

29th Mini-workshop on the frontier of LHC
2024-12-15 @福州

Baryogenesis via symmetry non-restoration

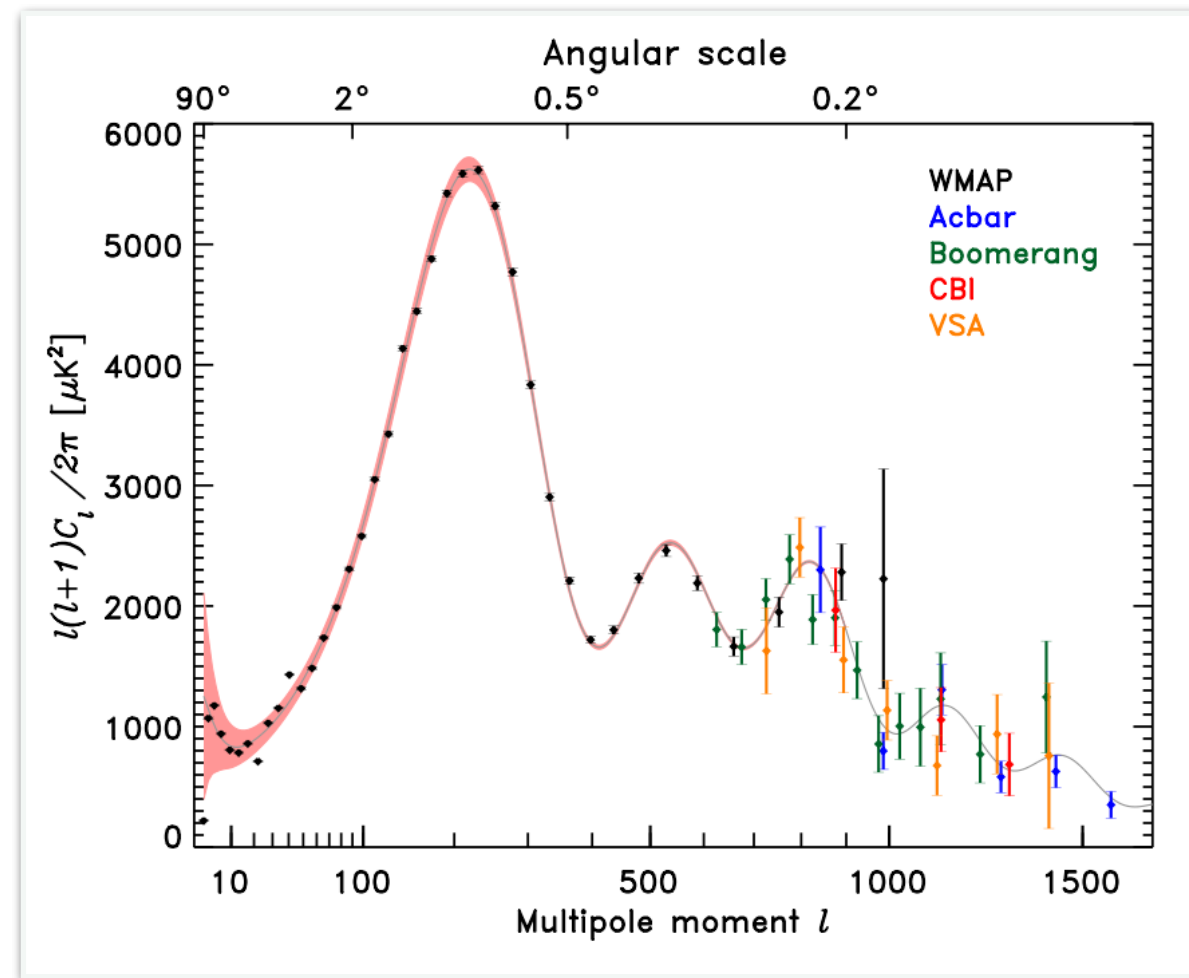
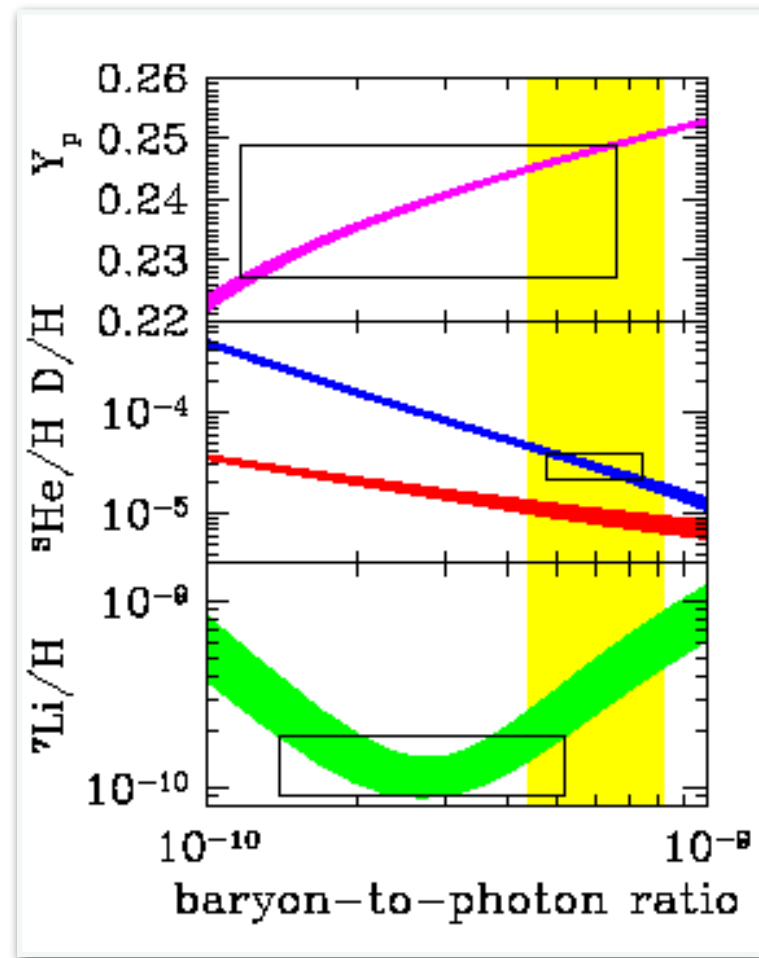
晁 伟

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New physics—The Baryon asymmetry

Matter-antimatter asymmetry

- 没有观测到反物质星系，否则光学望远镜会观测到星系湮灭的射线
- 元素的原初丰度以及CMB功率谱的形状都依赖于重子数与光子数之比



Baryon asymmetry:

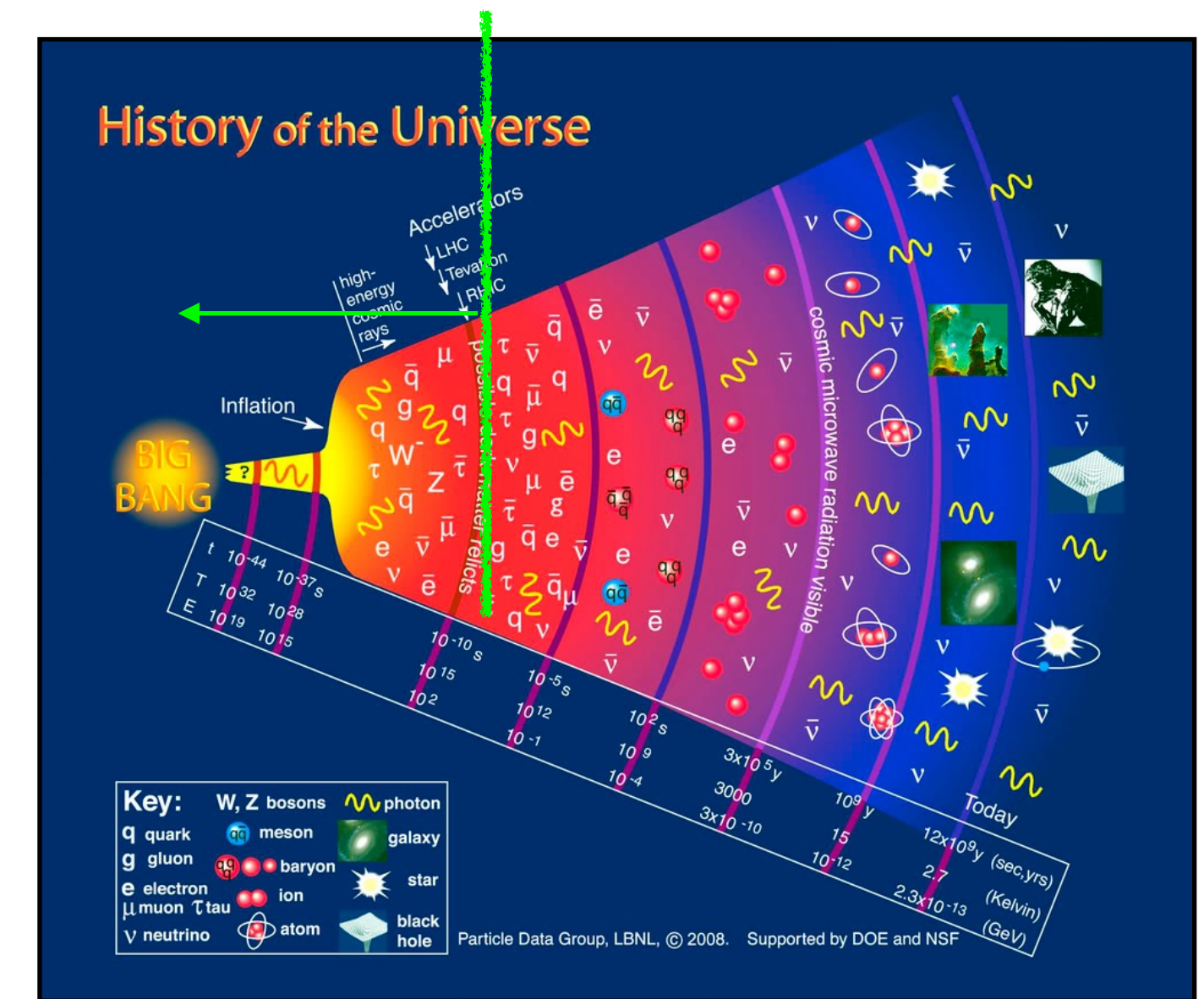
$$Y_B = \frac{\rho_B}{s} = (8.59 \pm 0.11) \times 10^{-11}$$

Planck

Baryogenesis



- ★ Baryon number violating
- ★ C&CP violation
- ★ Departure from equilibrium



- Leptogenesis
- Electroweak Baryogenesis
- GUT Baryogenesis
- Affleck-Dine Baryogenesis
- Post-sphaleron baryogenesis

History and development: Leptogenesis

轻子数破坏与 Leptogenesis (type-I seesaw case)

物理图像

Neutrino physics

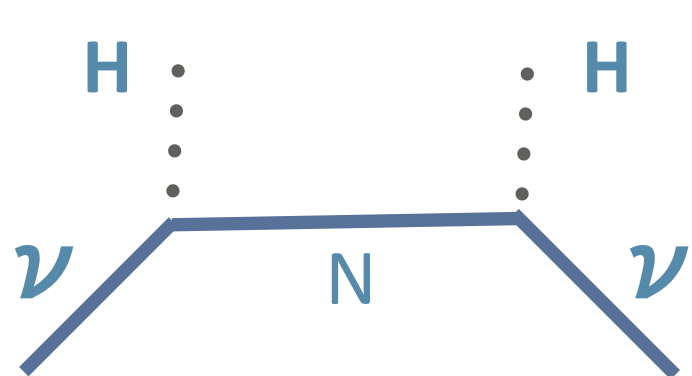
Leptogenesis via seesaw

BAU

Lagrangian:

$$-\mathcal{L} = Y_\nu \bar{\ell}_L \tilde{H} N_R + \frac{1}{2} \overline{N_R^C} M_R N_R + \text{h.c.}$$

Seesaw Mechanism

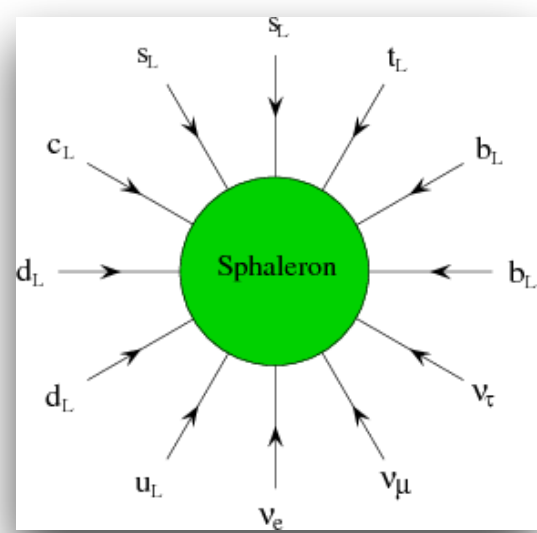


Neutrino mass:

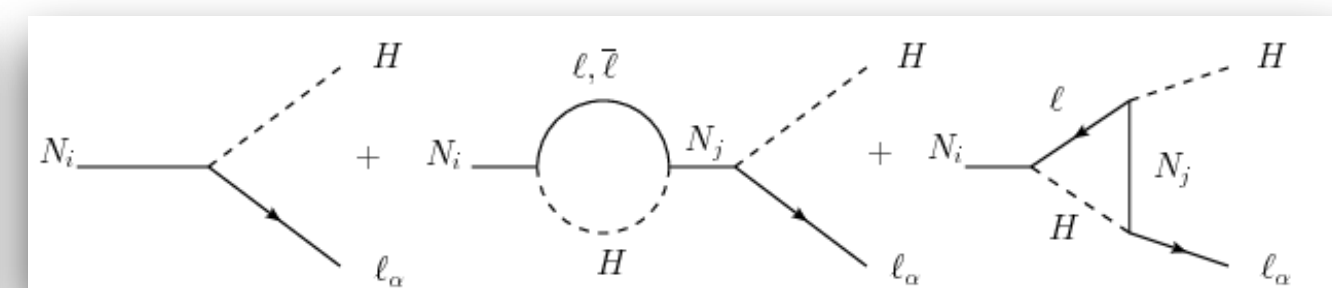
$$M_\nu = -M_D M_R^{-1} M_D^T$$

Elegant but losing testability at colliders!

BV



CPV

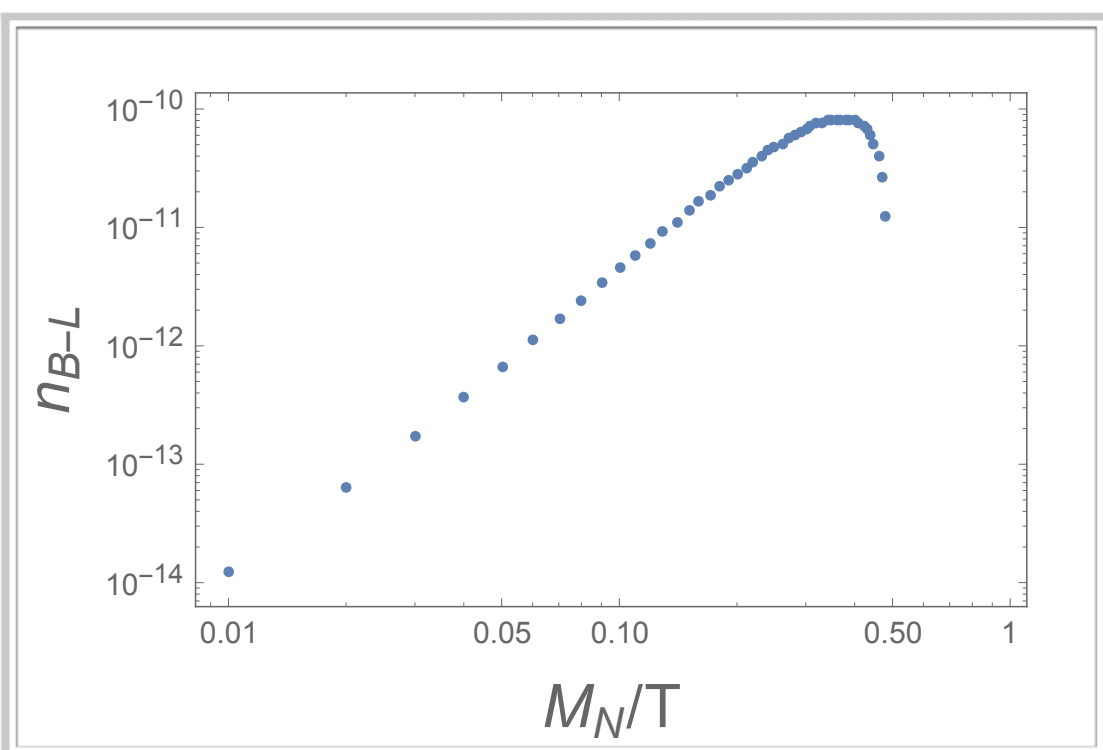


Out of EQ

Heavy neutrinos decouple

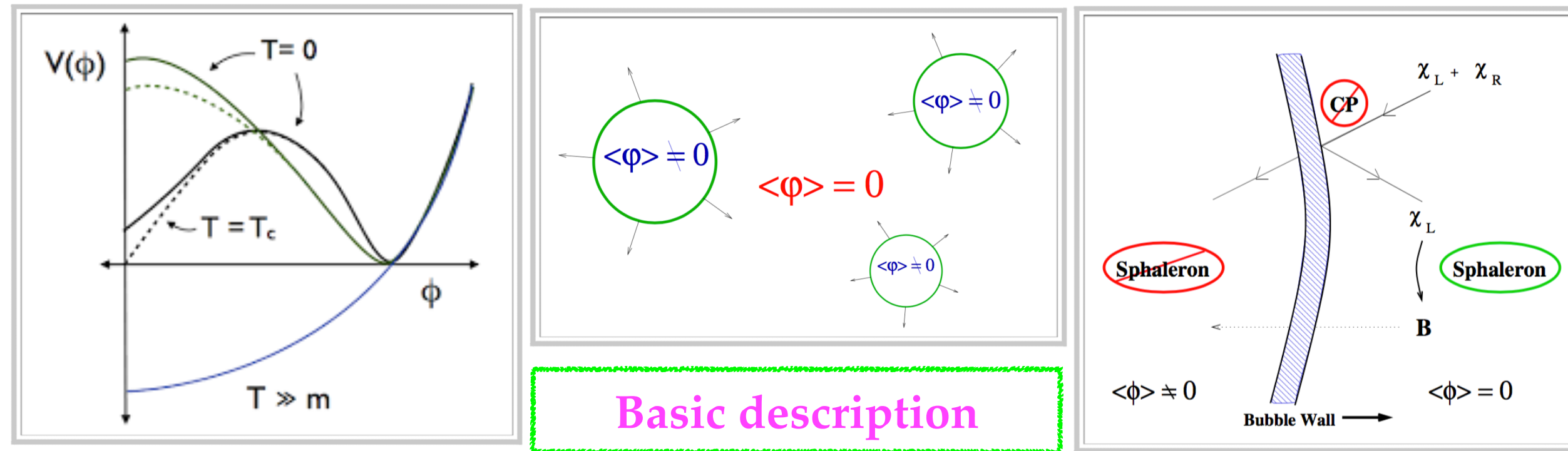
Boltzmann EQ

$$\frac{dN}{dz} = -D(N - N^{\text{eq}})$$

$$\frac{dN_{B-L}}{dz} = -\epsilon D(N - N^{\text{eq}}) + W_{\text{ID}} N_{B-L}$$


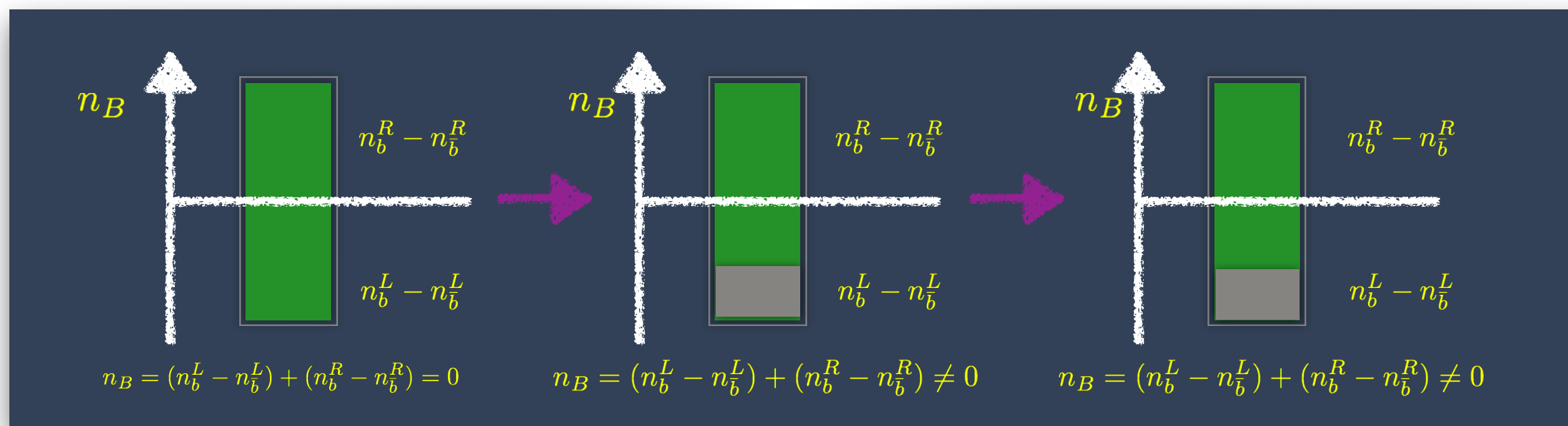
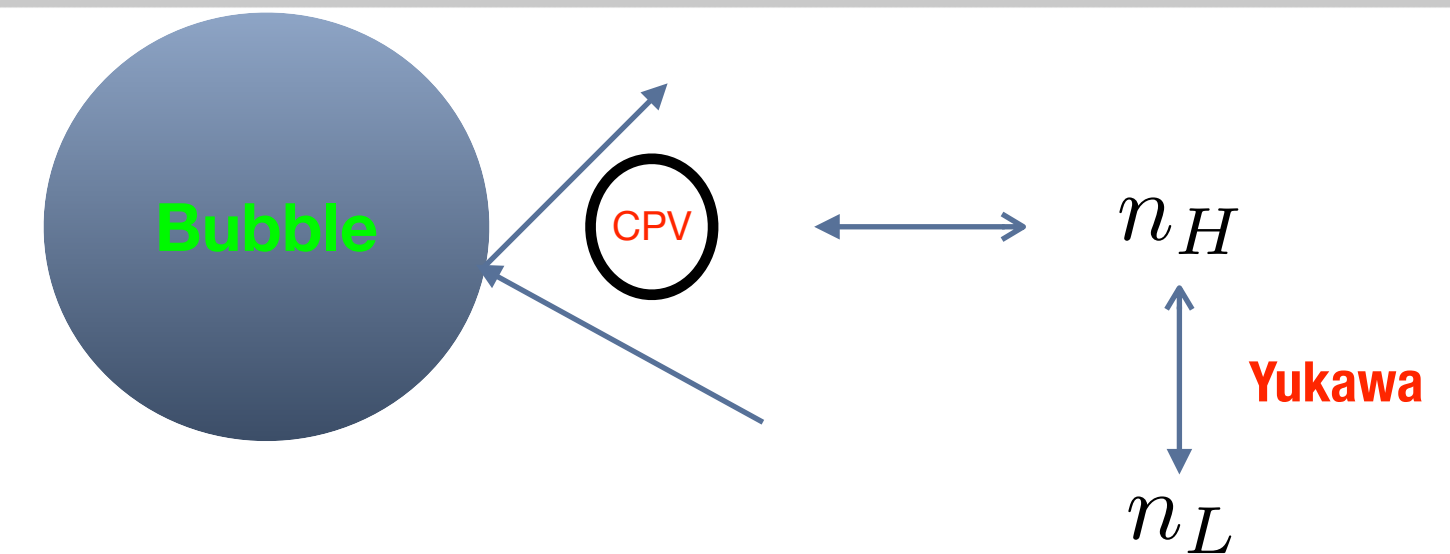
History and development: EW Baryogenesis

EWBG



输运方程

$$\frac{\partial n}{\partial t} + \nabla \cdot j(x) = - \int d^3z \int_{-\infty}^{x_0} dz^0 \text{Tr}[\Sigma^>(x, z)S^<(z, x) - S^>(x, z)\Sigma^<(z, x) + S^<(x, z)\Sigma^>(z, x) - \Sigma^<(x, z)S^>(z, x)]$$



$$\begin{aligned} \partial_\mu \psi_\mu &= +\Gamma_\psi^+ \left(\frac{\chi}{k_\chi} + \frac{\psi}{k_\psi} \right) + \Gamma_\psi^- \left(\frac{\chi}{k_\chi} - \frac{\psi}{k_\psi} \right) + \left(\sum_i \Gamma_{y_i} \right) \left(\frac{\chi}{k_\chi} - \frac{H}{k_H} - \frac{\psi}{k_\psi} \right) + S_{\text{CP}}^\psi \\ \partial_\mu \chi_\mu &= -\Gamma_\psi^+ \left(\frac{\chi}{k_\chi} + \frac{\psi}{k_\psi} \right) - \Gamma_\psi^- \left(\frac{\chi}{k_\chi} - \frac{\psi}{k_\psi} \right) - \left(\sum_i \Gamma_{y_i} \right) \left(\frac{\chi}{k_\chi} - \frac{H}{k_H} - \frac{\psi}{k_\psi} \right) - S_{\text{CP}}^\psi \\ \partial_\mu H_\mu &= \Gamma_{Y_t} \left(\frac{T}{k_T} - \frac{H}{k_H} - \frac{Q}{k_Q} \right) + \left(\sum_i \Gamma_{y_i} \right) \left(\frac{\chi}{k_\chi} - \frac{H}{k_H} - \frac{\psi}{k_\psi} \right) - \Gamma_h \frac{H}{k_H}, \end{aligned}$$

History and development: Afleck-Dine

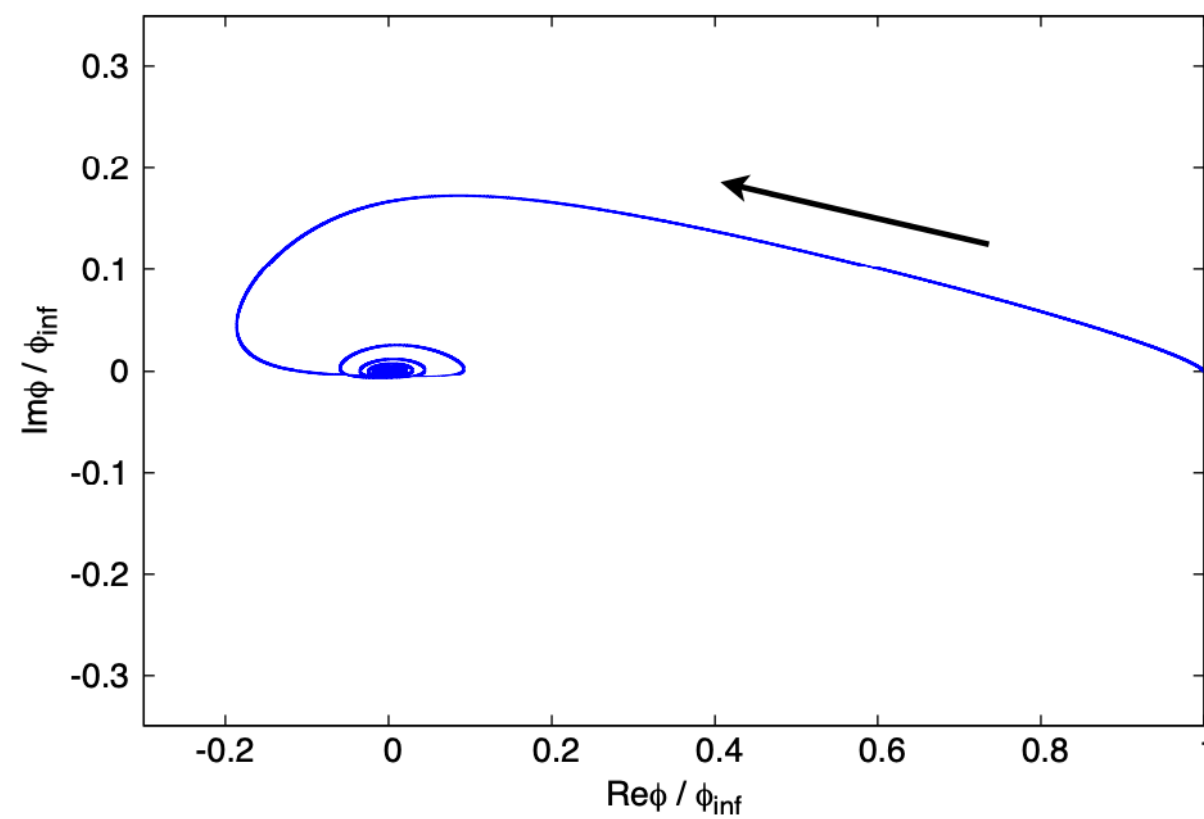
Affleck-Dine Mechanism

Scalars carrying non-zero U(1) charges

Flat directions (AD fields)

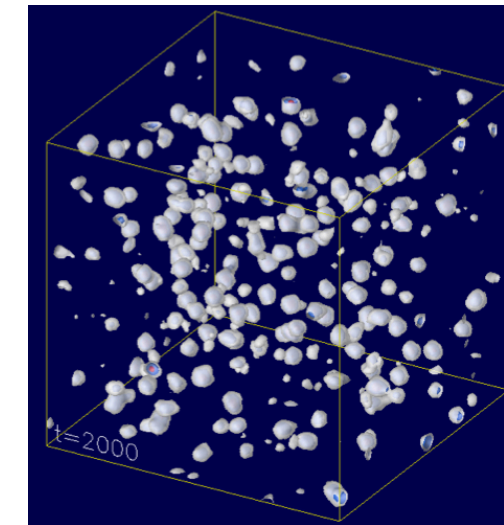
Lifting the potential via B/L violation operators

$$V = (m^2 - cH^2) |\phi|^2 + \lambda |\phi|^4 + \left(\frac{\phi^n}{M^{n-4}} + \text{h.c.} \right)$$



$$\dot{n}_{B,L} + 3Hn_{B,L} = 2\beta \text{Im} \left[\frac{\partial V}{\partial \phi} \phi \right]$$

Q-ball formation (Non-topological soliton in scalar field theory)



Oscillation of AD field

Q-ball formation

Long lived Q-ball

DM candidate

Evaporation

BAU when sphaleron erase is irrelevant

History and development: Axionogenesis

QCD Axion

Strong CP problem

Peccei-Quinn mechanism

Axion

$$d_n = 5.2 \times 10^{-16} \bar{\theta} e \cdot \text{cm}$$

$$d_n \leq 10^{-26} e \cdot \text{cm}$$

Pseudo-scalar particle

[Preskill, Wise, Wilczek (1983)]

[Abbott, Sikivie (1983)]

[Dine, Fischler (1983)]

Misalignment mechanism : $\phi_0 \neq 0, \dot{\phi} = 0$

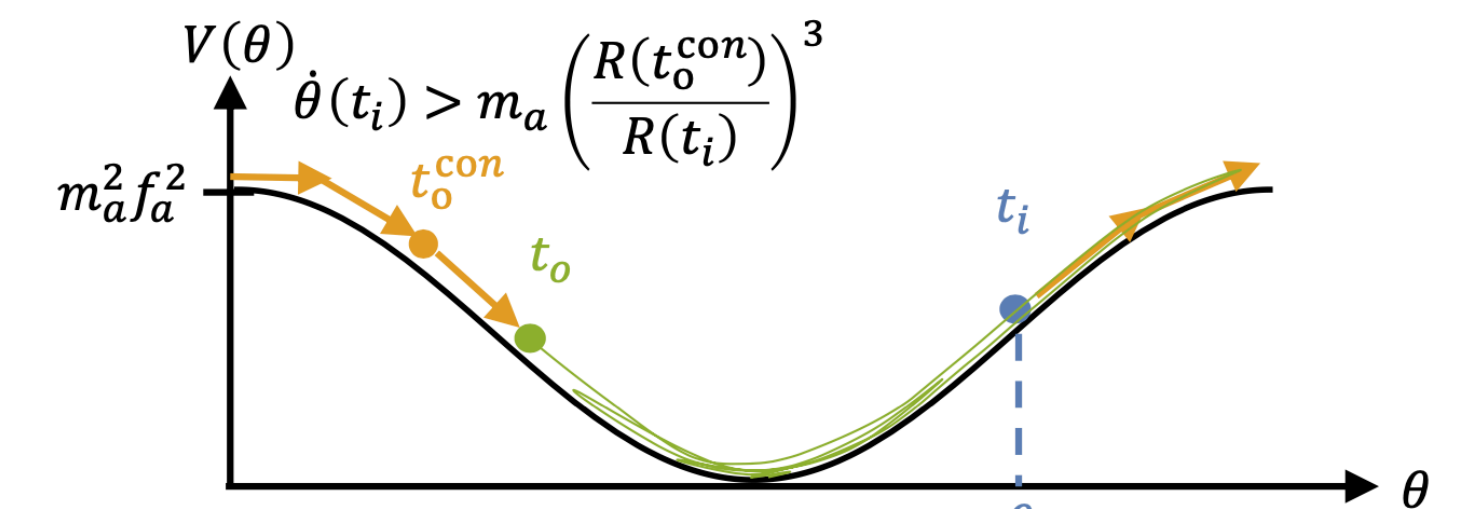
$$\ddot{\phi} + 3H\dot{\phi} + \frac{\Lambda^4(T)}{f^2}\phi = 0$$

$$\rho_{\text{DM}} \sim m_\phi \left[\frac{a(T_{\text{osc}})}{a_0} \right]^3 \left[\frac{\Lambda(T_{\text{osc}})^4 \theta_i^2}{m_\phi(T_{\text{osc}})} \right]$$

Axiogenesis

Kinetic Misalignment mechanism : $\phi_0 \neq 0, \dot{\phi} \neq 0$

[Co, Hall, Harigaya (2019)]
[Chang, Cui (2019)]



Non-zero Peccei-Quinn number \rightarrow Axionogenesis!

Strong sphaleron

$$n_{PQ} = S^2 \dot{\phi}$$

EWsphaleron

Quark chiral asymmetry

EWsphaleron

Baryon asymmetry

PRL, 124, 111602

Main point of this talk about Leptogenesis

**Traditional
Leptogenesis
mechanism:**

- **There must be primordial B-L violation**
- **There must exist right-handed neutrinos**

Eogenesis via the High-scale Electroweak Symmetry Restoration

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2412.03902

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Beijing Normal University, Beijing, 100875, China*

In this paper, we propose a novel electron-assisted Baryogenesis scenario that does not require explicit B-L violation, which is essential for the traditional Leptogenesis mechanism. This scenario

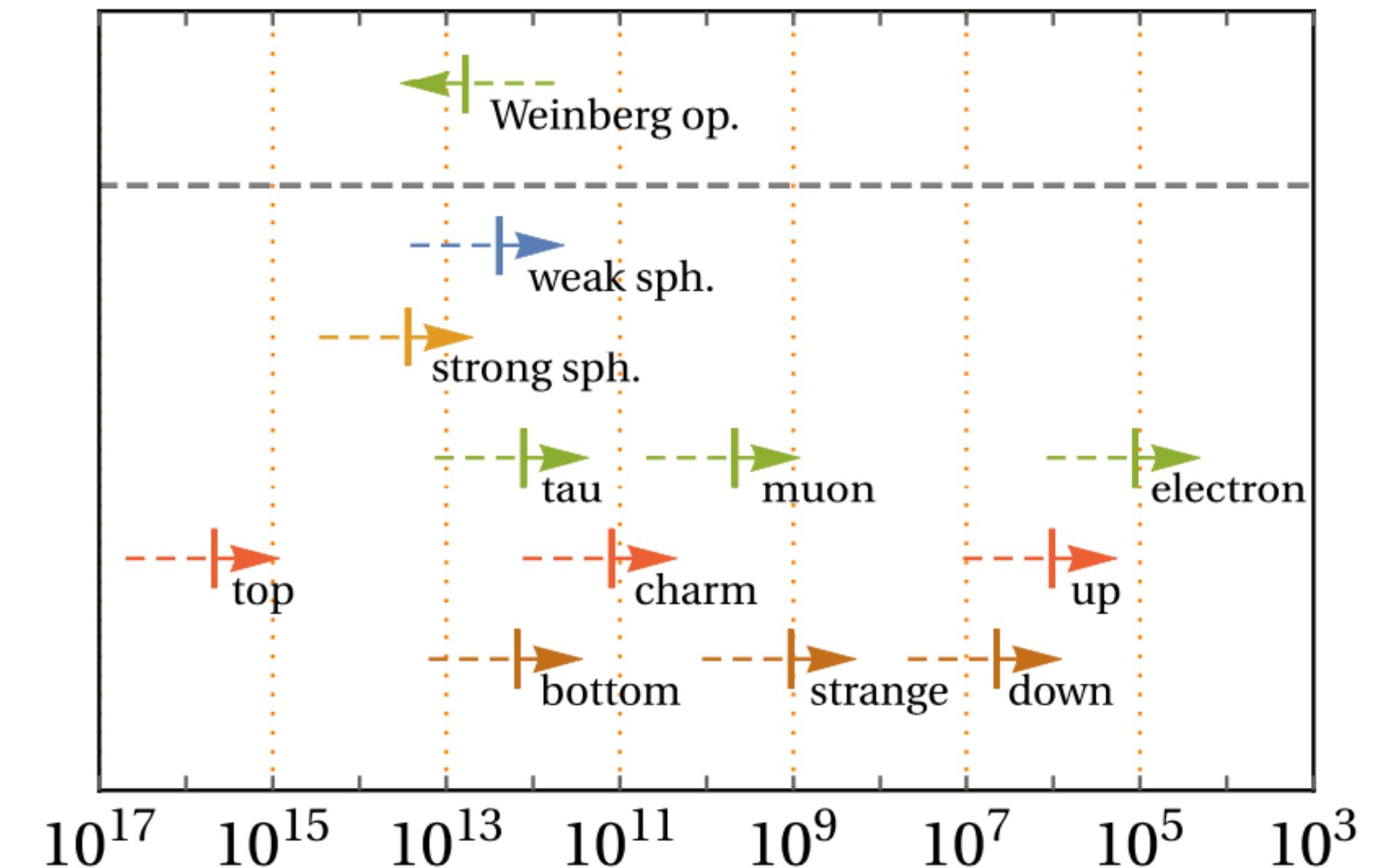
Not necessary!

**Sphaleron quenches before
electron Yukawa interaction
entering thermal equilibrium in
the early universe!**

Parameters for various interactions

- Key Point: Sphaleron may quench before the electron Yukawa interaction entering the equilibrium**

Interaction	Weinberg	WS	SS	Y_e	Y_μ	Y_τ
Γ_α/T^4	$\kappa_W \frac{m_\nu^2 T^2}{v_{EW}^4}$	$\frac{1}{2} \kappa_{WS} \alpha_2^5$	$\frac{1}{2} \kappa_{SS} \alpha_3^5$	$\kappa_{Y_e} y_e^2$	$\kappa_{Y_\mu} y_\mu^2$	$\kappa_{Y_\tau} y_\tau^2$
T_α [GeV]	6.0×10^{12}	2.5×10^{12}	2.8×10^{13}	1.1×10^5	4.7×10^9	1.3×10^{12}
Interaction	Y_u	Y_c	Y_t	Y_d	Y_s	Y_b
Γ_α/T^4	$\kappa_{Y_u} y_u^2$	$\kappa_{Y_c} y_c^2$	$\kappa_{Y_t} y_t^2$	$\kappa_{Y_d} y_d^2$	$\kappa_{Y_s} y_s^2$	$\kappa_{Y_b} y_b^2$
T_α [GeV]	1.0×10^6	1.2×10^{11}	4.7×10^{15}	4.5×10^6	1.1×10^9	1.5×10^{12}

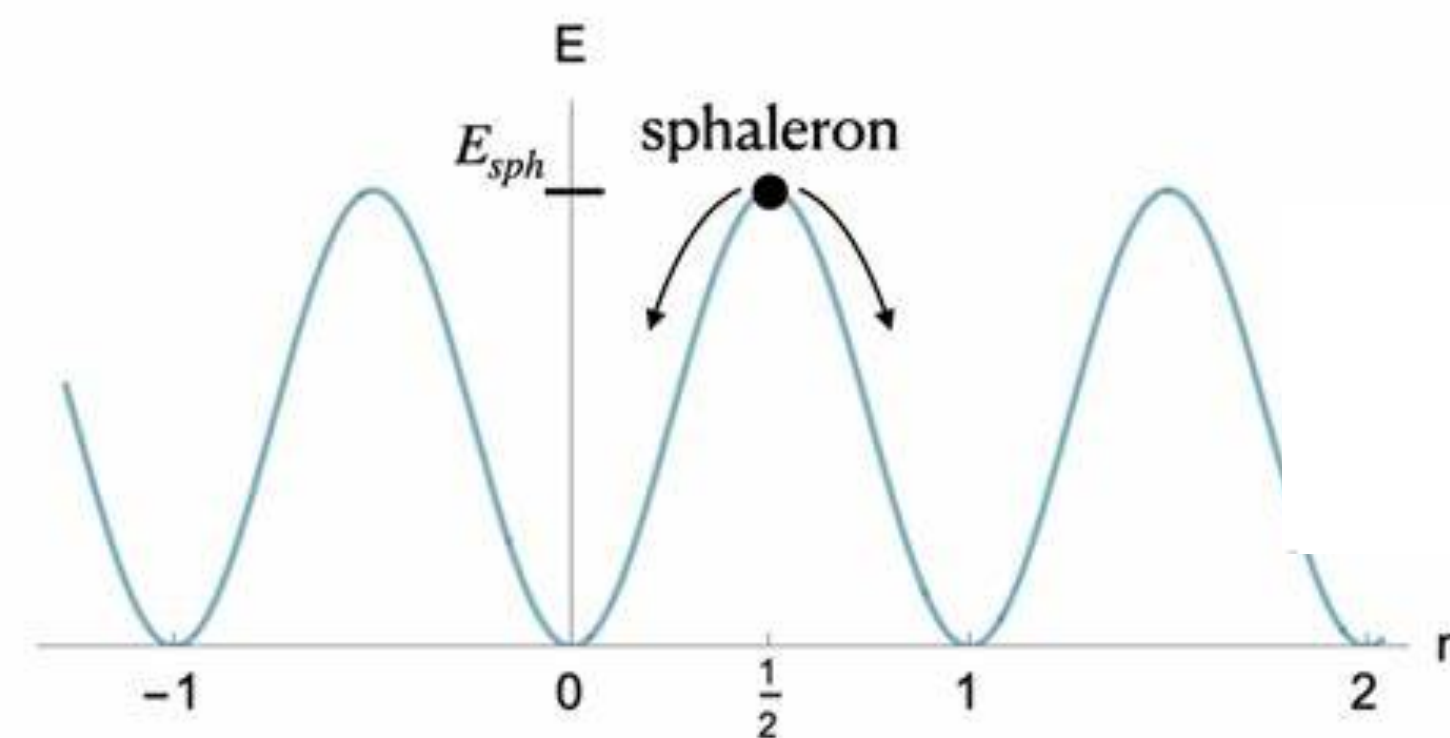
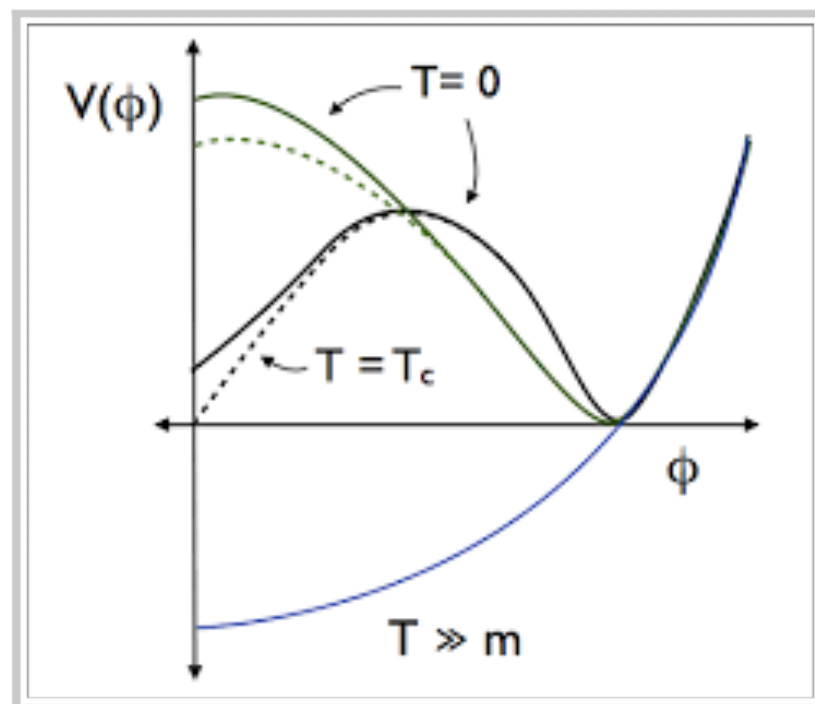


Sphaleron

- **Typical temperature:**

Sphaleron quench temperature: $T = 130 \text{ GeV}$

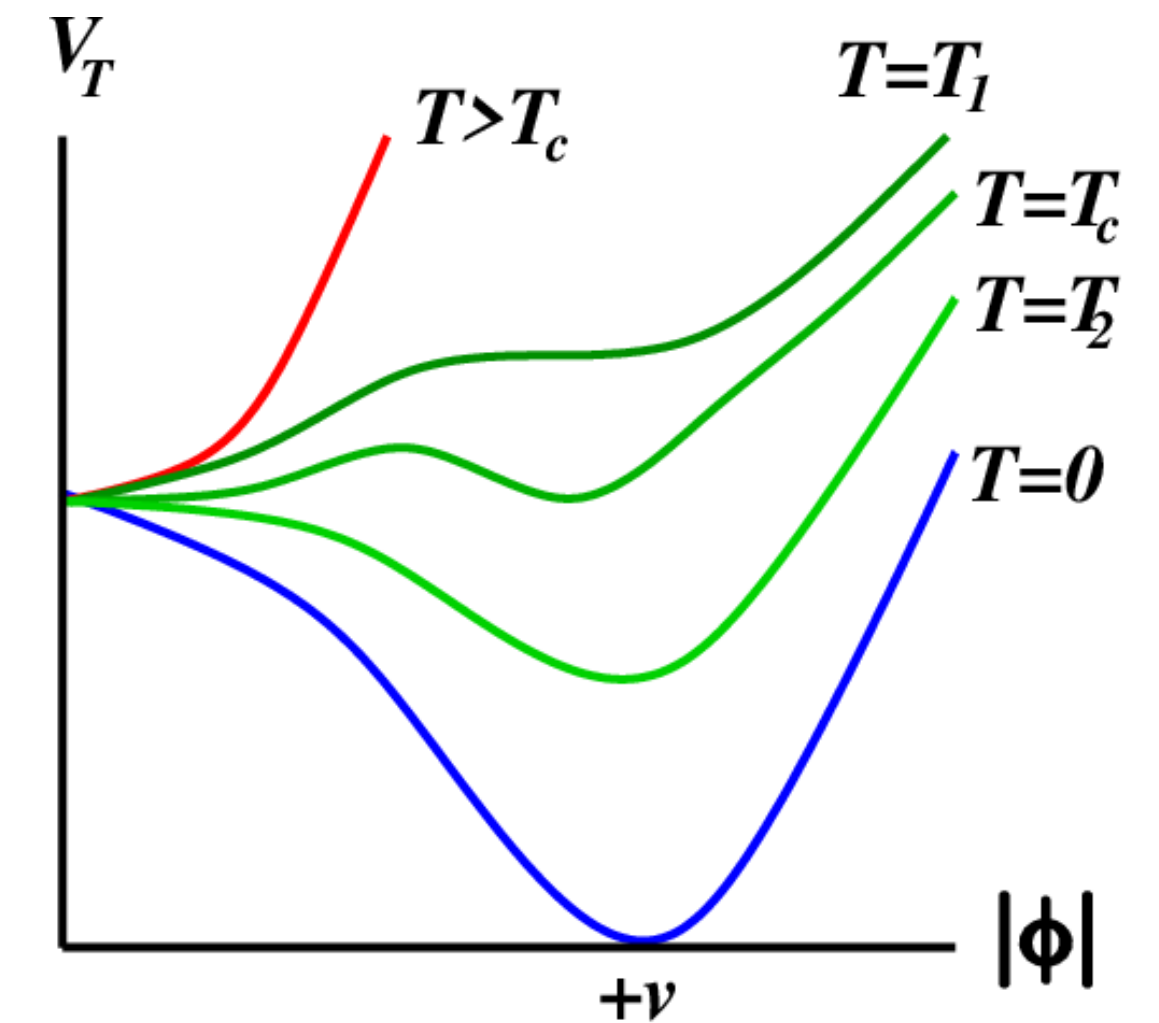
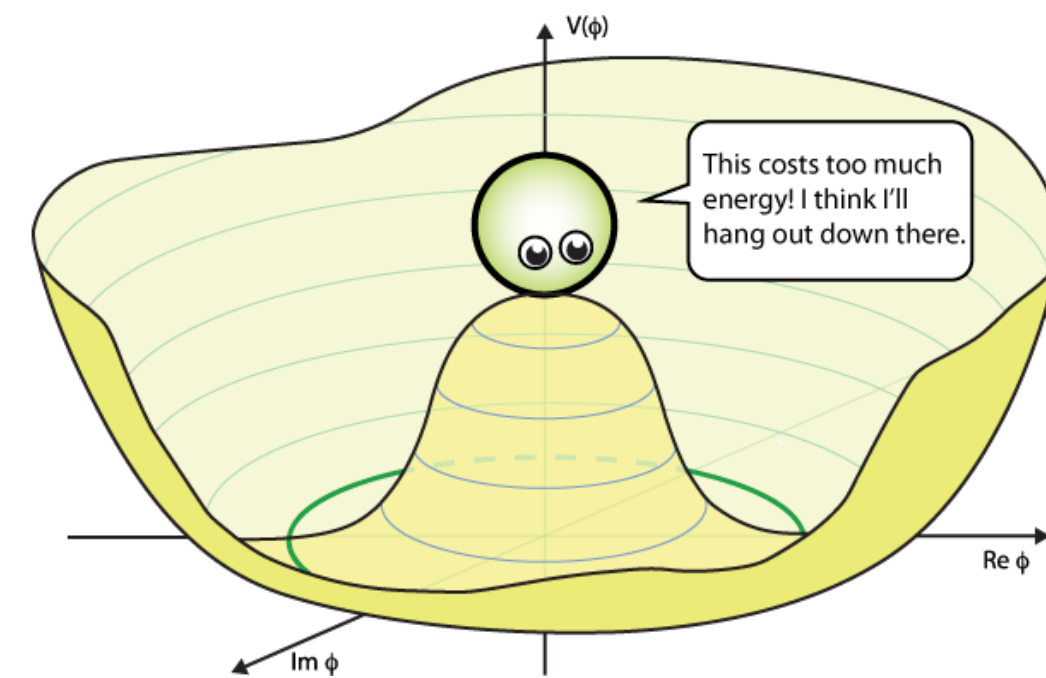
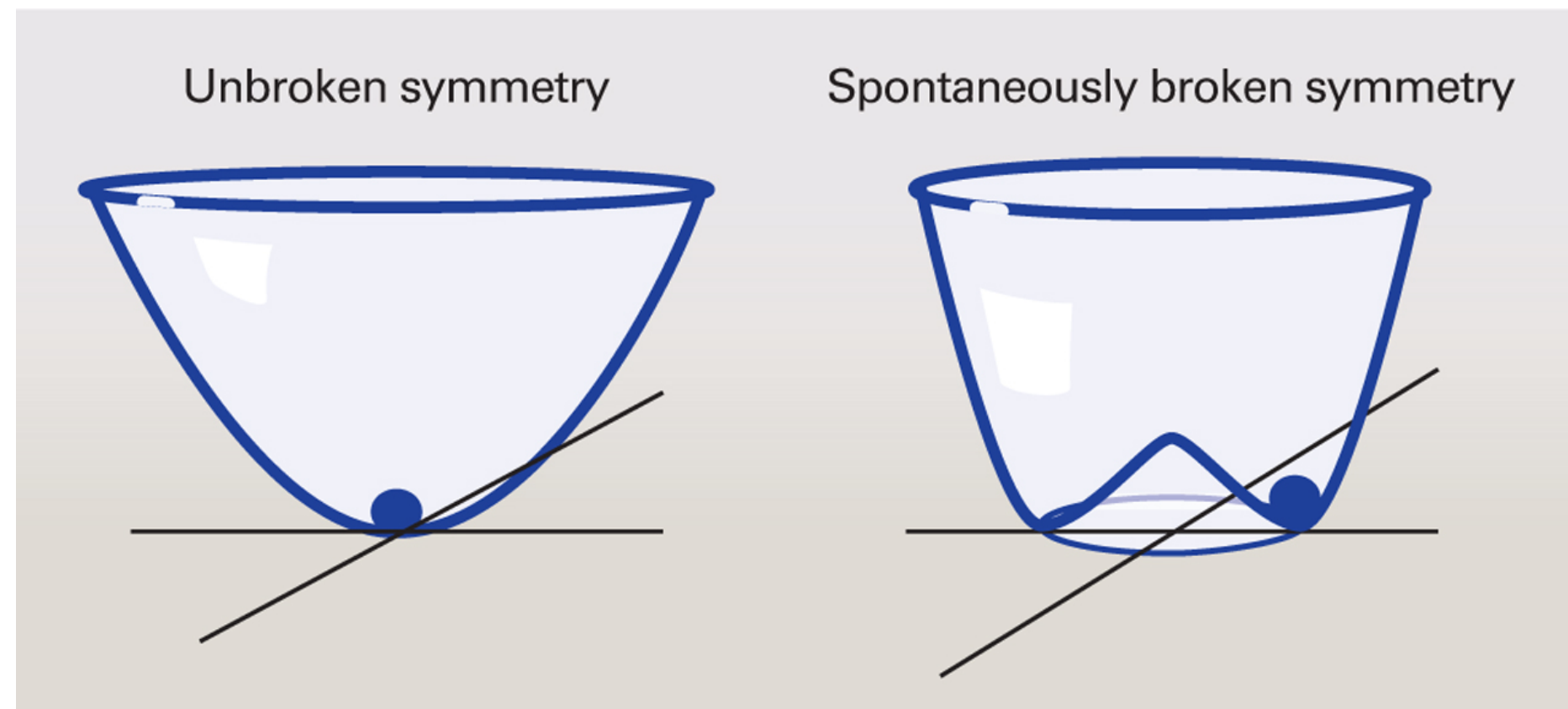
Electroweak symmetry restored temperature: $T = 160 \text{ GeV}$



$$\Gamma_{sph}^{brok}(T) = \kappa_{brok} \alpha_W^4 T^4 \exp\left(-\frac{E_{sph}}{T}\right) \quad (4)$$

Symmetry non-restoration

- EW symmetry non-restoration!**



$$V = -\frac{\mu^2}{2}h^2 + \frac{\lambda}{4}h^4$$

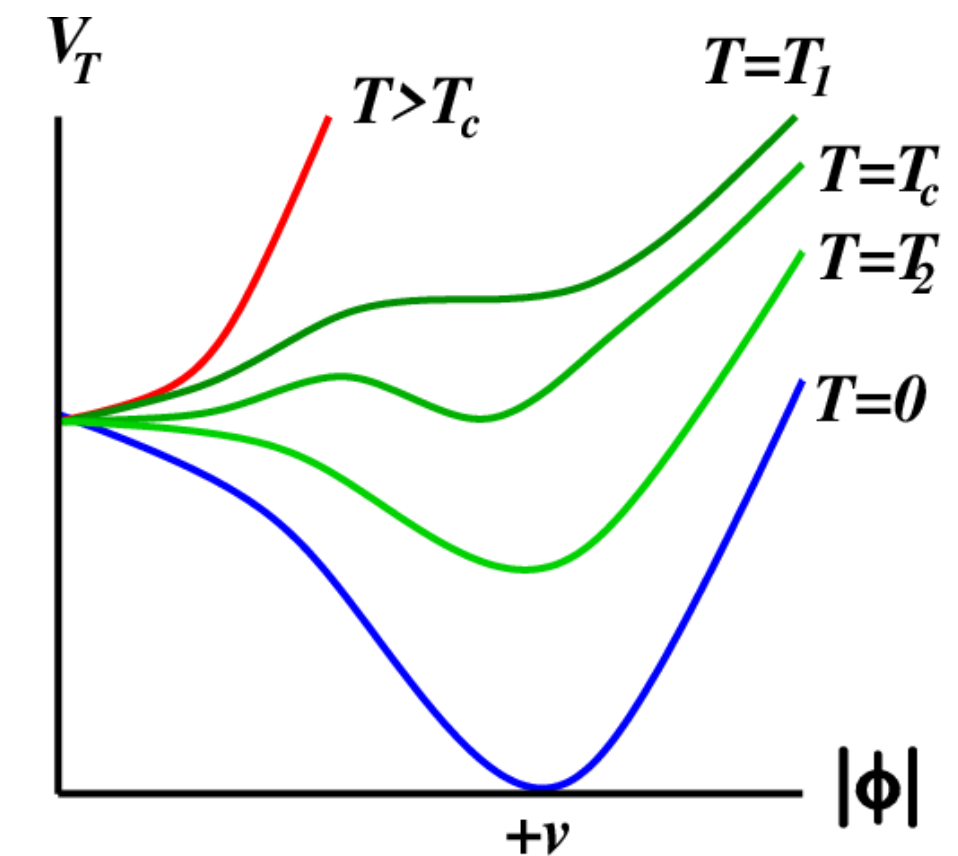
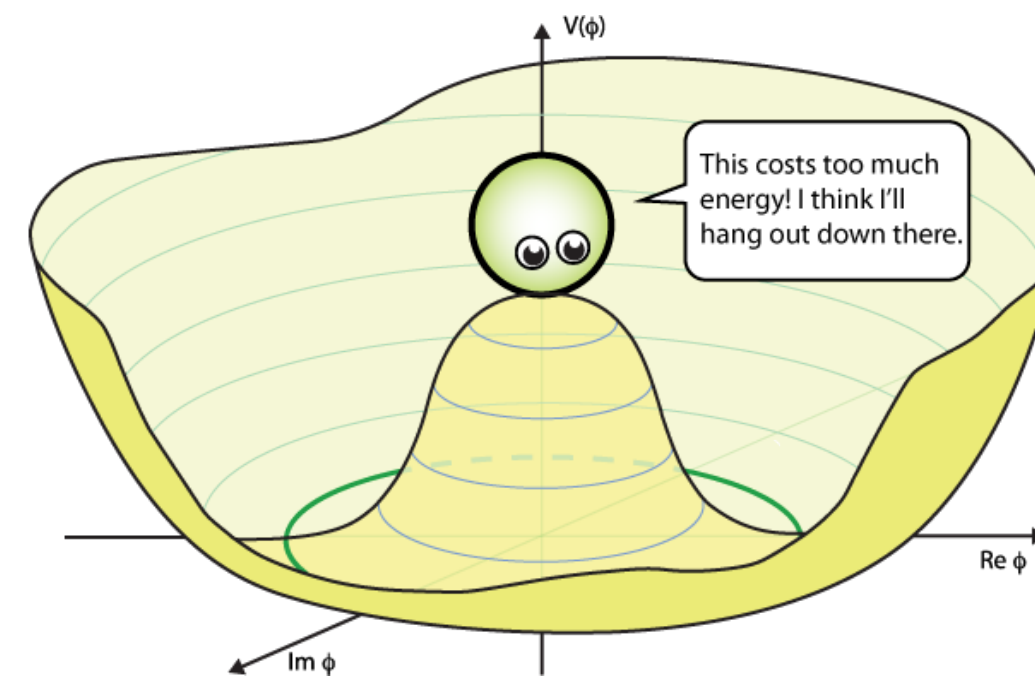
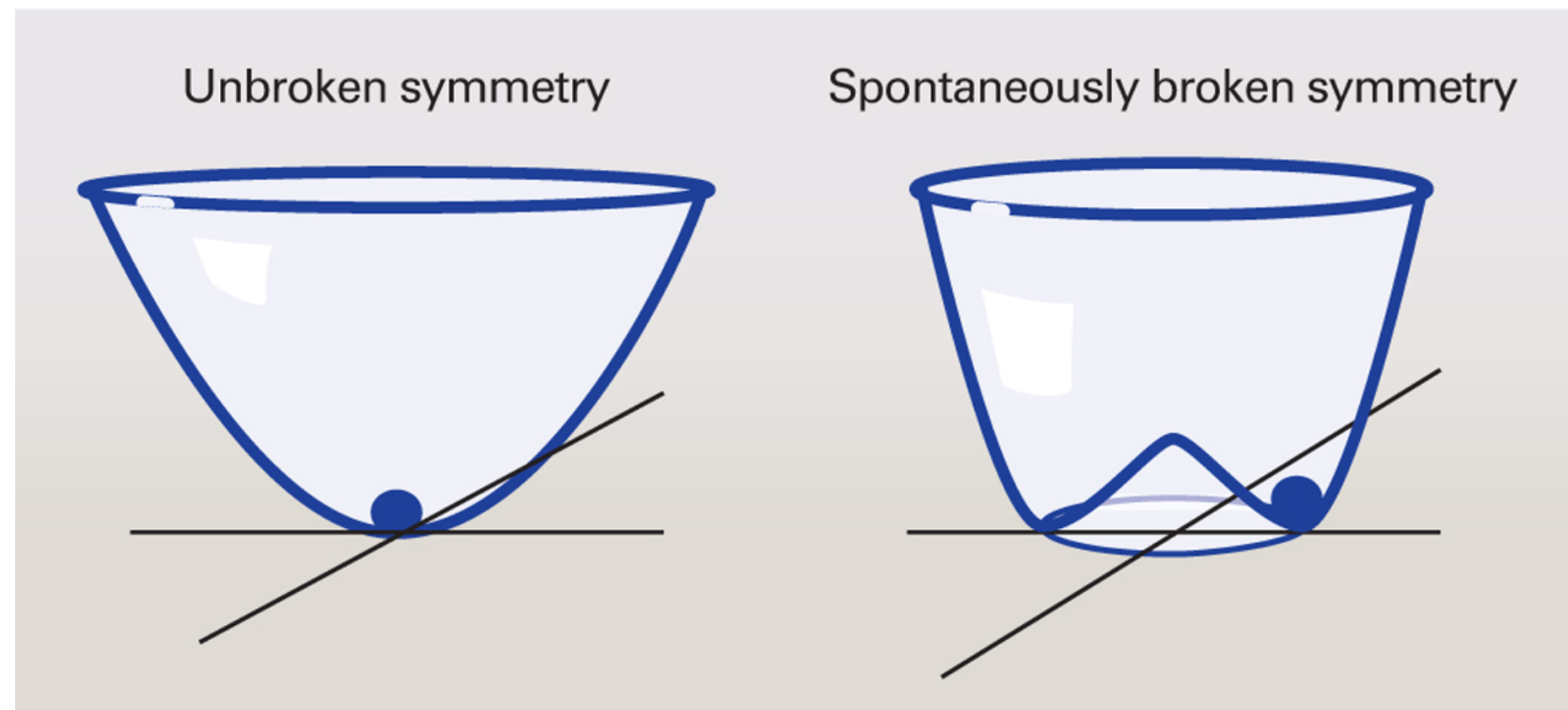
Tree-level

$$\Delta V_1(\phi, T) = \sum_F \frac{g_F T^4}{2\pi^2} \left[\sum_{n=1}^{\infty} \frac{(-1)^n}{n^2} (\beta m_F)^2 K_2(\beta m_F n) \right] - \sum_B \frac{g_B T^4}{2\pi^2} \left[\sum_{n=1}^{\infty} \frac{1}{n^2} (\beta m_B)^2 K_2(\beta m_B n) \right].$$

Thermal corrections

Symmetry non-restoration

- **EW symmetry non-restoration? The answer is **Yes!****



$$V = -\frac{\mu^2}{2}h^2 + \frac{\lambda}{4}h^4$$

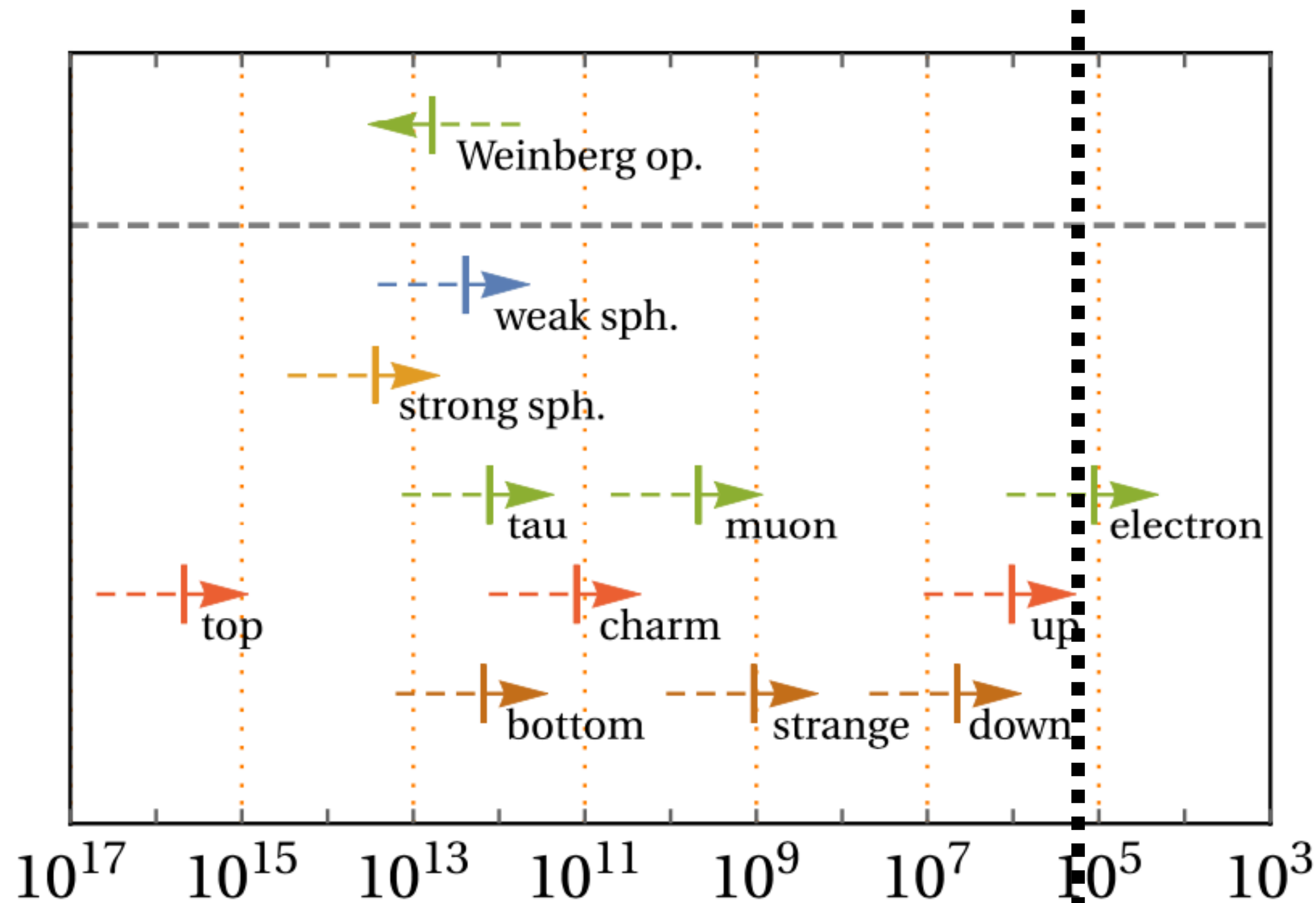
Tree-level

$$\Pi_h = T^2 \left(\frac{\lambda_t^2}{4} + \frac{3g^2}{16} + \frac{g'^2}{16} + \frac{\lambda}{2} + N_s \frac{\lambda_{hs}}{12} \right) \quad (3)$$

BSM corrections can be negative!

Symmetry non-restoration

- **EW symmetry non-restoration? The answer is Yes!**



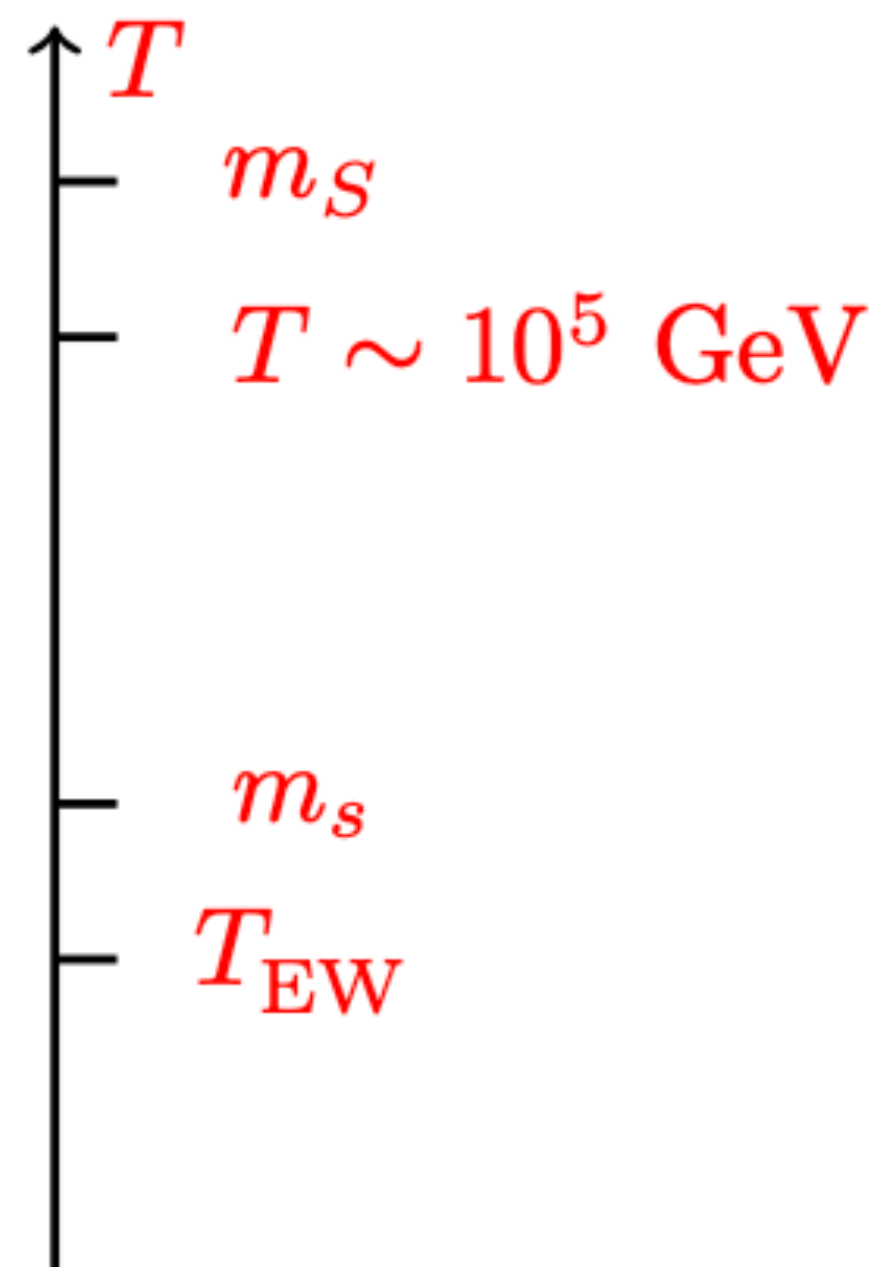
Sphaleron quenches at 10^5 GeV.

Electron asymmetry can be transported to the BAU.

No B-L violation is needed!

Symmetry non-restoration

- **EW symmetry non-restoration? The answer is Yes!**



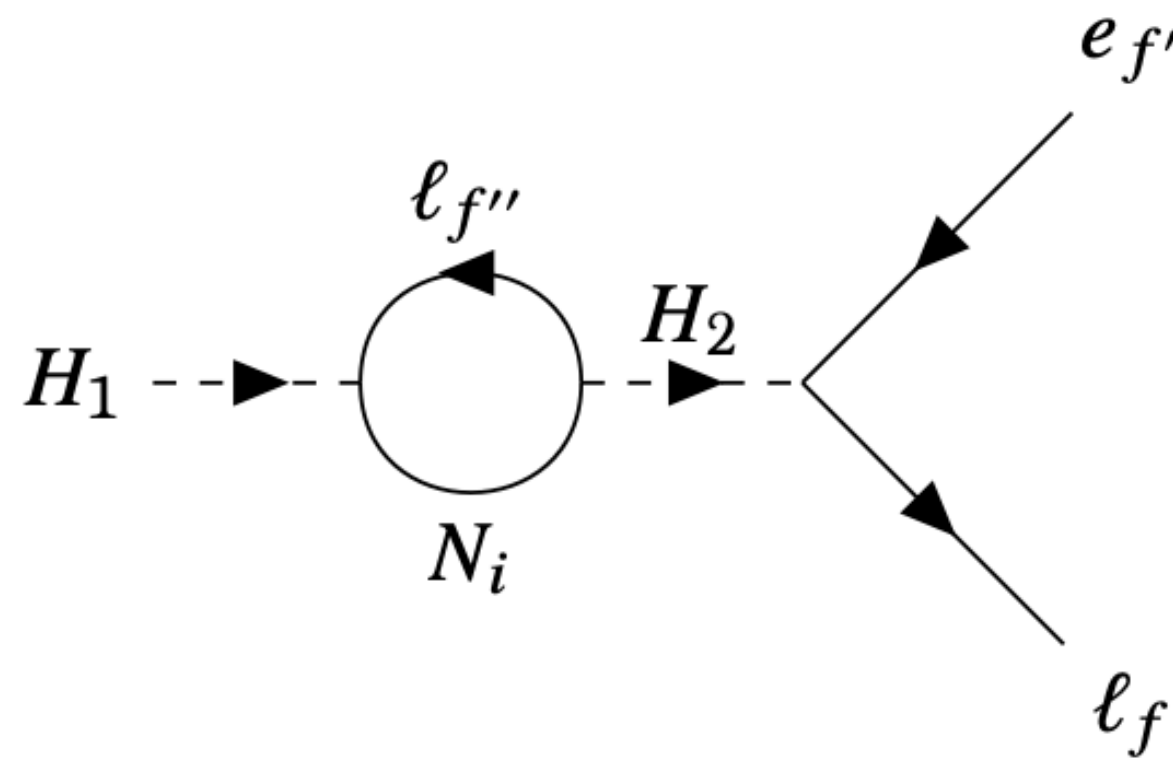
$$\Gamma_{\text{sph}}^{\text{brok}}(T) = \kappa_{\text{brok}} \alpha_W^4 T^4 \exp\left(-\frac{E_{\text{sph}}}{T}\right) \quad E_{\text{sph}} > T \log\left(\frac{\kappa_{\text{brok}} \alpha_W^4 M_P}{3T}\right)$$

$$E_{\text{sph}} = \frac{4\pi v}{g} \int_0^\infty d\xi \left[4(f')^2 + \frac{8}{\xi^2} f^2(1-f)^2 + \frac{\xi^2}{2} (h')^2 + h^2(1-f)^2 + \frac{\xi^2}{16} \sigma^2 (h^2 - 1)^2 \right]$$

$$\rightarrow N_s \lambda_{hs} < -4.82$$

Eogenesis

- **Higgs decay into chiral electrons**



$$\varepsilon = \frac{1}{8\pi} \frac{\text{Im} \left[Y_{fg}^1 Y_{gf}^{2\dagger} \text{tr}(Y^{1\dagger} Y^2) \right]}{\text{tr}(Y^{1\dagger} Y^1)} f \left(\frac{M_{\Phi_2}}{M_{\Phi_1}} \right)$$

- **Chiral asymmetries is generated.**
- **No primordial B-L violation is generated!**

Eogenesis

- **Transport equations**

$$-\frac{dY_{\Sigma}}{d \ln T} = -\frac{\gamma_D}{H} \left[Y_{\Sigma} - Y_{\Sigma}^{\text{EQ}} \right]$$

$$-\frac{d}{d \ln T} \left(\frac{\mu_{L_k}}{T} \right) = -\frac{1}{g_{L_k}} \frac{\gamma_{WS}}{H} \left[\sum_{i=1}^3 \left(\frac{\mu_{L_i}}{T} + 3 \frac{\mu_{Q_i}}{T} \right) \right]$$

$$-\frac{d}{d \ln T} \left(\frac{\mu_{\Phi}}{T} \right) = -2 \frac{\gamma_D}{H} \left(\frac{\mu_{\Phi}}{T} - \frac{4}{3} \frac{\mu_{L_k}}{T} \right)$$

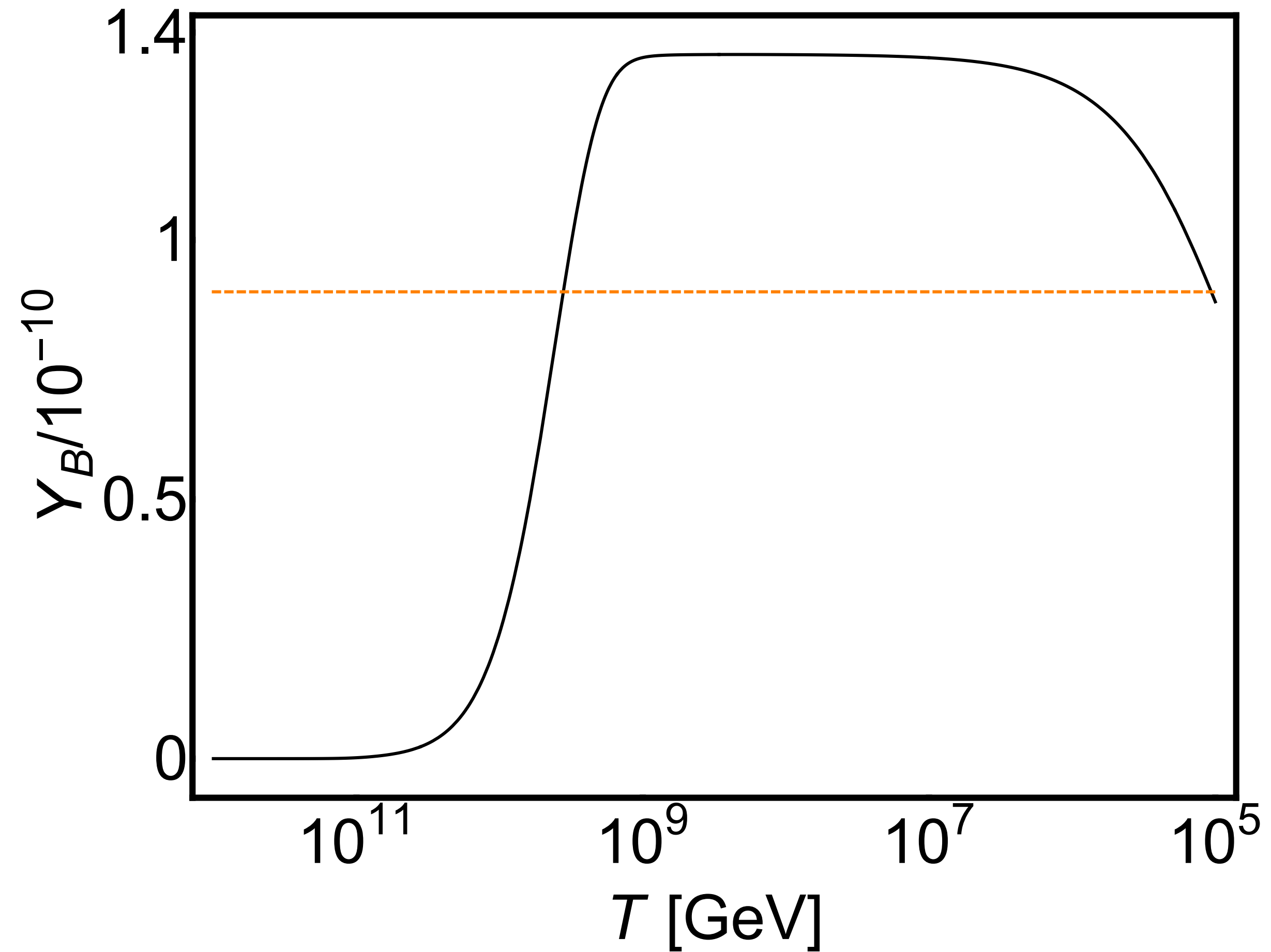
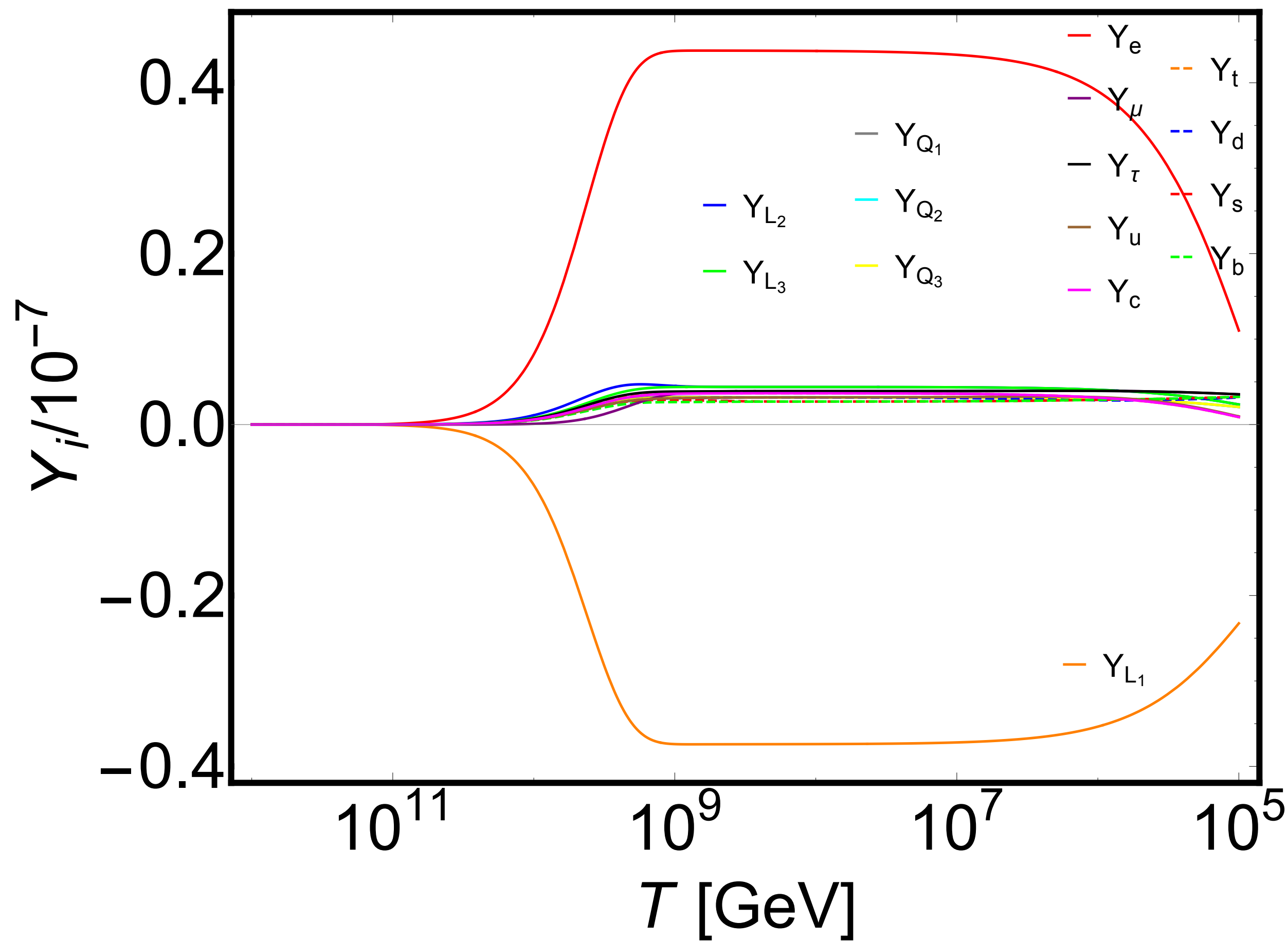
$$\gamma_D = \frac{K_1(z)}{K_2(z)} \Gamma_{\Phi}$$

$$-\frac{1}{g_{L_k}} \frac{\gamma_{Y_{E_k}}}{H} \left(-\frac{\mu_{E_k}}{T} + \frac{\mu_{L_k}}{T} - \frac{\mu_H}{T} \right)$$

$$+\frac{1}{g_{L_k}} \frac{\gamma_D}{H} \varepsilon \left[\frac{4\pi^2 g_*^S}{15} Y_{\Sigma} - \frac{48\zeta(3)}{\pi^2} \right]$$

$$-\frac{2}{g_{L_k}} g_{\Phi} \frac{\gamma_D}{H} \left(\frac{4}{3} \frac{\mu_{L_k}}{T} - \frac{\mu_{\Phi}}{T} \right)$$

Eogenesis



Conclusion

A new type of Leptogenesis is proposed, which is dubbed
as **Eogenesis**

- No primordial B-L is needed
- Not depends on the nature of neutrinos

Thank you for your attention!