

北京航空航天大学
BEIHANG UNIVERSITY

Probing first-order electroweak phase transitions at the LHC and future colliders

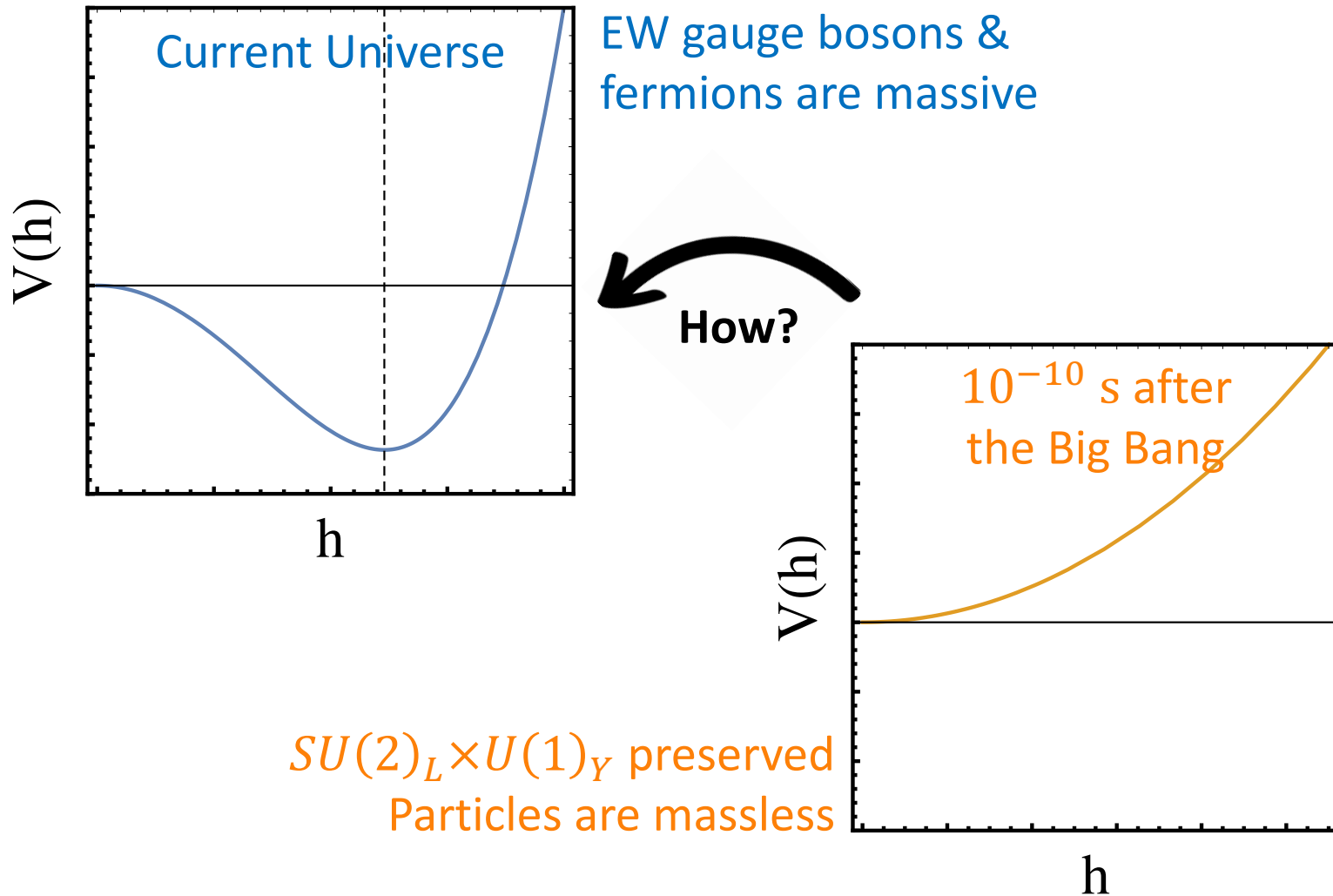
Ke-Pan Xie (谢柯盼)

Beihang University

2024.11.14 @ Qingdao CLHCP2024

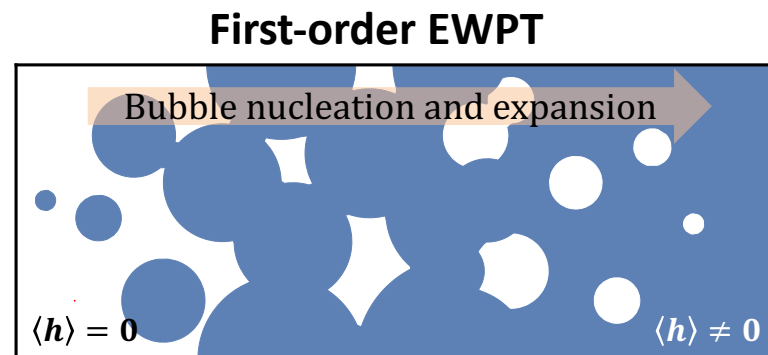
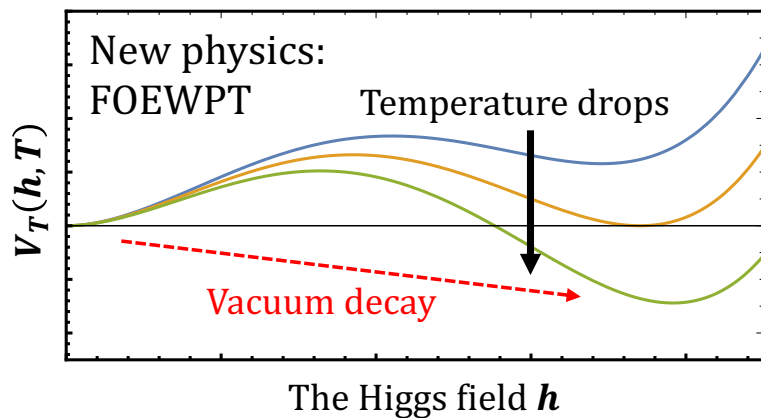
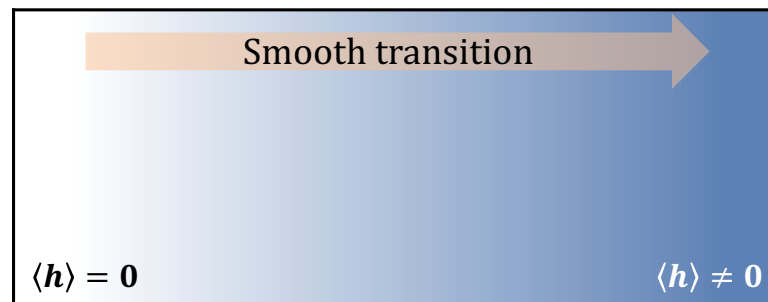
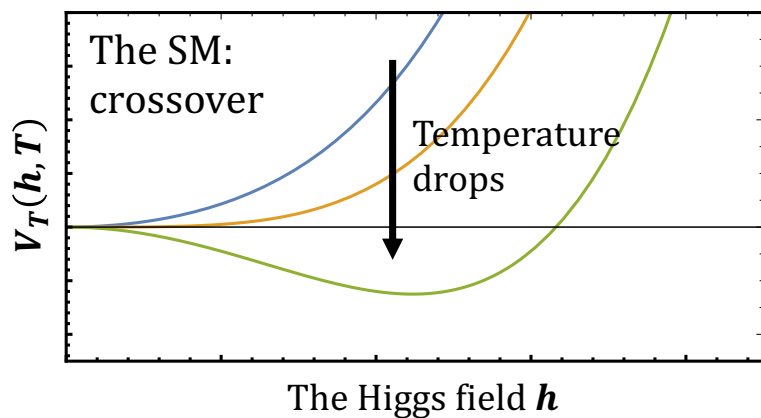
The electroweak phase transition

The evolution of Higgs potential & vacuum expectation value



The feature of electroweak phase transition

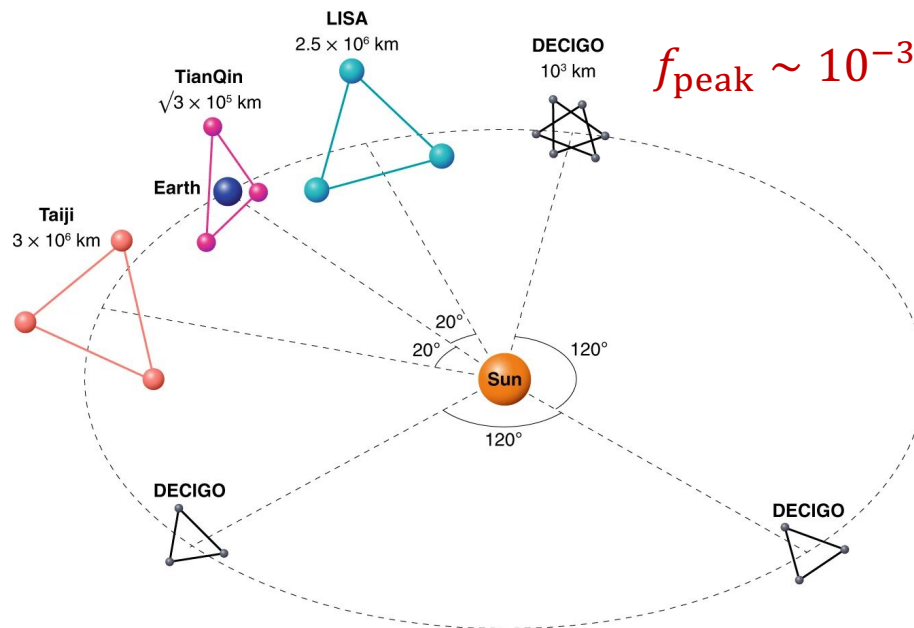
Which pattern?



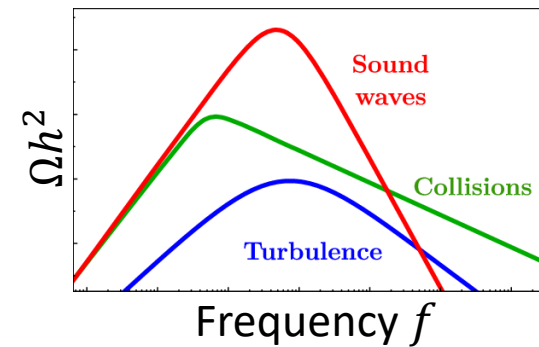
New physics is needed for FOEWPT

3 main reasons for studying FOEWPT

- **If** it has occurred, we may detect its stochastic *gravitational waves* in the next decade [Caprini et al, JCAP 04 \(2016\) 001; JCAP 03 \(2020\) 024; Athron, Wu et al, Prog.Part.Nucl.Phys. 135 \(2024\) 104094](#)



$$f_{\text{peak}} \sim 10^{-3} \text{ Hz} \times \left(\frac{1}{v_w} \right) \left(\frac{\beta/H_*}{100} \right) \left(\frac{T_*}{100 \text{ GeV}} \right)$$



- It is the *consequence* of many BSM models
- It is the *essential* condition of many BSM mechanisms solving the puzzles in the SM

Why we like FOEWPTs

Very **classic** motivation: explaining the matter-antimatter asymmetry via **EW baryogenesis**: since 1980s, still active

Joyce *et al*, PRL 75 (1995) 1695-1698; PRD 53 (1996) 2958-2980;
 Ramsey-Musolf *et al*, New J.Phys. 14 (2012) 125003 [review]
 Cline *et al*, PRD 101 (2020) 6, 063525; **KPX**, JHEP 02 (2021) 090;
KPX, Bian, Wu, JHEP 12 (2020) 047; ...

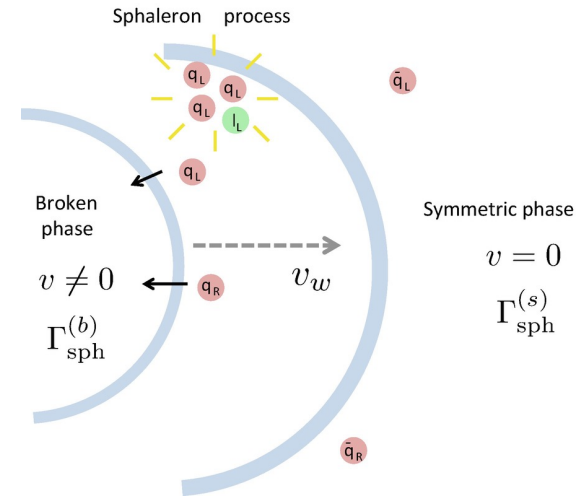


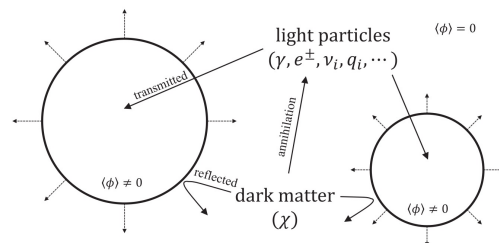
Figure from Fuyuto, PhD thesis (2016)

Recent progress

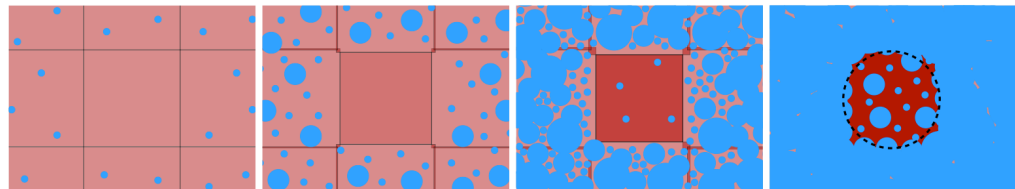
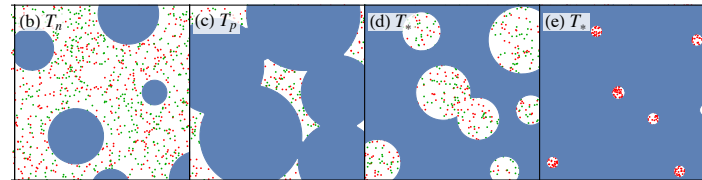
Novel dark matter and/or baryogenesis scenarios based on FOPTs

Filtered dark matter

Baker *et al*, PRL 125 (2020) 15, 151102
 Chway *et al*, PRD 101 (2020) 9, 095019



Solitons Hong, Jung, **KPX**, PRD 102 (2020) 7, 075028



Primordial black holes

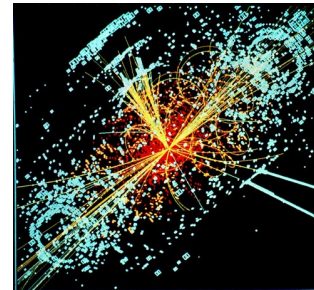
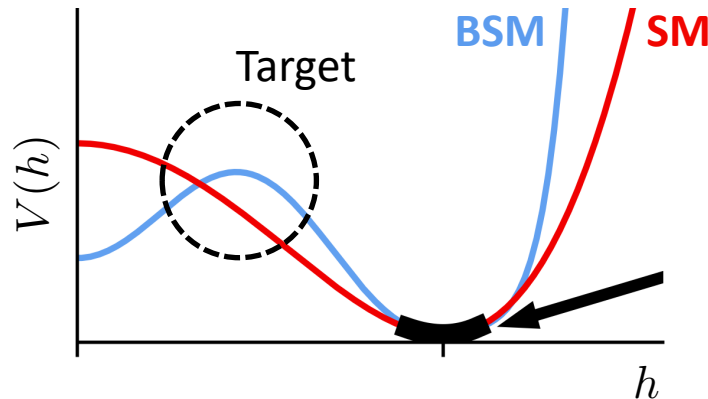
Liu *et al*, PRD 105 (2022) 2, L021303; Kanemura, Tanaka, **KPX**, JHEP 06 (2024) 036;
 Cai *et al*, SCPMA 67 (2024) 9, 290411

Probing FOEWPT at colliders

Searching for the underlying physics of the **potential barrier**

In general, this is

- *Challenging*, attempting to probe the global feature of the potential via phenomenology around vacuum
- *Model-dependent*, various signal channels



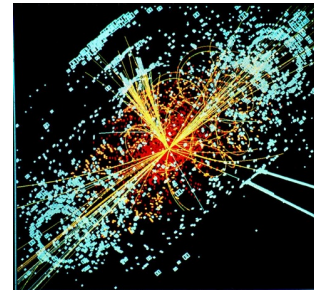
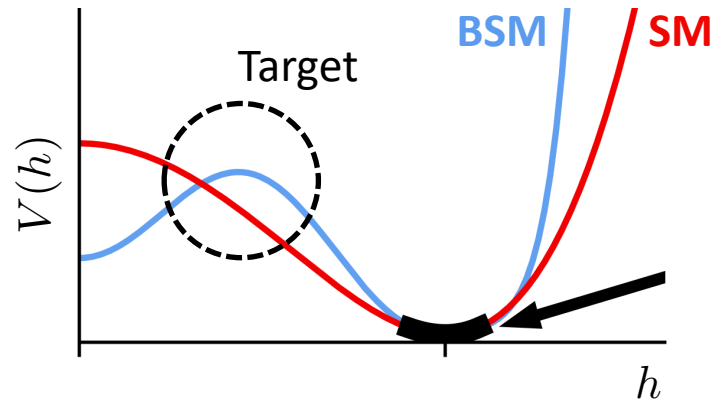
LHC → HL-LHC
HE-LHC? LHeC?
CEPC? ILC? FCC?
Muon collider?

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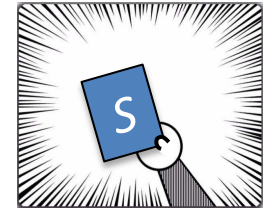
However, it also has

- General and representative features, can be systematically summarized and classified
- Complementarity and crosscheck with the GW astronomy

A prototype model

Scalar sector: Higgs (h) + scalar (s)

- **Prototype** of many well-motivated models



Singlet-extended SM [Cline et al, JCAP 01 \(2013\) 012](#); [Alanne et al, NPB 889 \(2014\) 692](#); [Chiang et al, PLB 789 \(2019\) 154](#); [Jiang et al, PRD 93 \(2016\) 6, 065032](#); [Alves et al, JHEP 04 \(2019\) 052, JHEP 12 \(2018\) 070, JHEP 03 \(2020\) 053, PLB 818 \(2021\) 136377](#); [Carena et al, JHEP 08 \(2020\) 107](#); [Liu, KPX, JHEP 04 \(2021\) 015](#); [Huang, KPX, PRD 105 \(2022\) 11, 115033](#), [Liu et al, PRD 105 \(2022\) 11, 115040](#); etc

2HDM [Cline et al, JHEP 11 \(2011\) 089](#); [Dorsch et al, JHEP 10 \(2013\) 029](#); [Basler et al, JHEP 02 \(2017\) 121](#); [Dorsch et al, JHEP 12 \(2017\) 086](#); [Bian, Jiang et al, JHEP 05 \(2018\) 151](#); [Wang et al, PLB 788 \(2019\) 519](#); [Wang et al, PRD 101 \(2020\) 015015](#); [Su et al, JHEP 04 \(2021\) 219](#); etc

Left-right model [Brdar et al, JCAP 12 \(2019\) 027](#); [Li et al, JHEP 03 \(2021\) 267](#); etc

Georgi-Machacek model [Bian et al, JHEP 01 \(2019\) 216](#); [Chen et al, PRD 106 \(2022\) 5, 055019](#); etc

Supersymmetry [Lee et al, PRD 71 \(2005\) 075010](#); [Balazs et al, PRD 71 \(2005\) 075002](#); [Huang et al, PRD 91 \(2015\) 2, 025006](#); [Bi et al, PRD 92 \(2015\) 023507](#); [Bian, Guo, Shu, CPC 42 \(2018\) 9, 093106](#); [Athron et al, JHEP 11 \(2019\) 151](#); [Wang, KPX, Wu, Yang, EPJC 82 \(2022\) 12, 1120](#); etc

Composite Higgs [Espinosa et al, JCAP 01 \(2012\) 012](#); [Bian, Wu, KPX, JHEP 12 \(2019\) 028, JHEP 12 \(2020\) 047](#); [De Curtis et al, JHEP 12 \(2019\) 149](#); [Angelescu et al, JHEP 10 \(2022\) 019](#); etc

and more...

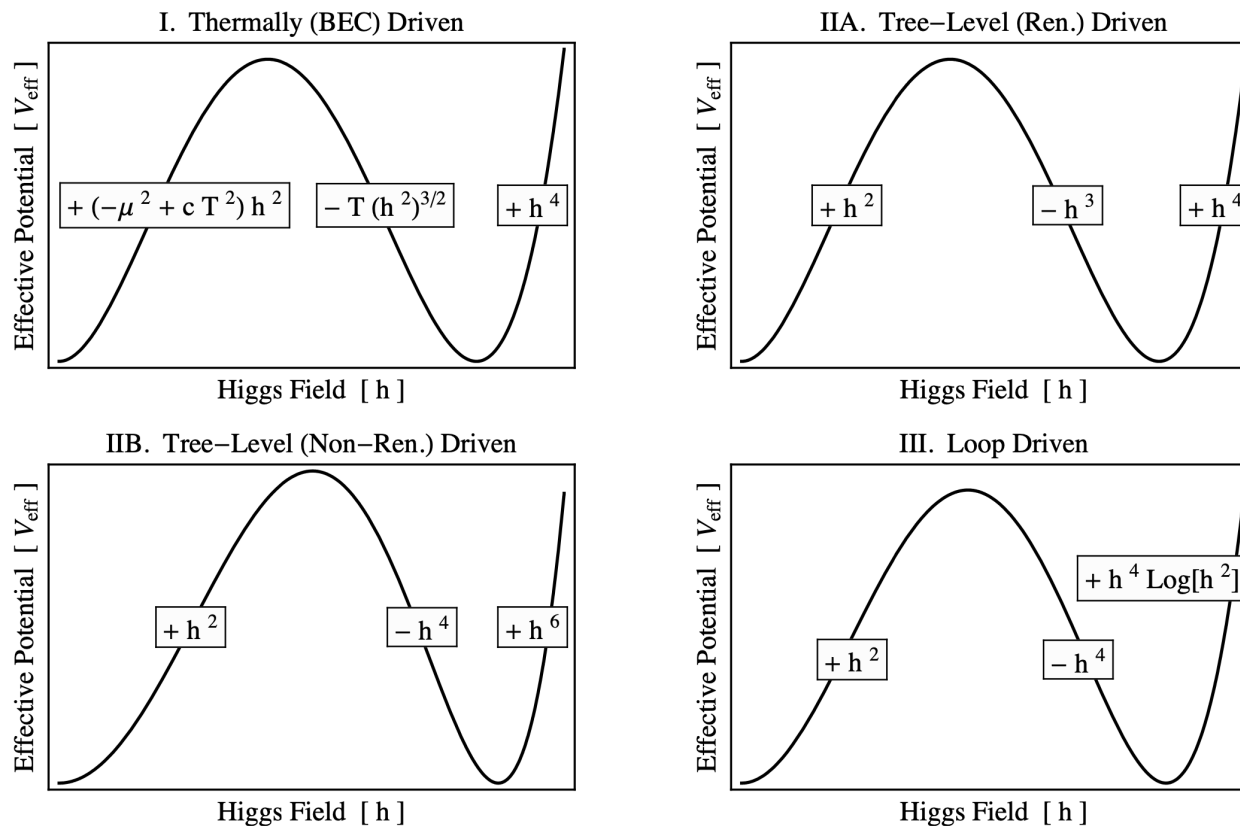
Underlying physics of the barrier

The full one-loop finite-temperature potential [Quiros, hep-ph/9901312]

$$V_T(h, T) = V_0(h) + V_1(h) + V_{1,T}(h, T) + V_{\text{daisy}}(h, T)$$

Tree level 1-loop CW 1-loop thermal Daisy resummation

A seminal classification [Chung, Long, and Wang, PRD 87 (2013) 2, 023509]



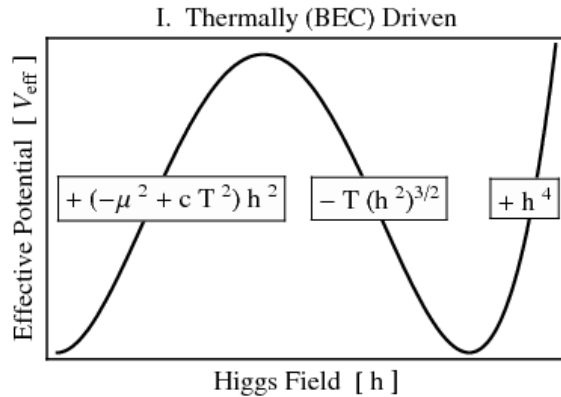
Type I: thermally loop-driven

Relevant $\mathcal{L} \supset -m_s^2 s^2/2 - \lambda_{hs} h^2 s^2/4$

IF $m_s^2 \ll \lambda_{hs} h^2/2 \lesssim T^2$,

$$V_T \approx \frac{-\mu^2 + cT^2}{2} h^2 - \frac{T}{12\pi} \left(\frac{\lambda_{hs}}{2}\right)^{3/2} h^3 + \frac{\lambda}{4} h^4$$

otherwise Boltzmann suppression $e^{-m_s/T}$



A strong FOEWPT requires

$$\frac{v_c}{T_c} \approx \frac{1}{6\pi\lambda} \left(\frac{\lambda_{hs}}{2}\right)^{3/2} \gtrsim 1$$

Usually, $\lambda = m_h^2/2v^2 \approx 0.13$ fixed

\Rightarrow need a sizable λ_{hs}

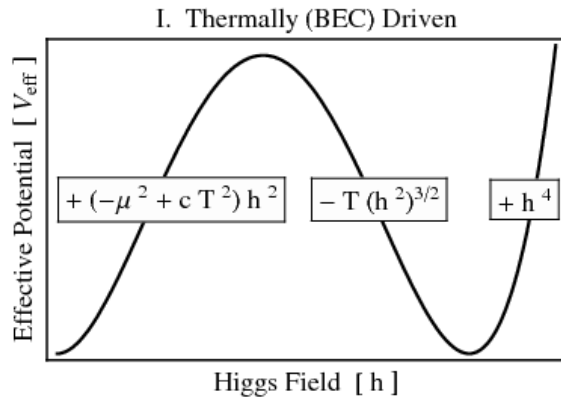
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 \Rightarrow needs a sizable λ_{hs}

Collider signal: a **light** scalar with **significant** coupling to the Higgs

- Candidates: stop (SUSY), h_2 (2HDM), etc
- Severely constrained by $\sigma(gg \rightarrow h)$, $h \rightarrow \gamma\gamma$, etc, but still rooms for model-building [\[Cao et al, JHEP 01 \(2022\) 001\]](#)

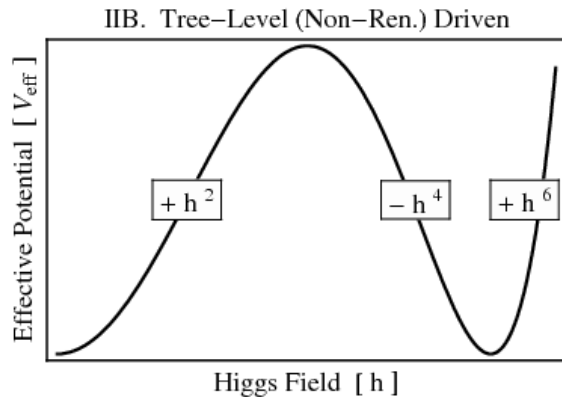
Type IIB: tree-level non-renormalizable operator

Relevant $\mathcal{L} \supset -m_S^2 s^2/2 - \lambda_{hs} h^2 s^2/4$

IF $m_S \gg 100$ GeV, EFT

$$V_T \approx \frac{\mu^2 + cT^2}{2} h^2 - \frac{\lambda}{4} h^4 + \frac{\lambda_{hs}^3}{48m_S^2} h^6$$

Barrier exists at tree-level at $T = 0$ via operator $\mathcal{O}_6 \equiv |H|^6/\Lambda^2$



A strong FOEWPT requires

$$0.55 \text{ TeV} < \Lambda < 0.89 \text{ TeV}$$

[Huang et al, PRD93, 103515 \(2016\)](#)

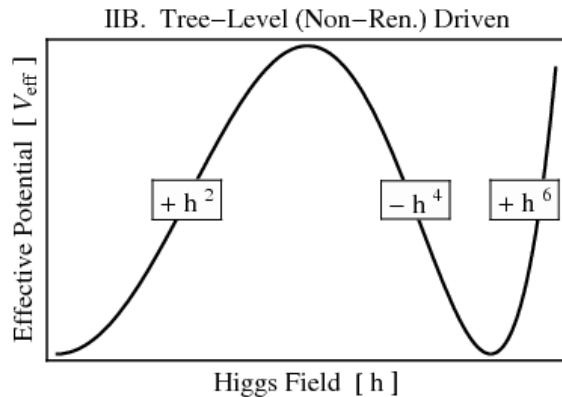
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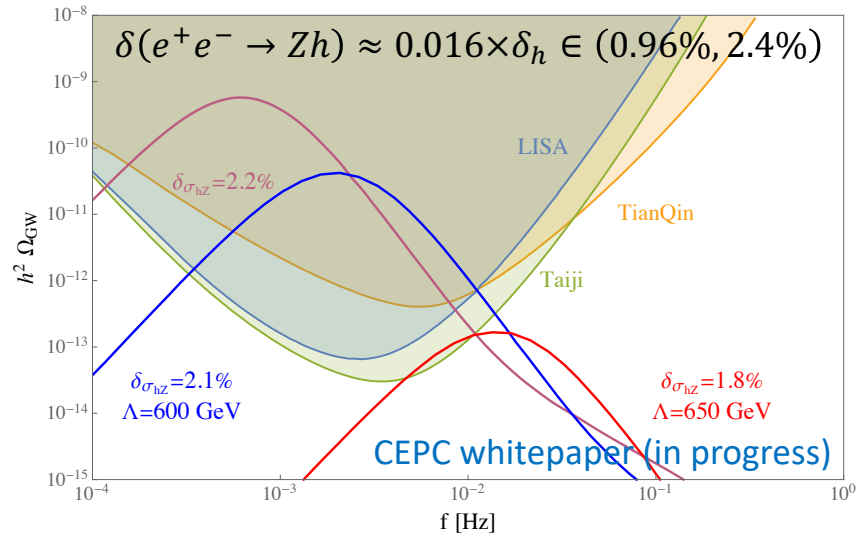
A strong FOEWPT requires

$$0.55 \text{ TeV} < \Lambda < 0.89 \text{ TeV}$$

Huang *et al*, PRD93, 103515 (2016)

Collider signal: h^3 deviation

- $\delta_h = \left| \lambda_{h^3} / \lambda_{h^3}^{\text{SM}} - 1 \right| \approx 0.468 \times (\text{TeV}/\Lambda)^2$



Type IIA: tree-level renormalizable operator

The most general case

$$V = -\frac{\mu^2}{2}h^2 + \frac{\lambda}{4}h^4 + \frac{a_1}{4}h^2s + \frac{a_2}{4}h^2s^2 + \frac{b_2}{2}s^2 + \frac{b_3}{3}s^3 + \frac{b_4}{4}s^4$$

Fixing Higgs mass & VEV, 5 free parameters

$$\begin{pmatrix} h \\ s \end{pmatrix} = \begin{pmatrix} \cos \theta & -\sin \theta \\ \sin \theta & \cos \theta \end{pmatrix} \begin{pmatrix} h_1 \\ h_2 \end{pmatrix} \begin{array}{l} \text{Higgs-like, 125 GeV} \\ \text{Singlet-like} \end{array}$$

Type IIA: tree-level renormalizable operator

The most general case

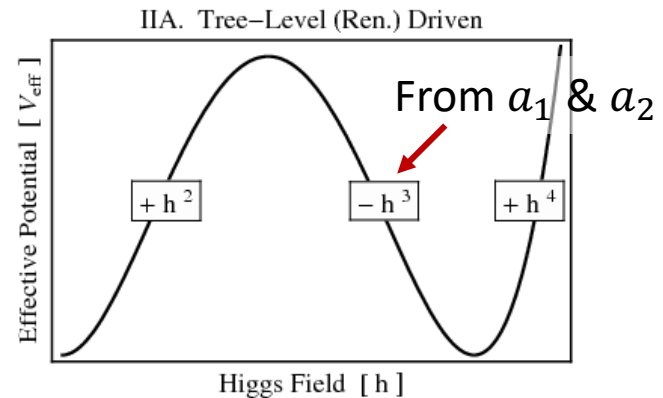
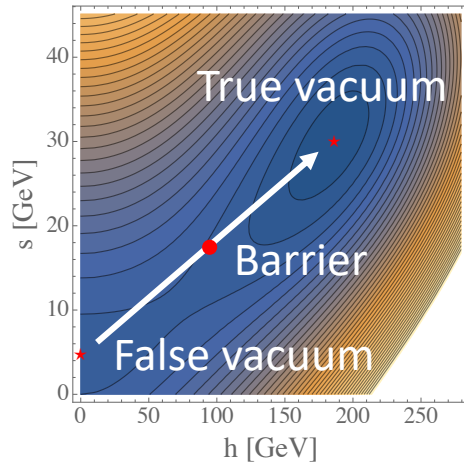
$$V = -\frac{\mu^2}{2}h^2 + \frac{\lambda}{4}h^4 + \frac{a_1}{4}h^2s + \frac{a_2}{4}h^2s^2 + \frac{b_2}{2}s^2 + \frac{b_3}{3}s^3 + \frac{b_4}{4}s^4$$

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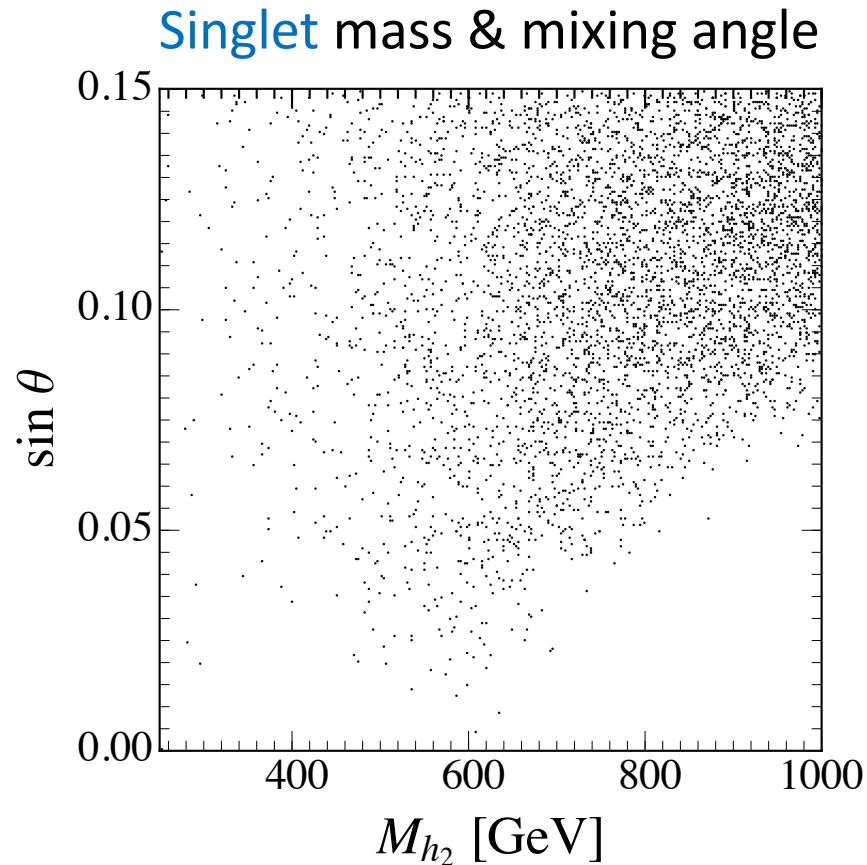
Thermal corrections

$$\Delta V_T = \frac{c_h T^2}{2} h^2 + m_1 T^2 s + \frac{c_s T^2}{2} s^2$$



FOEWPT parameter space

Varying the 5 free input parameters [Liu and KPX, JHEP 04 (2021) 015]

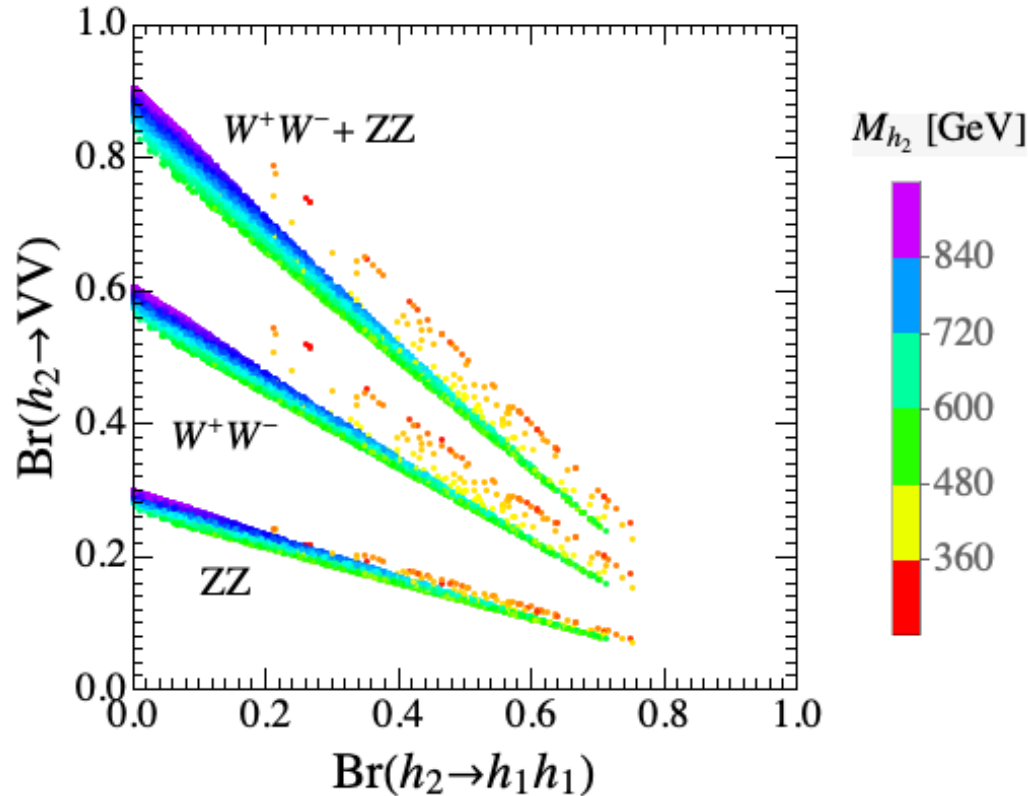


Sizable mixing required for a FOEWPT

Resonant heavy h_2 search

FOEWPT parameter space

- Produced via mixing, e.g. $gg \rightarrow h_2$, cross section $\propto \sin^2 \theta$

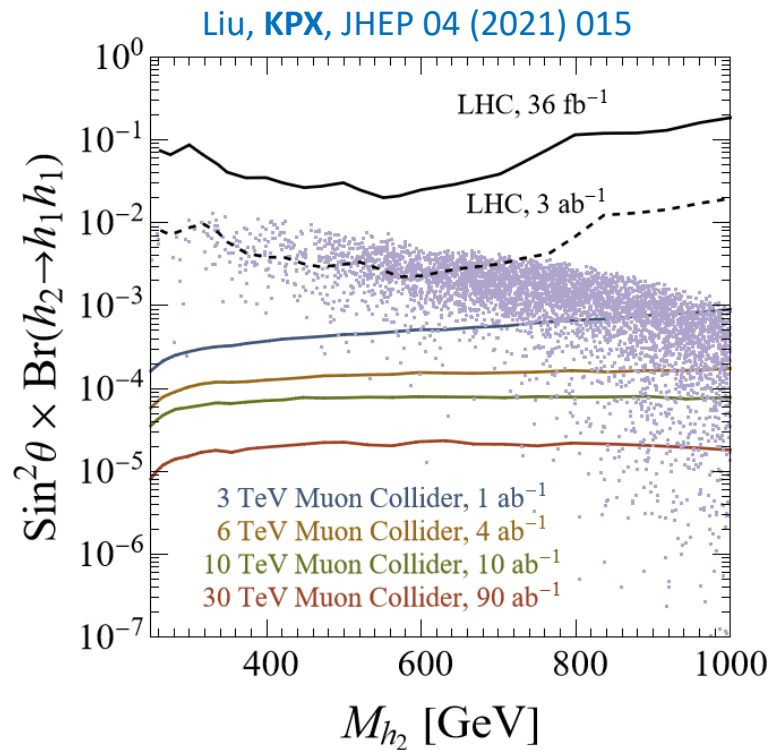
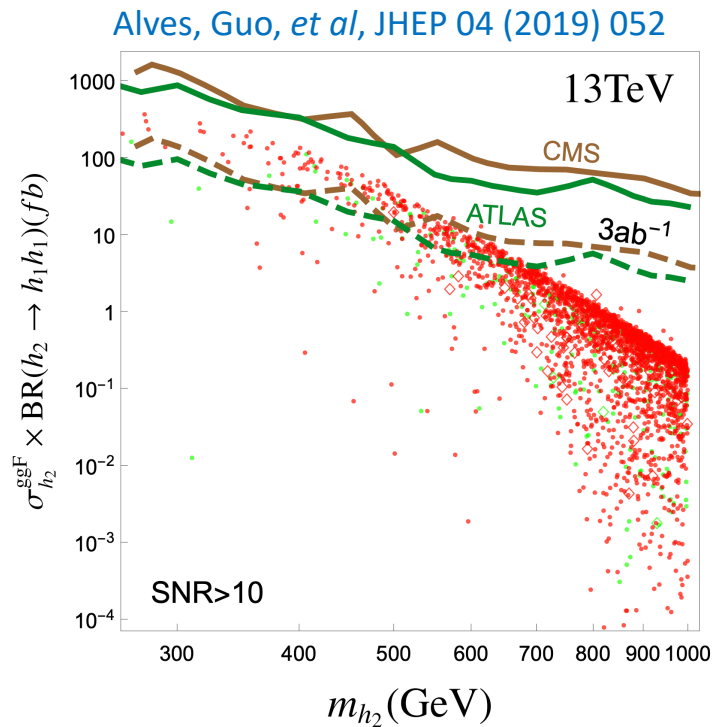


- Dominant decays $h_2 \rightarrow h_1 h_1$ and $h_2 \rightarrow W^+W^-$ or ZZ
- $\text{Br}(h_2 \rightarrow t\bar{t}) \lesssim 20\%$ subdominant

Resonant heavy h_2 search

FOEWPT parameter space

- HL-LHC or multi-TeV muon colliders
- Complementary to gravitational wave search



Recent progress [Biermann *et al*, 2408.08043; Aboudonia *et al*, 2410.22700, etc]

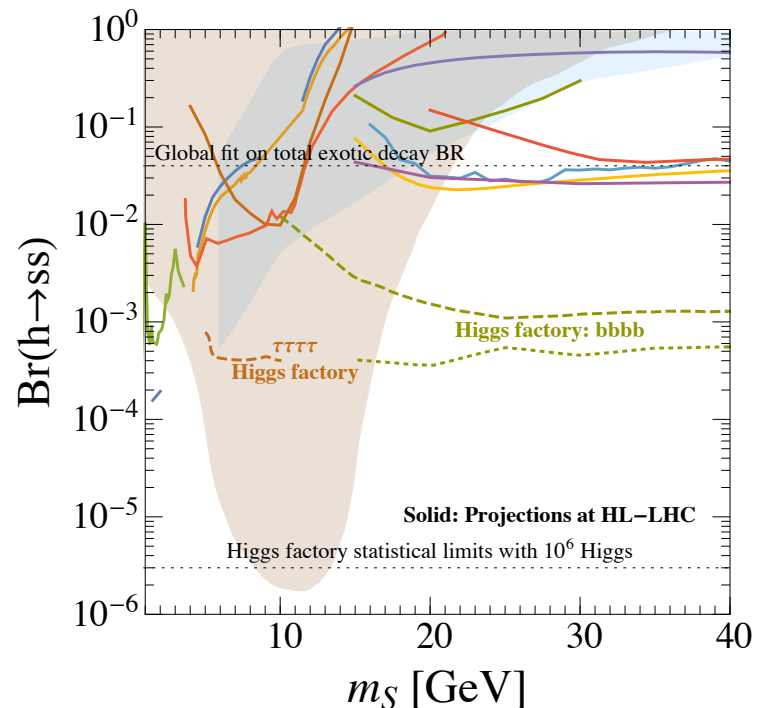
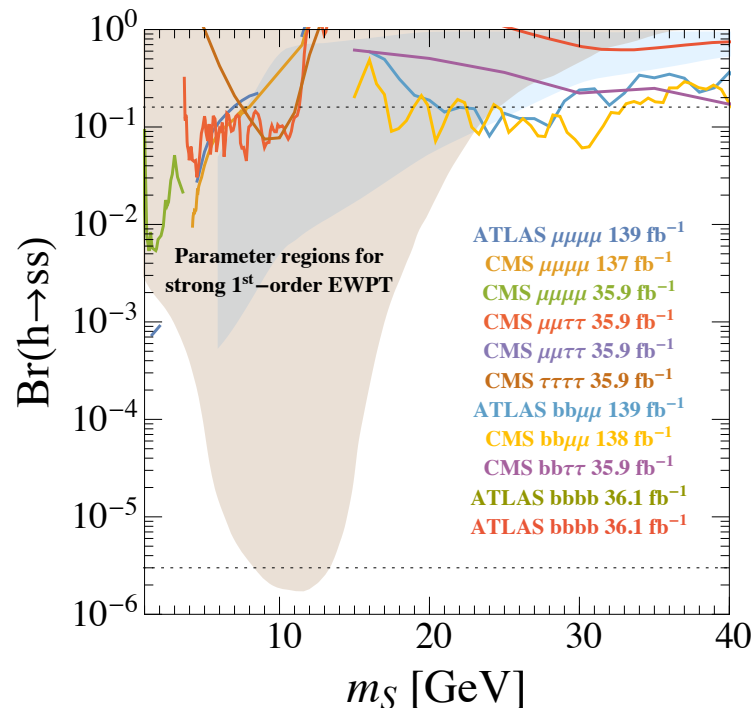
Light h_2 search

Higgs exotic decay $h_1 \rightarrow h_2 h_2 \rightarrow XXYY$ [SM light particles]

Carena, *KPX, et al*, LHEP 2023 (2023) 432

FOEWPT parameter space

- Light blue: fixing mixing angle $\sin \theta = 0.01$
Kozaczuk et al, PRD 101 (2020) 11, 115035
- Brown: spontaneous broken Z_2 of S
Carena et al, JHEP 08 (2020) 107

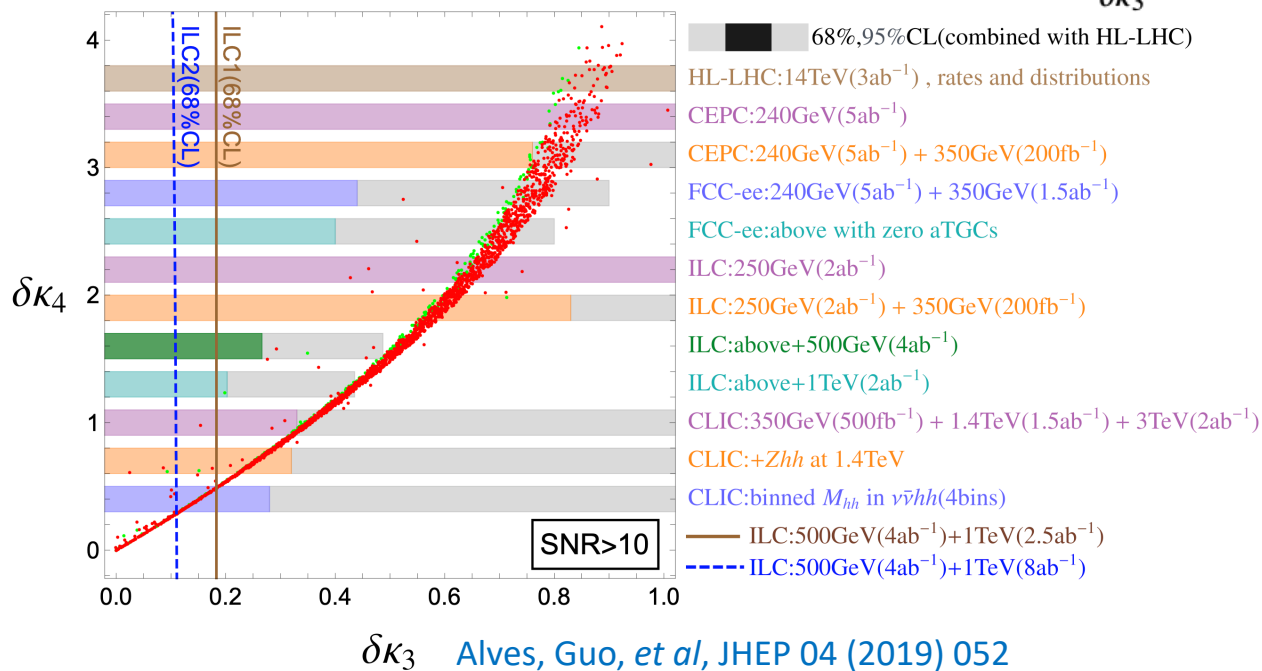
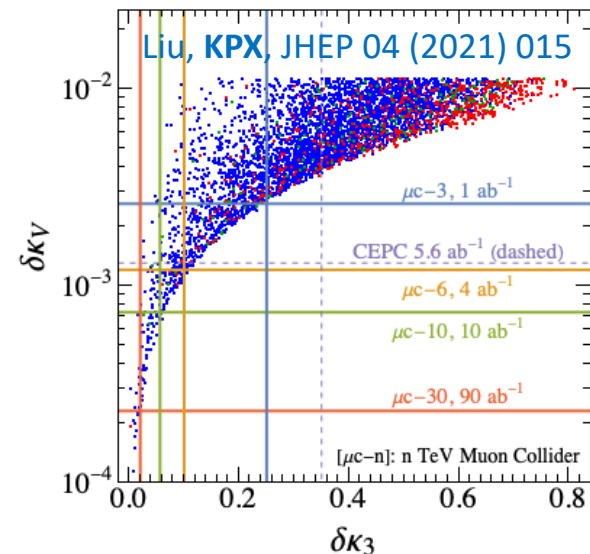


Higgs precision measurement

Higgs coupling deviations

$$\delta\kappa = |g/g_{\text{SM}} - 1|$$

- $\delta\kappa_V \rightarrow h_1 VV \propto |1 - \cos\theta|$
- $\delta\kappa_3 \rightarrow h_1^3$ and $\delta\kappa_4 \rightarrow h_1^4$
- CEPC, ILC, FCC-ee, CLIC, or muon colliders



Type III: loop driven

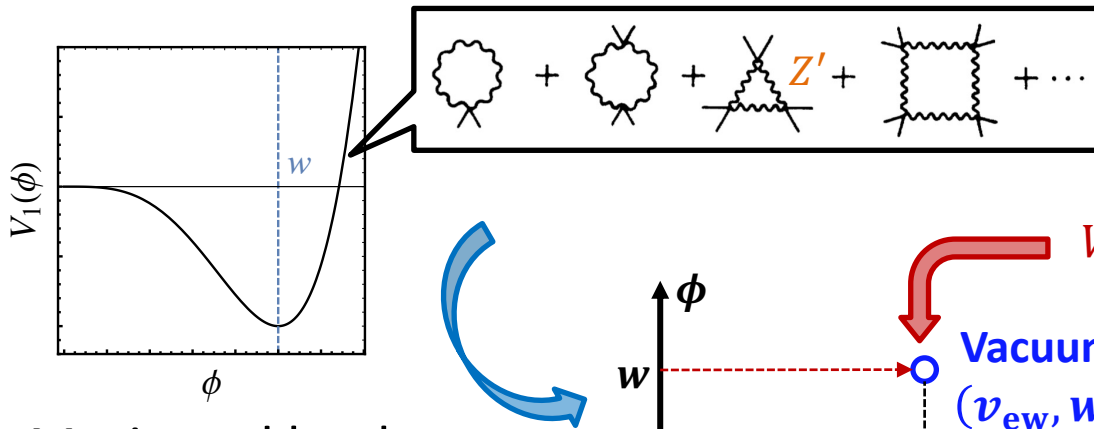
Based on **classically conformal principle**, tree-level

$$V_0(H, s) = \lambda_h |H|^4 + \lambda_s |s|^4 - \lambda' |H|^2 |s|^2$$

$s = (\phi + i\eta)/\sqrt{2}$ is charged under a dark $U(1)_X$

One-loop level [Coleman and Weinberg, PRD 7 (1973) 1888-1910]

$$V_1(h, \phi) \approx \frac{3g_X^4}{32\pi^2} \phi^4 \left(\log \frac{\phi}{w} - \frac{1}{4} \right) - \frac{\lambda'}{4} h^2 \phi^2 + \frac{\lambda_h}{4} h^4$$



Motivated by the **hierarchy problem**

Hempfling, PLB 379, 153 (1996); Iso *et al*, PLB 676, 81 (2009); Chun *et al*, PLB 725, 158 (2013), etc

After $\langle \phi \rangle \approx w$,

$$V_1 \rightarrow -\frac{\lambda' w^2}{4} h^2 + \frac{\lambda_h}{4} h^4$$

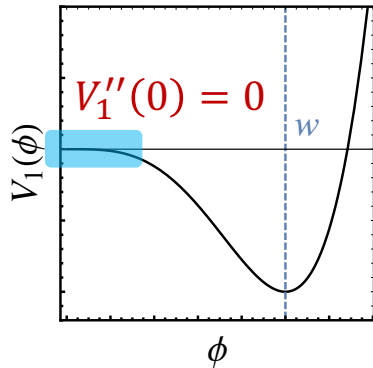
↓

EWSB with

$$v_{ew} \approx \sqrt{\frac{\lambda'}{\lambda_h}} w$$

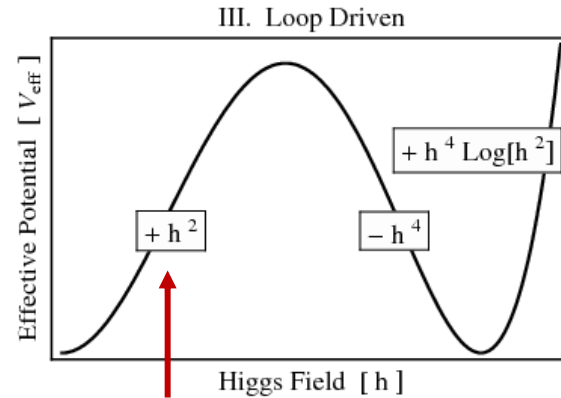
FOEWPT in classically conformal model

Very special feature of logarithmic potential



$$V_1(\phi) \sim \frac{B}{4} \phi^4 \left(\log \frac{\phi}{w} - \frac{1}{4} \right)$$

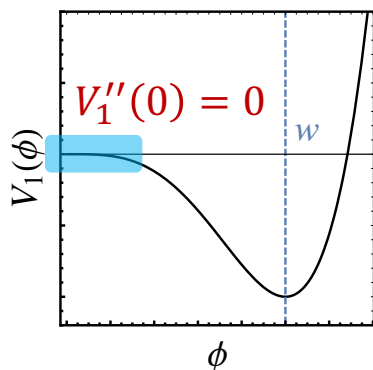
When $T > 0$



$\sim \lambda' g_X^2 T^2 h^2$ generated by Z' in hidden sector

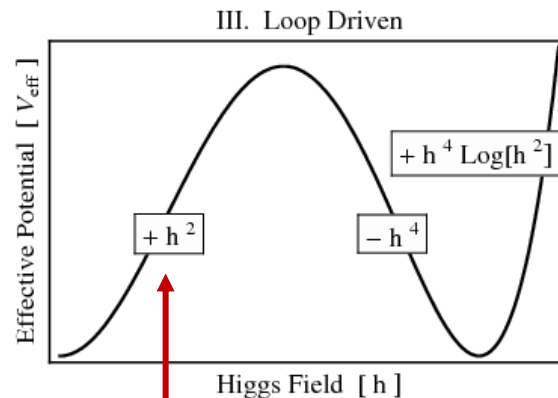
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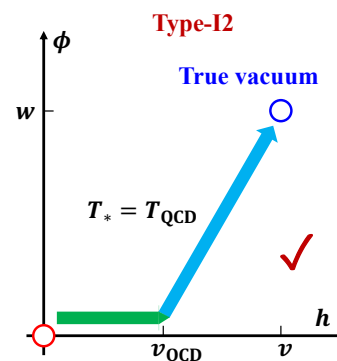
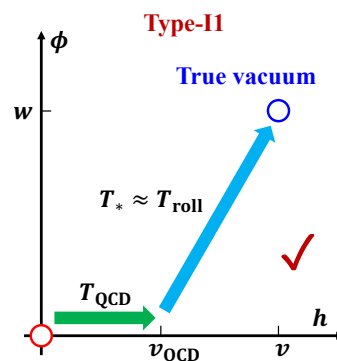
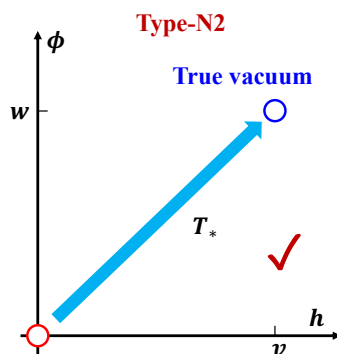
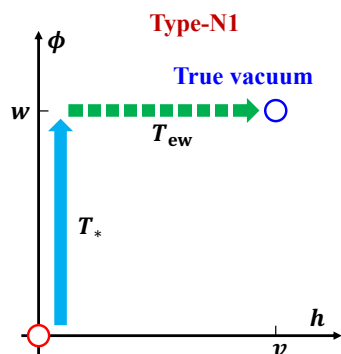
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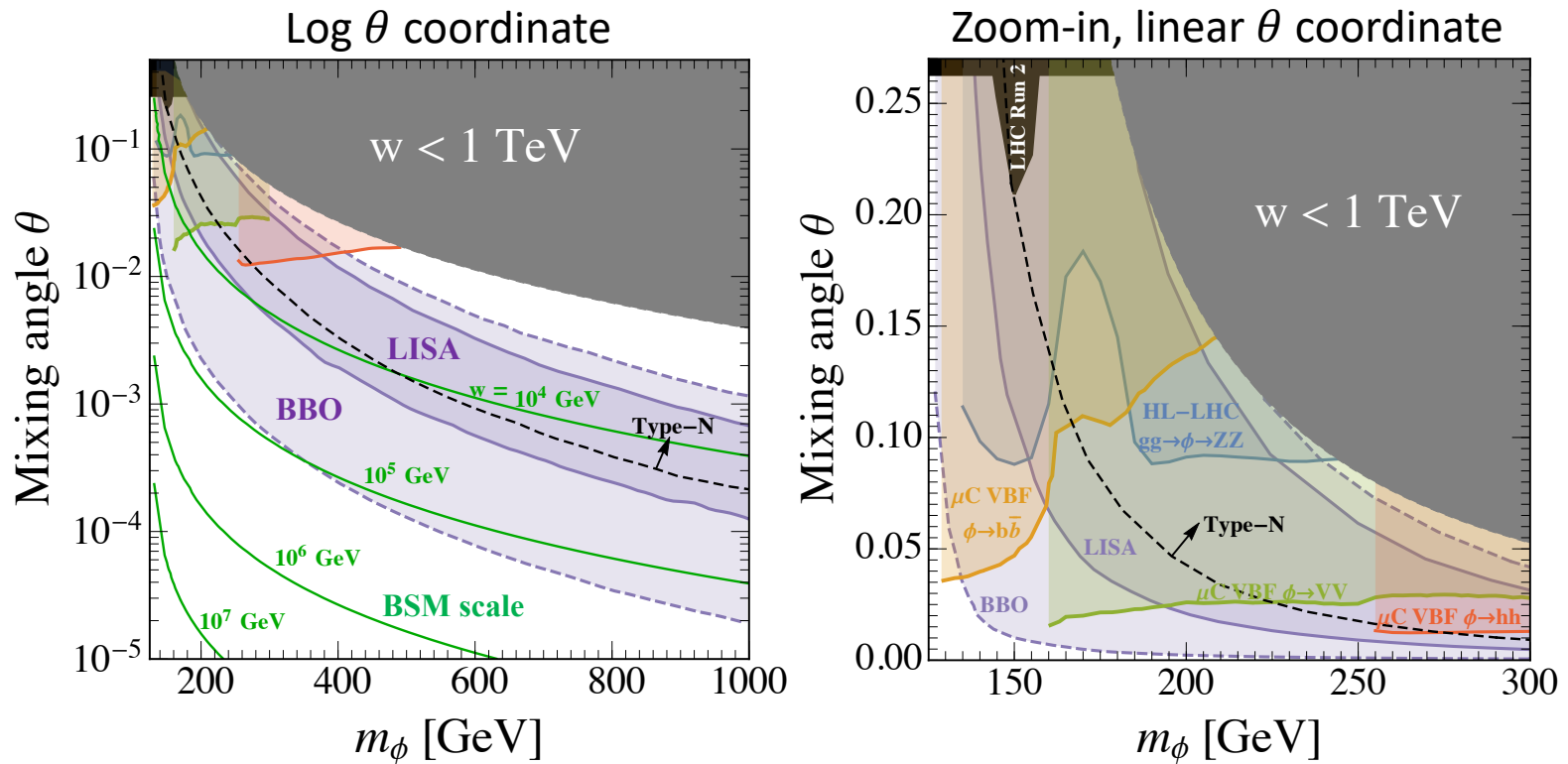
4 phase transition patterns (3 have FOEWPT) [Liu and KPX, 2408.03649]



Collider signals: for a heavy singlet

Collider search [Liu and KPX, 2408.03649]

- HL-LHC $gg \rightarrow \phi \rightarrow ZZ$
- 10 TeV muon collider VBF $\phi \rightarrow b\bar{b}, VV, hh$

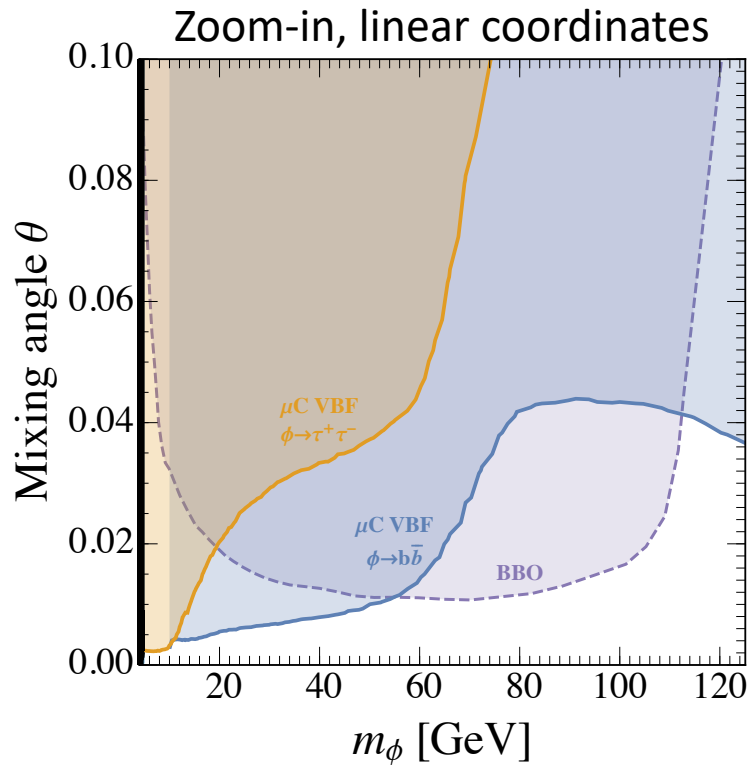
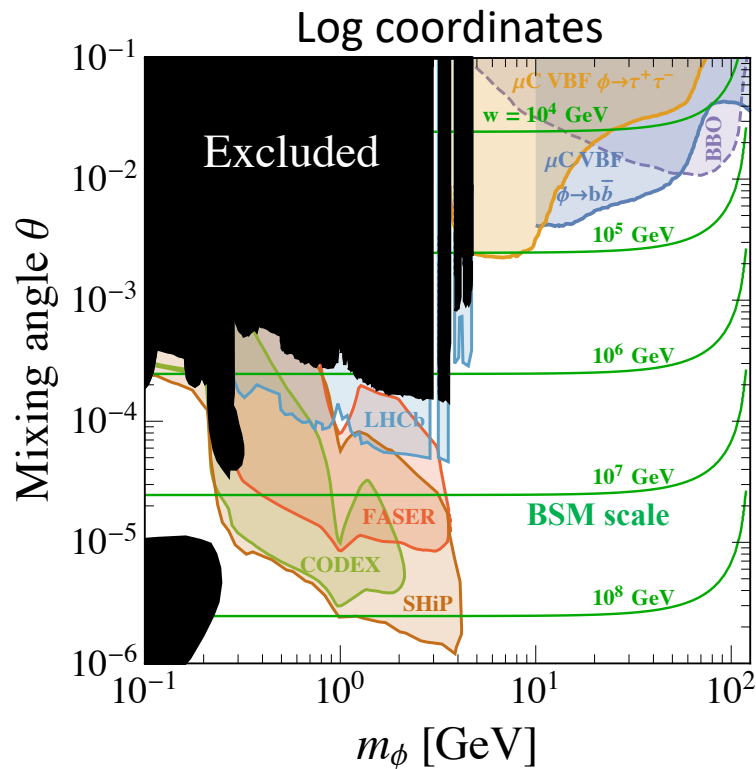


Gravitational wave search: LISA (also TianQin, Taiji) and BBO

Collider signals: for a light singlet

Collider search [Liu and KPX, 2408.03649]

- Long-lived particle search; muon collider
- NO Higgs exotic decay signal, as $\text{Br}(h \rightarrow \phi\phi) < 10^{-10}$



GW search NOT sensitive, as the FOPTs complete too fast

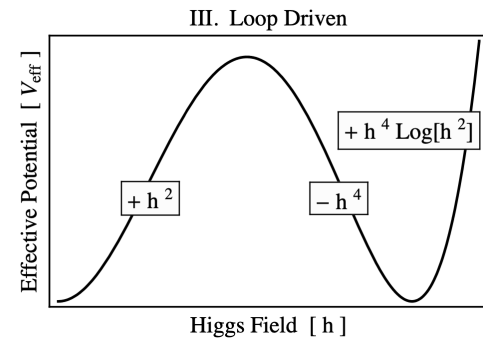
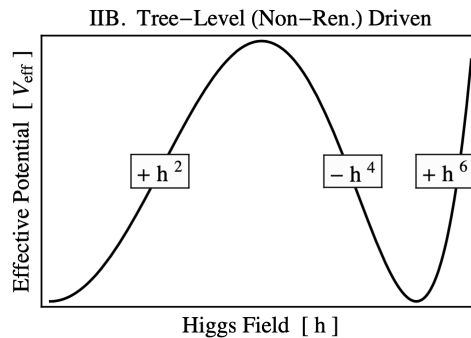
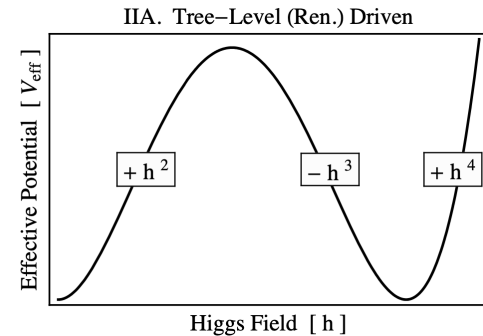
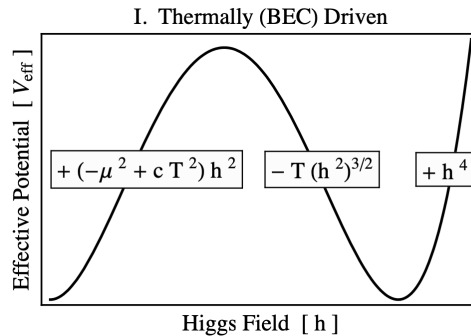
Summary

FOEWPT { A good probe for **physics beyond the SM**
 A key ingredient for solving **puzzles** in the SM

Origin of barrier:

1. Thermally driven
2. Tree-level ren.
3. Tree-level non-ren.
4. Loop driven

[Chung, Wang et al, PRD 87 \(2013\) 2, 023509](#)



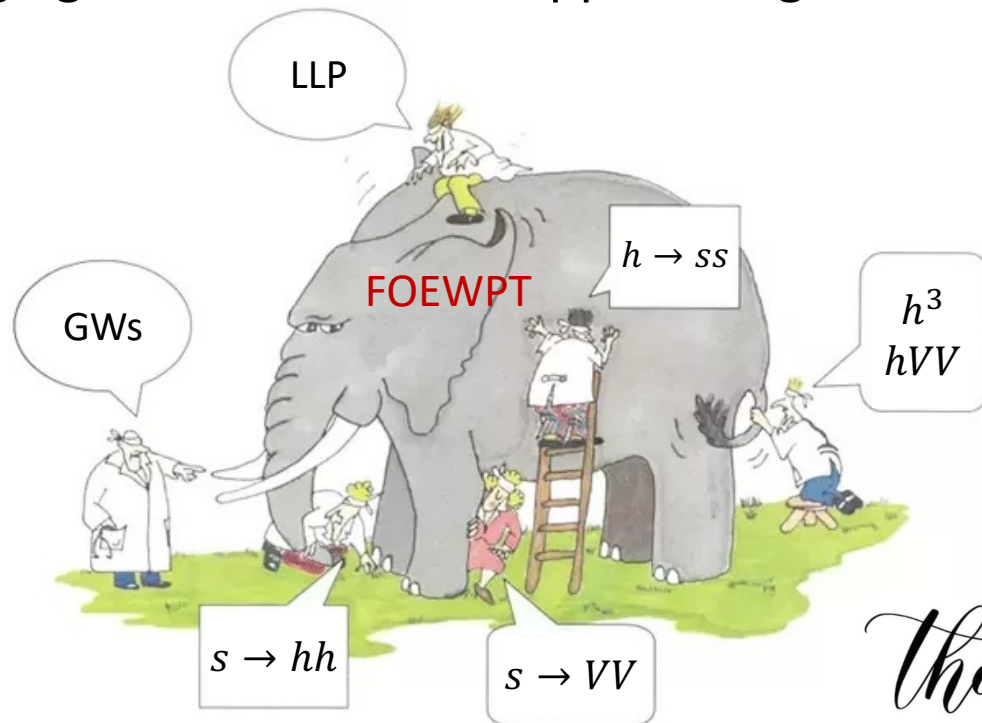
Typical signals

- Resonant $s \rightarrow hh$, $s \rightarrow W^+W^-/ZZ$, or $h \rightarrow ss \rightarrow XXYY$
- Higgs couplings deviation, especially h^3 or hVV

Hopefully we can detect **something** in the next decade!

Conclusion and outlook

It's challenging to confirm what happened right after the Big Bang!



Thank you

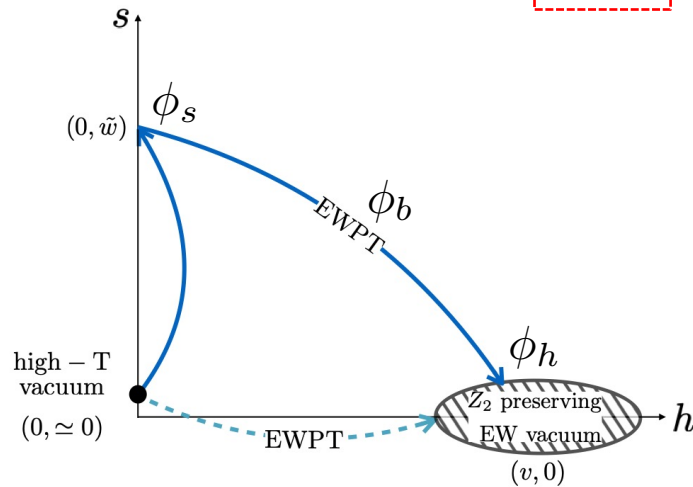
Future exploration

- More detailed simulations (e.g. boosted by AI)
- More sensitive channels (e.g. $h \rightarrow ss \rightarrow \mu^+ \mu^- \mu^+ \mu^-$ @ CEPC)
- Distinguishing the 4 scenarios by combining all channels
- ...

Backup: type IIA light singlet general case

The potential at finite temperature

$$V = -\frac{\mu^2 - c_h T^2}{2} h^2 + \frac{\lambda}{4} h^4 + \frac{a_1}{4} h^2 S + \frac{a_2}{4} h^2 S^2 + b_1 S + \frac{b_2 + c_S T^2}{2} S^2 + \frac{b_3}{3} S^3 + \frac{b_4}{4} S^4$$



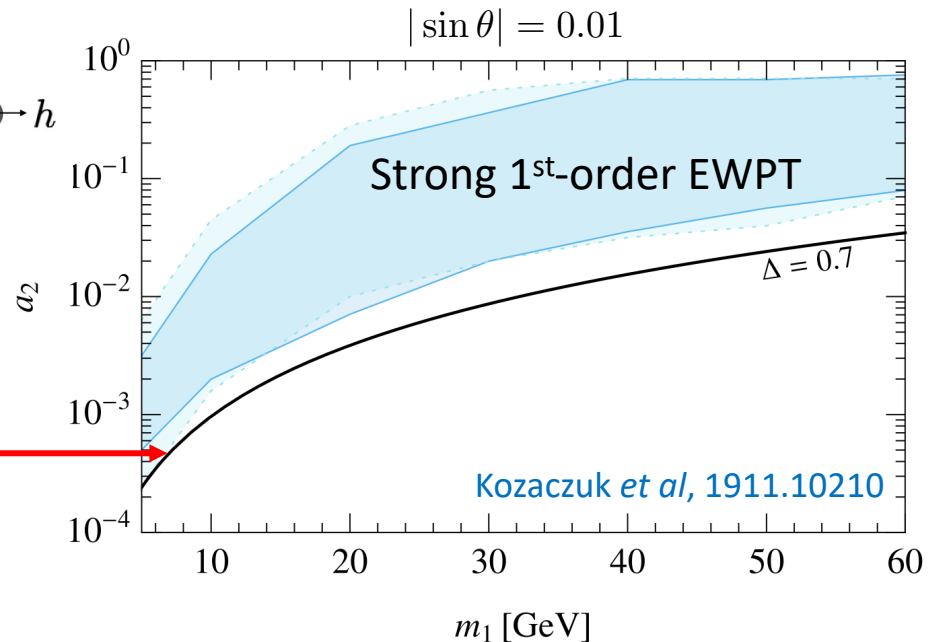
$$c_h = \frac{1}{48} (24a_2 + 9g^2 + 3g'^2 + 24\lambda + 12y_t^2),$$

$$c_S = \frac{1}{12} (2a_2 + 3b_4)$$

Analytical estimation

$$\frac{V(\phi_s, T_*) - V(\phi_h, T_*)}{V(\phi_b, T_*) - V(\phi_h, T_*)} > \Delta$$

$$a_2 \gtrsim \frac{m_1^2}{4v_{EW}^2} \frac{\Delta}{1 - \Delta}$$



Backup: type IIA Z_2 spontaneous breaking case

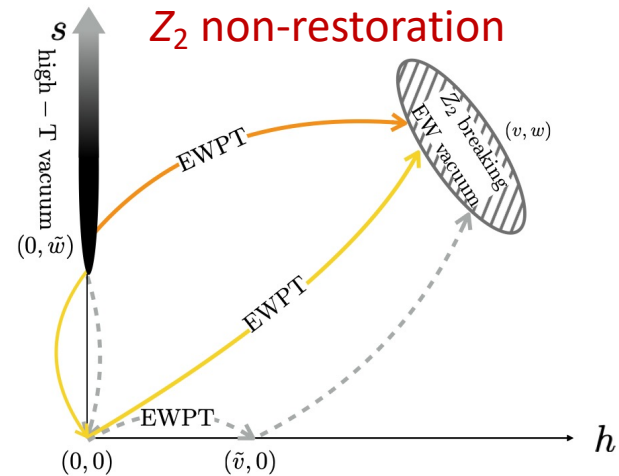
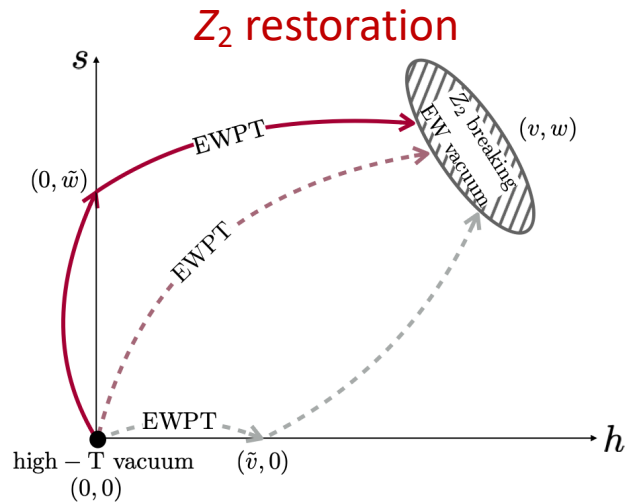
Motivated by dark symmetry breaking models [Carena *et al*, 1911.10206]

$$V = -\frac{\mu^2}{2}h^2 + \frac{\lambda}{4}h^4 + \frac{a_2}{4}h^2S^2 + \frac{b_2}{2}S^2 + \frac{b_4}{4}S^4$$

A complete one-loop level analysis

$$V_{\text{CW}} = \frac{1}{64\pi^2} \left(\sum_B n_B m_B^4(h, S) \left[\ln \frac{m_B^2(h, S)}{Q^2} - c_B \right] - \sum_F n_F m_F^4(h, S) \left[\ln \frac{m_F^2(h, S)}{Q^2} - \frac{3}{2} \right] \right)$$

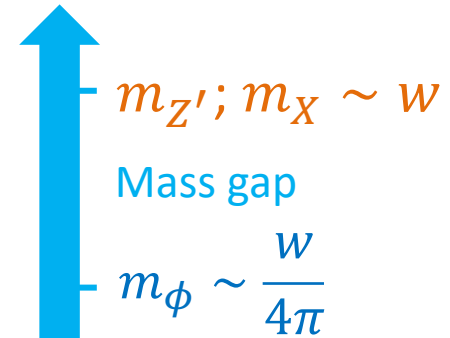
$$V^T(h, s, T) = \frac{T^4}{2\pi^2} \left[\sum_B n_B J_B \left(\frac{m_B^2(h, S)}{T^2} \right) + \sum_F n_F J_F \left(\frac{m_F^2(h, S)}{T^2} \right) \right] + \text{Daisy resummation}$$



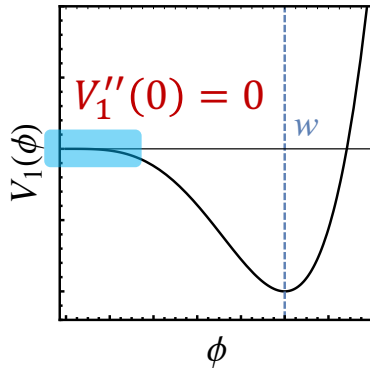
Backup: the classically conformal model

Particle content

- A scalar ϕ mixing with the Higgs
- Heavy particle(s) that generates the Coleman-Weinberg potential

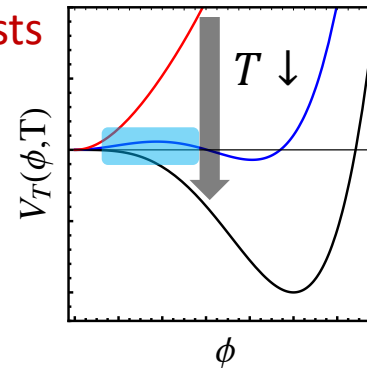


Early Universe evolution of the logarithmic-shaped potential



$$V_1(\phi) \sim \frac{B}{4} \phi^4 \left(\log \frac{\phi}{w} - \frac{1}{4} \right)$$

Barrier always exists



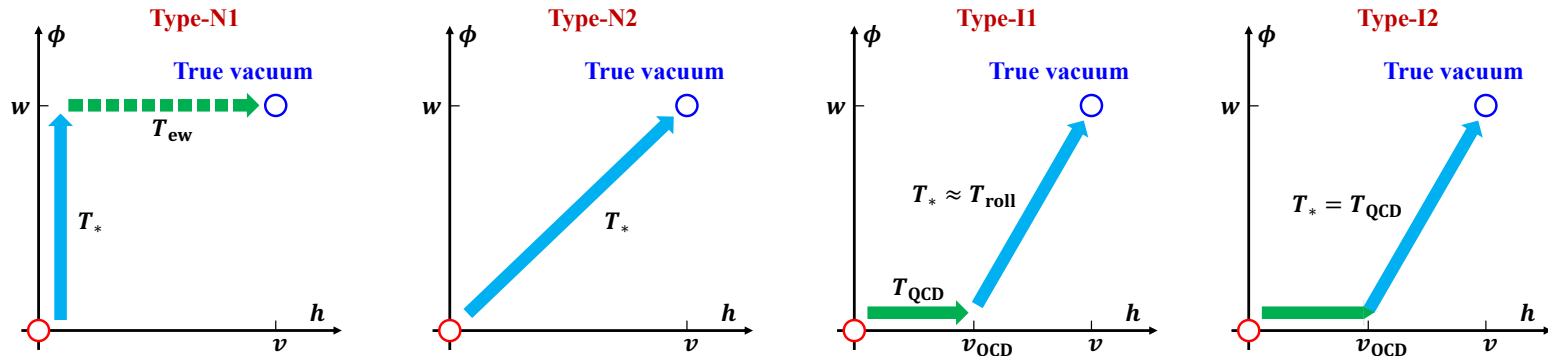
$$V_T(\phi, T) \sim g_X^2 T^2 \phi^2 + \frac{B}{4} \phi^4 \left(\log \frac{\phi}{w} - \frac{1}{4} \right)$$

First-order phase transition along the ϕ -direction **guaranteed**

Backup: thermal history patterns of CC theories

Normal pattern: high scale ϕ transition, then low scale h transition

- **N1:** ϕ -FOPT \rightarrow EW crossover
- **N2:** joint ϕ -EW-FOPT



Inverted pattern: low scale h transition, then high scale ϕ transition

- **I1:** QCD-EW-FOPT \rightarrow ϕ -FOPT
- **I2:** joint QCD-EW- ϕ -FOPT

Inverted pattern first proposed by Witten [NPB 177, 477 (1981)]

Transition rate $\Gamma(T) \sim e^{-S}$ with $S \propto g_X^{-3}$ [Iso et al, PRL 119 (2017) 14, 141301]

Small $g_X \rightarrow$ Universe trapped at $(h, \phi) = (0,0)$

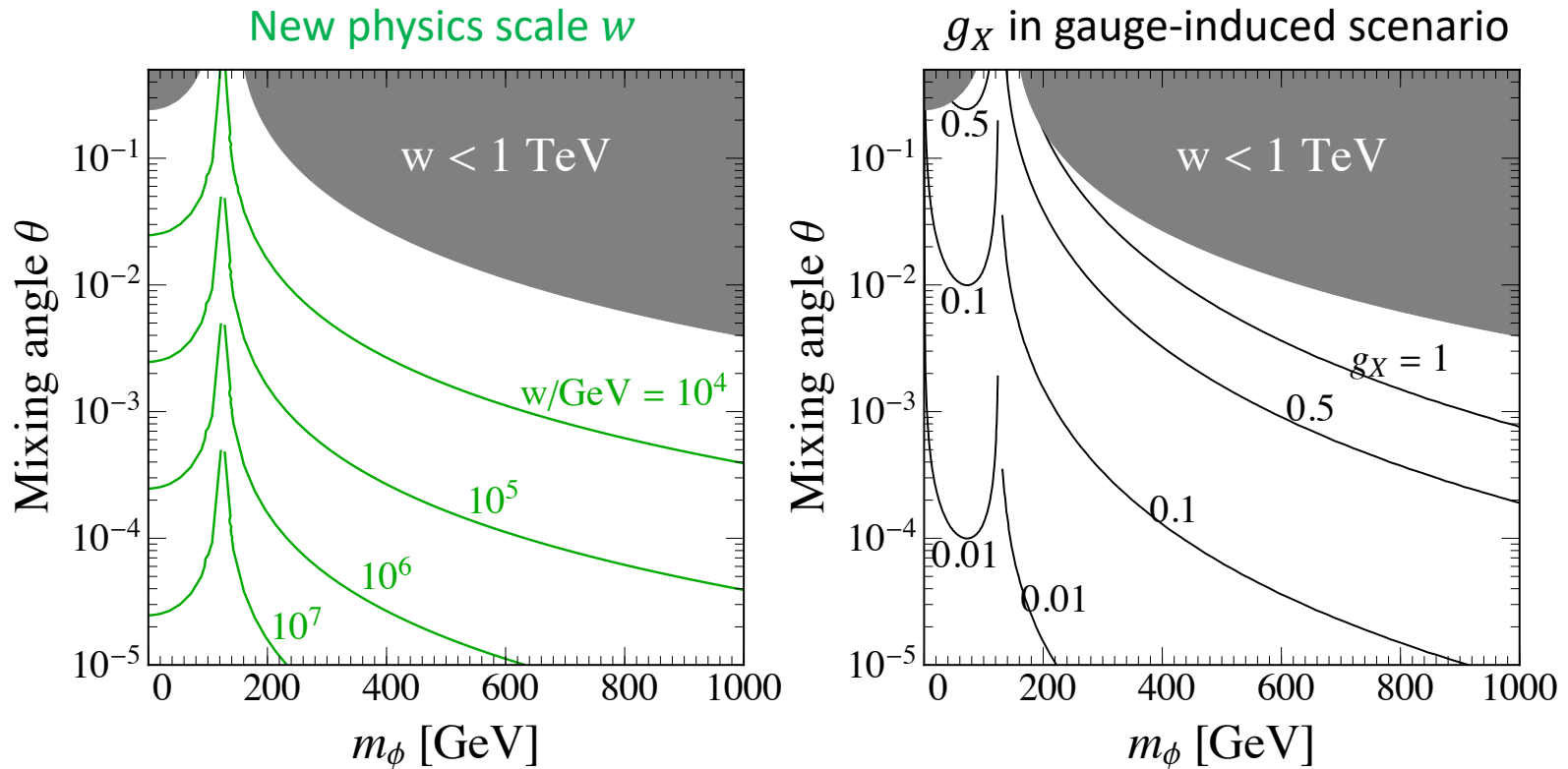
QCD transition occurs first! A FOPT with 6 massless quarks

- $-y_t h \langle \bar{t}t \rangle$ yields an **EWPT** $h: 0 \mapsto v_{\text{QCD}} \approx 100 \text{ MeV}$

Backup: parametrization of the CC model

Fixing $m_h = 125$ GeV and $v_{ew} = 246$ GeV --

- Only **2** free parameters
- First thorough and **rigorous** analysis in [\[Liu and KPX, 2408.03649\]](#)



- $m_\phi \neq 125$ GeV: ϕ cannot be degenerate with the Higgs

Backup: phase diagram of the CC model

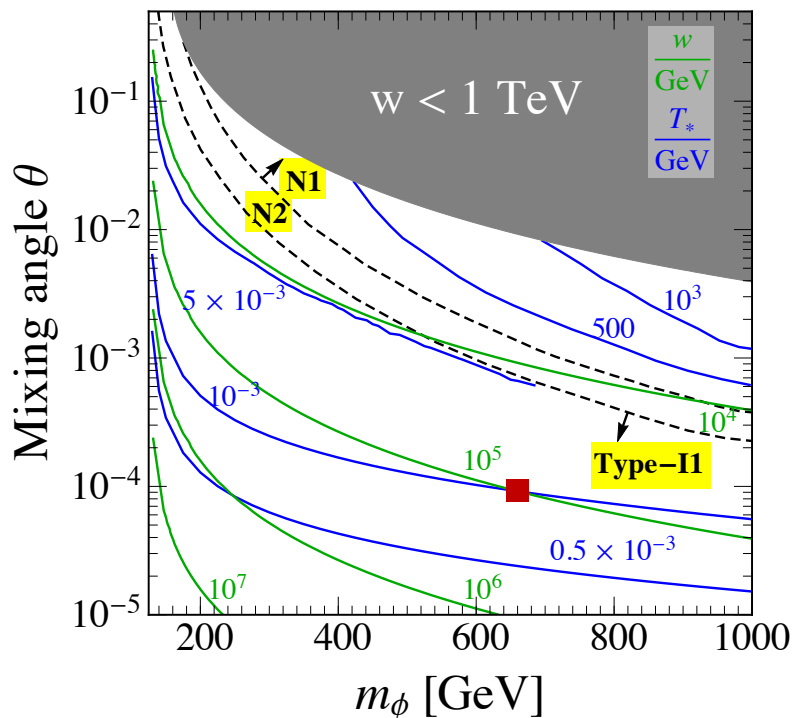
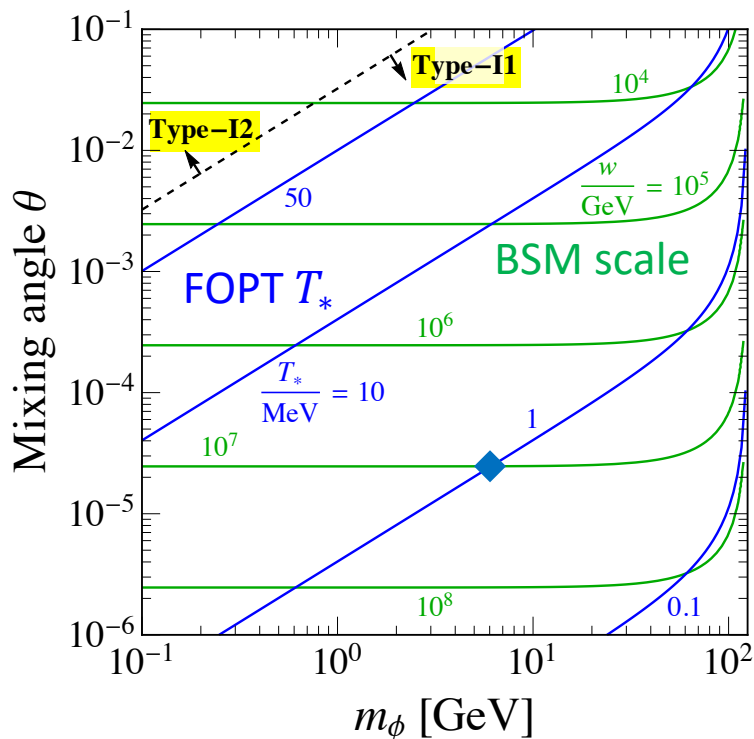
Full one-loop calculation & completeness check in [\[Liu and KPX, 2408.03649\]](#)

Type-N1: ϕ -FOPT \rightarrow EW crossover

Type-N2: joint ϕ -EW-FOPT

Type-I1: QCD-EW-FOPT \rightarrow ϕ -FOPT

Type-I2: joint QCD-EW- ϕ -FOPT



◆ $w = 10^7 \text{ GeV}, T_* = 1 \text{ MeV}$

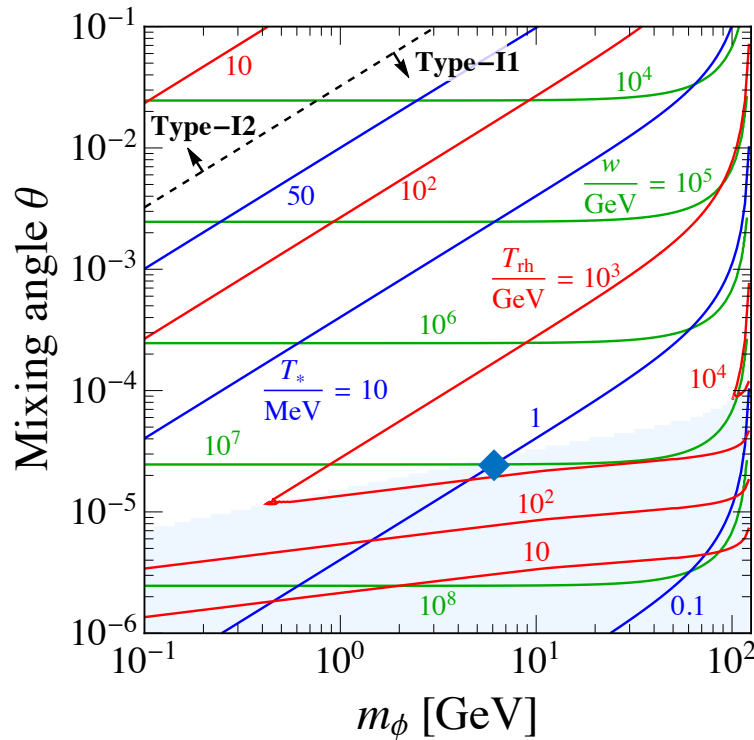
■ $w = 10^5 \text{ GeV}, T_* = 1 \text{ MeV}$

- Ultra-supercooled phase transitions!

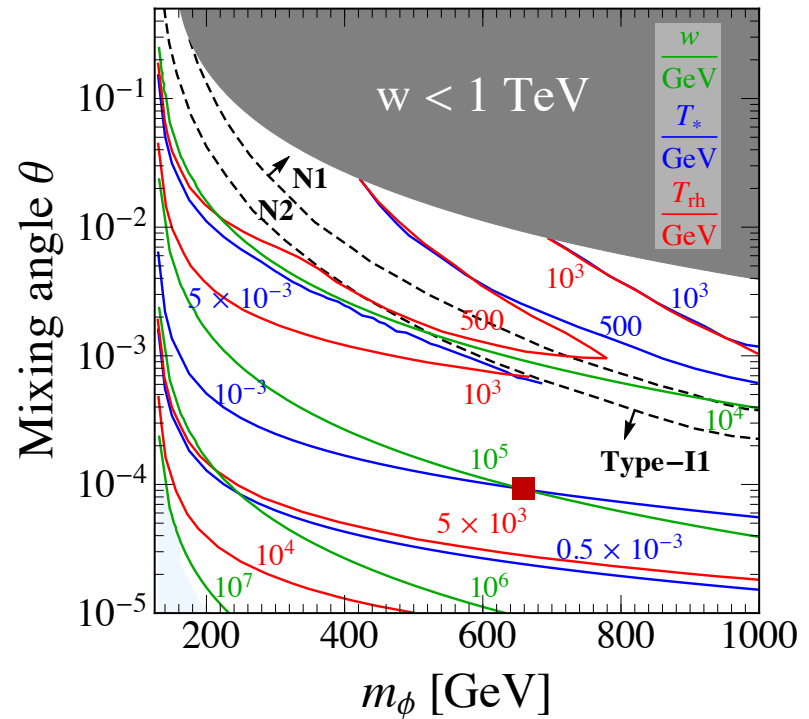
Backup: reheating in the CC model

The Universe is reheated to $T_{\text{rh}} \gtrsim T_*$ [Liu and KPX, 2408.03649]

- Possible EW symmetry **restoration** and a **second EWPT**
- Early matter era, if reheating is slow

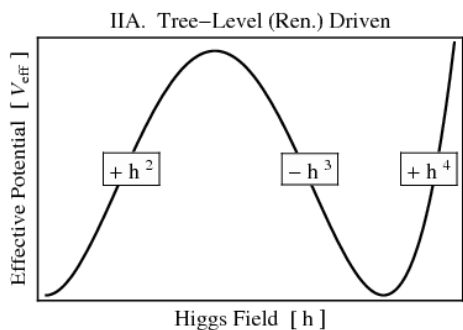


◆ $w = 10^7 \text{ GeV}, T_* = 1 \text{ MeV}, T_{\text{rh}} \sim 1 \text{ TeV}$

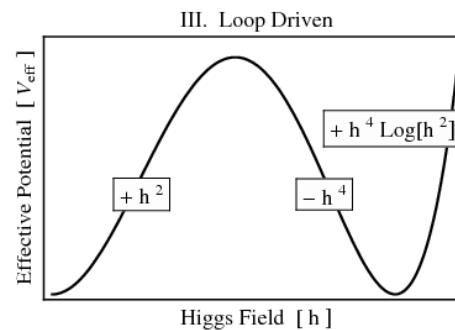


■ $w = 10^5 \text{ GeV}, T_* = 1 \text{ MeV}, T_{\text{rh}} \sim 3 \text{ TeV}$

Comparison between type IIA and type III



VS



	Type IIA	Type III
Motivation	Prototype of many BSM models	Hierarchy problem
Potential	Polynomial with mass terms	Logarithmic
# of parameters	5	2
Origin of barrier	h - s interaction	Z' from $U(1)_X$
FOEWPT	In some parameter space	Almost guaranteed
PT strength	Moderate	Ultra-supercooling
Pheno of heavy s	Resonant $s \rightarrow hh, VV$	
Pheno of light s	Higgs exotic decay	Long-lived particle