# SHiP : a new facility with a dedicated detector for studying tau neutrino properties

M. Komatsu (Nagoya Univ. JAPAN) on behalf of the SHiP collaboration



The 14<sup>th</sup> International Workshop on Tau Lepton Physics (Tau2016) IHEP, Beijing, 18-23 September 2016

### The SHiP experiment





CERN-SPSC-2015-016 SPSC-P-350 8 April 2015

#### Search for Hidden Particles

Streamed user-couthwest, and ancountered a hearier sea than they had not with before in the whole voyage. Saw gandeler and a prean ruch war the vessel. The crew of the Paka caw a case and a log, they also gicked up a stick which appeared to have been carried with an iron tool, a give of cana, a glast which prove on low, and a board. The crew of the Nina caw other signs of law, and a stick lowed with roce berries. These signs encountered ther, and they all preve cheerful Saled this day till curvet, twenty cover leques.

After canset steeved their oripinal course cast and called truelve miles on hour till two hours after mithight, poinp minety miles, which are truenty-two leagues and a helf and as the Parta can the suffrest caller, and kept ahead of the Adviral.

she discovered land

 Proposal for a new beam dump facility at the CERN SPS

- Hidden sector detector
- Tau neutrino detector
- 235 authors from 45 institutes and 15 countries + CERN
  - Technical Proposal is submitted in April 2015 (arxiv.org/abs/1504.04956)
  - Physics Proposal signed by 85 theorists (arxiv.org/abs/1504.0855)
- SPSC positive recommendation in January 2016

**Technical Proposal** 

### Primary motivation



- Look for Beyond Standard Model in :
  - High Energy Frontier : LHC
  - Intensity Frontier : SHiP
- Look for long lived neutral (hidden) particle from beam dump
   Known Physics
- $v_{\tau}$  beam source
  - $D_s \rightarrow \tau v_{\tau}$



### SHiP (Search for Hidden Particles) Physics objectives



- Explore hidden portals of the SM using > 2 × 10<sup>20</sup> p.o.t.
   (>10<sup>17</sup> D, >10<sup>15</sup> τ)
  - Heavy neutral lepton in various states
  - Dark photon
  - ✓ SUSY neutralino ....
  - See more detail on http://ship.web.cern.ch/
- ✓ Neutrino interactions (expect ~3500  $v_{\tau}$  interactions identified in 9.6 tons emulsion target)
  - $\checkmark$   $v_{\tau}$  and anti-  $v_{\tau}$  physics, cross-section
  - Physics in  $v_{\tau}$  scattering, structure function, magnetic moment.
    - Charm physics in neutrino and anti-neutrino interactions

### Physics programs



#### Full physics program

#### 2 Vector portal

- 2.1 Classification of vector portals
- 2.1.1 Kinetic mixing
- 2.1.2 Anomaly-free gauge groups  $(B L_{\gamma} L_{\mu} L_{\tau} \text{ etc})$
- 2.1.3 Other froms of vector portals.
- 2.1.4 Chern-Simons portal
- 2.2 Matter states charged under new U(1)
  - 2.2.1 Higgs mechanism in the dark sector
  - 2.2.2 Supersymmetric U(1)<sup>r</sup> models
- 2.3 Physics motivation for light mass (less than weak scale) vector parti
  - 2.3.1 Putative solution to the muon g 2 discrepancy
  - 2.3.2 Mediator of interaction with DM and possible connection positron excess
  - 2.3.3 Self-Intereaction of dark matter via light mediators
- 2.4 Main features of vector portal phenomenology.
  - 2.4.1 Deeay rates, modes, branchings, er for dark photon
  - 2.4.2 Other vector candidates
- 2.4.3 Higgsstrahlung process for U(1)<sup>i</sup> and delayed decays of h<sup>i</sup>.
- 2.5 Summary of the existing constraints on light vector and light DM st
  - 2.5.1 Current status of experimental constraints on exotic vector s
  - 2.5.2 Production and detection of light vector portal DM
  - 2.5.3 Cosmological and astrophysical constraints on vector portals
- 2.6 Case studies for SHIP
  - 2.6.1 Production and detection of kinetically mixed dark photor vectors.
- 2.8.2 Production and detection of other unstable particles (h<sup>2</sup> H)

#### 3 Scalar portal

- 3.1 The scalar sector of the Standard Model and Beyon
  - 3.1.1 Scalar portal effective Lagrangian
  - 3.1.2 Hidden Valleys
  - 3.1.3 Light scalars in supersymmetry
- 3.2 Linear scalar portals: Higgs-scalar mixing
- 3.2.1 Existing experimental limits
- 3.2.2 What SHiP can do
- 3.3 Z<sub>2</sub> scalar portals: pair-production of light hidd
  - 3.3.1 Probing Exotic Higgs Decays at SHiP
  - 3.3.2 What SHiP can do
- 3.4 Pseudoscalar portals
- 3.5 Scalar portals and Dark Matter
  - 3.5.1 Scalar as a mediator between DM and the SM
  - 3.5.2 Scalar as a DM candidate
- 3.6 Dark pions
  - 3.6.1 The model and scales
  - 3.6.2 Dark pion lifetime and decay modes
  - 3.6.3 What SHiP could do?
- 3.7 Scalar portals and inflation
  - 3.7.1 Light inflatons

#### 4 Neutrino portal

#### 4.1 Heavy neutral leptons

- 4.2 Active neutrino phenomenology
  - 4.2.1 Three-flavour neutrino oscillations. A theoretical overview
  - 4.2.2 Prosent experimental status of neutrino masses and mixings
  - 4.2.3 Short-Baseline neutrino anomalies
  - 4.2.4 Future neutrino experiments
- 4.3 HNLs and neutrino masses
  - 4.3.1 Seasaw formula 4.3.2 Seasaw scales
  - 4.3.3 Beyond the minimal sensaw model
  - 4.3.4 Possible origins of the keV-MeV-GeV scale of HNL masses
- 4.4 Direct HNL searches
  - 4.4.1 Direct signatures of HNL: kink searches; peak searches
  - 4.4.2 Direct signatures of HNL: fixed target experiments
  - 4.4.3 Direct Signatures of HNL: Collider Searches
- 4.5 Indirect HNL probes
  - 4.5.1 Neutrinoloss double beta docay for non-somew HNL
  - 4.5.2 Neutrinoloss double beta decay for two seasaw HNLs
  - 4.5.3 Charged lepton flavour violating processes
  - 4.5.4 HNL and primordial nucleosynthesis
- 4.6 HNL and baryon asymmetry of the Universe
  - 4.6.1 Sakharov conditions. Leptogenesis
  - 4.6.2 Thermal leptogenesis
  - 4.6.3 Resonant Leptogenesis
  - 4.6.4 Leptogenesis via HNL oscillations
- 4.7 HNL and dark matter
  - 4.7.1 Bounds on HNL as dark matter. No assumptions on production mechanise
  - 4.7.2 Bounds on DM HNL if produced via mixing with active neutrinos only
  - 4.7.3 3.5 keV line
- 4.8 ⊮MSM

#### 400 Fundamental and the instantion of the AMOM

- 8 Tau neutrino physics and other precision measurements in SHiP 8.1 Tau neutrino physics
  - 8.1.1 Flux of tau neutrinos
  - 8.1.1 Flux of the neutrinos
  - 8.1.2 Expected sensitivity
  - 8.2 Deep inelastic muon and electron neutrino scattering
  - 8.2.1 Status of perturbative QCD calculations
  - 8.2.2 Strangeness from heavy-quark DIS in CC interactions
  - 8.2.3 Nuclear effects in  $\nu N$  DIS and global analyses of nuclear PDFs
  - 8.2.4  $\alpha_S$  measurement via Gross-Llewellin Smith sum rule
  - 8.2.5 Precise Ratios for Neutrino Nucleon Interactions
- 8.3 Limit on Tau neutrino magnetic moment
- 8.4 Charmod pentaquark searches
- 8.5 Summary

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#### 9 Searches of lepton flavour violating processes $\tau \rightarrow 3\mu$

- 9.1 Motivation as a null-test of the standard model
- 9.2  $\tau \rightarrow 3\mu$  in seesaw scenarios
- 9.3 Supersymmetric models
- 9.4 Relation to two-body LVF decays of Z boson, neutral pseudoscalar and
- 9.5 Current and future experimental sensitivities
- 9.6 Proposal for a fixed-target facility

#### 5 ALPs and other PNGBs at SHiP 5.1 ALPs and why they are interesting

5.1.2 Connection to Dark Matter

5.3 ALPs coupled to two gauge bosons

5.3.1 Prospects for SHiP

5.4 ALPs coupled to SM fermions

5.4.2 Prospects at SHiP

6.2.2 R-parity Violation

6.2.5 Concluding remarks

6.3.3 Sgoldstinos at SHiP

6.3.4 Concluding remarks

6.4.4 Detection at SHiP

6.5.1 Motivation

Setup

6.7.1 Motivation

6.8 Additional Possibilities

6.6.2 SHiP Sensitivity

6.5.3 Existing bounds

6.5.4 What SHiP can do

6.5.5 Concluding remarks

6.6.3 Concluding remarks

6.7.3 Concluding remarks

6.7 Axinos and saxions, ALPinos and sALPs

6.5.2 Evolutes

6.4.5 Concluding remarks

Light Dirac gauginos

6.3.1 Origin of light sgoldstinos

5.5 Concluding remark

6.1 Introduction

6.2.3

6.2.4

6.3.2

6.4.1

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6.6.1

6 SUSY

5.2 Interactions, phenomenological features and existing limits

5.4.1 Interactions, phenomenological features and existing limits

6.2 A Very Light Supersymmetric Neutralino and R-Parity Violation

Finding Neutralinos at SHiP via R-Parity violation

6.2.1 Motivation for a very light neutralino

6.3 Light particles from the SUSY breaking sector

Comparison with Previous Bounds

Origins of Pseudo-Dirac fermions

Origin of the effective model

SUSY vector portal I: Hidden Photinos

Sgoldstino couplings and phenomenology

Effective model, phenomenological features

SUSY vector portal II: Novel Hidden Photon decays

6.7.2 Phenomenology of saxions and axinos and possibilities at SHiP

5.1.1 ALP origins

Search for Hidden Particles



#### Tau neutrino (Main in this talk)

- We started look for tau neutrino since 1994 in CERN WA95 CHORUS (SBL  $v_{\mu} \rightarrow v_{\tau}$  oscillation).
- Fermilab E872 DONUT (1997)
  - First observation : Phys. Lett. B504 (2001) 218-224
  - Cross-section : Phys. Rev. D78 (2008) 052002
    - 9(7.5) tau neutrino candidate events
    - $\sigma^{\text{const}}(v_{\tau}) = (0.39 \pm 0.13 \pm 0.13) \times 10^{-38} \text{ cm}^2 \text{ GeV}^{-1}$
- CERN CNGS1 OPERA (2008-2012) LBL  $v_{\mu} \rightarrow v_{\tau}$ 
  - <u>5 tau neutrino</u>, <u>5.1</u>σ : Phys. Rev. Lett. <u>115</u> (2015) <u>121802</u>
- Only 14 tau neutrino events ever observed.

## Experiment site

#### North Area



#### North area

0

SHiP

Search for Hidden Particles



## Beam parameters for SHiP

#### Proton beam

- Momentum : 400GeV/c
- Beam intensity : 4-5 x 10<sup>13</sup> /cycle
- Cycle length : 7.2 s
- Spill duration : 1 s (slow spill)
- Average power : 400kW (during spill ~3MW)
- Expected spot size (H/V) : 6mm/6mm
- > 4x10<sup>19</sup> pot / year  $\rightarrow$  2x10<sup>20</sup> pot for 5 years
  - Very same with CNGS performance
    - Plan was 2.25×10<sup>20</sup> but 1.8×10<sup>20</sup> was delivered.

## Beam dump target







- In case of no sweep, the target would not melt but will fail by pressure.
- 1.2 DPA (displacement per atom) with 2×10<sup>20</sup> pot

Segmented Mo and W target actively cooled with water.

Beam on target
 Sweep is necessary like LHC





## Comparison with DONUT



 Charm production by 400GeV, detector acceptance at 60m and tau neutrino cross-section

- ✓ DONUT/SHiP →  $1/(0.36 \times 0.2 \times 0.52) \sim 27$
- Proton on target for SHiP and DONUT
  - ✓ SHiP/DONUT →  $2 \times 10^{20}$  /  $3.6 \times 10^{17}$  ~ 560

✓ Target mass

✓ SHiP/DONUT → 9600kg/260kg ~ 37

✓ Overall advantage against DONUT  $\rightarrow$  560\*37/27 ~ 770

✓ Assuming OPERA like brick (8.3kg) → 1155 bricks

✓ 1155/150 000 = 0.8% of OPERA experiment

### SHiP neutrino detector





### The neutrino target





- Emulsion Cloud Chamber (ECC) technology used in OPERA and DONUT
- Lead plates (high density material for the interaction) interleaved with emulsion films (tracking devices with µm resolution)



## Detector design with CES

![](_page_13_Figure_1.jpeg)

![](_page_13_Picture_2.jpeg)

Target region: 11 mini-walls One wall contains 15x7 bricks Mass ~ 8.3kg x 15x7 x 11 ~ 9.6 ton

Charge determination not only for muonic channel.

MC simulation of CES provide 53% charge determination for hadrons

Statistical gain due to 
$$\frac{\sum_{i=1}^{N} br_i \varepsilon_i}{br_\mu \varepsilon_\mu} \simeq \frac{18 \cdot 0.95 + 50 \cdot 0.53 + 15 \cdot 0.53^2}{18 \cdot 0.9} \simeq 3$$

![](_page_14_Picture_0.jpeg)

## **Compact Emulsion Spectrometer**

![](_page_14_Picture_2.jpeg)

![](_page_14_Figure_3.jpeg)

Three emulsion films interleaved with 1.5cm air gap in magnetic field (~1T), 3cm thick compact spectrometer. H. Shibuya et al, NIM A592 (2008) 56

![](_page_14_Figure_5.jpeg)

# SHIP Search for Hidden Particles

# Separation $v_{\tau}$ /anti- $v_{\tau}$

- three emulsion films interleaved with two, 15-mm thick, Rohacell layers
- 90% efficiency for hadronic τ daughters reaching the end of ECC brick in a 1T field
- sagitta method used to discriminate between positive and negative charge

#### Performance

- electric charge can be determined with better than 3 σ level up to 10 GeV/c
- Momentum estimated from the sagitta Δp/p < 20% up to 12 GeV/c</li>

![](_page_15_Figure_8.jpeg)

![](_page_15_Figure_9.jpeg)

![](_page_16_Figure_0.jpeg)

![](_page_17_Picture_0.jpeg)

# **Objective lens**

Resolution : ~420nm N.A.: 0.65 Light source : G-line (436nm) Magnification : ×12.2 F.O.V : 5.1 (H) × 5.1 (V)mm

#of image plane 6 (by Beam splitter)

Weight : 90kg Total length : 844mm

![](_page_18_Picture_4.jpeg)

# Speed and Coverage of Mosaic Imager

Divide FOV into 72 parts. Need the sensor of 2M pixel and 340fps.

![](_page_19_Figure_2.jpeg)

#### Specially ordered Mosaic Imager

![](_page_19_Picture_4.jpeg)

![](_page_20_Figure_0.jpeg)

# SHIP Search for Hidden Particles

# Neutrino Flavor Identification

- $v_{\mu}$  : muon reconstruction in the magnetic spectrometer
- $v_e$  : EM shower detection in ECC
- $v_{\tau}$  : tau decay topological detection and kinematics
  - Also charge determination by CES

![](_page_21_Figure_6.jpeg)

### Neutrino interaction

- Rich tau neutrino content
  - o.45% relative to muon neutrino
  - 3.0% relative to electron neutrino

![](_page_22_Figure_4.jpeg)

	Ф	<e> (GeV)</e>
$\mathbf{v}_{\mu}$	1.7X10 <sup>6</sup>	29
V <sub>e</sub>	2.5X10 <sup>5</sup>	46
ν <sub>τ</sub>	7.6x10 <sup>3</sup>	59
Anti- $v_{\mu}$	6.7x10 <sup>5</sup>	28
Anti-v <sub>e</sub>	9.0X10 <sup>4</sup>	46
Anti-ν <sub>τ</sub>	3.9x10 <sup>3</sup>	58

Rates for five years of nominal operation with 2 x 10<sup>20</sup> protons on target

![](_page_22_Picture_7.jpeg)

gnal and ba	ackgr	ounc				Search for Hidd
SIGNAL EXPECTATIO	N BA(	TKGROI	R:	=S/B RATI	0	
decay channel	$N^{exp}$	$     \frac{     \nu_{ au}}{N^{bg}} $	R	Nexp	$\overline{ u}_{ au} N^{bg}$	R
$ au  ightarrow \mu$	570	30	19	290	140	2
au  ightarrow h	990	80	12	500	380	1.3
au  ightarrow 3h	210	30	7	110	140	0.8
total	1770	140	13	900	660	14

Main background source: charm production in  $v_{\mu}^{CC}$  (anti- $v_{\mu}^{CC}$ ) and  $v_{e}^{CC}$  (anti- $v_{e}^{CC}$ ) interactions, when the primary lepton is not identified

![](_page_24_Picture_0.jpeg)

### Physics with tau neutrino DIS

Structure function only accessible by tau neutrino

$$\begin{aligned} \frac{d^2 \sigma^{\nu(\overline{\nu})}}{dx dy} &= \frac{G_F^2 M E_{\nu}}{\pi (1 + Q^2 / M_W^2)^2} \left( (y^2 x + \frac{m_\tau^2 y}{2E_{\nu} M}) F_1 + \left[ (1 - \frac{m_\tau^2}{4E_{\nu}^2}) - (1 + \frac{M x}{2E_{\nu}}) \right] F_2 \\ &\pm \left[ xy(1 - \frac{y}{2}) - \frac{m_\tau^2 y}{4E_{\nu} M} \right] F_3 + \frac{m_\tau^2 (m_\tau^2 + Q^2)}{4E_{\nu}^2 M^2 x} F_4 + \frac{m_\tau^2}{E_{\nu} M} F_5 \right), \end{aligned}$$

#### Dependent on the lepton mass

![](_page_24_Figure_5.jpeg)

![](_page_24_Figure_6.jpeg)

## Summary and prospect

![](_page_25_Picture_1.jpeg)

#### Unique opportunity to study tau neutrino physics

- We have only 9 (DONUT)+5 (OPERA) tau neutrino CC interactions.
   Study with 3500 tau neutrino interaction can be done in SHiP.
- Unique chance to study tau and anti tau neutrino cross-section and anti neutrino charm production.
  - ✓ More than one order of magnitude.
- Technical Proposal is submitted (April 2015)
  - $\checkmark$  Physics run from 2026.
  - ✓ Detector design is also under way, CDR in 2018.
- SPSC gave positive statement. Proceed to Comprehensive Design Report (CDR) in 2016-2018.
- "Physics Beyond Colliders" kickoff workshop just took place at CERN on 6-7 September.

#### SHiP is in CERN Courier : March 2016

>>>> Energy Frontier LHC Unknown physics Known physics Neutrino physics **Flavour physics** Intensity Frontier

SHiP is a new experiment at the intensity frontier aimed at exploring the hidden sector.

#### SHiP sets a new course in intensity-frontier exploration

SHiP (Search for Hidden Particles) is a newly proposed experiment for CERN's Super Proton Synchrotron accelerator. Its challenging goals include the direct search for hidden non-Standard Model particles.

#### A Golutvin, Imperial College London/CERN, and R Jacobsson, CERN, on behalf of SHiP.

SHiP is an experiment aimed at exploring the domain of very weakly interacting particles and studying the properties of tau neutrinos. It is designed to be installed downstream of a new beam-dump facility at the Super Proton Synchrotron (SPS). The CERN SPS and PS experiments Committee (SPSC) has recently completed a review of the SHiP Technical and Physics Proposal, and it recommended that the SHiP collaboration proceed towards preparing a Comprehensive Design Report, which will provide input into the next update of the European Strategy for Particle Physics, in 2018/2019.

have now observed all the particles of the Standard Model, however it is clear that it is not the ultimate theory. Some yet unknown particles or interactions are required to explain a number of observed phenomena in particle physics, astrophysics and cosmology, the so-called beyond-the-Standard Model (BSM) problems, such as dark matter, neutrino masses and oscillations, baryon asymmetry, and the expansion of the universe.

CERN Courler March 2016

New physics

While these phenomena are well-established observationally, they give no indication about the energy scale of the new physics. The analysis of new LHC data collected at  $\sqrt{=13}$  TeV will soon have directly probed the TeV scale for new particles with couplings at O(%) level. The experimental effort in flavour physics, and searches for charged lepton flavour violation and electric dipole moments, will continue the quest for specific flavour symmetries to complement direct exploration of the TeV scale.

However, it is possible that we have not observed some of the particles responsible for the BSM problems due to their extremely feeble interactions, rather than due to their heavy masses. Even in the scenarios in which BSM physics is related to high-mass scales, many models contain degrees of freedom with suppressed couplings that stay relevant at much lower energies.

Given the small couplings and mixings, and hence typically Why is the SHiP physics programme so timely and attractive? We long lifetimes, these hidden particles have not been significantly

![](_page_26_Picture_12.jpeg)

# Backup

# Planning schedule of the SHiP

![](_page_28_Picture_1.jpeg)

![](_page_28_Figure_2.jpeg)

Form SHiP collaboration Technical Proposal Comprehensive Design Report Production Readiness Review Construction / production Data taking of 2x10<sup>20</sup>pot

- $\rightarrow$  2014 Done
- $\rightarrow$  April 2015 Done
- $\rightarrow$  2016 2018
- $\rightarrow$  End of 2019
- → 2021-
- → 2026-

#### Estimated cost

#### Detector breakdown

Item	$\mathbf{Cost}$	(MCHF)
Tau neutrino detector		11.6
Active neutrino target	6.8	
Fibre tracker	2.5	
Muon magnetic spectrometer	2.3	
Hidden Sector detector		46.8
HS vacuum vessel	11.7	
Surround background tagger	2.1	
Upstream veto tagger	0.1	
Straw veto tagger	0.8	
Spectrometer straw tracker	6.4	
Spectrometer magnet	5.3	
Spectrometer timing detector	0.5	
Electromagnetic calorimeter	10.2	
Hadronic calorimeter	4.8	
Muon detector	2.5	
Muon iron filter	2.3	
Computing and online system		0.2
Total detectors		58.7

#### Overall cost of SHiP facility

Item	$\mathbf{Cost}$	(MCHF)
Facility		135.8
Civil engineering	57.4	
Infrastructure and services	22.0	
Extraction and beamline	21.0	
Target and target complex	24.0	
Muon shield	11.4	
Detector		58.7
Tau neutrino detector	11.6	
Hidden Sector detector	46.8	
Computing and online system	0.2	
Grand total		194.5

#### Impact on $v_{\tau}$ YIELD

![](_page_30_Figure_1.jpeg)

### Charm production vs energy

![](_page_31_Figure_1.jpeg)

arXiv: 1504.04855 SHiP Physics Proposal

## TAU NEUTRINO MAGNETIC MOMENT

**IN SHiP** 

A massive neutrino may interact e.m.

 $\rightarrow$  magnetic moment proportional to its mass  $_{\nu}$ 

$$\mu_{\nu} = \frac{3 e G_F m_{\nu}}{8 \pi^2 \sqrt{2}} \simeq (3.2 \times 10^{-19}) \left(\frac{m_{\nu}}{1 \text{ eV}}\right) \mu_B$$
  
Current  $\left[ \begin{pmatrix} \nu_e \end{pmatrix} & \mu_{\nu} < 2.9 \cdot 10^{-11} \mu_B \\ (\nu_{\mu}) & \mu_{\nu} < 6.9 \cdot 10^{-10} \mu_B \end{pmatrix} \right]$ 

$$\frac{\sigma_{(\nu e, \overline{\nu} e)}}{dT}\Big|_{\mu_{\nu}} = \frac{\pi \alpha_{em}^2 \mu_{\nu}^2}{m_e^2} \left(\frac{1}{T} - \frac{1}{E_{\nu}}\right)$$

 $W^+$ 

No interference as it involves a spin flip of the neutrino

$$n_{evt} = \frac{\mu_{\nu}^2}{\mu_B^2} \int \Phi_{\nu_{\tau}} \sigma^{\mu} N_{nucl} dE = 4.3 \times 10^{15} \frac{\mu_{\nu}^2}{\mu_B^2}$$

Assuming 5% systematics from DIS measurements

SHiP can explore a region down to

$$\mu_{\nu} = 1.5 \times 10^{-7} \mu_B$$

$$\begin{array}{cccc}
BACKGROUND PROCESSES \\
\nu_x + e^- & \rightarrow & \nu_x + e^- & \text{NC} \\
\nu_e(\bar{\nu}_e) + e^- & \rightarrow & e^- + \nu_e(\bar{\nu}_e) & \text{CC} \\
\nu_e + n & \rightarrow & e^- + p & \text{QE} \\
\bar{\nu}_e + p & \rightarrow & e^+ + n & \text{QE} \\
\nu_e(\bar{\nu}_e) + N & \rightarrow & e^-(e^+) + X & \text{DIS} & 730 \end{array}$$

 $\theta_{\nu-e}^2 < 2m_e/E_e$ 

 $\int_{E_{e}}^{\theta_{\nu-e}} < 30 \, mrad$   $E_{e} > 1 \, \text{GeV}$ 

SIGNAL SELECTION

## CHARM PHYSICS @SHIP

![](_page_33_Figure_1.jpeg)

Expected charm exceeds the statistics available in previous experiments by more than one order of magnitude

	Expected events
$\nu_{\mu}$	$\frac{6.8 \cdot 10^4}{6.8 \cdot 10^4}$
$\nu_e$	$1.5 \cdot 10^{4}$
$ar{ u_{\mu}}$	$2.7 \cdot 10^4$
$ar{ u_e}$	$5.4 \cdot 10^{3}$
total	$1.1 \cdot 10^5$

![](_page_33_Figure_4.jpeg)

In NuTeV ~5100  $v_{\mu}$ ~ 1460 anti- $v_{\mu}$ In CHORUS ~2000  $v_{\mu}$ 32 anti- $v_{\mu}$ 

No charm candidate from  $v_e$  and  $v_{\tau}$  interactions ever reported!