

# **LSTM Peak finding algorithm for Cluster Counting on Real Data and ML based Algorithm for Particle Identification**

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Meeting

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# ① Analysis of Real Test Beam Data

- ① Configuration Setup for channel 5
- ① Optimization of LSTM Model by Search Grid
- ① Peak Finding Algorithm

# ① ML based Algorithm on for Particle Identification(Ongoing)

- ① Simulation Parameters
- ① Number of Primary Cluster (MC)
  - I) Muon
  - ii) Kaon
- ① NN (CNN) Reconstruction Results
- ① Particle Identification

## Main Goal of the Presentation

**Task 1:** The first task is related to apply the best trained LSTM Model (Highest Auc value among all configurations) on 180 GeV muon real data in order to detect peaks in comparison with RTA Algorithm on the behalf of different selection cuts(0.2-0.3 0.4)

**Task 2:** The second task is related to particle identification of muon and kaon. That is why first we generated some simulated samples from 2-10 GeV momenta for Kaon. Then we apply best LSTM model to detect peaks in each waveform and then in the second step best CNN regression Model was applied to reconstruct the number of primary cluster on the behalf of detected peaks.

**Note:** I already assumed that all of you know about the two steps of cluster counting techniques. In the first step the best LSTM model find the peaks (Primary + Secondary electrons) in each waveform against the noise. While in the second step best CNN model estimate the number of primary clusters based on the detected peaks. That is why I showed the final results of Number of primary Clusters MC & CNN regression Model.

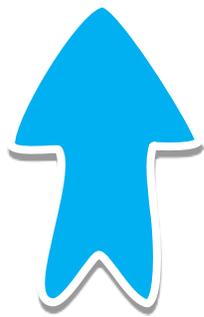
# LSTM Peak Finding Algorithm on Real Data

# Configuration Setup for the Real Test Beam data (180 GeV & Channel 5)

## Confiurations and setup information

### Test-Beam Setup Parameters for Channel 5

Particle	muon
Gas mixture	He 90% isobutane 10%
Cell size	1 cm
Momentum	180 GeV
Sampling rate	1.5 GHz
Angle	45°
Voltage at sense wire	1450 V



### Channel 5 Setup Configuration

DRS16 channels	HV channels	Tubes
0	0	1.0cm-20μm
1	1	1.0cm-20μm
2	2	1.0cm-20μm
3	3	1.0cm-20μm
4	4	1.0cm-20μm
5	5	1.0cm-20μm
6	12	1.5cm-20μm
7	13	1.5cm-20μm
8	14	1.5cm-20μm
9	15	1.5cm-20μm
10	-	-
11	-	-
12	-	-
13	-	-
14	-	Sipm Scintillator upstream
15	-	Sipm Scintillator downstream

Tubes setup with DRS

Tubes	Channels	HV channels	HV tag	
			1	2
OSC 1cm-20um	5,6,7,8	8,6,9,10	1450	
DRS16 1cm-20um	0,1,2,3,4,5	0,1,2,3,4,5	1450	
OSC 1.5cm-20um	1,2,3,4	16,17,18,19	1550	
DRS16 1.5cm-20um	6,7,8,9	12,13,14,15	1550	

HV tags with 90/10 gas mix runs

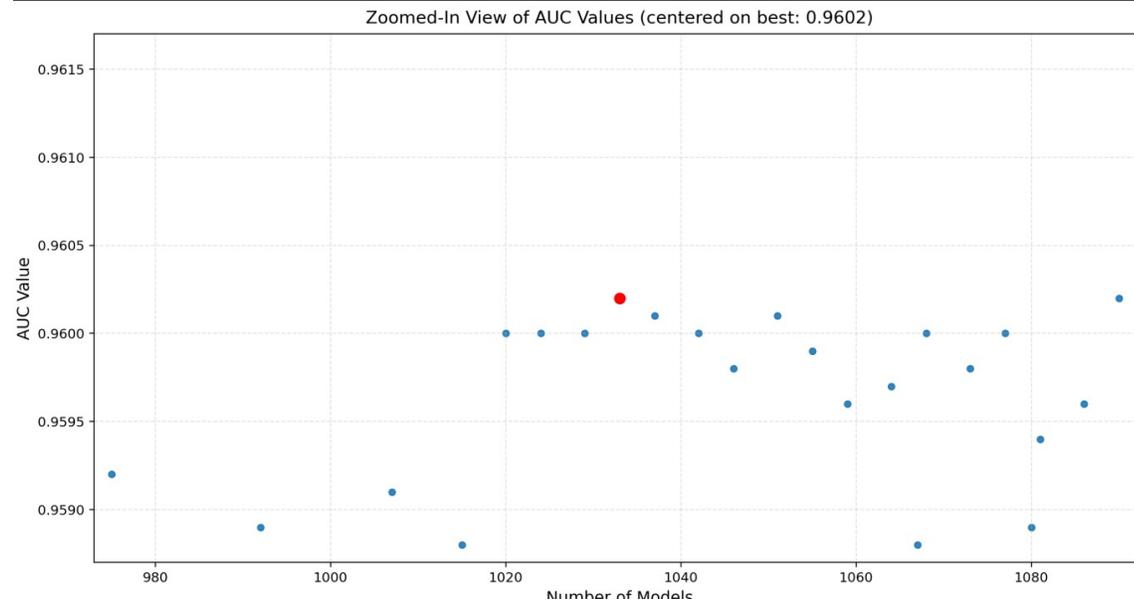
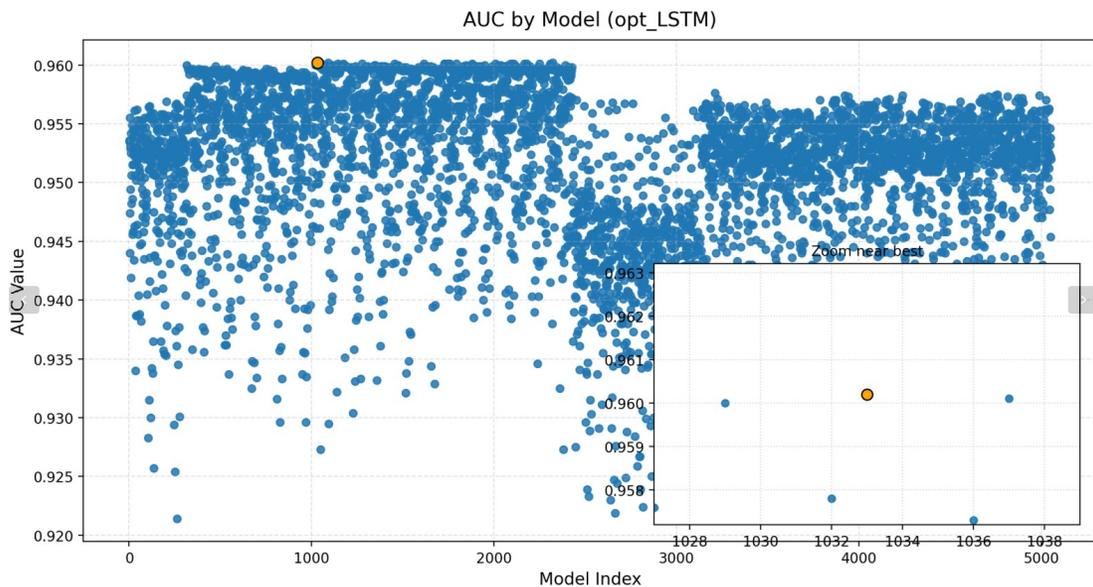
Oscilloscope	HV channels	Tubes
1	16	1.5cm-20μm
2	17	1.5cm-20μm
3	18	1.5cm-20μm
4	19	1.5cm-20μm
5	8	1.0cm-20μm
6	6	1.0cm-20μm
7	9	1.0cm-20μm
8	10	1.0cm-20μm

Tubes setup with OSC

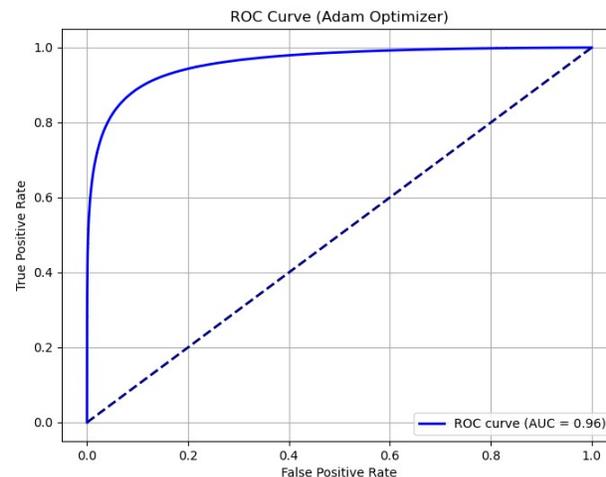
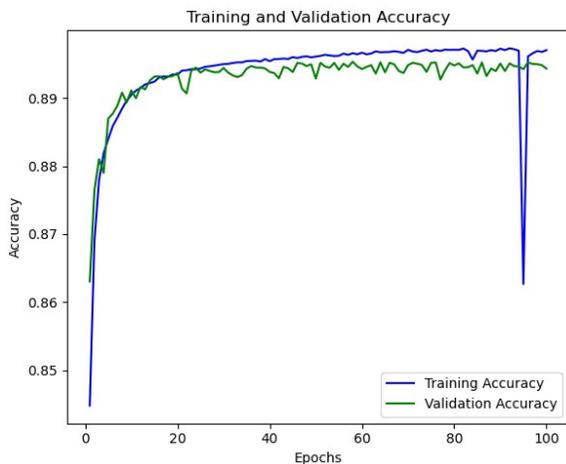
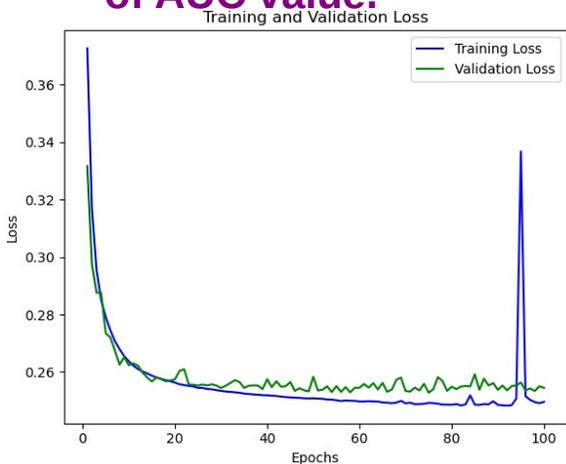
Tubes	Channels	HV channels	HV tag	
			1	2
OSC 1cm-20um	5,6,7,8	8,6,9,10	1450	1630
DRS16 1cm-20um	0,1,2,3,4,5	0,1,2,3,4,5	1450	1630
OSC 1.5cm-20um	1,2,3,4	16,17,18,19	1550	1730
DRS16 1.5cm-20um	6,7,8,9	12,13,14,15	1550	1730

HV tags with 85/15 gas mix runs

# Search Grid Optimization: BEST LSTM MODEL ON TUNE SIMULATED WAVEFORMS (50000)



- We used different sets of hyper-parameters like activation functions, optimizers etc to train LSTM peak finding model on 50000 tuned simulated waveforms on the simulation parameter (shown in the previous table) and then we select the best model with highest area under the curve value (0.96) among all configurations showed by red dot in left Zoom view of AUC value.

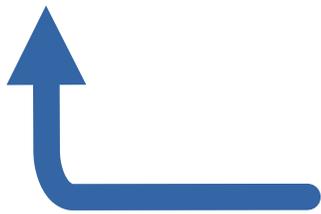
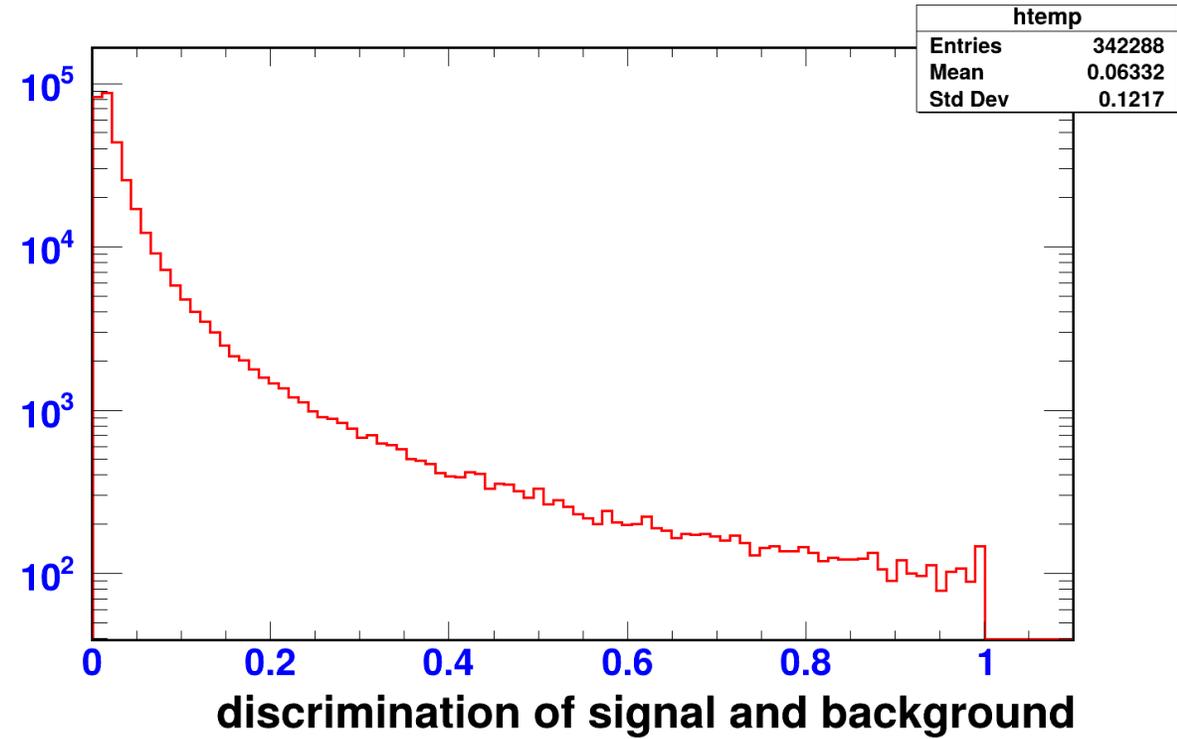
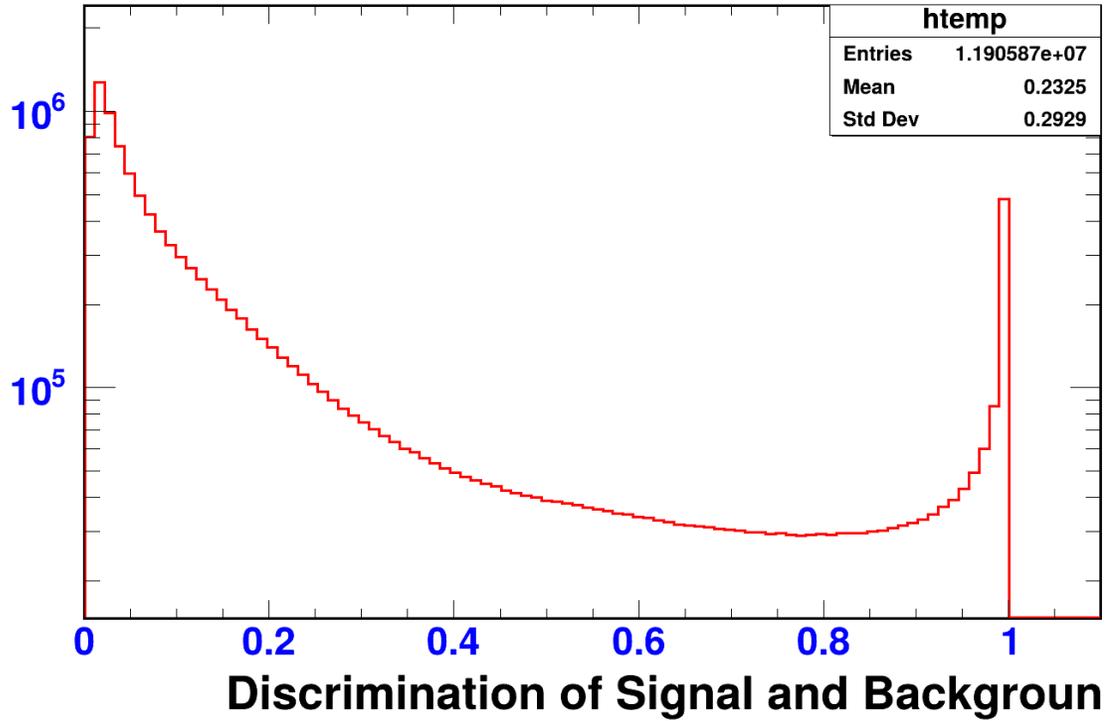


## Training configuration

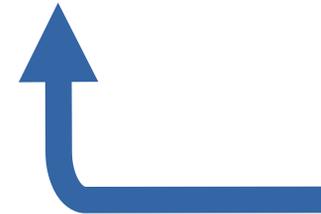
Optimizer	Adam
Network topology	[64, 32, 1]
Activation functions	gelu, sigmoid
Dropout rates	[0.0]
Batch size	16
Train/validation split	0.7 / 0.3
Number of epochs	100

## Hyperparameters of Best LSTM Model

# TUNE SIMULATED VS Data **Testing Part** by using Best LSTM Model (AUC =0.96)



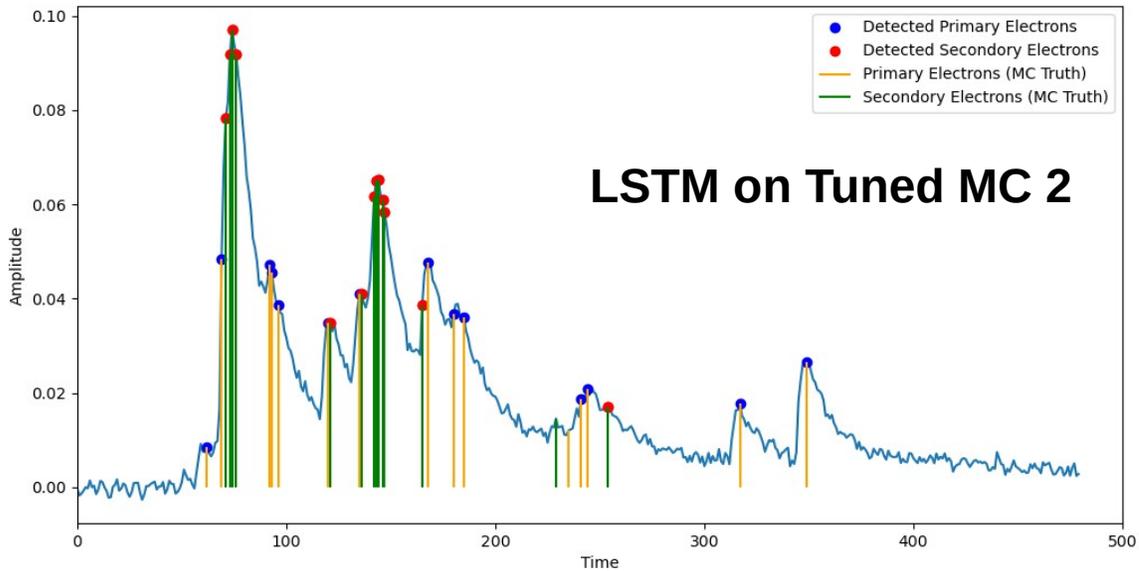
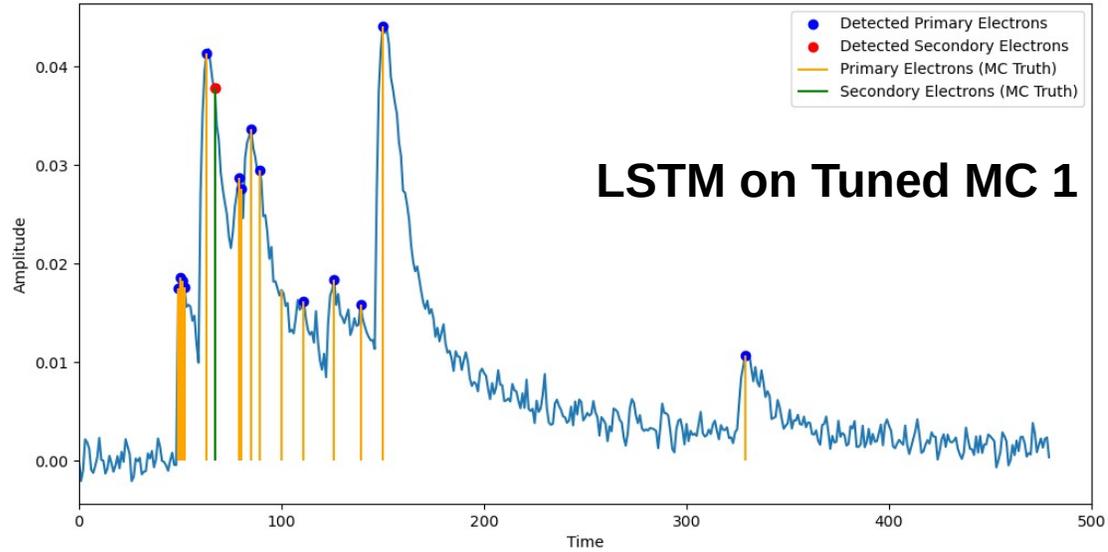
We applied best LSTM model on tune MC samples to discriminate signal and Background



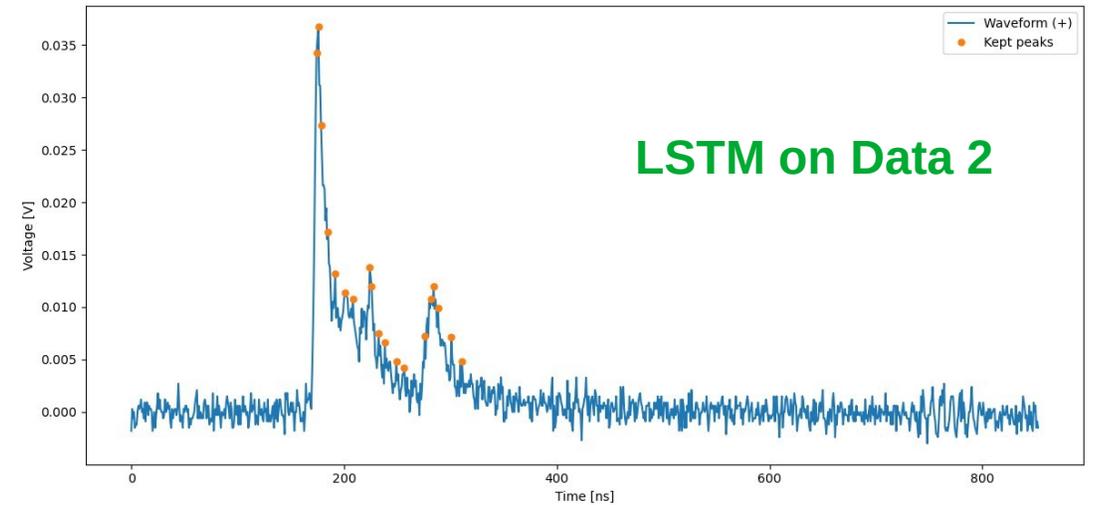
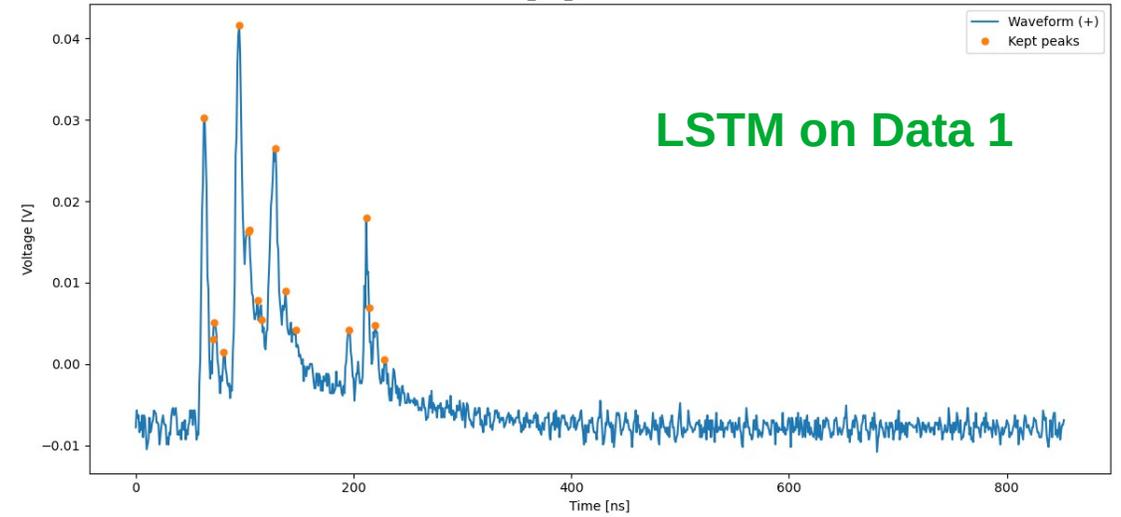
We applied best LSTM model on data (0.3) to discriminate signal and BackgroundData with Log Scale

# LSTM Peak Finding Waveforms Results of MC & Data In Comparison with RTA/Derivative Algorithm

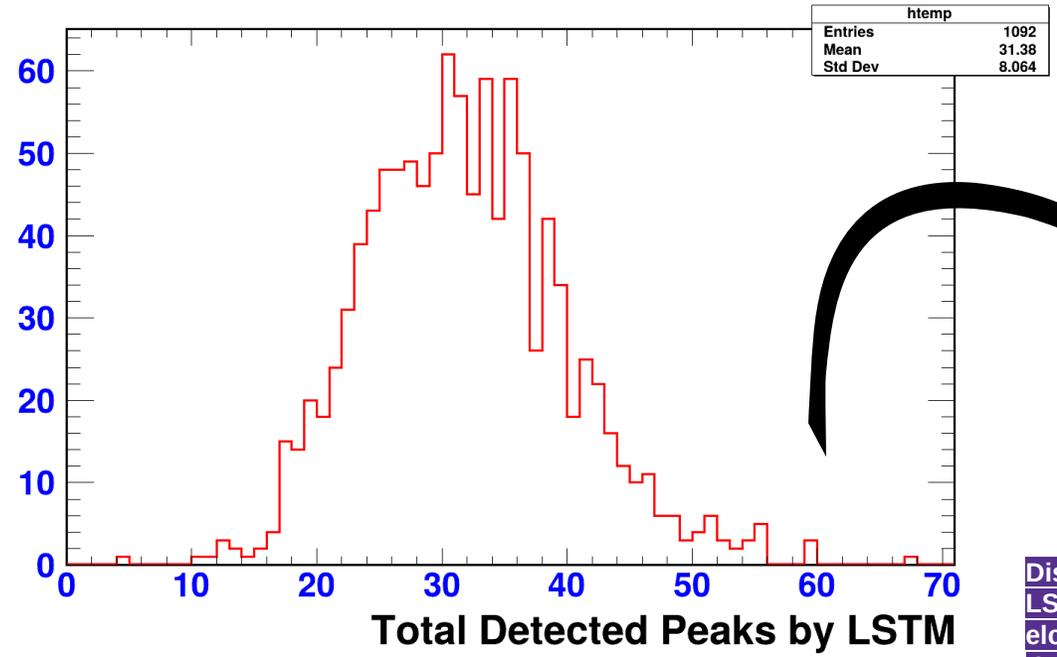
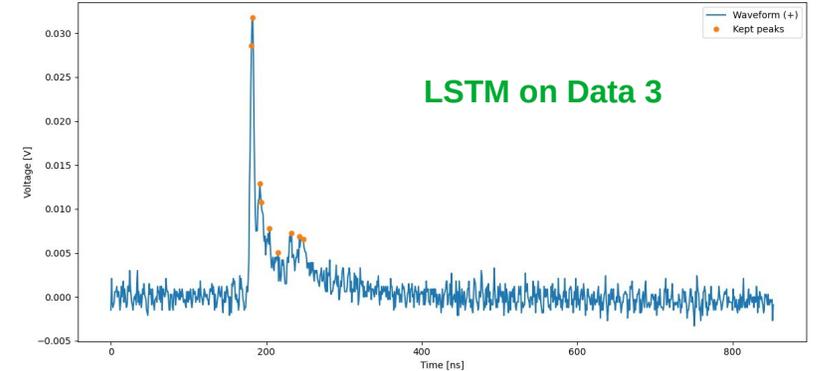
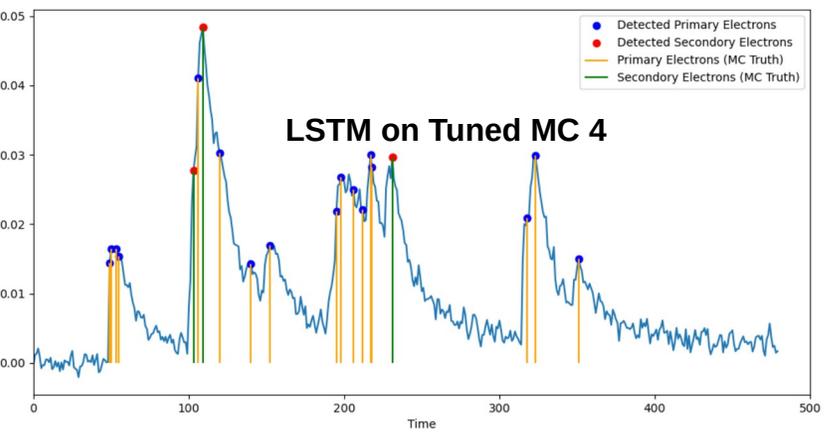
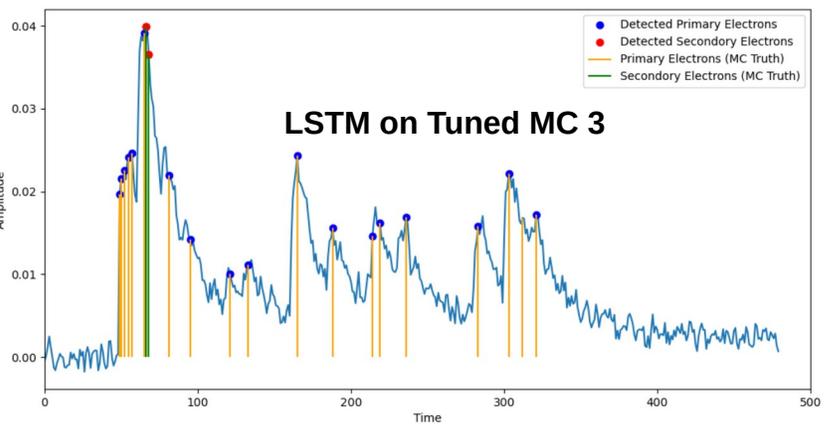
**Monte Carlo (MC)**



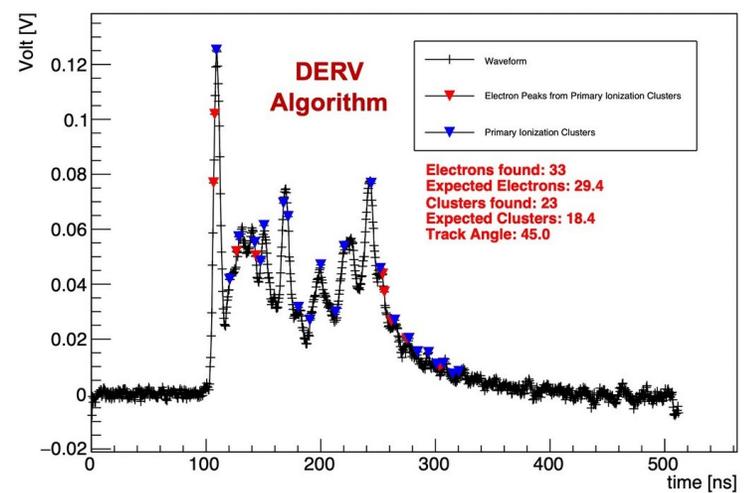
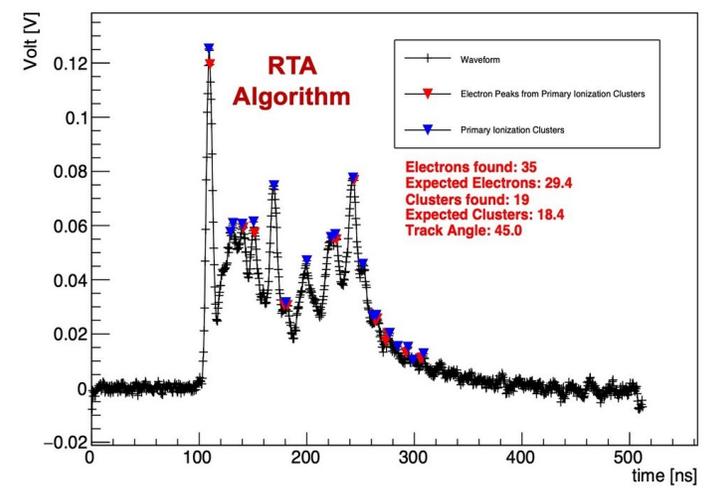
**180 GeV Real Test**



# LSTM Peak Finding Results of MC & Data In Comparison with RTA/Derivative Algorithm



Distribution of total detected peaks using LSTM arises from primary and secondary electrons (31.38) for channel 5 data with threshold 0.3



# Conclusion

Algorithm Type	Value
Algorithm	RTA (Template)
Mean detected peaks	35
Algorithm	Derivative (DERV)
Mean detected peaks	33
Algorithm	Machine Learning (LSTM)
Mean detected peaks	31.38

**ML Based Algorithm for  
Particle Identification (Simulated  
samples + NN) (Ongoing)**

## ML Based Algorithm for Particle Identification(Ongoing): Simulation Parameters

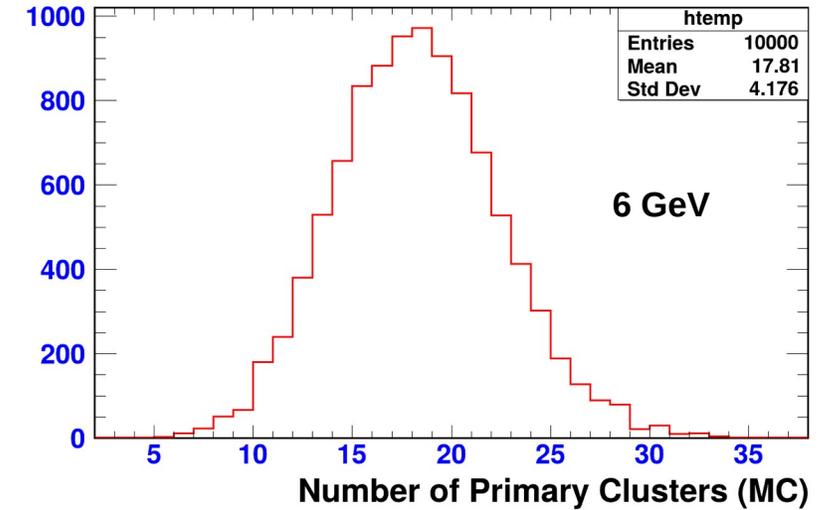
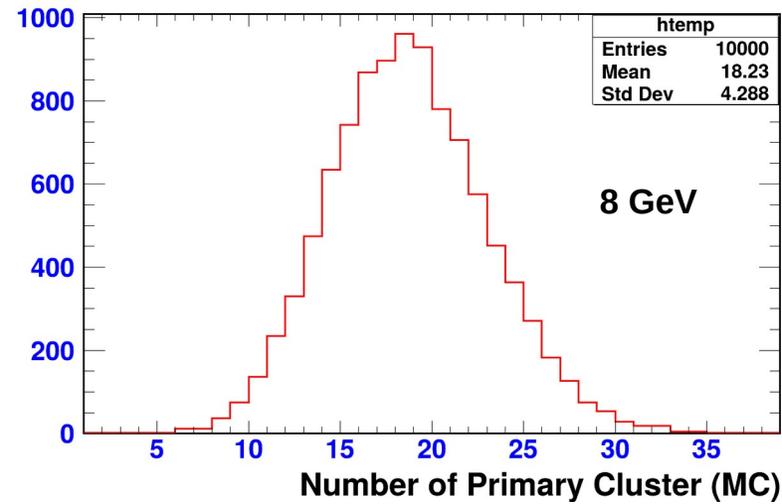
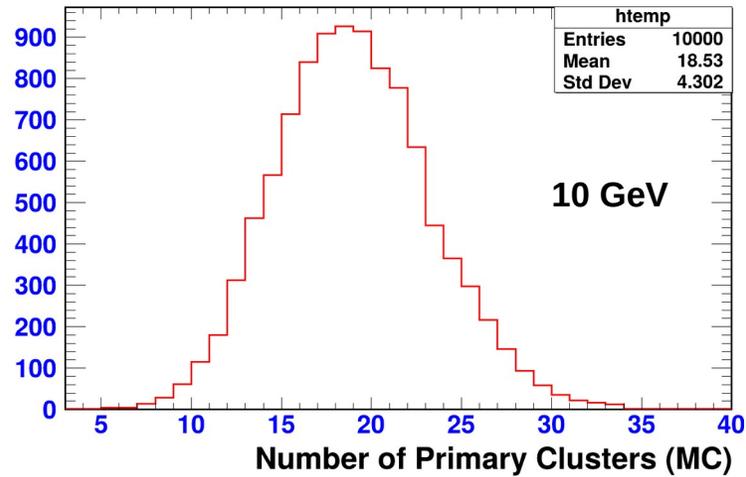
Simulation Parameters	
Particle	muon and Kaon
Gas mixture	He 90% + isobutane 10%
Cell size	1 cm
Momentum	2, 4, 6, 8, 10 GeV
Sampling rate	1.5 GHz
Angle	45°
Voltage at sense wire	1450 V



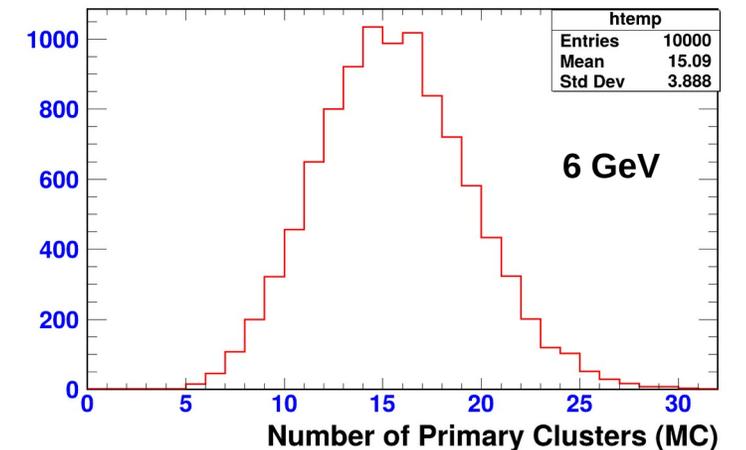
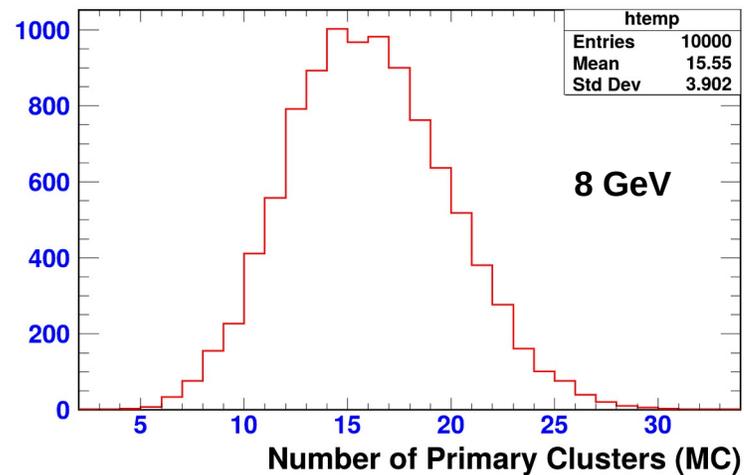
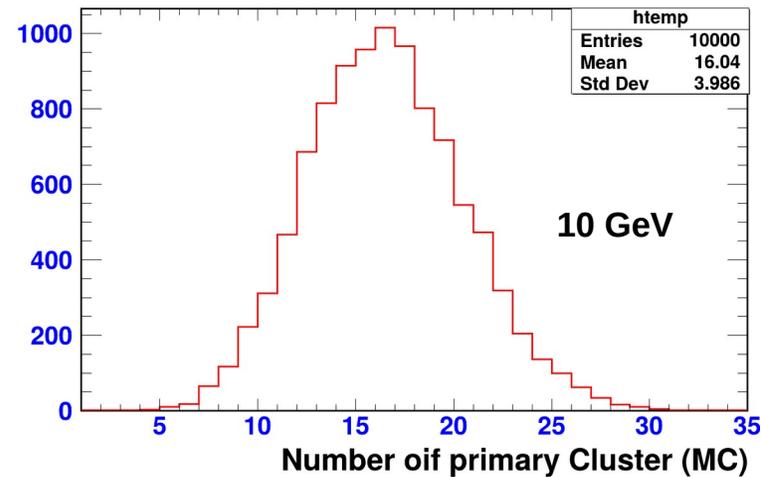
**Above are Simulation parameters  
were used which matched the real  
test beam data.**

# Number of Primary Cluster (MC)

## For Muon

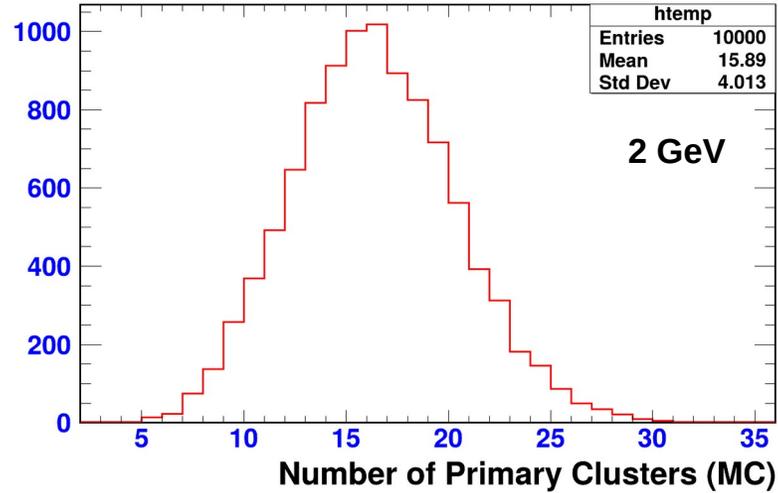
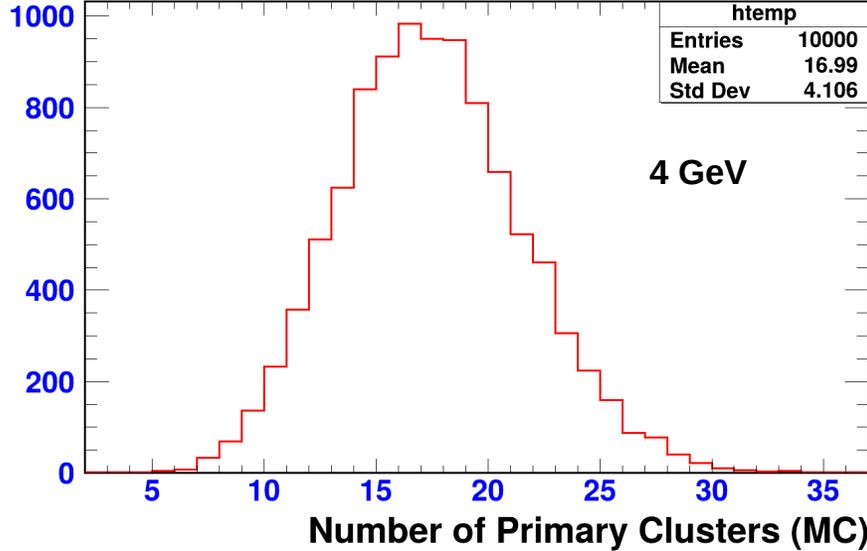


## For Kaon

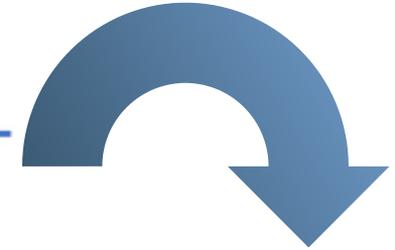
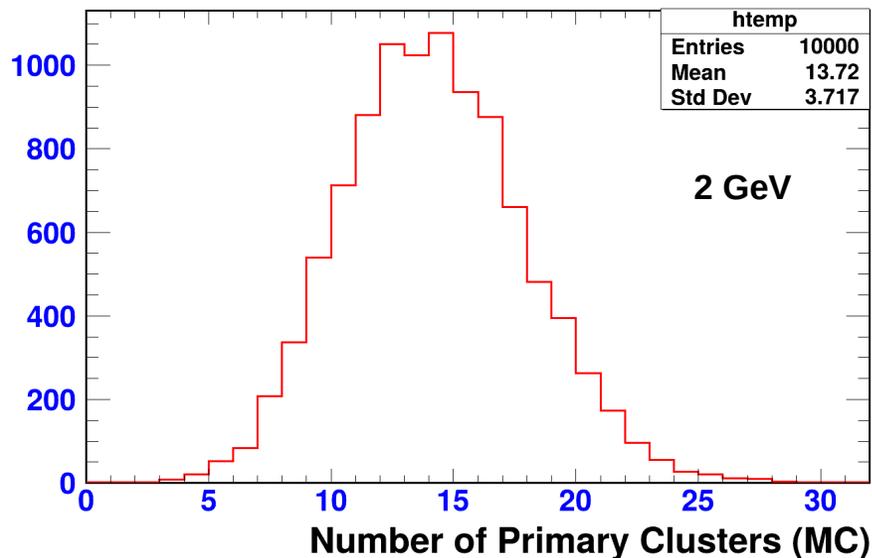
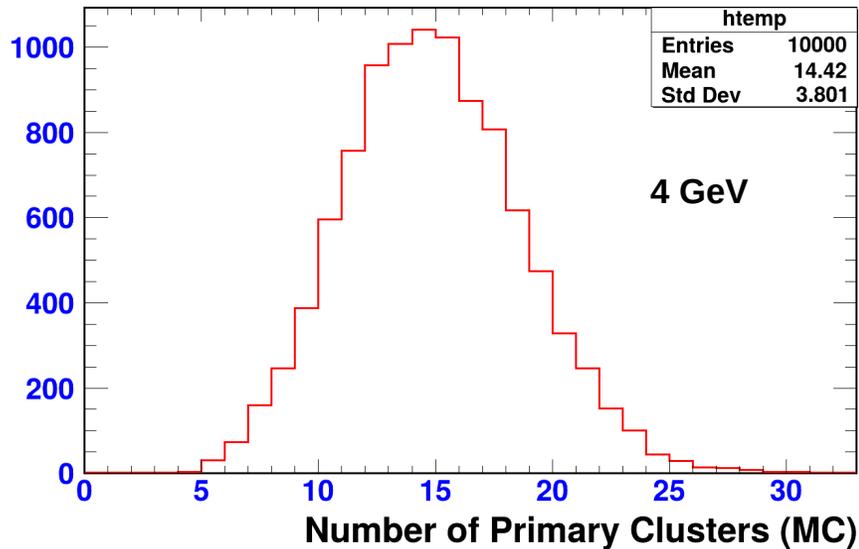


# Number of Primary Cluster (MC)

## For Muon



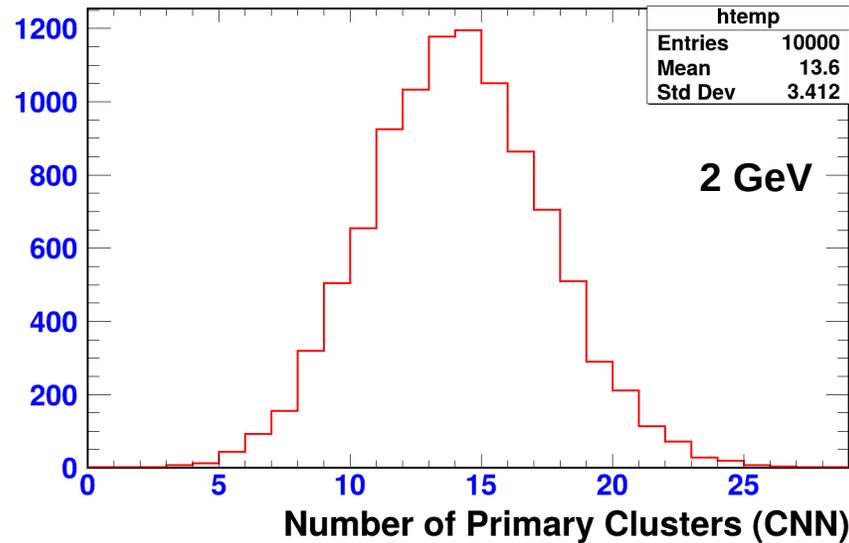
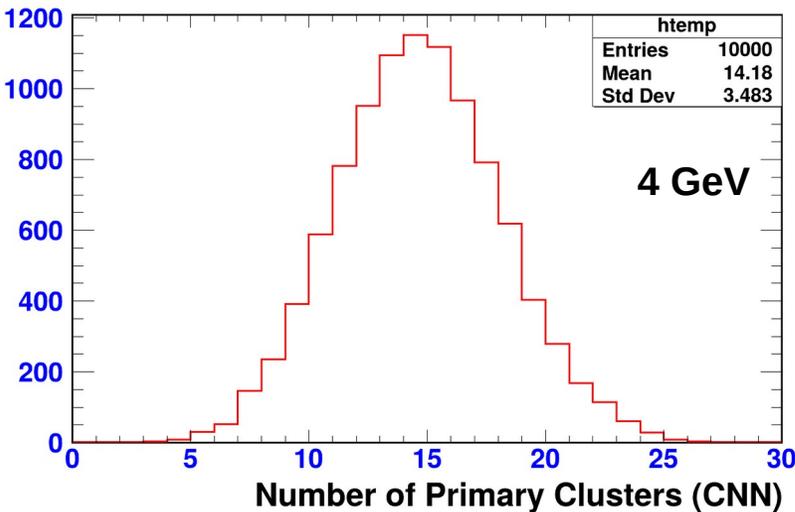
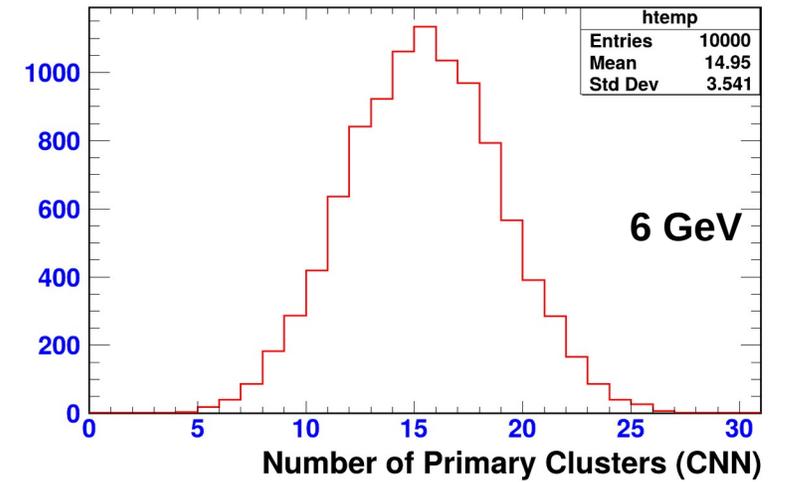
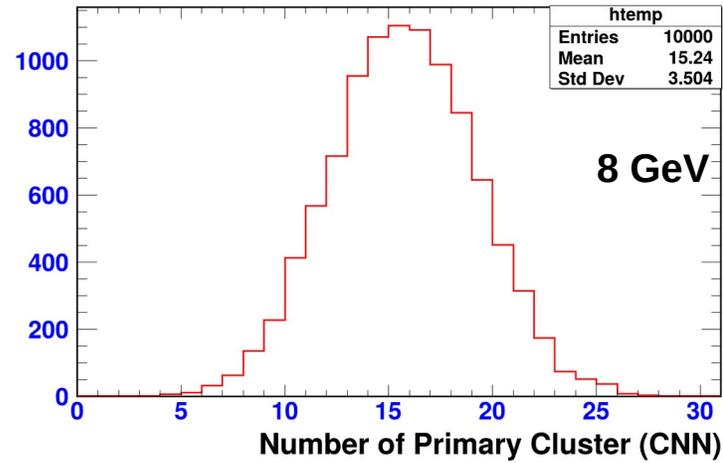
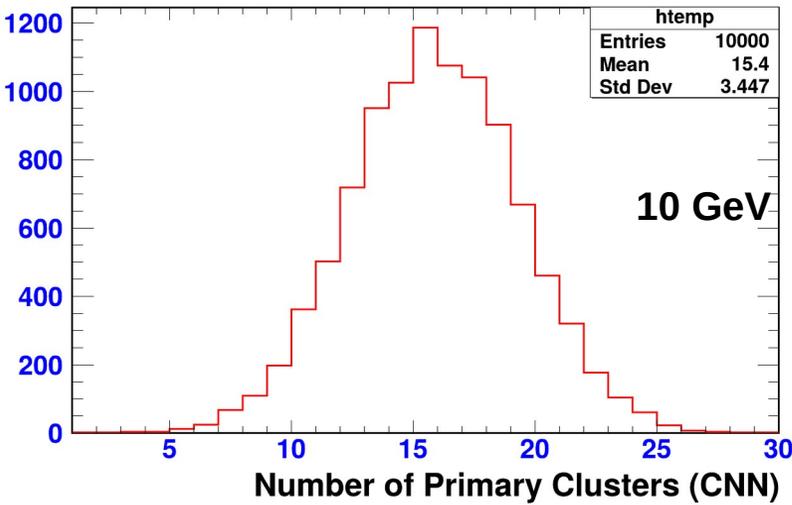
## For Kaon



All these distributions are related to the number of primary clusters Monte Carlo (MC) for the momentum (2-10 GeV) of muon and Kaon. We also need to generate it for 180 GeV momentum

# Reconstruction of Number of Primary Clusters (CNN)

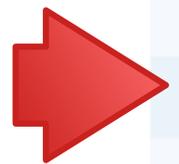
## For Kaon



All these distributions are related to the number of primary clusters reconstructed by CNN Model for the momentum (2-10 GeV) of Kaon

# MC truth and Final Results of Reconstructions for Kaon By CNN Clusterization Model

Number of primary Clusters MC Truth (Muon and Kaon)

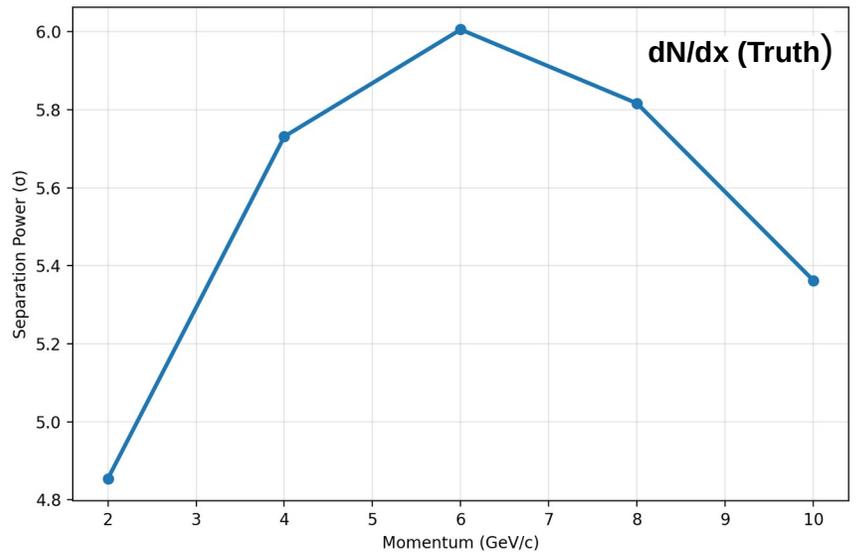


Momentum (GeV)		Muon		Kaon	
$p$		Mean	$\sigma$	Mean	$\sigma$
2		15.89	4.013	13.72	3.717
4		16.99	4.106	14.42	3.801
6		17.81	4.176	15.09	3.888
8		18.23	4.288	15.96	3.966
10		18.53	4.302	16.04	3.986

CNN Clusterization Results for Kaon (2-10 GeV). The mean number of primary clusters (CNN) matches with the MC

Momentum (GeV)	2	4	6	8	10
Mean	13.600	14.180	14.960	15.240	15.400
Std. Dev.	3.412	3.483	3.541	3.504	3.447

# Super Preliminary and Incomplete PID performances for MC and NN Reconstruction



$$S = \frac{|\mu_{\mu} - \mu_K|}{(\sigma_{\mu} + \sigma_K)/2}$$

The above plot is just the separation power for MC truth of 2, 4, 6, 8 and 10 GeV momenta of muon and kaon while for the NN reconstruction, I will do soon

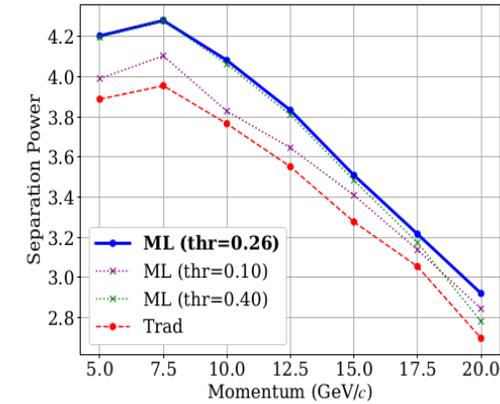
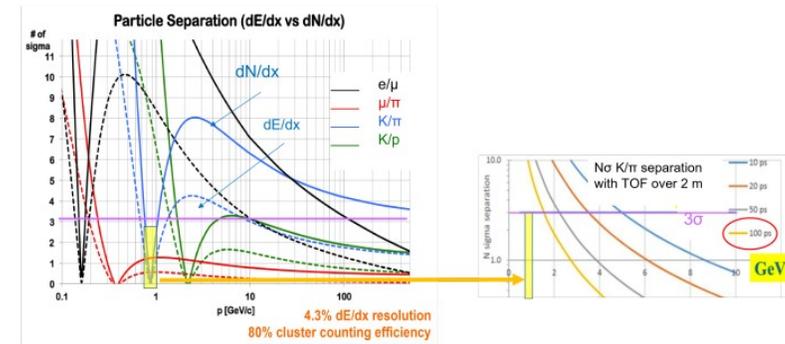
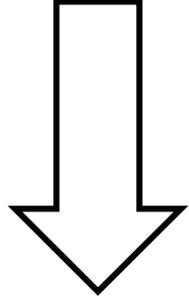


Fig. 9. The  $K/\pi$  separation power as a function of track momentum for 1 m track length. The red dashed line is from the traditional algorithm. The blue solid, violet dotted and green dotted lines are from the ML-based algorithm with a threshold of 0.26, 0.10 and 0.40, respectively. The blue solid line with a threshold of 0.26 achieves the overall best performance, which has a roughly 10% better  $K/\pi$  separation than the traditional algorithm.



**Figure 1:** Left: Analytic evaluation of particle separation capabilities achievable with dE/dx (solid curves) and dN/dx (dashed curves). The region between 0.85 GeV/c and 1.05 GeV/c where a different technique is needed is highlighted in yellow. Right: PID performance as a function of the time resolution by using a time of flight technique over 2 m to recover the particle identification around 1 GeV.

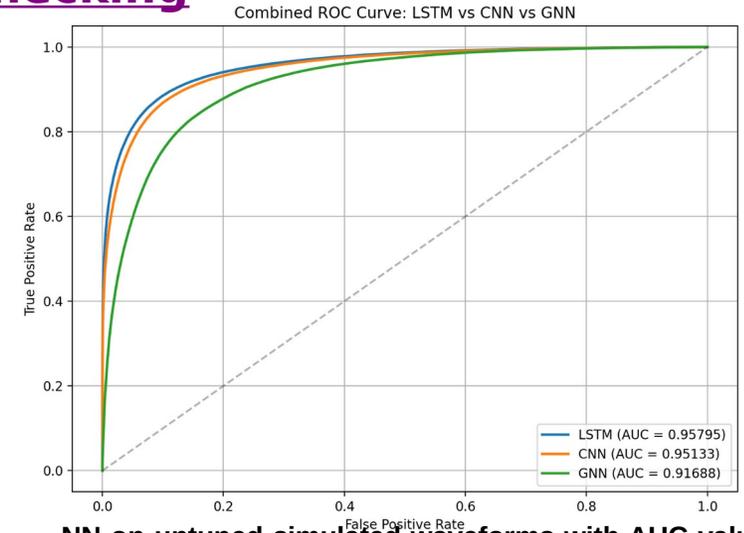
# What to do Next



- ◆ Now we will train CNN clusterization model and apply the model on MC and data to estimate the number of primary clusters based on the detected peaks
- ◆ All the results should be compare with RTA algorithm
- ◆ ML based Algorithm for PID should also be completed for muon and kaon

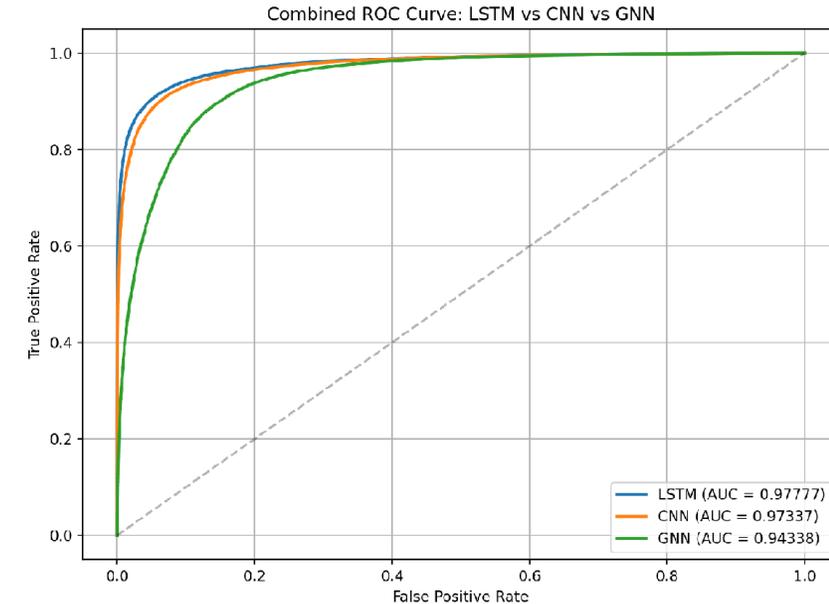
## Models optimization Just for Checking Which I did. Best is LSTM

AUC Value: LSTM VS CNN VS GNN for 180 GeV Tune Simulated waveforms (50000)



NN on untuned simulated waveforms with AUC value. It is also done for other three models too like DGCNN but the best one I found LSTM everywhere

AUC Value: LSTM VS CNN VS GNN for 180 GeV UnTune Simulated waveforms (50000)



NN on tuned simulated waveforms with AUC value. It is also done for other three models too like DGCNN but the best one I found LSTM everywhere

*Thank  
you*



# 1) THESIS (Priority)

**January 19 – February 5**

**→ Complete the following thesis chapters:**

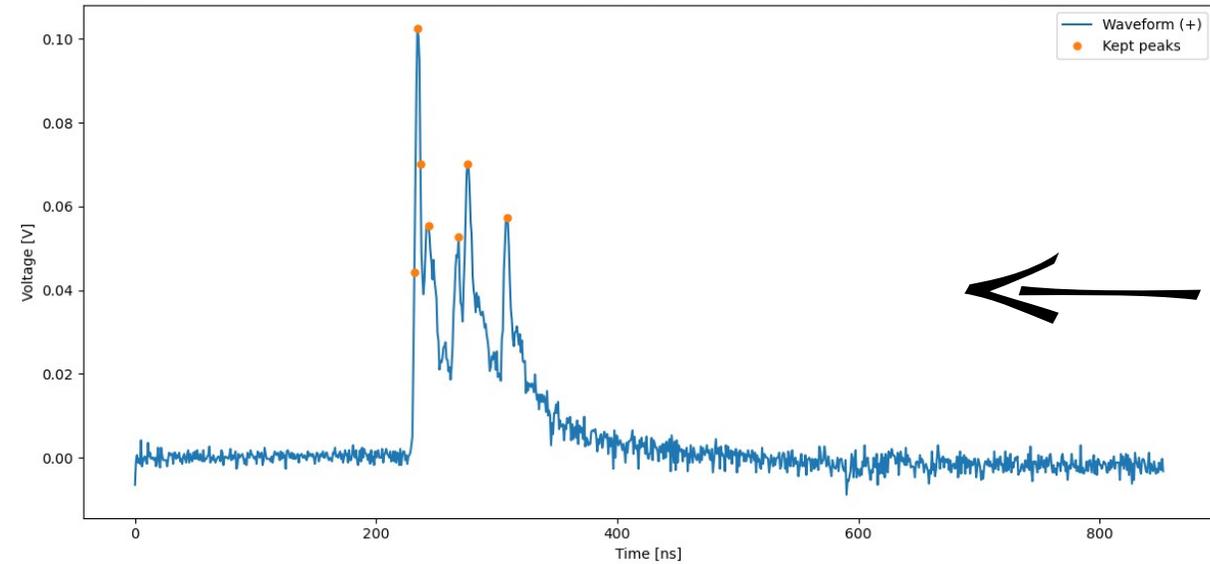
- Introduction to the ICSC Project (Completed)
- Introduction to FCC and IDEA Detector (Completed)
- Full Simulation of the IDEA Drift Tube Using Garfield++ (Completed)
- Deep Neural Network Models in Machine Learning (45 % completed)
- Machine-Learning-Based Cluster Counting for Particle Identification

**February 6 – February 28**

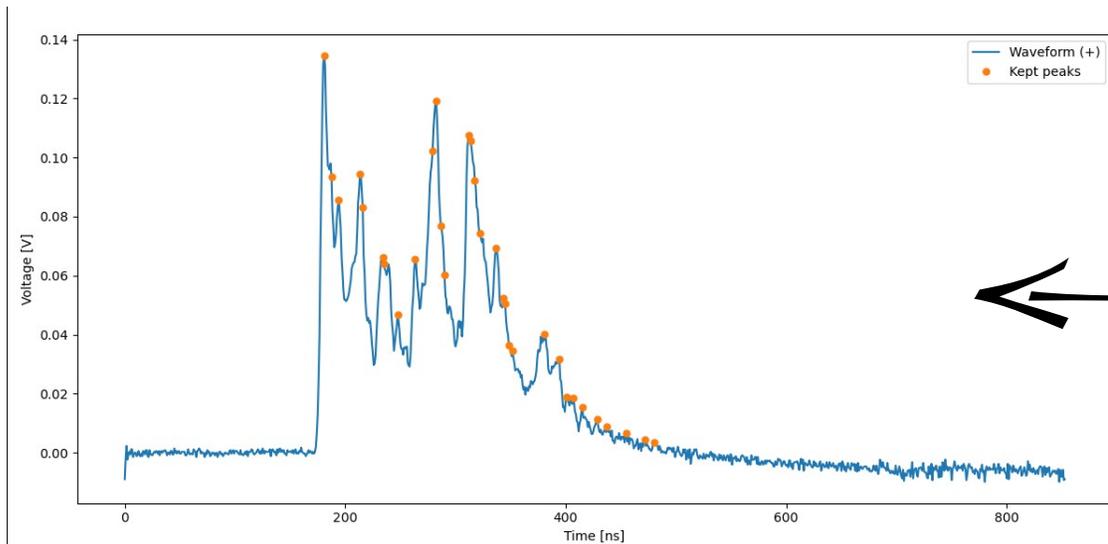
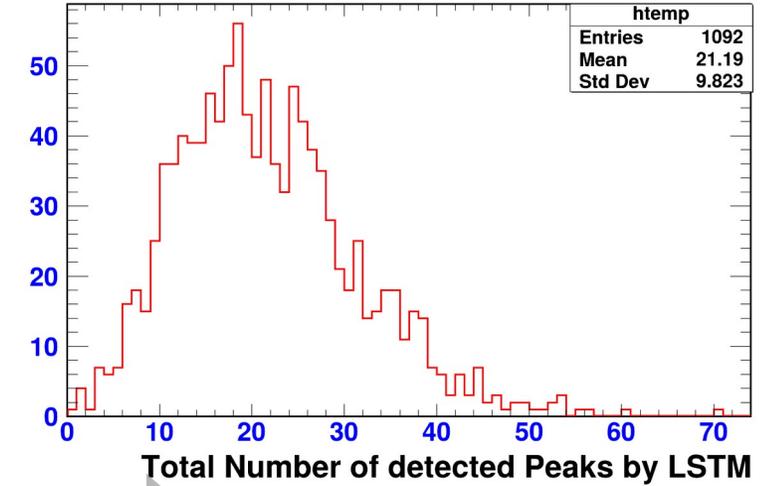
**→ Complete the remaining two thesis chapters:**

- Hyperparameter Optimization Using High-Performance Computing Resources and Final Reconstructed Results of Simulations and Neural Networks

# Selection cuts on real data (Selection cut /Threshold)

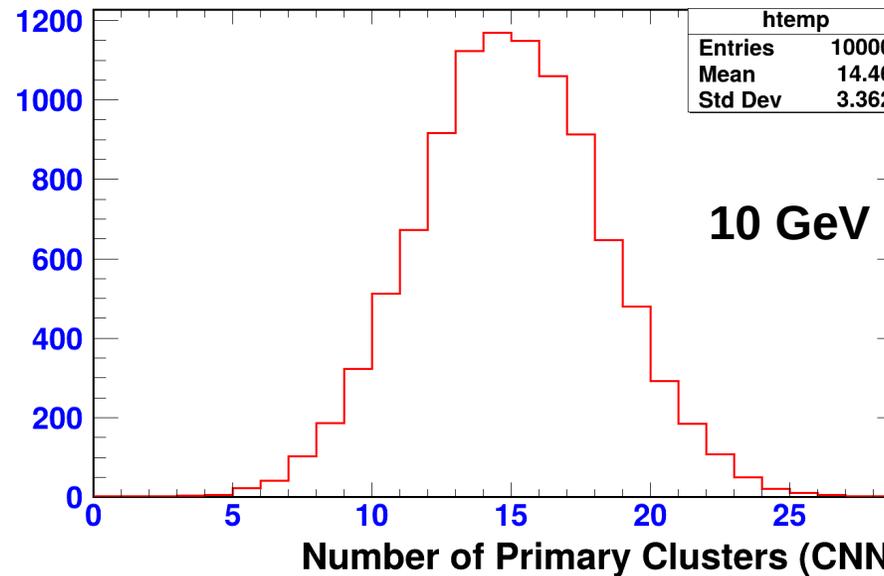
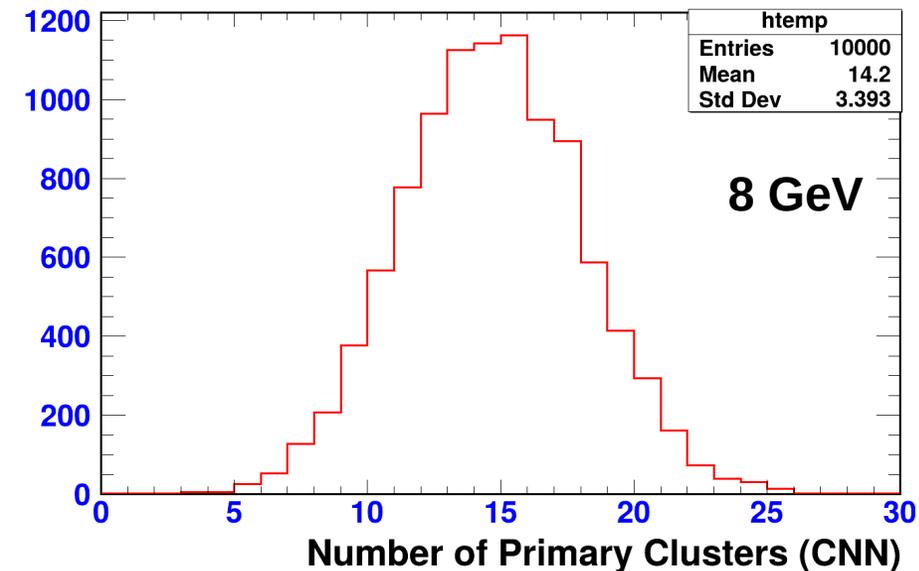
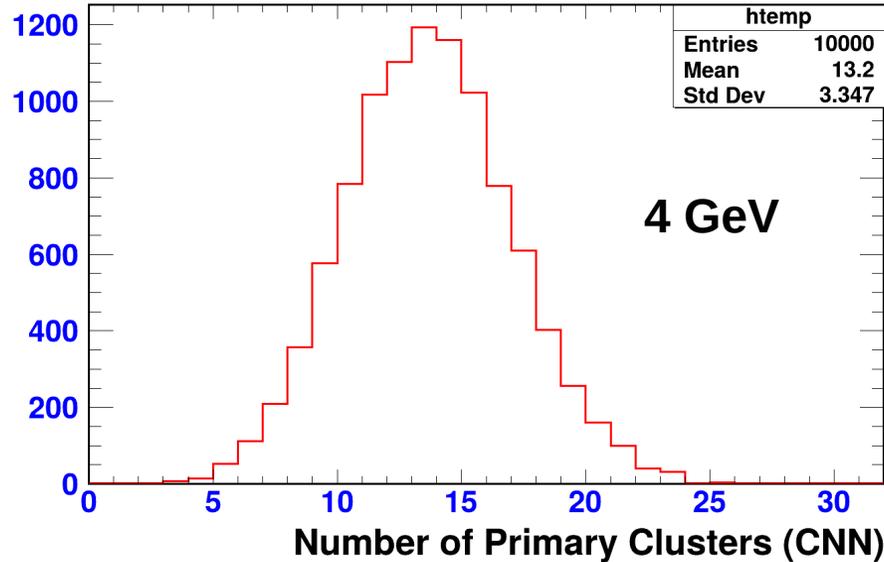
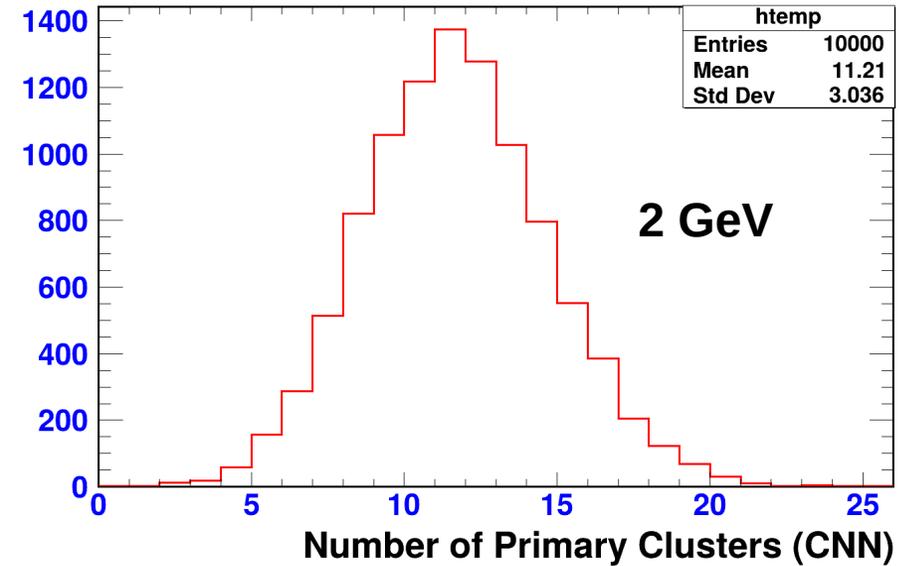


At Selection cut/Threshold = 0.4, less peaks appear as result found less mean value of the detected peaks in the distribution



At Threshold = 0.25, some peak appear in the region of noise amplitude peaks appear as result found some fake peaks in mean value of the detected peaks in the distribution

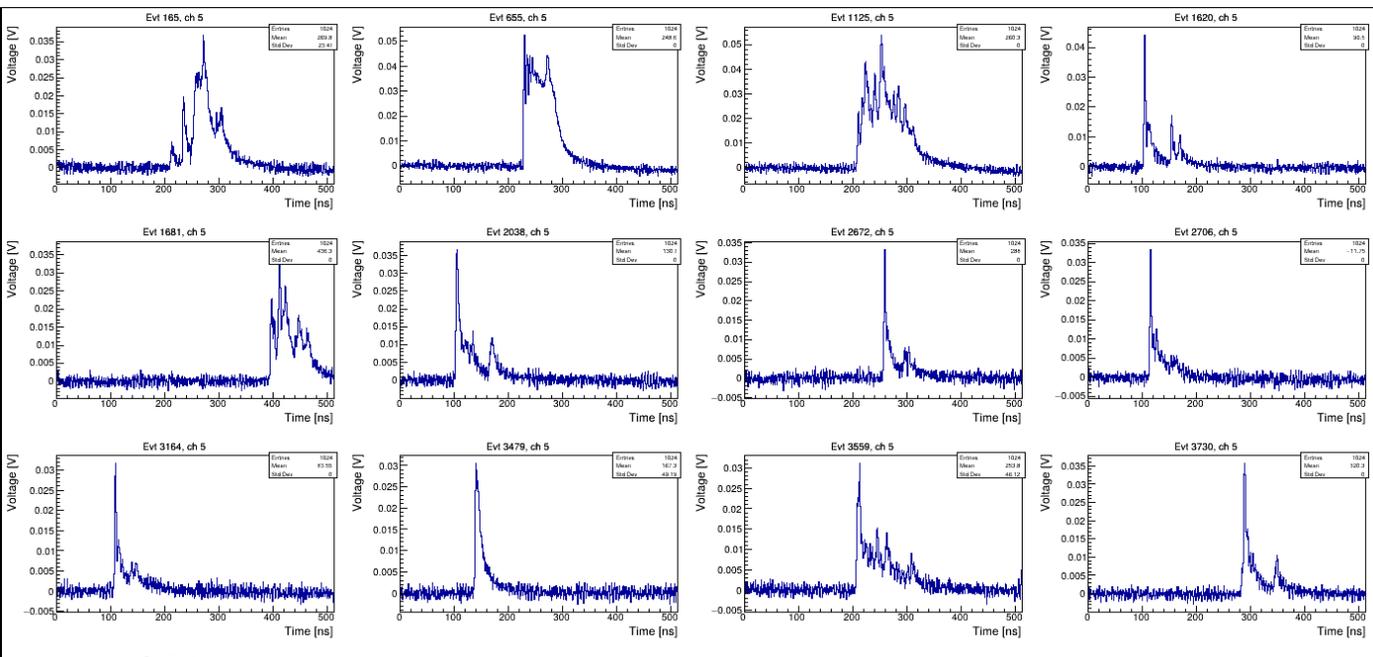
# Selection cut (0.9) on Simulated samples of Kaon



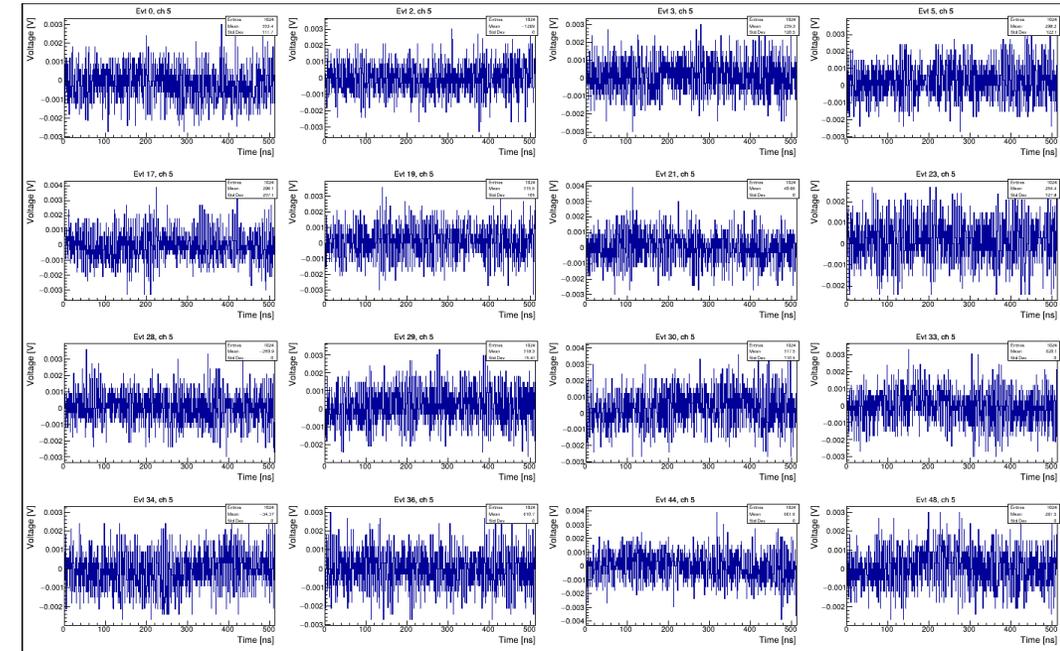
These distributions are related to primary clusters by CNN using threshold (0.9). As a results the mean number of primary clusters and standard deviation less compare to primary clusters (MC)

# Beam test data selection for Channel 5

## Signal-like data



## Noise-like data



- Signal-like data: peak finding
- Noise-like data: noise information

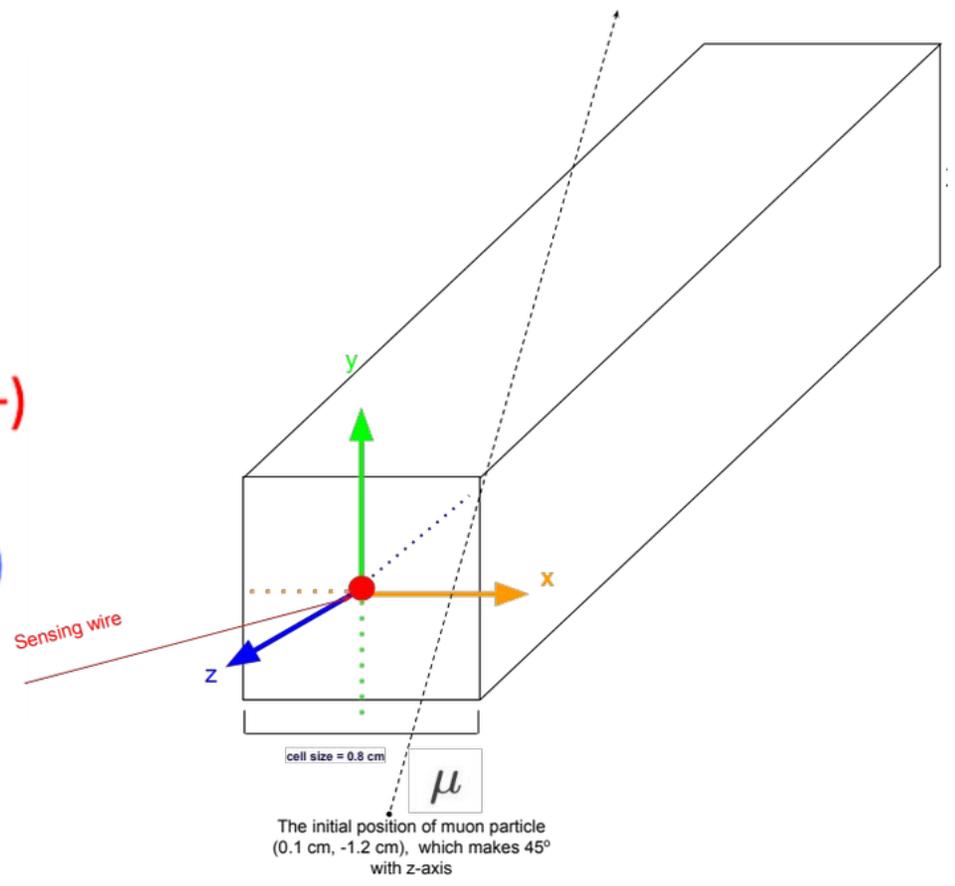
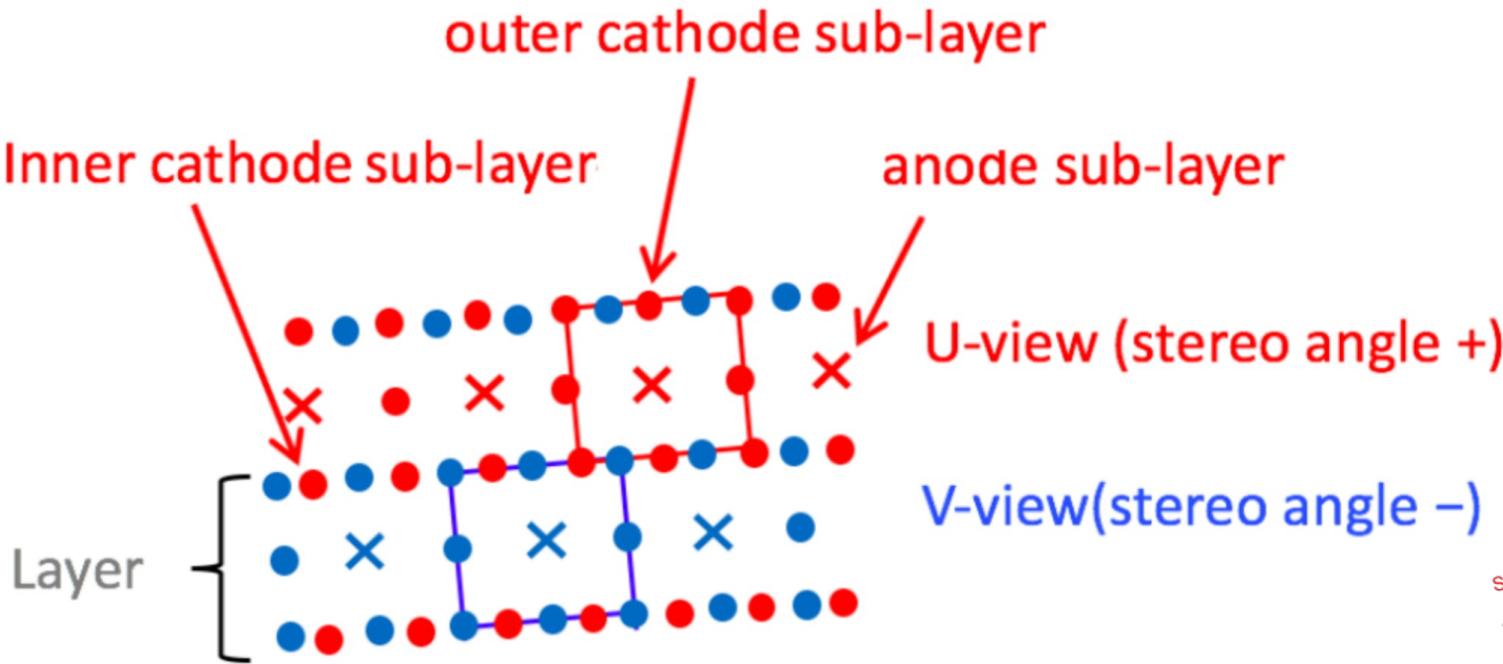
# Waveform Selection Logic

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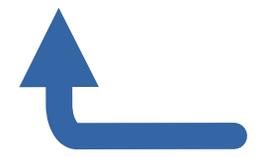
Quantity	Value
Total entries in TChain (all channels)	58,002
Entries for channel 5 (within range)	4,832
<b>Noise RMS cut</b>	$\sigma_{\text{noise}} < 2 \text{ mV}$
<b>Amplitude cut (VALID)</b>	$A_{\text{max}} > 0.03 \text{ V (30 mV)}$

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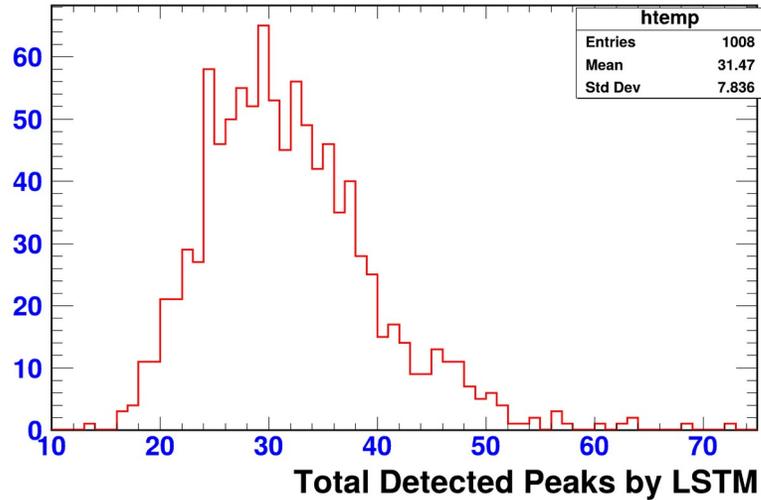
of 56,448 drift cells. Each cell is designed with a ratio of field to sense wires equal to 5:1 to ensure the proper electrostatic configuration, and is composed by one anode and two cathode sub-layers, as sketched in Fig. 3-Top. The anodes are 20 μm diameter tungsten wires, while the cathodes are 40 and 50 μm light aluminum alloy wires. In total, the CDCH is made with 56,448 sense wires,



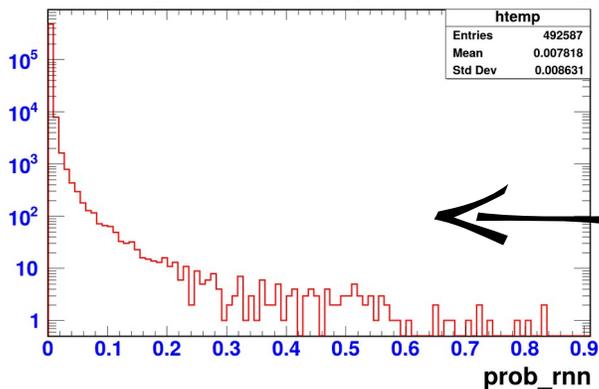
One cell is considered here as one sense wire surrounded by eight field wires... (Uniform Field)

# Applying Untrained NN on beam test data 2022 (180 GeV) Just for Checking

## Preliminary Results of Peak Finding distribution

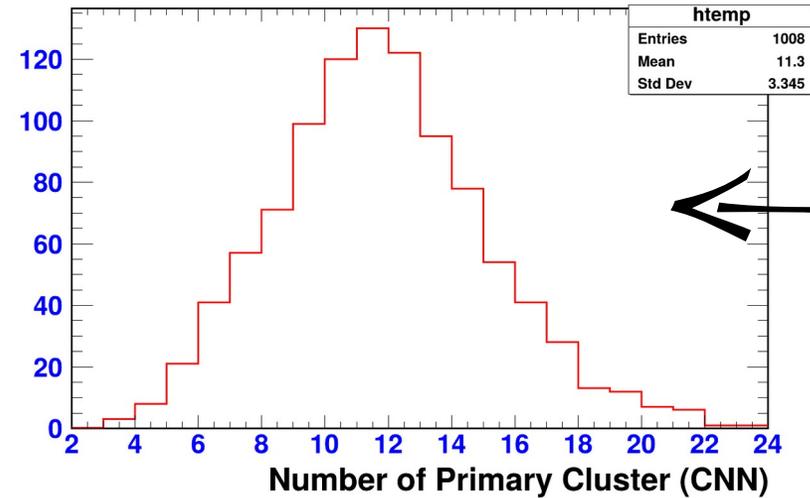


- The above distribution shows us total number of detected peaks (primary and secondary electrons) the

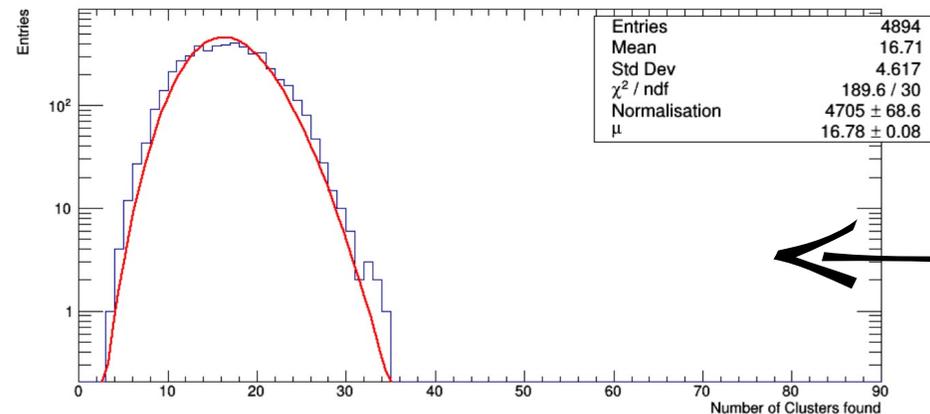


Bad Discrimination of signal and background by using trained LSTM model on untuned simulated waveforms. Because we did not see two main peaks one close to zero represent mostly background events and other close to one represent mostly signal events

## Clusterization

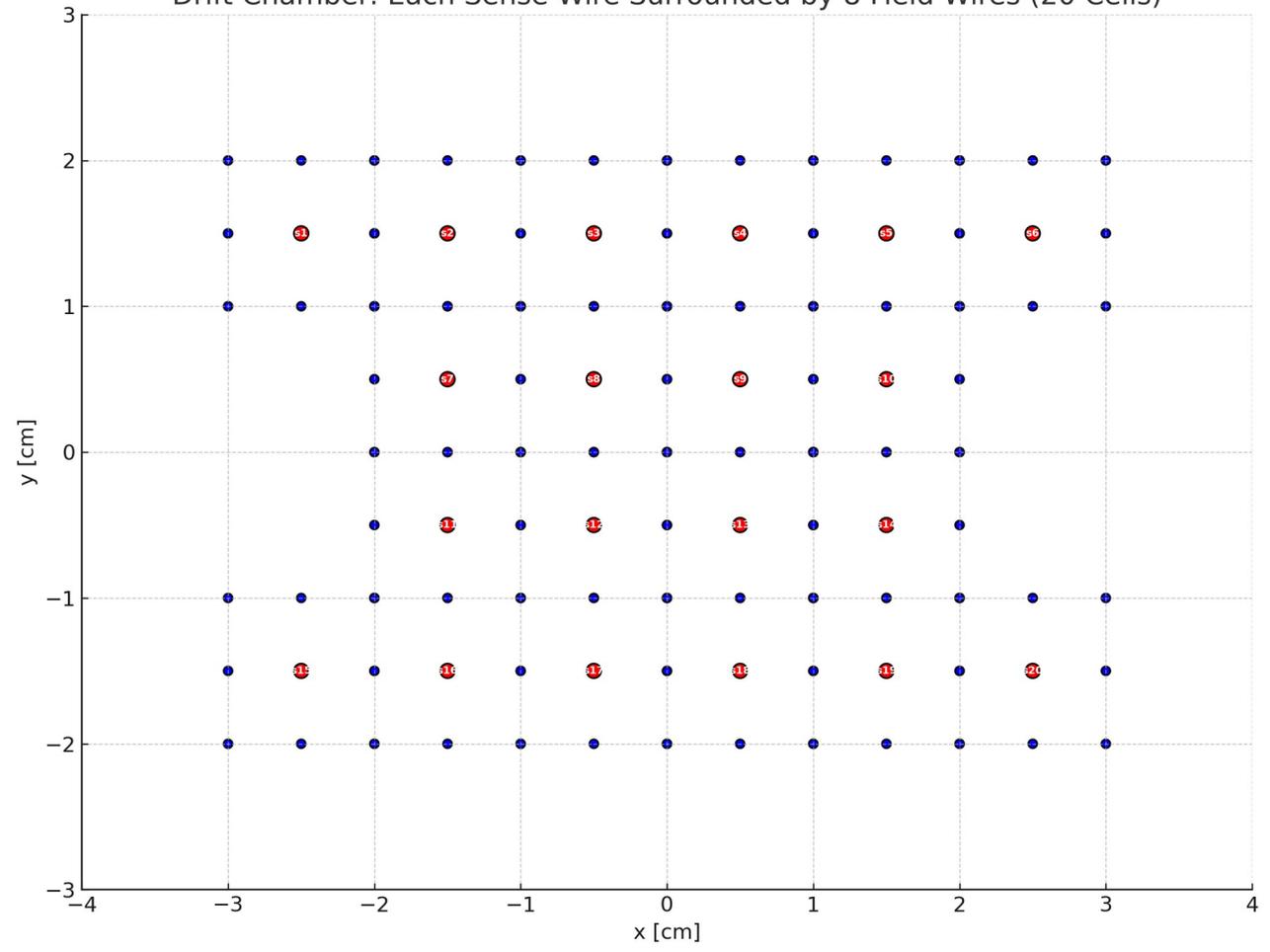


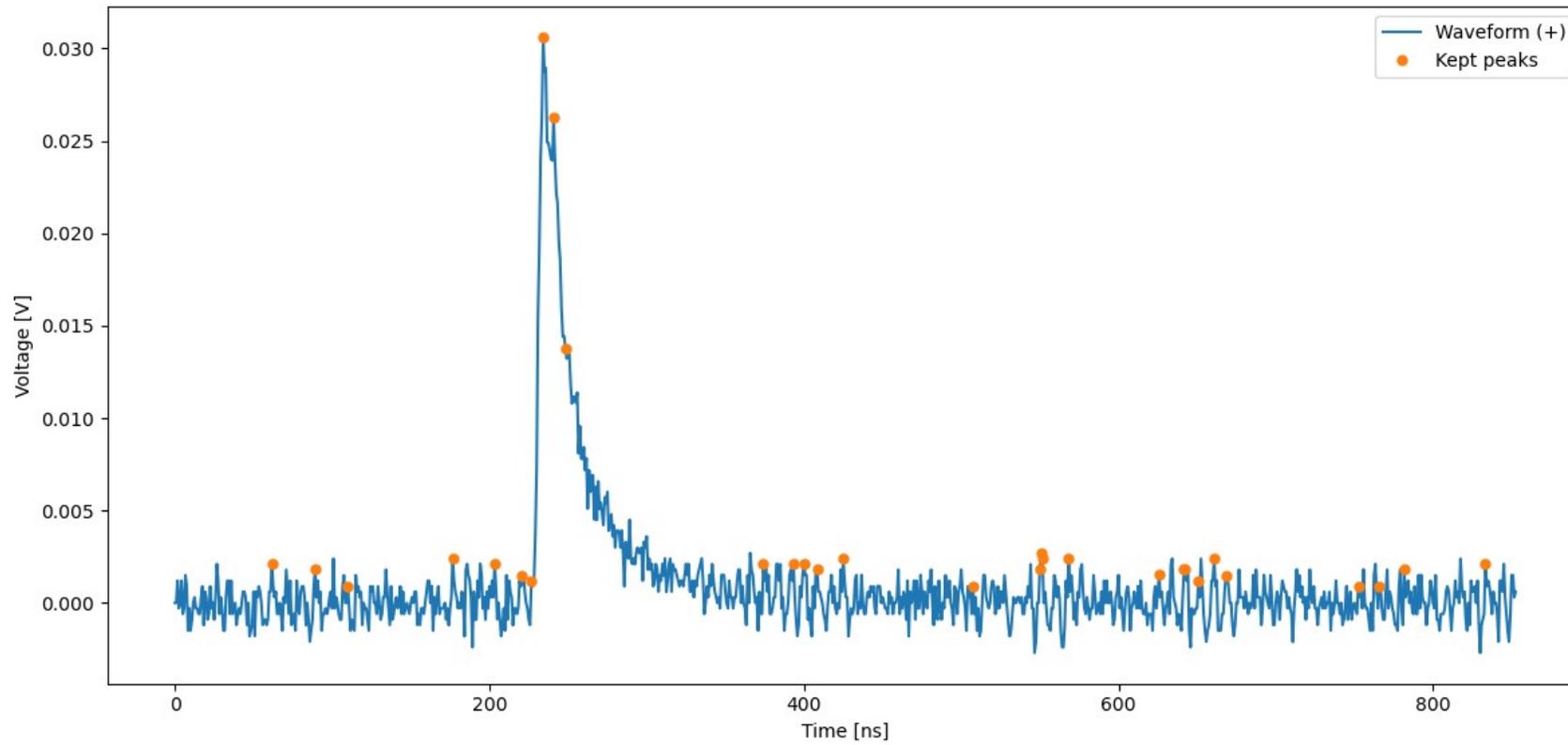
← Preliminary CNN Clusterization



← RTA Algorithm

Drift Chamber: Each Sense Wire Surrounded by 8 Field Wires (20 Cells)





### 1. Peak Detection Based on Probability Cut:

The detection process involves looping over all events and applying a probability cut to decide whether a peak is considered a valid detection:

**Looping Over Events:** The script iterates over all entries (events) in the probability file.

**Applying the Cut:** For each event, it checks if the predicted probability (prob\_ml) exceeds the cut threshold (cut), which is set to 0.95/0.65.

**Storing Detected Peaks:** If the probability exceeds the threshold, the corresponding peak time is stored in the detected\_time dictionary, keyed by event number (evtno)

### 2. Matching Detected Peaks with Truth Data

After detecting peaks, the script matches these peaks with the Monte Carlo (MC) truth data to classify them as primary or secondary:

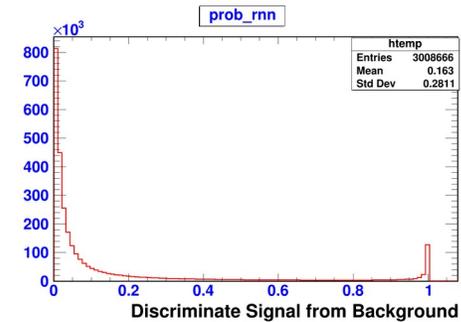
**Truth Data:** The truth data (truth\_time, truth\_tag) contains the actual times of primary and secondary peaks, labeled by truth\_tag (1 for primary, 2 for secondary).

**Matching Function:** The match function compares detected peak times with the truth peak times. For each detected peak, it finds the closest truth peak and assigns the corresponding tag (primary or secondary) based on the truth data.

**ID Assignment:** The id\_list array stores the classification of each detected peak as primary (1) or secondary (2)

**3. Counting Primary Peaks:** After classifying the detected peaks, the script counts how many of them are primary peaks:

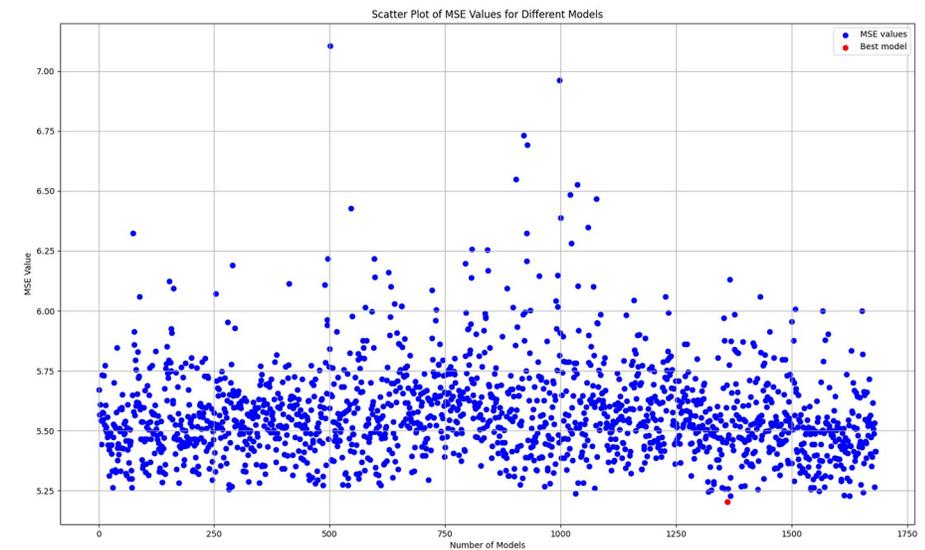
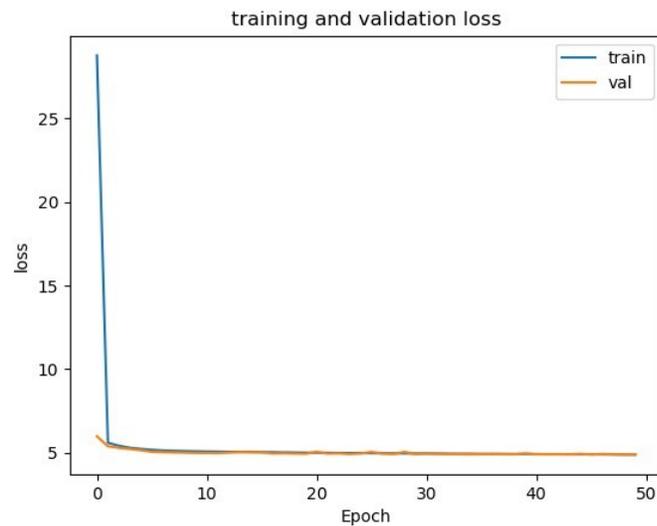
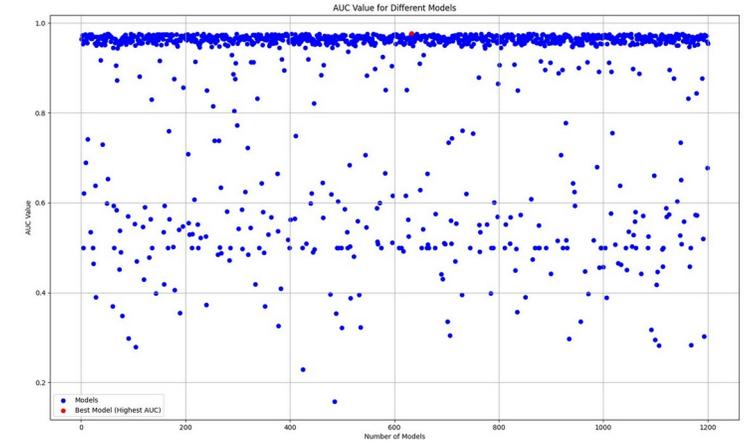
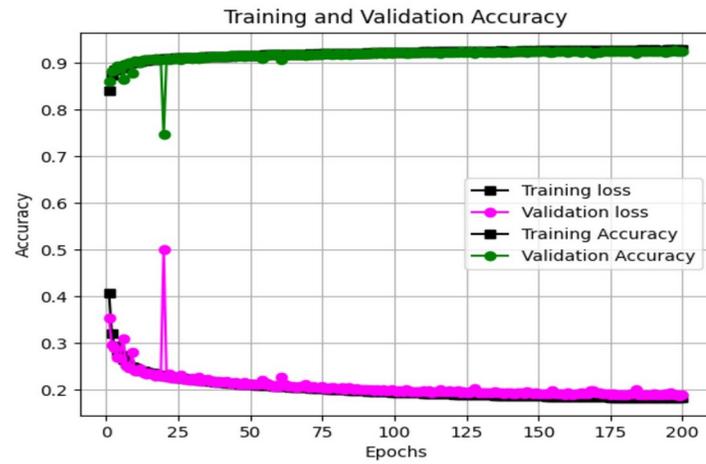
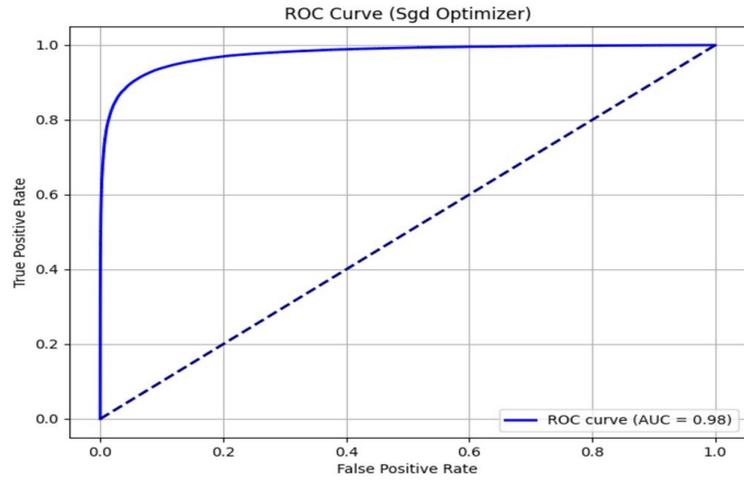
**Counting:** The script iterates over the id\_list and increments ncount\_pri for every primary peak (tag 1)



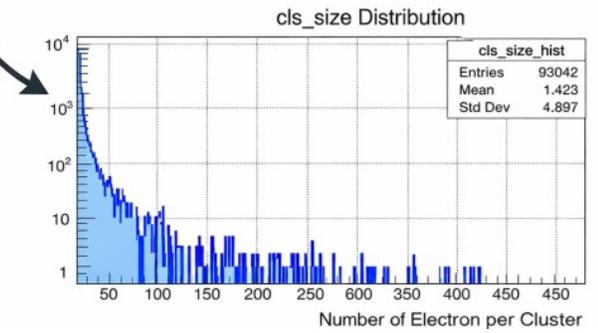
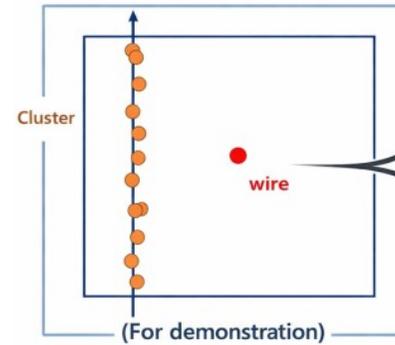
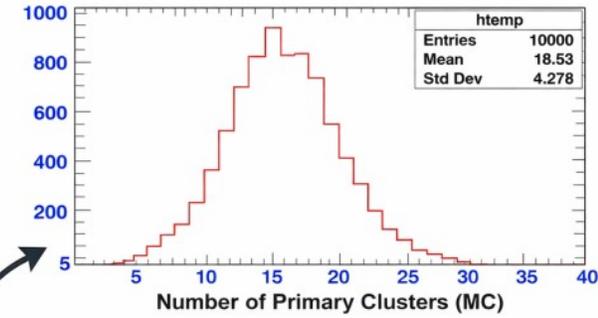
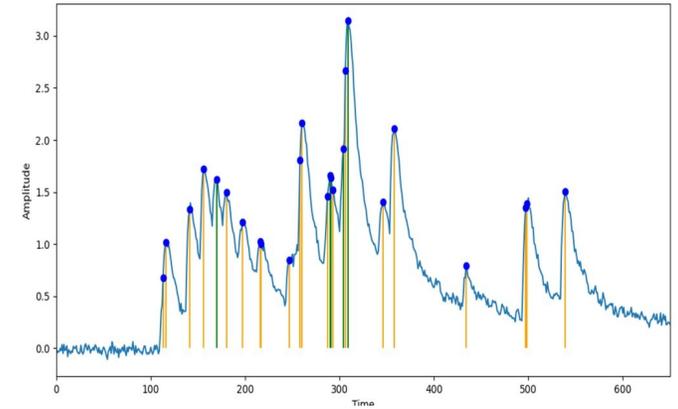
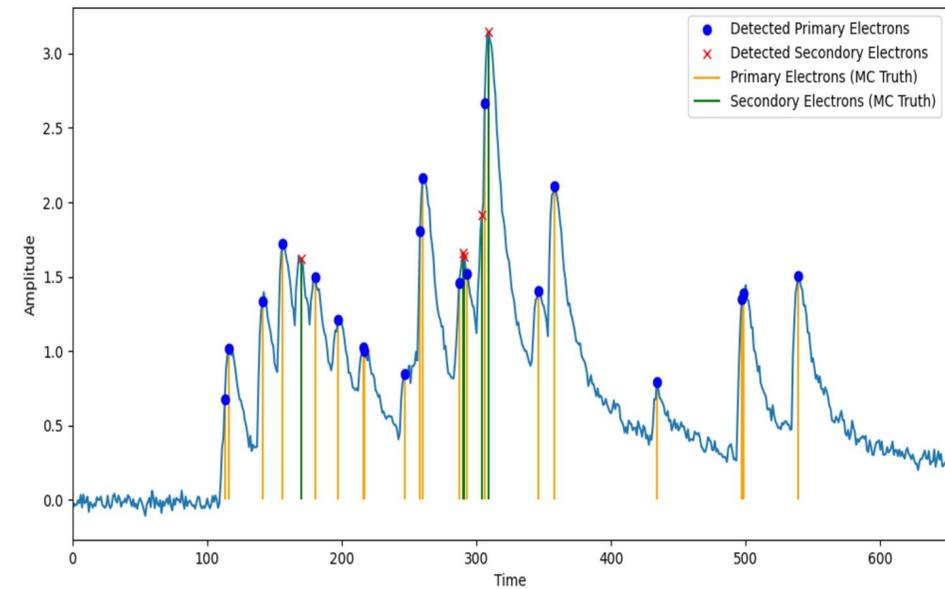
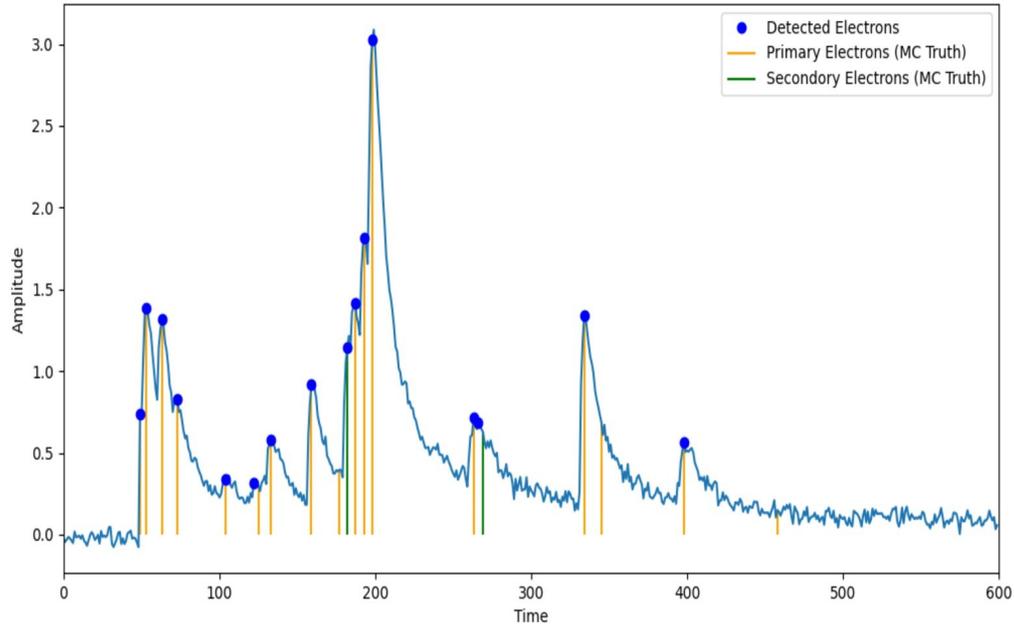
Different Momenta of Muon (GeV)	2	4	6	8	10	180
<b>Monte Carlo (MC)</b>						
Primary Cluster (MC)	15.89	16.99	17.81	18.28	18.53	19.10
Std. Deviation (MC)	4.01	4.10	4.12	4.30	4.20	4.30
<b>LSTM Model</b>						
Primary Cluster (LSTM)	14.45	15.37	16.06	16.34	16.49	17.30
Std. Deviation (LSTM)	3.77	3.84	3.90	3.90	3.90	4.02
<b>CNN Model</b>						
Primary Cluster (CNN)	14.38	15.00	15.38	15.77	16.29	16.76
Std. Deviation (CNN)	3.37	3.20	3.20	3.10	3.30	3.20

**Table 1:** Primary cluster means and standard deviations from MC, LSTM, and CNN across different muon momenta.

# Best LSTM Peak Finding (Above 1<sup>st</sup> row) and CNN Regression Model (Below 2<sup>nd</sup> Row)



# Simulated Peak Finding Waveforms



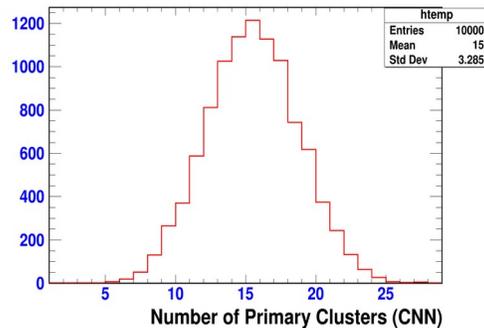
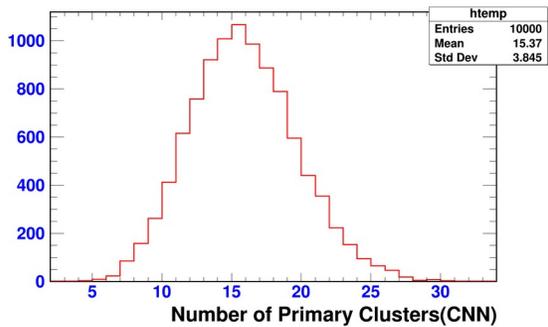
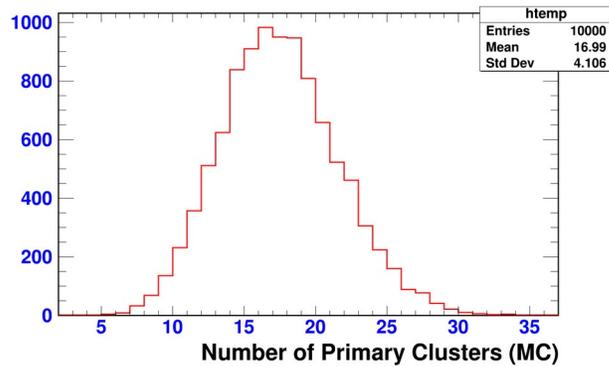
$$N_{\text{primary clusters}} \sim \text{Poisson}(\bar{N}_p)$$

with the mean

$$\bar{N}_p = \lambda L$$

- $\lambda$ : average primary-cluster density (clusters per cm)
- $L$ : track length in the gas (cm)

# Final results of the reconstruction for 4 GeV on the base of different selection cuts



Different selection cuts for 4 GeV	MC	$\sigma$ of MC	Primary Cluster( LSTM)	$\sigma$ of (LSTM)	CNN	$\sigma$ of (CNN)
0.55	16.99	4.1	15.37	3.84	15.	3.2
0.65	16.99	4.1	15.05	3.78	14.9	3.3
0.85	16.99	4.1	14.05	3.5	13.77	3.2
0.95	16.99	4.1	12.74	3.32	12.19	2.99