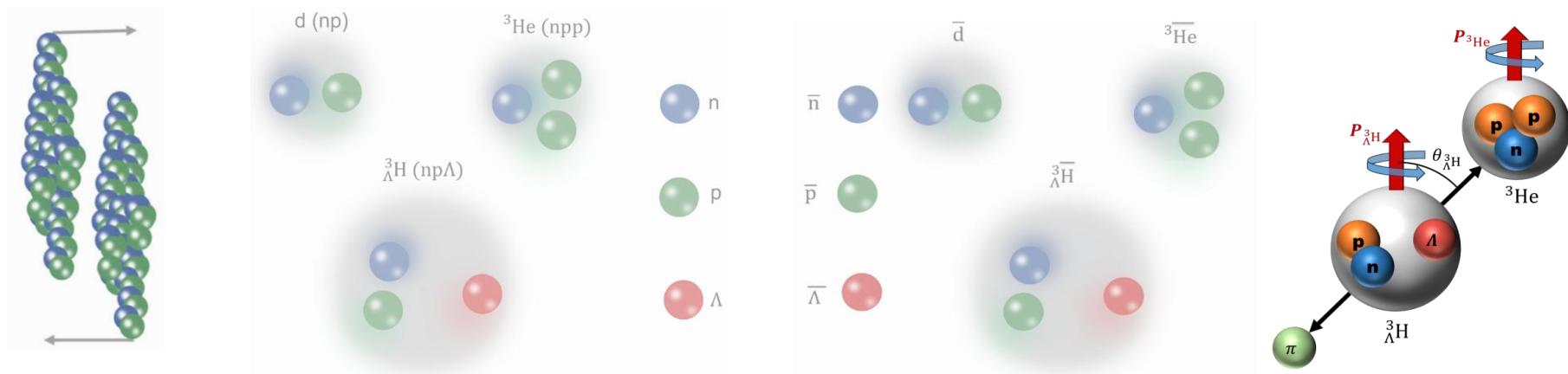


Production and Polarization of Hypernuclei in Heavy-Ion Collisions

KJ Sun, Dai-Neng Liu, Yun-Peng Zhen, Jin-Hui Chen, Che Ming Ko, Yu-Gang Ma [arXiv:2405.12015](https://arxiv.org/abs/2405.12015)



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Oct. 24, 2024

Outline

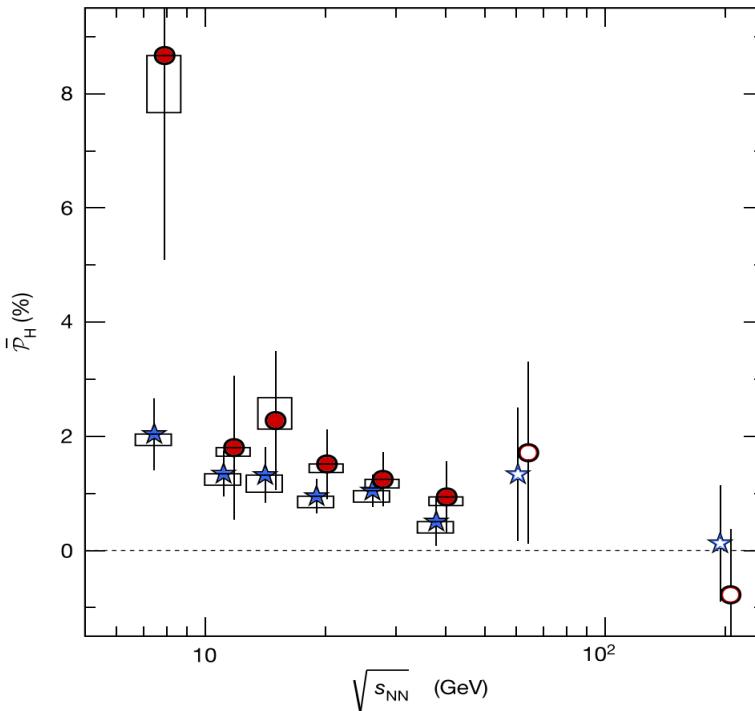
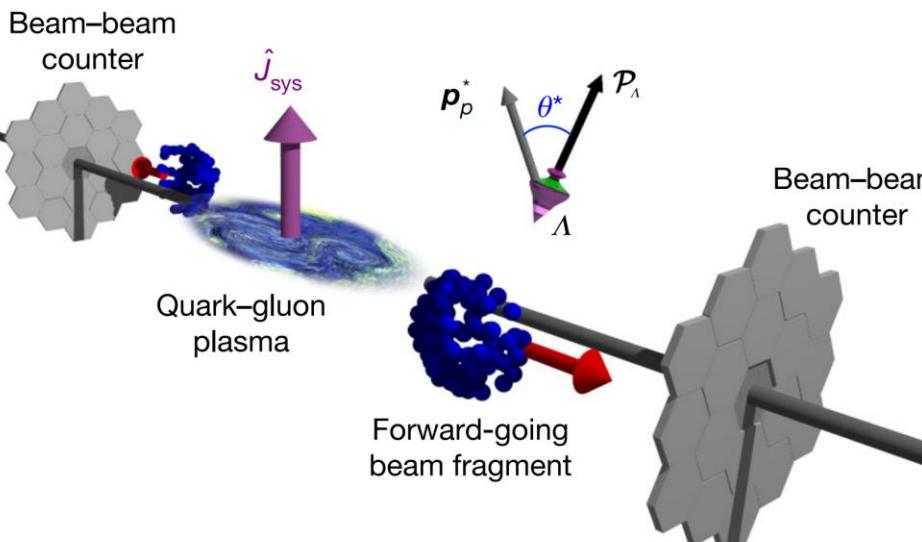
- 1. Polarization in heavy-ion collisions**
- 2. The property and production of (anti-)hypertriton**
- 3. Results: (Anti-)Hypertriton polarization and its spin structure**
- 4. Discussions: Effects of baryon spin correlations**
- 5. Summary and outlook**

1. Polarization of hadrons in relativistic heavy-ion collisions (1)

STAR, Nature 548, 62 (2017)

Z. T. Liang and X. N. Wang PRL 94, 102301 (2005)

F. Becattini, F. Piccinini, and J. Rizzo, PRC 77, 024906 (2008)



$$\frac{dN}{d \cos \theta^*} = \frac{1}{2} (1 + \alpha_H |\mathcal{P}_H| \cos \theta^*)$$

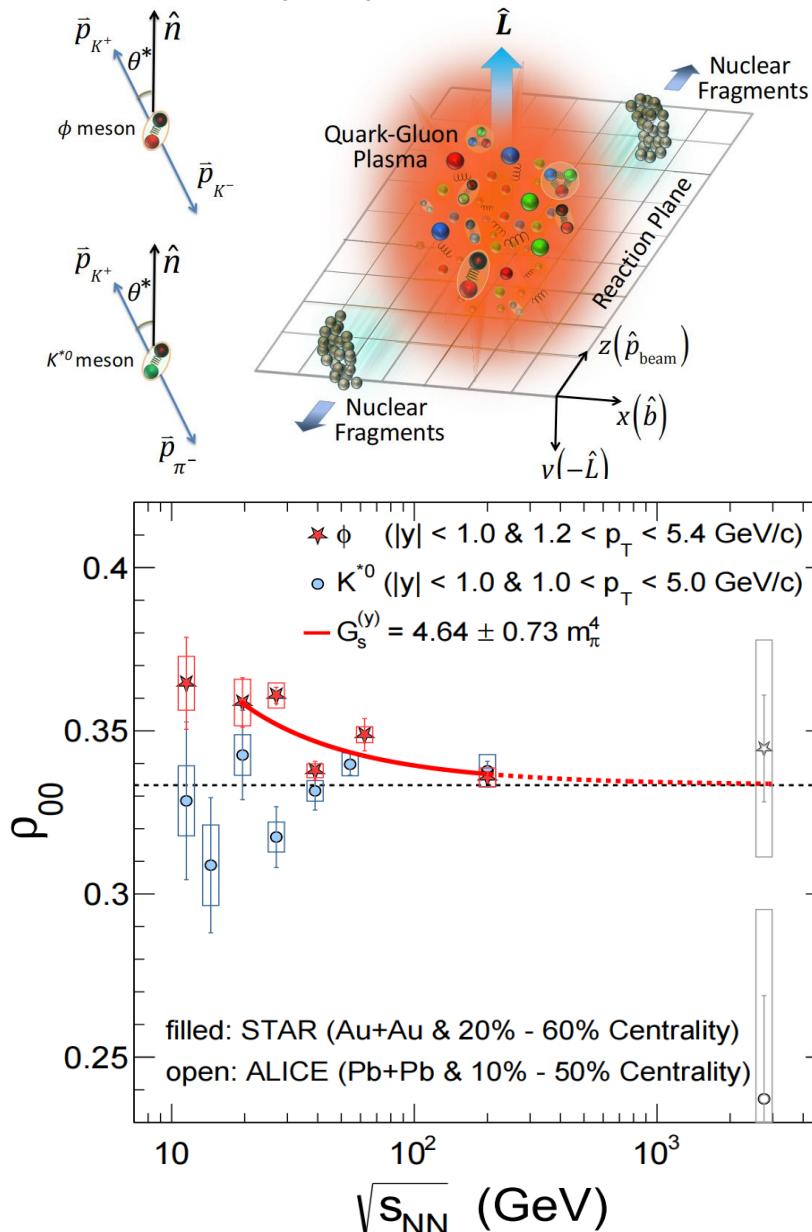
$$P_H = \frac{8}{\pi \alpha_H} \frac{\langle \sin(\Psi_1^{\text{obs}} - \phi_p^*) \rangle}{\text{Res}(\Psi_1)}$$

$$\omega \approx k_B T (\bar{\mathcal{P}}_\Lambda' + \bar{\mathcal{P}}_{\bar{\Lambda}}') / \hbar$$

Spin polarization of Lambda hyperon → Vorticity of QGP

1. Polarization of hadrons in relativistic heavy-ion collisions (2)

STAR, Nature 614, 7947 (2023)



Spin alignment of mesons

$$\frac{dN}{d(\cos\theta^*)} \propto (1 - \rho_{00}) + (3\rho_{00} - 1)\cos^2\theta^*$$



Fluctuation/correlation of strong force field

X. L. Sheng et al., PRL 131, 042304 (2023)

$$G_s^{(y)} \equiv g_\phi^2 \left[3\langle B_{\phi,y}^2 \rangle + \frac{\langle \mathbf{p}^2 \rangle_\phi}{m_s^2} \langle E_{\phi,y}^2 \rangle - \frac{3}{2} \langle B_{\phi,x}^2 + B_{\phi,z}^2 \rangle - \frac{\langle \mathbf{p}^2 \rangle_\phi}{2m_s^2} \langle E_{\phi,x}^2 + E_{\phi,z}^2 \rangle \right]$$

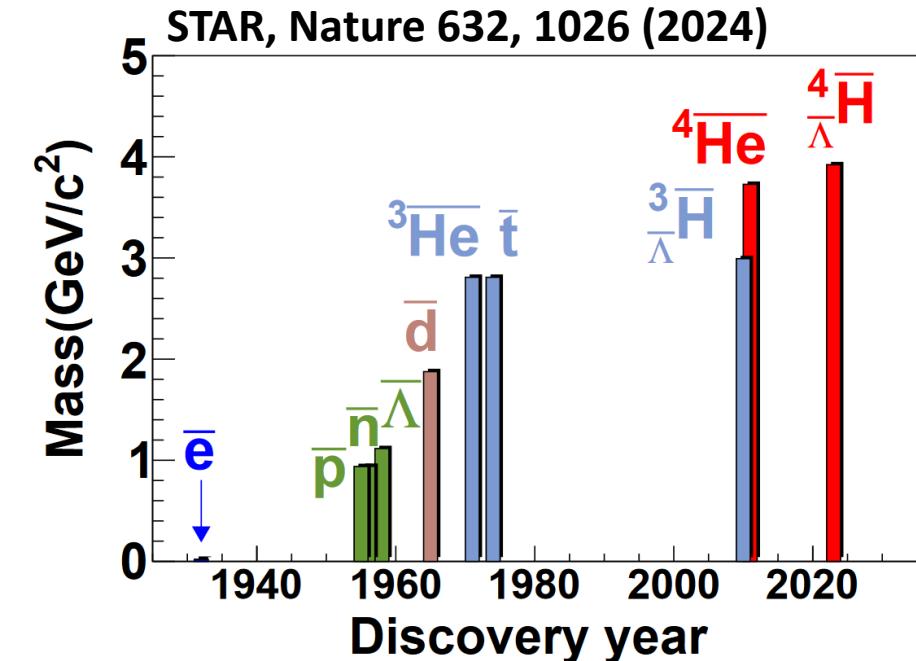
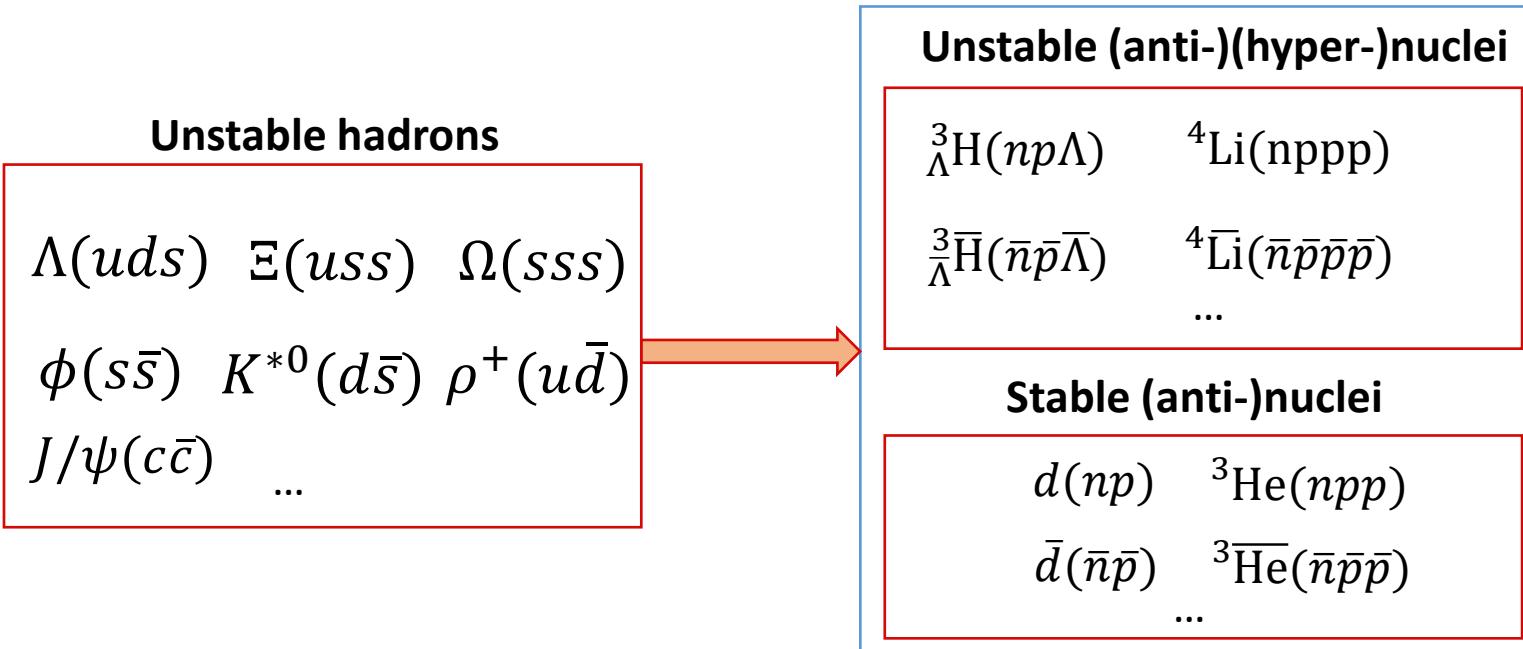
Quark-antiquark spin correlation

J. P. Lv et al., arXiv:2402.13721

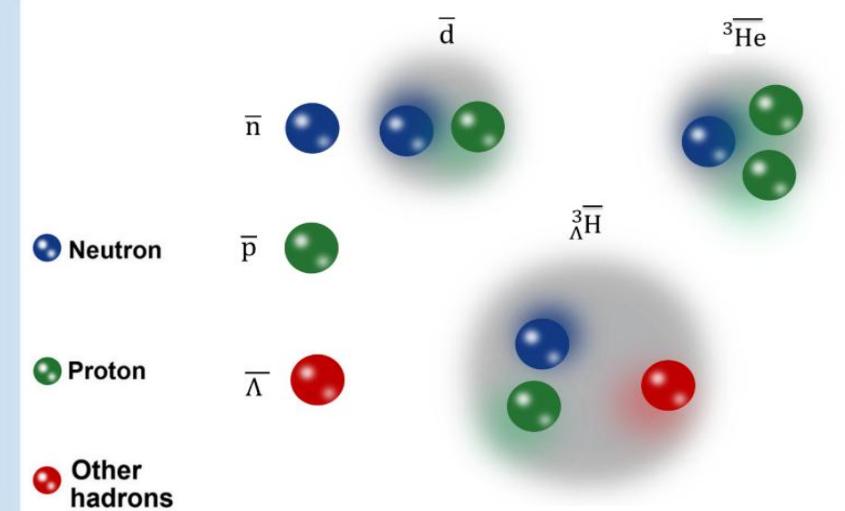
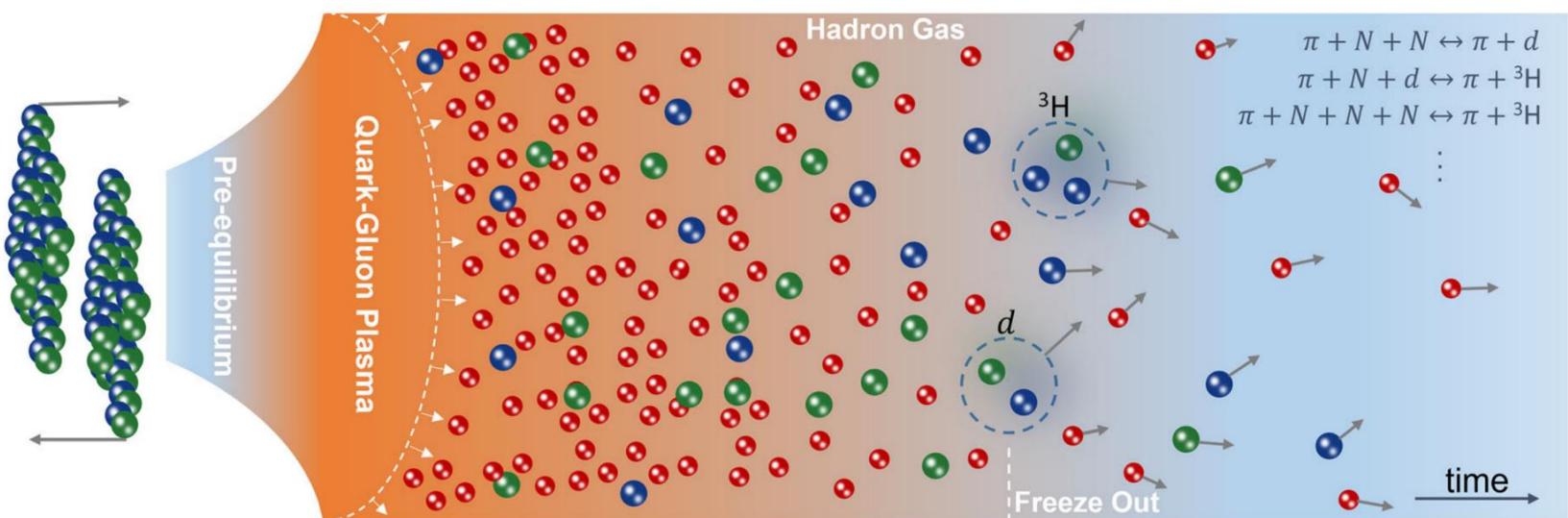
Meson spectral property

F. Li and S. Liu, arXiv:2206.11890

1. Polarization of light (anti-)(hyper-)nuclei (3)



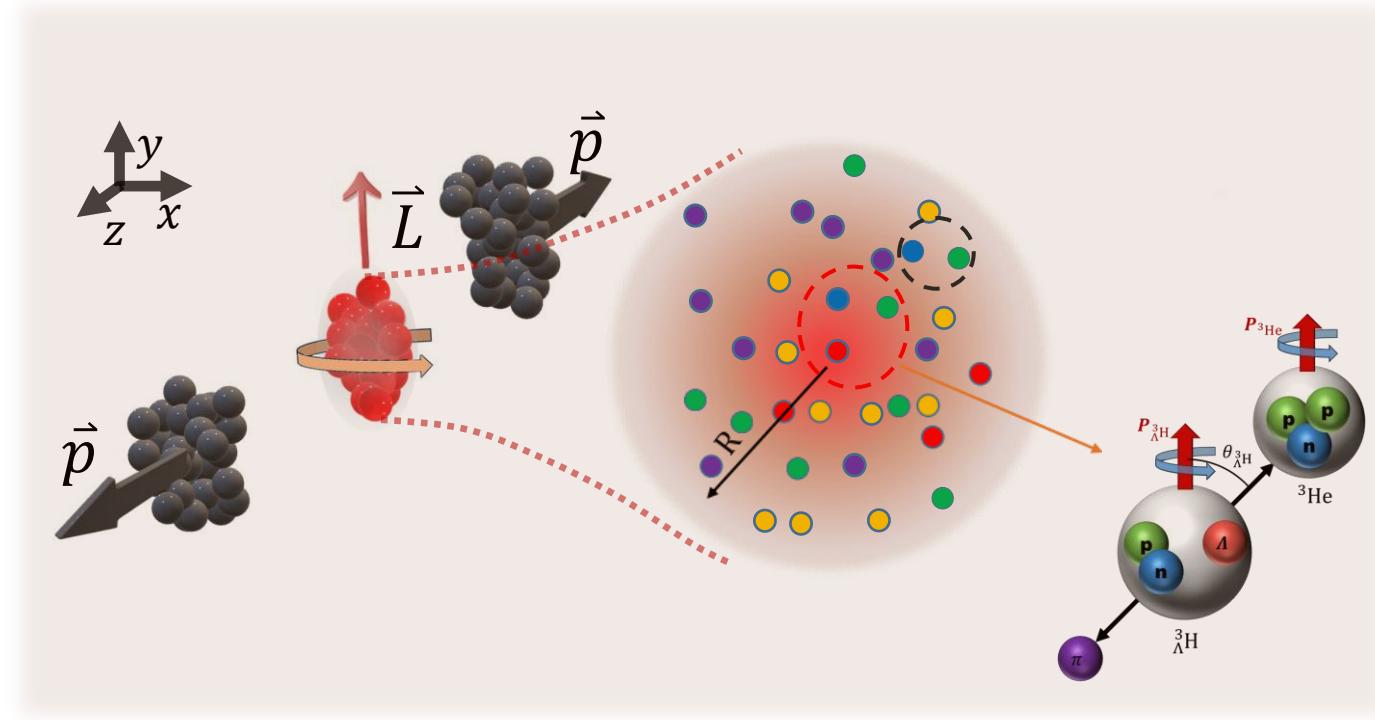
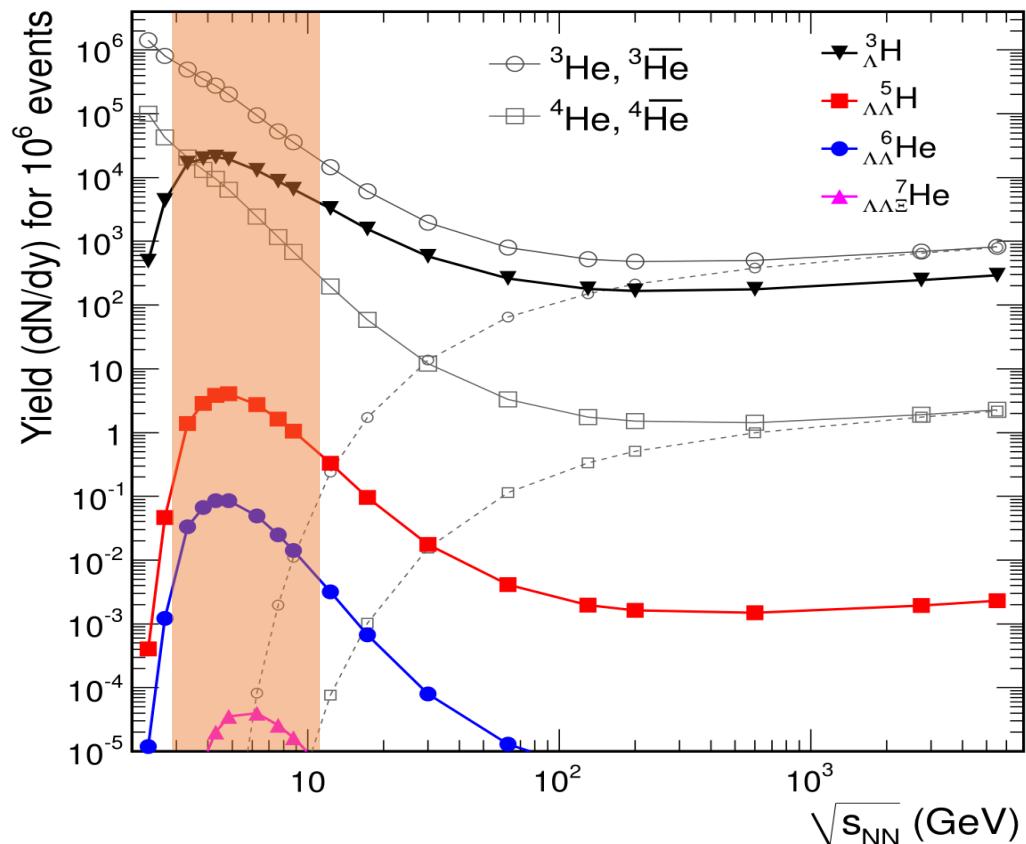
K. J. Sun, R. Wang, C. M. Ko, Y. G. Ma, C. Shen, Nature Commun. 15, 1074 (2024)



1. Polarization of light (anti-)(hyper-)nuclei

(4)

A. Andronic et al., Phys. Lett. B 697, 203-207 (2011)

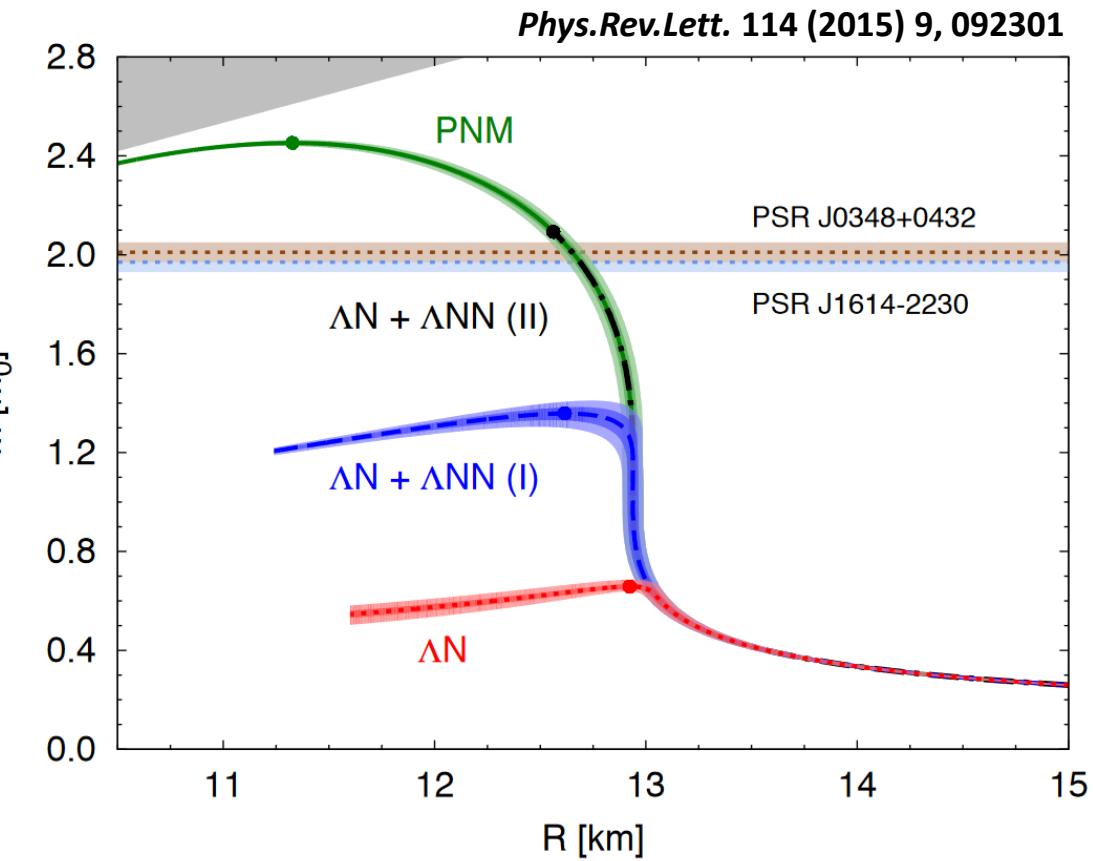
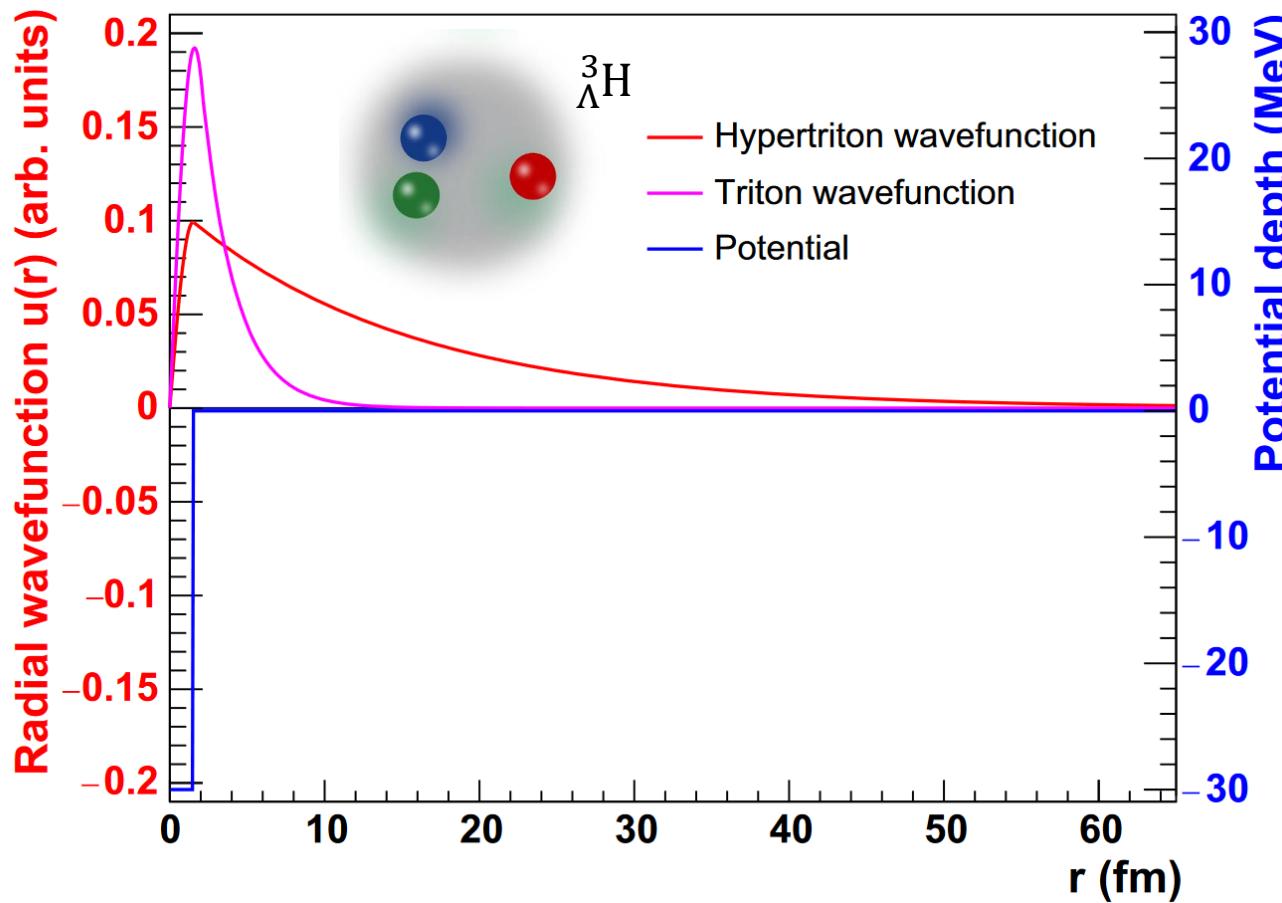


- FAIR/CBM (2.4-4.9 GeV)
- HIAF/CEE (2.1-4.5 GeV)
- NICA/MPD (4-11 GeV)

A novel tool to study the evolution of strongly-interacting matter at high-baryon density region

2. The halo-like nucleus: (anti-)hypertriton

(5)

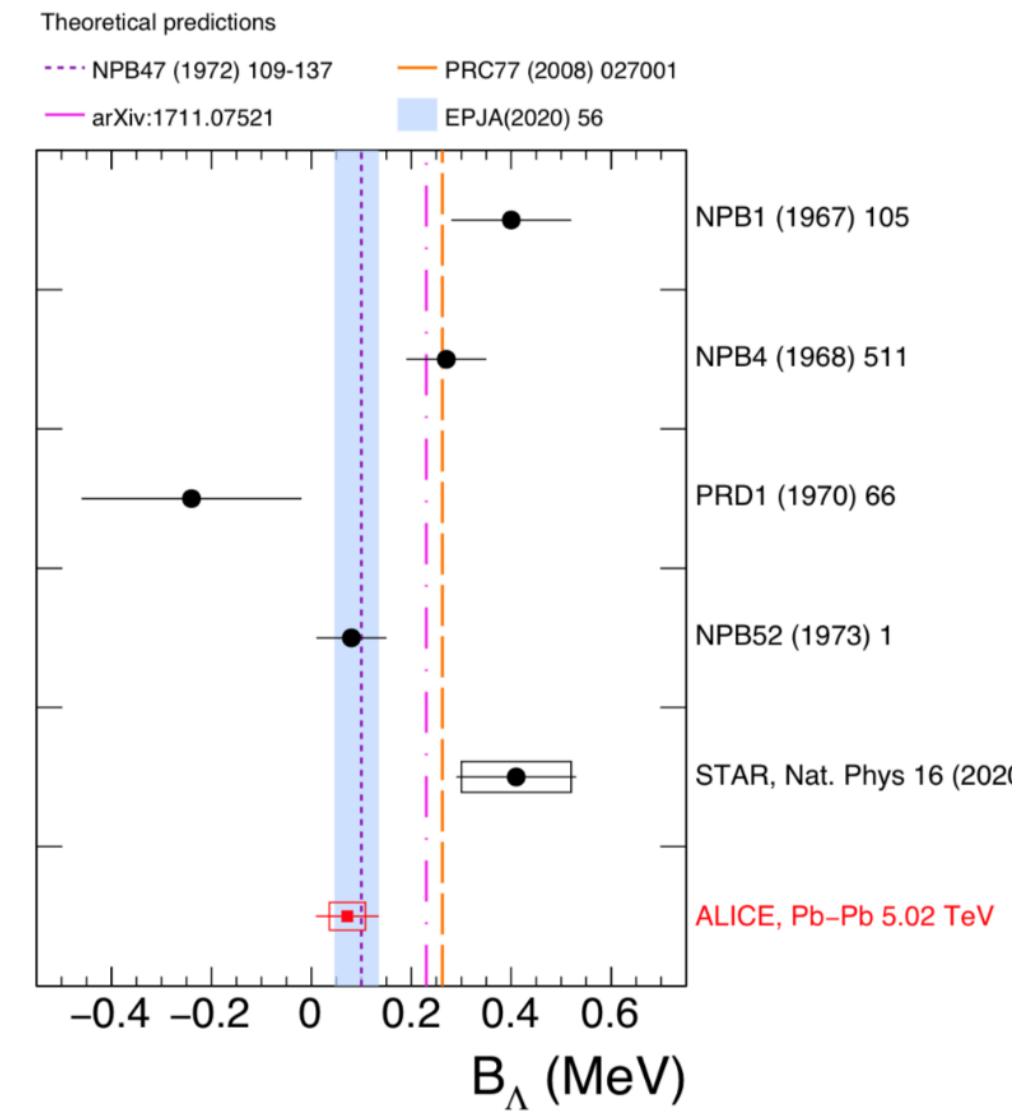
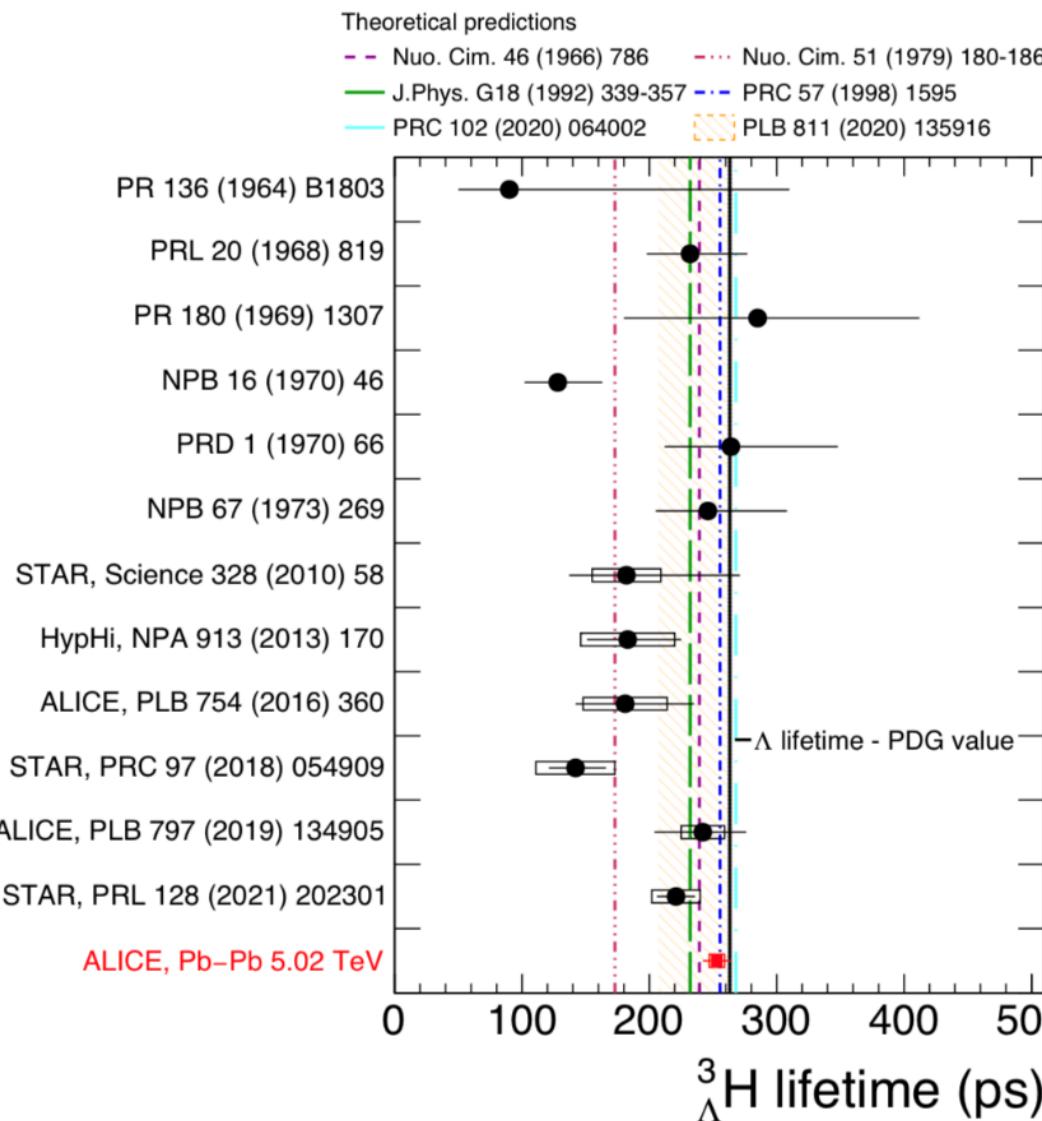


2. Binding energy and lifetime

(6)

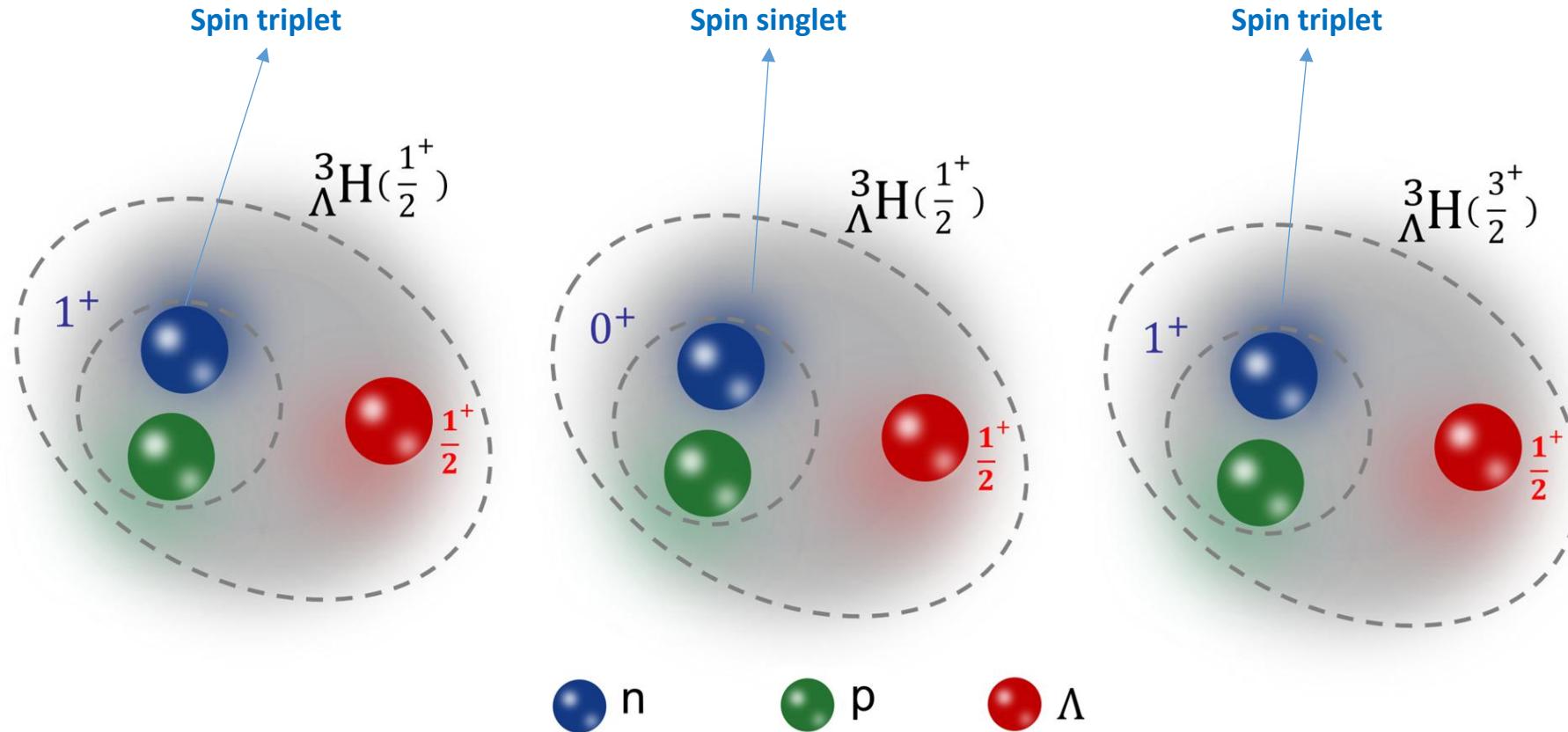
ALICE, PRL 131, 102302 (2023)

Y. G. Ma, Nucl. Sci. Tech. 3497 (2023)



2. Spin of (anti-)hypertriton ?

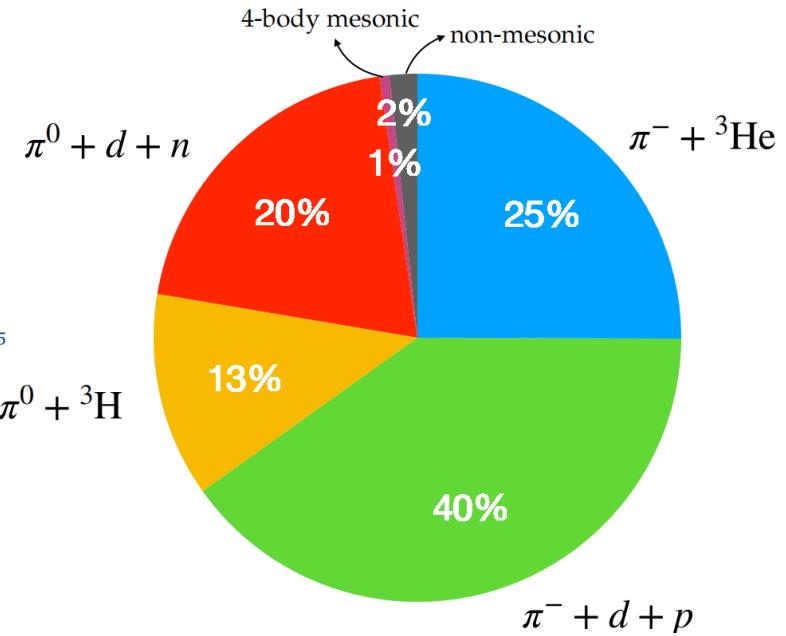
(7)



2. Spin of (anti-)hypertriton ?

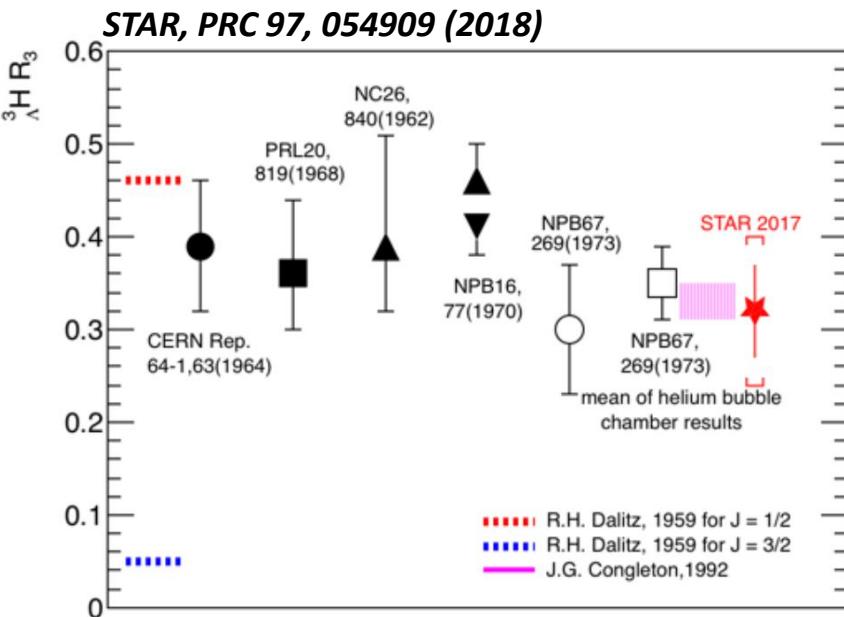
(8)

$$\begin{aligned} {}^3_{\Lambda}\text{H} &\rightarrow \pi^- + {}^3\text{He}, & {}^3_{\Lambda}\text{H} &\rightarrow \pi^0 + {}^3\text{H}, \\ {}^3_{\Lambda}\text{H} &\rightarrow \pi^- + d + p, & {}^3_{\Lambda}\text{H} &\rightarrow \pi^0 + d + n, \\ {}^3_{\Lambda}\text{H} &\rightarrow \pi^- + p + n + p, & {}^3_{\Lambda}\text{H} &\rightarrow \pi^0 + p + n + n. \end{aligned}$$



Relative branching ratio:

$$R_3 = \frac{\text{B.R.}({}^3_{\Lambda}\text{H} \rightarrow {}^3\text{He}\pi^-)}{\text{B.R.}({}^3_{\Lambda}\text{H} \rightarrow {}^3\text{He}\pi^-) + \text{B.R.}({}^3_{\Lambda}\text{H} \rightarrow d\pi^-)}$$



Favors spin 1/2

PHYSICAL REVIEW D 87, 034506 (2013)

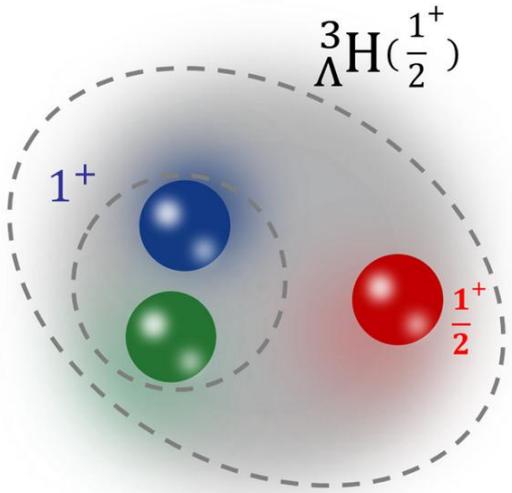
Light nuclei and hypernuclei from quantum chromodynamics in the limit of SU(3) flavor symmetry

S. R. Beane,¹ E. Chang,² S. D. Cohen,³ W. Detmold,^{4,5} H. W. Lin,³ T. C. Luu,⁶ K. Orginos,^{4,5} A. Parreño,² M. J. Savage,³ and A. Walker-Loud^{7,8}

| Label | <i>A</i> | <i>s</i> | <i>I</i> | <i>J</i> ^{<i>π</i>} | Local SU(3) irreps | This work |
|--|----------|----------|----------|------------------------------|------------------------------|----------------------|
| <i>N</i> | 1 | 0 | 1/2 | 1/2 ⁺ | 8 | 8 |
| <i>Λ</i> | 1 | -1 | 0 | 1/2 ⁺ | 8 | 8 |
| <i>Σ</i> | 1 | -1 | 1 | 1/2 ⁺ | 8 | 8 |
| <i>Ξ</i> | 1 | -2 | 1/2 | 1/2 ⁺ | 8 | 8 |
| <i>d</i> | 2 | 0 | 0 | 1 ⁺ | 10 | 10 |
| <i>nn</i> | 2 | 0 | 1 | 0 ⁺ | 27 | 27 |
| <i>nΛ</i> | 2 | -1 | 1/2 | 0 ⁺ | 27 | 27 |
| <i>nΛ</i> | 2 | -1 | 1/2 | 1 ⁺ | 8_A, 10 | - |
| <i>n\Sigma</i> | 2 | -1 | 3/2 | 0 ⁺ | 27 | 27 |
| <i>n\Sigma</i> | 2 | -1 | 3/2 | 1 ⁺ | 10 | 10 |
| <i>n\xi</i> | 2 | -2 | 0 | 1 ⁺ | 8_A | 8_A |
| <i>n\xi</i> | 2 | -2 | 1 | 1 ⁺ | 8_A, 10, 10 | - |
| <i>H</i> | 2 | -2 | 0 | 0 ⁺ | 1, 27 | 1, 27 |
| ${}^3\text{H}, {}^3\text{H}$ | 3 | 0 | 1/2 | 1/2 ⁺ | 35 | 35 |
| ${}^3\text{H}(1/2^+)$ | 3 | -1 | 0 | 1/2 ⁺ | 35 | - |
| ${}^3_{\Lambda}\text{H}(3/2^+)$ | 3 | -1 | 0 | 3/2 ⁺ | 10 | 10 |
| ${}^3_{\Lambda}\text{He}, {}^3_{\Lambda}\text{H}, nn\Lambda$ | 3 | -1 | 1 | 1/2 ⁺ | 27, 35 | 27, 35 |
| ${}^3_{\Lambda}\text{He}$ | 3 | -1 | 1 | 3/2 ⁺ | 27 | 27 |
| ${}^4\text{He}$ | 4 | 0 | 0 | 0 ⁺ | 28 | 28 |
| ${}^4_{\Lambda}\text{He}$ | 4 | -1 | 1/2 | 0 ⁺ | 28 | - |
| ${}^4_{\Lambda\Lambda}\text{He}$ | 4 | -2 | 1 | 0 ⁺ | 27, 28 | 27, 28 |
| $\Lambda\Xi^0 pnn$ | 5 | -3 | 0 | 3/2 ⁺ | 10 + ⋯ | 10 |

Favors spin 3/2

3. (Anti-)hypertriton polarization and its spin structure (9)



$$\begin{aligned} | \frac{1}{2}, \uparrow \rangle_{{}^3\Lambda H} = & \frac{\sqrt{6}}{3} | \frac{1}{2}, \frac{1}{2} \rangle_n | \frac{1}{2}, \frac{1}{2} \rangle_p | \frac{1}{2}, -\frac{1}{2} \rangle_\Lambda \\ & - \frac{\sqrt{6}}{6} (| \frac{1}{2}, \frac{1}{2} \rangle_n | \frac{1}{2}, -\frac{1}{2} \rangle_p | \frac{1}{2}, \frac{1}{2} \rangle_\Lambda \\ & + | \frac{1}{2}, -\frac{1}{2} \rangle_n | \frac{1}{2}, \frac{1}{2} \rangle_p | \frac{1}{2}, \frac{1}{2} \rangle_\Lambda), \end{aligned}$$

$$\begin{aligned} | \frac{1}{2}, \downarrow \rangle_{{}^3\Lambda H} = & -\frac{\sqrt{6}}{3} | \frac{1}{2}, -\frac{1}{2} \rangle_n | \frac{1}{2}, -\frac{1}{2} \rangle_p | \frac{1}{2}, \frac{1}{2} \rangle_\Lambda \\ & + \frac{\sqrt{6}}{6} (| \frac{1}{2}, \frac{1}{2} \rangle_n | \frac{1}{2}, -\frac{1}{2} \rangle_p | \frac{1}{2}, -\frac{1}{2} \rangle_\Lambda \\ & + | \frac{1}{2}, -\frac{1}{2} \rangle_n | \frac{1}{2}, \frac{1}{2} \rangle_p | \frac{1}{2}, -\frac{1}{2} \rangle_\Lambda). \end{aligned}$$

Coalescence model for hypertriton production (without baryon spin correlation)

$$E_i \frac{d^3 N_{i,\pm\frac{1}{2}}}{d\mathbf{p}_i^3} = \int_{\Sigma^\mu} d^3 \sigma_\mu p_i^\mu w_{i,\pm\frac{1}{2}}(\mathbf{x}_i, \mathbf{p}_i) \bar{f}_i(\mathbf{x}_i, \mathbf{p}_i)$$

$$w_{i,\pm\frac{1}{2}} = \frac{1}{2} [1 \pm \mathcal{P}_i(\mathbf{x}_i, \mathbf{p}_i)]$$

$$\hat{\rho}_i = \text{diag} \left(\frac{1+\mathcal{P}_i}{2}, \frac{1-\mathcal{P}_i}{2} \right)$$

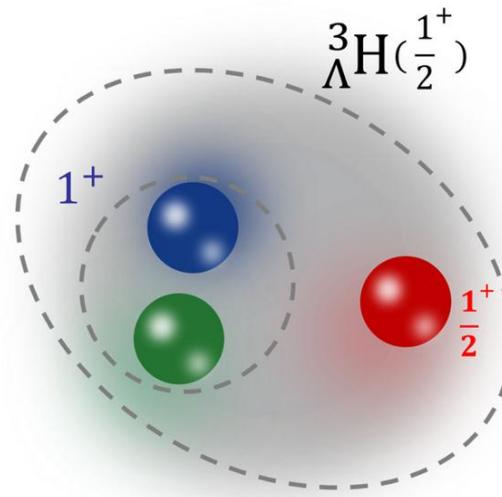
$$\bar{f}_i = \frac{g_i}{(2\pi)^3} \left[\exp(p_i^\mu u_\mu/T)/\xi_i + 1 \right]^{-1}$$

$$\hat{\rho}_{np\Lambda} = \hat{\rho}_n \otimes \hat{\rho}_p \otimes \hat{\rho}_\Lambda$$

$$\begin{aligned} E \frac{d^3 N_{{}^3\Lambda H, \pm\frac{1}{2}}}{d\mathbf{P}^3} = & E \int \prod_{i=n,p,\Lambda} p_i^\mu d^3 \sigma_\mu \frac{d^3 p_i}{E_i} \bar{f}_i(\mathbf{x}_i, \mathbf{p}_i) \\ & \times \left(\frac{2}{3} w_{n,\pm\frac{1}{2}} w_{p,\pm\frac{1}{2}} w_{\Lambda,\mp\frac{1}{2}} + \frac{1}{6} w_{n,\pm\frac{1}{2}} w_{p,\mp\frac{1}{2}} w_{\Lambda,\pm\frac{1}{2}} \right. \\ & \left. + \frac{1}{6} w_{n,\mp\frac{1}{2}} w_{p,\pm\frac{1}{2}} w_{\Lambda,\pm\frac{1}{2}} \right) \\ & \times W_{{}^3\Lambda H}(\mathbf{x}_n, \mathbf{x}_p, \mathbf{x}_\Lambda; \mathbf{p}_n, \mathbf{p}_p, \mathbf{p}_\Lambda) \delta(\mathbf{P} - \sum_i \mathbf{p}_i) \end{aligned}$$

3. (Anti-)hypertriton polarization and its spin structure

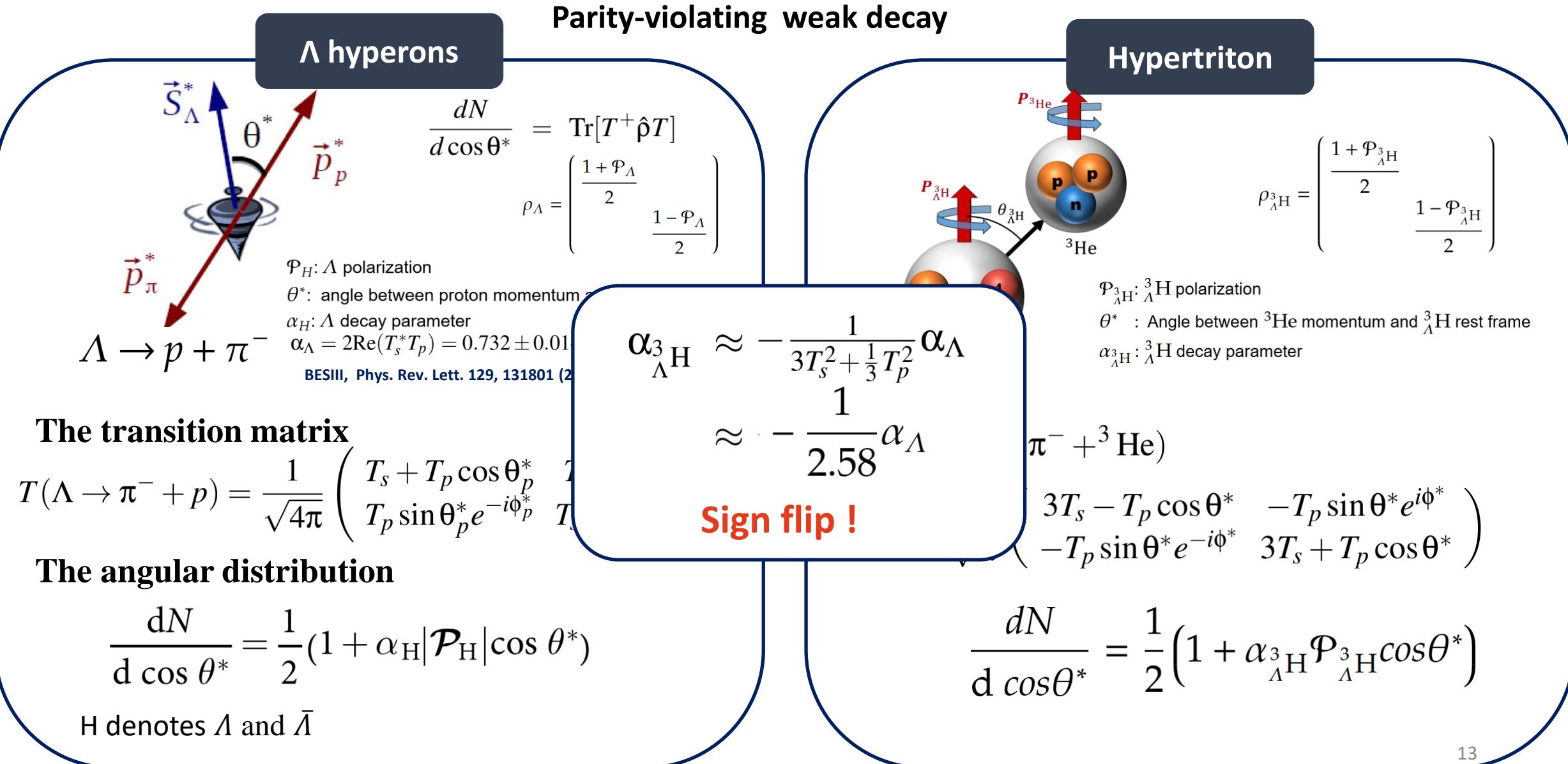
(10)



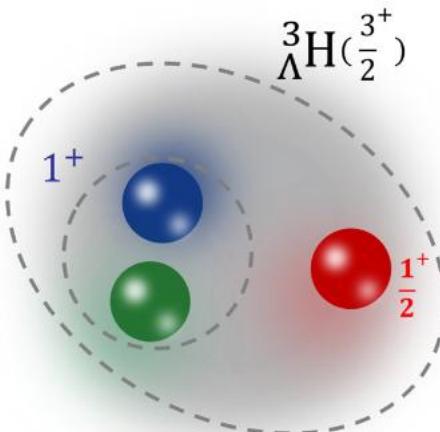
$$\begin{aligned}\mathcal{P}_{{}^3_{\Lambda}H} &\approx \frac{\frac{2}{3}\mathcal{P}_n + \frac{2}{3}\mathcal{P}_p - \frac{1}{3}\mathcal{P}_{\Lambda} - \mathcal{P}_n\mathcal{P}_p\mathcal{P}_{\Lambda}}{1 - \frac{2}{3}(\mathcal{P}_n + \mathcal{P}_p)\mathcal{P}_{\Lambda} + \frac{1}{3}\mathcal{P}_n\mathcal{P}_p} \\ &\approx \frac{2}{3}\mathcal{P}_n + \frac{2}{3}\mathcal{P}_p - \frac{1}{3}\mathcal{P}_{\Lambda} \\ &\approx \mathcal{P}_{\Lambda}\end{aligned}$$

3. (Anti-)hypertriton polarization and its spin structure (11)

K. J. Sun et al., arXiv:2405.12015(2024)



3. (Anti-)hypertriton polarization and its spin structure (12)

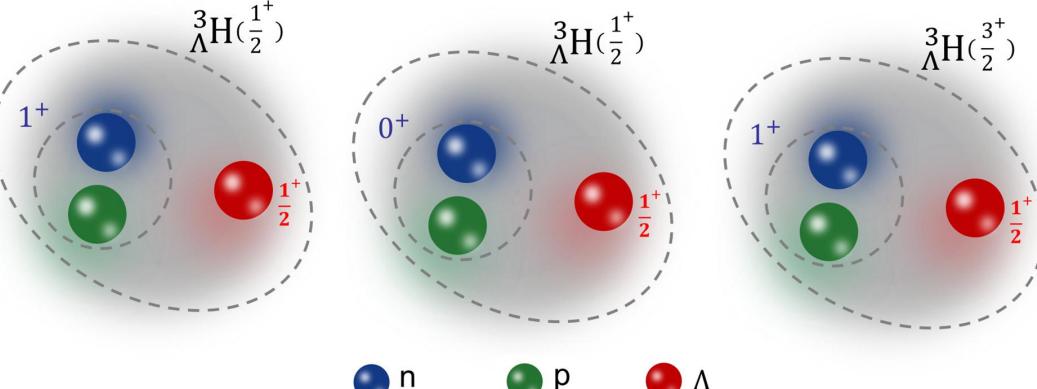


$$\hat{\rho}_{\Lambda^3H} \approx \text{diag} \left[\frac{(1 + \mathcal{P}_\Lambda)^3}{4(1 + \mathcal{P}_\Lambda^2)}, \frac{(1 - \mathcal{P}_\Lambda)(1 + \mathcal{P}_\Lambda)^2}{4(1 + \mathcal{P}_\Lambda^2)}, \right. \\ \left. \frac{(1 - \mathcal{P}_\Lambda)^2(1 + \mathcal{P}_\Lambda)}{4(1 + \mathcal{P}_\Lambda^2)}, \frac{(1 - \mathcal{P}_\Lambda)^3}{4(1 + \mathcal{P}_\Lambda^2)} \right]$$

$$T(\Lambda^3H \rightarrow \pi^- + {}^3He) = \frac{FT_p}{\sqrt{6\pi}} \begin{pmatrix} e^{i\phi^*} \sin \theta^* & 0 \\ -\frac{2}{\sqrt{3}} \cos \theta^* & \frac{e^{i\phi^*} \sin \theta^*}{\sqrt{3}} \\ -\frac{e^{-i\phi^*} \sin \theta^*}{\sqrt{3}} & -\frac{2}{\sqrt{3}} \cos \theta^* \\ 0 & -e^{-i\phi^*} \sin \theta^* \end{pmatrix}$$

$$\frac{dN}{d\cos \theta^*} = \frac{1}{2} \left[1 + \left(\hat{\rho}_{\frac{1}{2}, \frac{1}{2}} + \hat{\rho}_{-\frac{1}{2}, -\frac{1}{2}} - \frac{1}{2} \right) (3 \cos^2 \theta^* - 1) \right]$$

$$\hat{\rho}_{\frac{1}{2}, \frac{1}{2}} + \hat{\rho}_{-\frac{1}{2}, -\frac{1}{2}} - \frac{1}{2} \approx -\frac{\mathcal{P}_\Lambda^2}{1 + \mathcal{P}_\Lambda^2} \approx -\mathcal{P}_\Lambda^2$$



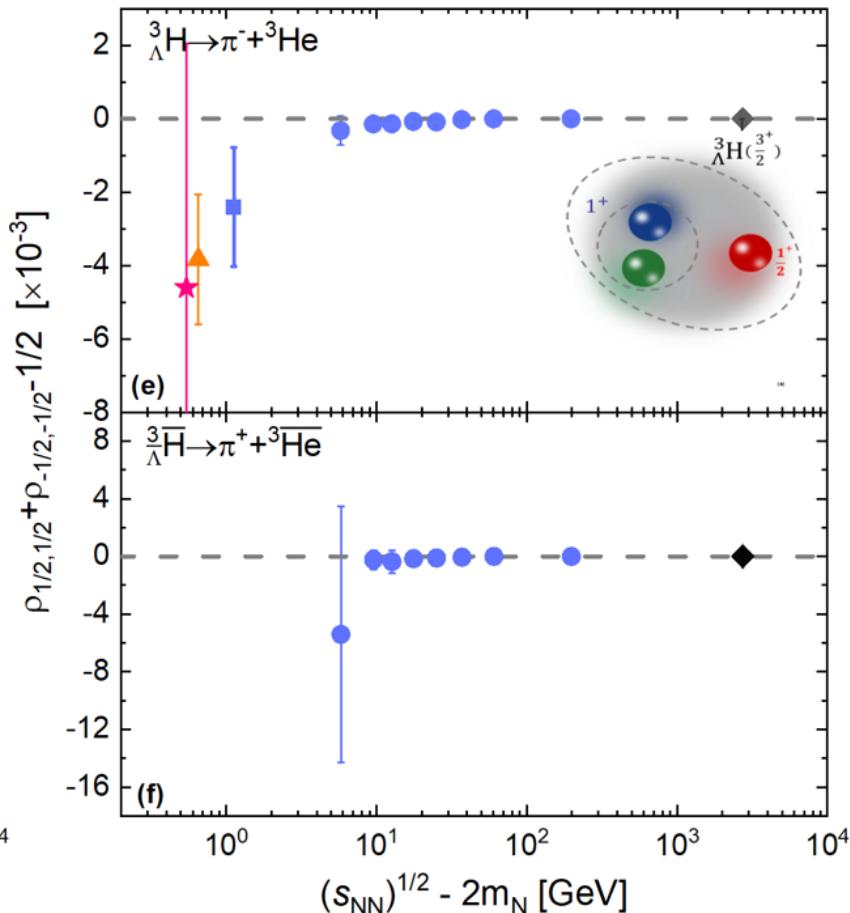
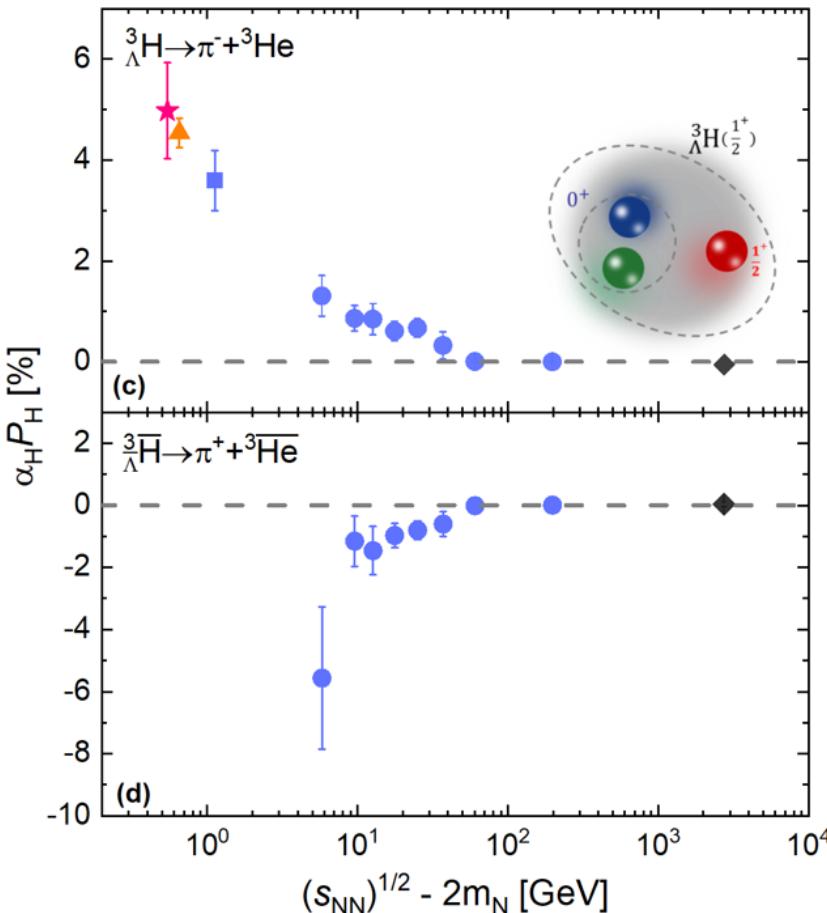
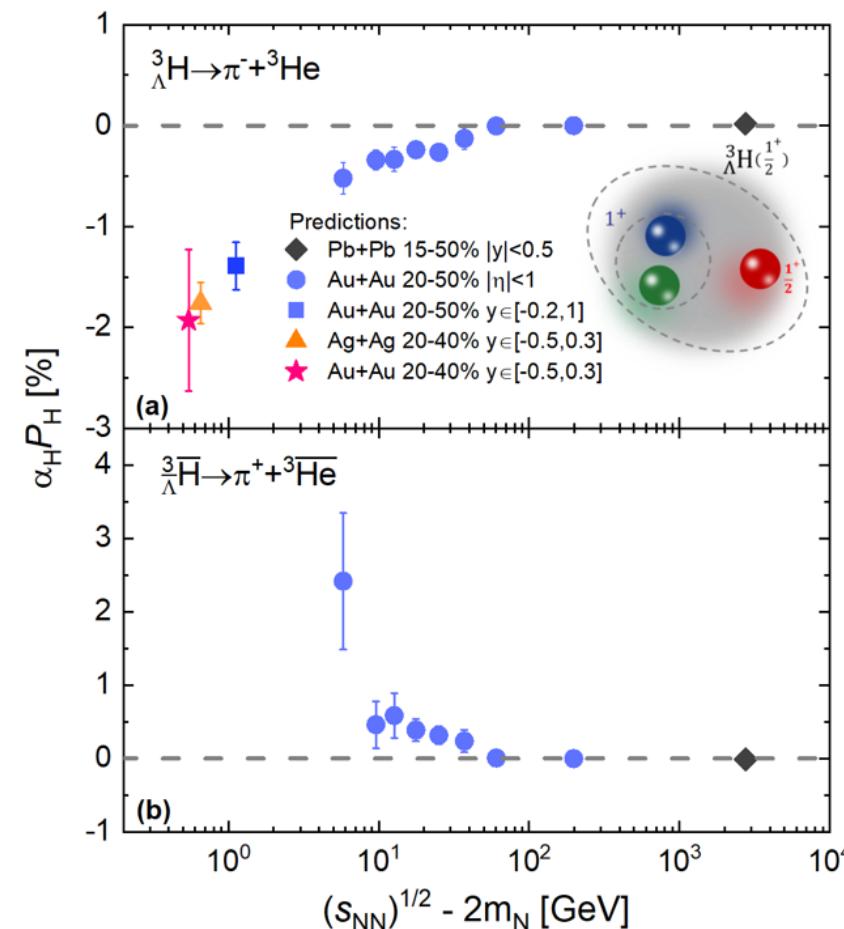
| J^P | structure | decay mode | $\frac{dN}{d\cos \theta^*}$ |
|-----------------|--|---|---|
| $\frac{1}{2}^+$ | $\Lambda(\frac{1}{2}^+) - np(1^+)$ | ${}^3_\Lambda H \rightarrow \pi^- + {}^3He$ | $\frac{1}{2}(1 - \frac{1}{2.58}\alpha_\Lambda \mathcal{P}_\Lambda \cos \theta^*)$ |
| $\frac{1}{2}^+$ | $\Lambda(\frac{1}{2}^+) - np(0^+)$ | ${}^3_\Lambda H \rightarrow \pi^- + {}^3He$ | $\frac{1}{2}(1 + \alpha_\Lambda \mathcal{P}_\Lambda \cos \theta^*)$ |
| $\frac{3}{2}^+$ | $\Lambda(\frac{1}{2}^+) - np(1^+)$ | ${}^3_\Lambda H \rightarrow \pi^- + {}^3He$ | $\frac{1}{2}(1 - \mathcal{P}_\Lambda^2(3 \cos^2 \theta^* - 1))$ |
| $\frac{1}{2}^-$ | $\bar{\Lambda}(\frac{1}{2}^-) - \bar{np}(1^-)$ | ${}^3_{\bar{\Lambda}} \bar{H} \rightarrow \pi^+ + {}^3\bar{He}$ | $\frac{1}{2}(1 - \frac{1}{2.58}\alpha_{\bar{\Lambda}} \mathcal{P}_{\bar{\Lambda}} \cos \theta^*)$ |
| $\frac{1}{2}^-$ | $\bar{\Lambda}(\frac{1}{2}^-) - \bar{np}(0^-)$ | ${}^3_{\bar{\Lambda}} \bar{H} \rightarrow \pi^+ + {}^3\bar{He}$ | $\frac{1}{2}(1 + \alpha_{\bar{\Lambda}} \mathcal{P}_{\bar{\Lambda}} \cos \theta^*)$ |
| $\frac{3}{2}^-$ | $\bar{\Lambda}(\frac{1}{2}^-) - \bar{np}(1^-)$ | ${}^3_{\bar{\Lambda}} \bar{H} \rightarrow \pi^+ + {}^3\bar{He}$ | $\frac{1}{2}(1 - \mathcal{P}_{\bar{\Lambda}}^2(3 \cos^2 \theta^* - 1))$ |

3. (Anti-)hypertriton polarization and its spin structure (13)

The measurement of hypertriton polarization provides a novel method to uniquely determine its internal spin structure

$$\alpha_{\Lambda^3 H} \approx -\frac{1}{2.58} \alpha_\Lambda$$

$$\alpha_{\Lambda^3 H} \approx \alpha_\Lambda$$



4. Effects of baryon spin correlation (14)

Z. T. Liang, Chirality 2023

$$\left| \rho_{00}^V - \frac{1}{3} \right| \gg P_\Lambda^2 \sim P_q^2$$

$$\rho_{00}^V - \frac{1}{3} \sim \langle P_q P_{\bar{q}} \rangle$$

The STAR data show that: $\langle P_q P_{\bar{q}} \rangle \neq \langle P_q \rangle \langle P_{\bar{q}} \rangle$ $\langle P_q P_{\bar{q}} \rangle \gg \langle P_q \rangle \langle P_{\bar{q}} \rangle$

By studying P_H , we study the average of quark polarization P_q ;
by studying ρ_{00}^V , we study the correlation between P_q and $P_{\bar{q}}$.

How to separate long range or local correlations

$$\rho_{10}^V = \frac{P_{qz}(1 + P_{\bar{q}y}) + (1 + P_{qy})P_{\bar{q}z} - iP_{qx}(1 + P_{\bar{q}y}) - i(1 + P_{qy})P_{\bar{q}x}}{\sqrt{2}(3 + \vec{P}_q \cdot \vec{P}_{\bar{q}})}$$

$$\rho_{0-1}^V = \frac{P_{qz}(1 - P_{\bar{q}y}) + (1 - P_{qy})P_{\bar{q}z} - iP_{qx}(1 - P_{\bar{q}y}) - i(1 - P_{qy})P_{\bar{q}x}}{\sqrt{2}(3 + \vec{P}_q \cdot \vec{P}_{\bar{q}})}$$

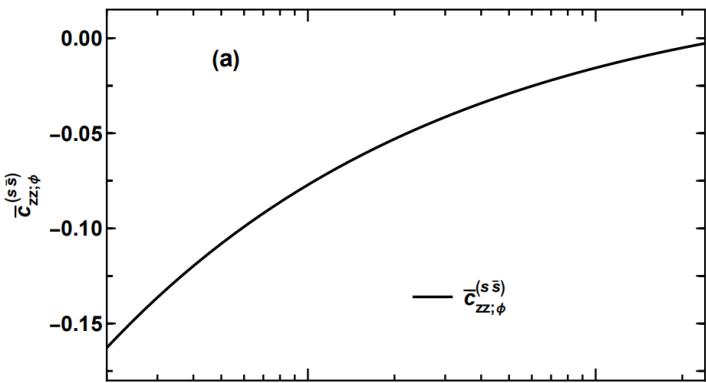
$$\rho_{1-1}^V = \frac{P_{qz}P_{\bar{q}z} - P_{qx}P_{\bar{q}x} + i(P_{qx}P_{\bar{q}y} + P_{qy}P_{\bar{q}x})}{3 + \vec{P}_q \cdot \vec{P}_{\bar{q}}}$$

sensitive to the long range correlation

They should be sensitive to the local correlations.

Global quark spin correlations in relativistic heavy ion collisions

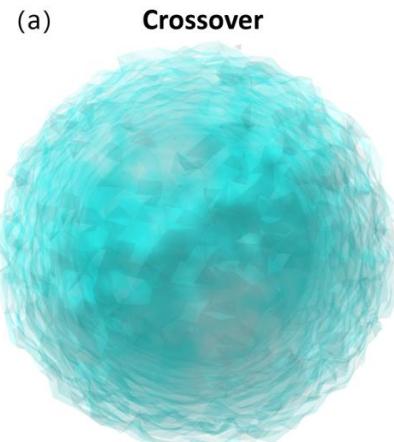
Ji-peng Lv,^{1,*} Zi-han Yu,^{1,†} Zuo-tang Liang,^{1,‡} Qun Wang,^{2,3,§} and Xin-Nian Wang^{4,¶}



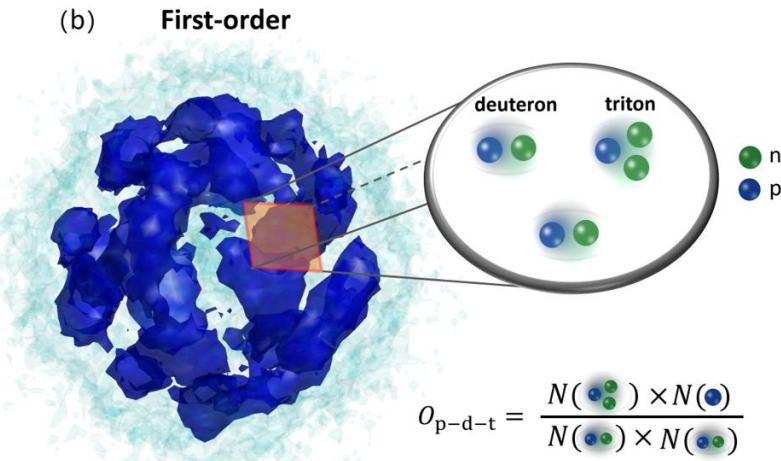
$\sqrt{s_{NN}}$ [GeV]

J. P. Lv et al., Phys. Rev. D 109, 114003 (2024)

C. M. Ko, NST 34, 80 (2023).



Crossover



First-order

$$O_{p-d-t} = \frac{N(\bullet\bullet) \times N(\bullet)}{N(\bullet\bullet) \times N(\bullet\bullet)}$$

$$N_d \approx N_d^{(0)}(1 + C_{np}) + \frac{3}{2^{1/2}} \left(\frac{2\pi}{mT} \right)^{3/2} \times \int d^3\mathbf{x}_1 d^3\mathbf{x}_2 C_2(\mathbf{x}_1, \mathbf{x}_2) \frac{e^{-\frac{(\mathbf{x}_1 - \mathbf{x}_2)^2}{2\sigma_d^2}}}{(2\pi\sigma_d^2)^{\frac{3}{2}}} \\ N_t \approx \frac{3^{3/2}}{4} \left(\frac{2\pi}{mT} \right)^3 \int d^3\mathbf{x}_1 d^3\mathbf{x}_2 d^3\mathbf{x}_3 \rho_{nnp}(\mathbf{x}_1, \mathbf{x}_2, \mathbf{x}_3) \times \frac{1}{3^{3/2}(\pi\sigma_t^2)^3} e^{-\frac{(\mathbf{x}_1 - \mathbf{x}_2)^2}{2\sigma_t^2} - \frac{(\mathbf{x}_1 + \mathbf{x}_2 - 2\mathbf{x}_3)^2}{6\sigma_t^2}},$$

$$\rho_{nnp}(\mathbf{x}_1, \mathbf{x}_2, \mathbf{x}_3) \approx \rho_n(\mathbf{x}_1)\rho_n(\mathbf{x}_2)\rho_p(\mathbf{x}_3) + C_2(\mathbf{x}_1, \mathbf{x}_2)\rho_p(\mathbf{x}_3) + C_2(\mathbf{x}_2, \mathbf{x}_3)\rho_n(\mathbf{x}_1) + C_2(\mathbf{x}_3, \mathbf{x}_1)\rho_n(\mathbf{x}_2) + C_3(\mathbf{x}_1, \mathbf{x}_2, \mathbf{x}_3)$$

$$\frac{N_t N_p}{N_d^2} \approx \frac{1}{2\sqrt{3}} \left[1 + \Delta\rho_n + \frac{\lambda}{\sigma} G\left(\frac{\xi}{\sigma}\right) \right]$$

K. J. Sun, L. W. Chen, C. M. Ko, and Z. Xu, Phys. Lett. B 774, 103 (2017);
K. J. Sun, C. M. Ko, and F. Li, PLB 816, 136258 (2021);

4. Effects of baryon spin correlation (15)

$$\begin{aligned}
\hat{\rho}_{np\Lambda} = & \hat{\rho}_n \otimes \hat{\rho}_p \otimes \hat{\rho}_\Lambda + \frac{1}{2^2} (c_{np}^{\alpha\beta} \hat{\sigma}_{n,\alpha} \otimes \hat{\sigma}_{p,\beta} \otimes \hat{\rho}_\Lambda \\
& + c_{p\Lambda}^{\alpha\beta} \hat{\sigma}_{p,\alpha} \otimes \hat{\sigma}_{\Lambda,\beta} \otimes \hat{\rho}_n + c_{n\Lambda}^{\alpha\beta} \hat{\sigma}_{n,\alpha} \otimes \hat{\sigma}_{\Lambda,\beta} \otimes \hat{\rho}_p) \\
& + \frac{1}{2^3} c_{np\Lambda}^{\alpha\beta\gamma} \hat{\sigma}_{n,\alpha} \otimes \hat{\sigma}_{p,\beta} \otimes \hat{\sigma}_{\Lambda,\gamma}, \\
\mathcal{P}_{\Lambda^3 H} \approx & \frac{\frac{2}{3}\langle \mathcal{P}_n \rangle + \frac{2}{3}\langle \mathcal{P}_p \rangle - \frac{1}{3}\langle \mathcal{P}_\Lambda \rangle - \langle \mathcal{P}_n \mathcal{P}_p \mathcal{P}_\Lambda \rangle + C_-}{1 - \frac{2}{3}(\langle (\mathcal{P}_n + \mathcal{P}_p) \mathcal{P}_\Lambda \rangle) + \frac{1}{3}\langle \mathcal{P}_n \mathcal{P}_p \rangle + C_+} \\
C_- = & -\frac{1}{4}(\langle c_{np}^{zz} \mathcal{P}_\Lambda \rangle + \langle c_{p\Lambda}^{zz} \mathcal{P}_n \rangle + \langle c_{n\Lambda}^{zz} \mathcal{P}_p \rangle) - \frac{1}{4}\langle c_{np\Lambda}^{zzz} \rangle, \quad \text{'genuine' correlation terms} \\
C_+ = & \frac{1}{12}(\langle c_{np}^{zz} \rangle - 2\langle c_{p\Lambda}^{zz} \rangle - 2\langle c_{n\Lambda}^{zz} \rangle).
\end{aligned}$$

Induced correlations

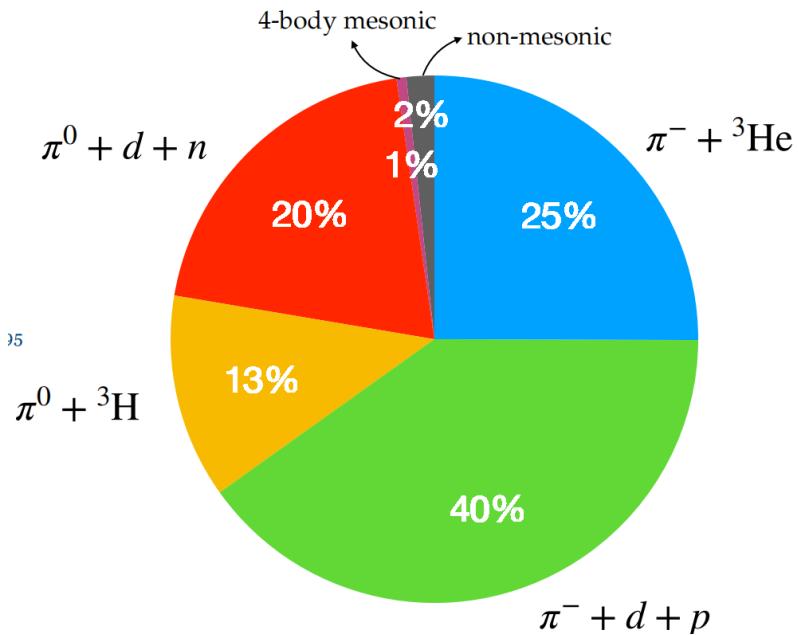
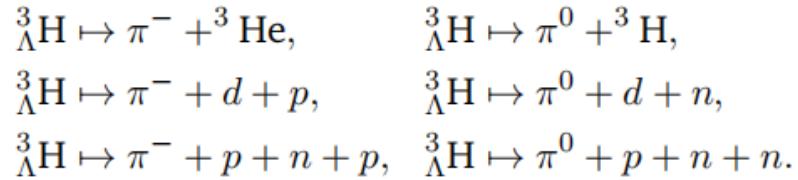
We can express the polarization of a particle as $\mathcal{P} = \langle \mathcal{P} \rangle + \delta\mathcal{P}$ with $\delta\mathcal{P}$ denoting its space and momentum dependent fluctuations, which leads to the relations $\langle \mathcal{P}_n \mathcal{P}_p \rangle = \langle \mathcal{P}_n \rangle \langle \mathcal{P}_p \rangle + \langle \delta\mathcal{P}_n \delta\mathcal{P}_p \rangle$ and $\langle \mathcal{P}_n \mathcal{P}_p \mathcal{P}_\Lambda \rangle = \langle \mathcal{P}_n \rangle \langle \mathcal{P}_p \rangle \langle \mathcal{P}_\Lambda \rangle + \langle \delta\mathcal{P}_n \delta\mathcal{P}_p \rangle \langle \mathcal{P}_\Lambda \rangle + \langle \delta\mathcal{P}_n \delta\mathcal{P}_\Lambda \rangle \langle \mathcal{P}_p \rangle + \langle \delta\mathcal{P}_p \delta\mathcal{P}_\Lambda \rangle \langle \mathcal{P}_n \rangle + \langle \delta\mathcal{P}_n \delta\mathcal{P}_p \delta\mathcal{P}_\Lambda \rangle$. Assuming again $\langle \mathcal{P}_n \rangle \approx \langle \mathcal{P}_p \rangle \approx \langle \mathcal{P}_\Lambda \rangle$ and neglecting the three-body correlation, we then have

$$\mathcal{P}_{\Lambda^3 H} \approx (1 - \langle \delta\mathcal{P}_n \delta\mathcal{P}_p \rangle - \langle \delta\mathcal{P}_p \delta\mathcal{P}_\Lambda \rangle - \langle \delta\mathcal{P}_n \delta\mathcal{P}_\Lambda \rangle) \langle \mathcal{P}_\Lambda \rangle.$$

This result suggests that it is possible to extract the information on the spin-spin correlations among nucleons and Λ hyperons from the measurement of hypertriton polarization in heavy-ion collisions, although it is non-trivial in practice.

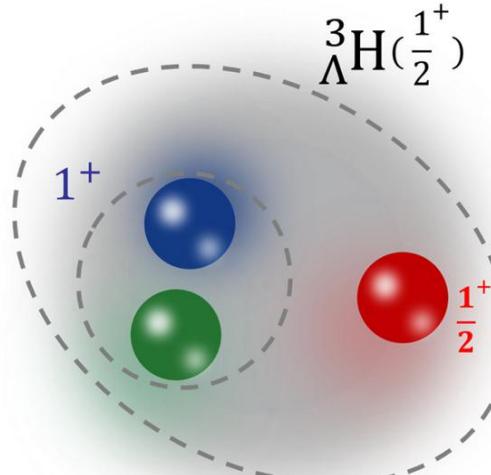
4. Three-body decay

(16)



Relative branching ratio:

$$R_3 = \frac{\text{B.R.}({}^3_{\Lambda}\text{H} \rightarrow {}^3\text{He}\pi^-)}{\text{B.R.}({}^3_{\Lambda}\text{H} \rightarrow {}^3\text{He}\pi^-) + \text{B.R.}({}^3_{\Lambda}\text{H} \rightarrow dp\pi^-)}$$



$$\frac{dN}{d \cos\theta^*} = \frac{1}{2} \left(1 + \alpha_{{}^3_{\Lambda}\text{H}} \mathcal{P}_{{}^3_{\Lambda}\text{H}} \cos\theta^* \right)$$

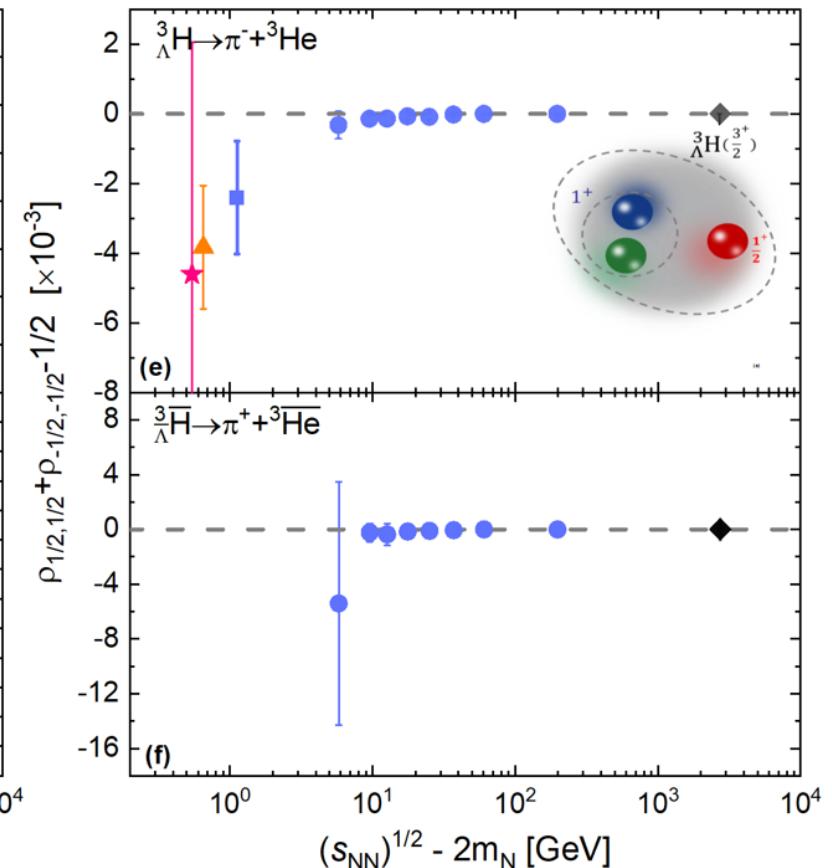
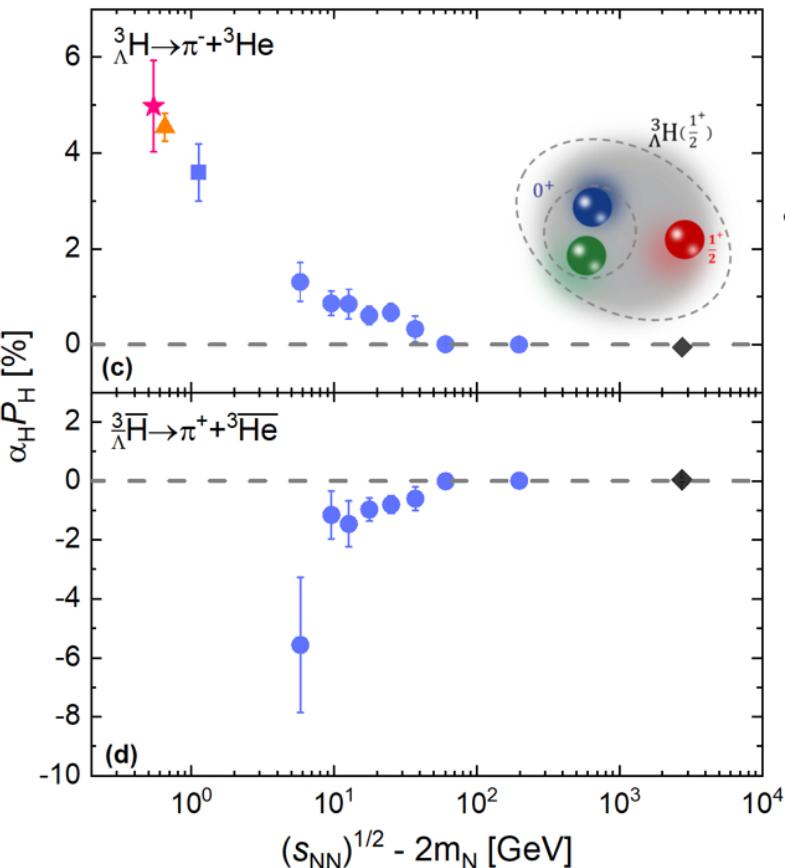
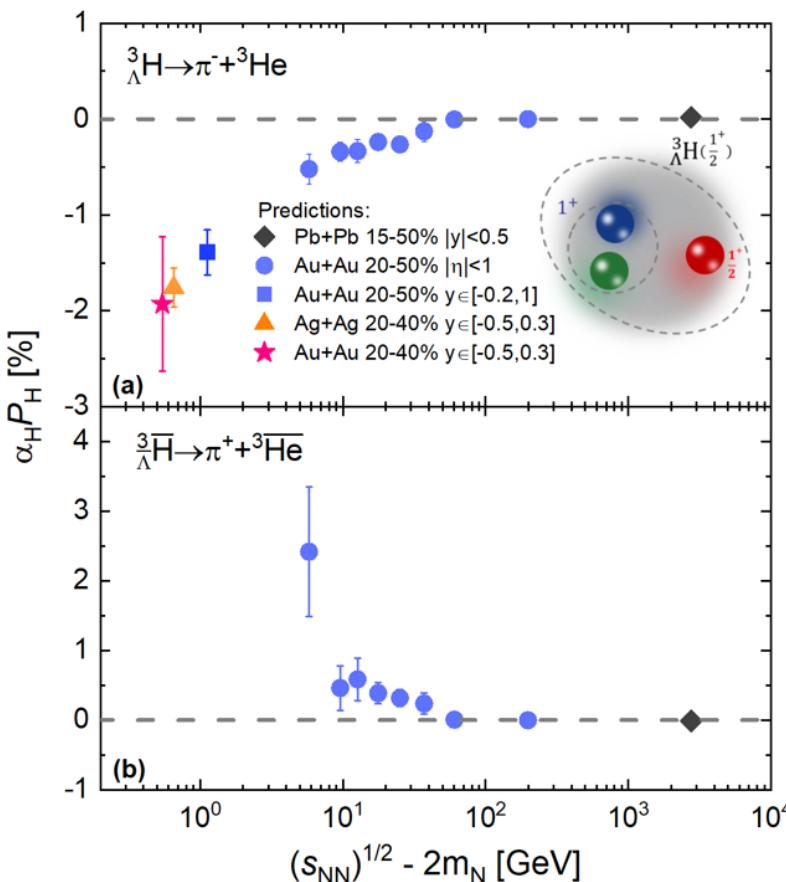
$$\begin{aligned} \alpha_{{}^3_{\Lambda}\text{H}} &\approx -\frac{1}{3T_s^2 + \frac{1}{3}T_p^2} \alpha_{\Lambda} \\ &\approx -\frac{1}{2.58} \alpha_{\Lambda} \end{aligned}$$

$$\begin{aligned} \alpha_{{}^3_{\Lambda}\text{H}} &\approx -\frac{1}{3.2} \alpha_{\Lambda} \\ \text{with } R_3 &= 0.25 \end{aligned}$$

Summary and outlook

(17)

1. (Anti-)hypertriton is globally polarized in non-central heavy-ion collisions.
2. (Anti-)hypertriton polarization and its decay pattern provide a novel method to uniquely determine the spin structure of its wavefunction.

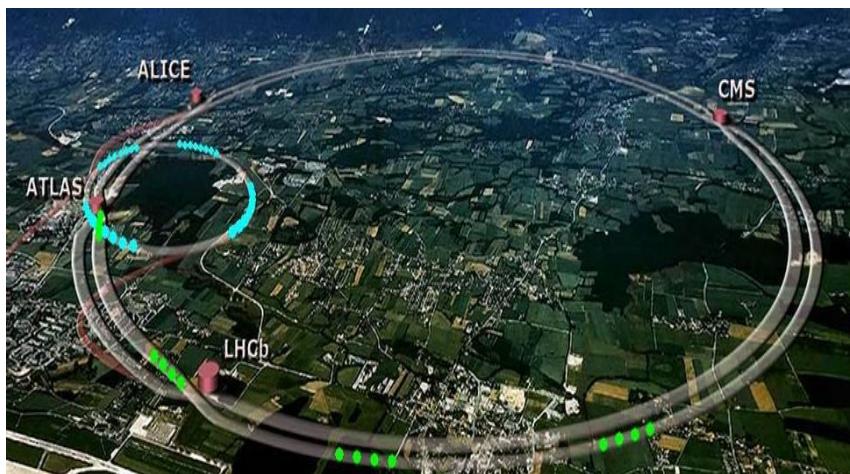
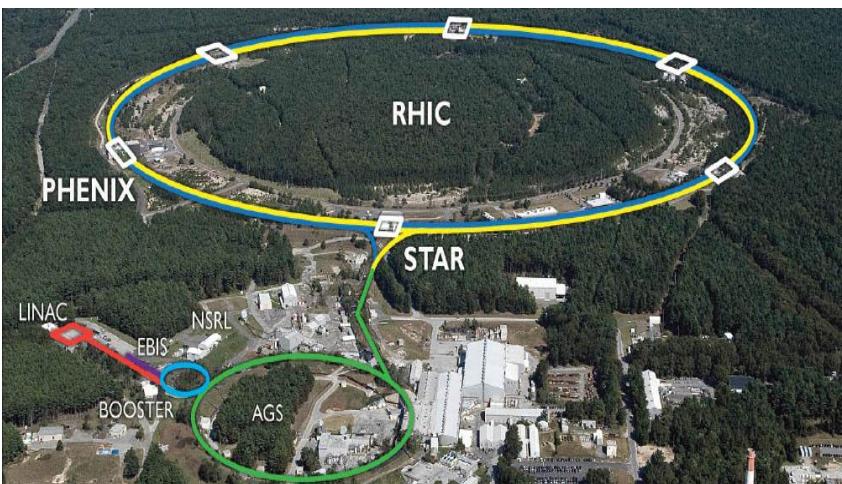


Backup

Little Bang Nucleosynthesis

(2)

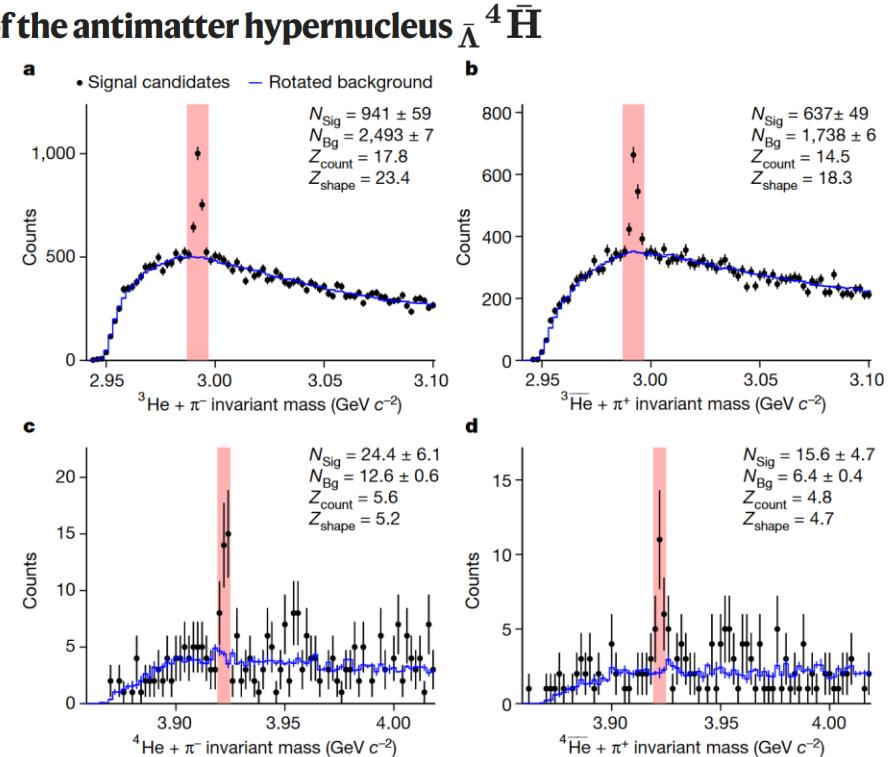
Antimatter factory



Observation of the antimatter hypernucleus $\bar{\Lambda}^4\bar{H}$

STAR Collaboration

[Nature \(2024\)](#) | [Cite this ar](#)



PHYSICAL REVIEW C 93, 064909 (2016)

Antimatter $\bar{\Lambda}^4\bar{H}$ hypernucleus production and the ${}^3\text{H}/{}^3\text{He}$ puzzle in relativistic heavy-ion collisions

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We show that the measured yield ratio ${}^3\text{H}/{}^3\text{He}$ (${}^3\bar{\Lambda}\bar{H}/{}^3\text{He}$) in Au + Au collisions at $\sqrt{s_{NN}} = 200$ GeV and in Pb + Pb collisions at $\sqrt{s_{NN}} = 2.76$ TeV can be understood within a covariant coalescence model if (anti-) Λ particles freeze out earlier than (anti-)nucleons but their relative freeze-out time is closer at $\sqrt{s_{NN}} = 2.76$ TeV than at $\sqrt{s_{NN}} = 200$ GeV. The earlier (anti-) Λ freeze-out can significantly enhance the yield of (anti)hypernucleus ${}^4\bar{\Lambda}\bar{H}$ (${}^4\bar{H}$), leading to that ${}^4\bar{\Lambda}\bar{H}$ has a comparable abundance with ${}^4\bar{\Lambda}\bar{H}$ and thus provides an easily measured antimatter candidate heavier than ${}^4\text{He}$. The future measurement on ${}^4\text{H}/{}^4\bar{\Lambda}\bar{H}$ would be very useful to understand the (anti-) Λ freeze-out dynamics and the production mechanism of (anti)hypernuclei in relativistic heavy-ion collisions.

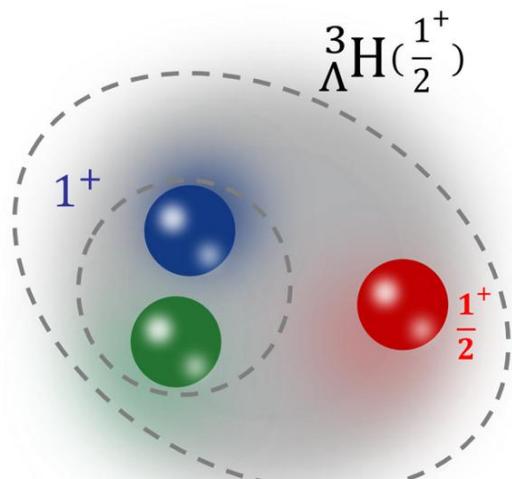
(Anti-)hypertriton polarization and its spin structure

Parity-violating weak decay:

$$T(\Lambda \rightarrow \pi^- + p) = \frac{1}{\sqrt{4\pi}} \begin{pmatrix} T_s + T_p \cos \theta_p^* & T_p \sin \theta_p^* e^{i\phi_p^*} \\ T_p \sin \theta_p^* e^{-i\phi_p^*} & T_s - T_p \cos \theta_p^* \end{pmatrix}$$

$$T(\Lambda^3 \text{H} \rightarrow \pi^- + ^3 \text{He})$$

$$= \frac{F}{6\sqrt{\pi}} \begin{pmatrix} 3T_s - T_p \cos \theta^* & -T_p \sin \theta^* e^{i\phi^*} \\ -T_p \sin \theta^* e^{-i\phi^*} & 3T_s + T_p \cos \theta^* \end{pmatrix}$$



Sign flip !

The normalized angular distribution of the ${}^3\text{He}$ in the decay $\Lambda^3 \text{H} \rightarrow \pi^- + {}^3 \text{He}$ is given by

$$\frac{dN}{d \cos \theta^*} = \text{Tr}[T^+ \hat{\rho} T] = \frac{1}{2}(1 + \alpha_{\Lambda^3 \text{H}} \mathcal{P}_{\Lambda^3 \text{H}} \cos \theta^*), \quad (7)$$

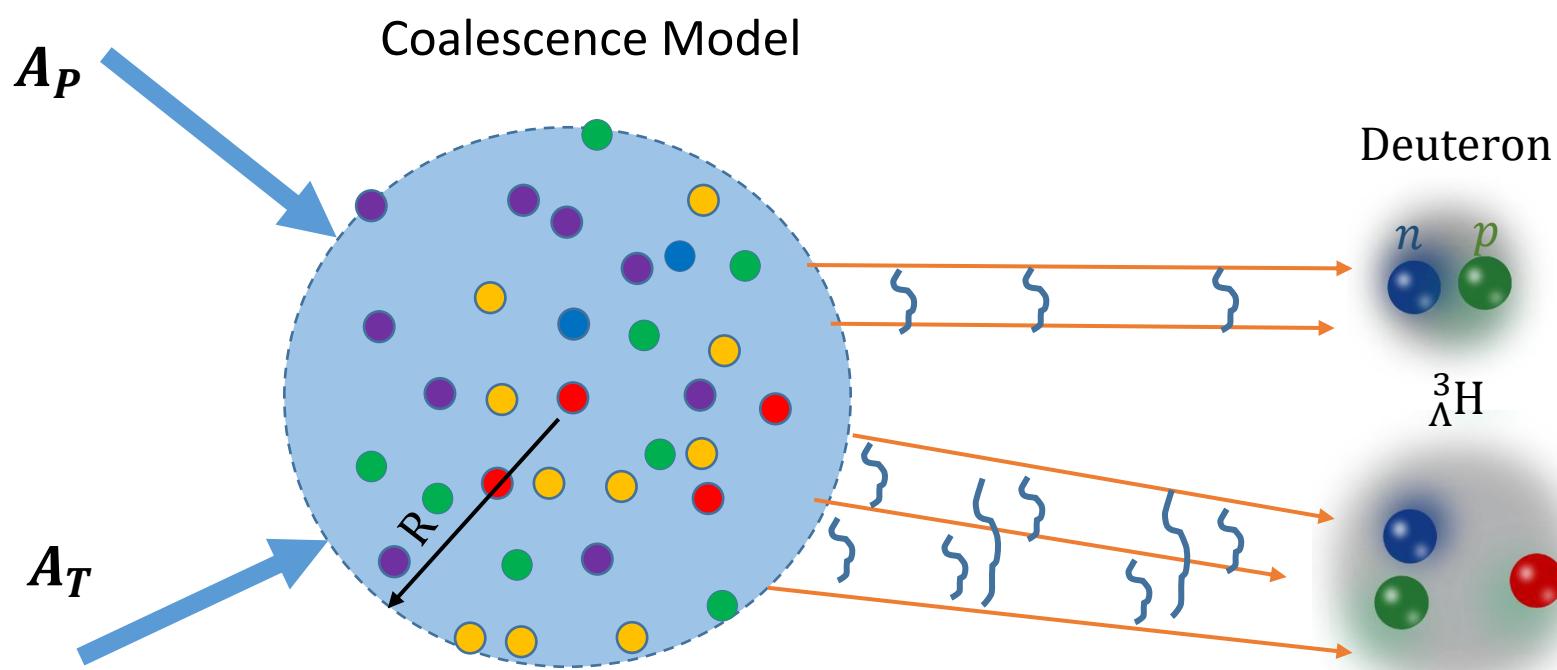
in terms of the hypertriton decay parameter $\alpha_{\Lambda^3 \text{H}} \approx -\frac{1}{3T_s^2 + \frac{1}{3}T_p^2} \alpha_\Lambda \approx -\frac{1}{2.58} \alpha_\Lambda$. The angular distribution of ${}^3\text{He}$ in the decay $\Lambda^3 \text{H} \rightarrow \pi^- + {}^3 \text{He}$ can thus be further expressed as

$$\frac{dN}{d \cos \theta^*} \approx \frac{1}{2}(1 - \frac{1}{2.58} \alpha_\Lambda \mathcal{P}_\Lambda \cos \theta^*). \quad (8)$$

Compared to the angular distribution of the proton in the Λ decay, which has the form

$$\frac{dN}{d \cos \theta_p^*} = \frac{1}{2}(1 + \alpha_\Lambda \mathcal{P}_\Lambda \cos \theta_p^*), \quad (9)$$

the ${}^3\text{He}$ in $\Lambda^3 \text{H}$ decay has an opposite sign in its angular dependence.

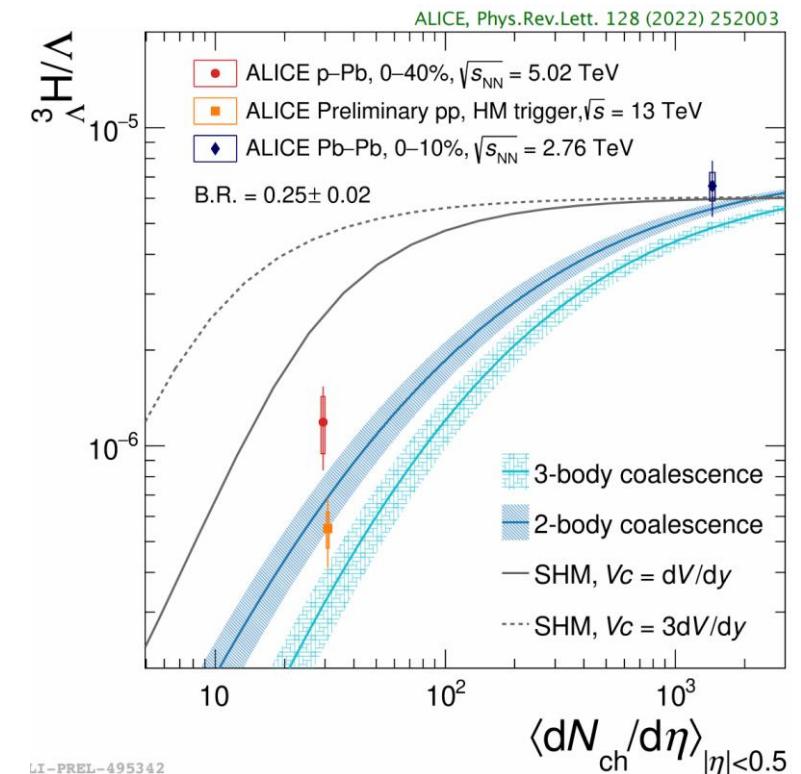


Density Matrix Formulation
(**sudden** approximation)

$$N_A = \text{Tr}(\hat{\rho}_s \hat{\rho}_A) \\ = g_c \int d\Gamma \rho_s(\{x_i, p_i\}) \times W_A(\{x_i, p_i\})$$

Wigner function of light cluster

Overlap between source distribution function and Wigner function of light nuclei



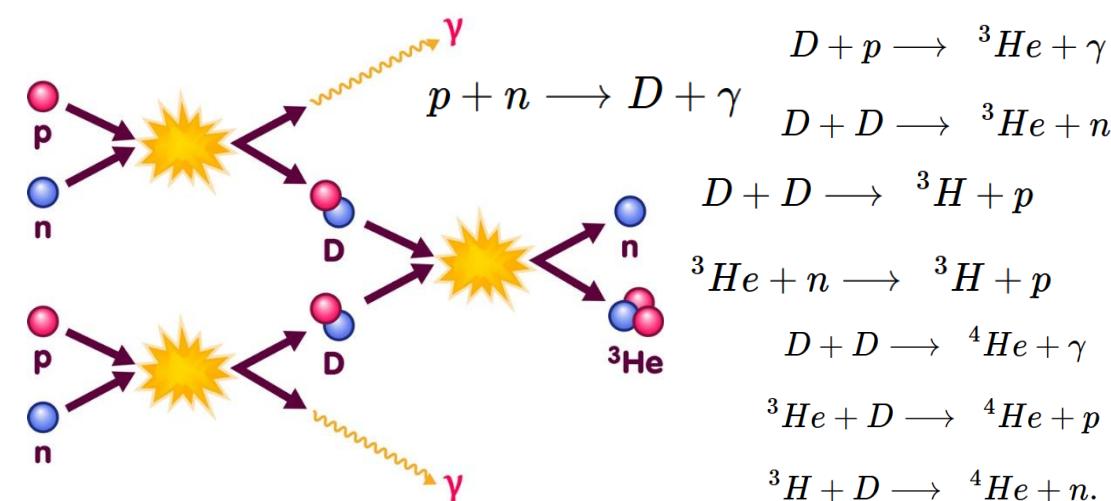
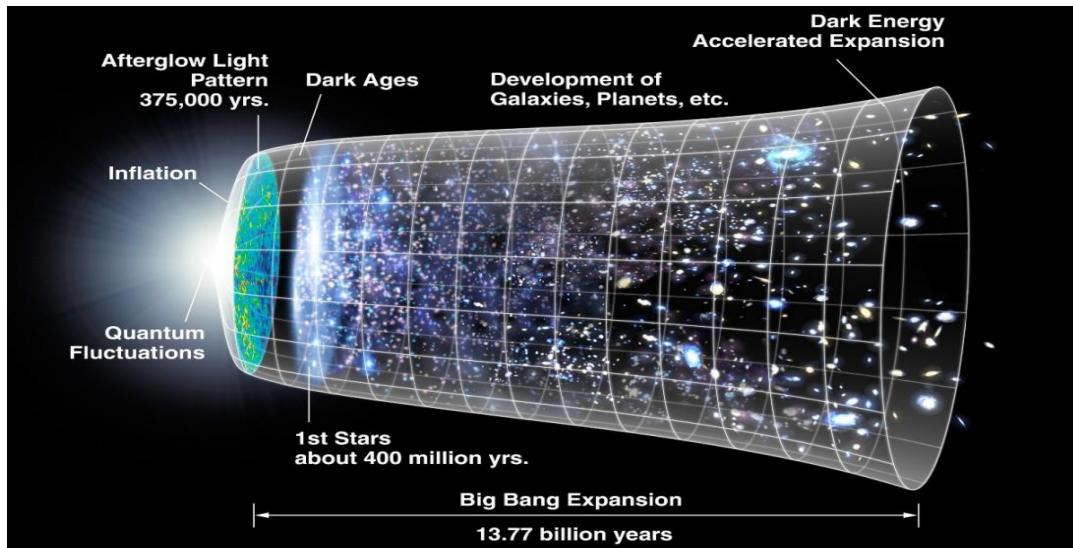
Little Bang Nucleosynthesis

Big-bang nucleosynthesis is responsible for the formation of light nuclei in our Universe.

$$t \sim 100 \text{ s}, kT < 1 \text{ MeV}$$

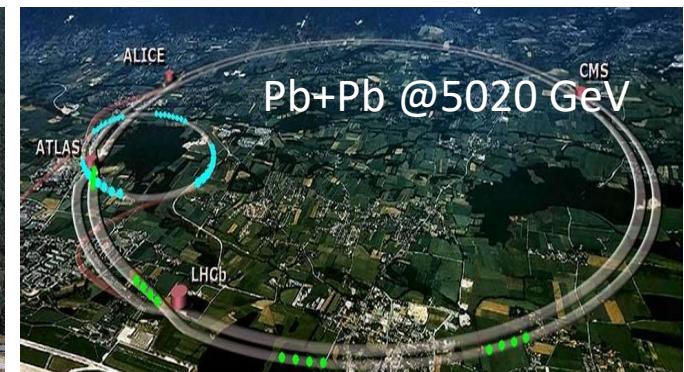
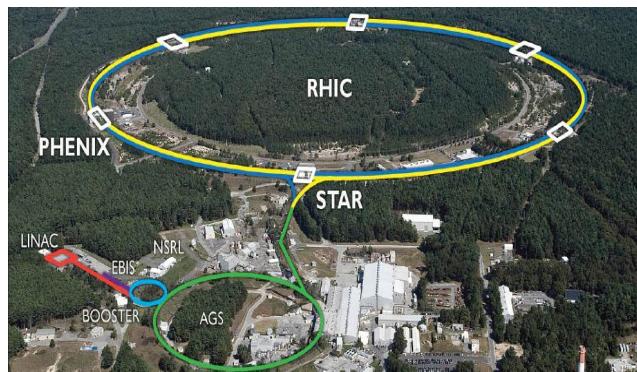
K. A. Olive et al., Phys. Rept. 333, 389–407 (2000);

《The First Three Minutes》 S. Weinberg



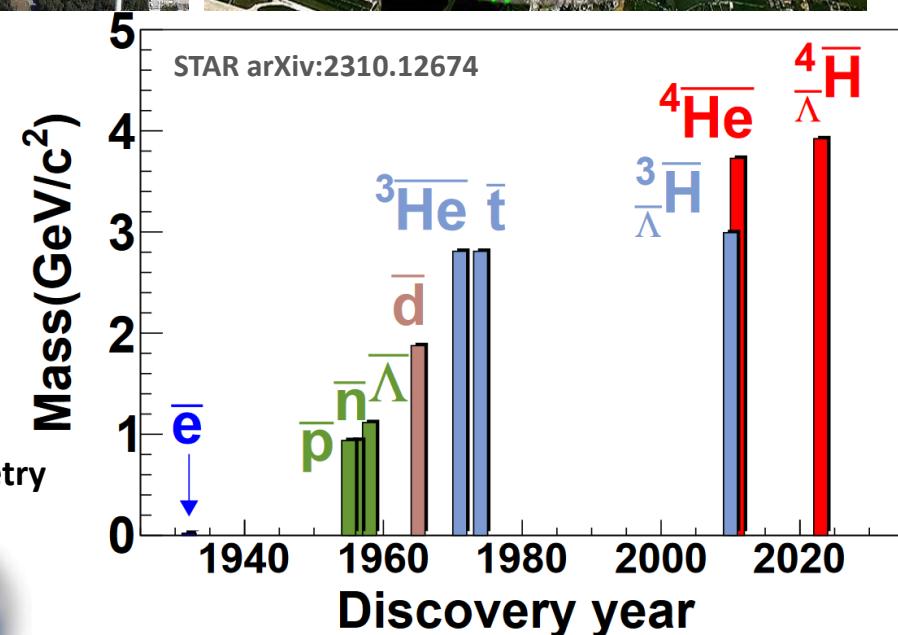
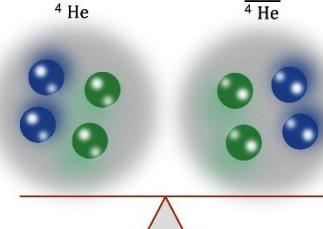
Synthesis of antimatter nuclei in little bangs of relativistic heavy-ion collisions

$$t \sim 10^{-22} \text{ s}, kT \sim 100 \text{ MeV}$$



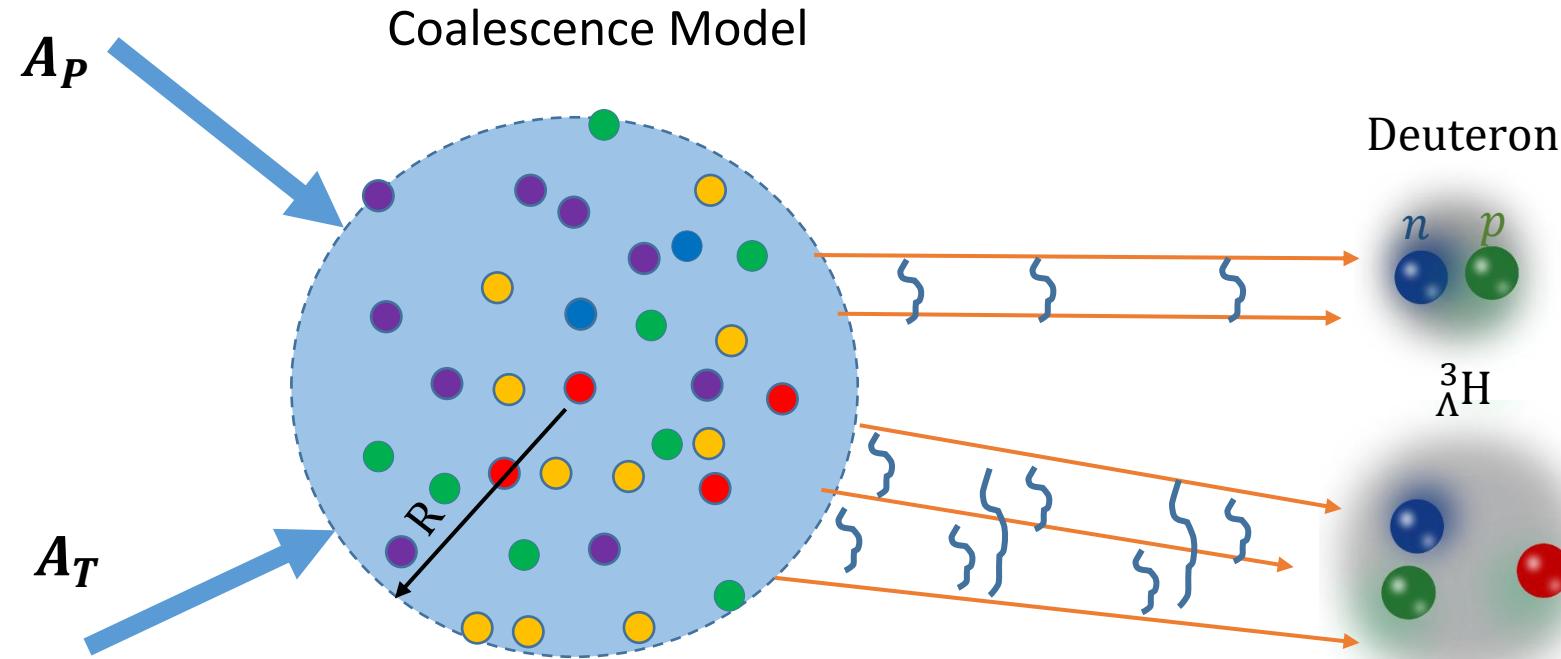
Antimatter factory

Matter-antimatter asymmetry



J. Chen et al., Phys. Rep. 760, 1 (2018);
P. Braun-Munzinger and B. Donigus NPA987, 144 (2019)

Final-state coalescence



Two-body coalescence $a + b \rightarrow c$:

$$N_c = \frac{2J_c + 1}{(2J_a + 1)(2J_b + 1)} \int \frac{dx_a dk_a}{(2\pi)^3} \frac{dx_b dk_b}{(2\pi)^3} f_a(x_a, k_a) f_b(x_b, k_b) W_c(x, k)$$

$$\approx \frac{2J_c + 1}{(2J_a + 1)(2J_b + 1)} \frac{N_a N_b}{(\frac{m_a m_b T}{m_a + m_b} (R_a^2 + R_b^2))^{3/2}} \times \frac{1}{(1 + \frac{\sigma^2}{R_a^2 + R_b^2})^{3/2}}$$

$$f_a = \frac{N_a}{(m_a T R_a^2)^{3/2}} e^{-\frac{k_a^2}{2m_a T} - \frac{x_a^2}{2R_a^2}} \quad W_c = 8e^{-x^2/\sigma^2 - \sigma^2 k^2}$$

$$N_a = \int \frac{dx_a dk_a}{(2\pi)^3} f_a(x_a, k_a) \quad 1 = \int \frac{dx dk}{(2\pi)^3} W_c(x, k)$$

“Quantum mechanical correction”

Production

Structure

$N_d \propto \frac{1}{\left[1 + \left(\frac{2r_d^2}{3R^2}\right)\right]^{3/2}}$

$N_{^3H} \propto \frac{1}{\left[1 + \left(\frac{r_{^3H}^2}{2R^2}\right)\right]^3}$

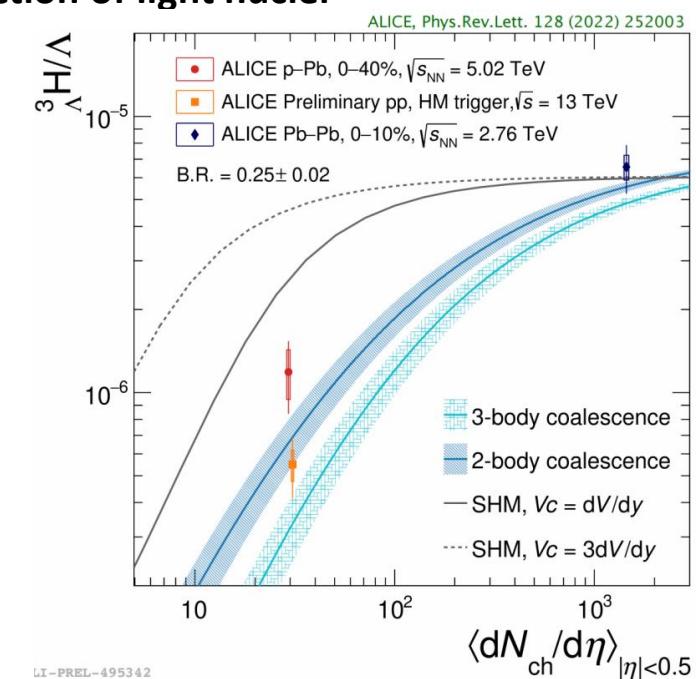
can be inferred from Femtoscopy

Density Matrix Formulation
(sudden approximation)

$$N_A = Tr(\hat{\rho}_s \hat{\rho}_A) \\ = g_c \int d\Gamma \rho_s(\{x_i, p_i\}) \times W_A(\{x_i, p_i\})$$

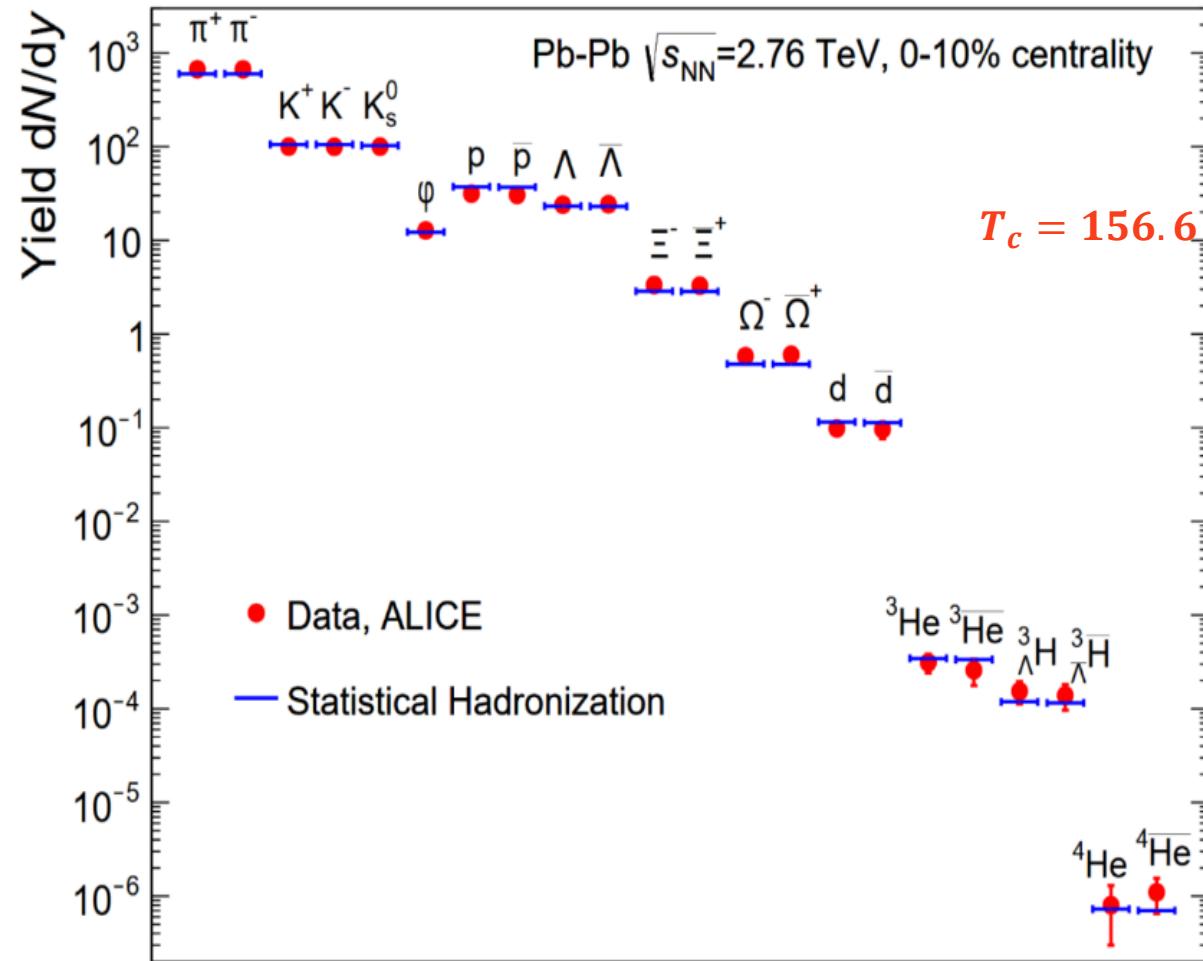
Wigner function of light cluster

Overlap between source distribution function and Wigner function of light nuclei



Statistical hadronization

Andronic, Braun-Munzinger, Redlich, Stachel, Nature 561, 321 (2018)



$$N_h \approx \frac{g_h V_C}{2\pi^2} m_h^2 T_C K_2\left(\frac{m_h}{T_C}\right)$$
$$\approx g_h V_C \left(\frac{m_h T_C}{2\pi}\right)^{3/2} e^{-m_h/T_C}$$

T_C : Chemical freeze-out temperature, which is close to the chiral transition temperature (LQCD)

All (stable) particles including light (hyper)nuclei are produced at the QCD phase boundary and share a common chemical freeze-out

Compact multi-quark states

loosely bound states

$\tau = 0 \text{ fm}/c$
Hadronization

$\tau > 10 - 20 \text{ fm}/c$
Wave function fully developed