

12 April, 2026

第十八届粒子物理、
核物理和宇宙学交叉学科前沿问题研讨
桂林市

Radiative corrections to inverse beta decay

O.T., arXiv:2512.07956, arXiv:2512.07957, arXiv:2604.07113



托马拉克

Sasha Tomalak

PMNS oscillation matrix

neutrinos produced and detected in flavor basis

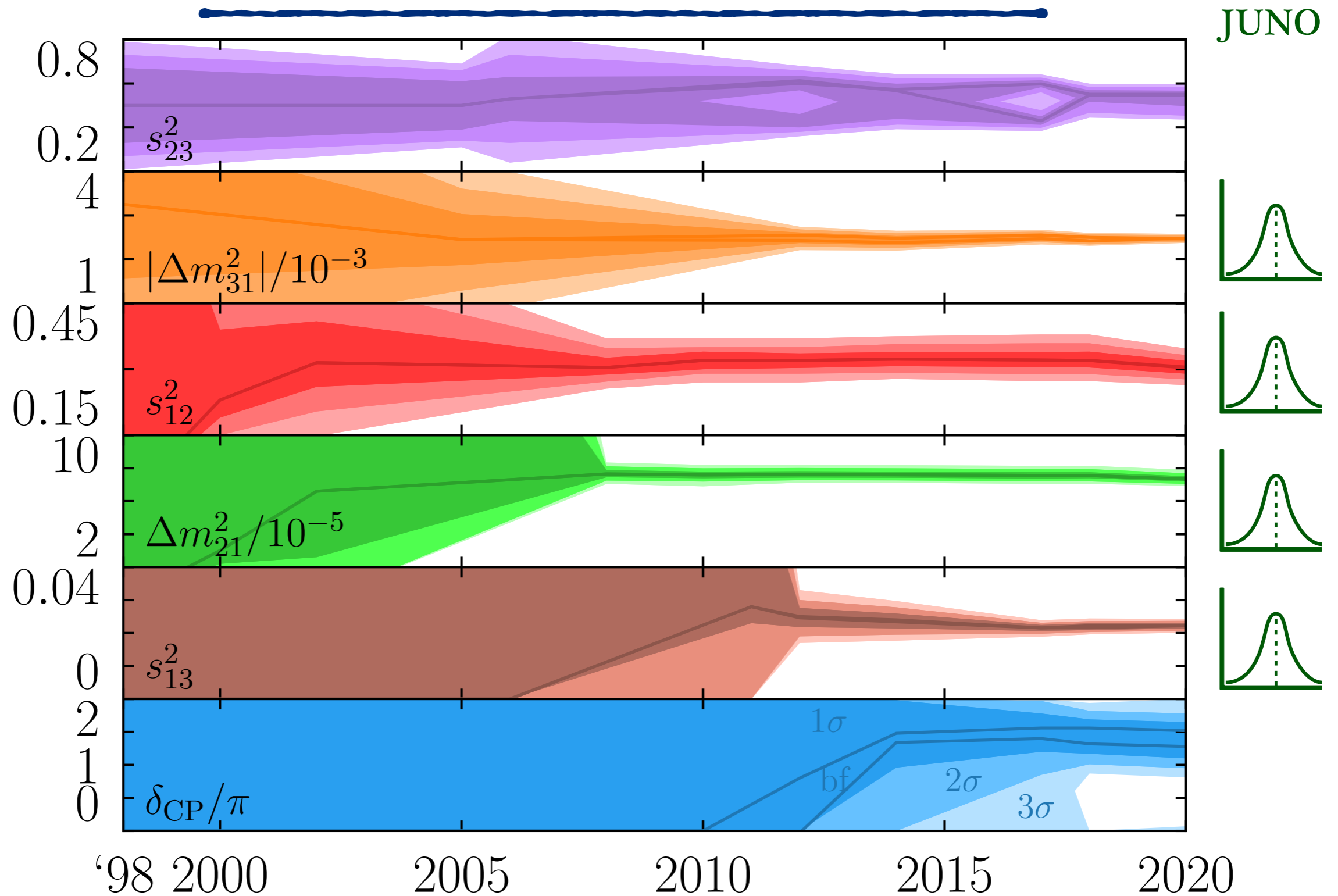
$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{bmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{bmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

neutrinos propagate in mass basis

Pontecorvo-Maki-Nakagawa-Sakata matrix relates two bases

- oscillations are described by **PMNS** mixing matrix

PMNS oscillation matrix

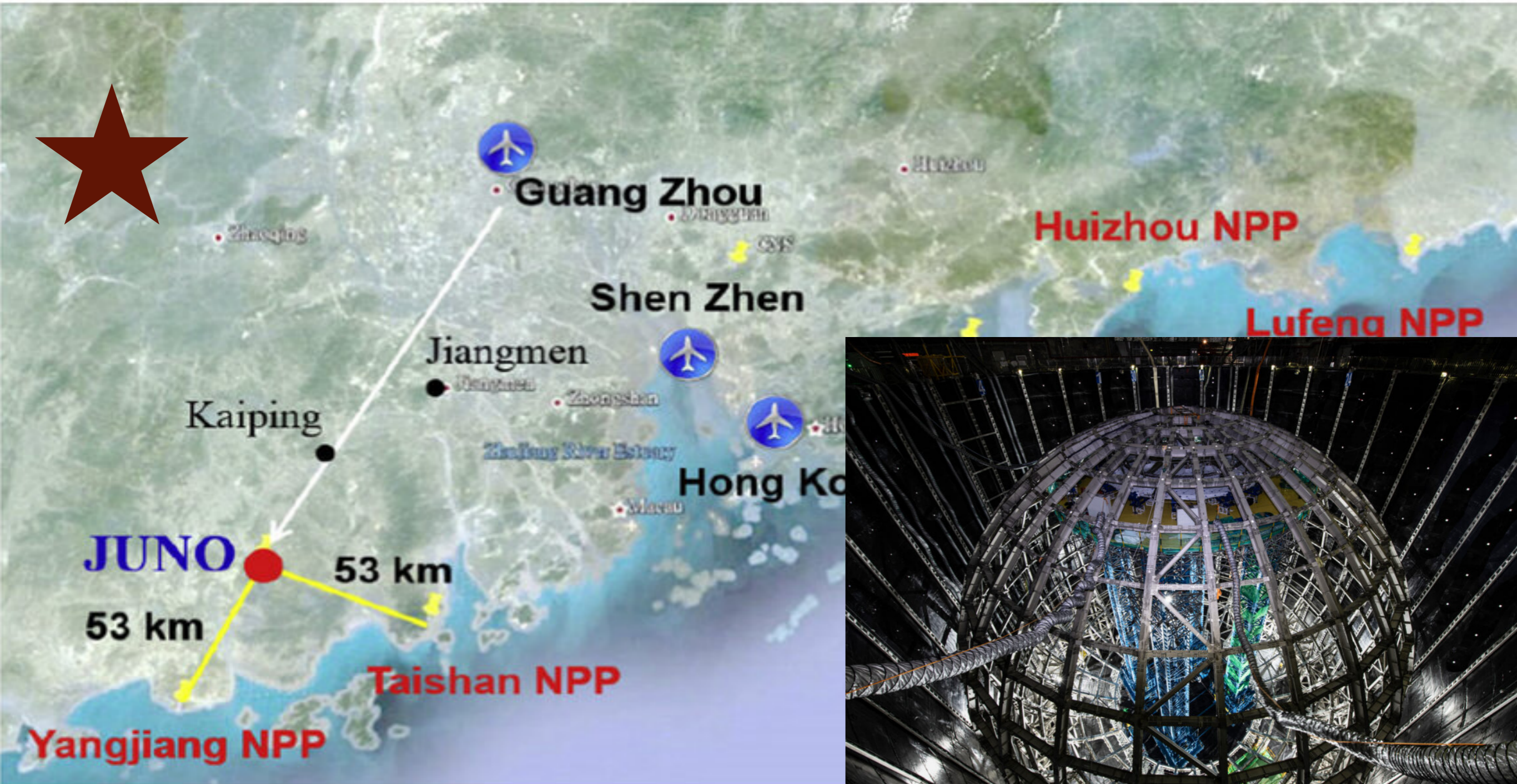


- tremendous progress in our lifetime

Snowmass 2021 NFO1 group report

CP violation and mass hierarchy@laboratory

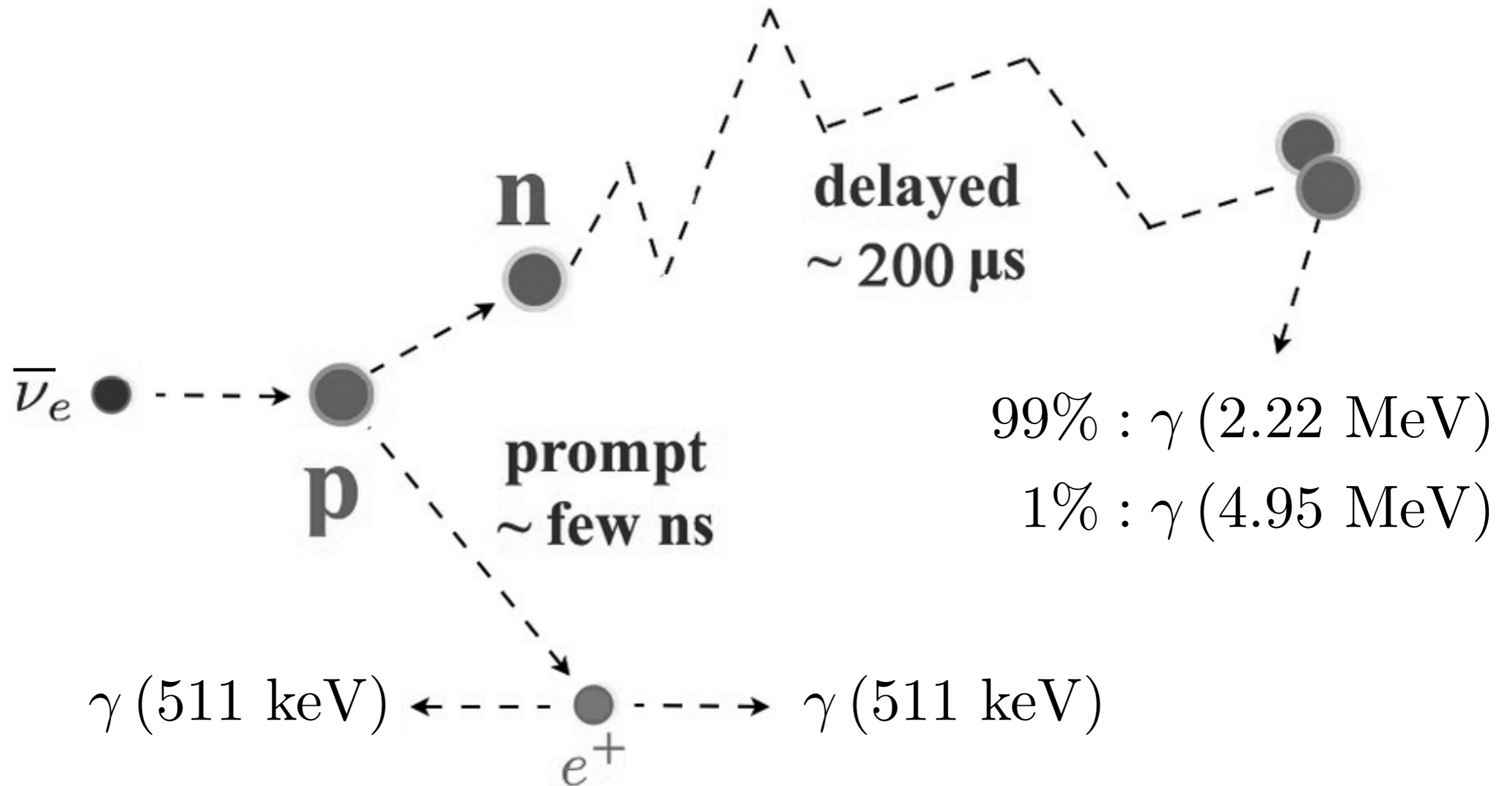
3σ determination of mass hierarchy



- 4 precise oscillation parameters from **JUNO** !!!

35 m, 20000 t

JUNO detection channel

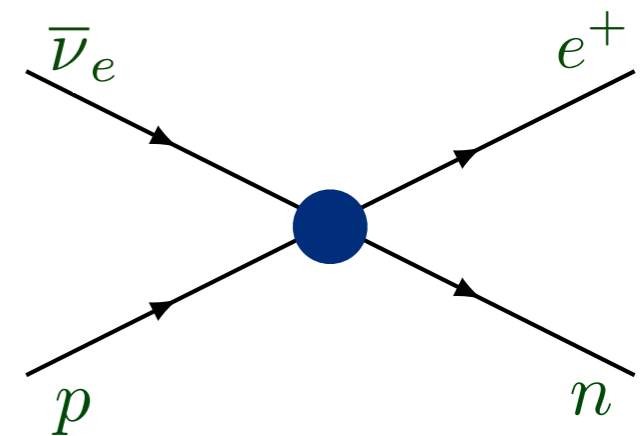
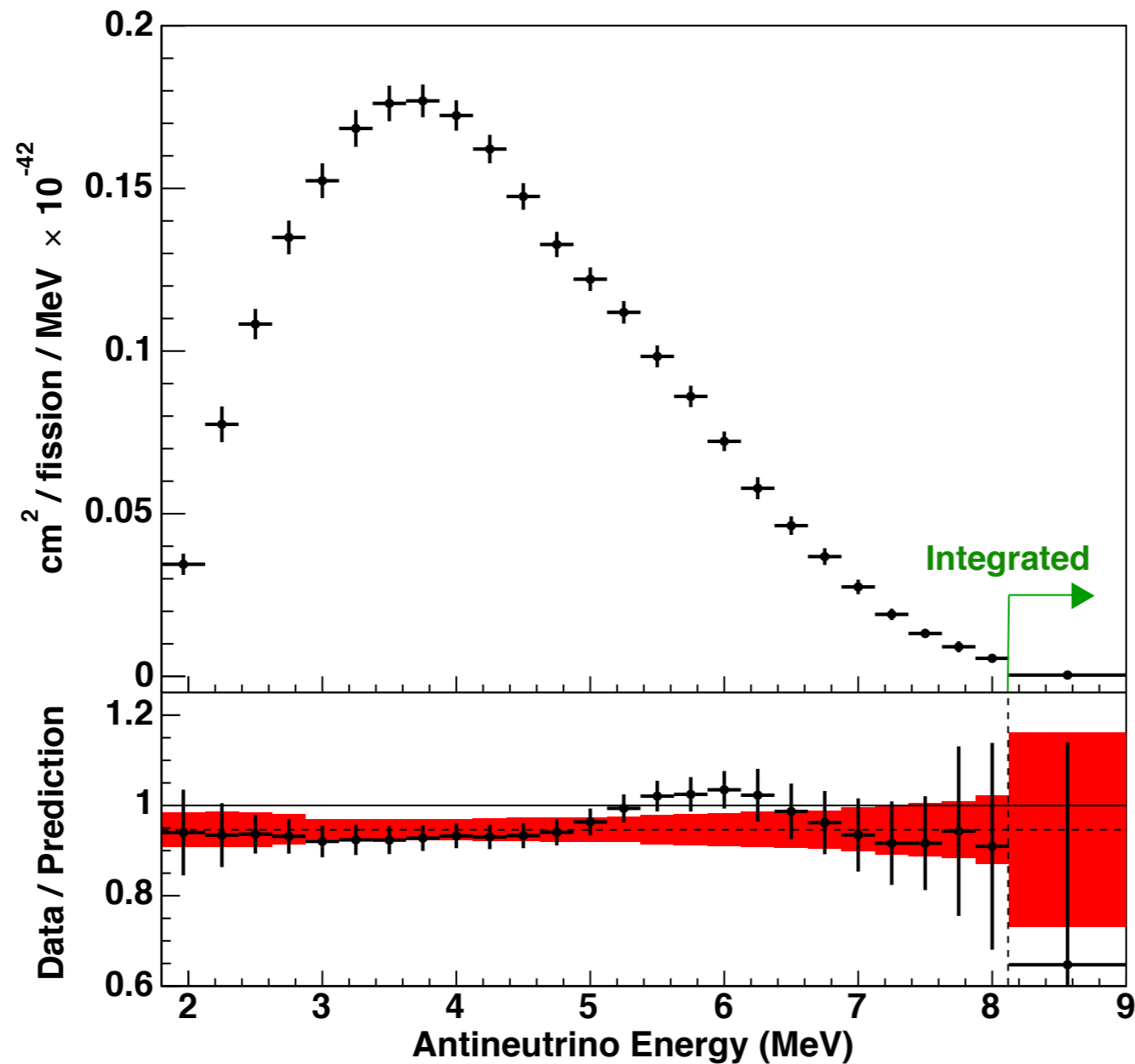


Teng Li et al., ChPC (2017)

- prompt-delayed time signature: background discrimination

IBD with reactor antineutrinos

- neutron is heavier than proton by 1.3 MeV

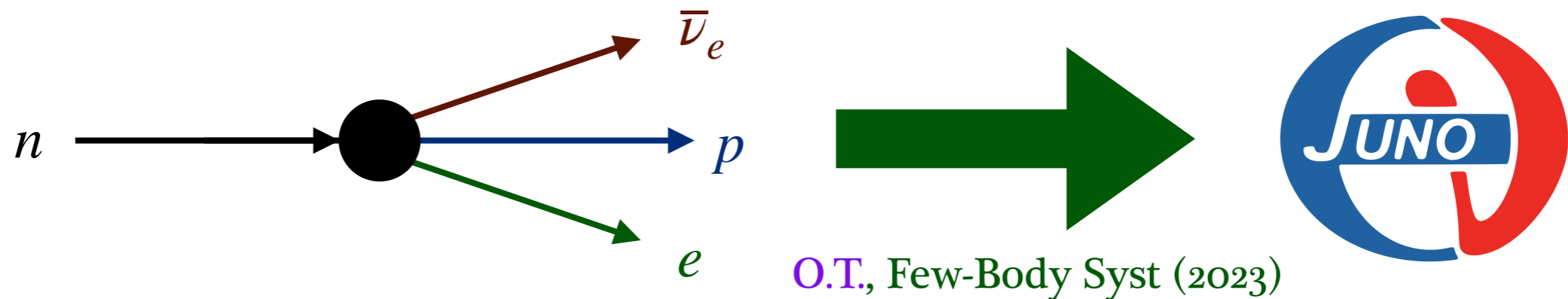


Daya Bay, PRL (2016)

- MeV-scale physics like in neutron decay

Complete EFT approach

- neutron measurements for inverse beta decay



O.T., Few-Body Syst (2023)

Vincenzo Cirigliano, Wouter Dekens, Emanuele Mereghetti, and O.T.,
PRD (2023) and PRD (2025)

- four-fermion interaction between leptons and heavy nucleons

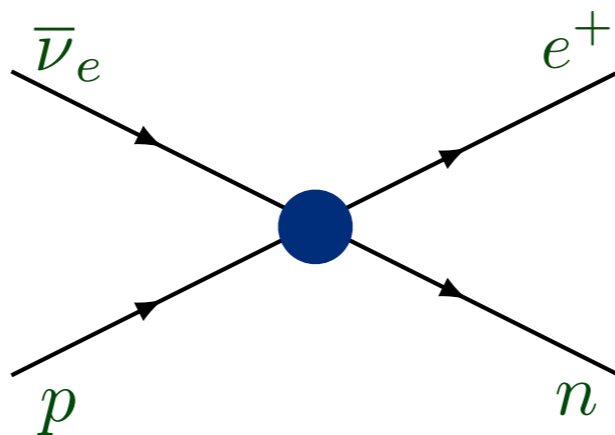
$$\mathcal{L}_{\text{eff}} = -\sqrt{2}G_{\text{F}}V_{ud}\bar{e}\gamma_{\mu}P_{\text{L}}\nu_e \cdot \bar{N}(g_V v^{\mu} - 2g_A S^{\mu})\tau^{+}N$$

g_V : theoretical prediction from Standard Model

g_A : beta asymmetry in polarized neutron decay

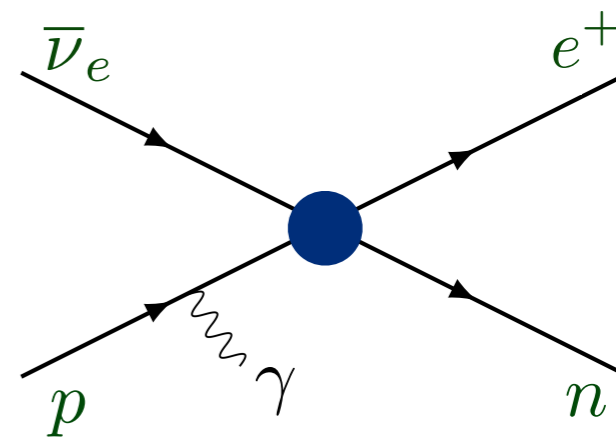
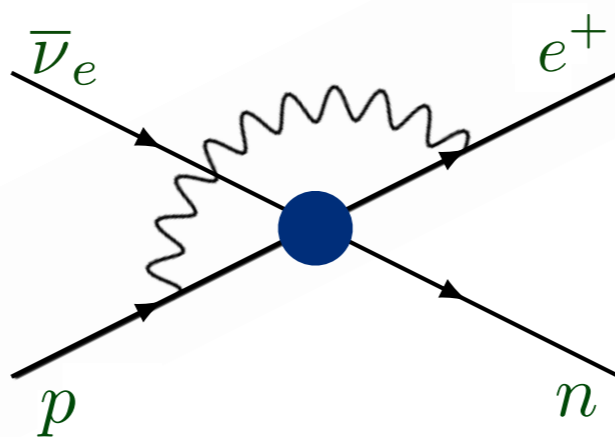
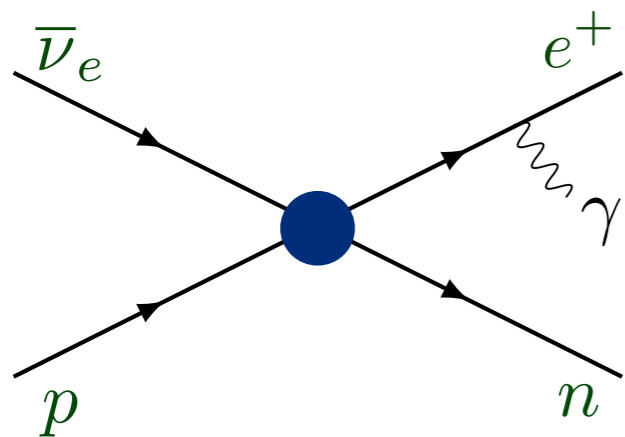
V_{ud} : superallowed transitions

- first consistent at 0.1%-level QED radiative corrections

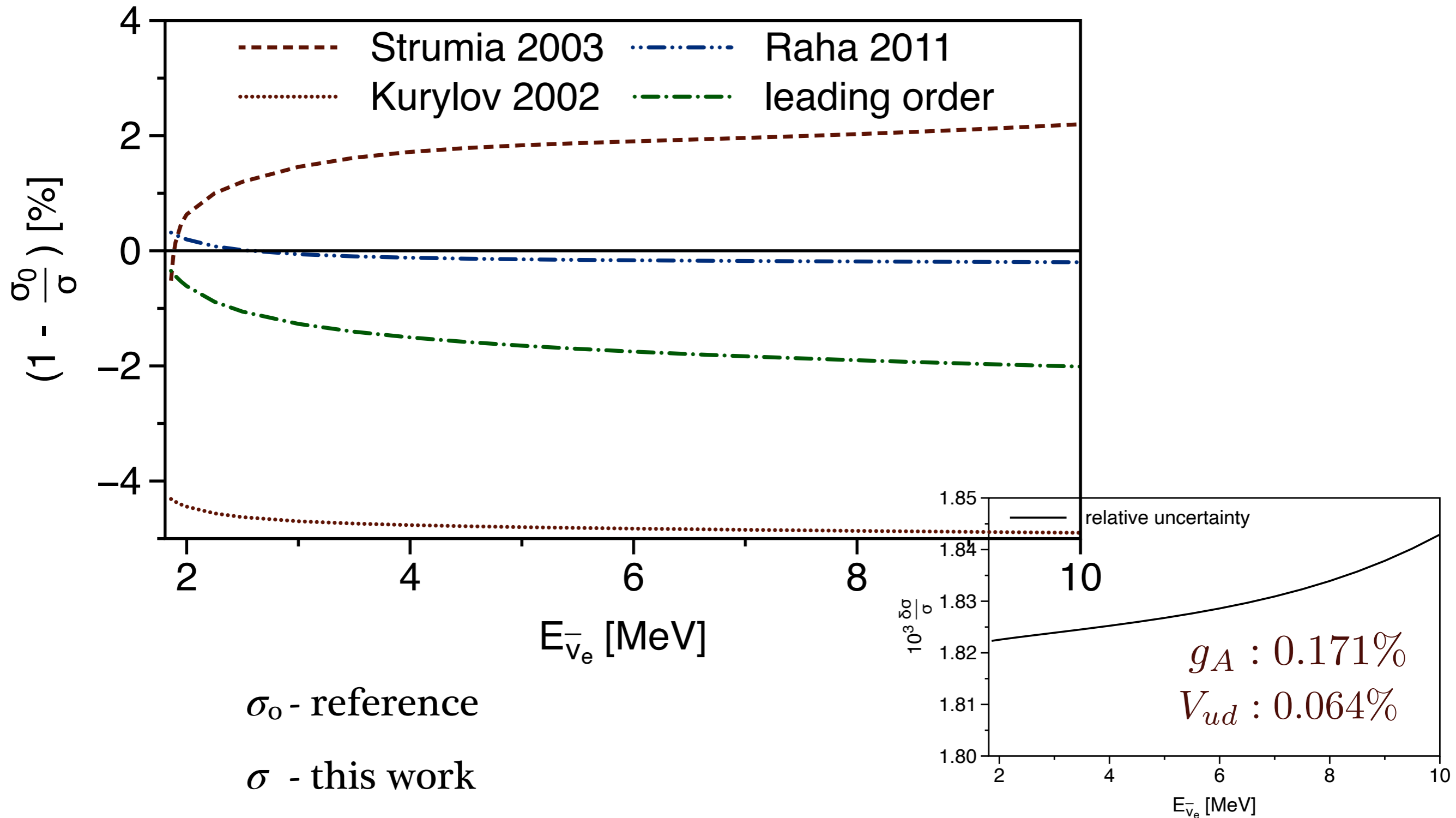


O.T., arXiv:2512.07956
arXiv:2512.07957

Theory of inverse beta decay for reactor antineutrinos

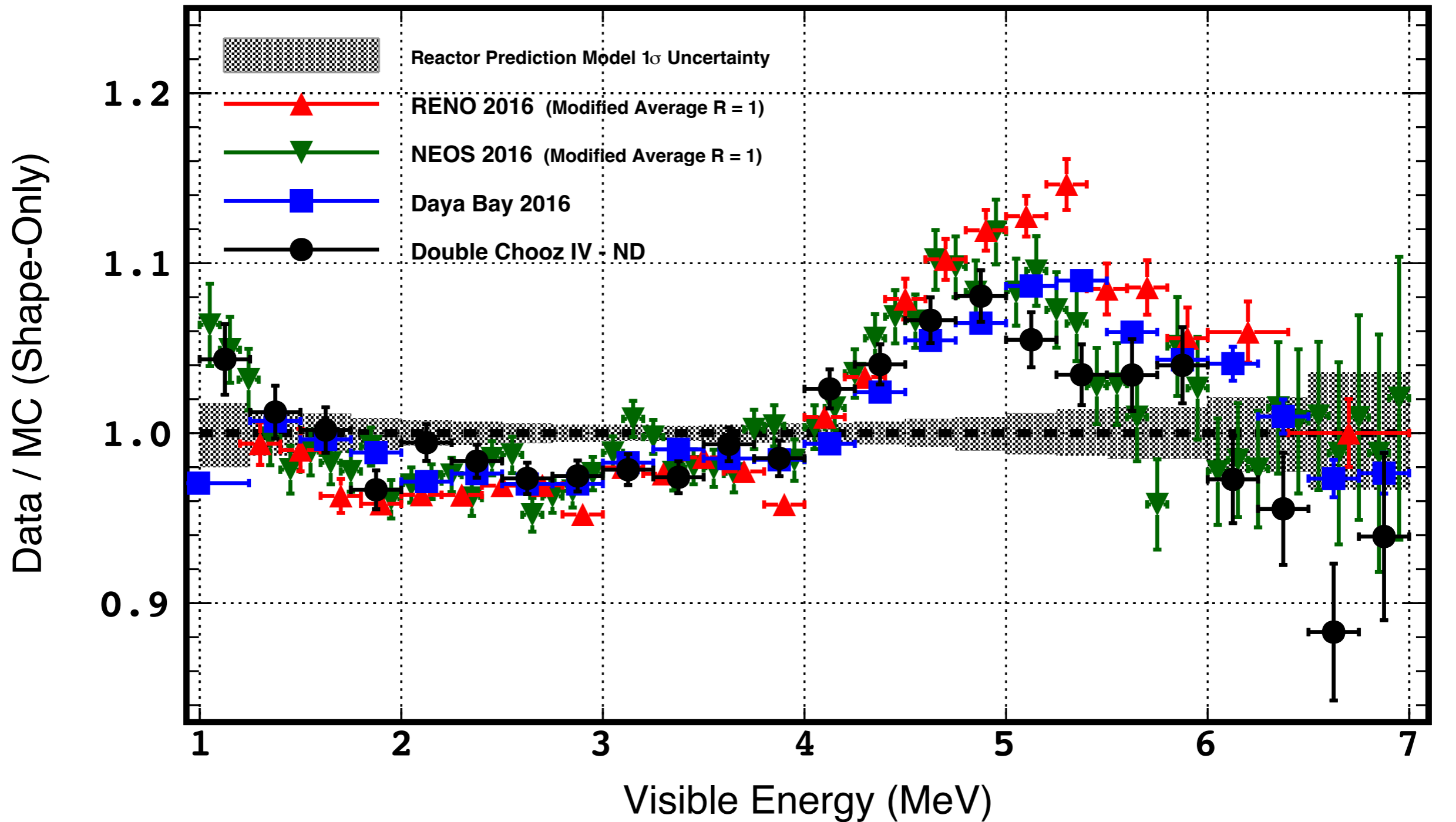


Total cross section



- %-level improvement to previous results; **0.18%** uncertainty

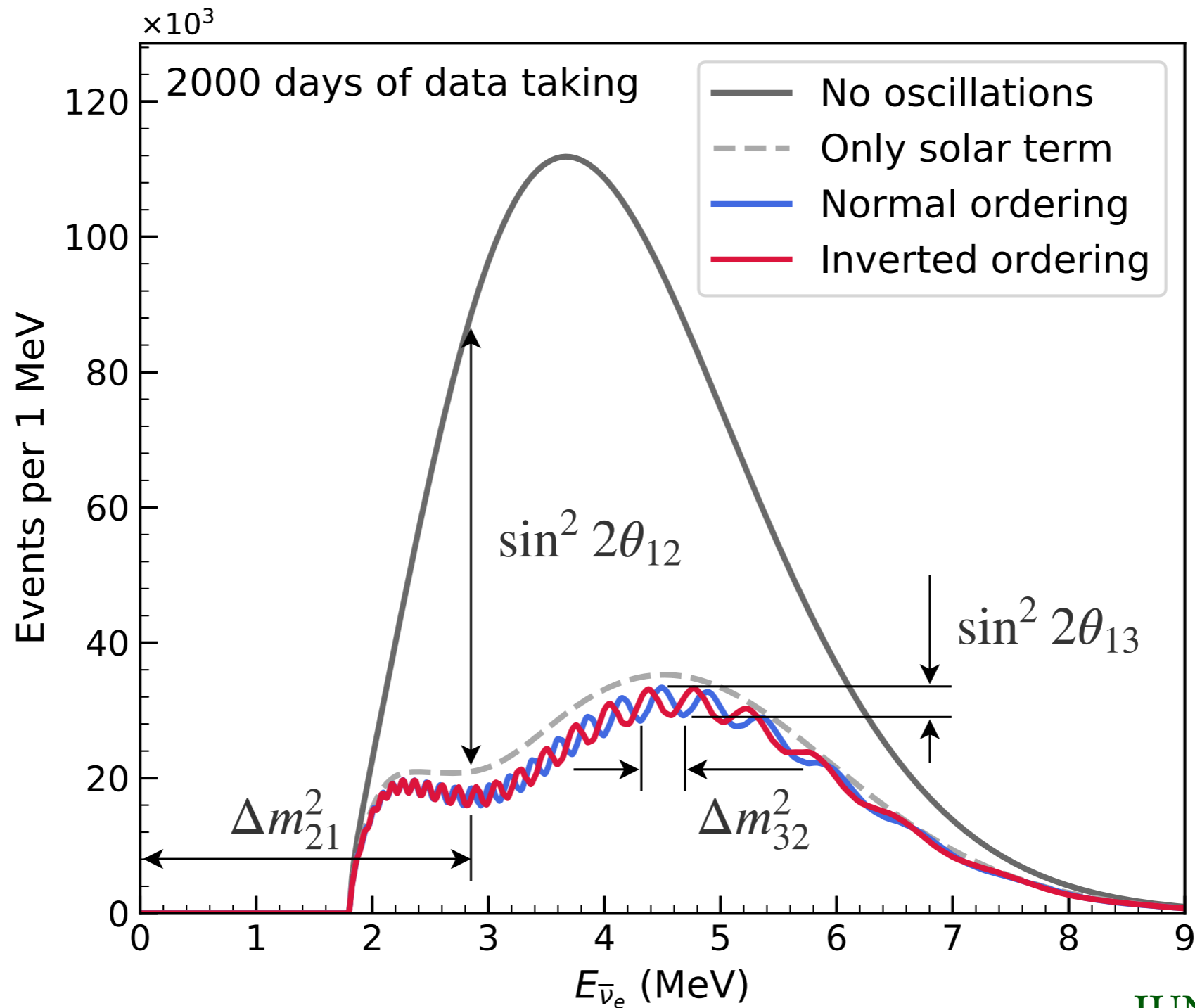
5-MeV flux excess



Double Chooz, Nature Phys (2019)

- update in cross section at 4-6 MeV: shift by 15-20% of bump

Energy reconstruction is key



JUNO, PPNP (2021)

- JUNO: 3% energy reconstruction for precision goals

Bremsstrahlung

- new distributions beyond E_{EM} spectrum

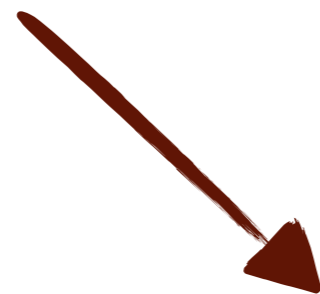
$$E_n, E_\gamma, \theta_n$$



$$E_n, \theta_n$$



$$E_n, E_{EM}$$



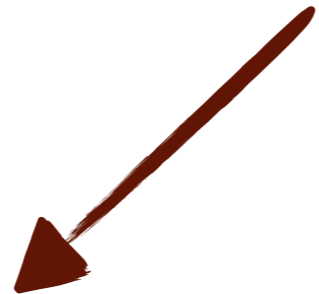
$$E_e, E_\gamma, \theta_e$$



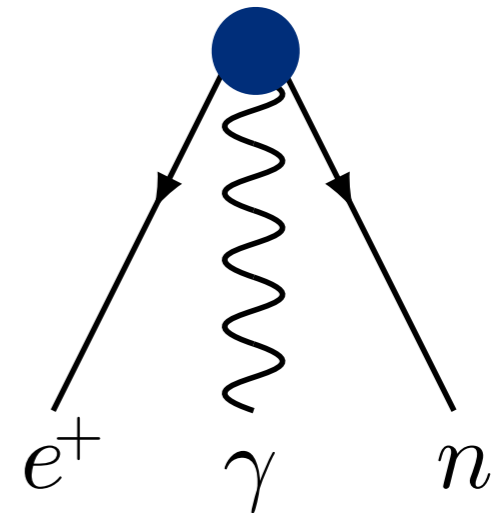
$$E_e, \theta_e$$



$$E_e$$



total cross section



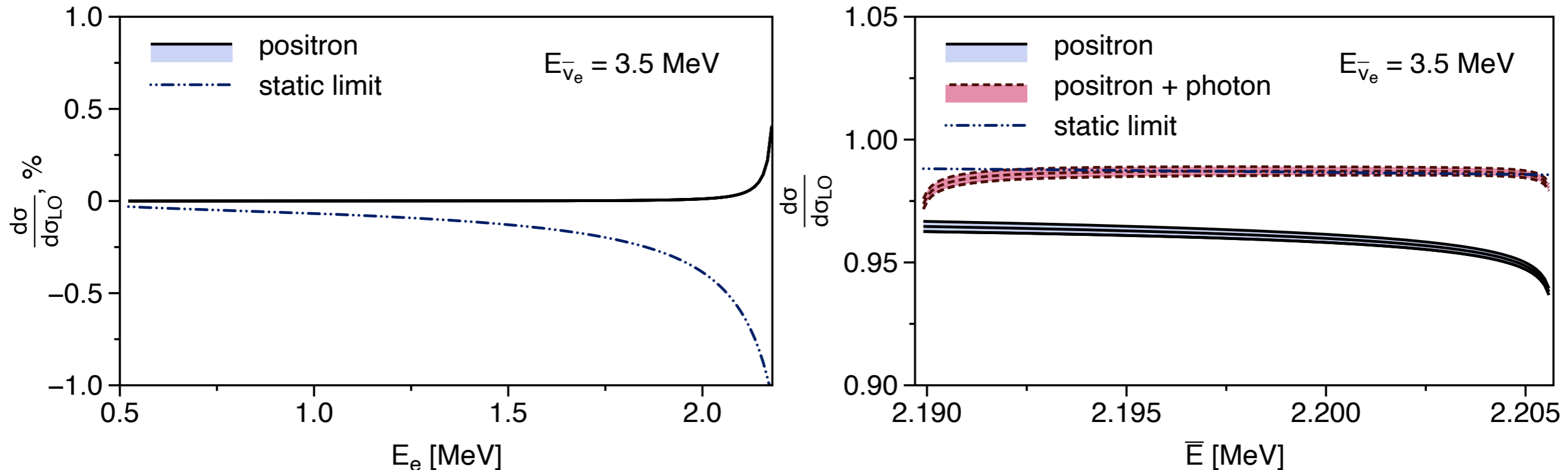
$$E_{EM} = E_e + E_\gamma$$

consistency at cross-section level

- analytical positron energy spectrum, 2D/3D distributions

Access each particle kinematics

- evaluate radiative process w/o no-recoil approximation

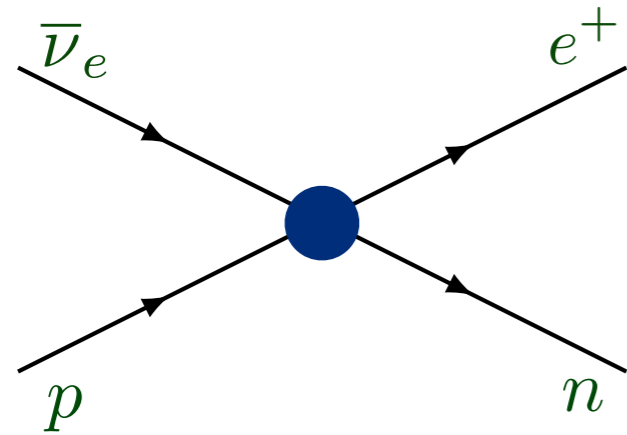


- 1st analytic expression for **positron** energy spectrum

$$E_{\bar{\nu}_e} \approx m_n - m_p + E_{e^+} + E_\gamma + O\left(\frac{1}{M}\right)$$

- control over each particle is necessary to achieve precision

Conclusions



radiative corrections
in EFT framework
+ scheme-dependence

Standard Model \rightarrow LEFT \rightarrow HB χ PT \rightarrow $\not\pi$ EFT

- 1st consistent inclusion of all QCD/EW effects at **0.1%** level
- total cross section, energy spectra, 2D/3D distributions in IBD
- precision physics: **important** to control γ and e^+ kinematics in calibration of energy response in liquid scintillators

$$\nu e^- \rightarrow \nu e^- (\gamma)$$

$$\pi^+ \rightarrow \mu^+ \nu_\mu (\gamma)$$

$$K^+ \rightarrow \mu^+ \nu_\mu (\gamma)$$

$$\mu^+ \rightarrow e^+ \nu_e \bar{\nu}_\mu (\gamma)$$

Thanks for your attention !!!

$$\bar{\nu}_\ell p \rightarrow \ell^+ n (\gamma)$$

谢谢

$$\nu_\ell n \rightarrow \ell^- p (\gamma)$$

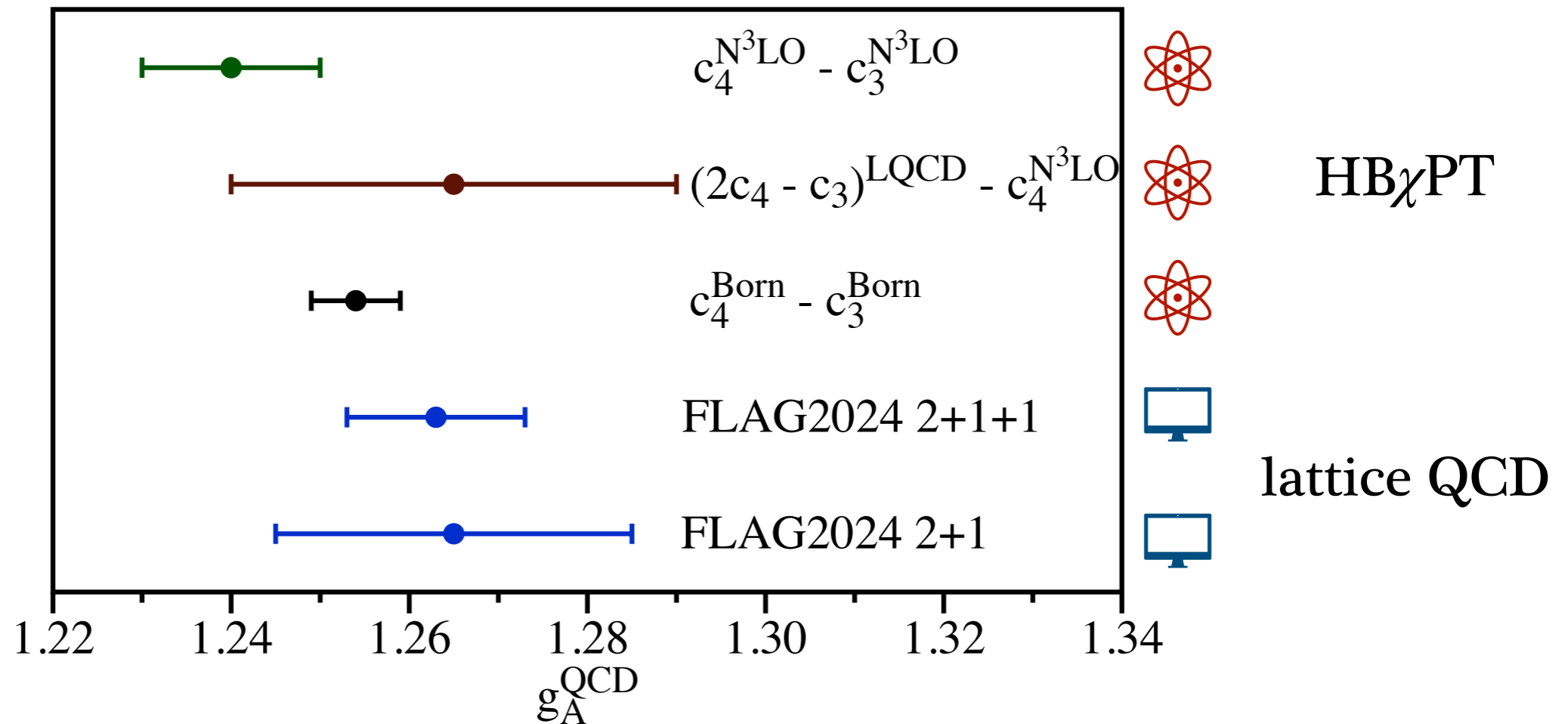
$$n \rightarrow p e^- \bar{\nu}_e$$

$$\nu \text{}^{40}\text{Ar} \rightarrow \nu \text{}^{40}\text{Ar} (\gamma)$$

Nucleon axial-vector charge

- effective field theory approach to low-energy charged currents

Standard Model \rightarrow LEFT \rightarrow HB χ PT \rightarrow π EFT



O.T. and Yi-Bo Yang, Universe (2026)

Vincenzo Cirigliano, Wouter Dekens, Emanuele Mereghetti, and O.T., PRD (2024) and PRD (2025)

V. Cirigliano, J. de Vries, L. Hayen, E. Mereghetti, and A. Walker-Loud, PRL (2022)

- first consistent at 0.1%-level QED radiative corrections

Update in nuclear/neutron decay

- evaluate radiative process w/o no-recoil approximation

$$m_n = m_p + E_e + E_\gamma + E_\nu + O\left(\frac{1}{M}\right)$$

- V_{ud} from superallowed transitions and neutron decay

$$V_{ud}^{0^+ \rightarrow 0^+} = 0.97373(31) \longrightarrow V_{ud}^{0^+ \rightarrow 0^+} = 0.97361(31)$$

Hardy and Towner, PRC (2020)

changes by 1 uncertainty of radiative corrections

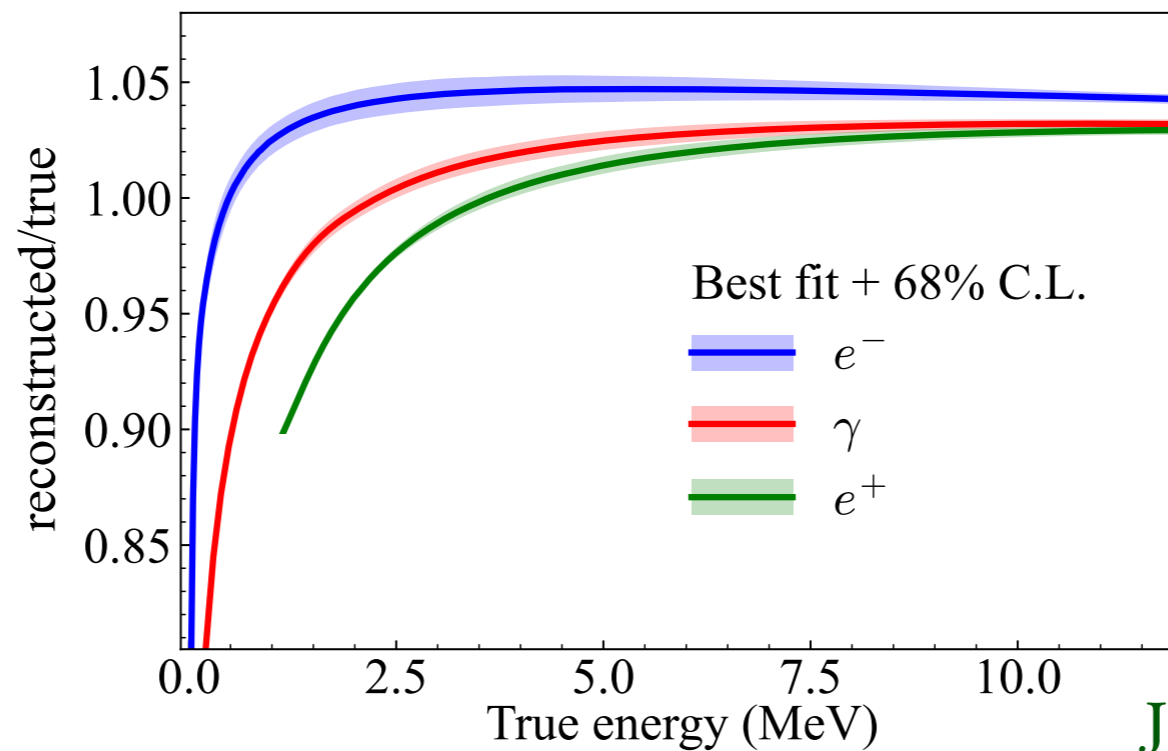
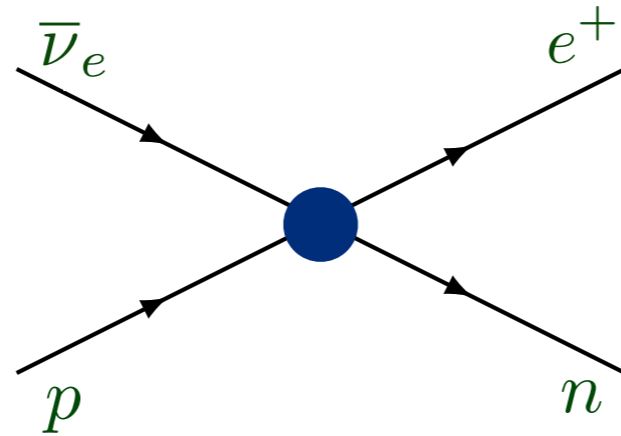
- g_A from beta asymmetry in polarized neutron decay

$$A = \frac{N_\uparrow - N_\downarrow}{N_\uparrow + N_\downarrow}$$

changes by 1/3 of PERKEO-III experimental uncertainty

- numerically quantified shifts in V_{ud} and $g_A \lesssim 1\sigma$

Energy reconstruction is key

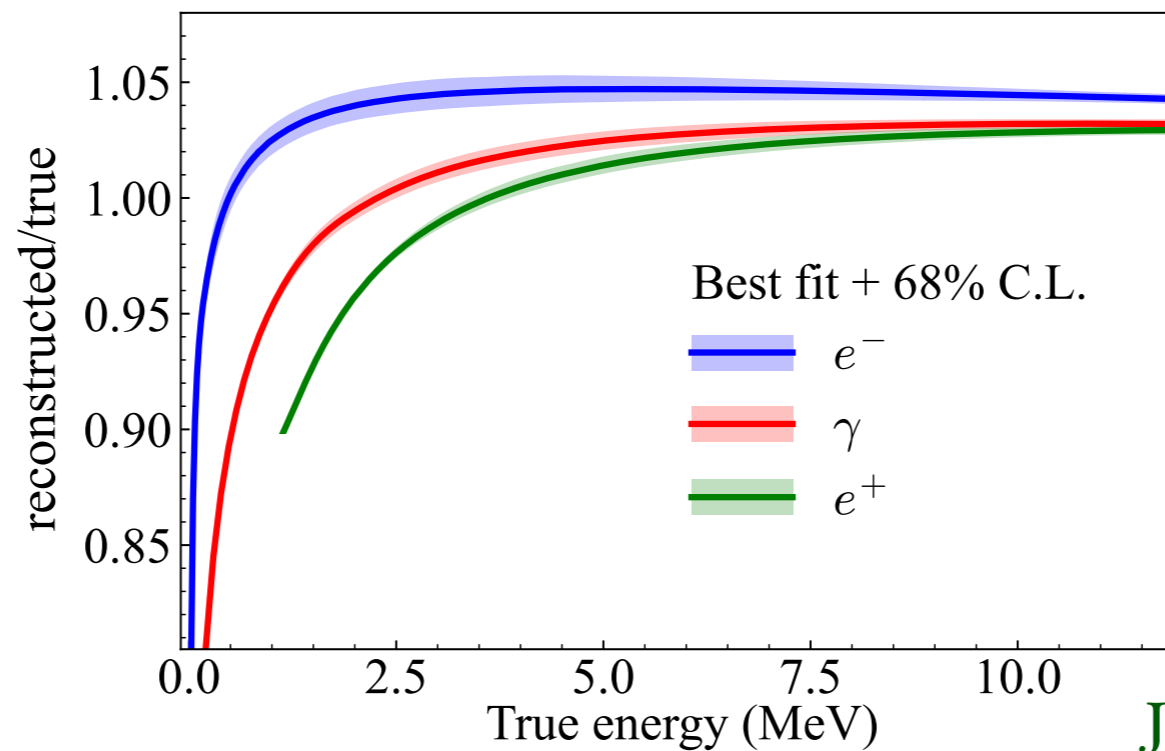
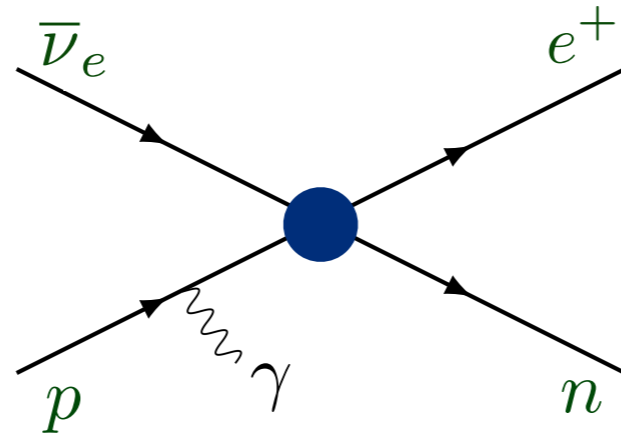


JUNO, arXiv:2511.14590

$$E_{\bar{\nu}_e} \approx m_n - m_p + E_{e^+} + O\left(\frac{1}{M}\right)$$

- control over each particle is necessary to achieve precision

Energy reconstruction is key



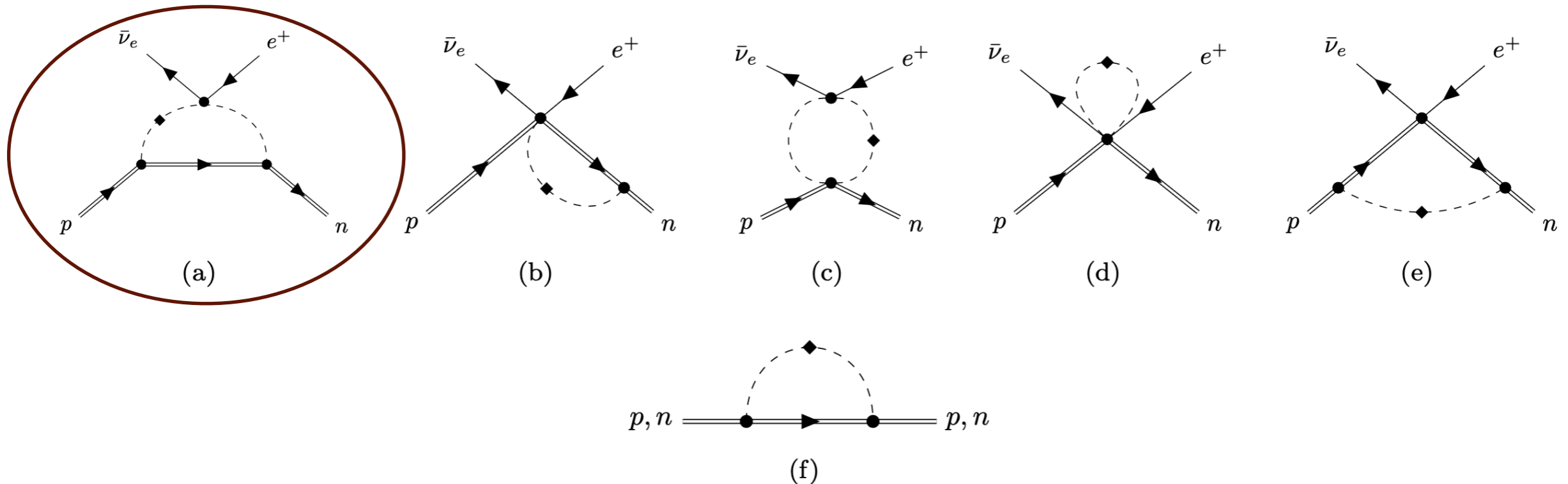
JUNO, arXiv:2511.14590

$$E_{\bar{\nu}_e} \approx m_n - m_p + E_{e^+} + E_\gamma + O\left(\frac{1}{M}\right)$$

- control over each particle is necessary to achieve precision

Supernova and π DAR energies

- include pion degrees of freedom to inverse beta decay



- work at leading and next-to-leading order in HBChPT

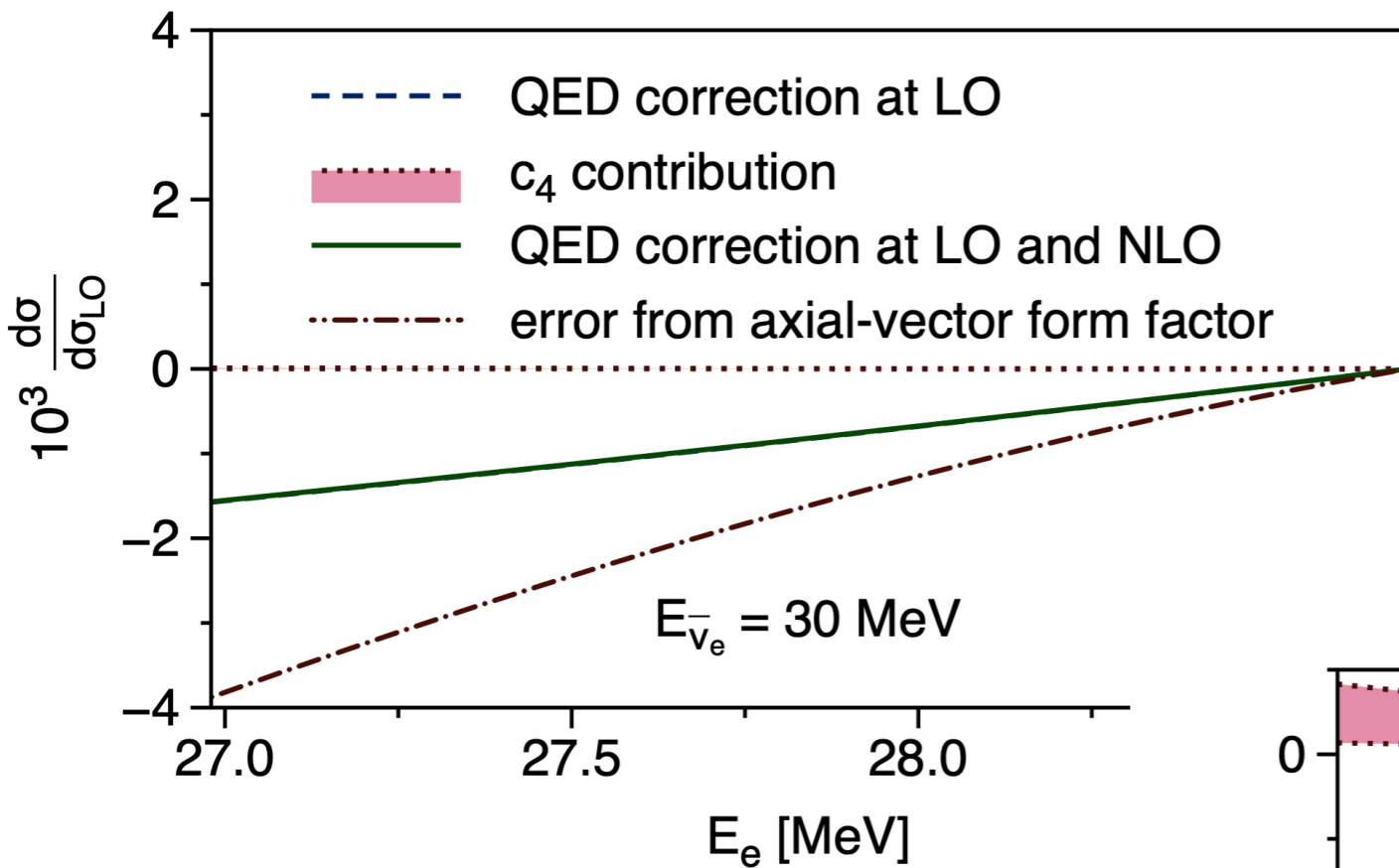
- neglect recoil diagrams without pions

O.T., arXiv:2604.07113

$$\frac{m_e^2}{m_\pi^2} \sim \frac{\alpha m_e}{\pi m_\pi} \ll \frac{\alpha m_e}{\pi E_{\bar{\nu}_e}} \lesssim \frac{\alpha E_{\bar{\nu}_e}}{\pi m_n} \lesssim \frac{\alpha E_{\bar{\nu}_e}^2}{\pi m_\pi^2} \ll \frac{\alpha m_\pi}{\pi m_n} \lesssim \frac{\alpha E_{\bar{\nu}_e}}{\pi m_\pi}$$

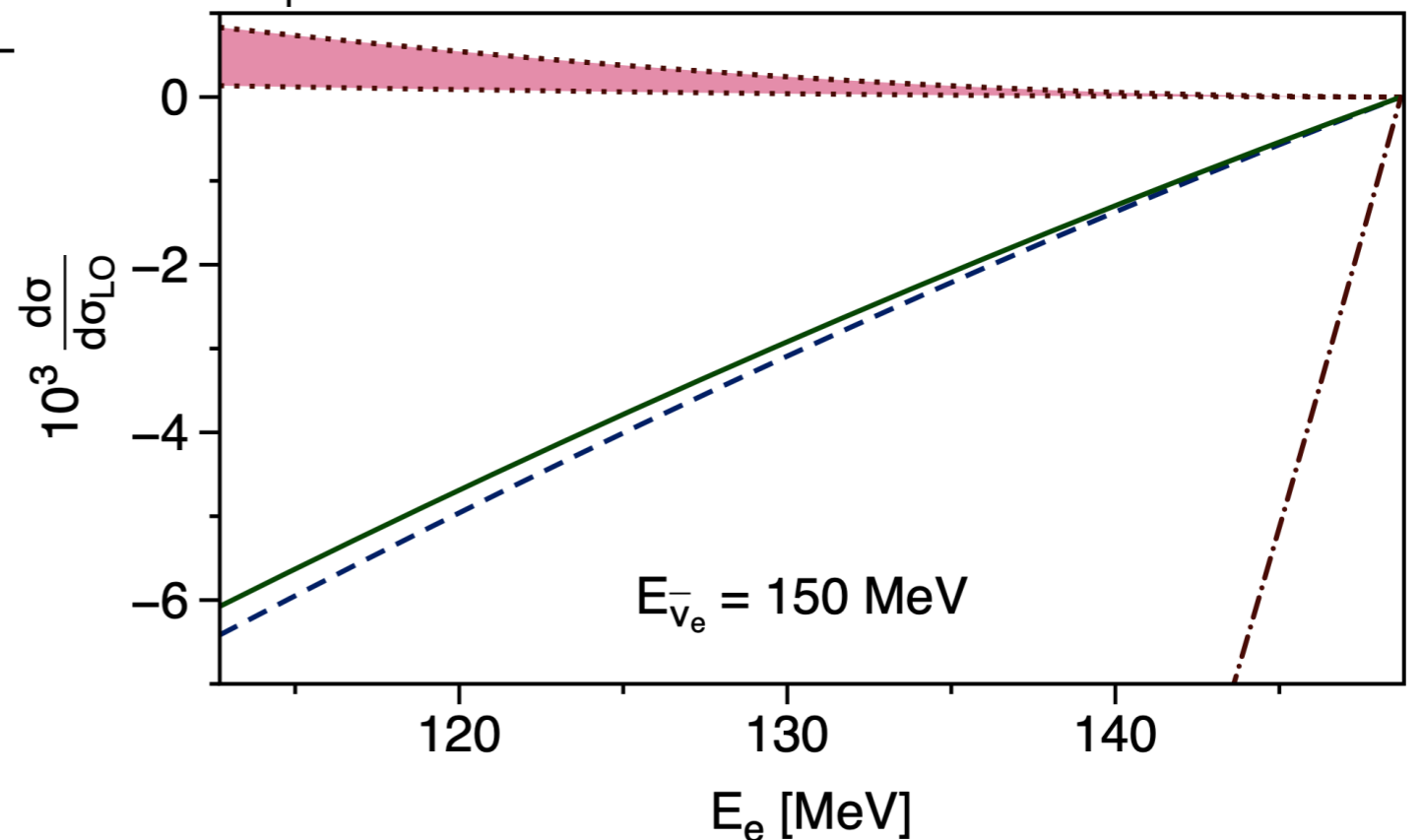
- extends **0.1%-level** predictions to energies above **10 MeV**

Supernova and π DAR energies



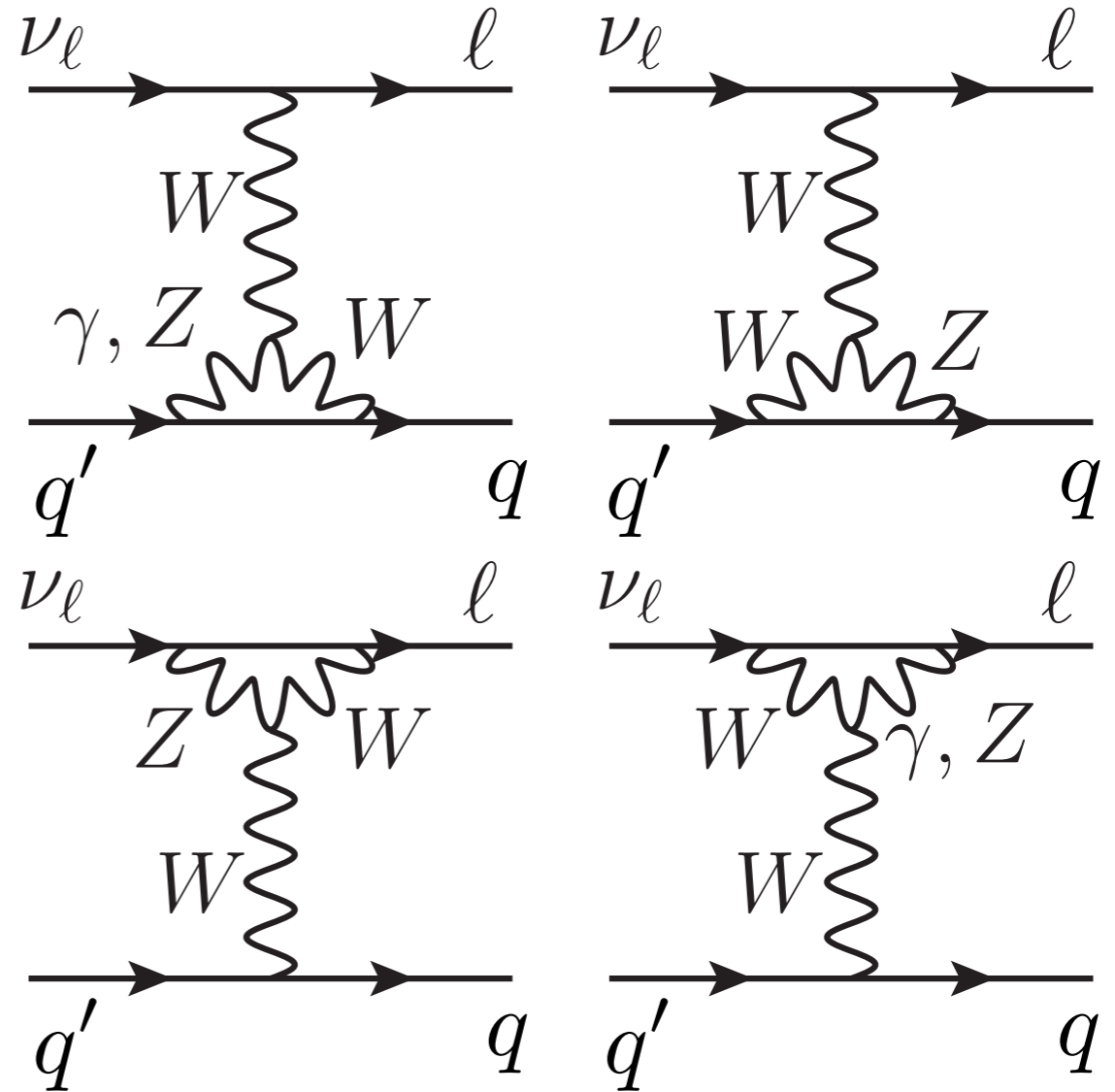
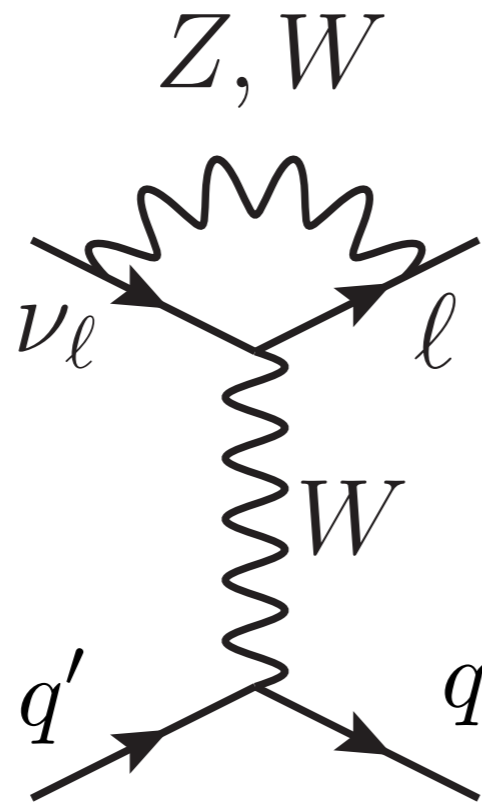
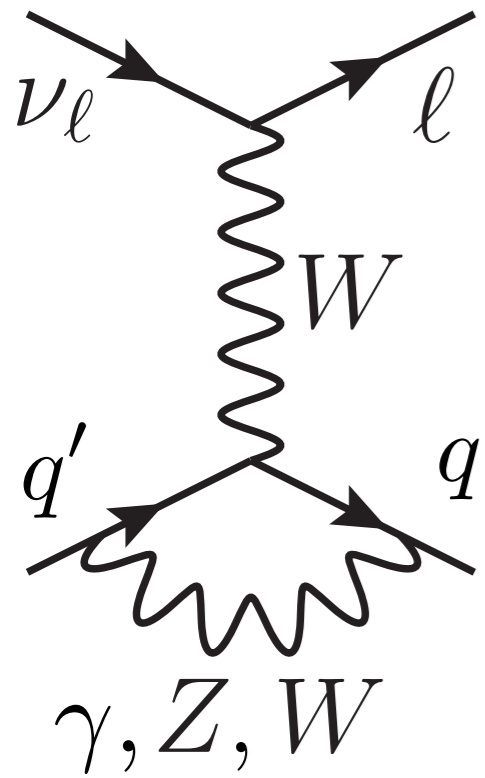
Wilson coefficient c_4 enters at higher energies

O.T., arXiv: 2604.07113



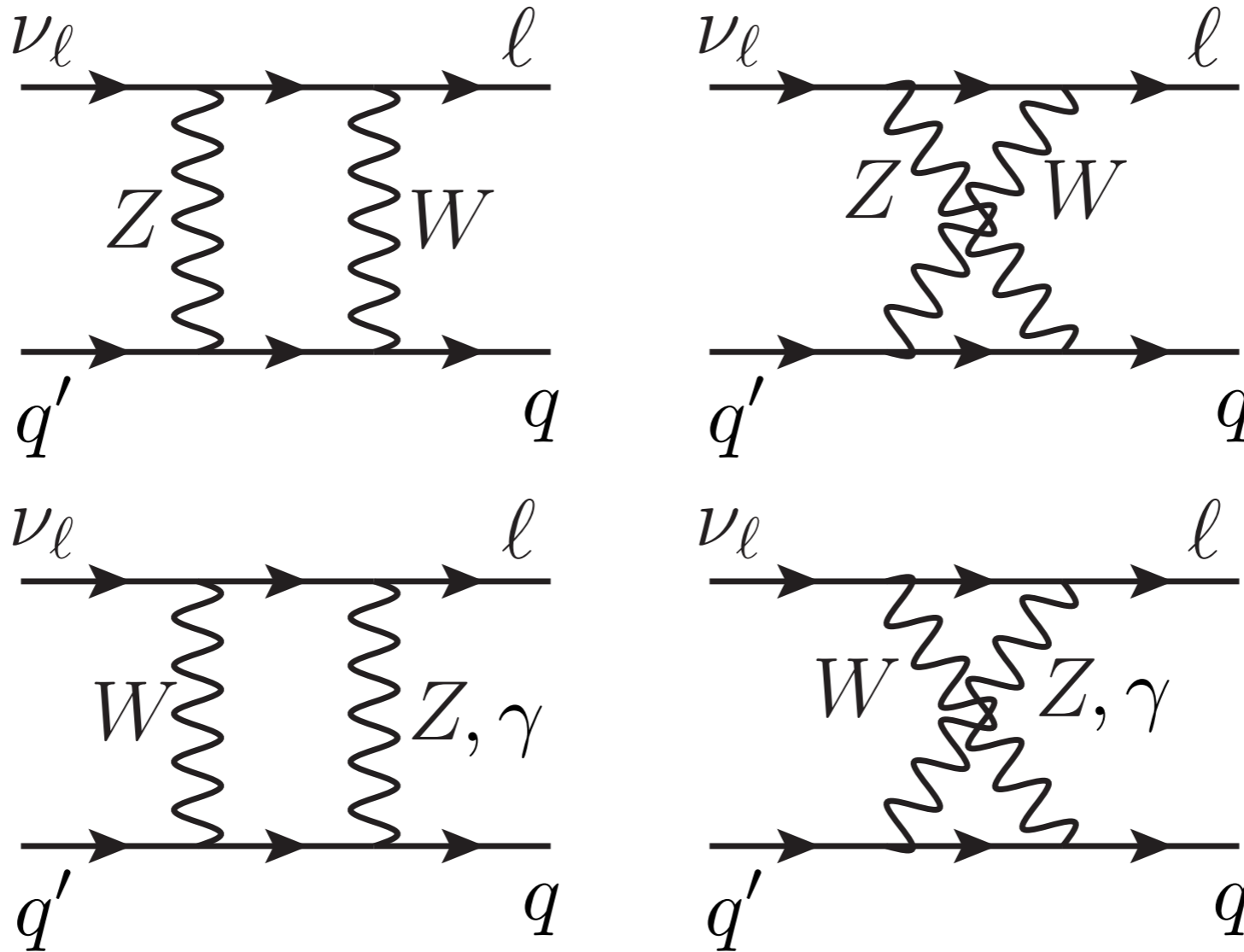
- extends **0.1%-level** predictions to energies above **10 MeV**

Charged current in SM. Vertexes



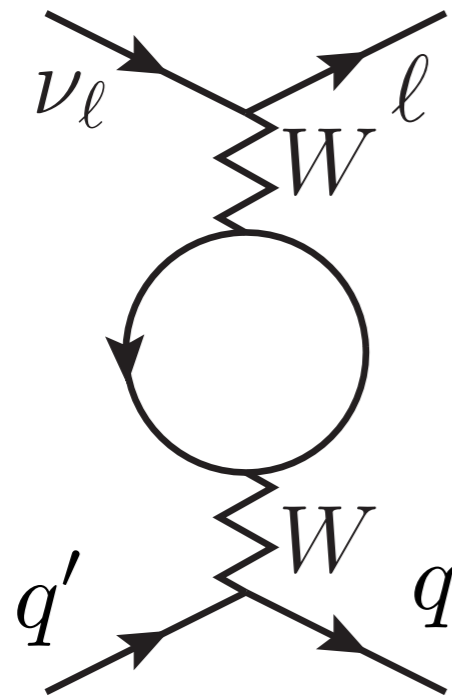
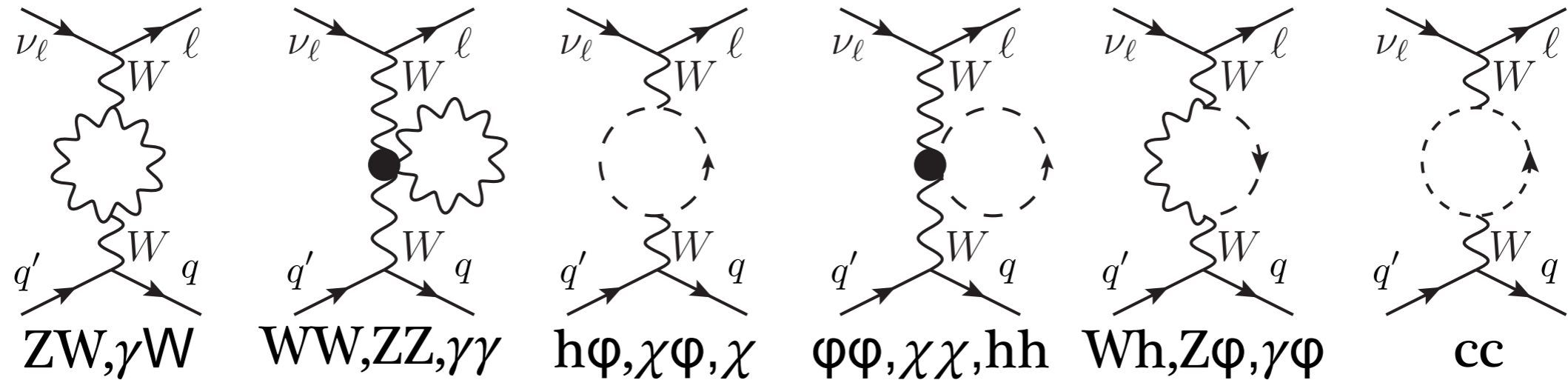
- contribution to effective couplings

Charged current in SM. Boxes



- contribution to effective couplings

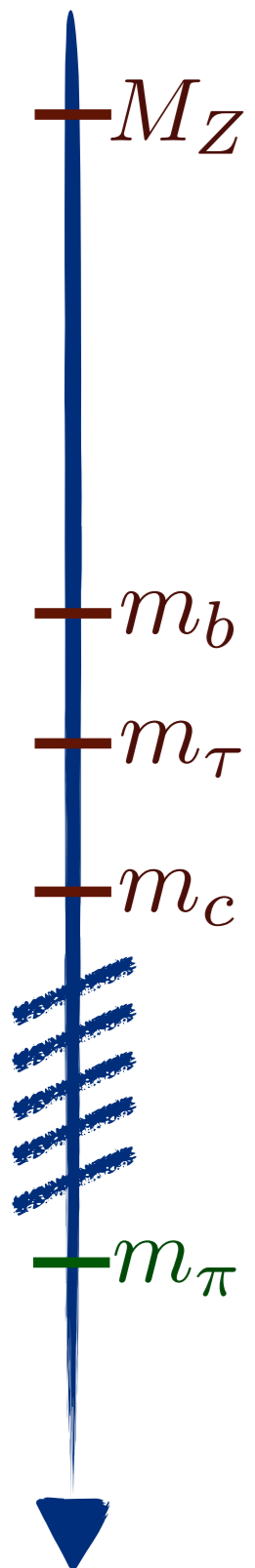
W self energy



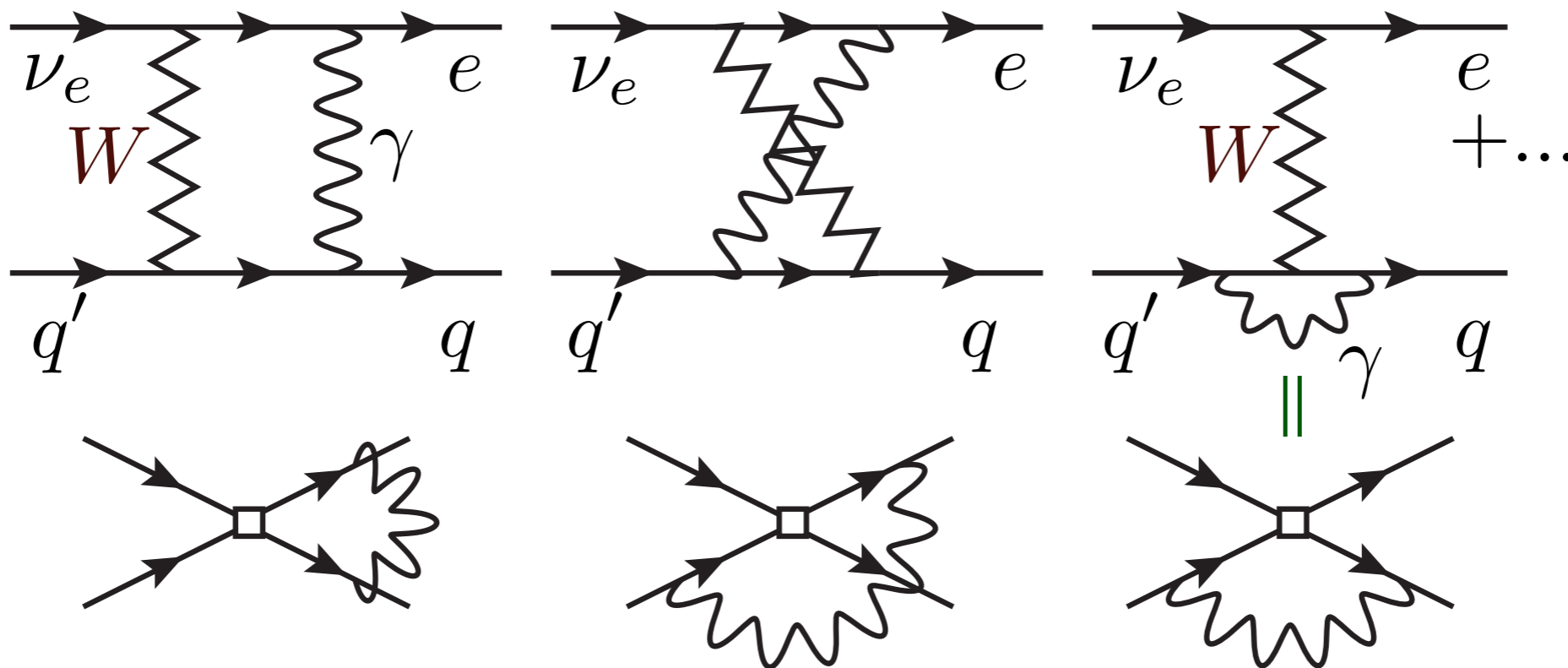
- vanishing contribution to matching besides loops with t quark

- contribution to effective couplings

LEFT Lagrangian



$$\mathcal{L}_{\text{LEFT}} \sim G_F V_{ud} C_\beta^r(a, \mu) \bar{e}_L \gamma_\rho \nu_{eL} \bar{u}_L \gamma^\rho d_L$$



scheme independent

scheme dependent for quarks

- NDR scheme for γ_5 for evanescent operators E

$$\gamma^\alpha \gamma^\beta \gamma^\mu P_L \otimes \gamma_\mu \gamma_\beta \gamma_\alpha P_L = 4(1 + a(4 - d)) \gamma^\mu P_L \otimes \gamma_\mu P_L + E(a)$$

Buras and Weisz, NPB (1990)

- scheme dependence of 1-loop matching and 2-loop running

Logarithms in neutron decay

M_Z

four-fermion interaction between leptons and heavy nucleons

$$\mathcal{L}_{\text{eff}} = -\sqrt{2}G_F V_{ud} \bar{e} \gamma_\mu P_L \nu_e \cdot \bar{N} (g_V v^\mu - 2g_A S^\mu) \tau^+ N$$

scale separation introduces large **logarithms**:

hadronic scale

$$\ln \frac{M_Z}{\Lambda_{\text{had}}}$$

$$\ln \frac{\Lambda_{\text{had}}}{m_e}$$

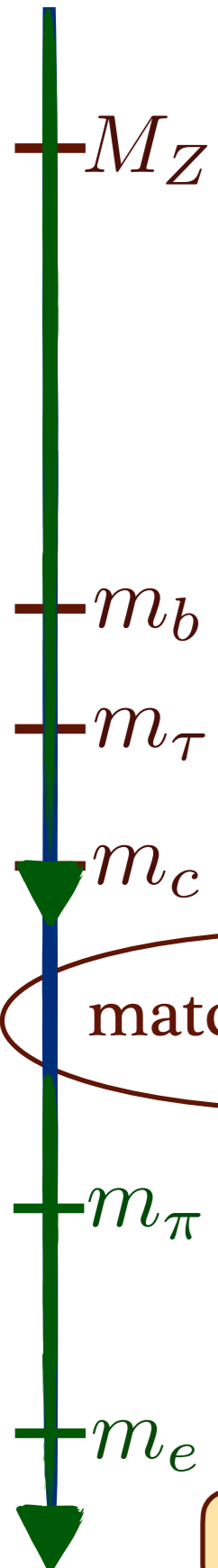
explain \sim half of radiative correction to **neutron lifetime**

$$\Delta = 7.756(27)\%$$

m_e

- g_V, g_A encode electroweak logarithms and hadronic corrections

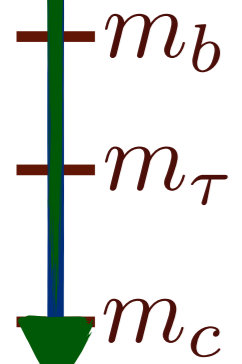
Effect of running



$$\mathcal{L}_{\text{LEFT}} \sim G_F V_{ud} C_\beta^r(a, \mu) \bar{e}_L \gamma_\rho \nu_{eL} \bar{u}_L \gamma^\rho d_L$$

- renormalization group equations

$$\mu \frac{dC_\beta^r(a, \mu)}{d\mu} = \left(\gamma_0 \frac{\alpha}{\pi} + \gamma_1 \left(\frac{\alpha}{\pi} \right)^2 + \gamma_{se} \frac{\alpha}{\pi} \frac{\alpha_s}{4\pi} \right) C_\beta^r(a, \mu)$$



$$\mu \frac{dg_V(\mu)}{d\mu} = \left(\tilde{\gamma}_0 \frac{\alpha}{\pi} + \tilde{\gamma}_1 \left(\frac{\alpha}{\pi} \right)^2 \right) g_V(\mu)$$

- resummed result with leading $\alpha\alpha_s$ corrections

matching scale

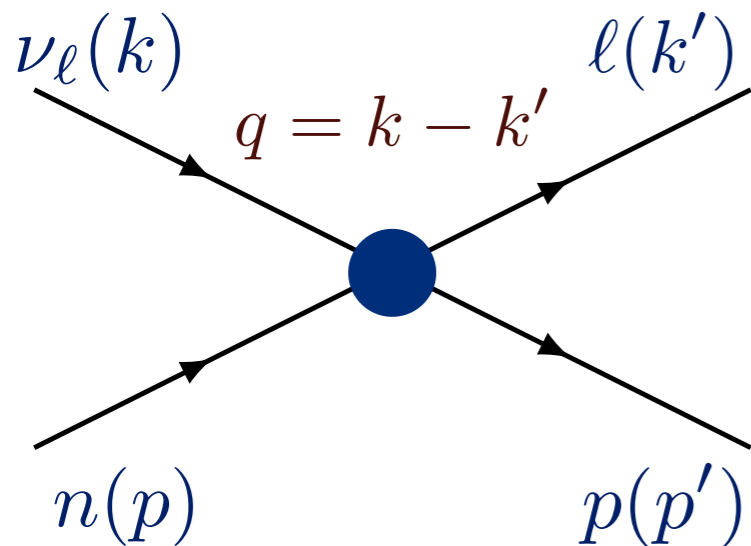
$$g_V(\mu_\chi = m_e) - 1 = (2.499 \pm 0.013) \% \text{ full error}$$

- one-loop result without $\alpha\alpha_s$ contributions

$$g_V^{1\text{-loop}}(\mu_\chi = m_e) - 1 = (2.430 \pm 0.012) \% \text{ hadronic error}$$

- cancellation of various contributions \rightarrow 0.07 % effect

Invariant amplitudes



averaged lepton momentum

$$K_\mu = \frac{k_\mu + k'_\mu}{2}$$

averaged nucleon momentum

$$P_\mu = \frac{p_\mu + p'_\mu}{2}$$

- **four amplitudes** for massless charged lepton

$$T_{\nu_\ell n \rightarrow \ell^- p}^{m_\ell=0} = \sqrt{2}G_F V_{ud} \bar{\ell}^- \gamma^\mu P_L \nu_\ell \bar{p} \left(\gamma_\mu (g_M + f_A \gamma_5) - (f_2 + f_A^3 \gamma_5) \frac{K_\mu}{M} \right) n$$

- **four extra amplitudes** for massive charged lepton

$$T_{\nu_\ell n \rightarrow \ell^- p}^{m_\ell \neq 0} = -\sqrt{2}G_F V_{ud} \frac{m_\ell}{M} \bar{\ell}^- P_L \nu_\ell \bar{p} \left(f_3 + f_P \gamma_5 - \frac{f_R}{4} \frac{P_\mu}{M} \gamma^\mu \gamma_5 \right) n$$

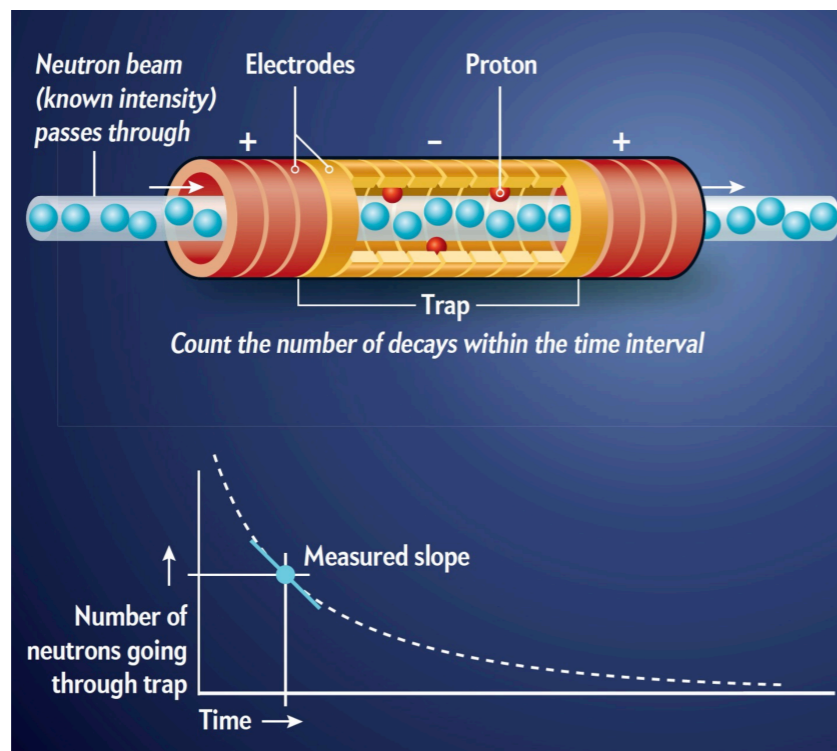
$$+ \sqrt{2}G_F V_{ud} \frac{m_\ell}{M} \frac{f_T}{4} \bar{\ell}^- \sigma^{\mu\nu} P_L \nu_\ell \bar{p} \sigma_{\mu\nu} n$$

- **8 invariant amplitudes** for charged-current elastic scattering

Neutron lifetime measurements



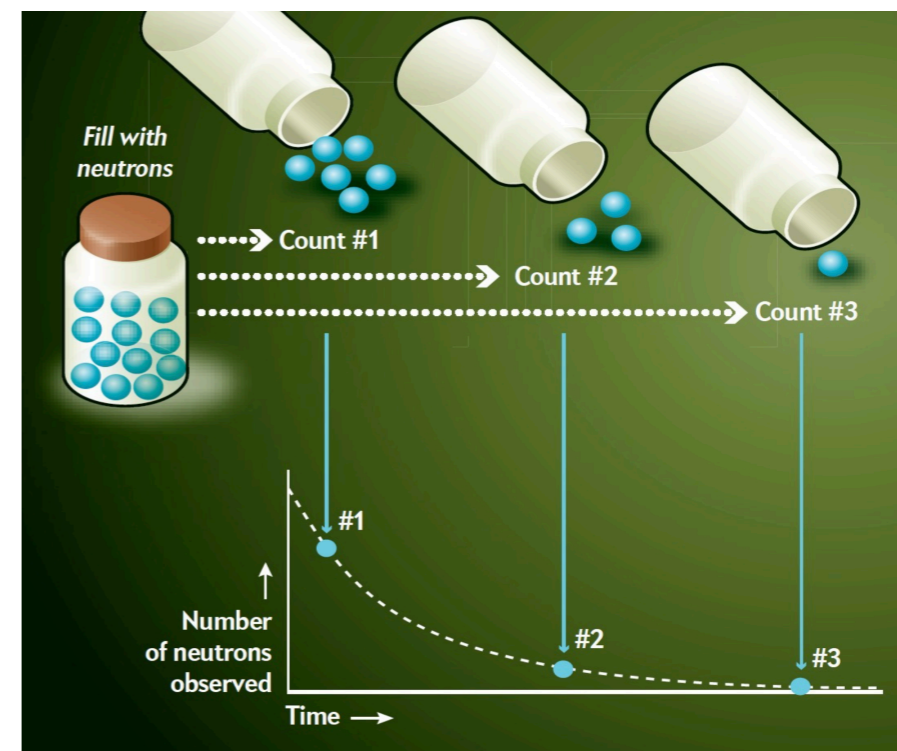
beam method



how many neutrons pass?

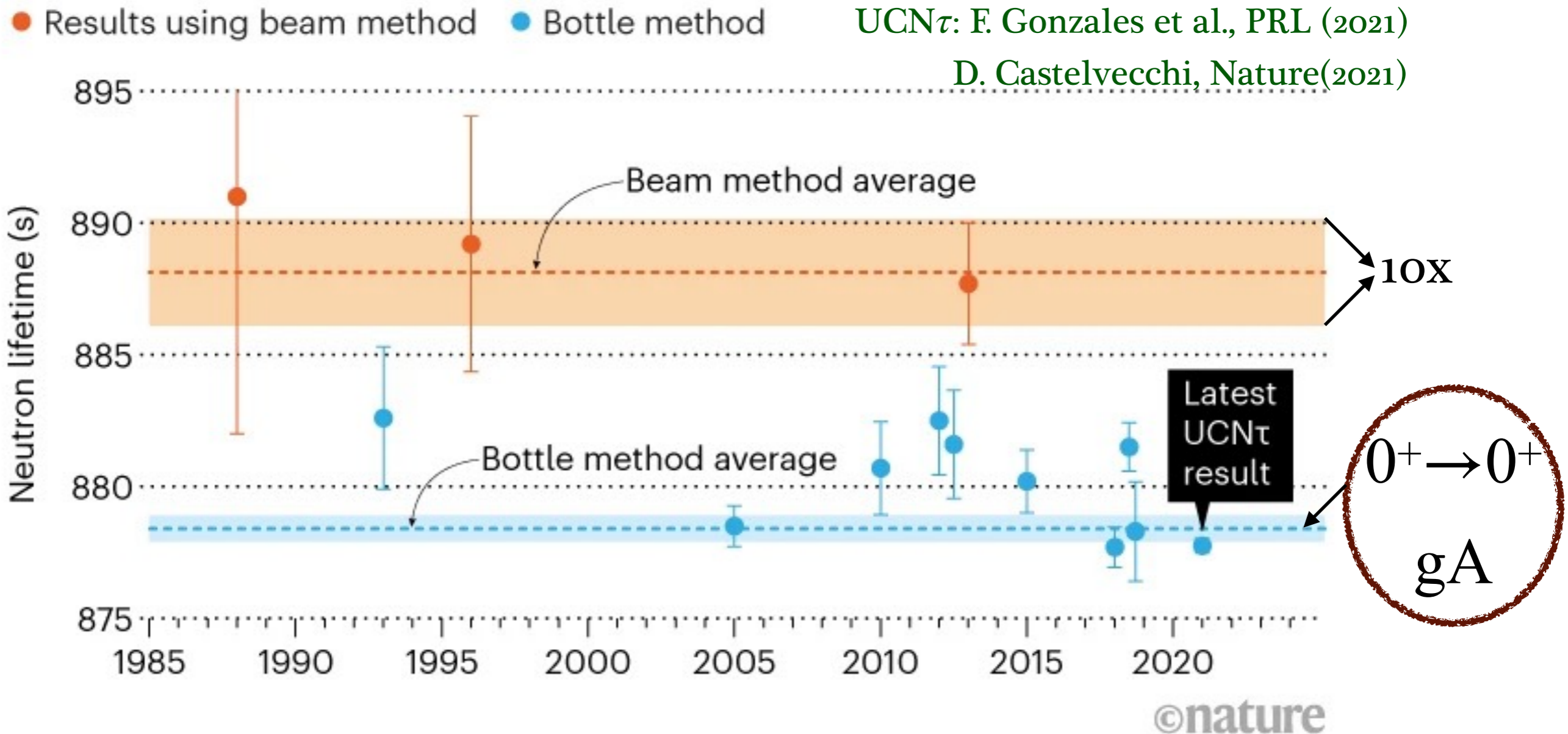


bottle method



how many neutrons survive?

Neutron lifetime puzzle

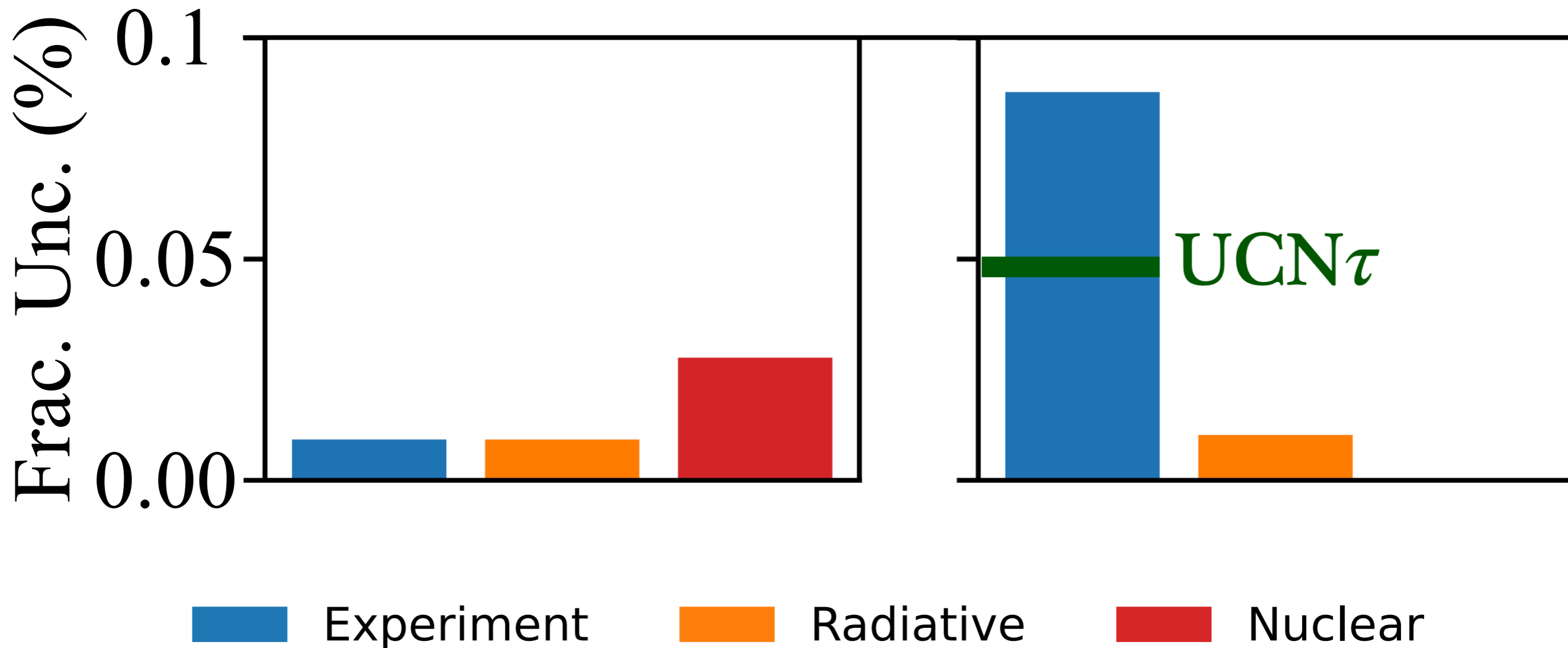


- 8-9 seconds discrepancy beam vs bottle method : $3-5\sigma$
- V_{ud} from $0^+ \rightarrow 0^+ + gA$ from β asymmetry \equiv bottle
- 0.3 seconds uncertainty of UCN τ @LANL : $(3 - 4) \times 10^{-4}$

V_{ud} determinations

$0^+ \rightarrow 0^+$

Neutron

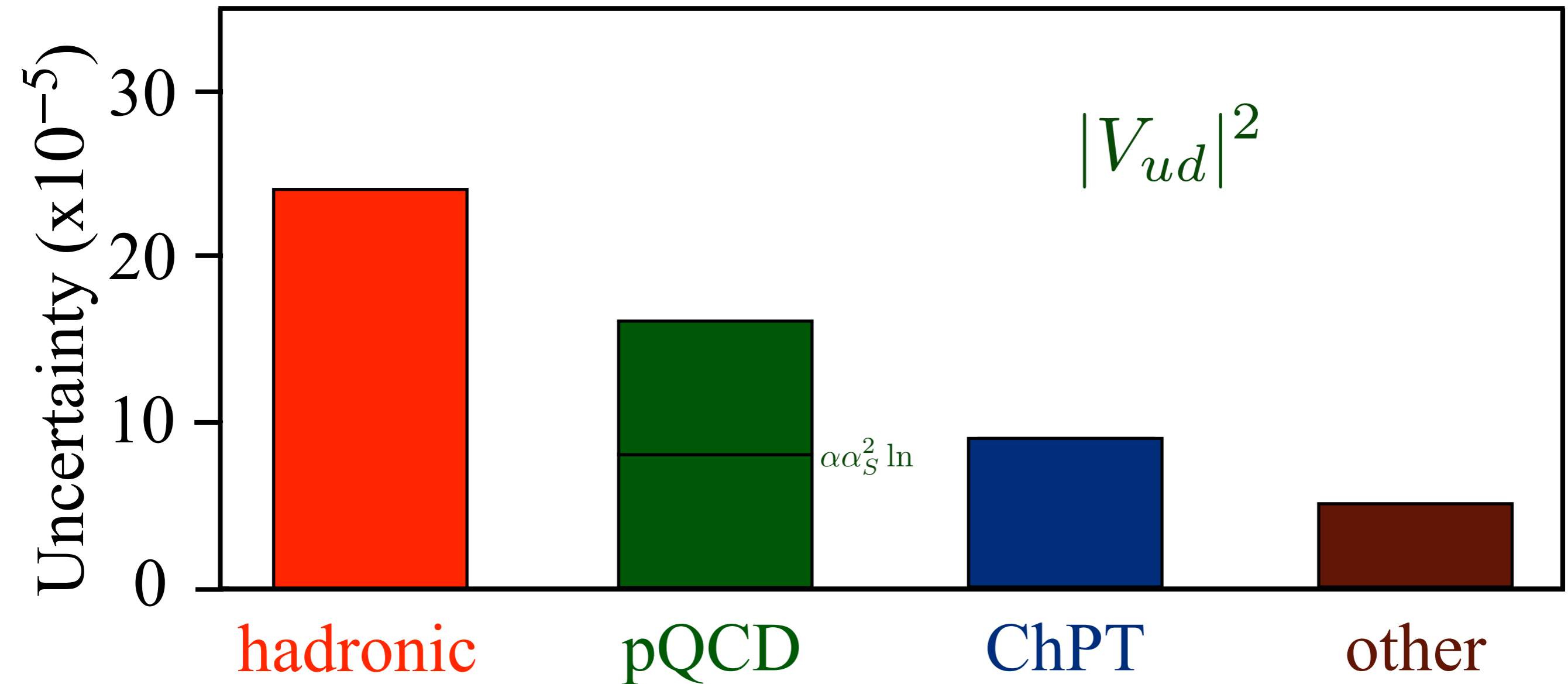


2022 Fundamental Symmetries, Neutrons, and Neutrinos (FSNN) white paper

- neutron decay becomes competitive with $0^+ \rightarrow 0^+$

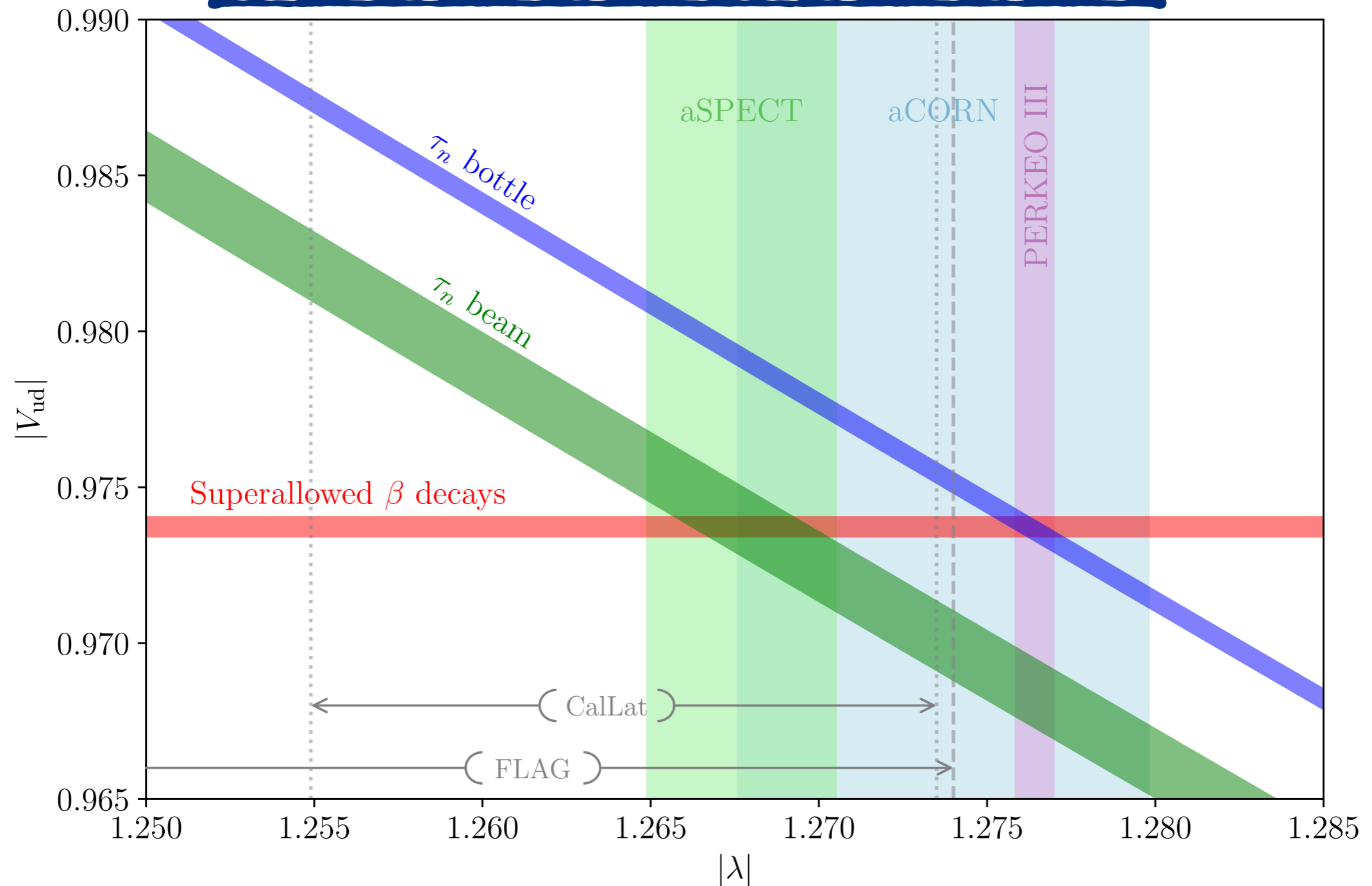
Our work

- detailed error budget and perturbative corrections



- shift of central value by ~ 1 error of radiative corrections
- quantification of perturbative/nonperturbative uncertainties

Neutron lifetime puzzle

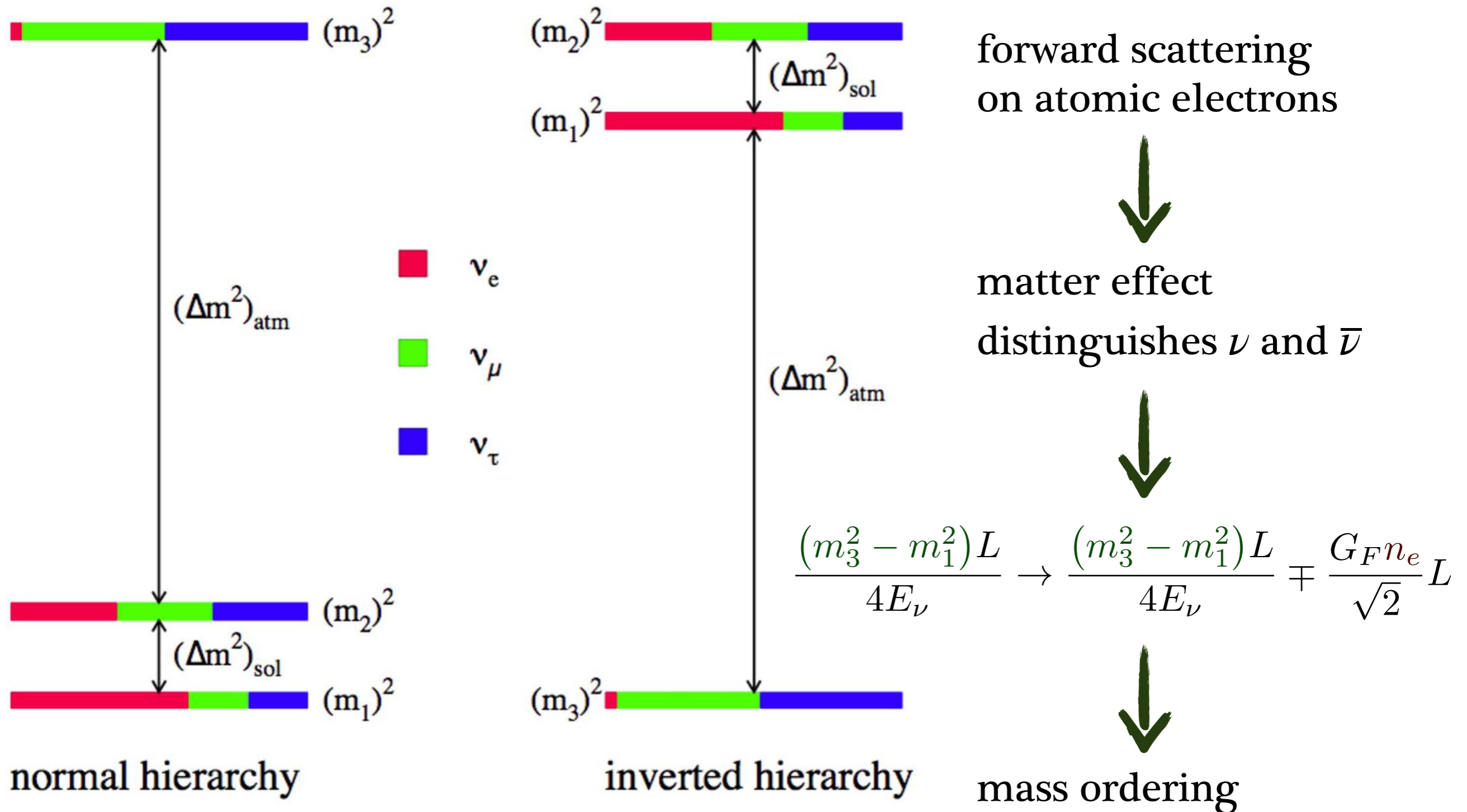


Andrzej Czarnecki, William J. Marciano, Alberto Sirlin, PRL (2018)

Susan Gardner, Mohammadreza Zakeri, Universe (2024)

- tensions in g_A data: puzzle is not resolved

Neutrino mass hierarchy



- matter effects can resolve neutrino mass hierarchy