

Revealing the Origin of Neutrino Masses through Displaced Shower Searches in the CMS Muon System

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Neutrino Mass



The masses of the neutrinos are much smaller than the other fermions!

Neutrinos are the only neutral fermions.

The nature of the neutrino masses can be different!

Figure from Hitoshi Murayama

Type-I Seesaw Mechanisms









$$L_{\text{Dirac}} = -y\overline{I_L}\widetilde{H}v_R$$

Tiny Yukawa coupling $\sim 10^{-12}$ for pure Dirac mass! Can the neutrino masses induced by both?

$$\begin{pmatrix} 0 & M_D \\ M_D & M_R \end{pmatrix}$$
 $m_1 \approx -\frac{M_D^2}{M_R} \sim eV$ $m_2 \approx M_R \gg GeV$

Seesaw Mechanism by hand? UMSM

Heavy Neutrino Searches

• GUT points out $M_N \approx \frac{M_D^2}{m_\nu} \sim 10^{14} \text{GeV}$

- Leptogenesis favors $M_N \sim 10^9 \text{ GeV}$
- Can we still probe heavy neutrinos at colliders?
- Such heavy degree of freedom leads to non-stable Higgs masses, $M_N < 10^7$ GeV.
- Larger CP violation can be produced if heavy neutrinos are degenerate, or the PMNS has CP violations, i.e. resonant leptogenesis or leptogenesis via oscillations. $M_N \sim 10^3 (1)$ GeV.
- The Yukawa couplings can be electron like,

$$y_D \sim 10^{-6}, M_N \sim 100 \text{ GeV}$$

- Inverse seesaw can yield natural Yukawa couplings.
- Colliders probes of heavy neutrinos are resonable!

Heavy Neutrino Searches



The main production processes of the heavy neutrinos at the LHC. Figures from P.S.Bhupal Dev, Apostolos Pilaftsis, Un-ki Yang, Phys.Rev.Lett. 112 (2014) 8, 081801. Our goal is to search the heavy neutrinos colliders!

Within the **UMSM** model, the heavy neutrinos can be produced either via the s and *t* channel.

The *s* channel production from the Drell-Yan processes are the dominant processes for $m_N \le 1$ TeV, otherwise by the *t* channel.

Both cannot reach low $V_{IN}!$ $\propto V_{IN}^{-2}$ $10^7 pb \times V_{IN}^2 < \text{fb}$

Heavy Neutrino Searches



Limits on the seesaw parameters (m_N, V_{IN}) from the collider searches for heavy neutrinos and other probes. Figure from Patrick D. Bolton, Frank F. Deppisch, P.S. Bhupal Dev, JHEP 03 (2020), 170.

We need to look for other production channels!

Our goal is to search the heavy neutrinos and test the type-I seesaw, the origin of the neutrino masses!

However, direct probes of the heavy neutrinos within the **UMSM** model **cannot reach the type-I seesaw** due to the low mixing at the EW scale

$$\begin{bmatrix} V_L \\ V_R \end{bmatrix} = \begin{bmatrix} V_{LL} & V_{RL} \\ V_{LR} & V_{RR} \end{bmatrix} \begin{bmatrix} V \\ N \end{bmatrix}$$

$$V_{IN}^2 \approx rac{m_v}{mN} < rac{eV}{GeV} < 10^{-10}$$

Natural Type-I Seesaw

The Majorana mass terms can be generated more naturally by the spontaneous breaking of a U(1) gauge.

$$1 \supset -y_D \overline{I_L} \widetilde{H} v_R - y_M v_R^{-c} \chi v_R.$$

Therefore, $m_N \approx y_M v_\chi \sim \text{TeV}$.

One of the simplest model UV complete and anomaly free model is the $U(1)_{B-L}$ model, other U(1) models such as $U(1)_{L_{\mu}-L_{\tau}}$ is also possible.

Baryon and Lepton numbers

Accidental symmetries in the SM, can be broken by anomalies

B-L number Anomaly free

$U(1)_{B-L}$ model

 $SU(3)_C \times SU(2)_L \times U(1)_Y \times U(1)_{B-L}$ R. N. Mohapatra and R. E. Marshak Phys. Rev. Lett. 44 (1980) 1316

Additional scalar singlet χ with the scalar potential $V(H, \chi) = m^2 H^+ H + \mu^2 |\chi|^2 + \lambda_1 (H^+ H)^2 + \lambda_2 |\chi|^4$

Additional B - L gauge boson Z' with interactions

 $1 \supset \sum ig_{B-L}Y_{B-L}Z'f\overline{f}$

The B - L gauge interactions can have suppressed kinetic mixings to the hypercharge in the loop level.

Experimental limits on Higgs mixing



Still have room for production of the additional scalars at colliders!

Higgs mixing $\begin{bmatrix}
 h_1 \\
 h_2
\end{bmatrix}
=
\begin{bmatrix}
 cos a & -sin a \\
 sin a & cos a
\end{bmatrix}
\begin{bmatrix}
 H \\
 \chi
\end{bmatrix}$

 $\lambda_{h_{x}XX} = sin\theta(cos\theta)\lambda_{HXX}$

$$\sigma_{h_x} = sin^2 \theta (cos^2 \theta) \sigma_{h_{SM}}$$

The most stringent limits on heavy scalar is the W boson mass measurement.

Heavy Neutrino, a Natural LLP







$$L \approx 3 \ cm \times (\frac{10^{-6}}{V_{\mu N}})^2 \times (\frac{100 \ GeV}{M_N})^5.$$
$$V_{IN}^2 \approx \frac{m_v}{mN} < \frac{eV}{GeV} < 10^{-10}.$$

LLP detectors at the LHC



- LLP detectors include CODEX-b, FASER, FACET, SHiP, MoEDAL-MAPP and MATHUSLA, etc.
- FASER, MoEDAL-MAPP collecting data since Run3,
- SHiP approved, run from 2031

LLP Search at the LHC



- CMS/ATLAS, LHCb can probe heavy LLPs, with high coverage, high background
- Forward LLP detectors for **light** LLPs, **high events** from meson decay
- Transverse LLP detectors can suffer in low coverage, low background

CMS Muon System for LLP





- We focus on heavy *N* as LLP
- LLP can also be searched with **CMS muon system** since **Run2**.
- No Dedicated trigger, require **high MET** to trigger, MET > 200 GeV
- New trigger since Run3, higher efficiency.

Higgs to pairs of **N**at the LHC





Heavy neutrinos can be pair-produced as exotic Higgs decays if the mixing is appreciable. This channel can still have 10s fb of cross section.

Higgs to pairs of **N**at the LHC



- ISR jets from **gluon fusion** to trigger MET
- Low efficiency from W/Z decays.
- **RHNs** as **LLP**s to **displaced shower** signals at CMS muon system.



- Higgs mixing at current limits, only require sin $\alpha > 0.02$ (HL-LHC)
- Existing CMS Search can test Seesaw!

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- Dedicated trigger since Run3, better sensitivity expected
- $Br(H \rightarrow NN) \approx 10^{-5}$ is required to test seesaw.

EFT

• More General Description in **N**_R**SMEFT(0806.0876)**.

$$\mathcal{L}_5 = \frac{(\alpha_W^{\dagger})_{ab}}{\Lambda} (\bar{L}_a \tilde{\phi}) (\phi^{\dagger} \tilde{L}_b^c) + \frac{(\alpha_{N\phi})_{ss'}}{\Lambda} (\phi^{\dagger} \phi) \,\bar{\nu}_{Rs} \,\nu_{Rs'}^c + \frac{(\alpha_{NB})_{ss'}}{\Lambda} \bar{\nu}_{Rs} \,\sigma^{\mu\nu} \nu_{Rs'}^c \,B_{\mu\nu} + h.c.$$

- Only three in Dim-5
- Weinberg Operator for Neutrino Mass, Dipole for W/Z decay to N



- Anisimov-Graesser Operator for HNN
- Current limits from Higgs signal strength (2210.16279, 2304.06772)

 $\Lambda \geq 30 \; {\rm TeV}$

- Large parameter space in EFTs to probe seesaw
- Limits from this work **surpass** current limits by **more than 10 times**

Leptogenesis to explain the BAU

 $10^{10} + 1$ 10^{10}

Figure from Kaori Fuyuto

Way too small to explain the observed BAU within the SM. We need new physics!

 $\frac{n_{\Delta B}}{s} \approx (8.59 \pm 0.11) \times 10^{-11}$ from Planck satellite and the CMB.

Sakhorov's criteria

1. Baryon number violating process,

2. C and CP violations, $\Gamma(X \rightarrow Y + b) \neq \Gamma(\overline{X} \rightarrow \overline{Y} + \overline{b}), L \text{ and } R.$

3. Out of equilibrium. $\Gamma(X \rightarrow Y + b) \neq \Gamma(Y + b \rightarrow X).$

Leptogenesis to explain the BAU

BAU from neutrino!

- 1. Lepton number is violated within the neutrino masses terms.
- 2. Additional CP violations can exists in the neutrino mass matrix.
- 3. Right-handed neutrinos decay out of equilibrium potentially.

And EW sphaleron to transfer $n_{\Delta L}$ into $n_{\Delta B}$ during EW phase transition, $Y_B = \frac{28}{79}Y_{B-L}$

$$Y_B \simeq 10^{-2} \times \epsilon \times \eta$$

- CP Asymmetry
- Thermal efficiency
- Consider B L gauge boson

Resonant Leptogenesis

• $\epsilon_{\alpha\alpha}$ is the CP asymmetry in N_1 decay,

comes from the interference between the tree-level and one loop amplitude.

• Hierarchical RH neutrinos,

 $\epsilon \lesssim 10^{-15} M_{N_1}, Y_{\Delta B} \simeq 10^{-2} \times \epsilon \times \eta \simeq 10^{-10}.$ As $\eta \sim 0.1, \epsilon \simeq 10^{-7}$, so $M_{N_1} \ge 10^9 \text{GeV}.$

Davidson-Ibarra Bound, no possible collider signatures.

 Resonant leptogenesis (what we focus on) if at least two of the RH neutrinos masses are degenerate, as ΔM ≤ Γ.
 ε ≤ ¹/₂, only needs M_{N1} ≥ T_{sph} ≈ 130 GeV.

CP Asymmetry

• ϵ_{aa} is the CP asymmetry in N_1 decay, without quantum effect

$$-\epsilon_{i\ell} = \sum_{j \neq i} \frac{\mathrm{Im}\left([\lambda_D^{\dagger}]_{i\ell} [\lambda_D]_{\ell j} [\lambda_D^{\dagger} \lambda_D]_{ij}\right) + \frac{M_{N_i}}{M_{N_j}} \mathrm{Im}\left([\lambda_D^{\dagger}]_{i\ell} [\lambda_D]_{\ell j} [\lambda_D^{\dagger} \lambda_D]_{ji}\right)}{[\lambda_D^{\dagger} \lambda_D]_{ii} [\lambda_D^{\dagger} \lambda_D]_{jj}} \left(f_{ij}^{\mathrm{mix}} + f_{ij}^{\mathrm{osc}}\right),$$

• Yukawa can be expressed by Casas-Ibarra paramterasation.

$$\begin{split} \lambda_D &= \frac{1}{v_{\rm EW}} U \sqrt{\hat{m_{\nu}}} R^T \sqrt{\hat{M_N}}, \\ U &= \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13} e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13} e^{i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} 1 & 0 & 0 \\ 0 & e^{i\frac{\alpha_{21}}{2}} & 0 \\ 0 & 0 & e^{i\frac{\alpha_{31}}{2}} \end{pmatrix} \\ R &= \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{\omega_1} & s_{\omega_1} \\ 0 & -s_{\omega_1} & c_{\omega_1} \end{pmatrix} \begin{pmatrix} c_{\omega_2} & 0 & s_{\omega_2} \\ 0 & 1 & 0 \\ -s_{\omega_2} & 0 & c_{\omega_2} \end{pmatrix} \begin{pmatrix} c_{\omega_3} & s_{\omega_3} & 0 \\ -s_{\omega_3} & c_{\omega_3} & 0 \\ 0 & 0 & 1 \end{pmatrix}, \end{split}$$

CP Asymmetry

• Connection to seesaw

$$|V|^2 \equiv \Sigma_{i=1,2,\ell=e,\mu,\tau} |V_{\ell N_i}|^2 = \frac{\Sigma_i m_{\nu_i}}{M_N} \cosh(2y_2) \approx 2\Sigma_{\ell=e,\mu,\tau} |V_{\ell N_{1/2}}|^2.$$

- V^2 is tiny from seesaw \rightarrow LLPs
- Tiny $V^2 \rightarrow \text{tiny } y_2$
- Large $\epsilon_{\alpha\alpha}$ require tiny y_2
- Both Leptogenesis and Seesaw prefer N as LLPs





Free Parameters

Consider neutrino oscillation data, -5 parameters

 $\begin{aligned} \theta_{12} &= 33.44^{\circ} \quad \theta_{13} = 8.57^{\circ}, \quad \theta_{23} = 49.20^{\circ}, \\ m_{\nu_2} &= \sqrt{\Delta m_{\rm sol}^2} = 8.6 \times 10^{-3} \text{ eV}, \quad m_{\nu_3} = \sqrt{\Delta m_{\rm atm}^2} = 5.0 \times 10^{-2} \text{ eV}. \end{aligned}$

Take the massless lightest neutrino, decouple the heaviest N, -6 parameters $\omega_{1,3} = \pi/2, \, \omega_2 = x_2 + i y_2$

One free parameter rotate away, -1

Consider couplings with B - L gauge boson Z' add two free parameters

 $m_{Z'}, g_{B-L}$

Free parameters,

$$3x3x2 - 5 - 6 - 1 + 2 = 8$$

Free Parameters

Benchmark for free parameters

 $M_N = 1 - 1000 \text{ GeV}, \quad \Delta M_N = 10^{-17} - 10^{-4} \text{ GeV},$ $M_{Z'} = 5 \text{ TeV}, \quad g_{B-L} = 0.025, 0.05, 0.15,$ $x_2 = \pi/4, \quad y_2 = 0 - 5, \quad \delta = 3\pi/2, \quad \alpha_{23} = \pi.$

$$U = \begin{pmatrix} 0.825 & 0.545i & 0.149i \\ -0.360 + 0.094i & 0.062 + 0.545i & 0.749 \\ 0.417 + 0.081i & 0.054 - 0.632i & 0.646 \end{pmatrix},$$

$$\begin{split} |V_{\ell N_i}|^2 &\approx 10^{-13} \times \frac{100 \text{ GeV}}{M_N} \\ &\times \begin{pmatrix} |0.59e^{-1.57i}\cosh y_2 + 0.12\sinh y_2|^2 & |0.12e^{1.57i}\cosh y_2 + 0.59\sinh y_2|^2 & 0 \\ |1.20e^{0.3i}\cosh y_2 + 1.28e^{1.29i}\sinh y_2|^2 & |1.28e^{-0.28i}\cosh y_2 + 1.20e^{-1.27i}\sinh y_2|^2 & 0 \\ |1.07e^{-0.4i}\cosh y_2 - 1.14e^{-1.20i}\sinh y_2|^2 & |1.14e^{0.37i}\cosh y_2 - 1.07e^{1.17i}\sinh y_2|^2 & 0 \end{pmatrix} \end{split}$$

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Thermal Efficiency

Corrections on the Boltzmann equations,

$$egin{aligned} rac{\mathrm{d}N_{N_i}}{\mathrm{d}z} &= -(D(K_i) + S_{h/A})(N_{N_i} - N_{N_i}^{\mathrm{eq}}) - 2\,S_{Z'}/N_{N_i}^{\mathrm{eq}}(N_{N_i}^2 - (N_{N_i}^{\mathrm{eq}})^2) \;, \ rac{\mathrm{d}N_{\Delta_\ell}}{\mathrm{d}z} &= \sum_i arepsilon_{i\ell} D(K_i)\,(N_{N_i} - N_{N_i}^{\mathrm{eq}}) - \sum_i W^0(K_{i\ell})N_{\Delta_\ell} \;, \end{aligned}$$

Large washout from Z' scattering Thermal efficiency can be analytically expressed

$$\kappa_{i\ell}(z, z_{\rm in}) \approx \int_{z_{\rm in}}^{z} dz' \frac{dN_{N_i}^{\rm eq}}{dz'} \frac{D(K_i, z')}{D(K_i, z') + S_{h/A}(z') + 4S_{Z'}(z')} \\ \times \exp\left[-\int_{z'}^{z} dz'' \sum_{i} W^0(K_{i\ell}, z'')\right],$$

Wash-out from Pair-production



The scattering mediated via Z[']

makes the *N* closer to the equilibrium. The BAU is diluted due to the scatterings.

Test Leptogenesis by LLP searches



Conclusion

- There are still a lot unknown in the neutrino physics. The origin of the neutrino masses are one of the most interesting.
- Right-handed neutrinos can explain the neutrino masses via the type-I seesaw, and the BAU via the leptogenesis.
- By adding a B-L gauge, the pair-production of right-handed neutrinos can be probed at colliders
- We show current CMS displaced shower search can test seesaw whenever there is tiny HNN interaction
- We show future LLP searches might test **both leptogenesis and seesaw** with **B-L gauge** interactions.