



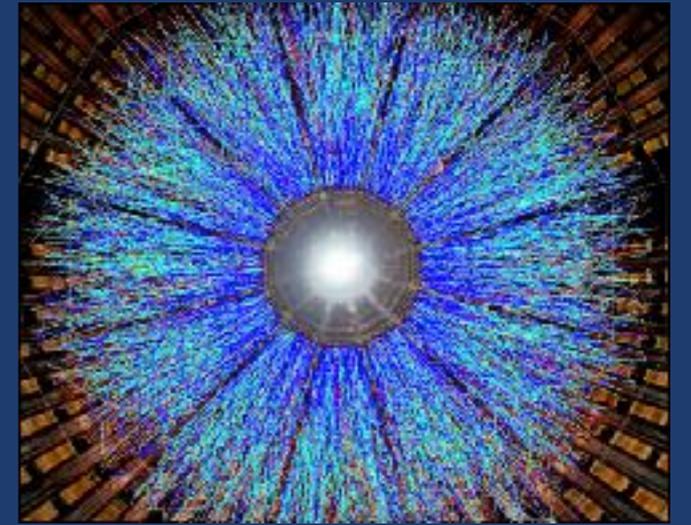
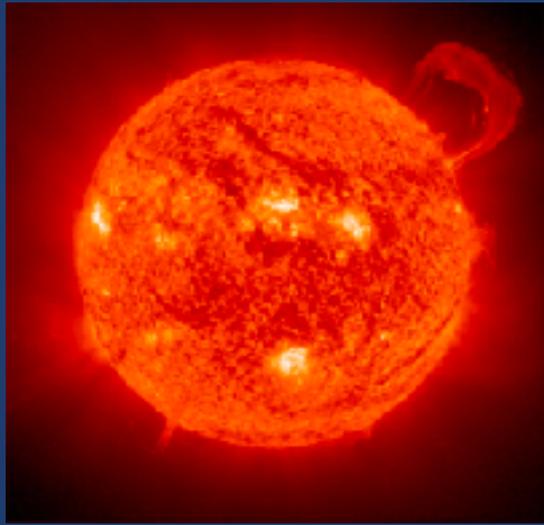
SAŠO GROZDANOV



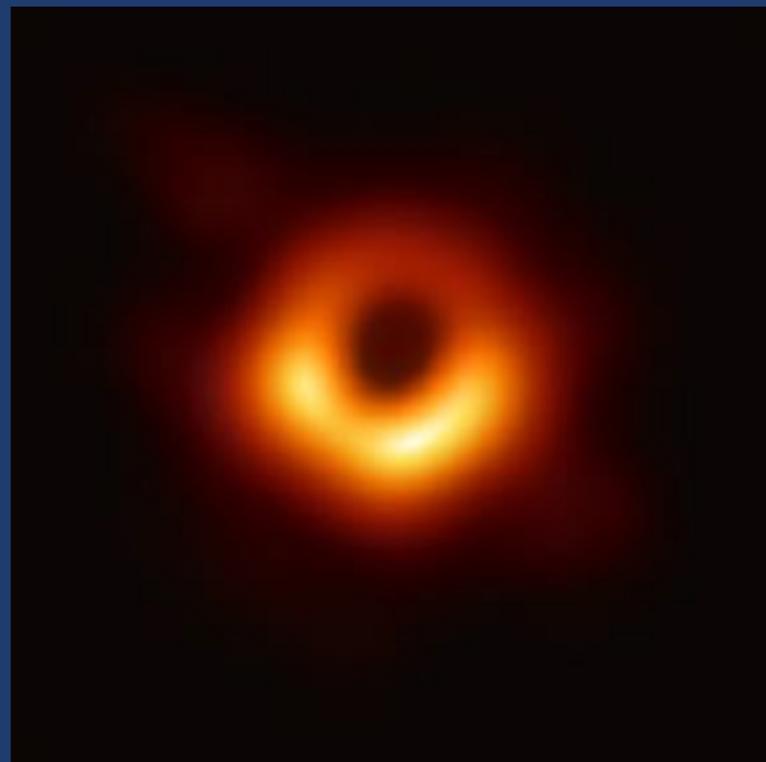
# THE SPECTRAL DUALITY RELATION

BEIJING, 15.7.2025

# THERMAL FIELD THEORY AND BLACK HOLES

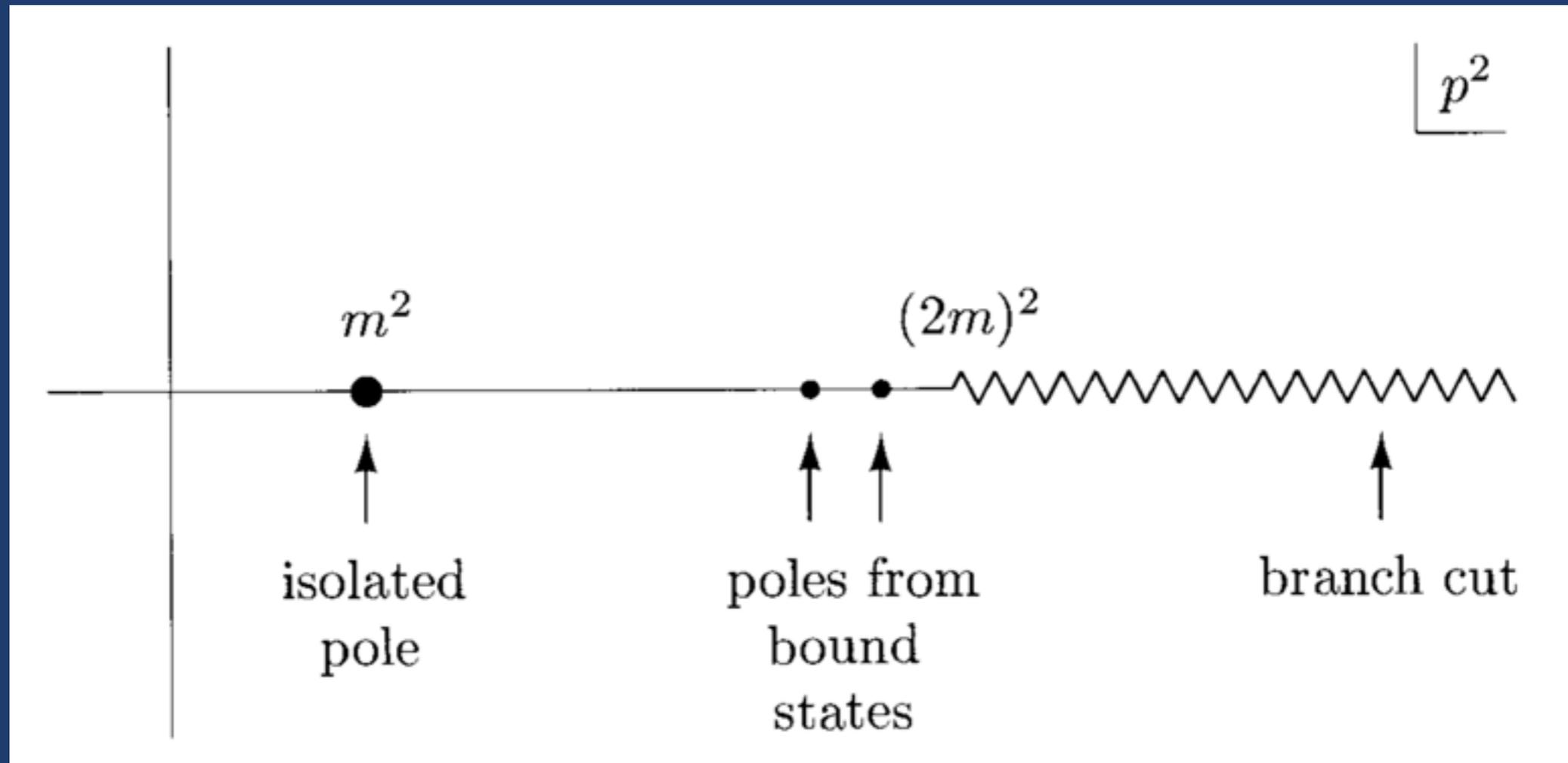


$$Z[\beta = 1/T] = \int \mathcal{D}\Phi e^{-\beta H} e^{\frac{i}{\hbar} \int d^d x \mathcal{L}(\Phi, \lambda)}$$



# SPECTRUM OF A SIMPLE $T=0$ CORRELATOR

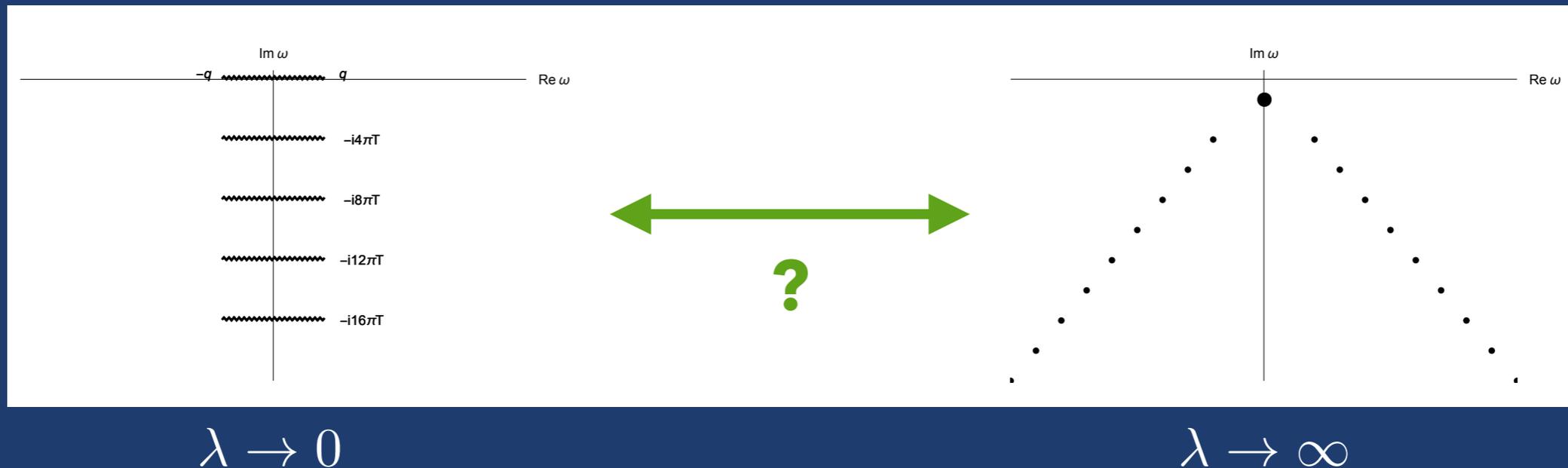
$$\langle \phi(p)\phi(-p) \rangle = \frac{Z(p^2)}{p^2 - m^2 + \Sigma(p^2)}$$



[from Peskin and Schroeder]

# ANALYTIC STRUCTURE OF THERMAL CORRELATORS

$$\langle T_{\mu\nu}(-\omega, -q), T_{\rho\sigma}(\omega, q) \rangle_R \sim \frac{b(\omega, q)}{a(\omega, q)}$$

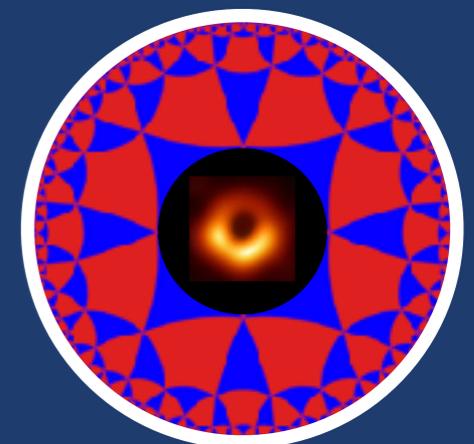


[Hartnoll, Kumar, (2005)]

holography ( $N=4$  SYM-type theories)

meromorphic momentum space correlator

quasinormal modes  
of black holes



methods:

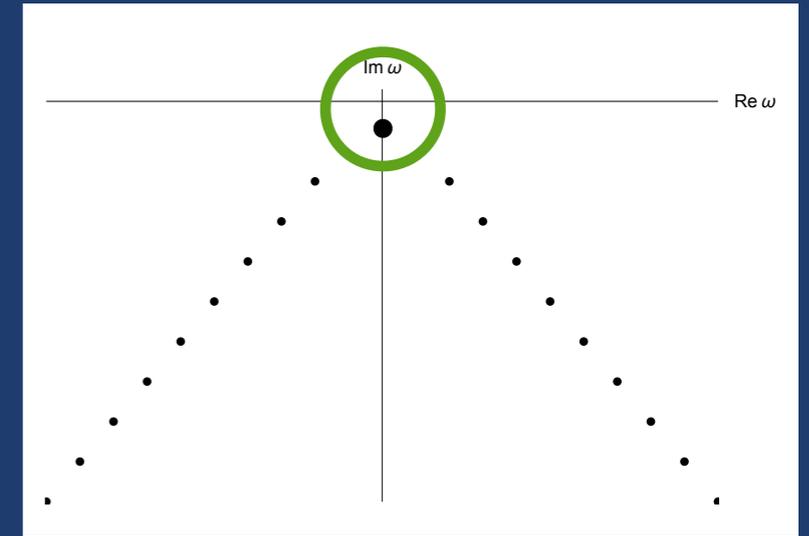
- perturbative QFT: calculation of Feynman diagrams
- holography: solving differential equations in black hole backgrounds

# LOW-ENERGY SPECTRUM AND HYDRODYNAMICS

- low-energy limit of (some) thermal QFTs is described **hydrodynamics**
- conservation laws and global **conserved operators**

$$\nabla_{\mu} T^{\mu\nu} = 0$$

- tensor structures** (symmetries, gradient expansions) and **transport coefficients** (QFT)



$$T^{\mu\nu} = \sum_{n=0}^{\infty} \left[ \sum_i^N \lambda_i^{(n)} \mathcal{T}_{(n)}^{\mu\nu} \right]$$

$$\partial u^{\mu} \sim \partial T \ll 1$$

$$\xrightarrow[\substack{\nabla_{\mu} T^{\mu\nu} = 0 \\ u^{\mu} \sim T \sim e^{-i\omega t + iqz}}]{}$$

convergent series

$$\omega(q) = \sum_{n=1}^{\infty} \alpha_n q^n$$

$$\omega/T \sim q/T \ll 1$$

- dispersion relations are poles:

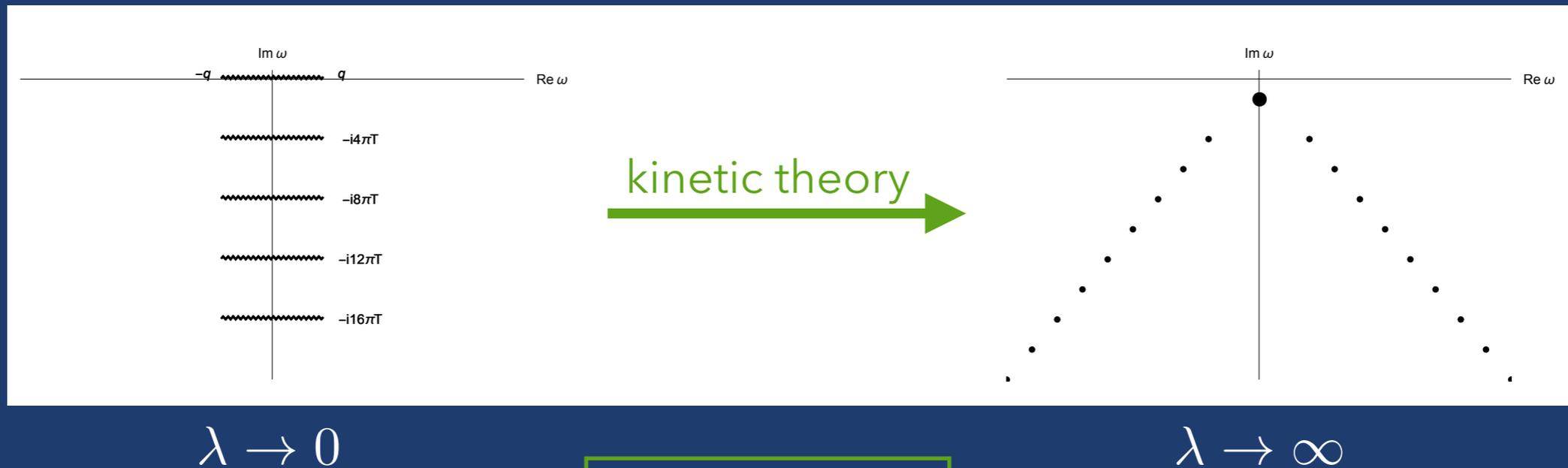
$$\begin{array}{cc} \text{diffusion} & \text{sound} \\ \omega = -iDq^2 & \omega = \pm v_s q - i\Gamma q^2 \end{array}$$

equilibrium temperature

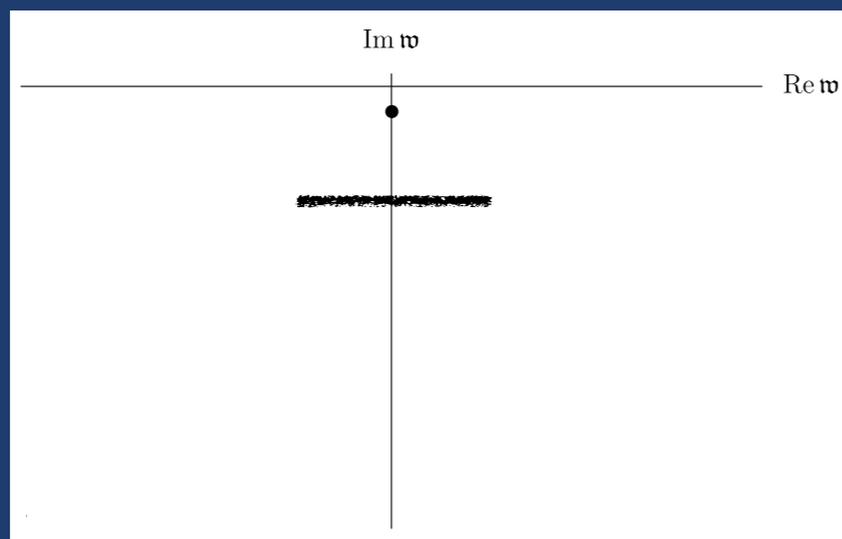
$$q = \sqrt{\mathbf{q}^2}$$

# KINETIC THEORY AND THERMAL SPECTRUM

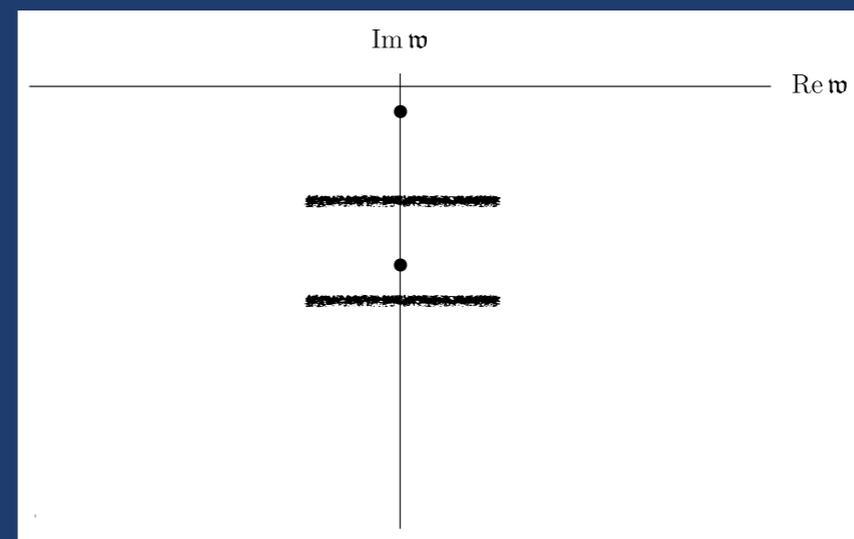
$$\langle T_{\mu\nu}(-\omega, -q), T_{\rho\sigma}(\omega, q) \rangle_R \sim \frac{b(\omega, q)}{a(\omega, q)}$$



poles from a cut?



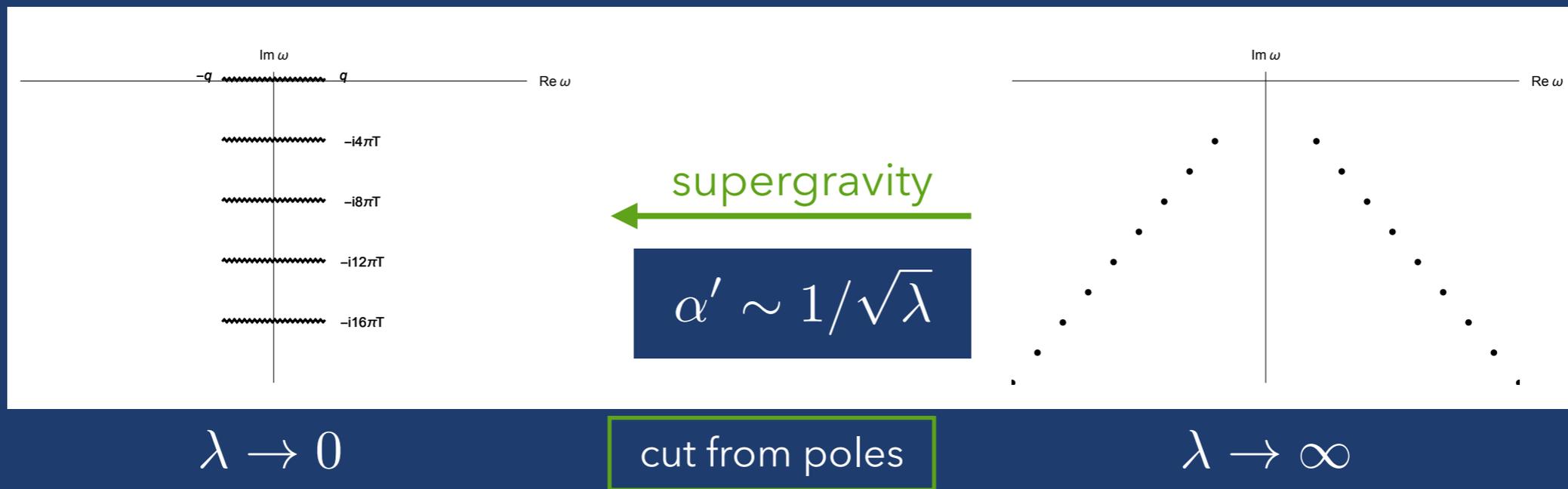
Boltzmann equation – RTA  
[Romatschke, (2016)]



BBGKY hierarchy – RTA-like  
truncations [SG, Soloviev (2025)]

# HOLOGRAPHY AND THERMAL SPECTRUM

$$\langle T_{\mu\nu}(-\omega, -q), T_{\rho\sigma}(\omega, q) \rangle_R \sim \frac{b(\omega, q)}{a(\omega, q)} = \frac{B(\omega, q)}{\prod_{i=0}^{\infty} (\omega - \omega_i(q))}$$

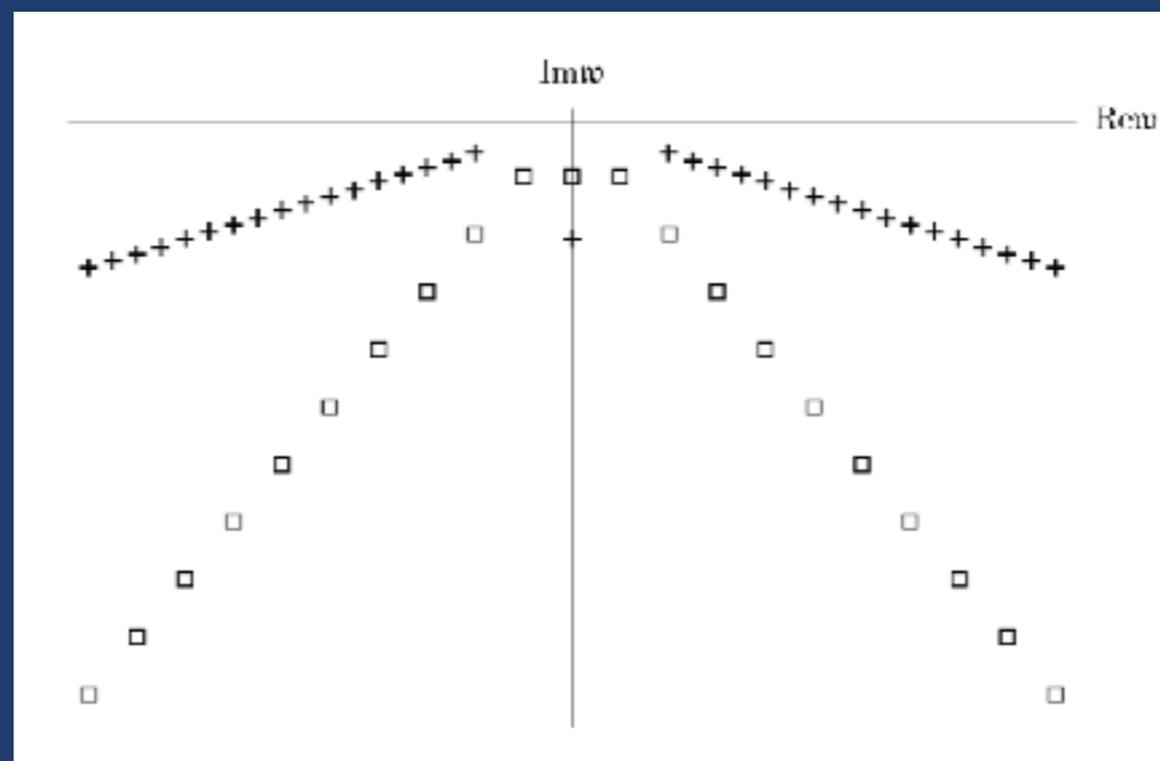


$\lambda \rightarrow 0$

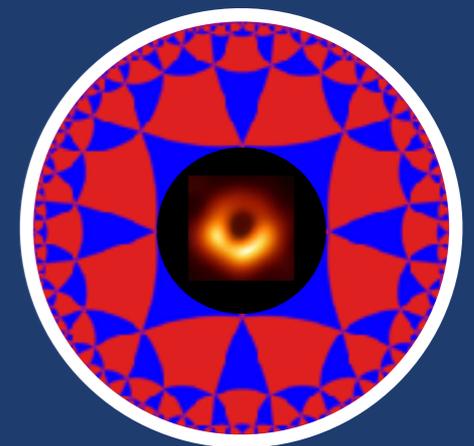
cut from poles

$\lambda \rightarrow \infty$

correlators remain  
meromorphic



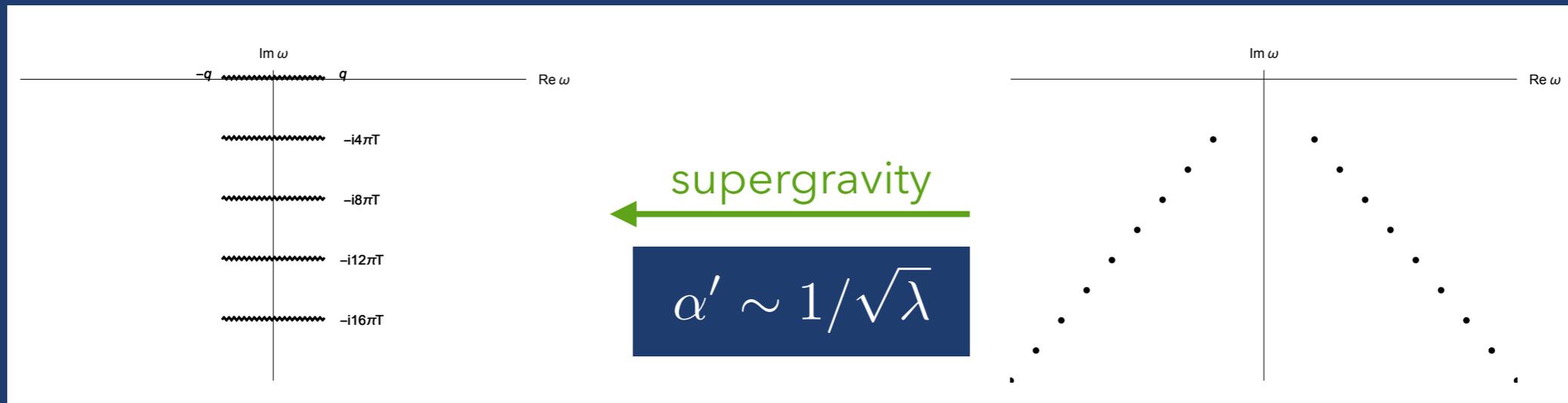
quasinormal modes  
of black holes



[SG, Starinets ..., several papers]

# HOLOGRAPHY AND THERMAL SPECTRUM

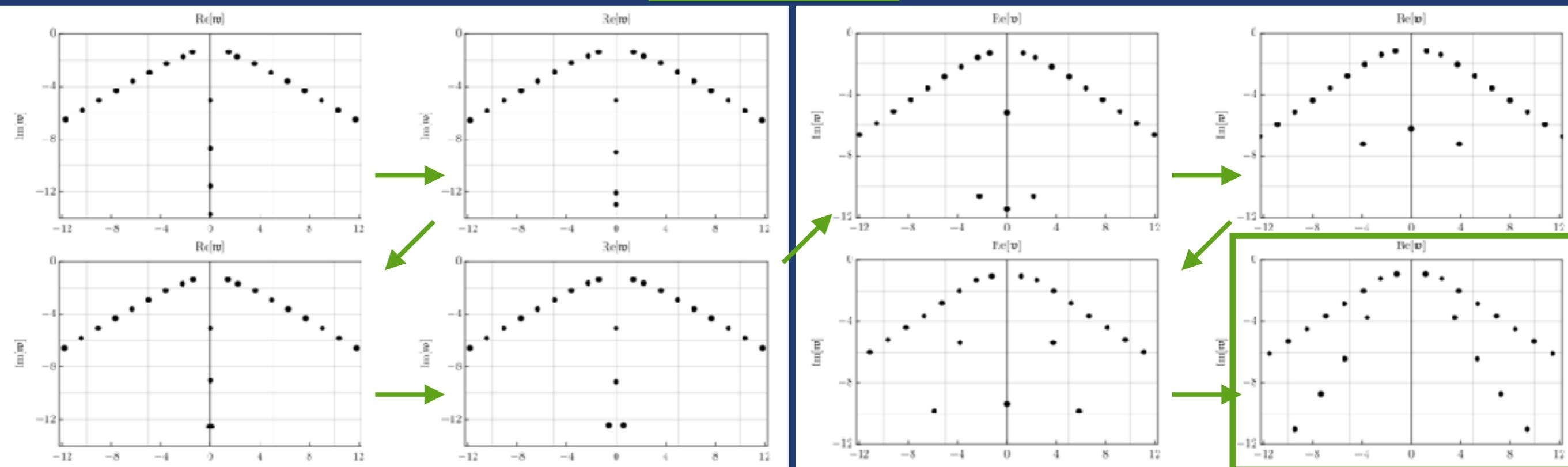
$$\langle T_{\mu\nu}(-\omega, -q), T_{\rho\sigma}(\omega, q) \rangle_R \sim \frac{b(\omega, q)}{a(\omega, q)} = \frac{B(\omega, q)}{\prod_{i=0}^{\infty} (\omega - \omega_i(q))}$$



$\lambda \rightarrow 0$

cut from poles

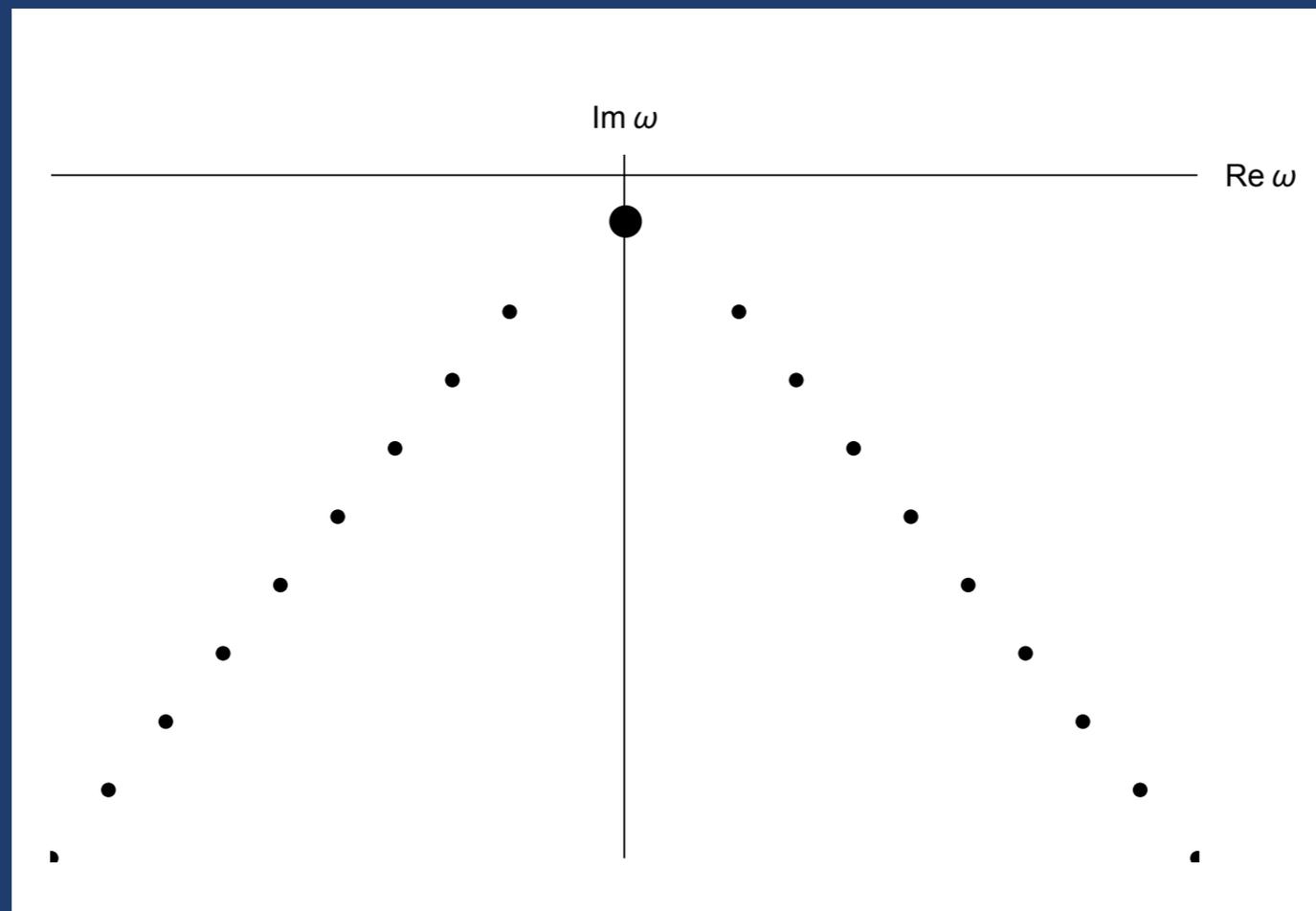
$\lambda \rightarrow \infty$



[SG, Starinets ..., several papers; see also Dodelson on SYK (2025)]

# HOLOGRAPHY AND THERMAL SPECTRUM

$$\langle T_{\mu\nu}(-\omega, -q), T_{\rho\sigma}(\omega, q) \rangle_R \sim \frac{b(\omega, q)}{a(\omega, q)} = \frac{B(\omega, q)}{\prod_{i=0}^{\infty} (\omega - \omega_i(q))}$$



for the rest of the talk:

correlators are meromorphic in  $\omega \in \mathbb{C}$

important recent result:

the thermal product formula [Dodelson, Iosco, Karlsson, Zhiboedov (2023)]

# OUTLINE

pole skipping and reconstruction of hydrodynamics

spectral duality relation

summary and future directions



# POLE SKIPPING

- pole skipping: ubiquitous feature of thermal correlators and black hole perturbations [SG, Schalm, Scopelliti, PRL (2017); Blake, Lee, Liu, JHEP (2018); Blake, Davison, SG, Liu, JHEP (2018); SG, JHEP (2019)]

- originally: all-order hydrodynamic sound mode  $\omega(q) = \sum_{n=1}^{\infty} \alpha_n (T, \mu_i, \langle \mathcal{O}_j \rangle, \lambda) q^n$

passes through a 'chaos point' at where the associated 2-pt function is '0/0':

$$\omega(q = i\lambda_L/v_B) = i\lambda_L = 2\pi T i$$

$$G_R = \frac{0}{0} = N(\delta\omega/\delta q)$$

- relation to maximal quantum chaos as measured by the out-of-time-ordered correlation functions and entanglement wedge [...; Chua, Hartman, Weng (2025)]
- triviality of the Einstein equation at the horizon
- infinite number of such '0/0' points at negative Matsubara frequencies for  $q \in \mathbb{C}$  [SG, Kovtun, Starinets, Tadić, JHEP (2019); Blake, Davison, Vegh, JHEP (2019)]

$$\omega_n(q_n) = -2\pi T i n \quad n \geq 0$$

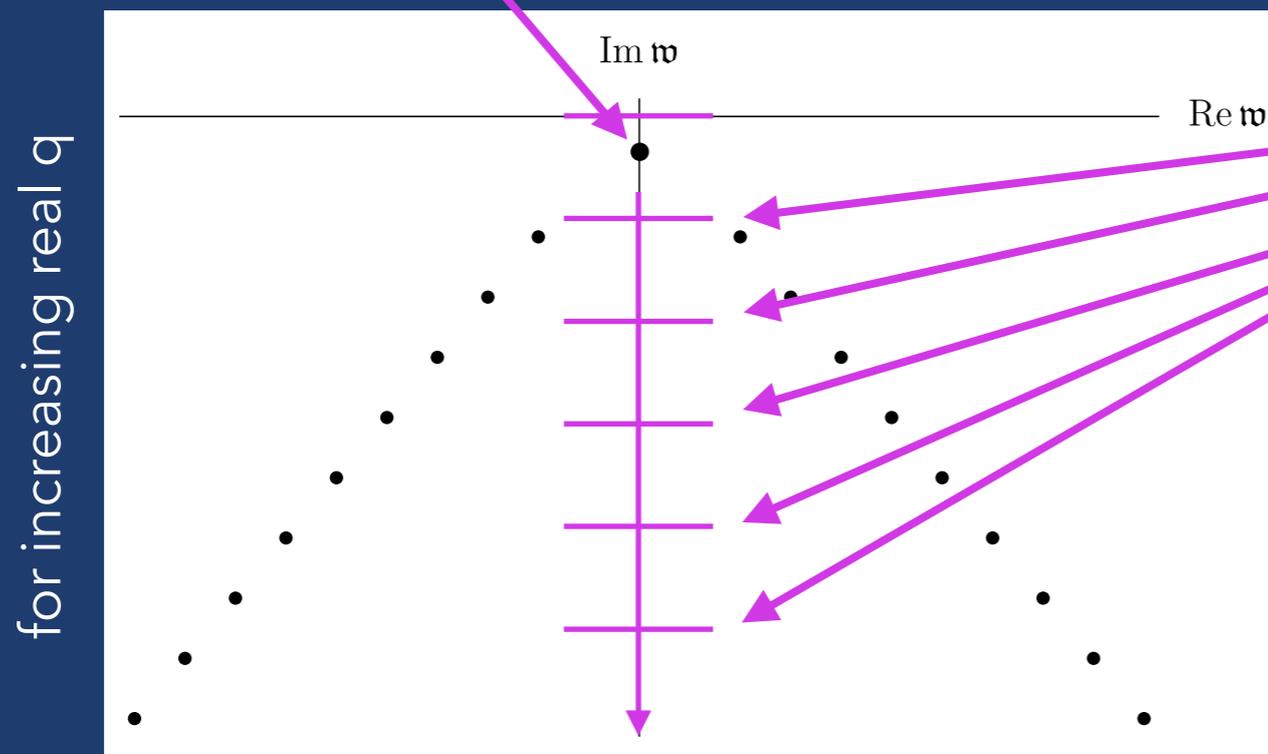
# POLE SKIPPING

- consider momentum diffusion in a neutral CFT dual to AdS-Schwarzschild black brane

$$\omega_{\text{shear}} = -i \sum_{n=1}^{\infty} c_n (q^2)^n = -i\mathcal{D}q^2 + \dots$$

convergent series

[SG, Kovtun, Starinets, Tadić, PRL (2019); ... ; Withers; JHEP (2018); Heller, et.al. (2020, ...)]



$$\omega_n(q_n) = -2\pi T i n$$

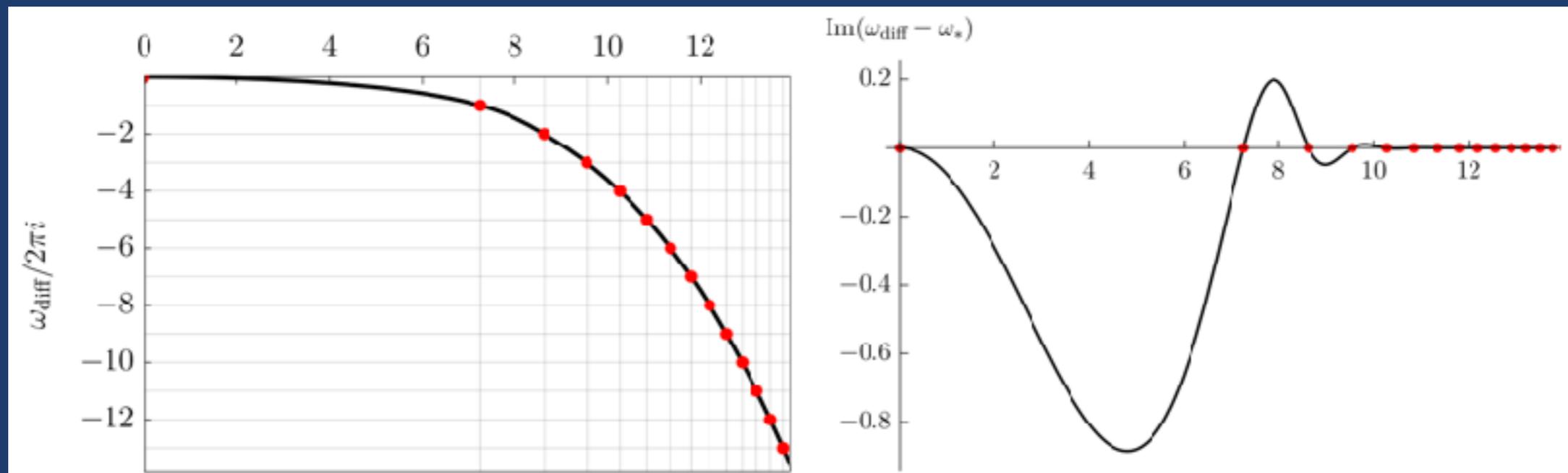
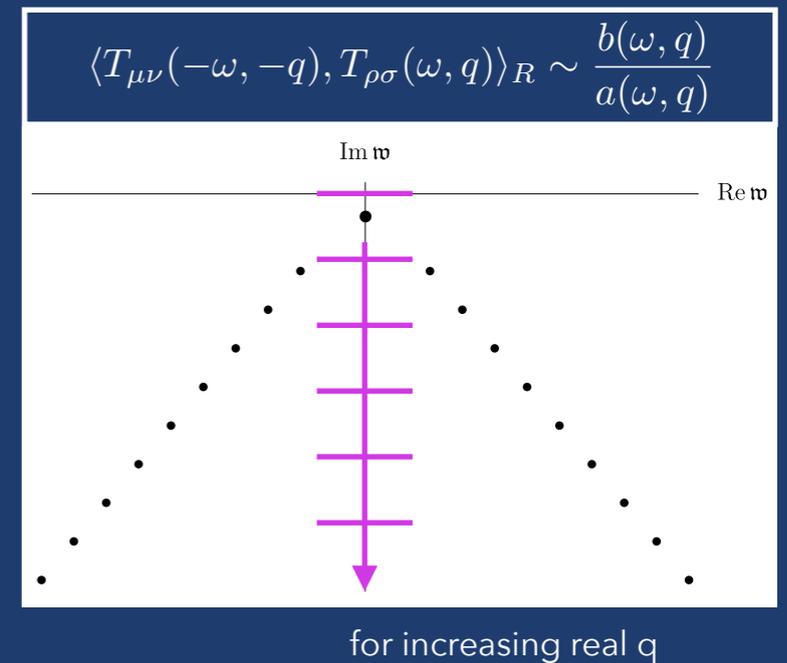
analytic result in AdS4/CFT3  
[SG, PRL (2021)]

$$q_n = \frac{4\pi T}{\sqrt{3}} n^{1/4}, \quad n = 0, 1, 2, \dots$$

# POLE SKIPPING AND HYDRODYNAMICS

- consider momentum diffusion in a neutral CFT dual to AdS-Schwarzschild black brane

$$\mathfrak{w}_{\text{shear}} = -i \sum_{n=1}^{\infty} c_n (q^2)^n = -i\mathcal{D}q^2 + \dots$$



**claim:** in holographic theories of the type discussed here (N=4 SYM, M2, M5, ...), the hydrodynamic dispersion relation follows from only a discrete set of pole-skipping points

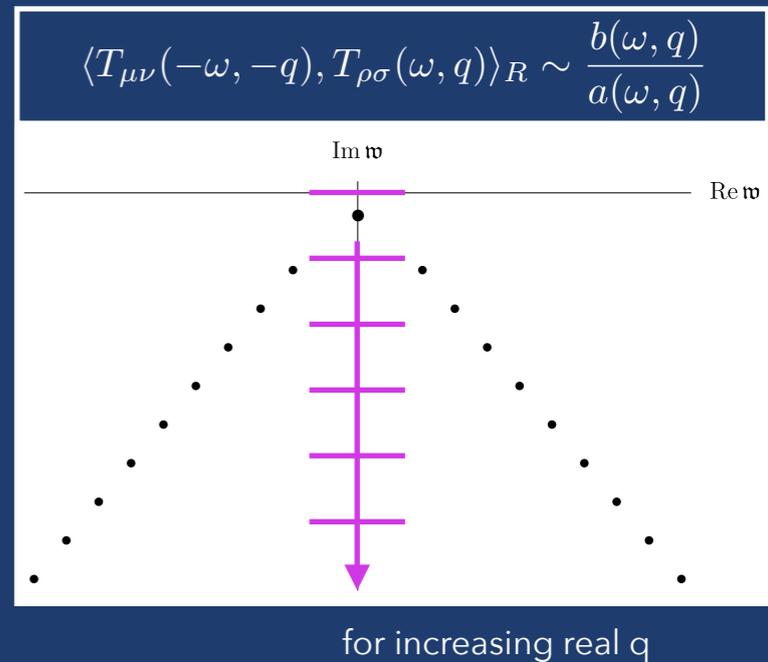
# POLE SKIPPING AND HYDRODYNAMICS

- interpolation problem:

$$\omega_n(q_n) = -2\pi T i n$$



$$\mathfrak{w}_{\text{shear}} = -i \sum_{n=1}^{\infty} c_n (q^2)^n = -i \mathfrak{D} q^2 + \dots$$

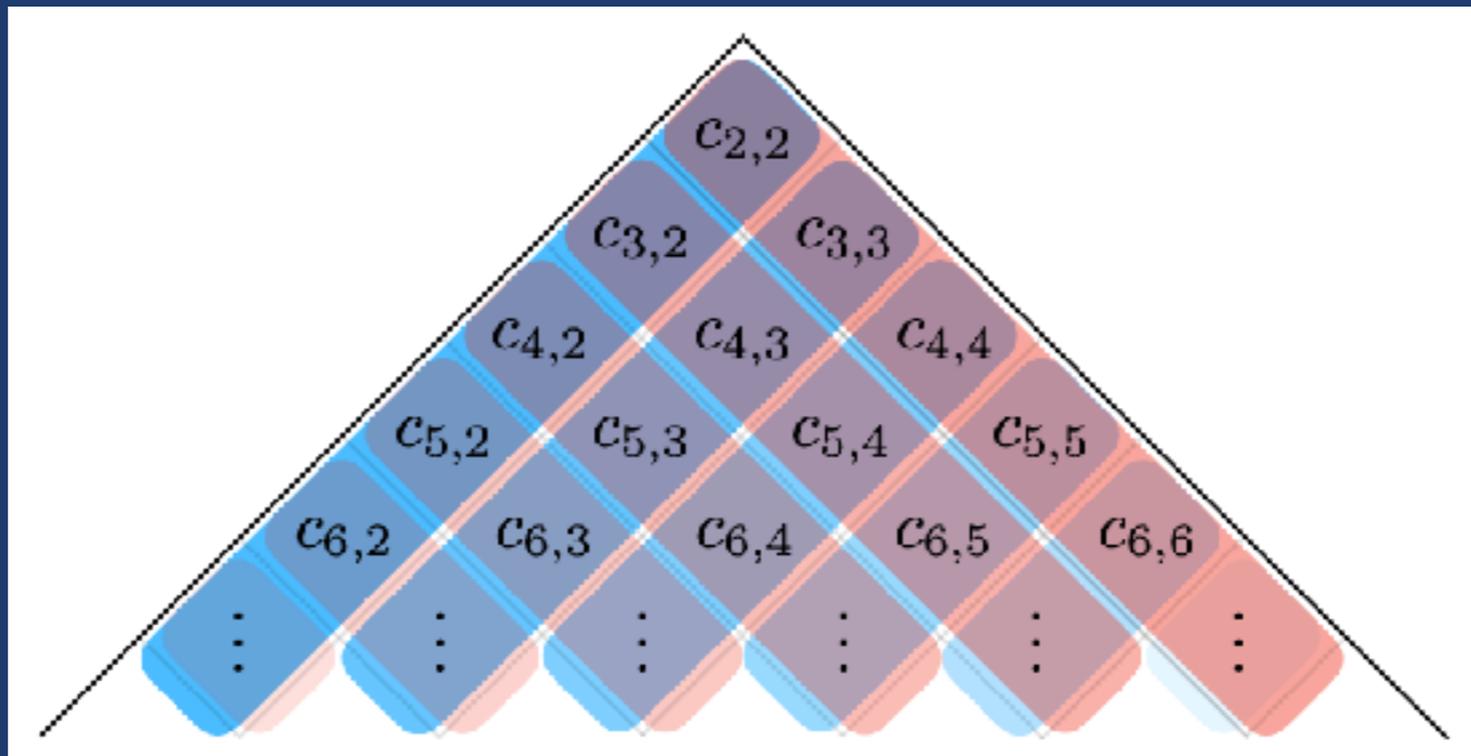


- unique solutions to interpolation problems are rare (Weierstrass-Hadamard, Nevanlinna-Pick)
- trick: 'analytic continuation' to  $d$  spacetime dimensions and expansion around infinite  $d$  [SG, Lemut, Pedraza, PRD (2023)]
- general relativity in large  $d$  drastically simplifies:  $V \sim 1/r^d$
- recall: large- $d$  limit of quantum mechanics is useful in atomic physics (e.g., for Helium)

# POLE SKIPPING AND HYDRODYNAMICS

- interpolation:  $\omega_n(q_n) = -2\pi T i n \longrightarrow \mathfrak{w}_{\text{shear}} = -i \sum_{n=1}^{\infty} c_n (q^2)^n = -i \mathcal{D} q^2 + \dots$
- 'analytic continuation' to  $d$  spacetime dimensions and expansion around infinite  $d$

$$\begin{aligned} \omega_0(q = Q\sqrt{d}) &= -iQ^2 - \frac{i}{d^2} (c_{2,2}Q^4) - \frac{i}{d^3} (c_{3,2}Q^4 + c_{3,3}Q^6) \\ &\quad - \frac{i}{d^4} (c_{4,2}Q^4 + c_{4,3}Q^6 + c_{4,4}Q^8) + \dots \end{aligned}$$



# POLE SKIPPING AND HYDRODYNAMICS

- interpolation:  $\omega_n(q_n) = -2\pi T i n \longrightarrow \mathfrak{w}_{\text{shear}} = -i \sum_{n=1}^{\infty} c_n (q^2)^n = -i \mathfrak{D} q^2 + \dots$
- 'analytic continuation' to  $d$  spacetime dimensions and expansion around infinite  $d$

$$\omega_0(q) = -i \left( \frac{q}{\sqrt{d}} \right)^2 - i \sum_{m=2}^{\infty} \frac{1}{d^m} \sum_{j=2}^m c_{m,j} \left( \frac{q}{\sqrt{d}} \right)^{2j}$$

$$\frac{q_n}{\sqrt{d}} = \sqrt{\frac{nd}{2}} \left( 1 + \sum_{m=1}^{\infty} \frac{b_{n,m}}{d^m} \right)$$

$$b_{n,1} = - \sum_{m=2}^{\infty} \frac{n^{m-1} c_{m,m}}{2^m}$$

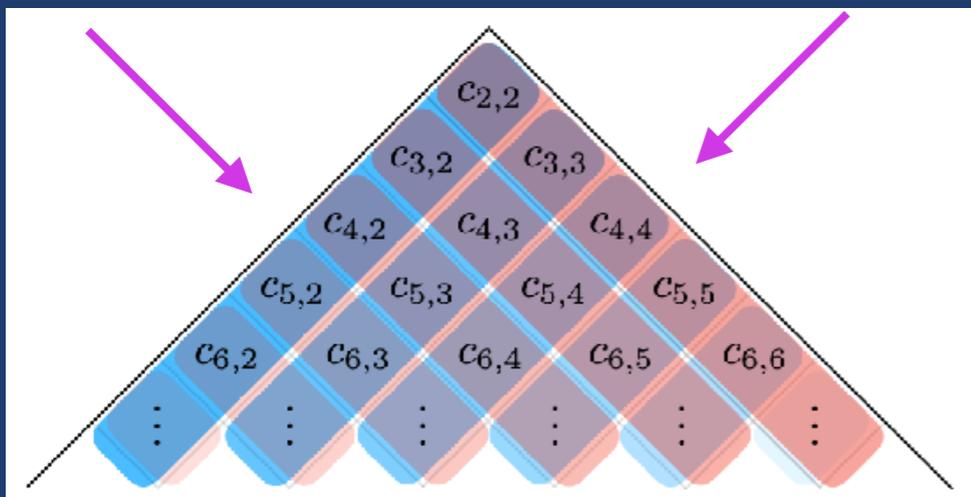
$$b_{n,2} = - \frac{b_{n,1}^2}{2} - \sum_{m=2}^{\infty} \frac{n^{m-1} (c_{m+1,m} + 2m b_{n,1} c_{m,m})}{2^m}$$

second analytic continuation

$$n \in \mathbb{Z} \rightarrow x \in \mathbb{R}$$

hydrodynamics

pole skipping



$$c_{m,m} = - \frac{2^m}{(m-1)!} \partial_x^{m-1} b_1(0)$$

$$c_{m+1,m} = - \frac{2^m \partial_x^{m-1} b_2(0)}{(m-1)!} + \sum_{j=2}^{m-1} \left( j - \frac{1}{4} \right) c_{j,j} c_{m-j+1, m-j+1}$$

generating functions

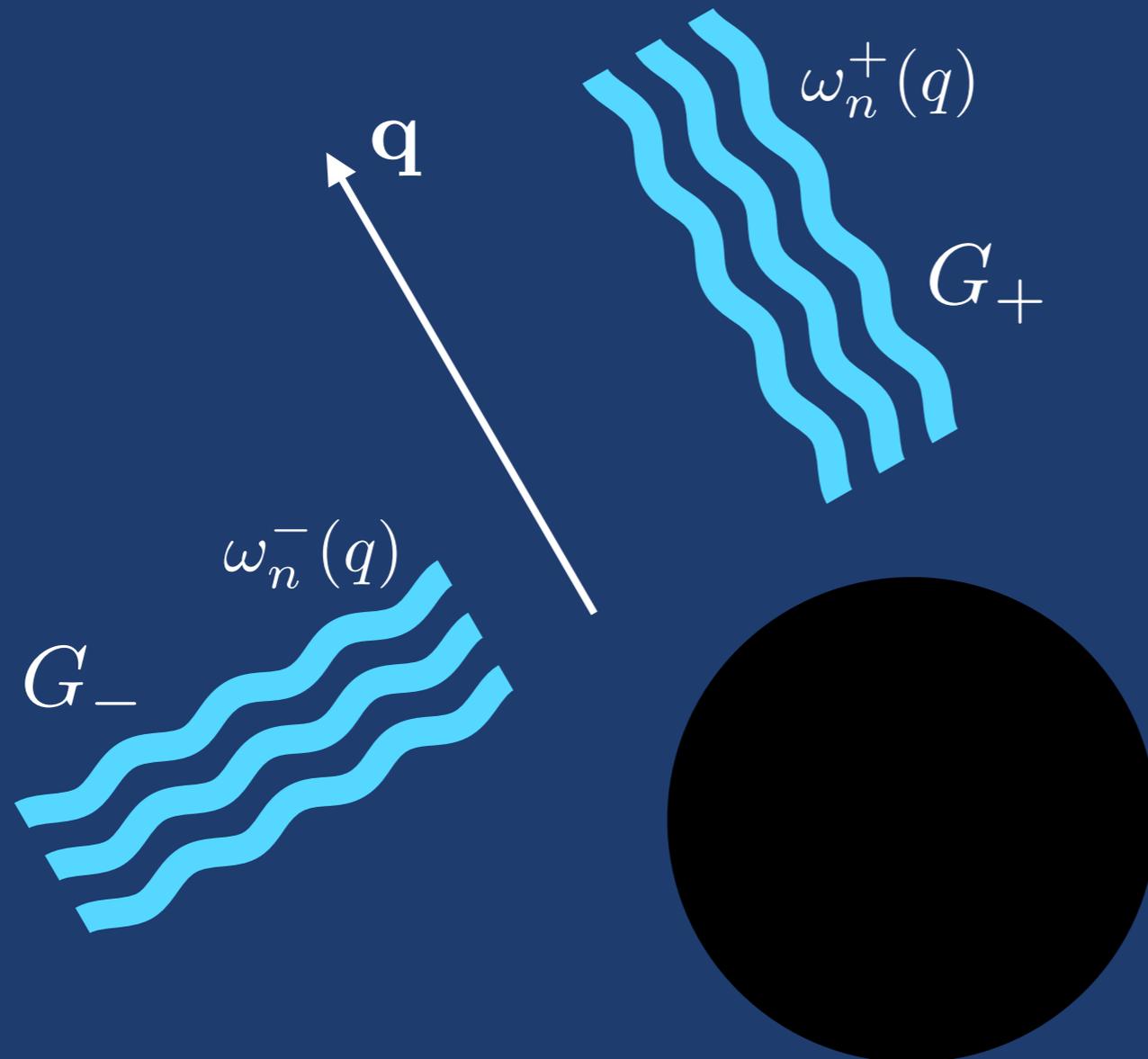
symmetry?

# SPECTRAL DUALITY RELATION

[SG, Vrbica, PRL (2024), 2505.14229  
and another paper *very soon*]

# SPECTRAL DUALITY RELATION

- two channels of black hole perturbations in AdS4/CFT3 cases:  
even (or sound) and odd (or shear)
- in each channel, we have QNMs  $\omega_n^\pm(q)$  with associated dual correlators  $G_\pm(\omega, q)$



isospectrality

in asymptotically flat black holes  
[Chandrasekhar, Detweiler (1975)]

$$\omega_n^+(q) = \omega_n^-(q)$$

spectral duality relation

in asymptotically AdS black holes

$$\omega_n^+(q) \leftrightarrow \omega_m^-(q)$$

# SPECTRAL DUALITY RELATION

- two channels of black hole perturbations in AdS4/CFT3 cases:  
even (or sound) and odd (or shear)
- two meromorphic CFT retarded correlators  $G_{\pm}(\omega, q)$  with QNMs  $\omega_n^{\pm}(q)$  (e.g. of  $T^{\mu\nu}$  or  $J^{\mu}$ )
- CFT3s have S-duality or particle-vortex duality  
gravity in 4d has Chandrasekhar duality, Darboux duality, EM duality

duality:

$$G_+(\omega, q)G_-(\omega, q) = \frac{\omega^2}{\omega_*^2(q)} - 1$$

self-duality:  
 $\omega_*(q) \rightarrow \infty$

$$G_+(\omega, q)G_-(\omega, q) = -1$$

algebraically special frequencies  
relation to pole skipping  
[SG, Vrbica, EPJC (2023)]  
easy to compute using the  
Robinson-Trautman ansatz

# SPECTRAL DUALITY RELATION

- define infinite convergent product

$$S(\omega, q) \equiv \left(1 + \frac{\omega}{\omega_*(q)}\right) \prod_n \left[1 - \frac{\omega}{\omega_n^+(q)}\right] \left[1 + \frac{\omega}{\omega_n^-(q)}\right]$$

- duality relations, the thermal product formula and details about QNM and Greens function asymptotics give a 'universal' relation:

$$S(\omega, k) - S(-\omega, k) = 2i\lambda(k) \sinh \frac{\beta\omega}{2}$$

$$\lambda(k) = \frac{2}{i\beta} \left[ \frac{1}{\omega_*(k)} + \sum_n \left( \frac{1}{\omega_n^-(k)} - \frac{1}{\omega_n^+(k)} \right) \right]$$

- infinite towers of constraints

$$e_{2j+1}(\mathcal{W}) = \frac{i\lambda}{(2j+1)!} \left(\frac{\beta}{2}\right)^{2j+1}$$

elementary symmetric  
polynomials with  $j \in \{0, 1, \dots, n\}$

$$\mathcal{W} = \{1/\omega_*, 1/\omega_1^-, \dots, 1/\omega_n^-, -1/\omega_1^+, \dots, -1/\omega_n^+\}$$

# SPECTRAL DUALITY RELATION

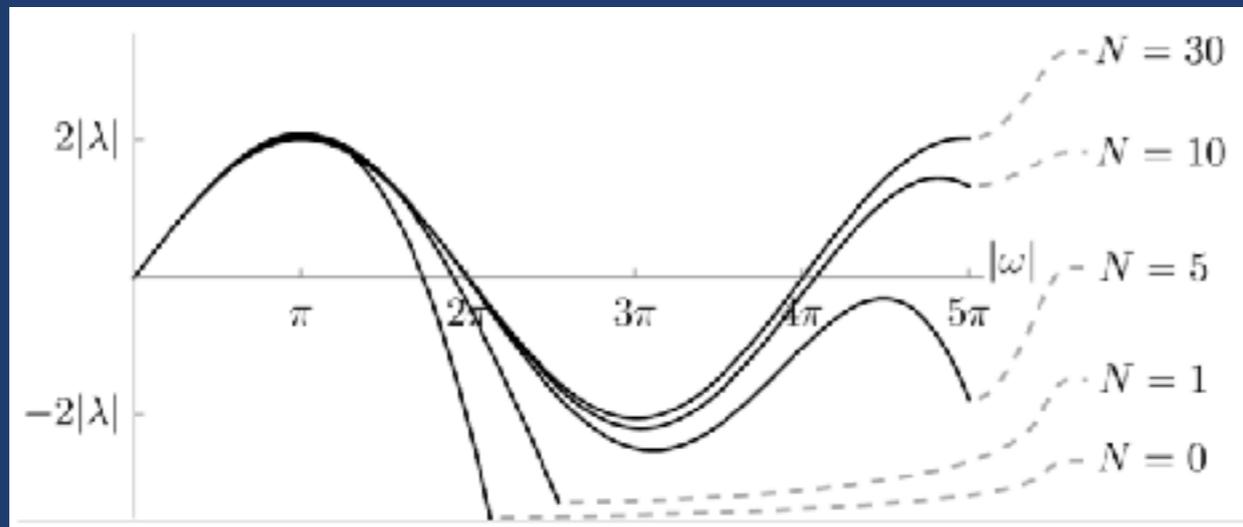
- AdS4-Schwarzschild black brane

- $\langle J, J \rangle_R$  is self-dual

poles must converge to Matsubara frequencies  $\frac{iD_c\beta}{2} \lim_{q \rightarrow 0} k^2 S(\omega) = \sinh \frac{\beta\omega}{2}$

- $\langle T, T \rangle_R$  has  $\omega_* = i \frac{\gamma q^4}{6\bar{\epsilon}}$

various hydro constraints follow, e.g.:  $D/\Gamma = 2$



- AdS4-Reissner-Nordström black brane

- channels are coupled with  $\omega_* = i \frac{\gamma q^4}{6\bar{\epsilon}} \left( \frac{1}{2} \pm \sqrt{\frac{1}{4} + \left( \frac{2Q\gamma}{3\bar{\epsilon}} \right)^2 q^2} \right)^{-1}$

# SPECTRAL DUALITY RELATION

- duality relation for any pair of meromorphic correlators:

$$G_+(\omega)G_-(\omega) = -1$$

$$S(\omega) = \prod_n \left(1 - \frac{\omega}{\omega_n^+}\right) \left(1 + \frac{\omega}{\omega_n^-}\right)$$

$$S(\omega) - S(-\omega) = 2i\lambda \sinh \frac{\beta\omega}{2}$$

- knowing one spectrum is sufficient for determining the other spectrum!
- spectral duality relation connects poles and zeros of a single correlator:

thermal product  
formula

$$-G(\omega)G^{-1}(\omega) = -1$$

$$S(\omega) = \prod_n \left(1 - \frac{\omega}{\omega_n}\right) \left(1 + \frac{\omega}{z_n}\right)$$

$$r_n = \lambda G(0) \frac{\omega_n \sinh \frac{\beta\omega_n}{2}}{2 \prod_{\substack{m \\ m \neq n}} \left(1 - \frac{\omega_n^2}{\omega_m^2}\right)}$$

# SPECTRAL DUALITY RELATION

- large- $N$  field theory and double-trace deformed RG flow

$$Z_f[J] = e^{\Gamma[J,f]} = \left\langle e^{\int \mathcal{O} J - \frac{f}{2} \int \mathcal{O}^2} \right\rangle \quad G(f) = \frac{G_0}{1 + fG_0}$$

$$[G(f_1)(f_2 - f_1) + 1][G(f_2)(f_2 - f_1) - 1] = -1$$

- using duality relations between UV and IR CFTs

$$Z_{f_-}[J_-] = e^{\Gamma_-[J_-,f_-]} = \left\langle e^{\int \mathcal{O}_- J_- - \frac{f_-}{2} \int \mathcal{O}_-^2} \right\rangle_-$$

$$Z_{f_+}[J_+] = e^{\Gamma_+[J_+,f_+]} = \left\langle e^{\int \mathcal{O}_+ J_+ - \frac{f_+}{2} \int \mathcal{O}_+^2} \right\rangle_+$$

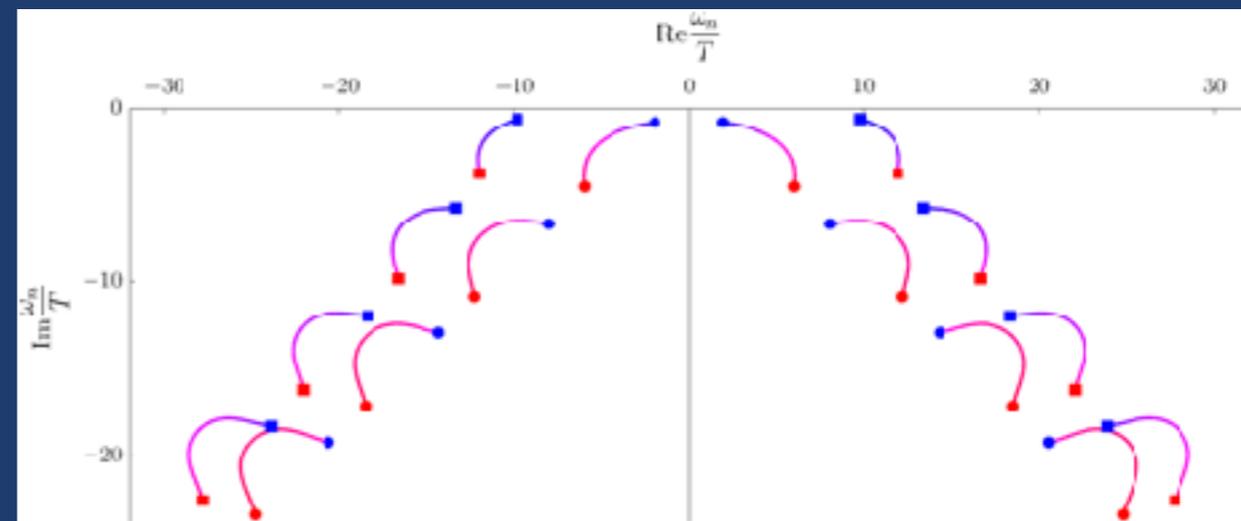
$$G_0^- G_0^+ = -1$$

$$\Delta_+ + \Delta_- = d$$

$$f_+ f_- = -1$$

$$G_-(f_-) = f_+ [f_+ G_+(f_+) - 1]$$

- same statement applies to correlators in any pair of theories related by Legendre transform



# SUMMARY AND FUTURE DIRECTIONS

# SUMMARY AND FUTURE DIRECTIONS

- QFTs and gravity exhibit extremely interesting and powerful (mathematical) structures
- large classes of large- $N$  QFTs and black hole spectra exhibit stringent constraints, in particular, the 'universal' **spectral duality relation**
- results apply directly to gravity in AdS and the spectra of quasinormal modes
- keep exploring...  
... ideally in realistic QFTs beyond the large- $N$

THANK YOU!