



Fermions





## **BSM SEARCHES** SUSY \_\_\_\_\_

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IHEP, Beijing, China

Aug. 11-19 iSTEP2015





Everyday life (and particle physics) are described by the Standard Model



希腊神话中的怪物"Uroboros"与格拉肖的"宇宙圈"



希腊神话中的怪物"Uroboros"与格拉肖的"宇宙圈"



### 希腊神话中的怪物"Uroboros"与格拉肖的"宇宙圈"











Higgs boson observed, SM is complete.
SM fits the experimental data very well
big success in EW scale



The Nobel Prize in Physics 2013 François Englert, Peter Higgs



## Unfortunately, there is a problem with the Higgs!



### While has problem in Planck scale:

- Naturalness and "hierarchy" problem
- Unification of gauge coupling
- Dark Matter

### Unfortunately, there is a problem with the Higgs!



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Need a more fundamental theory of which SM is only a low-energy approximation → New Physics

### While has problem in Planck scale:

- Naturalness and "hierarchy" problem
- Unification of gauge coupling
- Dark Matter

## Unfortunately, there is a problem with the Higgs!



## **SUSY can do help ?**



## (TeV-scale) Supersymmetry (SUSY)



## New Physics beyond the SM



→ Higgs 发现后,寻找超对称粒子将是LHC实验下一个最主要的物理目标!

# Outline

- SUSY Introduction
- The LHC and ATLAS
- SUSY search strategy
- Overview of SUSY search results
- Outlook and Summary







# Outline



- Overview of SUSY search results
- Outlook and Summary





### A symmetry which unified fermions (mater) and bosons (forces)

- Each particle has a super-partner
- Number of elementary particles doubled
- Spin differs by ½ between SUSY and SM particles
- Identical gauge numbers and couplings

A more fundamental theory: compatible with SM in EW scale, <sup>15-8</sup>solve most problems in Planck<sup>5</sup>scale

# **SUSY Introduction**



- Solve hierarchy problem without "fine tuning"
  - Fermion and boson loops contribute with different signs to the Higgs radiative corrections
  - Supersymmetric partner contributions to Higgs mass cancel SM contributions





## **SUSY Introduction**



### □ Unification of gauge couplings

- New particle content changes running of couplings
- requires SUSY masses below few TeV



## Provide Dark Matter candidate

天文学家发现宇宙中很 大一部分是我们看不见 的 暗物质(明物质只 占4.6%)

### 'Supersymmetric' particles ?





 Provide perfect dark mater candidate - WIMP(lightest neutralino in R-parity conserving models)
stable
electrically neutron
same density as DM

0.094 < Ω<sub>CDM</sub>h<sup>2</sup> < 0.136 (95% CL)

→ 通过寻找SUSY, 可以为 暗物质寻找提供实验证据!

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# Outline

## SUSY Introduction

The LHC and ATLAS

SUSY search strategy



- Overview of SUSY search results
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HCh-



- 世界最大,能量最高的加速器,进行最前沿的粒子物理研究
- 质心系能量14TeV (Tevatron的7倍),可以发现5TeV以下的较重的新粒子 22

822

• 积分亮度10<sup>34</sup> cm<sup>-2</sup> s<sup>-1</sup> (Tevatron 的100倍),可以发现微小衰变截面的稀有事例

## **Collisions at LHC**



## ATLAS and CMS detector @ LHC ATLAS and CMS: two multi-purpose detectors @LHC

### A Toroidal LHC ApparatuS

- 42m×22m, 7000 ton
- Solenoid + Toroidal magnet (2T)
- Fine granularity liquid Ar/Tile calorimeters

# Large Hadron Collider (LHC):

Proton-Proton synchrotron
World's highest and largest collider

### Compact Muon Spectrometer

ATLAS

- 21m×15m, 125000 ton
- All silicon trackers, 4T solenoid magnet
- PbWO4+Tile calorimeters

## **ATLAS and CMS**



#### Chinese muon chambers installed in the ATLAS detector



# Outline



Outlook and Summary





# Where do we start?

Huge parameter space, but guiding principles



Early analyses dominated by broad and inclusive searches for gluino and squark production

 Increasing luminosity gave access to rare production channels.
Additional motivation from *Natural SUSY* paradigm

 If 1<sup>st</sup> and 2<sup>nd</sup> squark and gluino is too heavy, EWK SUSY production may dominant in LHC

# SUSY searches strategy driven by cross section and luminosity



## How do we start? - SUSY Signature



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## **How do we search for SUSY?**



## **How do we search for SUSY?**

- SUSY search strategy: search for deviation from SM from the tails
- SUSY sensitive variables: Try to establish excess of events in some sensitive kinematic distribution
- SM background: the discovery of new physics can only be claimed when SM backgrounds are understood well or under control
  - SM bgs understood very well ③
  - No hints for new physics  $\otimes$
  - Slightly overshoot in WW cross section, but consistent with NNLO xsec.





## **SUSY Sensitive Variables**



- **E**<sub>T</sub><sup>miss</sup> from escaping LSP, to suppress bg from mismeasured jets and oth. SM BG
- Related to the sparticle mass scale, like effective mass (**M<sub>eff</sub>**)

$$M_{\text{eff}} \equiv \sum_{i=1}^{N_{\text{jets}}} p_{\text{T}}^{\text{jet},i} + \sum_{j=1}^{N_{\text{lep}}} p_{\text{T}}^{\text{lep},j} + E_{\text{T}}^{\text{miss}}$$

**mT, mT2** (stransverse mass): suppress BG with Ws

$$m_{\mathrm{T2}} = \min_{\mathbf{q}_{\mathrm{T}}} \left[ \max \left( m_{\mathrm{T}}(\mathbf{p}_{\mathrm{T}}^{\ell 1}, \mathbf{q}_{\mathrm{T}}), m_{\mathrm{T}}(\mathbf{p}_{\mathrm{T}}^{\ell 2}, \mathbf{p}_{\mathrm{T}}^{\mathrm{miss}} - \mathbf{q}_{\mathrm{T}}) \right) \right]$$

Many others ...

## How do we search for SUSY? -Analysis Procedure

### 1. Be aware of SUSY signature, design signal grid

 Pre-selection: select good objects (e, mu, tau, jet, ...), apply trigger depending on analysis, remove bad events (bad runs, not from pp collisions, in transition region ...)

### 3. SR definition and optimization

- Define signal regions based on decay topologies occurring in generic models
- Set final cut on discriminating variables (e.g. Meff) to optimize sensitivity to reference models with appropriate mass scale
- 4. SM Background estimations (data-driven + MC)
- 5. Compare SM predictions with data
- 6. If no excess, interpret results in different SUSY models

# Wide SUSY Signatures



# Wide SUSY Signatures



## **SUSY models: good sale in market**

### **Simplified Models:**

- Not really a model (Br~100%, most masses fixed at high scales)
- Important tool for signal region optimization & interpretation

### Phenomenological models:

- pMSSM: captures "most" of phenomenologic features of Rparity conserving MSSM
  - 19 free parameters: M1,M2,M3 ; tan β, μ and m<sub>A;</sub> 10 sfermion mass parameters;  $A_t$ ,  $A_b$  and  $A_\tau$
  - Comprehensive and computationally realistic approximation of the MSSM with neutralino LSP

**GGM** (gravitino)

## Complete SUSY models: mSUGRA, GMSB ...
## **One example**

#### JHEP04(2015)116





W → I nu W → qq

squarks or gluinos decaying via charginos and sleptons

Final states: ≥ 1 lepton + multi-jets + large E<sub>T</sub><sup>miss</sup>

IHEP

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# **Reconstructed Objects**

- Photons: no track but energy in el-m (and not in the hadronic) calorimeter
- Electrons: track and energy in el-m (and not in the hadronic) calorimeter
- Muons: track in inner tracker and muon chamber
- Jets: cluster in hadronic calorimeter



# **MET: Missing Transverse Energy**

- At the LHC an unknown proportion of the energy of the colliding protons escapes down the beam-pipe
- Invisible particles (neutrinos, neutralinos?) are created their momentum can be constrained in the plane transverse to the beam direction

 $\boldsymbol{E}_T^{ ext{miss}} = -\sum \boldsymbol{p}_T(i)$ 



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# **Triggering on Physics**



40MHz

~100kHz

~1kHz

Apply trigger depending on analysis
 Only pick up what we are interested events
 Single lepton or di-lepton trigger used here

Final states:

 $\geq$  1 lepton + multi-jets + large  $E_T^{miss}$ 

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# SR definition and optimization

	Single-bin (binned) hard single-lepton				
	3-jet	5-jet	6-jet		
$N_\ell$	1 electron or muon				
$p_{\mathrm{T}}^{\ell} \; [\mathrm{GeV}]$	> 25				
Lepton veto	$p_{\mathrm{T}}^{2^{\mathrm{nd}}\mathrm{lepton}} < 10 \mathrm{GeV}$				
$N_{ m jet}$	$\geq 3$	$\geq 5$	$\geq 6$		
$p_{\mathrm{T}}^{\mathrm{jet}}[\mathrm{GeV}]$	> 80, 80, 30	> 80, 50, 40, 40, 40	> 80, 50, 40, 40, 40, 40		
Jet veto	$(p_{\mathrm{T}}^{\mathrm{5^{th}jet}} < 40 \mathrm{~GeV})$	$(p_{\rm T}^{6^{\rm th} jet} < 40 { m GeV})$	_		
$E_{\rm T}^{\rm miss}   [{\rm GeV}]$	>500(300)	>300	>350(250)		
$m_{\rm T}  [{\rm GeV}]$	> 150	> 200 (150)	> 150		
$E_{ m T}^{ m miss}/m_{ m eff}^{ m excl}$	> 0.3	-	_		
$m_{ m eff}^{ m incl} \; [{ m GeV}]$	> 1400 (800)		> 600		
Binned variable	$(m_{\rm eff}^{\rm incl} \text{ in 4 bins})$		$(E_{\rm T}^{\rm miss} \text{ in } 3 \text{ bins})$		
Bin width	$(200 \text{ GeV}, 4^{\text{th}} \text{ is inclusive})$		$(100 \text{ GeV}, 3^{\text{rd}} \text{ is inclusive})$		



According to signal signature, select interested final states objects: number of jets, leptons requirement

Require lepton and jet pt cut due to triggers or optimization

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$p_{\mathrm{T}}^{\mathrm{jet}}[\mathrm{GeV}]$	> 80, 80, 30	> 80, 50, 40, 40, 40	> 80, 50, 40, 40, 40, 40		
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$E_{\rm T}^{\rm miss}$ [GeV]	>500 (300)	>300	>350(250)		
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# Suppress background using SUSY discriminating variablesThe cuts are from optimization with signal significance

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### 4. SM Background estimations (data-driven + MC)

- **5. Compare SM predictions with data**
- 6. If no excess, interpret results in different SUSY models

# **SM Background Estimation**

# SUSY searches rely primarily on the understanding of the SM BG

#### Standard Model

Top, multijets V, VV, VVV, Higgs & combinations of these

#### Reducible backgrounds

Determined from data Backgrounds and methods depend on analyses

#### Irreducible backgrounds

Dominant sources: normalise MC in data control regions Subdominant sources: MC

#### Validation

Validation regions used to cross check SM predictions with data

Signal regions

blinded

blinded

# **SM Background Estimation**

SUSY searches rely primarily on the understanding of the SM BG



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SUSY searches rely primarily on the understanding of the SM BG

Normalise MC prediction in SRs using dedicated CRs  $\rightarrow$  transfer factor: T



Standard Model

Top, multijets *V*, *VV*, *VVV*, Higgs combinations of these

#### Irreducible backgrounds

Dominant sources: normalise MC in data control regions Subdominant sources: MC

Validation

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Signal regions



E<sup>miss</sup><sub>T</sub> [GeV]



### How do we search for SUSY? -Analysis Procedure

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	Hard single-lepton			
	3-jet		5-jet	
	Single-bin	Binned	Single-bin	Binned
Observed events	9	75	5	16
Fitted background events	$8.5\pm1.4$	$82.5\pm7.2$	$7.3\pm1.7$	$17.7\pm4.0$
tt	$2.2\pm0.5$	$35.0\pm6.2$	$4.8\pm1.6$	$12.3\pm4.1$
Other top quarks	$0.79\pm0.35$	$7.6\pm3.0$	$0.71\pm0.18$	$2.1\pm0.5$
V+jets	$2.5\pm0.4$	$24.4\pm3.6$	$0.80\pm0.28$	$1.8\pm0.6$
Diboson	$2.9 \pm 1.0$	$14.3\pm4.3$	$0.96\pm0.69$	$1.5\pm1.0$
Fake leptons	$0.09^{+0.15}_{-0.09}$	$1.2^{+1.3}_{-1.2}$	$0.00^{+0.01}_{-0.00}$	$0.00^{+0.09}_{-0.00}$
Expected background events before the fit	10.1	104.4	9.5	23.2
$t\bar{t}$	3.1	49.3	6.5	16.5
Other top quarks	0.79	7.6	0.7	2.1
V+jets	3.3	32.0	1.3	3.1
Diboson	2.9	14.3	0.96	1.5
Fake leptons	0.09	1.2	0.00	0.00



JHEP04 (2015)116

No significant excess …



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#### squarks/gluinos via full hadronic decay : OL+2-6 jets + MET

JHEP 09 (2014) 176



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#### squarks/gluinos via stop/sbottom decay: 0-1I + 3 b-jets +MET

#### JHEP 10 (2014) 024



#### squarks/gluinos via long decay cha<u>in:</u> SS2L/3L+jets+MET IHEP



LC LC

ţ

give 95 °

1300

[q

m<sub>a</sub> [GeV]

1400  $m_{\tilde{a}}$  [GeV] 59



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### Inclusive search for squark and gluino production Summary



- m(~q) ~m(~g): m(~g) > 1.7 TeV
- M(~q) >> m(~g): m(~g)>1.35 TeV
- Conditional/indirect limit on LSP: m> 200-300 GeV

- No exclusion for  $M(LSP) \ge 700 \text{ GeV}$ 

Strongest limit:  $m(\sim g) \ge 1400 \text{ GeV}$ 

# **Direct stop/sbottom search**

# Light stop and sbottom is motivated by nature SUSY

strong physics case for third generation squarks.

#### **Generic signatures:**

n\_lepton(n=0,1, ≥2) + multi-jets (≥ 1 b-jets) + large MET



#### **1-lepton stop search: 2/3/4-body decay** to LSP

#### JHEP 11 (2014) 118



## **Direct stop pair production Summary**



 Strong search program on stop, covering range from low to heavy stop mass, various decay modes.
 Exclusion for

 Exclusion for m(~t1) < ~660GeV for massless LSP, exclusion up to m(LSP) ~250 GeV

# **Direct EWK-ino Search**

- Search for electroweak (EWK) SUSY below the TeV scale is motivated by naturalness arguments.
- EWK production has a low crosssection compared to strong production
  - Very challenging searches
  - But leads to multi-lepton signatures with very low SM background.
- If strong production is suppressed, EWK processes could be the dominant SUSY production at the LHC. (EWKino < 1 TeV)</p>



<u>Generic signature</u>: >=1 lepton(s) in the final state arising from the decay of charginos/neutralinos via sleptons/sneutrinos, gauge bosons or Higgs.

# C1N2 production – 2-3L(tau)

JHEP 04 (2014)169 JHEP 05 (2014) 071 JHEP 10 (2014) 096



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# C1C1 production – 2L(tau)

JHEP 05 (2014) 071 JHEP 10 (2014) 096





## ~I~I production – 2L(tau)

JHEP 05 (2014) 071 JHEP 10 (2014) 096



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## **Electroweak production – Wh**

[direct C1N2 production via higgs]



# **ElectroWeak Production Summary**



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## Outline

- SUSY Introduction
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- Overview of SUSY search results on run-1 data

Outlook and Summary



# **Outlook and Summary**

#### ATLAS developed a vast program to search for SUSY

- No significant excess seen so far
- All public results:

https://twiki.cern.ch/twiki/bin/view/AtlasPublic/SupersymmetryPublicResults

- In canonical scenarios, sensitivity is achieved to ~1.2 TeV gluinos, ~700 GeV stops and ~400 GeV for EWKinos
- The reach with SUSY is expected to increase significantly at run2 and run3 <sup>(i)</sup>


#### ATLAS SUSY Searches\* - 95% CL Lower Limits Status: July 2015

	Model	$e, \mu,  au, \gamma$	Jets	$E_{ m T}^{ m miss}$	∫ <i>L dt</i> [fb	Mass limit	$\sqrt{s} = 7 \text{ TeV}$	$\sqrt{s} = 8 \text{ TeV}$	Reference
Inclusive Searches	$ \begin{array}{l} \text{MSUGRA/CMSSM} \\ \tilde{q}\tilde{q}, \tilde{q} \rightarrow \tilde{q}\tilde{\lambda}_{1}^{0} \\ \tilde{q}\tilde{q}, \tilde{q} \rightarrow \tilde{q}\tilde{\lambda}_{1}^{0} (\text{compressed}) \\ \tilde{q}\tilde{q}, \tilde{q} \rightarrow q(\mathcal{U}[\ell_{V}\nu_{V}\nu)\tilde{\chi}_{1}^{0} \\ \tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_{1}^{0} \\ \tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_{1}^{0} \\ \tilde{g}\tilde{g}, \tilde{g} \rightarrow qq\tilde{\chi}_{1}^{+} \rightarrow qqW^{\pm}\tilde{\chi}_{1}^{0} \\ \tilde{g}\tilde{g}, \tilde{g} \rightarrow qq(\mathcal{U}[\ell_{V}\nu_{V}\nu)\tilde{\chi}_{1}^{0} \\ \text{GMSB} (\tilde{\ell} \text{ NLSP}) \\ \text{GGM (bino NLSP)} \\ \text{GGM (higgsino-bino NLSP)} \\ \text{GGM (higgsino NLSP)} \\ \text{GGM (higgsino NLSP)} \\ \text{Gravitino LSP} \end{array} $	$\begin{array}{c} 0{\text{-}3}\ e,\mu/1{\text{-}2}\ \tau \\ 0 \\ \text{mono-jet} \\ 2\ e,\mu\ (\text{off-}Z) \\ 0 \\ 0{\text{-}1}\ e,\mu \\ 2\ e,\mu \\ 1{\text{-}2}\ \tau+0{\text{-}1}\ \ell \\ 2\ \gamma \\ \gamma \\ 2\ e,\mu\ (Z) \\ 0 \end{array}$	2-10 jets/3 2-6 jets 1-3 jets 2-6 jets 2-6 jets 0-3 jets 0-2 jets - 1 <i>b</i> 2 jets 2 jets mono-jet	<ul> <li>b Yes</li> <li>Yes</li> </ul>	20.3 20.3 20.3 20.3 20 20 20 20.3 20.3 2	<ul> <li></li></ul>	1.33 TeV 1.26 TeV 1.32 TeV 1.6 1.29 TeV 1.3 TeV 1.25 TeV	<b>1.8 TeV</b> $m(\tilde{q})=m(\tilde{g})$ $m(\tilde{\chi}_{1}^{0})=0 \text{ GeV}, m(1^{st} \text{ gen.} \tilde{q})=m(2^{nd} \text{ gen.} \tilde{q})$ $m(\tilde{q})^{-m}(\tilde{\chi}_{1}^{0})=10 \text{ GeV}$ $m(\tilde{\chi}_{1}^{0})=0 \text{ GeV}$ $m(\tilde{\chi}_{1}^{0})=0 \text{ GeV}$ $m(\tilde{\chi}_{1}^{0})=0 \text{ GeV}$ $tan\beta > 20$ cr(NLSP)<0.1  mm $m(\tilde{\chi}_{1}^{0})<900 \text{ GeV}, cr(NLSP)<0.1 \text{ mm}, \mu<0$ $m(\tilde{\chi}_{1}^{0})<850 \text{ GeV}, cr(NLSP)<0.1 \text{ mm}, \mu>0$ $m(\tilde{\chi}_{1}^{0})<850 \text{ GeV}$ $m(\tilde{G})>1.8 \times 10^{-4} \text{ eV}, m(\tilde{g})=m(\tilde{q})=1.5 \text{ TeV}$	1507.05525 1405.7875 1507.05525 1503.03290 1405.7875 1507.05525 1501.03555 1407.0603 1507.05493 1507.05493 1507.05493 1503.03290 1502.01518
3 <sup>rd</sup> gen. ẽ med.	$\begin{array}{l} \tilde{g}\tilde{g}, \tilde{g} \rightarrow b \tilde{b} \tilde{\chi}_{1}^{0} \\ \tilde{g}\tilde{g}, \tilde{g} \rightarrow t \tilde{\chi}_{1}^{0} \\ \tilde{g}\tilde{g}, \tilde{g} \rightarrow t \tilde{\chi}_{1}^{0} \\ \tilde{g}\tilde{g}, \tilde{g} \rightarrow b \tilde{\iota} \tilde{\chi}_{1}^{+} \end{array}$	0 0 0-1 <i>e</i> , <i>µ</i> 0-1 <i>e</i> , <i>µ</i>	3 <i>b</i> 7-10 jets 3 <i>b</i> 3 <i>b</i>	Yes Yes Yes Yes	20.1 20.3 20.1 20.1	ğ ğ ğ ğ	1.25 TeV 1.1 TeV 1.34 TeV 1.3 TeV	$m(\tilde{k}_{1}^{0}) < 400 \text{ GeV}$ $m(\tilde{k}_{1}^{0}) < 350 \text{ GeV}$ $m(\tilde{k}_{1}^{0}) < 400 \text{ GeV}$ $m(\tilde{k}_{1}^{0}) < 300 \text{ GeV}$	1407.0600 1308.1841 1407.0600 1407.0600
3 <sup>rd</sup> gen. squarks direct production	$ \begin{array}{c} \tilde{b}_{1}\tilde{b}_{1}, \tilde{b}_{1} \rightarrow b\tilde{\chi}_{1}^{0} \\ \tilde{b}_{1}\tilde{b}_{1}, \tilde{b}_{1} \rightarrow \tilde{\chi}_{1}^{\tilde{x}_{1}} \\ \tilde{t}_{1}\tilde{t}_{1}, \tilde{t}_{1} \rightarrow b\tilde{\chi}_{1}^{\tilde{x}_{1}} \\ \tilde{t}_{1}\tilde{t}_{1}, \tilde{t}_{1} \rightarrow b\tilde{\chi}_{1}^{0} \\ \tilde{t}_{1}\tilde{t}_{1}, \tilde{t}_{1} \rightarrow \tilde{\chi}_{1}^{0} \\ \tilde{t}_{1}\tilde{t}_{1}, \tilde{t}_{1} \rightarrow \tilde{\chi}_{1}^{0} \\ \tilde{t}_{1}\tilde{t}_{1}, \tilde{t}_{1} \rightarrow \tilde{\chi}_{1}^{0} \\ \tilde{t}_{2}\tilde{t}_{2}, \tilde{t}_{2} \rightarrow \tilde{t}_{1} + Z \end{array} $	0 2 e, µ (SS) 1-2 e, µ 0-2 e, µ (C) 0 m 2 e, µ (Z) 3 e, µ (Z)	2 <i>b</i> 0-3 <i>b</i> 1-2 <i>b</i> 0-2 jets/1-2 nono-jet/ <i>c</i> -1 1 <i>b</i> 1 <i>b</i>	Yes Yes Yes 2 b Yes tag Yes Yes Yes	20.1 20.3 4.7/20.3 20.3 20.3 20.3 20.3 20.3	b1         100-620 GeV           b1         275-440 GeV           c1         110-167 GeV           c1         90-191 GeV           c1         90-240 GeV           c1         90-240 GeV           c1         20-600 GeV		$\begin{split} &m(\tilde{\chi}_{1}^{0}){<}90GeV \\ &m(\tilde{\chi}_{1}^{-1}){=}2m(\tilde{\chi}_{1}^{0}) \\ &m(\tilde{\chi}_{1}^{-1}){=}2m(\tilde{\chi}_{1}^{0}), m(\tilde{\chi}_{1}^{0}){=}55GeV \\ &m(\tilde{\chi}_{1}^{0}){=}1GeV \\ &m(\tilde{\chi}_{1}^{0}){=}1GeV \\ &m(\tilde{\chi}_{1}^{0}){>}150GeV \\ &m(\tilde{\chi}_{1}^{0}){<}200GeV \end{split}$	1308.2631 1404.2500 1209.2102, 1407.0583 1506.08616 1407.0608 1403.5222 1403.5222
EW direct	$ \begin{array}{l} \tilde{\ell}_{L,R} \tilde{\ell}_{L,R}, \tilde{\ell} \rightarrow \tilde{\ell} \tilde{\chi}^0_1 \\ \tilde{\chi}^+_1 \tilde{\chi}^1, \tilde{\chi}^+_1 \rightarrow \tilde{\ell} \nu(\ell \tilde{\nu}) \\ \tilde{\chi}^+_1 \tilde{\chi}^1, \tilde{\chi}^+_1 \rightarrow \tilde{\tau} \nu(\tau \tilde{\nu}) \\ \tilde{\chi}^+_1 \tilde{\chi}^0_2 \rightarrow \tilde{\ell}_L \nu^\ell_{L} \ell(\tilde{\nu}\nu), \ell \tilde{\nu} \tilde{\ell}_L \ell(\tilde{\nu}\nu) \\ \tilde{\chi}^+_1 \tilde{\chi}^0_2 \rightarrow W \tilde{\chi}^0_1 \tilde{\chi}^0_1 \\ \tilde{\chi}^+_1 \tilde{\chi}^0_2 \rightarrow W \tilde{\chi}^0_1 \tilde{\chi}^0_1, h \rightarrow b \tilde{b} / W W / \tau \\ \tilde{\chi}^+_2 \tilde{\chi}^3, \tilde{\chi}^0_2 \rightarrow \tilde{\ell}_R \ell \\ \end{array} $ GGM (wino NLSP) weak prod	$\begin{array}{c} 2 \ e, \mu \\ 2 \ e, \mu \\ 2 \ \tau \\ 3 \ e, \mu \\ 2 - 3 \ e, \mu \\ 2 - 3 \ e, \mu \\ \gamma \gamma \gamma  e, \mu, \gamma \\ 4 \ e, \mu \\ .  1 \ e, \mu + \gamma \end{array}$	0 0 0-2 jets 0-2 <i>b</i> 0	Yes Yes Yes Yes Yes Yes Yes	20.3 20.3 20.3 20.3 20.3 20.3 20.3 20.3	$\tilde{\ell}$ 90-325 GeV $\tilde{\chi}_{1}^{\pm}$ 140-465 GeV $\tilde{\chi}_{1}^{\pm}$ 100-350 GeV $\tilde{\chi}_{1}^{\pm}, \tilde{\chi}_{2}^{\pm}$ 700 GeV $\tilde{\chi}_{1}^{\pm}, \tilde{\chi}_{2}^{\pm}$ 420 GeV $\tilde{\chi}_{1}^{\pm}, \tilde{\chi}_{2}^{\pm}$ 250 GeV $\tilde{\chi}_{2,3}^{\pm}$ 620 GeV $\tilde{W}$ 124-361 GeV		$\begin{split} & m(\tilde{\chi}_{1}^{0}) {=} 0 \; GeV \\ & m(\tilde{\chi}_{1}^{0}) {=} 0 \; GeV \; m(\tilde{\ell}, \tilde{\nu}) {=} 0.5(m(\tilde{\chi}_{1}^{\pm}) {+} m(\tilde{\chi}_{1}^{0})) \\ & m(\tilde{\chi}_{1}^{0}) {=} 0 \; GeV, \; m(\tilde{\ell}, \tilde{\nu}) {=} 0.5(m(\tilde{\chi}_{1}^{\pm}) {+} m(\tilde{\chi}_{1}^{0})) \\ & m(\tilde{\chi}_{1}^{\pm}) {=} m(\tilde{\chi}_{2}^{0}), \; m(\tilde{\chi}_{1}^{0}) {=} 0, \; m(\tilde{\ell}, \tilde{\nu}) {=} 0, \; sleptons \; decoupled \\ & m(\tilde{\chi}_{1}^{\pm}) {=} m(\tilde{\chi}_{2}^{0}), \; m(\tilde{\chi}_{1}^{0}) {=} 0, \; sleptons \; decoupled \\ & m(\tilde{\chi}_{2}^{0}) {=} 0, \; m(\tilde{\ell}, \tilde{\nu}) {=} 0.5(m(\tilde{\chi}_{2}^{0}) {+} m(\tilde{\chi}_{1}^{0})) \\ & cr < 1 \; mm \end{split}$	1403.5294 1403.5294 1407.0350 1402.7029 1403.5294, 1402.7029 1501.07110 1405.5086 1507.05493
Long-lived particles	Direct $\tilde{\chi}_1^{\dagger} \tilde{\chi}_1^{-}$ prod., long-lived $\tilde{\lambda}$ Direct $\tilde{\chi}_1^{\dagger} \tilde{\chi}_1^{-}$ prod., long-lived $\tilde{\lambda}$ Stable, stopped $\tilde{g}$ R-hadron GMSB, stable $\tilde{r}$ , $\tilde{\chi}_1^{0} \rightarrow \tilde{\tau}(\tilde{e}, \tilde{\mu}) + \tau$ GMSB, $\tilde{\chi}_1^{0} \rightarrow \gamma \tilde{G}$ , long-lived $\tilde{\chi}_1^{0}$ $\tilde{g}_{\tilde{g}}, \tilde{\chi}_1^{0} \rightarrow eev/e\mu\nu/\mu\mu\nu$ GGM $\tilde{g}_{\tilde{g}}, \tilde{\chi}_1^{0} \rightarrow Z\tilde{G}$	$ \begin{array}{c} \overset{z_1^{\pm}}{\underset{1}{\overset{1}{\overset{1}{\overset{1}{\overset{1}{\overset{1}{\overset{1}{$	1 jet - 1-5 jets - - - μ - s -	Yes Yes - Yes - Yes -	20.3 18.4 27.9 19.1 19.1 20.3 20.3 20.3	$\tilde{k}_{1}^{\pm}$ 270 GeV $\tilde{k}_{1}^{\pm}$ 482 GeV $\tilde{g}$ 832 GeV $\tilde{g}$ 537 GeV $\tilde{k}_{1}^{0}$ 537 GeV $\tilde{k}_{1}^{0}$ 1.0 $\tilde{k}_{1}^{0}$ 1.0	1.27 TeV TeV TeV	$\begin{split} &m(\tilde{\chi}_{1}^{+})\!\cdot\!m(\tilde{\chi}_{1}^{0})\!\sim\!160\;MeV,\tau(\tilde{\chi}_{1}^{+})\!=\!0.2\;ns\\ &m(\tilde{\chi}_{1}^{+})\!\cdot\!m(\tilde{\chi}_{1}^{0})\!=\!100\;GeV,10\;\mus\!<\!\tau(\tilde{\chi}_{1}^{+})\!<\!15\;ns\\ &m(\tilde{\chi}_{1}^{0})\!=\!100\;GeV,10\;\mus\!<\!\tau(\tilde{g})\!<\!1000\;s\\ &10\!<\!tan\beta\!<\!50\\ &2\!<\!\tau(\tilde{\chi}_{1}^{0})\!<\!3\;ns,\;SPS8\;model\\ &7\!<\!c\tau(\tilde{\chi}_{1}^{0})\!<\!740\;nm,m(\tilde{g})\!=\!1.3\;TeV\\ &6\!<\!c\tau(\tilde{\chi}_{1}^{0})\!<\!480\;nm,m(\tilde{g})\!=\!1.1\;TeV \end{split}$	1310.3675 1506.05332 1310.6584 1411.6795 1411.6795 1409.5542 1504.05162 1504.05162
RPV	$ \begin{array}{c} LFV pp \rightarrow \tilde{v}_{\tau} + X, \tilde{v}_{\tau} \rightarrow e\mu/e\tau/\mu\tau \\ Bilinear \; RPV \; CMSSM \\ \tilde{X}_1^+ \tilde{X}_1^-, \tilde{X}_1^+ \rightarrow W \tilde{X}_1^0, \tilde{X}_1^0 \rightarrow ee\tilde{v}_{\mu}, e\mu \\ \tilde{X}_1^+ \tilde{X}_1^-, \tilde{X}_1^+ \rightarrow W \tilde{X}_1^0, \tilde{X}_1^0 \rightarrow \tau \\ \tilde{g} \tilde{g}, \tilde{g} \rightarrow q q \\ \tilde{g} \tilde{g}, \tilde{g} \rightarrow \tilde{q} \tilde{X}_1^0, \tilde{X}_1^0 \rightarrow q q \\ \tilde{g} \tilde{g}, \tilde{g} \rightarrow \tilde{t}_1 t, \tilde{t}_1 \rightarrow bs \\ \tilde{t}_1 \tilde{t}_1, \tilde{t}_1 \rightarrow bs \\ \tilde{t}_1 \tilde{t}_1, \tilde{t}_1 \rightarrow b\ell \end{array} $	$\begin{array}{cccc} r & e\mu, e\tau, \mu\tau \\ 2 & e, \mu  (\mathrm{SS}) \\ 5_e & 4 & e, \mu \\ 5_\tau & 3 & e, \mu + \tau \\ 0 & 0 \\ 2 & e, \mu  (\mathrm{SS}) \\ 0 \\ 2 & e, \mu \end{array}$	- 0-3 <i>b</i> - 6-7 jets 6-7 jets 0-3 <i>b</i> 2 jets + 2 2 <i>b</i>	- Yes Yes - - Yes b -	20.3 20.3 20.3 20.3 20.3 20.3 20.3 20.3	\$\vec{v}_r\$         \$\vec{v}_1\$         \$\vec{750}\$ GeV         \$\vec{v}_1\$         \$\vec{750}\$ GeV         \$\vec{v}_1\$         \$\vec{v}_1\$ <td>1.35 TeV 1.35 TeV V TeV</td> <td><b>7. TeV</b> <math>\lambda'_{311}=0.11, \lambda_{132/133/233}=0.07</math> <math>m(\tilde{q})=m(\tilde{g}), c\tau_{LSP}&lt;1 \text{ mm}</math> <math>m(\tilde{\chi}_1^0)&gt;0.2\times m(\tilde{\chi}_1^{\pm}), \lambda_{121}\neq 0</math> <math>m(\tilde{\chi}_1^0)&gt;0.2\times m(\tilde{\chi}_1^{\pm}), \lambda_{133}\neq 0</math> BR(t)=BR(b)=BR(c)=0% <math>m(\tilde{\chi}_1^0)=600 \text{ GeV}</math> <math>BR(\tilde{t}_1 \rightarrow be/\mu)&gt;20\%</math></td> <td>1503.04430 1404.2500 1405.5086 1405.5086 1502.05686 1502.05686 1404.250 ATLAS-CONF-2015-026 ATLAS-CONF-2015-015</td>	1.35 TeV 1.35 TeV V TeV	<b>7. TeV</b> $\lambda'_{311}=0.11, \lambda_{132/133/233}=0.07$ $m(\tilde{q})=m(\tilde{g}), c\tau_{LSP}<1 \text{ mm}$ $m(\tilde{\chi}_1^0)>0.2\times m(\tilde{\chi}_1^{\pm}), \lambda_{121}\neq 0$ $m(\tilde{\chi}_1^0)>0.2\times m(\tilde{\chi}_1^{\pm}), \lambda_{133}\neq 0$ BR(t)=BR(b)=BR(c)=0% $m(\tilde{\chi}_1^0)=600 \text{ GeV}$ $BR(\tilde{t}_1 \rightarrow be/\mu)>20\%$	1503.04430 1404.2500 1405.5086 1405.5086 1502.05686 1502.05686 1404.250 ATLAS-CONF-2015-026 ATLAS-CONF-2015-015
Other	Scalar charm, $\tilde{c} \rightarrow c \tilde{\chi}_1^0$	0	2 c	Yes	20.3	č 490 GeV		$m(\tilde{\chi}^0_1)$ <200 GeV	1501.01325
					10	-1	1	Mass scale [TeV]	

#### Mass scale [TeV]

\*Only a selection of the available mass limits on new states or phenomena is shown. All limits quoted are observed minus 1 $\sigma$  theoretical signal cross section uncertainty.

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ATLAS Preliminary

 $\sqrt{s} = 7.8 \text{ TeV}$ 



From LHCC Open meeting, 03.12.201 3 2015

## Excesses seen so far (Exotics, SUSY)

For a more detailed review, see the physics plenary talk (July 18th) by T. Golling

#### Mono-jet <u>CDS link</u>

- $~\sim \! 1.7\sigma$  /  $2.4\sigma$  excess above BG in the two highest MET SRs
- High MET SRs dropped for paper (very low stat in CR)
- **VV** $\rightarrow$ **JJ** <u>CDS link</u>
  - Mass of fat jets each consistent with W,Z mass
  - No issue found in cross checks
- Same-sign leptons / 3-leptons + b-jets <u>CDS link</u>
  - No issue found in cross checks, being made public
- 3-leptons + 3 b-jets <u>CDS link</u>
  - ~3.5 $\sigma$  in a validation region of the 3-lepton search
  - No issue found, but will not be made public as not a SR
  - Plan a dedicated SR for run2
- Z+jets + MET <u>CDS link</u>
  - Peaking at Z mass, BG dominated by non-Z (tt)
  - No issue found in cross checks, already public
- SUSY: <u>https://twiki.cern.ch/twiki/bin/viewauth/AtlasProtected/SusyDiscrepancies</u> Exotics: <u>https://twiki.cern.ch/twiki/bin/view/AtlasProtected/ExoticsExcessSummary</u>



ATLAS inter

## **Discovery and exclusion**

 P-value=probability that result is as/less compatible with the hypothesis

#### DISCOVERY:

- The <u>null hypothesis</u> H<sub>0</sub> describes <u>background only</u>
  - If the *p*-value of H<sub>0</sub> is found below a given threshold, one can consider looking for a better model
  - In HEP,  $Z \ge 5$  is conventionally required to claim a discovery
- The alternative hypothesis H<sub>1</sub> describes signal + background
  - The alternative hypothesis is supposed to fit the data very well for claiming a discovery

#### EXCLUSION:

- The <u>null hypothesis</u> H<sub>0</sub> describes <u>signal + background</u>
  - One is interested into setting an upper limit to the intensity of the signal alone
- The alternative hypothesis H<sub>1</sub> describes background only
  - No real need to test for it
  - The background-only model becomes important only in case of discovery

### Interpretation strategy



Likelihood function: L(μ,θ) μ: signal strength (POI); θ: nuisance parameters(NP) Profile Likelihood: constrain uncertainty (NP) as part of a likelihood fit



Construct test statistics  $t_{\mu}$  based on likelihood ratio  $\lambda$ :



Find the observed test statistic for tested µ: t<sub>µ,obs</sub>



If CLs<0.05: the value



iSTEP2015

### Simultaneous fit

- Background estimates in SRs are obtained by a simultaneous fit in each channel based on the profile likelihood method. Three dedicated fit for different purpose...
  - Background-only fit
    - Fit for all CRs, excluding SRs.
      - Get background-only estimates.
      - Also extrapolate to VRs (non used in fit, only for cross-check) and SRs.
  - Discovery fit
    - Fit for all CRs and SRs.
    - Signal contamination is turned off in CRs and set as a dummy number 1 in SR (so, the fitted non-SM signal strength = the excess in Nevents of SR)
      - Get model-independent upper limit on signal in SR.
  - Exclusion fit
    - Fit for all CRs and SRs.
    - Signal is turned on in all regions, according to model-dependent prediction.
      - Got signal model-dependent exclusion from all CRs+SRs →final exclusion contours for SUSY model
- The basic strategy is to share background information in all regions (CR,SR,VR). The background parameters are predominantly constrained by CRs with large statistics, which in turn reduces the impact of uncerts in SR.

### Data-driven background estimation

**One approach** to data-driven bg **estimation** is to use uncorrelated model-independent variables to *extrapolate* the background from a background-dominated control region to the signal region.

Normalization Region Other variable Signal Region (blinded) A(bg) D(S+bg) B(bg) C(bg) **ETmiss Control Region** 

### Key points:

# • <u>The two variables</u> should have good discrepancy and uncorrelated

Nbg in signal region  $D = (A/B)^*$ 

**Normalize Factor** 

- <u>Control Sample selection:</u> enough statistics;lower susy contamination; unbiased estimation of SM background
- <u>Normalization region selection</u>: enough statistics;lower susy contamination; flat ratio(A/B) distribution with ETmiss

"ABCD"

Method

**Control Sample** 

### **Background Estimation Strategy**

#### ATLAS-CONF-2013-062



Other small BGs (diboson, single top etc) are directly estimated from MC.

0(



Inner Detector: Highly segmented silicon strips, determine very accurately charged particles trajectories

Solenoid Magnet: Solenoid coil that generates a 2T magnetic field in the region of the Inner Detector

Electromagnetic Calorimeter: Electron and photon energies are measured through electromagnetic showers

- Hadronic Calorimeter: Hadrons interact with dense material and produce a shower of charged particles
- Toroid Magnets: 8 toroidal coils that create a 0,4T magnetic field in the area of the Muon Spectrometer
- Muon Spectrometer: Muons traverse the rest of the detector and are measured in its outer layers

### The Higgs mechanism, an analogy...

D. Miller (UC London)



The Higgs field fills all space



A 'particle' that moves in the Higgs field ...



... moves slower the more it attract attention (interacts with the Higgs field, generating its mass, the larger, the stronger its interactions...)

### The Higgs particle, an analogy...



Somebody whispers a rumour into the room...



... and the field starts to get excited and interact with itself giving birth to a massive particle