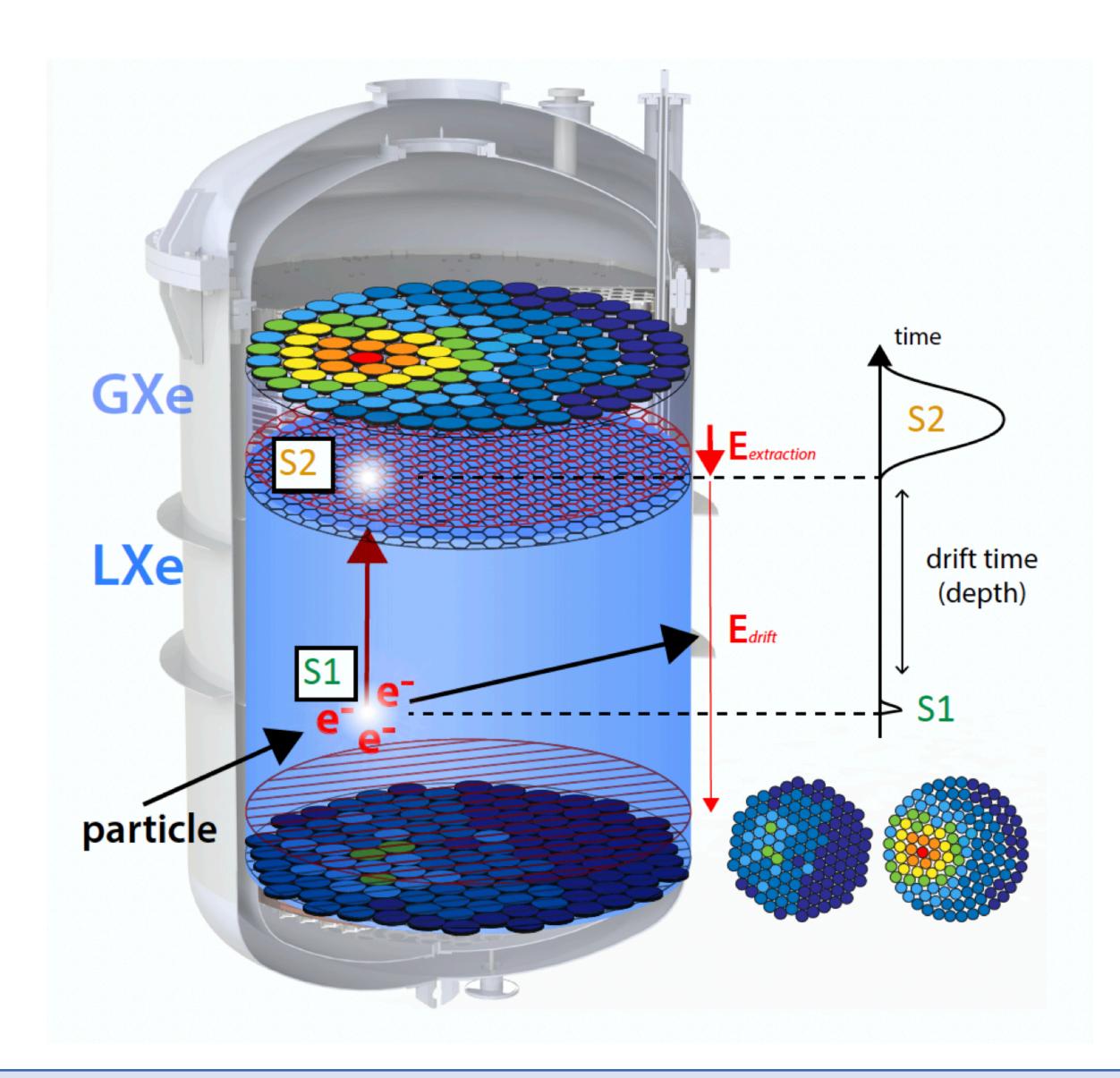
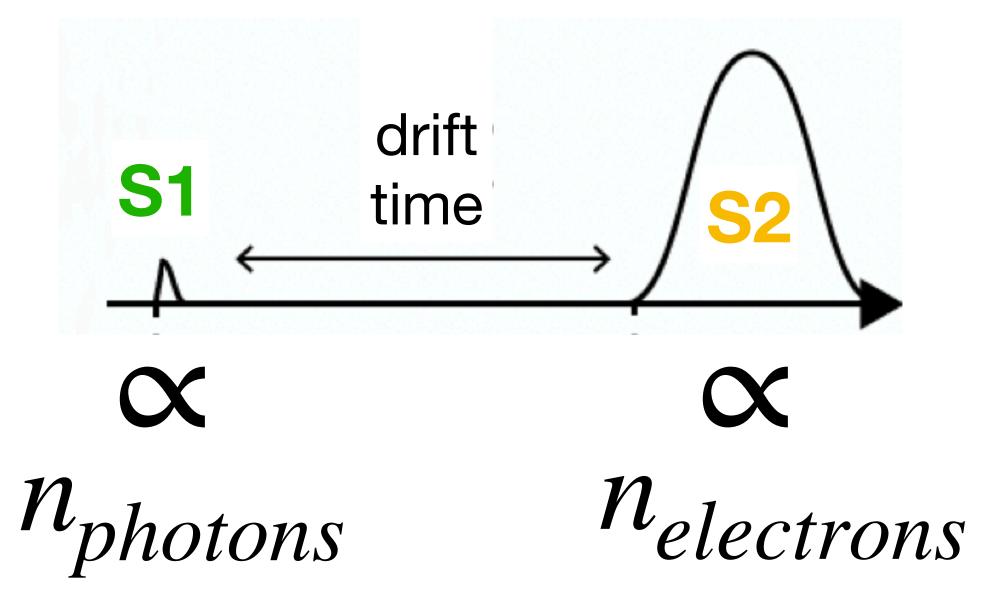


XENONnT Time Projection Chamber (TPC)



Signals in XENON:

- Each interaction generates two signals:
 \$1 from directly emitted photons,
 \$2 from drifted and extracted electrons
- Light and charge yields depend on:
 - Drift field (electron-ion recombination)
 - Total interaction energy



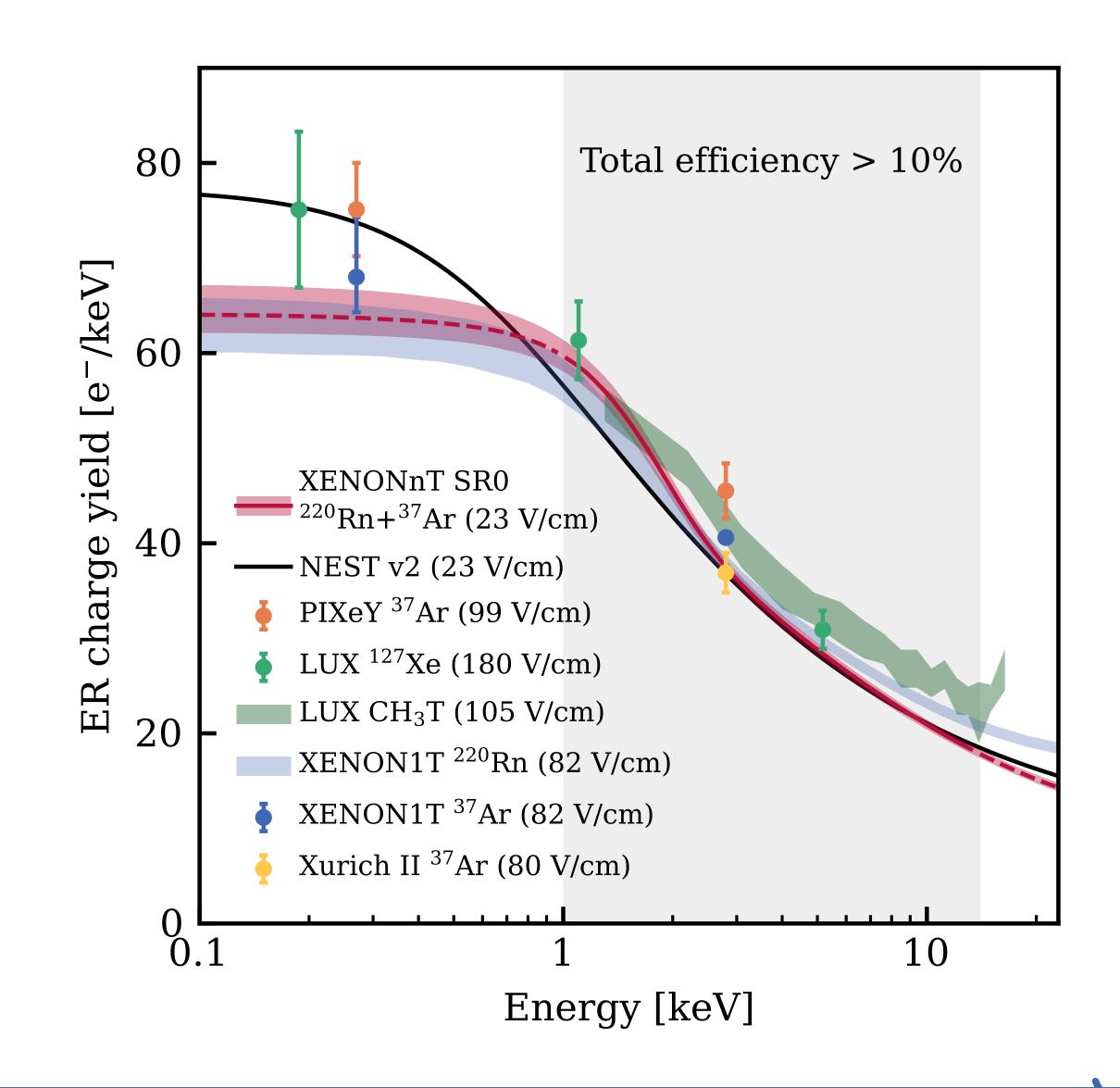


XENONnT LowEnergy ER Yields

Light and Charge yields:

- Precise modeling of yields is necessary for any search of rare events in xenon TPCs.
- The **focus** is often limited on a scale between 1 and 10 keV, **really low energies**!

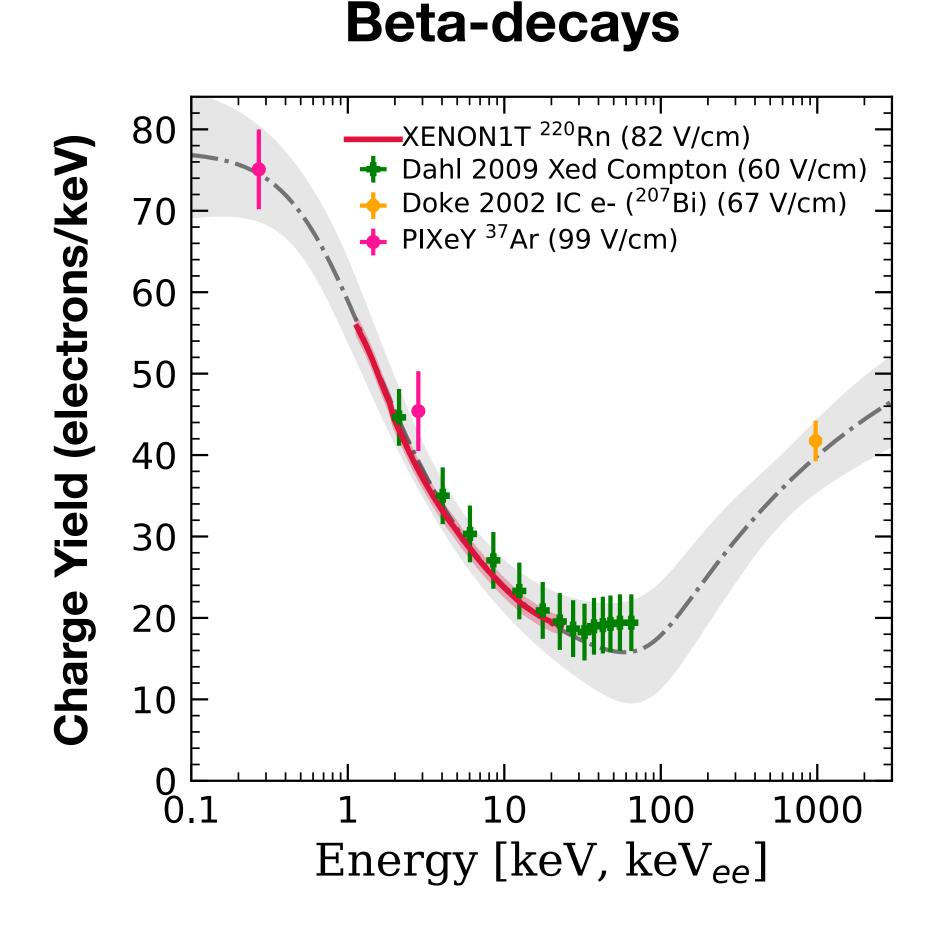
 Relevant scale for WIMP or CEvNS scattering, but we want to do much more with our data!
- Many high energy analyses like 2vbb/0vbb and many more, require characterisation of detector response upto few MeVs.



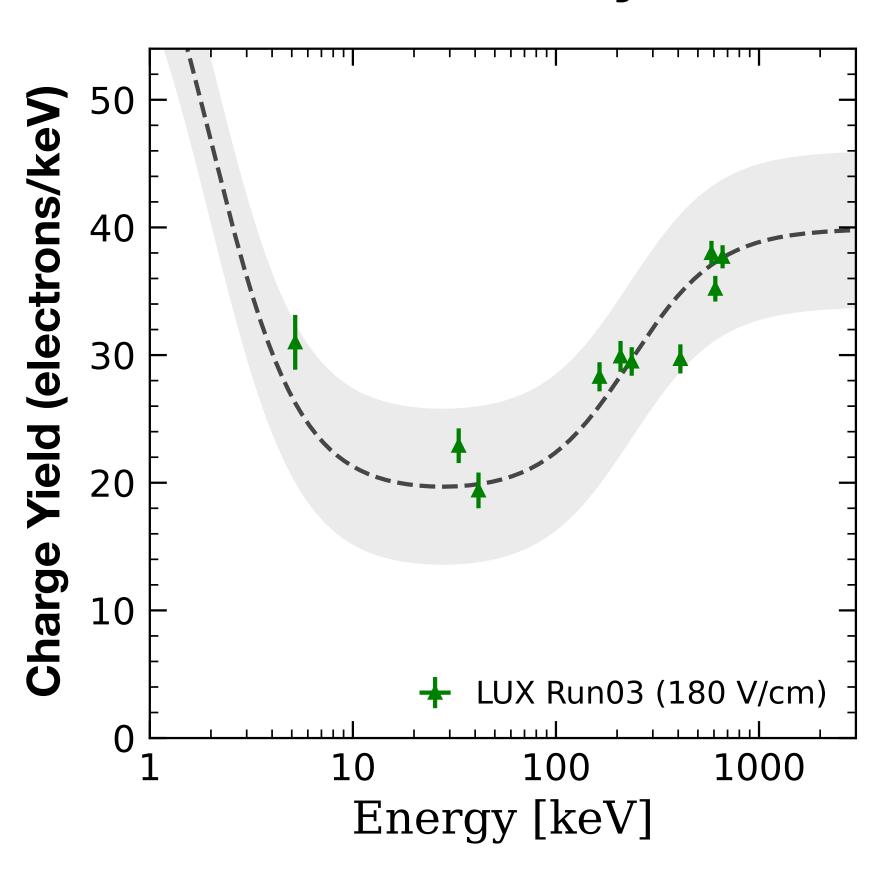
NEST Yields: builing the model from data

NEST Model:

- Fully empirical, based on combination of Thomas-Imel and Doke-Birks models.
- Aggregates data from many experiments, different fields.
- Separate model for beta and gamma. Compton makes multiple scatters, effective lower gamma charge yield.
- Missing data at high energies!



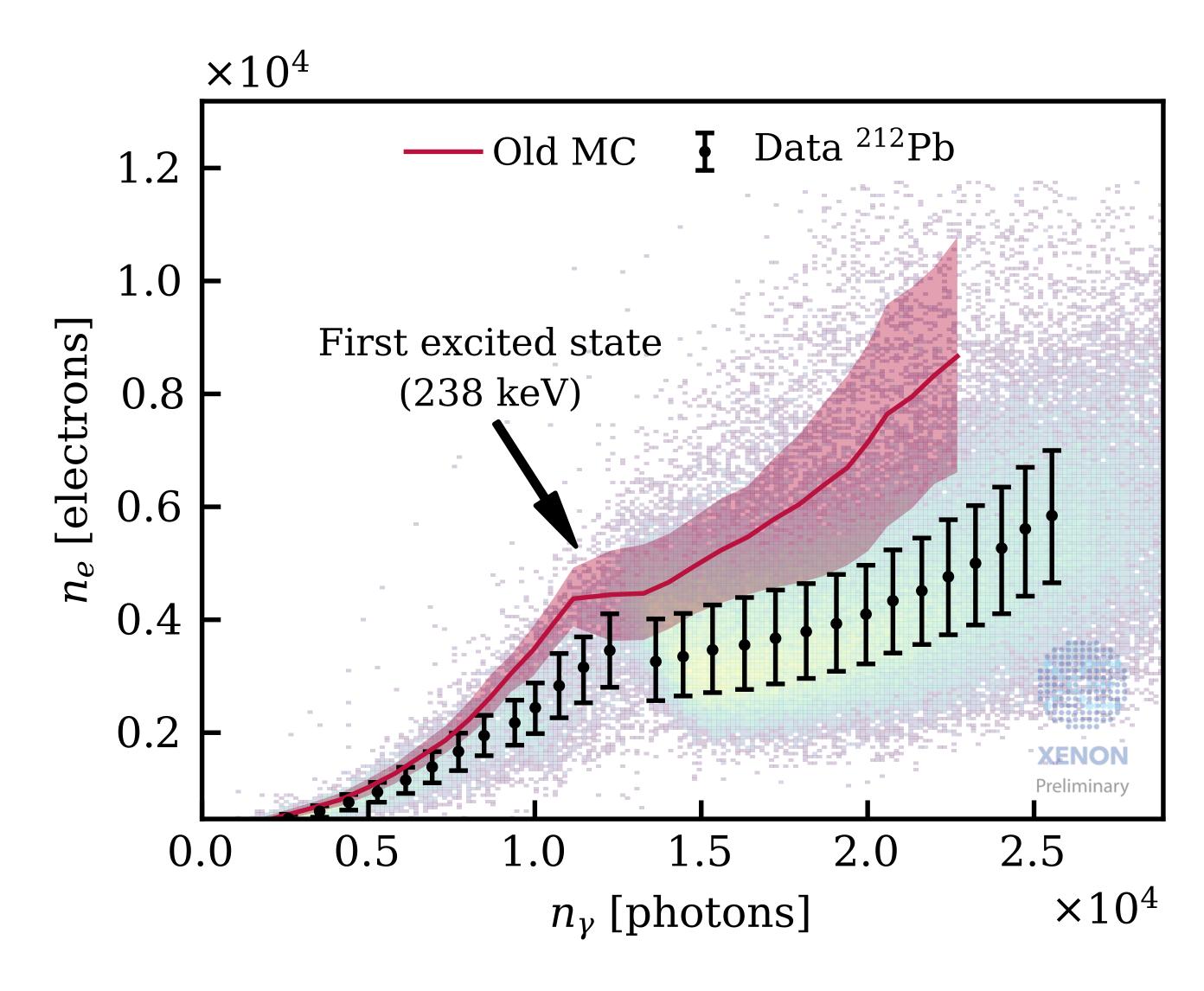
Gamma-decays



How do we model the response at higher energies?

Check on ²¹²Pb data:

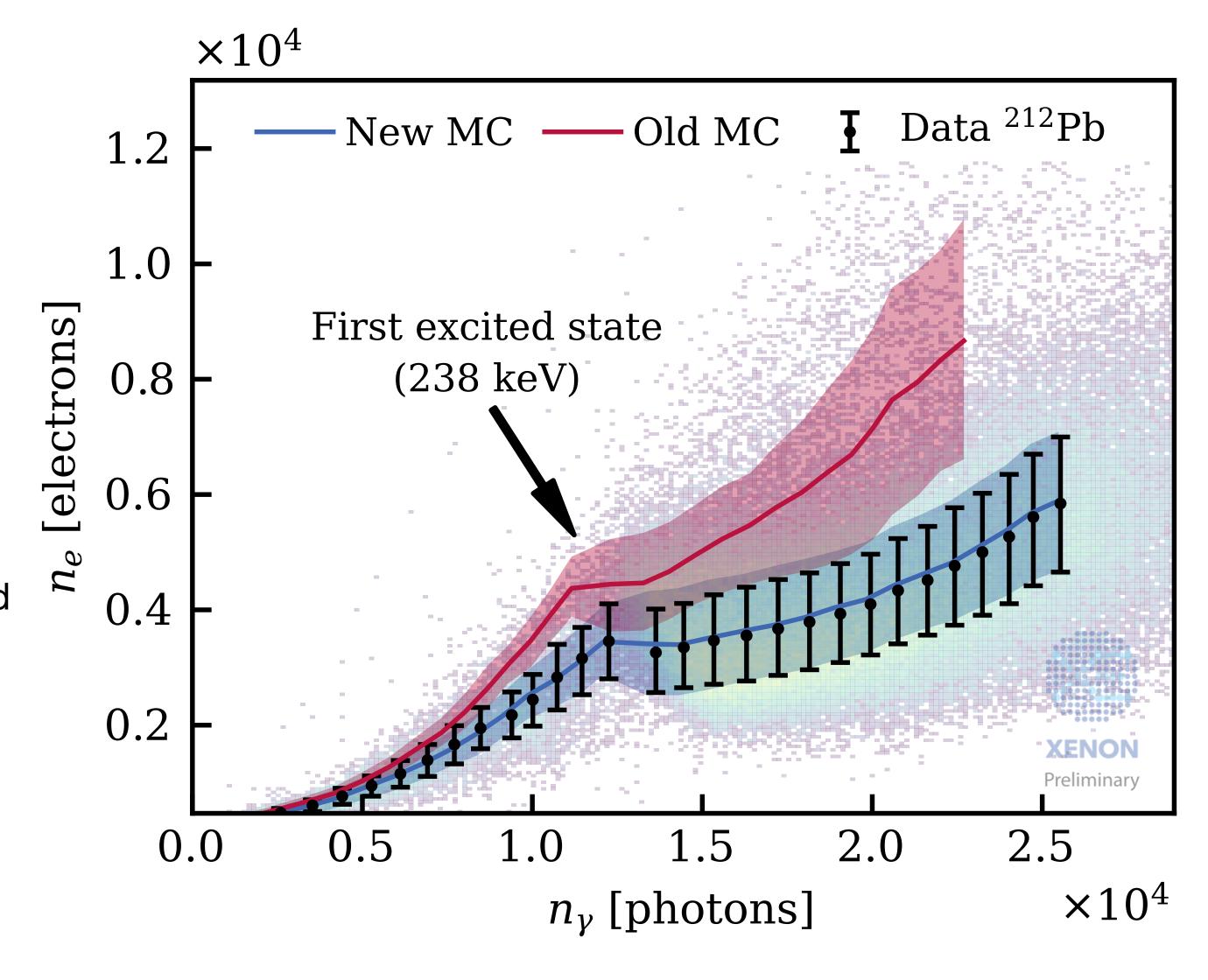
- Compare data and simulations...
- Already at a few hundreds of keV we are completely off!
- Why?
 - No reliable data at energies >100keV
 - Not a lot of data at field ~23V/cm
 - Energy deposits in simulations are grouped based on position only!
 - Inconsistency on how NEST yields are extracted and how they are used in simulations.



How do we model the response at higher energies?

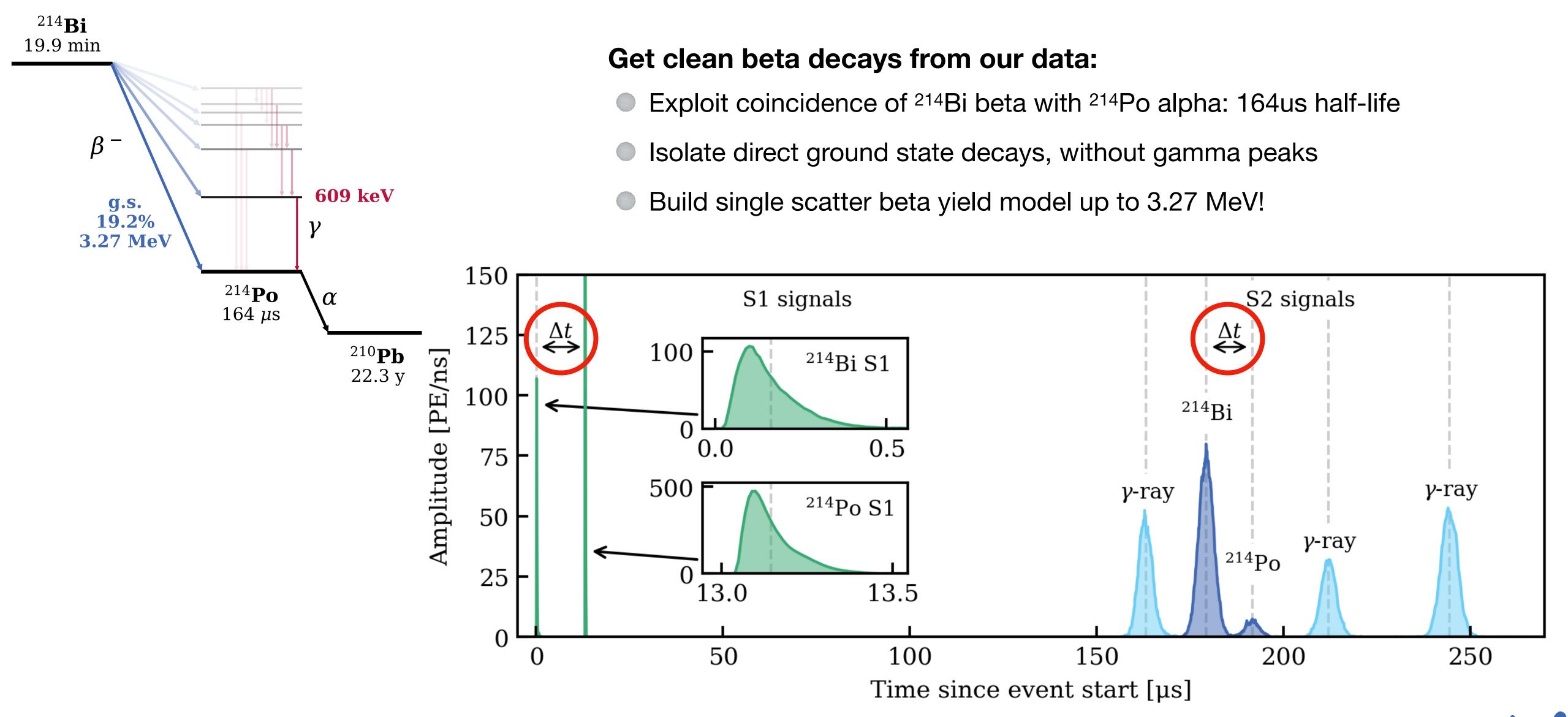
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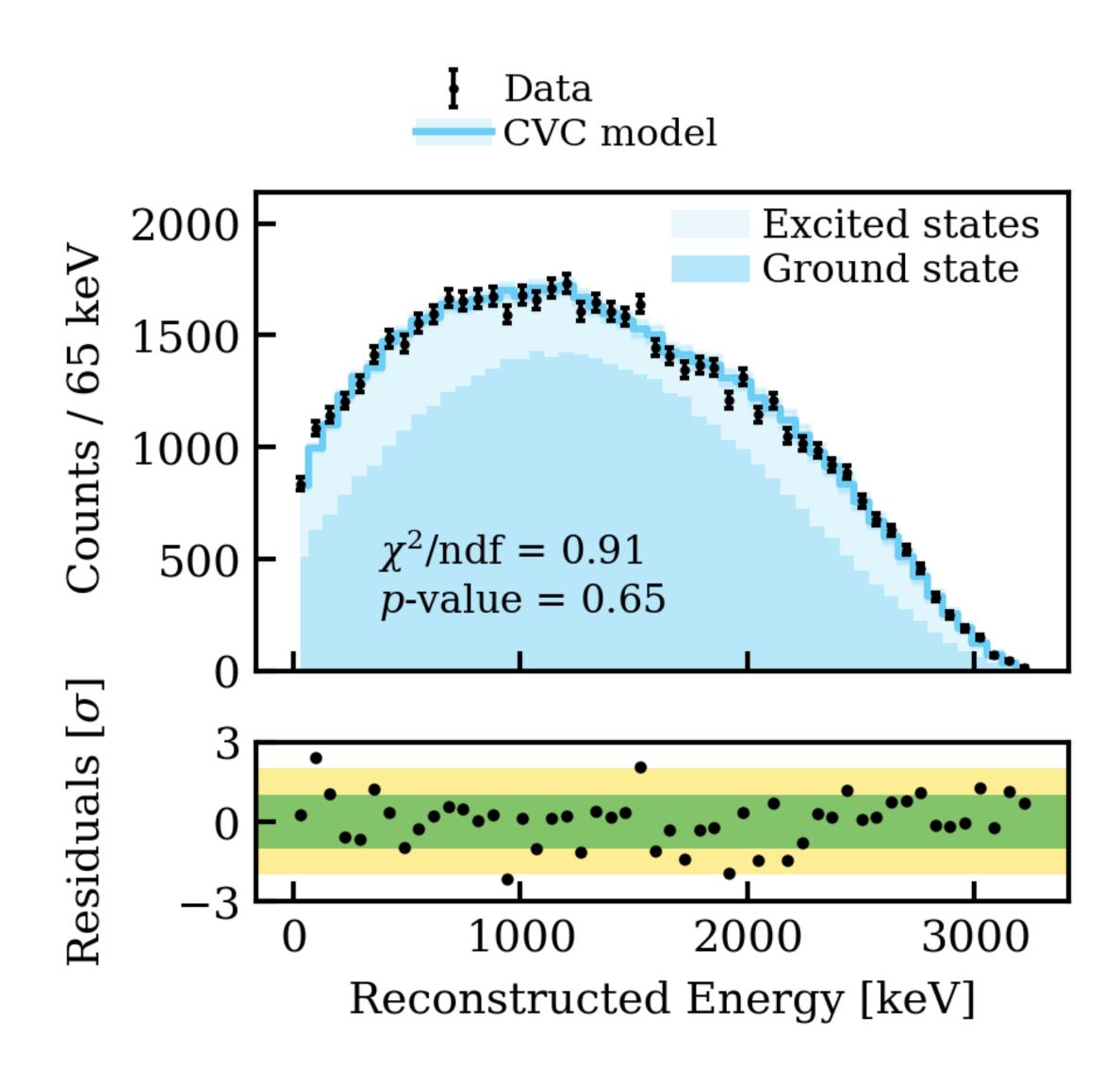




Extracting clean beta decay from background! XENON arxiv:2510.04846 (2025)

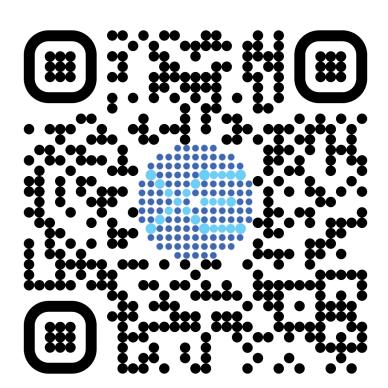


First measurement of ²¹⁴Bi g.s. beta-spectrum



THIS IS NOT AN AD.

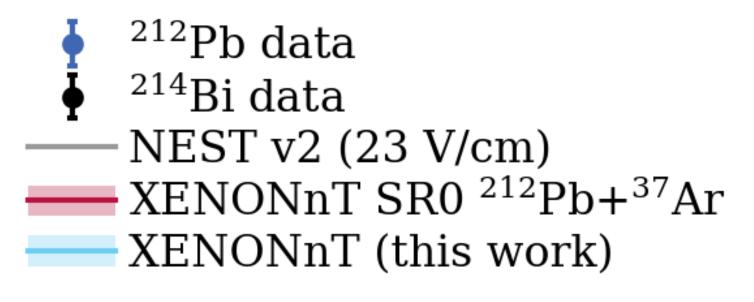
arXiv:2510.04846

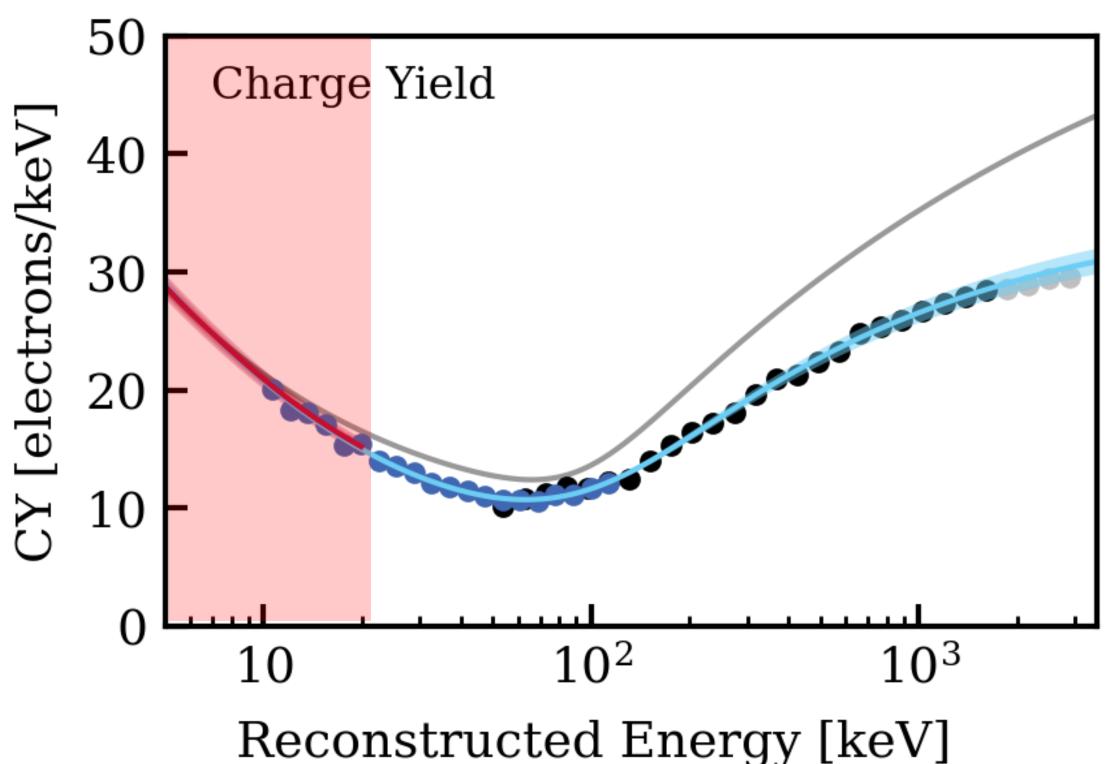


Result of the analysis:

- We can use the extracted sample to perform a measurement of the ²¹⁴Bi ground state decay spectrum.
- It is a first forbidden non-unique beta decay, meaning that nuclear structure impact the spectral shape.
- Compared spectrum with different theoretical models in recent XENONnT paper.

Extend yield model from 20keV to 3.2 MeV

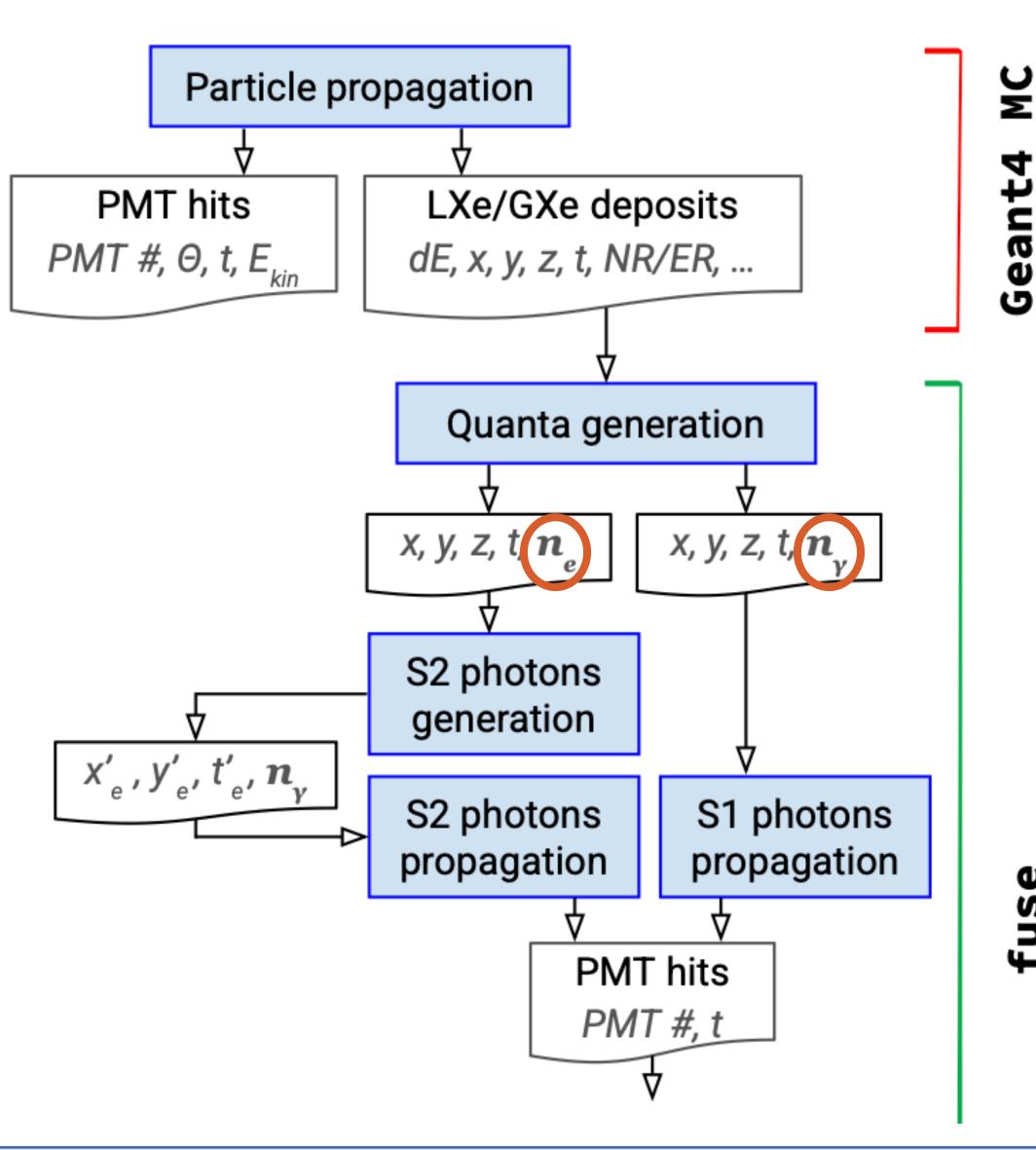




Data driven high energy yields:

- We obtain the charge (and light) yield from sample of clean single beta decays of ²¹⁴Bi to ground state, with Q-value of 3.27MeV.
- We estimate and correct for leakage of events containing gammas from excited states or from Brehmstralung emission via detailed simulations.
- Shows clear differences with respect to default NEST values, also because of the lack of data at the XENONnT operating drift field of 23V/cm.

XENONnT Simulations Framework: Geant4 + fuse



Simulations pipeline:

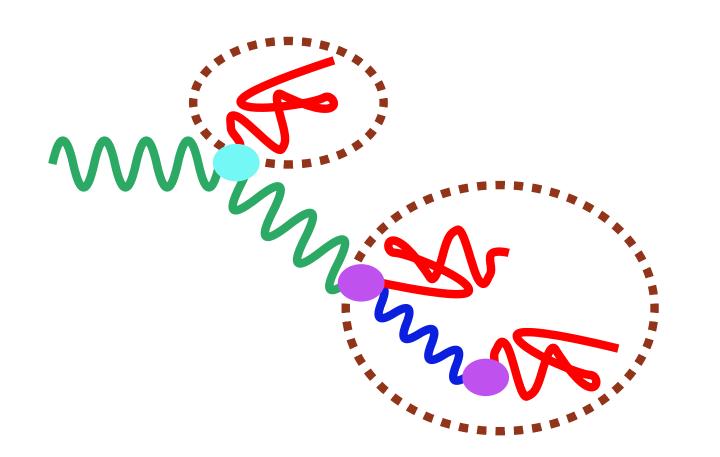
- Start from Geant4 simulation to get energy depositions, very verbose and detailed output
- Group energy deposit in clusters. We need to cluster them into meaningful interaction energies before assigning electrons and photons to each cluster.
- Clustering is now only position and time based algorithm (this works well at low energies, fails when multiple scatters!)
- Then, using NEST framework, assign quanta to each cluster based on interaction energy, interaction type and drift field.
- Propagate photons, drift electrons, make PMT waveforms... all the way to events!

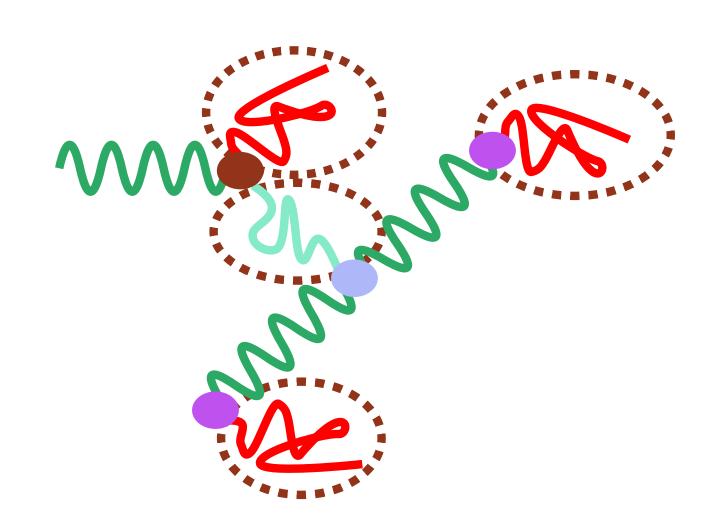


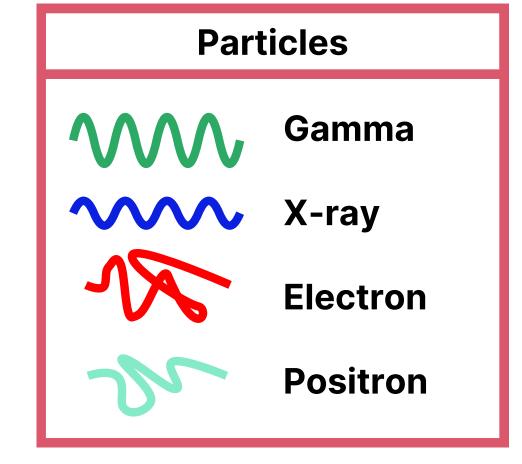
Lineage Clustering Algorithm

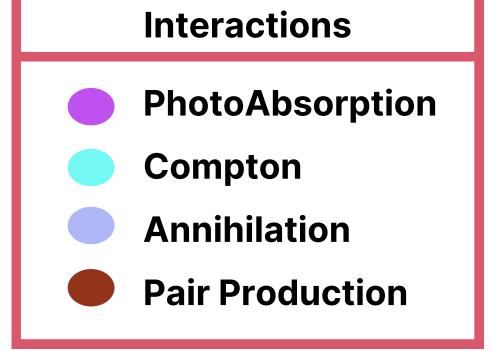
First change we need to make:

- We follow the "lineage" of each particle, splitting different clusters based on the interaction's physics.
- Rules:
 - beta-decay electrons are always one cluster (except Bremsstrahlung)
 - gamma-rays make a new cluster at each interaction step
 - photo-absorption is always one cluster, including extra gamma-rays and electrons emitted.
- Yields: we use only one yield model!











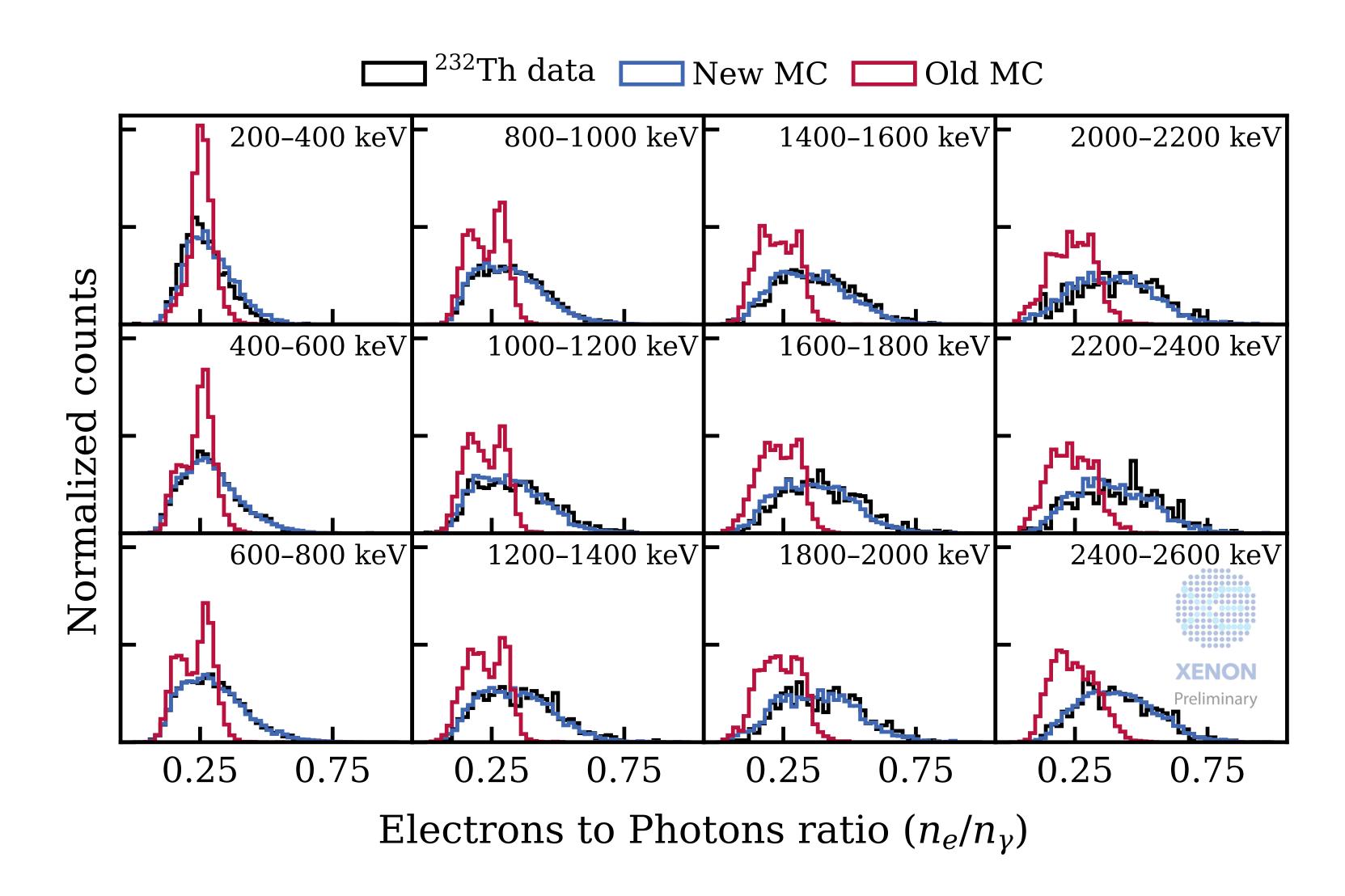




Matching Detector Response

High energy gamma-rays:

- We use ²³²Th data to validate the simulations at energies upto 2.6MeV. Population of gamma-rays with many compton scatters and interaction sites.
- We can now reproduce the response of the detector very well in the entire energy range with one model!
- We also adjusted the NEST recombination fluctuation model to fix the spread of the distribution.

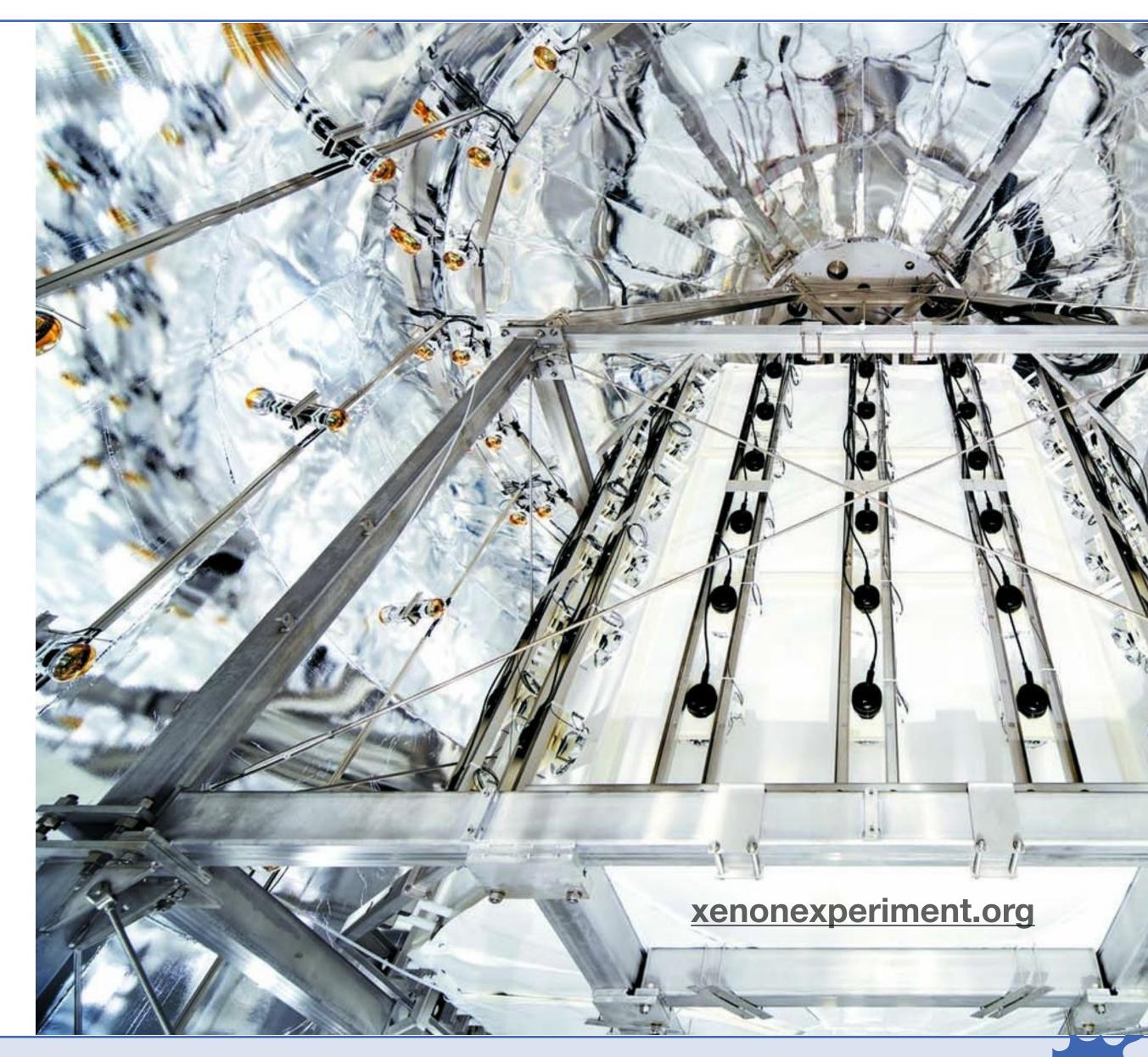




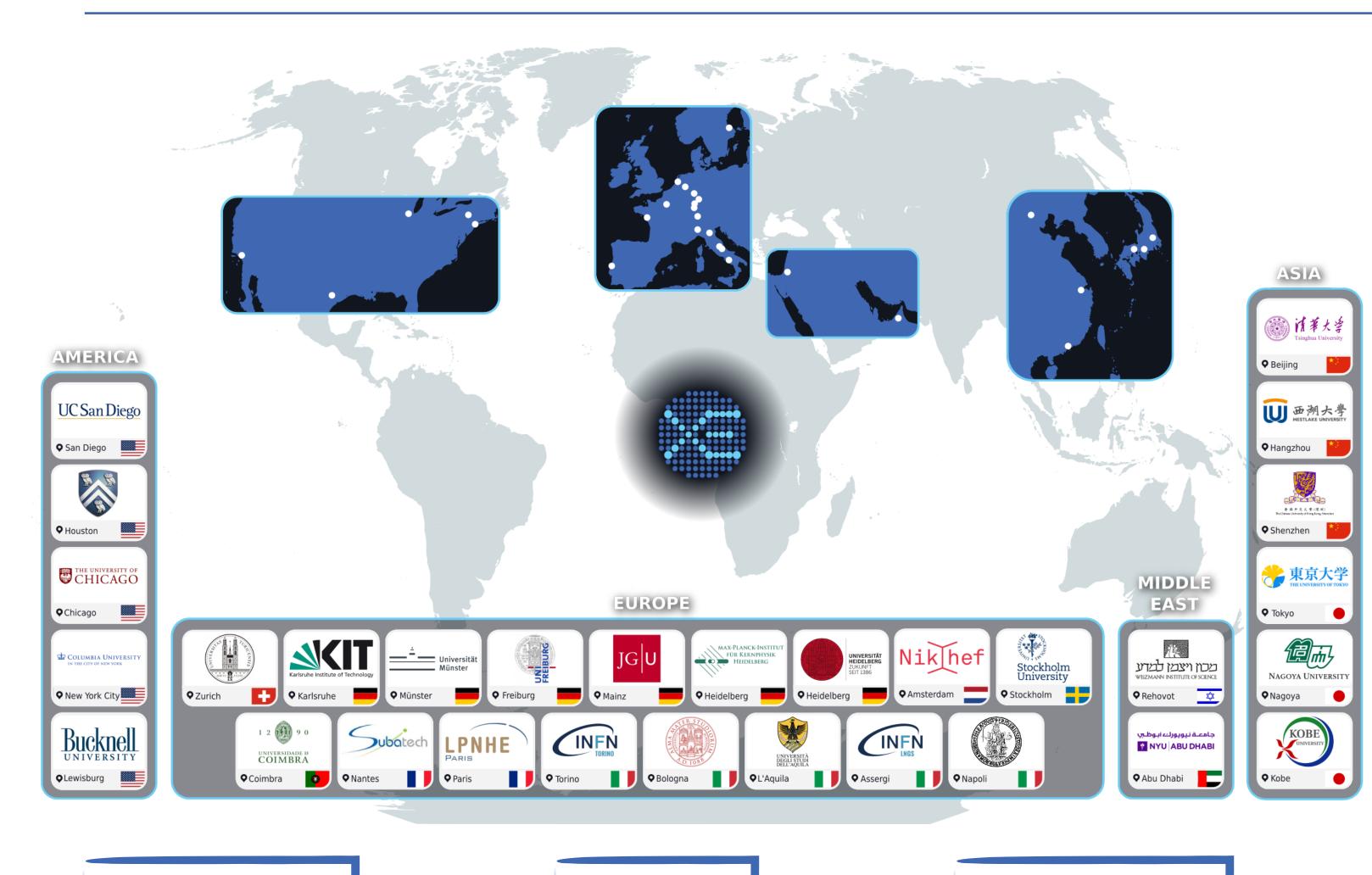
Thanks for your attention!

Summary:

- We extended the validity of the yield model of XENONnT from ~10keV to ~3MeV, using a clean sample of beta decays from ²¹⁴Bi.
- We introduced a physics based clustering method to properly account for multiple scatters at high energies.
- Combining the two, we are able to match data and simulations, including complicated topologies with mixed betas and gammas in the entire energy range of 0vbb searches.



The XENONnT Collaboration



Laboratori Nationali Gel Gran Sassonali

Running since 2020

8.6 tonnes LXe

Lowest Radon Background

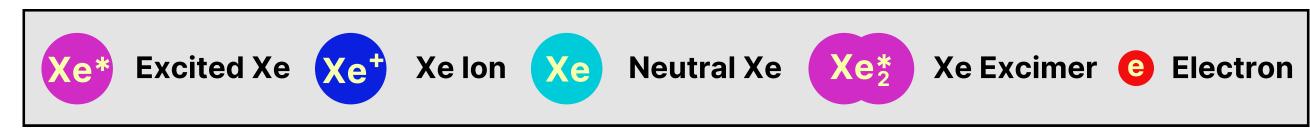
Process of scintillation

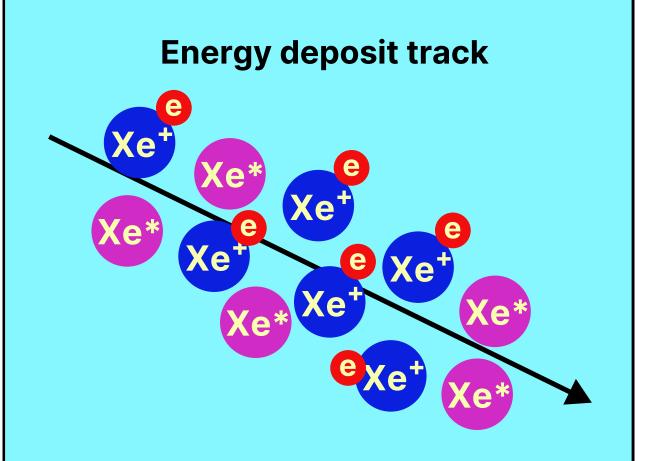
Quanta generation:

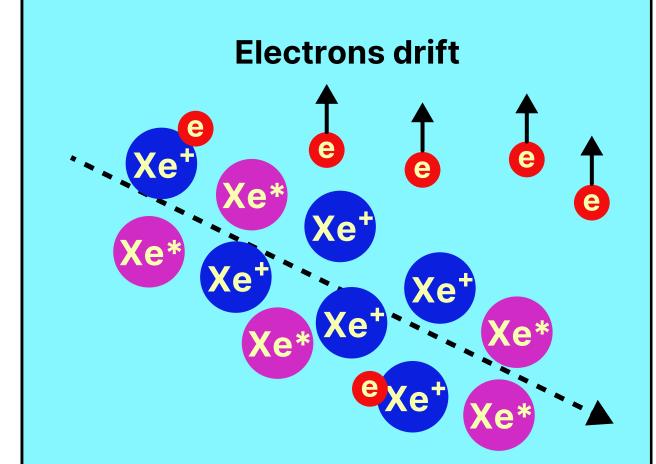
- Total number of excitons and ions created is~constant (depends slightly on density)
- Fraction of ions to excitons depends on the density and on the energy of the interaction.
- Fraction of electrons recombining depends strognly on drift field and interaction energy.
- Because with recombination we get one photon from one electron, still total number of quanta is conserved, and we can reconstruct energy as:

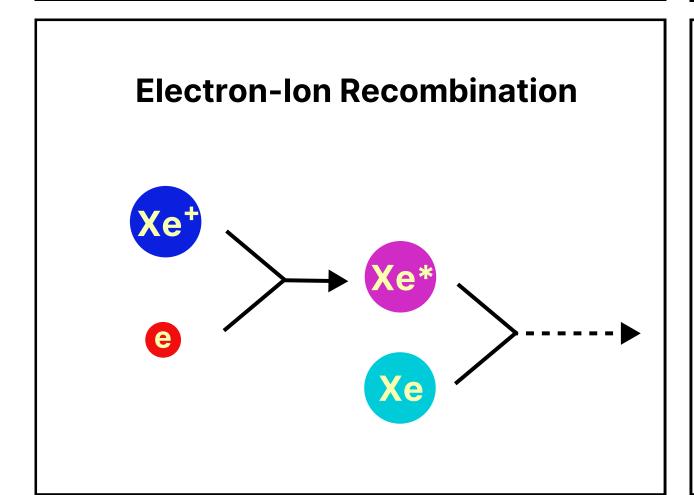
$$E = W(n_{photons} + n_{electrons})$$

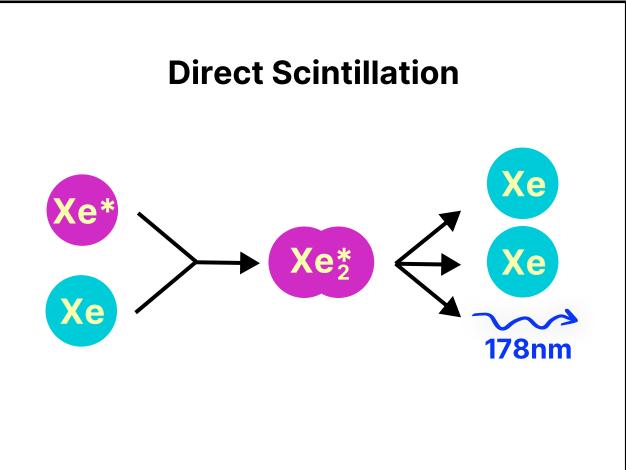
$$W \sim 13.7 \text{ eV/quanta}$$



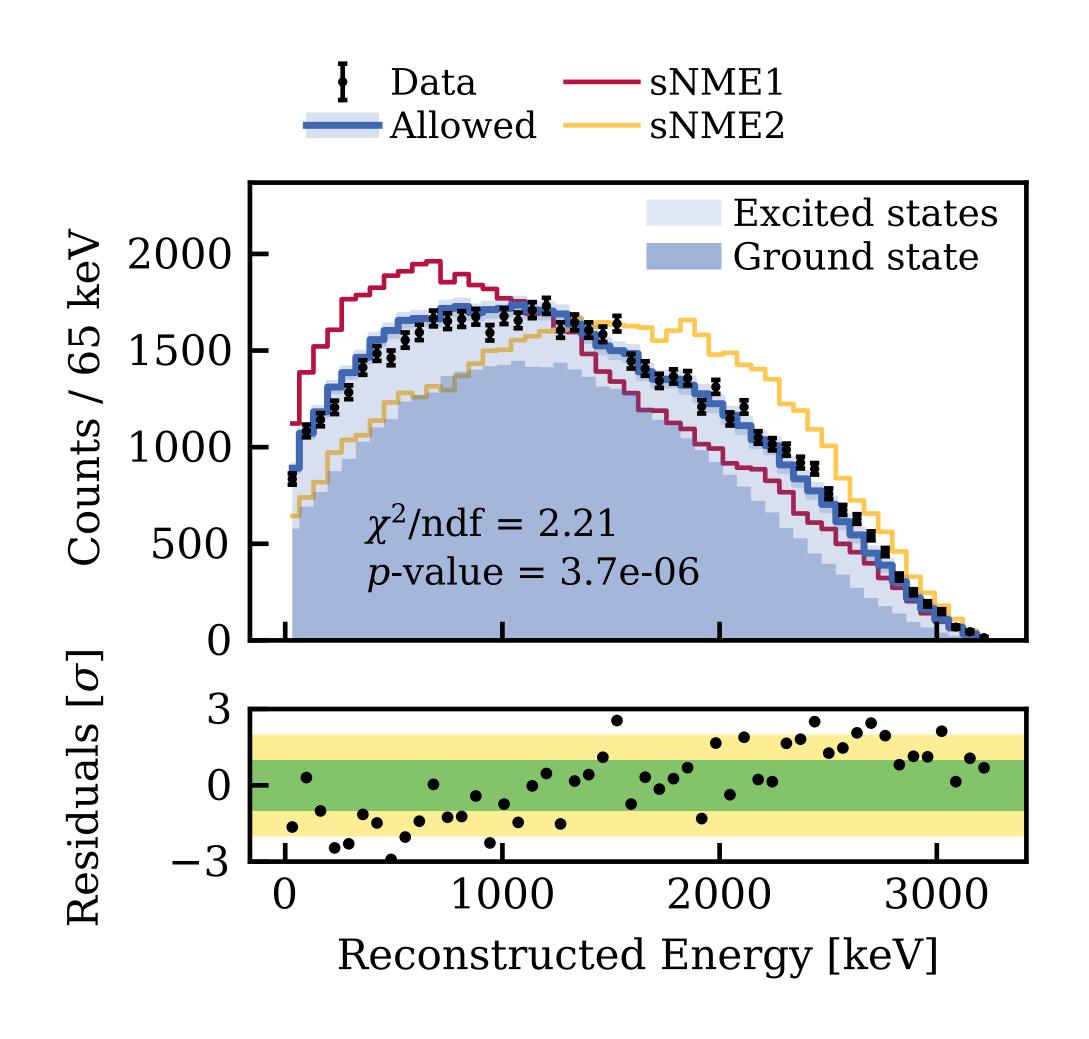


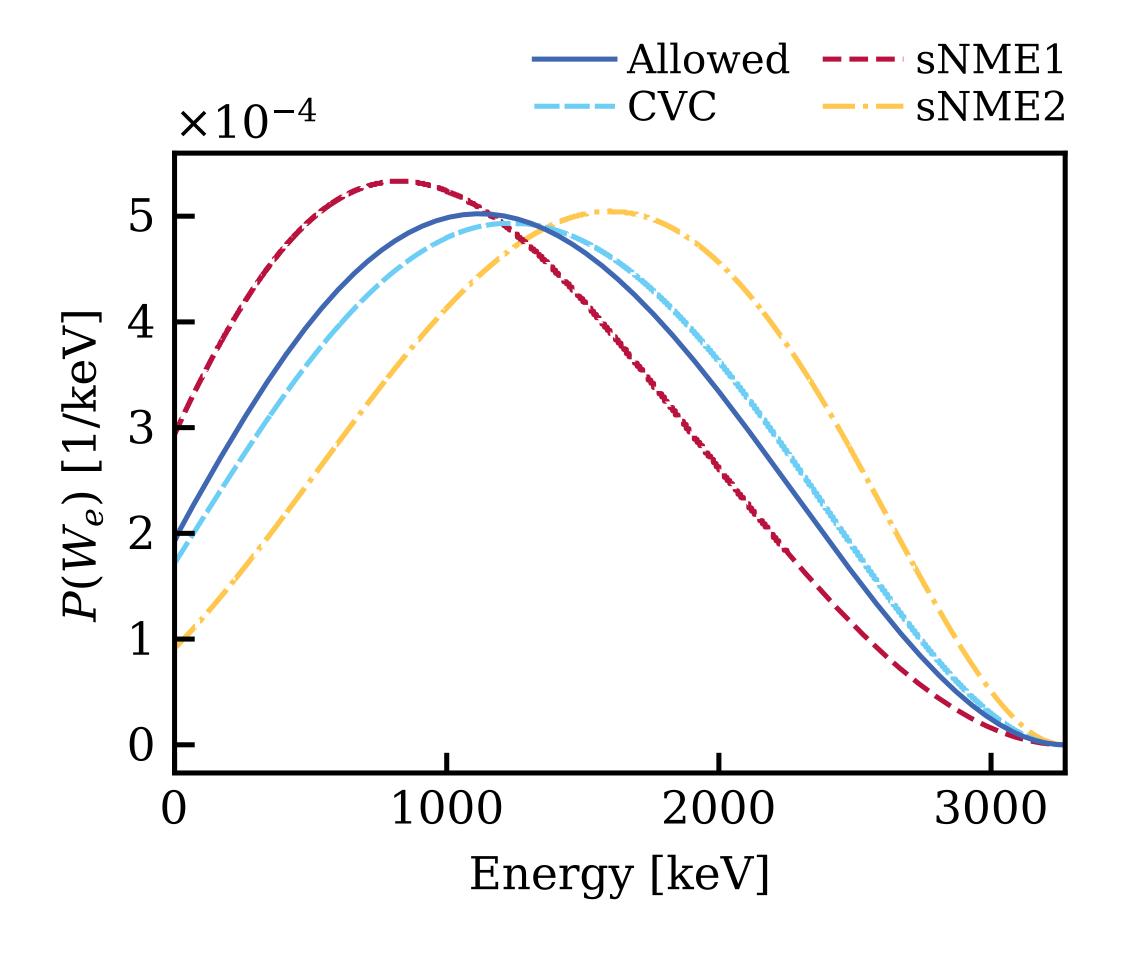






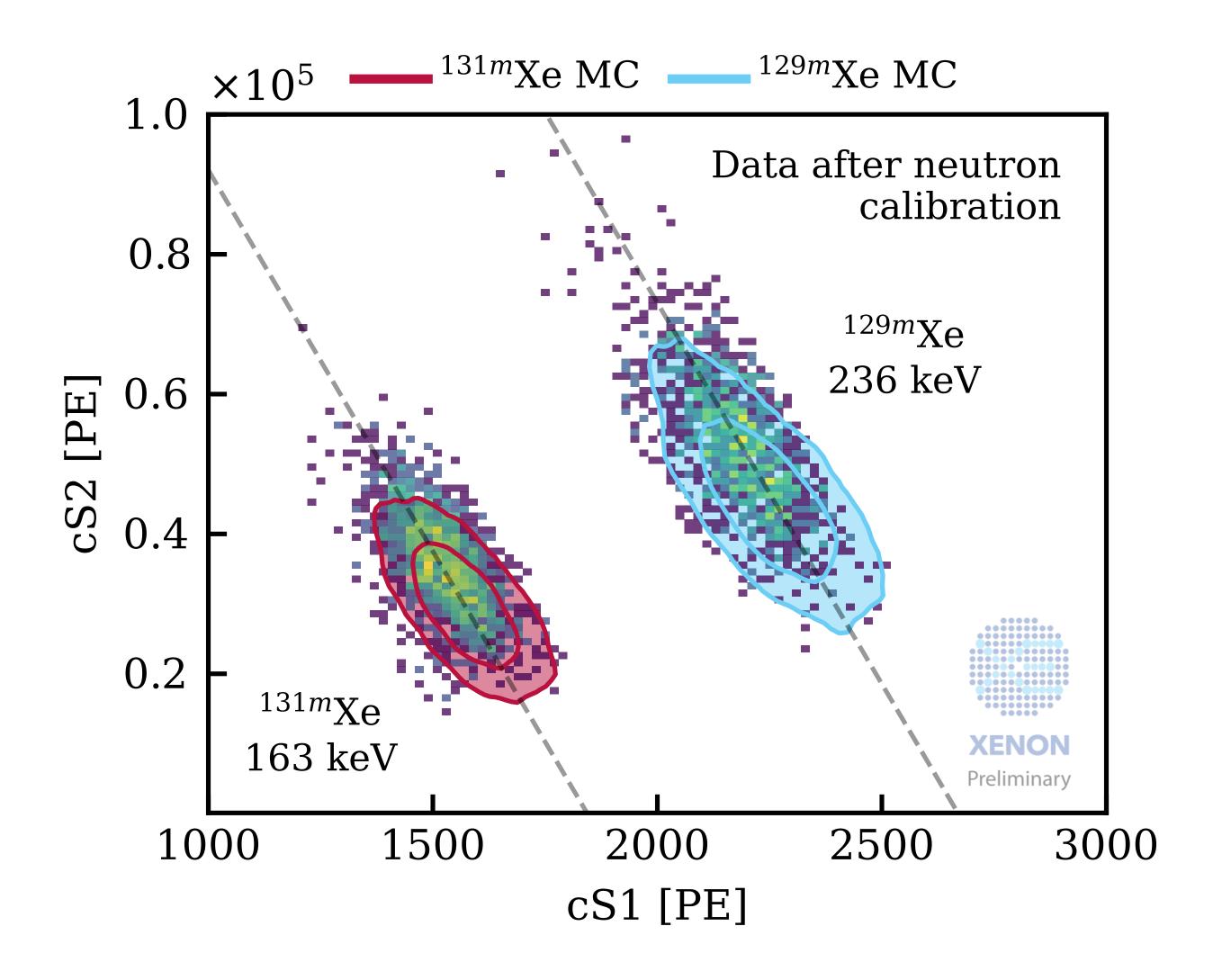
Other models



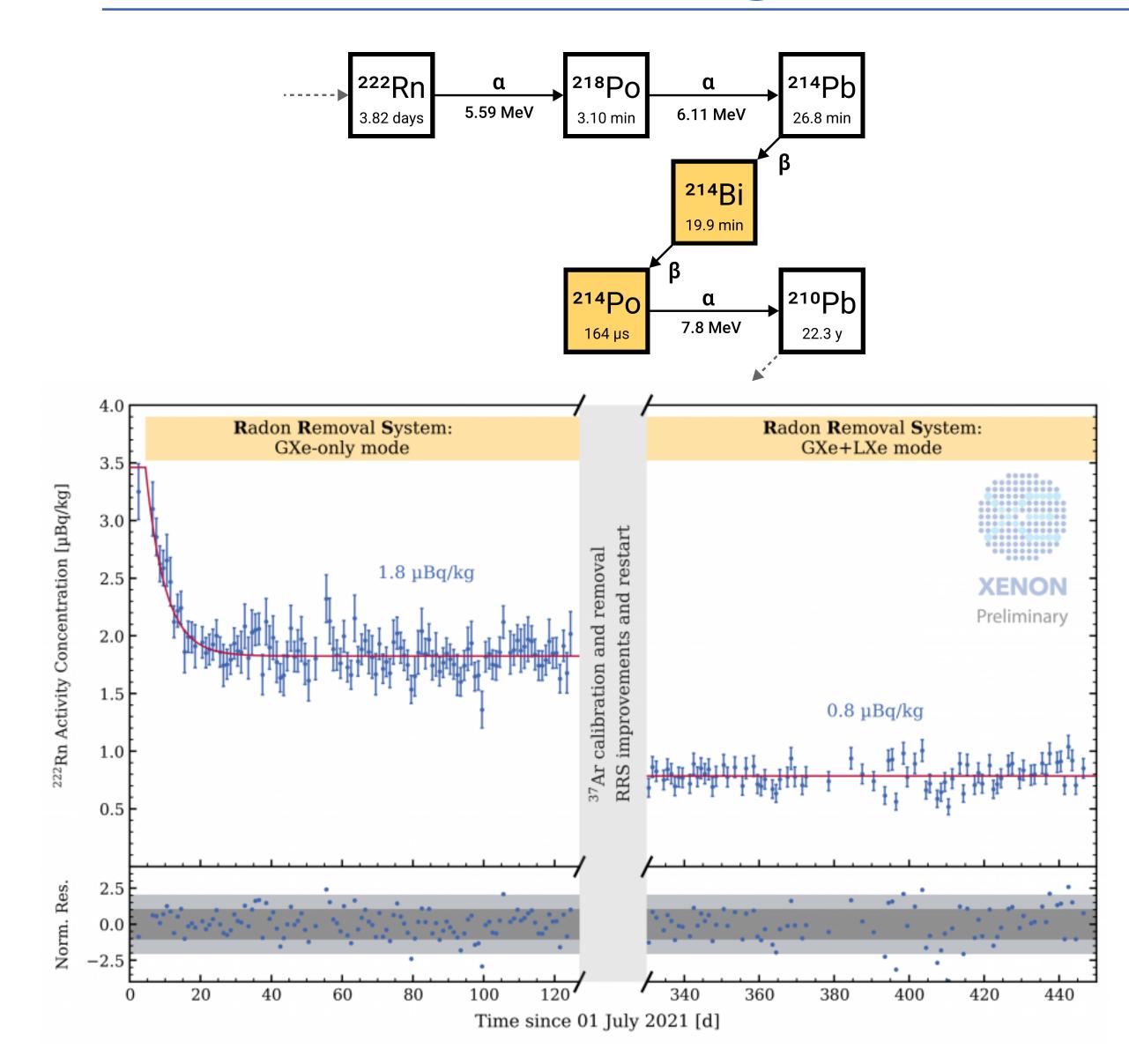


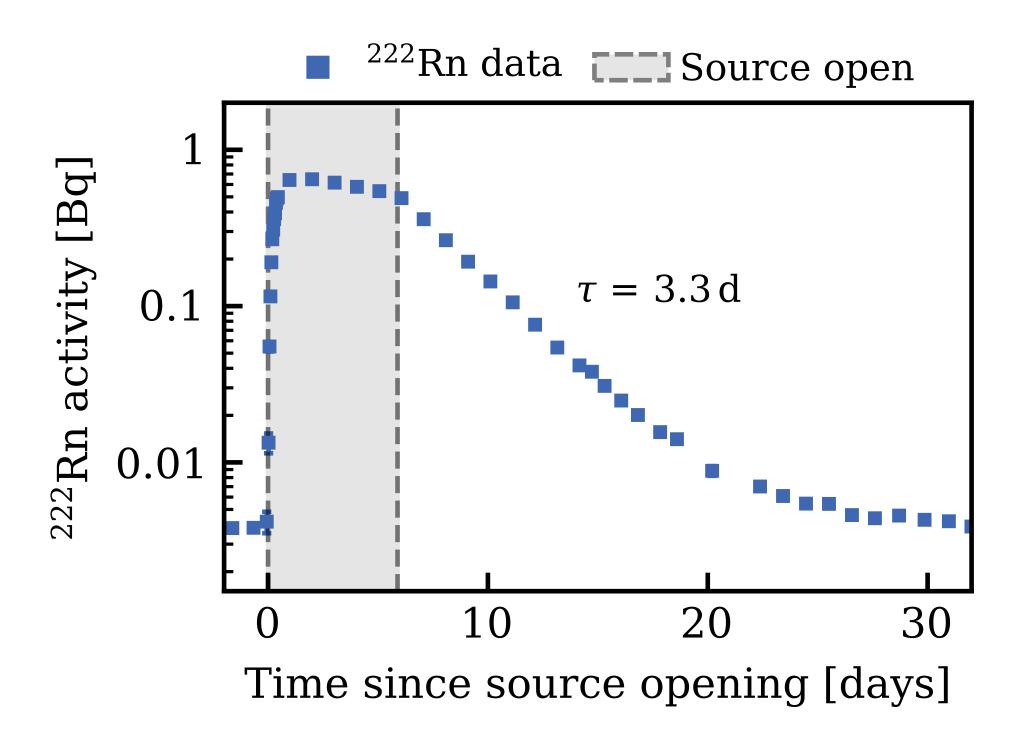


129Xe and 131Xe - IC

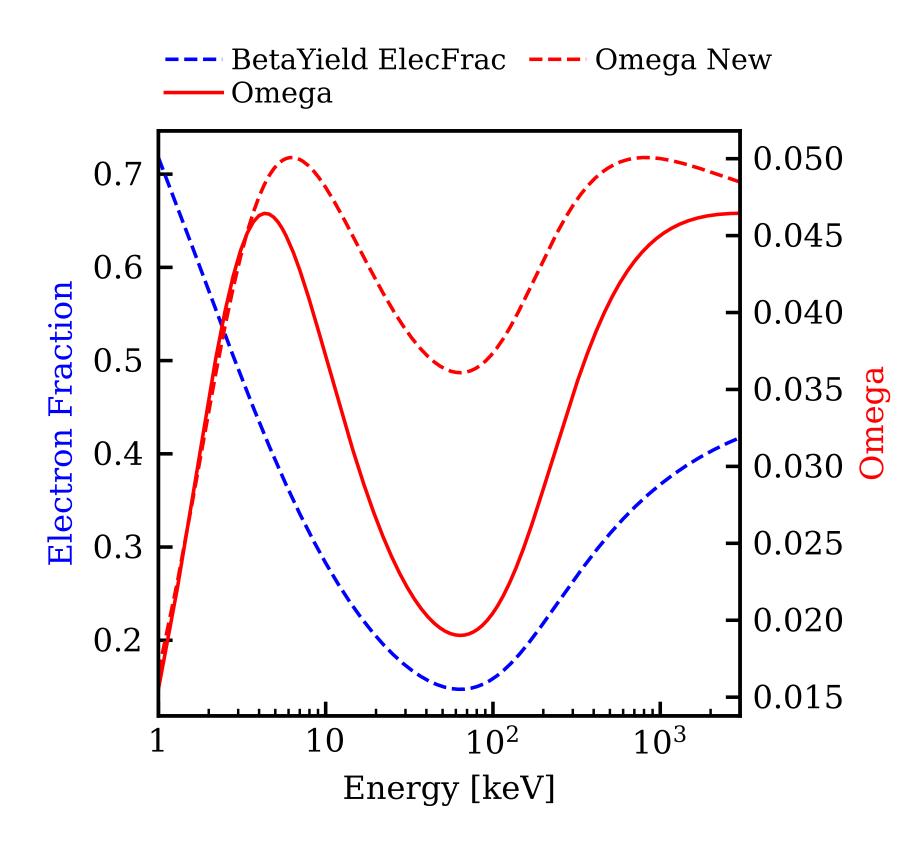


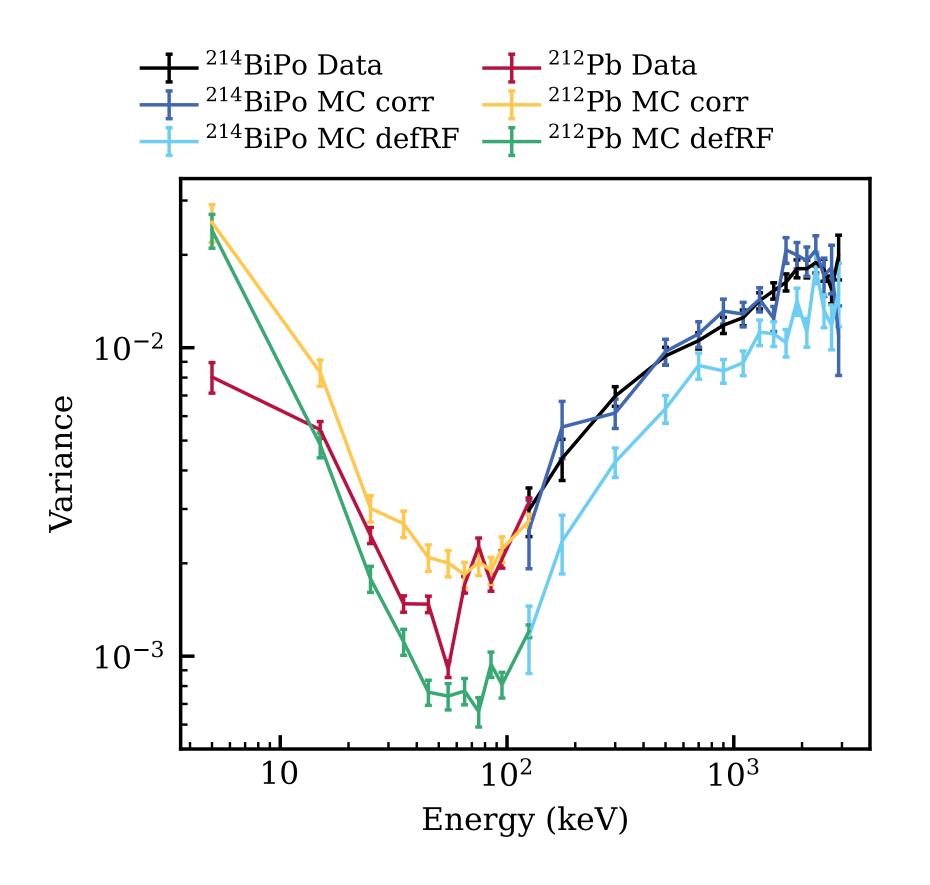
Radon chain backgrounds: the source of our signal





Recombination Fluctuation





NEST Yields: full beta model

