

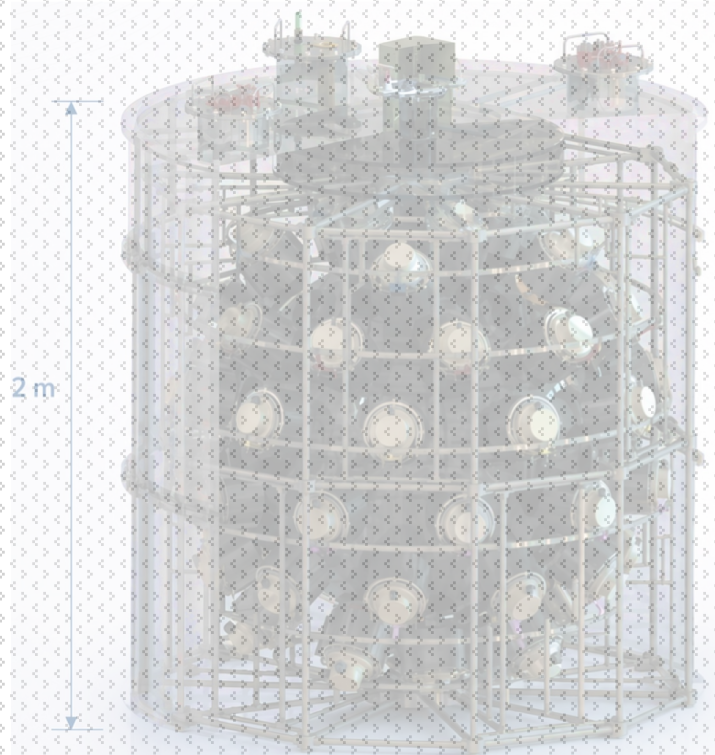


Tsinghua University



Jinping Neutrino Experiment

MeV Event-by-Event Direction Reconstruction in a Novel Cherenkov Liquid Scintillator Detector



Target mass

~1 ton liquid scintillator



Photodetectors

60 · 8-inch PMTs



Geometry

Cylindrical tank
Diameter ~1.6 m, Height ~2.0 m



Shielding

Acrylic tank + stainless steel frame

2026.7.16

Ling Liu, on behalf of the JNE Collaboration

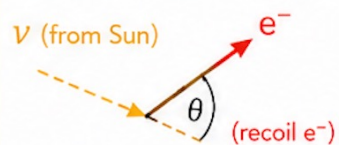
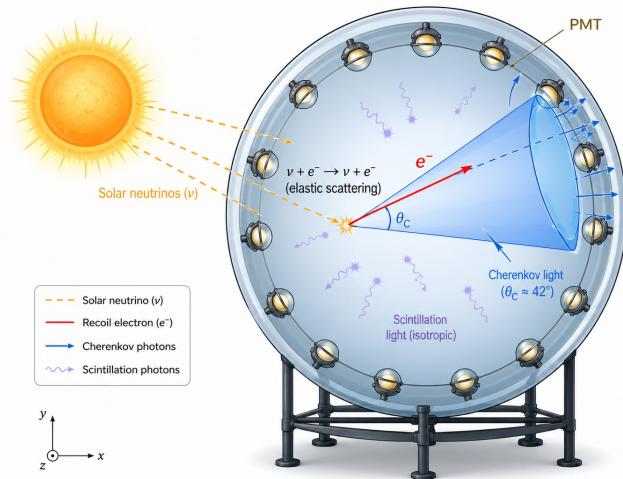
Center for High Energy Physics

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- Motivation
- Experimental Setup
 - Detector Overview
 - Americium–Beryllium (*AmBe*) Calibration Source
- Data Analysis
 - Event Selection
 - Readout and Waveform Analysis
 - Directional Reconstruction Results
- Summary
- References

Solar- ν

$$\nu + e \rightarrow \nu + e \text{ (Elastic Scattering)}$$



Recoil electron correlates with the incoming neutrino direction.

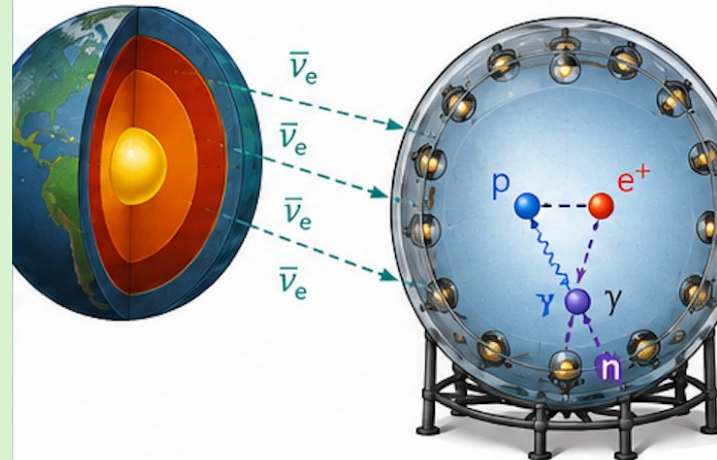
- ✓ Identify the solar direction (known source)
- ✓ Suppress isotopic radioactive background
- ✓ Improve signal significance

Directional information uniquely identifies the solar neutrino signal.

Geo- ν

$$\bar{\nu}_e + p \rightarrow e^+ + n \text{ (Inverse Beta Decay)}$$

$$\nu + e \rightarrow \nu + e \text{ (Elastic Scattering)}$$

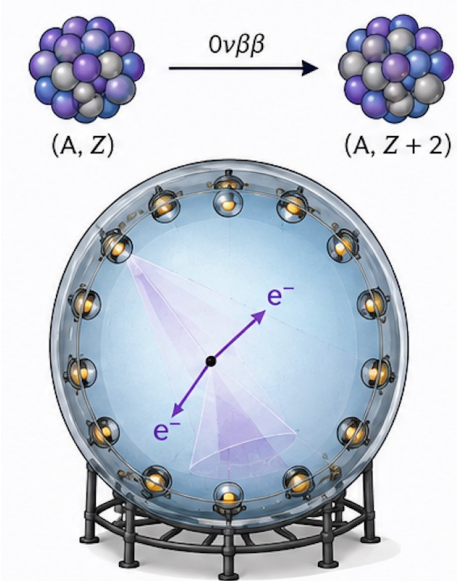


- ✓ Reject solar-neutrino background
- ✓ Separate different neutrino sources (crust / mantle)
- ✓ Improve geoneutrino flux precision

Directional reconstruction helps isolate geo-neutrino candidates.

$0\nu\beta\beta$ Decay

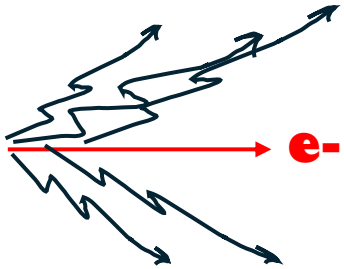
(Neutrinoless Double-Beta Decay)



- ✓ Two-electron topology information
- ✓ Powerful background discrimination

Direction and topology enhance signal identification.

Cherenkov Light

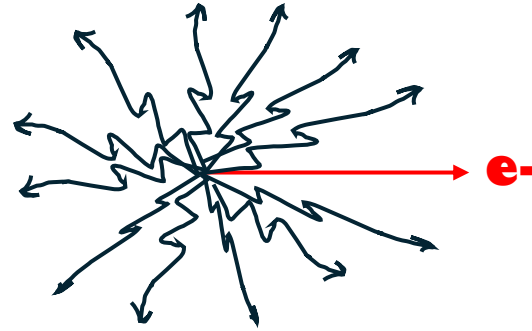


Carry Directional Information



Good Directional Resolution

Scintillator Light



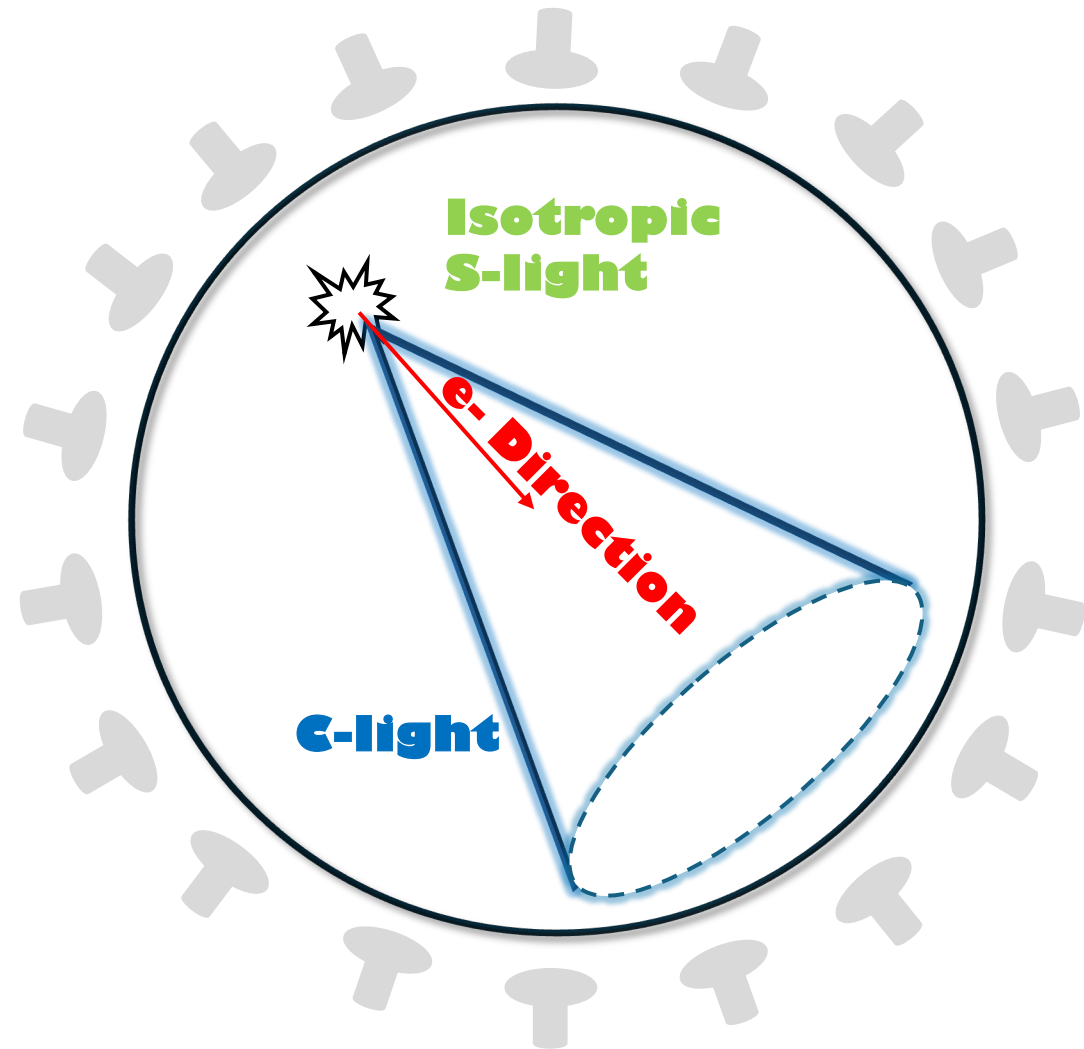
High Light Yield



Good Energy Resolution
Good Vertex Position Resolution



Cherenkov Liquid Scintillator Detector:
Slowing down S-light to separate two kinds of light



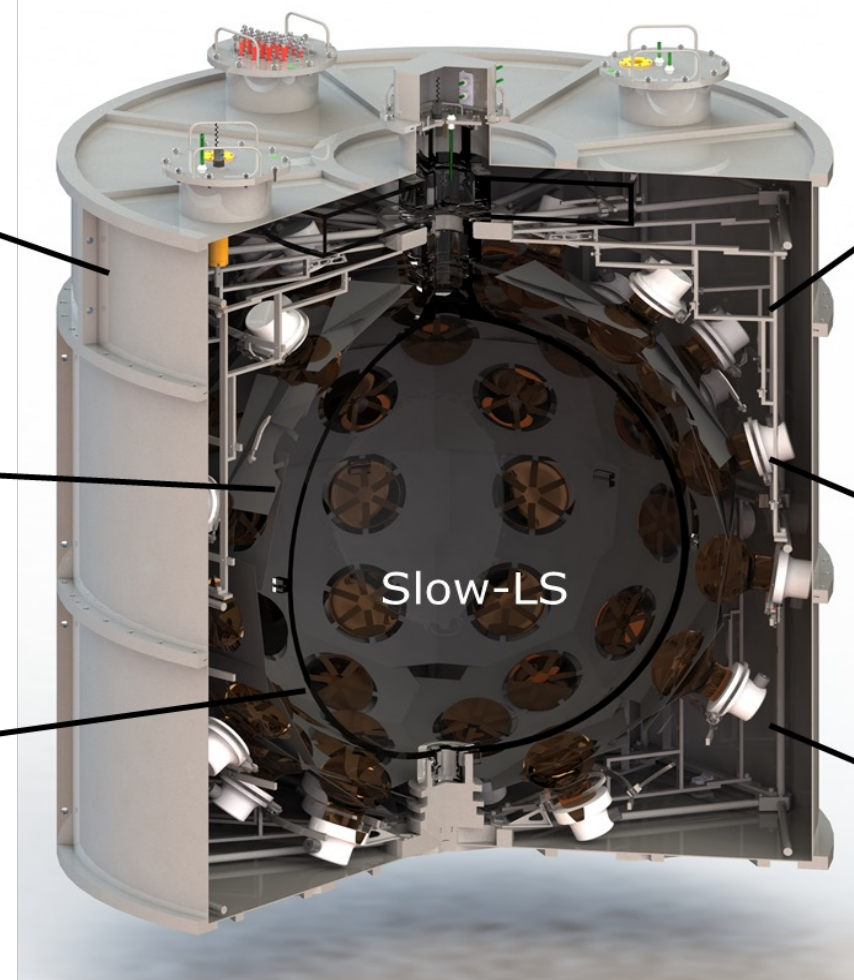
1-ton ProtoType

Stainless steel
container

Light
shielding

Acrylic sphere

645-650mm



PMT support

8-inch MCP-
PMT

- 60(52) channels
- ~34% coverage

Water
shielding

Slow-LS

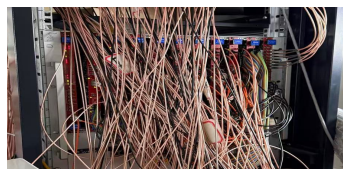
LS Composition	Scintillation light yield(photons/MeV)	Rise Time Constant (ns)	Decay Time Constant(ns)
LAB+0.07g/LPPO	4245	1.67	26.59

Key Factors for Time Resolution



Splitter

CAEN
electronics

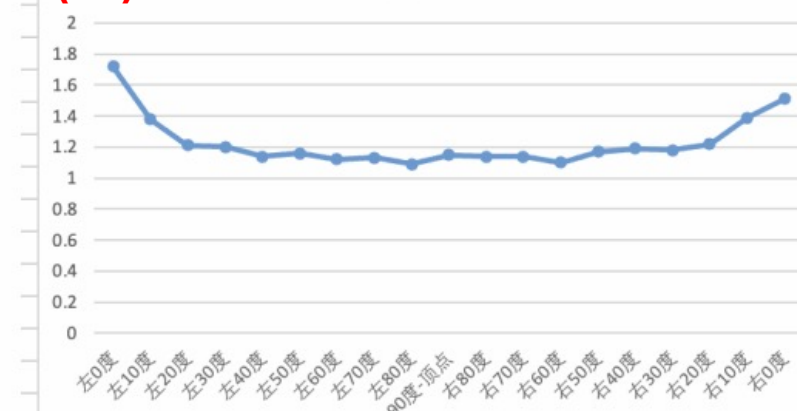


V1751+V1495

Sampling Rate=1GHz



TTS(ns)

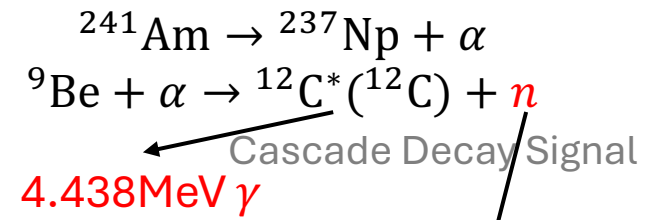
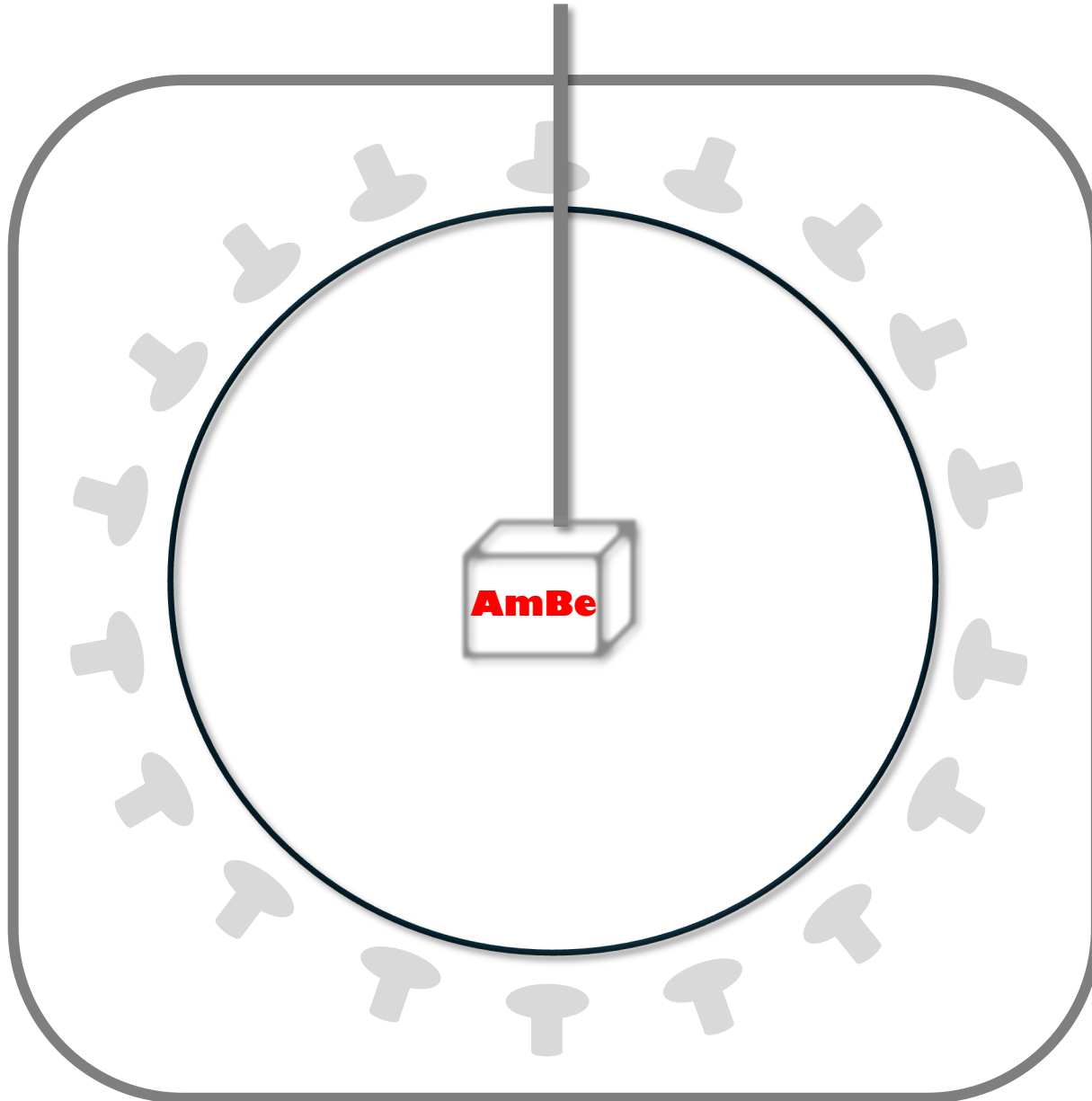


Angle of incident light (degree)

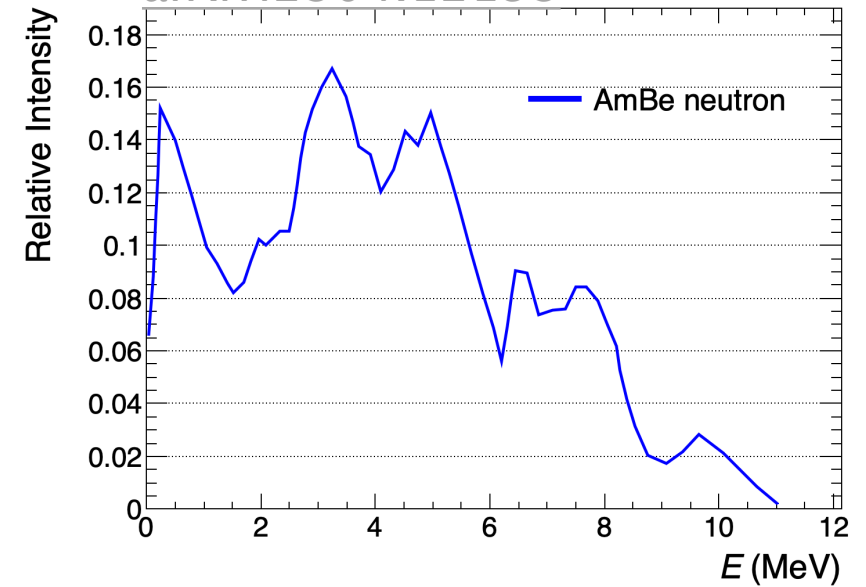
8-inch PhotoMultiplier Tubes:

Transit Time Spread ≈ 1.67 ns (full width at half maximum)

Experimental Setup: Americium–Beryllium (*AmBe*) Calibration Source

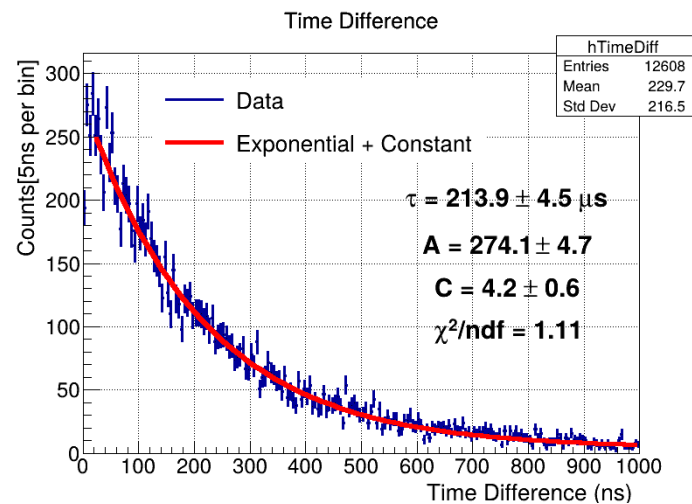
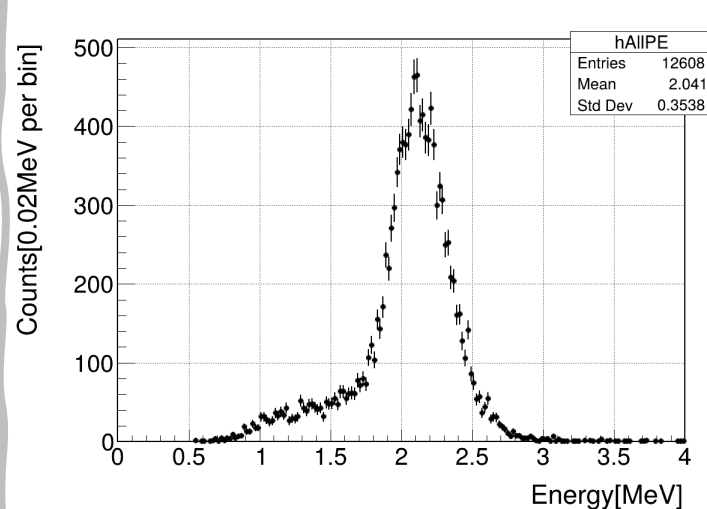


arXiv:2304.12153



Delayed Energy Spectrum Selection:

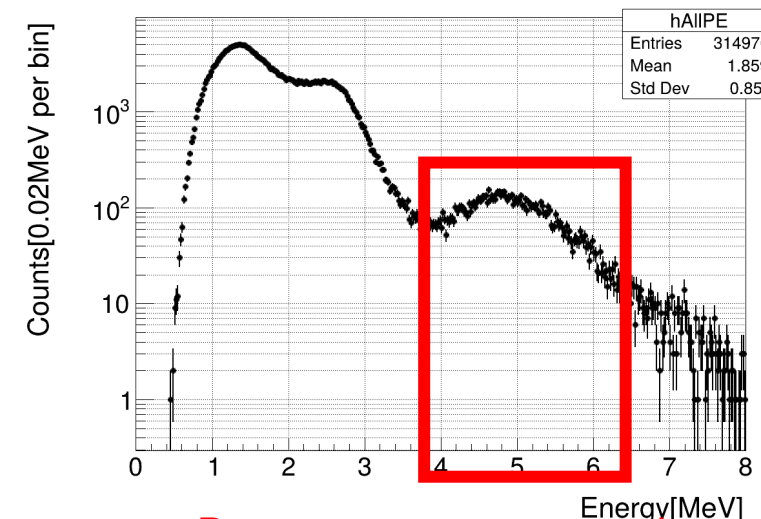
- Fiducial Volume Cut: < 500 mm
- Double coincidence within $1000 \mu\text{s}$
- Prompt energy $\in [3.8, 5.1]$ MeV



- Distinct monoenergetic peak & characteristic $\sim 200 \mu\text{s}$ lifetime
- Low accidental coincidence background

Prompt Energy Spectrum Selection:

- Fiducial Volume Cut: < 500 mm
- Double coincidence within $500 \mu\text{s}$
- Delayed energy $\in [1.7, 2.7]$ MeV



Reconstruct events selected within the red box

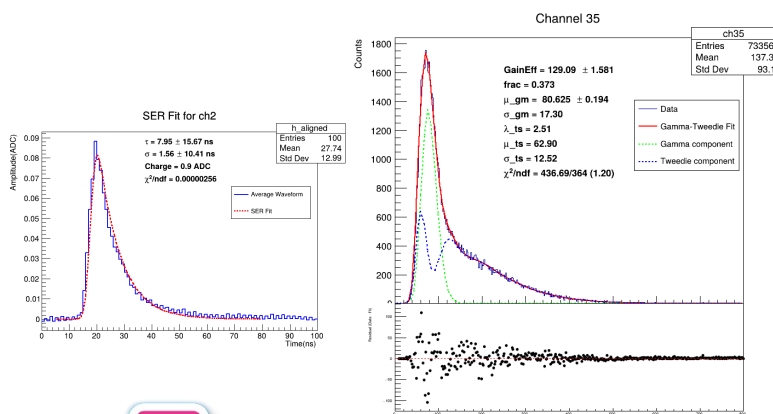
- Clear $4.438 \text{ MeV } \gamma$ signal
- High accidental coincidence background

PMT Gain spectrum

$$\mathcal{L}(n_i^{\text{Obs}}, t_{ij}, R_{ij}) = f_{\text{Poisson}}(n_i^{\text{Obs}}; n_{\text{average}}) \cdot \prod_{\text{windowlength}} \text{Gauss} \left(\left[V_{ik} - \sum_{j=1}^{n_i^{\text{Obs}}} R_{ij} \text{PE}_{\text{single}}(t - t_{ij}) \right], \sigma_{\text{base}} \right)$$

Residual Between True and Fitted Waveforms

$$\text{Gauss} \left(\left[V_{ik} - \sum_{j=1}^{n_i^{\text{Obs}}} R_{ij} \text{PE}_{\text{single}}(t - t_{ij}) \right], \sigma_{\text{base}} \right)$$



3

Fit by Minimizing the Negative Log-Likelihood :

- ✓ PE Number
- ✓ PE Time

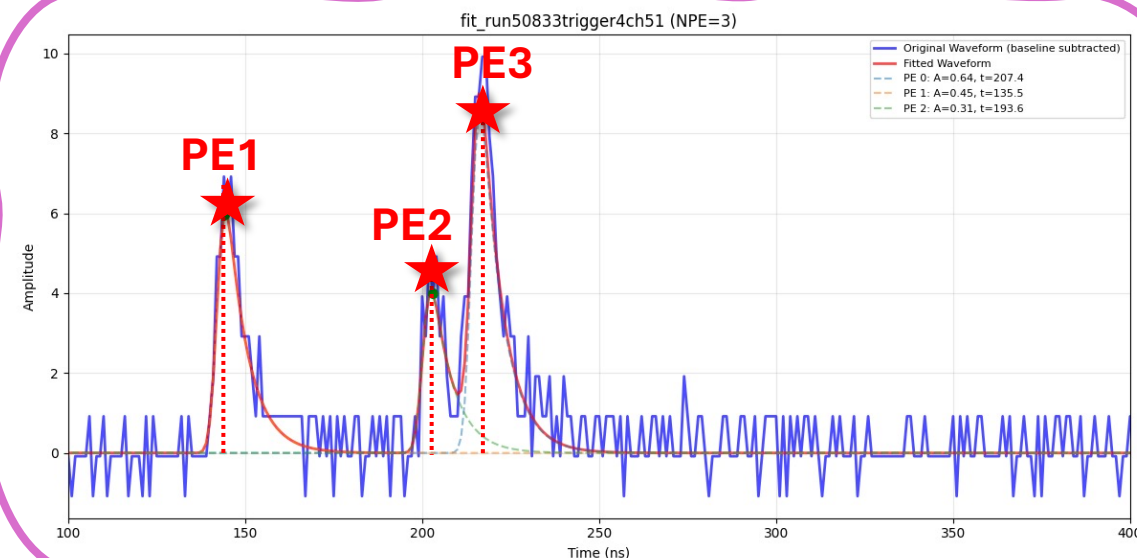
1

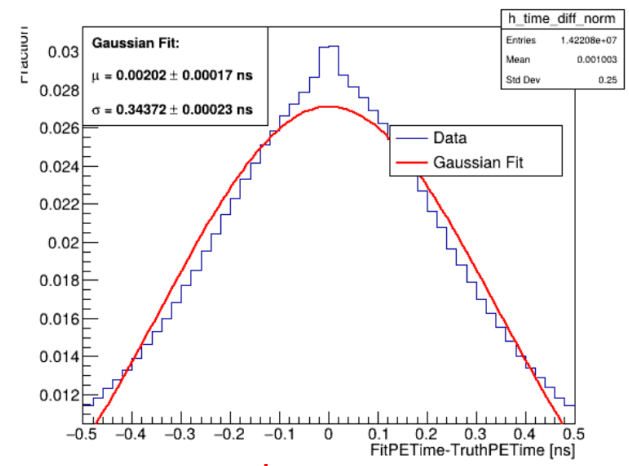
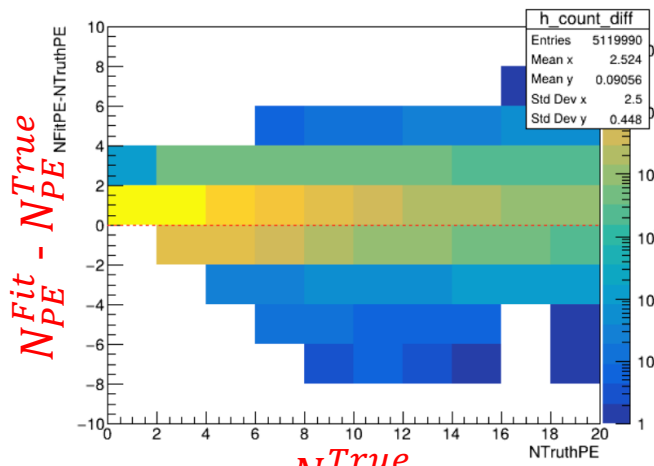
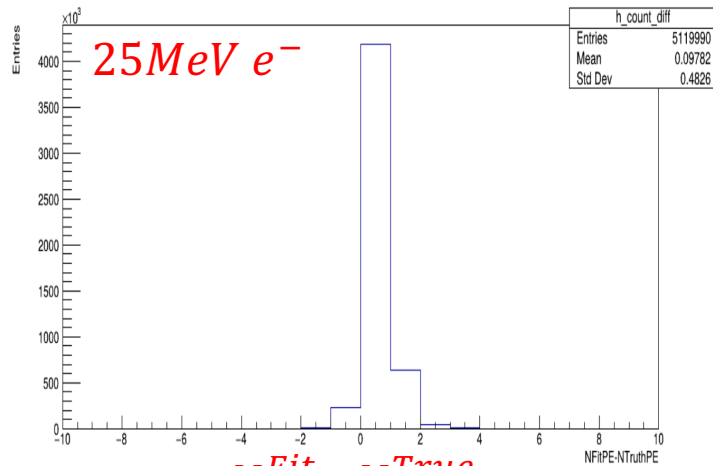
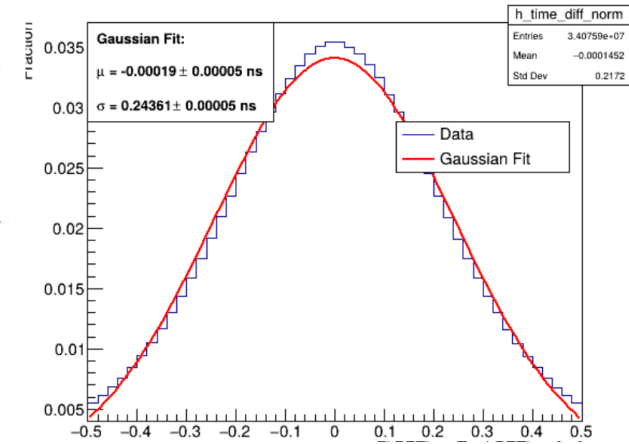
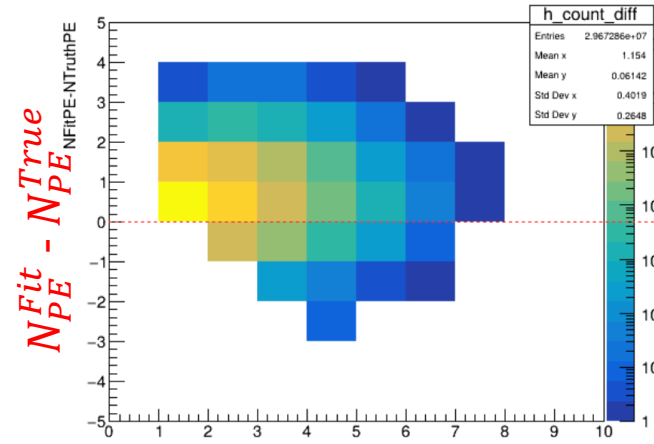
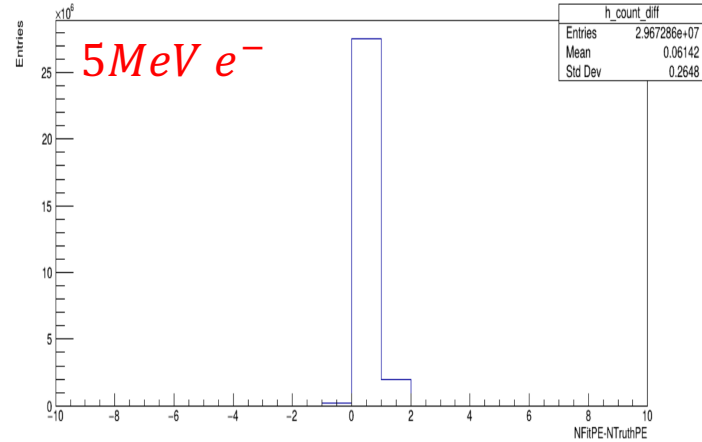
Get Single Electron Pulse
Get Gain Spectrum

2

Initialization

- Initial Photoelectron (PE) Count
(estimated from the integrated charge / effective gain)
- Initial PE Time
(estimated by waveform deconvolution)





$N_{PE}^{Fit} - N_{PE}^{True}$

N_{PE}^{True}

$t_{PE}^{Fit} - t_{PE}^{True}$

- PE **count** estimation error < **0.5 PE**, with most events concentrated around zero.
- PE **time** estimation error < **0.34 ns**.

$$\mathcal{L}(n_i^{\text{Obs}}, t_{ij} | E, x, y, z, t_{\text{event}}, \vec{d}_{\text{Fit}}) = \prod_i^{N_{\text{PMT}}} P_i^C \prod_j^{n_i^{\text{Obs}}} P_{ij}^T$$

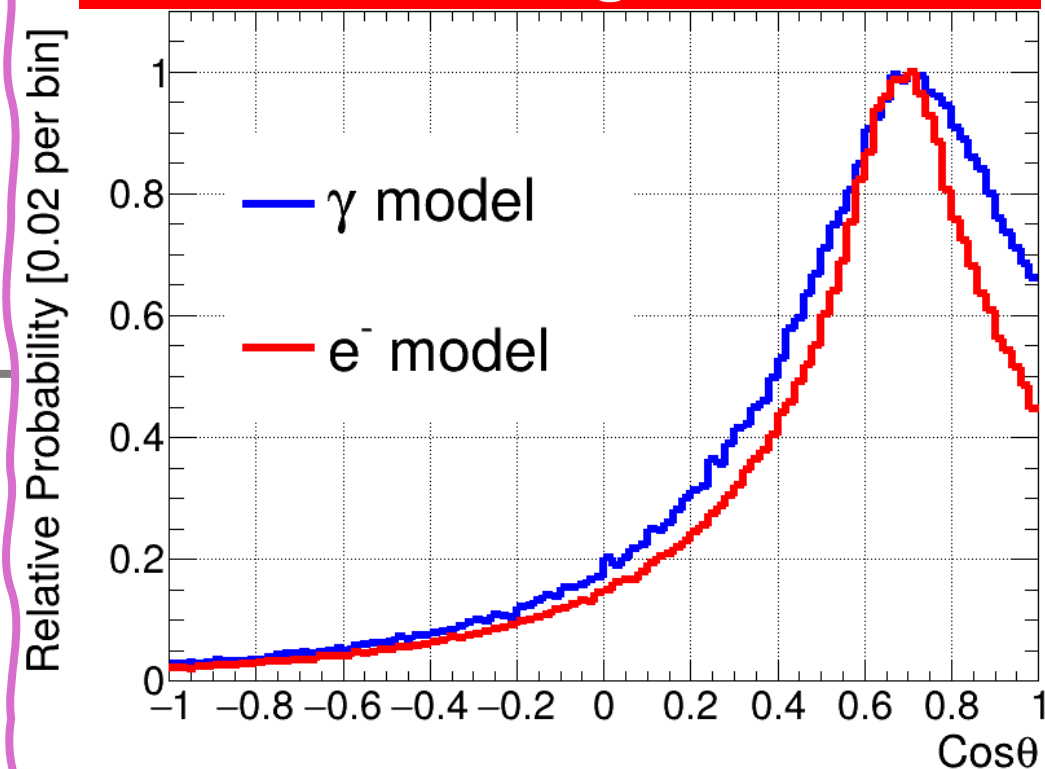
PE Term PETime Term

- PE Term
 - ✓ Photon emission
 - ✓ Photon propagation
 - ✓ Photon detection

- PETime Term
 - ✓ Photon path length in different materials
 - ✓ **Group velocity** in different materials

input

MC Cherenkov Angle Distribution



γ model differs from the e⁻ model

$$\mathcal{L}(n_i^{\text{Obs}}, t_{ij} | E, x, y, z, t_{\text{event}}, \vec{d}_{\text{Fit}}) = \prod_i^{N_{\text{PMT}}} P_i^C \prod_j^{n_i^{\text{Obs}}} P_{ij}^T$$

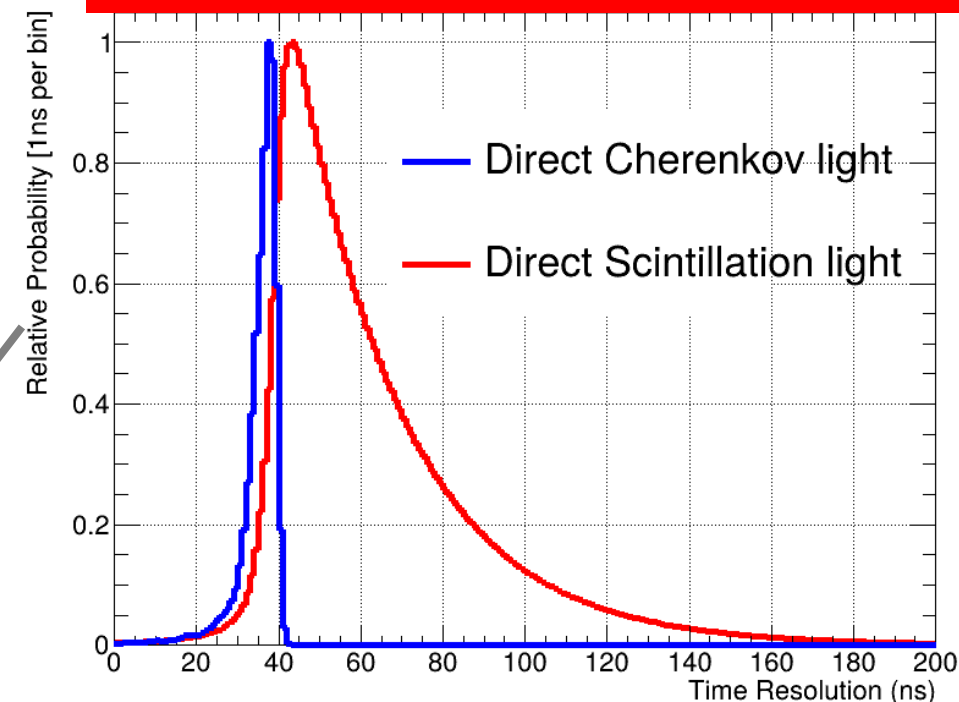
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- PE Term
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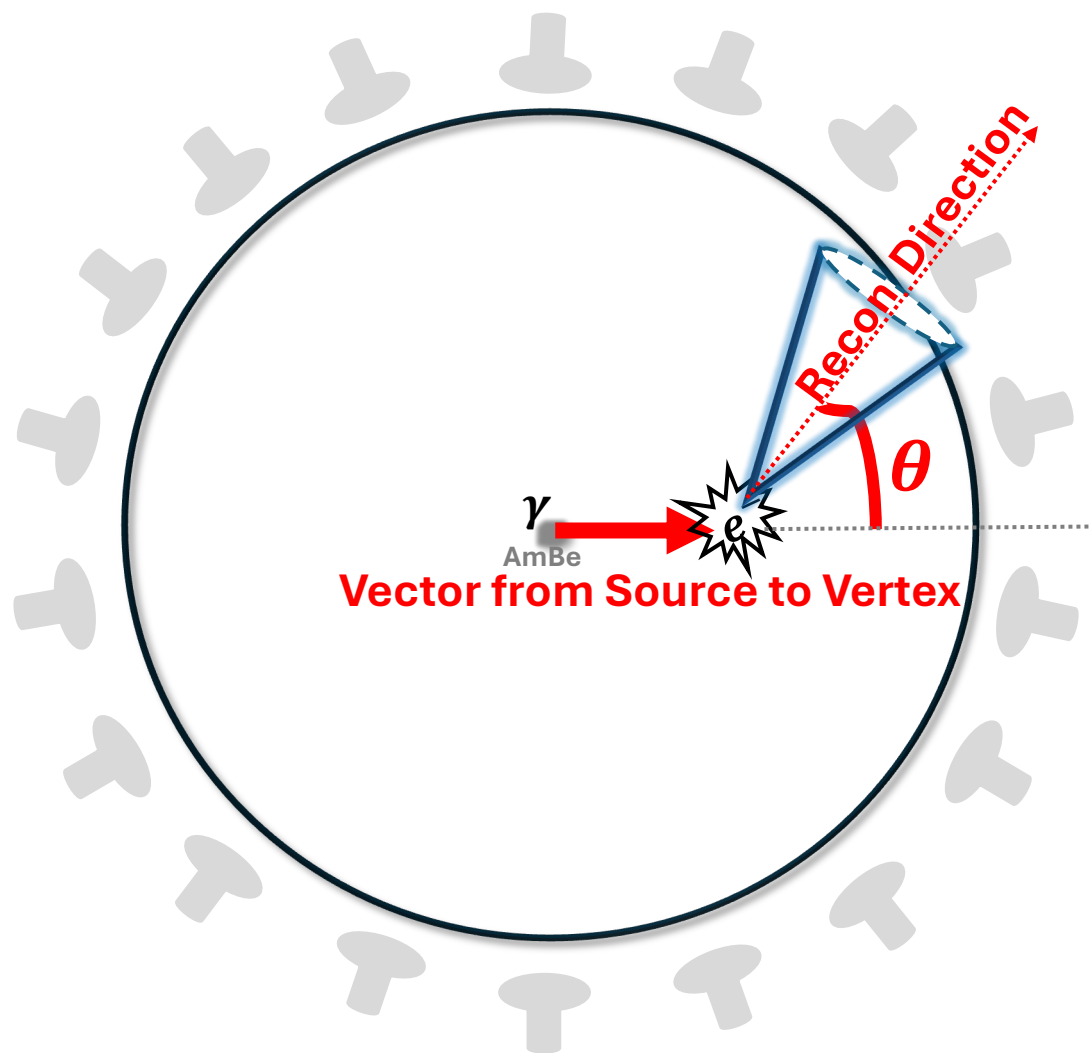
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 - ✓ Photon path length in different materials
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input

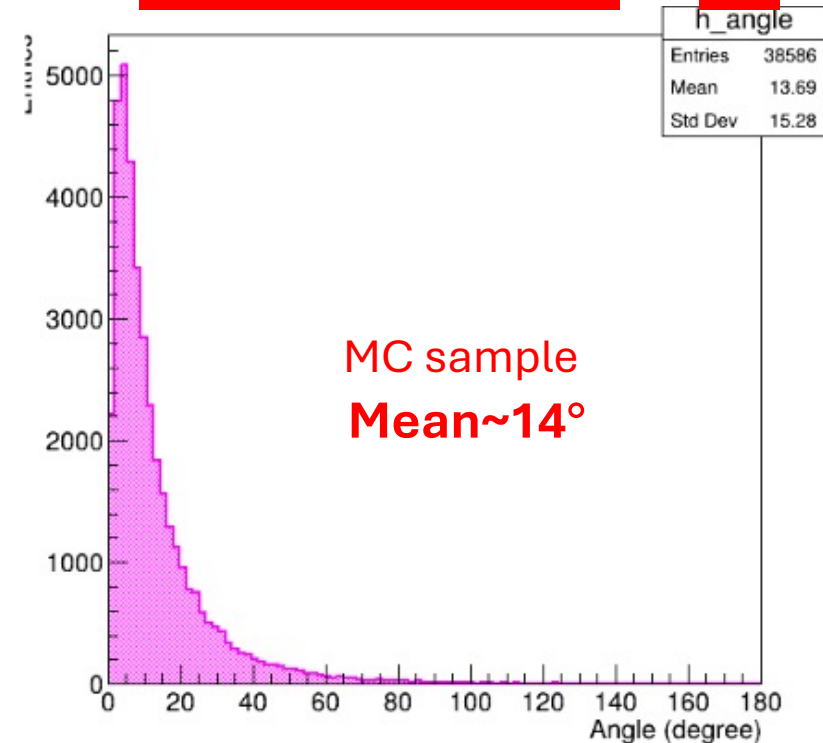
MC Time Residual Distribution



Independent modeling of Cherenkov and Scintillation light time residuals



Angle between **Vector from Source to vertex** and **True direction**

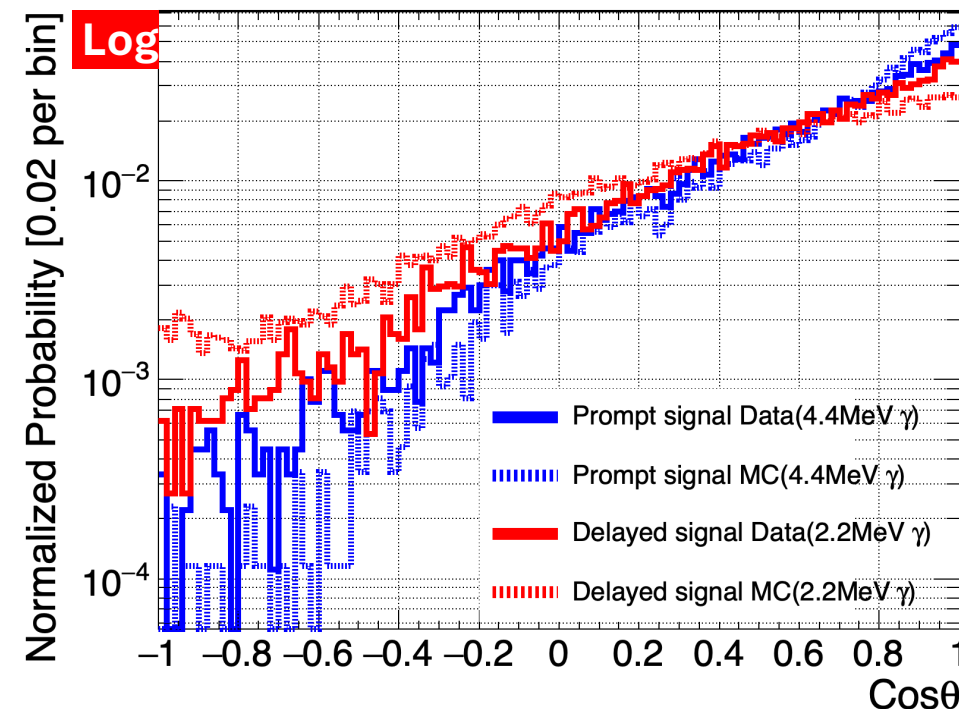
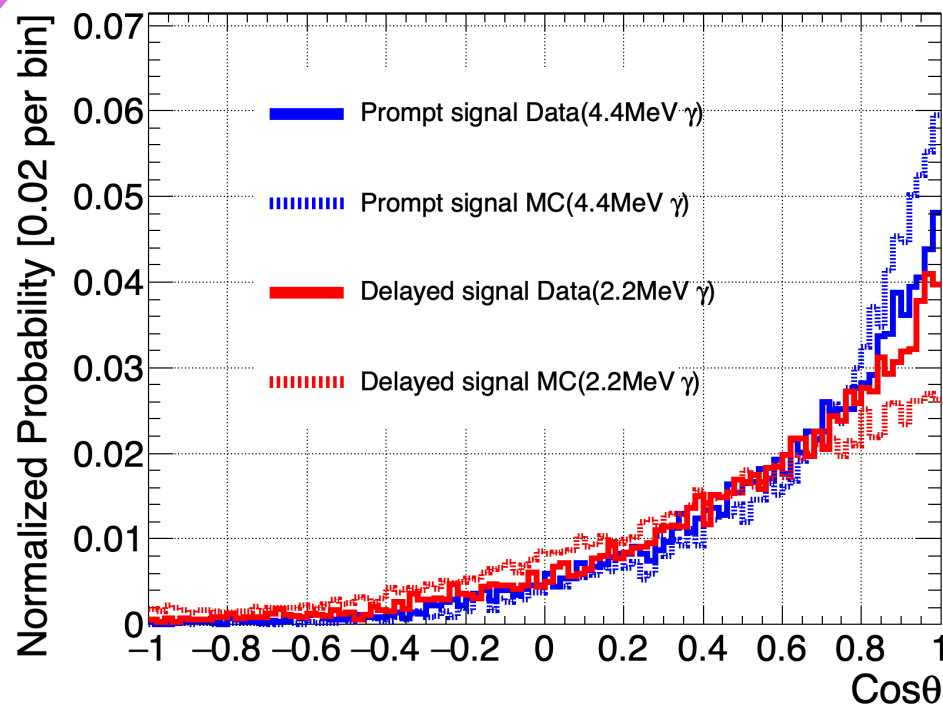


Direction Definition in DATA:

vector from the source position to the event's reconstructed position.

(SNO+ also used this method)

[arXiv:1812.03355](https://arxiv.org/abs/1812.03355)



θ : angle between **Vector from Source to Vertex** and **Recon** direction

- Clear directional dependence is observed, with good **DATA-MC** agreement.
- Prompt-signal** region shows **better** directional reconstruction performance than the Delayed-signal region.
- Angular resolution is evaluated using the **Real True** Direction for center-emitted γ events in simulation:

	Delayed-signal (@2.2MeV γ)	Prompt-signal (@4.4MeV γ)
Angular resolution (68.3% containment angle)	67°	43°

❑ Novel Cherenkov liquid scintillator detectors enable event direction reconstruction.

Requirements:

- **Long** Scintillation photons **Time Constant**
- **High Sampling Rate** electronics
- **Low** PMT **Transit Time Spread**

❑ Self-developed waveform analysis and event reconstruction algorithms perform well in simulation.

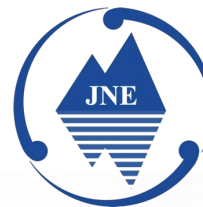
❑ **First demonstration of MeV event-by-event direction** reconstruction in a liquid scintillator detector.

❑ Calibration source events are selected from data. Good **DATA–MC agreement** is achieved.

- I. Luo, Wentai, et al. "Reconstruction algorithm for a novel Cherenkov scintillation detector." *Journal of Instrumentation* 18.02 (2023): P02004.
- II. Allega, A., et al. "Event-by-event direction reconstruction of solar neutrinos in a high light-yield liquid scintillator." *Physical Review D* 109.7 (2024): 072002.
- III. Anderson, M., et al. "Measurement of the B 8 solar neutrino flux in SNO+ with very low backgrounds." *Physical Review D* 99.1 (2019): 012012.
- IV. Ito, Hiroshi, et al. "Analyzing the neutron and γ -ray emission properties of an americium–beryllium tagged neutron source." *Nuclear Instruments and Methods in Physics Research Section A: Accelerators, Spectrometers, Detectors and Associated Equipment* 1057 (2023): 168701.
- V. Kostal, M., et al. "Validation of Heavy Water Cross Section Using AmBe Neutron Source". EPJ. Web of Conferences, vol. 239, 2020, p. vp., doi:10.1051/epjconf/202023918008.
- VI. Marsh, John W. et al. "High resolution measurements of neutron energy spectra from AmBe and AmB neutron sources." *Nuclear Instruments & Methods in Physics Research Section A-accelerators Spectrometers Detectors and Associated Equipment* 366 (1995): 340-348.
- VII. Zhang, Aiqiang, et al. "Performance evaluation of the 8-inch MCP-PMT for Jinping Neutrino Experiment." *Nuclear Instruments and Methods in Physics Research Section A: Accelerators, Spectrometers, Detectors and Associated Equipment* 1055 (2023): 168506.
- VIII. Kamiokande Collaboration. "Supernova Relic Neutrino Search with Neutron Tagging at Super-Kamiokande-IV." *arXiv preprint arXiv:1311.3738* (2013).
- IX. Kamiokande Collaboration. "First measurement of radioactive isotope production through cosmic-ray muon spallation in Super-Kamiokande IV." *arXiv preprint arXiv:1509.08168* (2015).



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Thanks



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~1 ton liquid scintillator



Photodetectors

60 8-inch PMTs



Geometry

Cylindrical tank
Diameter ~1.6 m, Height ~2.0 m



Shielding

Acrylic tank + stainless steel frame

2 m

