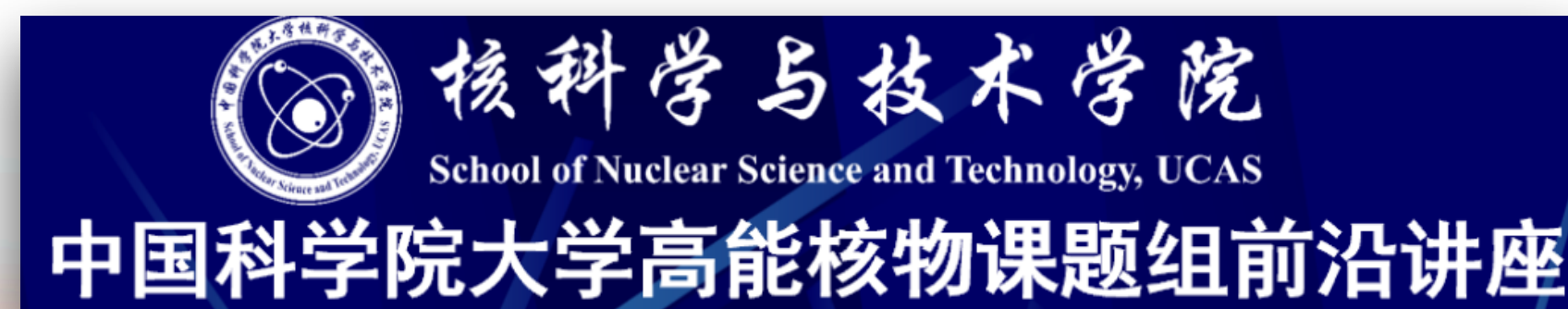


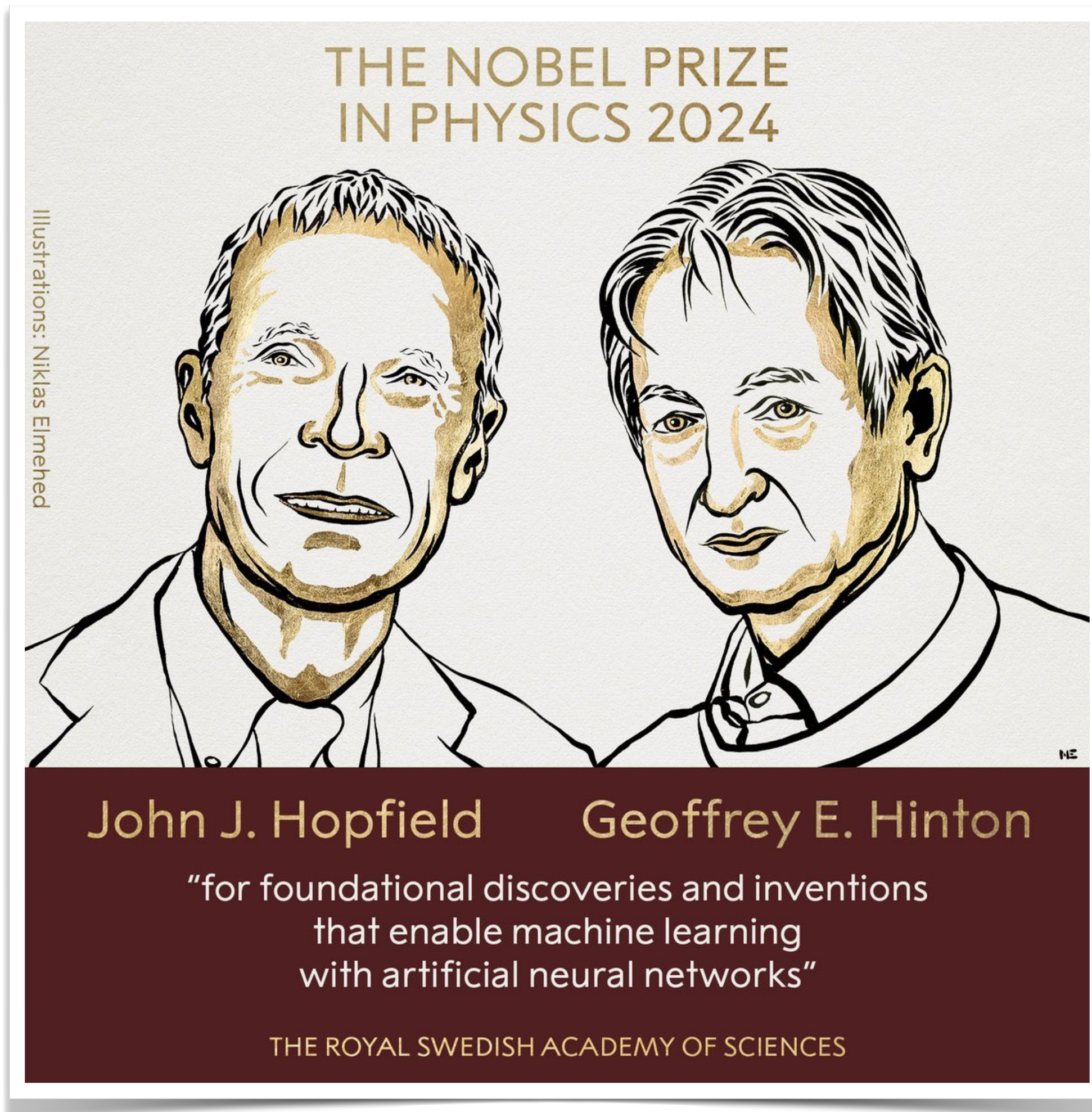
Physics-Driven Learning for Solving Inverse Problems towards QCD Physics

Lingxiao Wang(王凌霄)
RIKEN-iTHEMS



Arxiv:202412.xxxxx, arxiv:202410.03082, Prog.Part.Nucl.Phys. 104084(2023);
Phys. Rev. D 103, 116023, Phys. Rev. C 106, L051901, Phys. Rev. D 107, 083028, Phys. Rev. D 106, L051502;
Chin. Phys. Lett. 39, 120502, Phys. Rev. D 107, 056001, JHEP05(2024)060.

Nobel Prize in Physics 2024



For physicists,
**it is the best of times,
it is the worst of times.**

DEEP-IN Working Group

[Concept](#)
[Activities](#)
[Facilitators](#)
[Members](#)
[Contact](#)

DEEP-IN SCIENCE

深入 科学

CONCEPT

“**DEEP** learning for **IN**verse problems (**DEEP-IN**)” in Sciences Working Group

The essence of discovery in sciences has always been rooted in the reverse engineering of natural phenomena and observational data. This paradigm of deducing the underlying laws of nature from observable outcomes forms the cornerstone of our scientific inquiry. **The DEEP-IN working group is established with the recognition that the elucidation of such complex phenomena demands a fusion of physics insights and advanced deep learning methodologies.**

In response to the evolving landscape of scientific research, our objective is to integrate cutting-edge **deep learning techniques, alongside generative models and other advanced statistical learning methods**, into the toolkit of scientists.

The DEEP-IN working group at [RIKEN-iTHEMS](#) is dedicated to creating an interdisciplinary platform that harnesses the transformative potential of artificial intelligence(AI). This platform is designed to **tackle inverse problems that span a diverse spectrum of sciences, from biology to physics and more in the future.**

<https://sites.google.com/view/deep-in-wg/homepage>

iTHEMS

理化学研究所 数理創造プログラム
RIKEN Interdisciplinary Theoretical and Mathematical Sciences Program

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DEEP-IN Working Group

“DEEP learning for INverse problems (DEEP-IN) in Sciences” working group (April 1st, 2024 -)

Lattice Computations

Gert Aarts, Swansea U.
Takumi Doi, iTHEMS
Andreas Ipp, TU Wien
Tetsuo Hatsuda, iTHEMS
Yan Lyu, iTHEMS

Now mostly physicists -> Future more diverse scientists

BioPhysics: Catherine Beauchemin, iTHEMS
Condensed Matter Physics: Steffen Backes, iTHEMS
QCD Physics: Kenji Fukushima, UTokyo
Nuclear Physics: Haozhao Liang, UTokyo
Quantum Computing: Enrico Rinaldi, Quantinuum K.K./iTHEMS

Heavy-Ion Collisions

Long-Gang Pang, CCNU
Shuzhe Shi, THU
Kai Zhou, CUHK-ShenZhen

Astrophysics

Márcio Ferreira, Coimbra U.
Yuki Fujimoto, INT->iTHEMS
Akira Harada, NIT-Ibaraki
Zhenyu Zhu, TDLI->RIT

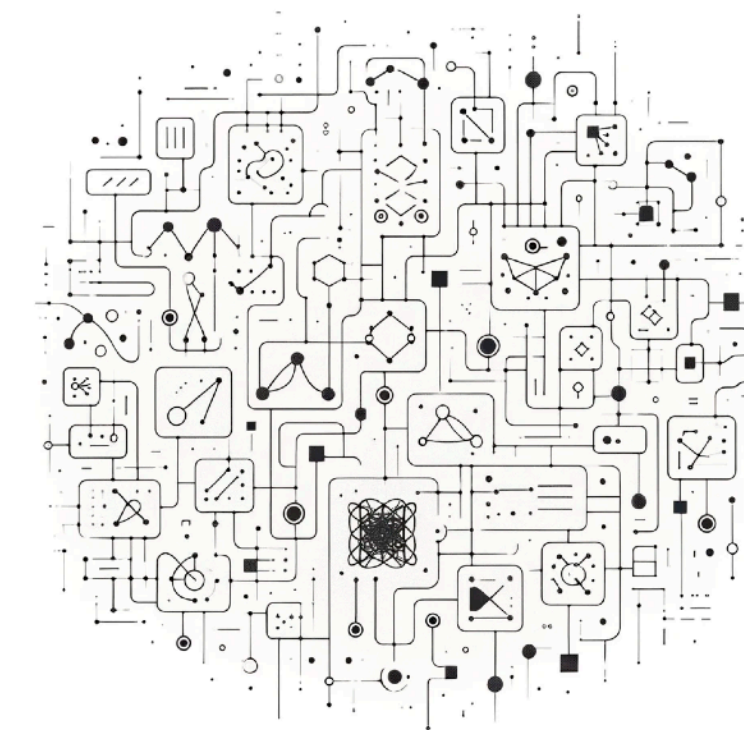
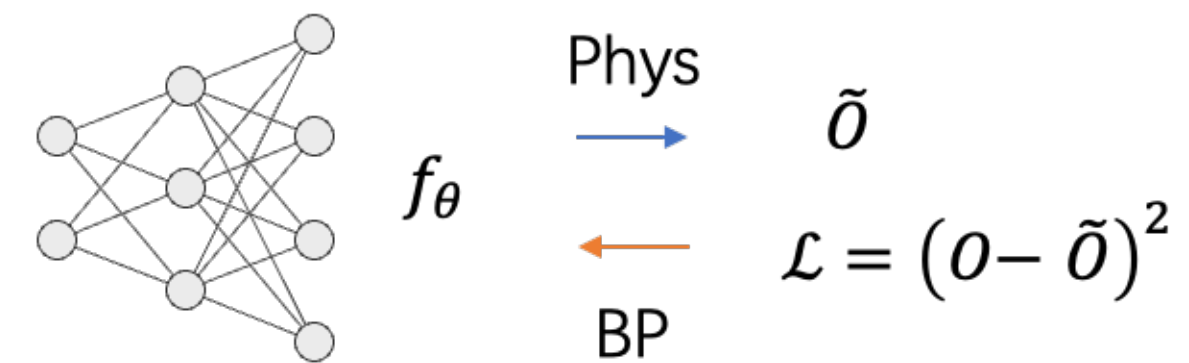
Machine Learning

Akinori Tanaka, AIP/iTHEMS
Lingxiao Wang, iTHEMS

Lingxiao Wang (RIKEN iTHEMS) *Contact at lingxiao.wang@riken.jp

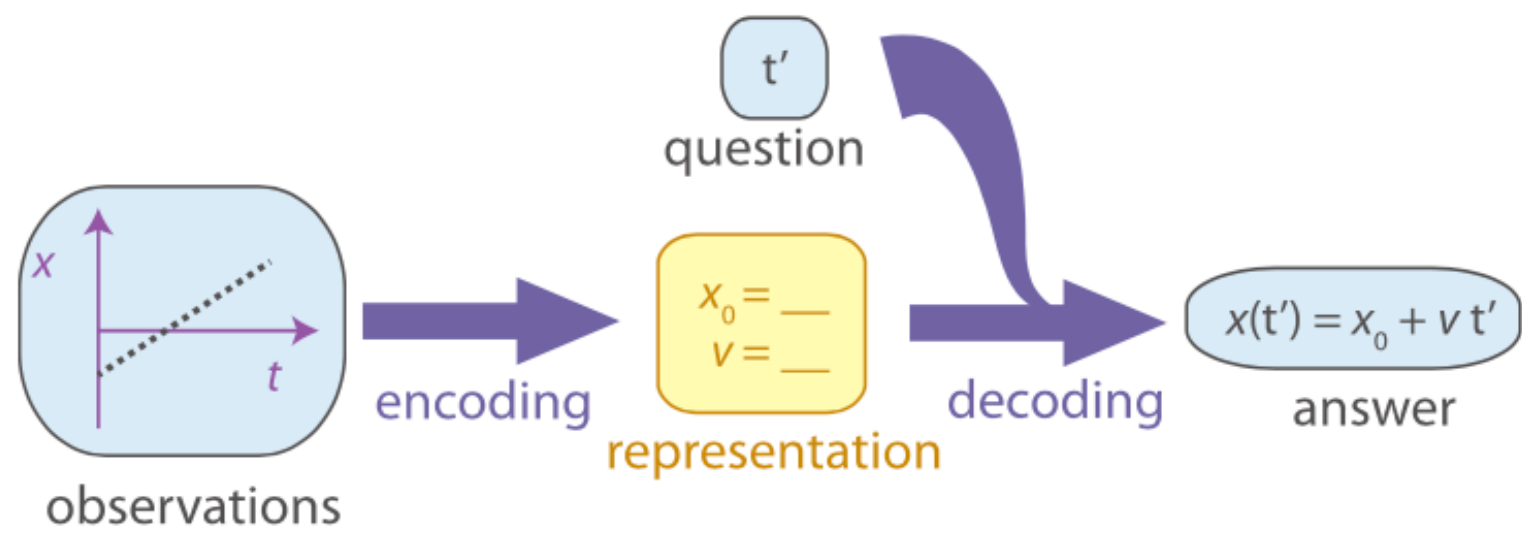
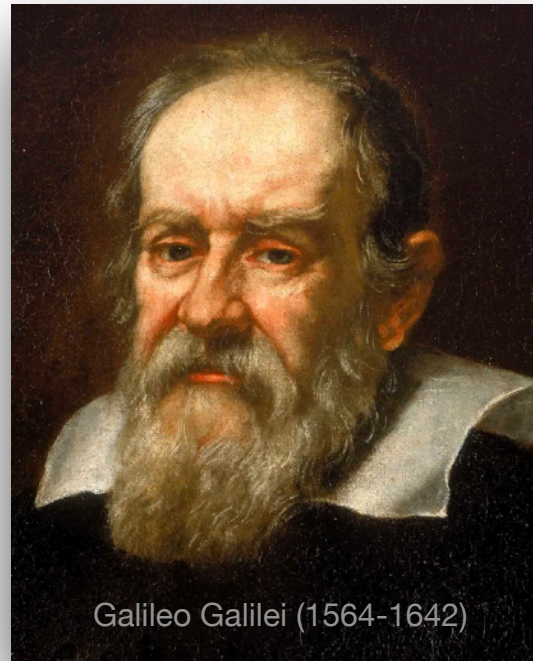
Outline

- **Physics for Machine Learning**
- **Inverse Problems**
 - Data-Driven Learning
 - Physics-Driven Learning
- **Generative Models**
 - Learn to Sample
 - Physics-Driven Generative Models
- **Outlooks**



Generated by ChatGPT-4 + DALL·E

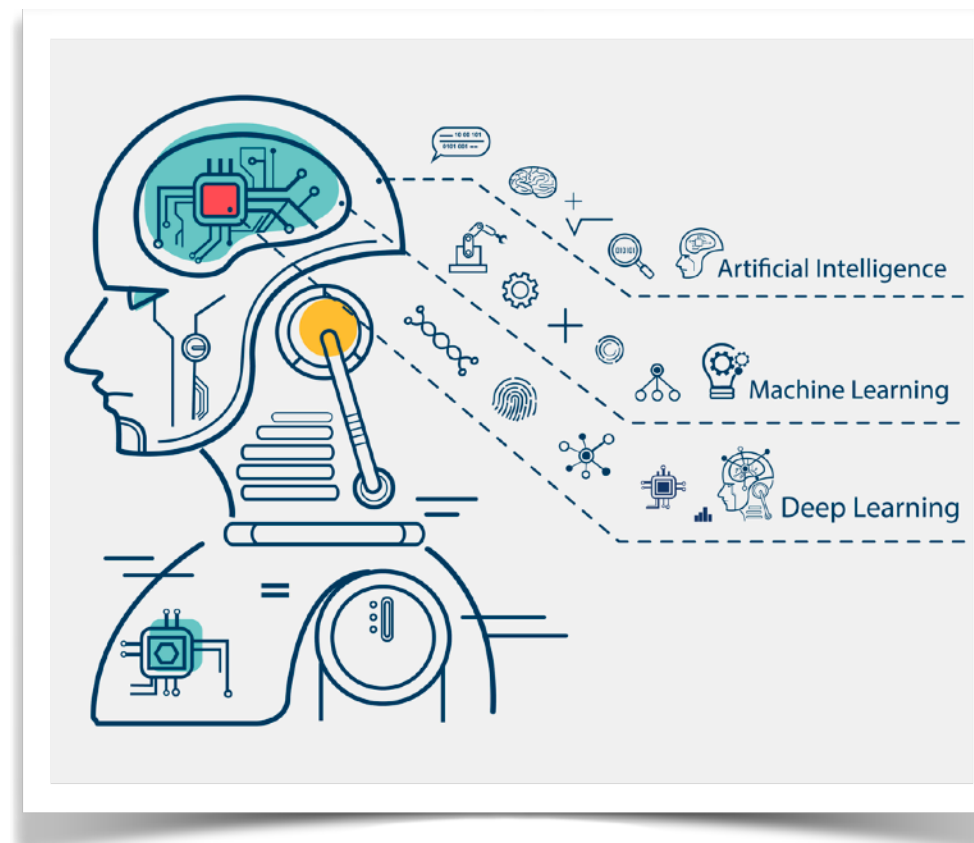
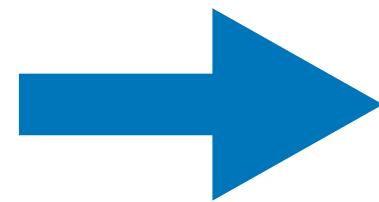
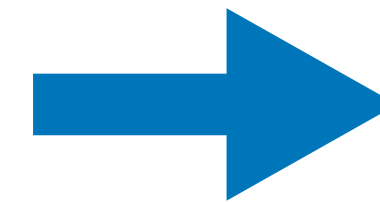
Machine Learning and Physics

Phys.Rev. Lett. **124**, 010508 (2020)

An **inverse problem** in science is the process of **inferring** from a set of **observations** the **causal factors** that produced them.



Data, X

Machine, $\{\theta\}$ 

Prediction $p(X | \theta)$

Estimation

$$\hat{\theta} = \mathbf{arg\ max}_{\theta} \{p(X | \theta)\}$$

Machine Learning and Inference

Maximum Likelihood Estimation(MLE)

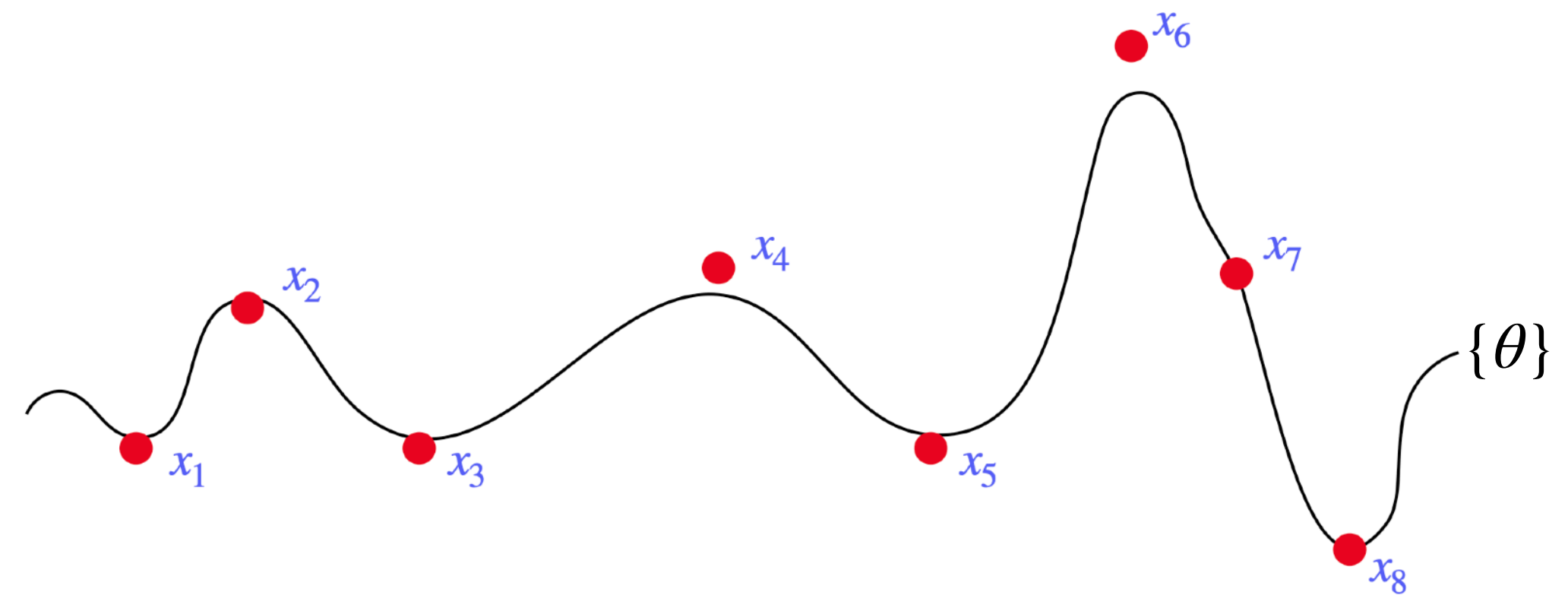
$$\max_{\theta} \prod_{i=1}^N p(\mathbf{x}_i | \theta)$$

Bayesian
Inference

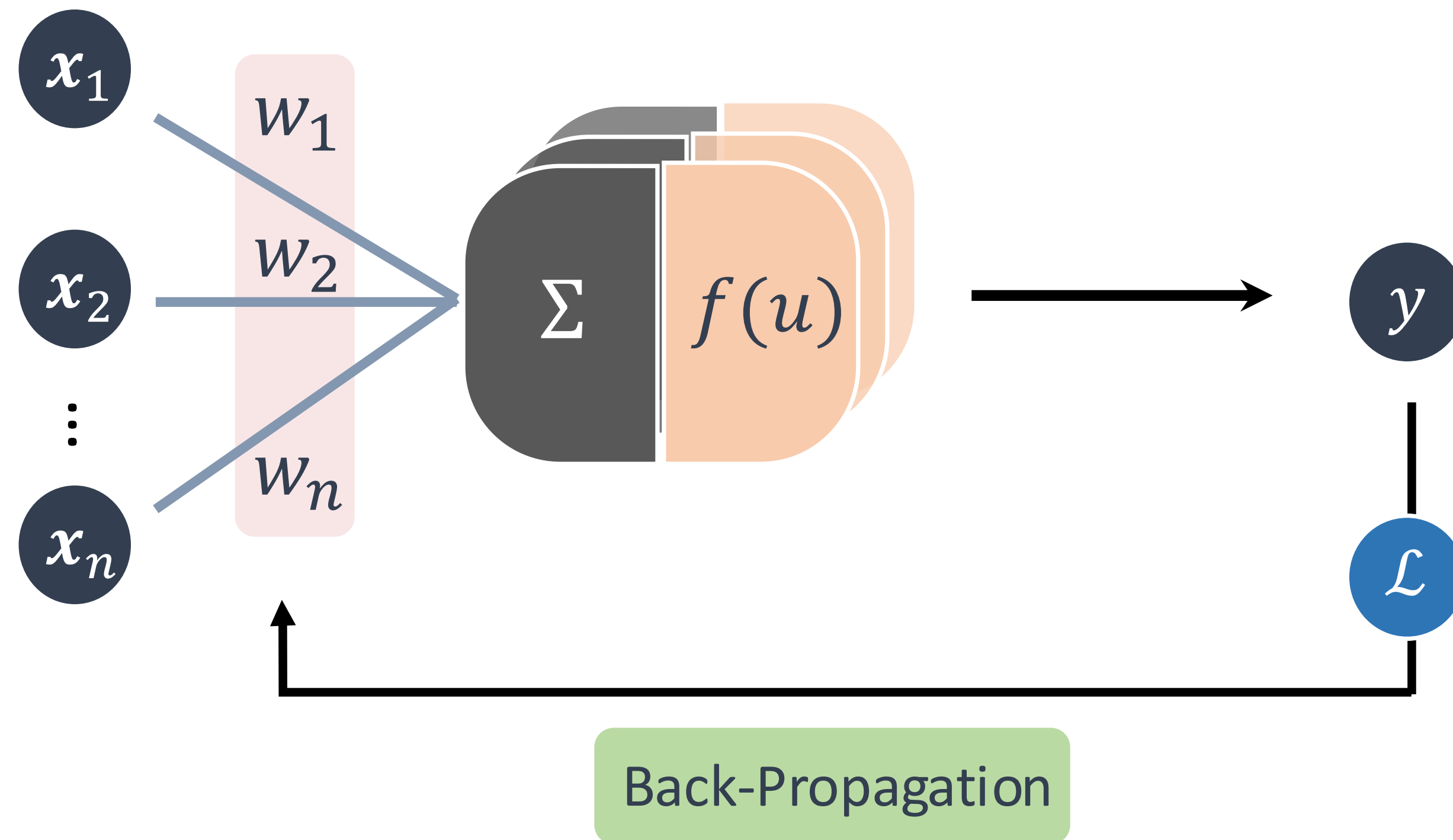
Maximum A Posterior(MAP)

$$p(\theta | X) = \frac{p(X | \theta)\pi(\theta)}{p(X)}$$

Posterior $p(\theta | X)$, Prior $\pi(\theta)$, Evidence $p(X)$



Deep Model as Machine



Deep (neural network) Model

- **Inputs**, $\{x\} = x_1, x_2, \dots, x_n$
- **Weights**, $\{w\} = w_1, w_2, \dots, w_n$
- **Outputs**, y
- **Summation**, $\Sigma(\cdot)$
- **Non-Linear Activation Functions**, $f(u)$
- **Single Layer** $y = f(\sum_{i=1}^n x_i w_i)$

Objective

- **Loss Function**, $\mathcal{L}(y, \hat{y})$
- **Data**, $\hat{y}$

Optimization Algorithm

- **Back-Propagation**, $\frac{\partial \mathcal{L}}{\partial \omega}$
- **Stochastical Methods**: SGD, Adam, ...

Physics-Driven Designs

Inputs

x_1

x_2

$\vdots$

x_n

w_1

w_2

w_n

Σ

$f(u)$

y

Outputs

$\mathcal{L}$

Symmetry

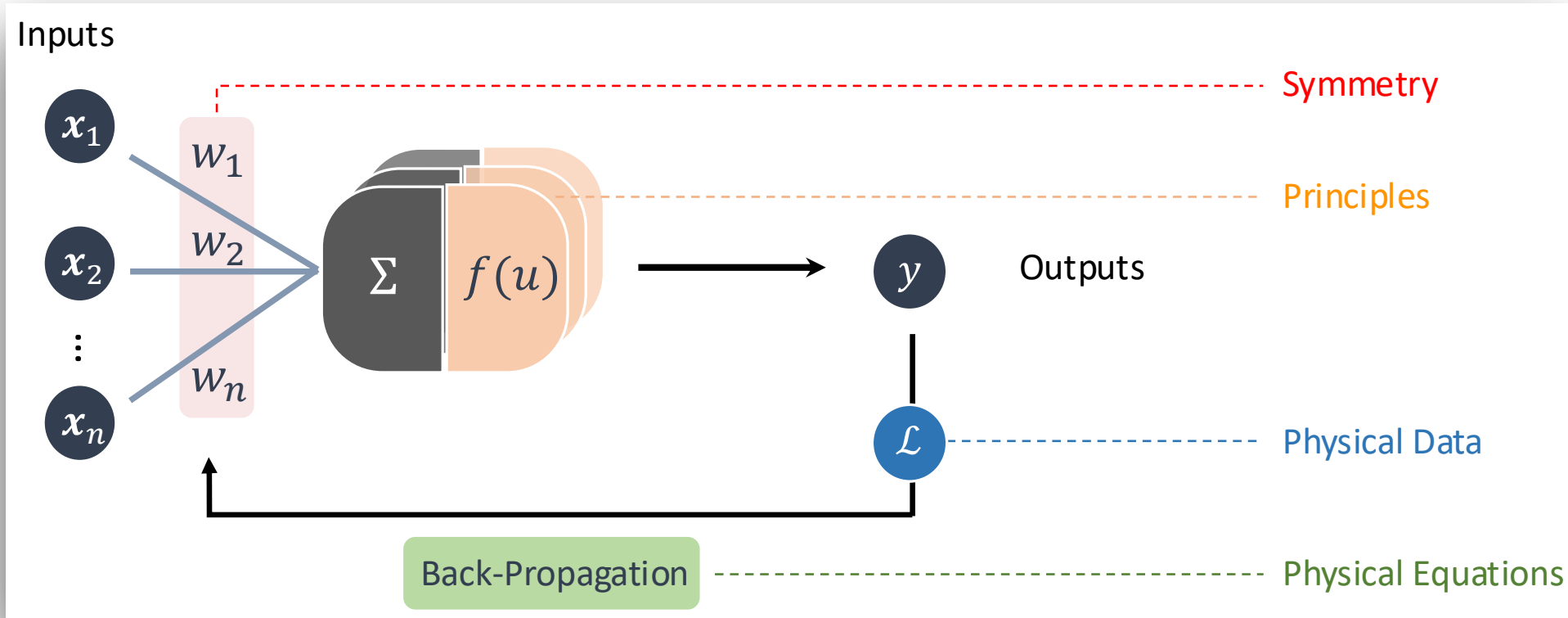
Principles

Physical Data

Back-Propagation

Physical Equations

Physics-Driven Designs: Symmetry

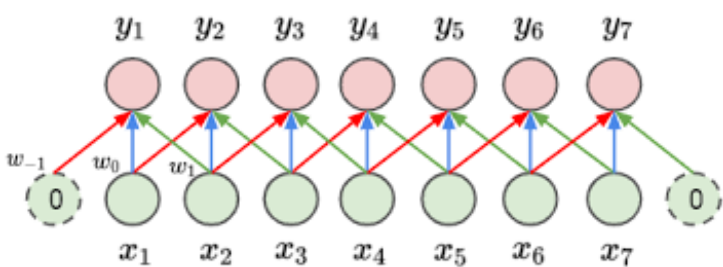


Embedding **symmetries** introduces a scheme for **sharing parameters** in deep models

Convolutional Neural Networks (CNNs)

$$Y_{i,j,c} = \sum_{m=1}^{k_h} \sum_{n=1}^{k_w} \sum_{p=1}^{C_{in}} X_{i+m,j+n,p} \cdot K_{m,n,p,c} + b_c$$

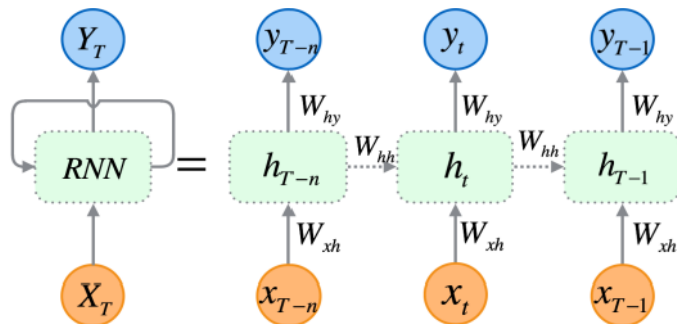
translational symmetry



Recurrent Neural Networks (RNNs)

$$h_t = \sigma(W_x x_t + W_h h_{t-1} + b)$$

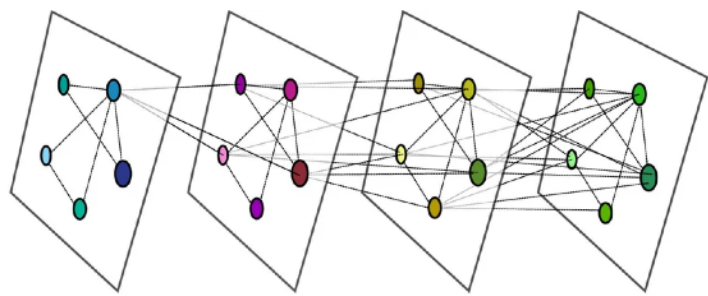
temporal symmetry



Graph Neural Networks (GNNs)

$$h_i^{(l+1)} = \sigma \left(W^{(l)} \cdot \sum_{j \in \mathcal{N}(i)} h_j^{(l)} + b^{(l)} \right)$$

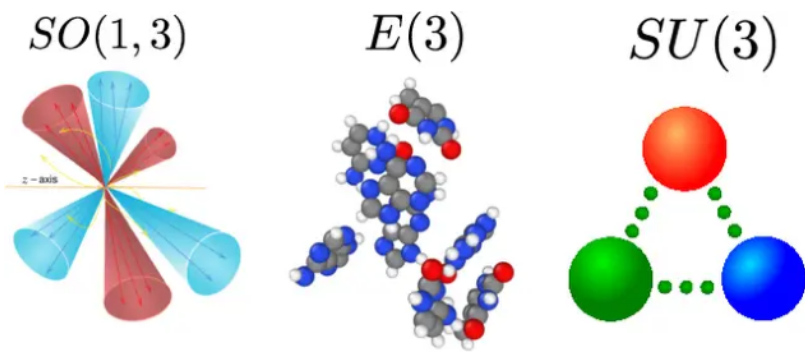
permutational symmetry



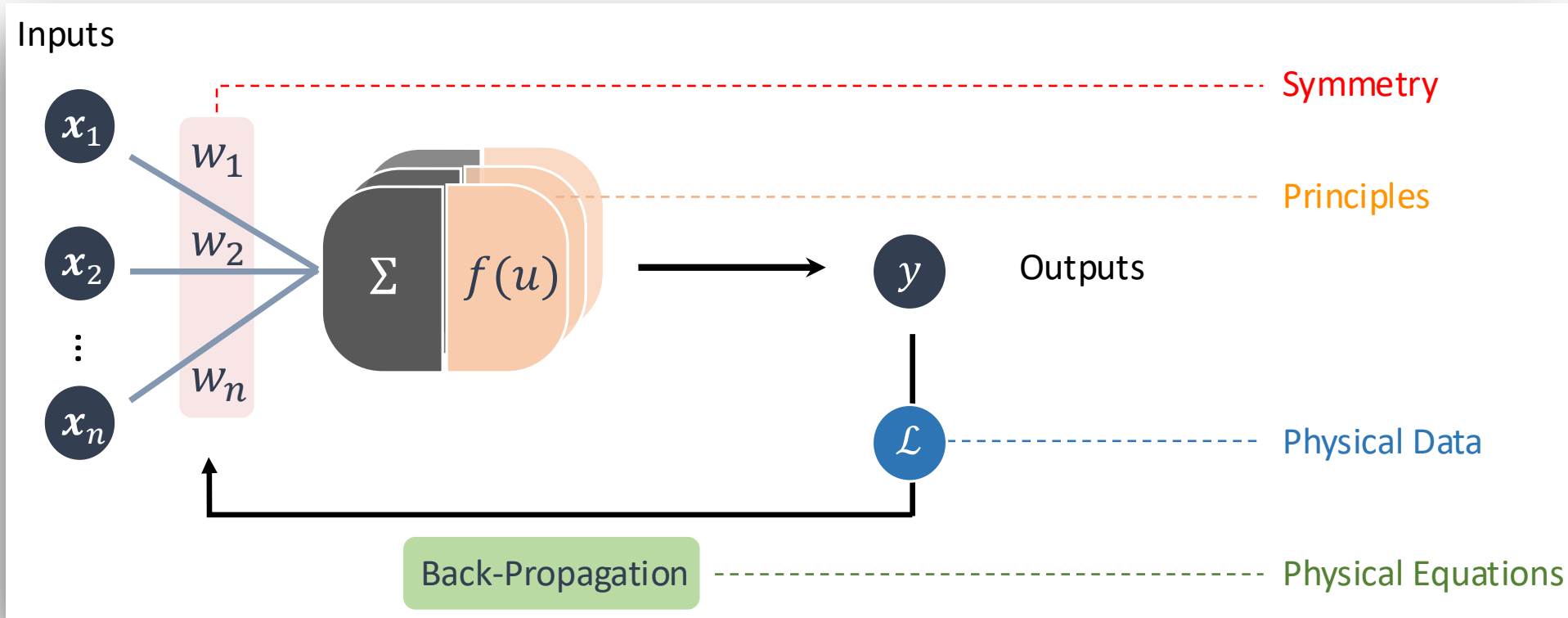
Equivariant Neural Networks (G-CNNs)

$$Y_g = \sum_{m=1}^{k_h} \sum_{n=1}^{k_w} X_g(i+m,j+n) \cdot K(m,n)$$

groups



Physics-Driven Designs: Symmetry



Embedding **symmetries** introduces a scheme for **sharing parameters** in deep models

Convolutional Neural Networks (CNNs)

$$Y_{i,j,c} = \sum_{m=1}^{k_h} \sum_{n=1}^{k_w} \sum_{p=1}^{C_{in}} X_{i+m,j+n,p} \cdot K_{m,n,p,c} + b_c$$

- X : Input feature map.
- K : Convolution filter.
- Y : Output feature map.

Recurrent Neural Networks (RNNs)

$$h_t = \sigma(W_x x_t + W_h h_{t-1} + b)$$

- x_t : Input at time t .
- h_t : Hidden state at time t .
- W_x and W_h : Shared weight matrices across time.

Graph Neural Networks (GNNs)

$$h_i^{(l+1)} = \sigma \left(W^{(l)} \cdot \sum_{j \in \mathcal{N}(i)} h_j^{(l)} + b^{(l)} \right)$$

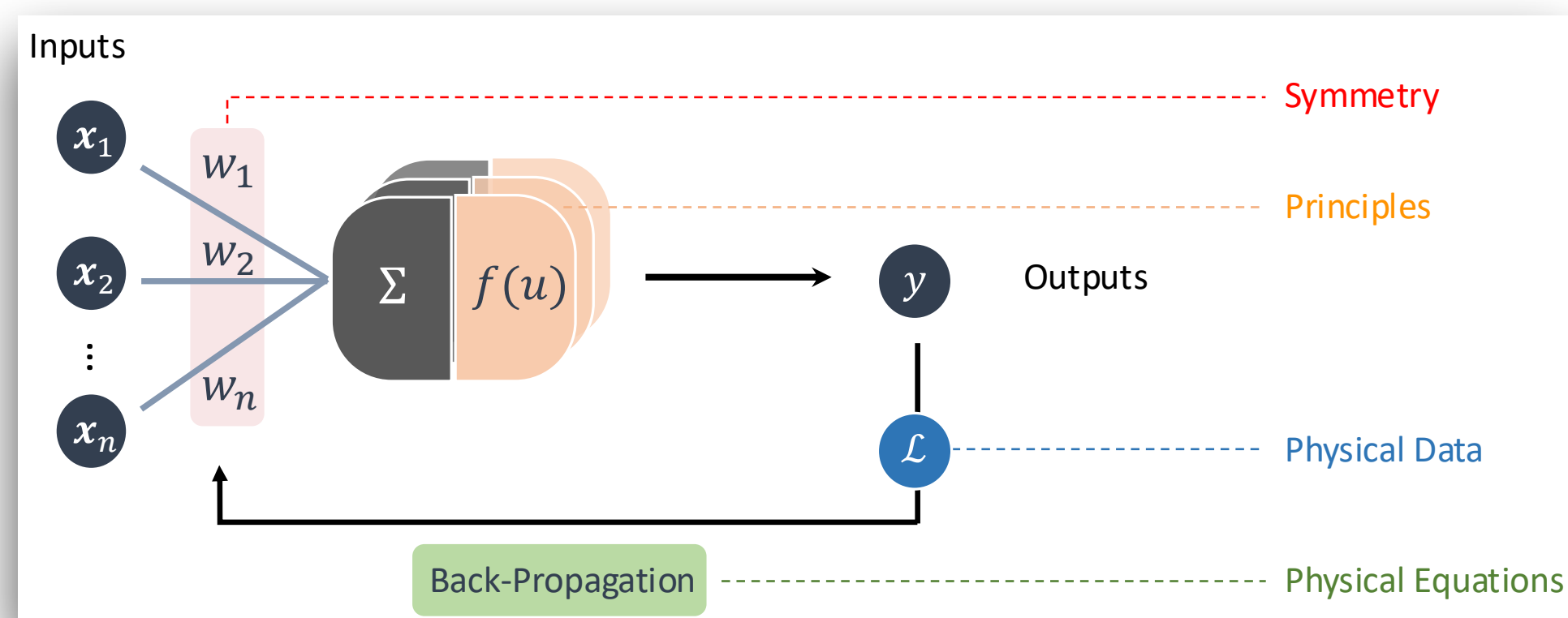
- $h_i^{(l)}$: Feature vector for node i at layer l .
- $W^{(l)}$: Shared weight matrix across all nodes.

Equivariant Neural Networks (G-CNNs)

$$Y_g = \sum_{m=1}^{k_h} \sum_{n=1}^{k_w} X_g(i+m, j+n) \cdot K(m, n)$$

- $g \in G$: Transformation from the group G (e.g., rotations).
- X_g : Input after applying transformation g .

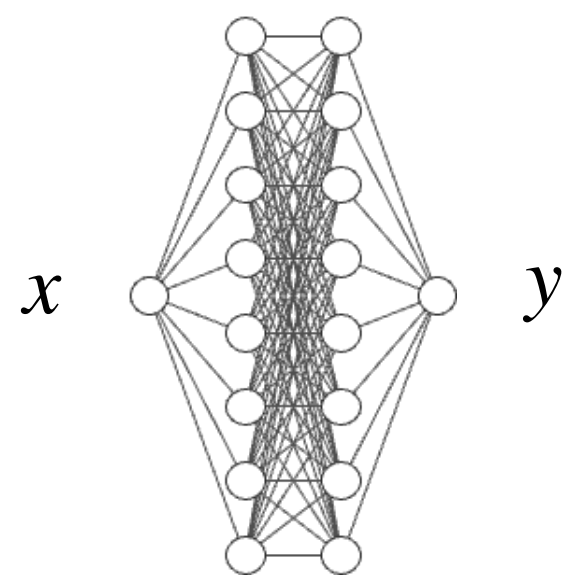
Physics-Driven Designs: Principles



Principles can enforce the outputs
are **physically meaningful**

Continuity

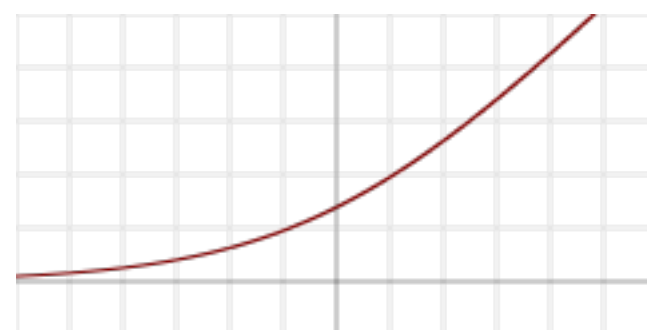
$$y := f(x)$$



Positive Definiteness

Softplus

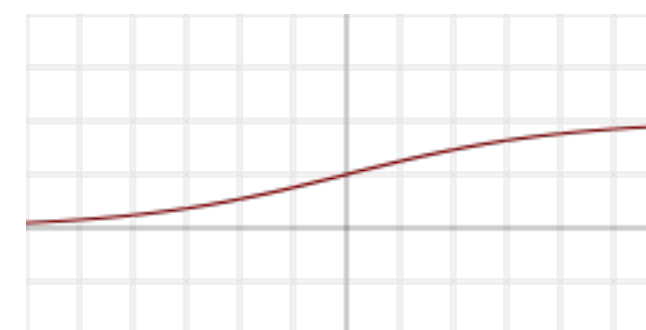
$$\ln(1 + e^x)$$



Causality

Sigmoid $\frac{1}{1 + e^{-x}}$

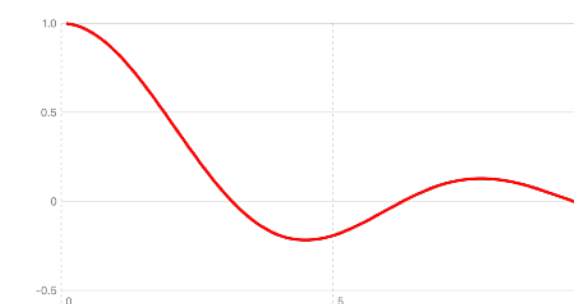
$$\frac{dp}{d\epsilon} = \frac{c_s^2}{c^2} < 1$$



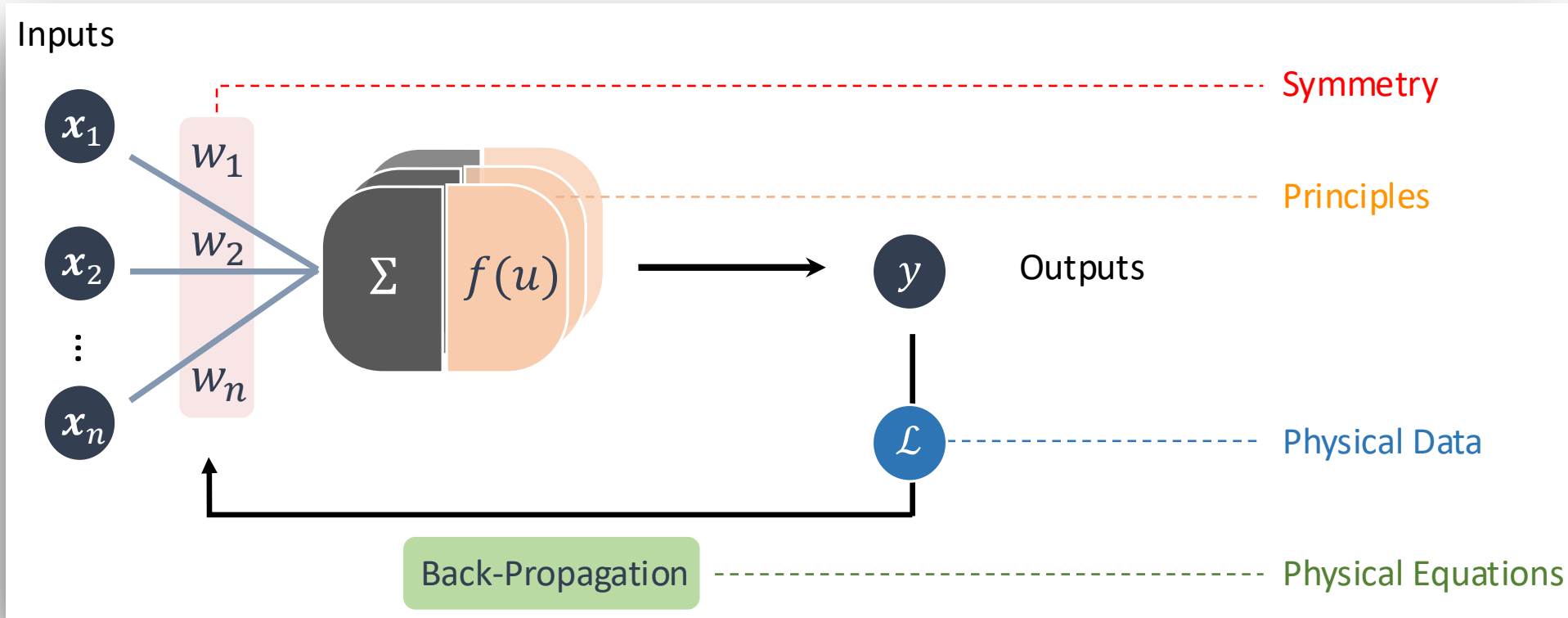
Asymptotic Behaviors

Scattering Wave Function

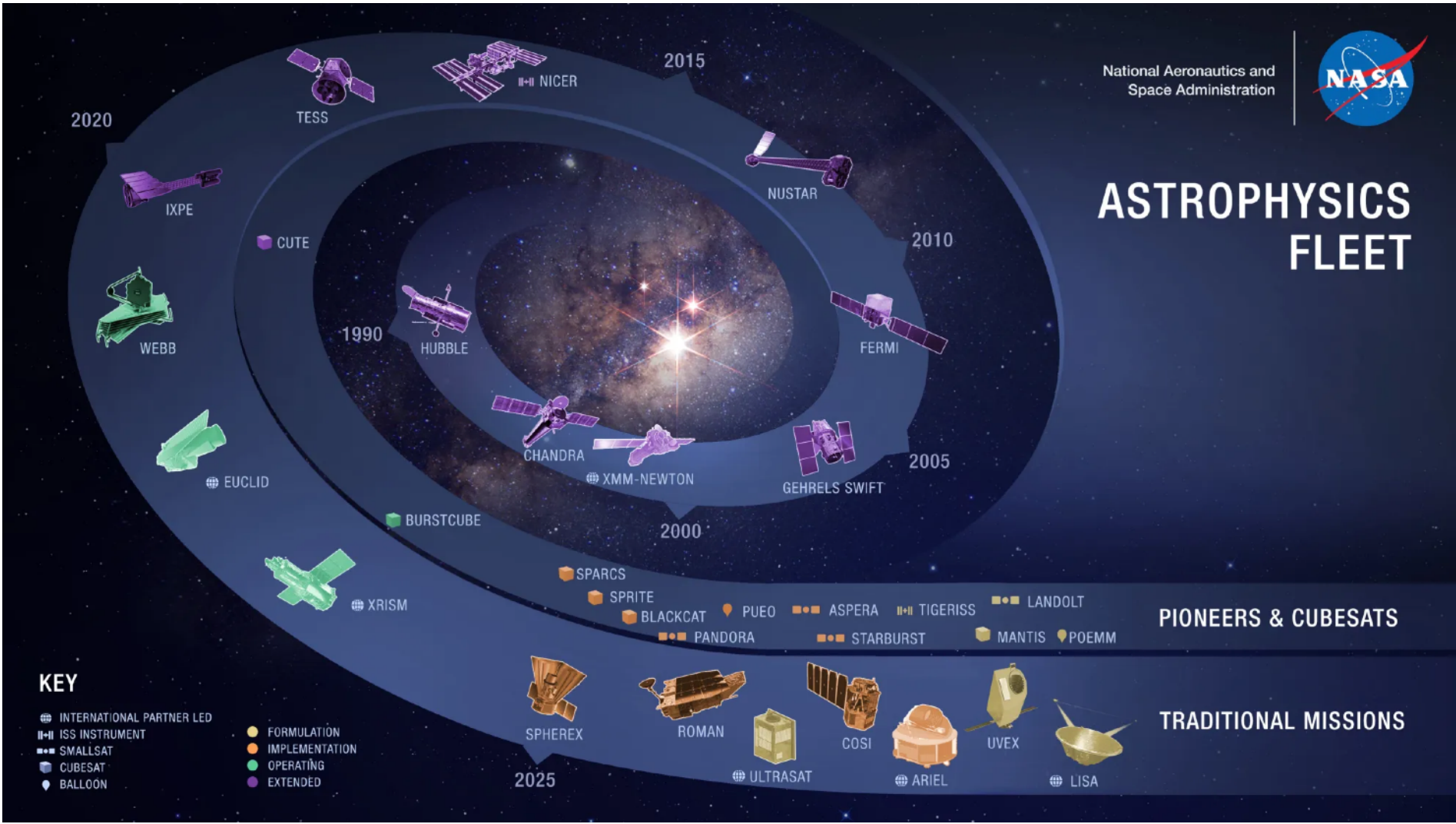
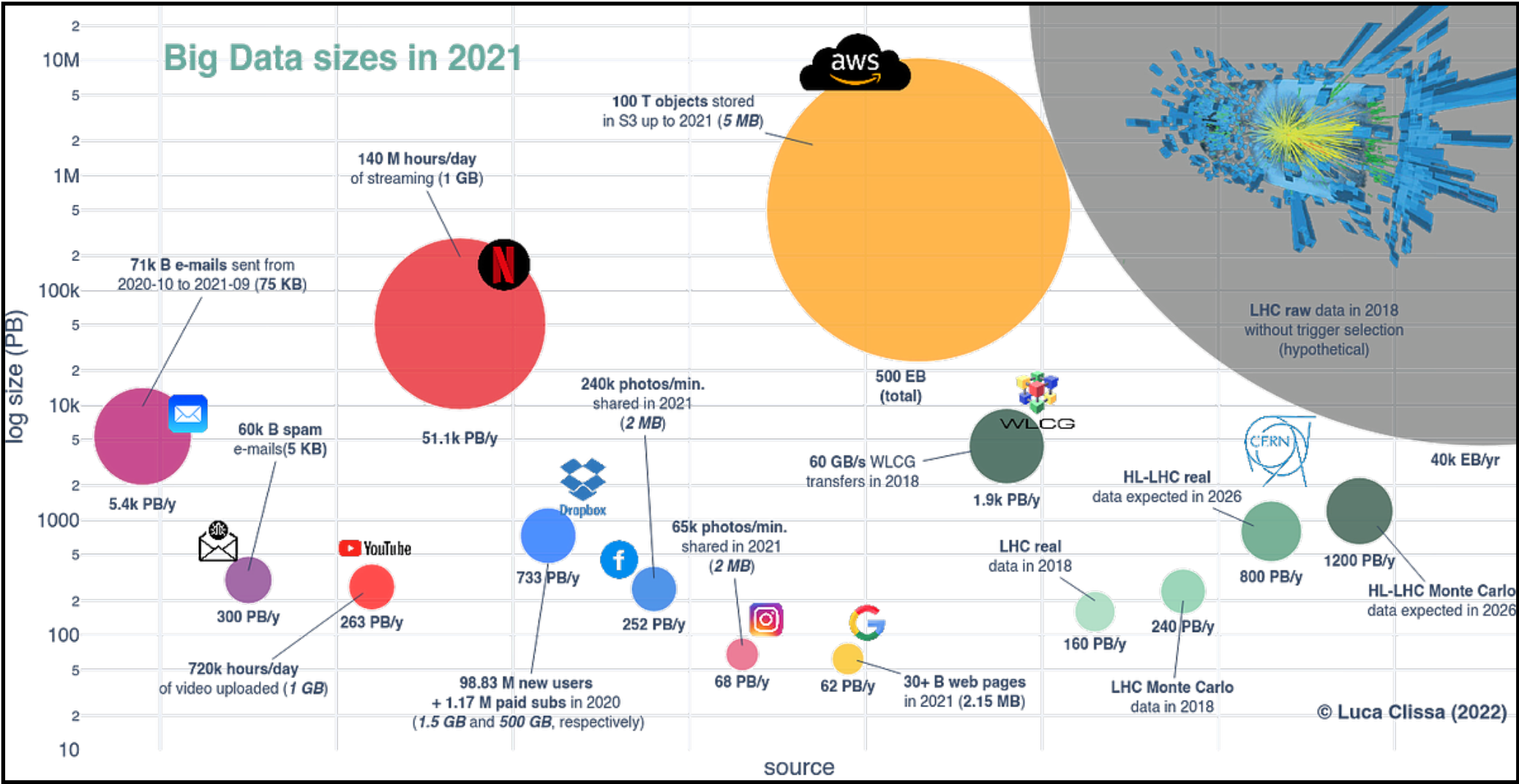
$$A \frac{\sin(kr)}{r} + f(r)$$



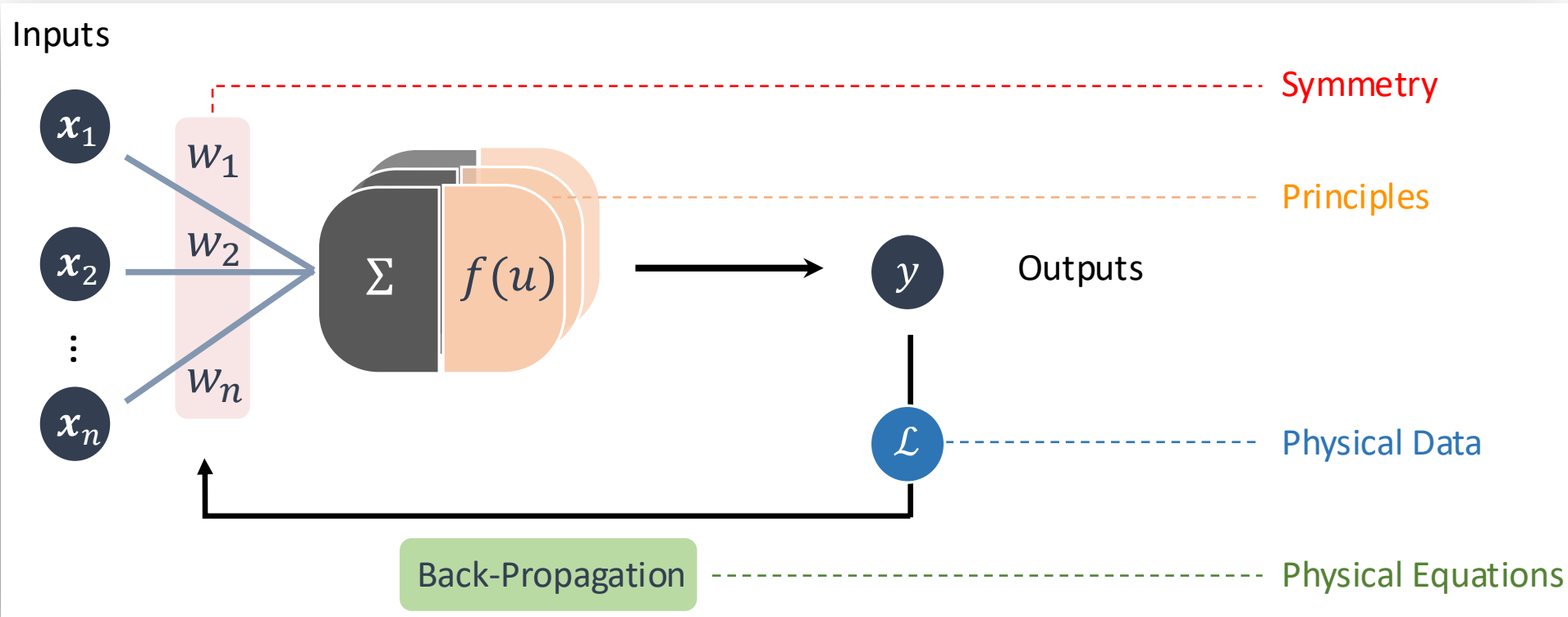
Physics-Driven Designs: Data



Physical data from experiments or simulations can align the model **outputs** with physical truths



Physics-Driven Designs: Equations



Physical equations can be encoded into **back-propagation** through automatic differentiation

Wave equation

$$\frac{\partial^2 u}{\partial t^2} = c^2 \frac{\partial^2 u}{\partial x^2}$$

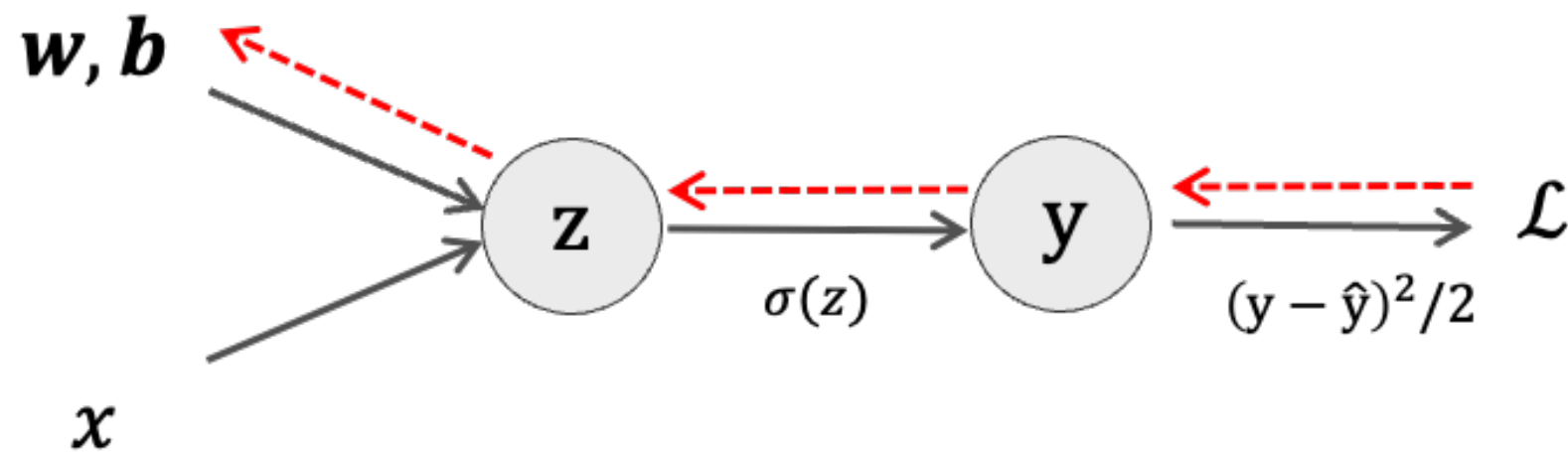
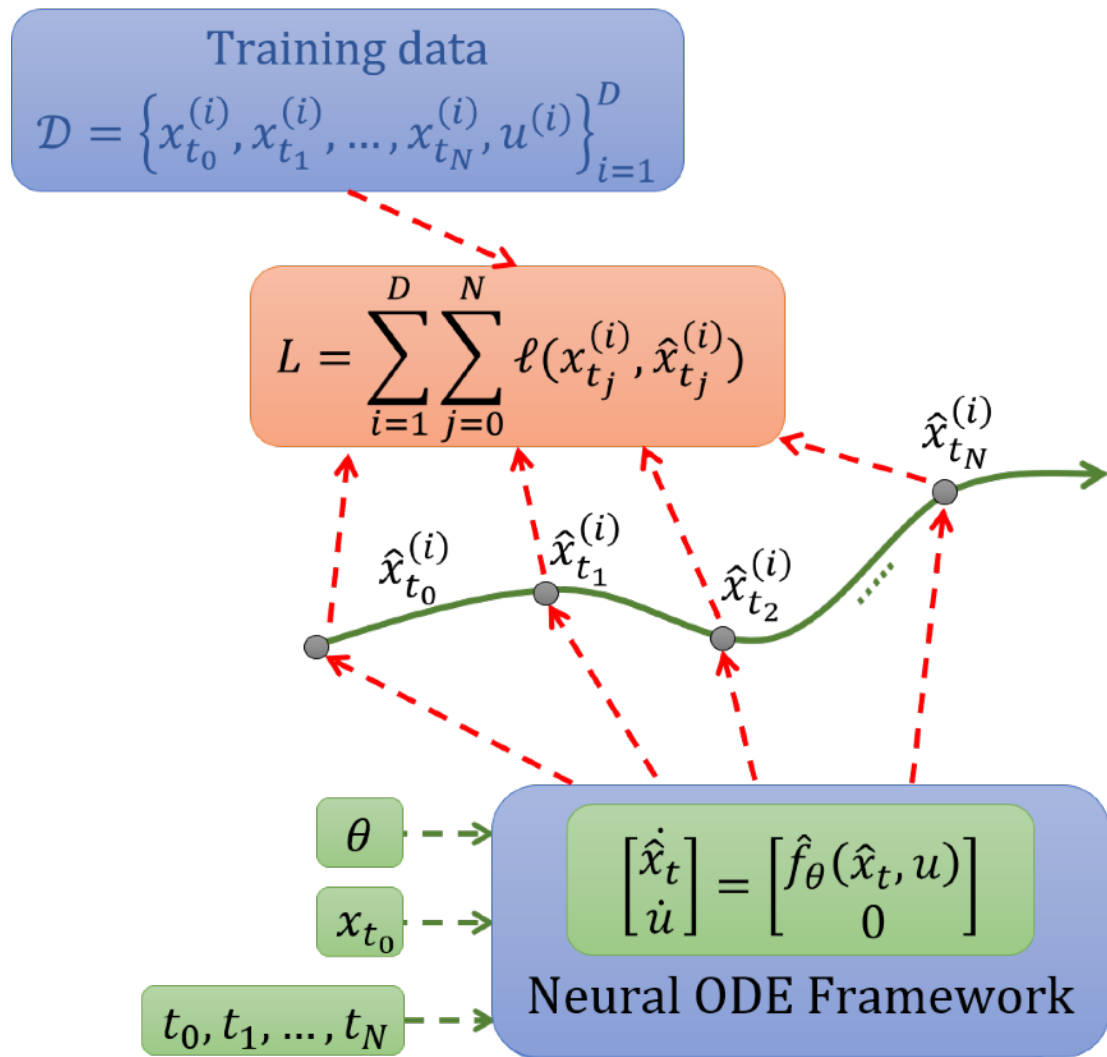
Maxwell's equations

$$\nabla \cdot \mathbf{E} = 0 \quad \nabla \times \mathbf{E} = -\frac{1}{c} \frac{\partial \mathbf{H}}{\partial t}$$

$$\nabla \cdot \mathbf{H} = 0 \quad \nabla \times \mathbf{H} = \frac{1}{c} \frac{\partial \mathbf{E}}{\partial t}$$

Schrödinger's equation

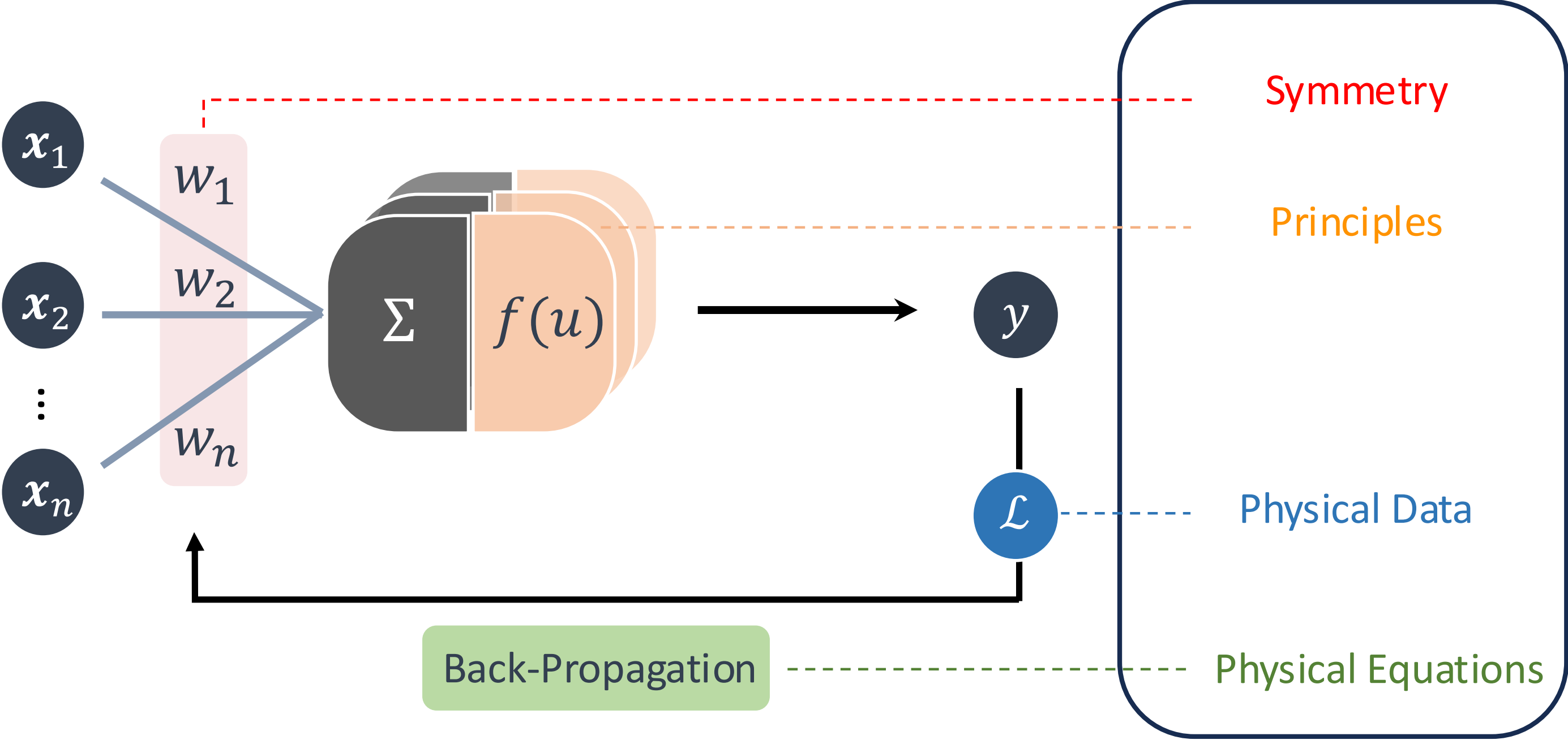
$$i\hbar \frac{\partial}{\partial t} \psi = \hat{H} \psi$$



Easy-To-Compute on GPUs

Physics-Driven Desgins

accepted at *Nature Reviews Physics*



Physics-Driven Learning for Solving Inverse Problems towards QCD Physics

Gert Aarts¹, Kenji Fukushima², Tetsuo Hatsuda³, Andreas Ipp⁴, Shuzhe Shi⁵, Lingxiao Wang^{3,*}, and Kai Zhou^{6,7}

¹Department of Physics, Swansea University, SA2 8PP, Swansea, United Kingdom
²Department of Physics, The University of Tokyo, 7-3-1 Hongo, Bunkyo-ku, Tokyo, 113-0033, Japan
³Interdisciplinary Theoretical and Mathematical Sciences Program (iTHEMS), RIKEN, Wako, Saitama 351-0198, Japan
⁴Institute for Theoretical Physics, TU Wien, Wiedner Hauptstraße 8-10/136, A-1040 Vienna, Austria
⁵Department of Physics, Tsinghua University, Beijing 100084, China
⁶School of Science and Engineering, The Chinese University of Hong Kong, Shenzhen (CUHK-Shenzhen), Guangdong, 518172, China
⁷Frankfurt Institute for Advanced Studies, Ruth Moufang Strasse 1, D-60438, Frankfurt am Main, Germany
*e-mail: lingxiao.wang@riken.jp

ABSTRACT

The integration of deep learning techniques and physics-driven designs is reforming the way we address inverse problems, in which accurate physical properties are extracted from complex data sets. This is particularly relevant for quantum chromodynamics (QCD), the theory of strong interactions, with its inherent limitations in observational data and demanding computational approaches. This perspective highlights advances and potential of physics-driven learning methods, focusing on predictions of physical quantities towards QCD physics, and drawing connections to machine learning (ML). It is shown that the fusion of ML and physics can lead to more efficient and reliable problem-solving strategies. Key ideas of ML, methodology of embedding physics priors, and generative models as inverse modelling of physical probability distributions are introduced. Specific applications cover first-principle lattice calculations, and QCD physics of hadrons, neutron stars, and heavy-ion collisions. These examples provide a structured and concise overview of how incorporating prior knowledge such as symmetry, continuity and equations into deep learning designs can address diverse inverse problems across different physical sciences.

Physics Knowledge for Designing
More Controllable and Reliable Deep Models

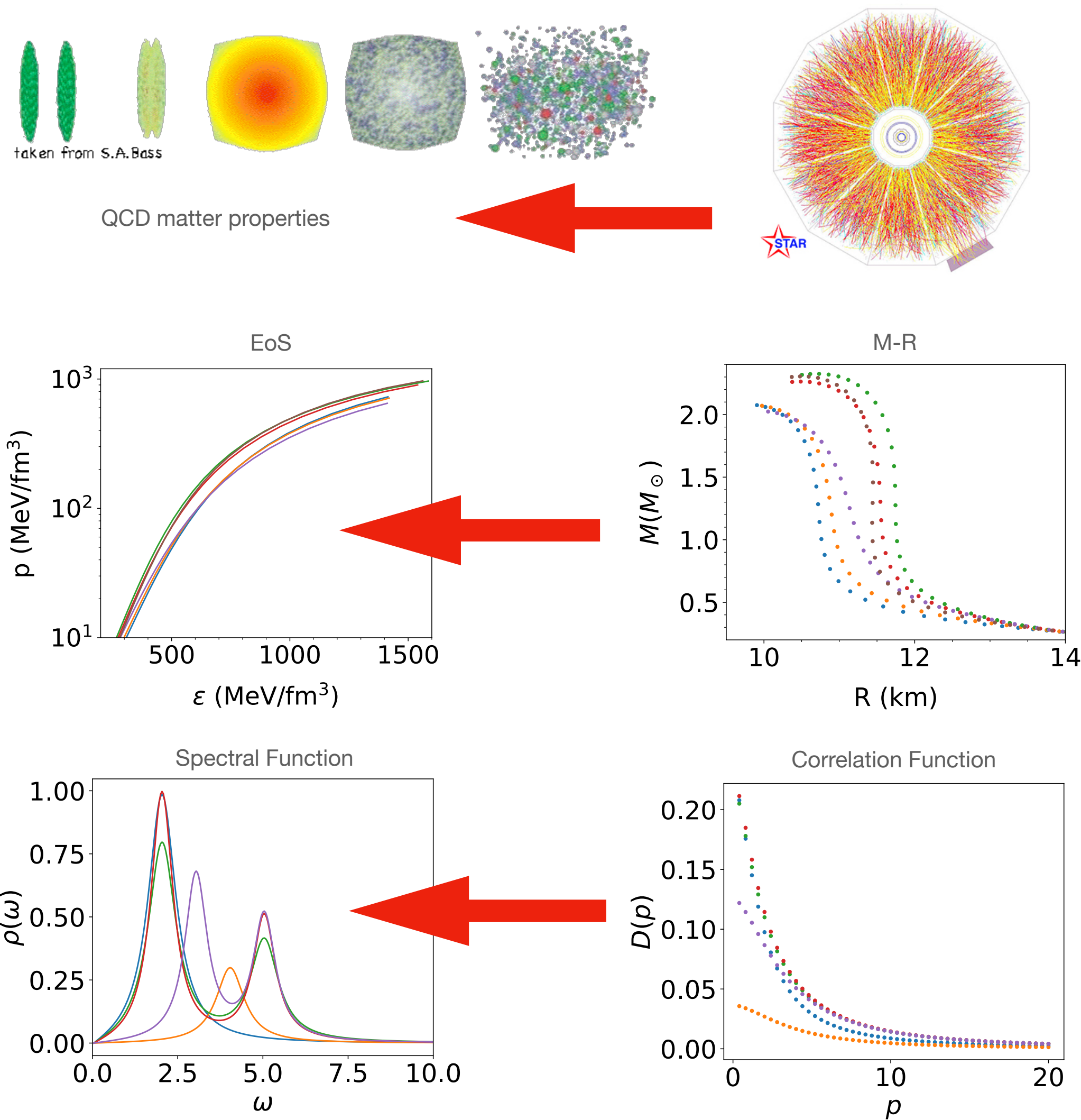
Inverse Problems in QCD Physics

Inverse Problems

Physics

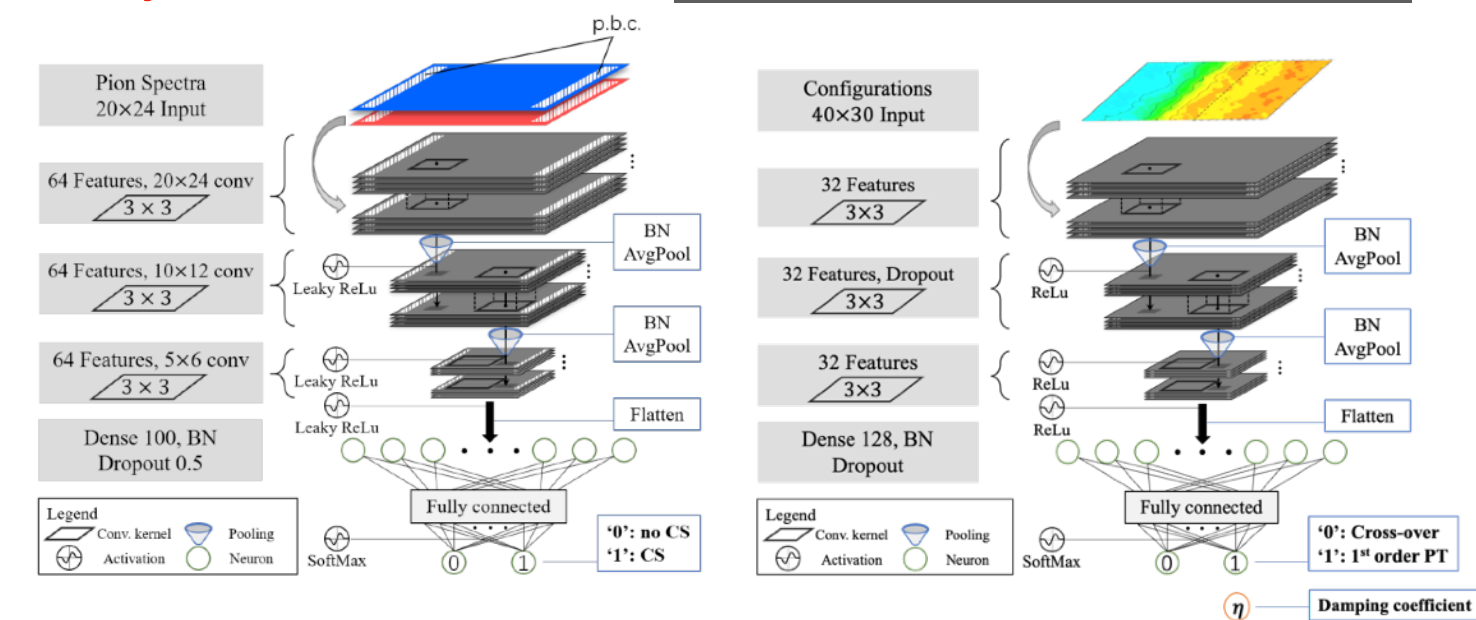


Data



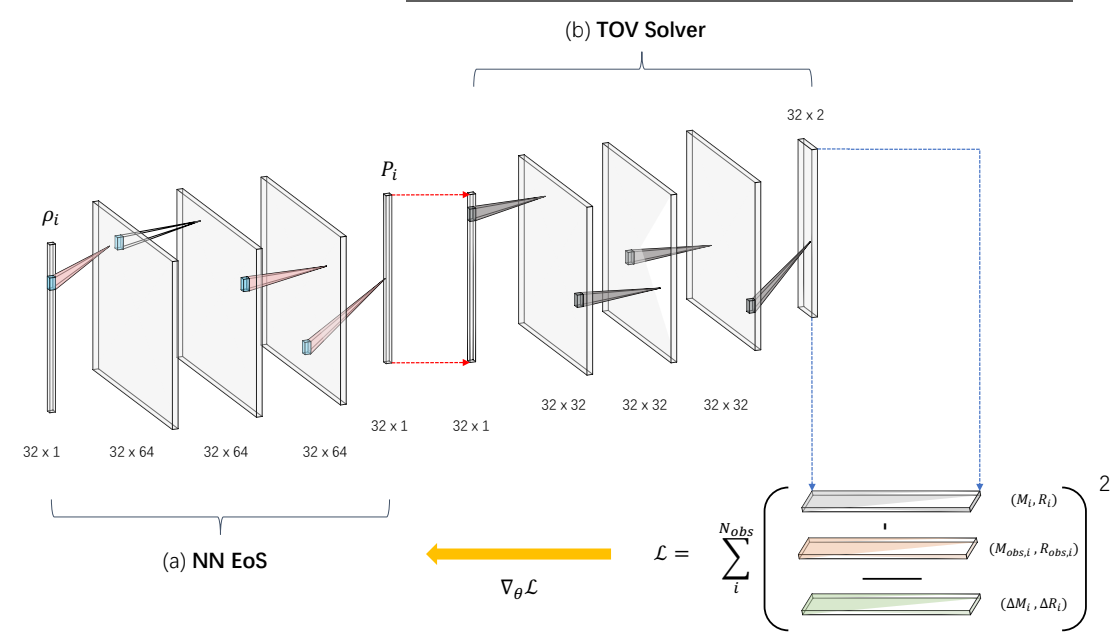
Heavy-Ion Collisions

Phys. Rev. C 106, L051901; Phys. Rev. D 103, 116023



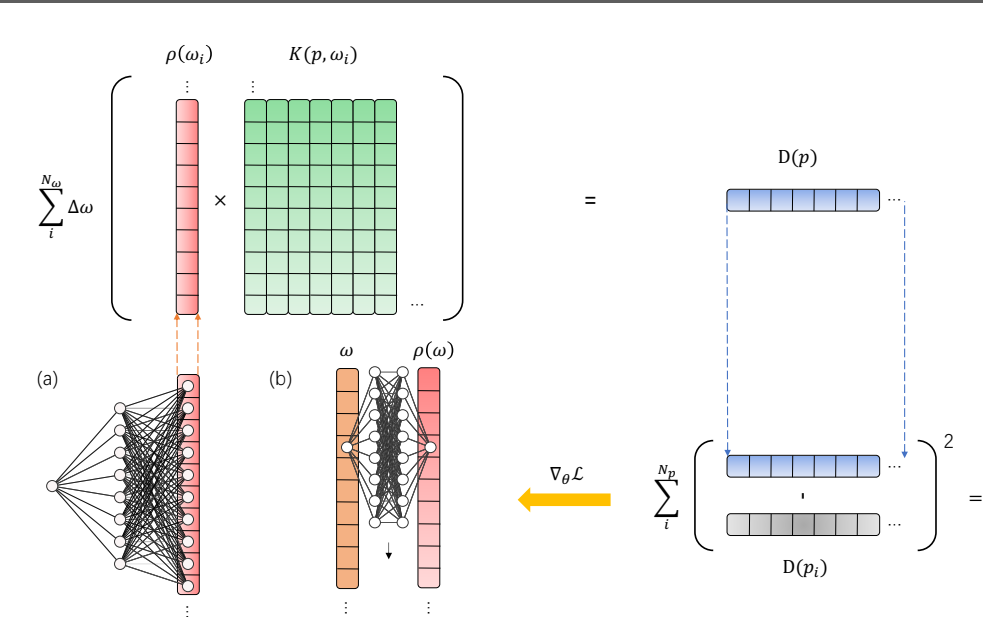
Neutron Star

Phys. Rev. D 107, 083028; JCAP08 (2022) 071



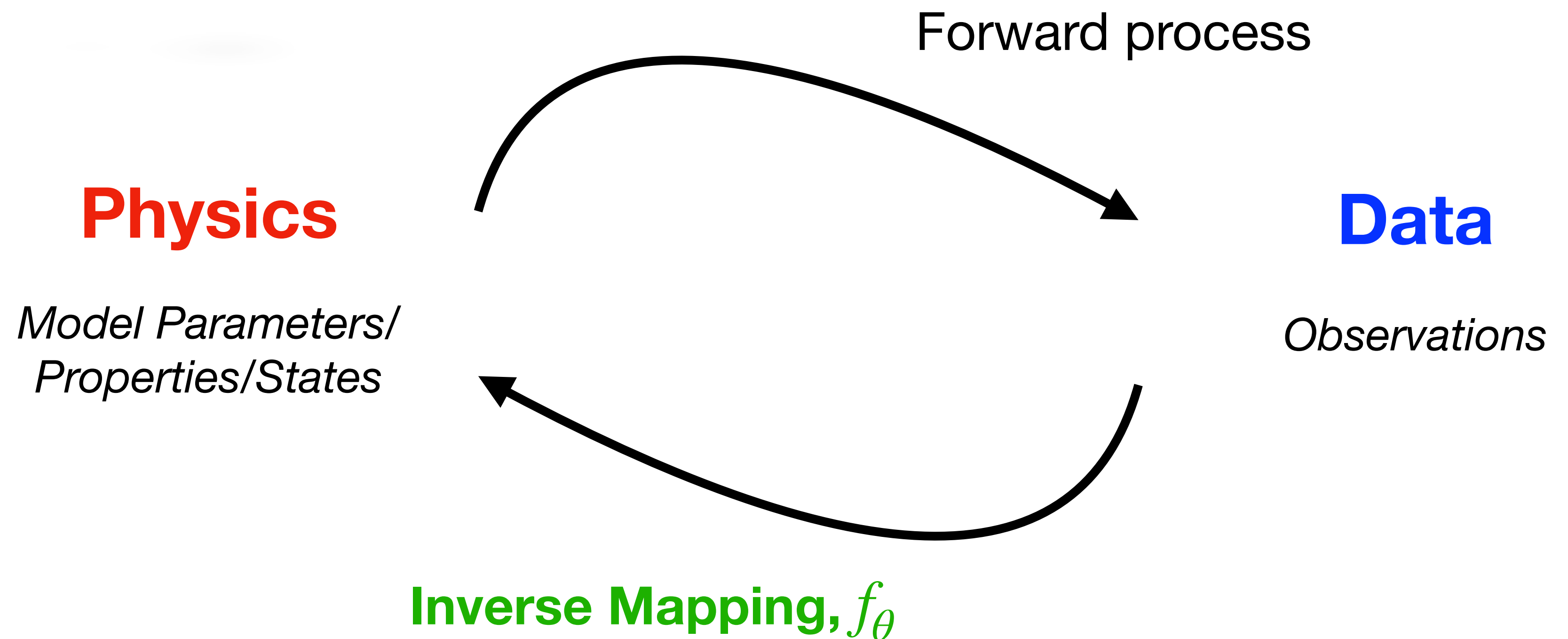
Lattice

Phys. Rev. D 106, L051502; Comput. Phys. Commun. 282, 108547 (2023);



Data-Driven Learning

$$f_{\theta} : X \rightarrow Y$$

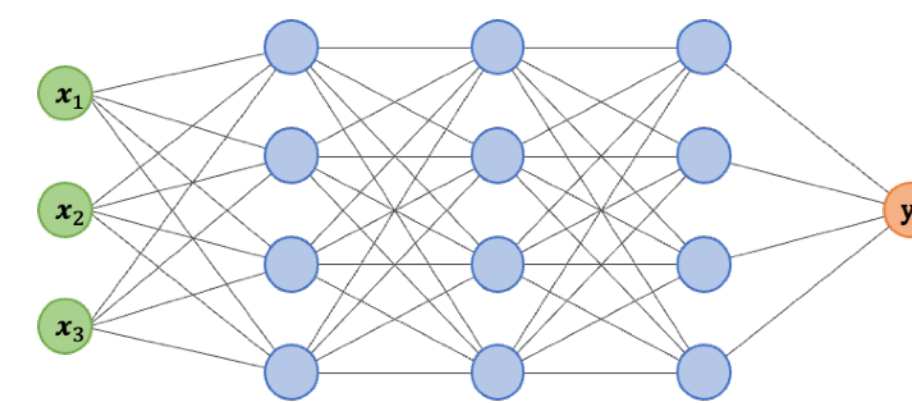


Data-Driven Learning

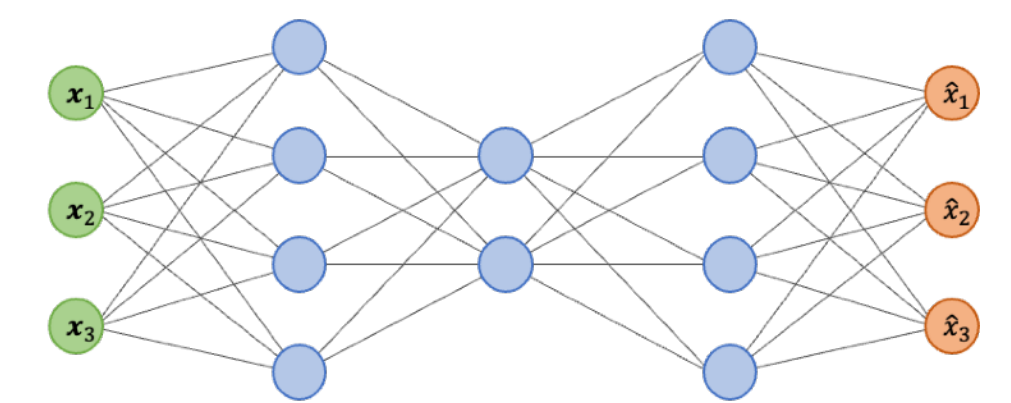
$$f_{\theta} : X \rightarrow Y$$

Universal Approximation Theorem (1989,1991)

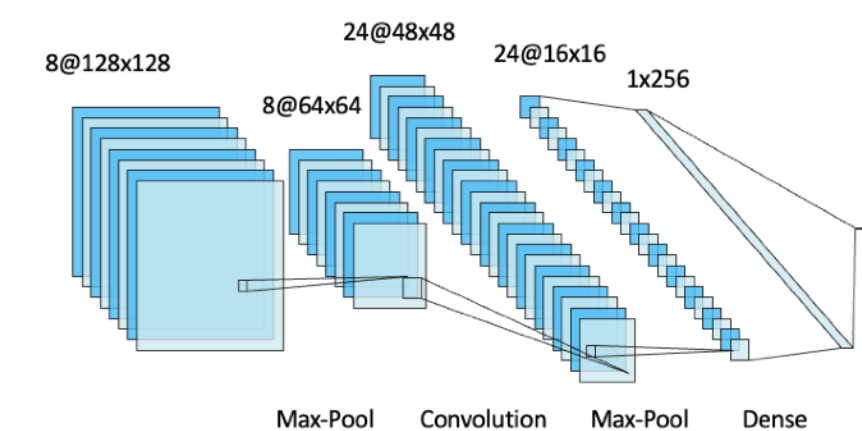
A feed-forward network with a single hidden layer containing a finite number of neurons can approximate arbitrary continuous functions.



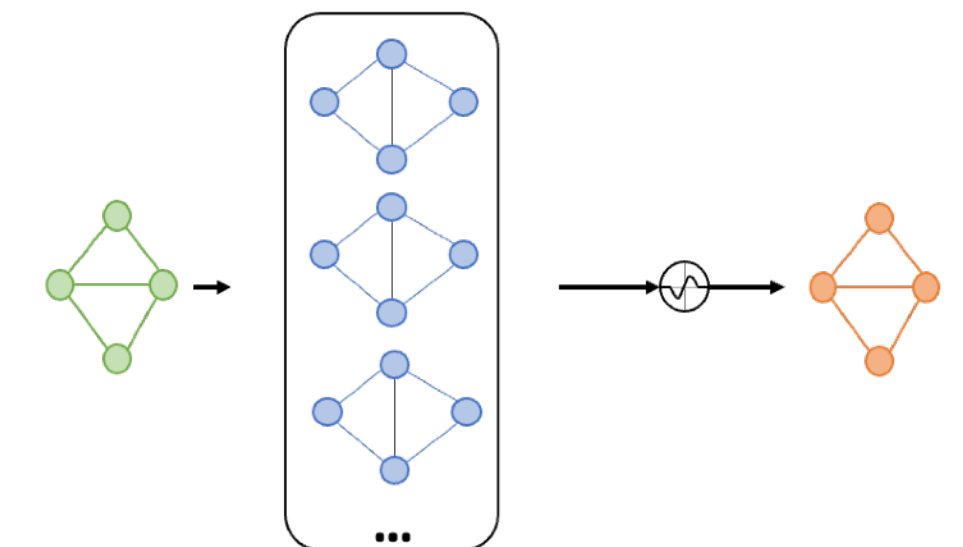
Deep Neural Network



AutoEncoder



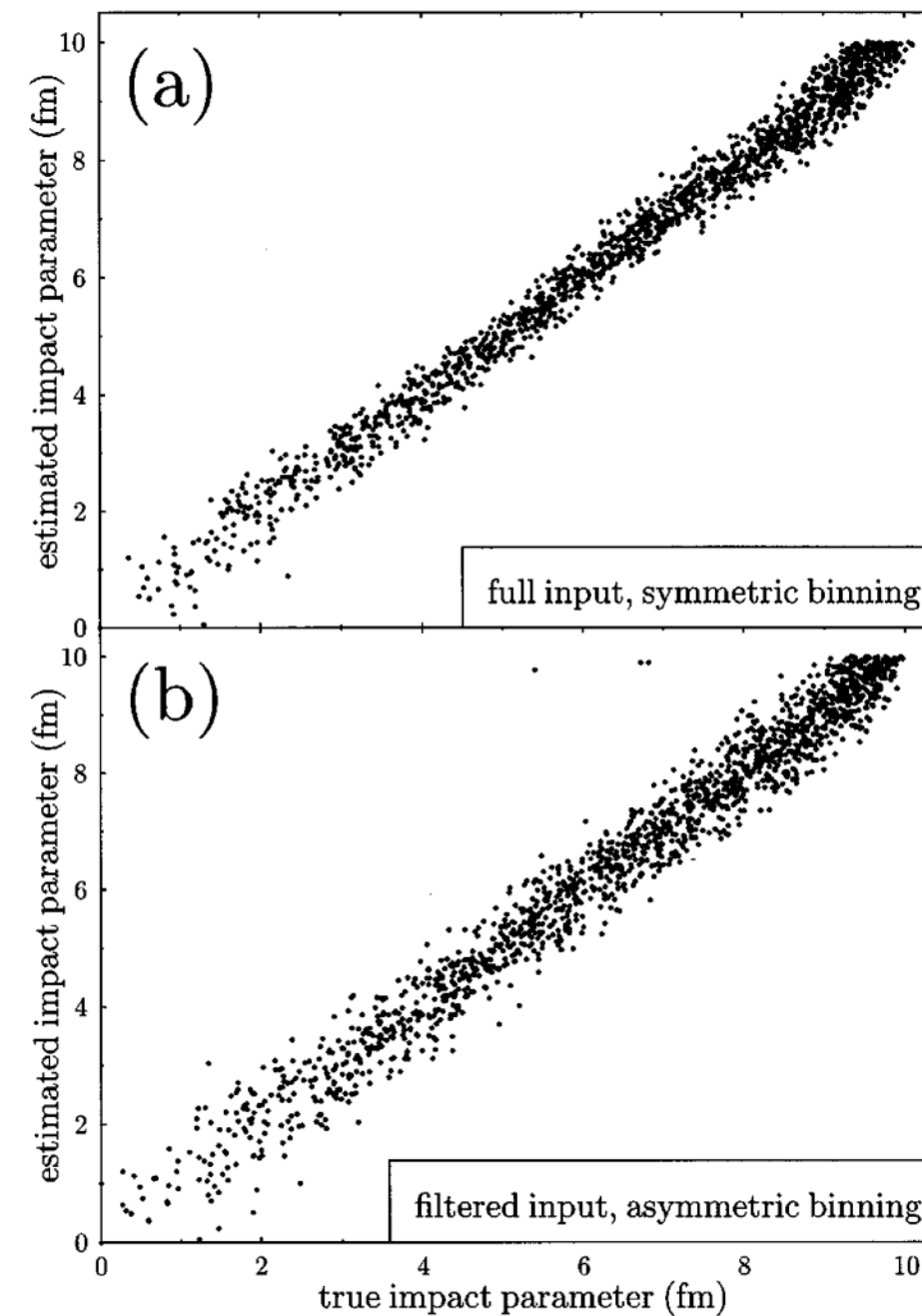
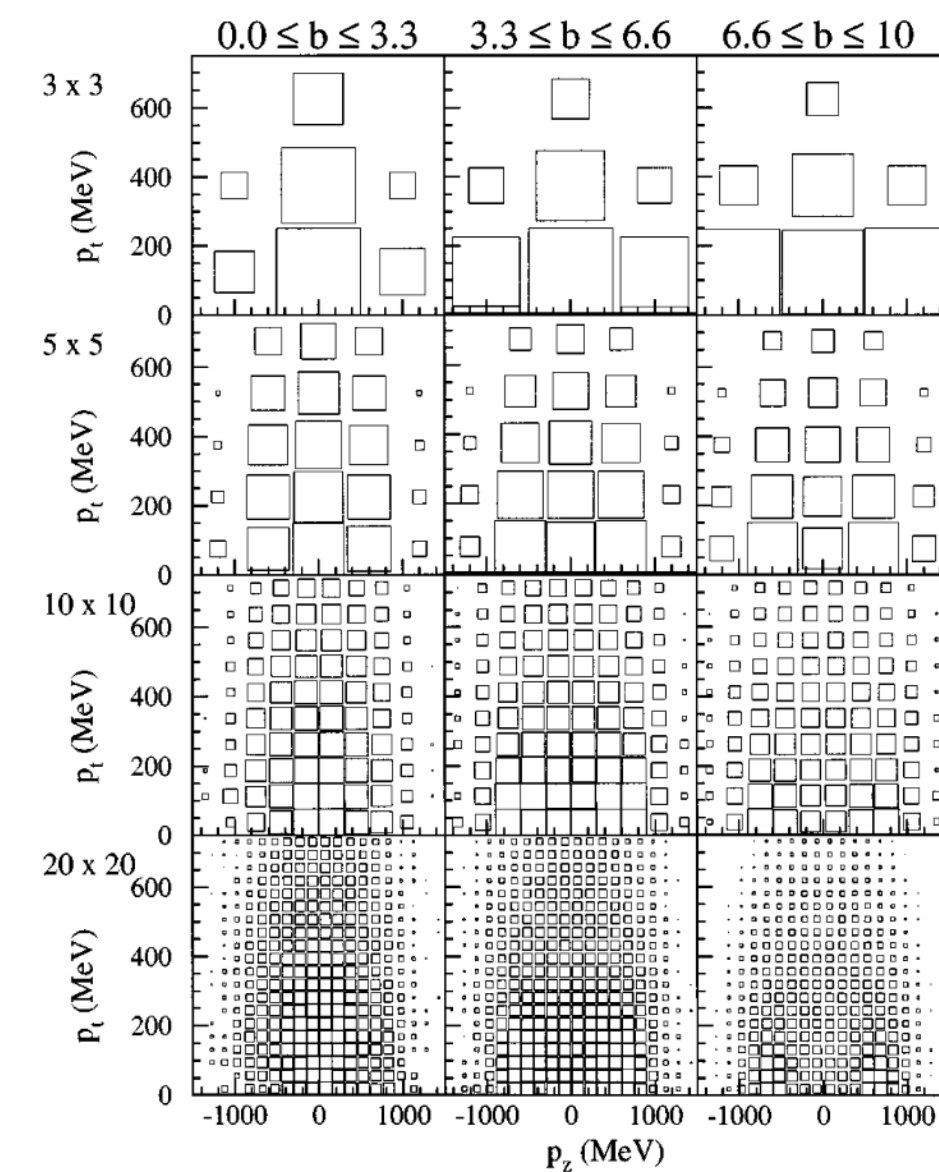
Convolutional Neural Network



Graph Neural Network

Data-Driven Learning

Determine Impact Parameter



S. A. Bass, A. Bischoff, J. A. Maruhn, H. Stöcker, and W. Greiner, Phys. Rev. C 53, 2358 (1996)

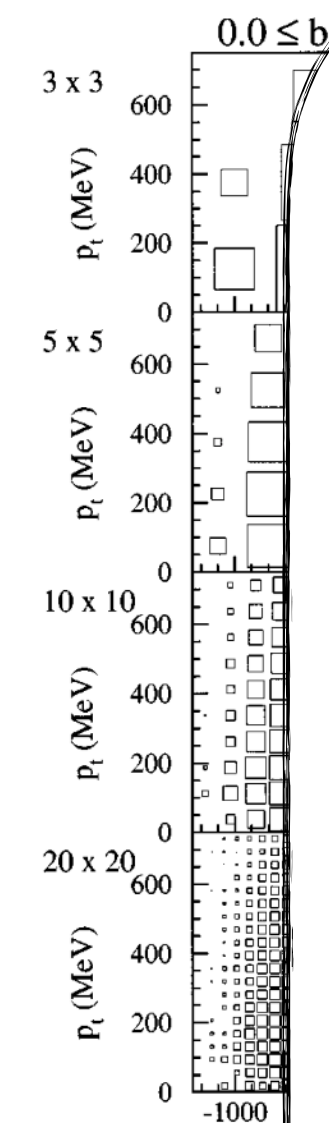
QMC data

1 hidden layer
with 20 neurons

Input 5X5

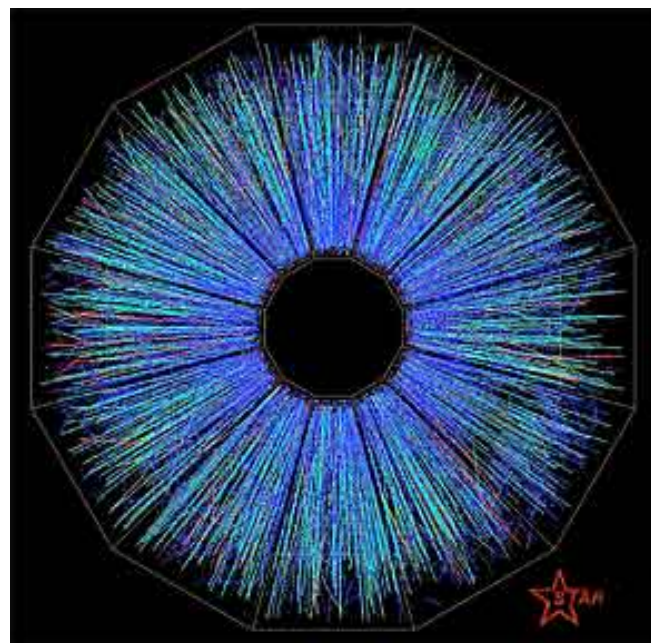
Data-Driven Learning

Determine Impact Parameter

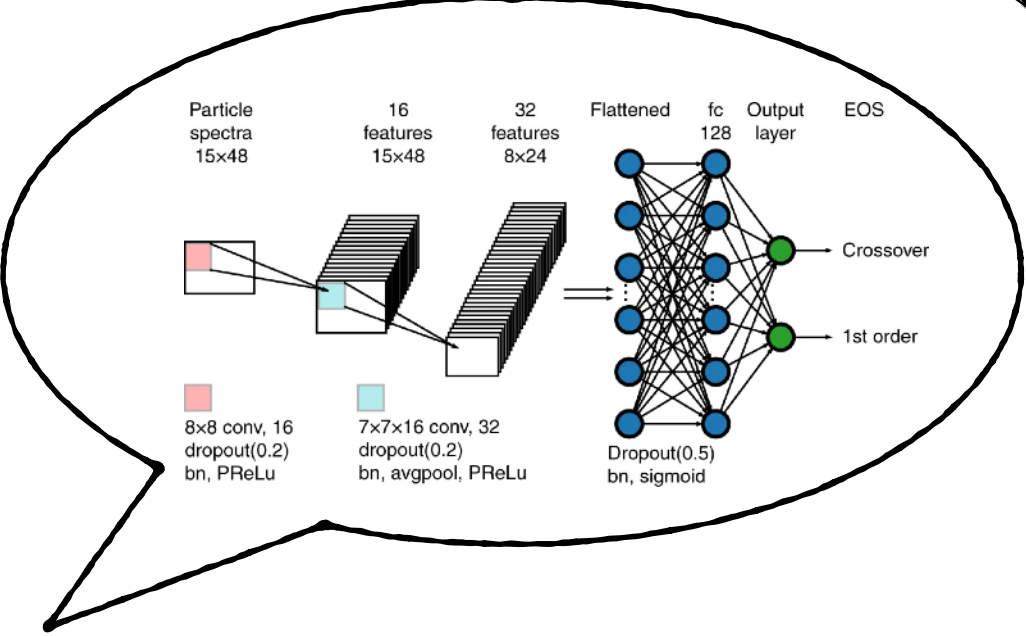


Recognizing QCD
Phase Transitions

$$\rho(p_T, \Phi)$$



Cross-Over
or
1st order PT



Hydro data

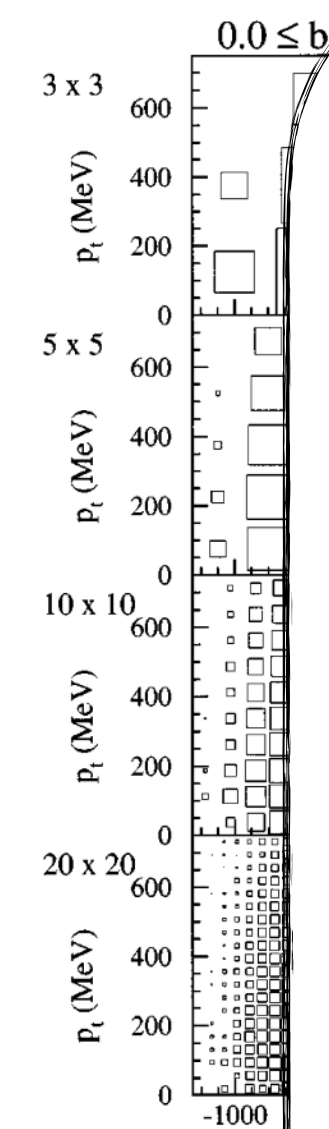
CNNs+DNNs

Input 15X48

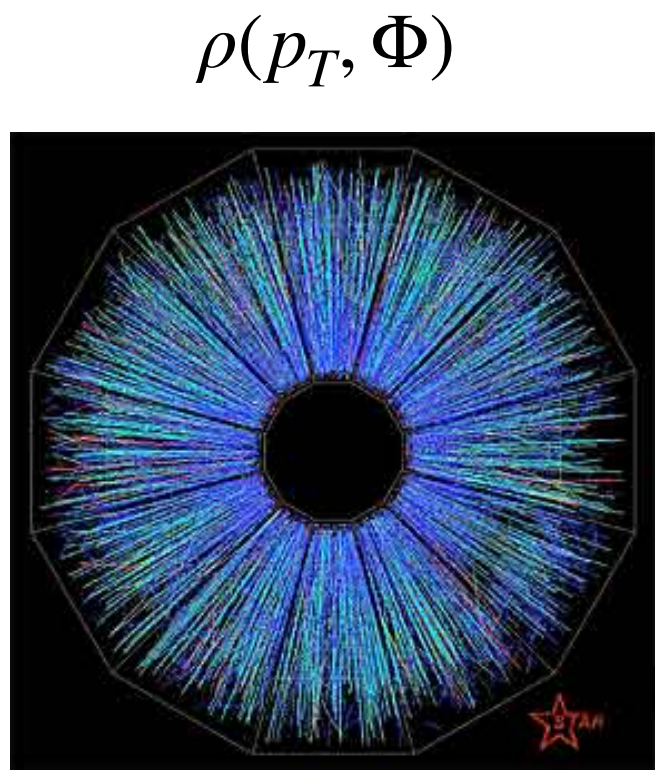
L.-G. Pang, K. Zhou, N. Su, H. Petersen, H. Stöcker, and X.-N. Wang, Nature Commun. 9, 210 (2018)

Data-Driven Learning

Determine Impact Parameter



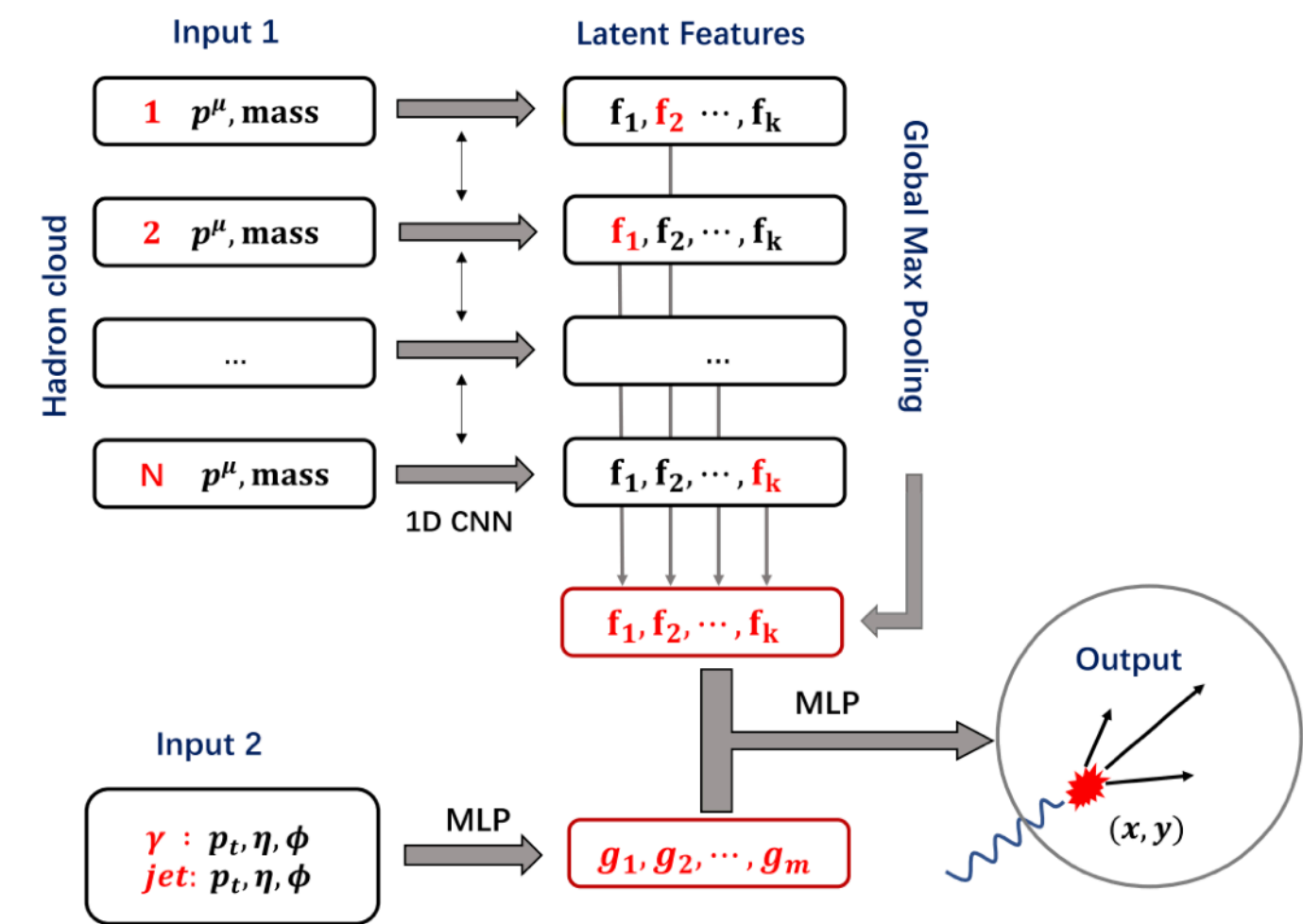
Recognizing QCD Phase Transitions



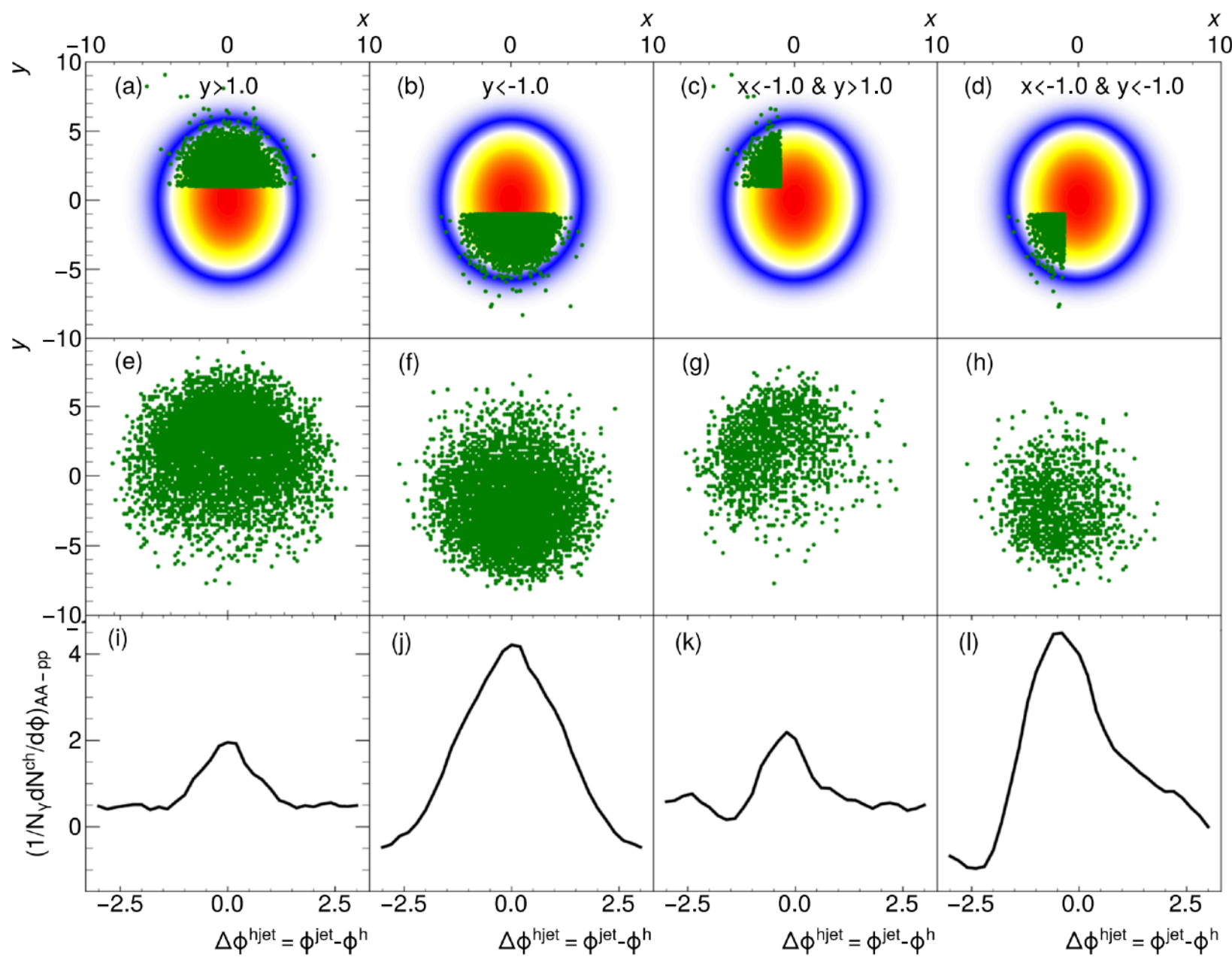
S. A.

L.-G. Pang, K. Zhou, N. S.

Point Net for Jet Tomography

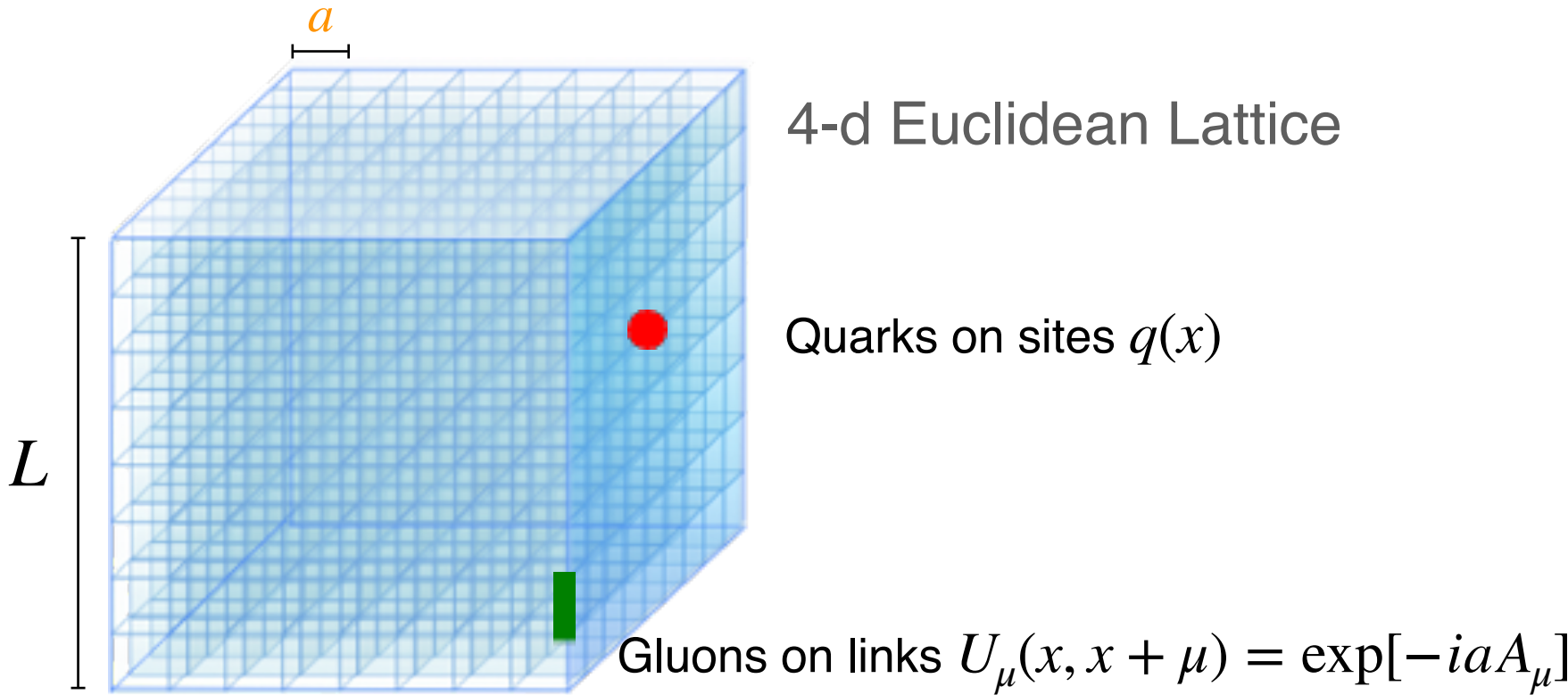


Z. Yang, Y. He, W. Chen, W.-Y. Ke, L.-G. Pang, and X.-N. Wang, Eur. Phys. J. C 83, 652 (2023).



Hadron Interactions

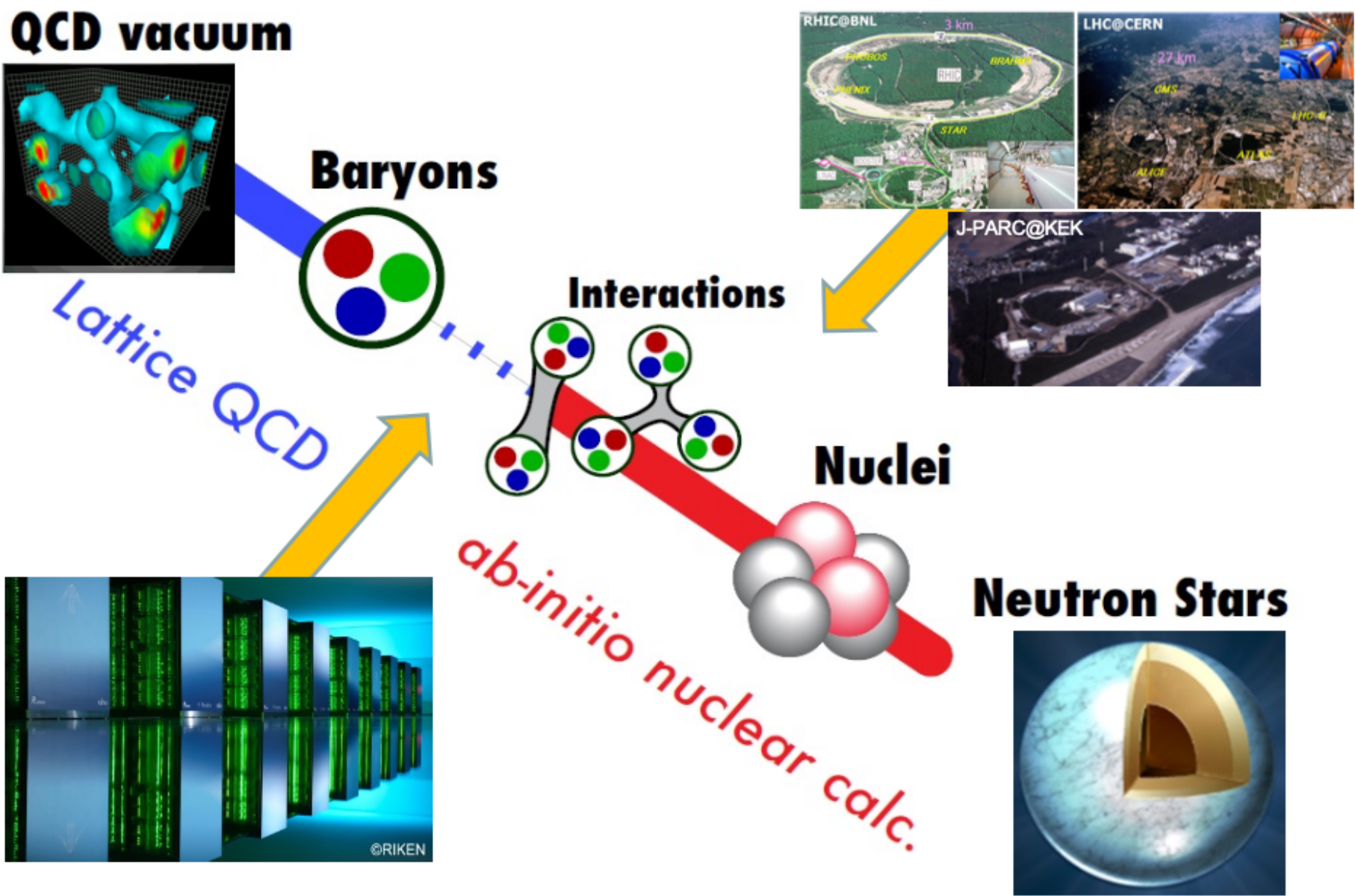
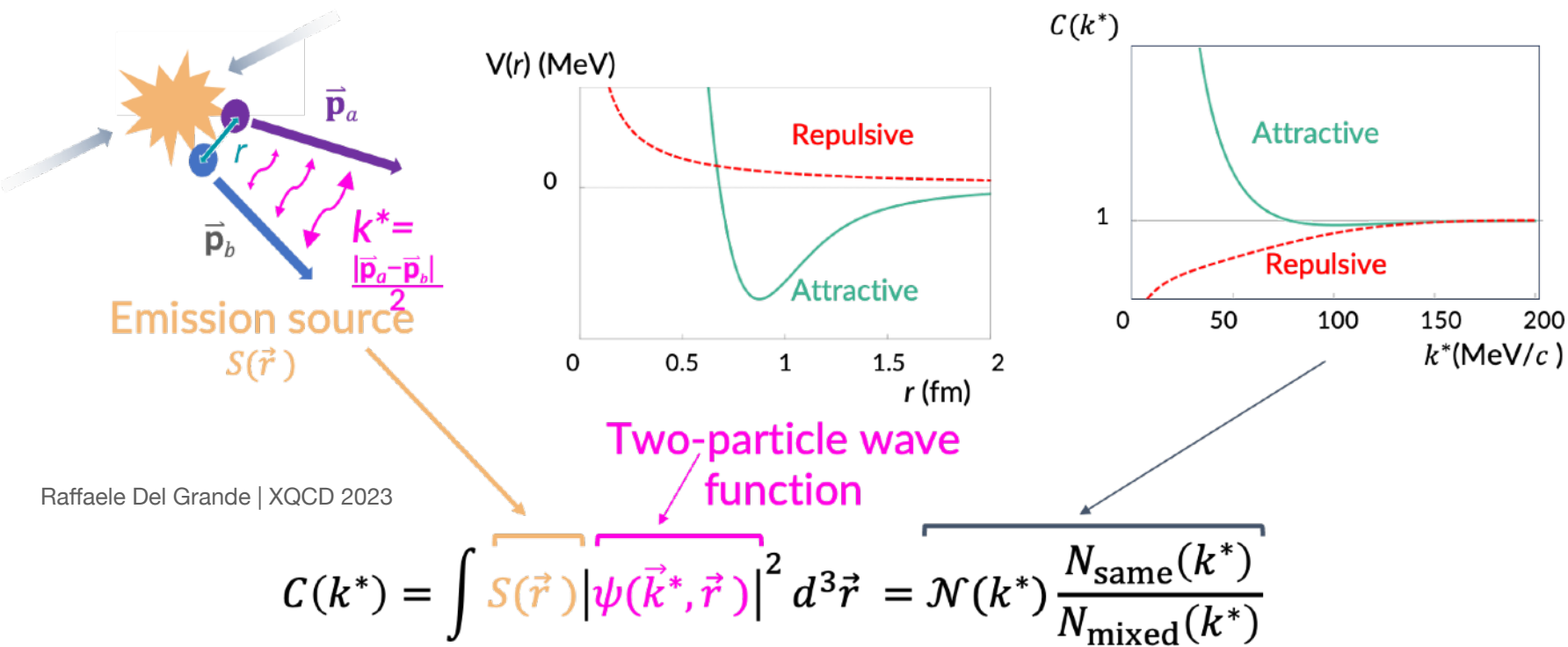
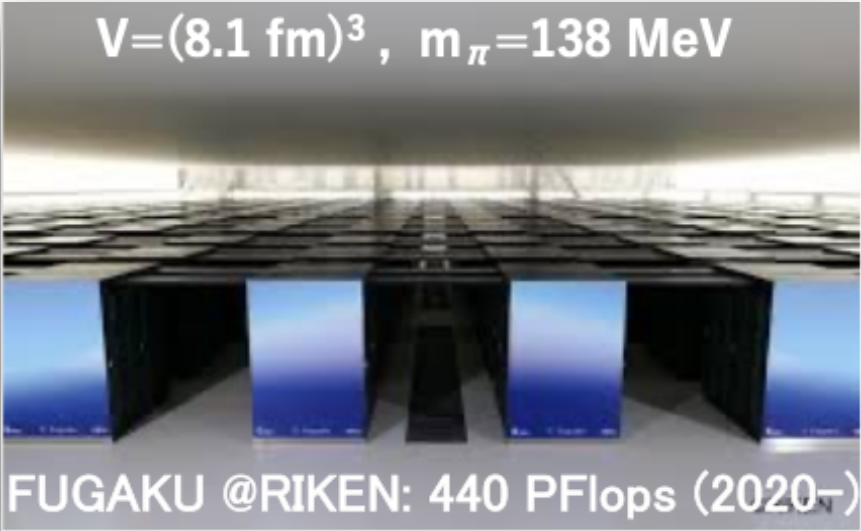
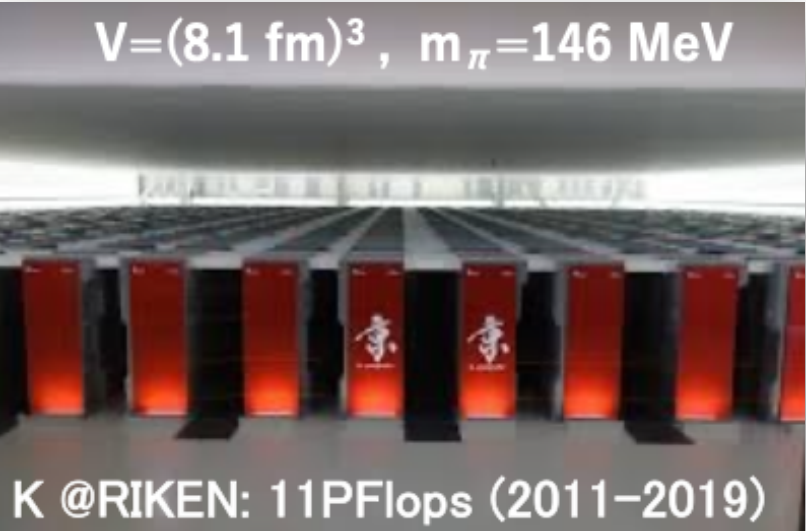
$$\mathcal{L} = -\frac{1}{4}G_{\mu\nu}^a G_a^{\mu\nu} + \bar{q}\gamma^\mu(i\partial_\mu - g t^a A_\mu^a)q - m\bar{q}q$$



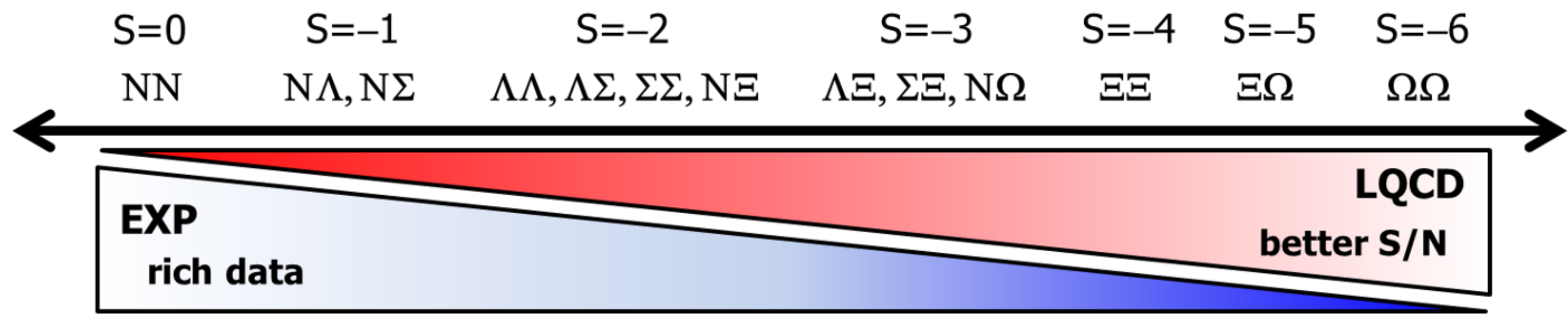
Huge integration variables
 $\sim 10^{9-10}$ for 96^4 lattice, ~ 50 GB/config

Importance Sampling
Hybrid MC = MD + Metropolis

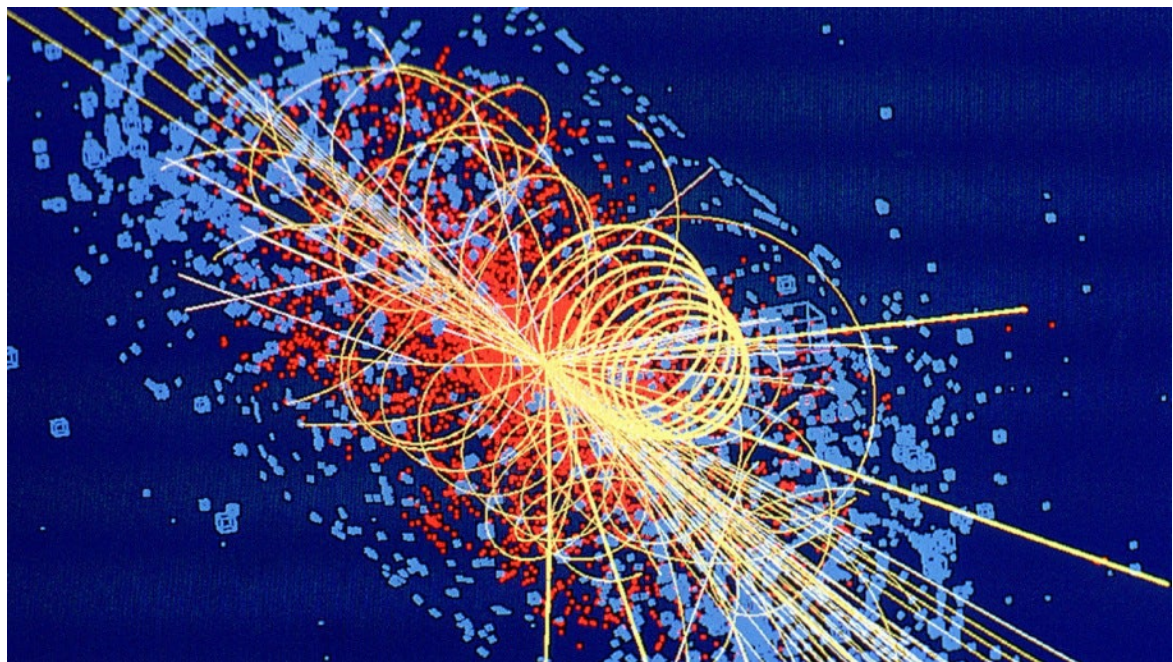
Continuum & Thermodynamic Limits
 $a \rightarrow 0, L \rightarrow \infty$



Hadron Interactions



LHC@CERN, RHIC@BNL
J-PARC@KEK, FAIR@GSI, HIAF



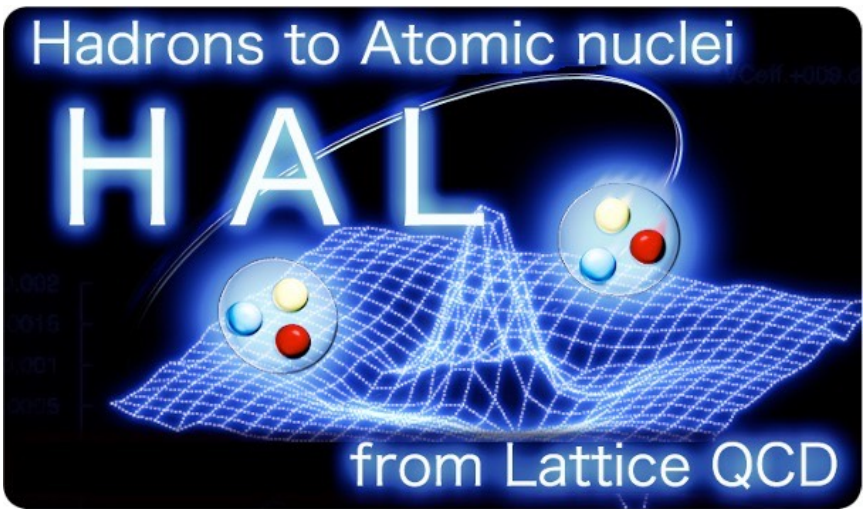
Femtoscscopy
HAL QCD method

+

Deep Learning

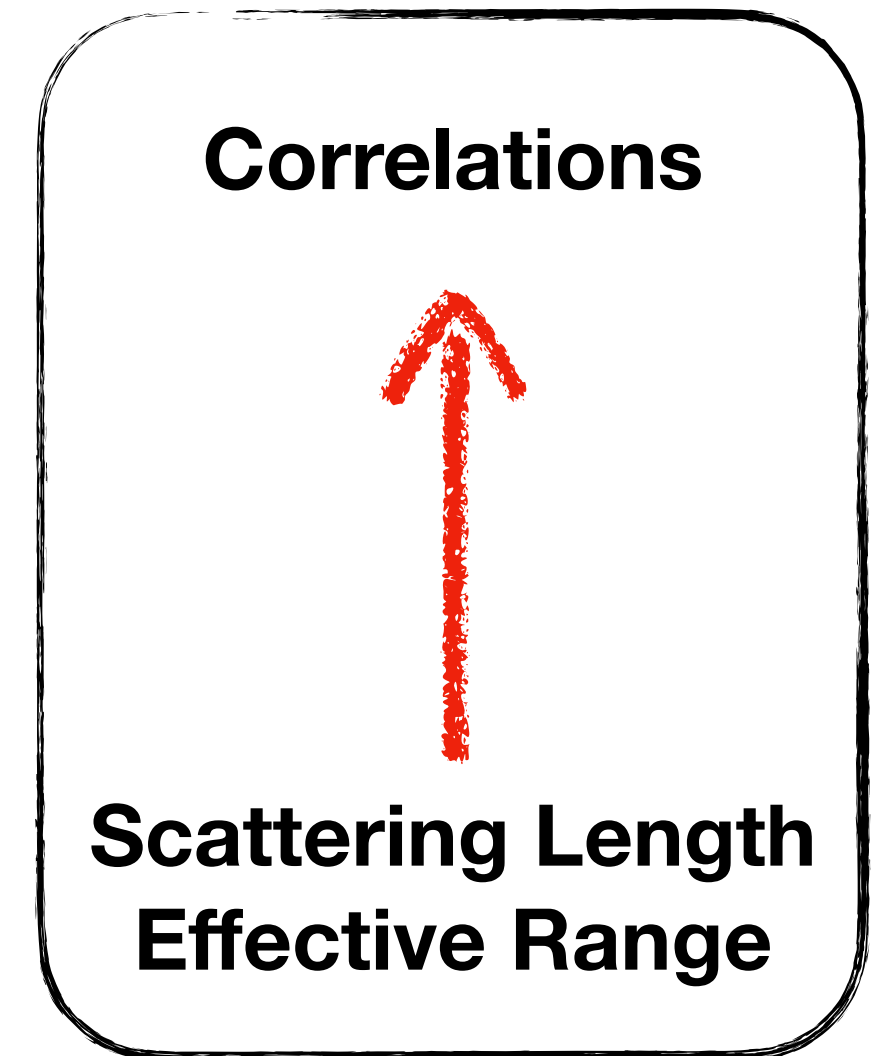
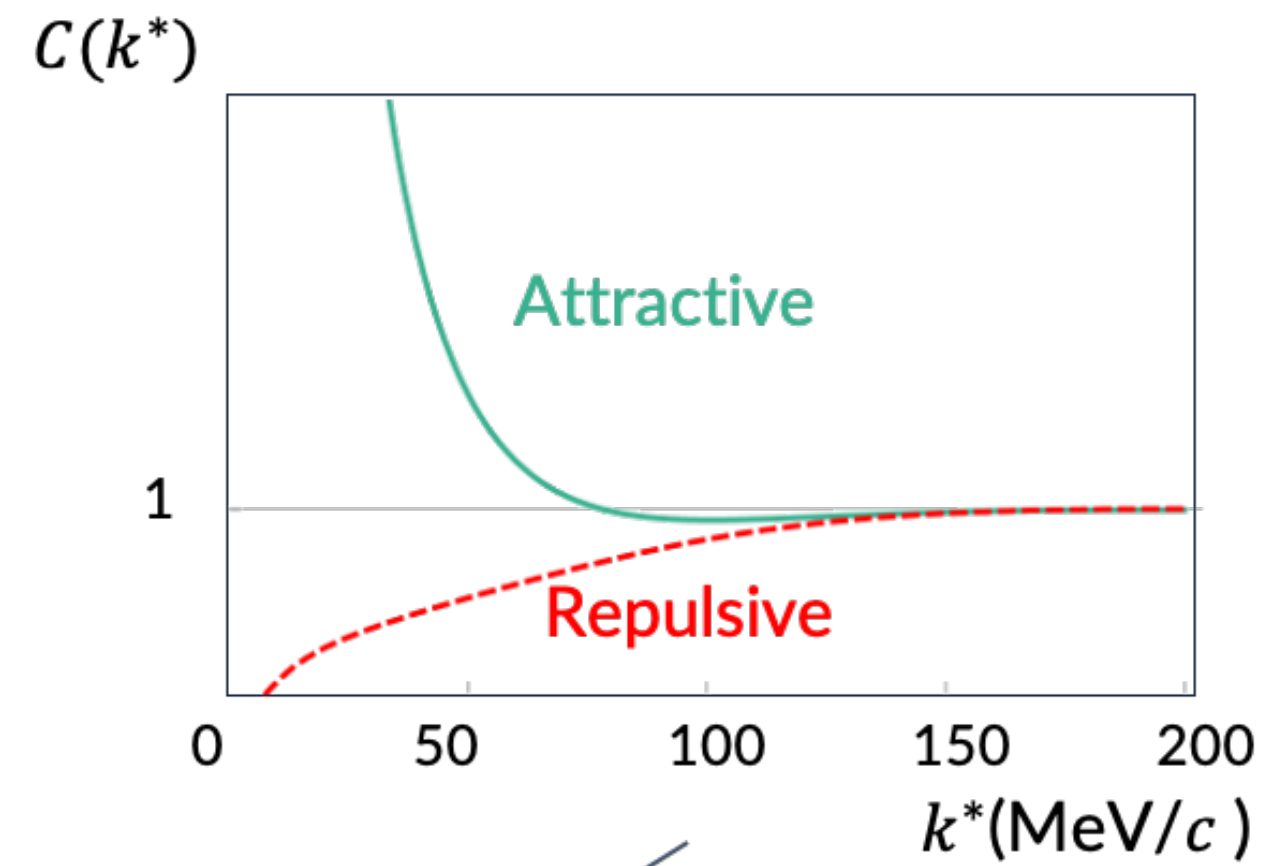
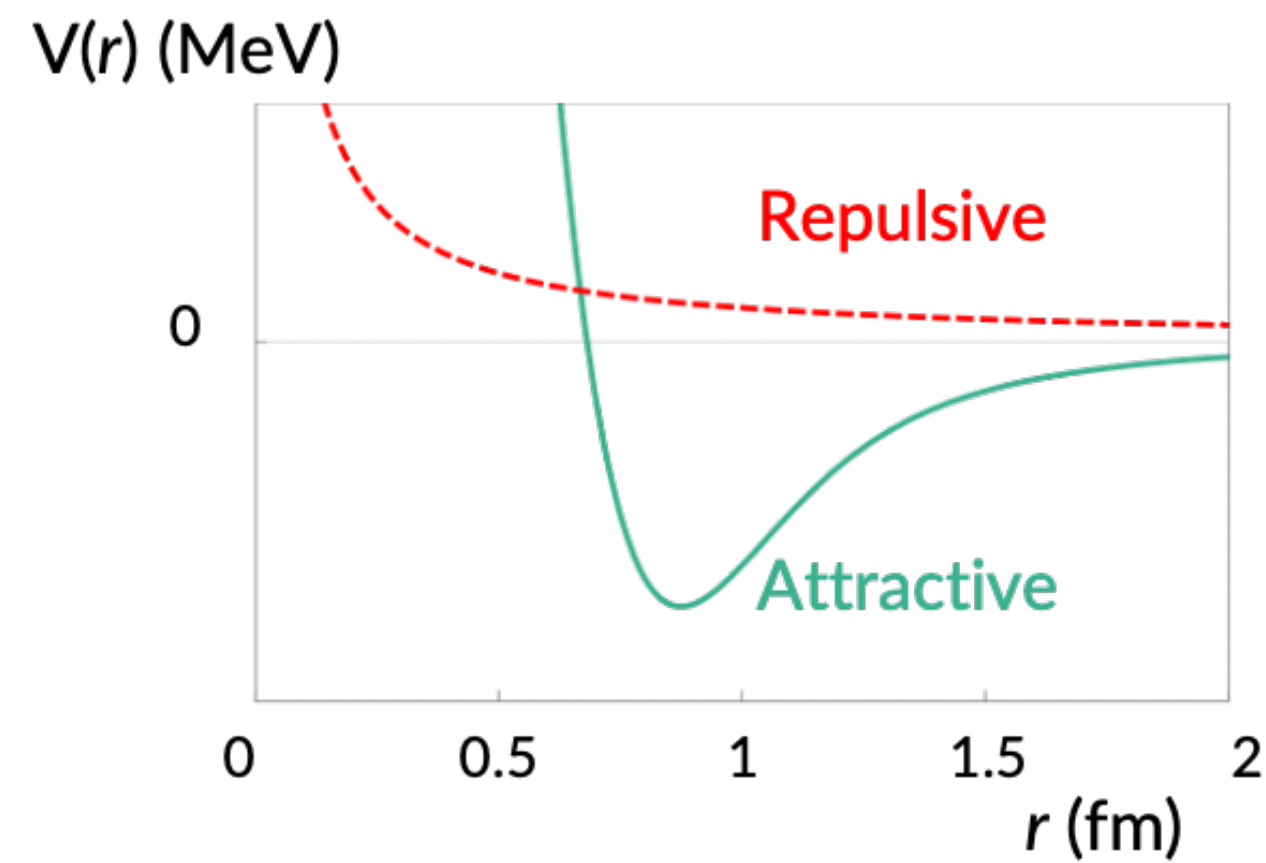
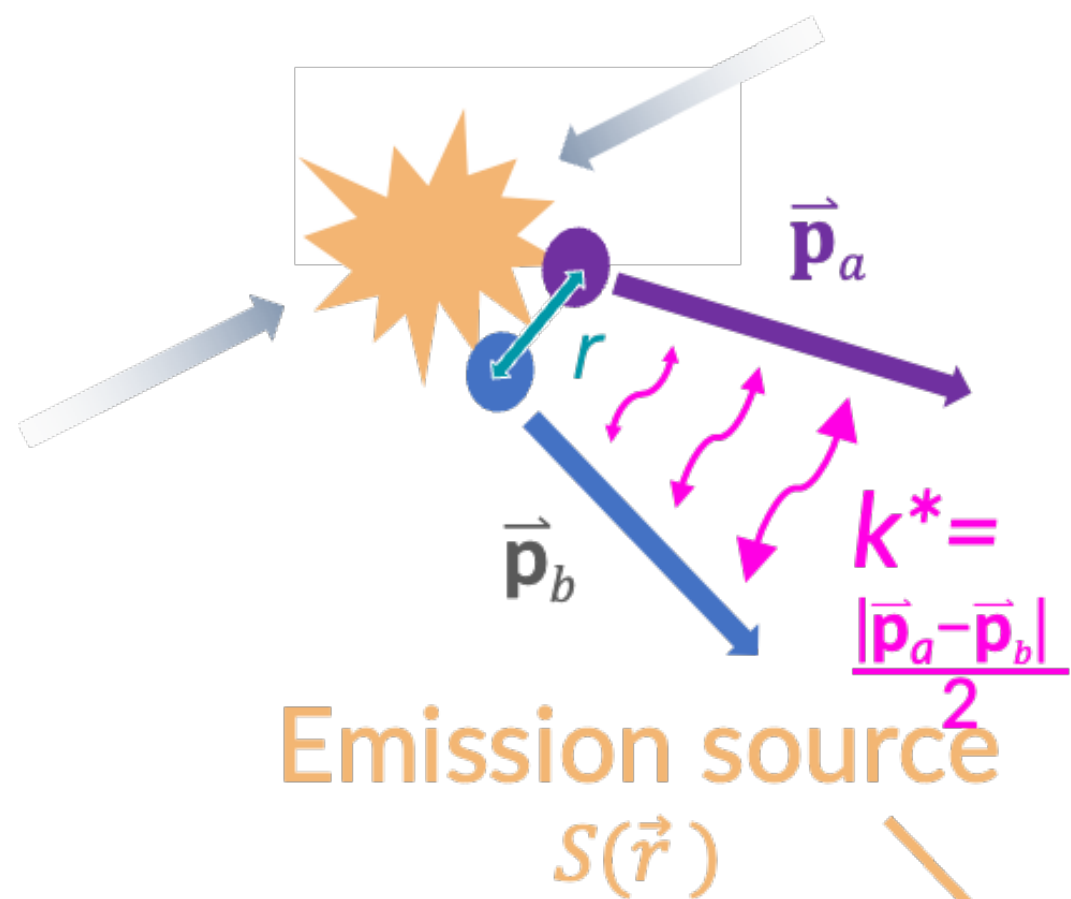


Hadrons to Atomic nuclei from Lattice QCD
(HAL QCD Collaboration)



S. Aoki, T. M. Doi, E. Itou (Kyoto U.) T. Aoyama (ISSP)
T. Doi, T. Hatsuda, Y. Lyu, L. Wang, R. Yamada, L. Zhang (RIKEN)
Y. Ikeda, N. Ishii, P. Junnarkar, H. Nemura, K. Sasaki (Osaka U.)
T. Inoue (Nihon U.) K. Murakami (TITech)
K. Murase (Tokyo Metropolitan U.) F. Etminan (U. of Birjand)
T. Sugiura (Rissho U.) H. Tong (U. of Bonn)

Femtoscopy



$$C(k^*) = \int \underbrace{S(\vec{r})}_{\text{Emission source}} \underbrace{|\psi(\vec{k}^*, \vec{r})|^2}_{\text{Two-particle wave function}} d^3\vec{r} = \mathcal{N}(k^*) \frac{N_{\text{same}}(k^*)}{N_{\text{mixed}}(k^*)}$$

Lednicky-Lyuboshits(LL) analytic model
 (Asymptotic wave-function+
 Effective range correlation+
 Gaussian source)

Raffaele Del Grande | XQCD 2023

Lednicky, Lyuboshits, Sov.J.Nucl.Phys. 35 (1982) 770

Femtoscscopy

Asymptotic wave-function

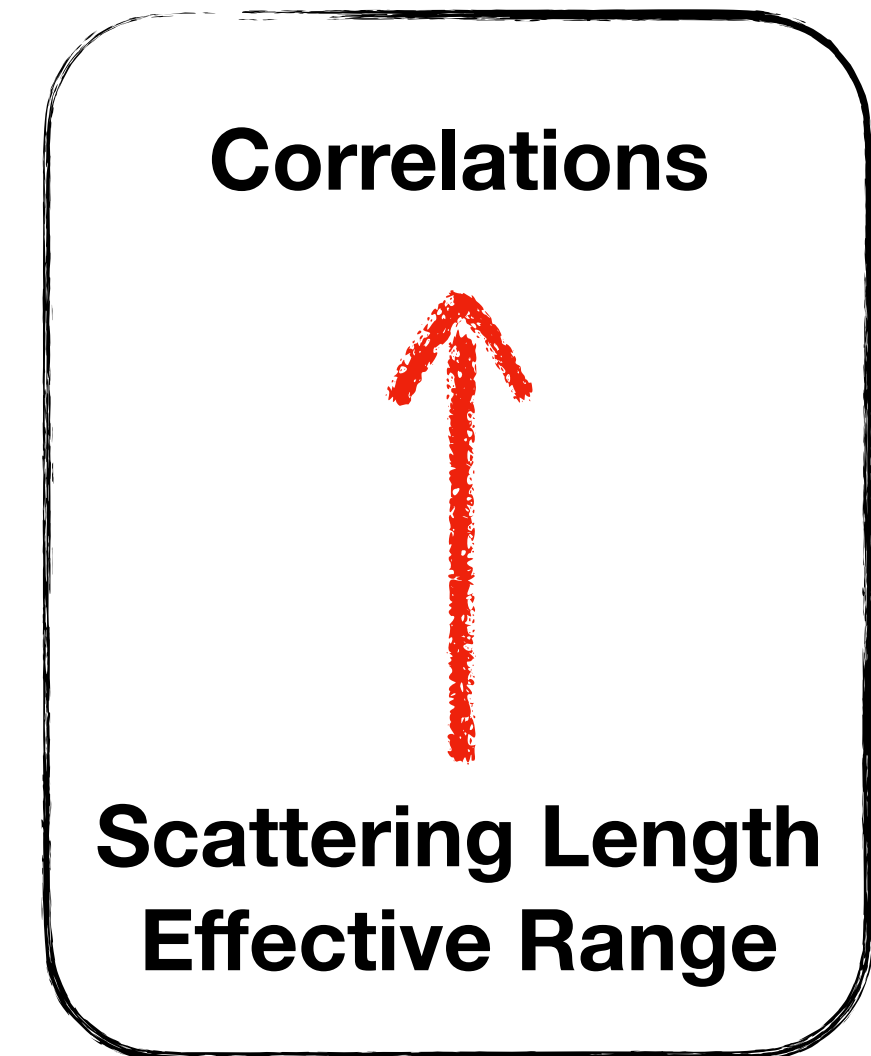
$$\psi_0(r) \rightarrow \psi_{\text{asy}}(r) = \frac{e^{-i\delta}}{qr} \sin(qr + \delta) = \mathcal{S}^{-1} \left[\frac{\sin qr}{qr} + f(q) \frac{e^{iqr}}{r} \right]$$

$$\begin{aligned} C_{\text{LL}}(q) &= 1 + \int dr S_{12}(r) \left(|\psi_{\text{asy}}(r)|^2 - |j_0(qr)|^2 \right) \\ &= 1 + \frac{|f(q)|^2}{2R^2} F_3 \left(\frac{r_{\text{eff}}}{R} \right) + \frac{2\text{Re}f(q)}{\sqrt{\pi}R} F_1(2x) - \frac{\text{Im}f(q)}{R} F_2(2x) \end{aligned}$$

$x = qR$, R is Gaussian Size, F_1, F_2, F_3 are known functions

Scattering amplitude at low energies

$$q \cot \delta = -\frac{1}{a_0} + \frac{1}{2} r_{\text{eff}} q^2 + O(q^4) \rightarrow f(q) = (q \cot \delta - iq)^{-1}$$



Lednicky-Lyuboshits(LL) analytic model

(Asymptotic wave-function+
Effective range correlation+
Gaussian source)

Lednicky, Lyuboshits, Sov.J.Nucl.Phys. 35 (1982) 770

Inverse Femtoscopy

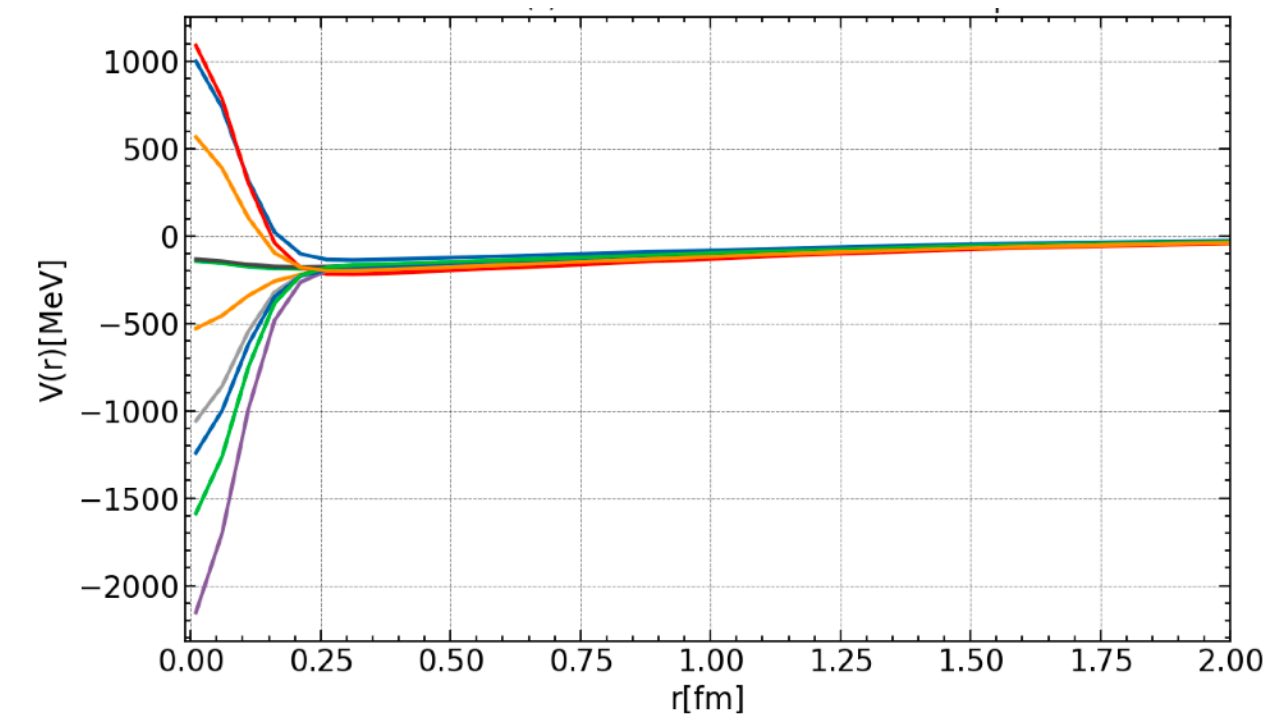
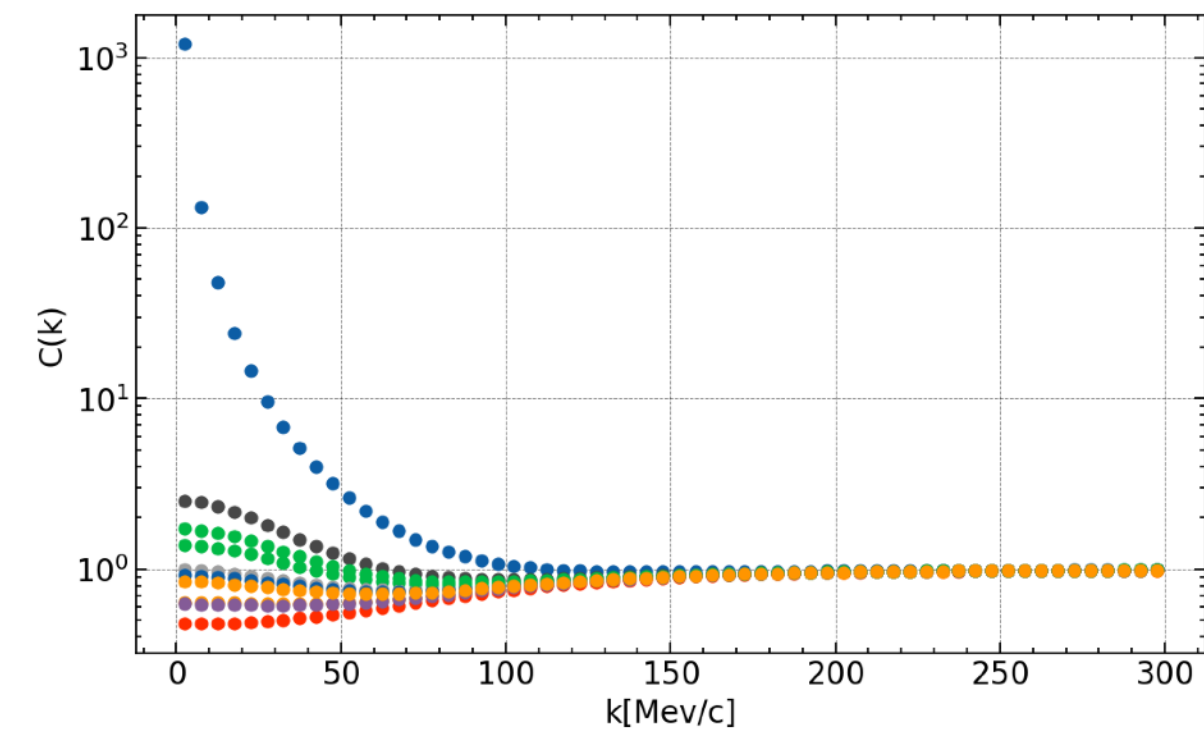
in Preparation

with Jiaxing Zhao, etc.

$$C(k^*) = \int S(\vec{r}) \left| \psi(\vec{k}^*, \vec{r}) \right|^2 d^3\vec{r} = \mathcal{N}(k^*) \frac{N_{\text{same}}(k^*)}{N_{\text{mixed}}(k^*)}$$

Does this inverse mapping exist?

$V(r)$



Correlations

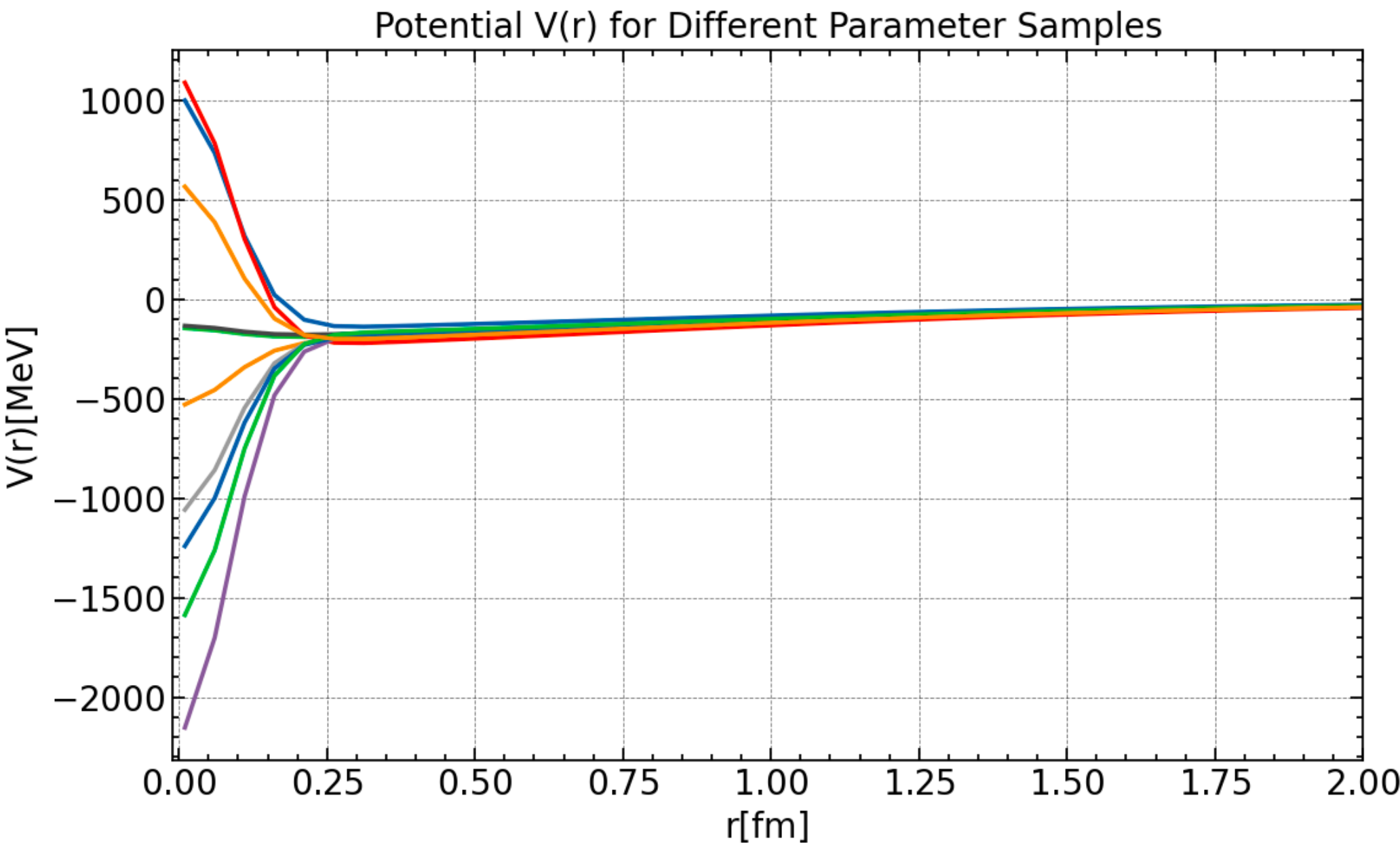


Potentials

Inverse Femtoscopy

Potential Functions

$$V(r) = b_1 e^{-b_2 r^2} + b_3 (1 - e^{-b_4 r^2}) \left(\frac{e^{(-m_\pi r)}}{r} \right)^{n_\pi}$$



Deep
Neural
Network
(DNN)

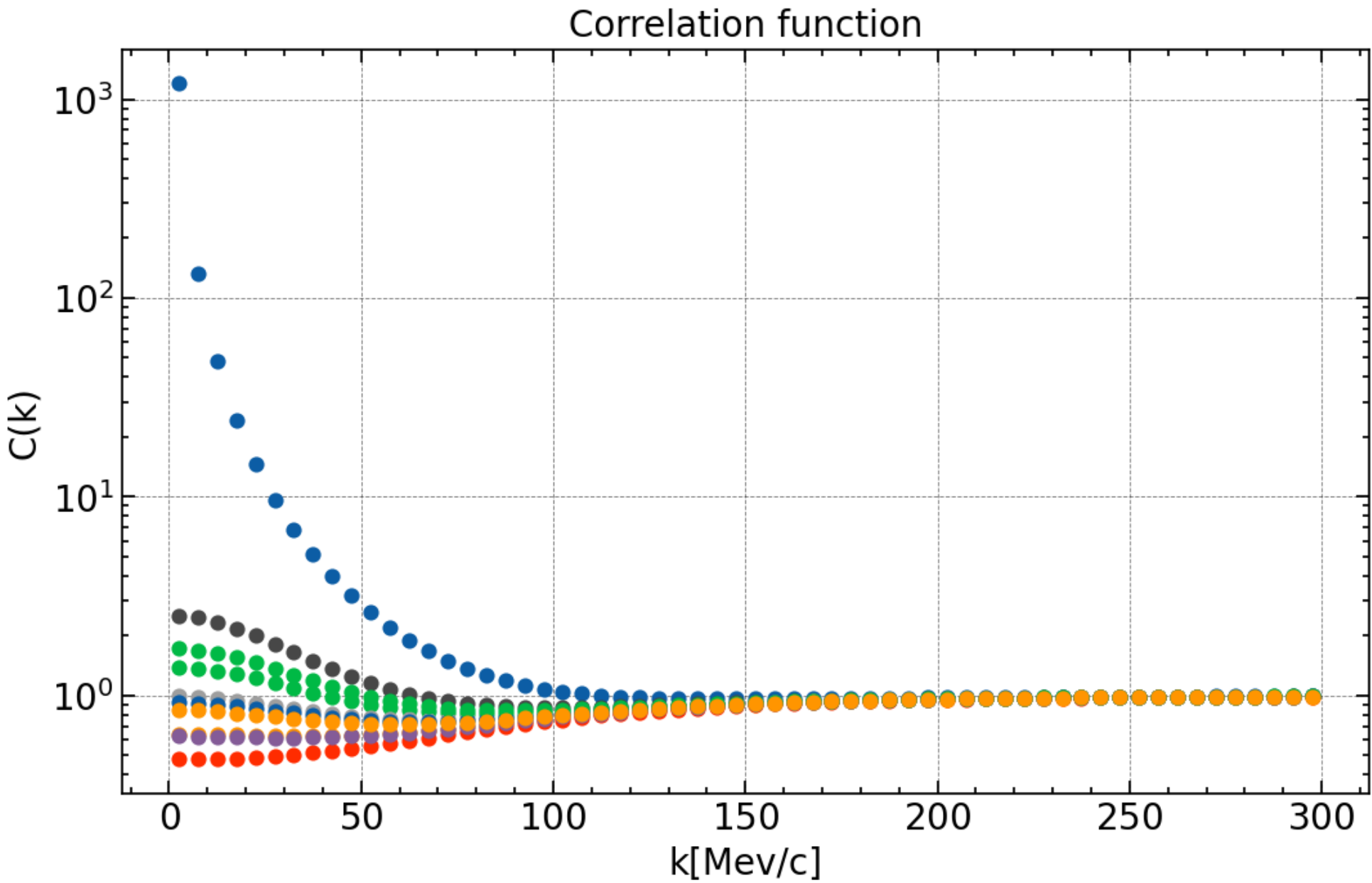
Schrödinger eq.

CATS Framework: D. Mihaylov et al.,
Eur. Phys. J. C78 (2018) 394



in Preparation

with Jiaxing Zhao, etc.



Source Function

$$S(r) = (4\pi r_0^2)^{-3/2} e^{-\frac{r^2}{4r_0^2}} \quad r_0 = 1.3 \text{ fm}$$

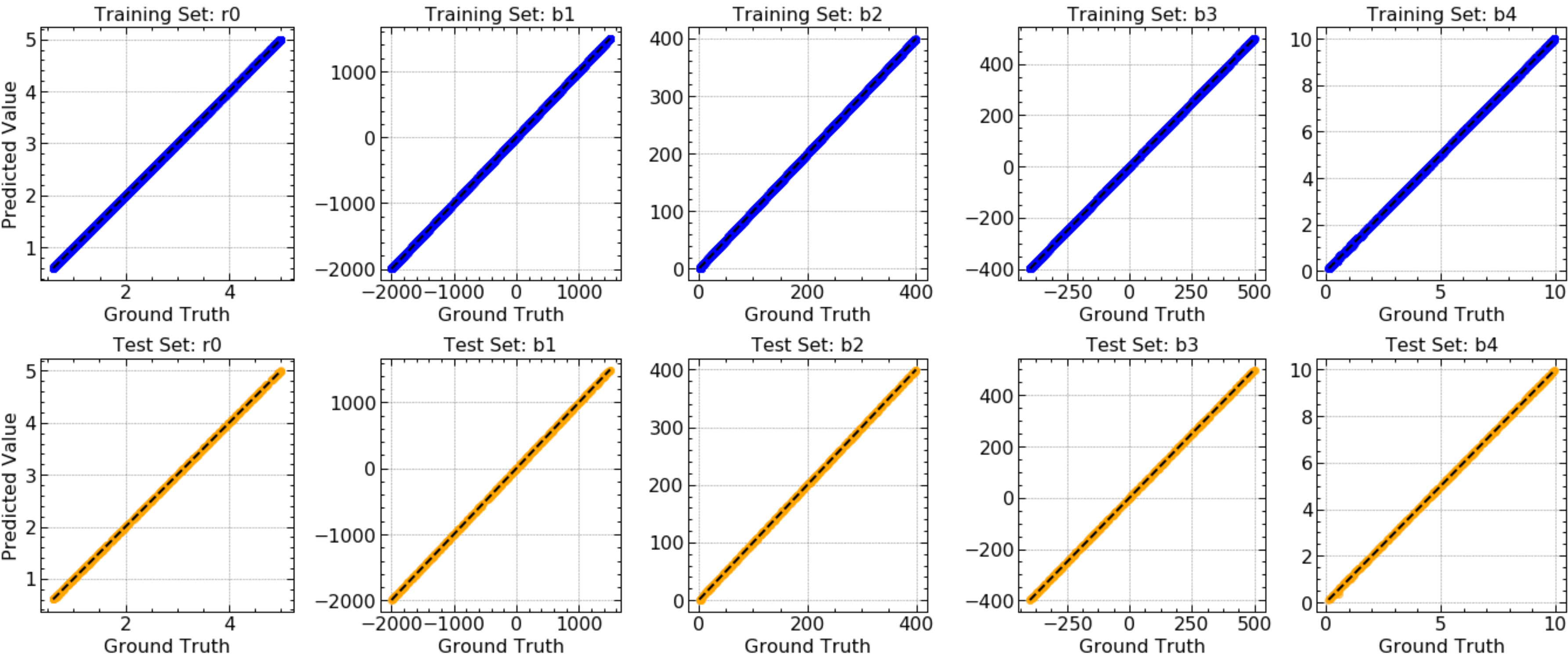
60 points(k), N_C correlations

Inverse Femtoscopy

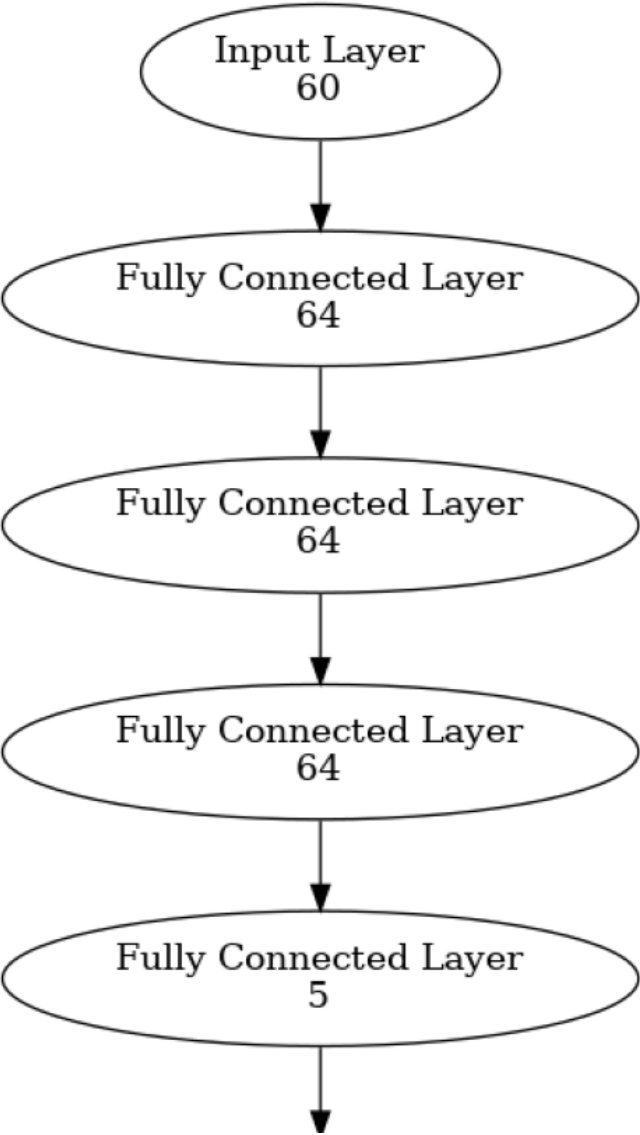
$N_c = 6400$ $r_0 = 1.3 \text{ fm}, b_1 = -306.5, b_2 = 200, b_3 = -266, b_4 = 0.78, n_\pi = 2$

in Preparation

with Jiaxing Zhao, etc.



60 points

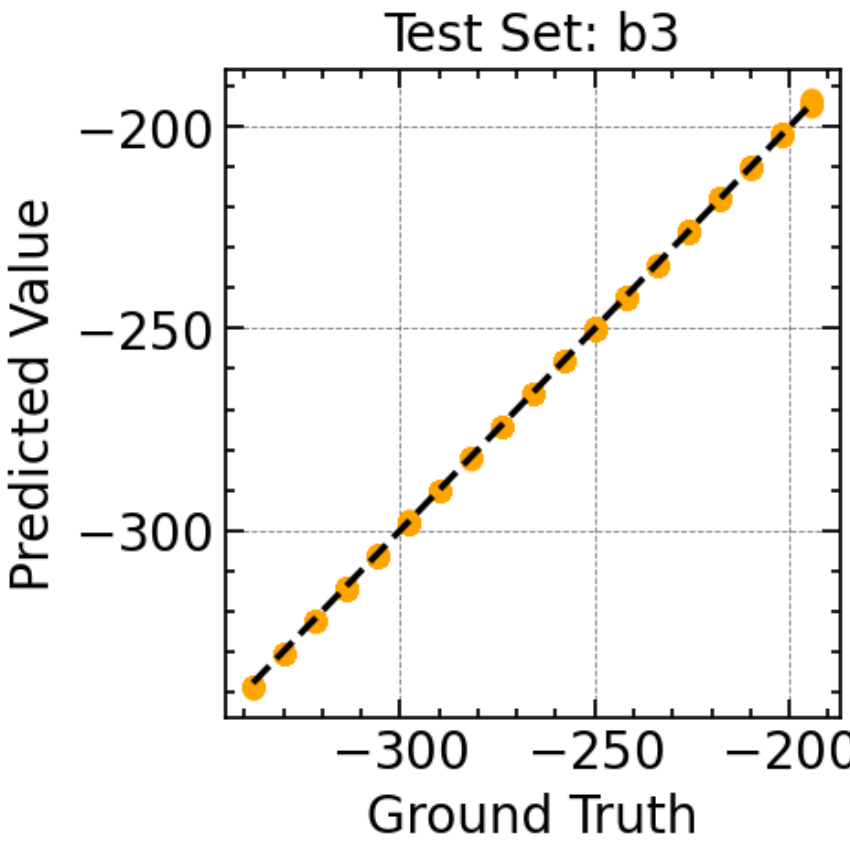
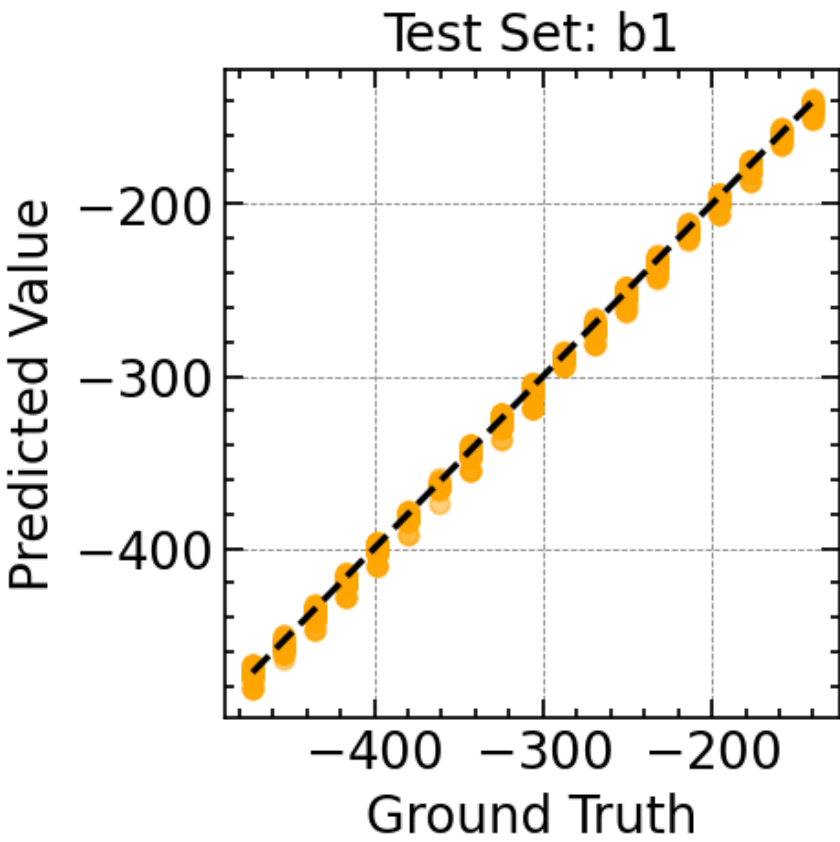
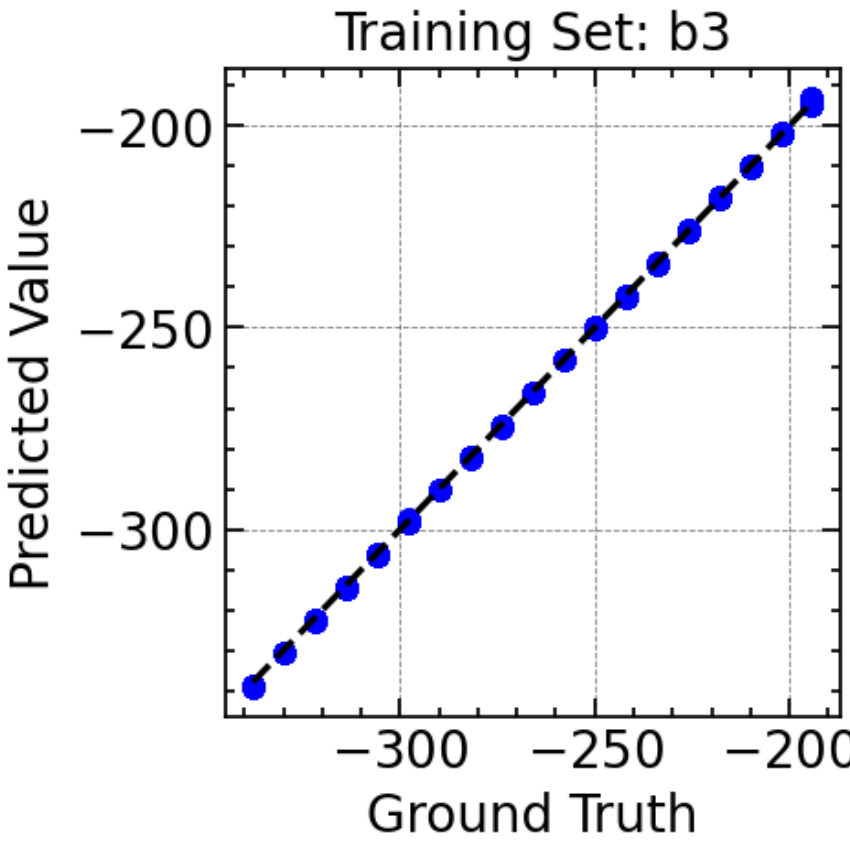
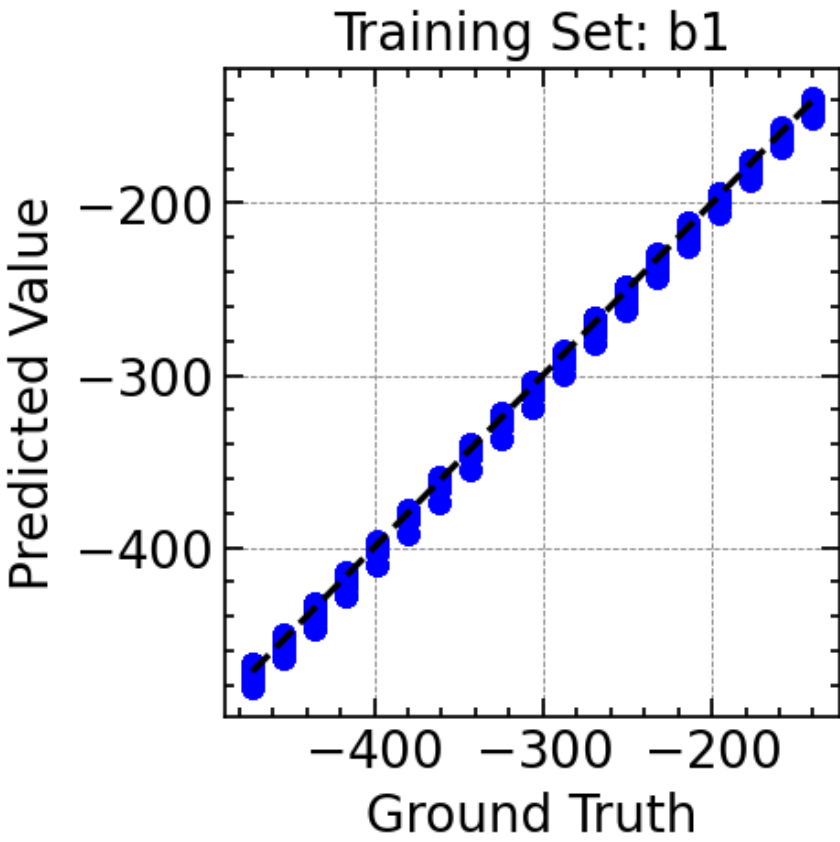


1 parameter

$$R^2 = 1 - \frac{\sum_{i=1}^n (y_i - \hat{y})^2}{\sum_{i=1}^n (y_i - \bar{y})^2}$$

R-squared: 0.99, 0.99, 0.99, 0.99, 0.99

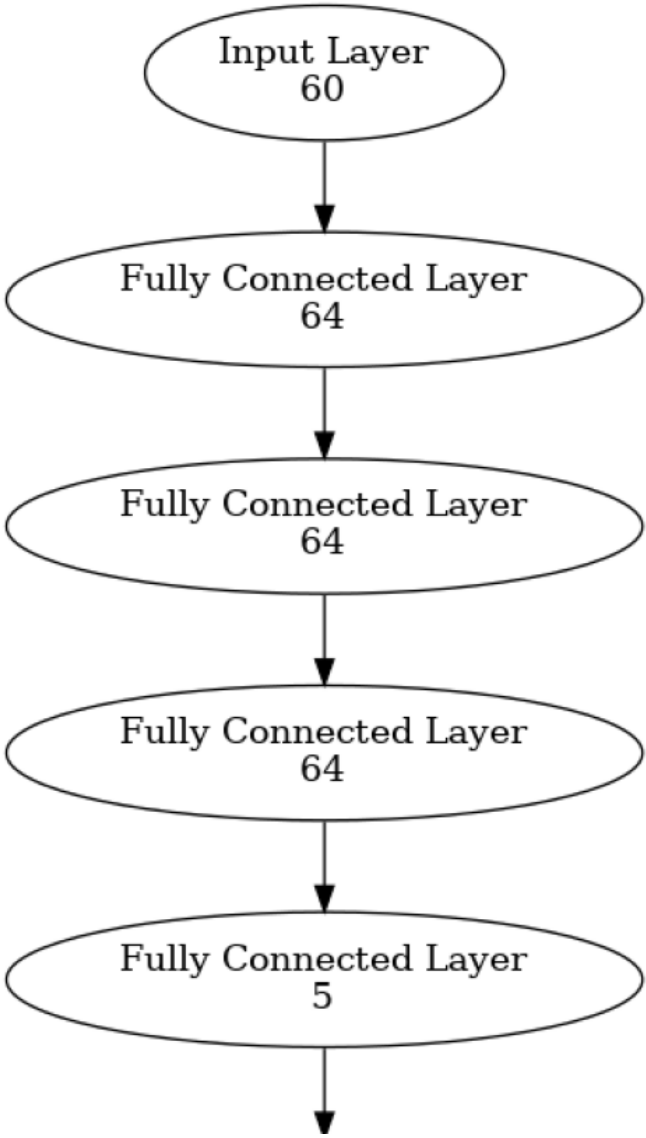
Inverse Femtoscopy



R-squared	b1	b3
Training	0.99	0.99
Testing	0.99	0.99

in Preparation

with Jiaxing Zhao, etc.



60 points

b_1, b_3
2 parameters

$r_0 = 1.3 \text{ fm}, b_2 = 73.9, b_4 = 0.78, n_\pi = 2$

$N_c = 10000$

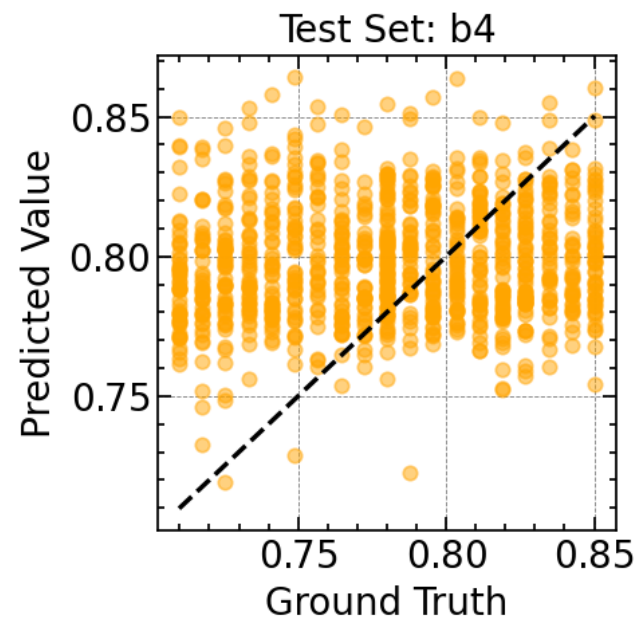
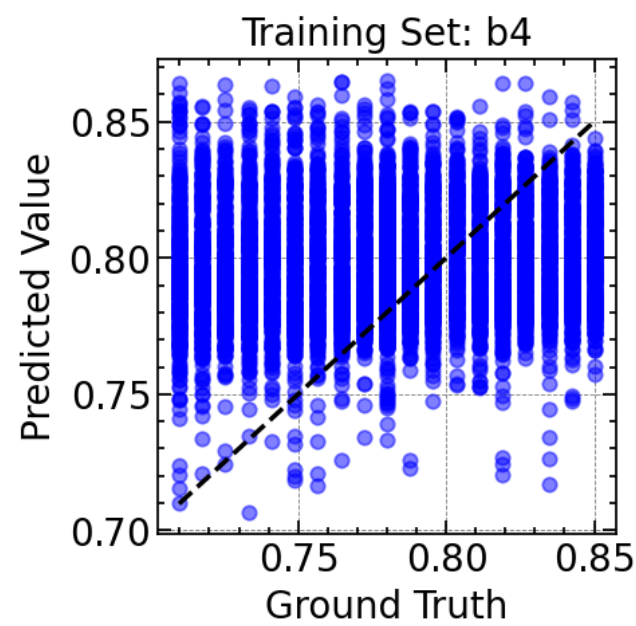
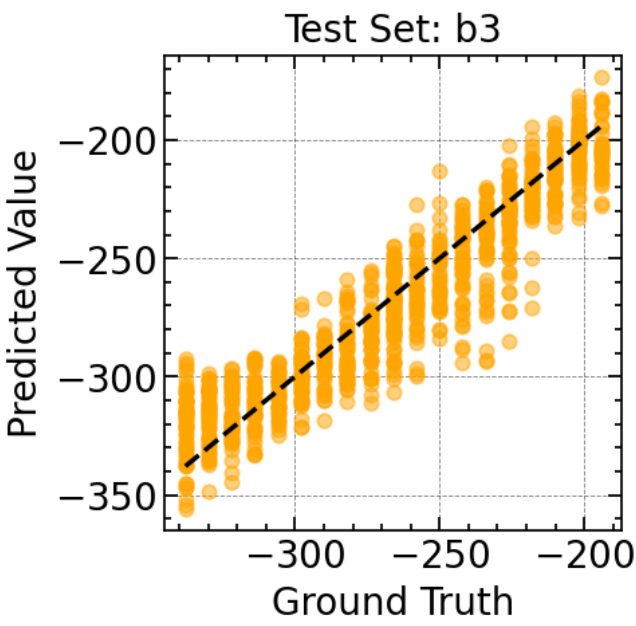
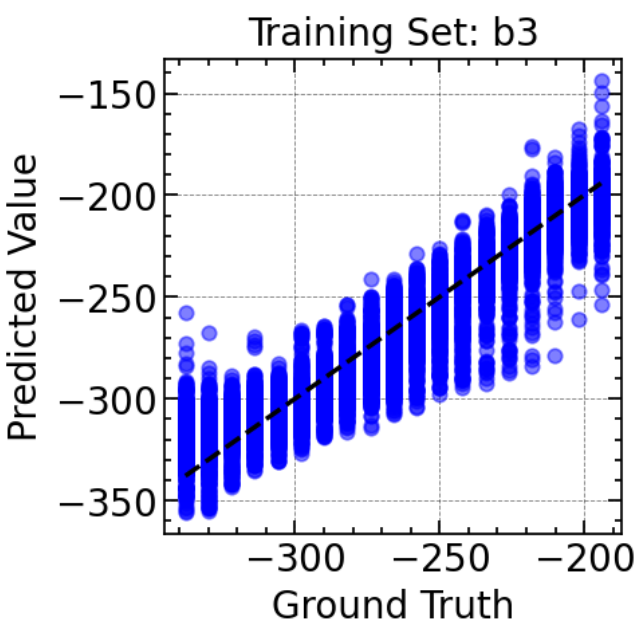
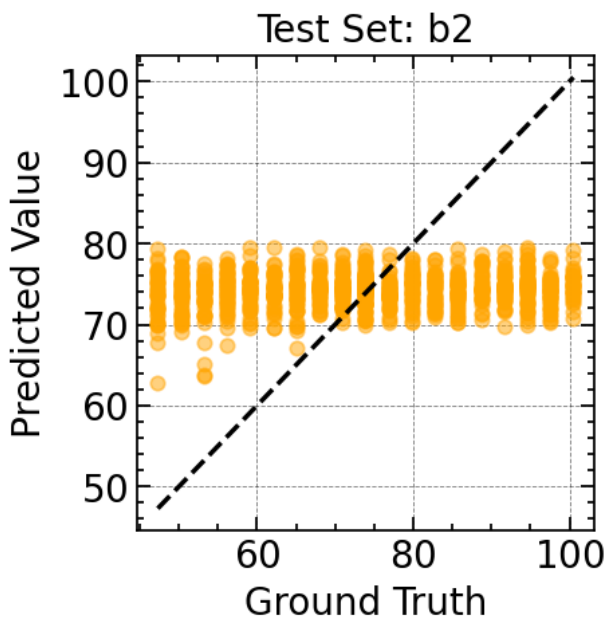
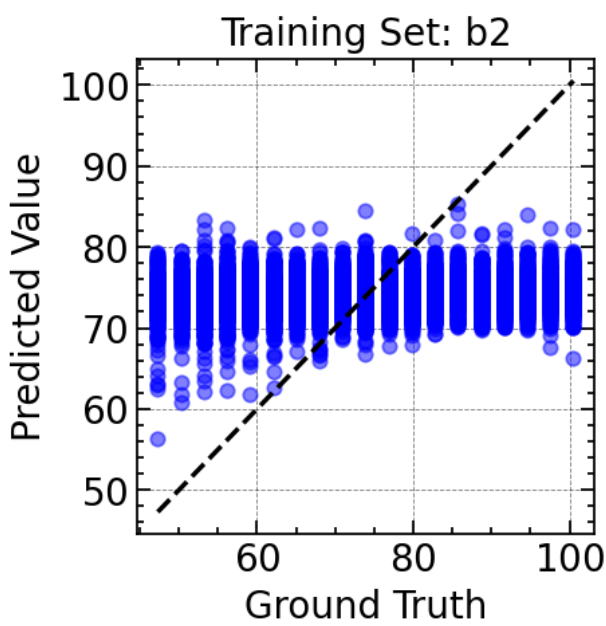
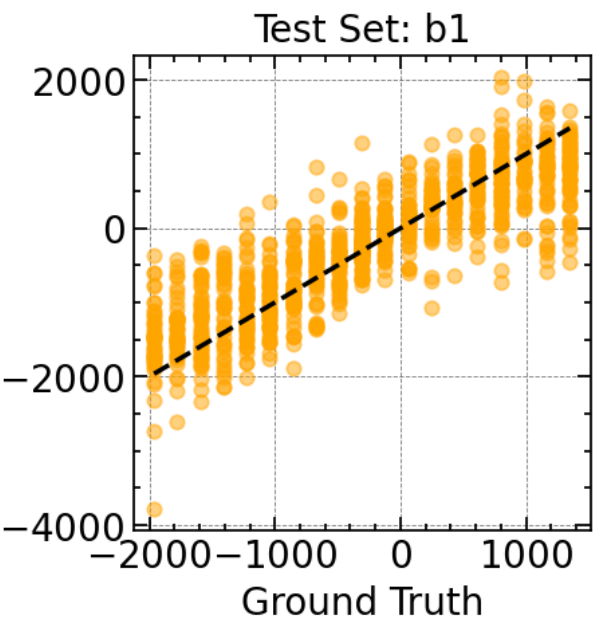
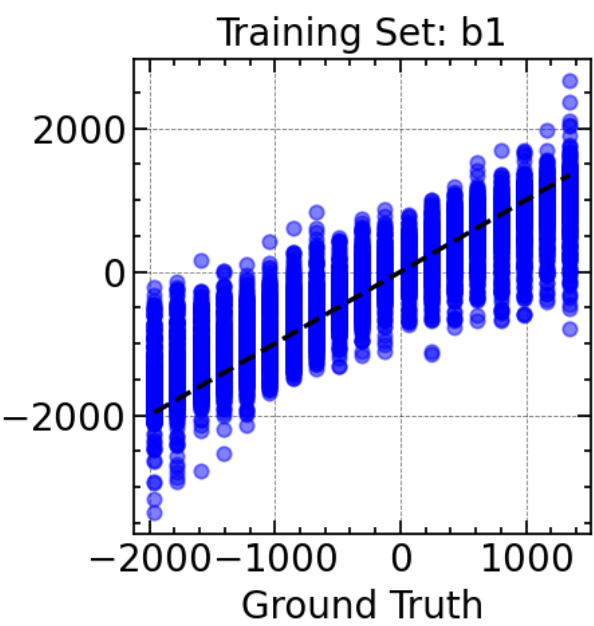
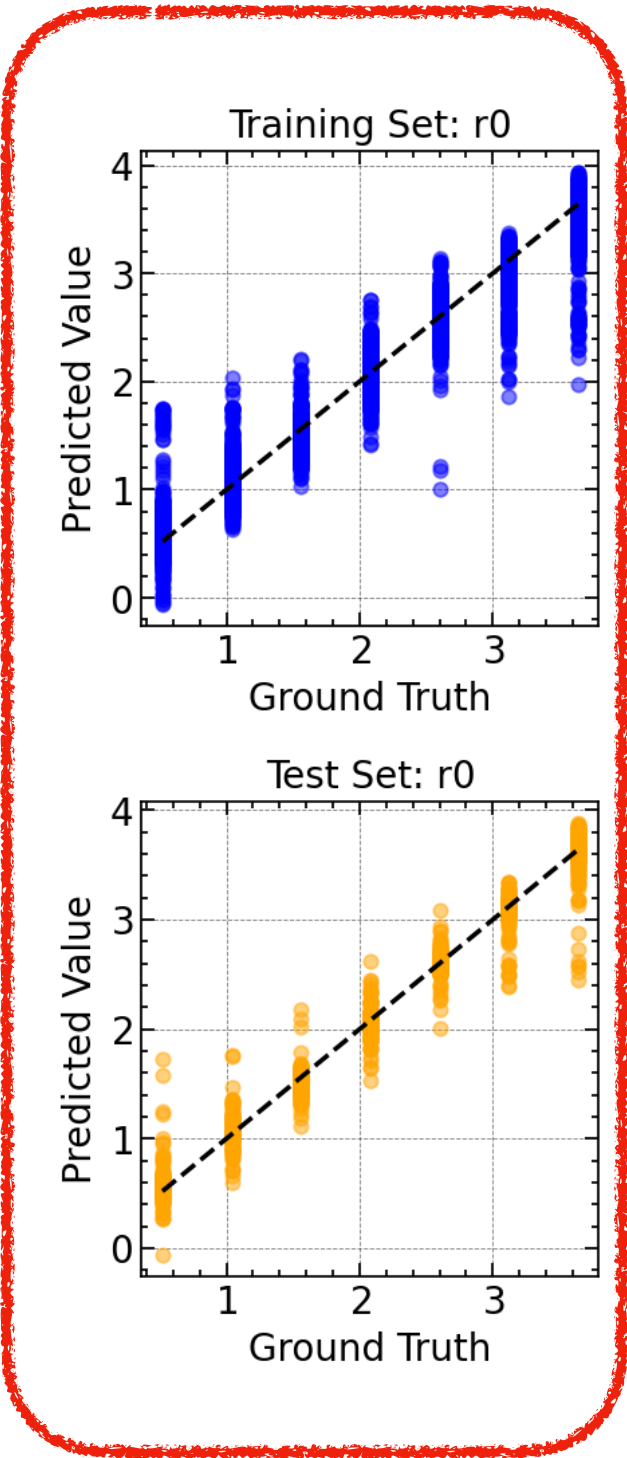
Inverse Femtoscopy

$N_c = 10000$

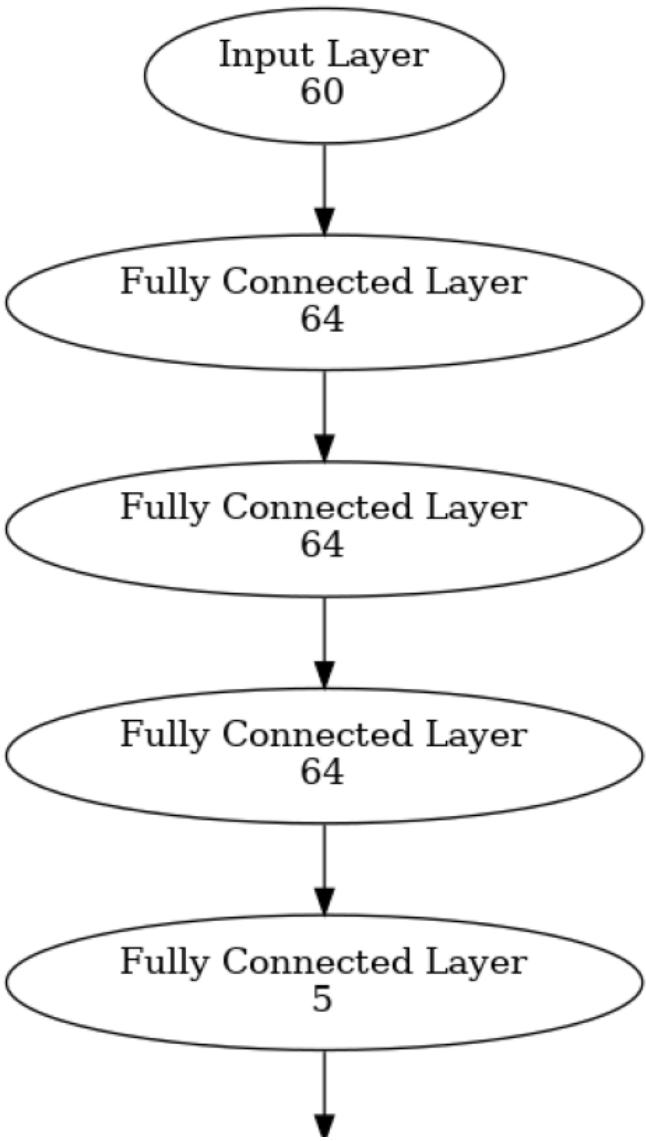
$n_\pi = 2$

in Preparation

with Jiaxing Zhao, etc.



R-squared	r0	b1	b2	b3	b4
Training	0.99	0.86	0.02	0.90	0.00
Testing	0.99	0.86	0.02	0.90	0.00



60 points

5 parameters

Potential Functions

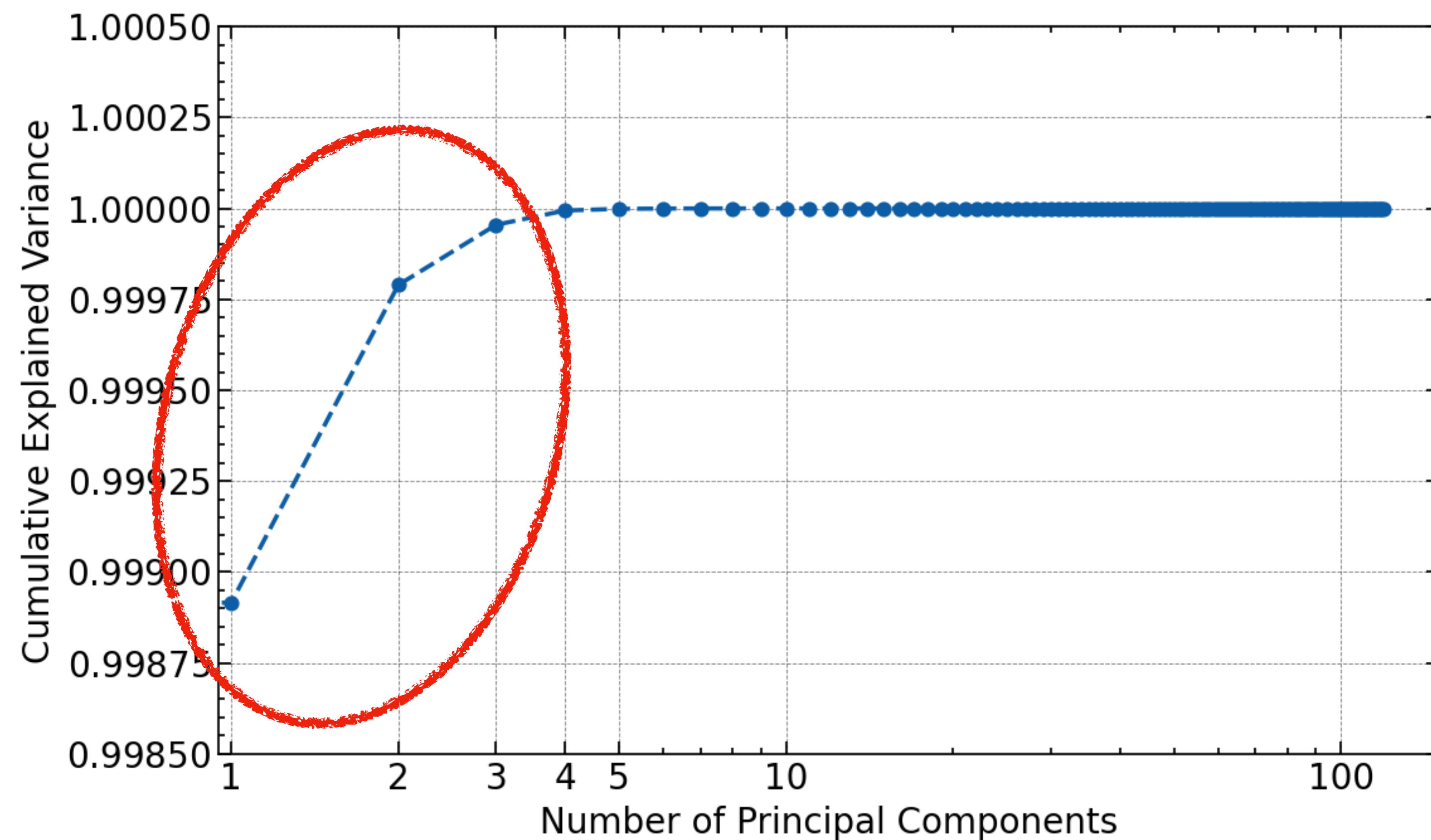
$$V(r) = b_1 e^{-b_2 r^2} + b_3 (1 - e^{-b_4 r^2}) \left(\frac{e^{(-m_\pi r)}}{r} \right)^{n_\pi}$$

Inverse Femtoscopy

in Preparation

with Jiaxing Zhao, etc.

Principal Component Analysis(PCA)



$$N_c = 25000$$

$$r_0: [0.52, 4.16] \text{ fm}$$

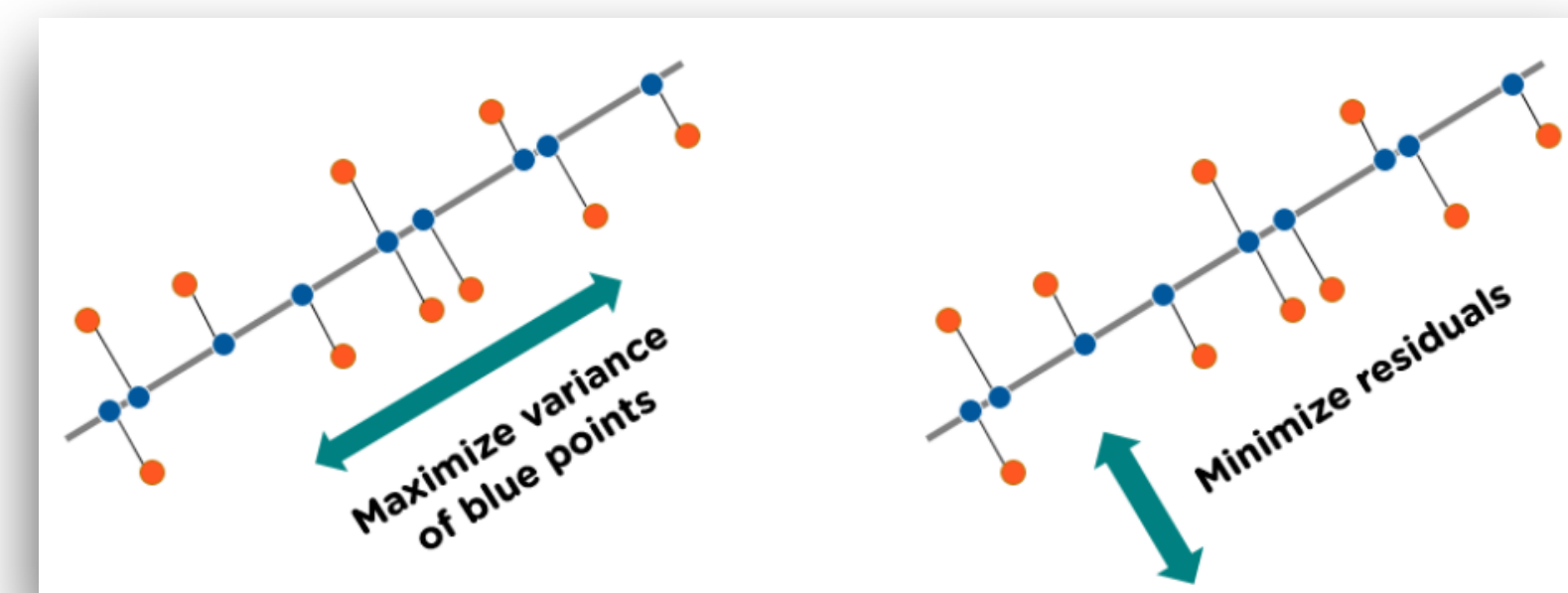
$$b_1: [-2145.5, 1532.5] \text{ MeV}$$

$$b_2: [0.739, 147.761] \text{ fm}^{-2}$$

$$b_3: [-1064, 532] \text{ MeV} \cdot \text{fm}^2$$

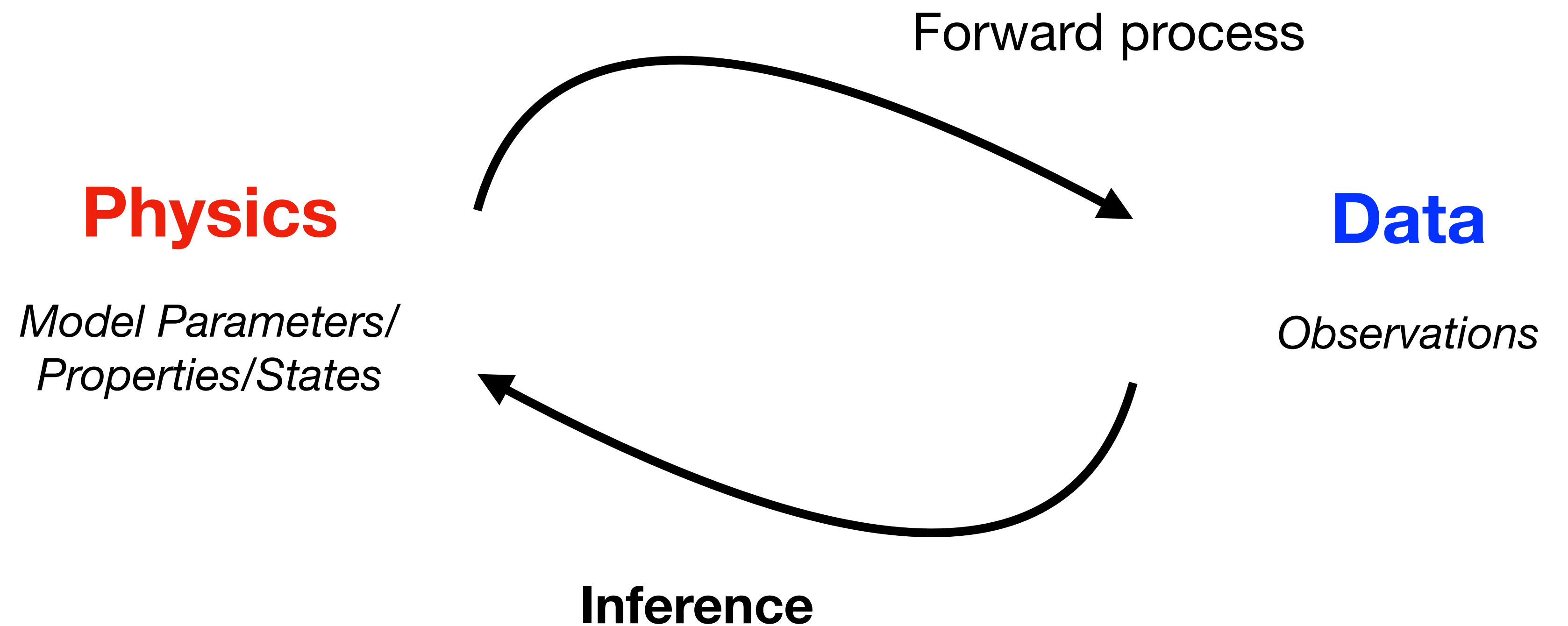
$$b_4: [0.078, 154.422] \text{ fm}^{-2}$$

$$n_\pi: 2$$



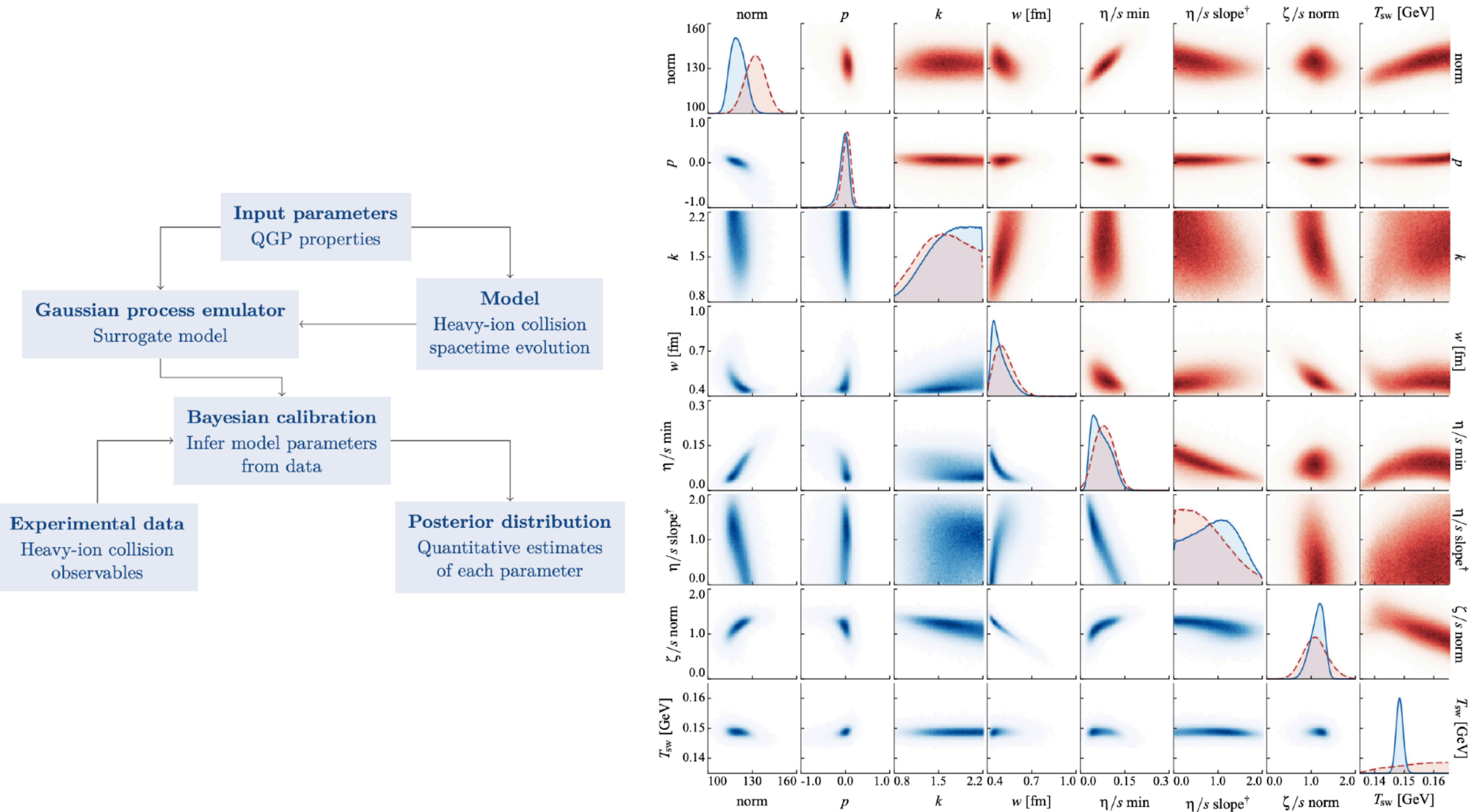
Bayesian Inference

$$\hat{\theta} = \arg \max_{\theta} \{p(X | \theta)\}$$



Bayesian Inference

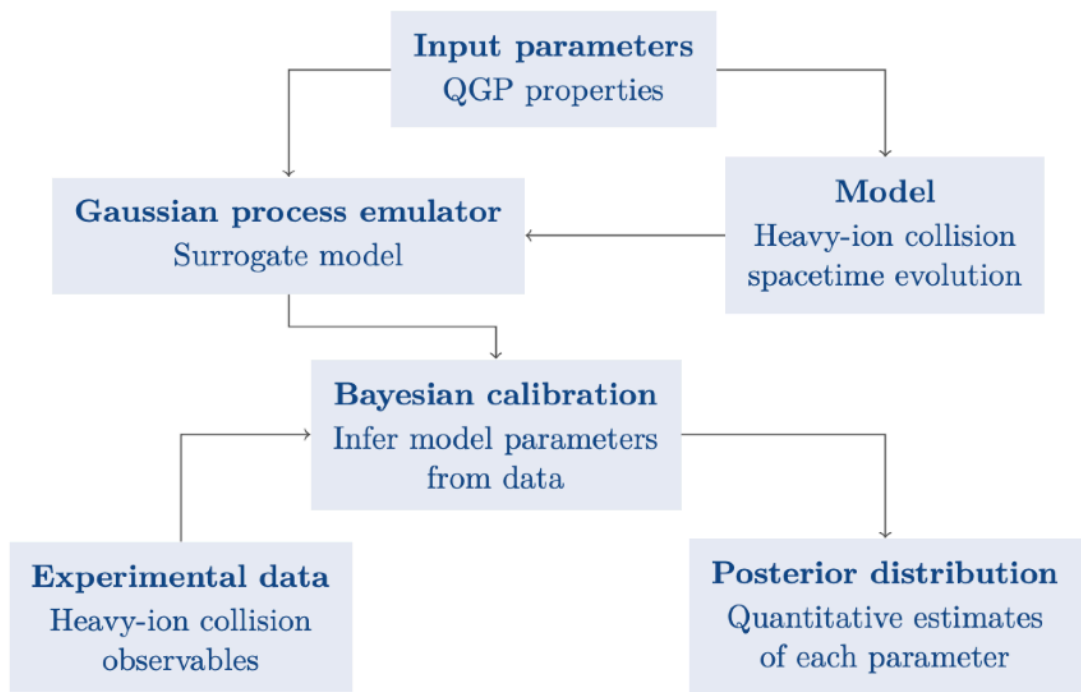
Inferring Dynamical Information



J. E. Bernhard, etc.(Duke Group), [arXiv:1804.06469](https://arxiv.org/abs/1804.06469) and Nat Phys **15**, 1113 (2019).

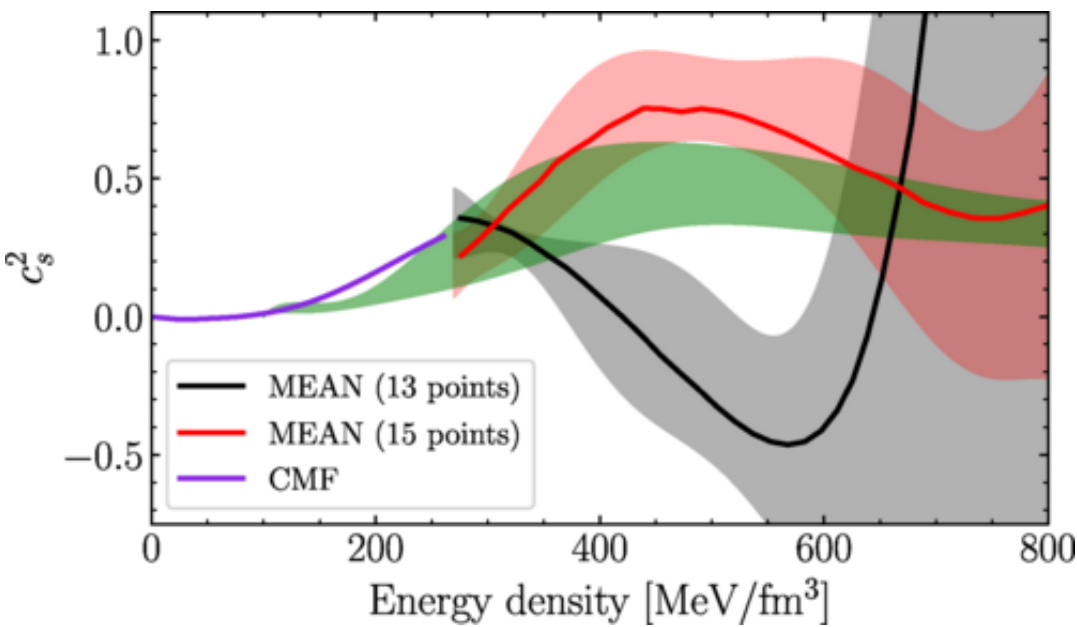
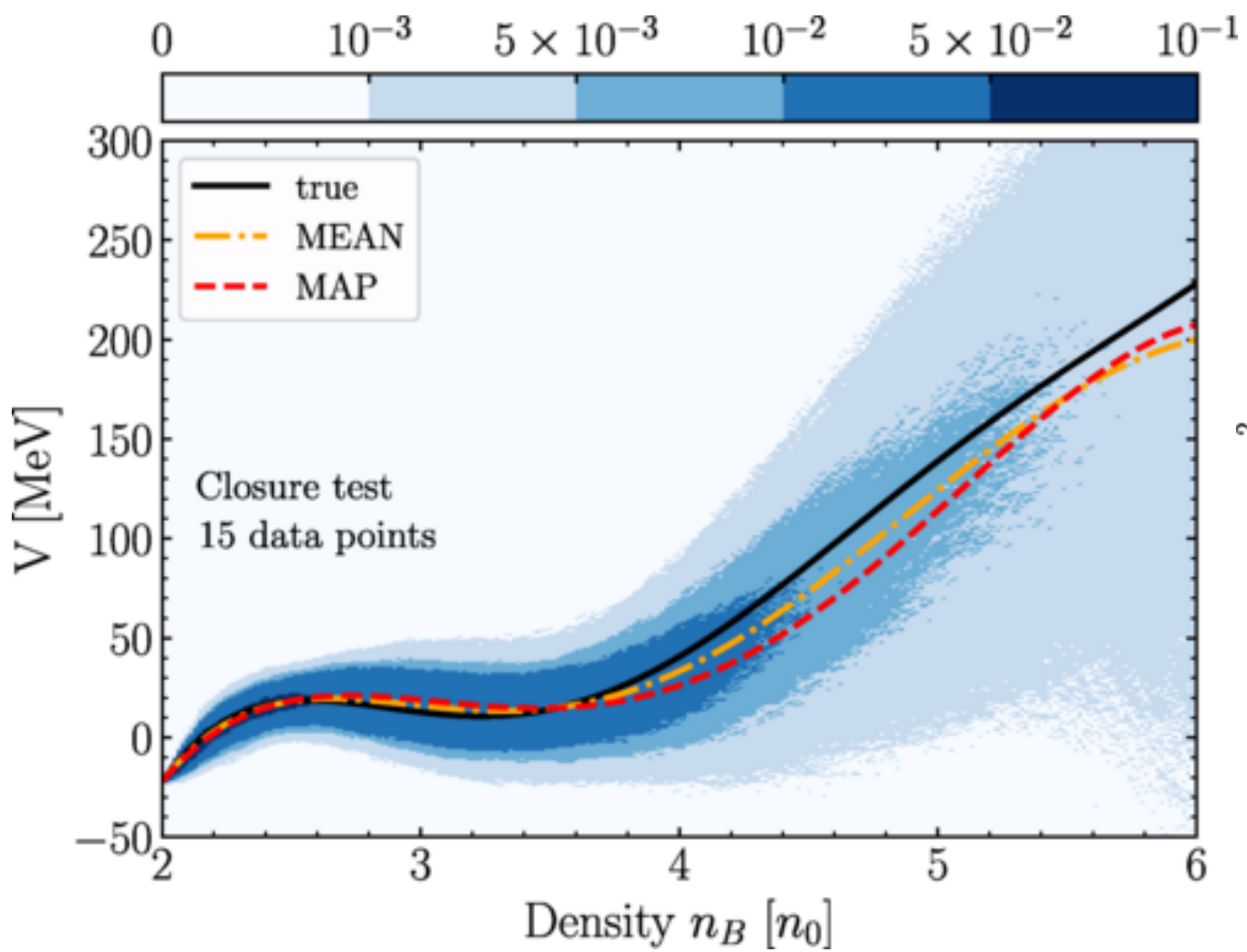
Bayesian Inference

Inferring Dynamical Information



J. E. Bernhard, etc.(Duke Group), [arXiv:1804.06469](#)

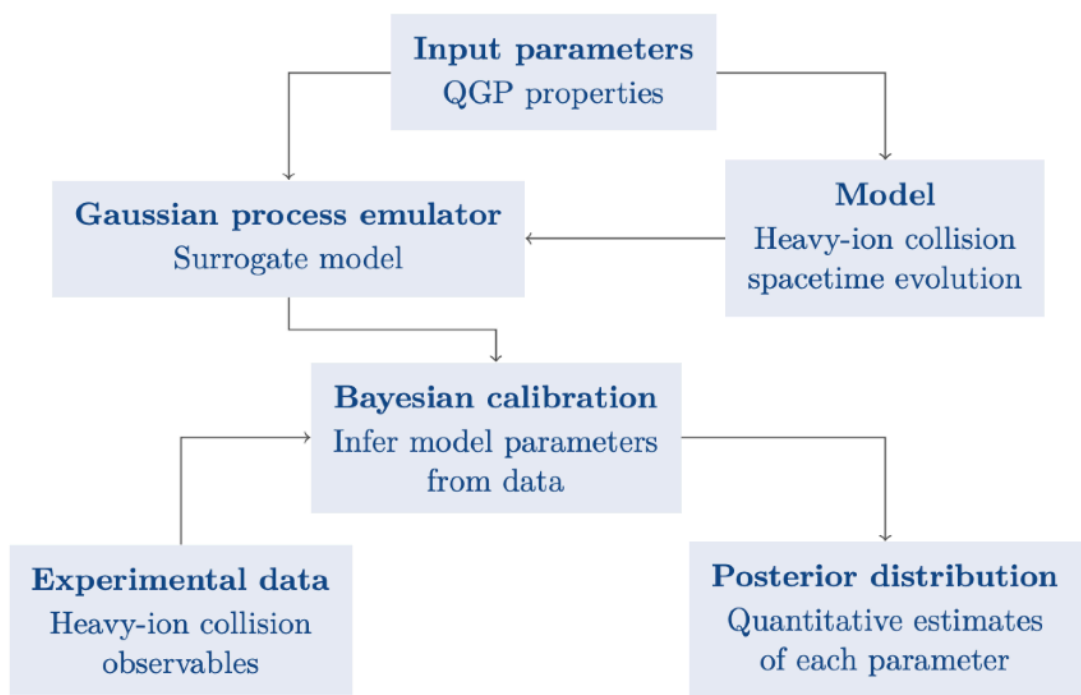
Ectracting Dense Matter EoSs



M. Omana Kuttan, J. Steinheimer, K. Zhou, and H. Stoecker, Phys. Rev. Lett. **131**, 202303 (2023).

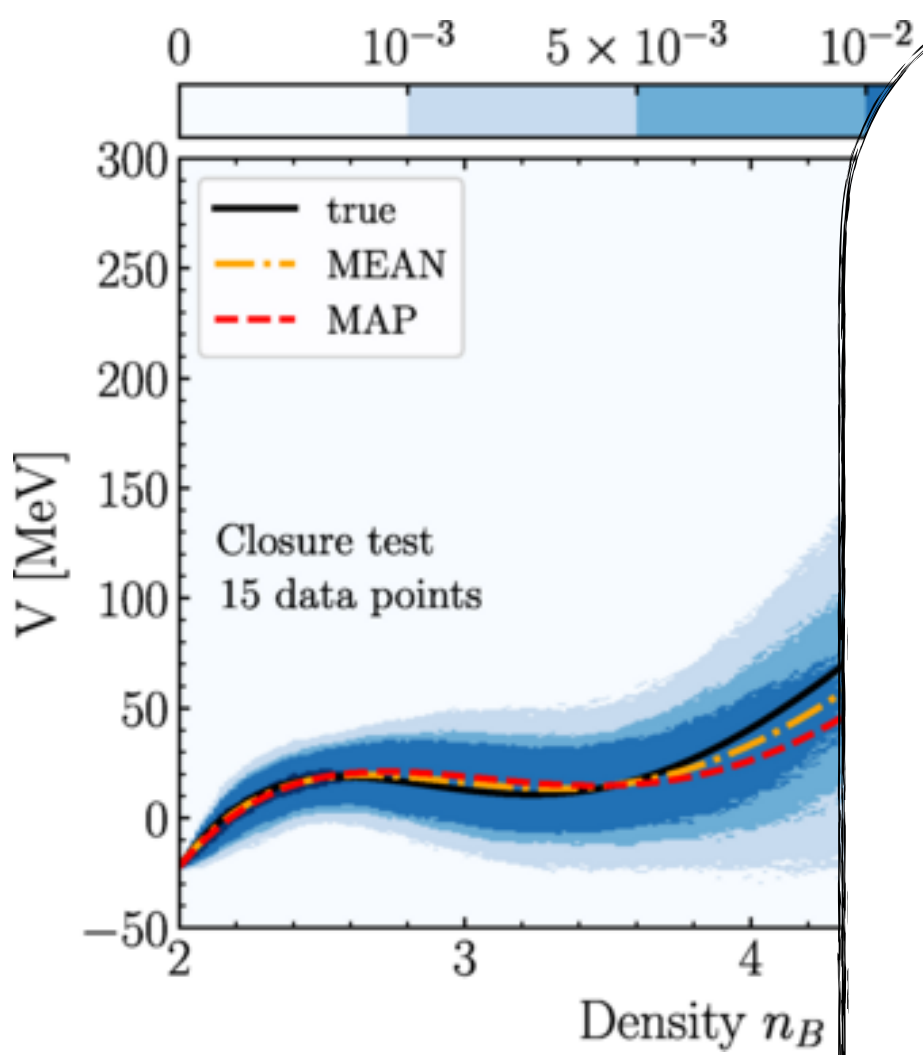
Bayesian Inference

Inferring Dynamical Information



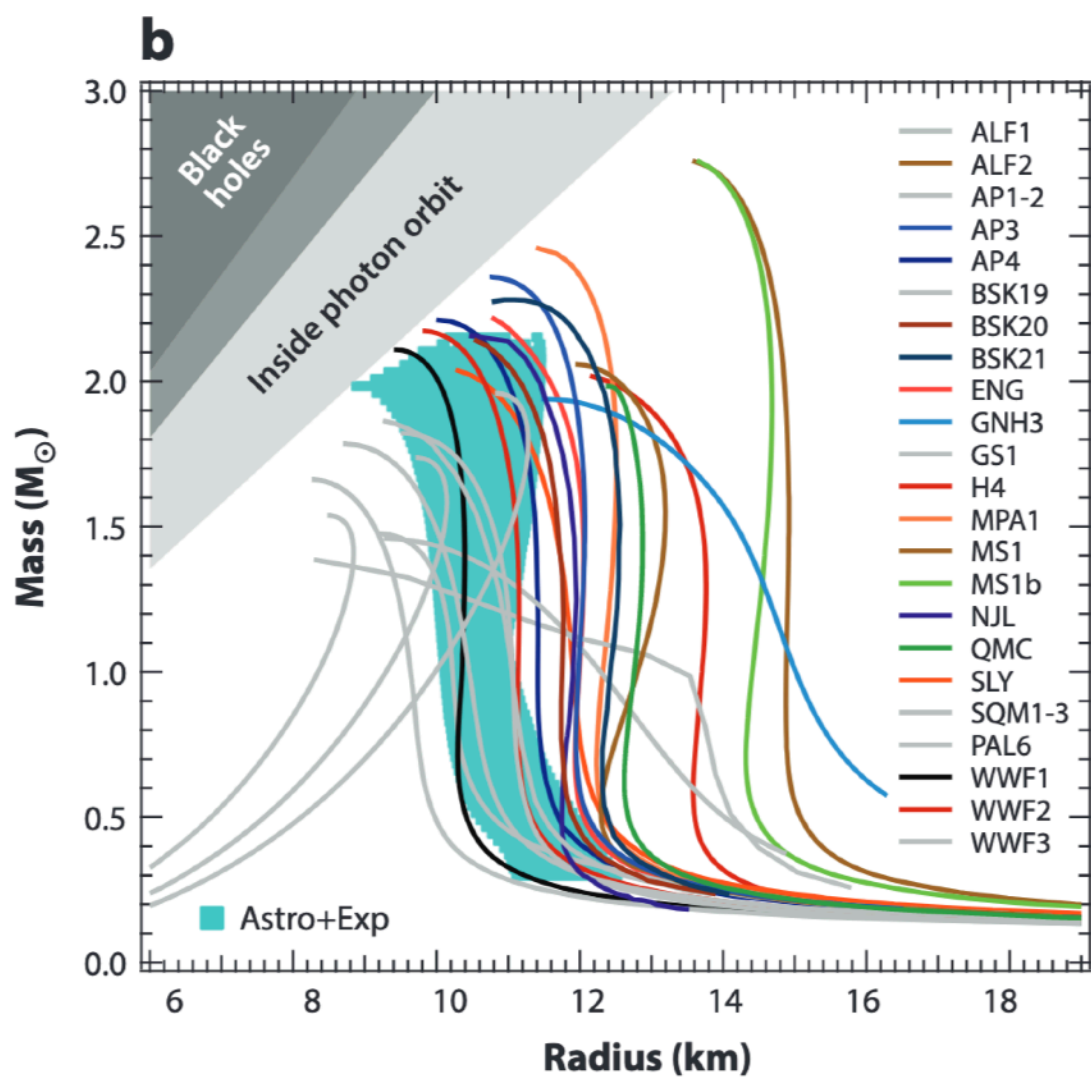
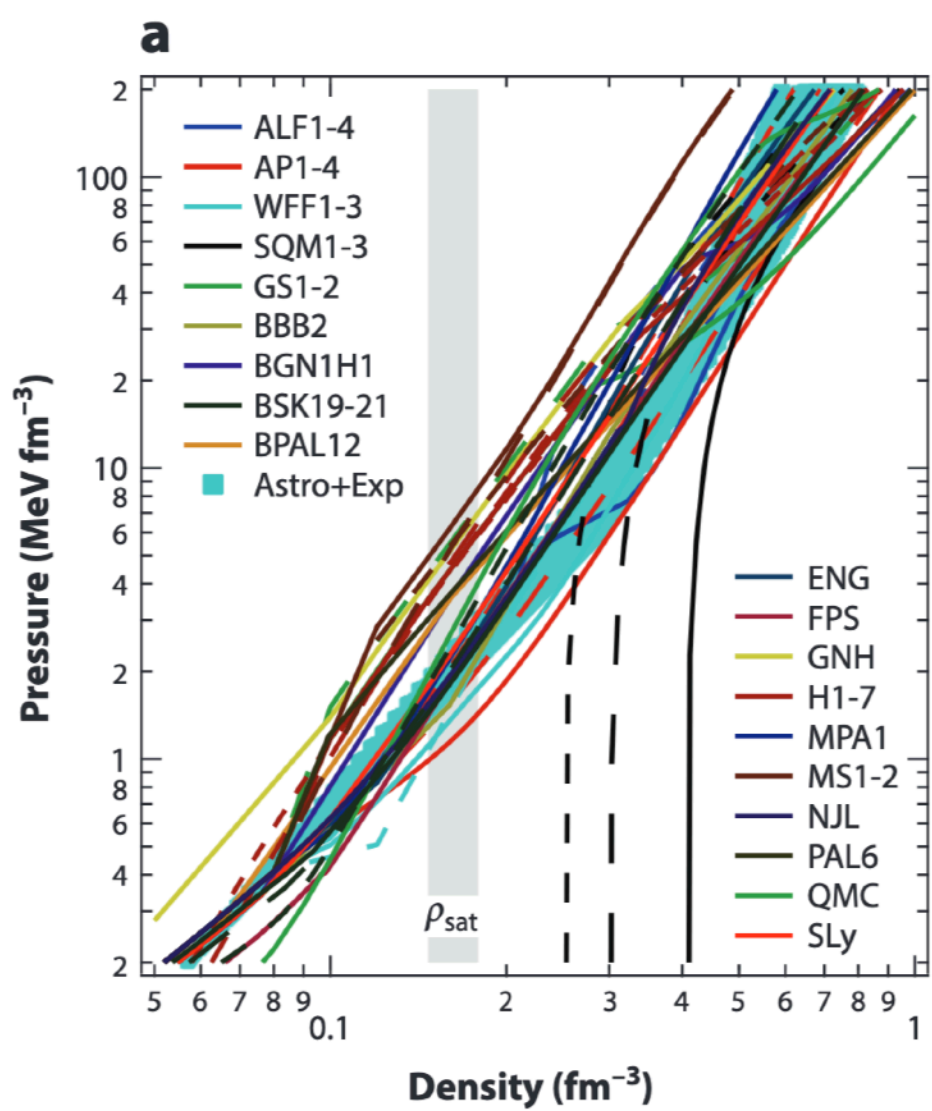
J. E. Bernhard, etc.(Duke Group), [arXiv:1804.06469](#)

Ectracting Dense Matter EoSs



M. Omana Kuttan, J. Ste

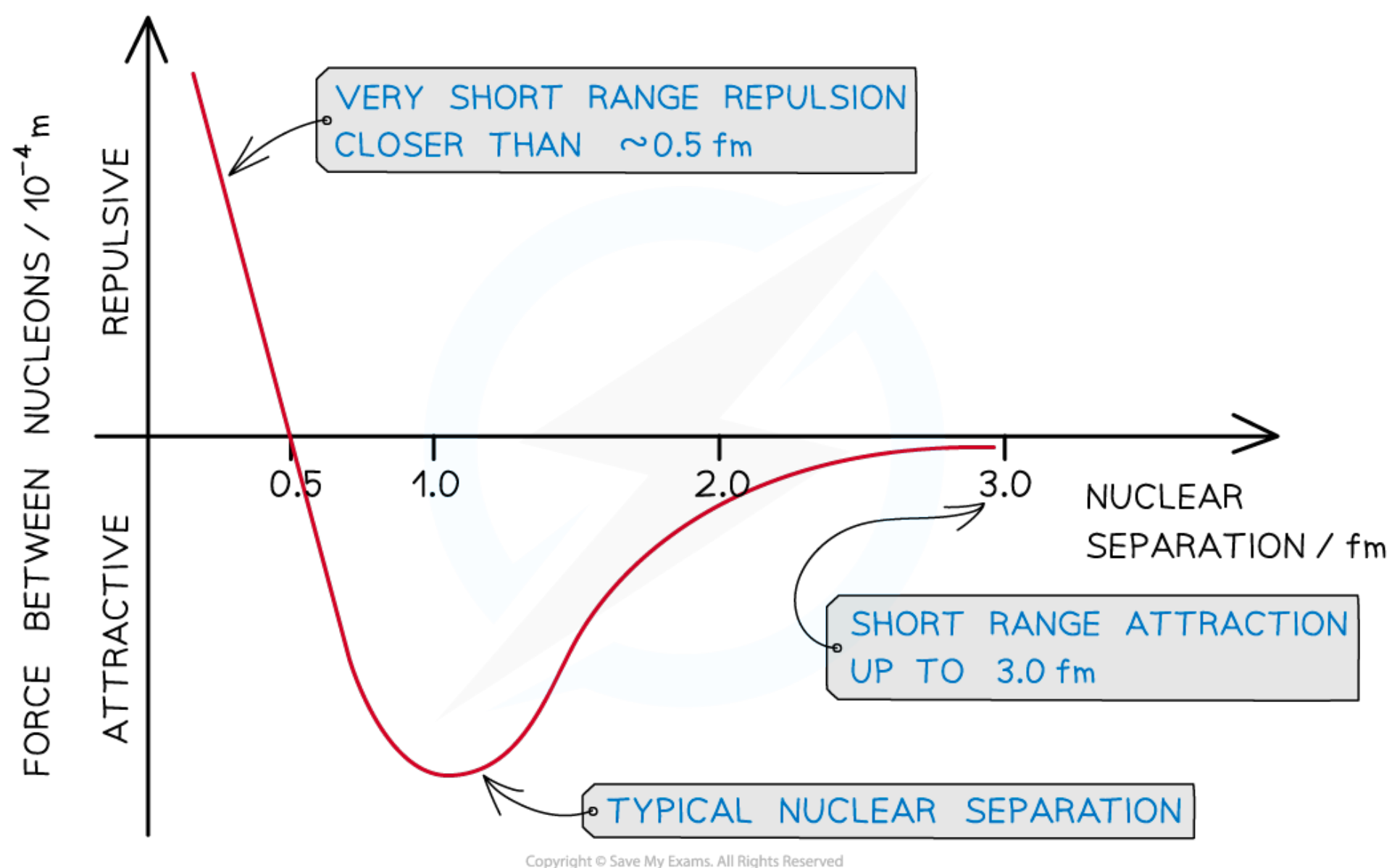
Building Nuclear Matter EoS



F. Özel and P. Freire, Annu. Rev. Astron. Astrophys. **54**, 401 (2016)

Bayesian Inference

Building Nuclear Forces(Interactions)

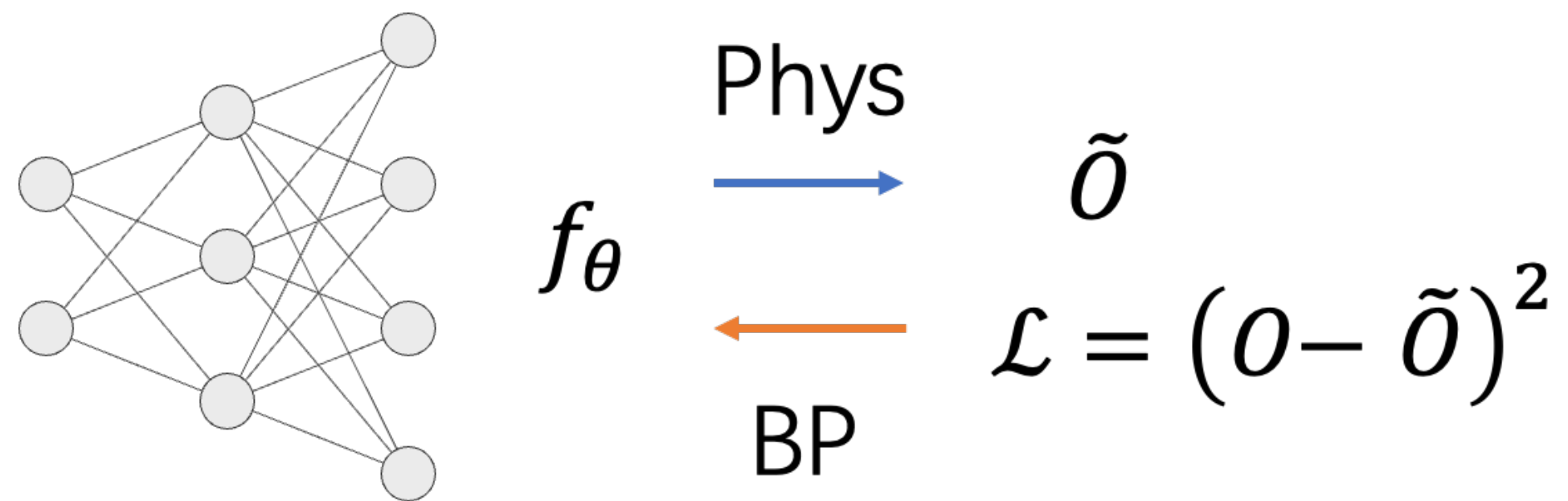


Physics Parameters are Finite
EoS, Wave-Function, Potential, **X**

Inference is Easy-To-Compute
ODEs, PDEs, Simulations, ... **X**

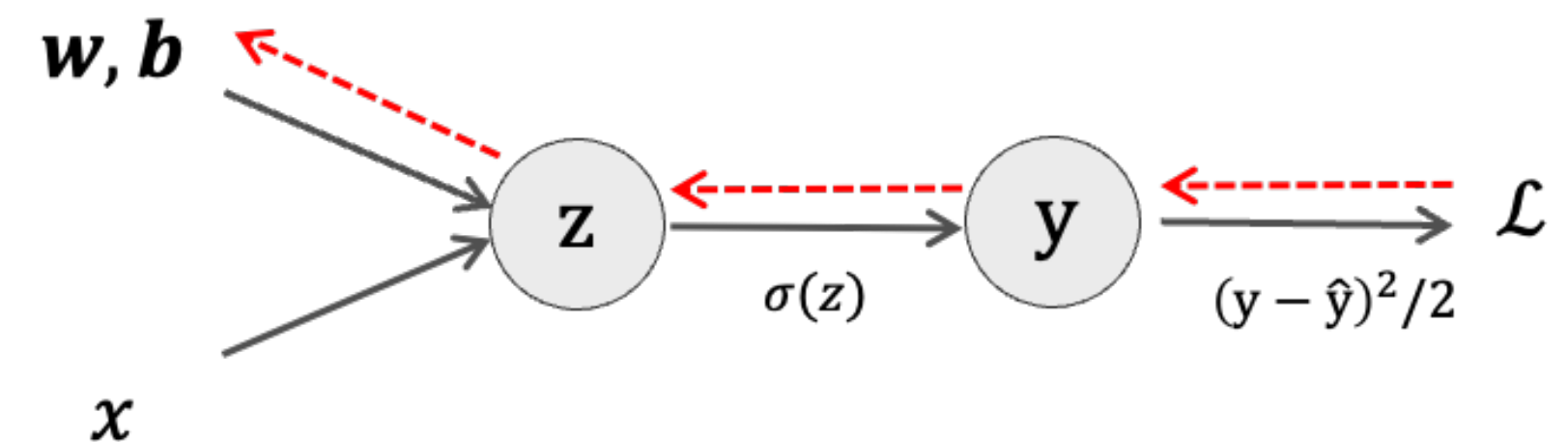
Physics-Driven Deep Learning

$$\hat{\theta} = \arg \max_{\theta} \{p(X | \theta)\}$$



Deep Neural Network represented Physics, f_{θ}

Flexible Representation

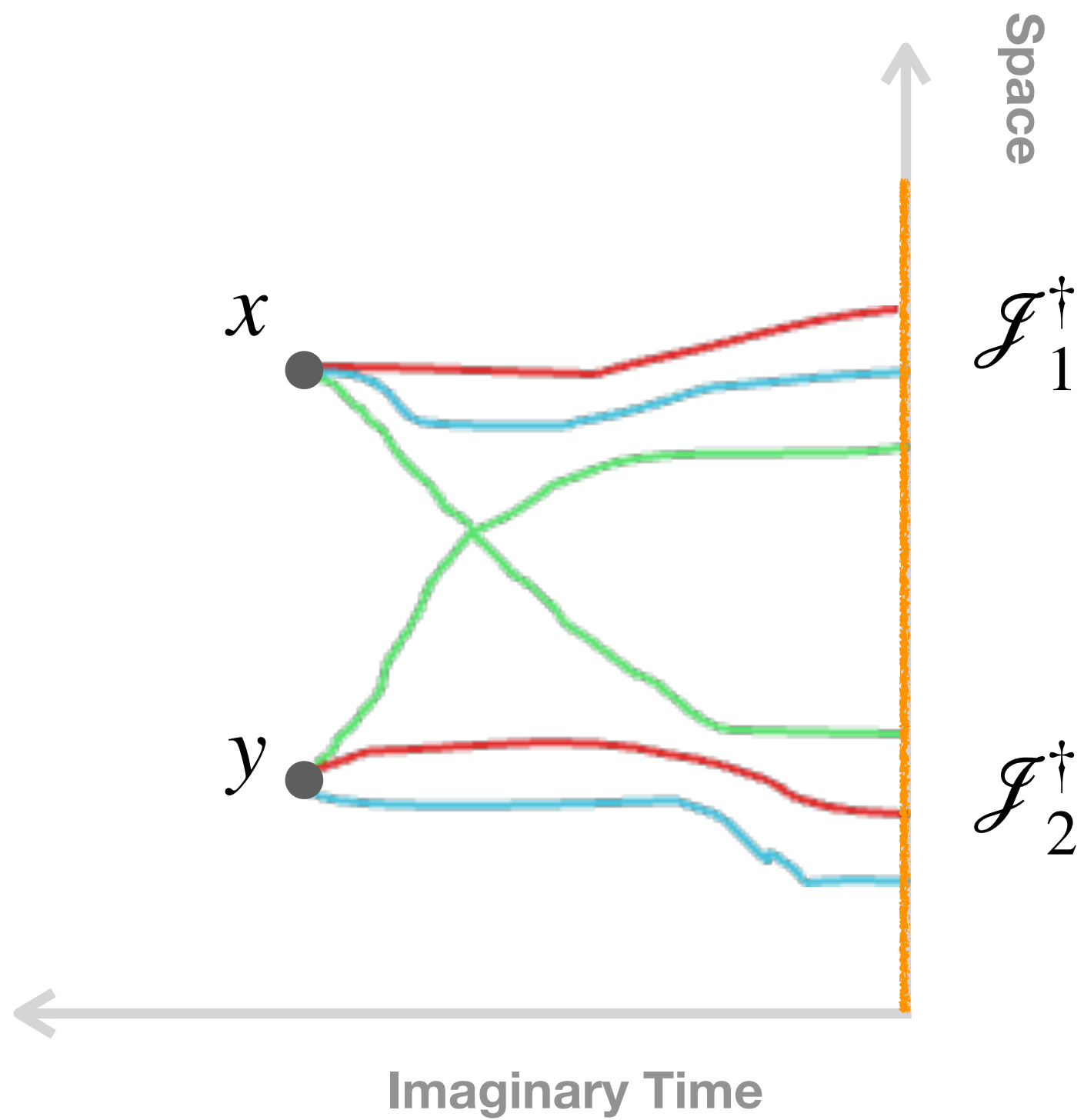


Back-Propagation

Easy-To-Compute on GPUs

HAL QCD method

N. Ishii, S. Aoki, and T. Hatsuda, Phys. Rev. Lett. 99, 022001 (2007)
S. Aoki, T. Hatsuda, and N. Ishii, Prog. Theor. Phys. 123, 89 (2010)
Aoki, S., Doi, T., Front. in Phys. 8, 307 (2020)
S. Aoki and T. Doi, in Handbook of Nuclear Physics(2023), pp. 1–31

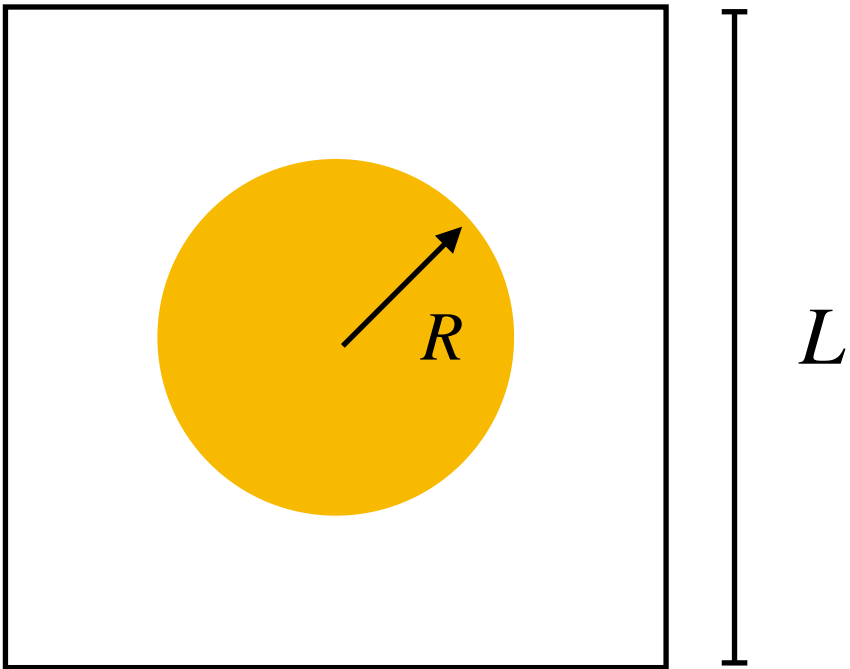


$$\langle N_1(\mathbf{x}, t) N_2(\mathbf{y}, t) \mathcal{J}_1^\dagger(0) \mathcal{J}_2^\dagger(0) \rangle$$
$$= \sum_n \langle 0 | N_1(\mathbf{x}) N_2(\mathbf{y}) | n \rangle a_n e^{-E_n t}$$
$$\xrightarrow{t > t^*} \phi(\mathbf{r}, t) = \sum_{n < n^*} b_n \phi_n(\mathbf{r}) e^{-E_n t}$$

$$(E_k - H_0) \phi_{\mathbf{k}}(\mathbf{r}) = \int d^3 r' U(\mathbf{r}, \mathbf{r}') \phi_{\mathbf{k}}(\mathbf{r}'), \quad r < R$$

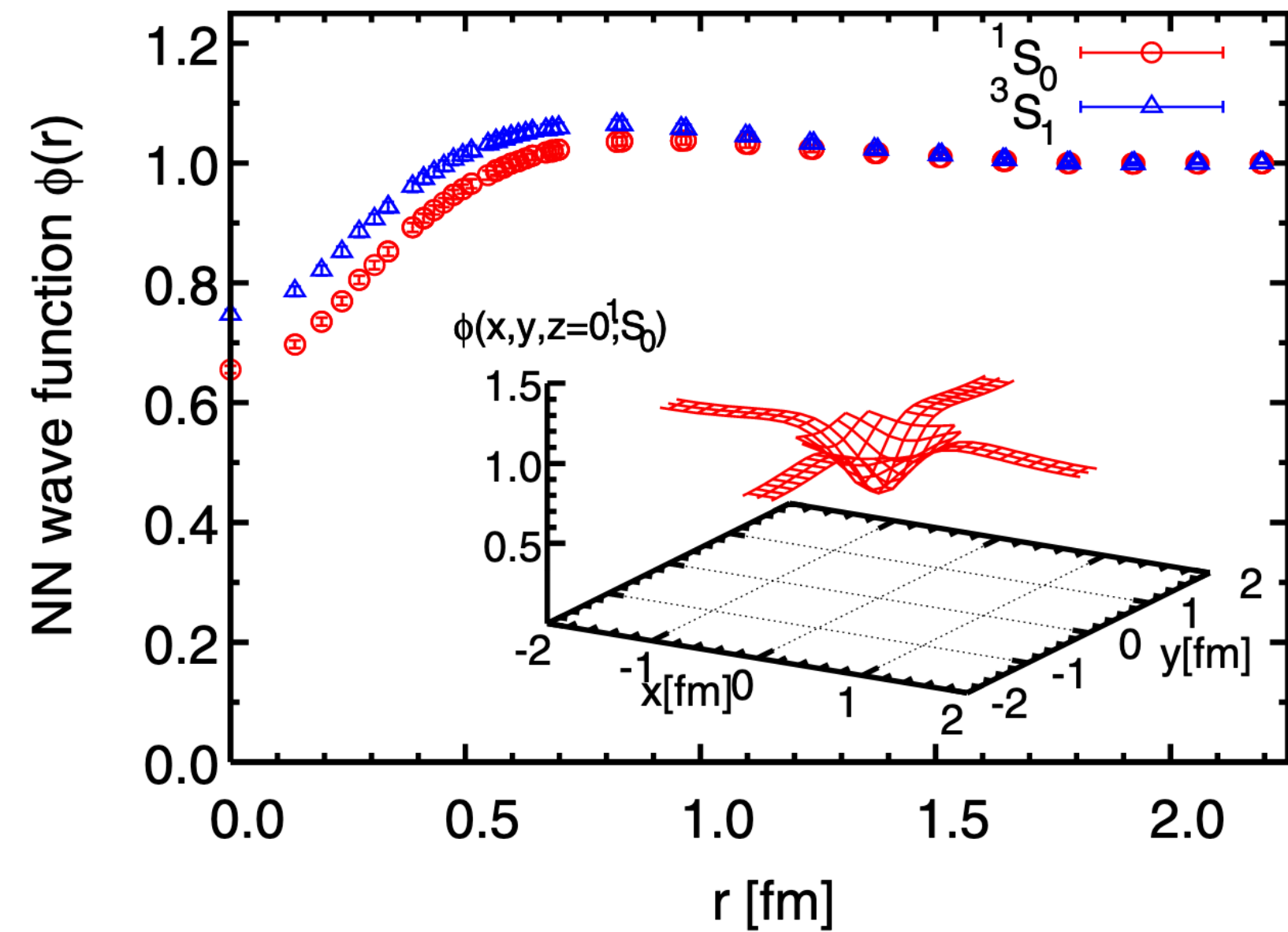
Consider the wave function at “interacting region” → Phase shift, Binding energy

$$\phi(\mathbf{r}, t) \rightarrow \text{2 PI Kernel}$$



HAL QCD method

N. Ishii, S. Aoki, and T. Hatsuda, Phys. Rev. Lett. 99, 022001 (2007)
S. Aoki, T. Hatsuda, and N. Ishii, Prog. Theor. Phys. 123, 89 (2010)
Aoki, S., Doi, T., Front. in Phys. 8, 307 (2020)
S. Aoki and T. Doi, in Handbook of Nuclear Physics(2023), pp. 1–31

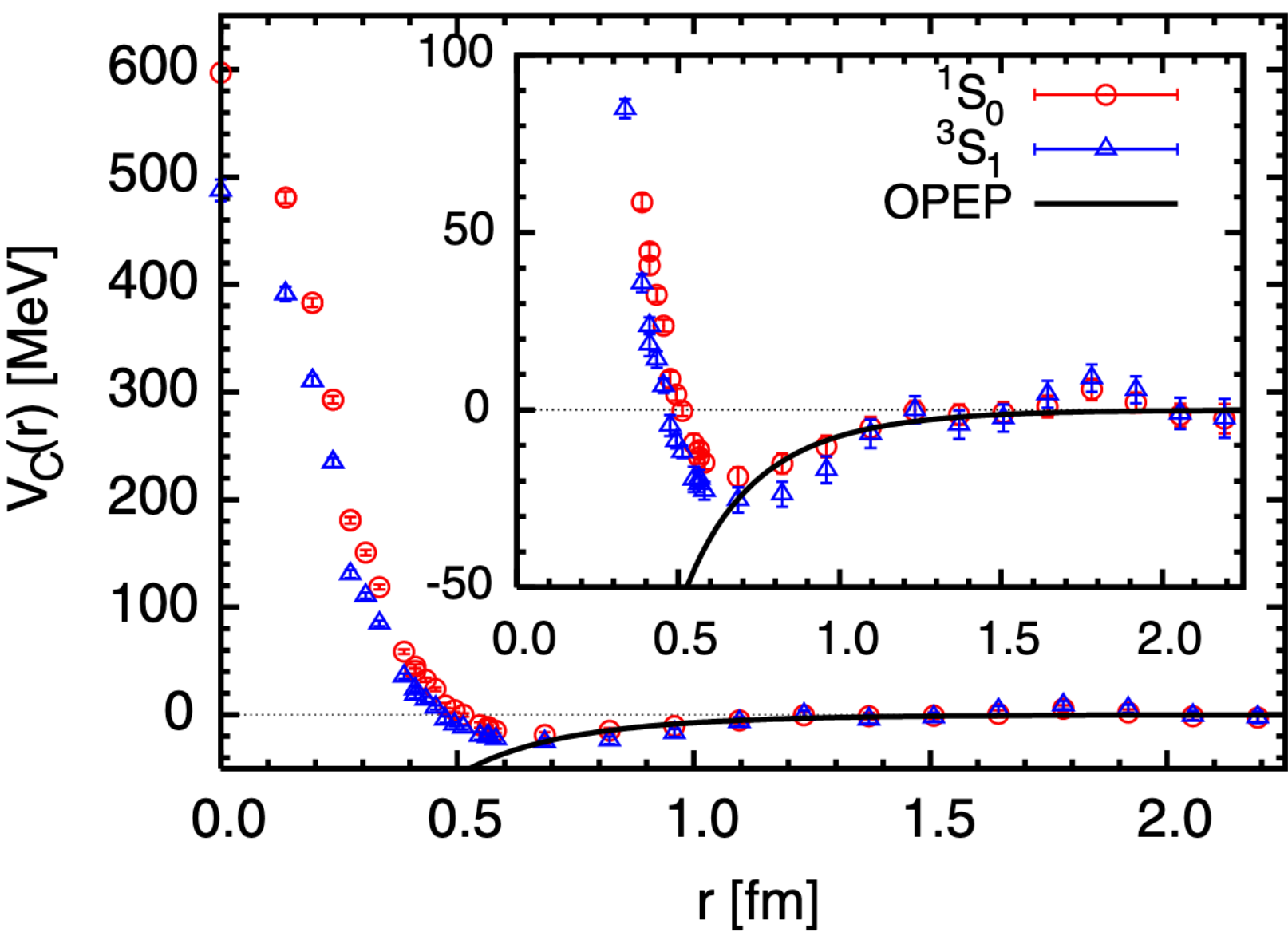


Nambu-Bethe-Salpeter (NBS) wave function

$$\begin{aligned}\psi_{NBS}(\vec{r}) &= \langle 0 | N(\vec{r}) N(\vec{0}) | N(\vec{k}) N(-\vec{k}), in \rangle \\ &\simeq e^{i\delta_l(k)} \sin(kr - l\pi/2 + \delta_l(k)) / (kr) \\ &\quad \text{(at asymptotic region)}\end{aligned}$$

Local Approx.
Derivative Expansion

→
HAL QCD method

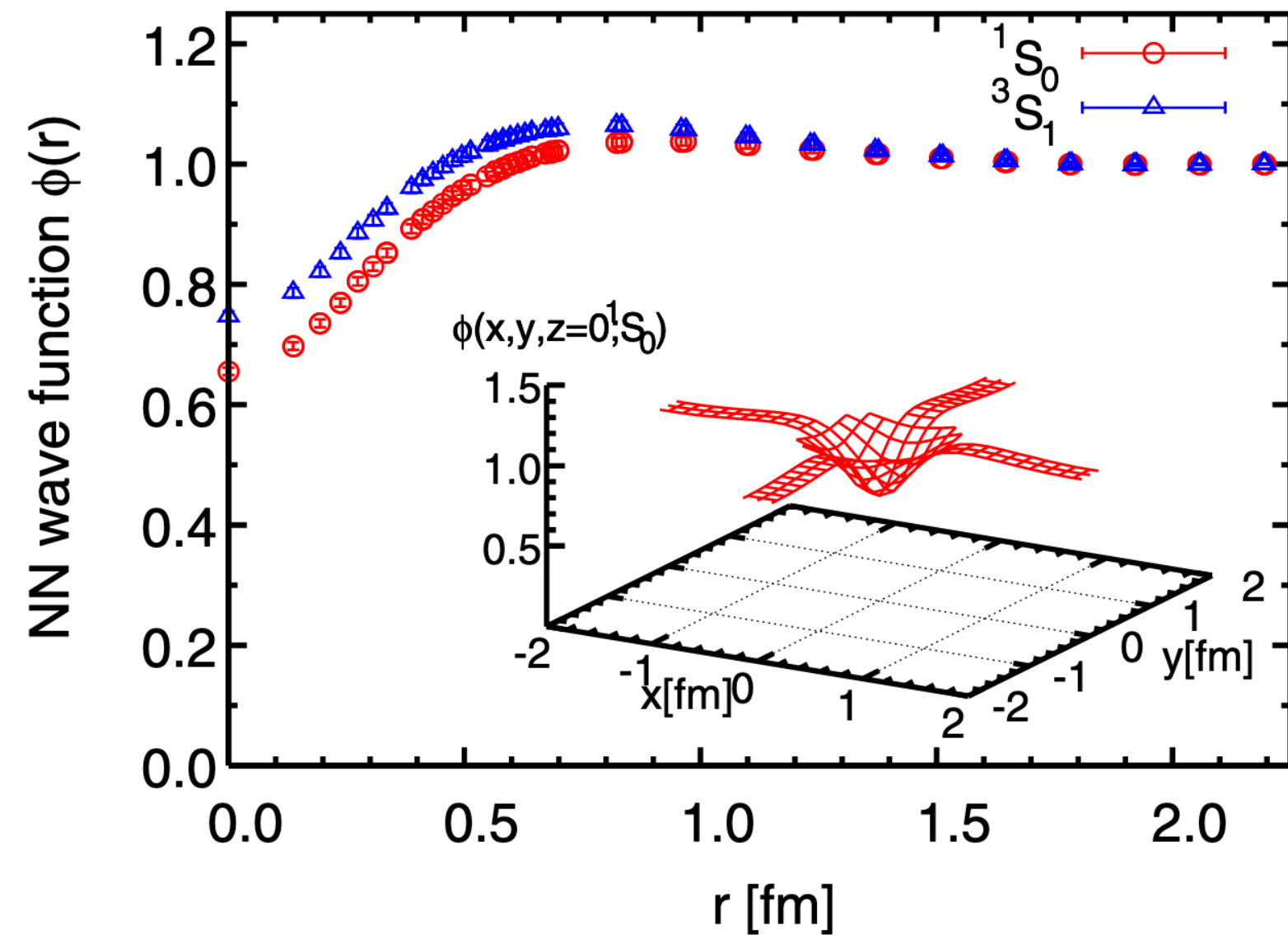


Nuclear Force

$$\begin{aligned}&(k^2/m_N - H_0) \psi_{NBS}(\vec{r}) \\ &= \int d\vec{r}' U(\vec{r}, \vec{r}') \psi_{NBS}(\vec{r}') \\ &\quad \text{(Schrodinger eq.)}\end{aligned}$$

HAL QCD: Inverse Problem Perspective

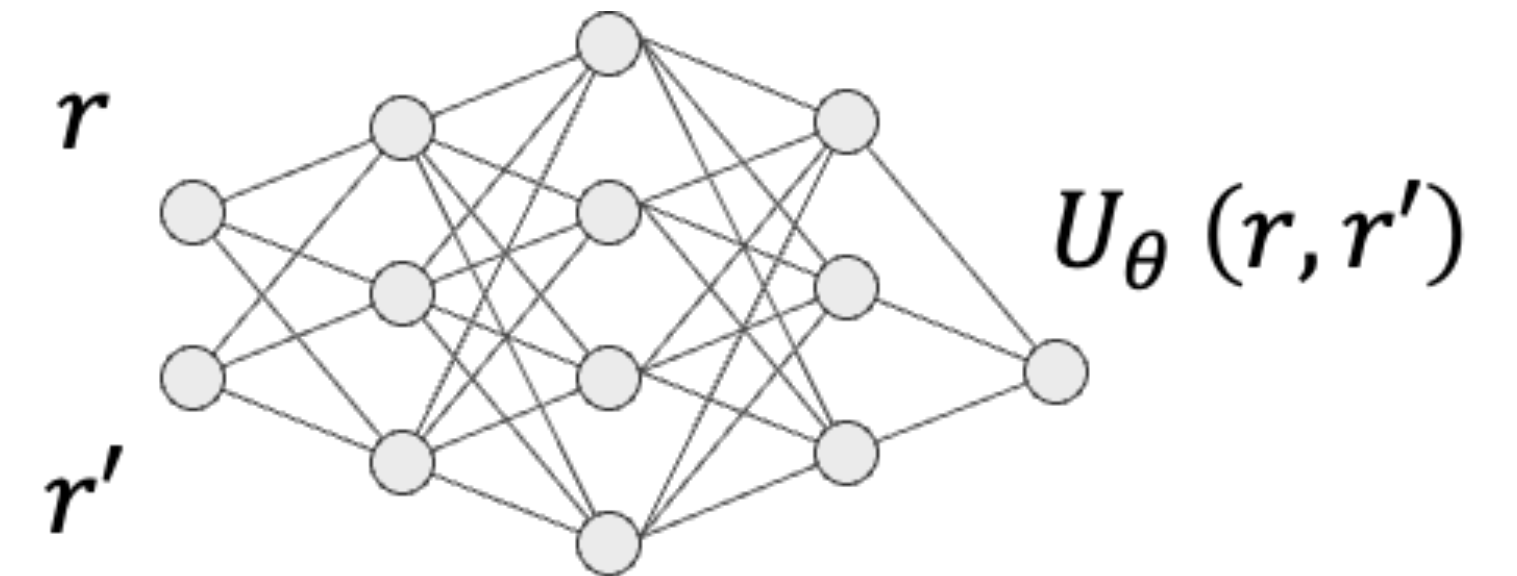
arXiv:2410.03082 (with HAL QCD)



NBS wave function
Data(Observations)

$$\theta_{i+1} \rightarrow \theta_i + \frac{\partial \mathcal{L}}{\partial U_\theta(r, r')} \frac{\partial U_\theta(r, r')}{\partial \theta}$$

Gradient Decent



Universal Approximation Theorem (1989,1991)

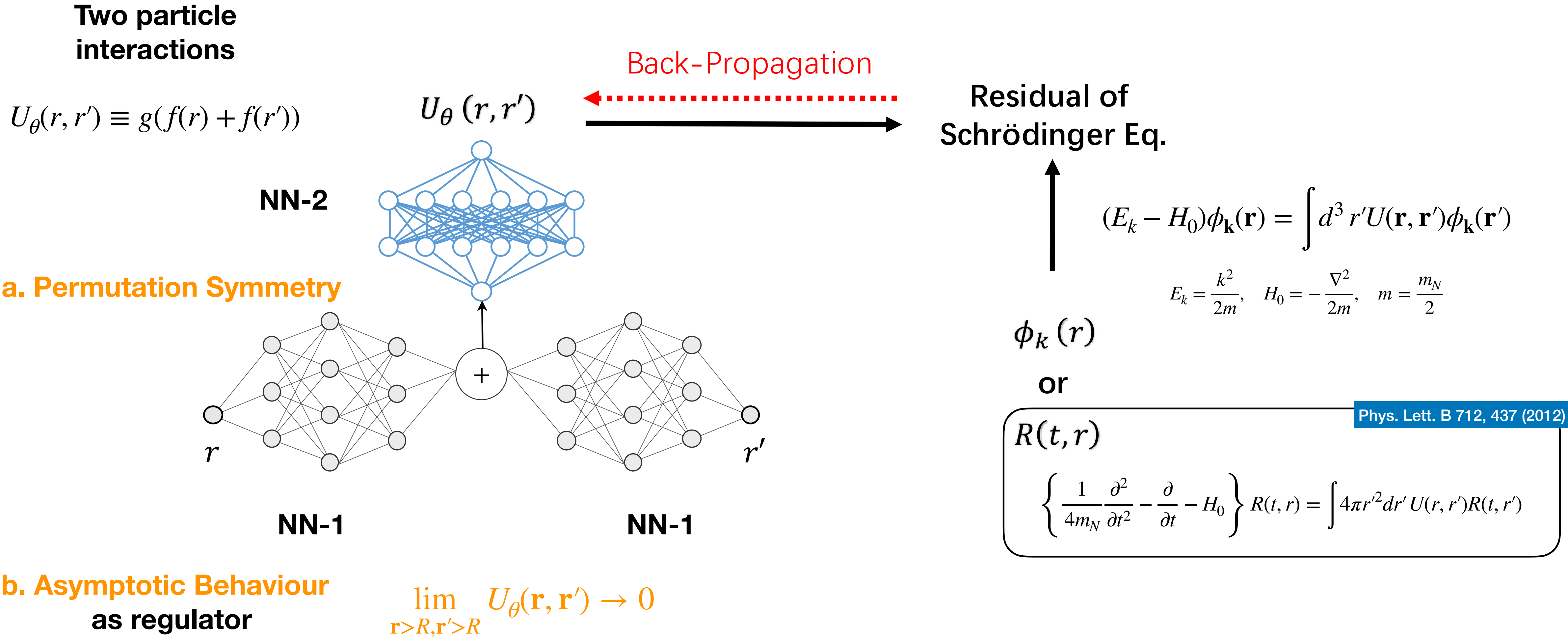
Potential Function
Physics Properties

Maximize Likelihood Estimation

$$\min_{\theta} \mathcal{L} = \sum_k \int d^3 \mathbf{r} \left[(E_k - H_0) \phi_k(\mathbf{r}) - \int d^3 \mathbf{r}' U_\theta(\mathbf{r}, \mathbf{r}') \phi_k(\mathbf{r}') \right]^2$$

Physics-Driven Deep Learning

arXiv:2410.03082 (with HAL QCD)



Mock Test: Separable Potential

As a numerical example, we take $\mu = 1.0, \omega = -0.017\mu^4, m = 3.30\mu, R = 2.5/\mu$

arXiv:2410.03082 (with HAL QCD)

$$U(\mathbf{r}, \mathbf{r}') \equiv \omega \nu(\mathbf{r}) \nu(\mathbf{r}'), \quad \nu(\mathbf{r}) \equiv e^{-\mu r}$$

$$U_{\text{NN}}(r, r') = \omega f_{\theta}(r, r')$$

$$\phi_k^0(r) = \frac{e^{i\delta_0(k)}}{kr} \left[\sin\{kr + \delta_0(k)\} - \sin\delta_0(k)e^{-\mu r} \left(1 + \frac{r(\mu^2 + k^2)}{2\mu} \right) \right]$$

$$k \cot \delta_0(k) = -\frac{1}{4\mu^2} \left[2\mu(\mu^2 - k^2) - \frac{3\mu^2 + k^2}{4\mu^3} (\mu^2 + k^2)^2 + \frac{(\mu^2 + k^2)^4}{8\pi m \omega} \right]$$

Nambu-Bethe-Salpeter (NBS) wave function

$$\phi_{\mathbf{k}}(\mathbf{r}) e^{-W_{\mathbf{k}} t} \equiv \langle 0 | N(\mathbf{x} + \mathbf{r}, t) N(\mathbf{x}, t) | NN, W_{\mathbf{k}} \rangle$$

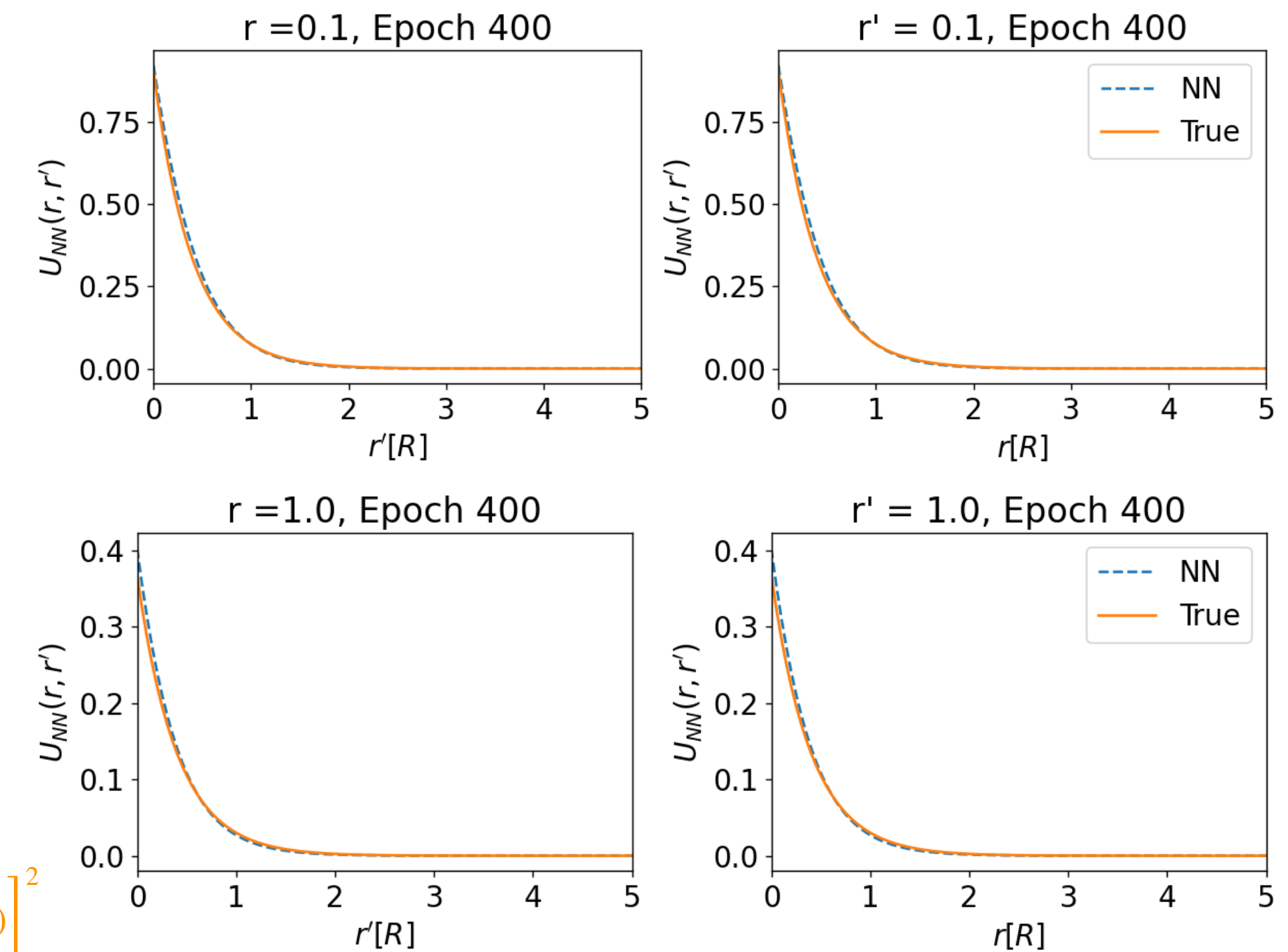
$$(E_k - H_0) \phi_{\mathbf{k}}(\mathbf{r}) = \int d^3 r' U(\mathbf{r}, \mathbf{r}') \phi_{\mathbf{k}}(\mathbf{r}')$$

$$E_k = \frac{k^2}{2m}, \quad H_0 = -\frac{\nabla^2}{2m}, \quad m = \frac{m_N}{2}$$

BP

$$\mathcal{L} = \sum_{\mathbf{k}} \int d^3 r \left[(E_k - H_0) \phi_{\mathbf{k}}(\mathbf{r}) - \int d^3 r' U_{\theta}(\mathbf{r}, \mathbf{r}') \phi_{\mathbf{k}}(\mathbf{r}') \right]^2$$

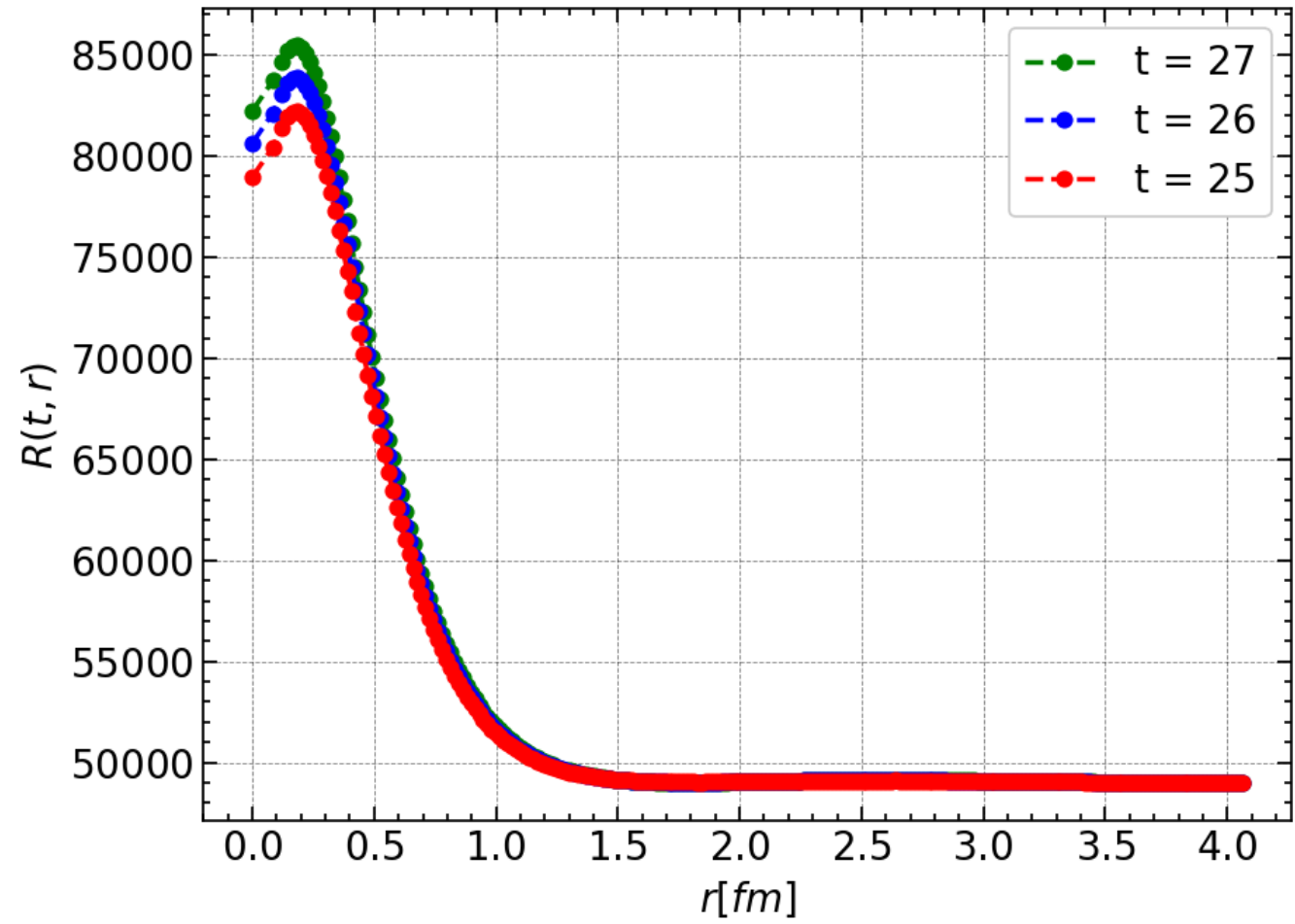
$$U(r > 3R, r' > 3R) \rightarrow 0$$



$\Omega_{ccc}\Omega_{ccc}$ Interaction: 1S_0

Y. Lyu, etc., Phys. Rev. Lett. 127, 072003 (2021)

$$m_N = 2.073, a^{-1} = 2333.0 \text{ MeV}$$



$$\left\{ \frac{1}{4m_N} \frac{\partial^2}{\partial t^2} - \frac{\partial}{\partial t} - H_0 \right\} R(t, r) = \int 4\pi r'^2 dr' U(r, r') R(t, r')$$

$$R2 = R_{t+1} - 2R_t + R_{t-1}, R1 = (R_{t+1} - R_{t-1})/2, Rr = \nabla^2 R(t, r)$$

Nambu-Bethe-Salpeter (NBS) wave function →

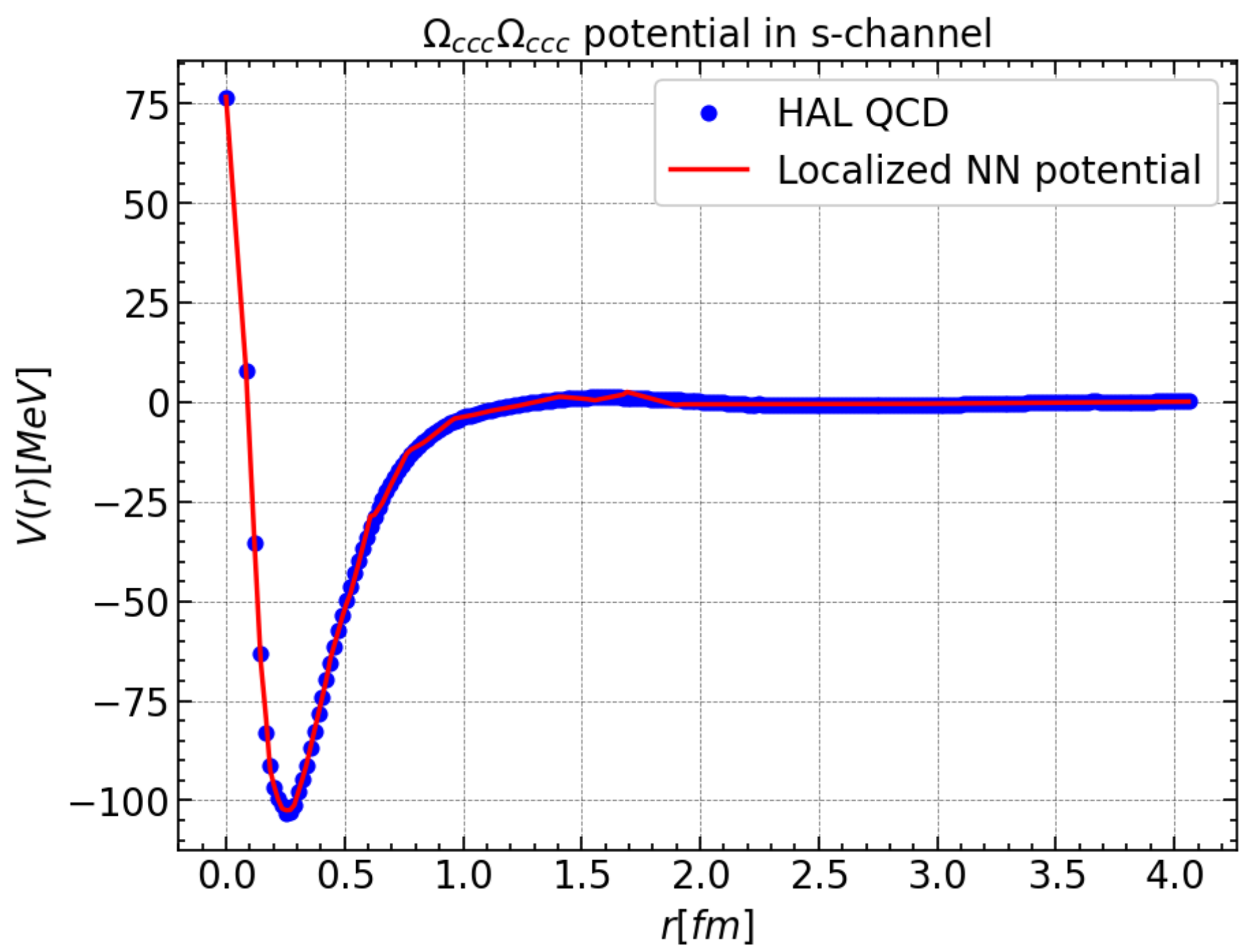
BP

$$\mathcal{L} = \sum_t \left\{ \frac{1}{4m_N} R2(t, r) - R1(t, r) + \frac{1}{m_N} Rr(t, r) - \int 4\pi r'^2 dr' U_\theta(r, r') R(t, r') \right\}$$

$U(r > 3\text{fm}, r' > 3\text{fm}) \rightarrow 0$

arXiv:2410.03082 (with HAL QCD)

Neural Network Hadron Force

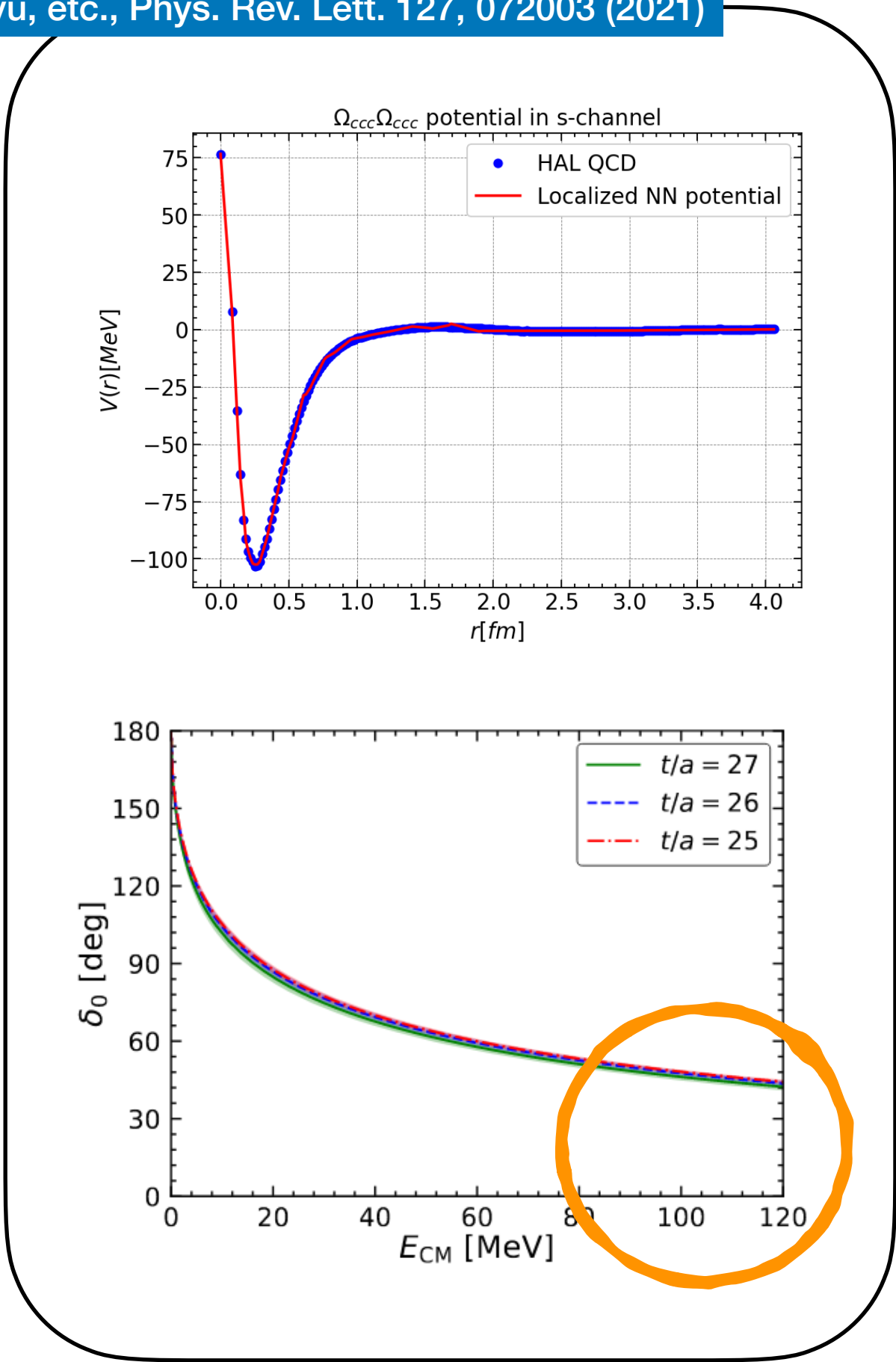


$$V_\theta(r) \equiv \frac{\sum_{r'} \Delta r' 4\pi r'^2 U_\theta(r, r') R(t, r')}{R(t, r)}$$

$\Omega_{ccc}\Omega_{ccc}$ Interaction: Non-Local Potential

arXiv:2410.03082 (with HAL QCD)

Y. Lyu, etc., Phys. Rev. Lett. 127, 072003 (2021)



“3D Map”

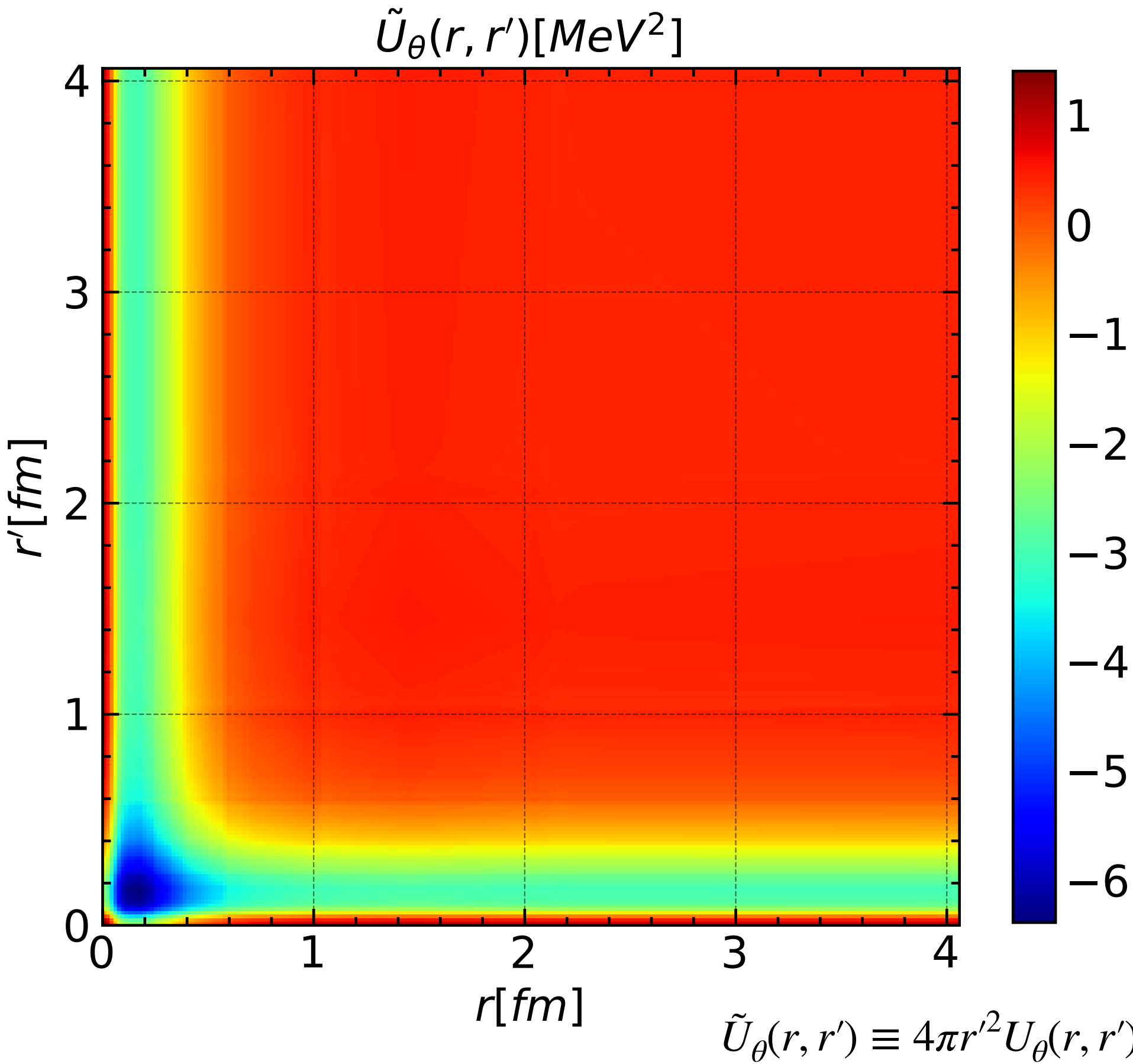
$$(E_k - H_0)\phi_{\mathbf{k}}(\mathbf{r}) = \int d^3 r' U_{\theta}(\mathbf{r}, \mathbf{r}') \phi_{\mathbf{k}}(\mathbf{r}')$$

$$E_k = \frac{k^2}{2m}, \quad H_0 = -\frac{\nabla^2}{2m}, \quad m = \frac{m_N}{2}$$



Affect ?

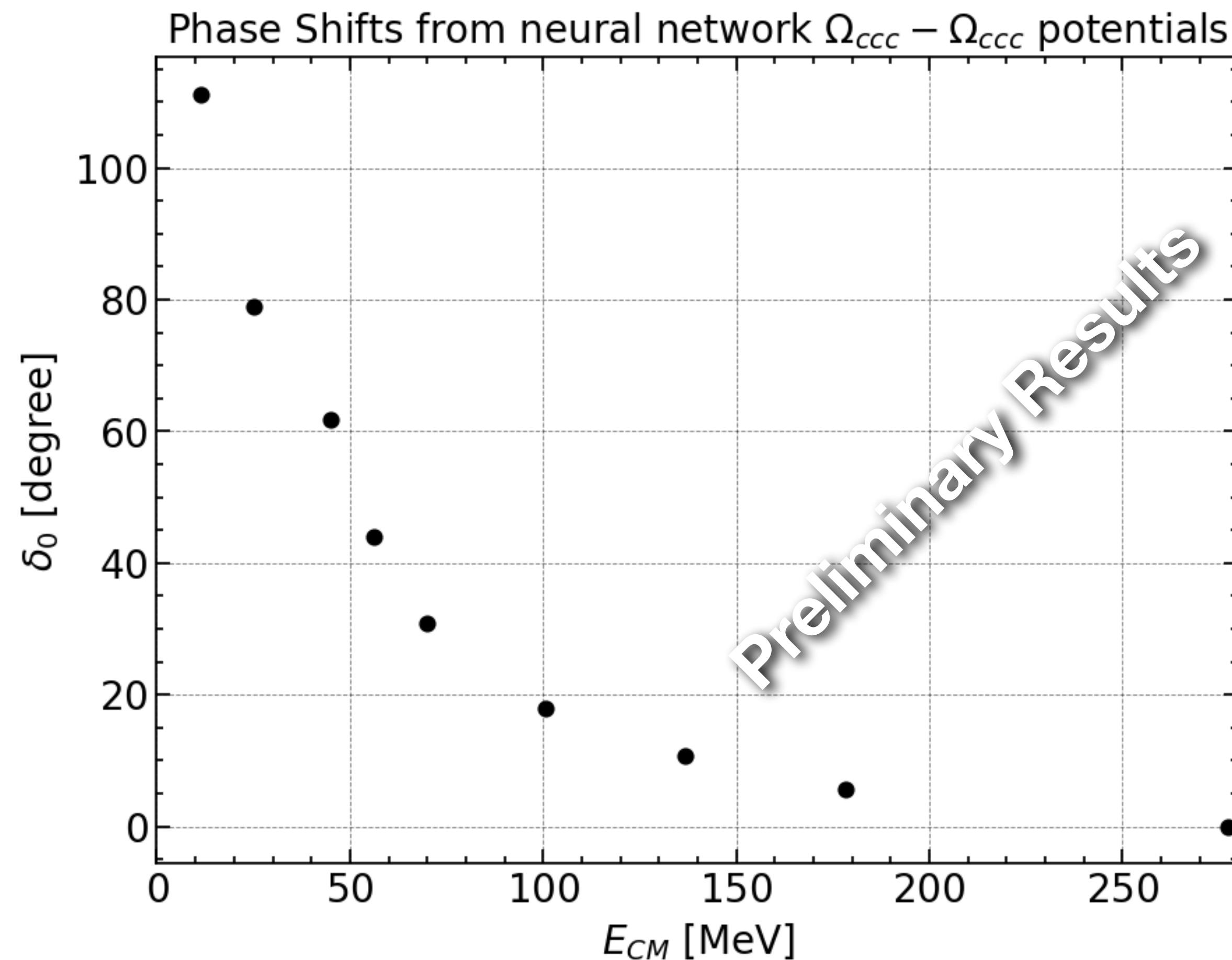
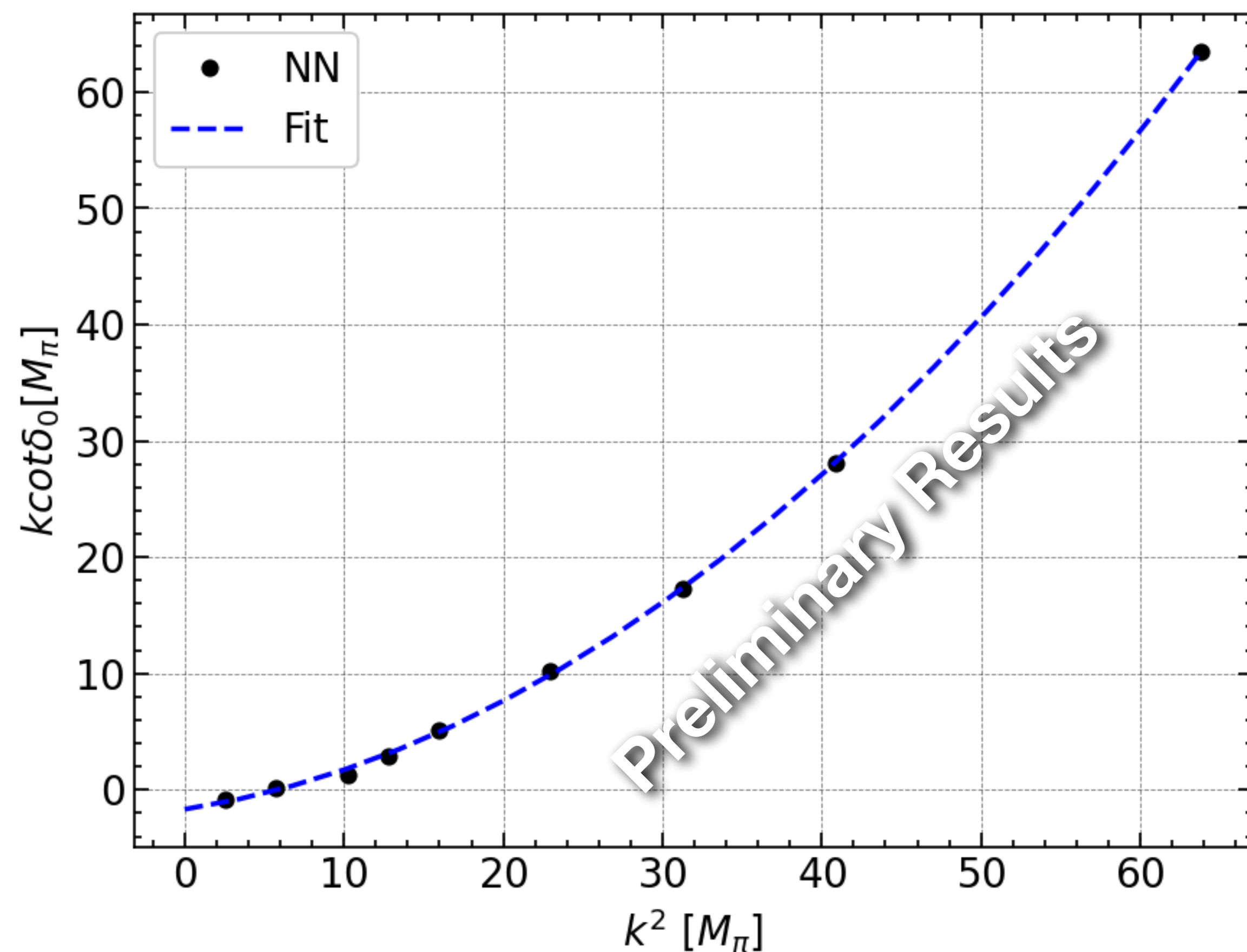
Neural Network Hadron Force



Scattering Behaviors

in Preparation

$$a_0 = 0.86 \text{ fm}, r_{\text{eff}} = 0.57 \text{ fm}$$



Summary I

- **Physics-Driven Desgins**
 - Symmetry
 - Physics Principles
 - Physical Data
 - Physical Equations
- **Inverse Problems**
 - Data-Driven Learning
 - Limited data set
 - Physics-Driven Deep Learning
 - Neural network representations
 - Embed physics priors explicitly
 - **Exchange Symmetry** and **Asymptotic Behavior** for HH interactions

Physics-Driven Learning for Solving Inverse Problems towards QCD Physics

Nat. Rev. Phys.

Gert Aarts¹, Kenji Fukushima², Tetsuo Hatsuda³, Andreas Ipp⁴, Shuzhe Shi⁵, Lingxiao Wang^{3,*}, and Kai Zhou^{6,7}

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⁴Institute for Theoretical Physics, TU Wien, Wiedner Hauptstraße 8-10/136, A-1040 Vienna, Austria

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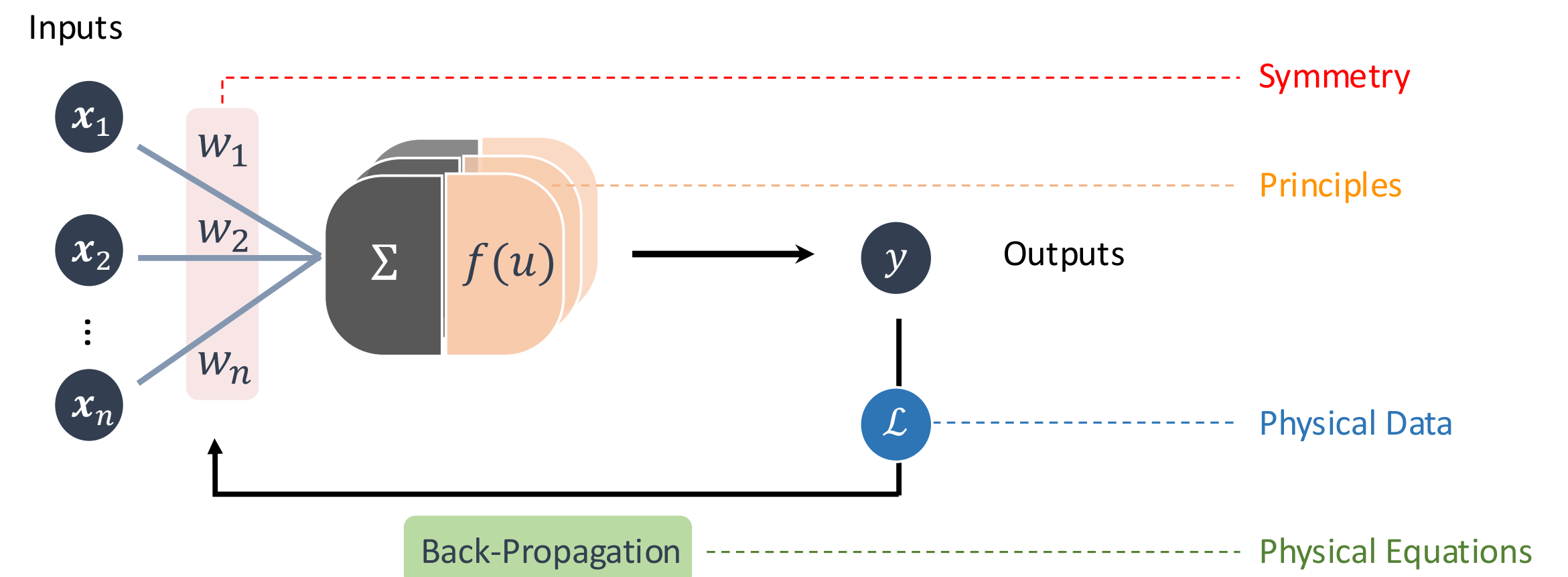
⁶School of Science and Engineering, The Chinese University of Hong Kong, Shenzhen (CUHK-Shenzhen), Guangdong, 518172, China

⁷Frankfurt Institute for Advanced Studies, Ruth Moufang Strasse 1, D-60438, Frankfurt am Main, Germany

*e-mail: lingxiao.wang@riken.jp

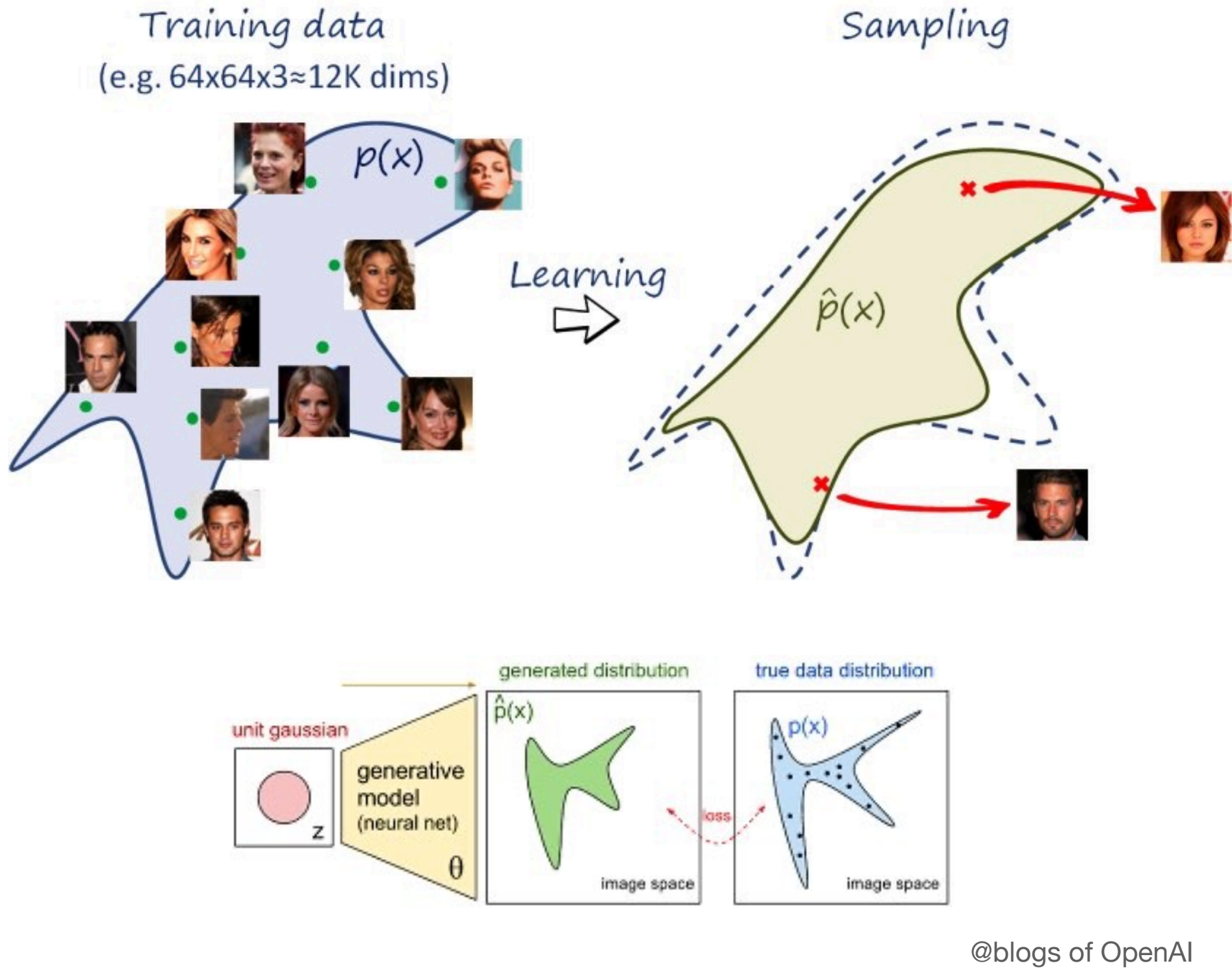
ABSTRACT

The integration of deep learning techniques and physics-driven designs is reforming the way we address inverse problems, in which accurate physical properties are extracted from complex data sets. This is particularly relevant for quantum chromodynamics (QCD), the theory of strong interactions, with its inherent limitations in observational data and demanding computational approaches. This perspective highlights advances and potential of physics-driven learning methods, focusing on predictions of physical quantities towards QCD physics, and drawing connections to machine learning (ML). It is shown that the fusion of ML and physics can lead to more efficient and reliable problem-solving strategies. Key ideas of ML, methodology of embedding physics priors, and generative models as inverse modelling of physical probability distributions are introduced. Specific applications cover first-principle lattice calculations, and QCD physics of hadrons, neutron stars, and heavy-ion collisions. These examples provide a structured and concise overview of how incorporating prior knowledge such as symmetry, continuity and equations into deep learning designs can address diverse inverse problems across different physical sciences.



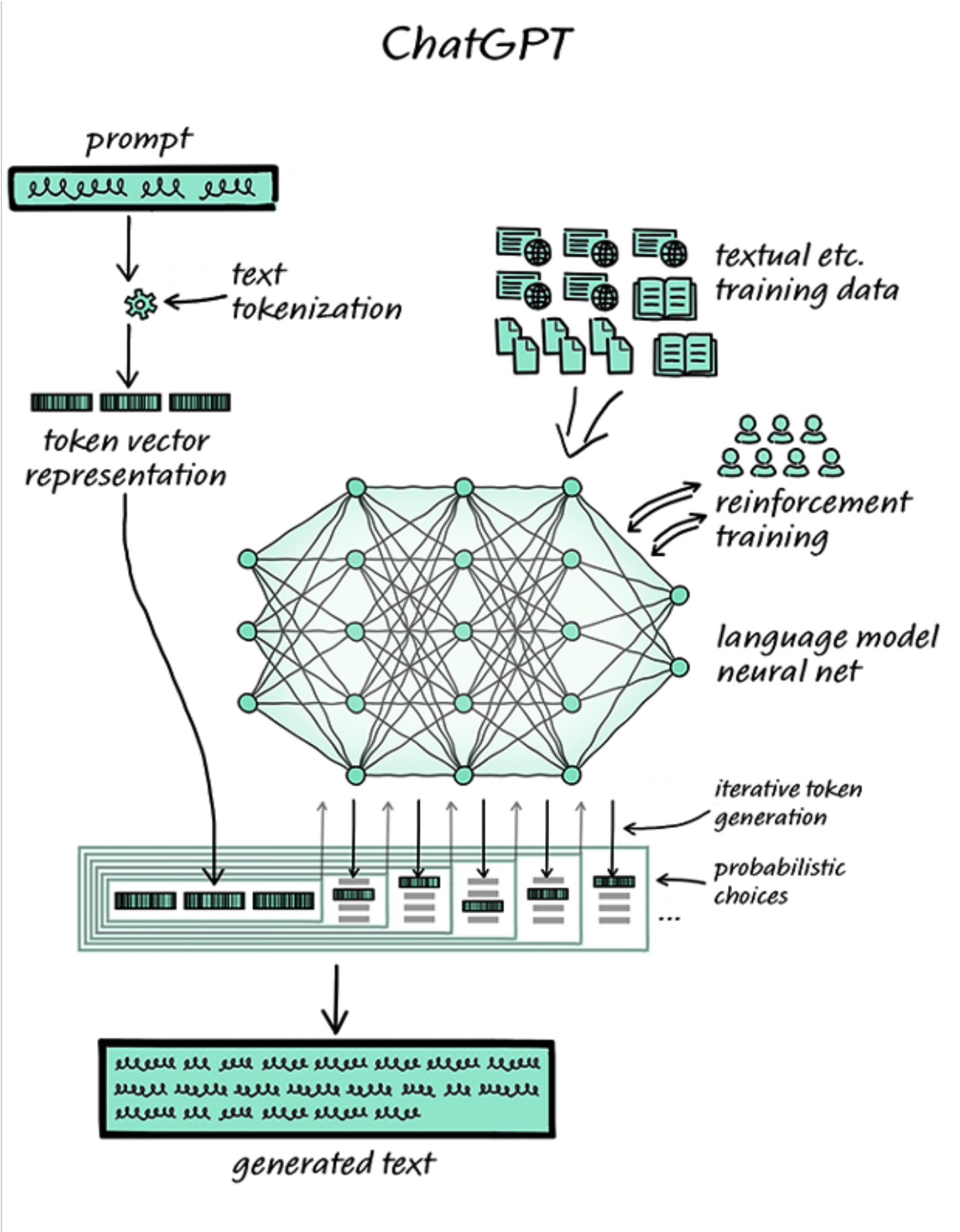
Generative Models as Inverse Modeling

Generative Models



Generative models
→ **Underlying Distributions** in Data

$$\max_{\theta} \prod_{i=1}^N p_{\theta}(\mathbf{x}_i)$$



High-Dimensional Distribution

$$p(\phi) = e^{-S(\phi)}/Z$$

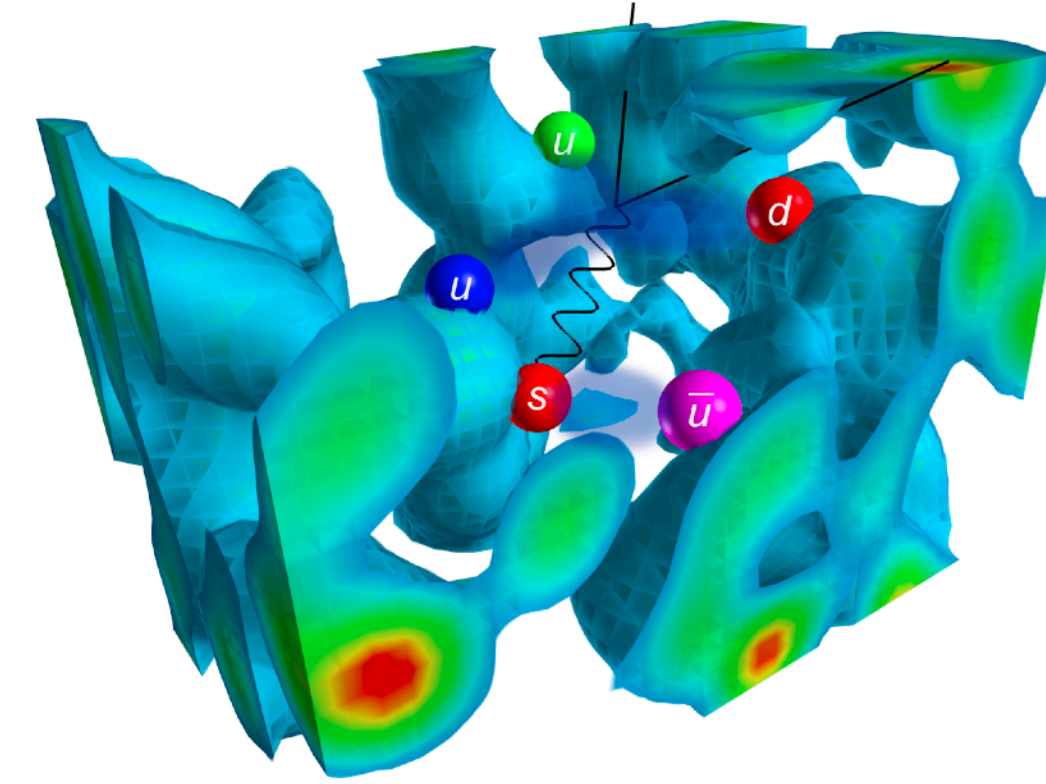
$$\langle O \rangle \approx \frac{1}{N} \sum_i O_i$$

→ **Physical Distribution, Sampling**
via Generative Models

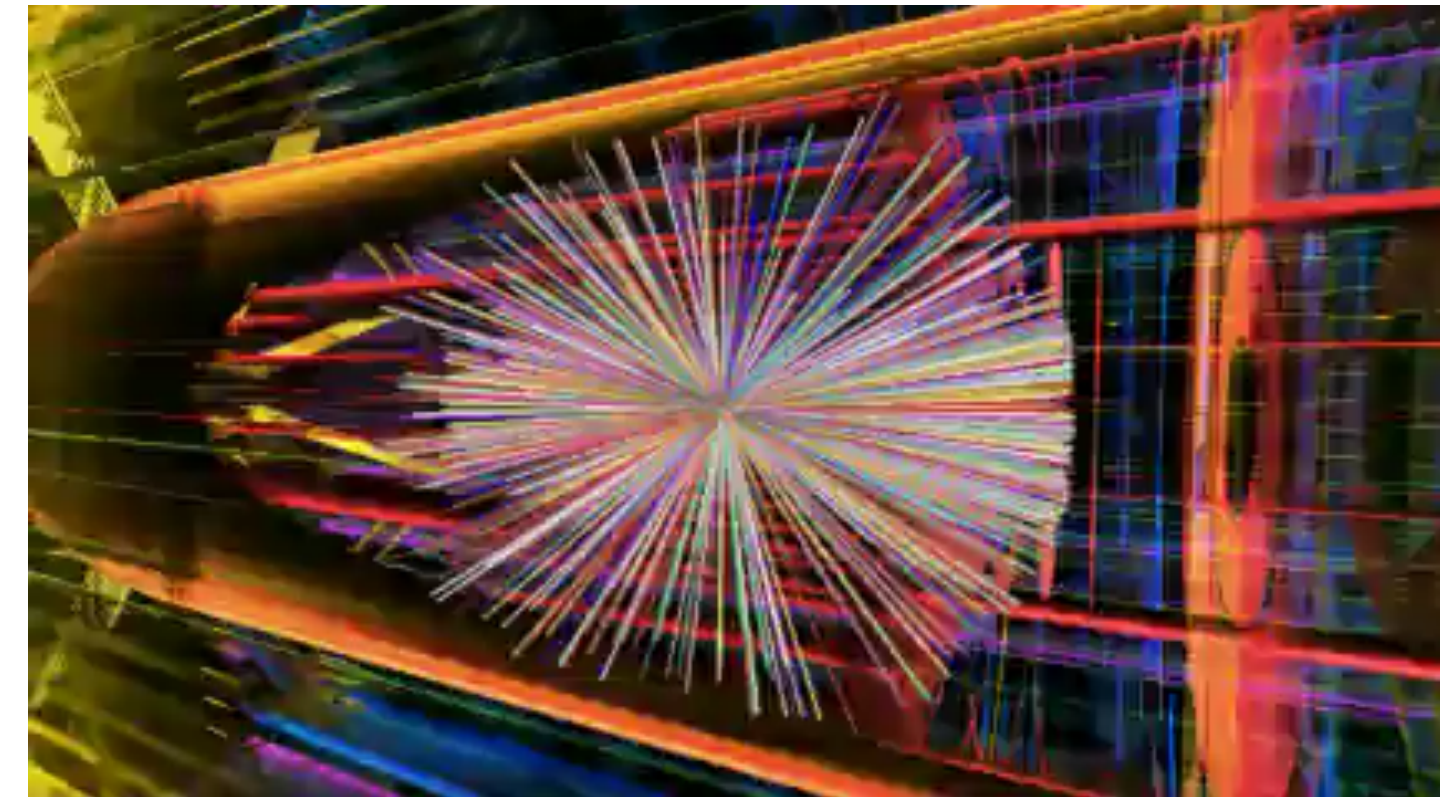
Global Sampling

Fast and Independent Sampler

Prog.Part.Nucl.Phys. 104084(2023)



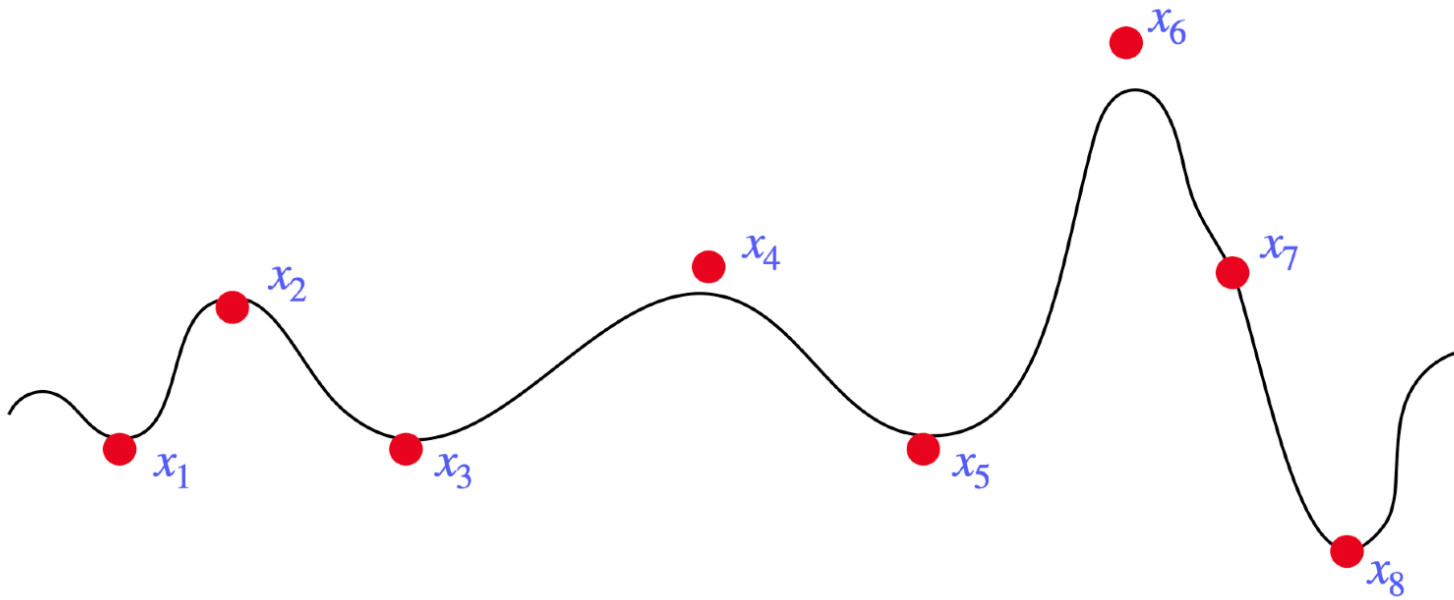
Lattice QCD © Derek Leinweber/CSSM/University of Adelaide



Heavy-Ion Collisions © 2010 CERN

Probabilistic Models

$$\mathbf{x} \sim p_{data}(\mathbf{x})$$

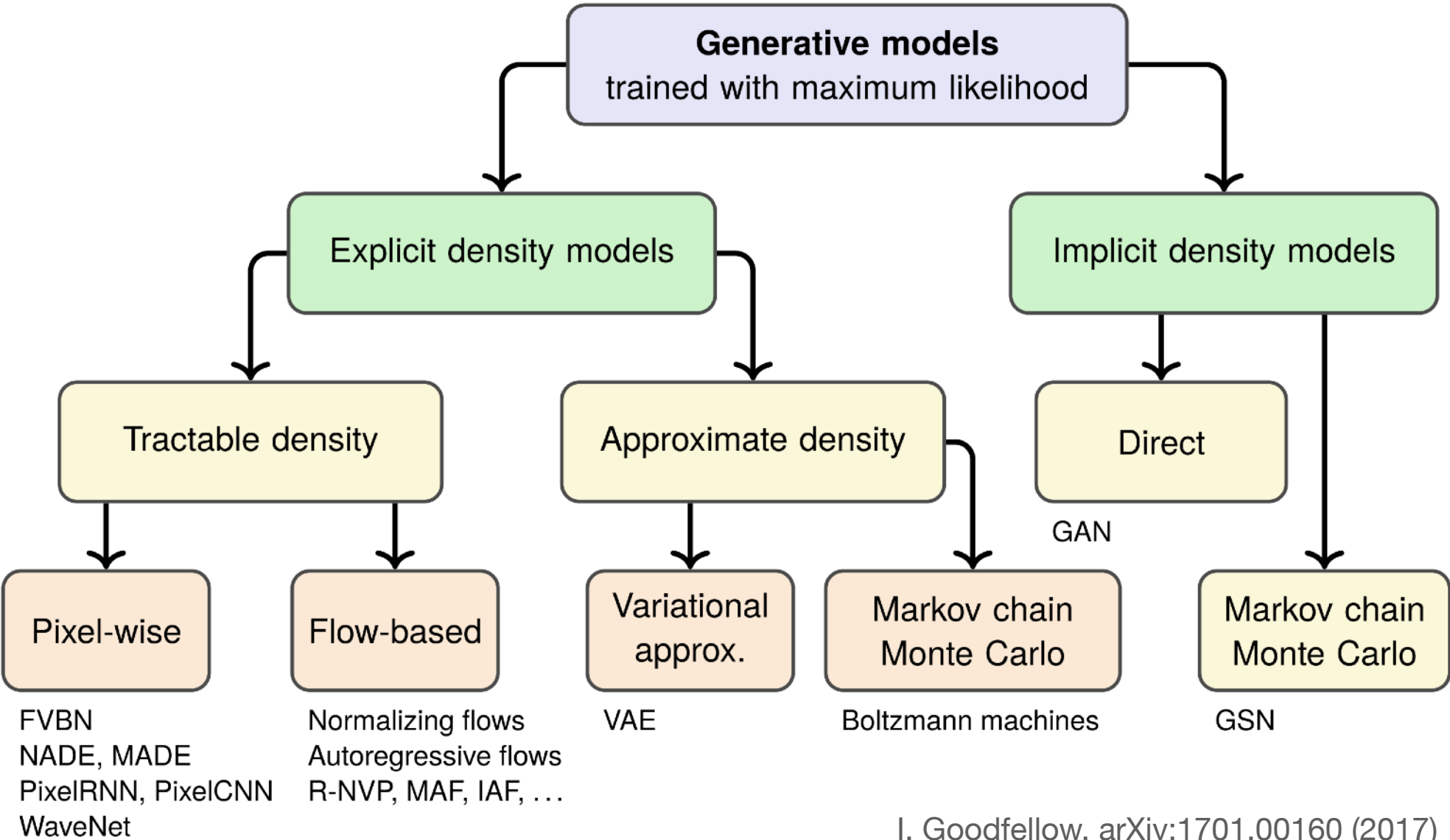


$$\max_{\theta} \prod_{i=1}^N p_{\theta}(\mathbf{x}_i)$$

Maximum likelihood estimation(MLE)

$$\max_{\theta} \sum_{i=1}^N \log p_{\theta}(\mathbf{x}_i)$$

Maximum log-likelihood estimation



Bishop, C. M. & Bishop, H. *Deep Learning: Foundations and Concepts*. doi:[10.1007/978-3-031-45468-4](https://doi.org/10.1007/978-3-031-45468-4).

I. Goodfellow, arXiv:1701.00160 (2017)

Learn to Sample

$$p(\phi) = e^{-S(\phi)}/Z$$

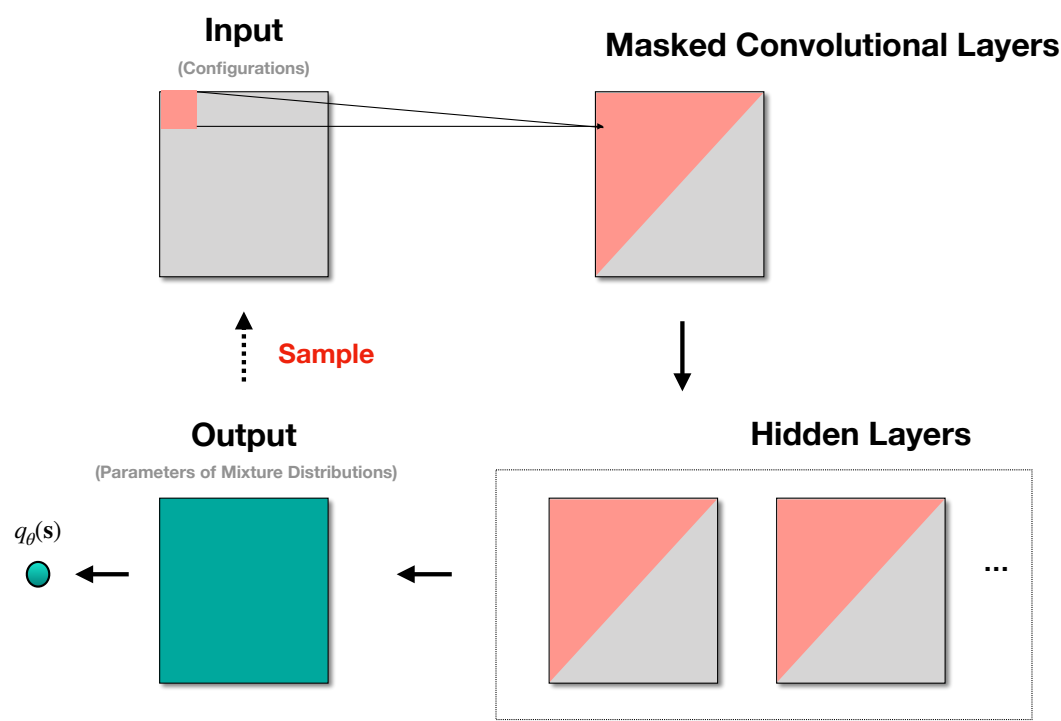
$$\langle O \rangle \approx \frac{1}{N} \sum_i O_i$$

→ **Physical Distribution, Sampling**
via Generative Models

Global Sampling

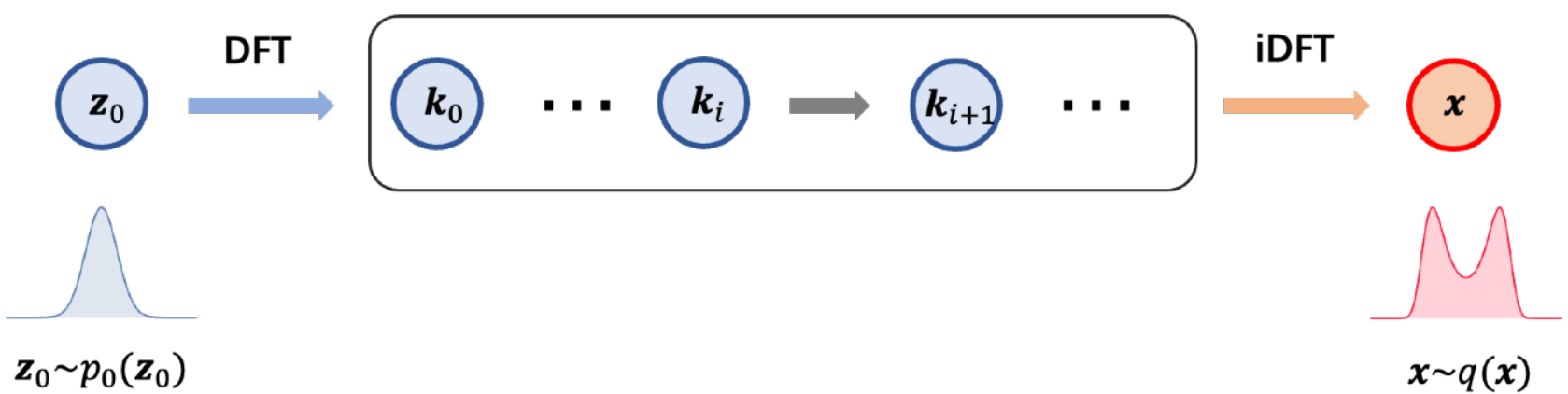
Fast and Independent Sampler

Prog.Part.Nucl.Phys. 104084(2023)



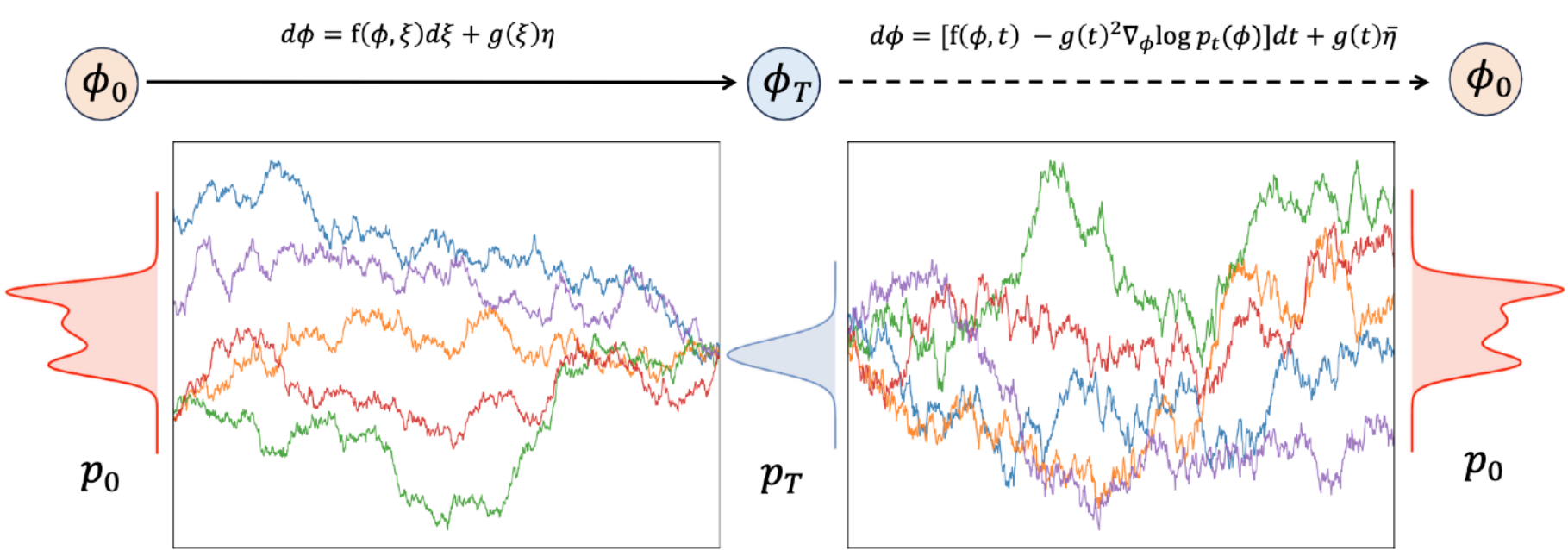
Continuous Autoregressive Models

Chinese Phys. Lett. 39, 120502 (2022)



Fourier Flow-based Model

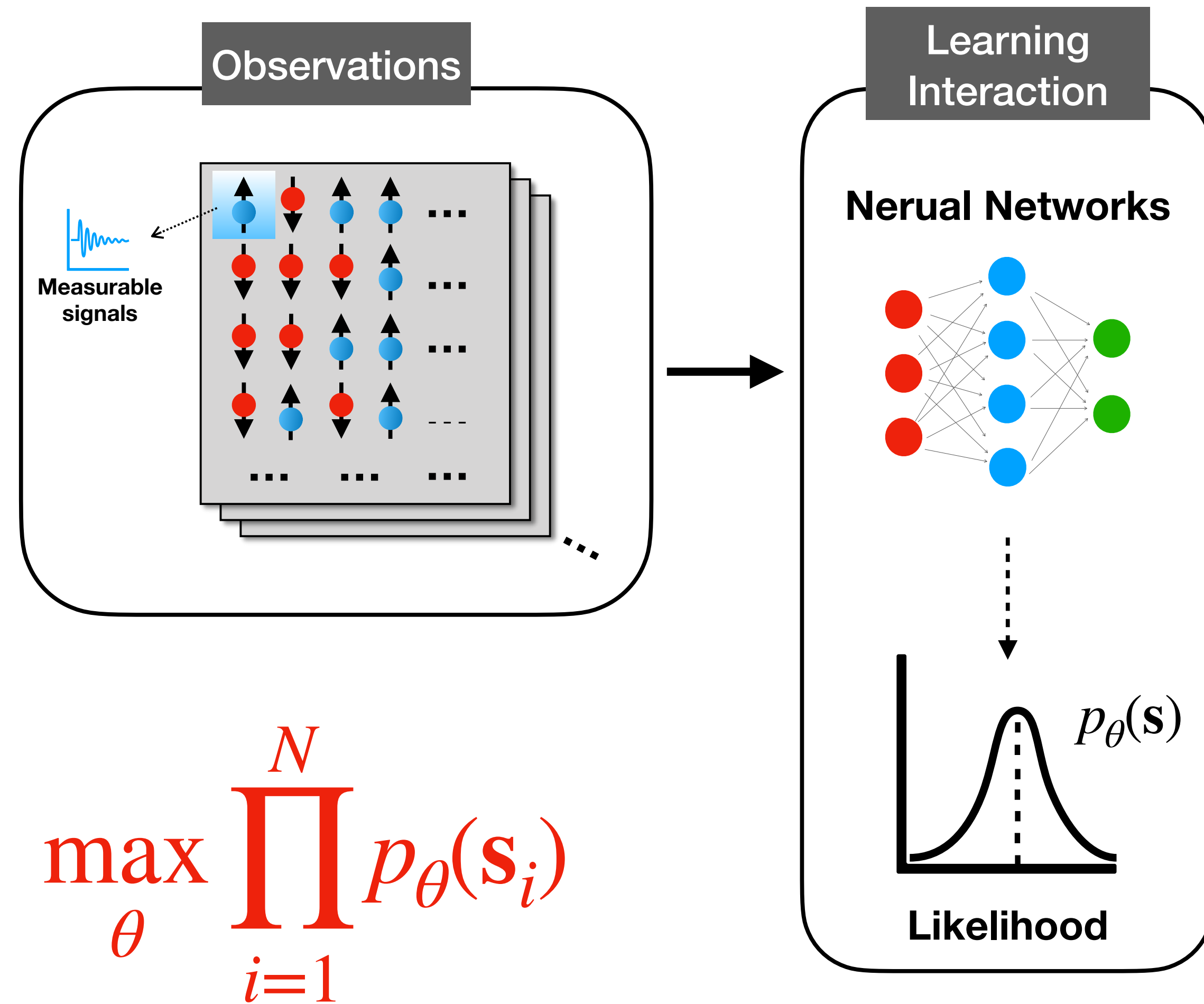
Phys. Rev. D 107, 056001



Diffusion Models

JHEP 05(2024)060

Learn Micro-Interactions from Observations

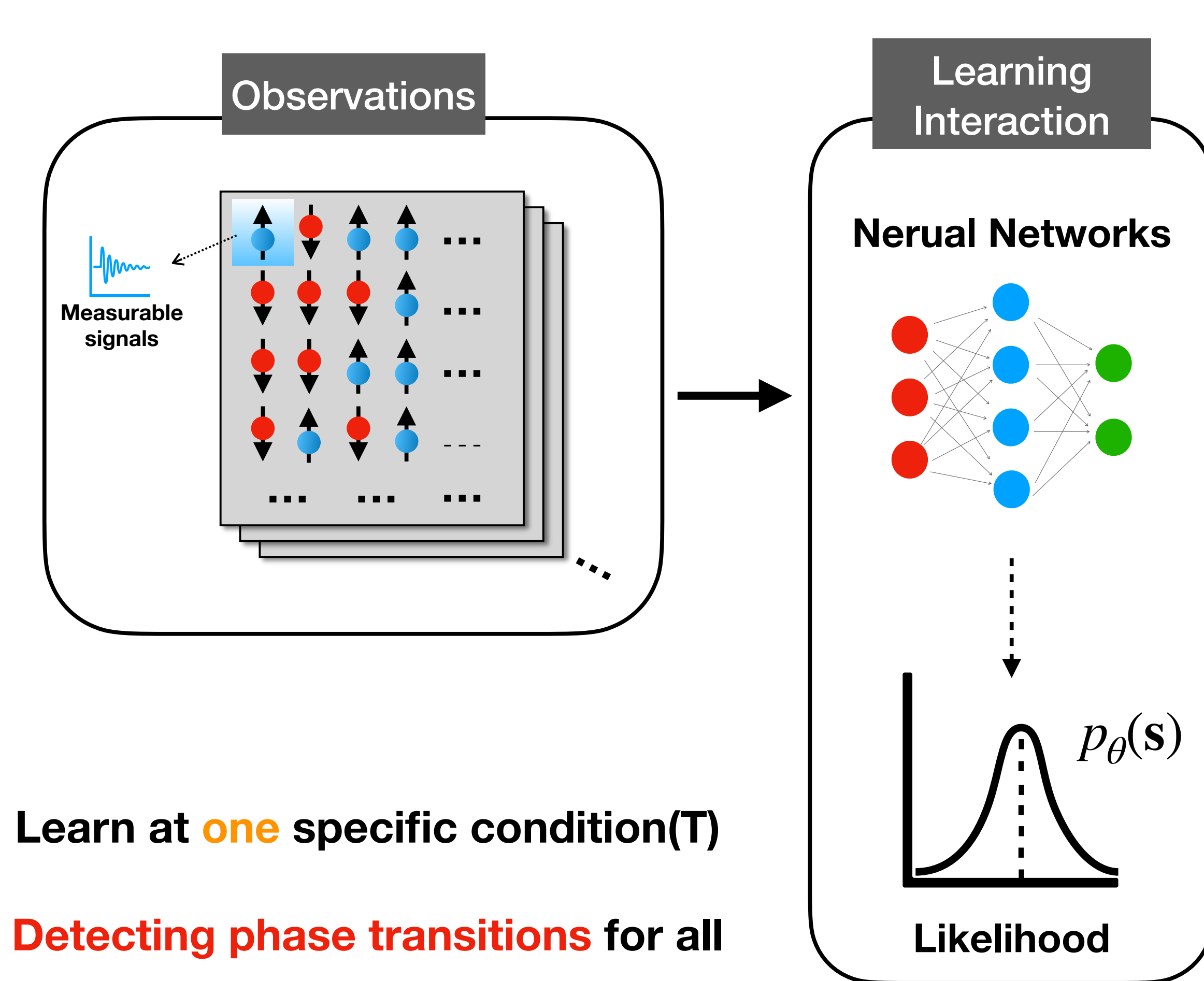


$$p(s | \mathbf{T}) \leftarrow p_{\theta}(s | \mathbf{T})$$

Learn at **one** specific condition(T)

Detecting phase transitions for all ?

Learn Micro-Interactions from Observations



$$\max_{\theta} \prod_{i=1}^N p_{\theta}(s_i)$$

$$p(s) \leftarrow p_{\theta}(s) \equiv \frac{e^{-\frac{H_{\theta}(s)}{T}}}{Z}$$

$$H_{\theta}(s, T) = -T \ln p_{\theta}(s) - T \ln Z$$

$$\frac{\Delta H_{\theta}(s, T)}{T'} \equiv -\frac{T}{T'} (\ln p_{\theta}(s + \delta s) - \ln p_{\theta}(s))$$

Autoregressive Networks

$$p_{\theta}(s) = \prod_{i=1}^N p(s_i | s_1, \dots, s_{i-1})$$

- **Example (L=3) for 1D spin model**

$$p_{\theta}(s) = p(s_3 | s_2, s_1) p(s_2 | s_1) p(s_1)$$

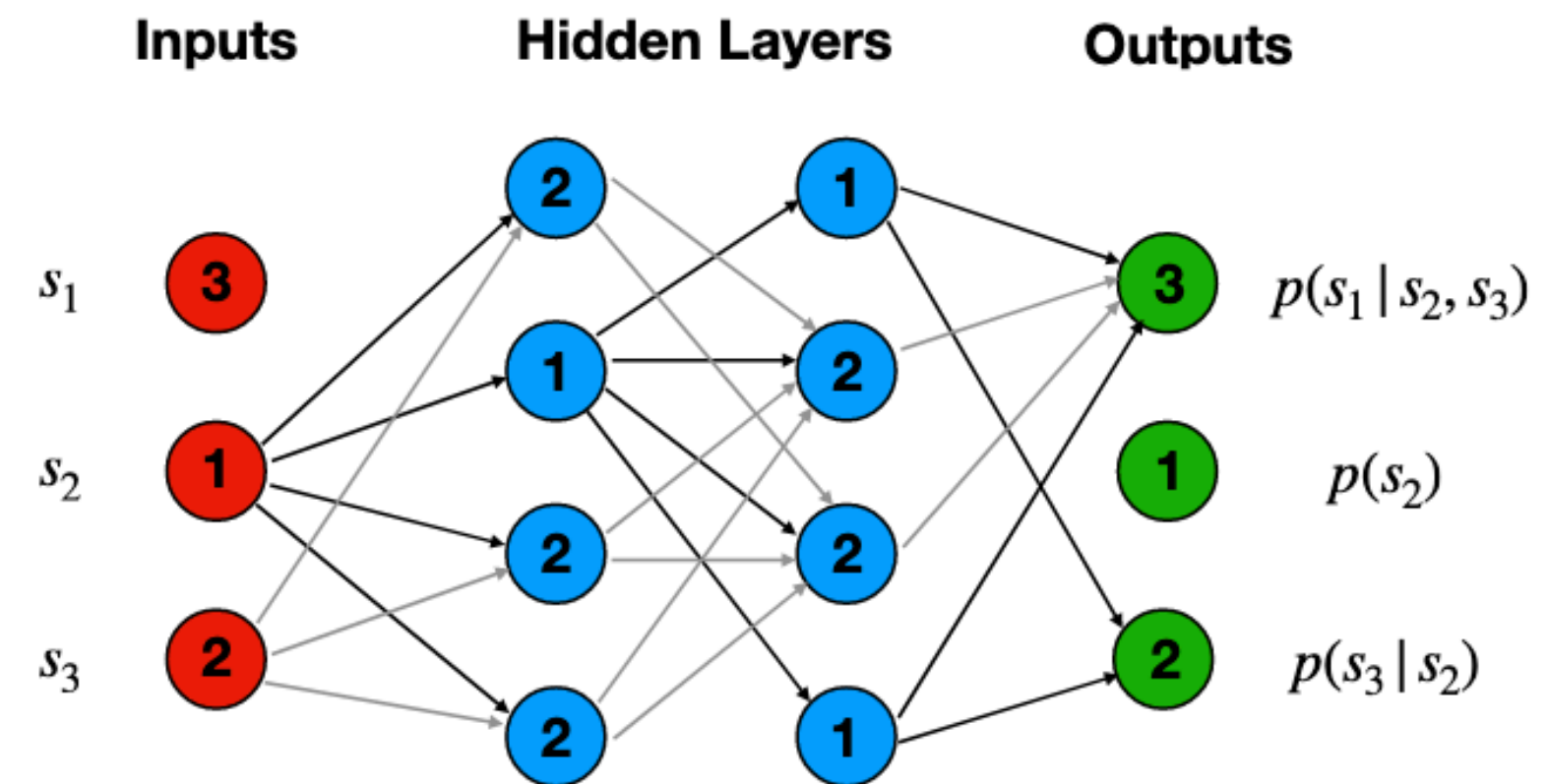
- $p(s_i | s_{<i})$ can be any naive distribution

Bernoulli distribution for Ising model

$$p(s_i | s_{<i}) = q_i \delta_{s_i, +1} + (1 - q_i) \delta_{s_i, -1}$$

$$q_1 = f(s_1 = +1), q_2 = f(s_2 = +1 | s_1), q_3 = f(s_3 = +1 | s_2, s_1)$$

discrete d.o.f.s



Chinese Phys. Lett. 39, 120502 (2022)

Beta distribution for continuous d.o.f., $X \sim \text{Beta}(a, b)$

$$p_{\theta}(s_i | s_1, \dots, s_{i-1}) = \frac{\Gamma(a_i + b_i)}{\Gamma(a_i)\Gamma(b_i)} s_i^{a_i-1} (1 - s_i)^{b_i-1}$$

$\Gamma(a)$ is gamma function, $s_i = \theta_i/2\pi \in [0, 1)$, $(a_i, b_i) > 0$

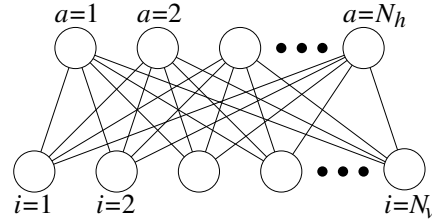
continuous d.o.f.s

Autoregressive Networks

$$p_{\theta}(s) = \prod_{i=1}^N p(s_i | s_1, \dots, s_{i-1})$$

Network is parametrized by a triangular matrix L , which ensures that s_i is independent with s_j when $j \geq i$. This is named as autoregressive property in machine learning.

Gaussian scalar field RBM



- induced distribution on visible layer

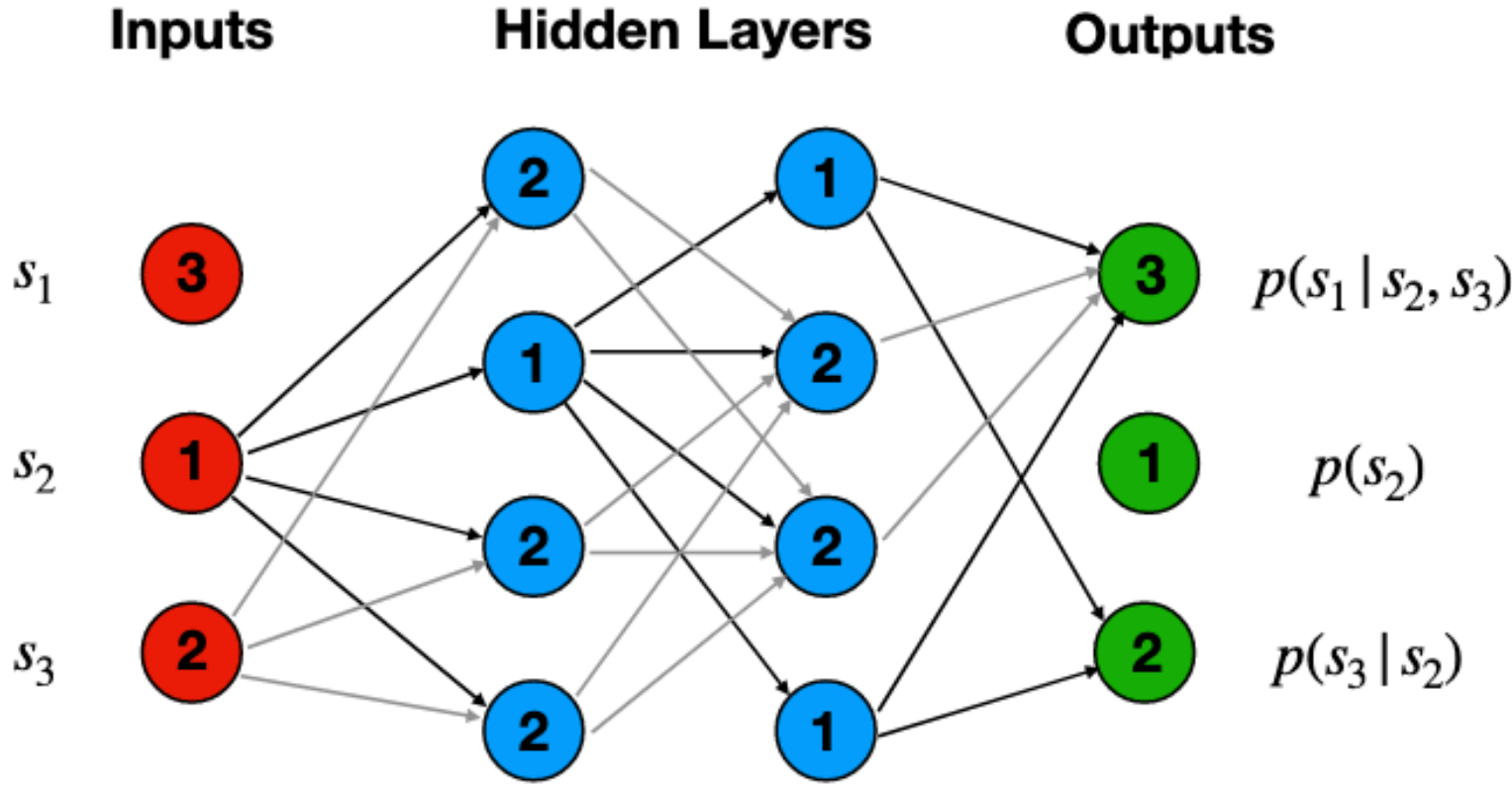
$$p(\phi) = \int Dh p(\phi, h) = \frac{1}{Z} \exp \left(-\frac{1}{2} \sum_{i,j} \phi_i K_{ij} \phi_j + \sum_i J_i \phi_i \right)$$

- scalar field with kinetic (all-to-all) term $K_{ij} = \mu_i^2 \delta_{ij} - \sigma^2 \sum_a w_{ia} w_{aj}^T$

and source $J_i = \sum_a w_{ia} \eta_a$

- unusual Gaussian LFT: what is the weight matrix W and bias η ?

Gert's slides@XQCD2023



$$WW^T = \frac{1}{\sigma^2} (\mu^2 \mathbb{1} - K^\phi) \equiv \mathcal{K}$$

Exact results for $N_h = N_v$

(infinitely) many solutions for weight matrix: $\mathcal{K}$ is symmetric and positive-definite

- Cholesky decomposition $\mathcal{K} = LL^T$: $W = L$ triangular
- diagonalisation $\mathcal{K} = ODO^T = O\sqrt{D}O^T O\sqrt{D}O^T$: $W = W^T = O\sqrt{D}O^T$
- non-uniqueness: internal symmetry $W \rightarrow WO_R \rightarrow \phi^T W h \rightarrow \phi^T WO_R h = \phi^T W h'$

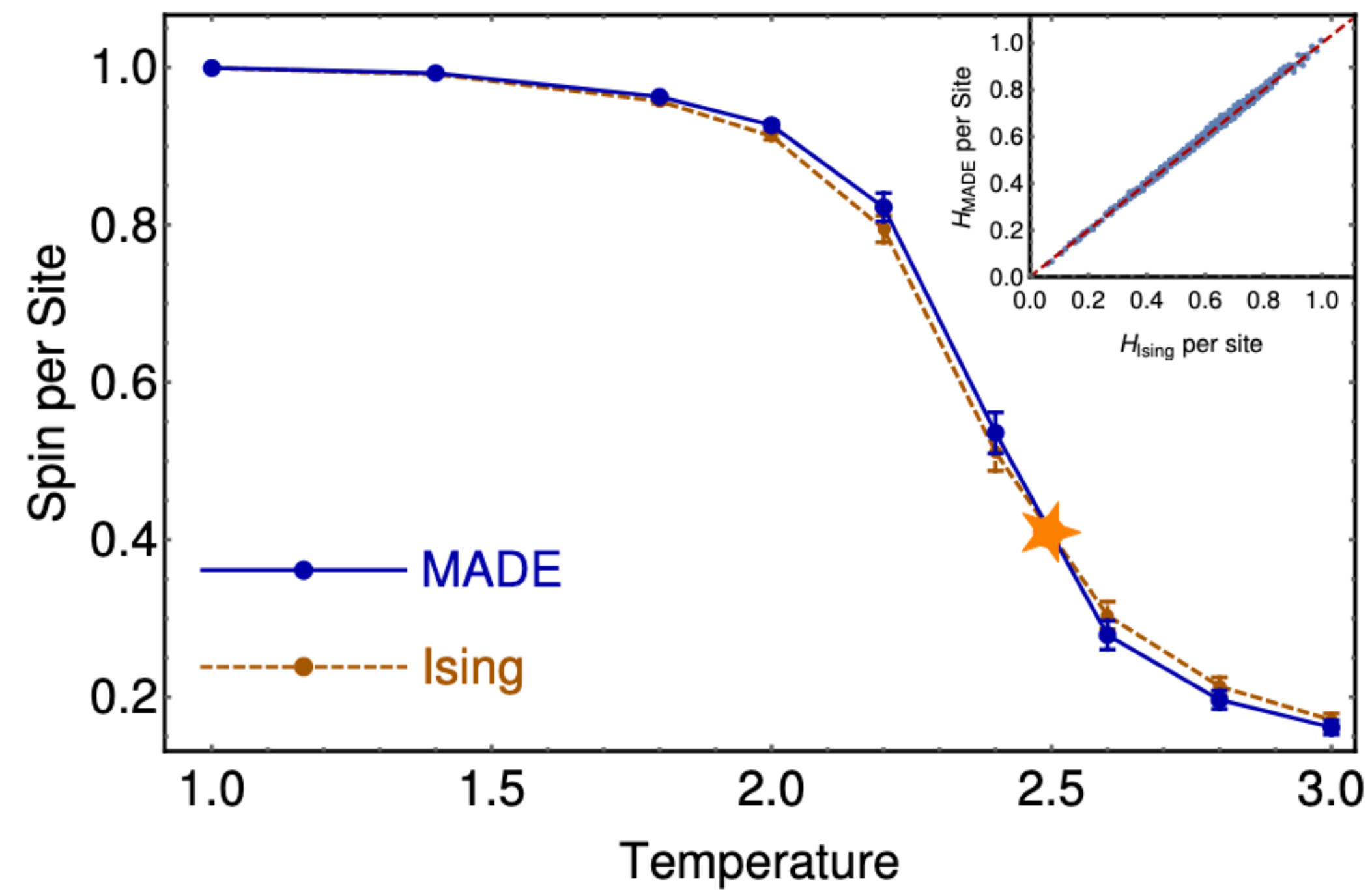
in practice

- all equally valid, realisation depends on initialisation
- non-observable degeneracy due to internal symmetry on hidden layer

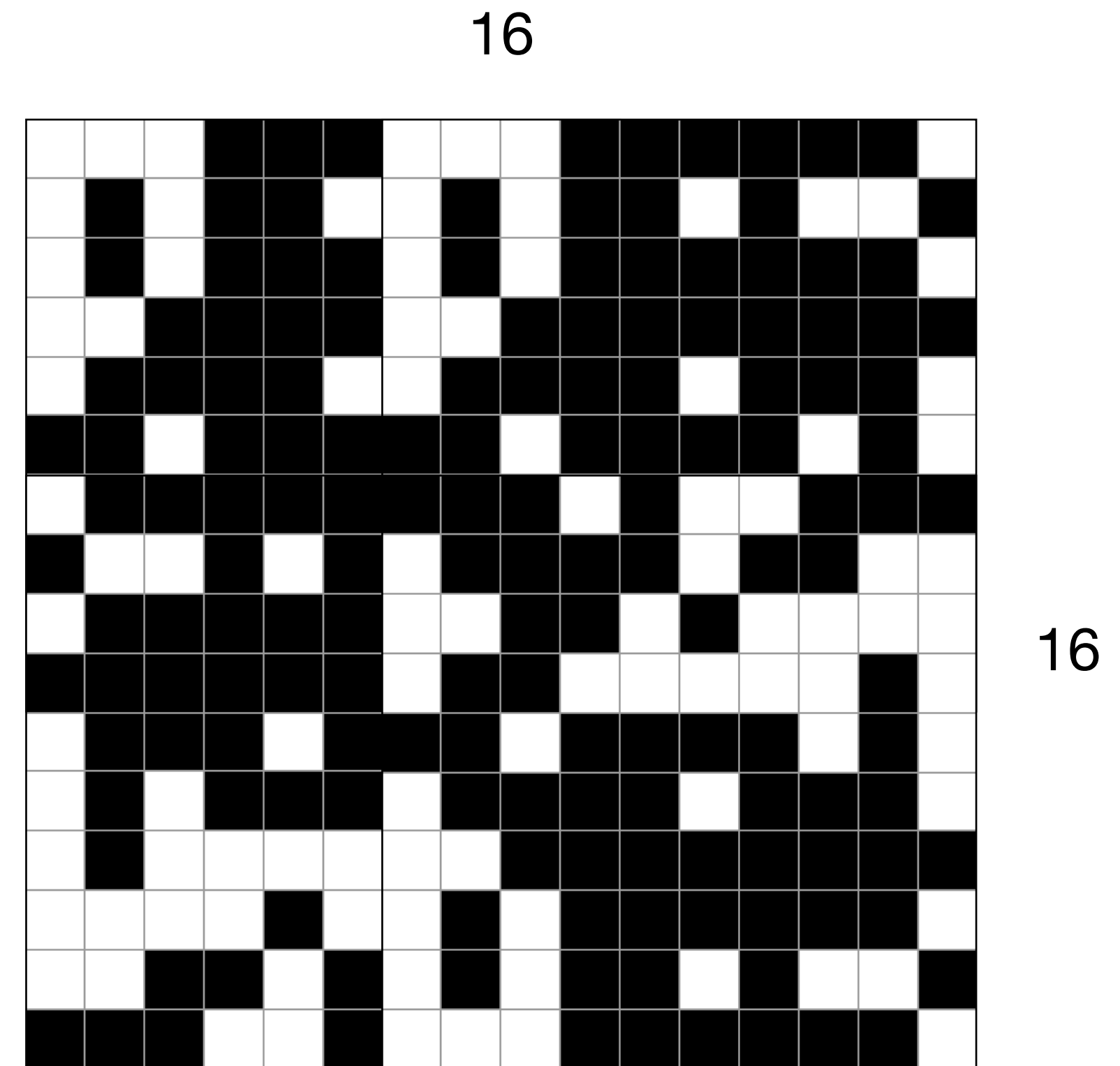
Ferromagnetic Phase Transition

2D Ising Model

arXiv:2007.01037



Masked Autoencoder for Distribution Estimation (**MADE**)

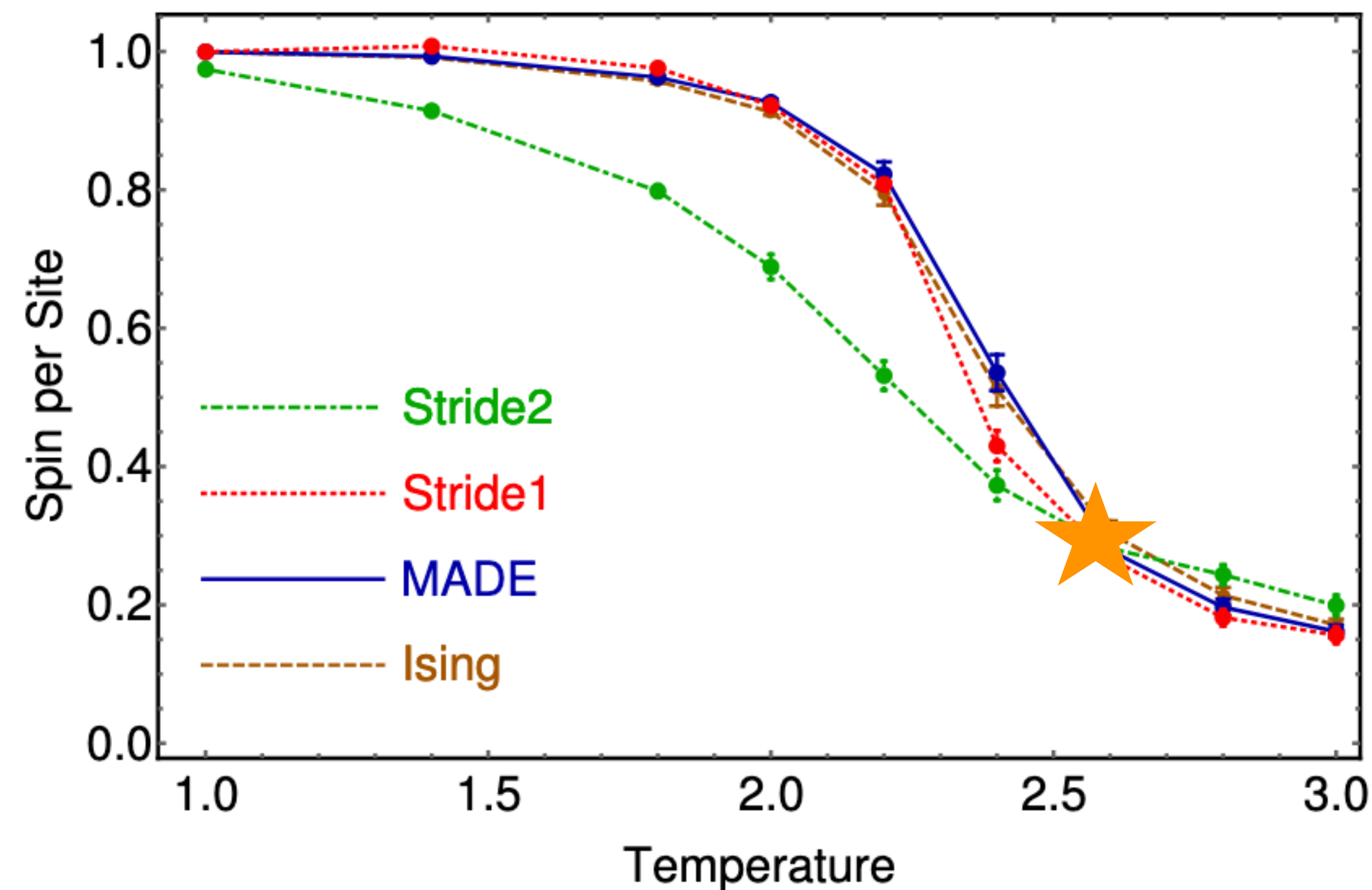


$$H(\mathbf{s}) = - \sum_{\langle i,j \rangle} s_i s_j$$

Ferromagnetic Phase Transition

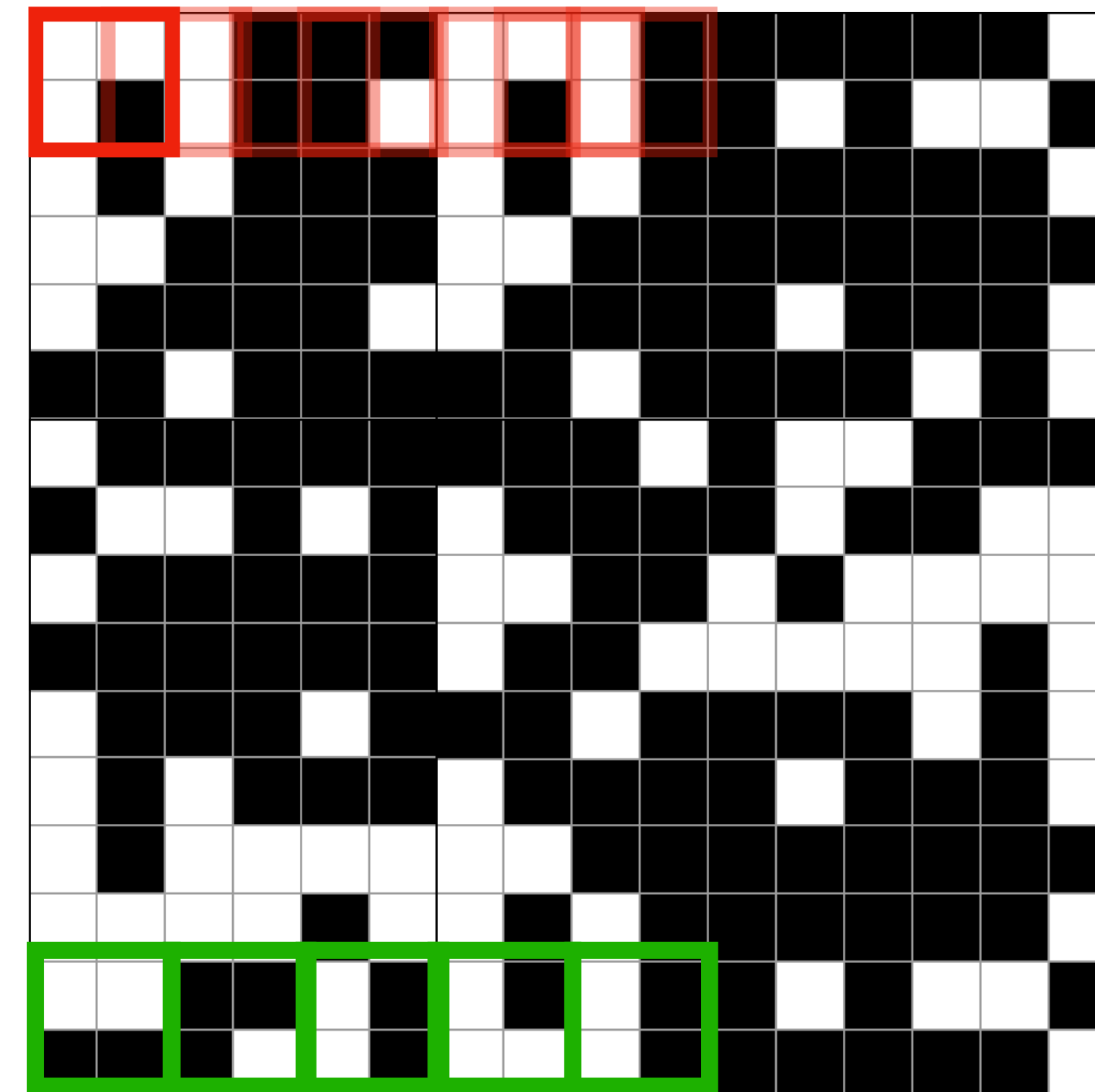
2D Ising Model

[arXiv:2007.01037](https://arxiv.org/abs/2007.01037)



Masked Autoencoder for Distribution Estimation (**MADE**)

Stride = 1



Stride = 2

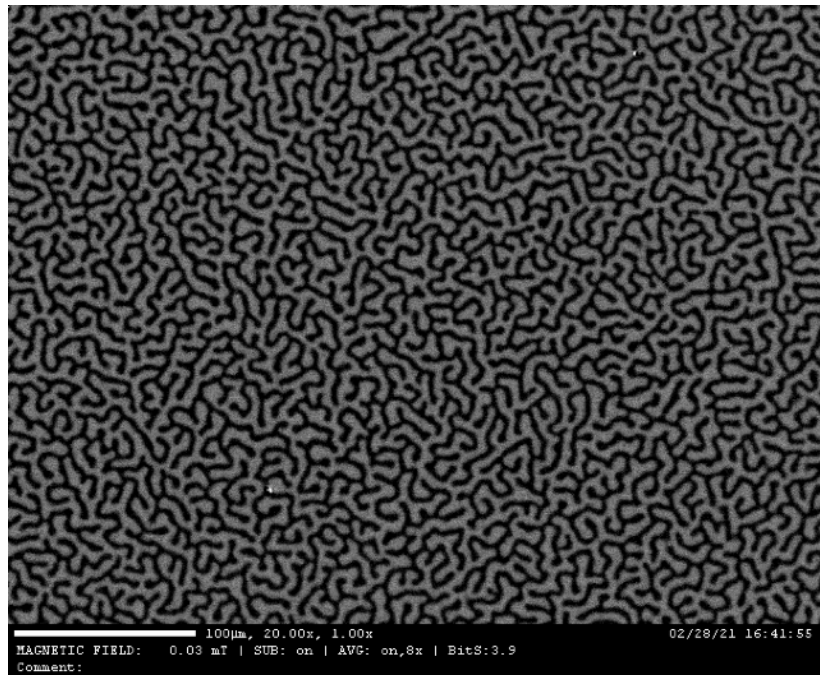


2×2 spin-block

Learn to Detect Phase Transitions

Ferromagnetic Materials

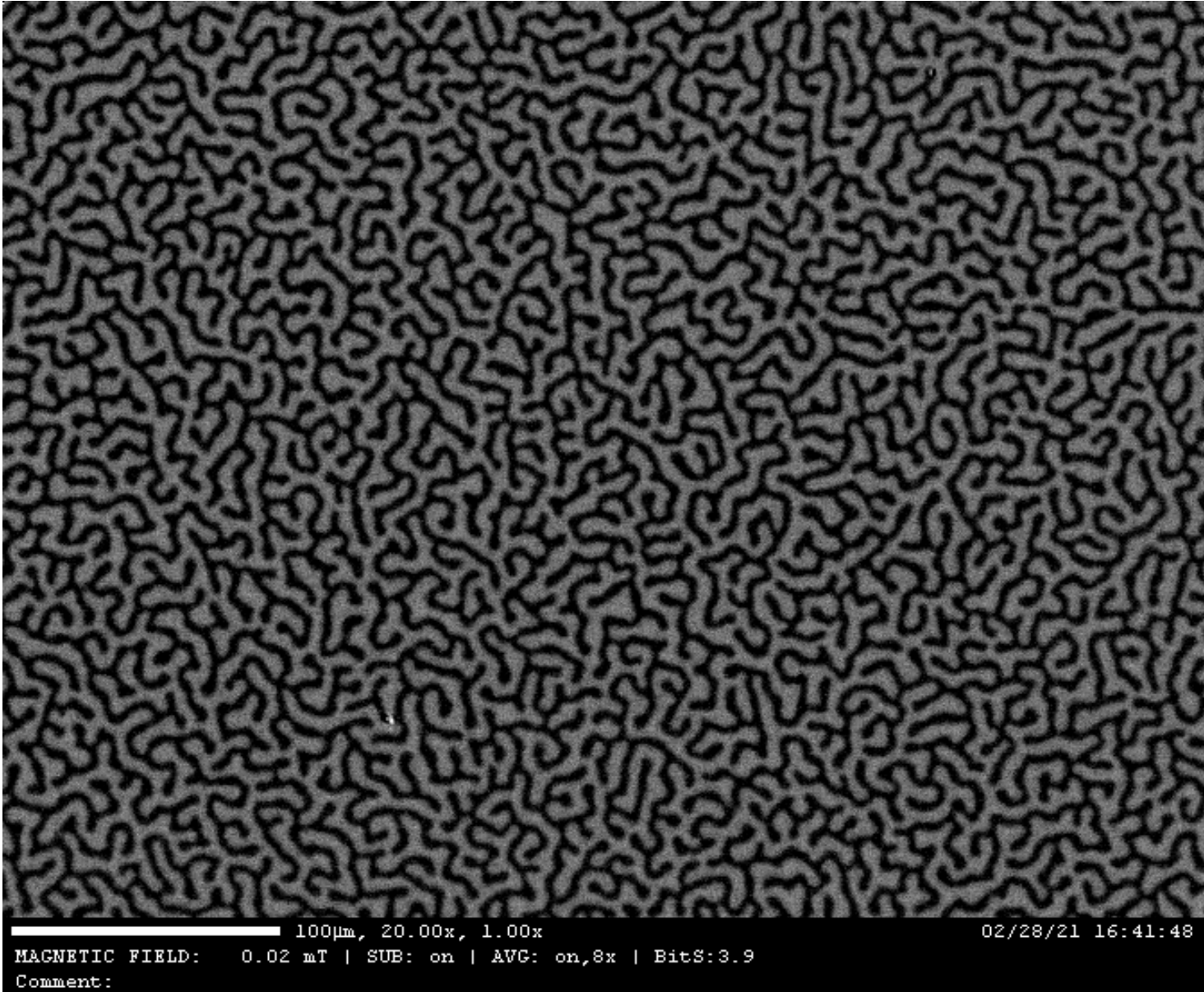
In preparation



...

3000 images

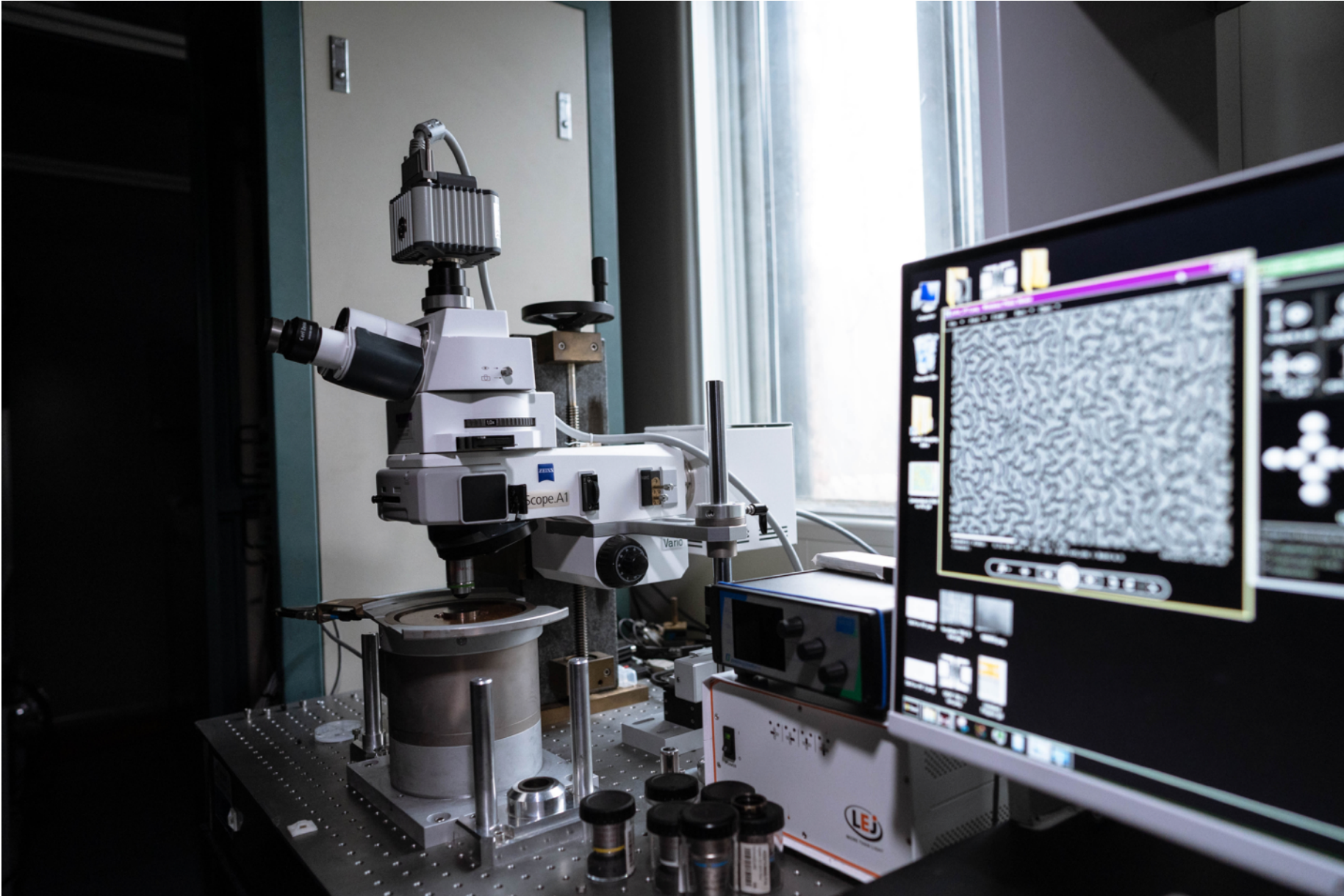
Resolution: 0.5 µm



PhysRevLett.125.027206

with Le Zhao and Wan-Jun Jiang

Magneto-optic Kerr effect (MOKE) microscope to capture images
for the magnetic domains appearing inside a Ta/CoFeB/TaO_x thin film
at room temperature $T = 296\text{ K}$

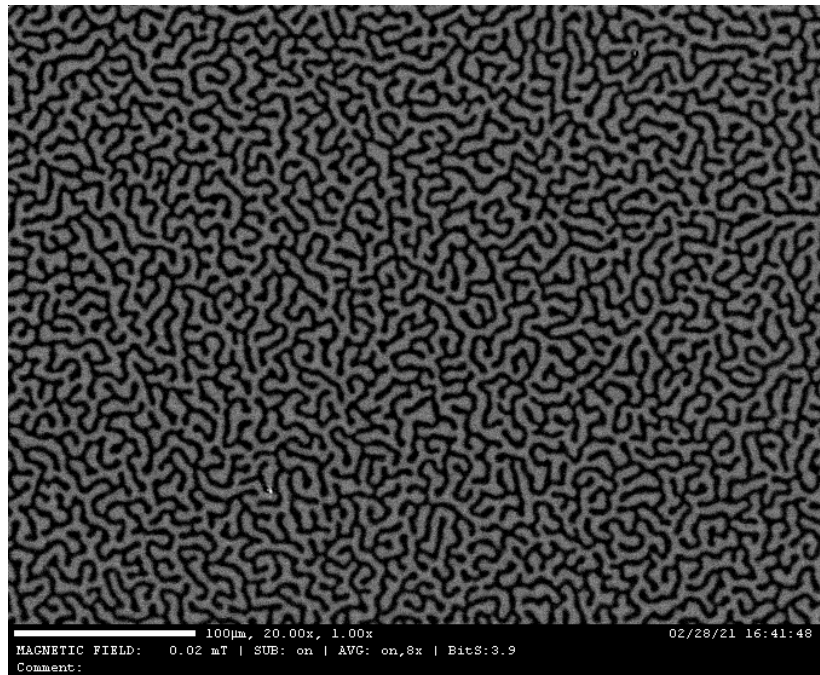
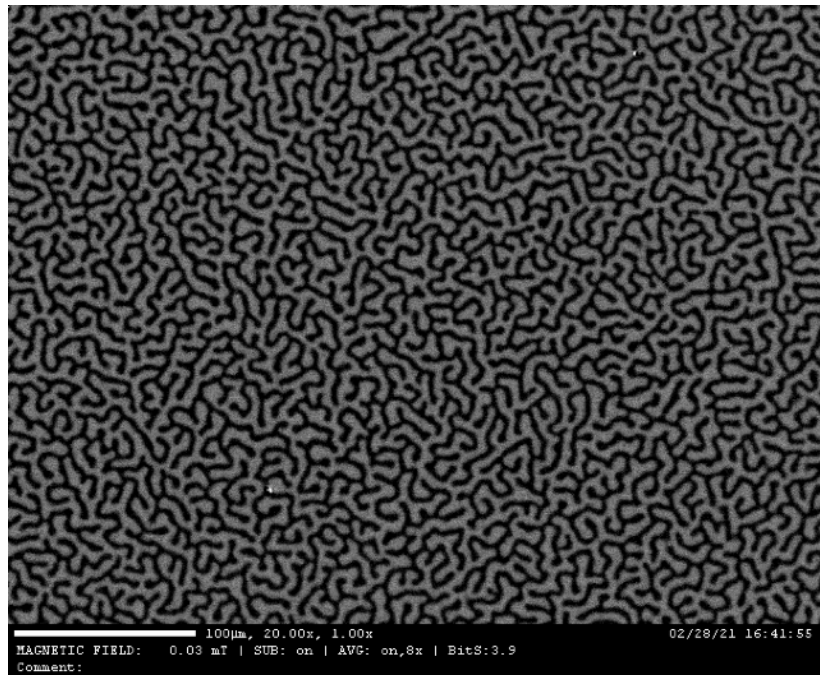


MOKE microscope@THU

Learn to Detect Phase Transitions

Ferromagnetic Materials

In preparation



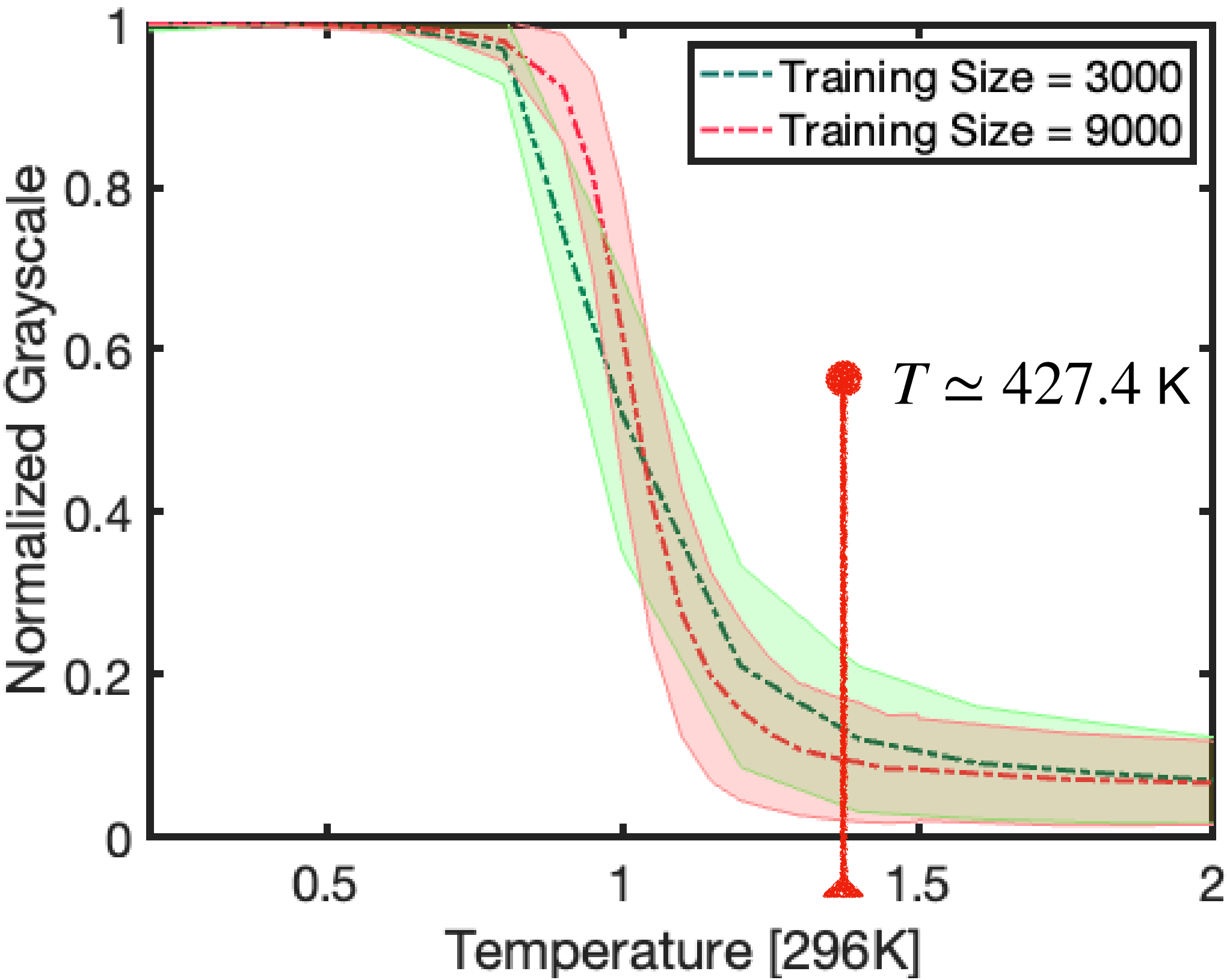
PhysRevLett.125.027206

3000 images

T [K]	50	100	150	200	250	300
M_s [emu/cc]	961.16	925.25	877.73	818.17	753.31	673.427

$$M_s(T) = M_s(0)(1 - T/T_C)^{1/3}$$

$$T_C = 427.4 \pm 2.9 \text{ K}$$



Finite-Temperature Fields

$$Z = \int D\Phi \exp(-S[\Phi])$$

$$S[\Phi] = \sum \Delta\tau(\Delta x)^3 \left[\left(\frac{\Delta\Phi}{\Delta\tau} \right)^2 + (\nabla\Phi)^2 + V(\Phi) \right]$$

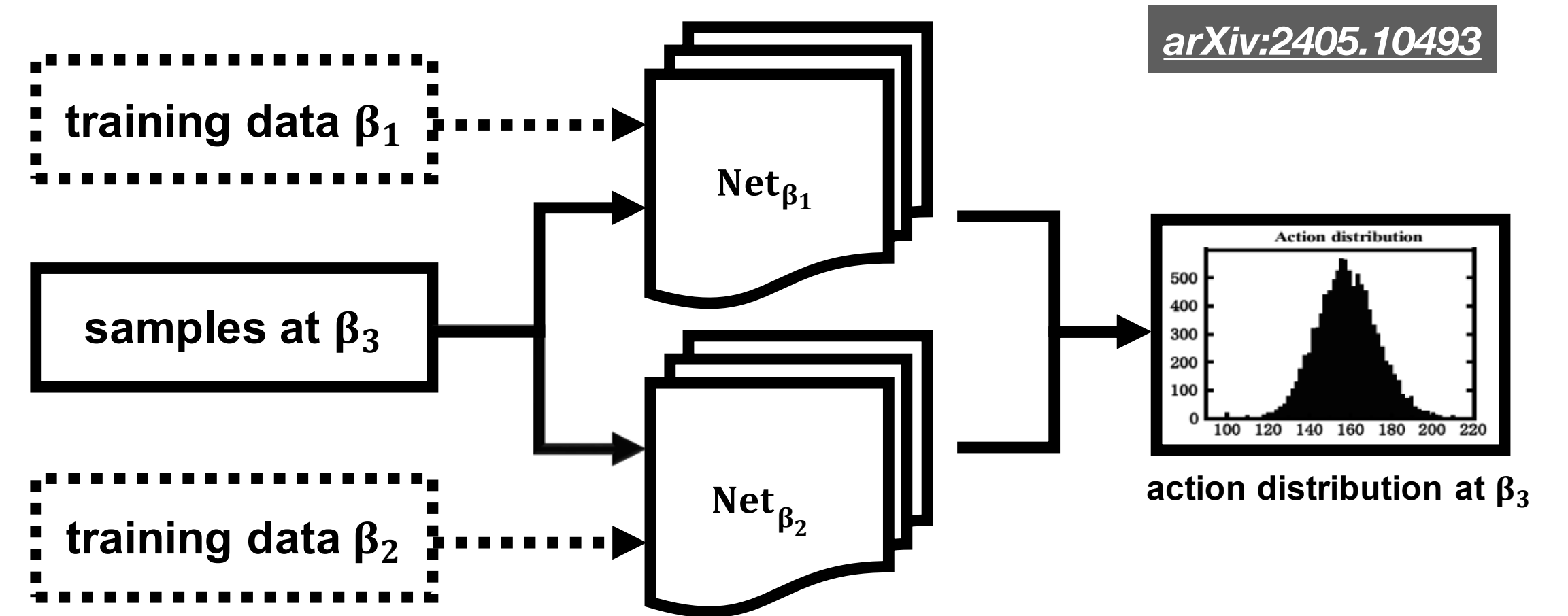
$$= \sum (\Delta x)^3 \left[\frac{(\Delta\Phi)^2}{\Delta\tau} + \Delta\tau((\nabla\Phi)^2 + V(\Phi)) \right]$$

$$= \beta^{-1}K + \beta V$$

$$\Delta\tau = \beta/N_\tau$$

$$K \equiv N_\tau \sum (\Delta x)^3 (\Delta\Phi)^2$$

$$V \equiv N_\tau^{-1} \sum (\Delta x)^3 [(\nabla\Phi)^2 + V(\Phi)]$$



$$S_1[\Phi] = \beta_1^{-1}K[\Phi] + \beta_1 V[\Phi] + C_1$$

$$S_2[\Phi] = \beta_2^{-1}K[\Phi] + \beta_2 V[\Phi] + C_2,$$

$$S_3[\Phi] = \frac{\beta_1(\beta_3^2 - \beta_2^2)}{\beta_3(\beta_1^2 - \beta_2^2)} S_1 + \frac{\beta_2(\beta_1^2 - \beta_3^2)}{\beta_3(\beta_1^2 - \beta_2^2)} S_2 + C_3$$

$$C_3 = \frac{\beta_1(\beta_2^2 - \beta_3^2)}{\beta_3(\beta_1^2 - \beta_2^2)} C_1 + \frac{\beta_2(\beta_3^2 - \beta_1^2)}{\beta_3(\beta_1^2 - \beta_2^2)} C_2$$

Finite-Temperature Fields

0+1 D Quantum Field

arXiv:2405.10493

$$\mathcal{L} = \frac{1}{2} \left(\frac{dx}{d\tau} \right)^2 + V_k(x) \quad V_k(x) = \frac{\lambda_k}{4} \left(x^2 - \frac{\mu_k^2}{2k} \right)^2$$

$$x(\tau) = \pm \frac{\mu_k}{\sqrt{\lambda_k}} \tanh \left[\frac{\mu_k}{\sqrt{2}} (\tau - \tau_0) \right]$$

$$Z = \int_{x(\beta)=x(0)} Dx e^{-S_E[x(\tau)]}$$
$$= \int \prod_{j=-N+1}^{N+1} \frac{dx_j}{\sqrt{2\pi a}} \times \exp \left\{ - \sum_{i=-N+1}^{N+1} \left[\frac{(x_{i+1} - x_i)^2}{2a} + a V_k(x_i) \right] \right\}$$

Kink/Anti-Kink solutions reach

$$\pm \mu_k / \sqrt{\lambda_k} \text{ at } \tau = \pm \infty$$

Numerical Simulations

$$\lambda_k = 4 \quad \mu_k / \sqrt{\lambda_k} = 1.4$$

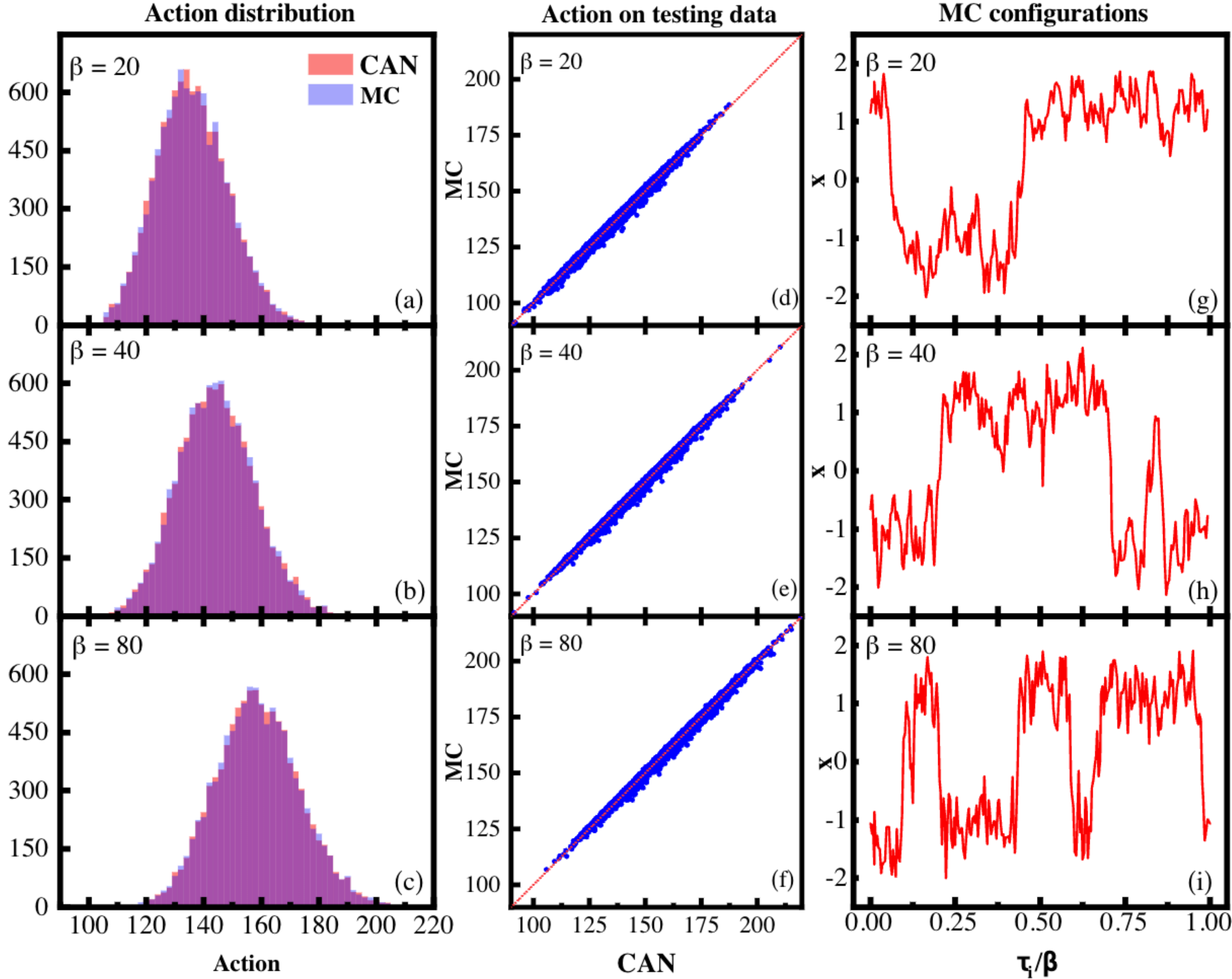
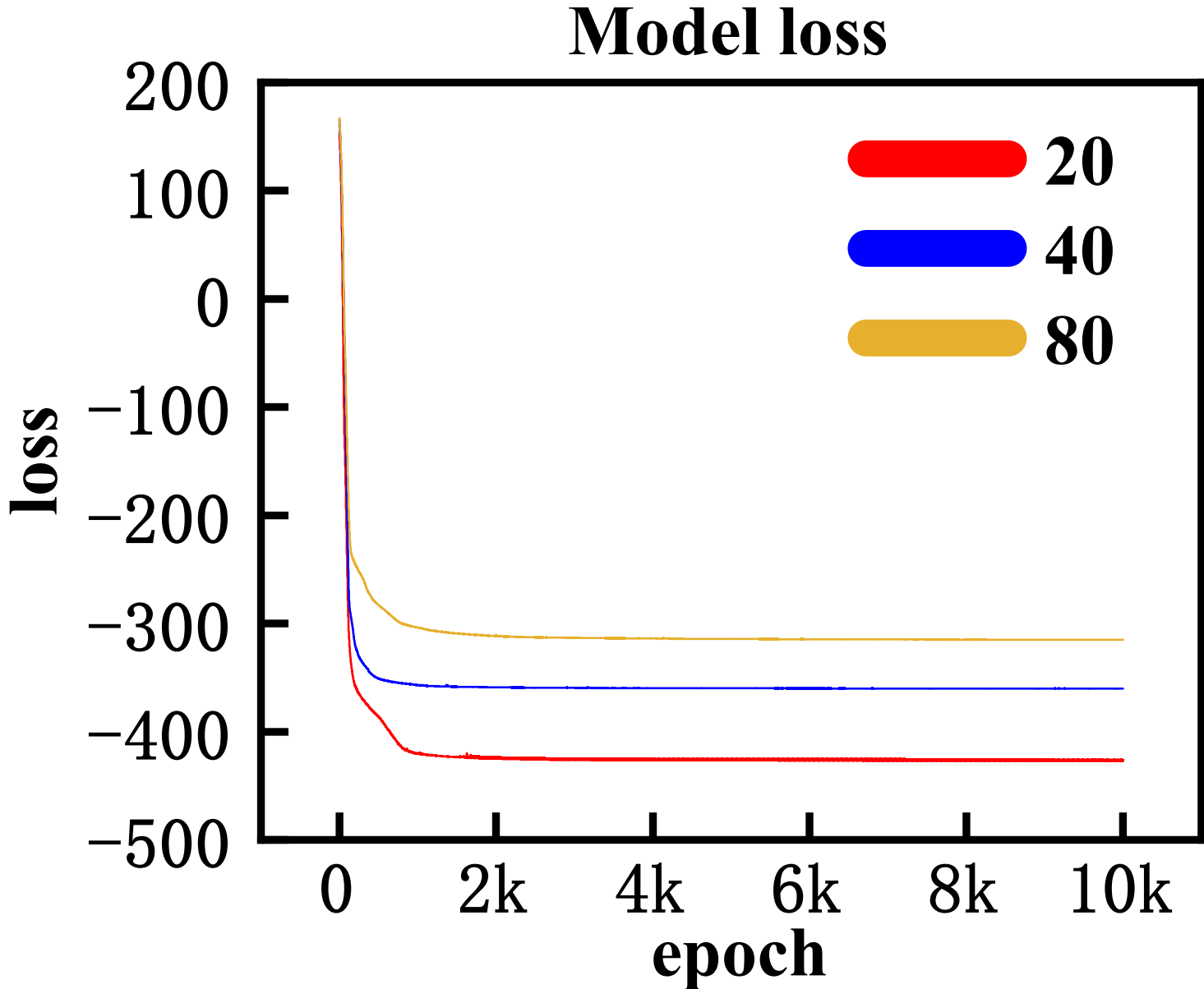
$$S_E[x(\tau)] = \int_0^\beta d\tau \mathcal{L}_E[x(\tau)] = \int_0^\beta d\tau \left[\frac{1}{2} \left(\frac{dx}{d\tau} \right)^2 + V_k(x) \right]$$

$$\beta = T^{-1} = 80, 40, 20 \quad N_{MC} = 5 \times 10^6$$

Finite-Temperature Fields

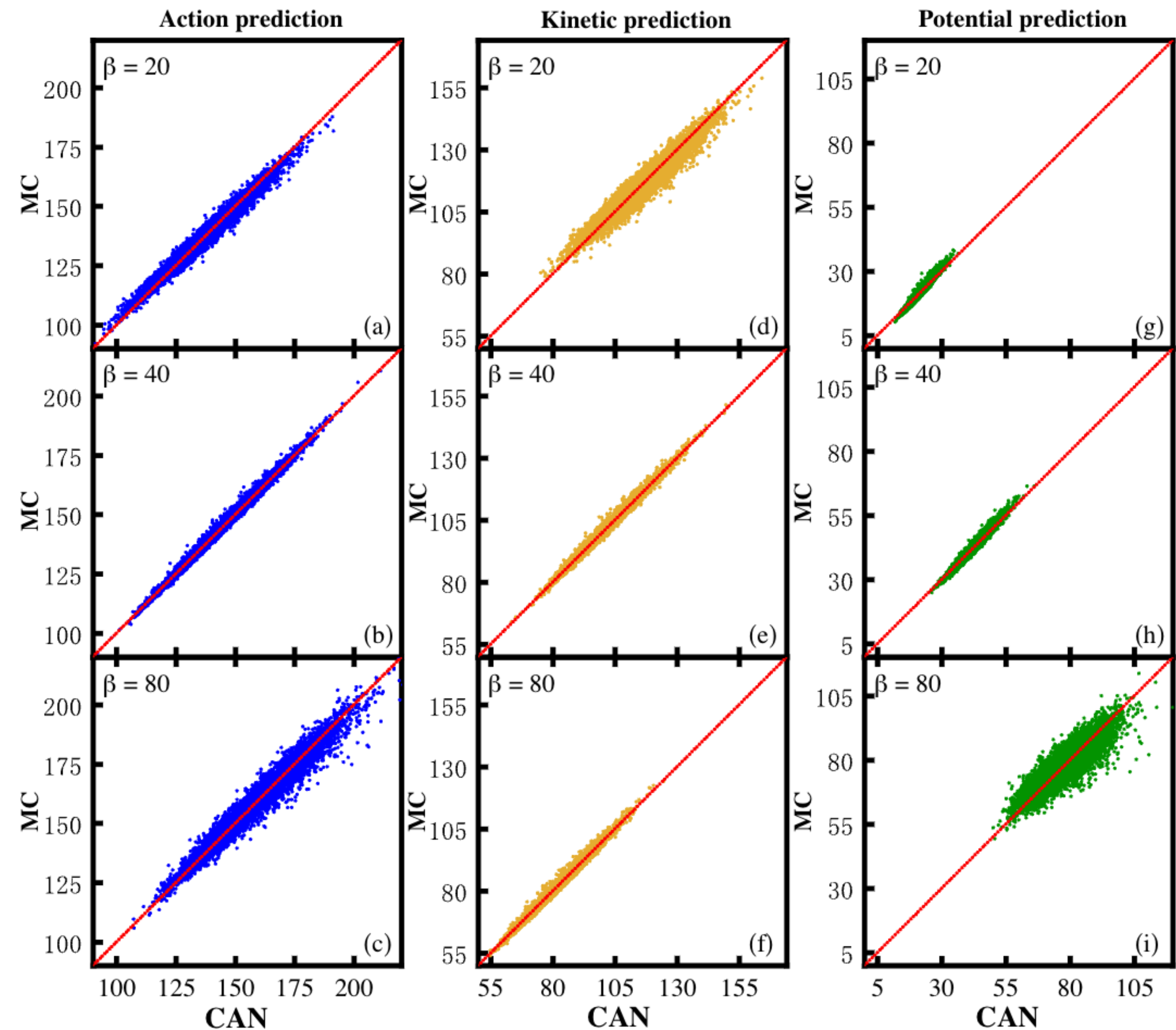
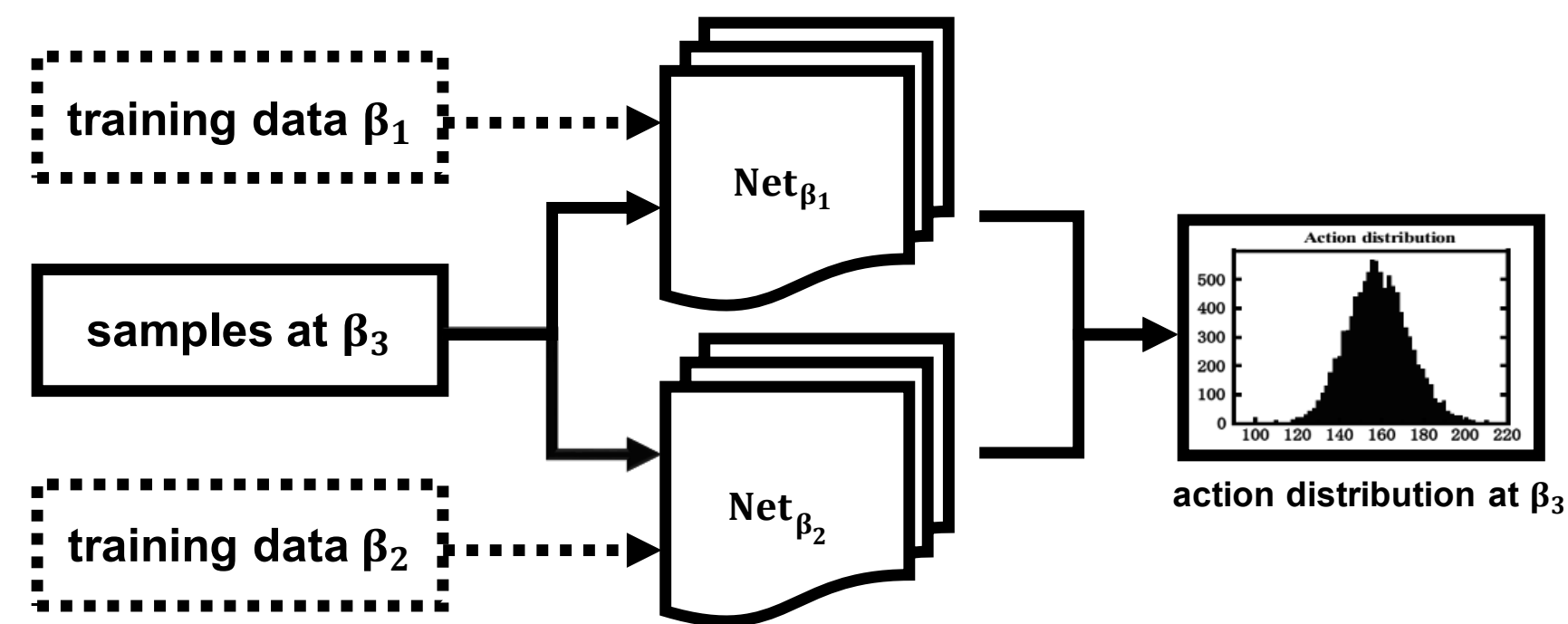
0+1 D Quantum Field

iv:2405.10493



Finite-Temperature Fields

0+1 D Quantum Field



[arXiv:2405.10493](https://arxiv.org/abs/2405.10493)

Diffusion Models

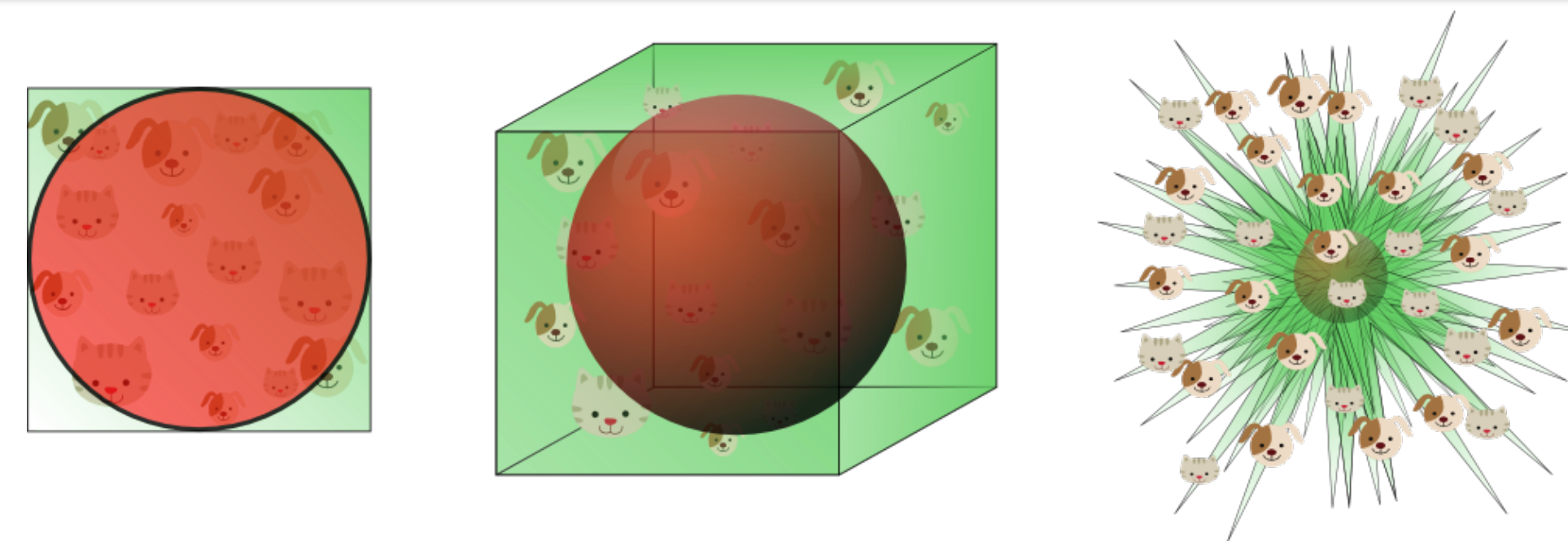
$$\mathbf{x} \sim p_{data}(\mathbf{x})$$

$$p_{\theta}(\mathbf{x}) = \frac{e^{-f_{\theta}(\mathbf{x})}}{\textcircled{Z_{\theta}}} \text{ 😵 }$$

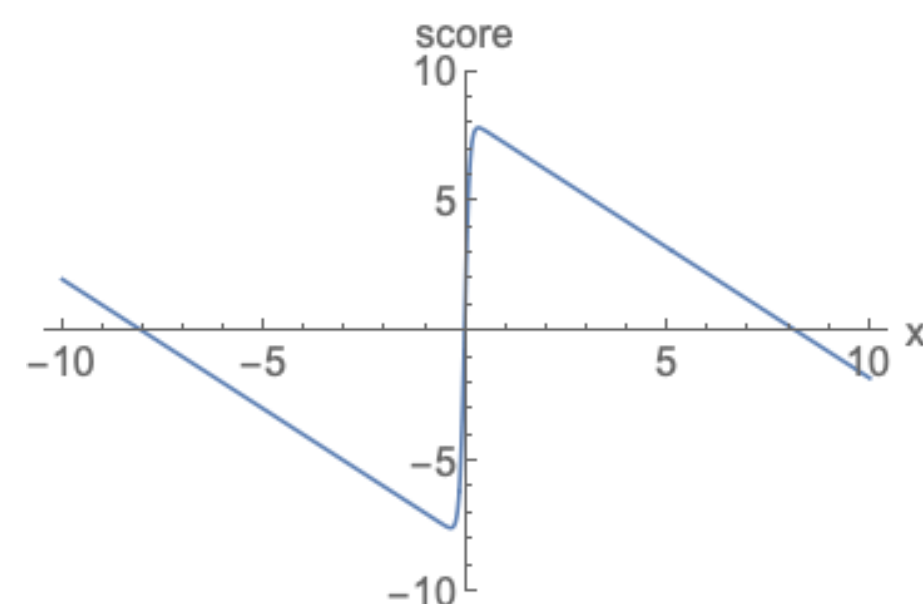
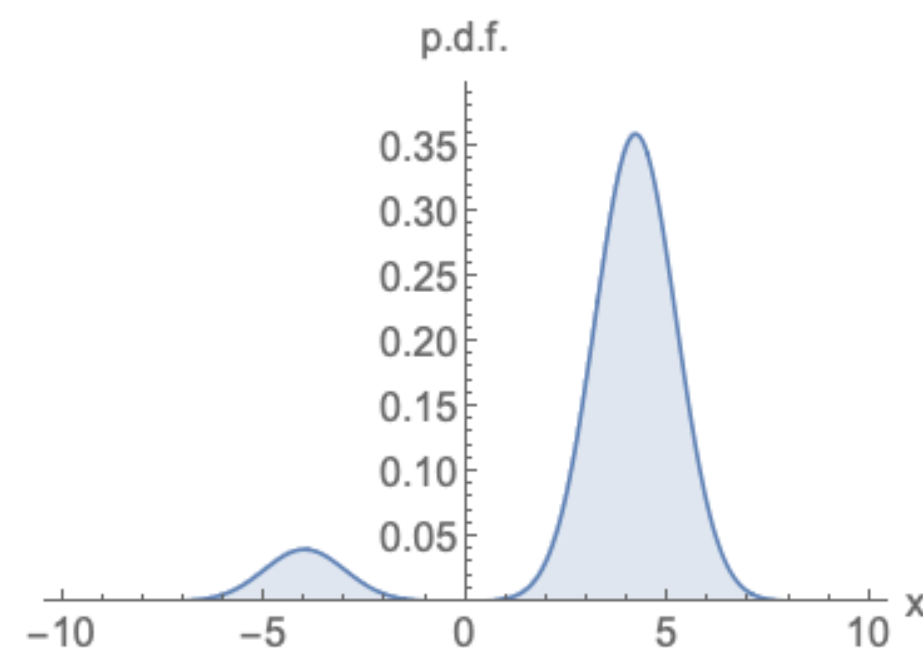
from a **statistical physics** perspective

$$\nabla_{\mathbf{x}} \log p(\mathbf{x})$$

$$\mathbf{s}_{\theta}(\mathbf{x}) \rightarrow \nabla_{\mathbf{x}} \log p(\mathbf{x})$$



Curse of Dimensionality



Approaching **Score Function**

$$\begin{aligned} \mathbf{s}_{\theta}(\mathbf{x}) &= \nabla_{\mathbf{x}} \log p_{\theta}(\mathbf{x}) \\ &= -\nabla_{\mathbf{x}} f_{\theta}(\mathbf{x}) - \underbrace{\nabla_{\mathbf{x}} \log Z_{\theta}}_{=0} = -\nabla_{\mathbf{x}} f_{\theta}(\mathbf{x}) \end{aligned}$$

Diffusion Models

Song et al., Score-Based Generative Modeling through Stochastic Differential Equations, ICLR 2021

- **Forward Diffusion SDE**

- **Drift term**: pulls towards mode
- **Diffusion term**: injects noise

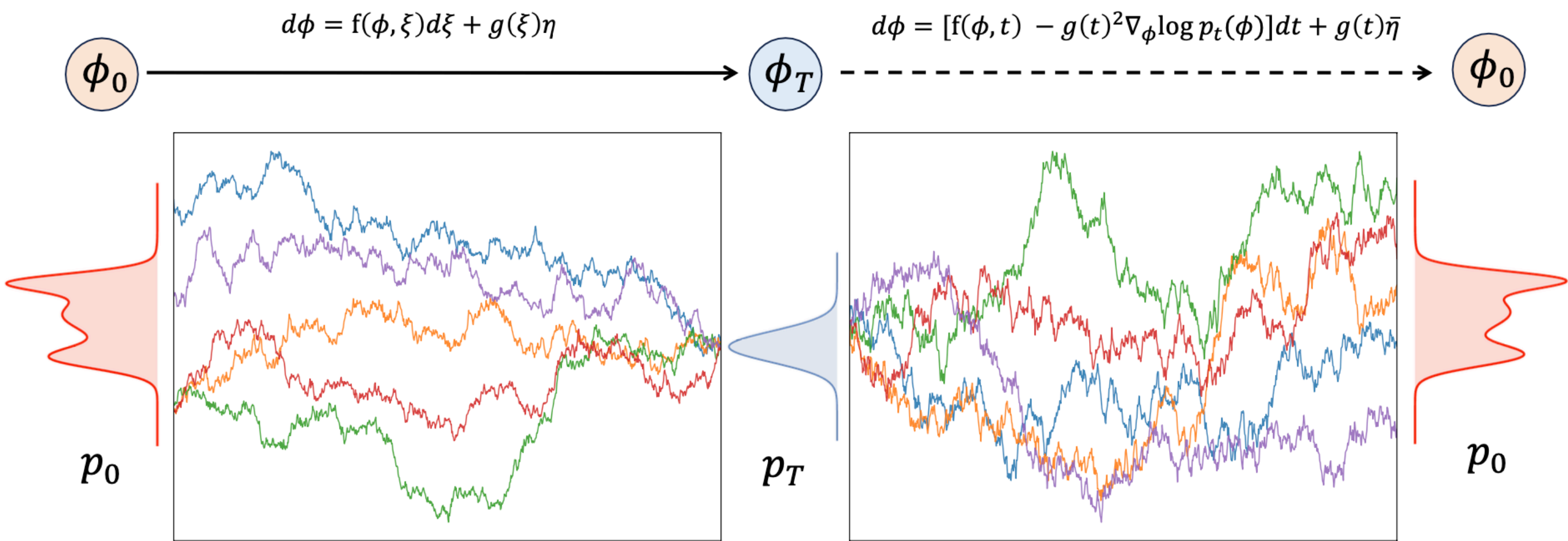
$$\frac{d\phi}{d\xi} = f(\phi, \xi) + g(\xi)\eta(\xi)$$

Anderson, in Stochastic Processes and their Applications, 1982

- **Reverse Generative Diffusion SDE**

- Drift term is adjusted with a “**Score Function**”
- Represent the score function with **neural networks**

$$\frac{d\phi}{dt} = \left[f(\phi, t) - g^2(t) \nabla_{\phi} \log p_t(\phi) \right] + g(t)\bar{\eta}(t)$$



Diffusion Models

Stochastic Quantization

$$\frac{\partial \phi(x, \tau)}{\partial \tau} = - \frac{\delta S_E[\phi]}{\delta \phi(x, \tau)} + \eta(x, \tau)$$

$$\langle \eta(x, \tau) \rangle = 0, \quad \langle \eta(x, \tau) \eta(x', \tau') \rangle = 2\alpha \delta(x - x') \delta(\tau - \tau')$$

τ : fictitious time, α : diffusion constant

- Fokker-Planck equation

$$\frac{\partial P[\phi, \tau]}{\partial \tau} = \alpha \int d^n x \left\{ \frac{\delta}{\delta \phi} \left(\frac{\delta}{\delta \phi} + \frac{\delta S_E}{\delta \phi} \right) \right\} P[\phi, \tau]$$

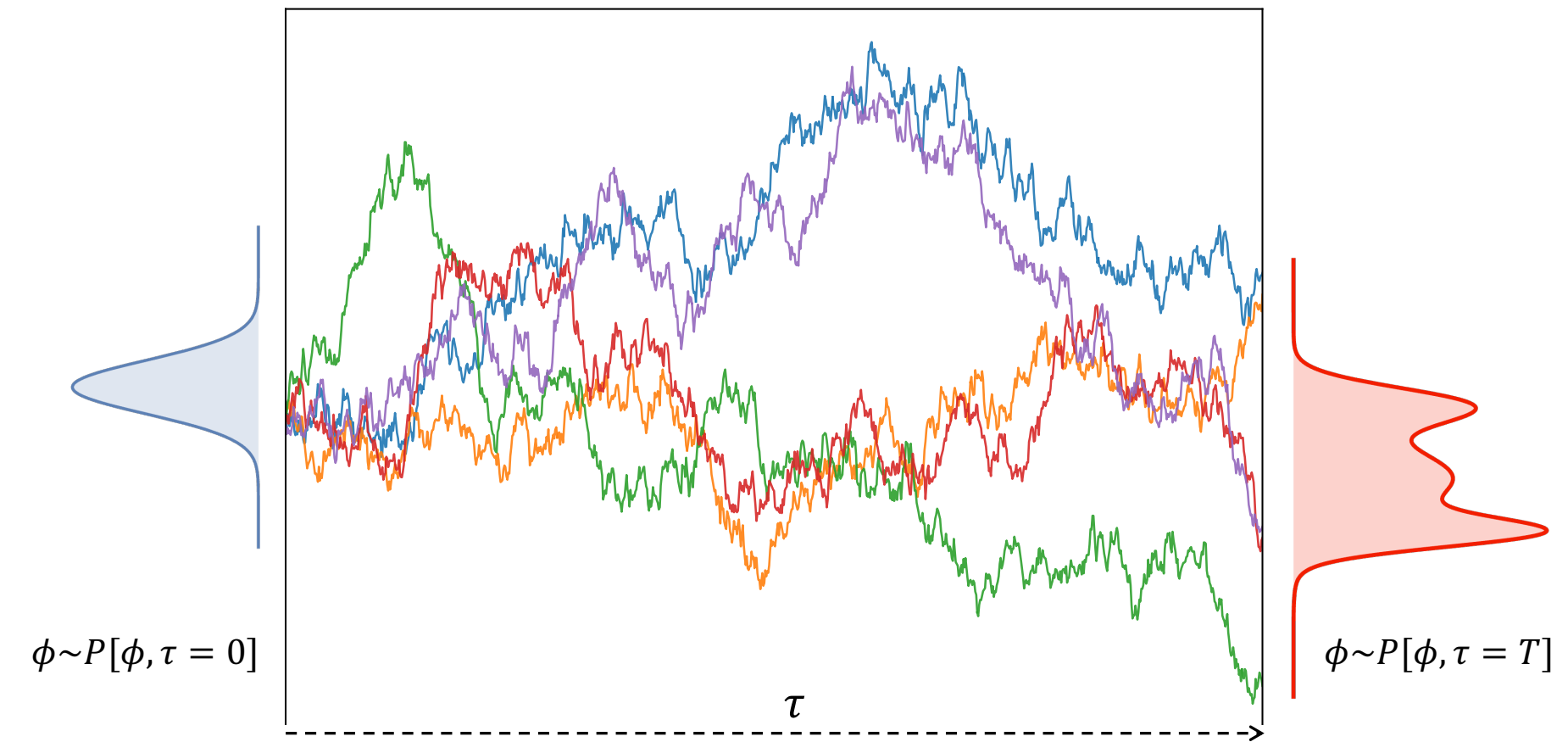
Equilibrium solution (long-time limit),

$$P_{\text{eq}}[\phi] \propto e^{-\frac{1}{\alpha} S_E[\phi]}$$

- Set the diffusion constant as $\alpha = \hbar$

$$P_{\text{eq}}[\phi] \sim e^{-\frac{1}{\hbar} S_E[\phi]} = P_{\text{quantum}}[\phi]$$

Parisi G. and Wu Y. S., Sci. China, A **24**, ASITP-80-004 (1980).



Thermal equilibrium limit → Quantum distribution

1. No need gauge-fixing!
2. Can handle fermionic fields naturally
→ (Complex Langevin method)

....

P. H. Damgaard and H. Hüffel, Stochastic Quantization, Phys. Rept. 152, 227 (1987).

M. Namiki, Basic Ideas of Stochastic Quantization, PTPS 111, 1 (1993).

G. Aarts, L. Bongiovanni, E. Seiler, D. Sexty, and I.-O. Stamatescu, Eur. Phys. J. A **49**, 89 (2013).

- Diffusion models(Reverse SDE):

$$\frac{d\phi}{dt} = -g(t)^2 \nabla_{\phi} \log p_t(\phi) + g(t) \bar{\eta}$$

- Define: $\tau \equiv T - t (d\tau \equiv -dt)$

$$\frac{d\phi}{d\tau} = g_{\tau}^2 \nabla_{\phi} \log q_{\tau}(\phi) + g_{\tau} \bar{\eta}$$

$$\phi(\tau_{n+1}) = \phi(\tau_n) + g_{\tau}^2 \nabla_{\phi} \log q_{\tau_n}[\phi(\tau_n)] \Delta\tau + g_{\tau} \sqrt{\Delta\tau} \bar{\eta}(\tau_n)$$

introducing **Noise scale**: $\langle \bar{\eta}^2 \rangle \equiv 2\bar{\alpha}$, **time scale**: $g_{\tau}^2 \Delta\tau$

- FP equation

$$\frac{\partial p_{\tau}(\phi)}{\partial \tau} = \int d^n x \left\{ g_{\tau}^2 \bar{\alpha} \frac{\delta}{\delta \phi} \left(\frac{\delta}{\delta \phi} + \frac{1}{\bar{\alpha}} \nabla_{\phi} S_{\mathbf{DM}} \right) \right\} p_{\tau}(\phi)$$

$$\nabla_{\phi} S_{\mathbf{DM}} \equiv -\nabla_{\phi} \log q_{\tau}(\phi)$$

$$p_{eq}(\phi) \propto e^{-\frac{S_{DM}}{\bar{\alpha}}}$$

$$p_{\tau=T}(\phi) \rightarrow P[\phi, T]$$

$$O(\bar{\alpha}) \sim O(\hbar)$$

The reverse mode of
a well-trained diffusion model at $\tau \rightarrow T$ serves as
the stochastic quantization for the input

Diffusion Models

DM for Scalar Field

JHEP 05(2024)060

- **Euclidean action on lattice**

$$S_E = \sum_x a^d \left[\sum_{\mu=1}^d \frac{(\phi_0(x + a\hat{\mu}) - \phi_0(x))^2}{a^2} + \frac{m_0^2}{2} \phi_0^2 + \frac{\lambda_0}{4!} \phi_0^4 \right]$$

- **Dimensionless form**

$$S_E = \sum_x \left[-2\kappa \sum_{\mu=1}^d \phi(x)\phi(x + \hat{\mu}) + (1 - 2\lambda)\phi(x)^2 + \lambda\phi(x)^4 \right]$$

$$a^{\frac{d-2}{2}} \phi_0 = (2\kappa)^{1/2} \phi$$
$$(am_0)^2 = \frac{1 - 2\lambda}{\kappa} - 2d, \quad a^{-d+4} \lambda_0 = \frac{6\lambda}{\kappa^2}$$

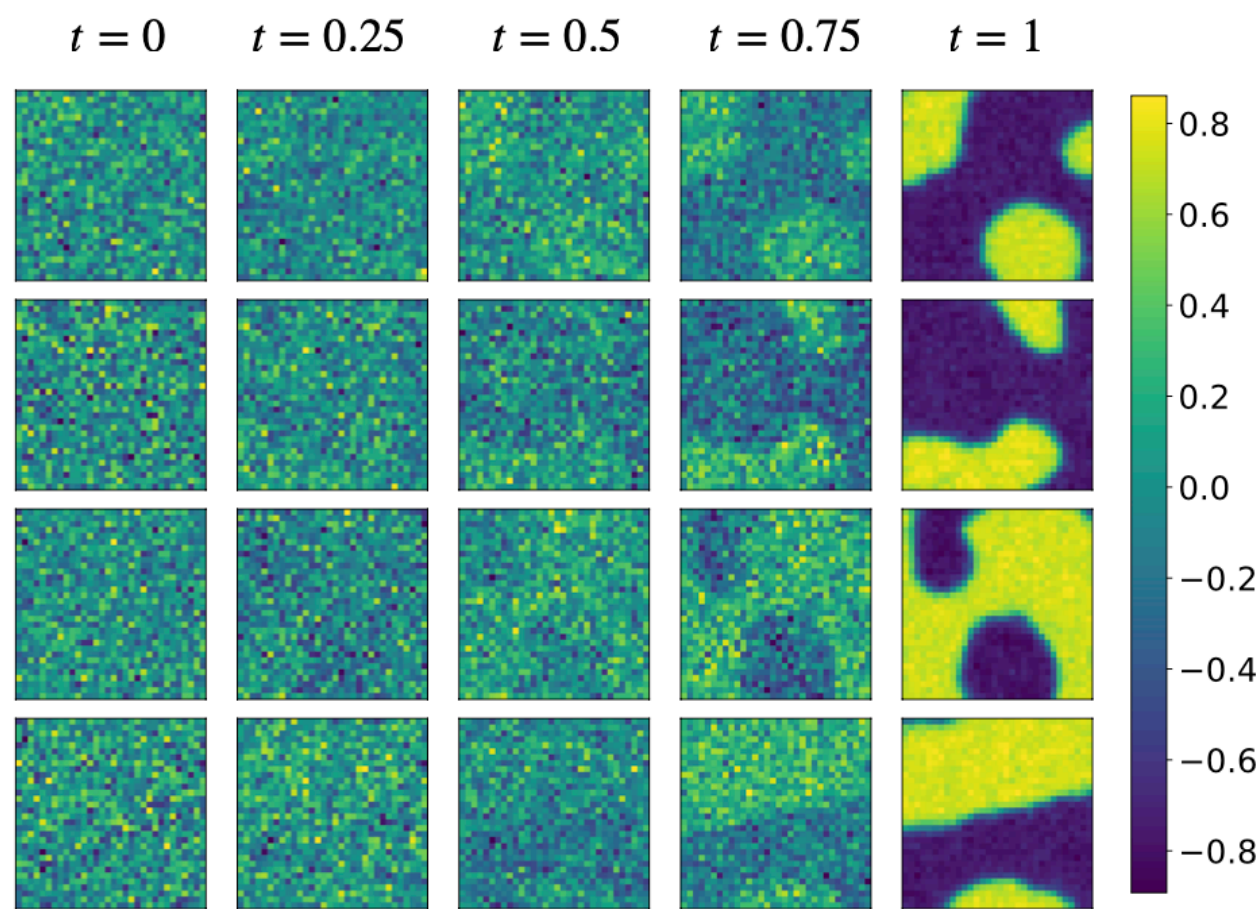
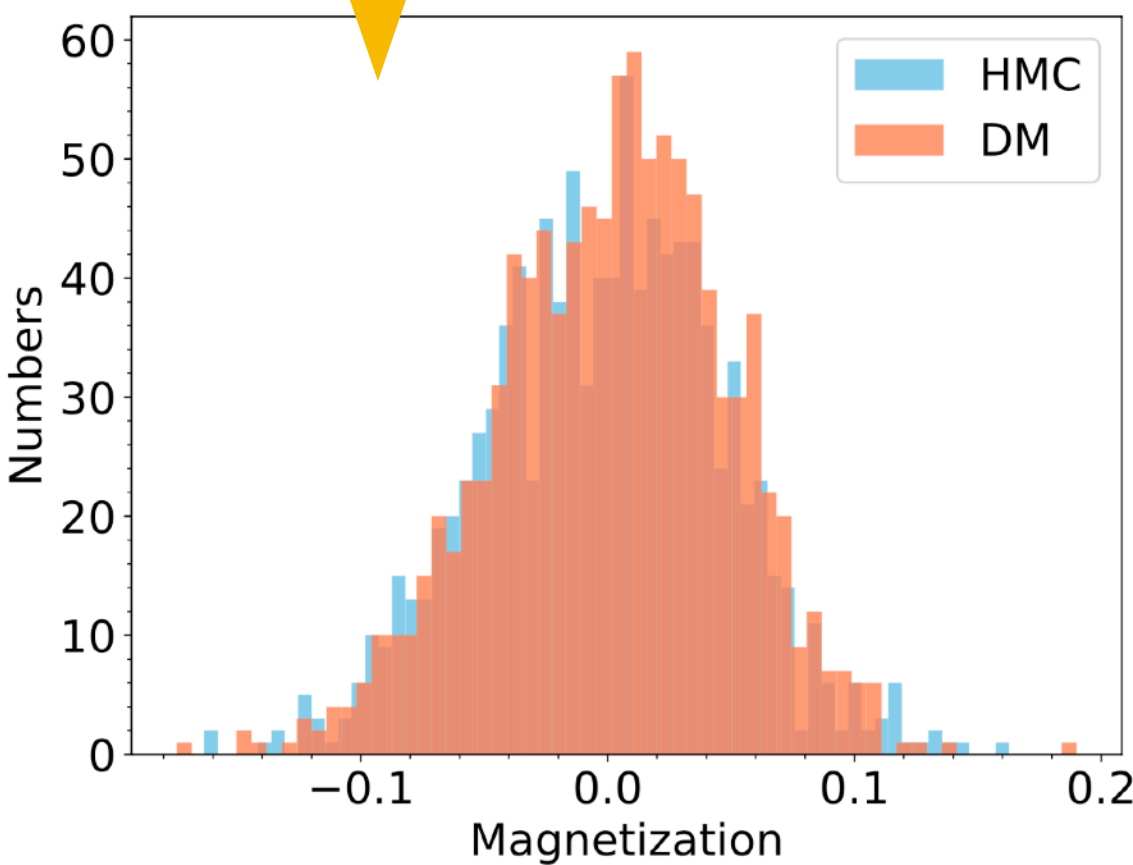
- Hopping parameter κ ,
Coupling constant λ

- **Diffusion models**

- $T = 1.0, \sigma = 25$

- **Data generation**

- 2-d 32×32 lattice; Hamiltonian Monte Carlo(HMC); 5120 configurations for training.
- Broken phase: $\kappa = 1.0, \lambda = 0.022$
- Symmetric phase: $\kappa = 0.21, \lambda = 0.022$

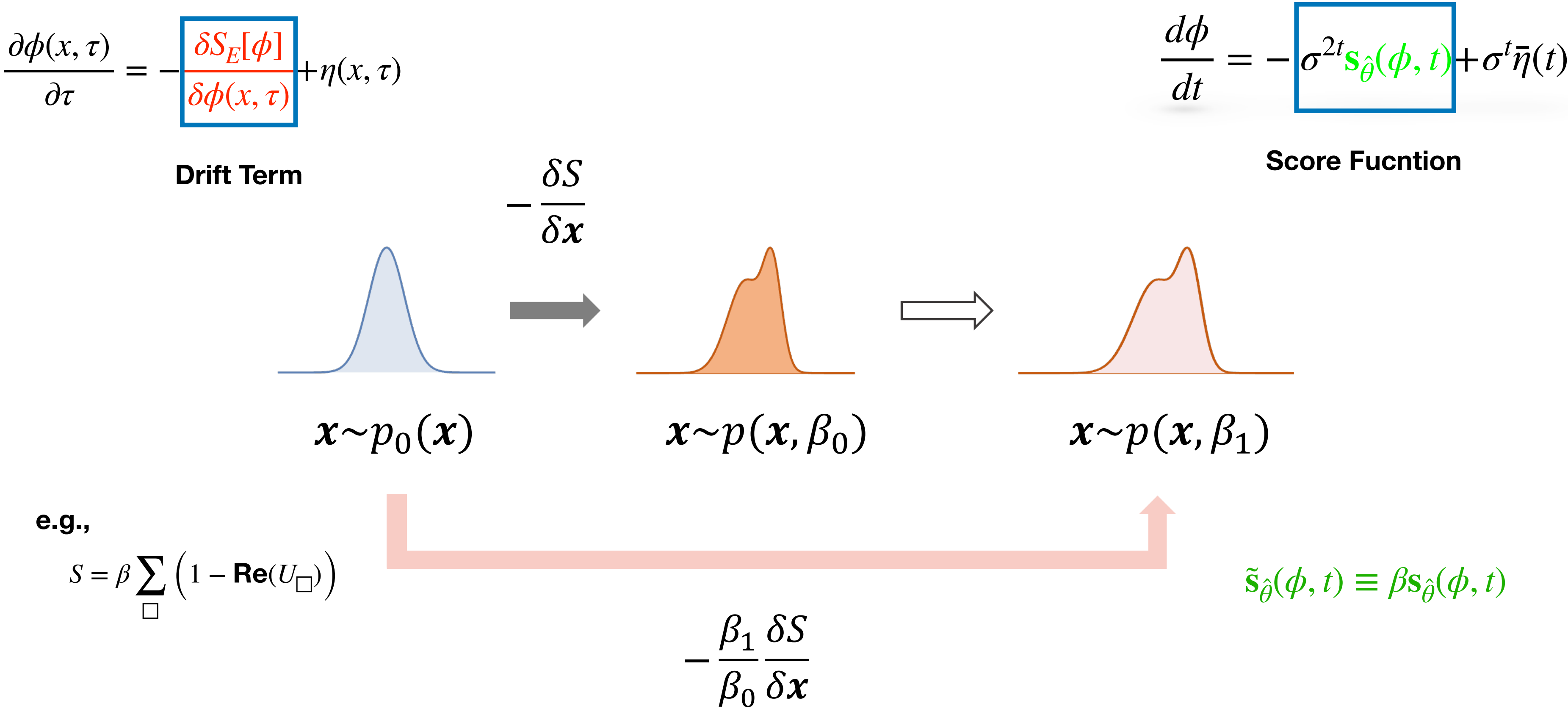


data-set	$\langle M \rangle$	χ_2	U_L
Training(HMC)	0.0012 ± 0.0007	2.5160 ± 0.0457	0.1042 ± 0.0367
Testing(HMC)	0.0018 ± 0.0015	2.4463 ± 0.1099	-0.0198 ± 0.1035
Generated(DM)	0.0017 ± 0.0015	2.4227 ± 0.1035	0.0484 ± 0.0959

Physics-Conditioned Diffusion Models

NeurIPS2024 ML4PS Workshop

with Qianteng Zhu, Gert Aarts, etc.



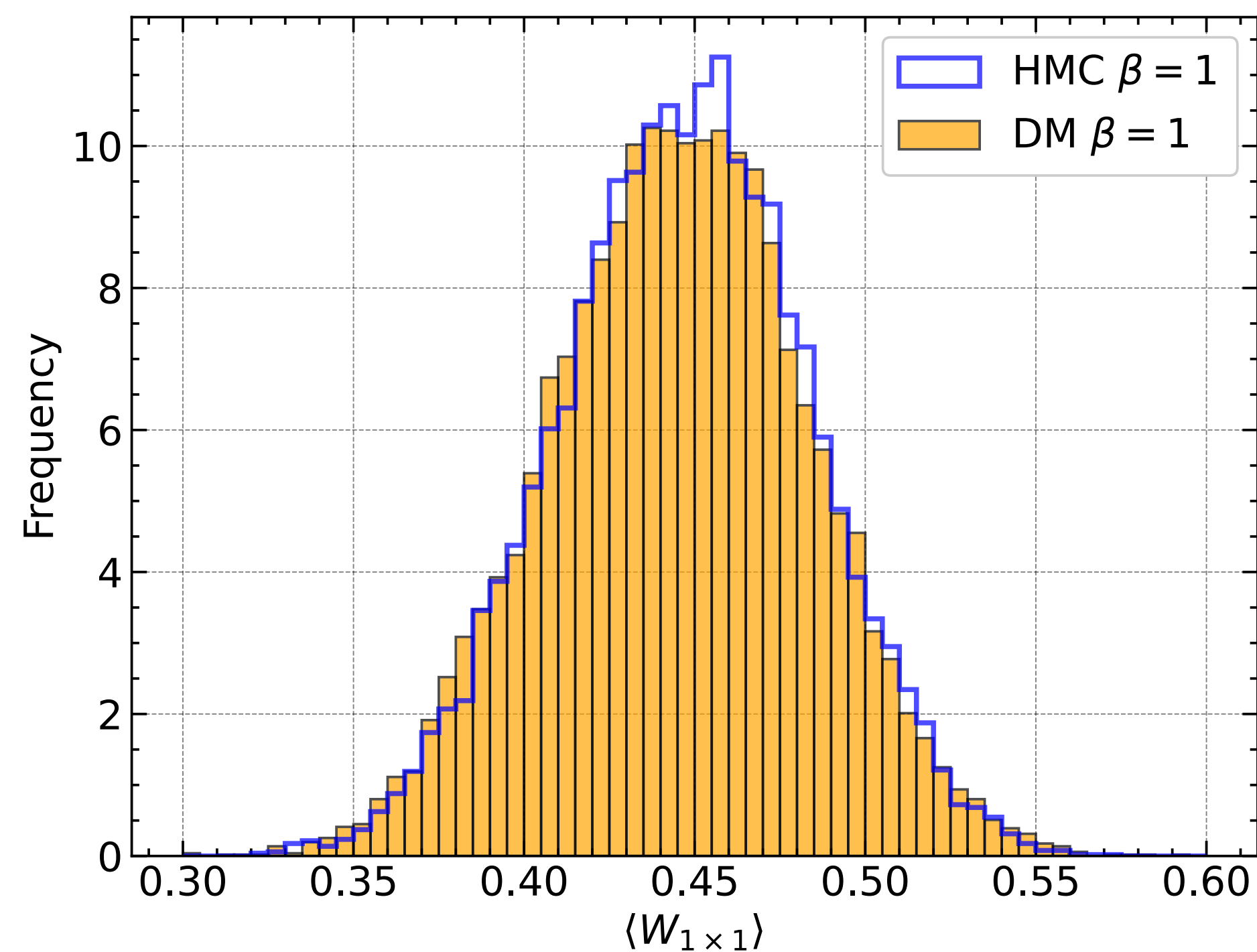
Gauge Fields

2D U(1) Gauge Field

$$S = \beta \sum_{\square} \left(1 - \mathbf{Re}(U_{\square}) \right)$$

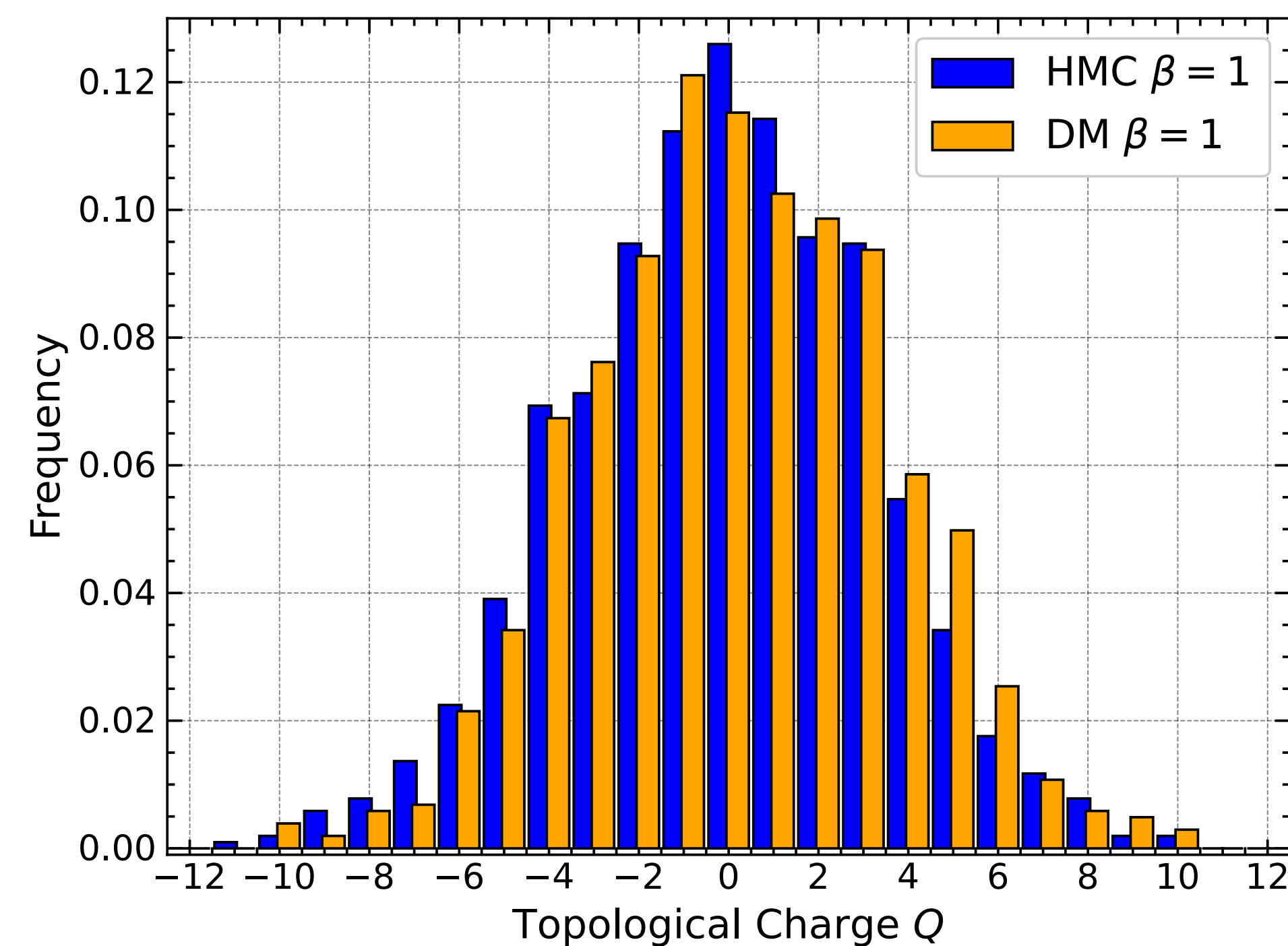
plaquette

$$U_{\square} = U_{x,\mu} U_{x+\hat{\mu},\nu} U_{x+\hat{\nu},\mu}^{\dagger} U_{x,\nu}^{\dagger}$$



$$W(C) = \text{Tr} \left(\prod_{(x,\mu) \in C} U_{x,\mu} \right)$$

Learned at $\beta = 1$ with **10,240** configurations, $L = 16$
Generated 1024 configs for testing

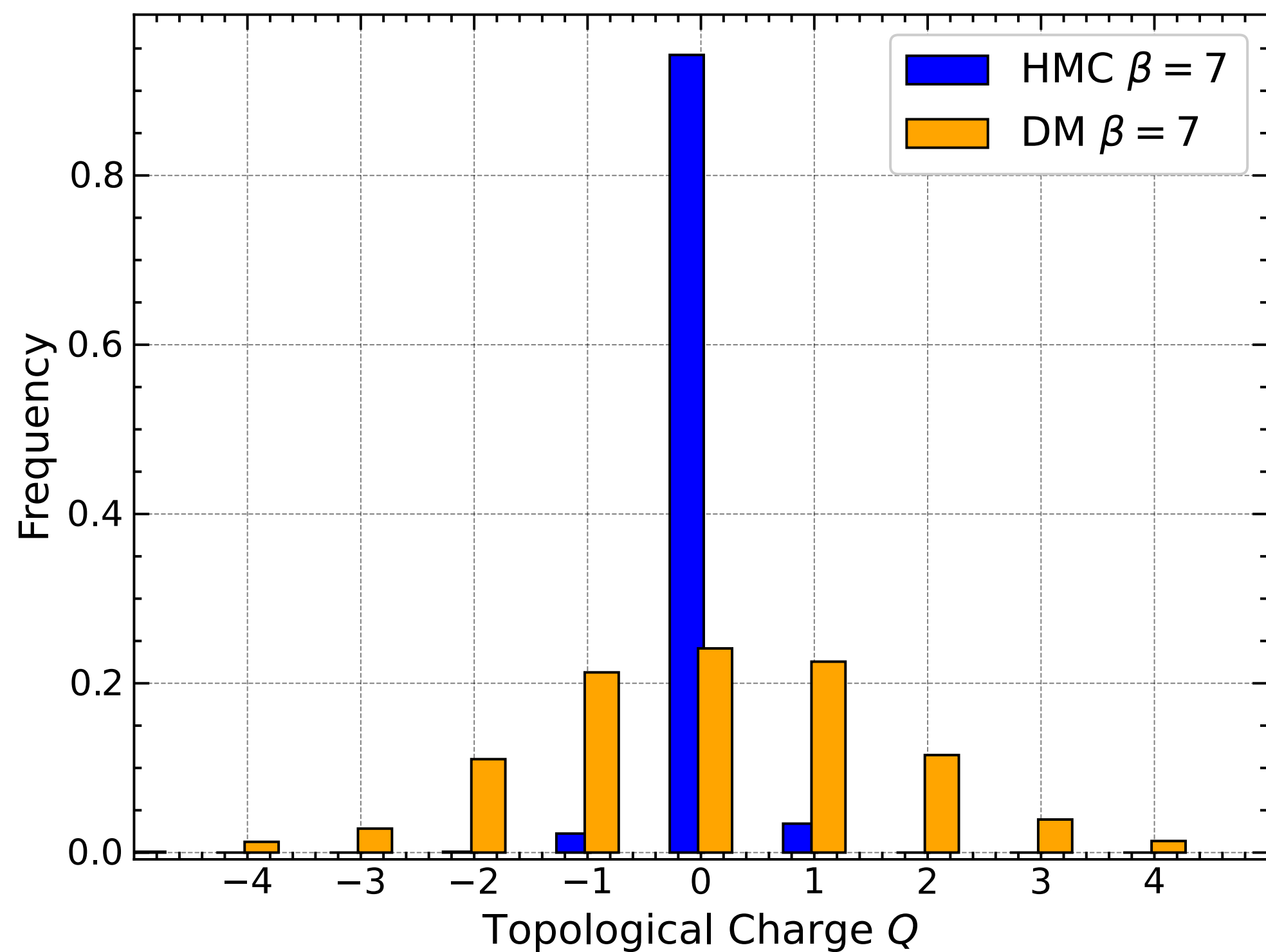


$$Q = \frac{1}{2\pi} \sum_x F_{01}(x)$$

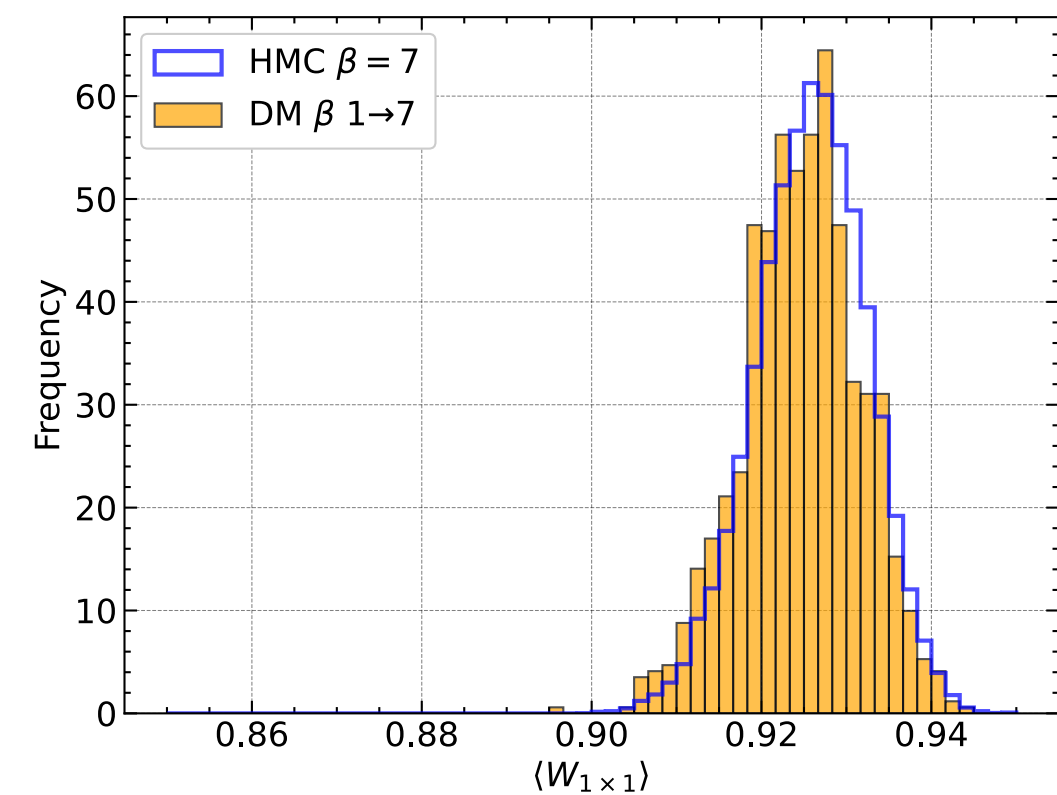
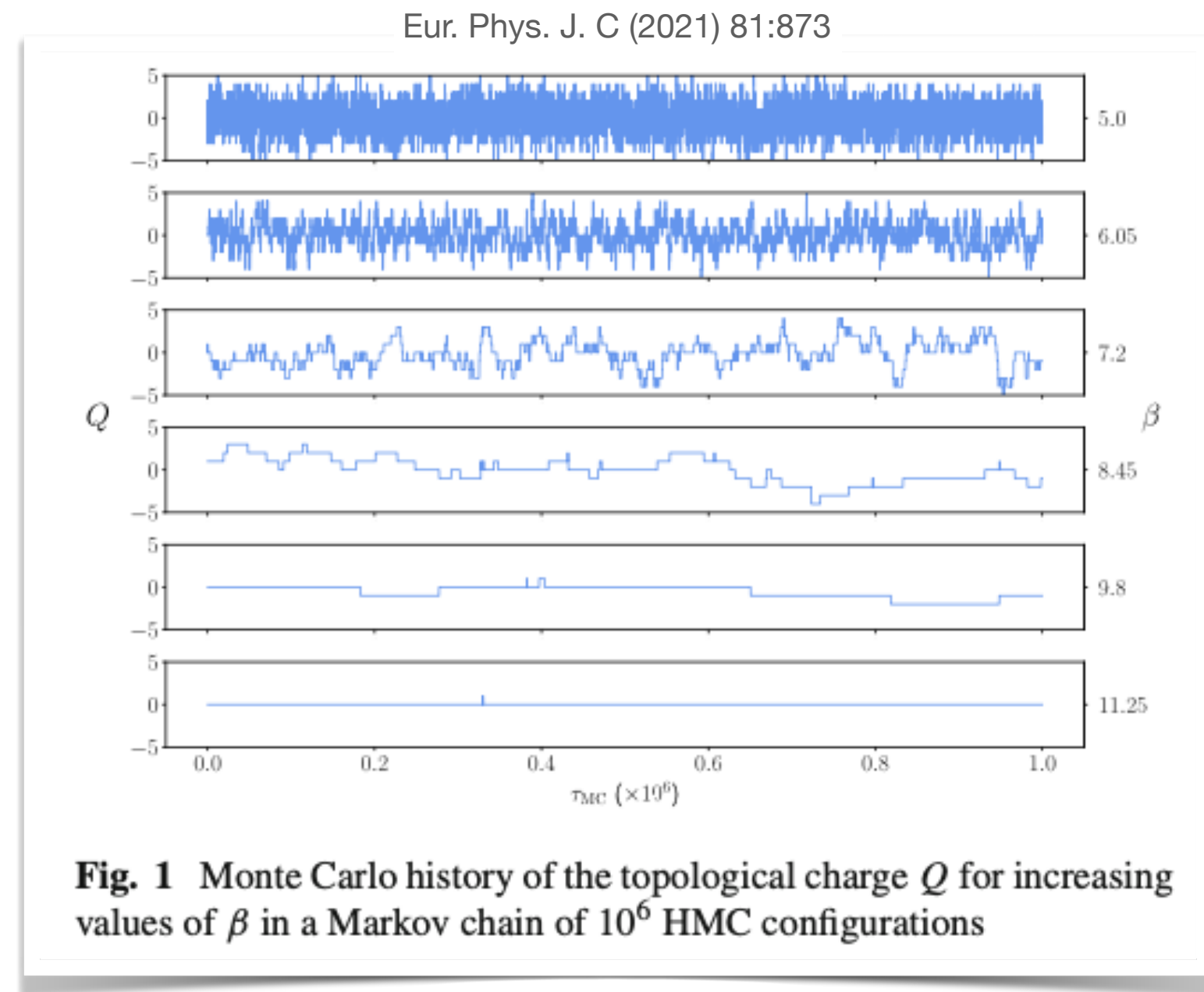
Gauge Fields

2D U(1) Gauge Field

Topological Freezing?

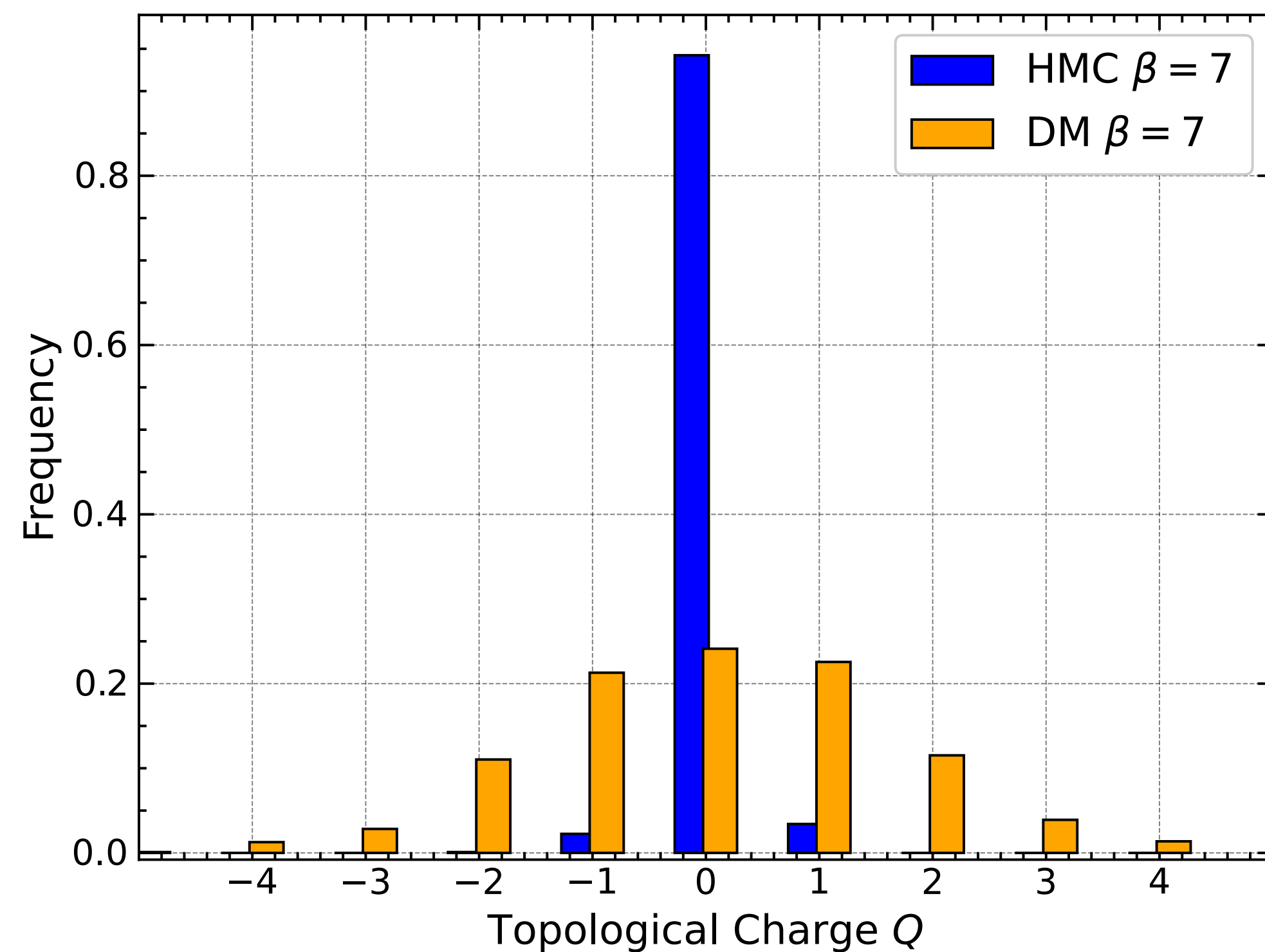


Generated at $\beta = 7$ with 1024 configurations, $L = 16$

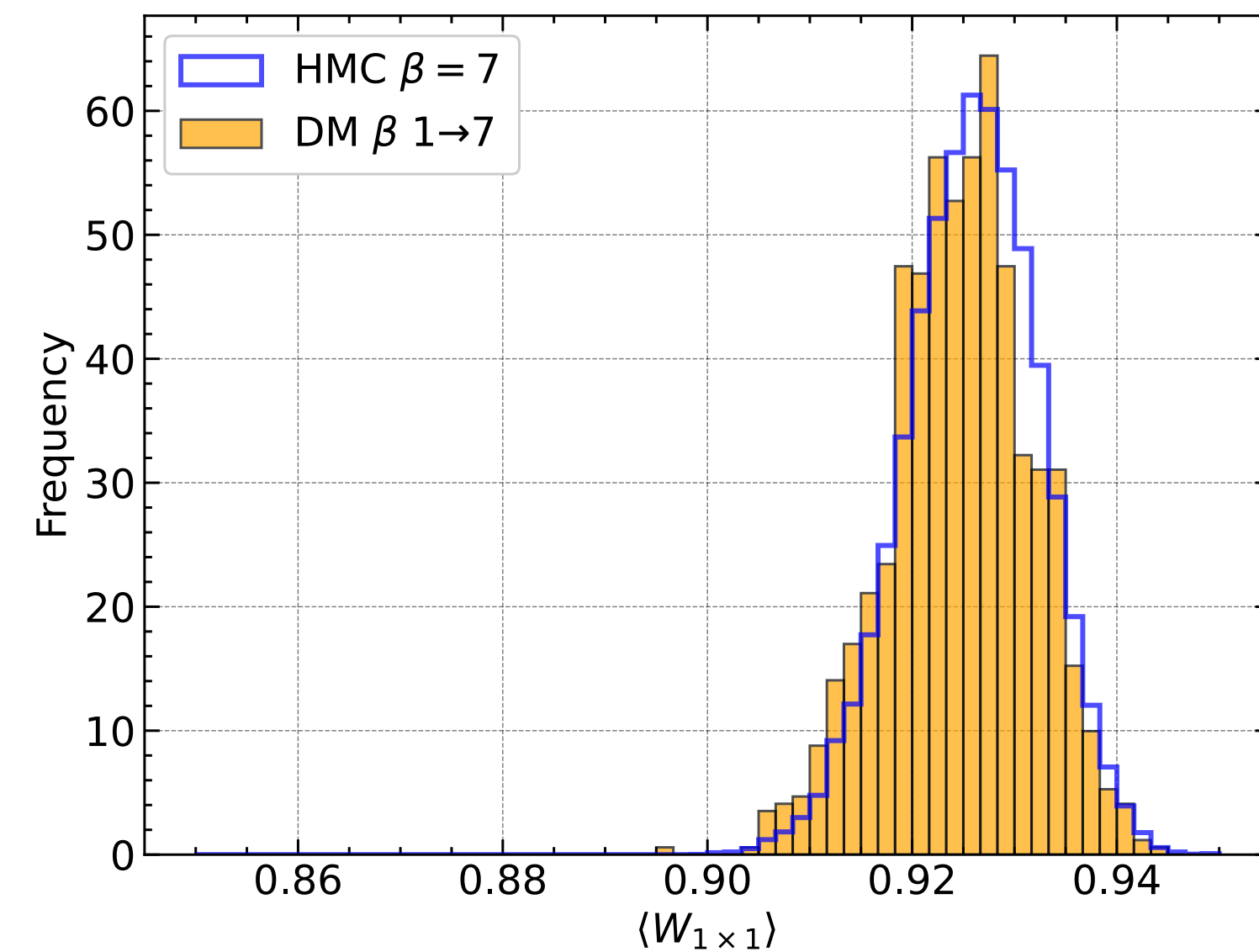


Gauge Fields

2D U(1) Gauge Field



Generated at $\beta = 7$ with **1024** configurations, $L = 16$

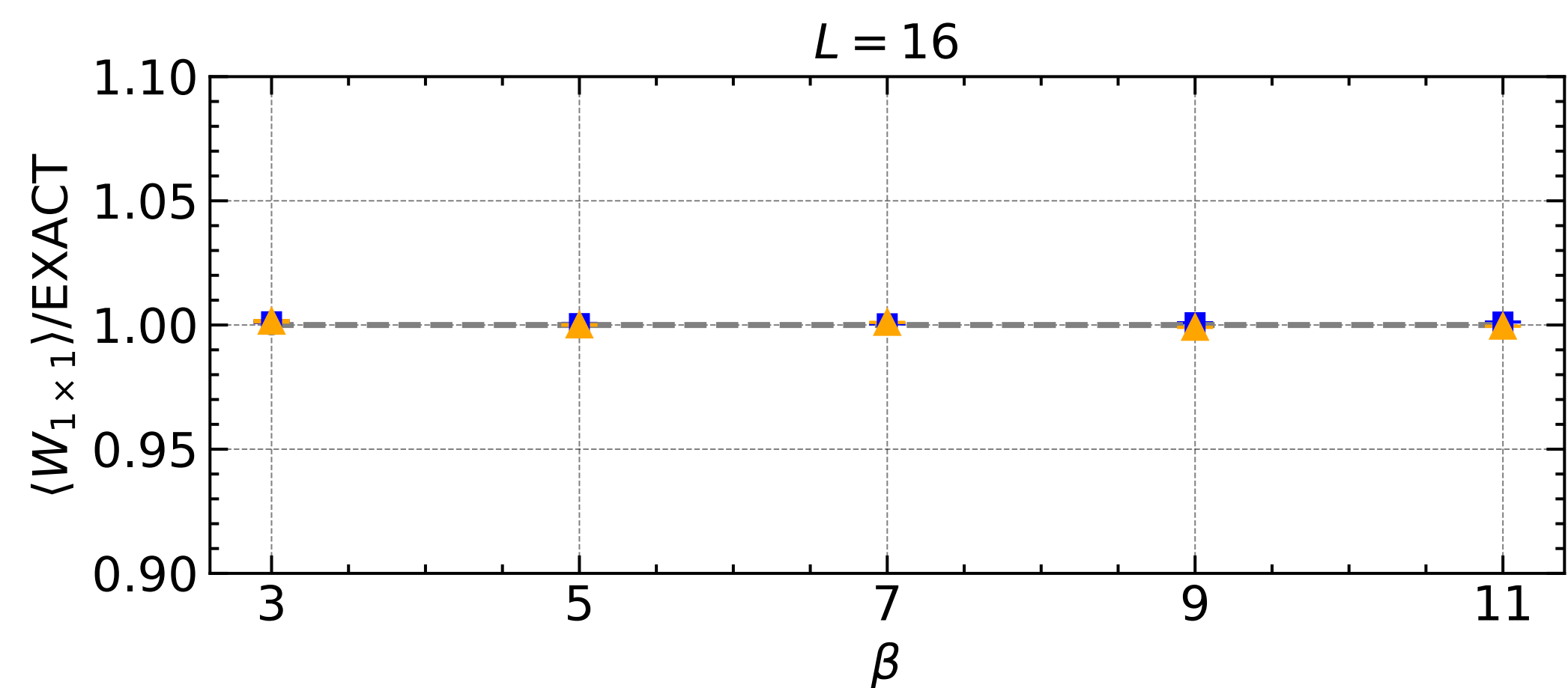
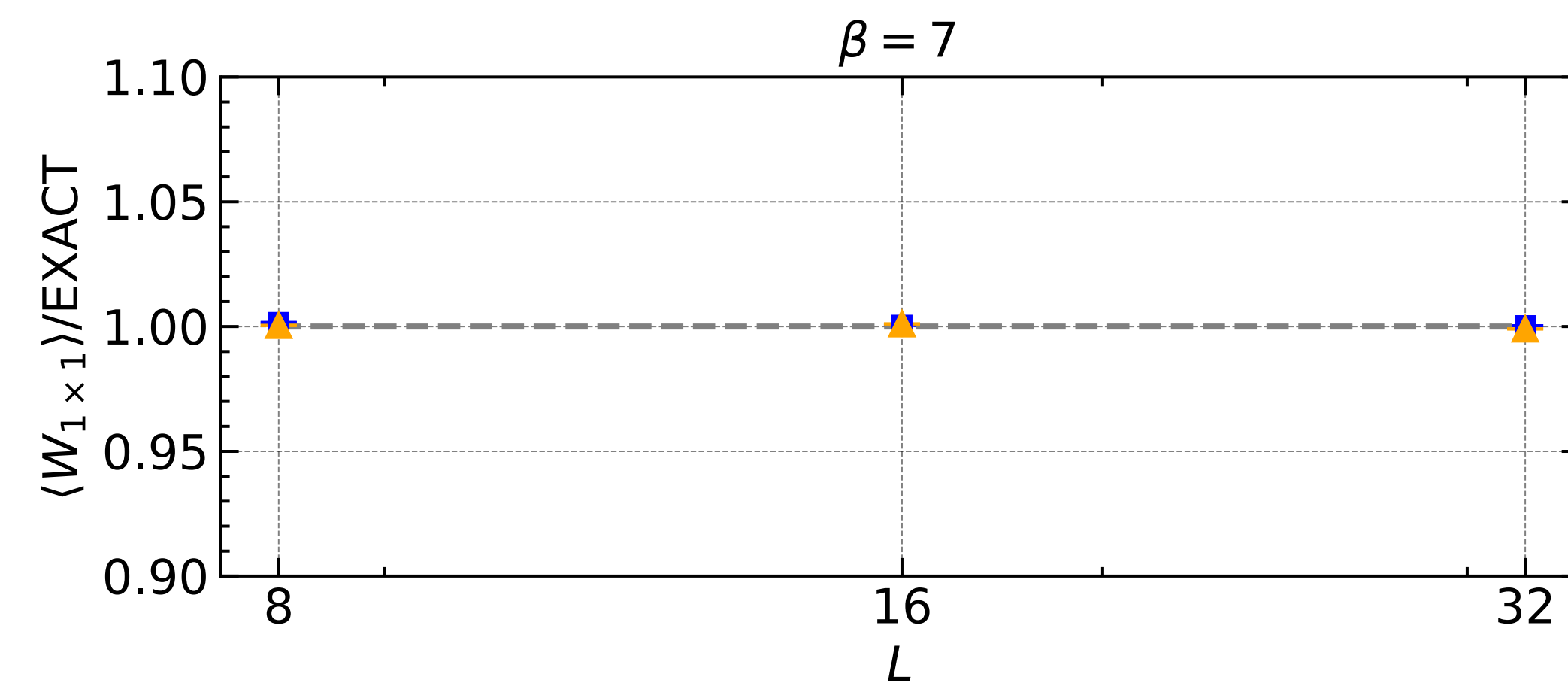
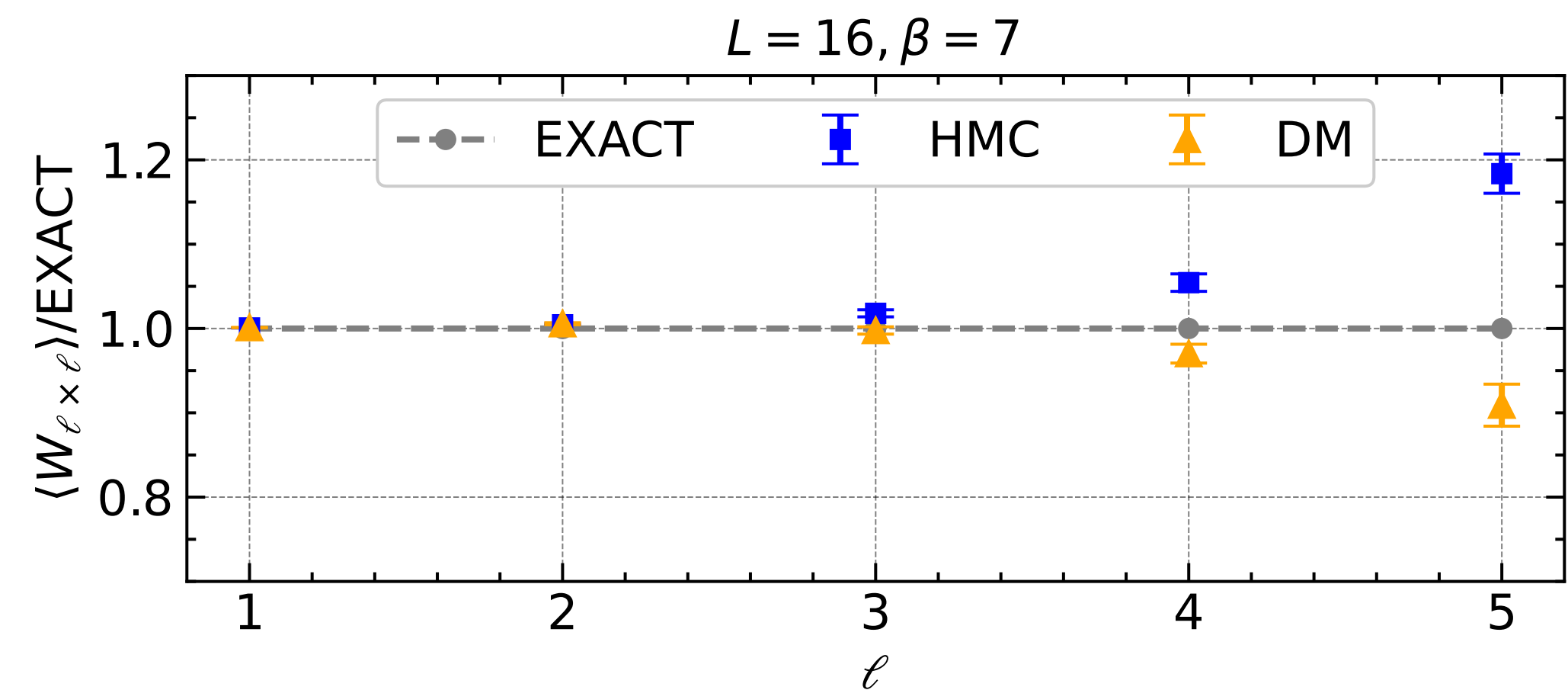
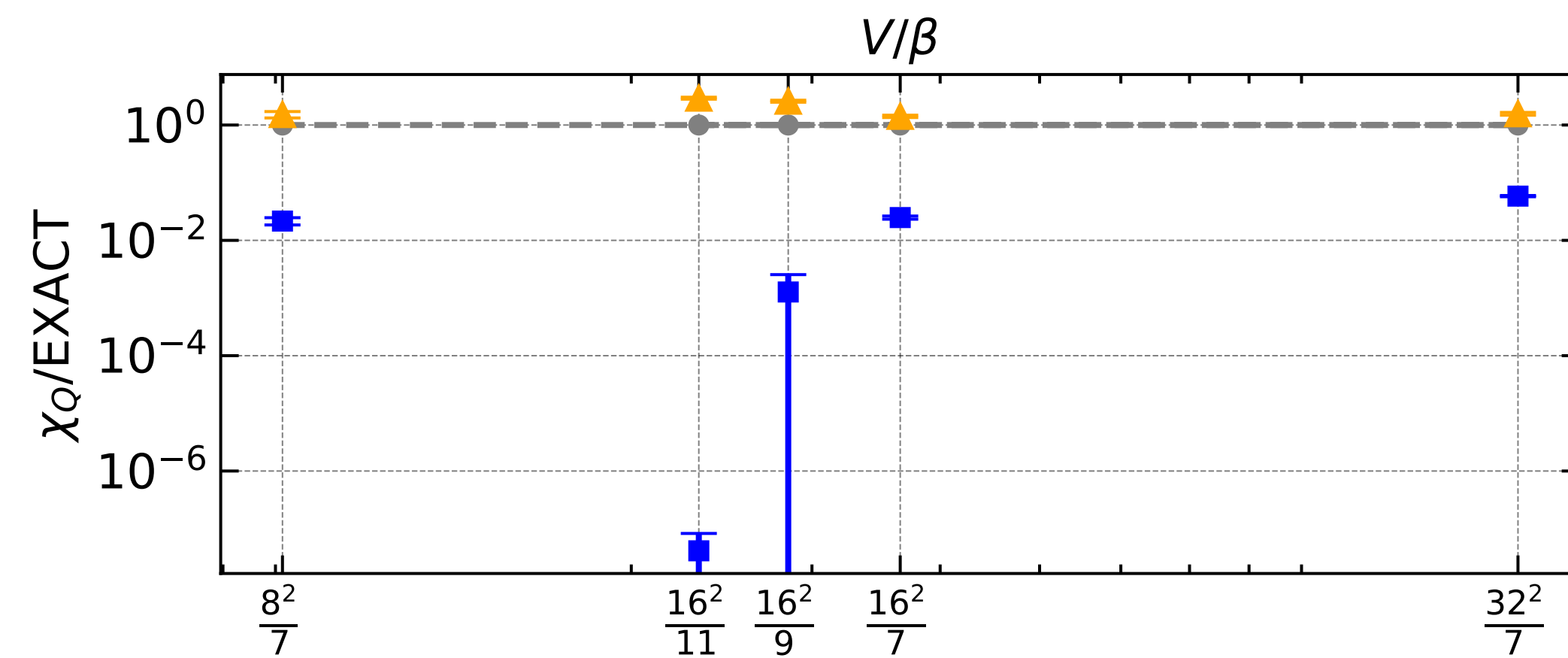


**Physics-Conditioned
Diffusion Models**

**Training at one,
Transfer and Generate at all**

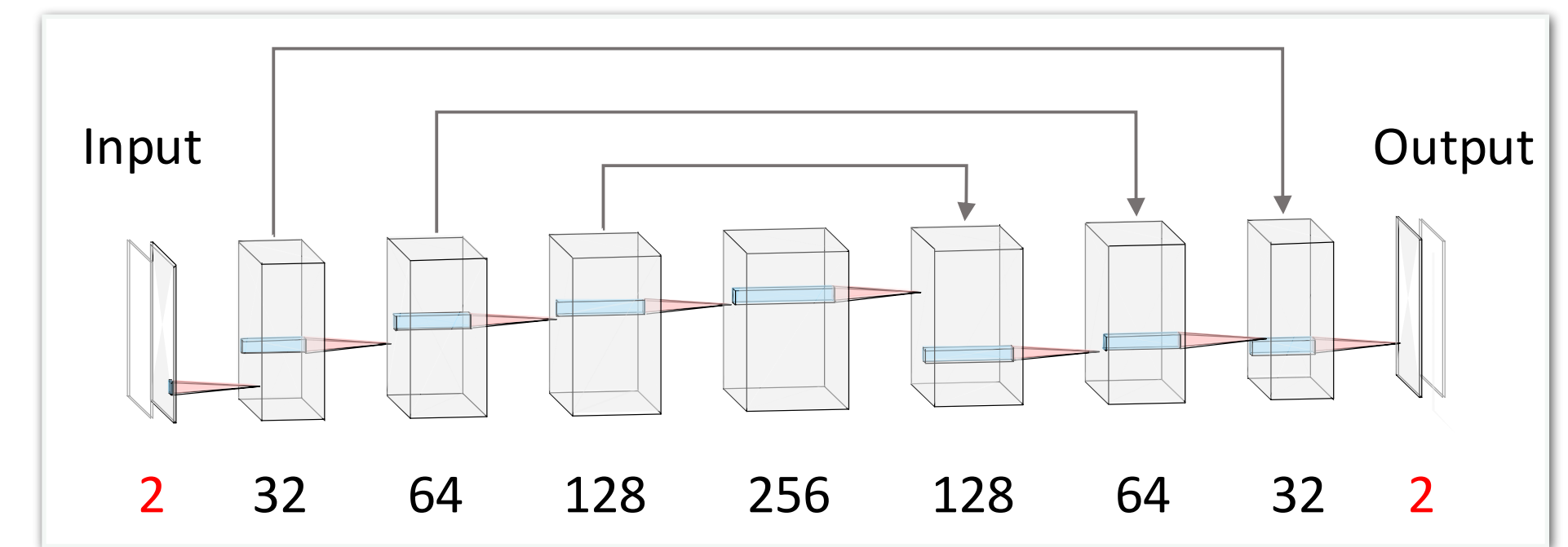
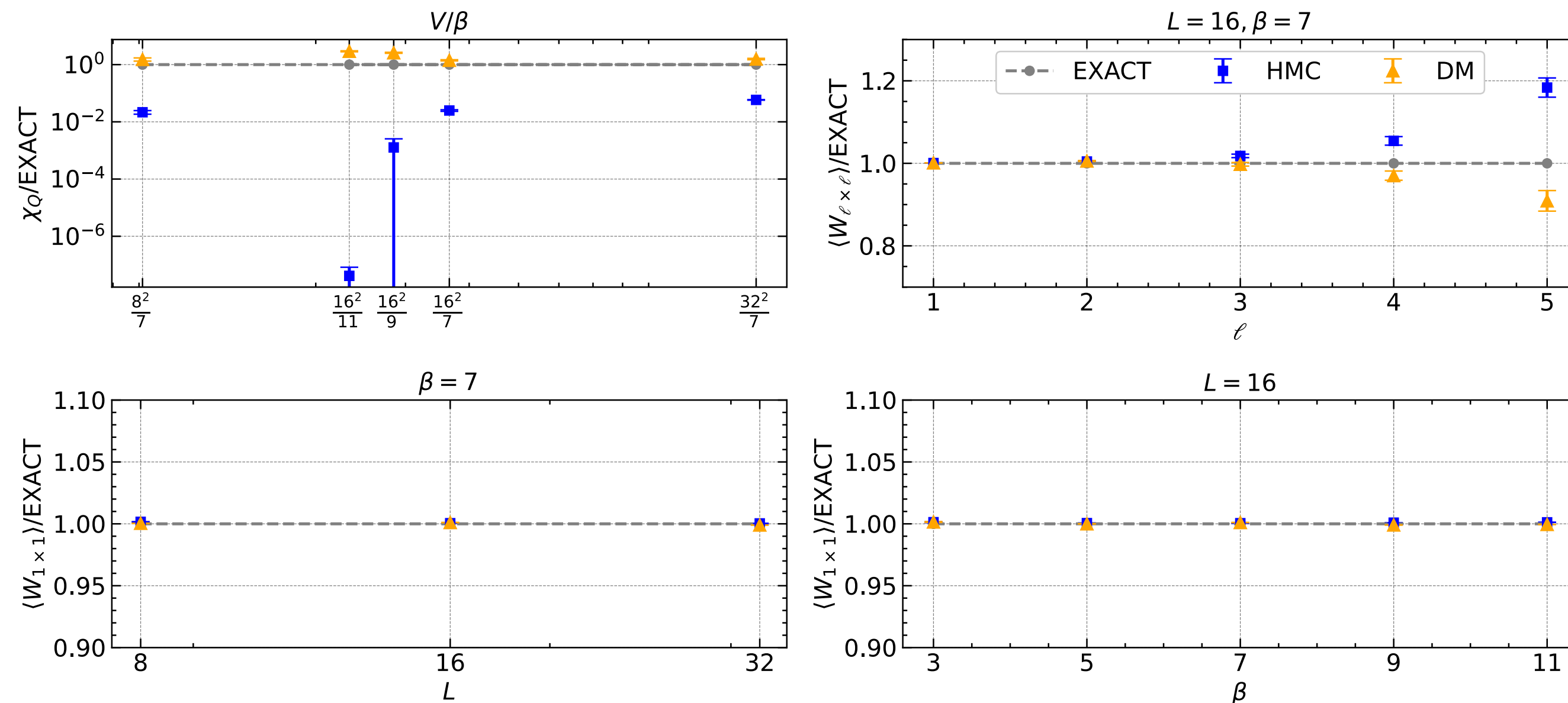
Gauge Fields

2D U(1) Gauge Field



Gauge Fields

2D U(1) Gauge Field



$$\tilde{s}_{\hat{\theta}}(\phi, t) \equiv \beta s_{\hat{\theta}}(\phi, t)$$

$$\frac{\partial \phi(x, \tau)}{\partial \tau} = - \frac{\delta S_E[\phi]}{\delta \phi(x, \tau)} + \eta(x, \tau)$$

Drift Term

Different Lattice Sizes, Inverse Coupling Constants

Training at one,
Transfer and Generate at all

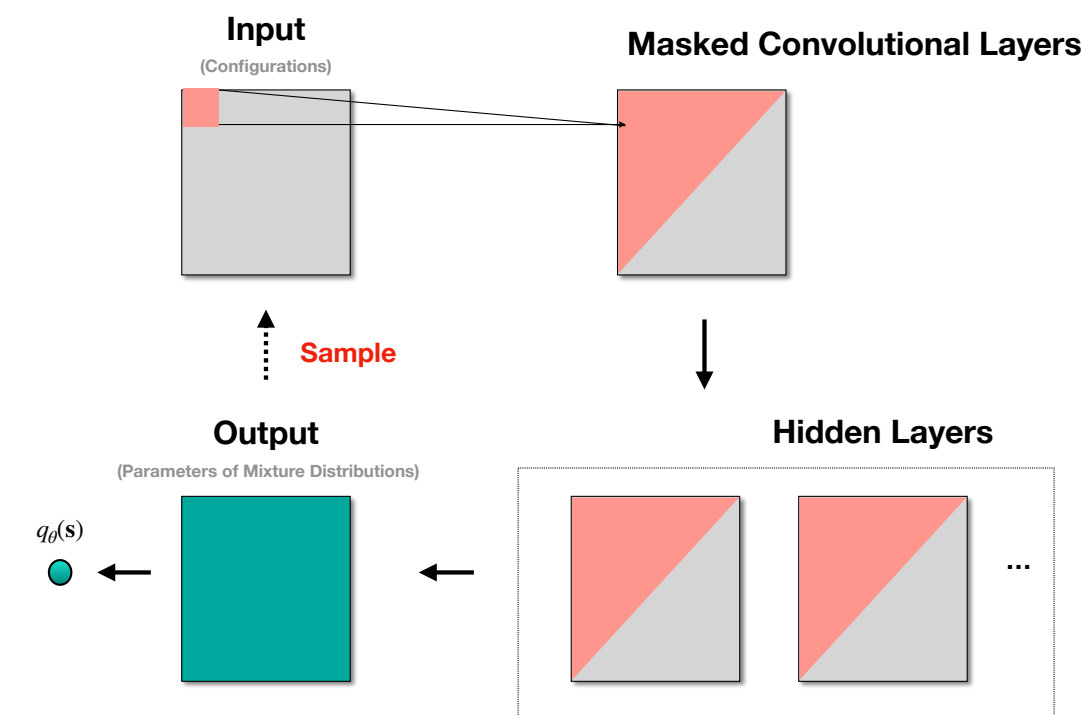
Summary II

- **Generative Models**

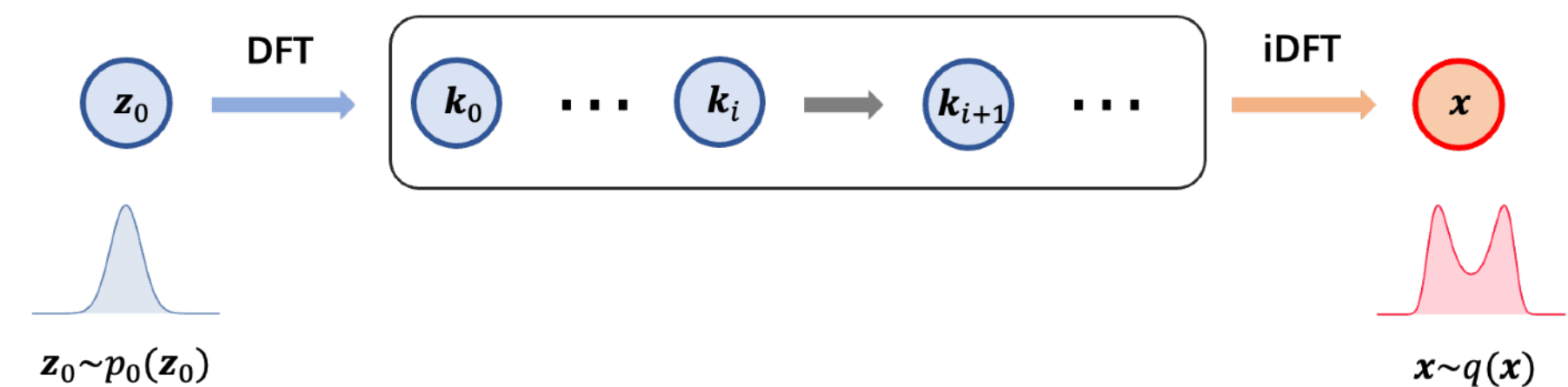
- Probabilistic Models
- Learn to Sample
- Learn to Detect Phase Transition
- Learn for Generating All

- **Future works**

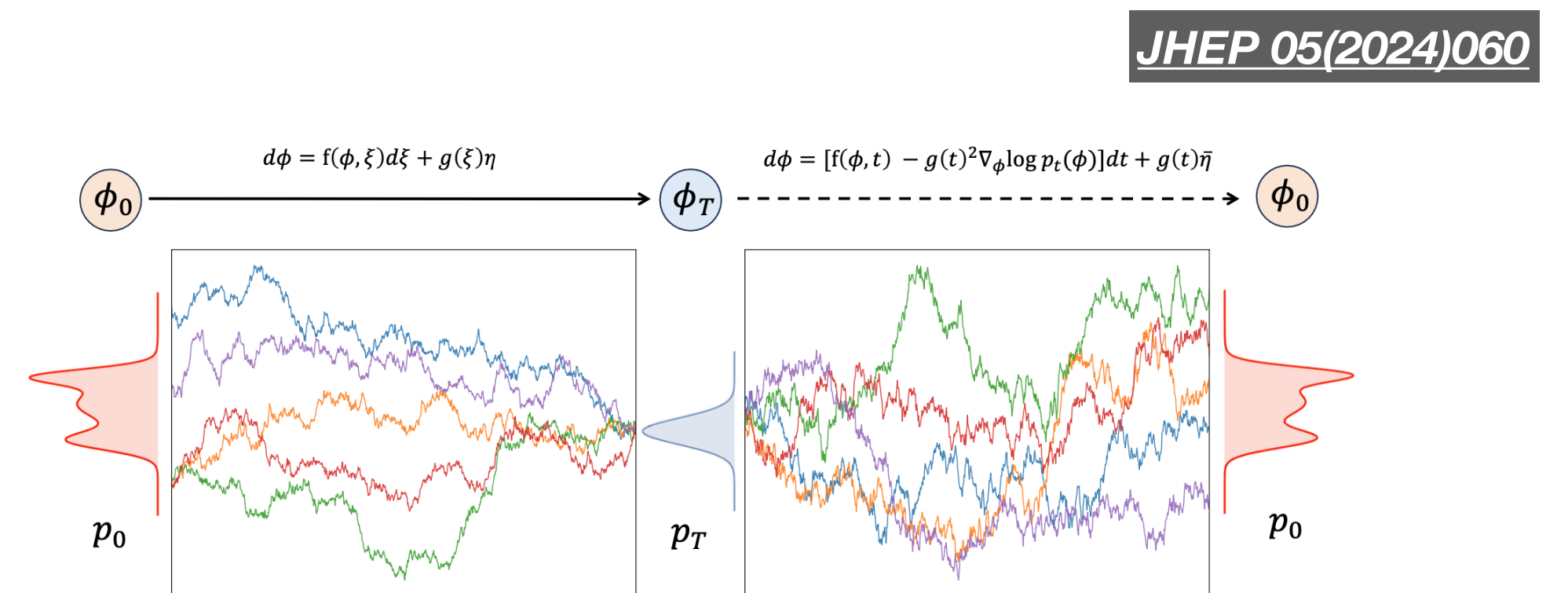
- 2D SU(2) Gauge Field
- Complex Langevin Method(CLM)
- DM for full QCD



Chinese Phys. Lett. 39, 120502 (2022)



Phys. Rev. D 107, 056001



JHEP 05(2024)060

Thank You !

ML meets Physics, Opportunities and Challenges



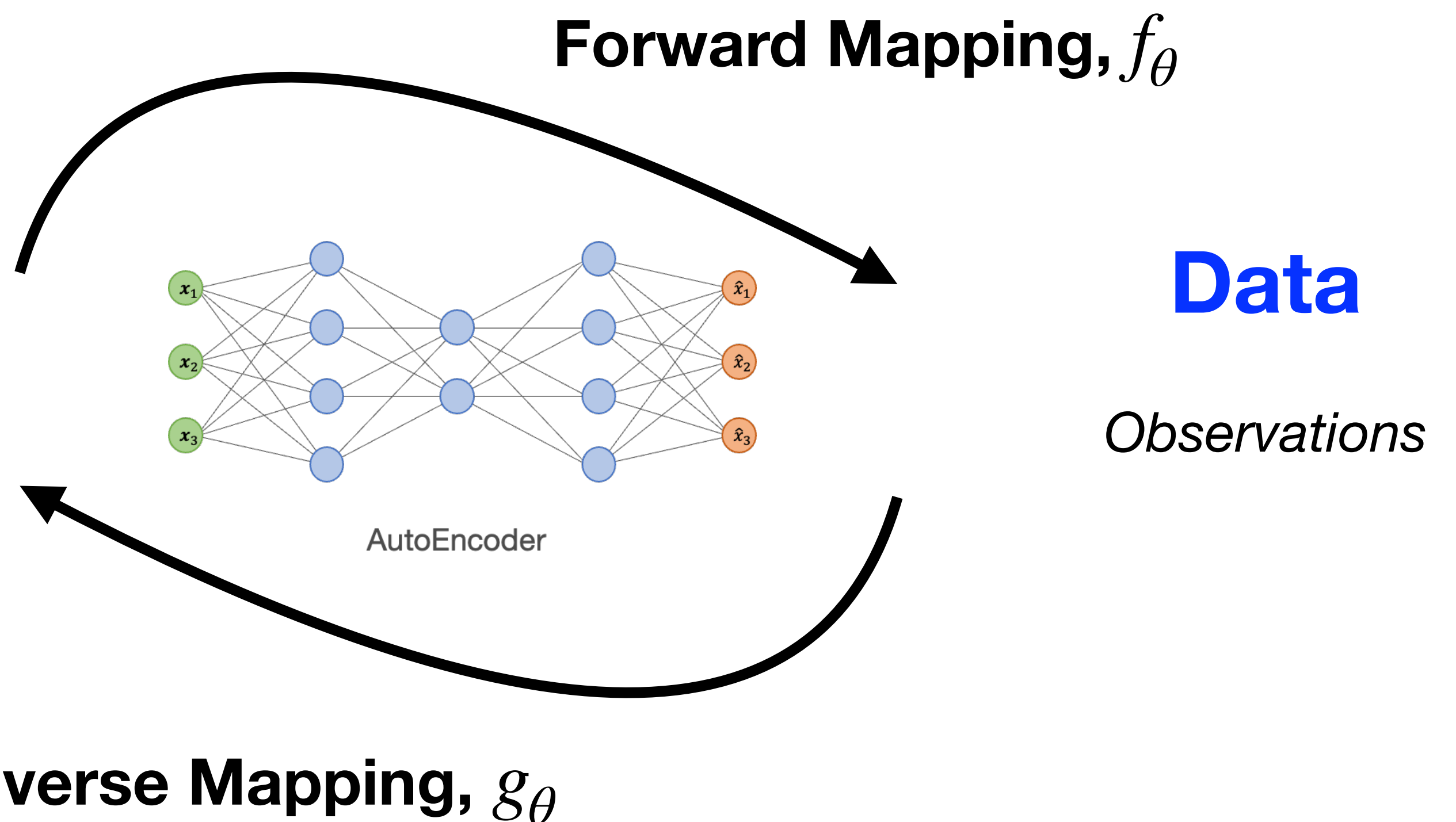
Representation Learning

$$g_{\theta} : X \rightarrow Y$$

$$f_{\theta} : Y \rightarrow X$$

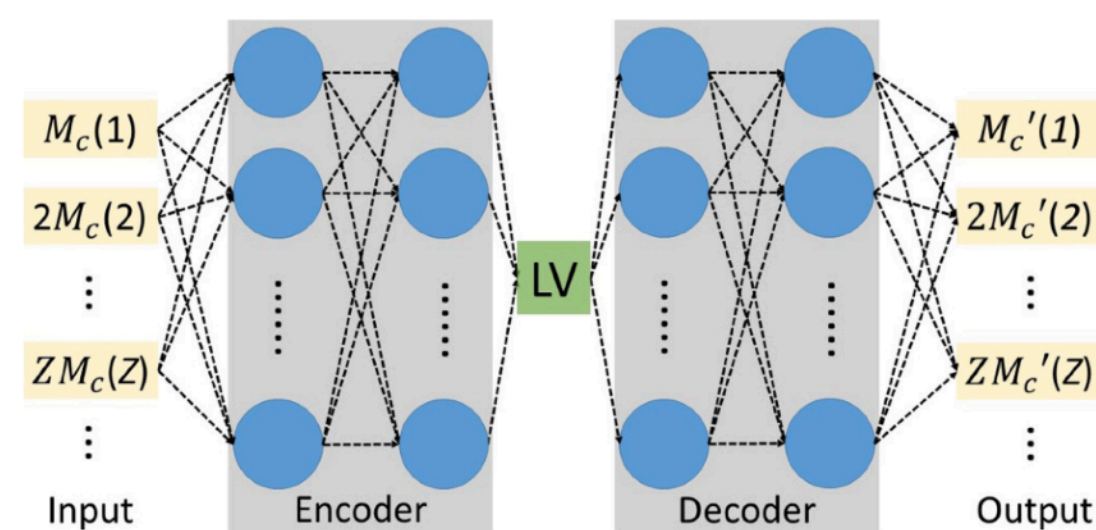
Physics

*Model Parameters/
Properties/States*

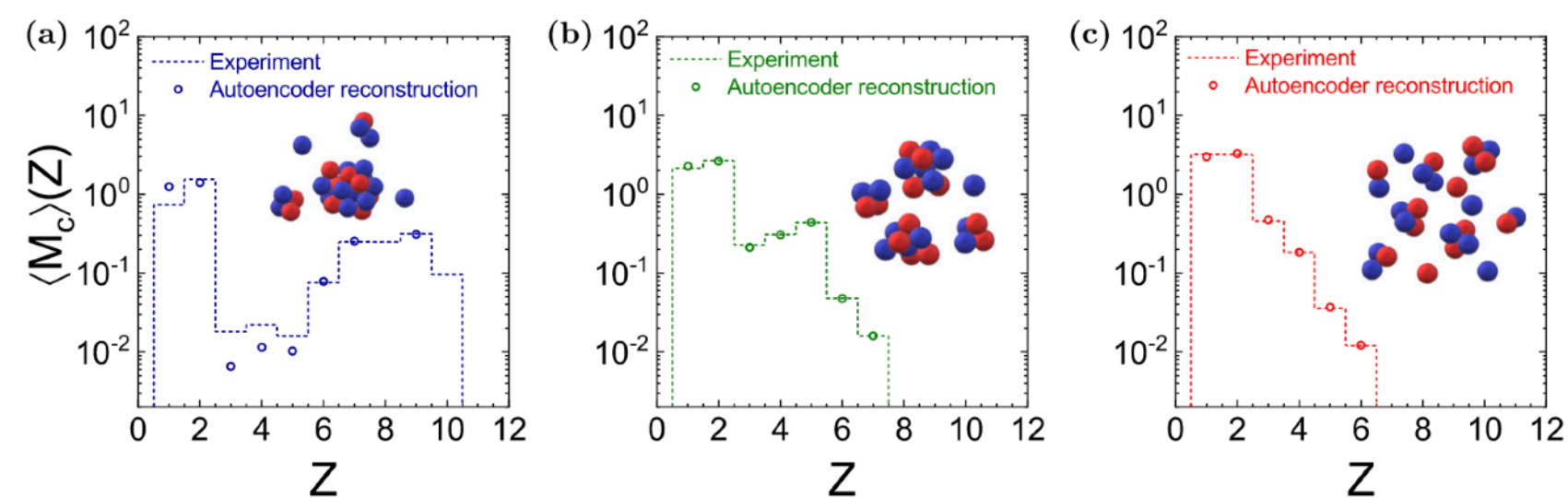


Representation Learning

Recognizing nuclear liquid-gas phase transition



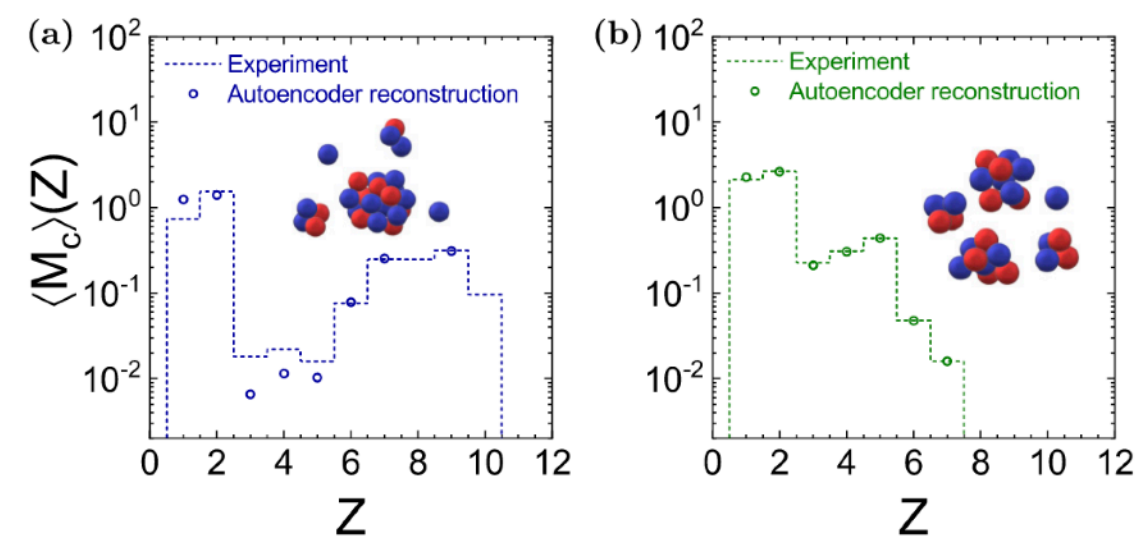
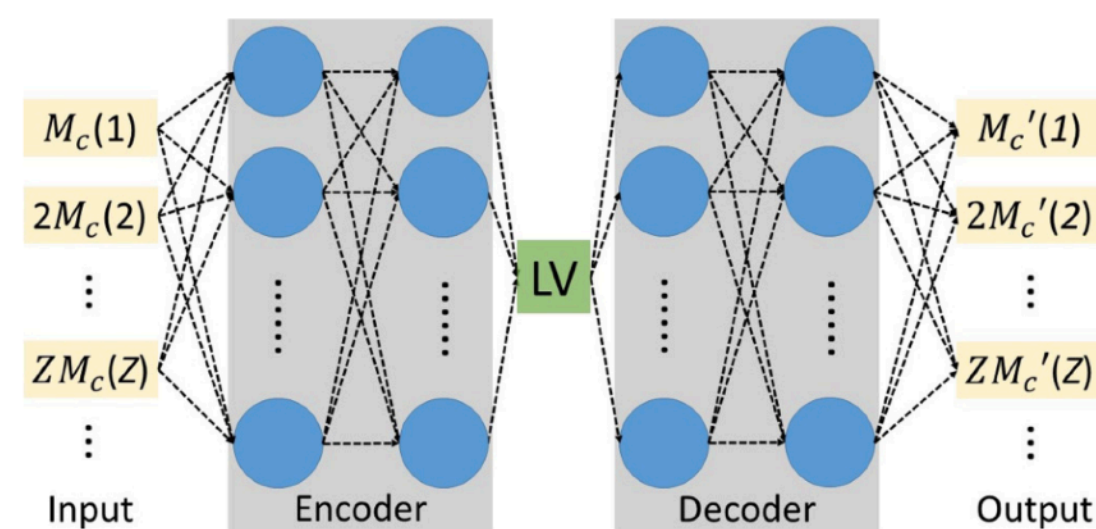
Anomaly Detection
Latent Variable Extraction



R. Wang, Y.-G. Ma, R. Wada, L.-W. Chen, W.-B. He, H.-L. Liu, and K.-J. Sun, Phys. Rev. Res. **2**, 043202 (2020)

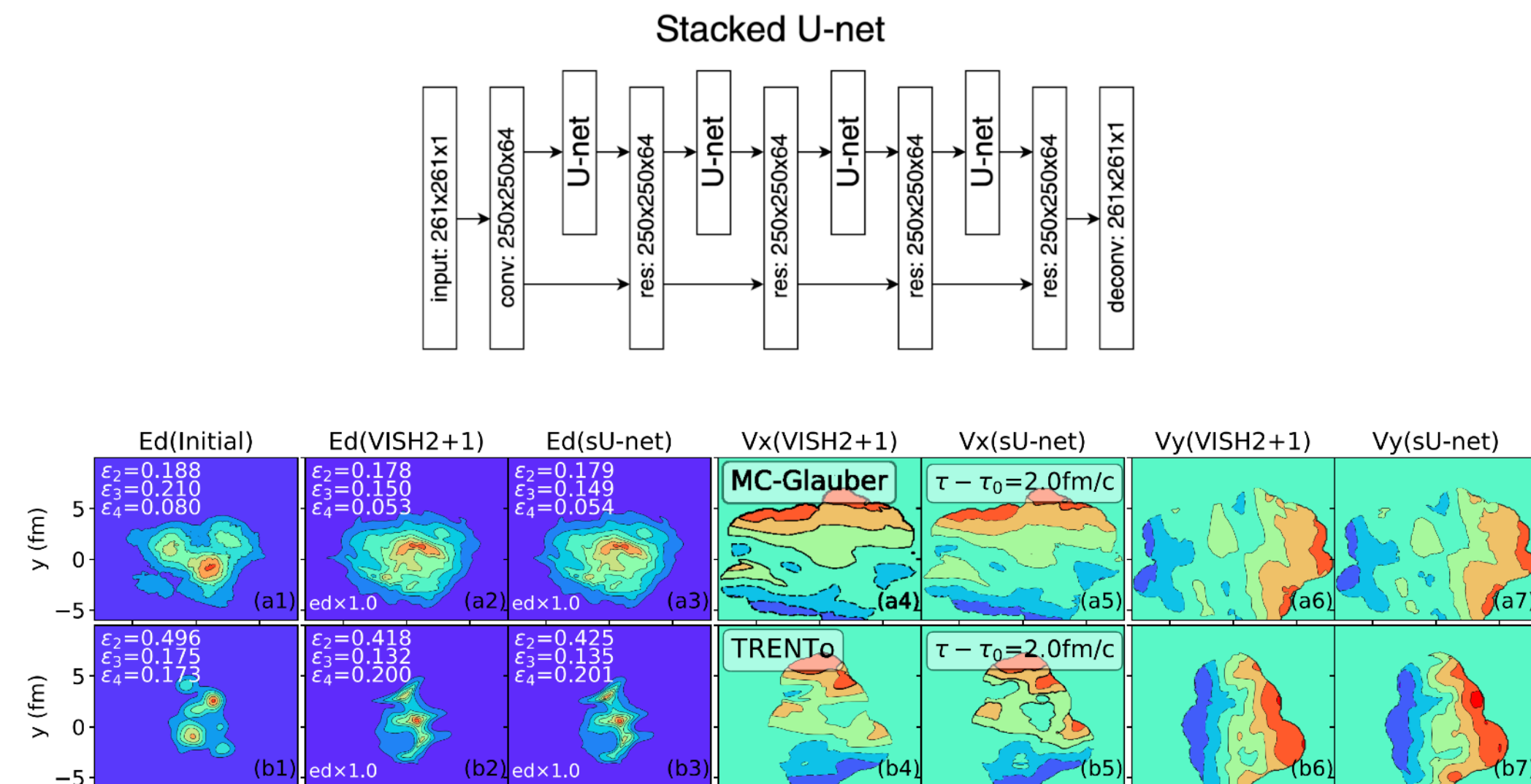
Representation Learning

Recognizing nuclear liquid-gas phase transition



R. Wang, Y.-G. Ma, R. Wada, L.-W. Chen, W.-B. He, H.-L. Liu, and K.-J. Sun,

Predict relativistic hydrodynamics



H. Huang, B. Xiao, Z. Liu, Z. Wu, Y. Mu, and H. Song, Phys. Rev. Res. **3**, 023256 (2021)

Backups I

Reconstructing Spectral Function

Comput. Phys. Commun. 282, 108547

- In practice, the Euclidean correlations have **finite number of points** and **with finite precision**;
- The ill-posedness of the spectral reconstruction **fundamentally exists even for continuous correlation functions(infinite observations)**;
- It's caused by the **numerical inaccuracy** of the correlation measurements (induced high degeneracy in solution space).

$$K_{ij}, i \in N_x, j \in N_\omega, N_x < N_\omega$$
$$\vec{D} \equiv K \vec{\rho} \quad \text{highly rectangular}$$

vectorization

$$D(x) \equiv \int_0^\infty K(x, \omega) \rho(\omega) d\omega$$

eigenvalue problem

$$\int_0^\infty \psi_s(\omega) K(x, \omega) d\omega = \lambda_s \psi_s(x)$$

J. Phys. A: Math. Gen., Vol. 11, No. 9, 1978. Printed in Great Britain.

Backups II: Autoregressive Networks

$$\max_{\theta} \prod_{i=1}^N p_{\theta}(s_i)$$

arXiv:2007.01037

1. Prepare data-set from **observations**

$$\mathbf{s} \sim q_{\text{data}}$$

2. Put them into the **deep autoregressive network(DAN)**

$$p_{\theta}(\mathbf{s}) = \prod_{i=1}^N p(s_i | s_1, \dots, s_{i-1}, \theta)$$

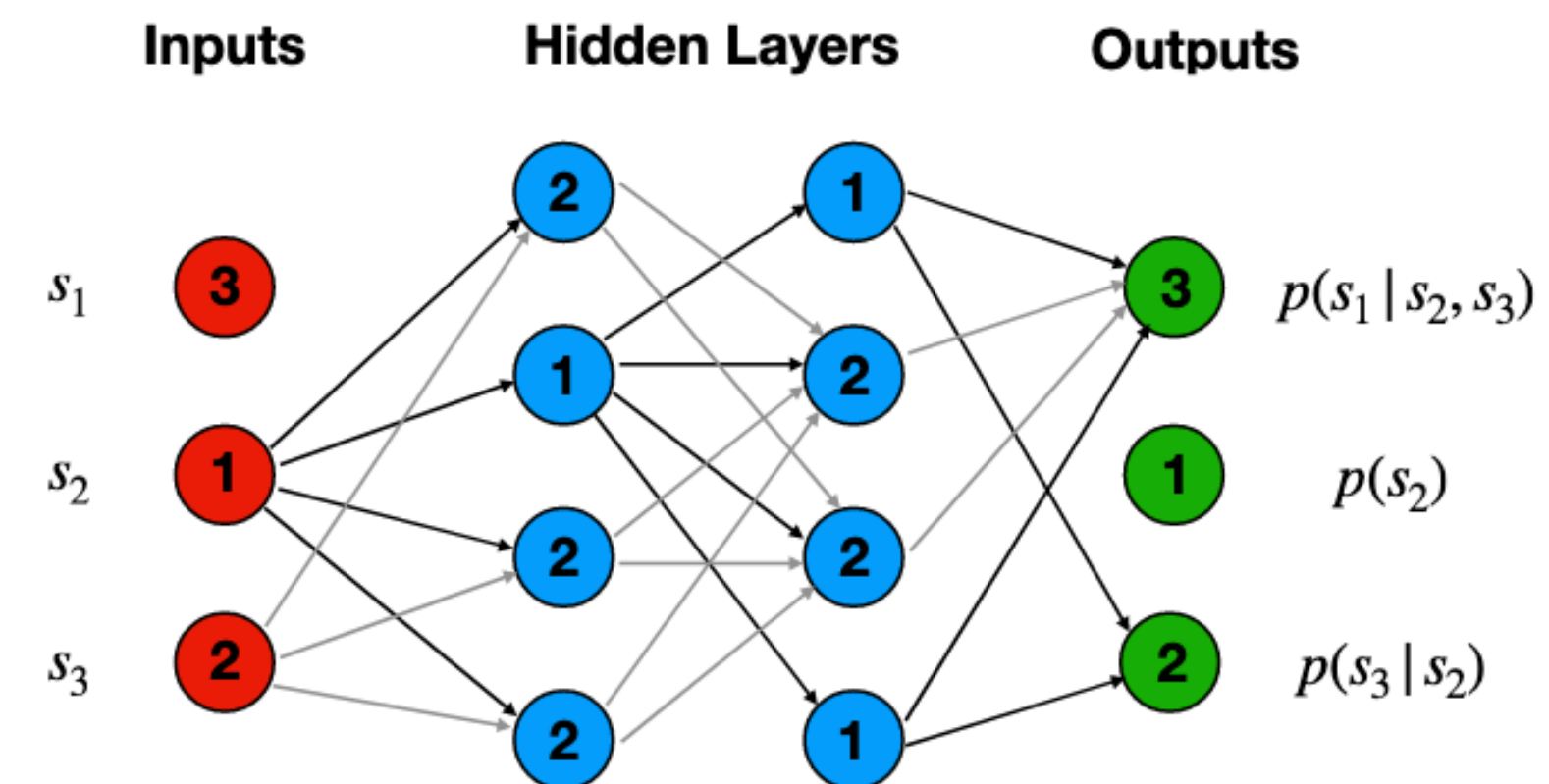
3. Minimize the negative log-likelihood(NLL)

$$\mathcal{L} = - \sum_{\mathbf{s} \sim q_{\text{data}}} \sum_{d=1}^N \log(p(s_d | \mathbf{s}_{<d}, \theta))$$

4. Get your DAN represented Hamiltonian

$$H_{\theta}(\mathbf{s}, T) = -T \ln p_{\theta}(\mathbf{s})$$

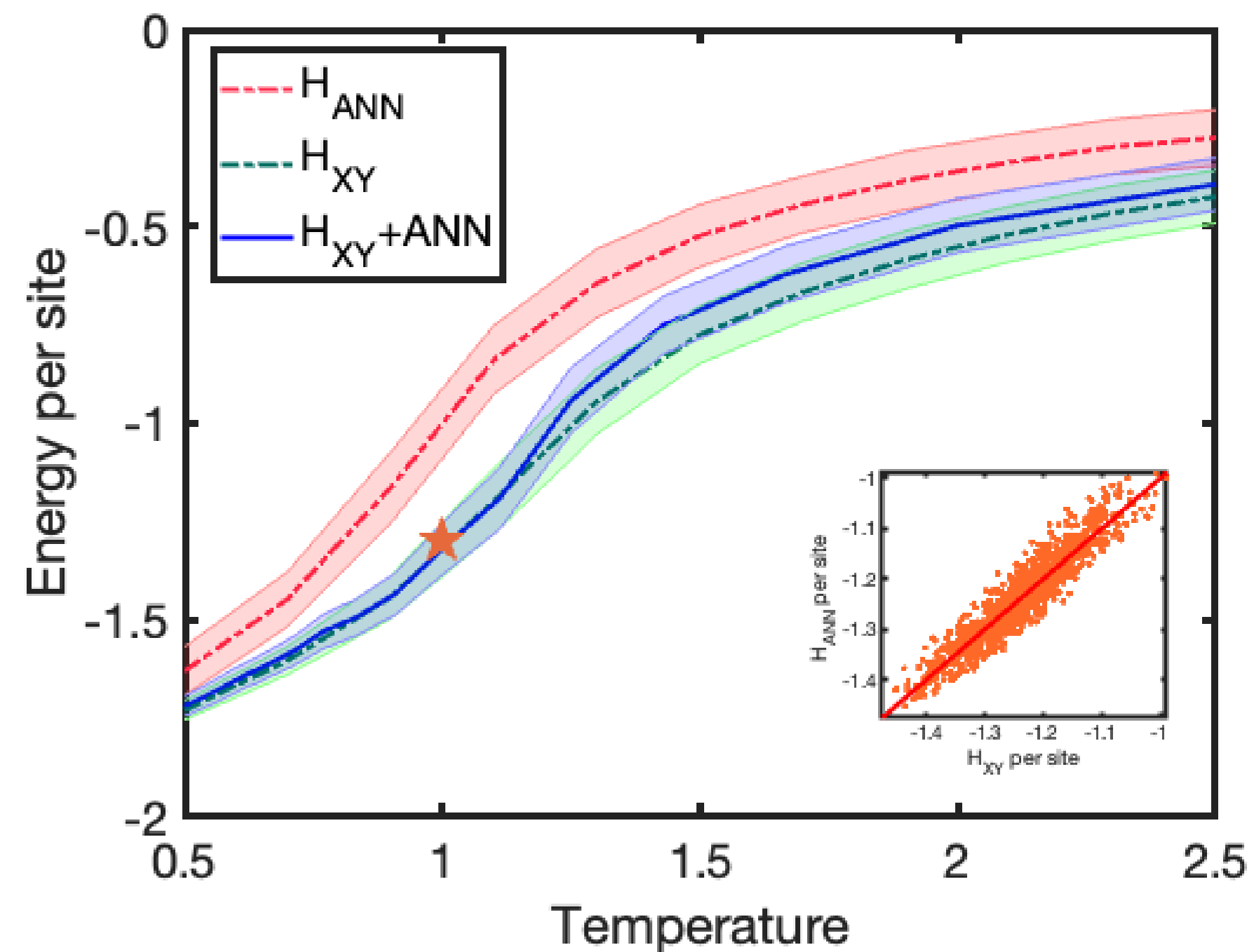
on specific degrees of freedom(*d.o.f.s*)



Backups II: Topological Phase Transition

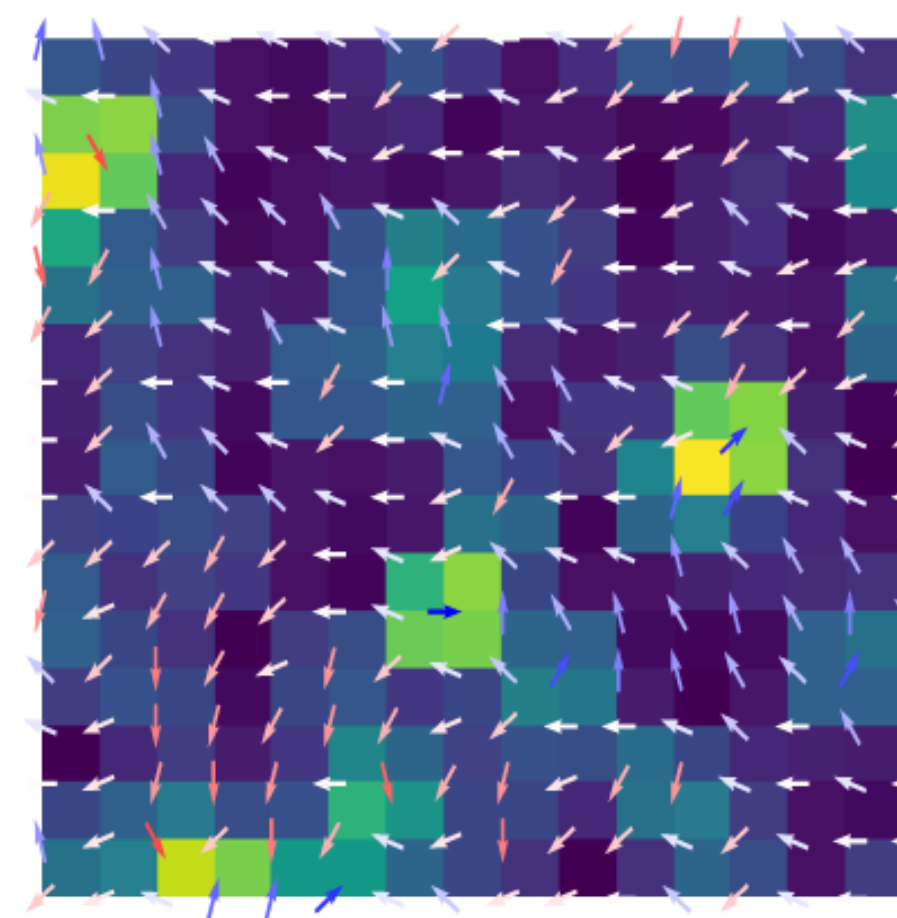
2D XY Model

$$H_{XY} = - \sum_{\langle i,j \rangle} s_i s_j = - \sum_{\langle i,j \rangle} \cos(\phi_i - \phi_j)$$

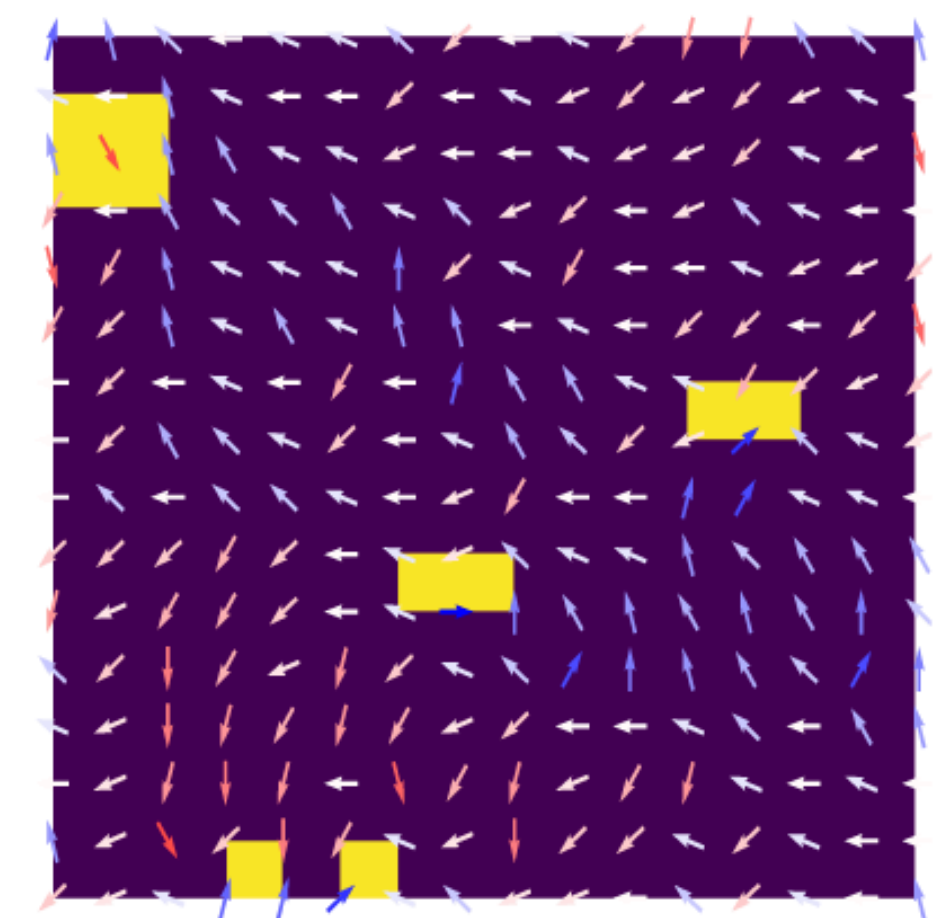


9000 training configurations with $L = 16$

Chinese Phys. Lett. 39, 120502 (2022)



Probability distributions from CANs



Vortices

$$\delta q_{[i,j]} \equiv \sum_{[i,j]} q_{\theta}(s_{ij})$$

$q_{\theta}(s_{ij})$ conditional probability differences in the same given direction

$$v = (1/2\pi) \oint_C \nabla \phi(\mathbf{r}) \cdot d\mathbf{r}$$

Diffusion Models

Adding noise

$$\mathbf{x} \sim p_{data}(\mathbf{x}) \equiv p(\mathbf{x})$$

$$\mathbb{E}_{p(\mathbf{x})}[\|\nabla_{\mathbf{x}} \log p(\mathbf{x}) - \mathbf{s}_{\theta}(\mathbf{x})\|_2^2]$$

Minimizing the Fisher divergence

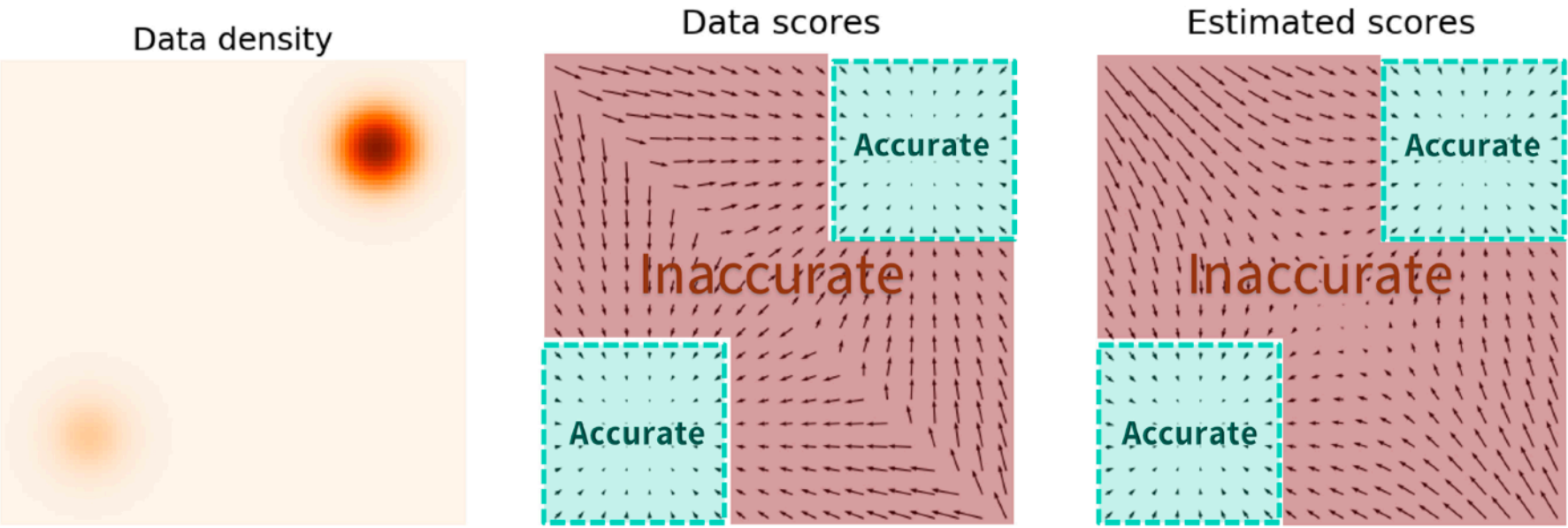
1-D

$$\left\{ \begin{aligned} \tilde{x} &= x + \sigma \epsilon \\ p_{\sigma}(\tilde{x}) &\equiv \int p_{\sigma}(\tilde{x} | x) p_{data}(x) dx \\ p_{\sigma}(\tilde{x} | x) &\sim N(\tilde{x}; x, \sigma^2 I), \epsilon \sim N(0, I) \\ \frac{\partial \log(p_{\sigma}(\tilde{x}))}{\partial \tilde{x}} &\equiv \frac{\partial \log(p_{\sigma}(\tilde{x} | x) p_{data}(x))}{\partial \tilde{x}} = \frac{\partial \log(p_{\sigma}(\tilde{x} | x))}{\partial \tilde{x}} \end{aligned} \right.$$

Score-matching a noise-perturbed distribution

$$\mathbb{E}_{p_{\sigma}(\mathbf{x})}[\|\nabla_{\mathbf{x}} \log p_{\sigma}(\mathbf{x}) - \mathbf{s}_{\theta}(\mathbf{x})\|_2^2]$$

Matching a noise-perturbed distribution



No enough data in low-density region!
But we sample from the low-density region using Langevin dynamics...



Diffusion Models

$$\mathbb{E}_{p_{\sigma}(\mathbf{x})}[\|\nabla_{\mathbf{x}} \log p_{\sigma}(\mathbf{x}) - \mathbf{s}_{\theta}(\mathbf{x})\|_2^2]$$

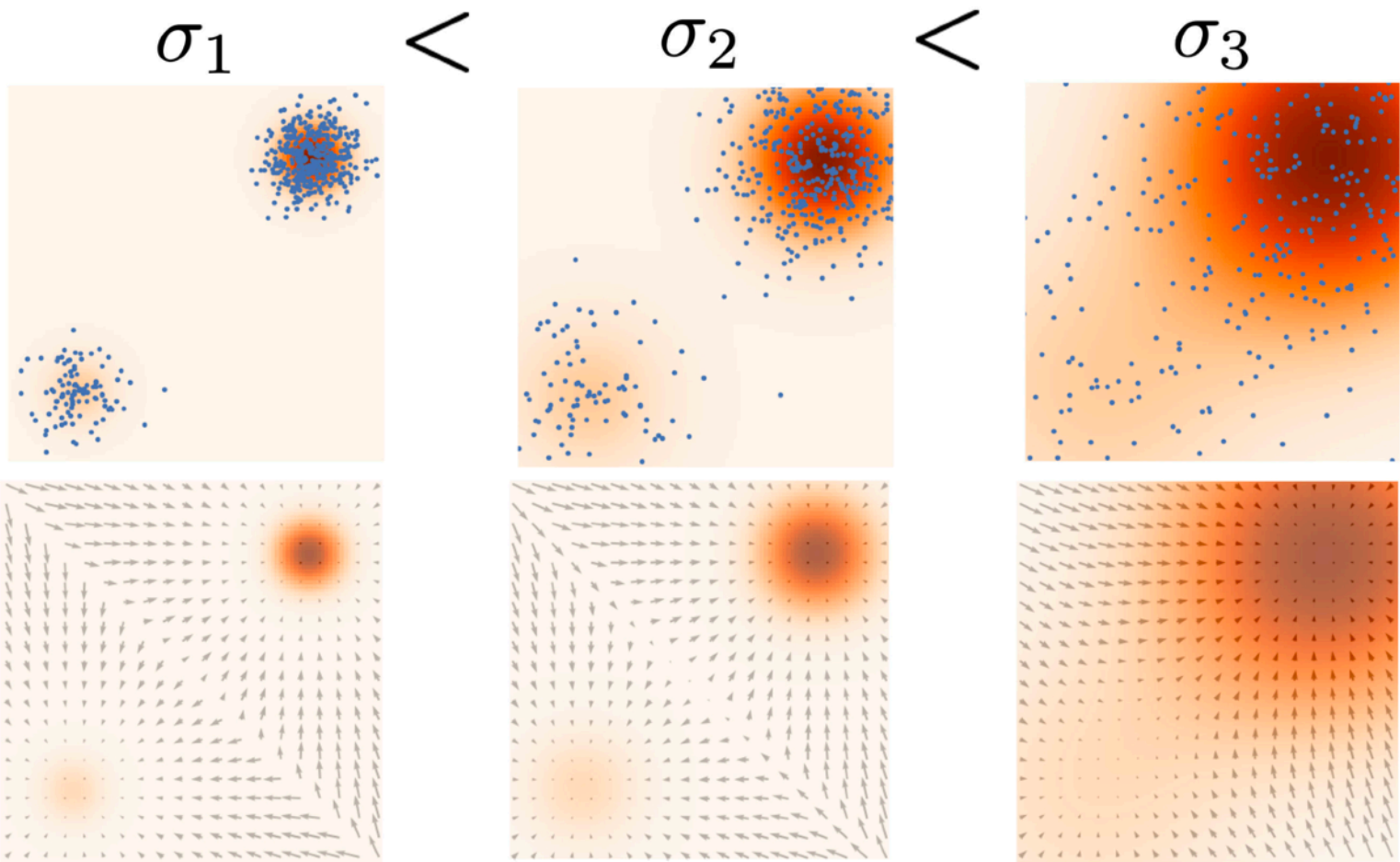
Matching a noise-perturbed distribution

Small noise → Good approximation to data, poor in low-density region!
Large noise → Poor approximation to data, good in low-density region!

$$\sum_{i=1}^L \lambda_i \mathbb{E}_{p_{\sigma_i}(\mathbf{x})}[\|\nabla_{\mathbf{x}} \log p_{\sigma_i}(\mathbf{x}) - \mathbf{s}_{\theta}(\mathbf{x}, i)\|_2^2]$$

Matching multi-noise perturbed distributions

Multiple noise perturbations! 🎉



- Choose the noise scheme as a geometric progression, $\sigma_1/\sigma_2 = \sigma_{i-1}/\sigma_i = \dots = \sigma_{L-1}/\sigma_L > 1$, with σ_1 being sufficiently small and σ_L comparable to the maximum pairwise distance between all training data points. L is typically on the order of hundreds or thousands.
- Parameterize the score-based model, $\mathbf{s}_{\theta}(\mathbf{x}, i)$, with U-Net skip connections.

Multiple scales of Gaussian noise to perturb the data distribution (first row), and jointly estimate the score functions for all of them (second row).

[Ho et al., Denoising Diffusion Probabilistic Models, NeurIPS 2020](#)

[Song et al., Score-Based Generative Modeling through Stochastic Differential Equations, ICLR 2021](#)

Diffusion Models

$$\sum_{i=1}^L \lambda_i \mathbb{E}_{p_{\sigma_i}(\mathbf{x})} [\|\nabla_{\mathbf{x}} \log p_{\sigma_i}(\mathbf{x}) - \mathbf{s}_{\theta}(\mathbf{x}, i)\|_2^2]$$

Matching multi-noise perturbed distributions

$$x_{i+1} = x_i + \sigma_i \epsilon$$

$$\begin{array}{l} \frac{d\mu}{dt} = 0 \\ \frac{d\sigma^2}{dt} = \sigma^{2t} \\ t \in [0, \tau] \end{array}$$

$$\nabla_{\mathbf{x}} \log p_{\sigma_i}(\mathbf{x}) = \nabla_{\mathbf{x}} \log(p_{\sigma_i}(\mathbf{x}_{\tau} | \mathbf{x}) p(\mathbf{x})) = \nabla_{\mathbf{x}} \log p_{\sigma_i}(\mathbf{x}_{\tau} | \mathbf{x})$$

$$p_{\sigma_i}(\mathbf{x}_{\tau} | \mathbf{x}) = \mathcal{N}\left(\mathbf{x}_{\tau}; \mathbf{x}, \frac{1}{2 \log \sigma} (\sigma^{2\tau} - 1) \mathbf{I}\right)$$

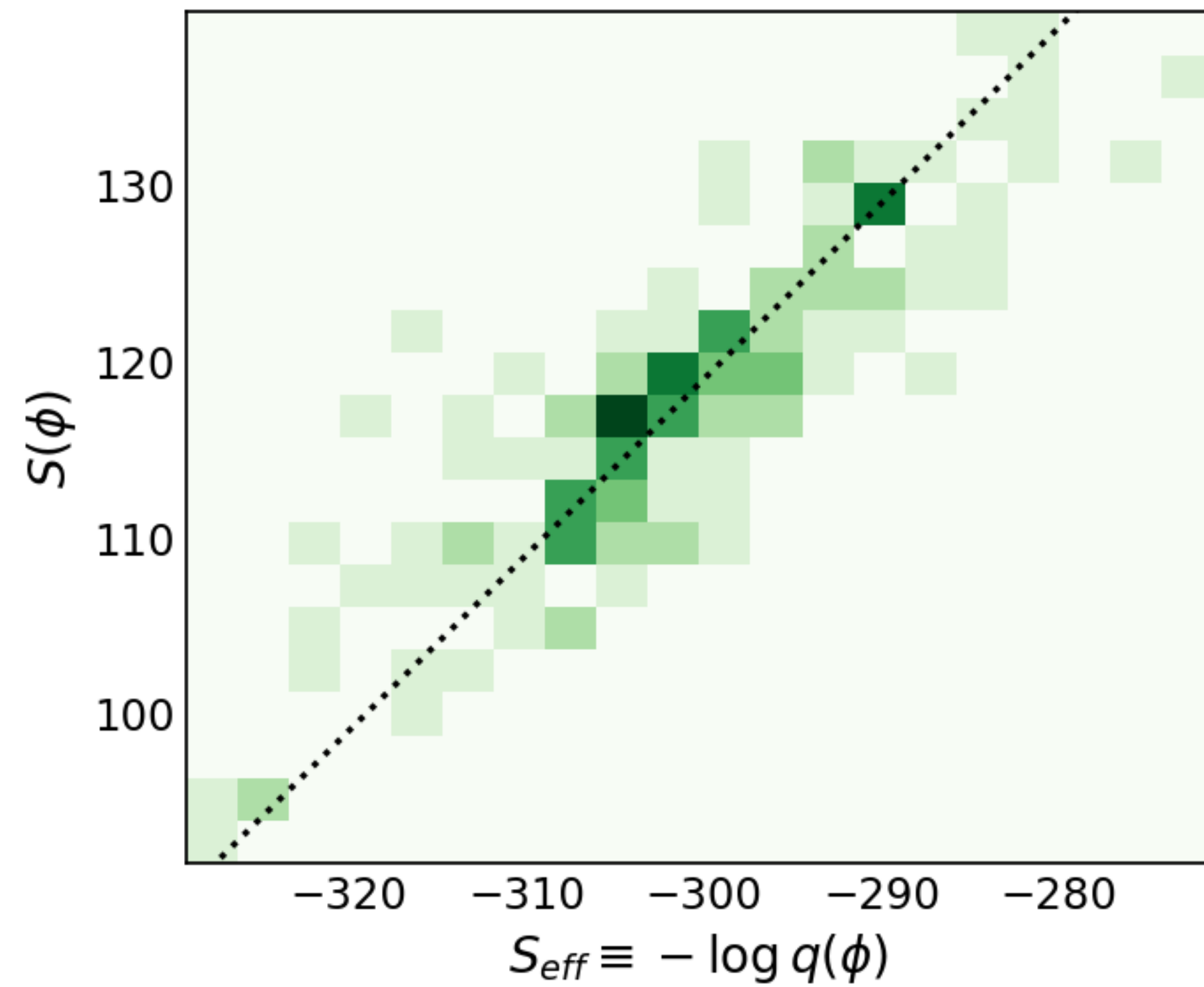
- Choose the noise scheme as, $\sigma_i \equiv \sigma^{\tau}$, where τ indicates “time-step” for adding noise.

H. Risken, The Fokker-Planck Equation: Methods of Solution and Applications

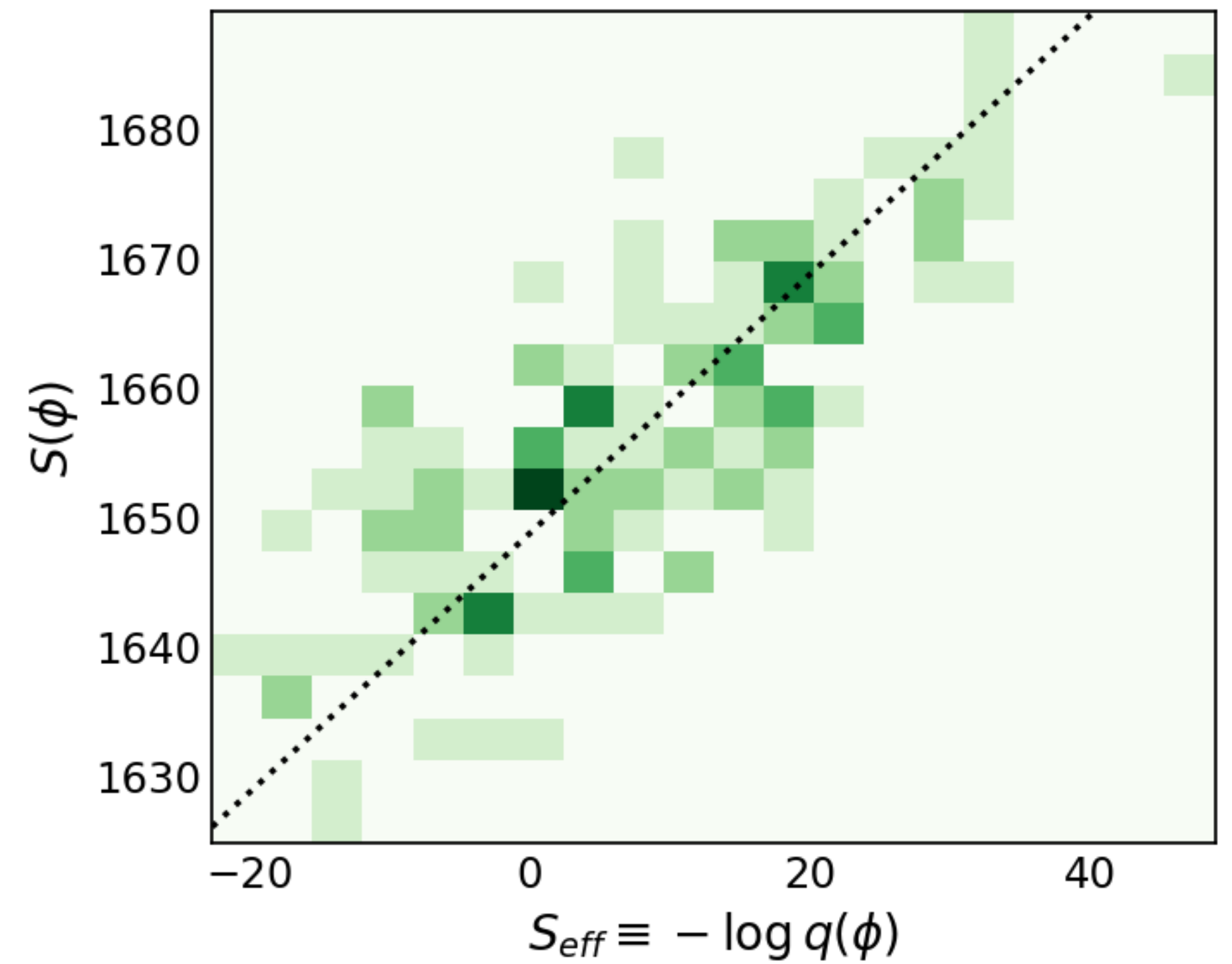
Training is Matching

Backups III: Likelihood in DM

$$\beta = 1$$



$$\beta = 7$$



Learned at $\beta = 1$ with **10,240** configurations, $L = 16$
Generated 128 configs for estimating Likelihood