

Two-Loop Electroweak Radiative Corrections to Low-Energy Electron-Electron Scattering

Yong Du (杜勇)

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The 2025 International Workshop on the High Energy Circular Electron
Positron Collider

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Based on

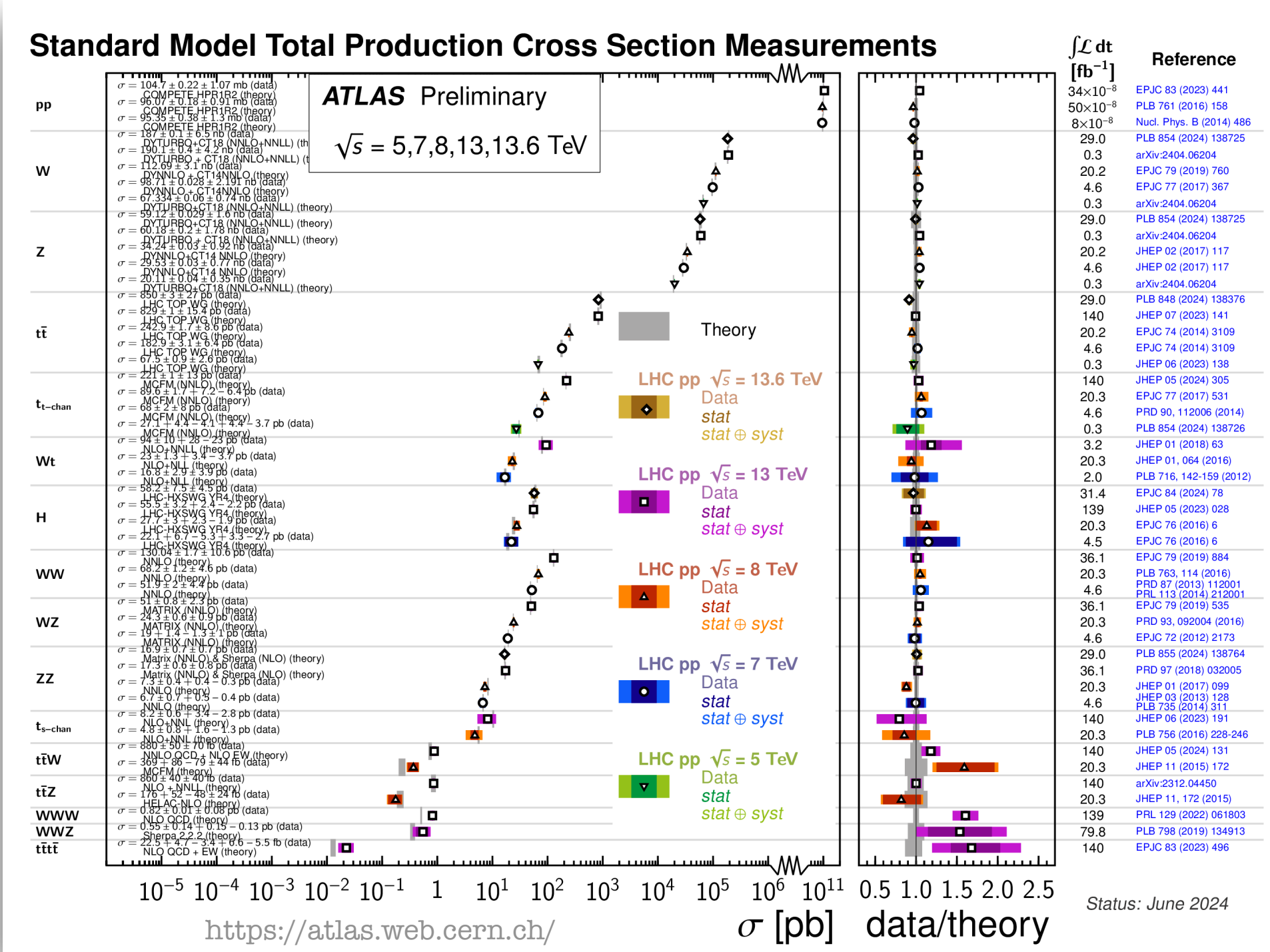
PRL 126 (2021) 13, 131801, with Ayres Freitas, Hiren Patel, Michael Ramsey-Musolf



中国科学院近代物理研究所
Institute of Modern Physics, Chinese Academy of Sciences

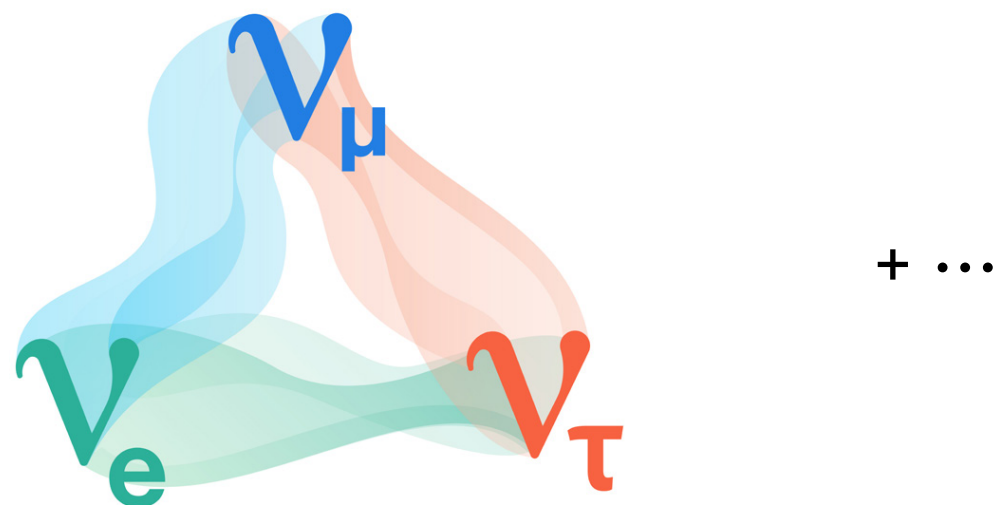
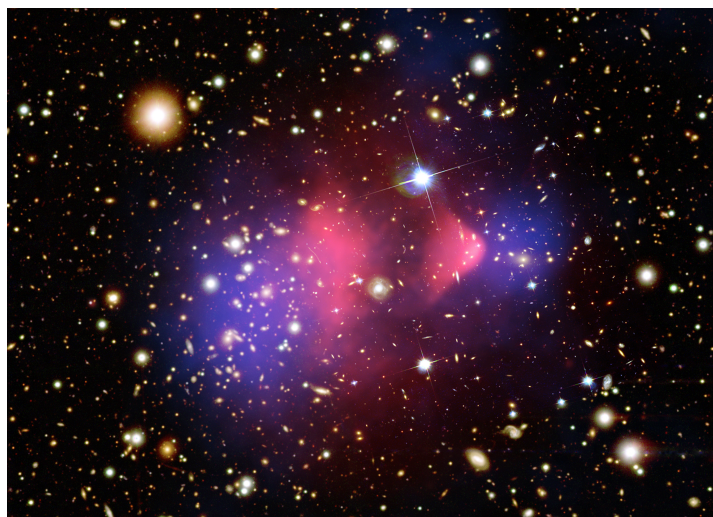
Introduction

The standard model of particle physics is quite *successful*



Introduction

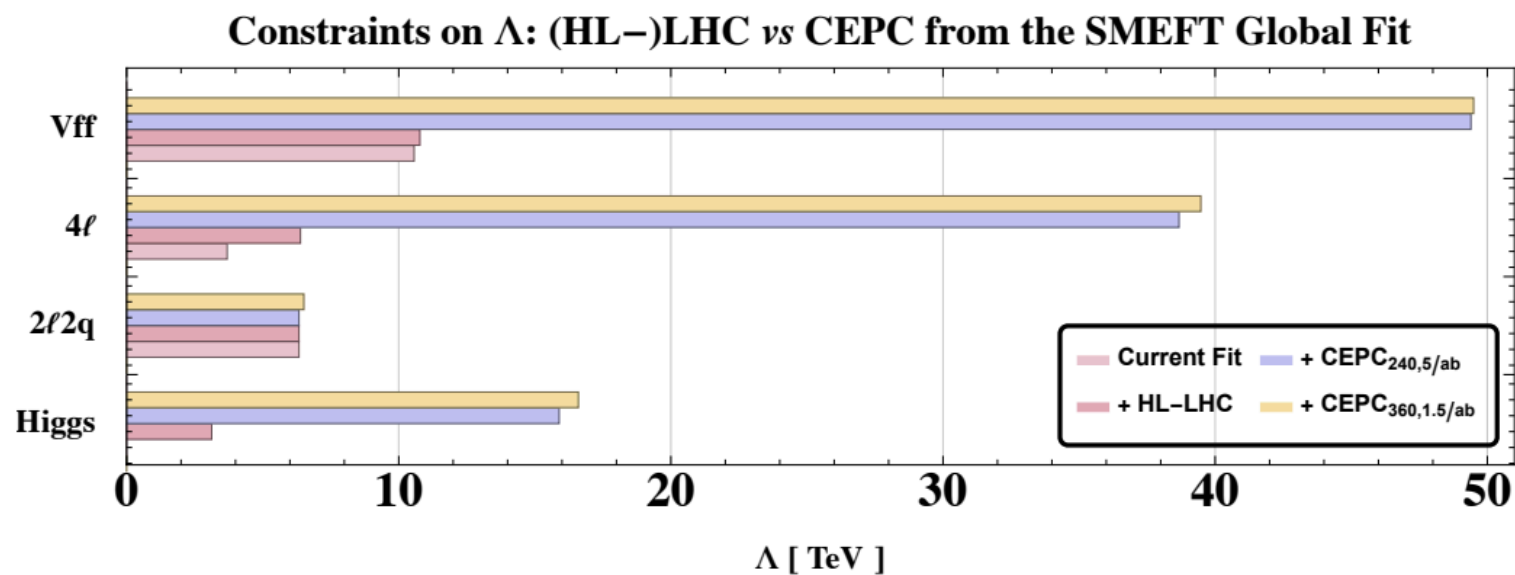
yet we know new physics beyond the SM is needed



CEPC: a powerful machine — broad reach + precision



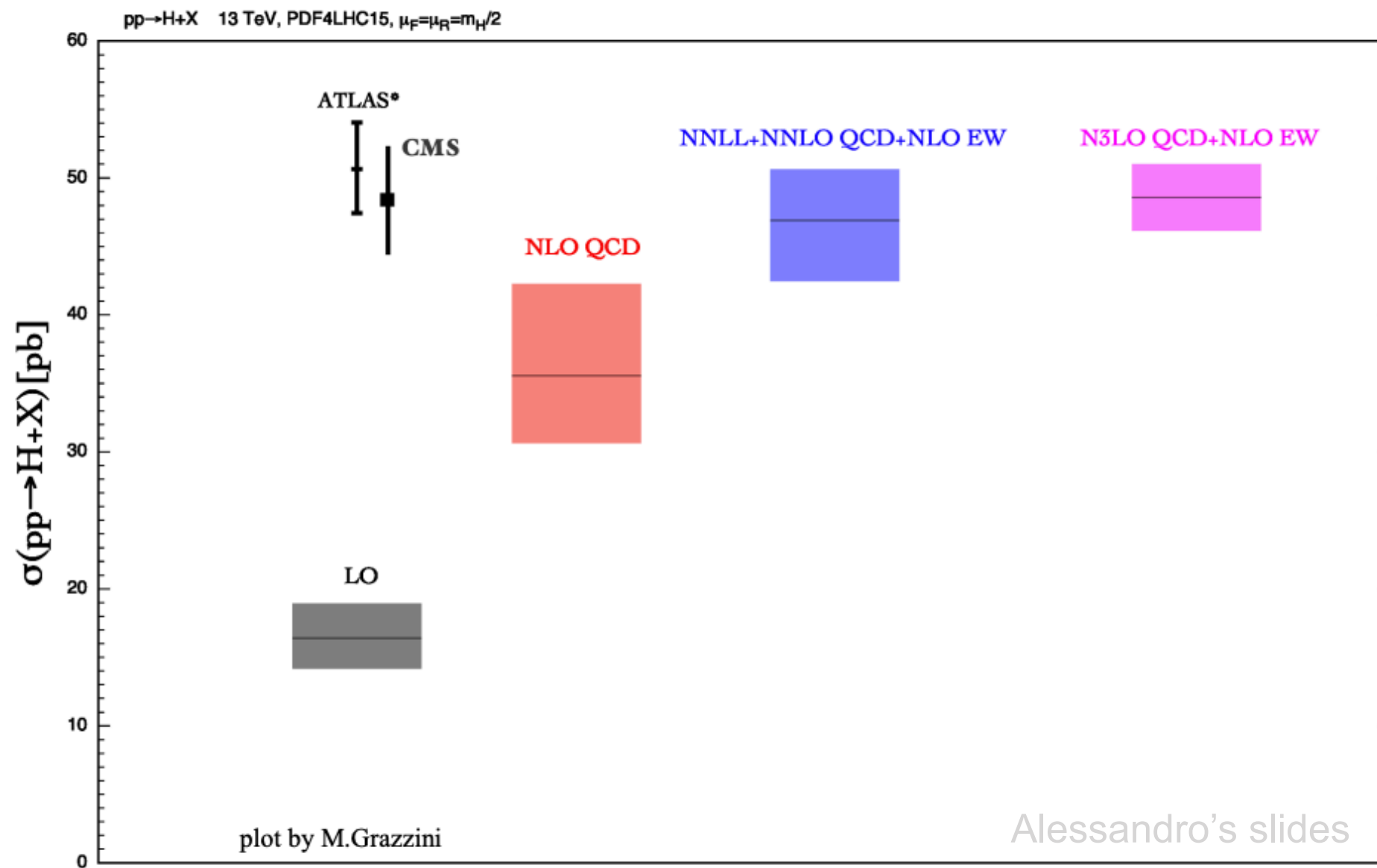
CEPC study group, 2505.24810



See Jorge's talk on Thursday for more

Introduction

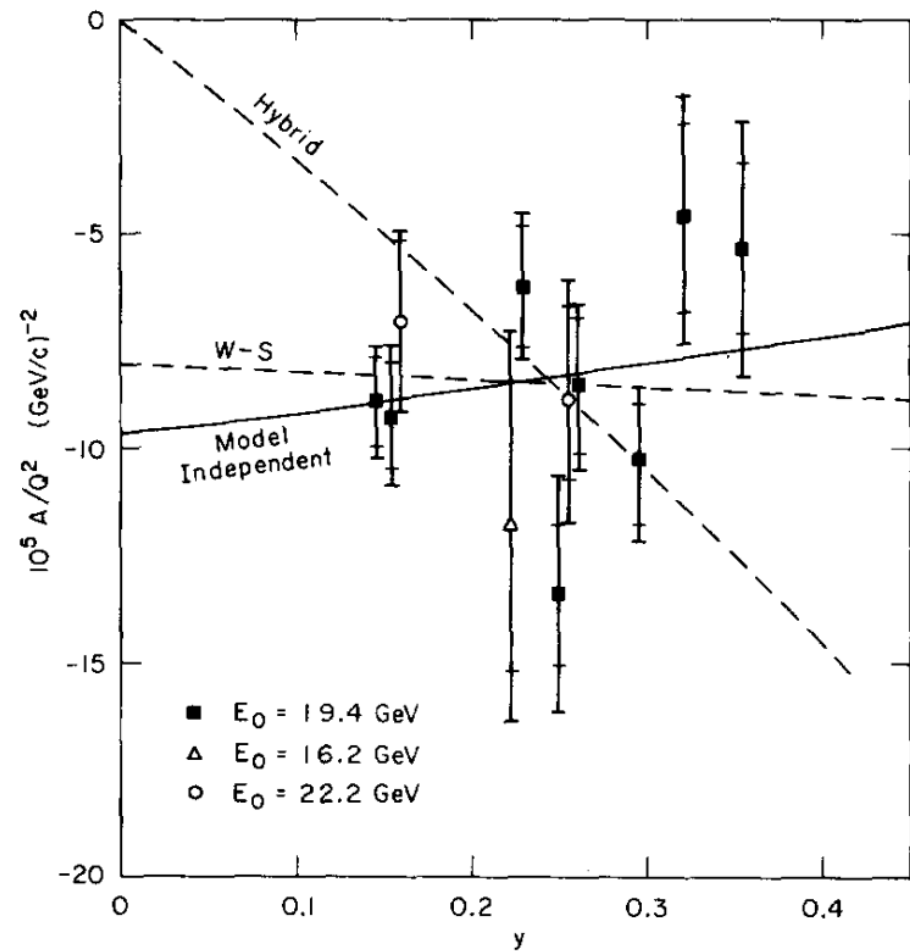
the statement on new physics needs carefulness — precision match + convergence (pert)



See Alessandro's talk on Thursday

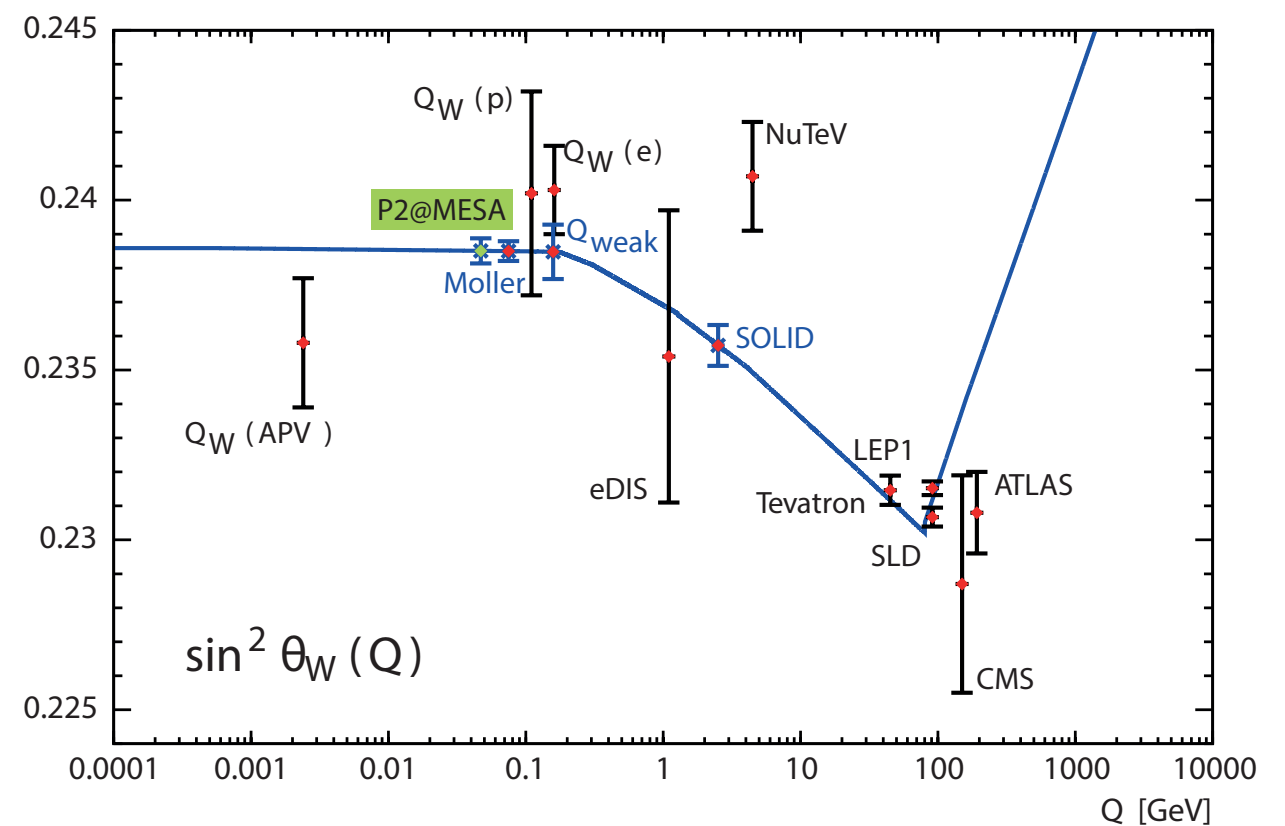
Introduction

Precision $\sin^2 \theta_W$ from parity-violating electron scattering is my focus today



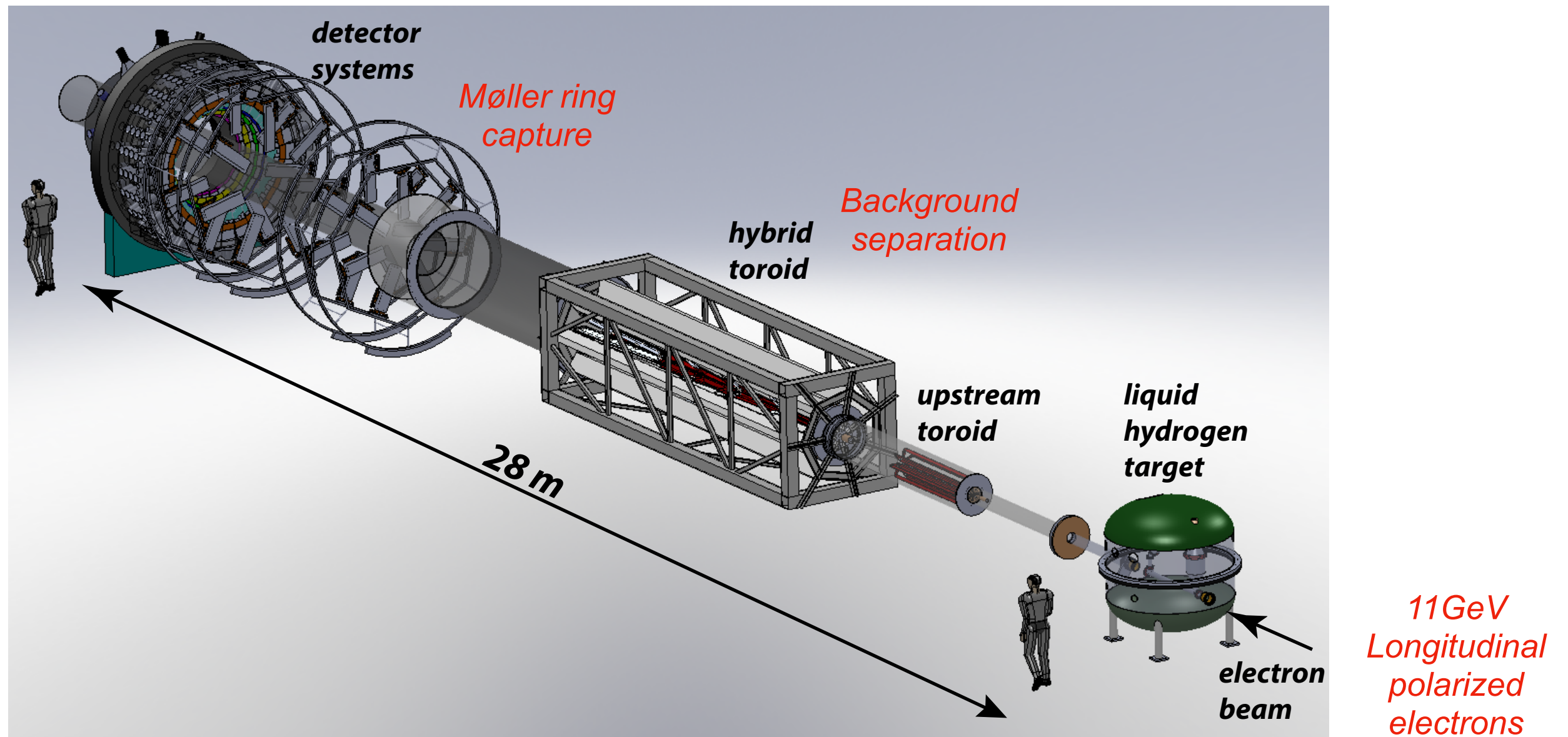
Prescott et al, 1978, 1979

(past) model selection



(future) precision test + ?

Introduction

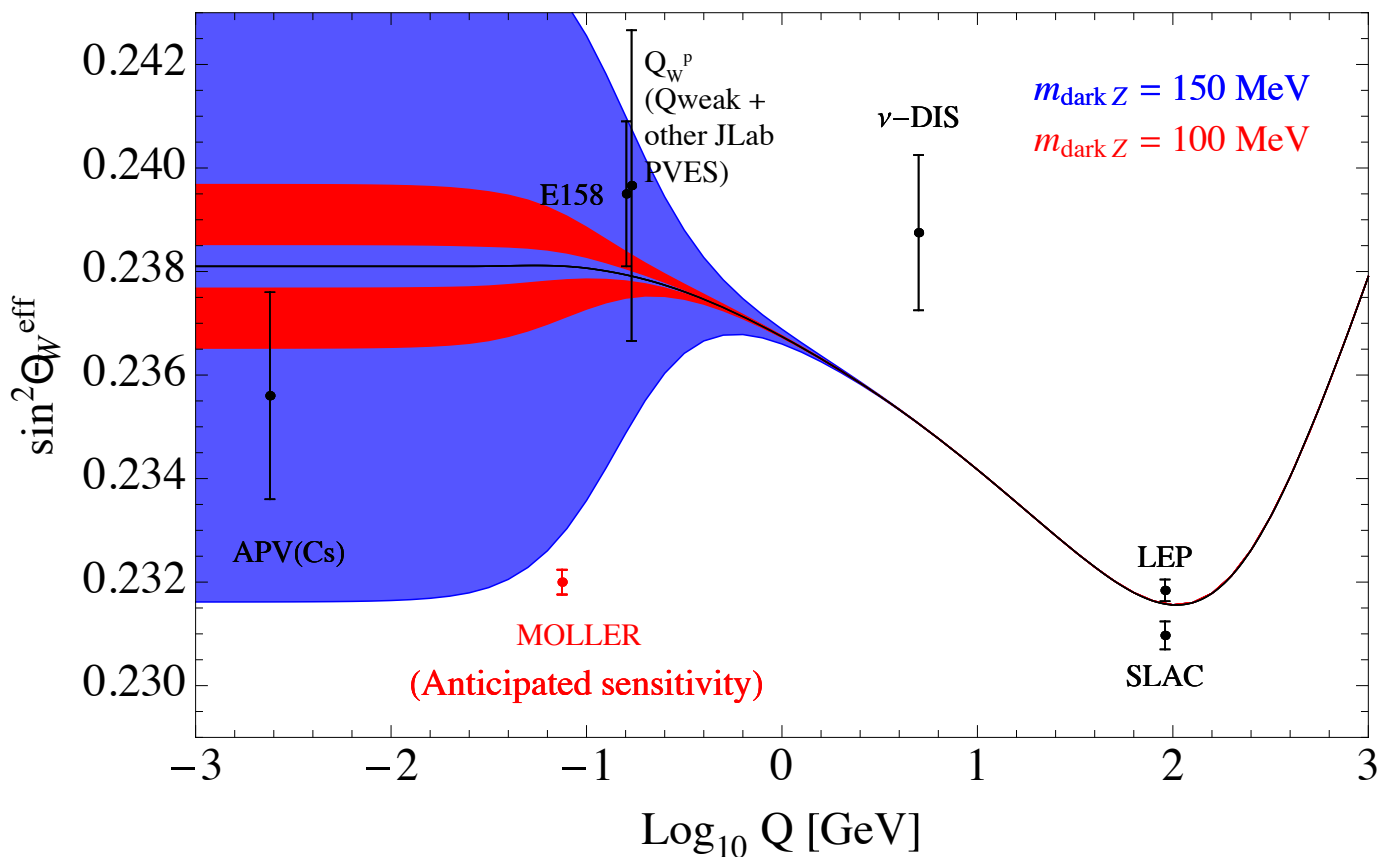


MOLLER Collaboration, 1411.4088

target precision: $\Delta Q_W^e = 1.1 \times 10^{-3}$

Introduction

Also good opportunity for new physics searches + complementary to CEPC!



MOLLER Collaboration, 1411.4088

$$\mathcal{L}_{\text{SM}}^{\text{PV}} = \frac{G_F}{2\sqrt{2}} g_A^e \bar{e} \gamma_\mu \gamma_5 e \sum_f g_V^f \bar{f} \gamma^\mu f$$

$$\mathcal{L}_{\text{NP}}^{\text{PV}} = \frac{4\pi\kappa^2}{\Lambda^2} \bar{e} \gamma_\mu \gamma_5 e \sum_f h_V^f \bar{f} \gamma^\mu f$$

1% relative precision in the nucleon/
electron weak charge leads to

$$\frac{\Delta Q_w}{Q_w} \leq 1\% \iff \Lambda \geq \left(\frac{8\sqrt{2}\pi\kappa^2}{0.01 G_F} \right)^{1/2} \approx 20 \kappa \text{ TeV}$$

Ramsey-Musolf, [hep-ph/9903264](https://arxiv.org/abs/hep-ph/9903264)

Similarly, we can ask: theoretical prediction precise enough and convergent?

Theoretical status: LO

LO calculation:

$[Q_W^e]_{\text{LO}}$ Derman and Marciano, 1979

$$A_{\text{PV}} = \frac{G_\mu Q^2}{\sqrt{2}\pi\alpha_{\text{EM}}} \frac{1-y}{1+y^4+(1-y)^4} Q_W^e$$

$$[Q_W^e]_{\text{LO}} = 1 - 4 \sin^2 \theta_W$$

Radiative corrections are parameterized as

$$Q_W^e = [Q_W^e]_{\text{LO}} + [Q_W^e]_{\text{NLO}} + [Q_W^e]_{\text{NNLO}} + \dots = \sum_n \left(\frac{\alpha_{\text{EM}}}{\pi} \right)^n c_n$$

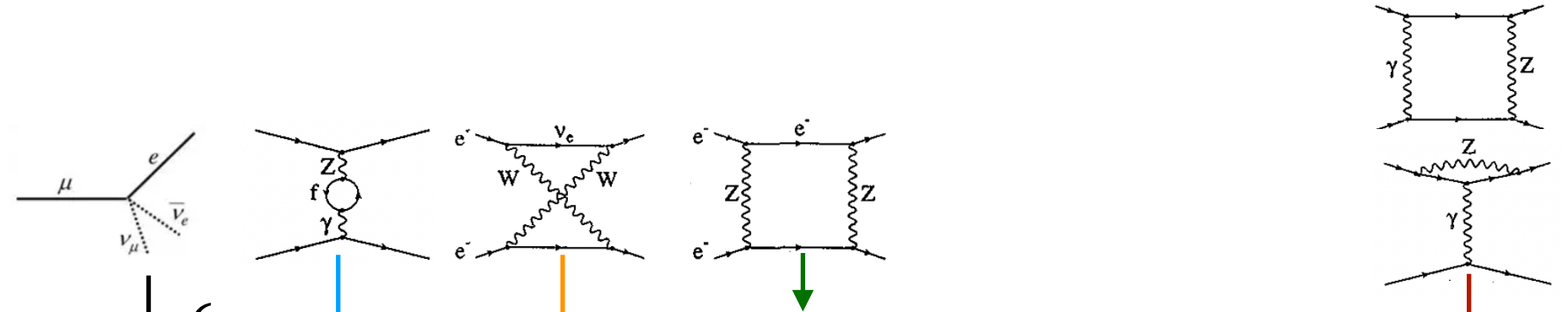
where terms proportional to G_μ in c_n are ignored. Plugging in $\sin^2 \theta_W(m_Z) = 0.2314$:

$$[Q_W^e]_{\text{LO}} = 1 - 4 \sin^2 \theta_W = 0.0744$$

Theoretical status: NLO

NLO calculation:

$[Q_W^e]_{\text{NLO}}$ Czarnecki and Marciano, 1996; Denner and Pozzorini, 1998; Petriello, 2003; Zykunov, 2004; Kolomensky et al, 2005; Zykunov et al, 2005; Zykunov, 2009; Aleksejevs et al, 2010, 2011, 2012

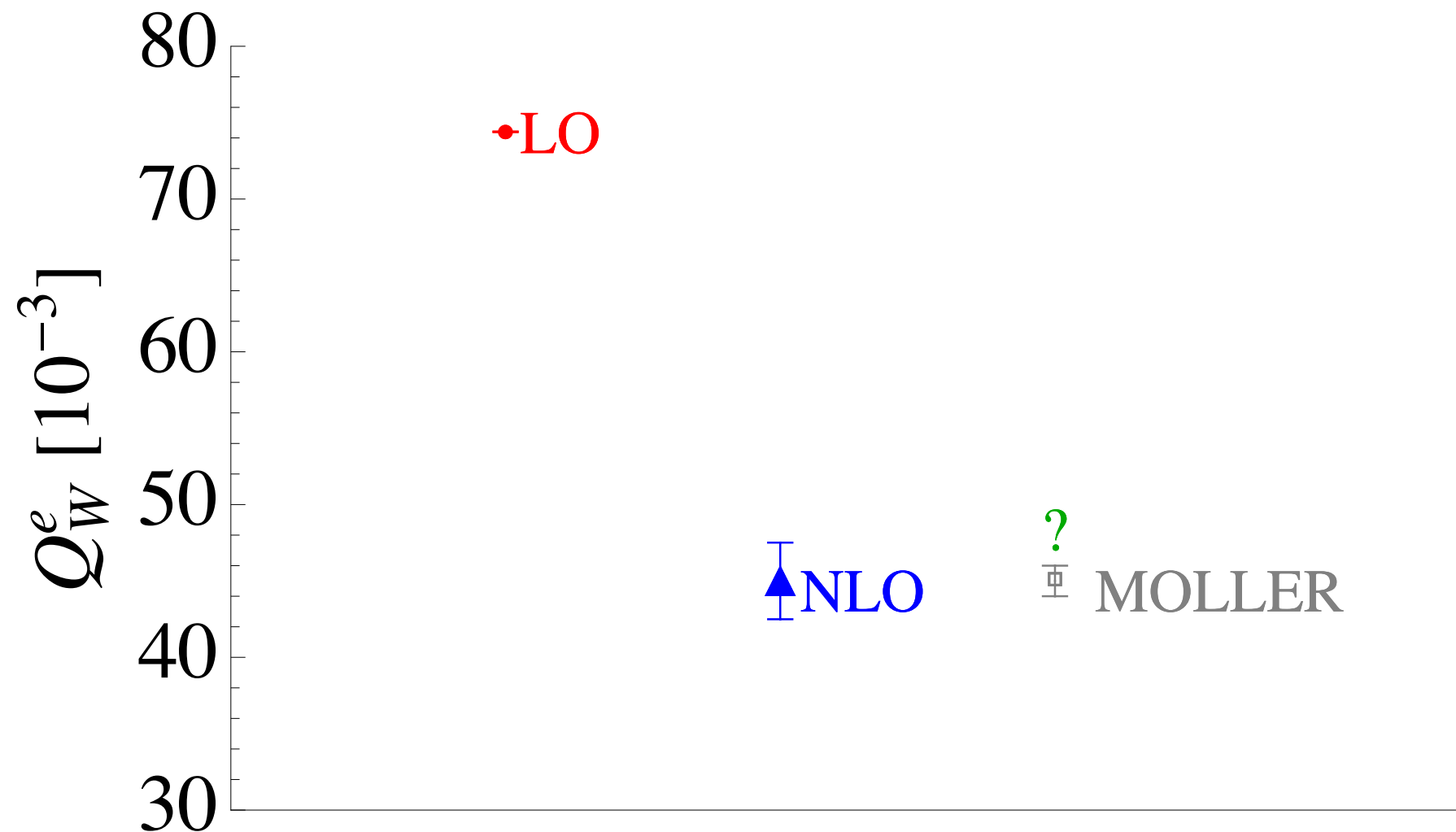


$$[Q_W^e]_{\text{LO+NLO}} = \rho \left\{ 1 - 4\kappa(0)s_W^2 + \frac{\alpha_{\text{EM}}}{4\pi s_W^2} - \frac{3\alpha_{\text{EM}}}{32\pi s_W^2 c_W^2} (1 - 4s_W^2)[1 + (1 - 4s_W^2)^2] + F_1(y, Q^2) + F_2(y, Q^2) \right\}$$

$$= 0.0450 \pm 0.0023 \pm 0.0010$$

Theoretical status: NLO

NLO calculation:

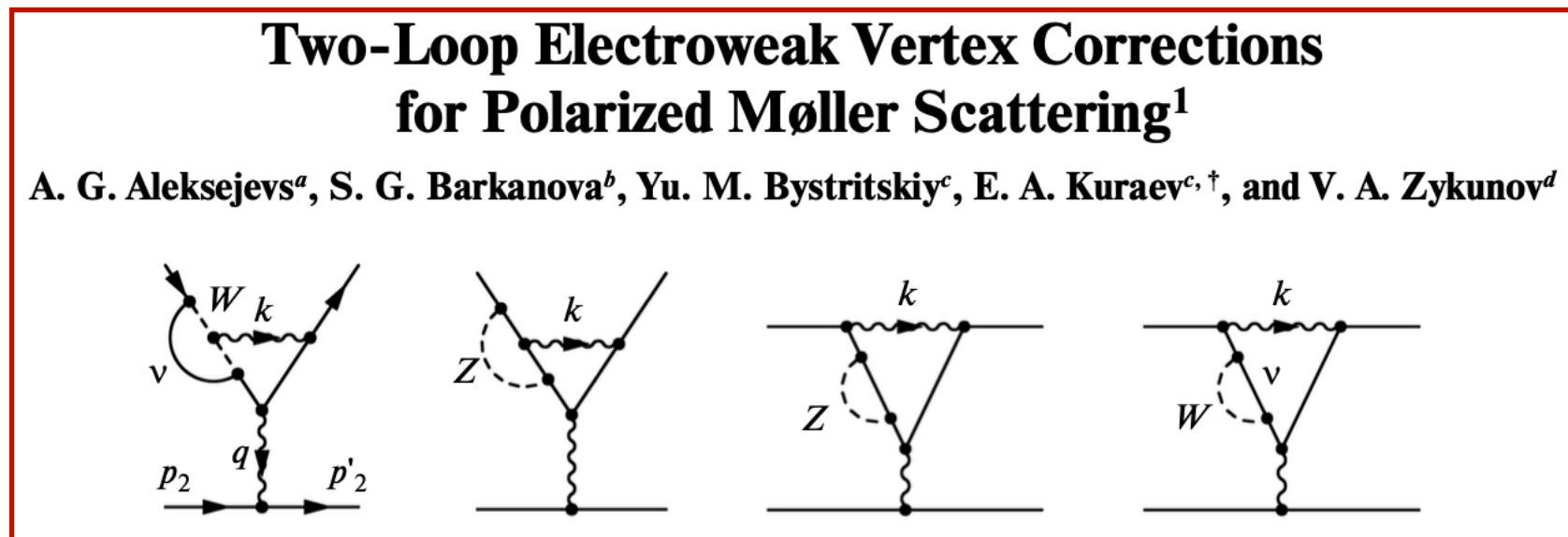
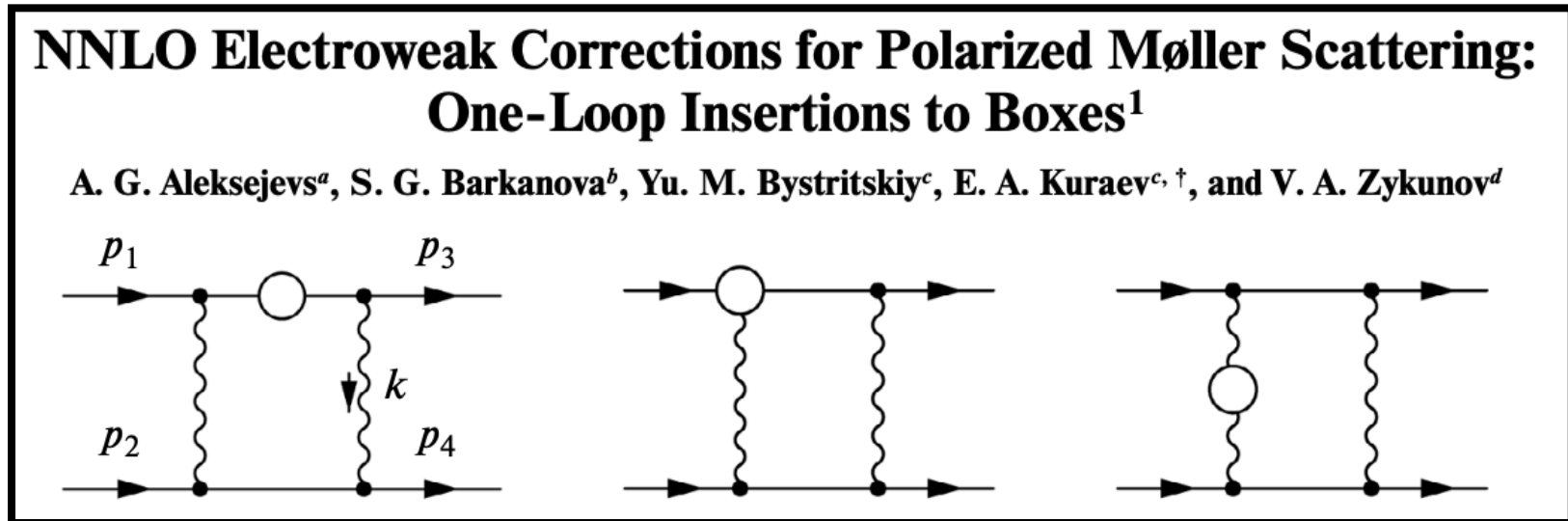


40% reduction! Not precise enough for future experiment, neither good convergence.

Theoretical status: NNLO

NNLO calculation: earlier efforts

$[Q_W^e]_{\text{NNLO}}$ Aleksejevs et al, 2011, 2012, 2015 (partial estimation)

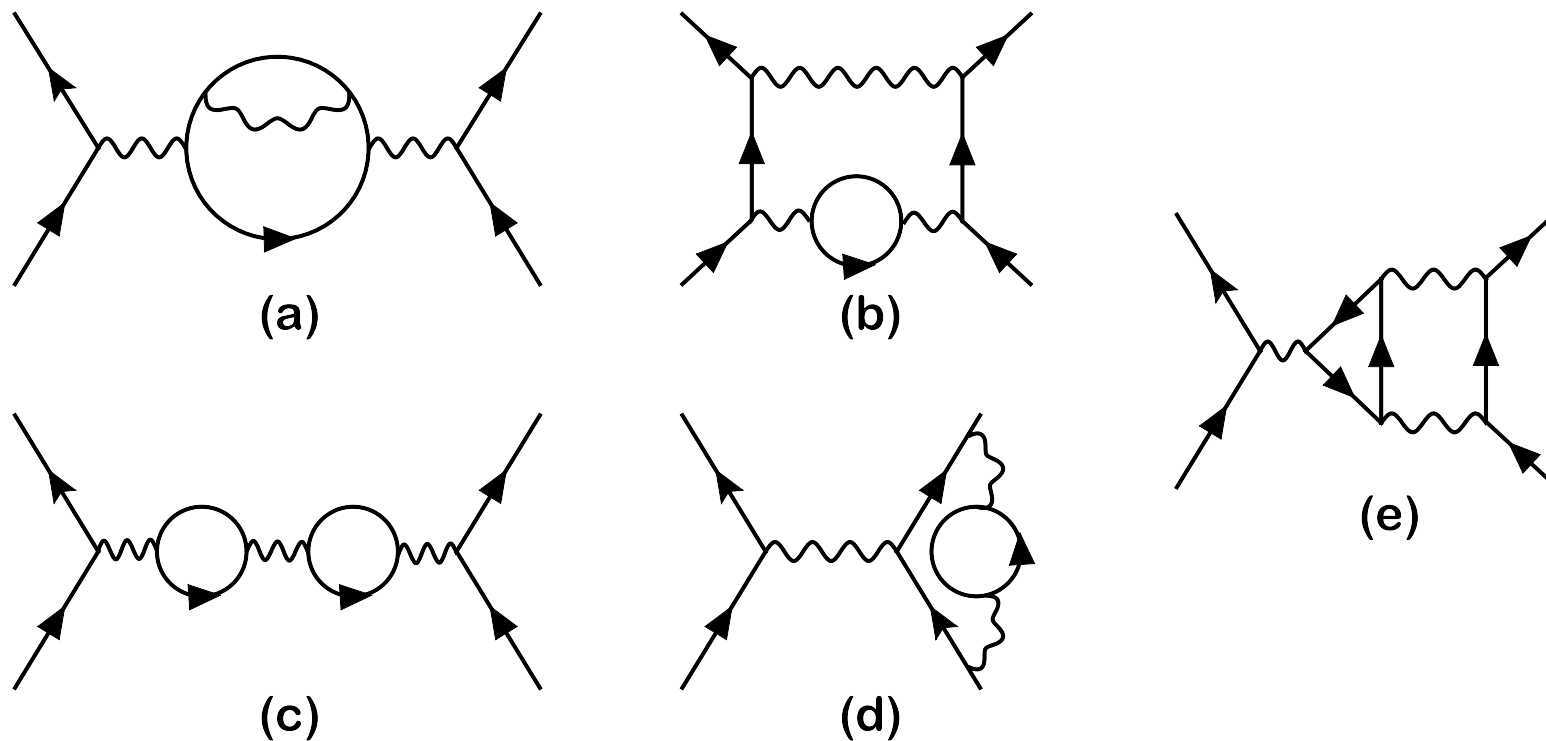


Take-aways: A full NNLO computation is needed for MOLLER.

Theoretical status: Our work @ NNLO

Our strategy: firstly focus on NNLO topologies with closed fermion loops

- Gauge-invariant subset
- Dominant by fermion numbers/flavors **vs** bosonic loops more EW scale suppression

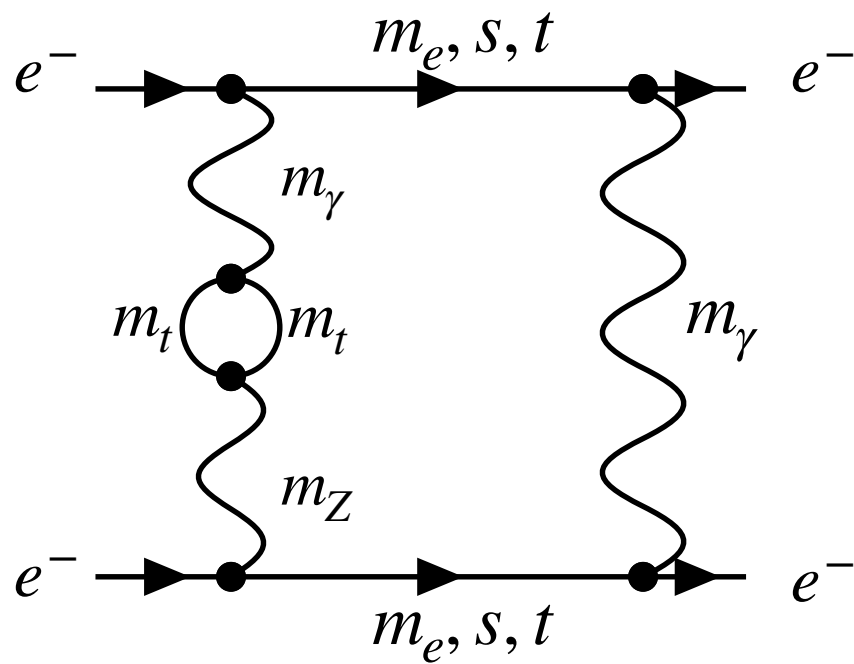


Work flow: FeynArts $\rightarrow$ FeynCalc/Package-X $\rightarrow$ code in-house $\rightarrow$ FIRE/Kira

See Qian's talk on yesterday

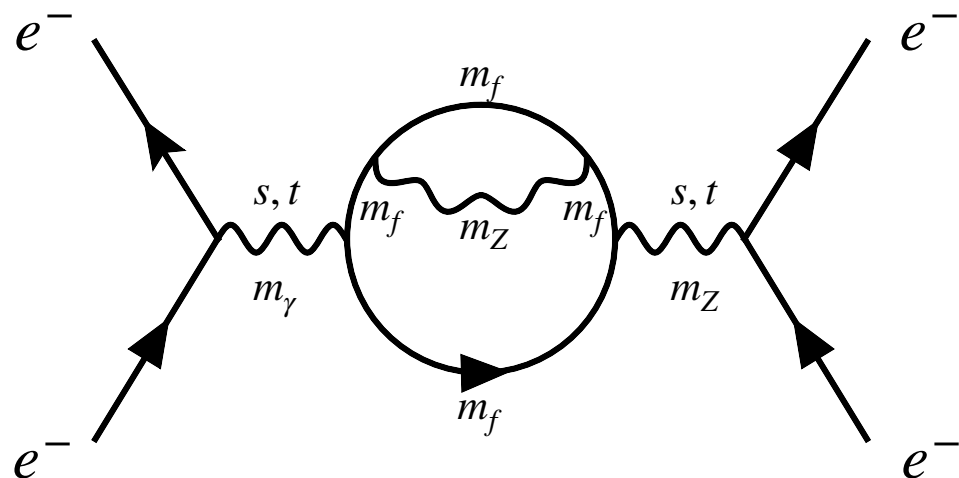
NNLO Methods: EBRs

Generically speaking, integrals at NNLO and beyond contain *multiple scales*, which complicate the analytic behavior of the loop integrand



$$I = \int \tilde{d}q_1 \tilde{d}q_2 f(q_1, q_2, m_t, m_Z, m_\gamma, m_e, s, t)$$

6 scales



$$I = \int \tilde{d}q_1 \tilde{d}q_2 f(q_1, q_2, m_Z, m_f, m_e, s, t)$$

5 scales

NNLO Methods: EBRs

For us, we aim for analytical results:

to gain stability through avoiding large number cancellation

to provide instantaneous evaluation with arbitrary precision

to track individual UV/IR divergence

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Is this (technically) doable to systematically handle $\mathcal{O}(10^3 \sim 10^4)$ Feynman diagrams ?

Yes! MOLLER is a low-energy scattering process, we have a natural hierarchy:

$$\begin{array}{ccccc} m_f & \ll & s, t & \ll & m_{W,Z,H,t} \\ \mathcal{O}(10^{-4}) \text{ GeV} & & \mathcal{O}(10^{-1}) \text{ GeV} & & \mathcal{O}(10^2) \text{ GeV} \end{array}$$

Asymptotic expansion based on this hierarchy renders it possible to systematically compute the integrals and to have analytical results in the end.

Another gain from this hierarchy, as we'll see shortly, is to oftentimes reform the two-loop integral into a one-loop one.

NNLO Methods: EBRs

This asymptotic strategy for *multiple scale* integrals is based on the method of regions (recall that we are only interested in $A_{PV} \propto G_\mu$)

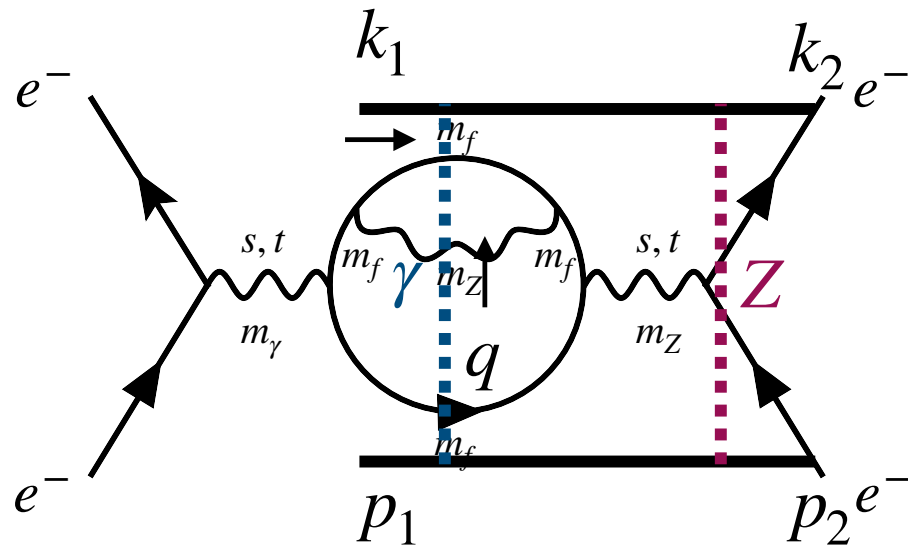


Fig. 1

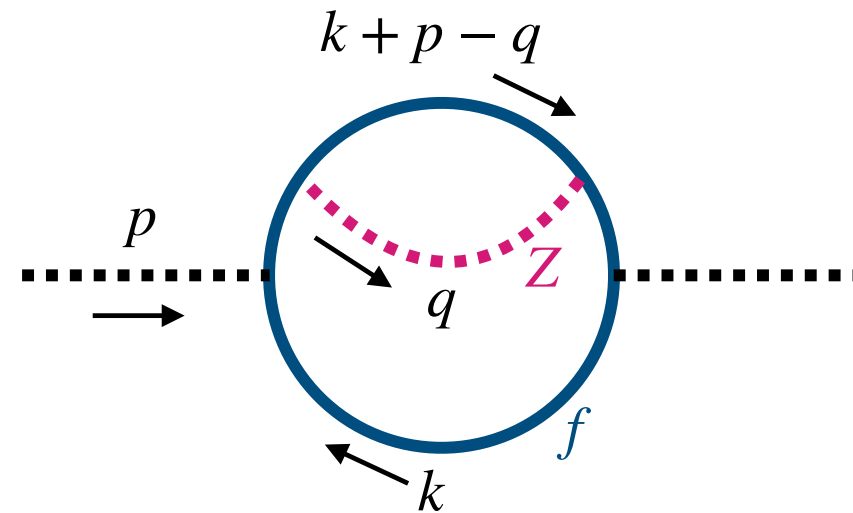


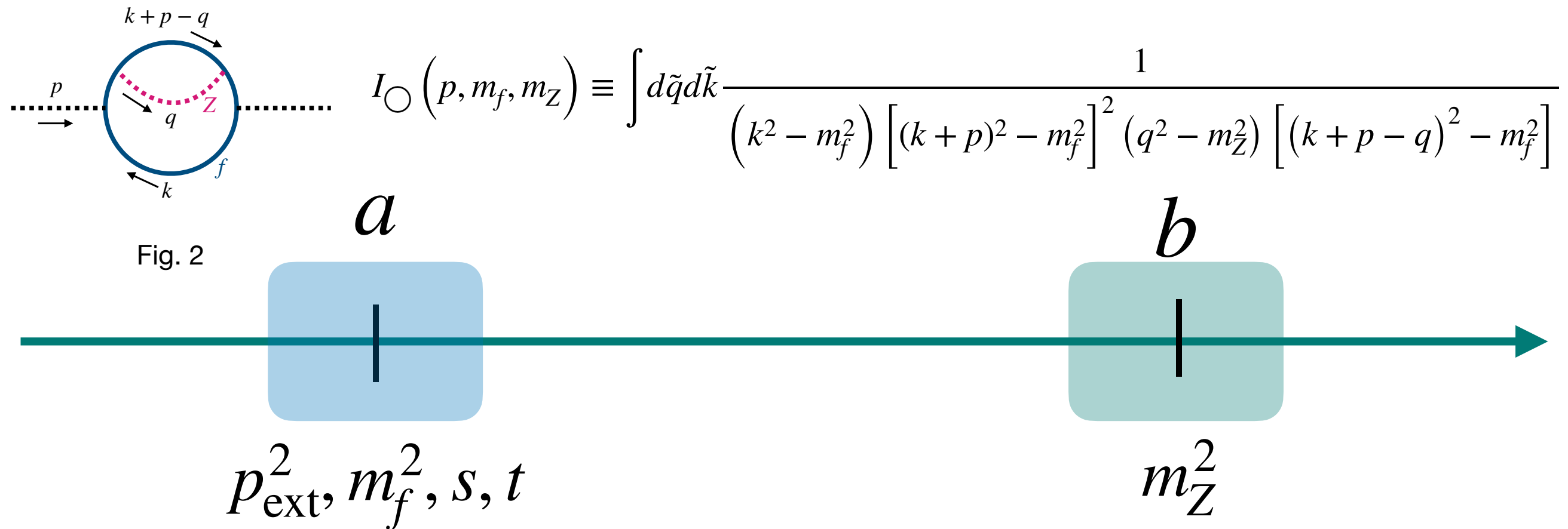
Fig. 2

Let's focus on the scalar integral (always achievable through Chisholm expansion)

$$I_{\bigcirc} \left(p, m_f, m_Z \right) \equiv \int d\tilde{q} d\tilde{k} \frac{1}{\left(k^2 - m_f^2 \right) \left[(k + p)^2 - m_f^2 \right]^2 (q^2 - m_Z^2) \left[(k + p - q)^2 - m_f^2 \right]}$$

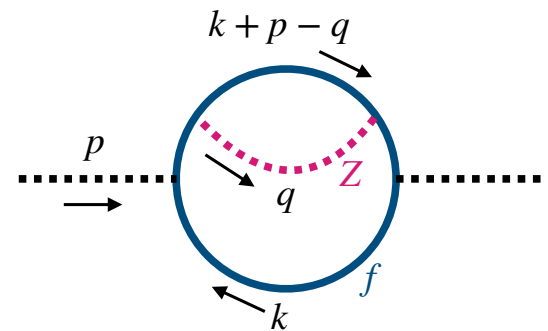
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Fig. 2

a

$$p_{\text{ext}}^2, m_f^2, s, t$$

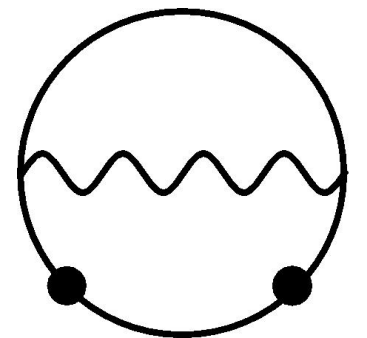
b

$$k^2, q^2, m_Z^2$$

hard-hard region: $\{m_f^2, p^2, s, t, u\} \ll k^2 \simeq q^2 \simeq (k-q)^2 \simeq m_Z^2$

$$I_{\bigcirc}^{hh} = \int d\tilde{q} d\tilde{k} \frac{1}{(k^2)^3 (q^2 - m_Z^2) (k-q)^2} + \dots$$

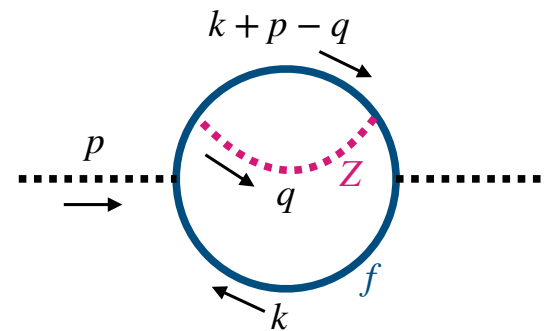
Power counting: $4 \times 2 - 2 \times 3 - 2 - 2 = -2$



single-scale vacuum

NNLO Methods: EBRs

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$$I_{\bigcirc}(p, m_f, m_Z) \equiv \int d\tilde{q} d\tilde{k} \frac{1}{(k^2 - m_f^2) [(k+p)^2 - m_f^2]^2 (q^2 - m_Z^2) [(k+p-q)^2 - m_f^2]}$$

Fig. 2

a

b

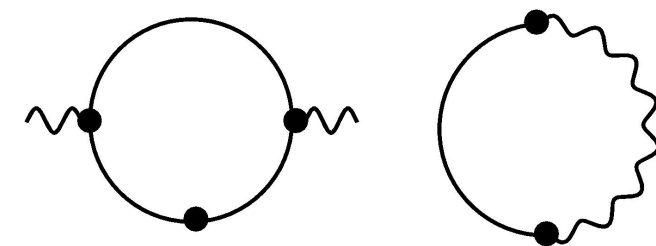
$$p_{\text{ext}}^2, m_f^2, s, t, k^2$$

$$q^2, m_Z^2$$

soft-hard region: $\{m_f^2, p^2, k^2, s, t, u\} \ll q^2 \simeq m_Z^2$

$$I_{\bigcirc}^{sh}(p, m_f, m_Z) = \int d\tilde{q} d\tilde{k} \frac{1}{(k^2 - m_f^2) [(k+p)^2 - m_f^2]^2 (q^2 - m_Z^2) q^2} + \dots$$

Power counting: $4 - 2 - 2 = 0$



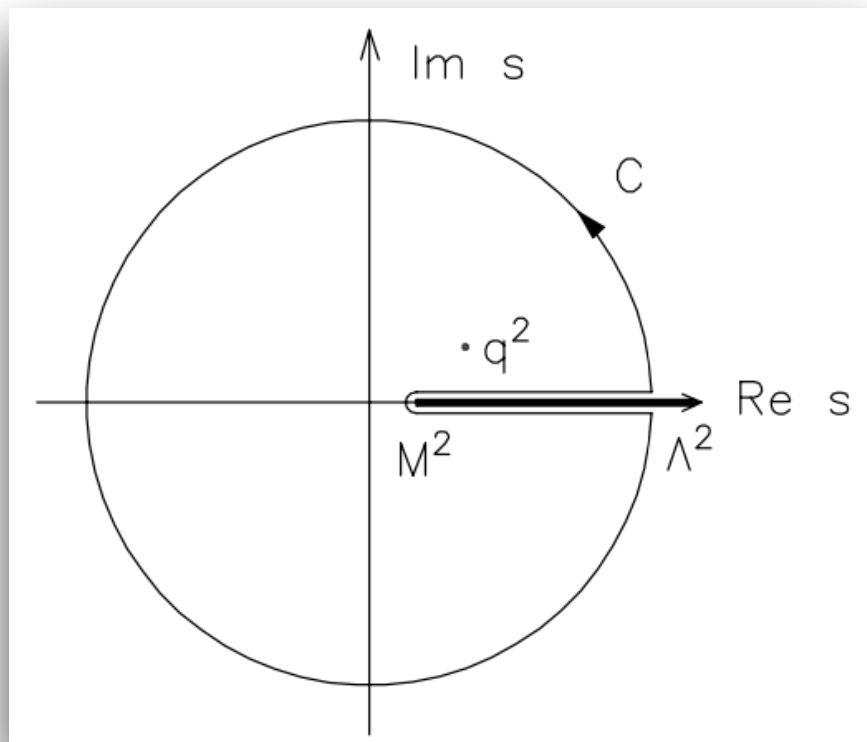
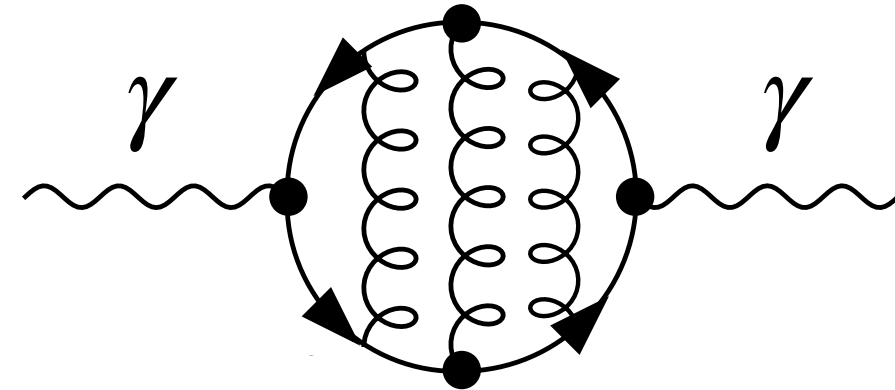
decoupled

NNLO Methods: DRs

This method of regions similarly applies to the other topologies, while for box diagrams, *dispersion relation* is another trick to gain control over hadronic uncertainties

This idea is mature:

The amplitude develops a non-vanishing absorptive part above some threshold, allowing the reconstruction of the full amplitude by integrating along the branch cut



$$\begin{aligned} F(q^2) &= \frac{1}{2\pi i} \oint_{\infty} ds \frac{F(s)}{s - q^2} \\ &= \frac{1}{\pi} \int_{M^2}^{\infty^2} ds \frac{\text{Im } F(s)}{s - q^2 - i\epsilon} + \frac{1}{2\pi i} \oint_{|s|=\infty} ds \frac{F(s)}{s - q^2} \end{aligned}$$

0

as long as analytic is guaranteed at infinity

NNLO Methods: DRs

Otherwise, we reconstruct the full result from the *n_{th} -subtracted dispersion relations*:

$$F(q^2) = \sum_{j=0}^{n-1} \frac{(q^2 - q_0^2)^j}{j!} F^{(j)}(q_0^2) + \frac{(q^2 - q_0^2)^n}{\pi} \int_{M^2}^{\infty} ds \frac{\text{Im}F(s)}{(s - q_0^2)^n (s - q^2 - i\epsilon)}$$

with n (typically) the *minimal* number that guarantees analyticity at infinity

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For one-loop $\gamma\gamma$ and γZ self energies, the physical condition is $\Pi_T^{\gamma\gamma, \gamma Z}(0) = 0$ and $n = 1$ is enough since $\Pi_T^{\gamma\gamma, \gamma Z}(s) \rightarrow \log s$ for $s \rightarrow \infty$. Thus the once-subtracted DR is given by

$$\Pi_T^{\gamma\gamma, \gamma Z}(q^2) = \frac{q^2 - q_0^2}{\pi} \int_{M^2}^{\infty} ds \frac{1}{s - q_0^2} \frac{\text{Im} \Pi_T^{\gamma\gamma, \gamma Z}(s)}{s - q^2 - i\epsilon}$$

where the UV divergence in $\Pi_T^{\gamma\gamma, \gamma Z}(q^2)$ is treated in the standard way

NNLO Methods: DRs

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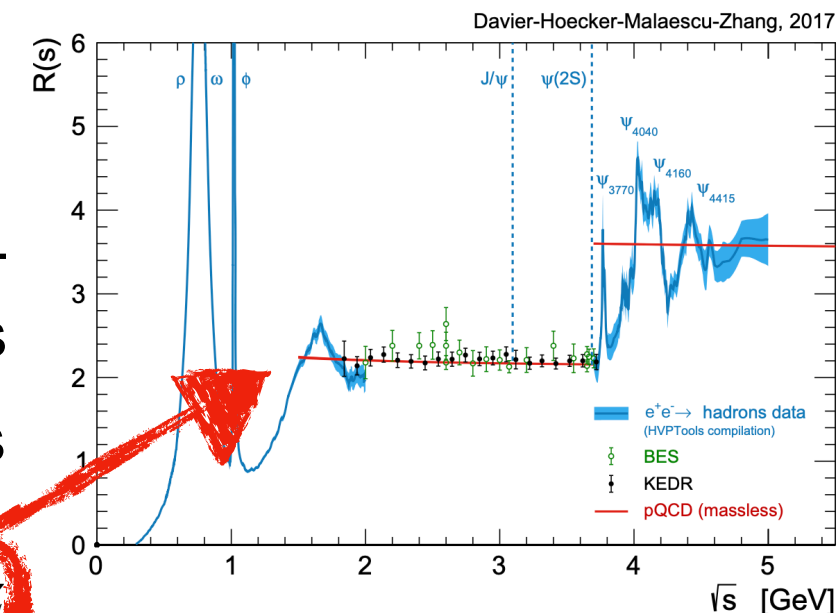
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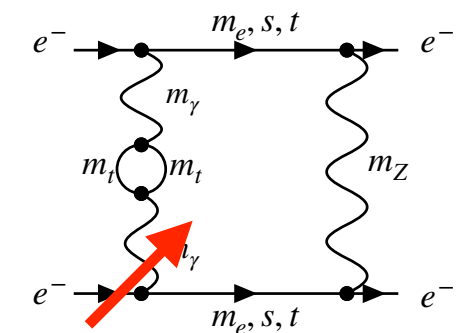
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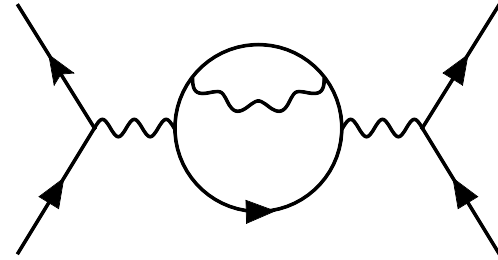
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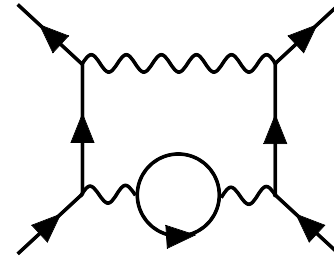


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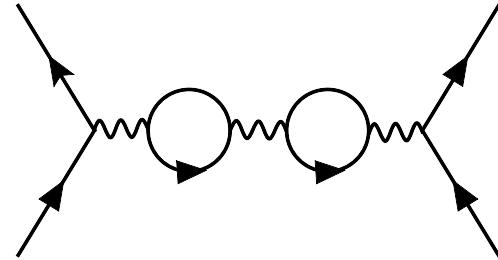
NNLO Results



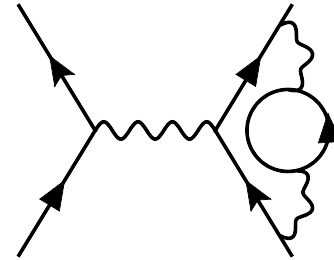
(a)



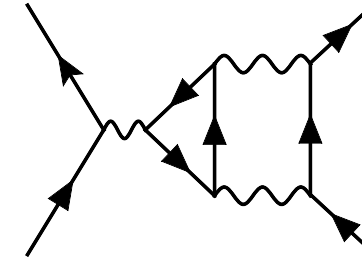
(b)



(c)



(d)



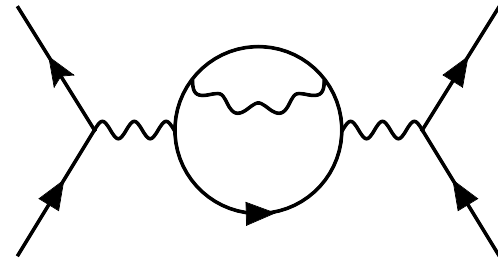
(e)

$$\begin{aligned}
 m_Z &= 91.1876 \text{ GeV}, & s_W^2 &= 0.2314, \\
 m_H &= 125.1 \text{ GeV}, & m_t &= 173.0 \text{ GeV}, \\
 m_\tau &= 1.777 \text{ GeV}, & m_b &= 3.99 \text{ GeV}, \\
 m_\mu &= 105.7 \text{ MeV}, & m_c &= 1.185 \text{ GeV}, \\
 m_e &= 0.511 \text{ MeV}, & m_s &= 0.342_{-0.053}^{+0.048} \text{ GeV}, \\
 m_{u,d} &= 0.246_{-0.057}^{+0.054} \text{ GeV}, \\
 \Delta\alpha &= 0.02761_{\text{had}} + 0.0314976_{\text{lep}},
 \end{aligned}$$

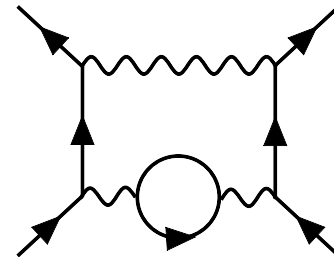
Quantity	Contribution ($\times 10^{-3}$)
1–4 $\sin^2\theta_W$	+74.4
$\Delta Q_{W(1,1)}^e$	–29.0
$\Delta Q_{W(1,0)}^e$	+3.1
$\Delta Q_{W(2,2)}^e$	–0.18 $_{-0.0040}^{+0.0024}$
$\Delta Q_{W(2,1)}^e$	+1.18 $_{-0.010}^{+0.015}$
$\Delta Q_{W(2,0)}^e$	± 0.13 (estimate)

YD, Freitas, Patel, Ramsey-Musolf, PRL 2021

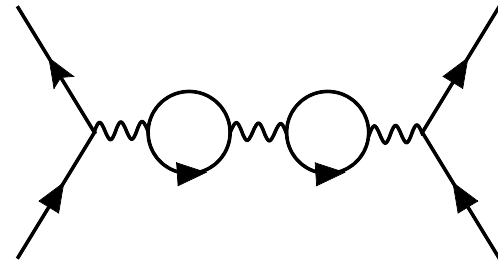
NNLO Results



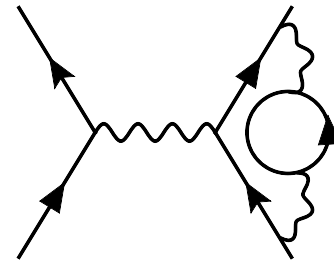
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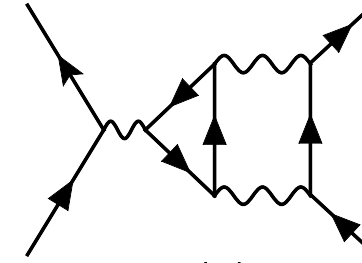
(b)



(c)



(d)



(e)

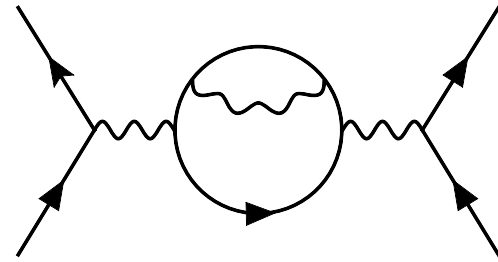
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$\Delta Q_{W(2,0)}^e$	± 0.13 (estimate)

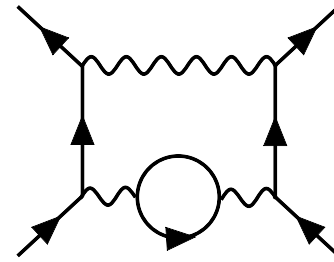
YD, Freitas, Patel, Ramsey-Musolf, PRL 2021

Q: Uncertainties?

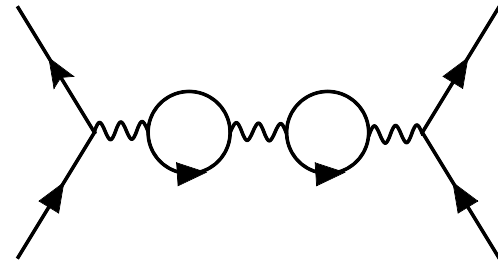
NNLO Results



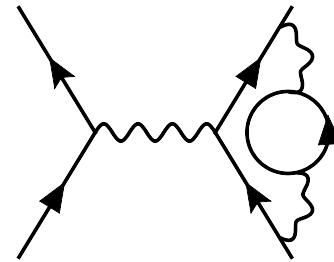
(a)



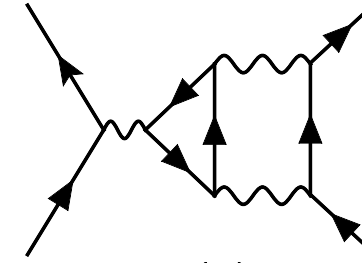
(b)



(c)



(d)



(e)

$$m_Z = 91.1876 \text{ GeV},$$

$$s_W^2 = 0.2314,$$

$$m_H = 125.1 \text{ GeV},$$

$$m_t = 173.0 \text{ GeV},$$

$$m_\tau = 1.777 \text{ GeV},$$

$$m_b = 3.99 \text{ GeV},$$

$$m_\mu = 105.7 \text{ MeV},$$

$$m_c = 1.185 \text{ GeV},$$

$$m_e = 0.511 \text{ MeV},$$

$$m_s = 0.342^{+0.048}_{-0.053} \text{ GeV},$$

$$m_{u,d} = 0.246^{+0.054}_{-0.057} \text{ GeV},$$

$$\Delta\alpha = 0.02761_{\text{had}} + 0.0314976_{\text{lep}},$$

Quantity	Contribution ($\times 10^{-3}$)
1–4 $\sin^2\theta_W$	+74.4
$\Delta Q_{W(1,1)}^e$	–29.0
$\Delta Q_{W(1,0)}^e$	+3.1
$\Delta Q_{W(2,2)}^e$	–0.18 $^{+0.0024}_{-0.0040}$
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YD, Freitas, Patel, Ramsey-Musolf, PRL 2021

Q: Uncertainties?

NNLO Results

Our input of $\sin^2 \theta_W(m_Z) = 0.2314$ only takes into account LO large logs like $\log m_f^2/m_Z^2$

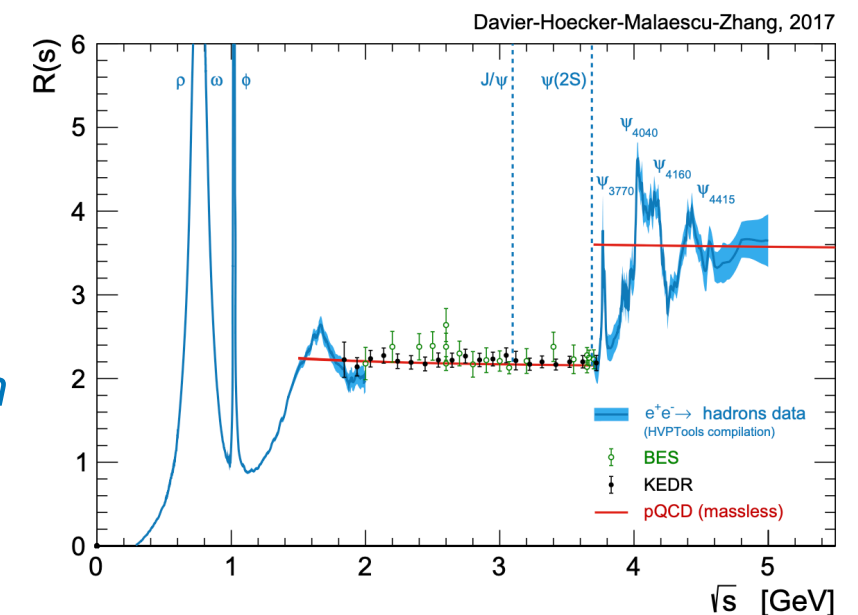
To resume all the large logs, one uses $\sin^2 \theta_W(0)$ through the RGE ($\overline{\text{MS}}$ scheme):

Jens & Michael, 2005

$$\hat{s}^2(\mu) = \hat{s}^2(\mu_0) + \frac{\hat{\alpha}(\mu_0)}{\pi} \left(\beta_0 \lambda_1 + \frac{\lambda_2}{3} - \beta_0 \hat{s}^2(\mu_0) \right) \ln \frac{\mu^2}{\mu_0^2} + \frac{\hat{\alpha}^2(\mu_0)}{\pi^2} [\dots]$$

Perturbative regime: straightforward computation

Non-Perturbative regime: included in $\hat{\alpha}_{\text{had}}$ from the dispersion relation



Reevaluating A_{PV} with $\sin^2 \theta_W(0) = 0.23861$ leads to an error of $\Delta Q_W^e = 0.08 \times 10^{-3}$

Erler, Ferro-Hernández, Freitas, 2023

NNLO Results

Parametric uncertainties in quark masses:

$$\begin{aligned} m_Z &= 91.1876 \text{ GeV}, & s_W^2 &= 0.2314, \\ m_H &= 125.1 \text{ GeV}, & m_t &= 173.0 \text{ GeV}, \\ m_\tau &= 1.777 \text{ GeV}, & m_b &= 3.99 \text{ GeV}, \\ m_\mu &= 105.7 \text{ MeV}, & m_c &= 1.185 \text{ GeV}, \\ m_e &= 0.511 \text{ MeV}, & m_s &= 0.342^{+0.048}_{-0.053} \text{ GeV}, \\ m_{u,d} &= 0.246^{+0.054}_{-0.057} \text{ GeV}, \\ \Delta\alpha &= 0.02761_{\text{had}} + 0.0314976_{\text{lep}}, \end{aligned}$$

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$\Delta Q_{W(2,2)}^e$	$-0.18^{+0.0024}_{-0.0040}$
$\Delta Q_{W(2,1)}^e$	$+1.18^{+0.015}_{-0.010} \approx 0.02$
$\Delta Q_{W(2,0)}^e$	± 0.13 (estimate)

This introduces another error of $\Delta Q_W^e = 0.03 \times 10^{-3}$ when evaluating A_{PV} at $\sin^2\theta_W(0)$. This is further enlarged by a factor 2 to be conservative

Erler, Ferro-Hernández, Freitas, 2023

NNLO Results

Parametric uncertainties in quark masses:

$m_Z = 91.1876 \text{ GeV}, \quad s_W^2 = 0.2314,$ $m_H = 125.1 \text{ GeV}, \quad m_t = 173.0 \text{ GeV},$ $m_\tau = 1.777 \text{ GeV}, \quad m_b = 3.99 \text{ GeV},$ $m_\mu = 105.7 \text{ MeV}, \quad m_c = 1.185 \text{ GeV},$ $m_e = 0.511 \text{ MeV}, \quad m_s = 0.342^{+0.048}_{-0.053} \text{ GeV},$ $m_{u,d} = 0.246^{+0.054}_{-0.057} \text{ GeV},$ $\Delta\alpha = 0.02761_{\text{had}} + 0.0314976_{\text{lep}},$

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This introduces another error of $\Delta Q_W^e = 0.03 \times 10^{-3}$ when evaluating A_{PV} at $\sin^2\theta_W(0)$. This is further enlarged by a factor 2 to be conservative — *could be improved by dispersive method and lattice*

Source	$\delta\sin^2\hat{\theta}_W(0) \times 10^5$
LQCD (Mainz)	2.3
pQCD	0.1
Condensates	0.2
Total	2.3

Erler, Ferro-Hernández, Freitas, 2023

Ce et al, 2203.08676

Erler, Ferro-Hernández, Kuberski, PRL, 2024

NNLO Results

Scheme dependence: *trading* $\alpha(0)$ with G_μ , one can estimate the order of NNNLO

$$\frac{G_\mu}{\sqrt{2}} = \frac{\pi\alpha}{2\hat{s}_Z^2\hat{c}_Z^2 m_Z^2} (1 + \Delta r)$$

Freitas, Hollik, Walter, Weiglein, 2002

Freitas, Hollik, Walter, Weiglein, 2003

Awramik, Czakon, 2003

This uncertainty is then computed by comparing A_{PV} obtained in these two schemes, leading to an error of $\Delta Q_W^e = 0.23 \times 10^{-3}$ (*dominant!*)

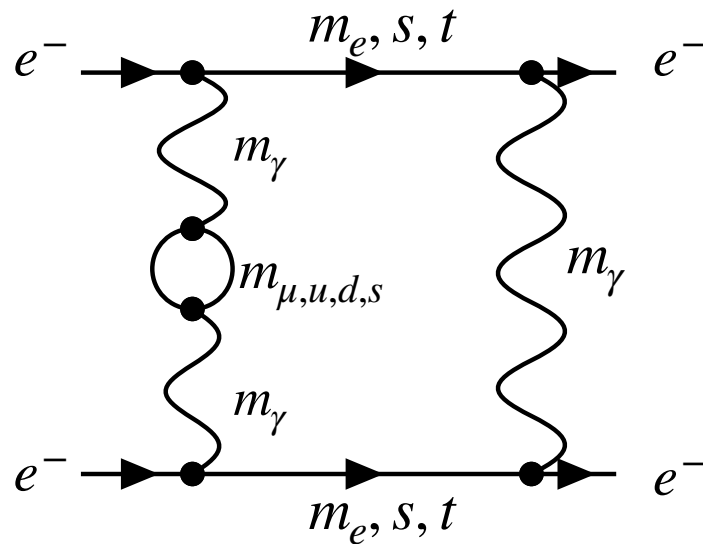
$$\begin{aligned} \Delta Q_W^e &= \left(\pm 0.08_{\hat{s}(0)} \pm 0.06_{\Delta Q_{W(2,1)}^{e,X}(\text{had})} \pm 0.13_{\Delta Q_{W(2,0)}^{e,X}(\text{missing})} \pm 0.23_{\text{scheme}} \right) \times 10^{-3} \\ &= \pm 0.28_{\text{theory}} \times 10^{-3} \end{aligned}$$

5 times better than the MOLLER precision goal

Erler, Ferro-Hernández, Freitas, 2023

Finite Q^2 effects?

Is the hierarchical expansion based on $m_f \ll s, t$ justified?



*In particular, here $|Q| \simeq 0.1 \text{ GeV} \simeq m_{\mu, u, d, s}$
such that $\frac{m_{\mu, u, d, s}^2}{s, t}$ is not small!*

$$A_{\text{PV}} = \frac{G_\mu Q^2}{\sqrt{2}\pi\alpha_{\text{EM}}} \frac{1-y}{1+y^4+(1-y)^4} [1 - 4\sin^2\theta_W + F_2(Q^2, y)]$$

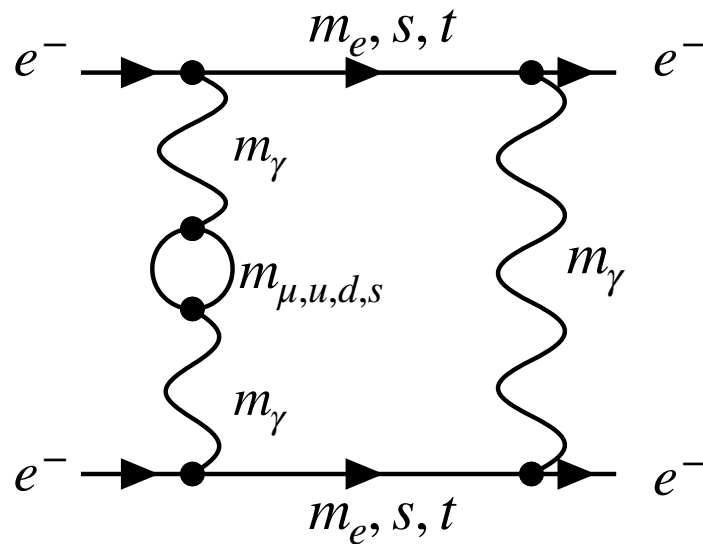
At NLO, accidental cancellation leads to

$$F_2(Q^2, 0.5) \approx 2 \times 10^{-5}$$

Czarnecki, Marciano, 1996

Finite Q^2 effects?

Is the hierarchical expansion based on $m_f \ll s, t$ justified?



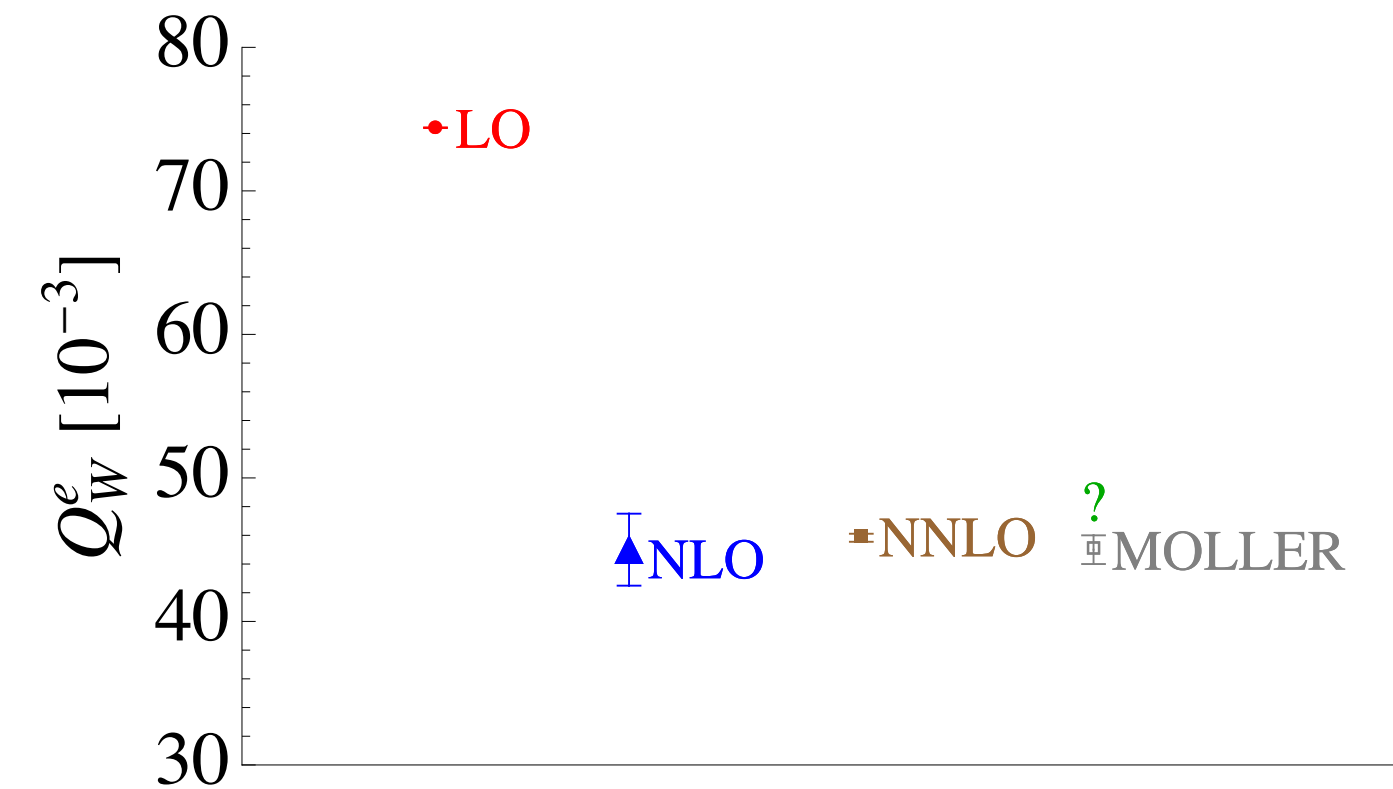
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y	$F_2^{(1,1)}$ [10^{-5}]	$F_2^{(2,2)}$ [10^{-5}]	$F_2^{(2,1)}$ [10^{-5}]
0.25 (0.75)	5.01	-0.54	0.00
0.30 (0.70)	4.11	-0.25	-0.18
0.35 (0.65)	3.39	-0.02	-0.29
0.40 (0.60)	2.86	0.15	-0.36
0.45 (0.55)	2.55	0.25	-0.40
0.50	2.44	0.29	-0.41

Erler, Ferro-Hernández, Freitas, 2023

Results summary

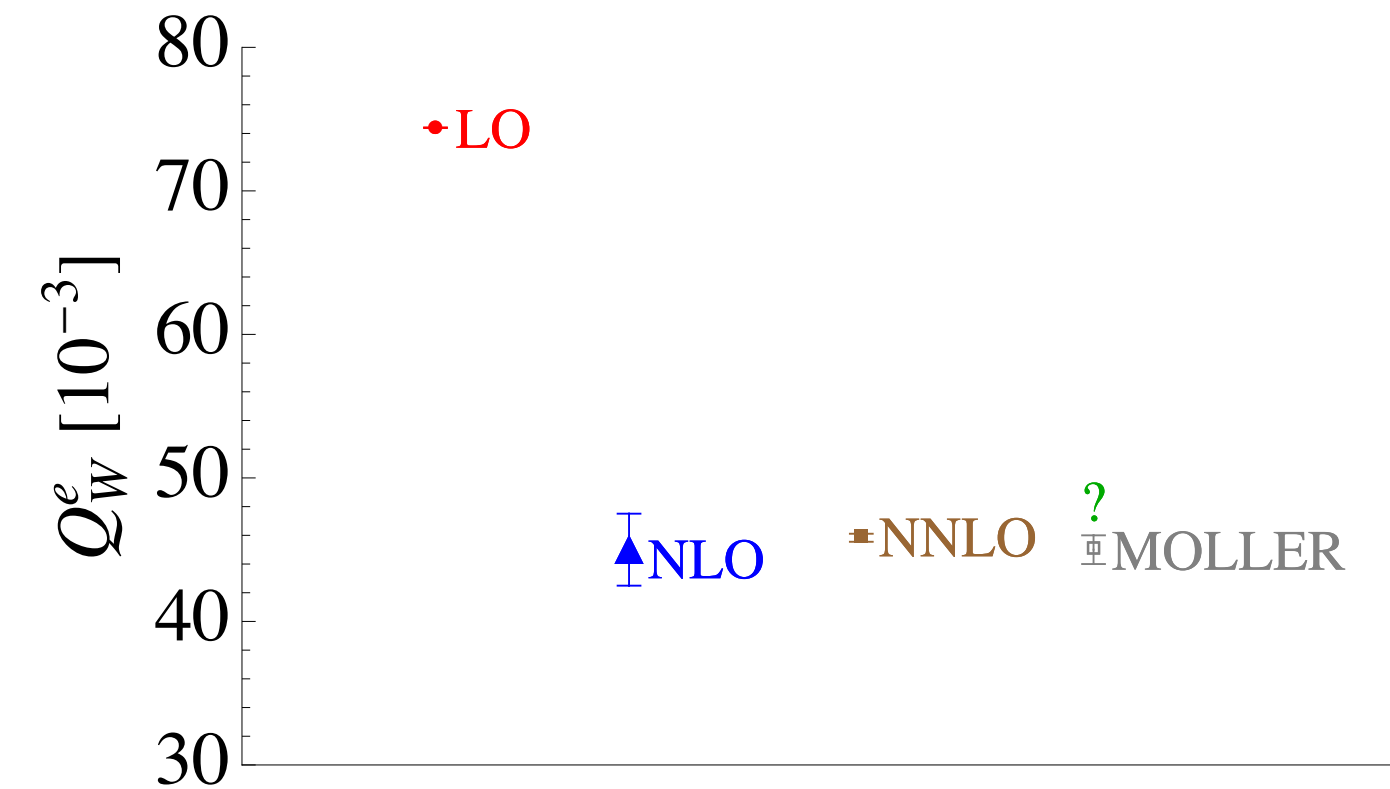


prediction stabilized

theoretical uncertainty under control

catalyzed approval of MOLLER from DOE

Results summary



prediction stabilized

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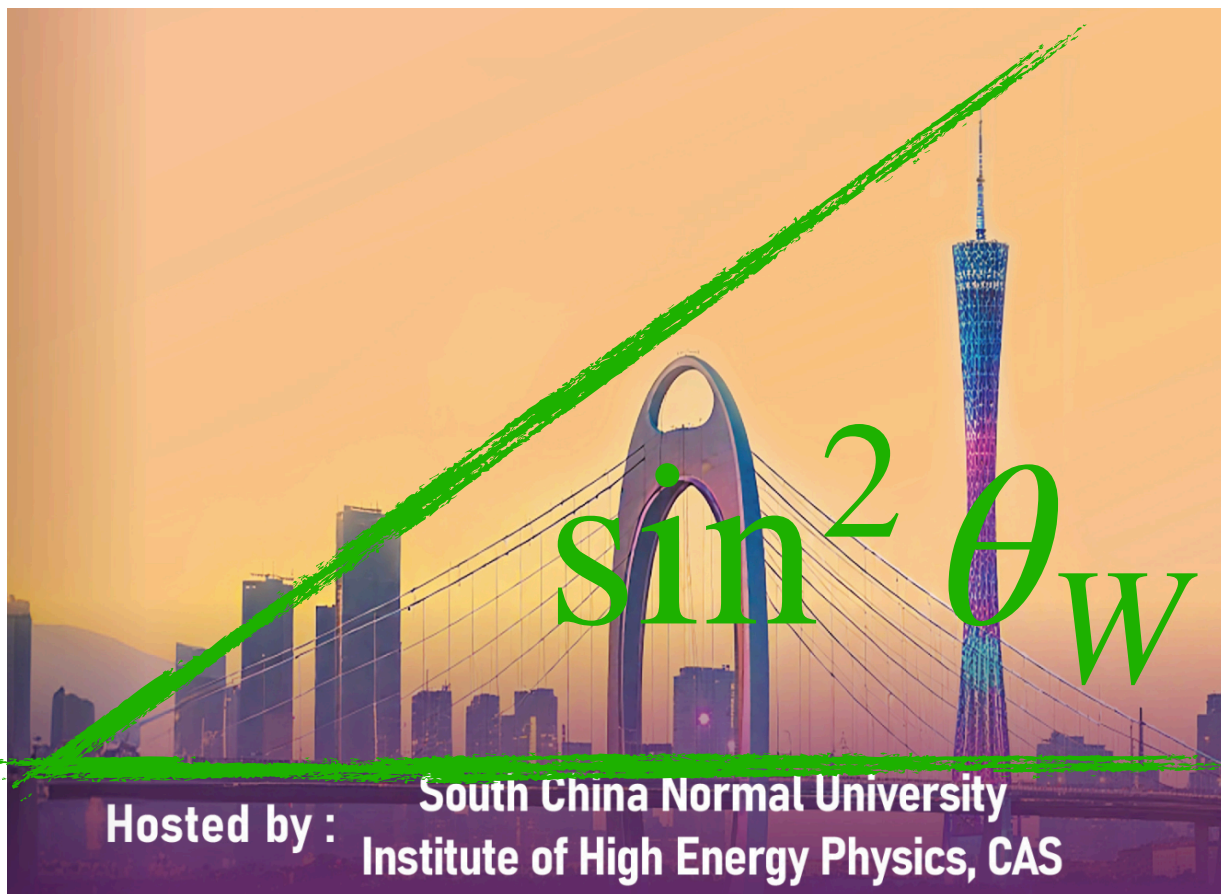
Now time for MOLLER to fit the central value of the electron weak charge

- Engineering Design complete: final technical design report is public
- Acquisitions ongoing and Construction has begun
- Installation: 2025-2026
- Commissioning: Early 2027
- Physics Analysis through 2029 and beyond

Huong Nguyen, SPIN2025

Summary

- ❖ MOLLER will measure $\sin^2 \theta_W$ at low scale with a precision comparable to $\sin^2 \theta_W(m_Z)$
 - ❖ *NNLO corrections with closed fermion loop: comparable to exp precision target*
 - ❖ *Theoretical uncertainties $\sim 1/5$ of exp precision target*
 - ❖ *Construction at JLab underway, data analysis expected in 2029*
- ❖ *Room for surprise from NNLO bosonic corrections? Ongoing efforts, stay tuned.*



Loop calculations are fun

CEPC is another machine besides the Canton Tower (+ more than measuring θ_W)

Postdoc hiring (EFT + precision) ongoing, check out [here](#) and [here](#). Very welcome to get in touch or please help spread the news

Backup