

Nuclear deformation effects of vector mesons in UPC

AND

New nonlinear response of anisotropic flow to eccentricity

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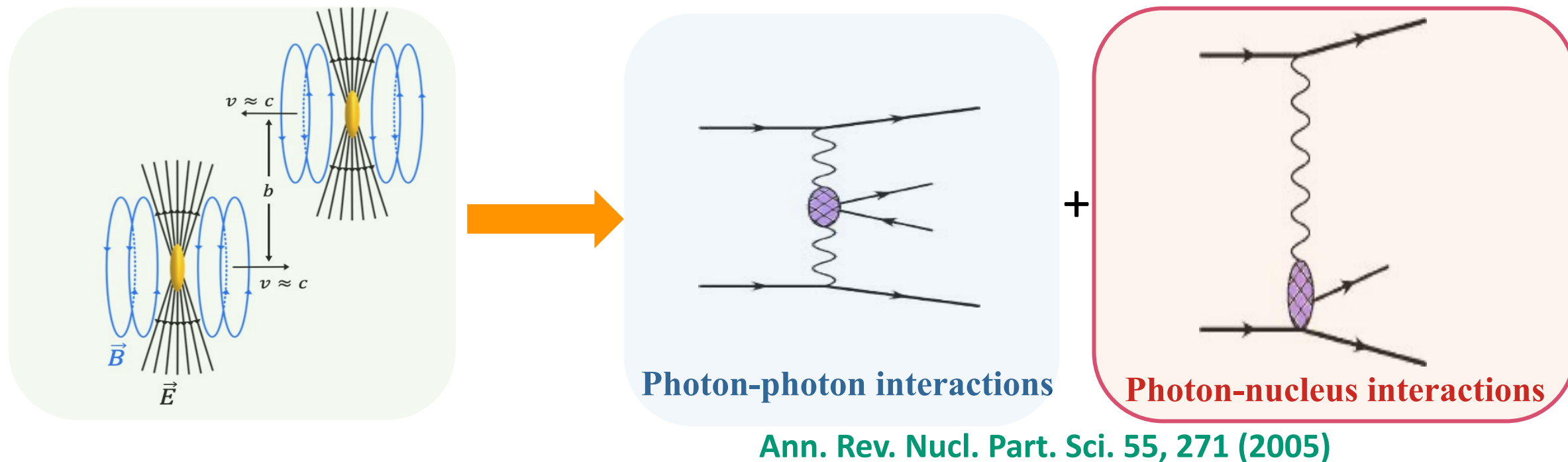
Outline

- **Nuclear deformation effects in UPC**
- **New nonlinear response of anisotropic flow to eccentricity (in hydrodynamics)**
- **Summary**

Nuclear deformation effects in UPC

Lin, Hu, Xu, SP, Wang, PRD (2024)

Ultra-Peripheral Collisions



- **Ultra-Peripheral Collisions (UPC):**

The impact parameter is larger than 2 times the radius of a nucleus

- Since the QCD effects are higher orders and QED effects are enhanced by the $Z\alpha$, UPC provides a nice platform to study the strong EB effects.

$Z\alpha \approx 1 \rightarrow$ High photon density

- Because the relativity, the photon (EB fields) are almost real. Photon-photon, photon-nuclear interactions

Measure mass distribution

- Charge distribution:

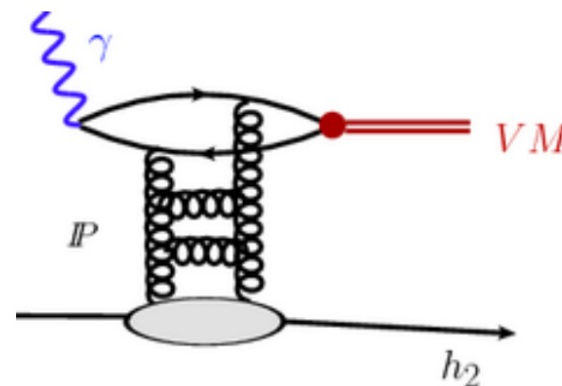
$$\gamma + A \rightarrow \gamma + A$$

to get the form factor $F(q^2)$ and the charge radius

$$\sqrt{\langle r^2 \rangle} = -6 \lim_{q^2 \rightarrow 0} \frac{\partial F(q^2)}{\partial q^2}$$

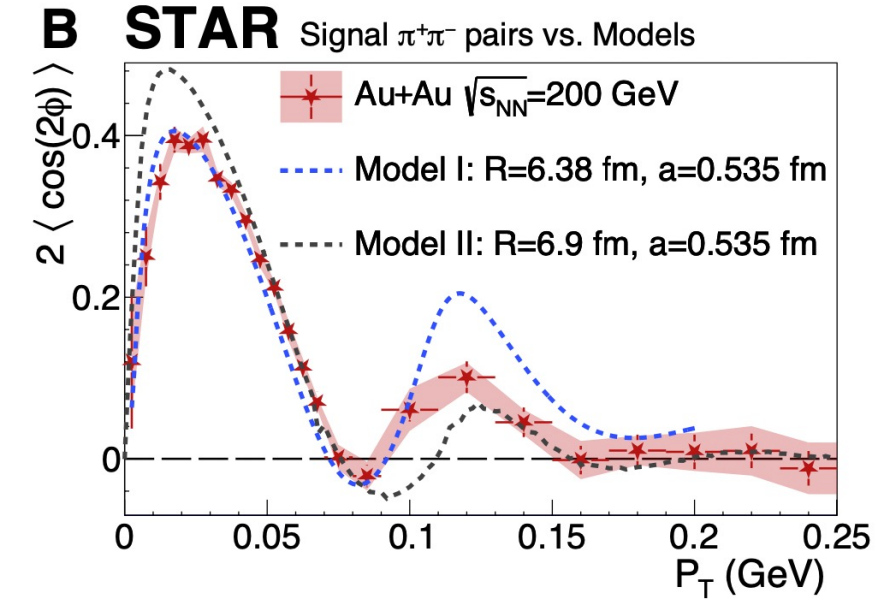
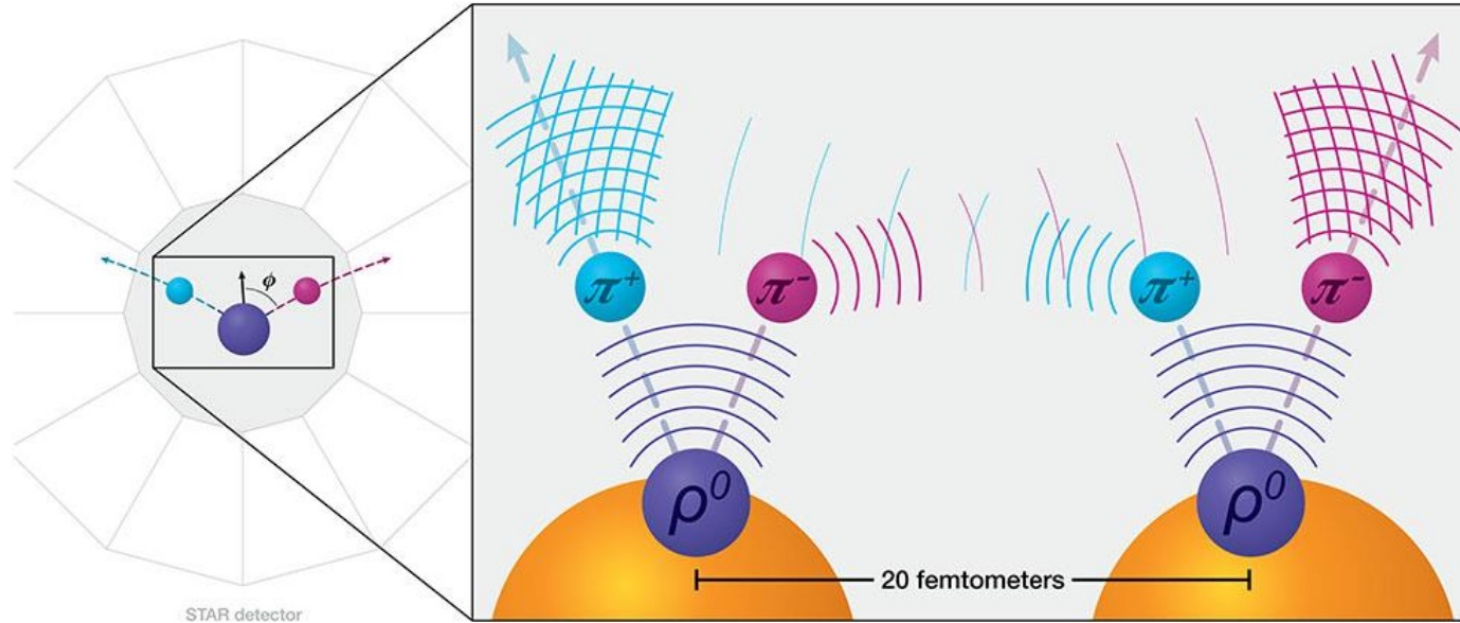
- Mass distribution:

$$\begin{aligned} \gamma + A &\rightarrow A + \rho^0, \\ \rho^0 &\rightarrow \pi^+ + \pi^- \end{aligned}$$

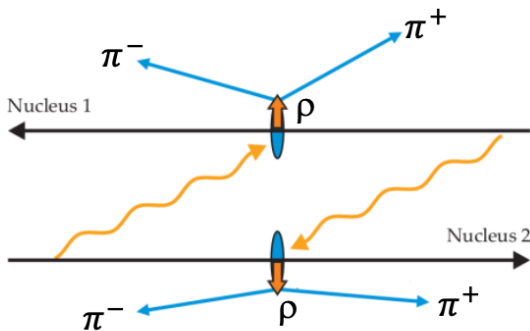


The cross section includes the information of mass distribution.

Double-slit experiment in Fermi scale

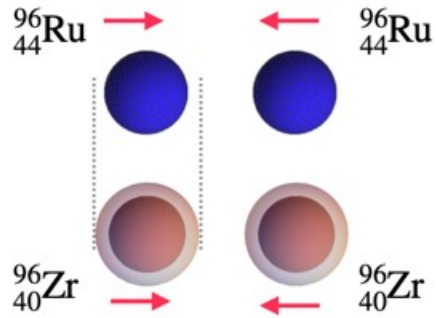


STAR: [Sci. Adv. 9, abq3903 \(2023\)](#)



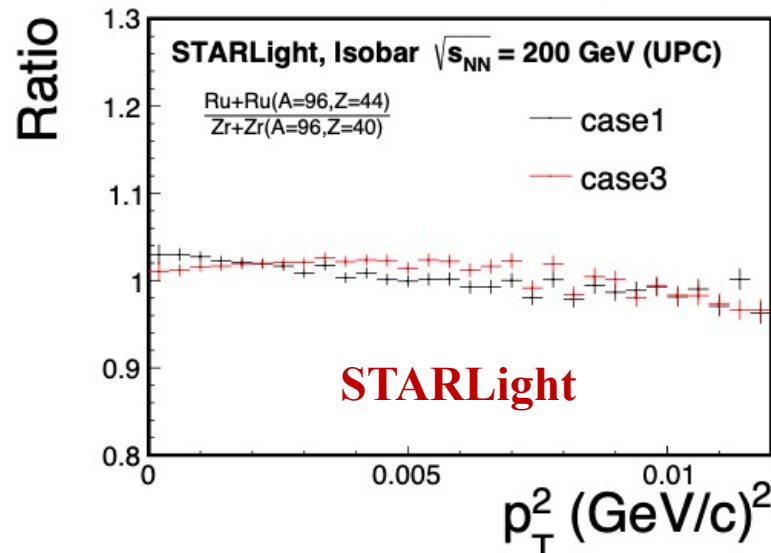
- Azimuthal asymmetries $\cos(2\phi)$ in diffractive vector meson production in UPC
 - Model I: [Zha, Brandenburg, Ruan, Tang, Xu, PRD 2021](#)
 - Model II: [Xing, Zhang, Zhou, Zhou, JHEP 2020](#)
- For $\cos(\phi)$ and $\cos(3\phi)$ related to ρ^0 : [Hagiwara, Zhang, Zhou, Zhou, PRD 2021](#)
- Also see studies for J/ψ : [Brandenburg, Xu, Zha, Zhang, Zhou, Zhou, PRD 2022](#)

Puzzle in isobaric collisions



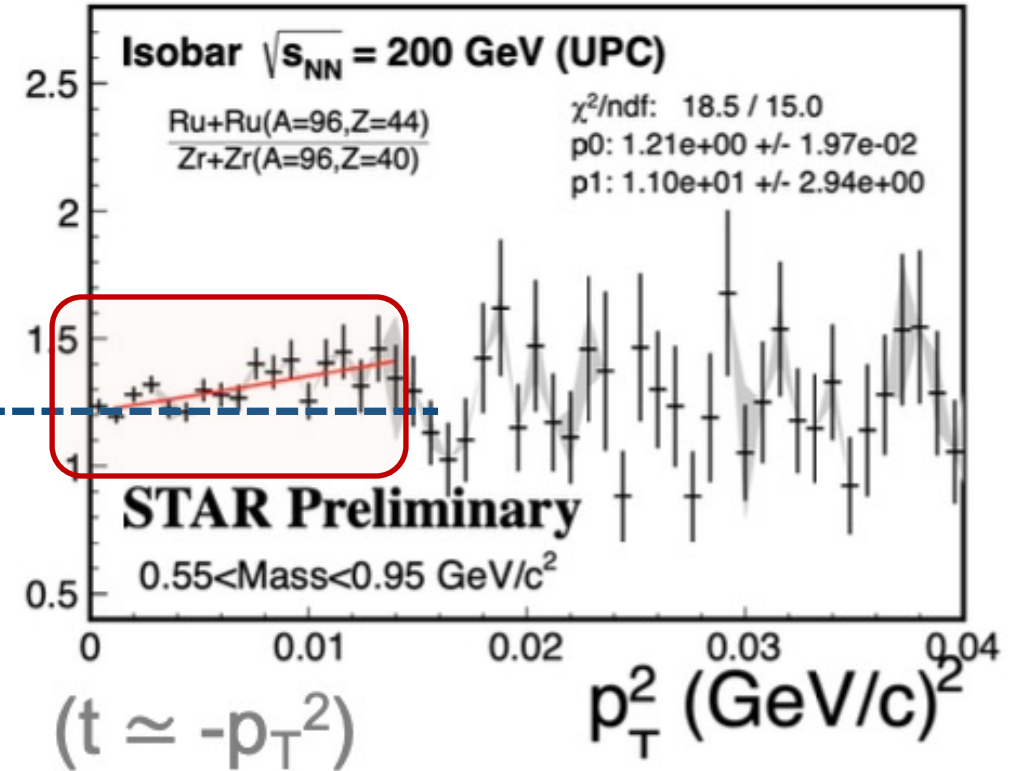
$$\gamma + A \rightarrow A + \rho^0,$$

$$\rho^0 \rightarrow \pi^+ + \pi^-$$



Ratio

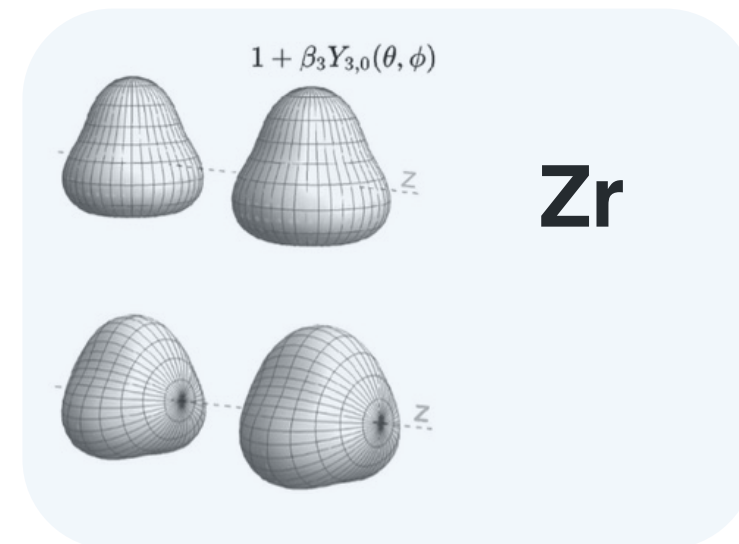
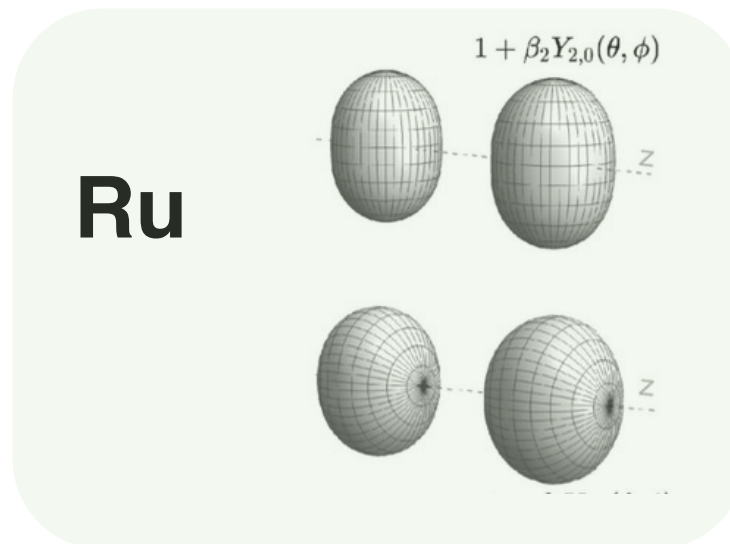
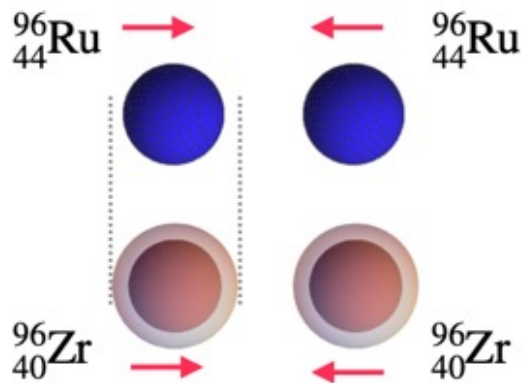
Theoretical
expectation
1.2



Theoretical expectation:

Without deformation effects, Ru and Zr are almost the same, therefore the ratio of the differential cross section for Ru+Ru and Zr+Zr should be a constant and equal to $(44/40)^2=1.2$

Nuclear deformation effects



$$\rho(r, \theta, \phi) = \frac{\rho_0}{1 + \exp\{[r - R_0(\theta, \phi)]/a\}}, \quad R_0(\theta) = R [1 + \beta_2 Y_{2,0}(\theta) + \beta_3 Y_{3,0}(\theta) + \dots],$$

(a) with deformation	R	a	β_2	β_3
Ru	5.093 fm	0.471 fm	0.16	0
Zr	5.021 fm	0.517 fm	0	0.20

PRC 105, 014905 (2022)

Theoretical model

- For simplicity, we take the **equivalent photon approximation**

$$\sigma = \int dy [n_1(y) \sigma_{\gamma A_2 \rightarrow \rho A_2}(y) + n_2(-y) \sigma_{\gamma A_1 \rightarrow \rho A_1}(-y)],$$

- We implement the **dipole model** to study this process

Ryskin, Z. Phys. C 57, 89 (1993); Brodsky, Frankfurt, Gunion, Mueller, Strikman, PRD 50, 3134 (1994); Kowalski, Teaney, PRD 68, 114005 (2003); Kowalski, Motyka, Watt, PRD 74, 074016 (2006)

$$\frac{d\sigma^{\gamma A \rightarrow \rho^0 A}}{dq_T^2} = \frac{1}{16\pi} |\mathcal{A}|^2$$

In the amplitude $\mathcal{A}$, we input the the nuclear mass distribution to the thickness function.

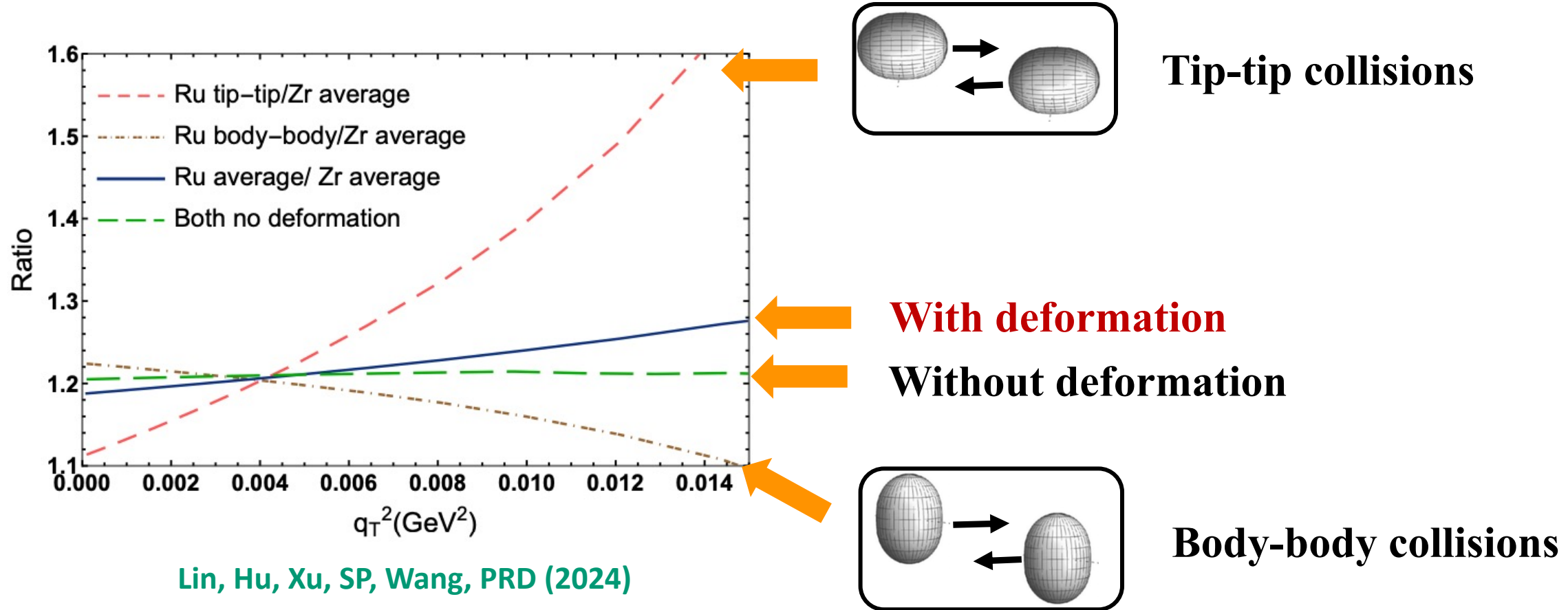
$$T_A(\mathbf{b}_T) = \int dz \rho(z, \mathbf{b}_T), \quad \rho(r, \theta, \phi) = \frac{\rho_0}{1 + \exp\{[r - R_0(\theta, \phi)]/a\}},$$

Also see more comprehensive studies beyond EPA in:

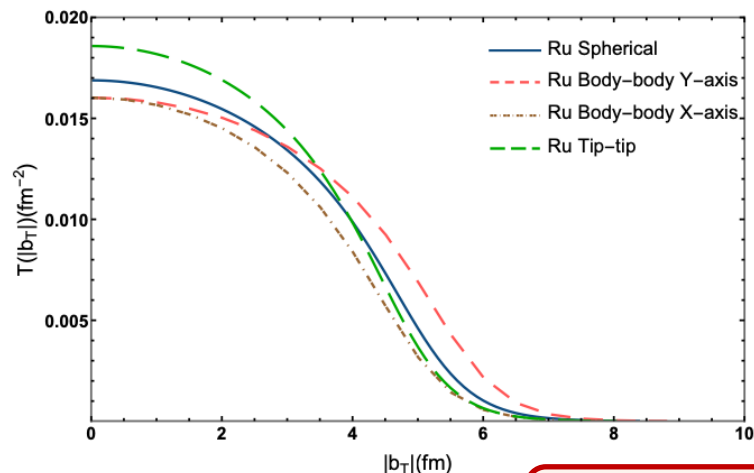
Mäntysaari, B. Schenke, C. Shen, and W. Zhao, PRL 131, 062301 (2023)

Results

Nuclear deformation effect is the key to understand the data!



How the nuclear deformation influence the ratio?



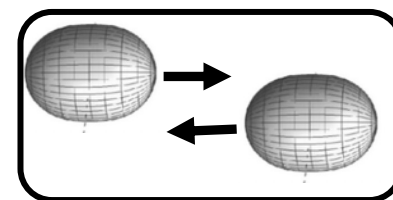
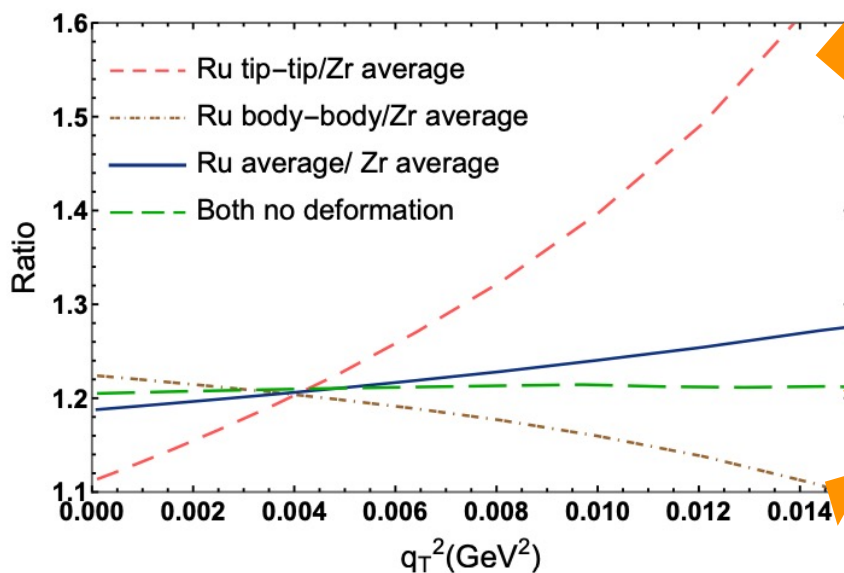
We parameterize the thickness function as

$$T_A(\mathbf{b}_T) \sim \exp\left(-\frac{\mathbf{b}_T^2}{w_T^2}\right)$$

where w_T is a parameter that can be interpreted as the **effective width of the nucleus**.

$$\frac{d\sigma_{Ru}/dq_T^2}{d\sigma_{Zr}/dq_T^2} \propto e^{\delta w_T q_T^2}, \quad \delta w_T \equiv -\frac{1}{2} [(w_T^{Ru})^2 - (w_T^{Zr})^2]$$

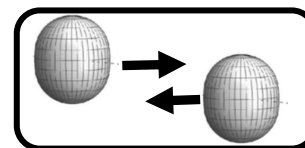
$$w_{T,tip}^{Ru} < w_{T,spherical}^{Ru} \approx w_{T,spherical}^{Zr} < w_{T,body}^{Ru}$$



Tip-tip collisions

With deformation

Without deformation



Body-body collisions

New nonlinear response of anisotropic flow to eccentricity (in preparation)

Xiang Ren, Jin-Yu Hu and Hao-Jie Xu

Eccentricity vs anisotropic flow

- Eccentricity

$$\epsilon_n \equiv -\frac{\int r^n e^{in\phi} \epsilon(r, \phi) r dr d\phi}{\int r^n \epsilon(r, \phi) r dr d\phi}$$

- Anisotropic flow

$$(2\pi)^3 \frac{dN}{dY p_T dp_T d\phi_p} = A_0 \left[1 + \sum_{n=1}^{\infty} 2v_n(p_T) \cos n\phi \right]$$

$$v_2 \propto \epsilon_2$$

$$v_4 = \chi v_2^2 + v_{4L} \propto \epsilon_2^2 + \dots$$

How to understand this non-linear response from the hydrodynamics analytically?

Conservation equations in hydrodynamics

- Energy-momentum conservation equation

$$\partial_\mu T^{\mu\nu} = 0$$

- Constitution equation

$$T^{\mu\nu} = (\epsilon + P)u^\mu u^\nu - P g^{\mu\nu} + \pi^{\mu\nu}$$
$$\pi^{\mu\nu} = -\Pi g^{\mu\nu} + 2\eta_s \sigma^{\mu\nu}$$

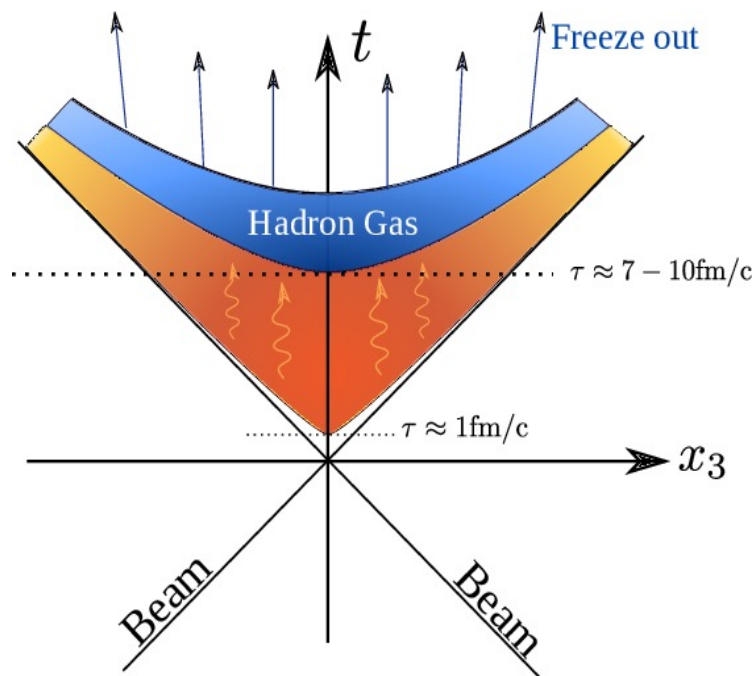
- Equation of state

$$P = P(\epsilon)$$

Pressure as a function of energy density

Bjorken flow

- **Bjorken flow:** a well-known 0+1 dimensional analytic solution for the relativistic hydrodynamics with **longitudinal boost invariance**



- **Physical picture:**

- All the particles are generated at the same time in the rest frame of the fluid cell.
- The system has longitudinal boost invariance, all thermodynamic variables depend on proper time only.

- **Solutions**

- in **(t, x, y, z)** coordinates:

Scaling law for Energy density $\frac{e(t)}{e(t=0)} = \frac{1}{\tau^{4/3}}$
 τ : proper time

Fluid velocity $u^\mu = \left(\frac{t}{\tau}, 0, 0, \frac{z}{\tau} \right)$

- in **(τ, x, y, η)** coordinates (η space rapidity):

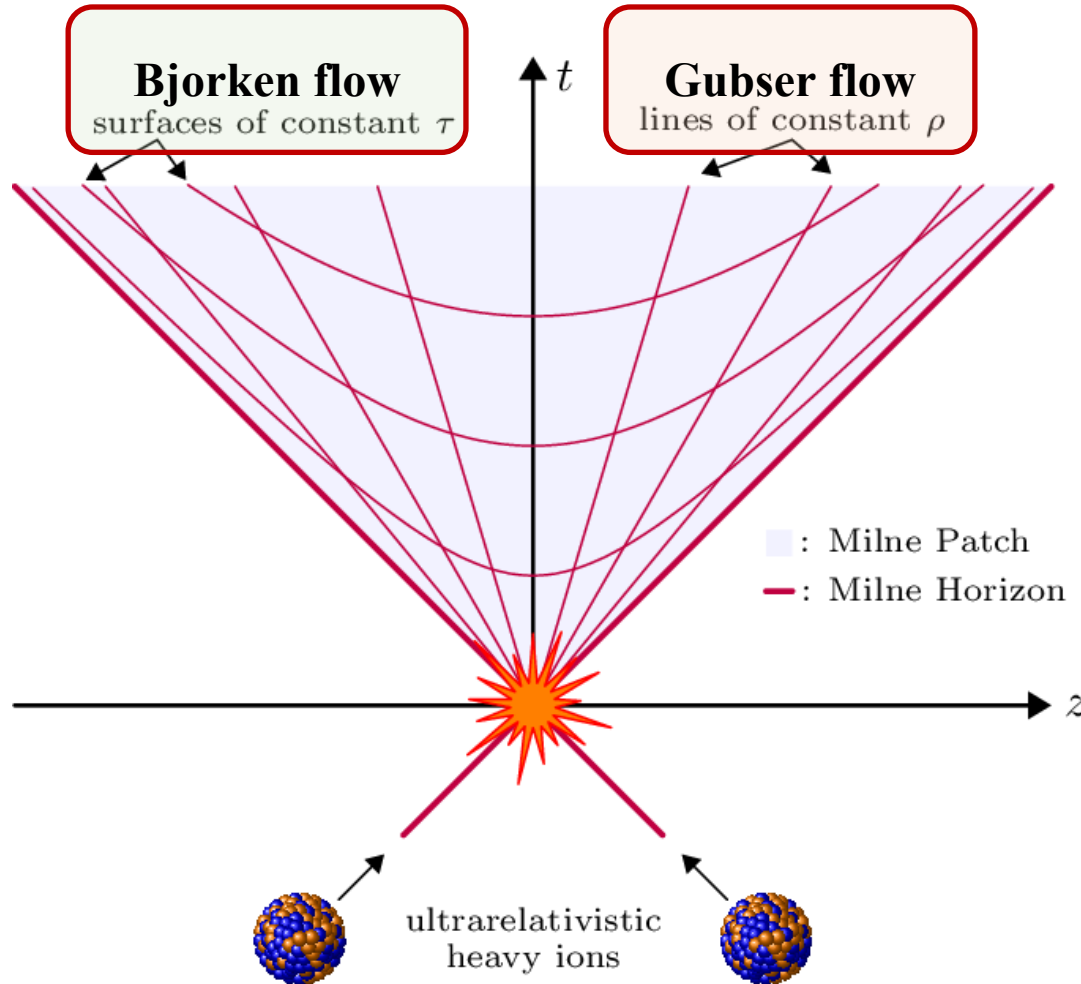
$$u^\mu = (1, 0, 0, 0)$$

Bjorken flow: the fluid cell is at rest in (τ, x, y, η) coordinates

No radial flow in a Bjorken flow!

Gubser flow (I)

- **Gubser flow:** a 1+1 dimensional analytic solution for relativistic hydrodynamics with **transverse flows**



- **Motivation:**
Extended Bjorken flow with radial flow
- **Basic idea:**
Search for the analytic solution as $u^\mu = (1, 0, 0, 0)$ in special coordinates?
- **Answer:**
Taking conformal (Weyl) transformation

Gubser, PRD (2010); Gubser, Yarom, NPB (2011)

Gubser flow (II)

- **Gubser flow:** a 1+1 dimensional analytic solution for relativistic hydrodynamics with **transverse flows**

$$(\tau, x_{\perp}, \phi, \eta)$$

τ : proper time

$x_{\perp}$: transverse length

ϕ : azimuthal angle in transverse plane

η : space rapidity

Scaling law for Energy density

$$\varepsilon = \frac{CL^{8/3}}{\tau^{4/3}} \frac{1}{[L^4 + (x_{\perp}^2 - \tau^2)^2 + 2L^2(x_{\perp}^2 + \tau^2)]^{4/3}}$$

C: integration constant
L: size of the system

Fluid velocity

$$u_{\mu} = \left(-\frac{1}{\cosh \rho} \frac{L^2 + x_{\perp}^2 + \tau^2}{2L\tau}, \boxed{-\frac{1}{\cosh \rho} \frac{x_{\perp}}{L}}, 0, 0 \right)$$
$$\cosh \rho = \frac{1}{2L\tau} [L^4 + (x_{\perp}^2 - \tau^2)^2 + 2L^2(x_{\perp}^2 + \tau^2)]^{1/2}$$

The radial flow survive! But, no ϕ dependence!

Perturbation to Gubser flows

- Perturbation in temperature:

$$T_0 \rightarrow T = T_0 [1 + \delta(\rho) S(\Theta, \phi)],$$

Accordingly, from the conservation equations, one can find that the fluid velocity becomes,

$$\begin{aligned} u^\mu &\rightarrow u^\mu + \delta u^\mu \\ \delta u^\mu &= \left(0, \frac{2L\tau}{L^2 + x_\perp^2} \nu_s(\rho) \partial_\Theta S(\Theta, \phi), \tau \nu_s(\rho) \partial_\phi S(\Theta, \phi), 0 \right) \\ &\quad + \mathcal{O}(\tau^2) \end{aligned}$$

ϕ dependence is introduced through the perturbations

where

$$\begin{aligned} \delta &= \delta_0 + \delta_0 \frac{\eta_0}{2\lambda C} (2e^\rho)^{-2/3} + \mathcal{O}(e^{2\rho}), \\ \nu_s &= -\frac{3}{2}\delta_0 + k_0 \frac{\eta_0}{\lambda C} (2e^\rho)^{-2/3} + \mathcal{O}(e^{2\rho}), \end{aligned}$$

Hatta, Noronha, Torrieri, Xiao PRD (2014); Hatta, Monnai, Xiao, PRD (2015)

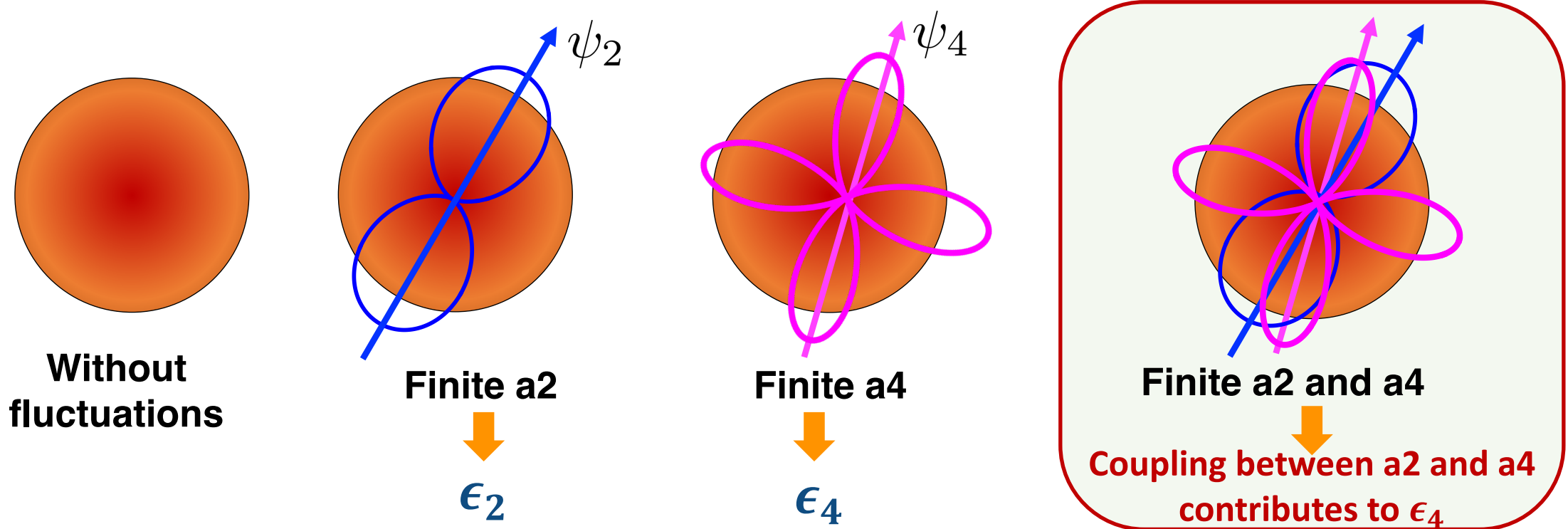
Perturbation and eccentricity (I)

- **Fluctuations:** In this work, we introduce the fluctuations as

$$S(\Theta, \phi) = a_2 \left(\frac{2x_{\perp} L}{L^2 + x_{\perp}^2} \right)^2 \cos 2\phi + a_4 \left(\frac{2x_{\perp} L}{L^2 + x_{\perp}^2} \right)^4 \cos 4\phi,$$

For simplicity, here we set $\psi_2 = \psi_4 = 0$

- The parameters a_2 and a_4 are related to eccentricity,

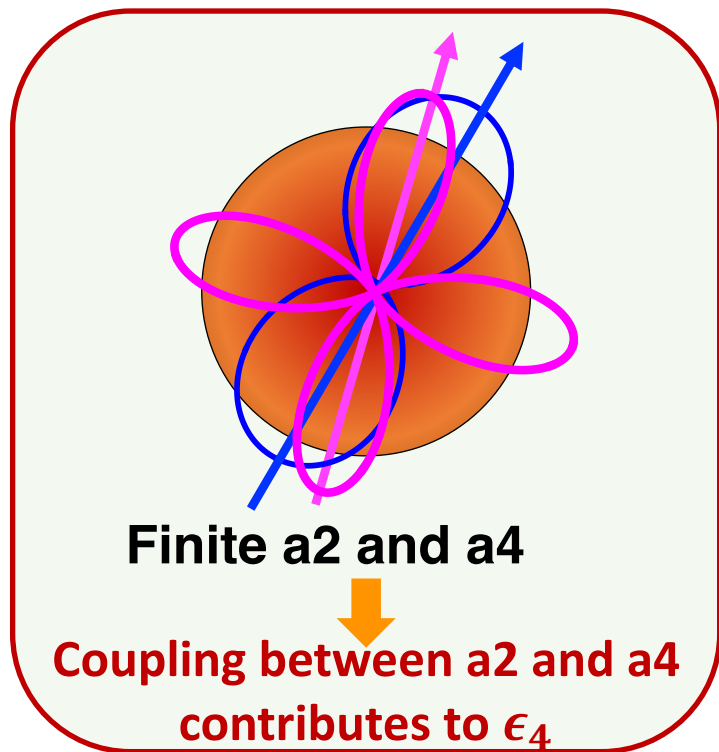


Perturbation and eccentricity (II)

- **Fluctuations:** In this work, we introduce the fluctuations as

$$S(\Theta, \phi) = a_2 \left(\frac{2x_{\perp} L}{L^2 + x_{\perp}^2} \right)^2 \cos 2\phi + a_4 \left(\frac{2x_{\perp} L}{L^2 + x_{\perp}^2} \right)^4 \cos 4\phi, \quad \text{For simplicity, here we set } \psi_2 = \psi_4 = 0$$

- The parameters a_2 and a_4 are related to eccentricity,



$$\epsilon_n \propto - \frac{\int d^2 x_{\perp} \epsilon^{3/4} \frac{x_{\perp}^n}{(L^2 + x_{\perp}^2)^{n-1}} \cos n\phi}{\int d^2 x_{\perp} \epsilon^{3/4} \frac{x_{\perp}^n}{(L^2 + x_{\perp}^2)^{n-1}}}$$

$$\epsilon_2 = -a_2 - \frac{24}{35} a_2 a_4 + \dots$$

$$\epsilon_4 = -\frac{36}{35} a_4 - \frac{18}{35} a_2^2 + \dots$$

$$a_2 = -\epsilon_2 + \mathcal{O}(\epsilon_2^2, \epsilon_4^2, \epsilon_2 \epsilon_4),$$

$$a_4 = -\frac{35}{36} \epsilon_4 - \frac{1}{2} \epsilon_2^2 + \mathcal{O}(\epsilon_2^2, \epsilon_4^2, \epsilon_2 \epsilon_4).$$

Cooper-Frye formula and v_2, v_4

- **Cooper-Frye formula:**

For a given freezeout proper time, we insert the temperature and fluid velocity into the Cooper-Frye formula

$$(2\pi)^3 \frac{dN}{dY p_T dp_T d\phi_p} = - \int_{\Sigma} p^{\mu} d\Sigma_{\mu} f \left(\frac{p^{\mu} u_{\mu}}{T} \right) = A_0 \left[1 + \sum_{n=1}^{\infty} 2v_n(p_T) \cos n\phi \right]$$

- **Anisotropic flow:**

$$\begin{aligned} v_2 &= A_2 \epsilon_2, \\ v_4 &= A_4 \epsilon_4 + C_4 \epsilon_2^2. \end{aligned}$$

A2, A4, C4 are functions of the p_T , freezeout temperature T and $x_{\perp}$.

Small and large p_T limits

- In small p_T limit,

$$v_2 = \frac{2^{10} p_T^2 (K_1 m_T + 4K_0 T)}{21 B^6 K_1 m_T T^2} \epsilon_2,$$

$$v_2 / \epsilon_2 \propto p_T^2$$

$$v_4 = \frac{2^{16} p_T^4 (7K_1 m_T + 56K_0 T)}{663 B^{12} K_1 m_T T^4} \epsilon_4 + \frac{2^{18} p_T^4 (67K_1 m_T + 549K_0 T)}{7735 B^{12} K_1 m_T T^4} \epsilon_2^2.$$

$$K_n \left(\frac{m_T}{T} \right) = \int_0^\infty dy \exp \left(-\frac{m_T}{T} \cosh(y - Y) \right) \cdot \cosh(ny - nY)$$

- In large p_T limit,

$$v_2 \approx \frac{p_T}{2T} \delta u_{\perp 2}^* = \epsilon_2 \frac{p_T}{T} \frac{500\sqrt{5}}{B^3 81}.$$

$$v_2 / \epsilon_2 \propto p_T$$

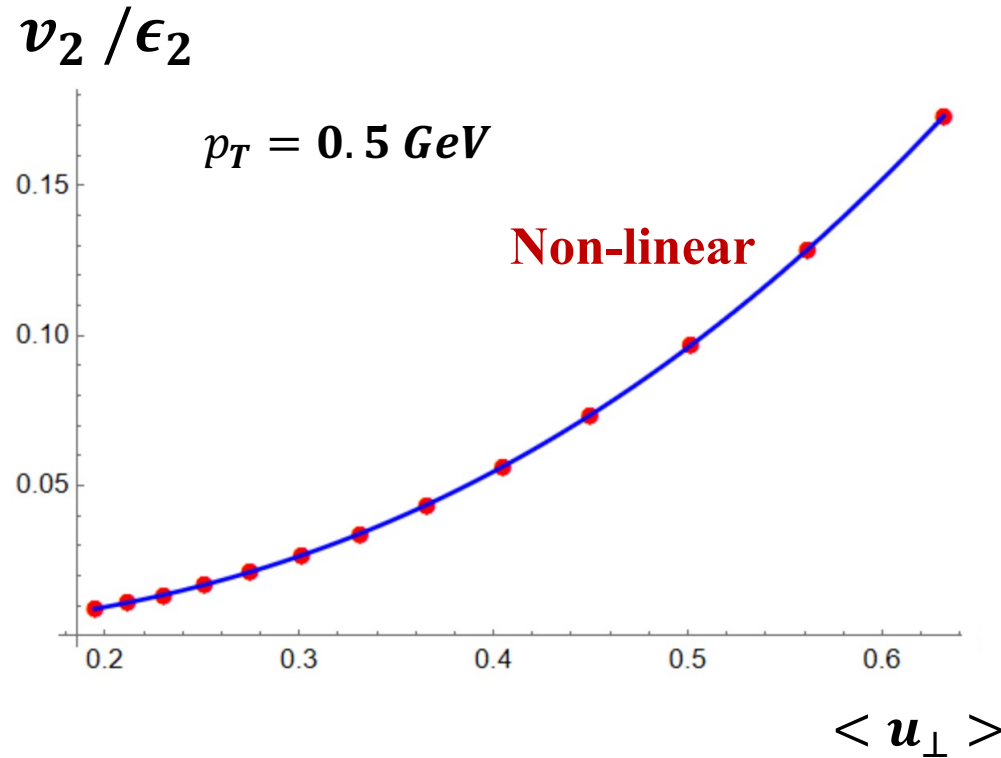
$$v_4 \approx \frac{1}{2} \left(\frac{p_T}{2T} \right)^2 \delta u_{\perp 2}^{*2} = \frac{1}{2} v_2^2.$$

$$v_4 / v_2^2 = 1/2$$

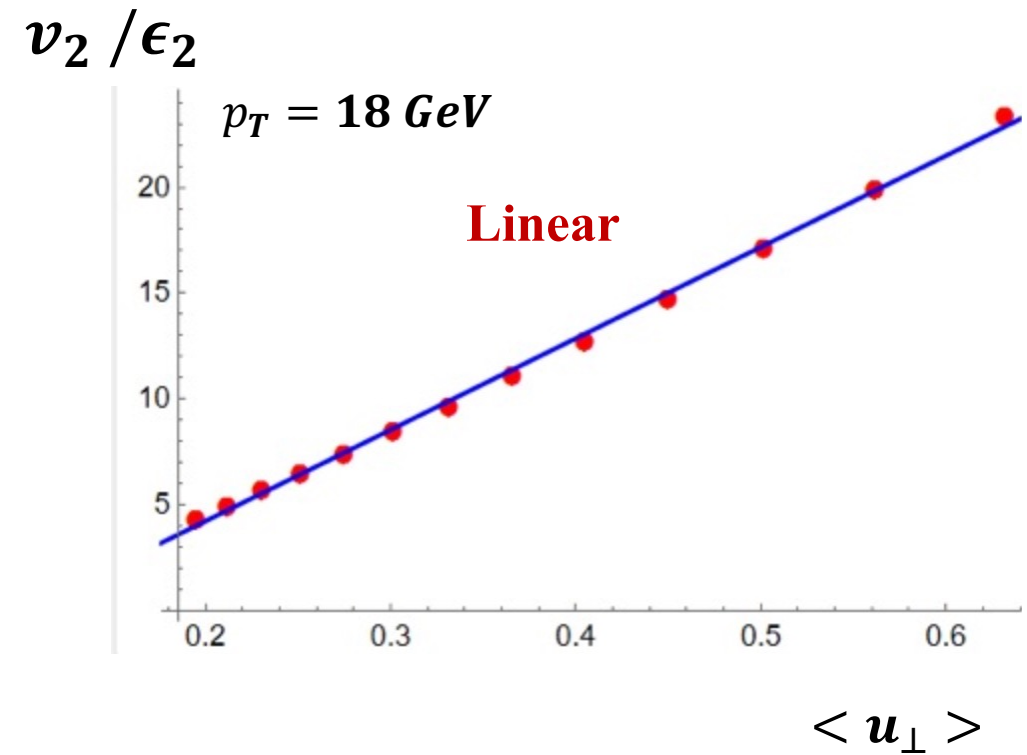
Agree with Teaney, Li, PRC (2011)

v_2 / ϵ_2 as function of averaged transverse velocity $\langle u_\perp \rangle$

Small p_T



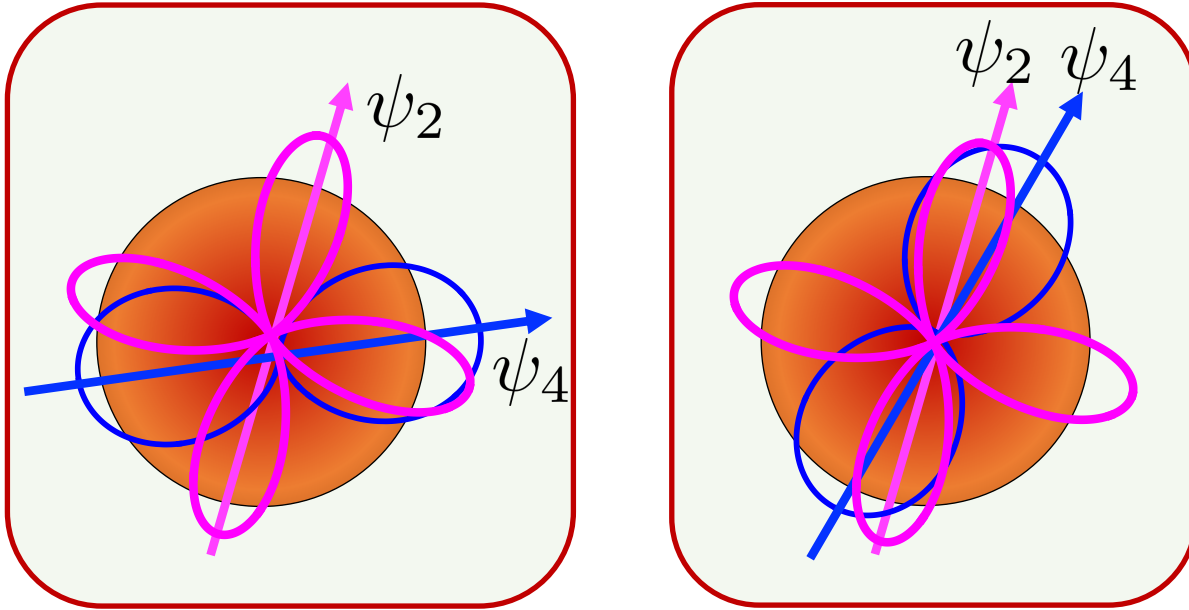
Large p_T



Averaged transverse velocity: $\langle u_\perp \rangle = \frac{\int x_\perp dx_\perp d\theta u_\perp}{\int x_\perp dx_\perp d\theta}$

Effects of different participant and event planes

- Different participant planes: finite and different ψ_2 and ψ_4



$$\begin{aligned}\epsilon_2 &= -a_2 - \frac{24}{35}a_2a_4 + \dots \\ \epsilon_4 &= -\frac{36}{35}a_4 - \frac{18}{35}\cos(\psi_2 - \psi_4)a_2^2 + \dots\end{aligned}$$

- Introduce the event plane Ψ_2 and Ψ_4 in Cooper-Frye formula:

$$(2\pi)^3 \frac{dN}{dY p_T dp_T d\phi_p} = - \int_{\Sigma} p^\mu d\Sigma_\mu f \left(\frac{p^\mu u_\mu}{T} \right) = A_0 \left[1 + \sum_{n=1}^{\infty} 2v_n(p_T) \cos[n(\phi - \Psi_n)] \right]$$

Nonlinear response for anisotropic flows

- Ψ_2 and Ψ_4 : measured from experiments
- ψ_2 and ψ_4 : correspond to ϵ_2 and ϵ_4

$$\psi_2 \approx \Psi_2 \quad \psi_4 \not\approx \Psi_4$$

Commonly used in experiments:

$$v_4\{\Psi_4\}^2 = v_{4L}^2 + v_4\{\Psi_2\}^2 \approx v_{4L}^2 + (\chi v_2^2\{\psi_2\})^2$$

Our results in Gubser flow:

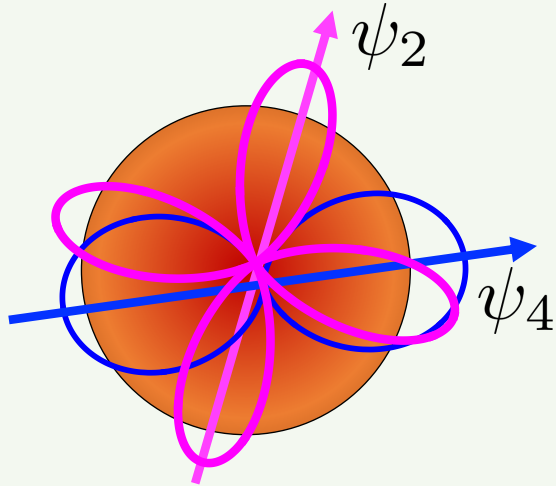
$$v_4\{\Psi_4\} = v_{4L} + \chi \cos[4(\psi_2 - \psi_4)] v_2^2\{\psi_2\}$$

Nonlinear response coefficient is related to the initial participant planes.

Nonlinear response for anisotropic flows

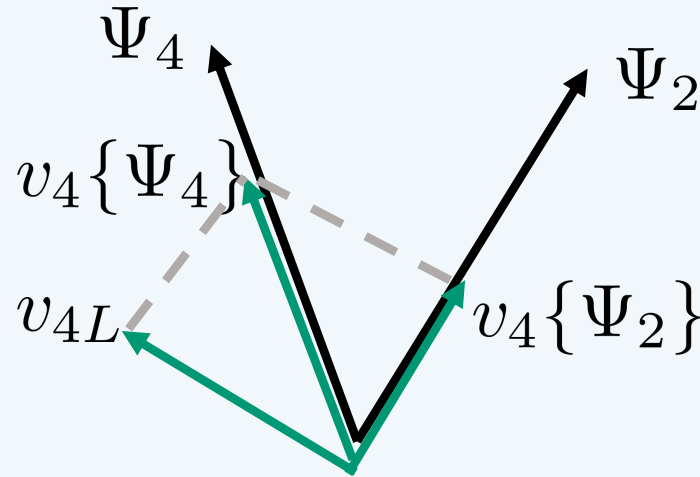
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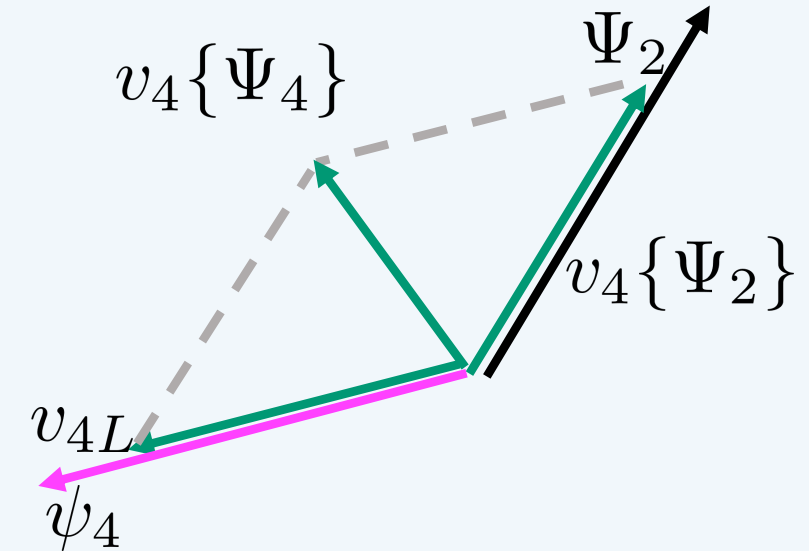
perturbations given by initial conditions, e.g. deformations, fluctuations

Commonly used in experiments:



$$\begin{aligned} v_4\{\Psi_4\}^2 &= v_{4L}^2 + v_4\{\Psi_2\}^2 \\ &\approx v_{4L}^2 + \chi v_2^2\{\psi_2\} \end{aligned}$$

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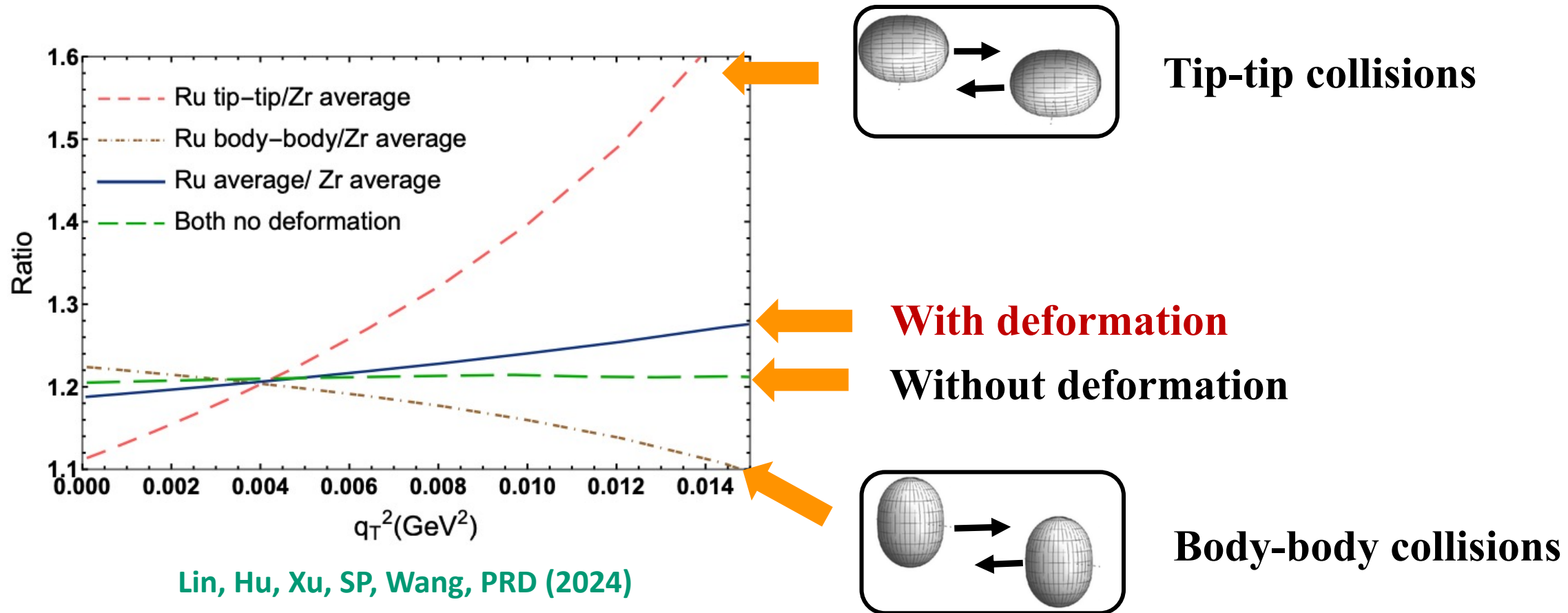


$$\begin{aligned} v_4\{\Psi_4\} &= v_{4L} \\ &+ \chi \cos[4(\psi_2 - \psi_4)] v_2^2\{\psi_2\} \end{aligned}$$

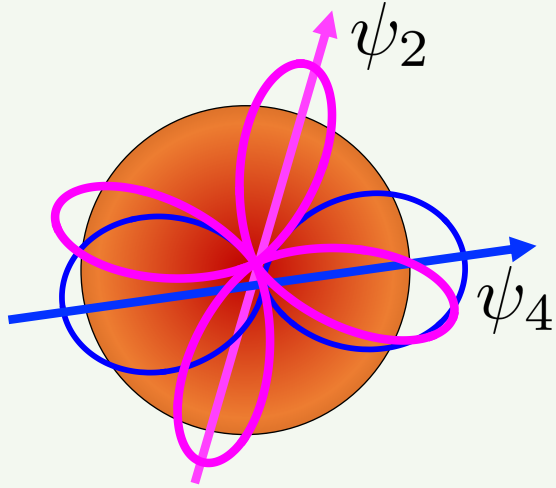
Summary

Summary (I)

- Nuclear deformation effect is the key to understand photon-nuclear interaction in isobaric collisions.

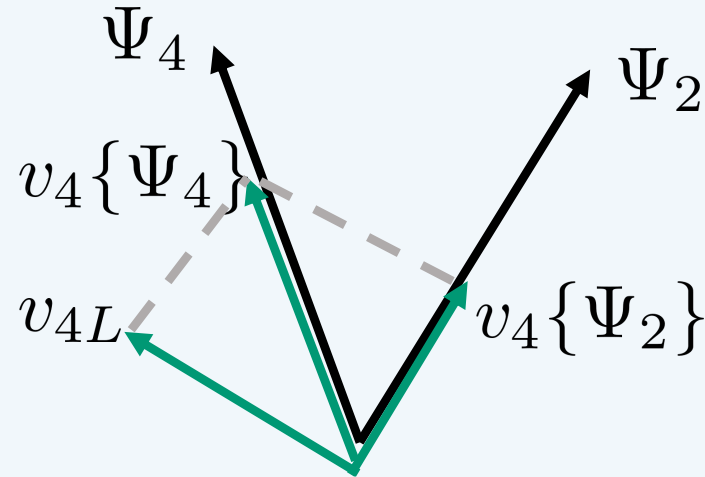


Summary (II)



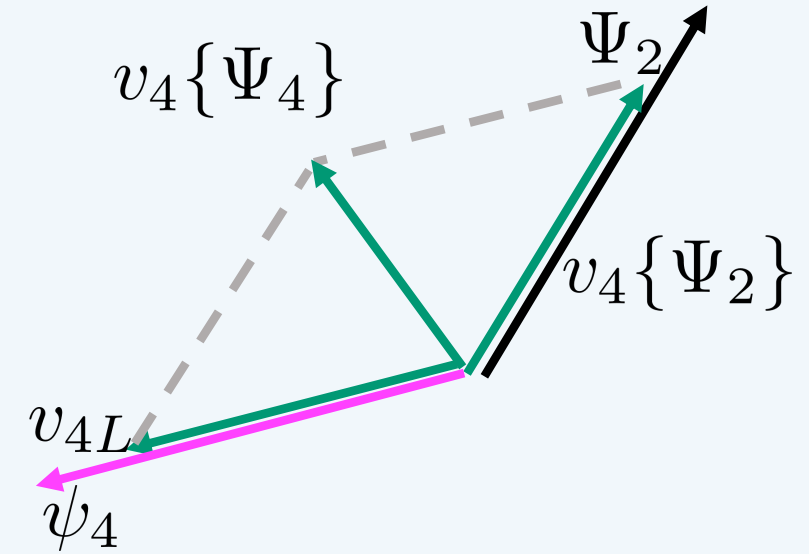
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Commonly used in experiments:



$$\begin{aligned} v_4\{\Psi_4\}^2 &= v_{4L}^2 + v_4\{\Psi_2\}^2 \\ &\approx v_{4L}^2 + \chi v_2^2\{\psi_2\} \end{aligned}$$

Our results in Gubser flow:



$$\begin{aligned} v_4\{\Psi_4\} &= v_{4L} \\ &+ \chi \cos 4(\psi_2 - \psi_4) v_2^2\{\psi_2\} \end{aligned}$$

Nonlinear response coefficient is related to the initial participant planes.

Thank you for listening!

Any comments and suggestions are welcome!