

Probing New Physics with Double Beta Decay

Frank Deppisch

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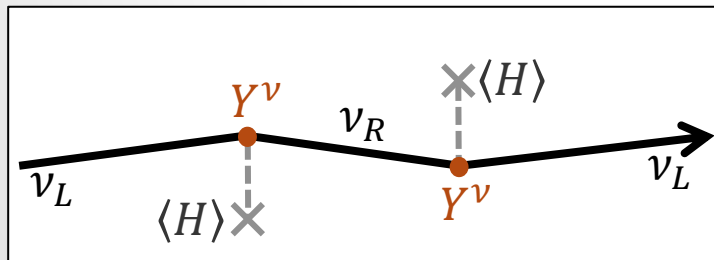
University College London

Dirac versus Majorana

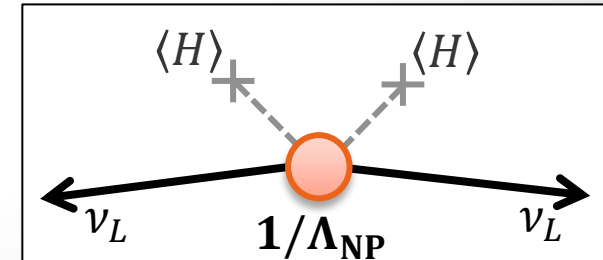
- ▶ Origin of neutrino masses beyond the Standard Model
- ▶ Two possibilities to define neutrino mass



Dirac mass analogous to other fermions but with $m_\nu / \Lambda_{EW} \approx 10^{-12}$ couplings to Higgs



Majorana mass, using only a left-handed neutrino
→ Lepton Number Violation



Beta Decay and Neutrino Mass

- ▶ Single beta decay

$$(A, Z) \rightarrow (A, Z + 1) + e^- + \bar{\nu}_e$$

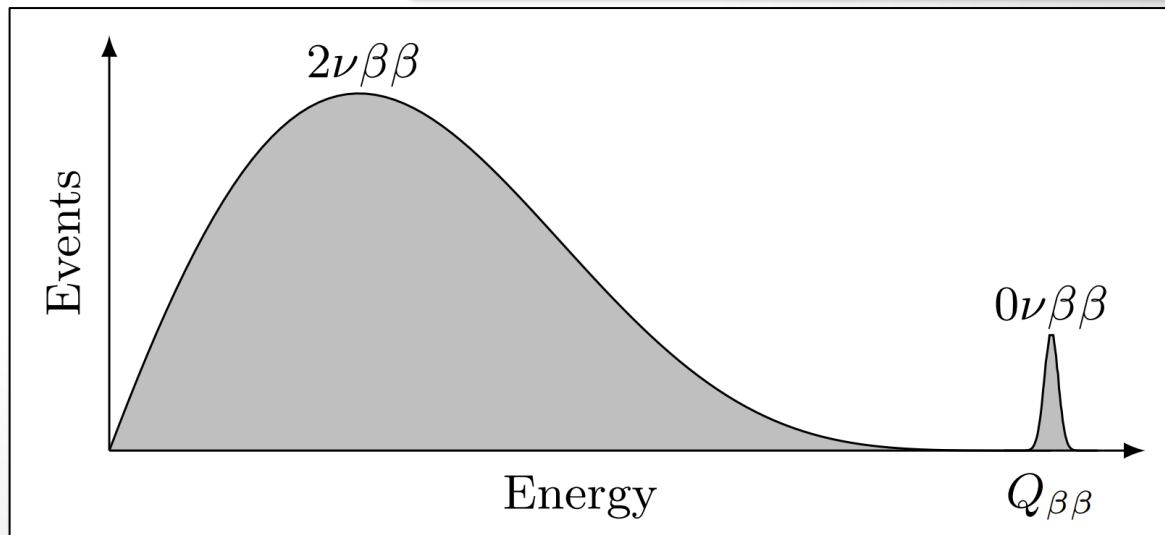
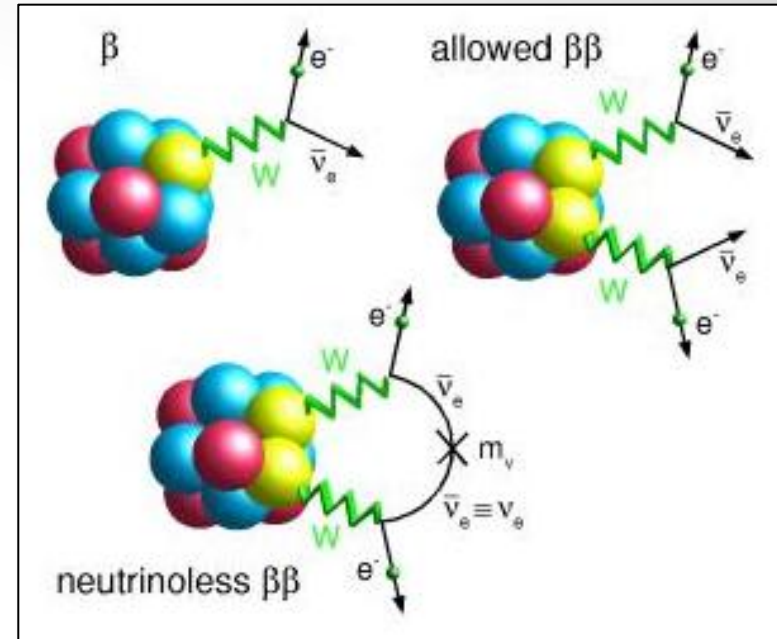
- ▶ Allowed double beta ($2\nu\beta\beta$) decay

$$(A, Z) \rightarrow (A, Z + 2) + 2e^- + 2\bar{\nu}_e$$

- ▶ Neutrinoless double beta ($0\nu\beta\beta$) decay

$$(A, Z) \rightarrow (A, Z + 2) + 2e^-$$

- Violation of lepton number
- Mediated by massive Majorana neutrinos

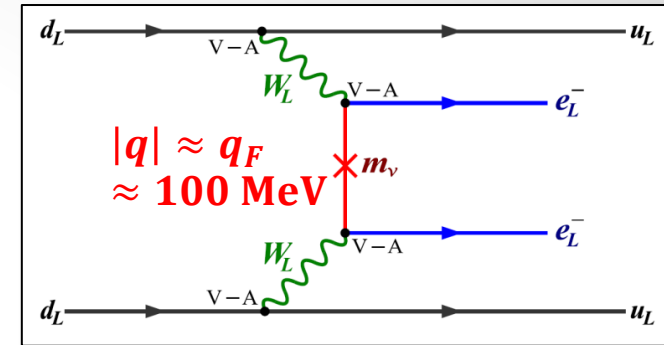


Neutrinoless Double β Decay

▶ Half-life

$$T_{1/2}^{-1} = |\mathbf{m}_{\beta\beta}|^2 \mathbf{G}^{0\nu} |\mathbf{M}^{0\nu}|^2$$

▶ Particle Physics



$$\mathcal{A}_{\mu\nu}^{lep} = \frac{1}{4} \sum_{i=1}^3 U_{ei}^2 \gamma_\mu (1 + \gamma_5) \frac{\not{q} + m_{\nu_i}}{q^2 - m_{\nu_i}^2} \gamma_\nu (1 - \gamma_5) \approx \frac{\gamma_\mu (1 + \gamma_5) \gamma_\nu}{4q^2} \sum_{i=1}^3 U_{ei}^2 m_{\nu_i} \rightarrow \mathbf{m}_{\beta\beta}$$

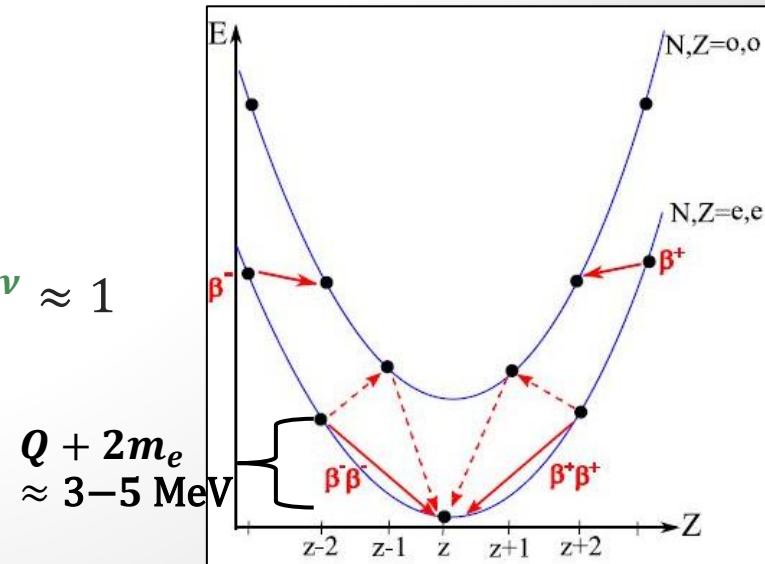
▶ Atomic Physics

- Leptonic phase space $\mathbf{G}^{0\nu} \propto Q^5$

▶ Nuclear Physics

- Nuclear transition matrix element $\mathbf{M}^{0\nu} \approx 1$ but large uncertainties, factor 2-3

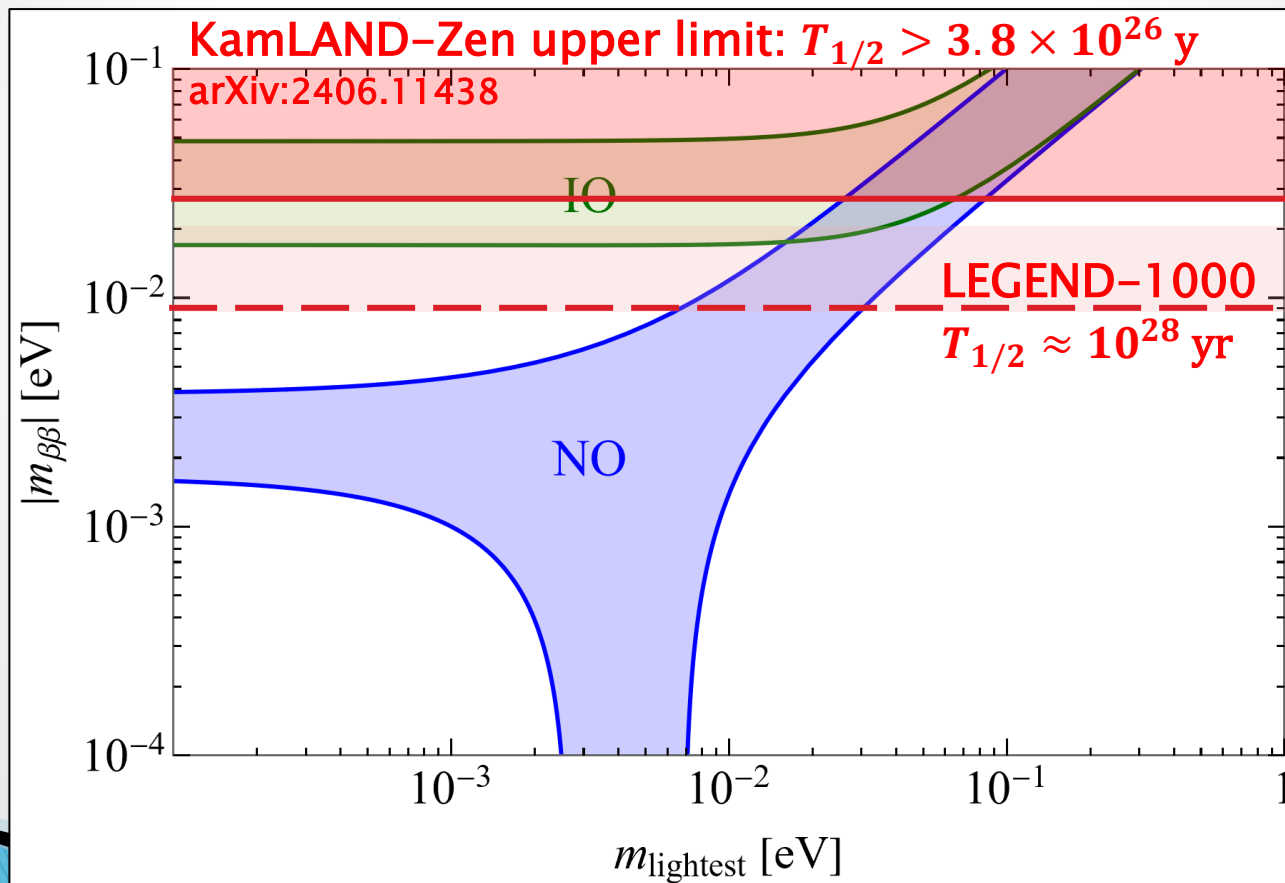
$$\frac{10^{25} \text{ y}}{T_{1/2}} \approx \left(\frac{|\mathbf{m}_{\beta\beta}|}{0.1 \text{ eV}} \right)^2$$



Three Active Neutrinos

► Effective $0\nu\beta\beta$ Mass degenerate ν & $\theta_{13} \approx 0$

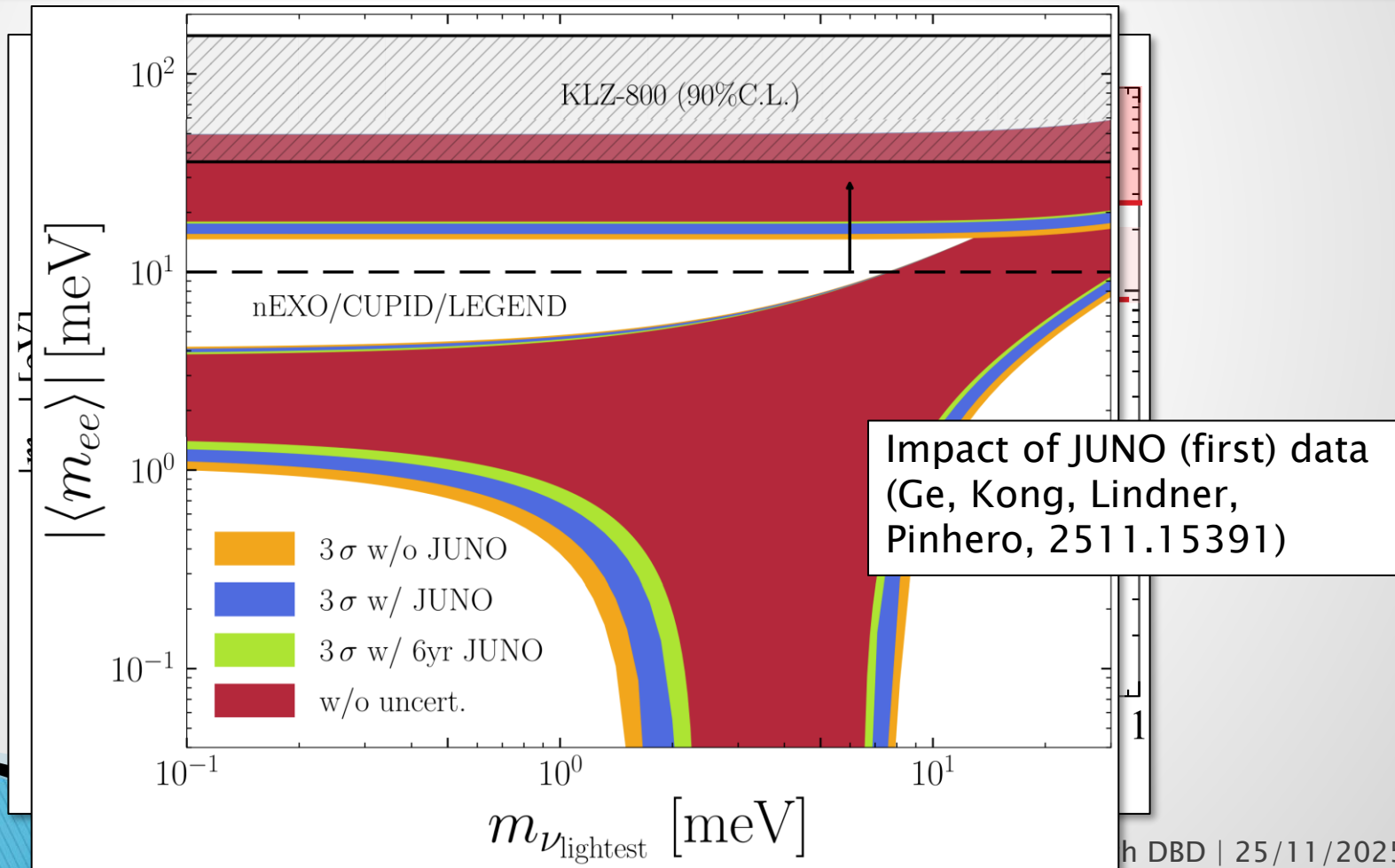
$$|m_{\beta\beta}| = |c_{12}^2 c_{13}^2 m_{\nu_1} + s_{12}^2 c_{13}^2 m_{\nu_2} e^{i\phi_{12}} + s_{13}^2 m_{\nu_3} e^{i\phi_{13}}| \approx m_{\nu} \sqrt{1 - \sin^2(2\theta_{12}) \sin^2(\phi_{12}/2)}$$



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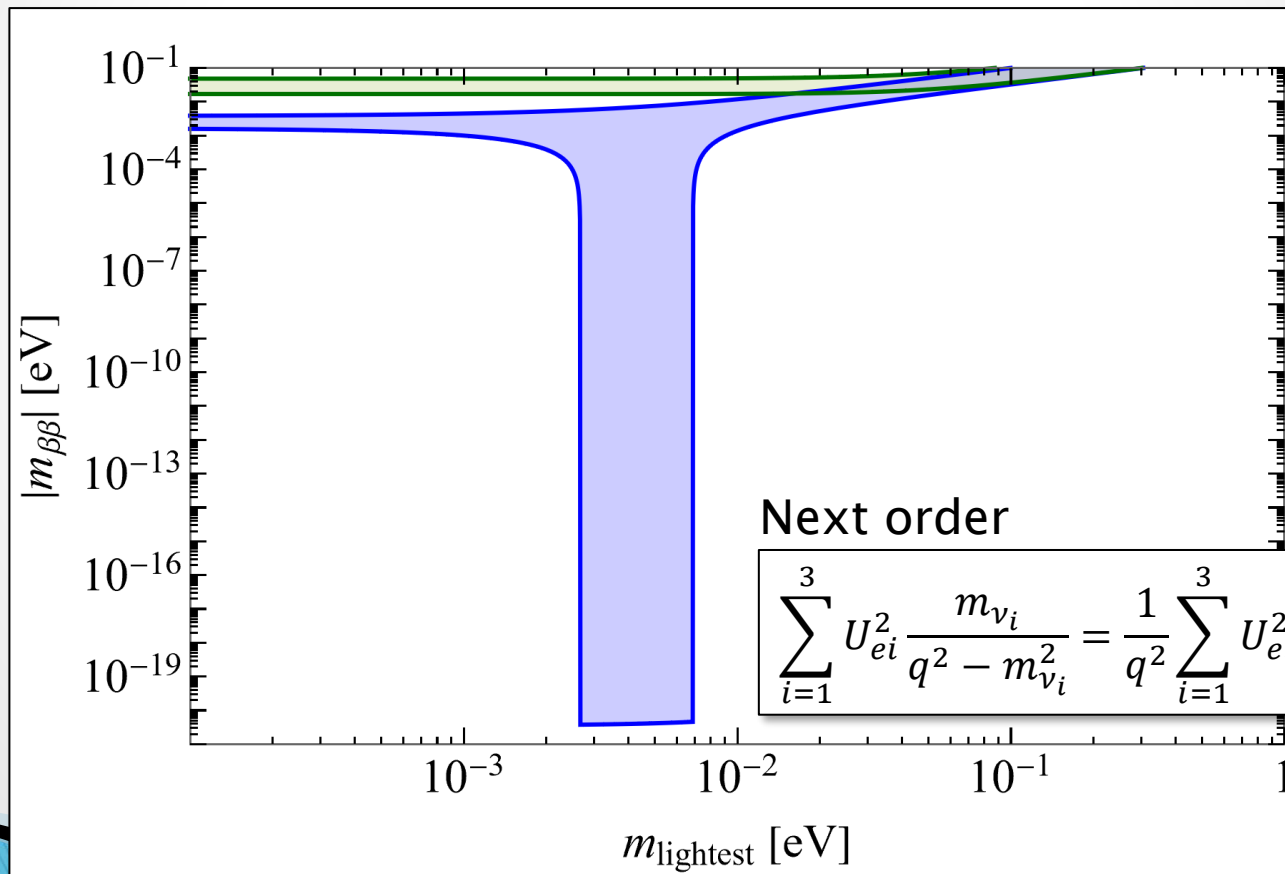
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Three Active Neutrinos

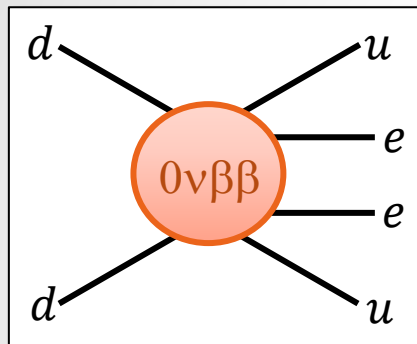
▶ Effective $0\nu\beta\beta$ Mass

$$m_{\beta\beta} = 0 \not\Rightarrow \Gamma_{0\nu\beta\beta} = 0$$

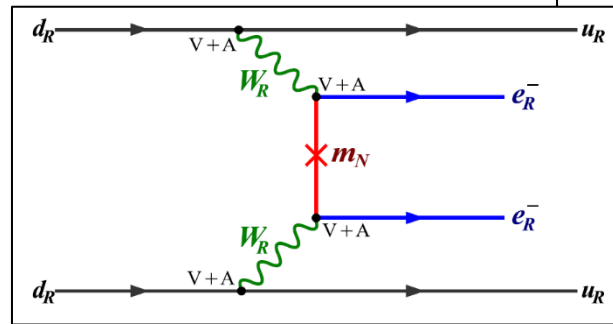


New Physics and $0\nu\beta\beta$

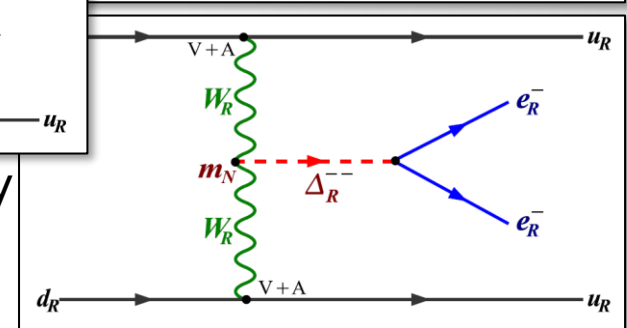
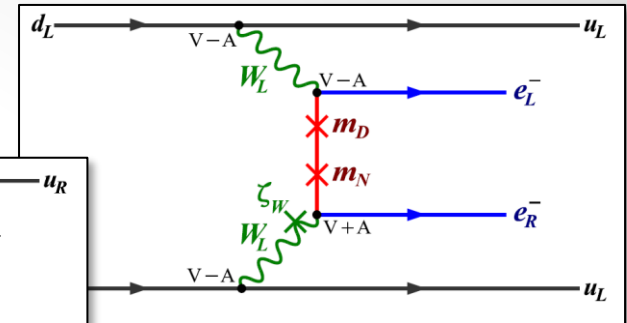
► Plethora of New Physics scenarios



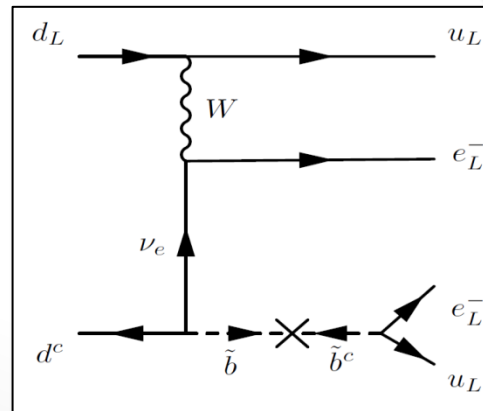
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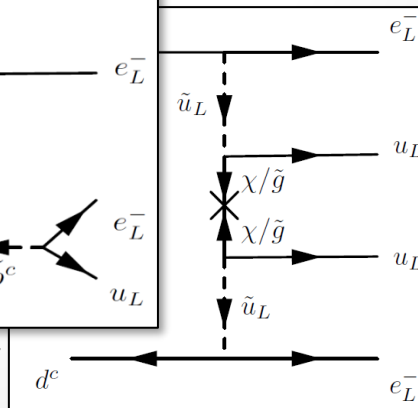
Left-Right Symmetry



$$T_{1/2}^{-1} = \epsilon_{NP}^2 G_{NP}^{0\nu} |M_{NP}^{0\nu}|^2$$



R-Parity
Violating SUSY



Extra
Dimensions

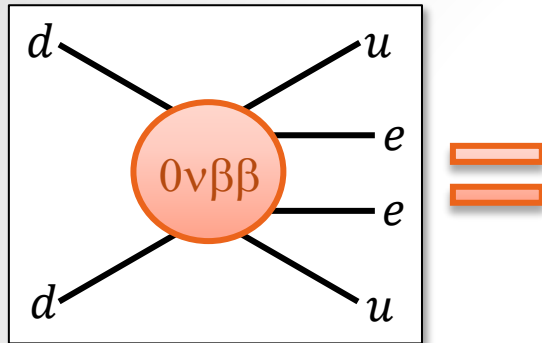
Majorons

Leptoquarks

...

New Physics and $0\nu\beta\beta$

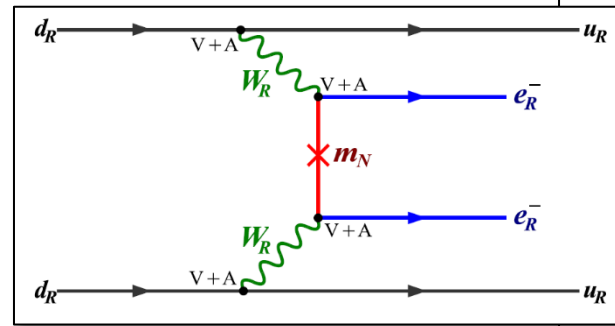
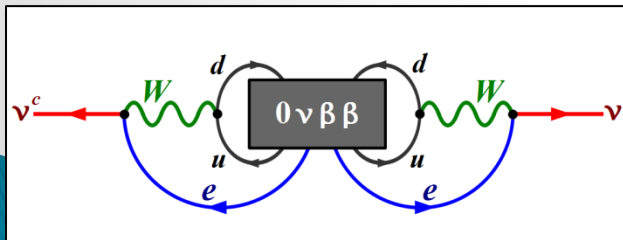
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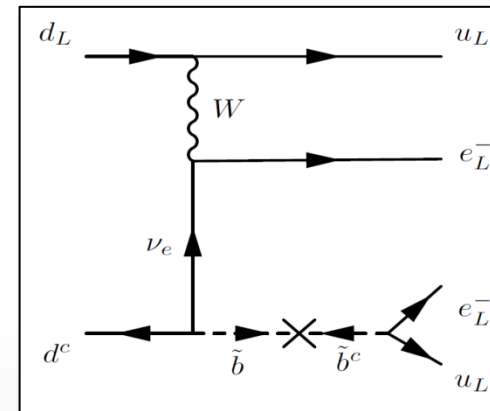
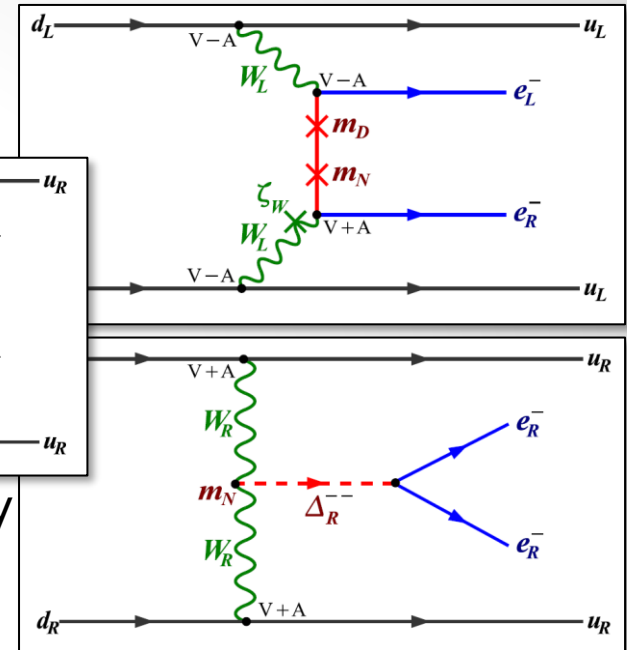
$$T_{1/2}^{-1} = \epsilon_{NP}^2 G_{NP}^{0\nu} |M_{NP}^{0\nu}|^2$$

► Neutrinos still Majorana

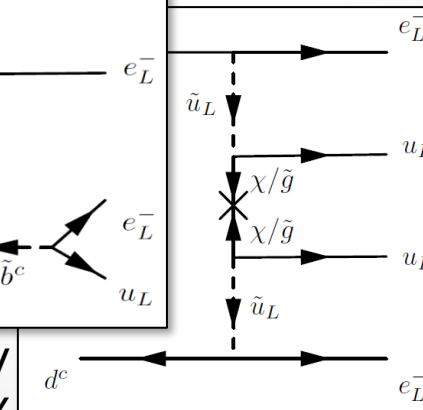
Schechter, Valle, Phys.Rev.D
25 (1982) 2951



Left-Right Symmetry



R-Parity
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New Physics and $0\nu\beta\beta$

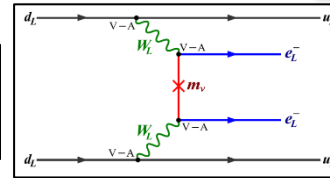
► Decay modes, rates and NP sensitivity

- $2\nu\beta\beta$ (SM)

$$\Gamma_{SM}^{2\nu} = G_F^4 Q^{11} p_F^{-2} \approx (10^{19} \text{ yr})^{-1}$$

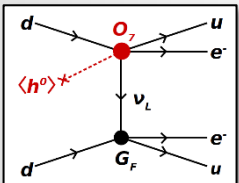
- $0\nu\beta\beta$ via Majorana neutrinos

$$\Gamma_{m_\nu}^{0\nu} = m_\nu^2 G_F^4 Q^5 p_F^2 \approx \left(\frac{m_\nu}{0.1 \text{ eV}} \right)^2 (10^{25} \text{ yr})^{-1}$$



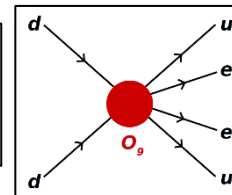
- $0\nu\beta\beta$ via BSM 7-dim LNV operator

$$\Gamma_{LR}^{0\nu} = \Lambda^{-6} G_F^4 Q^5 p_F^4 \approx \left(\frac{20 \text{ TeV}}{\Lambda} \right)^6 (10^{25} \text{ yr})^{-1}$$



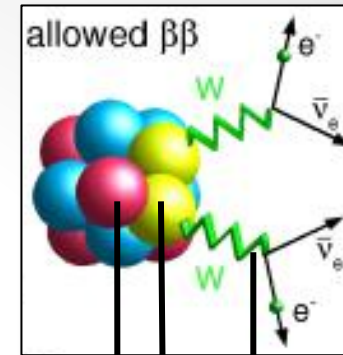
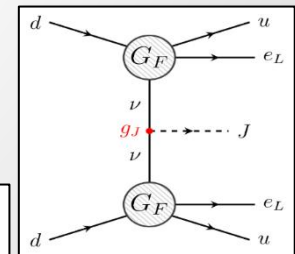
- $0\nu\beta\beta$ via BSM 9-dim LNV operator

$$\Gamma_{SR}^{0\nu} = \Lambda^{-10} Q^5 p_F^6 \approx \left(\frac{4 \text{ TeV}}{\Lambda} \right)^{10} (10^{25} \text{ yr})^{-1}$$



- $0\nu\beta\beta J$ ordinary Majoron emission

$$\Gamma^{0\nu J} = g^2 G_F^4 Q^7 p_F^2 \approx \left(\frac{g}{10^{-5}} \right)^2 (10^{21} \text{ yr})^{-1}$$



$Q \sim 1 \text{ MeV}$
energy
release

$G_F \sim 10^{-5} \text{ GeV}^{-2} \sim v_{EW}^{-2}$
electroweak scale

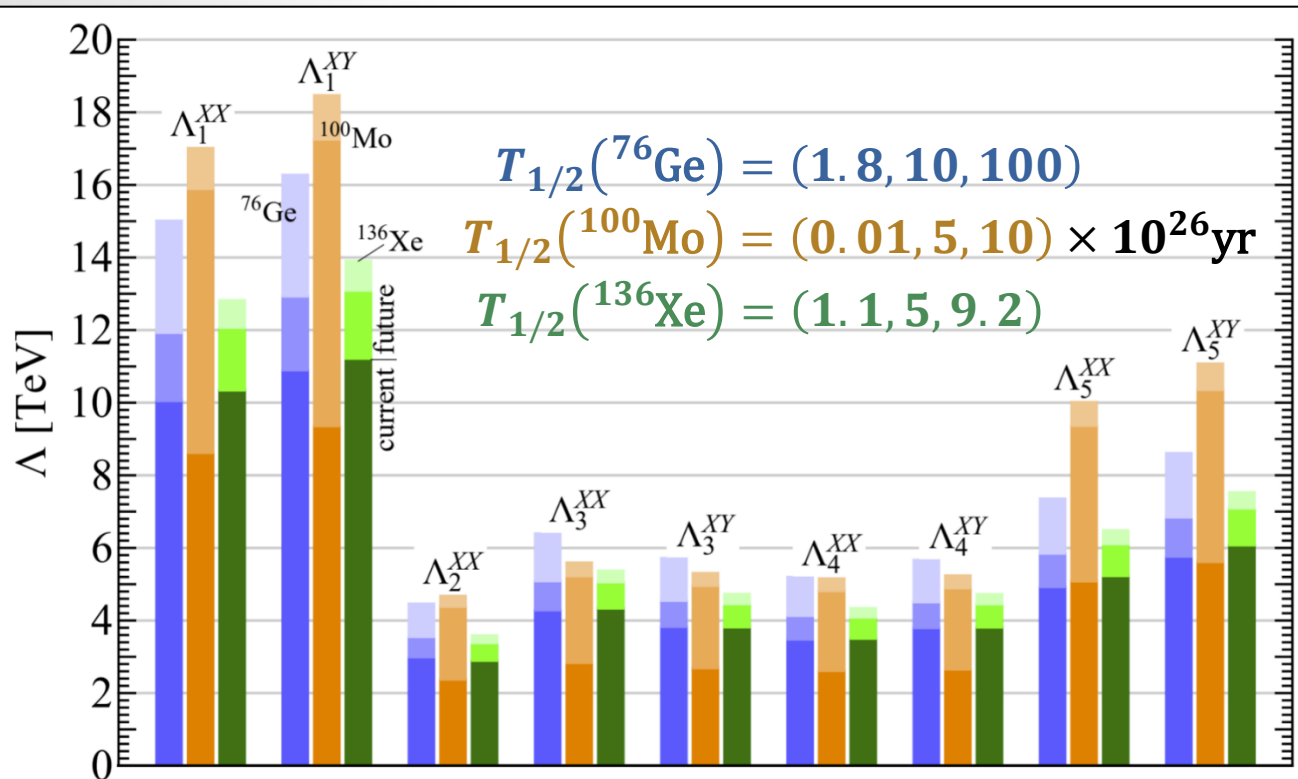
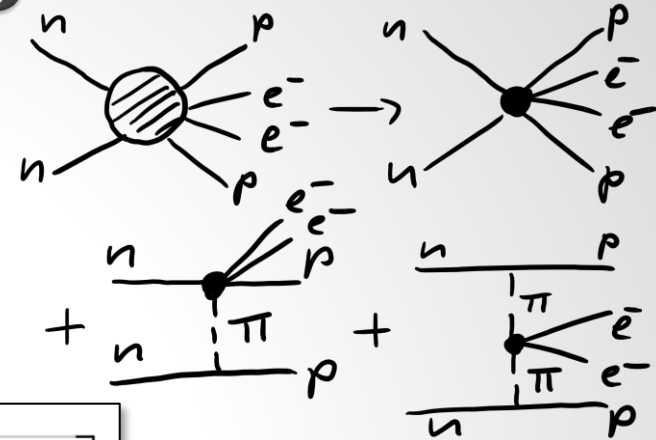
$p_F \sim 100 \text{ MeV} \sim 1 \text{ fm}^{-1}$
inter-nucleon distance

New Physics and $0\nu\beta\beta$

FFD, Graf, Iachello, Kotila, PRD 102 (2020)

► Limits on short-range operators

- NMEs from IBM-2 with $g_A = 1.0$ and short-range correlations in Argonne parametrization



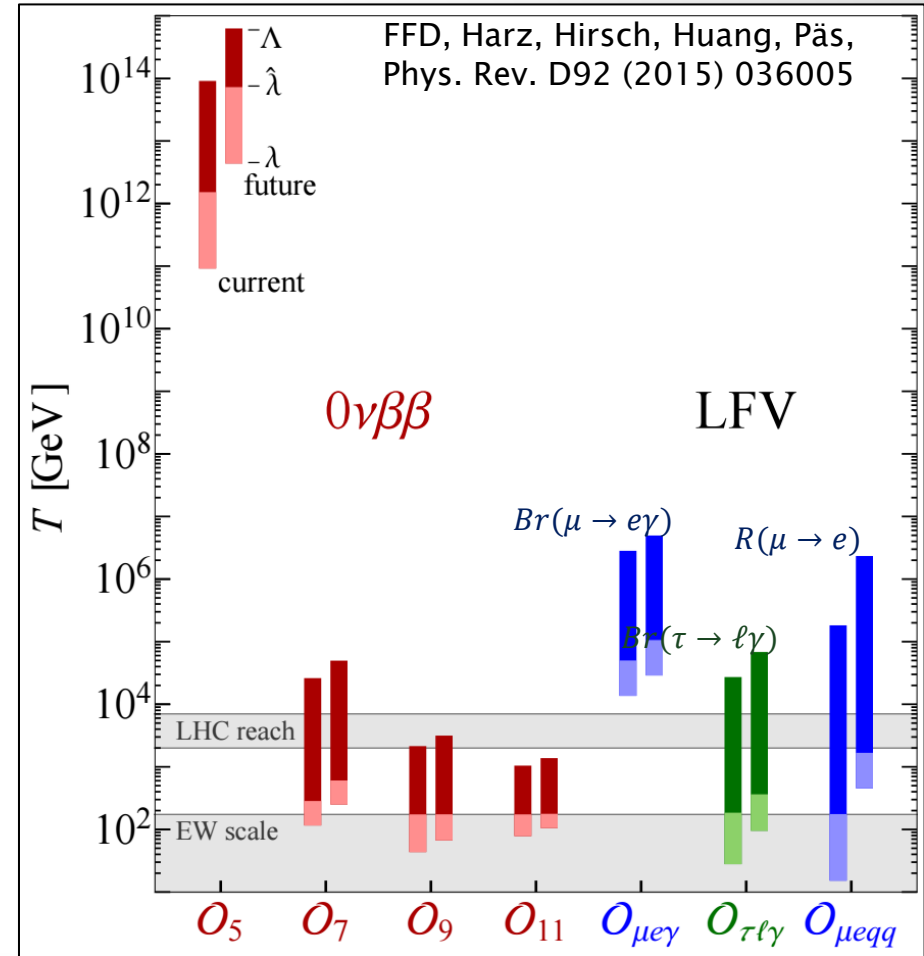
Pion-mediated contributions

- R-parity violating SUSY (Faessler, Kovalenko, Simkovic, Schwieger, Phys.Rev.Lett. 78 (1997) 183)
- Chiral EFT with Pion operators from Lattice QCD (Cirigliano, Dekens, de Vries, Graesser, Mereghetti, JHEP 1812 (2018) 097)

Falsifying Baryogenesis

FFD, Harz, Hirsch, PRL 112 (2014) | FFD, Harz, Hirsch, Huang, Päs, PRD 92 (2015)

- ▶ LNV processes lead to washout of lepton asymmetry in early universe
- ▶ Observation of LNV
 - gives information at what temperatures operators are in equilibrium
 - **can falsify high-scale baryogenesis scenarios**



New Physics in $2\nu\beta\beta$

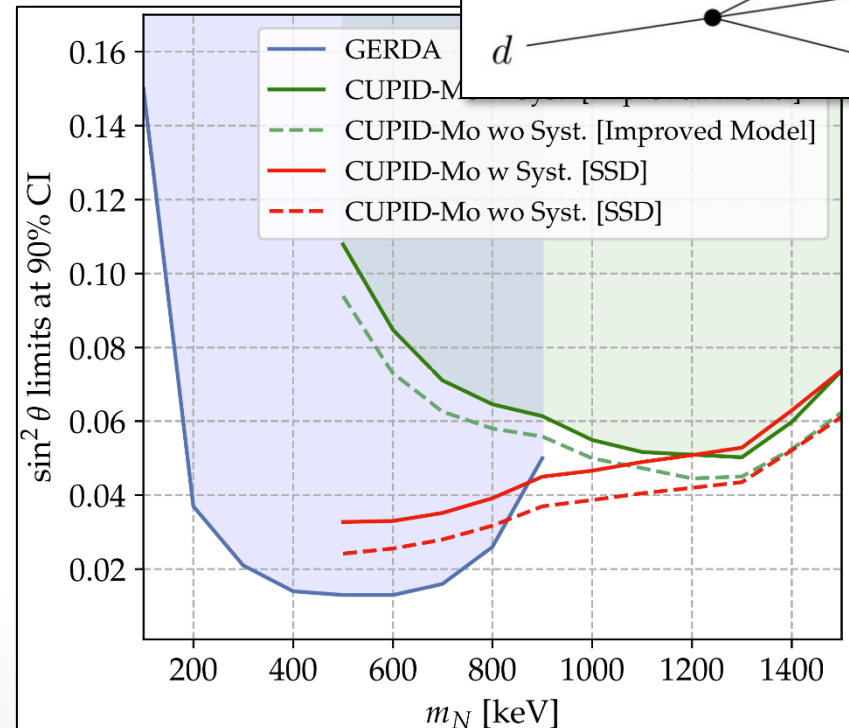
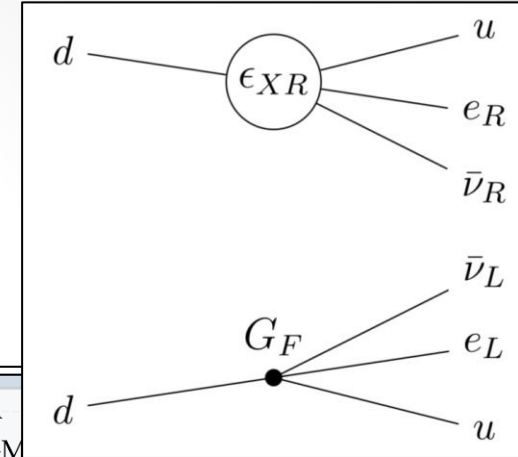
FFD, Graf, Simkovic, PRL 125 (2020) | FFD, Graf, Rodejohann, Xu, PRD 102 (2020)

► Sterile neutrino search through energy endpoint

(also: Agostini, Bossio, Ibarra, Marcano, PLB 815 (2021))

- Emission of one sterile neutrino in double beta decay: $\nu N\beta\beta$
- Observed limits:

- GERDA (JCAP 12 (2022) 012)
- CUPID-Mo (Eur.Phys.J.C 84 (2024) 9)



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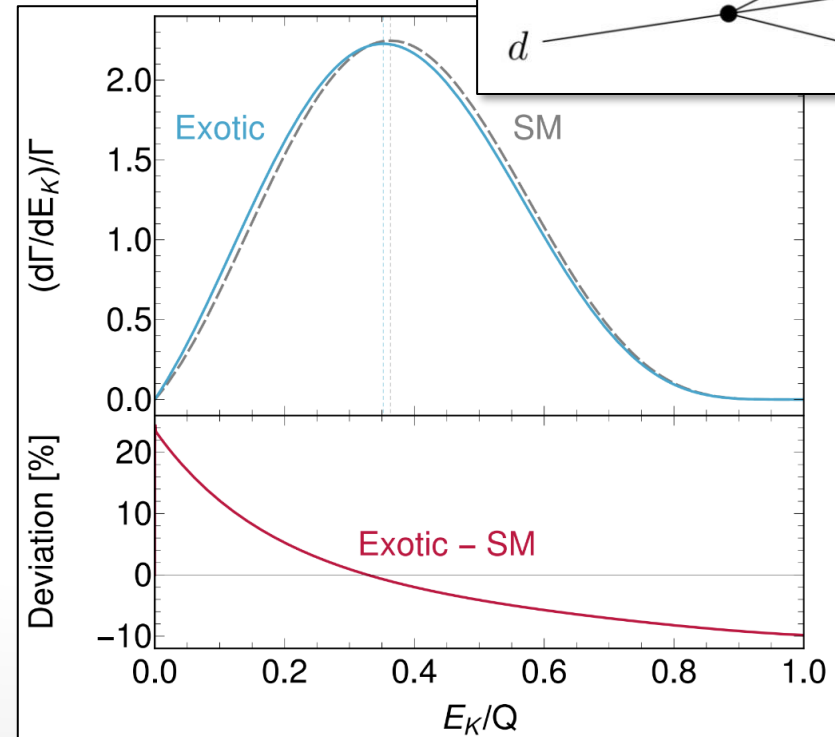
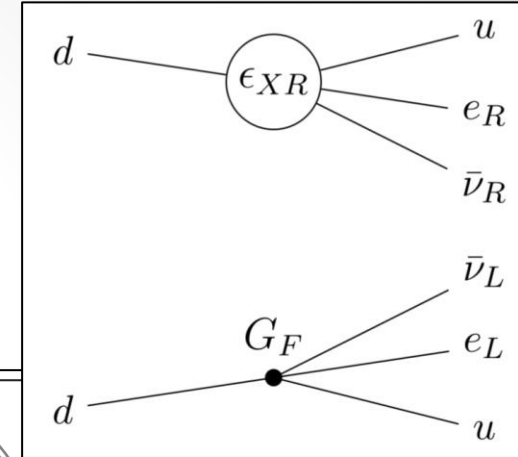
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► Lepton number conserving RH currents

- LEGEND analysis ongoing



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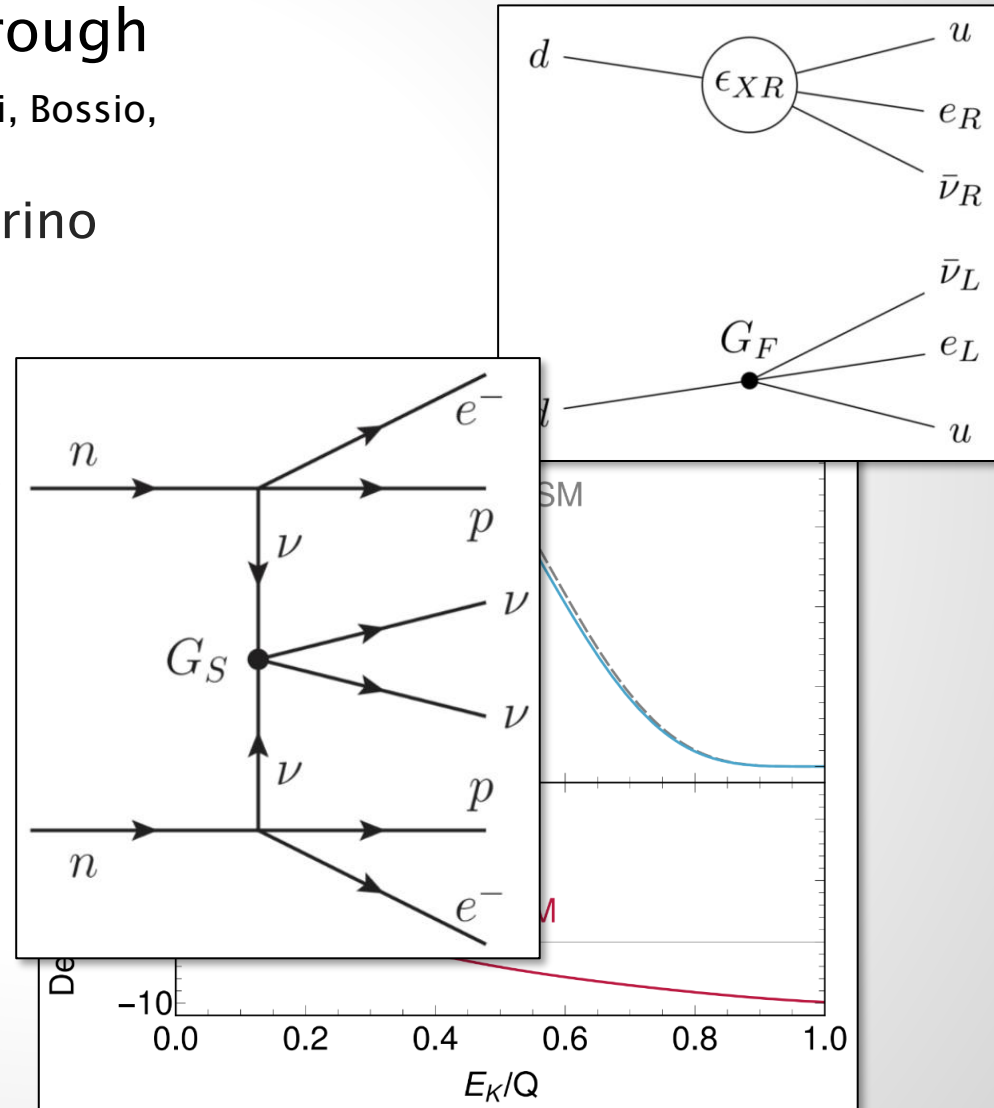
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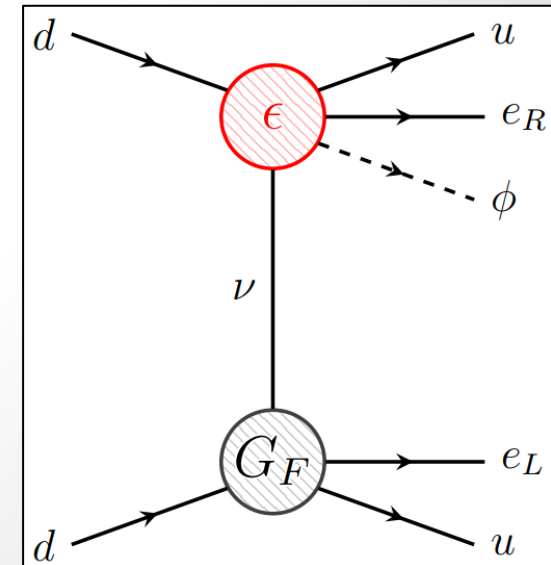
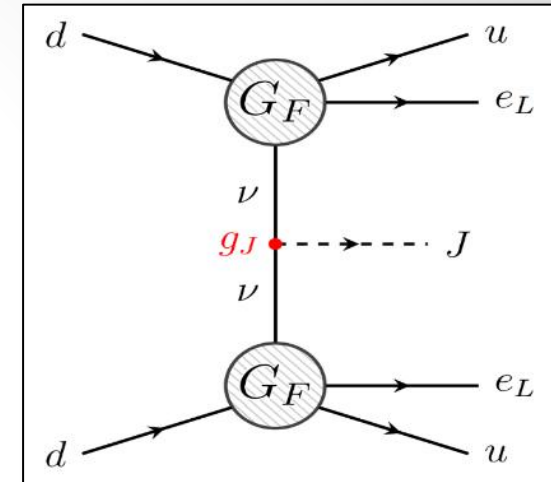
► Neutrino self-interactions



Exotic Particle Emission

Cepedello, FFD, González, Hati, Hirsch, PRL 122 (2019) | Boudjema, FFD, Herrero, Majumdar, Senapati, 2511.13606

- ▶ Majoron(-like) J emission
- ▶ Majoron-like ϕ emission assisted by RH current
 - Non-standard energy distribution
 - Searched for in EXO-200: $T_{1/2} > 4 \times 10^{24}$ y (PRD 104 (2021) 11, 112002)



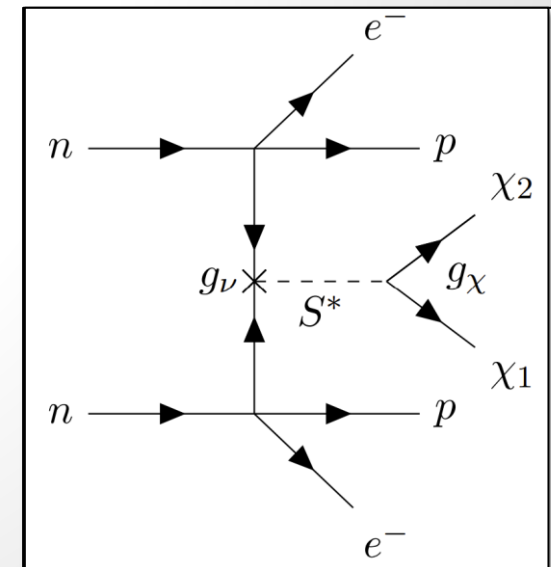
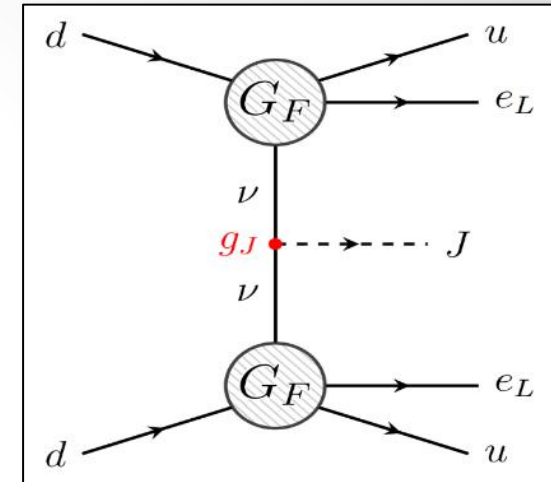
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- ▶ Dirac neutrinos coupled to massive, L-charged “Majoron” S and dark fermions χ

$$-\mathcal{L} = \bar{\nu}^c (a_\nu P_L + b_\nu P_R) \nu S^* + \bar{\chi}^c (a_\chi P_L + b_\chi P_R) \chi S^* + \text{h.c.}$$

- Z_2 symmetry to forbid mixing
- Self-interacting neutrinos and self-interacting (potentially) DM



Neutrinos coupled to DS

Boudjema, FFD, Herrero, Majumdar, Senapati, 2511.13606

- ▶ Model I': Dirac neutrinos coupled to massive, L-charged scalar S and dark fermions χ

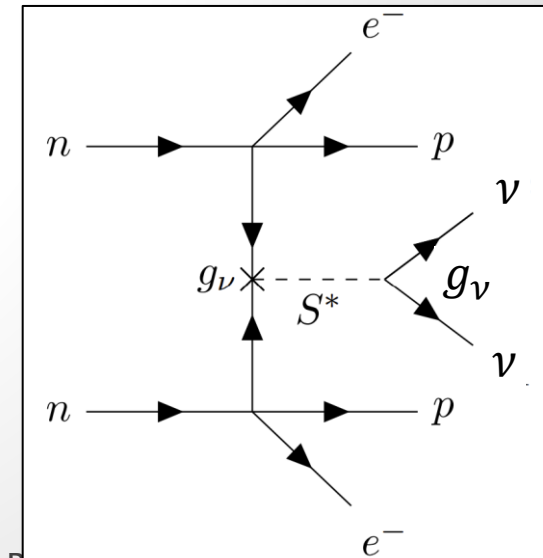
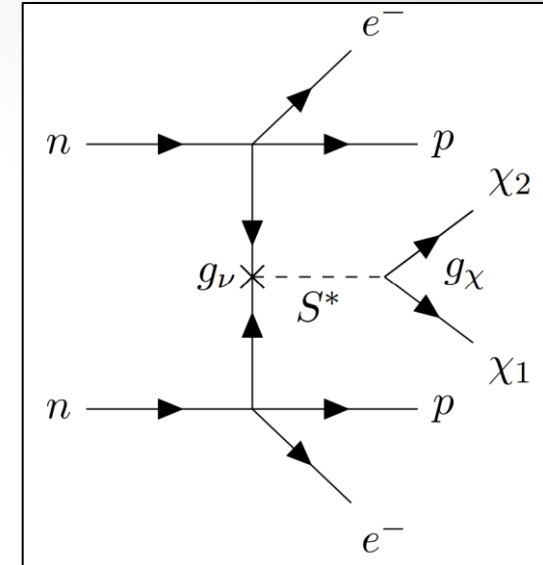
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- Z_2 symmetry to forbid mixing
- Self-interacting neutrinos and self-interacting (potentially) DM

- ▶ Model I: No Z_2 , mixing terms allow

$$-\mathcal{L} = m_\nu \bar{\nu}'_R \nu'_L + m_{\nu\chi} \bar{\chi}'_R \nu'_L + m_{\chi\nu} \bar{\nu}'_R \chi'_L + M \bar{\chi}'_R \chi'_L \\ + (a_\nu \bar{\nu}'^c_L \nu'_L S^* + b_\nu \bar{\nu}'^c_R \nu'_R S^* + a_\chi \bar{\chi}'^c_L \chi'_L S^* + b_\chi \bar{\chi}'^c_R \chi'_R S^* + \text{h.c.})$$

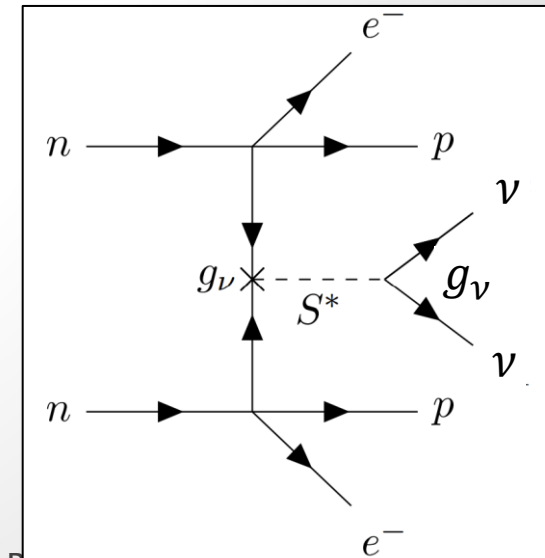
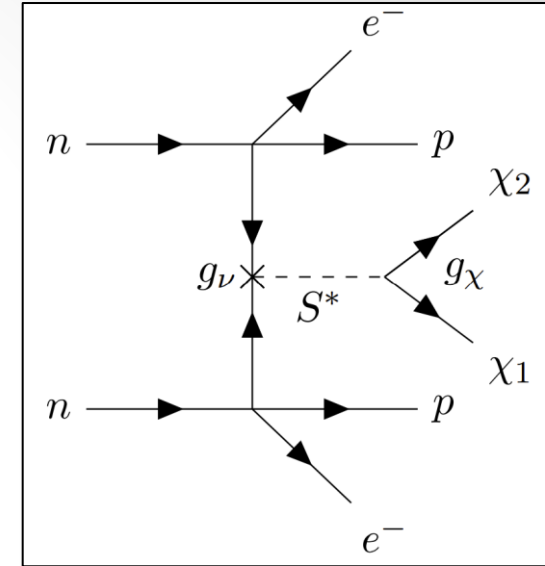
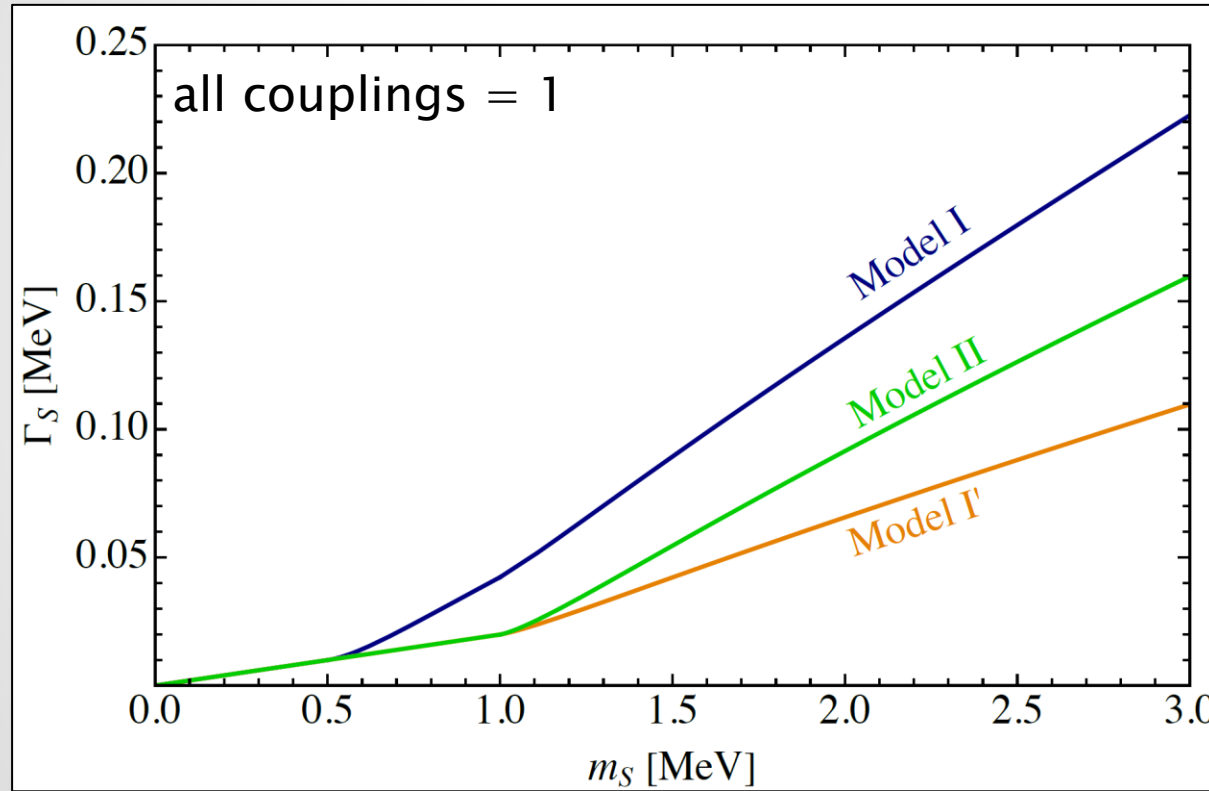
- ▶ Dirac neutrinos not required, but here, no constraints from $0\nu\beta\beta$



Neutrinos coupled to DS

Boudjema, FFD, Herrero, Majumdar, Senapati, 2511.13606

- ▶ Small but sizeable S decay width possible



Neutrinos coupled to DS

Boudjema, FFD, Herrero, Majumdar, Senapati, 2511.13606

► Rich double beta decay phenomenology

- $2\chi_S\beta\beta$

$$\frac{d\Gamma_S^{2\chi}}{dT} = \kappa_{0\nu} \left(\frac{m_e}{2R}\right)^2 |\mathcal{M}_{0\nu}|^2 |a_\nu|^2 \left\{ (|a_\chi|^2 + |b_\chi|^2) I_1(T) - 2\text{Re}[a_\chi^* b_\chi] I_2(T) \right\} g(T) \\ \times \Theta(Q - m_{\chi_1} - m_{\chi_2} - T),$$

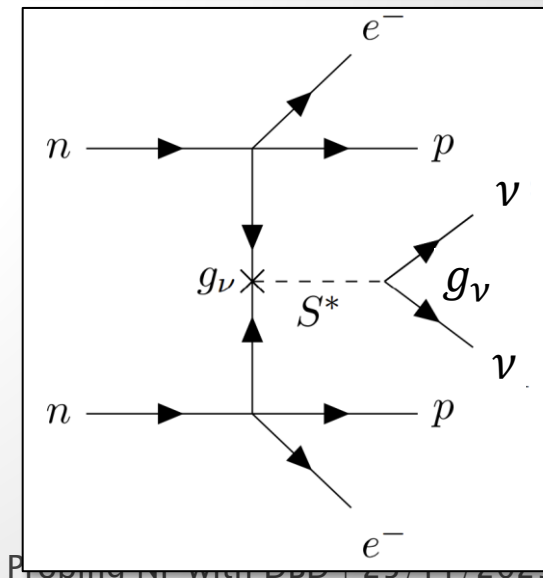
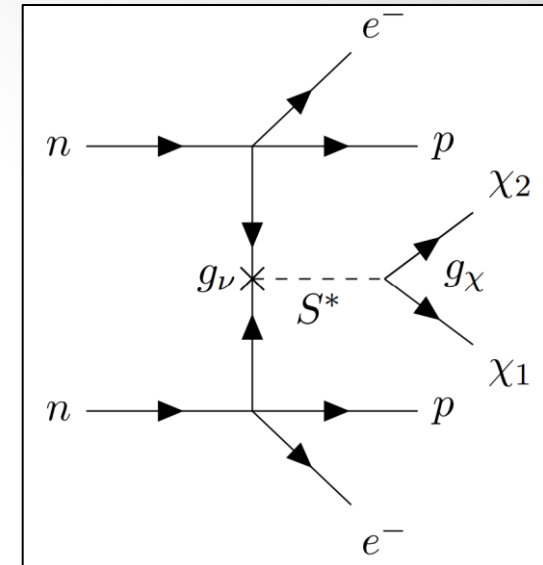
$$I_1(T) = \frac{1}{4(2\pi)^4} \int_{m_{\chi_1}}^{Q-T-m_{\chi_2}} dE_3 \left\{ \frac{m_S^2 - m_{\chi_1}^2 - m_{\chi_2}^2}{2m_S\Gamma_S} [\arctan \zeta_+(E_3) + \arctan \zeta_-(E_3)] \right. \\ \left. + \frac{1}{4} \log \left[\frac{\zeta_-^2(E_3) + 1}{\zeta_+^2(E_3) + 1} \right] \right\},$$

$$I_2(T) = \frac{1}{4(2\pi)^4} \frac{m_{\chi_1} m_{\chi_2}}{m_S \Gamma_S} \int_{m_{\chi_1}}^{Q-T-m_{\chi_2}} dE_3 [\arctan \zeta_+(E_3) + \arctan \zeta_-(E_3)].$$

- $2\nu_S\beta\beta$

- Including interference with SM $2\nu\beta\beta$

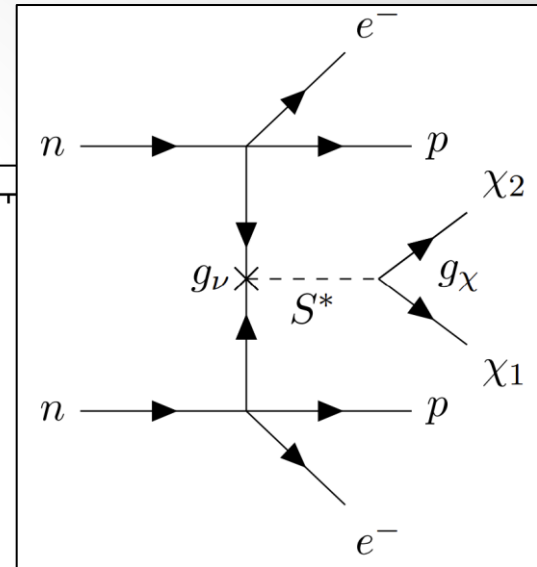
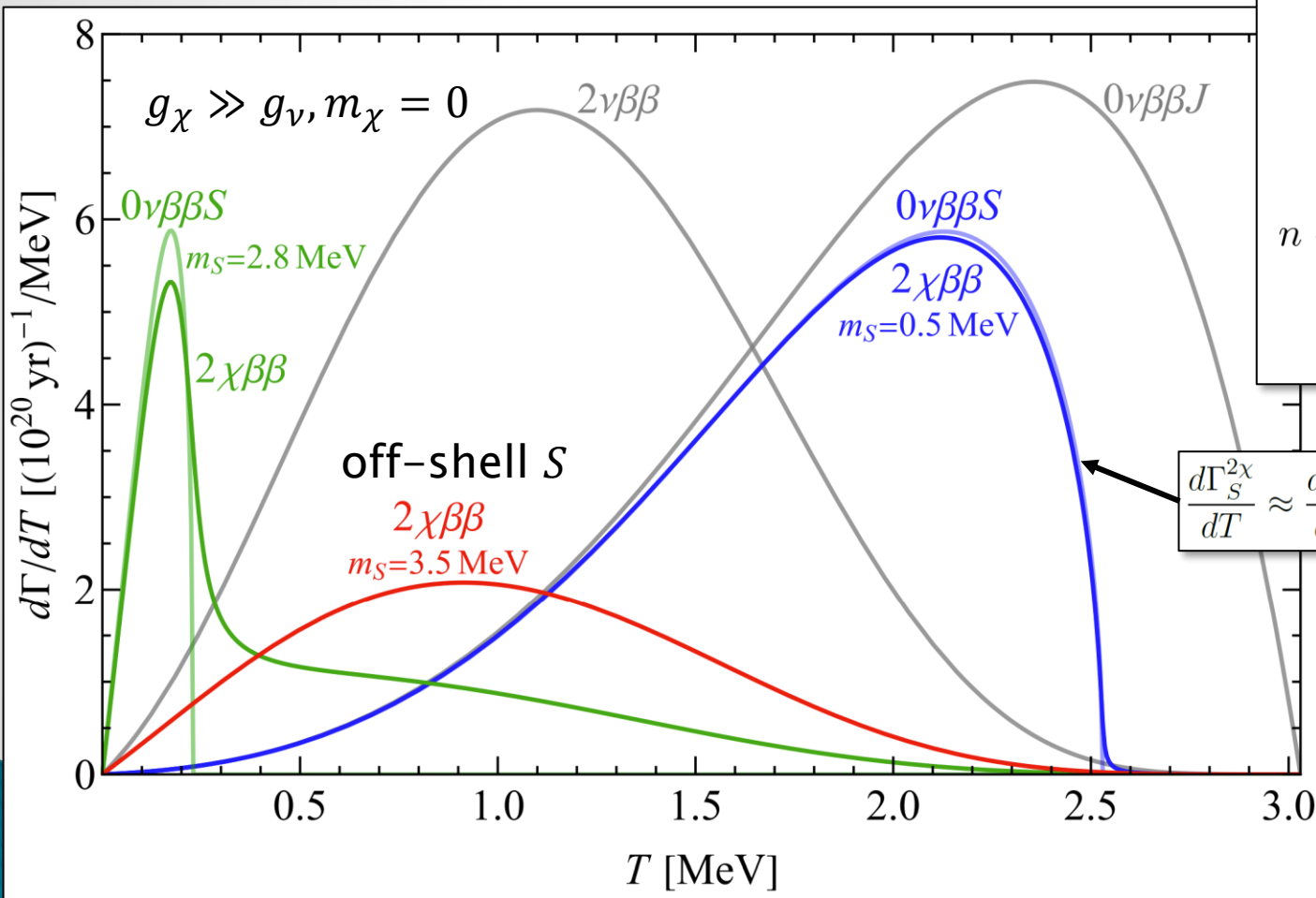
$$\zeta_\pm(E_3) = \frac{2|\mathbf{p}_3||\mathbf{p}_4|(x \pm 1)}{m_S \Gamma_S x}$$



Neutrinos coupled to DS

Boudjema, FFD, Herrero, Majumdar, Senapati, 2511.13606

► Finite S width/off-shell S



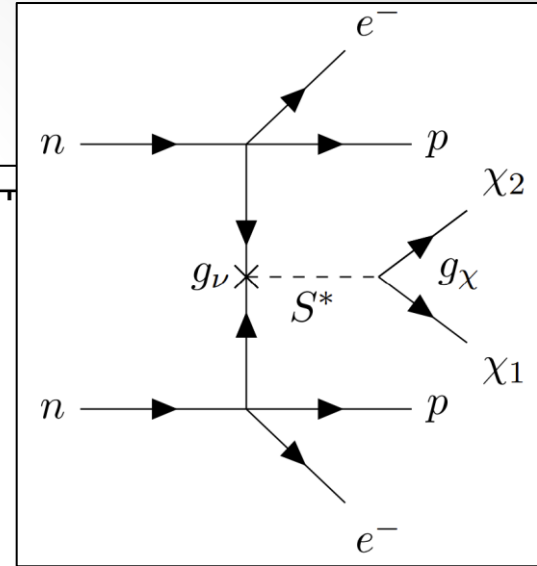
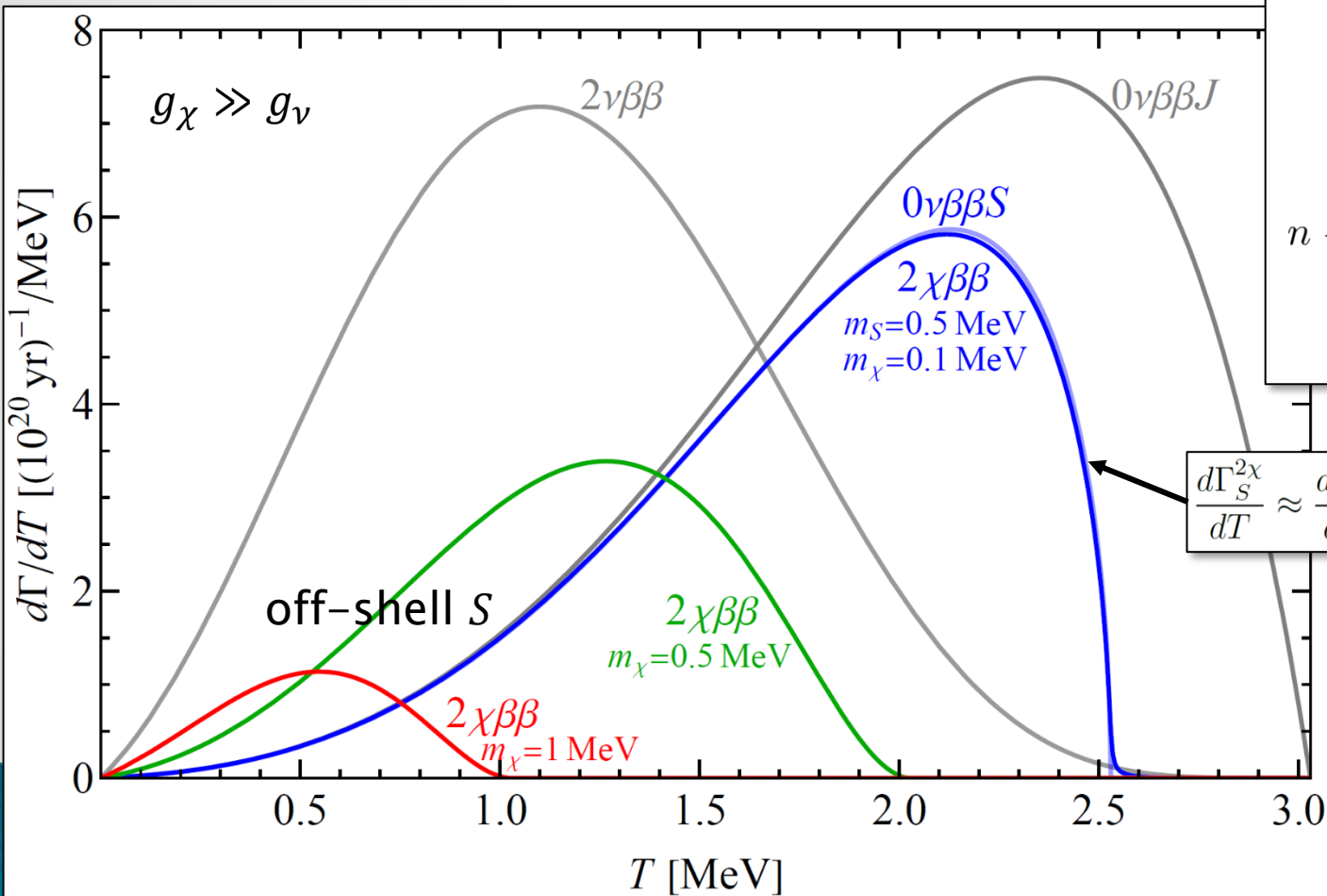
$$\frac{d\Gamma_S^{2\chi}}{dT} \approx \frac{d\Gamma_S}{dT} \times \text{Br}(S \rightarrow \chi\chi)$$

On-shell S in narrow-width approximation

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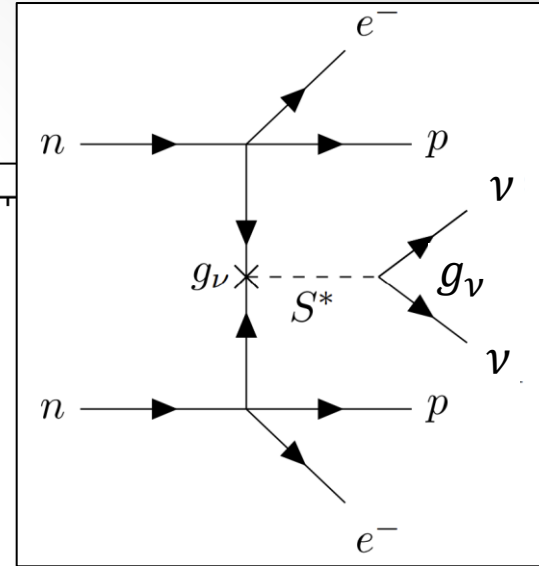
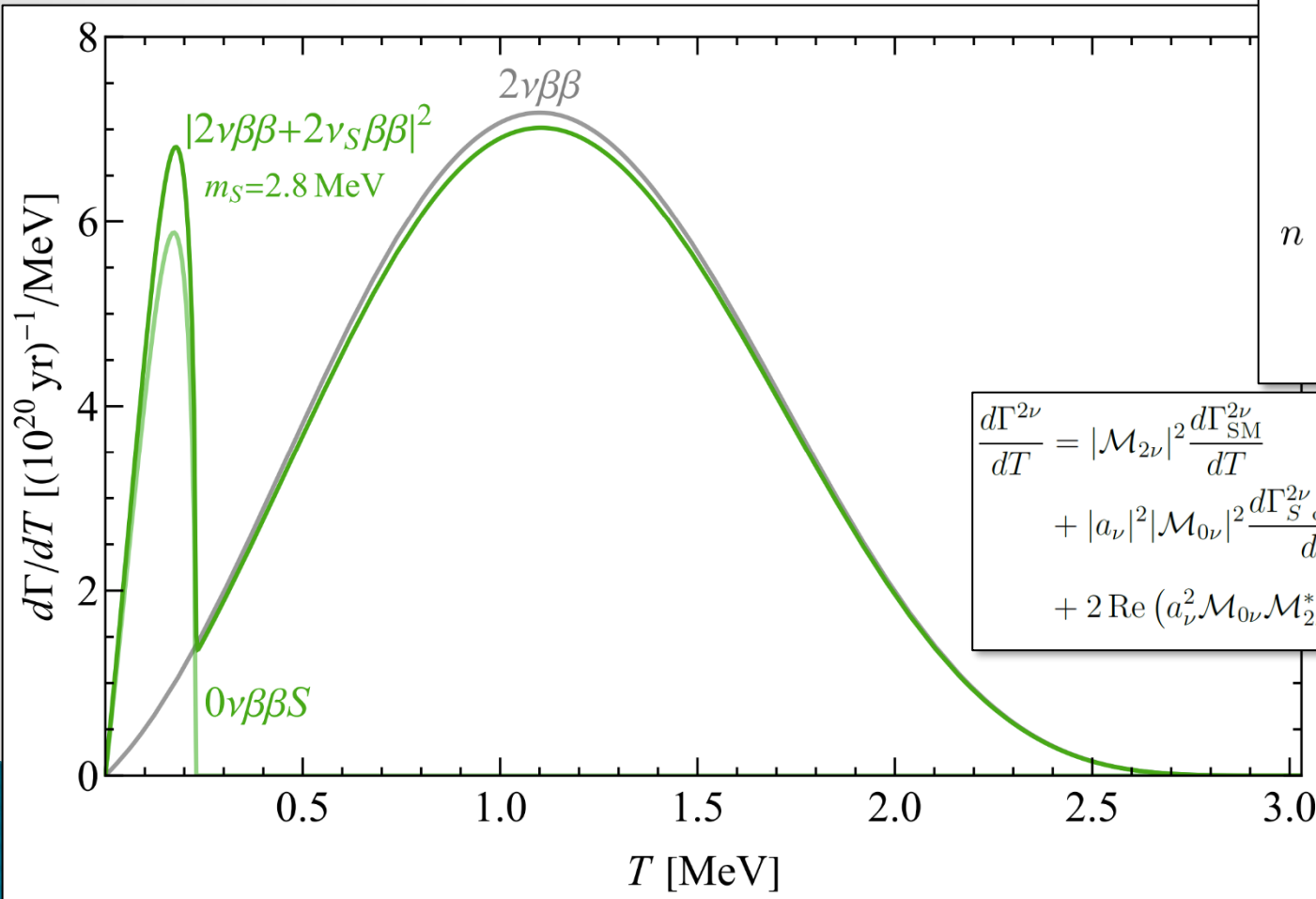


On-shell S in narrow-width approximation

Neutrinos coupled to DS

Boudjema, FFD, Herrero, Majumdar, Senapati, 2511.13606

► $2\nu_S\beta\beta$ and interference with SM $2\nu\beta\beta$



$$\begin{aligned} \frac{d\Gamma^{2\nu}}{dT} = & |\mathcal{M}_{2\nu}|^2 \frac{d\Gamma_{\text{SM}}^{2\nu}}{dT} \\ & + |a_\nu|^2 |\mathcal{M}_{0\nu}|^2 \frac{d\Gamma_S^{2\nu} \text{ on-shell}}{dT} + |a_\nu|^4 |\mathcal{M}_{0\nu}|^2 \frac{d\Gamma_S^{2\nu} \text{ off-shell}}{dT} \\ & + 2 \text{Re} (a_\nu^2 \mathcal{M}_{0\nu} \mathcal{M}_{2\nu}^*) \frac{d\Gamma_{\text{int}}^{2\nu}}{dT}, \end{aligned}$$

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► Sensitivity of current and future double beta decay searches

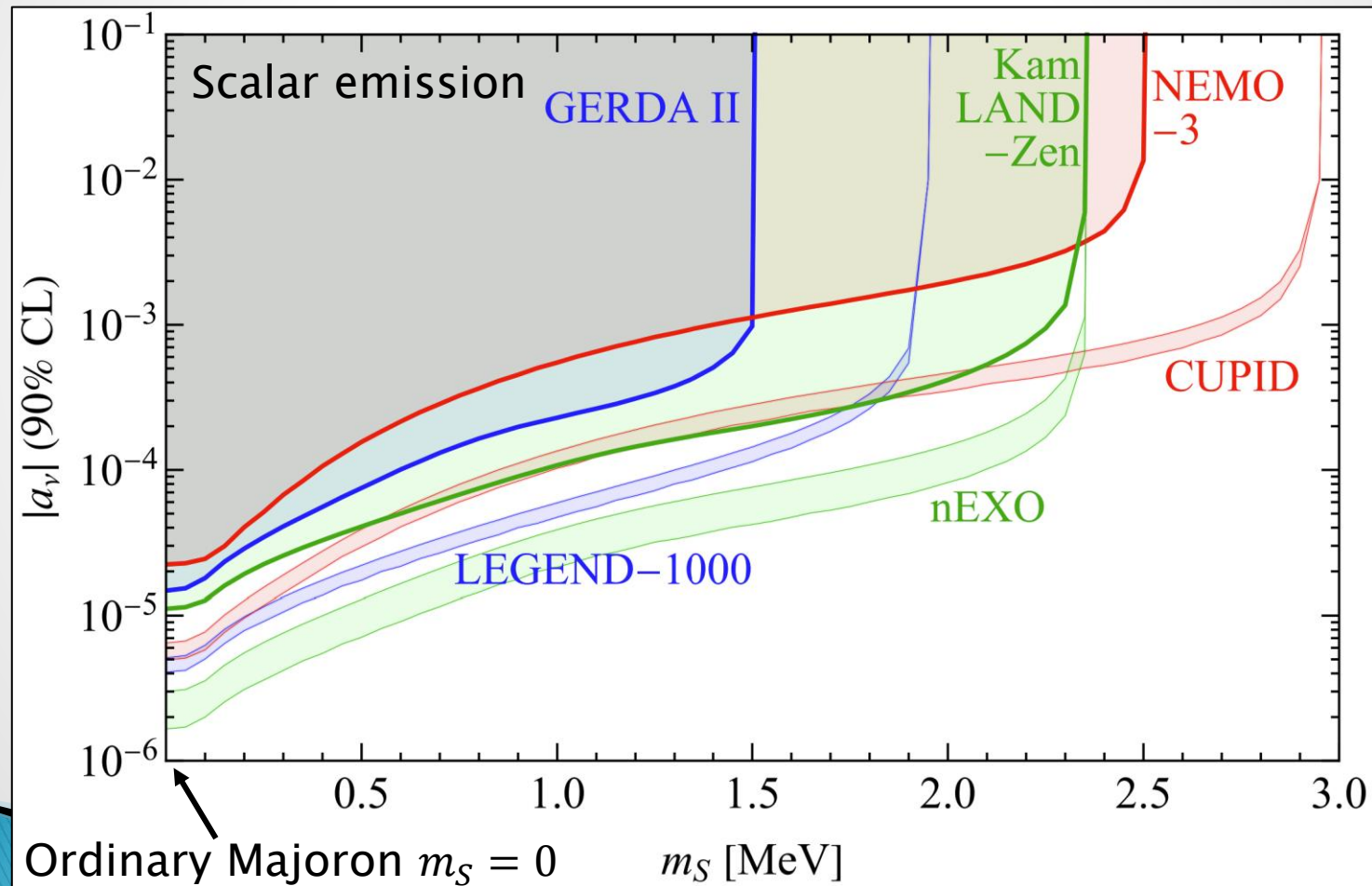
Isotope	Experiment	$\mathcal{E}$ [kg · yr]	$T_{\min}$ [keV]	ΔT [keV]	σ_f [%]
^{76}Ge	GERDA II [67]	104	500	15	1.9
	LEGEND-1000 [68]	7.0×10^3	100	2.5	0.5
^{100}Mo	NEMO-3 [10]	34	500	100	1.8
	CUPID-1T [69]	1.7×10^3	100	5	0.5
^{136}Xe	KamLAND-Zen [5]	126	100	50	0.3
	nEXO [70]	1.9×10^4	100	5	0.5

Isotope	m_{iso} [g/mol]	Q [MeV]	$(T_{1/2}^{2\nu})_{\text{exp}}$ [yr]	$\langle \mathcal{M}_{2\nu} \rangle$	$\delta \mathcal{M}_{2\nu}$	$\langle \mathcal{M}_{0\nu} \rangle$	$\delta \mathcal{M}_{0\nu}$
$^{76}_{32}\text{Ge}$	72.6	2.04	1.9×10^{21}	0.10	0.22	4.3	1.2
$^{100}_{42}\text{Mo}$	96.0	3.03	7.1×10^{18}	0.19	0.27	4.9	1.5
$^{136}_{54}\text{Xe}$	131	2.46	2.2×10^{21}	0.016	0.19	2.6	1.1

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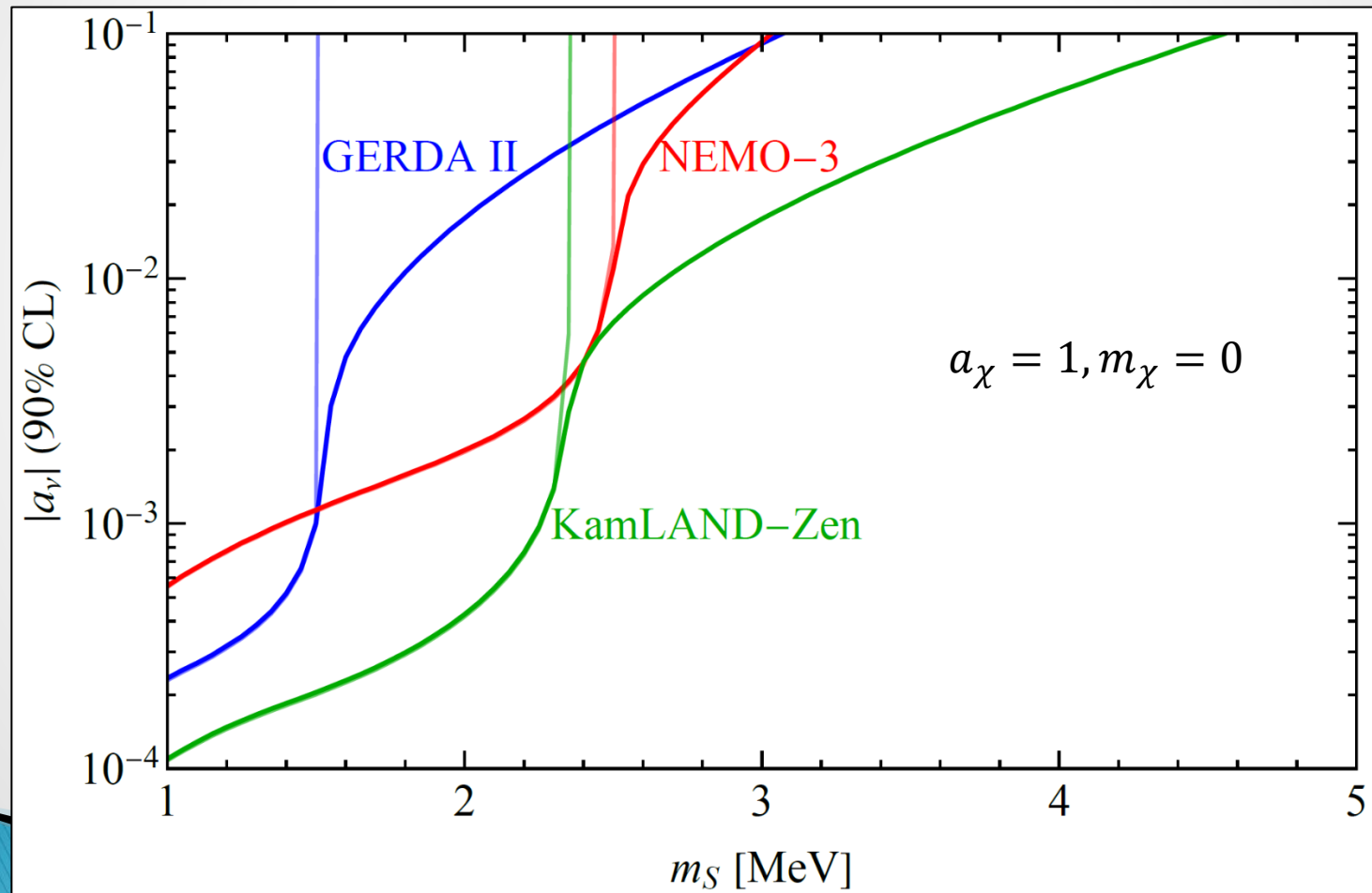
- Sensitivity of current and future double beta decay searches



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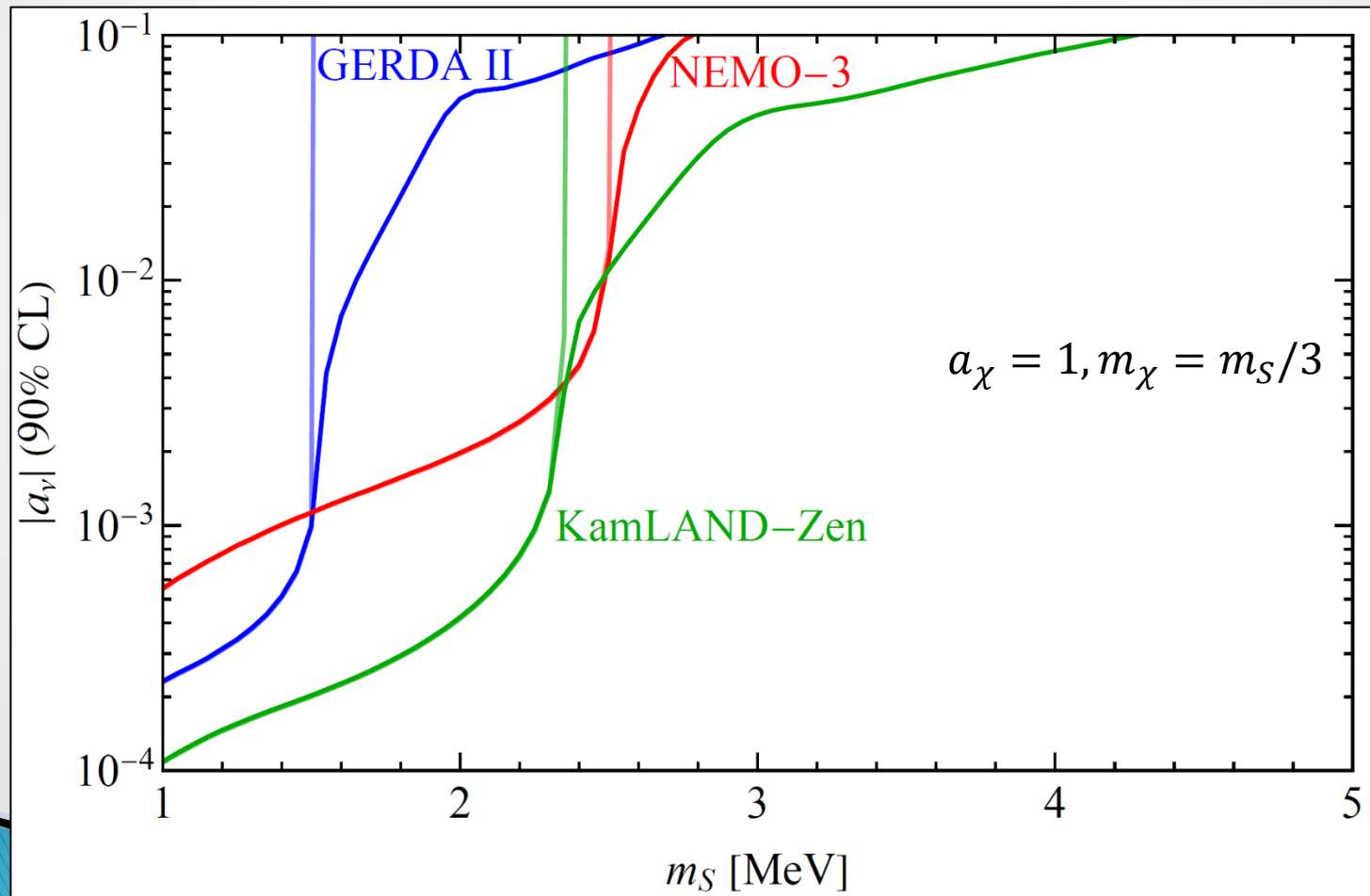
- Modification from interference and off-shell S



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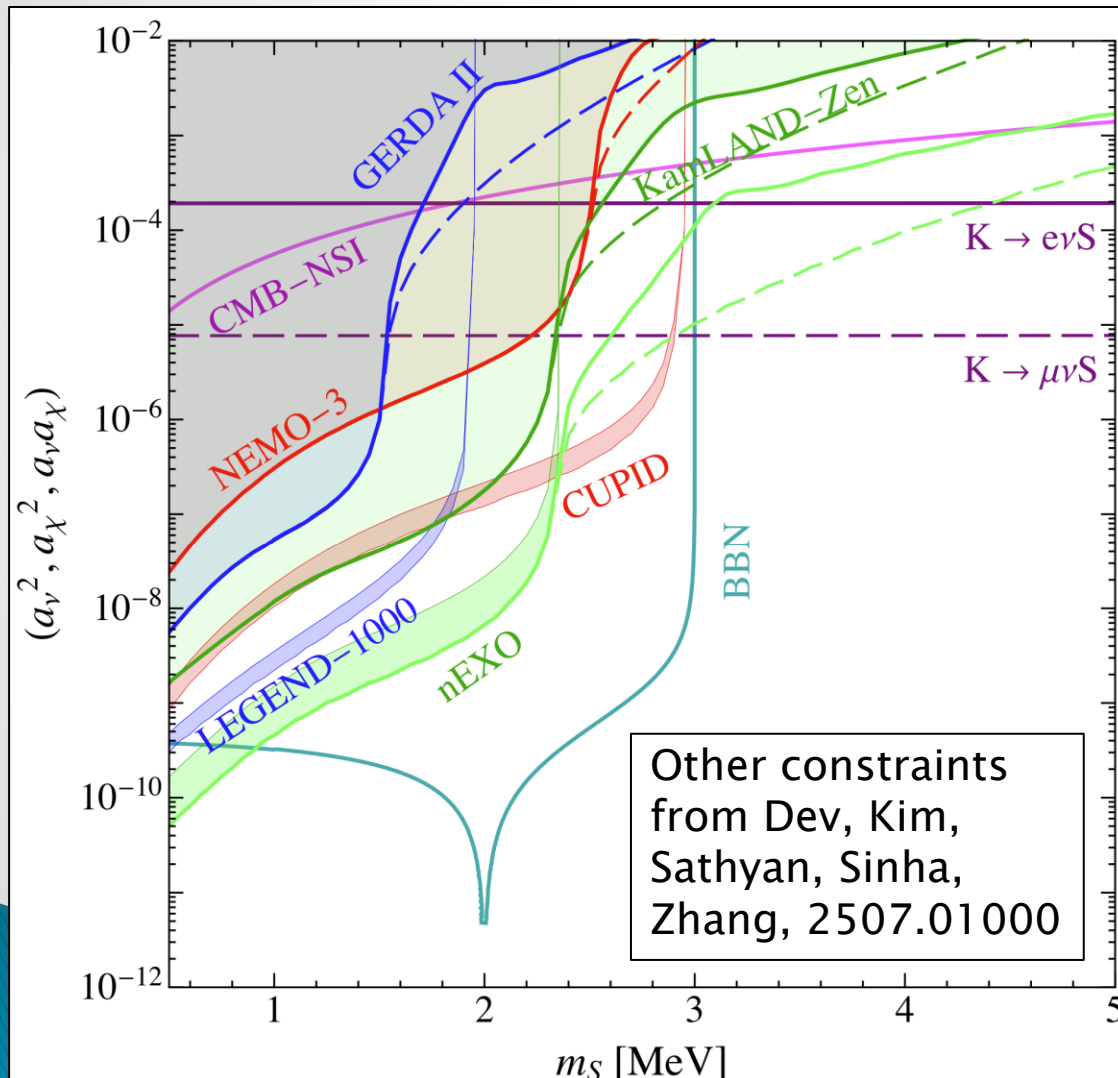
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Neutrinos coupled to DS

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► Comparison with other constraints



► Cosmology

- Thermal relic χ ?
- DM- ν collisional damping
- BBN
- Neutrino reheating from DM annihilations
- Impact of ν NSI on free-streaming (CMB)

► Astrophysics

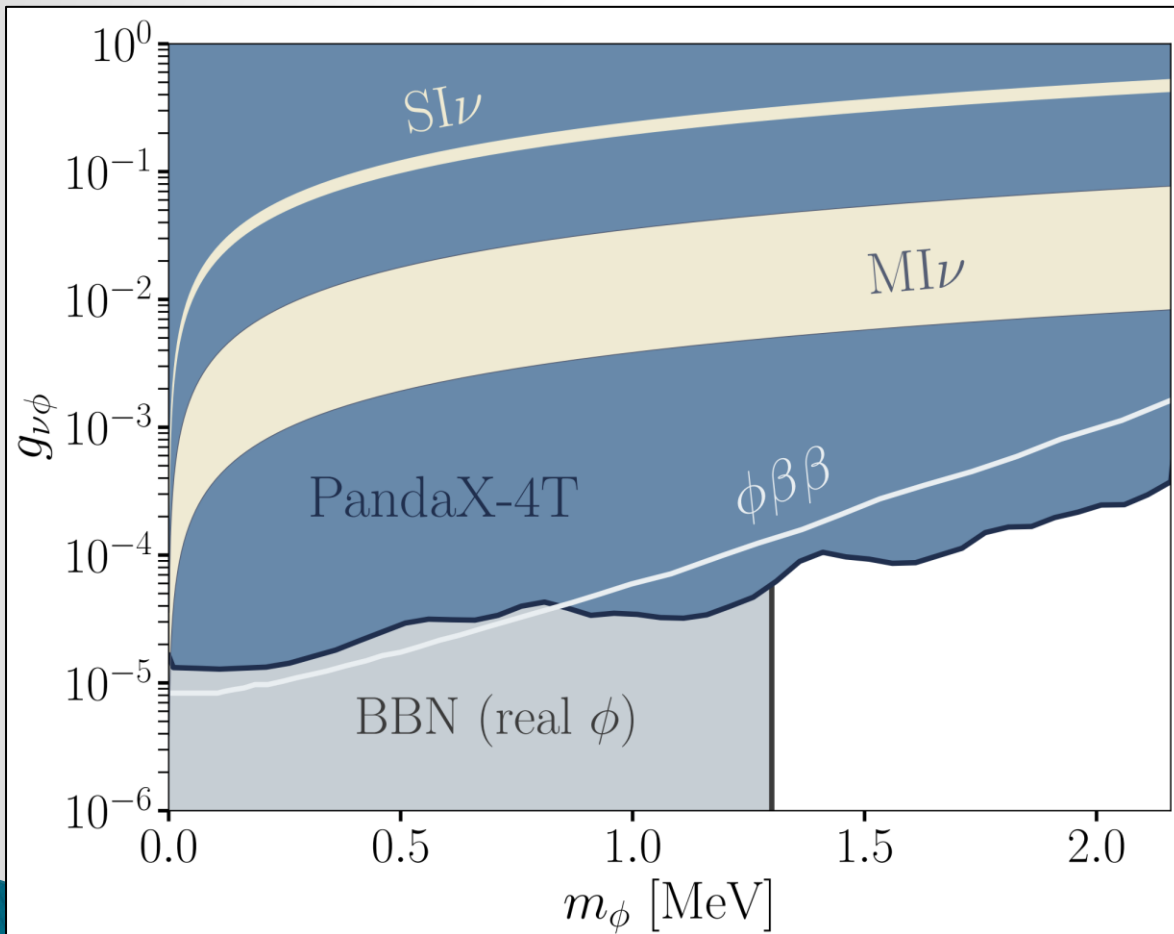
- Supernovae cooling

► Terrestrial

- Kaon decays $K \rightarrow e\nu S$
- Z-width, ... for higher m_S

Neutrinos coupled to DS

- ▶ Recent search at PandaX-4T (PandaX collaboration, arXiv:2511.13515)



- ▶ **Cosmology**

- Thermal relic χ ?
- DM- ν collisional damping
- BBN
- Neutrino reheating from DM annihilations
- Impact of ν NSI on free-streaming (CMB)

- ▶ **Astrophysics**

- Supernovae cooling

- ▶ **Terrestrial**

- Kaon decays $K \rightarrow e\nu S$
- Z-width, ... for higher m_S

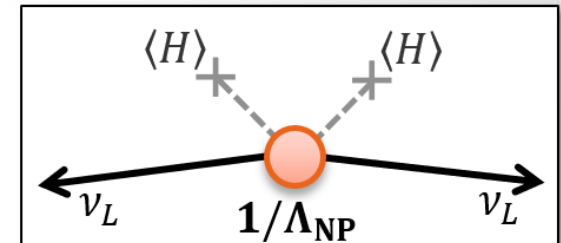
Conclusion

▶ Neutrinos much lighter than other fermions

- Dirac or Majorana?
- Mechanism of neutrino mass generation?

▶ $0\nu\beta\beta$ is crucial probe for BSM physics

- Prove neutrino Majorana nature
- LNV physics near GUT scale and at scales $m_N \approx 1 \text{ eV} - 100 \text{ TeV}$
- Falsify high-scale baryogenesis scenarios



$$\frac{T_{1/2}^{0\nu\beta\beta}}{10^{28} \text{ y}} \approx \left(\frac{\Lambda_{NP}}{10^{15} \text{ GeV}} \right)^2$$

▶ $2\nu\beta\beta$ spectrum sensitive to New Physics

- Non-standard ν interactions, sterile ν
- Neutrino self-interactions and DS interactions
 - Strongest terrestrial limits
 $|g_{\nu e S}| < 10^{-5} - 10^{-1}$ for $m_S < 3 - 5 \text{ MeV}$
 - Synergy and overlap with cosmological considerations

