

### Review of MC generators

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Shandong University

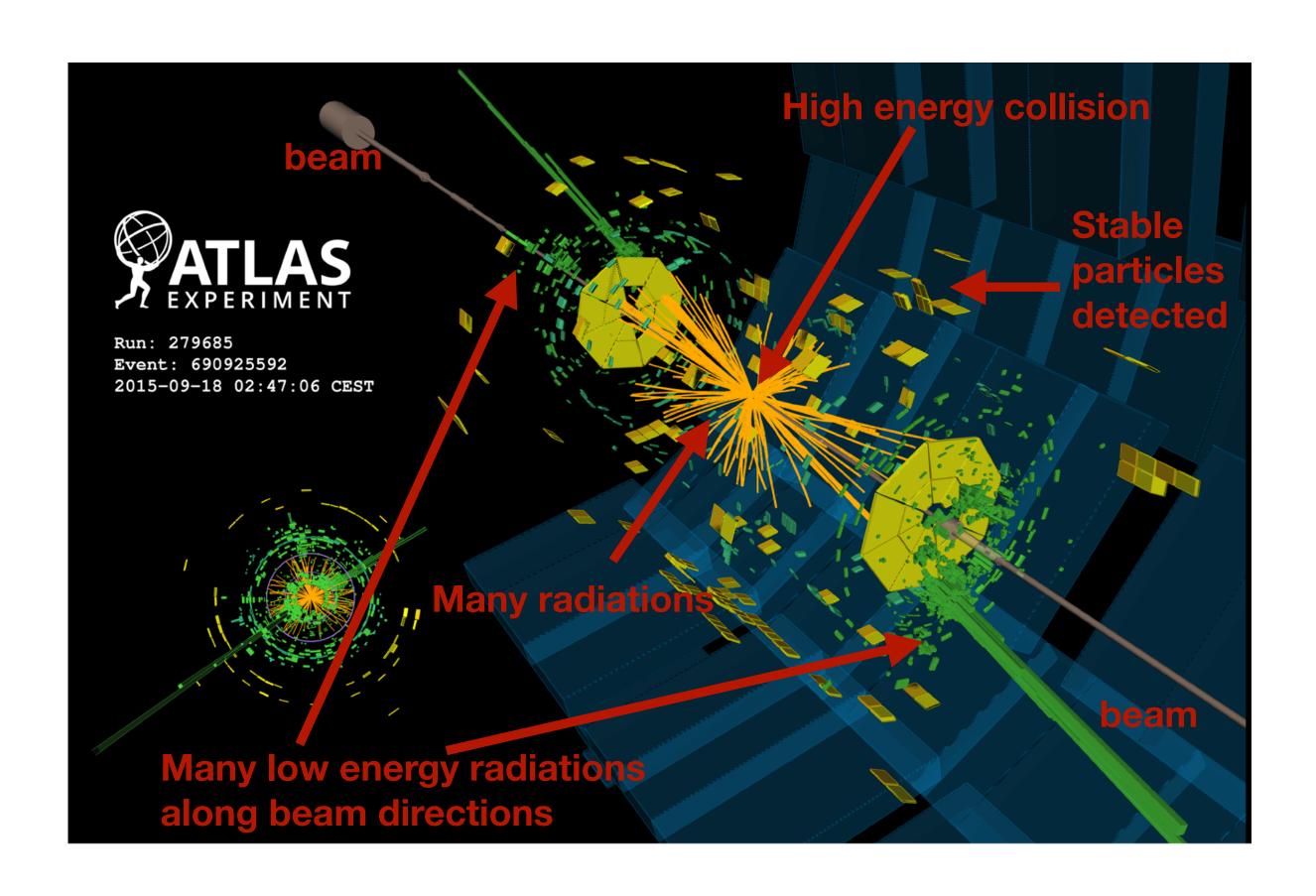
TeV物理前沿专题研讨会暨第31届LHC Mini-Workshop

2025-10-11

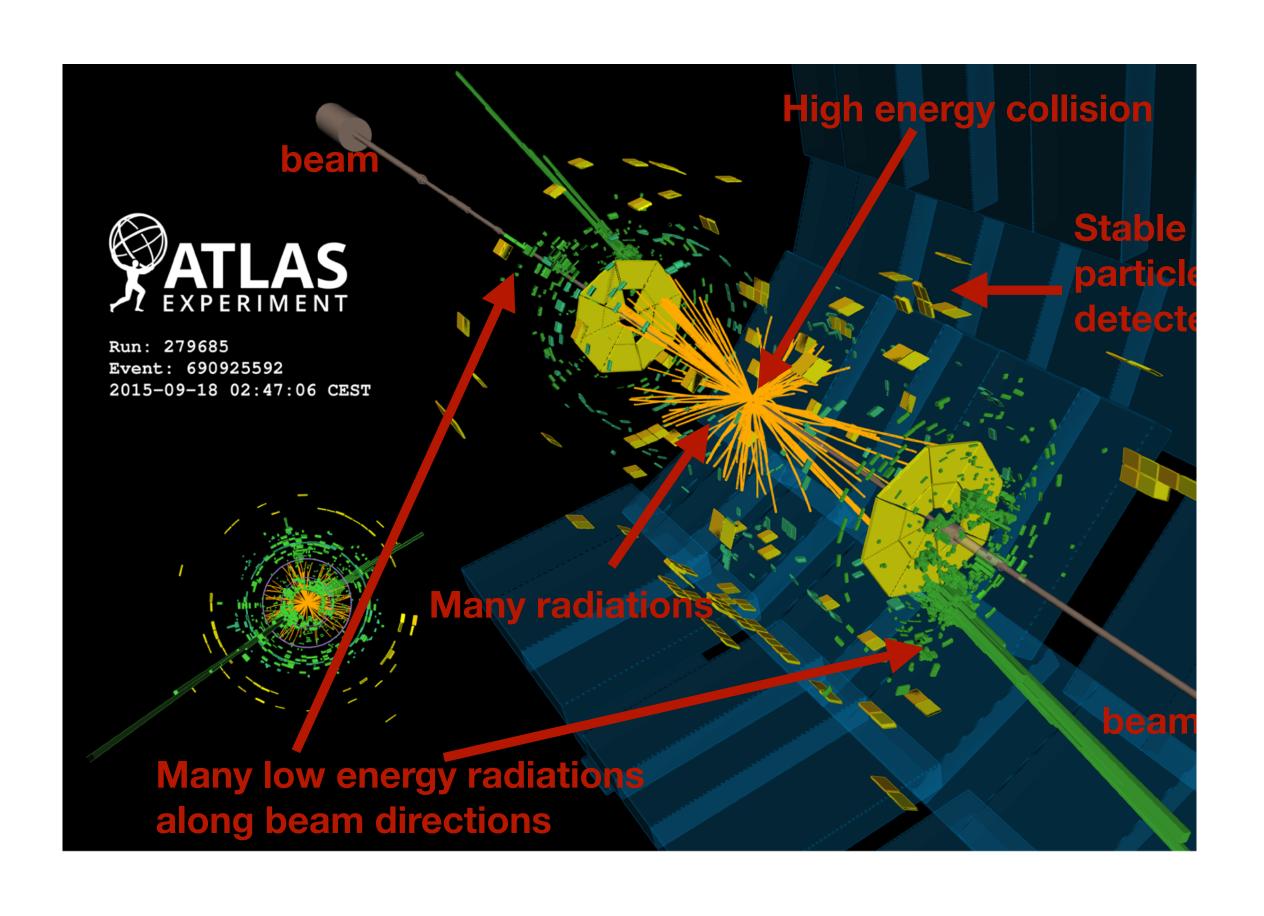
## Outline

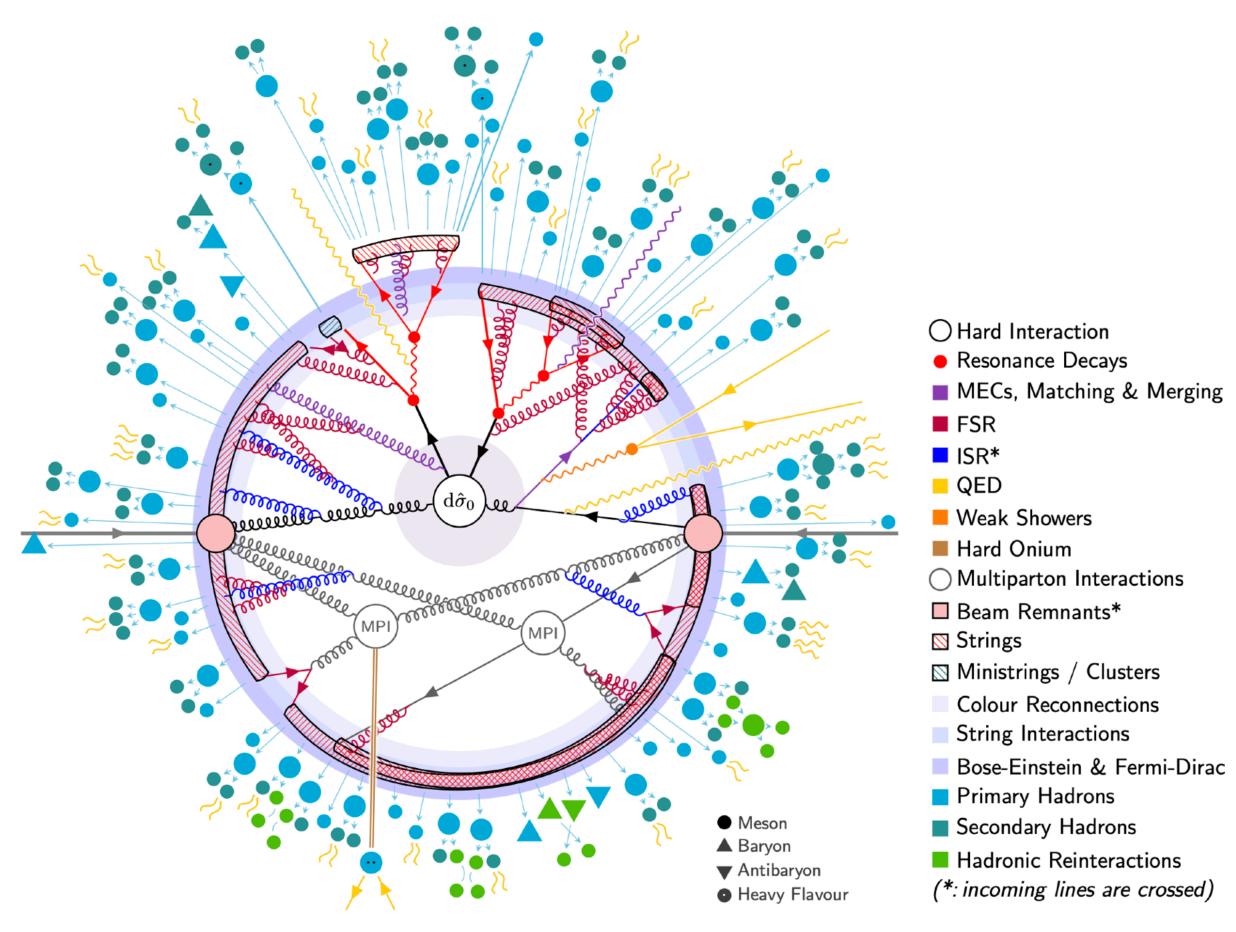
- 1. Introduction to MCEGs
- 2. Parton Showers
- 3. Hadronization models
- 4. Summary

# 1. Introduction to MCEGs



### 1. Introduction to MCEGs





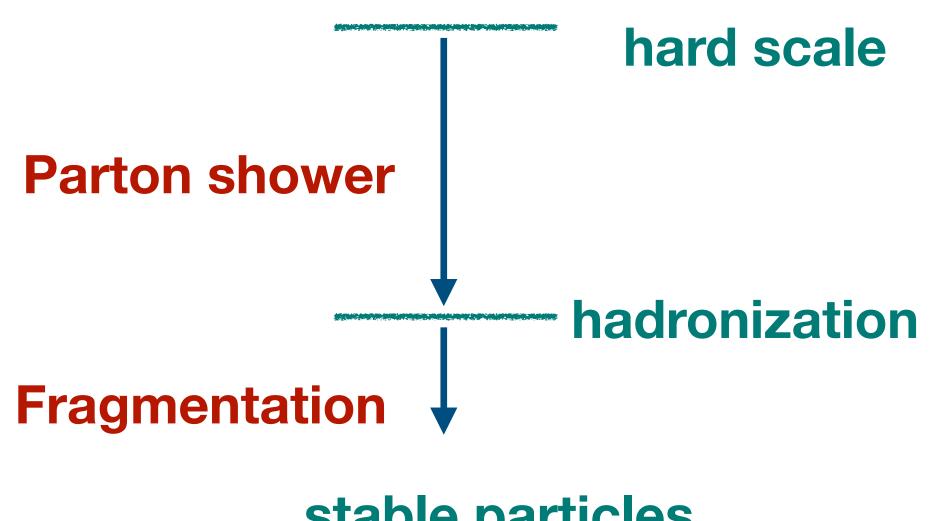
From PYTHIA 8.3

### 1. Introduction to MCEGs

The purpose of Monte Carlo event generators is to generate events in as much details as nature (generate average and fluctuation right)

$$\mathcal{P}_{\text{event}} = \mathcal{P}_{\text{Hard}} \otimes \mathcal{P}_{\text{Decay}} \otimes \mathcal{P}_{\text{ISR}} \otimes \mathcal{P}_{\text{FSR}} \otimes \mathcal{P}_{\text{MPI}} \otimes \mathcal{P}_{\text{Had}} \cdots$$

- ☐ Hard process in high energy
- ☐ Transition from high energy to low energy
  - —parton shower
- ☐ Low energy soft regime
  - —fragmentation



stable particles

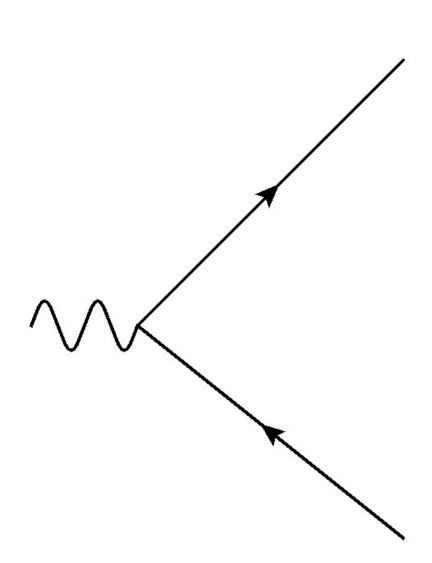
Parton shower: a model for the evolution from high scale to hadronization scale

Parton showers approximate higher-order real-emission corrections to the hard scattering process

- ☐ Generate cascades of radiation automatically
- ☐ Locally conserved four momentum
- ☐ Locally conserved flavor
- ☐ Unitarity by construction

Parton showers

- ☐ sample infrared configurations
- simulate the evolution of parton (resummation)

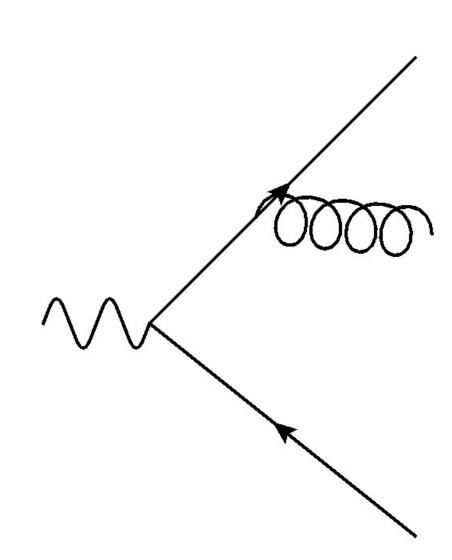


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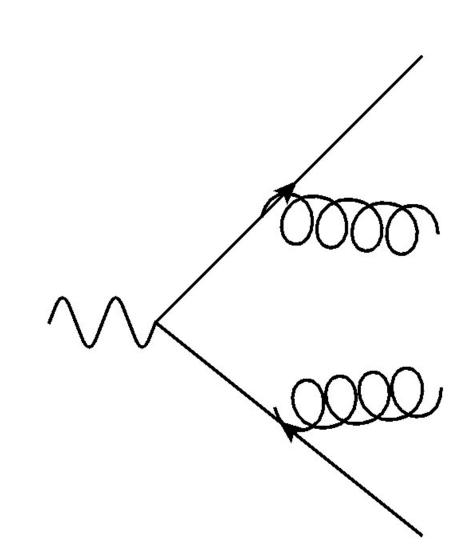


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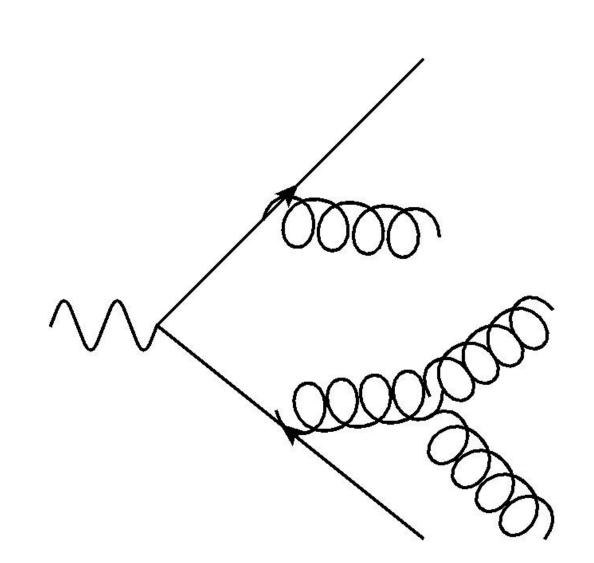


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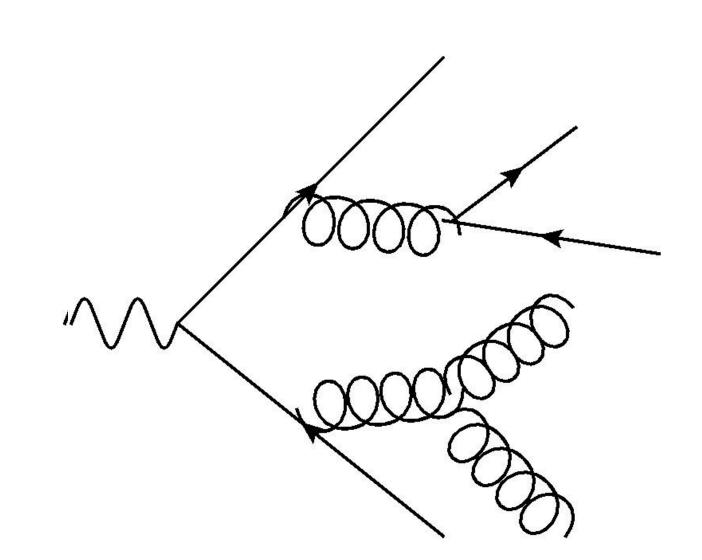


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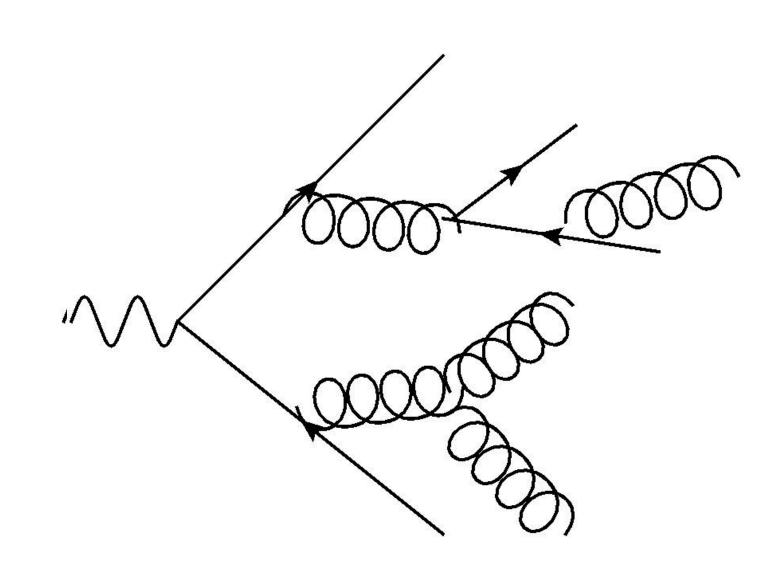


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Sudakov form factor: Non-branching probability  $\exp \left| \int d\phi_{n \to n+1} \frac{|M_{n+1}|^2}{|M_n|^2} \right|$ 

Probability that there is no branching from Q to q is  $\Delta_i\left(Q^2,q^2\right)$ 

choose kinematic variable as the evolution scale

$$\Delta (Q^{2}, q^{2}) = \exp \left\{ \int_{Q^{2}}^{q^{2}} d\phi_{n \to n+1} \frac{|M_{n+1}|^{2}}{|M_{n}|^{2}} \right\}$$

Probability for one observed branching

$$1-\Delta\left(Q^2,q^2\right)$$

Probability one branching between the scale  $q^2$  to  $q^2 + dq^2$ 

$$\frac{d}{dq^2} \Delta (Q^2, q^2) = \Delta (Q^2, q^2) \times d\phi_{n \to n+1} \frac{|M_{n+1}|^2}{|M_n|^2}$$

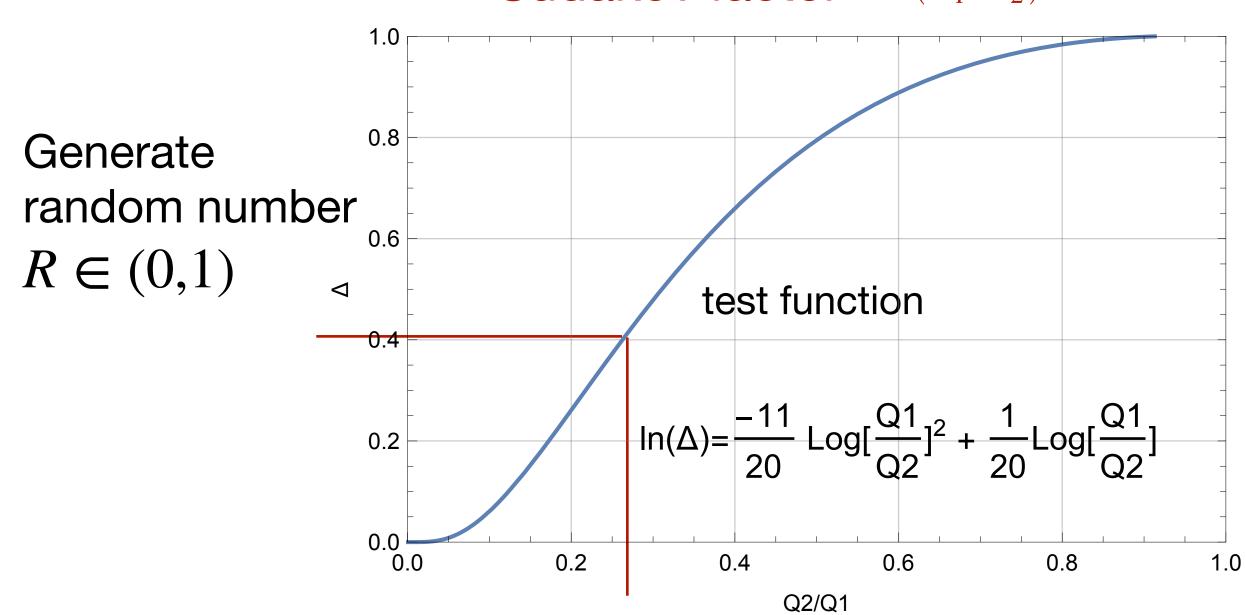
Additional radiations can be added according to the function  $\Delta\left(Q^{2},q^{2}\right)$ 

Phase space mapping 
$$\int \! d\phi_{n \to n+1} \frac{|M_{n+1}|^2}{|M_n|^2} = \int_{q^2}^{\mathcal{Q}^2} \frac{\mathrm{d}k^2}{k^2} \frac{\alpha_{\mathrm{s}}}{2\pi} \int_{\mathcal{Q}_0^2/k^2}^{1-\mathcal{Q}_0^2/k^2} \mathrm{d}z P_{ji}(z) \qquad \qquad \frac{d\theta^2}{\theta^2} = \frac{dq^2}{q^2} = \frac{dk_\perp^2}{k_\perp^2}$$
 many choices for the evolution variables

$$\frac{d\theta^2}{\theta^2} = \frac{dq^2}{q^2} = \frac{dk_\perp^2}{k_\perp^2}$$

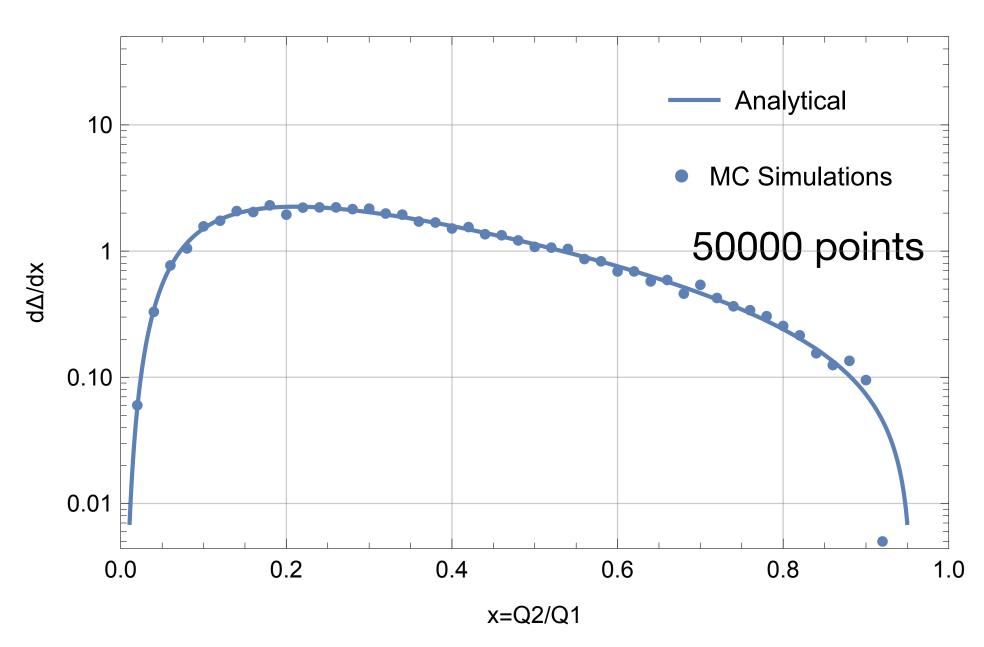
### Monte-Carlo Technique and resummation

### Sudakov factor $\Delta(Q_1^2, Q_2^2)$



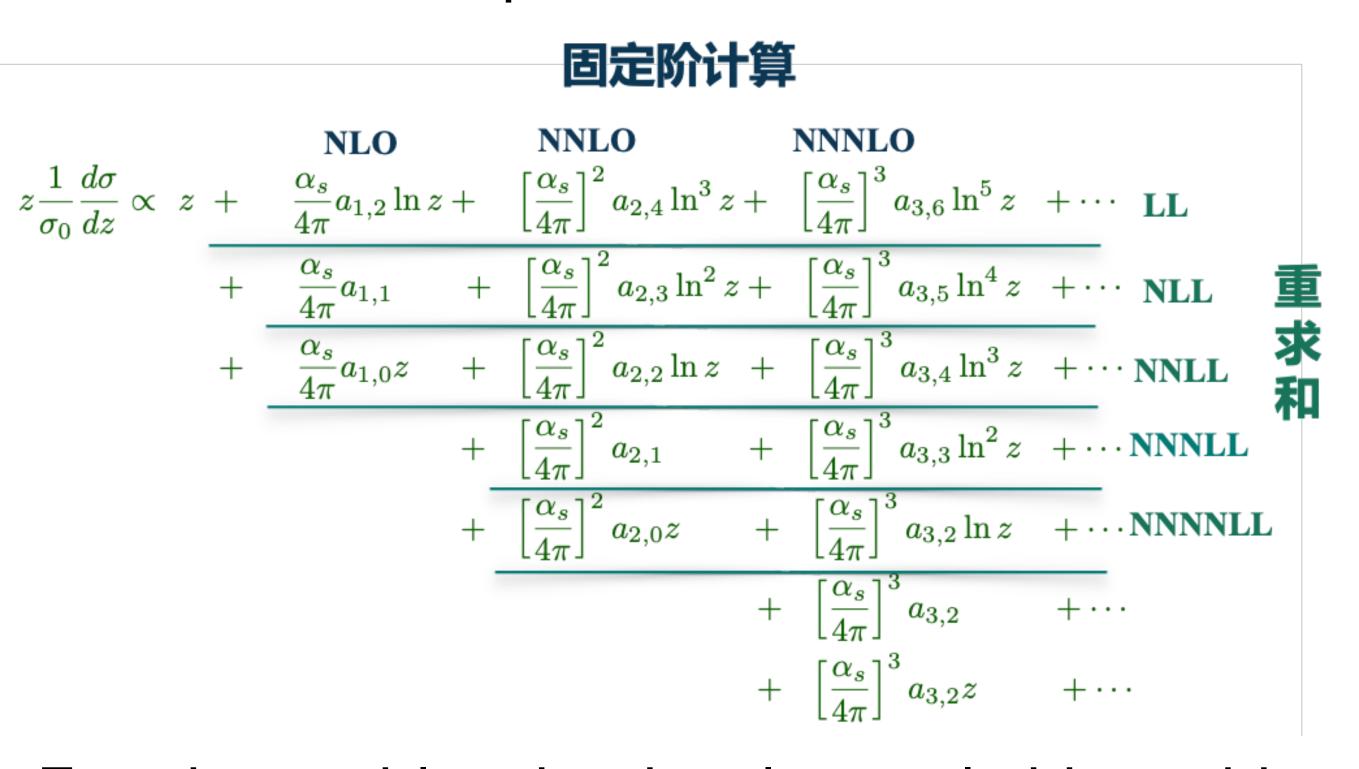
Solve  $R = \Delta$  for Q1/Q2

$$Q_2/Q_1$$
 distribution generated by 
$$\frac{d}{dq^2} \Delta \left(Q^2, q^2\right) = \Delta \left(Q^2, q^2\right) \times d\phi_{n \to n+1} \frac{\left|M_{n+1}\right|^2}{\left|M_n\right|^2}$$



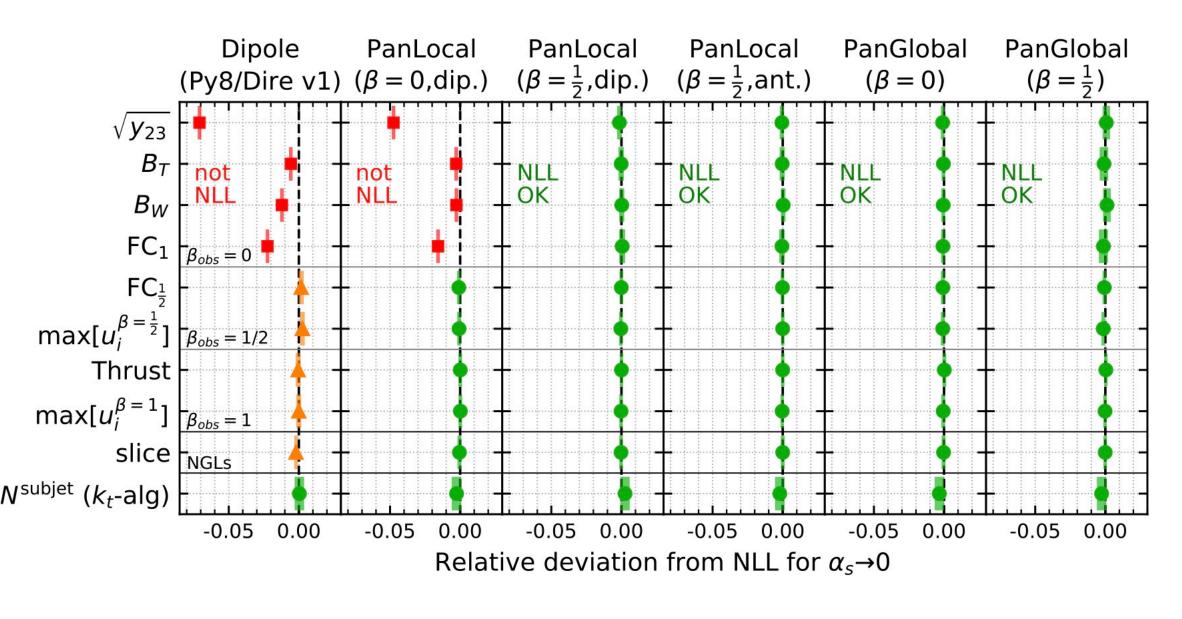
new phase space point generated according to the new scales

### For multi-scale problem



NLL: PanScales, Alaric, Herwig et al with higher order effects: Vincia, DIRE et al

### PanScales:



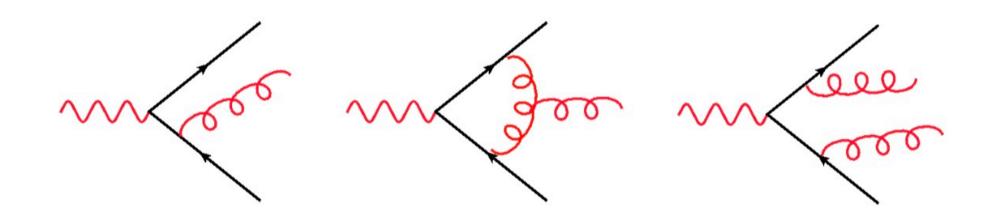
For observables that involve scale hierarchies resummation is required

arXiv:2002.11114

### To which order can Parton Showers do?

NLO corrections to resummation kernel

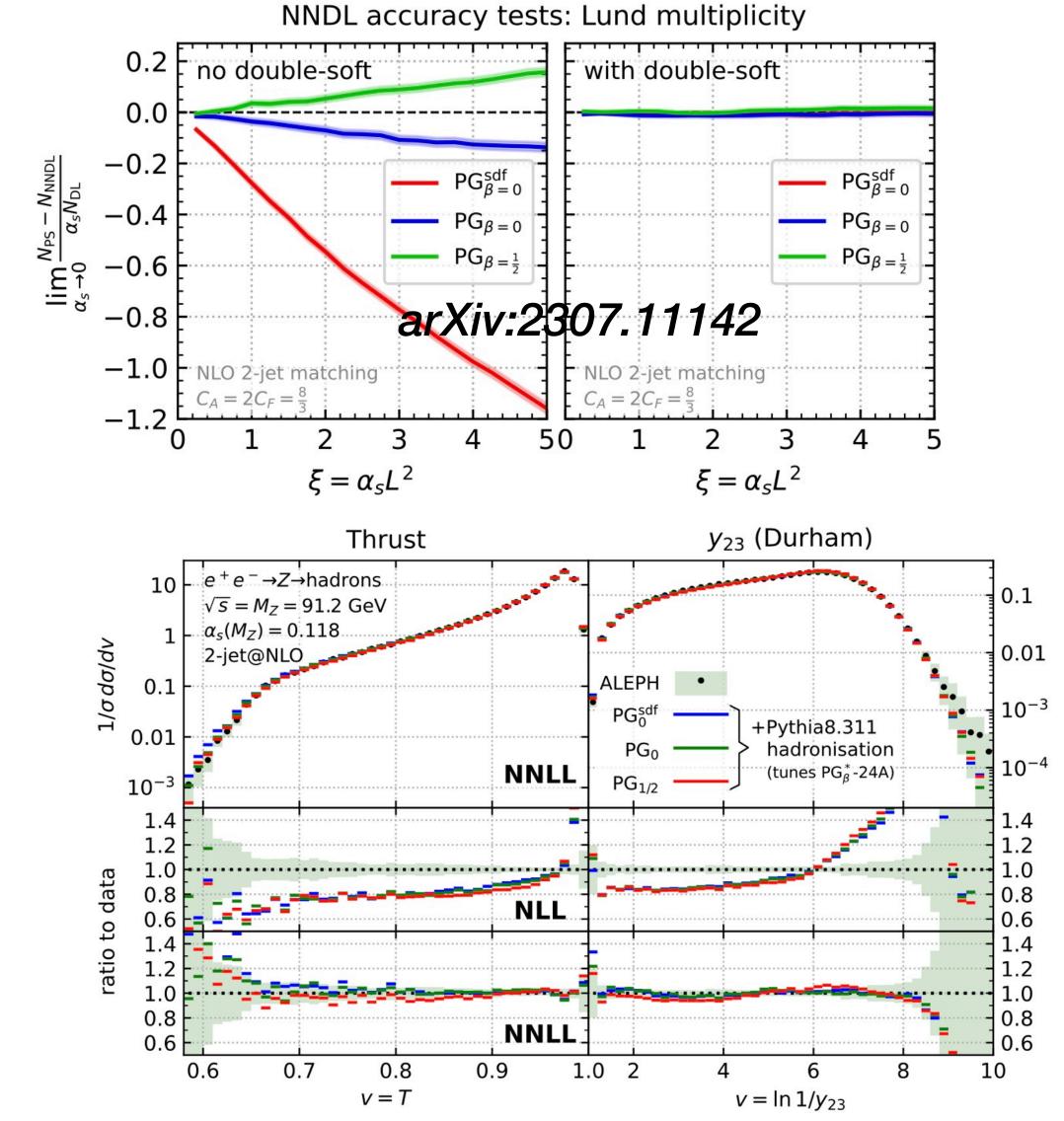
### What we expect for NLO showers



#### **NLO** parton shower

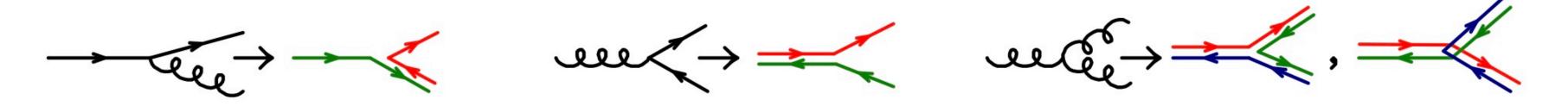
$$\frac{d}{dQ^{2}} \underbrace{\left(1-\Delta\left(Q_{0}^{2},Q^{2}\right)\right)}_{\text{branching probability}} = -\underbrace{\int \frac{d\Phi_{3}}{d\Phi_{2}} \delta\left(Q^{2}-Q^{2}\left(\Phi_{3}\right)\right)\left(a_{3}^{0}+a_{3}^{1}\right) \Delta\left(Q_{0}^{2},Q^{2}\right)}_{\text{born and virtual correction}} \\ -\underbrace{\int \frac{d\Phi_{4}}{d\Phi_{2}} \delta\left(Q^{2}-Q^{2}\left(\Phi_{4}\right)\right) a_{4}^{0} \Delta\left(Q_{0}^{2},Q^{2}\right)}_{\text{real correction}}$$

$$+TL, Skands, arXiv:1611.00013$$

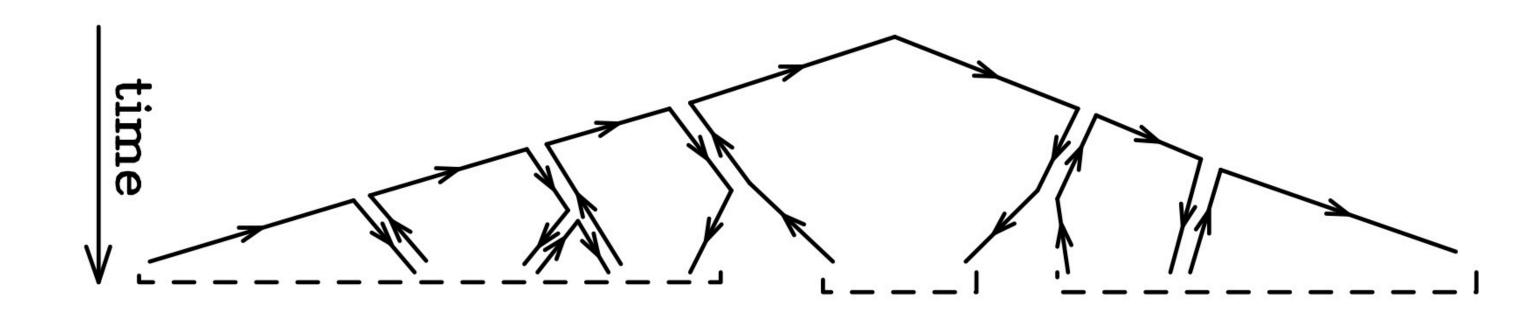


arXiv:2406.02661

Leading Color Approximation: Dipole Shower



QCD radiation in this approximation is always simulated as the radiation from a single color dipole, rather than a coherent sum from a color multipole.

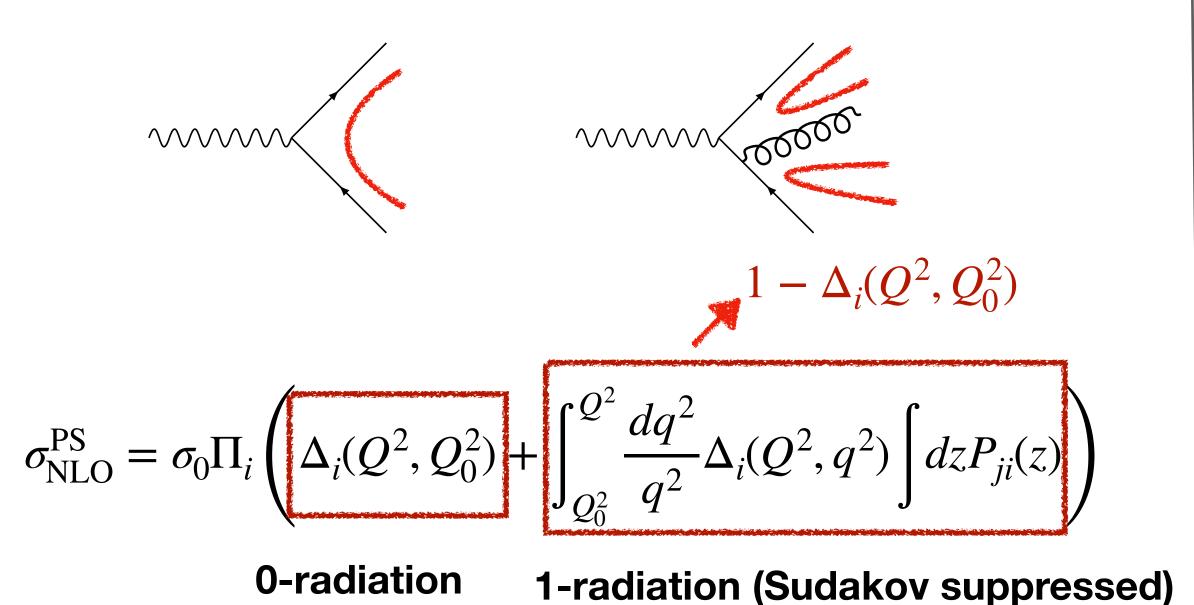


a color density operator Deductor, arXiv:1902.02105 simulates parton showers at the amplitude level with full color information CVolver, arXiv:2502.12133

$$\rightarrow \operatorname{Tr} \left( \begin{array}{c} \\ \\ \\ \end{array} \right) \qquad \mathbf{A}_n(E) = \mathbf{V}_{E,E_n} \mathbf{D}_n^{\mu} \mathbf{A}_{n-1}(E_n) \mathbf{D}_{n\mu}^{\dagger} \mathbf{V}_{E,E_n}^{\dagger} \Theta(E \leq E_n),$$

LO parton shower

### From parton shower



From the definition of Sudakov factor, we have

$$\mathcal{P}(\text{unresolved}) + \mathcal{P}(\text{resolved}) = 1$$

probability conservation from the definition of  $\Delta$ 

**Resummation from Showers +** 

#### From NLO calculations

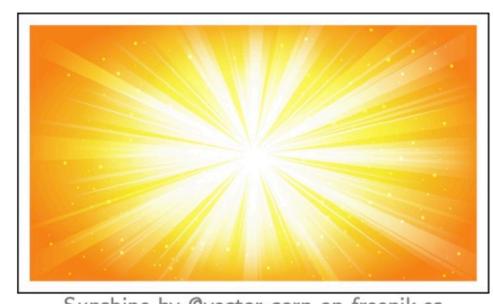
$$\sigma_{\rm NLO} = \sigma_0 + \left(\int d\Phi_n V + \int d\Phi_{n+1} S\right) \mathcal{O}_n + \int d\Phi_{n+1} (R\mathcal{O}_{n+1} - S\mathcal{O}_n)$$
 virtual integrated subtraction subtraction

$$\sigma_{\text{NLO}} = \sigma_0^n + \int_0^{t_n} d\sigma_{(1)}^n + \int_{t_n} d\sigma_{(1)}^{n+1}$$

 $t_n$  as the resolution scale for 1-radiation

LO parton showers reproduce the NLO singular behavior of the underlying hard process with unitarity assumption  $V + \mid R = 0$ .

Hard emissions From fixed orders



Sunshine by @vector\_corp on freepik.es

### Sunshine

**Sudakov Nesting of Hard Integrals** 

Using generalized parton shower to generate fixed order corrections

Fixed order should look like

$$\frac{\mathrm{d}\mathcal{P}}{\mathrm{d}\Phi_n} = |M_0|^2 \prod_{i=0}^{n-1} \mathrm{ant}_{i \mapsto i+1}$$

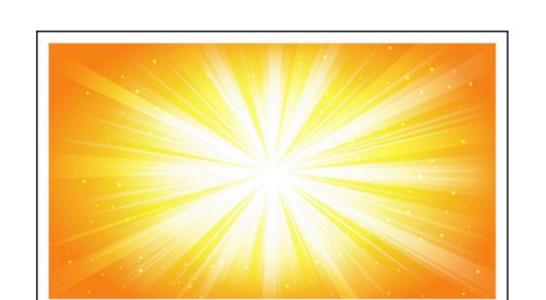
matrix element ratio

$$(0 \rightarrow 1) \times (1 \rightarrow 2) \times \cdots \times (n-1 \rightarrow n)$$

Usually showers will give  $(0 \rightarrow n)$ 

$$\frac{\mathrm{d}\mathcal{P}_{00...0}}{\mathrm{d}\Phi_n} = |M_0|^2 \prod_{i=0}^{n-1} \mathrm{ant}_{i \mapsto i+1} \Delta_i(t_i, t_{i+1})$$

Sudakov factor from showers  $\Delta_0 \times \Delta_1 \times \cdots \times \Delta_{(n-1)}$ 



Sunshine by @vector\_corp on freepik.es

### Sunshine

**Sudakov Nesting of Hard Integrals** 

Using generalized parton shower to generate fixed order corrections

keep the parent events after branching, and ask the event branches mtimes at stage  $0 \rightarrow 1$ , then shower them afterwards

$$\frac{\mathrm{d}\mathcal{P}_{m0...0}}{\mathrm{d}\Phi_n} = \frac{\mathrm{d}\mathcal{P}_{00...0}}{\mathrm{d}\Phi_n} \prod_{j=1}^m \int_{t_1}^{\tilde{t}_{j-1}} \mathrm{ant}_{0\mapsto 1}(\tilde{t}_j) \,\mathrm{d}\tilde{t}_j$$

keep all the intermediate states and shower them  $m_k$  times from k-1 partons to k partons

$$\frac{\mathrm{d}\mathcal{P}_{m_1 m_2 \dots m_n}}{\mathrm{d}\Phi_n} = \frac{\mathrm{d}\mathcal{P}_{00\dots 0}}{\mathrm{d}\Phi_n} \prod_{k=1}^n \prod_{j=1}^{m_k} \int_{t_k}^{\tilde{t}_{k_{j-1}}} \mathrm{ant}_{k-1 \mapsto k}(\tilde{t}_{k_j}) \, \mathrm{d}\tilde{t}_{k_j}$$

sum 
$$m_k$$
 to infinity 
$$\sum_{m_k\geq 0} \frac{\mathrm{d}\mathcal{P}_{m_1m_2...m_n}}{\mathrm{d}\Phi_n} \ = \ \frac{\mathrm{d}\mathcal{P}_{00...0}}{\mathrm{d}\Phi_n} \prod_{k=1}^n \frac{1}{\Delta_k(t_{k-1},t_k)}$$

SUNSHINE: 
$$\sum_{m_k \ge 0} \frac{d\mathcal{P}_{m_1 m_2 ... m_n}}{d\Phi_n} = |M_0|^2 \prod_{i=0}^{n-1} \operatorname{ant}_{i \mapsto i+1}$$

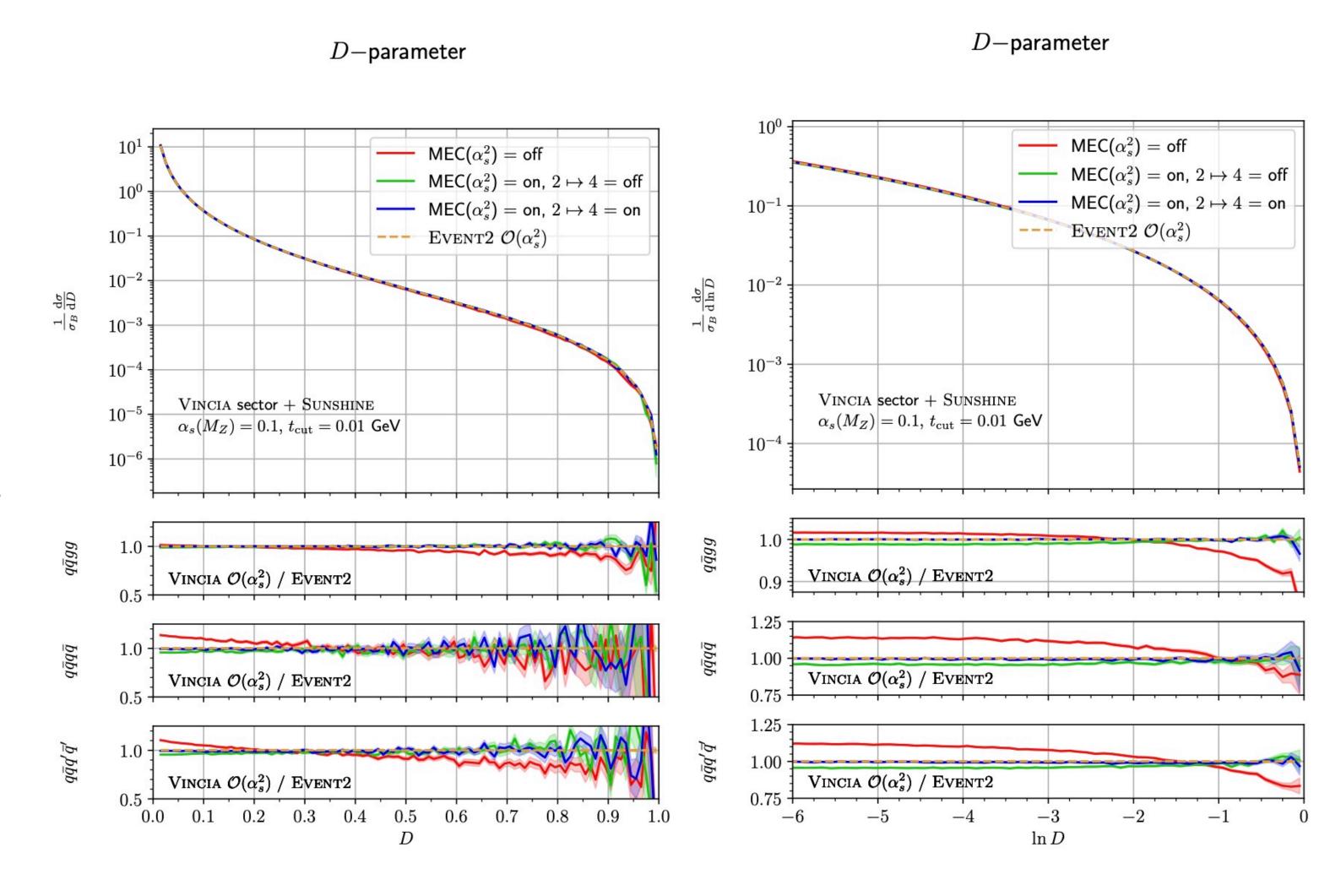
Altmann, HTL, Scyboz, Skands, arXiv:2507.00111



### Sunshine

**Sudakov Nesting of Hard Integrals** 

Using generalized parton shower to generate fixed order corrections



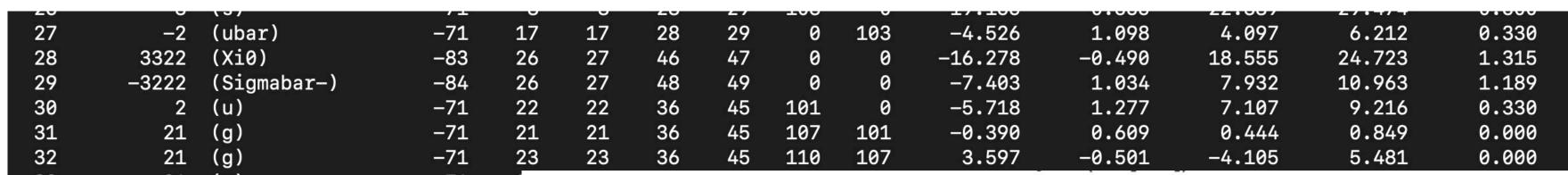
Altmann, HTL, Scyboz, Skands, 2025

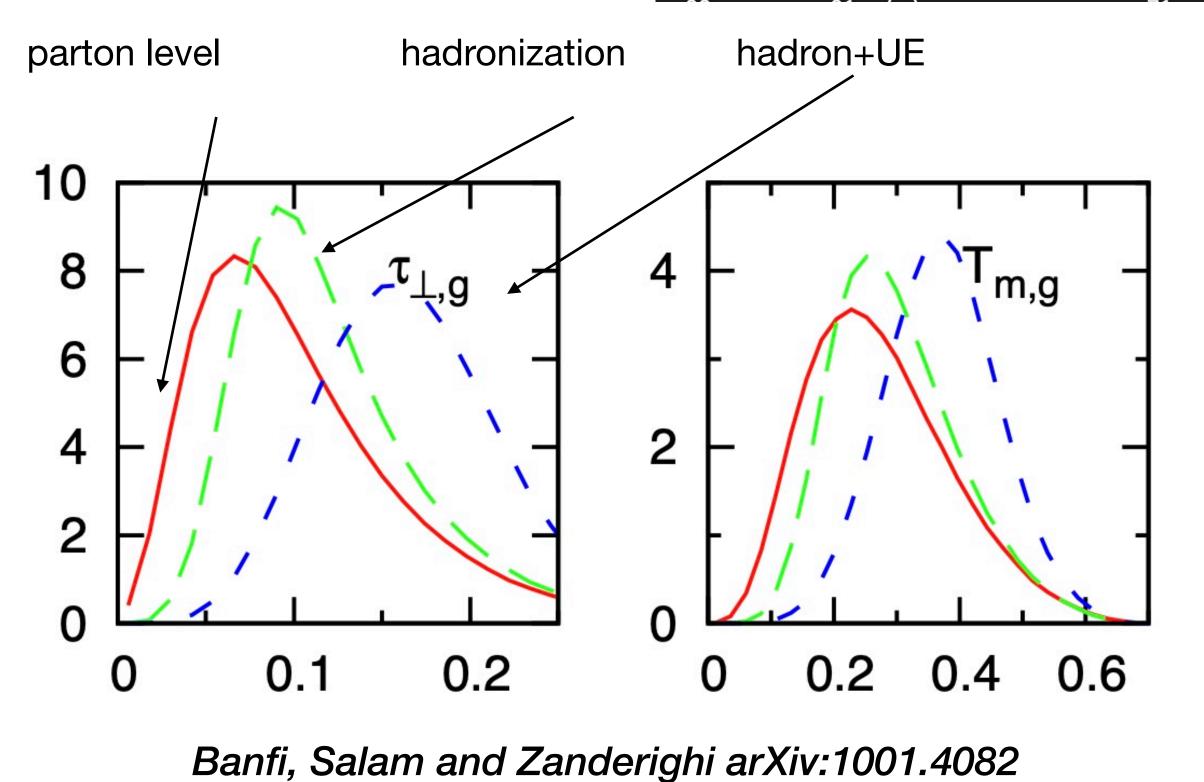
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hadronization effects

20		(3)	, _							1/1100	0.000	<i>LL</i> • 00 /	<i>4</i> /• <b>7</b> / <b>7</b>	0.000
27	-2	(ubar)	-71	17	17	28	29	0	103	-4.526	1.098	4.097	6.212	0.330
28	3322	(Xi0)	-83	26	27	46	47	0	0	-16.278	-0.490	18.555	24.723	1.315
29	-3222	(Sigmabar-)	-84	26	27	48	49	0	0	-7.403	1.034	7.932	10.963	1.189
30	2	(u)	-71	22	22	36	45	101	0	-5.718	1.277	7.107	9.216	0.330
31	21	(g)	-71	21	21	36	45	107	101	-0.390	0.609	0.444	0.849	0.000
32	21	(g)	-71	23	23	36	45	110	107	3.597	-0.501	-4.105	5.481	0.000
33	21	(g)	-71	24	24	36	45	106	110	1.334	0.455	-1.320	1.932	0.000
34	21	(g)	-71	25	25	36	45	105	106	7.964	-0.514	-7.933	11.253	0.000
35	-3	(sbar)	-71	11	11	36	45	0	105	16.893	-1.870	-20.679	26.772	0.500
36	111	(pi0)	-83	30	35	50	51	0	0	-3.511	0.738	4.002	5.377	0.135
37	211	pi+	83	30	35	0	0	0	0	0.002	0.218	0.085	0.273	0.140
38	-211	pi-	83	30	35	0	0	0	0	-1.767	-0.071	2.475	3.045	0.140
39	211	pi+	83	30	35	0	0	0	0	-0.182	0.285	0.651	0.747	0.140
40	-211	pi-	83	30	35	0	0	0	0	0.016	0.232	0.209	0.342	0.140
41	211	pi+	83	30	35	0	0	0	0	-0.413	0.450	-0.145	0.643	0.140
42	-211	pi-	84	30	35	0	0	0	0	2.478	-0.473	-2.622	3.642	0.140
43	2212	p+	84	30	35	0	0	0	0	6.374	-0.009	-6.640	9.252	0.938
44	111	(pi0)	-84	30	35	52	53	0	0	0.270	0.111	-0.364	0.486	0.135
45	-3122	(Lambdabar0)	-84	30	35	54	55	0	0	20.414	-2.024	-24.136	31.696	1.116
46	3122	(Lambda0)	-91	28	0	56	57	0	0	-14.222	-0.534	16.090	21.510	1.116
47	111	(pi0)	-91	28	0	58	59	0	0	-2.056	0.043	2.465	3.213	0.135
48	-2112	nbar0	91	29	0	0	0	0	0	-5.613	0.671	6.203	8.445	0.940
49	-211	pi-	91	29	0	0	0	0	0	-1.790	0.363	1.728	2.518	0.140
50	22	gamma	91	36	0	0	0	0	0	-3.222	0.667	3.613	4.887	0.000
51	22	gamma	91	36	0	0	0	0	0	-0.289	0.071	0.388	0.490	0.000
52	22	gamma	91	44	0	0	0	0	0	0.028	-0.020	-0.008	0.036	0.000
53	22	gamma	91	44	0	0	0	0	0	0.242	0.131	-0.356	0.450	0.000
54	-2212	pbar-	91	45	0	0	0	0	0	18.123	-1.732	-21.512	28.198	0.938
55	211	pi+	91	45	0	0	0	0	0	2.291	-0.292	-2.624	3.498	0.140
56	2212	p+	91	46	0	0	0	0	0	-10.893	-0.393	12.398	16.535	0.938
57	-211	pi-	91	46	0	0	0	0	0	-3.329	-0.140	3.692	4.975	0.140
58	22	gamma	91	47	0	0	0	0	0	-0.678	0.003	0.911	1.136	0.000
59	22	- <del></del>	91	47	0	0	0	0	0	-1.378	0.040	1.554	2.077	0.000
1 121 121 121 1	31 131 131 131 131 131	. 8%	21 Marie 121 121 121 121		1 2 185						1 1 1 1 1 1 22 1 1 14 88 15 1 1 1	THE RESERVE OF THE PARTY OF THE	1 1 1 10 1 1 1 100 101 1 1 1 1	the state of the s

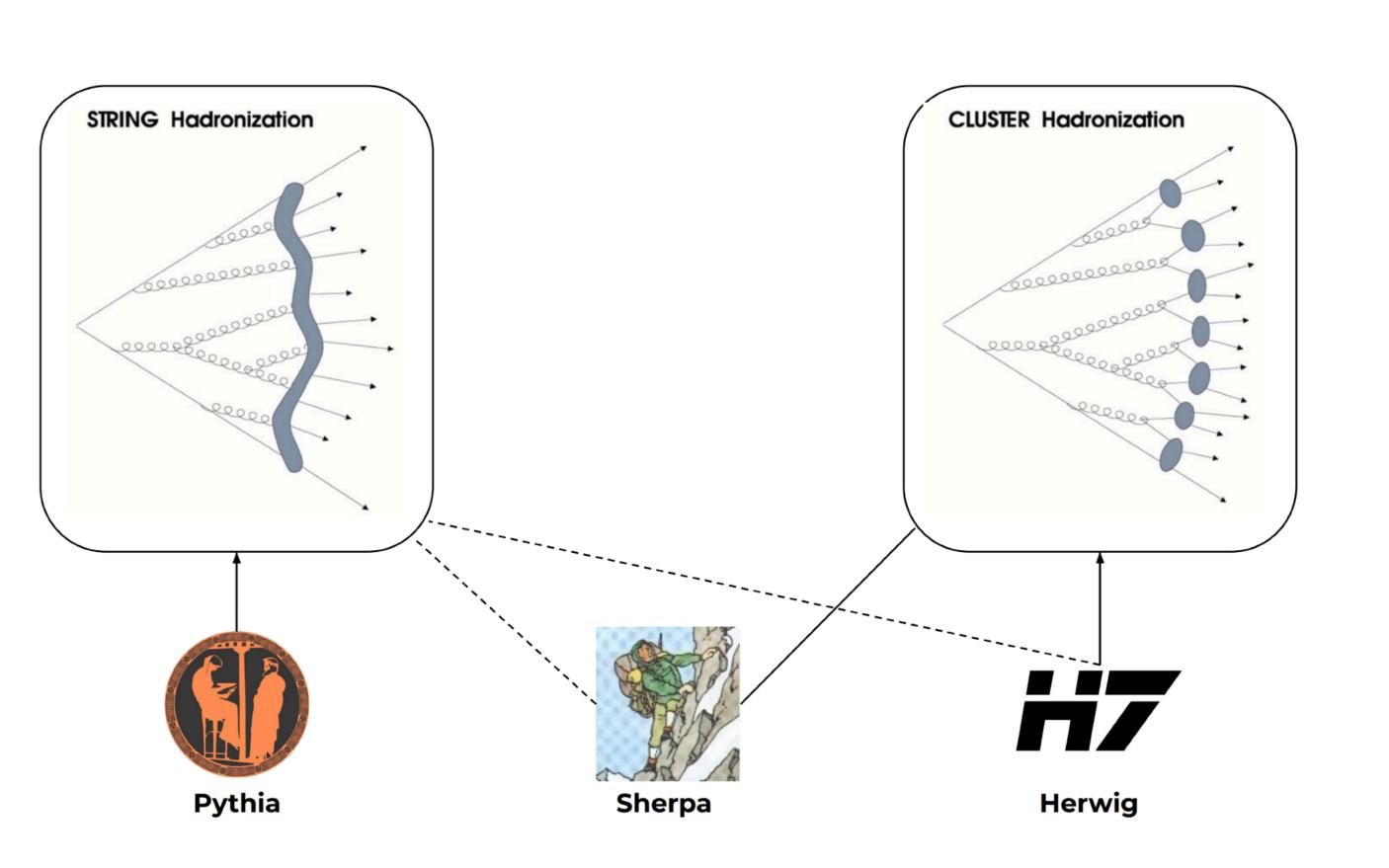


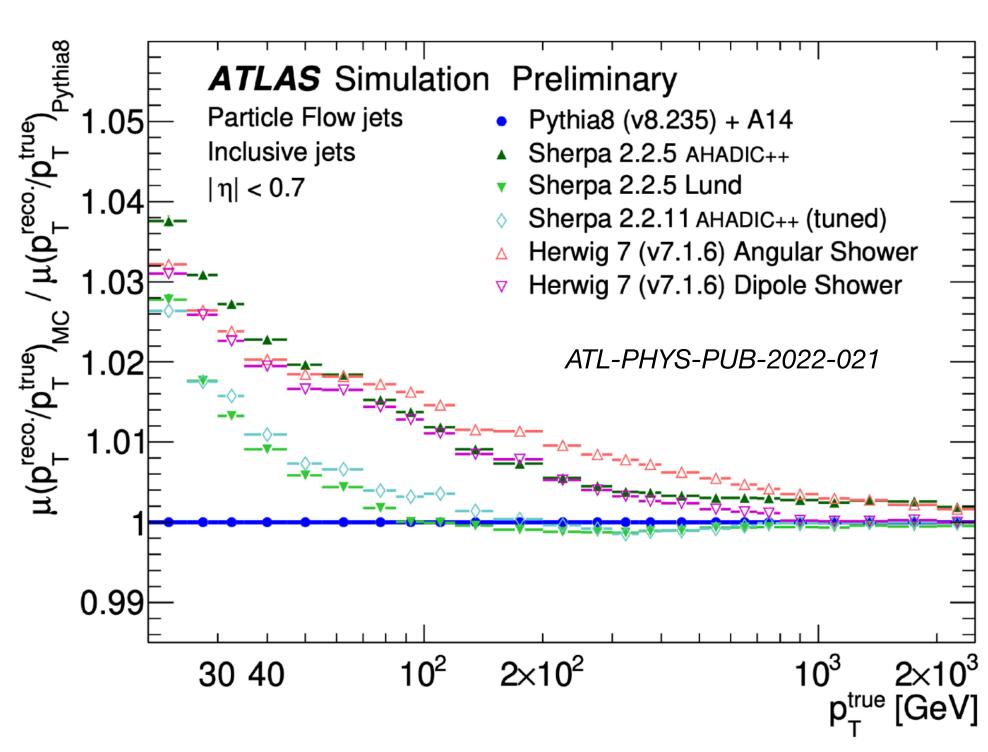




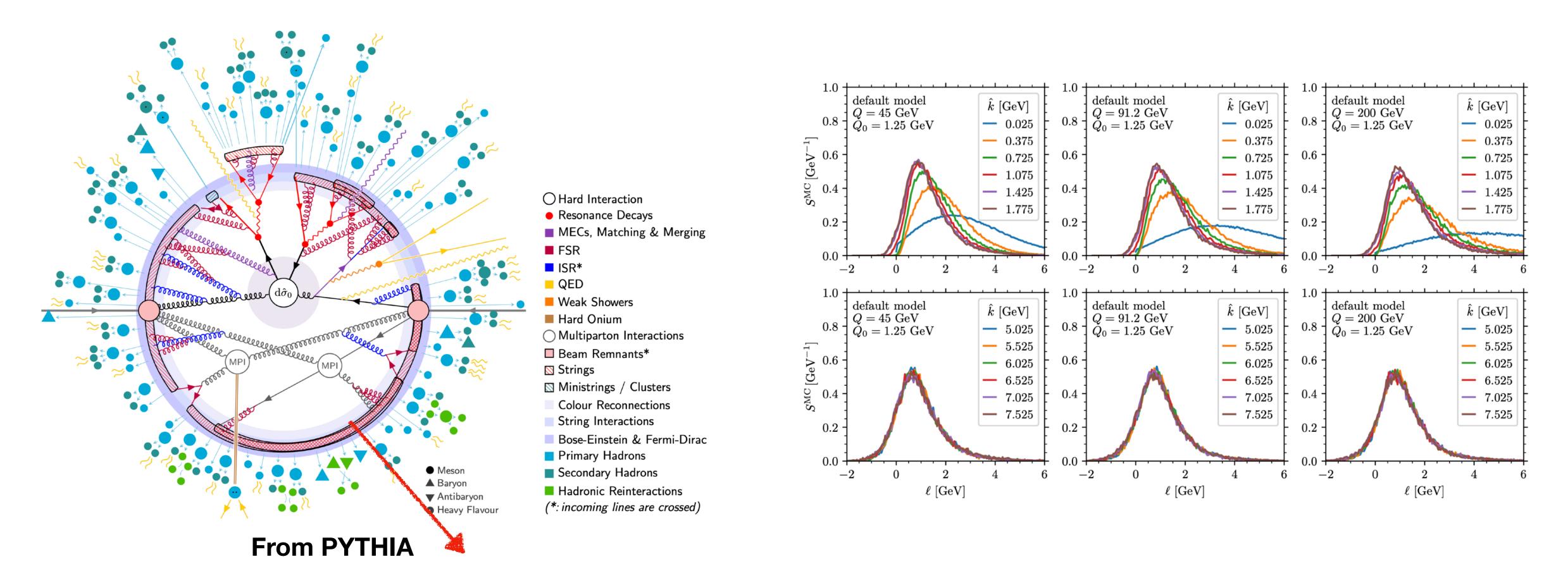
TFR resum free hadron gas transition hard 10<sup>1</sup>  $\Sigma_{N}/dy(pb)$ CEBAF: e(22GeV)+p(2GeV), Q=3GeV pythia free particle pythia had. free hadron 0  $y=ln(tan[\theta/2])$ 

Cao, HTL, Mi, arXiv:2312.07655



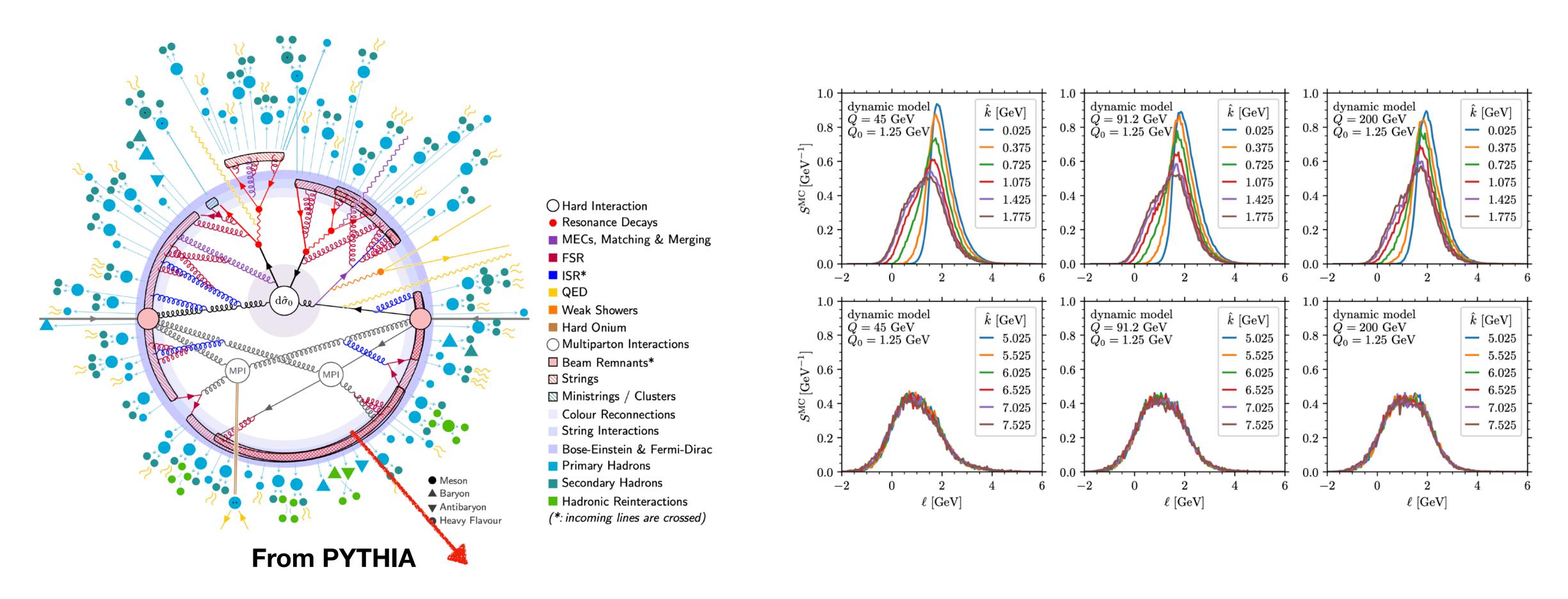


differences between different models and tunes



Physics should be independent on the transition scales

Matching the evolution of the perturbative evolution with hadronization arXiv:2404.09856



Physics should be independent on the transition scales

Matching the evolution of the perturbative evolution with hadronization arXiv:2404.09856

# 4. Summary

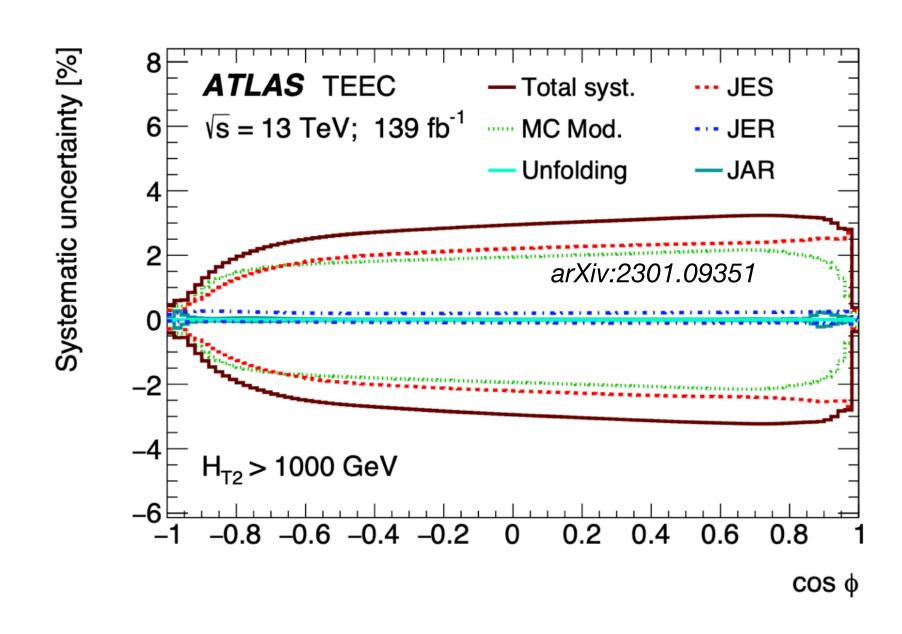
MCEGs are essential computational tools for experimentalists and theorists

Starting from hard processes to generate the perturbative and nonperturbative QCD radiations

Recently, a lot progresses on improving the logarithmic resummation order of Parton Showers

Also, subleading color effects are discussed

Hadronization model, multiple parton interactions (MPI), and underlying event descriptions introduce uncertainties



# 4. Summary

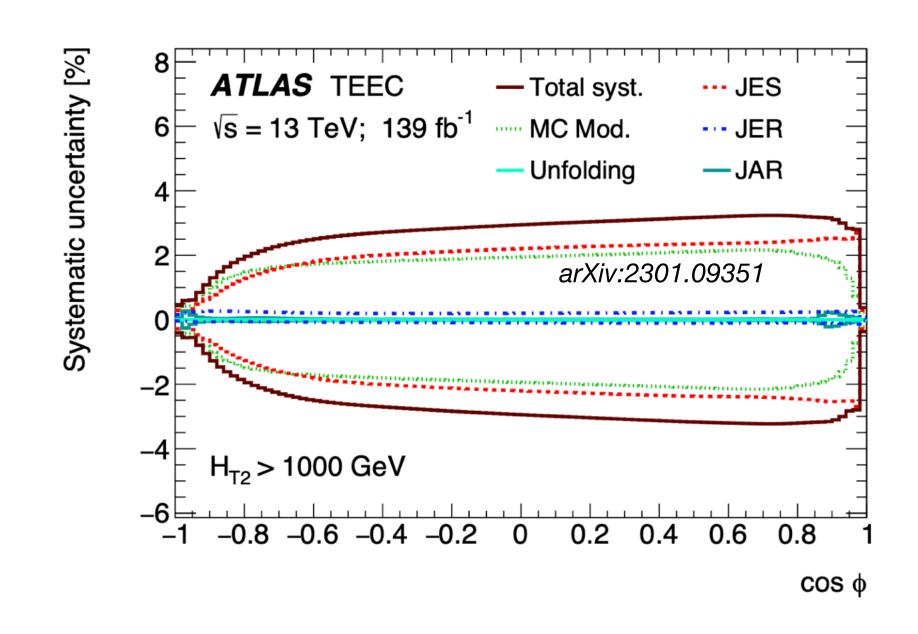
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# Thank you!