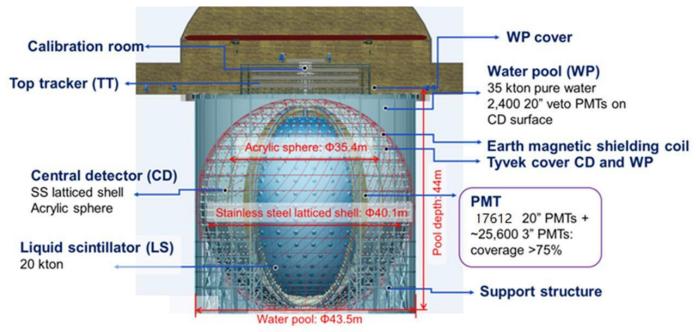


#### **JUNO Detector overview**

- Central detector:
  - Acrylic vessel with liquid scintillator
  - 17612 large PMTs (20-inch)
  - 25600 small PMTs (3-inch)
  - ~ 78% PMT coverage
  - PMTs in water buffer
- Veto detector for cosmic muon detection and background reduction.
  - Water Cherenkov Detector :
  - Top Tracker



# The veto system

#### Water Cherenkov Veto:

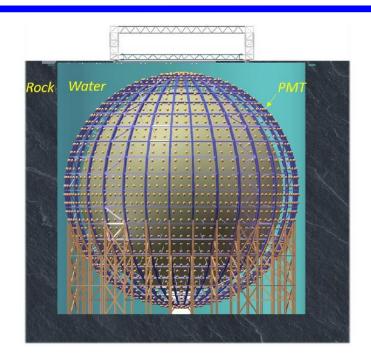
- Muon event tagging.
  - Outer of the detector;
  - 35 kton ultrapure water as medium;
  - Fast neutrons background rejection
    - Muon tagging+ passive shielding;
- Radioactivity from rock
  - Passive shielding by water

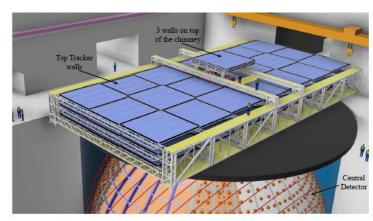
#### Compensation coils:

- Compensates Earth's magnetic field (~0.45 Gs), which otherwise reduces PMT efficiency by ~60%.
- Ensures uniform magnetic shielding;
- Tyvek reflector: increase light collection for PMT

#### Top tracker(detail see JP's talk):

- On top of water pool, cover ½ of pool;
- A precise muon track reconstruction;
- Cosmogenic muon induced isotopes reduction (<sup>9</sup>Li/<sup>8</sup>He and other isotopes).





### **Challenging for detector**

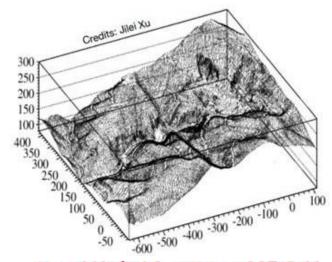
- Detector requirement
  - Large statistics
    - Large target mass;
    - Powerful nuclear power plants (NPPs)
  - Very good energy resolution
    - Very high PMT coverage + High transparency of LS+ High PMT efficiency
  - Cosmic muon induced background reduction
    - ~700 m rock overburden+ Veto system with >99% efficiency
  - Radioactivity background(reactor neutrinos, solar neutrinos)
    - Material background control + Installation procedure & clean environment control
  - Precise reference spectra of NPPs
    - Satellite detector → JUNO-TAO

Experiment	Daya Bay	Borexino	KamLAND	JUNO
Target mass [tons]	8 x 20	~300	~1,000	20,000
Photo electron[p.e./MeV]	~160	~500	~250	>1345
Energy resolution	~8.5%	~5%	~6%	~3%
Photocathode coverage	12%	34%	34%	~78%
Energy calibration uncertainty	0.5%	1%	2%	<1%

# Cosmogenic background

- Cosmic muons
  - ~700m rock overburden;
- Muon related background
  - 9Li/8He unstable isotopes produced by muon spallation on <sup>12</sup>C and decay beta-neutron;
    - ~127 <sup>9</sup>Li+40 <sup>8</sup>He isotope/day(IBD signal ~60/day);
    - Untagged muon induced fast neutron background.
  - Reduce the background to low level:
    - Good veto detector are required;
  - With the current veto strategy, the muoninduced background
    - <sup>9</sup>Li+ <sup>8</sup>He ->0.8/day;
    - Fast neutron->0.1/day

#### Digitalized mountain profile of JUNO site

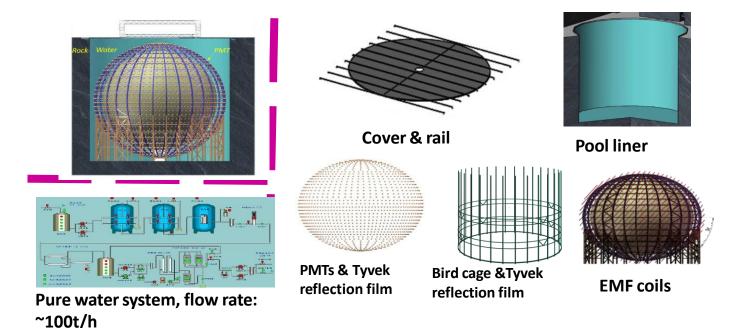


$$R_{\mu}$$
 = 4 Hz in LS,  $<$   $E_{\mu}$  $>$  = 207 GeV

Background	Rate $(day^{-1})$
Geoneutrinos	1.2
World reactors	1.0
Accidentals	0.8
$^9\mathrm{Li}/^8\mathrm{He}$	0.8
Atmospheric neutrinos	0.16
Fast neutrons	0.1
$^{13}{\rm C}(\alpha,{\rm n})^{16}{\rm O}$	0.05

### Water veto sub-systems/components

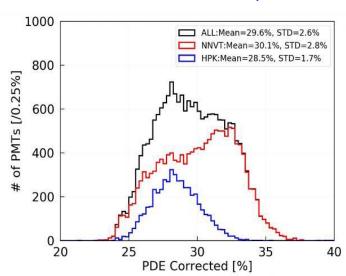
- Sub-systems/components
  - PMTs:2400 20-inch PMTs
  - **EMF coils**: shielding the detector to ensure the 20 inch PMT performance
  - Water system: 100 t/h water system
  - Pool lining: covering the pool wall as Rn barrier
  - Tyvek reflector: increase light collection for PMT
  - Cover: light/gas tight cover for the detector
  - Support structure



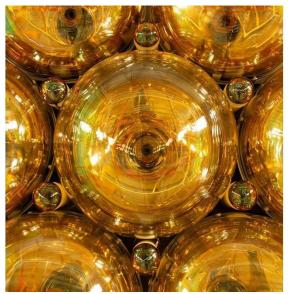
#### **PMTs**

- Three types of PMTs used in JUNO
  - Central detector
    - 17612 large PMTs (20-inch)
      - 12612 MCP-PMTs from NNVT
      - 5000 dynode PMTs from Hamamatsu
    - 25600 small PMTs (3-inch) from HZC
  - Water veto
    - 2400 MCP-PMTs from NNVT

#### **Photon Detection Efficiency**



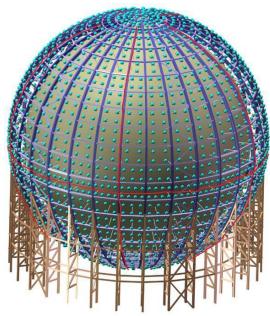




### PMT placement in water veto

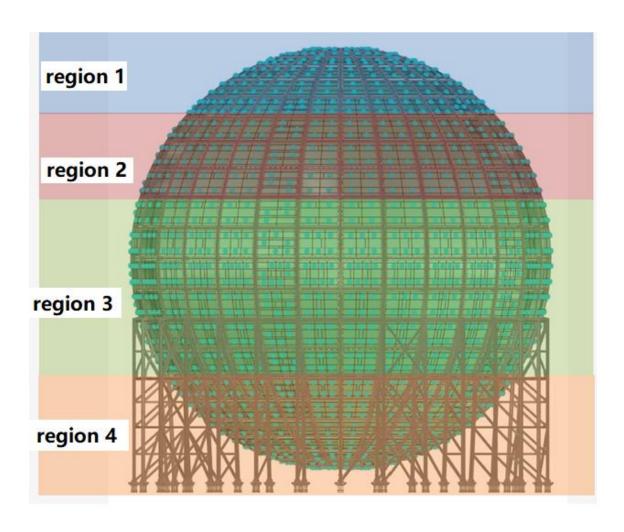
#### Optimized PMT placement

- Original design had some PMTs in wall/floor pointing inwards;
  - In this case, PMTs facing muon exiting CD in lower part of detector
  - Positions are too close to the compensation coils/outside the coils;
    - effected by the magnetic field by the EMF coil;
- Move PMTs on the sphere of the stainless
  - PMTs put on the surface of the sphere and facing outside to get better performance.
  - Trigger & efficiency
    - Divide the detector into 10 pieces for local trigger;
    - Detector efficiency is expected to reach >99.0%.
  - Fast neutron background
    - With the high muon tagging efficiency, the fast neutron background is anticipated to be <0.1/day.

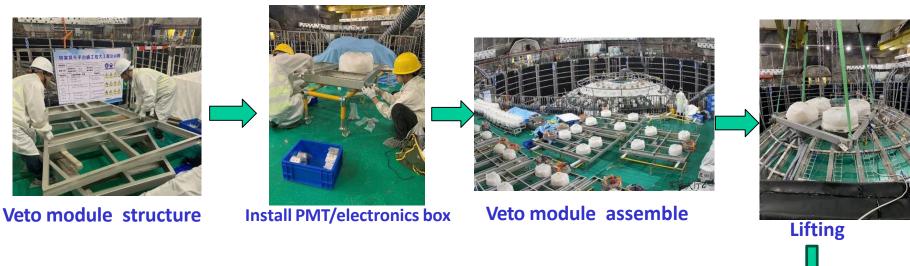


#### **Detector installation**

#### Total 4 regions for the whole detector installation

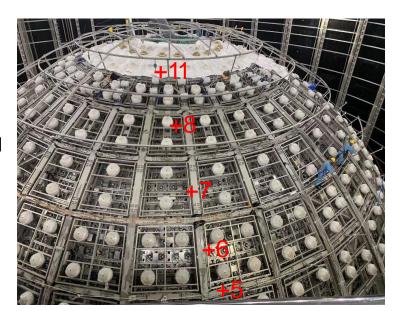


#### **Veto PMT/electronics installation(region 1, 2)**



Top half detector installation;

- PMT/module assemble in Exp. Hall.
- The module is mainly lifted by crane.

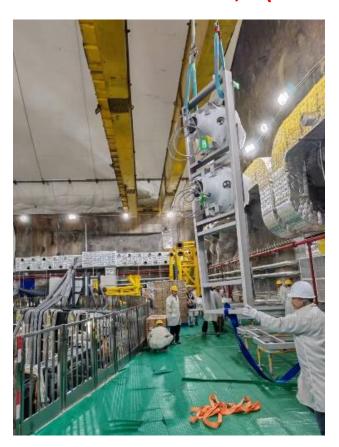




**Module fixation** 

### PMT module Installation(region 3)

- Lower hemisphere -3, -4(S3,S4) layer installation;
- Difficulty in installing slanted support window;
  - Large window changed to 2 small windows;
  - Small window PMT first, large window module, and then PMT installed.
    \$3,\$4(installation)





### The bottom half sphere installation(region 4)

#### Module







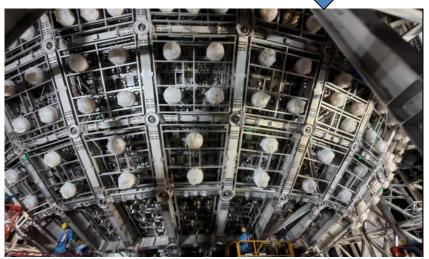
S5-S11 installation



**Install PMTs** 

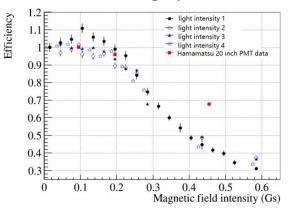
- Lower Hemisphere-5 -11 vetoPMT/Module installation:
- Initial Installation Options Scaffolding solution:
  - time consuming, inefficient installation.
- Final solution:
- Lift truck + winch:
  - Speed boosted installation of 20 modules/shift;
  - greatly improving installation efficiency and completing the task on time.

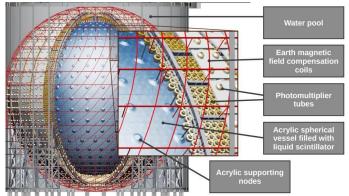




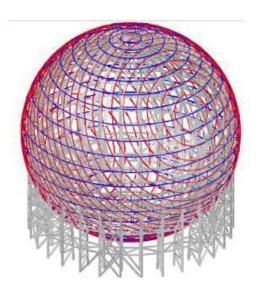
### **Compensation coils system**

- Earth Magnetic Field(EMF) intensity at JUNO site
  - Intensity ~0.45Gs
    - Big negative effect on the 20 inch PMT performance;
  - Need a shielding system for compensation EMF.

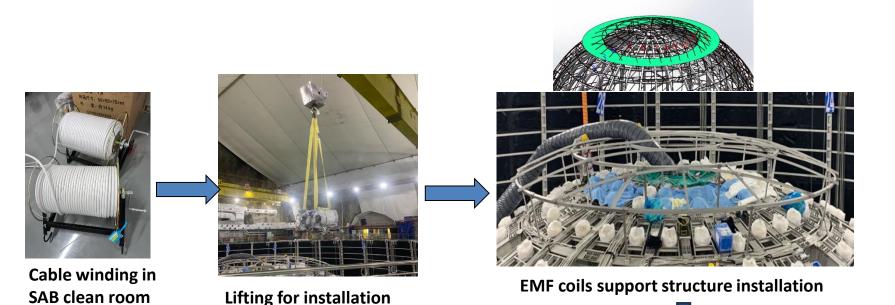




- Use one set of coils to generate the opposite direction of the geomagnetic field to compensate it.
  - 32 coils scheme;
  - Coils's uniformity in CD~0.05G.
- EMF direction change effect
  - The EMF direction change every year(<0.2deg/y).</p>
  - Set a compensation angle when the coils are installed.
  - Make the angle change < 1 degree within 10 years.</p>



# **EMF** shield coils installation(I)



#### EMF coils installation

- Winding coils is done in the SAB clean room, the length of the wire is numbered, and resistance is tested;
- Lifts spools of cable by crane for installation.



Coils fixation and spool placement

# **EMF** shield coils installation(II)

- EMF coils and support structure installation:
  - Top half sphere EMF colis structure installation and coil installation.
  - Mainly depend on crane;







- Bottom half sphere EMF colis structure/coils installation and coil installation.
  - More difficult with no crane;



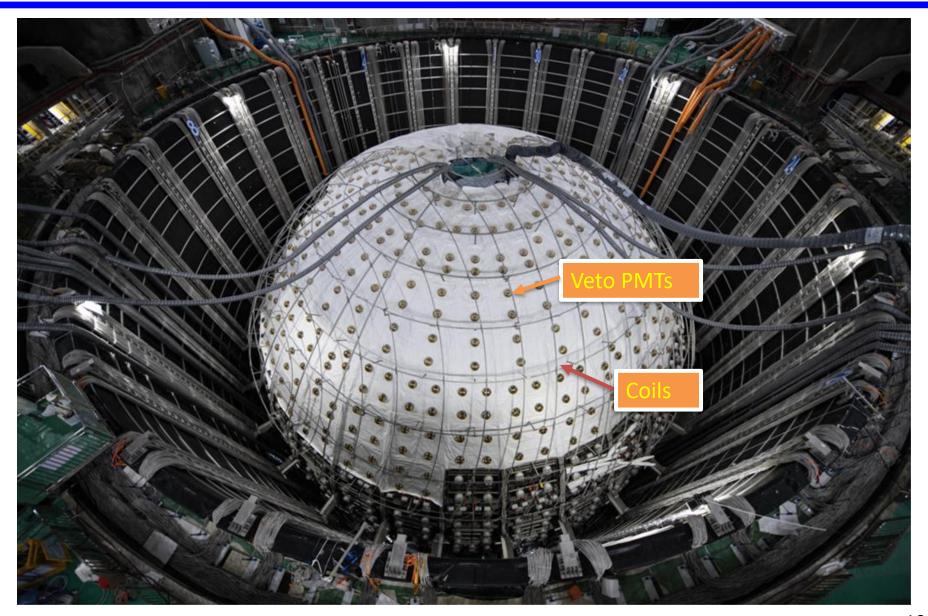
Put the coils on the right coils structure position



Fixed to the structure

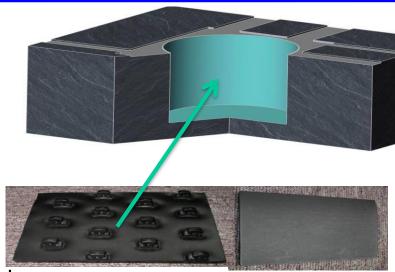


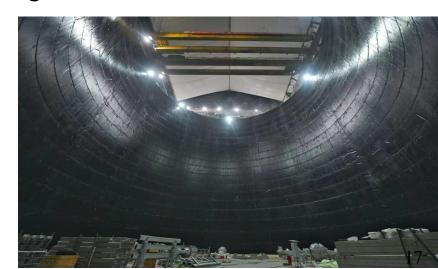
Final installation finished.



### **Pool lining**

- High Density Polyethylene (HDPE)
  - To separate pure water from the rock
  - To prevent rock radon from diffusing into the pool.
  - Two kinds of HDPE plate, with and without nails.
    - Thickness 5 mm.
  - The side wall lining installation is combine with civil wall construction.
  - The bottom HDPE film is installed at last stage.
  - Dimension:43.5 m diameter\*44m height;
  - >6000 m<sup>2</sup> lining.





### **Pool bottom HDPE lining installation**



**Spread the HDPE** 



Fixed on injection ring



Welding on the wall



Welding



**Cleaning** 



**Leakage detection** 

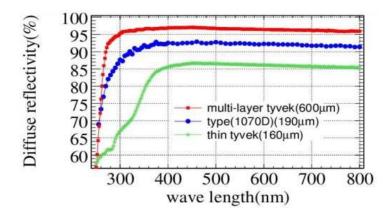
# **Tyvek reflector**

#### Tyvek reflection film

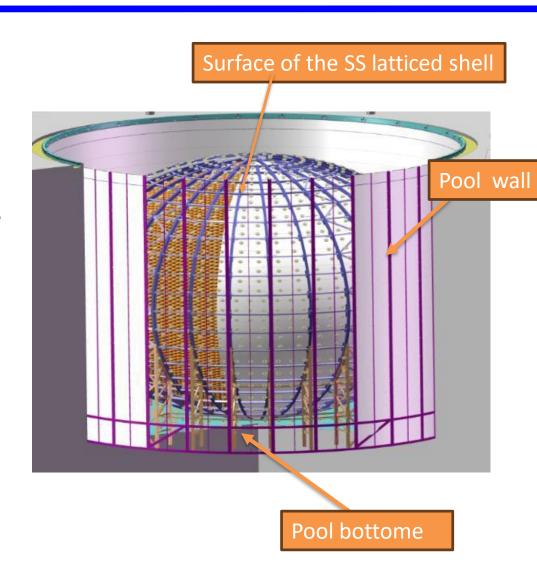
To be installed on

- Surface of the SS latticed shell;
- Pool wall, bottom;
- Cover the whole inner surface of the pool to improve light collection.





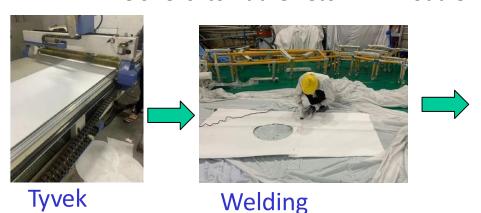
Reflectivity larger than 95% for wavelengths > 300nm

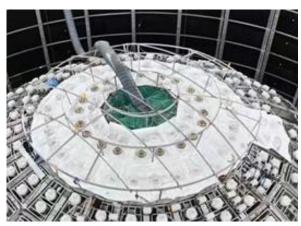


# **Tyvek installation**

#### CD SS structure surface Tyvek installation

- Installed Tyvek window by window;
- The size is to fit the veto PMT module.

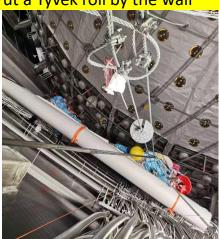




Installation

#### **Pool wall Tyvek installation**

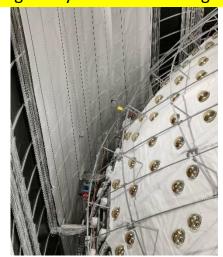
Put a Tyvek roll by the wall



Unroll the Tyvek along the wall



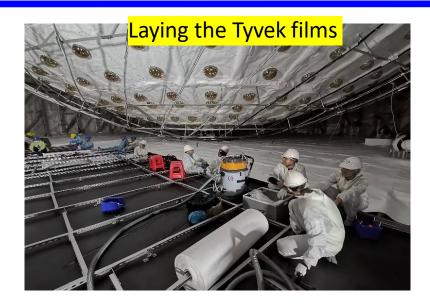
Fixing the Tyvek on the birdcage structure



### **Pool bottom Tyvek installation**

Assembling the Unistrut support structure





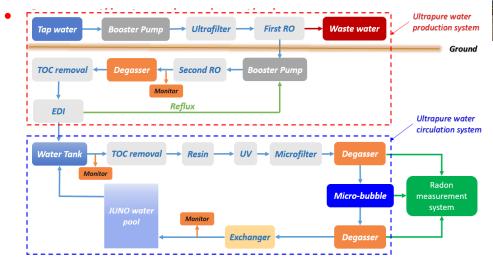




#### The ultrapure water production and circulation system

#### Water system

- Keep water quality with good transparency for detector performance
- Flow rate: ~100 t/hour
- Ground system:
  - Water production
- Underground system:
  - Purification and circulation;
- Connection ground and underground system by 1300 m stainless steel pipe in slope tunnel
- Water circulation
  - 2-3 weeks one volume.

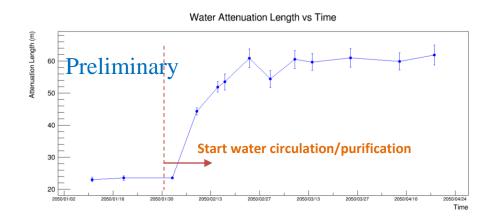


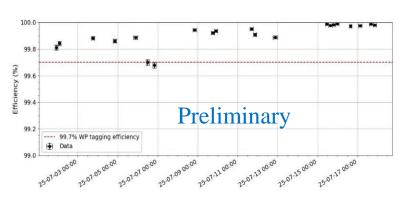




### **Detector operation and performance**

- Dec.18,2024,whole detector installation finished and detector start filling.
- Feb. 2025, the pool is full and start LS/water change and LS filing in CD.
- Water Cherenkov detector start commission and operation and tuning.
  - Threshold tunning, trigger rate, PMT HV, trigger window scan...

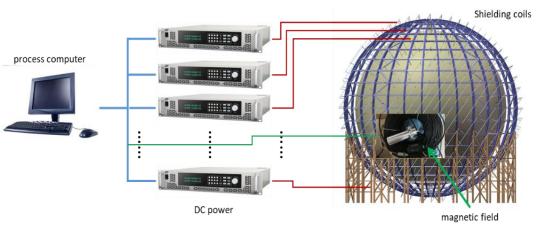




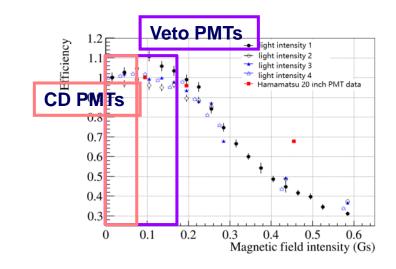
- The water attenuation length now can reach 60m;
  - Target 40m, better than anticipation.
- For CD muon detection efficiency >99.5%
- Now the detector is still under commissioning, we try to improve the WCD performance at the best condition.

### **Veto EMF shielding coils performance**





- 32 set of EMF coils, highly effective for 20" PMT detection efficiency increase
  - Residual EMF for CD PMT: max. < 0.1 Gauss (average ~0.05 Gauss); same as anticipation.
  - Residual EMF for Veto PMT: max. < 0.2 Gauss (average ~0.1 Gauss); same as anticipation.



monitoring sensor

# **Summary**

- JUNO will measure neutrino mass ordering, and a rich physics potential with supernova neutrinos, geo-neutrinos, solar neutrinos, atmospheric neutrinos and other oscillation physics such as searches for proton decay, etc.
- JUNO water Cherenkov detector is designed for muon detection and background reduction.
  - The detector assembly/installation is finished.
  - Now the detector is under commissioning and tunning.
  - The preliminary results shows that the water Cherenkov has good performance.

# Thanks!