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Particle acceleration in turbulent medium

Recent advances in MHD turbulence call for fundamental revisions in the paradigm of cosmic ray transport and acceleration. I would like to clarify some outstanding issues related to particle transport and

acceleration in realistic turbulent astrophysical environments. I shall discuss

both the transport and acceleration of CRs, and demonstrate that compressible

fast modes dominate the interactions. I shall address effects arising from the

preexisting turbulence and waves generated by CR instabilities and provide implications

for solar flares and SNRs. I shall also discuss how Gamma Ray Bursts (GRBs) properties

can be explained by a model based on turbulent reconnection.

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