

# A new de-excitation tool: GEMINI++4 $\nu$

**Yujie Niu** <sup>1,2</sup> (牛玉杰)

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<sup>1</sup>Institute of High Energy Physics, Beijing, China

<sup>2</sup> University of Chinese Academy of Sciences, Beijing, China

E-mail: niuyj@ihep.ac.cn

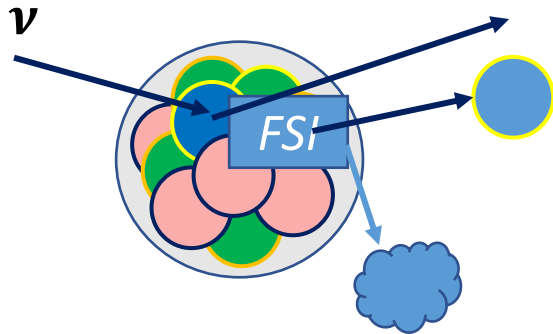
# Outline

- What is / why de-excitation?
- De-excitation generators status and experiments
- GEMINI++ / Performance of GEMINI++
- GEMINI++4 $\nu$ : Modified version of GEMINI++ for **neutrino** exp.
- Summary

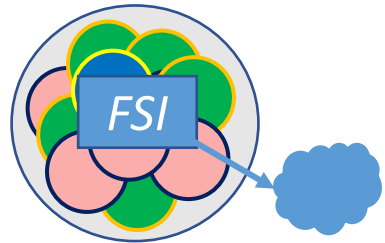
GEMINI++/4 $\nu$ : **Y.J. Niu**, W.L. Guo, M. He, J. Su, arXiv: 2408.14955

# What is / why de-excitation?

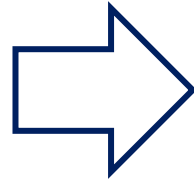
Nucleus de-excitation takes a significant part in neutrino experiment



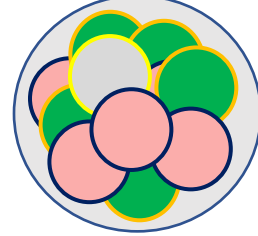
$\nu - N$  interaction



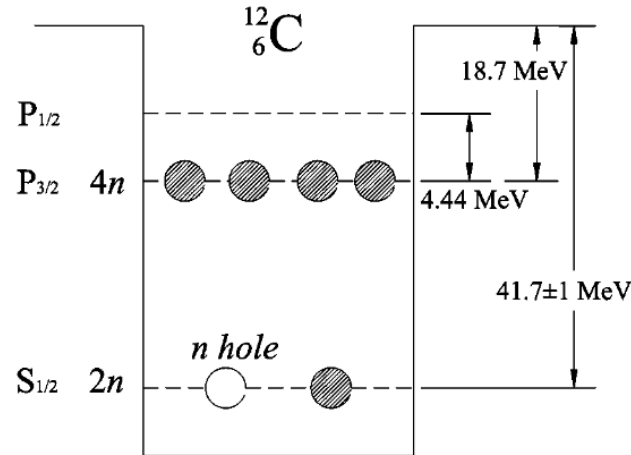
Nucleon decay



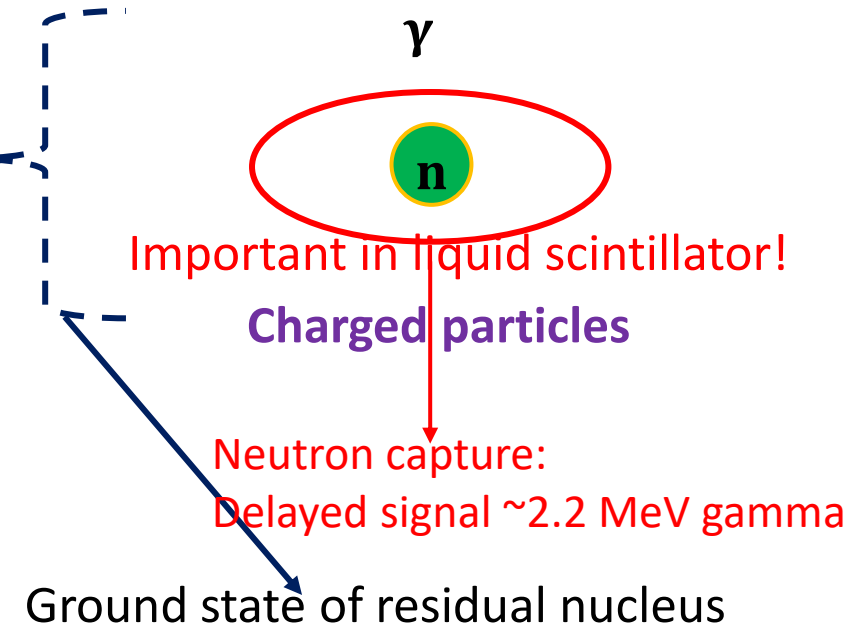
Excited nucleus



De-excite



PHYSICAL REVIEW D 67, 076007 (2003)



- Emitting particles, especially neutrons from de-excitation contribute to **signal-background ratio** like **nucleon decay searching, DSNB, atmospheric  $\nu$**  and so on
- But de-excitation is not considered well in widely-used  $\nu$  generator such as GENIE and NuWro. De-excitation in GEANT4 is not reasonable compared to experimental result of light nucleus <sub>3</sub>

# De-excitation generators

*How about the generator of de-excitation? Good performance?*

Models used to calculate de-excitation probability of emitting particles :

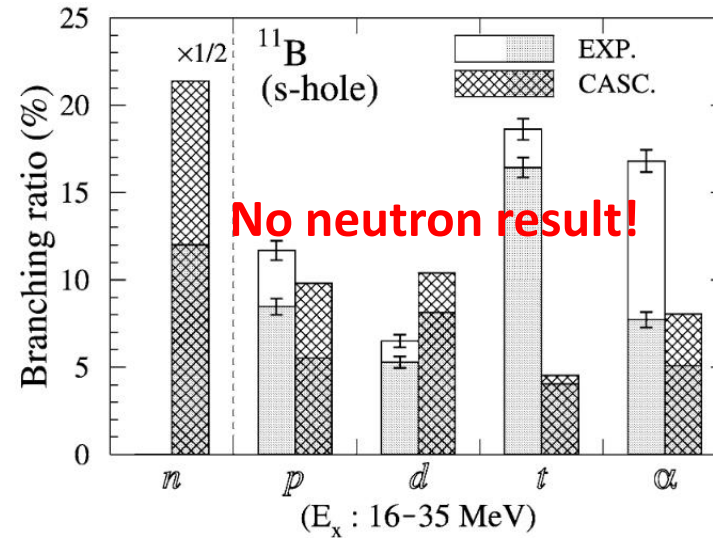
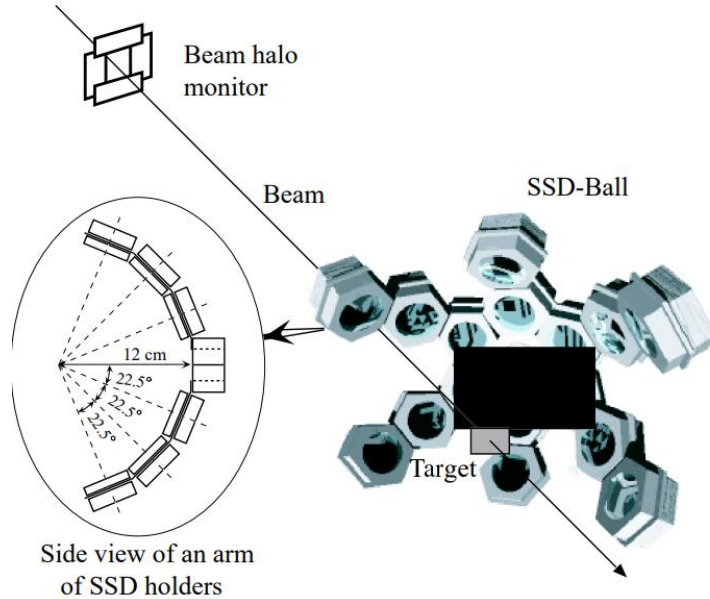
- **Weisskopf-Ewing (WE):** Angular momentum is not considered
- **Hauser-Feshbach (HF):** Angular momentum is considered

<i>Generator</i>	<i>Comments</i>	<i>Model</i>
TALYS	$p \rightarrow \bar{\nu} K^+$ , DSNB, strange axial coupling constant	HF
ABLA	Energy resolution of accelerator neutrinos	WE
SMOKER	Neutron invisible decays	WE
CASCADE	Experimental data of $^{11}\text{B}^*$ and $^{15}\text{N}^*$ de-excitations	HF
<b>GEMINI++</b>	<b>Not used in neutrino experiments!</b>	HF / WE

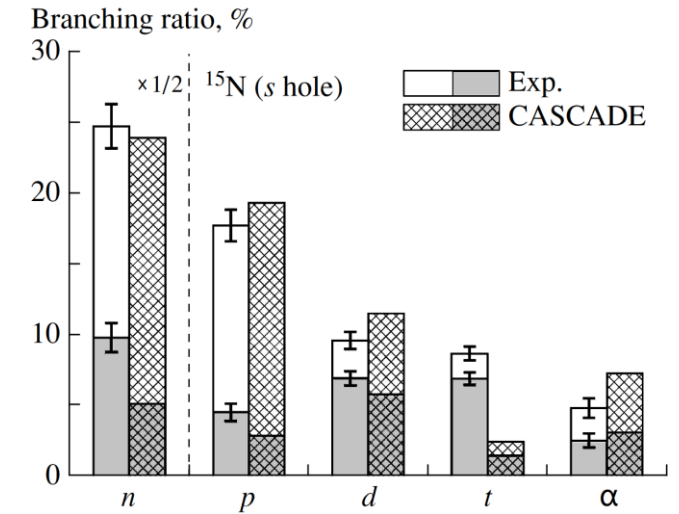
These generators must be validated with **experimental data!**

# De-excitation experiments

## (Yosoi et al.) Exp.1



M. Yosoi et al., Phys. Lett. B 551 (2003);

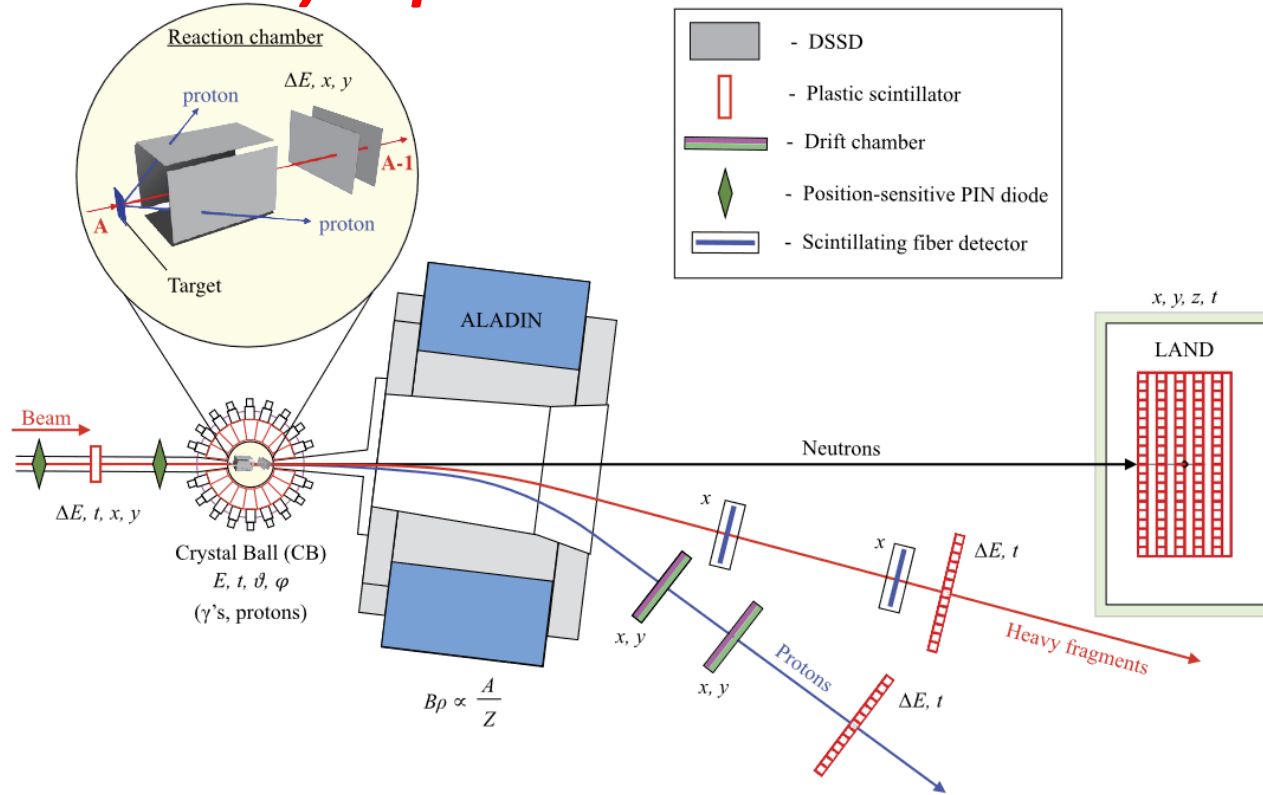


Phys. Atom. Nucl. 67, (2004) 1810

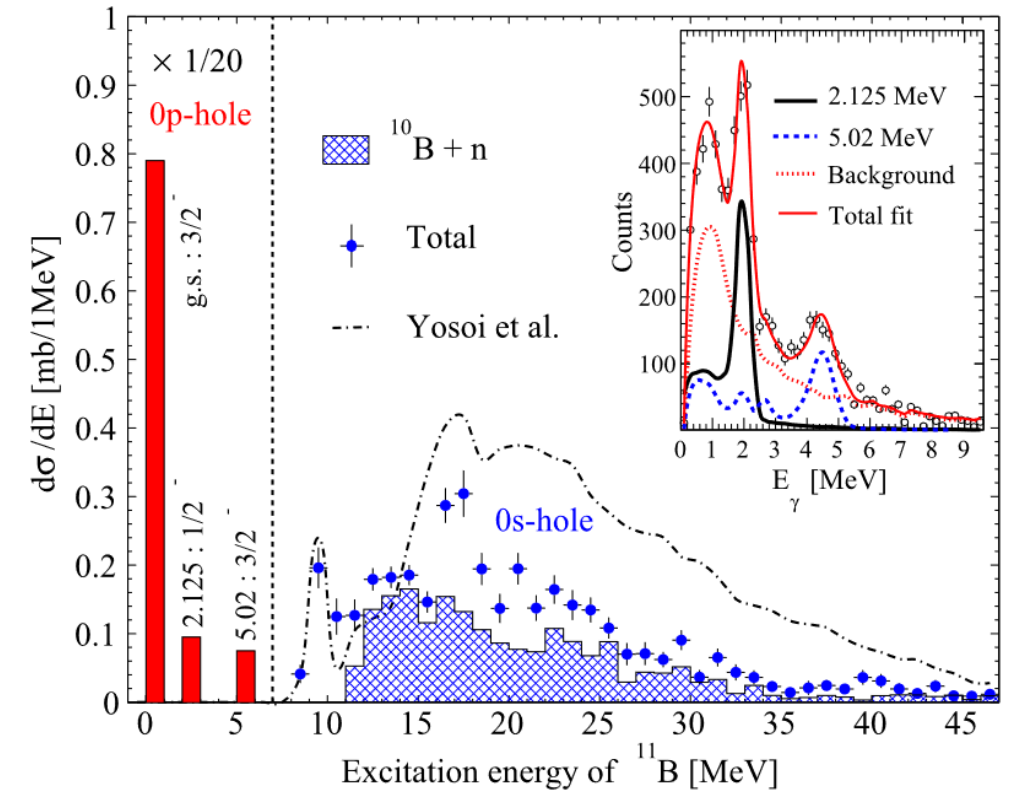
- Quasi-free (p, 2p) reaction, targets are carbon ( $^{12}\text{C} \rightarrow ^{11}\text{B}^*$ ) and ice ( $^{16}\text{O} \rightarrow ^{15}\text{N}^*$ )
- JUNO (Liquid-scintillator  $\rightarrow ^{12}\text{C}$ ), Super-Kamiokando / Hyper-Kamiokando (Water  $\rightarrow ^{16}\text{O}$ )
- Different threshold for particles from de-excitation
- Results shown above, darker color for '2-body decay' and lighter for '3-body decay' In a decay

# De-excitation experiments

## (Panin et al.) Exp.2



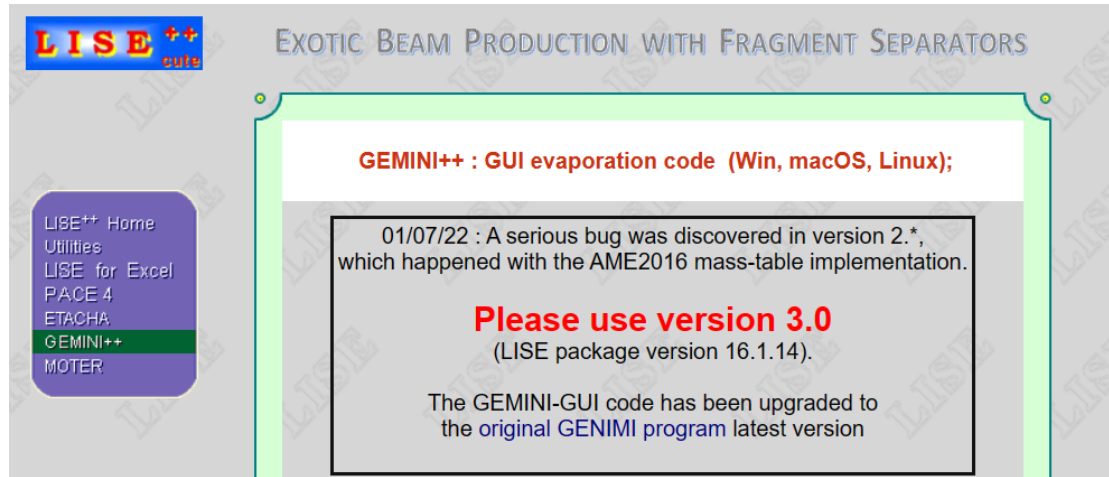
V. Panin, et al., Phys. Lett. B 753 (2016) 204



- $^{12}\text{C}(p, 2p) ^{11}\text{B}^*$
- Good energy resolution, **no threshold** for particle identification, even residue
- Only **three two-body decay channels** of  $^{11}\text{B}$  were analyzed:  $^{10}\text{B} + \text{n}$ ,  $^9\text{Be} + \text{d}$  and  $^7\text{Li} + \alpha$

# Generator GEMINI++

**GENIMI++: a Monte Carlo code, is an improved C++ version based on GEMINI**



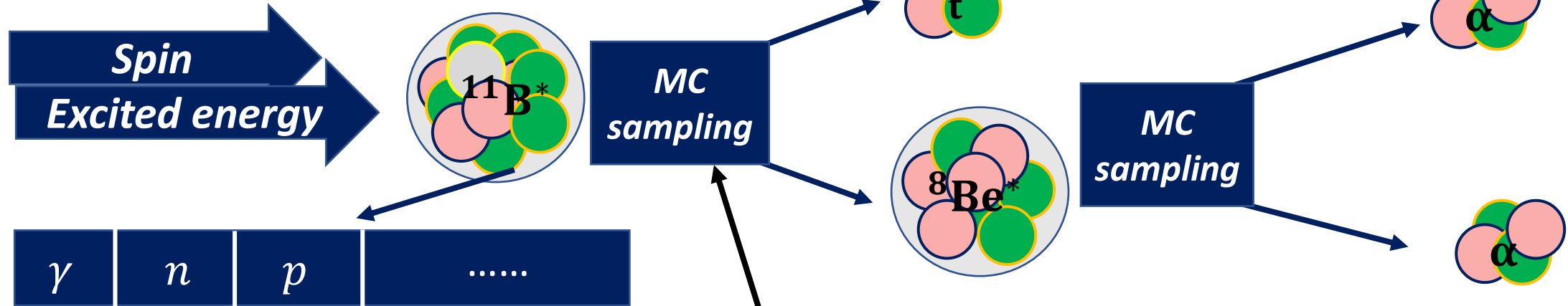
R. J. Charity, PRC 82 (2010), 014610; PRC 82 (2010) 044610  
<https://lise.frib.msu.edu/gemini.html>

<i>Code</i>	<i>GEMINI++</i>
Input	Nucleus, Excited energy, Spin
Formulism of width $\Gamma_i$	HF or WE
Output	Complete de-excitation cascade information
Convenience	Event-by-event, Convenience

- GEMINI++ is designed for **heavy nucleus** process, like fission, fusion and so on
- GEMINI++ had been widely used in nuclear physics area, getting **cheerful** achievements!

# Generator GEMINI++

Take  $^{11}\text{B}^*$  as an example



WE formulism

$$\Gamma_i^W(E_i) = \frac{2s_v + 1}{2\pi\rho_i(E_i)} \int \sum_{l=0}^{\infty} T_l(\epsilon) \rho(U) d\epsilon$$

$i$  means different emitting particles  
 $\Gamma_i^{WE}$ : the partial decay width for evaporation of particle  $i$   
 $s_i, l$ : the spin and orbital angular momenta  
 $T_l(\epsilon)$ : the transmission coefficient,  $\epsilon$ : the kinetic energy of  $i$   
 $U$ :  $E_x - B_i - E_{rot} - \epsilon$   
 $\rho^0, \rho$ : the level densities of the parent and daughter nuclei

$$\rho(U) \propto \frac{\exp\left[2\sqrt{a(U - E_1)}\right]}{a^{1/4}(U - E_1)^{5/4}}$$

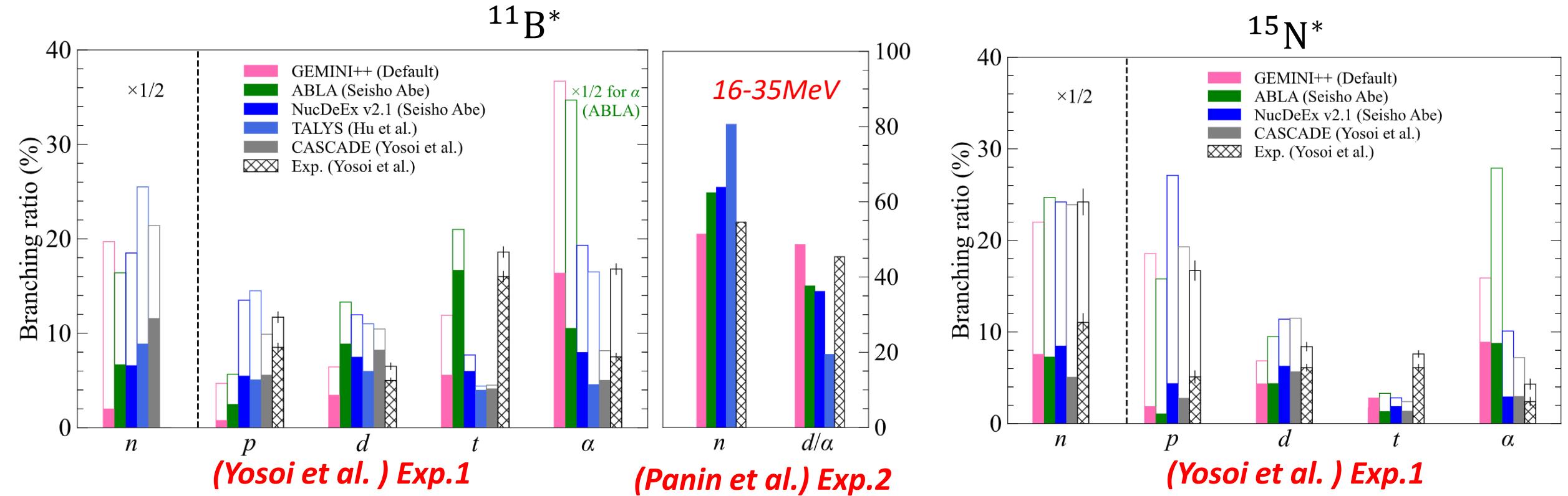
Back-shifted term  $E_1$ : modification of  $U$

Final width  $\leftarrow \Gamma_i = F_s \times \Gamma_i^W$

Suppression factor: Tuning width



# Performance of GEMINI++



- Bad agreement with Exp.1
- Good prediction to Exp.2, but not enough
- GEMINI++ is designed for heavy nucleus, for light nucleus, modifications are required

# Generator GEMINI++4v

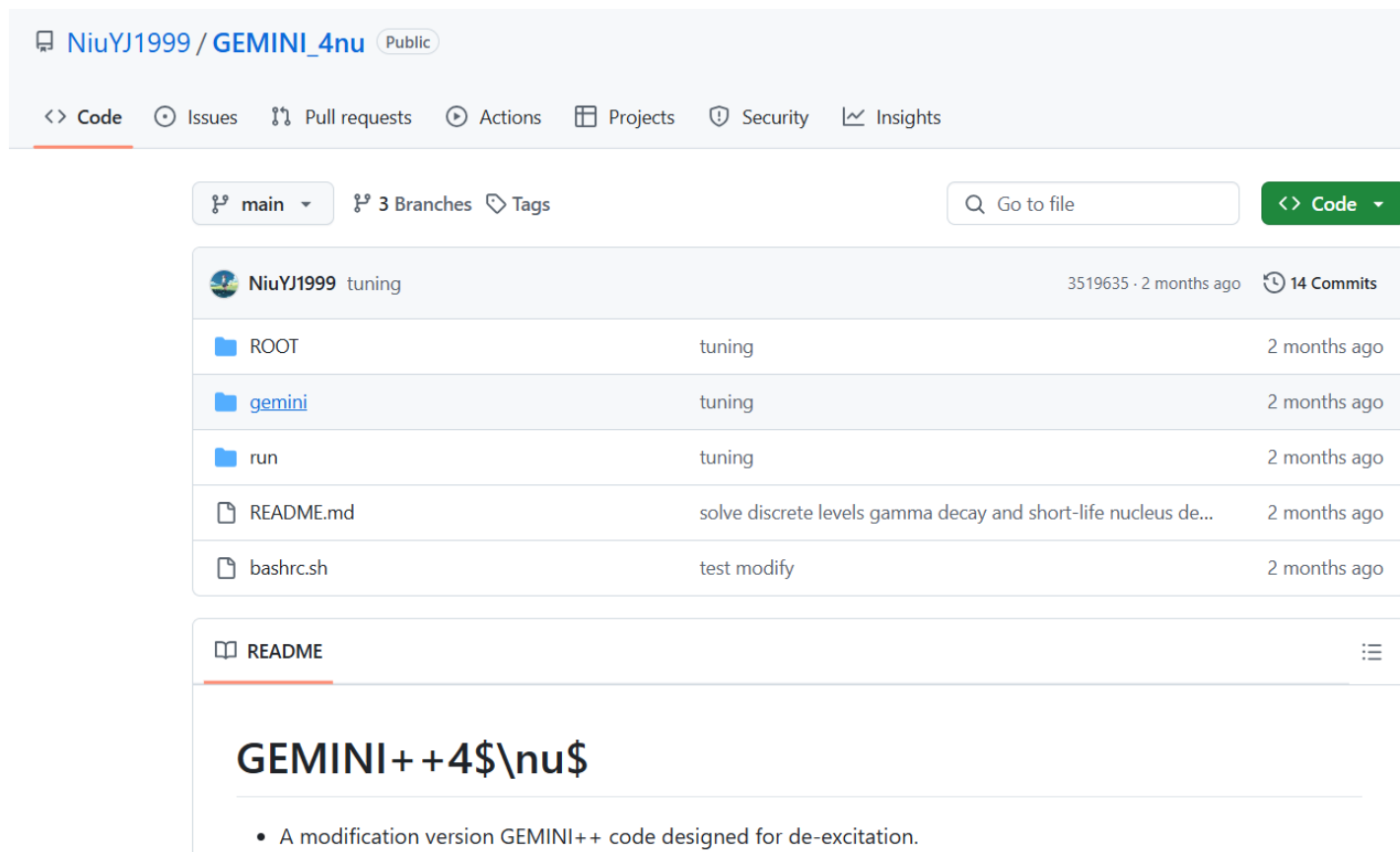
GEMINI++4v is developed for **neutrino** experiments to handle de-excitations of residual nuclei associated with **v** interaction and nucleon decay based on GEMINI++ code

## Open source:

[https://github.com/NiuYJ1999/GEMINI\\_4nu](https://github.com/NiuYJ1999/GEMINI_4nu)

Three modifications are carried:

- ***Removing Back-shifted term***
- ***Add discrete levels***
- ***Adjust suppression factor***



NiuYJ1999 / GEMINI\_4nu Public

<> Code Issues Pull requests Actions Projects Security Insights

main 3 Branches Tags Go to file Code

NiuYJ1999 tuning 3519635 · 2 months ago 14 Commits

ROOT	tuning	2 months ago
gemini	tuning	2 months ago
run	tuning	2 months ago
README.md	solve discrete levels gamma decay and short-life nucleus de...	2 months ago
bashrc.sh	test modify	2 months ago

README

## GEMINI++4 $\nu$

- A modification version GEMINI++ code designed for de-excitation.

# Removing Back-shifted term

5000 events for every interval of 0.1 MeV are simulated ( $0 \text{ MeV} \leq E_x \leq 50 \text{ MeV}$ ) to determine the critical energy  $E_c$  of particle-emitting for 2-body decay

Modes	Separation energy		critical energy
	Theory	GEMINI++	GEMINI++ $4\nu$
$n + {}^{10}\text{B}$	11.5	12.2	12.2
$p + {}^{10}\text{Be}$	11.2	17.1	12.2
$d + {}^9\text{Be}$	15.8	16.6	16.6
$t + {}^8\text{Be}$	11.2	20.5	12.1
${}^3\text{He} + {}^8\text{Li}$	27.2	28.6	28.6
$\alpha + {}^7\text{Li}$	8.7	10.1	10.1

Improved!

$$\rho(U) \propto \frac{\exp\left[2\sqrt{a(U - E_{\mp})}\right]}{a^{\frac{1}{4}}(U - E_{\mp})^{5/4}}$$

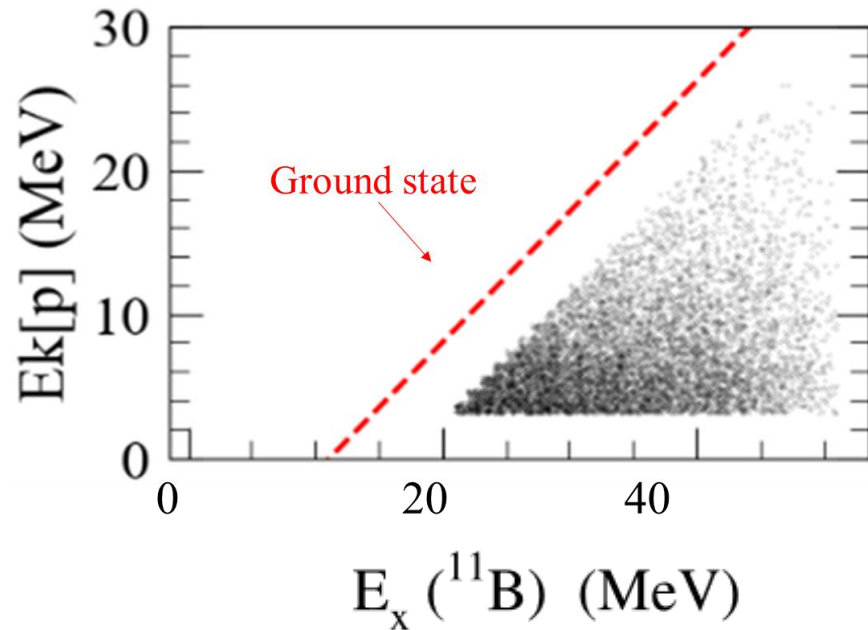
Modification of  $E_1$  equals to tune excited energy:

- $E_c$  can be changed due to this modification
- Winner in competition can be changed

- **No available experimental  $E_1$**  for  ${}^{11}\text{B}^*$  and  ${}^{15}\text{N}^*$  and their daughter nucleus can be used (only  $E_1$  of nucleus mass  $> {}^{19}\text{F}$  can be found)
- $\rho(U)_{E_1 \neq 0}$  and  $\rho(U)_{E_1 = 0}$  are used to do  $\chi^2$  calculation to fit existing discrete levels
- $\rho(U)_{E_1 = 0}$  gives a better result, this modification is reasonable

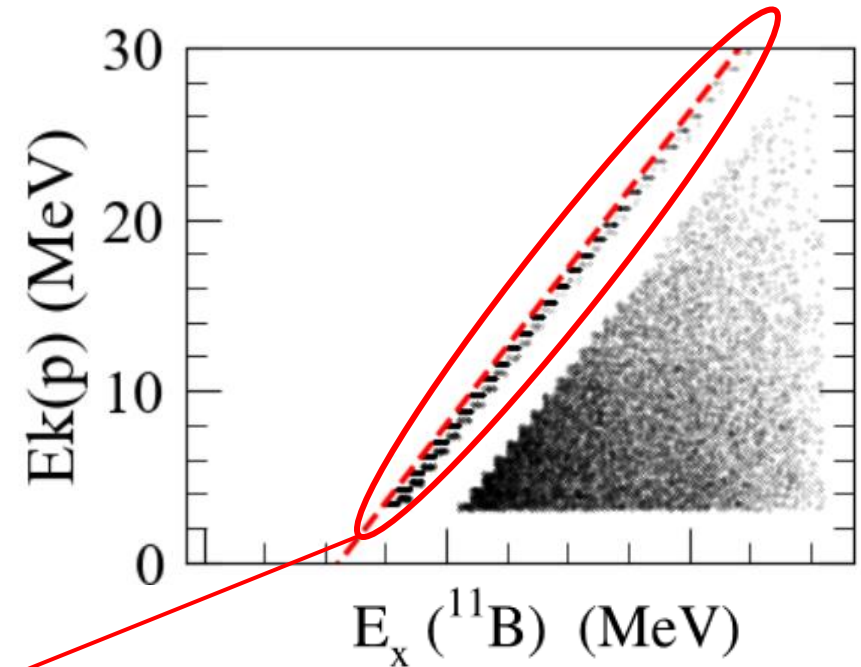
# Add discrete levels

- **GEMINI++ does not consider discrete levels**
- Decayed nucleus always has low excited energy at the last decay
- Modifying the de-excitation result by adding the clearly known discrete levels in NNDC



**Absence of discrete levels**

**Add discrete levels**



**Discrete levels occur**

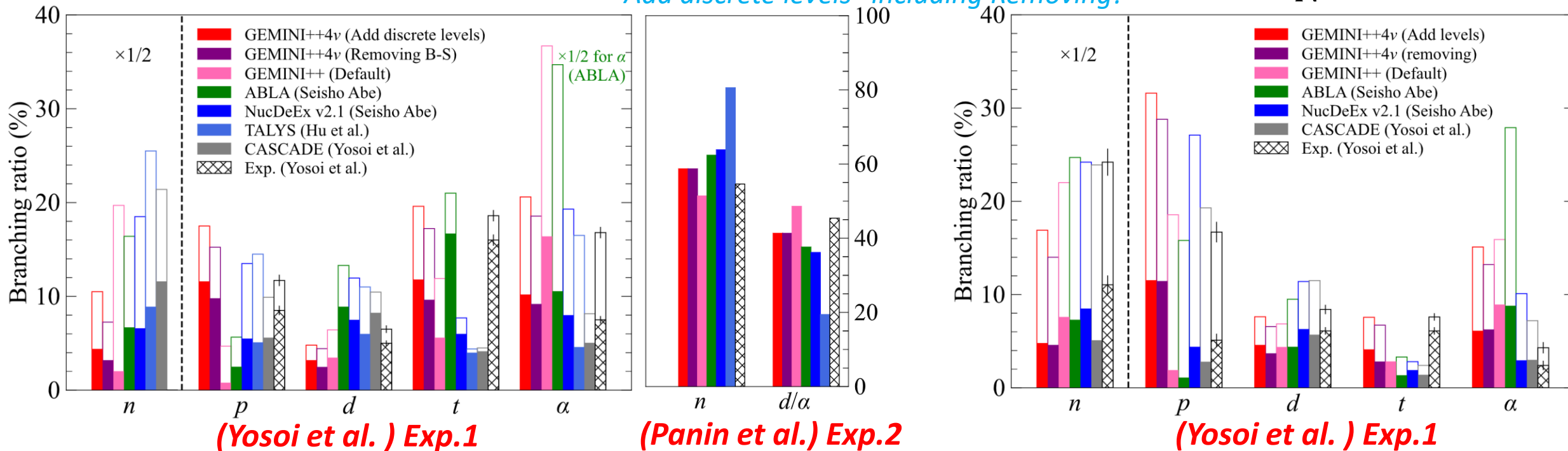
Kinematic energy of some particles can be **increased** to cross the detection threshold in Exp.1

# Impact of modification 1 and 2

$^{11}\text{B}^*$

“Add discrete levels” including Removing!

$^{15}\text{N}^*$



- Besides more reasonable  $E_c$  can be given, removing  $E_1$  obviously provides a better result
- Adding discrete levels provides higher “2-body decay” and “3-body decay” ratios for almost all particles

# Modification of GEMINI++

Default  $F_s$  settings originate from the de-excitations of heavy nuclei

Settings	$n$	$p$	$d$	$t$	${}^3\text{He}$	$\alpha$
Default	1.0	1.0	0.5	0.5	0.5	1.0
$F_s = 1.0$	1.0	1.0	1.0	1.0	1.0	1.0
$F_s = 0.5$	1.0	0.5	0.5	0.5	0.5	0.5

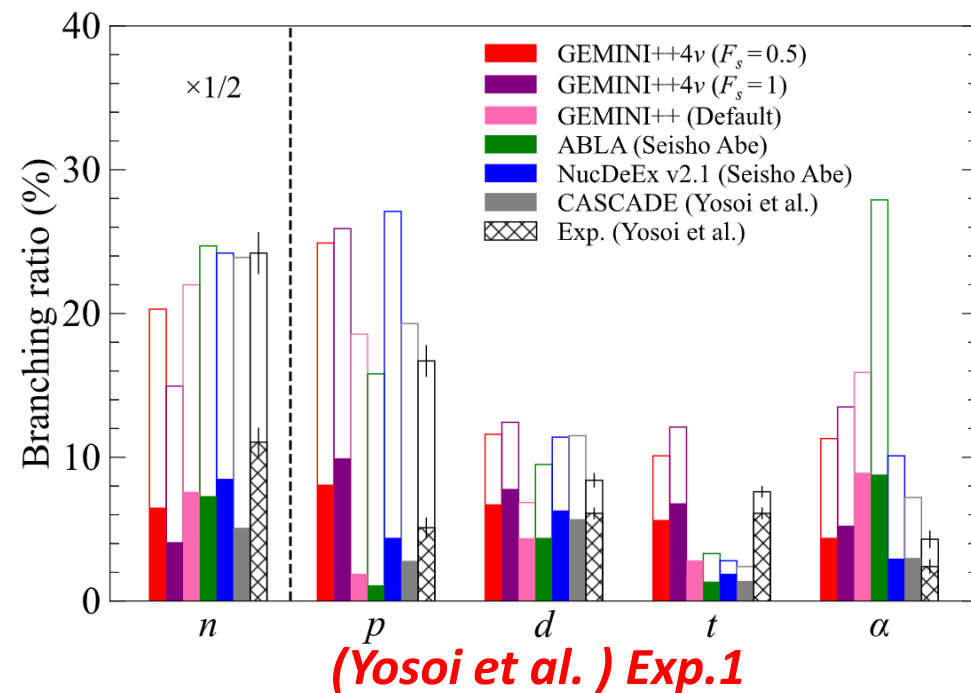
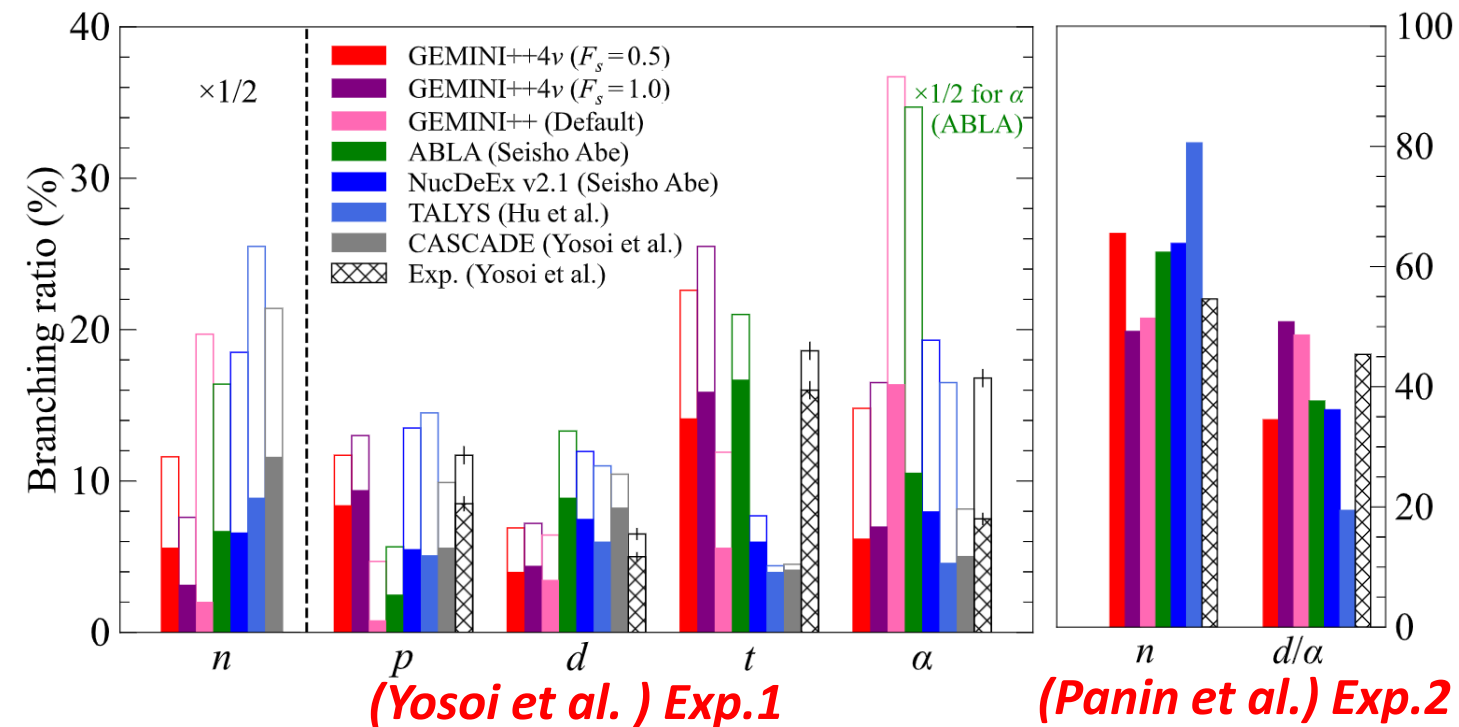
Are default settings reasonable for light nuclei?

- ➔ Don't use suppression factor to adjust results, namely  $F_s = 1.0$  for all particles
- ➔  $F_s = 0.5$  for all charged particles. Compared with default, only two changes

# Modification of GEMINI++

$^{11}\text{B}^*$

$^{15}\text{N}^*$



**GEMINI++4v with  $F_s = 1.0$ :**

- Good agreement with  $^{11}\text{B}^*$  data
- Can't account for  $^{15}\text{N}^*$  data well

**GEMINI++4v with  $F_s = 0.5$  (Recommend!)**

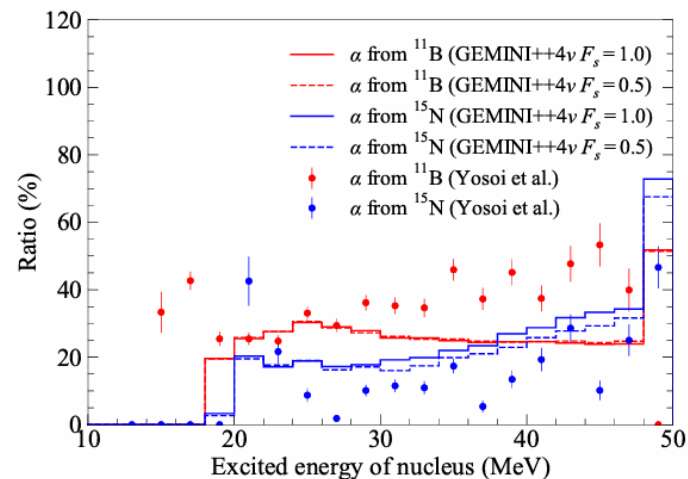
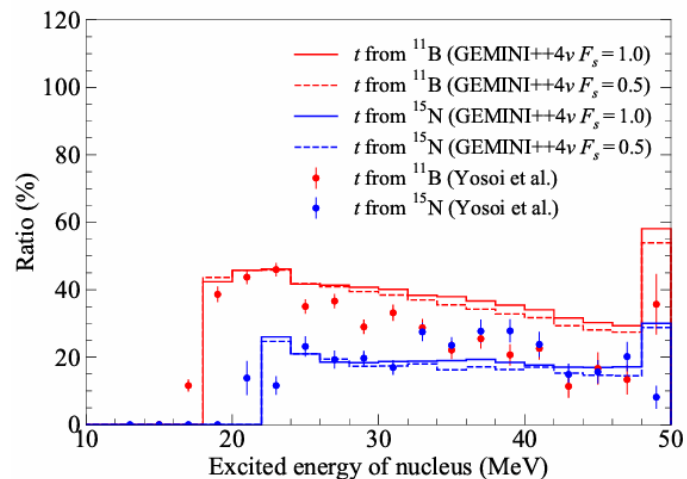
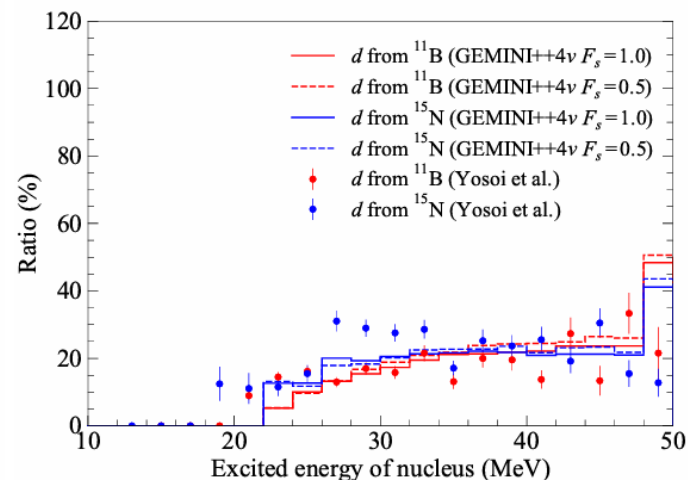
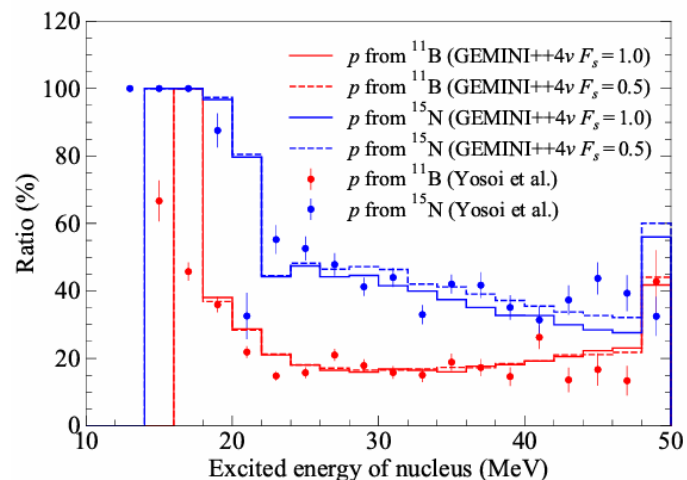
- Better agreement with  $^{11}\text{B}^*$  data
- Partially account for  $^{15}\text{N}^*$  data, include  $n$

**This is the first time that a code can basically reproduce both  $^{11}\text{B}^*$  and  $^{15}\text{N}^*$  data**

# Accidental coincidence check

Fixed energy ranges of  $16 \leq E_x \leq 35$  MeV for  $^{11}\text{B}^*$  and  $20 \leq E_x \leq 40$  MeV for  $^{15}\text{N}^*$

➔ Compare the ratio of each type of charged particle emission among four types for every energy bin



- $F_s = 1.0$  and  $F_s = 0.5$  differences are relatively small
- Predicted shapes are basically consistent with data except  $\alpha$
- Discrepancy maybe come from  $^{11}\text{B}^* \rightarrow t + \alpha + \alpha$

**Not coincidental!**



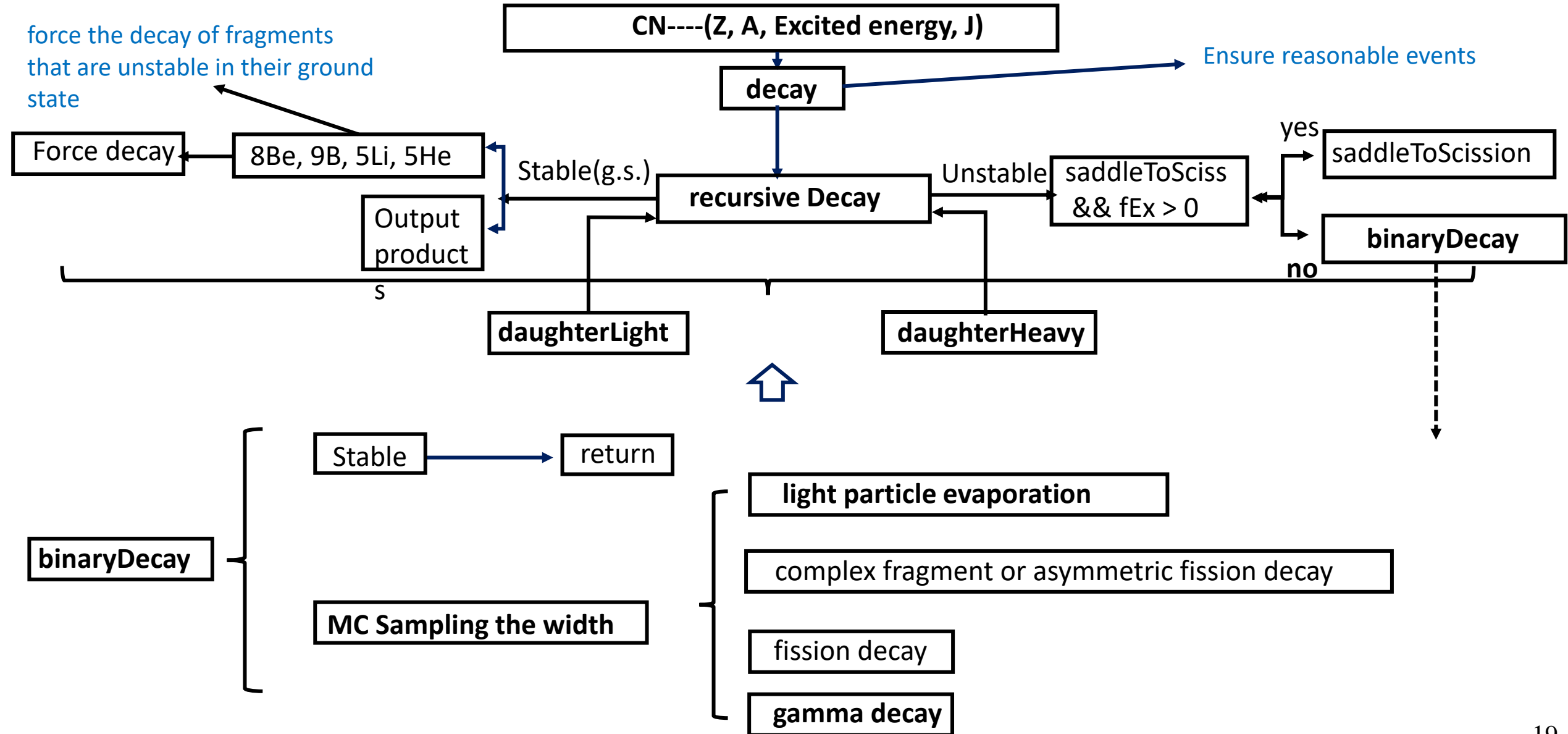
# Summary

- De-excitation plays an more and more important role in  $\nu$  experiments
- GEMINI++ is a potential event-by-event de-excitation generator
- Three modifications provide the best agreement with experiments in  $^{11}\text{B}^*$  and  $^{15}\text{N}^*$ , the modified generator is named **GEMINI++4 $\nu$**
- More work is on-going, such as the predictions of gamma emitting from de-excitation
- Plan to combine GEMINI++4 $\nu$  into widely-used  $\nu$  generator such as GENIE and NuWro

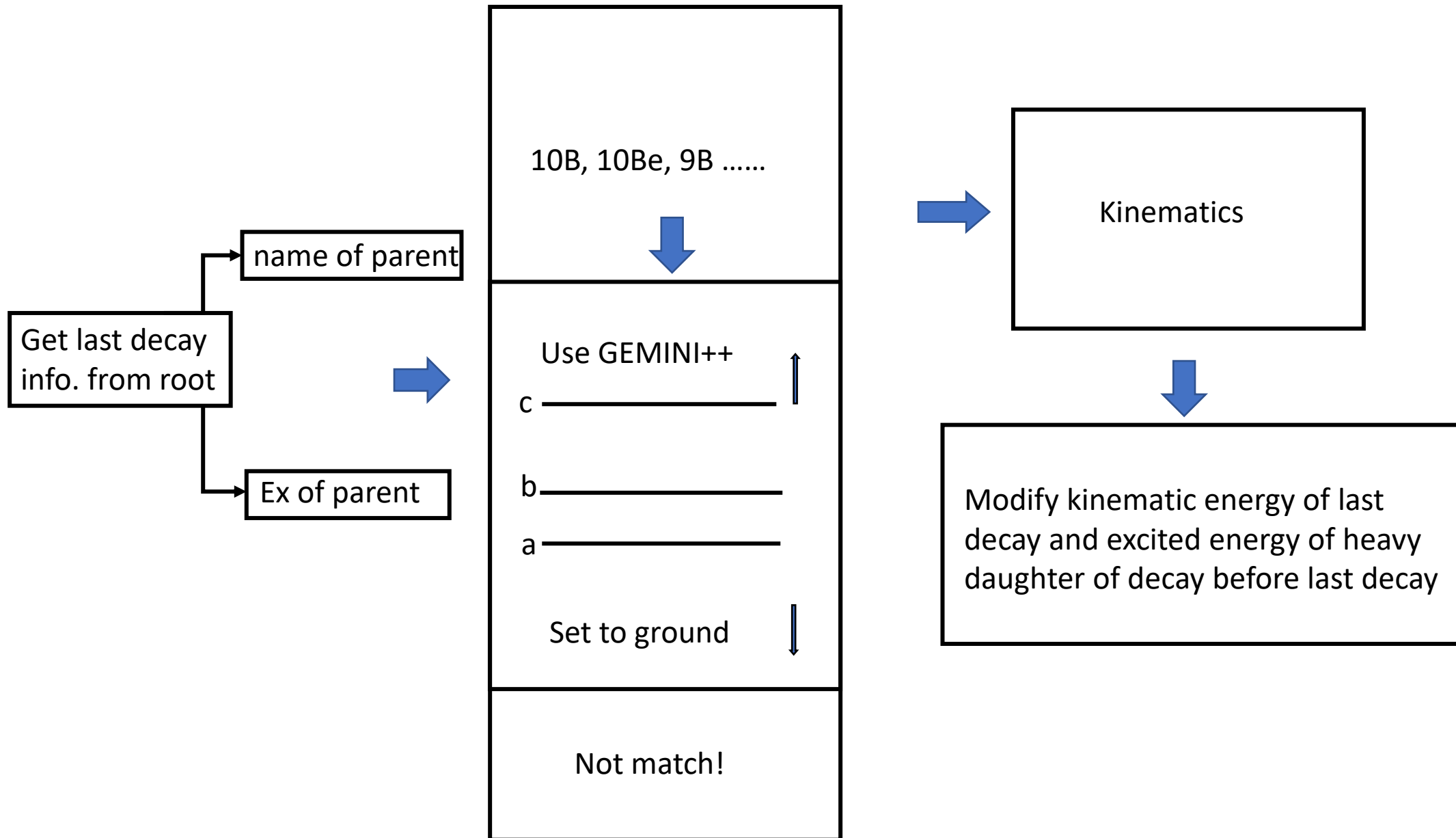
*Thanks for your attention!*

**Back up**

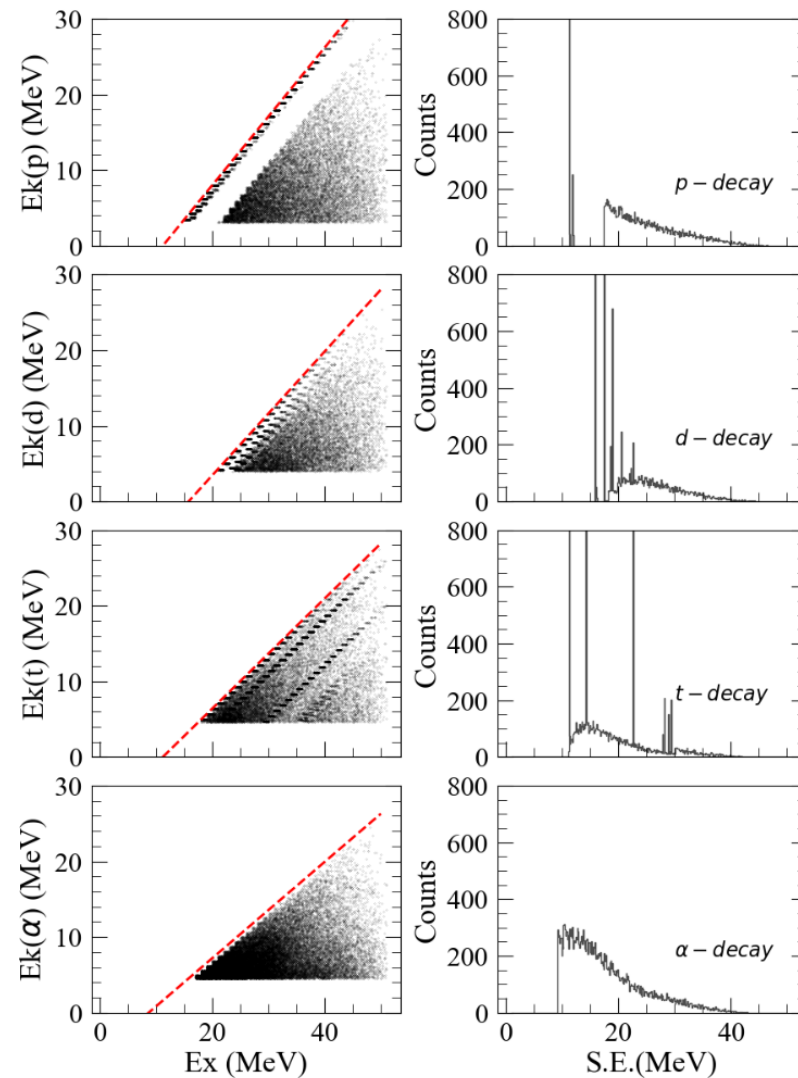
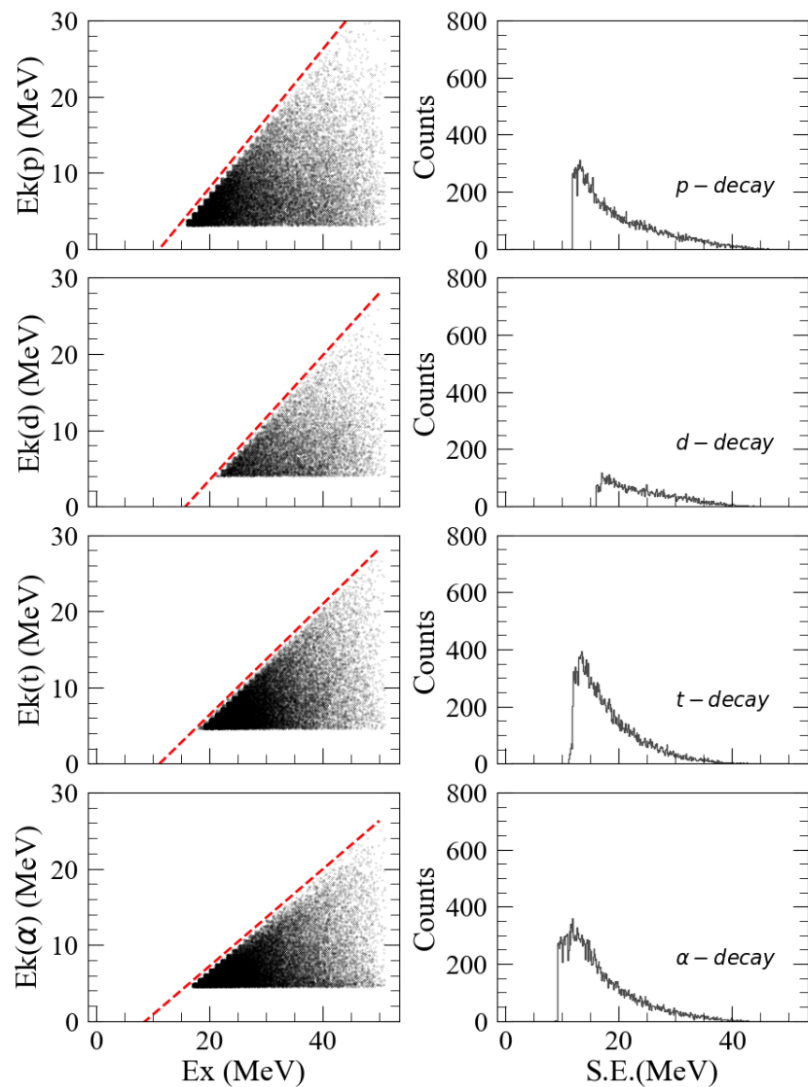
# Code detail (picture)



# Modification of the last decay



# Modification of the last decay



Modify some residues to G.S.



Modify with p



Usually the decay for  $^8\text{Be}$

# Width calculation

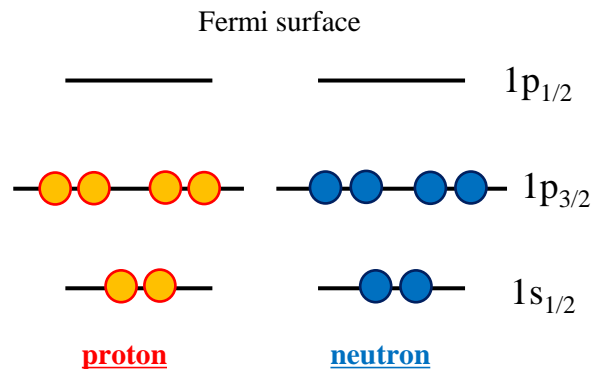
## Hauser-Feshbach formalism

$$\Gamma_i^{HF} = \frac{1}{2\pi\rho(E^*, S_{CN})} \int d\epsilon \sum_{S_d=0}^{\infty} \sum_{J=|S_{CN}-S_d|}^{S_{CN}+S_d} \sum_{l=|J-S_i|}^{J+S_i} T_l(\epsilon) \rho(E^* - B_i - \epsilon, S_d)$$

- $S_i, J, l$  for the evaporated particle,  $T_l$  is transmission coefficient,  $\epsilon$  for kinematic energy,  $B_i$  for separation energy
- $S_d$  is the spin of residue,  $S_{CN}$  is the spin of CN,  $\rho$  for level density
- Evaporation channels include n, p, d, t, 3He,  $\alpha$  and so on
- **Ignore the effect of angular momentum, gives Weisskopf formalism**

$$\Gamma_i^W(E_i) = \frac{2s_v + 1}{2\pi\rho_i(E_i)} \frac{2m_v}{\pi\hbar^2} \int_0^{E_i - S_v - B_v} \sigma_c(\epsilon_v) \rho_f(E_f) \times (\epsilon_v - B_v) d\epsilon$$

## Nuclear effects considered



- **Shell correction**
- **Pairing**

## Back-shifted Fermi Gas model

$$\rho_F(E_x, J, \Pi) = \frac{1}{2} \frac{2J+1}{2\sqrt{2\pi}\sigma^3} \exp\left[-\frac{(J+\frac{1}{2})^2}{2\sigma^2}\right] \frac{\sqrt{\pi} \exp[2\sqrt{aU}]}{12 a^{1/4} U^{5/4}}$$

- $U = E^* - \Delta_{BFM}$ ,  $\Delta_{BFM}$  is the parameter affected by shell correction and pairing, in fact, an adjustable parameter to fit experiment data in theory
- Parameter  $a$  has relationship with shell correction and A in GEMINI++

