# Solar and atmospheric V's

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> ICFA2014 Seminar at Beijing October 27, 2014

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#### 3-flavor oscillation scheme

Parameterized by 4 (mixing matrix) and 2 (difference of squared masses)



Solar&Atmospheric V's played pioneering roles in the past V oscillation studies and would also play important roles in future.



Wide range of V energy (0.1 GeV ~ 10<sup>4</sup> GeV and beyond)
Wide range of V baseline (10km downward ~ 13,000km upward)
V<sub>μ</sub>:V<sub>e</sub> ~ 2:1 at production
V oscillation study by high statistical data (>40,000 events in Super-K) and all three flavors (V<sub>e</sub>, V<sub>μ</sub>, V<sub>T</sub>)
Unique tests of V's exotic property (4th V, Lorentz violation, etc)

#### Early history of atmv's oscillations



- Dominant effect is  $v_{\mu}$  disappearance (discovered in 1998)
- Oscillatory signature (evidence in 2004)

#### Dominant effect: $v_{\mu}$ disappearance



Though consistent with accelerator V experiments atmV allows more parameter space.

 $\sin^2\theta_{23} = 0.5$  (maximal mixing) or <0.5 or >0.5 is still a open question ( $\theta_{23}$  octant problem).



Good constraints on anti-neutrino parameters. Data is consistent with CPT conservation.

### Evidence for $\tau$ neutrino appearance



 $N_{\tau}^{DATA}/N_{\tau}^{exp} = 1.42 \pm 0.35(stat) + 0.14_{-0.12}(syst)$ 

3.8 σ significance for null T 3 τ observation @ OPERA (PRD89,051102(R) (2014))

> $v_{\mu} \rightarrow v_{\tau}$  channel has been established. Atmospheric v anomaly (problem) is finally concluded.

PRL 110, 181802 (2013)

#### Zenith Distribution of T-like events



#### Atmospheric $V_e$ oscillation



Through the matter effect in the Earth, we study on

- Mass hierarchy : resonance in multi-GeV ve or ve
- CP δ
- $\theta_{23}$  octant
- : magnitude of the resonance

: interference btw two  $\Delta m^2$  driven oscill.



### MH and $\delta_{CP}$ study



- Preference of normal hierarchy:  $\chi^{2}_{IH}-\chi^{2}_{NH}=0.9$  (SK only) I.2 (T2K combined).
- Preference of  $\delta_{CP}$  near  $3\pi/2$  is also strengthened by T2K • CP conservation (sin $\delta_{CP}$ =0) is still allowed at 90% CL
- Need more data

### Sterile neutrino oscillations

4th v(Sterile v) is indicated by LSND, MiniBooNE, reactor, solar  $v \exp{}^{2}s$ 

$$U = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} & U_{e4} & \cdots \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} & U_{\mu 4} & \cdots \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} & U_{\tau 4} & \cdots \\ U_{s1} & U_{s2} & U_{s3} & U_{s4} & \cdots \\ \vdots & \vdots & \vdots & \vdots & \ddots \end{pmatrix}$$

- independent on the number of sterile Vs and  $\Delta m^2_{sterile}$ 
  - constraints for 3+1 framework can be extended to 3+N
  - constant constraints as long as  $\Delta m^2_{sterile} > 0.0 \, I \, eV^2$
- U<sub>µ4</sub>→uniform µ deficit in angle/ energy
- $U_{\tau 4} \rightarrow$  shape distortion in angular distributions of high energy  $\mu$



#### Limits on sterile V contribution

arXiv:1410.2008, submitted to PRD



Atmospheric V data is consistent with 3-active V framework (no sterileV contribution), and provides limits on mixings.

### Lorentz invariance violating oscillations

arXiv:1410.4267, submitted to PRD



Provided 3~7 orders of magnitude better constraints than past

# Future prospect of atmv studies: mass hierarchy determination



- Proton decay searches extends to discovery region
- Studies toward understanding full picture of neutrino oscillations
  - CPV
  - MH determination for any  $\delta$
  - $\theta_{23}$  octant determination
- Astrophysical neutrino observatory
  - Supernova V
  - Solar neutrino (~200v ev/day)



Whole allowed parameter space is covered by atmv only study

#### More for hierarchy study

#### IceCube+DeepCore+PINGU ~2020-E threshold down to O(I)GeV Ê10 ≻ IceCube DeepCore 50 PINGU 75m 125m • 20m -50 -100 -150 200 X (m) -50 50 100 150 -100 n For more details see the PINGU Lol, arXiv:1401.2046 Tracks Cascades Multichannel NMH significance **IH** True Preliminary 0 2 8 10

PINGU livetime [vrs]

#### INO ~2020-50kton Iron Calorimeter



More than 3  $\sigma$  sensitivities in both programs

#### Solar neutrino studies

- $\bullet$  Pure  $\nu_e$  source from pp-chain reactions and CNO cycle
- High statistical data available (>70,000 events in Super-K)
- Study on neutrino properties
- Search for new physics beyond the standard model
- Unique probe of solar structure & solar system formation

### Latest oscillation results (global fit)

#### Combined solar fit w/ KamLAND

Without reactor  $\theta_{13}$  constraint



#### ~2 $\sigma$ tension in $\Delta m^2_{21}$ between solar and KamLAND

Non-zero θ<sub>13</sub> at ~2σ by solar and KamLAND only
Good agreement with sin<sup>2</sup>θ<sub>13</sub>=0.0221±0.0012 by Daya Bay, RENO, and DC



First indication (at 2.8~3.0  $\sigma$ ) of terrestrial matter effects

could be important achievement for future LBL and atmospheric V studies in which the matter effect is relevant

### Survival probability (spectrum)



Gabriel D. Orebi Gann @Neutrino2014

### Probing the Unknown

#### Non-standard physics effects can alter the shape / position of the "MSW rise"

#### Non-standard interactions (flavour changing NC) **Sterile Neutrinos** Mass varying 8 0.7 C MA-Opep MA-Ipep neutrinos (MaVaNs) LMA-0 B MA-I B 0.6 no sterile $\sin^2 2\alpha = 1 \times 10^7$ 0.4 $\sin^2 2\alpha = 2 \times 10^7$ 0.5 0,2 $\sin^2 2\alpha = 5 \times 10$ R.=0.08 م**ٹ** ہے ا 0.4 survival probabiliy $\Delta m_{0.21}^2 = 0.000175 \text{ eV}^2$ tan<sup>2</sup> v<sub>12</sub>=0.5 0.8 0.3 <sup>7</sup>Be B SND 0.7 bep pр $= R_{A} = 0.2$ 0.6 0.2 α's=0 0.5 10-1 10 0.4 E./MeV 0.4 α's≠0 0.3 0.2 0.2 R\_=0.25 Friedland, Lunardini, Peña-Garay, 0.1 neutrino energy (MeV) 0 PLB 594, (2004) 10 1 10 $E_{\nu}$ (MeV) Holanda & Smirnov PRD 83 (2011) 113011

M.C. Gonzalez-Garcia, M. Maltoni Phys Rept 460:1-129 (2008)

### Lower threshold in Super-K



High-speed front-end electronics "record every hits" + Realtime reconst.+reduction

"software trigger"







More data at the transition region is coming soon.

#### SNO+







#### measurement errors

	рер	<sup>8</sup> B	<sup>7</sup> Be	pp	CNO
l yr	9%	7.5%	4%	~ a	a. 15.0/
2 yr	6.5%	5.4%	2.8%	few %	~ 13 %

Also to cover  $\beta\beta$  decay search, geo-V, SNV, proton decays...

### More on future prospect

∆chi<sup>2</sup> profile for fixed pep and CNO rates



• Borexino (phase-II) continues the hunt for CNO cycle v aiming to solve metallicity problem of the Sun.



- Clean (liquid Neon), XMASS (liquid Xenon) to measure pp-neutrino with 1%-level precision
- LENA for CNO and low energy <sup>8</sup>B
- JUNO, RENO precision measurements of parameters

### Summary of atm & solar v studies

#### • Oscillation studies toward full understanding of v properties

- Solar  $\theta_{12}$ ,  $\Delta m^2_{21}$  measurements consistent with reactor results
- Indication of terrestrial matter effect
- Atmospheric V provides complemental measurements of  $\theta_{23}$ ,  $\Delta m^{2}_{32}$
- $\overline{\nu}_{\mu}$  disappearance consistent with  $\nu_{\mu}$  (CPT conservation)
- Need more data for mass hierarchy determination,  $\theta_{23}$  octant, and  $\delta_{CP}$

#### • $v_{\mu} \rightarrow v_{\tau}$ channel as a solution of "atmospheric v anomaly"

• No indication of exotic effects such as neutrino decays, sterile V, Lorentz invariance violation, etc.

### • More solar neutrino experiments and data for oscillation studies, tests of exotic scenarios, and astrophysical studies