

Strange Hadron Production in Au+Au Collisions from STAR Fixed-Target Experiment

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QPT-2025, Guilin, Guangxi, China

Oct., 24-28, 2025



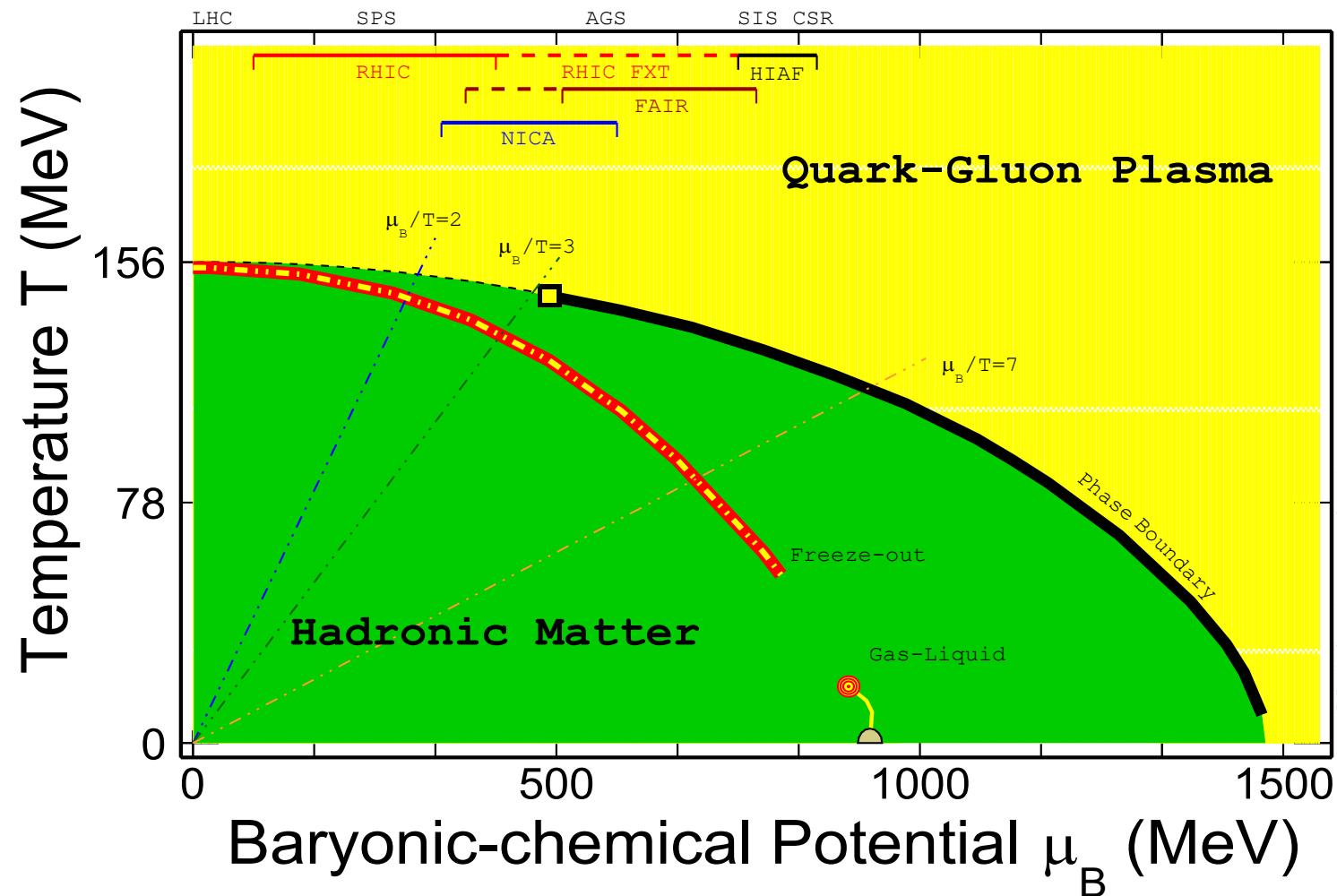
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Outline

- **Motivation**
- **Experimental Setup**
- **Measurement of Strange Hadron Production**
 - **Yield and yield ratio**
 - **Kinetic freeze-out**
- **Summary and Outlook**

Exploring the QCD Phase Diagram

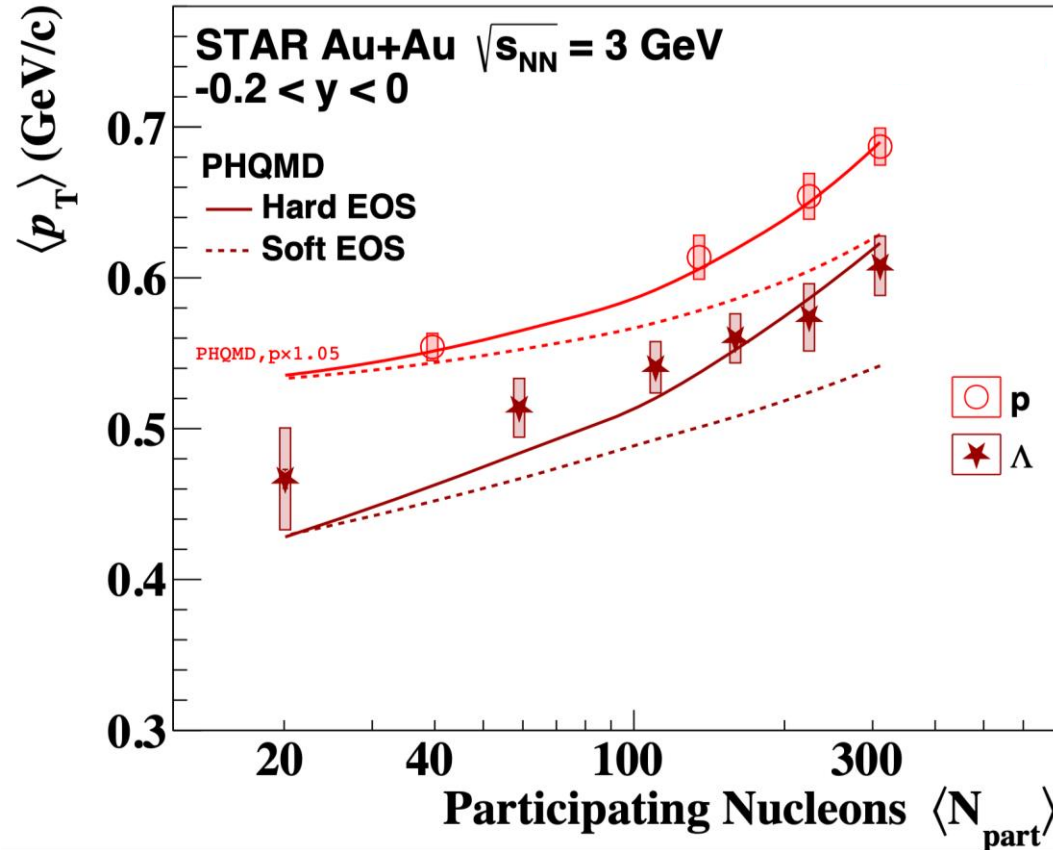


- Study the nature of QGP and QCD phase boundary using high energy heavy-ion collisions
 - At small μ_B , LQCD predicts smooth crossover phase transition
 - At large μ_B , QCD effective models predict 1st order phase transition
 - QCD critical point ?
- Experimental exploration of QCD phase diagram in high μ_B region

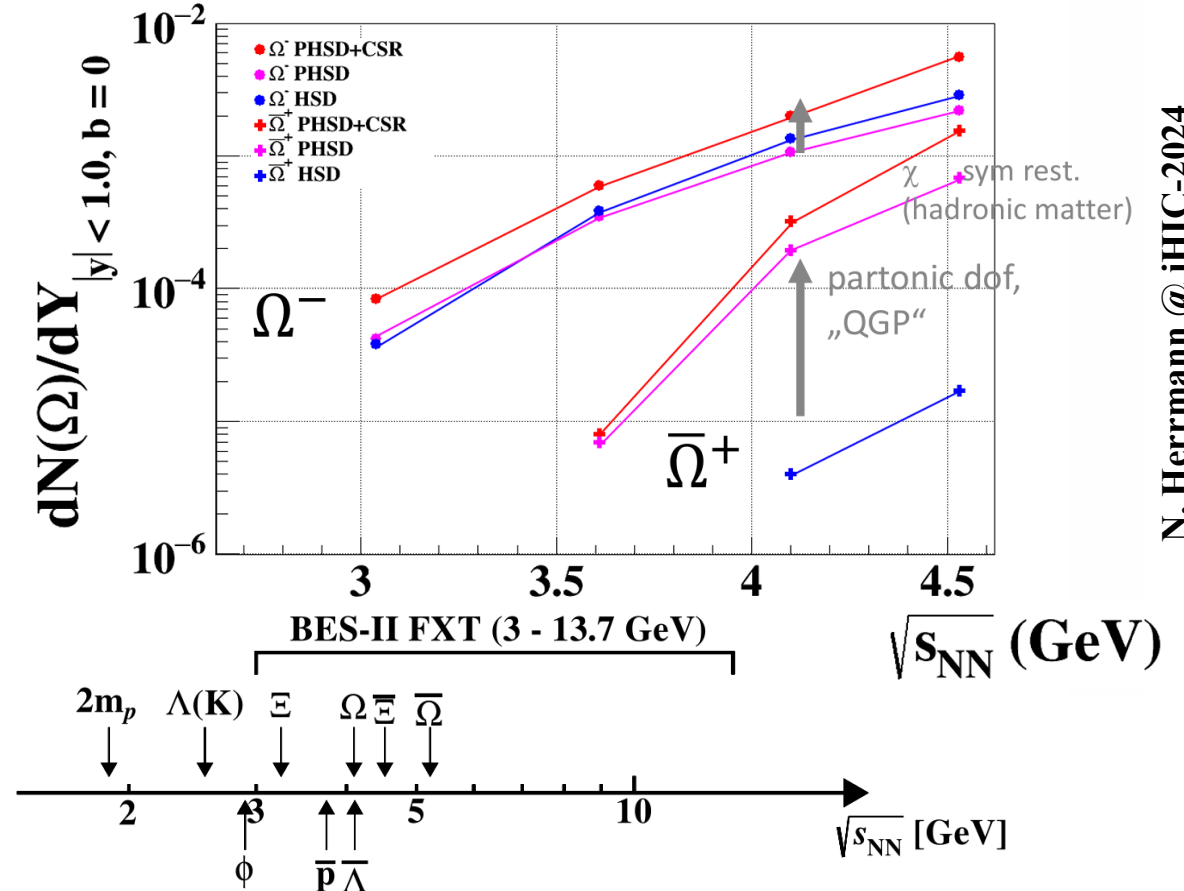
N. Xu @ SQM2022

Strange Hadron Production at Near/sub-threshold Energies

STAR Collaboration. Phys. Rev. C 110, 054911 (2024);
JHEP 2024, 139 (2024)
PHQMD: J. Aichelin, et.al. Phys. Rev. C 101, 044905 (2020)



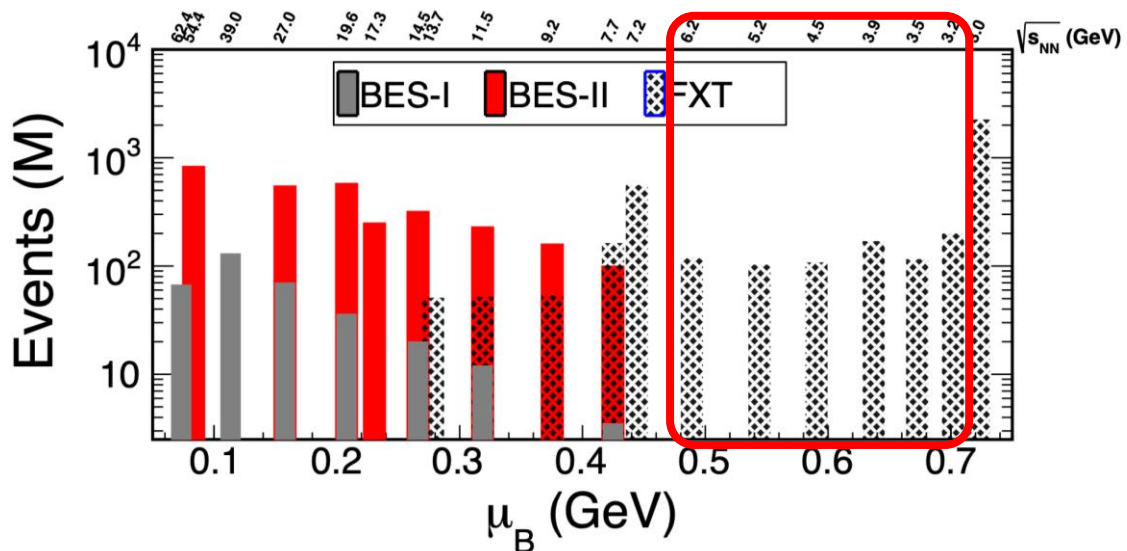
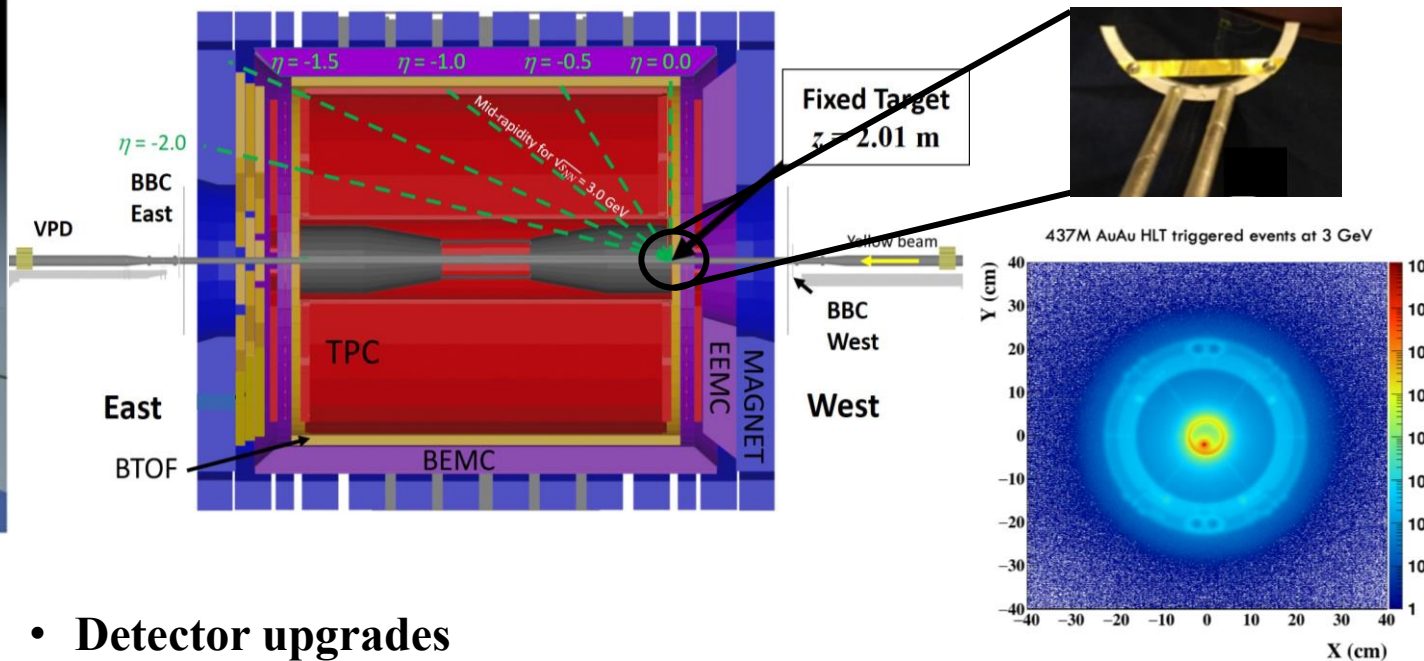
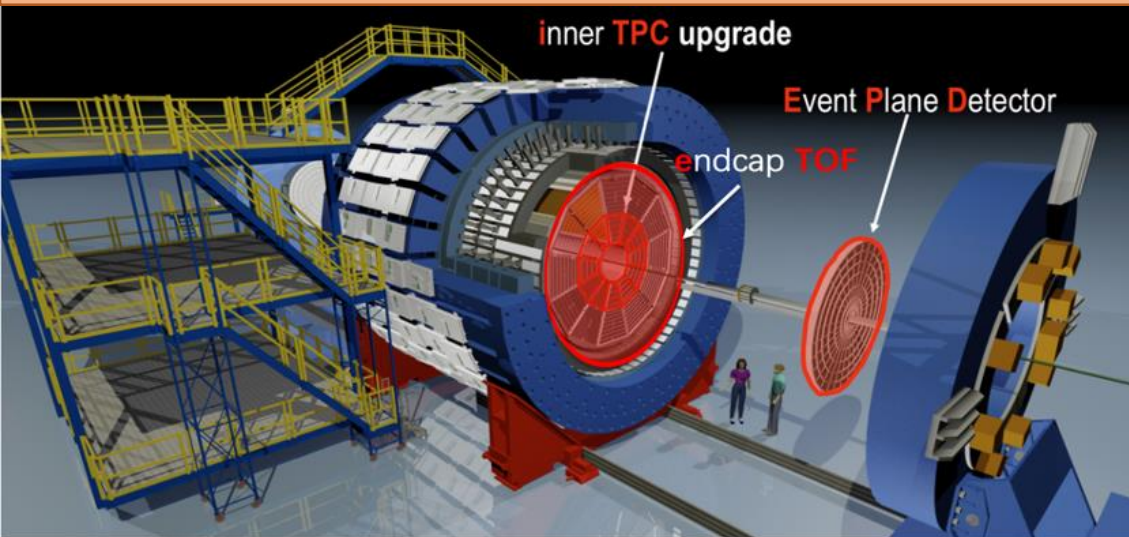
Y.J. Zhou @ CPOD-2024



N. Herrmann @ iHIC-2024

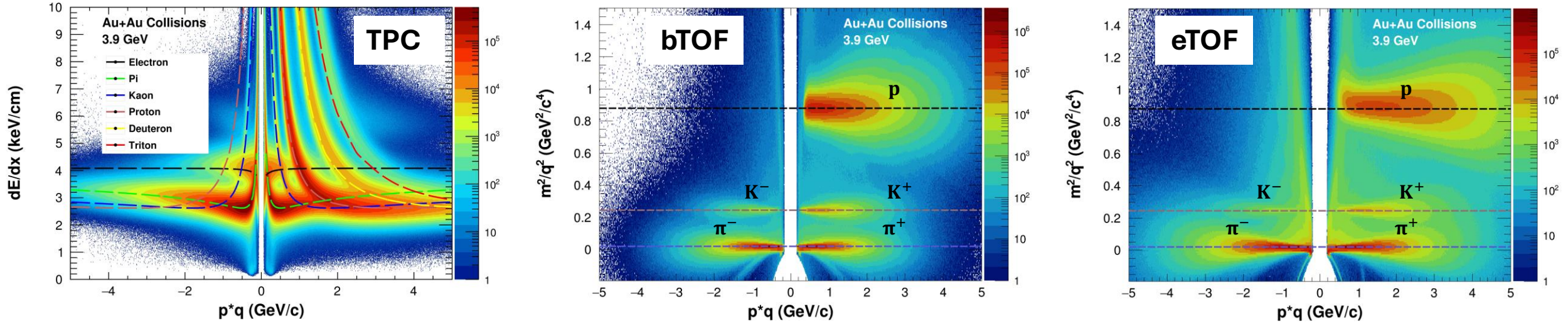
- Strange hadron production at high baryon density is a good probe to study medium properties
- The STAR BES-II FXT experiment provides unique opportunity to study strange hadron production at near or sub-threshold energies

STAR Detector and Fixed Target Setup



- **Detector upgrades**
 - Time Projection Chamber (TPC & inner TPC: $-2.4 < \eta < 0$)
 - Time-of-Flight
 - Barrel TOF: $-1.45 < \eta < 0$ & end-cap TOF: $-2.15 < \eta < -1.55$
- **Fixed target mode (Au + Au collisions at $\sqrt{s_{NN}} = 3.0 - 13.7$ GeV)**
 - 10× statistics compared to BES-I
 - **This talk: $\sqrt{s_{NN}} = 3.2 - 6.2$ GeV**

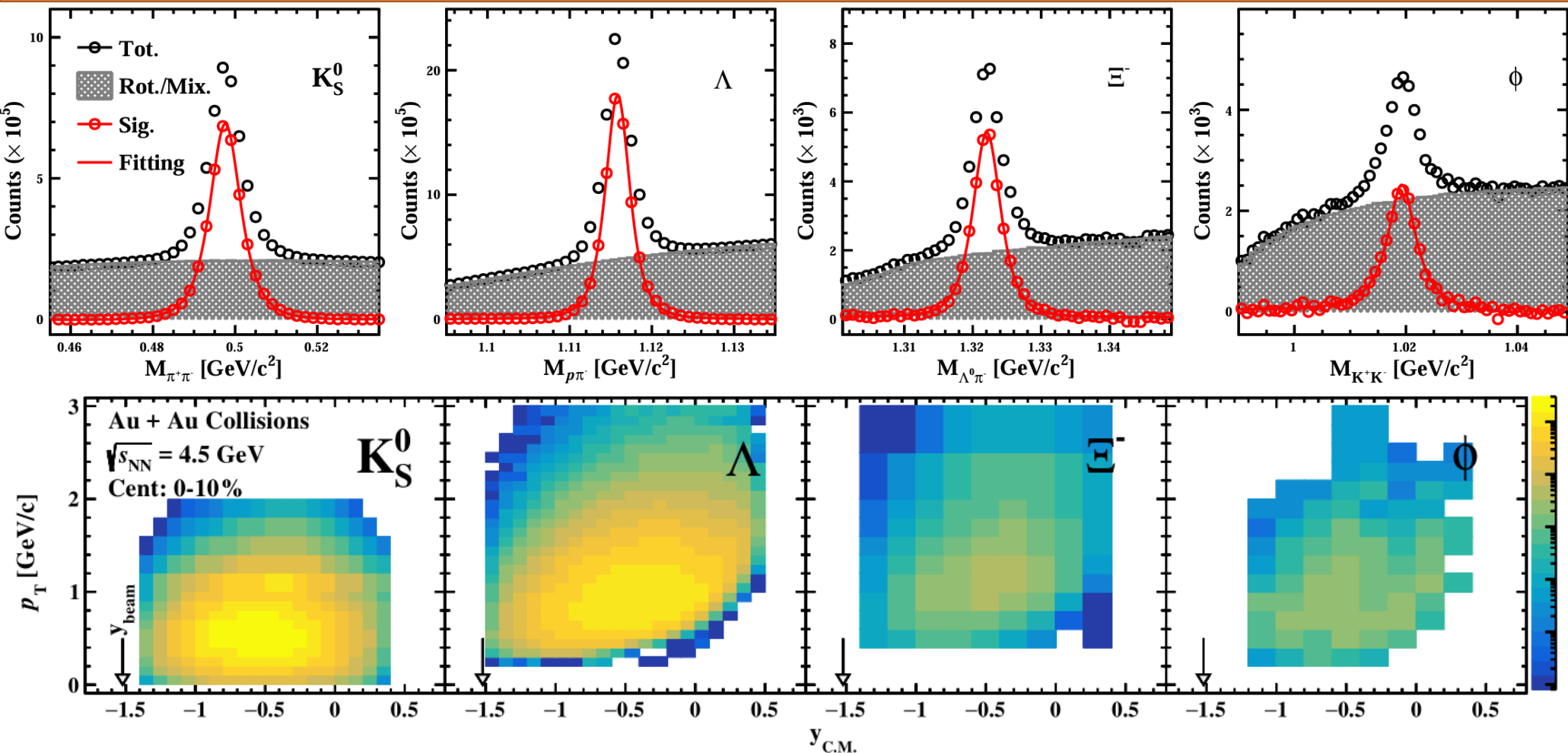
Particle Identification



- TPC (dE/dx) and TOF (β) for charged pion, kaon and proton identification
 - TOF $m^2 = p^2 \left(\frac{1}{\beta^2} - 1 \right)$
- Invariant mass method is used to reconstruct decay strange hadrons (K_S^0 , ϕ , Λ and Ξ^-) from their decay daughters

$$\text{➤ } m_{inv.} = \sqrt{(E_{dau,1} + E_{dau,2})^2 - (\vec{p}_{dau,1} + \vec{p}_{dau,2})^2}$$

Signal Extraction



Decay Channel:

$K_S^0 \rightarrow \pi^+ + \pi^-$

$\phi \rightarrow K^+ + K^-$

$\Lambda \rightarrow p + \pi^-$

$\Xi^- \rightarrow \Lambda + \pi^-$

Fitting function:

Sig: Student-t for K_S^0 , Λ , Ξ^-

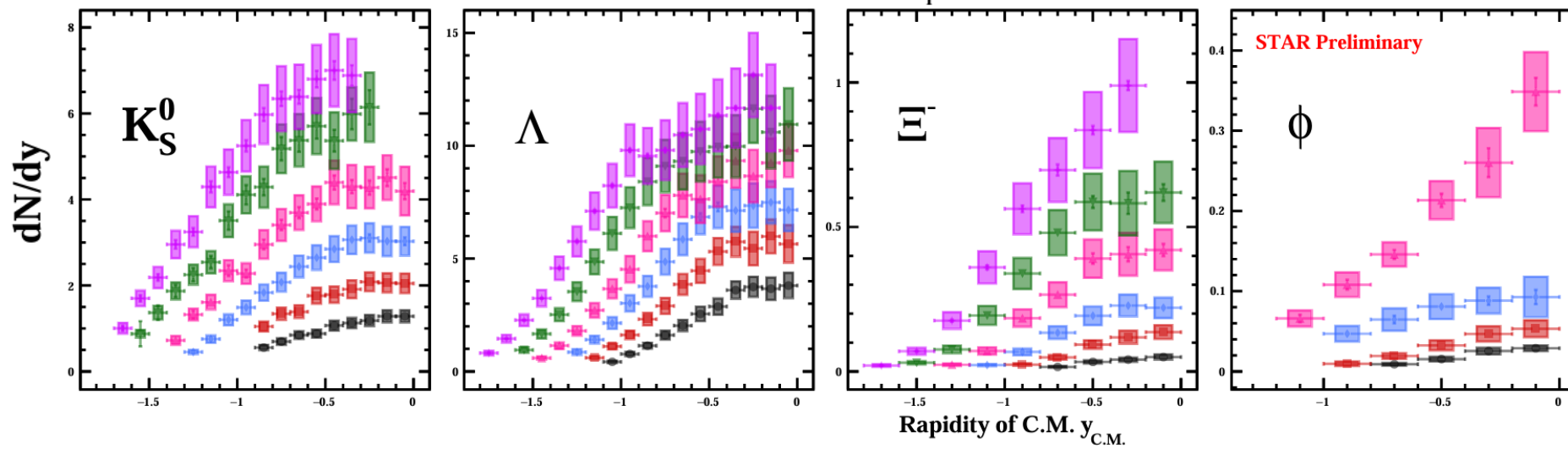
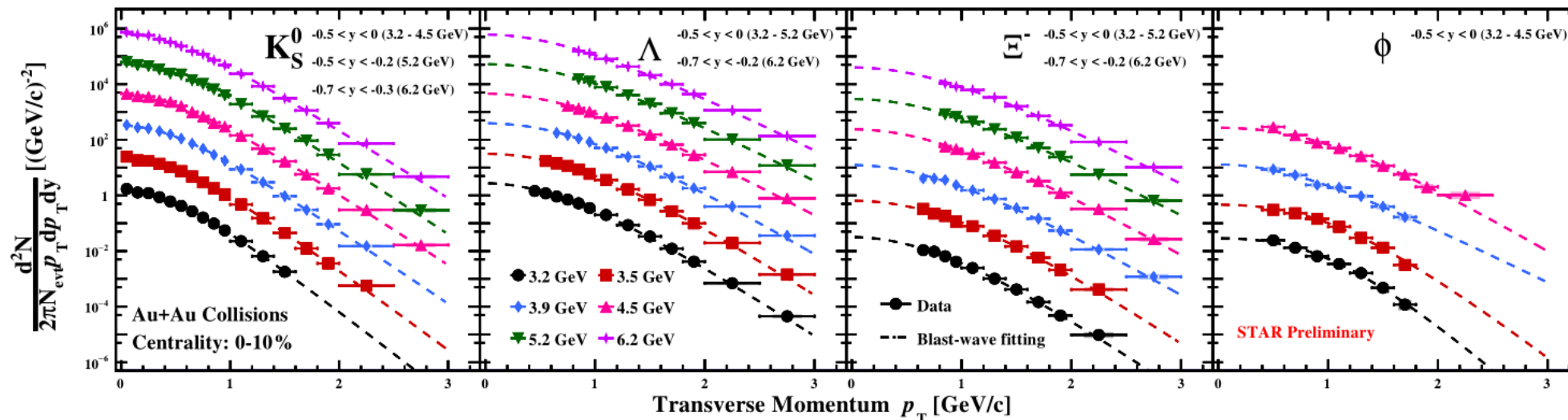
Breit-Wigner for ϕ

Bkg: Polynomial

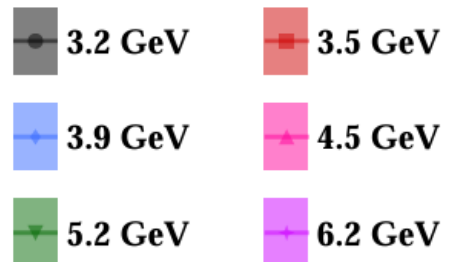
$\sqrt{s_{NN}}$	y_{beam}
3.2	1.14
3.5	1.25
3.9	1.38
4.5	1.52
5.2	1.68
6.2	1.87

- Combinatorial background is reconstructed by mixed-event or track rotation method
- Good coverage from beam-rapidity to mid-rapidity for K_S^0 , Λ , Ξ^- and ϕ

p_T Spectra and Rapidity Density Distributions



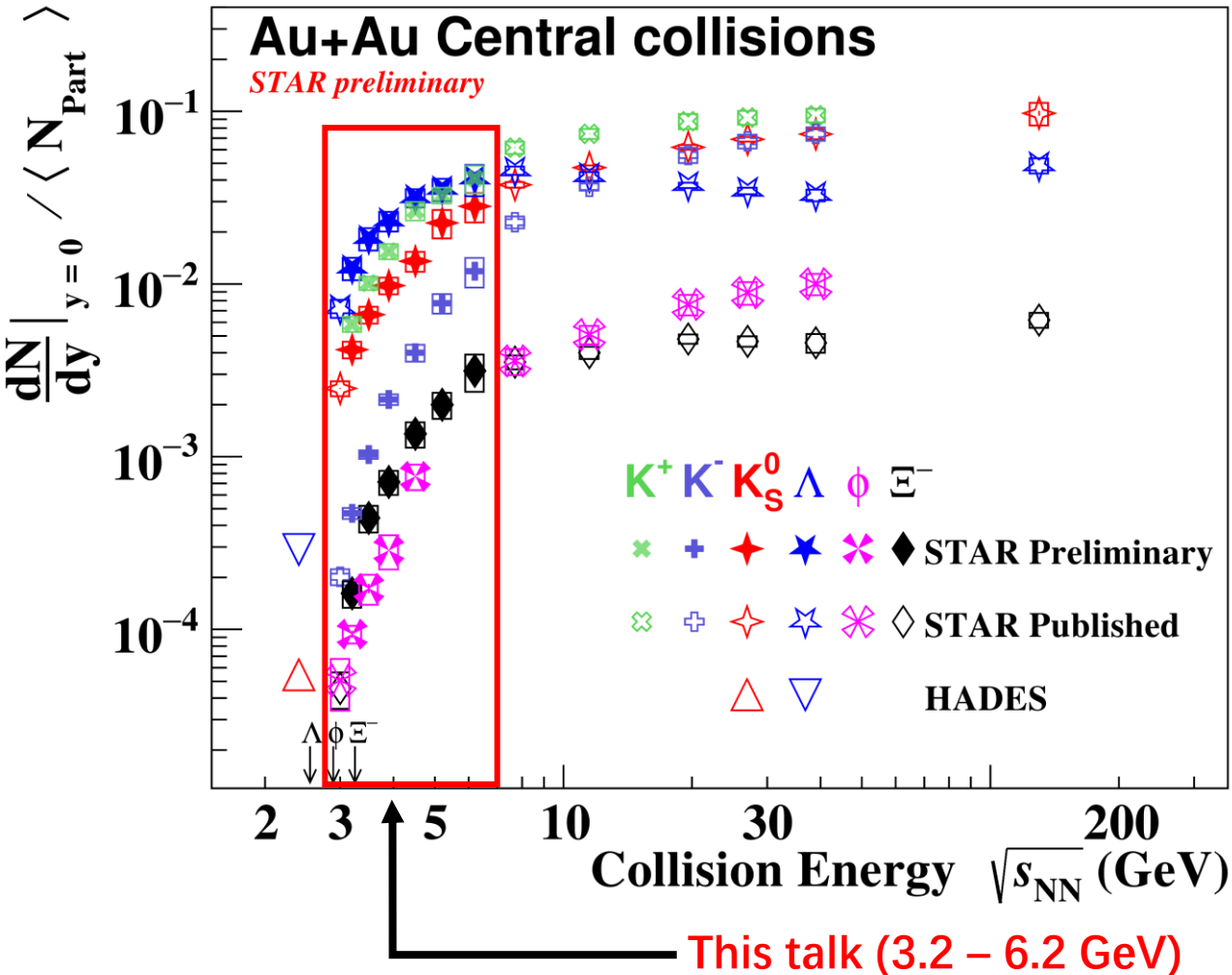
Au + Au Collision @ 0-10%



- Comprehensive measurements for strange hadron production at STAR FXT energies

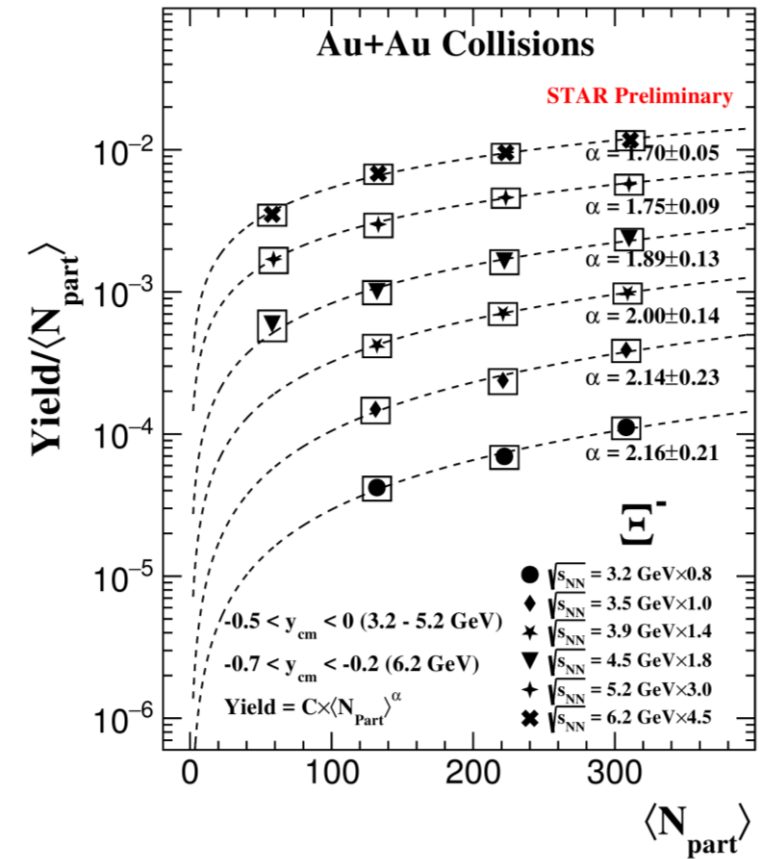
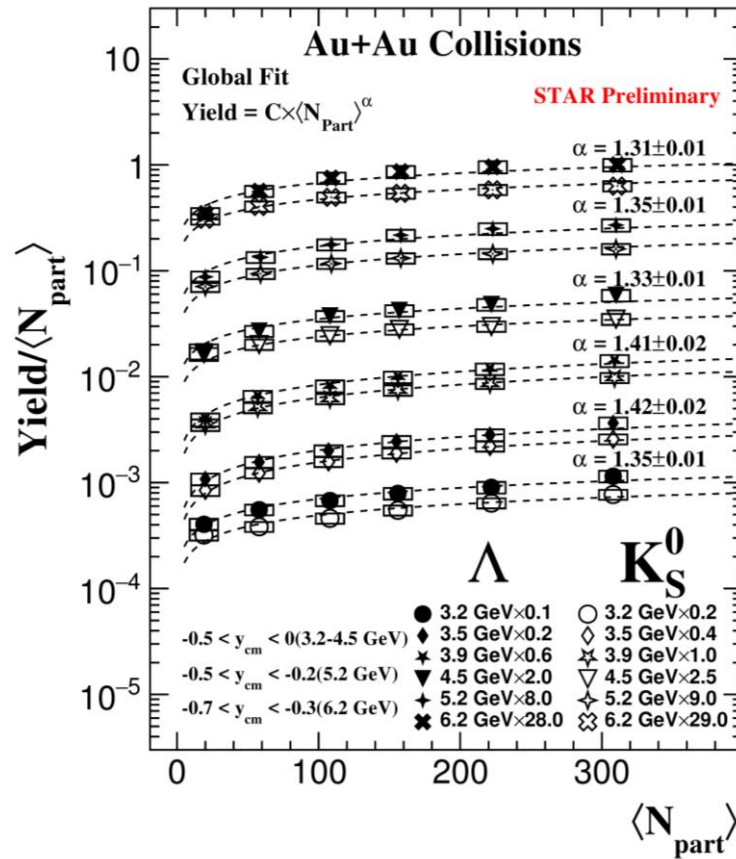
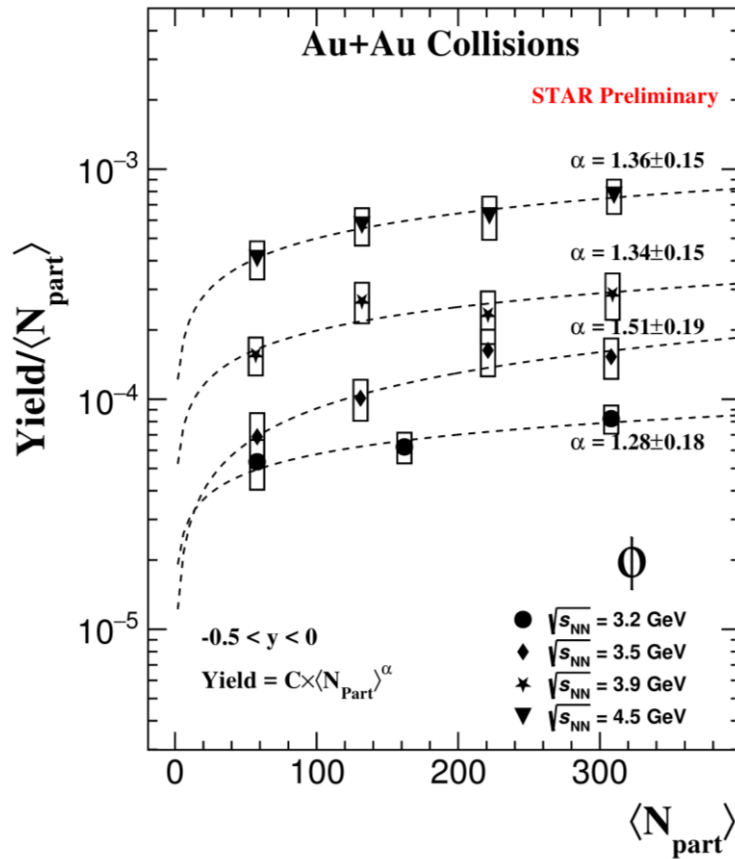
Energy Dependence of Mid-Rapidity Yields

STAR Collaboration. Phys. Rev. C 96 (2017) 044904, 2017; Phys. Rev. C 102, 034909 (2020);
Phys. Lett. B 831, 137152 (2022); JHEP 2024, 139 (2024)
HADES Collaboration. Phys. Lett. B 793, 457 (2019)



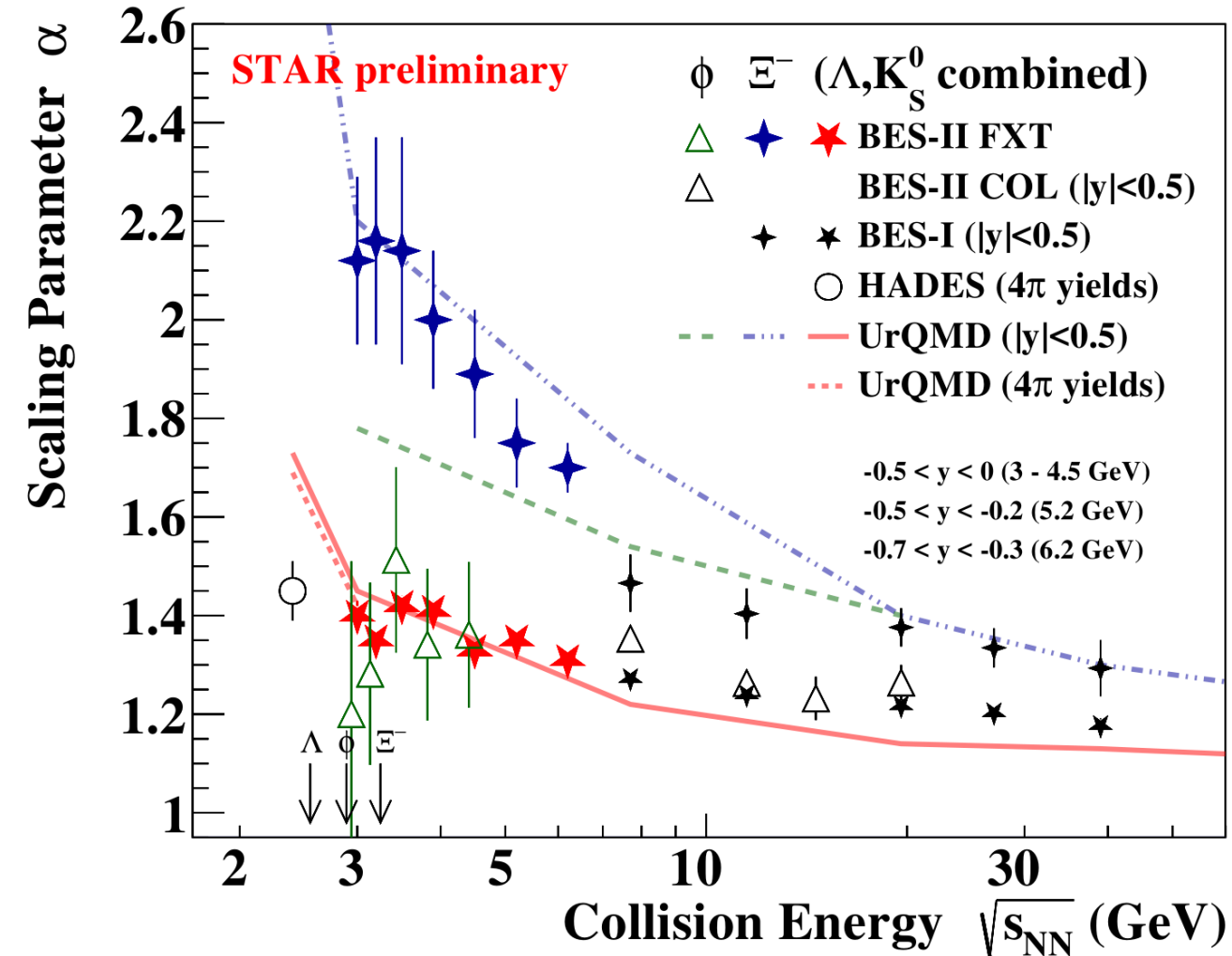
- Mid-rapidity yields increases rapidly at low energy and approximately saturate at high energy
 - Partonic interaction (pair production)
 $gg \rightarrow s\bar{s}$ or $q\bar{q} \rightarrow s\bar{s}$
 - Hadronic interaction (associated production)
 $BB \rightarrow BYK$ or $BB \rightarrow BEKK$
B: N, p, Δ , etc. Y: Λ , Σ , etc. K: K^+ , K^0
- First measurement of Ξ^- at sub-threshold energies in Au+Au collisions
- Λ yields exceed those of K_S^0 below $\sqrt{s_{\text{NN}}} \sim 8$ GeV
 - Due to higher baryon density at low collision energies

Centrality Dependence of Mid-Rapidity Yields



- Centrality dependence can be described by power law scaling: Yield = $c \times \langle N_{\text{part}} \rangle^\alpha$
- Strange hadron yields increase faster than $\langle N_{\text{part}} \rangle$ from peripheral to central collisions

Energy dependence of Scaling Parameter α

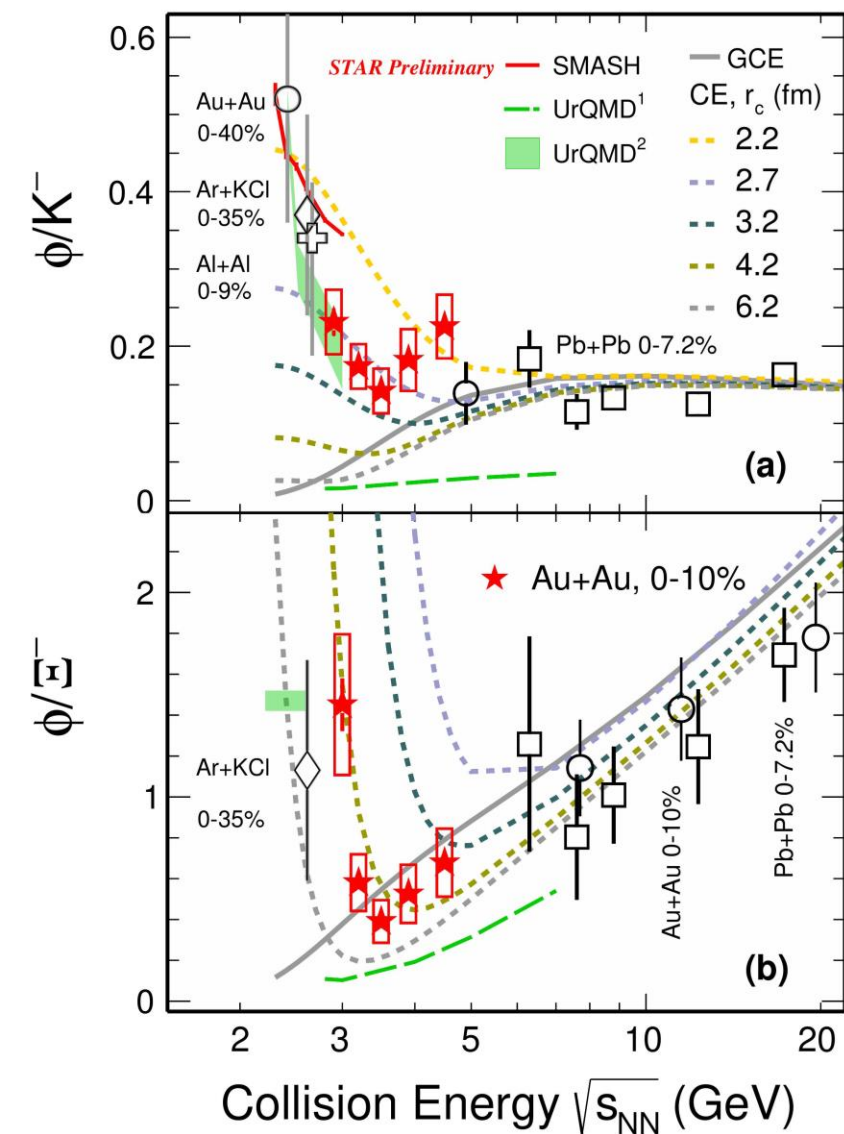


- Centrality dependence of near-threshold production may be sensitive to EoS
- Scaling parameter α decreases with energy
 - $\Lambda(K_S^0)$ and ϕ have similar α
 - Ξ^- has significantly larger α compared to Λ, K_S^0 and ϕ below $\sqrt{s_{NN}} \sim 7$ GeV
- Likely due to Ξ^- mainly produced via multi-step hadronic interactions

e.g.: $NN \rightarrow NN^*$ & $N^*N \rightarrow NEKK$
 $NN \rightarrow N\Lambda K$ & $\Lambda\Lambda \rightarrow N\Xi$
- Transport model simulations UrQMD
 - Qualitatively reproduces the energy dependence
 - Overestimates α for ϕ meson

UrQMD: S.A. Bass, et.al. Prog. Part. Nucl. Phys. 41 (1998)

Energy Dependence of Yield Ratios

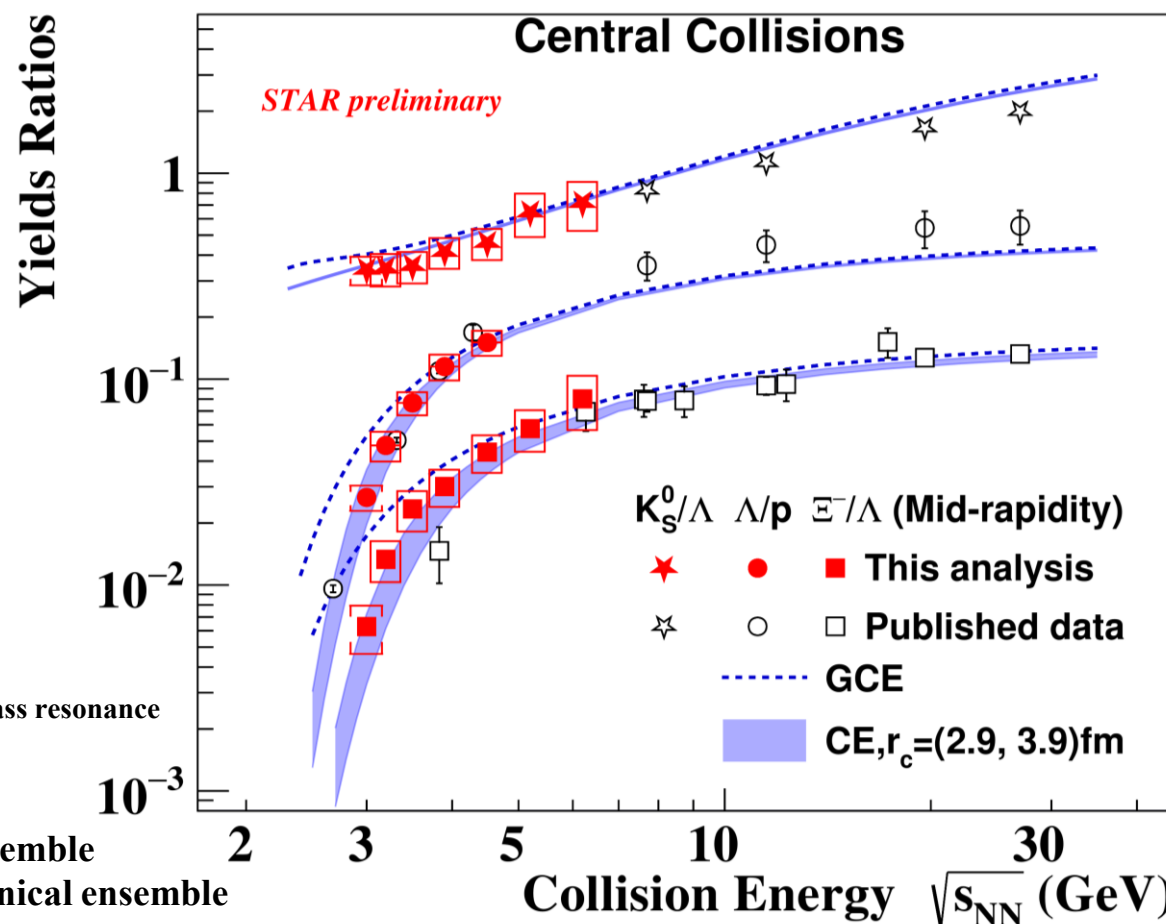


STAR Collaboration. Phys. Lett. B 831, 137152 (2022)
 UrQMD¹: Prog. Part. Nucl. Phys. 41 (1998) 225-370
 UrQMD²: J. Phys. G: Nucl. Part. Phys. 43 015104

UrQMD¹: Default
 UrQMD²: with high mass resonance

CE: canonical ensemble
 GCE: grand canonical ensemble

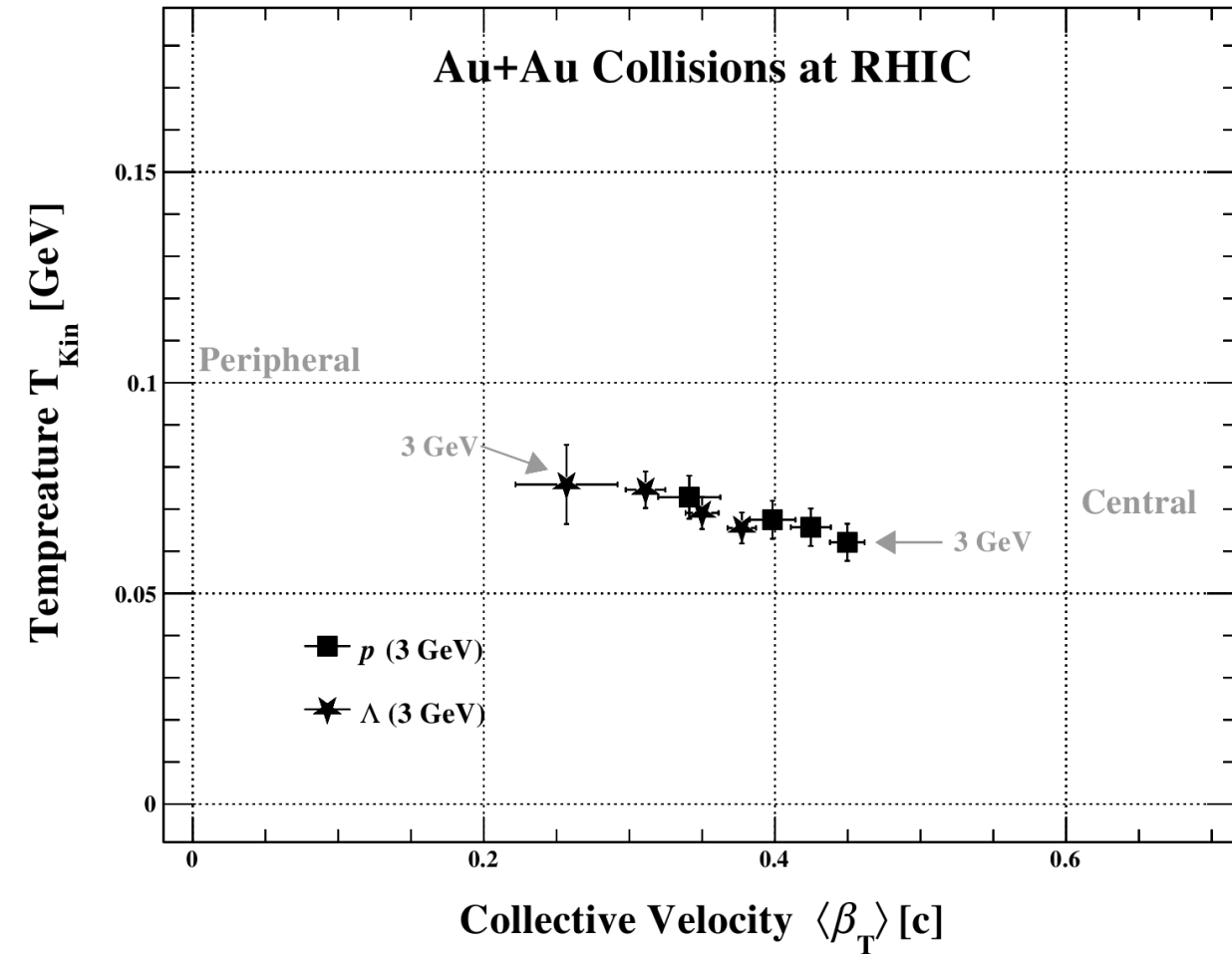
- CE describe yield ratios with $r_c \sim 3 - 4$ fm, but GCE fails below $\sqrt{s_{NN}} \sim 5$ GeV
- Local strangeness conservation is important in high baryon density region
- Medium properties may change in high baryon density region



STAR Collaboration. Phys. Rev. C 102, 034909 (2020);
 Phys. Rev. C 110, 054911 (2024);
 JHEP 2024, 139 (2024)
 V. Vovchenko, et al. Phys. Rev. C 93, 064906 (2016)
 S. Wheaton, et al. Comput. Phys. Commun. 180 (2009)

Kinetic Freeze-out

STAR Collaboration. Phys. Rev. C 110, 054911 (2024); JHEP 2024, 139 (2024)



- **Blast-wave model**

$$\frac{d^2N}{2\pi p_T dp_T dy} \propto \int_0^R r dr m_T I_0 \left(\frac{p_T \sinh \rho(r)}{T_{Kin}} \right) K_1 \left(\frac{m_T \cosh \rho(r)}{T_{Kin}} \right)$$

$$\rho(r) = \frac{1}{2} \ln \frac{1 + \beta_T(r)}{1 - \beta_T(r)}, \quad \beta_T(r) = \beta_{T,Max} \left(\frac{r}{R} \right)^n$$

- **Kinetic freeze-out temperature: T_{Kin}**

- **Collective velocity: $\langle \beta_T \rangle = \frac{2}{n+2} \beta_{T,Max}$**

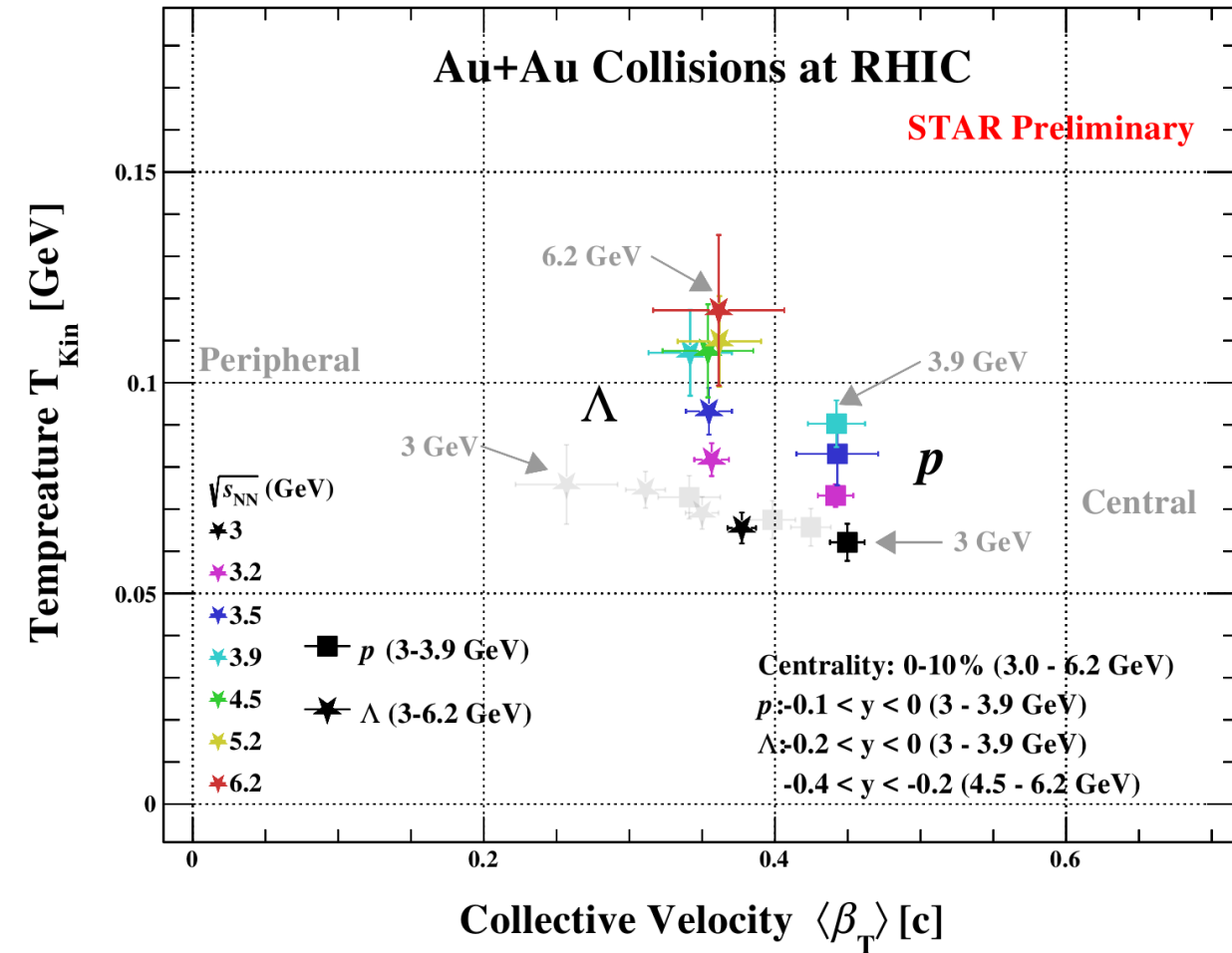
- **Apply Blast-wave fits to proton and Λ spectra separately**

- **$\langle \beta_T \rangle$ decreases while T_{Kin} slightly increases from central to peripheral collisions at $\sqrt{s_{NN}} = 3$ GeV**

- **Smaller fireball and weaker pressure in peripheral collisions**

Kinetic Freeze-out

STAR Collaboration. Phys. Rev. C 110, 054911 (2024); JHEP 2024, 139 (2024)



- Clear energy dependence

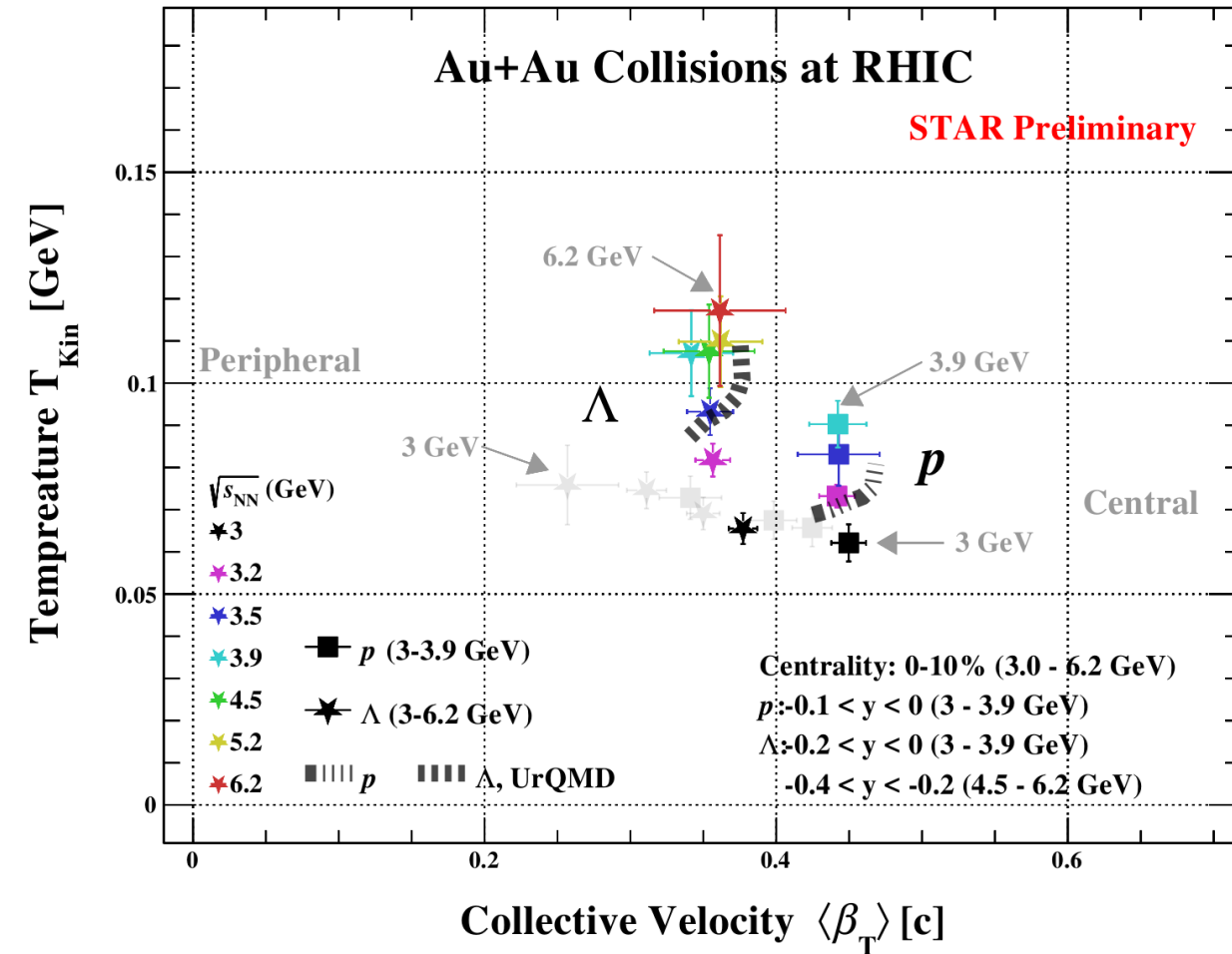
- T_{Kin} increases while $\langle \beta_T \rangle$ remains almost constant from $\sqrt{s_{\text{NN}}} = 3 - 6.2$ GeV for Λ

- Different freeze-out parameters between proton and Λ from $\sqrt{s_{\text{NN}}} = 3 - 3.9$ GeV

- May be due to different production mechanisms

Kinetic Freeze-out

STAR Collaboration. Phys. Rev. C 110, 054911 (2024); JHEP 2024, 139 (2024)



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- Hadronic transport model UrQMD qualitatively reproduces the trend at STAR FXT energies

UrQMD: S.A. Bass, et.al. Prog. Part. Nucl. Phys. 41 (1998)

Summary and Outlook

- **Summary**

- Strange hadron measurements (K^+ , K^- , K_S^0 , ϕ , Λ and Ξ^-) in Au+Au collisions from $\sqrt{s_{NN}} = 3.2 - 6.2$ GeV
 - ❑ CE is mandatory to describe strange hadron yields below $\sqrt{s_{NN}} \sim 5$ GeV $\rightarrow$ implying local strangeness conservation is important in high baryon density region
 - ❑ Significantly larger α for Ξ^- compared to Λ , K_S^0 and ϕ below $\sqrt{s_{NN}} \sim 7$ GeV $\rightarrow$ likely due to production from multi-step hadronic interaction
 - ❑ Different freeze-out parameters for proton and Λ from $\sqrt{s_{NN}} = 3 - 3.9$ GeV $\rightarrow$ likely due to different production mechanisms

- **Outlook**

- Measurements of $\bar{\Lambda}$, $\bar{\Xi}^+$, Ω^- and $\bar{\Omega}^+$ production at near/sub-threshold energy from BES-II

Thanks for your attention!

Back up

NN Collision Threshold Energies

NN Collision Threshold Energy

$NN \rightarrow N\Lambda K \sim 2.548 \text{ GeV}$	$NN \rightarrow NNK\bar{K} \sim 2.864 \text{ GeV}$
$NN \rightarrow NN\phi \sim 2.896 \text{ GeV}$	$NN \rightarrow N\Xi KK \sim 3.247 \text{ GeV}$
$NN \rightarrow NNN\bar{N} \sim 3.753 \text{ GeV}$	$NN \rightarrow N\Omega KKK \sim 4.096 \text{ GeV}$
$NN \rightarrow N\Lambda\bar{\Lambda} \sim 4.107 \text{ GeV}$	$NN \rightarrow N\Xi\bar{\Xi} \sim 4.520 \text{ GeV}$
$NN \rightarrow NN\Omega\bar{\Omega} \sim 5.221 \text{ GeV}$	

