



中山大學 物理与天文学院
SUN YAT-SEN UNIVERSITY SCHOOL OF PHYSICS AND ASTRONOMY

Indirect probes of new physics with non-standard interactions of leptons

李刚

中山大学物理与天文学院

GL, Chuan-Qiang Song, Feng-Jie Tang, Jiang-Hao Yu, 2409.04703
Yuxuan He, GL, Jia Liu, Xiao-Ping Wang, Xiang Zhao, 2407.06523

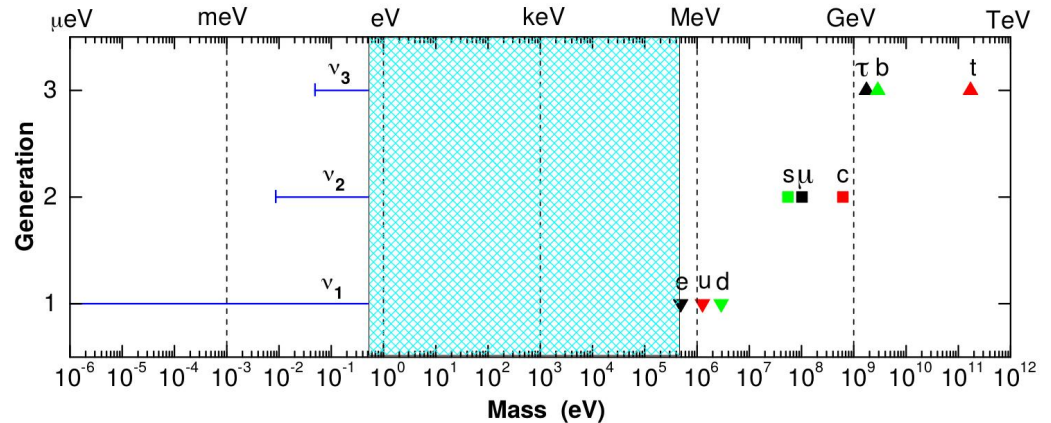
第三届高能物理理论与实验融合发展研讨会

辽宁师范大学, 2024年11月3日

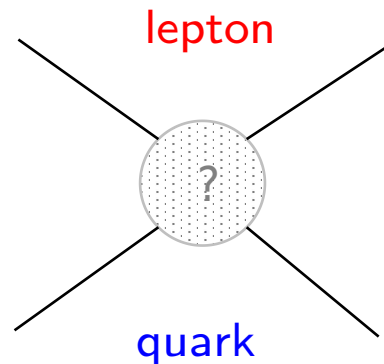
New physics

The Standard Model is successful but incomplete:

- neutrino masses
- baryon asymmetry
- dark matter
- strong CP problem
- flavor structure
-



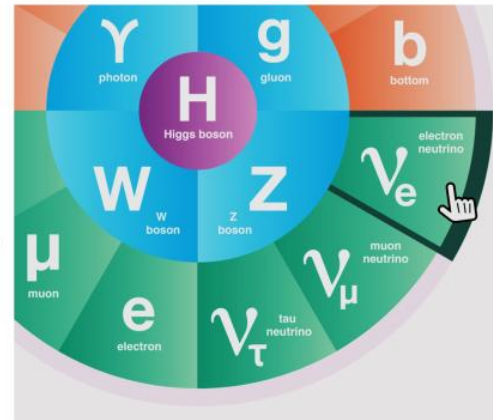
Z.-z. Xing, Phys.Rept. 854 (2020) 1-147



Neutrino physics

Open questions:

- Normal or Inverted (sign of Δm_{31}^2 ?)
- Leptonic CP Violation ($\delta = ?$)
- Octant of θ_{23} ($>$ or $<$ 45° ?)
- Absolute Neutrino Masses ($m_{\text{lightest}} = 0$?)
- Majorana or Dirac Nature ($\nu = \nu^c$?)
- Majorana CP-Violating Phases (how?)



- Extra Neutrino Species
- Exotic Neutrino Interactions
- Various LNV & LFV Processes
- Leptonic Unitarity Violation

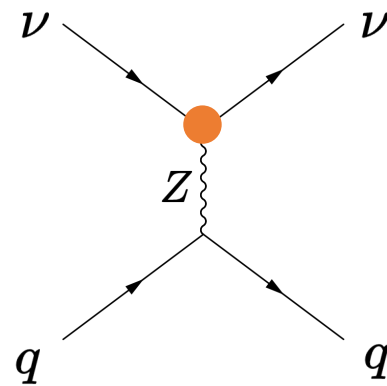
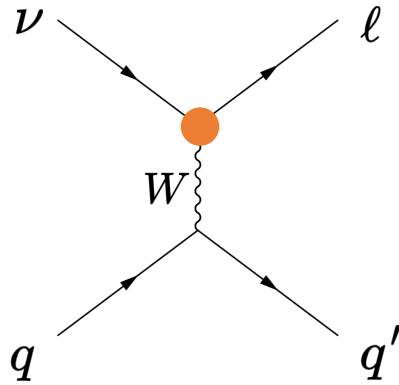


- Origin of Neutrino Masses
- Flavor Structure (Symmetry?)
- Quark-Lepton Connection
- Relations to DM and/or BAU

credit: Shun Zhou

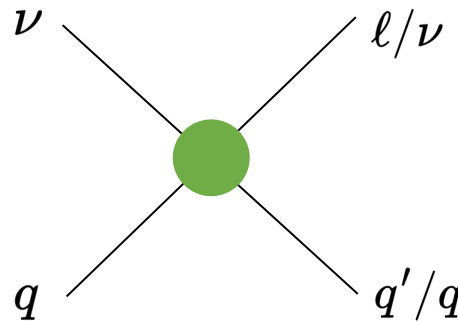
Neutrino Interactions

Neutrino Non-Standard Interactions:



or other contributions
from BSM particles
(W' , Z' , new scalar,
leptoquark)

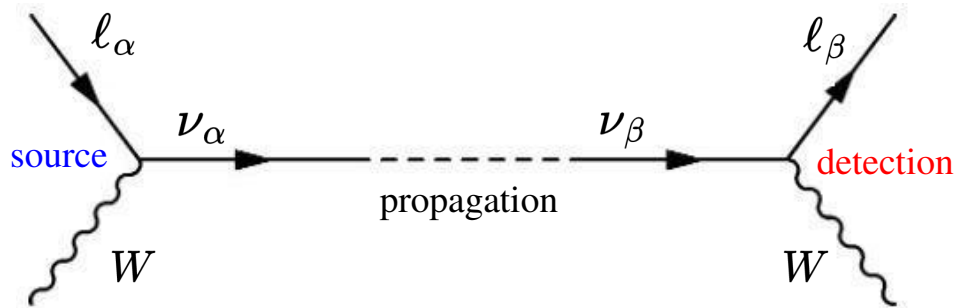
$\mu < M_W, M_Z$



**semileptonic four-fermion
operators**

Neutrino Interactions

Charged-current interactions:



Neutrino CC NSIs affect the **source (production)** and **detection (scattering)** of neutrinos in oscillation experiments

$$\mathcal{L}_{CC} \supset -2\sqrt{2}G_F V_{ud}^{\text{SM}} \left\{ [\mathbf{1} + \epsilon_L]_{\alpha\beta}^{ij} (\bar{u}_i \gamma^\mu P_L d_j) (\bar{\ell}_\alpha \gamma_\mu P_L \nu_\beta) + [\epsilon_R]_{\alpha\beta}^{ij} (\bar{u}_i \gamma^\mu P_R d_j) (\bar{\ell}_\alpha \gamma_\mu P_L \nu_\beta) \right. \\ \left. + \frac{1}{2} [\epsilon_S]_{\alpha\beta}^{ij} (\bar{u}_i d_j) (\bar{\ell}_\alpha P_L \nu_\beta) - \frac{1}{2} [\epsilon_P]_{\alpha\beta}^{ij} (\bar{u}_i \gamma_5 d_j) (\bar{\ell}_\alpha P_L \nu_\beta) + \frac{1}{4} [\epsilon_T]_{\alpha\beta}^{ij} (\bar{u}_i \sigma^{\mu\nu} P_L d_j) (\bar{\ell}_\alpha \sigma_{\mu\nu} P_L \nu_\beta) + \text{H.c.} \right\}$$

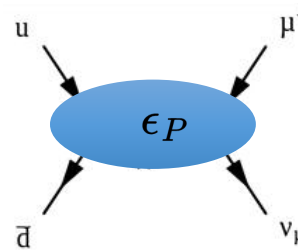
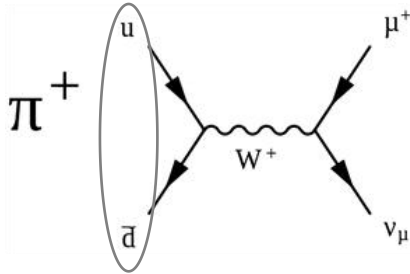
Neutrino Interactions

Neutrino production:

- Neutrinos from pion decays

$$J_{\pi\mu} = \frac{m_\pi^2}{m_\mu (m_u + m_d)} \sim 20$$

$$\delta\Gamma (\pi^+ \rightarrow \mu^+ \nu_\mu) \simeq \frac{(m_\pi^2 - m_\mu^2)^2 m_\mu^2}{64\pi m_\pi^3} f_\pi^2 \left| J_{\pi\mu} \left(\frac{V_{ud}}{v^2} [\epsilon_P]_{22}^{11} \right) \right|^2$$



Neutrino Interactions

Charged-current interactions:

- Chirality-flip four-fermion operator

$$O_{ledq}^{\alpha\beta 11} = (\bar{L}_\alpha^j e_{R\beta}) (\bar{d}_R Q^j)$$

$$\frac{C_{ledq}^{2211*}}{\Lambda^2} = \frac{-2V_{ud}}{v^2} [\epsilon_P]_{22}^{11}$$



	T2HK limit	DUNE limit	JUNO limit	T2HK and DUNE limit	JUNO and TAO limit
Operator (TeV)	(TeV)	(TeV)	(TeV)	(TeV)	(TeV)
$\mathcal{O}_{ledq2211}$	9.1	11.2	0.7	12.3	0.7

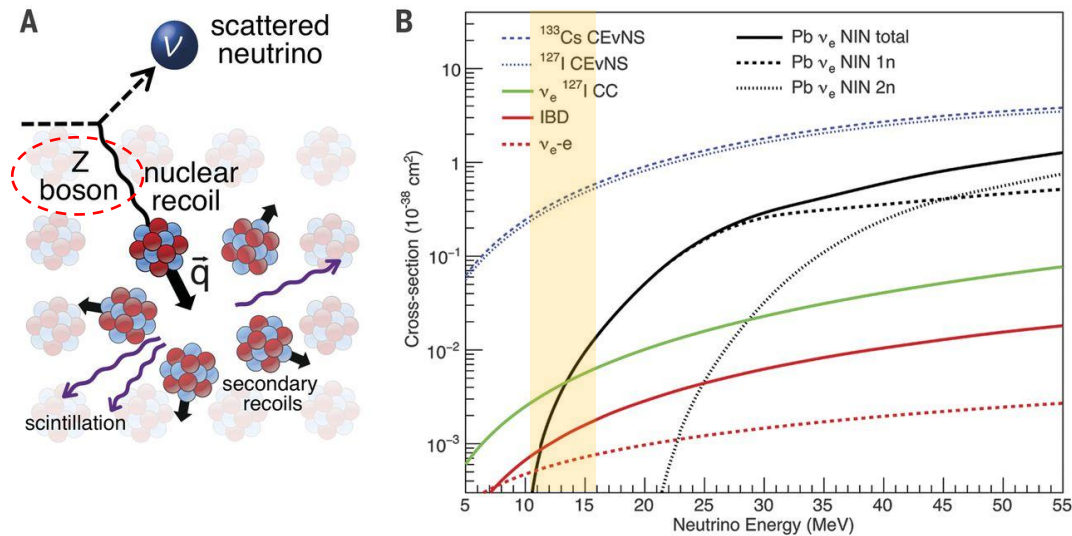
Best sensitivity among the flavor-conserving semileptonic four-fermion operators

Y. Du, H.-L. Li, J. Tang, S. Vihonen, J.-H. Yu 2106.15800 (PRD)

For the study of chirality-flip four-fermion operator at the EIC, see H.-L. Wang, X.-K. Wen, H. Xing, B. Yan, 2401.08419 (PRD)

Neutrino Interactions

Neutral-current interactions:



Science 357 (2017) 6356, 1123

Coherent Neutrino-nucleus scattering (CEvNS)

Neutrino Interactions

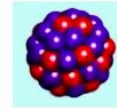
Neutrino scattering:

CEvNS: what's it good for? ① So ② Many ③ Things ! (not a complete list!)

CEvNS as a **signal**
for signatures of *new physics*



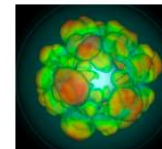
CEvNS as a **signal**
for understanding of “old” physics



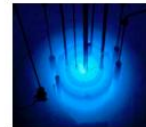
CEvNS as a **background**
for signatures of new physics



CEvNS as a **signal** for *astrophysics*



CEvNS as a **practical tool**

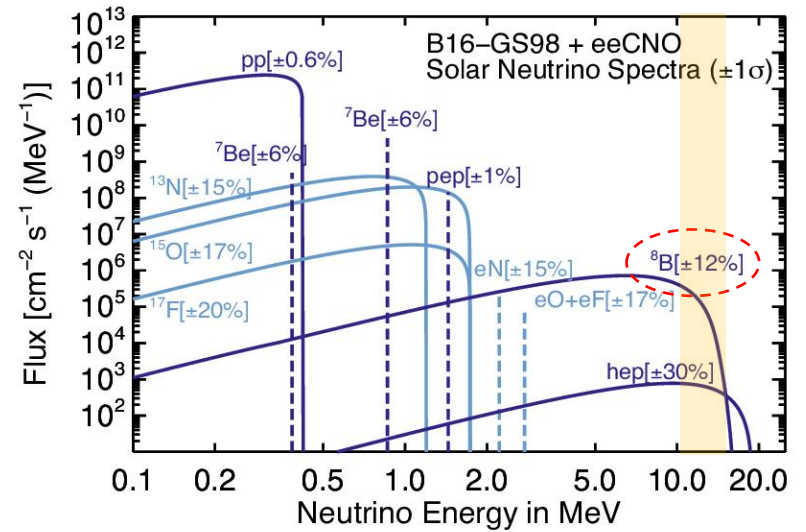
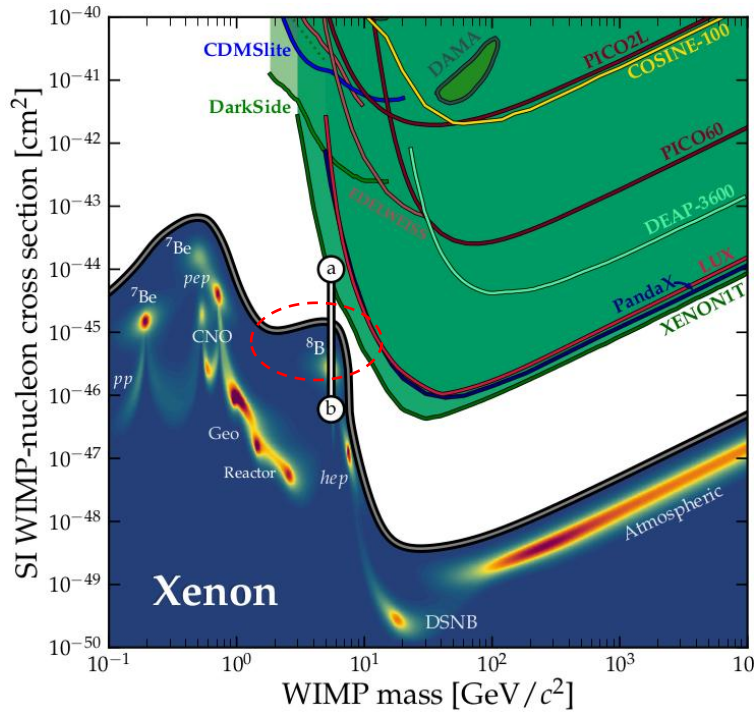


credit: Kate Scholberg

Neutrino NC NSIs affect the CEvNS

Neutrino Interactions

Neutrino scattering:

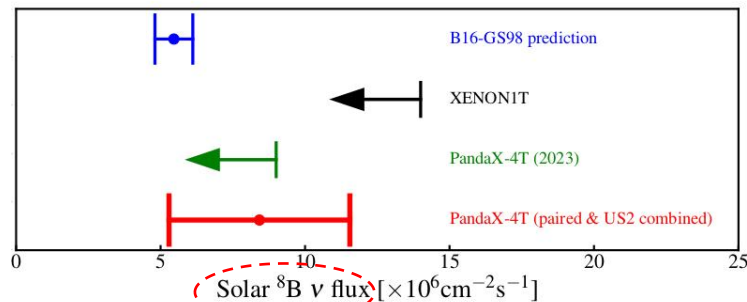


neutrino floor/frog

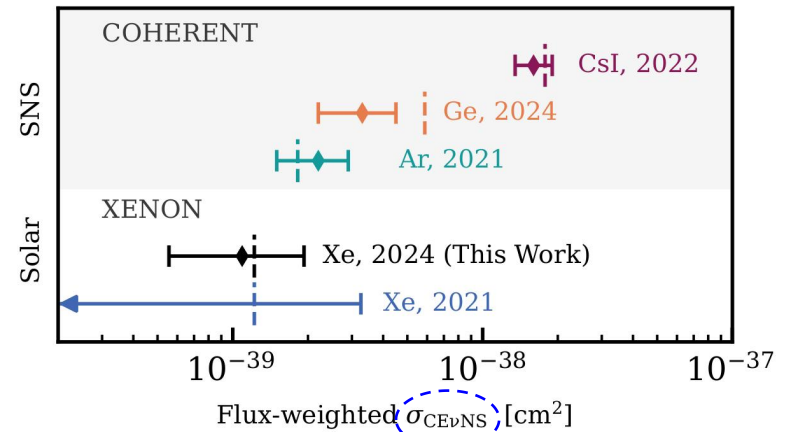
C. A. J. O'Hare, 2109.03116 (PRL)
 J. Tang, B.-L. Zhang, 2304.13665 (PRD)

Neutrino Interactions

Neutrino scattering:



PandaX, 2407.10892 (PRL)



XENONnT, 2408.02877

Number of signal events

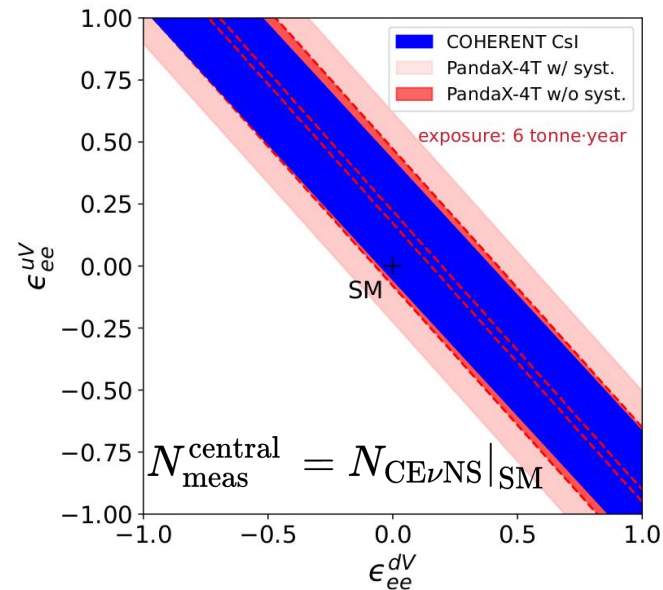
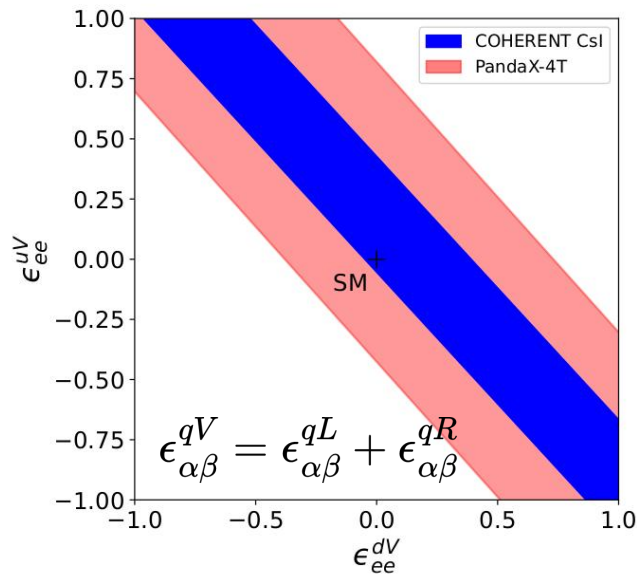
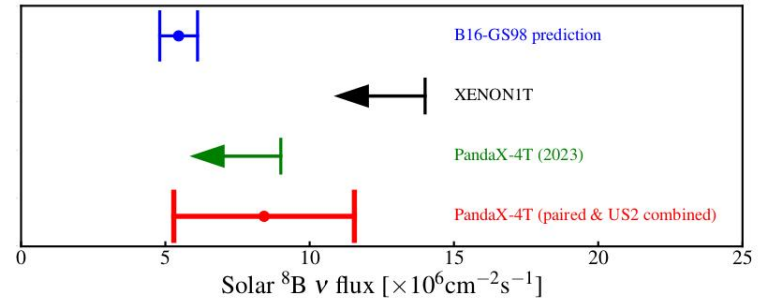
= solar ^8B neutrino flux \otimes CEvNS cross section

Neutrino NC NSIs

Neutrino Interactions

Neutral-current interactions:

$$\mathcal{L}_{\text{NC}} \supset -2\sqrt{2}G_F \left[\epsilon_{\alpha\beta}^{qL} (\bar{\nu}_\alpha \gamma^\mu P_L \nu_\beta) (\bar{q} \gamma_\mu P_L q) + \epsilon_{\alpha\beta}^{qR} (\bar{\nu}_\alpha \gamma^\mu P_L \nu_\beta) (\bar{q} \gamma_\mu P_R q) \right]$$



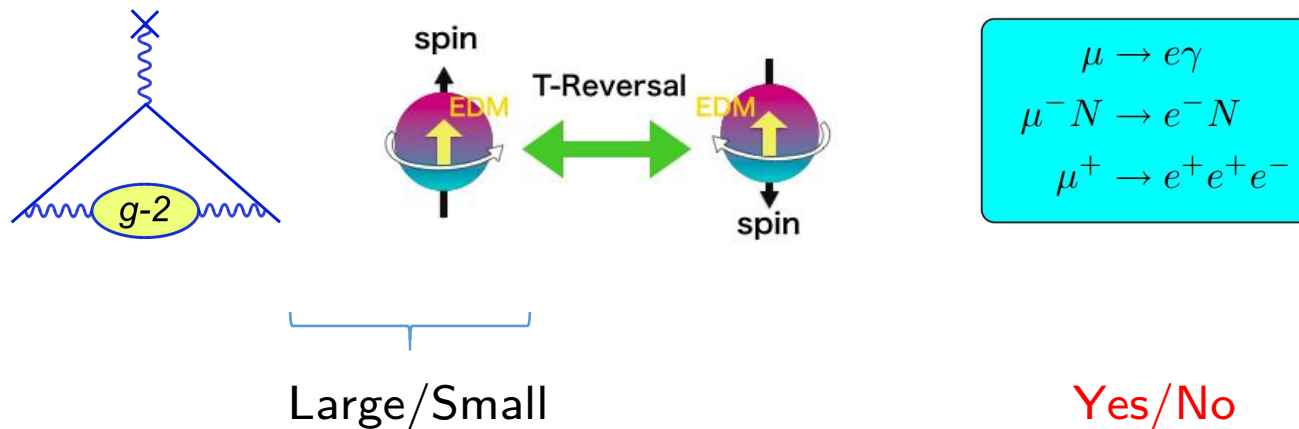
GL, C.-Q. Song, F.-J. Tang, J.-H. Yu, 2409.04703

Charged lepton physics

Open questions:

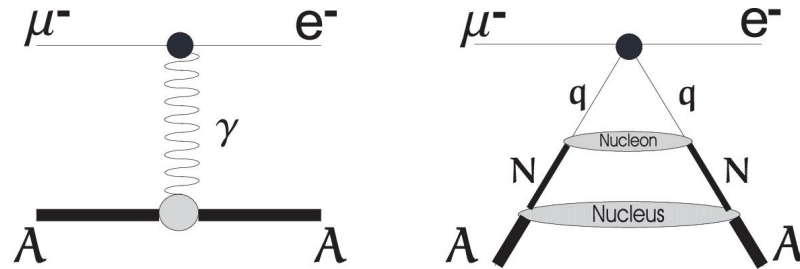
- muon anomalous magnetic moment ($g-2$)
- electron electric dipole moment (EDM)
- charged lepton flavor violation (CLFV)
- ...

see also Qiang Li's talk



CLFV Interactions

$\mu \rightarrow e$ conversion:



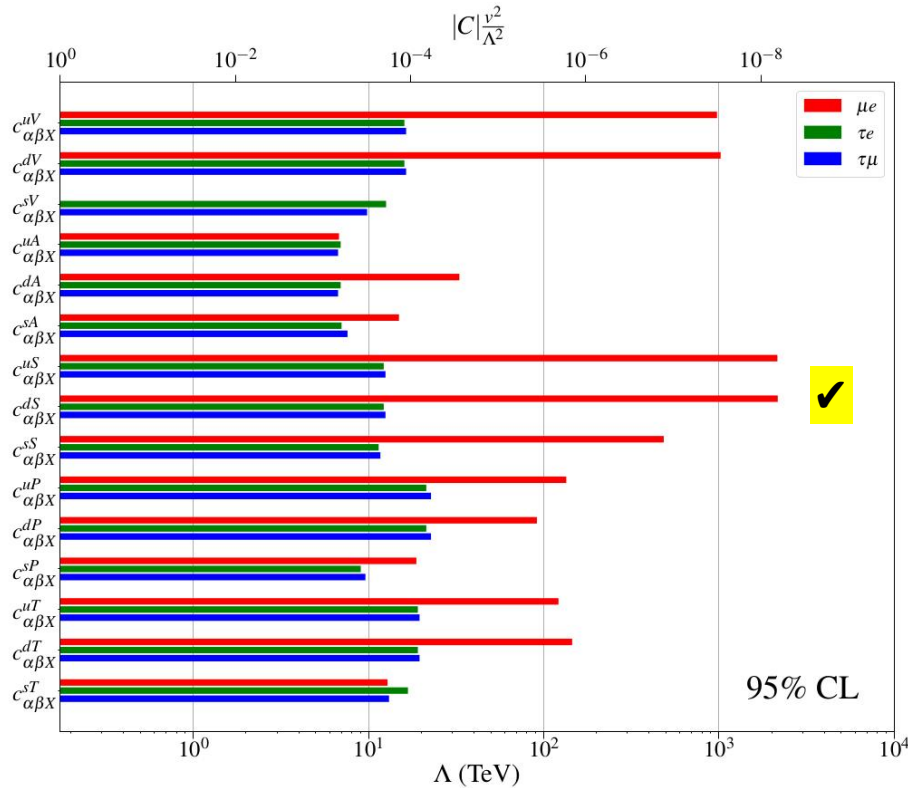
$$\text{CR}(\mu^- + (A, Z) \rightarrow e^- + (A, Z)) \equiv \frac{\Gamma(\mu^- + (A, Z) \rightarrow e^- + (A, Z))}{\Gamma(\mu^- + (A, Z) \rightarrow \text{capture})}$$

cLFV obs.	Present upper bounds (90% CL)	
$\text{CR}(\mu \rightarrow e, \text{S})$	7.0×10^{-11}	Badertscher <i>et al.</i> (1982)
$\text{CR}(\mu \rightarrow e, \text{Ti})$	4.3×10^{-12}	SINDRUM II (1993)
$\text{CR}(\mu \rightarrow e, \text{Pb})$	4.6×10^{-11}	SINDRUM II (1996)
✓ $\text{CR}(\mu \rightarrow e, \text{Au})$	7.0×10^{-13}	SINDRUM II (2006)

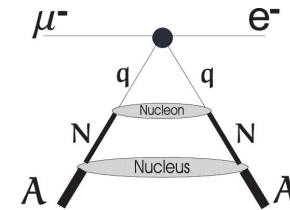
New experiments to start: COMET, Mu2e

CLFV Interactions

$\mu \rightarrow e$ conversion to probe semileptonic four-fermion operators



Fernández-Martínez, et al., 2403.09772 (EPJC)



Chirality-flip four-fermion operator:

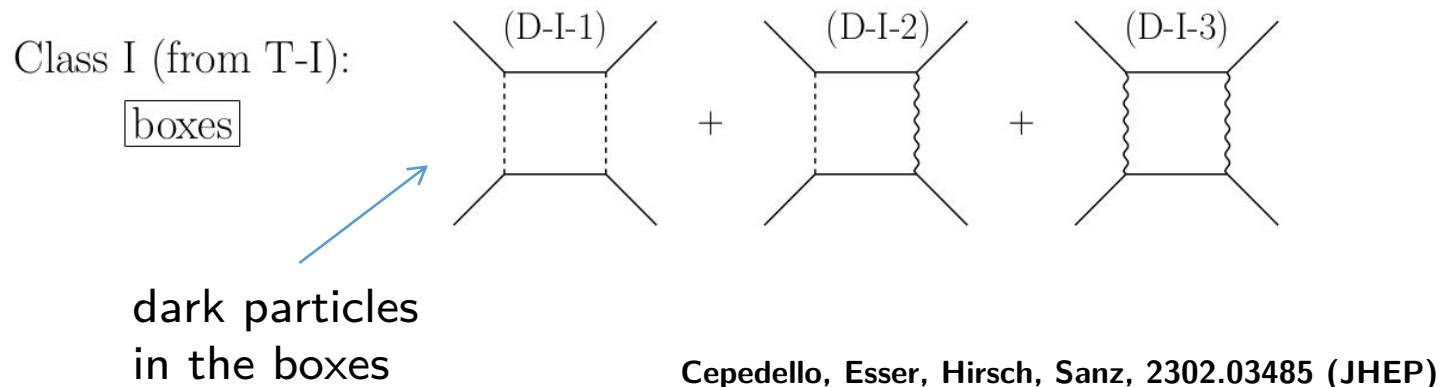
$$O_{ledq}^{\alpha\beta 11} = (\bar{L}_\alpha^j e_{R\beta}) (\bar{d}_R Q^j)$$

$$c_{\alpha\beta R}^{dS} = \frac{v^2}{2\Lambda^2} C_{ledq}^{\alpha\beta 11}$$

$$C_{ledq}^{1211} / \Lambda^2 < (2.2 \times 10^3 \text{ TeV})^{-2}$$

CLFV Interactions

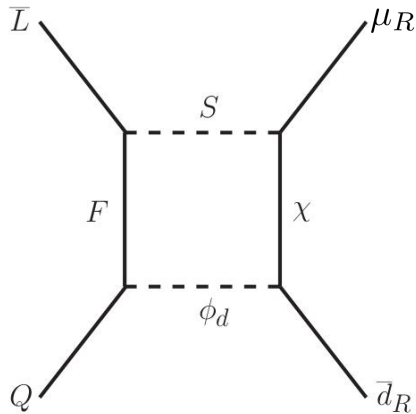
- Given the severe CLFV constraints, is there *natural* scenario in which new physics is expected?
- Dark loop paradigm:



The dark symmetry guarantees that **CLFV is naturally suppressed**, which is forbidden at tree-level

Fermion portal dark matter

One-loop realization of the semileptonic four-fermion operator:



new fields	SU(3) _C	SU(2) _L	U(1) _Y	Z ₂
χ	1	1	0	-1
F	1	2	$\frac{1}{2}$	-1
S	1	1	1	-1
ϕ_d	3	1	$-\frac{1}{3}$	-1

$$\mathcal{L} = f_{LS} (\bar{L} F_R) S^* + f_{\chi S} (\bar{\chi}_L \mu_R) S + f_{FQ} (\bar{F}_R Q) \phi_d^* + f_{d\chi} (\bar{d}_R \chi_L) \phi_d + \text{h.c.}$$

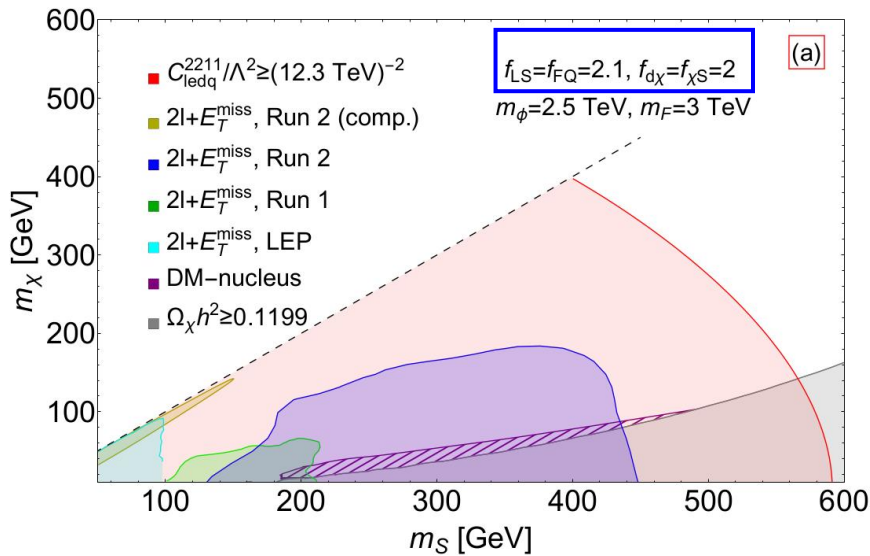
Majorana DM: χ
mediators: S, ϕ_d
lepton: F

$$O_{ledq}^{\alpha\beta 11} = (\bar{L}_\alpha^j e_{R\beta}) (\bar{d}_R Q^j)$$

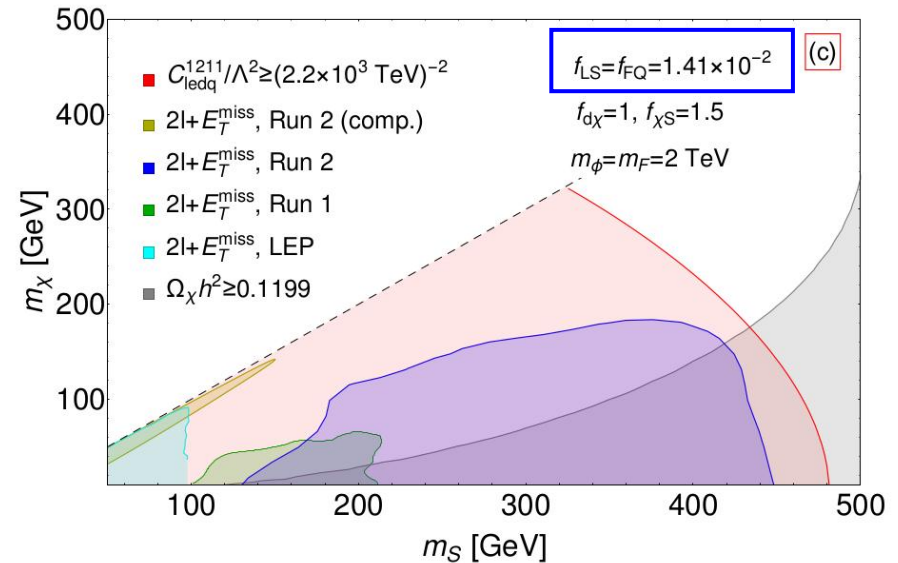
Cepedello, Esser, Hirsch, Sanz, 2302.03485 (JHEP)
An, Wang and Zhang, 1308.0592 (PRD)
Bai, Berger, 1308.0612 (JHEP); 1402.6696 (JHEP)
DiFranzo, Nagao, Rajaraman, Tait, 1308.2679 (JHEP)

Fermion portal dark matter

Indirect probes of dark matter with the NSIs of leptons:



Red: Neutrino CC NSI



Red: CLFV interaction

Y. He, GL, J.Liu, X.-P. Wang, X. Zhao, 2407.06523

Summary

- Non-standard interactions (NSIs) of leptons are sensitive to new physics
- We investigate constraint on the **neutrino NC NSI** using recent measurement of solar ${}^8\text{B}$ neutrinos via CEvNS by PandaX
- We show that both the **neutrino CC NSI** and **charged-lepton-flavor-violation interaction** can effectively probe dark matter

Thank you