Dark matter signals at future high energy colliders

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DM detections





DM phenomenology



- Particle: DM-SM interaction, DM mass, spin, quantum numbers, symmetry...
- Cosmology: why observed relic density, cold or warm, distribution, relation to baryon, dark energy, structure formation...
- Ways:(1)Use available data to constrain known DM model
 (2)Build self-consistent model to explain the experiment results
 (3)Predict possible signals in the future DM detection experiments.









Search for DM at e⁺e⁻ colliders

Advantage: No large QCD background;
 precise beam energy; polarized beams;
 can precisely measure the mass, spin, and other quantum numbers of DM

***** Search for interactions between the DM and electrons/EW gauge bosons

Disadvantage: Low CM energy;
 difficult to directly detect heavy DM



Gamma-ray line and monophoton

- Gamma-ray line is a critical evidence of DM annihilation/decay
- ~3-4 σ signals reported by some groups
- Not confirmed by the Fermi-LAT collaboration; possible instrumental effect
- Search for corresponding mono-photon signals at future e⁺e⁻ colliders









Effective operators

Dirac Fermion DM

$$\mathcal{O}_F = \frac{1}{\Lambda^3} \bar{\chi} i \gamma_5 \chi F_{\mu\nu} \tilde{F}^{\mu\nu}$$

$$\langle \sigma_{\rm ann} v \rangle_{\chi \bar{\chi} \to 2\gamma} \simeq \frac{4m_{\chi}^4}{\pi \Lambda^6} = 10^{-27} \text{ cm}^3 \text{ s}^{-1} \left(\frac{m_{\chi}}{130 \text{ GeV}}\right)^4 \left(\frac{1272 \text{ GeV}}{\Lambda}\right)^6$$

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Mass scale for ~100 GeV and a detectable cross section is ~TeV Similar for scalar DM

Signature: mono-photon

Backgrounds: mainly ννγ







Kinematical cuts





Sensitivities

- Consider possible e⁺e⁻ colliders with CM energies of 250, 500, 1000 and 3000 GeV
- **Φ** 3 σ reaches for mass scale and annihilation cross section
- Require large luminosities





Beam polarization



- Polarized beams would be available at the future e⁺e⁻ collider
- e.g at the ILC (P_{e-},P_{e+})~(0.8,-0.3)
- Main EW backgrounds are induced by chiral interactions ; can be reduced by polarized beams
- Using polarized beams ~ increasing luminosity by a factor of one order





Results for DM interactions with electron/positron

Consider operators

$$\mathcal{O}_e = \frac{1}{\Lambda^2} \bar{\chi} \Gamma_{\chi} \chi \bar{e} \Gamma_e e \quad \Gamma_{\chi}, \Gamma_e \in \{1, \gamma_5, \gamma^{\mu}, \gamma^{\mu} \gamma_5, \sigma^{\mu\nu}\}$$

photon is emitted from initial state radiation





Mono-Z signals

- DM can interact with both the photon and Z boson
- Consider effective operators

$$\mathcal{O}_{F1} = \frac{1}{\Lambda_1^3} \bar{\chi} \, \chi B_{\mu\nu} B^{\mu\nu} + \frac{1}{\Lambda_2^3} \bar{\chi} \, \chi W^a_{\mu\nu} W^{a\mu\nu}$$

$$\supset \bar{\chi} \, \chi (G_{ZZ} Z_{\mu\nu} Z^{\mu\nu} + G_{AZ} A_{\mu\nu} Z^{\mu\nu})$$

$$\mathcal{O}_{F2} = \frac{1}{\Lambda_1^3} \bar{\chi} \, i\gamma_5 \chi B_{\mu\nu} \tilde{B}^{\mu\nu} + \frac{1}{\Lambda_2^3} \bar{\chi} \, i\gamma_5 \chi W^a_{\mu\nu} \tilde{W}^{a\mu\nu}$$

$$\supset \bar{\chi} \, i\gamma_5 \chi (G_{ZZ} Z_{\mu\nu} \tilde{Z}^{\mu\nu} + G_{AZ} A_{\mu\nu} \tilde{Z}^{\mu\nu})$$

$$\mathcal{O}_{FH} = \frac{1}{\Lambda^3} \bar{\chi} \, \chi (D_\mu H)^\dagger D_\mu H \rightarrow \frac{m_Z^2}{2\Lambda^3} \bar{\chi} \, \chi Z_\mu Z^\mu$$



+ Z boson can also com from initial state radiation

$$\mathcal{O}_{\rm FP} = \frac{1}{\Lambda^2} \bar{\chi} \gamma_5 \chi \bar{e} \gamma_5 e,$$
$$\mathcal{O}_{\rm FA} = \frac{1}{\Lambda^2} \bar{\chi} \gamma^\mu \gamma_5 \chi \bar{e} \gamma_\mu \gamma_5 e$$





Reconstruct Z->jj

- Small QCD background at e⁺e⁻ colliders ; reconstruct Z boson via hadronic channel
- Main background: jjvv, jjlv, tt
- Require two jets satisfying 40 < m_{ii} < 95 GeV
- **Reconstruct recoil mass** $m_{\rm rec} = \sqrt{(p_{e^+} + p_{e^-} p_{j_1} p_{j_2})^2}$
- Similar for Z->ll





Reaches

- \bullet 3 σ reaches for interactions between the DM and gauge bosons
- Assume: 1000 fb⁻¹ of data; $\Lambda = \Lambda_1 = \Lambda_2$
- Compare with the limit from Fermi observations of dwarf galaxies





Reaches

- **Φ** 3 σ reaches for interactions between the DM and electrons/positrons
- **•** Compare with the limit from Fermi observations of dwarf galaxies



Search for DM at pp colliders



- Advantage: large energy and luminosity for DM production
- Disadvantage: large backgrounds
- Reaches for effective operator
- Better for light DM and spin-dependent interactions





Breakdown of EFT

Assume a Z' model



- **•** EFT can not describe resonance effect
- Unitarily bound; small width limit;
 t-channel mediator....



e.g. Buchmueller at al, 1308.6799



From EFT to simplified model

(minimal) simplified models including Z'

$$\mathcal{L}_{\rm FV} = \sum_{q} g_{q} Z'_{\mu} \bar{q} \gamma^{\mu} q + g_{\chi} Z'_{\mu} \bar{\chi} \gamma^{\mu} \chi$$
$$\mathcal{L}_{\rm FA} = \sum_{q} g_{q} Z'_{\mu} \bar{q} \gamma^{\mu} \gamma_{5} q + g_{\chi} Z'_{\mu} \bar{\chi} \gamma^{\mu} \gamma_{5} \chi$$
$$\mathcal{L}_{\rm SV} = \sum_{q} g_{q} Z'_{\mu} \bar{q} \gamma^{\mu} q + i g_{\chi} Z'_{\mu} [\chi^{*} \partial^{\mu} \chi - (\partial^{\mu} \chi^{*}) \chi]$$

- Require $\Gamma_{Z'} < m_{Z'}$
- For simplicity, assume g_q = g_χ
 Note that g_q can be limited by di-jet resonance searches.
 In the resonance region, Z' width would affect the DM production rate





Sensitivities of SppC



Collider reaches and relic density







Difficult to directly detect DM signals at e⁺e⁻ colliders

Collider has strong capability to search for the light DM

- EFT is not valid at future hadron colliders; simplified models should be taken into account
- Complementary to direct and indirect detections





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