

# NEUTRINOS AND BEYOND THE STANDARD MODEL SEARCHES

**Michele Lucente**

XVII International Conference on  
Heavy Quarks and Leptons

*18<sup>th</sup> September 2025, Peking University*

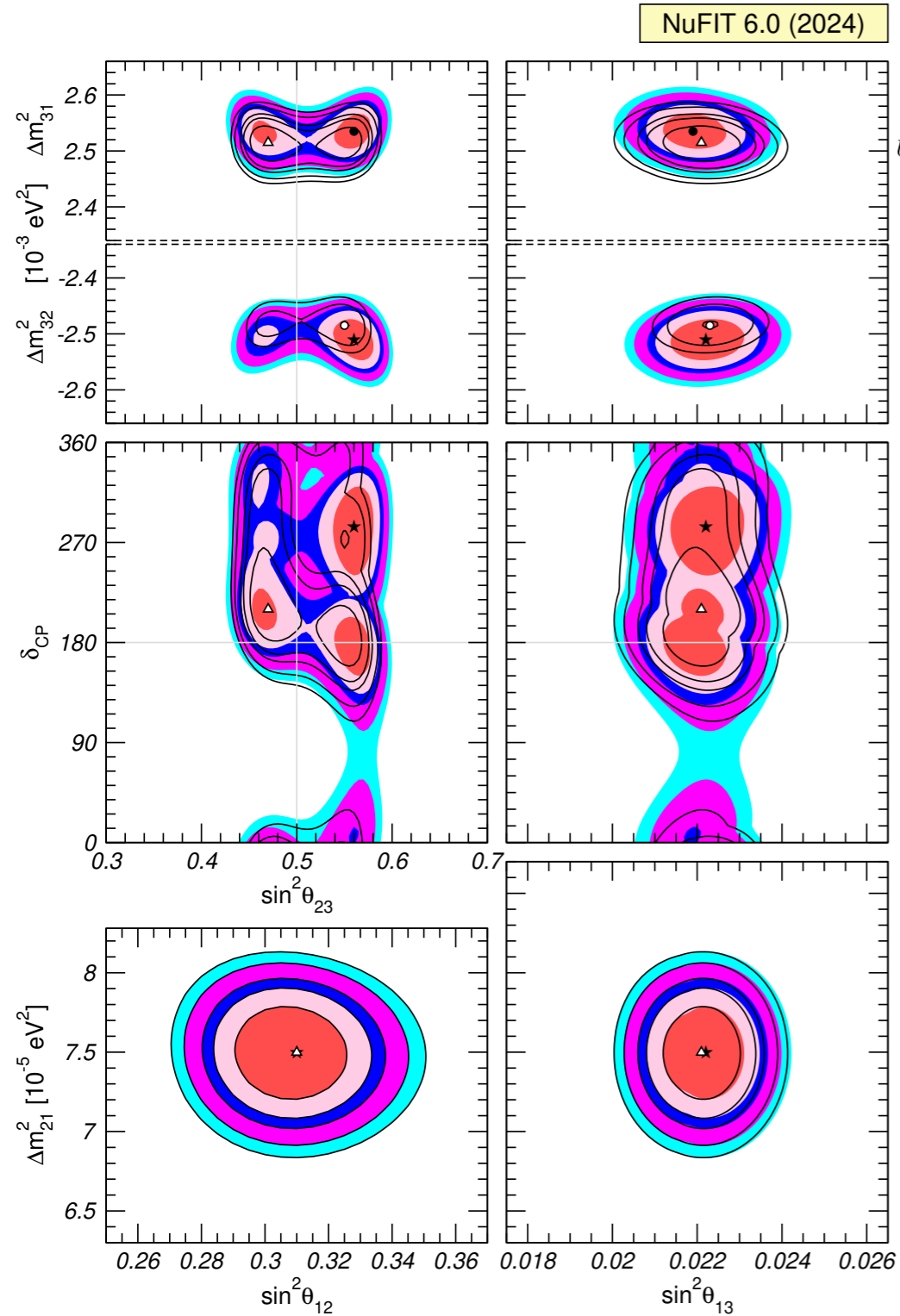


ALMA MATER STUDIORUM  
UNIVERSITÀ DI BOLOGNA



**Funded by  
the European Union**

# Neutrinos are massive and leptons mix



$$U^{\text{NO}} = \begin{pmatrix} 0.7838 \rightarrow 0.8442 & 0.5135 \rightarrow 0.6004 & (-0.1568 \rightarrow 0.1489) + i(-0.1182 \rightarrow 0.1520) \\ (-0.4831 \rightarrow -0.2394) + i(-0.0749 \rightarrow 0.0962) & (0.4636 \rightarrow 0.6749) + i(-0.0521 \rightarrow 0.0668) & 0.6499 \rightarrow 0.7719 \\ (0.3068 \rightarrow 0.5391) + i(-0.0643 \rightarrow 0.0933) & (-0.6897 \rightarrow -0.4821) + i(-0.0446 \rightarrow 0.0644) & 0.6161 \rightarrow 0.7434 \end{pmatrix}$$

P. F. de Salas, D. V. Forero, S. Gariazzo, P. Martínez-Miravé, O. Mena, C. A. Ternes, M. Tórtola and J. W. F. Valle, arXiv:2006.11237 [hep-ph]

**Evidence for BSM physics!**

Parameter	Ordering	Best fit	1 $\sigma$ range	2 $\sigma$ range	3 $\sigma$ range	"1 $\sigma$ " (%)
$\delta m^2 / 10^{-5} \text{ eV}^2$	NO, IO	7.37	7.21 – 7.52	7.06 – 7.71	6.93 – 7.93	2.3
$\sin^2 \theta_{12} / 10^{-1}$	NO, IO	3.03	2.91 – 3.17	2.77 – 3.31	2.64 – 3.45	4.5
$ \Delta m^2  / 10^{-3} \text{ eV}^2$	NO	2.495	2.475 – 2.515	2.454 – 2.536	2.433 – 2.558	0.8
	IO	2.465	2.444 – 2.485	2.423 – 2.506	2.403 – 2.527	0.8
$\sin^2 \theta_{13} / 10^{-2}$	NO	2.23	2.17 – 2.27	2.11 – 2.33	2.06 – 2.38	2.4
	IO	2.23	2.19 – 2.30	2.14 – 2.35	2.08 – 2.41	2.4
$\sin^2 \theta_{23} / 10^{-1}$	NO	4.73	4.60 – 4.96	4.47 – 5.68	4.37 – 5.81	5.1
	IO	5.45	5.28 – 5.60	4.58 – 5.73	4.43 – 5.83	4.3
$\delta / \pi$	NO	1.20	1.07 – 1.37	0.88 – 1.81	0.73 – 2.03	18
	IO	1.48	1.36 – 1.61	1.24 – 1.72	1.12 – 1.83	8
$\Delta \chi^2_{\text{IO-NO}}$	IO-NO	+5.0				

I. Esteban, M. C. Gonzalez-Garcia, M. Maltoni, I. Martinez-Soler, J. P. Pinheiro and T. Schwetz, arXiv:2410.05380 [hep-ph]

F. Capozzi, W. Giarè, E. Lisi, A. Marrone, A. Melchiorri and A. Palazzo, arXiv:2503.07752 [hep-ph]



# Naturally small neutrino masses

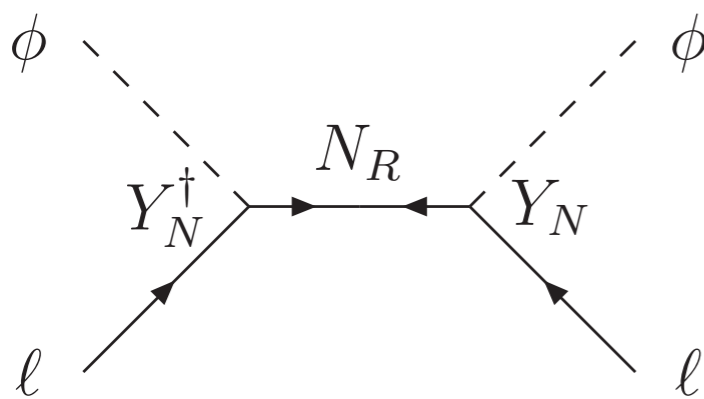
There are only 3 tree-level realisations of neutrino masses in the minimal SM

E. Ma, arXiv:hep-ph/9805219 [hep-ph]

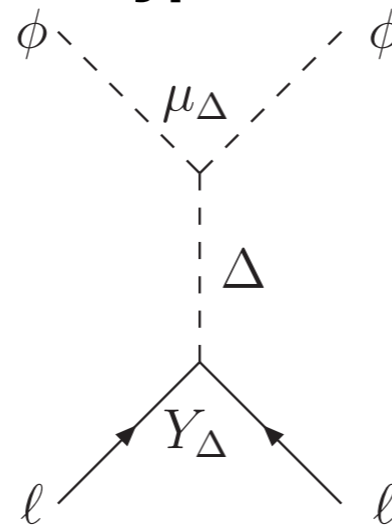
A. Abada, C. Biggio, F. Bonnet, M. B. Gavela and T. Hambye, arXiv:0707.4058 [hep-ph]

Generate  $\frac{1}{2} \frac{c_{\alpha\beta}}{\Lambda} \left( \overline{l_{L\alpha}^c} \tilde{\Phi}^* \right) \left( \tilde{\Phi}^\dagger l_L^\beta \right) + h.c.$  from seesaw mechanism

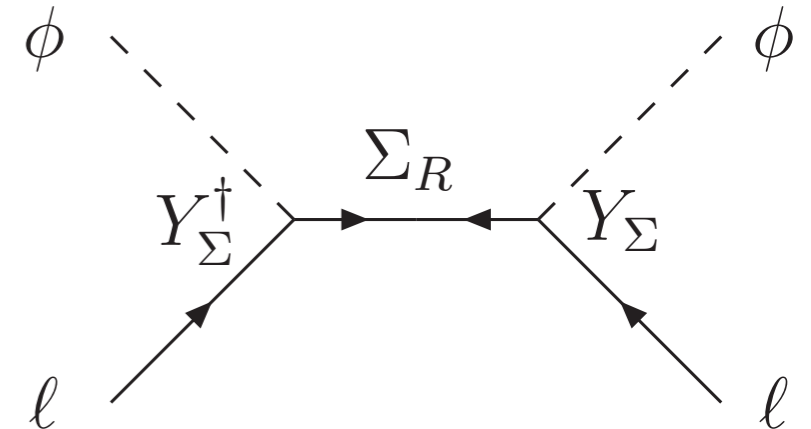
**Type-I**



**Type-II**



**Type-III**



R. Foot, H. Lew, X. G. He and G. C. Joshi, Z. Phys. C 44 (1989), 441; E. Ma and D. P. Roy, arXiv:hep-ph/0206150 [hep-ph]; T. Hambye, Y. Lin, A. Notari, M. Papucci and A. Strumia, arXiv:hep-ph/0312203 [hep-ph]; B. Bajc and G. Senjanovic, arXiv:hep-ph/0612029 [hep-ph]; B. Bajc, M. Nemevsek and G. Senjanovic, arXiv:hep-ph/0703080 [hep-ph]; I. Dorsner and P. Fileviez Perez, arXiv:hep-ph/0612216 [hep-ph]; P. Fileviez Perez, arXiv:hep-ph/0702287 [hep-ph]; P. Fileviez Perez, arXiv:0705.3589 [hep-ph]

P. Minkowski, Phys. Lett. B 67 (1977), 421-428; M. Gell-Mann, P. Ramond and R. Slansky, arXiv:1306.4669 [hep-th]; T. Yanagida in Proceedings of the Workshop on the Unified Theories and the Baryon Number in the Universe (1979); R. N. Mohapatra and G. Senjanovic, Phys. Rev. Lett. 44 (1980), 912

M. Magg and C. Wetterich, Phys. Lett. B 94 (1980), 61-64; J. Schechter and J. W. F. Valle, Phys. Rev. D 22 (1980), 2227; C. Wetterich, Nucl. Phys. B 187 (1981), 343-375; G. Lazarides, Q. Shafi and C. Wetterich, Nucl. Phys. B 181 (1981), 287-300; R. N. Mohapatra and G. Senjanovic, Phys. Rev. D 23 (1981), 165

# Type-I seesaw mechanism

Complete the SM field pattern with **right-handed neutrinos**

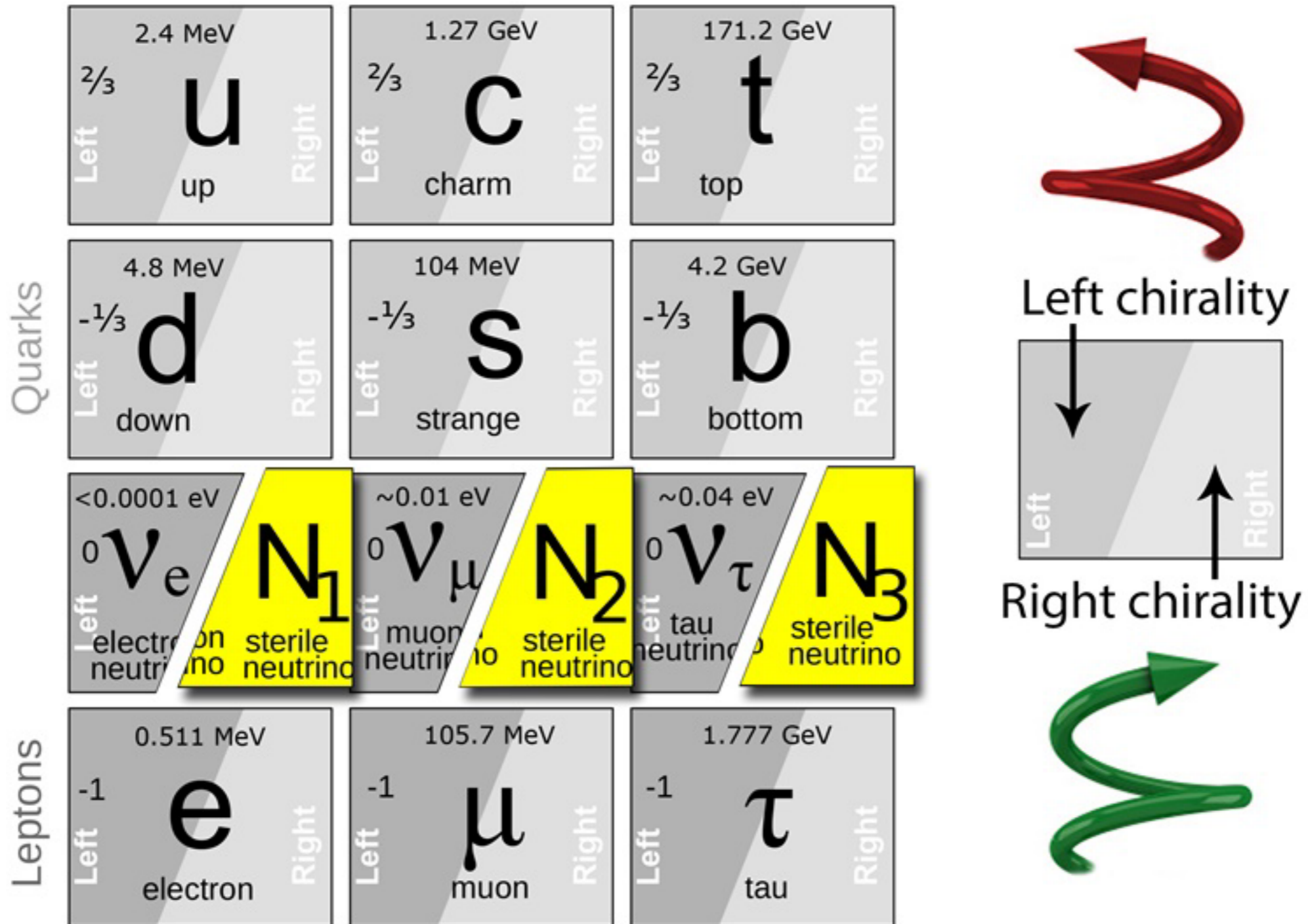


Figure from S. Alekhin *et al.*, arXiv:1504.04855 [hep-ph]

# Neutrino masses from sterile singlets

**Type-I seesaw mechanism:** SM + gauge singlet fermions  $N_I$

$$\mathcal{L} = \mathcal{L}_{\text{SM}} + i\overline{N_I}\not{\partial}N_I - \left( F_{\alpha I} \overline{\ell}_L^\alpha \tilde{\phi} N_I + \frac{M_{IJ}}{2} \overline{N_I^c} N_J + h.c. \right)$$

After electroweak phase transition  $\langle \Phi \rangle = v \approx 174 \text{ GeV}$

$$m_\nu = -v^2 F \frac{1}{M} F^T$$

**$m_\nu$  is much smaller than EW scale**

Laboratory

$$m_\nu < 1.1 \text{ eV}$$

Cosmology

$$\sum m_\nu < 0.12 \text{ eV}$$

KATRIN collaboration, arXiv:1909.06048 [hep-ex]

Planck collaboration, arXiv:1807.06209 [astro-ph.CO]

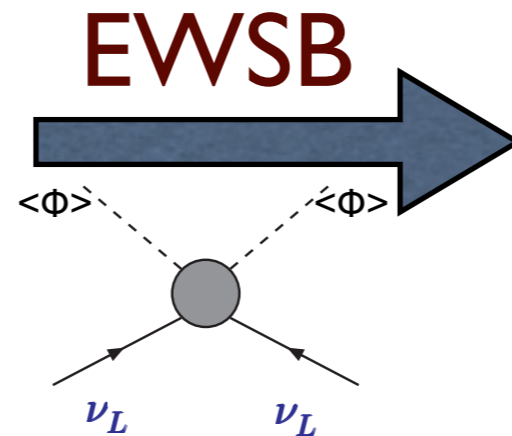
# SM as an effective theory

Relaxing the renormalizability condition there is only one dim=5 gauge invariant operator  
(Weinberg operator) S. Weinberg, Phys. Rev. Lett. 43 (1979) 1566

## Lepton number violation

$$\frac{1}{2} \frac{c_{\alpha\beta}}{\Lambda} \left( \overline{l_{L\alpha}^c} \tilde{\Phi}^* \right) \left( \tilde{\Phi}^\dagger l_L^\beta \right) + h.c.$$

$\Delta L = 2$



## Neutrino masses and mixing

$$\frac{v^2}{2} \frac{c_{\alpha\beta}}{\Lambda} \overline{\nu_{L\alpha}^c} \nu_{L\beta} + h.c.$$

New physics scale

$$m_{\alpha\beta}^\nu = c_{\alpha\beta} \frac{v}{\Lambda} v \lesssim \text{eV} \ll v$$

**Why are neutrinos so light?**

- Suppression mechanisms
- $\frac{v}{\Lambda} \ll 1$     High NP scale
  - $c_{\alpha\beta} \ll 1$     Symmetry (Lepton number)
  - $c_{\alpha\beta} \ll 1$     Accidental cancellations

# Unveiling neutrino mass generation mechanism

$$\mathcal{L}_{eff} = \mathcal{L}_{SM} + \frac{c_5}{\Lambda} \mathcal{O}^{d=5} + \frac{c_6^i}{\Lambda^2} \mathcal{O}_i^{d=6} + \dots$$

v masses and mixing  
common to all SM  
extensions with Majorana v

New physics effects

**If only  $\Lambda$  at work**

$$\frac{c_6^i}{\Lambda^2} \approx \left(\frac{c_5}{\Lambda}\right)^2 \approx \left(\frac{m_\nu}{v^2}\right)^2$$

New physics effects  
strongly suppressed  
by the v mass scale

**If symmetry at work**

$$c_5 \ll 1 \quad \text{and} \quad c_6^{\text{LNV},i} \ll 1$$

$$c_6^{\text{LNC},i} \approx \mathcal{O}(1)$$

possible for L  
conserving operators

**If accidental cancellation**

$$c_5 \ll 1$$

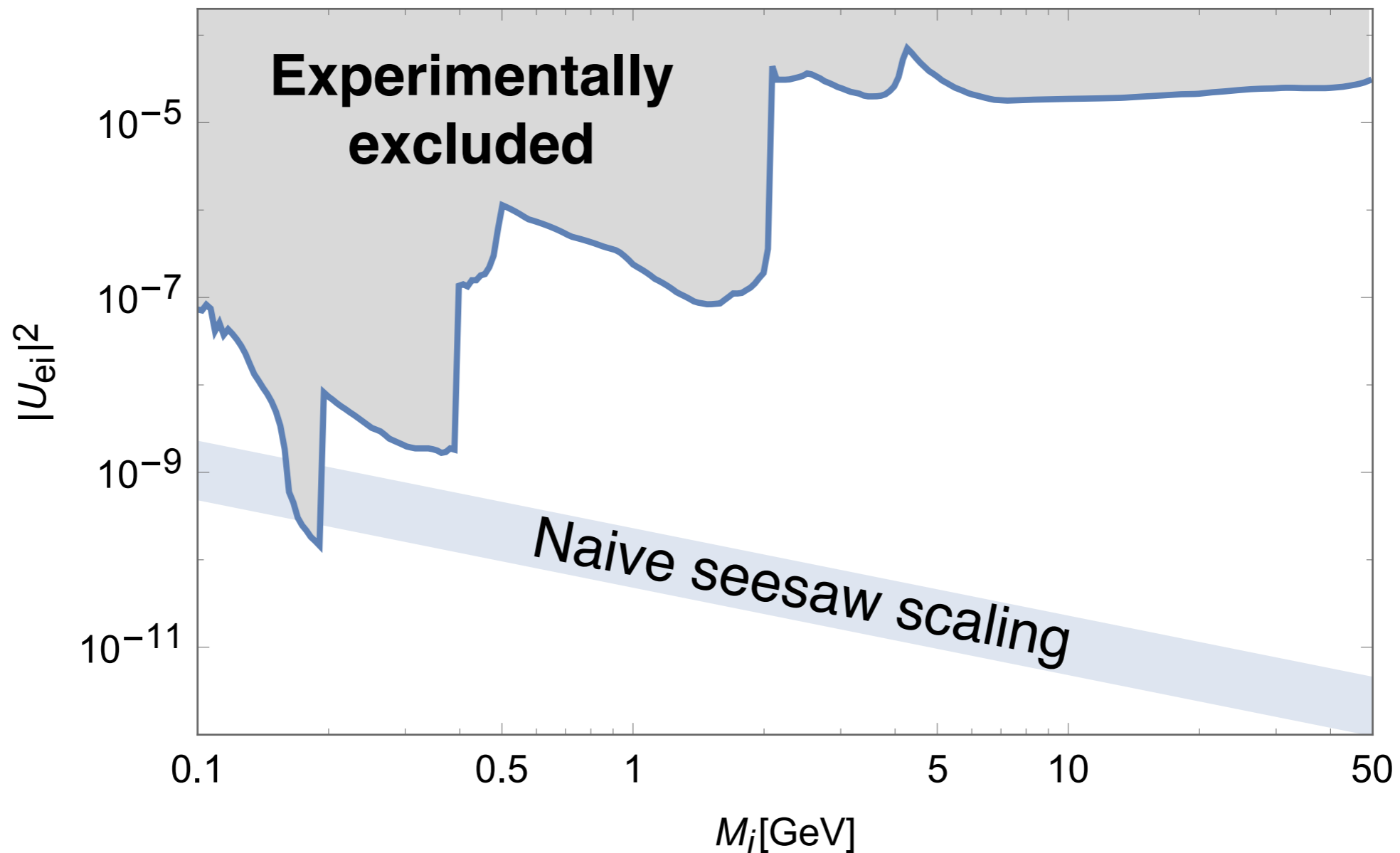
$$c_6^i \approx \mathcal{O}(1)$$

possible for all operators

# Λ suppression: naive Seesaw scaling

Seesaw scaling  $m_\nu = -v^2 F \frac{1}{M} F^T$

In the **absence** of any **structure** in the  $F$  and  $M$  matrices  $|U_{\alpha i}| \lesssim \sqrt{\frac{m_\nu}{M}} \lesssim 10^{-5} \sqrt{\frac{\text{GeV}}{M}}$





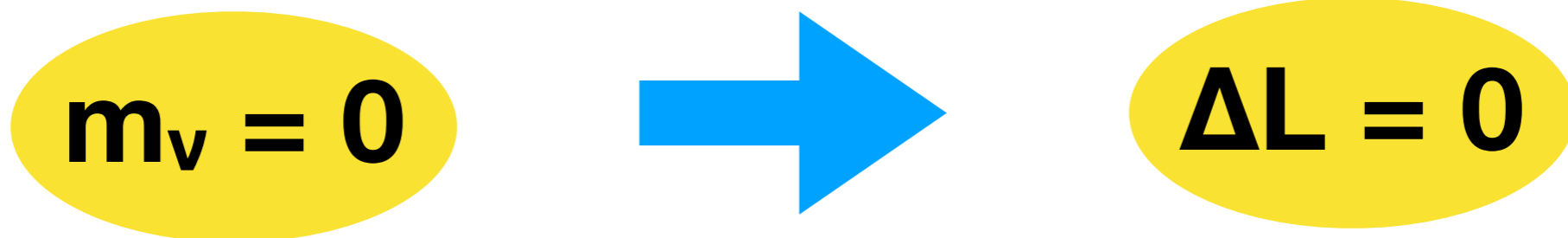
# Symmetries: L number has a special role

## Theorem: SM + fermionic gauge singlets

K. Moffat, S. Pascoli and C. Weiland, arXiv:1712.07611 [hep-ph]

*“The most general gauge-singlet neutrino extensions of the SM with no cancellation between different orders of the seesaw expansion, no fine-tuned cancellations between different radiative orders and which lead to three massless neutrinos are lepton number conserving”*

**In the SM extended with fermionic gauge singlets (e.g. Right-Handed neutrinos)**



Unless there are accidental cancellations in  $m_\nu$ , the rate for Lepton number violating events is proportional to the small active neutrino masses

**The theorem extends and generalises previous results:** G. Ingelman and J. Rathsman, Z. Phys. C 60 (1993) 243; J. Gluza, hep-ph/0201002; J. Kersten and A. Y. Smirnov, arXiv:0705.3221 [hep-ph]

# L symmetry and Majorana fields

Majorana fermions violate all global symmetries, including L

**How to preserve lepton number with Majorana states?**

	Pair two states to form a <b>Dirac</b> state ( <i>equal masses, maximal mixing, opposite CP</i> )	<b>Decouple</b> a state	Have a <b>massless</b> state
<b>Exact symmetry</b>	$M_1 = M_2$ $\mathcal{U}_{\alpha 1} = i \mathcal{U}_{\alpha 2}$	$\mathcal{U}_{\alpha i} = 0$	$M_i = 0$
<b>Approximate symmetry</b>	$\frac{M_2 - M_1}{M_1 + M_2} \ll 1$ $\mathcal{U}_{\alpha 1} \simeq i \mathcal{U}_{\alpha 2}$	$ \mathcal{U}_{\alpha, i}  \ll  \mathcal{U}_{\alpha, j \neq i} $	$M_i \ll M_{j \neq i}$

# NEUTRINOLESS DOUBLE BETA DECAY

# Neutrinoless double beta decay: $\Delta L = 2$

W. H. Furry, Phys. Rev. 56 (1939) 1184

If neutrinos are Majorana particles  $0\nu 2\beta$  is possible

$$\mathcal{N}(A, Z) \rightarrow \mathcal{N}(A, Z + 2) + 2e^-$$

Clear experimental signature

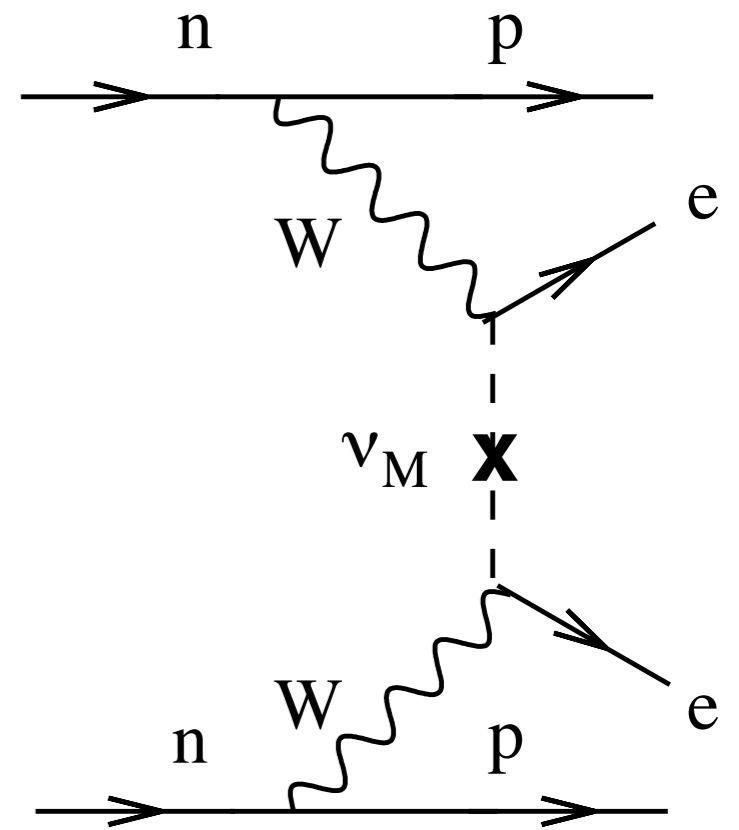
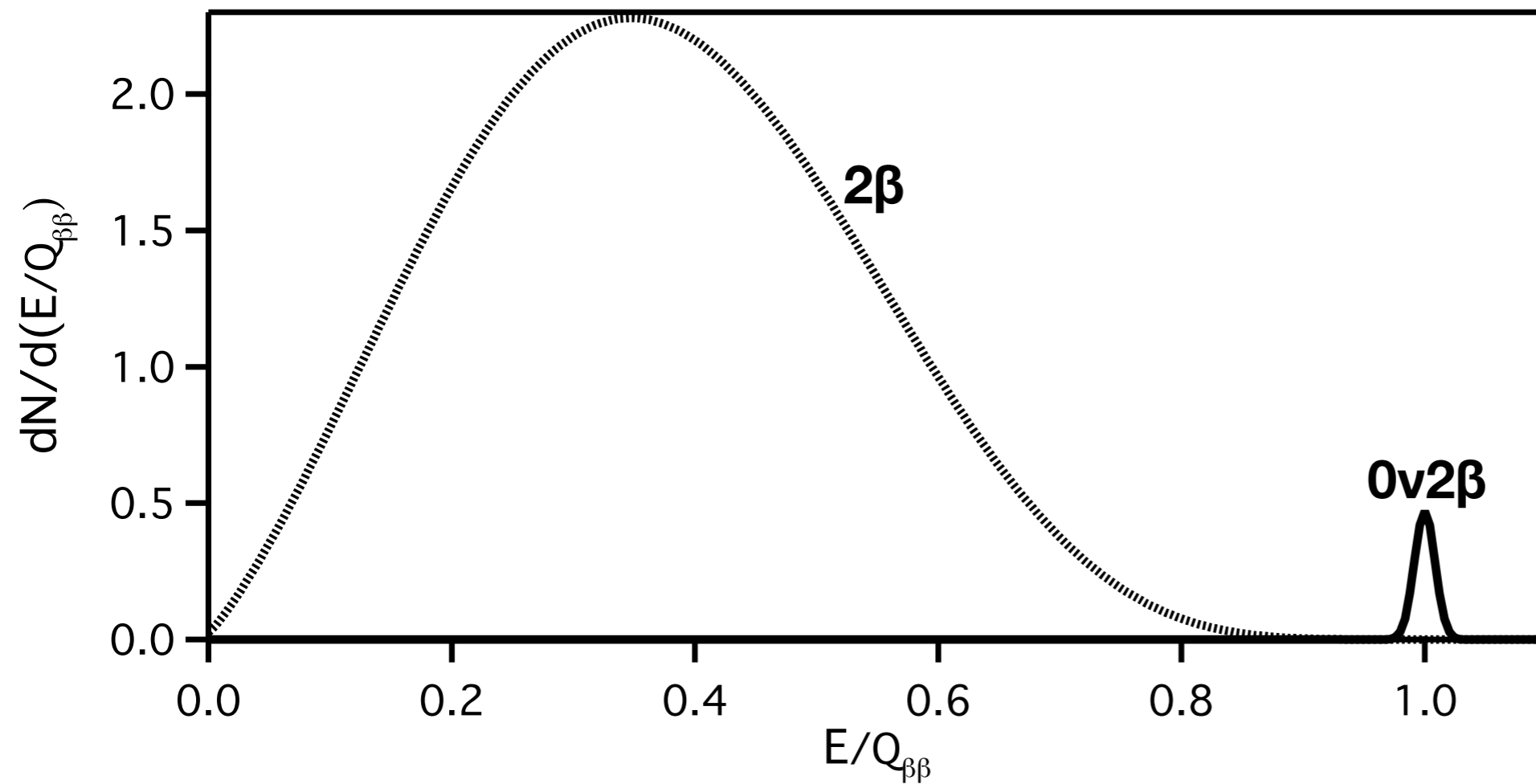


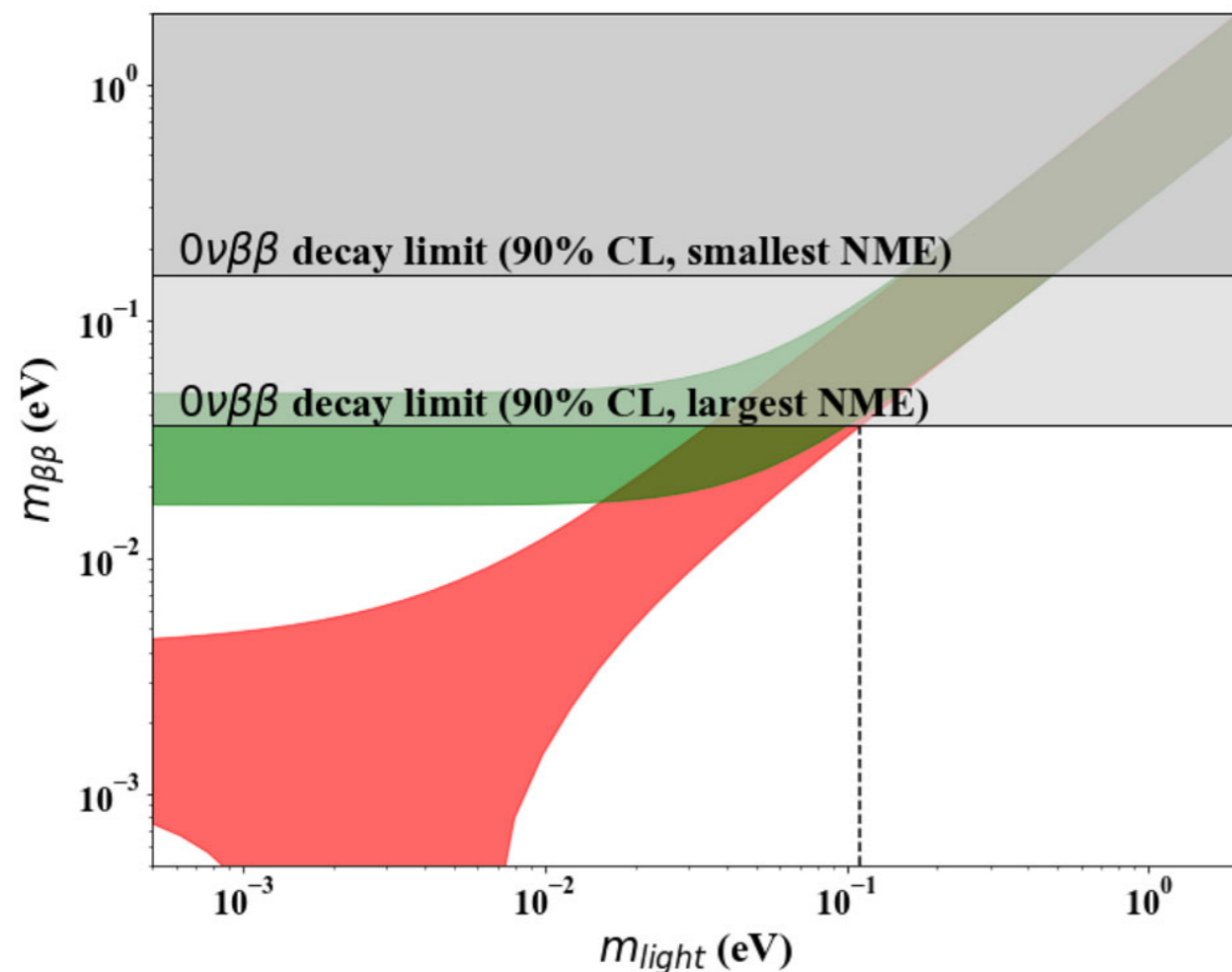
Figure modified from F. T. Avignone III, S. R. Elliott and J. Engel, arXiv:0708.1033 [nucl-ex]

# Experimental status: minimal SM

The amplitude for light neutrino exchange is proportional to

$$m_{2\beta} = \left| \sum_i U_{ei}^2 m_i \right|$$

Isotope	$T_{1/2}^{2\nu}$ (year)	Experiments
$^{48}\text{Ca}$	$(5.3^{+1.2}_{-0.8}) \times 10^{19}$	Irvine TPC [44], TGV [45], NEMO-3 [46]
$^{76}\text{Ge}$	$(1.88 \pm 0.08) \times 10^{21}$	HEIDELBERG-MOSCOW [47], GERDA [48]
$^{82}\text{Se}$	$(0.87 \pm^{+0.02}_{-0.01}) \times 10^{20}$	NEMO-3 [49], CUPID-0 [50], Irvine TPC [51], NEMO-2 [52]
$^{96}\text{Zr}$	$(2.3 \pm 0.2) \times 10^{19}$	NEMO-2 [53], NEMO-3 [54]
$^{100}\text{Mo}$	$(7.06^{+0.15}_{-0.13}) \times 10^{18}$	NEMO-3 [55], CUPID-Mo [56], NEMO-2 [57], Irvine TPC [58], ZnMoO <sub>4</sub> bolometers [59]
$^{116}\text{Cd}$	$(2.69 \pm 0.09) \times 10^{19}$	NEMO-3 [60], Aurora [61], ELEGANT [62], Solotvina [63], NEMO-2 [64]
$^{130}\text{Te}$	$(7.91 \pm 0.21) \times 10^{20}$	CUORE-0 [65], CUORE [66], CUORICINO [67], NEMO-3 [68]
$^{136}\text{Xe}$	$(2.18 \pm 0.05) \times 10^{21}$	EXO-200 [69], KamLAND-Zen [70]
$^{150}\text{Nd}$	$(9.34 \pm 0.65) \times 10^{18}$	NEMO-3 [71]



From current knowledge on neutrino oscillation parameters it is possible to compute  $m_{2\beta}$  as a function of unknown lightest neutrino mass, ordering and CP phases

J. J. Gómez-Cadenas, J. Martín-Albo, J. Menéndez, M. Mezzetto, F. Monrabal and M. Sorel,  
Riv. Nuovo Cim. 46 (2023) no.10, 619-692

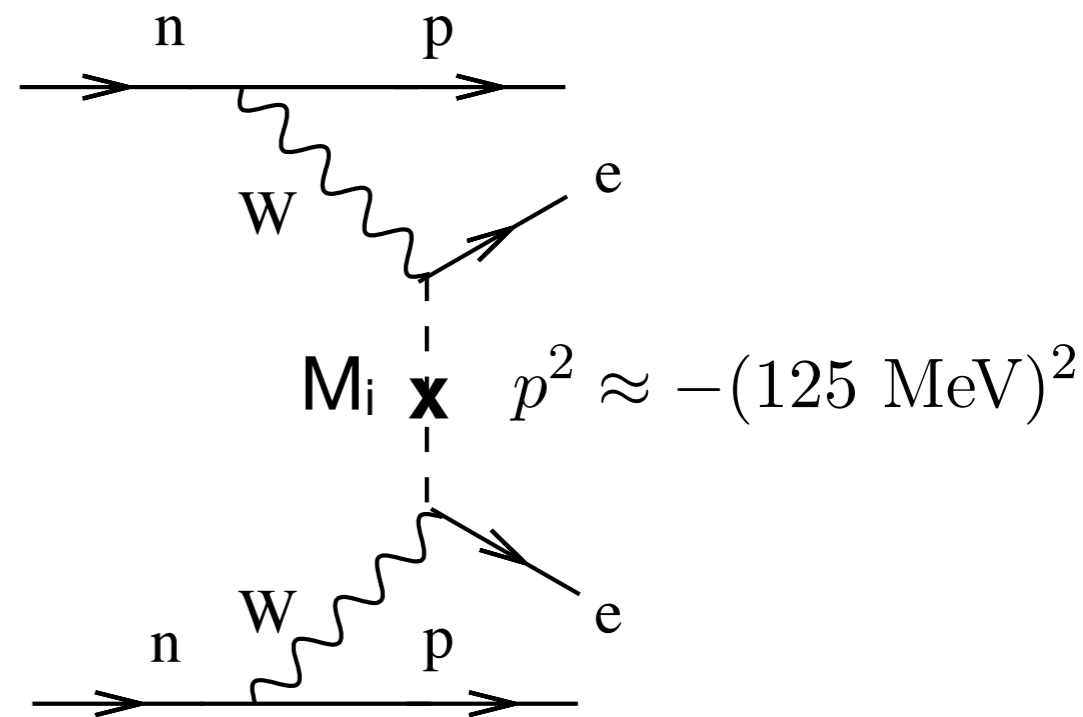
# Contribution of heavy neutrinos

## Heavy Majorana neutrinos contribute as well to $0\nu 2\beta$ amplitude

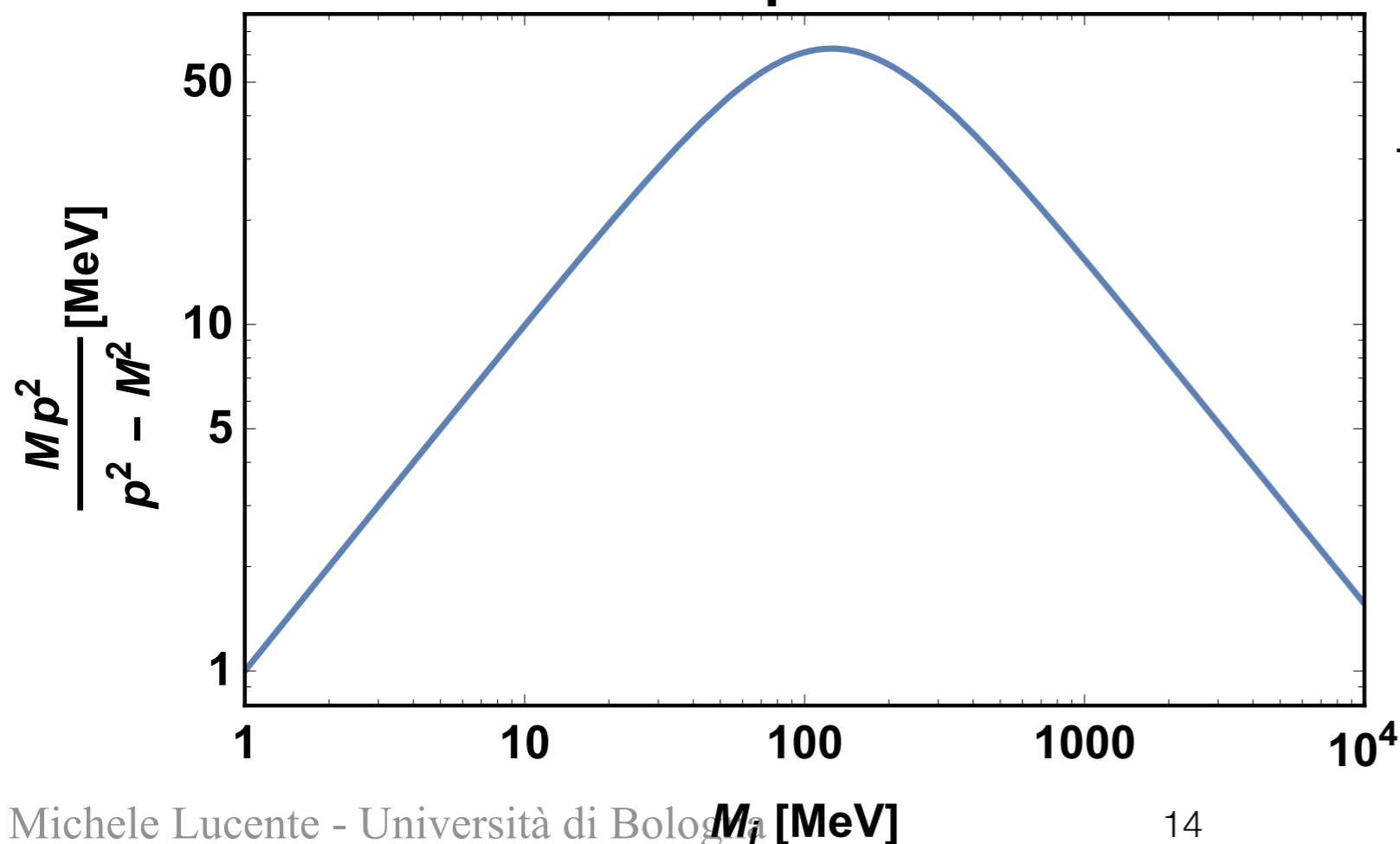
F. L. Bezrukov, hep-ph/0505247; M. Blennow, E. Fernandez-Martinez, J. Lopez-Pavon and J. Menendez, arXiv:1005.3240 [hep-ph]; A. Abada and M.L., arXiv:1401.1507 [hep-ph]; A. Faessler, M. González, S. Kovalenko and F. Šimkovic, arXiv:1408.6077 [hep-ph]; A. Abada, V. De Romeri, M.L., A. M. Teixeira and T. Toma, arXiv:1712.03984 [hep-ph]; A. Babič, S. Kovalenko, M. I. Krivoruchenko and F. Šimkovic, arXiv:1804.04218 [hep-ph]

$$A^{0\nu 2\beta} \propto \sum_i M_i U_{ei}^2 M^{0\nu 2\beta}(M_i)$$

$$M^{0\nu 2\beta}(M_i) \simeq M^{0\nu 2\beta}(0) \frac{p^2}{p^2 - M_i^2}$$



Mass dependence



If pseudo-Dirac

$$M_1 \simeq M_2$$

$$U_{e1} \simeq i U_{e2}$$

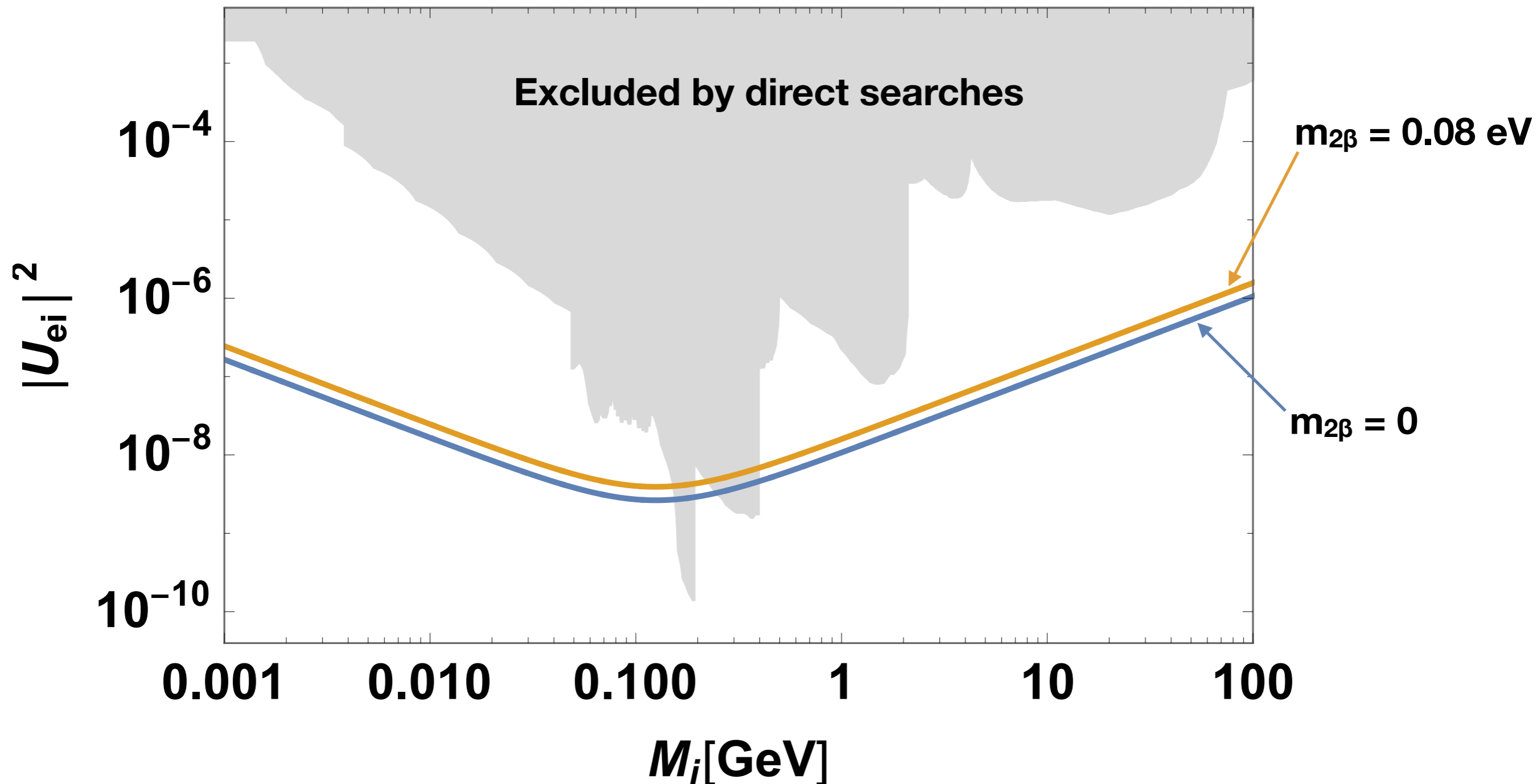
cancellation between contributions of single Majorana states



# Extracting constraints on heavy neutrinos

$$A^{0\nu 2\beta} \propto \sum_{i=1}^{3+n} M_i \mathcal{U}_{ei}^2 M^{0\nu 2\beta}(M_i)$$

$0\nu 2\beta$  constraints depend on the full mass spectrum (light + heavy)



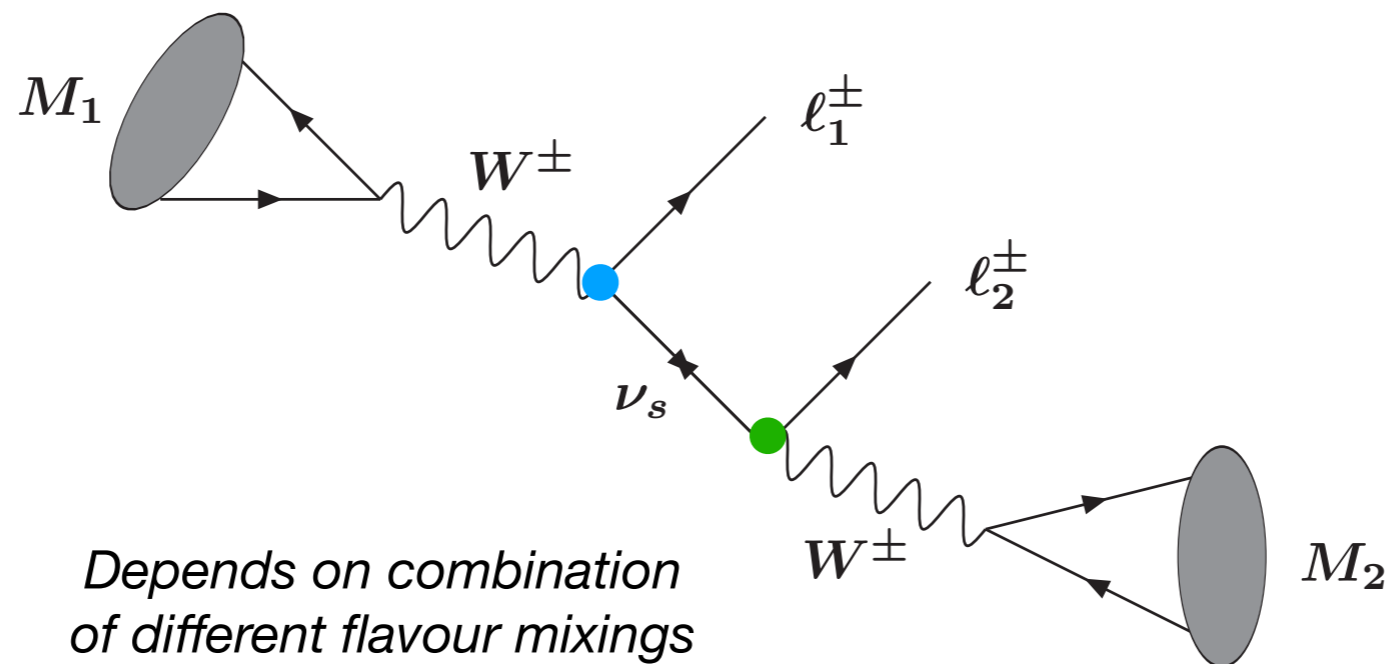
These constraints do not apply to (pseudo-)Dirac particles

# TAU AND MESON DECAY

# L-violating $\tau$ and meson decay

Heavy Majorana neutrinos can mediate L-violating decays of pseudo-scalar mesons and  $\tau$  lepton

$$M_1(p, m_{M_1}) \rightarrow \ell_\alpha(k_1, m_{\ell_\alpha}) \ell_\beta(k_2, m_{\ell_\beta}) M_2(k_3, m_{M_2})$$



Depends on combination of different flavour mixings

$$i\mathcal{M}_P \equiv i\mathcal{M}_{P1} + i\mathcal{M}_{P2} = 2i G_F^2 V_{M_1} V_{M_2} U_{\ell_\alpha 4} U_{\ell_\beta 4} m_4 f_{M_1} f_{M_2} \left[ \frac{\bar{u}(k_1) \not{k}_3 \not{p} P_R v(k_2)}{m_{31}^2 - m_4^2 + im_4 \Gamma_4} + \frac{\bar{u}(k_1) \not{p} \not{k}_3 P_R v(k_2)}{m_{23}^2 - m_4^2 + im_4 \Gamma_4} \right]$$

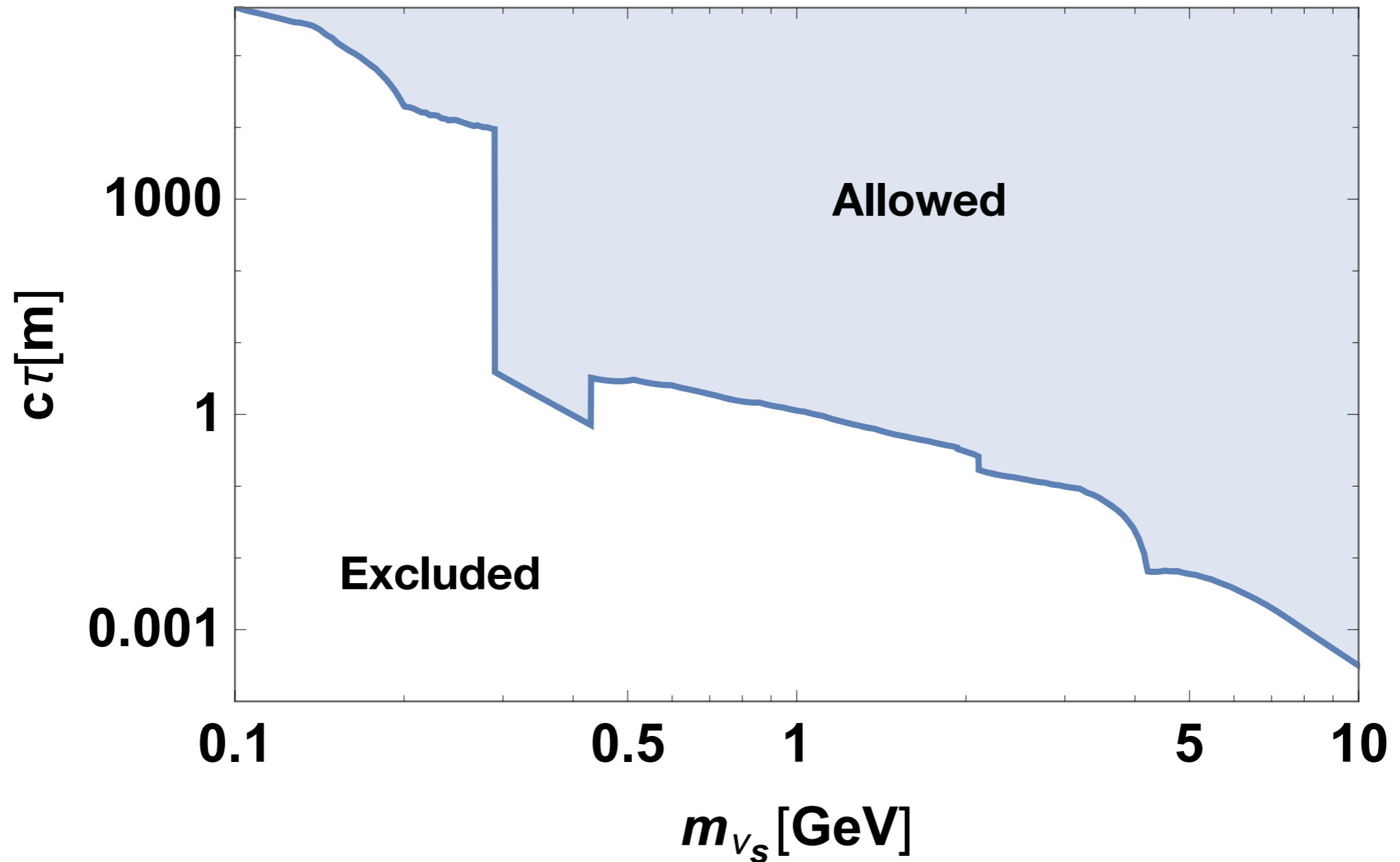
**Negligible amplitude unless the intermediate state can go on-shell**

$$\frac{1}{(m_{ij}^2 - m_4^2)^2 + m_4^2 \Gamma_4^2} \rightarrow \frac{\pi}{m_4 \Gamma_4} \delta(m_{ij}^2 - m_4^2)$$

# Lifetime limitations

In the resonant regime  $i\mathcal{M} \propto \frac{M_{\nu_s}}{\Gamma_{\nu_s}} \equiv M_{\nu_s} \tau_{\nu_s}$

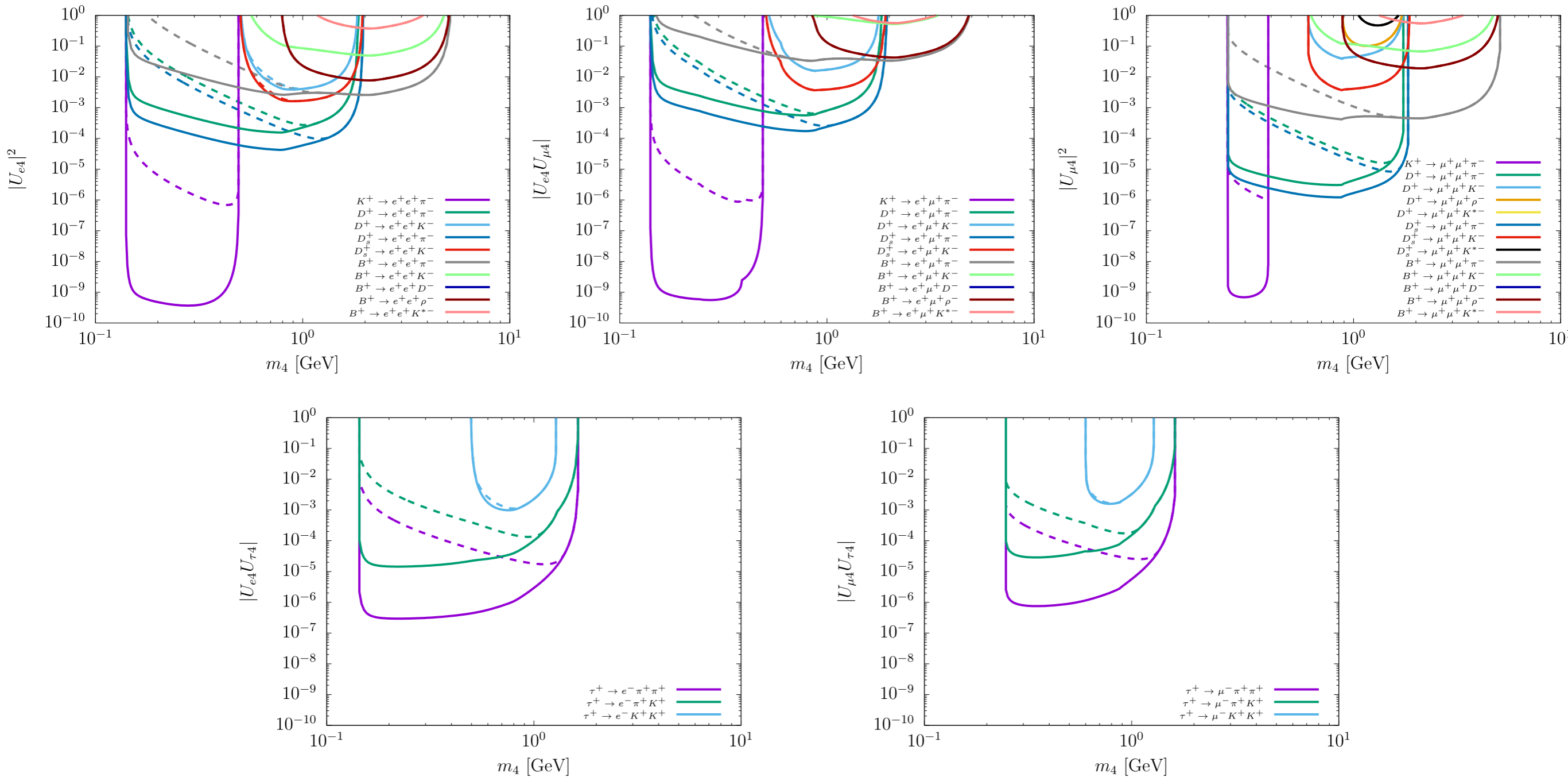
**But too long-lived heavy neutrinos decay outside the detector**



Asking for observable (inside detector) decays imposes a further constraint

# Constraints: single intermediate state

Figures from A. Abada, V. De Romeri, M.L., A. M. Teixeira and T. Toma, arXiv:1712.03984 [hep-ph];  
see also A. Atre, T. Han, S. Pascoli and B. Zhang, arXiv:0901.3589 [hep-ph]



**Dashed lines:** the on-shell heavy neutrino travels for less than 10 m

# Multiple intermediate states: interference

A. Abada, C. Hati, X. Marcano and A. M. Teixeira, arXiv:1904.05367 [hep-ph]

If more than one heavy neutrino mediate the process, and

$$\Delta M \ll M \quad \text{and} \quad \Delta M < \Gamma_N$$

interference effects arise due to the CP-violating phases

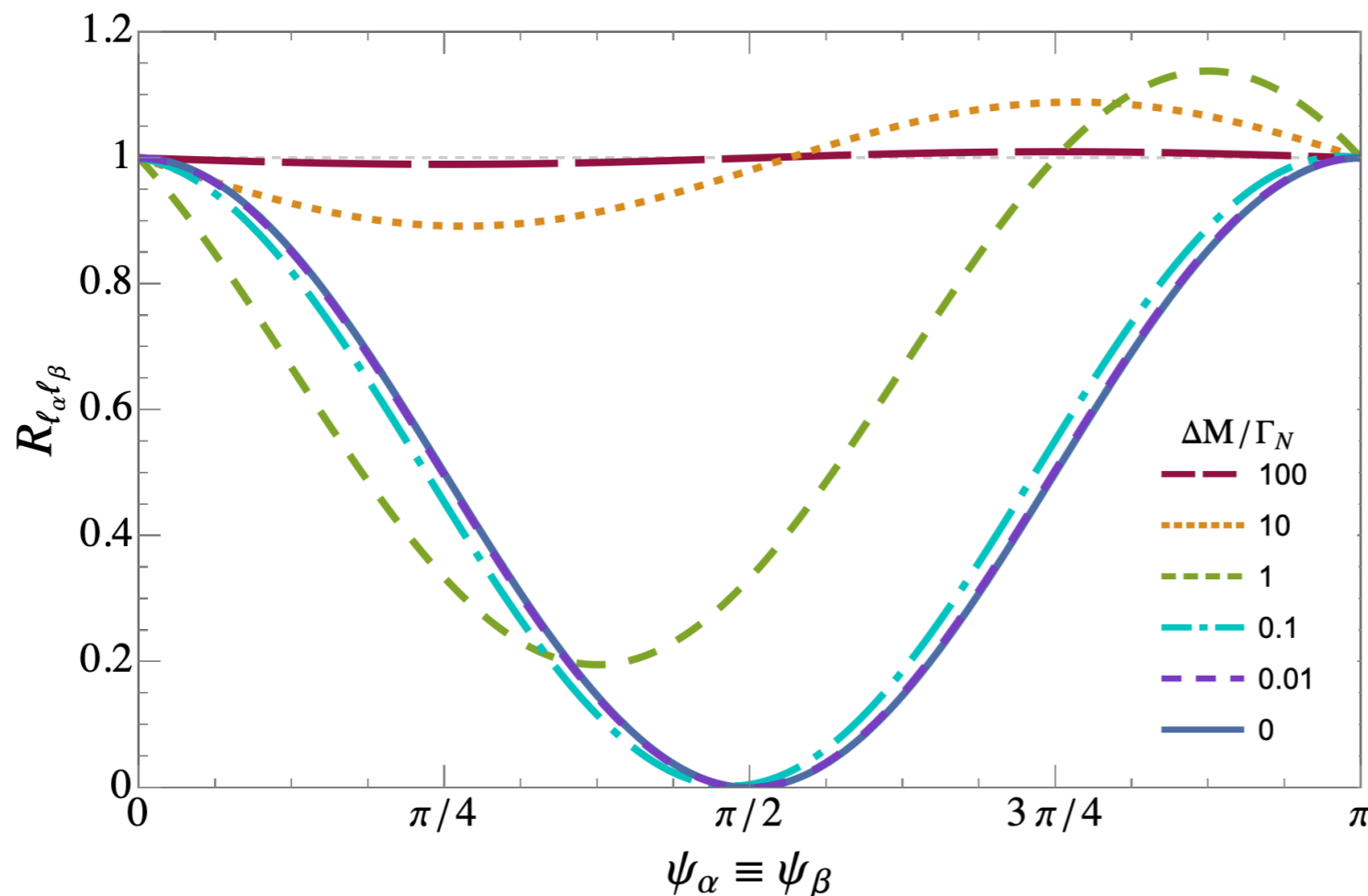
$$\left| \mathcal{A}_{M \rightarrow M' l_\alpha^+ l_\beta^-}^{\text{LNC}} \right|^2 \propto \left| U_{\alpha 4} U_{\beta 4}^* g(m_4) + U_{\alpha 5} U_{\beta 5}^* g(m_5) \right|^2,$$

$$\left| \mathcal{A}_{M \rightarrow M' l_\alpha^+ l_\beta^+}^{\text{LNV}} \right|^2 \propto \left| U_{\alpha 4} U_{\beta 4} f(m_4) + U_{\alpha 5} U_{\beta 5} f(m_5) \right|^2,$$

$$R_{l_\alpha l_\beta} \equiv \frac{\Gamma_{M \rightarrow M' l_\alpha^\pm l_\beta^\pm}^{\text{LNV}}}{\Gamma_{M \rightarrow M' l_\alpha^\pm l_\beta^\mp}^{\text{LNC}}}$$

$$U_{\alpha i} = e^{-i\phi_{\alpha i}} |U_{\alpha i}|$$

$$\psi_\alpha \equiv \phi_{\alpha 5} - \phi_{\alpha 4}$$



**Dirac limit**

$$\frac{\Delta M}{\Gamma_N} = 0$$

$$\psi_\alpha = \frac{\pi}{2}$$



# LHC SEARCHES

# LNV at LHC

Heavy neutrinos in pp collisions produced through a variety of mechanisms

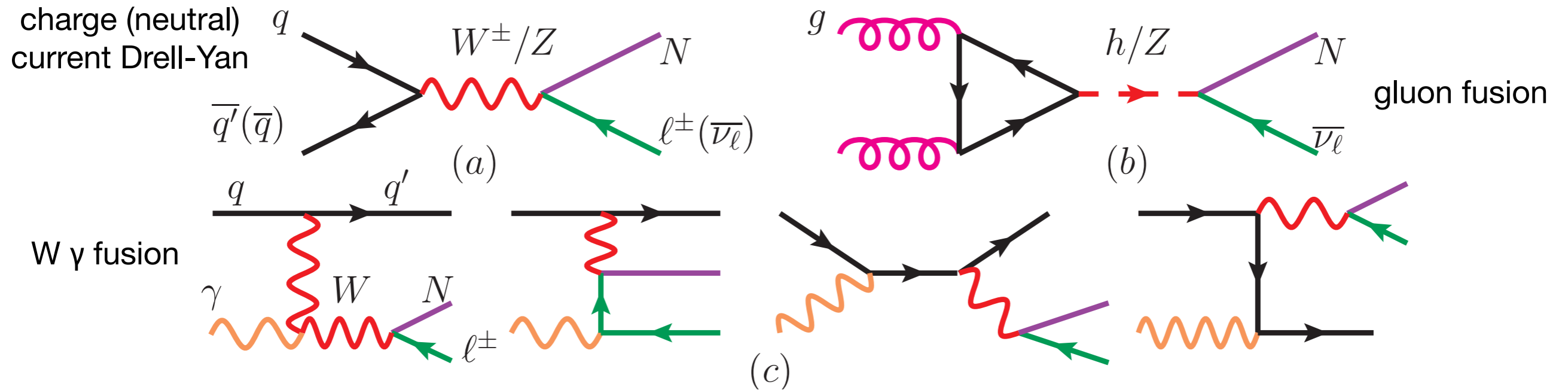


Figure from C. Degrande, O. Mattelaer, R. Ruiz and J. Turner, arXiv:1602.06957 [hep-ph]; see also Y. Cai, T. Han, T. Li and R. Ruiz, arXiv:1711.02180 [hep-ph]

**LNV can manifest with clean experimental signatures:**

e.g. two same-sign leptons (any flavour combination of e and  $\mu$ ) and at least one jet

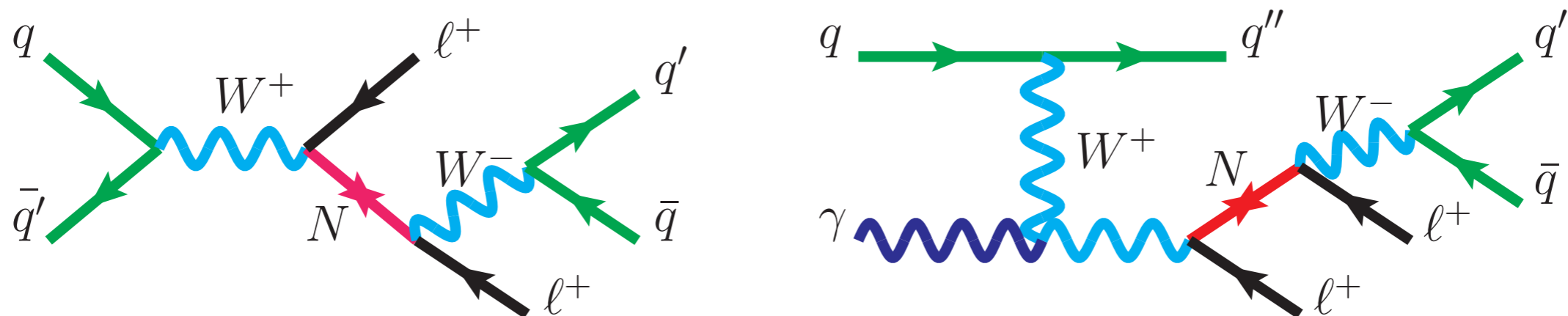


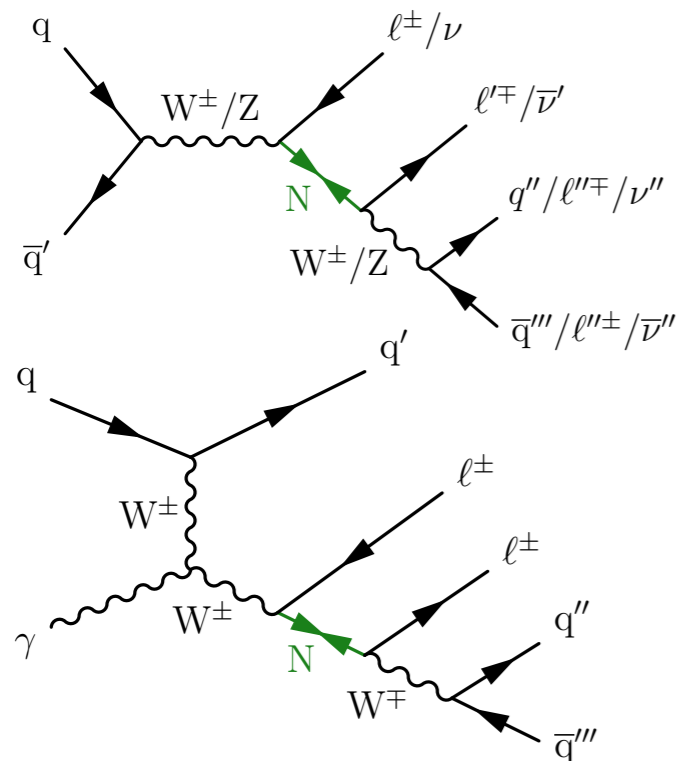
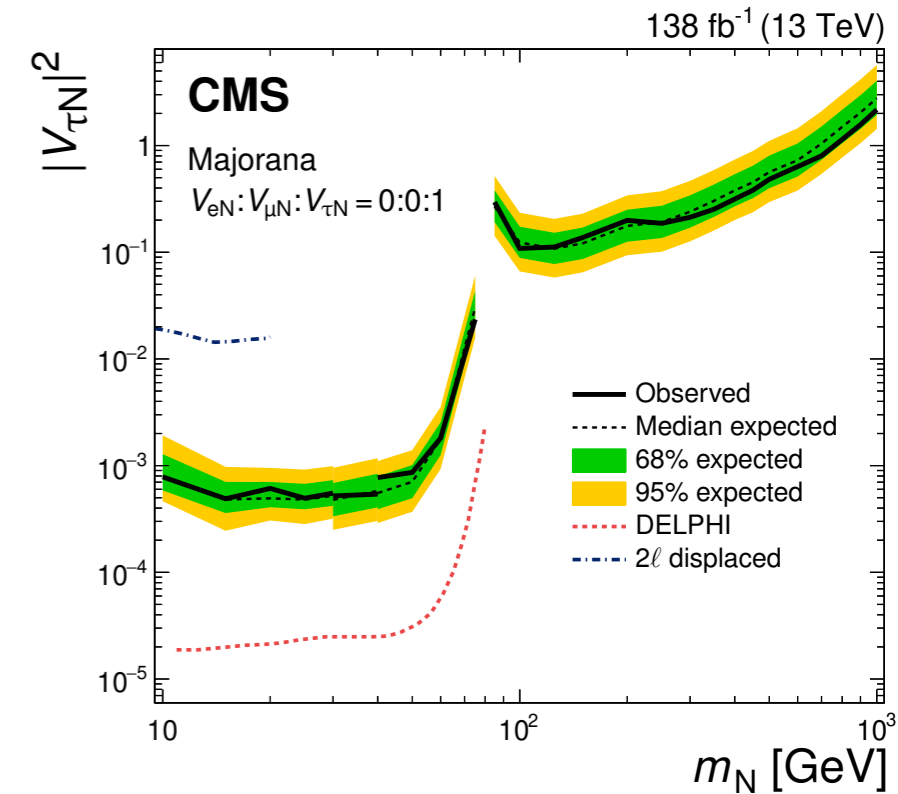
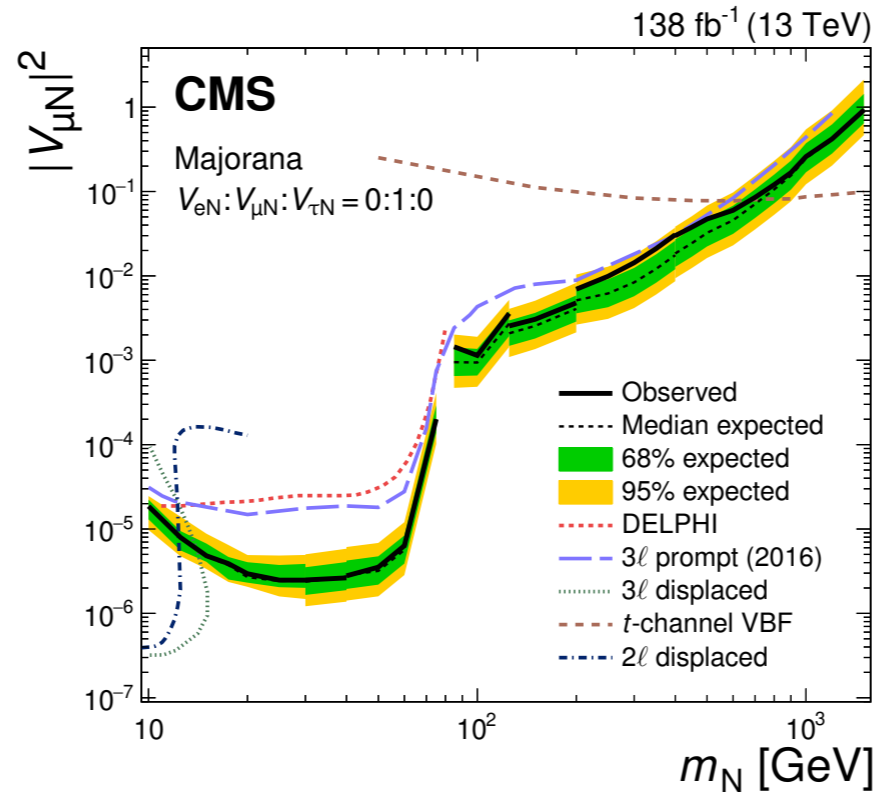
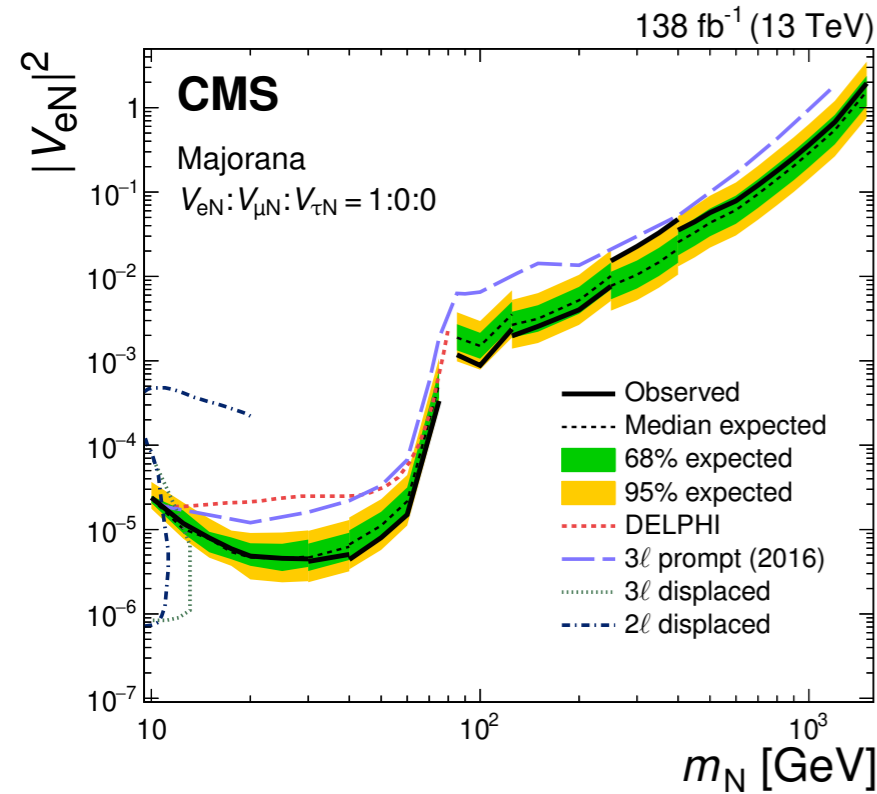
Figure from CMS Collaboration, arXiv:1806.10905 [hep-ex]

# Prompt decays @ CMS

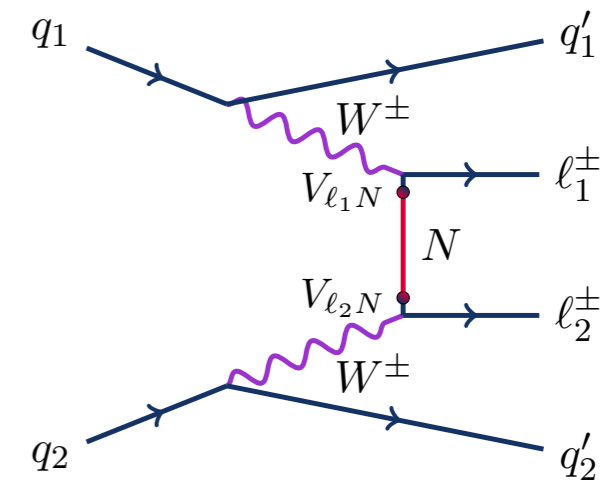
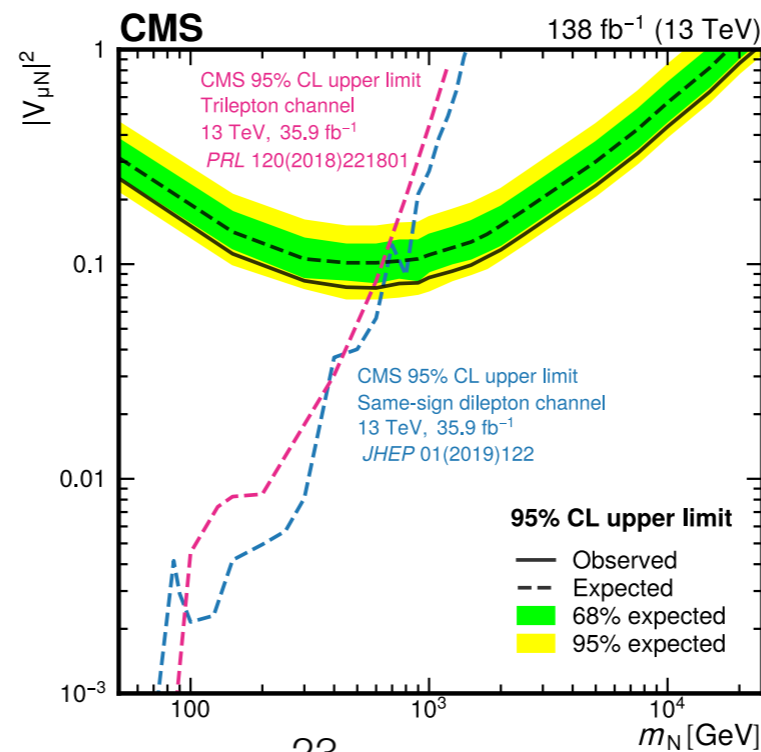
CMS collaboration, arXiv:2405.17605 [hep-ex]

## Prompt decays

arXiv:2403.00100 [hep-ex]



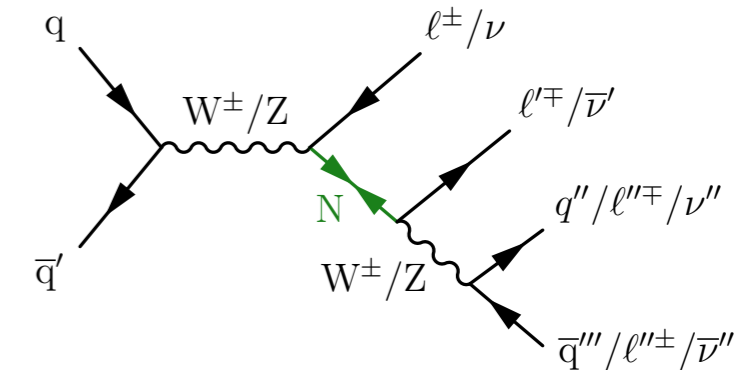
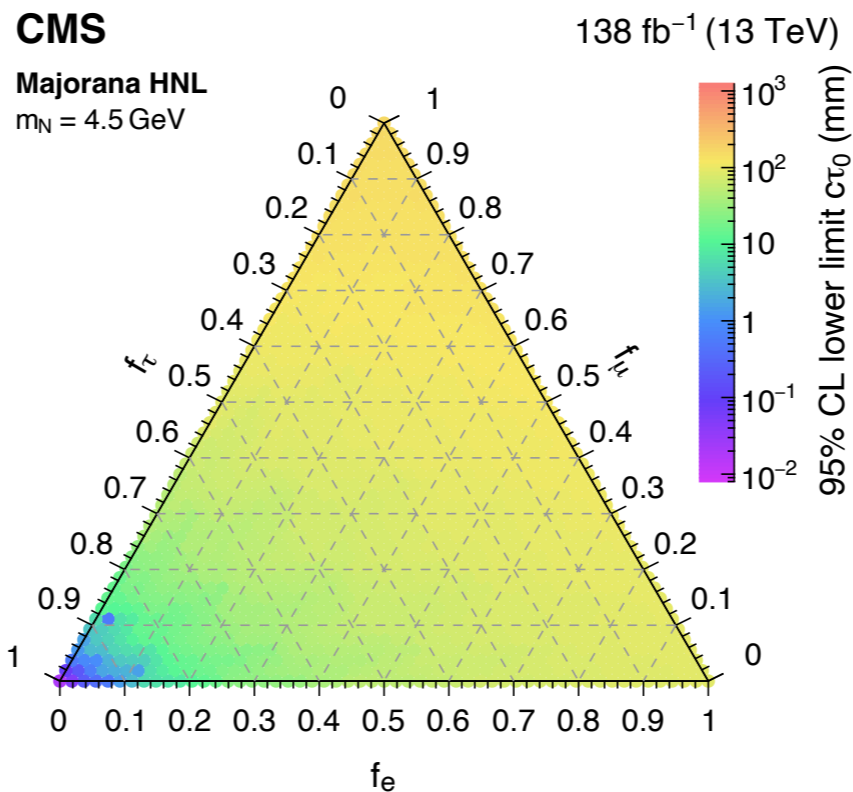
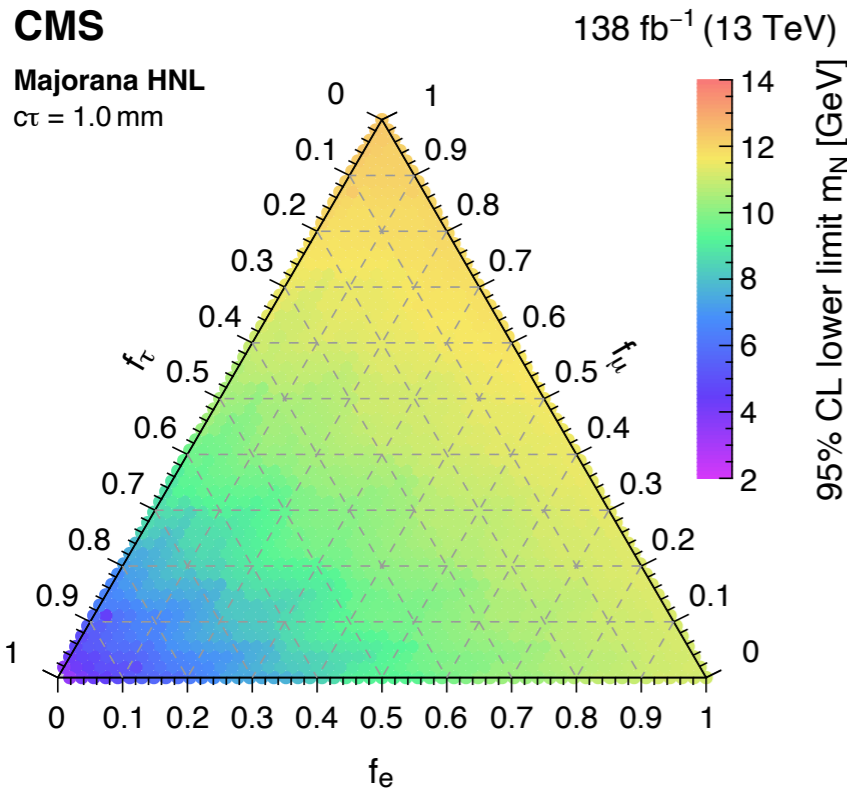
## Vector boson fusion



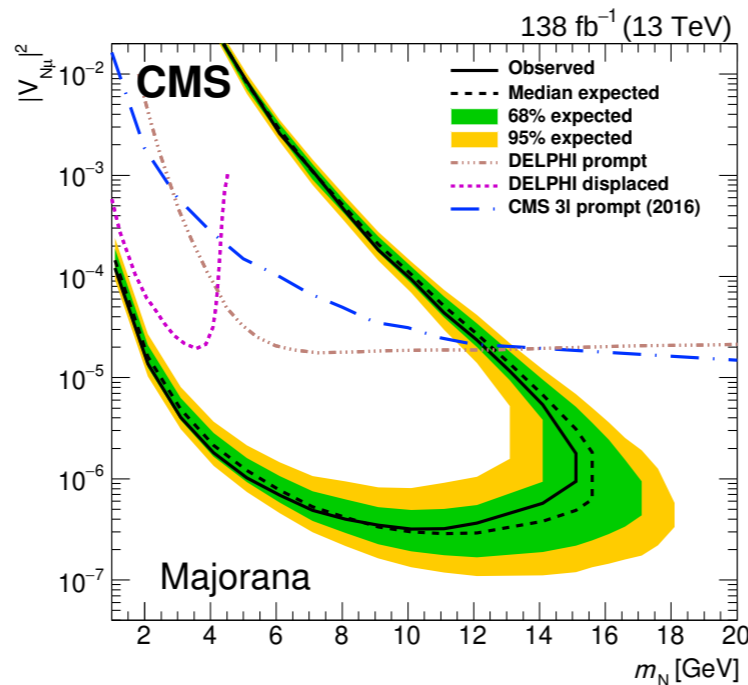
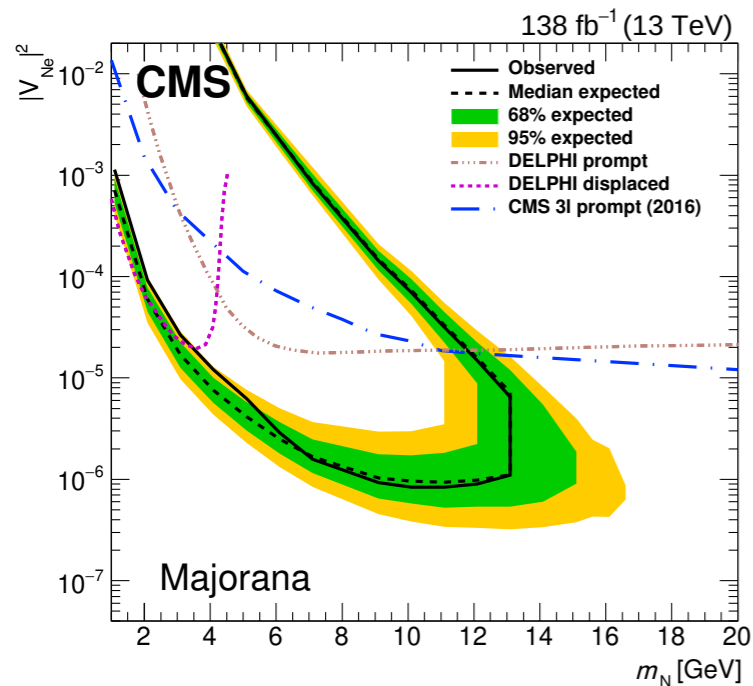
arXiv:2206.08956 [hep-ex]

# Long-lived HNL @ CMS

CMS collaboration, arXiv:2405.17605 [hep-ex]

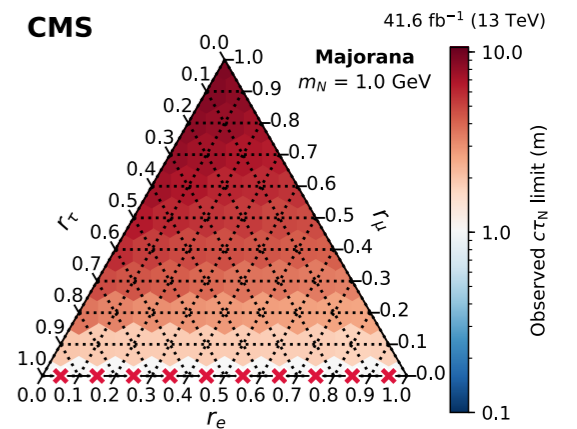
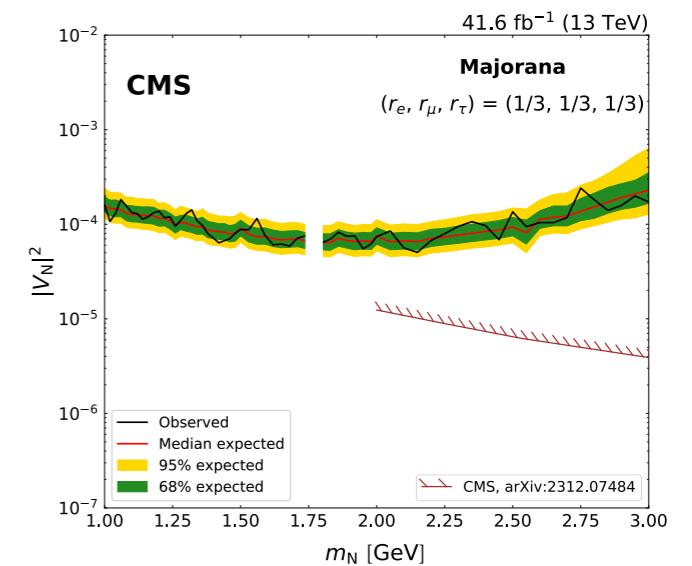


arXiv:2312.07484 [hep-ex]



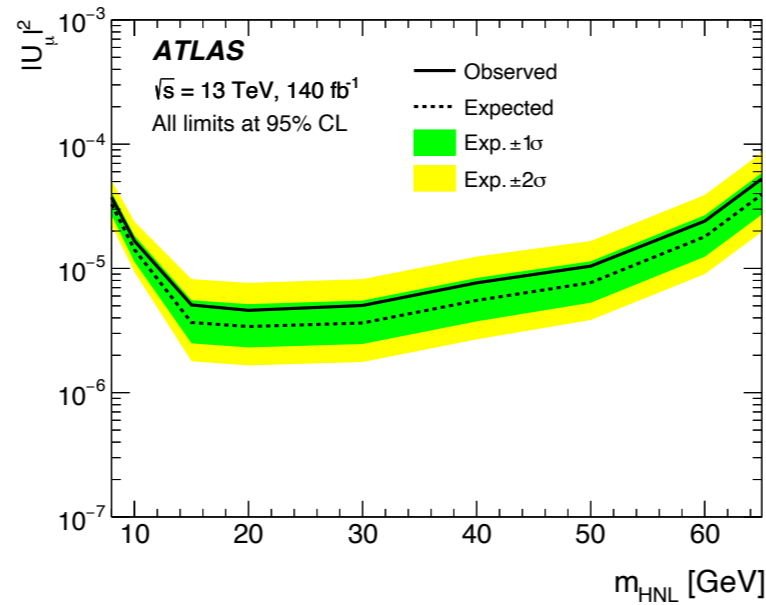
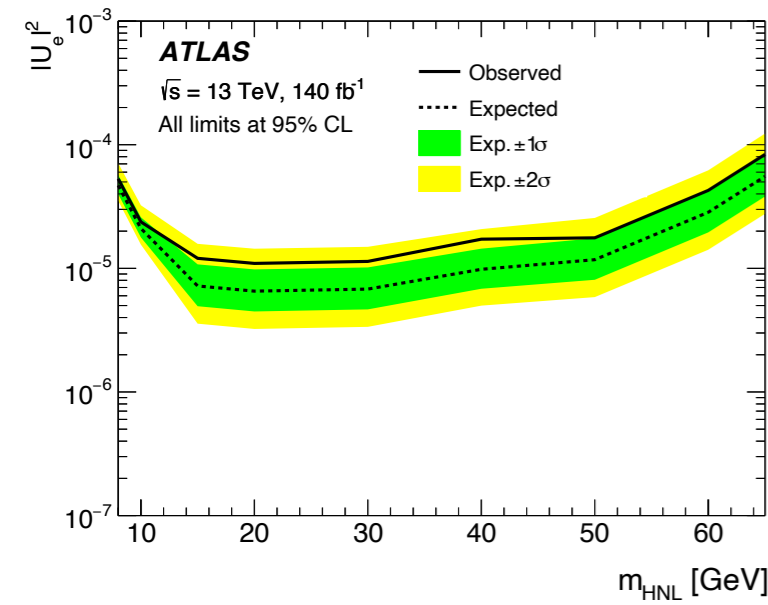
arXiv:2201.05578 [hep-ex]

## B meson decay

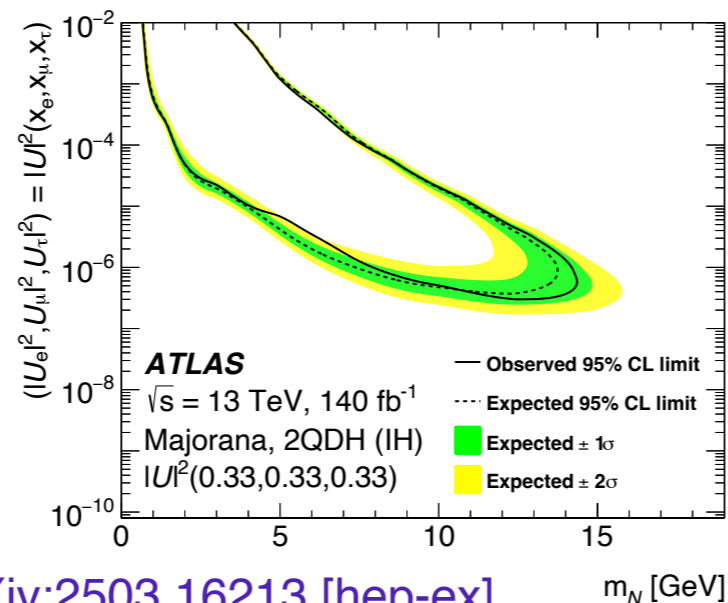
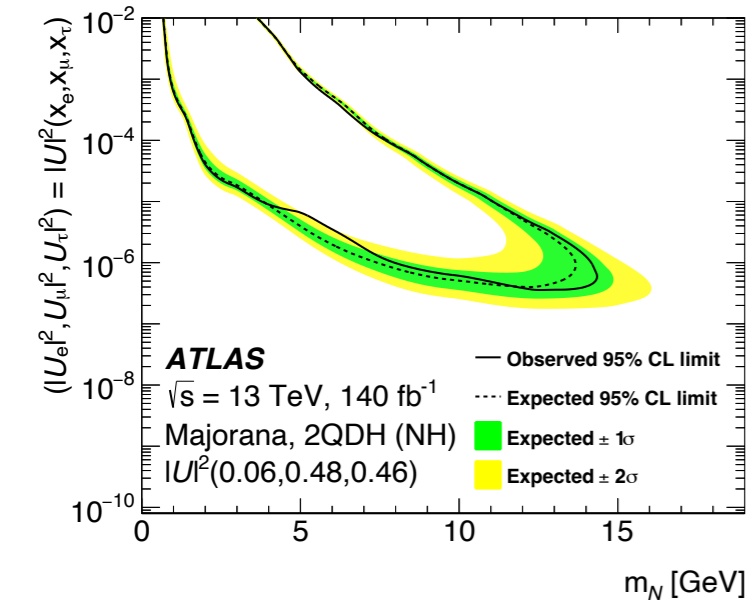
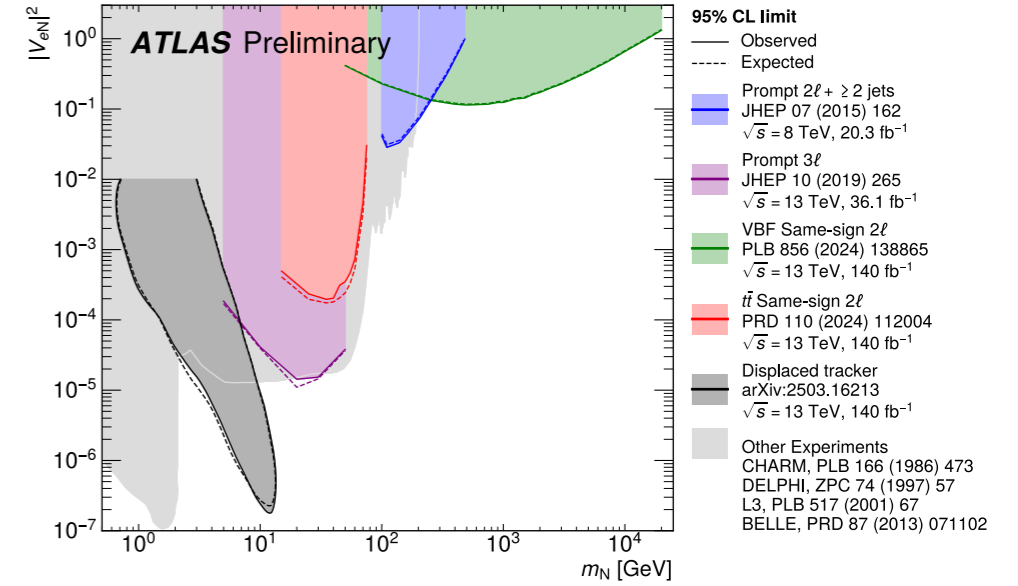


arXiv:2403.04584 [hep-ex]

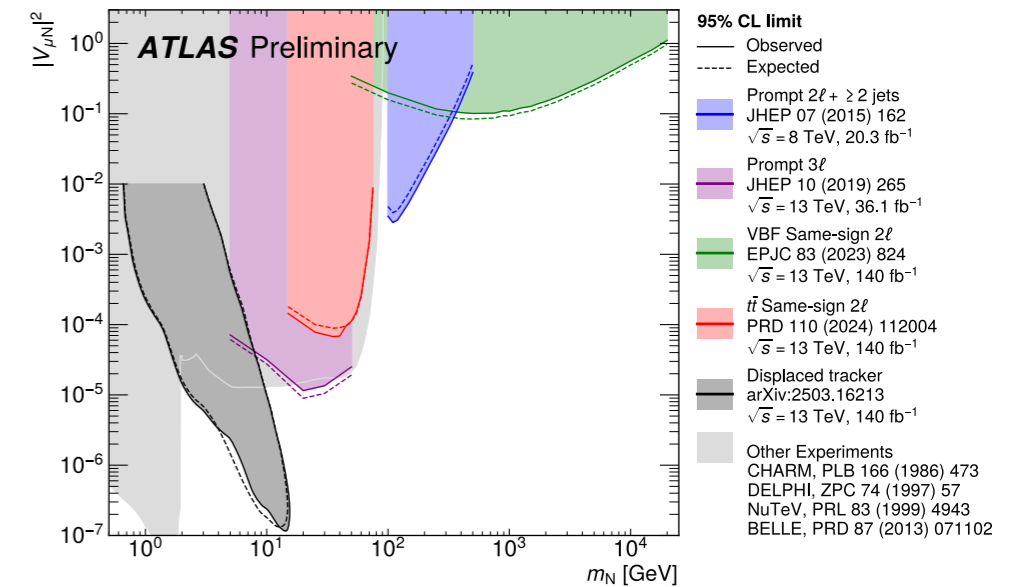
# HNL ATLAS searches



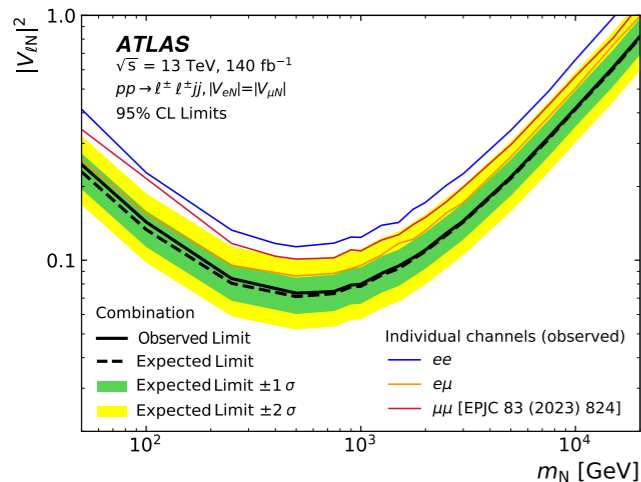
arXiv:2508.20929 [hep-ex]



arXiv:2503.16213 [hep-ex]



ATL-PHYS-PUB-2025-008



PoS LHCP2024 (2025), 032

# LNV/LNC oscillations

Y. Nir, Conf. Proc. C9207131, 81 (1992); G. Anamiati, M. Hirsch and E. Nardi, arXiv:1607.05641 [hep-ph]

**Flavour eigenstate = coherent superposition of mass eigenstates**

$$\left\{ \begin{array}{l} N_\ell = \frac{1}{\sqrt{2}}(N_+ - iN_-) \\ N_{\bar{\ell}} = \frac{1}{\sqrt{2}}(N_+ + iN_-) \end{array} \right. \xrightarrow{\text{evolution}} \left\{ \begin{array}{l} N_\ell(t) = g_+(t)N_\ell + g_-(t)N_{\bar{\ell}} \\ N_{\bar{\ell}}(t) = g_-(t)N_\ell + g_+(t)N_{\bar{\ell}} \end{array} \right.$$

$$g_+(t) = e^{-iMt} e^{-\frac{\Gamma}{2}t} \cos\left(\frac{\Delta M}{2}t\right)$$

$$g_-(t) = i e^{-iMt} e^{-\frac{\Gamma}{2}t} \sin\left(\frac{\Delta M}{2}t\right)$$

$$\Delta M = M^+ - M^-$$

## Timescales

$\Delta M \gg \Gamma$  **decay after decoherence (Majorana limit)**

$\Delta M \approx \Gamma$  **oscillations**

$\Delta M \ll \Gamma$  **oscillations do not develop (Dirac limit)**

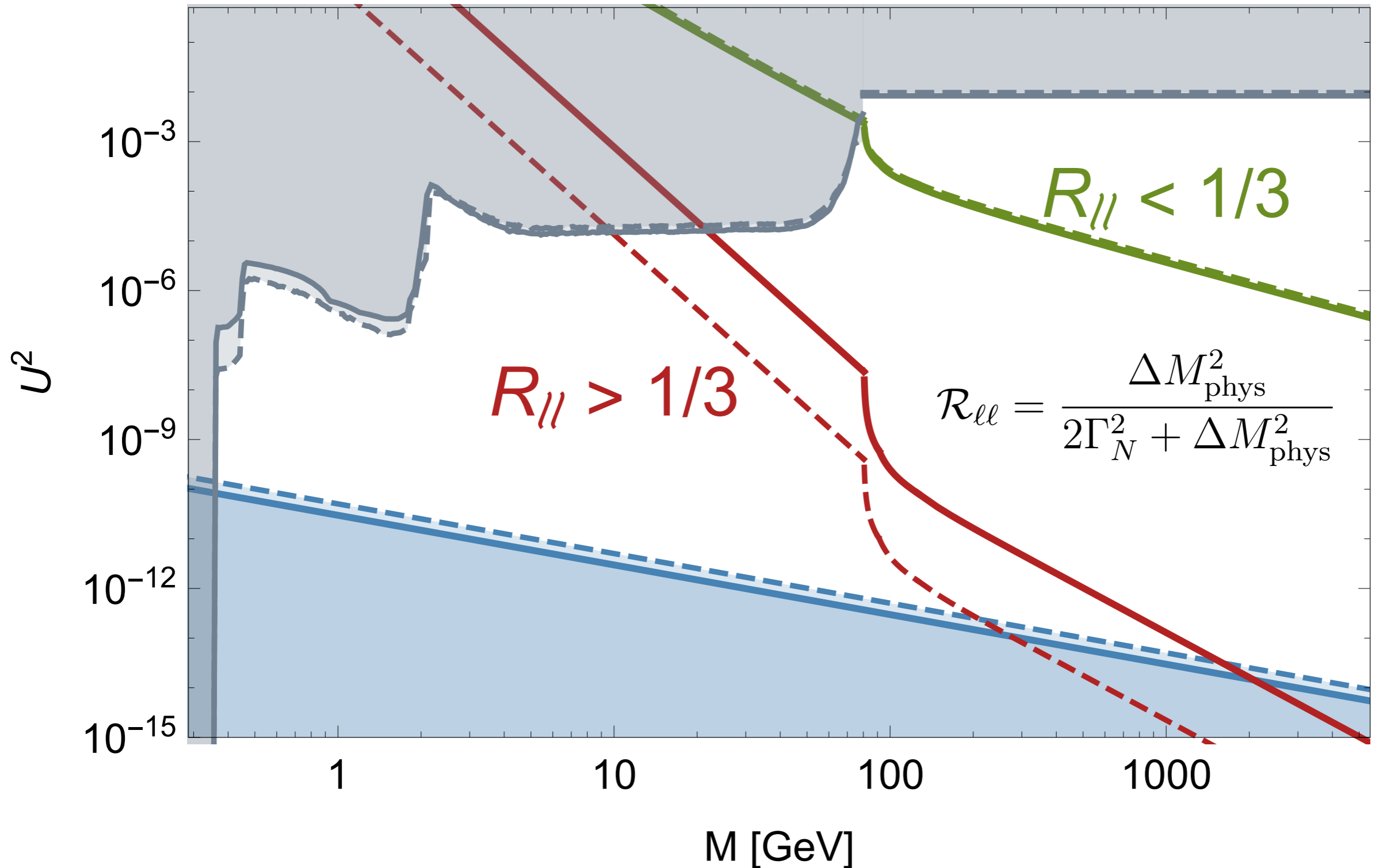
$$R_{\ell\ell}(t_1, t_2) = \frac{\int_{t_1}^{t_2} |g_-(t)|^2 dt}{\int_{t_1}^{t_2} |g_+(t)|^2 dt} = \frac{\#(\ell^+ \ell^+) + \#(\ell^- \ell^-)}{\#(\ell^+ \ell^-)}$$

$$R_{ll}(0, \infty) = \frac{\Delta M^2}{2\Gamma^2 + \Delta M^2}$$



# LNV observability at colliders

Type-I seesaw with  $n = 2$  RHN reproducing neutrino oscillation data



M. Drewes, J. Klarić and P. Klose, arXiv:1907.13034 [hep-ph]

# ADDITIONAL SEARCHES

# A wealth of opportunities

**Neutrino Masses and Mixing: Current Status and Open Questions**

*Yufeng Li*

14:30 - 14:55

**Long-Baseline Neutrino Experiments: A Journey from Discovery to Precision and Beyond**

*Lucio Ludovici*

09:00 - 09:25

**Neutrinos in the Multi-Messenger Era**

*Piera Sapienza*



15:45 - 16:10

**Status of the Hyper-Kamiokande experiment**

*Jan Kisiel*



14:55 - 15:20

**The DUNE Experiment: Status and Outlook**

*Jianming Bian et al.*

15:20 - 15:45

**JUNO status and prospects**

*Jie ZHAO*

10:15 - 10:40

**FASERv2, FLArE and other future nu exp. at collider**

*Wenjie Wu*

15:45 - 16:10

**LFV/LFU theory**

*Marco Ardu*



16:40 - 17:05

**Muon LFV/LFU measurements at JPARC, PSI, FNAL**

*Yusuke Uchiyama*



17:05 - 17:30

**Tau LFV/LFU measurements**

*Fabian Becherer*



17:30 - 17:55

# CONNECTION TO FURTHER BSM QUESTIONS

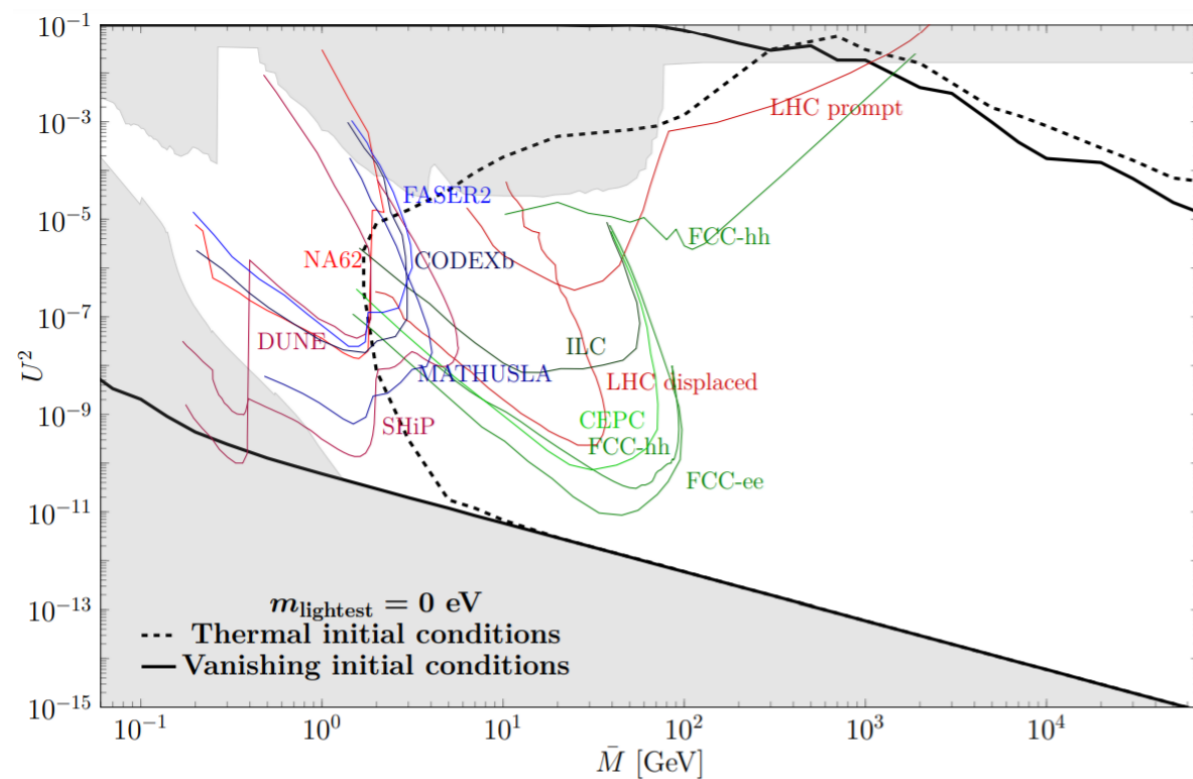
# One stone, three birds

Other than  $\nu$  masses, HNL can provide a solution to other SM observational issues

**New source of CP violation**



**Leptogenesis**

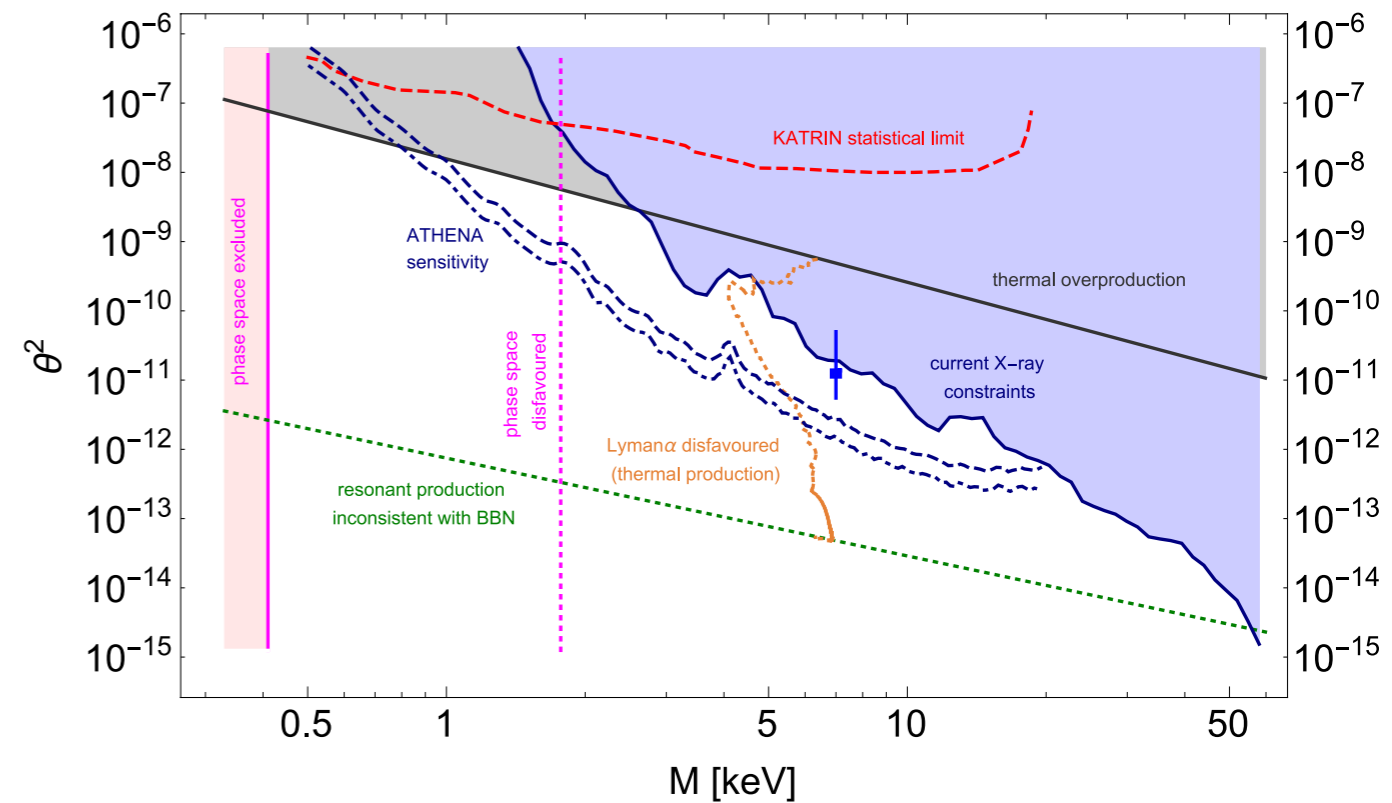


M. Drewes, Y. Georis and J. Klarić,  
arXiv:2106.16226 [hep-ph]

**Long-lived particles**

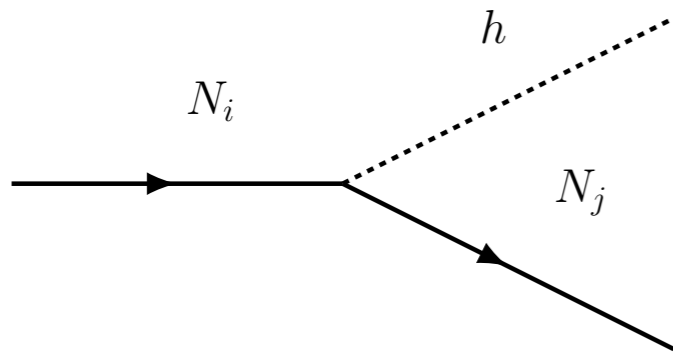


**Dark matter**



A. Boyarsky, M. Drewes, T. Lasserre, S. Mertens  
and O. Ruchayskiy, arXiv:1807.07938 [hep-ph]

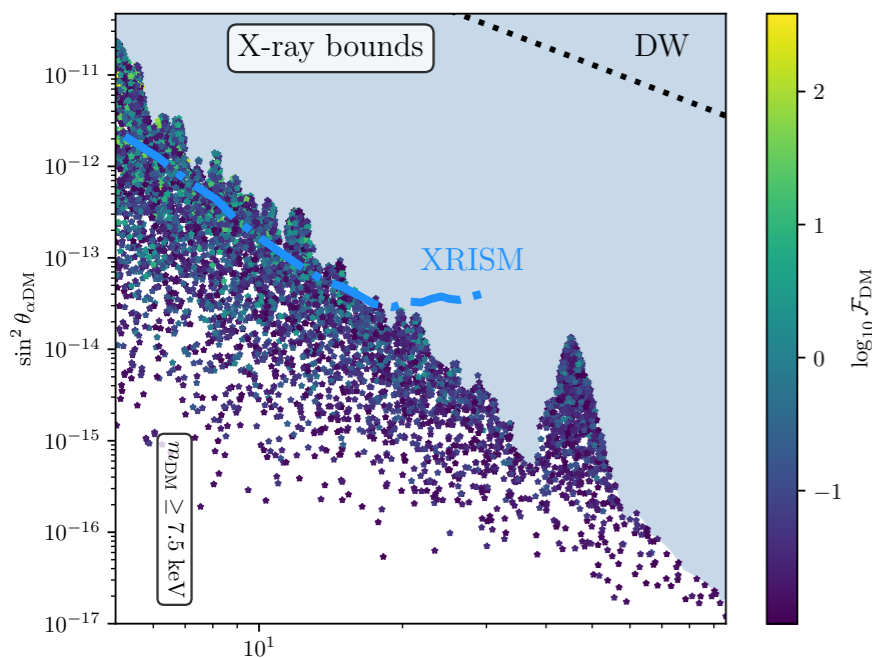
# A new testable dark matter solution



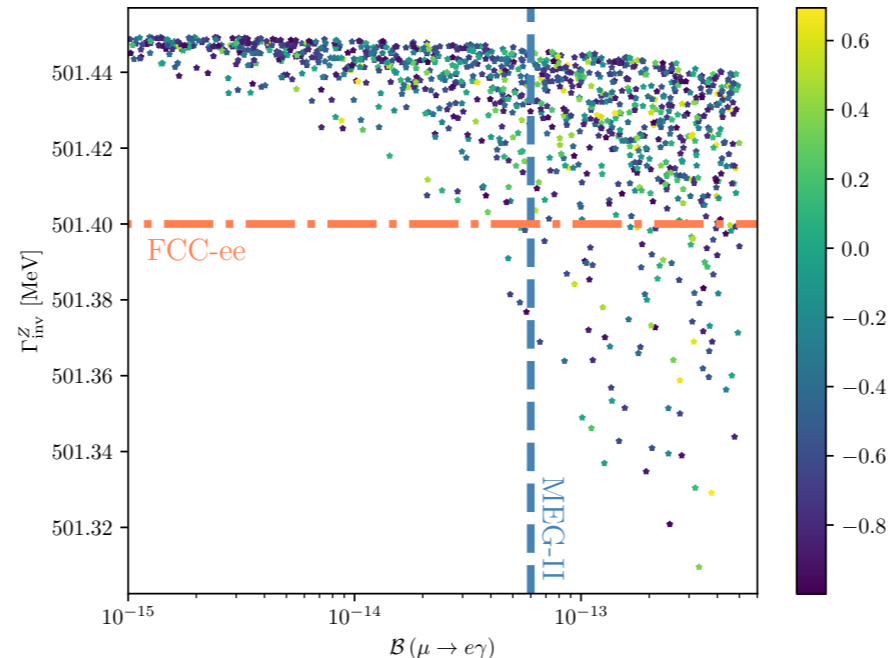
We computed the thermal correction in the production of DM from the decay of heavy states

The DM solutions exist and are testable by different experimental searches

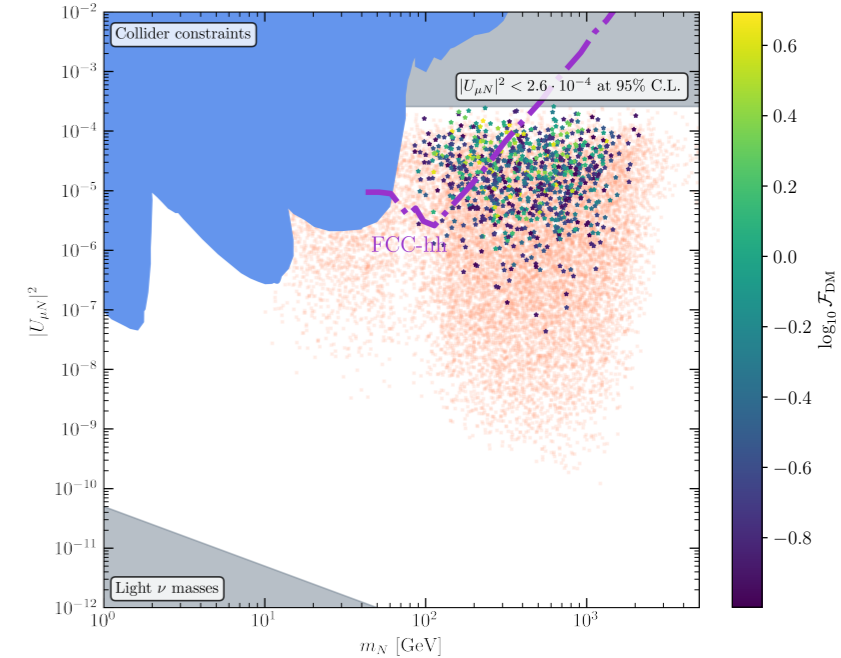
## Indirect



## Precision



## Collider



A. Abada, G. Arcadi, M. Lucente and S. Rosauero-Alcaraz, arXiv:2503.20017 [hep-ph]; A. Abada, G. Arcadi, M. Lucente, G. Piazza and S. Rosauero-Alcaraz, arXiv:2308.01341 [hep-ph]; M. Lucente, arXiv:2103.03253 [hep-ph]

# Conclusion

**Massive neutrinos call for BSM physics**

**Multiple possibilities, we focus on the simple RHN extension**

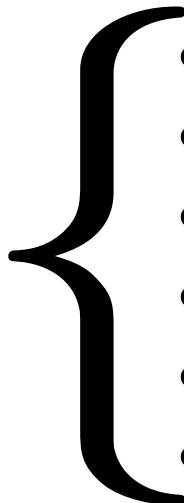
**HNL phenomenology is generally connected with  $\nu$  mass generation mechanism**

**Lepton number symmetry allows for low scale NP and sizeable couplings**

**But this symmetry generally suppresses LNV processes**

**LNV rates depend in general on the interference of multiple virtual states**

**Multiple testability opportunities**

- 
- **neutrinoless  $2\beta$  decay**
  - **meson and tau decay**
  - **collider and fixed target**
  - **LFV/LNV and precision**
  - **neutrino telescopes**
  - **...**

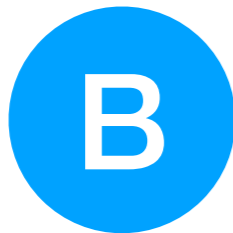
**HNL can also provide simultaneous solution to the BAU and DM problems**

# Backup



# Accidental symmetries of the SM

The Standard Model has accidental perturbative symmetries, arising from:  
**gauge group + field content + renormalizability**



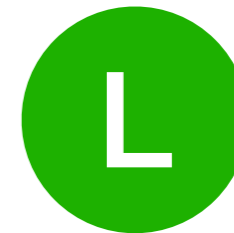
**Baryon number**

*(Individual quark flavour numbers  
are violated by CKM mixing)*



**Flavour numbers**

$\alpha = e, \mu, \tau$



**Lepton number**

$L = \sum_{\alpha} L_{\alpha}$

**Non perturbative effects violate both B and L, but preserve**



$$\partial_{\mu} J_B^{\mu} = \partial_{\mu} J_L^{\mu} = \frac{N_f}{32\pi^2} \epsilon^{\mu\nu\sigma\tau} \left( -g_W^2 \text{Tr} W_{\mu\nu} W_{\sigma\tau} + g_Y^2 B_{\mu\nu} B_{\sigma\tau} \right)$$

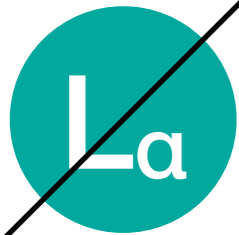
G. 't Hooft, Phys. Rev. Lett. 37 (1976) 8; Phys. Rev. D 14 (1976) 3432

# Accidental symmetries: experimental status



**No evidence of violation**

E.g. proton mean life  $> 3.6 \times 10^{29}$  years CL=90%  
PDG, Prog. Theor. Exp. Phys. 2022, 083C01 (2022)



**Violated in neutrino oscillations**



**New physics BSM**

$$|U|_{3\sigma}^{\text{with SK-atm}} = \begin{pmatrix} 0.801 \rightarrow 0.845 & 0.513 \rightarrow 0.579 & 0.144 \rightarrow 0.156 \\ 0.244 \rightarrow 0.499 & 0.505 \rightarrow 0.693 & 0.631 \rightarrow 0.768 \\ 0.272 \rightarrow 0.518 & 0.471 \rightarrow 0.669 & 0.623 \rightarrow 0.761 \end{pmatrix}$$

I. Esteban, M. C. Gonzalez-Garcia, M. Maltoni, T. Schwetz and A. Zhou, arXiv:2007.14792 [hep-ph]



**No evidence of violation**

Massive neutrinos violate it if they are Majorana particles

# Fermionic singlet extensions of the SM

**SM +  $n$  gauge singlet fermions  $N_I$**

$$\mathcal{L} = \mathcal{L}_{\text{SM}} + i\overline{N}_I \not{\partial} N_I - \left( F_{\alpha I} \overline{\ell}_L^\alpha \tilde{\phi} N_I + \frac{M_{IJ}}{2} \overline{N}_I^c N_J + h.c. \right)$$

$\nearrow$   $3 \times n$  matrix  
Yukawa couplings

$\nwarrow$   $n \times n$  matrix  
Majorana mass couplings

After electroweak phase transition  $\langle \Phi \rangle = v \approx 174$  GeV

$$-\mathcal{L}_m^\nu = \frac{1}{2} \begin{pmatrix} \overline{\nu}_L & \overline{N}^c \end{pmatrix} \underbrace{\begin{pmatrix} \delta m_\nu^{\text{loop}} & vF \\ vF^T & M \end{pmatrix}}_{\mathcal{M}} \begin{pmatrix} \nu_L^c \\ N \end{pmatrix} + h.c.$$

$\nwarrow$   $(3+n)$  dimensional  
mass matrix

# Phenomenology of fermionic singlets

$$\mathcal{U}^T \mathcal{M} \mathcal{U} = \hat{\mathcal{M}}_{\text{diag}} \quad \rightarrow \quad \begin{cases} 3 \text{ light (mostly active) states} \\ n \text{ heavy (mostly sterile) states} \end{cases}$$

PMNS matrix:  
neutrino oscillations

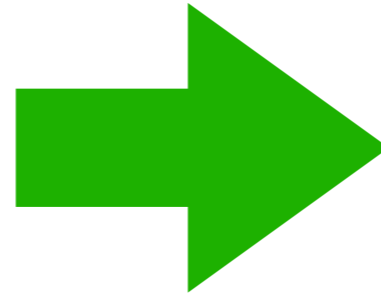
Couples the heavy states  
with SM gauge bosons

$$\mathcal{U} = \begin{pmatrix} \mathcal{U}^{\alpha, i=1,2,3}_{\text{active-active}} & \mathcal{U}^{\alpha, i \geq 4}_{\text{active-sterile}} \\ \vdots & \ddots \end{pmatrix}$$

↑  
Unobservable

# Accidental cancellations: quantify fine tuning

If a symmetry is present in the Lagrangian, it will be manifest at any order in perturbation theory



The **neutrino mass scale** is **stable** under **radiative corrections**

Compute neutrino masses  $m_\nu$  at 1-loop, and quantify the level of fine-tuning of a solution as

$$f.t.(m_\nu) = \sqrt{\sum_{i=1}^3 \left( \frac{m_i^{\text{loop}} - m_i^{\text{tree}}}{m_i^{\text{loop}}} \right)^2}$$

$m_i^{\text{loop}}$

1-loop neutrino mass spectrum

$m_i^{\text{tree}}$

tree-level neutrino mass spectrum

# Double beta decay

2 $\beta$  decay: 2<sup>nd</sup> order weak process  $\mathcal{N}(A, Z) \rightarrow \mathcal{N}(A, Z + 2) + 2e^- + 2\bar{\nu}_e$

**Only relevant when the single  $\beta$  decay is kinematically forbidden**  $^{48}\text{Ca}, ^{76}\text{Ge}, ^{82}\text{Se}, ^{96}\text{Zr}, ^{100}\text{Mo}, ^{116}\text{Cd}, ^{130}\text{Te}, ^{136}\text{Xe}, ^{150}\text{Nd}$

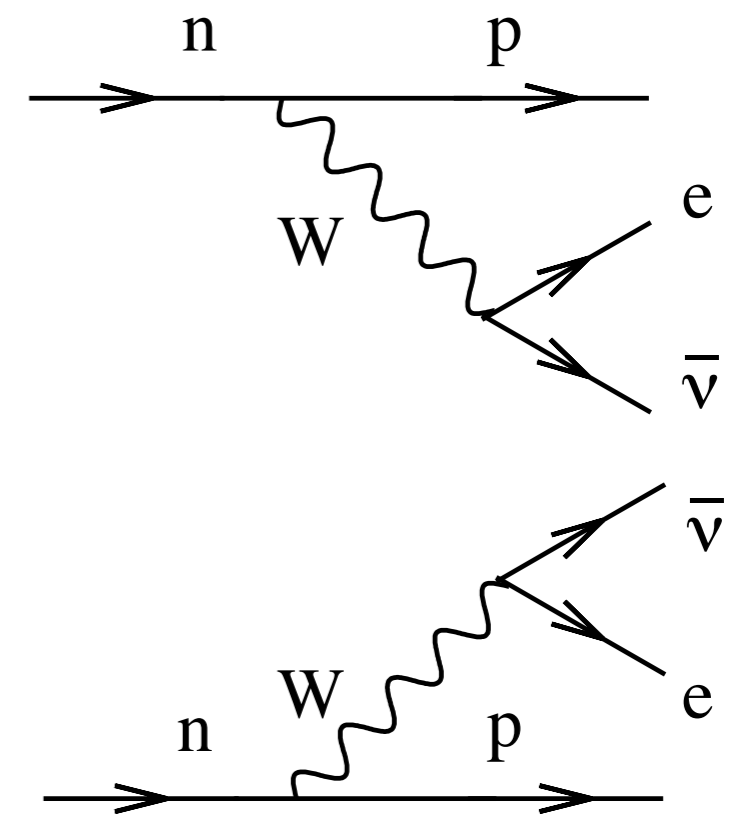
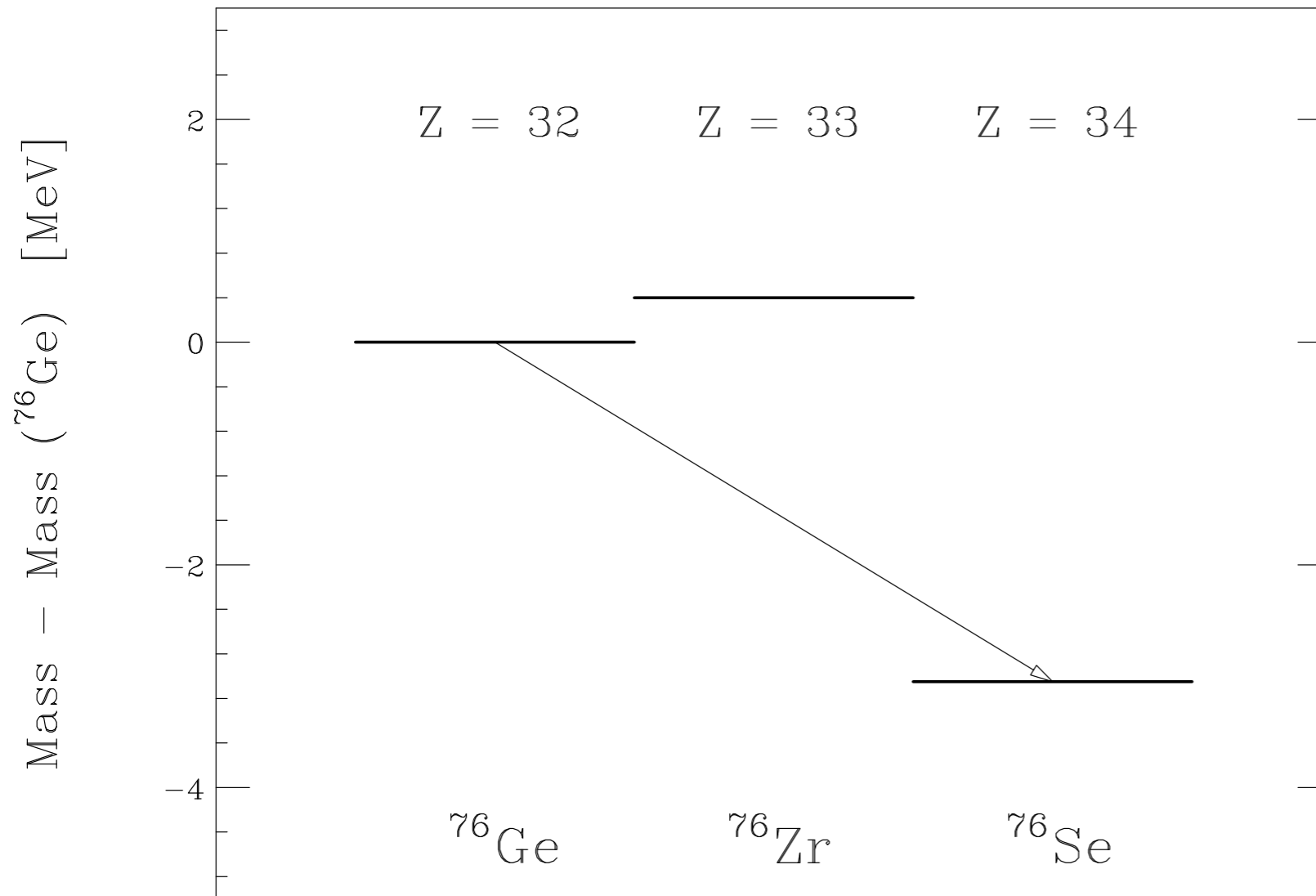


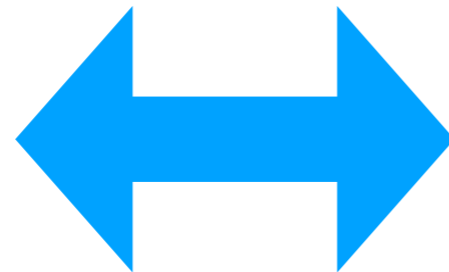
Figure from P. Lipari, Introduction to neutrino physics, in 2001 CERN-CLAF School of high-energy physics

# The black box theorem

J. Schechter and J. W. F. Valle, Phys. Rev. D 25 (1982) 2951; E. Takasugi, Phys. Lett. 149B (1984) 372;  
 see also M. Duerr, M. Lindner and A. Merle, arXiv:1105.0901 [hep-ph]

Non-vanishing  
 $0\nu 2\beta$  amplitude

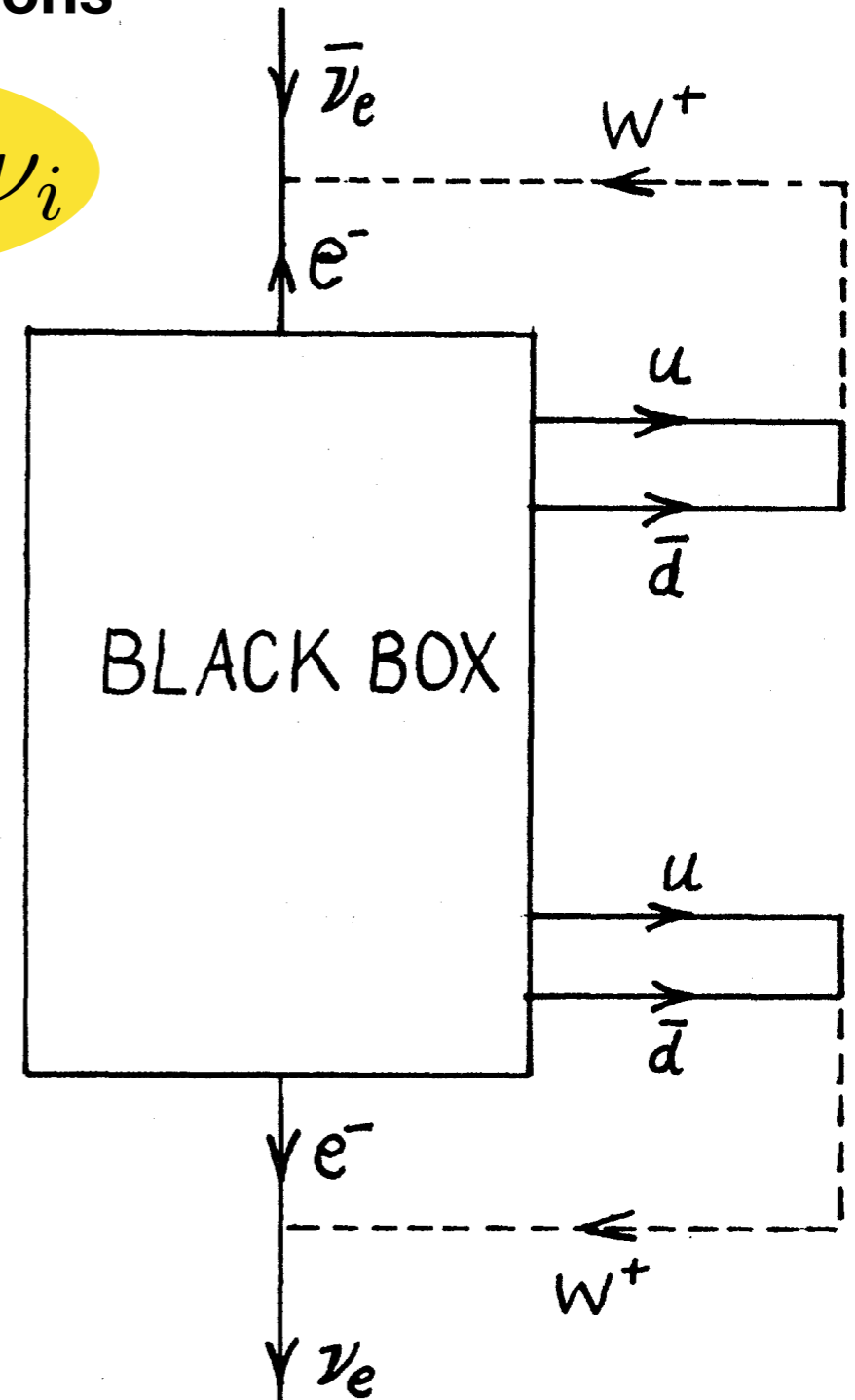
$$\Gamma_{0\nu 2\beta} \neq 0$$



Neutrinos are  
 Majorana fermions

$$\nu_i^c = e^{i\phi} \nu_i$$

Irrespectively of the underlying mechanism, a non-vanishing  $0\nu 2\beta$  amplitude generates a Majorana mass term for the SM neutrinos

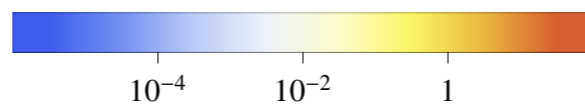




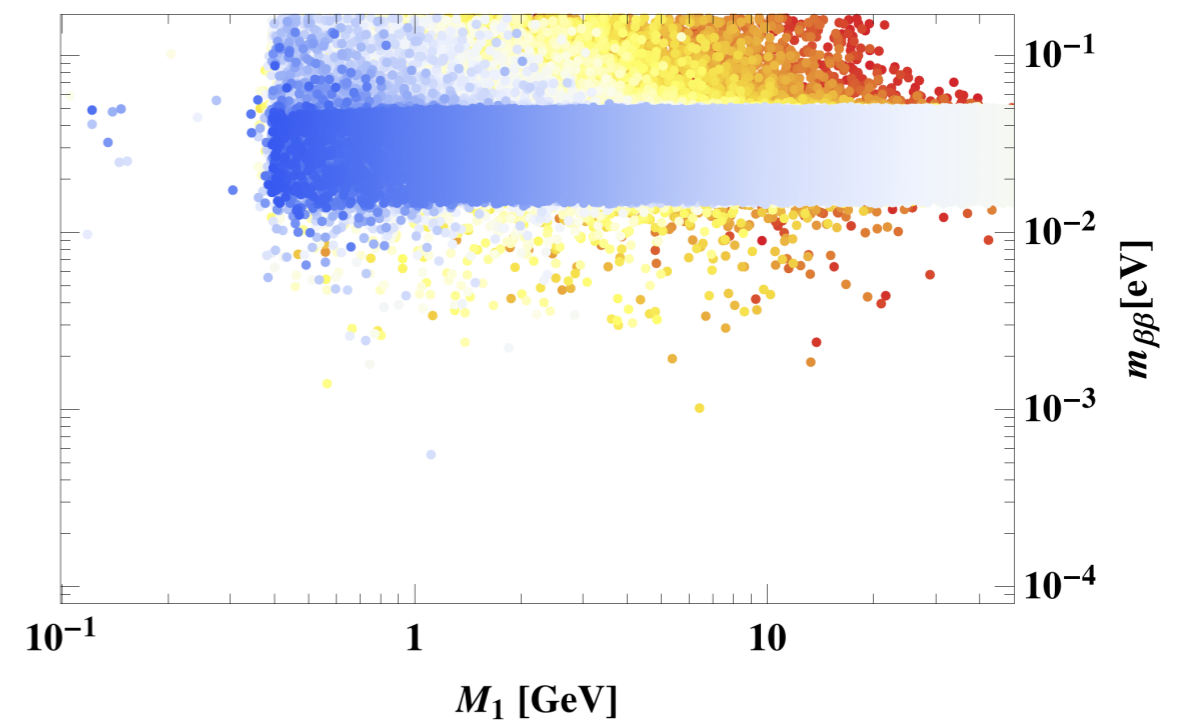
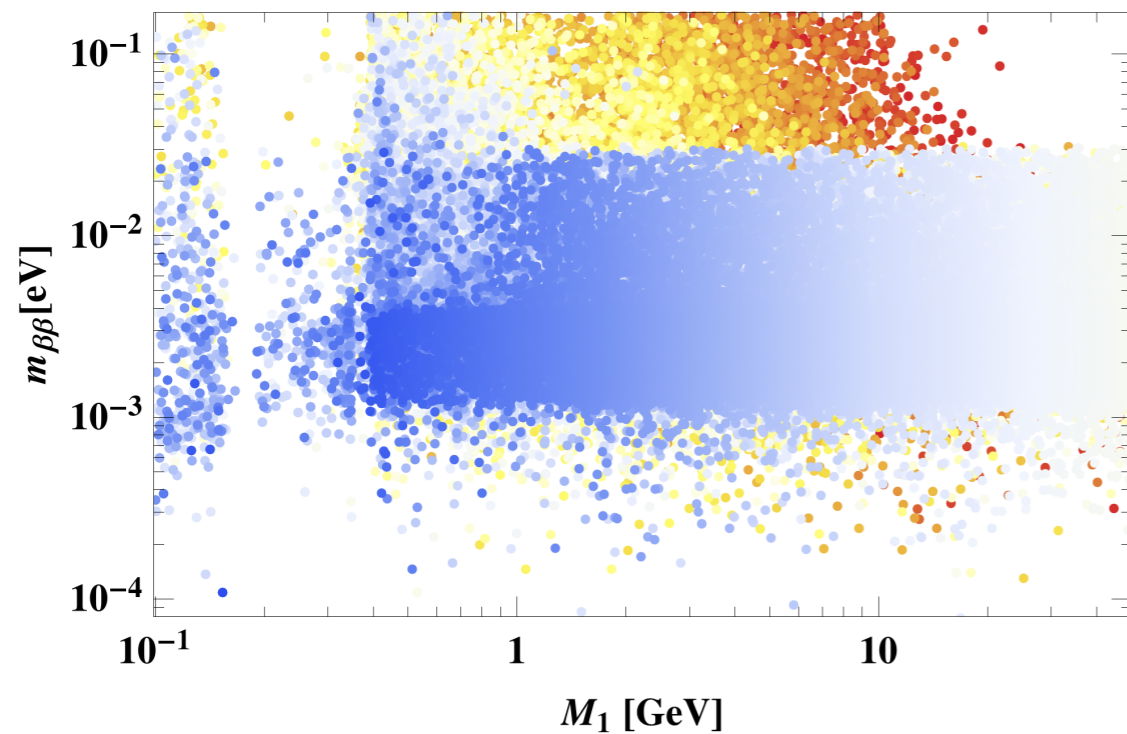
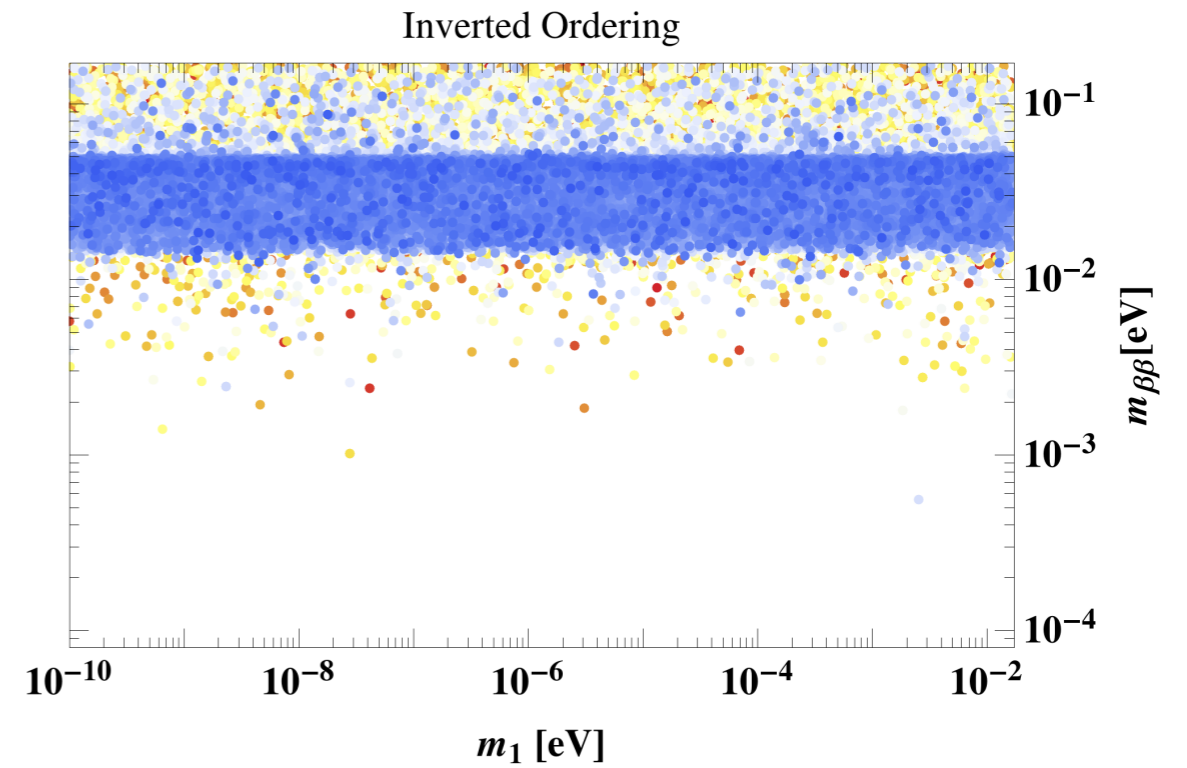
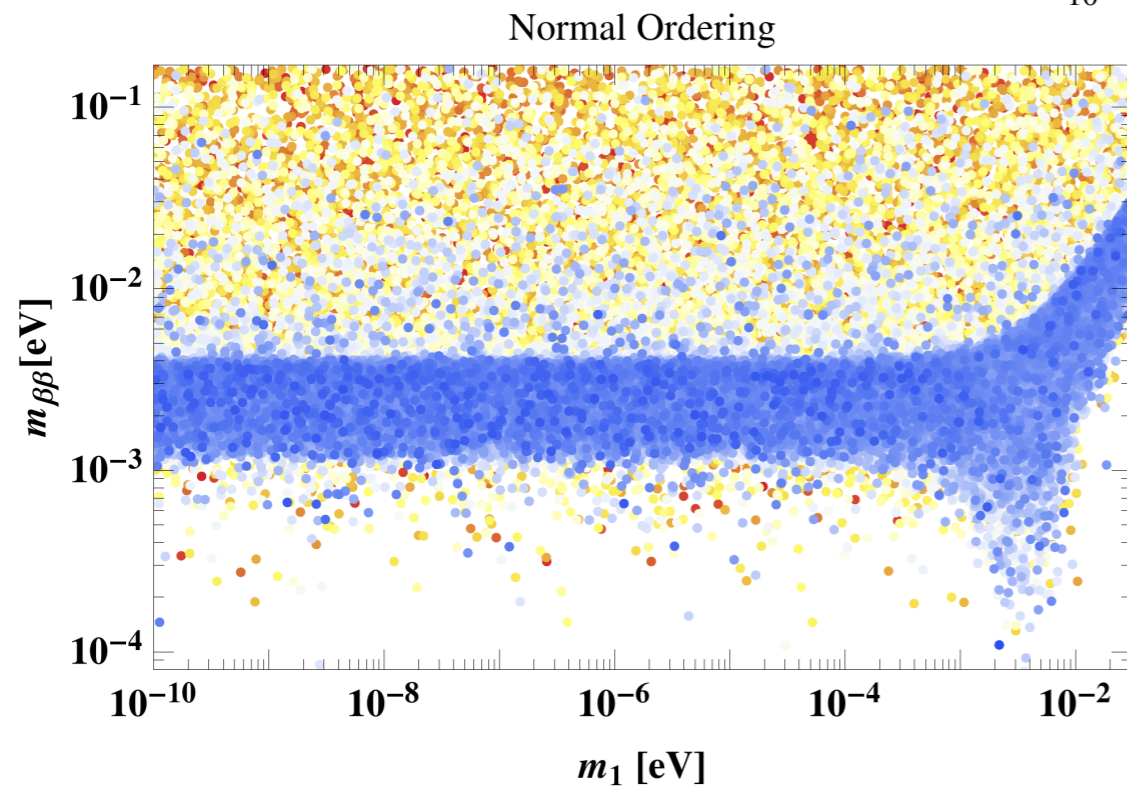
# Heavy neutrinos at GeV scale

f.t.

Blue points: not fine tuned



Red points: fine tuned



Figures from A. Abada, G. Arcadi, V. Domcke, M. Drewes, J. Klaric and M. Lucente, arXiv:1810.12463 [hep-ph];  
 see also J. Lopez-Pavon, S. Pascoli and C. f. Wong, arXiv:1209.5342 [hep-ph]; J. Lopez-Pavon, E. Molinaro  
 and S. T. Petcov, arXiv:1506.05296 [hep-ph]



# Meson and tau LNV: current bounds

Tables (and list of references) from A. Abada, V. De Romeri, M.L., A. M. Teixeira and T. Toma, arXiv:1712.03984 [hep-ph]

## Meson decay

LNV decay	Current bound		
	$l_\alpha = e, l_\beta = e$	$l_\alpha = e, l_\beta = \mu$	$l_\alpha = \mu, l_\beta = \mu$
$K^- \rightarrow l_\alpha^- l_\beta^- \pi^+$	$6.4 \times 10^{-10}$ [41]	$5.0 \times 10^{-10}$ [41]	$1.1 \times 10^{-9}$ [41]
$D^- \rightarrow l_\alpha^- l_\beta^- \pi^+$	$1.1 \times 10^{-6}$ [41]	$2.0 \times 10^{-6}$ [78]	$2.2 \times 10^{-8}$ [79]
$D^- \rightarrow l_\alpha^- l_\beta^- K^+$	$9.0 \times 10^{-7}$ [78]	$1.9 \times 10^{-6}$ [78]	$1.0 \times 10^{-5}$ [78]
$D^- \rightarrow l_\alpha^- l_\beta^- \rho^+$	—————	—————	$5.6 \times 10^{-4}$ [41]
$D^- \rightarrow l_\alpha^- l_\beta^- K^{*+}$	—————	—————	$8.5 \times 10^{-4}$ [41]
$D_s^- \rightarrow l_\alpha^- l_\beta^- \pi^+$	$4.1 \times 10^{-6}$ [41]	$8.4 \times 10^{-6}$ [78]	$1.2 \times 10^{-7}$ [79]
$D_s^- \rightarrow l_\alpha^- l_\beta^- K^+$	$5.2 \times 10^{-6}$ [78]	$6.1 \times 10^{-6}$ [78]	$1.3 \times 10^{-5}$ [78]
$D_s^- \rightarrow l_\alpha^- l_\beta^- K^{*+}$	—————	—————	$1.4 \times 10^{-3}$ [41]
$B^- \rightarrow l_\alpha^- l_\beta^- \pi^+$	$2.3 \times 10^{-8}$ [80]	$1.5 \times 10^{-7}$ [81]	$4.0 \times 10^{-9}$ [82]
$B^- \rightarrow l_\alpha^- l_\beta^- K^+$	$3.0 \times 10^{-8}$ [80]	$1.6 \times 10^{-7}$ [81]	$4.1 \times 10^{-8}$ [83]
$B^- \rightarrow l_\alpha^- l_\beta^- \rho^+$	$1.7 \times 10^{-7}$ [81]	$4.7 \times 10^{-7}$ [81]	$4.2 \times 10^{-7}$ [81]
$B^- \rightarrow l_\alpha^- l_\beta^- D^+$	$2.6 \times 10^{-6}$ [84]	$1.8 \times 10^{-6}$ [84]	$6.9 \times 10^{-7}$ [85]
$B^- \rightarrow l_\alpha^- l_\beta^- D^{*+}$	—————	—————	$2.4 \times 10^{-6}$ [41]
$B^- \rightarrow l_\alpha^- l_\beta^- D_s^+$	—————	—————	$5.8 \times 10^{-7}$ [41]
$B^- \rightarrow l_\alpha^- l_\beta^- K^{*+}$	$4.0 \times 10^{-7}$ [81]	$3.0 \times 10^{-7}$ [81]	$5.9 \times 10^{-7}$ [81]
LNV matrix $m_\nu$	$m_\nu^{ee}$	$m_\nu^{e\mu}$	$m_\nu^{\mu\mu}$

Results from

Belle [84],  
BABAR [78,80,81] and  
LHCb [79,82,83,85];

summarised in PDG [41]

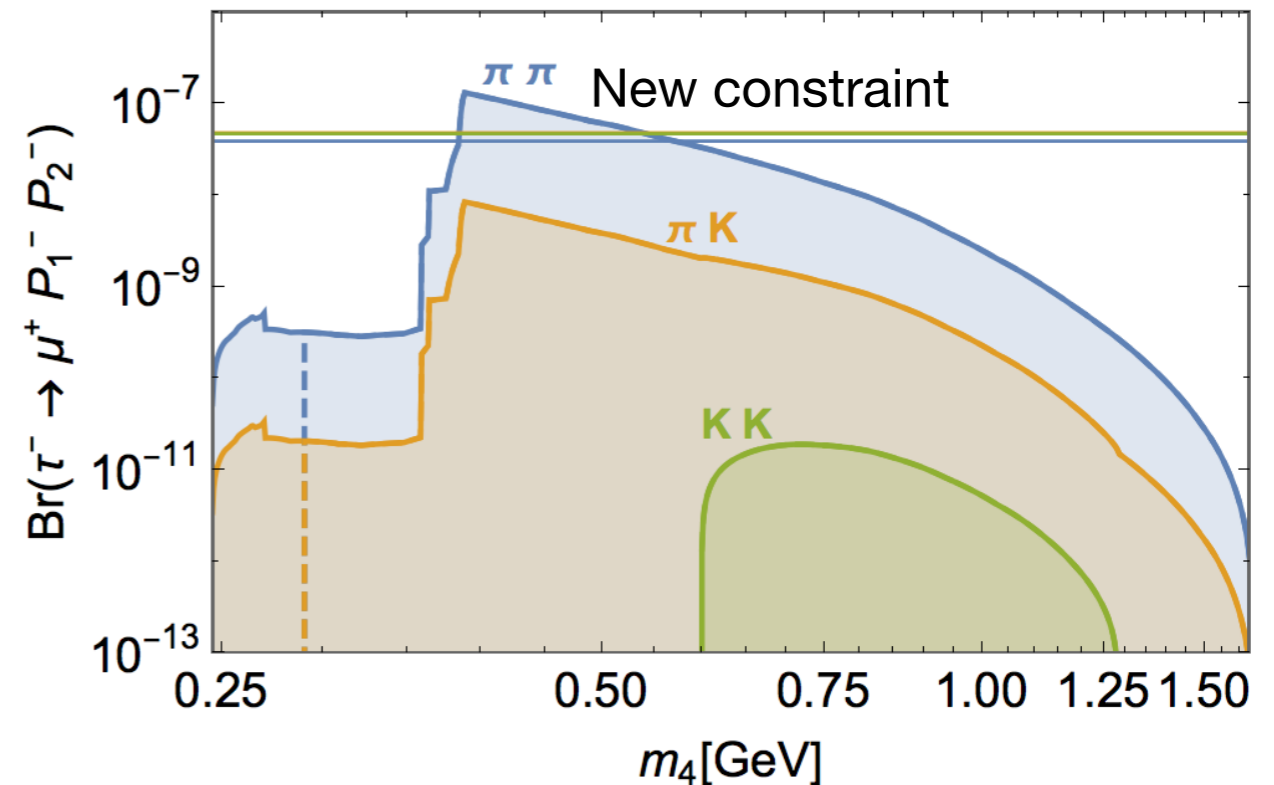
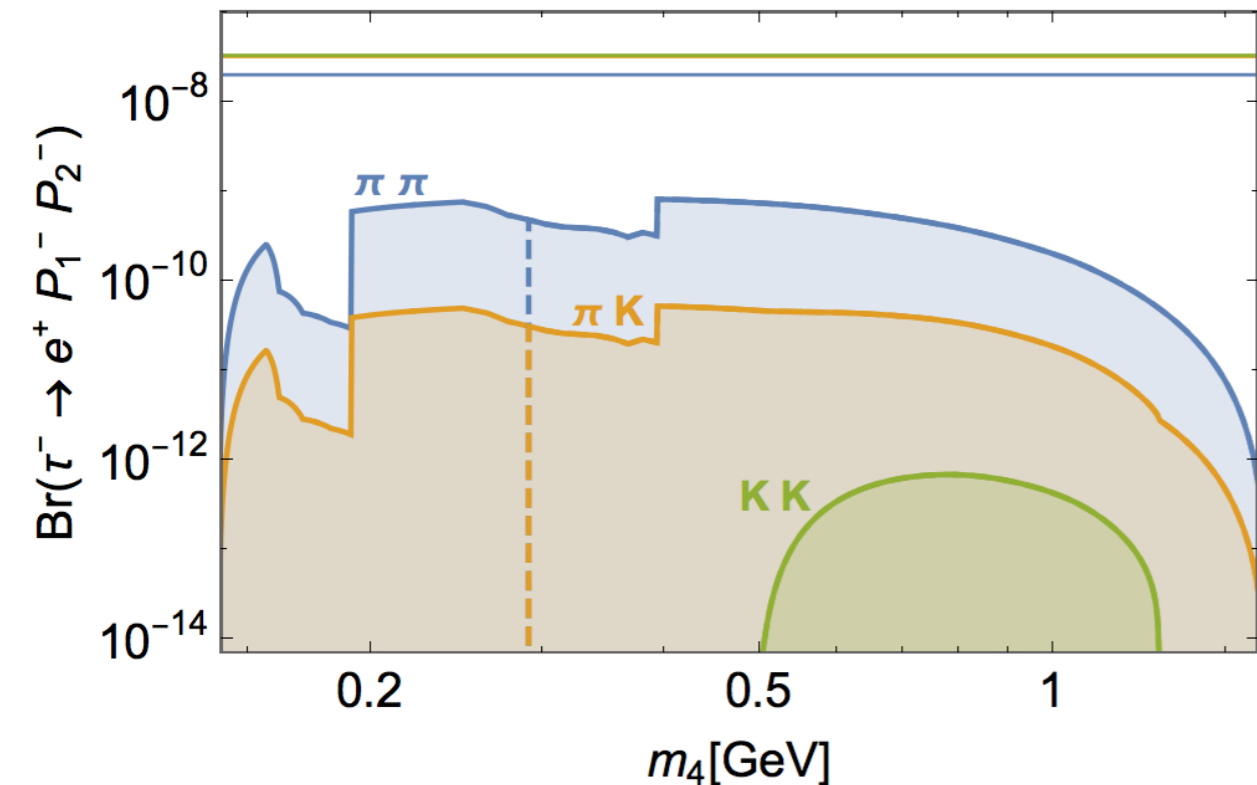
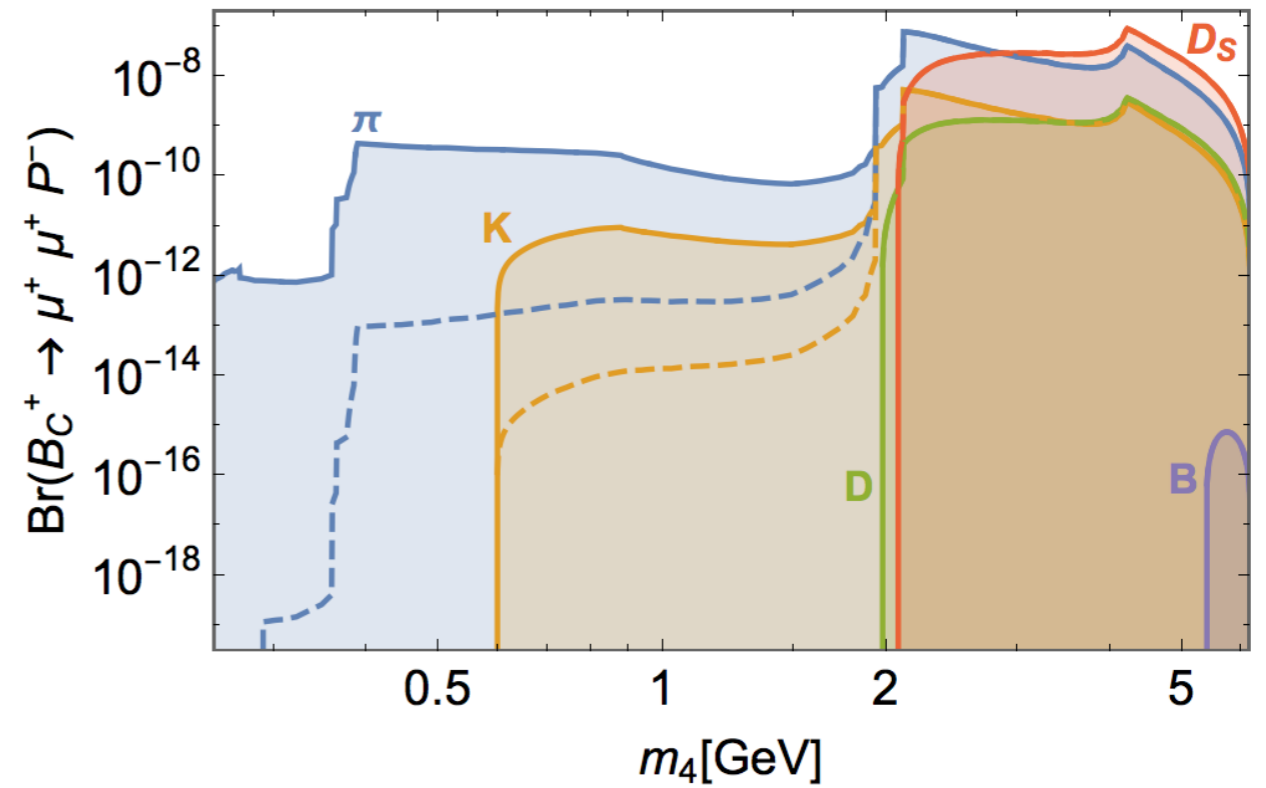
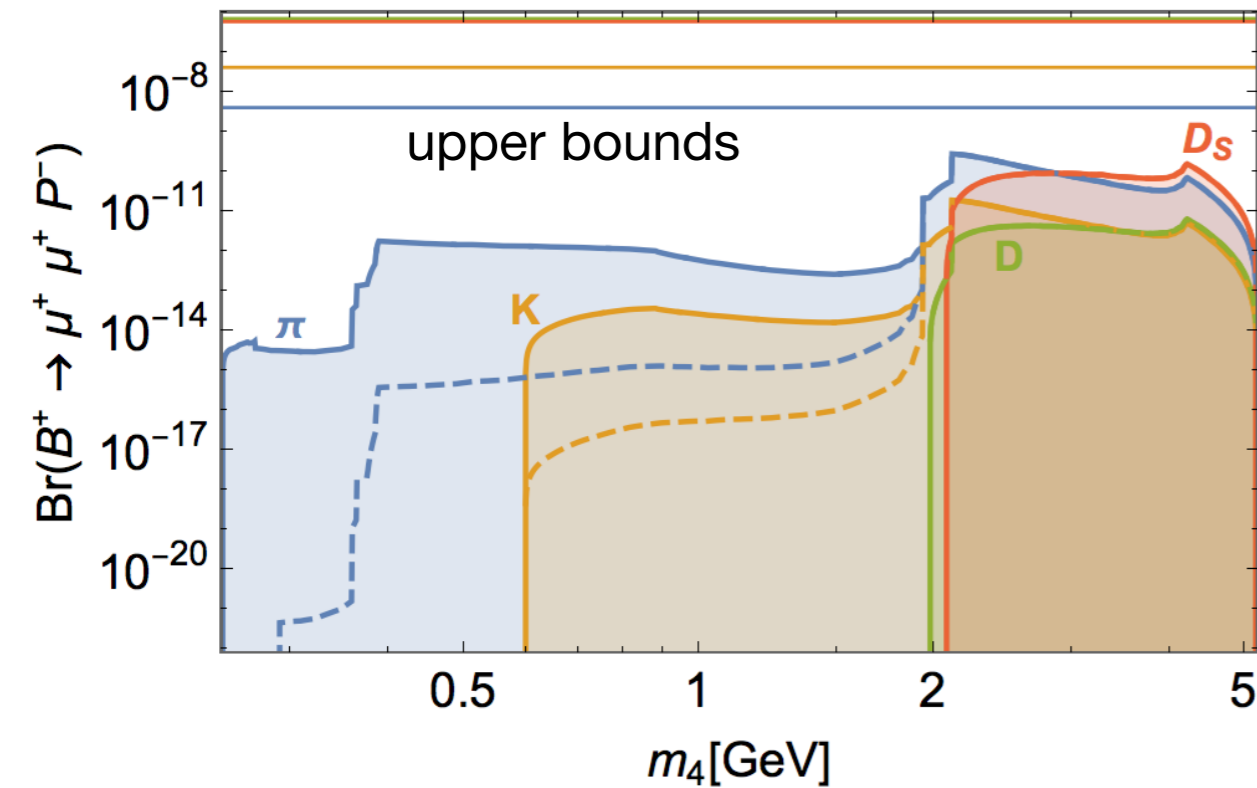
## $\tau$ decay

LNV decay	Current bound	
	$l = e$	$l = \mu$
$\tau^- \rightarrow l^+ \pi^- \pi^-$	$2.0 \times 10^{-8}$	$3.9 \times 10^{-8}$
$\tau^- \rightarrow l^+ \pi^- K^-$	$3.2 \times 10^{-8}$	$4.8 \times 10^{-8}$
$\tau^- \rightarrow l^+ K^- K^-$	$3.3 \times 10^{-8}$	$4.7 \times 10^{-8}$
LNV matrix $m_\nu$	$m_\nu^{e\tau}$	$m_\nu^{\mu\tau}$

upper bounds from the Belle

# Some predictions: single intermediate state

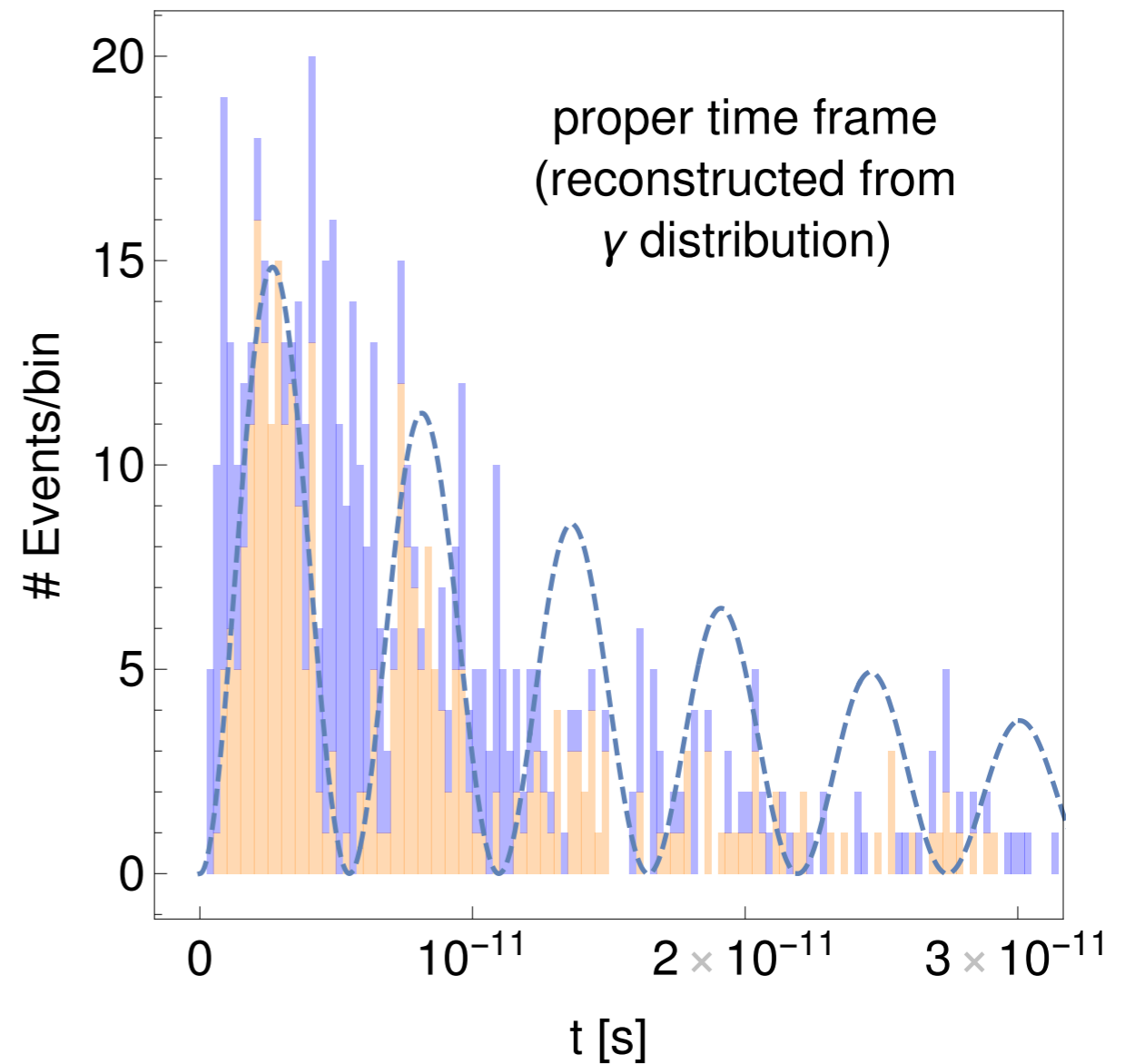
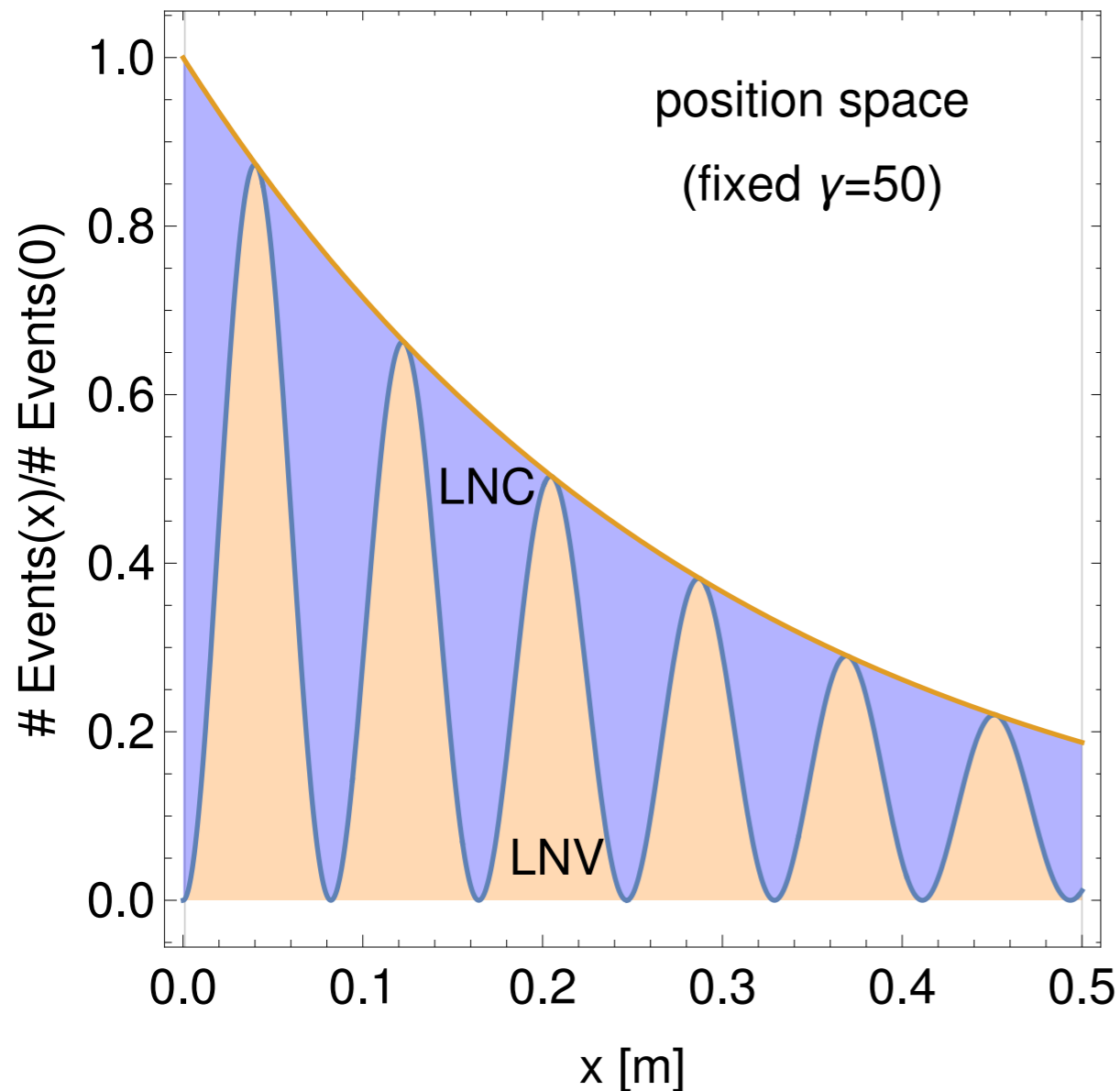
Comprehensive analysis for  $\tau$  and pseudo-scalar mesons in 1712.03984  
(all possible initial and 3-body final states)



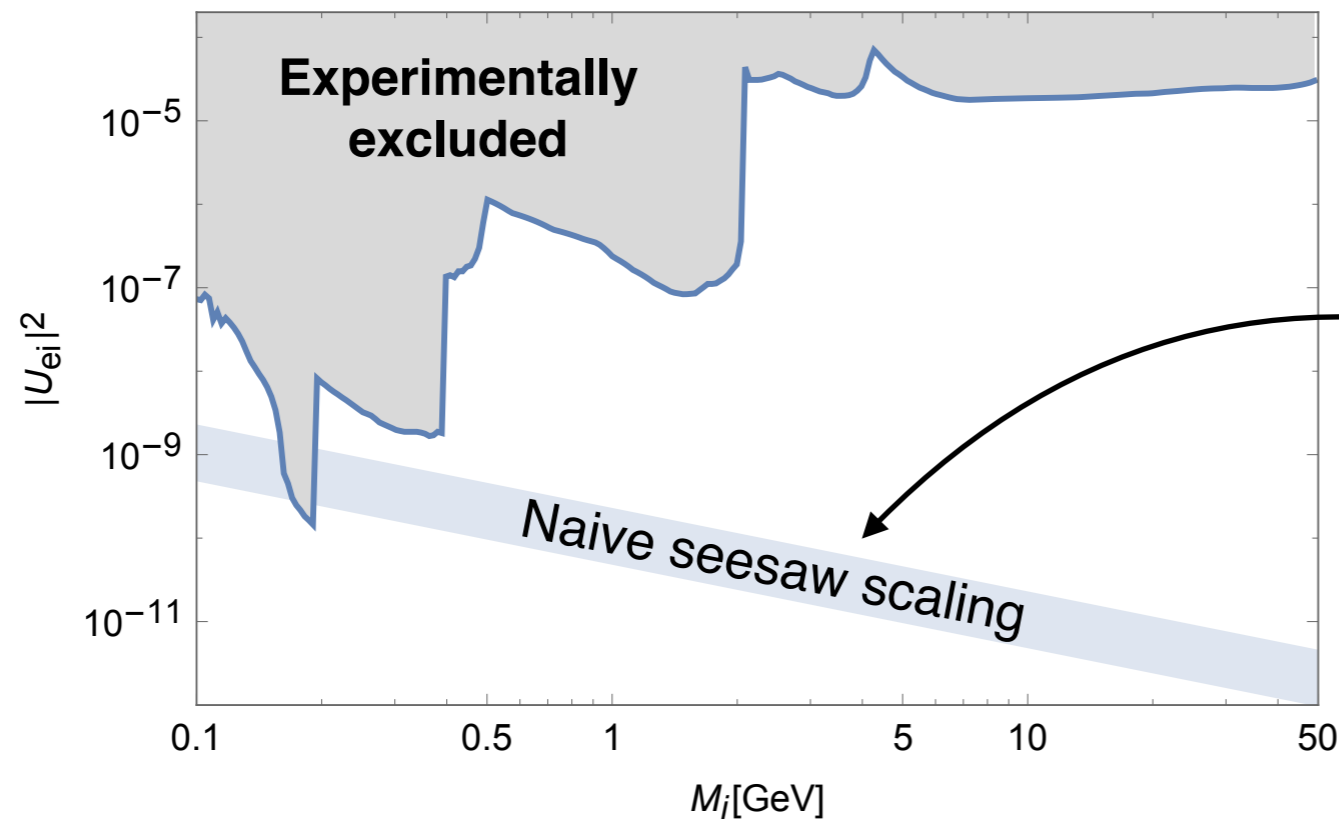
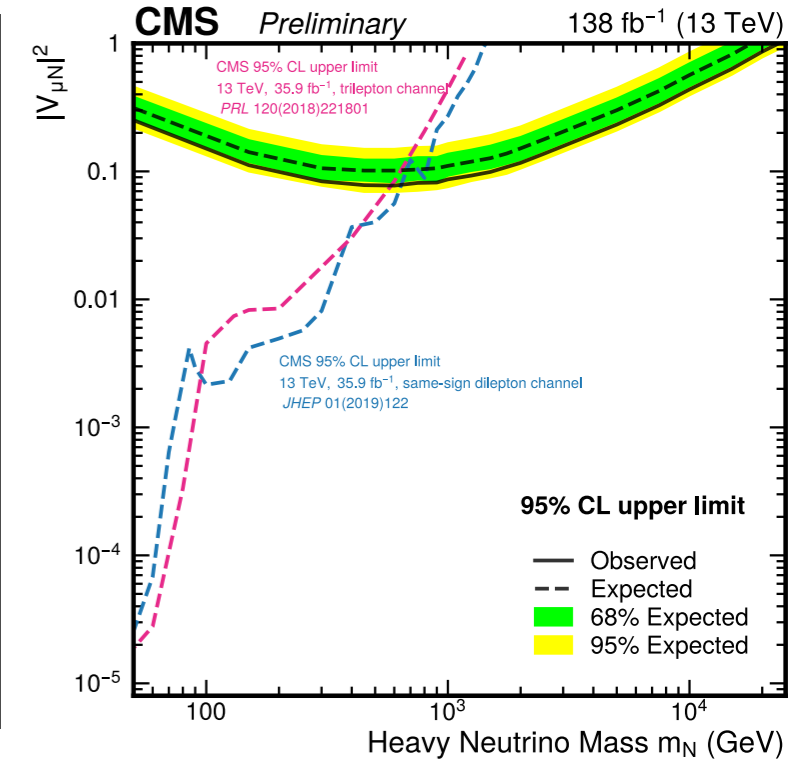
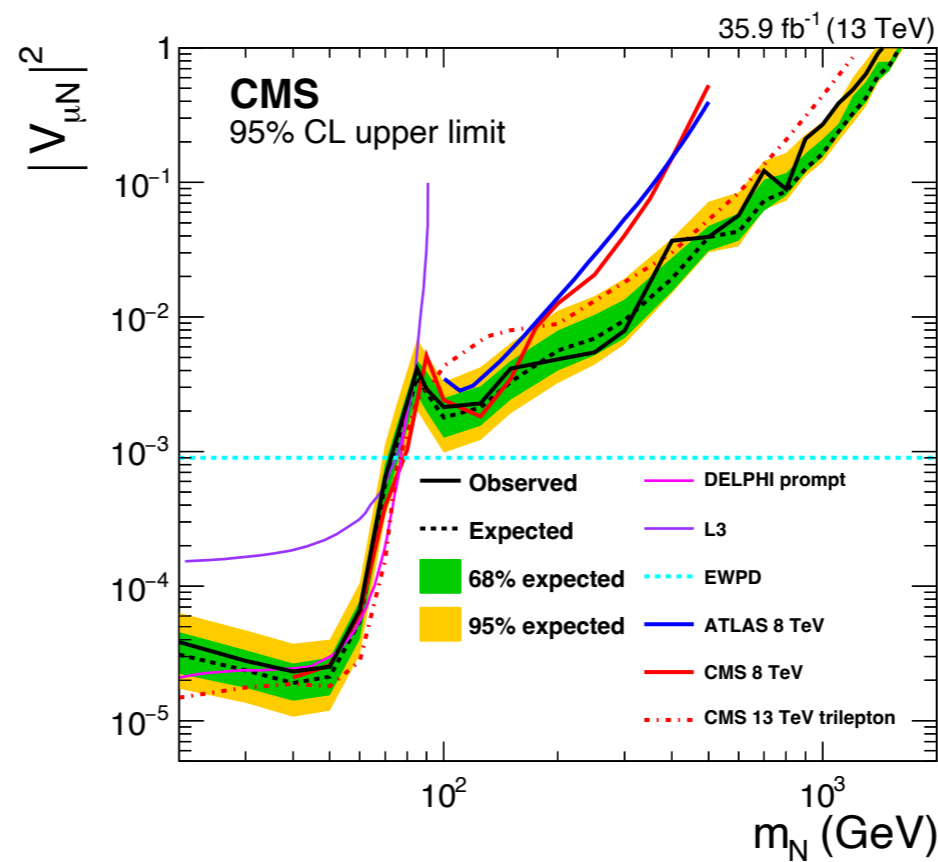
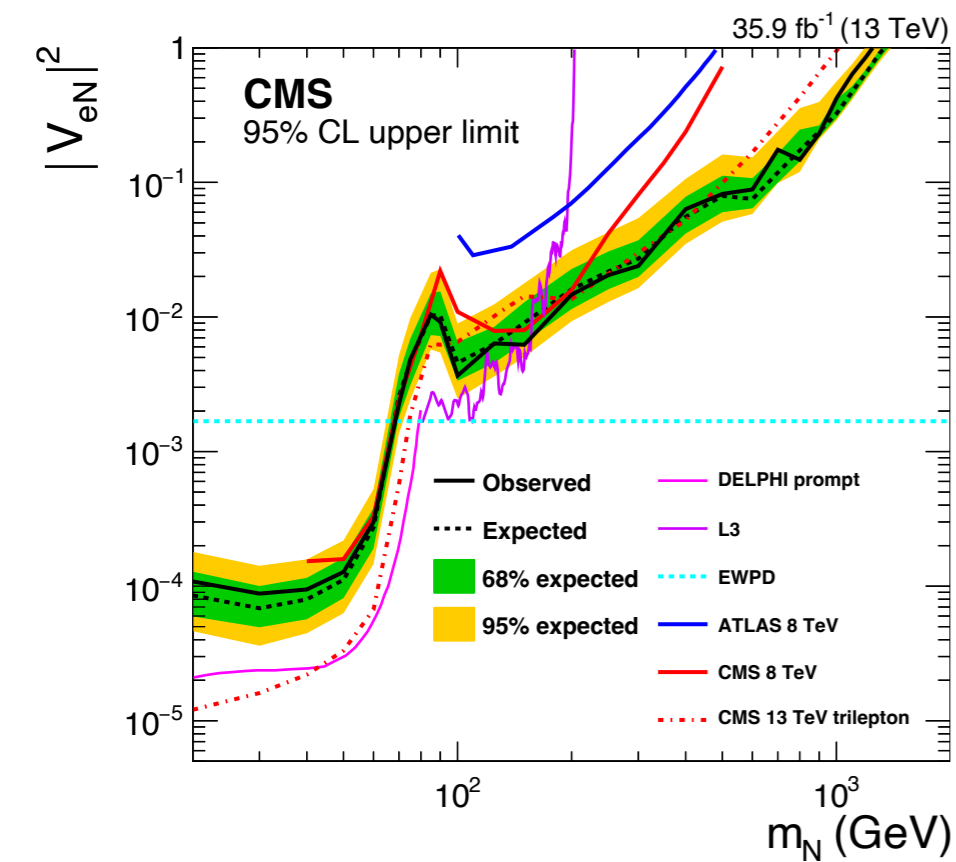
# HNL oscillations at colliders

S. Antusch, E. Cazzato and O. Fischer, arXiv:1709.03797 [hep-ph]

**E.g. LHCb experiment for  
Linear Seesaw with  $M = 7$  GeV,  $U^2 = 10^{-5}$ , Inverted Ordering**



# LNV searches: challenging from model building



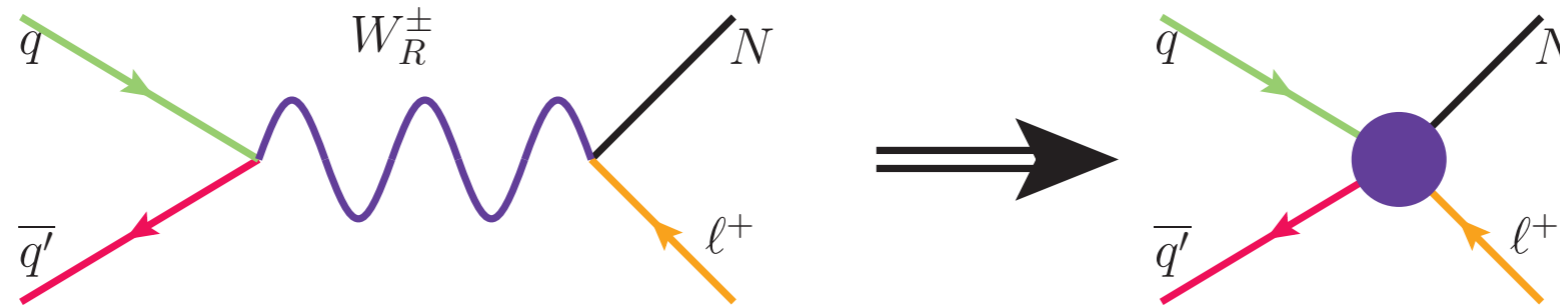
**Typical LNV  
coupling size  
if no accidental  
cancellations**

# Why to look for LNV if $m_\nu \simeq 0$ ?

Equivalence between L conservation and massless neutrinos only holds in SM + singlet fermions

## E.g. Left-right symmetric model

If new gauge mediators are too heavy, light  $N$  are still accessible



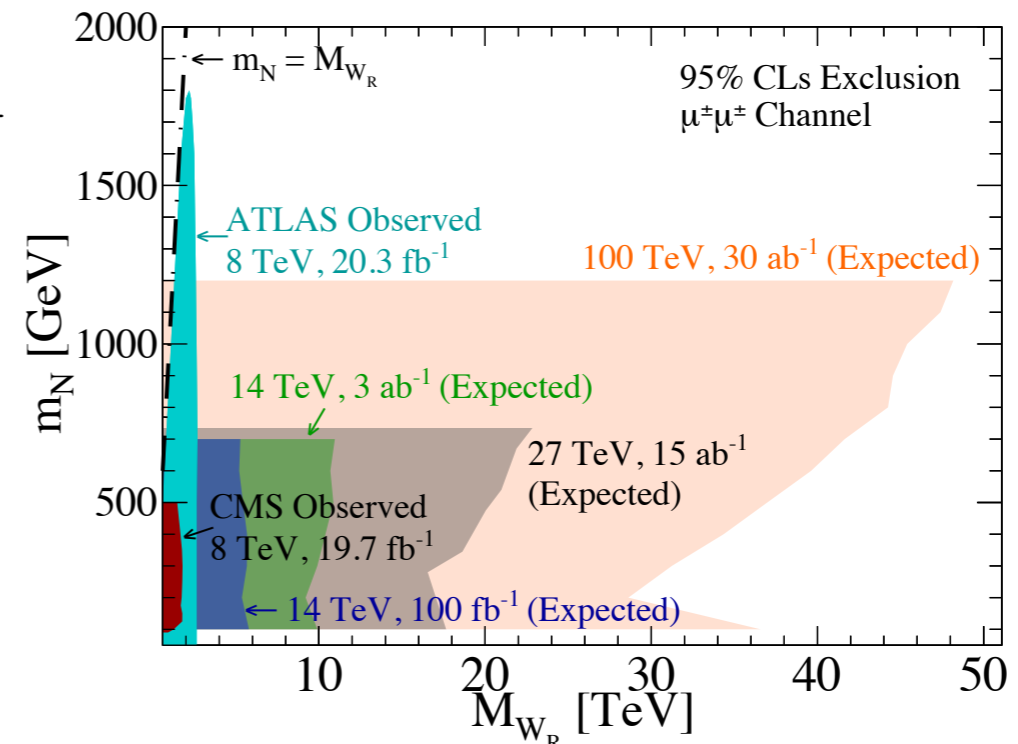
Courtesy of Richard Ruiz

When  $M_{W_R} \gg \sqrt{\hat{s}}$  but  $m_N \lesssim \mathcal{O}(1)$  TeV,  $pp \rightarrow N\ell + X$  production in the LRSM and minimal Type I Seesaw are not discernible<sup>11</sup>

- **Signature:**  $pp \rightarrow \ell^\pm \ell^\pm + nj + X + p_T^\ell \gtrsim \mathcal{O}(m_N) + \text{no MET}$

- At 14 (100) TeV with  $\mathcal{L} = 1$  (10)  $\text{ab}^{-1}$ ,  $M_{W_R} \lesssim 9$  (40) TeV probed

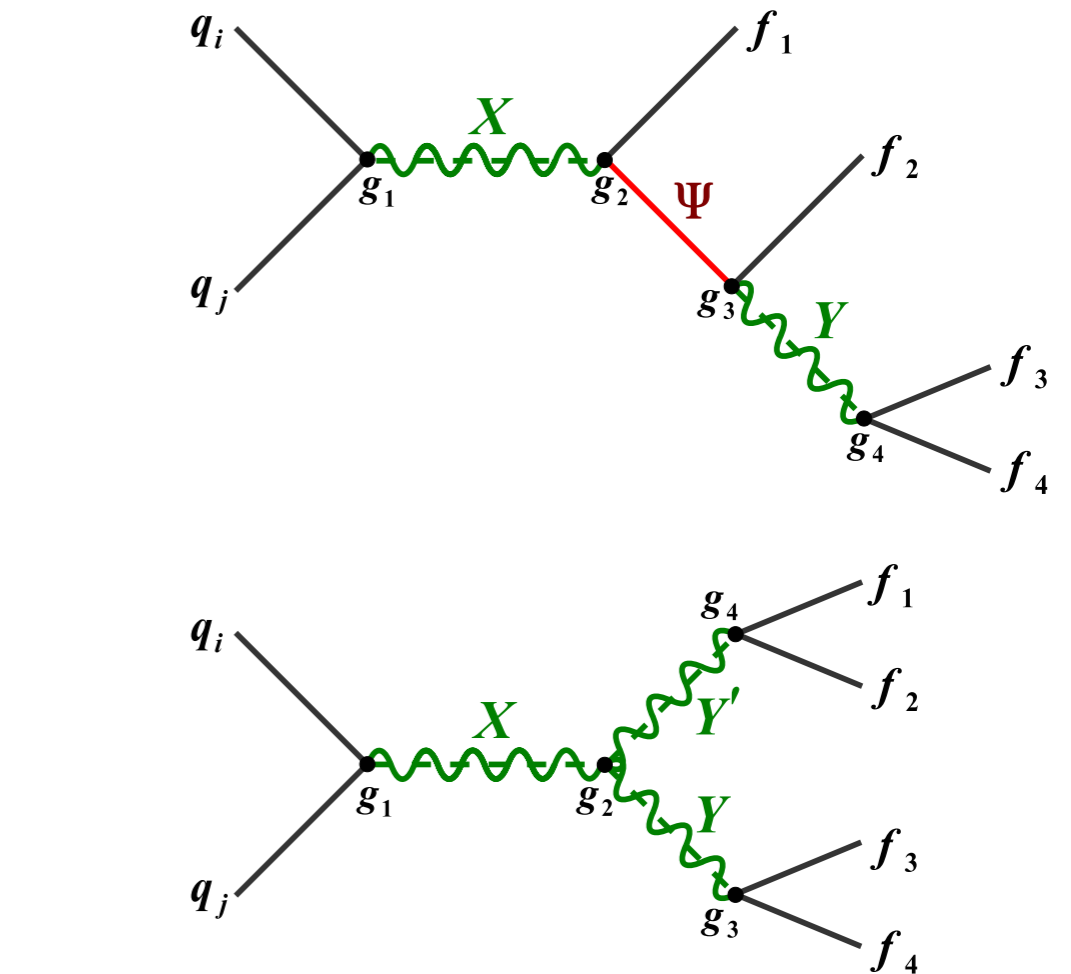
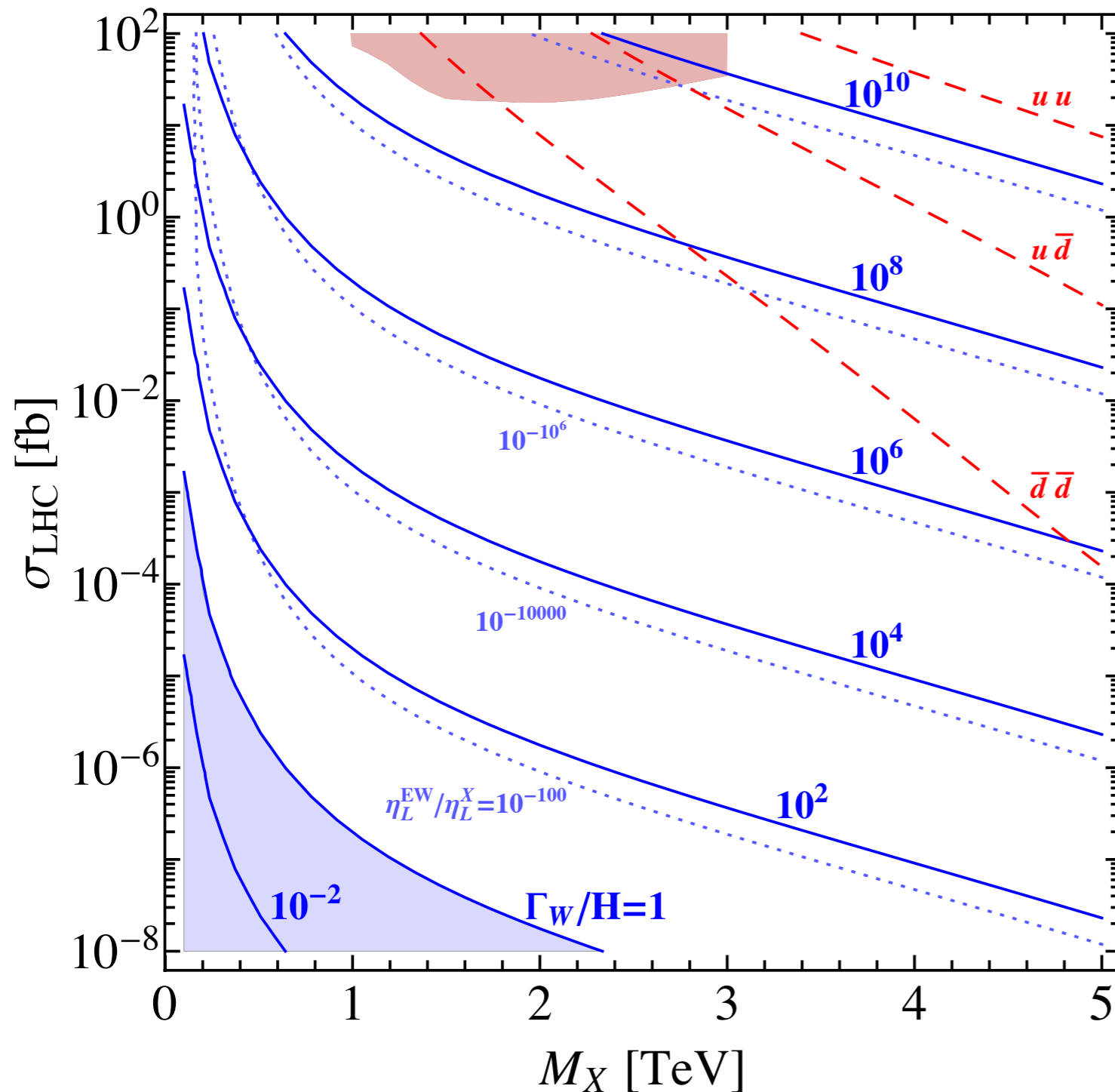
- **DO NOT STOP SEARCHING FOR TYPE I LNV**



<sup>11</sup>Han, Lewis, RR, Si, [1211.6447]; RR, [1703.04669]

# Falsify high-scale leptogenesis with LNV

J. M. Frere, T. Hambye and G. Vertongen, arXiv:0806.0841 [hep-ph];  
 F. F. Deppisch, J. Harz and M. Hirsch, arXiv:1312.4447 [hep-ph]



$$\frac{\Gamma_W}{H} = \frac{0.028}{\sqrt{g_*}} \frac{M_P M_X^3}{T^4} \frac{K_1(M_X/T)}{f_{q_1 q_2}(M_X/\sqrt{s})} \times (s\sigma_{\text{LHC}})$$

**A LNV observation at LHC likely falsifies high-scale leptogenesis**