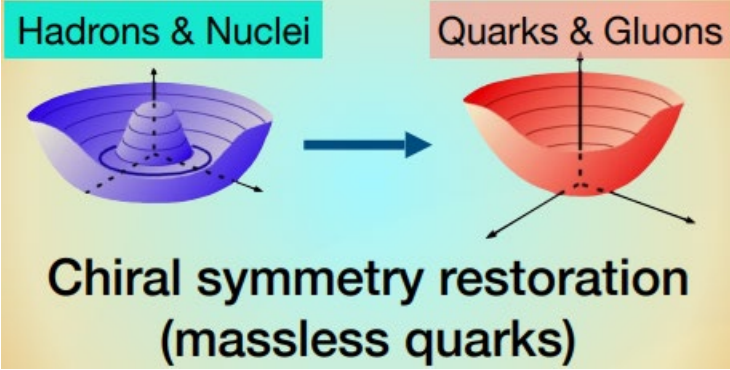
Thanks to Gang Wang, Zhiwan Xu, Jinfeng Liao, Jinhui Chen and Diyu Shen

Outline

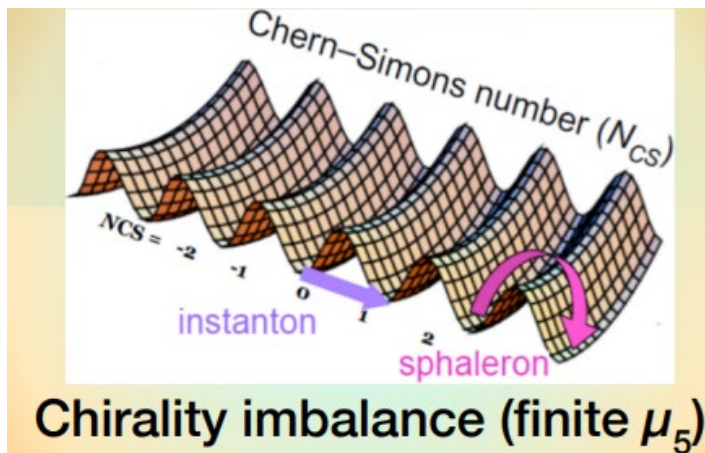
- ❖ Introduction
- ❖ Lessons from Previous Results
- ❖ Novel Approach of Event Shape Selection (ESS)
- ❖ STAR ESS Results from BES-II and 200 GeV Data
- ❖ Summary and Future Outlook

Chiral Magnetic Effect

A Rare Opportunity to Experimentally Access Key Intrinsic Properties of the QCD



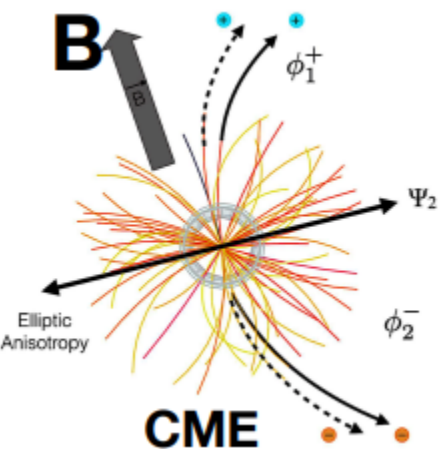
Strong magnetic field (B)



Chiral Magnetic Effect ($\mathbf{J} \parallel \mathbf{B}$)

$$\vec{J} = \frac{e^2}{2\pi^2} \mu_5 \vec{B}$$

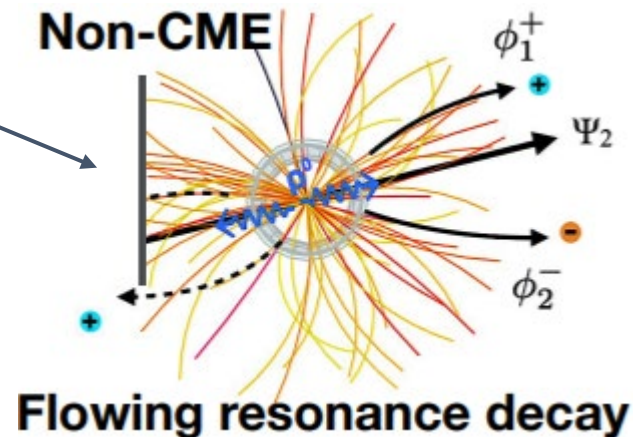
CME Observables



$$\frac{dN_{\pm}}{d\varphi} \propto 1 + 2v_1 \cos(\varphi - \Psi_{RP}) + \boxed{2a_1^{\pm}} \sin(\varphi - \Psi_{RP}) + \boxed{2v_2} \cos(2\varphi - 2\Psi_{RP}) + \dots$$

$\propto \mu_5 B$

Parity odd, can not directly observe



Popular CME-sensitive observables:

- γ correlator

S.A. Voloshin, Phys. Rev. C70(2004)057901

- R correlator

N. N. Ajitanand et al., Phys. Rev. C83(2011)011901(R)

- Signed balance functions

A. H. Tang, Chin. Phys. C44, No.5 (2020)054101

Model studies show that these methods have **similar sensitivities** to the CME signal and to the background. (Best Paper Award 2023)

S. Choudhury et al.(STAR), Chin. Phys. C46(2022)014101

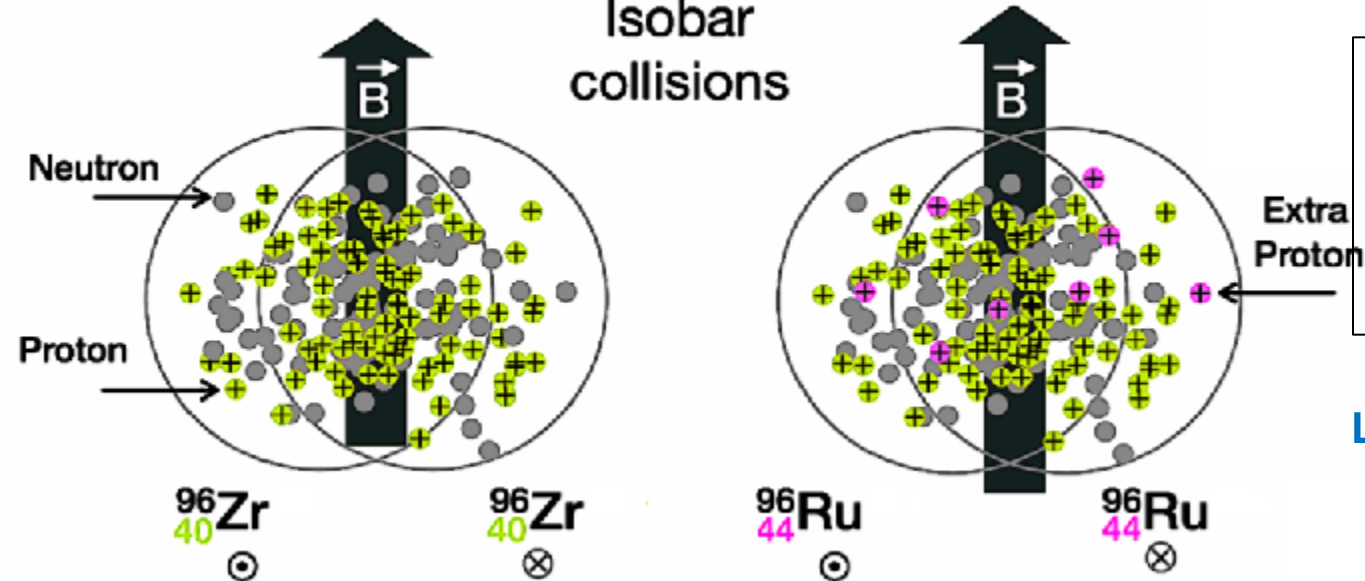
Here, we focus on $\gamma^{112} \equiv \langle \cos(\varphi_{\alpha} + \varphi_{\beta} - 2\Psi_{RP}) \rangle$

The CME causes $\Delta\gamma^{112} \equiv \gamma_{OS}^{112} - \gamma_{SS}^{112} > 0$

Background indicator $\gamma^{132} \equiv \langle \cos(\varphi_{\alpha} - 3\varphi_{\beta} + 2\Psi_{RP}) \rangle$

Lessons from Isobar Collisions

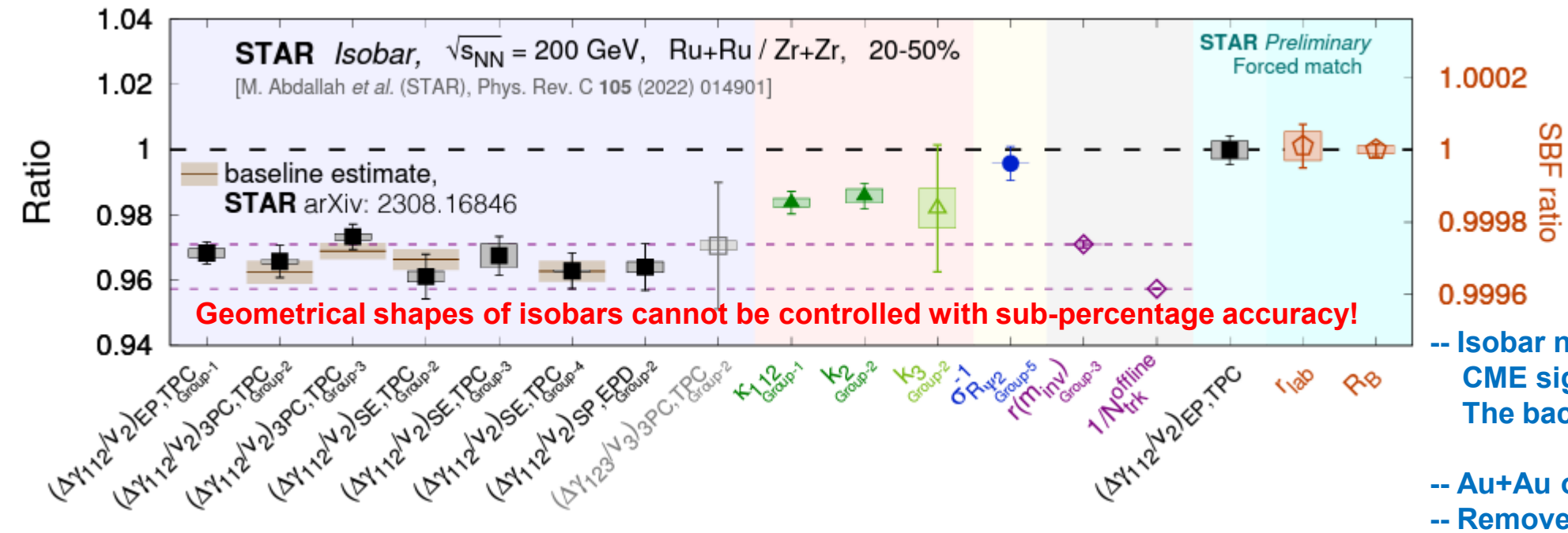
Isobar collisions



Compare the two isobaric systems:

- CME: $B\text{-field}^2$ is $\sim 15\%$ larger in Ru+Ru
- Flow-related Background: utilize $\Delta\gamma_{112}/v_2$
- Nonflow-related Background: almost same

Looking for difference in the signal due to Z difference
 $< 15\%$ of the signal strength !



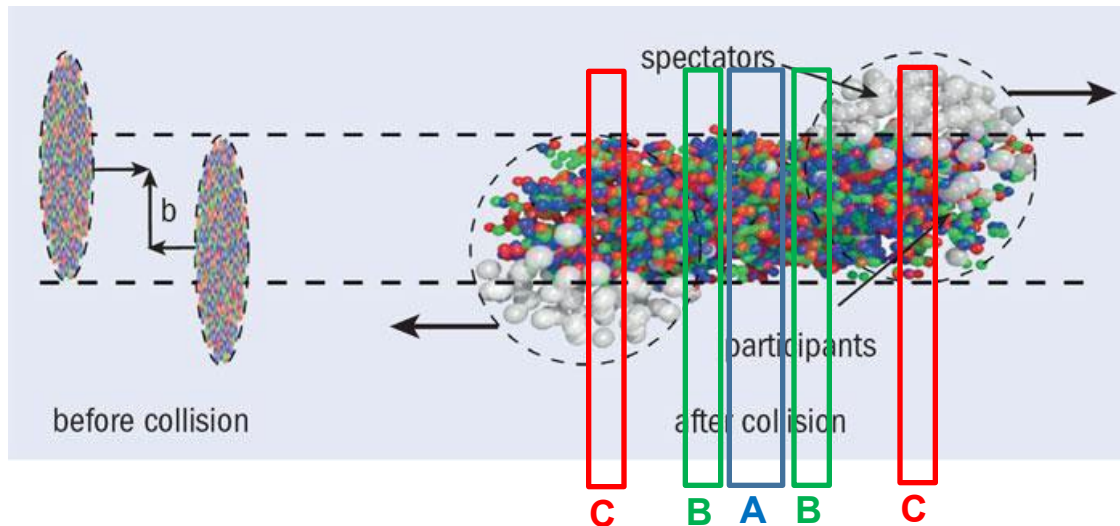
- Isobar not sensitive enough
- CME signal is smaller than expected
- The background fluctuation cannot be controlled well
- Au+Au or Pb+Pb better system
- Remove background essential

Previous Event Shape Method

Previous Event Shape Engineering (ESE) Approach

“Standard” ESE splits an event into 3 sub-events

- (A) particles of interest (POI)
- (B) particles to construct q_n shape
- (C) particles to reconstruct EP



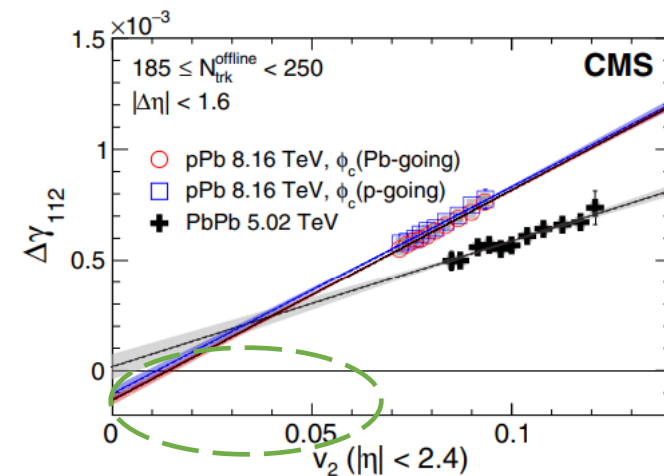
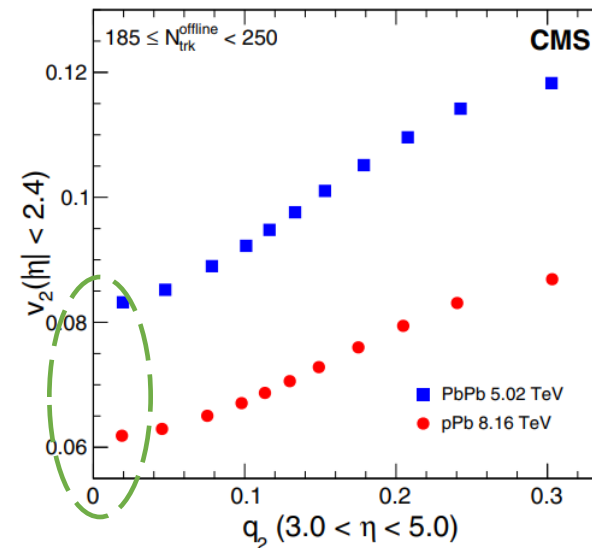
Sensitive to eccentricity which limited range of variation for a given centrality !
CME background – overall particle emission !

We found

Shape Observable flow vector q_n in region B not effective in selecting shape for particles A

Flow vector q from B correlated to $\langle v_2 \rangle$

Extreme shape fluctuations are largely local, not global feature!



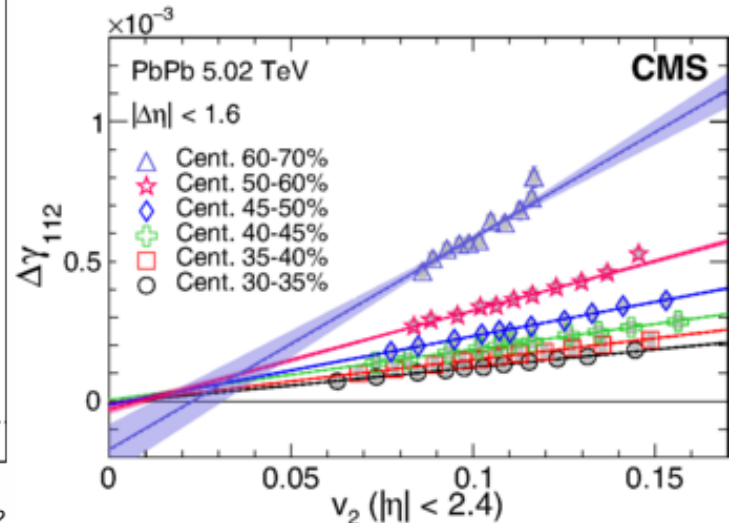
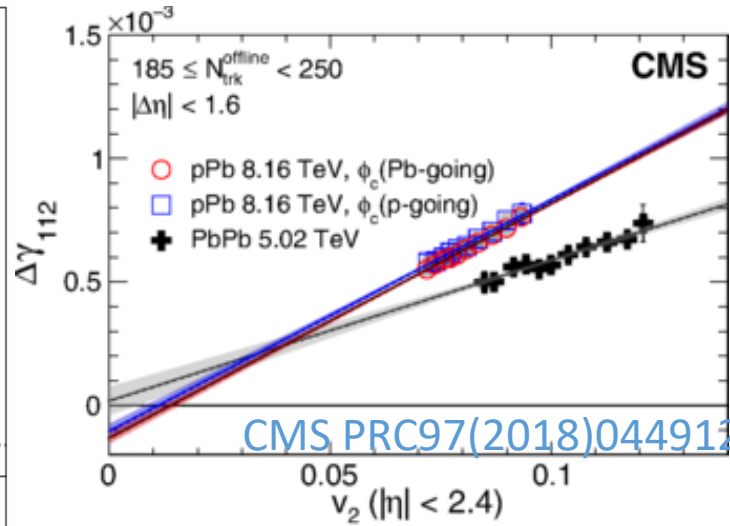
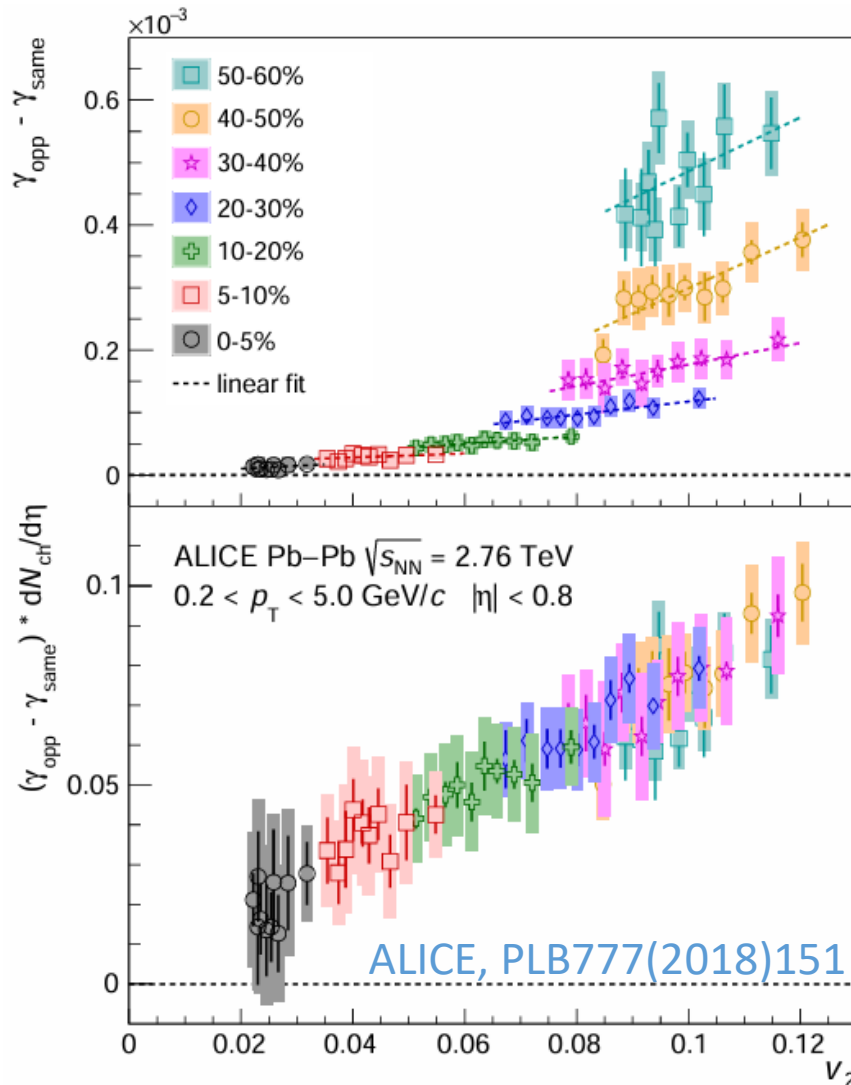
CMS, PRC 97(2018)044912

“Standard” Event Shape Engineering

Three sub-events are used: one for POI, one for event plane, and one for event shape variable, q_2 , the modulus of the flow vector.

$$q_x \equiv \frac{1}{\sqrt{N}} \sum_i^N \cos(2\phi_i)$$

$$q_y \equiv \frac{1}{\sqrt{N}} \sum_i^N \sin(2\phi_i)$$



- Measure $\Delta\gamma_{112}$ vs q_2 and v_2 vs q_2 , then plot $\Delta\gamma_{112}$ vs v_2 , and finally extrapolate $\Delta\gamma_{112}$ to zero v_2 .
- At LHC energies, all the ESE results are consistent with zero. (too short duration of the B field?)
- Since **particles of interest (POI) are excluded from q_2** , the lever arm on v_2 is very weak, making the extrapolation **unstable**.

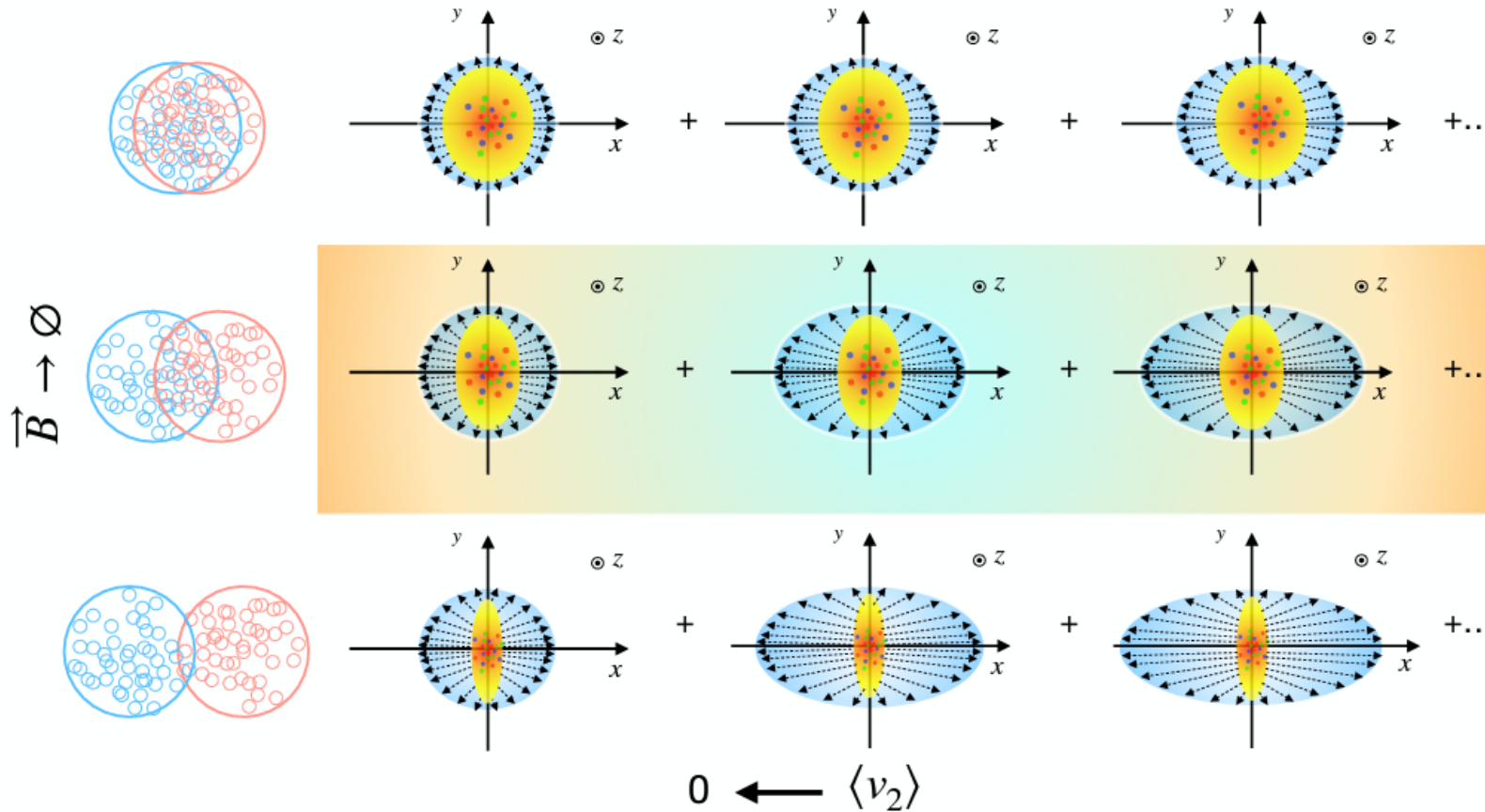
Event Shape Selection (ESS)

Ideally, if we control eccentricity, we control flow for everything.
But large event-by-event fluctuations could dominate the observable.

- participant zone geometry: expected to be long ranged in rapidity emission
- pattern fluctuations: more localized, less correlated over rapidity

H. Petersen and B. Müller,
Phys. Rev. C 88, 044918

Geometry Variation



Event shape variables based on **particles of interest (POI)** are sensitive to both geometry and emission pattern.

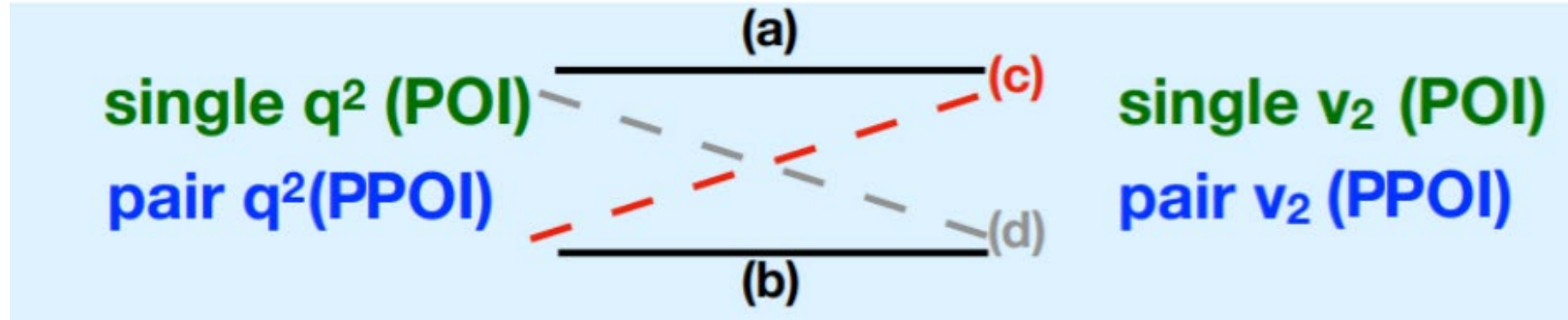
CME background e-by-e comes from combined eccentricity and emission patterns

Emission pattern fluctuation

Shape Variable and v2 Control

Event shape variable

Elliptic flow variable



$$q_2^2 = \frac{1}{N} \left[\left(\sum_{i=1}^N \sin 2\varphi_i \right)^2 + \left(\sum_{i=1}^N \cos 2\varphi_i \right)^2 \right]$$

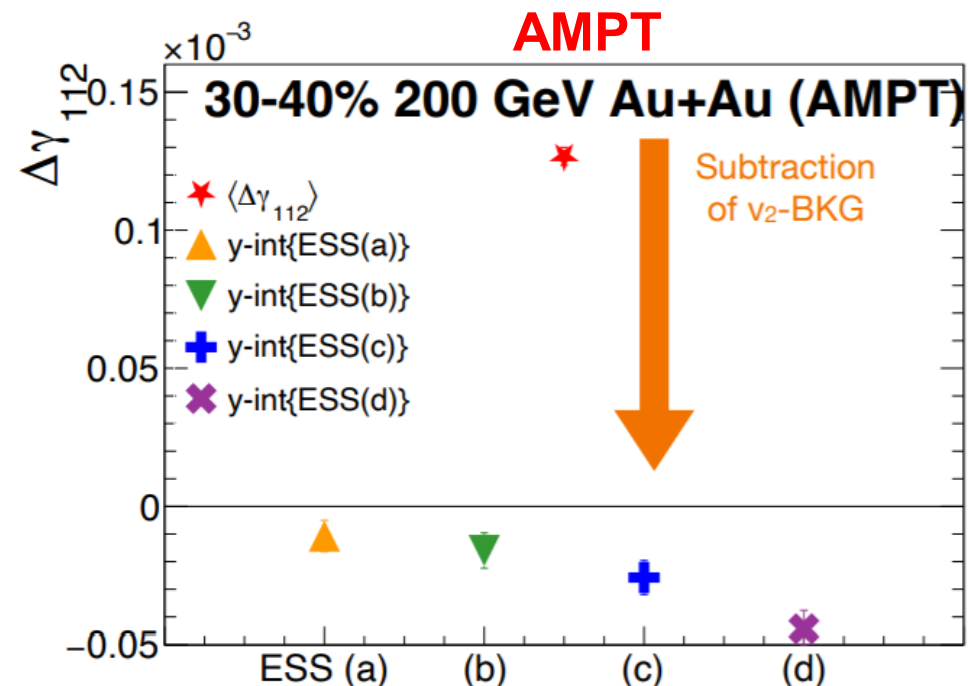
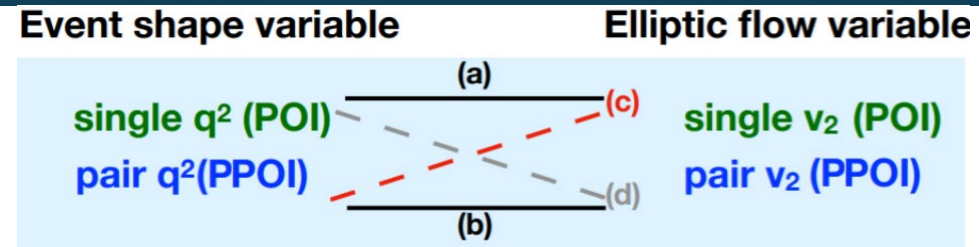
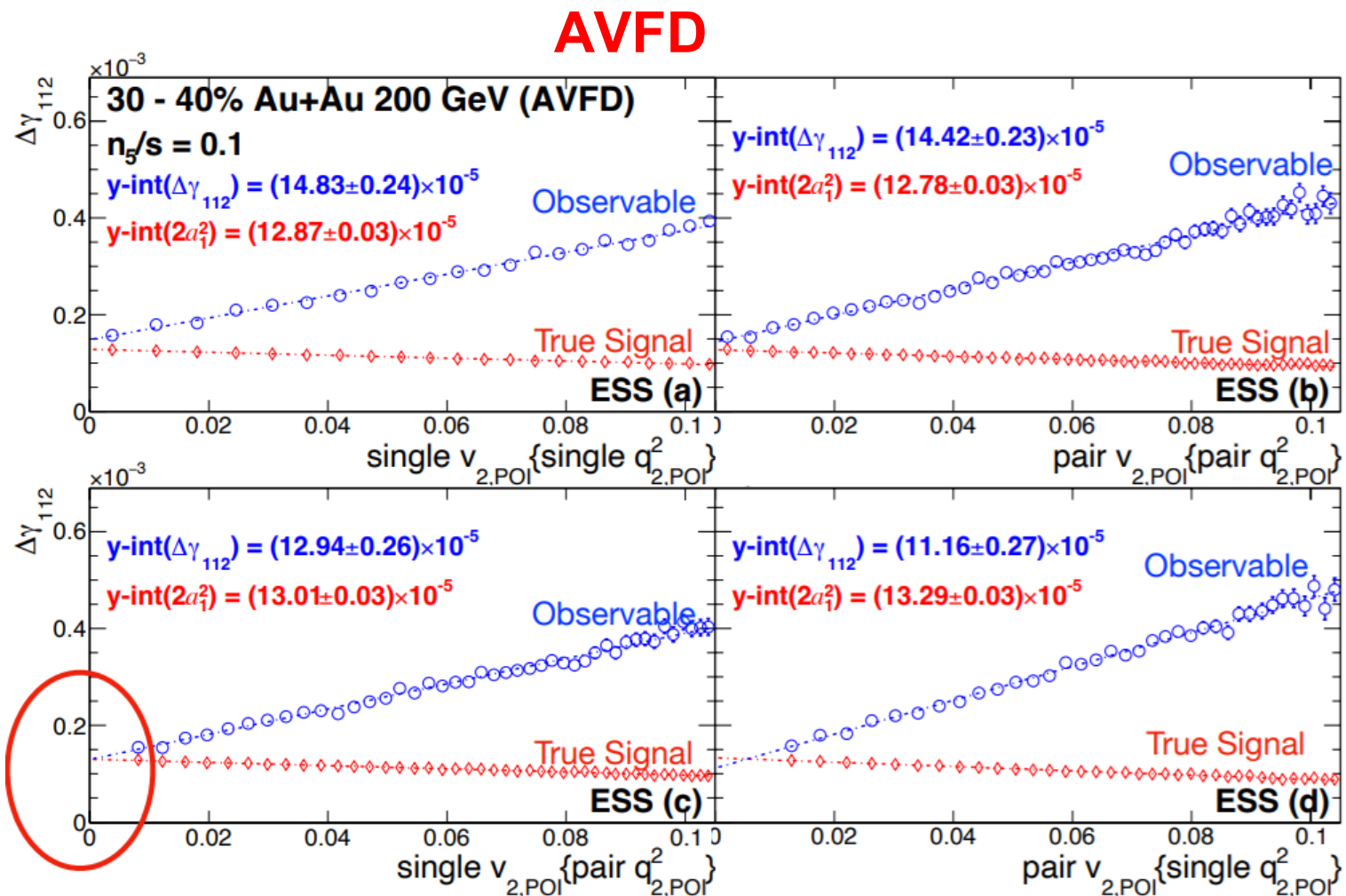
$$= 1 + \frac{1}{N} \sum_{i \neq j} \cos[2(\varphi_i - \varphi_j)],$$

- ESS recipes (a) and (b) involve direct event-by-event correlations between q_2^2 and v_2 , which will cause under-subtraction of background.
- We should use “mixed” recipes, (c) or (d).
- Redefine q_2^2 with an extra normalization.
- Pair q_2^2 and pair v_2 are based on φ_p .

$$\langle q_2^2 \rangle \approx 1 + N v_2^2 \{2\}$$

$$q_2^2 = \frac{\left(\sum_{i=1}^N \sin 2\varphi_i \right)^2 + \left(\sum_{i=1}^N \cos 2\varphi_i \right)^2}{N(1 + N v_2^2 \{2\})}$$

Simulations



Z. Xu *et al*, PLB 848(2024)138367

- AVFD: the optimal ESS recipe (c) accurately matches the input CME signal.
- Intercepts follow an ordering (a)>(b)>(c)>(d).
- AMPT: all ESS recipes over-estimate the BKG (with the same ordering as AVFD).

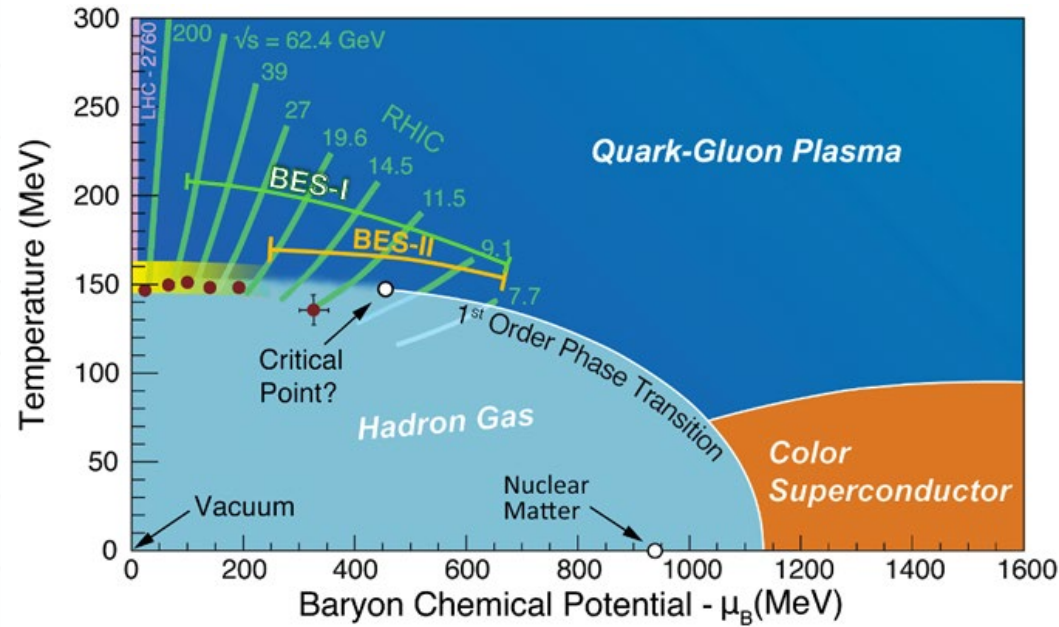
Application to Real Data

BES-I

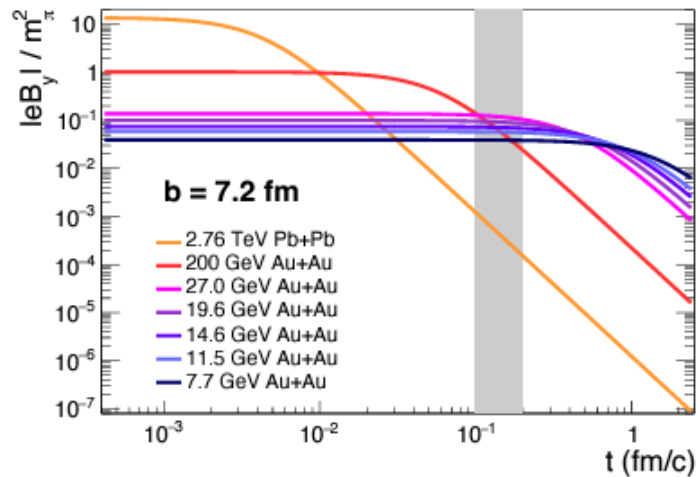
“Events” represents good events after quality cuts.

BES-II

$\sqrt{s_{NN}}$ (GeV)	Events (10^6)	Year
62.4	46	2010
39	86	2010
27	30	2011
19.6	15	2011
14.6	13	2014
11.5	7	2010
9.2	0.3	2008
7.7	4	2010



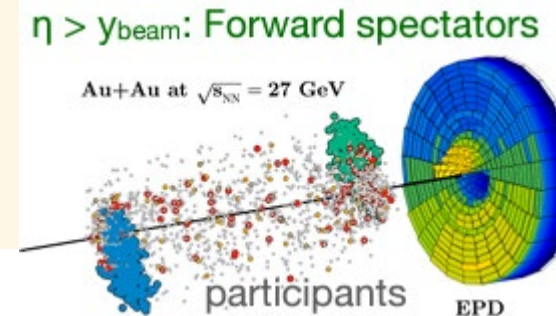
$\sqrt{s_{NN}}$ (GeV)	Events (10^6)	Year
27	555	2018
19.6	478	2019
14.6	324	2019
11.5	230	2020
9.2	160	2020
7.7	101	2021



Event Shape Selection **Spectator Ψ_1**

$$\Delta\gamma^{112} = \Delta\gamma^{\text{CME}} + \cancel{k \frac{v_2}{N}} + \cancel{\Delta\gamma^{\text{non-flow}}}$$

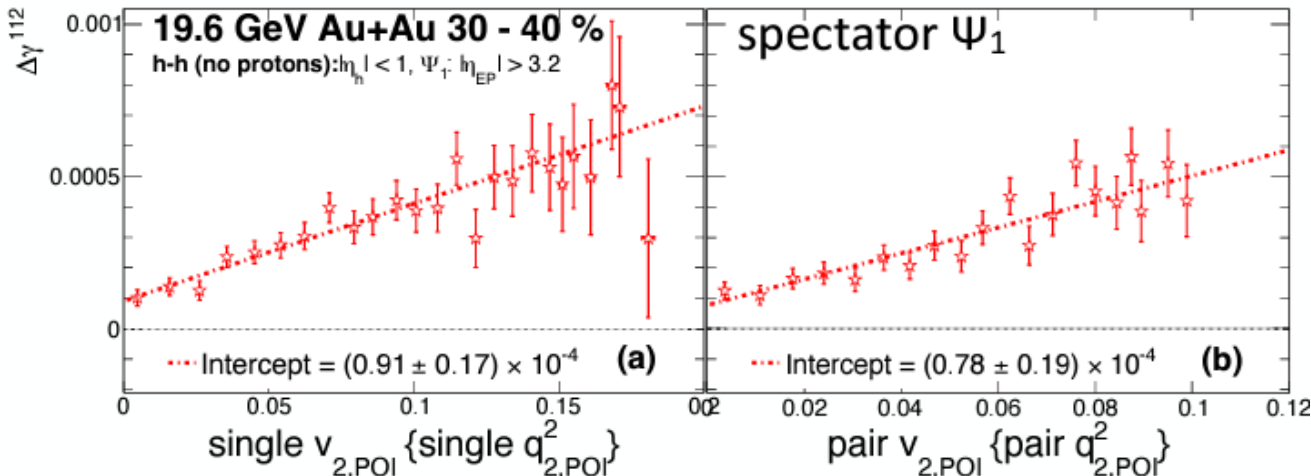
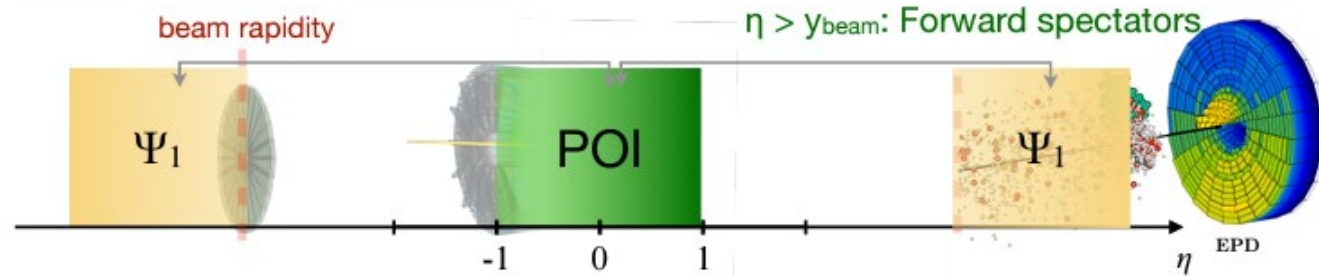
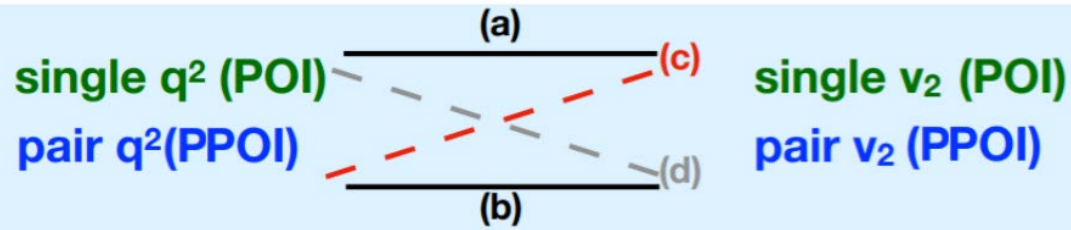
$\Delta\gamma^{112}$ → **Measured** $\Delta\gamma^{\text{CME}}$ → **Signal** $k \frac{v_2}{N}$ → **Backgrounds** $\Delta\gamma^{\text{non-flow}}$ → **Backgrounds**



Au+Au at 19.6 GeV

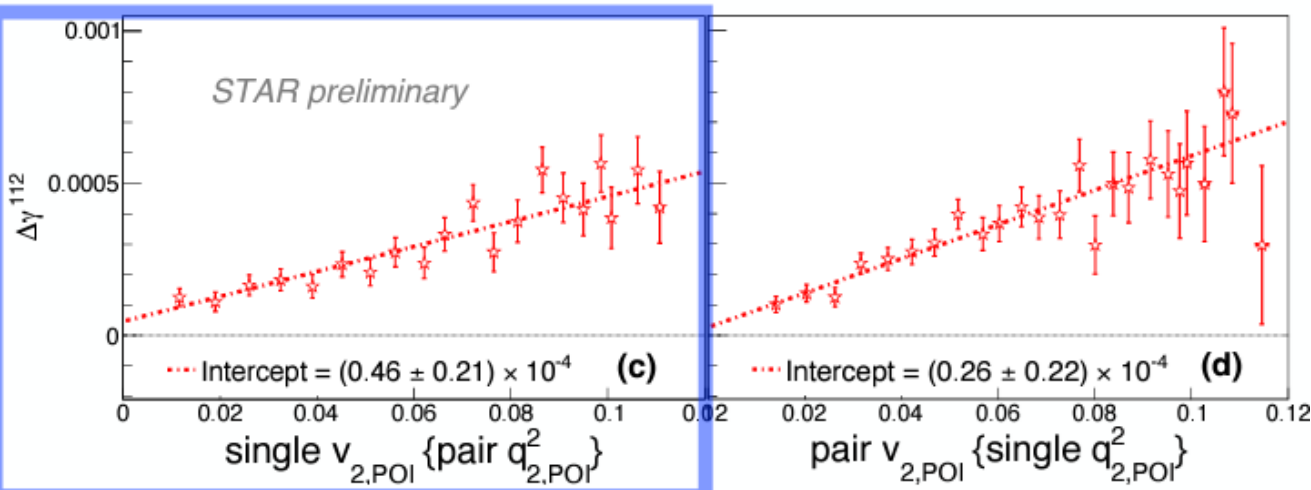
Event shape variable

Elliptic flow variable

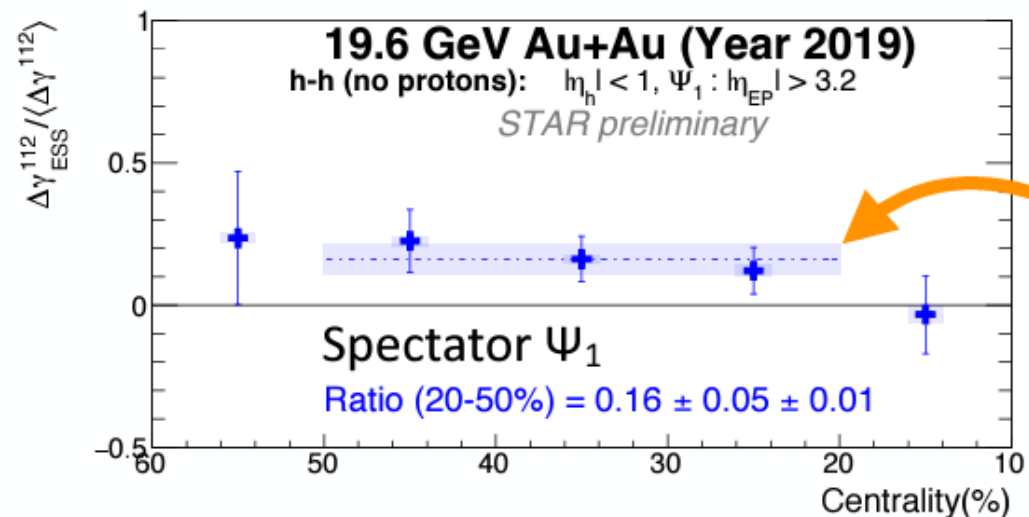
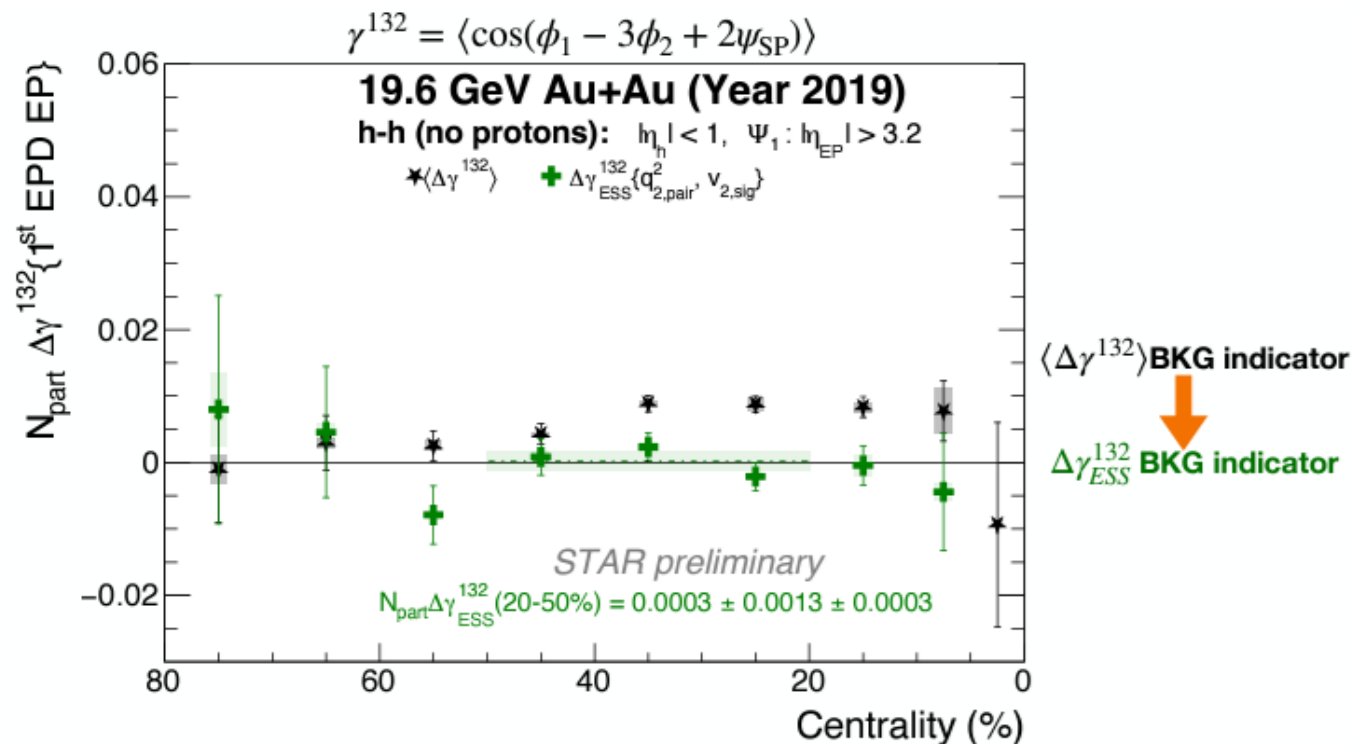
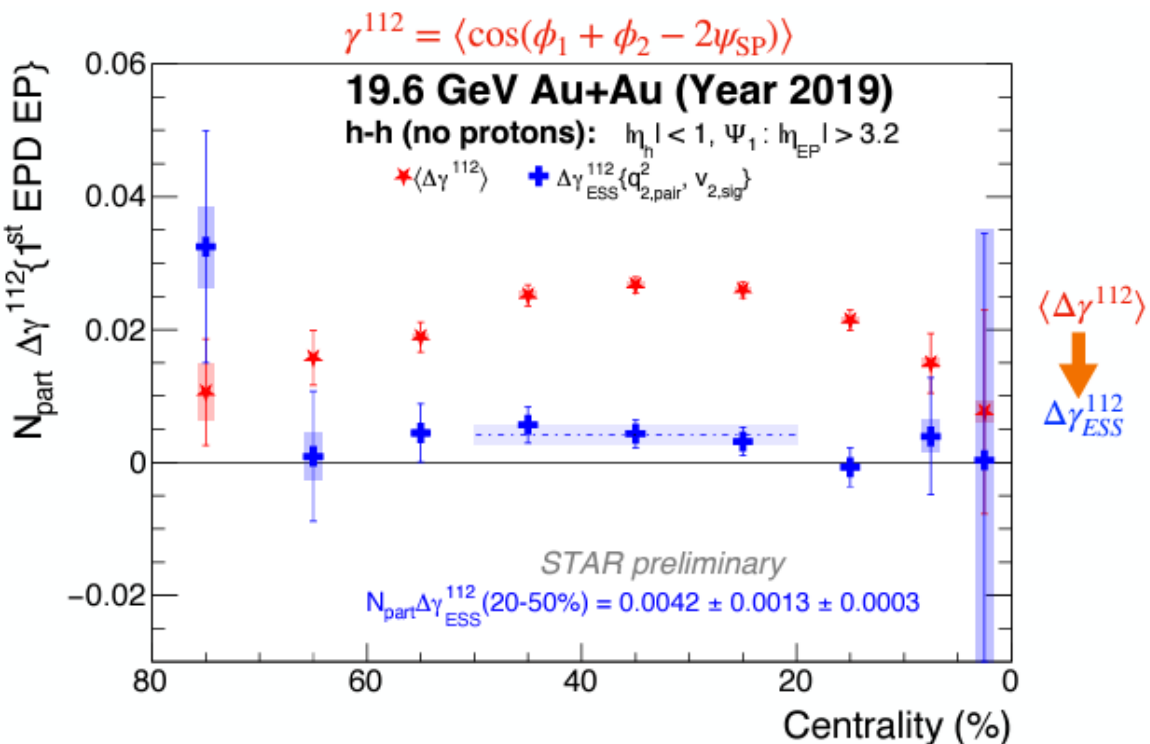


- ESS using POI allows much shorter extrapolation to zero v_2 .
- The ordering of y -intercepts follows predictions from both AVFD and AMPT.
- The y -intercept requires a small conversion to restore the unbiased signal:
$$\Delta\gamma_{ESS}^{112} = \text{Intercept} \times (1 - v_2)^2$$

Z. Xu *et al.*, Phys. Rev. C 107, L061902

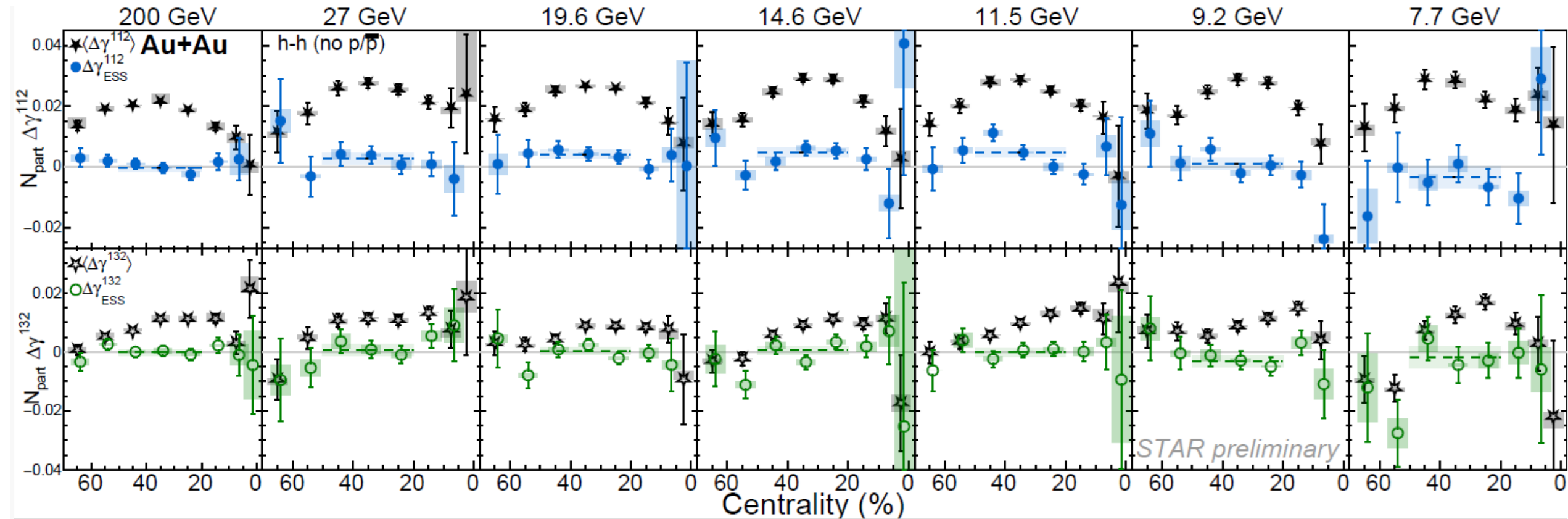


Au+Au at 19.6 GeV



- After v_2 -BKG subtraction, a finite signal in mid-central (20-50%) events.
- Ratio from the optimal ESS (c), pair q_2 and single v_2 , yields a 3σ significance in the 20-50% centrality.
- From the BKG indicator $\Delta\gamma^{132}$, ESS successfully suppresses v_2 -BKG.

Au+Au at 7.7 -- 200 GeV

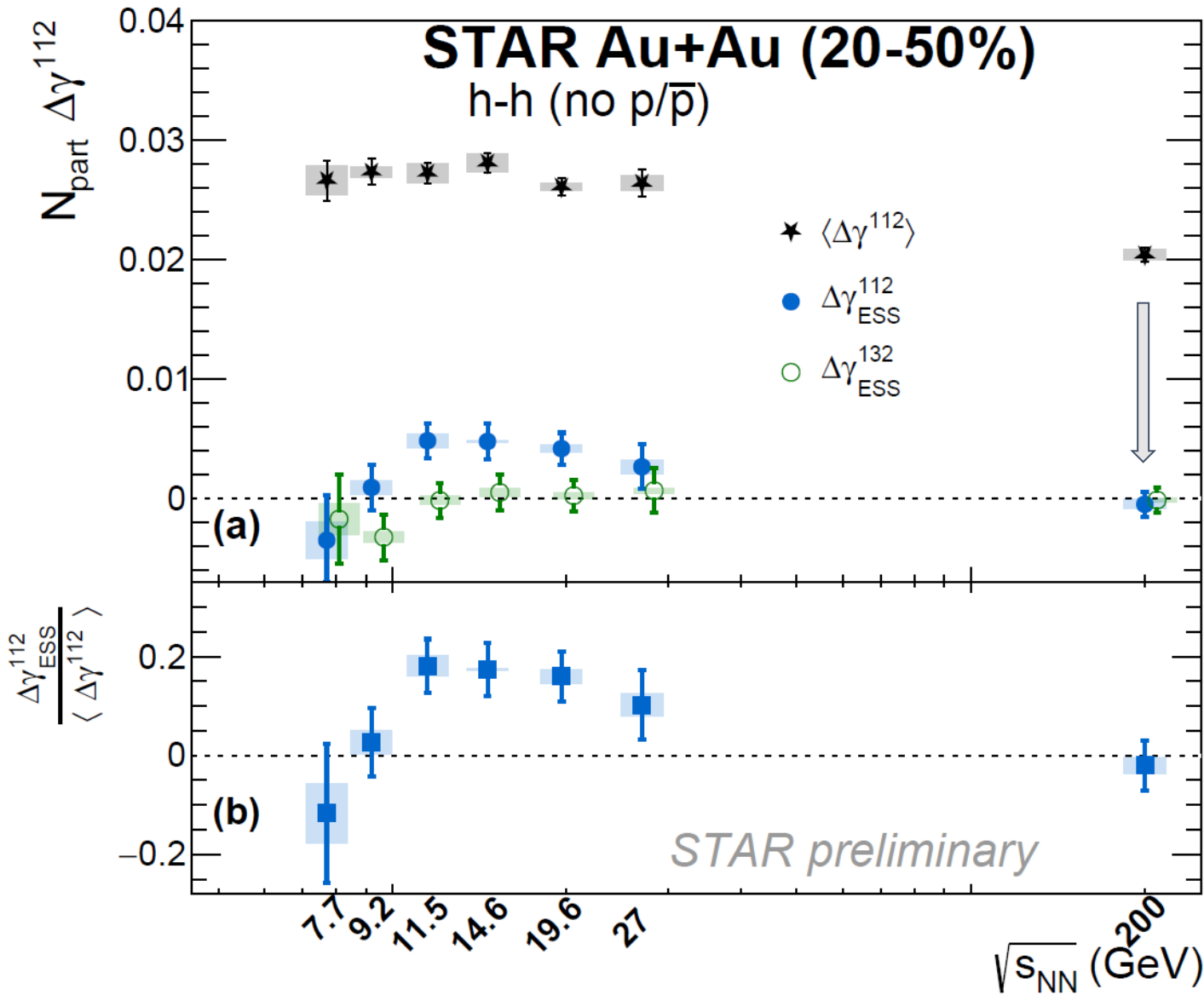


$\Delta\gamma^{112}_{\text{ESS}}$ from the optimal ESS (c), pair q_2 and single v_2 :

- At 200 GeV, using ZDC-SMD planes, no signal is observed.
- At 19.6, 14.6 and 11.5 GeV, a finite $\Delta\gamma^{112}_{\text{ESS}}$ (3σ significance) in the 20-50% centrality.
- At 9.2 and 7.7 GeV, data favor the zero-CME scenario.

$\Delta\gamma^{132}_{\text{ESS}}$ is consistent with zero.

Beam Energy Dependence



- $\Delta\gamma_{\text{ESS}}^{132}$ consistent with zero.
- At least 80% of the measured $\Delta\gamma^{112}$ comes from BKG.
- At 200 GeV,
 - ratio is $(-2 \pm 5.1 \pm 1.6)\%$
 - upper limit of $f_{\text{CME}} \sim 10\%$ in Au+Au
 - upper limit of $f_{\text{CME}} \sim 5\%$ in **isobars** using participant planes: 0.7% difference, too small to detect!
- If we combine three points at 19.6, 14.6 and 11.5 GeV, the literal average of the ESS results reaches an over 5σ significance (assuming similar physics conditions between 10 and 20 GeV).
- The ESS results approach zero around 9.2 and 7.7 GeV.

STAR ESS CME Search Summary

- The novel Event Shape Selection effectively suppresses flow-related backgrounds.
 - At 200 GeV, upper limit of $f_{\text{CME}} \sim 10\%$.
 - At each of 11.5, 14.6 and 19.6 GeV, a positively finite $\Delta\gamma^{112}_{\text{ESS}} (>3\sigma)$. Over 5σ if combined.
 - Around 7.7 GeV, approaches zero CME limited with large uncertainties.
- More theoretical insights are needed:
 - The remaining B field effect too weak at 200 GeV?
 - Chiral symmetry breaking around 7.7 GeV?
 - The chance of the CME occurrence is enhanced near the critical point?

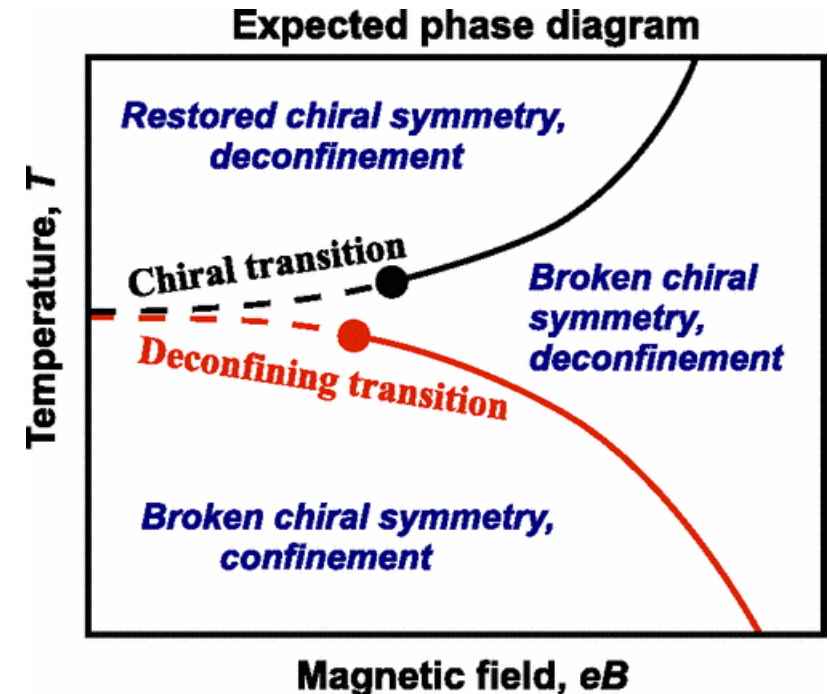
$$\Delta\gamma^{112} = \Delta\gamma^{\text{CME}} + k \frac{v_2}{N} + \Delta\gamma^{\text{non-flow}}$$

Measured Signal Backgrounds

Event Shape Selection Spectator Ψ_1

Note: In the original image, the terms $k \frac{v_2}{N}$ and $\Delta\gamma^{\text{non-flow}}$ are crossed out with red lines, and arrows point from the labels 'Event Shape Selection' and 'Spectator Ψ_1 ' to these terms respectively.

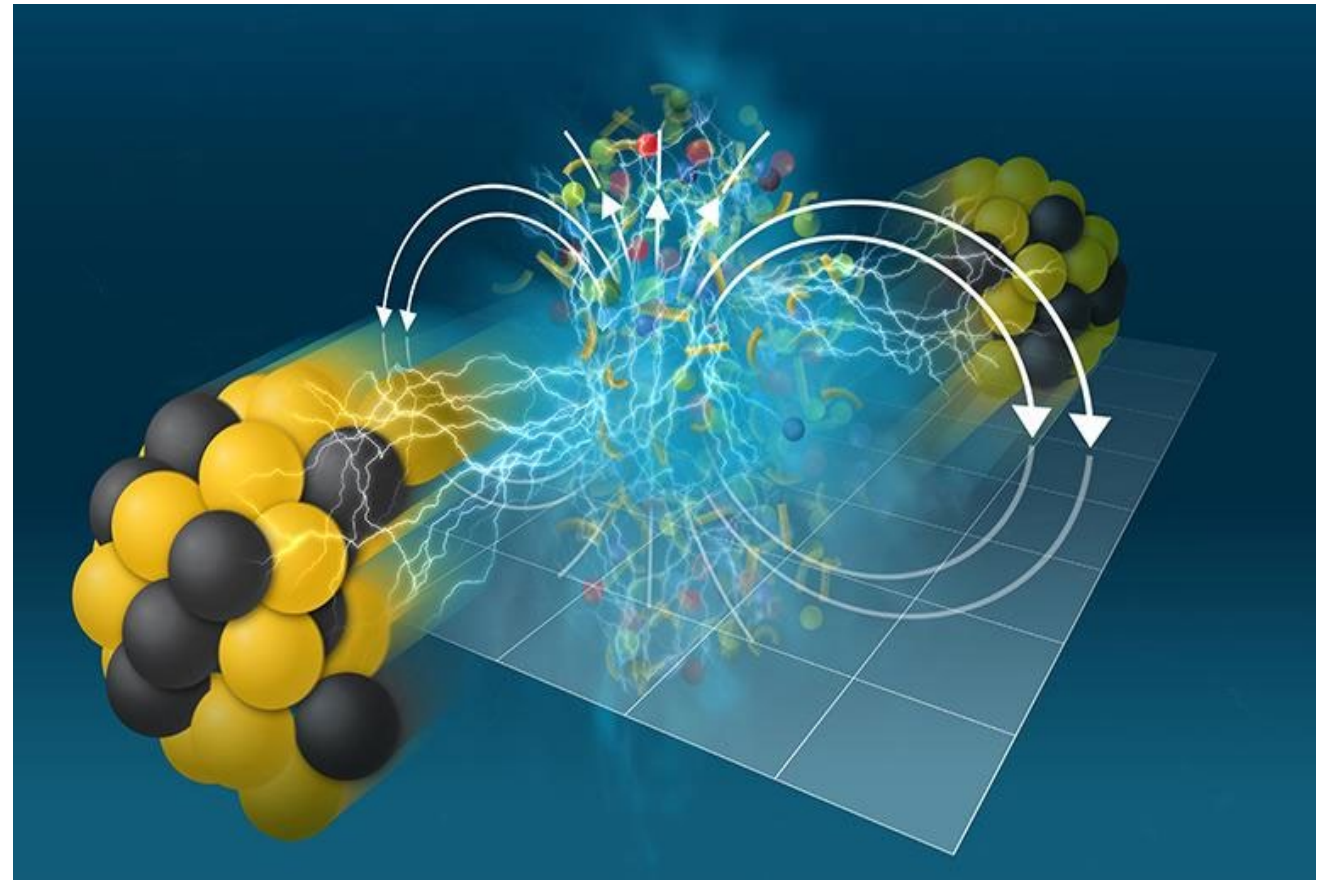
A. J. Mizher, M. N. Chernodub, and E. S. Fraga, PRD 82 (2010) 105016



Super Strong Magnetic fields' Imprint

Analysis of electrical charge dependent deflections in quark-gluon plasma by the STAR Collaboration at the BNL Relativistic Heavy Ion Collider(RHIC)

- Data confirm that super strong magnetic fields ($\sim 10^{18}$ Gauss) generated in off-center collisions could induce an electric current in the quark-gluon plasma
- The findings offer a measure that could relate to the electrical conductivity of the quark-gluon plasma to learn about nature's fundamental building blocks



Observation of the Electromagnetic Field Effect via Charge-Dependent Directed Flow in Heavy-Ion Collisions at the Relativistic Heavy Ion Collider, [Phys. Rev. X 14 \(2024\) 011028](#)

Future Prospect

Impact of Model Dependence on Event Shape Approaches

All event shape methods will have some model dependence –

event shape – observable measured from final state particles in momentum space

shape -- preferably in the coordinate system (initial eccentricity or emission source)

What shape selection most related to CME background contributions

Event-Shape Engineering (ESE) – more sensitive to initial collision eccentricity

Event-Shape Selection (ESS) – Sensitive to combination of eccentricity and particle emission pattern

For preferred mid-centrality for CME searches (20-50% for example)

ESE – limited range of eccentricity variation --- cannot reach the v_2 approach zero round shape to minimize CME bkgd

-- extrapolation to v_2 zero limit – model dependent

-- if the extrapolation follows the eccentricity variation, then initial eccentricity zero corresponds to the most central collisions – small B field and no CME!

ESS – with limited range of eccentricity the approach to v_2 zero is mostly due to emission pattern fluctuations

What CME background at the v_2 approaches zero limit – the intercept point

Depends on shape observable versus v_2 control method

For hydro-induced background, the optimized approach $q_2(\text{pair})$ vs $v_2(\text{single})$

What Dynamics at RHIC 200 GeV and LHC

With ESS method we found the $\Delta\gamma^{112}_{\text{ESS}}$ close to ZERO in Au+Au 200 GeV !!
Expect $\Delta\gamma^{112}_{\text{ESS}}$ to be small or near ZERO at the LHC energy ?!

The magnetic field B magnitude at these energies are certainly
larger at the initial collision $t = 0$!!

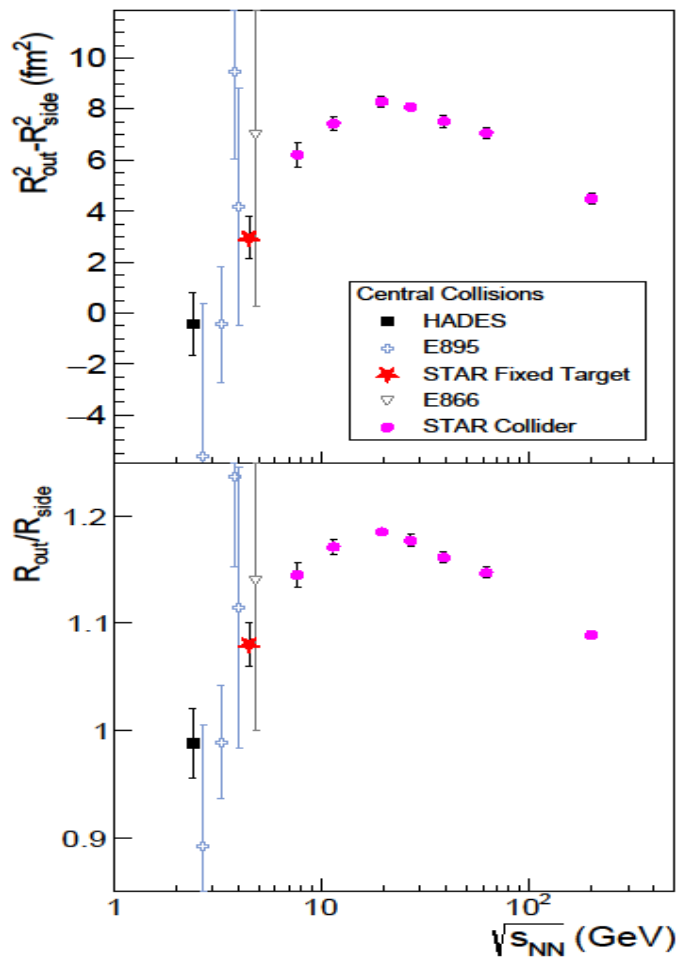
Why? (Professor Pengfei Zhuang gave an answer during his talk--
B field effect disappears at high Temperatures !)

Please measure $\Delta\gamma^{112}_{\text{ESS}}$ at the LHC energy !

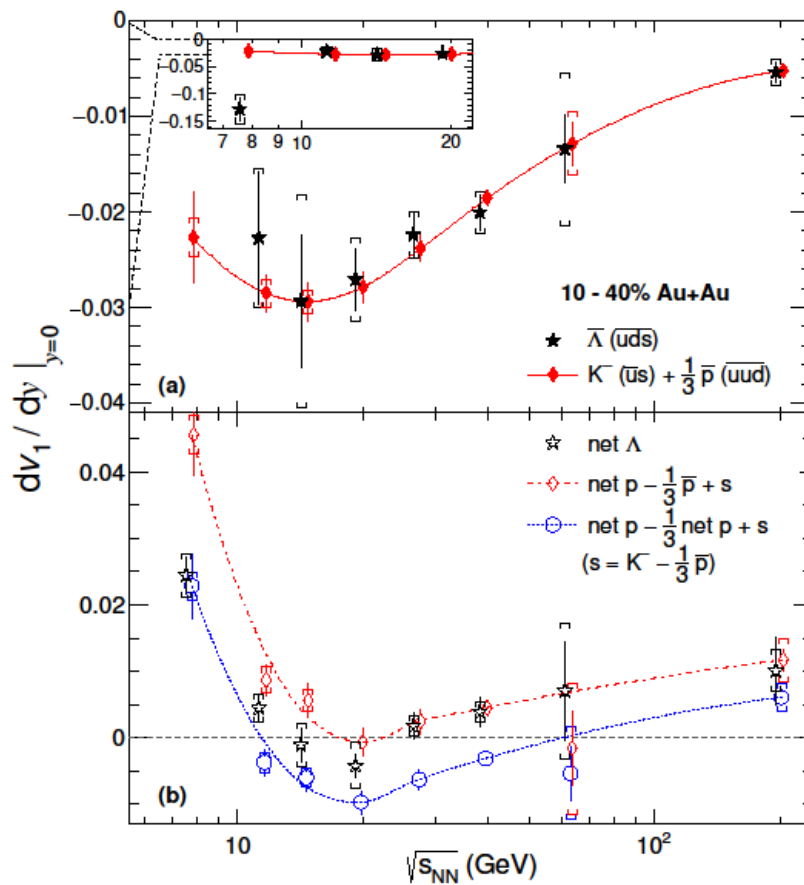
Please measure $v_2\delta$ background correlation as well !

What so Special for Collisions at 10-30 GeV

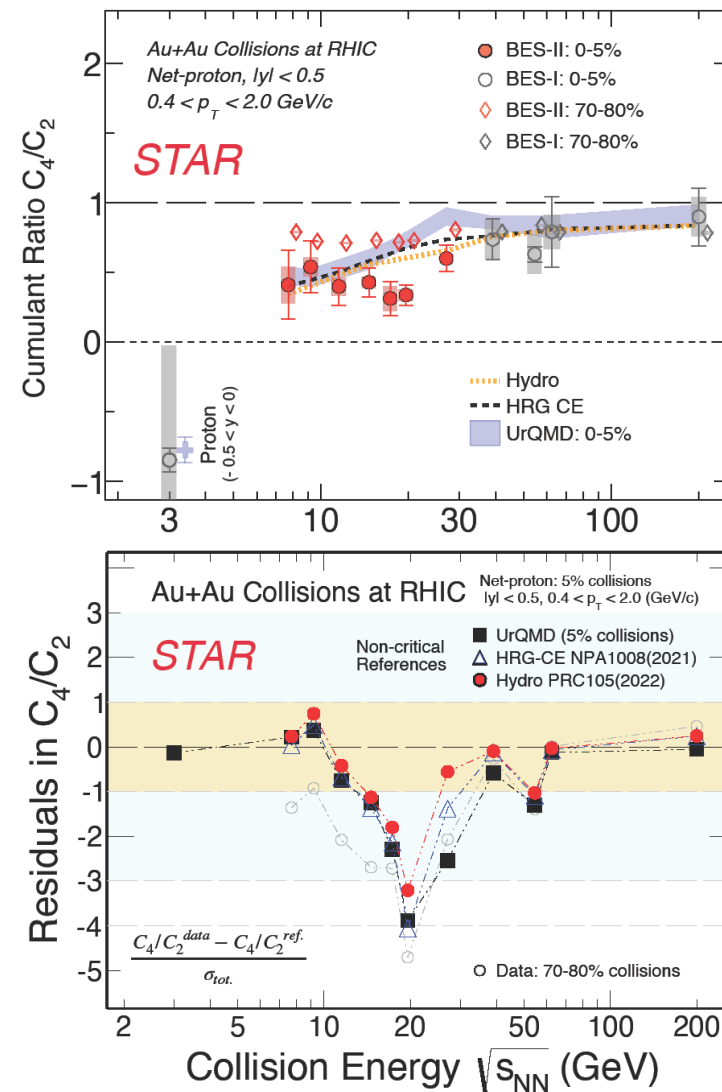
HBT Rout/Rside



v_1 slope dv_1/dy



Critical Point: C_4/C_2



Future of Experimental CME Searches

Improve understanding background contributions !

Improve CME search approach !

We improved ESS approach and we are open to more optimizations

Understand magnetic field effect !

Theoretical insights !

Be Critical, but also Be Truthful !

2407.14489v1

The goal of the ESE was to approach $v_2 = 0$ limit, it is clear that the ESE method has a problem here !

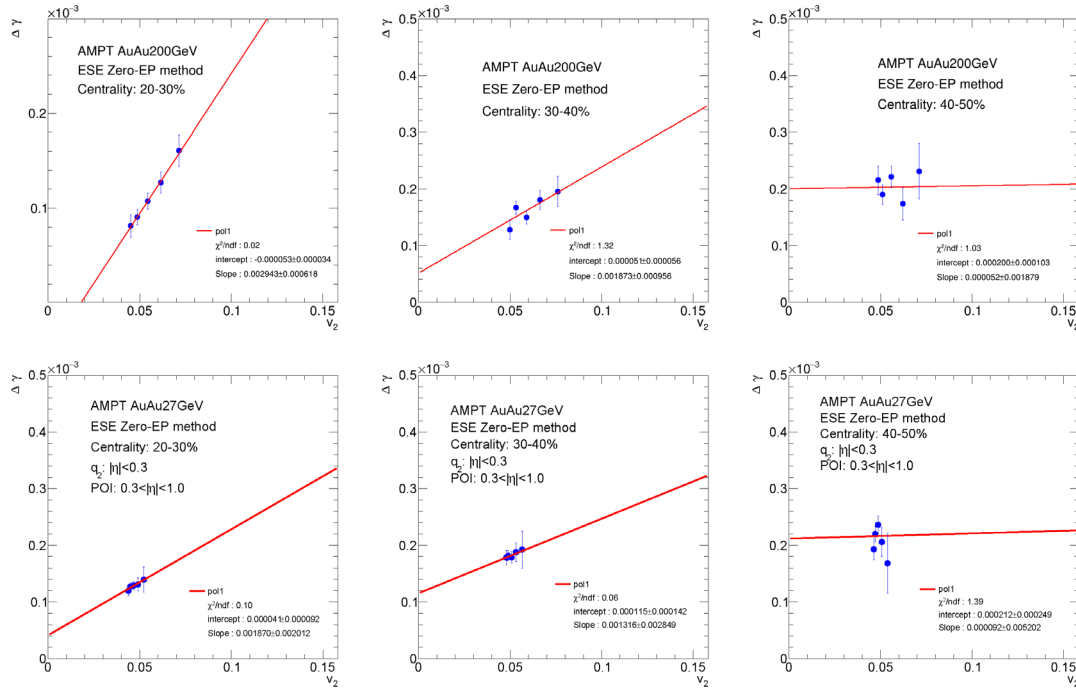


FIG. A.4. AMPT ESE results. Shown are three centralities of Au+Au collisions at $\sqrt{s_{NN}} = 200$ GeV (upper panels) and at 27 GeV (lower panels) simulated by AMPT, with approximately 5×10^6 events for each centrality at each energy. The $\Delta\gamma$ is plotted as a function of v_2 in events binned in $q_2^2\{2\}$ (Eqs. 5,6). POIs are from acceptance $0.3 < |\eta| < 1$, and the event selection variable $q_2^2\{2\}$ is computed from particles in $|\eta| < 0.3$, both with $0.2 < p_T < 2$ GeV/c. The model's known impact parameter direction $\psi = 0$ is taken as the EP in calculating $\Delta\gamma$ (Eqs. 2,3) and $\langle v_2 \rangle$ (Eq. 10).

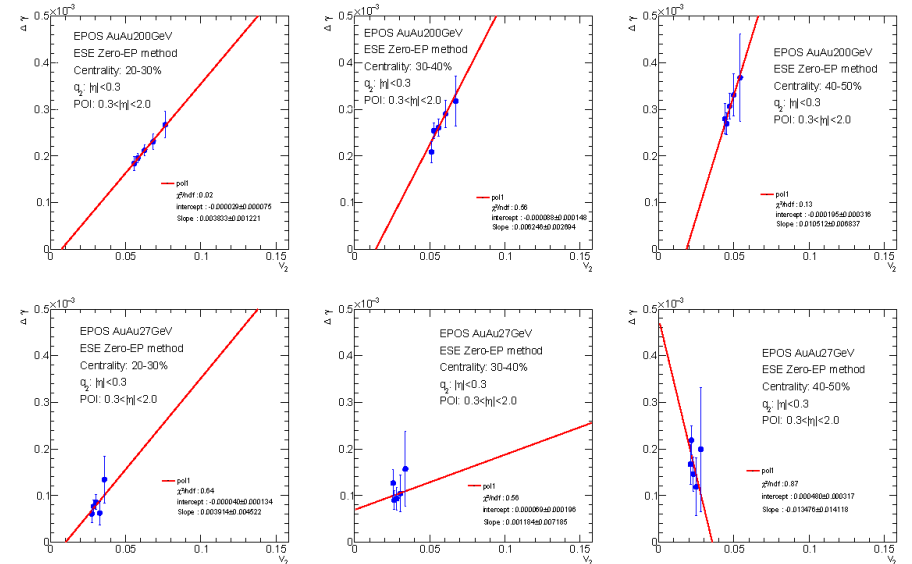


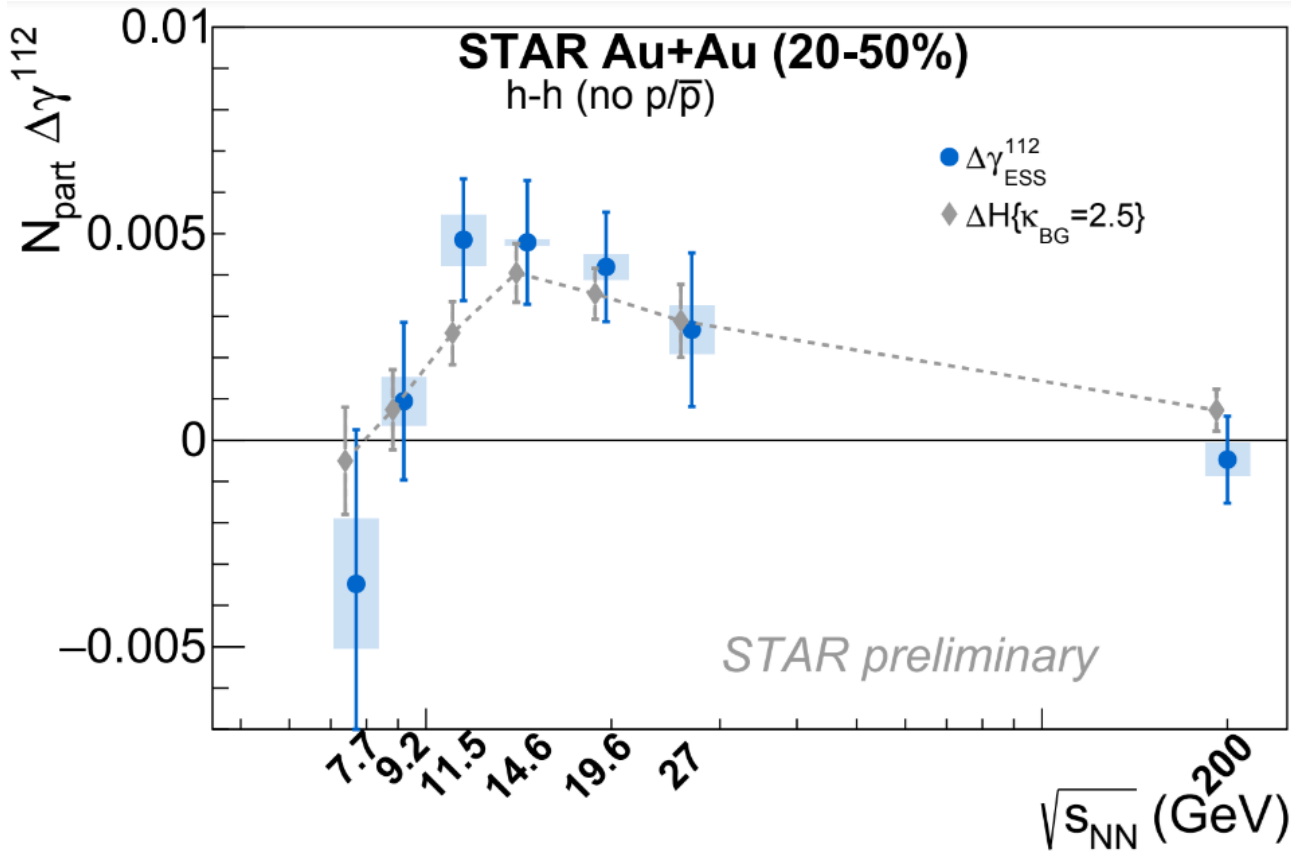
FIG. A.6. EPOS ESE results. Shown are three centralities of Au+Au collisions at $\sqrt{s_{NN}} = 200$ GeV (upper panels) and at 27 GeV (lower panels) simulated by EPOS4, with approximately 1.6×10^6 and 8×10^5 events for each centrality, respectively. The $\Delta\gamma$ is plotted as a function of v_2 in events binned in $q_2^2\{2\}$ (Eqs. 5,6). POIs are from acceptance $0.3 < |\eta| < 2$ and $0.2 < p_T < 2$ GeV/c, and the event selection variable $q_2^2\{2\}$ is computed from particles in $|\eta| < 0.3$, both with $0.2 < p_T < 2$ GeV/c. The model's known impact parameter direction $\psi = 0$ is taken as the EP in calculating $\Delta\gamma$ (Eqs. 2,3) and $\langle v_2 \rangle$ (Eq. 10).

Event Shape Analysis cannot solve all our physics problems – need to find the best approach for your particular physics

- 1) **Some toy models are indeed just toys, avoid playing “garbage in, garbage out” game !**
- 2) **Respect statistics: when you get 1+-1, result is consistent with zero, but is also consistent with many other scenarios -- it does not mean your method working, it could also mean that your method does not have sensitivity**

Thank You !

Connection between ESS and the H correlator



- In dealing with the BES-I data, we introduced the H correlator to subtract the flow BKG:

$$H(\kappa_{bg}) \equiv (\kappa_{bg} v_2 \delta - \gamma^{112}) / (1 + \kappa_{bg} v_2)$$

$$\Delta \bar{H} \equiv H_{\text{SS}} - H_{\text{OS}} \quad \delta = \cos(\phi_1 - \phi_2)$$

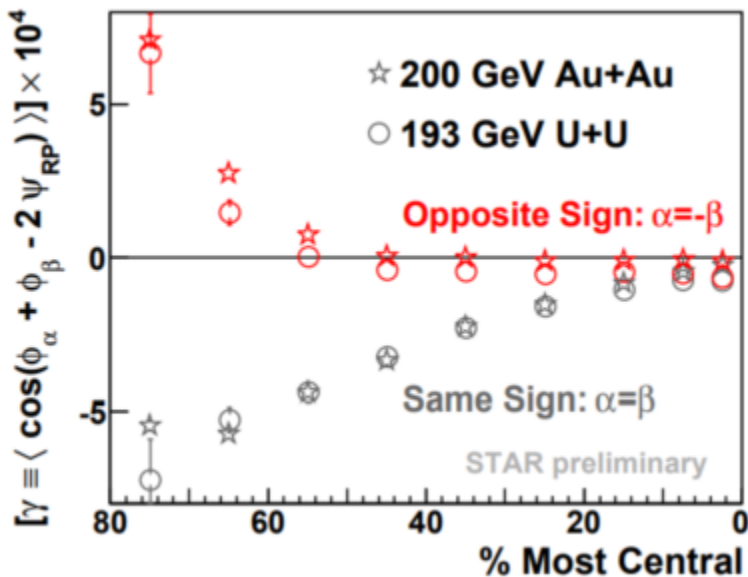
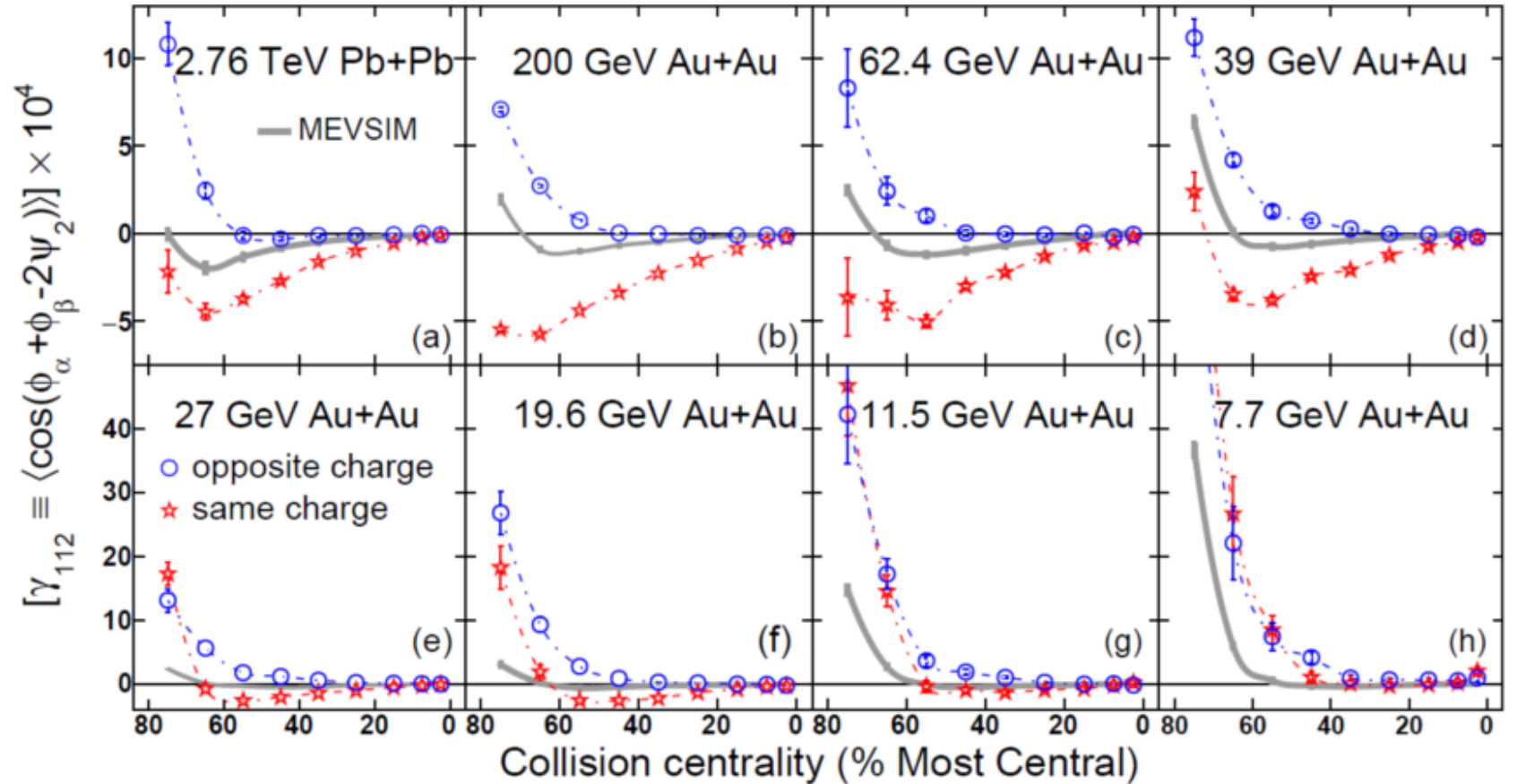
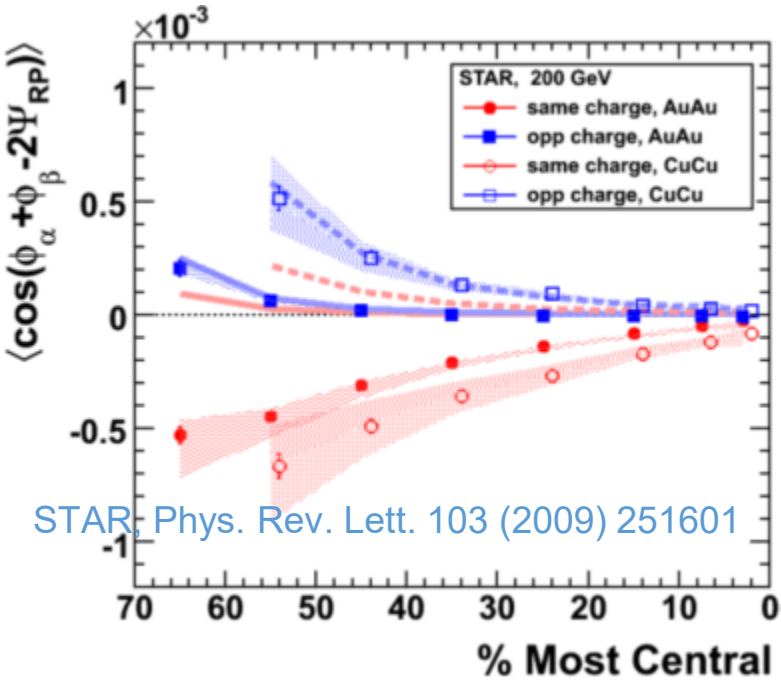
$$\gamma = \kappa v_2 \mathbf{B} - \mathbf{H}$$

$$\delta = \mathbf{B} + \mathbf{H}$$

- κ_{bg} is an adjustable parameter, unknown a priori. It quantifies the coupling between elliptic flow and other mechanisms manifested in the two-particle correlation.

- With κ_{bg} set to 2.5, ΔH agrees with the ESS result at all beam energies under study.
- The flow background can be reasonably well described by a universal coupling between v_2 and the two-particle correlation.

Initial Evidence

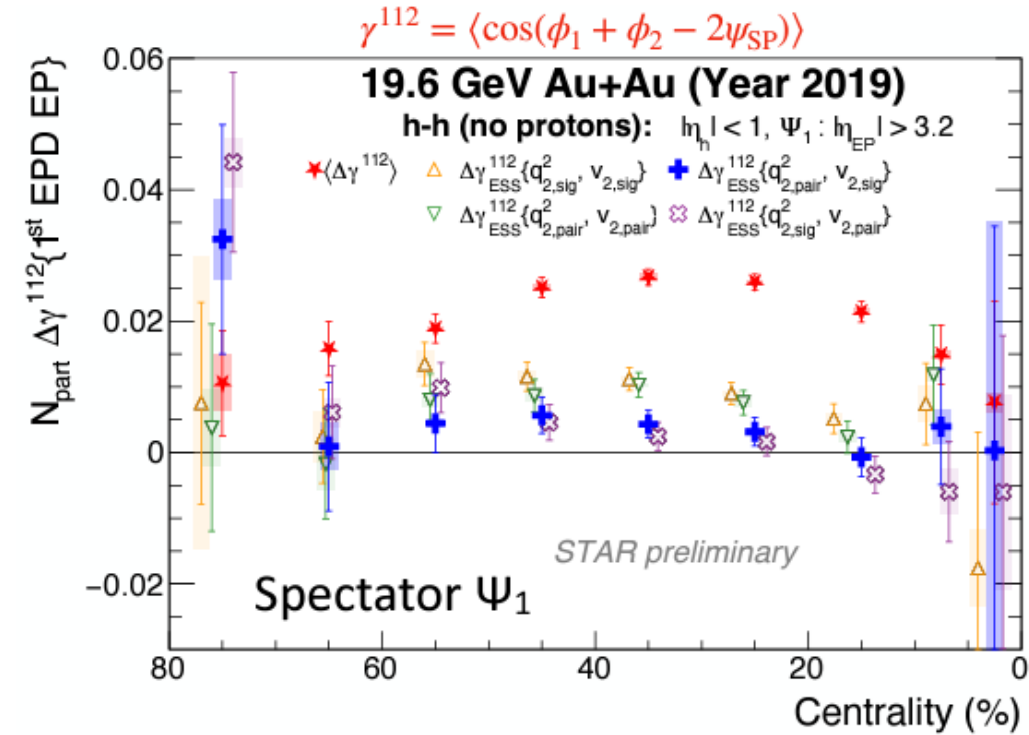


ALICE, Phys. Rev. Lett. 110(2013)012301. STAR, Phys. Rev. Lett. 113(2014)52302

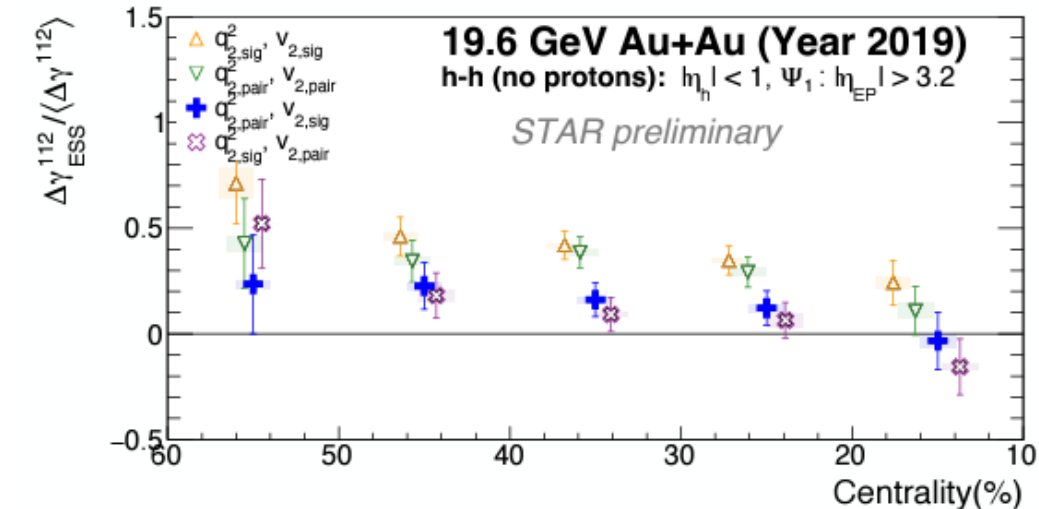
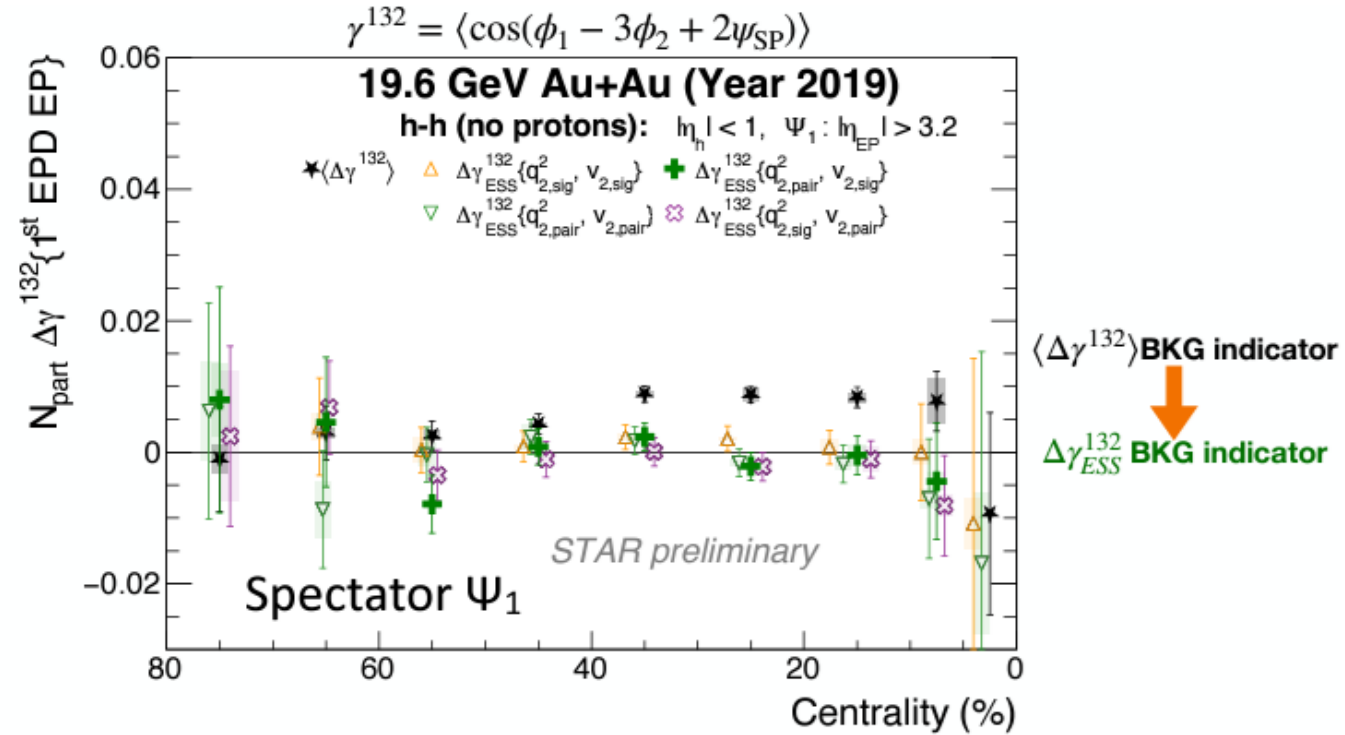
In various collision systems and at different beam energies, positively finite $\Delta\gamma_{112}$ meets the CME expectation, but could contain contributions from:

- Flow-related background $\propto v_2$ (elliptic flow)
- Nonflow-related background (di-jets)

Au+Au at 19.6 GeV



$\langle \Delta\gamma^{112} \rangle$
↓
 $\Delta\gamma_{ESS}^{112}$



- The ordering of γ -intercepts follows predictions from both AVFD and AMPT.

Not all event shape selections are equal, there is some model dependence
We need to optimize the method to suppress the hydro-related CME background

Also event shape selection optimized for CME search only, is not universally best!
Approach for hydro comparisons, for example, the ESE method would be better!