



The Belle II Experiment and SuperKEKB Upgrade

Wang, Boqun (University of Cincinnati)

On Behalf of Belle II Collaboration

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PhiPsi2015, USTC, Hefei, China

The B Factories

A e^+e^- collider runs at $\Upsilon(4S)$ resonance to produce B meson pairs.



Belle/KEKB at KEK (Japan)

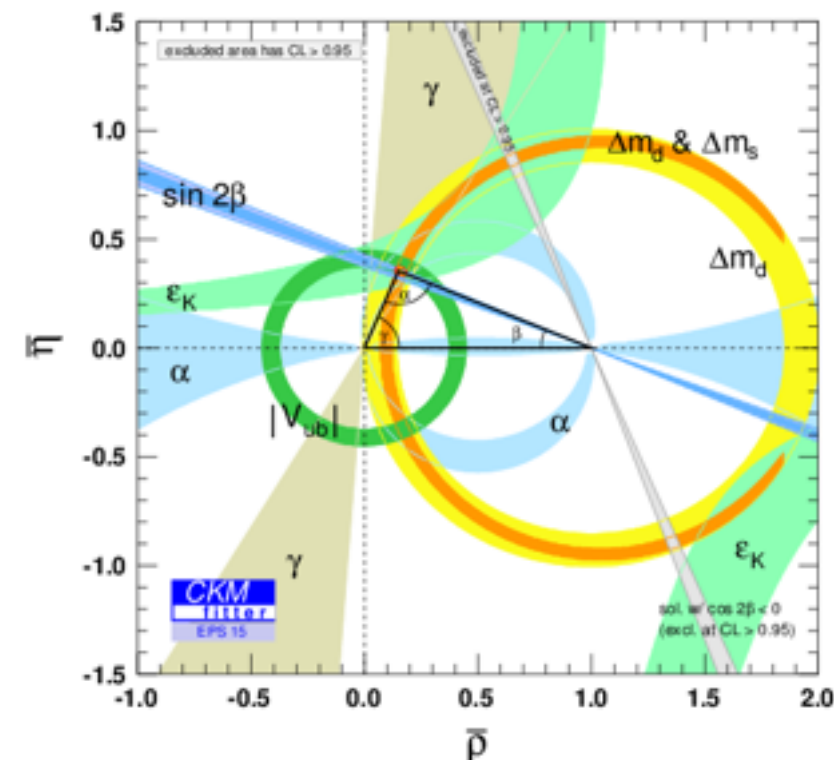
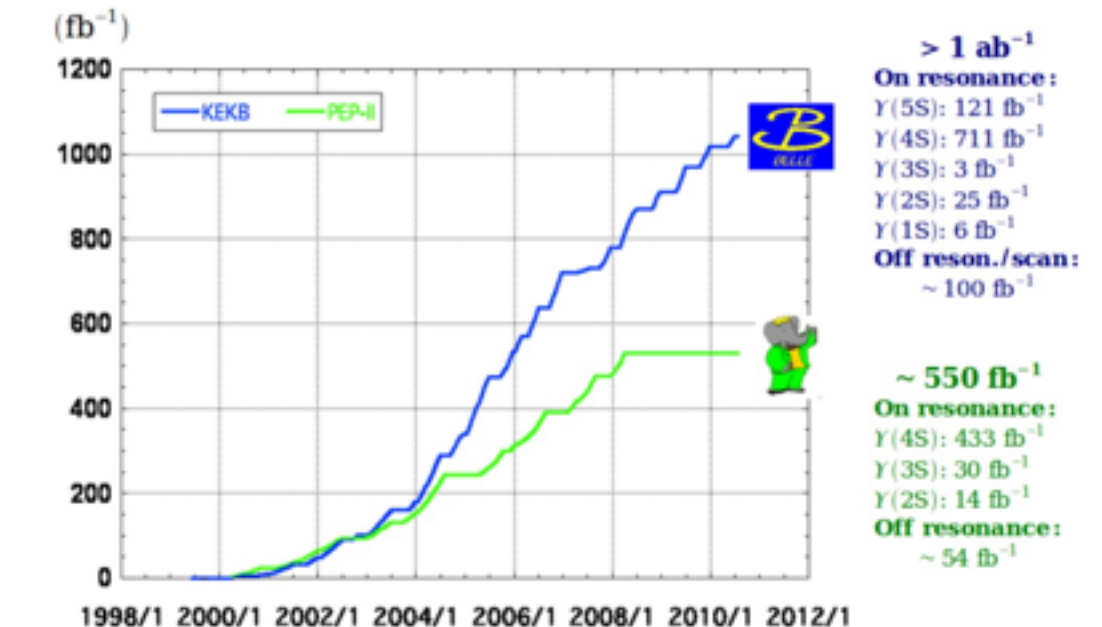


BaBar/PEP-II at SLAC (USA)

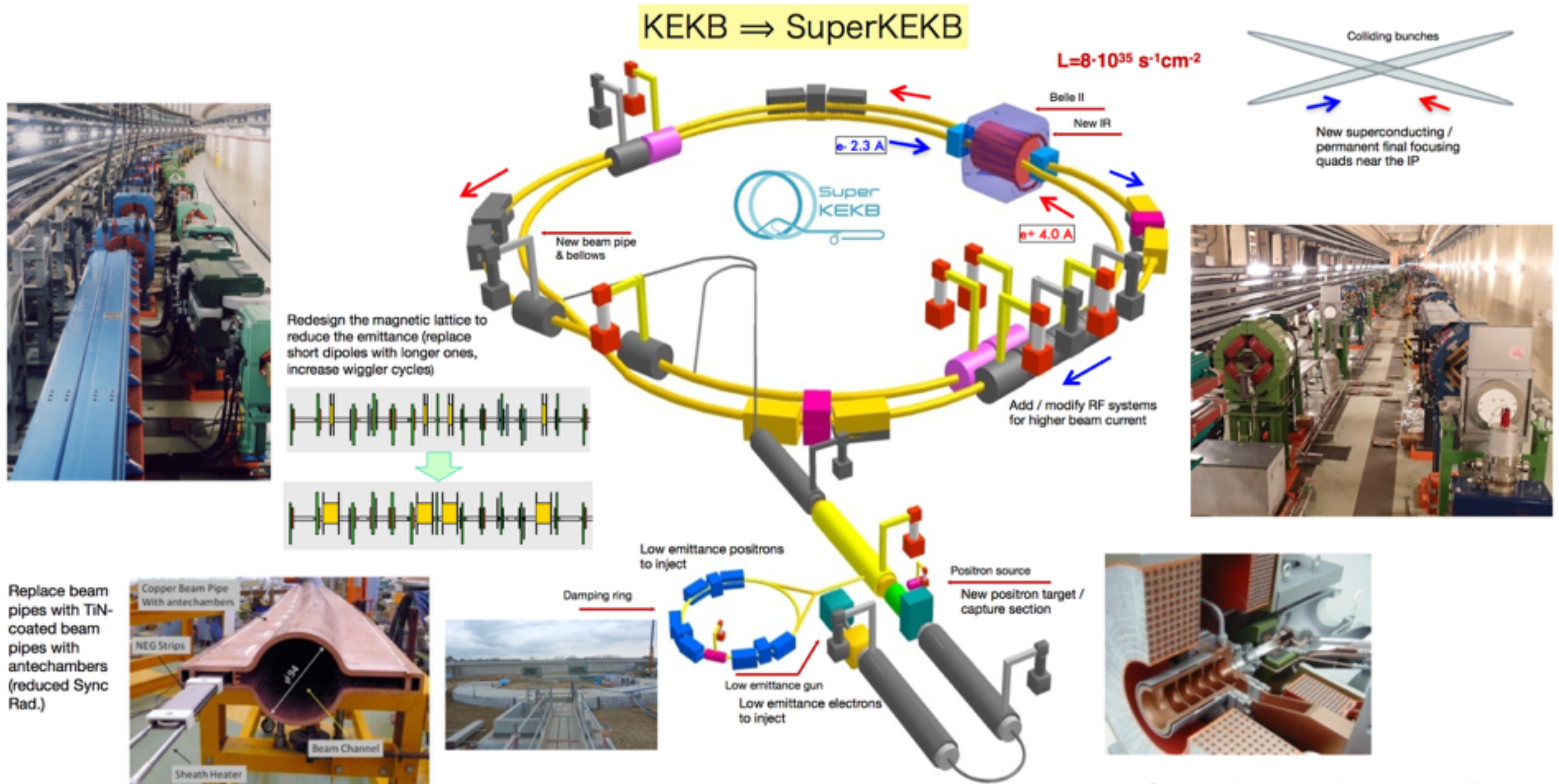
Physics Highlights

- Collected $\sim 1.5 \text{ ab}^{-1}$ of data
- Measurement of the Unitarity triangle, and CKM parameters
- Observation of direct CP violation in B decays
- Observation of D meson mixing
- Observation of new (X, Y, Z) hadrons
- Search for rare tau decays
- Constraints on new physics
- Direct searches for light Higgs, dark photon, etc.

Integrated luminosity of B factories



Super-KEKB



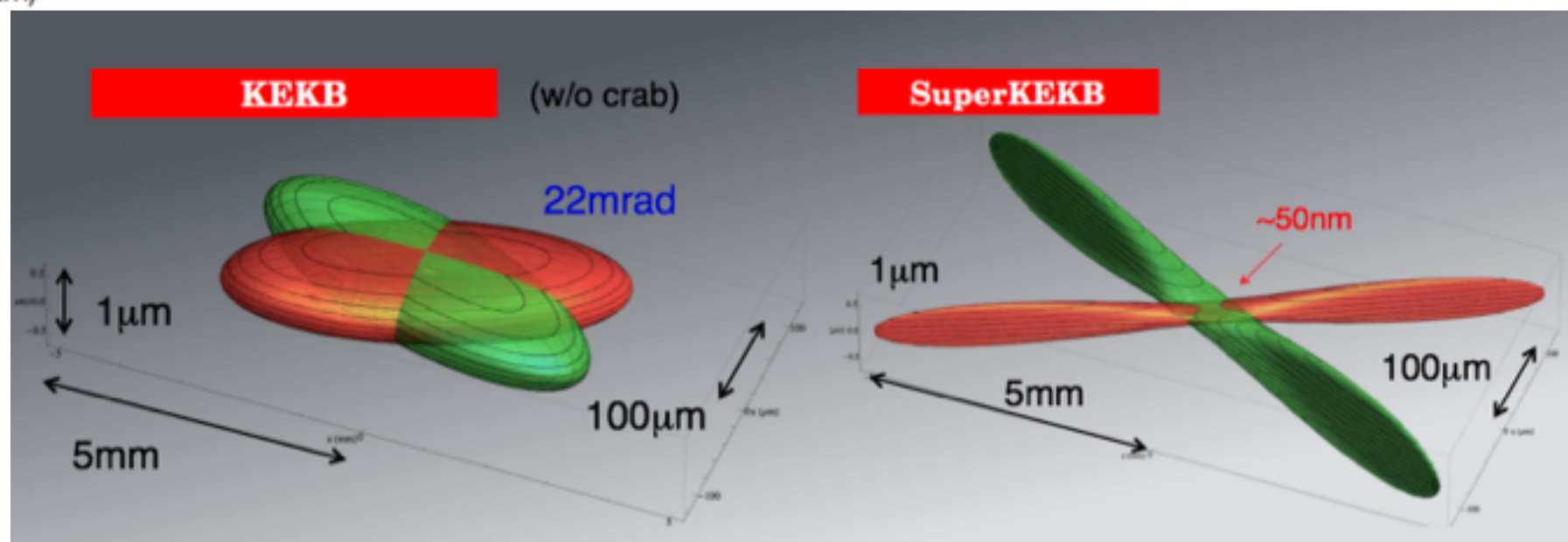
Target integrated luminosity for Belle II/SuperKEKB: $\sim 50 \text{ ab}^{-1}$

How to achieve 40x luminosity

$$L = \frac{\gamma_{e\pm}}{2er_c} \left(1 + \frac{\sigma_y^*}{\sigma_x^*}\right) \left(\frac{I_{e\pm} \xi_{e\pm}}{\beta_y^*}\right) \left(\frac{R_L}{R_{\xi_y}}\right)$$

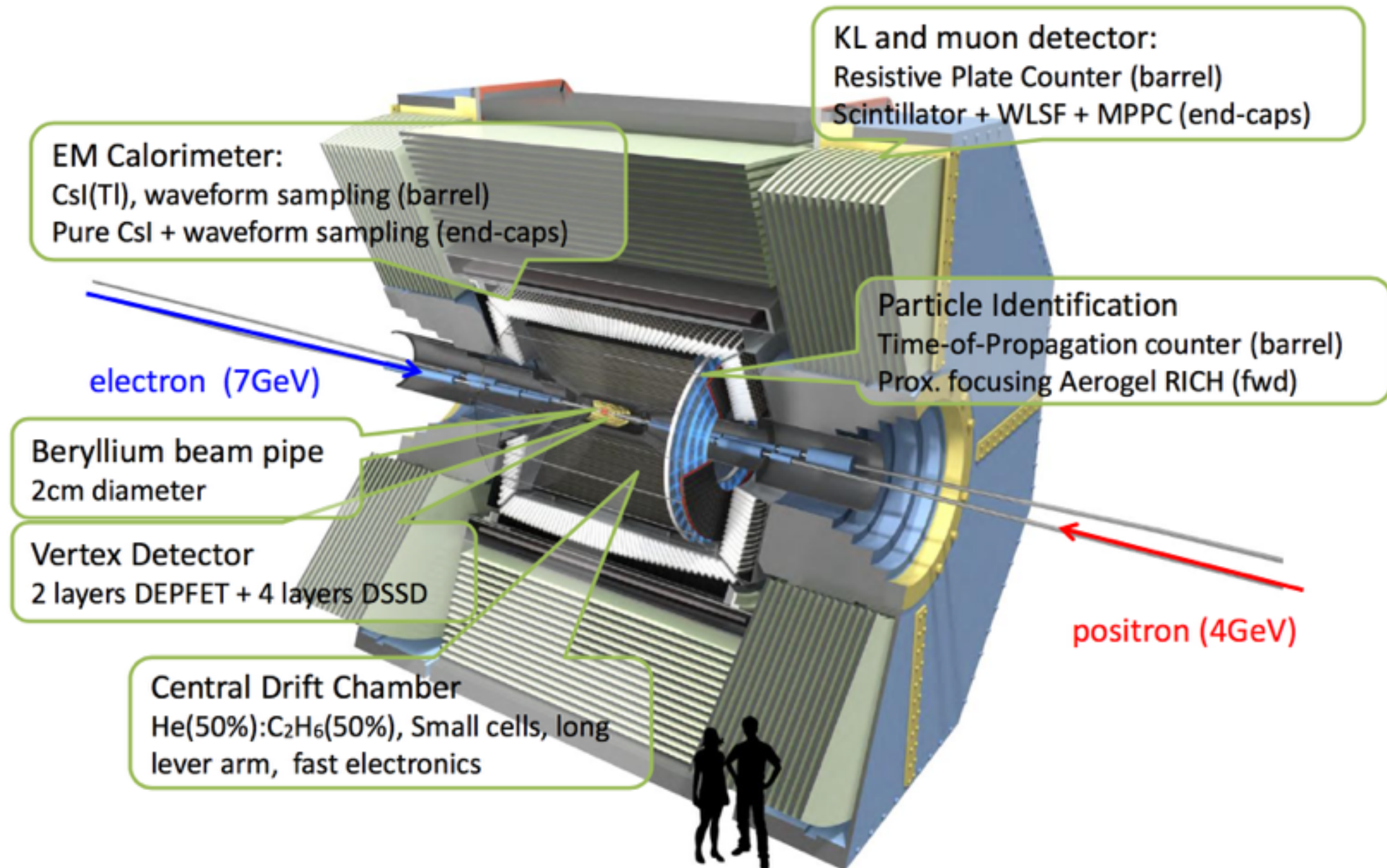
Lorentz factor
 Beam current
 Beam-beam parameter
 Classical electron radius
 Beam size ratio@IP
 1 ~ 2 % (flat beam)
 Vertical beta function@IP
 Lumi. reduction factor (crossing angle) & Tune shift reduction factor (hour glass effect)
 0.8 ~ 1 (short bunch)

- Nano beam:
 - Small current increase (2-3x)
 - Smaller β_y^* (20x) via superconducting focus magnets
- Challenge of beam backgrounds

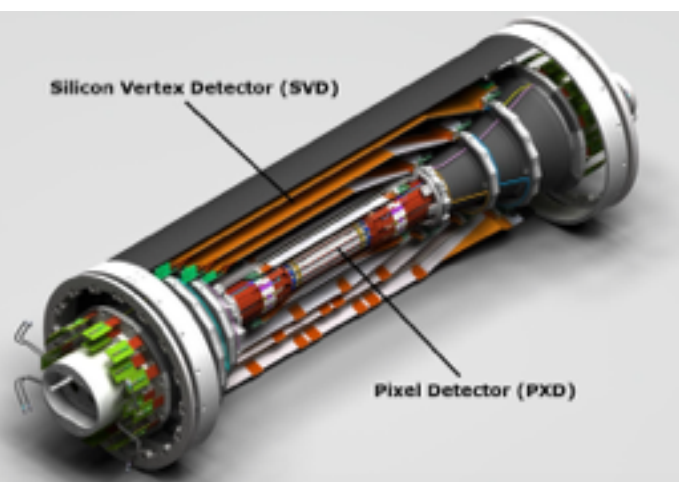


	E (GeV) LER/HER	β_y^* (mm) LER/HER	β_x^* (cm) LER/HER	ϕ (mrad)	I (A) LER/HER	L (cm ⁻² s ⁻¹)
KEKB	3.5/8.0	5.9/5.9	120/120	11	1.6/1.2	2.1 x 10 ³⁴
SuperKEKB	4.0/7.0	0.27/0.30	3.2/2.5	41.5	3.6/2.6	80 x 10 ³⁴

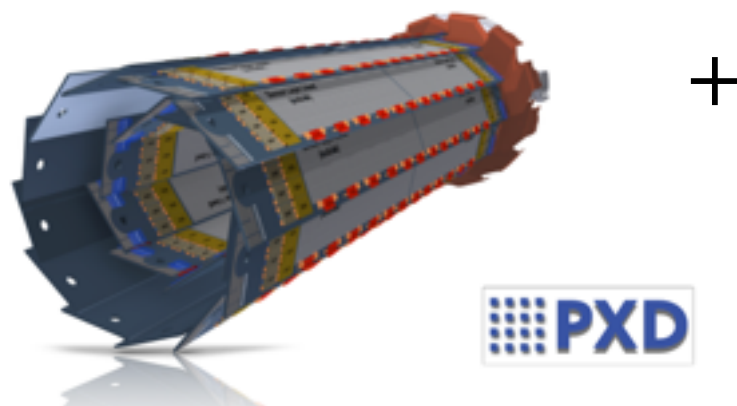
Belle II Detector



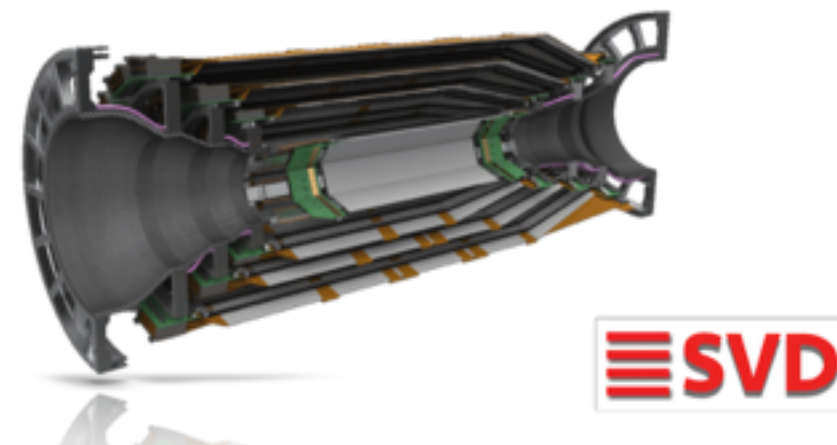
Belle II: VXD



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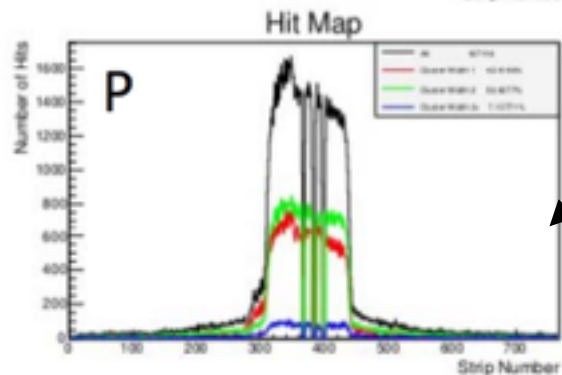
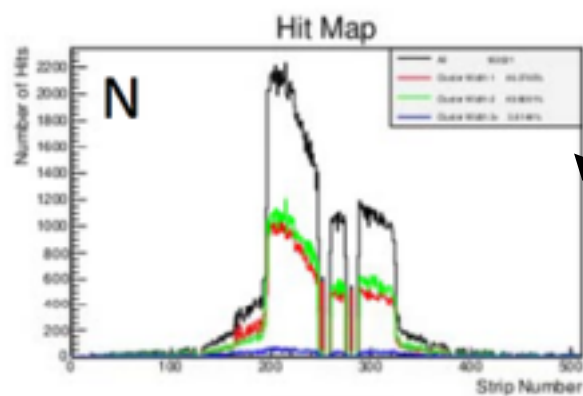


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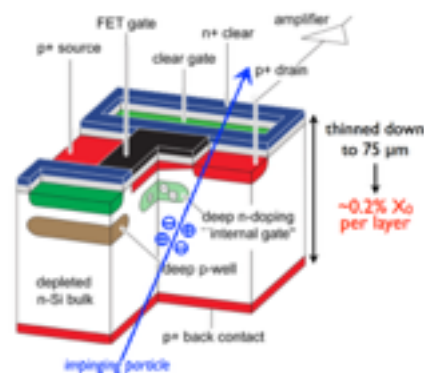
VXD

Hit map example (CE)



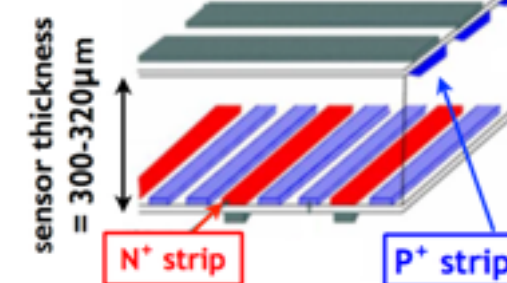
SVD Beam Test Results

DEPLETED p-channel Field Effect Transistor ~ DEPFET



PXD: 2 layers of DEPFET

Double Sided Strip Detectors DSSD

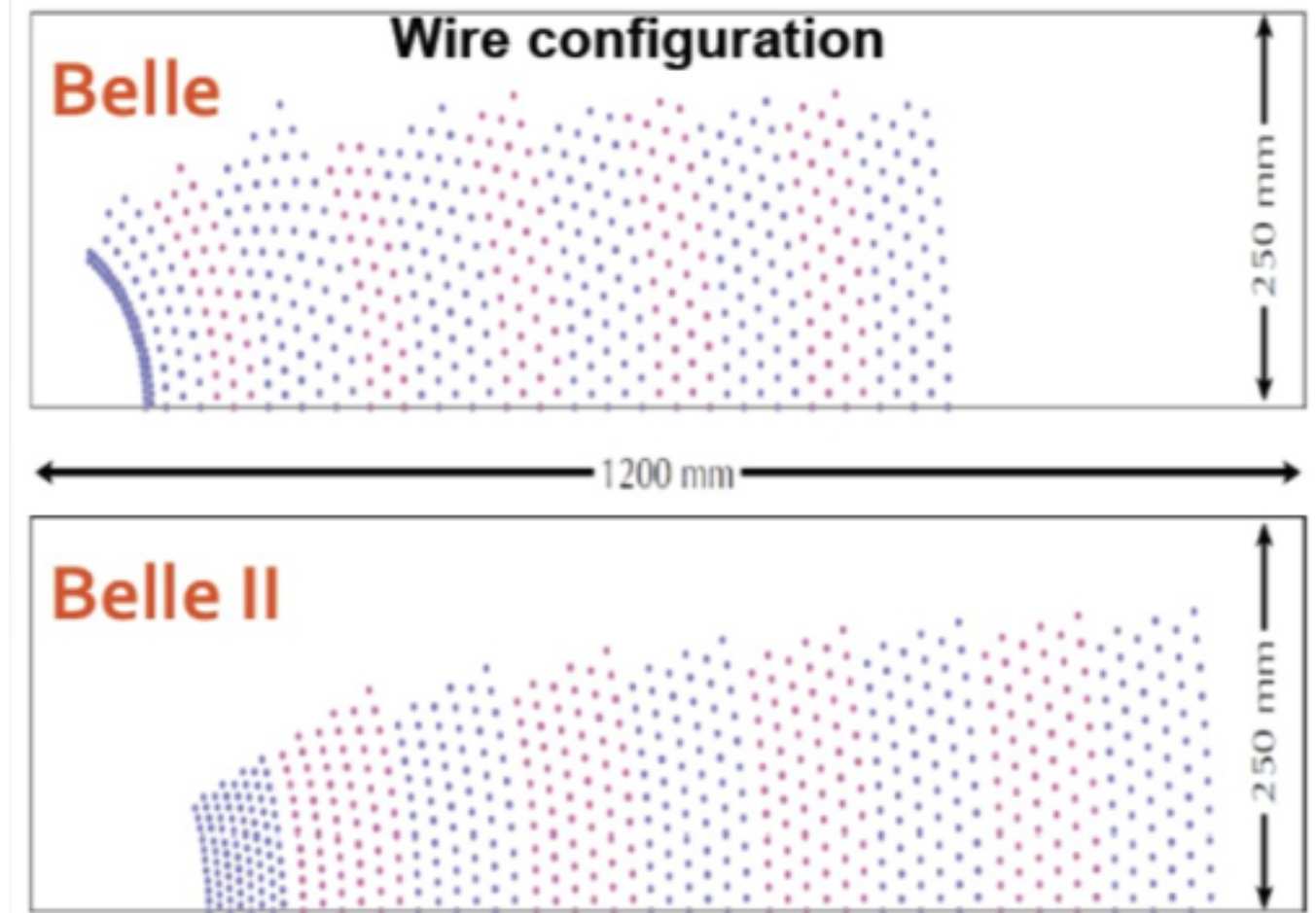
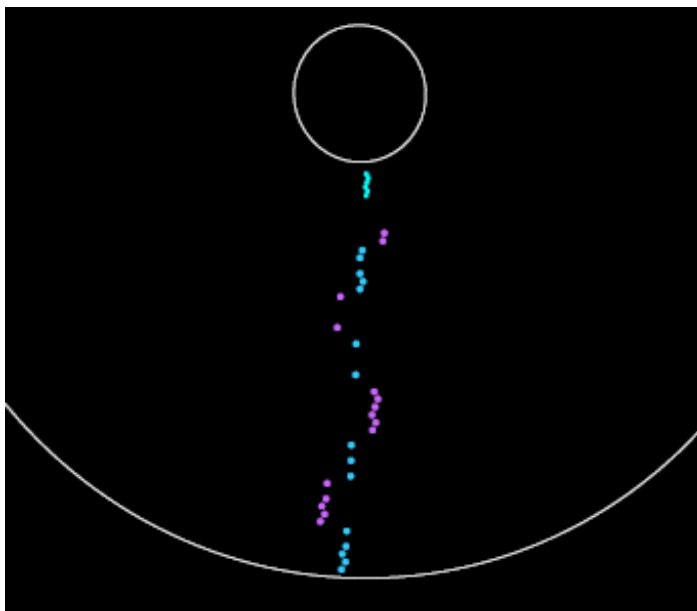


SVD: 4 layers of DSSD

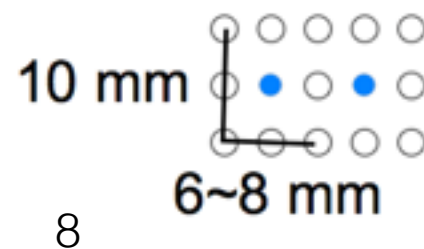


Belle II: CDC

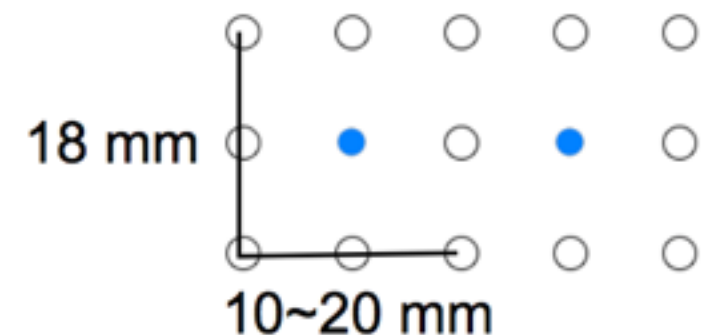
- Belle II CDC will be larger than Belle CDC with smaller cells
- Improved p and dE/dx resolution
- Stringing completed in January 2014 with 51456 wires
- Commissioning with cosmic rays



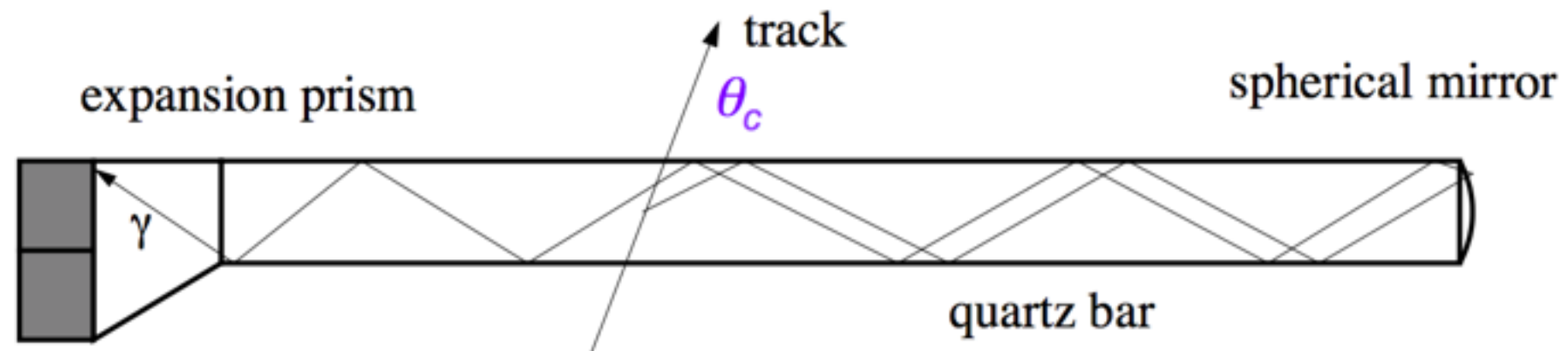
Small cell



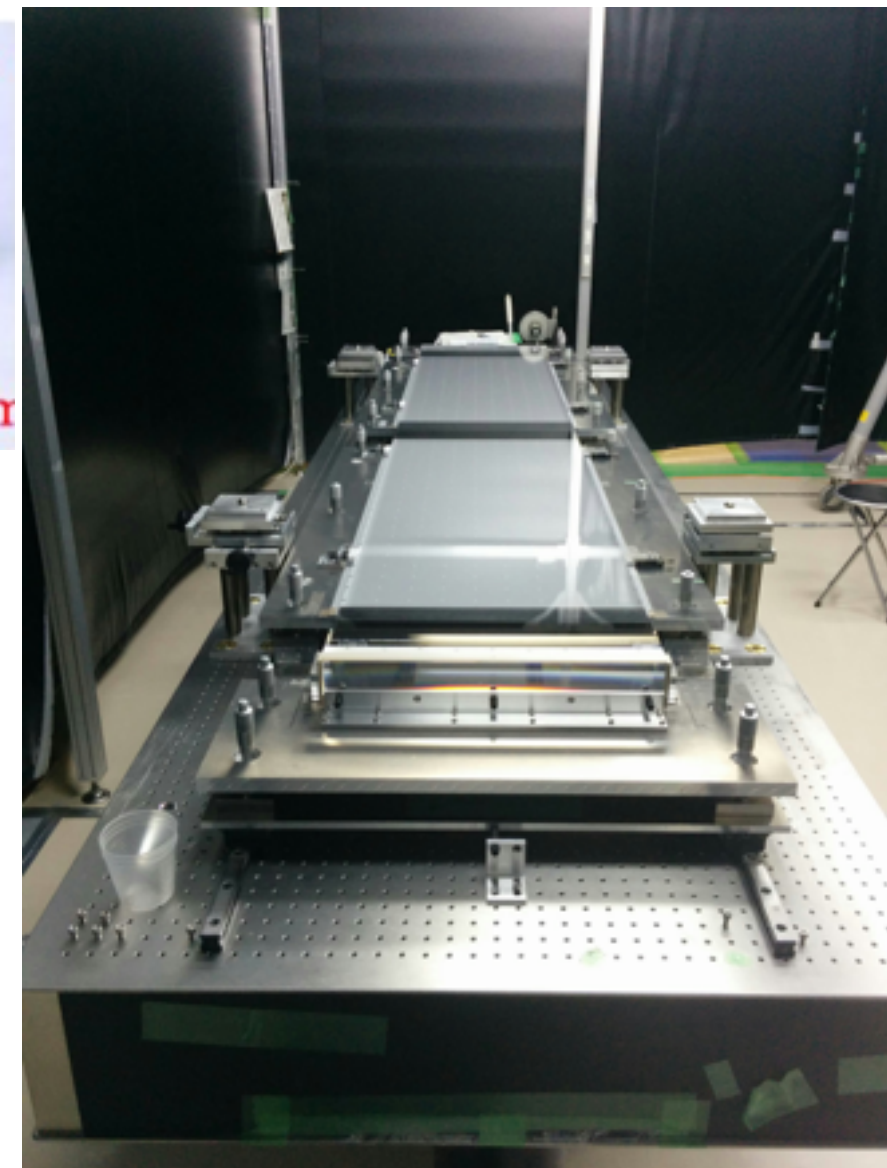
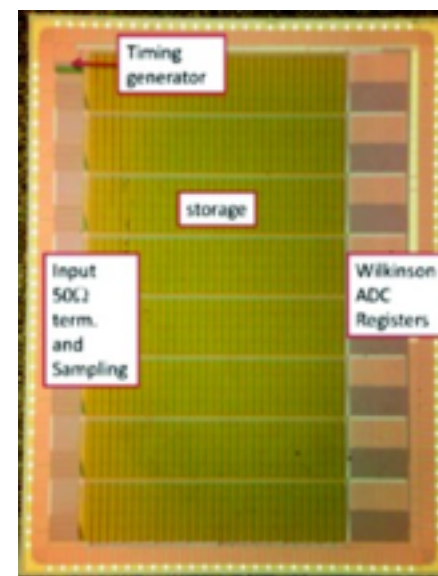
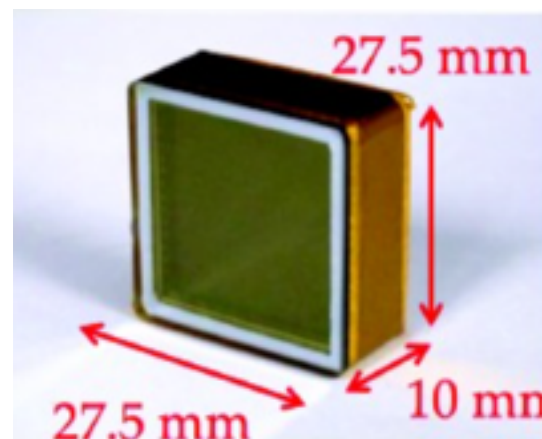
Normal cell



Belle II: TOP

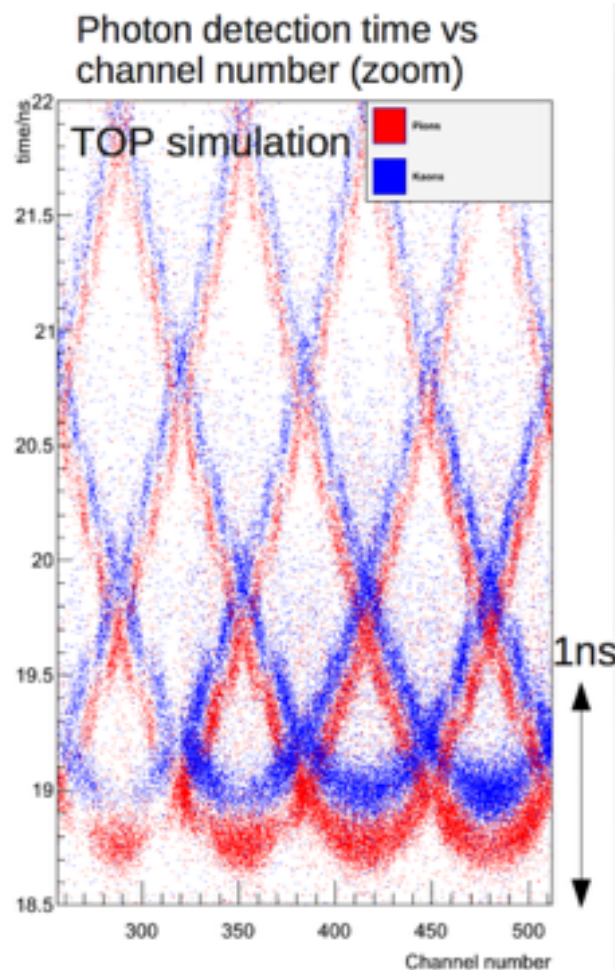


- The imaging Time of Propagation sub-detector (TOP or iTOP) will be used for particle identification in the barrel region of Belle II
- Each TOP module consists of two quartz bars, one mirror, one prism, and an array of photo-detectors to collect Cerenkov photons from charged tracks
- To distinguish between kaons and pions, the photo-detectors should have excellent position and timing resolution
- This is achieved by using MCP-PMTs and new waveform sampling electronics

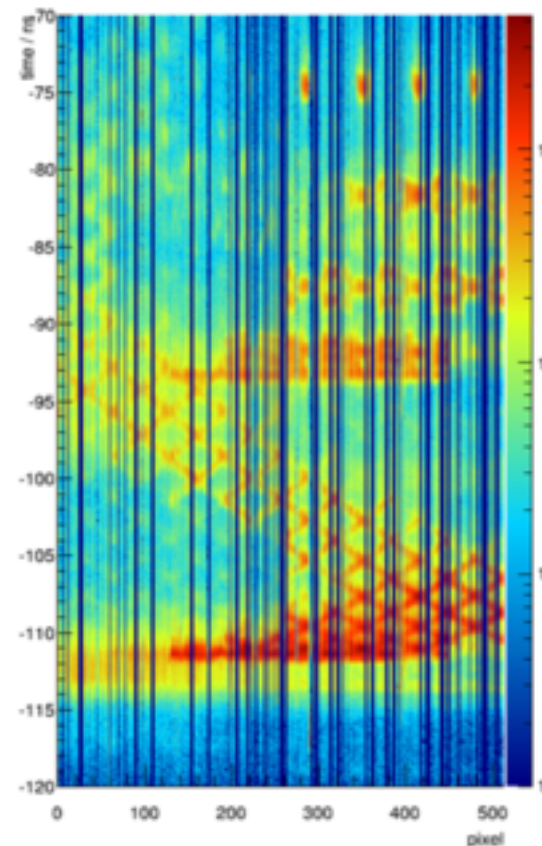


Belle II: TOP

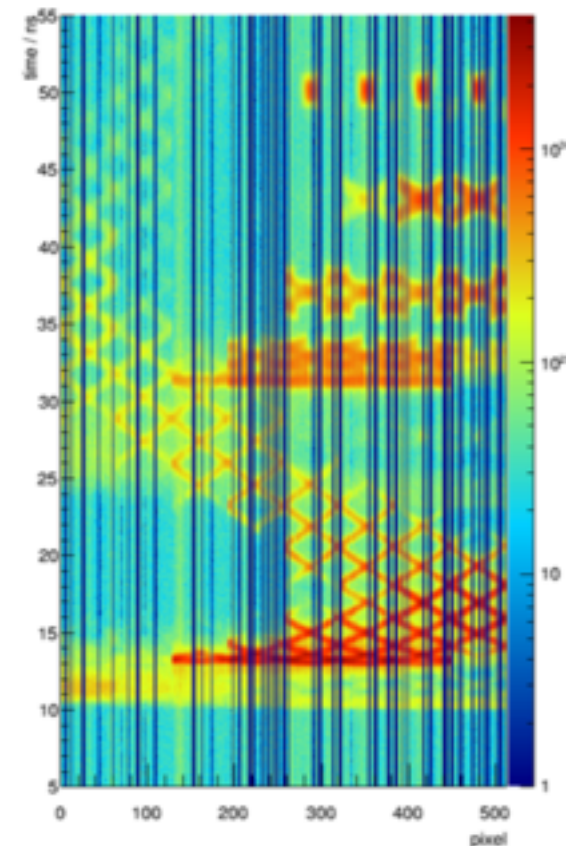
Beam Test at LEPS (June 2013)



Data ring image for $\cos\theta = 0.00$



Simulated ring image for $\cos\theta = 0.00$

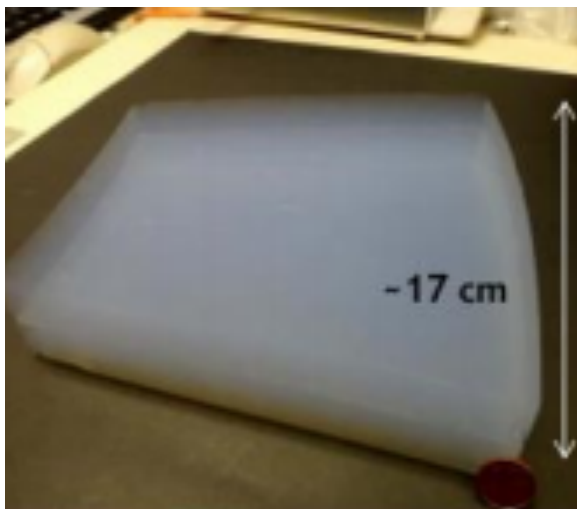


- TOP modules have been tested at beam test at SPring-8 LEPS in 2013, and good agreement between data and MC simulation has been obtained, with timing requirement $\sim O(100\text{ps})$
- 9 out of 16 modules have been assembled (the optical and mechanical parts). The assembly of the 10th module is on going in KEK
- Commissioning with cosmic rays is under way

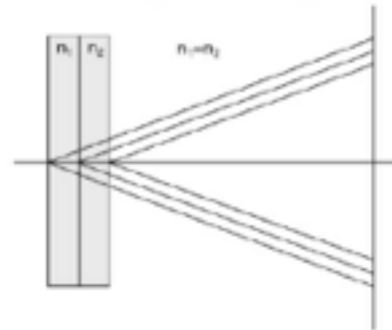
Belle II: ARICH

- Aerogel Ring Imaging Cerenkov (ARICH) detector will be used for particle identification in the forward end cap
- 420 Hybrid Avalanche Photo Detectors (HAPD), each with 144 channels
- Two layers of aerogel lead to better photon yield, while not affecting resolution

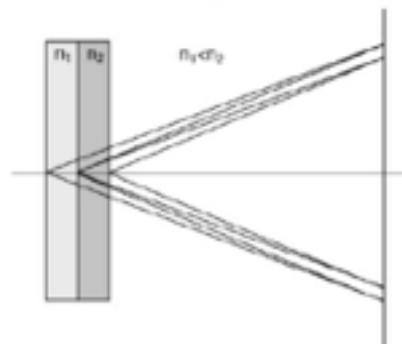
Aerogel



4cm aerogel single index

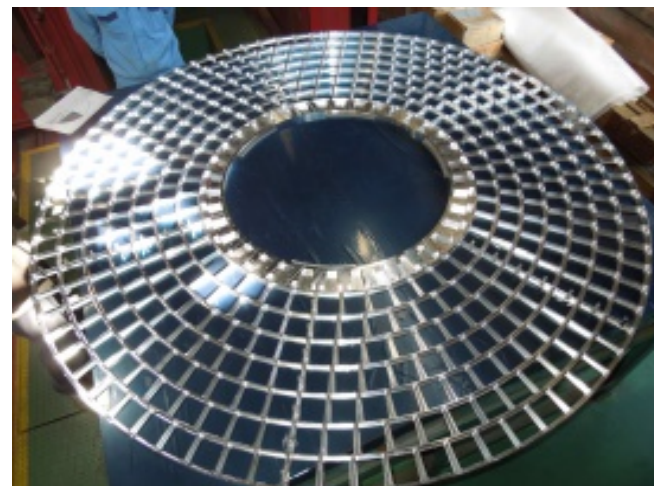


2+2cm aerogel

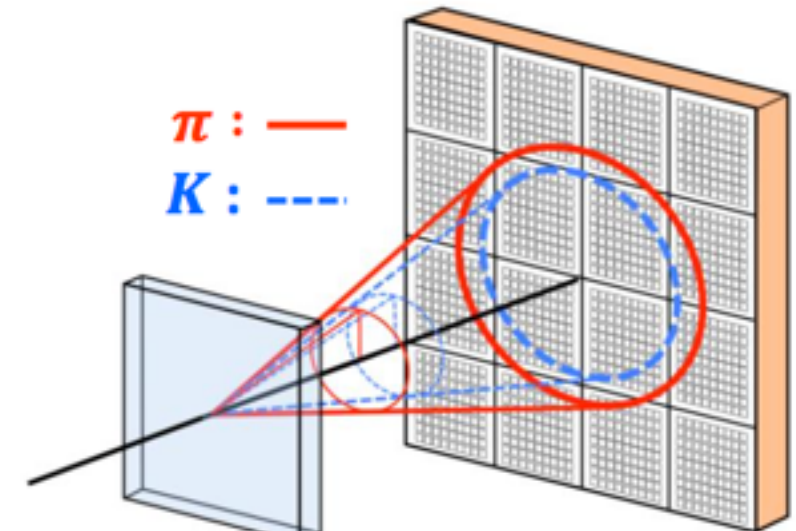


→NIM A548 (2005) 383

Structure



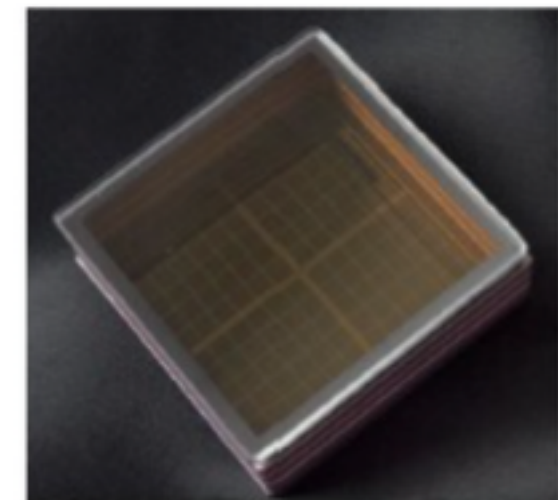
Hamamatsu HAPD



Aerogel radiator

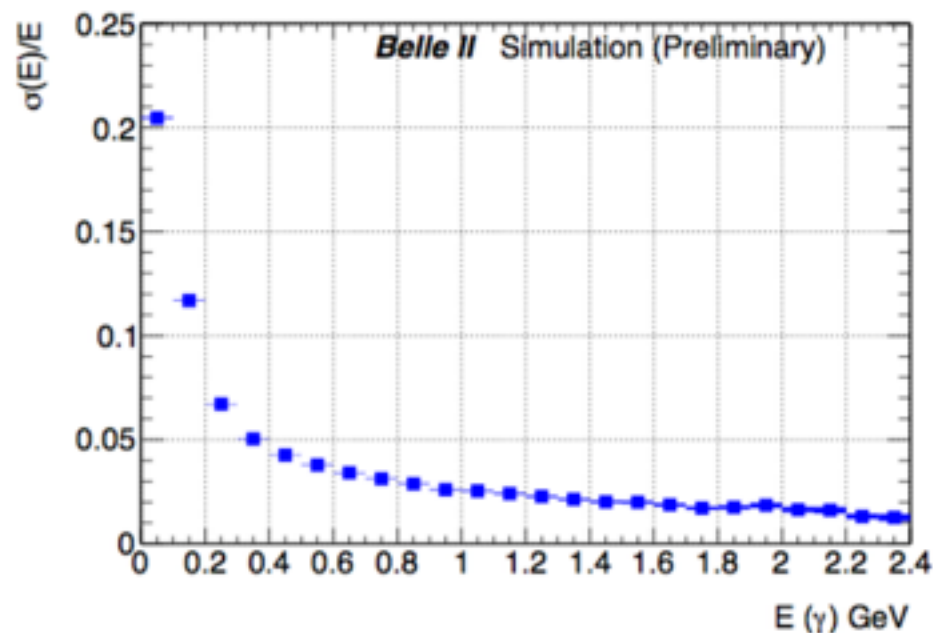
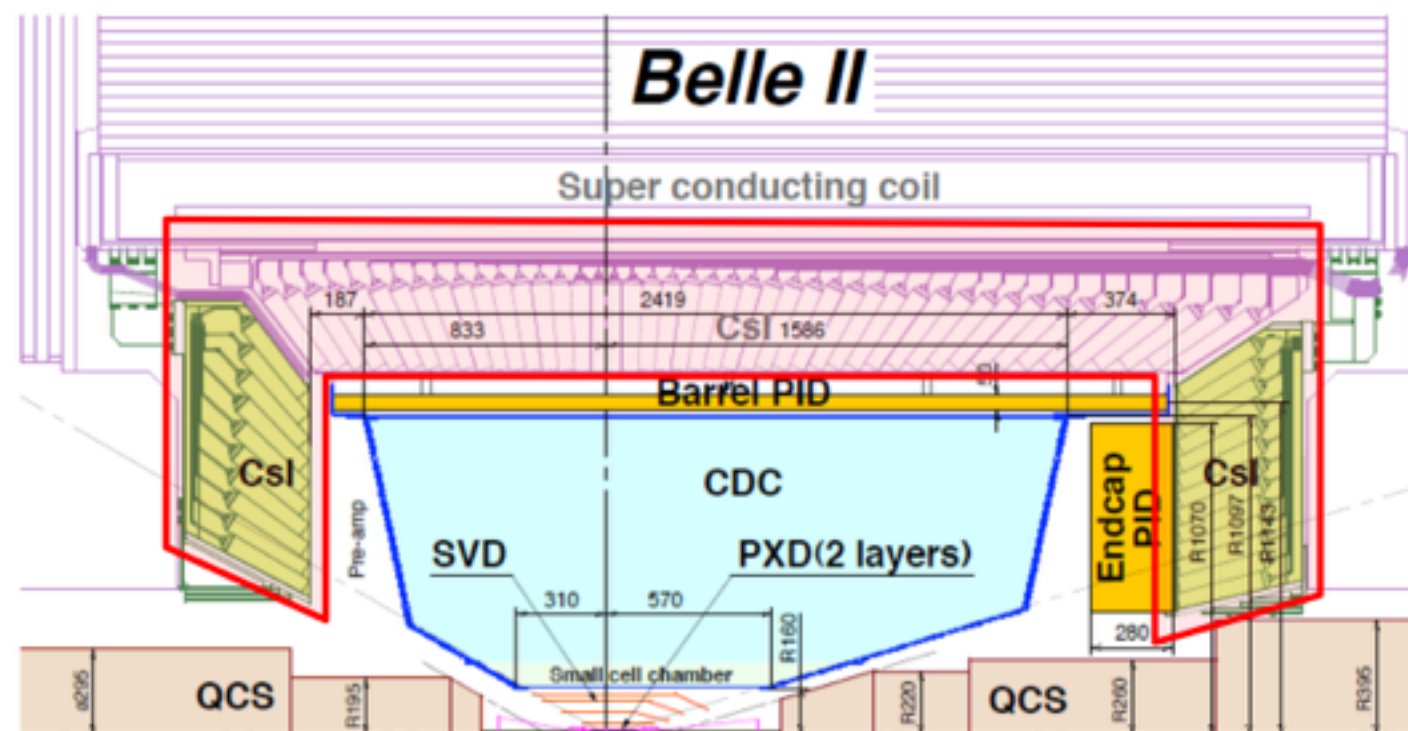
End-cap PID: Aerogel RICH (ARICH)

HAPD



Belle II: ECL

- Upgrades for high backgrounds:
 - Barrel: CsI(Tl) crystals reused, new electronics for waveform sampling
 - Endcaps: old crystals refurbished, bias filter is modified
- Cosmic ray test is on going



Expected Performance

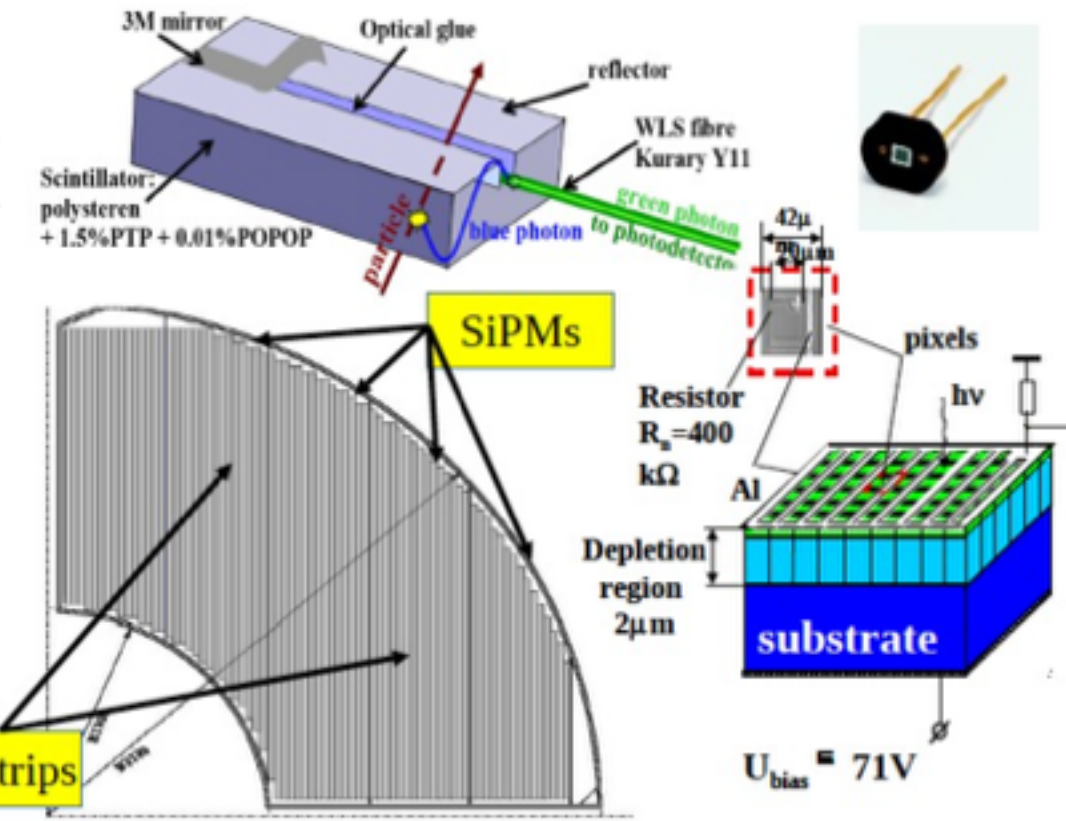


Belle II: KLM

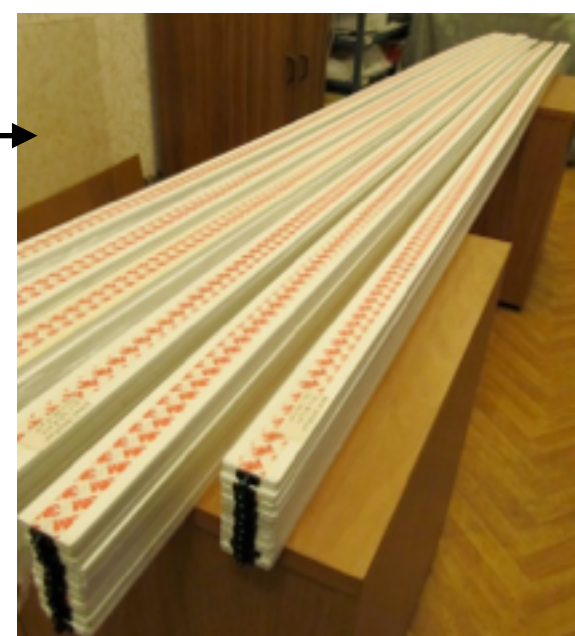
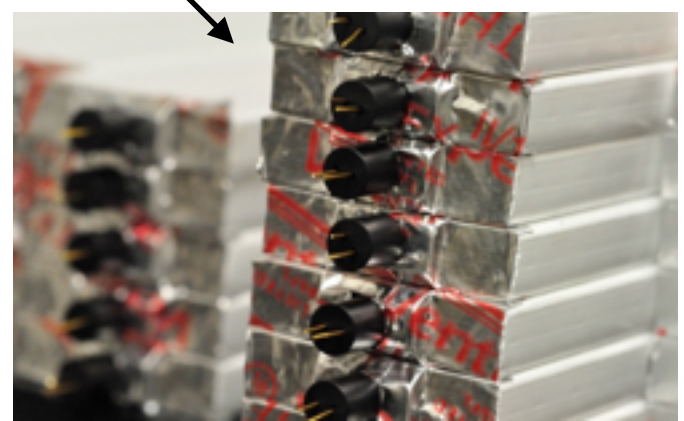
- Endcaps and parts of the barrel KLM RPCs of Belle will be replaced with scintillators due to increased backgrounds expected in Belle II
- Barrel KLM was the first sub-detector to be installed in Belle II

TDR efficiencies for RPC

Layer	Endcap forward	
	KEKB	SuperKEKB
0	0.91	0.0
1	0.93	0.0
2	0.94	0.0
3	0.94	0.0
4	0.94	0.0
5	0.92	0.0
6	0.93	0.0
7	0.92	0.0
8	0.92	0.0
9	0.90	0.0
10	0.87	0.0
11	0.82	0.0
12	0.78	0.0
13	0.77	0.0
14	N/A	N/A

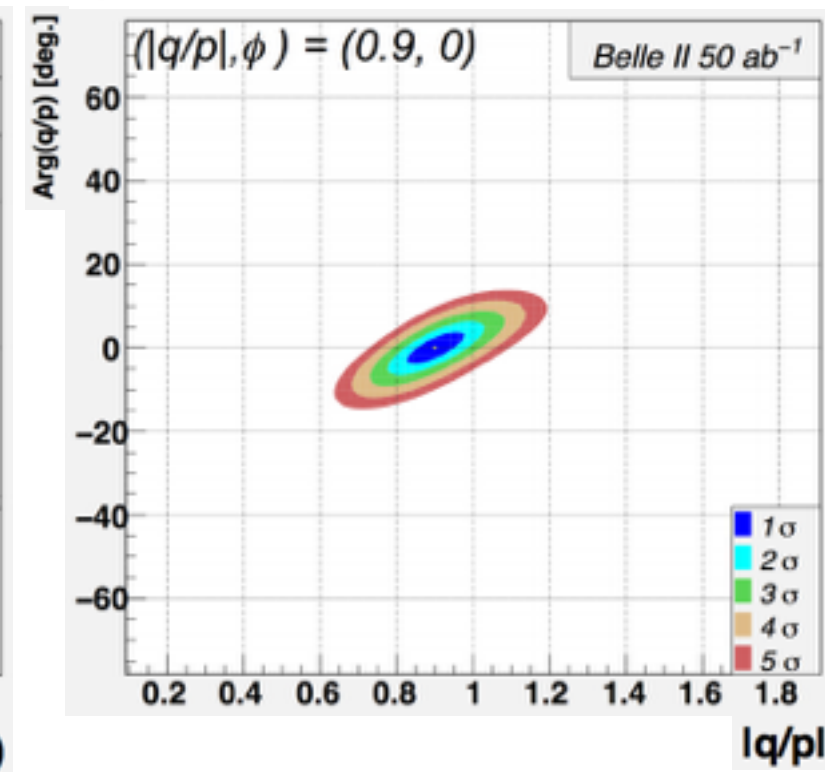
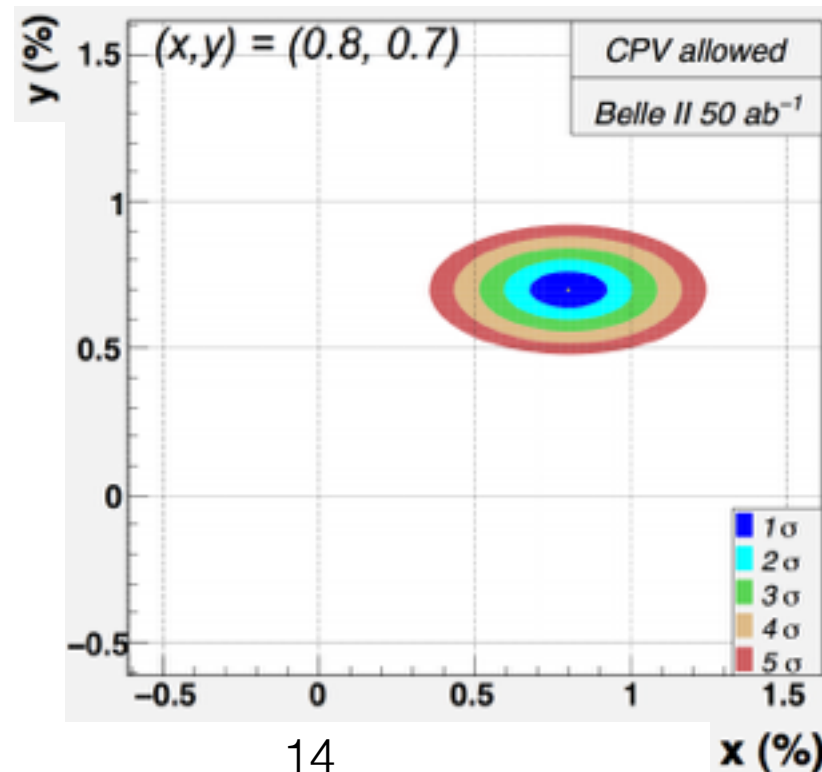
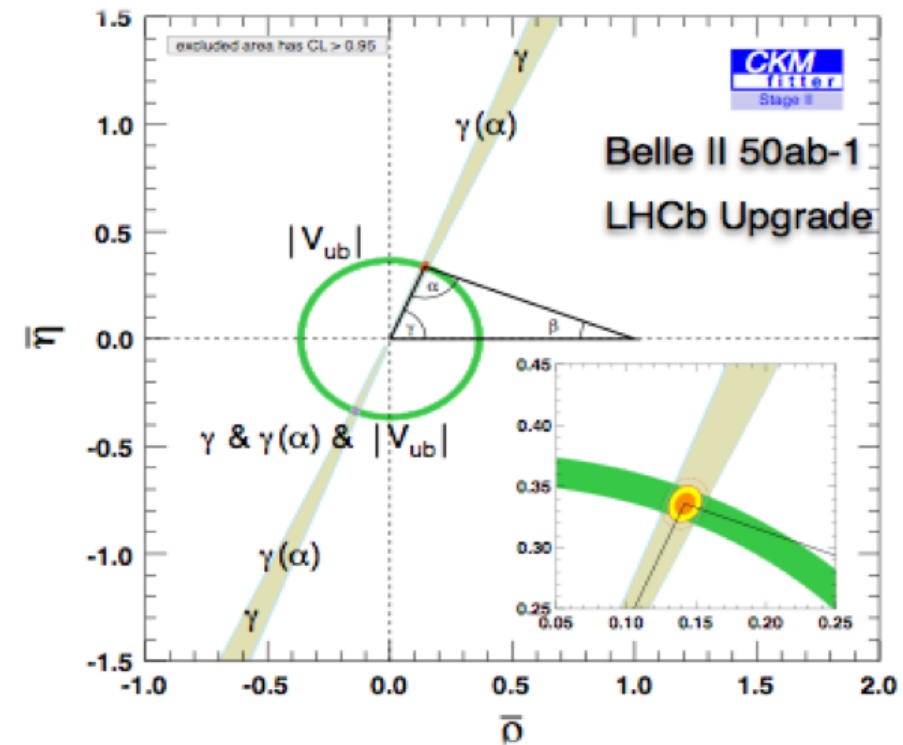


Scintillators for barrel and endcap



Physics Opportunities

- Potential signals for new physics:
 - Flavor changing neutral currents
 - Probing charged Higgs
 - New sources of CPV
 - Lepton Flavour Violation decays
 - Dark sectors
- Belle II physics programme will be complementary with LHCb.



Summary of Belle II Physics

Observables	Belle	Belle II		\mathcal{L}_s [ab ⁻¹]
	(2014)	5 ab ⁻¹	50 ab ⁻¹	
$\sin 2\beta$	$0.667 \pm 0.023 \pm 0.012$	± 0.012	± 0.008	6
α		$\pm 2^\circ$	$\pm 1^\circ$	
γ	$\pm 14^\circ$	$\pm 6^\circ$	$\pm 1.5^\circ$	
$S(B \rightarrow \phi K^0)$	$0.90^{+0.09}_{-0.19}$	± 0.053	± 0.018	>50
$S(B \rightarrow \eta' K^0)$	$0.68 \pm 0.07 \pm 0.03$	± 0.028	± 0.011	>50
$S(B \rightarrow K_S^0 K_S^0 K_S^0)$	$0.30 \pm 0.32 \pm 0.08$	± 0.100	± 0.033	44
$ V_{cb} $ incl.	$\pm 2.4\%$	$\pm 1.0\%$		< 1
$ V_{cb} $ excl.	$\pm 3.6\%$	$\pm 1.8\%$	$\pm 1.4\%$	< 1
$ V_{ub} $ incl.	$\pm 6.5\%$	$\pm 3.4\%$	$\pm 3.0\%$	2
$ V_{ub} $ excl. (had. tag.)	$\pm 10.8\%$	$\pm 4.7\%$	$\pm 2.4\%$	20
$ V_{ub} $ excl. (untag.)	$\pm 9.4\%$	$\pm 4.2\%$	$\pm 2.2\%$	3
$\mathcal{B}(B \rightarrow \tau\nu)$ [10 ⁻⁶]	96 ± 26	$\pm 10\%$	$\pm 5\%$	46
$\mathcal{B}(B \rightarrow \mu\nu)$ [10 ⁻⁶]	< 1.7	5σ	$\gg 5\sigma$	>50
$R(B \rightarrow D\tau\nu)$	$\pm 16.5\%$	$\pm 5.6\%$	$\pm 3.4\%$	4
$R(B \rightarrow D^*\tau\nu)$	$\pm 9.0\%$	$\pm 3.2\%$	$\pm 2.1\%$	3
$\mathcal{B}(B \rightarrow K^{*+}\nu\bar{\nu})$ [10 ⁻⁶]	< 40		$\pm 30\%$	>50
$\mathcal{B}(B \rightarrow K^+\nu\bar{\nu})$ [10 ⁻⁶]	< 55		$\pm 30\%$	>50
$\mathcal{B}(B \rightarrow X_s\gamma)$ [10 ⁻⁶]	$\pm 13\%$	$\pm 7\%$	$\pm 6\%$	< 1
$A_{CP}(B \rightarrow X_s\gamma)$		± 0.01	± 0.005	8
$S(B \rightarrow K_S^0\pi^0\gamma)$	$-0.10 \pm 0.31 \pm 0.07$	± 0.11	± 0.035	> 50
$S(B \rightarrow \rho\gamma)$	$-0.83 \pm 0.65 \pm 0.18$	± 0.23	± 0.07	> 50
$C_7/C_9 (B \rightarrow X_s\ell\ell)$	$\sim 20\%$	10%	5%	
$\mathcal{B}(B_s \rightarrow \gamma\gamma)$ [10 ⁻⁶]	< 8.7	± 0.3		
$\mathcal{B}(B_s \rightarrow \tau^+\tau^-)$ [10 ⁻³]		< 2		

Observables	Belle	Belle II		\mathcal{L}_s [ab ⁻¹]
	(2014)	5 ab ⁻¹	50 ab ⁻¹	
$\mathcal{B}(D_s \rightarrow \mu\nu)$	$5.31 \times 10^{-3} (1 \pm 0.053 \pm 0.038)$	$\pm 2.9\%$	$\pm (0.9\%-1.3\%)$	> 50
$\mathcal{B}(D_s \rightarrow \tau\nu)$	$5.70 \times 10^{-3} (1 \pm 0.037 \pm 0.054)$	$\pm (3.5\%-4.3\%)$	$\pm (2.3\%-3.6\%)$	3-5
y_{CP} [10 ⁻²]	$1.11 \pm 0.22 \pm 0.11$	$\pm (0.11-0.13)$	$\pm (0.05-0.08)$	5-8
A_Γ [10 ⁻²]	$-0.03 \pm 0.20 \pm 0.08$	± 0.10	$\pm (0.03-0.05)$	7 - 9
$A_{CP}^{K^+K^-}$ [10 ⁻²]	$-0.32 \pm 0.21 \pm 0.09$	± 0.11	± 0.06	15
$A_{CP}^{\pi^+\pi^-}$ [10 ⁻²]	$0.55 \pm 0.36 \pm 0.09$	± 0.17	± 0.06	> 50
$A_{CP}^{\phi\gamma}$ [10 ⁻²]	± 5.6	± 2.5	± 0.8	> 50
$x^{K_S\pi^+\pi^-}$ [10 ⁻²]	$0.56 \pm 0.19 \pm \begin{smallmatrix} 0.07 \\ 0.13 \end{smallmatrix}$	± 0.14	± 0.11	3
$y^{K_S\pi^+\pi^-}$ [10 ⁻²]	$0.30 \pm 0.15 \pm \begin{smallmatrix} 0.05 \\ 0.08 \end{smallmatrix}$	± 0.08	± 0.05	15
$ q/p ^{K_S\pi^+\pi^-}$	$0.90 \pm \begin{smallmatrix} 0.16 \\ 0.15 \end{smallmatrix} \pm \begin{smallmatrix} 0.08 \\ 0.06 \end{smallmatrix}$	± 0.10	± 0.07	5-6
$\phi^{K_S\pi^+\pi^-}$ [°]	$-6 \pm 11 \pm \begin{smallmatrix} 4 \\ 5 \end{smallmatrix}$	± 6	± 4	10
$A_{CP}^{\pi^0\pi^0}$ [10 ⁻²]	$-0.03 \pm 0.64 \pm 0.10$	± 0.29	± 0.09	> 50
$A_{CP}^{K_S^0\pi^0}$ [10 ⁻²]	$-0.10 \pm 0.16 \pm 0.09$	± 0.08	± 0.03	> 50
$Br(D^0 \rightarrow \gamma\gamma)$ [10 ⁻⁶]	< 1.5	$\pm 30\%$	$\pm 25\%$	2
$\tau \rightarrow \mu\gamma$ [10 ⁻⁹]		< 45	< 14.7	< 4.7
$\tau \rightarrow e\gamma$ [10 ⁻⁹]		< 120	< 39	< 12
$\tau \rightarrow \mu\mu\mu$ [10 ⁻⁹]		< 21.0	< 3.0	< 0.3

BELLE2-NOTE-0021

Direct CPV in $D^0 \rightarrow \Phi\gamma, \rho^0\gamma$

- Direct CPV in radiative decays can be enhanced to exceed 1%:

- $D^0 \rightarrow \Phi\gamma$: A_{CP} up to 2%

- $D^0 \rightarrow \rho^0\gamma$: A_{CP} up to 10%

- $D^0 \rightarrow \Phi\gamma$: first observation by Belle with 78 fb^{-1} :

- Yield: $27.6^{+7.4+0.5}_{-6.5-1.0}$

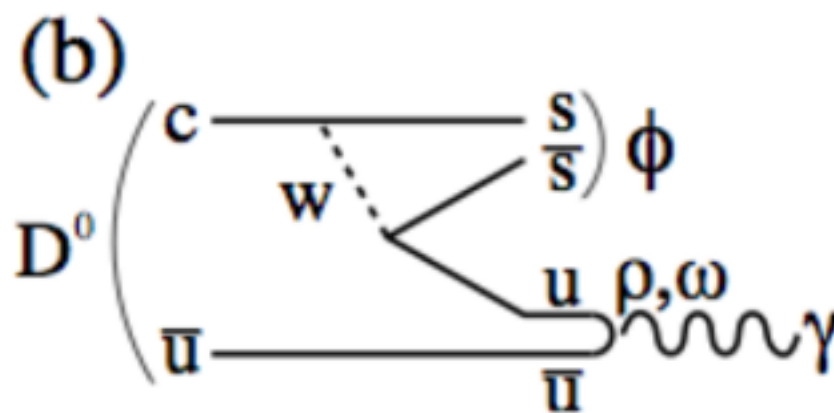
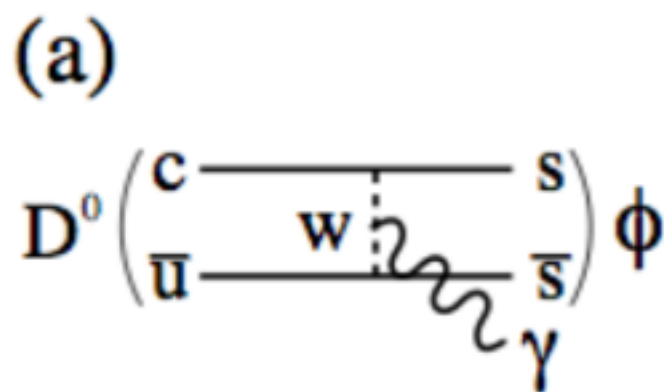
- relative error on yield $\sim 25\%$

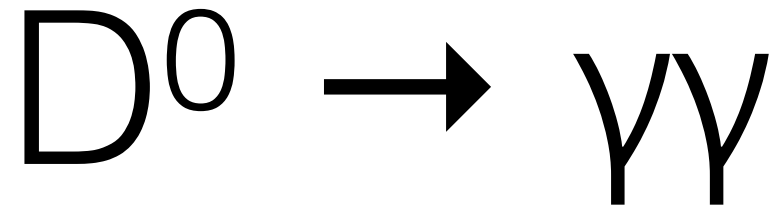
- A_{CP} sensitivity at 50 ab^{-1} : $\sim 1\%$

Refs:

G. Isidori and J. F. Kamenik, PRL
109, 171801 (2012)

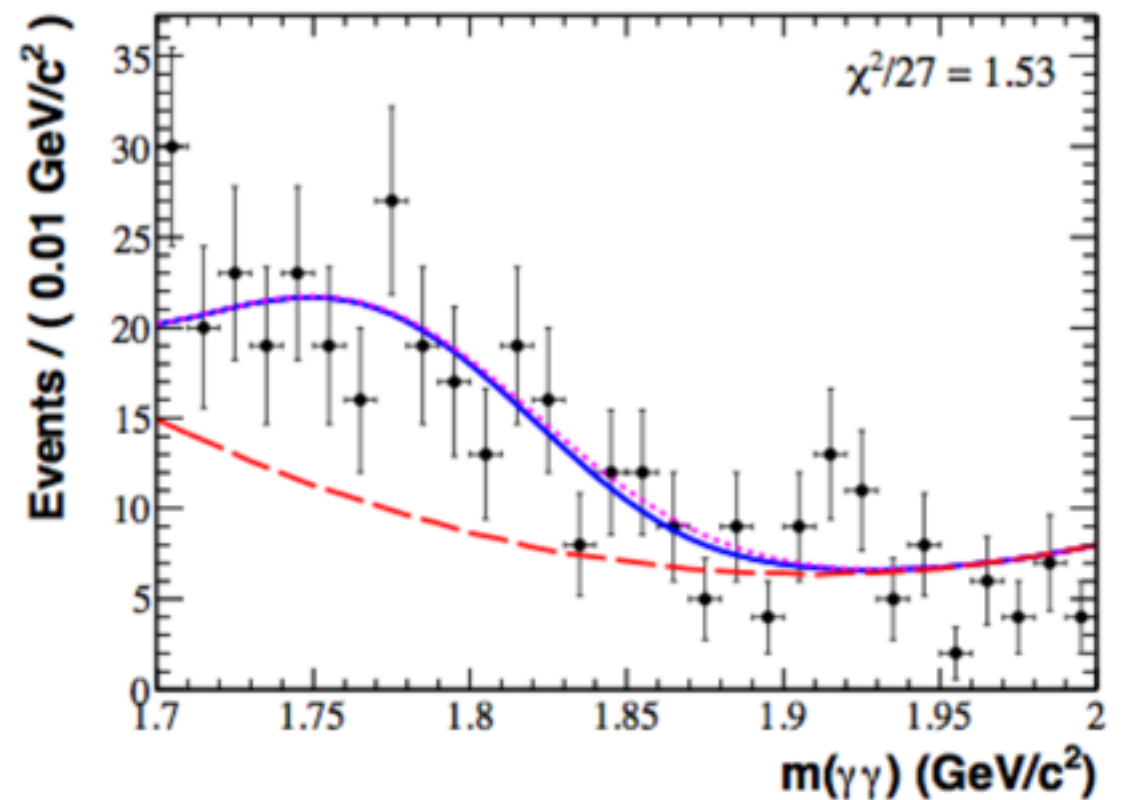
O. Tajima et al. (Belle Collaboration),
PRL 92, 101803 (2004)





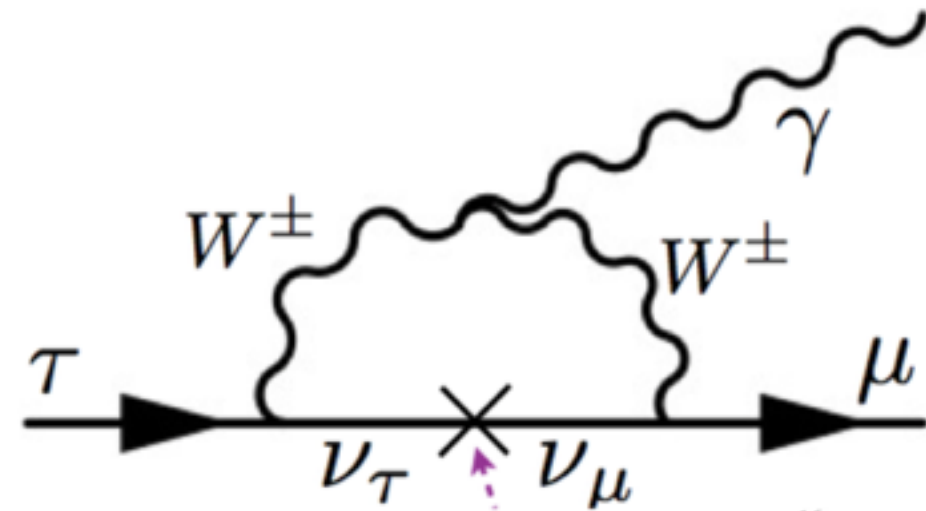
- SM predictions with long distance effects domination: $BR \sim 10^{-8}$
- BaBar with 470 fb^{-1} : $BR < 2.2 \times 10^{-6}$ @ 90% CL
- Belle II with 50 ab^{-1} :
 - depends on how background behaves
 - $UL \sim 2 \times 10^{-8}$, if UL scales with L
 - $UL \sim 2 \times 10^{-7}$, if UL scales with \sqrt{L}

PRD 85 (2012) 091107



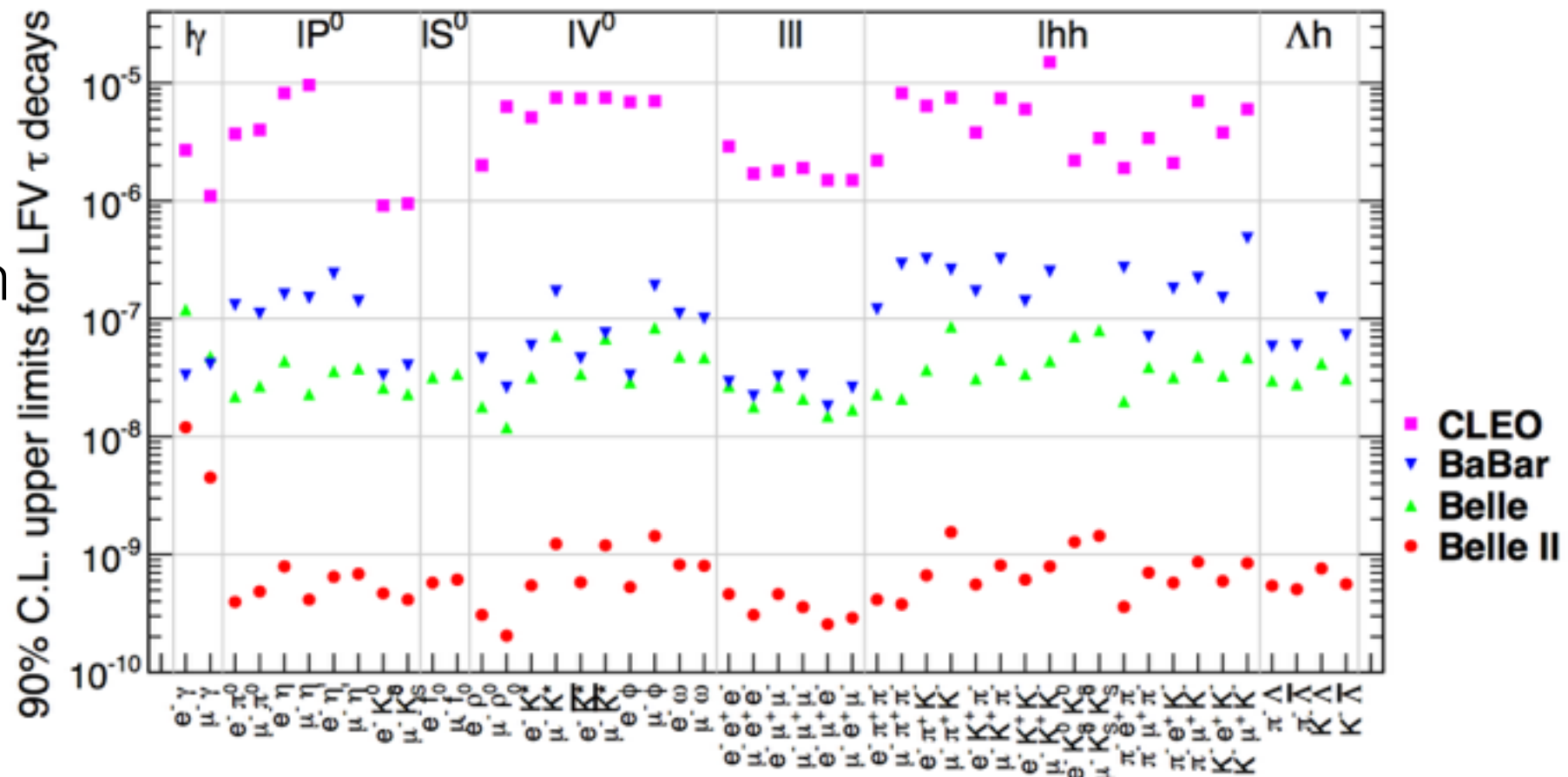
τ Lepton Flavour Violation

SM prediction:
BR(LFV) $\sim 10^{-25}$



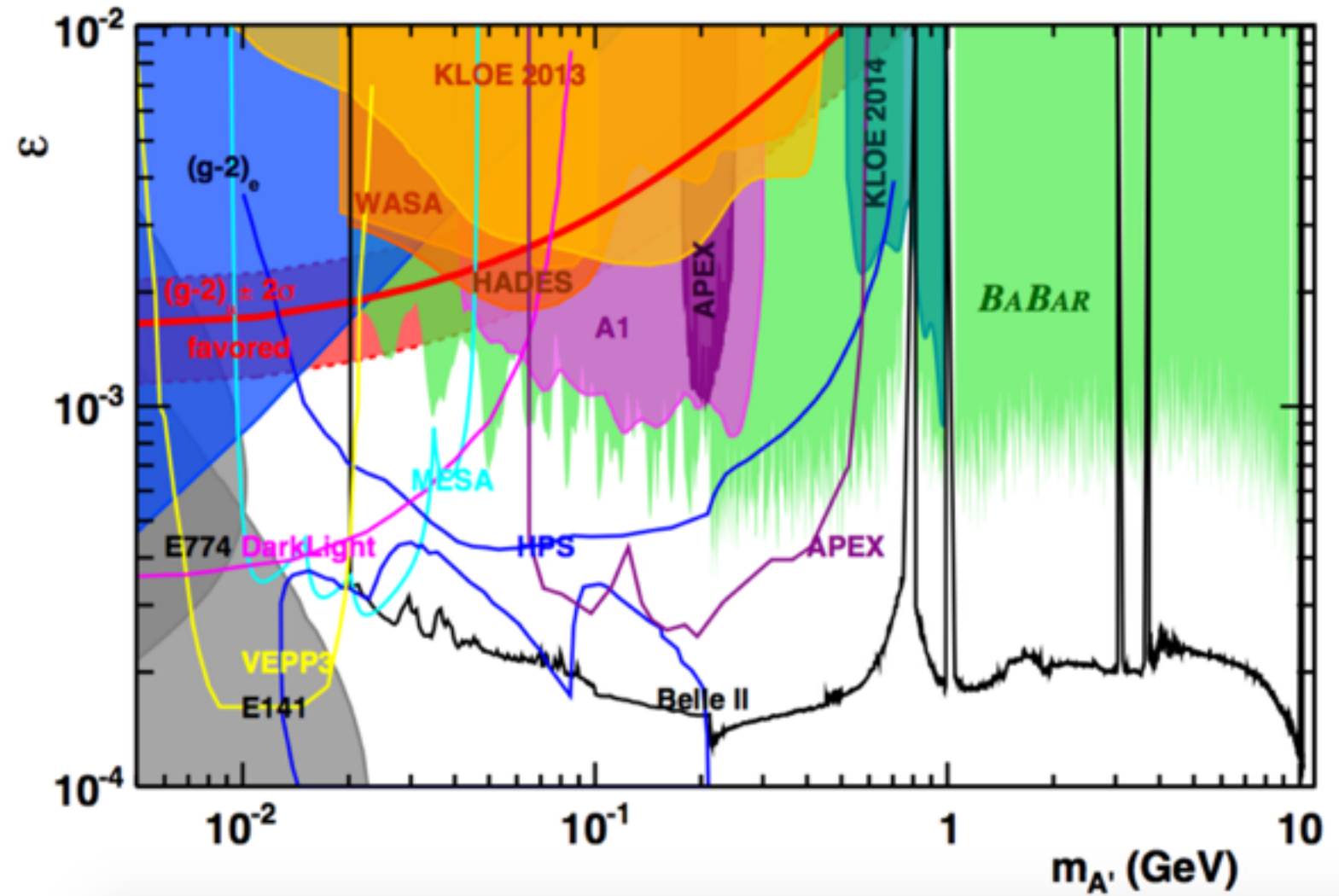
Possible NP in LFV:

- slepton mixing
- H^{++} Zee-Babu models
- Neutral Higgs boson
- Majorana neutrinos
- Seesaw mechanisms



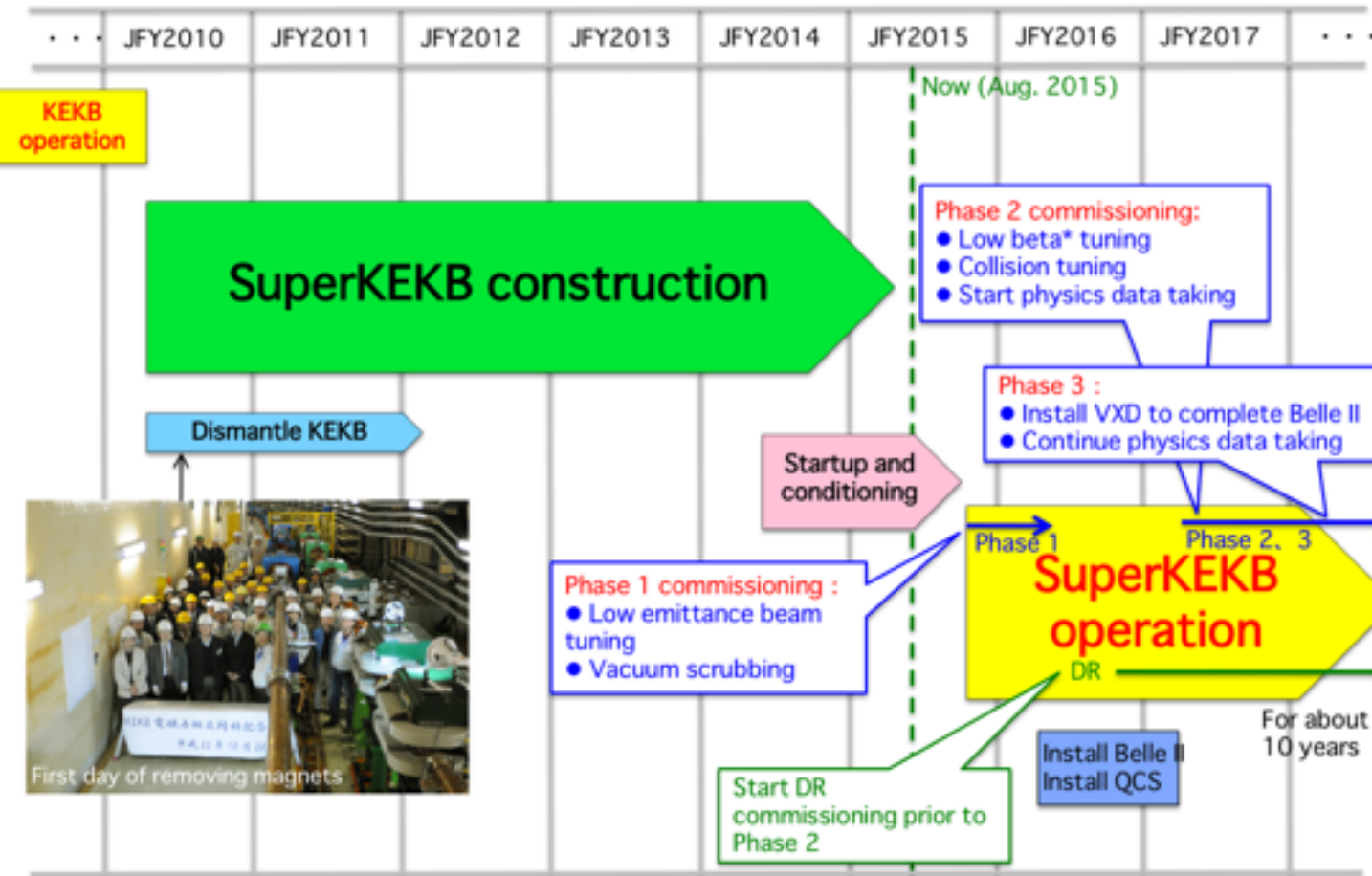
Dark Sector

- Dark matter suggests dark sector.
- Dark photon: A' , to be in MeV \sim GeV mass.
- Probing method:
 - Leptonically decaying dark photons through mixing.
 - Sub-GeV dark matter in invisible decays.



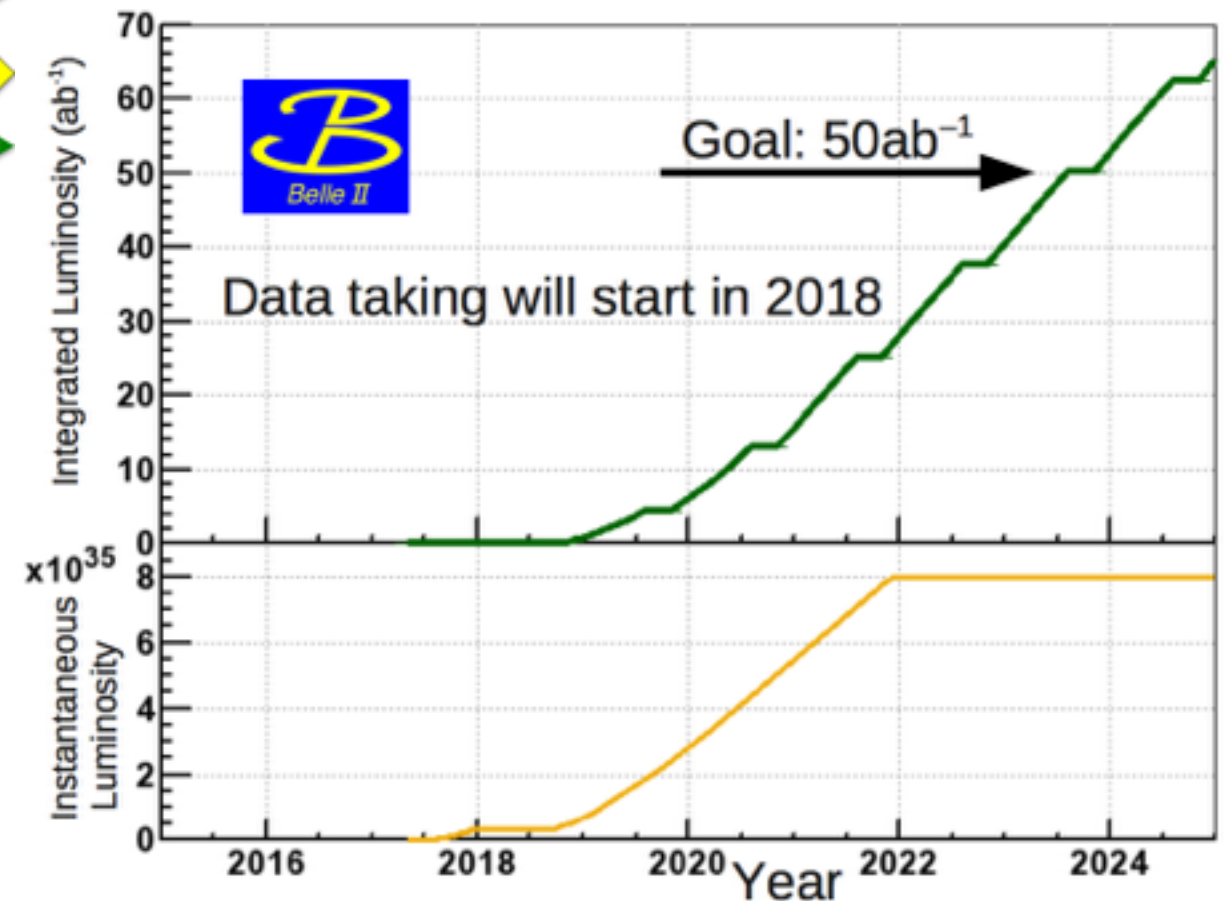
Current and projected limits, radiative production of dark photon, decay to SM particles (C. Hearty, B2TIP 2014)

Schedule

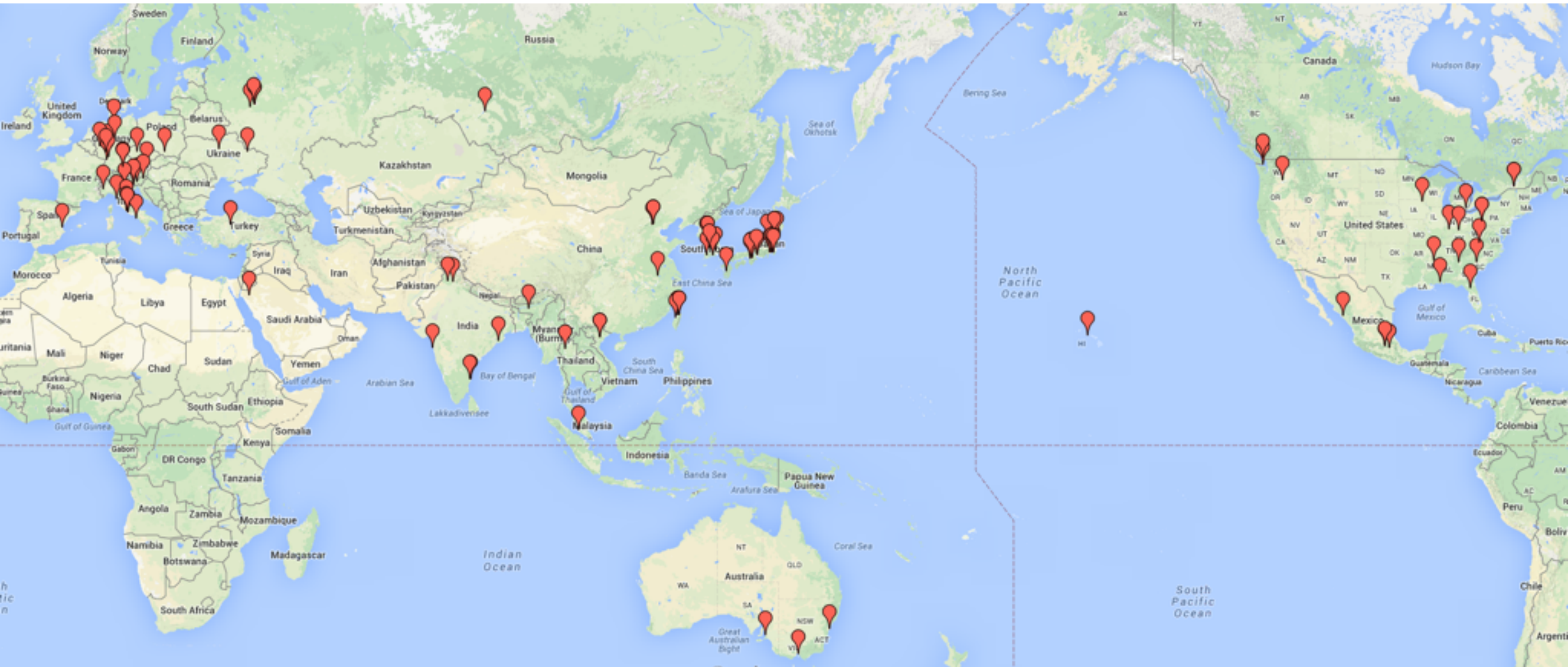


Installation and commissioning plan

Luminosity schedule



Belle II Collaboration



~600 collaborators, 97 institutions, 23 countries

Summary

- B-factories had some rich and successful physics results and many hints of new physics.
- Belle will be upgraded to Belle II to further explore these opportunities.
- Belle II will start operation in 2016 and start taking physics data in 2018.
- The target integrated luminosity for physics data is 50 ab^{-1} , which is much larger than the current data set.