#### 第二届武汉高能物理青年论坛

#### The review of QCD axion

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**\star** The  $\theta$ -vacua and strong CP problem

★ The Peccei Quinn mechanism

★ PQWW, DFSZ and KSVZ model

★ Axion effective Lagrangian and mass

★ Axion and other open issues of the SM

#### QCD vacuum structure

The vacuum configurations of gluon is gauge equivalent to  $G_{\mu}=0$ . In temporal gauge, this means

$$G_j = ig^{-1}U\partial_j U^{-1}$$

$$U = U(\mathbf{x}) \in SU(3)_C$$

 $U(\mathbf{x})$  is a map from  $\mathbb{R}^3$  to  $SU(3)_C$ . Such maps, and hence the vacuum configurations, can be classified by elements of  $\pi_3(SU(3)_C) \cong \mathbb{Z}$ .

These integers are called winding number

By a **path** in a space *X* we mean a continuous map  $f: I \rightarrow X$  where *I* is the unit interval [0,1]. The idea of continuously deforming a path, keeping its endpoints fixed, is made precise by the following definition. A **homotopy** of paths in *X* is a family  $f_t: I \rightarrow X$ ,  $0 \le t \le 1$ , such that

- (1) The endpoints  $f_t(0) = x_0$  and  $f_t(1) = x_1$  are independent of *t*.
- (2) The associated map  $F: I \times I \rightarrow X$  defined by  $F(s, t) = f_t(s)$  is continuous.



#### ★ QCD vacuum structure



Weinberg EJ. *Classical Solutions in Quantum Field Theory: Solitons and Instantons in High Energy Physics*. Cambridge University Press; 2012.

#### ★ QCD vacuum structure

The winding number can be calculated by a homomorphism

$$N[U] = \frac{1}{24\pi^2} \epsilon^{ijk} \int d^3x \operatorname{tr} U^{-1} \partial_i U U^{-1} \partial_j U U^{-1} \partial_k U$$

This homomorphism is related to chiral anomaly

$$\partial_{\mu}J^{\mu}_{A} = \frac{g^{2}}{16\pi^{2}} \operatorname{Tr} G_{\mu\nu}\tilde{G}^{\mu\nu}$$

$$\operatorname{Tr} G_{\mu\nu}\tilde{G}^{\mu\nu} = \partial_{\mu}\epsilon^{\mu\nu\rho\sigma}\operatorname{Tr} \left(G_{\nu}G_{\rho\sigma} + \frac{2\mathrm{i}g}{3}G_{\nu}G_{\rho}G_{\sigma}\right)$$

#### ★ QCD vacuum structure

The charge associated with the axial current is  

$$Q_A = \int d^3x J_A^0 = \frac{g^2}{16\pi^2} \int d^3x \epsilon^{ijk} \operatorname{tr} \left( G_i G_{jk} + \frac{2ig}{3} G_i G_j G_k \right)$$

Now let  $G_{\mu}(t, \mathbf{x})$  be a path parameterized by t, interpolate between two vacuum configurations with winding numbers  $N_1$  and  $N_2$ 

 $\Delta N$  is gauge invariant while N is not

$$\frac{g^2}{16\pi^2} \int d^4x \,\mathrm{Tr}\,G_{\mu\nu}\tilde{G}^{\mu\nu} = N_2 - N_1 \,\checkmark$$

#### 🖈 θ-vacua

There exists one quantum vacuum state for each winding number, called n-vacua. Large gauge transformation takes one n-vacuum to the next

$$T|n\rangle = |n+1\rangle$$

The true vacuum is also a eigenvector of T

$$|\theta\rangle = \sum_{n=-\infty}^{\infty} e^{-in\theta} |n\rangle \qquad \qquad T|\theta\rangle = e^{i\theta} |\theta\rangle$$

★ θ-vacua

The  $\theta$ -vacua to  $\theta'$ -vacua amplitude is

$$\langle \theta' | \theta \rangle = 2\pi \delta(\theta - \theta') \sum_{k} e^{ik\theta} \langle k | 0 \rangle$$
$$\langle k | 0 \rangle = \int [dG] e^{i \int d^4 x \mathcal{L}}$$

The factor of  $e^{ik\theta}$  is equivalent to adding to the Lagrangian density a term of

$$e^{ik\theta} = \exp\left(\frac{i\theta g^2}{16\pi^2} \int d^4x \operatorname{Tr} G_{\mu\nu} \tilde{G}^{\mu\nu}\right)$$
  
$$\theta \text{ term}$$

★ Strong CP problem

The  $\theta$  term is CP odd

$$CP: \operatorname{Tr} G_{\mu\nu}\tilde{G}_{\mu\nu} \to -\operatorname{Tr} G_{\mu\nu}\tilde{G}_{\mu\nu}$$

The  $\theta$  term can be transformed to the phase of the mass term for the quarks by  $U(1)_A$ 

$$U(1)_A : \bar{q}_L M q_R + h.c. \to \bar{q}_L M e^{-2i\alpha} q_R + h.c.$$
$$\frac{\theta g^2}{32\pi^2} G^{a\mu\nu} \tilde{G}^a_{\mu\nu} \to (\theta + 2N\alpha) \frac{g^2}{32\pi^2} G^{a\mu\nu} \tilde{G}^a_{\mu\nu}$$

#### ★ Strong CP problem

Now the CP odd term is the imaginary part of mass term for quark

$$\mathcal{L}_{\mathcal{SP}} = -\mathrm{i}\bar{\theta} \frac{m_u m_d}{m_u + m_d} (\bar{u}\gamma_5 u + \bar{d}\gamma_5 d) \longrightarrow \bar{\theta} = \theta + \arg\det M$$

This term will generate a CP-violating pion-nucleon interaction

$$\mathcal{L}_{\pi NN} = g^{\theta}_{\pi NN} \bar{N} \tau^a N \pi^a$$

#### ★ Strong CP problem

It contributes to neutron's electric dipole moment



David Bailin and Alexander Love. *Introduction to Gauge Field Theory*. Taylor and Francis Groupr, 1993.

★ Strong CP problem

The experimental data gives<sup>1,2</sup>

 $|\bar{\theta}| < 10^{-10}$ 

#### The strong CP problem: why $\bar{\theta}$ is so small?

1. C. A. Baker *et al., Phys. Rev. Lett.* **97** (2006) 131801 2. J. Engel et al., *Prog. Part. Nucl. Phys.* **71** (2013) 21–74,

#### ★ Peccei and Quinn's proof

Peccei and Quinn proved that  $\bar{\theta}$  is vanished when the Lagrangian possesses a chiral U(1) symmetry.<sup>1,2</sup>

In single-flavor toy model  $\mathcal{L} = -\frac{1}{4}G^a_{\mu\nu}G^{\mu\nu a} + i\bar{\psi}D_{\mu}\gamma^{\mu}\psi + G\bar{\psi}_L\psi_R\varphi + G^*\bar{\psi}_R\psi_L\varphi^*$   $-|\partial_{\mu}\varphi|^2 - \mu^2|\varphi|^2 - h|\varphi|^4; \quad \mu^2 < 0.$ 

1. Roberto D Peccei and Helen R Quinn, *Phys. Rev. Lett*, 38(25):1440, 1977. 2. Roberto D Peccei and Helen R Quinn, *Phys. Rev. D*, 16(6):1791, 1977.

★ Peccei and Quinn's proof

The CP violating phase is

$$\bar{\theta} = \theta + \arg G + \arg \langle \varphi \rangle$$

The LO of effective potential can be calculated

$$A = \arg \varphi - \arg \langle \varphi \rangle$$

$$V_{\text{eff}} = U(|\varphi|) - K|G\varphi|\cos\left(\bar{\theta} + A\right)$$

It implies that  $\ \bar{\theta} = 0$ 

The phase of scalar  $a = A/|\langle \varphi \rangle|$  is called axion

#### ★ Vafa-Witten theorem

Parity cannot be spontaneously broken in QCD<sup>1</sup>

$$E(0) \le E(\lambda) = 1 + \lambda \langle G\tilde{G} \rangle + O(\lambda^2)$$
$$E(\bar{\theta} + A) \ge E(\bar{\theta}) \Rightarrow \bar{\theta} = 0$$

1. C. Vafa, E. Witten, Phys. Rev. Lett. 53 (1984) 535.

#### ★ PQ quality problem

PQ symmetry may be violated by effective operator by gravitational corrections. For example

$$g_5 \, \frac{|\Phi|^4 \, (\Phi + \Phi^*)}{M_{\rm Pl}}$$

These violations result in a non-vanished  $\bar{\theta}$ 

### PQWW, DFSZ and KSVZ model

#### ★ PQWW model

The first realistic axion model is PQWW model<sup>1</sup>

$$-\mathcal{L}_Y = y_{ij}^{(u)} \bar{Q}_i H_u u_{jR} + y_{ij}^{(d)} \bar{Q}_i H_d d_{jR} + h.c.$$

 $H_u$  is independent with  $H_d$  for the independence of  $U(1)_{PQ}$  and  $U(1)_Y$ .



1. Weinberg PhysRevLett.40.223, Wilczek PhysRevLett.40.279, 1978

### PQWW, DFSZ and KSVZ model

#### ★ DFSZ model

DFSZ model requires the additional SM-singlet scalar to PQWW model<sup>1</sup>  $V(H_u, H_d, \phi) = \tilde{V}_{moduli}(|H_u|, |H_d|, |\phi|, |\phi_u^{\dagger} \tilde{\phi}_d|) + \lambda H_u^{\dagger} \tilde{H}_d \phi^2 + h.c$ 

The axion is the combination of three phases

$$v_a^2 \equiv \mathcal{X}_u^2 v_u^2 + \mathcal{X}_d^2 v_d^2 + \mathcal{X}_\phi^2 v_\phi^2, \quad a = \frac{1}{v_a} (\mathcal{X}_u v_u a_u + \mathcal{X}_d v_d a_d + \mathcal{X}_\phi v_\phi a_\phi)$$

In the limit  $v_{\phi} \gg v_u, v_d$ ,

$$f_a \simeq v_\phi/6, \quad a \simeq a_\phi$$

1. M. Dine et al., Physics Letters B 104, 199 (1981)

## PQWW, DFSZ and KSVZ model

#### ★ KSVZ model

The KSVZ model extends the SM field content with a heavy quark and a SM-singlet<sup>1</sup>

$$-\mathcal{L}_Y = y_{\mathcal{Q}} \bar{\mathcal{Q}}_L \mathcal{Q}_R \phi + h.c.$$

The axion is just the phase of  $\phi$  with  $f_a = 2v$ 

1. J. E. Kim, Phys. Rev. Lett. 43, 103 (1979)

#### ★ Axion effective Lagrangian

All of the axion model contain the quark-scalar Yukawa term

$$-\mathcal{L}_Y \supset y_q \bar{q}_L q_R \phi + h.c.$$

With a local quark  $U(1)_A$  rotation

$$q \to \mathrm{e}^{-\mathrm{i}\gamma_5 a/f_a} q$$

Now the axion is disentangled from the Yukawa term with some additional term

★ Axion effective Lagrangian

$$\Delta \mathcal{L} = \frac{g_s^2}{32\pi} \frac{a}{f_a} G^a_{\mu\nu} \tilde{G}^{a\mu\nu} + \frac{e^2}{32\pi} \frac{E}{N} \frac{a}{f_a} F^a_{\mu\nu} \tilde{F}^{a\mu\nu} + c_q^0 \frac{\partial_\mu a}{2f_a} \bar{q} \gamma^\mu \gamma_5 q$$

The axion effective Lagrangian (in quark level) reads

$$\mathcal{L}_{a} = \frac{1}{2} \left(\partial_{\mu} a\right)^{2} + \frac{a}{f_{a}} \frac{g_{s}^{2}}{32\pi^{2}} G\tilde{G} + \frac{1}{4} g_{a\gamma}^{0} aF\tilde{F} + \frac{\partial_{\mu} a}{2f_{a}} \bar{q} c_{q}^{0} \gamma^{\mu} \gamma_{5} q - \bar{q}_{L} M_{q} q_{R} + \text{ h.c.}$$

#### ★ Axion effective Lagrangian

#### The axion can be probed by $g_{a\gamma}$ and $g_{af}$ $f_a[\text{GeV}]$ $10^{19} \ 10^{18} \ 10^{17} \ 10^{16} \ 10^{15} \ 10^{14} \ 10^{13} \ 10^{12} \ 10^{11} \ 10^{10} \ 10^{9} \ 10^{8} \ 10^{7}$ CAST $10^{-10}$ ALPS-II **HB** hint **SN87A** $10^{-11}$ BabyIAXO IAXO Fermi-LAT $10^{-12}$ MADMAX **HAYSTACK** (galactic SN) $g_{a\gamma}[{\rm GeV}^{-1}]$ $10^{-13}$ superradiance $10^{-14}$ ABRA broad (PH. 1) SPEr $10^{-16}$ BH ABRA res. (PH. 1 Hadronic E/N=5/3 - 44/3 Hadronic E/N<170/3 $10^{-17}$ DFSZ I $10^{-18}$ DFSZ II 10-19 $10^{-12}10^{-11}10^{-10}10^{-9}10^{-8}10^{-7}10^{-6}10^{-5}10^{-4}10^{-3}10^{-2}10^{-1}$ $m_a[eV]$



#### ★ Axion effective Lagrangian



arXiv:2003.01100v4

**★** Axion effective Lagrangian

 $K^+ \to \pi^+ + a, \quad J/\psi \to a + \gamma$ 

The bound from particle physics experiment is  $f_a > 10^4 \text{GeV}$ 

However, for an axion being so light, the stringent bounds emerge from astrophysics and cosmology

 $10^9 \text{Gev} \le f_a \le 10^{12} \text{Gev}$ 

#### ★ Axion mass

The mass of axion comes from the anomaly term, it can be calculated by CHPT<sup>1</sup>

$$\mathcal{L}_{\text{eff}} = v^3 \text{Tr}(M\tilde{\Sigma} + M\tilde{\Sigma}^{\dagger}) + K\cos(\sqrt{6}\pi_9/f_9 + a/f_a)$$
$$\tilde{\Sigma} = \exp[i(2\pi_9/f_9\sqrt{6})I]\Sigma$$

The physical axion is the mixing among  $a, \pi^3, \pi^8, \pi^9$ , with mass

$$m_A^2 = \frac{K}{f_a^2} \frac{1}{1 + K \text{Tr} M^{-1} / 2v^3}$$

1. Kiwoon Choi et al., Physics Letters B, 181(1):145-149, 1986.

★ Axion mass

In QCD,  $\Lambda_{QCD} \sim 100 \text{Mev} \gg m_u, m_d \sim 1 \text{Mev}$ 

$$m_A^2 = \frac{2v^3}{f_a^2 \text{Tr} M^{-1}} = \frac{m_u m_d}{(m_u + m_d)^2} \frac{f_\pi^2 m_\pi^2}{f_a^2}$$

$$2(m_u + m_d)v^3 = f_\pi^2 m_\pi^2$$

This is the well-known formula for the axion mass.

### Axion and other open issues

#### ★ Axions and neutrino masses

The RH neutrino and the PQ symmetry breaking scales naturally fall in the same intermediate range  $M_R \sim f_a \sim 10^9 \text{Gev} - 10^{12} \text{Gev}$ 

A common origin for mass of RH neutrino and axion decay constant:  $N_R N_R \Phi$ 

Classical article: M. Shin, Phys. Rev. Lett. 59, 2515 (1987)

Composite axion and sterile neutrino: arXiv:2310.08557v1

Flavor violated axion neutrino model: arXiv:2408.05903v1

 $0.0011 \text{eV} \lesssim m_3 \lesssim 0.0029 \text{eV}$ 

### Axion and other open issues

#### **★** Axions and the baryon asymmetry

Axion carries PQ charge and generates net PQ charge, and these charge can be transformed to net baryon number

Axiogenesis: 10.1103/PhysRevLett.124.111602

Affleck-Dine baryogenesis: arXiv:1906.05286

### Axion and other open issues

#### ★ Axions and inflation

The potential of axion is slow, it can play the role as inflaton

Natural Inflation: 10.1103/PhysRevLett.114.151303

SMASH model : arXiv:1610.01639