

第一届强相互作用物理自旋研讨会, 青岛, 2026.07.26-31



北京航空航天大学
BEIHANG UNIVERSITY

相对论手征核力进展与展望

耿立升 北京航空航天大学

- ① Jun-Xu Lu, Chun-Xuan Wang, Yang Xiao, **LSG***, Jie Meng, and Peter Ring, PRL128 (2022)142002
- ② Wei-Jiang Zou, Jun-Xu Lu, Peng-Wei Zhao, **LSG***, Jie Meng, PLB 854 (2024)138732
- ③ Shihang Shen, Jun-Xu Lu, **LSG***, Jie Meng, Wei-Jiang Zou, 2507.01257, to appear in PRC
- ④ Ru-You Zheng, Zhi-Wei Liu, **LSG***, Jin-Niu Hu, Sibao Wang, PLB 864 (2025) 139416
- ⑤ Xuan Zeng, Ru-You Zheng, Zhi-Wei Liu, **LSG***, Xian-Rong Zhou*, PRC 113 (2026) 014310
- ⑥ Qing-Yu Zhai, Dan-Yang Pang, Wen-Di Chen, O. A. Rubtsova, Rui-Rui Xu, Jun-Xu Lu*, Haozhao Liang,[†] and **LSG[‡]**, [2512.02475](#), to appear in PRC
- ⑦ Jun-Xu Lu and **LSG***, PRC 113 (2026) 054002
- ⑧ Qing-Yu Zhai, Jun-Xu Lu, and **LSG***, in preparation

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- 👉 **Why relativistic chiral nuclear force**
- 👉 **The first high-precision relativistic chiral nuclear force**
- 👉 **Recent developments and works in progress**
 - *Progress in constructing higher-order NN and A_y puzzle*
 - *Progress in studies of nuclear matter and finite nuclei*
- 👉 **Summary and outlook**

Nuclear force—always at the frontiers of nuclear physics

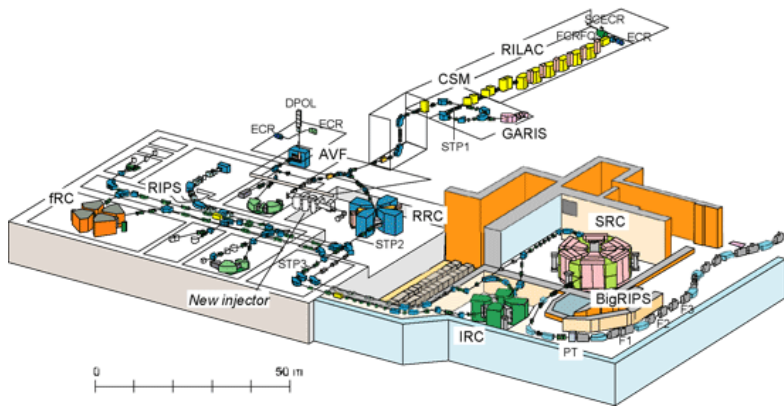
RIBF



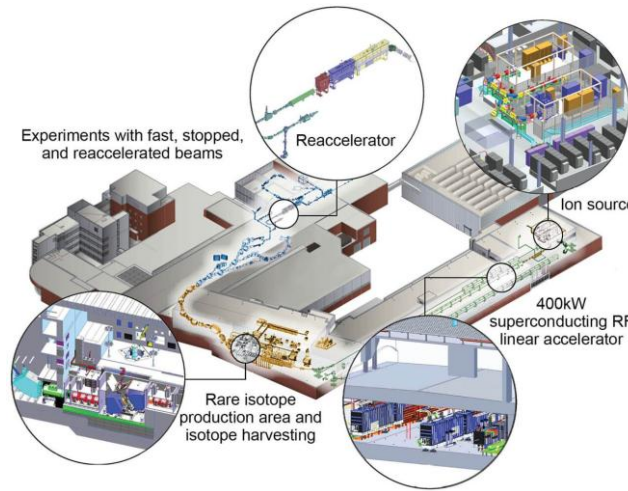
RIKEN Nishina Center for Accelerator-Based Science

Introduction to RI Beam Factory and Users' Information

RIBF Accelerators



Facility for Rare Isotope Beams
at Michigan State University



Three pillars

- **understanding of the nuclear force, especially the 3NF**
- understanding of the physics of the continuum
- understanding of the structure and dynamics of heavy element RIs,

Experimental nuclear physics

- map the nuclear landscape
- **understand the forces that bind nucleons into nuclei**
- answer questions about the astrophysical origin of nuclear matter
- address societal needs related to nuclear science and technology

HIAF scientific goals

- **认识原子核内有效相互作用。**
- 探索宇宙中从铁到铀元素的来源
- 粒子辐照应用研究

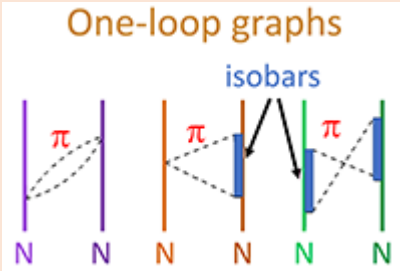
Central to developments in the field of nuclear physics



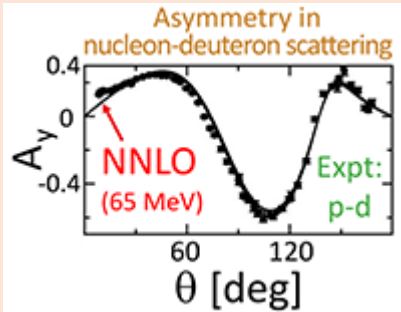
Remain central to developments in the field of nuclear physics
Announcing major discoveries/opening up new avenues of research

7/41

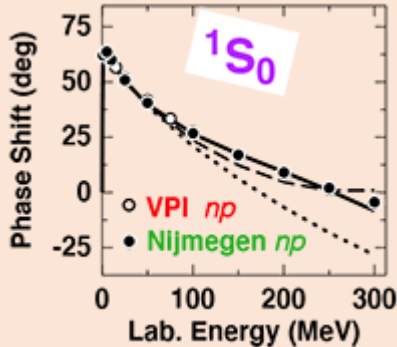
chiral EFT



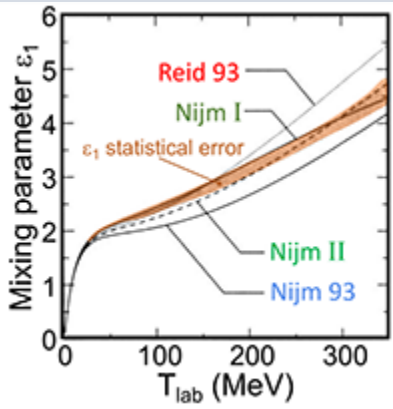
Ordóñez, Ray, van Kolck
Phys. Rev. C **53**, 2086 (1996)



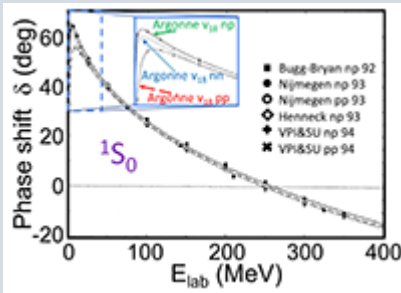
Epelbaum, Nogga, Glöckle,
Kamada, Meißner, Wiśniewski
Phys. Rev. C **66**, 064001 (2002)



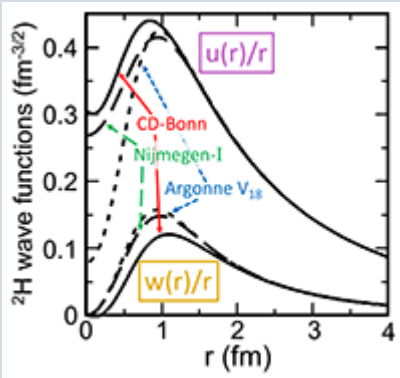
Entem, Machleidt
Phys. Rev. C **68**, 041001 (2003)



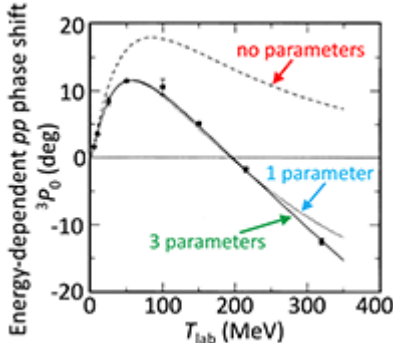
Stoks, Klomp, Terheggen, Swart
Phys. Rev. C **49**, 2950 (1994)



Wiringa, Stoks, Schiavilla
Phys. Rev. C **51**, 38 (1995)



R. Machleidt
Phys. Rev. C **63**, 024001 (2001)

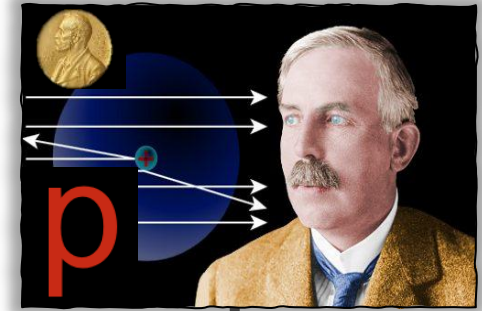


Stoks, Klomp, Rentmeester, Swart
Phys. Rev. C **48**, 792 (1993)

realistic

Milestones in nuclear force theories

Philos. Mag. 1919.



Neutron

1918

Proton



Nature, 1932.

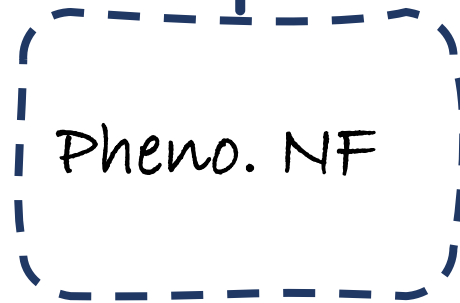
Proc. Phys. Math. Soc. Jap., 1935.



Meson theory

1935

1960s



AP, N.Y. 1968; NPA, 1976; ...

Phys. Rev. D, 1974.



Lattice QCD

1974

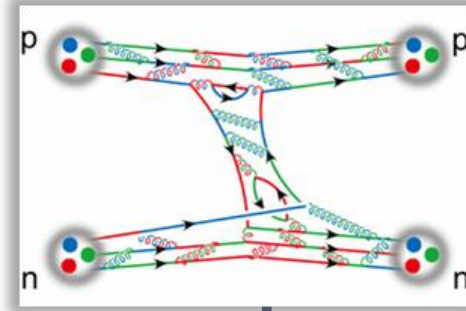


PLB, 1990; NPB, 1991; PLB, 1992.

ChEFT

1990

Phys. Rev. Lett 2022



rChEFT

2022

Advantages of chiral forces—compared with conventional ones

1

Closer connection with QCD

2

Systematically improvable

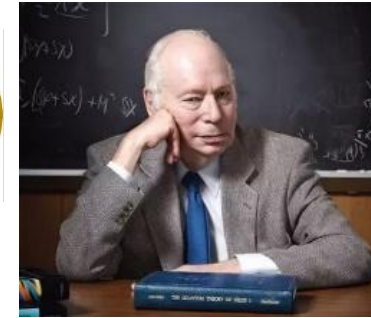
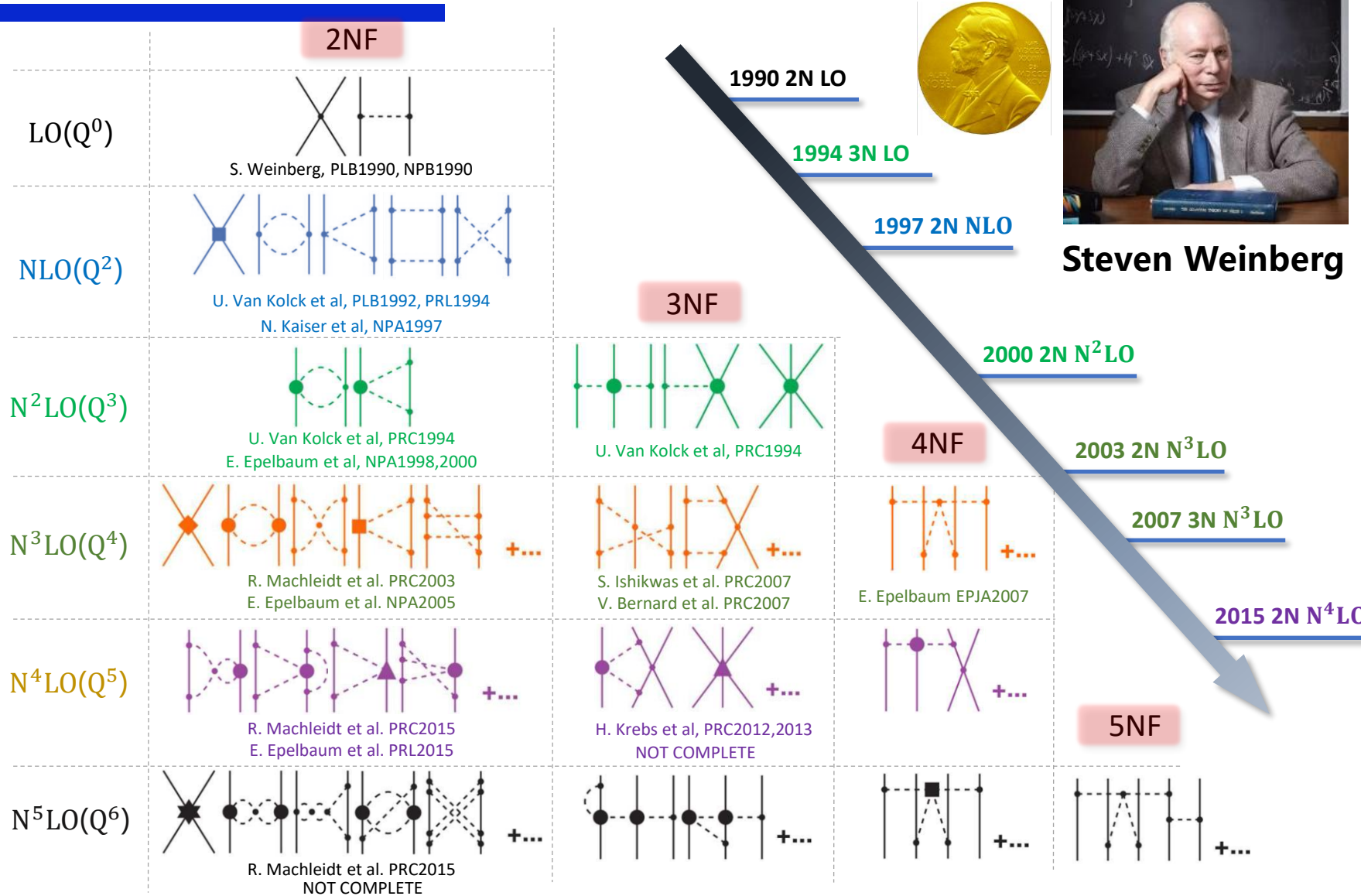
3

Self-consistent treatment of many-body forces

4

Uncertainties quantifiable

Modern de facto standard: chiral nuclear forces



Steven Weinberg



van Kolck



Kaiser



Epelbaum



Machleidt



Philipps

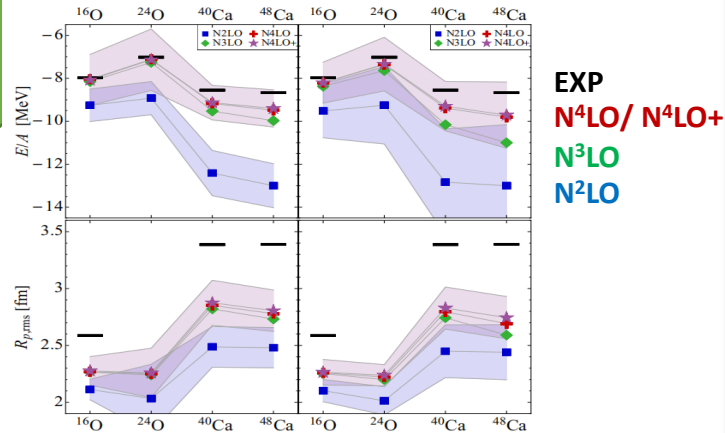


Savage

NR chiral 2N and 3N forces still need to refined

Slow convergence

$N^4LO?$



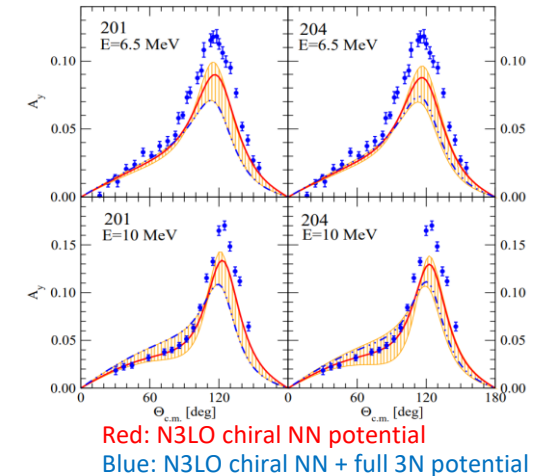
P. Maris et al. PRC2021

Ay puzzle

N-d polarization?

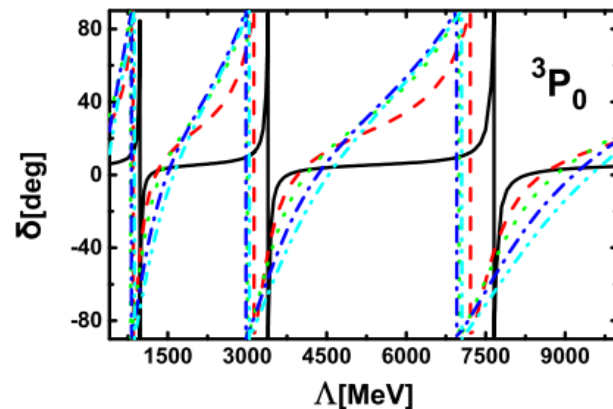
L. Girlanda et al. PRC2019

J. Golak et al. EPJA50,177



RG invariance

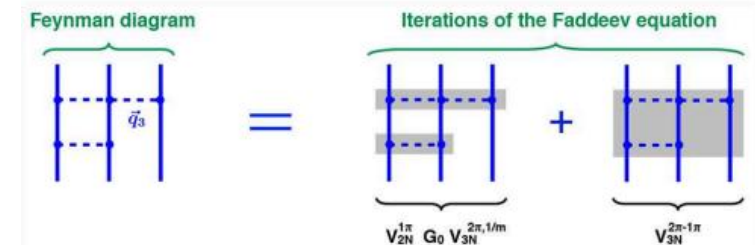
Cutoff dependent?



C.X. Wang et al. CPC2021

Self-consistency in 3N

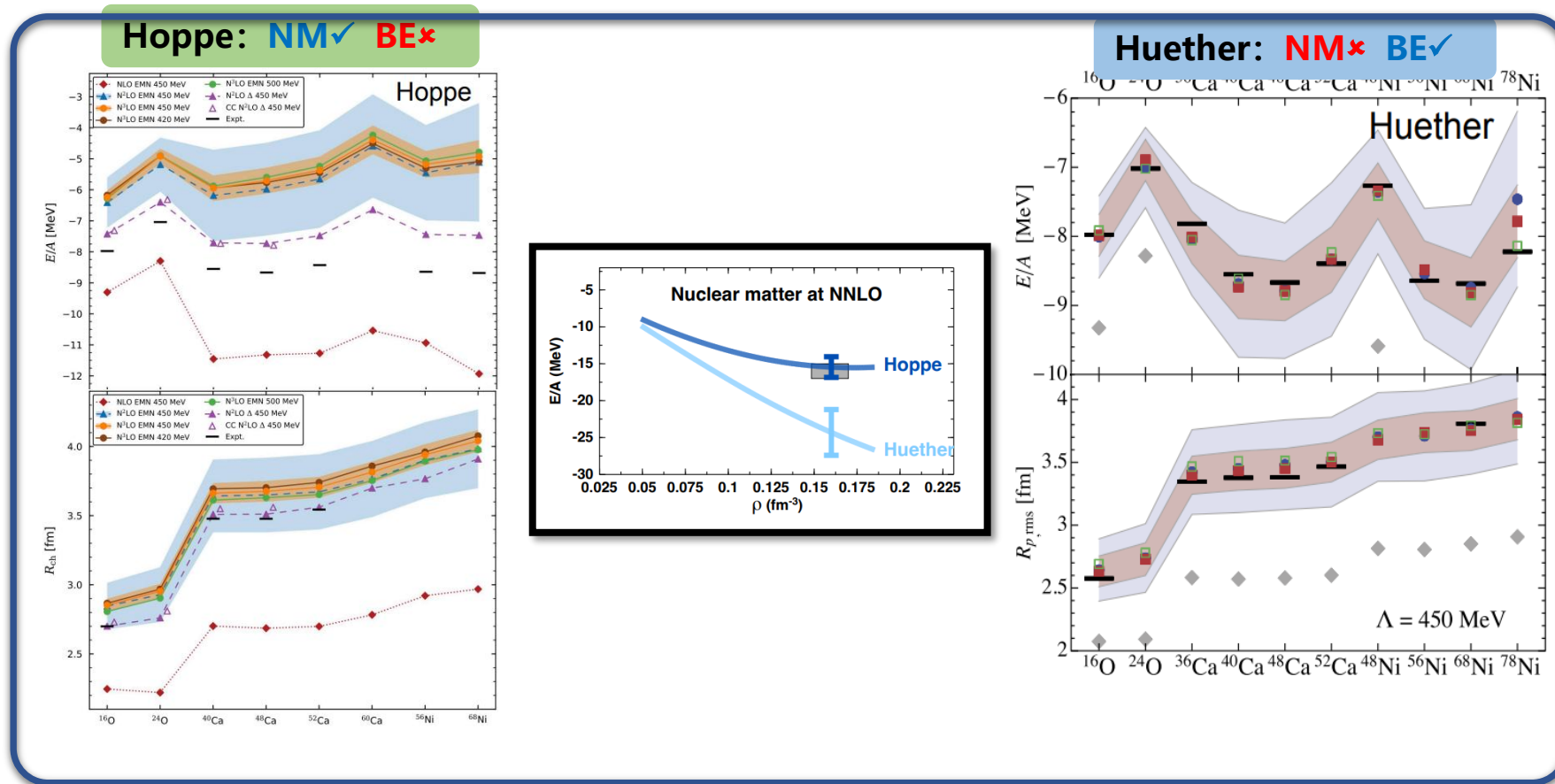
E. Epelbaum et al. fphy2020



N^3LO 3N force breaks chiral symmetry

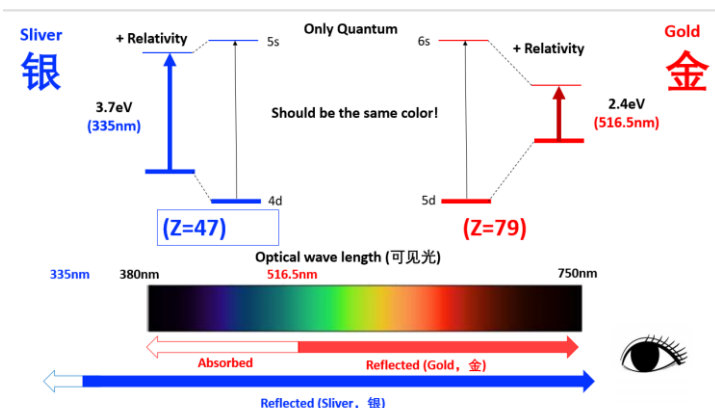
Many-body studies face self-consistency challenges

- Most are not truly ab initio, with two exceptions, but they can not simultaneously describe medium mass nuclei and nuclear matter

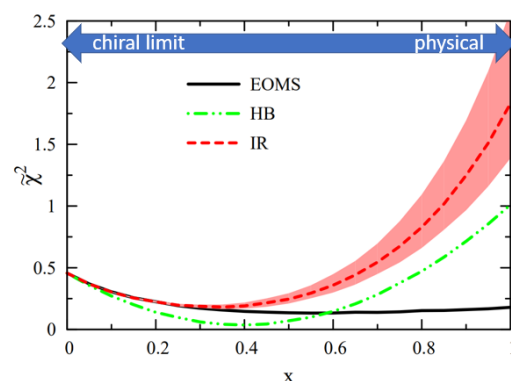


Why relativistic/covariant theories

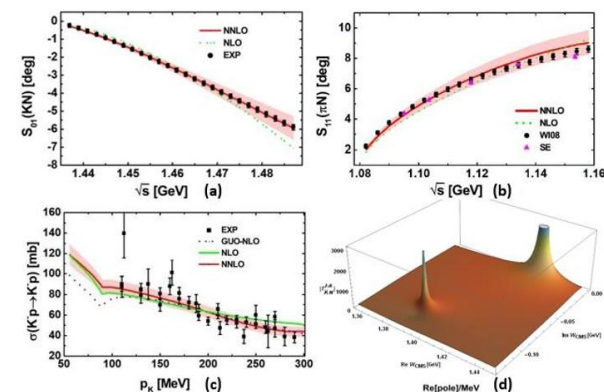
- ❑ Lorentz invariance is one of **the most essential symmetries** in Nature, where **SPIN naturally emerges**
- ❑ Both kinematical and dynamical relativistic corrections **self-consistently included**
- ❑ Relativistic approaches have been successful in explaining many **fine structures**
 - ✓ Atomic and molecular systems: why gold is yellow
 - ✓ Nuclear system: spin-orbit splitting, pseudospin symmetry
 - ✓ One-baryon sector: magnetic moments, masses, sigma terms
 - ✓ Meson-baryon scattering: πN 、 $\bar{K} N$ 、 $\Lambda(1405)$



Why Gold is yellow, and Silver is white



LSG et al., PRL101 (2008) 222002



LSG* et al, PRL130 (2023)071902

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How to become relativistic/covariant

□ **Dirac spinors and Clifford algebra** (instead of nonrelativistic wave functions and Pauli matrices)

$$u(\mathbf{p}, s) = N_p \left(\frac{1}{\epsilon_p} \begin{pmatrix} \boldsymbol{\sigma} \cdot \mathbf{p} \\ \epsilon_p \end{pmatrix} \right) \chi_s, N_p = \sqrt{\frac{\epsilon_p}{2M_N}}$$

	$\mathbb{1}$	γ_5	γ_μ	$\gamma_5 \gamma_\mu$	$\sigma_{\mu\nu}$	$\epsilon_{\mu\nu\rho\sigma}$	$\overleftrightarrow{\partial}_\mu$	∂_μ
$\mathcal{P}$	+	-	+	-	+	-	+	+
$\mathcal{C}$	+	+	-	+	-	+	-	+
h.c.	+	-	+	+	+	+	-	+
$\mathcal{O}$	0	1	0	0	0	-	0	1

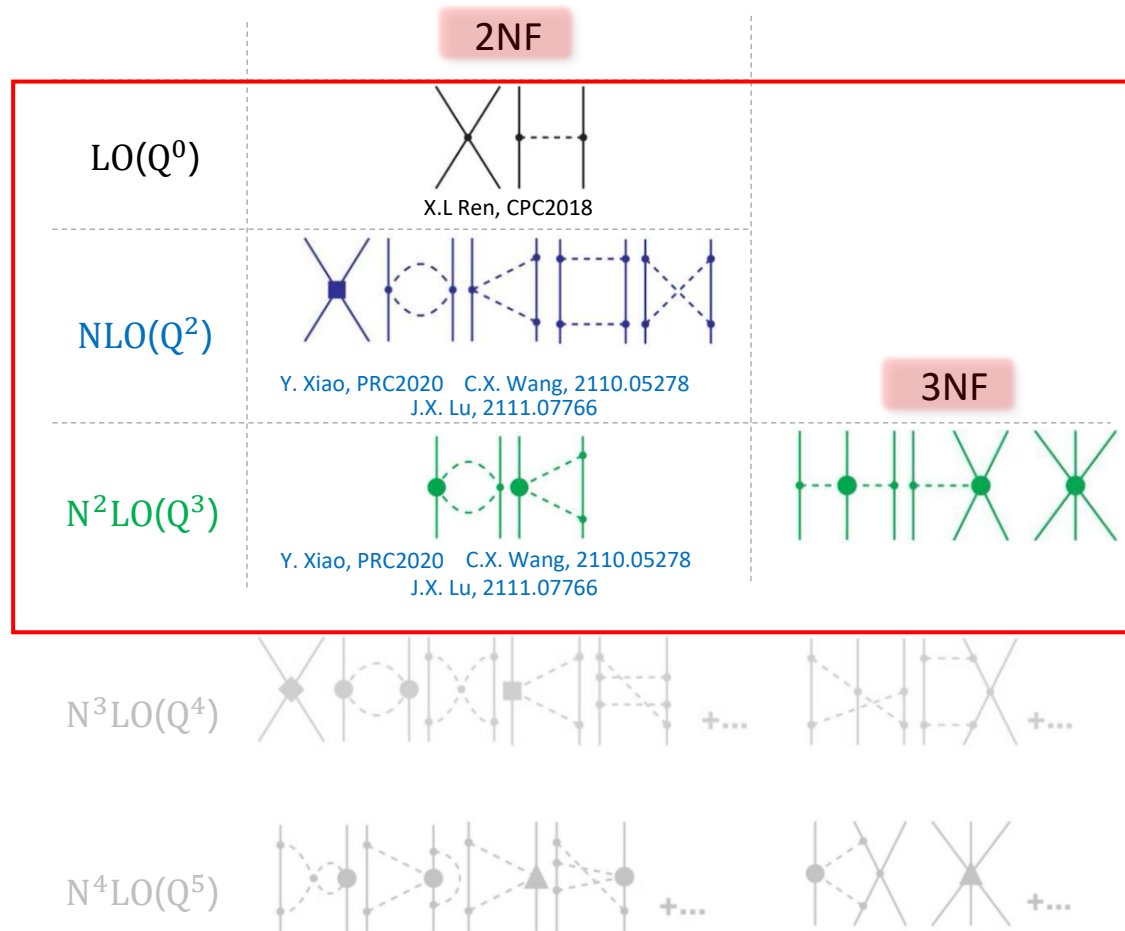
□ **Covariant power counting** (vs. the Weinberg PC)

□ **Covariant scattering equation** (instead of Lippmann-Schwinger eq.)

$$\mathcal{T}(p', p | W) = \mathcal{A}(p', p | W) + \int \frac{d^4 k}{(2\pi^4)} \mathcal{A}(p', k | W) G(k | W) \mathcal{T}(k, p | W)$$

$$G(k | W) = \frac{i}{[\gamma^\mu (W + k)_\mu - m_N + i\epsilon]^{(1)} [\gamma^\mu (W - k)_\mu - m_N + i\epsilon]^{(2)}}$$

Accurate relativistic chiral NN interaction up to **NNLO**



Three key inputs

1. Contact Lagrangians:

✓ **PRC99(2019)024004**

2. Meson-baryon vertices:

✓ **PRD99(2019)054024**

✓ **PRL130(2023)071902**

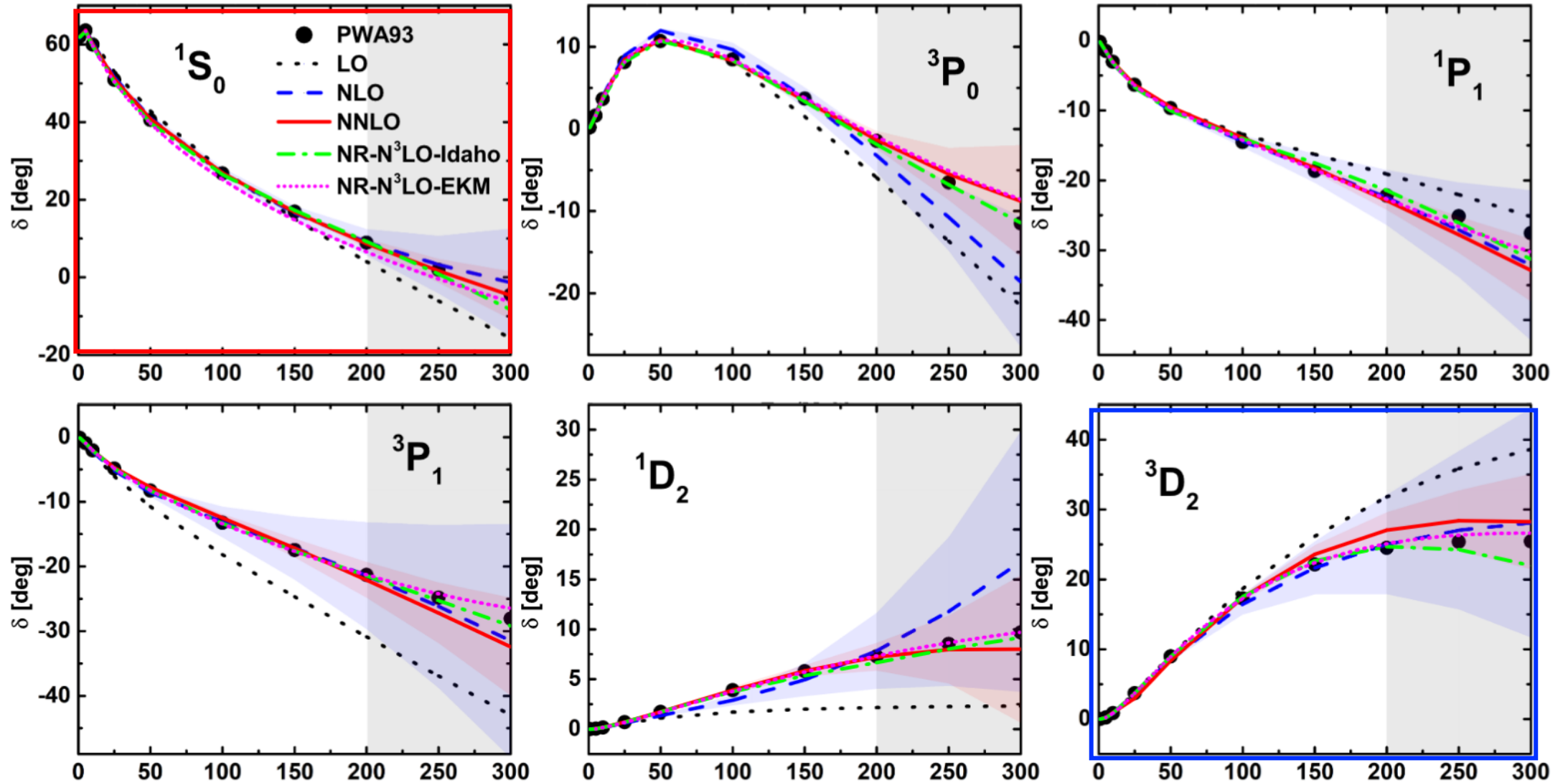
3. Two-meson exchanges

✓ **PRC102(2020)054001**

✓ **PRC105(2022)014003**

Fit results for $J \leq 2$ partial waves

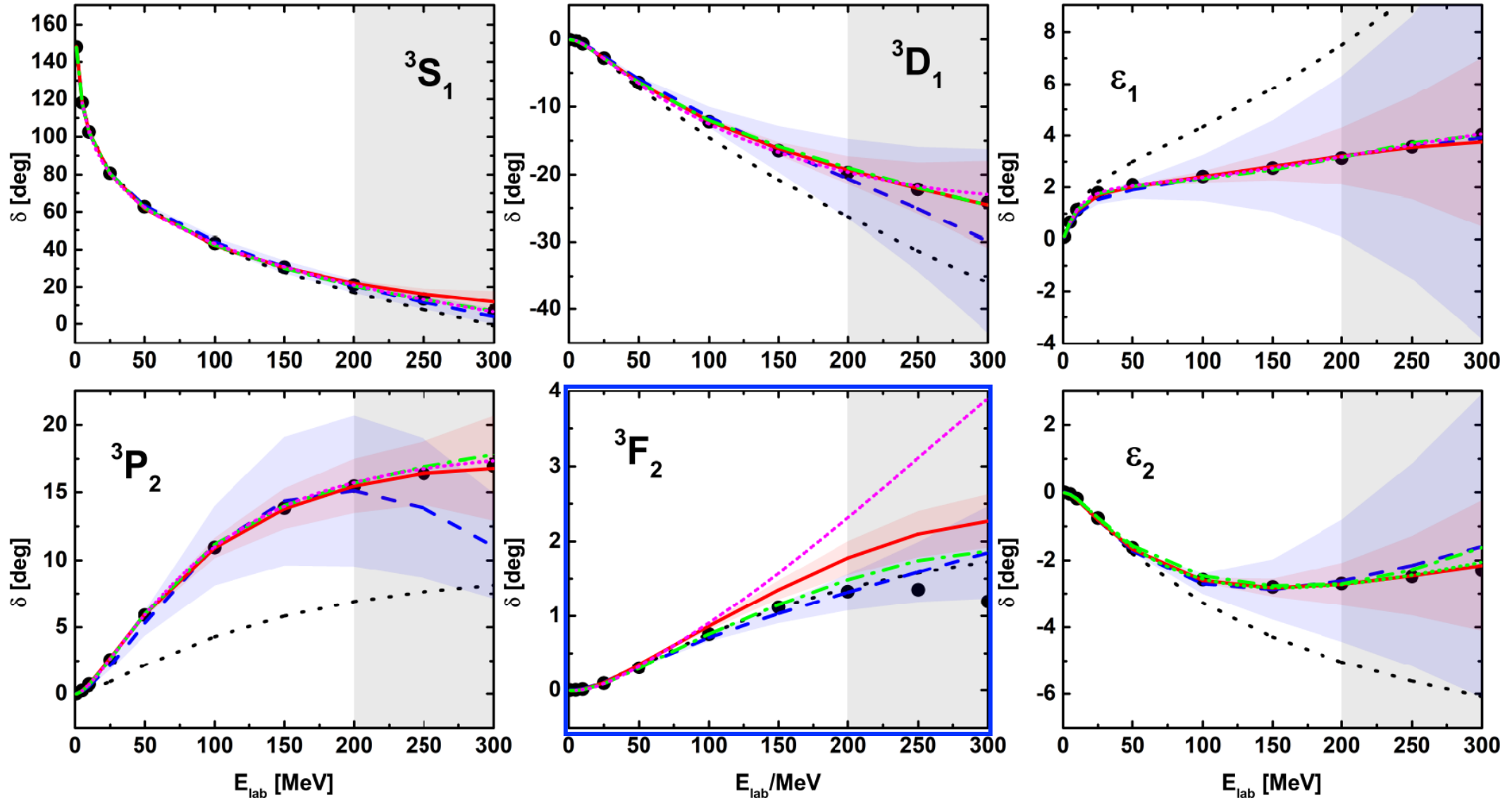
Jun-Xu Lu, Chun-Xuan Wang, Yang Xiao, LSG*, Jie Meng, and Peter Ring,
PRL128 (2022)142002



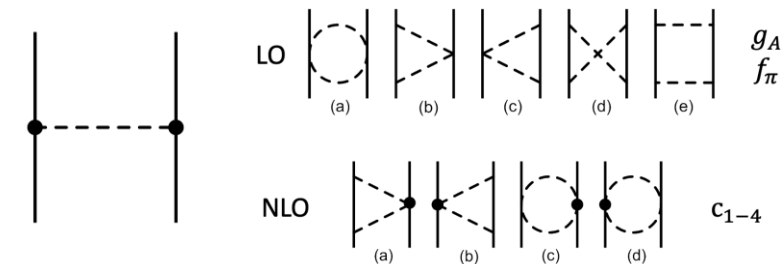
NR-N3LO-Idaho: Machleidt and D. R. Entem, *Phys. Rept.*503, 1 (2011); NR-N3LO-EKM: E. Epelbaum, H. Krebs, and U. G. Meißner, *Eur. Phys. J.*A51,53 (2015)

Fit results for $J \leq 2$ partial waves

Jun-Xu Lu, Chun-Xuan Wang, Yang Xiao, LSG*, Jie Meng,
and Peter Ring, PRL128 (2022)142002

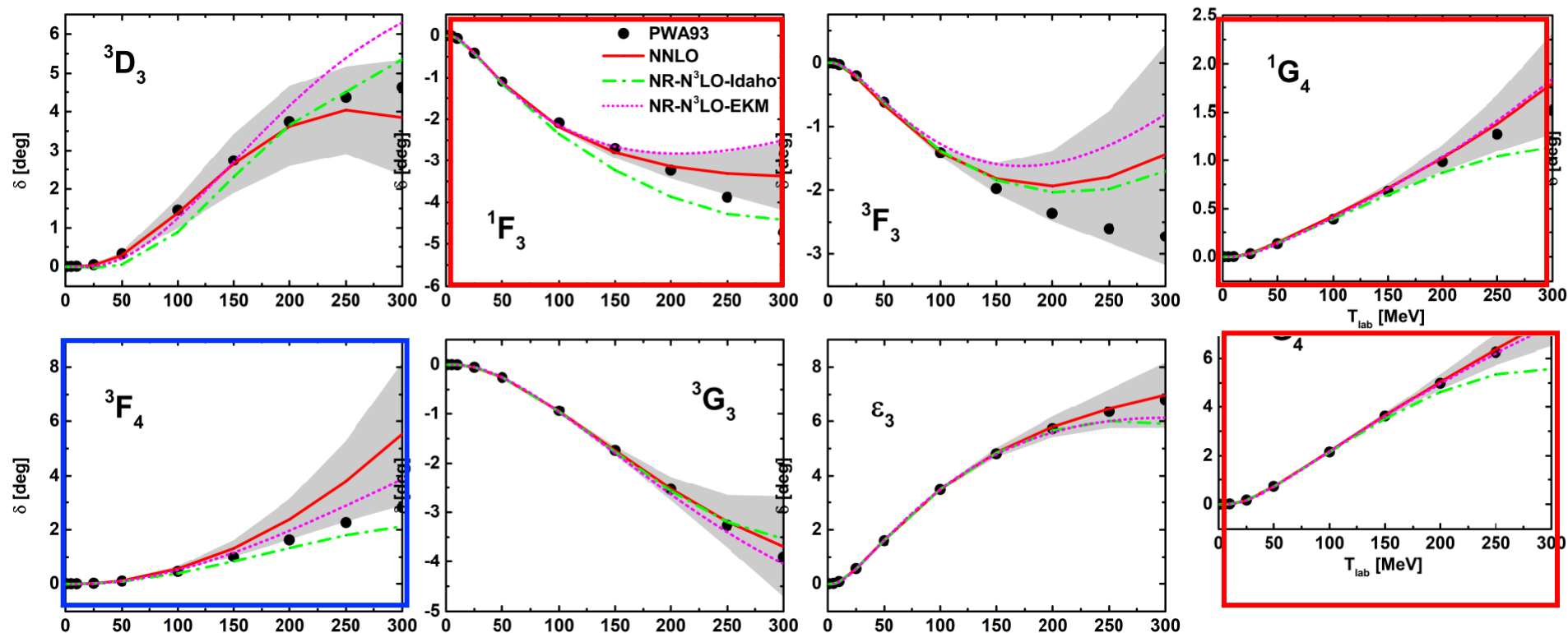


Higher partial waves: chiral dyn. & rela.

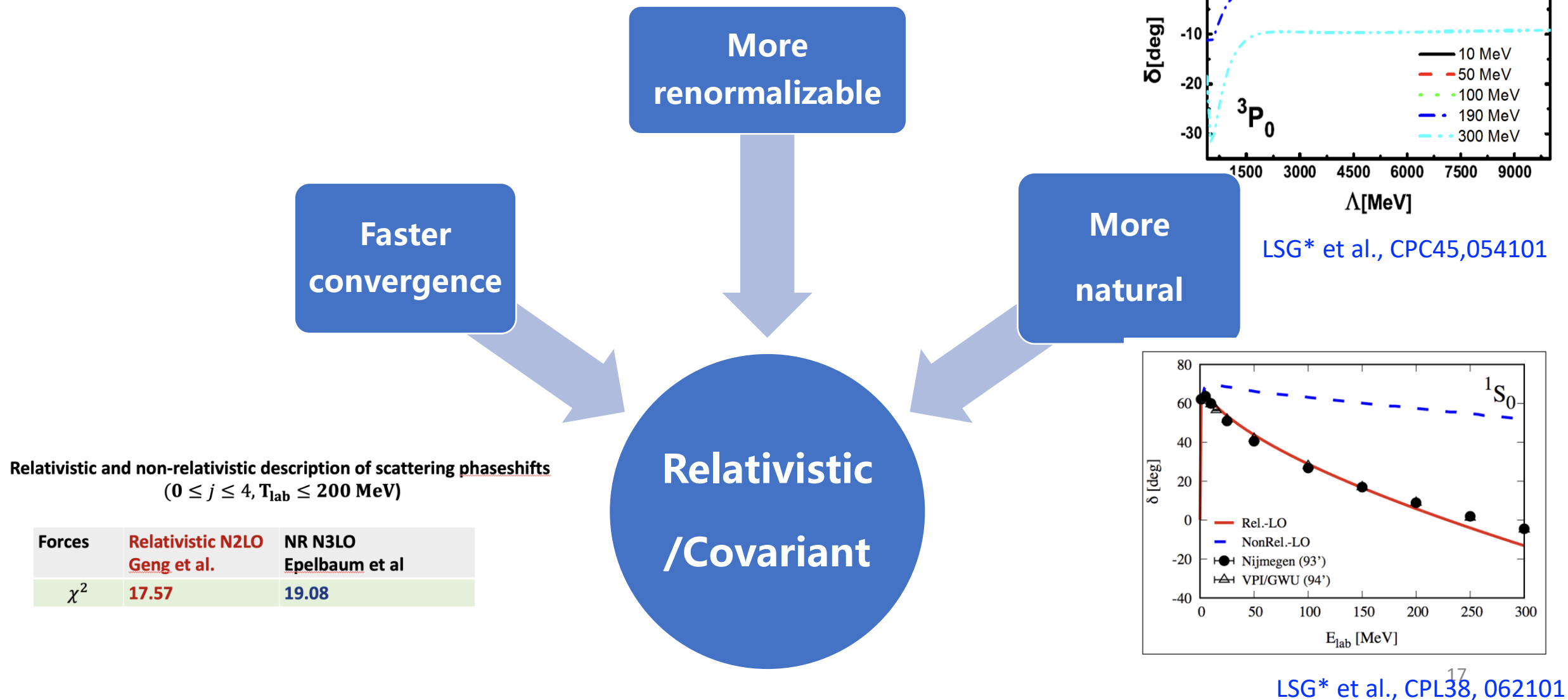


Test of relativistic treatment

	Total	3D_3	1F_3	3F_3	3F_4	3G_3	ϵ_3	1G_4	3G_4
NNLO	0.98	0.03	0.03	0.21	0.70	0.00	0.01	0.00	0.00
NR-N ³ LO-EKM	3.00	0.56	1.44	0.28	0.54	0.01	0.01	0.03	0.13



Three features of the relativistic force:



Status quo

We have constructed the first high-precision relativistic chiral nuclear force



1990/91 Weinberg



1994, Van Kolck



2003, Machleidt



2015, Epelbaum

Beijing
2022

非相对论手征核力

LO NN, Weinberg,
PLB251 (1990) 288

NN potential with Delta at **N2LO**, Ordonez,
Ray, **U. van Kolck**, PRL72 (1994) 1982

Accurate NN potential at **N3LO**,
Entem and Machleidt, PRC68(2003)041001



Precision NN potential at **N4LO**, **Epelbaum**,
Krebs, Meißner, PRL115 (2015)122301

相对论手征核力

Relativistic NN potential at **N2LO**, Lu, Wang, Xiao,
Geng*, Meng, Ring, PRL128 (2022) 142002

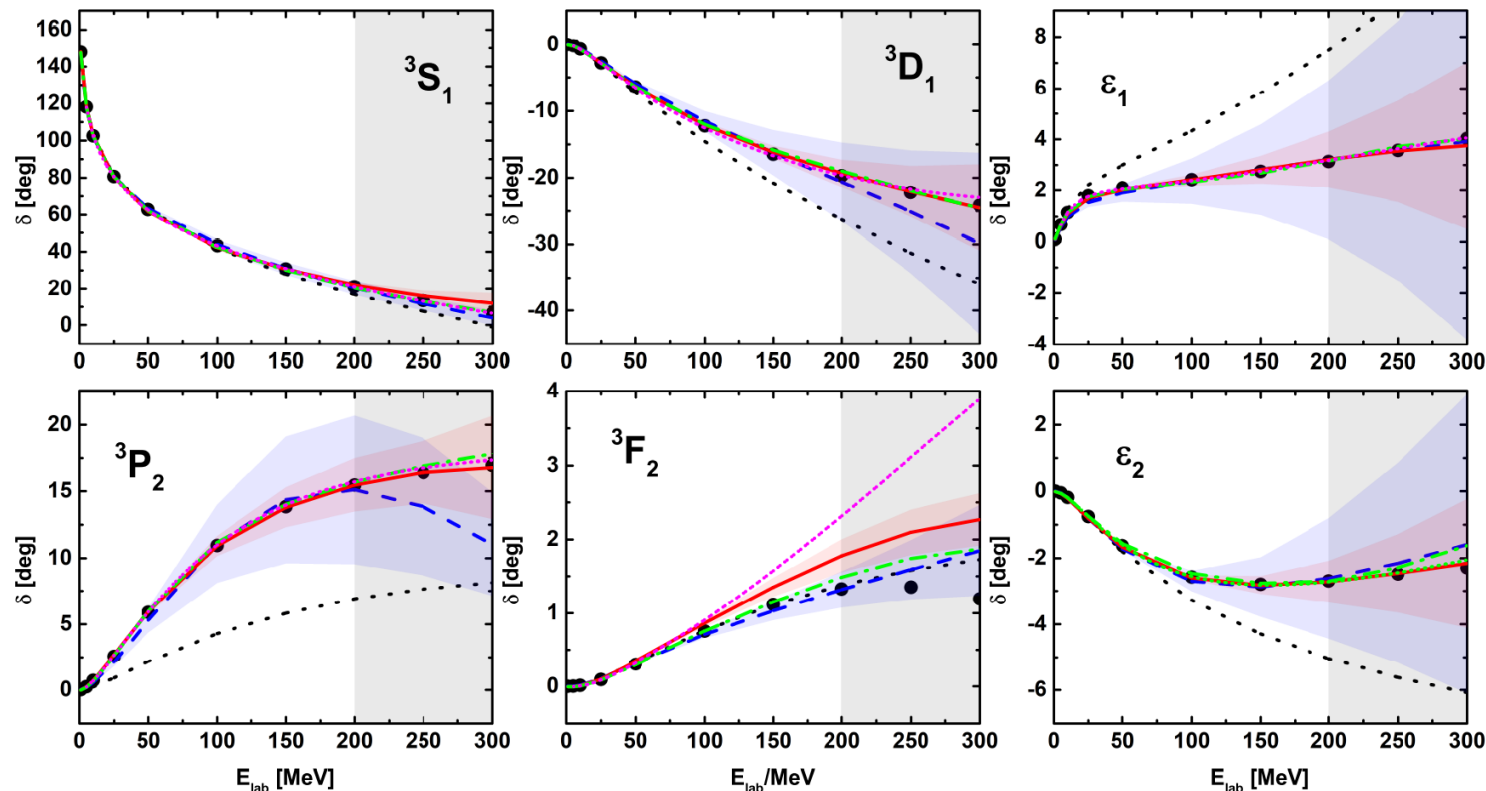
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Why even higher orders

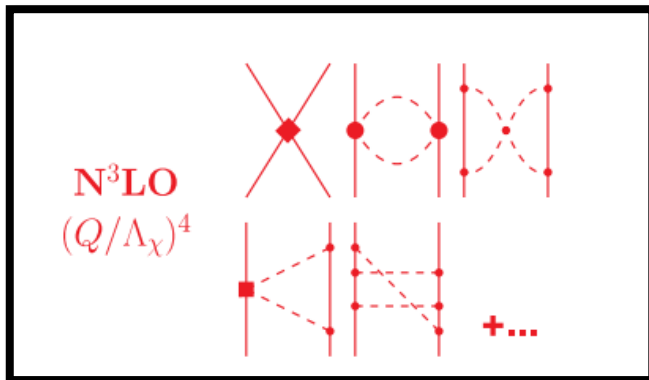
□ NR-ChNF: N⁴LO is complete and partial N5LO already exists!

□ R-ChNF: N3LO is needed to check the chiral convergence and describe better high-energy scattering data.



e.g., 3F2

Up to $N^3\text{LO}$



- ✓ Contact term (23 terms)
- ✓ One Pion completely absorbed
- ✓ Two Pion One + Two loop
- ✓ Three Pion negligible

$$V_{1\pi} = V_{1\pi}^{(0)} + V_{1\pi}^{(2)} + V_{1\pi}^{(3)} + V_{1\pi}^{(4)} + \dots$$

$$V_{2\pi} = V_{2\pi}^{(2)} + V_{2\pi}^{(3)} + V_{2\pi}^{(4)} + \dots$$

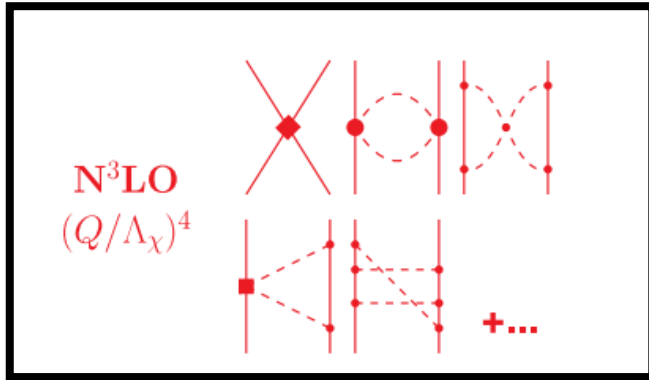
$$V_{3\pi} = V_{3\pi}^{(4)} + \dots,$$

+all isospin breaking effects

➤ Contact terms (23 terms)

$\tilde{O}_1$	$(\bar{\psi}\psi)(\bar{\psi}\psi)$	$\tilde{O}_{21}$	$\frac{1}{16m^4}(\bar{\psi}i\overleftrightarrow{\partial}^\mu\psi)\partial^2\partial^\nu(\bar{\psi}\sigma_{\mu\nu}\psi)$
$\tilde{O}_2$	$(\bar{\psi}\gamma^\mu\psi)(\bar{\psi}\gamma_\mu\psi)$	$\tilde{O}_{22}$	$\frac{1}{16m^4}(\bar{\psi}\sigma^{\mu\alpha}\psi)\partial^2\partial_\alpha\partial^\nu(\bar{\psi}\sigma_{\mu\nu}\psi)$
$\tilde{O}_3$	$(\bar{\psi}\gamma_5\gamma^\mu\psi)(\bar{\psi}\gamma_5\gamma_\mu\psi)$	$\tilde{O}_{23}$	$\frac{1}{16m^4}(\bar{\psi}\sigma^{\mu\nu}i\overleftrightarrow{\partial}^\alpha\psi)\partial^\beta\partial_\nu(\bar{\psi}\sigma_{\alpha\beta}i\overleftrightarrow{\partial}_\mu\psi)$
$\tilde{O}_4$	$(\bar{\psi}\sigma^{\mu\nu}\psi)(\bar{\psi}\sigma_{\mu\nu}\psi)$	$\tilde{O}_{24}$	$\frac{1}{16m^4}(\bar{\psi}\psi)\partial^4(\bar{\psi}\psi)$
$\tilde{O}_5$	$(\bar{\psi}\gamma_5\psi)(\bar{\psi}\gamma_5\psi)$	$\tilde{O}_{25}$	$\frac{1}{16m^4}(\bar{\psi}\gamma^\mu\psi)\partial^4(\bar{\psi}\gamma_\mu\psi)$
$\tilde{O}_6$	$\frac{1}{4m^2}(\bar{\psi}\gamma_5\gamma^\mu i\overleftrightarrow{\partial}^\alpha\psi)(\bar{\psi}\gamma_5\gamma_\alpha i\overleftrightarrow{\partial}_\mu\psi)$	$\tilde{O}_{26}$	$\frac{1}{16m^4}(\bar{\psi}\gamma_5\gamma^\mu\psi)\partial^4(\bar{\psi}\gamma_5\gamma_\mu\psi)$
$\tilde{O}_7$	$\frac{1}{4m^2}(\bar{\psi}\sigma^{\mu\nu}i\overleftrightarrow{\partial}^\alpha\psi)(\bar{\psi}\sigma_{\mu\alpha}i\overleftrightarrow{\partial}_\nu\psi)$	$\tilde{O}_{27}$	$\frac{1}{16m^4}(\bar{\psi}\sigma^{\mu\nu}\psi)\partial^4(\bar{\psi}\sigma_{\mu\nu}\psi)$
$\tilde{O}_8$	$\frac{1}{4m^2}(\bar{\psi}i\overleftrightarrow{\partial}^\mu\psi)\partial^\nu(\bar{\psi}\sigma_{\mu\nu}\psi)$	$\tilde{O}_{28}$	$\frac{1}{4m^2}(\bar{\psi}\gamma_5 i\overleftrightarrow{\partial}^\alpha\psi)(\bar{\psi}\gamma_5 i\overleftrightarrow{\partial}_\alpha\psi) - \tilde{O}_5$
$\tilde{O}_9$	$\frac{1}{4m^2}(\bar{\psi}\sigma^{\mu\alpha}\psi)\partial_\alpha\partial^\nu(\bar{\psi}\sigma_{\mu\nu}\psi)$	$\tilde{O}_{29}$	$\frac{1}{16m^4}(\bar{\psi}\gamma_5\gamma^\mu i\overleftrightarrow{\partial}^\alpha i\overleftrightarrow{\partial}^\beta\psi)(\bar{\psi}\gamma_5\gamma_\alpha i\overleftrightarrow{\partial}_\mu i\overleftrightarrow{\partial}_\beta\psi) - \tilde{O}_6$
$\tilde{O}_{10}$	$\frac{1}{4m^2}(\bar{\psi}\psi)\partial^2(\bar{\psi}\psi)$	$\tilde{O}_{30}$	$\frac{1}{16m^4}(\bar{\psi}\sigma^{\mu\nu}i\overleftrightarrow{\partial}^\alpha i\overleftrightarrow{\partial}^\beta\psi)(\bar{\psi}\sigma_{\mu\alpha}i\overleftrightarrow{\partial}_\nu i\overleftrightarrow{\partial}_\beta\psi) - \tilde{O}_7$
$\tilde{O}_{11}$	$\frac{1}{4m^2}(\bar{\psi}\gamma^\mu\psi)\partial^2(\bar{\psi}\gamma_\mu\psi)$	$\tilde{O}_{31}$	$\frac{1}{16m^4}(\bar{\psi}i\overleftrightarrow{\partial}^\mu i\overleftrightarrow{\partial}^\beta\psi)\partial^\alpha(\bar{\psi}\sigma_{\mu\alpha}i\overleftrightarrow{\partial}_\beta\psi) - \tilde{O}_8$
$\tilde{O}_{12}$	$\frac{1}{4m^2}(\bar{\psi}\gamma_5\gamma^\mu\psi)\partial^2(\bar{\psi}\gamma_5\gamma_\mu\psi)$	$\tilde{O}_{32}$	$\frac{1}{16m^4}(\bar{\psi}\sigma^{\mu\alpha}i\overleftrightarrow{\partial}^\beta\psi)\partial_\alpha\partial^\nu(\bar{\psi}\sigma_{\mu\nu}i\overleftrightarrow{\partial}_\beta\psi) - \tilde{O}_9$
$\tilde{O}_{13}$	$\frac{1}{4m^2}(\bar{\psi}\sigma^{\mu\nu}\psi)\partial^2(\bar{\psi}\sigma_{\mu\nu}\psi)$	$\tilde{O}_{33}$	$\frac{1}{16m^4}(\bar{\psi}i\overleftrightarrow{\partial}^\alpha\psi)\partial^2(\bar{\psi}i\overleftrightarrow{\partial}_\alpha\psi) - \tilde{O}_{10}$
$\tilde{O}_{14}$	$\frac{1}{4m^2}(\bar{\psi}i\overleftrightarrow{\partial}^\alpha\psi)(\bar{\psi}i\overleftrightarrow{\partial}_\alpha\psi) - \tilde{O}_1$	$\tilde{O}_{34}$	$\frac{1}{16m^4}(\bar{\psi}\gamma^\mu i\overleftrightarrow{\partial}^\alpha\psi)\partial^2(\bar{\psi}\gamma_\mu i\overleftrightarrow{\partial}_\alpha\psi) - \tilde{O}_{11}$
$\tilde{O}_{15}$	$\frac{1}{4m^2}(\bar{\psi}\gamma^\mu i\overleftrightarrow{\partial}^\alpha\psi)(\bar{\psi}\gamma_\mu i\overleftrightarrow{\partial}_\alpha\psi) - \tilde{O}_2$	$\tilde{O}_{35}$	$\frac{1}{16m^4}(\bar{\psi}\gamma_5\gamma^\mu i\overleftrightarrow{\partial}^\alpha\psi)\partial^2(\bar{\psi}\gamma_5\gamma_\mu i\overleftrightarrow{\partial}_\alpha\psi) - \tilde{O}_{12}$
$\tilde{O}_{16}$	$\frac{1}{4m^2}(\bar{\psi}\gamma_5\gamma^\mu i\overleftrightarrow{\partial}^\alpha\psi)(\bar{\psi}\gamma_5\gamma_\mu i\overleftrightarrow{\partial}_\alpha\psi) - \tilde{O}_3$	$\tilde{O}_{36}$	$\frac{1}{16m^4}(\bar{\psi}\sigma^{\mu\nu}i\overleftrightarrow{\partial}^\alpha\psi)\partial^2(\bar{\psi}\sigma_{\mu\nu}i\overleftrightarrow{\partial}_\alpha\psi) - \tilde{O}_{13}$
$\tilde{O}_{17}$	$\frac{1}{4m^2}(\bar{\psi}\sigma^{\mu\nu}i\overleftrightarrow{\partial}^\alpha\psi)(\bar{\psi}\sigma_{\mu\nu}i\overleftrightarrow{\partial}_\alpha\psi) - \tilde{O}_4$	$\tilde{O}_{37}$	$\frac{1}{16m^4}(\bar{\psi}i\overleftrightarrow{\partial}^\alpha i\overleftrightarrow{\partial}^\beta\psi)(\bar{\psi}i\overleftrightarrow{\partial}_\alpha i\overleftrightarrow{\partial}_\beta\psi) - 2\tilde{O}_{14} - \tilde{O}_1$
$\tilde{O}_{18}$	$\frac{1}{4m^2}(\bar{\psi}\gamma_5\psi)\partial^2(\bar{\psi}\gamma_5\psi)$	$\tilde{O}_{38}$	$\frac{1}{16m^4}(\bar{\psi}\gamma^\mu i\overleftrightarrow{\partial}^\alpha i\overleftrightarrow{\partial}^\beta\psi)(\bar{\psi}\gamma_\mu i\overleftrightarrow{\partial}_\alpha i\overleftrightarrow{\partial}_\beta\psi) - 2\tilde{O}_{15} - \tilde{O}_2$
$\tilde{O}_{19}$	$\frac{1}{16m^4}(\bar{\psi}\gamma_5\gamma^\mu i\overleftrightarrow{\partial}^\nu\psi)\partial^2(\bar{\psi}\gamma_5\gamma_\nu i\overleftrightarrow{\partial}_\mu\psi)$	$\tilde{O}_{39}$	$\frac{1}{16m^4}(\bar{\psi}\gamma_5\gamma^\mu i\overleftrightarrow{\partial}^\alpha i\overleftrightarrow{\partial}^\beta\psi)(\bar{\psi}\gamma_5\gamma_\mu i\overleftrightarrow{\partial}_\alpha i\overleftrightarrow{\partial}_\beta\psi) - 2\tilde{O}_{16} - \tilde{O}_3$
$\tilde{O}_{20}$	$\frac{1}{16m^4}(\bar{\psi}\sigma^{\mu\nu}i\overleftrightarrow{\partial}^\alpha\psi)\partial^2(\bar{\psi}\sigma_{\mu\alpha}i\overleftrightarrow{\partial}_\nu\psi)$	$\tilde{O}_{40}$	$\frac{1}{16m^4}(\bar{\psi}\sigma^{\mu\nu}i\overleftrightarrow{\partial}^\alpha i\overleftrightarrow{\partial}^\beta\psi)(\bar{\psi}\sigma_{\mu\nu}i\overleftrightarrow{\partial}_\alpha i\overleftrightarrow{\partial}_\beta\psi) - 2\tilde{O}_{17} - \tilde{O}_4$

Up to $N^3\text{LO}$



- ✓ Contact term (23 terms)
- ✓ One Pion completely absorbed
- ✓ Two Pion One + Two loop
- ✓ Three Pion negligible

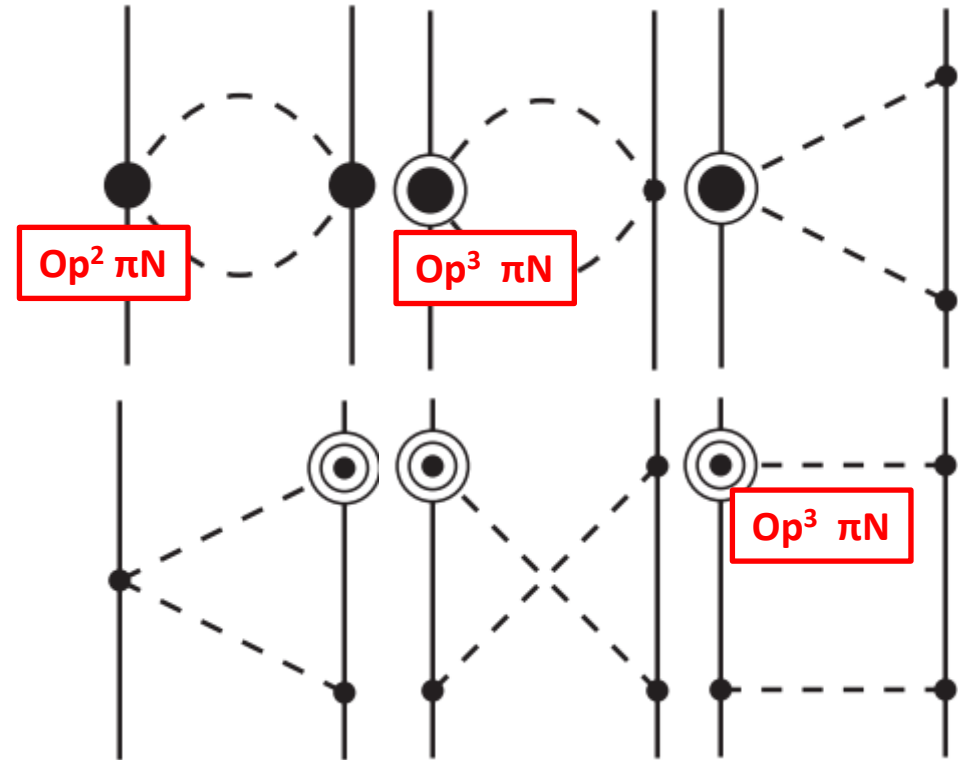
$$V_{1\pi} = V_{1\pi}^{(0)} + V_{1\pi}^{(2)} + V_{1\pi}^{(3)} + V_{1\pi}^{(4)} + \dots$$

$$V_{2\pi} = V_{2\pi}^{(2)} + V_{2\pi}^{(3)} + V_{2\pi}^{(4)} + \dots$$

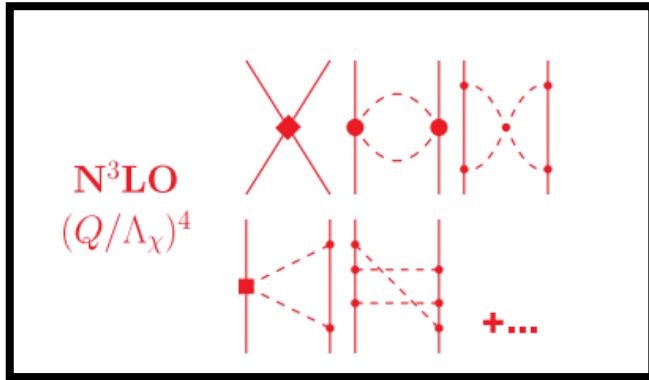
$$V_{3\pi} = V_{3\pi}^{(4)} + \dots,$$

+all isospin breaking effects

➤ TPE (One-loop part)



Up to $N^3\text{LO}$



- ✓ Contact term **(23 terms)**
- ✓ One Pion **completely absorbed**
- ✓ Two Pion **One + Two loop**
- ✓ Three Pion **negligible**

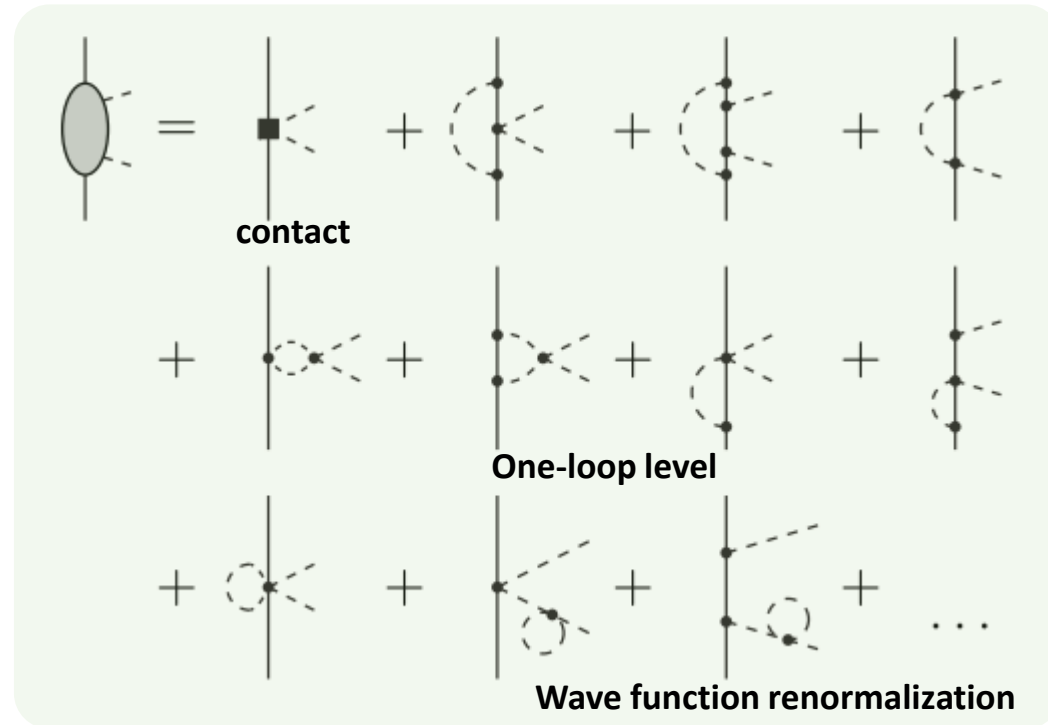
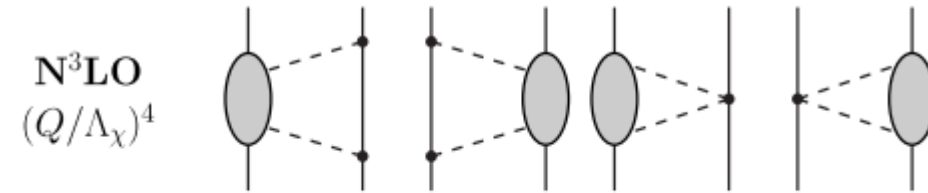
$$V_{1\pi} = V_{1\pi}^{(0)} + V_{1\pi}^{(2)} + V_{1\pi}^{(3)} + V_{1\pi}^{(4)} + \dots$$

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$$V_{3\pi} = V_{3\pi}^{(4)} + \dots,$$

+all isospin breaking effects

➤ TPE (two-loop part)



Relativistic two-loop diagrams—the most tough task

□ Dimensional regularization

- Convenient / straightforward
- **Full amplitude**—non-analytical part and polynomial part
- The loop momentum—pion momentum are integrated up to infinity
- Too complicated to be used for two-loop, especially in NF involving massive fermions

So far, the two-loop contributions to NF with the DR method remain untouched

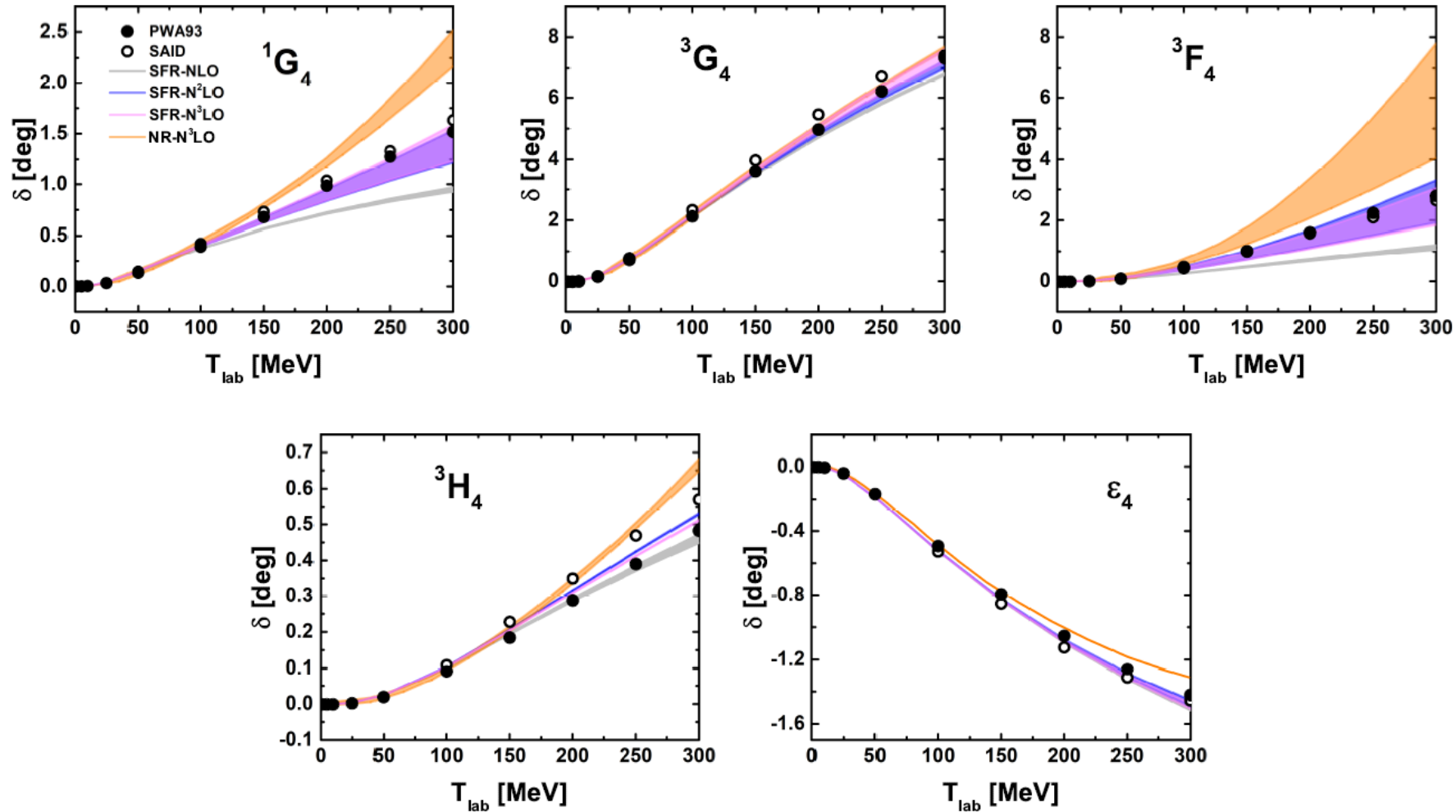
□ Spectral representation

- **Dispersion relation** / Mandelstam representation
- Mostly **only** non-analytical parts due to subtraction
- Convenient to introduce cutoffs to the loop momentum
- From the experience of NR, it should be applicable for two-loops in the relativistic case as well.

It has become the common choice in the newly developed NR nuclear forces

- **Semi-local N⁴LO NR-NF** Reinert et al. *Eur.Phys.J.A* 54, 86 (2018)
- **Local position-space N³LO NR-NF** Saha et al. *Phys.Rev.C* 107, 034002 (2023)

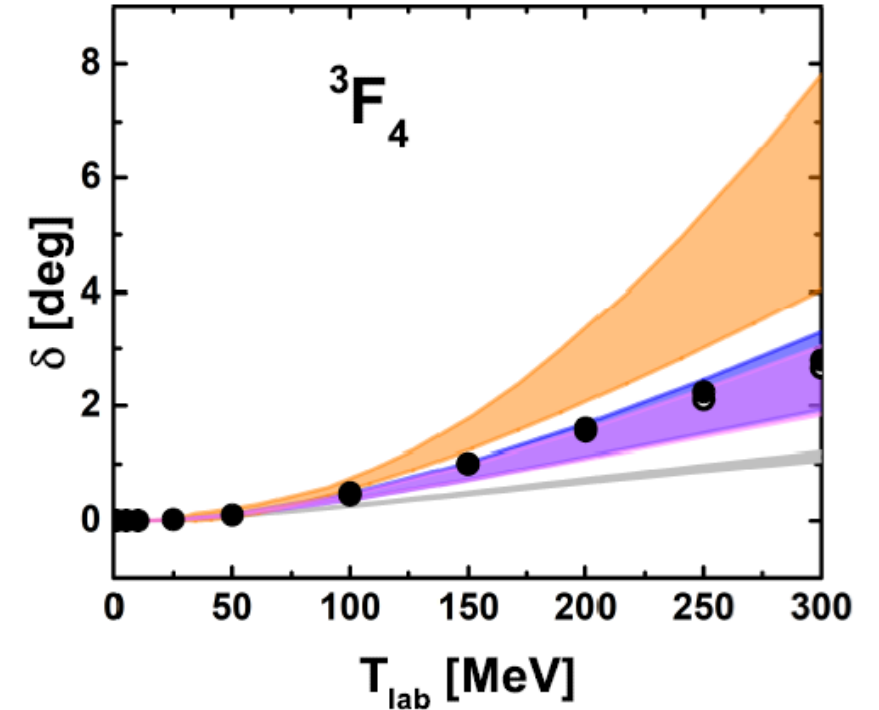
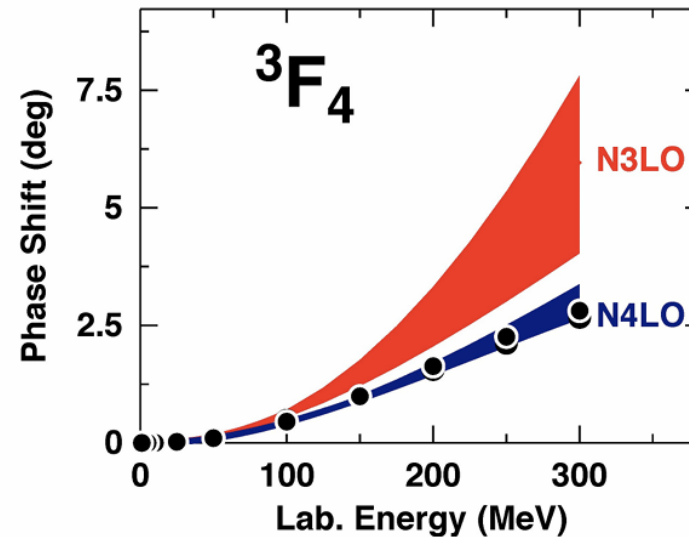
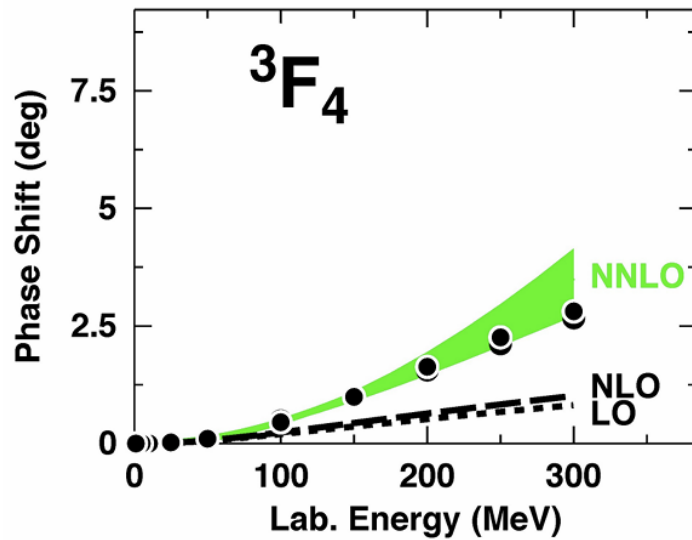
NLO&N2LO&N3LO SFR results—J=4



The convergence is very good! Better than NR

In comparison with the NR results

The Relativistic result



Contents

- ☞ Why relativistic chiral nuclear force
- ☞ The first high-precision relativistic chiral nuclear force
- ☞ Recent developments and works in progress
 - *Progress in constructing higher-order NN and **Ay puzzle***
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- ☞ Summary and outlook

The A_y Puzzle in Nucleon–Deuteron Scattering

➤ Emergence of the A_y puzzle: failure of high-precision NN forces [A. Kievsky et al, *PRC* 52 \(1995\) R15](#)

➤ Benchmark calculations with realistic NN interactions

[W. Glöckle et al, *Physics Reports* 274 \(1996\) 107](#)

[A. Kievsky et al, *PRC* 58 \(1998\) 3085](#)

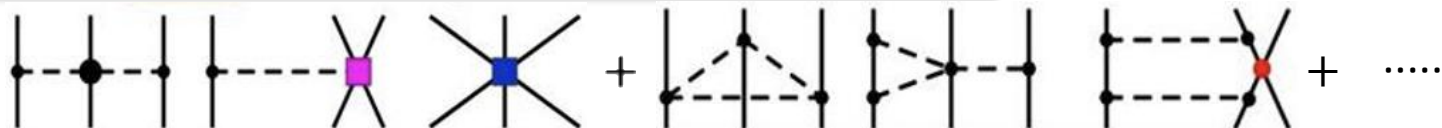
➤ Identification of the origin: 3P_J problem [D. Hüber et al, *PRC* 58, 674 \(1998\)](#)

➤ Attempts with conventional three-nucleon forces [C.R. Brune et al, *PLB* 428 \(1998\) 13-17](#)

- Tucson–Melbourne (TM/TM99)
- Urbana IX
- Brazil force

➤ Chiral three-nucleon forces and remaining A_y discrepancy

[L. Girlanda et al, *PRC* 99 \(2019\) 5, 054003](#)

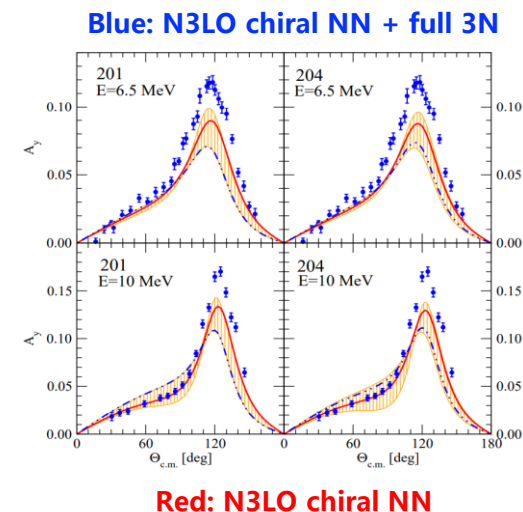


➤ Beyond conventional 3NF: new mechanisms

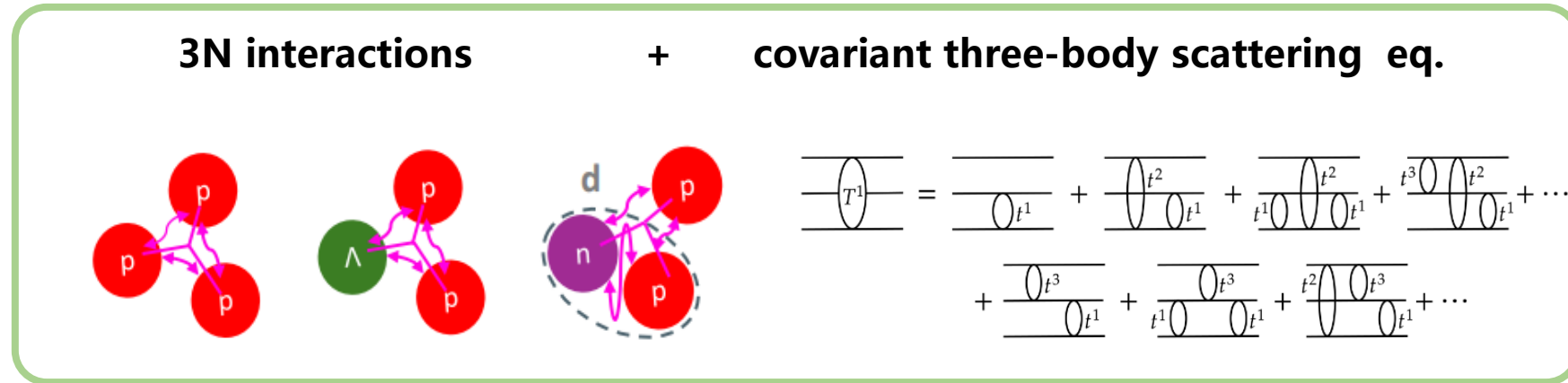
- Modified NN P -waves [D.R. Entem et al, *PRC* 65 \(2002\) 064005](#)

- Relativistic effects [H. Witała et al, *PRC* 77 \(2008\) 034004](#)

- Higher-order chiral three-nucleon forces [E. Epelbaum et al, *EPJA* 56 \(2020\) 92](#)



3B theoretical framework – the nonrelativistic Faddeev equation



$$t_i = v_i + v_i G_0 t_i$$

$$\langle \vec{p}_i \vec{q}_i | t_i(E) | \vec{p}_i' \vec{q}_i' \rangle = \delta(\vec{q}_i - \vec{q}_i') \langle \vec{p}_i | t_i(E - q_i^2 / 2\tilde{\mu}_i) | \vec{p}_i' \rangle$$

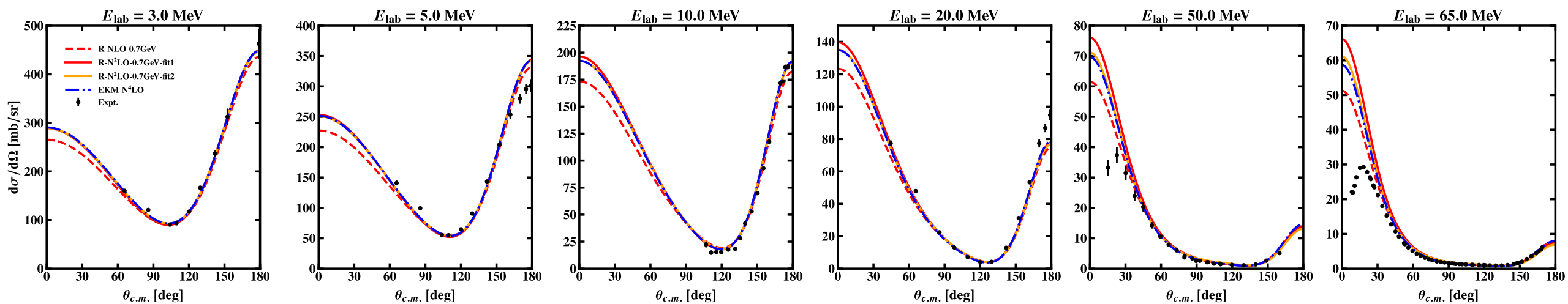
$$\begin{pmatrix} T^{(i)} \\ T^{(j)} \\ T^{(k)} \end{pmatrix} = \begin{pmatrix} t_i \\ t_j \\ t_k \end{pmatrix} + \begin{pmatrix} 0 & t_i & t_i \\ t_j & 0 & t_j \\ t_k & t_k & 0 \end{pmatrix} G_0 \begin{pmatrix} T^{(i)} \\ T^{(j)} \\ T^{(k)} \end{pmatrix}$$



L.D. Faddeev
23 Mar. 1934 – 26 Feb. 2017

First study of the nD scattering in the relativistic framework (preliminary)

□ The differential cross section —only two-body force up to J=2



✓ Results from R-NNLO and EKM-N4LO almost overlap. Differences emerge at extremely small angles.

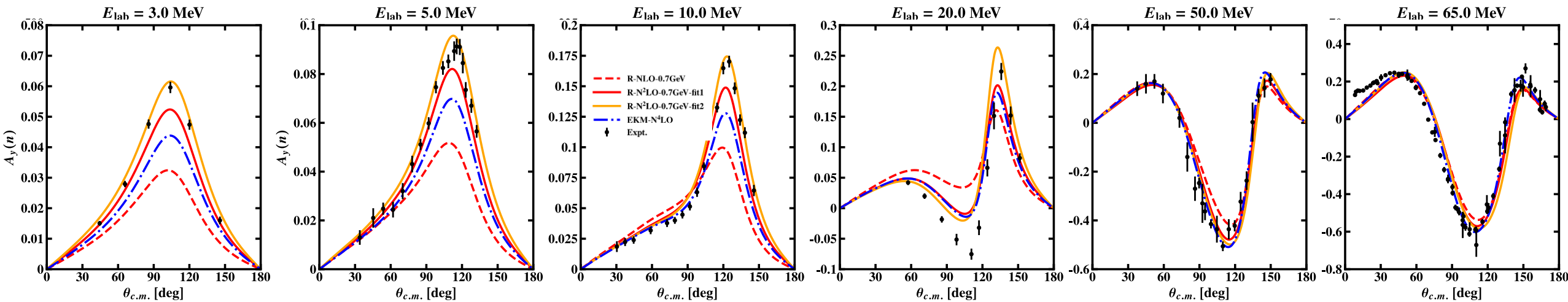
□ Further comparison for tritium binding energy

EKM – N ³ LO	isospin breaking	7.37145 MeV
	isospin symmetry	7.53719 MeV
CChNF – N ² LO	isospin symmetry	8.30648 MeV

**Exp.
8.4818**

First study of the nD scattering in the relativistic framework (**preliminary**)

□ The polarization A_y for n-d scattering—only two-body force up to $J=2$



- ✓ The description of A_y is significantly improved up to NNLO.
- ✓ Sensitive to the P-wave NN force

Covariance may indeed play a crucial role in the description of polarization A_y and related observables of few-body system.

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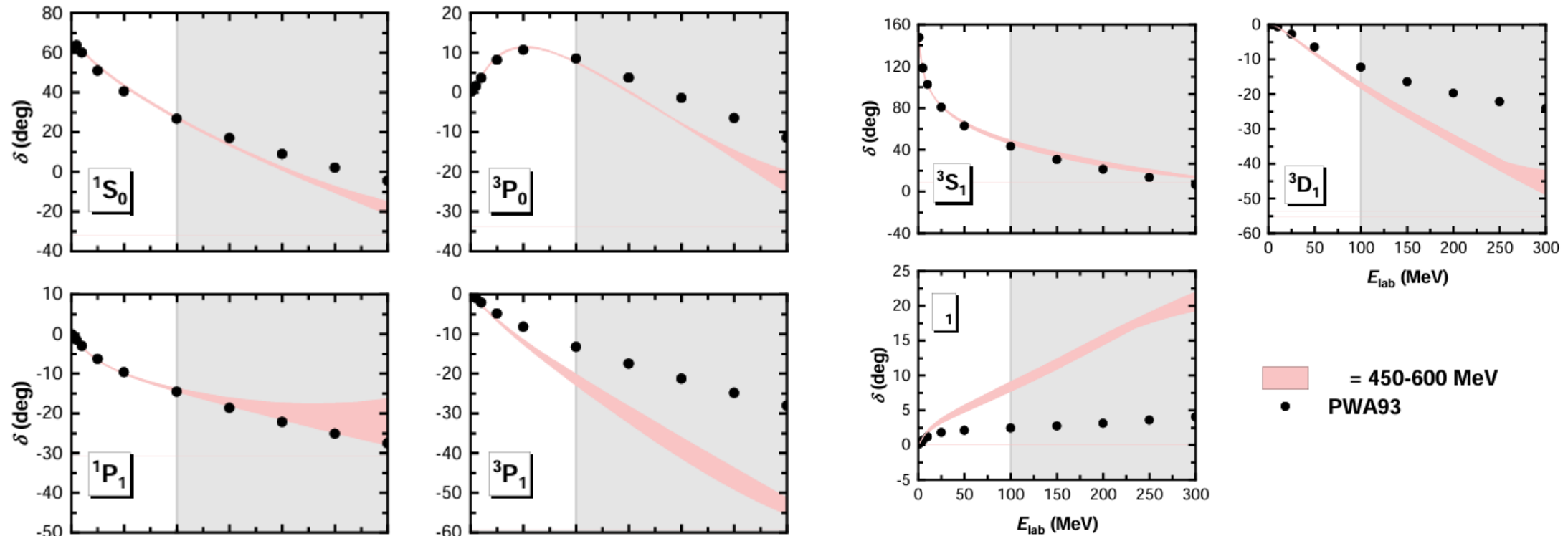
- ☞ Why relativistic chiral nuclear force
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- ☞ **Why relativistic chiral nuclear force**
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- ☞ **Summary and outlook**

SNM in the DBHF theory (vs. BHF)

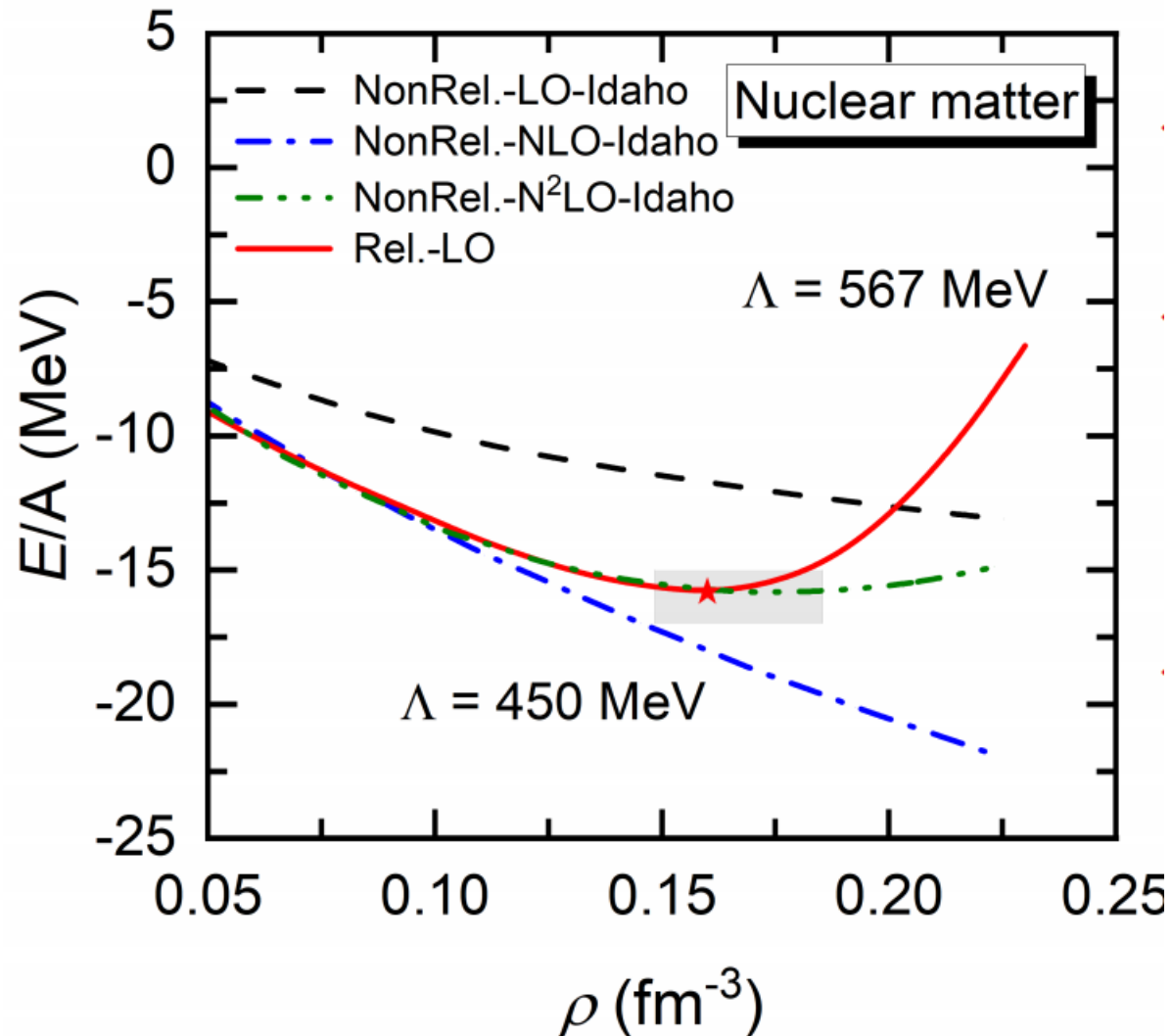
Wei-Jiang Zou, Jun-Xu Lu, Peng-Wei Zhao, LSG*, Jie Meng,
Phys.Lett.B 854 (2024)138732



Only four LECs

SNM in the DBHF theory (vs. BHF)

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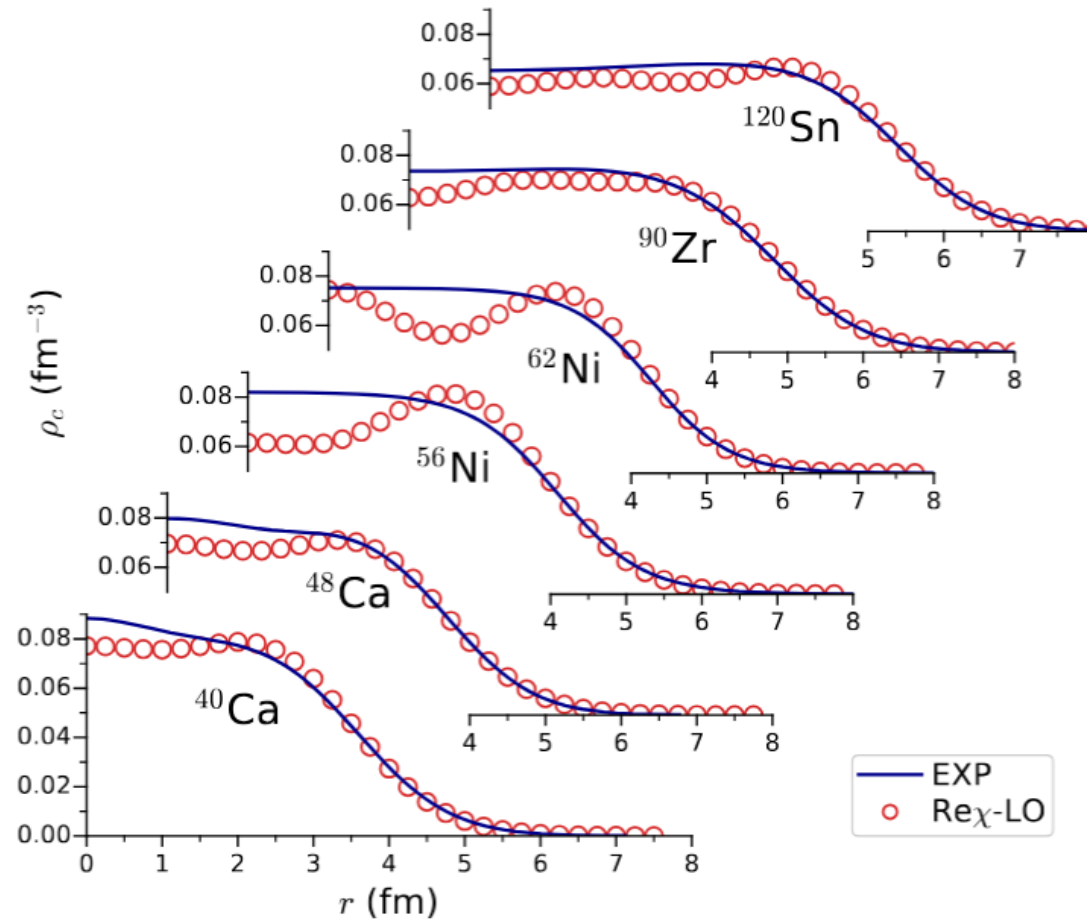
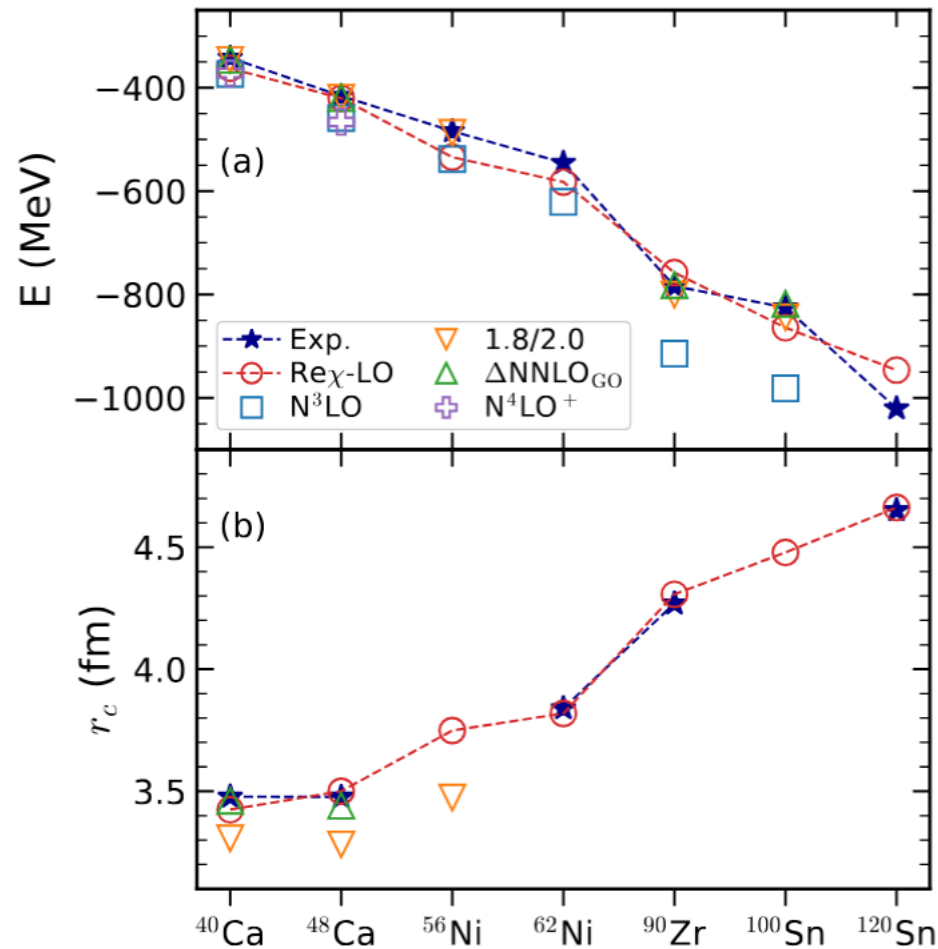
- No saturation in the BHF up to NLO.
- Only at N²LO and with 3NF considered, saturation can be achieved in the BHF.
- Below the saturation point, results are similar

F. Sammarruca et al, PRC104, 064312

Only four LECs

Finite nuclei in the DBHF theory

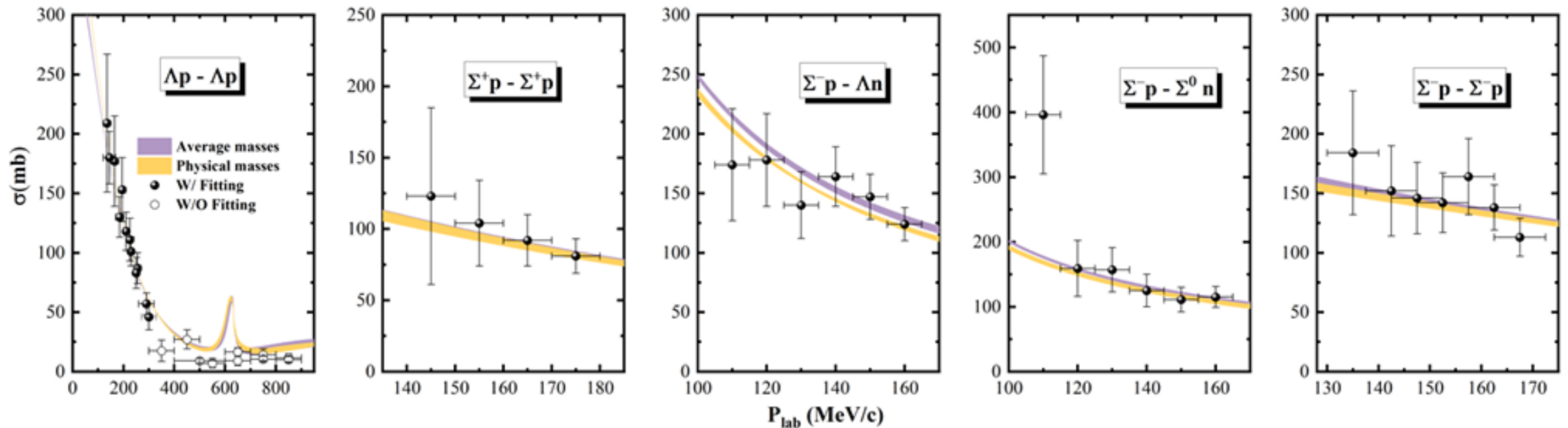
Shi-Hang Shen, Jun-Xu Lu, LSG*, Jie Meng, Wei-Jiang Zou, 2507.01257, to appear in PRC



Only four LECs

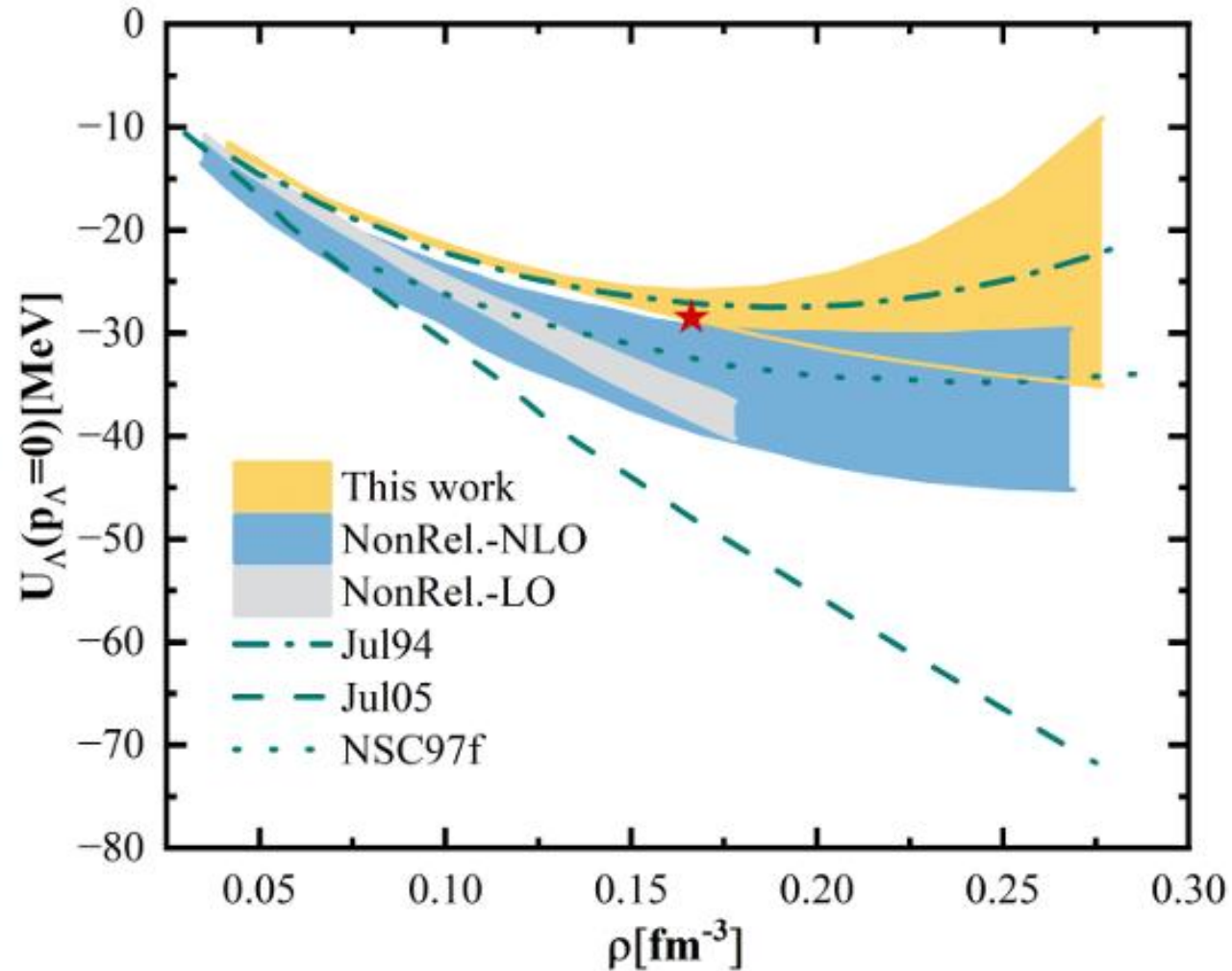
Relativistic chiral hyperon-nuclear force

Ru-You Zheng, Zhi-Wei Liu, LSG*, Jin-Niu Hu, Sibao Wang,
Phys.Lett.B 864 (2025) 139416



The Λ single-particle potential

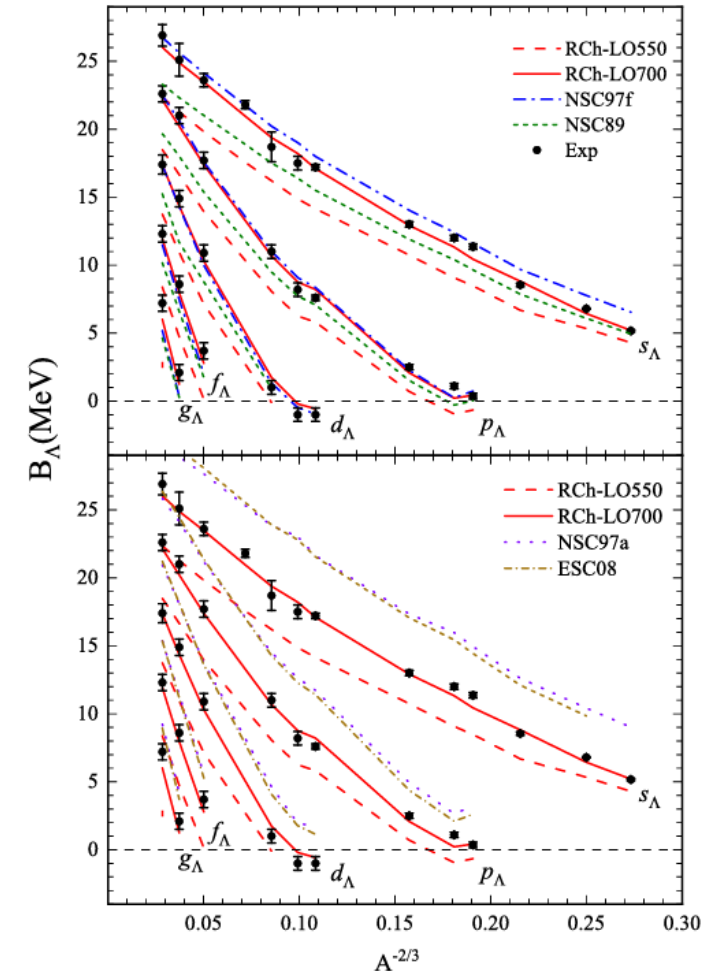
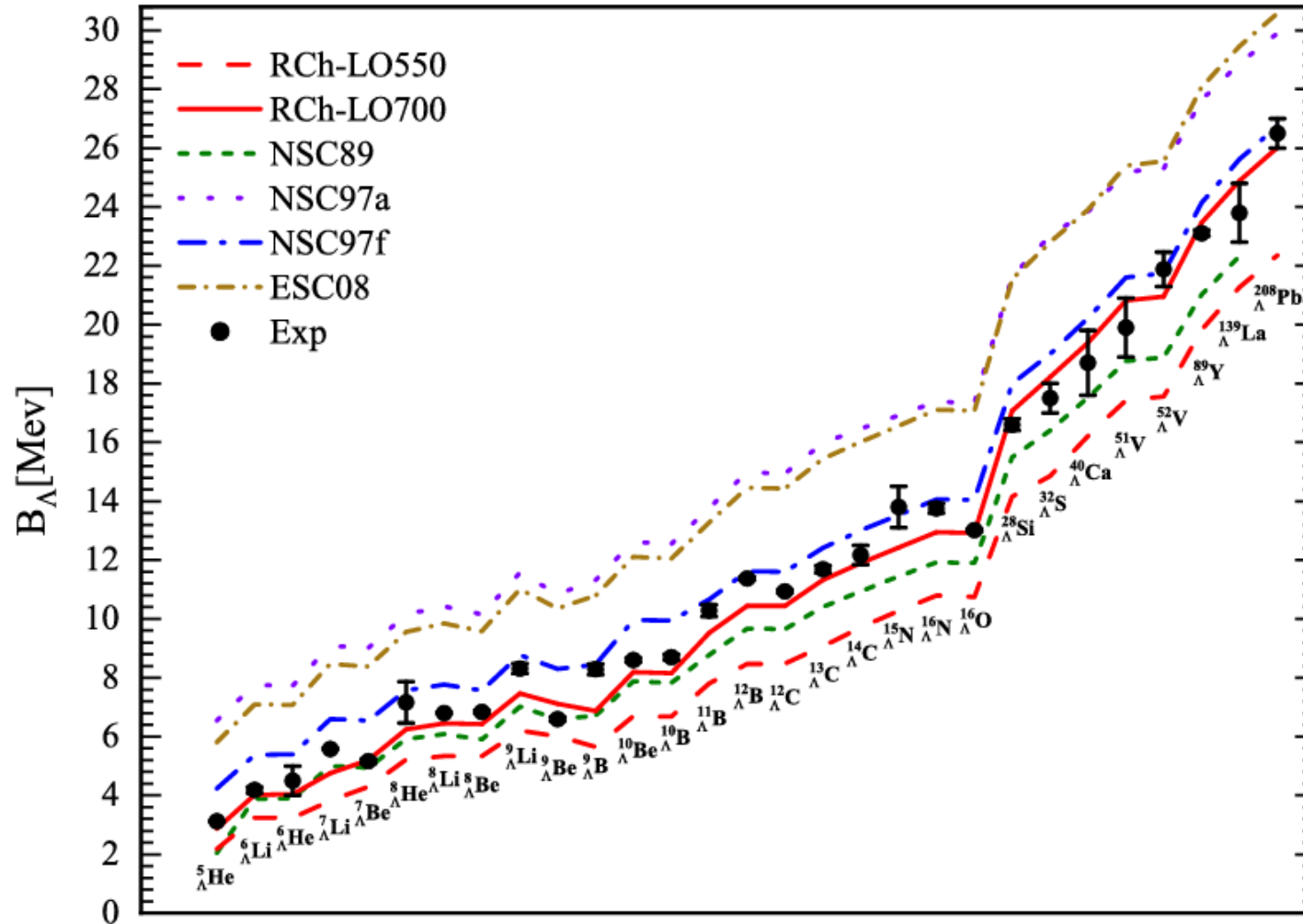
Ru-You Zheng, Zhi-Wei Liu, LSG*, Jin-Niu Hu, Sibo Wang,
Phys.Lett.B 864 (2025) 139416



Hypernuclear structure

Xuan Zeng, Ru-You Zheng, Zhi-Wei Liu, LSG*, Xian-Rong Zhou,
PRC 113 (2026) 014310

Λ hyperon binding energies



Energy levels of the Λ single-particle major shells

Summary and Outlook

- We have developed a high-precision covariant chiral nuclear force up to the NNLO.
- We are now working on N³LO covariant NN forces and covariant three-body forces, and have worked out all the technical complexities
- We have applied the leading order NN and YN forces to study symmetric **nuclear matter, finite nuclei**, as well as **hypernuclear systems**. The results are very encouraging.
- More to come. Remain tuned.

Collaborators

- **Beihang University:** Jun-Xu Lu, Yang-Xiao, Zhi-Wei Liu, Qing-Yu Zhai, Ye-bin Wen, Ru-You Zheng, Shi-Hang Shen
- **Peking University:** Jie Meng, Peng-Wei Zhao, Wei-Jiang Zou
- **Institute of Modern Physics:** Ju-jun Xie
- **Chinese Academy for Atomic Energy:** Rui-Rui Xu
- **University of Tokyo:** Hao-Zhao Liang
- **East China Normal University:** Xian-Rong Zhou

Funding agencies

