

# **Hhh Combination update**

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**Xiaohu SUN, IHEP, Beijing, 18-02-2014**

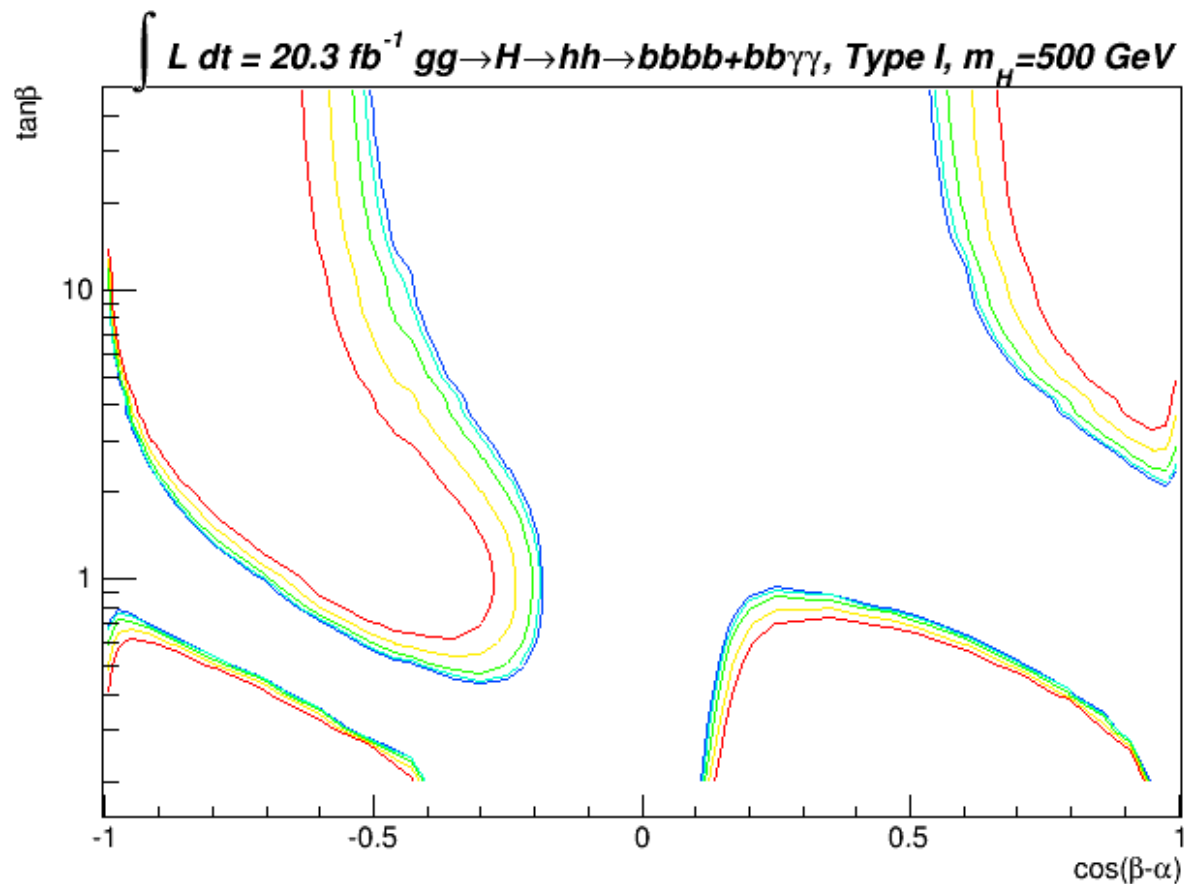
# **Part I**

## **Check width and VBF production**

All made from v160 grid file

# Review on the exclusion

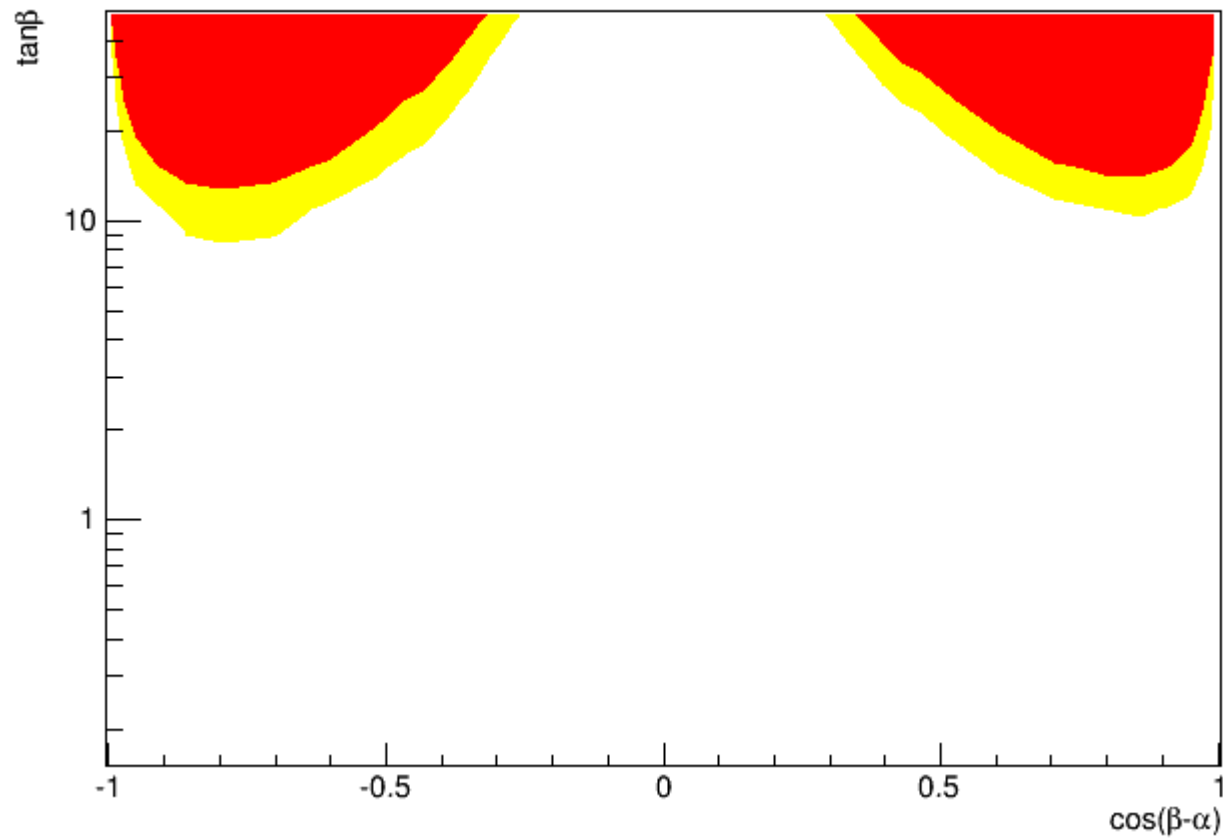
- The exclusion plot made for  $m_H=500\text{GeV}$  ( $ggH>0.37\text{pb}$ )
- Based on a temp upper limit set with graviton signal in bbbb and pythia 2HDM signal in bbyy
- <https://indico.cern.ch/event/288214/contribution/2/material/slides/0.pdf>



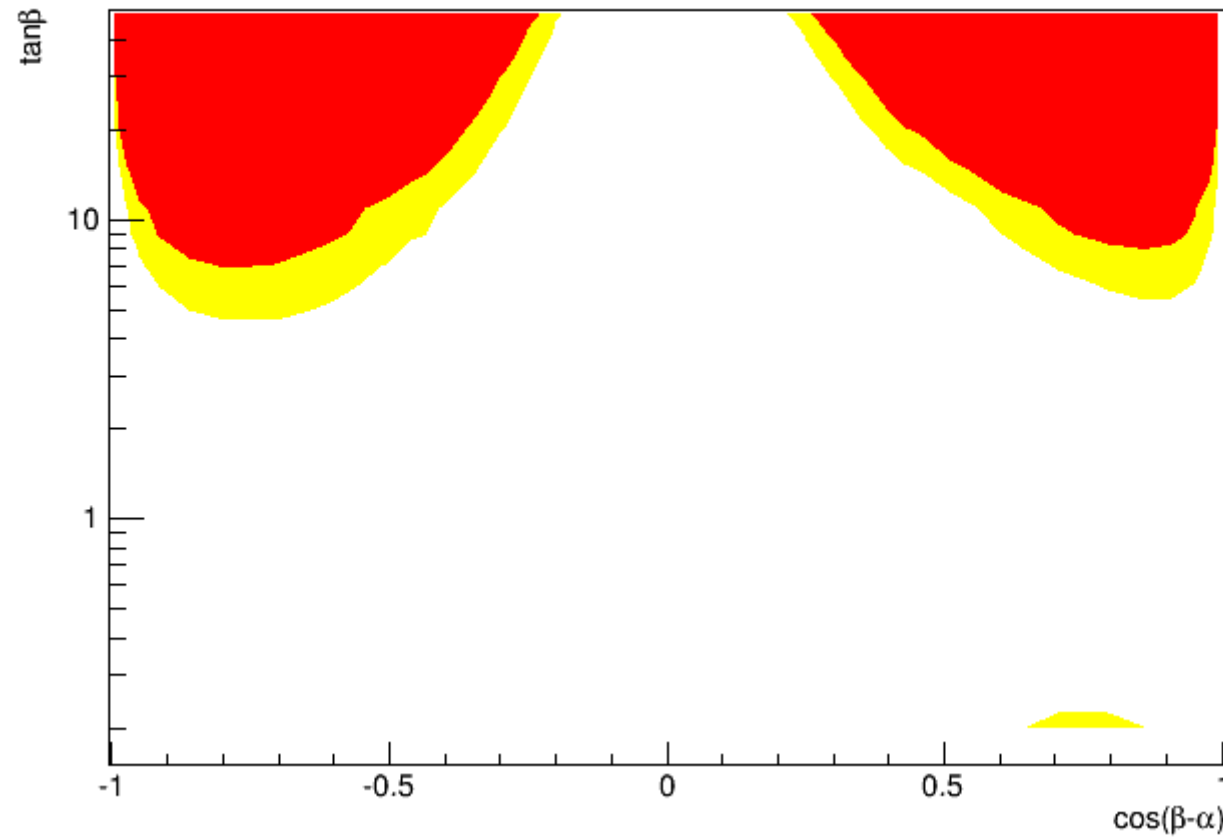
# Check on H width

- Make contours of
  - (  $\text{width\_H} / \text{mH}$  )
- Legends
  - Set contour area **yellow**:  $\text{widthH} > 10\% \text{ mH}$
  - Set contour area **red**:  $\text{widthH} > 20\% \text{ mH}$
- Checked for type I and type II
  - Type I will be shown, since type II is the same

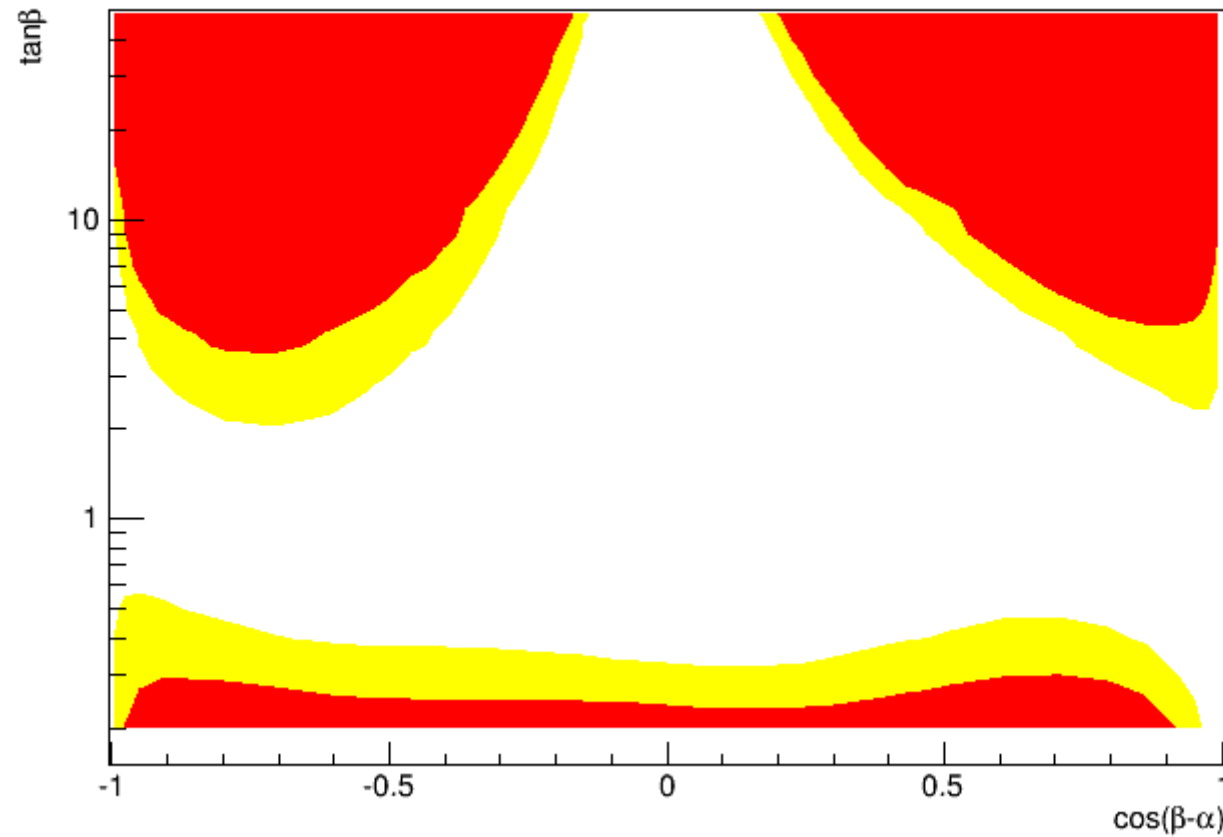
# Check width @ $mH=260$



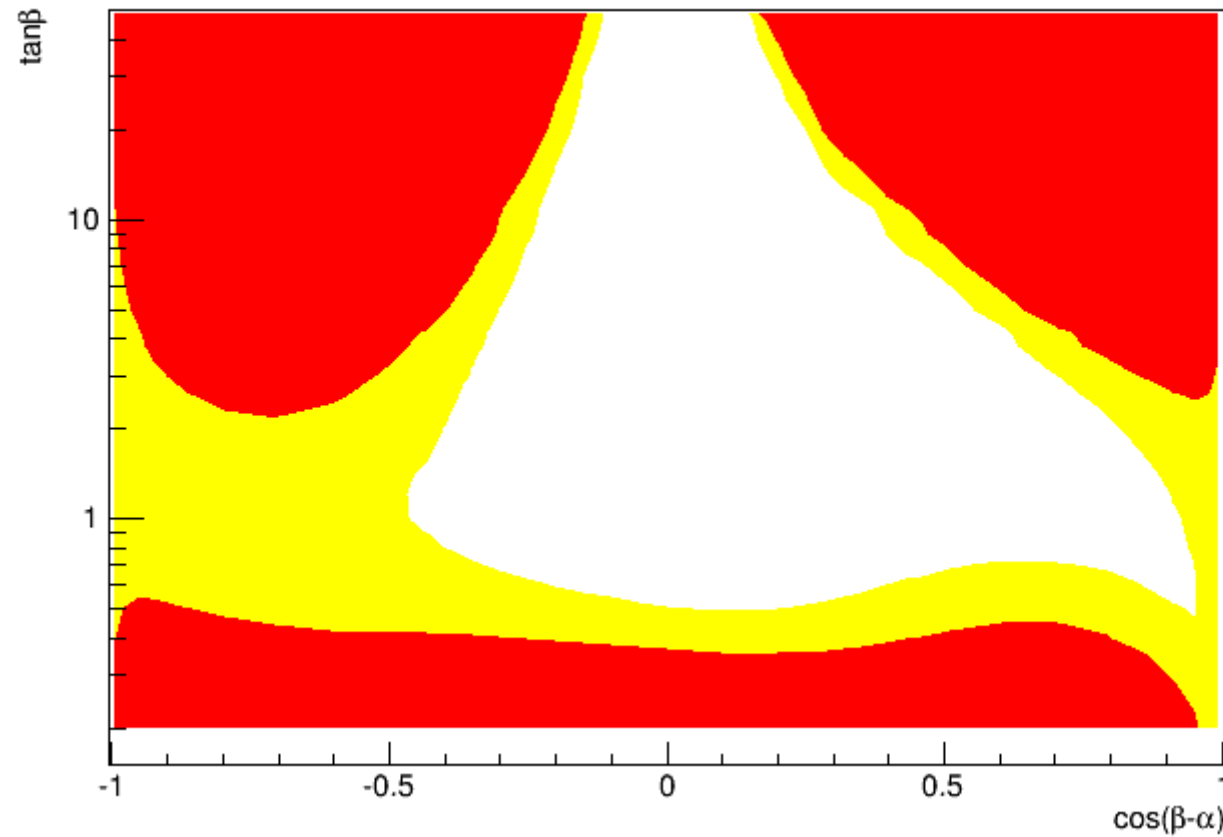
# Check width @ $mH=300$



# Check width @ $mH=400$

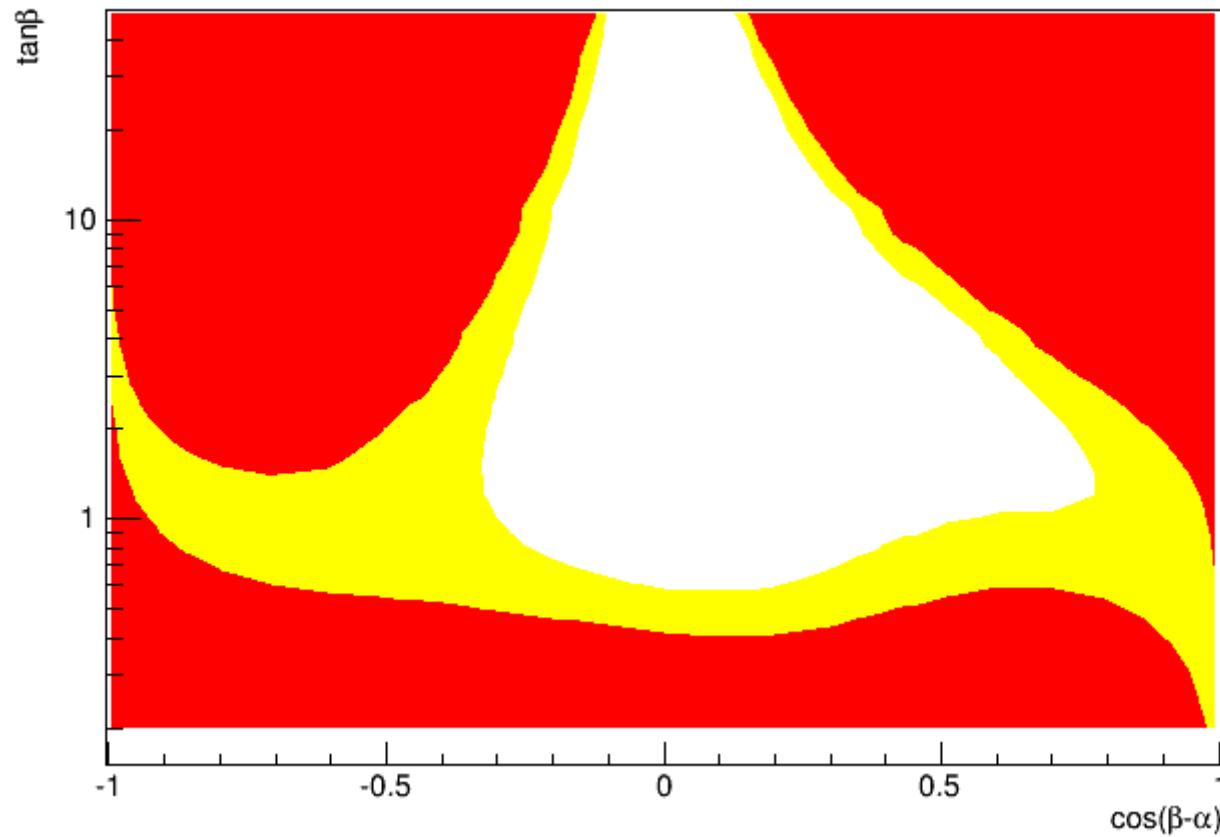


# Check width @ $mH=500$

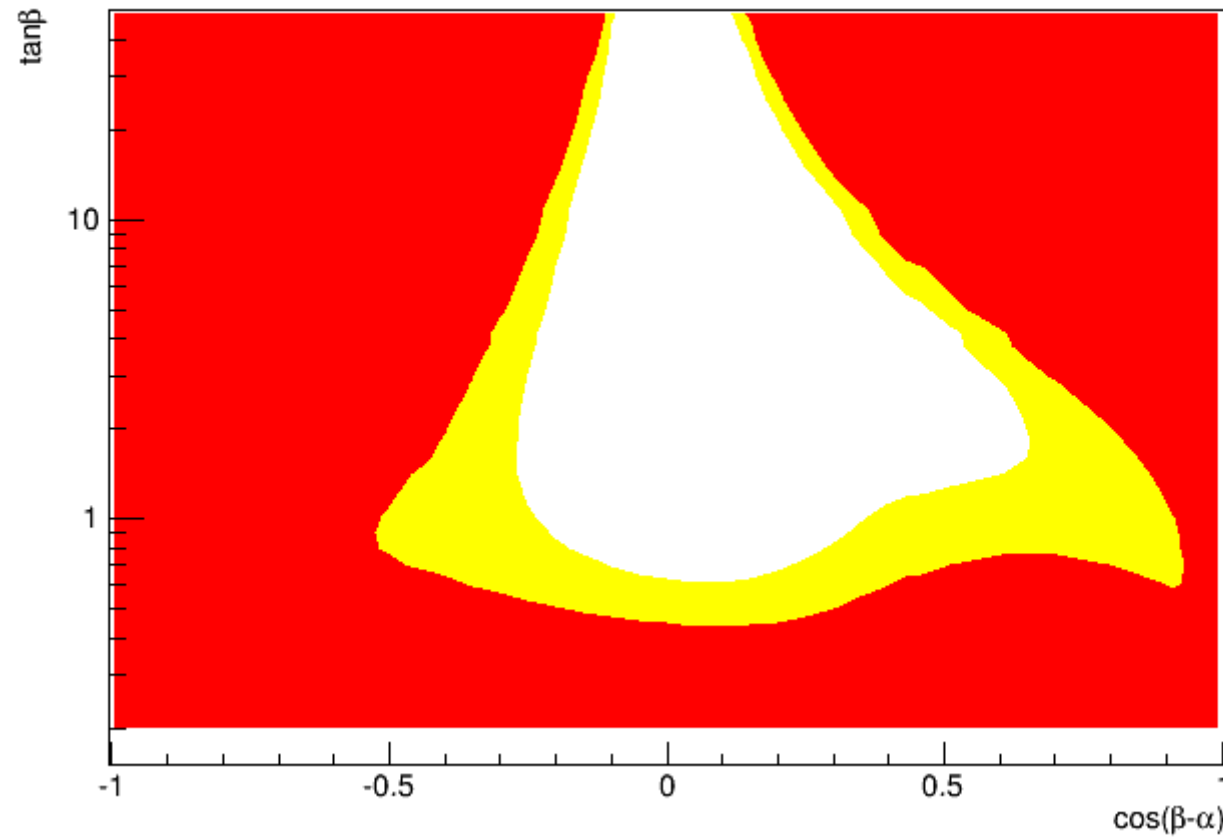




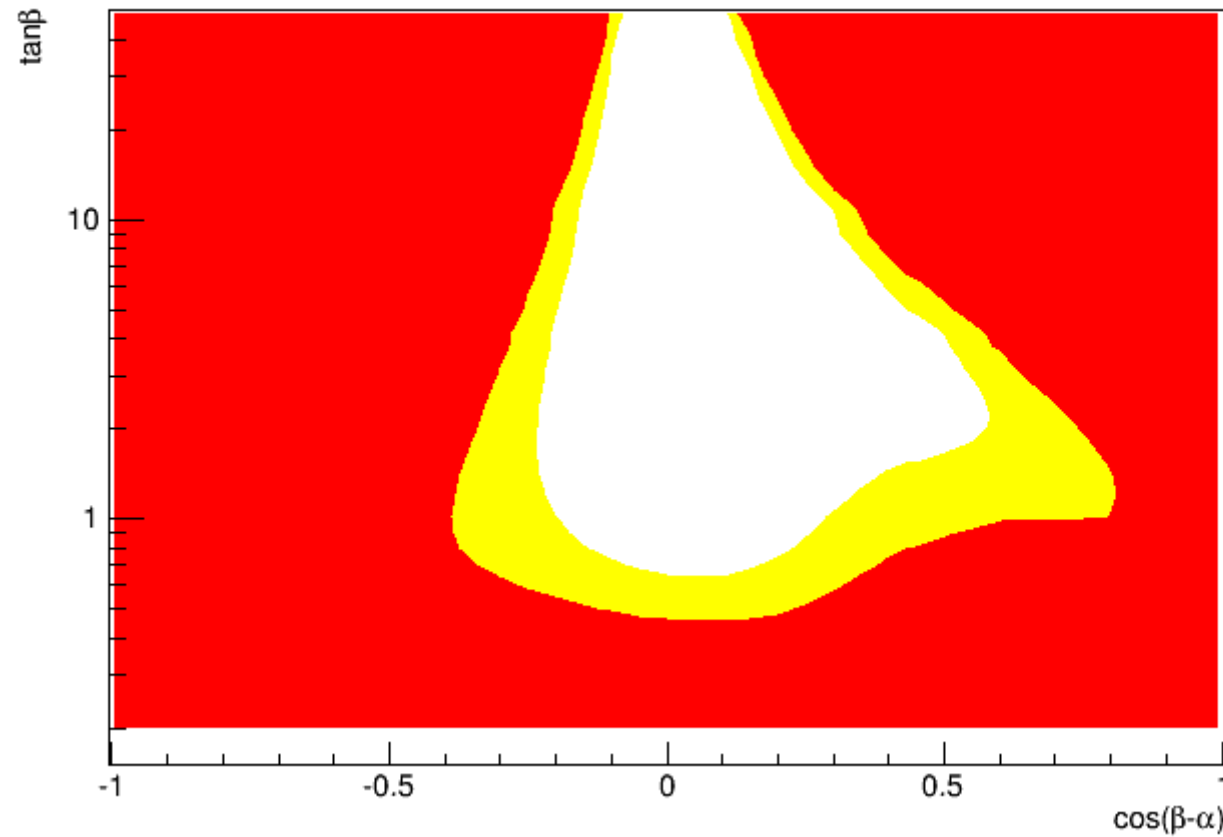
# Check width @ $mH=600$



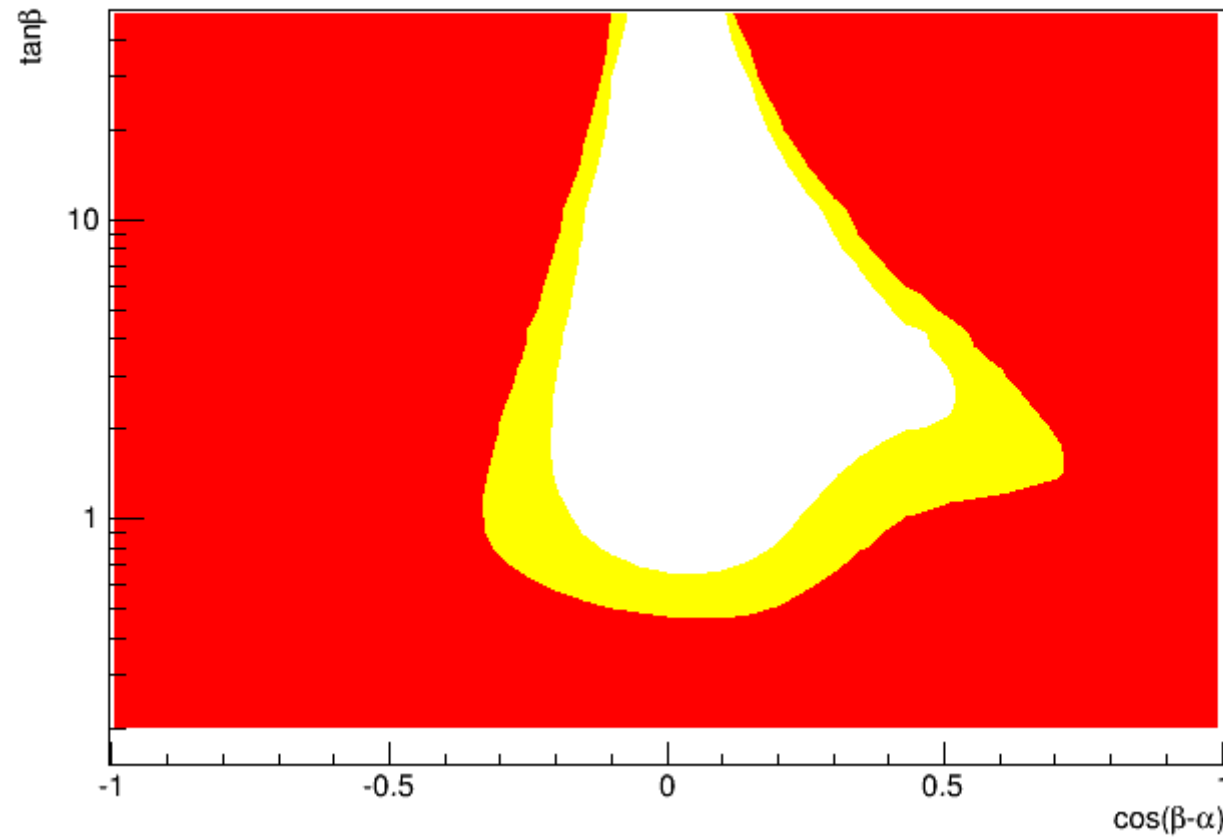
# Check width @ $mH=700$



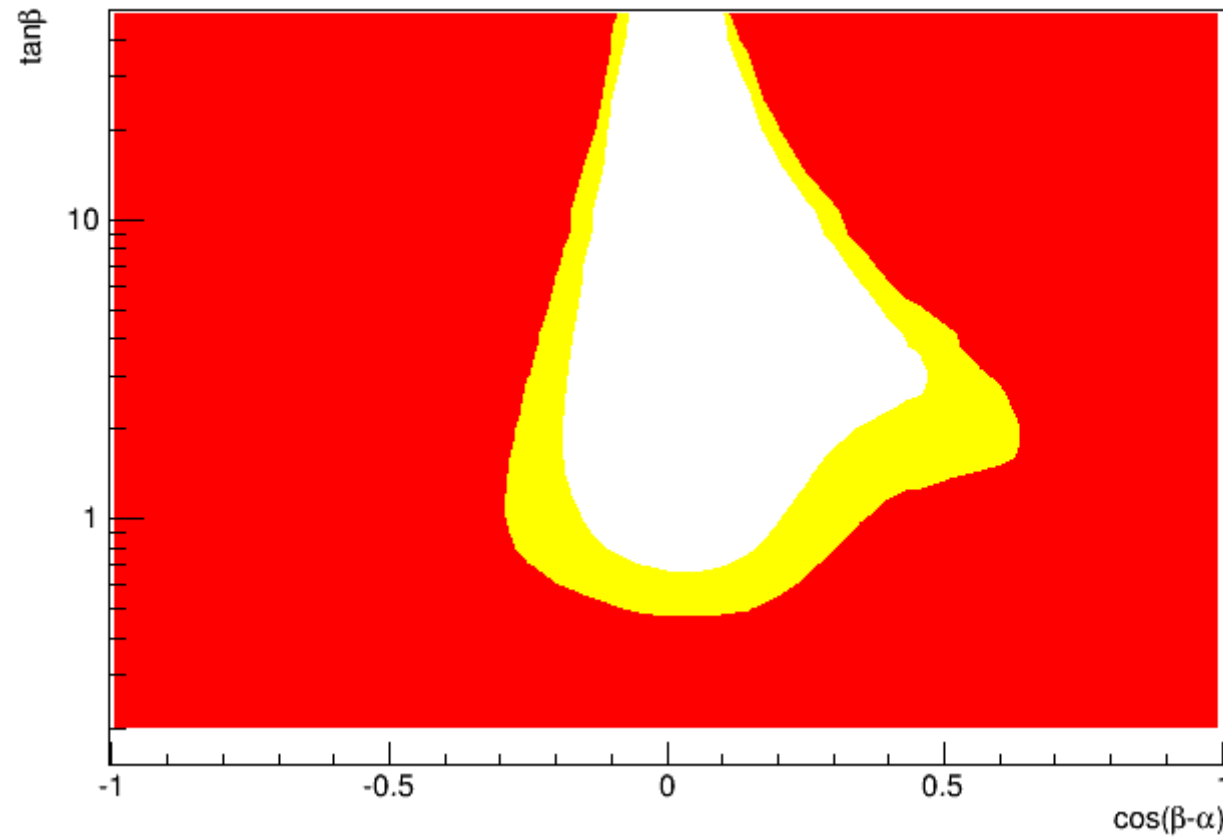
# Check width @ $mH=800$



# Check width @ $mH=900$

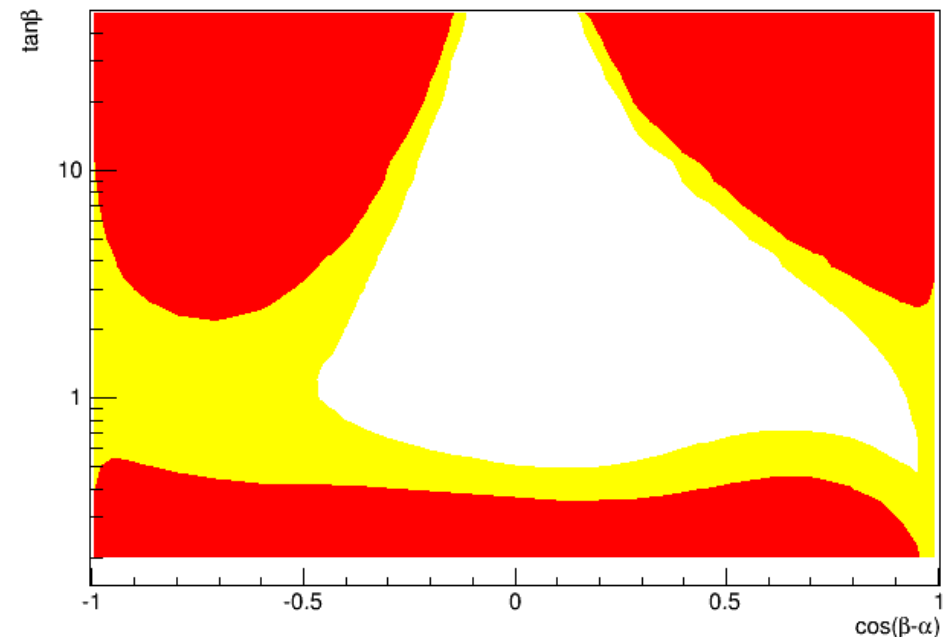
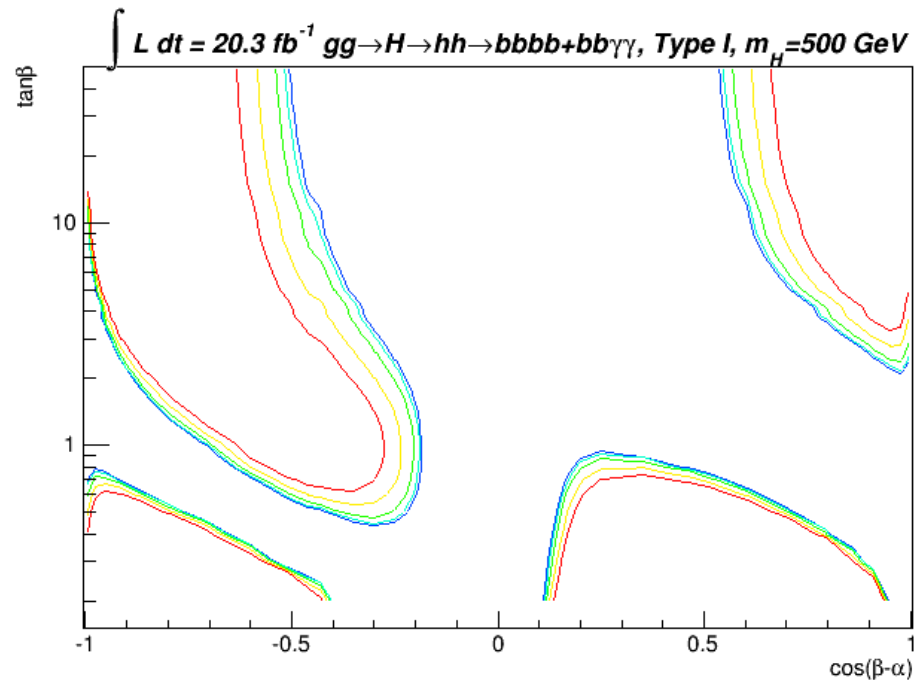


# Check width @ $mH=1000$



# Summary on H width

- Ask both of the analysis groups bbbb and bbyy have to provide studies on the effects from the H width



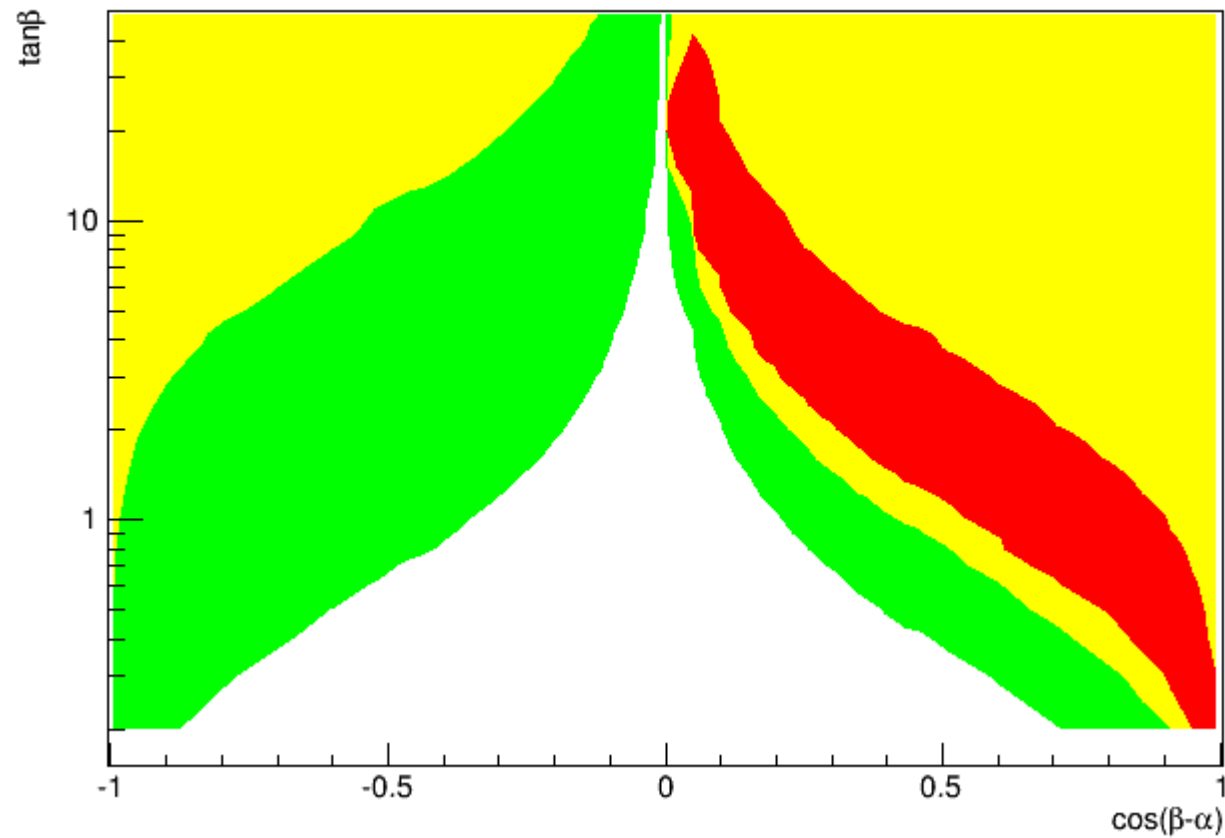
# Check on VBF

- As the second largest production after ggf, one needs to understand how much it will get in our analyses
- Make contours of
  - $(\text{xsec\_H\_VBF} / \text{xsec\_H\_gg})$
- Legends:
  - Set contour area **green**:  $\text{VBF/ggH} > 1\%$
  - Set contour area **yellow**:  $\text{VBF/ggH} > 10\%$
  - Set contour area **red**:  $\text{VBF/ggH} > 50\%$
- Checked for type I and type II
- They will be both shown

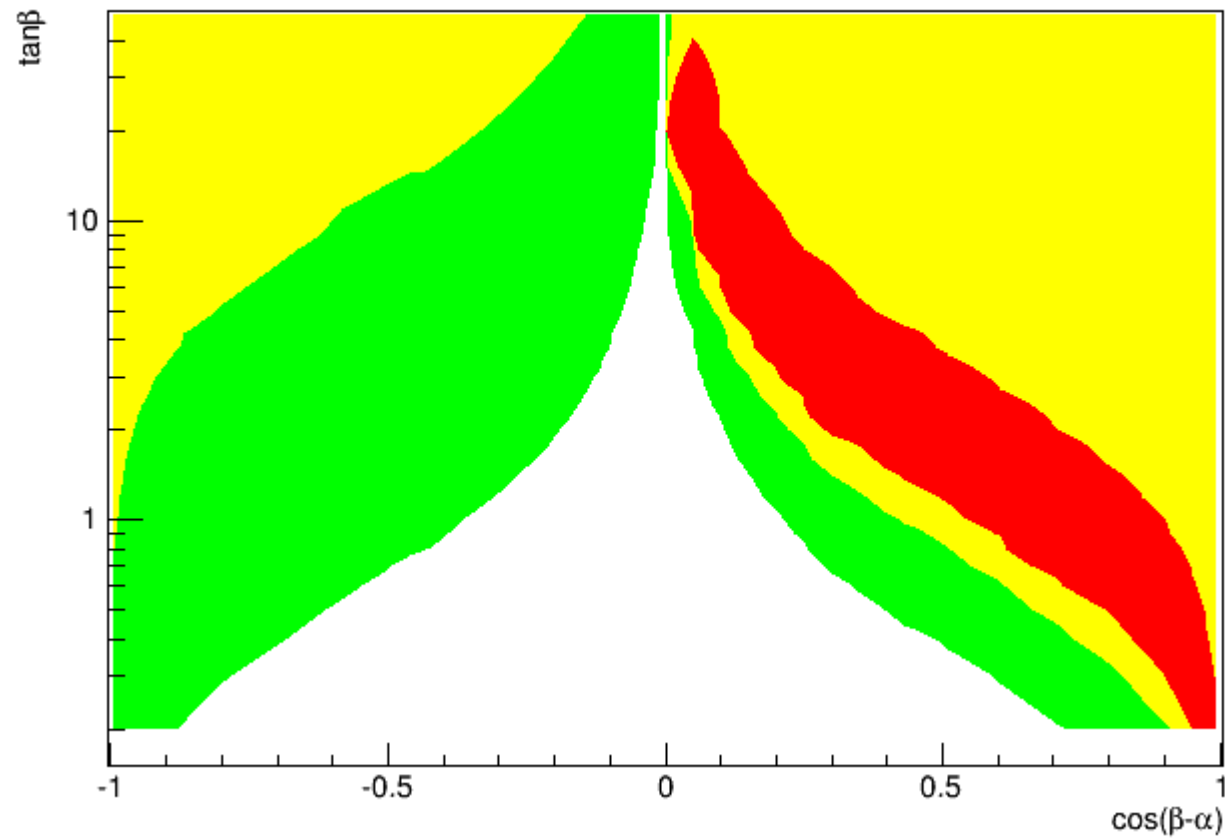
**Type I**



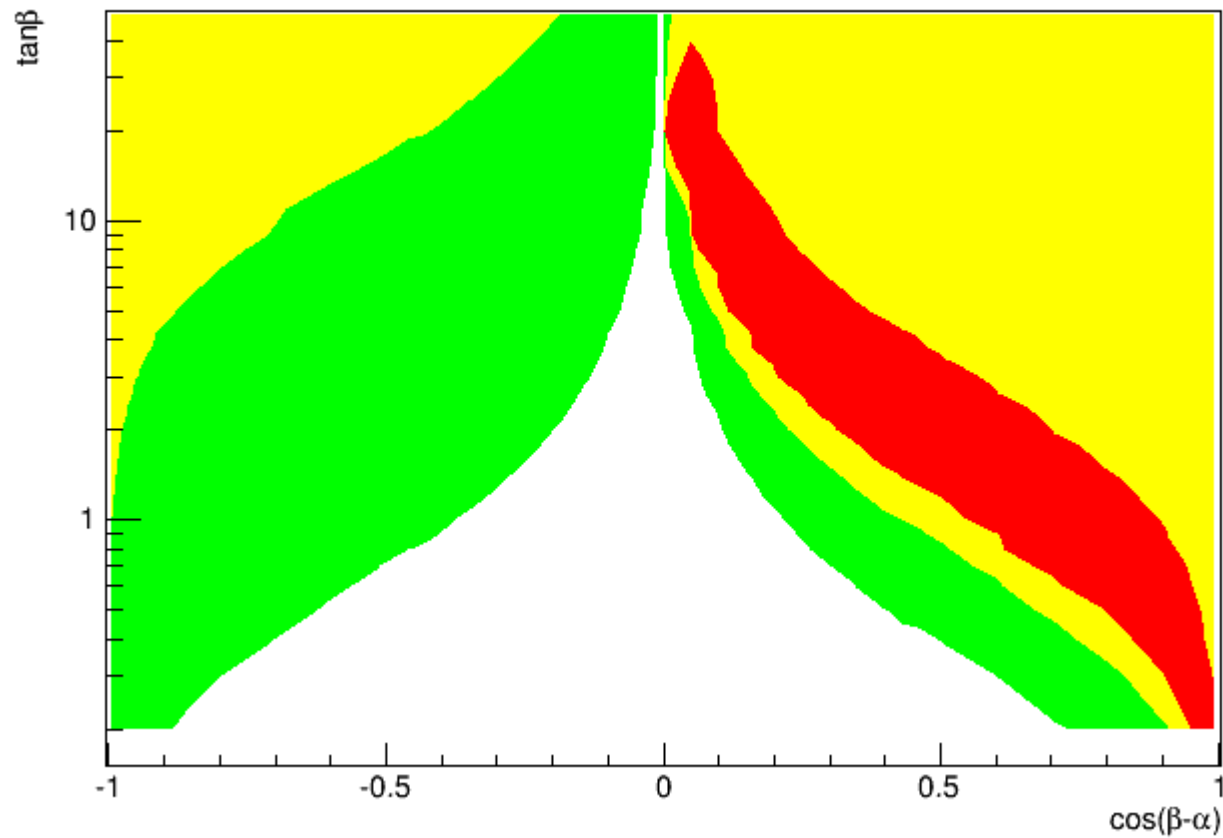
# Check width @ $mH=260$



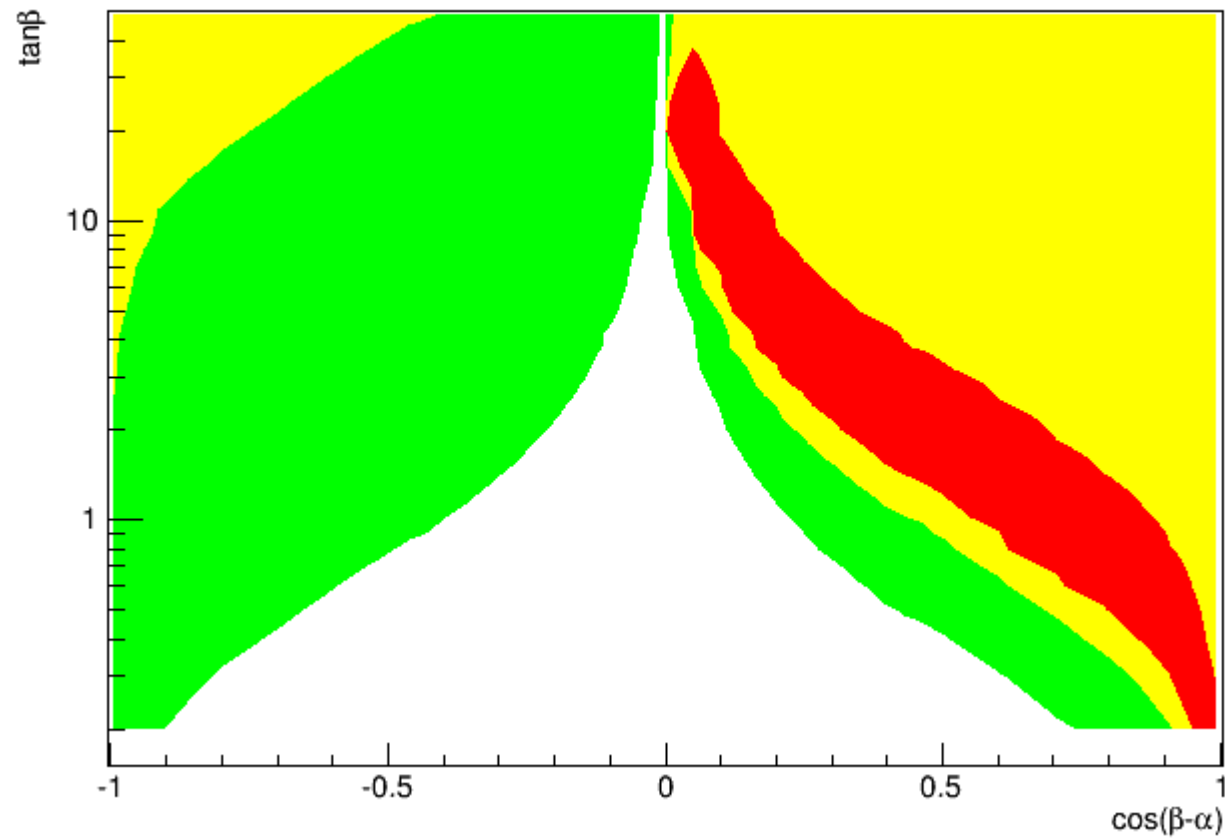
# Check width @ $mH=300$



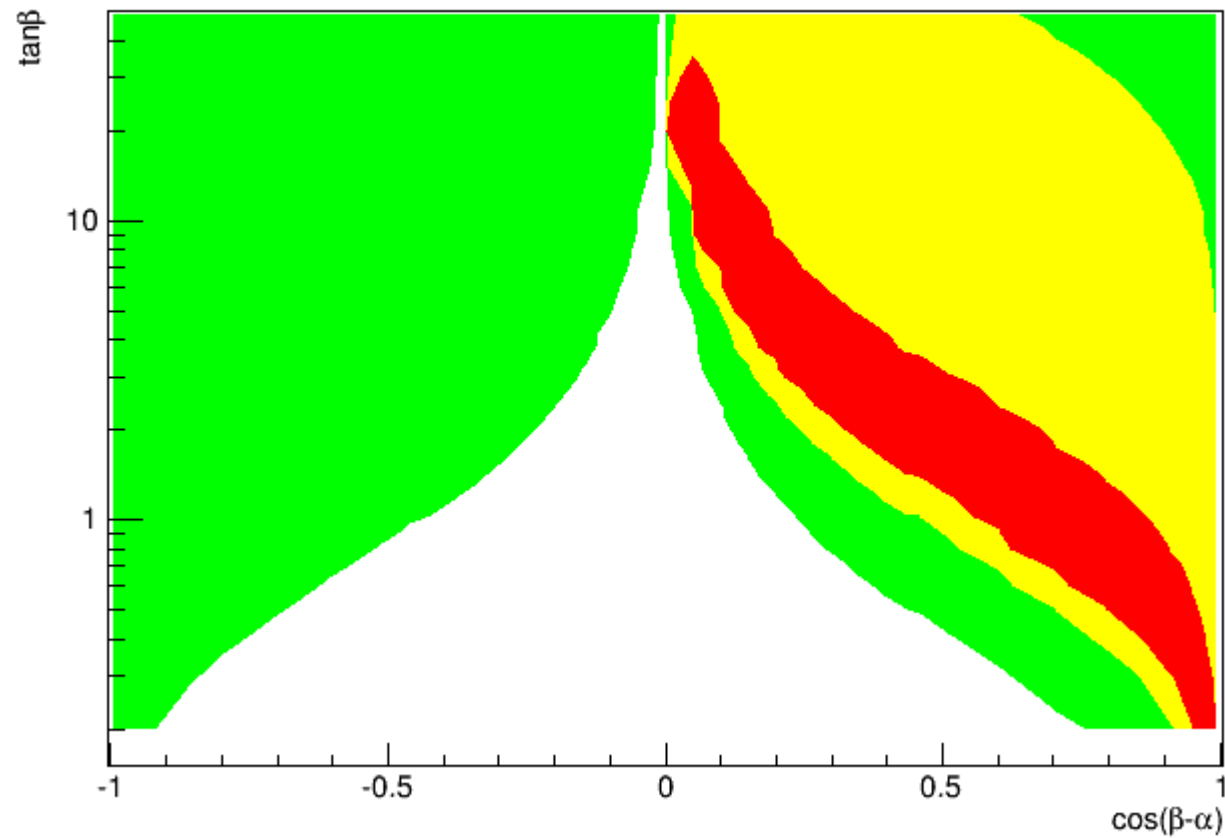
# Check width @ $mH=320$



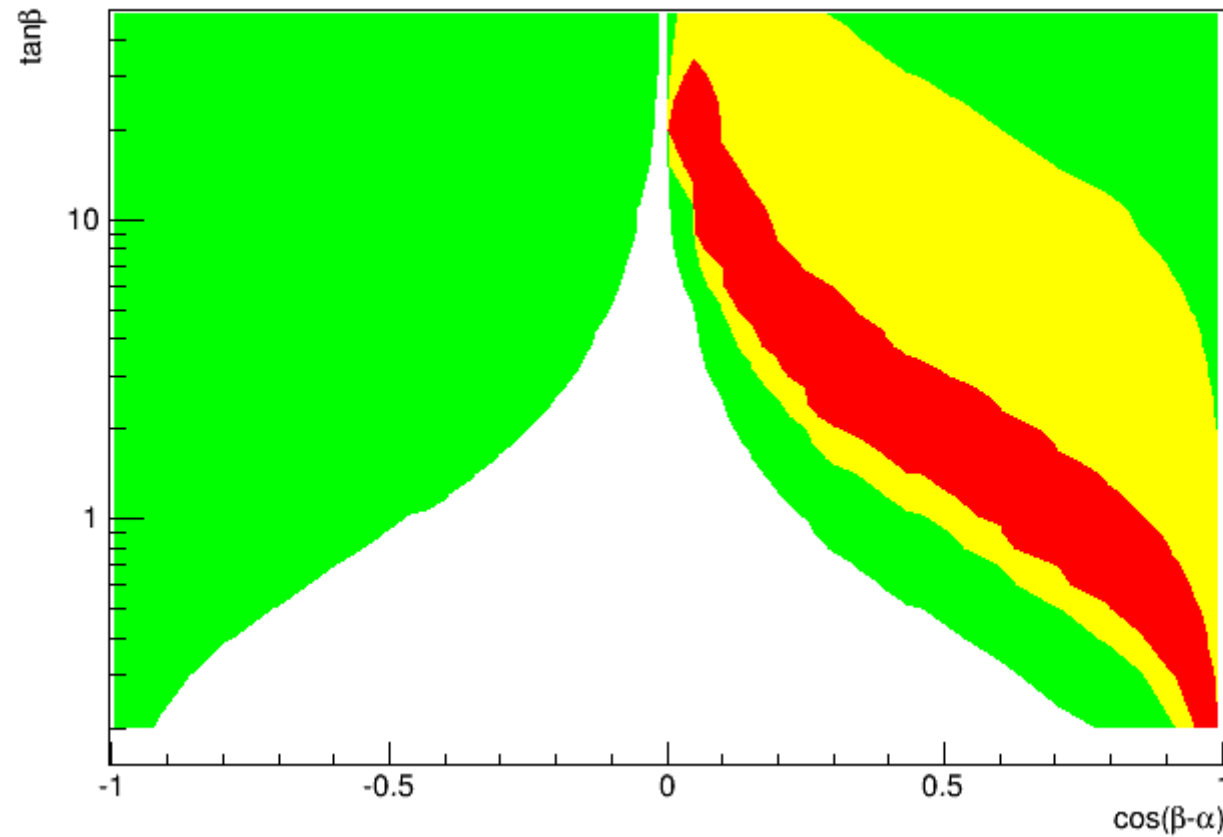
# Check width @ $mH=340$



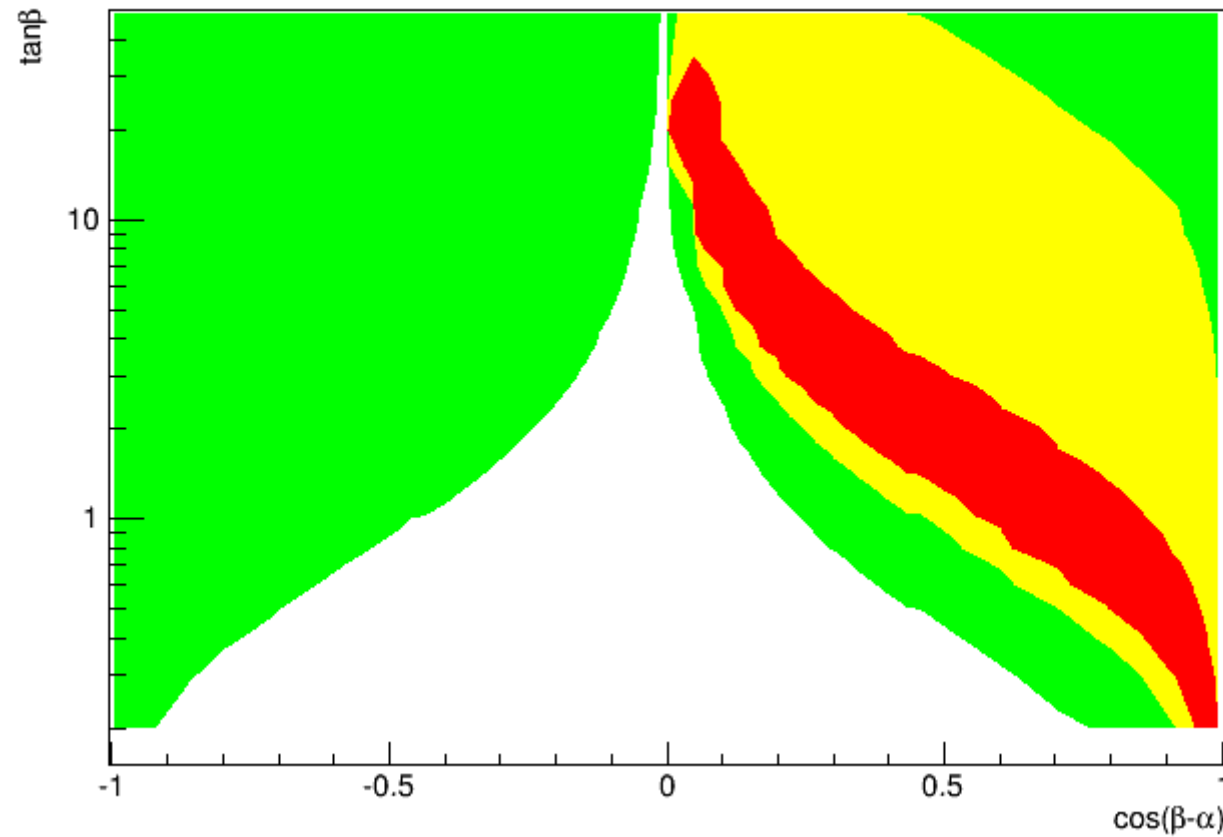
# Check width @ $mH=350$



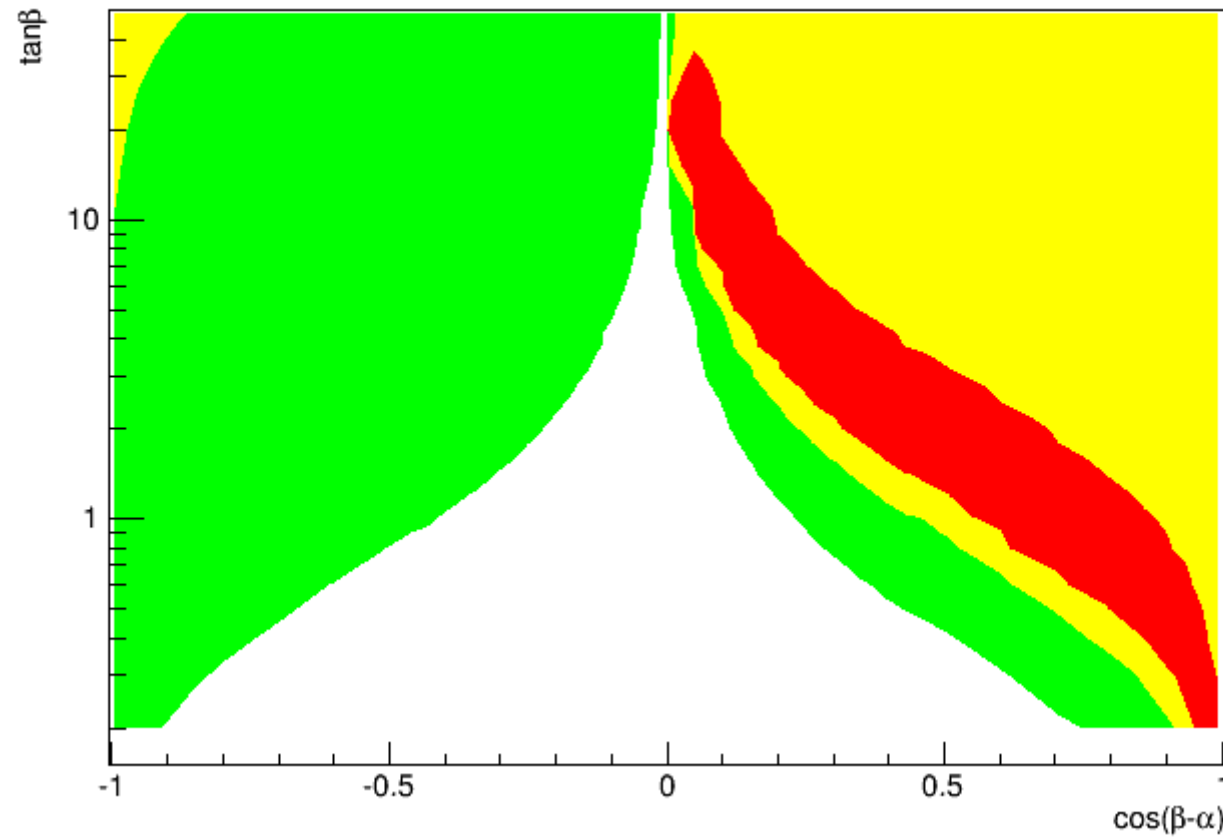
# Check width @ $mH=400$



# Check width @ $mH=420$

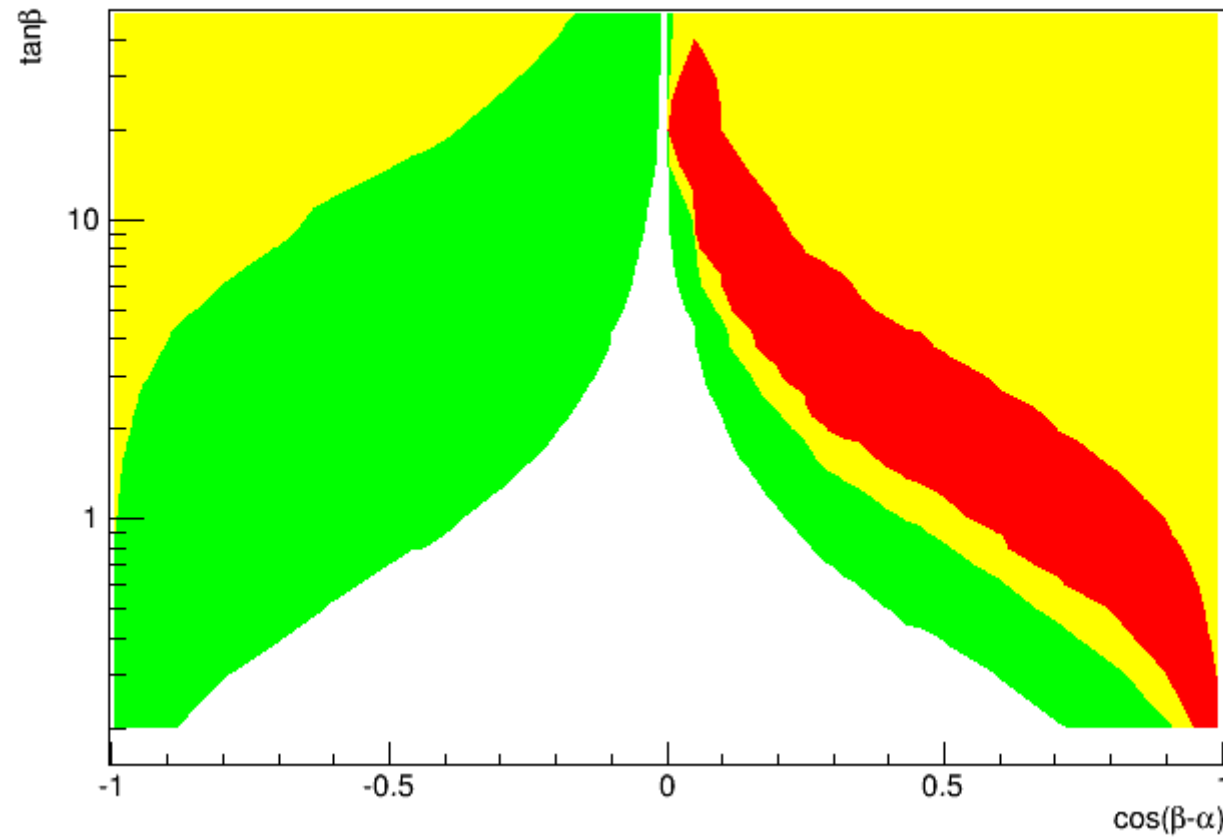


# Check width @ $mH=450$

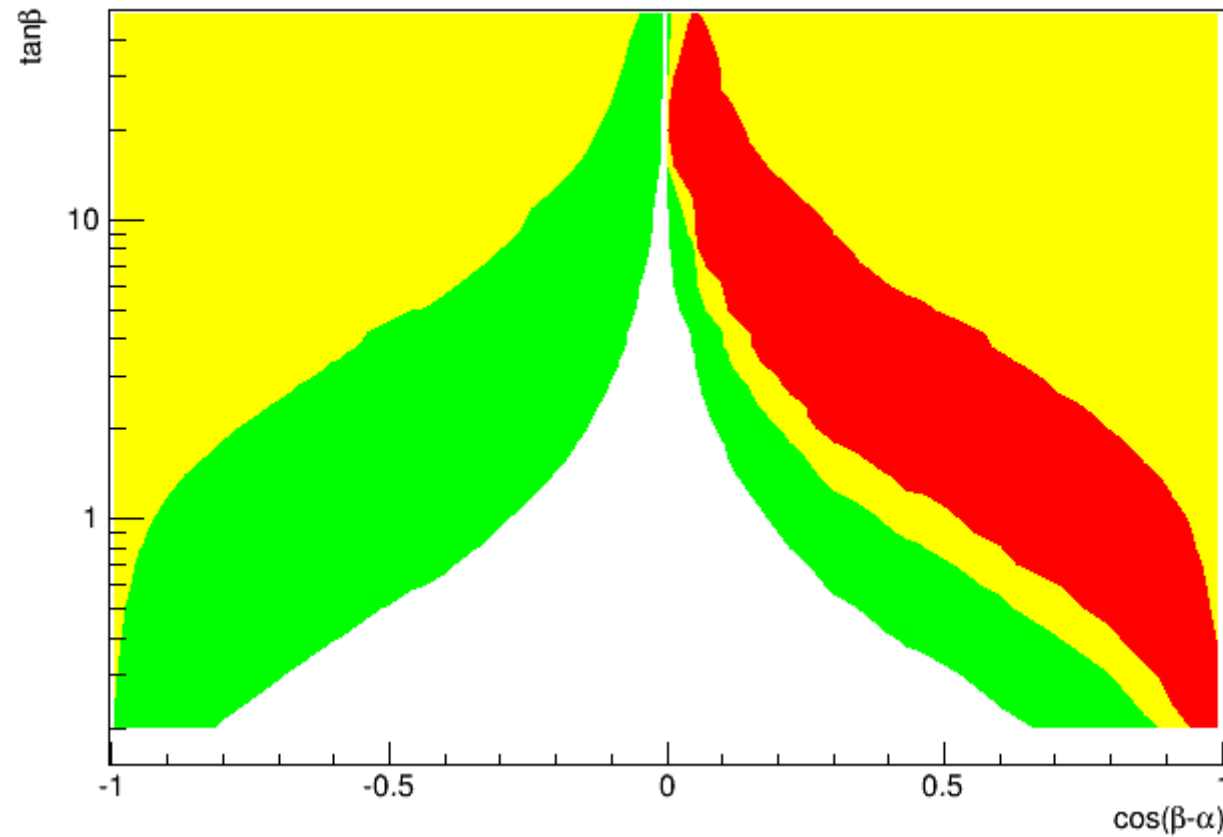




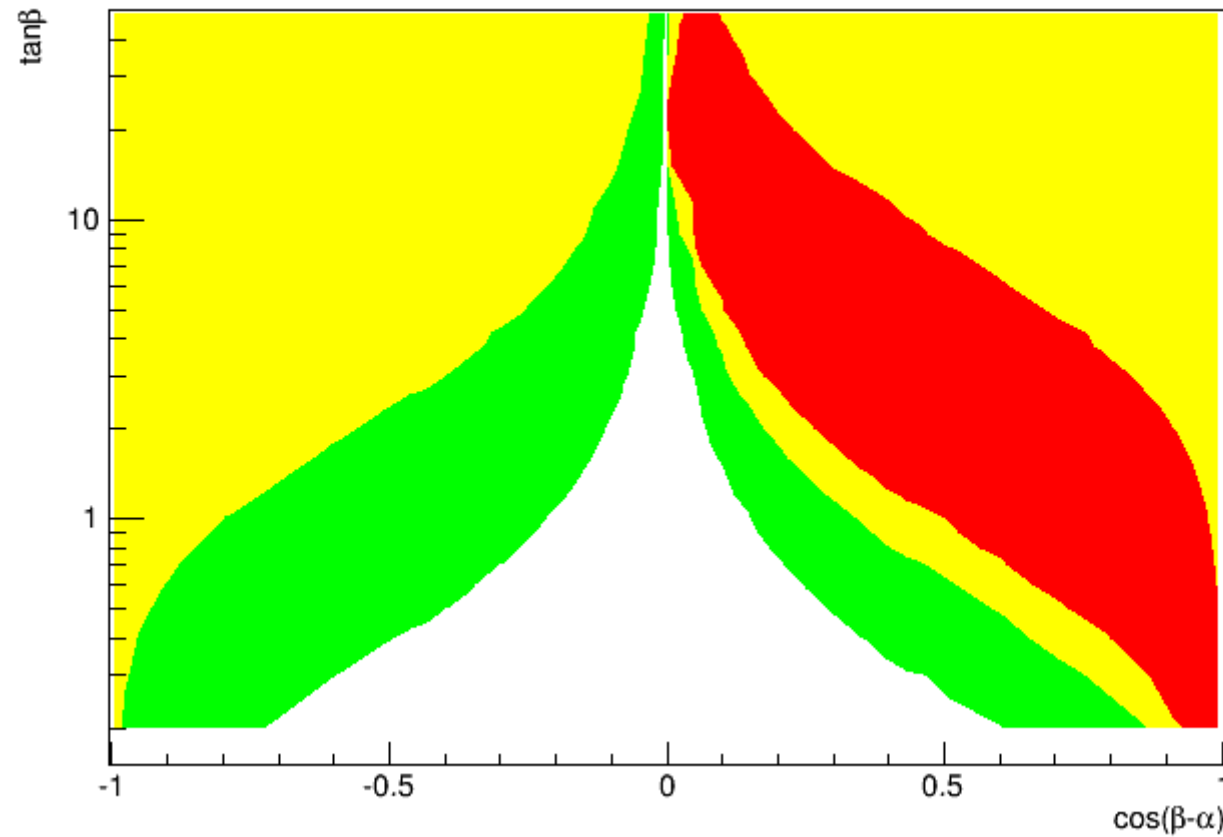
# Check width @ $mH=500$



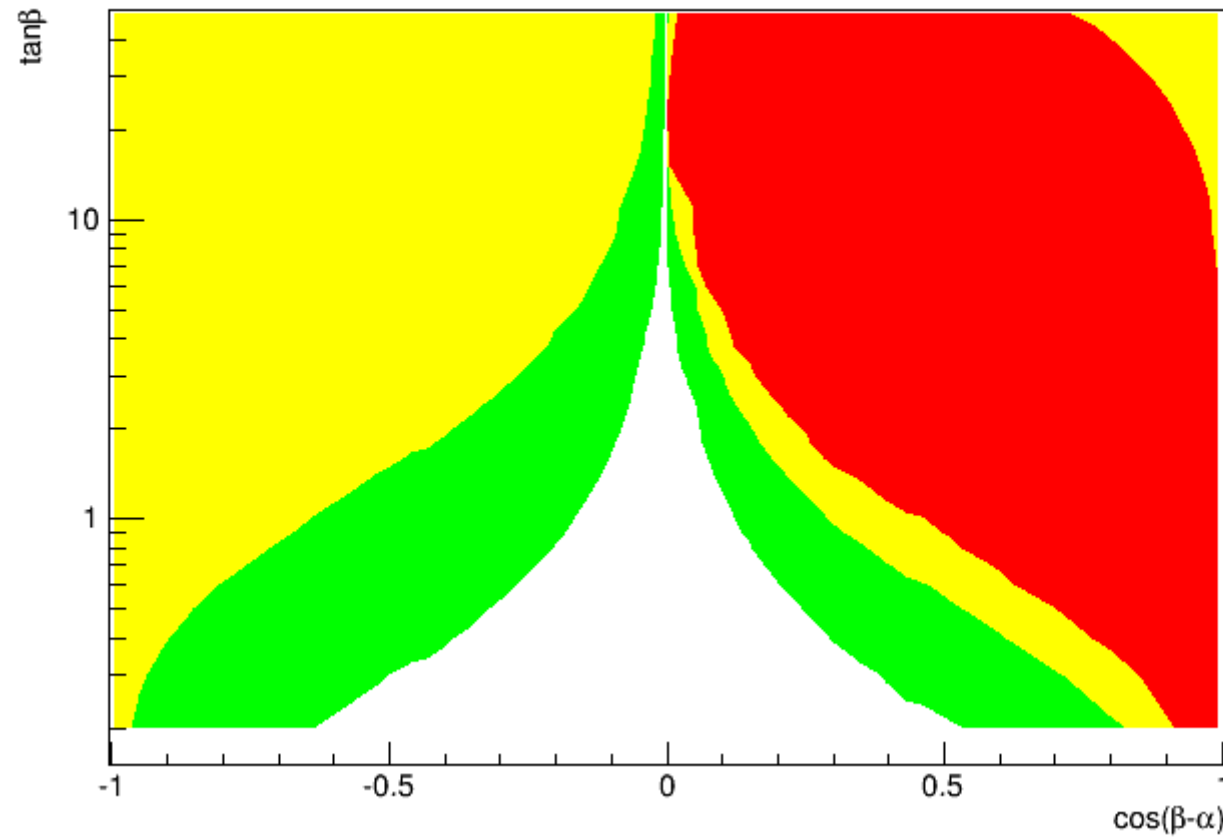
# Check width @ $mH=600$



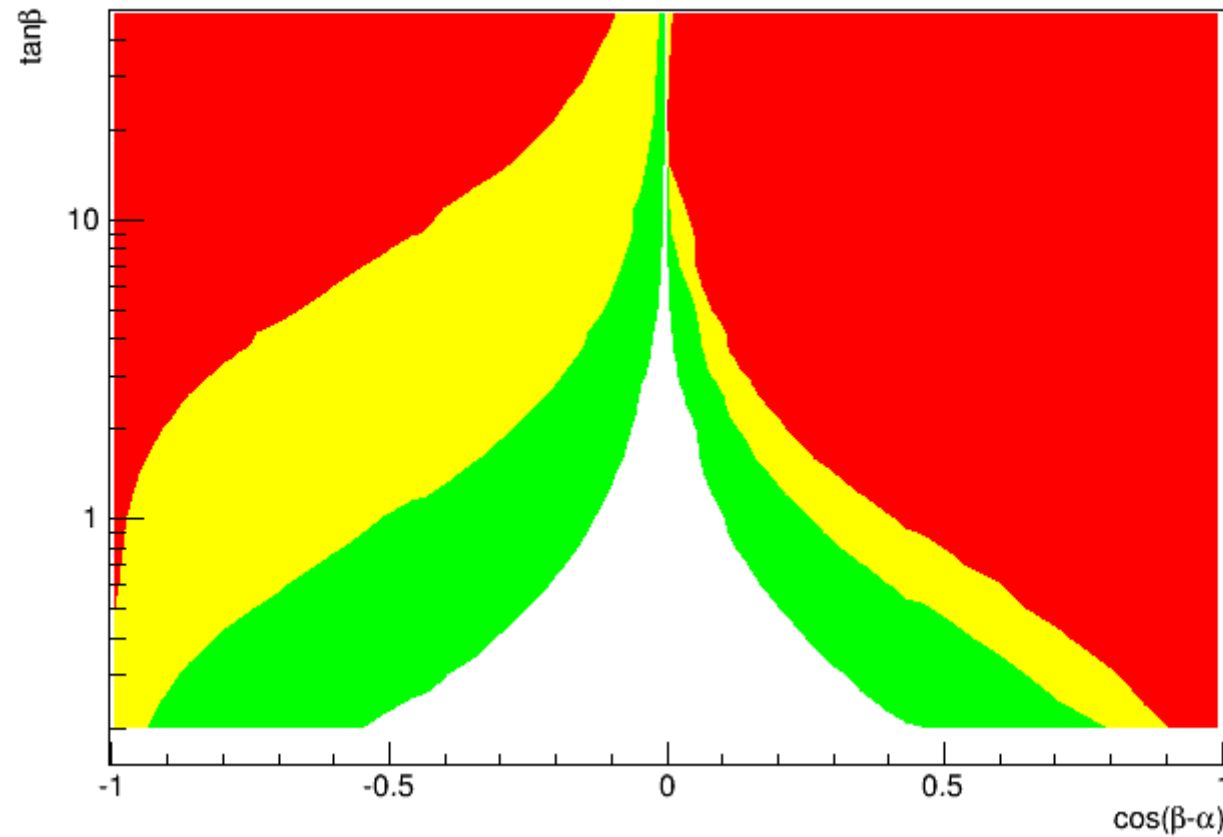
# Check width @ $mH=700$



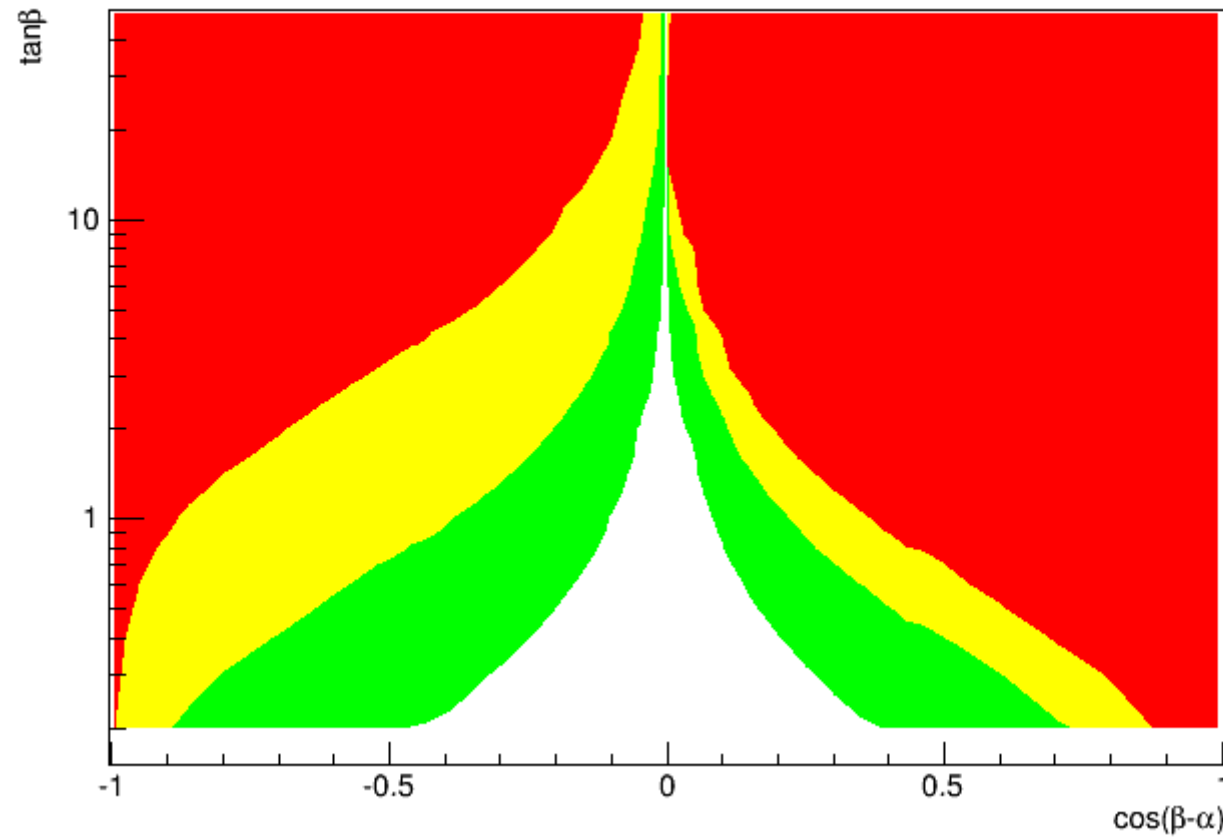
# Check width @ $mH=800$



# Check width @ $mH=900$

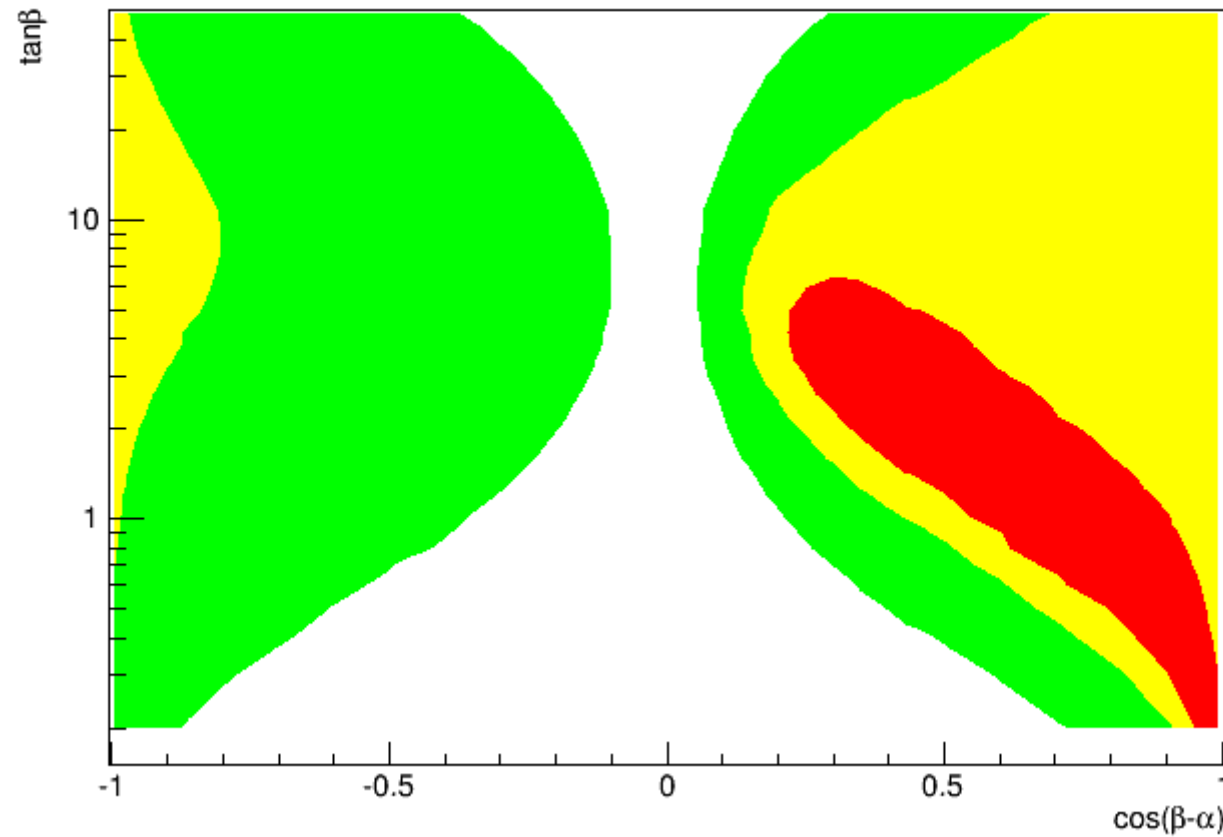


# Check width @ $mH=1000$



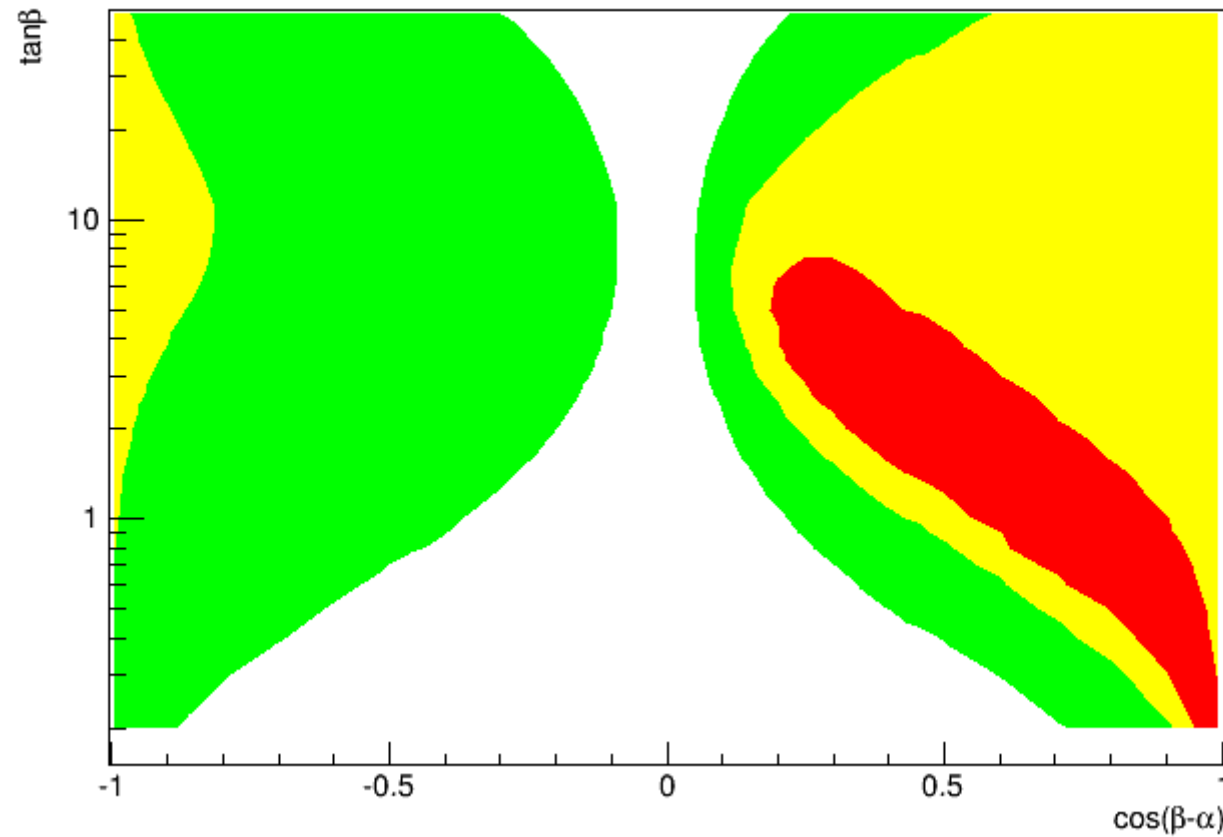
**Type II**

# Check width @ $mH=260$

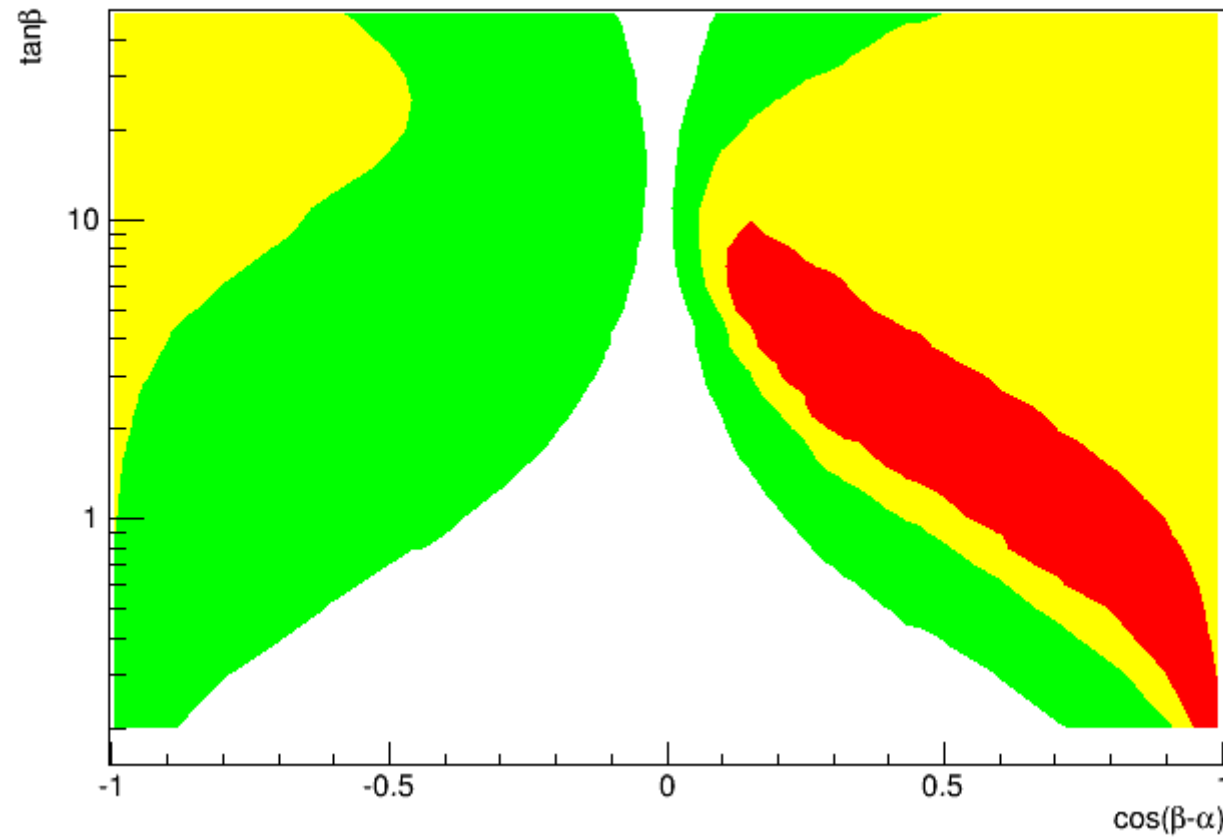




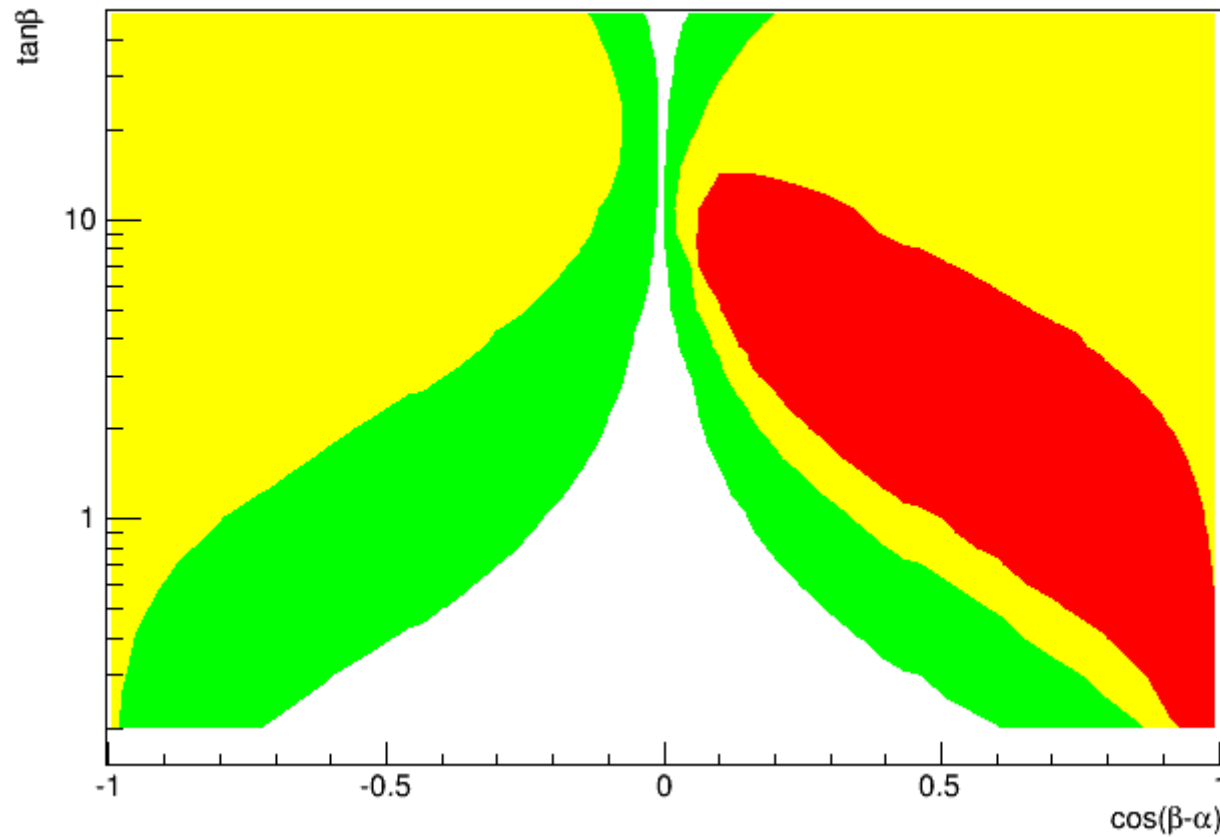
# Check width @ $mH=300$



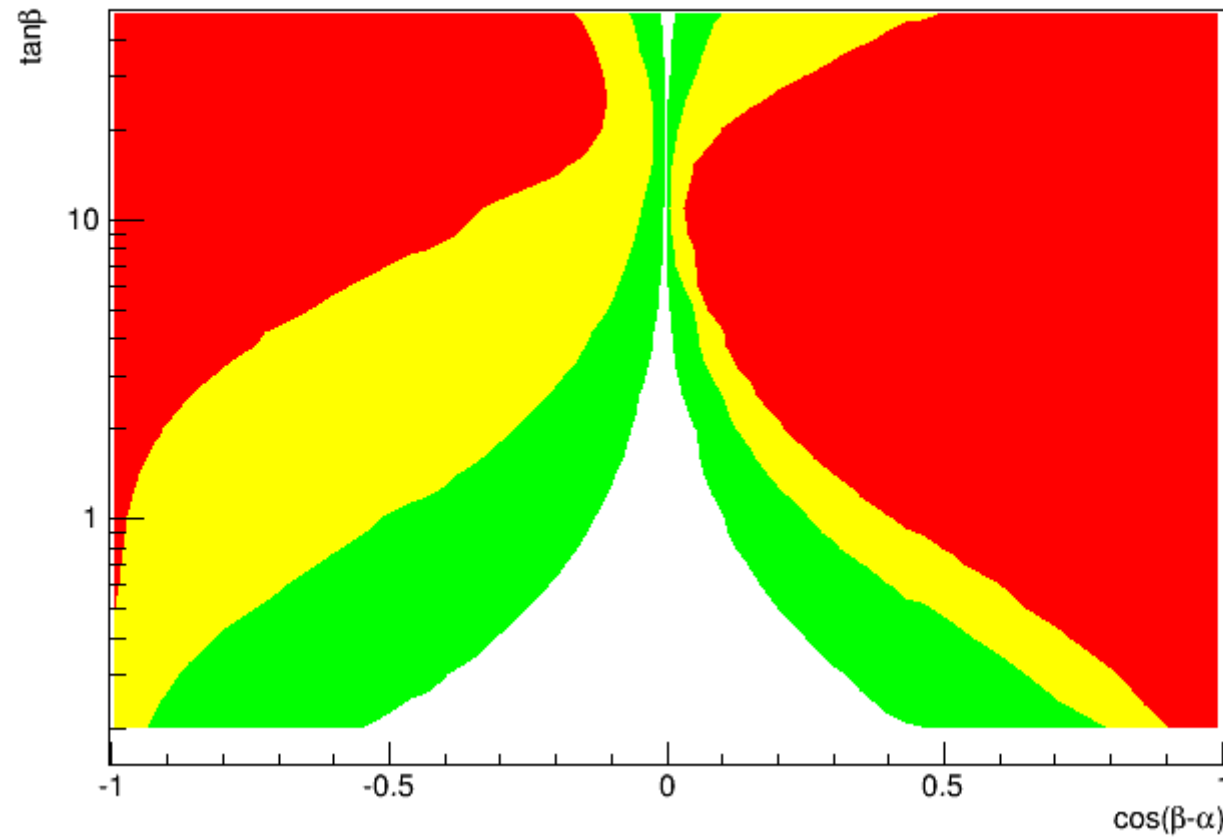
# Check width @ $mH=500$



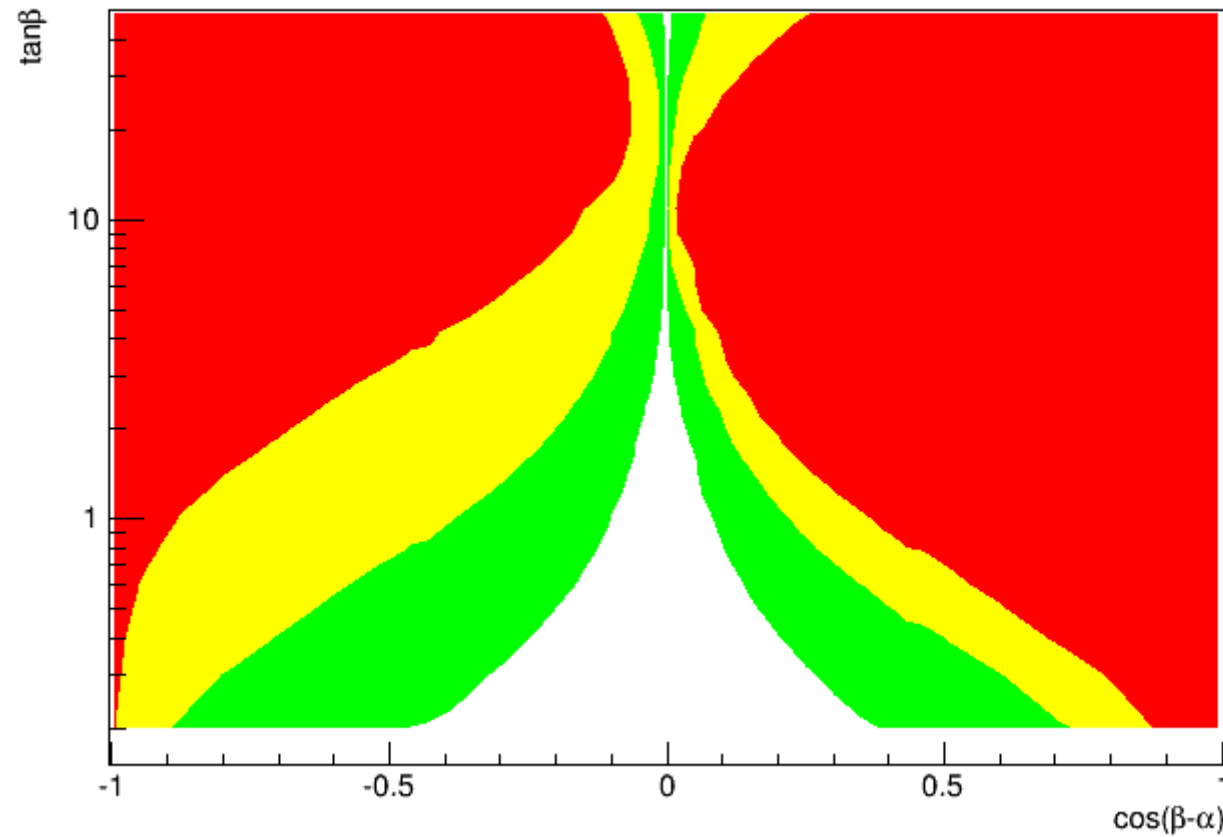
# Check width @ $mH=700$



# Check width @ $mH=900$

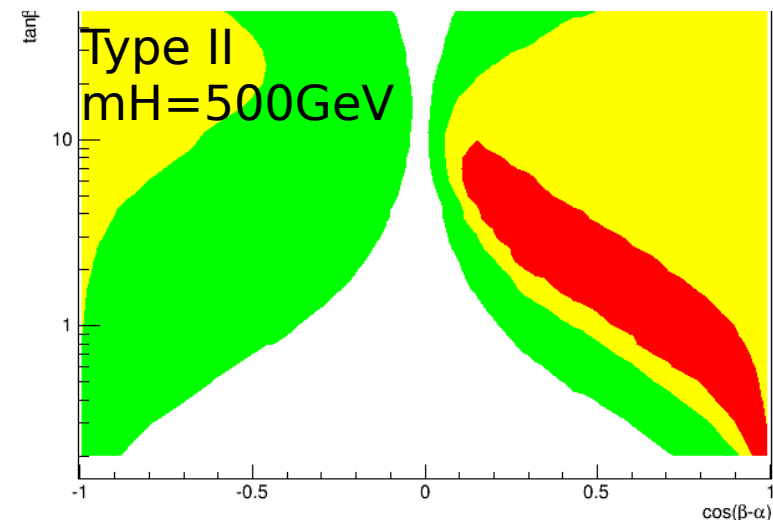
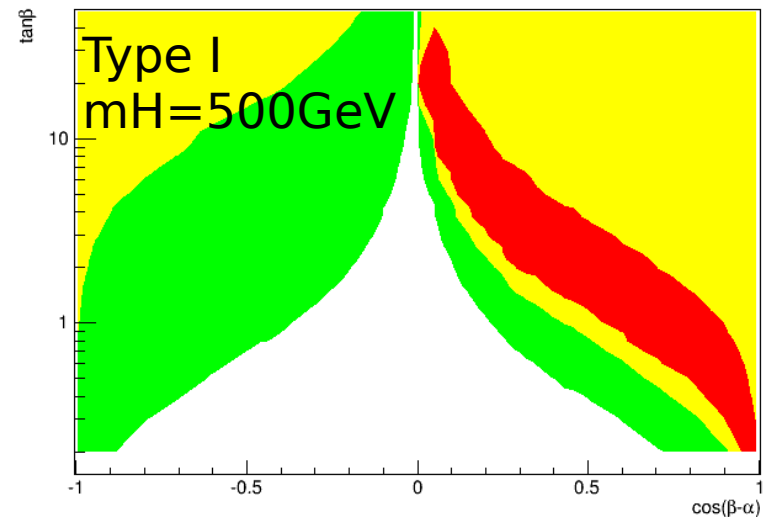
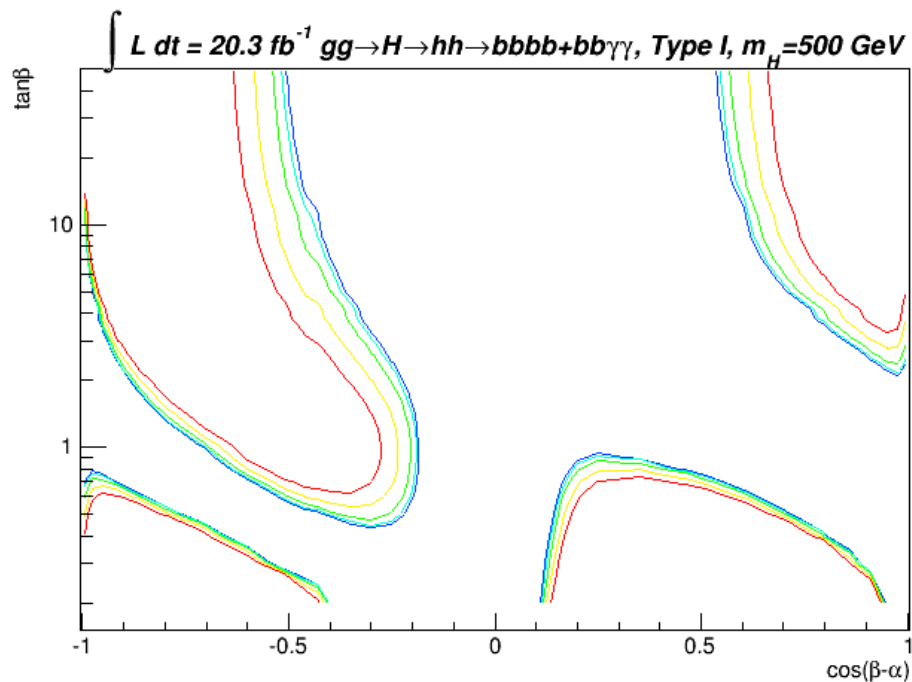


# Check width @ $m_H=1000$



# Summary on VBF production

- Seeming to be safe, the region affected by VBF the most is not in our exclusion region
- Still need to check in bbbb/bbyy analyses after including VBF samples

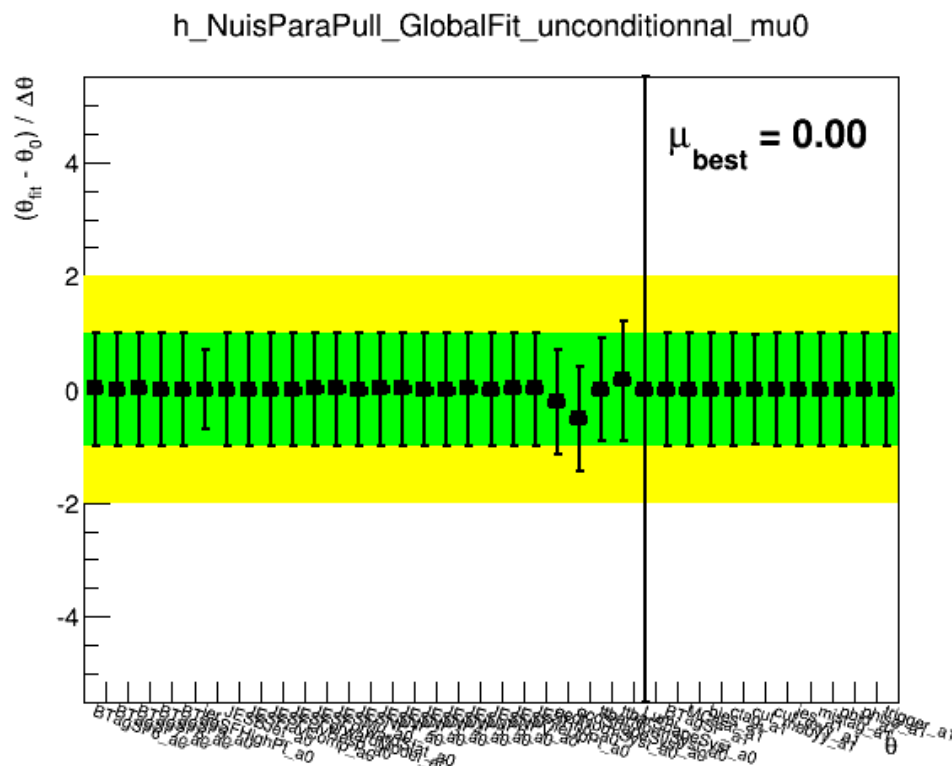


# **Part II**

## **Nuisance parameters**

# Current status

- The two groups: bbbb is finalizing graviton CONF note for Moriond; bbyy is stuck in Hfitter since last week
- bbbb ws is v1 (updated just before chinese new year); no bbyy
- The inputs are still the same since HSG7 meeting 6 Feb
- From the last time, the nuisance parameter “Lumi” did not pass our pull check, since it gave a huge error bar

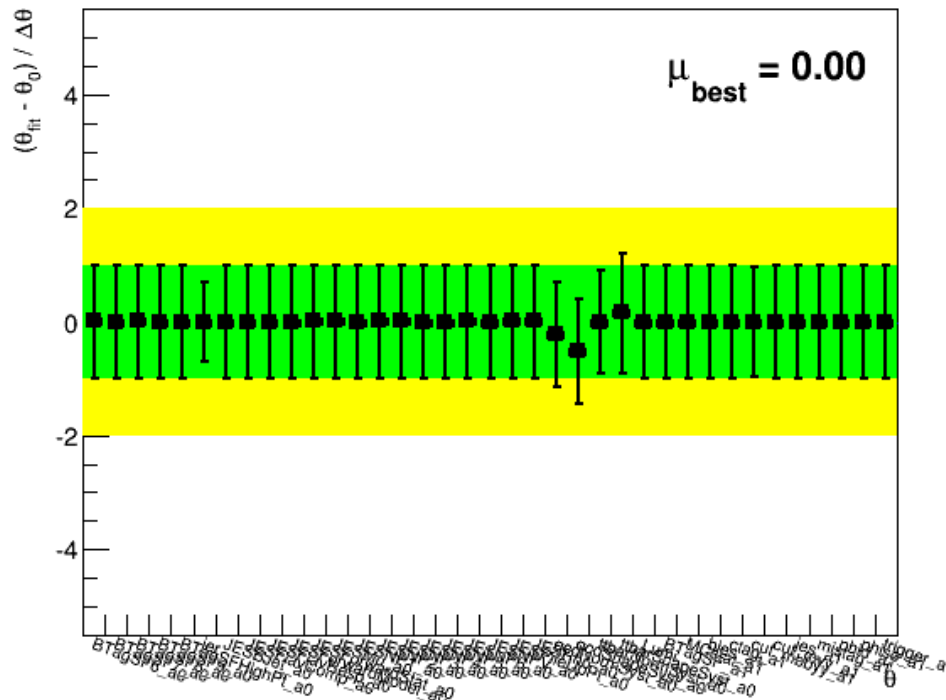




# Solve the “Lumi” problem – pull check

- Many checks were done for the combination machinery to search for the reason of the problematic Lumi
- In the end, it turned out that the problem came from the input workspace of bbbb analysis, where they set “Lumi” as a “Constant” in the fit, surprisingly to me
  - Strange, being contacting with them for an explanation
  - Solution, manually reset LUMI as a floating parameter in the combination

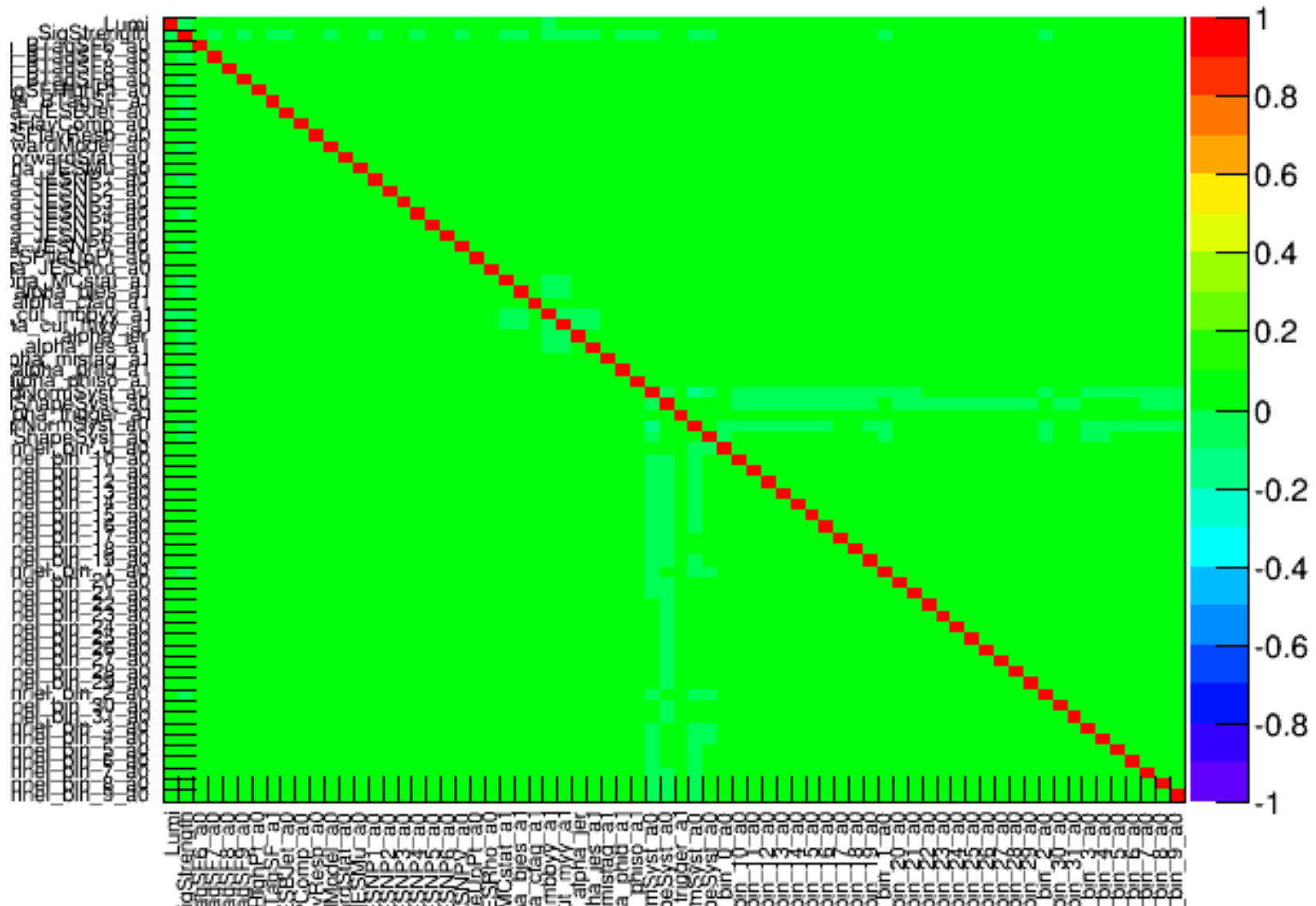
### h\_NuisParaPull\_GlobalFit\_unconditionnal\_mu0



# Correlation check

- Correlation matrix seems good

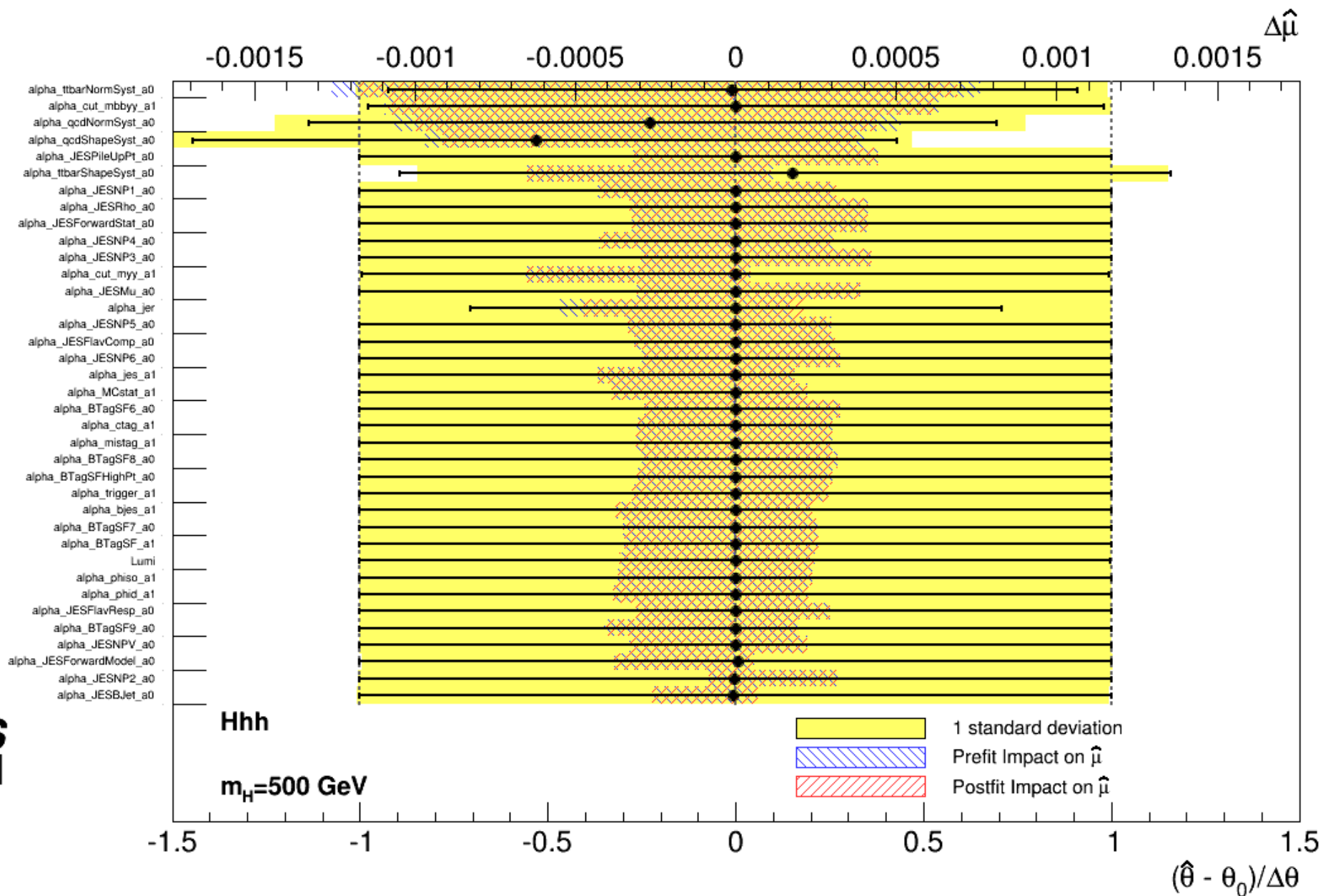
correlation\_matrix



# Importance check

- [Strongly proposed by Wouter](#), after checking the correlations and the pull distributions, we need to rank all the nuisance parameters
- The ranking follows the magnitude of the importance of nuisance parameters
- The importance is defined by the influences on POI from a certain parameter
- The importance is calculated for two types
  - Prefit importance by varying  $nui \pm 1.0$
  - Postfit importance by varying  $nui \pm \text{one sigma}$ 
    - where the sigma is estimated by assuming the NLL as a parabolic function of this nuisance parameter

# Importance check



**ATLAS**  
Internal

# Summary

- Width check and VBF check are being reported to bbbb/bbyy, and we need their studies on the effects from wide width as well as non-negligible vbf production
- For the combination, this week the focus is on nuisance parameters checks, which are being implemented in a various ways
  - Correlation check
  - Pull distribution check
  - Importance check

# COMING SOON!

- *The next time, I will start a series of lectures on the introduction to statistics, probably 2 to 3 lectures*
  - *The basics of statistics*
  - *Hypothesis testing*
  - *Limit setting*