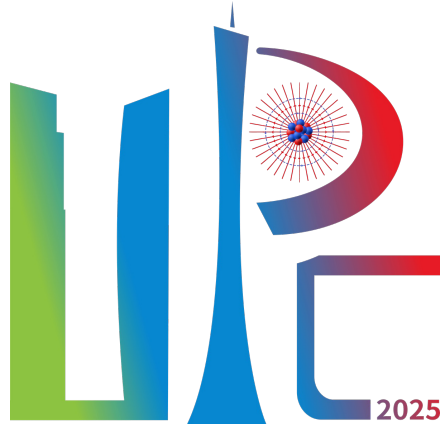




跨能量尺度原子核结构实验研究新进展及展望

张春健

复旦大学

2025年11月24日，华南师范大学，广州

Outline

I. Nuclear structure across energy scales

II. Nuclear deformation in ground-state ^{238}U nucleus

STAR, Nature 635, 67-72 (2024)
STAR, Rep. Prog. Phys. 88, 108601 (2025)
arXiv: 2504.15245
arXiv: 2509.09376

III. Nucleonic cluster pattern in light ^{16}O nucleus

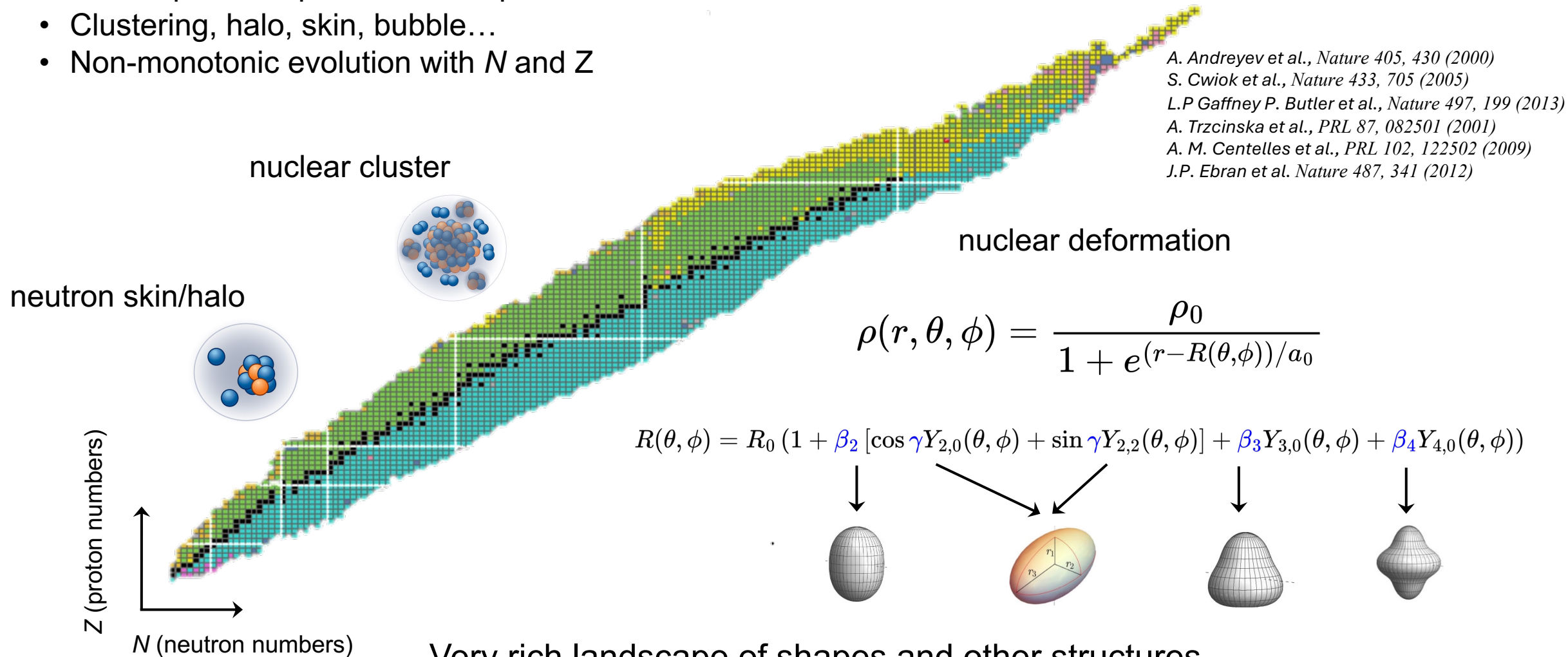
STAR, arXiv: 2510.19645
PLB 862, 139322 (2025)
PLB 870, 139926 (2025)

IV. Conclusions and outlooks

Shape of atomic nuclei

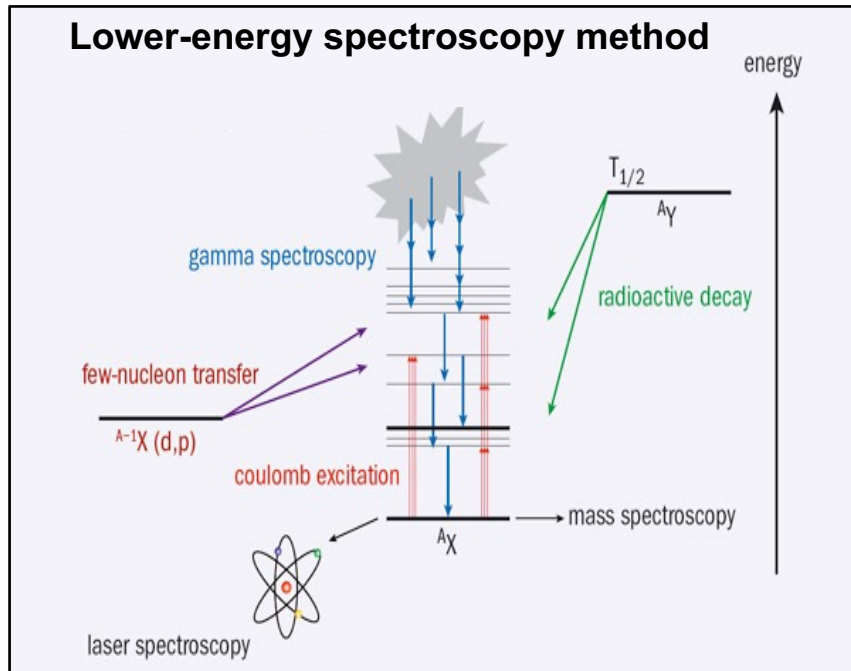
Emergent phenomena of the many-body quantum system, govern by short-range strong nuclear force

- Quadrupole/octupole/hexadecapole deformations
- Clustering, halo, skin, bubble...
- Non-monotonic evolution with N and Z



A. Andreyev et al., *Nature* 405, 430 (2000)
 S. Cwiok et al., *Nature* 433, 705 (2005)
 L.P Gaffney P. Butler et al., *Nature* 497, 199 (2013)
 A. Trzcinska et al., *PRL* 87, 082501 (2001)
 A. M. Centelles et al., *PRL* 102, 122502 (2009)
 J.P. Ebran et al. *Nature* 487, 341 (2012)

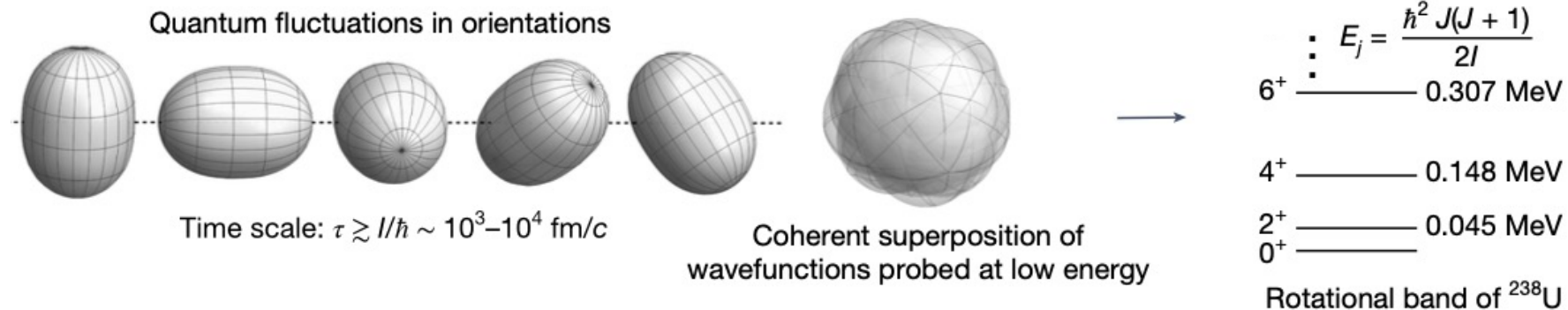
Nuclear shape at low energy: long exposure



Traditional imaging method taken before destruction

- Low energy spectroscopic methods probe a superposition of these fluctuations.
- Instantaneous shapes not directly seen, but inferred from model comparison.

Each DOF has zero-point fluctuations within certain timescales.



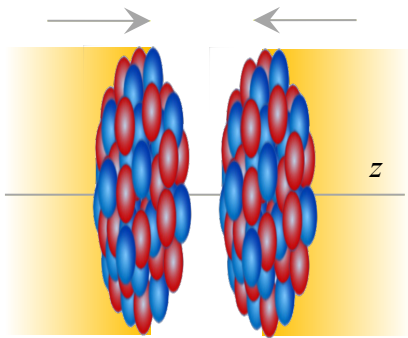
Imaging by smashing method

Take a snapshot

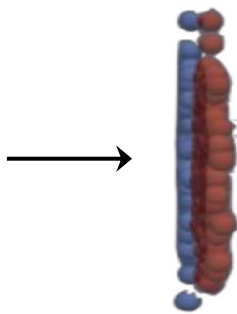
Evolution

Measurement

Nuclei collisions



$$T_{\mu\nu}(\tau = 0)$$

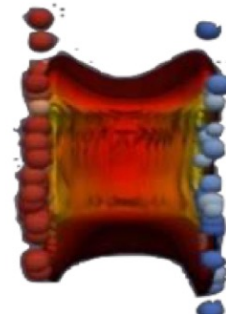


$$\tau \sim 2R_0/\Gamma \sim 0.1\text{fm}/c$$

exposure

Pressure-driven expansion of
Quark-gluon plasma (QGP)

$$\partial_\mu T^{\mu\nu} = 0, \text{ EOS, viscosity...}$$

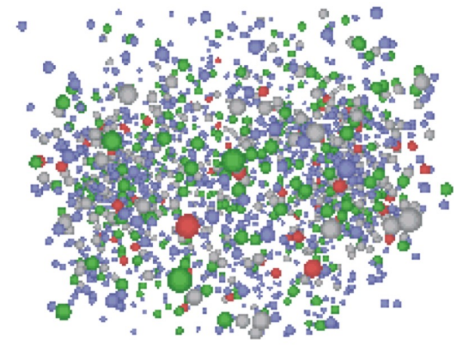


$$\tau \sim 10\text{fm}/c$$

expansion

free-streaming

$$T_{\mu\nu}(\tau = \infty)$$



$$\tau \sim 10^{15}\text{fm}/c$$

detection

$$\rho(r, \theta, \phi) = \frac{\rho_0}{1 + e^{(r-R(\theta, \phi))/a_0}}$$

$\beta_2 \rightarrow$ quadrupole deformation

$\beta_3 \rightarrow$ octupole deformation

$\gamma \rightarrow$ triaxiality

$a_0 \rightarrow$ surface diffuseness

$R_0 \rightarrow$ nuclear size

ab initio theory/shell model/DFT

J. Jia et al., Nucl. Sci. Tech. 35, 220 (2024)

Imaging by smashing method

Take a snapshot

Evolution

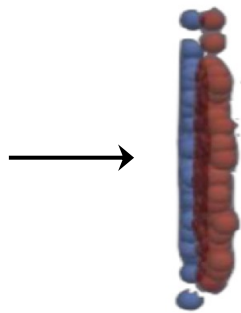
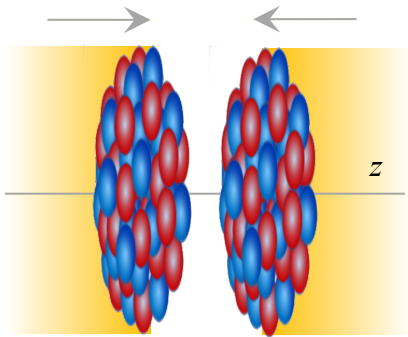
Measurement

Nuclei collisions

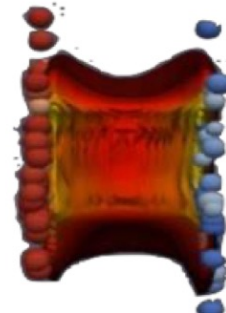
$$T_{\mu\nu}(\tau = 0)$$

$$\partial_\mu T^{\mu\nu} = 0, \text{ EOS, viscosity...}$$

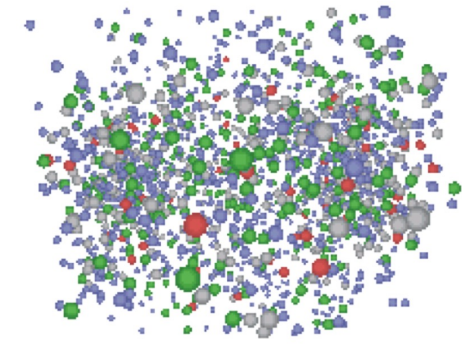
$$T_{\mu\nu}(\tau = \infty)$$



Pressure-driven expansion of
Quark-gluon plasma (QGP)



free-streaming



$\tau \sim 2R_0/\Gamma \sim 0.1\text{fm}/c$
exposure

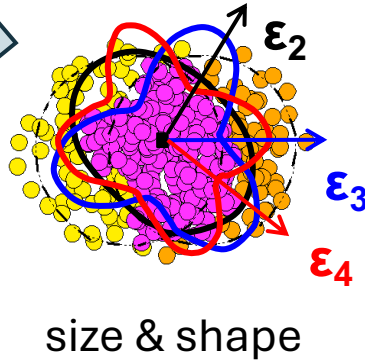
$\tau \sim 10\text{fm}/c$
expansion

$\tau \sim 10^{15}\text{fm}/c$
detection

$$\rho(r, \theta, \phi) = \frac{\rho_0}{1 + e^{(r-R(\theta, \phi))/a_0}}$$

- $\beta_2 \rightarrow$ quadrupole deformation
- $\beta_3 \rightarrow$ octupole deformation
- $\gamma \rightarrow$ triaxiality
- $a_0 \rightarrow$ surface diffuseness
- $R_0 \rightarrow$ nuclear size

ab initio theory/shell model/DFT



size & shape

J. Jia et al., Nucl. Sci. Tech. 35, 220 (2024)

$$R_\perp^2 \propto \langle r_\perp^2 \rangle \quad \mathcal{E}_n \propto \langle r_\perp^n e^{in\phi} \rangle$$

Diagram showing relationships between R_0 , a_0 , and β_n with arrows indicating dependencies.

$$\text{Observables } \frac{d^2 N}{d\phi dp_T} = N(p_T) \left(\sum_n V_n e^{-in\phi} \right)$$

Event-by-event linear responses:

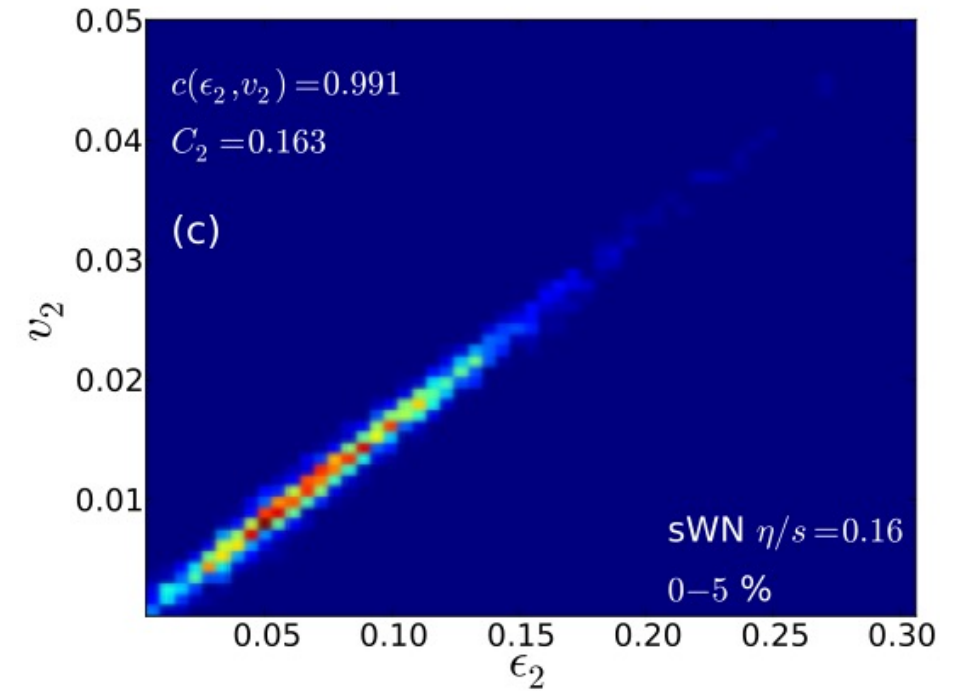
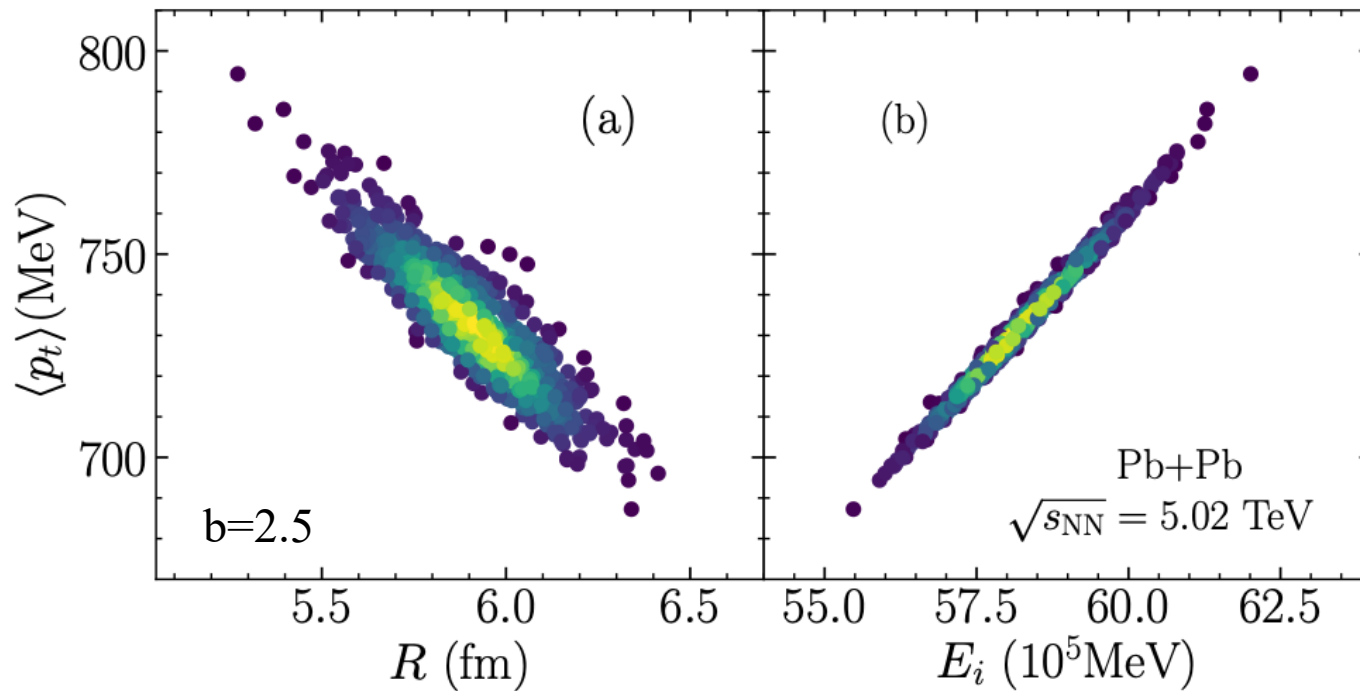
$$\frac{\delta[p_T]}{[p_T]} \propto -\frac{\delta R_\perp}{R_\perp} \quad V_n \propto \mathcal{E}_n$$

Key: 1) fast snapshot, 2) linear response, 3) large multiplicity for many-body correlation

Linear response in ultra-central collisions

$$\frac{\delta[p_T]}{[p_T]} \propto -\frac{\delta R_{\perp}}{R_{\perp}}$$

$$V_n \propto \mathcal{E}_n$$

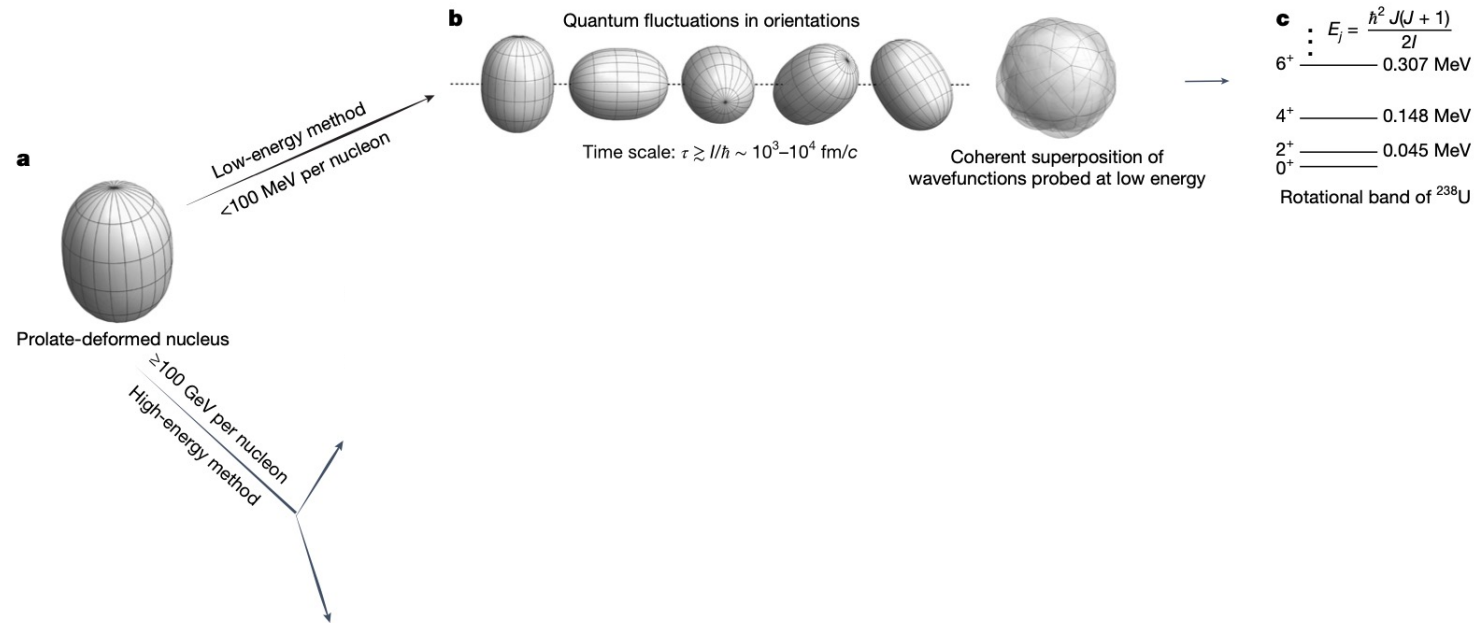


G. Giacalone, F. Gardim, J. Hostler, J. Ollitrault, et al, PRC 103, 024909 (2021)

H. Niemi, G. Denicol, H. Holopainen, P. Huovinen, PRC 87, 054901 (2013)

Imaging nuclear shape in high-energy snapshot

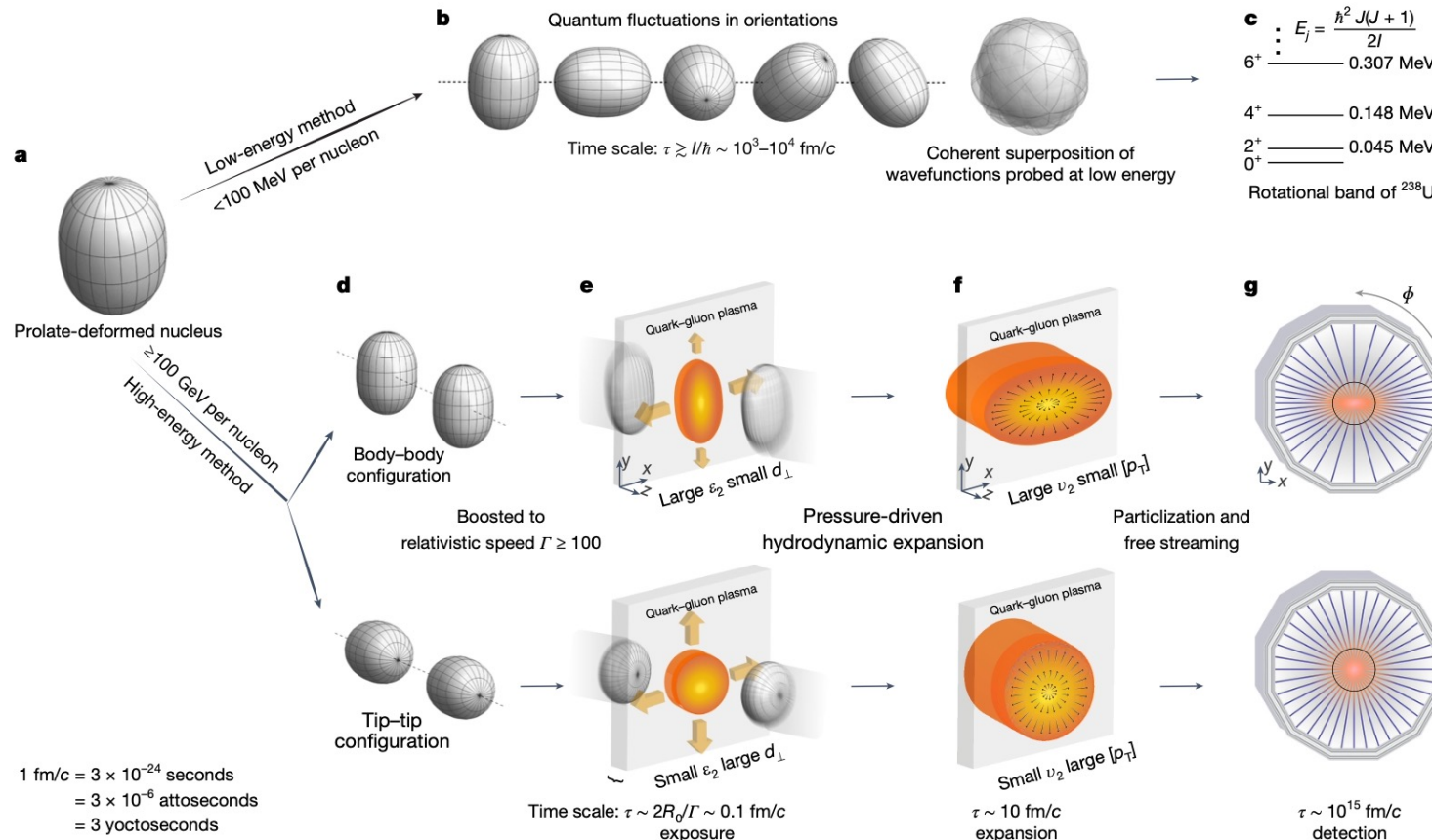
- Nuclear shape in intrinsic (body-fixed) frame not directly visible in the lab frame
--Mainly inferred from non-invasive spectroscopy methods.



$1 \text{ fm/c} = 3 \times 10^{-24} \text{ seconds}$
 $= 3 \times 10^{-6} \text{ attoseconds}$
 $= 3 \text{ yoctoseconds}$

Imaging nuclear shape in high-energy snapshot

- Nuclear shape in intrinsic (body-fixed) frame not directly visible in the lab frame
--Mainly inferred from non-invasive spectroscopy methods.



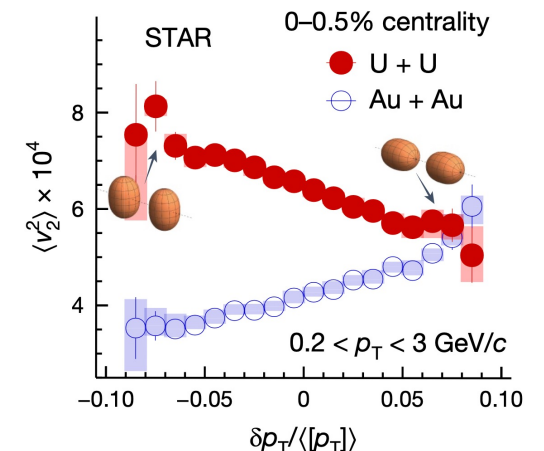
STAR, Nature 635, 67-72 (2024)
<https://www.nature.com/articles/s41586-024-08097-2>

Body-body: large-eccentricity large-size

$v_2 \nearrow$ $p_T \searrow$

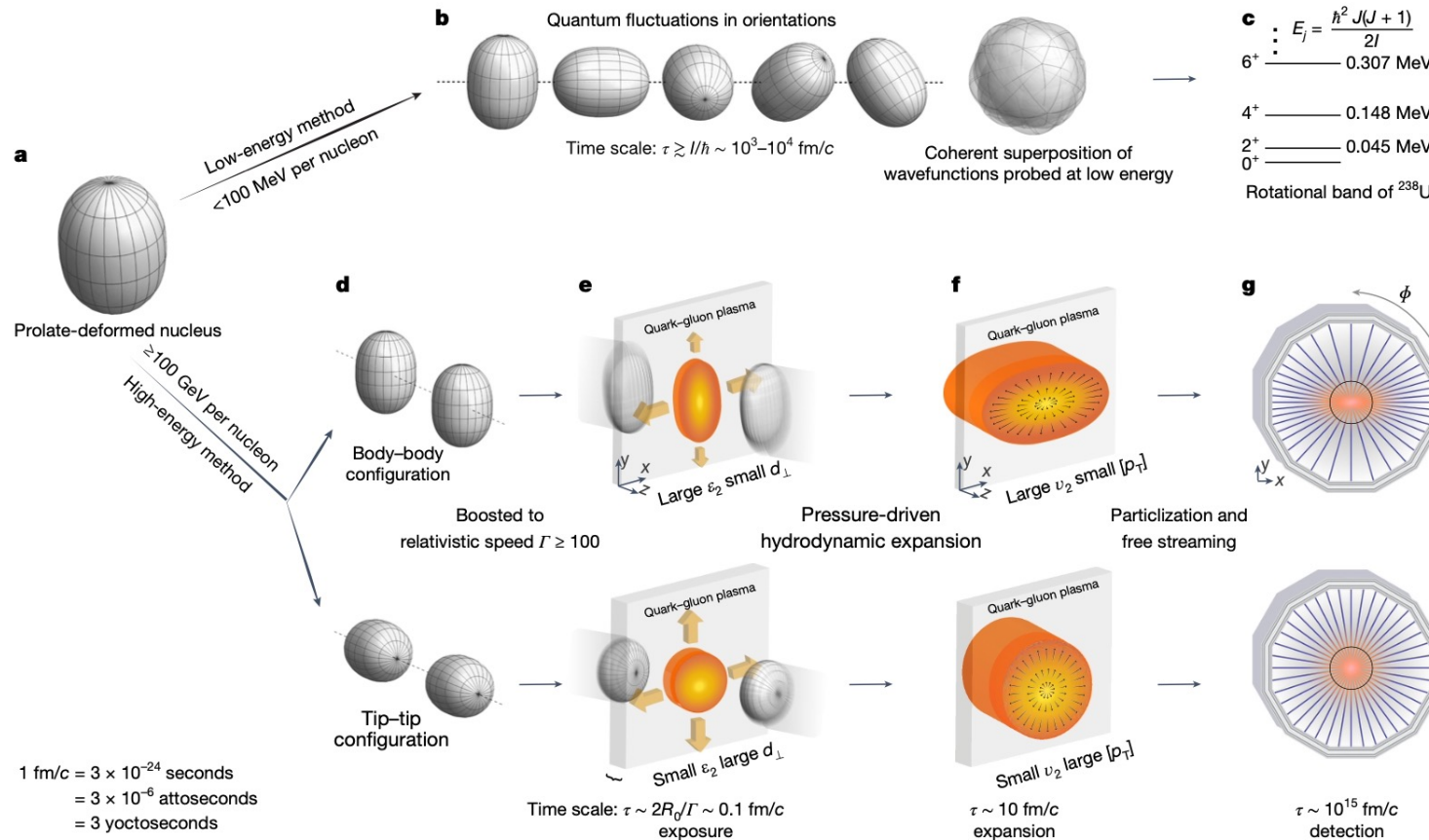
Tip-tip : small-eccentricity small-size

$v_2 \searrow$ $p_T \nearrow$



Imaging nuclear shape in high-energy snapshot

- Nuclear shape in intrinsic (body-fixed) frame not directly visible in the lab frame
--Mainly inferred from non-invasive spectroscopy methods.



STAR, Nature 635, 67-72 (2024)
<https://www.nature.com/articles/s41586-024-08097-2>

across energy scales

Body-body: large-eccentricity large-size

$v_2 \nearrow \quad p_T \searrow$

Tip-tip : small-eccentricity small-size

$v_2 \searrow \quad p_T \nearrow$

$$\begin{aligned} \langle v_2^2 \rangle &= a_1 + b_1 \beta_2^2, \\ \langle (\delta p_T)^2 \rangle &= a_2 + b_2 \beta_2^2, \\ \langle v_2^2 \delta p_T \rangle &= a_3 - b_3 \beta_2^3 \cos(3\gamma). \end{aligned}$$

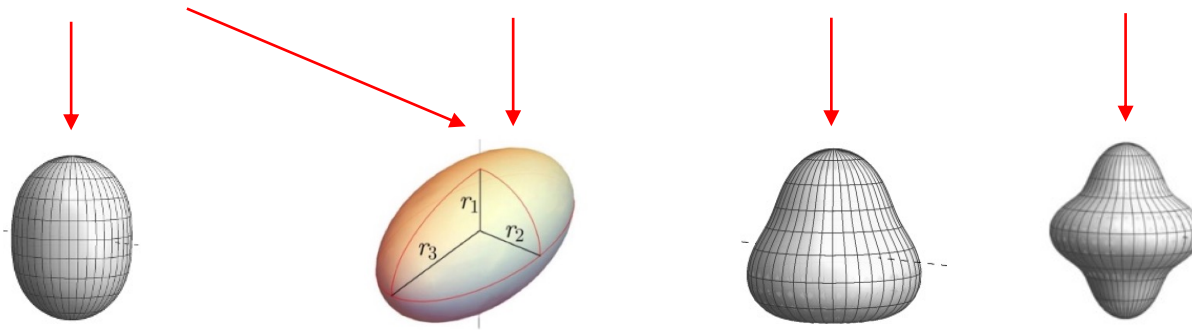
G. Giacalone, J. Jia, C. Zhang, PRL 127, 242301(2021)

Shape-frozen like a snapshot during nuclear crossing ($10^{-25}\text{s} \ll$ rotational time scale 10^{-21}s)
 probe entire mass distribution in the intrinsic frame via multi-point correlations

II. Nuclear deformation in heavy ^{238}U nucleus

$$\rho(r, \theta, \phi) = \frac{\rho_0}{1 + e^{(r-R(\theta, \phi))/a_0}}$$

$$R(\theta, \phi) = R_0 (1 + \beta_2 [\cos \gamma Y_{2,0}(\theta, \phi) + \sin \gamma Y_{2,2}(\theta, \phi)] + \beta_3 Y_{3,0}(\theta, \phi) + \beta_4 Y_{4,0}(\theta, \phi))$$



W. Ryssens, G. Giacalone, B. Schenke, C. Shen, *PRL*130, 212302(2023)

DFT calculations predict a slightly small WS deformation

$$\beta_{2\text{U}} \approx 0.28 \rightarrow \beta_{2\text{U,WS}} \approx 0.25$$

corresponding to a larger volume deformation in presence of $\beta_{4\text{U}} \sim 0.1$ $\beta_{2,\text{body}} = \frac{4\pi}{3R_0^2 A} \int d^3r \rho(\mathbf{r}) r^2 Y_{20}$

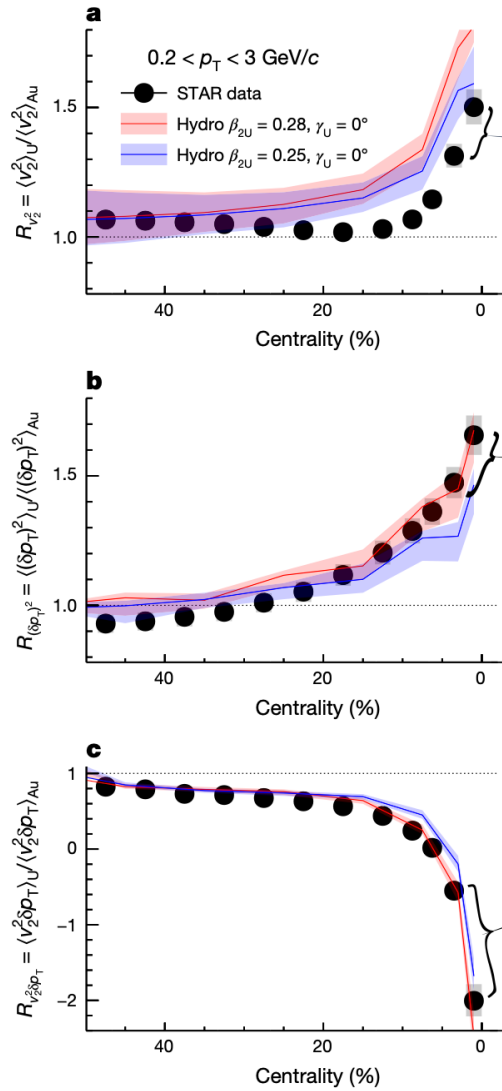
Low-energy estimate with rigid rotor assumption from B(E2) data $\beta_{2,\text{LD}} = \frac{4\pi}{5R_0^2 Z} \sqrt{\frac{B(\text{E2})}{e^2}}$

$$\beta_{2\text{U,LD}} = 0.287 \pm 0.007 \quad \gamma_{\text{U,LD}} = 6^\circ - 8^\circ$$

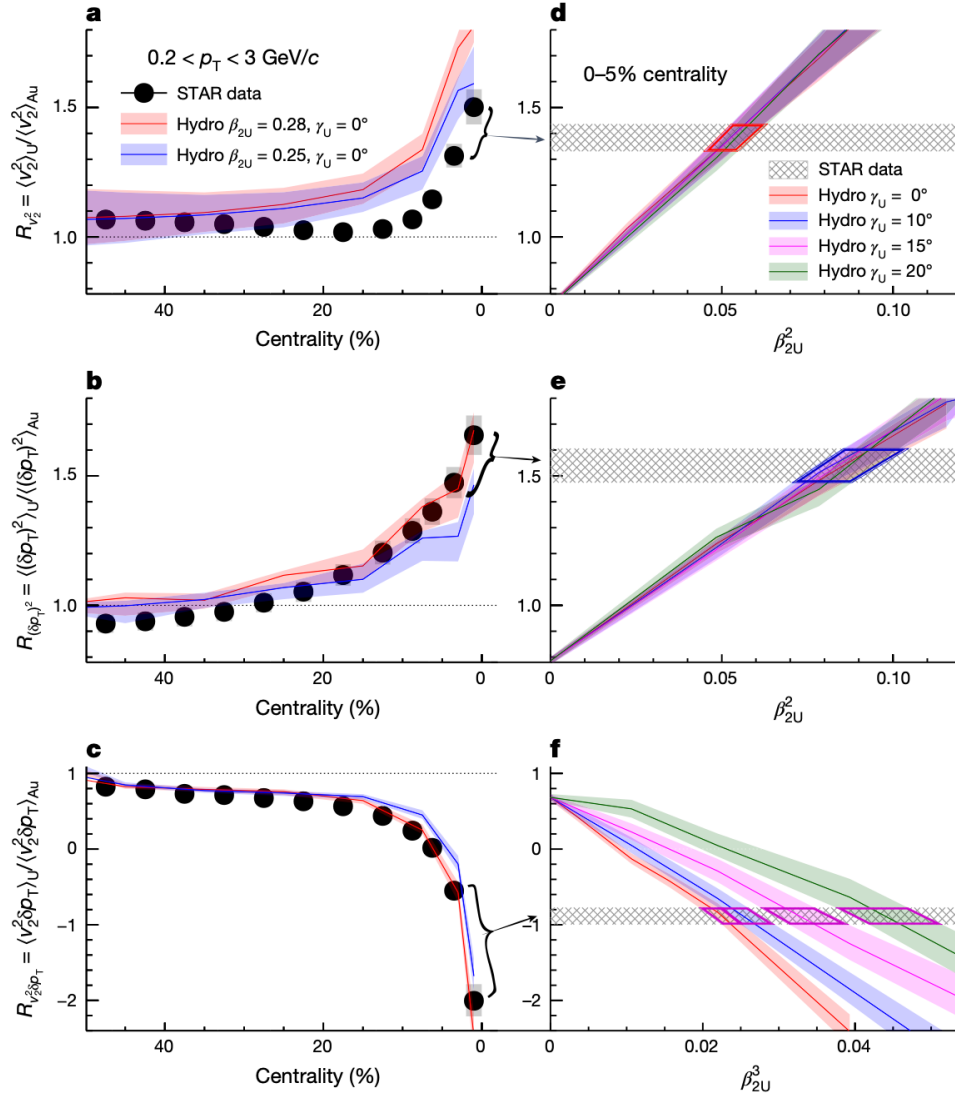
$$\beta_{3\text{U}} \sim \beta_{4\text{U}} \sim 0.1$$

B. Pritychenko et al., *J.ADT*107, 1(2016)
C. Y. Wu et al., *PRC*54, 2356(1996)

Constraining the ground-state ^{238}U : $\beta_{2,U}$ and γ_U



Constraining the ground-state ^{238}U : $\beta_{2,U}$ and γ_U



Sufficient precision is achieved from ratios in ultra-central collisions

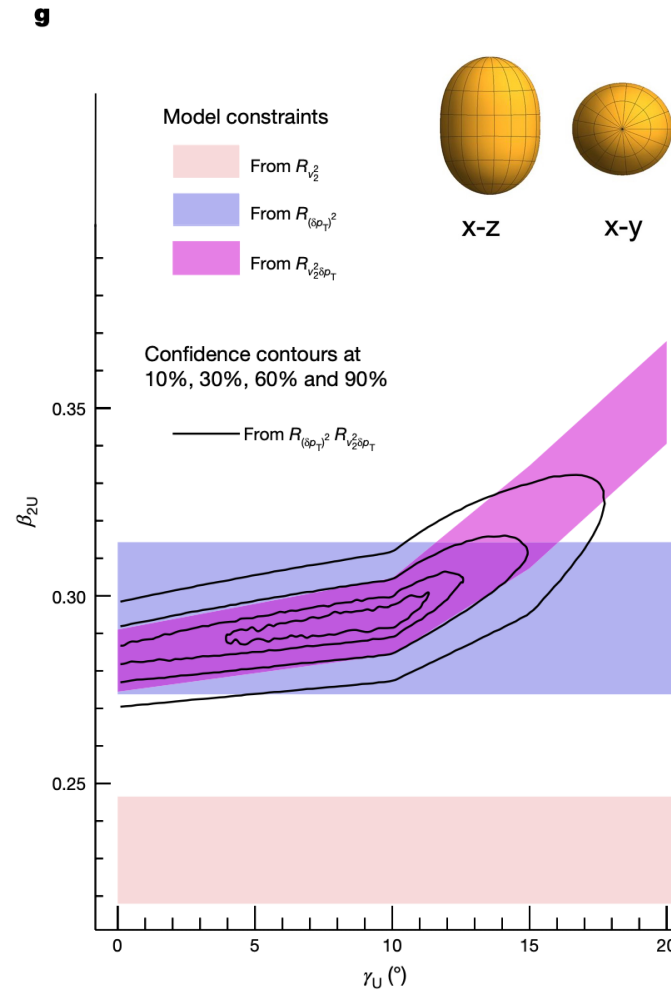
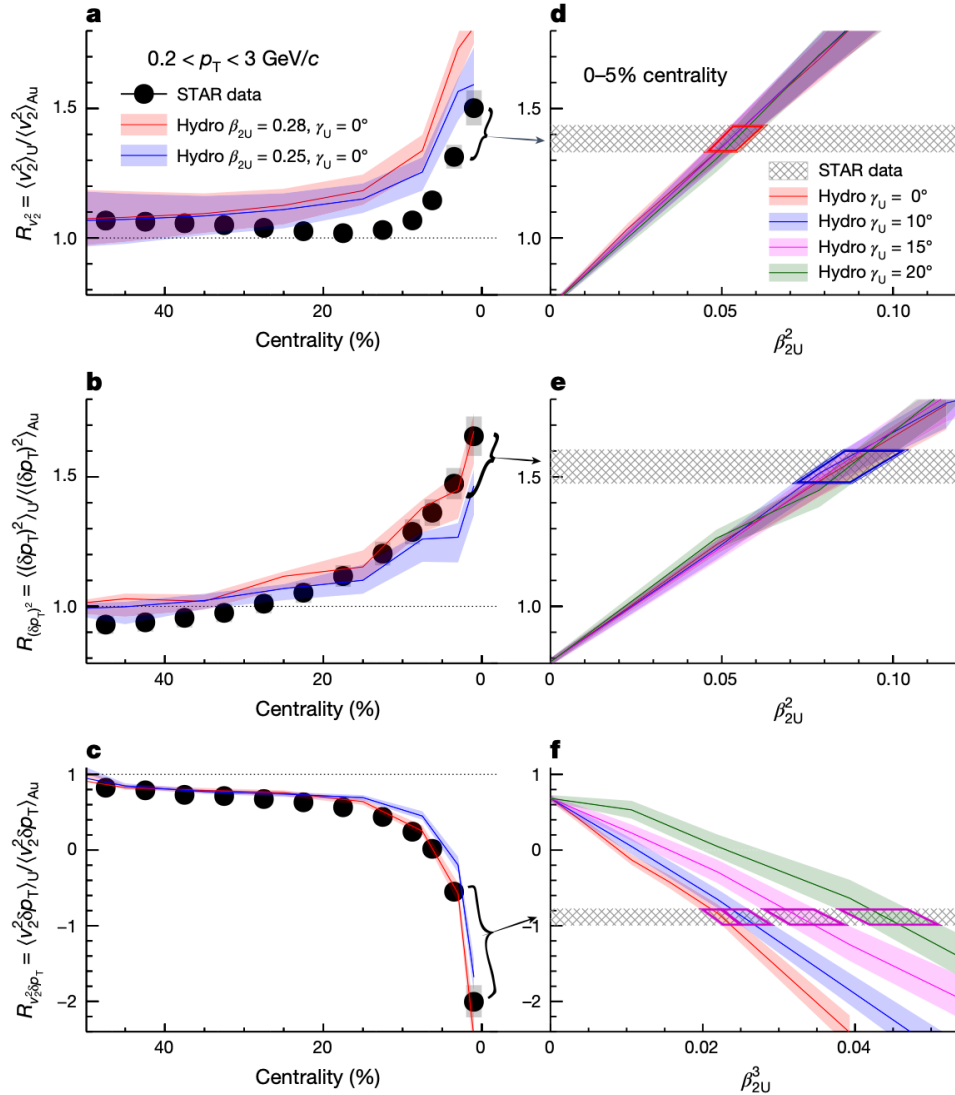
Relation confirmed from hydro

$$R_{(v_2^2)} \approx 1 + \frac{b_1}{a_1} \beta_{2,U}^2,$$

$$R_{\langle (\delta p_T)^2 \rangle} \approx 1 + \frac{b_2}{a_2} \beta_{2,U}^2,$$

$$R_{\langle v_2^2 \delta p_T \rangle} \approx 1 - \frac{b_3}{a_3} \beta_{2,U}^3 \cos(3\gamma)$$

Constraining the ground-state ^{238}U : $\beta_{2,U}$ and γ_U



Sufficient precision is achieved from ratios in ultra-central collisions

Relation confirmed from hydro

$$R_{(v_2^2)} \approx 1 + \frac{b_1}{a_1} \beta_{2,U}^2,$$

$$R_{\langle (\delta p_T)^2 \rangle} \approx 1 + \frac{b_2}{a_2} \beta_{2,U}^2,$$

$$R_{\langle v_2^2 \delta p_T \rangle} \approx 1 - \frac{b_3}{a_3} \beta_{2,U}^3 \cos(3\gamma_U)$$

High-energy estimate (Ip-glasma+MUSIC and Trajectum)

$$\beta_{2,U} = 0.286 \pm 0.025$$

$$\gamma_U = 8.5^\circ \pm 4.8^\circ$$

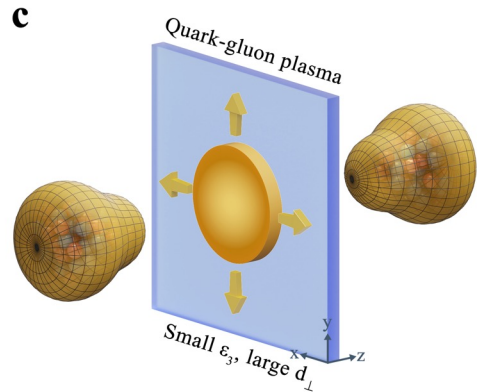
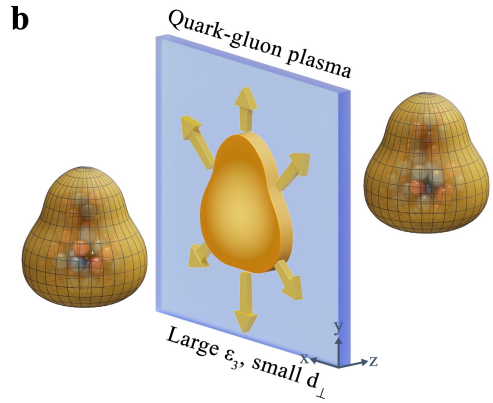
low-energy estimate:

$$\beta_{2,U} = 0.287 \pm 0.007$$

$$\gamma_U = 6^\circ - 8^\circ$$

A large deformation with a slight deviation from axial symmetry in the nuclear ground-state

Evidence of octupole deformation $\beta_{3,U}$



However, v_3 is fluctuation driven, expect in central

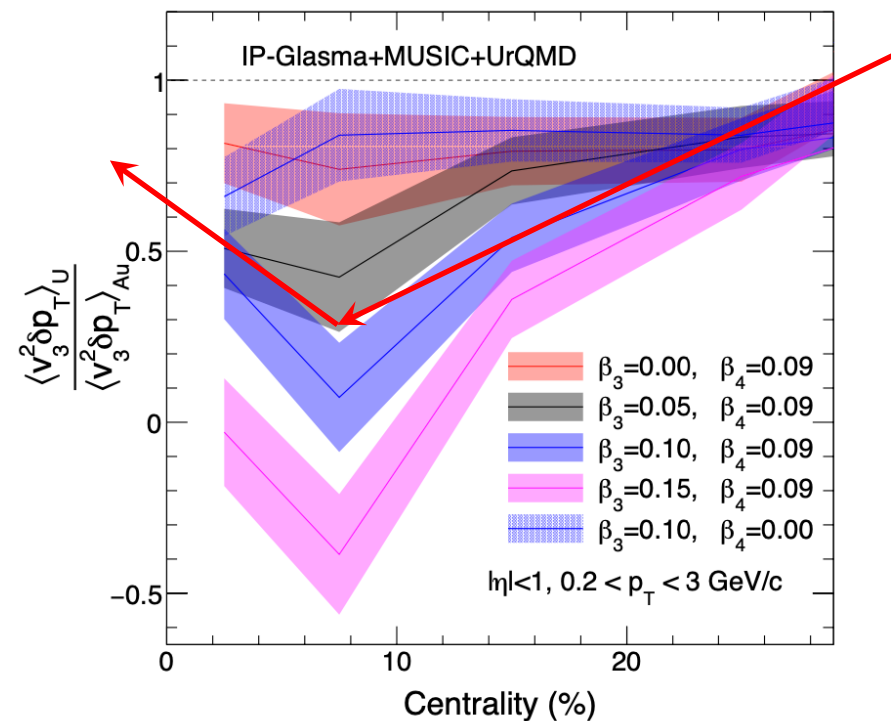
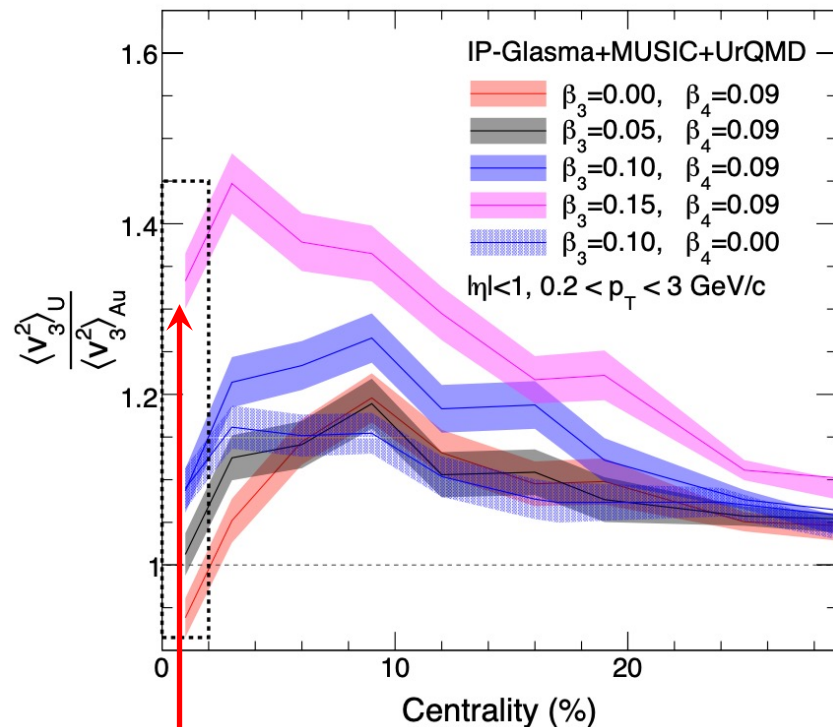
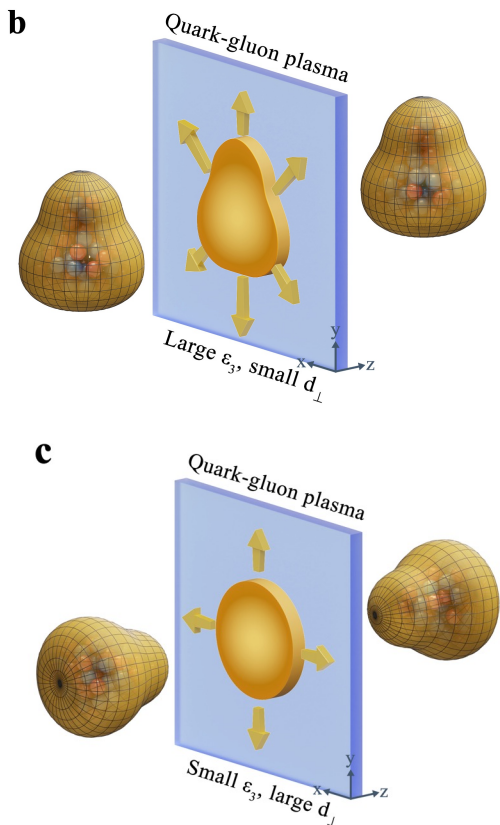
$$\langle v_3^2 \rangle \propto \langle \epsilon_3^2 \rangle \sim 1/A$$

mass number ↗

Evidence of octupole deformation $\beta_{3,U}$

IP-Glasma+MUSIC calculations

C. Zhang, J. Jia, J. Chen, C. Shen, L. Liu, 2504.15245



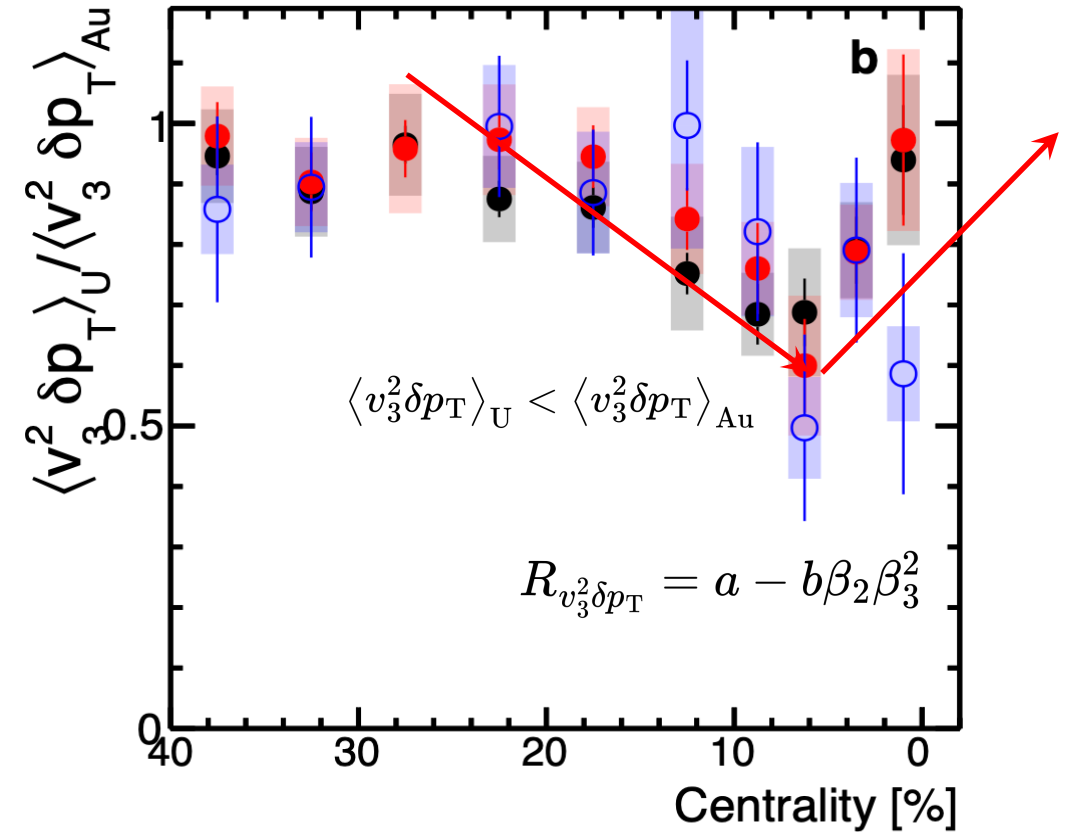
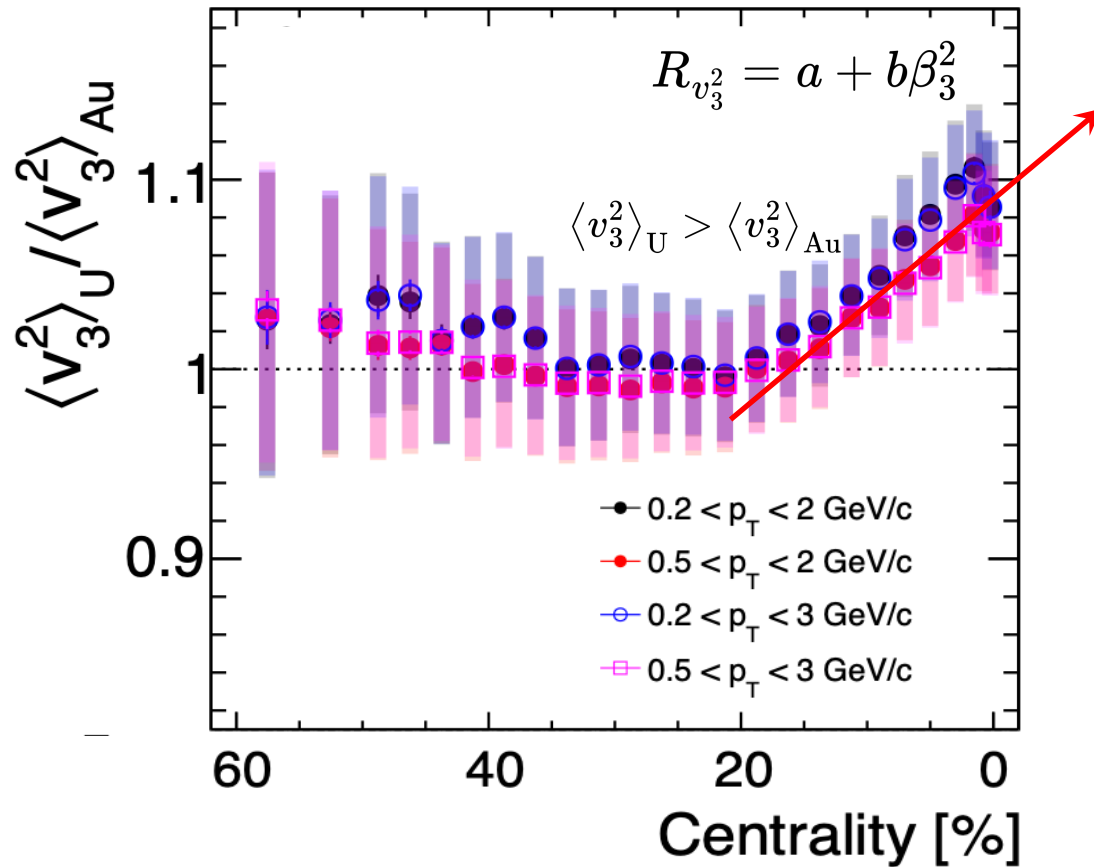
However, v_3 is fluctuation driven, expect in central

$$\langle v_3^2 \rangle \propto \langle \epsilon_3^2 \rangle \sim 1/A$$

mass number

$\langle v_3^2 \rangle$ follows a linear increase with β_3^2 ,

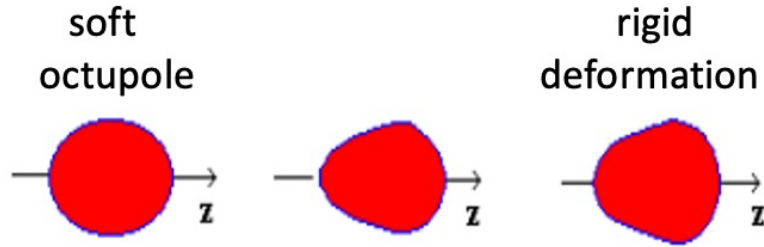
Characteristic anticorrelation in $\langle v_3^2 \delta p_T \rangle$ shows a pronounced β_3 -dependent suppression.



- No apparent p_T dependence
- Order of v_3 and v_3 - p_T reversed by considering non-zero $\beta_{3,U}$, $\beta_{4,U}$.
- An evidence and modest $\beta_{3,U} \sim 0.08$ - 0.10 are confirmed and $\beta_{3U} \sim \beta_{4U}$

Probe $\beta_{3,U}$ and its fluctuation

Octupole collectivity



$$\langle \beta_3^2 \rangle = \bar{\beta}_3^2 + \sigma_{\beta_3}^2$$

$$c_{n,\epsilon}\{2\} = \langle \epsilon_n^2 \rangle \approx \langle \epsilon_{n,0}^2 \rangle + \langle p_n p_n^* \rangle \langle \beta_n^2 \rangle$$

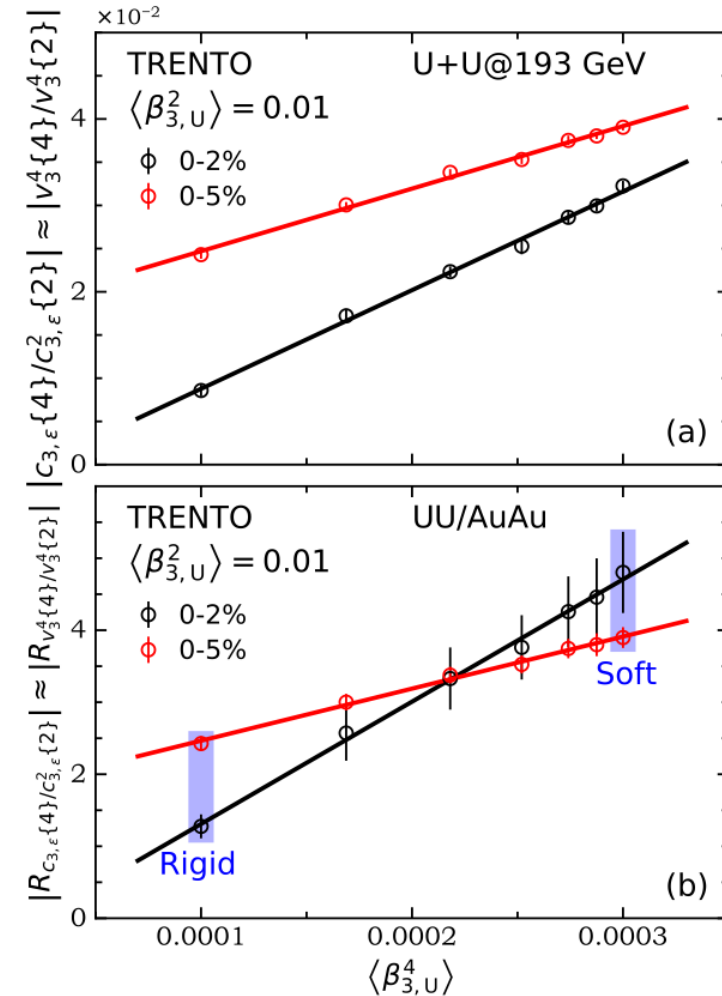
$$c_{n,\epsilon}\{4\} = \langle \epsilon_n^4 \rangle - 2\langle \epsilon_n^2 \rangle^2$$

$$\approx \langle \epsilon_{n,0}^4 \rangle - 2\langle \epsilon_{n,0}^2 \rangle^2 + \langle p_n^2 p_n^{2*} \rangle \langle \beta_n^4 \rangle - 2\langle p_n p_n^* \rangle^2 \langle \beta_n^2 \rangle^2$$

Four-particle correlation is linearly scaled to $\langle \beta_{3,U}^4 \rangle$.

A way to discriminate between static and dynamic collective modes in high-energy nuclear collisions

L. Liu, C. Zhang, J. Chen, J. Jia, X. Huang, Y.-G. Ma, 2509.09376



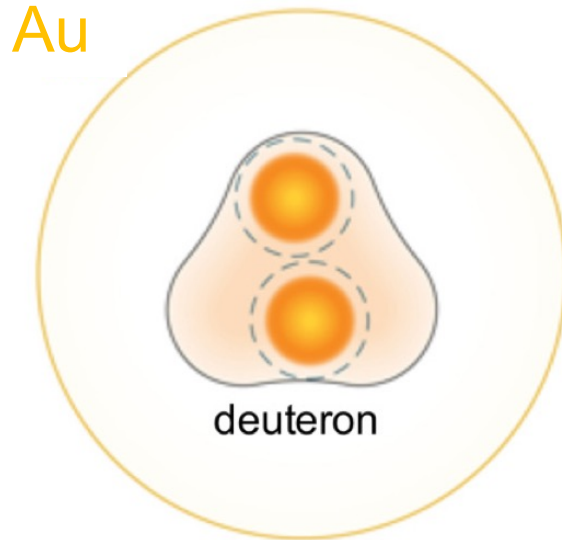
III. Nucleonic cluster pattern in light ^{16}O nucleus

--- from one-body distribution to many-body nucleon correlations

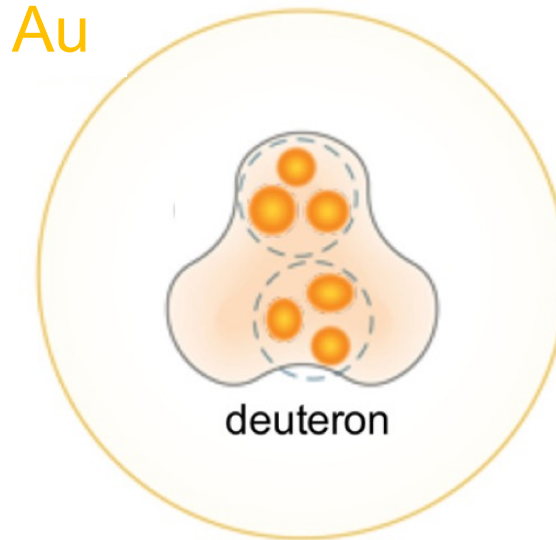
$$\rho(r) \propto \frac{1 + w(r^2/R^2)}{1 + e^{(r-R)/a_0}} \longrightarrow \text{Modern } ab \text{ initio first-principle calculations}$$

Light nuclear geometry and nucleon-nucleon correlations

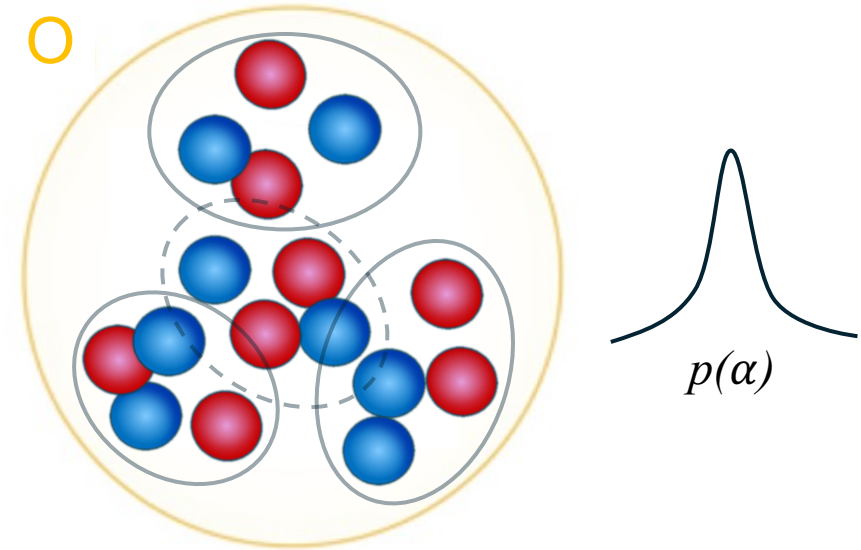
Nucleon fluctuation



subnucleon fluctuation



NN correlation (cluster pattern)



STAR, PRL 130, 242301 (2023)

STAR, PRC 110, 064902 (2024)

W. He, Y.-G. Ma et al., PRL 113, 032506 (2014)

Y.-G. Ma, S. Zhang, Handbook of Nuclear Physics (2022)

...

d+Au collision, 200 GeV

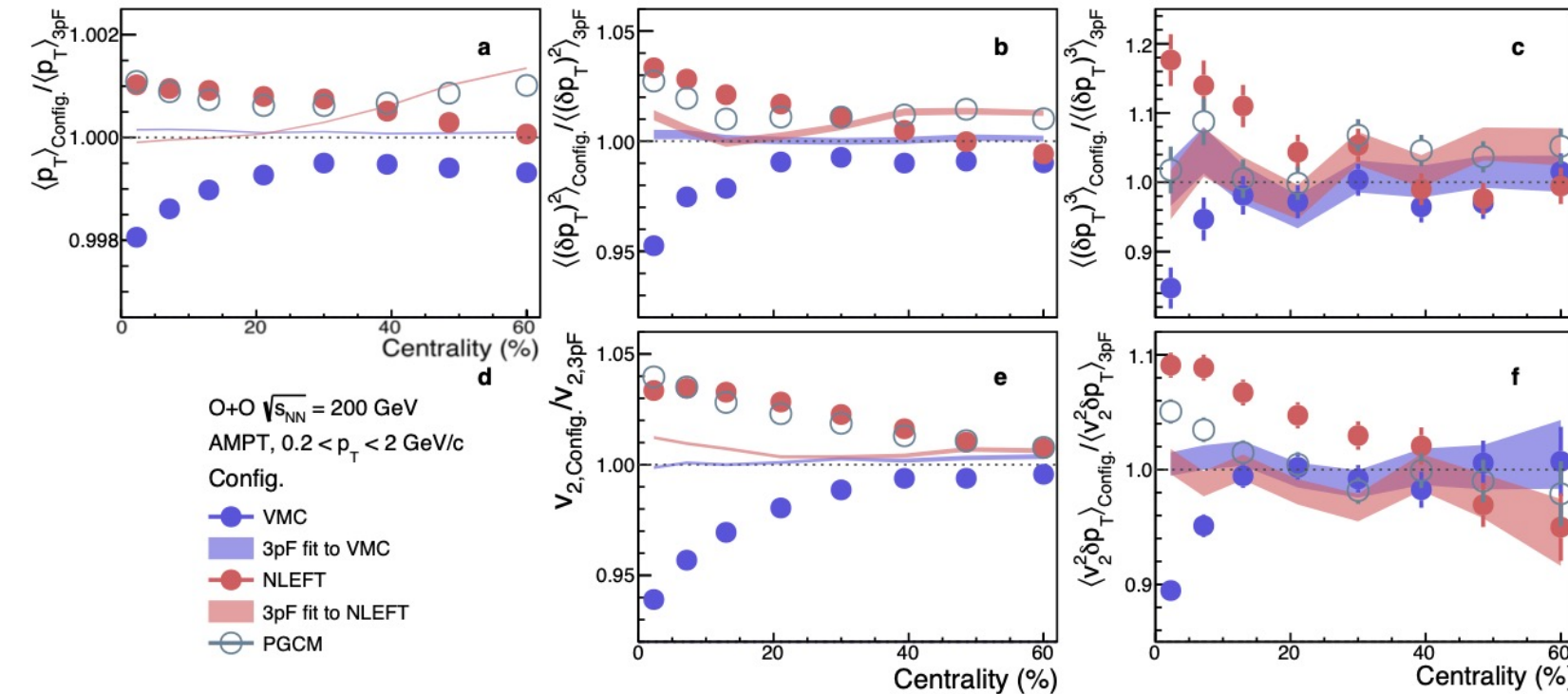
Nearly same N_{ch}

O+O collision, 200 GeV

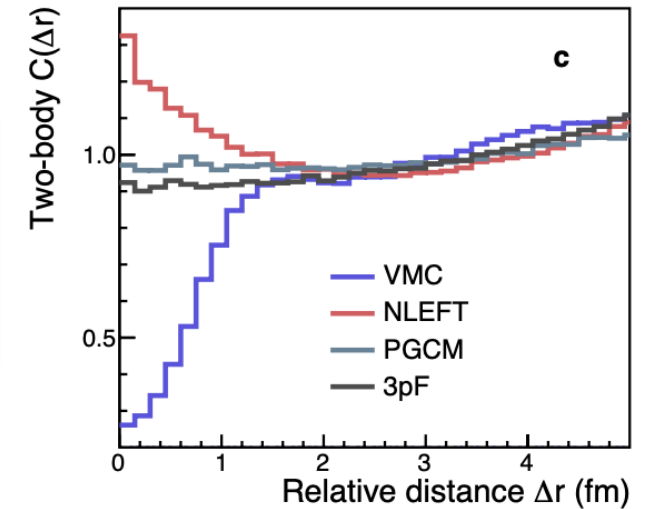
Ab-initio nucleon-nucleon correlations and their impact on high-energy

C. Zhang, J. Chen, G. Giacalone, S. Huang, J. Jia, Y.-G. Ma, *PLB* 862, 139322 (2025)

Multiple final-state observables



Two-body correlations in coordinate space



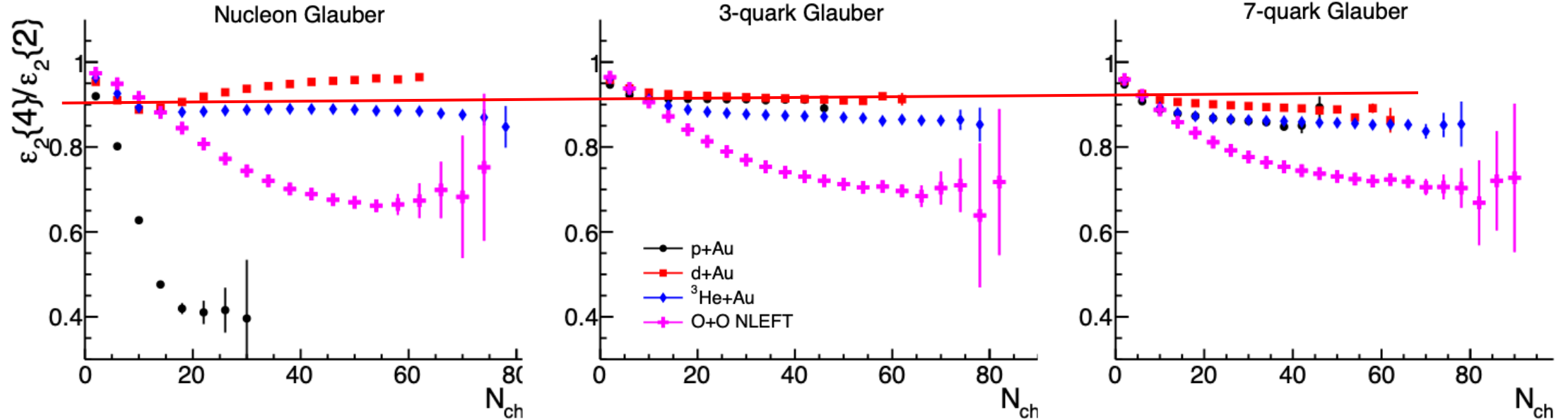
$$C(\Delta r) = \frac{\int d^3s f(\mathbf{s} + \frac{\mathbf{r}}{2}, \mathbf{s} - \frac{\mathbf{r}}{2})}{\int d^3s f(\mathbf{s} + \frac{\mathbf{r}}{2}) f(\mathbf{s} - \frac{\mathbf{r}}{2})}$$

- 3pF as the baseline can NOT capture the NN correlations.
- Distinct differences are observed between VMC and NLEFT/PGCM.
- Central collision is prominent.

Data analysis is ongoing and will be presented in *sQM*.

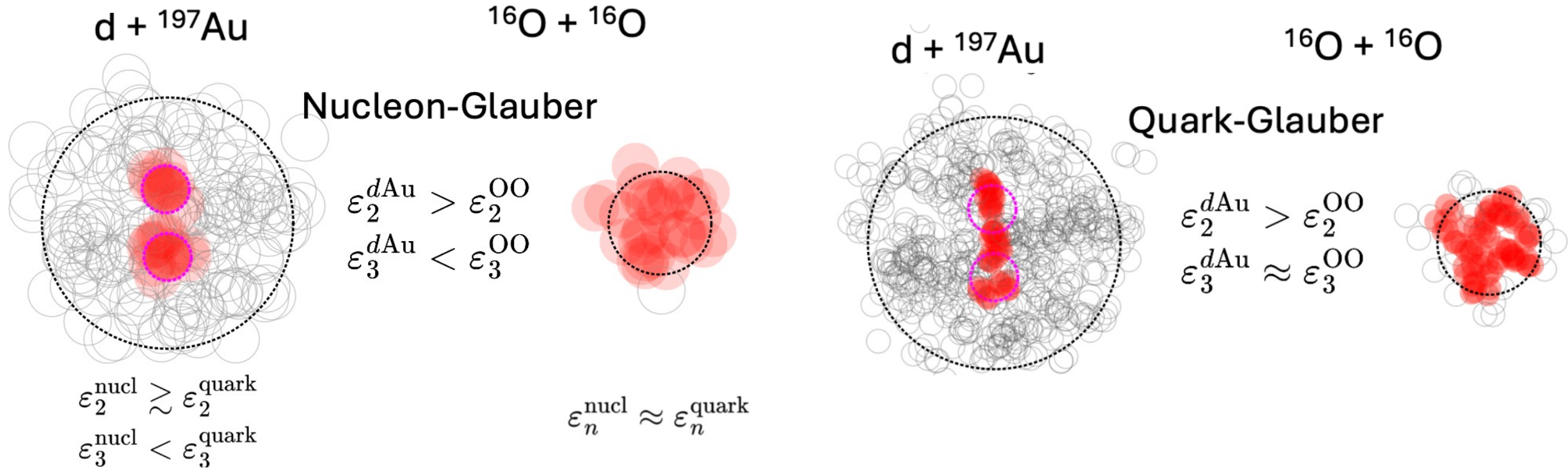
Disentangling nuclear structure and subnucleonic structure effects

S. Huang, J. Jia, C. Zhang, *PLB* 870, 139926 (2025)



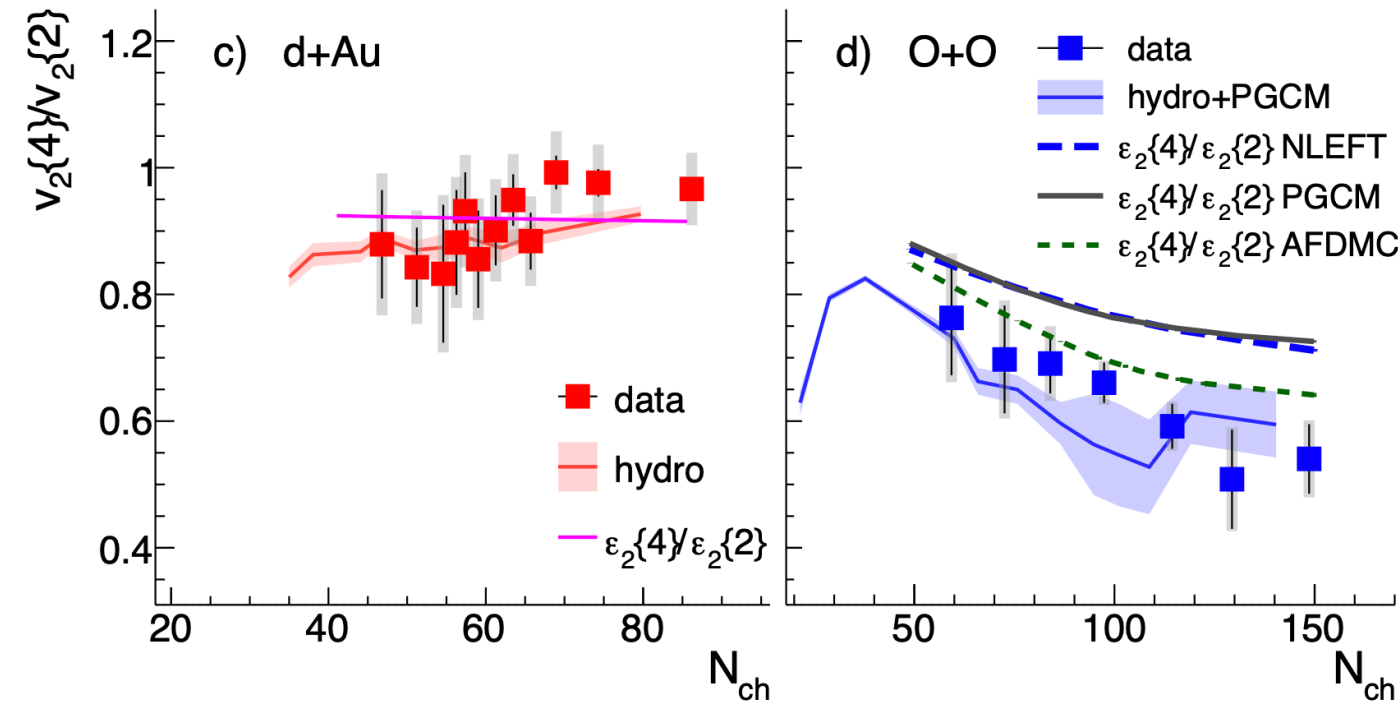
- Quark Glauber is more flattened in d+Au, indicating subnucleon hints
- No clear difference beyond 3 quarks inside of proton.
- Similar behavior in p_T fluctuations.

STAR, 2510.19645



Nucleon-Glauber and quark-Glauber yield different triangular v_3

STAR, 2510.19645



$\varepsilon_2\{4\}/\varepsilon_2\{2\}$ from three models:

AFDMC vs. EFT/PGCM have a visible difference.

Can many-nucleon correlations significantly impact the eccentricity fluctuations? YES!

C. Zhang et al., PLB 862, 139322 (2025)

The interplay between sub-nucleon fluctuation and many-nucleon correlation.

STAR, PRL 130, 242301 (2023)

S. Huang, J. Jia, C. Zhang, PLB 870, 139926 (2025)

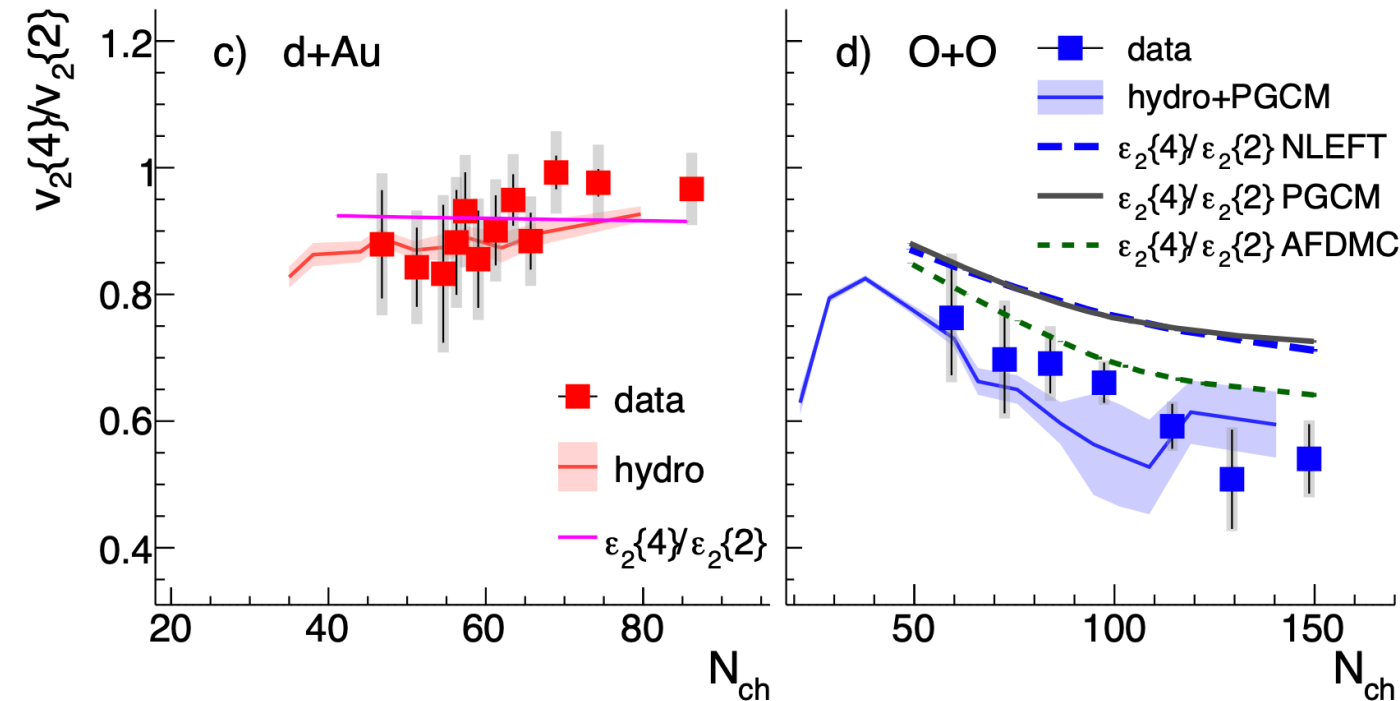
$$\begin{aligned} (v_n\{2\})^2 &= c_n\{2\} = \langle v_n^2 \rangle & \varepsilon_2\{2\}^2 &= \langle \varepsilon_2^2 \rangle \\ (v_n\{4\})^4 &= -c_n\{4\} = 2\langle v_n^2 \rangle^2 - \langle v_n^4 \rangle & \varepsilon_2\{4\}^4 &= 2\langle \varepsilon_2^2 \rangle^2 - \langle \varepsilon_2^4 \rangle \end{aligned}$$

Y.-G. Ma, S. Zhang, Handbook of Nuclear Physics (2022); W. He, Y.-G. Ma et al, PRL 113, 032506 (2014)

G. Giacalone, G. Nijs et al., PRL135, 012302 (2025), G. Giacalone, W. Zhao et al., PRL134, 082301 (2025); P. Li, B. Zhou, G.-L. Ma, 2504.04688

Y. Wang, S. Zhao, B. Cao, H. Xu, H. Song, PRC 109, L051904 (2024); X. Zhao, G.-L. Ma. Y. Zhou, Z. Lin, C. Zhang, 2404.09780; S. Jahan, Roch, C. Shen, 2507. 11394

STAR, 2510.19645



$\varepsilon_2\{4\}/\varepsilon_2\{2\}$ from three models:

AFDMC vs. EFT/PGCM have a visible difference.

Can many-nucleon correlations significantly impact the eccentricity fluctuations? YES!

C. Zhang et al., PLB 862, 139322 (2025)

The interplay between sub-nucleon fluctuation and many-nucleon correlation.

STAR, PRL 130, 242301 (2023)

S. Huang, J. Jia, C. Zhang, PLB 870, 139926 (2025)

Geometric scan elucidates nuclear tomography and strong nuclear force?

more low-energy model inputs & revisit

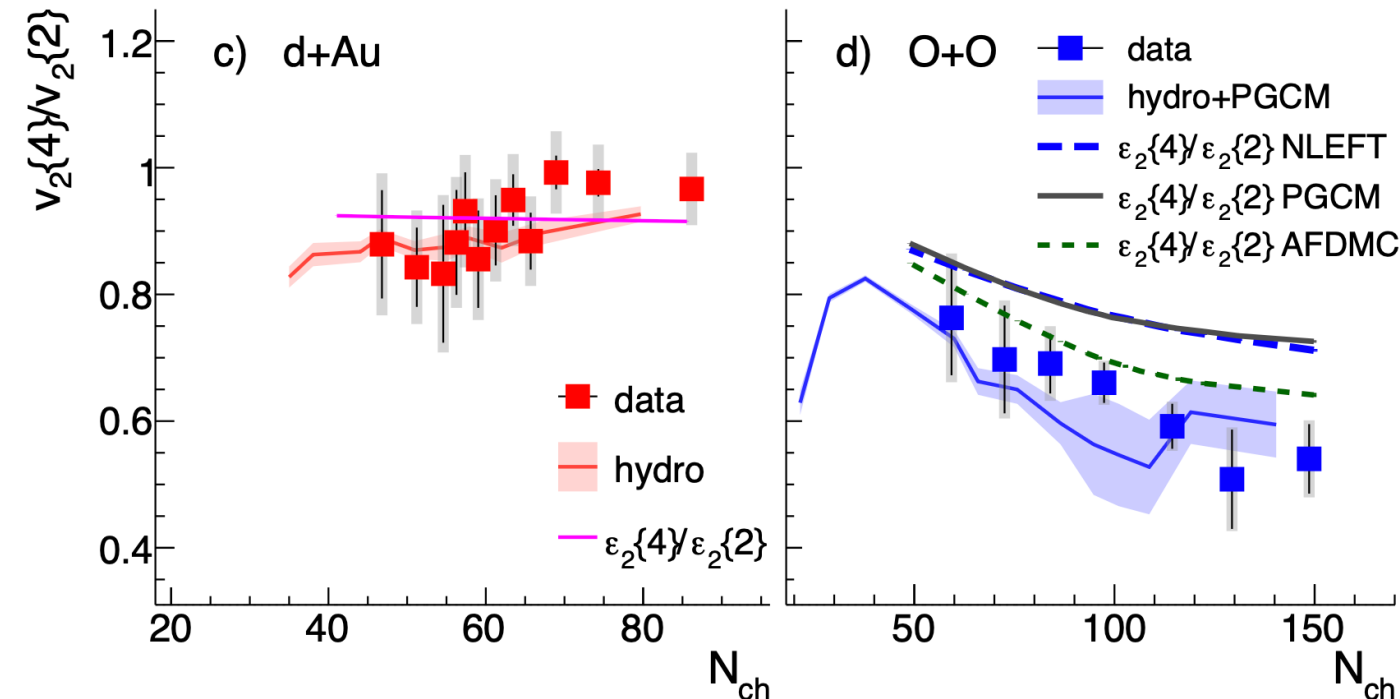
$$\begin{aligned} (v_n\{2\})^2 &= c_n\{2\} = \langle v_n^2 \rangle & \varepsilon_2\{2\}^2 &= \langle \varepsilon_2^2 \rangle \\ (v_n\{4\})^4 &= -c_n\{4\} = 2\langle v_n^2 \rangle^2 - \langle v_n^4 \rangle & \varepsilon_2\{4\}^4 &= 2\langle \varepsilon_2^2 \rangle^2 - \langle \varepsilon_2^4 \rangle \end{aligned}$$

Y.-G. Ma, S. Zhang, Handbook of Nuclear Physics (2022); W. He, Y.-G. Ma et al, PRL 113, 032506 (2014)

G. Giacalone, G. Nijs et al., PRL135, 012302 (2025), G. Giacalone, W. Zhao et al., PRL134, 082301 (2025); P. Li, B. Zhou, G.-L. Ma, 2504.04688

Y. Wang, S. Zhao, B. Cao, H. Xu, H. Song, PRC 109, L051904 (2024); X. Zhao, G.-L. Ma. Y. Zhou, Z. Lin, C. Zhang, 2404.09780; S. Jahan, Roch, C. Shen, 2507. 11394

STAR, 2510.19645



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$$(v_n\{2\})^2 = c_n\{2\} = \langle v_n^2 \rangle$$

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LHC (ALICE, ATLAS, CMS, LHCb) also released the O+O/Ne+Ne.
Shape imaging method for light ions (New avenue)

Y.-G. Ma, S. Zhang, Handbook of Nuclear Physics (2022); W. He, Y.-G. Ma et al, PRL 113, 032506 (2014)

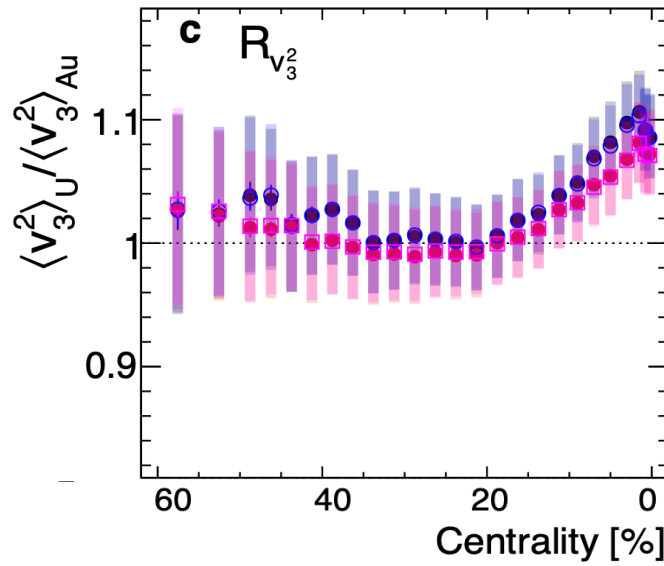
G. Giacalone, G. Nijs et al., PRL135, 012302 (2025), G. Giacalone, W. Zhao et al., PRL134, 082301 (2025); P. Li, B. Zhou, G.-L. Ma, 2504.04688

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IV. Conclusions and Outlooks

1. The signatures of nuclear structure in nuclear collisions are ubiquitous:

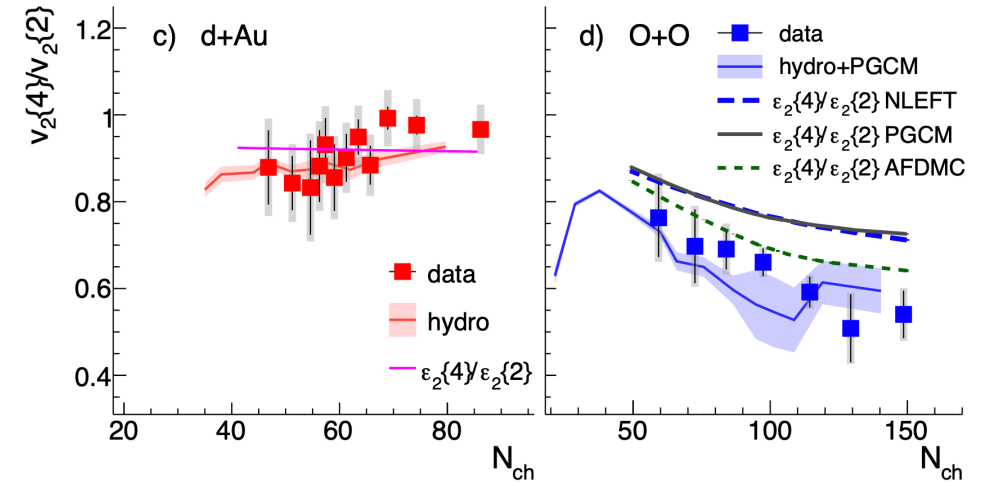
STAR, Rep. Prog. Phys. **88**, 108601 (2025)



$$\beta_{3U} = 0.08 - 0.1; \quad \beta_{3U} \sim \beta_{4U}$$

from large to small nuclei

STAR, 2510.19645

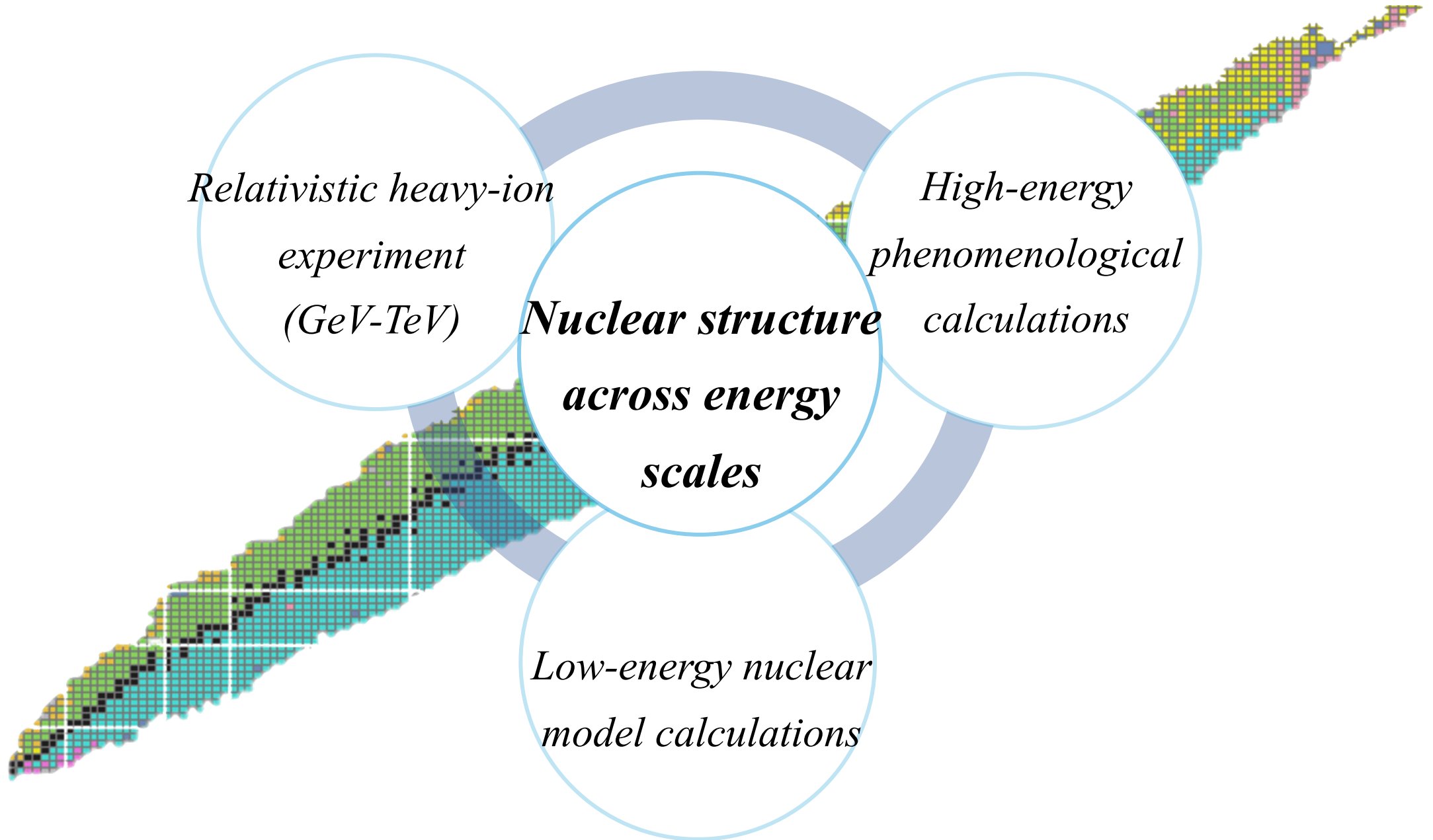


NN correlations is remarkable

2. Many potential applications from large to small collision systems :

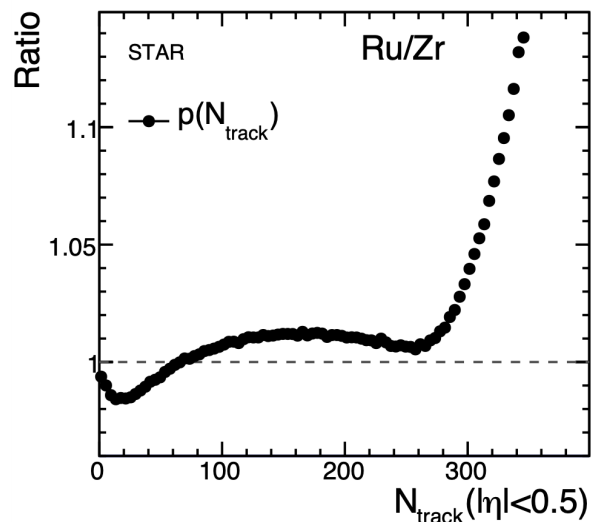
- Rigid and soft β_n and γ (shape fluctuations/coexistence)
- Neutron skin & symmetry energy constraints
- Energy evolution between GeV/TeV and MeV in even-and odd-A nuclei
- Confronted with ab initio calculations and more light-ion studies
- Intersection with other research: Nucleosynthesis, nuclear fission, $0\nu\beta\beta$

IV. Conclusions and Outlooks

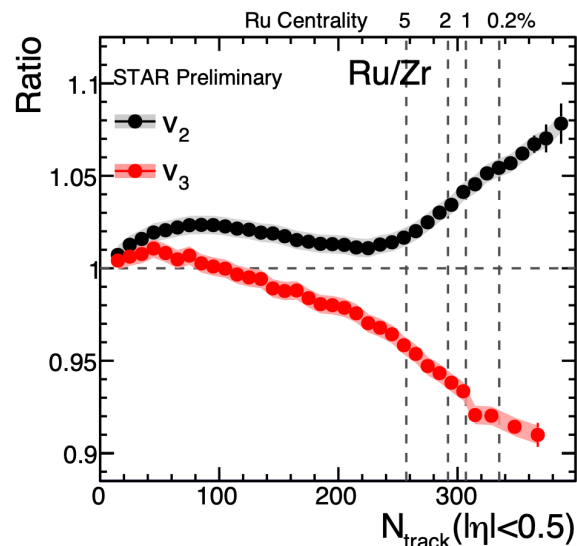


Nuclear structure is inherent of heavy-ion probes

One-body distribution



Two-body correlations



Three-body correlations

