

Latest Results from the LUX-ZEPLIN () Experiment



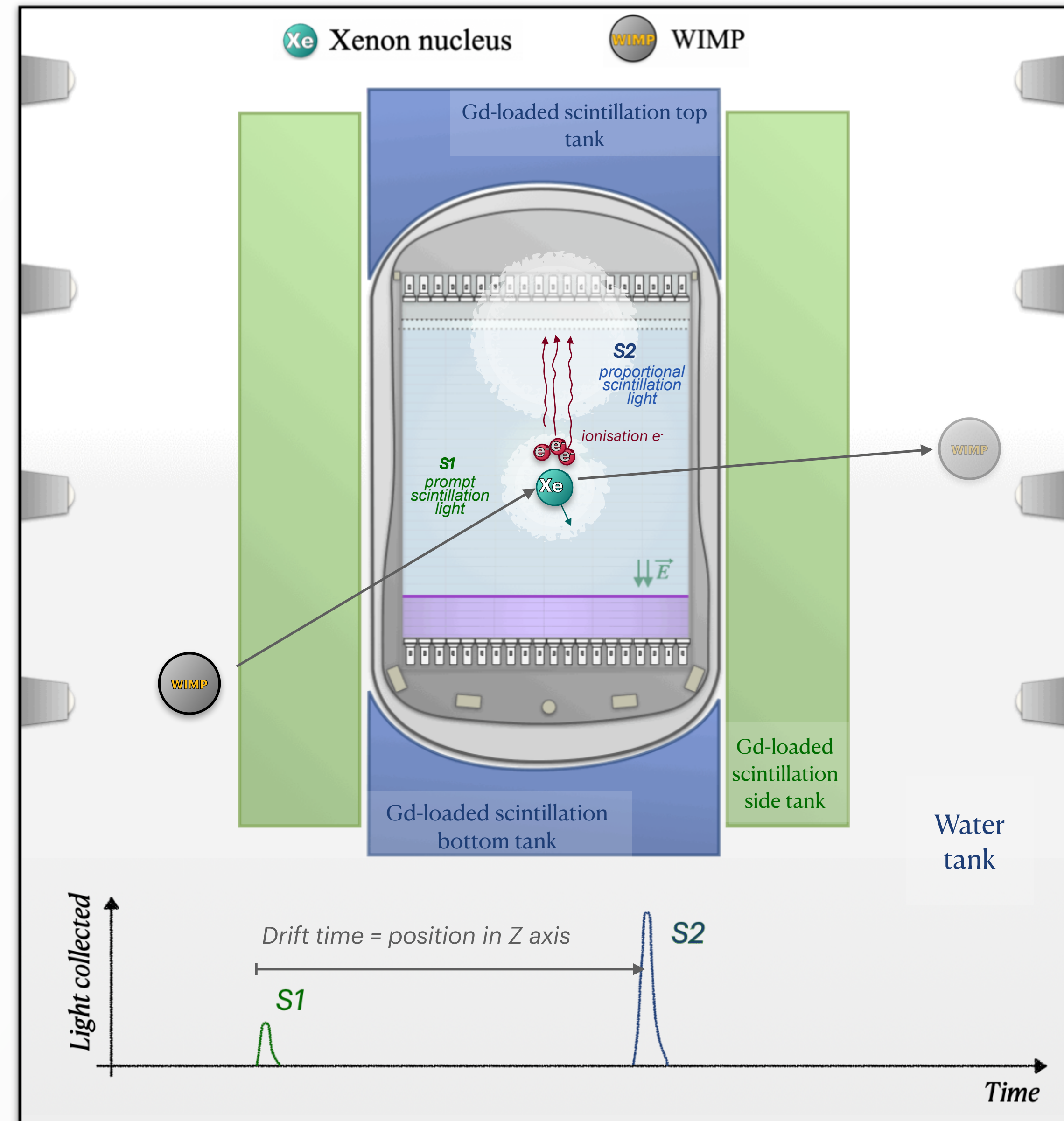
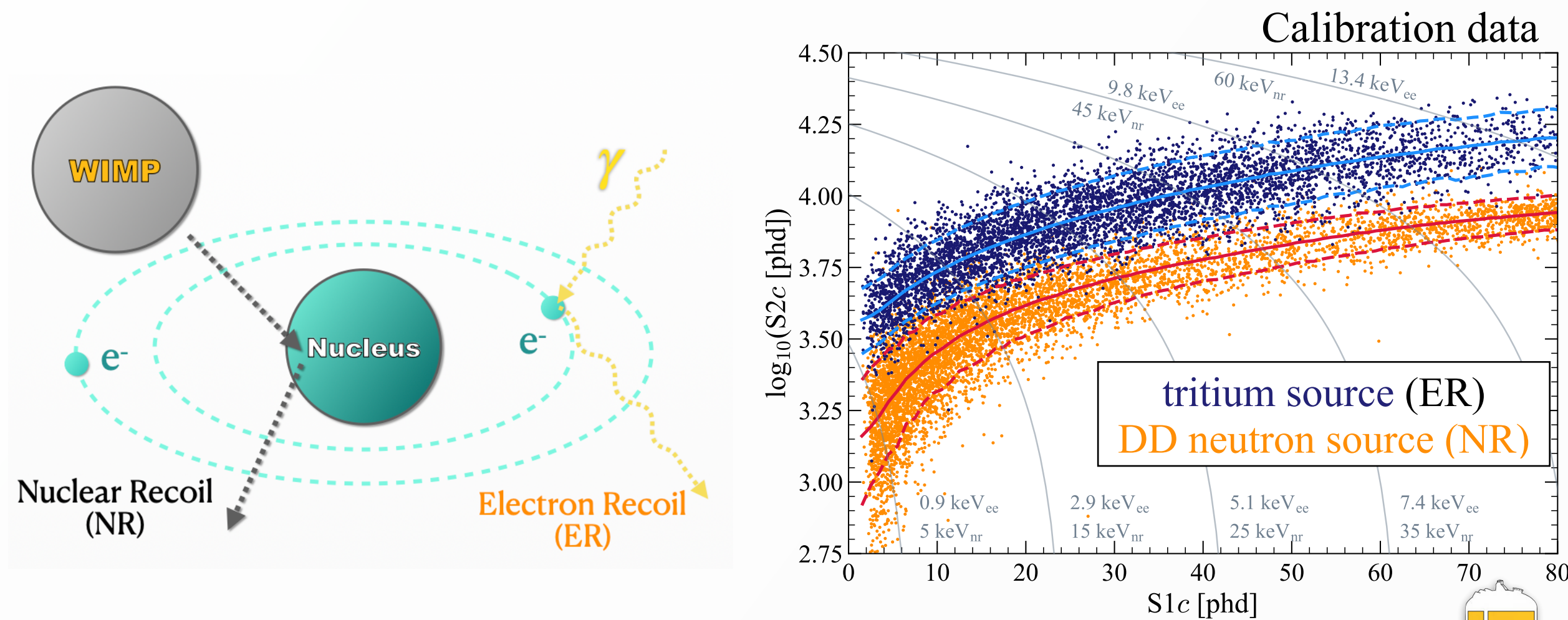
LIDINE 2025

Alberto Uson
on behalf of the LZ Collaboration
21/10/2025



The LUX-ZEPLIN (LZ) detector

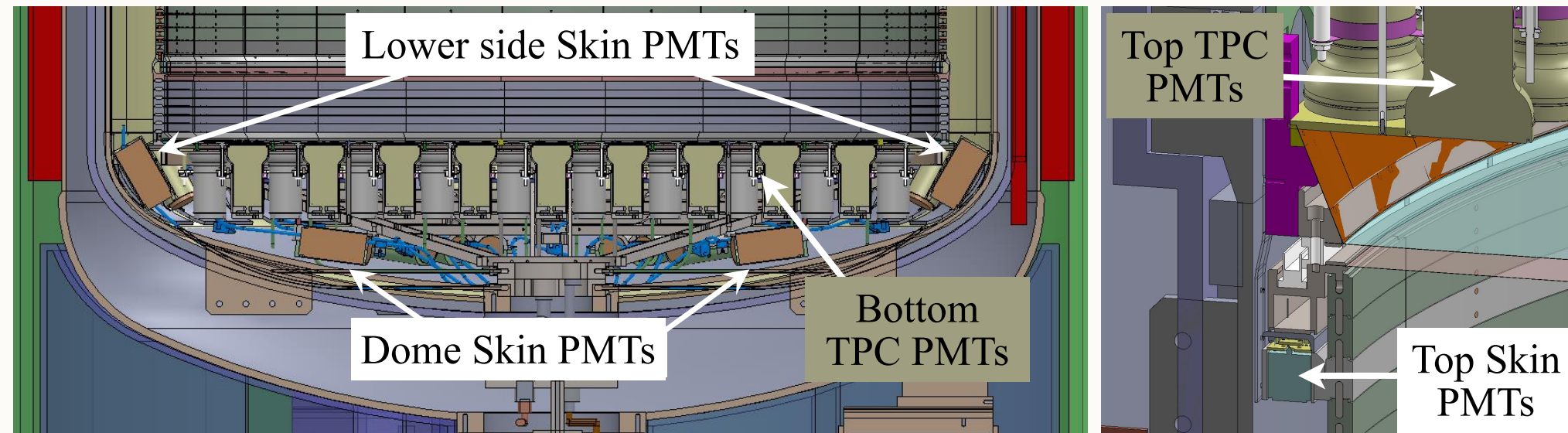
- Optimised for searching **WIMPs** with masses above $\sim 3 \text{ GeV}/c^2$
- Sensitive to additional processes, including BSM physics: $0\nu\beta\beta$, axions, etc
- 7t active mass **dual-phase xenon electroluminescent-TPC** at the Sanford Underground Research Facility (SURF), Lead, South Dakota
- Events reconstructed combining scintillation (S1) and ionisation (S2)
- Excellent 3D position reconstruction ($\sim \text{mm}$)
- Single Scatter (SS) vs Multiple Scatter (MS) interactions discrimination
- S2/S1 ratio indicates the type of interaction:
 - Nuclear recoils (**NR**) in xenon mostly from WIMPs
 - Most backgrounds are Electron Recoils (**ER**)



The LZ veto system: Skin+OD

Skin detector

Main purpose: veto Compton recoils from $\sim$ MeV radiogenic γ -rays



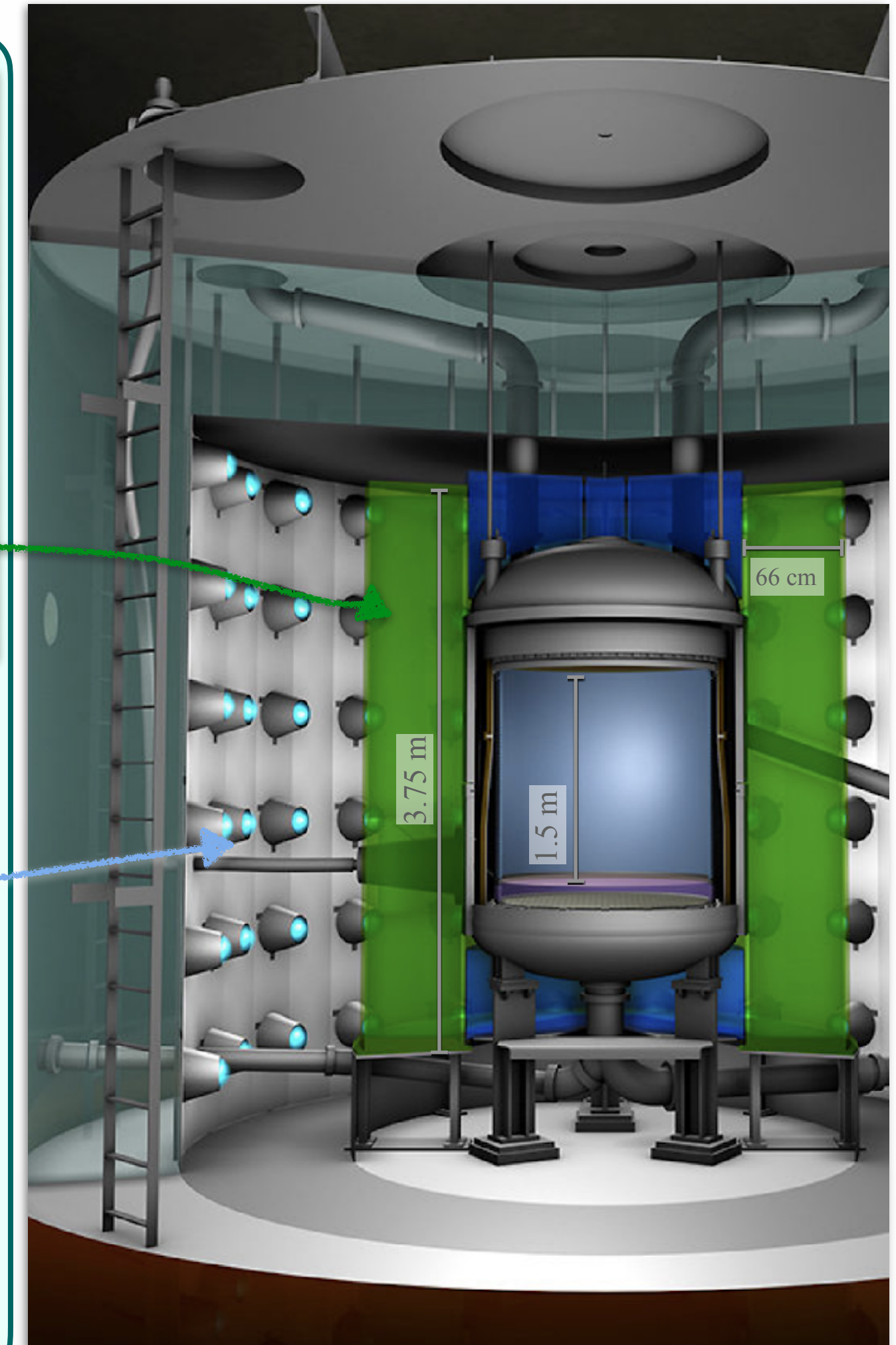
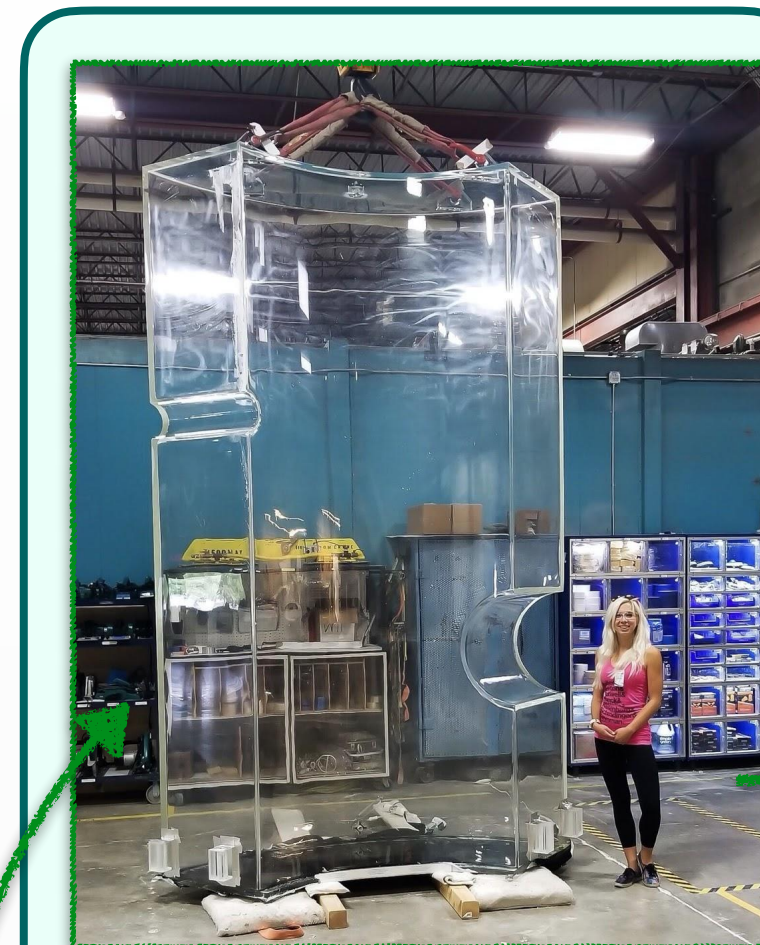
Comprises: 2t-LXe region between TPC outer walls and inner cryostat vessel: $\rightarrow$ 93 PMTs at the top and 38 PMTs at the bottom

Outer detector (OD)

Main purposes: $\left\{ \begin{array}{l} \text{active muon and internal neutrons veto} \\ \text{passive external neutron veto} \end{array} \right.$

Comprises:

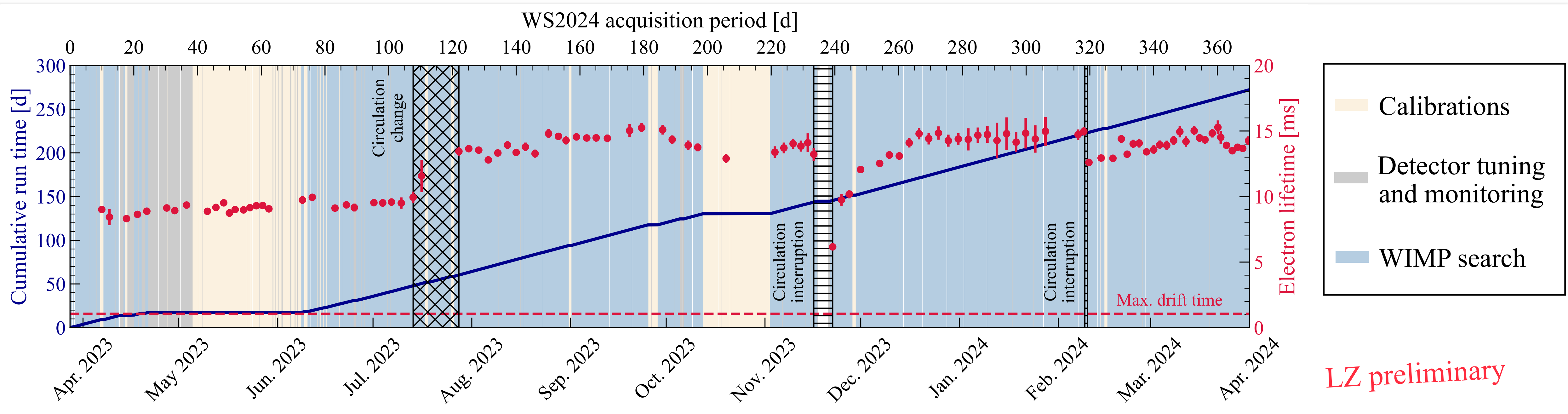
1. 10 UV-transparent acrylic tanks with 17t of Gd-loaded liquid scintillator to increase n-capture efficiency
2. A tank filled with $\sim$ 238t of ultra-pure water instrumented with 120 8-inch PMTs



WIMP Search 2024 (WS2024)

[PRL 135, 011802 \(2025\)](#)
Announced on 26/08/24

- Pre-WS2024: successful (60d) first run: **WS2022**, provided world-leading Dark Matter result [PRL 131, 041002 \(2023\)](#)
- After detector optimisation and calibration: **WS2024**
- **220 live-day exposure**: data from March 2023 to end March 2024
 - Several calibrations before and during run to correct for spatial/temporal effects
 - Very stable run conditions: e.g. >8 ms e⁻ lifetime
 - $>95\%$ data-taking uptime during science operation (before data-quality and lifetime exclusions)

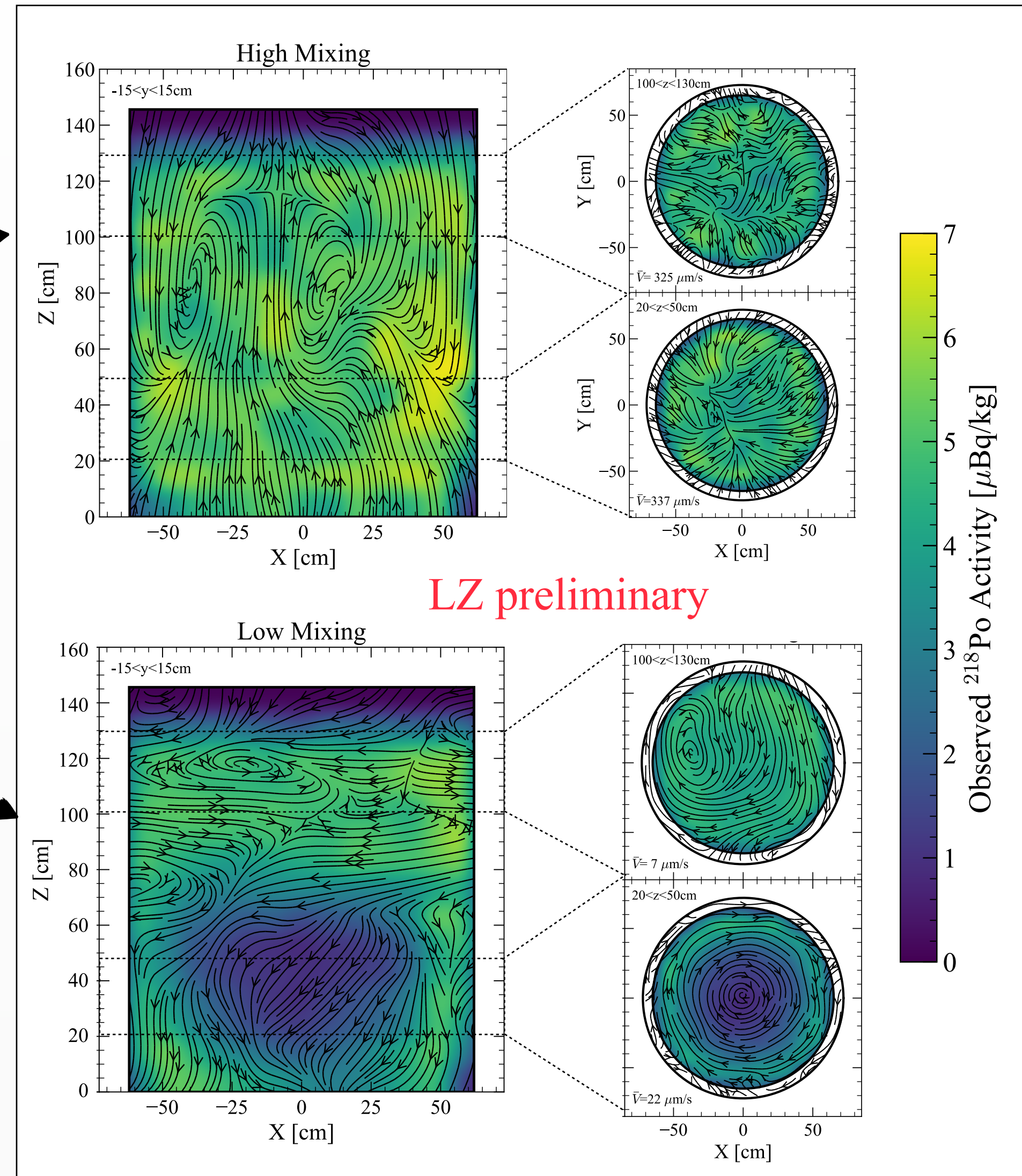
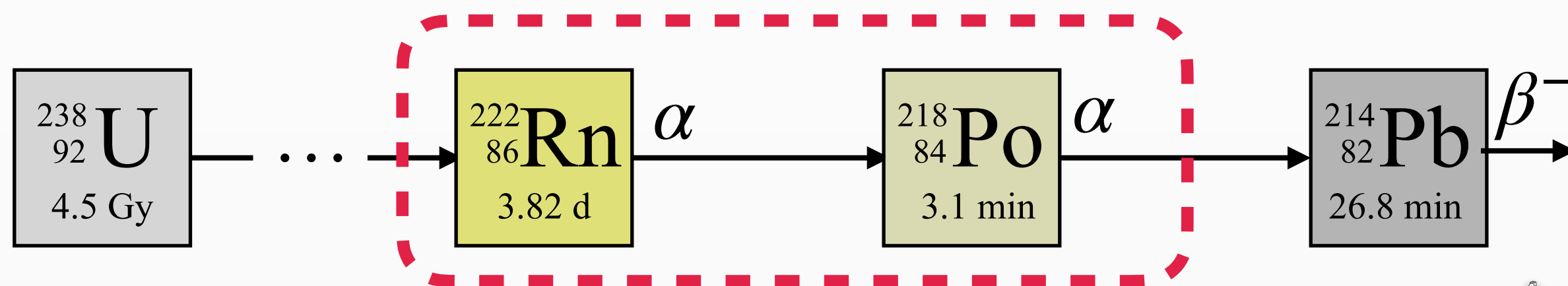


Xenon flow in the TPC: Radon tag

- “Naked” ^{214}Pb β decay is the biggest ER background
- Excellent control of temperature and xenon flow
- WS2024 data acquired in two different flow states:

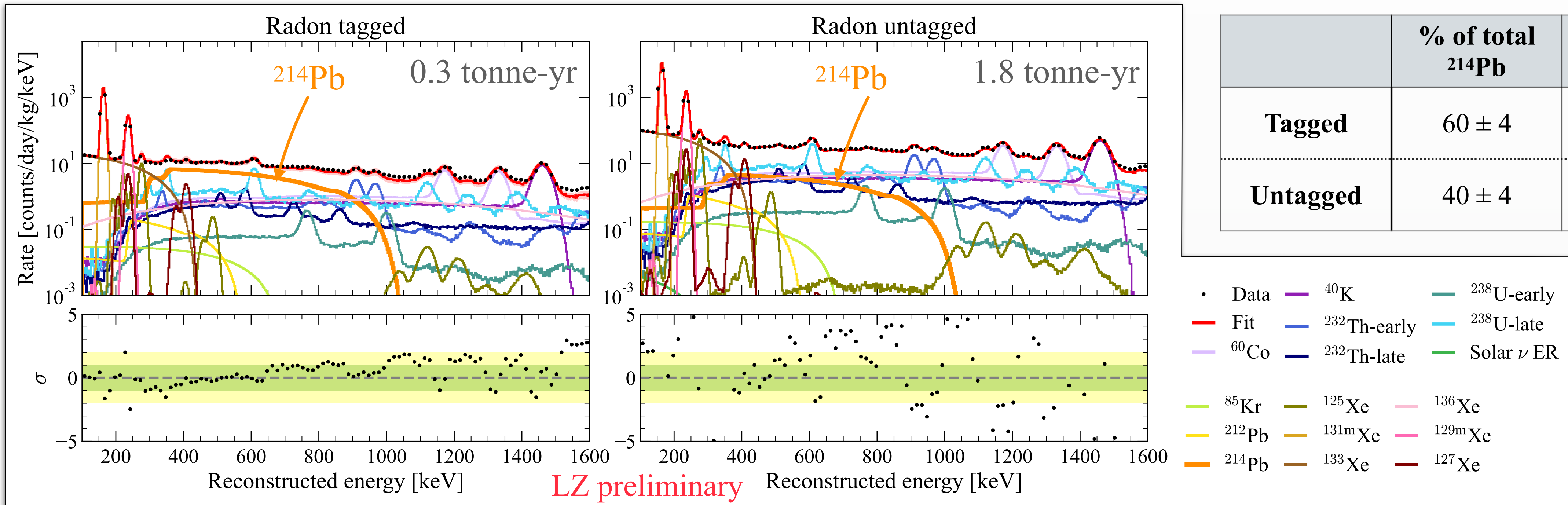
1. **High-mixing:** turbulent flow $\rightarrow$ uniform distribution of injected calibration sources
2. **Low-mixing:** distinct regions of laminar flow $\rightarrow$ possibility to map the fluid velocity field

- Coincident ^{222}Rn (^{238}U chain) - ^{218}Po α decays (^{218}Po $T_{1/2} = 3.1$ min) to map the fluid velocity field



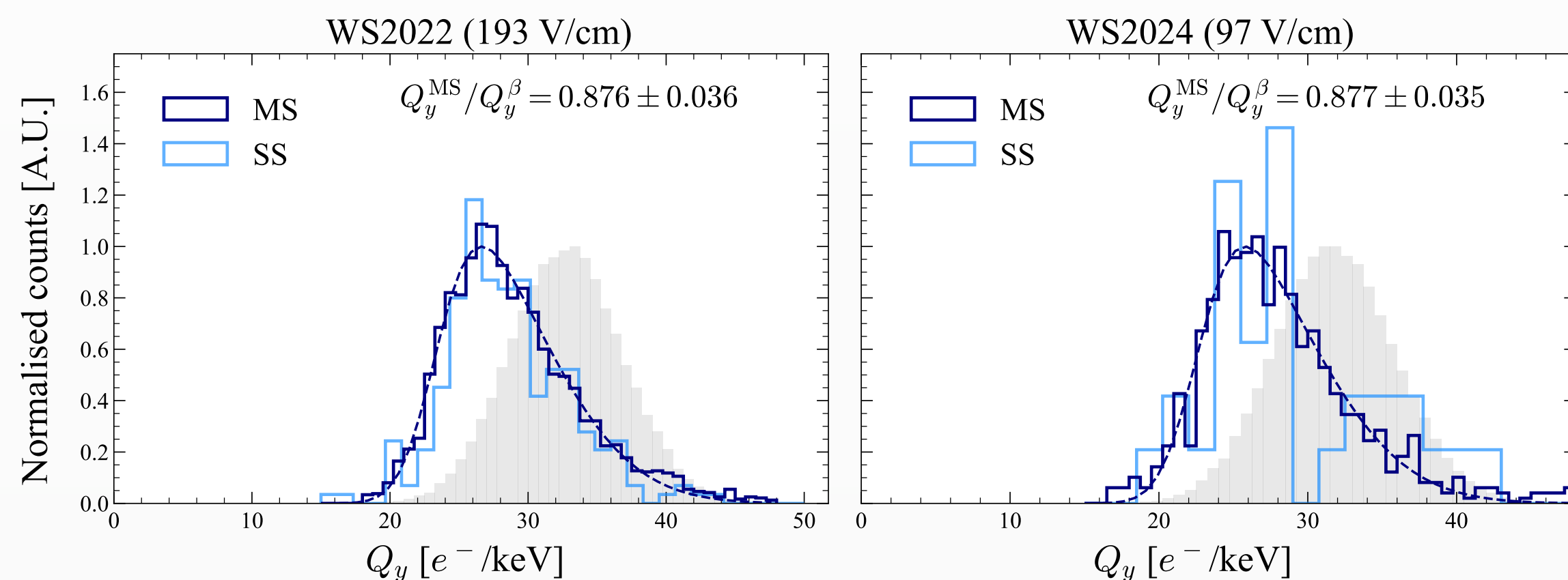
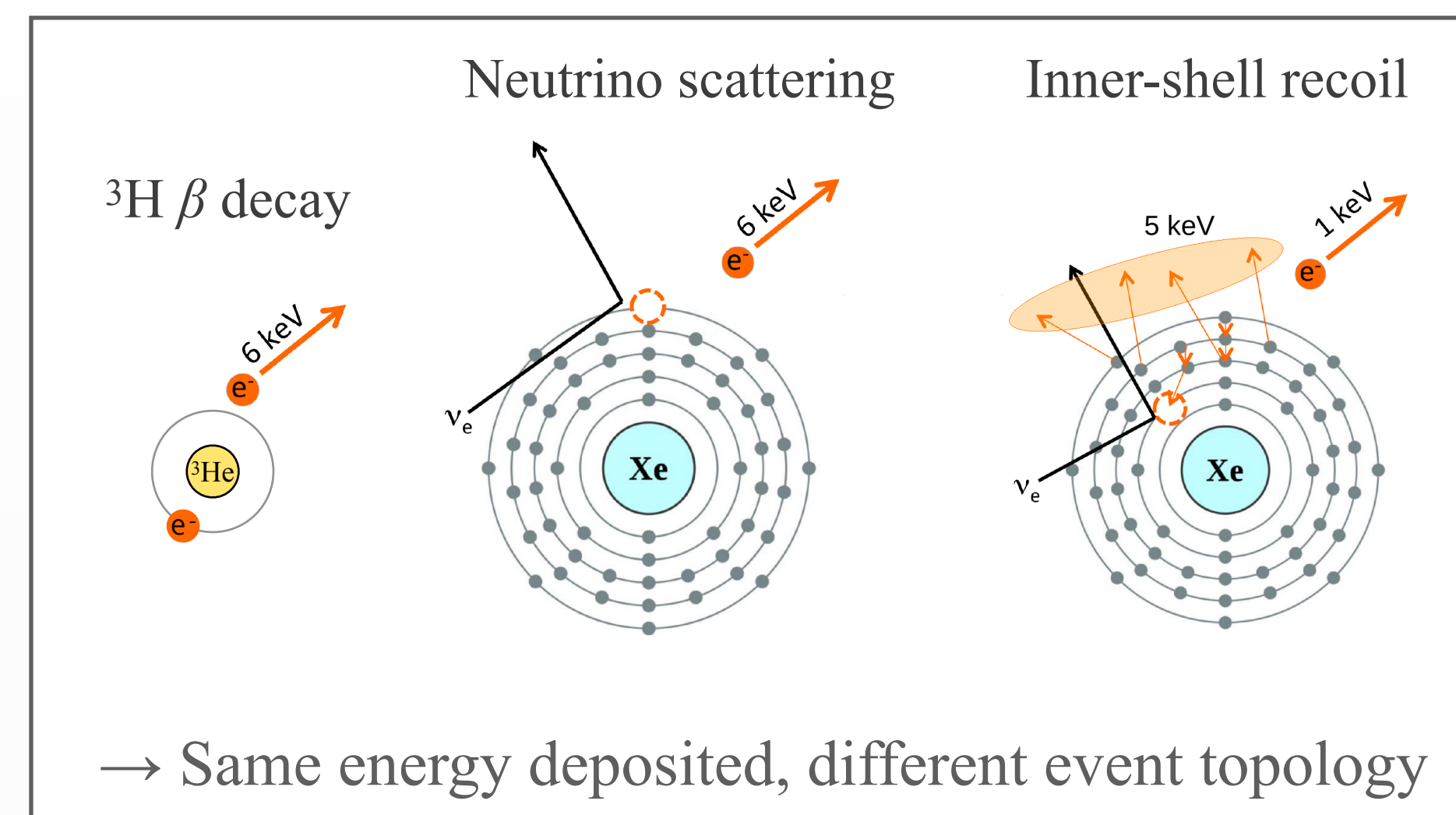
Xenon flow in the TPC: Radon tag

- “**Radon tag**” to identify events with high likelihood of being ^{214}Pb decays. Based on simulations of neutral and charged progenies using these velocity fields during low-mixing state
 - ➔ Tag volume defined for each identified ^{218}Po α decay along streamlines for $3 \times T_{1/2}(^{214}\text{Pb})$
- ^{214}Pb in untagged subsample $1.8 \pm 0.3 \mu\text{Bq/kg}$ ($3.9 \pm 0.6 \mu\text{Bq/kg}$ in total exposure $\rightarrow \sim 60\%$ reduction)
- Both tagged and untagged data included in WIMP analysis $\rightarrow$ no “loss of exposure”



Electron Capture (EC) Decay Backgrounds

- Events with X-ray/Auger electrons subject to **charge yield attenuation** due to higher ionisation density → **more NR-like response**
- In LZ, EC from ^{127}Xe and ^{125}Xe (activated by calibration neutrons, or cosmogenically-produced)
- Especially relevant xenon L-shell @5.2 keV



- In-situ LZ measurement with WS2024 data (at 97 V/cm)
- Using MS event selection:

$$Q_L/Q_{\beta} = 0.877 \pm 0.035$$



NEW FEATURE

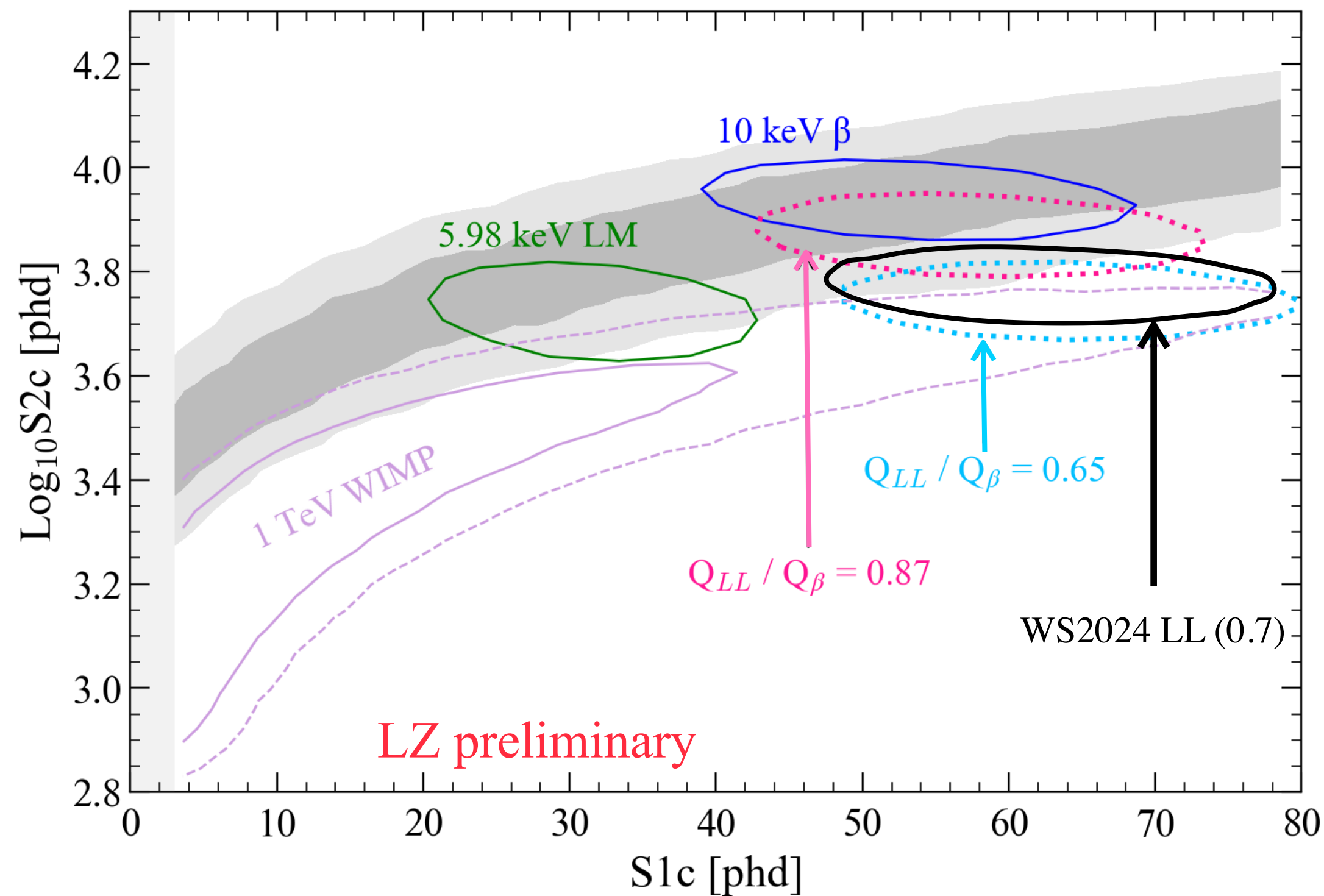
PRD 112, 012024 (2025)

EC Decay Backgrounds: ^{124}Xe DEC

- ^{124}Xe two-neutrino DEC measured by LZ: $T_{1/2} = (1.09 \pm 0.15(\text{stat}) \pm 0.05(\text{syst}) \times 10^{22})$ yr
- LM (6.0 keV) & LL (10.0 keV) shells relevant for WS

[J. Phys. G: Nucl. Part. Phys. 52 015103 \(2025\)](#)

- Expect 7.1 (LM) + 12.3 (LL) = 19.4 counts
- LM modelled similarly as single L-shell charge suppression
- Expect LL captures to display further charge yield suppression
- Background model allows LL-capture to float in the final fit:



$0.65 < Q_{LL}/Q_{\beta} < 0.87$

Q_L/Q_{β}

2x L-shell ionisation density

Best fit value: 0.70 ± 0.04

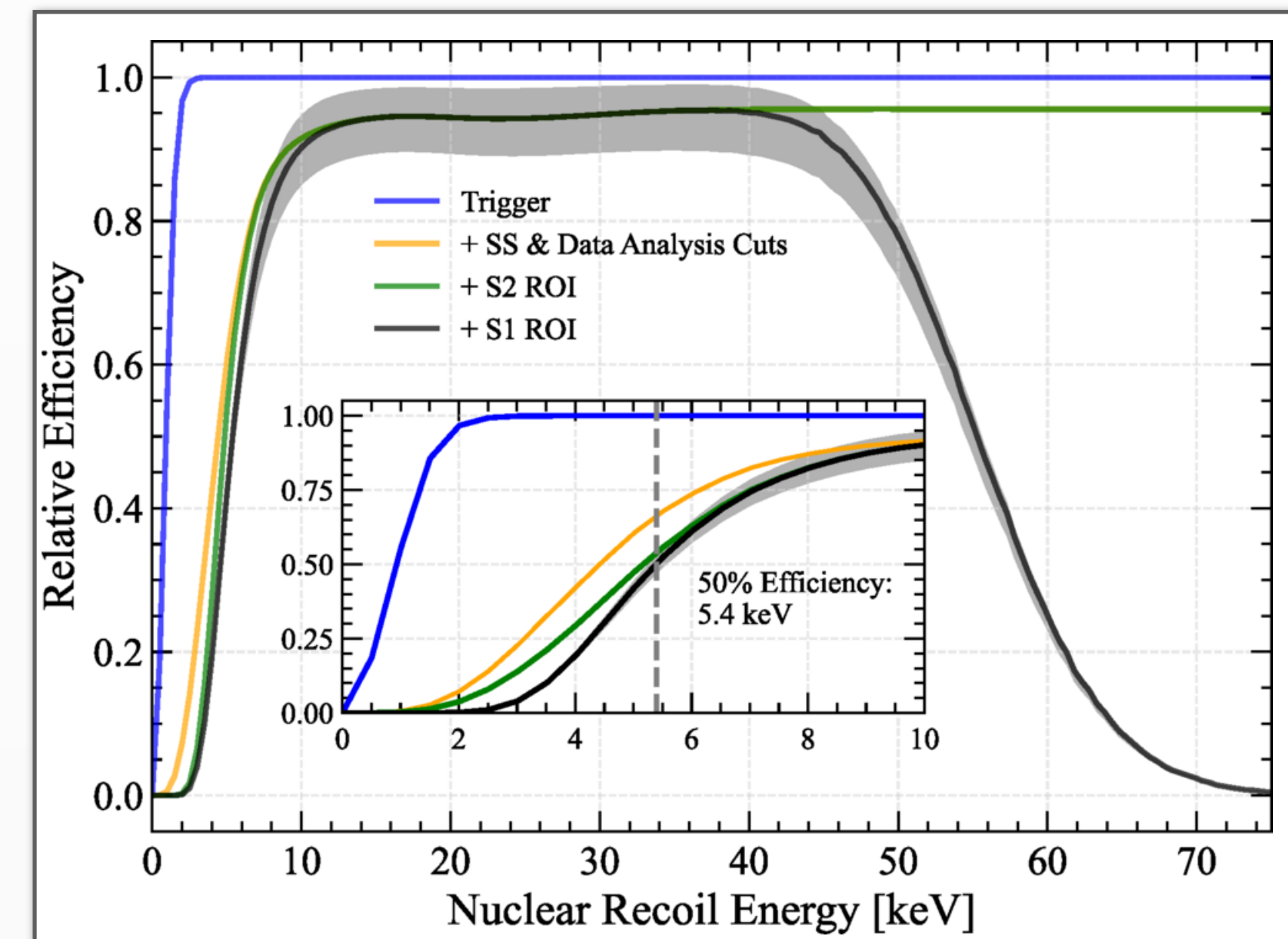
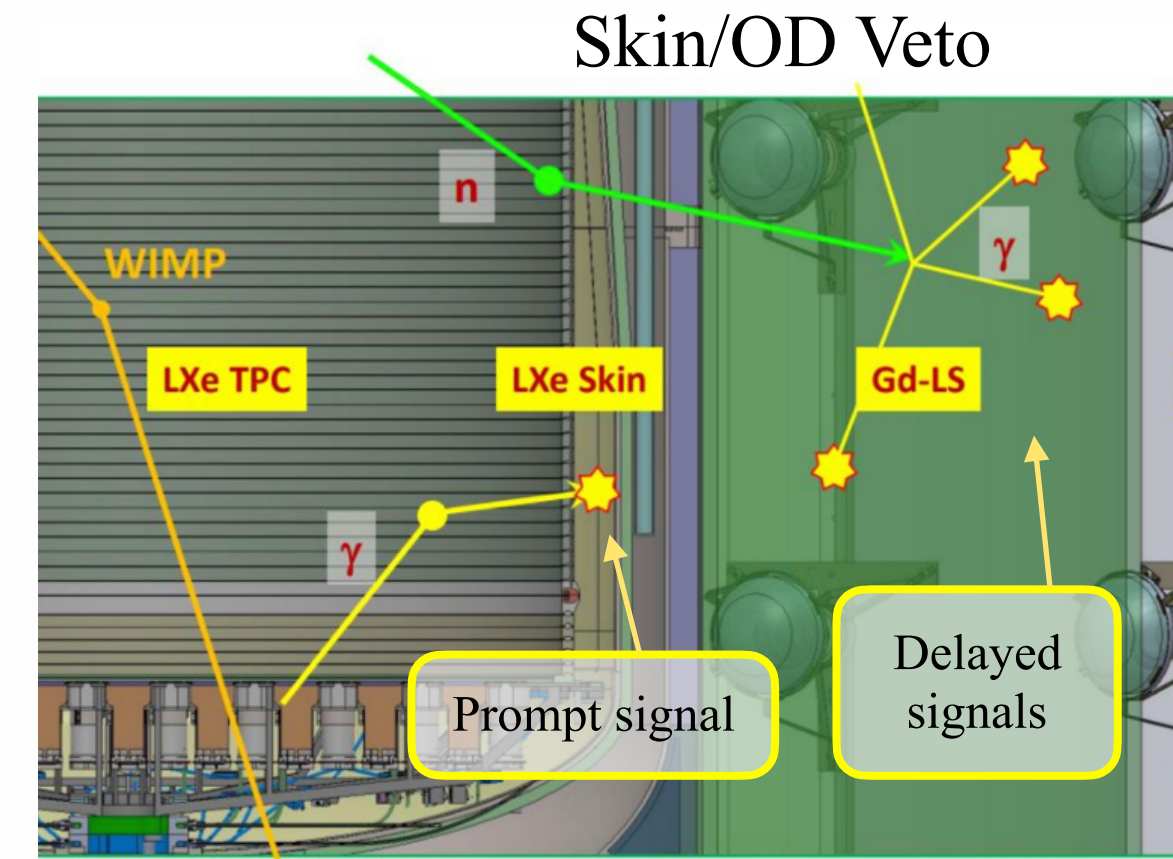
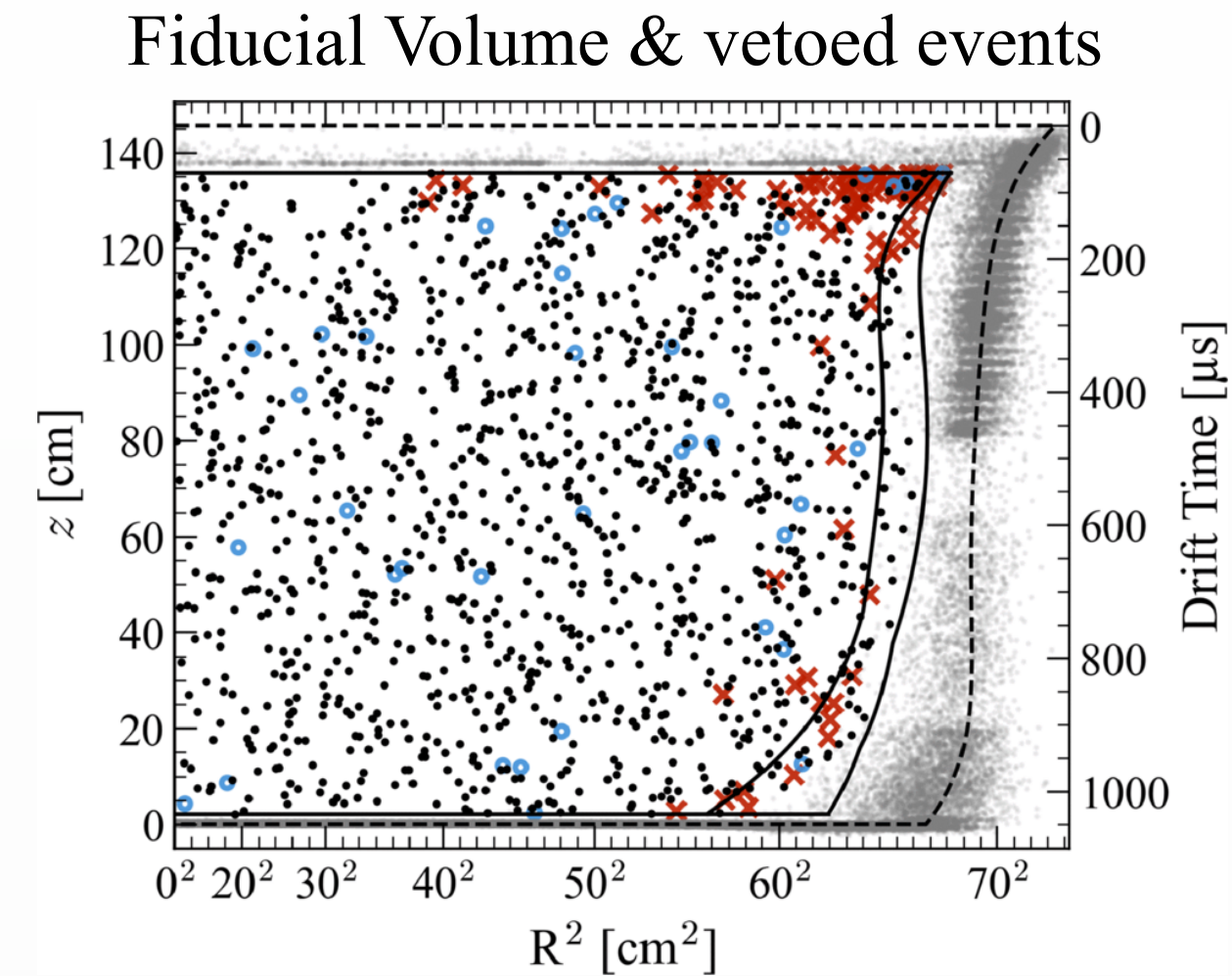


Analysis cut summary

- Cuts developed on non-WIMP ROI data. Include:

- Single Scatter (SS) events
- Region of interest (ROI) used:
 - S1 (>3 PMT) ROI: $3 < S1c < 80$ phd
 - S2 ROI: $\begin{cases} 645 \text{ phd } (\equiv 14.5 \text{ electrons}) < S2 \\ \log_{10}(S2c) < 4.5 \end{cases}$
- Time exclusions**: periods of high-rate, detector instability, hold-off after large S2s $\rightarrow$ 86% livetime retention
- S1- and S2-based cuts** (target pulse accidentals)
- Fiducial Volume cut** (5.5 ± 0.2 t)
- Events not **tagged** by **Skin** (x) or **OD** (o) in prompt/delayed time window optimised to tag γ -rays and neutrons

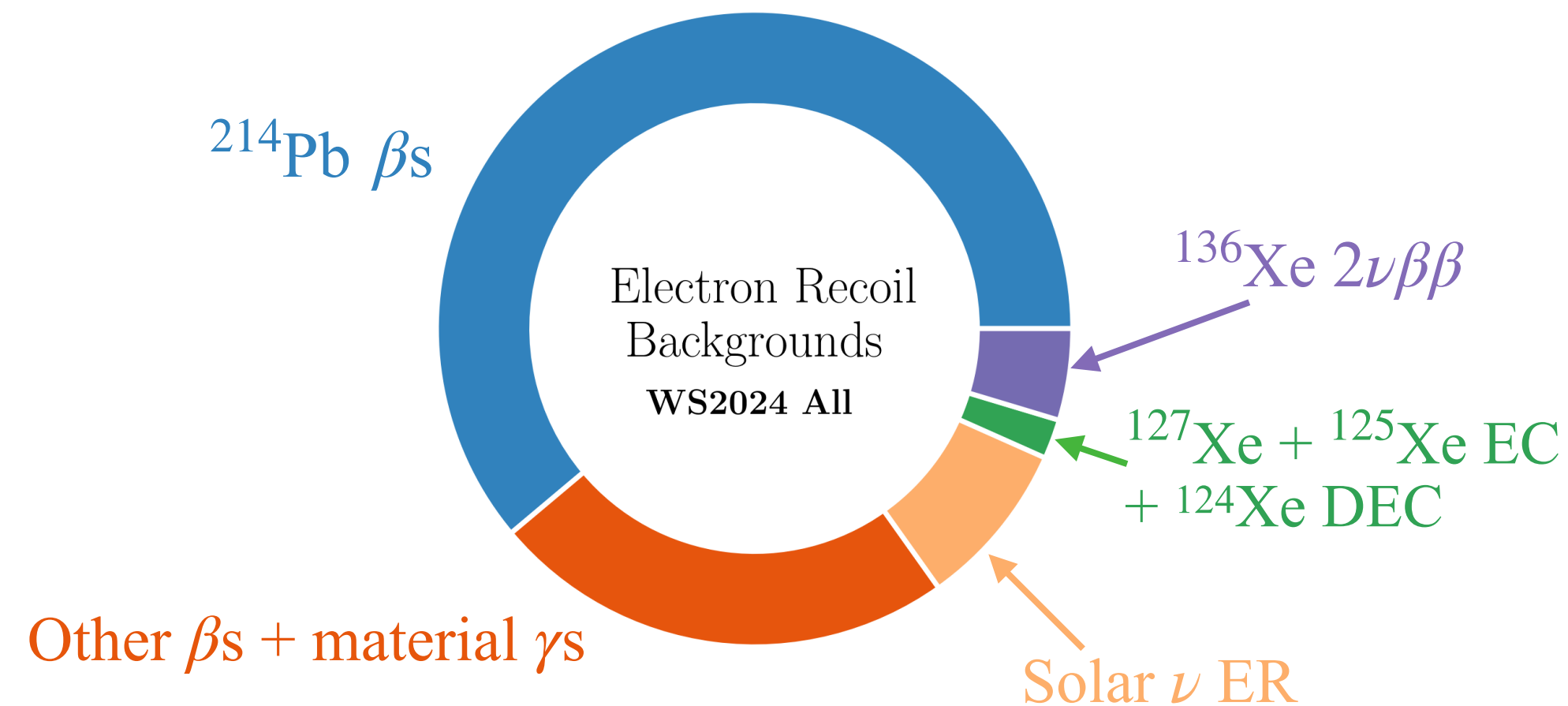
- Efficiency of cuts vs NR energy:



Background model

ER backgrounds

- ▶ Dissolved β emitters:
 - **^{214}Pb** (^{222}Rn), ^{212}Pb (^{220}Rn), ^{85}Kr , ^{136}Xe ($\beta\beta$)
- ▶ Dissolved **EC decays** (x-rays/Auger cascades):
 - $^{127/125}\text{Xe}$ from neutron calibration activation
 - ^{124}Xe double EC
- ▶ Solar pp and ^7Be neutrinos
- ▶ Long-lived gamma emitters in detector materials:
 - ^{238}U chain, ^{232}Th chain, ^{40}K , ^{60}Co



Expected ER counts in WS2024 exposure: **1207**

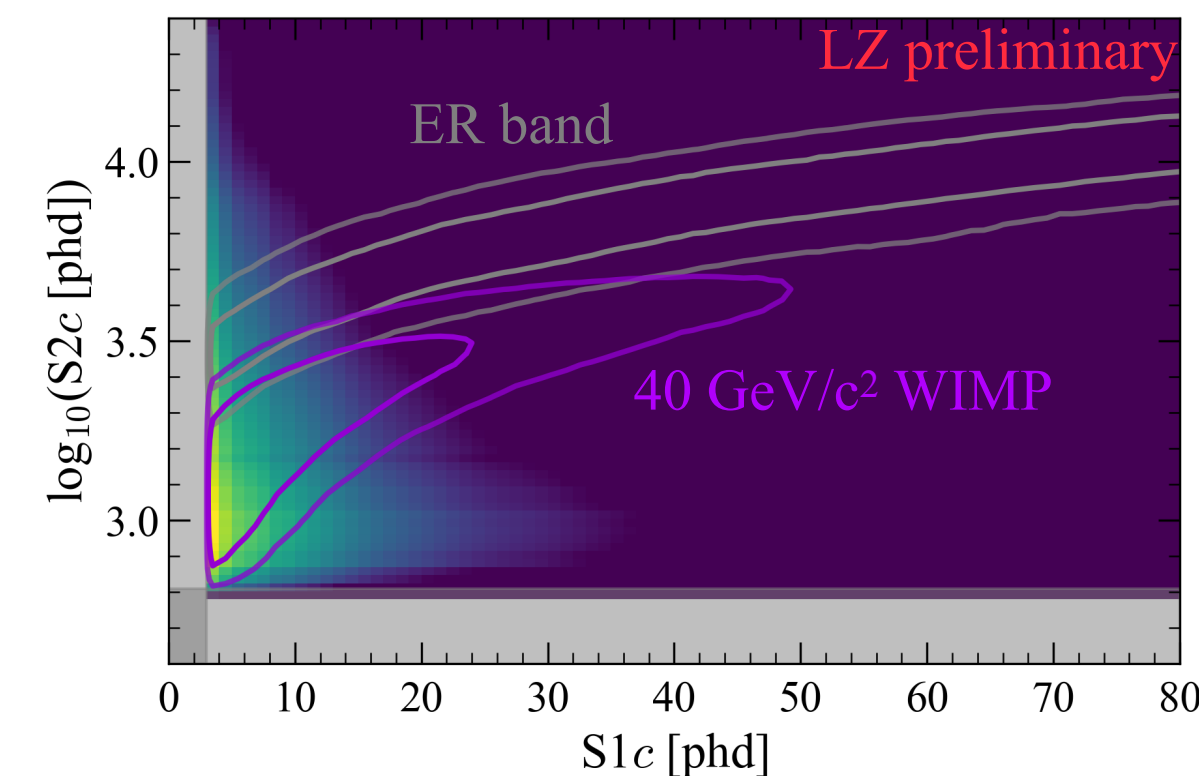
NR backgrounds

- ▶ Neutrons from spontaneous fission (SF) and (α, n) reactions in detector materials
- ▶ Solar ^8B and hep neutrinos (CEvNS)

Expected NR counts in WS2024 exposure: **0.18**

Instrumental backgrounds

- ▶ Accidental coincidences (uncorrelated S1/S2 pulses),

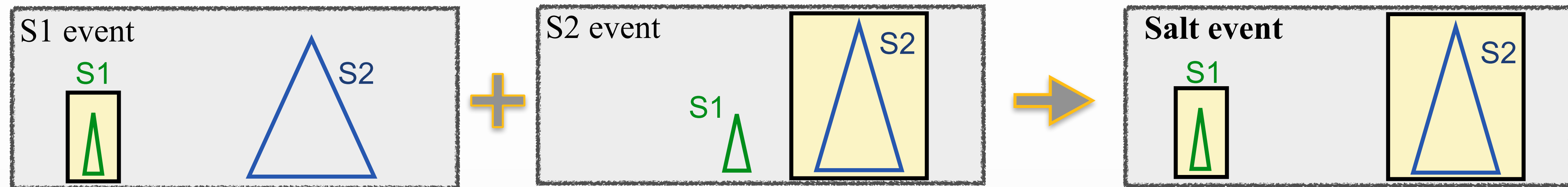


Data-driven
accidentals PDF
Rate derived from
Unphysical Drift
Time events

Expected ER counts in WS2024 exposure: **2.8**

Experimenter bias mitigation: salting

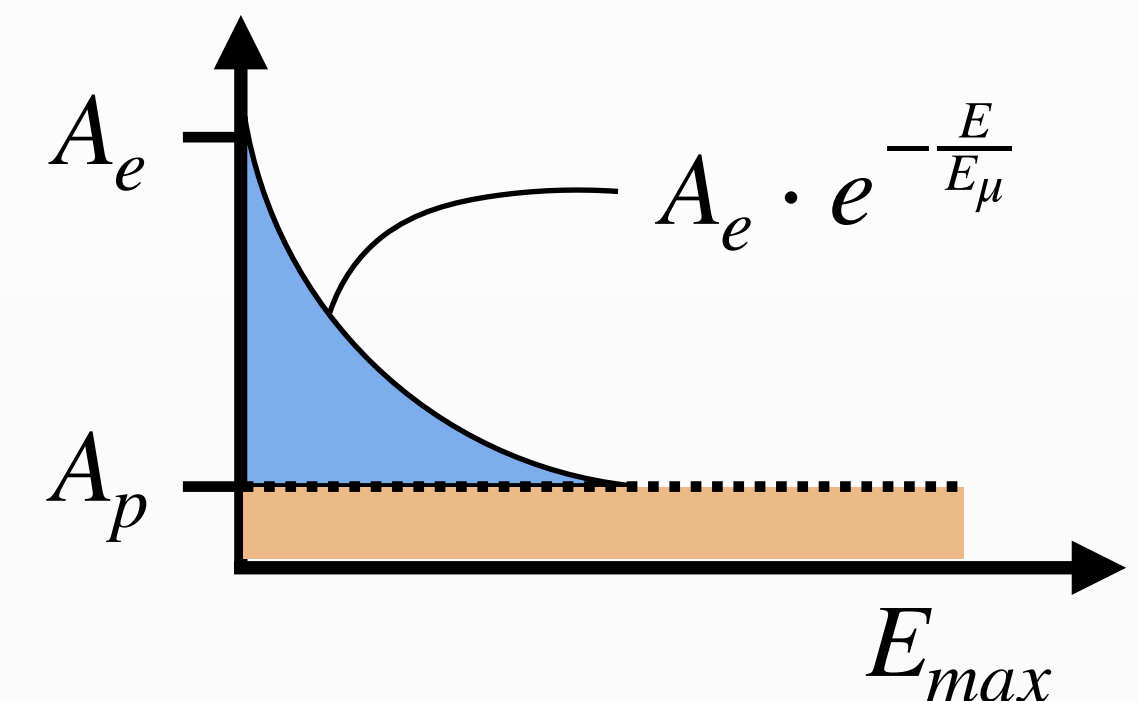
- “*Salting*”: injecting a random number of fake signal events during science data collection
 - How? Events manufactured by combining S1s & S2s signals from sequestered calibration data



- Which energy? Exponential distribution that represents WIMP scattering (of randomly selected mass $\rightarrow E_\mu \propto m_\chi$) + flat pedestal to higher energies (A_p) for higher NR energies analyses
- How many? Enough to obscure any bias, but not so many as to overwhelm any pathological events in the real data: random number between

$[0, N]$, N from WS2022 upper limit

Identity revealed after analysis input are finalised for final statistical inference



$$A_e = \alpha \times N$$

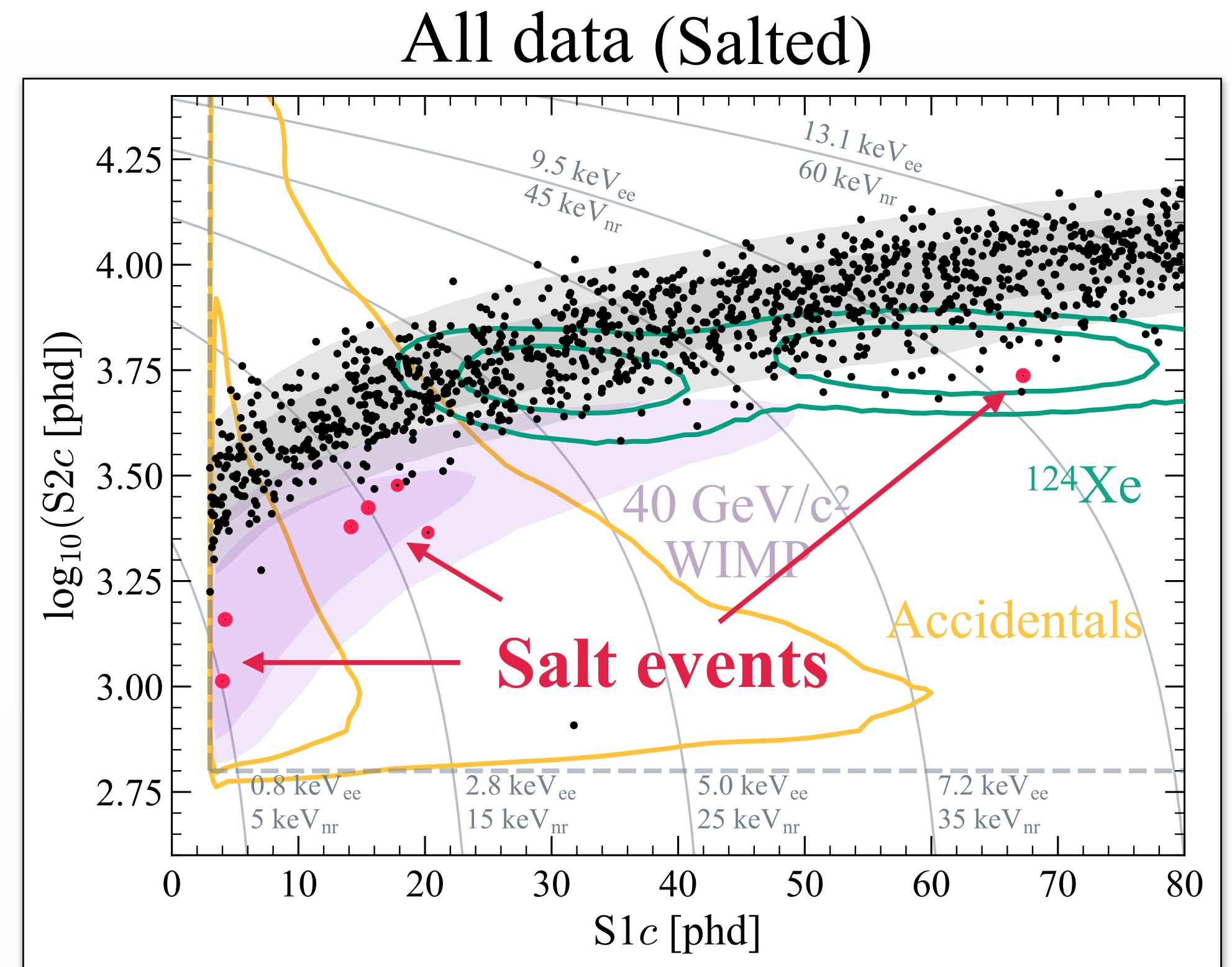
$$A_p = 1 - A_e$$



WS2024 Final Dataset

- 8 salt events revealed → **7 of which pass all analysis cut** (consistent with evaluated signal efficiency)
- **1220** events remain after unsalting
- No post-unsalting changes to model
- Final exposure:
220 live days x 5.5 tonnes = **3.3 tonne yr**

- Statistical analysis in observed $S1c$ - $\log_{10}(S2c)$ space
- Likelihood fed with 6 subsamples: WS2024 (with 4 Rn-tag configurations), Skin/OD veto sideband, WS2022



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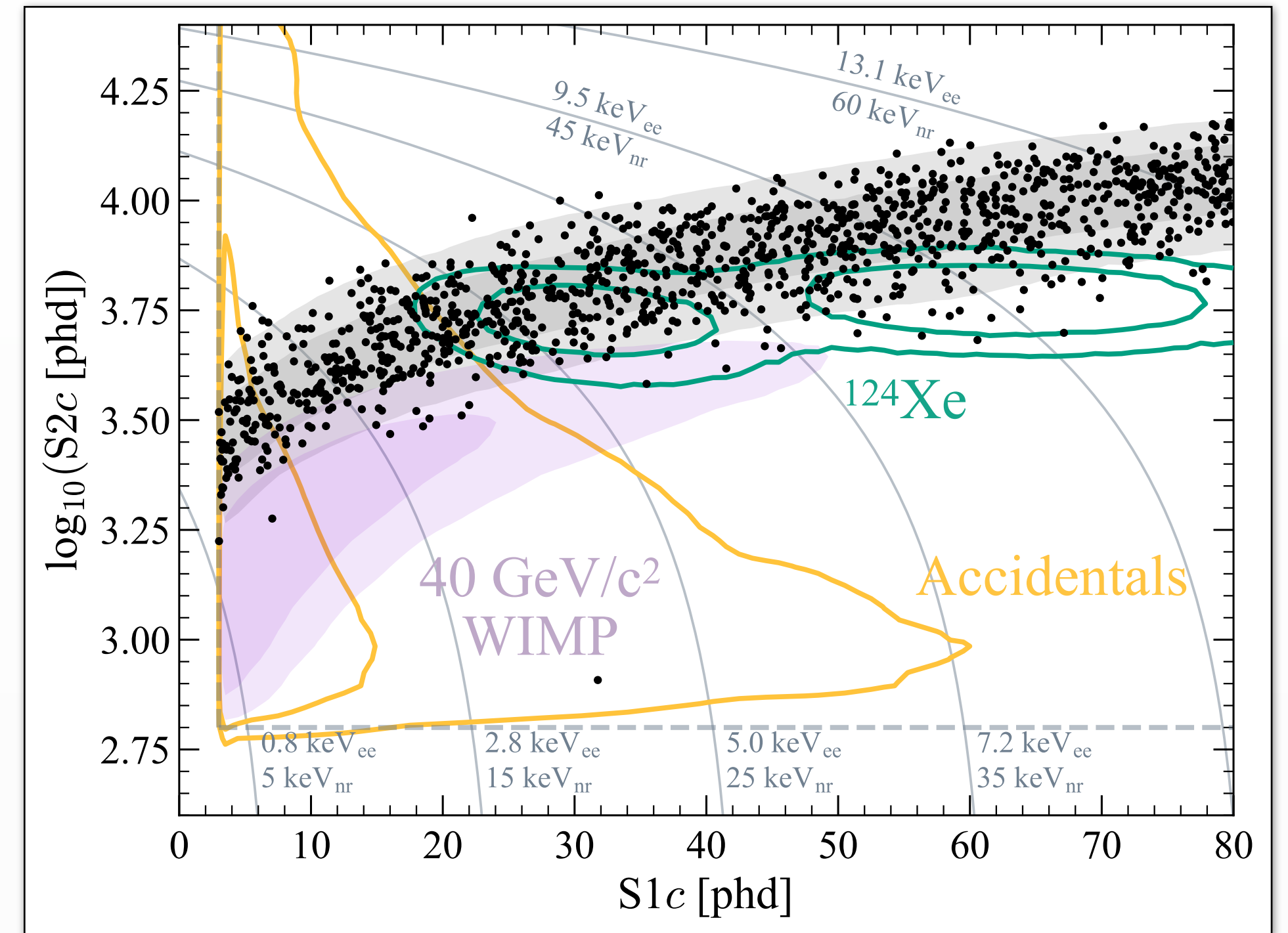
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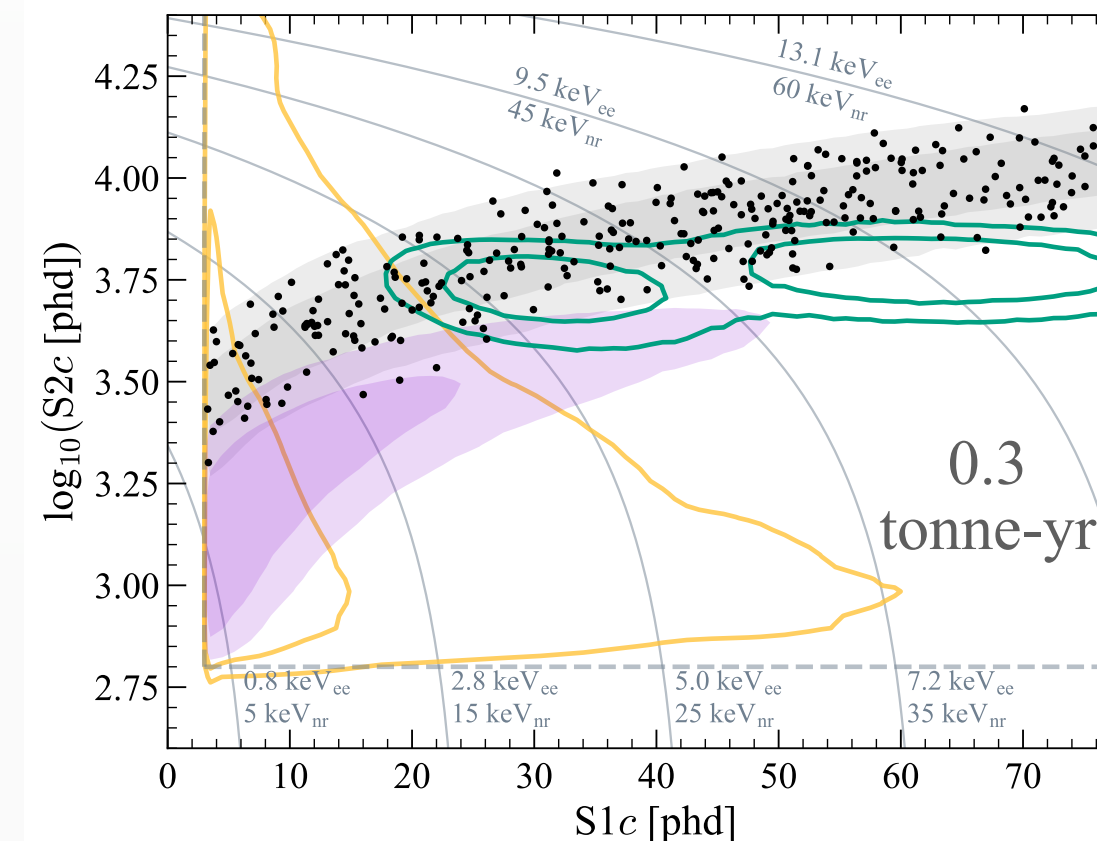
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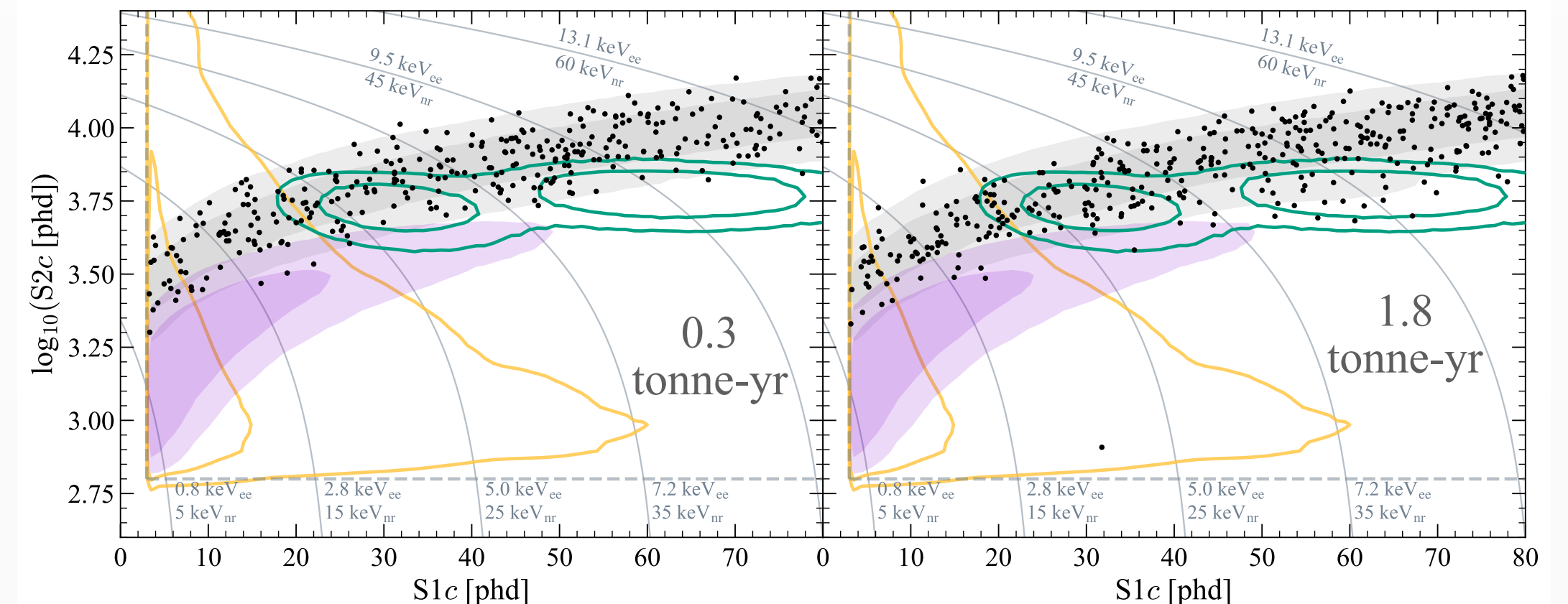
All data



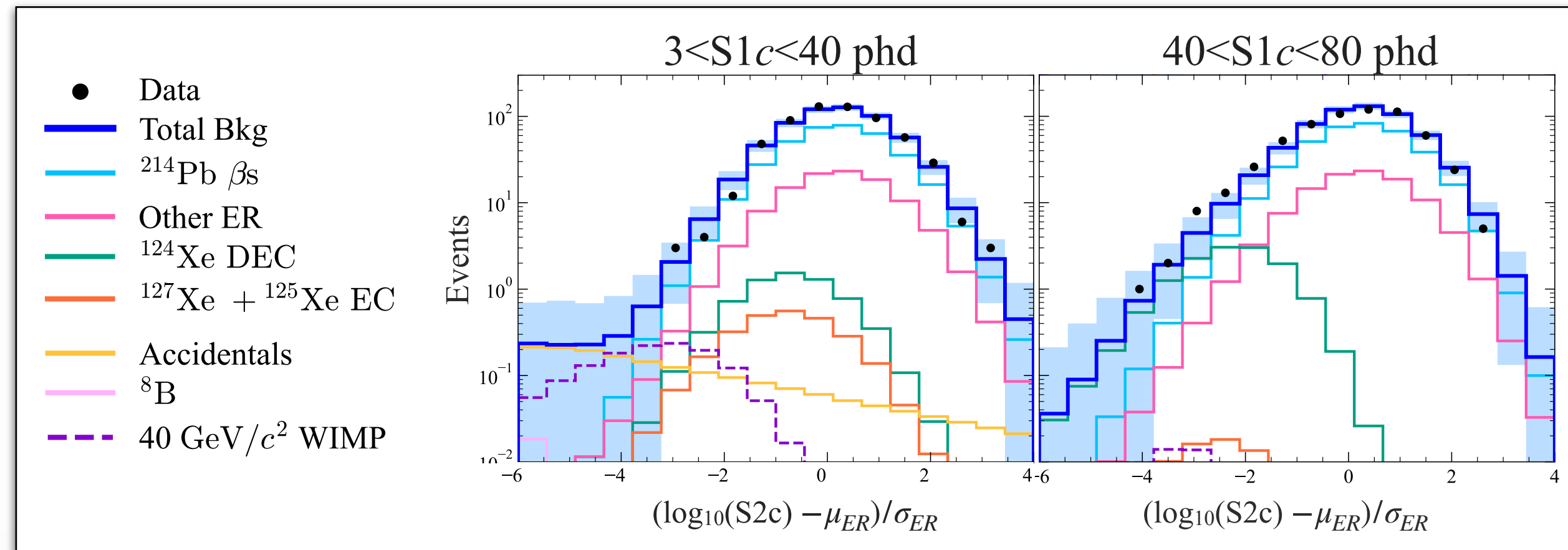
Radon-Tagged subsample



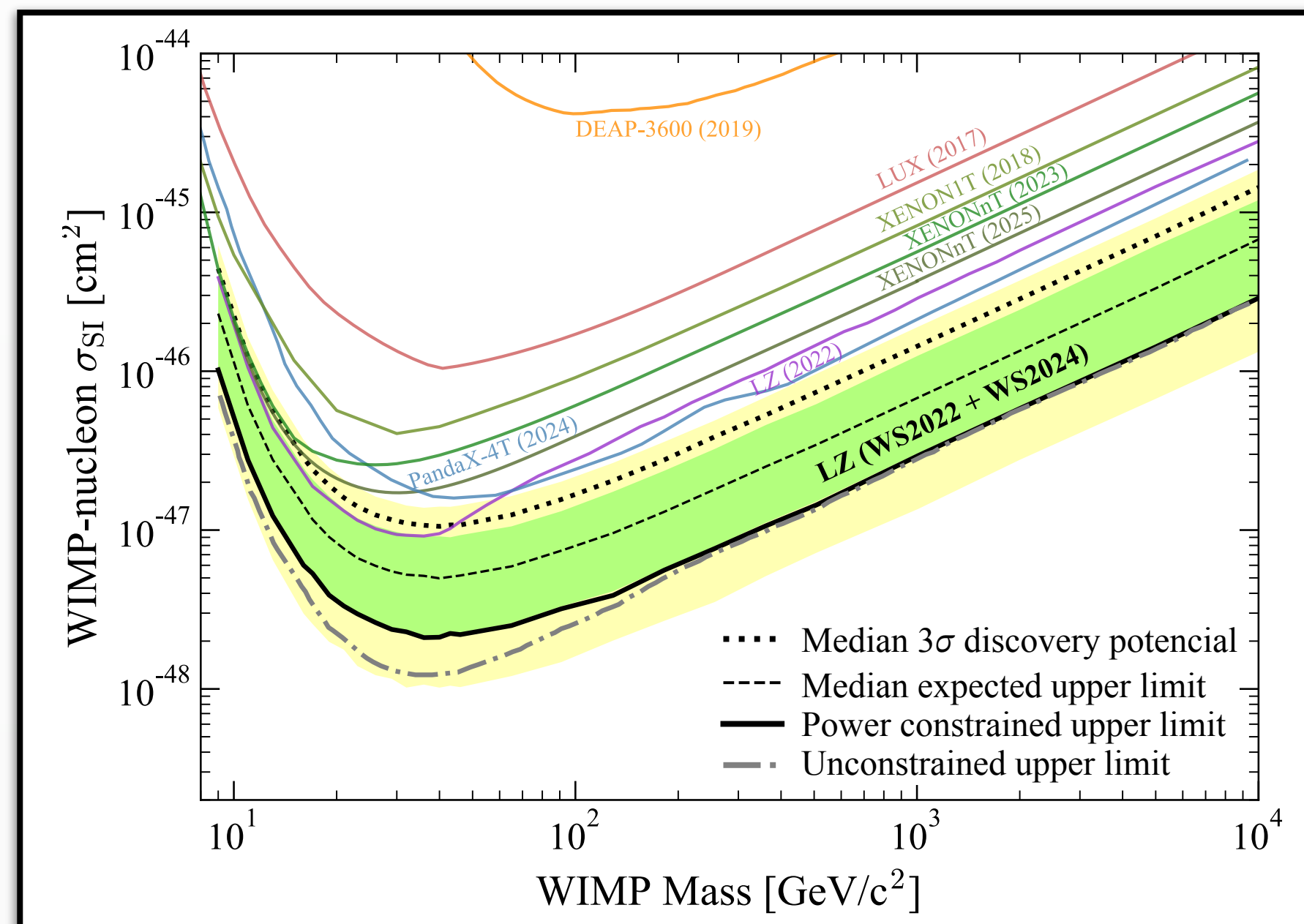
Radon-Untagged subsample



WS2024+WS2022 Fit result and SI limits



- Good data/model agreement in all spaces
- Best fit of **zero WIMPs** at all tested masses (9 GeV/c² - 100 TeV/c²)
- Two-sided unbinned profile likelihood ratio test statistic
- Upper limit power constrained at -1σ sensitivity band (as DM conventions [EPJC 81 907 \(2021\)](#))
- Under fluctuations from distribution of accidental events in WIMP region (in WS2024) + ER under fluctuation in WS2022

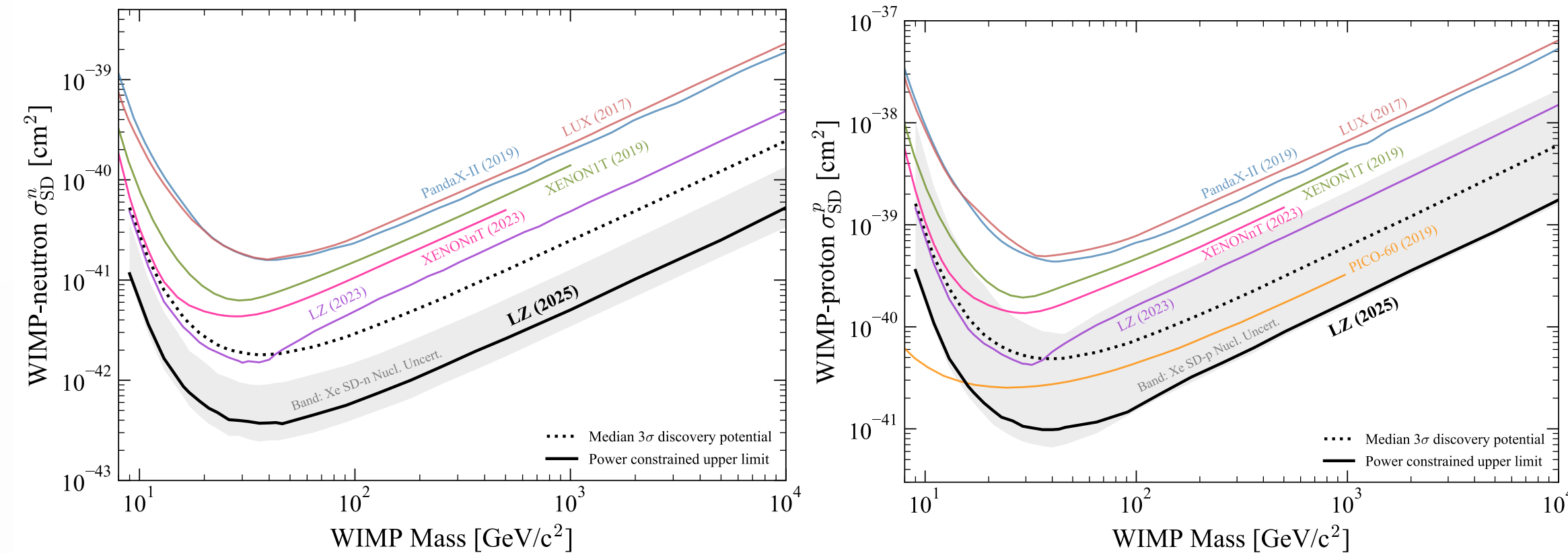


WS2024+WS2022 combined minimum
limit: **2.2 x 10⁻⁴⁸ cm² at 43 GeV/cm²**



Other LZ results and prospects

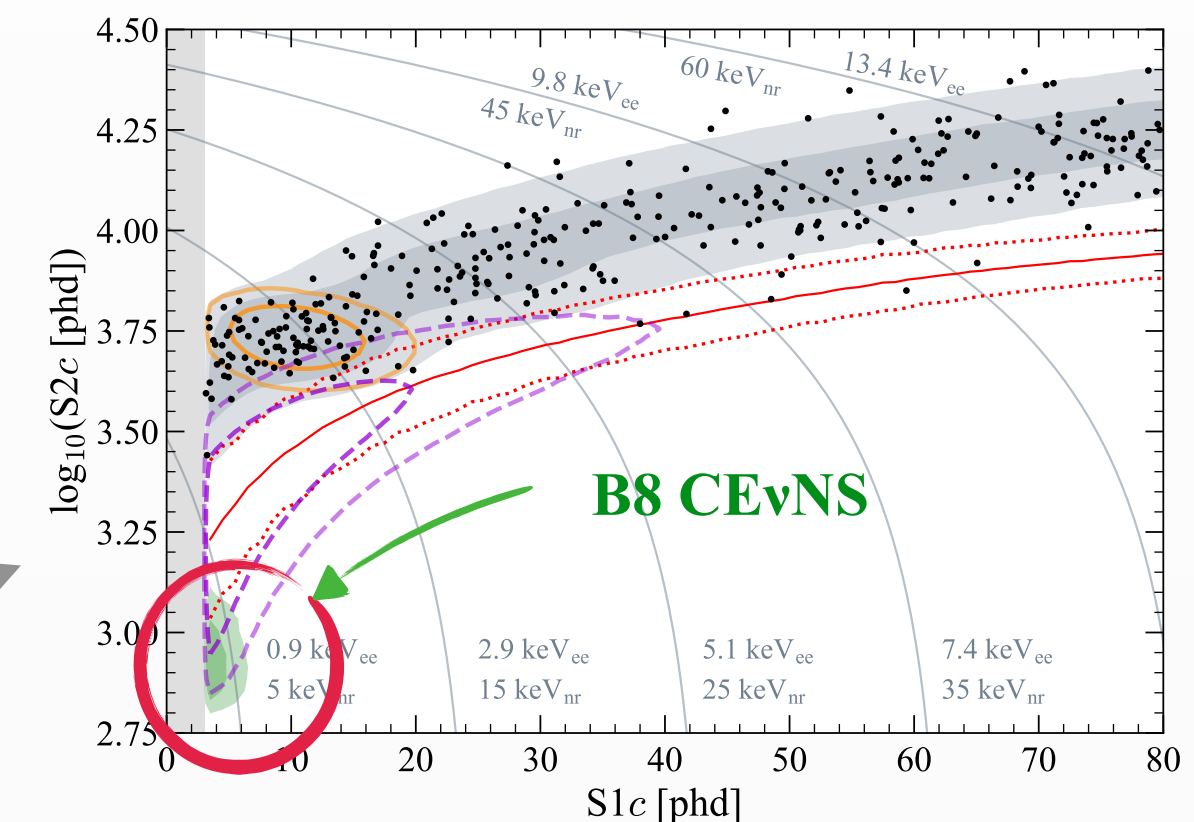
- Similar analysis to search for Spin-Dependent (SD) WIMP-nucleon interactions. Two limiting cases considered: WIMP-neutron, and WIMP-proton
- Atmospheric millicharge particles [PRL 134, 241802 \(2025\)](#)
- WIMP-pion [Commun Phys 7, 292 \(2024\)](#)
- Cosmic ray-boosted Dark Matter [PRL 134, 241801 \(2025\)](#)
- Covariant EFT [PRL 133, 221801 \(2024\)](#)
- And more on the horizon... exploring other energy ranges:



- Low-mass WIMPs / B8 neutrinos via CEvNS

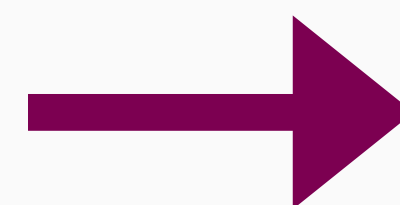
- Neutrinoless Double Beta Decay

See E. Jacquet talk on Wednesday!!



[arXiv:2410.17137](#)

- Further into the future, much larger exposures needed to fully cover the space parameter until the neutrino fog ...



60-80 t LXe TPC
LZ+XENON+DARWIN
 ~500 people / 72 institutions



Conclusions & Prospects



WIMP search 2024

- LZ is the world's most sensitive WIMP direct detection experiment:

New world-leading WIMP search limits using WS2024+WS2022 4.2 tonne year exposure

- **$2.2 \times 10^{-48} \text{ cm}^2$ at 43 GeV/cm^2** , exceeding previous best constraints by **>4** times
- Radon-tag technique developed and exploited for the first time in a dark matter result, reducing 60% the primary ER background
- First observation of suppressed charge yield from LL-shell captures of ^{124}Xe

Future

- LZ to keep taking data until 2028, with the goal of 1000 live days
- Multiple other physics searches on the horizon: ^8B CEvNS, neutrino less double beta decay, ER-based searches, and more!



Thank you for listening!

And to our sponsors and 38 participating institutions:



@lzdarkmatter



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LZ Collaboration Meeting at SURF, June 2023