

One-Loop Renormalization of GraviChPT with Heat-Kernel Technique

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What is GraviChPT? Why Important?

- ChPT in Curved Spacetime^[1,2]
- Incorporate Also Interactions between Pion, Nucleon and Gravitational Field

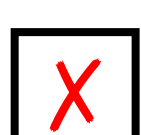
$$\mathcal{L}_{\pi N}^{(2)} = \frac{c_8}{8} R \bar{\Psi} \Psi + \frac{i c_9}{m} R^{\mu\nu} \left[\bar{\Psi} \gamma_\mu (\vec{\nabla}_\nu - \vec{\nabla}_\nu) \Psi \right]$$

$$\begin{aligned} \mathcal{L}_{\pi N}^{(3)} = & \tilde{d}_{g1} R \bar{\Psi} \gamma^\mu \gamma_5 u_\mu \Psi + \tilde{d}_{g2} R^{\mu\nu} \bar{\Psi} \gamma_\mu \gamma_5 u_\nu \Psi \\ & + \tilde{d}_{g3} R^{\alpha\beta\mu\nu} \left[\bar{\Psi} \sigma_{\alpha\beta} \gamma_5 u_\mu (\vec{\nabla}_\nu + \vec{\nabla}_\nu) \Psi \right] \\ & + \tilde{d}_{g4} \nabla_\nu R^{\alpha\beta\mu\nu} \left[\bar{\Psi} \sigma_{\alpha\beta} \gamma_5 (\vec{\nabla}_\mu - \vec{\nabla}_\mu) \Psi \right] \end{aligned}$$

- New LECs Present in Nucleon's GFFs Even in Flat Spacetime

$$A(t) = 1 - \frac{2c_9}{m_N} t + \dots \quad J(t) = \frac{1}{2} - \frac{c_9}{m_N} t + \dots \quad D(t) = c_8 m_N + \dots$$

How to Renormalize at One-Loop Level?



Feynman-Diagram Technique

$$\text{div}[W_{L=1}] = \text{div} \left[\sum_{\text{One-Loop 1PI}} \text{Feynman Diagram} \right]$$

Compute One-Loop 1PI Feynman Diagrams

Cumbersome



Heat-Kernel Technique

$$\text{div}[W_{L=1}] \propto \frac{1}{\epsilon} \int d^d x \sqrt{g} \text{str} [A^{(2)}(x, x)]$$

Compute Second-Order Heat-Kernel Expansion Coefficient

Algebraic, Systematic

- The left panel of the figure below illustrates the general workflow of the heat-kernel technique^[3,4].
- The right panel details the calculation of the heat-kernel expansion coefficients, for which we propose a novel iterative method in momentum space.

Step 1: Start with Action

$$S = \int d^4 x \sqrt{g} \mathcal{L}(\phi)$$

Step 2: Split Fields into Classical and Quantum Parts

$$\phi = \phi_{\text{cl}} + \xi, \quad \left. \frac{\delta S}{\delta \phi} \right|_{\phi_{\text{cl}}} = 0$$

Step 3: Perform Path Integral at One-Loop Level

$$e^{iW_{L=1}} = \int \mathcal{D}[\xi] e^{iS_{\text{bil}}}$$

$$S_{\text{bil}} = \int d^4 x \sqrt{g} \frac{1}{2} \xi^T \mathbf{M} \xi, \quad \mathbf{M} = -\frac{1}{\sqrt{g}} \partial_\mu \sqrt{g} (g^{\mu\nu} \mathbf{I} + \mathbf{V}^{\mu\nu}) \partial_\nu - \mathbf{X}^\mu \partial_\mu + \mathbf{Q}$$

Step 4: Apply Gaussian Quadrature Formula

$$W_{L=1} = \frac{i}{2} \ln \{ \text{Sdet}[\mathbf{M}] \} = \frac{i}{2} \text{Str} \{ \ln[\mathbf{M}] \}$$

Step 5: Apply Proper-Time Representation

$$W_{L=1} = -\frac{i}{2} \int_0^\infty \frac{d\tau}{\tau} \text{Str} [e^{-\tau \mathbf{M}}] = -\frac{i}{2} \int_0^\infty \frac{d\tau}{\tau} \int d^4 x \sqrt{g} \text{str} [\langle x | e^{-\tau \mathbf{M}} | x \rangle]$$

Step 6: Apply Heat Kernel Expansion

$$\langle x | e^{-\tau \mathbf{M}} | y \rangle \equiv \mathbf{K}(x, y | \tau) = \frac{1}{(4\pi\tau)^{d/2}} e^{-\frac{\sigma^2(x, y)}{4\tau}} \sum_{n=-\infty}^\infty \tau^n \mathbf{A}^{(n)}(x, y)$$

$$W_{L=1} = -\frac{i}{2(4\pi)^{d/2}} \int_0^\infty \frac{d\tau}{\tau^{d/2+1}} \int d^d x \sqrt{g} \sum_{n=-\infty}^\infty \tau^n \text{str} [\mathbf{A}^{(n)}(x, x)]$$

Step 7: Extract Divergent Part

$$\text{div}[W_{L=1}] = -\frac{i}{(4\pi)^{d/2}} \frac{4}{d(d-2)} \frac{\Gamma(2-d/2)}{\mu^{4-d}} \int d^d x \sqrt{g} \text{str} [\mathbf{A}^{(2)}(x, x)]$$

Step 6.1: Start with Heat Equation and Boundary Condition

$$\partial_\tau \mathbf{K}(x, y | \tau) = \{ (\eta^{\mu\nu} \mathbf{I} + \mathbf{V}^{\mu\nu}) \partial_\mu \partial_\nu + \mathbf{X}^\mu \partial_\mu - \mathbf{Q} \} \mathbf{K}(x, y | \tau)$$

$$\mathbf{K}(x, y | \tau = 0) = \delta^{(d)}(x - y) \mathbf{I}$$

Step 6.2: Heat Kernel Expansion in Coordinate Space

$$\mathbf{K}(x, y | \tau) = \frac{1}{(4\pi\tau)^{d/2}} e^{-\frac{(x-y)^2}{4\tau}} \sum_{n=-\infty}^\infty \tau^n \mathbf{A}^{(n)}(x, y)$$

Step 6.3: Fourier Transform to Momentum Space

$$\mathbf{K}(x, y | \tau) = \int \frac{d^d k}{(2\pi)^d} e^{ik \cdot (x-y)} e^{-\tau k^2} \sum_{r=0}^\infty \tau^r \mathbf{B}^{(r)}(k; x, y)$$

Step 6.4: Iterative Solution

$$\mathbf{B}^{(r)}(k; x, y) = \sum_{s=0}^{r+1} k_{\mu_1} \cdots k_{\mu_s} \mathbf{B}^{(r,s)\mu_1 \cdots \mu_s}(x, y) + \mathcal{O}(V^2)$$

Step 6.5: Perform Momentum Integral and Take Coincidence Limit

$$\lim_{x \rightarrow y} \int \frac{d^d k}{(2\pi)^d} e^{ik \cdot (x-y)} e^{-\tau k^2} k_{\mu_1} \cdots k_{\mu_s} = \begin{cases} 0 & \text{if } s \text{ is odd} \\ \eta_{\mu_1 \mu_2} \cdots \eta_{\mu_{s-1} \mu_s} + \cdots & \text{if } s \text{ is even} \end{cases}$$

Step 6.6: Match from Momentum to Coordinate Space

$$\mathbf{B}^{(r)}(k; x, y) \xrightarrow{\text{contribute to}} \begin{cases} \mathbf{A}^{(r - \text{floor}[r+1/2])}(x, x) \\ \vdots \\ \mathbf{A}^{(r)}(x, x) \end{cases}$$

$$\left. \begin{matrix} \mathbf{B}^{(2)}(k; x, y) \\ \vdots \\ \mathbf{B}^{(5)}(k; x, y) \end{matrix} \right\} \xrightarrow{\text{contribute to}} \mathbf{A}^{(2)}(x, x)$$

Preliminary Results

$$c_8 = c_8^R - \frac{15g_A^2 m}{2} \Lambda \quad c_9 = c_9^R + \frac{25g_A^2 m}{864F^2} \Lambda \quad \tilde{d}_{g1} = \tilde{d}_{g1}^R + \left(\frac{10g_A^3}{23F^2} - \frac{g_A}{3F^2} \right) \Lambda \quad \tilde{d}_{g2} = \tilde{d}_{g2}^R + \frac{g_A^3}{72F^2} \Lambda \quad \tilde{d}_{g3} = \tilde{d}_{g3}^R \quad \tilde{d}_{g4} = \tilde{d}_{g4}^R$$

$$\Lambda = \frac{1}{32\pi^2} \left(\frac{2}{d-4} - 1 + \gamma - \ln(4\pi) \right)$$

References

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