



中國科學院高能物理研究所
Institute of High Energy Physics
Chinese Academy of Sciences

Review of solar neutrino in 2028

Xue-Feng DING

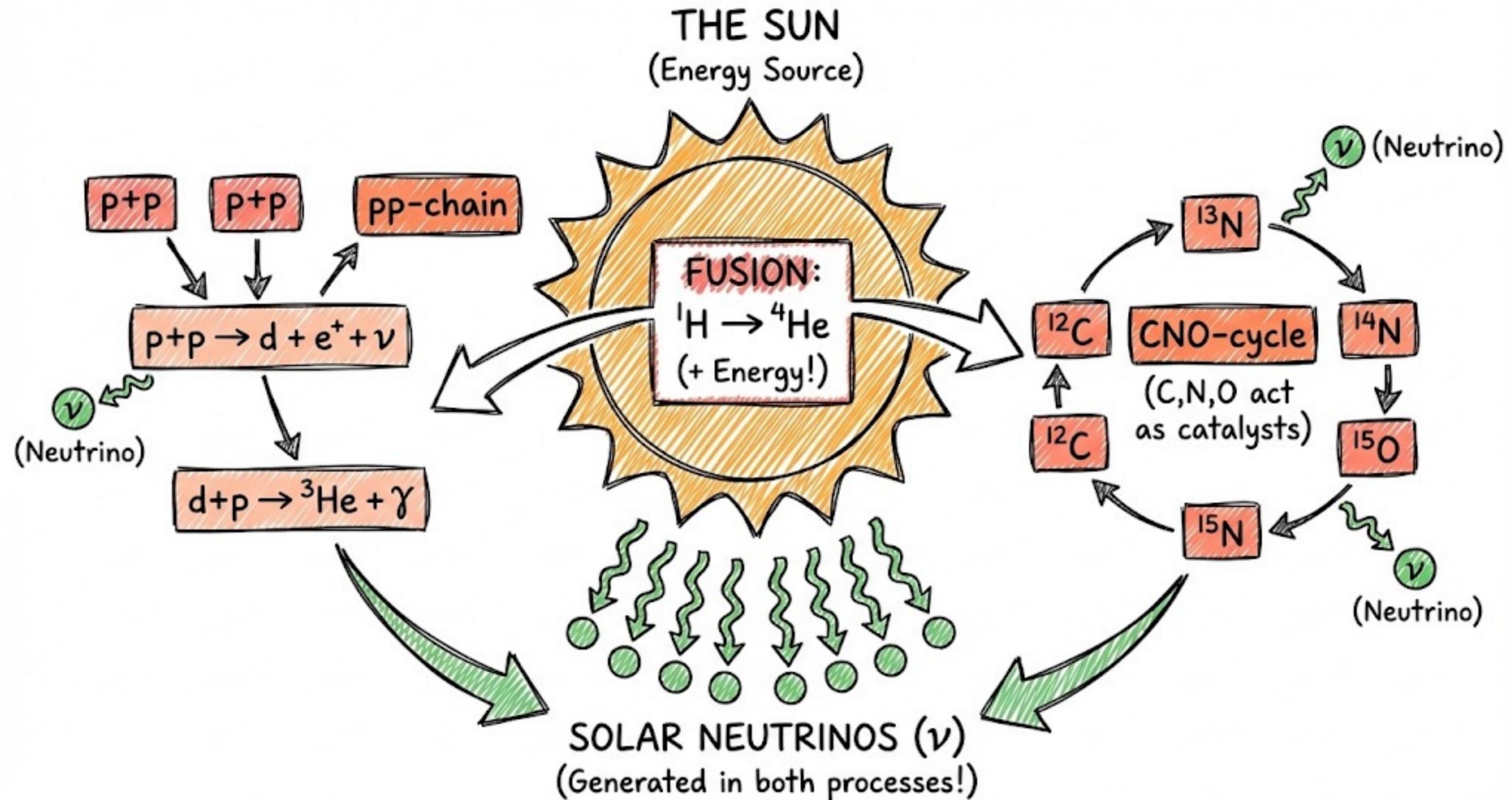
IHEP, CAS

2026 January 28th

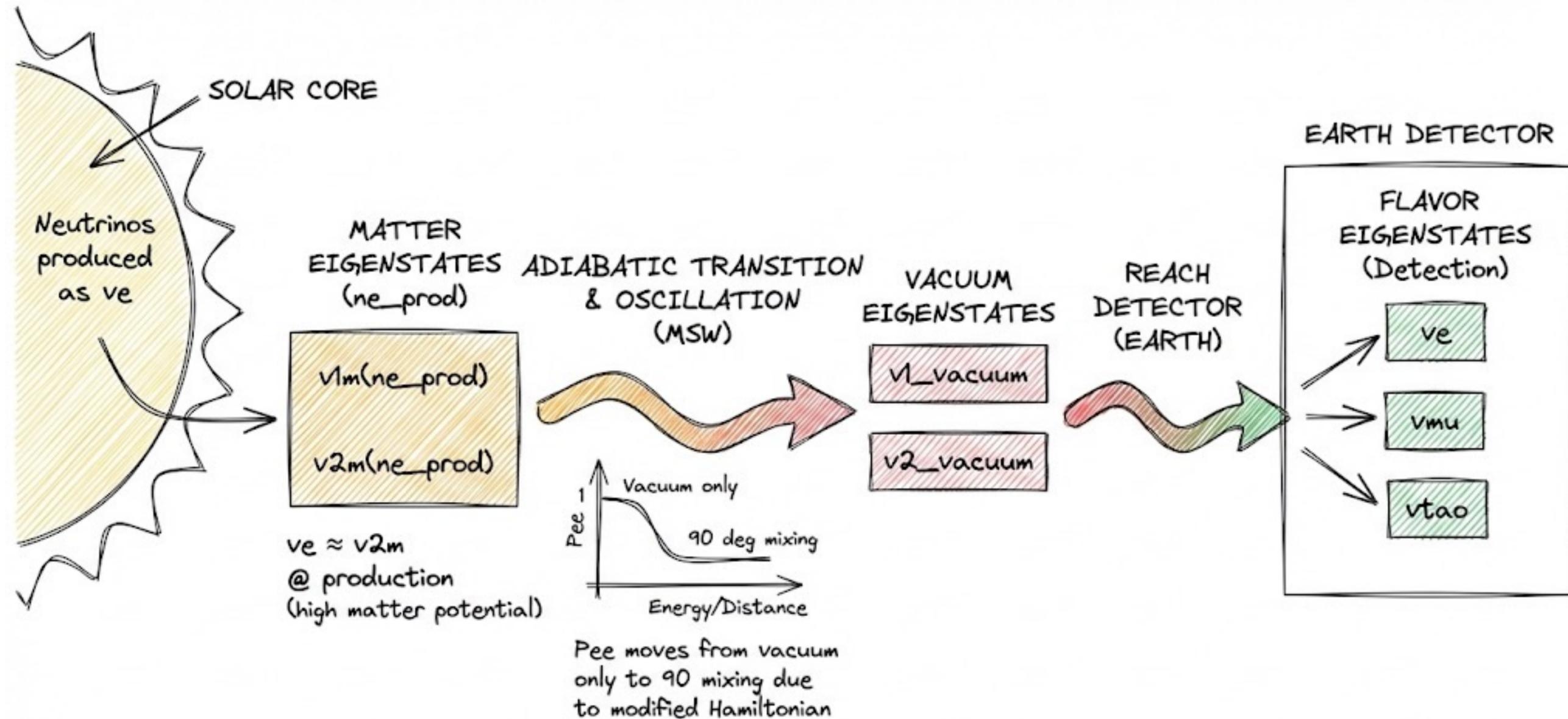
Outline

- Introduction
- Standard solar model
- Neutrino oscillation
- Conclusions

Energy production in the SUN

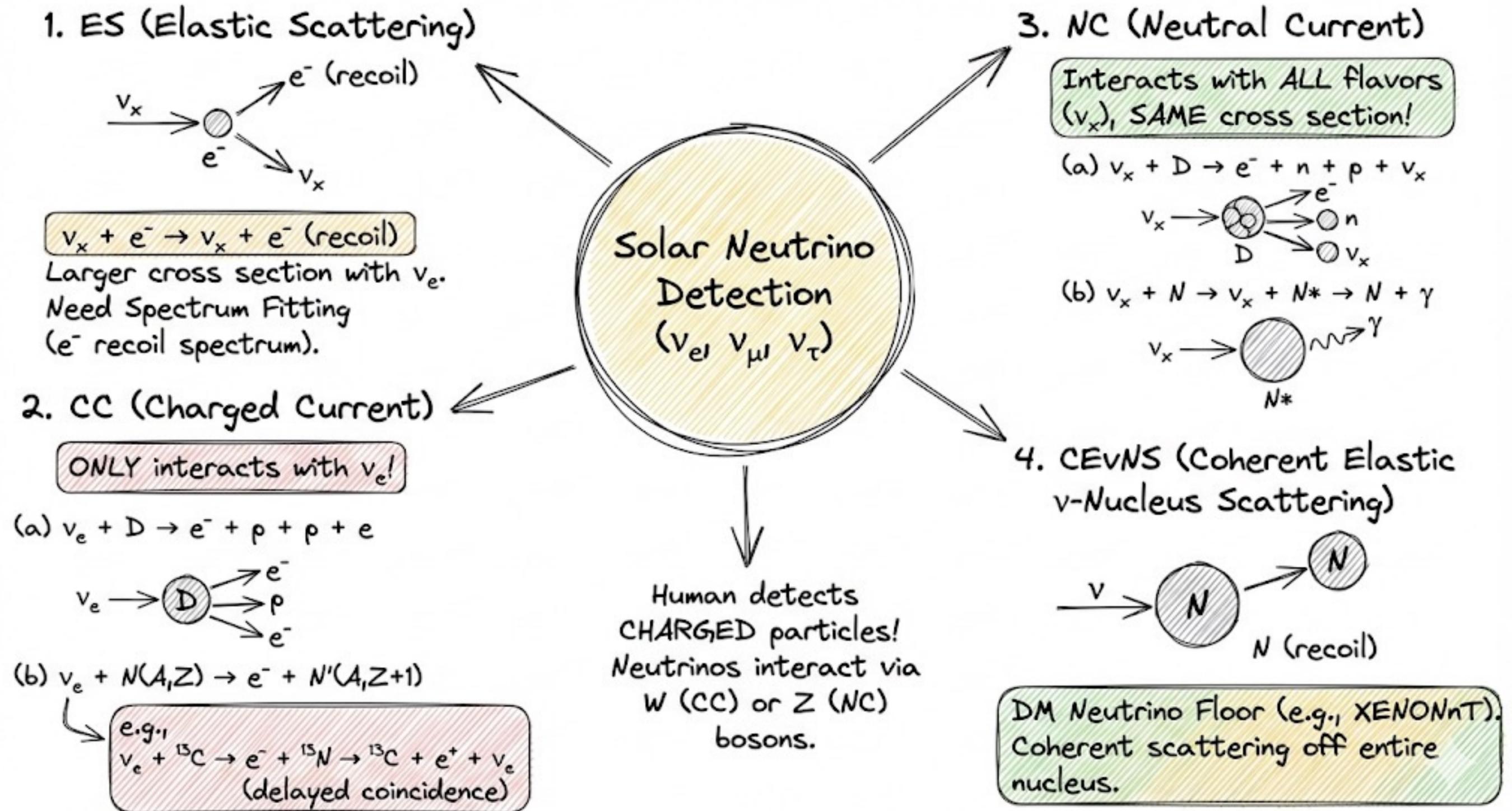


Adiabatic transition: no oscillation

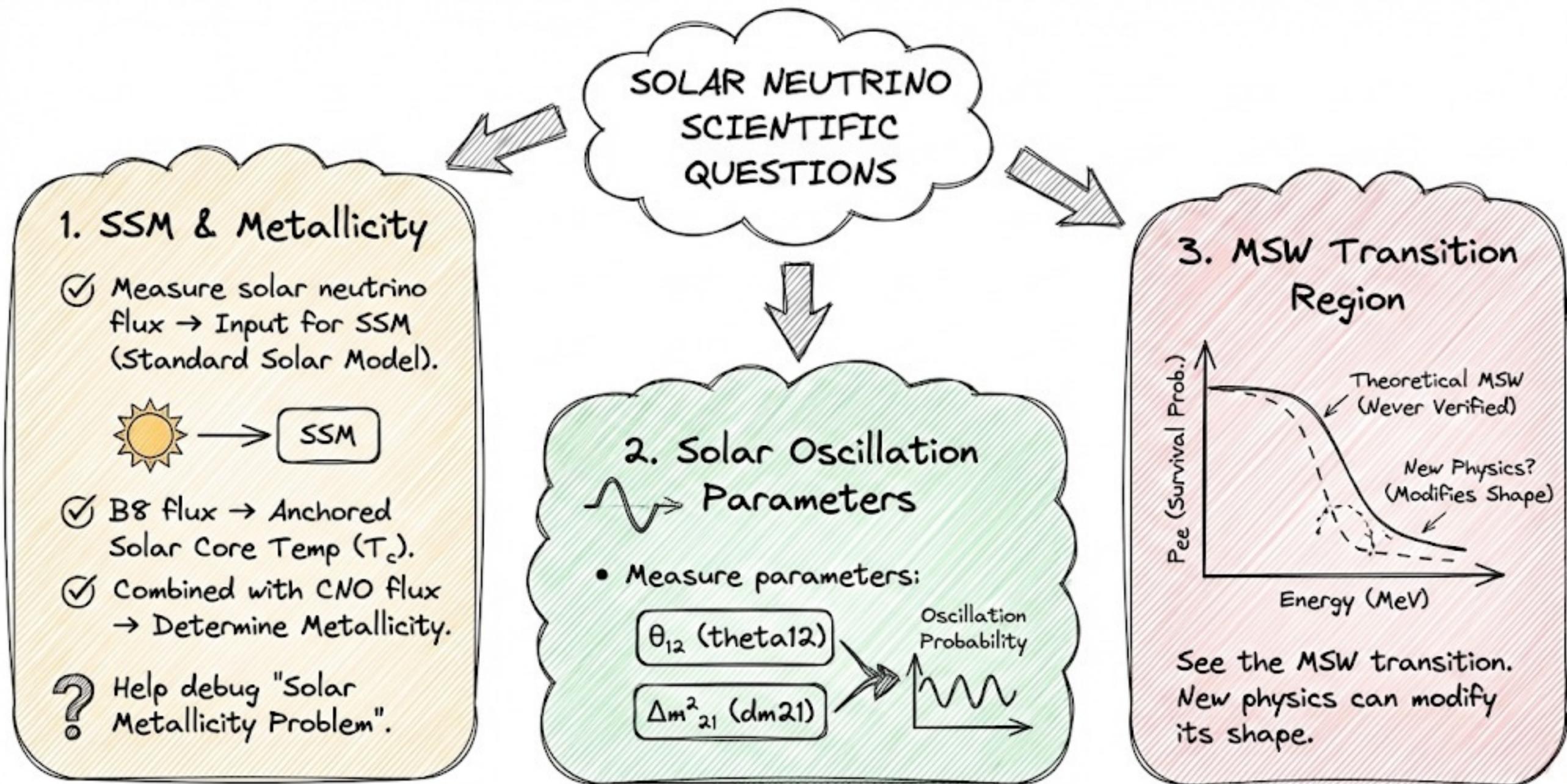


Matter potential modifies Hamiltonian and thus ν_{1m} and ν_{2m} .

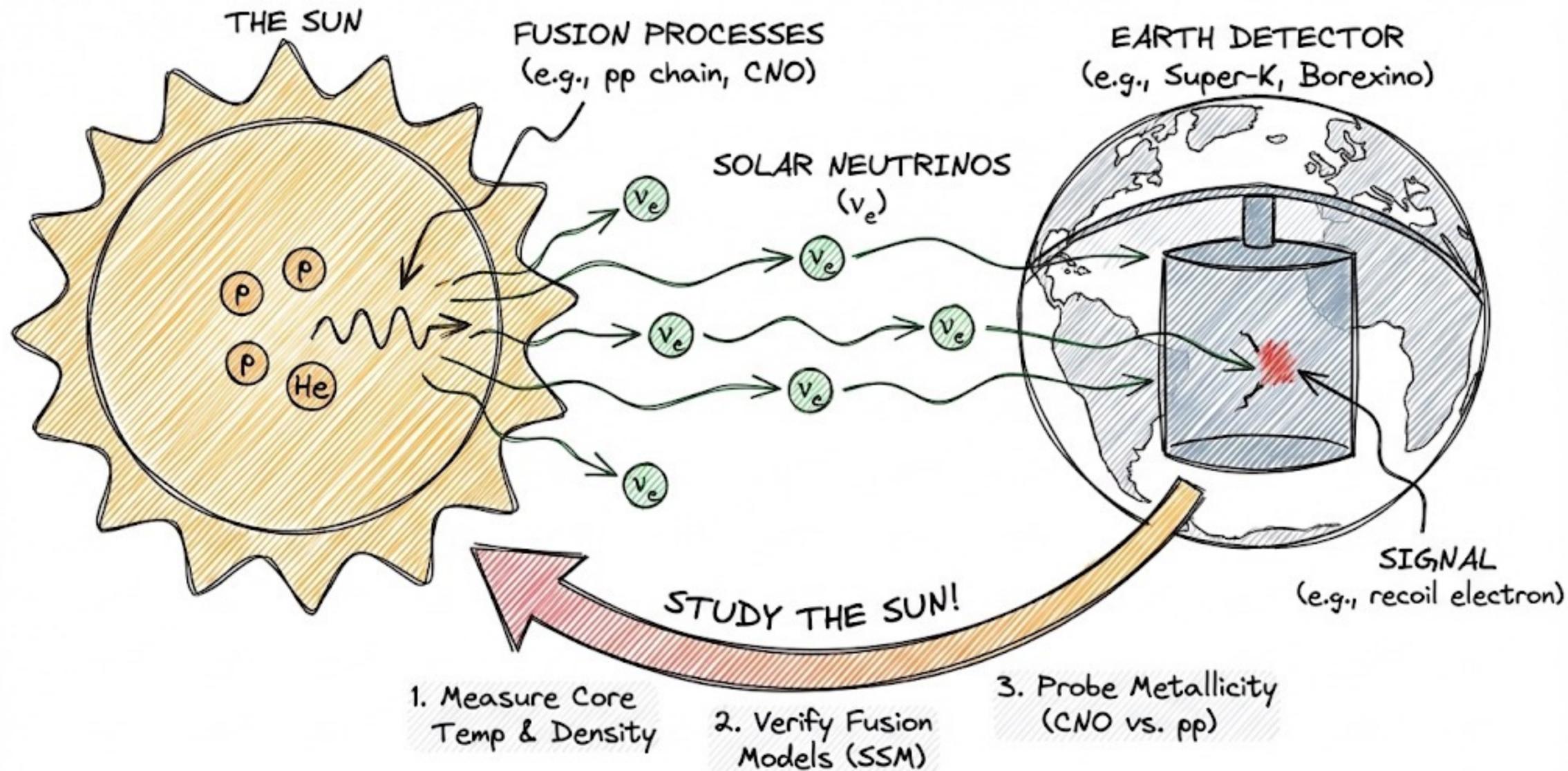
Detection of solar neutrino



Scientific question of solar ν

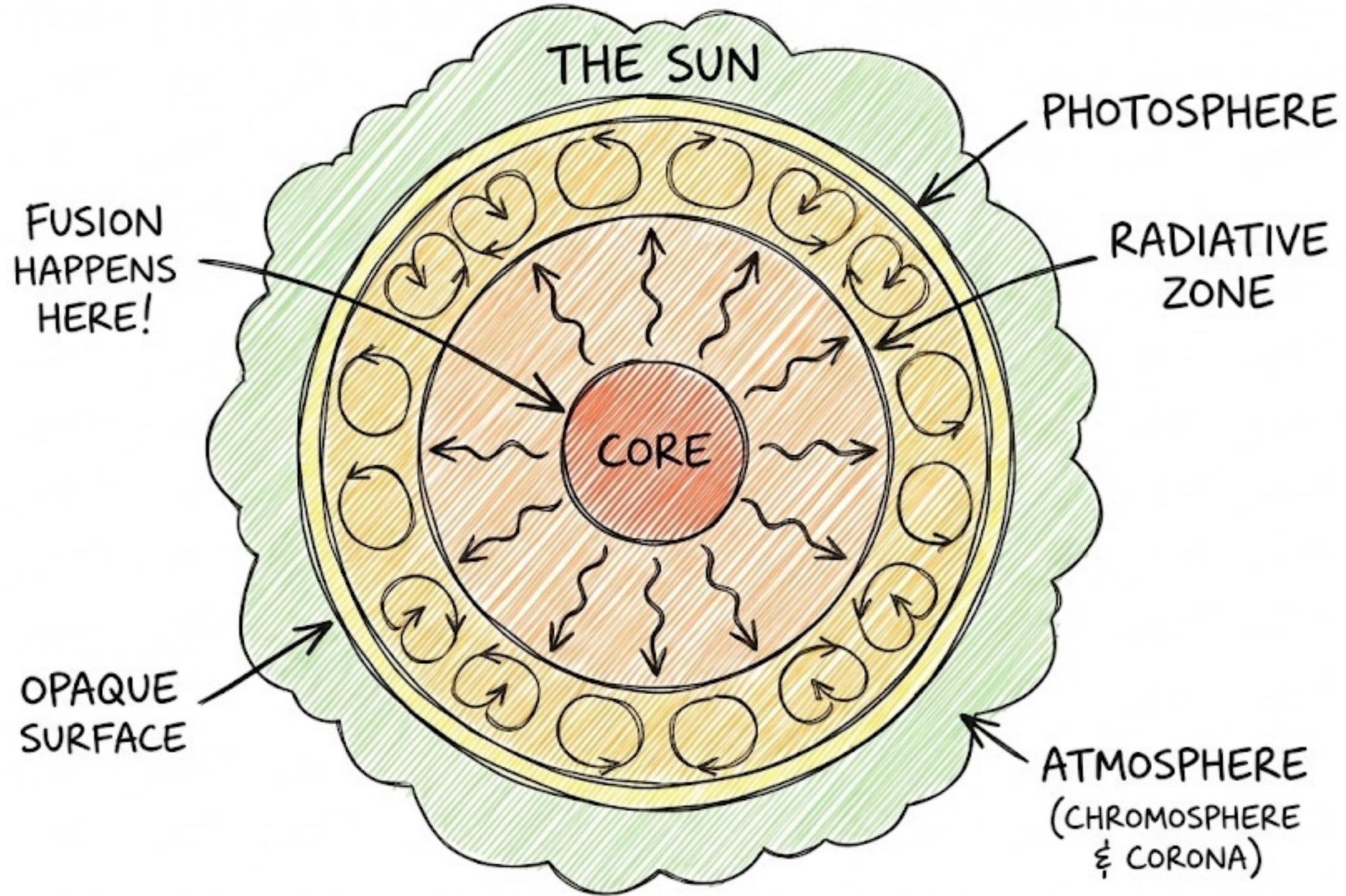


Part 1: study the Sun

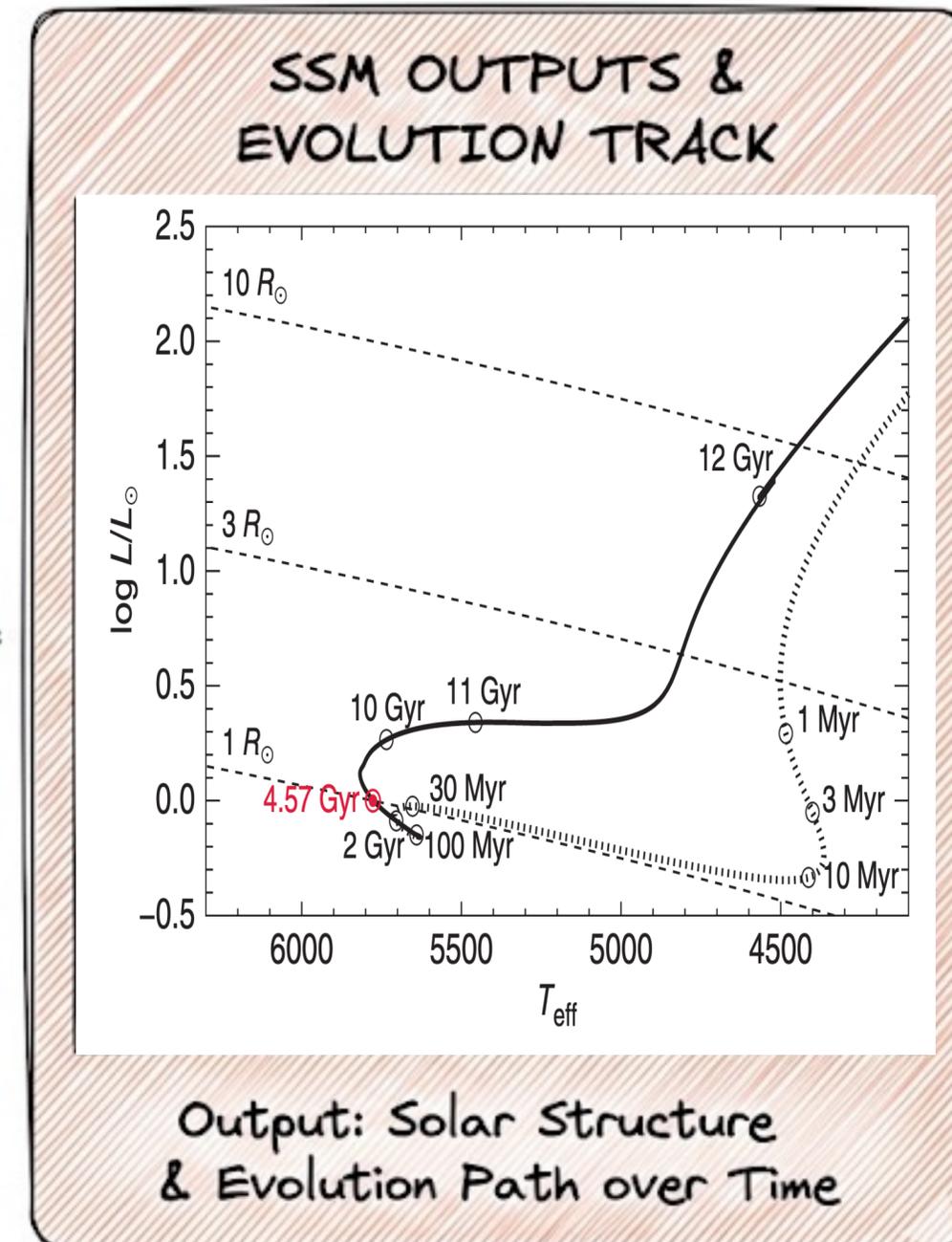
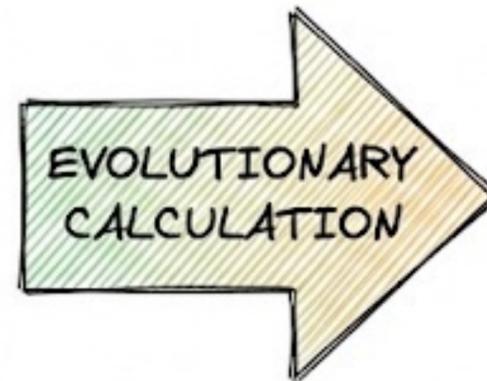
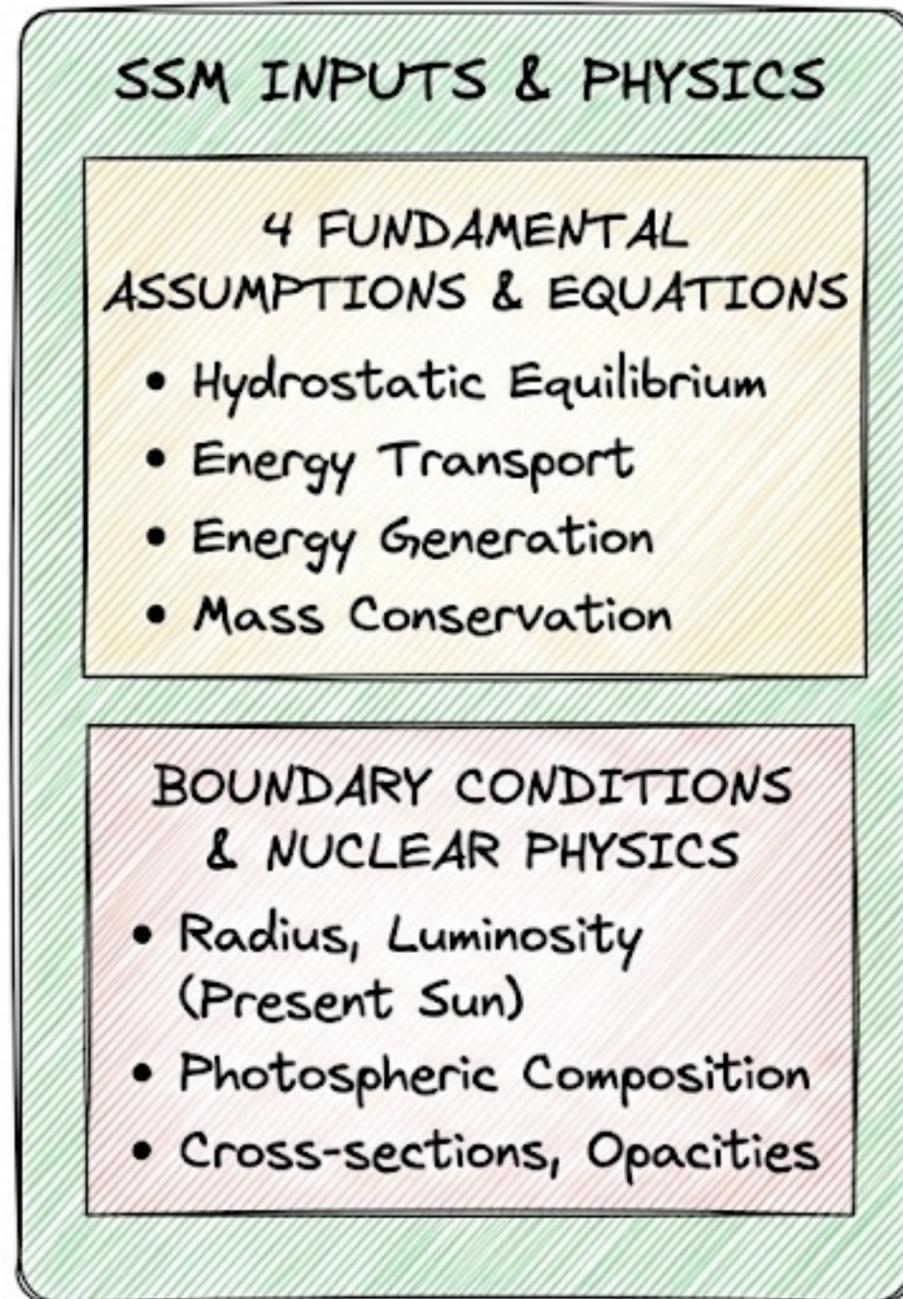


Neutrinos: Direct messengers from the solar core!

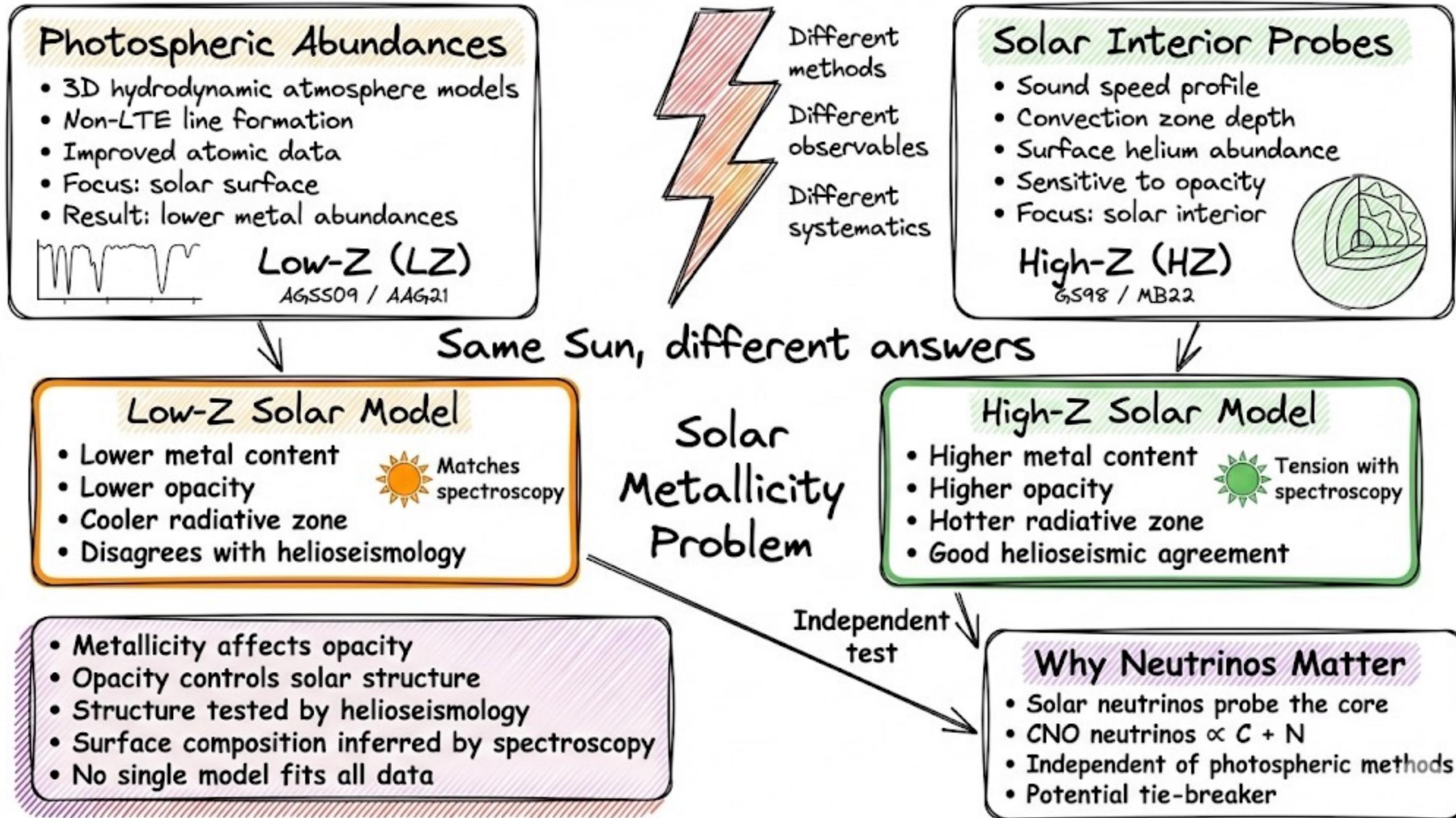
The Sun



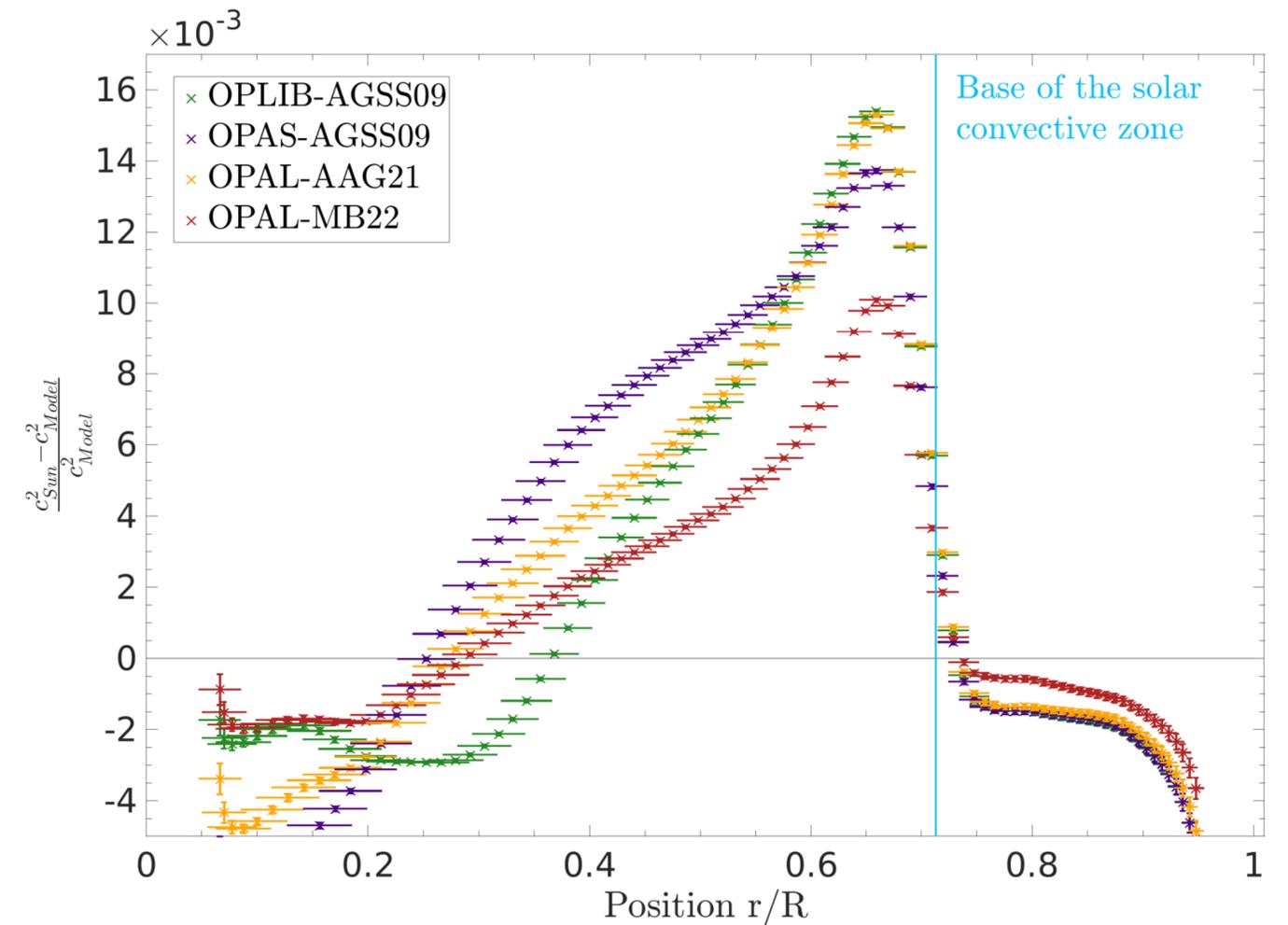
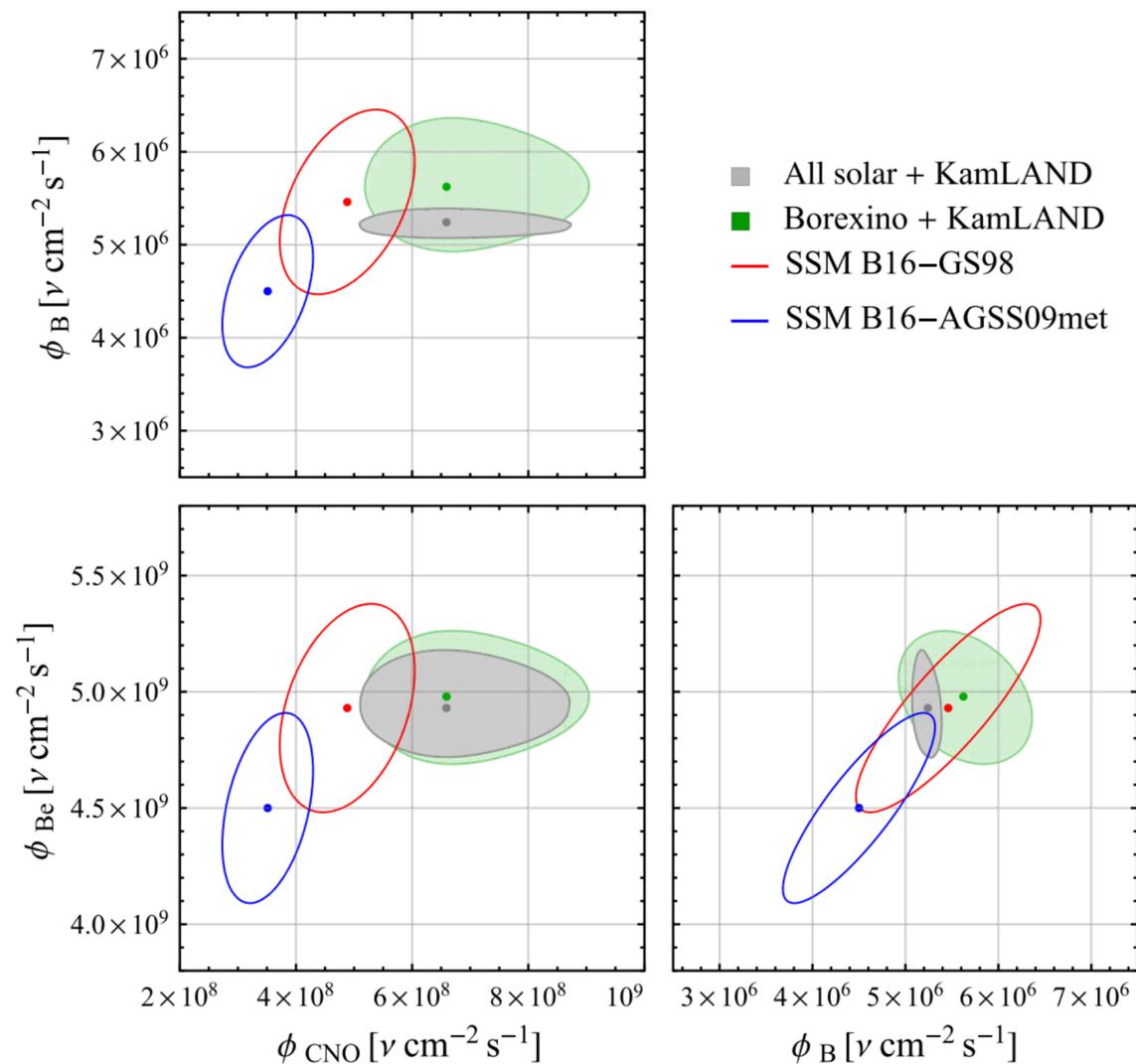
Standard Solar Model



Photosphere vs Helioseismology



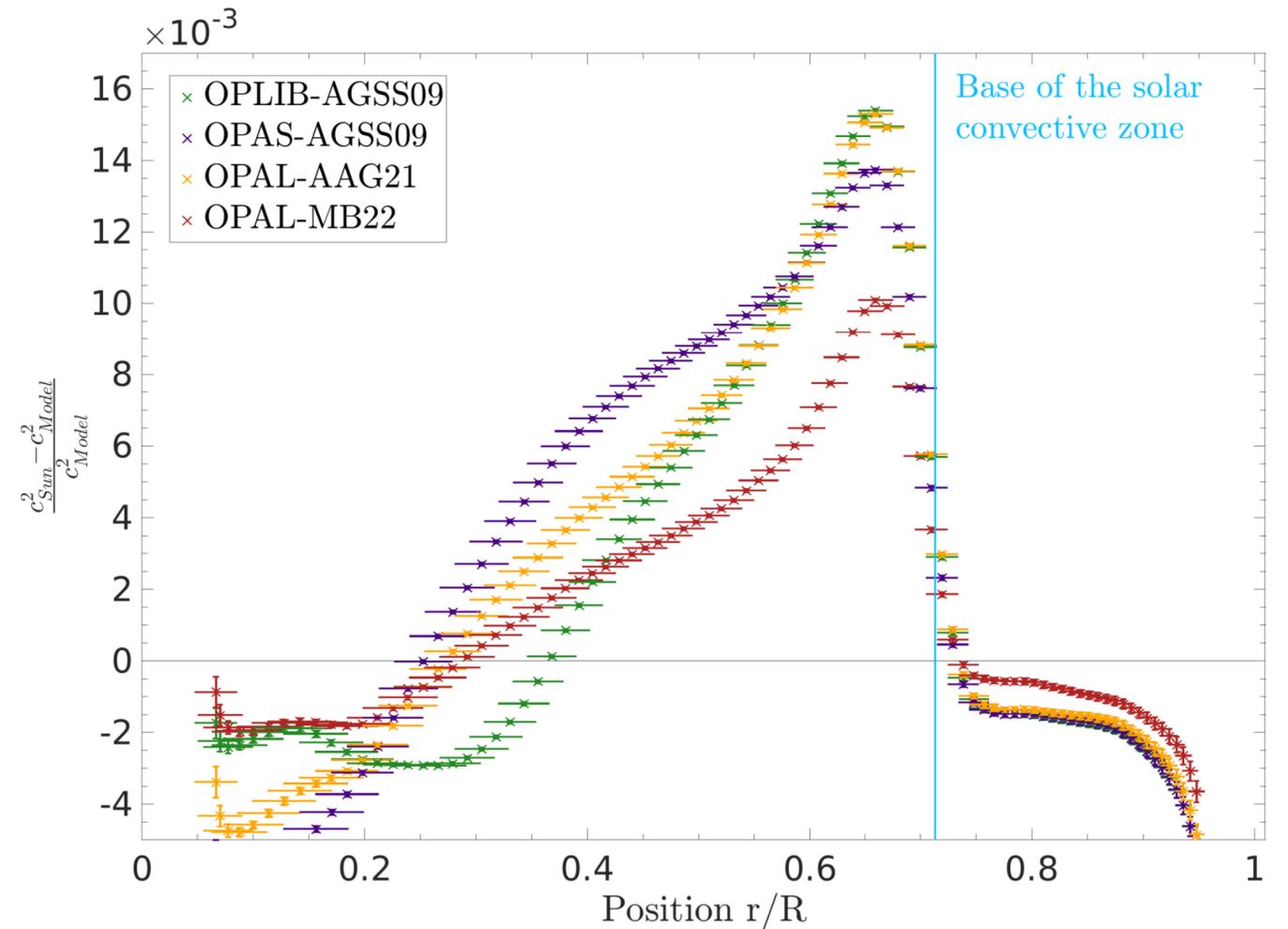
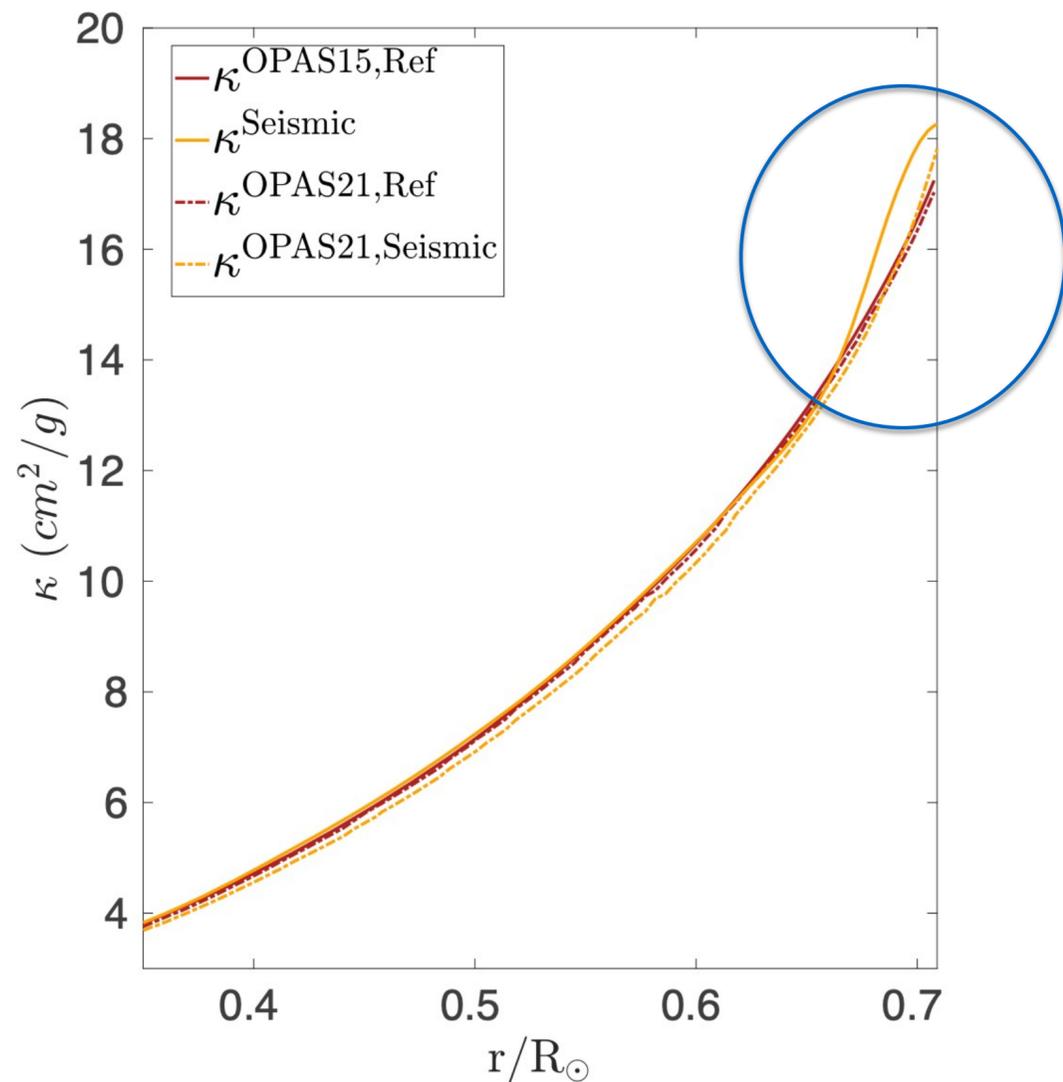
Solar metallicity problem



Appel, S., Z. Bagdasarian, D. Basilico, et al. "Improved Measurement of Solar Neutrinos from the Carbon-Nitrogen-Oxygen Cycle by Borexino and Its Implications for the Standard Solar Model." *Physical Review Letters* 129, no. 25 (2022): 252701. <https://doi.org/10.1103/PhysRevLett.129.252701>.

Buldgen, Gael, Gloria Canocchi, Arthur Le Saux, et al. "The Future of Solar Modelling: Requirements for a New Generation of Solar Models." arXiv:2506.14514. Preprint, arXiv, June 17, 2025. <https://doi.org/10.48550/arXiv.2506.14514>.

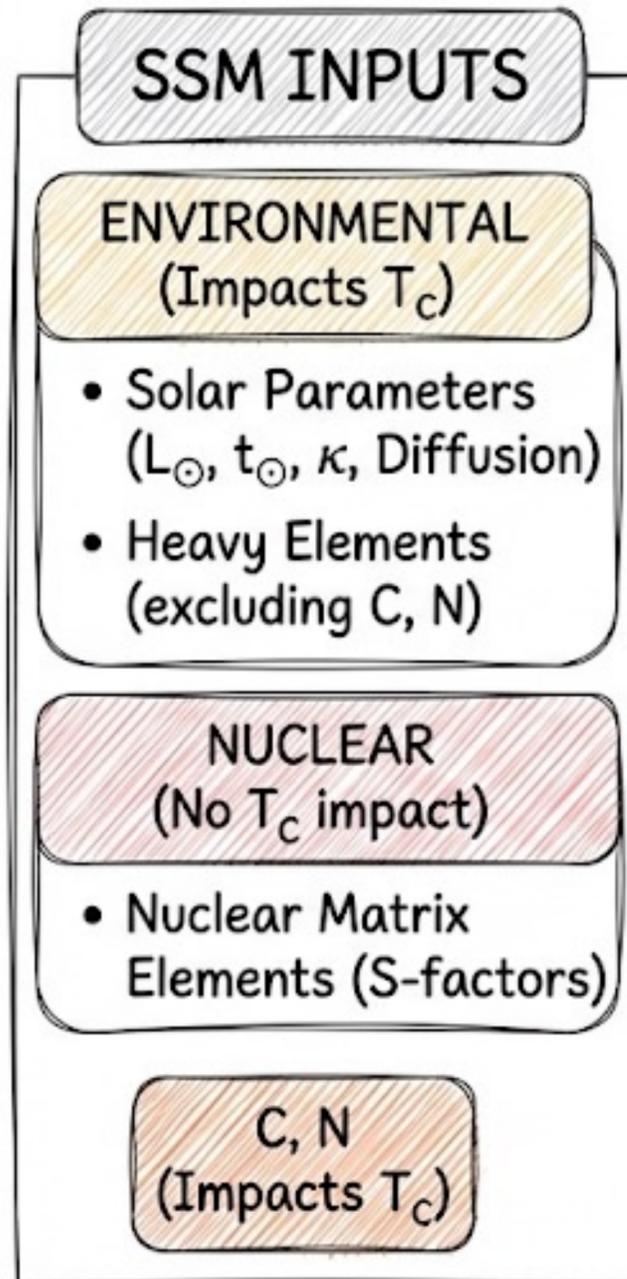
What next? Opacity



Buldgen, Gaël, Jean-Christophe Pain, Philippe Cossé, et al. "Helioseismic Inference of the Solar Radiative Opacity." *Nature Communications* 16, no. 1 (2025): 693. <https://doi.org/10.1038/s41467-024-54793-y>.

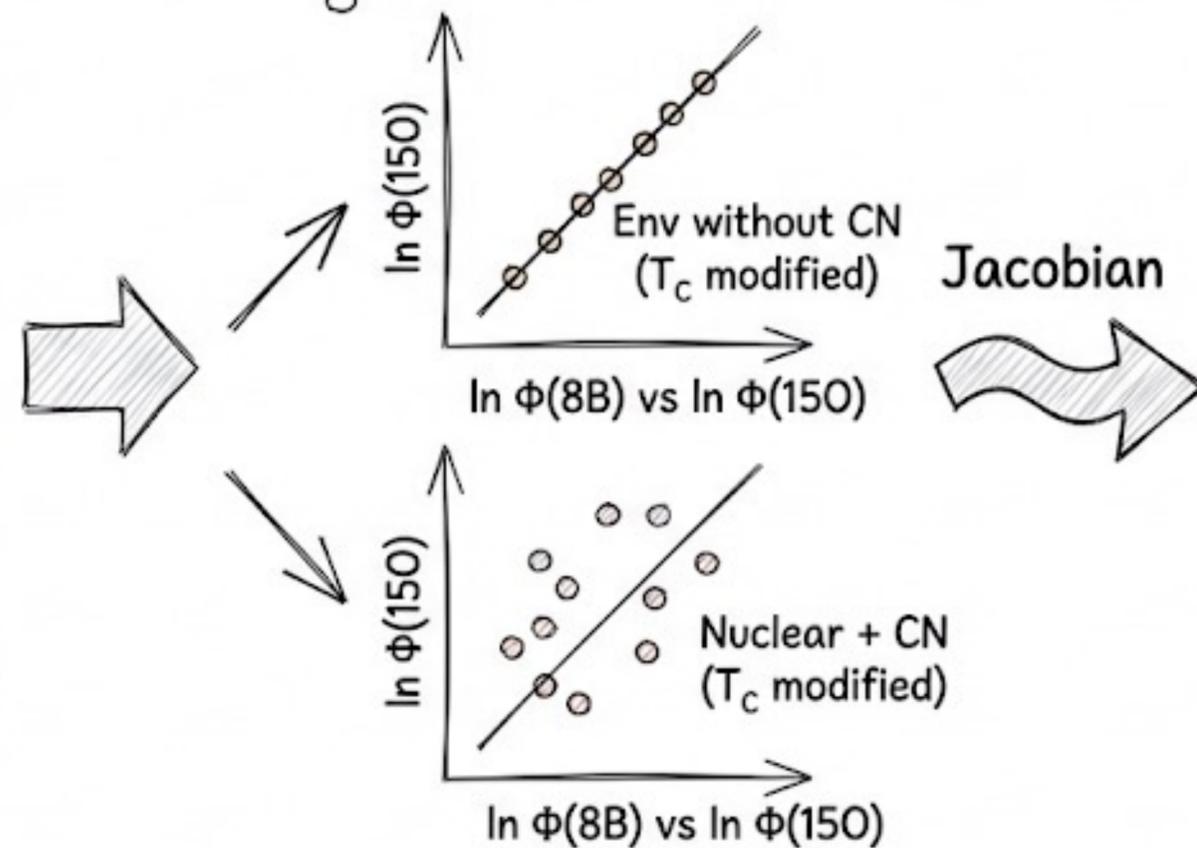
Buldgen, Gaël, Gloria Canocchi, Arthur Le Saux, et al. "The Future of Solar Modelling: Requirements for a New Generation of Solar Models." arXiv:2506.14514. Preprint, arXiv, June 17, 2025. <https://doi.org/10.48550/arXiv.2506.14514>.

Solar neutrino as a probe



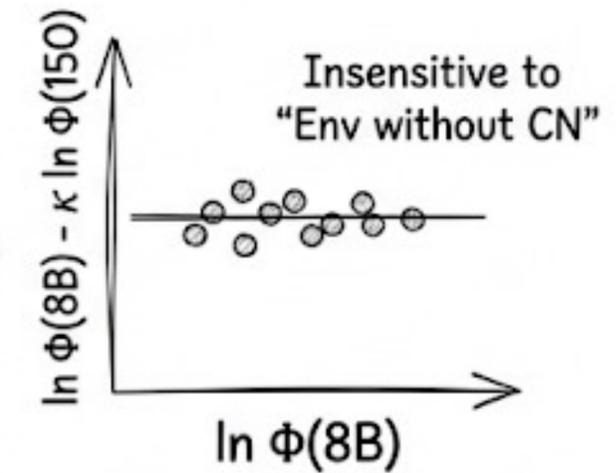
B8 + CNO Neutrino Flux
 → Determine C+N Abundance

⇒ Solar Metallicity Problem



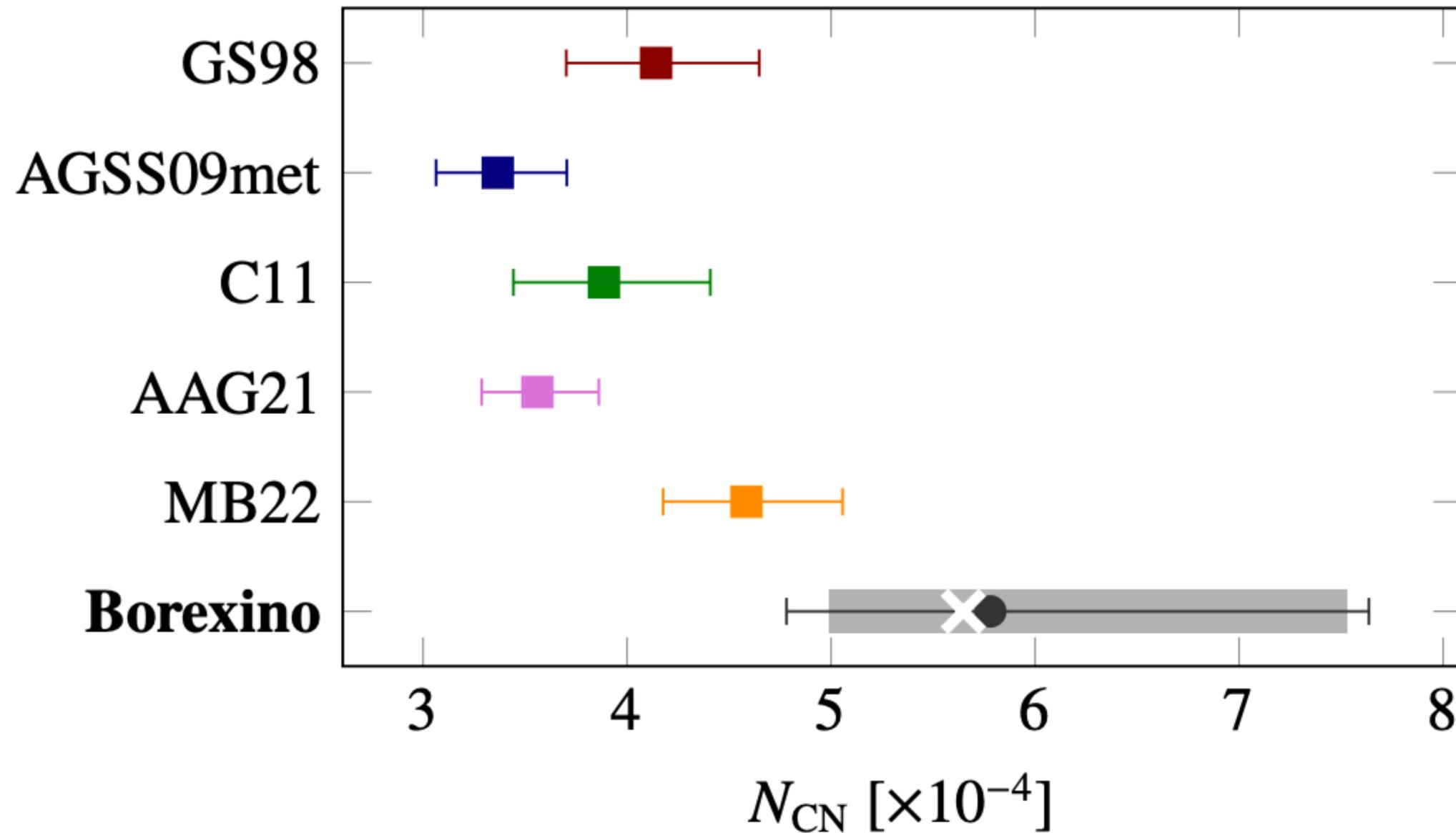
$$\ln \Phi(8B) - \kappa * \ln \Phi(150)$$

In sensitive to Env



(Almost) only depends on NUCLEAR + CN

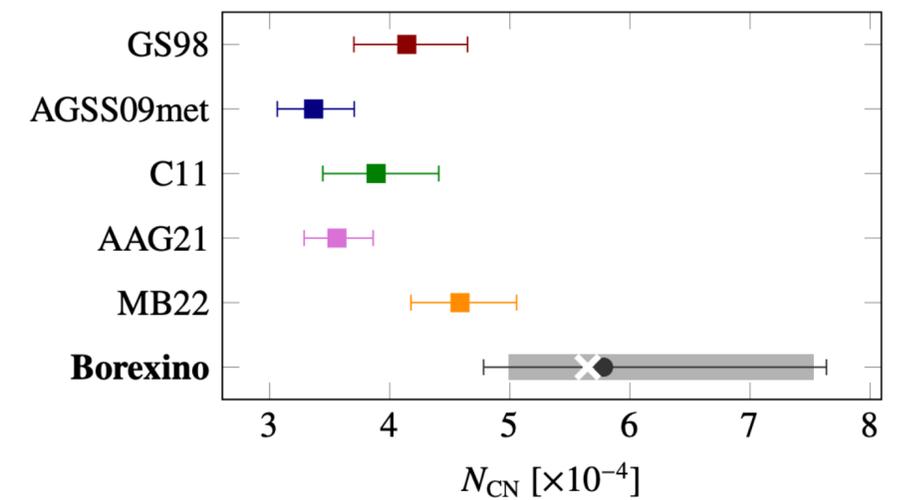
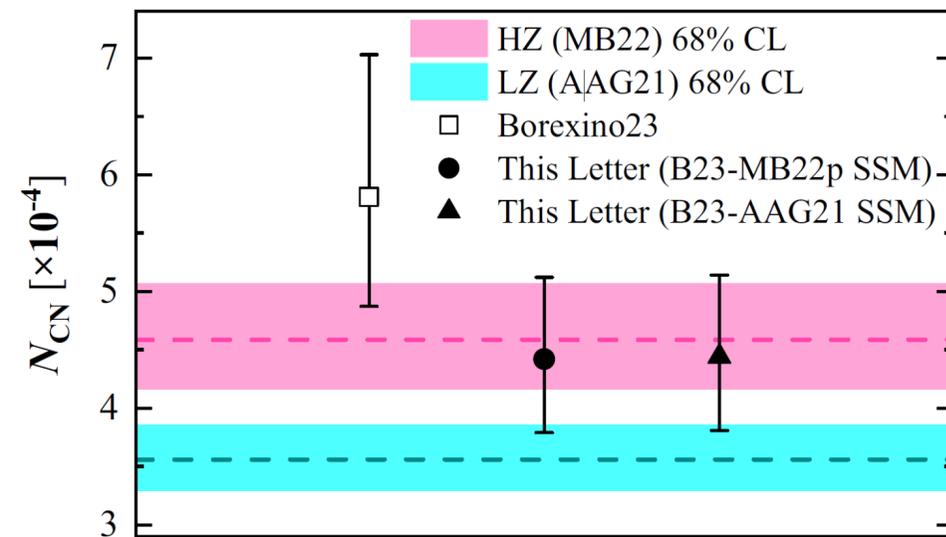
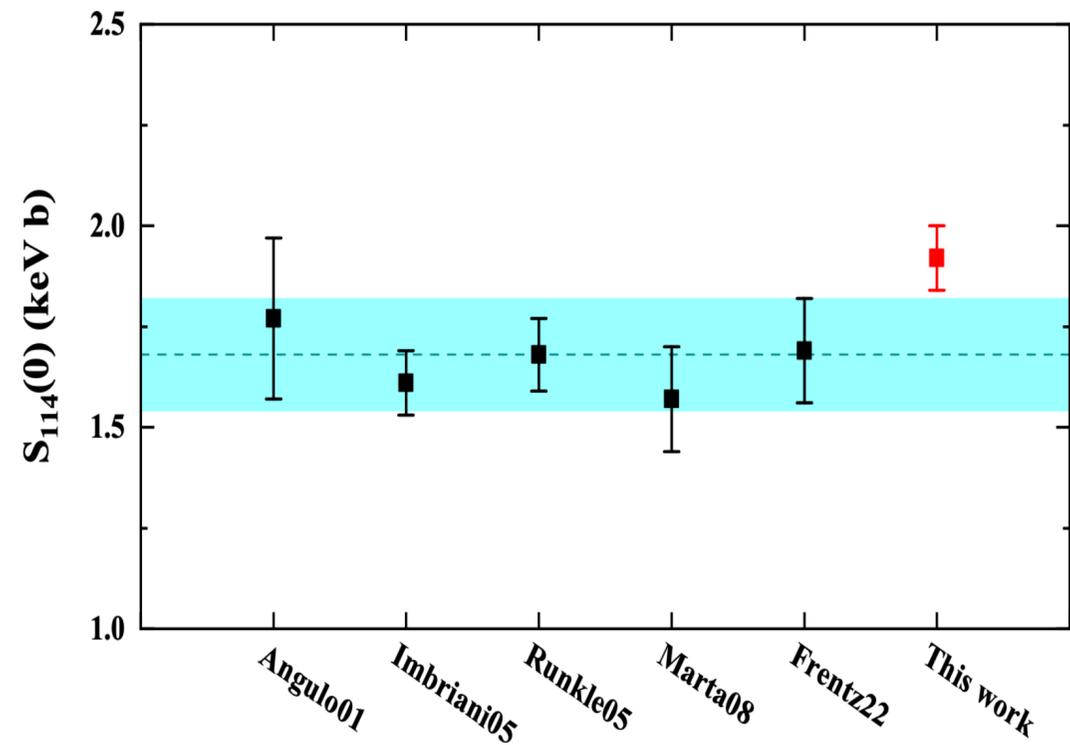
Measured CN abun. by Borexino



Appel, S., Z. Bagdasarian, D. Basilico, et al. "Improved Measurement of Solar Neutrinos from the Carbon-Nitrogen-Oxygen Cycle by Borexino and Its Implications for the Standard Solar Model." *Physical Review Letters* 129, no. 25 (2022): 252701.

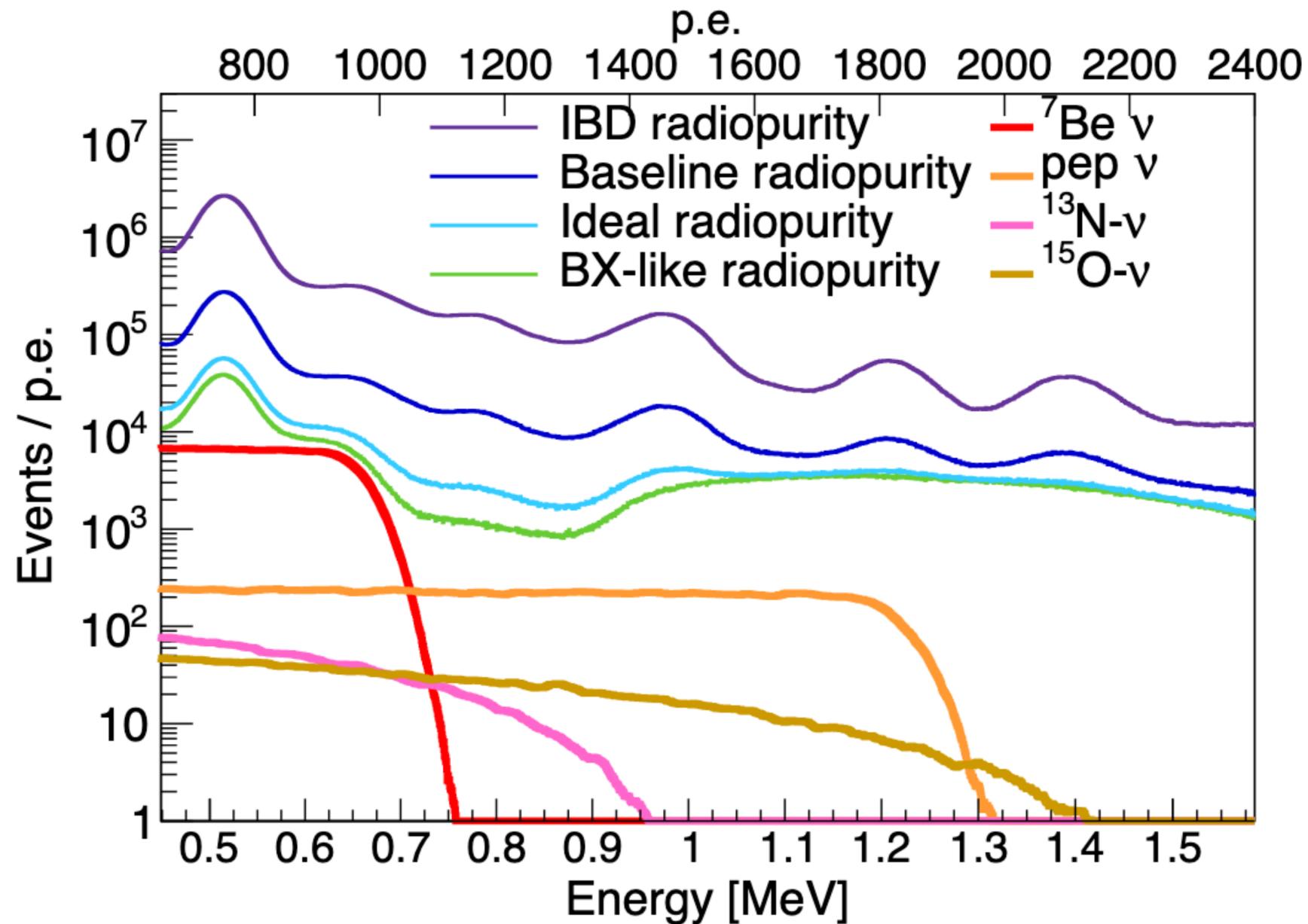
<https://doi.org/10.1103/PhysRevLett.129.252701>.

S factor improved

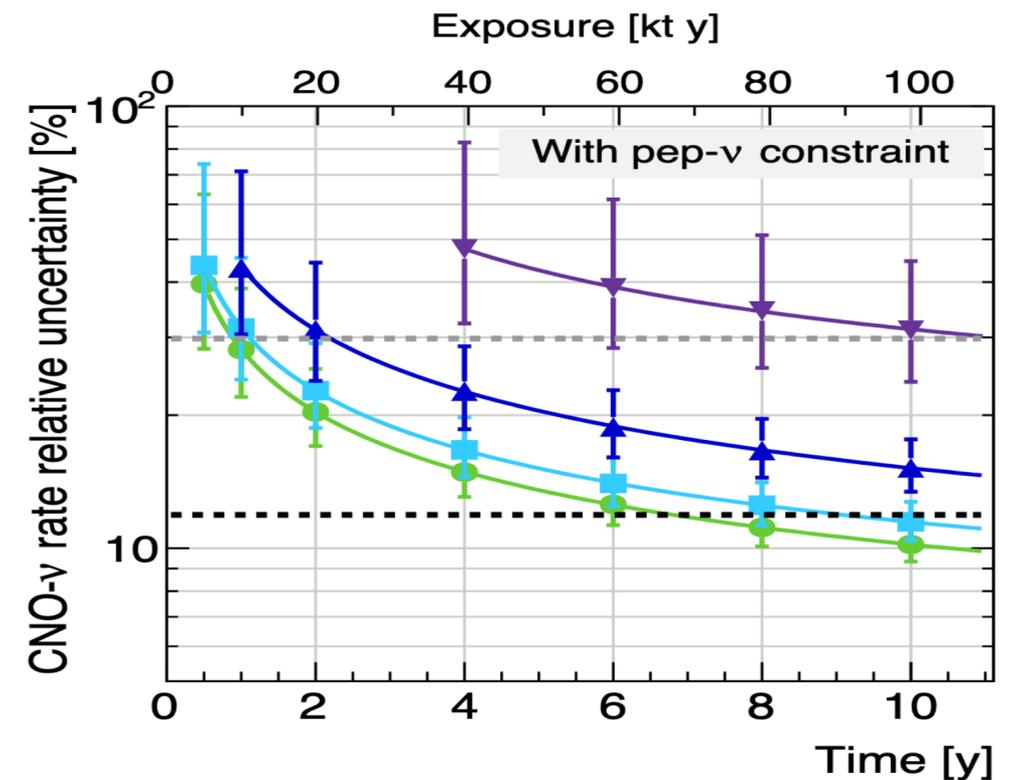


Chen, X., J. Su, Y. P. Shen, et al. "Enhanced S -Factor for the $^{14}\text{N}(p,\gamma)^{15}\text{O}$ Reaction and Its Impact on the Solar Composition Problem." arXiv:2410.16086. Preprint, arXiv, October 21, 2024. <https://doi.org/10.48550/arXiv.2410.16086>.

JUNO's prospects to CNO



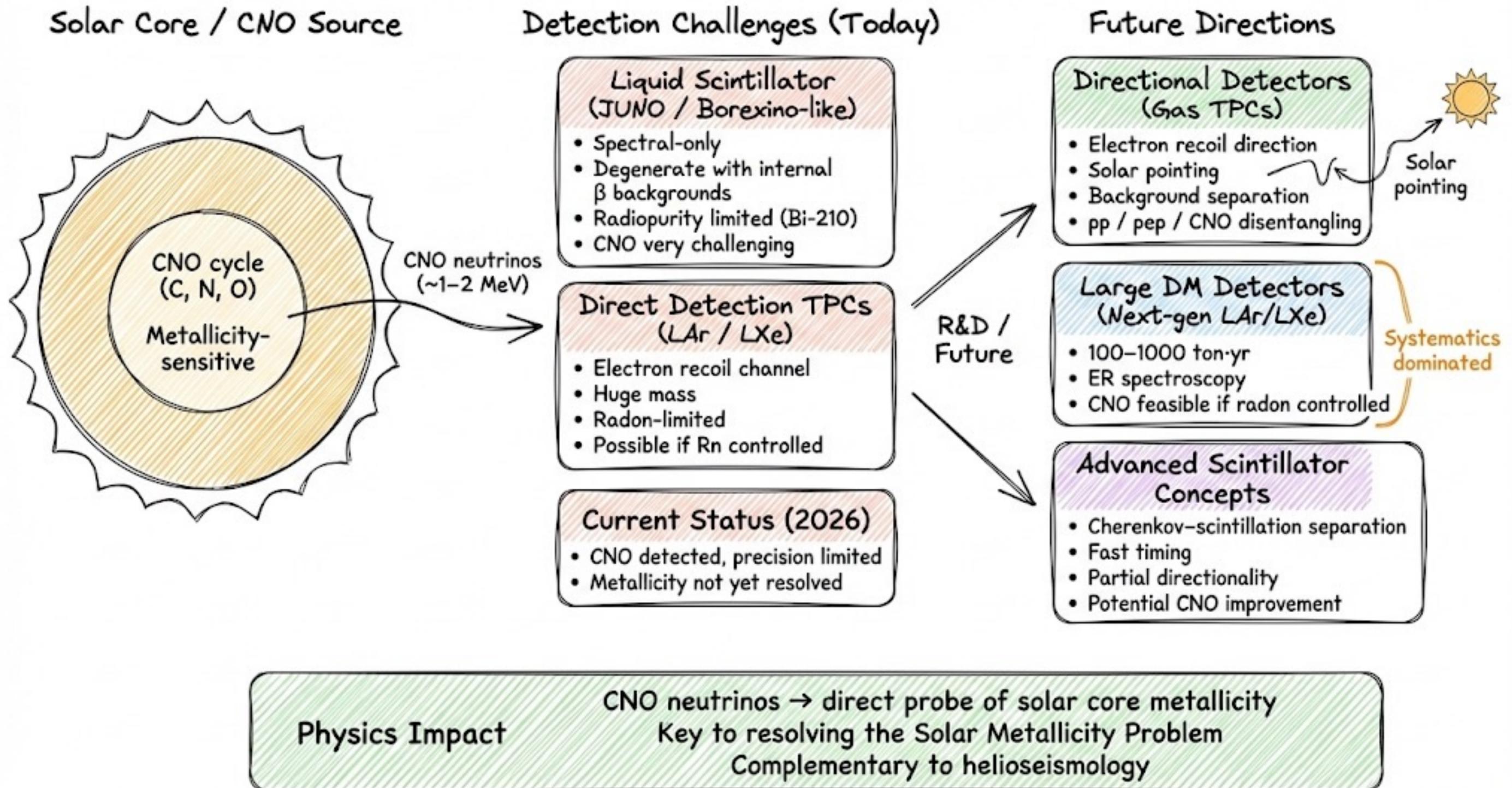
- Limited by internal bkg.
- Direction may help
- Challenging



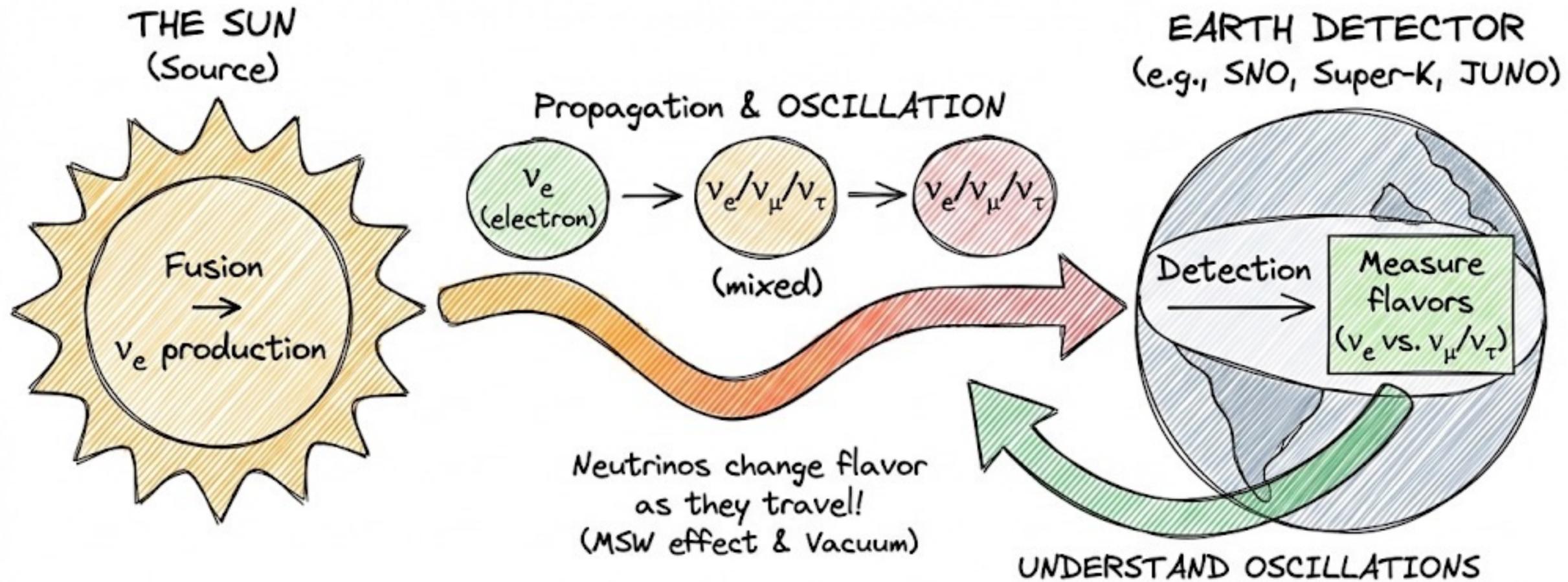
Abusleme, Angel and others. *JUNO Sensitivity to ⁷Be, Pep, and CNO Solar Neutrinos*. March 2023. 2639017. INSPIRE.

<https://arxiv.org/abs/2303.03910>.

CNO in next years



Part 2: study the oscillation

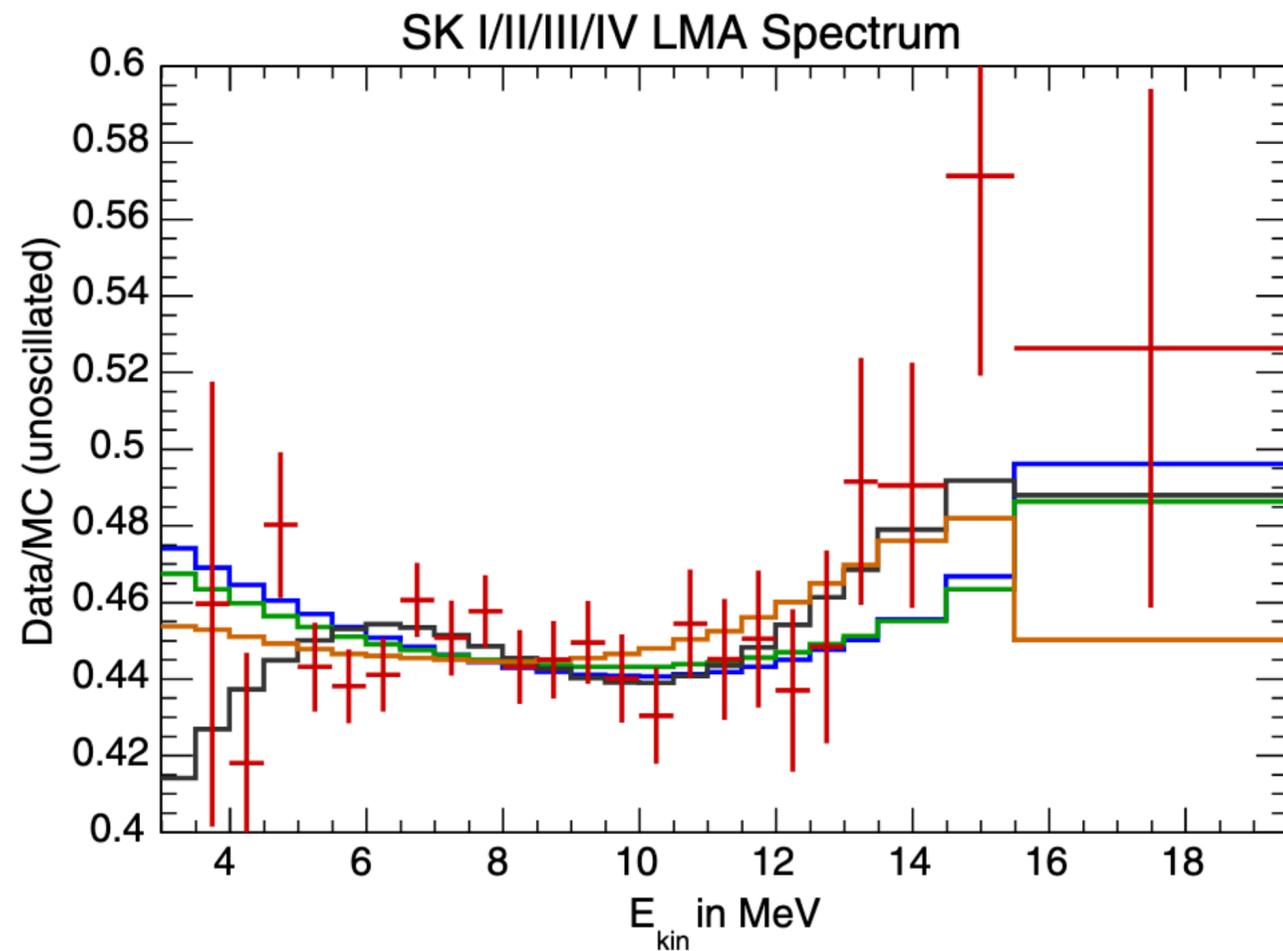


PHYSICS IMPACT: Study Neutrino Oscillations!

- $\nu_e \rightarrow \nu_\mu, \nu_\tau$ conversion
- Verify MSW Effect (Matter)
- Measure Osc. Parameters ($\theta_{12}, \Delta m_{21}^2$)

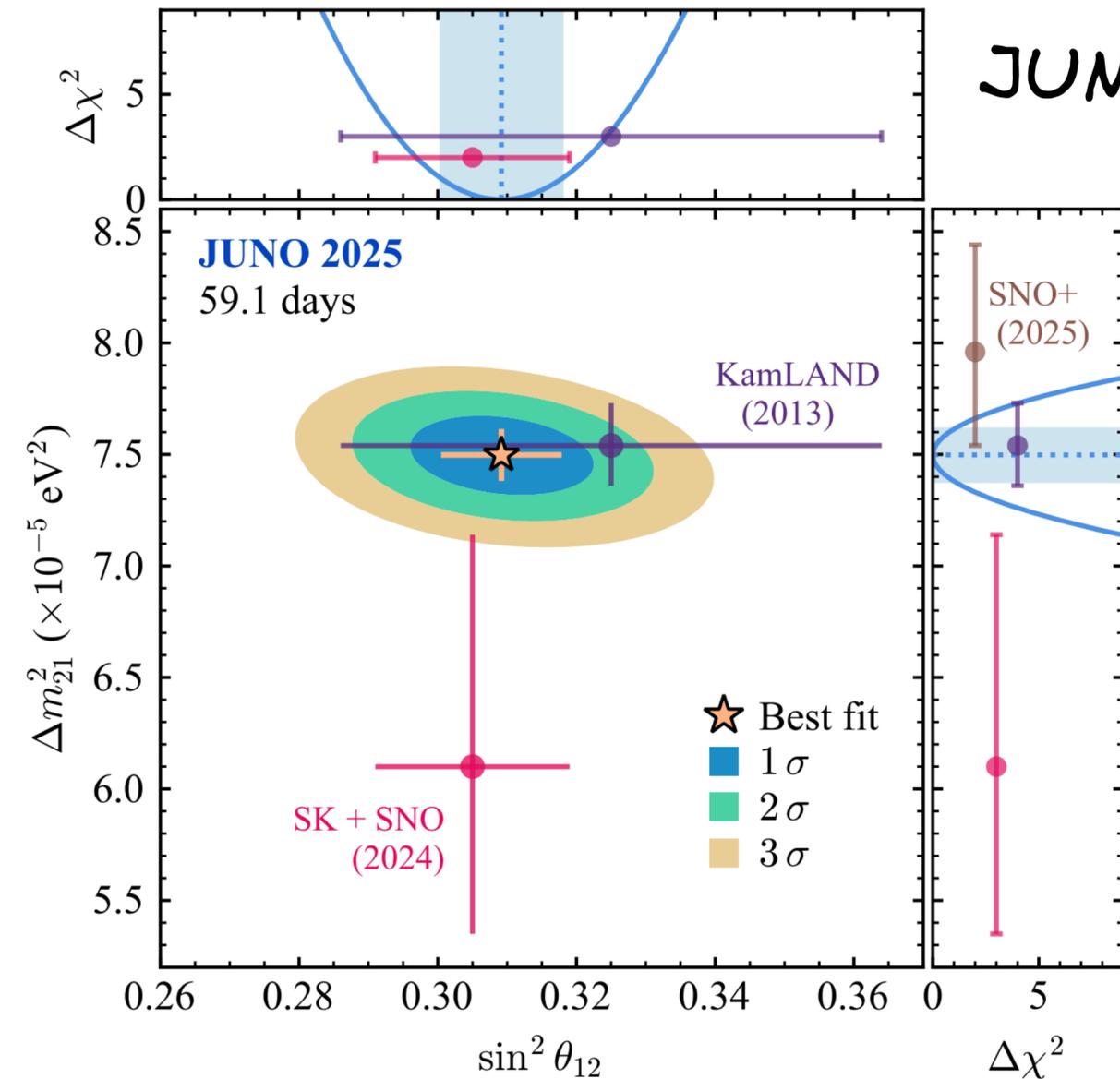
Δm^2_{21} "tension"

SK-I-IV-2016



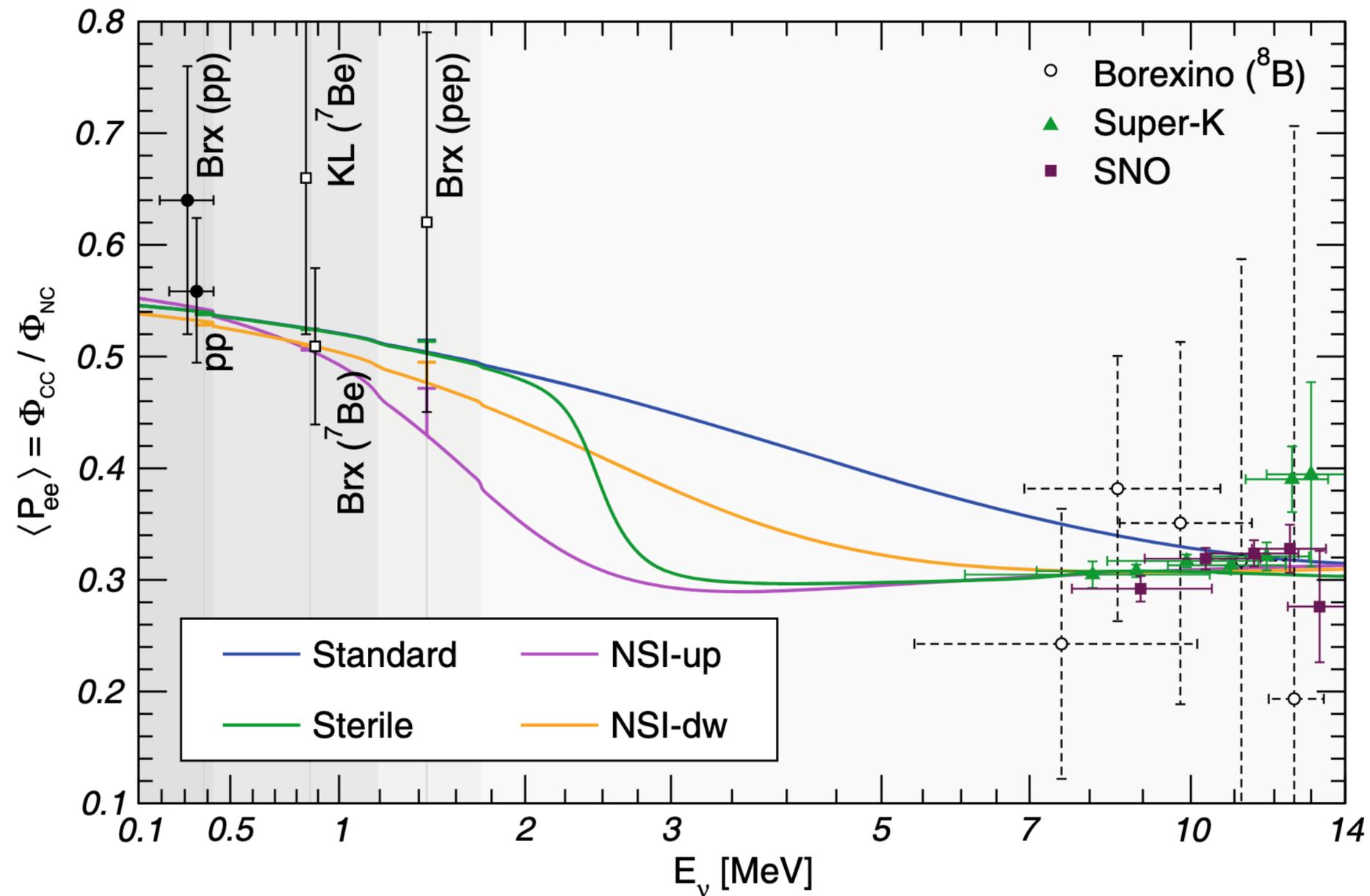
Abe, K., Y. Haga, Y. Hayato, et al. "Solar Neutrino Measurements in Super-Kamiokande-IV." *Physical Review D* 94, no. 5 (2016): 052010. 1472086, p. 052010. INSPIRE. <https://doi.org/10.1103/PhysRevD.94.052010>.

JUNO-2025



Abusleme, Angel, Thomas Adam, Kai Adamowicz, et al. "First Measurement of Reactor Neutrino Oscillations at JUNO." arXiv:2511.14593. Preprint, arXiv, November 18, 2025. <https://doi.org/10.48550/arXiv.2511.14593>.

MSW transition region

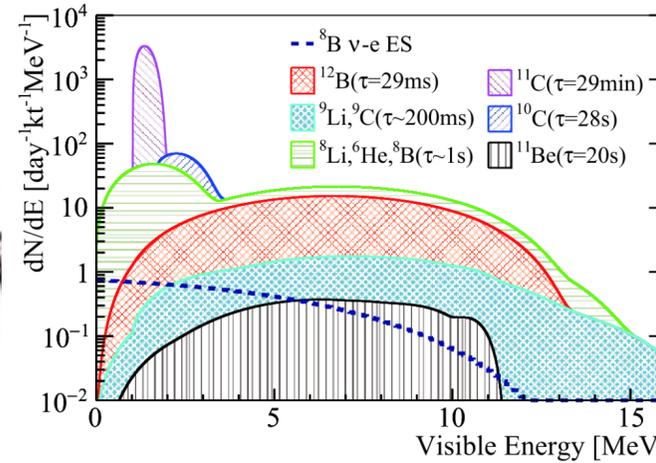


Maltoni, Michele, and Alexei Yu. Smirnov. "Solar Neutrinos and Neutrino Physics." *The European Physical Journal A* 52, no. 4 (2016): 87. 1383834, p. 87. INSPIRE. <https://doi.org/10.1140/epja/i2016-16087-0>.

JUNO: 5 MeV ^8B ES

SUPER SHORT LIVED ($\sim 29\text{ms}$)

B12: 29.1 ms | 1552 cpd/20kt
(65x signal)



LONG LIVED ($>1\text{s}$)

Li8: 1.21 s | 433 cpd/20kt
(20x signal)

B8: 1.11 s | 363 cpd/20kt
(15x signal)

SHORT LIVED ($<1\text{s}$)

Li9: 257.2 ms | 79 cpd/20kt
(3x signal)

C9: 182.5 ms | 143 cpd/20kt
(6x signal)

**^8B Solar ν ES
Channel Bkg
(5 MeV Analysis)
- Cosmogenic**

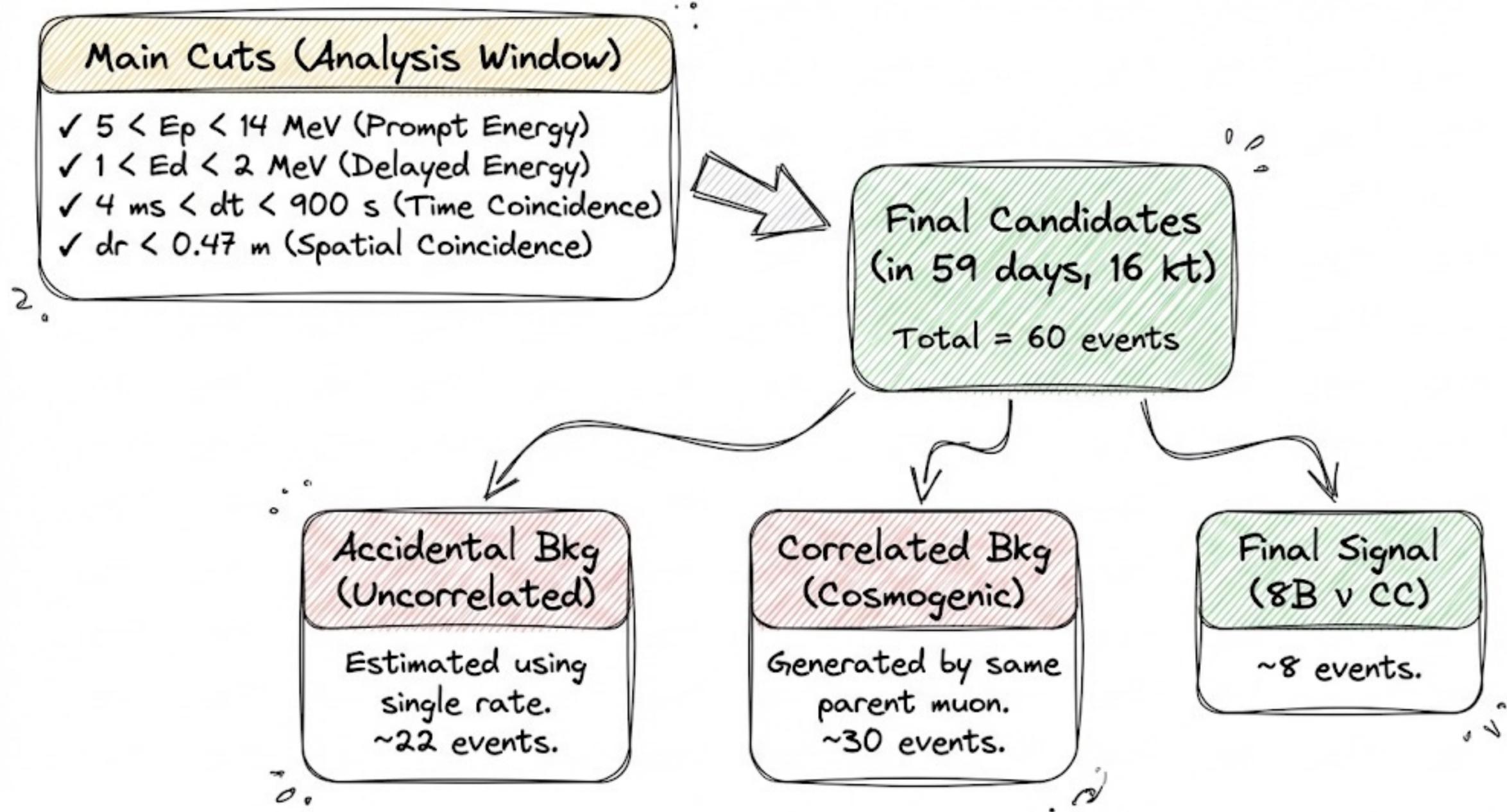
SUPER LONG LIVED ($\sim 20\text{s}$)

Be11: 19.9 s | 40 cpd/20kt
(2x signal)

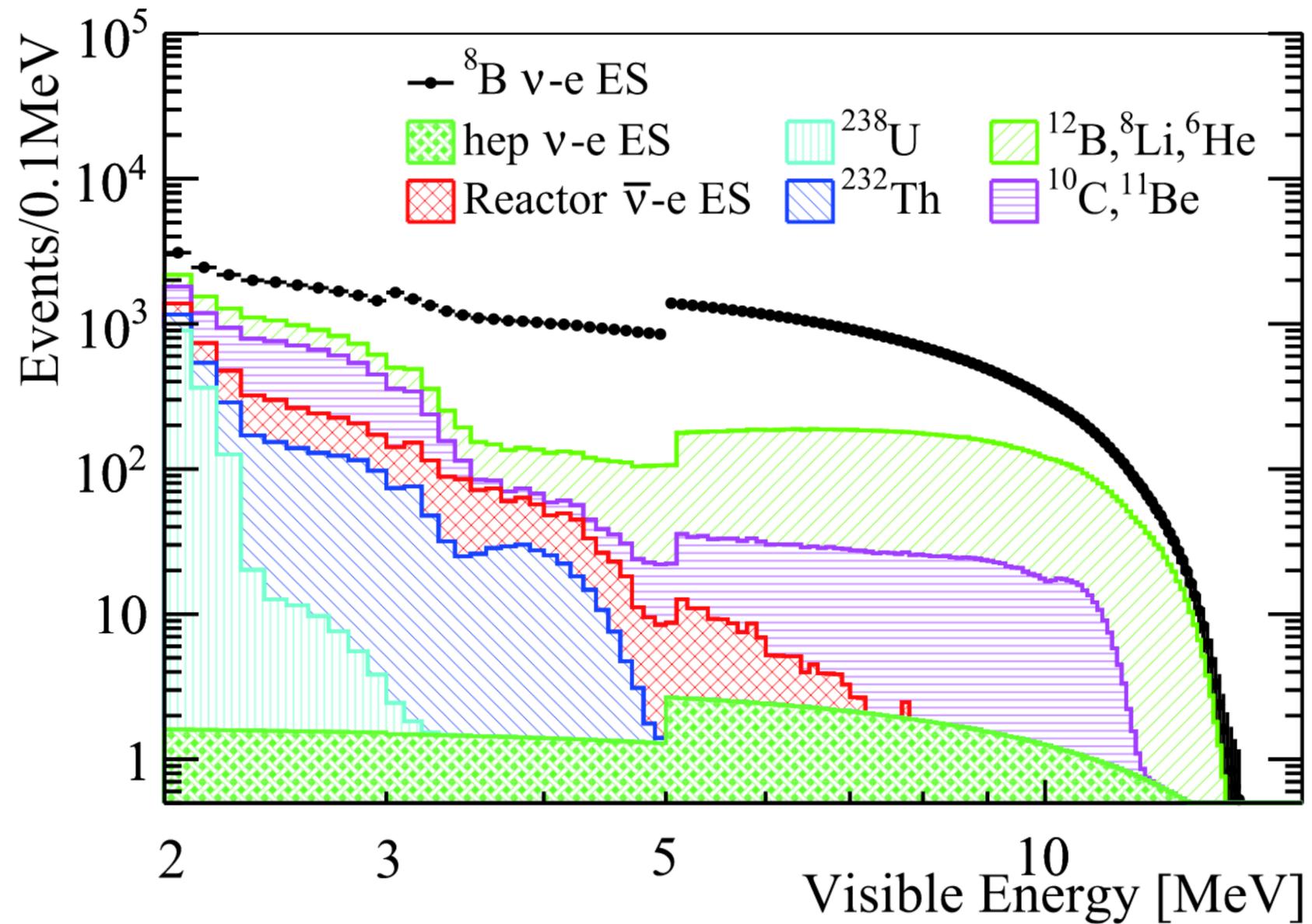
**Overall Bkg Rate:
 ~ 2610 cpd/20kt, 100x signal**

Context: Analysis cut at 5 MeV to suppress these

JUNO: 5 MeV ^8B CC



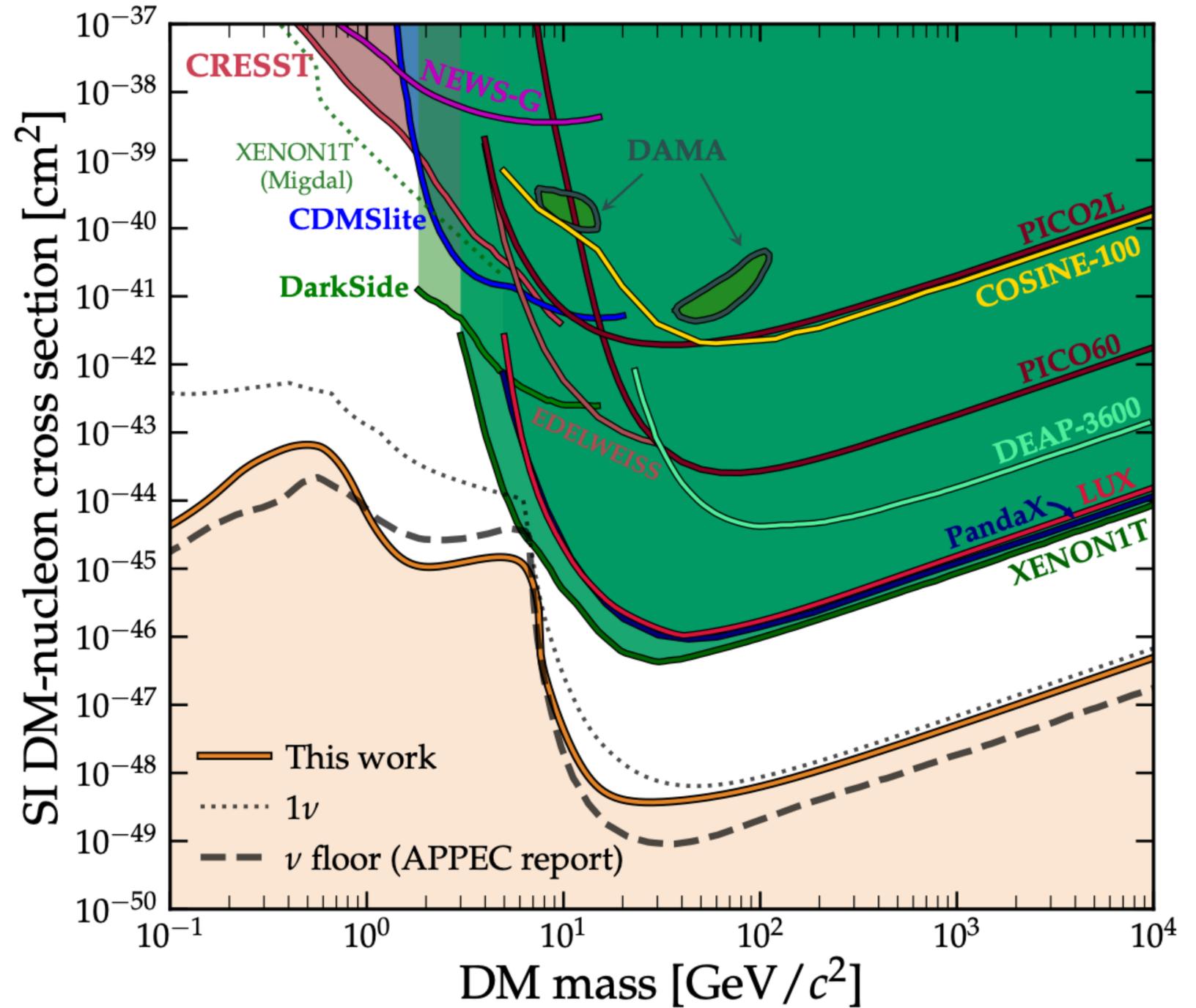
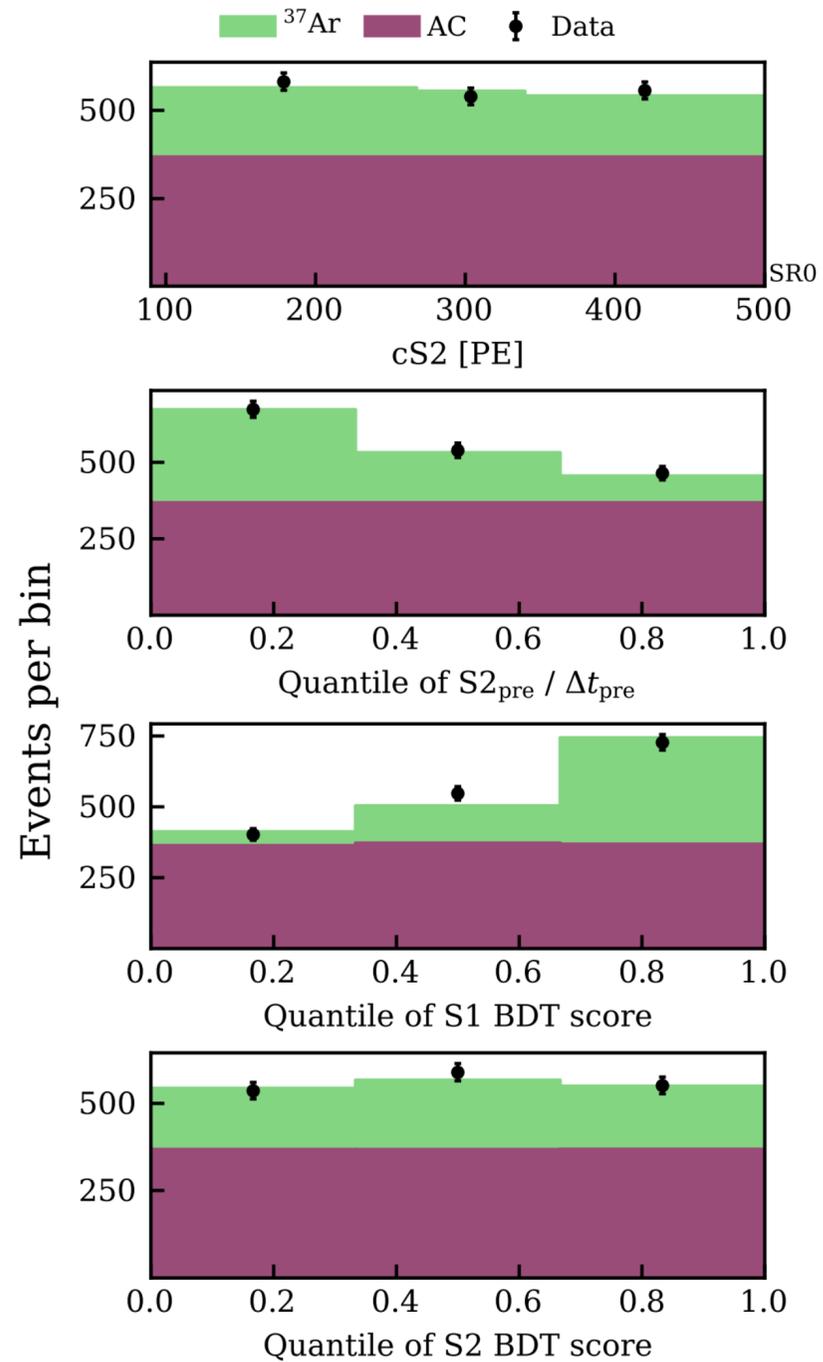
Next: 2 MeV JUNO



Abusleme, Angel, Thomas Adam, Shakeel Ahmad, et al. "Feasibility and Physics Potential of Detecting ^8B Solar Neutrinos at JUNO *." *Chinese Physics C* 45, no. 2 (2021): 023004.

<https://doi.org/10.1088/1674-1137/abd92a>.

Dark matter: ER and NR



Conclusions

- Solar neutrinos can study both the Sun and oscillations
- For solar metallicity problem, precise CNO flux is essential
- For solar oscillations, JUNO may push to 2 MeV
- Dark matter is also evolving fast, yet not world leading.