

QCD相变临界点及奇异双重子态的实验寻找

“重子数涨落与重子-重子相互作用：一场重子的狂欢盛宴”



罗晓峰

华中师范大学

2026年4月25日

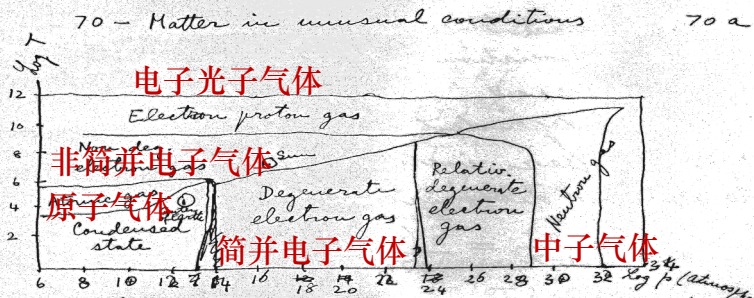


极端条件下的物质形态



E. Fermi

E. Fermi: "Notes on Thermodynamics and Statistics" (1953)



Start from ordinary condensed matter with ordinary equation of state controlled by ordinary chemical forces.

a) Increase pressure at $T < 1000$ until deg. electron energies exceeds 20 eV —

Condition
$$\bar{w} = \frac{3}{40} \left(\frac{6}{\pi} \right)^{2/3} \frac{h^2 n^{2/3}}{2^{2/3} m} \quad \rho = \frac{2}{3} \bar{w} n$$

$$\bar{w} = 3.6 \times 10^{-27} n^{2/3} = 3.2 \times 10^{-11}$$

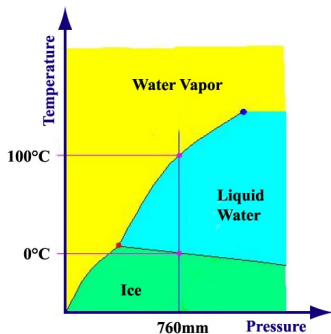
$$n \approx 10^{24} \quad \rho = \frac{2}{3} 3.2 \times 10^{-11} \times 10^{24} \approx 2 \times 10^{13} \text{ at}$$

As pressure increases beyond this point

$$\rho = 3.6 \times 10^{-27} n^{2/3} \quad n \times \frac{2}{3} = 2.4 \times 10^{-27} n^{5/3}$$

$$n = 6 \times 10^{23} \frac{\rho}{A} Z \quad \rho = 10^{13.01} \left(\frac{\rho Z}{A} \right)^{5/3} \approx 3.2 \times 10^{12} \rho^{5/3}$$

- 171 -



极端高温高密环境中的物质形态及相结构是什么？

水

$\sim 10^2 \text{ }^\circ\text{C}$



太阳

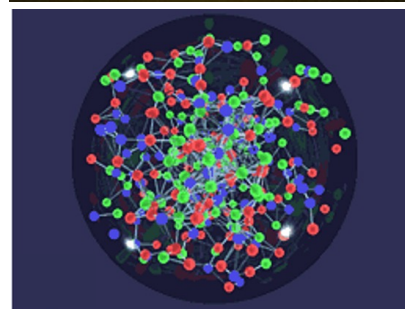
$\sim 10^7 \text{ }^\circ\text{C}$



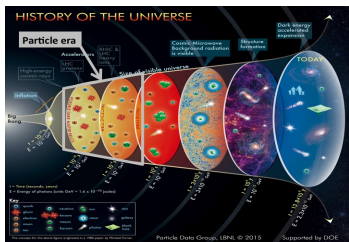
QGP

$\sim 10^{12} \text{ }^\circ\text{C}$

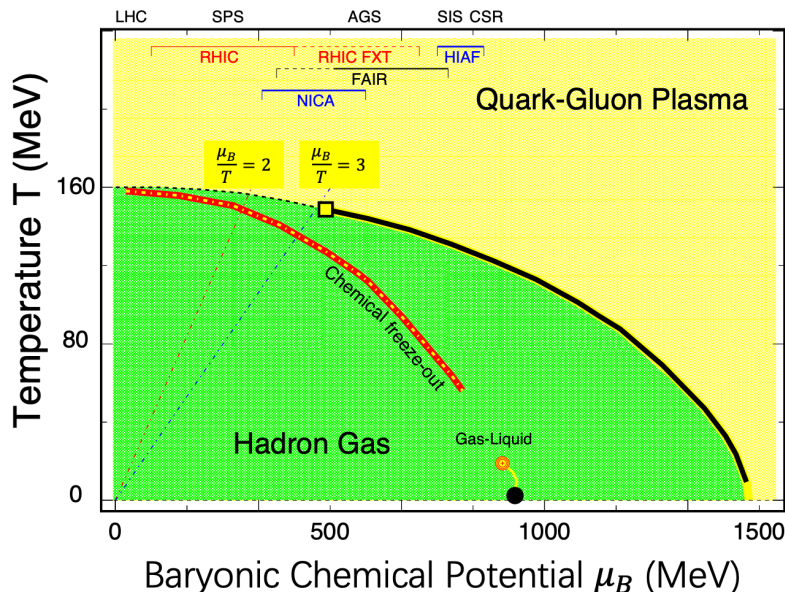
(Early Universe)



QCD Phase Diagram



宇宙演化



关键科学问题：高重子密度区是否存在QCD临界点？

它是国际核物理领域待解决的关键科学问题之一，也是我国大科学装置强流离子加速器(HIAF)的一个主要物理目标(2025年在广东惠州建成)。

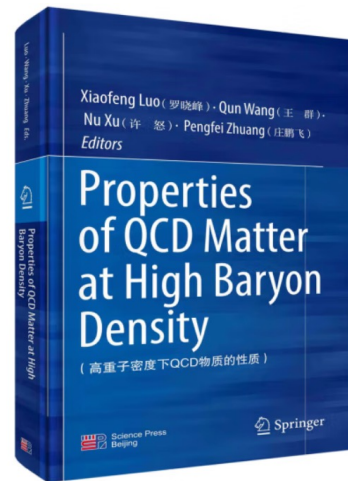
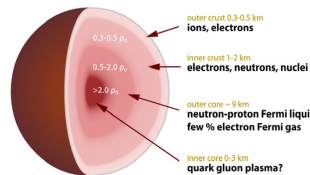
马余刚、许怒、刘峰，基于HIAF集群的QCD相结构研究，*中国科学:物理学 力学 天文学*,2020,50(11):124

赵红卫,从HIAF到CNUF, *现代物理学知识*, 第36卷(2024年), 第1期

Lattice QCD : at $\mu_B = 0$, smooth crossover. $T_c \sim 156$ MeV
Large μ_B : 1st order phase transition and QCD critical point ?

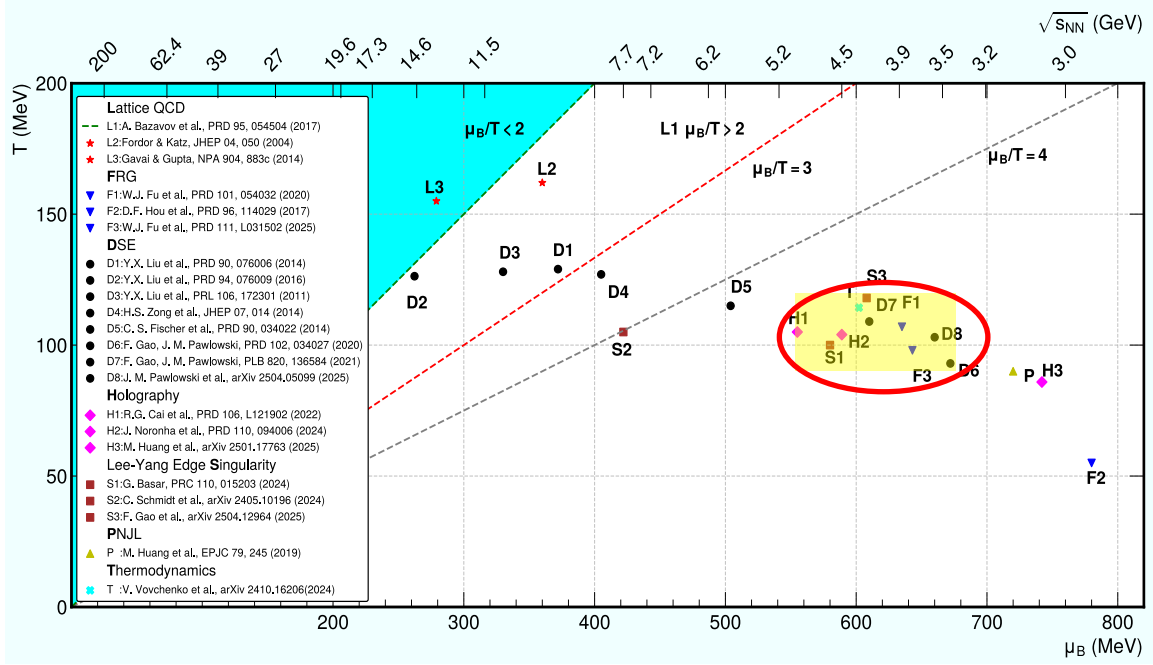
Y. Aoki et al., *Nature* 443, 675 (2006) ;
A. Bazavov et al (HotQCD), *PRD* 85, 054503 (2012).
K. Fukushima and C. Sasaki, *PPNP*, 72, 99 (2013).
A. Bzdak et al., *Phys. Rep.* 853, 1 (2020).

中子星



<https://doi.org/10.1007/978-981-19-4441-3>

Theoretical Estimations (2004-2025)



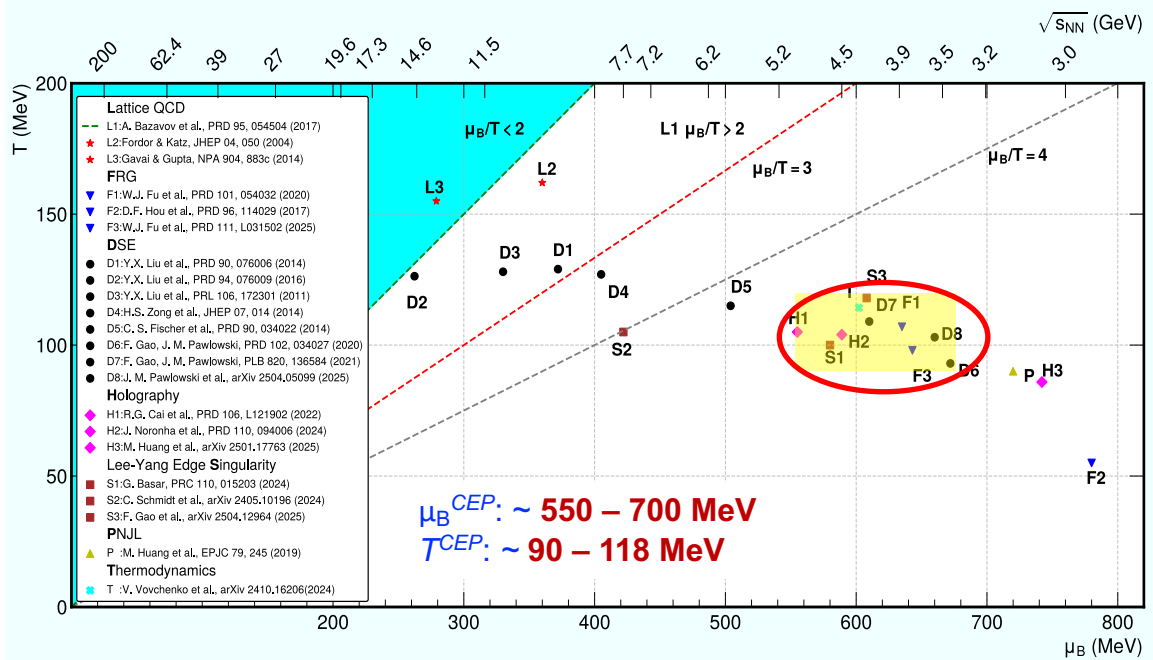
“Converge region” : $\mu_B^{CEP} : \sim 550 - 700 \text{ MeV}$
 $T^{CEP} : \sim 90 - 118 \text{ MeV}$

FRG : Wei-jie Fu, J. M. Pawłowski, and F. Rennecke, Phys. Rev. D 101, 054032 (2020) [441 citations]
 Wei-jie Fu, Chuang Huang, Jan M. Pawłowski, Fabian Rennecke, Rui Wen, Shi Yin, 2603.13455

DSE : Fei Gao, J. M. Pawłowski, Phys. Lett. B 820, 136584 (2021) [179 citations]
 Yi Lu, Fei Gao, Yu-Xin Liu, Jan M. Pawłowski, Phys. Rev. D 113, 054019 (2026)

See Wei-jie Fu's talk

Theoretical Estimations (2004-2025)

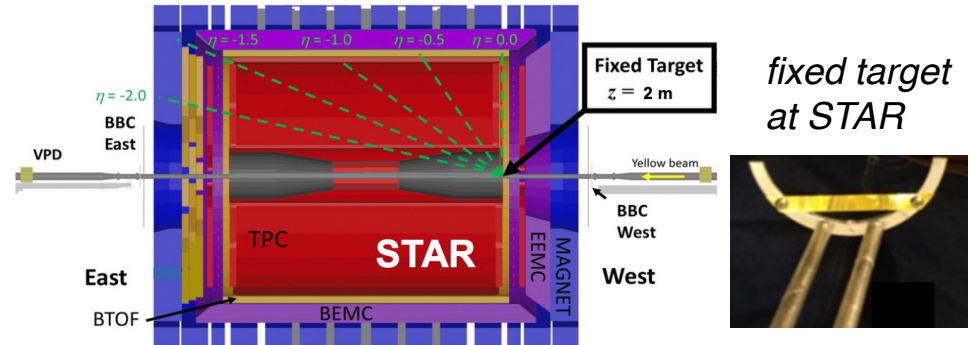
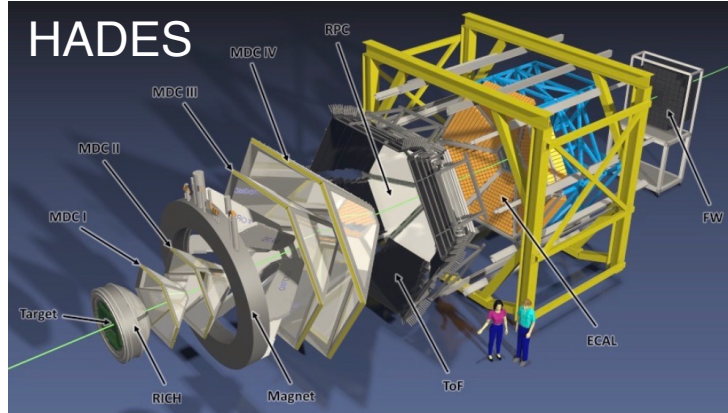
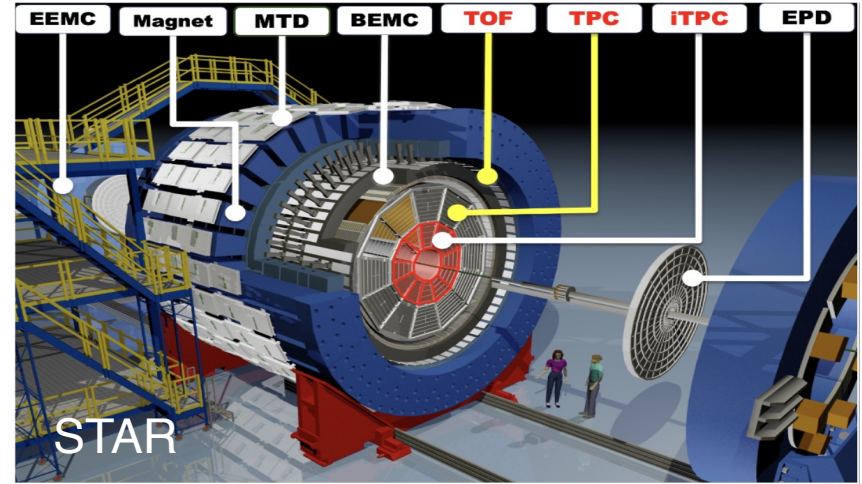
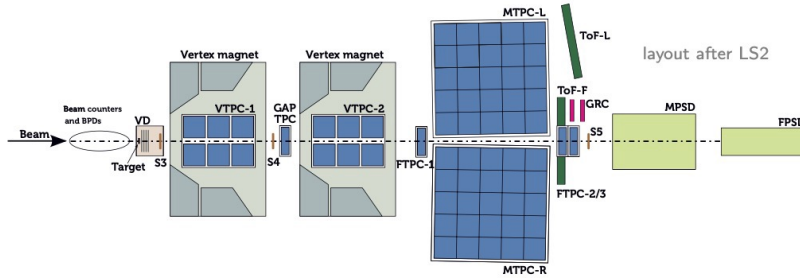


“...experimental measurements are essential to determine whether a QCD critical point exists.” *2023 US Long Range Plan for Nuclear Science*

Y. Zhang, Z. Wang, X. Luo, N. Xu, arXiv : 2602.08356
 Christian S. Fischer, Jan M. Pawłowski, arXiv : 2603.11135

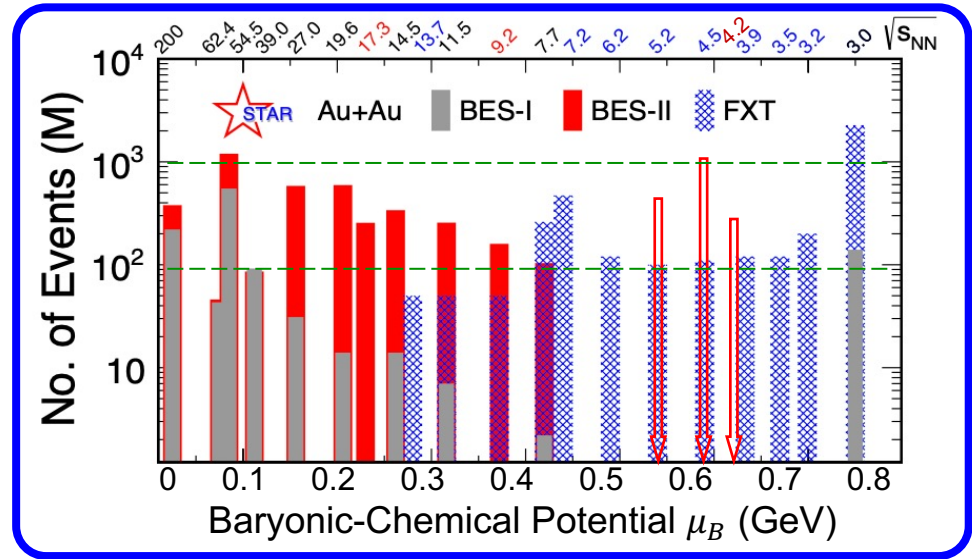
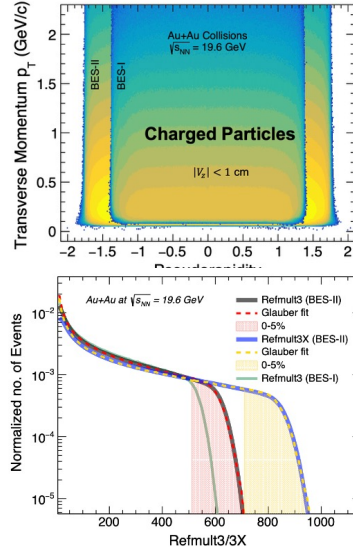
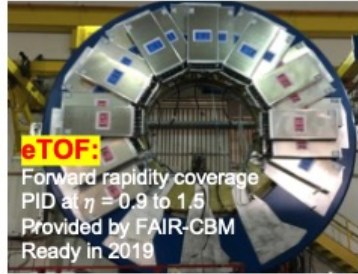
Search for QCD Critical Point in Heavy-ion Collisions Experiments

NA61/SHINE



RHIC Beam Energy Scan Program (2010 – 2026)

Detector Upgrades BES-II



(1) Enlarge rapidity acceptance; (2) Improve particle identification; (3) Better centrality/event plane resolution

- 1) BES-I (2010 – 2014): 7.7, 11.5, 14.5, 19.6, 27, 39, 62.4, 200 GeV;
- 2) BES-II (2018 – 2026): Collider mode (7.7, 9.2, 11.5, 14.6, 17.3, 19.6, 27 GeV),
FXT mode: (3.0, 3.2, 3.5, 3.9, 4.2, 4.5, 5.2, 6.2, ..., 13.7 GeV)
- 3) μ_B coverage : $25 < \mu_B < 750$ MeV

X. Luo and N. Xu, Nucl. Sci. Tech. 28, 112 (2017)
Jinhui Chen, et al., Nucl. Sci. Tech. 35, 214 (2024)

Observables : Fluctuations of Conserved Quantities

➤ **At critical point with an infinite system**

- correlation length should diverge
- susceptibilities should diverge

$$\chi_q^{(n)} = \frac{1}{VT^3} \times C_{n,q} = \frac{\partial^n (p/T^4)}{\partial (\mu_q)^n}, q = B, Q, S$$

Conserved Charges q : Net Baryon Number (B), Net Charge (Q), Net Strangeness (S)

➤ **Experimental Observables: Cumulants of Conserved Quantities Dis.**

- 1) Sensitive to correlation length (ξ)
- 2) Directly related to the susceptibility ratios

$$\langle\langle (\delta N)^3 \rangle\rangle_c \approx \xi^{4.5}, \quad \langle\langle (\delta N)^4 \rangle\rangle_c \approx \xi^7$$

Notation: Cumulants: $C_{1,2,\dots}$
Factorial Cumulants: $\kappa_{1,2,\dots}$

$$\begin{aligned} \kappa_1 &= C_1 = \langle N \rangle, \\ \kappa_2 &= -C_1 + C_2, \\ \kappa_3 &= 2C_1 - 3C_2 + C_3, \\ \kappa_4 &= -6C_1 + 11C_2 - 6C_3 + C_4, \end{aligned}$$

Measured multiplicity N , $\langle \delta N \rangle = N - \langle N \rangle$

mean: $M = \langle N \rangle = C_1$

variance: $\sigma^2 = \langle (\delta N)^2 \rangle = C_2$

skewness: $S = \langle (\delta N)^3 \rangle / \sigma^3 = C_3 / C_2^{3/2}$

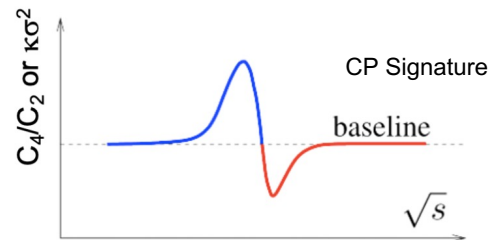
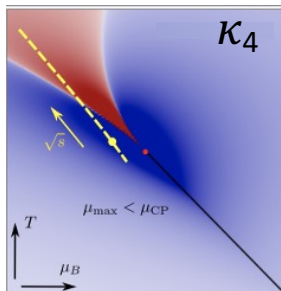
kurtosis: $\kappa = \langle (\delta N)^4 \rangle / \sigma^4 - 3 = C_4 / C_2^2$

Moments, cumulants and susceptibilities:

2nd order: $\sigma^2 / M \equiv C_2 / C_1 = \chi_2 / \chi_1$

3rd order: $S \sigma \equiv C_3 / C_2 = \chi_3 / \chi_2$

4th order: $\kappa \sigma^2 \equiv C_4 / C_2 = \chi_4 / \chi_2$

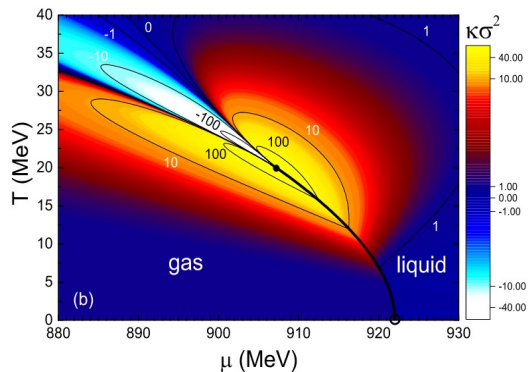


M. A. Stephanov, PRL 102 (2009) 032301

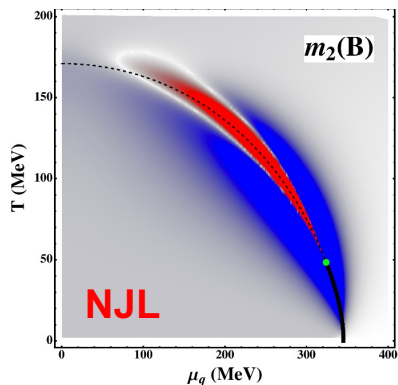
M. A. Stephanov, Phys. Rev. Lett. 102, 032301 (2009); 107, 052301 (2011). M. Asakawa, S. Ejiri and M. Kitazawa, Phys. Rev. Lett. 103, 262301 (2009). Cheng et al, PRD (2009) 074505. F. Karsch and K. Redlich, PLB 695, 136 (2011). B. Friman et al., EPJC 71 (2011) 1694. S. Gupta, et al., Science, 332, 1525(2012). A. Bazavov et al., PRL109, 192302(12) // S. Borsanyi et al., PRL111, 062005(13)

Critical Contributions to Fourth-order Fluctuations ($\kappa\sigma^2$)

van der Waals (VDW)



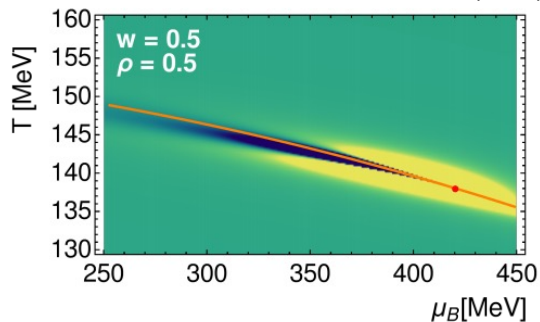
Vovchenko et al., PRC92, 054901 (2015)



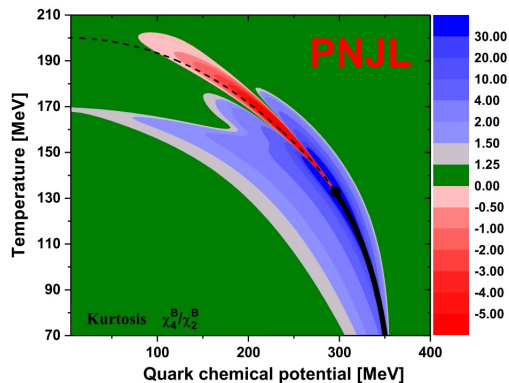
邓建等, PRD93, 034037 (2016)

3D Ising Mapping

D. Mroczek et al, PRC 103, 034901 (2021)



Symmetry and Universality

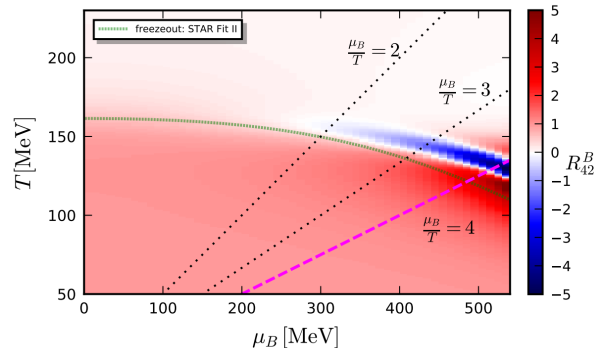


邵国运等, EPJC 78, 138 (2018)

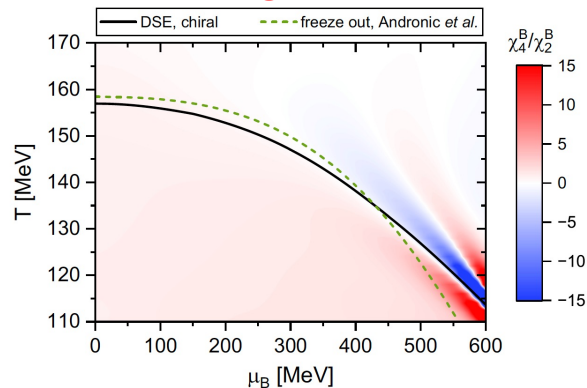
李志宾、许坤、王昕扬、黄梅, EPJC 79, 245 (2019).

FRG

付伟杰等 PRD 104 (2021) 9, 094047



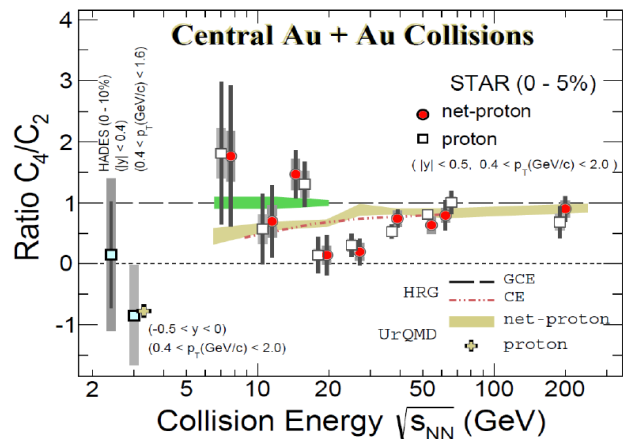
DSE



陆易、高飞、刘玉鑫等, PRD 113 (2026), 054019

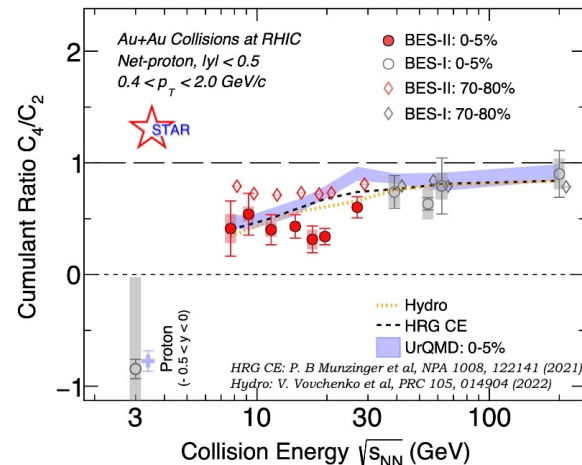
Net-Proton Fluctuations in BES-I and BES-II

BES-I (3 – 200 GeV)



$\sqrt{s_{NN}}$ (GeV)	Events BES-I (10 ⁶)	Events BES-II (10 ⁶)
7.7	3	45
9.2	-	78
11.5	7	110
14.5	20	178
17.3	-	116
19.6	15	270
27	30	220

BES-II (7.7 – 27 GeV)

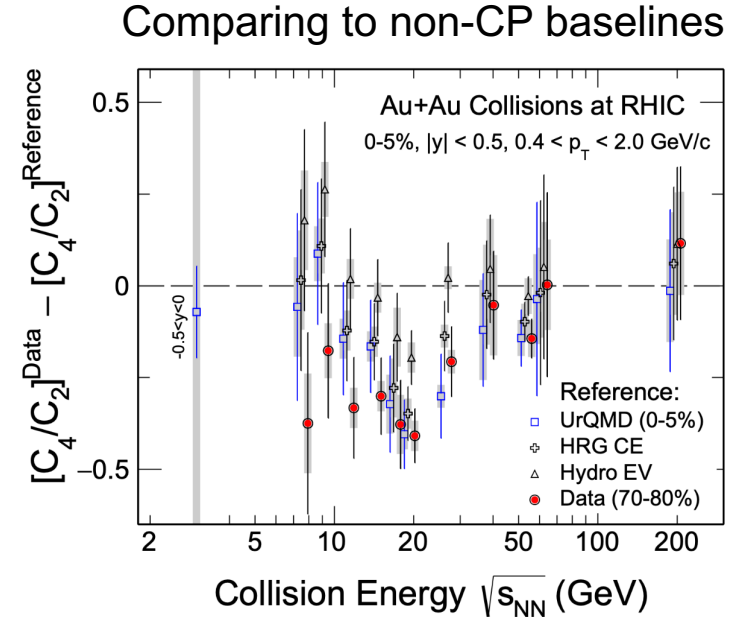
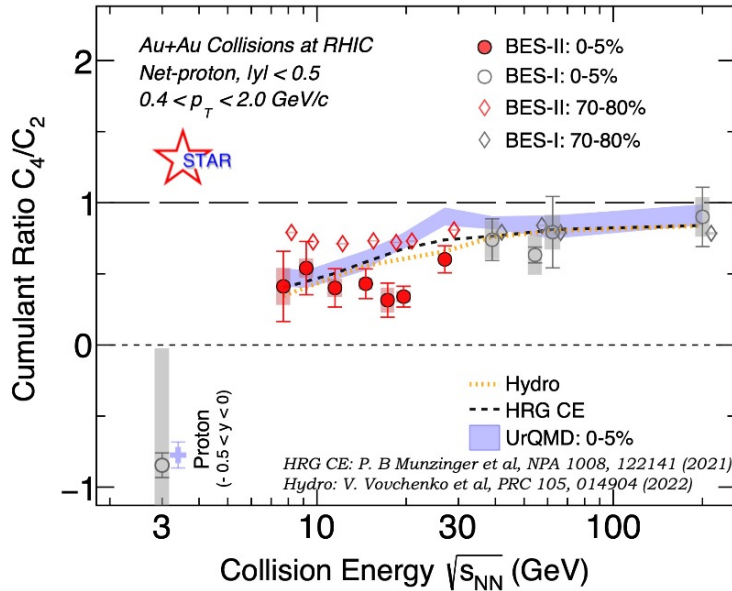


STAR: Phys. Rev. Lett. 128, 202303 (2022)
 Phys. Rev. Lett. 126, 092301 (2021)
 Phys. Rev. C 104, 024902 (2021)
 Phys. Rev. C 107, 024908 (2023)
 HADES: Phys. Rev. C 102, 024914 (2020)

BES-II and BES-I results are consistent
 BES-II : Better statistical precision
 Larger Acceptance ($|\eta| < 1.6$)
 Better control on systematics

STAR : Phys. Rev. Lett. 135, 142301 (2025)
[Editor Suggestion, Featured in Physics]

Net-Proton Fluctuations in BES-II (7.7 – 27 GeV)



STAR : Phys. Rev. Lett. 135, 142301 (2025)
[Editor Suggestion, Featured in Physics]

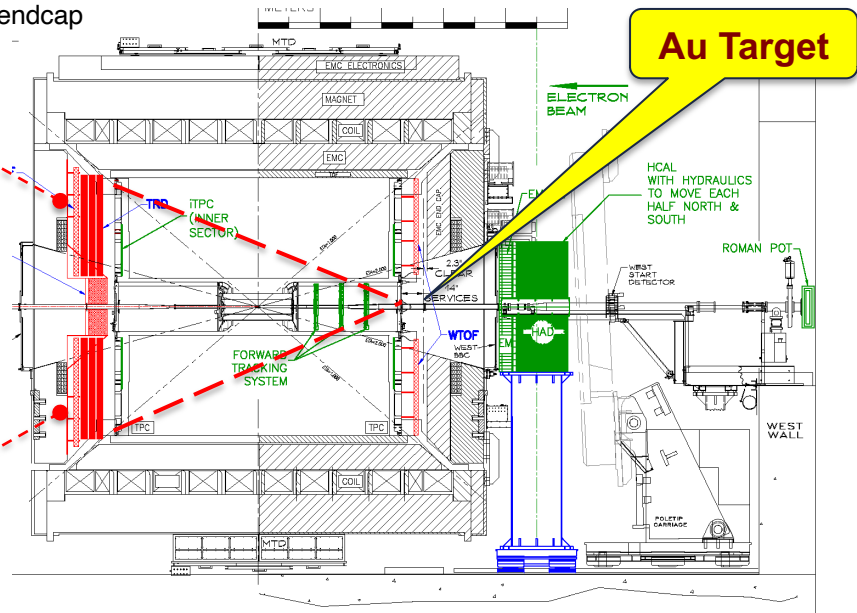
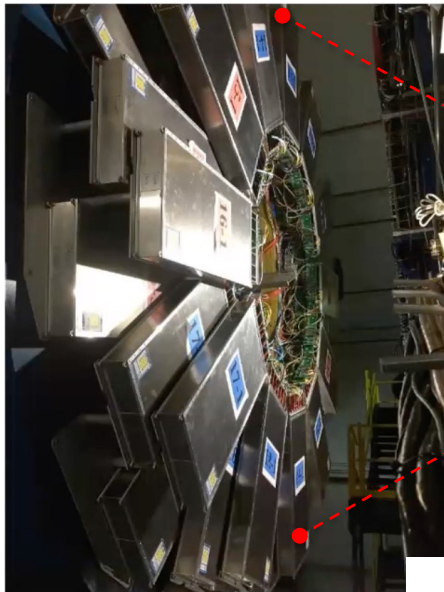
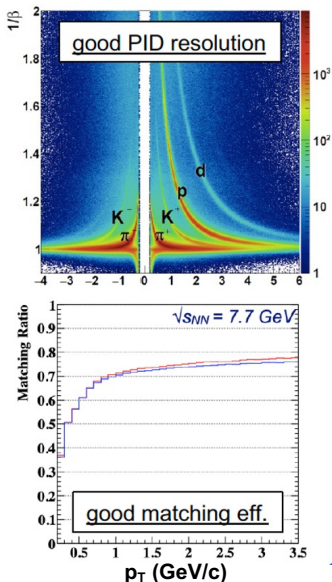
STAR: PRL126, 92301(2021); PRC104, 024902 (2021)
PRL128, 202303(2022); PRC107, 024908 (2023)
HADES: PRC102, 024914(2020)

- 2-5 σ deviations to non-CP baselines at 19.6 GeV
- Data from 3 – 7.7 GeV are crucial to establish the oscillation pattern of possible CP signal.

STAR Fixed-Target Setup

eTOF (2019+)

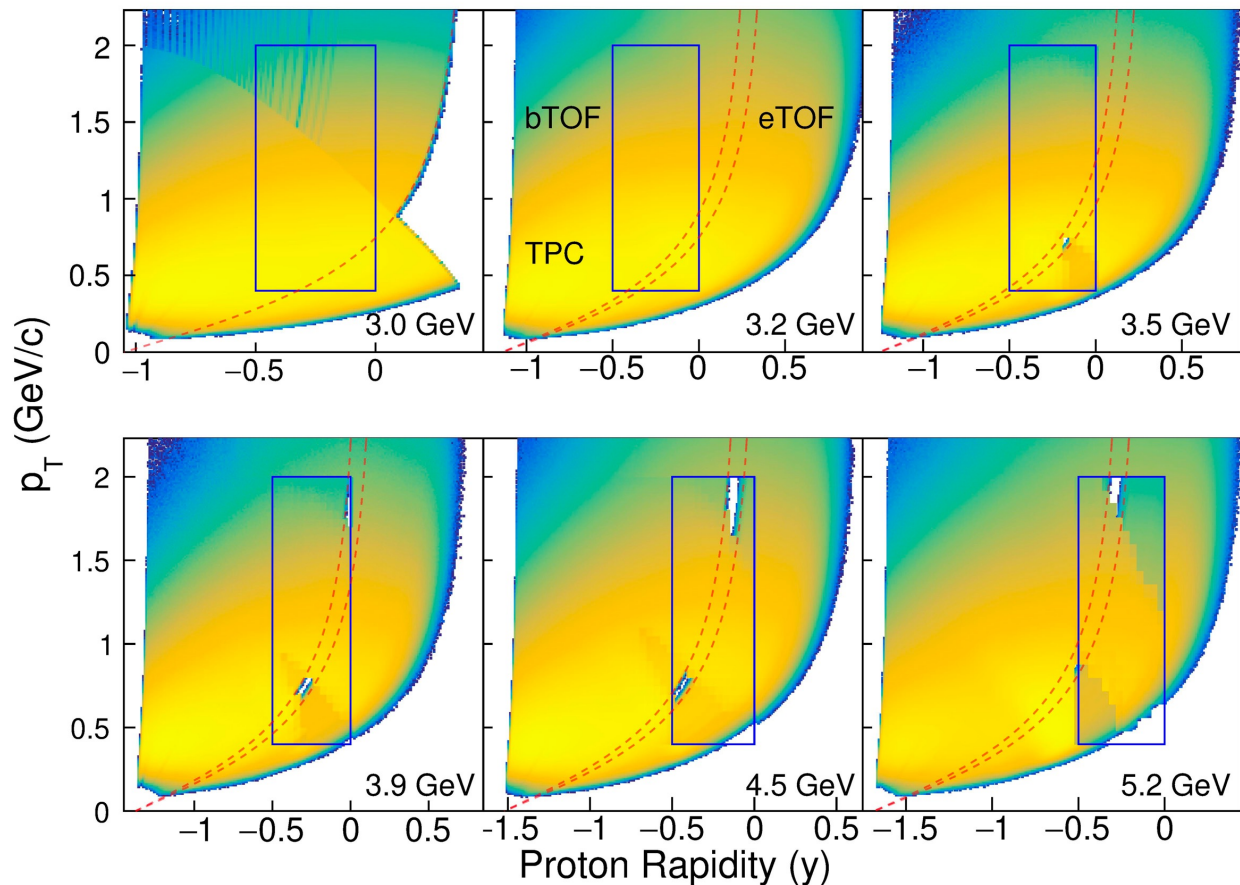
10% of the CBM TOF modules installed at STAR endcap



CBM participates in RHIC BES-II in 2019 – 2021:

- Fixed-target Au+Au collisions : $\sqrt{s_{NN}} = 3 - 7.7 \text{ GeV}$ ($750 \geq \mu_B \geq 420 \text{ MeV}$)
- Study the properties of QCD matter at high baryon density region

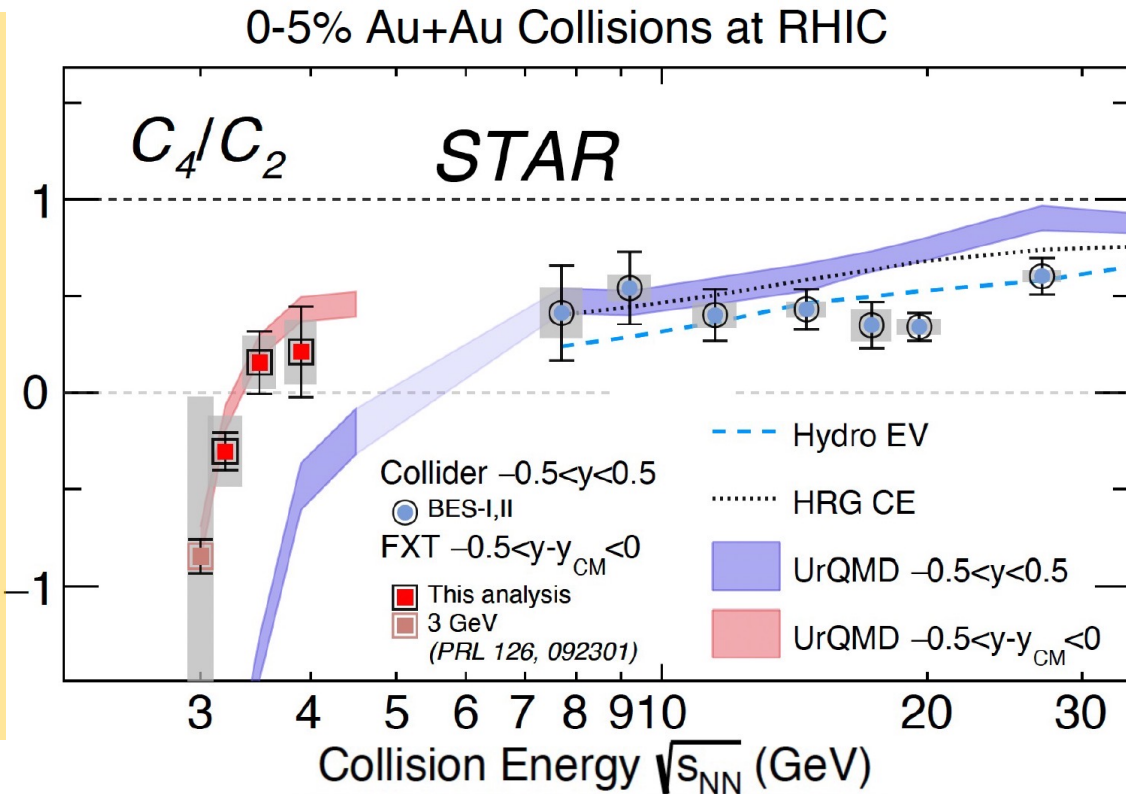
Acceptance of Protons in FXT Program



- 1) TPC, iTPC, bTOF, eTOF are used in the FXT program;
- 2) Lossing mid-y acceptance for energy higher than 4.5 GeV!
- 3) Analysis is focused on the mid-y region;

(Net-) Protons (C_4/C_2) from FXT Program

- 1) Collider mode: $|y| < 0.5$
FXT mode: $-0.5 < y < 0.0$
- 2) 3.0, 3.2, 3.5 and 3.9 GeV data were shown at QM2025
- 3) Data and UrQMD calculations are consistent for the FXT results. However, at the lowest energies, values of C_4/C_2 become negative: Centrality resolution issue



STAR 3 GeV data (collected in 2018):
Phys. Rev. Lett. 128, 202303 (2022)
Phys. Rev. C107, 024908 (2023)

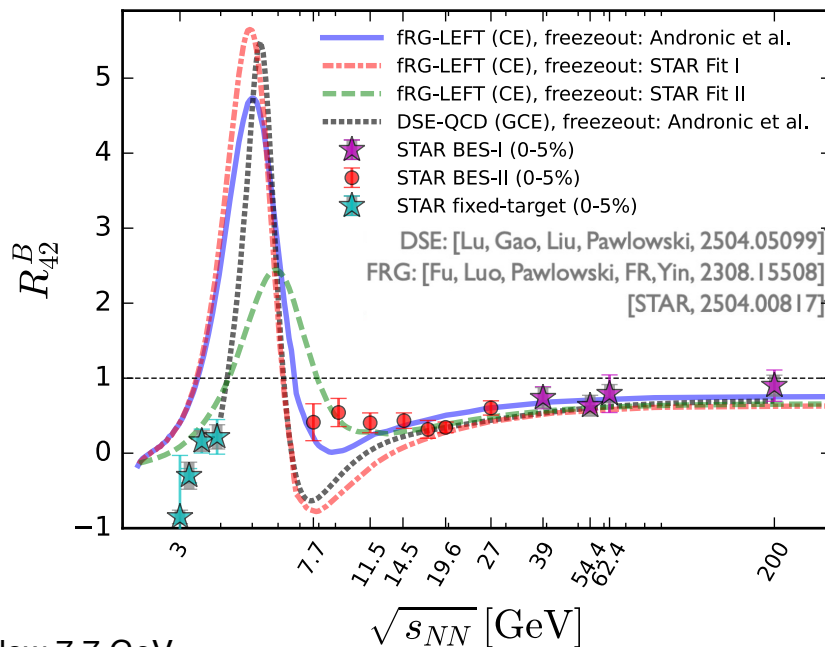
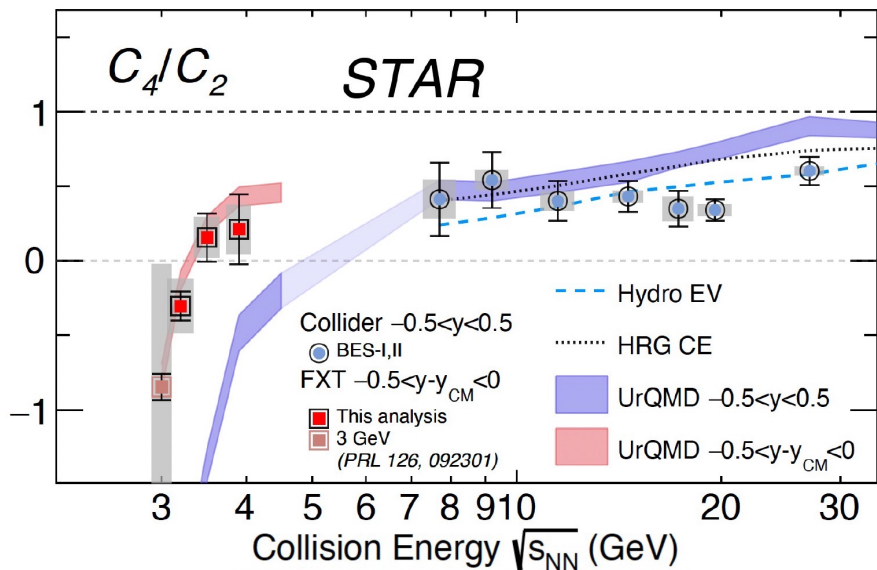
STAR, QM2025

Proton Number Fluctuations (C_4/C_2) from STAR FXT Program

STAR, QM2025

Fabian @ QM2025, plenary
Wei-jie Fu @ QPT2025, plenary

0-5% Au+Au Collisions at RHIC

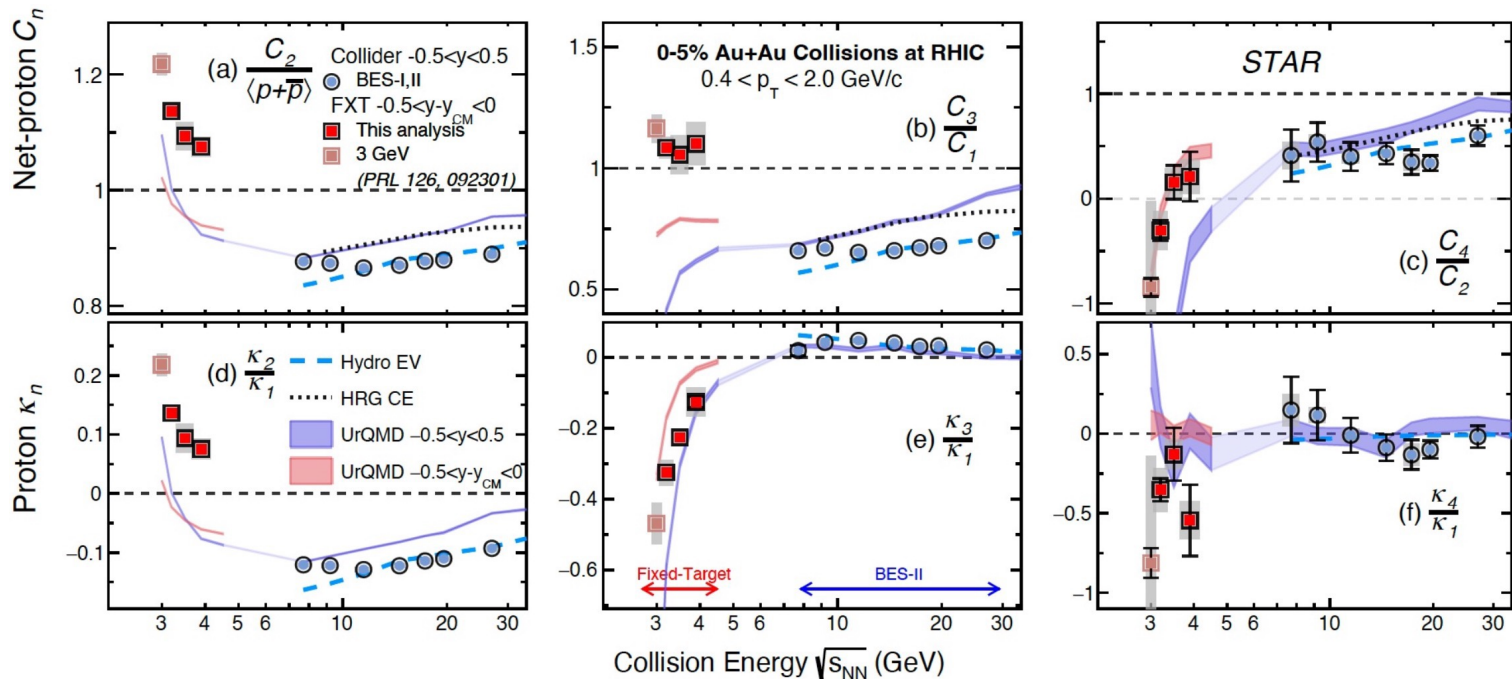


- Search for a maximum in data relative to non-CP baseline below 7.7 GeV.
- The STAR FXT results from 3, 3.2, 3.5 and 3.9 GeV are consistent with UrQMD calculations.
- Analysis of high statistics 3, 4.5 and 5.2 GeV from Run21 are ongoing.
In addition, STAR took more data of 4.2, 4.5 and 5.2 GeV during Jan. 22-28, 2026.

Non-equilibrium effects ?

See Wei-jie Fu's talk

How about Lower Order Fluctuations ?



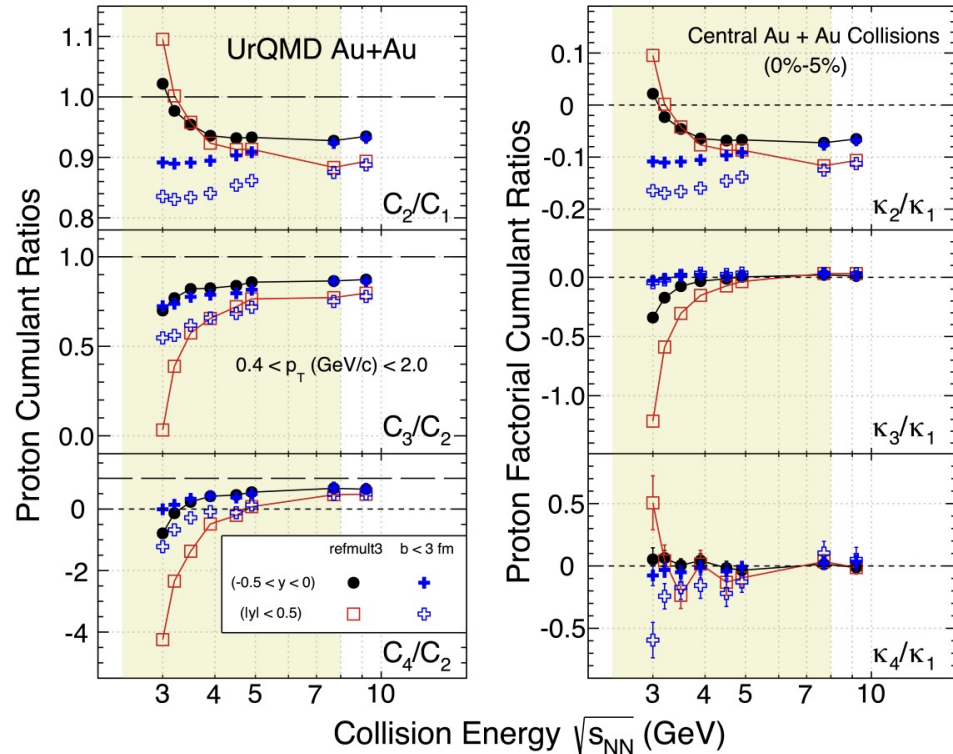
STAR, QM2025

At 3.0 – 3.9 GeV central collisions (0-5%, $-0.5 < y < 0$):

- 1) Change of trends and increase above unity for 2nd and 3rd order cumulants;
- 2) Deviations from UrQMD model calculations is seen

UrQMD Simulations – I : Effects of Volume Fluctuations on Cumulants

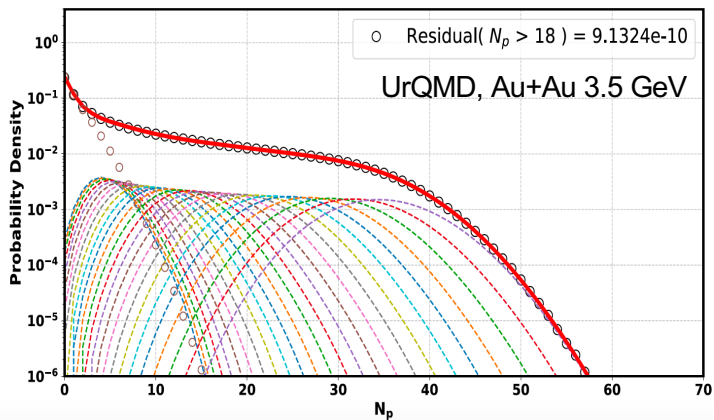
Centrality Bin With Correction (CBWC) is applied.



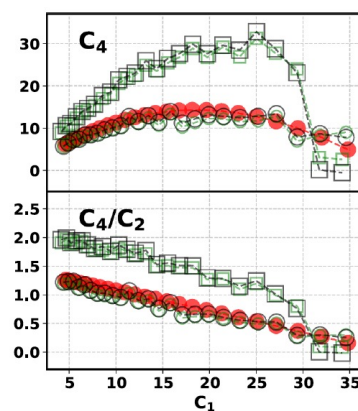
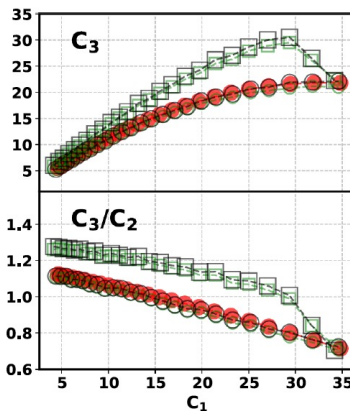
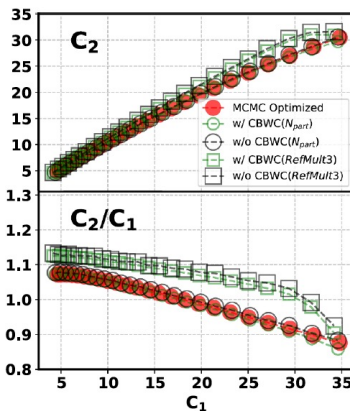
- 1) Initial Volume Fluctuation:** The collision centrality is determined by charged particle multiplicities. Even in the same centrality bin, the collision volume varies event by event;
- 2) Simulations indicate that effects of the initial volume fluctuations becomes stronger at lower collision energies, which leads to the observed rapidly increase or decrease of κ_2/κ_1 and κ_3/κ_1 ;
- 3) At lower energies, as the centrality resolution becomes very poorer, initial volume fluctuations become more significant

X. Zhang, Y. Zhang, XFL, N. Xu, Chin. Phys. C 50, 011003 (2026)
B. Friman, V. Koch, arXiv : 2511.11869

A New Approach



UrQMD, Au+Au 3.5 GeV



Centrality-Independent Genuine cumulant Analysis Framework (CIGAR)

Zhaohui Wang(汪朝辉), XFL, Phys. Lett. B 871, 139984 (2025)

Edgeworth expansion :

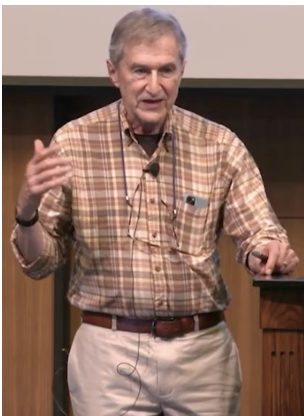
$$\sigma p(\sigma x) = Z(x) \left\{ 1 + \sum_{s=1}^{\infty} \left[\sigma^s \sum_{\{k_m\}} H e_{s+2r}(x) \prod_{m=1}^s \frac{1}{k_m!} \left(\frac{S_m + 2}{m + 2} \right)^{k_m} \right] \right\}$$

where Chebyshev-Hermite polynomials H_{e_n} is defined as: $H_{e_n}(x) = (-1)^n \exp(x^2/2) \frac{d^n}{dx^n} \exp(-x^2/2)$

- Using Edgeworth Expansion to reconstruct the proton number distribution with cumulant as parameters
- New Method (CIGAR) has been verified to effectively suppress the volume fluctuations.

Stay Tune !!

奇异双重子态的预言和寻找



Perhaps a Stable Dihyperon*

R. L. Jaffe†

Stanford Linear Accelerator Center, Stanford University, Stanford, California 94305, and Department of Physics and Laboratory of Nuclear Science, ‡ Massachusetts Institute of Technology, Cambridge, Massachusetts 02139

(Received 1 November 1976)

In the quark bag model, the same gluon-exchange forces which make the proton lighter than the $\Delta(1236)$ bind six quarks to form a stable, flavor-singlet (with strangeness of -2) $J^P=0^+$ dihyperon (H) at 2150 MeV. Another isosinglet dihyperon (H^*) with $J^P=1^+$ at 2335 MeV should appear as a bump in $\Lambda\Lambda$ invariant-mass plots. Production and decay systematics of the H are discussed.

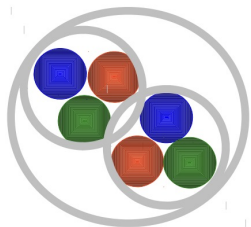
R. L. Jaffe, Phys. Rev. Lett. 38, 617 (1977)

1977年, Jaffe基于夸克袋模型预言 dihyperon (H): deeply bound $\Lambda\Lambda$ system (uuddss), ($J=0, S=-2, m\sim 2150$ MeV), 81 MeV below $\Lambda\Lambda$ 阈值(2231)。

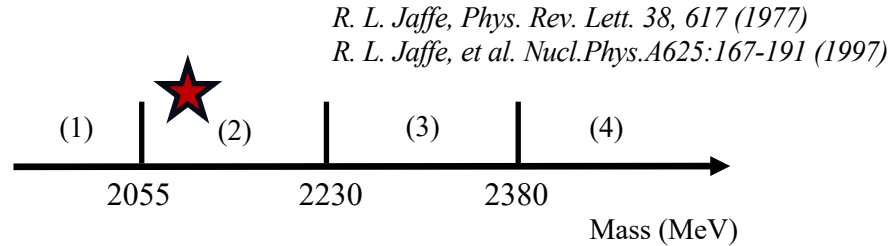
之后的40多年间, 理论及实验寻找H-dibaryon的热潮从未间断。多个实验组尝试寻找, 如日本JPARC E42实验, 美国RHIC-STAR/AGS实验, 欧洲LHC-ALICE/SPS实验等等。

综述: H. Clement, Progress in Particle and Nuclear Physics 93, 195 (2017)

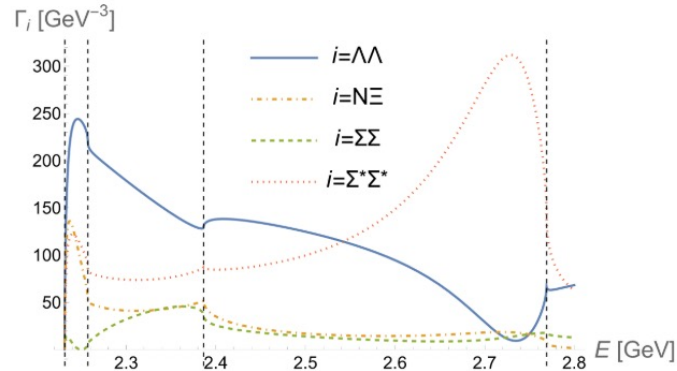
(1946-)
2023年在MIT,
“50 years of QCD”



To bracket the possibilities we consider four possible mass ranges for the H : (1) $M(H) < 2055$ MeV; (2) $2055 < M(H) < 2230$; (3) $2230 < M(H) < 2380$; and (4) $M(H) > 2380$ MeV. In case (4), the H is above all two-baryon thresholds to which it couples strongly. It represents, therefore, a repulsive interaction in all two-baryon channels to which it couples. It is improbable that the bag-mass calculation is so much in error that case (4) applies. In case (3), the H is a $\Sigma\Sigma$ bound state [or if $M(H) < 2260$ MeV, also an $N\Xi$ bound state] decaying strongly into $\Lambda\Lambda$. It would appear as a bump in $\Lambda\Lambda$ invariant-mass plots (e.g., in $\Xi p \rightarrow \Lambda\Lambda$). Were the H very light [case (1)], it would be stable except against double weak decay ($H \rightarrow nn$). Cases (1) and (3) are not favored by the bag-mass estimates but must be retained as possibilities in light of the crude nature of the model.



The $\Lambda\Lambda$, $N\Xi$, $\Sigma\Sigma$, and $\Sigma^*\Sigma^*$ systems are coupled, and their interplay governs the structure of possible dibaryon states and near-threshold phenomena.



胡陶然, 郭奉坤, arXiv:2601.14922

Strangeness -3 $N\Omega$ and Strangeness -6 $\Omega\Omega$ Dibaryons

$N\Omega$

MIT bag model: in 1987 MIT bag model predicted $N\Omega$ dibaryon which is stable with respect to strong decay

Quark-Delocalization color-screening model (QDCSM) and chiral quark model (ChQM)

VOLUME 59, NUMBER 6

PHYSICAL REVIEW LETTERS

10 AUGUST 1987

Strangeness -3 Dibaryons

T. Goldman, K. Maltman,^(a) and G. J. Stephenson, Jr.
Los Alamos National Laboratory, Los Alamos, New Mexico 87545

K. E. Schmidt
Chemistry Department and Courant Institute of Mathematical Sciences, New York University, New York, New York 10012

and

Fan Wang^(b)
University of California at Los Angeles, Los Angeles, California 90024
(Received 12 May 1987)

We demonstrate, using two different quark models of hadrons, that there should be isodoublets of dibaryons with strangeness -3 and $J=1,2$, which are stable with respect to strong decay.

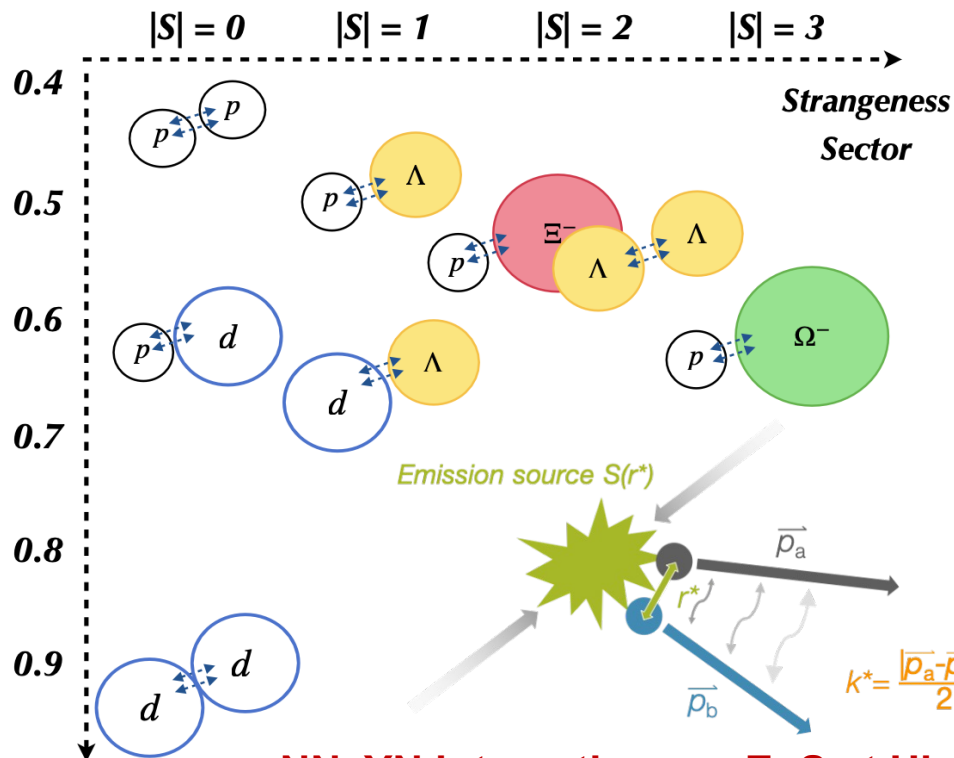
PACS numbers: 14.20.Pt, 12.40.Aa

H. Pang, J. Ping, F. Wang, T. Goldman, E. Zhao, PRC-69-065207 (2004);
H. Pang, J. Ping, L. Chen, F. Wang, T. Goldman, PRC-70-035201 (2004);
M. Chen, H. Huang, J. Ping, F. Wang, PRC-83-015202 (2011);
H. Huang, J. Ping, F. Wang, PRC-92-065202 (2015)

$\Omega\Omega$

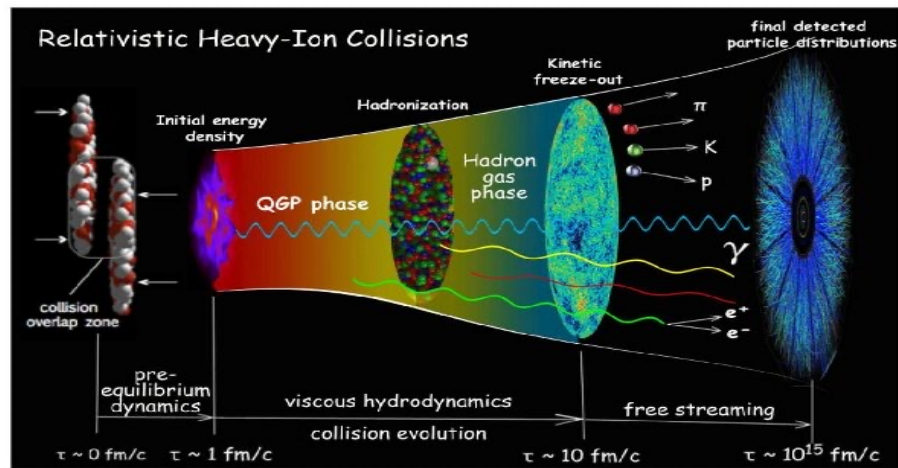
Binding energy 100 MeV, lifetime 2 times of free Omega's
More likely a six-quark particle with large binding energy and short relative distance (RMS=0.84 fm) between two Omegas
Z. Y. Zhang et al., NPA625, 59 (1997); PRC61, 065204 (2000)

Femtoscscopy correlations : NN and YN Interactions



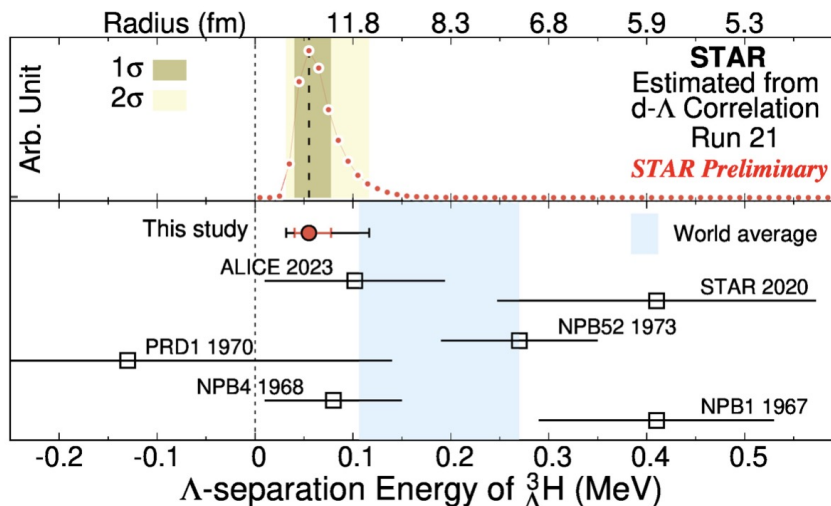
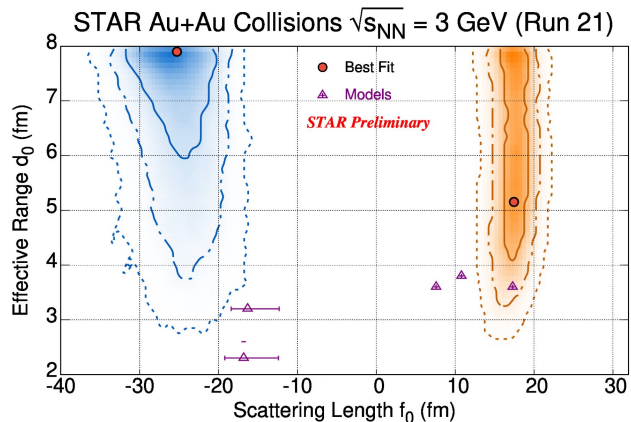
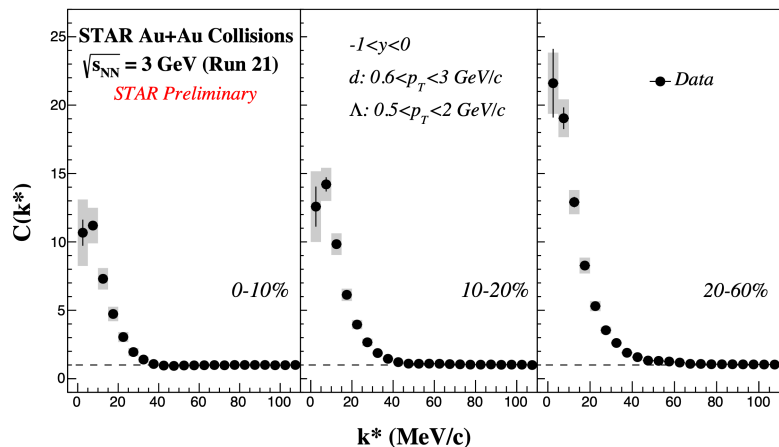
NN, YN interactions : EoS at High Baryon Density

- 1) Light Nuclei ($A \leq 3$) – Lambda at Au+Au 3 GeV
- 2) Search for strange Di-baryon in isobar collisions 200 GeV



Heavy-ion collisions produce a large number of nucleons and hyperons, making it an ideal setting for studying the properties of strong interactions.

$d - \Lambda$ Correlation functions at 3.0 GeV (Run21)

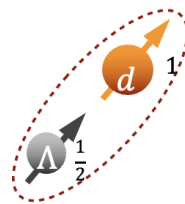


- 1) The most accurate results of binding energy and radius of ${}^3\Lambda\text{H}$.
Open a new way to constrain ${}^3\Lambda\text{H}$ properties.

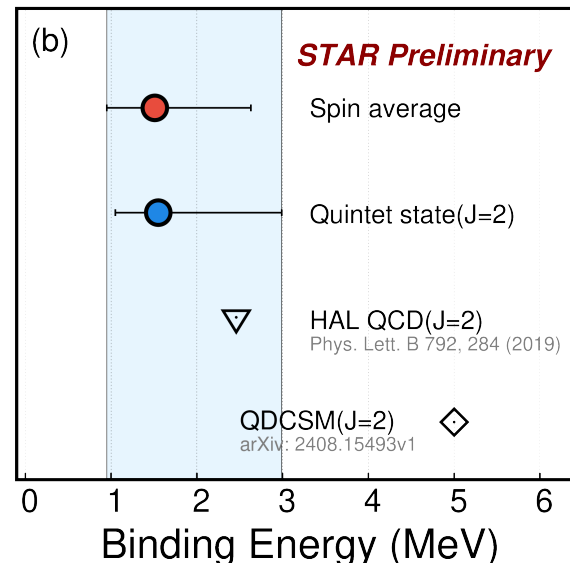
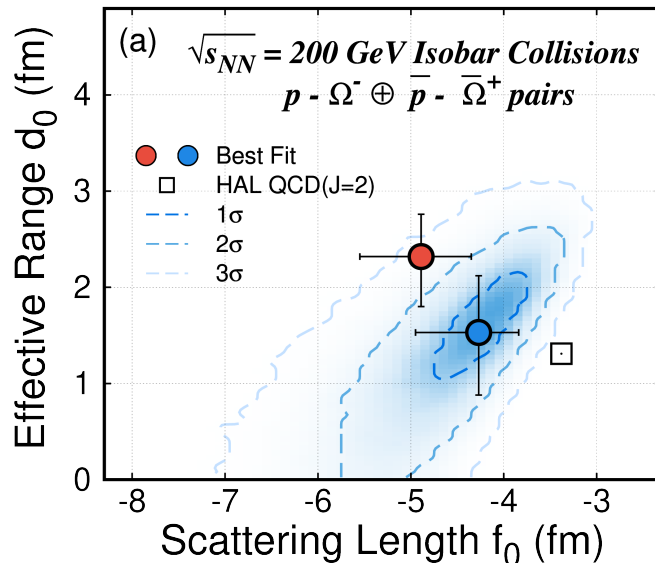
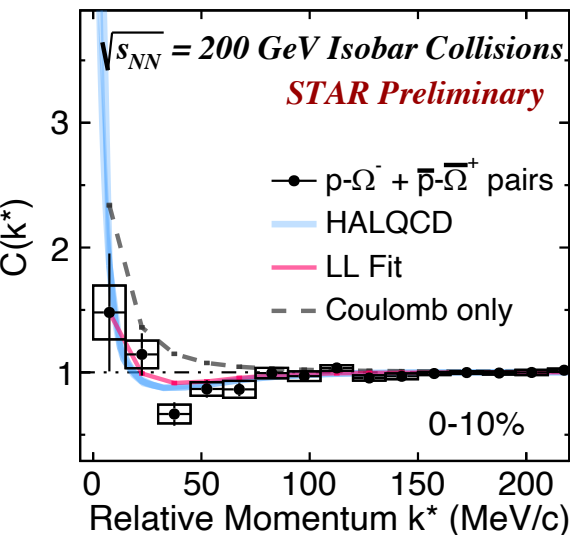
$${}^3\Lambda\text{H } B_\Lambda = 0.06^{+0.06}_{-0.03} \text{ (MeV) @ 95\% CL}$$

Extracted physical parameters:
 $f_0 \text{ (D)} = -25.3 \pm 3.3 \text{ fm}$
 $f_0 \text{ (Q)} = 17.5 \pm 1.6 \text{ fm}$

Xialei Jiang (STAR), QM2025
 STAR, arXiv: 2511.15493 (PRL in press)



p- Ω^- Correlation Function ($|S|=3$)



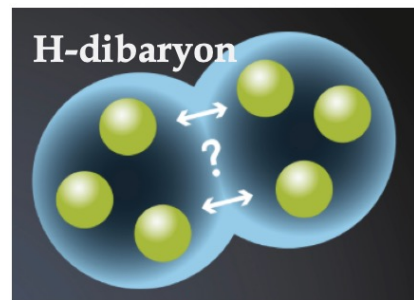
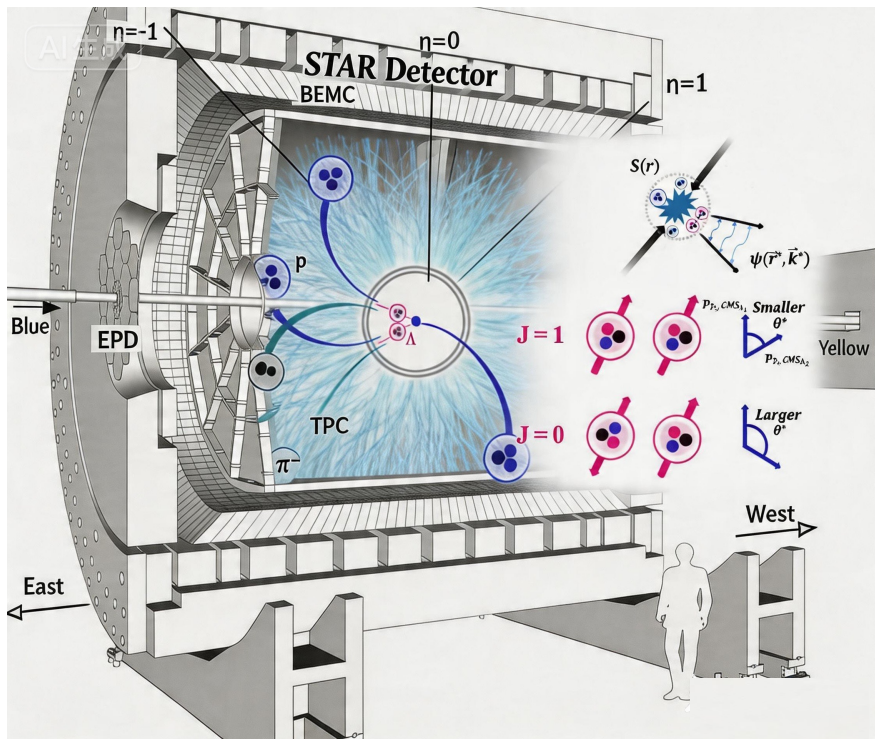
Kehao Zhang (STAR), QM2025

- I. First experimental constraints in heavy-ion collisions of strong interaction parameters in p- Ω^- pair
- II. Extracted **negative** f_0 ($|f_0| > 2d_0$) by Spin average method and Quintet method

□ **First experimental evidence of Strange Dibaryon**

	Spin ave.	Quintet	HAL QCD
f_0 (fm)	$-4.89^{+0.54}_{-0.66}$	$-4.27^{+0.43}_{-0.68}$	-3.38
d_0 (fm)	$2.32^{+0.44}_{-0.52}$	$1.53^{+0.54}_{-0.66}$	1.31
BE (MeV)	$1.51^{+1.12}_{-0.56}$	$1.55^{+1.44}_{-0.50}$	2.27

H-Dibaryon Searches at STAR



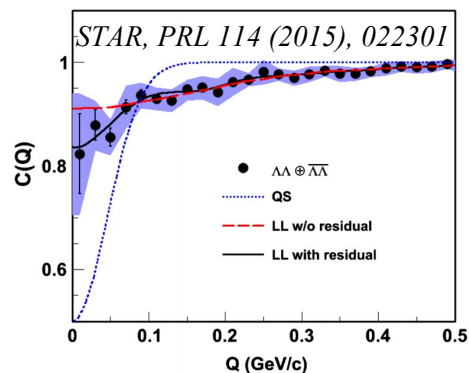
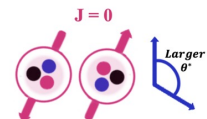
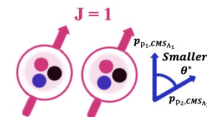
Possible channel:

$$\text{H-dibaryon} \Leftrightarrow p + \Xi$$

$$\text{H-dibaryon} \Leftrightarrow \Lambda + \Lambda$$

H → Λ + Λ

利用Λ自相似衰变行为挑选
衰变末态ΛΛ系统的自旋!



Analysis of high stat.
Iso-bar data finished
Paper proposal soon.

Summary and Outlook

- STAR reported precision measurement of (net-)proton fluctuations in Au+Au collisions from BES-II:

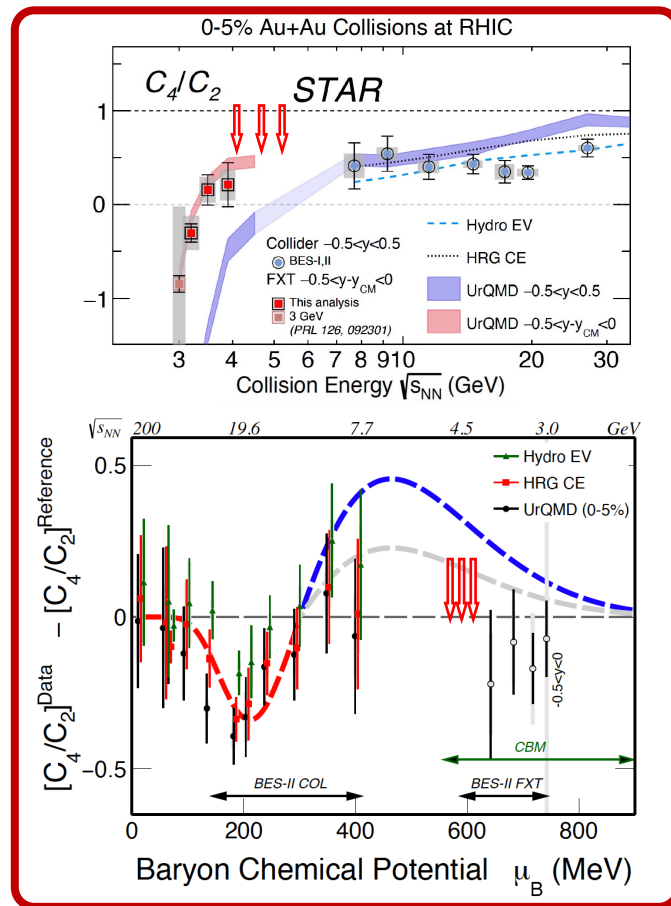
Collider energies: for net-proton C_4/C_2 in central collisions, deviations of 2-5 σ are observed w.r.t non-CP reference at 19.6 GeV

FXT energies: proton C_4/C_2 in central collisions are consistent with UrQMD calculations at 3.0 – 3.9 GeV while low orders changing trends and deviate from the transport model calculations

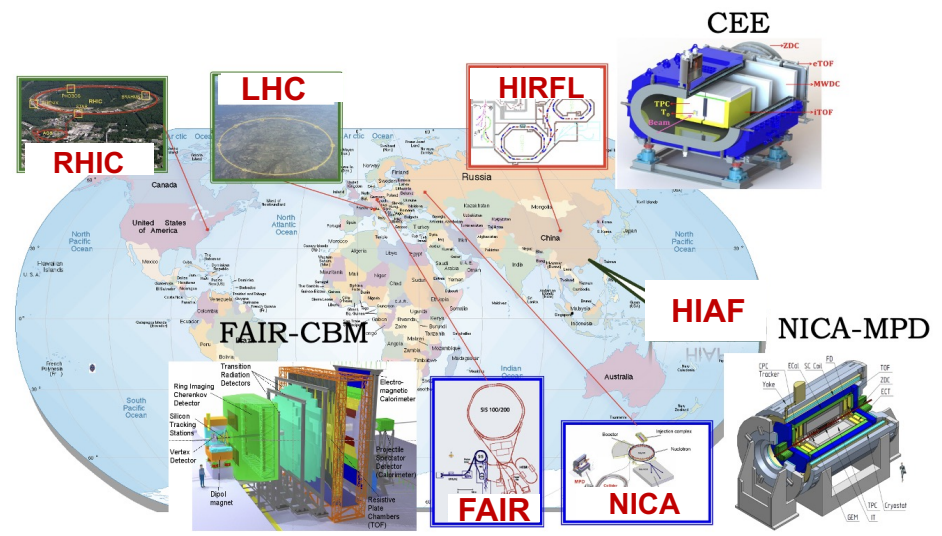
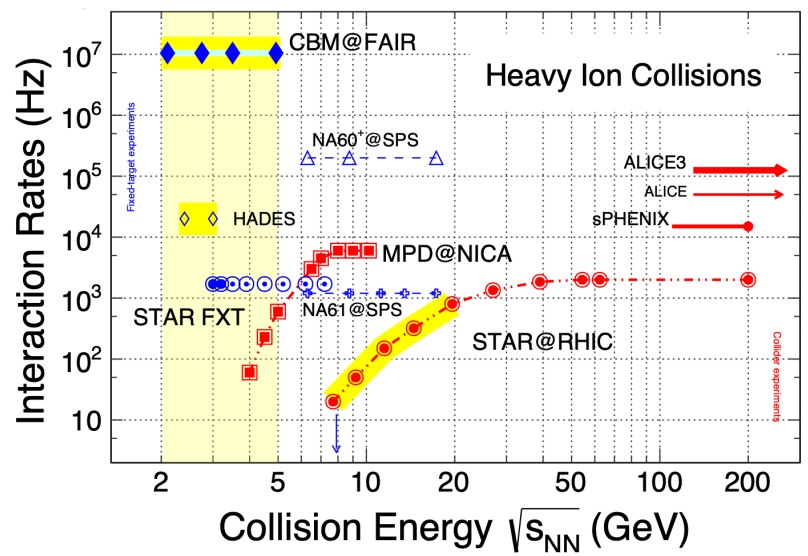
- STAR is working on a new analysis framework to determine the collision centrality, to suppress the initial volume fluctuation. Reanalyzing all data sets including 4.2, 4.5 and 5.2 GeV Au+Au collisions is under way;

- Modeling with dynamical effects including critical fluctuations and non-equilibrium effects are crucial to understand the experimental data

Both CEP and Strange Dibaryon Search are ongoing !!



Summary and Outlook



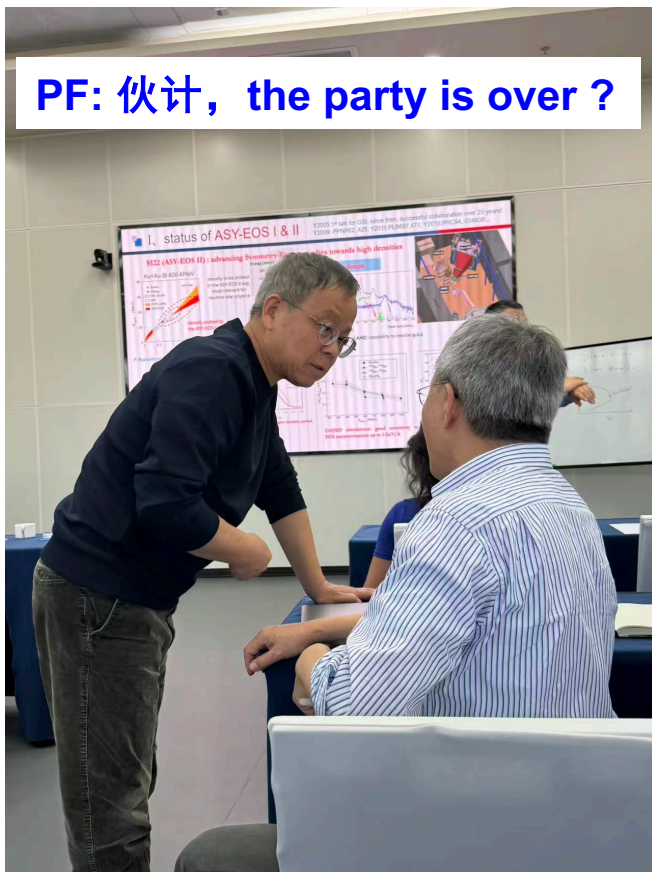
Future Low Energy Heavy-ion Experiments :

- FAIR-CBM (2.4 - 4.9 GeV, 2028 -)
- CSR-CEE (2.1 GeV, 2026-)
- HIAF-U/H-NS (2.1- 4.5 GeV, 203?)
- NICA-MPD (4 - 11 GeV, 2027 -)
- JPARC-HI (2-5 GeV, planning)

Key Physics at High Baryon Density :

- Critical point and phase boundary;
- Nuclear matter EOS at high baryon density;
- Y-N interactions, inner structure of compact stars.

PF: 伙计, the party is over ?



NU: Come to Huizhou, the party just begin !



International Conference on Nucleus-Nucleus Collisions (NN2027)

co-Chairs: Xiaofeng Luo¹ and Bing Guo²

¹Central China Normal University

²China Institute of Atomic Energy

Wuhan, China
Oct. 10 - 15, 2027



Many Thanks to:

members of STAR Collaboration and
Theory Colleagues

Thank you for your attention!