



- What is the absolute scale of neutrino mass?
- Are neutrinos their own antiparticles?
- •Are there *more* than 3 mass eigenstates?
  - •Are there non-weakly-interacting "sterile" neutrinos?
    - •What are the neutrino magnetic and electric dipole moments?

•How close to maximal (45°) is  $\theta_{23}$ ?

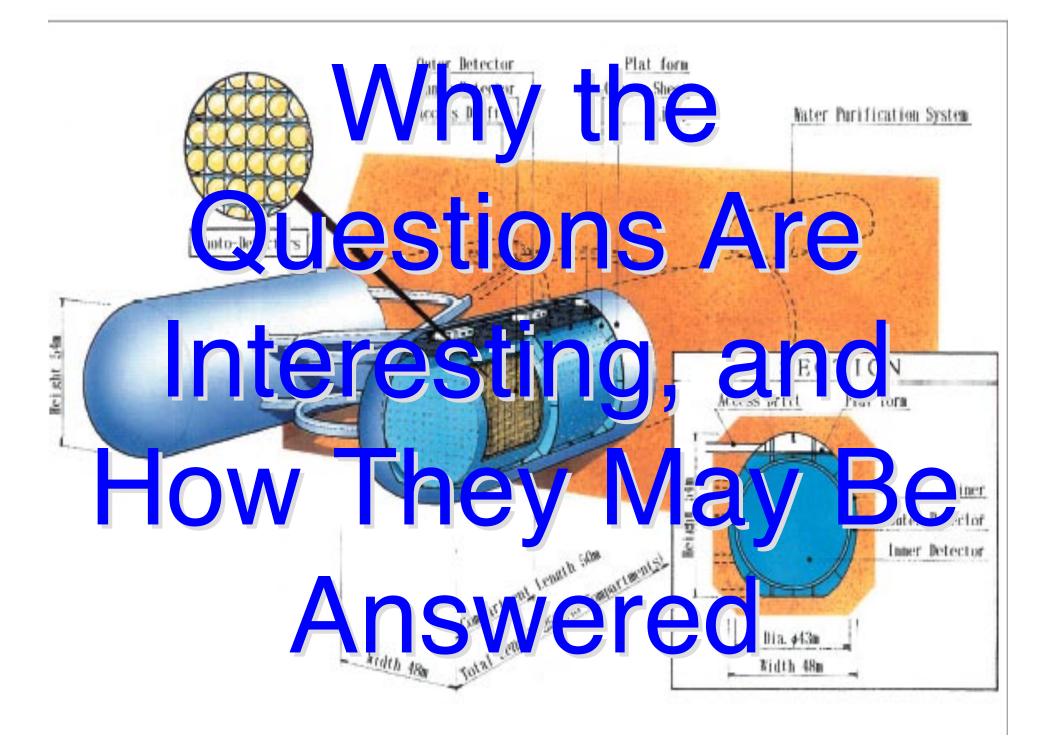
•Is the spectrum like  $\equiv$  or  $\equiv$ ?

•Do neutrino interactions violate CP? Is  $P(\bar{\nu}_{\alpha} \to \bar{\nu}_{\beta}) \neq P(\nu_{\alpha} \to \nu_{\beta})$ ? • What can neutrinos and the universe tell us about one another?

• Is CP violation involving neutrinos the key to understanding the matter — antimatter asymmetry of the universe?

•Where does neutrino mass come from?

•What surprises are in store?



### Does $\overline{\mathbf{v}} = \mathbf{v}$ ?

#### What Is the Question?

For each mass eigenstate  $\nu_i$  , and given helicty h, does —

• 
$$\overline{v_i}(h) = v_i(h)$$
 (Majorana neutrinos)

or

• 
$$\overline{v_i}(h) \neq v_i(h)$$
 (Dirac neutrinos)?

Equivalently, do neutrinos have *Majorana masses*? If they do, then the mass eigenstates are *Majorana neutrinos*.

#### **Dirac Masses**

Dirac neutrino masses are the neutrino analogues of the SM quark and charged lepton masses.

To build a Dirac mass for the neutrino v, we require not only the left-handed field  $v_L$  in the Standard Model, but also a right-handed neutrino field  $v_R$ .

The Dirac neutrino mass term is —

$$m_{D}\overline{\mathbf{v}}_{L}\,\mathbf{v}_{R}$$
 $\stackrel{(\overline{\mathbf{v}}_{R})}{\longrightarrow} \underbrace{\mathbf{v}}_{L}$ 
 $m_{D}$ 

Dirac neutrino masses do not mix neutrinos and antineutrinos.

#### Majorana Masses

Out of, say, a left-handed neutrino field,  $v_L$ , and its charge-conjugate,  $v_L^c$ , we can build a Left-Handed Majorana mass term —

$$m_L \overline{\nu}_L \nu_L^c$$
 $(\overline{\nu})_R \qquad \nu_L \qquad m_1$ 

Majorana masses do mix  $\nu$  and  $\bar{\nu}$ , so they do not conserve the Lepton Number L defined by —

$$L(\mathbf{v}) = L(\ell^{-}) = -L(\overline{\mathbf{v}}) = -L(\ell^{+}) = 1.$$

A Majorana mass for any fermion f causes  $f \leftrightarrow \overline{f}$ .

Quark and charged-lepton Majorana masses are forbidden by electric charge conservation.

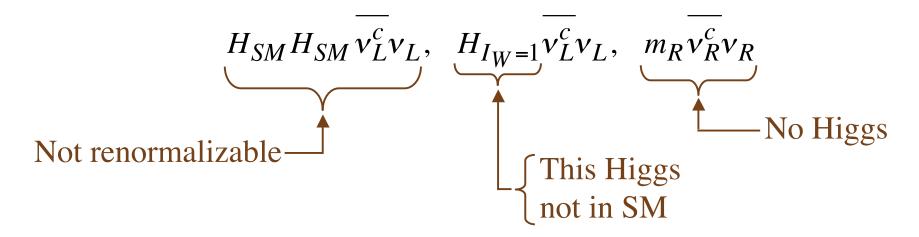
Neutrino Majorana masses would make the neutrinos very distinctive.

Majorana v masses cannot arise via the Higgs mechanism:

$$\mathcal{L}_{SM} = fH_{SM}\overline{v}_Lv_R \Rightarrow f\left\langle H_{SM}\right\rangle_0 \overline{v}_Lv_R \equiv m_v\overline{v}_Lv_R$$
SM Higgs field—Vacuum expectation value

This, the  $\nu$  analogue of the mechanism that produces the q and  $\ell$  masses, leads only to a **Dirac**  $\nu$  mass term.

Possible (Weak-Isospin-Conserving) couplings that can lead to Majorana mass terms:



Majorana neutrino masses must have a different origin than the masses of quarks and charged leptons.

#### Why Majorana Masses — Majorana Neutrinos

The objects  $\mathbf{v}_L$  and  $\mathbf{v}_L^c$  in  $\mathbf{m}_L \overline{\mathbf{v}_L} \mathbf{v}_L^c$  are not the mass eigenstates, but just the neutrinos in terms of which the model is constructed.  $\mathbf{v}_L$  and  $\mathbf{v}_L^c$  are distinct.

$$m_L \overline{v_L} v_L^c$$
 induces  $v_L \leftrightarrow v_L^c$  mixing.

As a result of  $K^0 \longleftrightarrow \overline{K^0}$  mixing, the neutral K mass eigenstates are —

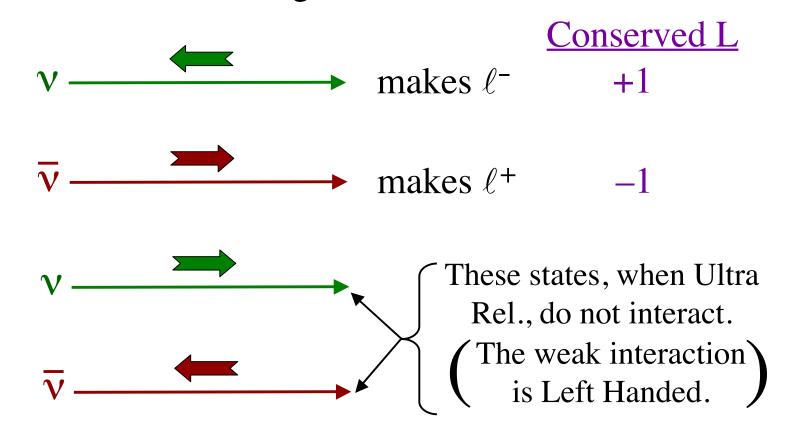
$$K_{S,L} \cong (K^0 \pm \overline{K^0})/\sqrt{2}$$
.  $\overline{K_{S,L}} = K_{S,L}$ .

As a result of  $v_L \leftrightarrow v_L^c$  mixing, the neutrino mass eigenstate is —

$$v_i = v_L + v_L^c = "v + \overline{v}". \overline{v_i} = v_i.$$

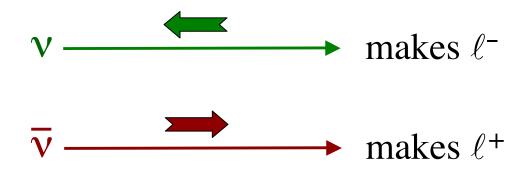
#### SM Interactions Of A Dirac Neutrino

We have 4 mass-degenerate states:



#### SM Interactions Of A Majorana Neutrino

We have only 2 mass-degenerate states:



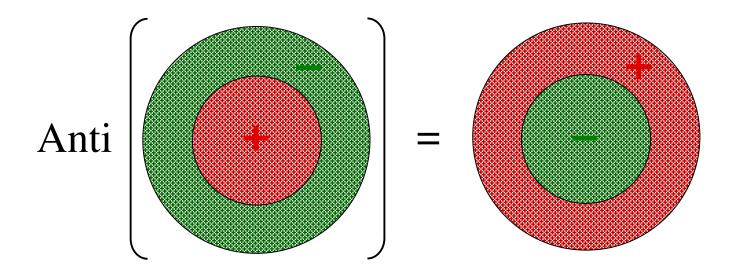
The weak interactions violate *parity*. (They can tell *Left* from *Right*.)

An incoming left-handed neutral lepton makes  $\ell^-$ .

An incoming right-handed neutral lepton makes  $\ell^+$ .

## Can a Majorana Neutrino Have an Electric Charge *Distribution*?

No!

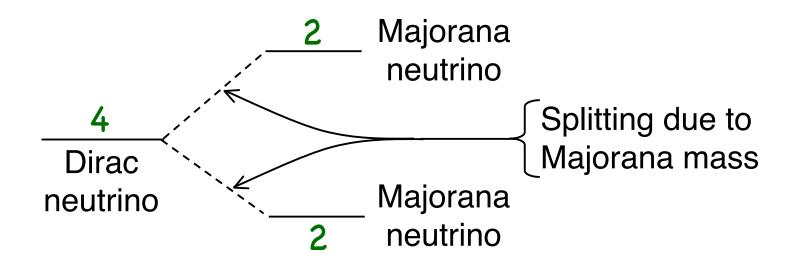


But for a Majorana neutrino —

Anti 
$$(v) = v$$

### Majorana Masses Split Dirac Neutrinos

A Majorana mass term splits a Dirac neutrino into two Majorana neutrinos.



### Why Most Theorists Expect Majorana Masses

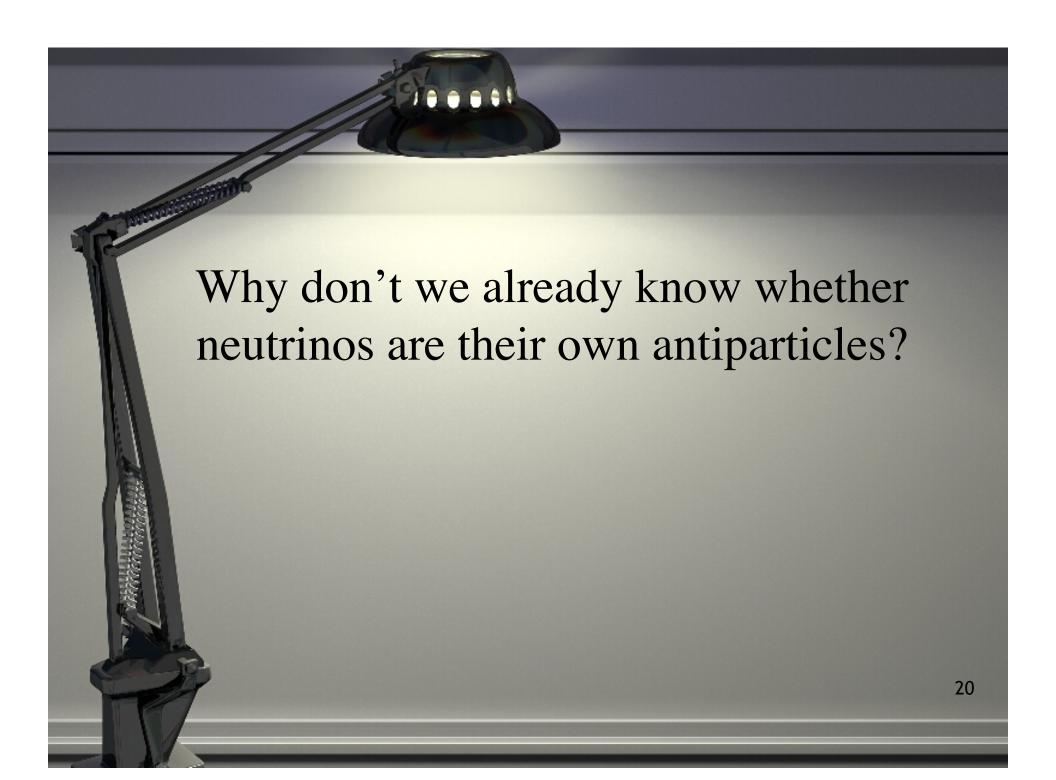
The Standard Model (SM) is defined by the fields it contains, its symmetries (notably weak isospin invariance), and its renormalizability.

Leaving neutrino masses aside, anything allowed by the SM symmetries occurs in nature.

Right-Handed Majorana mass terms are allowed by the SM symmetries.

Then quite likely Majorana masses occur in nature too.

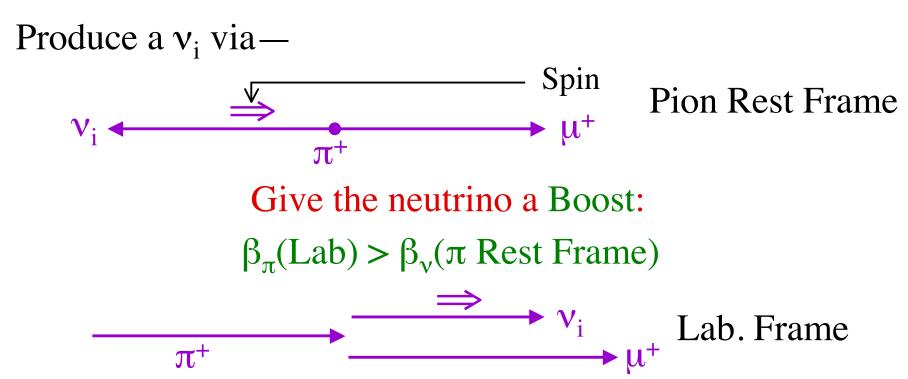
### To Determine Whether Majorana Masses Occur in Nature, So That $\overline{\mathbf{v}} = \mathbf{v}$



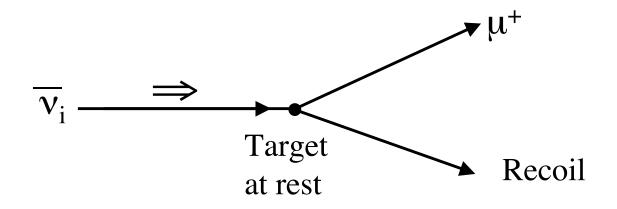
We assume neutrino interactions are correctly described by the SM. Then the interactions conserve L ( $\mathbf{v} \to \ell^-; \bar{\mathbf{v}} \to \ell^+$ ).

It is the Majorana masses that do not conserve L.

### An Idea that Does Not Work [and illustrates why most ideas do not work]



#### The SM weak interaction causes—



$$v_i = \overline{v_i}$$
 means that  $v_i(h) = \overline{v_i}(h)$ .

helicity

If 
$$v_i \xrightarrow{\Longrightarrow} = \overline{v_i} \xrightarrow{\Longrightarrow}$$
,

our 
$$v_i \longrightarrow will$$
 make  $\mu^+$  too.

#### Minor Technical Difficulties

$$\beta_{\pi}(\text{Lab}) > \beta_{\nu}(\pi \text{ Rest Frame})$$

$$\Rightarrow \frac{E_{\pi}(\text{Lab})}{m_{\pi}} > \frac{E_{\nu}(\pi \text{ Rest Frame})}{m_{\nu_{i}}}$$

$$\Rightarrow E_{\pi}(\text{Lab}) \gtrsim \frac{10^{5} \text{ TeV if } m_{\nu_{i}} \sim 0.05 \text{ eV}}{m_{\nu_{i}}}$$

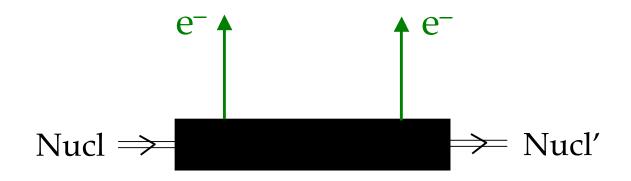
Fraction of all  $\pi$  – decay  $\nu_i$  that get helicity flipped

$$\approx \left(\frac{m_{v_i}}{E_v(\pi \text{ Rest Frame})}\right)^2 \sim 10^{-18} \text{ if } m_{v_i} \sim 0.05 \text{ eV}$$

Since L-violation comes only from Majorana neutrino *masses*, any attempt to observe it will be at the mercy of the neutrino masses.

(B.K. & Stodolsky)

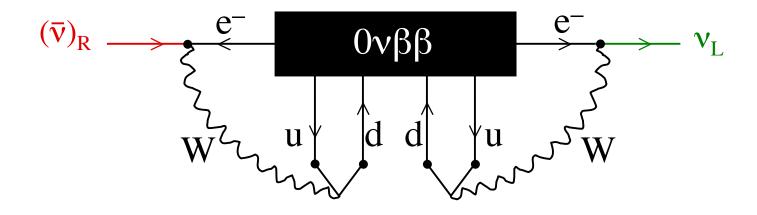
## The Promising Approach — Seek Neutrinoless Double Beta Decay [0νββ]



We are looking for a *small* Majorana neutrino mass. Thus, we will need *a lot* of parent nuclei (say, one ton of them).

Whatever diagrams cause  $0\nu\beta\beta$ , its observation would imply the existence of a Majorana mass term:

(Schechter and Valle)



$$(\bar{\mathbf{v}})_{R} \rightarrow \mathbf{v}_{L} : A \text{ (tiny) Majorana mass term}$$
  

$$\therefore 0 \mathbf{v} \beta \beta \longrightarrow \overline{\mathbf{v}}_{i} = \mathbf{v}_{i}$$

# What Is the Mass Ordering?

#### Is The Mass Spectrum $\equiv$ or $\equiv$ ?

Generically, grand unified models (GUTS) favor —

GUTS relate the Leptons to the Quarks.

However, *Majorana masses*, with no quark analogues, could turn \_\_\_ into \_\_\_ .

If the spectrum looks like  $\equiv$ , and neutrinos are Majorana particles, there is a lower bound on the rate for  $0\nu\beta\beta$ .

## How To Determine If The Spectrum Is Normal Or Inverted

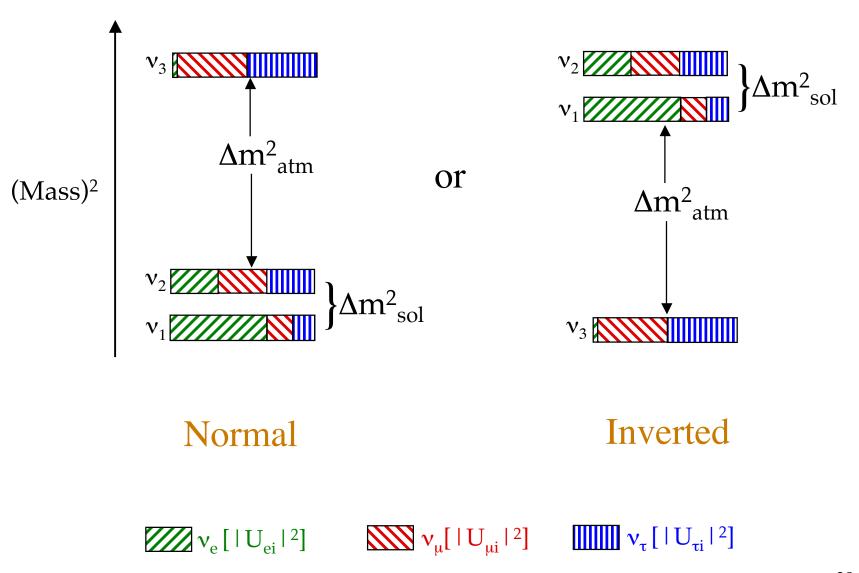
One way: exploit the *matter effect* on accelerator neutrinos.

Recall that the matter effect *raises* the effective mass of  $v_e$ , but *lowers* that of  $\overline{v}_e$ . Thus, it affects v and  $\overline{v}$  oscillation *differently*, leading to:

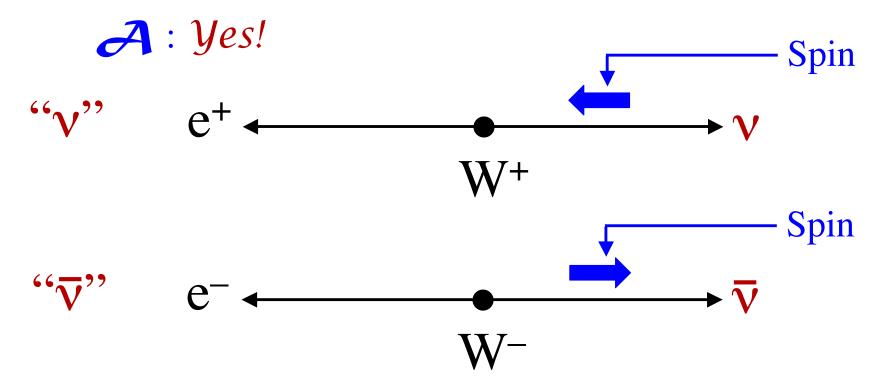
$$\frac{P(\mathbf{v}_{\mu} \rightarrow \mathbf{v}_{e})}{P(\overline{\mathbf{v}_{\mu}} \rightarrow \overline{\mathbf{v}_{e}})} \begin{cases} > 1 ; \\ < 1 ; \end{cases} \qquad \text{Note fake CP}$$

Note dependence on the mass ordering

### The matter effect depends on whether the spectrum is Normal or Inverted.



②: Does matter still affect v and  $\overline{v}$  differently when  $\overline{v} = v$ ?



The weak interactions violate *parity*. Neutrino – matter interactions depend on the neutrino *polarization*.

## PViolation In General, In Neutrino Oscillation, and In Leptogenesis

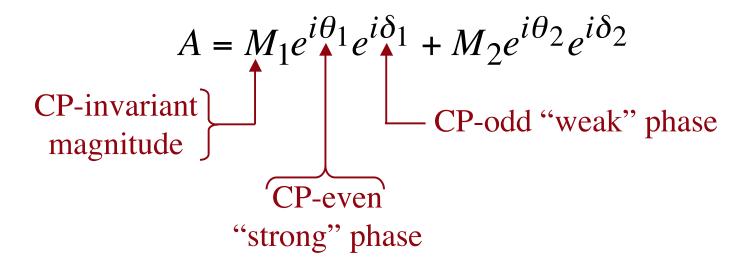
# How CP Violation **Comes About** 32

A always comes from phases.

Therefore, P always requires an *interference* between (at least) two amplitudes.

For example, an interference between two Feynman diagrams.

Let us consider how a CP-violating rate difference between two CP-mirror-image processes, such as  $B^+ \to D^0 K^+$  and  $B^- \to \overline{D}^0 K^-$ , arises. Suppose some process P has the amplitude —



Then the CP-mirror-image process  $\overline{P}$  has the amplitude —

$$\overline{A} = M_1 e^{i\theta_1} e^{-i\delta_1} + M_2 e^{i\theta_2} e^{-i\delta_2}$$

Then the rates for  $\overline{P}$  and P differ by —

$$\overline{\Gamma} - \Gamma = |\overline{A}|^2 - |A|^2 = 4M_1M_2\sin(\theta_1 - \theta_2)\sin(\delta_1 - \delta_2)$$

Assumes equal amounts of the initial states of 
$$\overline{P}$$
 and  $P$ .
$$\overline{\Gamma} - \Gamma = |\overline{A}|^2 - |A|^2 = 4M_1M_2\sin(\theta_1 - \theta_2)\sin(\delta_1 - \delta_2)$$

A CP-violating rate difference requires 3 ingredients:

- Two interfering amplitudes
- •These two amplitudes must have different CP-even phases
- •These two amplitudes must have different CP-odd phases

# Mixing Can Lead to Unequal Amounts of the CP-Mirror-Image Initial States

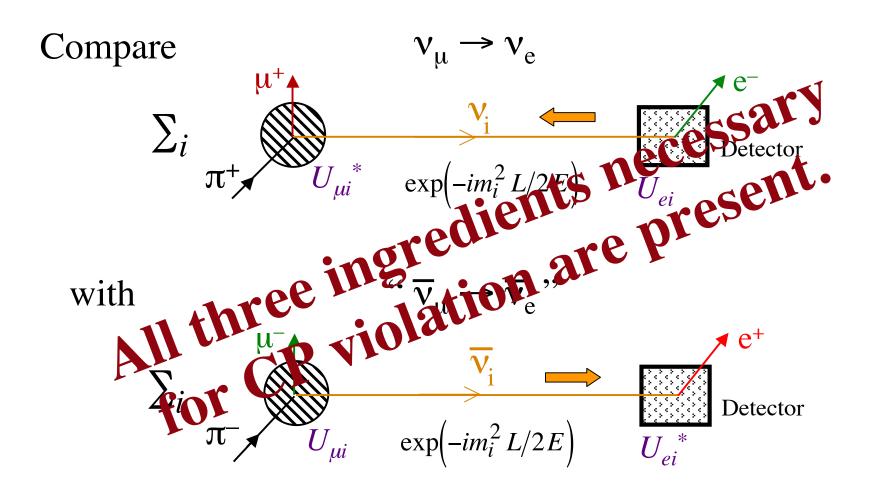
From 
$$K^0$$
 only 
$$\Gamma(K_L \to \pi^- \ell^+ \nu) - \Gamma(K_L \to \pi^+ \ell^- \overline{\nu}) \neq 0 \text{ violates CP.}$$
But  $K_L \propto (1+\varepsilon)K^0 - (1-\varepsilon)\overline{K}^0$ ,

where  $\varepsilon$  is the usual CP-violating parameter arising from  $K^0 - \overline{K}^0$  mixing, and  $|1 + \varepsilon|^2 \neq |1 - \varepsilon|^2$ .

This rate difference comes from  $\varepsilon$ , rather than the 3 ingredients just listed.

# **CP Violation In** Neutrino Oscillation 37

# The ingredients for CP in neutrino oscillation, even if $\overline{v} = v$



## There Is Nothing Special About $\theta_{13}$

All mixing angles must be nonzero for P in oscillation.

For example —
$$P(\bar{v}_{\mu} \rightarrow \bar{v}_{e}) - P(v_{\mu} \rightarrow v_{e}) = 2\cos\theta_{13}\sin2\theta_{13}\sin2\theta_{12}\sin2\theta_{23}\sin\delta$$

$$= 2\cos\theta_{13}\sin(2\theta_{13}\sin2\theta_{12}\sin2\theta_{23}\sin\delta)$$

$$= 2\cos\theta_{13}\sin(2\theta_{13}\sin2\theta_{12}\sin2\theta_{23}\sin\delta)$$

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$$= 2\cos\theta_{13}\sin(2\theta_{13}\sin2\theta_{12}\sin2\theta_{23}\sin\delta)$$

In the factored form of U, one can put  $\delta$  next to  $\theta_{12}$  instead of  $\theta_{13}$ .

$$P(\overline{v}_{\mu} \to \overline{v}_{e}) - P(v_{\mu} \to v_{e}) = 2\cos\theta_{13}\sin2\theta_{13}\sin2\theta_{12}\sin2\theta_{23}\sin\delta$$

$$\times \sin\left(\Delta m^{2}_{31}\frac{L}{4E}\right)\sin\left(\Delta m^{2}_{32}\frac{L}{4E}\right)\sin\left(\Delta m^{2}_{21}\frac{L}{4E}\right)$$

This, and everything else about neutrino oscillation, is independent of whether neutrinos are Dirac or Majorana particles.

# **CP Violation In** Leptogenesis

Leptogenesis explains the matter-antimatter asymmetry of the universe in terms of CP violation in the early-universe decays of super-heavy neutrinos.

(Silvia Pascoli)

Imagine in the early universe 3 heavy neutrinos  $N_i = \overline{N}_i$ , coupled to the 3 light lepton families  $(v_\alpha, \ell_\alpha)$ , and the SM Higgs doublet  $(\overline{H^0}, H^-)$ , by a "Yukawa" coupling:

$$\mathcal{L}_{Y} = \sum_{\substack{\alpha = e, \mu, \tau \\ i = 1, 2, 3}} y_{\alpha i} \left[ \overline{v_{\alpha}} \overline{H^{0}} - \overline{\ell_{\alpha}} H^{-} \right] N_{i} + h.c.$$
Yukawa coupling matrix

Imagine that this coupling is the only way the heavy neutrinos  $N_i$  interact with the rest of the world.

#### The Yukawa interaction —

$$\mathcal{L}_{Y} = \sum_{\substack{\alpha = e, \mu, \tau \\ i = 1, 2, 3}} y_{\alpha i} \left[ \overline{v_{\alpha}} \overline{H^{0}} - \overline{\ell_{\alpha}} H^{-} \right] N_{i} + h.c.$$

causes the decays —

$$N \rightarrow \ell^{\mp} + H^{\pm}$$
 and  $N \rightarrow (\overline{\nu}) + \overline{H^0}$ 

Phases in the matrix y will lead to —

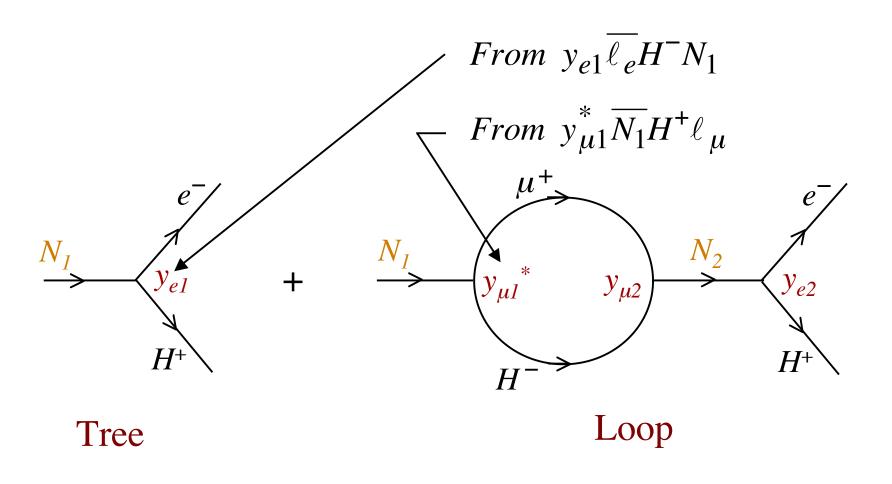
and 
$$\Gamma(N \to \ell^- + H^+) \neq \Gamma(N \to \ell^+ + H^-)$$

$$\Gamma(N \to \nu + H^0) \neq \Gamma(N \to \overline{\nu} + \overline{H^0})$$

### How Do Such of Inequalities Come About?

Let us look at an example.

This example illustrates that **P** in **any decay** always involves amplitudes **beyond** those of lowest order in the Hamiltonian.



$$\Gamma(N_1 \to e^- + H^+) = \left| y_{e1} K_{\text{Tree}} + y_{\mu 1}^* y_{\mu 2} y_{e2} K_{\text{Loop}} \right|^2$$
Kinematical factors

$$\Gamma(N_1 \rightarrow e^- + H^+) = \left| y_{e1} K_{\text{Tree}} + y_{\mu 1}^* y_{\mu 2} y_{e2} K_{\text{Loop}} \right|^2$$

When we go to the CP-mirror-image decay,  $N_1 \rightarrow e^+ + H^-$ , all the coupling constants get complex conjugated, but the kinematical factors do not change.

$$\Gamma(N_1 \to e^+ + H^-) = |y_{e1}^* K_{\text{Tree}} + y_{\mu 1} y_{\mu 2}^* y_{e2}^* K_{\text{Loop}}|^2$$

All three ingredients needed for LF are present.

$$\Gamma(N_1 \to e^- + H^+) - \Gamma(N_1 \to e^+ + H^-)$$
=  $4 \operatorname{Im}(y_{e1}^* y_{\mu 1}^* y_{e2} y_{\mu 2}) \operatorname{Im}(K_{\text{Tree}} K_{\text{Loop}}^*)$ 

The P inequalities —

$$\Gamma\left(N \to \ell^{-} + H^{+}\right) \neq \Gamma\left(N \to \ell^{+} + H^{-}\right)$$

and

$$\Gamma(N \rightarrow v + H^0) \neq \Gamma(N \rightarrow \overline{v} + \overline{H^0})$$

will produce a universe with unequal numbers of leptons ( $\ell^-$  and  $\nu$ ) and antileptons ( $\ell^+$  and  $\bar{\nu}$ ).

In this universe the lepton number L, defined by  $L(\ell^-) = L(v) = -L(\ell^+) = -L(\overline{v}) = 1$ , is not zero.

A further step in leptogenesis then converts part of this nonzero lepton number into a nonzero baryon number.

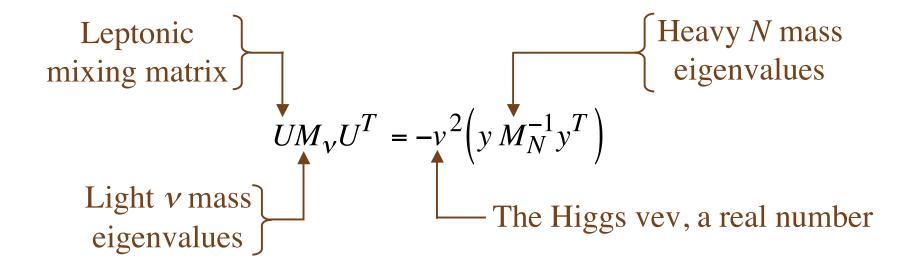
## The Connection Between Leptogenesis and Light-Neutrino &

This is being discussed by Silvia Pascoli.

I would just like to add a comment about the connection being generic.

Leptogenesis is an outgrowth of the *See-Saw picture* of the origin of neutrino masses.

#### The See-Saw Relation



$$\left(\begin{array}{c}
UM_{\nu}U^{T} = -v^{2}\left(yM_{N}^{-1}y^{T}\right) \\
\text{Outputs}
\end{array}\right)$$
Inputs, in  $\mathcal{L}$ 

Through U, the phases in y lead to  $\mathbb{CP}$  in light neutrino oscillation.

$$P(\stackrel{\leftarrow}{\nu_{\alpha}} \rightarrow \stackrel{\leftarrow}{\nu_{\beta}}) = \text{Distance}$$

$$= \delta_{\alpha\beta} - 4 \sum_{i>j} \Re(U_{\alpha i}^* U_{\beta i} U_{\alpha j} U_{\beta j}^*) \sin^2(\Delta m_{ij}^2 \frac{L}{4E})$$

$$+ 2 \sum_{i>j} \Im(U_{\alpha i}^* U_{\beta i} U_{\alpha j} U_{\beta j}^*) \sin(\Delta m_{ij}^2 \frac{L}{2E})$$
Neutrino (Mass)<sup>2</sup> splitting
$$\text{Energy}$$

Generically, leptogenesis and light-neutrino LY imply each other.

Experiments to look for LY
in neutrino oscillation are being
developed in Japan and in the US.

CP is a fundamental symmetry.

Is its nonconservation special to quark mixing?

Or, does it occur in both quark and lepton mixing, as suggested by Grand Unified Theories, which unify the quarks and the leptons?