



Modular TM_1 mixing in light of precision measurement in JUNO

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Based on (2511.16348 with Wen-Hao Jiang and Ye-Ling Zhou)
NuPhyR 2025, Zhuhai, China



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1. Introduction



The origin of lepton masses and mixings remains an open question in neutrino theory (See also Talk by Yu-Feng Li): three mixing angles and two mass differences are observed, leaving the CP phase and mass ordering to be determined. In particular, **JUNO** just released their first measurement for normal ordering (NO) scenario:

$$\sin^2 \theta_{12} = 0.3092 \pm 0.0087, \quad \Delta m_{21}^2 = (7.50 \pm 0.12) \times 10^{-5} \text{ eV}^2.$$

$$U_{\text{PMNS}} = \begin{pmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta} \\ -s_{12}c_{23} - c_{12}s_{13}s_{23}e^{i\delta} & c_{12}c_{23} - s_{12}s_{13}s_{23}e^{i\delta} & c_{13}s_{23} \\ s_{12}s_{23} - c_{12}s_{13}c_{23}e^{i\delta} & -c_{12}s_{23} - s_{12}s_{13}c_{23}e^{i\delta} & c_{13}c_{23} \end{pmatrix} \begin{pmatrix} 1 & 0 & 0 \\ 0 & e^{i\frac{\alpha_{21}}{2}} & 0 \\ 0 & 0 & e^{i\frac{\alpha_{31}}{2}} \end{pmatrix}.$$



The origin of lepton masses and mixings remains an open question in neutrino theory (See also Talk by Yu-Feng Li): three mixing angles and two mass differences are observed, leaving the CP phase and mass ordering to be determined. In particular, **JUNO** just released their first measurement for normal ordering (NO) scenario:

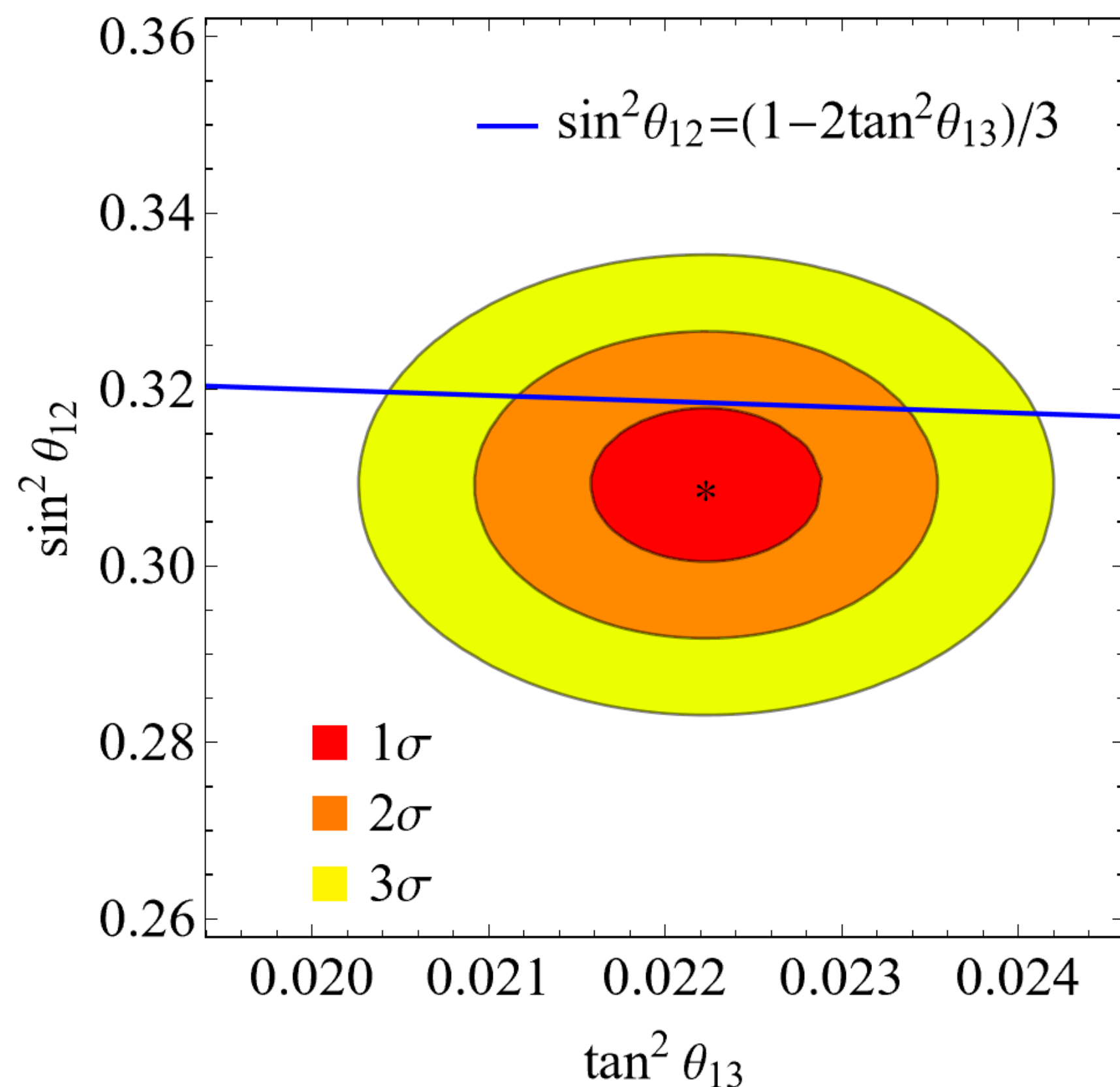
$$\sin^2 \theta_{12} = 0.3092 \pm 0.0087, \quad \Delta m_{21}^2 = (7.50 \pm 0.12) \times 10^{-5} \text{ eV}^2.$$

The key question is: can we use symmetry to explain these parameters?

- To explain the observed neutrino oscillation parameters, various of simple mixing ansatzes for the lepton flavor mixing were proposed, such as Tri-bimaximal mixing (**excluded**, Harrison *et al.* '02; Xing, '02) and Trimaximal mixing (TM₁) proposed by (Xing & Zhou, '06) and further discussed in (Lam, '06; Albright & Rodejohann, '08).
- The latest studies show that **TM₁ is most favored**, while the TM₂ is disfavored (Zhang, '25; He, '25).



TM₁ mixing pattern



TM₁ mixing (blue) and JUNO data
(S. Zhou) https://ihep.cas.cn/xwdt2022/gnxw/hotnews/2025/202511/t20251125_8016598.html

- The TM₁ is a partially constant mixing which inherits the first column the TBM form:

$$U_{\text{TM}_1} = \begin{pmatrix} \frac{2}{\sqrt{6}} & \times & \times \\ \frac{-1}{\sqrt{6}} & \times & \times \\ \frac{-1}{\sqrt{6}} & \times & \times \end{pmatrix}.$$

- It predicts **the sum rule**: $\sin^2 \theta_{12} = 1 - \frac{2}{3(1 - \sin^2 \theta_{13})}$,
 $\theta_{23} \approx 45^\circ + \sqrt{2} \theta_{13} \cos \delta$.
- It was showed that the spontaneous breaking of a non-Abelian discrete **S₄ flavor symmetry** can **lead to TM₁** (de Medeiros Varzielas & Lavoura, '13; Luhn, '13).
- However, to realize the desired symmetry breaking pattern, many flavons with particular vev alignments are introduced.



- The modular symmetry approach (See also Talk by Gui-Jun Ding) provides a new way to realize the TM_1 without introducing many flavons.
- The question we ask: how to realize TM_1 from S_4 symmetry with the least number of free parameters?
Can these models well-fit the latest JUNO data?

Therefore, we investigate the landscape of models based on modular S_4 symmetry that predicts the TM_1 mixing pattern, and explores their parameter spaces with constraints from the latest high-precision measurement on θ_{12} and Δm_{21}^2 given by JUNO.



2. Realizing TM_1 mixing from S_4 flavor symmetry



- Let's assume the lepton sector is invariant under a flavor symmetry G_f at a high energy scale, which is broken spontaneously by either the VEV of a flavon or by the VEV of a modular field. We can consider the **different residual flavor symmetries** G_ℓ and G_ν as subgroups of G_f in the **lepton sector** and **neutrino sector**, respectively.
- After the spontaneous breaking of G_f , the effective Lagrangian at a low energy scale describing neutrino masses is

$$-\mathcal{L}_{\text{mass}} = \overline{\ell}_L M_\ell \ell_R + \frac{1}{2} \overline{\nu}_L M_\nu \nu_L^c + \text{h.c.},$$

which transforms under G_ℓ and G_ν as:

$$\begin{cases} \ell_L \rightarrow \rho_L(g_\ell) \ell_L \\ \ell_R \rightarrow \rho_R(g_\ell) \ell_R \end{cases} \text{ under } g_\ell \longrightarrow \overline{\ell}_L M_\ell \ell_R = \overline{\ell}_L \left(\rho_L^\dagger(g_\ell) M_\ell \rho_R(g_\ell) \right) \ell_R$$

$$\nu_L \rightarrow \rho_L(g_\nu) \nu_L \text{ under } g_\nu \longrightarrow \overline{\nu}_L M_\nu \nu_L^c = \overline{\nu}_L \left(\rho_L^\dagger(g_\nu) M_\nu \rho_L^*(g_\nu) \right) \nu_L^c$$



Charged lepton sector G_ℓ

- In the charged lepton sector, diagonalizing the Hermitian matrix $H_\ell \equiv M_\ell M_\ell^\dagger$ introduces an unitary matrix U_ℓ satisfying $U_\ell^\dagger H_\ell U_\ell = \hat{H}_\ell$. Applying the g_ℓ transformation $\rho_L^\dagger(g_\ell) H_\ell \rho_L(g_\ell) = H_\ell$ implies:

$$[U_\ell^\dagger \rho_L(g_\ell) U_\ell]^\dagger \hat{H}_\ell [U_\ell^\dagger \rho_L(g_\ell) U_\ell] = \hat{H}_\ell.$$

- To recover non-degenerate charged lepton masses, $U_\ell^\dagger \rho_L(g_\ell) U_\ell$ must be diagonal $U_\ell^\dagger \rho_L(g_\ell) U_\ell = \text{diag}\{e^{i\alpha_1}, e^{i\alpha_2}, e^{i\alpha_3}\}$. Therefore, G_ℓ must be an **Abelian symmetry** and values of phases α_i depend on the selected residual symmetry.
- If the flavor symmetry is a S_4 , it is convenient to choose the residual symmetry G_ℓ to be Z_3^T so that a diagonal U_ℓ is guaranteed. (Jiang, RO, Zhou '25)



- Similarly, in the neutrino sector, a unitary U_ν diagonalizing the mass matrix $U_\nu^\dagger M_\nu U_\nu^* = \hat{M}_\nu$ implies

$$[U_\nu^\dagger \rho_L(g_\nu) U_\nu]^\dagger \hat{M}_\nu [U_\nu^\dagger \rho_L(g_\nu) U_\nu]^* = \hat{M}_\nu$$

- The condition requires $[U_\nu^\dagger \rho_L(g_\nu) U_\nu]$ to be real and diagonal, and thus the only choices could be

$$[U_\nu^\dagger \rho_L(g_\nu) U_\nu] = \text{diag}\{(-1)^{k_1}, (-1)^{k_2}, (-1)^{k_3}\},$$

where $k_{1,2,3} = 0, 1$.

- This conclusion leaves only Z_2 or $Z_2 \times Z_2'$ allowed as residual symmetry G_ν in the Majorana neutrino sector.
- Therefore, there is a partial dependence of U_ℓ and U_ν on $\rho_L(g_\ell)$ and $\rho_L(g_\nu)$ which eventually restricts $U_{\text{PMNS}} = U_\ell^\dagger U_\nu$.
- In the flavor basis $\rho_L(g_\ell)$ (and also U_ℓ) is diagonal, so U_{PMNS} appears to be the unitary matrix diagonalizing $\rho_L(g_\nu)$:

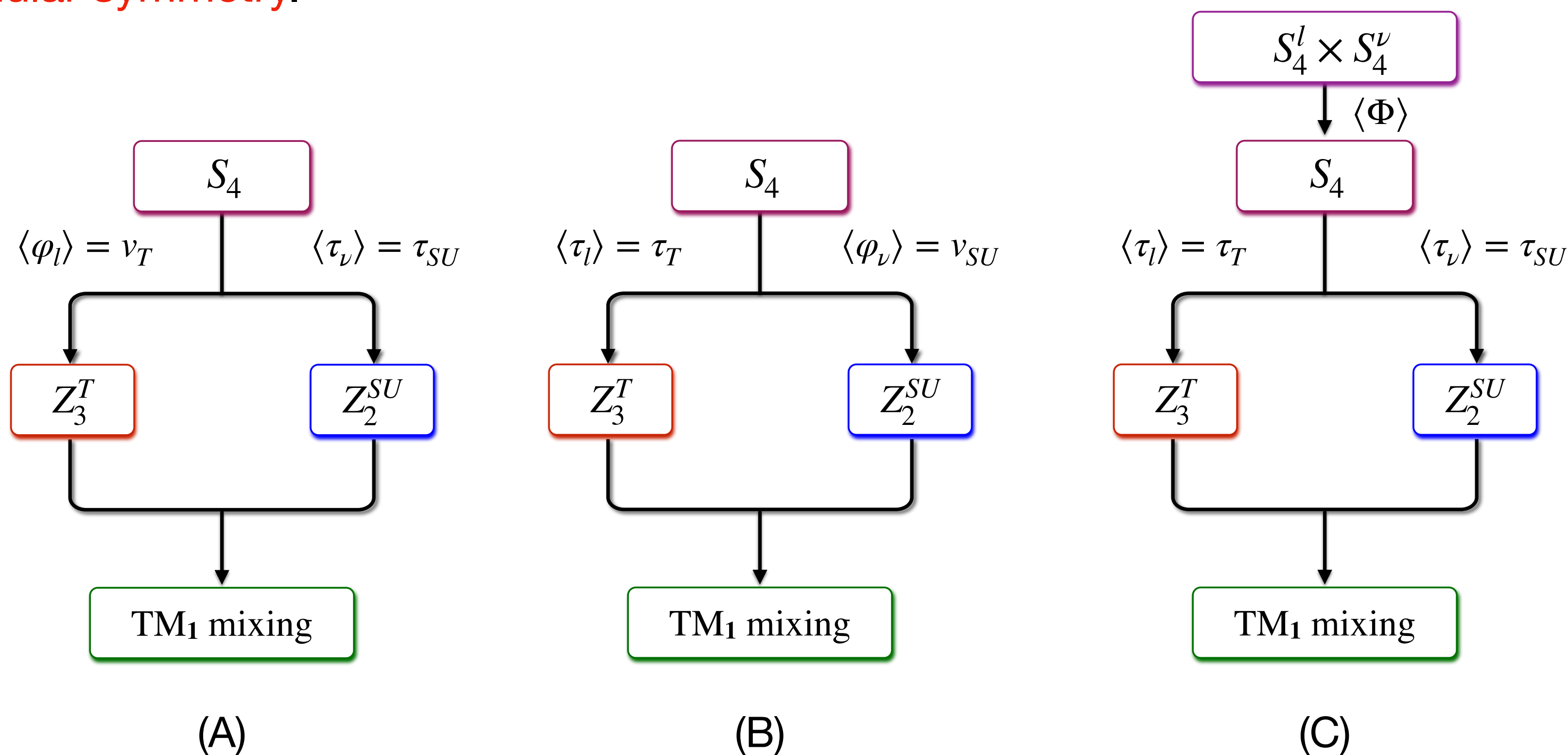
$$U_{\text{PMNS}}^\dagger \rho_L(g_\nu) U_{\text{PMNS}} = \text{diag}\{(-1)^{k_1}, (-1)^{k_2}, (-1)^{k_3}\}.$$

- Different choices of G_ν clearly results in different mixing patterns. For S_4 symmetry, if we take $G_\nu = Z_2^{SU}$, we can find the representations $\rho_{3,3}(SU)$ and its' eigenvector $(2, -1, -1)^T$ for the non-degenerate eigenvalues. (Jiang, RO, Zhou '25)



Realizing TM_1 mixing from S_4

- Therefore, the prediction of TM_1 mixing pattern is thus obtained via the structure of residual symmetries assigned as: $G_\ell = Z_3^T$ and $G_\nu = Z_2^{SU}$, which is **independent of model details** and **regardless of realizations in traditional flavor symmetry or modular symmetry**.



(Jiang, RO, Zhou '25)



3. Models based on modular S_4 symmetry



- Modular invariance can be used for lepton model building (Feruglio, '17), where the finite modular group $\Gamma(N)$ can be used which is isomorphic to a permutation group at the first few levels, i.e. $\Gamma_2 \simeq S_3$, $\Gamma_3 \simeq A_4$, $\Gamma_4 \simeq S_4$ for $N = 2, 3, 4$. (See also Talk by Gui-Jun Ding)

- Under the modular invariance Γ_N , a chiral superfield $\phi_i(\tau)$ transforms non-linearly as function of τ as

$$\phi_i(\tau) \rightarrow \phi_i(\gamma\tau) = (c\tau + d)^{-2k_i} \rho_{I_i}(\gamma) \phi_i(\tau)$$

- Due to the holomorphicity of the superpotential, the Yukawa coupling transforms as (where k_Y is non-negative):

$$Y_{I_Y}(\tau) \rightarrow Y_{I_Y}(\gamma\tau) = (c\tau + d)^{2k_Y} \rho_{I_Y}(\gamma) Y_{I_Y}(\tau).$$

- The modulus τ can be stabilized at a fixed point ($\gamma\tau_\gamma = \tau_\gamma$) such that $\langle \tau \rangle = \tau_\gamma$. In this case, an Abelian residual symmetry generated by γ is preserved, leaving $Y_I(\gamma\tau_\gamma) = Y_I(\tau_\gamma)$:

$$\rho_I(\gamma) Y_I(\tau_\gamma) = (c\tau + d)^{-2k} Y_I(\tau_\gamma).$$

- Given a representation matrix of S_4 and a fixed point τ_γ , we can determine the **eigenvectors** for $Y_I(\tau_\gamma)$.



For example, at the fixed point $\tau_\gamma = \tau_T = (-1 + i\sqrt{3})/2$, the modular forms are: (de Medeiros Varzielas, King, Zhou, '19)

$$Y_{3^{(+)}}^{(6j+2)}(\tau_T) \propto \begin{pmatrix} 0 \\ 1 \\ 0 \end{pmatrix}, \quad Y_{3^{(+)}}^{(6j+4)}(\tau_T) \propto \begin{pmatrix} 0 \\ 0 \\ 1 \end{pmatrix}, \quad Y_{3^{(+)}}^{(6j+6)}(\tau_T) \propto \begin{pmatrix} 1 \\ 0 \\ 0 \end{pmatrix},$$

- The advantage of using modular S_4 symmetry: **it reduces physical degrees of freedom very efficiently.**
- For example, in the minimal setup (Penedo & Petcov, '18; Novichkov, Penedo, Petcov, Titov, '18), it only needs one modular field, which has only two real degrees of freedom, acquiring a VEV, **without the necessity of introducing many flavon fields aligned in some specific ways** (e.g. King, Luhn, '13)



Fermions	S_4	$2k$	Scalars	S_4	$2k$	Modular forms	S_4	$2k$
e^c	1	-5	φ_ℓ	3	+2	$Y_2(\tau_\nu)$	2	+2
μ^c	1'	-2	η	2	-1	$Y_{3'}(\tau_\nu)$	3'	+2
τ^c	1	-2	ξ	1'	0			
L	3	+1	$H_{u,d}$	1	0			
ν^c	3	-1						

- The superpotential is:
$$W = \frac{y_e}{\Lambda^2} (L\varphi_l^2)_1 e^c H_d + \frac{y_\mu}{\Lambda^2} (L(\varphi_l\eta)_{3'})_{1'} \mu^c H_d + \frac{y_\tau}{\Lambda^2} (L(\varphi_l\eta)_3)_1 \tau^c H_d$$

$$+ y_D L\nu^c H_u + \frac{1}{2} Y_2(\tau_\nu) (\nu^c \nu^c)_2 \xi + \frac{1}{2} Y_3(\tau_\nu) (\nu^c \nu^c)_3 \xi.$$
- In the **lepton sector**, the flavon φ_l breaks S_4 to the residual symmetry Z_3^T by the vev $\langle \varphi_l \rangle = v_T \begin{pmatrix} 1 \\ 0 \\ 0 \end{pmatrix}$.
- In the **neutrino sector**, the modular field τ_ν stabilized at $\langle \tau_\nu \rangle = \tau_{SU} = -\frac{1}{2} + \frac{i}{2}$ breaks S_4 to the residual Z_2^{SU} symmetry.



- Expanding the superpotential gives the charged lepton mass matrix: $M_\ell = \frac{v_d v_T^*}{\Lambda^2} \begin{pmatrix} y_e v_T & 0 & 0 \\ 0 & y_\mu v_{\eta_1} & y_\tau v_{\eta_1} \\ 0 & -y_\mu v_{\eta_2} & y_\tau v_{\eta_2} \end{pmatrix}^*$,

the Dirac mass matrix in the flavor basis: $M_D = y_D^* v_u U_\ell^\dagger \begin{pmatrix} 1 & 0 & 0 \\ 0 & 0 & 1 \\ 0 & 1 & 0 \end{pmatrix}$,

and the Majorana mass matrix: $M_R = a_2 \begin{pmatrix} 0 & 1 & 1 \\ 1 & 1 & 0 \\ 1 & 0 & 1 \end{pmatrix} + a_3 \begin{pmatrix} 2 & -1 & -1 \\ -1 & 2 & -1 \\ -1 & -1 & 2 \end{pmatrix} - a_3 \sqrt{6} \begin{pmatrix} 0 & 1 & -1 \\ 1 & 2 & 0 \\ -1 & 0 & -2 \end{pmatrix}$,

where $a_2 = [\lambda_2 Y_{2,1}^{(2)}(\tau_{SU})]^*$, $a_3 = [\lambda_3 Y_{3,1}^{(2)}(\tau_{SU})]^*$ are complex coefficients.

- The standard neutrino seesaw formula $M_\nu = -M_D M_R^{-1} M_D^T$ can then be used to find the mass matrix for (active) neutrinos and the $U_{\text{PMNS}} = U_\ell^\dagger U_\nu$.



Fermions	S_4	$2k$	Scalars	S_4	$2k$	Modular forms	S_4	$2k$
e^c	$\mathbf{1}'$	-6	φ_ν	$\mathbf{3}$	0	$Y_e(\tau_l)$	$\mathbf{3}'$	+6
μ^c	$\mathbf{1}'$	-4	φ'_ν	$\mathbf{3}'$	0	$Y_\mu(\tau_l)$	$\mathbf{3}'$	+4
τ^c	$\mathbf{1}'$	-2	η	$\mathbf{1}'$	0	$Y_\tau(\tau_l)$	$\mathbf{3}'$	+2
L	$\mathbf{3}$	0	$H_{u,d}$	$\mathbf{1}$	0			
ν^c	$\mathbf{3}$	0						

- The superpotential reads:
$$W = Y_e(\tau_l) L e^c H_d + Y_\mu(\tau_l) L \mu^c H_d + Y_\tau(\tau_l) L \tau^c H_d + y_D (L \nu^c)_1 H_u + \frac{1}{2} M_1 (\nu^c \nu^c)_1 + \frac{\lambda}{2} \varphi_\nu (\nu^c \nu^c)_3 + \frac{\lambda'}{2\Lambda} \varphi'_\nu (\nu^c \nu^c \eta)_{3'}.$$
- In the **lepton sector**, the modular field τ_l stabilized at $\langle \tau_l \rangle = \tau_T$ preserves Z_3^T and diagonalizes charged leptons.
- In the **neutrino sector**, two flavons are required as one flavon only gives TBM (Luhn '13). To preserve the Z_2^{SU} with two flavons $\varphi_\nu \sim \mathbf{3}$ and $\varphi'_\nu \sim \mathbf{3}'$, their vevs must be aligned as:

$$\langle \varphi_\nu \rangle = v_{SU} \begin{pmatrix} 1 \\ 1 \\ 1 \end{pmatrix}, \quad \langle \varphi'_\nu \rangle = v'_{SU} \begin{pmatrix} 0 \\ 1 \\ -1 \end{pmatrix}.$$



- The charged lepton mass matrix $M_\ell = \begin{pmatrix} Y_e(\tau_T) & 0 & 0 \\ 0 & Y_\mu(\tau_T) & 0 \\ 0 & & Y_\tau(\tau_T) \end{pmatrix}^*$.
- The Dirac mass matrix $M_D = y_D^* v_u \begin{pmatrix} 1 & 0 & 0 \\ 0 & 0 & 1 \\ 0 & 1 & 0 \end{pmatrix}$.
- The Majorana mass matrix $M_R = b_1 \begin{pmatrix} 1 & 0 & 0 \\ 0 & 0 & 1 \\ 0 & 1 & 0 \end{pmatrix} + b_2 \begin{pmatrix} 2 & -1 & -1 \\ -1 & 2 & -1 \\ -1 & -1 & 2 \end{pmatrix} + b_3 \begin{pmatrix} 0 & 1 & -1 \\ 1 & 2 & 0 \\ -1 & 0 & -2 \end{pmatrix},$

where $b_1 = M_1^*$, $b_2 = (\lambda v_{SU})^*$ and $b_3 = (\frac{1}{\Lambda} \lambda' v'_{SU})^*$. Due to the suppression of higher-dimensional operator, b_3 is in general much smaller than b_2 .



- The superpotential is given by (King, Zhou '19):

$$W = Y_e(\tau_l) L e^c H_d + Y_\mu(\tau_l) L \mu^c H_d + Y_\tau(\tau_l) L \tau^c H_d \\ + \frac{y_\nu}{\Lambda} L \Phi H_u \nu^c + \frac{1}{2} M_1(\tau_\nu) (\nu^c \nu^c)_1 + \frac{1}{2} M_2(\tau_\nu) (\nu^c \nu^c)_2 + \frac{1}{2} M_3(\tau_\nu) (\nu^c \nu^c)_3.$$

- The model is based on multiple modular symmetry (de Medeiros Varzielas, King, Zhou, '19), where the original modular symmetry $S_4^l \times S_4^\nu$ is broken to a diagonal subgroup S_4 by the vev of a flavon $\Phi \sim (3, 3)$ assigned as

$$\langle \Phi \rangle = v_\Phi \begin{pmatrix} 1 & 0 & 0 \\ 0 & 0 & 1 \\ 0 & 1 & 0 \end{pmatrix}$$

- The subgroup S_4 is then broken to two distinct residual symmetries $G_\ell = Z_3^T$ and $G_\nu = Z_2^{SU}$, by the modular field τ_l stabilized at $\langle \tau_l \rangle = \tau_T$ in the **charged lepton sector**, and the modular field τ_ν fixed at $\langle \tau_\nu \rangle = \tau_{SU}$ in the **neutrino sector**, respectively.

Fields	S_4^l	S_4^ν	$2k_l$	$2k_\nu$
e^c	1'	1	-6	-2
μ^c	1'	1	-4	-2
τ^c	1'	1	-2	-2
L	3	1	0	+2
ν^c	1	3	0	-2
Φ	3	3	0	0
$H_{u,d}$	1	1	0	0

(King, Zhou '19)



- This scenario gives the same Yukawa matrices as model B for charged leptons and thus will not be repeated.

- The Dirac mass matrix is given by $M_D = \frac{y_\nu^* v_\Phi^* v_u}{\Lambda} \begin{pmatrix} 1 & 0 & 0 \\ 0 & 0 & 1 \\ 0 & 1 & 0 \end{pmatrix}.$

- The Majorana mass matrix is $M_R = c_1 \begin{pmatrix} 1 & 0 & 0 \\ 0 & 0 & 1 \\ 0 & 1 & 0 \end{pmatrix} + c_2 \begin{pmatrix} 0 & 1 & 1 \\ 1 & 1 & 0 \\ 1 & 0 & 1 \end{pmatrix} + c_3 \sqrt{2} \begin{pmatrix} 2 & -1 & -1 \\ -1 & 2 & -1 \\ -1 & -1 & 2 \end{pmatrix} - c_3 \sqrt{3} \begin{pmatrix} 0 & 1 & -1 \\ 1 & 2 & 0 \\ -1 & 0 & -2 \end{pmatrix},$

where $c_1 = [M_1(\tau_{SU})]^*$, $c_2 = [M_{2,1}(\tau_{SU})]^*$ and $c_3 = \frac{1}{\sqrt{2}} [M_{3,1}(\tau_{SU})]^*$.



4. Experimental predictions in light of JUNO data



- All three discussed models can be parameterized using the TBM matrix U_{TBM}

$$U_{\text{TBM}}^T M_R U_{\text{TBM}} = \begin{pmatrix} \eta & 0 & 0 \\ 0 & \alpha & \gamma \\ 0 & \gamma & \beta \end{pmatrix},$$

$$U_{\text{TBM}} = \begin{pmatrix} \frac{2}{\sqrt{6}} & \frac{1}{\sqrt{3}} & 0 \\ \frac{-1}{\sqrt{6}} & \frac{1}{\sqrt{3}} & \frac{1}{\sqrt{2}} \\ \frac{-1}{\sqrt{6}} & \frac{1}{\sqrt{3}} & \frac{-1}{\sqrt{2}} \end{pmatrix}$$

where it can be shown that

Model A : $\eta = -\alpha + \beta$, $\gamma = \alpha - 2\beta$, with $\alpha = 2a_2$, $\beta = a_2 + 3a_3$,

Model B : $\eta = 2\alpha + \beta$, with $\alpha = b_1$, $\beta = -b_1 + 3b_2$, $\gamma = \sqrt{6}b_3$,

Model C : $\eta = -\beta - 2\gamma$, with $\alpha = c_1 + 2c_2$, $\beta = -c_1 + c_2 + 3\sqrt{2}c_3$, $\gamma = -3\sqrt{2}c_3$.

- This parameterization allows us to express the analytical predictions for our models in closed forms.

- The seesaw formula gives $M_\nu = -m_0^2 U_\ell^\dagger P_{23} M_R^{-1} P_{23}^T U_\ell^*$, where $m_0^2 = (y_D^*)^2 v_u^2$.

- The M_ν can be diagonalized by the PMNS matrix $U_{\text{PMNS}} = U_\ell^\dagger U_{\text{TM}_1}$ where $U_{\text{TM}_1} \equiv U_{\text{TBM}} \begin{pmatrix} e^{-i\alpha'_3} & 0 & 0 \\ 0 & \cos \theta_R e^{i\alpha_1} & \sin \theta_R e^{-i\alpha_2} \\ 0 & -\sin \theta_R e^{i\alpha_2} & \cos \theta_R e^{-i\alpha_1} \end{pmatrix}$.



- All three models contain **five free parameters**, including two complex coefficients and a norm, which can be fit by three mixing angles and two mass differences.
- A straightforward result from the previous parameterization gives the predictions:

$$\sin \theta_{13} = \frac{\sin \theta_R}{\sqrt{3}},$$

$$\tan \theta_{12} = \frac{\cos \theta_R}{\sqrt{2}}.$$

- These correlations recover the **TM₁** sum rules between mixing angles $\sin^2 \theta_{12} = 1 - \frac{2}{3(1 - \sin^2 \theta_{13})}$!



- For model A, we get the following correlations:

$$\sin(2\theta_R) \cos(\alpha_1 - \alpha_2) = 1 + \frac{(5m_2^2 - 2m_1^2) m_3^2}{m_1^2 (m_3^2 - m_2^2)},$$
$$1 = - \left[\frac{\Delta m_{21}^2}{4m_1^2} - \frac{\Delta m_{32}^2 (2 - 5 \sin^2 \theta_R)}{8m_3^2} + \frac{1}{8} \right]^2 + \frac{m_3^4}{m_1^4} \frac{20 (\Delta m_{21}^2)^2 + 6m_1^4}{(\Delta m_{32}^2)^2 \sin^2(2\theta_R)}$$
$$+ \frac{m_3^2}{m_1^2} \frac{22m_3^2 \Delta m_{21}^2 + 2\Delta m_{32}^2 [2\Delta m_{21}^2 - m_2^2 (3 \sin^2 \theta_R - 4)]}{(\Delta m_{32}^2)^2 \sin^2(2\theta_R)}.$$

- This shows **model A is over-constrained**. Modifying to higher modular weight of τ_ν can fit the data with the price of introducing more free parameters, rendering model A less predictive and thus will not be considered.



- For model B and C, the universal predictions are (King, Zhou '19):

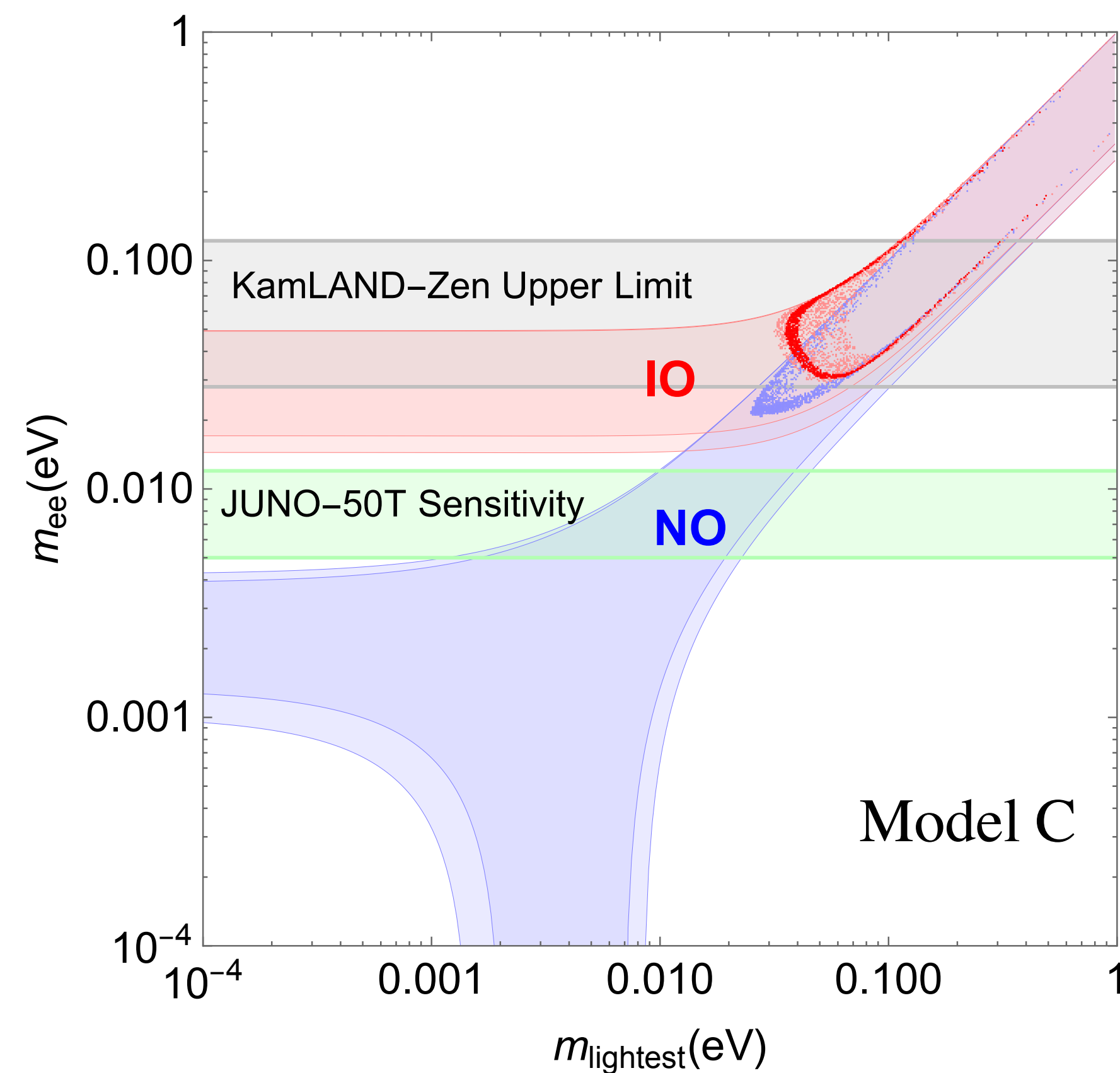
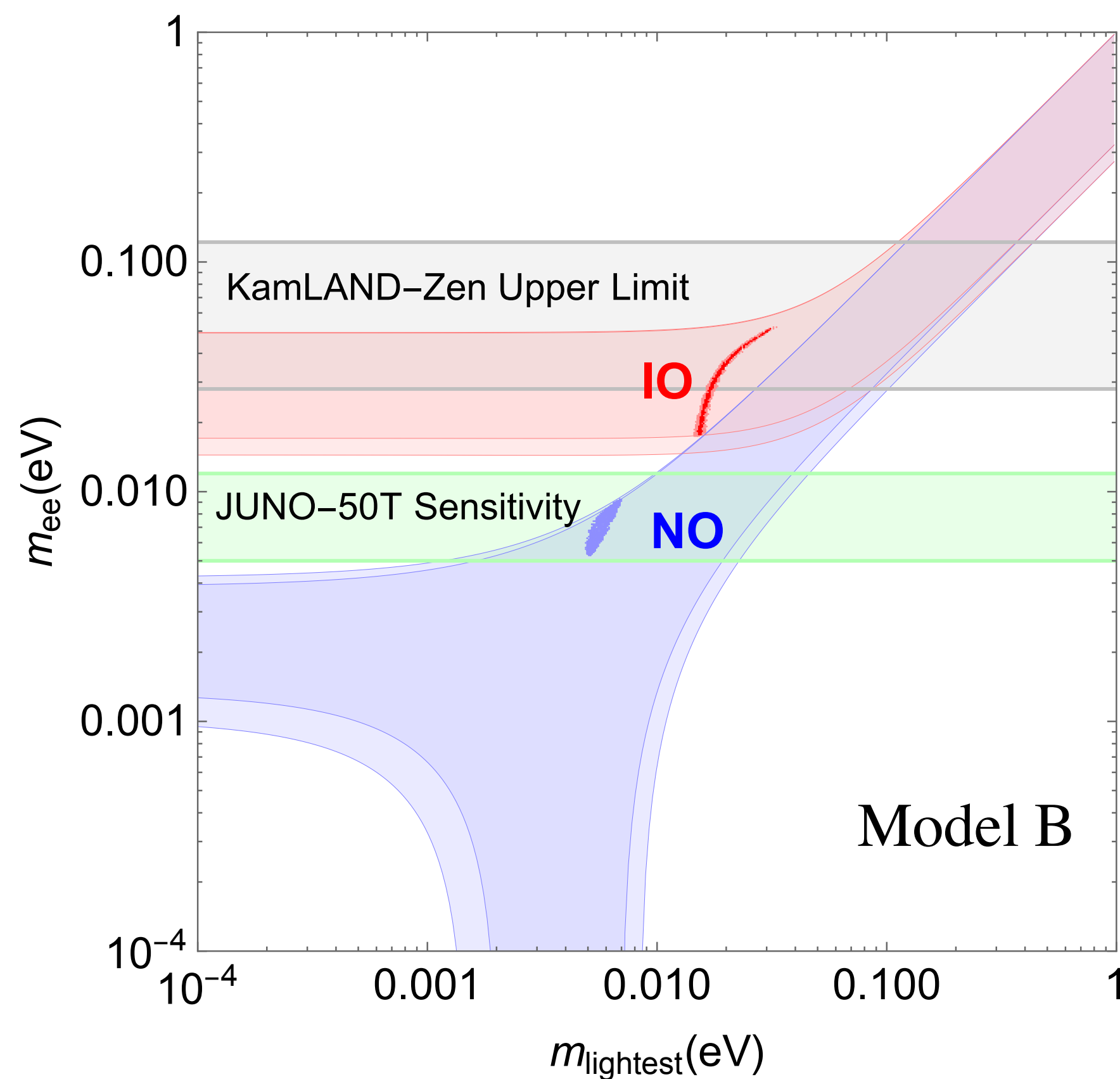
$$\tan \theta_{23} = \left| \frac{\cos \theta_R + \sqrt{\frac{2}{3}} e^{i(\alpha_1 - \alpha_2)} \sin \theta_R}{\cos \theta_R - \sqrt{\frac{2}{3}} e^{i(\alpha_1 - \alpha_2)} \sin \theta_R} \right| ,$$

$$\delta = \arg[(5(\cos 2\theta_R + 1)\cos(\alpha_1 - \alpha_2) - i(\cos 2\theta_R + 5)\sin(\alpha_1 - \alpha_2)] .$$

- In particular for model B, we found (Jiang, RO, Zhou '25):

$$\frac{1}{m_1} = \frac{1}{|2\alpha + \beta|} = \left| \frac{2e^{-2i\alpha_1} \cos^2 \theta_R + e^{-2i\alpha_2} \sin^2 \theta_R}{m_2} + \frac{e^{2i\alpha_1} \cos^2 \theta_R + 2e^{2i\alpha_2} \sin^2 \theta_R}{m_3} \right| .$$

$$m_{ee} = m_0^2 \left| \frac{2}{3(2\alpha + \beta)} + \frac{\beta}{3(\alpha\beta - \gamma^2)} \right| = \left| \frac{2m_2m_3}{3m_2(e^{2i\alpha_1} \cos^2 \theta_R + 2e^{2i\alpha_2} \sin^2 \theta_R) + 3m_3(e^{-2i\alpha_2} \sin^2 \theta_R + 2e^{-2i\alpha_1} \cos^2 \theta_R)} - \frac{1}{3} (m_2 e^{2i\alpha_1} \cos^2 \theta_R + m_3 e^{-2i\alpha_2} \sin^2 \theta_R) \right| .$$



(Jiang, RO, Zhou '25)



5. Summary



- We investigate the landscape of models based on modular S_4 symmetry that realizes TM_1 and compare them with the latest JUNO data.
- We reviewed how the TM_1 mixing pattern arises from the residual symmetries G_ℓ and G_ν after the spontaneous breaking of the flavor symmetry S_4 .
- We show **three different models** that realize the TM_1 in three approaches **with the same symmetry structure**.
- The **predictions** of different models on the relations between the **lightest neutrino mass** and the **effective Majorana mass in neutrinoless double beta decay** are different, making them distinguishable.



Thank you for your attention!



- At the fixed point $\tau_\gamma = \tau_{SU} = -\frac{1}{2} + \frac{i}{2}$, one can solve for the eigenvector of $\rho_3(SU)$ with respect to the degenerate eigenvalue 1, to obtain the modular forms with weight ≤ 4 as (King, Zhou '19)

$$Y_2^{(2)}(\tau_{SU}) \propto \begin{pmatrix} 1 \\ -1 \end{pmatrix}, \quad Y_{3'}^{(2)}(\tau_{SU}) \propto \begin{pmatrix} 1 \\ 1 - \sqrt{6} \\ 1 + \sqrt{6} \end{pmatrix},$$

$$Y_{3'}^{(4)}(\tau_{SU}) \propto \begin{pmatrix} 2 \\ -1 \\ -1 \end{pmatrix}, \quad Y_3^{(4)}(\tau_{SU}) \propto \begin{pmatrix} \sqrt{2} \\ \sqrt{2} - \sqrt{3} \\ \sqrt{2} + \sqrt{3} \end{pmatrix}.$$



- The predictions of model C is given by (King, Zhou '19)

$$\frac{1}{m_1} = \frac{1}{|\beta + 2\gamma|} = \left| \frac{e^{-2i\alpha_2} \sin^2 \theta_R + e^{-i(\alpha_1 + \alpha_2)} \sin 2\theta_R}{m_2} + \frac{e^{2i\alpha_1} \cos^2 \theta_R - e^{i(\alpha_1 + \alpha_2)} \sin 2\theta_R}{m_3} \right|.$$

$$\begin{aligned} m_{ee} &= m_0^2 \left| (M_{R,C}^{-1})_{(1,1)} \right| = m_0^2 \left| \frac{2}{3\beta + 2\gamma} - \frac{\beta}{3(\alpha\beta - \gamma^2)} \right| \\ &= \left| \frac{2m_2m_3}{3m_2(e^{i2\alpha_1} \cos^2 \theta_R - e^{i(\alpha_1 + \alpha_2)} \sin 2\theta_R) + 3m_3(e^{-i2\alpha_2} \sin^2 \theta_R + e^{-i(\alpha_1 + \alpha_2)} \sin 2\theta_R)} \right. \\ &\quad \left. + \frac{1}{3} (m_2 e^{2i\alpha_1} \cos^2 \theta_R + m_3 e^{-2i\alpha_2} \sin^2 \theta_R) \right|. \end{aligned}$$