



東南大學
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Classical Gravitational Wave Observables from Quantum Field Theory

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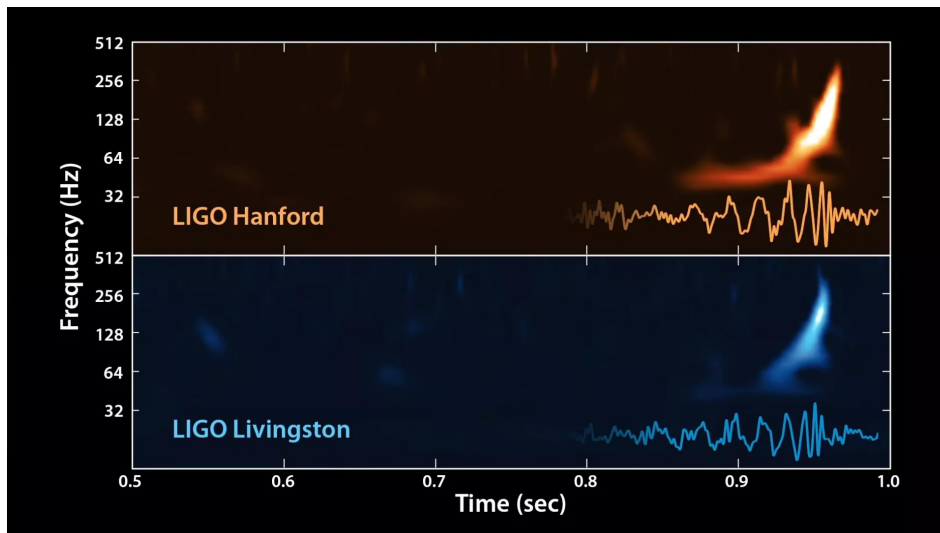
JHEP 08 (2023) 109 PRL 132 (2024) 221401 PRL 130 (2023) 101401 PRL 128 (2022) 161104
PLB 831 (2022) 137203 PRL 125 (2020) 261103 PRD 102 (2020) 124025 arXiv:2506.20665

第五届量子场论及其应用研讨会

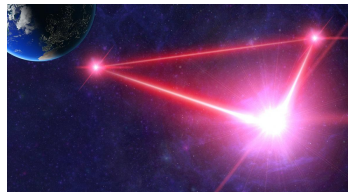
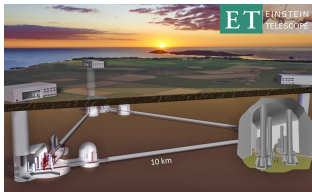


Gravitational wave science era has begun!

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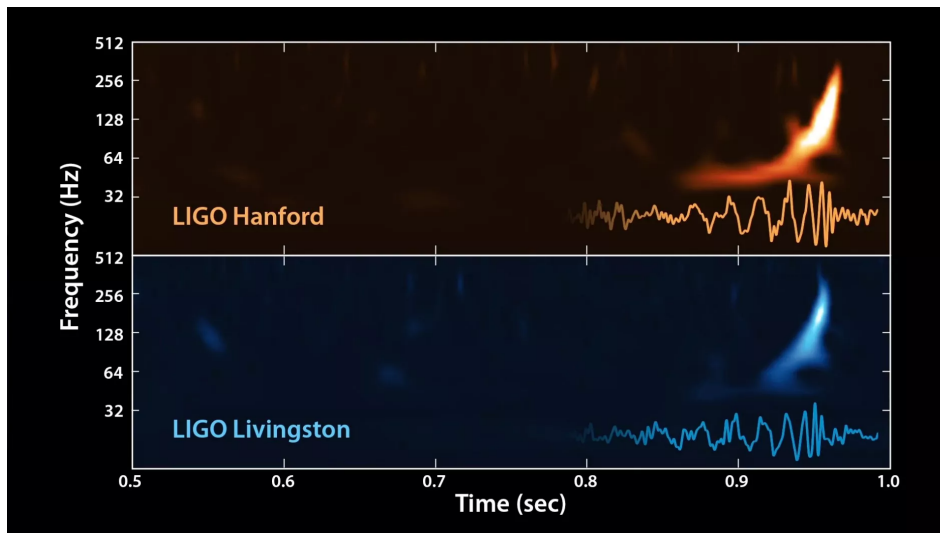


Opened a new window on the Universe!

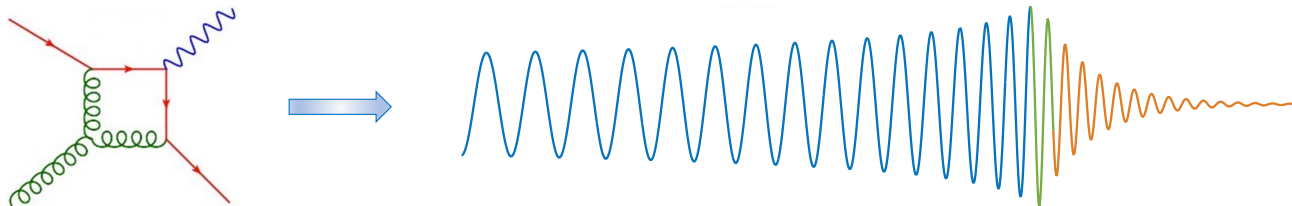


Gravitational wave science era has begun!

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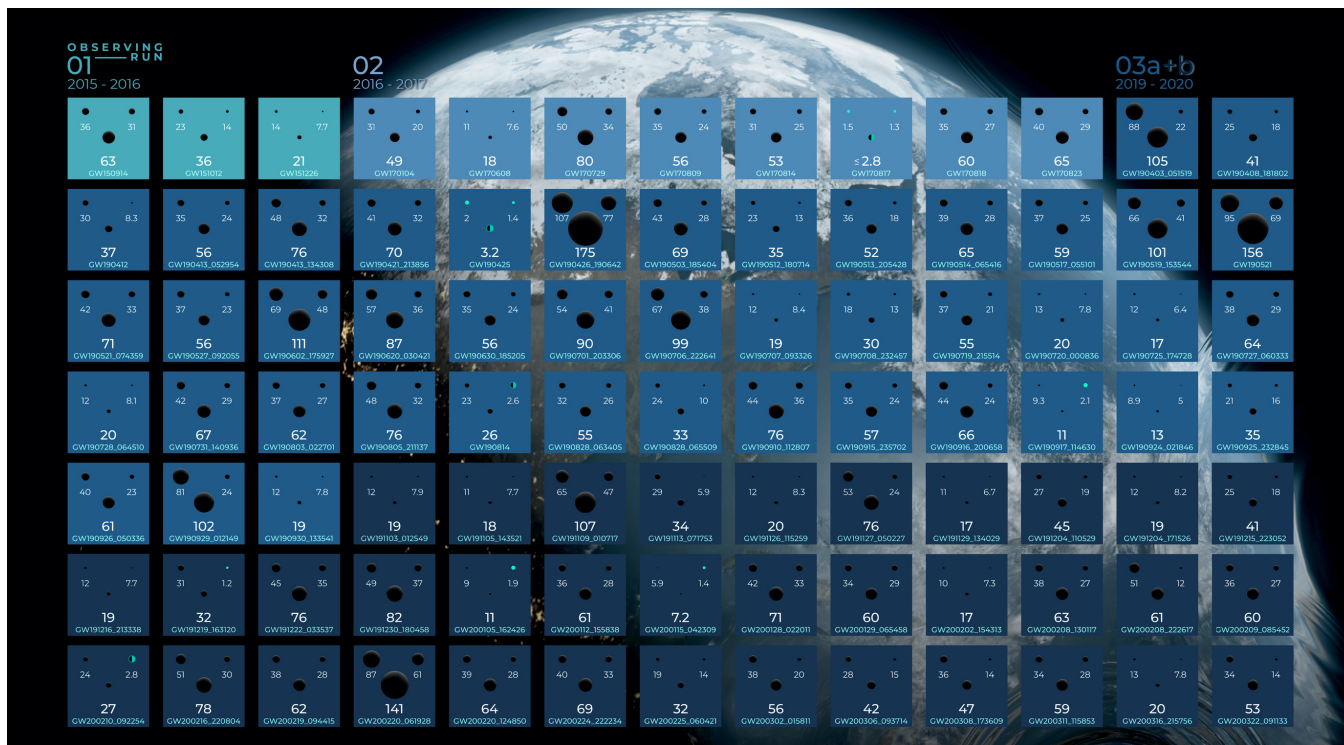


Modern Quantum Field Theory techniques are playing a crucial role in precision GW physics!



Gravitational waves from binary coalescences

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Credit: Carl Knox (OzGrav, Swinburne)

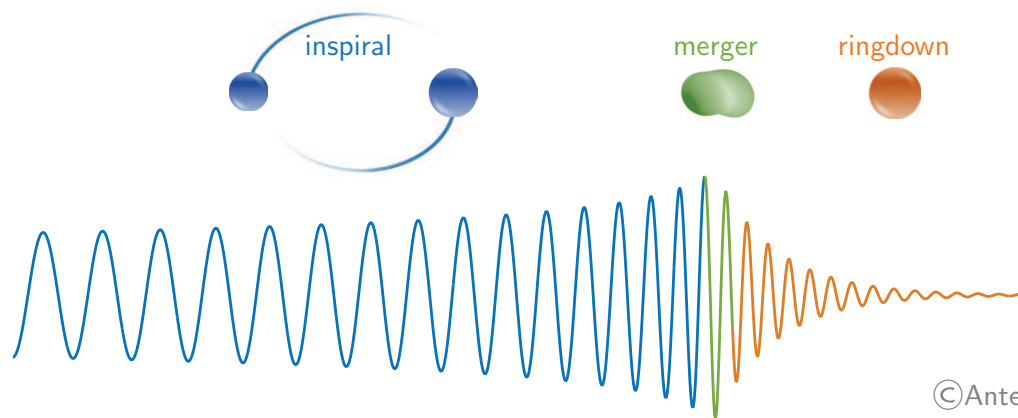


GWTC-3: 90 GW events—the majority are binary black holes (BH), but also several binary neutron stars (NS) and mixed NS-BHs.

gwosc.org

Gravitational waves from binary coalescences

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Merger: Numerical Relativity

Ringdown: black hole perturbation theory

Inspiral: the interaction between two bodies is weak

$$v^2 \sim \frac{GM}{r} \ll 1$$

- Numerical Relativity: accurately, but computationally expensive
- Analytic methods: corrections in v or G are perturbatively calculable

Post-Newtonian/post-Minkowskian expansion

- ▶ QFT methodology (EFT & Feynman loop techniques) shown great power!

Effective Field Theory

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- Gravitational binary system

$$S_{\text{WL}} = \sum_{i=1,2} \left[-\frac{m_i}{2} \int dt g_{\mu\nu} \dot{x}_i^\mu \dot{x}_i^\nu + \dots \right] \quad S_{\text{GR}} = \frac{-1}{16\pi G} \int d^4x \sqrt{-g} R + \dots$$

- Effective action for gravitational binary systems

Goldberger-Rothstein hep-th/0409156

$$g_{\mu\nu} = \eta_{\mu\nu} + h_{\mu\nu}$$

$$e^{iS_{\text{eff}}[x_a(\tau)]} = \int \mathcal{D}h_{\mu\nu} e^{iS_{\text{WL}} + iS_{\text{GR}}}$$

- Post-Minkowskian expand in powers of G

$$L_{\text{eff}} = L_0 + GL_1 + G^2L_2 + \dots \quad L_0 = -\sum_i \frac{m_i}{2} \eta_{\mu\nu} \dot{x}_i^\mu \dot{x}_i^\nu$$

- The equations of motion for trajectories:

Kälin-Porto 2006.01184

$$m_i \ddot{x}_i^\mu = -\eta^{\mu\nu} \sum_{n=1}^{\infty} G^n \left(\frac{\partial L_n}{\partial x_i^\nu} - \frac{d}{d\tau_i} \frac{\partial L_n}{\partial \dot{x}_i^\nu} \right) \quad x_i^\mu = b_i^\mu + u_i^\mu \tau_i + \delta x_i^\mu(\tau_i) + \dots$$

- Physical observables:

$$\Delta p_i^\mu = p_i^\mu(+\infty) - p_i^\mu(-\infty) = -\eta^{\mu\nu} \sum_{n=1}^{\infty} G^n \int_{-\infty}^{\infty} d\tau_i \left(\frac{\partial L_n}{\partial x_i^\nu} \right)$$

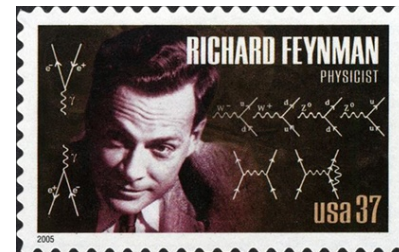
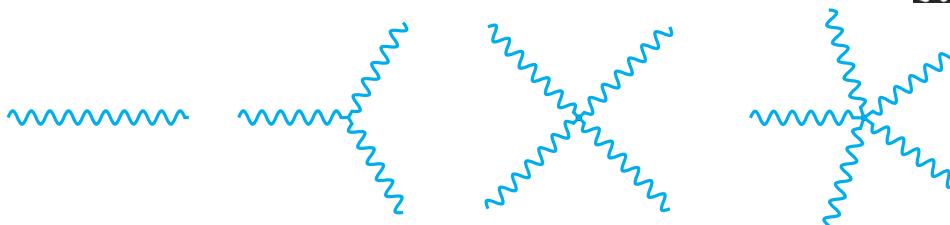
Effective Field Theory

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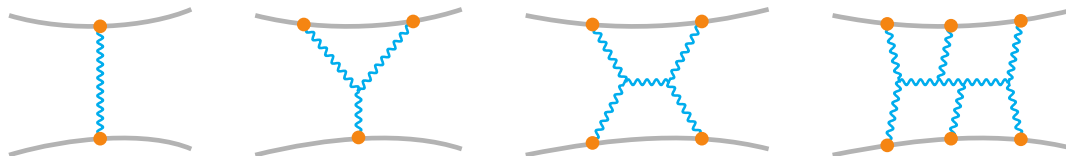
- Worldlines as classical sources in path integral:



- Hilbert-Einstein: $\mathcal{L}_{\text{HE}} = \mathcal{L}_{hh} + \mathcal{L}_{hhh} + \mathcal{L}_{hhhh} + \dots$



- Classical physics:** we use the saddle-point approximation in path integrals.



- Enjoy the advantages of **quantum field theory methods** and **classical physics**
powerful and systematic & **purely classical at all steps (simplicity)**

Effective Field Theory

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The in-in effective action is obtained by performing a closed-time-path integral

$$e^{i\mathcal{S}_{\text{eff}}[x_{a,1}, x_{a,2}]} = \int \mathcal{D}h_1 \mathcal{D}h_2 e^{i(S_{\text{GR}}[h_1] - S_{\text{GR}}[h_2] + S_{\text{WL}}[h_1, x_{a,1}] - S_{\text{WL}}[h_2, x_{a,2}])}$$

It is convenient to use the Keldysh basis

Galley PRL 110 (2013) 174301

$$\begin{aligned} h_{\mu\nu}^- &= \frac{1}{2}(h_{1\mu\nu} + h_{2\mu\nu}) & x_{a,+}^\alpha &= \frac{1}{2}(x_{a,1}^\alpha + x_{a,2}^\alpha) \\ h_{\mu\nu}^+ &= h_{1\mu\nu} - h_{2\mu\nu} & x_{a,-}^\alpha &= x_{a,1}^\alpha - x_{a,2}^\alpha \end{aligned}$$

for which the matrix of (classical) propagators for gravitons becomes

$$i \begin{pmatrix} 0 & -\Delta_{\text{adv}}(x-y) \\ -\Delta_{\text{ret}}(x-y) & 0 \end{pmatrix}$$

The worldline equations of motion:

Kälin-Neef-Porto JHEP 01 (2023) 140

$$m_i \ddot{x}_i^\mu(\tau) = -\eta^{\mu\nu} \frac{\delta S_{\text{eff, int}}[x_{a,\pm}]}{\delta x_{i,-}^\nu(\tau)} \Big|_{\text{PL}} \quad \Delta p_i^\mu = -\eta^{\mu\nu} \int_{-\infty}^{\infty} d\tau \frac{\delta S_{\text{eff, int}}[x_{a,\pm}]}{\delta x_{i,-}^\nu(\tau)} \Big|_{\text{PL}}$$

Physical Limit (PL): $x_{a,-} \rightarrow 0$, $x_{a,+} \rightarrow x_a$.

Closed-time path integrals

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EQUILIBRIUM AND NONEQUILIBRIUM FORMALISMS MADE UNIFIED

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周光召



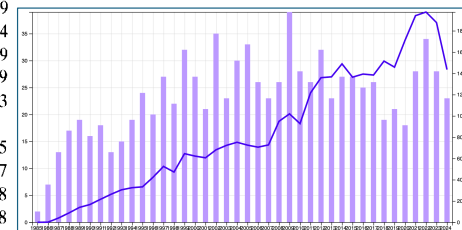
郝柏林



苏肇冰



于渌



Effective Field Theory

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- Observables at $\mathcal{O}(G^N)$

Kälin-ZL-Porto PRL2020

Dlapa-Kälin-ZL-Porto PRL2022 JHEP2024

$$\Delta p_i^\mu \sim \int d^D q \frac{e^{iq \cdot b} \delta(q \cdot u_1) \delta(q \cdot u_2)}{|q^2|^\sharp} \int \left(\prod_{i=1}^{N-1} d^D \ell_i \frac{\delta(\ell_j \cdot u_a)}{(\ell_j \cdot u_b - i0)^{\nu_i}} \right) \frac{\mathcal{N}^\mu(q, u_a)}{D_1 D_2 D_3 \cdots}$$

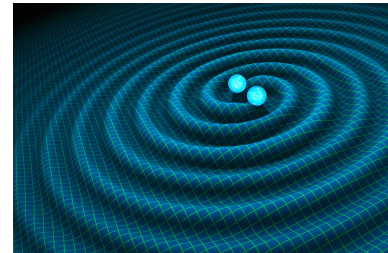
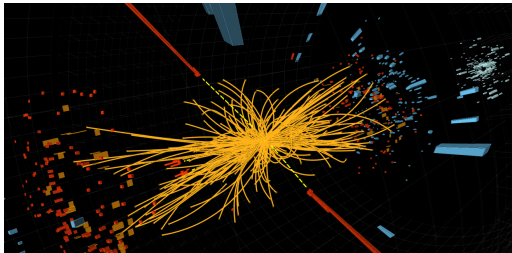
- ▶ Graviton propagators:

$$\frac{1}{D_i} \longrightarrow \frac{1}{\ell^2 + i0} \quad \text{or} \quad \frac{1}{(\ell^0 \pm i0)^2 - \vec{\ell}^2}$$

- ▶ **Cut**: always one delta function $\delta(\ell_j \cdot u_a)$ for each loop

- ▶ Kinematics: $q \cdot u_a = 0$, $u_a^2 = 1$, $u_1 \cdot u_2 = \gamma \implies$ **single scale** γ to all orders

- Multi-loop technology from QFT can be used to solve classical gravitational problems!

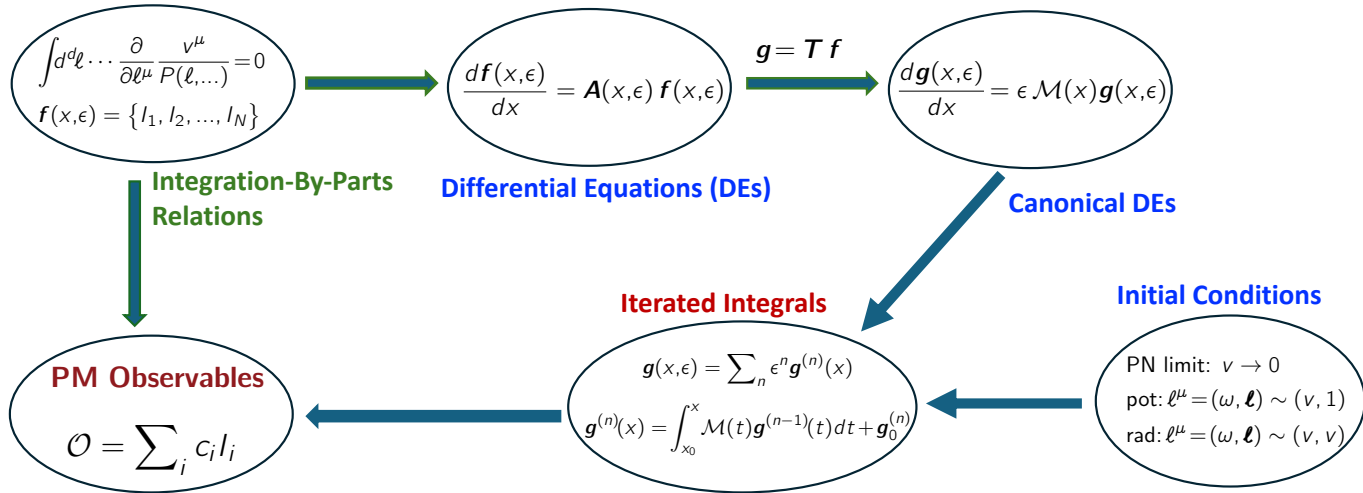


- Observables at $\mathcal{O}(G^N)$ Kälin-ZL-Porto PRL 2020 Dlapa-Kälin-ZL-Porto PRL 2022 Dlapa-Kälin-ZL-Neef-Porto PRL 2023

$$\Delta p_i^\mu \sim \int d^D q \frac{e^{iq \cdot b} \delta(q \cdot u_1) \delta(q \cdot u_2)}{|q^2|^\#} \int \left(\prod_{i=1}^{N-1} d^D \ell_i \frac{\delta(\ell_j \cdot u_a)}{(\ell_i \cdot u_b - i0)^{\nu_i}} \right) \frac{\mathcal{N}^\mu(q, u_a)}{D_1 D_2 D_3 \dots}$$

- Perturbative QFT toolbox:

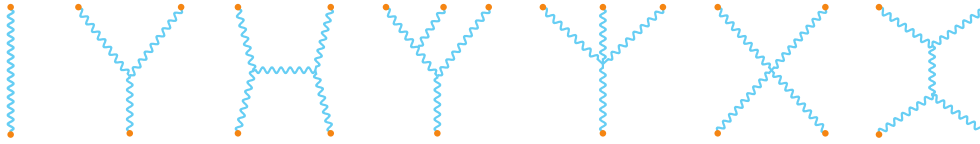
Dlapa-Kälin-ZL-Porto JHEP 2024



- Post-Minkowskian physics can be bootstrapped from Post-Newtonian data using DEs!

NNLO: 3PM

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- $\mathcal{O}(G^3)$: two-loop integrals

Kälin-ZL-Porto PRL 2020 PRD 2020

$$\int \frac{d^D \ell_1 d^D \ell_2 \delta(\ell_1 \cdot u_1) \delta(\ell_2 \cdot u_2)}{[\ell_1 \cdot u_2]^{a_1} [\pm \ell_2 \cdot u_1]^{a_2}} \frac{1}{[\ell_1^2]^{a_3} [\ell_2^2]^{a_4} [(\ell_1 + \ell_2 - q)^2]^{a_5} [(\ell_1 - q)^2]^{a_6} [(\ell_2 - q)^2]^{a_7}}$$

- The reduction and evaluation of integrals can be performed in standard techniques.

$$\partial_x \vec{f}(x, \epsilon) = \epsilon \left(\frac{A}{x} + \frac{B}{x+1} + \frac{C}{x-1} \right) \vec{f}(x, \epsilon)$$

- Conservative dynamics at $\mathcal{O}(G^3)$:

Kälin-ZL-Porto PRL 2020

$$\Delta p_1^\mu = \frac{G^3 b^\mu}{|b|^2} \left(\frac{8m_1^2 m_2^2 (4\gamma^4 - 12\gamma^2 - 3)}{(\gamma^2 - 1)} \log(\gamma - \sqrt{\gamma^2 - 1}) - \frac{2m_1 m_2 (m_1^2 + m_2^2) (16\gamma^6 - 32\gamma^4 + 16\gamma^2 - 1)}{(\gamma^2 - 1)^{5/2}} \right. \\ \left. - \frac{4m_1^2 m_2^2 \gamma (20\gamma^6 - 90\gamma^4 + 120\gamma^2 - 53)}{3(\gamma^2 - 1)^{5/2}} \right) + \frac{3\pi (2\gamma^2 - 1) (5\gamma^2 - 1)}{2(\gamma^2 - 1)^2} \frac{G^3 m_1 m_2 (m_1 + m_2)}{|b|^2 |b|^{3/2}} \left((m_1 + \gamma m_2) u_2^\mu - (m_2 + \gamma m_1) u_1^\mu \right)$$

- We provided the first confirmation for the result from a scattering amplitude calculation.

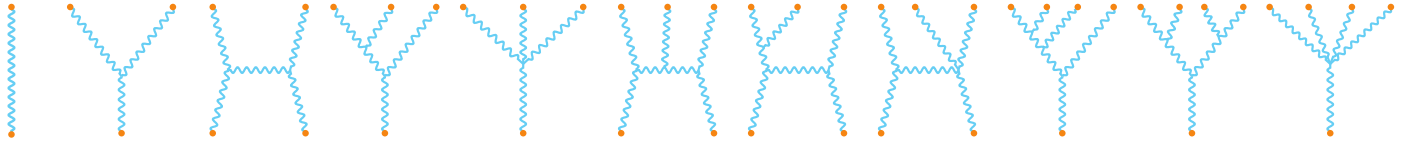
Bern-Cheung-Roiban-Shen-Solon-Zeng 2019

- We computed quadrupolar and octupole tidal corrections at $\mathcal{O}(G^3)$.

Kälin-ZL-Porto PRD 2020

NNNLO: 4PM

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$\mathcal{O}(G^4)$: three-loop integrals

Dlapa-Kälin-ZL-Porto PRL 2022 PLB 2022 Dlapa-Kälin-ZL-Neef-Porto PRL 2023

$$\int d^D \ell_1 d^D \ell_2 d^D \ell_3 \frac{\delta(\ell_1 \cdot u_1) \delta(\ell_2 \cdot u_2) \delta(\ell_3 \cdot u_2)}{[\ell_1 \cdot u_2]^{\alpha_1} [\ell_2 \cdot u_1]^{\alpha_2} [\ell_3 \cdot u_1]^{\alpha_3}} \frac{D_8^{-\nu_8} D_9^{-\nu_9}}{D_1^{\nu_1} D_2^{\nu_2} \dots D_7^{\nu_7}} \left\{ \ell_1^2, \ell_2^2, (\ell_1 - q)^2, (\ell_2 - q)^2, (\ell_3 - q)^2, \right. \\ \left. \ell_3^2, (\ell_1 - \ell_2)^2, (\ell_2 - \ell_3)^2, (\ell_3 - \ell_1)^2 \right\}$$

IBP reduction:

conservative: $\mathcal{O}(10^2)$ master integrals

full: $\mathcal{O}(10^3)$ master integrals

Differential Equations

Dlapa-Kälin-ZL-Porto PRL 2022 PLB 2022

$$\frac{d\vec{f}(x, \epsilon)}{dx} = \epsilon \mathcal{M}(x) \vec{f}(x, \epsilon)$$

- The majority can be solved in terms of **multiple polylogarithms**.
- Elliptic integrals appear in post-Minkowskian gravity for the first time.

NNNLO: 4PM

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The full impulse at $\mathcal{O}(G^4)$: Dlapa-Kälin-ZL-Porto PRL 2022 JHEP 2023 Dlapa-Kälin-ZL-Neef-Porto PRL 2023

$$\Delta p_1^\mu|_{\text{NNNLO}} = \frac{G^4}{|b|^4} \left(c_b \frac{b^\mu}{|b|} + c_1 \frac{\gamma u_2^\mu - u_1^\mu}{\gamma^2 - 1} + c_2 \frac{\gamma u_1^\mu - u_2^\mu}{\gamma^2 - 1} \right)$$

$$\begin{aligned} \frac{c_b}{\pi} = & -\frac{3h_{34}m_2m_1(m_1^3+m_2^3)}{64v_\infty^5} + \frac{m_1^2m_{12}m_2^2}{4} \left[\frac{3h_6K^2(w_2)}{4v_\infty^3} - \frac{3h_8K(w_2)E(w_2)}{4v_\infty^3} + \frac{21h_5w_3E^2(w_2)}{8v_\infty^3} - \frac{\pi^2h_{16}v_\infty}{4(\gamma+1)} + \frac{3\gamma h_{10}(Li_2(w_2) - 4Li_2(\sqrt{w_2}))}{w_3v_\infty^2} \right. \\ & + \log(v_\infty) \left(\frac{h_{32}}{2v_\infty^3} - \frac{3h_{14}\log(\frac{w_3}{2})}{v_\infty} - \frac{3\gamma h_{22}\log(w_1)}{2v_\infty^4} \right) \left. \right] + m_2^2m_1^3 \left[\frac{h_{52}}{48v_\infty^6} - \frac{h_{63}}{768\gamma^9w_3v_\infty^5} - \frac{3v_\infty(h_{40}Li_2(w_2) + 2w_3h_{33}Li_2(-w_2))}{64w_3} \right. \\ & + \frac{3h_{14}\log(\frac{w_3}{2})\log(w_3)}{4v_\infty} + \frac{\gamma h_{39}\log(w_1)}{8w_3^3v_\infty^2} + \frac{3\gamma h_{22}\log(w_3)\log(w_1)}{8v_\infty^4} - \frac{h_{35}\log(\frac{w_3}{2})}{32v_\infty^5} + \frac{h_{56}\log(2) - h_{57}\log(w_3) + 2\gamma h_{55}\log(\gamma)}{32v_\infty^5} - \frac{\gamma h_{51}\log(w_1)}{16v_\infty^7} \left. \right] \\ & + m_1^2m_2^3 \left[\frac{h_{58}}{192\gamma^7v_\infty^5} + \frac{h_{53}}{48v_\infty^6} + \frac{\gamma h_{49}\log(w_1)}{16v_\infty^6} - \frac{2\gamma h_{50}\log(w_1) + 3\gamma^2h_{13}\log^2(w_1)}{32v_\infty^7} - \frac{h_{41}\log(\frac{w_3}{2})}{8v_\infty^4} + \frac{3\gamma\log(w_1)(5h_{26}\log(2) + 8h_{12}\log(w_3))}{8v_\infty^4} \right. \\ & - \frac{h_{36}\log(w_3)}{4v_\infty^3} + \frac{\gamma h_{30}\log(\gamma)}{2v_\infty^3} + \frac{h_{37}\log(2)}{8v_\infty^3} + \frac{3(h_{17}w_3Li_2(w_2) - 2h_{23}Li_2(-w_2) + h_{15}\log^2(w_3) - h_9\log^2(2))}{8v_\infty} - \frac{3h_7\log(2)\log(w_3)}{v_\infty} \left. \right] \\ c_1 = & m_1m_2^2 \left(\frac{2h_{46}m_{12s}}{v_\infty^6} + \frac{9\pi^2h_1m_{12}^2}{32v_\infty^2} \right) + m_1^2m_2^3 \left(\frac{4\gamma h_{47}}{3v_\infty^6} - \frac{8\gamma h_2\log(w_1)}{v_\infty^6} + \frac{16h_{25}\log(w_1)}{v_\infty^3} - \frac{8h_3}{3v_\infty^5} \right) \\ c_2 = & -m_1^4m_2 \left(\frac{9\pi^2h_1}{32v_\infty^2} + \frac{2h_{46}}{v_\infty^6} \right) + m_2^2m_1^3 \left[\frac{h_{60}}{705600\gamma^8v_\infty^5} - \frac{4\gamma h_{48}}{3v_\infty^6} + \frac{3h_{38}(Li_2(w_2) - 4Li_2(\sqrt{w_2})) - \gamma h_{21}(Li_2(-w_1^2) + 2\log(\gamma)\log(w_1))}{16v_\infty^4} \right. \\ & + \frac{3\gamma h_{31}(2Li_2(-w_1) + \log(w_1)\log(w_3))}{8v_\infty^4} + \frac{h_{62}\log(w_1)}{6720\gamma^9v_\infty^6} + \frac{32\gamma^2h_{44}\log^2(w_1)}{v_\infty^7} + \frac{8\gamma(2h_4\log(2) - h_{27}\log(w_1))\log(w_1)}{v_\infty^4} - \frac{32h_{29}\log(w_1)}{3v_\infty^3} + \frac{\pi^2h_{42}}{192v_\infty^4} \left. \right] \\ & + m_2^3m_1^2 \left[\frac{h_{59}}{1440\gamma^7v_\infty^5} - \frac{h_{19}(Li_2(-w_1^2) + 2\log(\gamma)\log(w_1))}{8v_\infty^4} + \frac{h_{43}(Li_2(w_2) - 4Li_2(\sqrt{w_2}))}{32v_\infty^4} - \frac{h_{20}(2Li_2(-w_1) + \log(w_1)\log(w_3))}{4v_\infty^4} \right. \\ & - \frac{h_{61}\log(w_1)}{480\gamma^8v_\infty^6} - \frac{16\gamma^2h_{11}\log^2(w_1)}{v_\infty^4} - \frac{32\gamma h_{45}\log^2(w_1)}{v_\infty^7} + \frac{16\gamma h_{28}\log(w_1)}{5v_\infty^3} - \frac{32h_{24}\log(2)\log(w_1)}{v_\infty^4} - \frac{\pi^2h_{18}}{48v_\infty^4} - \frac{2h_{54}}{45v_\infty^6} \left. \right] \end{aligned}$$

with $\gamma \equiv u_1 \cdot u_2$, $v_\infty = \sqrt{\gamma^2 - 1}$, $w_1 = \gamma - v_\infty$, $w_2 = \frac{\gamma-1}{\gamma+1}$, $w_3 = \gamma + 1$, h_i = polynomial in γ .

$$Li_2(z) \equiv \int_0^1 \frac{dx \log(1-x)}{1-x}$$

$$K(z) \equiv \int_0^1 \frac{dx}{\sqrt{(1-x^2)(1-zx^2)}}$$

$$E(k) \equiv \int_0^1 dx \frac{\sqrt{1-zx^2}}{\sqrt{1-x^2}}$$

The full impulse at $\mathcal{O}(G^4)$: [Dlapa-Kälin-ZL-Porto JHEP 2023](#) [Dlapa-Kälin-ZL-Neef-Porto PRL 2023](#)

$$\Delta p_1^\mu|_{\text{NNNLO}} = \frac{G^4}{|b|^4} \left(c_b \frac{b^\mu}{|b|} + c_1 \frac{\gamma u_2^\mu - u_1^\mu}{\gamma^2 - 1} + c_2 \frac{\gamma u_1^\mu - u_2^\mu}{\gamma^2 - 1} \right)$$

- We obtained the full dynamics of binary inspirals at $\mathcal{O}(G^4)$ for the first time.
- Conservative part agrees perfectly with previous derivations.

[Bern-Parra-Martinez-Roiban-Ruf-Shen-Solon-Zeng 2022](#) [Dlapa-Kälin-ZL-Porto PRL 2022](#)

- Perfect agreement with the state-of-the-art PN computations

[Cho-Dandapat-Gopakumar 2021](#) [Cho 2022](#) [Bini-Geralico 2021 2022](#) [Bini-Damour 2022](#)

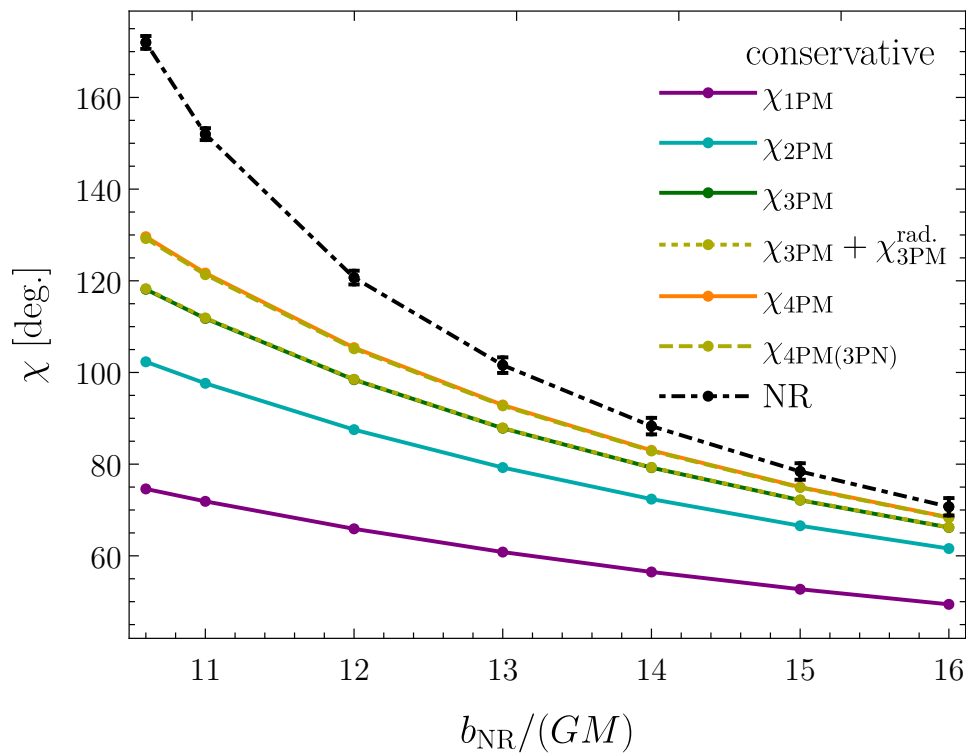
- Later, two new calculations confirmed our results.

[Damgaard-Hansen-Planté-Vanhove 2023](#) (exponentiation of amplitudes)

[Jakobsen-Mogull-Plefka-Sauer-Xu 2023](#) (worldline formalism)

Analytic vs Numerical Relativity

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Khalil-Buonanno-Steinhoff-Vines 2204.05047

NNNLO: local-in-time part

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- The full result does not describe generic elliptic-like motion due to nonlocal-in-time effects.

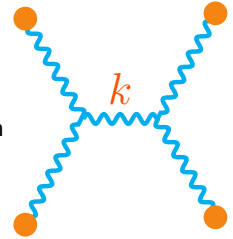
Damour-Jaranowski-Schäfer 2014 Galley-Leibovich-Porto-Ross 2015 Cho-Kälin-Porto 2021

- Nonlocal-in-time radial action:

$$\mathcal{S}_r^{(\text{nloc})} = -\frac{GE}{2\pi} \int_{-\infty}^{\infty} \frac{d\omega}{2\pi} \frac{dE}{d\omega} \log \left(\frac{4\omega^2}{\mu^2} e^{2\gamma_E} \right)$$

- The 4PM integrand can be built from 3PM diagrams.

$$\omega \equiv k \cdot u_{\text{com}}$$



$$\int d^D \ell_1 d^D \ell_2 \frac{\delta(\ell_1 \cdot u_1) \delta(\ell_2 \cdot u_2)}{[\ell_1 \cdot u_2][\ell_2 \cdot u_1]} \frac{\log(\omega^2)}{[\ell_1^2][\ell_2^2][(\ell_1 + \ell_2 - q)^2][(\ell_1 - q)^2][(\ell_2 - q)^2]}$$

- We managed to compute the integrals and obtained nonlocal-in-time contribution:

$$\begin{aligned} & \frac{\nu}{(\gamma^2 - 1)^2} \left[h_1 + \frac{\pi^2 h_2}{\sqrt{\gamma^2 - 1}} + h_3 \log \frac{\gamma + 1}{2} + \frac{h_4 \text{arccosh}(\gamma)}{\sqrt{\gamma^2 - 1}} + h_5 \log \frac{\gamma - 1}{8} + h_6 \log^2 \frac{\gamma + 1}{2} + \frac{h_8 \log(2) \text{arccosh}(\gamma)}{\sqrt{\gamma^2 - 1}} \right. \\ & \left. + h_7 \text{arccosh}(\gamma)^2 + h_9 \log \frac{\gamma - 1}{8} \log \frac{\gamma + 1}{2} + \frac{h_{10} \log \frac{\gamma^2 - 1}{16} \text{arccosh}(\gamma)}{\sqrt{\gamma^2 - 1}} + h_{11} \text{Li}_2 \frac{\gamma - 1}{\gamma + 1} + h_{12} \frac{\text{arccosh}^2(\gamma) + 4 \text{Li}_2(\sqrt{\gamma^2 - 1} - \gamma)}{\sqrt{\gamma^2 - 1}} \right] \end{aligned}$$

Coefficients h_i : exact- ν (iterated elliptic integrals) and SF-expanded (30SF) forms

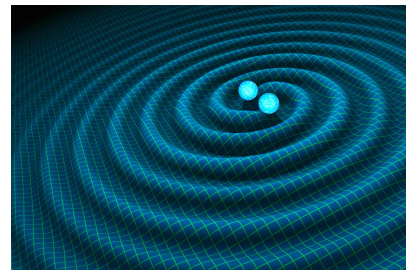
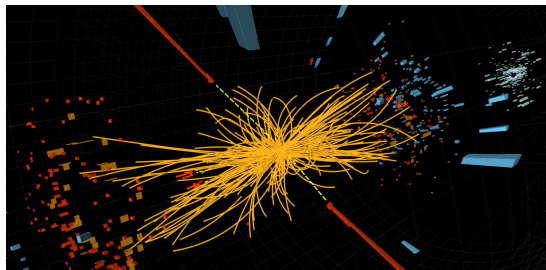
Dlapa-Kälin-ZL-Porto PRL 2024

- Using 6PN results in the literature, we constructed an improved bound Hamiltonian.

Conclusion & Outlook

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Modern techniques from Quantum Field Theory have already proven useful to improve theoretical predictions for gravitational-wave observables.



We have developed an efficient framework and made breakthroughs to NNNLO.

- Conservative spin & tidal effects at NLO [JHEP 06 \(2021\) 012](#) [PRD 102 \(2020\) 124025](#)
- Conservative dynamics at NNLO [PRL 125 \(2020\) 261103](#)
- Conservative dynamics at NNNLO [PLB 822 \(2021\) 136698](#) [PRL 128 \(2022\) 161104](#) [PRL 130 \(2023\) 101401](#)
- Local-in-time & nonlocal-in-time separation to $N^4\text{LO}$ [PRL 132 \(2024\) 221401](#) [arXiv:2506.20665](#)
- Novel techniques to evaluate classical loop integrals [JHEP 07 \(2023\) 181](#) [JHEP 08 \(2023\) 109](#)

Outlook: precision frontier

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GW science is just starting! discovery potential = precise theoretical predictions!

	0PN	1PN	2PN	3PN	4PN	5PN	6PN	...		
1PM	$\left(\frac{GM}{r}\right) \times$	$\left(1 + v^2 + v^4 + v^6 + v^8 + v^{10} + v^{12} + \dots\right)$	$=$	$\left(\frac{GM}{r}\right) \times (1)$						
2PM		$\left(\frac{GM}{r}\right)^2 \times$	$\left(1 + v^2 + v^4 + v^6 + v^8 + v^{10} + \dots\right)$	$=$	$\left(\frac{GM}{r}\right)^2 \times (1)$					
3PM			$\left(\frac{GM}{r}\right)^3 \times$	$\left(1 + v^2 + v^4 + v^6 + v^8 + \dots\right)$	$=$	$\left(\frac{GM}{r}\right)^3 \times (1 + \nu)$				
4PM				$\left(\frac{GM}{r}\right)^4 \times$	$\left(1 + v^2 + v^4 + v^6 + \dots\right)$	$=$	$\left(\frac{GM}{r}\right)^4 \times (1 + \nu)$			
5PM					$\left(\frac{GM}{r}\right)^5 \times$	$\left(1 + v^2 + v^4 + \dots\right)$	$=$	$\left(\frac{GM}{r}\right)^5 \times (1 + \nu + \nu^2)$		
6PM						$\left(\frac{GM}{r}\right)^6 \times$	$\left(1 + v^2 + \dots\right)$	$=$	$\left(\frac{GM}{r}\right)^6 \times (1 + \nu + \nu^2)$	
7PM							$\left(\frac{GM}{r}\right)^7 \times$	$\left(1 + \dots\right)$	$=$	$\left(\frac{GM}{r}\right)^7 \times (1 + \nu + \nu^2 + \nu^3)$
									0SF 1SF 2SF 3SF ...	

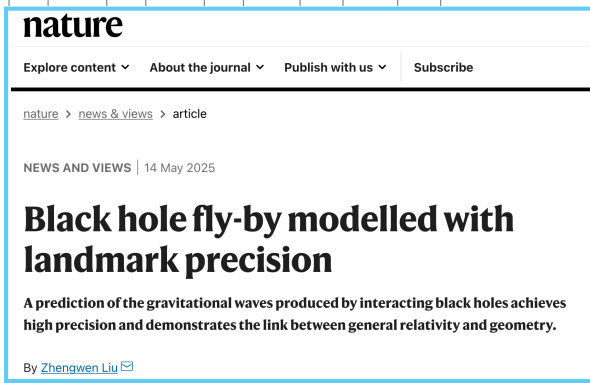
Outlook: precision frontier

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GW science is just starting! discovery potential = precise theoretical predictions!

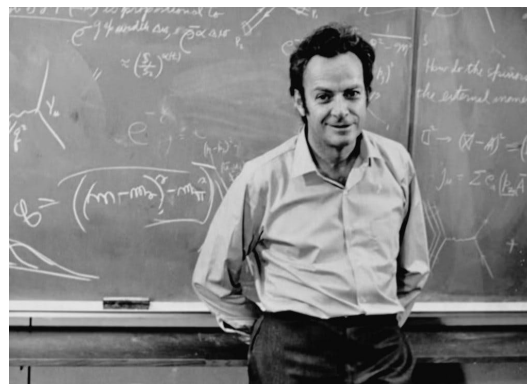
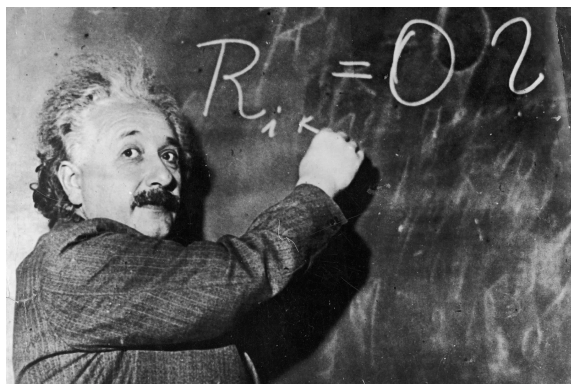
$$\begin{array}{l}
 \text{0PN} \quad \text{1PN} \quad \text{2PN} \quad \text{3PN} \quad \text{4PN} \quad \text{5PN} \quad \text{6PN} \quad \dots \\
 \text{1PM} \quad \left(\frac{GM}{r}\right) \times (1 + v^2 + v^4 + v^6 + v^8 + v^{10} + v^{12} + \dots) = \left(\frac{GM}{r}\right) \times (1) \\
 \text{2PM} \quad \left(\frac{GM}{r}\right)^2 \times (1 + v^2 + v^4 + v^6 + v^8 + v^{10} + \dots) = \left(\frac{GM}{r}\right)^2 \times (1) \\
 \text{3PM} \quad \left(\frac{GM}{r}\right)^3 \times (1 + v) \\
 \text{4PM} \quad \left(\frac{GM}{r}\right)^4 \times (1 + v) \\
 \text{5PM} \quad \left(\frac{GM}{r}\right)^5 \times (1 + v + v^2) \\
 \text{6PM} \quad \left(\frac{GM}{r}\right)^6 \times (1 + v + v^2) \\
 \text{7PM} \quad \left(\frac{GM}{r}\right)^7 \times (1 + v + v^2 + v^3)
 \end{array}$$

0SF 1SF 2SF 3SF ...



- ▶ 1SF: Driesse-Jakobsen-Mogull-Plefka-Sauer-Usovitsch 2024 D-J-Klemm-M-Nega-P-S-U 2024
- ▶ 1SF: local-in-time conservative dynamics: Dlapa-Kälin-ZL-Porto 2506.20665
- ▶ 2SF: nonplanar diagram, more complicated functions ...

Feynman's integrals solve Einstein's equations!



谢谢!