



Introduction to Past and Future MeV Missions

Xilu Wang (王夕露)

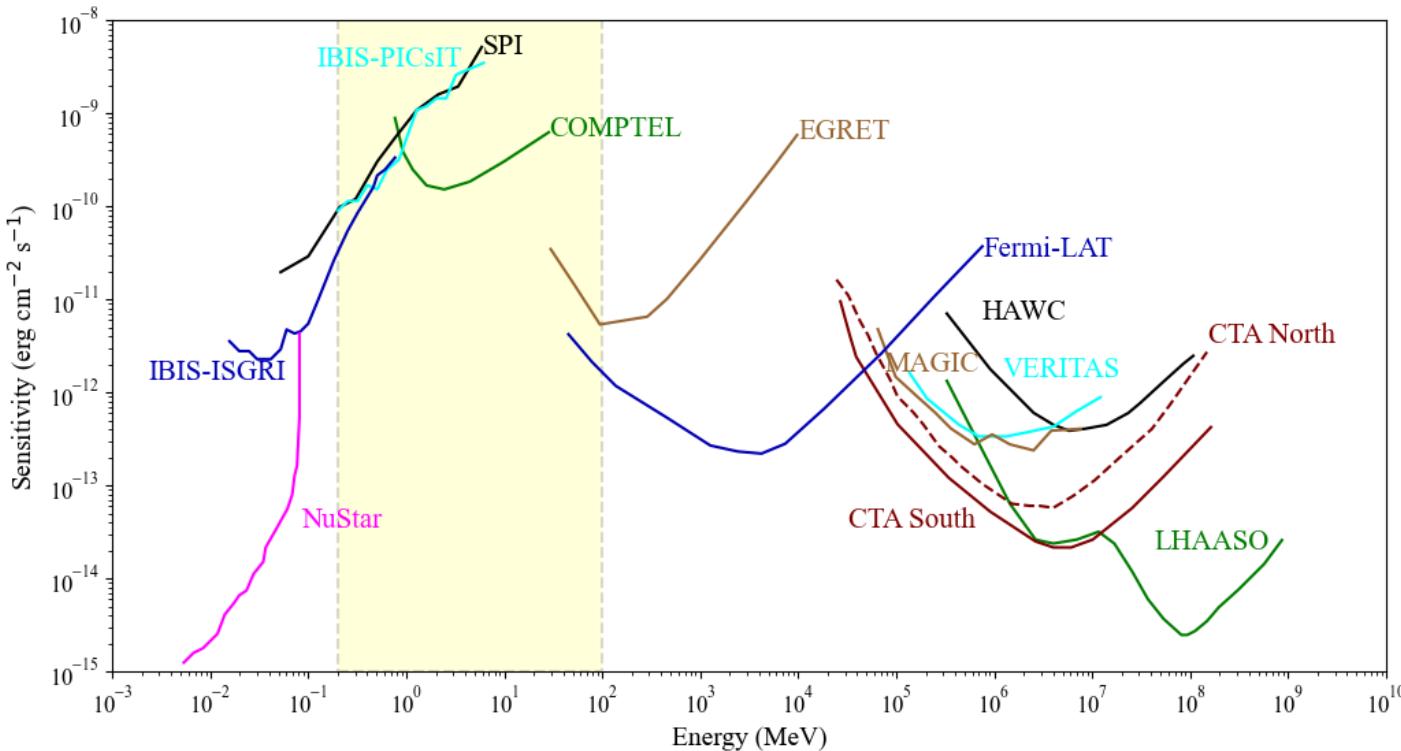
Institute of High Energy Physics, CAS

wangxl@ihep.ac.cn



中国科学院高能物理研究所
Institute of High Energy Physics
Chinese Academy of Sciences

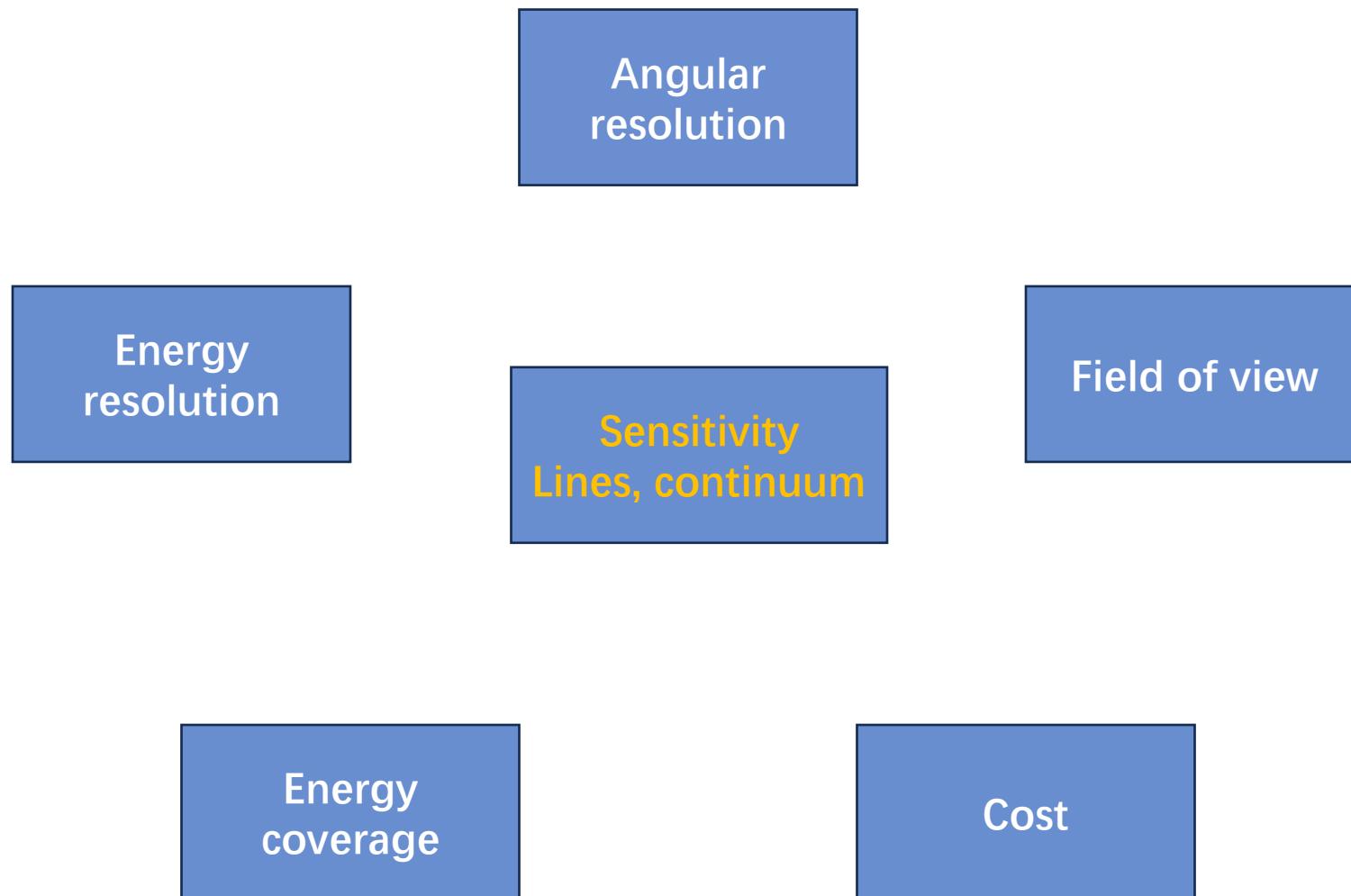
MeV Gamma Ray - An important window in multi-messenger observation



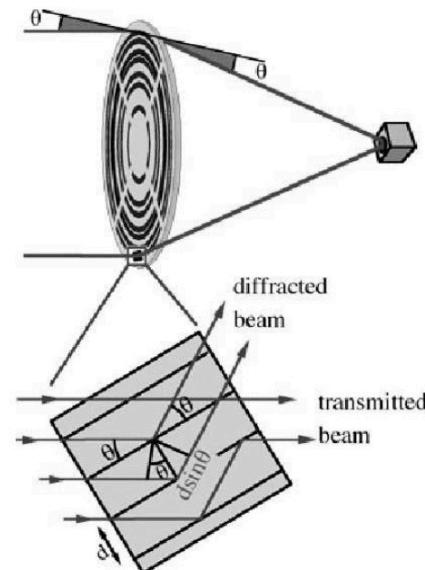
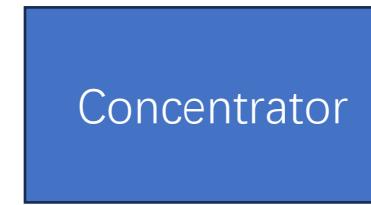
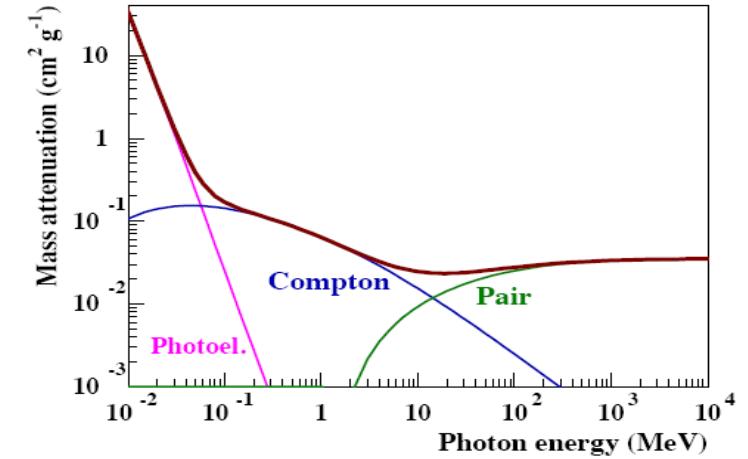
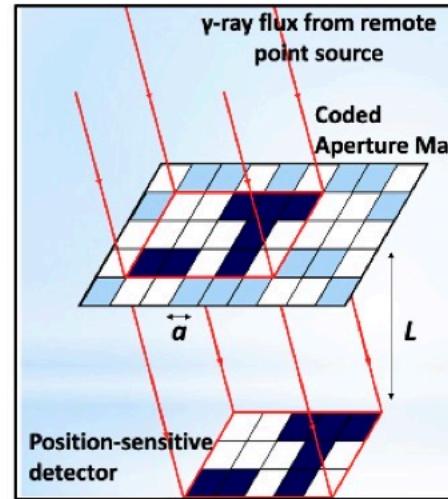
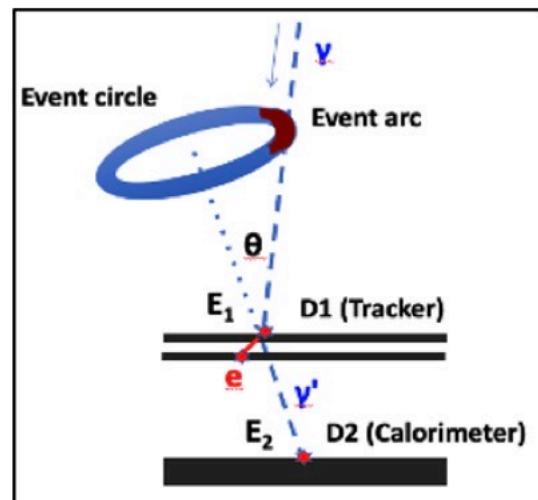
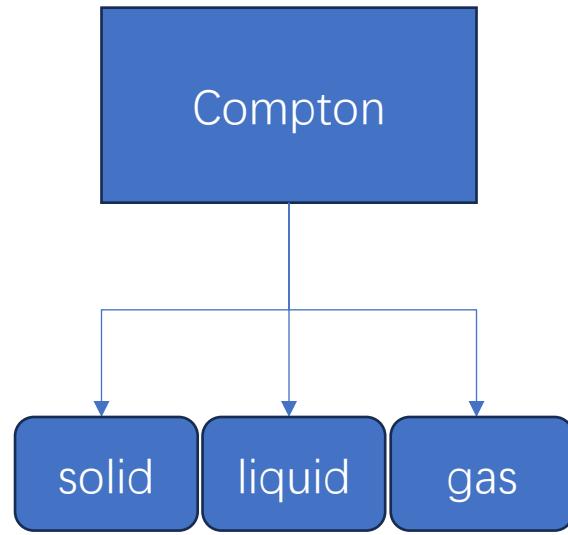
“MeV Gap”:
sensitivity gap in
MeV band (~ 0.2 -
100MeV)

- Many multi-messenger sources' **peak** radiation power is expected to be in the MeV band: blazars, pulsars, gamma-ray bursts (GRB) ...
- MeV gamma rays can provide unique information to help address the **frontiers**: the origin of cosmic rays, dark matter, the radiation mechanisms of **compact objects** ...
- MeV band is the region of **gamma-ray nuclear lines**, providing a unique and direct probe of the nuclear process in the universe → the origin of heavy elements & stellar evolutions

MeV missions --- important factors



MeV Detection Techniques

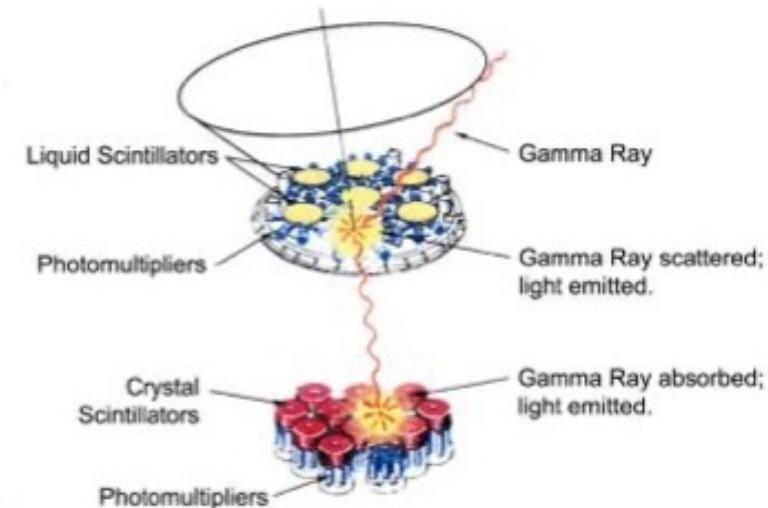


MeV Missions - Compton - COMPTEL

Compton Gamma Ray Observatory (CGRO)
1991/4/5– 2000/06/04



COMPTEL



Energy Range: 0.8-30 MeV

Energy Resolution: ~8.8% @ 1.27 MeV

Spatial Precision: ~1°

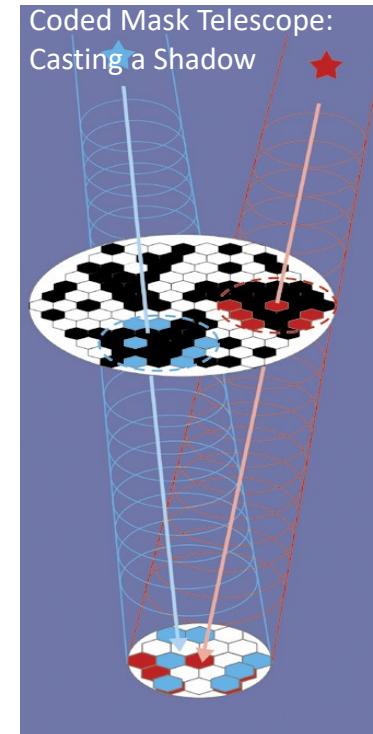
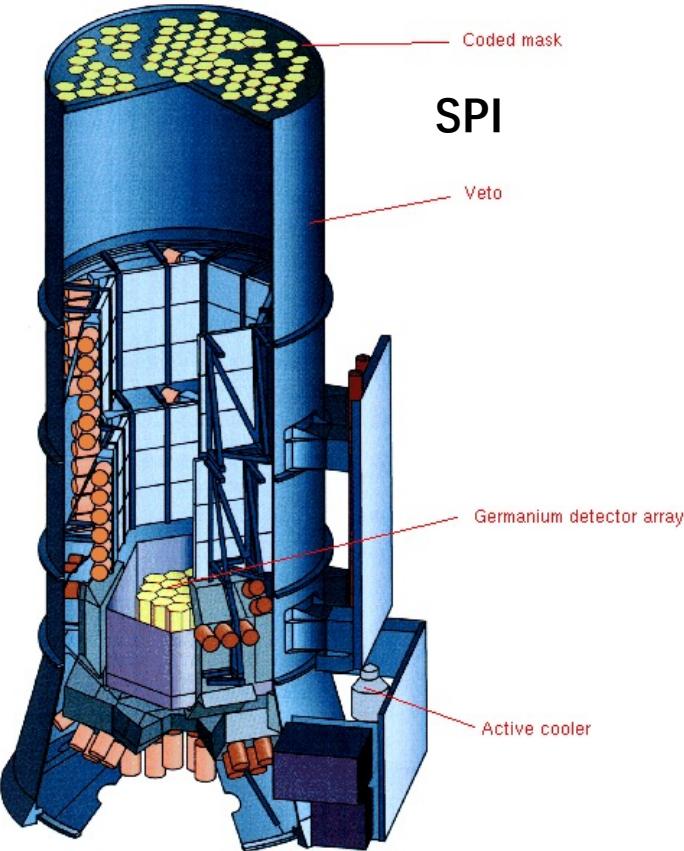
Field of View: 64°

Line sensitivity (5×10^5 sec): $\sim 1.5\text{-}6 \times 10^{-5} \text{ cm}^{-2} \text{ s}^{-1}$

MeV Missions - Coded Mask - SPI

INTEGRAL

2002/10/17– 2025/02/28



Energy Range: 15keV-8 MeV

Energy Resolution: ~ 2.2 keV @ 662 keV

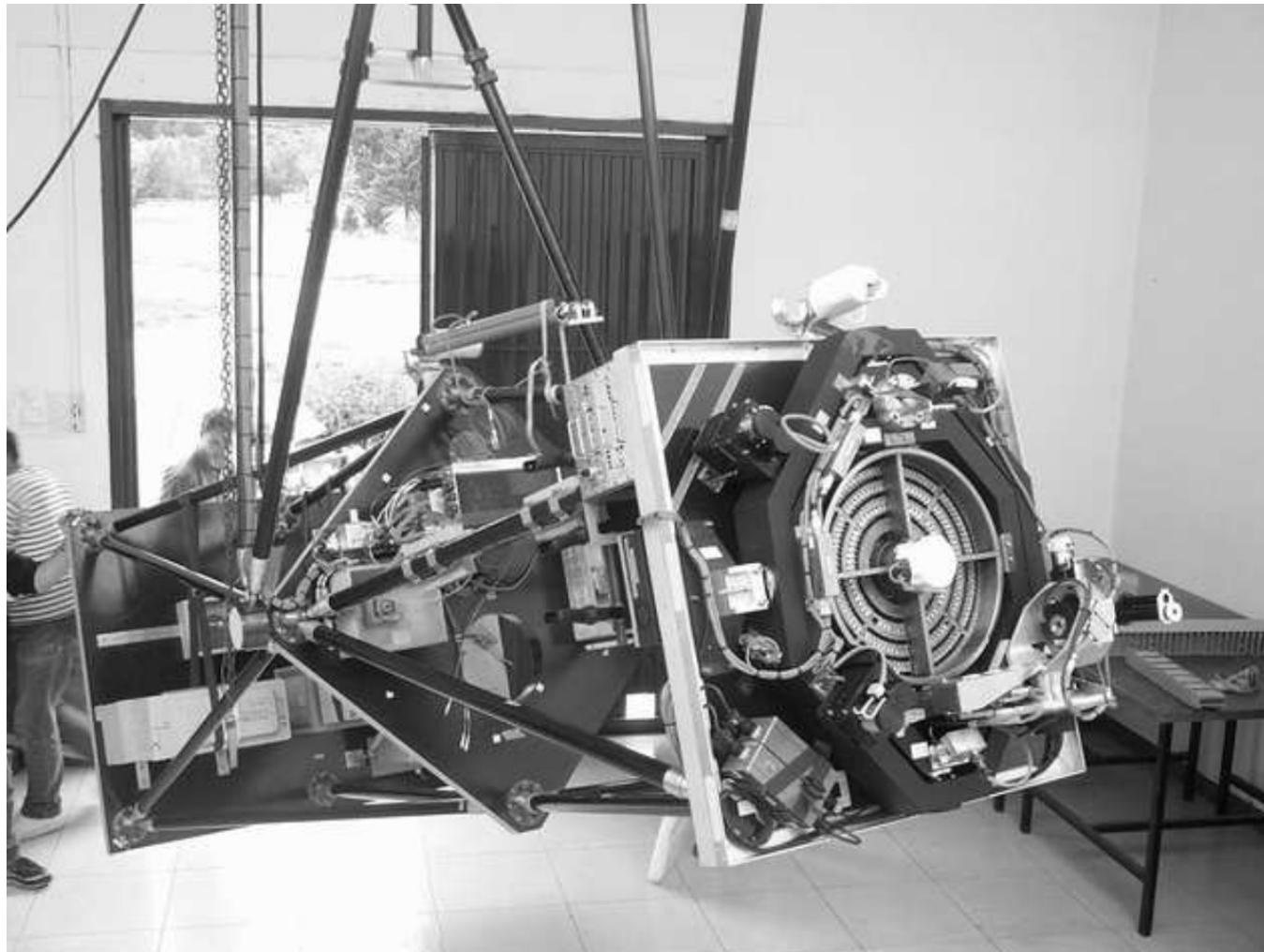
Spatial Precision: 2.6° / ~ 2 arcmin

Field of View: $16 \times 16^\circ$ (fully coded)

Line sensitivity (10^6 sec): $\sim 2.4 \times 10^{-5} \text{ cm}^{-2} \text{ s}^{-1}$ @ 1MeV

MeV Missions – Concentrator - CLAIRE

CLAIRE – balloon experiment
2001/06/14



- Lens: 556 Ge-Si crystals (45cm)
- Focal length: 2.77 meter
- focusing 170 keV gamma-ray photons onto a 3x3 matrix of HPGe detectors
- Angular resolution: 18 arcmin @ 170keV
- The float altitude was 41 km and lasted about 5 hours.

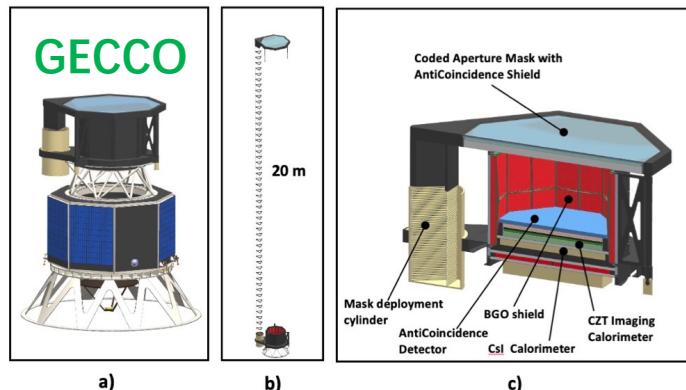
MeV Missions – Proposals

Observatory:

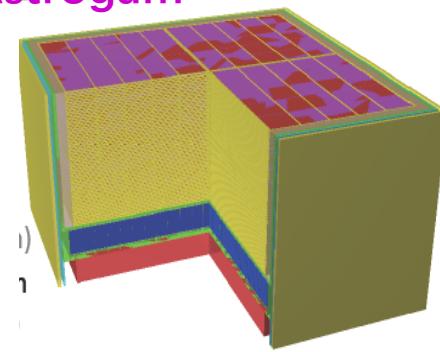
- eAstrogam (ESA)
- AMEGO/AMEGO-X (NASA)
- *MeVGRO*
- **VLAST**
- SRG-M (Russia)

Small-Mid:

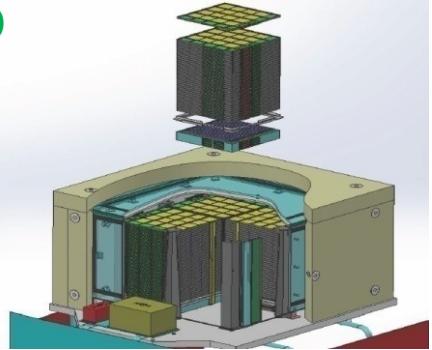
- MASS, COSI** (funded in 2021, 0.2~2.5MeV; launch in 2027)
- GECCO—Coded mask
- GRAMS, GammaTPC—LAr-TPC
- MeGaT**, SMILE, Adept — Gaseous-TPC
- DUAL, CLAIRE,511-cam, **PAIRS**-- Concentrator



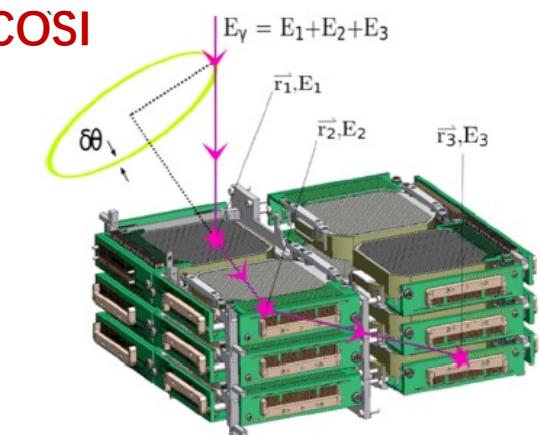
eAstrogam



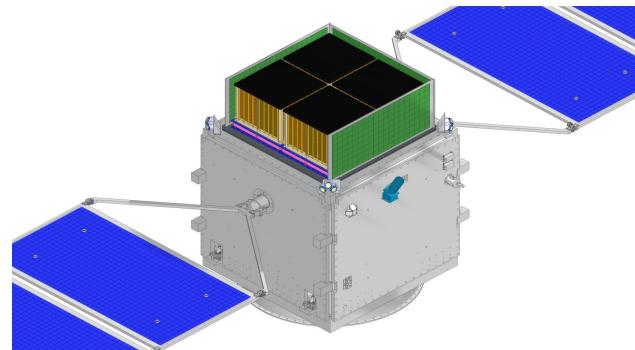
AMEGO



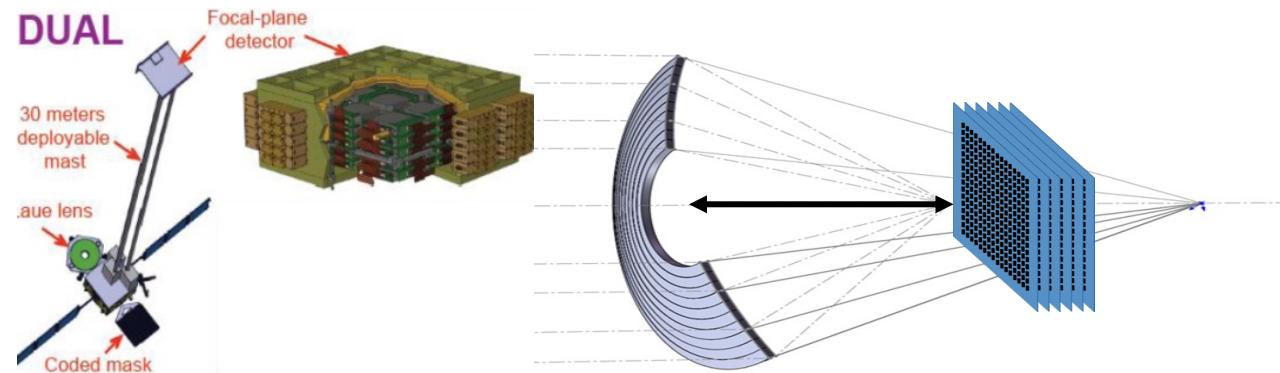
COSI



MeVGRO

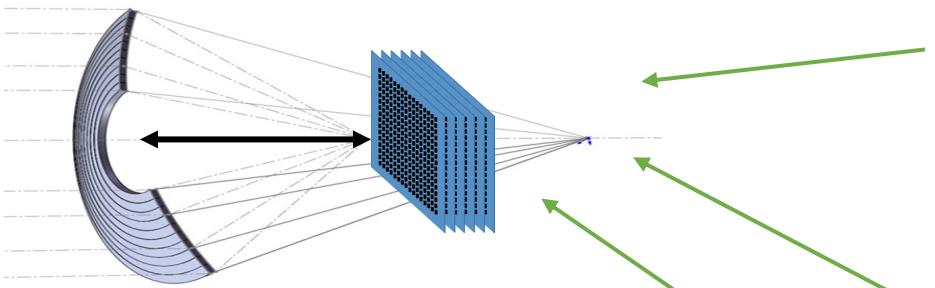


PAIRS



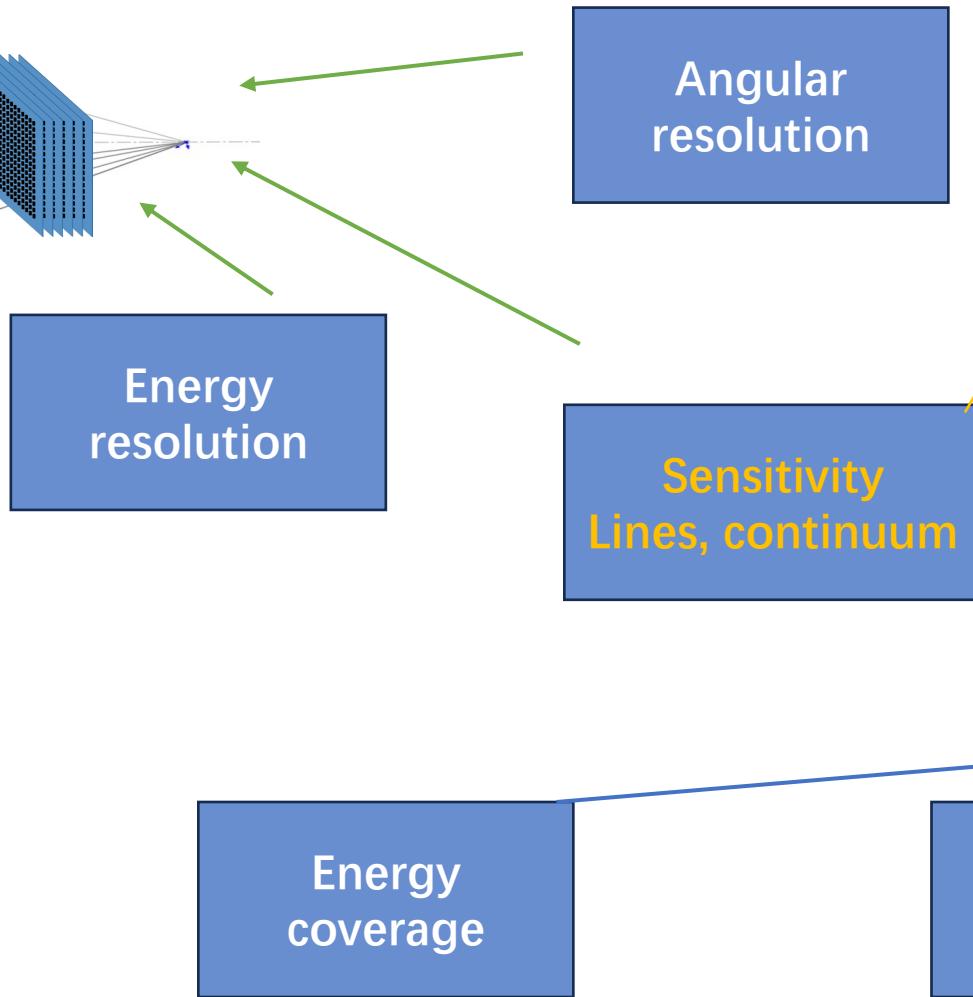
Future MeV missions

PAIRS



Line detection:
 $\leq 0.1\% @ 511\text{keV}$,
Angular resolution
 $\leq 2'$

Energy band:
450~550keV



Line sensitivity:
 $\leq 7.9 \times 10^{-6} \text{ ph/cm}^2/\text{s} @ 511\text{keV}$,
 $\leq 1.7 \times 10^{-6} \text{ ph/cm}^2/\text{s} @ 1.8\text{MeV}$
(2yrs survey of COSI)
 Continuum sensitivity:
 $\leq 2 \times 10^{-11} \text{ erg/cm}^2/\text{s} @ 1\text{MeV}$,
100MeV

Line detection:
 $\leq 1\% @ 662\text{keV}$,
Field of View:
 $\pm 65^\circ$,
Energy band:
0.2~5MeV

