

Non-Abelian Domain walls

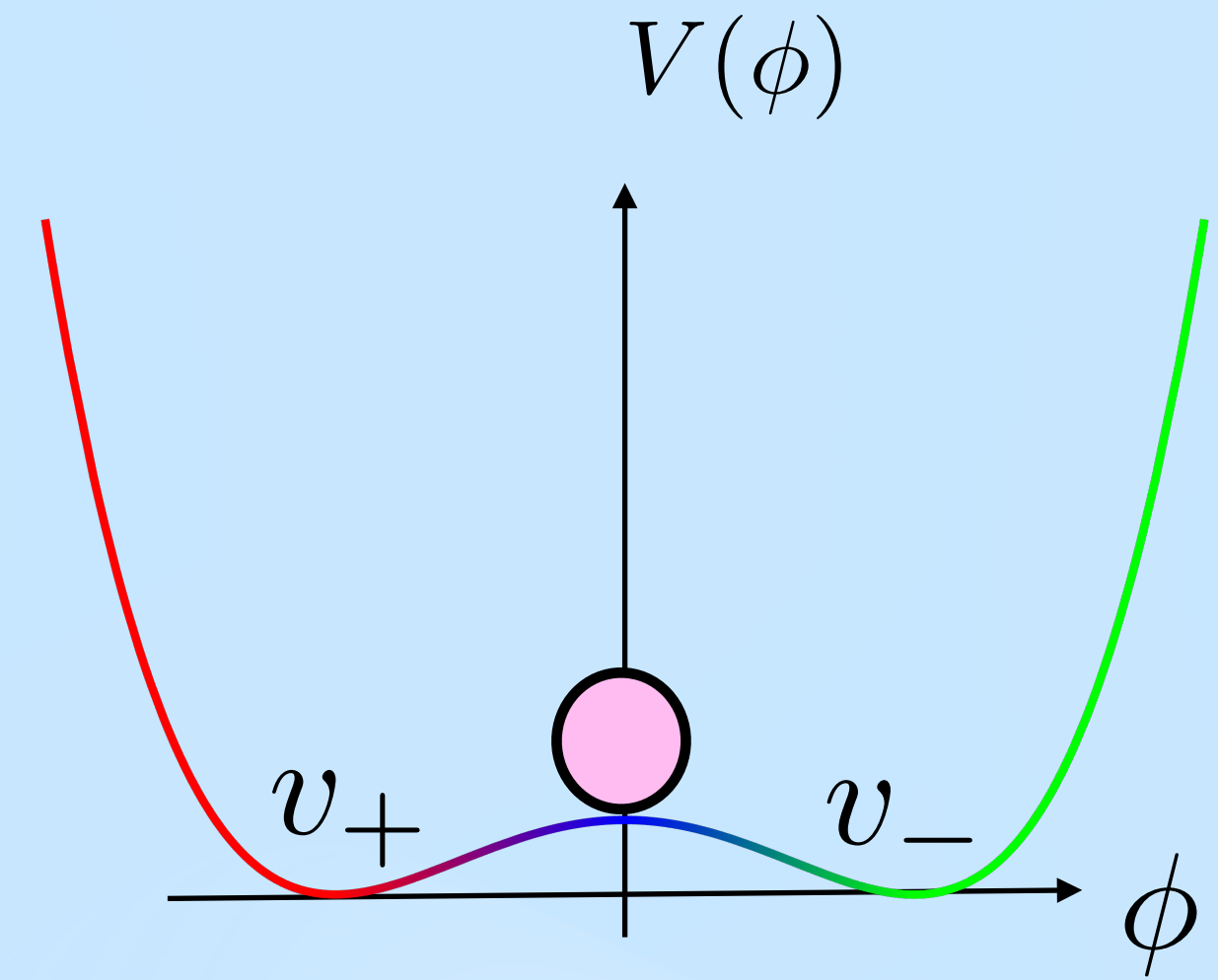
付博文 20 Sep 2025

高能物理理论与实验融合发展研讨会

Based on **BF**, S. F. King, L. Marsili, S. Pascoli, J. Turner, Y-L Zhou, 2409.16359

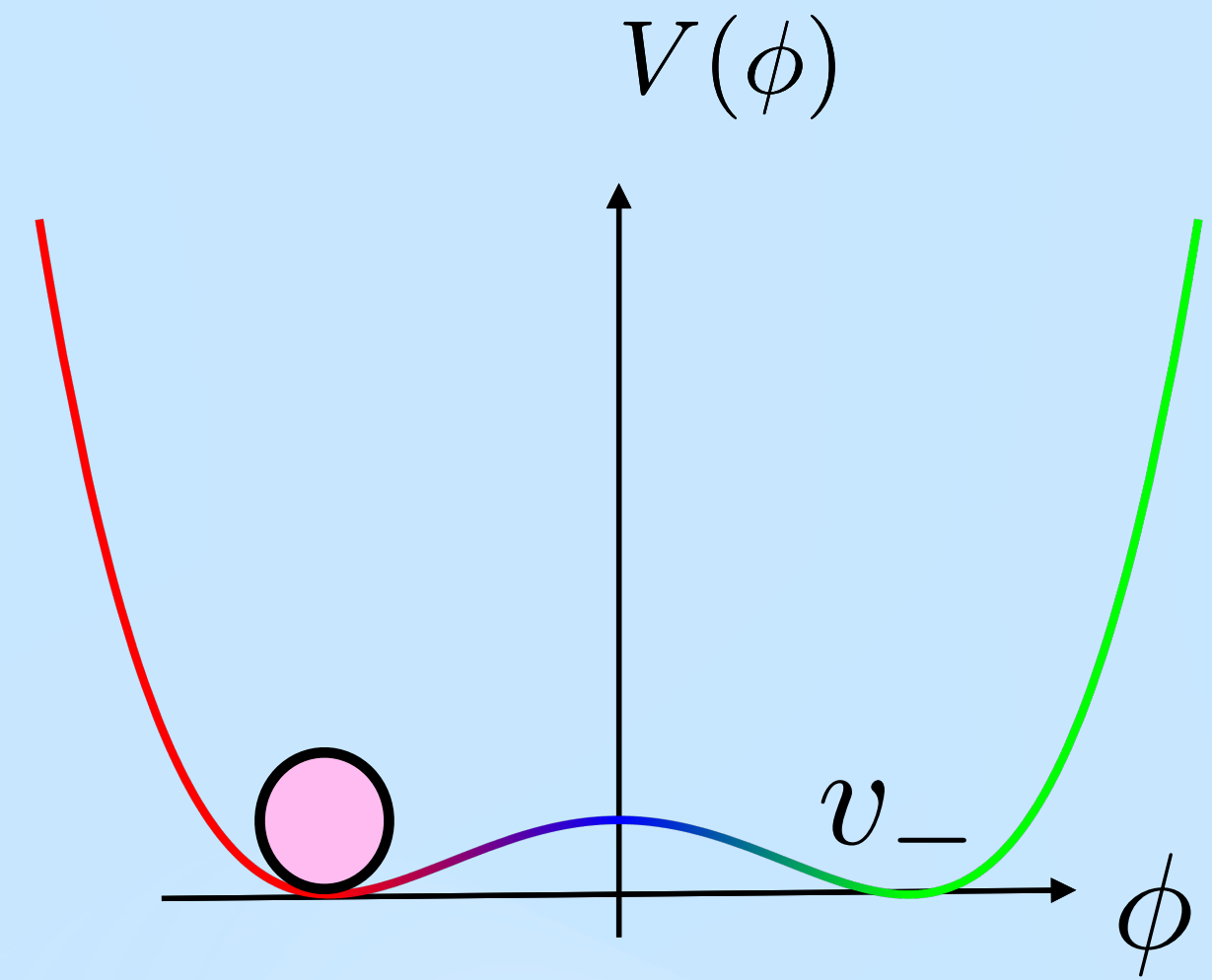
Domain walls

Consider a real scalar field with $\mathbb{Z}_2$ -symmetric potential

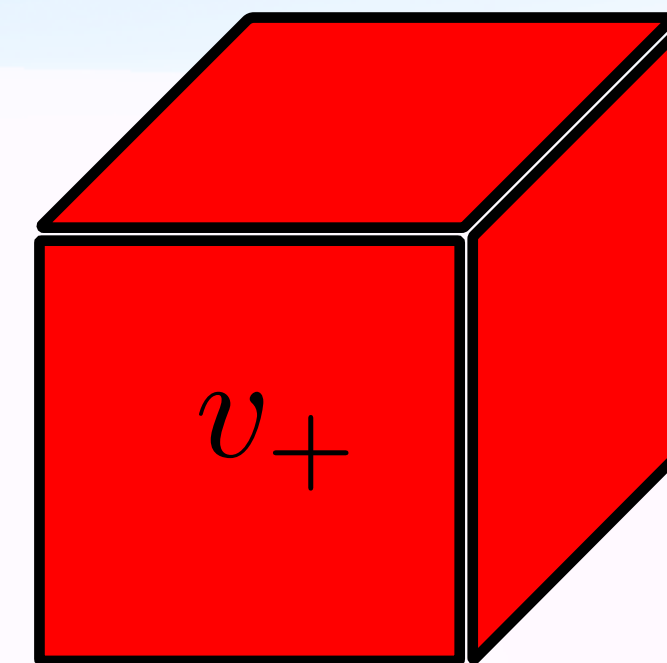


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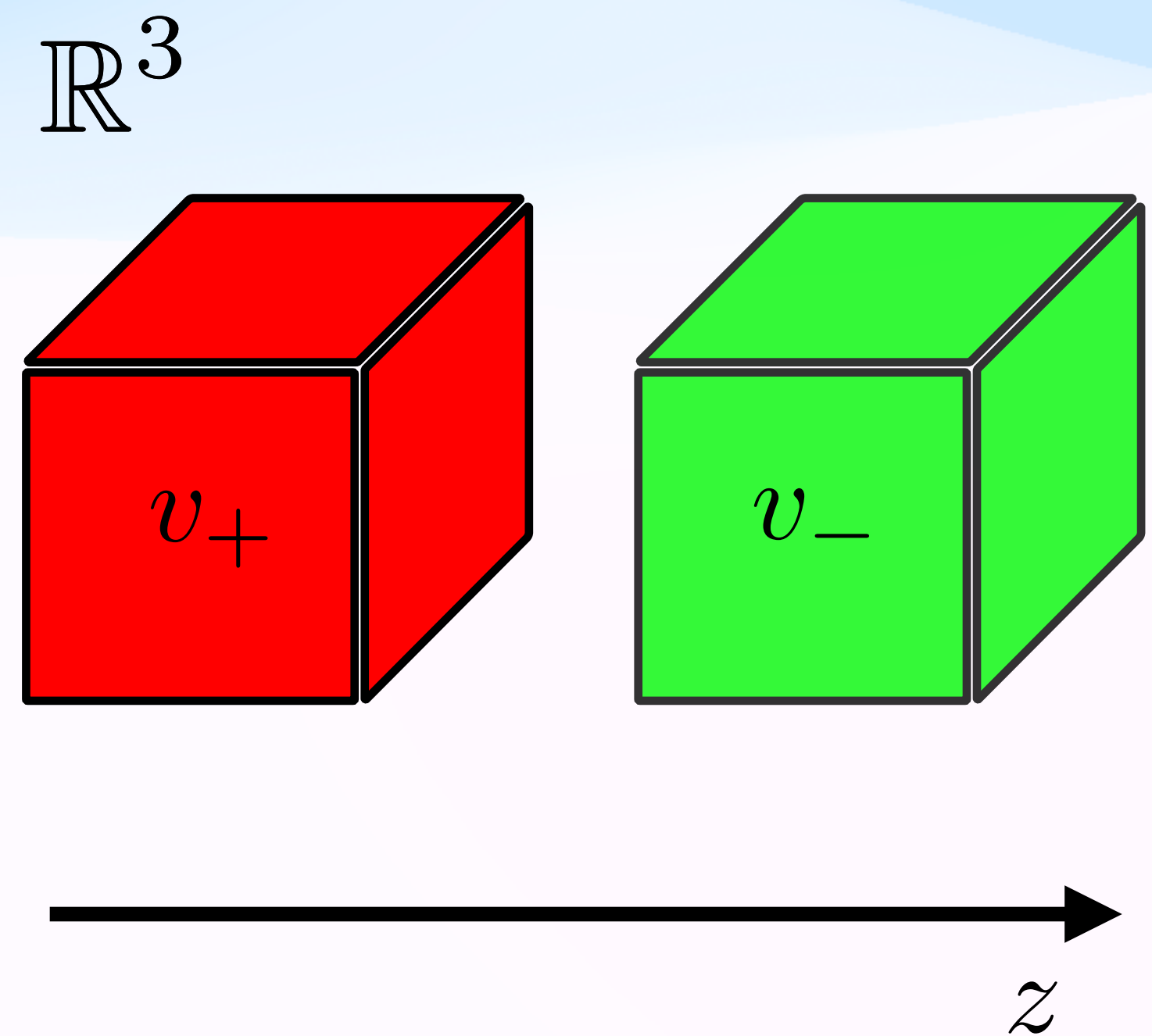
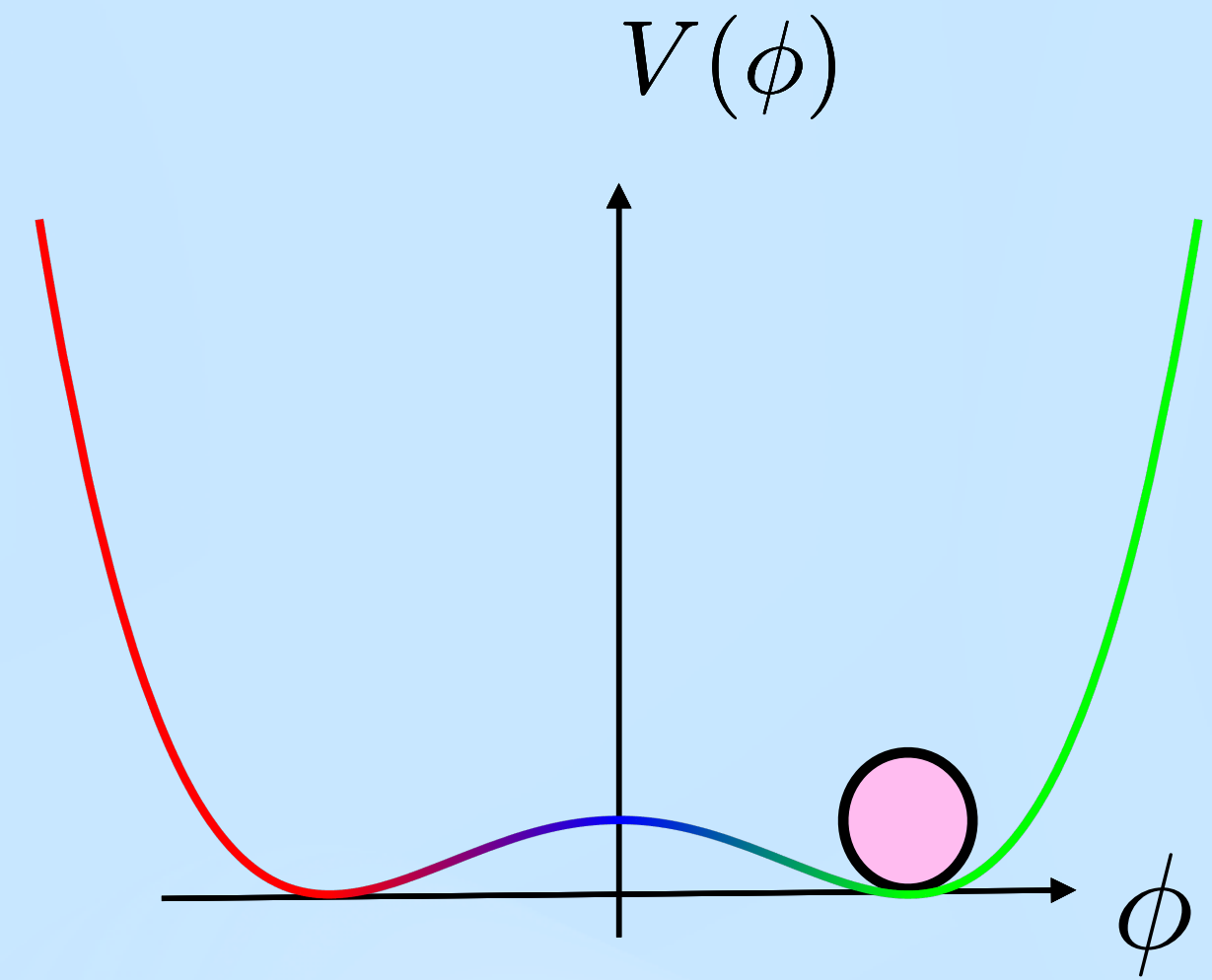


$\mathbb{R}^3$



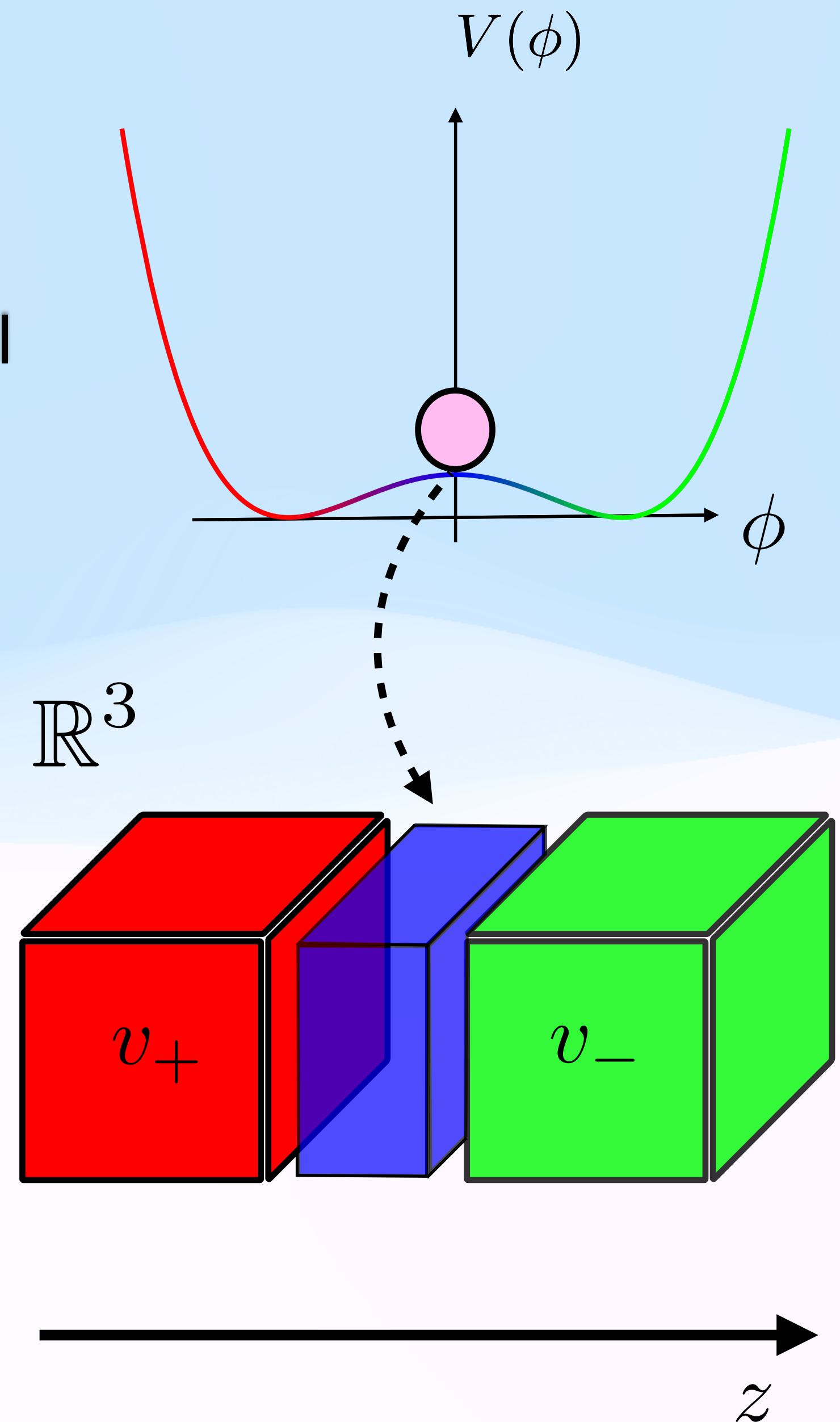
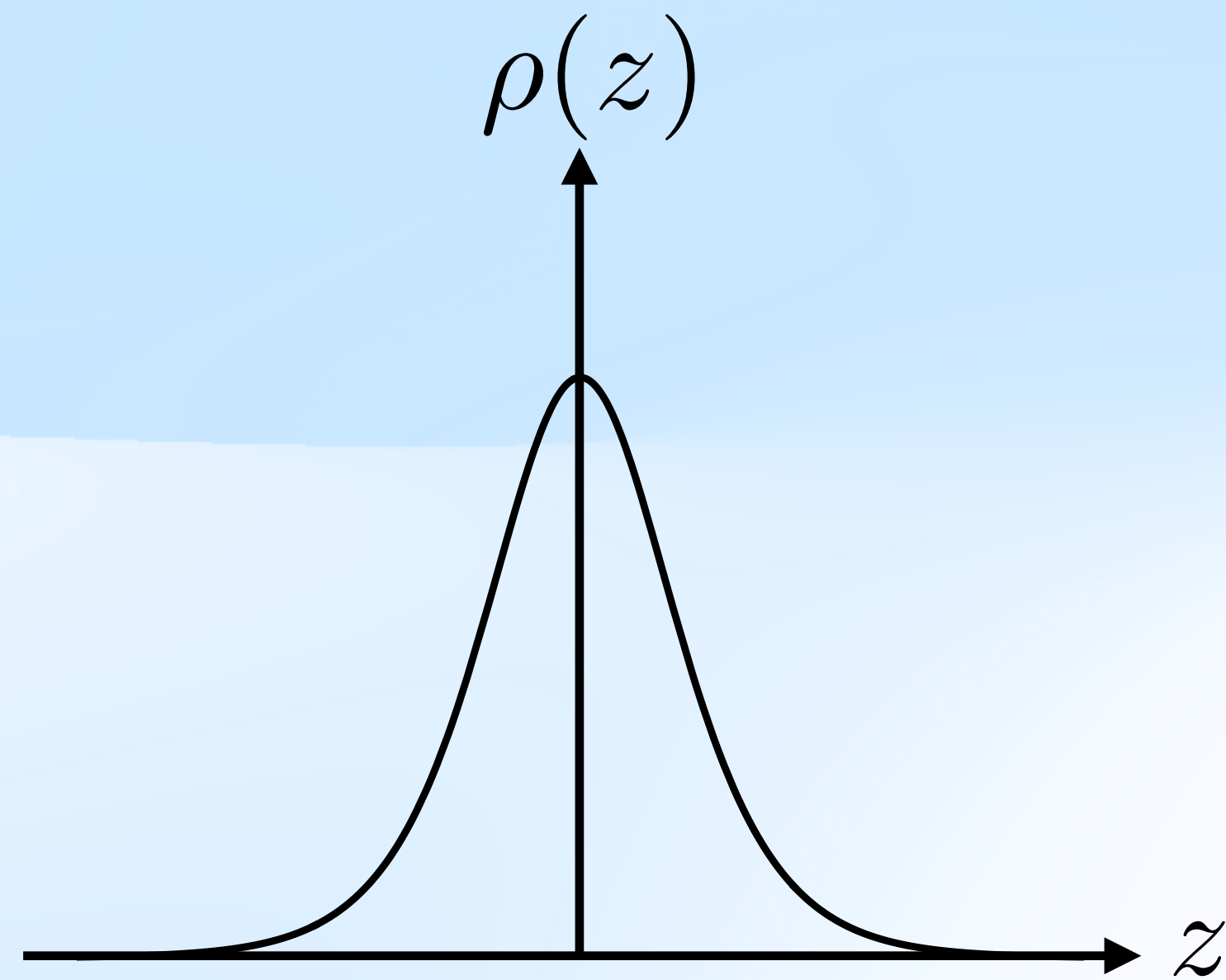
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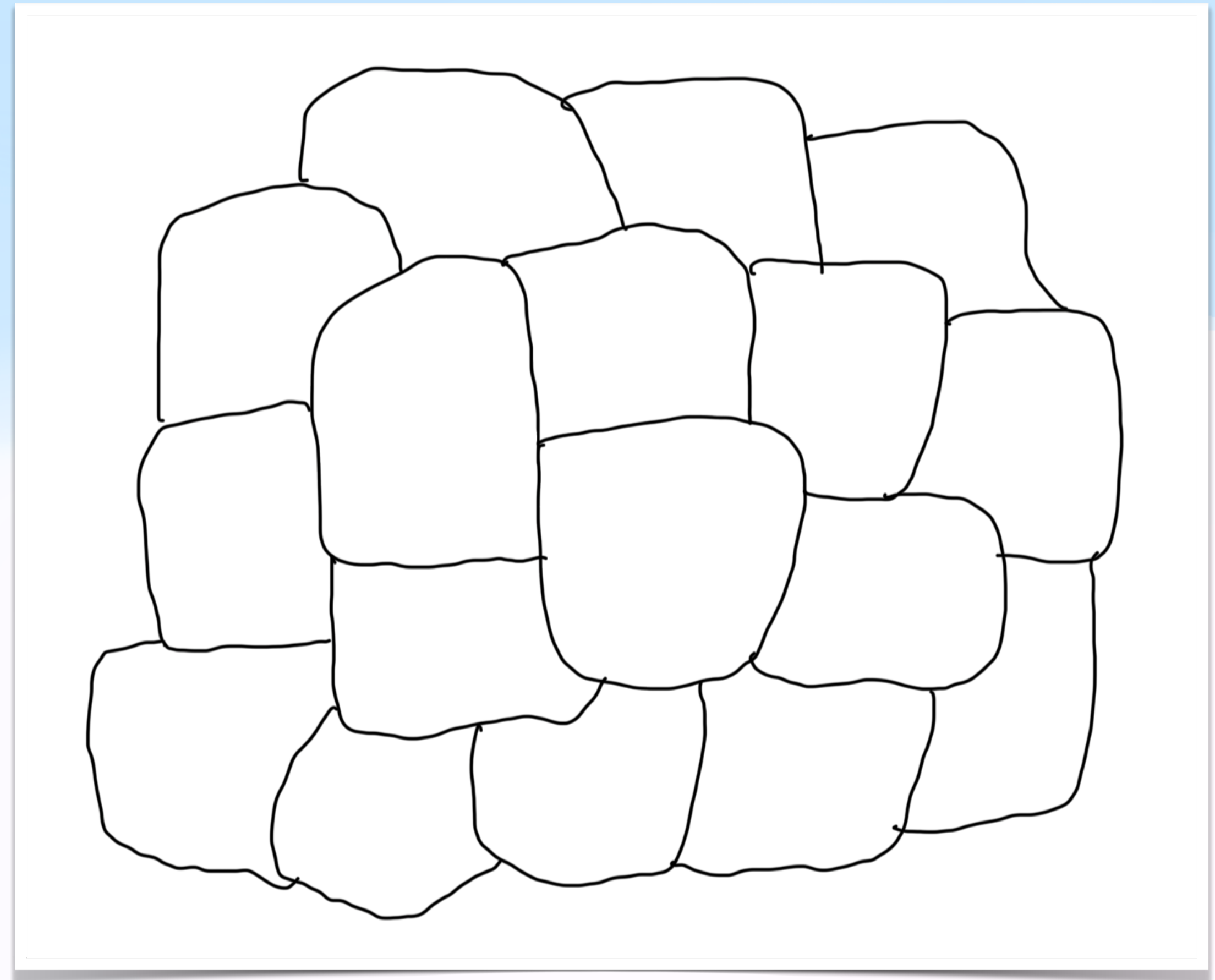
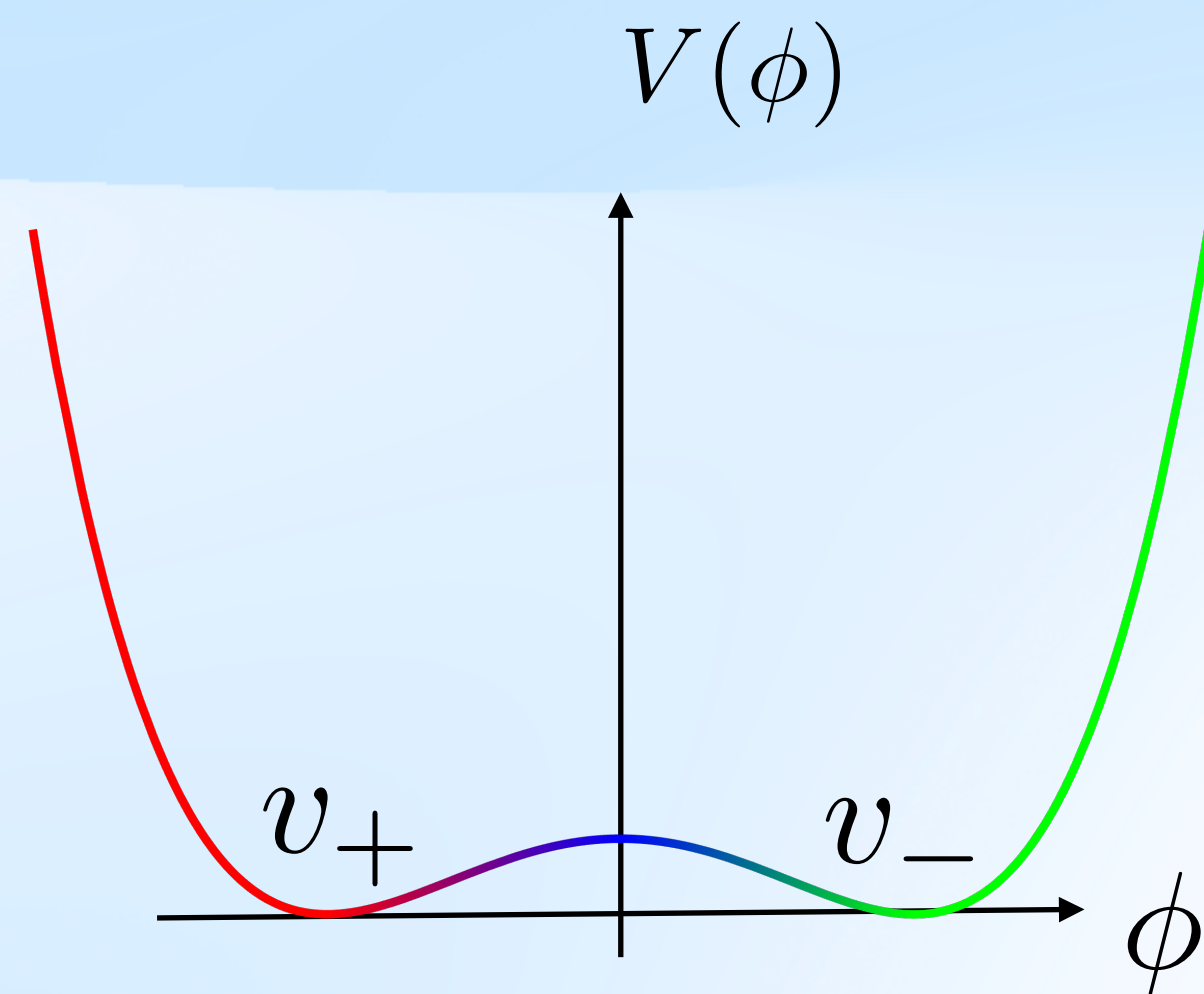
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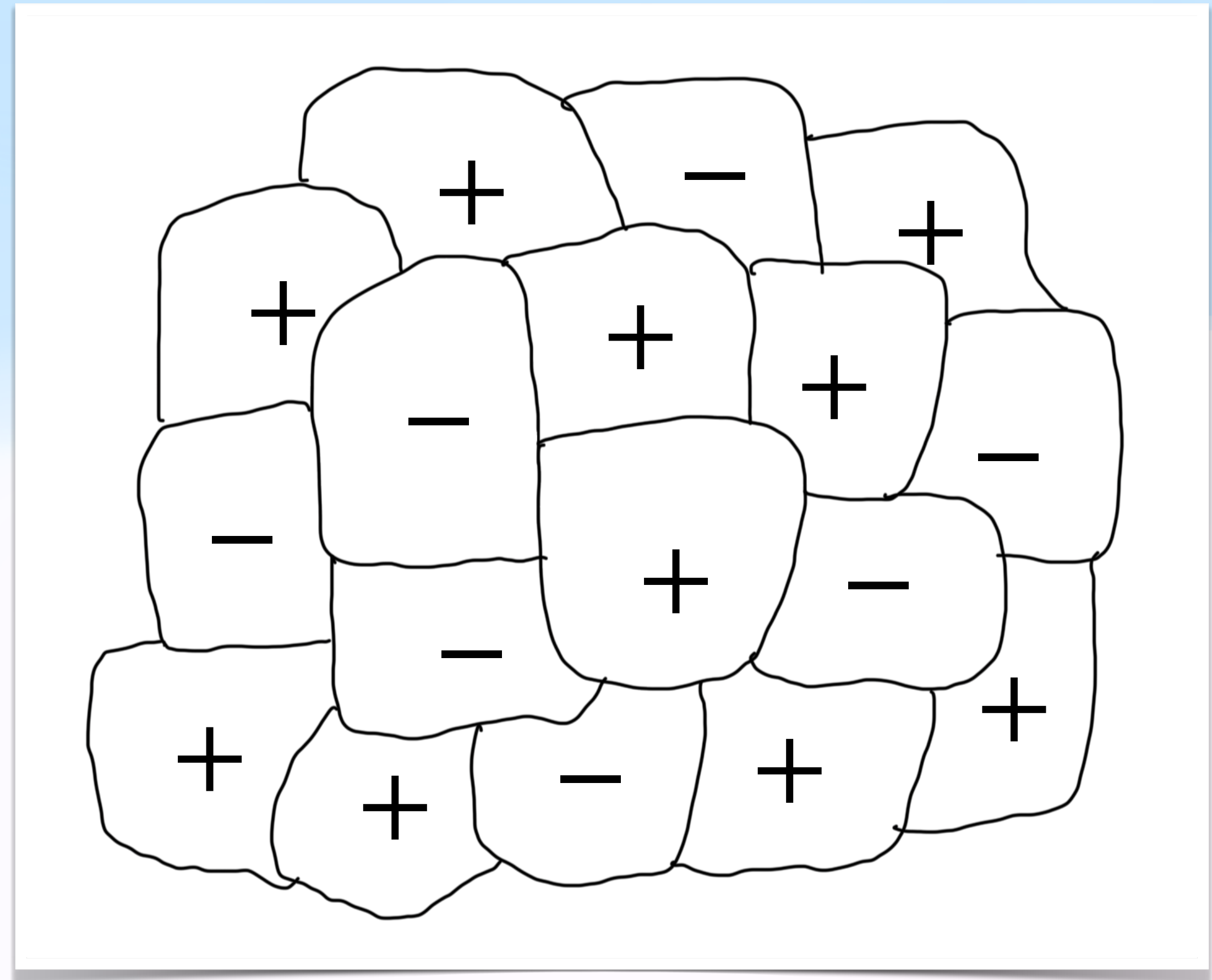
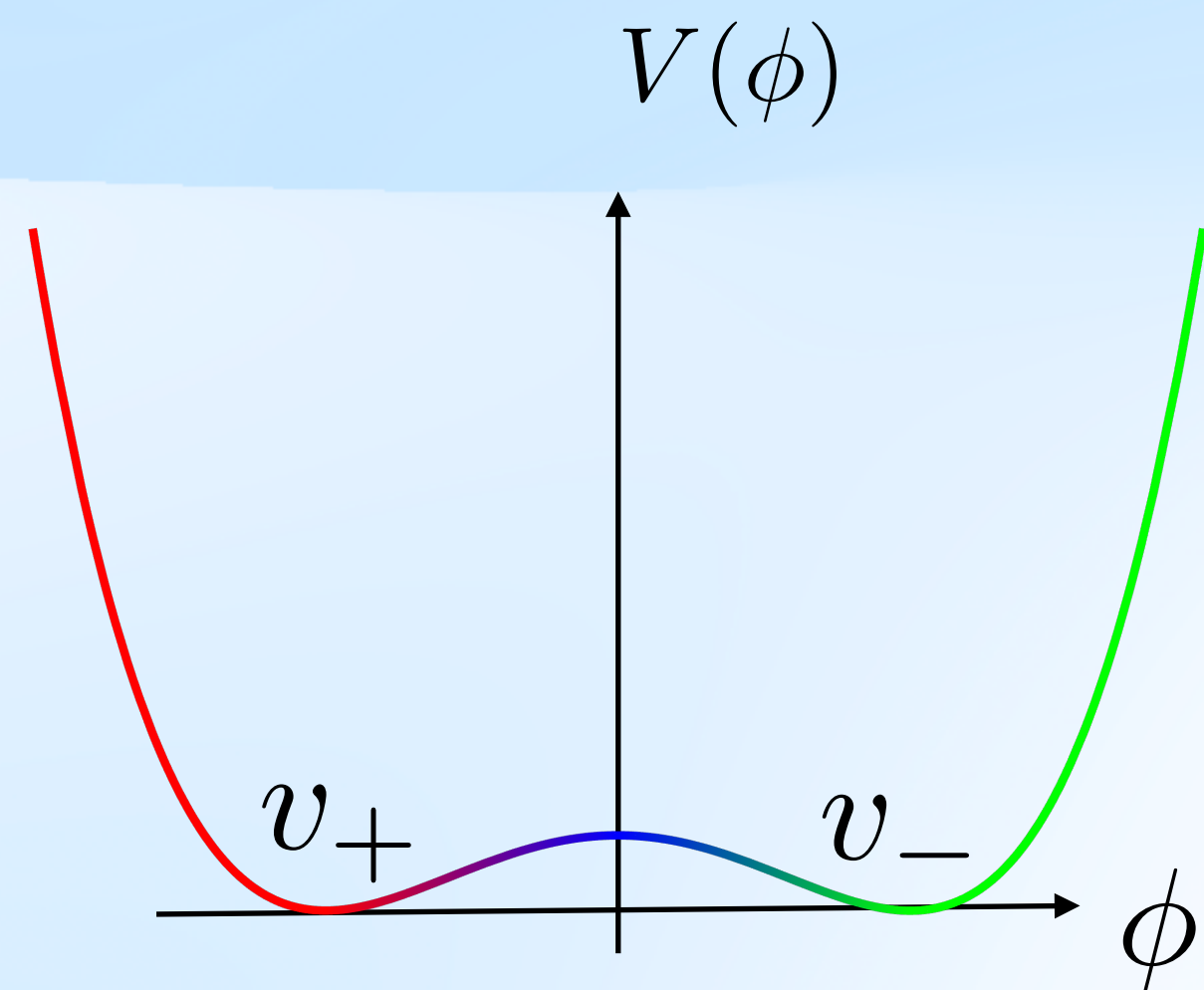
Domain wall formation

Kibble mechanism



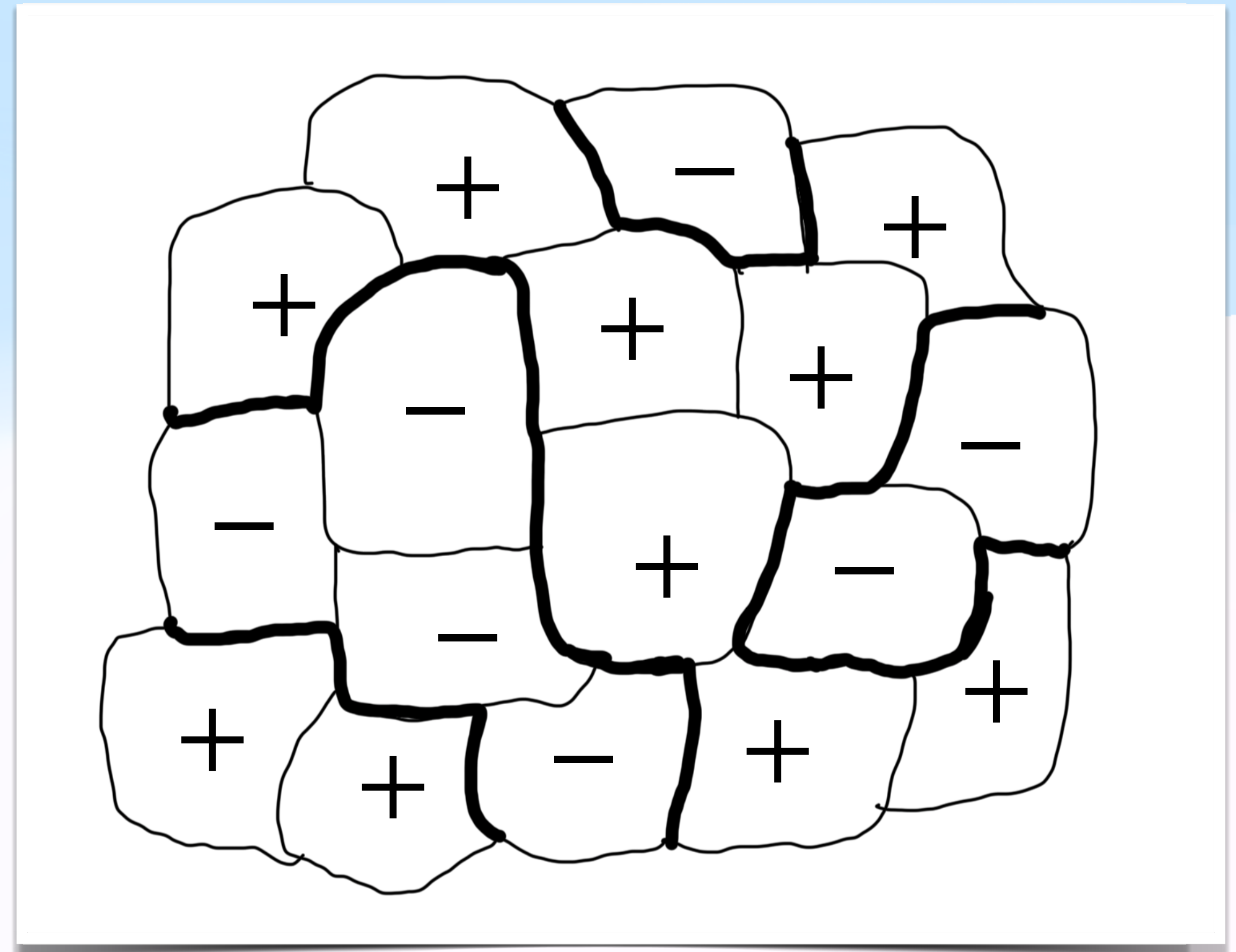
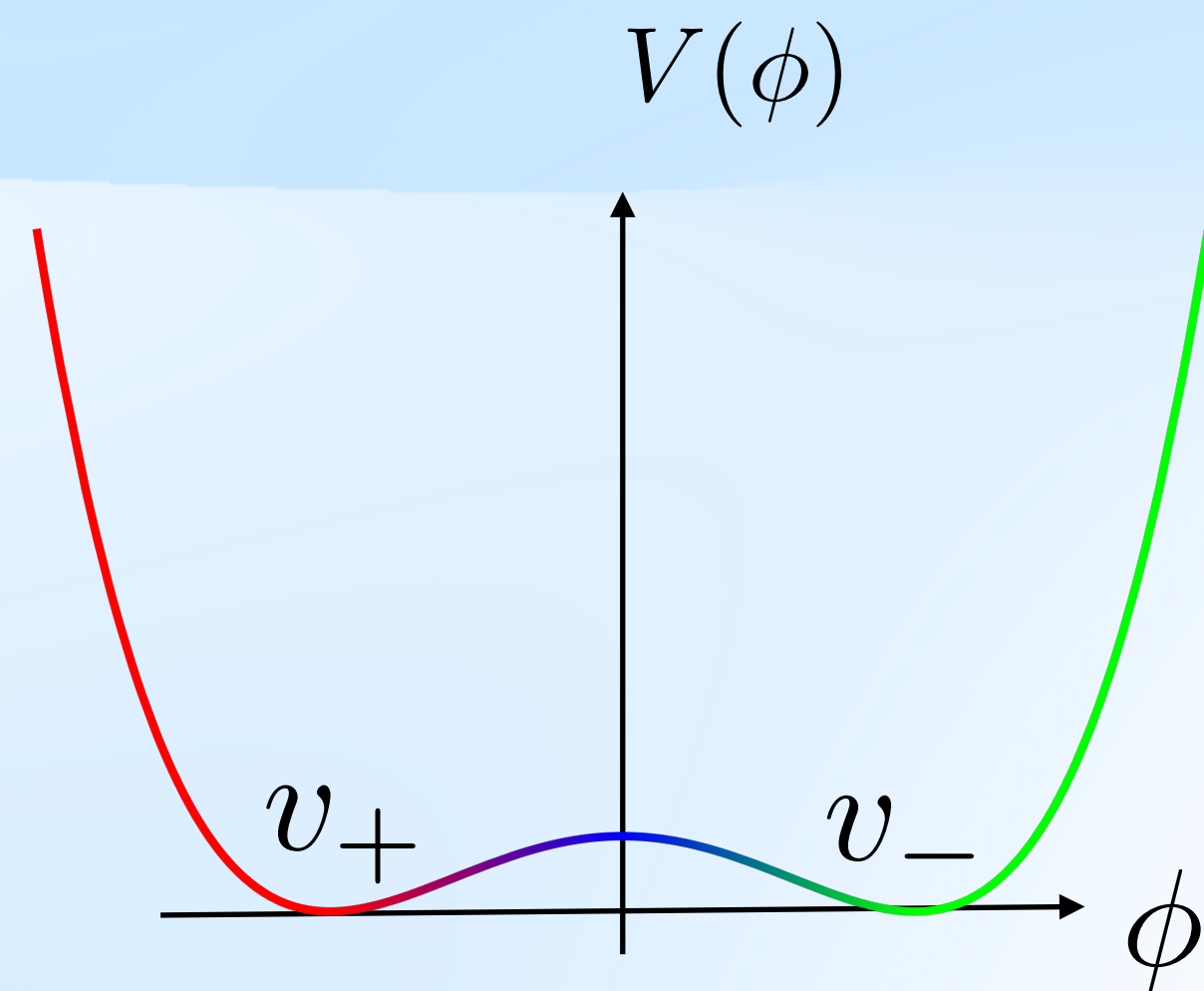
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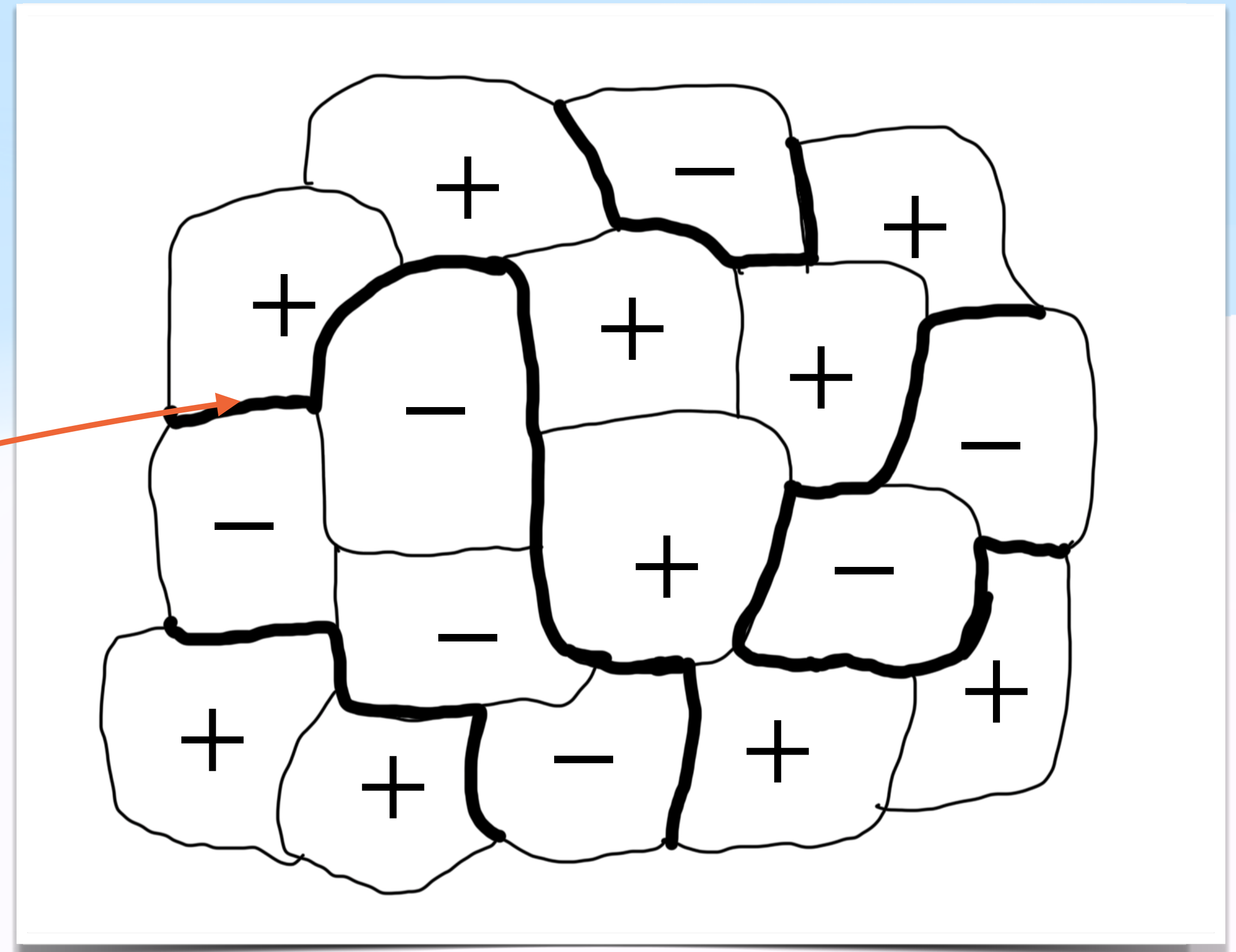
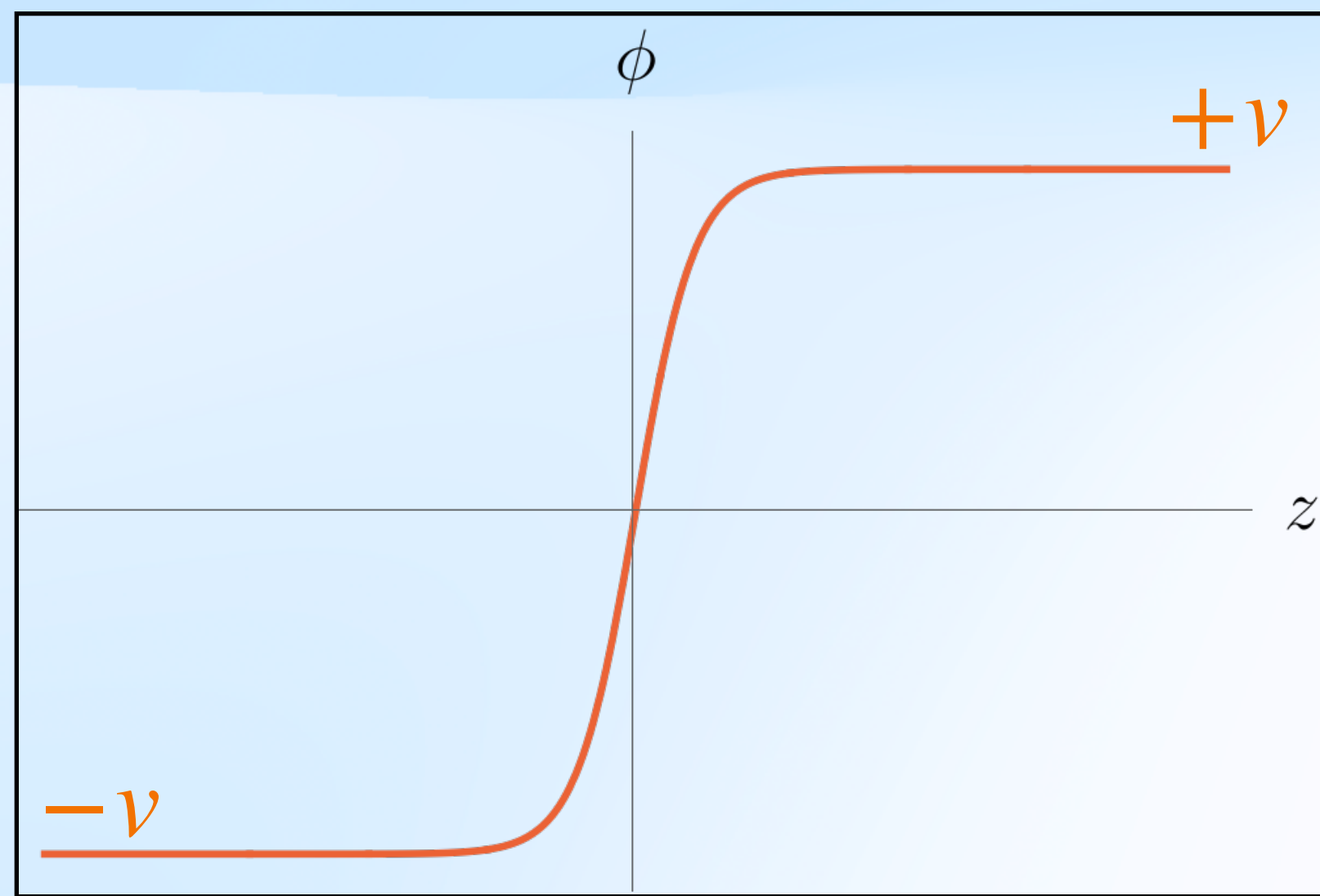
Kibble mechanism



Domain wall formation

$$\frac{\partial^2 \phi}{\partial z^2} = \frac{\partial V}{\partial \phi}, \quad \phi(\pm\infty) = \pm v$$

$$\phi(z) = v \tanh \frac{z}{\Delta} \quad \Delta = \sqrt{\frac{2}{\lambda v^2}}$$



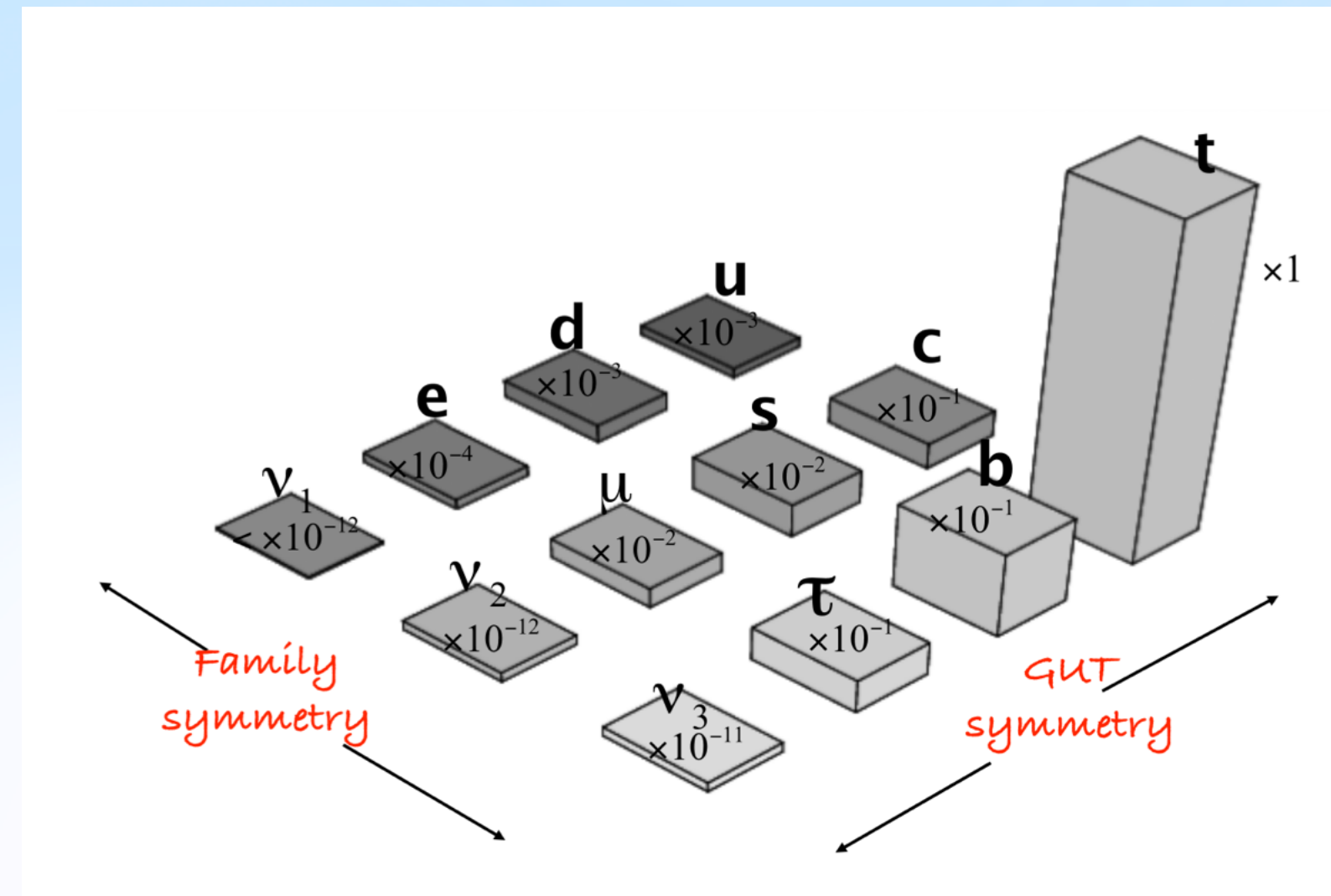
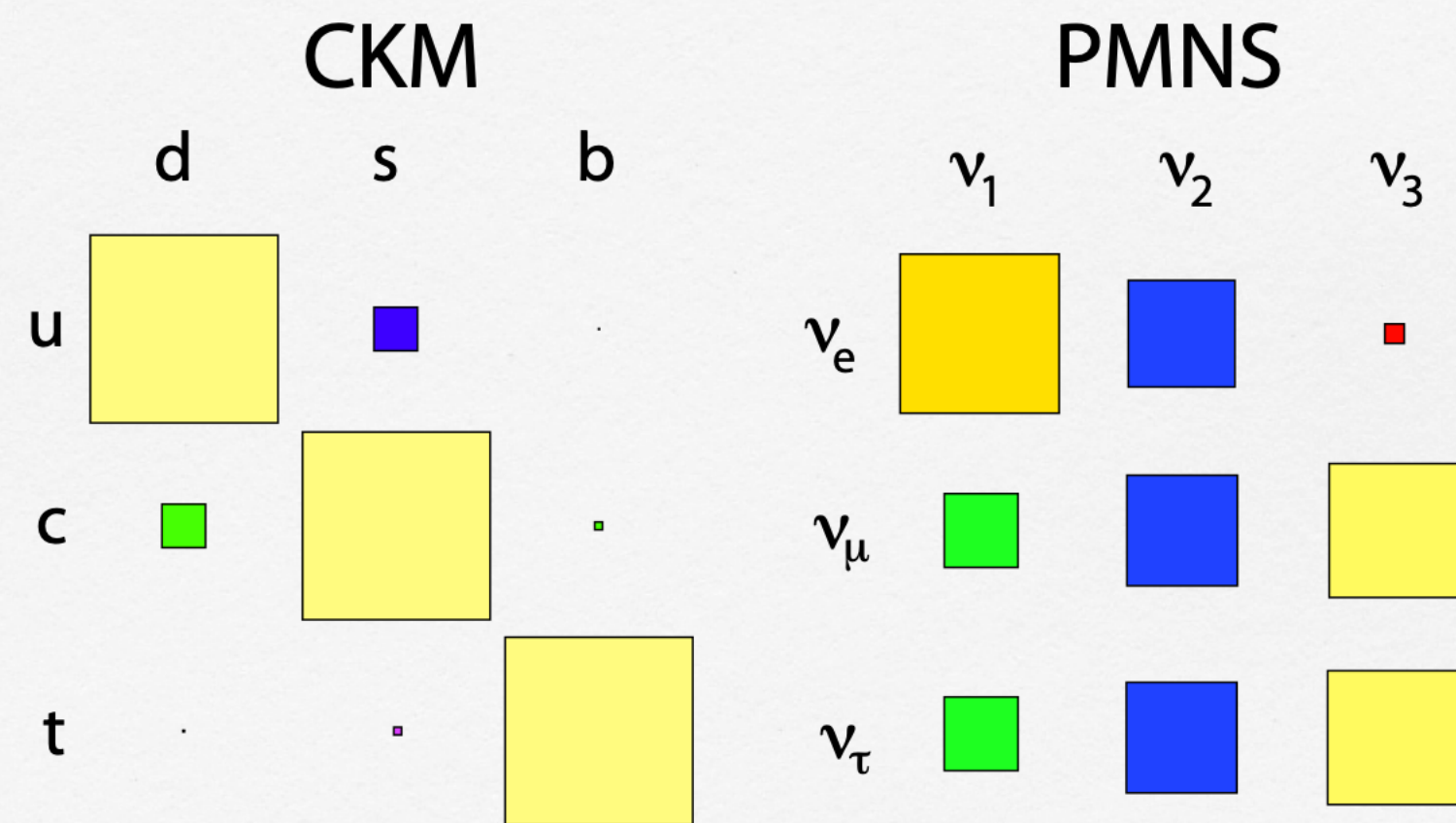
Discrete Symmetries

- Abelian: Z_n
- Non-Abelian: $A_n, S_n, \Delta(27).....$

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What is the origin of Quark and Lepton Mixing?

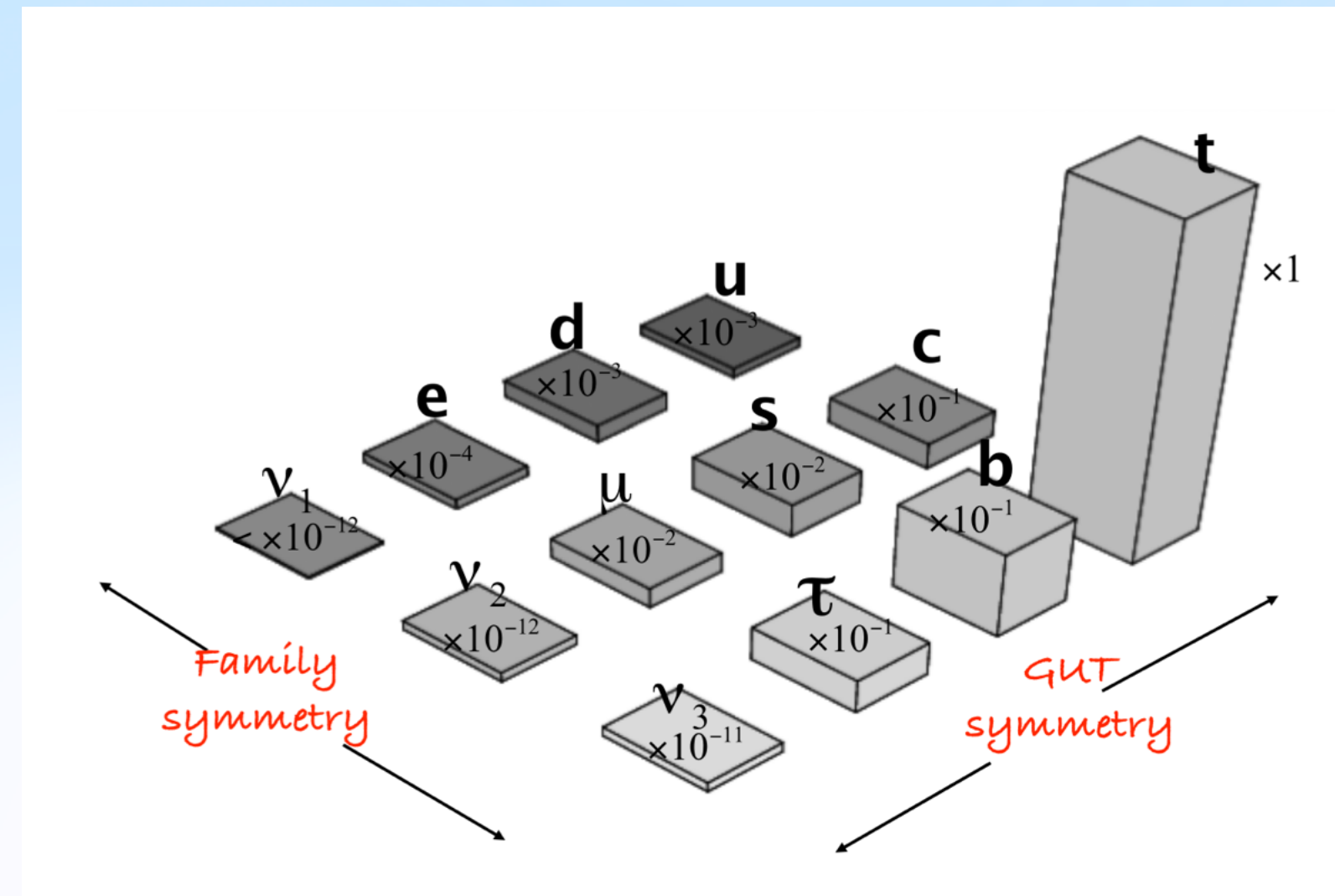
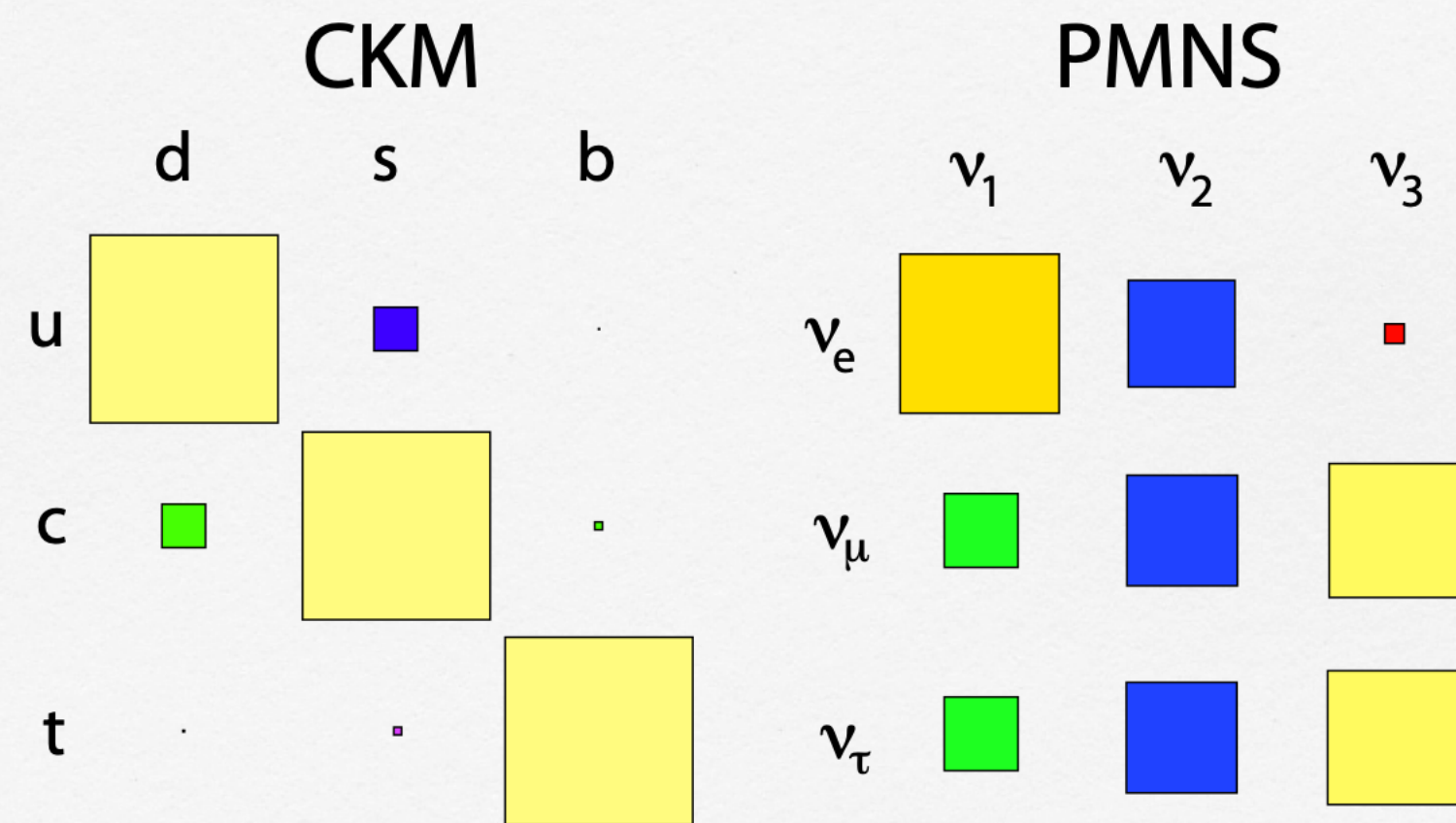


King 2015

Discrete Symmetries

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What is the origin of Quark and Lepton Mixing?



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flavour symmetries

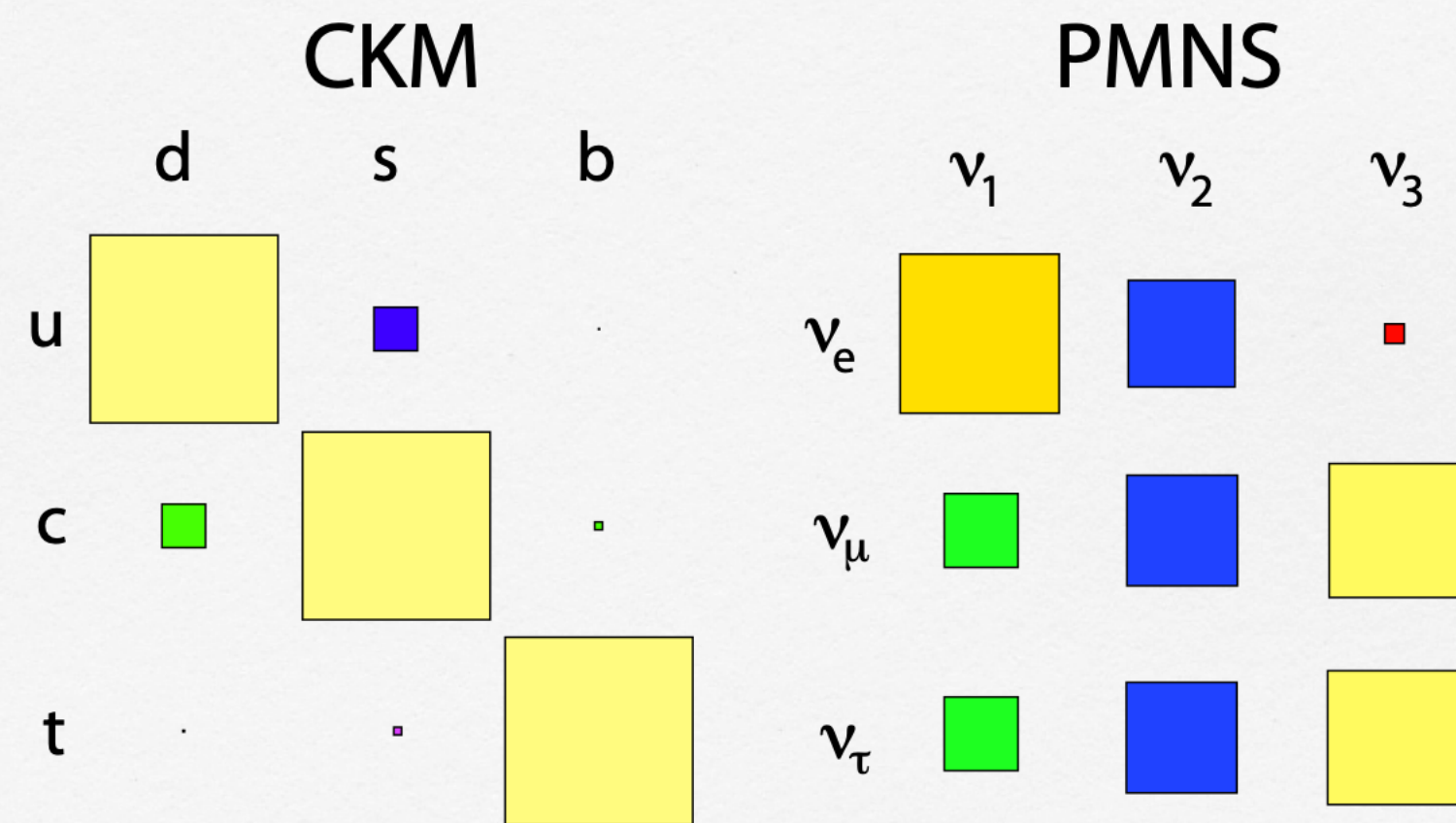
Altarelli&Feruglio 1002.0211

Petcov 1711.10806

Discrete Symmetries

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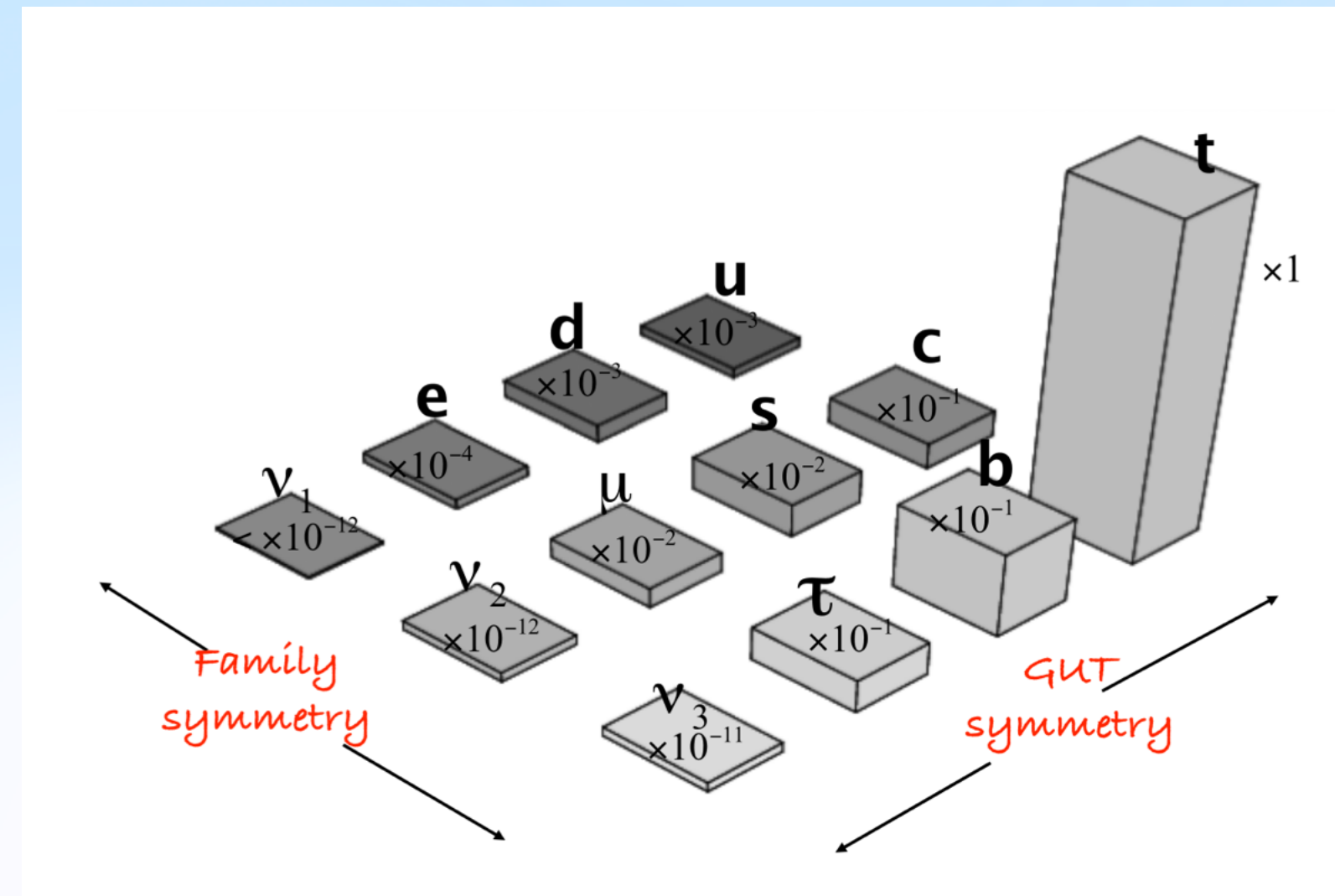
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flavour symmetries

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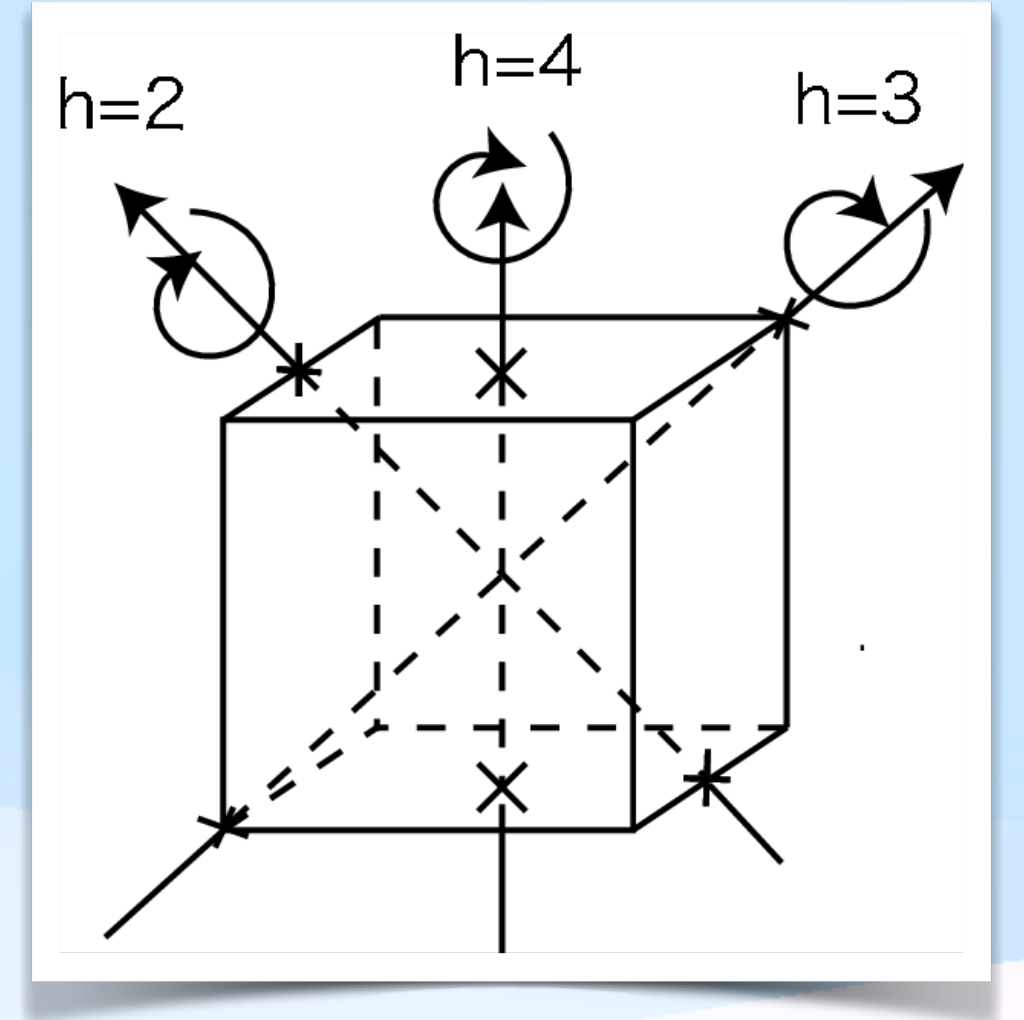
dark matter

Hirsch etc 1007.0871

S_4 scalar theory

- The octahedral/cube group S_4 :

$$T = \begin{pmatrix} 0 & 0 & 1 \\ 1 & 0 & 0 \\ 0 & 1 & 0 \end{pmatrix}, \quad S = \begin{pmatrix} 1 & 0 & 0 \\ 0 & -1 & 0 \\ 0 & 0 & -1 \end{pmatrix}, \quad U = \pm \begin{pmatrix} 1 & 0 & 0 \\ 0 & 0 & 1 \\ 0 & 1 & 0 \end{pmatrix}$$



Ishimori, etc 1003.3552

- The most general renormalisable flavon potential:

$$V(\phi) = -\frac{\mu^2}{2}I_1 + \frac{g_1}{4}I_1^2 + \frac{g_2}{2}I_2$$

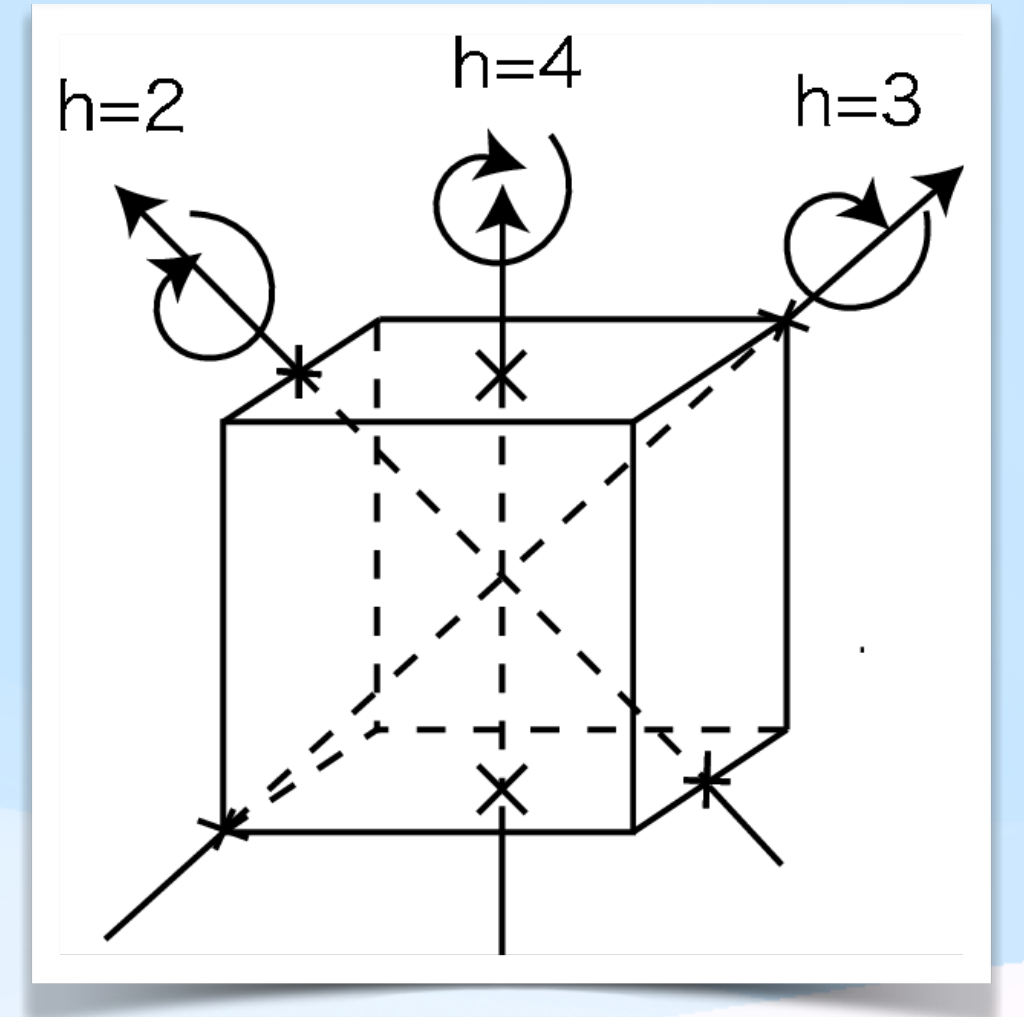
$$I_1 = \phi_1^2 + \phi_2^2 + \phi_3^2, \quad I_2 = \phi_1^2\phi_2^2 + \phi_2^2\phi_3^2 + \phi_3^2\phi_1^2$$

Also for $A_4 \times Z_2$

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$$\beta = \frac{g_2}{g_1}$$

$$I_1 = \phi_1^2 + \phi_2^2 + \phi_3^2, \quad I_2 = \phi_1^2\phi_2^2 + \phi_2^2\phi_3^2 + \phi_3^2\phi_1^2$$

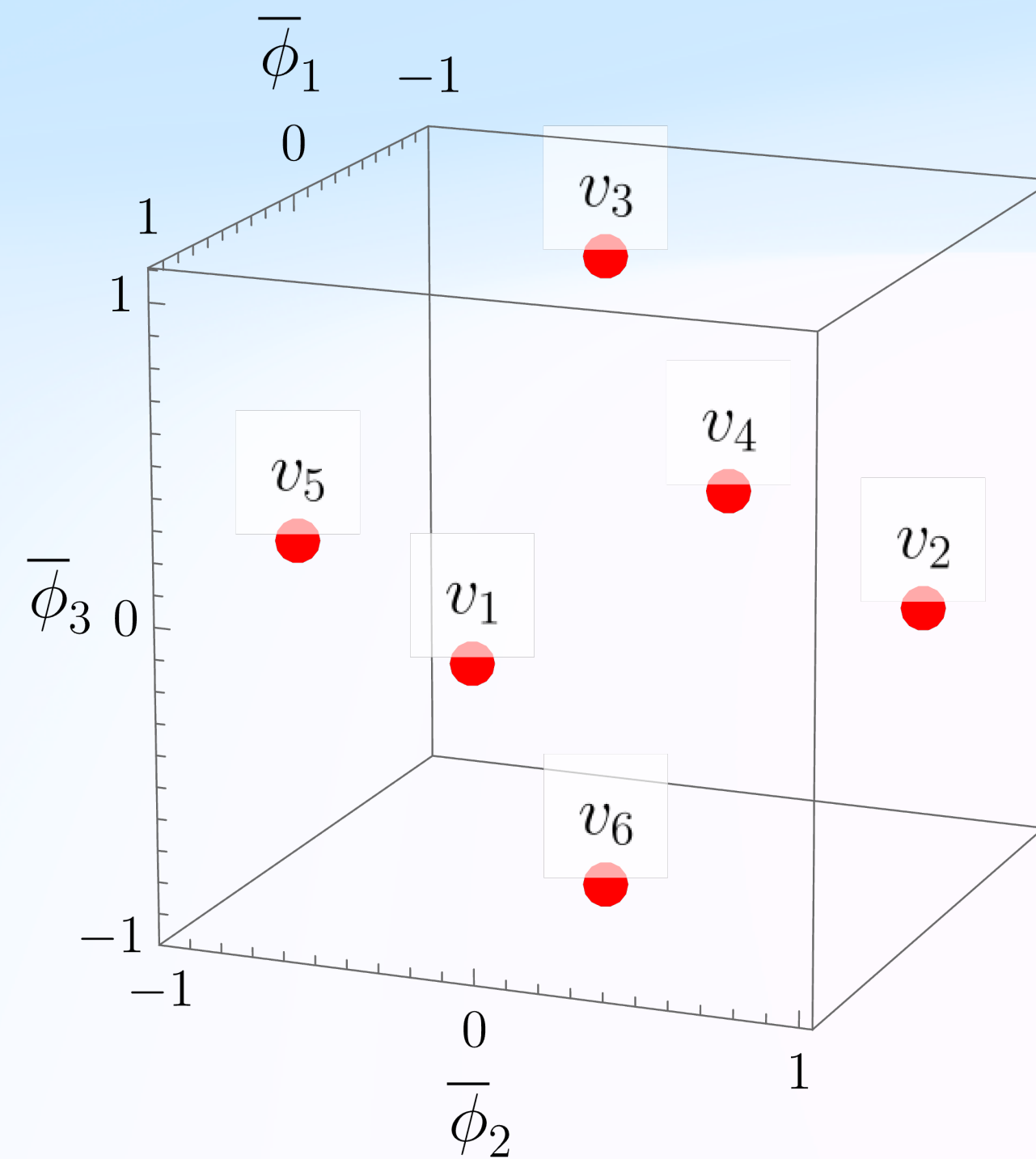
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S_4 vacuum structure

2409.16359 Fu, King, Marsili,
Pascoli, JT, Zhou

$$\left\{ \begin{pmatrix} 1 \\ 0 \\ 0 \end{pmatrix}, \begin{pmatrix} 0 \\ 1 \\ 0 \end{pmatrix}, \begin{pmatrix} 0 \\ 0 \\ 1 \end{pmatrix}, \begin{pmatrix} -1 \\ 0 \\ 0 \end{pmatrix}, \begin{pmatrix} 0 \\ -1 \\ 0 \end{pmatrix}, \begin{pmatrix} 0 \\ 0 \\ -1 \end{pmatrix} \right\} v$$

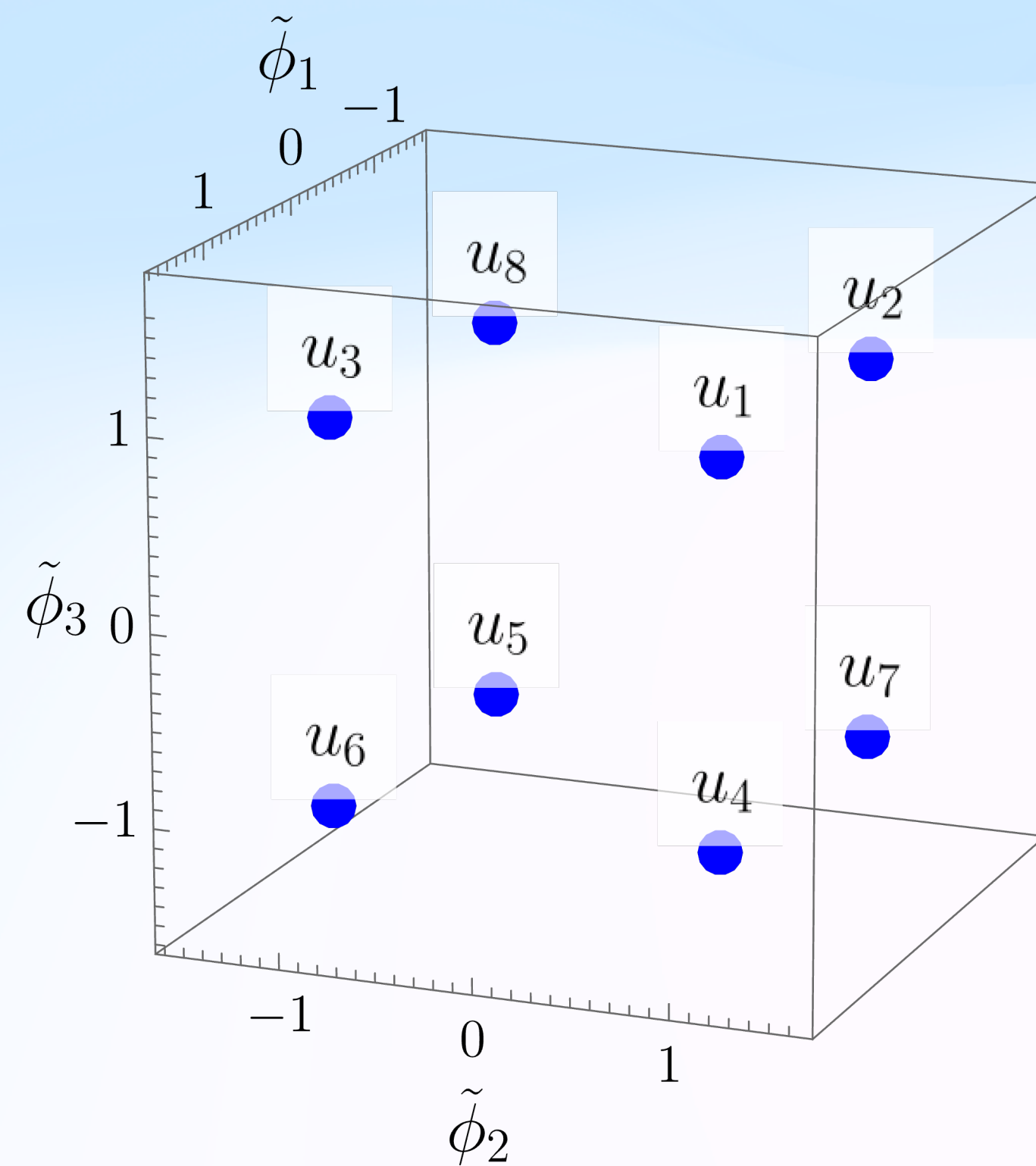
$$v = \frac{\mu}{\sqrt{g_1}}$$



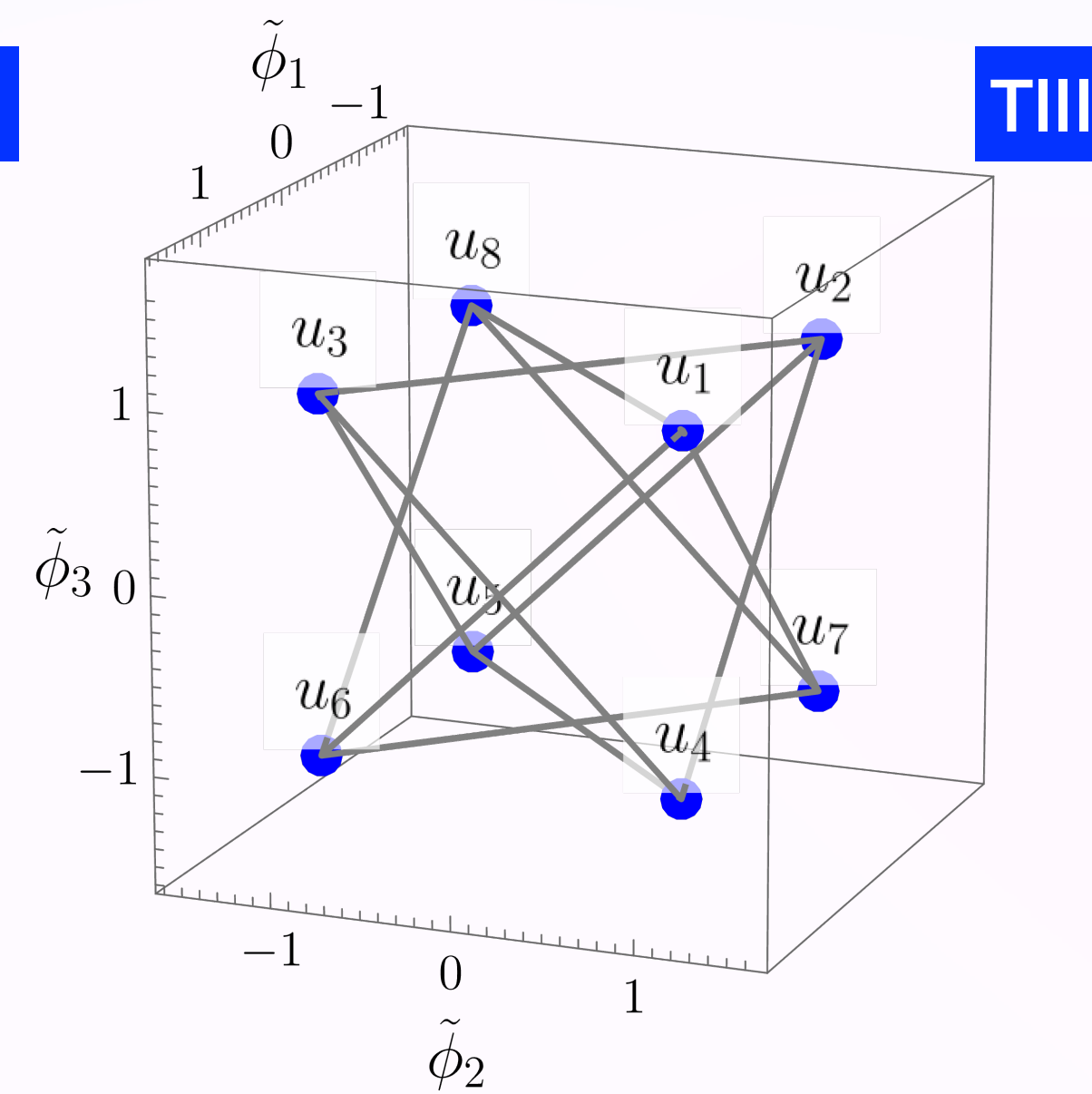
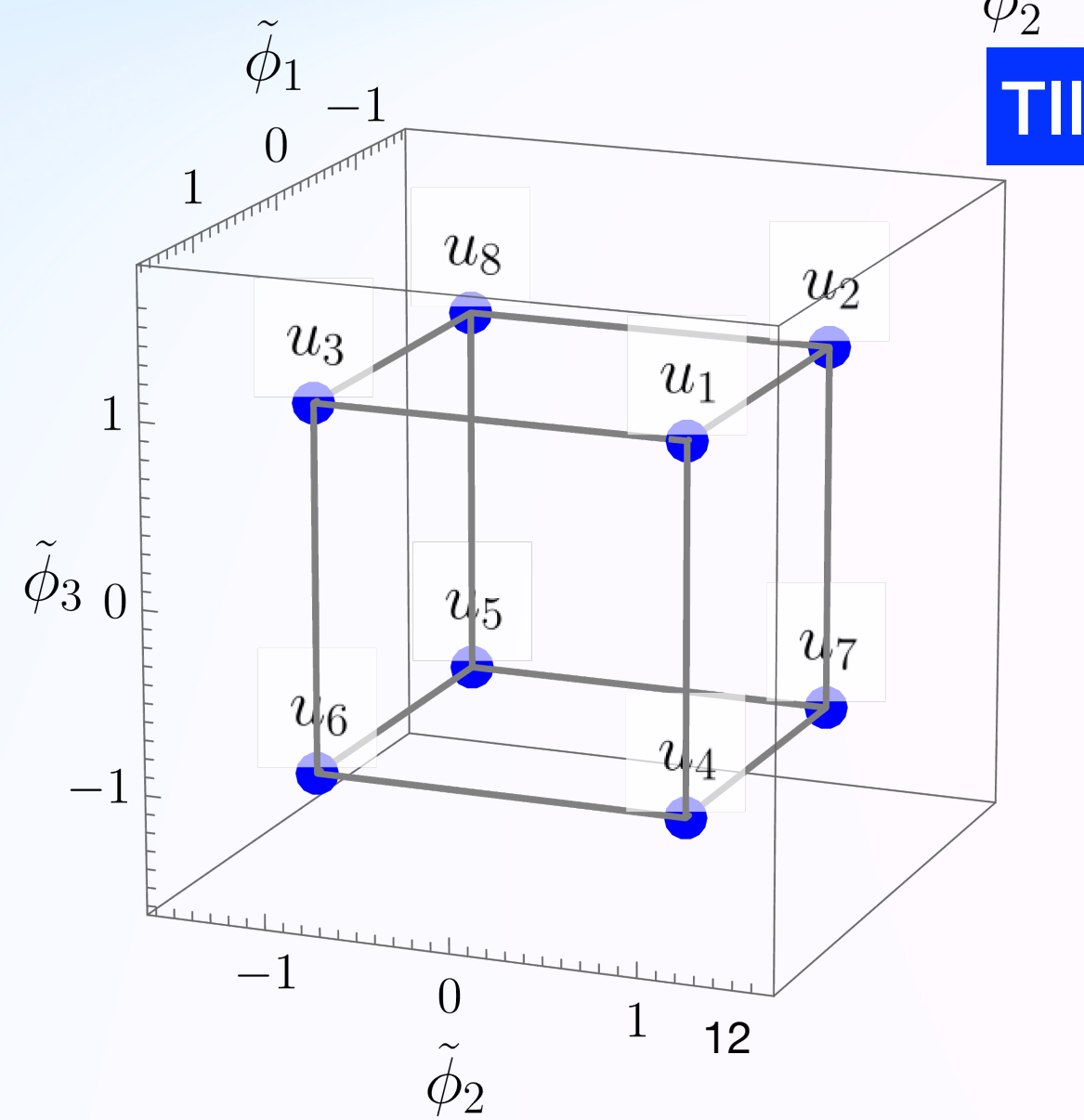
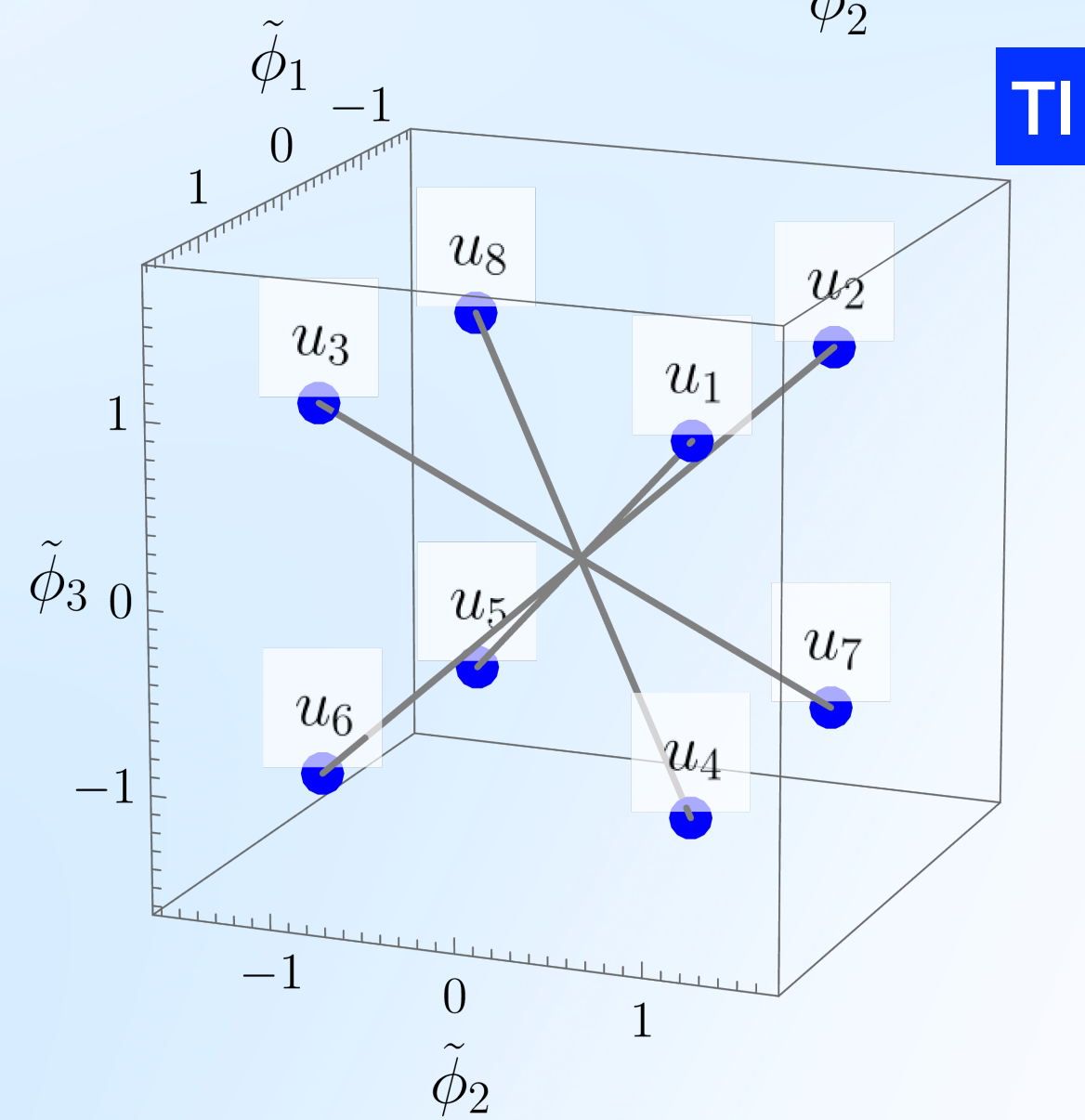
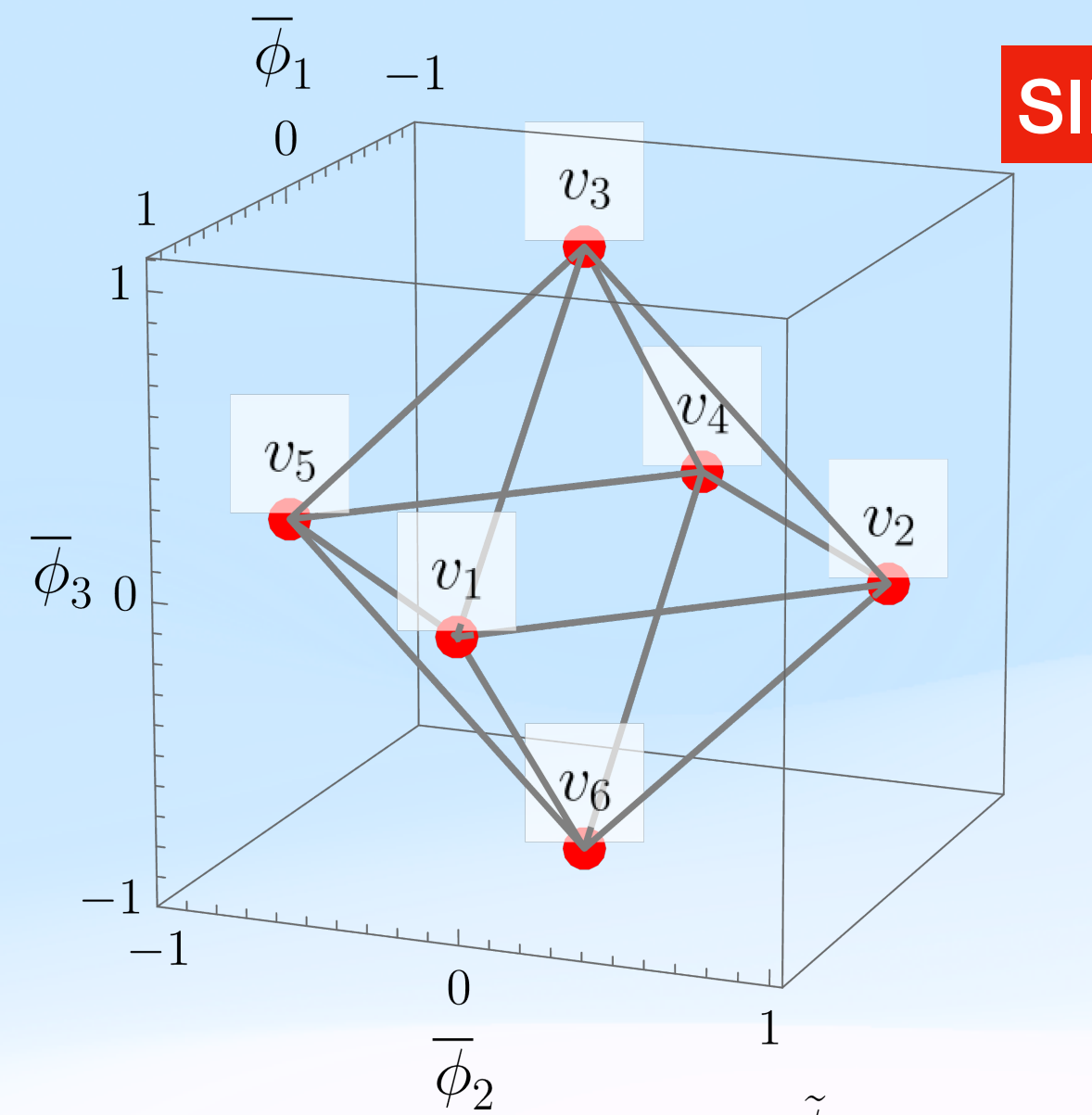
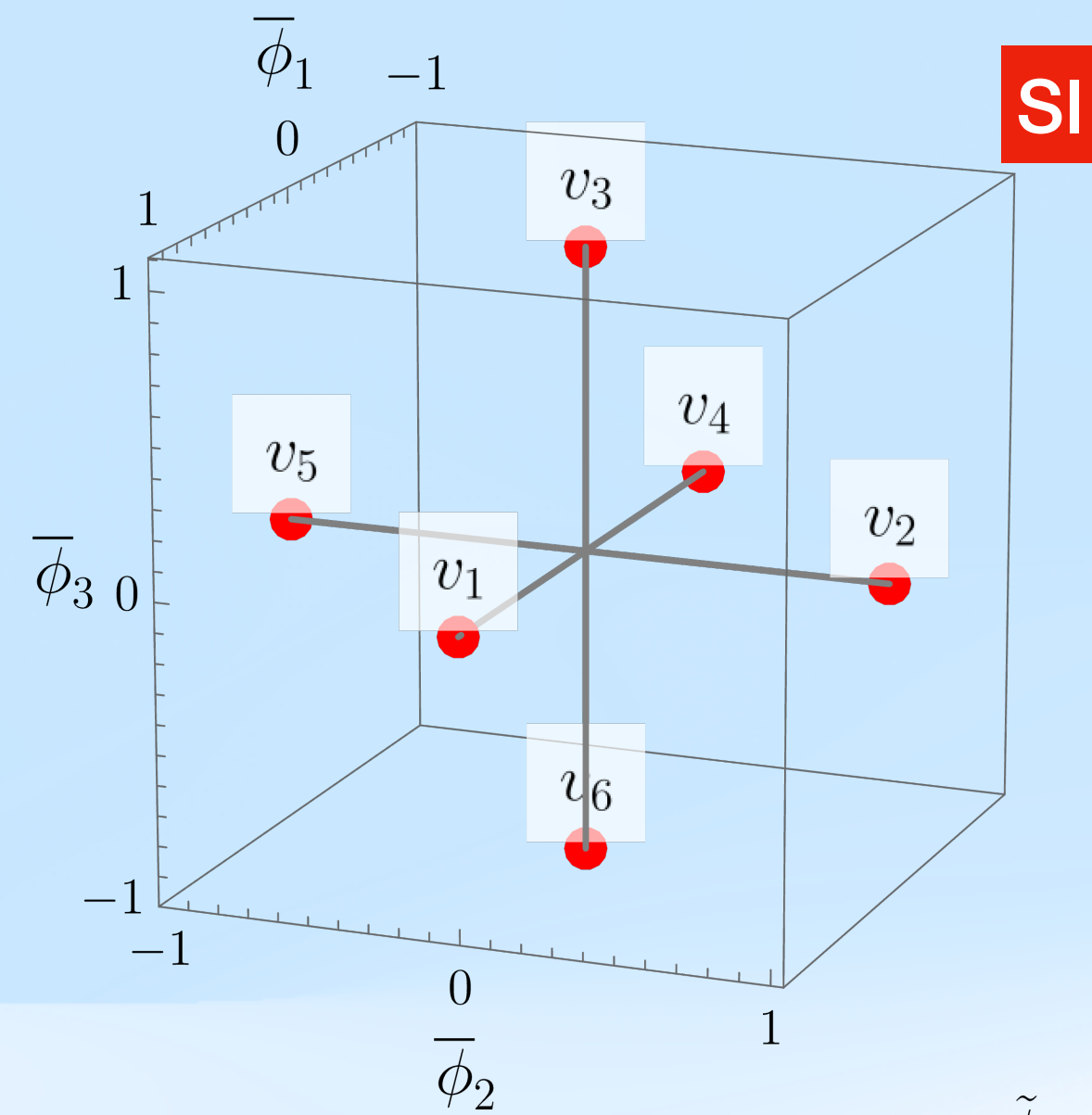
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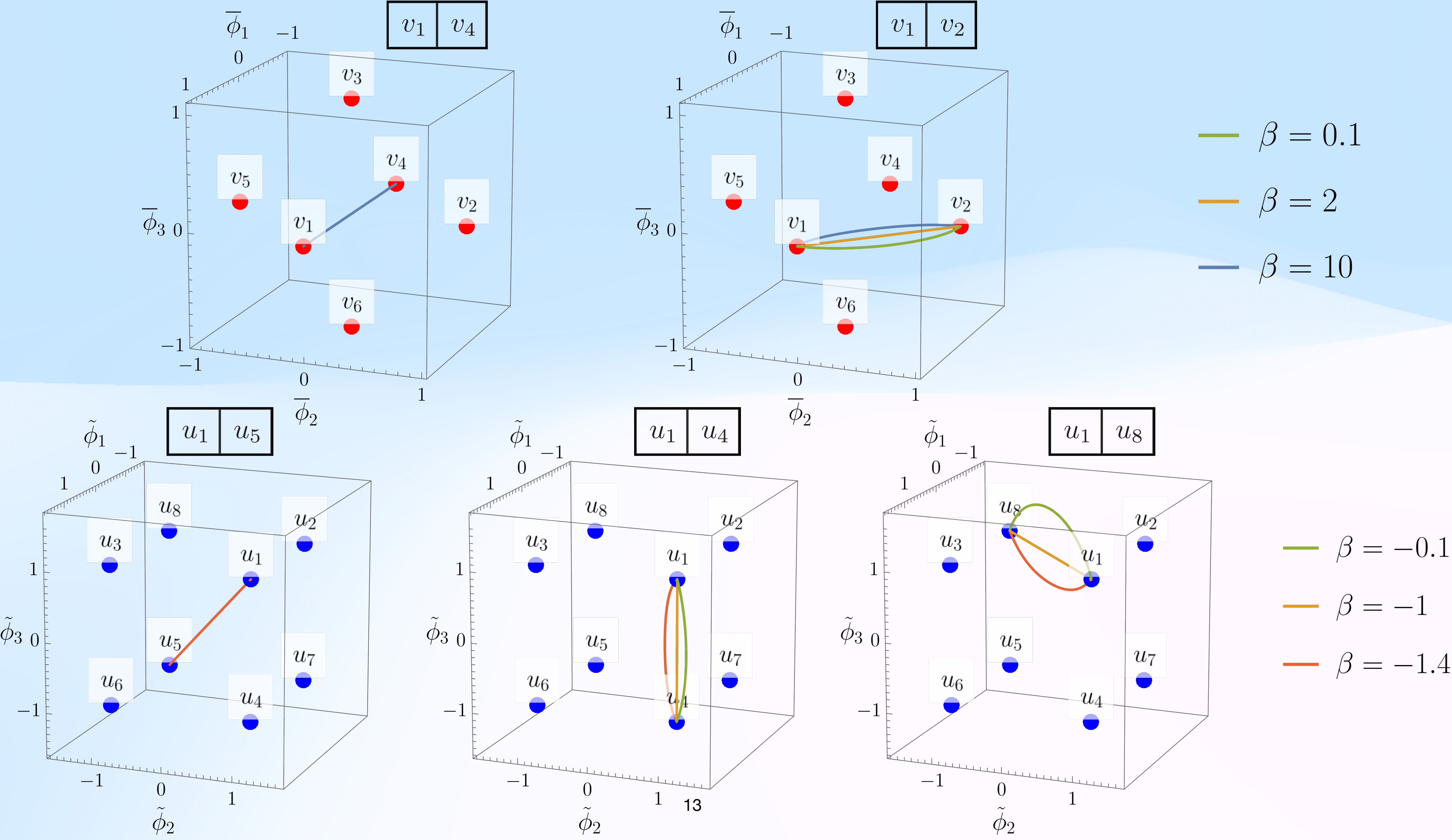
$$\left\{ \begin{pmatrix} 1 \\ 1 \\ 1 \end{pmatrix}, \begin{pmatrix} -1 \\ 1 \\ 1 \end{pmatrix}, \begin{pmatrix} 1 \\ -1 \\ 1 \end{pmatrix}, \begin{pmatrix} 1 \\ 1 \\ -1 \end{pmatrix}, \begin{pmatrix} -1 \\ -1 \\ -1 \end{pmatrix}, \begin{pmatrix} 1 \\ -1 \\ -1 \end{pmatrix}, \begin{pmatrix} -1 \\ 1 \\ -1 \end{pmatrix}, \begin{pmatrix} -1 \\ -1 \\ 1 \end{pmatrix} \right\}_u \quad u = \frac{\mu}{\sqrt{3g_1 + 2g_2}}$$



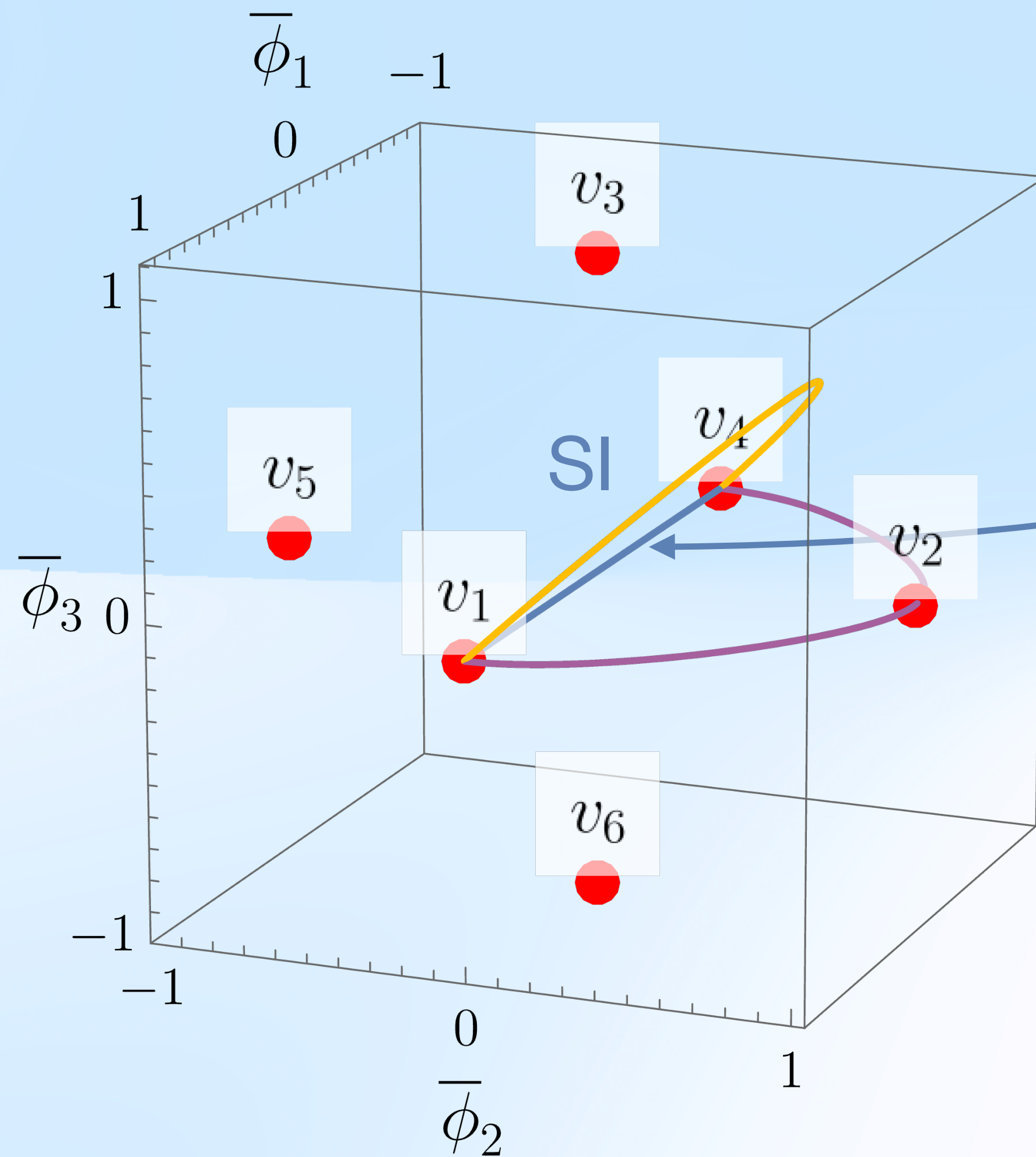
S_4 domain walls



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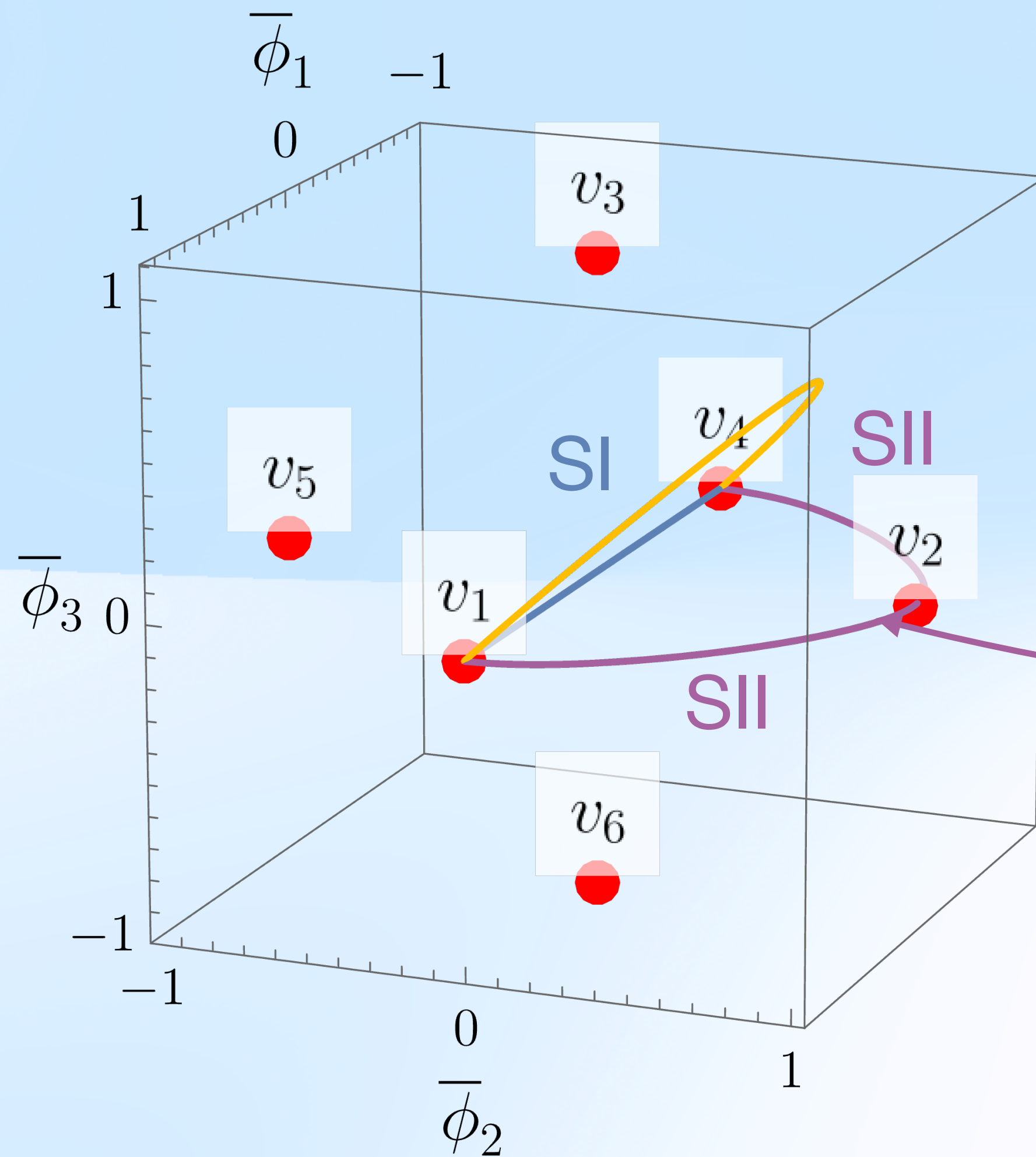


Stability of DWs



Straight line SI solution
Independent of $\beta = g_2/g_1$

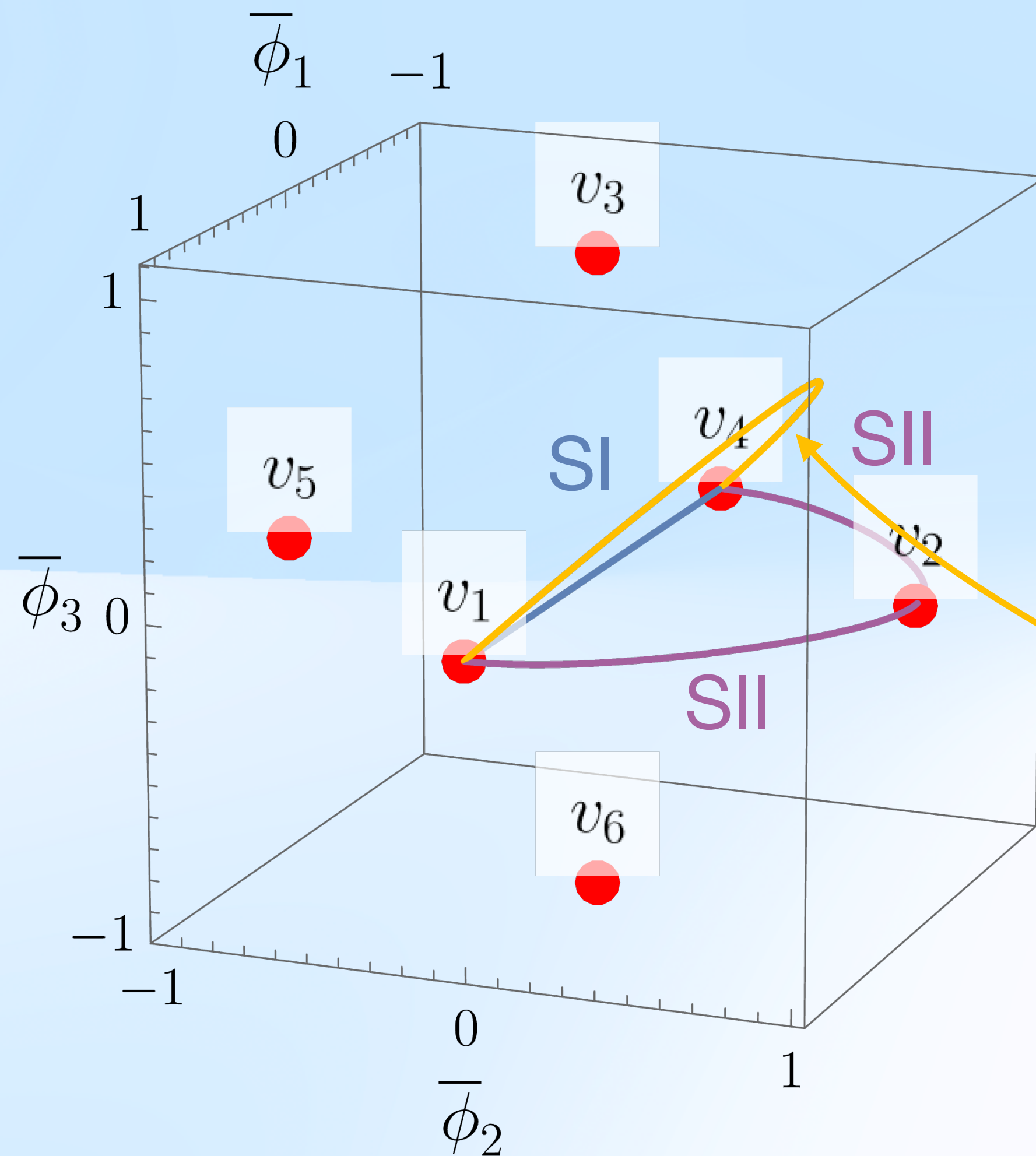
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Two SIII solutions with
pitstop at v_2

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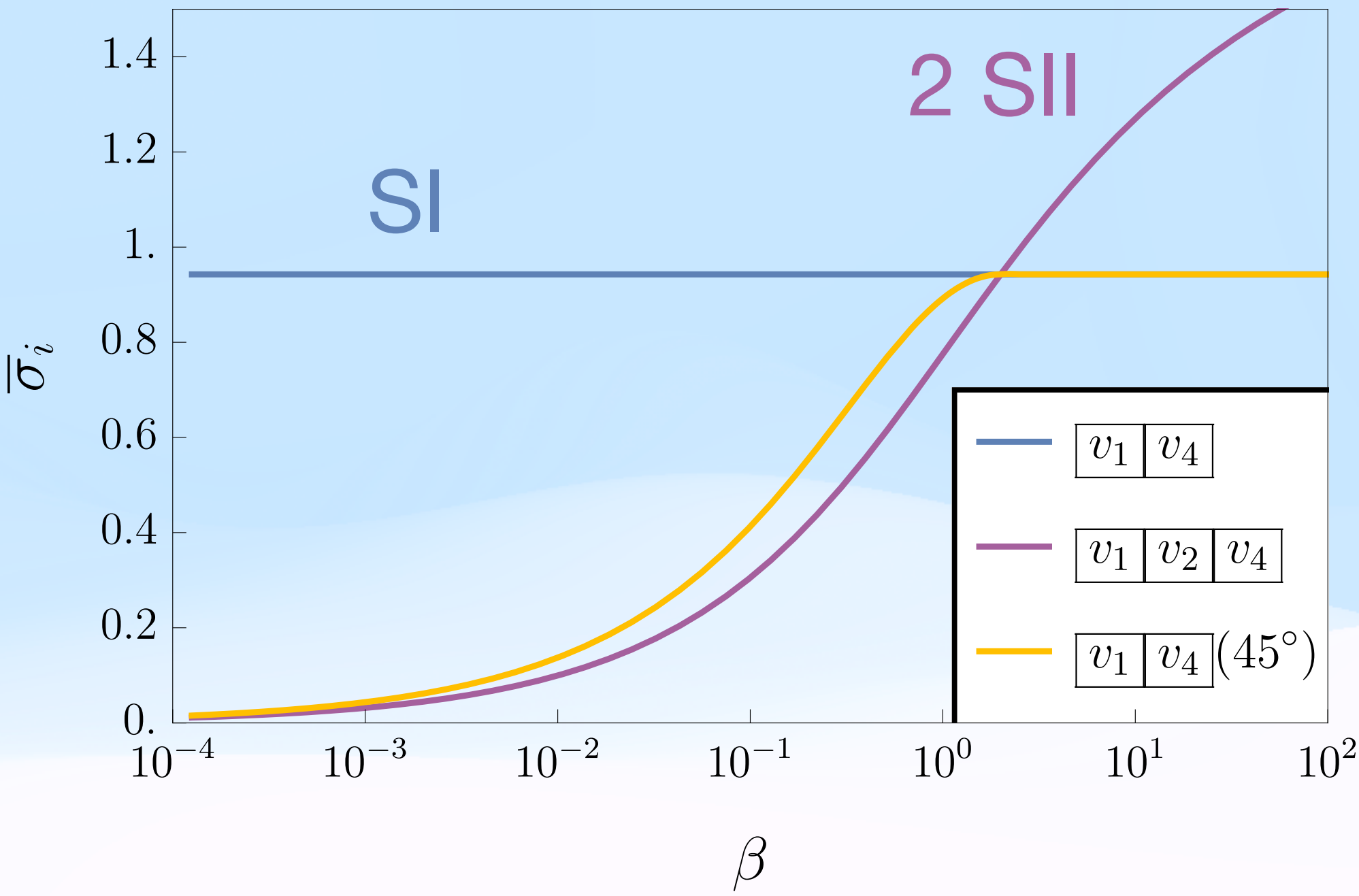
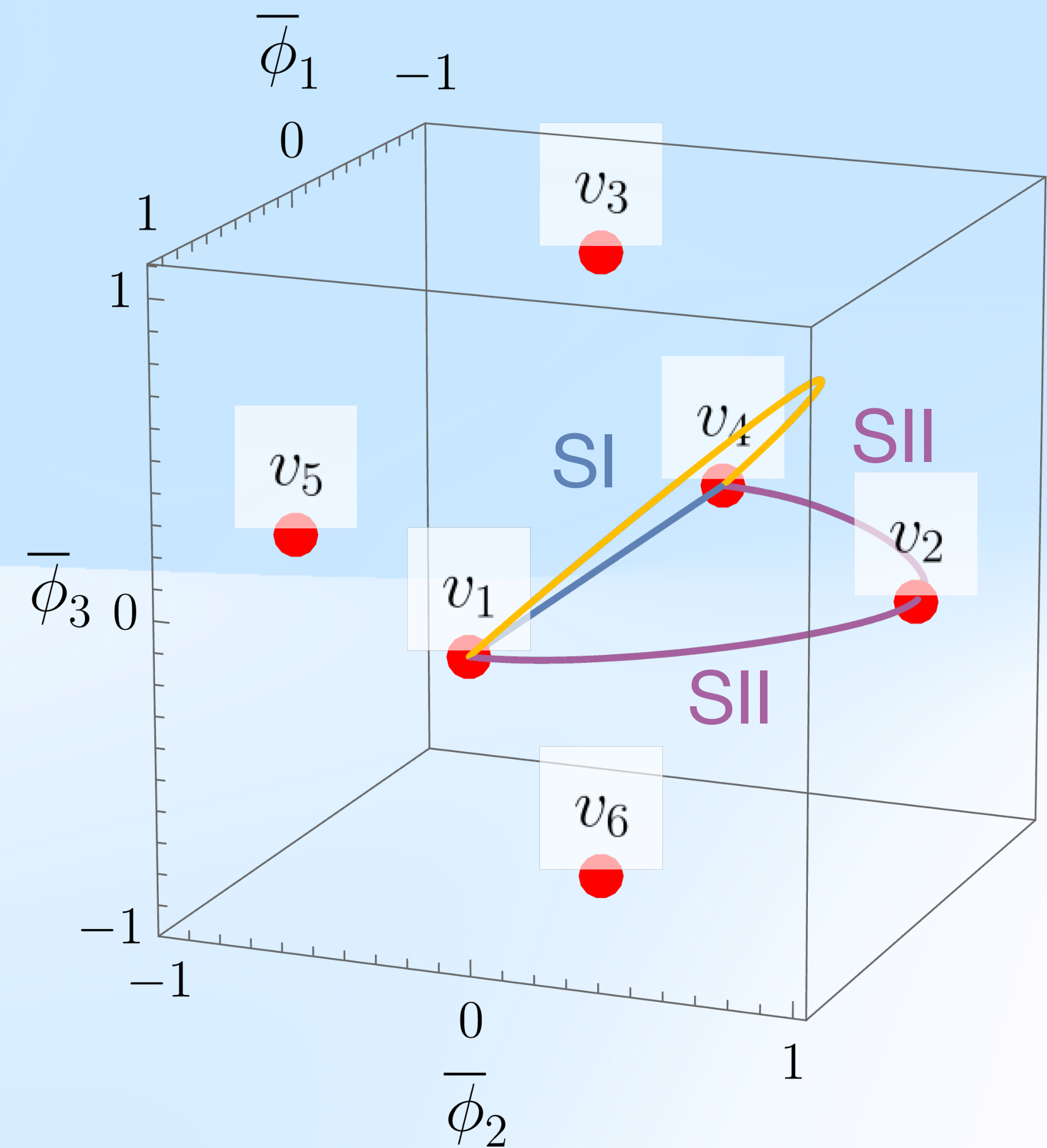


Straight line SI solution
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Two SII solutions with
pitstop at v_2

Intermediate solution (still
satisfies EoM)

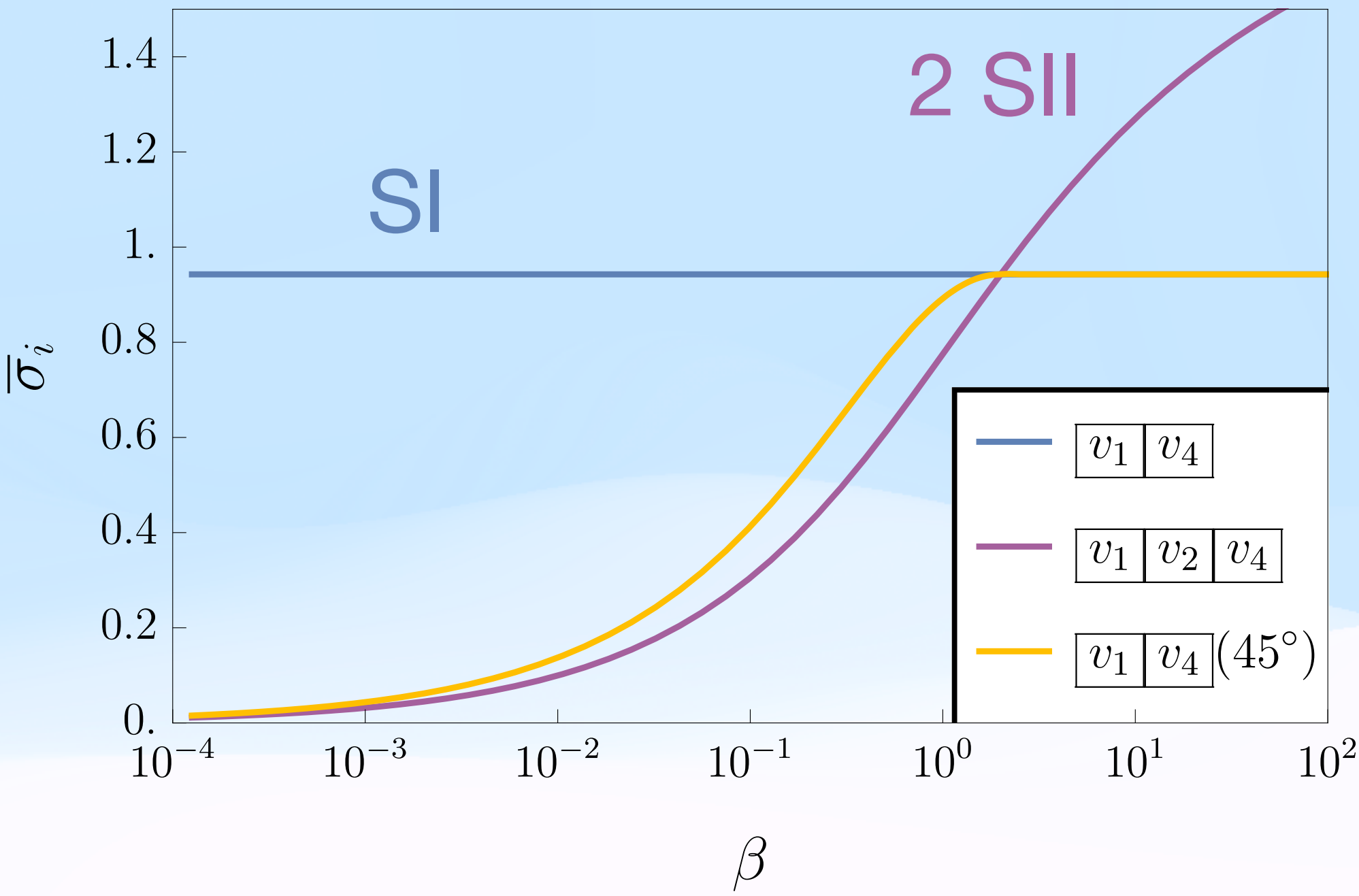
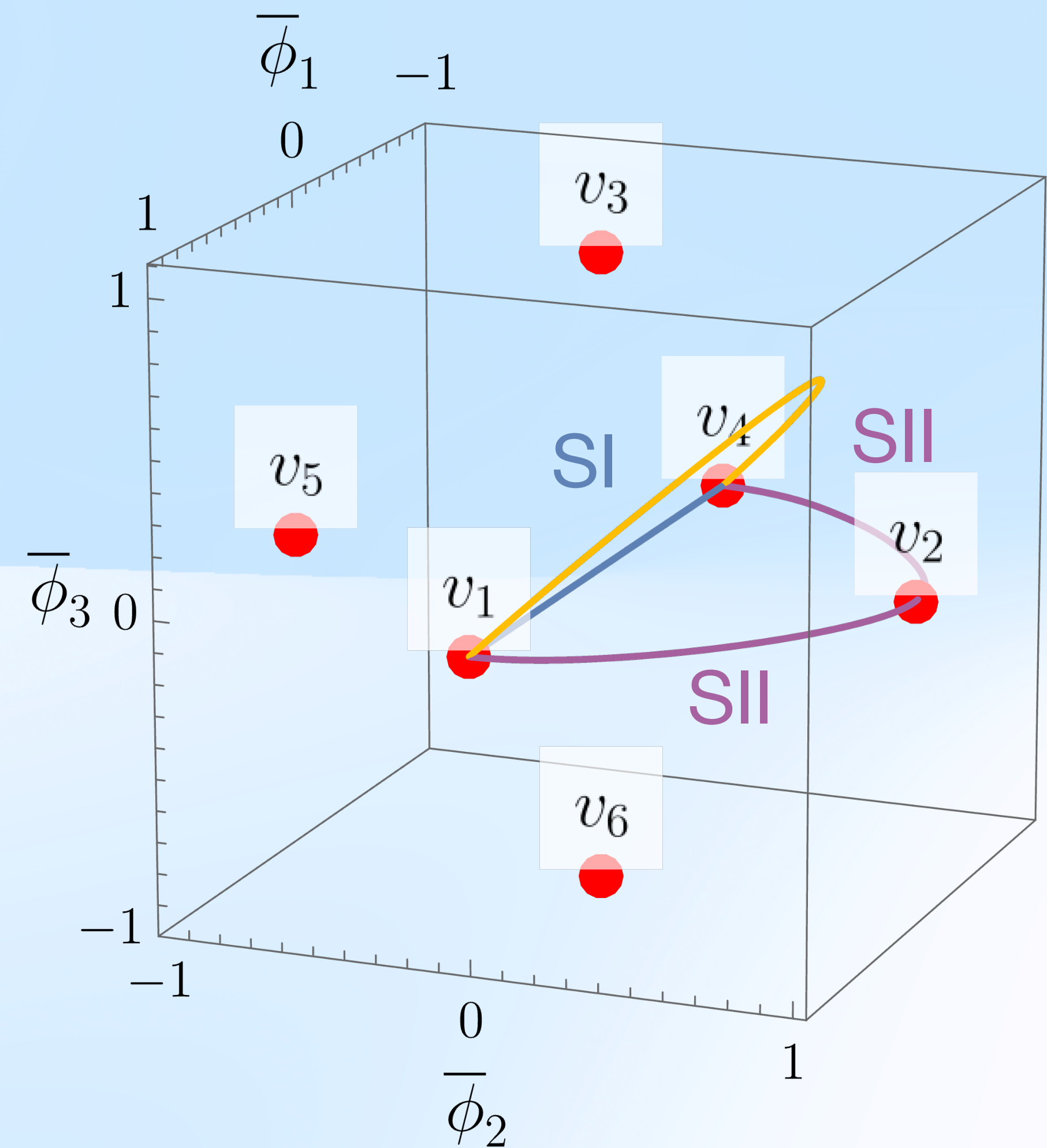
Stability of DWs



$\beta \ll 2 : \sigma(\text{SI}) > 2\sigma(\text{SII})$

SI DW unstable & would decay to SII

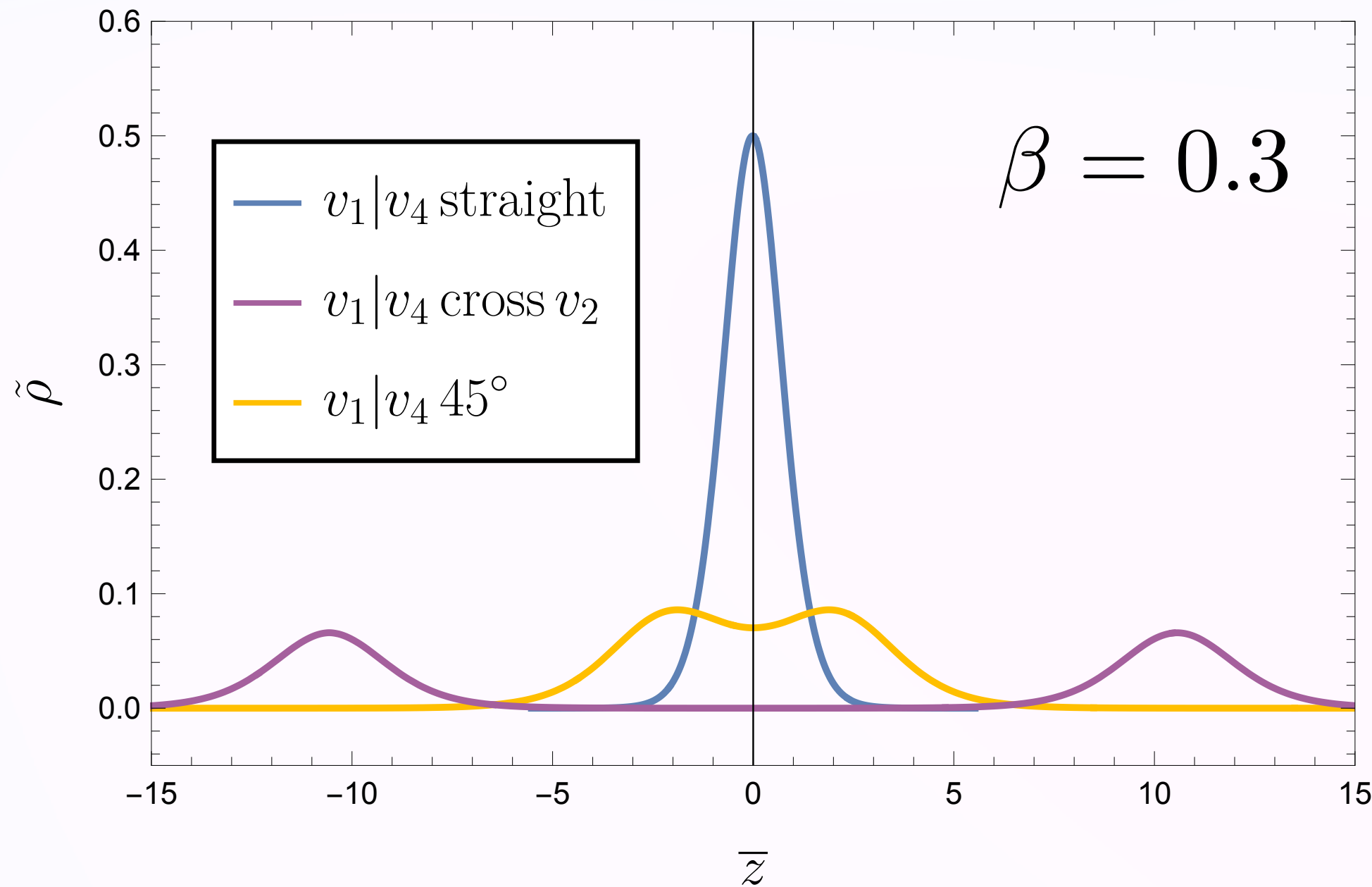
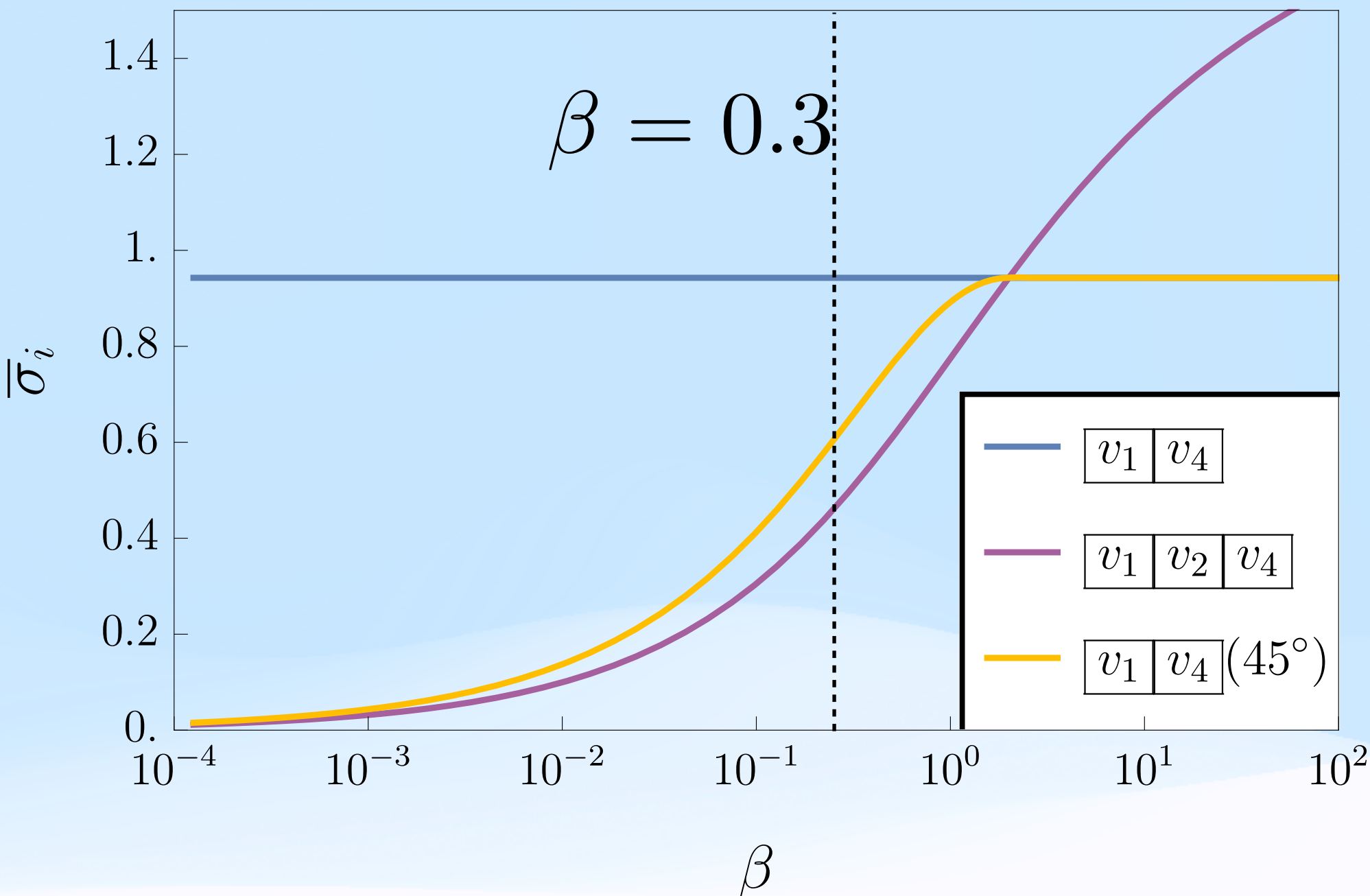
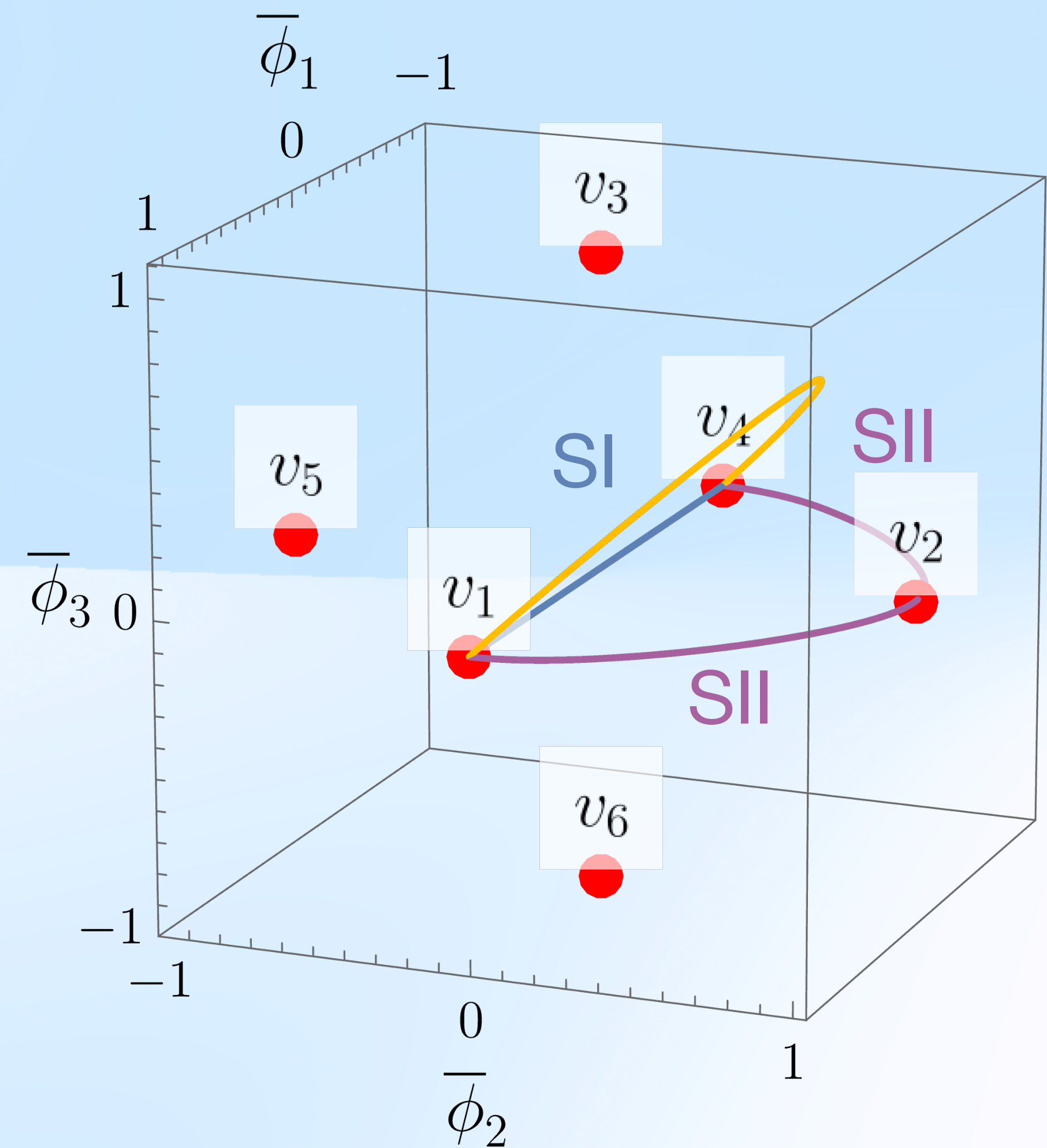
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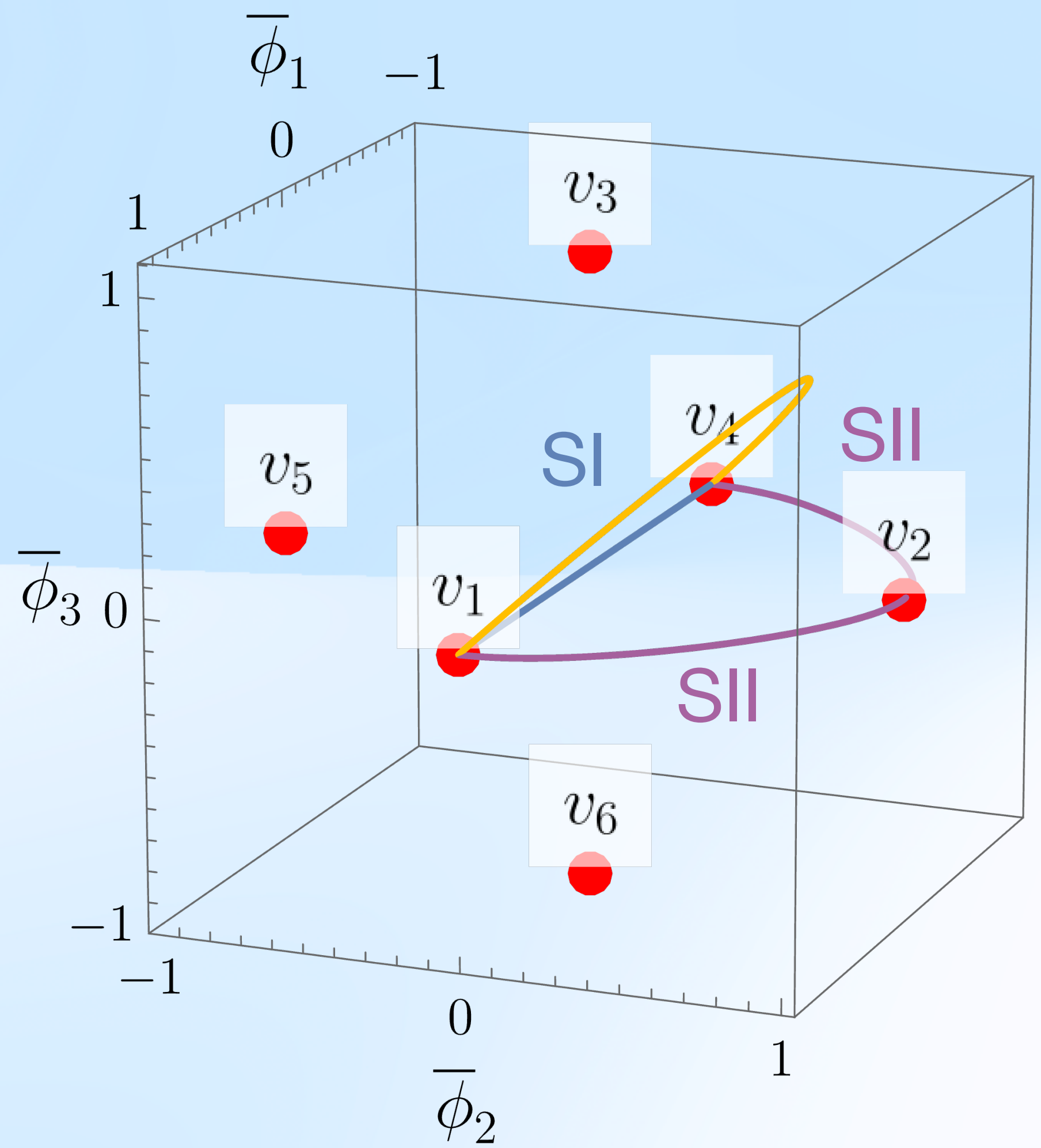
$\beta \gg 2 : \sigma(\text{SI}) < 2\sigma(\text{SII})$

SI DW stable

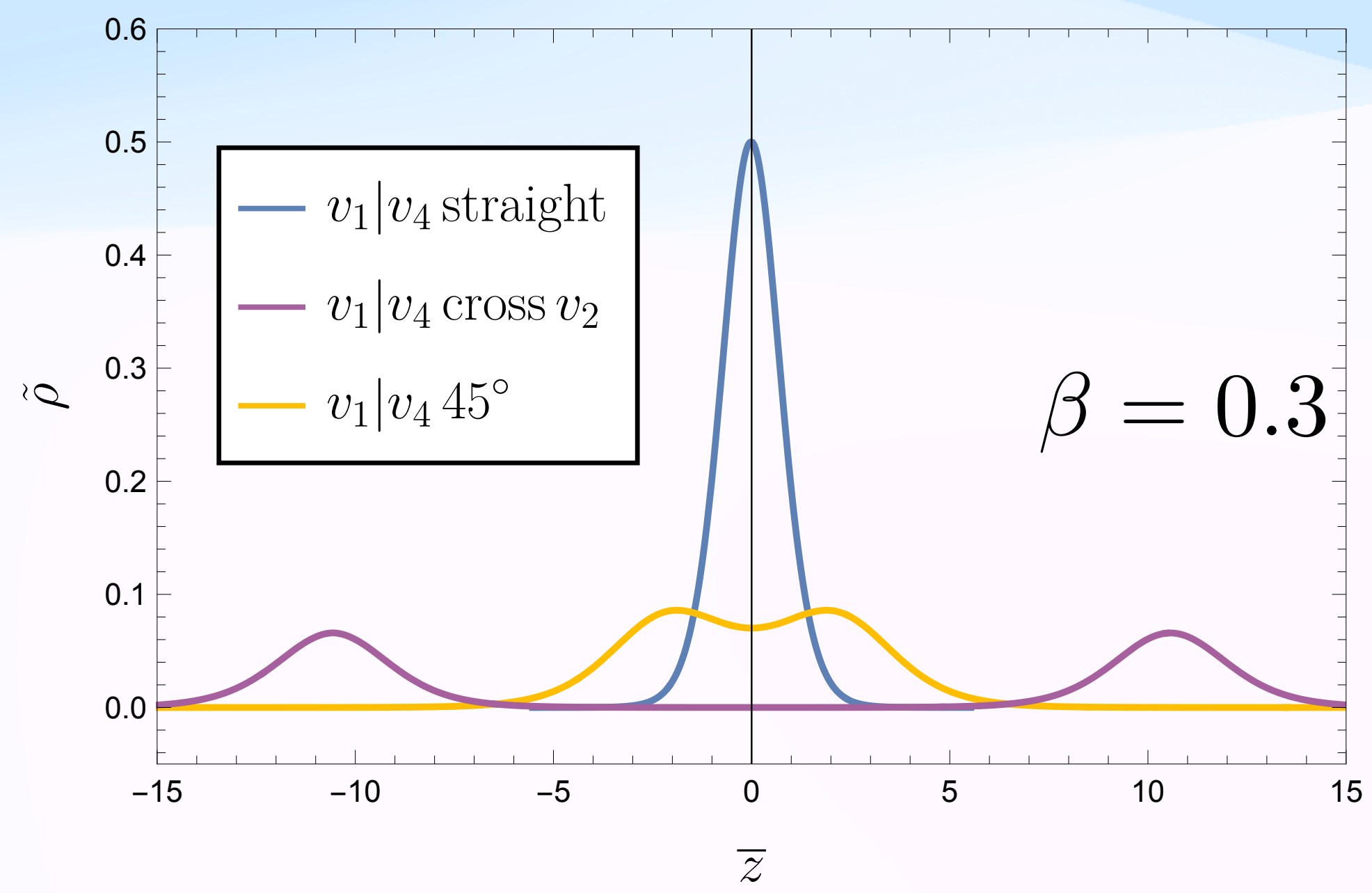
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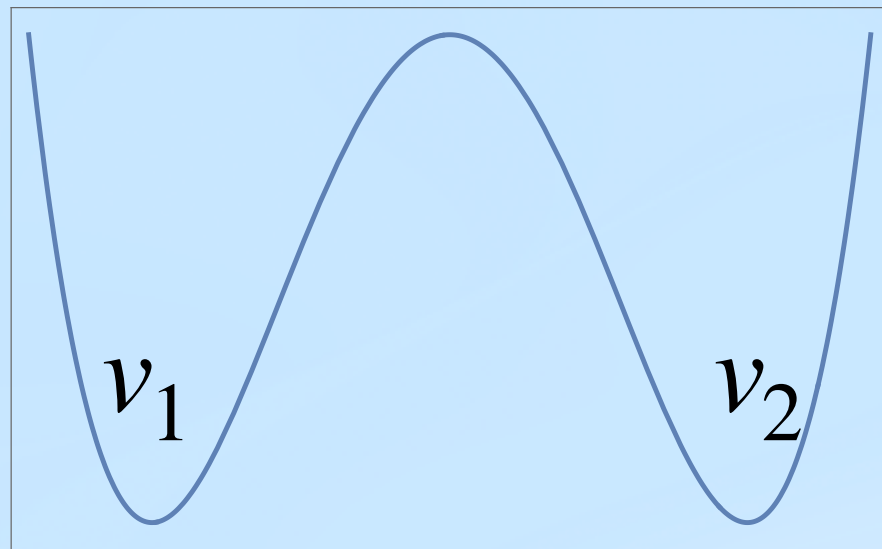


For $\beta = 0.3$, the SI-type DW will decay to two SII type DWs

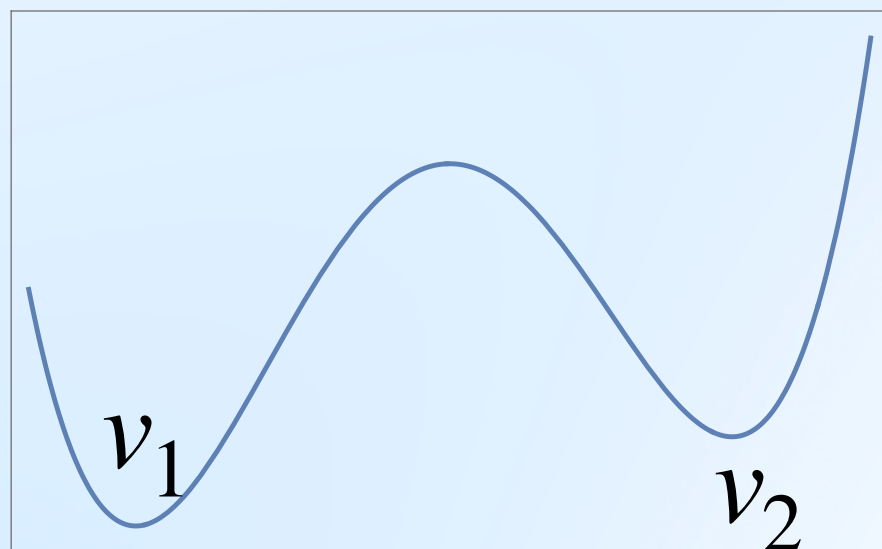


Gravitational wave from DWs

- Exact discrete symmetry $\implies$ stable DWs

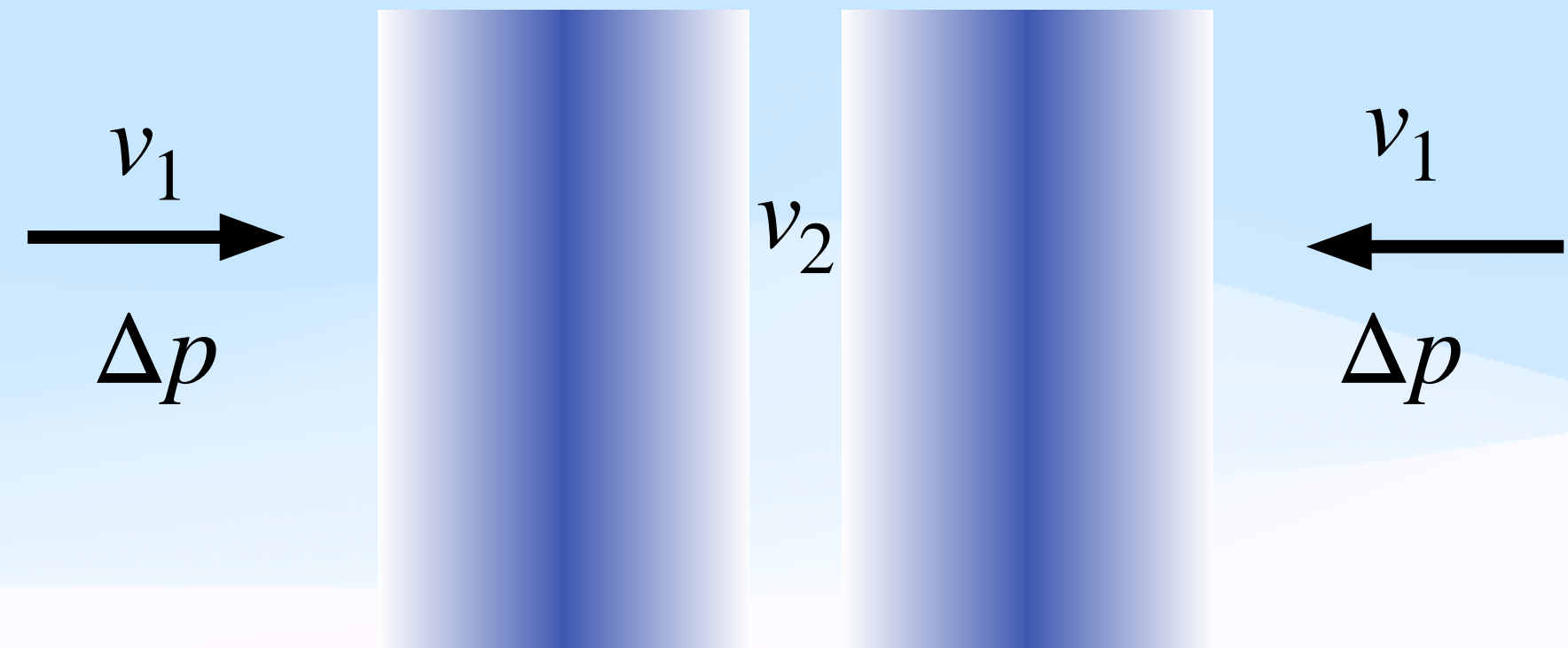


- Bias $\implies$ unstable DWs



$$V_{\text{bias}} \equiv V(v_2) - V(v_1)$$

Pressure difference $\Delta p \propto V_{\text{bias}}$

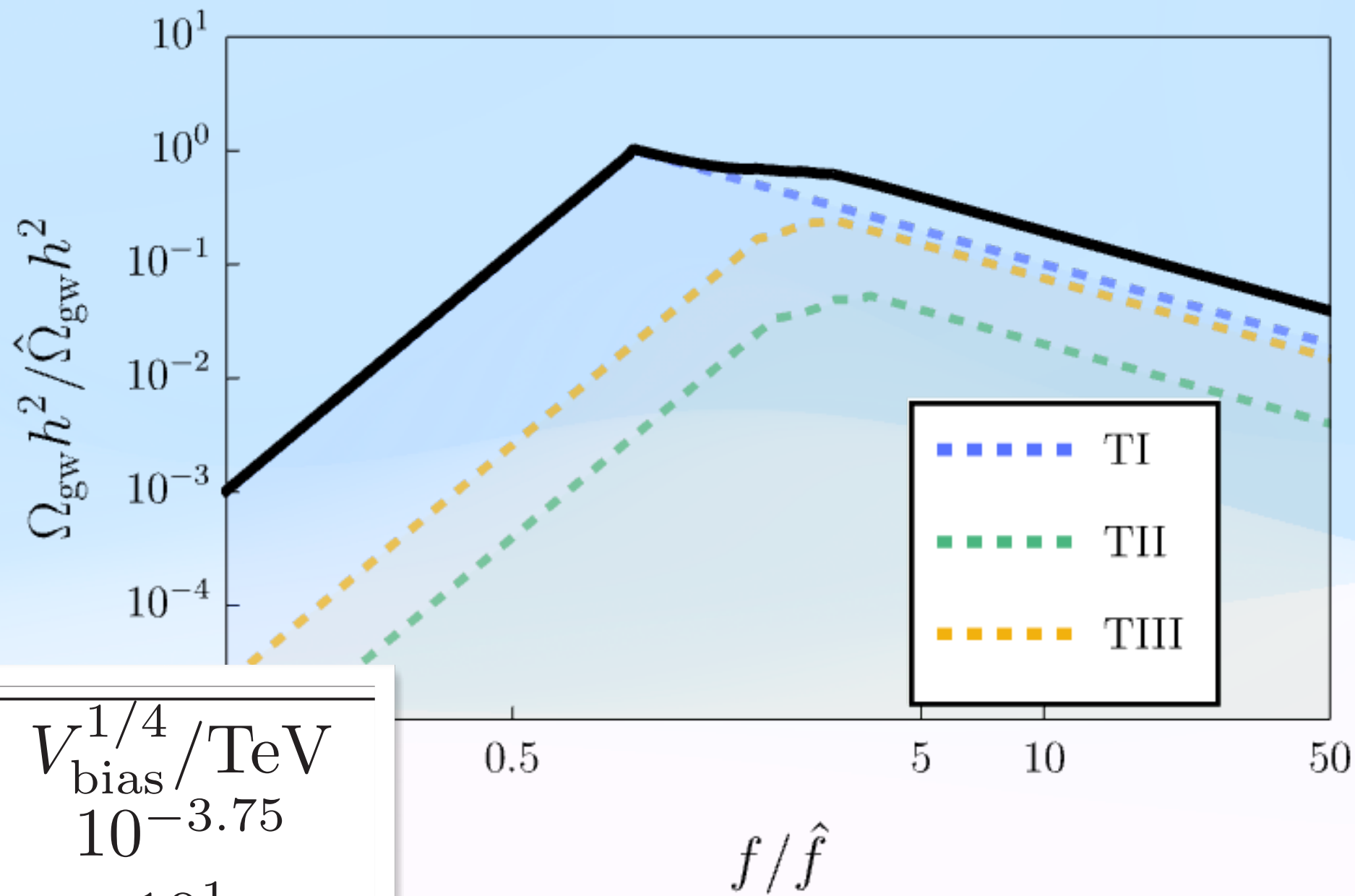
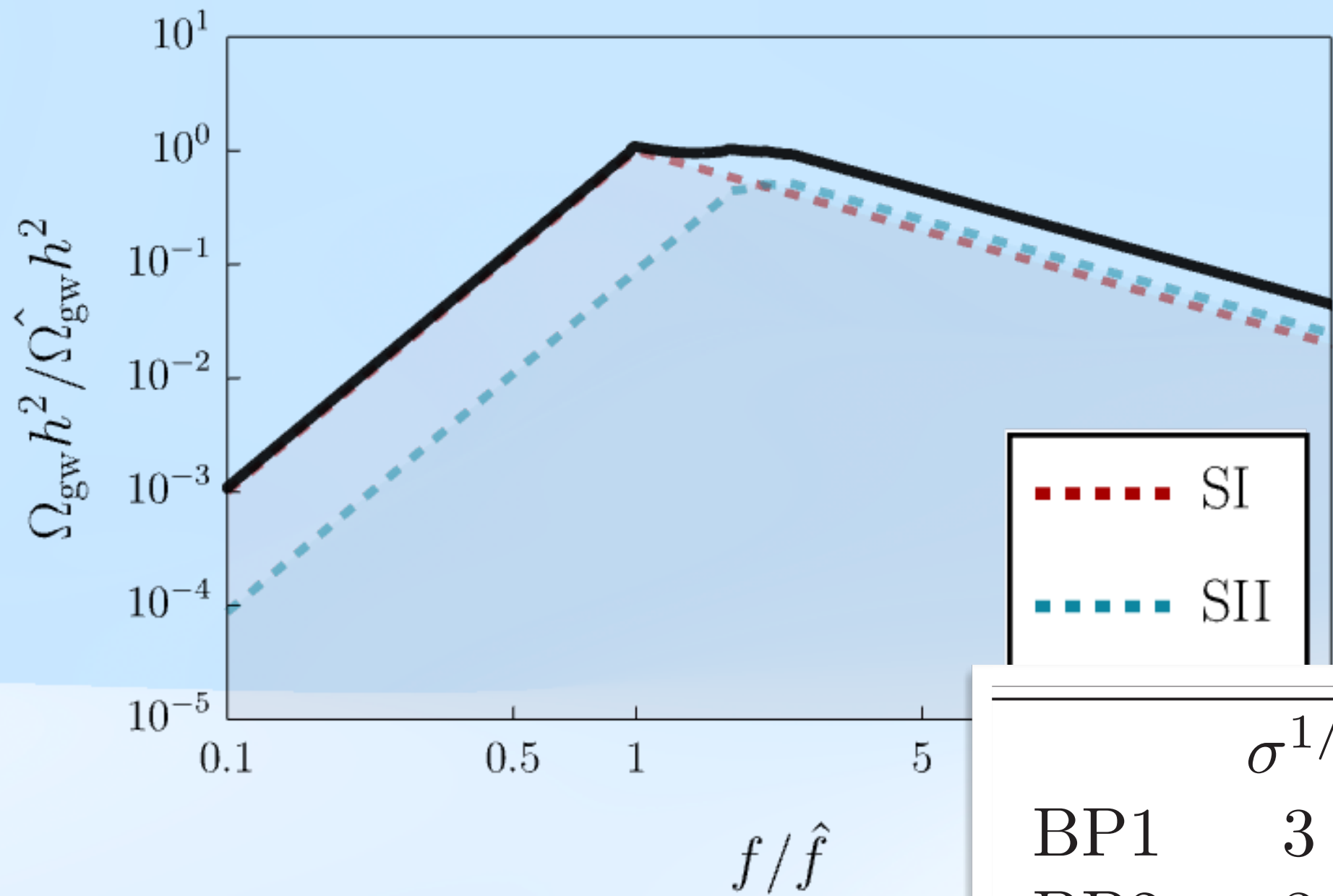


$$\Omega_{\text{GW}}^{\text{peak}} \propto \frac{\sigma^4}{V_{\text{bias}}^2} \quad f_{\text{peak}} \propto \sqrt{\frac{V_{\text{bias}}}{\sigma}}$$

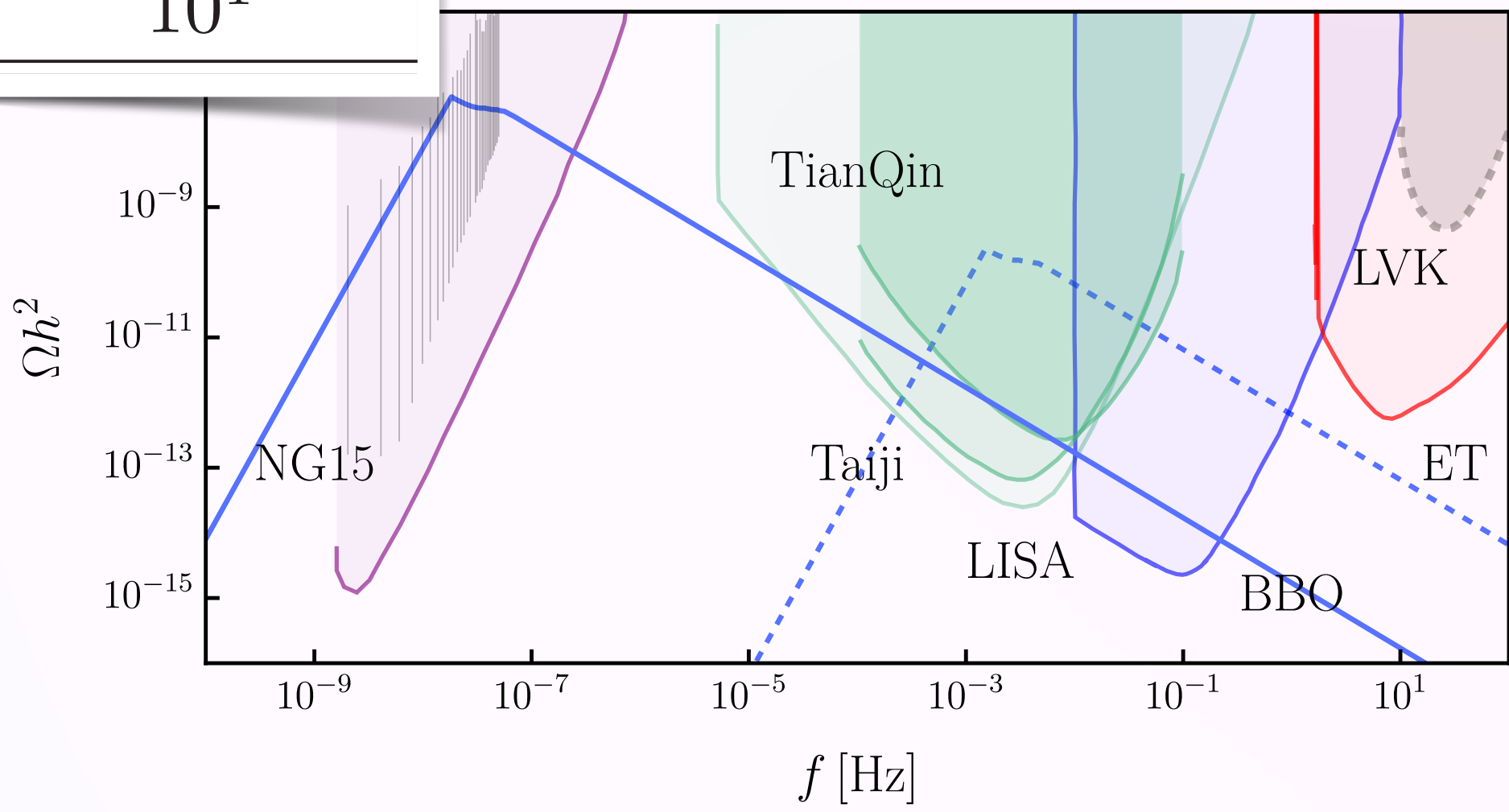
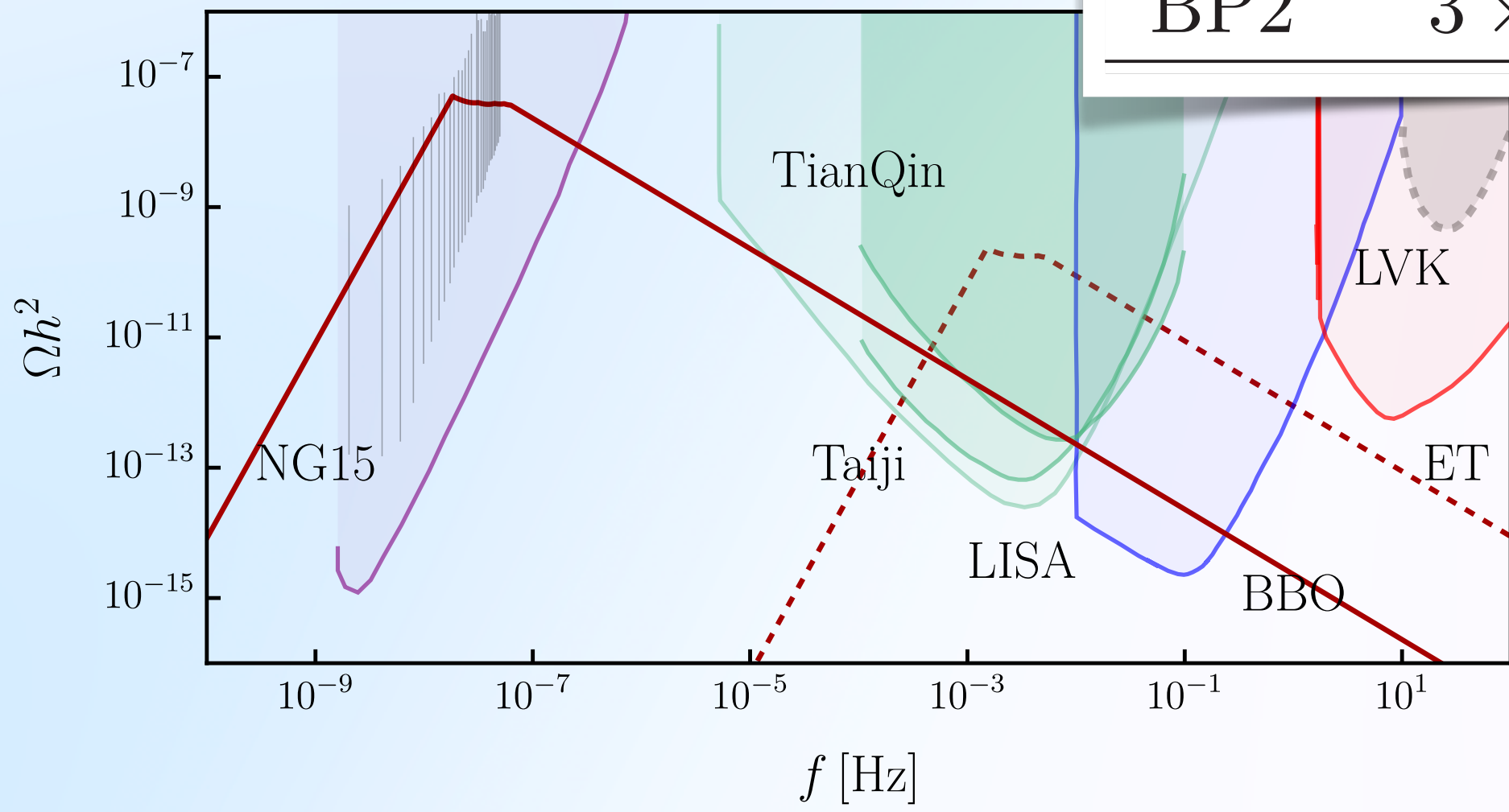
$$\begin{cases} f > f_{\text{peak}}, \Omega_{\text{GW}} \propto f^{-1} \\ f < f_{\text{peak}}, \Omega_{\text{GW}} \propto f^3 \end{cases}$$

Gravitational wave

ϵ_{12}^v	ϵ_{13}^v	ϵ_{14}^v	ϵ_{15}^v	ϵ_{16}^v	ϵ_{12}^u	ϵ_{13}^u	ϵ_{14}^u	ϵ_{15}^u	ϵ_{16}^u	ϵ_{17}^u	ϵ_{18}^u
$2\hat{\epsilon}$	$3\hat{\epsilon}$	$\hat{\epsilon}$	$4\hat{\epsilon}$	$5\hat{\epsilon}$	$2\hat{\epsilon}$	$4\hat{\epsilon}$	$6\hat{\epsilon}$	$\hat{\epsilon}$	$3\hat{\epsilon}$	$5\hat{\epsilon}$	$7\hat{\epsilon}$



	$\sigma^{1/3}/\text{TeV}$	$V_{\text{bias}}^{1/4}/\text{TeV}$
BP1	3×10^2	$10^{-3.75}$
BP2	3×10^5	10^1



Summary and Outlook

- Non-abelian DWs have more interesting and non-trivial structure and phenomena
- In certain range of parameter space, unstable DWs can show up
- If the DWs are stable, they can give rise to a unique multi-peak GW signal
- The signature of GW raised by unstable domain walls is still unexplored
- In realistic flavour models, the flavon fields are complex scalars and there may be more than one multiplets

